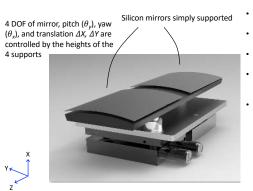
NEXT GENERATION X-RAY OPTICS (NGXO)

Alignment and Bonding of Silicon Mirrors for High-Resolution Astronomical X-ray Optics [10699-141]

Kai-Wing Chan^{a,c}, James R. Mazzarella^{b,c}, Timo T. Saha^c, William W. Zhang^c, Ryan S. McClelland^c, Michael P. Biskach^{b,c}, Peter M. Solly^{b,c}, Raul E. Riveros^{a,c}, Ai Numata^{b,c} ^aCenter for Research and Exploration in Space Science and Technology & University of Maryland, Baltimore County, Maryland, USA; ^bStinger Ghaffarian Technologies, Inc., Maryland, USA; ^cNASA/Goddard Space Flight Center, Maryland, USA

Mirror Alignment for Large X-ray Telescopes

- Future large x-ray astronomy missions require large area and high resolution
- Present consensus is to integrate thin, lightweight, segmented mirrors to form a compact but large telescope with high resolution (better than 1")
- Kev technologies
 - 1) Accurate mirror substrates: polishing high quality mono-crystalline silicon
 - 2) Stress-free reflective coating: stress-based distortion cancellation
 - 3) Precise alignment and integration: kinematic mounting and distortion-free
 - bonding
- · Four-point alignment for quasi-cylindrically symmetric mirrors (segments) 4-Point alignment: Pitch and yaw angles, image center (X, Y) are controlled by
 - heights of the 4 mount points
 - Alianment Precision: better than 1"
 - **Bonding error:** ~ 0.1µm (~ 1")
 - Current single mirror pair x-ray tested: 3" (Half-Power Diameter)
- Integration into (meta) shell
 - Shell structure has rotationally defined axis
 - Interlocking mirrors \Rightarrow lightweight, mechanically strong telescope



4-Point Alignment of Mirrors

- Pitch (θ_y) and yaw (θ_x): compactness of focus
- ΔX, ΔY: image center onto the optical axis of system Rotation (θ ,) is invariant from
- cylindrical symmetry **∆Z** is nearly invariant from
- small grazing angle (long focal length)
- Co-alignment of "primary" and "secondary" mirrors, and mirrors in the next [(n+1)th] shell, are done through a common optical axis reference

Silicon base

▲ Alignment setup. Sub-aperture images were taken for mirrors aligned in a collimated optical beam. Centroid the sub-aperture images at the focal plane determine the

quality of focusing and de-center errors. They, in turn,

heights needed for a better alignment.

determine the amount of iterative correction of spacer

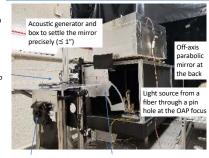
- Individual mirror statically determined by the 4 spacers Sub-aperture masks Silicon mirrors facing down. Optical axis horizontal

Focus and Precision of Alignment and Bonding



Silicon mirrors mounted on top of another

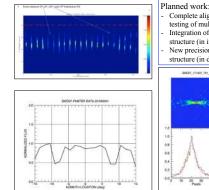
- Mirrors are acoustically settled Sub-aperture measurement of mirror images in a
- collimated beam qualifies the alignment
- Corrective spacer height is achieved by polishing Epoxy applied to round-top spacers bonds mirror in place



Optical axis reference defined as Movable part of the collimated beam masks ough a center aperture

Summary: X-ray Test Result

- Aligned and Bonded mirrors were tested at GSFC and MPE Panter
- Resolution of single pair of mirror is 3 arcsecond
- Flux is uniform (except at the spacers)
- Alignment of mirror in the second shell is achieved in optical beam at 1" (not x-ray tested yet)



structure (in implementation) New precision mirror positioning structure (in development) 9 9 9 1 3

Complete alignment, bonding, and testing of multiple pairs (in progress) Integration of mirrors onto meta-shell



For Further Information

Silicon Mirror Fabrication: Raul Riveros, 10699-23 [Monday, 2:00 PM] Raul.E.Riveros@nasa.gov Mirror Alignment and Bonding: Kai Chan, 10699-141 [this poster] Kai-Wing.Chan-1@nasa.gov Telescope Design and analysis: Peter Solly Peter.M.Solly@nasa.gov Optics design and analysis: 10699-179 [Wednesday, 6 PM Poster] Timo.T.Saha@nasa.gov Mirrors Technology: W. Zhang. 10699-22 [Monday, 1:40 PM] William.W.Zhang@nasa.gov

Acknowledgment

NGXO research is supported by NASA Strategic Astrophysics Technology (SAT) under Research Opportunities in Space and Earth Sciences (ROSES)

