# AN IMBRICATE MID CRUSTAL SUTURE ZONE: THE MOJAVE-YAVAPAI PROVINCE BOUNDARY IN GRAND CANYON, ARIZONA 

Mark Holland

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AN IMBRICATE MID CRUSTAL SUTURE ZONE: THE MOJAVE-YAVAPAI PROVINCE BOUNDARY IN GRAND CANYON, ARIZONA
by

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## THESIS

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# AN IMBRICATE MID CRUSTAL SUTURE ZONE: THE MOJAVE-YAVAPAI PROVINCE BOUNDARY IN GRAND CANYON, ARIZONA 

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#### Abstract

The Mojave and Yavapai provinces in southwestern Laurentia contain evolved and juvenile crust respectively, but the nature of their boundary remains uncertain. This paper analyzes the $\mathrm{U}-\mathrm{Pb}-\mathrm{Hf}$ isotopic composition of zircons from both the oldest plutons and metasedimentary rocks in Grand Canyon. My results show that the Vishnu Schist does not support suturing. Paradoxically, plutons east and west of the Crystal shear zone support suturing based on different $\mathrm{U}-\mathrm{Pb}-\mathrm{Hf}$ isotopic characteristics across the shear zone. I interpret the nature of this boundary to be the existence of an Archean substrate which is sampled by plutons west of the shear zone. The overlapping Vishnu Schist suggests a more complicated architecture. The Vishnu Schist was deposited across Mojave and Yavapai crust, and imbricated during the Yavapai orogeny. The ultimate architecture is a distributed boundary with slivers of plutons carrying the isotopic signature of their respective provinces imbricated within metasediments.


INTRODUCTION ..... 1
GEOLOGIC BACKGROUND ..... 6
METHODS ..... 11
U-PB AND HF ICPMS RESULTS FROM GRAND CANYON ..... 13
Interlaboratory Comparison ..... 13
Elves Chasm Gneiss ..... 16
Tuna Creek Pluton ..... 16
Comparison Summary ..... 17
Results from Supracrustal Rocks ..... 20
Vishnu Schist ..... 20
Rama Schist ..... 24
Results from Granodiorite Plutons ..... 25
East of Crystal Shear Zone ..... 25
Granodiorites West of Crystal Shear Zone ..... 27
DISCUSSION ..... 30
Comparison of Igneous and Metasedimentary Zircons ..... 30
Source of the Vishnu Schist ..... 34
Source of Inherited Zircons ..... 38
Tectonic Model for Crustal Architecture ..... 42
CONCLUSIONS ..... 49
APPENDIX A ..... 50
Table A 1: U-Pb geochronologic analyses of Vishnu Schist (ALC).....Error! Bookmark not defined.
Table A 2: Hf isotopic data for Vishnu Schist (ALC).......Error! Bookmark not defined.
Table A 3: U-Pb geochronologic analyses of Grand Canyon plutons (ALC) Error! Bookmark not defined.

Table A 4: Hf isotopic data for Grand Canyon plutons (ALC).....Error! Bookmark not defined.

Table A 5: U-Pb geochronologic data from Grand Canyon plutons (GEMOC) .Error! Bookmark not defined.

Table A 6: Hf isotopic data for Grand Canyon plutons (GEMOC)....Error! Bookmark not defined.
APPENDIX B ..... 78
Geochemical Evolution and Metallogeny of Continents (GEMOC) Key Centre Methods. ..... 78
U-Pb Geochronology ..... 78
Hf Isotopes ..... 79
Arizona Laserchron Center (ALC) Methods ..... 80
U-Pb Geochronology ..... 80
Hf Isotopes ..... 82
APPENDIX C ..... 85
Pluton Descriptions by Sample ..... 85
Grapevine Camp Pluton (K12-81L): ..... 85
Zoroaster Pluton (K12-85.3L): ..... 87
Horn Creek Pluton (K12-90.5R): ..... 88
Trinity Pluton (K12-91.5R): ..... 90
Boucher Pluton (K12-96.2L): ..... 91
Tuna Creek Pluton (13H-99R; K05-100.5): ..... 94
Elves Chasm Gneiss (K12-115L; K06-113): ..... 95
Ruby Pluton (K06-107): ..... 96
Diamond Creek Pluton (K06-228.3): ..... 97
238-Mile Pluton (K06-238-2): ..... 97
245-Mile Pluton (K06-245-2): ..... 98
REFERENCES ..... 100

## LIST OF FIGURES

Figure 1. Regional map of Proterozoic provinces of western Laurentia ..... 2
Figure 2. Geologic maps ..... 7
Figure 3. Geologic cross sections ..... 8
Figure 4. Comparison of results from the Elves Chasm Gneiss as analyzed at the
Geochemical Evolution and Metallogeny of Continents (GEMOC) Key Centre, andthe Arizona Laserchron Center (ALC)18
Figure 5. Comparison of ALC and GEMOC results for the Tuna Creek pluton ..... 19
Figure 6. Comparison and synthesis of new and previous Vishnu Schist detrital zircon U-
Pb age data. ..... 21
Figure 7. Paired U-Pb-Hf isotopic data for all supracrustal lithologies in Grand Canyon...23
Figure 8. Paired U-Pb-Hf results for zircons separated from plutons east of the Crystal
$\qquad$shear zone26
Figure 9. Paired U-Pb-Hf results for zircons separated from plutons west of the Crystal shear zone ..... 28
Figure 10. Compiled Vishnu Schist U-Pb-Hf data ..... 31
Figure 11. Synthesis of all U-Pb-Hf data from plutonic and detrital samples in Grand
Canyon ..... 33
Figure 12. U-Pb-Hf isotopic data from zircons separated from 1.78-1.74 Ga Mojave
province compared to the Vishnu Schist. ..... 37
Figure 13. Field photo of the Zoroaster pluton. ..... 41
Figure 14. Schematic cross section corresponding to Figure 2A ..... 44

Figure 15. Plate tectonic cartoons illustrating possible scenarios for the depositional setting of the Vishnu Schist and lithospheric formation in the Grand Canyon region.

## LIST OF TABLES

Table 1: Geochronologic summary of Paleoproterozoic Grand Canyon rocks ..... 14
Table A 1: U-Pb geochronologic analyses of Vishnu Schist (ALC) ..... 50
Table A 2: Hf isotopic data for Vishnu Schist (ALC) ..... 60
Table A 3: U-Pb geochronologic data of Grand Canyon plutons (ALC) ..... 64
Table A 4: Hf isotopic data for Grand Canyon plutons (ALC) ..... 68
Table A 5: U-Pb geochronologic analyses from Grand Canyon Plutons (GEMOC) ..... 72
Table A 6: Hf isotopic data for Grand Canyon plutons (GEMOC) ..... 75

## INTRODUCTION

The core of the North American continent was assembled by collision of Archean cratons during the 1.83-1.80 Ga Trans-Hudson system (Hoffman, 1988; Corrigan et al., 2005; 2009) followed by progressive southward (present coordinates) addition of lithosphere in accretionary orogens of southwestern Laurentia (Karlstrom and Bowring, 1988; Windley, 2003) beginning $\sim 1.8 \mathrm{Ga}$ and culminating with the Grenville orogeny and assembly of Rodinia (e.g. Whitmeyer and Karlstrom, 2007). In this model, southern Laurentia is an important field laboratory for studies of continent formation processes because it has been interpreted as an amalgamation of dominantly juvenile Paleoproterozoic terranes that were added to the Archean and 1.8 Ga nucleus (Figure 1 inset) (DePoalo, 1981; Hoffman, 1988; Karlstrom and Bowring, 1988; Bowring and Karlstrom, 1990; Duebendorfer, 2007). Further, this region may represent one of the largest additions of juvenile continental crust in Earth history (Reymer and Schubert, 1986). Alternatively, new models suggest that many orogens previously thought to be predominantly accretionary can contain hybridized and metasomatized lithosphere where cryptic substrates of older lower crust and mantle lithosphere underlie younger crust (Begg et al., 2007; 2009; Griffin et al., 2008; 2011; Belousova et al., 2009; 2010).

The Mojave Province of southwestern Laurentia (Figure 1) has long been known to include older crustal material (Bennett and DePaolo, 1987; Wooden and Miller, 1990; Chamberlain and Bowring, 1990; Wooden and DeWitt, 1991; Ramo and Calzia, 1998; Ilg et al., 1996; Hawkins et al., 1996; Iriondo et al., 2004; Barth et al., 2000, 2009; Wooden et al., 2012), but the age, origin, and distribution of this material remains uncertain. A parallel debate has involved the boundary between the Mojave province and the Yavapai


Figure 1. Regional map of Proterozoic provinces of western Laurentia modified after Karlstrom et al., 2004. Outcrops of Proterozoic rocks are shown in gray. Box shows location of Figure 2. Csz (Crystal shear zone), GCsz (Gneiss Canyon shear zone).
province, which contains 1.8-1.7 Ga juvenile crust (Wooden and DeWitt, 1991; Karlstrom and Bowring, 1993; Bennett and DePaolo, 1987; Duebendorfer et al., 2006). Previous models of the nature of the Mojave province's inherited signature and the Mojave-Yavapai boundary include: 1) subducted Archean detritus as the source of the evolved isotopic signature of the Mojave, with inherited signature decreasing with distance from the Archean Wyoming province (Bennett and DePaolo, 1987; Ramo and Calzia, 1998), 2) a $\sim 75 \mathrm{~km}$ wide boundary zone defined by whole-rock and feldspar Pb isotopic data that was interpreted to be the result of post 1.73 Ga modification of the Mojave-Yavapai boundary by deformation and plutonism after the provinces were tectonically juxtaposed (Wooden and DeWitt, 1991), 3) a distributed tectonic suture centered at the Crystal shear zone in Grand Canyon (Ilg et al., 1996; Hawkins et al., 1996) that extends $\sim 130 \mathrm{~km}$ to the Gneiss Canyon shear zone in western Grand Canyon (Karlstrom et al., 2003), 4) Duebendorfer et al. (2006) interpreted similar Pb isotopic data to represent a wide isotopically mixed zone resulting from rifting and hybridization of older crust, and 5) the presence of Archean crust in middle to lower crustal subcrops of the Mojave province that contributed detritus to the Vishnu Schist (Shufeldt et al., 2010).

A recent detrital zircon study was conducted on the Vishnu Schist in the Grand Canyon, which spans over 30 km on both sides of the proposed Crystal suture zone (Shufeldt et al., 2010). Laser ablation-multicollector-inductively coupled plasma-mass spectrometry (LA-MC-ICP-MS) analysis of >1000 grains separated from 12 spatially distributed samples along a $180-\mathrm{km}$-long cross-strike transect revealed a uniform bimodal detrital zircon age population with peaks at 1.8 Ga and 2.5 Ga . This surprising result lead Shufeldt et al. (2010) to conclude that: 1) any collision of Yavapai crust with Mojave
either pre-dated or was synchronous with Vishnu deposition at 1.75 Ga because the turbidite succession overlaps the proposed crustal boundary at Crystal shear zone, 2) these sediments were not derived from juvenile terranes as only $13 \%$ were 1.75 Ga grains, 3 ) the $\sim 1.85 \mathrm{Ga}$ peak in the metasediments was derived from the underlying 1.84 Ga Elves Chasm gneiss, and 4) that an older Mojave crustal substrate contributed detritus to the Vishnu Schist.

The purpose of this paper is to present new Hf and $\mathrm{U}-\mathrm{Pb}$ data from the oldest rocks in the Grand Canyon and use a synthesis of all available age and isotopic data to test models to explain the nature of the Mojave-Yavapai boundary. The goal is to resolve the nature of the Crystal shear zone by comparing the $\mathrm{U}-\mathrm{Pb}$ and Hf isotopic composition of zircons from the Vishnu metaturbidite with those from plutons that intrude on both sides of the Crystal shear zone. This methodology will test the hypotheses of Hawkins et al. (1996) and Shufeldt et al. (2010) that the crystal shear zone represents a MojaveYavapai crustal boundary, and that there exists an older crustal substrate in the Mojave province. The excellent exposure of a $180-\mathrm{km}$-long cross-strike basement transect, the uniform detrital zircon population of the Vishnu Schist, and the well described structural, geochronologic, and thermobarometric history of the Granite Gorges make the Grand Canyon an ideal field locality for exploring and resolving the timing and character of the Mojave-Yavapai crustal boundary.

New data includes paired $\mathrm{U}-\mathrm{Pb}$ and Hf isotopic analyses of 187 detrital zircons from the Vishnu Schist, and 233 igneous zircons from 10 plutons that intrude the Vishnu Schist, from east to west across the entire transect. Our approach of applying paired $\mathrm{U}-\mathrm{Pb}$ dating and Hf isotopic analysis to both the oldest metasedimentary rocks and the oldest
plutons in the orogen provides a powerful dataset on earliest evolution of the Mojave and Yavapai provinces. These data provide both "top-down" and "bottom-up" views of the crust in the region. $\mathrm{U}-\mathrm{Pb}$ and Hf composition of zircons collected from metasedimentary rocks provide information about the age and chemical maturity of crust exposed in the provenance region during the time of deposition. Conversely, the $\mathrm{U}-\mathrm{Pb}$ and Hf composition of zircons separated from intrusive lithologies provide information about the character of their lower crustal melt source regions. The concept of using plutons as probes of the lower crustal melt reservoirs has been applied to many orogens in the North American cordillera, (Farmer and DePaolo, 1983; Samson et al., 1989; 1991; Friedman et al., 1995) the Variscan orogen (Finger et al., 1997) Alpine orogeny (Kohút and Nabelek, 2008), and the Central Asian Orogenic Belt (Jahn et al., 2000; Kovalenko et al., 2004). Hf isotopes of plutonic zircons have been used to evaluate the participation of juvenile versus evolved crust in Proterozoic terranes (Andersen et al., 2002), and to discern terrane boundaries and infer crustal architecture (Cecil et al., 2011).

## GEOLOGIC BACKGROUND

Basement rocks are exposed from River Mile (RM measured downstream from Lees Ferry) 78 to 260 within both the Upper Granite Gorge (RM 78-116) and Lower Granite Gorge (RM 208-260) segments. The Upper Granite Gorge in particular is characterized by a block type architecture, consisting of six $\sim 10-\mathrm{km}$-scale tectonic blocks bounded by five discrete high strain zones, the deformational and metamorphic character of which have been studied extensively (Ilg et al., 1996 Hawkins et al., 1996; Karlstrom et al., 2003; Dumond et al., 2007). These six blocks, from east to west are the Mineral Canyon, Clear Creek, Trinity Creek, Topaz Canyon, Tuna Creek, and Walthenberg Canyon blocks. Two blocks are sampled in the Lower Granite Gorge: the Travertine Grotto and Spencer Canyon blocks, separated by the Gneiss Canyon shear zone (Karlstrom et al., 2003). These eight blocks all contain Vishnu Schist, and they share a similar $1.75-1.69 \mathrm{Ga}$ tectonic history as described below.

Upper Granite Gorge contains about half supracrustal and half plutonic rocks (Figure 2). The supracrustal Granite Gorge Metamorphic Suite is composed of three intimately interlayered but mappable units: the $1740 \pm 2$ Ma Rama Schist, the $1750 \pm 1$ Ma Brahma Schist, and Vishnu Schist. These felsic and mafic metavolcanic, and metaturbidite units respectively were interpreted to represent juvenile marine volcanicarc rocks (Hawkins et al., 1996; Ilg et al., 1996). These are in a transposed depositional contact with underlying basement of the $1840 \pm 1 \mathrm{Ma}$ Elves Chasm Gneiss, an orthogneiss that ranges from hornblende biotite tonalite to granodiorite (Hawkins et al., 1996). This is the oldest presently known rock in southwestern North America and is $\sim 90$ Ma older than all other lithologies in the Grand Canyon region.

GRANITE GORGES

| EXPLANATION |
| :--- |
| $\square$ |
| 1.4 Ga granite |
| $1.70-1.66 ~ G a ~ g r a n i t e ~ p e g m a t i t e ~$ |
| 1.74-1.71 Ga arc plutons |
| Ultramafics |
| 1.75-1.74 Ga Vishnu Schist |
| 1.75-1.74 Ga Rama and Brahma Schists |
| 1.84 Ga Elves Chasm Gneiss |
| Igneous Hf sample |
| Detrital zircon + Hf sample |
| Detrital zircon sample |
| Fault, ball on downthrown side |
| Axial Trace 参 Shear zone |



Figure 2. A: Digital elevation model of the Grand Canyon region. Paleoproterozoic outcrops are shown in white. Csz (Crystal shear zone), GCsz (Gneiss Canyon shear zone). Line of section A to A' is keyed to Figure 13. B: Geologic map of the Upper Granite Gorge after Ilg et al. (1996). Line of section B to B' is keyed to Figure 3A. Inset shows the Colorado River and outcrops of Proterozoic rocks in the Arizona Transition Zone are shown in gray. Sample locations are projected orthogonally from their location along the river corridor to the line of section. C: Geologic map of the Lower Granite Gorge after Karlstrom et al. (2003). Line of section C to C' is keyed to Figure 3B. Sample locations are projected orthogonally from their location along the river corridor to the line of section.


Figure 3. Geologic cross sections of the Granite Gorges in Grand Canyon. Sample locations are projected orthogonally from the river corridor to the line of section. The block-type architecture and block-bounding high strain zones are shown along with river mile downstream of Lee's Ferry.A: Geologic cross section of the Upper Granite Gorges modified after Dumond et al., 2007. B: Geologic cross section of the Lower Granite Gorge modified after Karlstrom et al., 2003 respectively.

The plutonic rocks of the Granite Gorges are calc-alkaline plutons of 1741-1713 Ma (Hawkins et al., 1996), interpreted as subduction-related arc plutons, are primarily granodiorite and commonly have gabbro-diorite enclaves that record varying degrees of magma mixing (Ilg et al., 1996; Karlstrom et al., 2003). Some contacts with the country rock are intrusive with cross cutting fabric relations, but many are tectonic contacts (Table 1). The largest plutons are ten-km-scale large $\mathrm{F}_{2}$-folded sheet-like bodies, while others are km-wide subvertical tabular bodies that suggest tectonic slices (Figure 3). Ultramafic cumulates also occur as lenses within turbidites and are interpreted as dismembered roots of arc magma chambers (Seaman et al., 1997; Low et al., in progress). In addition, several 1.70-1.66 Ga granite-pegmatite dike swarms are interpreted to be syncollisional granites derived from partial melting of a tectonically thickened crust (Ilg et al., 1996; Hawkins et al., 1996).

The tectonic history of the Upper Granite Gorge involves multi-stage deformation and metamorphism culminating in the 1.72-1.68 Ga Yavapai orogeny (Ilg et al., 1996; Dumond et al., 2007). Structural, geochronologic, and metamorphic studies suggest that the two main stages of deformation overlap locally, and peak metamorphism occurred synchronously with the transition between $D_{1}$ and $D_{2}$ (Ilg et al., 1996). $D_{1}$ involved early thrusting and isoclinal folding, preserved in domains of NW-striking $S_{1}$ fabric developed between 1730-1698 Ma (Ilg et al., 1996). $\mathrm{D}_{2}$ involved km -scale upright, isoclinal to open $\mathrm{F}_{2}$ folds and development of a penetrative NE striking subvertical $\mathrm{S}_{2}$ foliation from 17131685 Ma (Ilg et al., 1996; Dumond et al., 2007). This subvertical foliation is the dominant fabric throughout the transect. The block-bounding high strain zones (Figure 2) are in $\mathrm{D}_{2}$ orientations but probably accommodated multiple slip events, including earlier $\mathrm{D}_{1}$
movements and much later brittle displacement of Grand Canyon Supergroup (Huntoon, 1980; Elston, 1989). Metamorphic studies indicate variable peak temperatures of 520-770 ${ }^{\circ} \mathrm{C}$ with the high strain zones serving as $100-200^{\circ} \mathrm{C}$ thermal discontinuities between tectonic blocks at near isobaric $\sim 0.7$ GPa pressure ( $\sim 25 \mathrm{~km}$ depth; Ilg et al., 1996;

Dumond et al., 2007). Dumond et al. (2007) proposed that changes in peak metamorphic temperature across shear zones were the result of magma-enhanced metamorphic field gradients and that the entire transect decompressed from $\sim 0.7$ to $\sim 0.3-0.4 \mathrm{GPa}(\sim 12 \mathrm{~km}$ depths) mainly during the end of $\mathrm{D}_{2}$ by 1680 Ma , and certainly by 1.4 Ga .

Shear zones are primarily $\mathrm{D}_{2}$ shortening high strain zones, however some have been proposed to have earlier histories, like the Crystal shear zone. It is a $\sim 1 \mathrm{~km}$-wide, NE-striking zone of strong foliation with stretching lineations plunging steeply to the west (Ilg et al., 1996). Based on macroscopic fold asymmetries, and a $\sim 0.1$ GPa higher pressure to the west (Ilg et al., 1996; Dumond et al., 2007), the shear zone is interpreted as west side up. The presence of possible mélange-like rocks, a step to more radiogenic common Pb isotopic composition west of the shear zone, and the presence of xenocrystic >2.0 Ga zircons to the west of the shear zone suggest the Crystal shear zone represents an early $\left(D_{1}\right)$ structure that was reactivated and transposed during $D_{2}$ deformation (Ilg et al., 1996; Hawkins et al., 1996). However, as mentioned above, the uniform detrital zircon population of the Vishnu Schist across this boundary suggests that any juxtaposition of crustal blocks across this boundary must have occurred before or during Vishnu Schist deposition at 1750 Ma (Shufeldt et al., 2010).

## METHODS

Samples for this study were collected over the course of several field seasons from 2005-2014, and U-Pb and Hf isotopic analyses were conducted at two different laboratories. Plutonic samples from western Grand Canyon were taken by Dr. Graham Begg and Dr. Karl Karlstrom during two field seasons in 2005 - 2006 and analyzed at the Geochemical Evolution and Metallogeny of Continents (GEMOC) Key Centre at Macquarrie University in Sydney, Australia. Samples of Vishnu Schist from across the entire Grand Canyon transect were taken on field seasons in 2005-2006 and 2008 and detrital zircon U-Pb analyses were conducted at the Arizona Laserchon Center (ALC) at the University of Arizona in Tucson by Owen Schufeldt. Subsequent Hf isotopic analysis of the Vishnu Schist was carried out on these samples by Dr. George Gehrels at the ALC in 2010 and I carried out new U-Pb and Hf isotopic analyses from 2012-2014. I collected plutonic samples from the eastern Grand Canyon and performed $\mathrm{U}-\mathrm{Pb}$ and Hf isotopic analyses on them at the ALC during 2012-2014. Analytical methods at both laboratories are detailed in Appendix A (see also: Belousova et al., 2001; Griffin et al., 2000, 2002, 2004; Jackson et al., 2004; Gehrels et al., 2006, 2008; Cecil et al., 2011; Gehrels and Pecha, 2014).

As detailed in the data repository, both laboratories employ broadly similar sample preparation and analytical methods and report similar precision and accuracy in both $\mathrm{U}-\mathrm{Pb}$ and Hf isotopic analyses. Both laboratories perform standard heavy mineral separation techniques and mount unknowns together with zircon standards. Prior to analysis, backscattered electron (BSE) or cathodoluminescent (CL) images of all unknown grains were obtained to guide spot selection and identify potential xenocrystic
cores. All $\varepsilon H f$ values were calculated using the bulk silicate earth composition of Bouvier et al. (2008), and we compare the results to the depleted mantle composition of Vervoort and Blichert-Toft (1999).

## U-Pb AND Hf ICPMS RESULTS FROM GRAND CANYON

Table 1 provides a synthesis of new and published U-Pb geochronologic and Hf isotopic analyses from the basement rocks of the Grand Canyon. The results of all analyses are presented in Appendix A. Table A1 contains all new Vishnu Schist and Rama Schist U-Pb geochronologic analyses. Table A2 contains all Vishnu Schist and Rama Schist Hf isotopic analyses. Table A3 contains all plutonic U-Pb geochronologic analyses conducted at the ALC. Table A4 contains all plutonic Hf isotopic analyses conducted at the ALC. Table A5 contains all plutonic U-Pb analyses conducted at the GEMOC Key Centre. Table A6 contains all plutonic Hf isotopic analyses conducted at the GEMOC Key Centre. Supporting data are presented in more detail in Appendices B and C; analytical methods are detailed in Appendix B, and descriptions of individual plutonic sample zircon ages, morphology, and internal textures are provided in Appendix C.

## Interlaboratory Comparison

Samples of the Tuna Creek pluton, and the Elves Chasm gneiss were independently analyzed at both laboratories. Before discussion and interpretation of the $\mathrm{U}-\mathrm{Pb}-\mathrm{Hf}$ isotopic data discussed herein, we compare the results from these very different samples. The Tuna Creek pluton and the Elves Chasm gneiss provide an excellent means of inter-laboratory comparison; the Tuna Creek pluton contains a complex zircon population with varied $\mathrm{U}-\mathrm{Pb}$ ages and Hf isotopic compositions, and has never been precisely dated. In contrast, the age of the Elves Chasm gneiss is precisely known (Hawkins et al., 1996), and results from both laboratories show that it is juvenile at 1840 Ma. Therefore, as discussed in more detail below, similar U-Pb-Hf characteristics

Table 1. Geochronologic summary of Paleoproterozoic Grand Canyon rocks

|  | Rock Name | River <br> Mile | Weighted <br> Mean UPb Age (Ma) | MSWD | Deformational Context | Contact <br> Relations | Metamorphic Block | Lab |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K12-79.6R | Rama Schist | 79.6 | $\begin{gathered} \hline \hline 1751 \pm \\ 16 \\ 1741 \pm 1^{1} \end{gathered}$ | 0.2 | Pre $D_{1}$; from hinge of Sockdolager antiform | Interlayered with Vishnu Schist | Mineral Canyon | ALC |
| K12-81L | Grapevine Camp Pluton | 81 | $\begin{gathered} 1756 \pm \\ 16 \\ 1737 \pm 1^{1} \end{gathered}$ | 0.3 | Contains $\mathrm{S}_{1}$; alligned along $\mathrm{S}_{2}$ Vishnu shear zone. Pre $\mathrm{D}_{1}$ | Cross cuts compositional layering in Vishnu Schist on eastern contact; western contact is the Vishnu shear zone. | Mineral Canyon | ALC |
| K12-85.3L | Zoroaster Pluton | 85.3 | $\begin{gathered} 1755 \pm \\ 14 \\ 1740 \pm 2^{1} \end{gathered}$ | 0.6 | Contains $\mathrm{S}_{1}$, folded by $\mathrm{F}_{2}$. <br> Pre- or syn- $\mathrm{D}_{2}$ | Contains screens of Grand Canyon Metamorphic Suite | Clear Creek | ALC |
| K12-90.5R | Horn Creek Pluton | 90.5 | $\begin{gathered} 1719 \pm \\ 14 \\ 1713 \pm 1^{1} \end{gathered}$ | 0.3 | Contains magmatic $\mathrm{S}_{1}$, solid state $S_{2}$. Syn-D ${ }_{1}$ | Intrudes parallel to compositional layering in Vishnu Schist | Trinity Creek | ALC |
| $\begin{gathered} \text { K12- } \\ 901.5 R \end{gathered}$ | Trinity Gneiss | 91.5 | $\begin{gathered} 1755 \pm \\ 16 \\ 1730 \pm 3^{1} \end{gathered}$ | 0.4 | Contains $\mathrm{S}_{1}$ and $\mathrm{S}_{2}$. Pre- $\mathrm{D}_{1}$ | Contacts between Brahma amphibolites are folded and sheared; intrusive contacts locally preserved | Trinity Creek | ALC |
| K12-96.2L | Boucher <br> Pluton | 96.2 | $\begin{gathered} 1730 \pm \\ 15 \\ 1714 \pm 1^{2} \end{gathered}$ | 0.5 | Weakly foliated. Pre or syn-D ${ }_{1}$ | Eastern margin is 96-mile shear zone; western margin is intrusive into metasupracrustals | Topaz <br> Canyon | ALC |
| 13H-99R | Tuna Pluton | 99 | $\begin{gathered} 1751 \pm \\ 15 \\ <1750^{1} \end{gathered}$ | 1.1 | Contains $\mathrm{S}_{1}$ and $\mathrm{S}_{2}$. Pre- to syn- $D_{1}$ | Folded sheet-like body intrudes Vishnu Schist | Tuna Creek | ALC |
| $\begin{gathered} \text { K05- } \\ 100.5-105 \end{gathered}$ | Tuna Pluton | 100.5 | $\begin{aligned} & 1737 \pm 7 \\ & <1750^{1} \end{aligned}$ | 0.7 | Contains $\mathrm{S}_{1}$ and $\mathrm{S}_{2}$. Pre- to syn- $D_{1}$ | Folded sheet-like body intrudes Vishnu Schist | Tuna Creek | GEMOC |
| K05-107 | Ruby Pluton | 107 | $\begin{gathered} 1726 \pm \\ 15 \\ 1716.6 \pm \\ 0.5^{1} \end{gathered}$ | 0.1 | Contains $\mathrm{S}_{1}$ as magmatic layering. Syn$\mathrm{D}_{1}$ | Eastern margin is tectonized, but locally preserves intrusive relations to supracrustal rocks, western margin is cut by the Bass shear zone. | Tuna Creek | GEMOC |


| K06-113 | Elves Chasm Gneiss | 113 | $\begin{aligned} & 1842 \pm 5 \\ & 1840 \pm 1^{1} \end{aligned}$ | 0.9 | Contains $\mathrm{S}_{1}$ and $\mathrm{S}_{2}$. Pre- $\mathrm{D}_{1}$ basement to Vishnu Schist | Transposed depositional contact between Brahma Schist and Elves Chasm Gneiss | Walthenberg Canyon | GEMOC |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K12-115L | Elves Chasm Gneiss | 115 | $\begin{gathered} 1850 \pm \\ 18 \\ 1840 \pm 1^{1} \end{gathered}$ | 1.4 | Contains $\mathrm{S}_{1}$ and $\mathrm{S}_{2}$. Pre- $\mathrm{D}_{1}$ basement to Vishnu Schist | Transposed depositional contact between Brahma Schist and Elves Chasm <br> Gneiss | Walthenberg Canyon | ALC |
| K06-228.3 | Diamond Creek pluton | 228.3 | $\begin{gathered} 1738 \pm \\ 14 \\ 1736 \pm 1^{3} \end{gathered}$ | 0.2 | Contains magmatic $\mathrm{S}_{1}$ and solidstate $S_{2}$ shear zones. Syn-D | Intrudes Vishnu Schist on western margin, eastern margin covered by Phanerozoic rock | Travertine Block | GEMOC |
| K06-238-1 | Granitic Gneiss from Gneiss Canyon shear zone | 238 | $\begin{gathered} 1731 \pm \\ 14 \end{gathered}$ | 0.1 | Cuts $\mathrm{S}_{1}$, contains weak to strong and magmatic $S_{2}$. Syn- $D_{2}$ | Sheet like intrusions into mixed para and orthogneisses | Gneiss Canyon shear zone | GEMOC |
| K06-245-2 | 245-mile pluton | 245 | $\begin{gathered} 1741 \pm \\ 13 \\ 1720 \pm 5^{3} \end{gathered}$ | 0.6 | Contains $\mathrm{S}_{1}$ and $\mathrm{S}_{2}$. Pre- $\mathrm{D}_{2}$ | Intrudes into Vishnu Schist on western margin, eastern margin is intruded by the younger Seperation pluton | Spencer Canyon Block | GEMOC |

reported from each laboratory show consistent results and the data can be presented and interpreted together.

## Elves Chasm Gneiss

Sample K05-113 was analyzed by Dr. Graham Begg at the GEMOC Key Centre, and I analyzed sample K12-115L at the ALC. The results of each sample analysis are compared in Figure 4.20 grains separated from sample K05-113 (GEMOC lab) yields a weighted mean age of $1841 \pm 5 \mathrm{Ma}(\mathrm{MSWD}=0.7)$. Likewise, 12 ages from sample K12115 L (ALC lab) yielded a weighted mean age of $1855 \pm 18 \mathrm{Ma}(\mathrm{MSWD}=0.2)$. The results from both laboratories are in good agreement with the $1840 \pm 1$ Ma TIMS age of the Elves Chasm gneiss (Hawkins et al., 1996).

Both samples yield juvenile $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values for the Elves Chasm gneiss. The $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ of the depleted mantle at 1840 Ma is $+9.8 . \varepsilon \mathrm{Eff}_{(\mathrm{t})}$ values from sample $\mathrm{K} 05-113$ range from +12.5 to +8.7 with an average of +10.5 . Similarly, K12-115L yields $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values from +11.3 to +6.1 with an average of $+8.6 . \mathrm{\varepsilon Hf}_{(\mathrm{t})}$ values from the GEMOC sample (K05-113) are consistently higher than those from the ALC sample (K12-115L), but there is substantial overlap between the two data sets and all zircons yield $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values that are within analytical error of the depleted mantle array.

## Tuna Creek Pluton

Two samples of the Tuna Creek pluton were collected; K05-100.5 was analyzed by Dr. Graham Begg at the GEMOC Key Centre, and I analayzed 13H-99R at the ALC (Figure 5). Both samples show excellent agreement in terms of zircon ages; with a 1740 Ma age peak, and a xenocrystic population of $2480-2485 \mathrm{Ma}$. The presence of inheritance in the Tuna Creek pluton has been documented previously (Hawkins et al., 1996), but the
age constraints were limited to crystallization at 1750-1710 Ma with $>2.0 \mathrm{Ga}$ inheritance. These results corroborate prior findings, and provide more complete geochronologic constraints on the age of inherited grains.

There is somewhat more disparity between $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values in the Tuna Creek pluton than the Elves Chasm gneiss. GEMOC $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values for six $\sim 1.74 \mathrm{Ga}$ zircons plot in a tight cluster at $\sim+2.8$. Two additional grains yield lower values; one zircon yields an $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ value of -11.3 at 1853 Ma . This grain overlaps at $2 \sigma$ with the age of the Elves Chasm gneiss, but has a substantially lower $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ value. The 2.48 Ga xenocrystic population in the GEMOC sample also plots in a tight cluster, with an average $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ value of -1.2 for six grains. Conversely, the $\mathrm{ALC}_{\varepsilon \mathrm{Hf}_{(t)}}$ values for the 1740 Ma population show more of a vertical spread, with $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values of seven grains ranging from +2.8 to +12.7 and averaging +7.1 . Three additional grains that overlap with the age of the Elves Chasm gneiss were identified in sample 13H-99R; two yielded juvenile $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values, and one yielded an intermediate value of -3.7 at 1851 Ma . Finally, the $\sim 2.48$ Ga population of zircons yielded $\varepsilon \mathrm{Hf}_{(t)}$ values ranging from +0.9 to +3.7 and averaging at +2.2 .

## Comparison Summary

Hf isotopic data from each sample show substantial overlap, however average values consistently differ by $\sim 2$ espilon units. Typically these differences are within uncertainty of the $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values. $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values for the Elves Chasm gneiss were consistently higher from the GEMOC data, however the ALC data yielded higher $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values for the Tuna Creek pluton. As described by Gehrels and Pecha (2014), the uncertainties reported by the ALC are "not of ideal precision," due to analytical procedures that emphasis accuracy at the expense of internal precision. The result of these procedures is that the


Figure 4. Comparison of results from the Elves Chasm Gneiss as analyzed at the Geochemical Evolution and Metallogeny of Continents (GEMOC) Key Centre, and the Arizona Laserchron Center (ALC). Symbols with error bars represent the average 2 sigma uncertainties for $\mathrm{U}-\mathrm{Pb}$ and $\mathrm{Lu}-\mathrm{Hf}$ isotopic analyses from each data set.


Figure 5. Comparison of ALC and GEMOC results for the Tuna Creek pluton. Coloring and symbols as in Figure 4. Age probability plots are normalized such that the area underneath each curve is equal. These show excellent agreement between $\mathrm{U}-\mathrm{Pb}$ ages in primary and inherited grains.
$\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values reported by the GEMOC Key Centre are more precise. However, standard analyses from the ALC yielded uncertainties of $\pm 0.000035$ (May 2013) to 0.000040 (May 2014) at $1 \sigma$, which result in standard uncertainties of 2 to 3 epsilon units at the 2sigma level. The uncertainty in standard analyses is consistent with the $\sim 2$ epsilon unit disparity between the ALC and GEMOC Key Centre and suggests that all measurements are accurate to within 2-3 epsilon units. However, it must be emphasized that while sample pairs K05-113, K12-115L, and K05-100.5, 13H-99R were obtained from the same plutons, analyses were not conducted on the same zircons. It is possible that one or both plutons may not be isotopically homogenous, and that the observed variation in $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values reflect real isotopic variation within the plutons. In conclusion, we feel that the results from each laboratory are consistent and the entire data set can be compared within reasonable uncertainty.

## Results from Supracrustal Rocks

## Vishnu Schist

380 new U-Pb ages from Vishnu Schist were obtained for this study. Analyses were conducted on the same grain mounts prepared by Shufeldt et al. (2010) with the intention of obtaining paired $\mathrm{U}-\mathrm{Pb}$ and Hf isotopic data. I selected samples that were distributed across the entire length of the Grand Canyon transect (river mile 84 - 246) to obtain data that are representative of the entire population. The results of each analysis are presented in Appendix D, Table A1.

My new data are compiled with the entire dataset of Shufeldt et al. (2010) and presented on order of east to west (river mile downstream of Lee's Ferry) in Figure 6. New samples show the same bimodal Paleoproterozoic and early Archean populations


Figure 6. Comparison and synthesis of new and previous Vishnu Schist detrital zircon UPb age data. A: Composite age probability plot take from Shufeldt et al., 2010. B: New detrital zircon ages obtained for this study. C: New composite age probability plot with all current Vishnu Schist detrital zircon ages. D: Normalized age probability plots of Vishnu Schist detrital zircon samples arranged by river mile (OS0878-1 = Owen Shufeldt, 2008, river mile 78). Results from Shufeldt et al., 2010 are shown in blue, new results obtained in this study are shown in red. See also Table A1 for detrital zircon age data.
with peaks at 1780 Ma and 2480 Ma as identified by Shufeldt (2010), and confirm the observation that, whereas individual samples vary, there is no systematic difference in the relative proportions of the 2.5 and 1.8 Ga age populations from east to west across the Crystal shear zone. Shufeldt et al. (2010) reported that only 13\% of all grains overlapped at $2 \sigma$ with the 1750-1740 Ma depositional age of the Vishnu Schist (Hawkins et al., 1996); new data show more grains of this age (Figure 6), such that the new composite Vishnu distribution has 18\% 1740-1750 Ma first-cycle grains. 14\% of all ages overlap at $2 \sigma$ with the $1840 \pm 1 \mathrm{Ma}$ Elves Chasm gneiss, and $28 \%$ are >2500 Ma in age.

I obtained Hf isotopic information for 187 zircons from six spatially distributed samples of Vishnu Schist (Figure 7). The results of each individual analysis are presented in Appendix D, Table A2. As for the U-Pb ages, there is variation in which mode dominates a given spectrum, but there is no systematic variation in Hf composition with river mile (Figure 7). For example, sample K0598.6-104 did in fact yield Archean U-Pb ages (Table A1), but each of these grains were either too small to withstand the larger spot size and deeper pit depth of Hf analysis, or the Hf analyses did not pass the data reduction process. The apparent difference in detrital zircon age spectra in this case is probably the result of a reduced sample size for this sample ( $\mathrm{n}=32$ for $\mathrm{U}-\mathrm{Pb}$, and $\mathrm{n}=16$ for Hf$)$. However, I do not feel that the Hf isotopic data for the Archean population is underrepresented in the entire dataset.

Detrital zircons from the Vishnu Schist yield a broad range of $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values from +10.0 to -19.4 . For our purposes, only $\varepsilon \mathrm{Hf}_{(t)}$ values that overlap within analytical uncertainty $(2 \sigma)$ of the depleted mantle array are considered juvenile. Surprisingly, only 30 grains (16\%) of the entire Vishnu Schist dataset yielded juvenile values. However, due

## Supracrustal Rocks



Figure 7. Paired U-Pb-Hf isotopic data for all supracrustal lithologies in Grand Canyon. Each diamond represents a zircon with paired $\mathrm{U}-\mathrm{Pb}$ and $\mathrm{Lu}-\mathrm{Hf}$ isotopic data. $\mathrm{DM}-$ depleted mantle of Vervoort and Blichert-Toft, 1999; CHUR-chondritic uniform reservoir of Bouvier et al., 2008. Sample K12-79.6R is of the Rama Schist (shown in red), all other data are Vishnu Schist. Below are normalized detrital zircon age probability plots of grains for which we have obtained $\mathrm{Lu}-\mathrm{Hf}$ isotopic data.
to the time-integrated nature of Hf isotopic information, these terms refer to the zircons themselves compared to depleted mantle rather than compared to depositional age. Thus, for the purpose of this study, only zircons that yield crystallization ages of 1750-1740 Ma as well as $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values that overlap with the depleted mantle array, and therefore could have been derived from local arc rocks, are considered juvenile. Of the 187 grains for which I have paired $\mathrm{U}-\mathrm{Pb}-\mathrm{Hf}$ data 58 grains (31\%) fall in this age range, and only 14 of those grains (6\%) yield $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values that overlap with the depleted mantle array. Therefore, I infer that only $7 \%$ of grains could have been derived from juvenile (1.751.74 Ga ) plutonic sources (see below). The rest were derived from older crust.

Shufeldt et al. (2010) proposed that the 1.8 Ga age peak in the Vishnu Schist was derived from the Elves Chasm Gneiss, a hypothesis which is tested with these new Hf isotopic data. 23 grains for which I have paired $\mathrm{U}-\mathrm{Pb}-\mathrm{Hf}$ isotopic data overlap at $2 \sigma$ with the $1840 \pm 1 \mathrm{Ma}$ Elves Chasm gneiss. However, only 4 yield $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values that overlap with the depleted mantle array. Therefore the Vishnu Schist was also not significantly sourced by the juvenile Elves Chasm pluton. My data show that very little of the detritus that sourced the Vishnu Schist was derived from local crust. Instead, Vishnu Schist was overwhelmingly derived from older, isotopically evolved crust.

## Rama Schist

I collected a sample of the Rama Schist from the same locality sampled by
Hawkins et al. (1996). A weighted mean of 12 analyses yielded an age of $1751 \pm 16 \mathrm{Ma}$ (MSWD $=0.2$ ), in good agreement with the $1741 \pm 1$ Ma age of Hawkins et al. (1996). Interestingly, the 10 grains for which I obtained Hf isotopic data yielded juvenile $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values that ranged from +5.7 to +11.3 (Figure 7). All but one grain yielded $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$
values that overlap with the depleted mantle array. This is essentially the same range of $\varepsilon \mathrm{Hf}_{(t)}$ values shown by plutons in eastern Grand Canyon (see below), which suggests that the interlayered metavolcanics were locally derived juvenile arc products and that they became interlayered with far-traveled sedimentary rocks derived from older crustal sources. Interlayering is now parallel to strong $\mathrm{S}_{2}$ foliation such that it could have been primary or tectonic, an interpretation that would require additional study.

## Results from Granodiorite Plutons

## East of Crystal Shear Zone

I obtained 134 paired U-Pb-Hf isotopic analyses from zircons separated from five plutons east of the Crystal shear zone from river mile 81-96 (Figure 8). 84 grains (74\%) overlap at $2 \sigma$ with the 1741-1713 Ma timing of arc magmatism described by Hawkins et al. (1996). Although there are individual outliers, and ICPMS precision is considerably less than the previous ID-TIMS dates, weighted mean ages for each pluton are in good agreement with previously published ID-TIMS ages (Table 1).

Hf isotopic data for plutonic zircons east of Crystal shear zone are shown in Figure 8 . We obtained paired $\mathrm{U}-\mathrm{Pb}-\mathrm{Hf}$ isotopic data for 97 plutonic zircons. $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values range from +5.5 to 13.1 . Of the 97 analyses, $86(87 \%)$ of them yield $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values that overlap with the depleted mantle array. They form a very tight cluster that indicates these granodiorites were derived almost entirely from a juvenile source. However, the few grains that yield more evolved $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values suggest minor involvement of slightly older juvenile crust, such as the Elves Chasm Gneiss, or a crustal component as old as 2.0 Ga .

Two of the grains that do not overlap with the depleted mantle array are core analyses from the Boucher pluton: K12-96.2L-22C, which yielded an $\varepsilon \mathrm{Eff}_{(\mathrm{t})}$ value of -5.1


Figure 8. A: Paired U-Pb-Hf results for zircons separated from plutons east of the Crystal shear zone. Shown in green above are all results from igneous zircons east of Crystal shear zone. B: Results for each pluton, note change in x-axis scale. Average 2sigma uncertainty of all epsilon Hf values is shown in the top right.
at 1945 Ma , and $\mathrm{K} 12-96.2 \mathrm{~L}-17 \mathrm{C}$, which yielded an $\varepsilon \mathrm{Eff}_{(\mathrm{t})}$ value of -2.9 at 2483 Ma . Grain $\mathrm{K} 12-96.2 \mathrm{~L}-1 \mathrm{C}$ yielded a juvenile $\varepsilon \mathrm{Ef}_{(\mathrm{t})}$ value of +8.7 at 1839 Ma . The Boucher pluton at river mile 96 (K12-96.2L) contained zircons that displayed inherited cores in CL-texture. 25 core-rim pairs were analyzed, however only 8 core-rim pair analyses passed the data reduction process. Many apparent core-rim pairs ultimately yielded overlapping ages that contributed to the $1730 \pm 15 \mathrm{Ma}$ weighted mean age. Only 1 core-rim pair yielded nonoverlapping ages. Grain K12-96.2L-3 yielded a core age of $2598 \pm 22 \mathrm{Ma}$, and a rim age of $1727 \pm 37 \mathrm{Ma}$ (Figure 8). In addition, grain K12-96.2L-1 yielded core and rim ages of $1839 \pm 19$ and $1761 \pm 101$ Ma respectively. The large uncertainty associated with the rim analysis causes these grains to overlap at $2 \sigma$, however the similarity between the core age and the Elves Chasm gneiss is noteworthy. I infer that the rims are the same age as the 1730 Ma crystallization age of this pluton. Similarly, despite the lack of reliable rim age data, two additional core analyses yielded ages substantially older than the 1730 Ma population. Grain K12-96.2L-22C yielded an age of $1945 \pm 22 \mathrm{Ma}$, and grain K12-96.2L17C yielded an age of $2483 \pm 24 \mathrm{Ma}$. The latter age corresponds quite closely to the 2481 Ma age peak defined in the compiled data of the Vishnu Schist (Shufeldt et al., 2010).

## Granodiorites West of Crystal Shear Zone

Granodioritic plutons west of the Crystal shear zone are markedly different than those to the east. I obtained 134 paired U-Pb-Hf isotopic analyses from zircons separated from six plutons west of Crystal shear zone from river mile 99-245 (Figure 9). Excluding the 36 zircons from the Elves Chasm Gneiss, a majority of 64 grains (65\%) still overlap at $2 \sigma$ with the 1741-1713 Ma timing of arc-magmatism in the Upper Granite Gorge. However, the remainder of grains yield a semi-continuous spectrum of ages that range


Figure 9. Paired U-Pb-Hf results for zircons separated from plutons west of the Crystal shear zone. A: Composite diagram of all paired igneous U-Pb-Hf data west of Crystal shear zone. Elves Chasm Gneiss shown in purple. Average 2-sigma uncertainty of all epsilon Hf values is shown in top right. Gray arrow shows average crustal evolution of Vervoort and Patchett, 1996. Yellow column represents interpreted mixing of crust of varying age. Boxes show area of B and C. B: $\sim 1.7 \mathrm{Ga}$ population of zircons shown according to pluton name and sample number. Elves Chasm Gneiss not shown (see Figure 4). C: Inherited population of zircons by sample. BSE images of inherited grains and Paleoproterozoic grains with low epsilon Hf values are keyed to B and C.
from 1741-1867 Ma, and a substantially older inherited population that ranges from 2275-2936 Ma. Interestingly, 13 grains (10\%) yield ages that overlap at $2 \sigma$ with the 2480 Ma age peak in the compiled data for the Vishnu Schist, and 8 grains overlap at $2 \sigma$ with the age of the Elves Chasm gneiss.

Of the 64 grains that overlap with the $1.74-1.71 \mathrm{Ga}$ magmatism, 25 grains ( $39 \%$ ) yield juvenile $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values; the remaining 39 grains yield a spread of more evolved Hf isotopic compositions, with $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values as low as -2.8 (Figure 9B). Furthermore, the 8 grains that overlapped in age with the Elves Chasm gneiss yielded varied $\varepsilon \mathrm{Hf}_{(t)}$ values ranging from +10.3 to -11.3 , with four of them juvenile at 1.84 Ga .

The Elves Chasm gneiss itself yielded exclusively juvenile $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values. Attempts were made to date apparently distinct core and rim age domains in zircons, but all attempts yielded ages of $\sim 1.84 \mathrm{Ga}$. Along with the Elves Chasm Gneiss, the Ruby pluton, and Diamond Creek pluton yielded only juvenile grains, with the exception of one zircon from the Diamond Creek pluton that yielded a juvenile Elves Chasm-aged grain.

The 2275-2936 Ma population of inherited zircons is derived from the Tuna Creek, 238-Mile, and 245-Mile plutons. These xenocrystic grains are generally more evolved than most of the primary Paleoproterozoic grains (Figure 9).

## DISCUSSION

## Comparison of Igneous and Metasedimentary Zircons

My approach of applying paired $\mathrm{U}-\mathrm{Pb}$ dating and Hf isotopic analysis to both the oldest metasedimentary rocks and the oldest plutons in the orogen provides insight into the earliest evolution of the Mojave and Yavapai provinces. These two datasets provide initially contradictory information: metasedimentary rocks are the same across the proposed Crystal shear zone, whereas plutons are markedly different. The following discussion addresses this apparent conundrum.

Detrital zircons from the Vishnu Schist show a broad range of juvenile and evolved Hf isotopic compositions but are uniform across the entire Grand Canyon transect. The vertical spread of data in epsilon space suggests that the Vishnu Schist was derived from crust that experienced substantial mixing of older crustal material with juvenile magmas, including: $\sim 3.3-3.2$ Ga crust, $\sim 2.8-2.4 \mathrm{Ga}$ crust, and 2.0-1.7 Ga crust all characterized by mixing of juvenile material with older crust (Figure 10). Few zircons of any age are juvenile $\left(\varepsilon \mathrm{Ef}_{(\mathrm{t})}\right.$ values that overlap at 2-sigma with the depleted mantle array) at the time of their crystallization (16\%), and most late Archean grains are dominated by $\varepsilon H f$ values that suggest the involvement of $\sim 3.0$ Ga crust. The Paleoproterozoic population of zircons predominantly suggest the involvement of 2.0-2.2 Ga crust, but the lack of crust of that age in our data set, and globally, might suggest instead that the abundance of Paleoproterozoic grains with $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values between +5.0 and 0.0 are the result of mixing between juvenile $\sim 1.8$ and $\sim 2.5$ Ga crust.

Conversely, interlayered volcanic deposits of the Rama schist in eastern most Grand Canyon are juvenile at 1.74 Ga . The similarity between plutonic and volcanic


Figure 10. Compiled Vishnu Schist U-Pb-Hf data show that the Vishnu Schist was sourced by an Archean craton characterized by extensive crustal reworking at $\sim 1.8 \mathrm{Ga}$, $\sim 2.5 \mathrm{Ga}$, and $\sim 3.2 \mathrm{Ga}$.
$\varepsilon \mathrm{Hf}_{(t)}$ values suggests that the Rama Schist in the eastern part of the transect is locally derived from similarly aged, but slightly older, arc magmatism.

The age and Hf isotopic composition of plutonic zircons, summarized in Figure 11, are in stark contrast to the detrital grains of the Vishnu Schist. Plutons east of Crystal shear zone are uniformly juvenile at $1.74-1.71 \mathrm{Ga}$. Very few detrital zircons of this age and Hf isotopic composition are found in the Vishnu Schist suggesting a dearth of local juvenile Yavapai province derived crust contributing to the Vishnu basin. However, the xenocrystic grains and Paleoproterozoic grains with more evolved Hf isotopic compositions found in plutons west of the Crystal shear zone are more similar to Vishnu detritus, which suggests that these plutons interacted with crust similar to that which sourced the Vishnu basin. Moreover, this may suggest that the Vishnu Schist was sourced from the west (present coordinates).

Two important first-order observations are as follows: (1) granodioritic arc plutons include both juvenile and inherited zircons, (2) there is a dramatic change in igneous zircon ages and Hf composition across the Crystal shear zone: plutons east of Crystal shear zone yield almost exclusively juvenile $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values, and a lack of inherited grains. The only exceptions are the two inherited cores found in the Boucher pluton (sample K12-96.2L) at river mile 96, which is the western-most pluton still east of the Crystal shear zone at river mile 97 (see below). Conversely, west of the Crystal shear zone some plutons yield abundant xenocrystic zircons, inherited cores, and more evolved $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values, while other plutons are predominantly juvenile. Juvenile plutons include the 1.84 Ga Elves Chasm gneiss (K06-113, and K12-115L), and 1.74-1.71 Ga Grapevine Camp (K12-81L), Zoroaster (K12-85.3L), Horn (K12-90.5R), Trinity (K12-91.5R), Ruby


Figure 11. Synthesis of all U-Pb-Hf data from plutonic and detrital samples in Grand Canyon. Vishnu Schist is shown in blue, plutons east of the Crystal shear zone in green, and plutons west of the Crystal shear zone in red. The dramatic difference in plutonic zircons is apparent. Paradoxically, the plutonic data suggest a suture across the shear zone, however the detrital zircon data do not.
(K06-107), and Diamond Creek (K06-228) plutons. Plutons with inherited signatures include the 1.74-1.71 Ga Boucher (K12-96.2L), Tuna Creek (K06-100.5, and 13H-99R), 238-Mile (K05-238), and 245-Mile (K06-245) plutons.

These data describe a uniform metasedimentary basin that was sourced primarily by isotopically evolved 1.8 and 2.5 Ga crust, which was intruded by a suite of plutonic rocks with essentially the same crystallization ages but markedly different zircon populations and Hf isotopic compositions across a structural boundary.

## Source of the Vishnu Schist

An important insight gained from these data is the disparity between the juvenile isotopic signature of the Elves Chasm gneiss and the more evolved character of 1.84 Ga detrital grains. My new data show that Paleoproterozoic detrital zircons yield primarily more evolved $\varepsilon H f$ values than igneous Elves Chasm Gneiss grains, and preclude the Elves Chasm Gneiss from contributing a substantial volume of detrital zircons into the Vishnu basin, as suggested by Shufeldt et al. (2010). The presence of additional Penokean aged crust in central Colorado has been inferred from inherited zircons found in $\sim 1.75 \mathrm{Ga}$ granites and rhyolites (Hill and Bickford, 2001). Subsequent Hf isotopic analysis of many of the same samples discussed by Hill and Bickford (2001) include only one $>1800$ Ma grain (aged $1877 \pm 37$ ), which yielded an $\varepsilon \mathrm{Hf}_{(t)}$ value of 1.1. This value is substantially lower than $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values obtained from the Elves Chasm gneiss, more similar to $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values from the $\sim 1.85$ Ga population of zircons in the Vishnu Schist, and similar to the somewhat evolved Nd isotopic character of $1.88-1.84 \mathrm{Ga}$ Penokean plutonic rocks (Barovich et al., 1989). It is possible that Penokean aged crust further to the northeast sourced the $\sim 1.85$ Ga population of detrital zircons in the Vishnu schist,
however one zircon with similar U-Pb-Hf characteristics is not enough to suggest a central Colorado provenance.

The paucity of juvenile first-cycle $1.75-1.74 \mathrm{Ga}$ detrital zircons also preclude local juvenile Yavapai-type plutons from contributing substantial detritus to the Vishnu Schist. Of course, with the exception of the Elves Chasm gneiss, plutons exposed in the Grand Canyon intrude the Vishnu Schist and could not have contributed detritus to the Vishnu protolith, however my data suggest that very little of the Vishnu Schist was sourced by older juvenile arc terranes such as the 1.78 Ga Green Mountain arc (Jones et al., 2011). In the Gunnison-Salida area, the 1751 Ma Powderhorn Granite yields zircons with slightly evolved $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values ranging from +2 to +6 (Bickford et al., 2008). $35 \%$ of $1.74-1.78 \mathrm{Ga}$ Vishnu grains yielded similar $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values. In addition, the $1772-1754 \mathrm{Ma}$ Twilight Gneiss of the Needle Mountains is similarly juvenile, and is old enough to have contributed to the Vishnu Schist (Gonzales and Van Schmus, 2007). Therefore, a modest amount of more locally derived detritus could have been sourced from Colorado, however given the much more evolved values of some $\sim 1.85 \mathrm{Ga}$ grains, it is unlikely that any 1.85 Ga crust in Colorado provided the dominant source of Paleoproterozoic detritus to the Vishnu Schist, and even more unlikely that far traveled Penokean crust did so.

The inherited Archean age population and Paleoproterozoic grains with lower $\varepsilon \mathrm{Hf}$ values suggest that plutons west of the Crystal shear zone might have interacted with the same type of crust that provided detritus to the Vishnu Schist. Crust that sourced the Vishnu Schist might therefore lie to the west (present coordinates). It is likely that the Paleoproterozoic, and possibly Archean, detrital zircon population is derived from the proximal Mojave province, similar to the model favored by Shufeldt et al. (2010) for the

Archean component of the Vishnu Schist. At present no Elves Chasm gneiss or older aged crust has been identified in the Mojave province, but abundant 1.75-1.78 Ga granitoids with evolved Nd (Bennett and DePaolo, 1987; Ramo and Calzia 1998) and Hf (Wooden et al., 2012) compositions exist (Figure 12). In addition, an Archean component has been recognized in the Nd isotopic composition of the Mojave province in the Death Valley region (Ramo and Calzia, 1998), and detrital zircons of similar ages to the Vishnu Schist have been documented in the Mojave province (Barth et al., 2000; 2009; Strickland et al., 2012).

The source of the $>2.4$ Ga population of detrital zircons in the Vishnu Schist remains uncertain. Shufeldt et al. (2010) suggested that the Gawler craton of southern Australia had the most similar age spectrum to the Archean population in the Vishnu schist, however recent $\mathrm{U}-\mathrm{Pb}$ ages obtained from orthogneisses in the Farmington Canyon Complex are similar to the 2.48 Ga age peak in the Vishnu Schist (Mueller et al., 2011). This may provide a potential Laurentian source for the older Vishnu Schist detritus, however the dearth of voluminous Wyoming craton aged detritus in the Vishnu Schist (Shufeldt et al., 2010) still argues against a substantial Laurentian cratonic source. Mueller et al. (2011) recognized this dilemma and suggested that the Mojave province may have evolved independently at 2.5 Ga , which was similarly proposed by Whitmeyer and Karlstrom (2007), and might seem supported by the presence of $\sim 2.5 \mathrm{Ga}$ lithosphere in the Mojave province (Lee et al., 2001) and inherited zircons of the same age in plutons west of Crystal shear zone (see below).


Figure 12. U-Pb-Hf isotopic data from zircons separated from 1.78-1.74 Ga Mojave province plutonic rocks shown in yellow (Wooden et al., 2012) compared to the Vishnu Schist.

## Source of Inherited Zircons

Understanding the history of the inherited grains in the plutons is of key importance for evaluating crustal evolution models. The refractory nature of zircon gives it the tendency to retain valuable isotopic data through a wide variety of geologic processes and often results in the preservation of distinct zones within a single zircon. These internal heterogeneities can provide information regarding the petrogenesis of the host rock or even protolith. Alternatively, if these zones are not identified or misinterpreted they can obscure the true history of the zircon or provide misleading geochronologic and isotopic information.

The presence or absence of inherited zircons in an igneous rock depends on the temperature and composition of the melt; if the melt is sufficiently saturated in Zr then preexisting zircons may be preserved in the pluton as xenocrystic grains or cores rimed by younger zircon growth. However, if the melt is Zr undersaturated then any preexisting zircons can be dissolved (Watson and Harrison, 1983; Hanchar and Watson, 2003). If inherited xenocrystic zircon grains and cores exist in granodioritic plutons they have two possible sources: (1) they can be assimilated from the partly digested wall-rocks that the pluton intruded; or (2) they can be inherited from the melt-source region of the pluton as incompletely resorbed igneous or metasedimentary grains in the magma.

With the exception of the Elves Chasm basement, the plutons in Grand Canyon intrude the Vishnu Schist such that inherited grains could plausibly be assimilated from Vishnu Schist wall-rocks. The similarity between U-Pb-Hf characteristics of Vishnu detrital zircons and inherited grains in plutons might suggest that this is the case.

Alternatively, since granodiorites tend to form from differentiation of basaltic magma in
lower crustal magma chambers, inherited grains may reflect the age of older lower crustal igneous material similar to those that were the provenance for the metasedimentary grains. Recent models that emphasize the potential importance of sedimentary recycling via relamination (Hacker et al., 2011) allow for a hybridized possibility whereby xenocrystic grains are inherited from recycled lower crust composed of a relaminated Vishnu-like source.

The origin of zircon inheritance in granitic bodies is difficult to determine without ambiguity (Corfu et al., 2003). However, with the aid of well characterized tectonic histories and field observations, scrutiny of zircon morphology, internal texture, and composition can help elucidate the varied history of complex zircons (Whitehouse et al., 1999; Miller et al. 2000; Cavosie et al., 2004, 2006, 2007; Bea et al., 2007). With detailed structural, geochronological, and thermobarometric studies of the Grand Canyon transect already completed (Ilg et al., 1996; Hawkins et al., 1996; Dumond et al., 2007; Shufeldt et al., 2010), I attempt to interpret the nature of zircon inheritance west of the Crystal shear zone, and the lack of inheritance east of the shear zone, in the context of all available data.

Internal textures of inherited zircons as revealed by CL and BSE imaging primarily show igneous growth zoning, but textures vary even within the same pluton. The inherited cores of the Boucher pluton range from rounded and homogenously CLdark (K12-96.2L-1C) to almost euhedral with complex igneous zoning (K12-96.2L-17C). Similar to the Boucher pluton, the Tuna Creek pluton yielded an inherited core overgrown by a young rim (K05-100.5-1). In contrast, however, many of the inherited grains in plutons west of Crystal shear zone show no rims and range from euhedral to
rounded (Figure 9). Rounded cores or xenocrysts may suggest a detrital origin, however partial resorbtion of igneous grains can yield a similar morphology. In addition, many zircons found in igneous or meta-igneous granulite grade lower crustal xenoliths display rounded morphologies and irregular internal textures (Hanchar and Rudnick, 1995; Corfu et al., 2003; Crowley et al. 2006, 2008; Siebel et al., 2011; Sommer et al., 2013).

The varied morphology and internal texture of xenocrystic grains combined with the similarity in ages between inherited grains and the Vishnu Schist would seem to make a strong case for the assimilation of Vishnu wall-rocks by plutons west of the Crystal shear zone. However, this interpretation begs the question of why there are no inherited zircons in the plutons east of the Crystal shear zone? The Vishnu basin is uniform across the entire Grand Canyon transect, i.e. plutons on either side of the shear zone are intruding the same wall rocks. Why then, do some preserve inherited grains and not others? One explanation is that some plutons crystallized from melts that were initially undersaturated in Zr , and that any zircons assimilated from the Vishnu wall-rocks were dissolved rather than preserved. This explanation is unsatisfactory because resorption of the older and more compositionally evolved Vishnu Schist detrital zircons would substantially lower the $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values of any zircons that ultimately crystallized from the melt.

In addition, both juvenile plutons and those with inheritance display similar field relations with Vishnu Schist (Figure 13). While many of the plutons across the transect preserve intrusive contacts, the Zoroaster and Tuna plutons in particular are intimately interlayered with Vishnu Schist metaturbidites, yet have completely contrasting U-Pb-Hf systematics.


Figure 13. Field photo of the Zoroaster pluton. The pluton is intimately injected into Vishnu Schist and yet shows no evidence for assimilation of Vishnu Schist.

If the xenocrystic population of zircons west of Crystal shear zone is not derived from the Vishnu wall-rocks, then it is likely that they are derived from preexisting lower crust with which juvenile arc magmas interacted. This is my preferred interpretation. The implication of this alternative source for xenocrystic grains is that west of Crystal shear zone there is a lower crustal substrate, perhaps variably distributed, that is substantially older than the crust currently exposed in the Grand Canyon region. The Crystal shear zone has previously been proposed as a suture between the Yavapai and Mojave provinces, and the presence of Archean crust in the Mojave province has been postulated (Wooden et al., 1994; Barth et al., 2000; Whitmeyer and Karlstrom 2007; Shufeldt et al., 2010). These data support the interpretation that the Crystal shear zone represents a discrete crustal boundary between the juvenile Yavapai province and the more evolved Mojave province. Furthermore, it is my interpretation that the Mojave province contains an older lower crustal substrate, and that the boundary between the Mojave and Yavapai provinces can be defined by the presence of older crust to the west of the Crystal shear zone.

## Tectonic Model for Crustal Architecture

A long-standing model for the tectonic history of southwestern North America is that of successive accretion of juvenile arc terranes to the cratonic margin (Bennet and DePaolo, 1987; Karlstrom and Bowring, 1988; Whitmeyer and Karlstrom, 2007). The discovery of numerous locations that have evidence for pre-1.8 Ga crust in the orogeny (Hawkins et al., 1996; Hill and Bickford, 2001; Bickford and Hill, 2007; Bickford et al., 2008), as well as substantial Archean detrital zircons in the Vishnu Schist (Shufeldt et al., 2010), lead to the question of the extent and character of older crust involved in the
assembly of Proterozoic terranes of Laurentia (Karlstrom et al., 2007). Paired analysis of metasedimentary rocks and the igneous rocks that intrude them provides unique insight into lithospheric formation in the Grand Canyon region.

A summary of our observations and interpretations are presented as a schematic cross section in Figure 14. I envision the Grand Canyon transect as an example of a midcrustal section of an imbricated accretionary complex. The lithologies exposed in the transect are consistent with an arc setting. Interlayered bimodal metavolcanics and turbidites of the Grand Canyon Metamorphic Suite locally preserve pillow basalts, and Bouma sequences that suggest deposition in a marine environment (Ilg et al., 1996). The present geometry of the transect is characterized by a strong vertical foliation attributed to shortening during the 1.7 Ga Yavapai orogeny. Many plutons are aligned with this fabric, and outcrop as tectonic slivers cut by subvertical shear zones. Thus, I interpret many of these plutons to represent imbricated fragments of arc batholiths rather than the batholiths themselves. High pressure ultramafic cumulates (Low et al., in progress) are also imbricated amongst Vishnu turbidites. These cumulates represent the roots of arc magma chambers that crystallized at $\sim 30 \mathrm{~km}$ depths, but are now interleaved with turbidites, suggesting substantial relative transport (minimum 5 km ) of ultramafic slivers. The subvertical shear zones across the transect accommodated shortening during the Yavapai orogeny, and minimal vertical offset post 1.7 Ga . However, my view is that the Crystal shear zone (and perhaps others) has an earlier $\mathrm{D}_{1}$ history and has been transposed into a $\mathrm{D}_{2}$ orientation. Suturing happened either before or synchronously with deposition of a cratonically sourced Vishnu Schist across both types of crust. The only pluton east of Crystal shear zone to yield xenocrystic zircons is the Boucher pluton, which lies between


Figure 14. Schematic cross section corresponding to Figure 2A. The Grand Canyon transect is shown here as a mid-crustal imbricated accretionary complex. Red circles show the concept of plutons probing their melt source regions.
the 96 -Mile shear zone and the Crystal shear zone. The $96-\mathrm{Mile}$ shear zone is shown as a minor splay of the Crystal shear zone in our cross section (Figure 14).

Karlstrom et al. (2003) proposed that the area between the Crystal and Gneiss Canyon shear zones was an isotopically mixed boundary between the Mojave and Yavapai provinces, analogous to the boundary in western Arizona described by Wooden and DeWitt (1991). Our sample set from the Grand Canyon includes only one pluton west of the Gneiss Canyon shear zone, the 245-Mile pluton (K06-245), which yielded xenocrystic zircons as well as juvenile values in primary grains (Figure 9). Zircons separated from end member Mojave plutons yield substantially evolved $\varepsilon \mathrm{Hf}_{(t)}$ values at 1.78-1.75 Ga (Figure 12) (Wooden et al., 2012) that are in contrast to the juvenile values of Paleoproterozoic grains from the 245-Mile pluton. While I interpret the presence of xenocrystic zircons in plutons west of Crystal shear zone to be derived from older crustal fragments that are the source of an isotopically mixed zone, the Gneiss Canyon shear zone is not the western edge of this zone. The presence of the Elves Chasm Gneiss in Mojave crust west of Crystal shear zone provides direct evidence for the presence of older crust in the Mojave province, however our new data suggest that the crustal architecture is even more complex. I propose a heterogeneous crustal architecture west of Crystal shear zone that includes: 1) juvenile 1.74 Ga crust with an inherited Archean component, 2) the juvenile 1840 Ma Elves Chasm Gneiss, and 3) fragments of Archean crust which, combined with juvenile Paleoproterozoic melts, are responsible for the observed isotopically mixed zone. Furthermore, I propose that these older crustal fragments increase in volume to the west, and more extensive mixing of these fragments with Paleoproterozoic melts are the source of the inherited isotopic signature of the

Mojave Province. My interpretations of these data reconcile the model for an isotopically mixed boundary zone between Mojave and Yavapai crust (Wooden and DeWitt, 1991; Karlstrom et al., 2003; Duebendorfer et al., 2006) with the discrete boundary at Crystal shear zone, and provide an explanation for the source of the characteristic evolved isotopic signature of the Mojave province.

Figure 14 shows the current geometry of the transect. I propose two possible tectonic cartoons that explain all observations and might culminate in the present geometry of the transect (Figure 15). Deposition of the Vishnu Schist from 1750-1740 Ma occurred either after or during final suturing between the Mojave and Yavapai provinces. A shallowly dipping tectonic layering $\left(\mathrm{S}_{1}\right)$ began to form shortly after deposition, likely facilitated by thrust stacking of turbidites during thickening of arc crust in the early stages of the Yavapai orogeny (Ilg et al., 1996). Near synchronously with the development of $S_{1}$, calk-alkaline granodiorite plutons began to intrude the Granite Gorge Metamorphic Suite and become imbricated within thrust sheets. The transect achieved its current configuration by the end of the Yavapai orogeny, and was minimally reactivated during later Mesoproterozoic tectonism (Karlstrom et al., 1997). Deposition of the Vishnu Schist may have occurred in a trench fill sequence during final convergence between the Mojave and Yavapai provinces (Figure 15A). Alternatively, deposition of the Vishnu Schist may have occurred after pre-1750 Ma suturing of the Mojave and Yavapai provinces in a back arc basin subsequently deformed and metamorphosed during the Yavapai orogeny (Figure 15B).


Mojave provinces are already joined. While no >1750 Ma Yavapai crust exists in the Grand Canyon region, older Yavapai crust exists to the northeast in Colorado. D: The Vishnu Schist is deposited across both provinces in a back arc basin after the juxtaposition of older Yavapai crust against Mojave. E: By 1740 Ma the arc has migrated such that calk-alkaline plutons begin to intrude the Vishnu Schist. Thrust stacking imbricates Vishnu turbidites with arc plutons, and a shallow S1 foliation develops. The Crystal shear zone manifests as a boundary between Mojave and Yavapai crust. Arc plutons continue to intrude the Vishnu Schist until ~1710 Ma. F: From 1710-1690 Ma the transect undergoes protracted crustal shortening during the Yavapai orogeny, a penetrative subhorizontal fabric (S2) is developed throughout the transect, and the Crystal shear zone is transposed into this orientation. The transect is intruded by crustal melts, but no longer by calk-alkaline plutons, suggesting that the arc has either migrated or terminated. By 1700 Ma the transect resides at mid crustal levels. The box straddling the Mojave - Yavapai boundary shows the cross section of Figure 14 at 1700 Ma .

## CONCLUSIONS

Comparison of metasedimentary rocks with the plutons that intrude them provide unique insight into lithospheric formation. The Grand Canyon transect is underlain by supracrustal rocks that consist of interlayered juvenile arc volcanics, and turbidites sourced in part by evolved Paleoproterozoic crust of the Mojave province, and a yet undertermined Archean craton, but possibly the Gawler craton of Australia. The Vishnu Schist received little of its detritus from local Yavapai type crust, and the Elves Chasm Gneiss did not contribute substantial detritus to the Vishnu Schist. The depositional environment of the Vishnu Schist was certainly an arc setting, however may have been deposited in a back arc basin, or as an accretionary prism in the fore arc.

While the Vishnu Schist was deposited uniformly across both Mojave and Yavapai crust, and does not support the existence of a crustal boundary in Grand Canyon, arc plutons yield markedly different $\mathrm{U}-\mathrm{Pb}-\mathrm{Hf}$ characteristics across the Crystal shear zone. Plutons that intrude the Vishnu Schist and serve as probes of lower crustal meltsource regions across the entire transect are predominantly juvenile, but west of the Crystal shear zone xenocrystic grains suggest the presence of older lower crustal fragments, and more evolved Paleoproterozoic grains support this conclusion.

The Paleoproterozoic rocks of the Grand Canyon are an example of a $100 \%$ exposed distributed mid-crustal suture zone between the Mojave and Yavapai crustal provinces. The nature of this boundary is defined by the presence of early Archean ( $\sim 2.5$ $\mathrm{Ga})$ lower crustal fragments in the Mojave province.

## Appendix A

Table A 1: U-Pb geochronologic analyses of Vishnu Schist (ALC)

|  |  |  |  |  |  | Isotope ratios |  |  |  |  | Bestage | $\pm$ | Conc |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Analysis | U | 206Pb | $\begin{aligned} & \hline \mathrm{U} / \\ & \mathrm{Th} \\ & \hline \end{aligned}$ | $206 \mathrm{~Pb}^{*}$ | $\pm$ | 207Pb* | $\pm$ | $206 \mathrm{~Pb}^{*}$ | $\pm$ | error |  |  |  |
|  | (ppm) | 204 Pb |  | $207 \mathrm{~Pb} *$ | (\%) | 235U* | (\%) | 238 U | (\%) | corr. | (Ma) | (Ma) | (\%) |
| OS08842A-1 | 588 | 45658 | 3.3 | 8.8246 | 0.2 | 5.3112 | 0.8 | 0.3399 | 0.8 | 0.98 | 1853.3 | 2.9 | 101.8 |
| OS08842A-2 | 595 | 4986 | 7.5 | 9.0902 | 0.8 | 4.7549 | 1.6 | 0.3135 | 1.4 | 0.88 | 1799.5 | 13.8 | 97.7 |
| OS08842A-3 | 113 | 13212 | 1.5 | 5.5290 | 0.4 | 12.7337 | 1.1 | 0.5106 | 1.1 | 0.94 | 2660.8 | 6.6 | 99.9 |
| OS08842A-4 | 171 | 115575 | 2.1 | 5.4956 | 0.3 | 12.4391 | 2.6 | 0.4958 | 2.6 | 0.99 | 2670.9 | 4.6 | 97.2 |
| OS08842A-5 | 205 | 21671 | 1.9 | 8.2260 | 0.4 | 5.9110 | 2.6 | 0.3527 | 2.5 | 0.99 | 1979.3 | 7.0 | 98.4 |
| OS08842A-6 | 73 | 320950 | 2.1 | 5.7147 | 0.4 | 11.9465 | 1.1 | 0.4951 | 1.0 | 0.93 | 2605.9 | 6.7 | 99.5 |
| OS08842A-8 | 164 | 238621 | 2.3 | 6.1475 | 0.2 | 10.5619 | 0.7 | 0.4709 | 0.6 | 0.93 | 2483.6 | 4.2 | 100.2 |
| OS08842A-9 | 152 | 136824 | 2.8 | 9.1746 | 0.3 | 4.7772 | 1.3 | 0.3179 | 1.3 | 0.98 | 1782.7 | 5.2 | 99.8 |
| OS08842A-10 | 126 | 2176 | 2.7 | 9.0692 | 1.5 | 4.6245 | 7.6 | 0.3042 | 7.5 | 0.98 | 1803.7 | 28.1 | 94.9 |
| OS08842A-11 | 132 | 1717 | 1.7 | 8.9075 | 4.4 | 4.9522 | 5.8 | 0.3199 | 3.8 | 0.65 | 1836.4 | 80.0 | 97.4 |
| OS08842A-12 | 427 | 8353 | 3.6 | 5.6475 | 0.3 | 11.2544 | 3.4 | 0.4610 | 3.4 | 1.00 | 2625.6 | 5.0 | 93.1 |
| OS08842A-16 | 190 | 7979 | 1.7 | 8.7863 | 1.7 | 5.2399 | 3.0 | 0.3339 | 2.5 | 0.83 | 1861.2 | 30.4 | 99.8 |
| OS08842A-17 | 205 | 17628 | 2.5 | 6.1426 | 0.2 | 10.6348 | 0.8 | 0.4738 | 0.7 | 0.96 | 2484.9 | 3.5 | 100.6 |
| OS08842A-18 | 319 | 357901 | 3.0 | 8.7838 | 0.2 | 5.3482 | 1.9 | 0.3407 | 1.9 | 1.00 | 1861.7 | 3.5 | 101.5 |
| OS08842A-19 | 372 | 1973 | 4.0 | 9.0218 | 2.4 | 4.5185 | 3.9 | 0.2957 | 3.0 | 0.78 | 1813.3 | 43.5 | 92.1 |
| OS08842A-20 | 495 | 25846 | 1.9 | 5.8219 | 0.9 | 11.6090 | 1.1 | 0.4902 | 0.6 | 0.56 | 2574.9 | 15.8 | 99.9 |
| OS08842A-22 | 249 | 6719 | 1.7 | 9.0031 | 0.9 | 4.9700 | 2.1 | 0.3245 | 1.9 | 0.91 | 1817.0 | 15.8 | 99.7 |
| OS08842A-23 | 756 | 3964 | 0.8 | 5.4523 | 4.0 | 13.4333 | 7.9 | 0.5312 | 6.9 | 0.87 | 2683.9 | 65.5 | 102.3 |
| OS08842A-24 | 102 | 189339 | 1.4 | 5.4952 | 0.4 | 12.8493 | 1.2 | 0.5121 | 1.1 | 0.93 | 2670.9 | 7.0 | 99.8 |
| OS08842A-26 | 142 | 3647 | 4.1 | 8.6077 | 1.2 | 5.4435 | 3.8 | 0.3398 | 3.7 | 0.95 | 1898.2 | 21.0 | 99.4 |
| OS08842A-27 | 507 | 18533 | 1.2 | 8.8609 | 0.6 | 5.2128 | 1.1 | 0.3350 | 0.9 | 0.84 | 1845.9 | 10.8 | 100.9 |
| OS08842A-28 | 161 | 9874 | 3.1 | 9.1445 | 3.1 | 4.6816 | 3.2 | 0.3105 | 1.0 | 0.30 | 1788.7 | 56.2 | 97.5 |
| OS08842A-29 | 123 | 960 | 2.5 | 6.1510 | 2.5 | 10.0472 | 3.1 | 0.4482 | 1.8 | 0.59 | 2482.6 | 42.4 | 96.2 |
| OS08842A-30 | 874 | 5227 | 2.5 | 9.1396 | 0.7 | 3.7842 | 2.8 | 0.2508 | 2.7 | 0.97 | 1789.7 | 12.2 | 80.6 |
| OS08842A-32 | 322 | 56887 | 1.5 | 9.2180 | 0.3 | 4.7591 | 1.1 | 0.3182 | 1.1 | 0.97 | 1774.1 | 5.1 | 100.4 |
| OS08842A-33 | 376 | 5055 | 4.6 | 9.2201 | 0.8 | 4.3168 | 4.9 | 0.2887 | 4.8 | 0.99 | 1773.7 | 14.4 | 92.2 |
| OS08842A-34 | 463 | 10216 | 1.1 | 5.3927 | 0.4 | 12.0593 | 9.2 | 0.4717 | 9.2 | 1.00 | 2702.1 | 6.2 | 92.2 |
| OS08842A-35 | 302 | 1395 | 1.5 | 8.2050 | 1.0 | 5.9496 | 1.6 | 0.3540 | 1.2 | 0.76 | 1983.9 | 18.4 | 98.5 |
| OS08842A-36 | 220 | 4451 | 1.2 | 8.5327 | 2.4 | 5.3154 | 6.4 | 0.3289 | 5.9 | 0.92 | 1913.9 | 43.8 | 95.8 |
| OS08842A-37 | 256 | 3866 | 1.3 | 8.4832 | 1.7 | 5.6248 | 2.7 | 0.3461 | 2.1 | 0.78 | 1924.3 | 30.6 | 99.6 |
| OS08842A-38 | 485 | 23182 | 2.7 | 5.9428 | 0.3 | 11.5527 | 2.0 | 0.4979 | 2.0 | 0.99 | 2540.5 | 4.6 | 102.5 |
| OS08842A-39 | 459 | 2800 | 2.4 | 8.9282 | 1.0 | 4.6374 | 2.6 | 0.3003 | 2.4 | 0.92 | 1832.2 | 18.6 | 92.4 |
| OS08842A-40 | 335 | 92264 | 2.5 | 9.1600 | 0.2 | 4.8312 | 0.8 | 0.3210 | 0.8 | 0.96 | 1785.6 | 4.1 | 100.5 |
| OS08842A-42 | 210 | 37721 | 1.6 | 9.2027 | 0.6 | 4.7698 | 1.9 | 0.3184 | 1.8 | 0.95 | 1777.1 | 10.6 | 100.3 |
| OS08842A-43 | 304 | 6355 | 5.1 | 8.6851 | 0.7 | 5.0918 | 2.0 | 0.3207 | 1.8 | 0.93 | 1882.0 | 13.3 | 95.3 |


|  |  |  |  |  |  |  |  |  |  |  |  |  |
| :--- | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: |
| OS08842A-44 | 218 | 2317 | 1.4 | 8.5799 | 2.2 | 5.4938 | 2.6 | 0.3419 | 1.2 | 0.48 | 1904.0 | 40.2 | 99.6


| K05-986-104-1 | 830 | 1381 | 4.1 | 9.0766 | 1.3 | 2.9416 | 6.6 | 0.1936 | 6.5 | 0.98 | 1802.3 | 23.0 | 63.3 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K05-986-104-4 | 242 | 3028 | 1.3 | 5.9905 | 1.8 | 10.1160 | 2.7 | 0.4395 | 2.0 | 0.74 | 2527.1 | 31.1 | 92.9 |
| K05-986-104-5 | 192 | 22659 | 2.1 | 9.3791 | 0.4 | 4.6817 | 1.4 | 0.3185 | 1.4 | 0.97 | 1742.4 | 6.8 | 102.3 |
| K05-986-104-6 | 258 | 5032 | 1.4 | 6.1769 | 0.2 | 10.0099 | 0.9 | 0.4484 | 0.9 | 0.98 | 2475.5 | 2.6 | 96.5 |
| K05-986-104-7 | 244 | 7692 | 1.0 | 9.3142 | 0.7 | 4.5265 | 1.2 | 0.3058 | 1.0 | 0.79 | 1755.1 | 13.6 | 98.0 |
| K05-986-104-8 | 138 | 12048 | 1.0 | 9.3143 | 0.5 | 4.5003 | 0.7 | 0.3040 | 0.6 | 0.78 | 1755.1 | 8.4 | 97.5 |
| K05-986-104-9 | 313 | 5834 | 1.8 | 9.3542 | 0.4 | 4.1122 | 0.9 | 0.2790 | 0.8 | 0.89 | 1747.3 | 7.8 | 90.8 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 10 \end{aligned}$ | 369 | 2056 | 1.3 | 9.0653 | 3.5 | 4.5578 | 9.8 | 0.2997 | 9.1 | 0.93 | 1804.5 | 64.2 | 93.6 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 14 \\ & \hline \end{aligned}$ | 204 | 8778 | 1.0 | 5.6284 | 0.2 | 11.9633 | 0.7 | 0.4884 | 0.7 | 0.96 | 2631.2 | 3.3 | 97.4 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 15 \end{aligned}$ | 293 | 1408 | 1.3 | 8.8712 | 6.2 | 4.4617 | 6.4 | 0.2871 | 1.7 | 0.26 | 1843.8 | 111.5 | 88.2 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 16 \\ & \hline \end{aligned}$ | 91 | 150005 | 1.5 | 9.3946 | 0.6 | 4.5478 | 1.2 | 0.3099 | 1.0 | 0.88 | 1739.4 | 10.1 | 100.0 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 17 \\ & \hline \end{aligned}$ | 74 | 102799 | 2.5 | 9.3699 | 0.5 | 4.5814 | 1.3 | 0.3113 | 1.2 | 0.91 | 1744.2 | 10.0 | 100.2 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 18 \\ & \hline \end{aligned}$ | 128 | 22996 | 2.1 | 6.2827 | 1.9 | 11.1220 | 6.1 | 0.5068 | 5.8 | 0.95 | 2446.8 | 32.7 | 108.0 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 20 \end{aligned}$ | 120 | 1157 | 1.1 | 8.9314 | 3.6 | 4.5807 | 4.7 | 0.2967 | 3.0 | 0.63 | 1831.5 | 66.1 | 91.5 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 24 \\ & \hline \end{aligned}$ | 342 | 4979 | 1.8 | 8.0869 | 0.8 | 6.0475 | 4.7 | 0.3547 | 4.6 | 0.98 | 2009.6 | 14.7 | 97.4 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 26 \\ & \hline \end{aligned}$ | 170 | 194538 | 2.0 | 8.9960 | 0.4 | 5.1036 | 1.1 | 0.3330 | 1.0 | 0.94 | 1818.5 | 6.6 | 101.9 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 28 \end{aligned}$ | 251 | 8655 | 0.9 | 9.2958 | 1.1 | 4.5231 | 1.2 | 0.3049 | 0.5 | 0.45 | 1758.7 | 19.4 | 97.6 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 32 \end{aligned}$ | 280 | 314924 | 2.2 | 8.1661 | 0.1 | 6.1980 | 0.5 | 0.3671 | 0.5 | 0.99 | 1992.3 | 1.2 | 101.2 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 34 \\ & \hline \end{aligned}$ | 397 | 10724 | 1.5 | 9.3691 | 0.6 | 4.5634 | 1.1 | 0.3101 | 0.9 | 0.84 | 1744.4 | 10.7 | 99.8 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 35 \\ & \hline \end{aligned}$ | 209 | 132336 | 1.6 | 9.3386 | 0.4 | 4.6442 | 0.8 | 0.3146 | 0.7 | 0.87 | 1750.3 | 7.2 | 100.7 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 36 \end{aligned}$ | 250 | 1262 | 1.3 | 8.8793 | 4.0 | 4.6292 | 4.7 | 0.2981 | 2.5 | 0.53 | 1842.1 | 71.7 | 91.3 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 37 \\ & \hline \end{aligned}$ | 111 | 154798 | 1.4 | 9.3703 | 0.3 | 4.6279 | 0.9 | 0.3145 | 0.8 | 0.93 | 1744.1 | 6.1 | 101.1 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 38 \\ & \hline \end{aligned}$ | 87 | 279928 | 1.4 | 8.9715 | 0.6 | 5.0837 | 2.0 | 0.3308 | 1.9 | 0.95 | 1823.4 | 11.7 | 101.0 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 40 \\ & \hline \end{aligned}$ | 102 | 201577 | 0.9 | 9.4381 | 0.6 | 4.5528 | 0.9 | 0.3116 | 0.7 | 0.74 | 1730.9 | 11.0 | 101.0 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 41 \\ & \hline \end{aligned}$ | 215 | 13494 | 1.5 | 9.3215 | 0.5 | 4.5830 | 1.1 | 0.3098 | 1.0 | 0.90 | 1753.7 | 8.5 | 99.2 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 42 \\ & \hline \end{aligned}$ | 611 | 1166 | 0.9 | 9.0851 | 0.6 | 2.9318 | 6.0 | 0.1932 | 6.0 | 1.00 | 1800.6 | 10.3 | 63.2 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 43 \\ & \hline \end{aligned}$ | 105 | 14482 | 1.2 | 9.2984 | 0.9 | 4.6050 | 1.5 | 0.3106 | 1.3 | 0.81 | 1758.2 | 16.5 | 99.2 |
| $\begin{aligned} & \hline \text { K05-986-104- } \\ & 44 \\ & \hline \end{aligned}$ | 123 | 2799 | 0.8 | 9.2008 | 1.3 | 4.5540 | 1.7 | 0.3039 | 1.1 | 0.63 | 1777.5 | 23.9 | 96.2 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 47 \\ & \hline \end{aligned}$ | 193 | 357875 | 1.6 | 9.3762 | 0.2 | 4.5844 | 0.7 | 0.3118 | 0.7 | 0.95 | 1743.0 | 3.9 | 100.4 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 46 \\ & \hline \end{aligned}$ | 302 | 3219 | 1.5 | 9.3176 | 0.4 | 4.3013 | 4.7 | 0.2907 | 4.7 | 1.00 | 1754.4 | 8.0 | 93.8 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 49 \end{aligned}$ | 188 | 4038 | 2.4 | 9.0598 | 1.0 | 4.9456 | 1.2 | 0.3250 | 0.7 | 0.59 | 1805.6 | 18.1 | 100.5 |
| $\begin{aligned} & \text { K05-986-104- } \\ & 50 \end{aligned}$ | 104 | 10391 | 1.4 | 9.3402 | 0.7 | 4.5905 | 1.3 | 0.3110 | 1.2 | 0.87 | 1750.0 | 12.2 | 99.7 |
| OS081084A-2 | 318 | 15764 | 1.4 | 9.3770 | 0.3 | 4.3287 | 1.8 | 0.2944 | 1.7 | 0.98 | 1742.8 | 6.3 | 95.4 |
| OS081084A-3 | 717 | 99257 | 3.1 | 9.1754 | 0.1 | 4.8294 | 0.4 | 0.3214 | 0.4 | 0.95 | 1782.5 | 2.3 | 100.8 |
| OS081084A-4 | 135 | 127978 | 1.6 | 5.6327 | 0.3 | 12.3175 | 2.4 | 0.5032 | 2.3 | 0.99 | 2630.0 | 4.6 | 99.9 |
| OS081084A-5 | 230 | 82242 | 2.1 | 9.3230 | 0.5 | 4.7122 | 0.7 | 0.3186 | 0.5 | 0.67 | 1753.4 | 9.0 | 101.7 |
| OS081084A-6 | 149 | 64143 | 1.6 | 9.3705 | 0.5 | 4.6468 | 1.7 | 0.3158 | 1.6 | 0.96 | 1744.1 | 9.1 | 101.4 |
| OS081084A-7 | 154 | 9908 | 2.3 | 6.1768 | 0.4 | 10.3292 | 2.5 | 0.4627 | 2.5 | 0.98 | 2475.5 | 7.6 | 99.0 |
| OS081084A-8 | 218 | 12998 | 1.0 | 9.3711 | 0.5 | 4.6244 | 0.9 | 0.3143 | 0.8 | 0.81 | 1744.0 | 9.9 | 101.0 |
| OS081084A-9 | 165 | 121550 | 1.6 | 9.3511 | 0.8 | 4.5349 | 1.9 | 0.3076 | 1.7 | 0.91 | 1747.9 | 14.1 | 98.9 |
| $\begin{aligned} & \text { OS081084A- } \\ & 10 \end{aligned}$ | 176 | 19928 | 1.7 | 9.2803 | 0.9 | 4.7012 | 1.7 | 0.3164 | 1.5 | 0.85 | 1761.8 | 16.7 | 100.6 |


| $\begin{aligned} & \text { OS081084A- } \\ & 11 \\ & \hline \end{aligned}$ | 191 | 110670 | 1.2 | 9.3625 | 0.3 | 4.6376 | 0.8 | 0.3149 | 0.7 | 0.92 | 1745.6 | 5.8 | 101.1 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\begin{aligned} & \text { OS081084A- } \\ & 12 \end{aligned}$ | 186 | 135569 | 2.0 | 9.3866 | 0.6 | 4.6024 | 1.9 | 0.3133 | 1.8 | 0.95 | 1740.9 | 10.4 | 100.9 |
| $\begin{aligned} & \text { OS081084A- } \\ & 13 \\ & \hline \end{aligned}$ | 500 | 11617 | 1.9 | 9.3056 | 0.3 | 4.5772 | 1.4 | 0.3089 | 1.4 | 0.98 | 1756.8 | 4.6 | 98.8 |
| $\begin{aligned} & \text { OS081084A- } \\ & 15 \end{aligned}$ | 148 | 2770 | 1.7 | 9.2668 | 1.1 | 3.9392 | 3.2 | 0.2648 | 3.0 | 0.94 | 1764.4 | 20.4 | 85.8 |
| $\begin{aligned} & \text { OS081084A- } \\ & 16 \\ & \hline \end{aligned}$ | 99 | 253724 | 2.1 | 9.1073 | 0.7 | 4.9574 | 2.9 | 0.3274 | 2.8 | 0.97 | 1796.1 | 13.3 | 101.7 |
| $\begin{aligned} & \text { OS081084A- } \\ & 17 \\ & \hline \end{aligned}$ | 194 | 31939 | 2.5 | 9.2355 | 0.4 | 4.7554 | 2.5 | 0.3185 | 2.5 | 0.99 | 1770.6 | 7.7 | 100.7 |
| $\begin{aligned} & \text { OS081084A- } \\ & 18 \\ & \hline \end{aligned}$ | 500 | 61432 | 1.7 | 9.3695 | 0.2 | 4.6620 | 0.7 | 0.3168 | 0.7 | 0.97 | 1744.3 | 3.5 | 101.7 |
| $\begin{aligned} & \text { OS081084A- } \\ & 19 \\ & \hline \end{aligned}$ | 349 | 28714 | 2.1 | 9.3680 | 0.3 | 4.5977 | 0.9 | 0.3124 | 0.8 | 0.94 | 1744.6 | 5.4 | 100.5 |
| $\begin{aligned} & \text { OS081084A- } \\ & 20 \\ & \hline \end{aligned}$ | 210 | 14219 | 1.1 | 8.7618 | 0.4 | 5.2079 | 0.8 | 0.3309 | 0.8 | 0.90 | 1866.2 | 6.5 | 98.8 |
| $\begin{aligned} & \text { OS081084A- } \\ & 21 \\ & \hline \end{aligned}$ | 268 | 228678 | 2.4 | 9.3428 | 0.4 | 4.6342 | 0.9 | 0.3140 | 0.7 | 0.87 | 1749.5 | 7.7 | 100.6 |
| $\begin{aligned} & \text { OS081084A- } \\ & 22 \\ & \hline \end{aligned}$ | 323 | 4044 | 1.2 | 9.2513 | 1.8 | 4.4306 | 2.2 | 0.2973 | 1.2 | 0.54 | 1767.5 | 33.0 | 94.9 |
| $\begin{aligned} & \text { OS081084A- } \\ & 23 \end{aligned}$ | 257 | 150584 | 3.0 | 9.1794 | 0.3 | 4.8256 | 0.9 | 0.3213 | 0.9 | 0.95 | 1781.7 | 5.1 | 100.8 |
| $\begin{aligned} & \text { OS081084A- } \\ & 24 \\ & \hline \end{aligned}$ | 193 | 133971 | 1.0 | 9.4194 | 0.3 | 4.6109 | 0.6 | 0.3150 | 0.5 | 0.82 | 1734.5 | 6.1 | 101.8 |
| $\begin{aligned} & \text { OS081084A- } \\ & 26 \\ & \hline \end{aligned}$ | 158 | 67211 | 1.7 | 9.4024 | 0.3 | 4.7002 | 1.8 | 0.3205 | 1.8 | 0.98 | 1737.9 | 6.0 | 103.1 |
| $\begin{aligned} & \text { OS081084A- } \\ & 27 \\ & \hline \end{aligned}$ | 227 | 43036 | 1.7 | 9.3857 | 0.2 | 4.5293 | 0.6 | 0.3083 | 0.6 | 0.92 | 1741.1 | 4.4 | 99.5 |
| $\begin{aligned} & \text { OS081084A- } \\ & 28 \end{aligned}$ | 342 | 42528 | 1.4 | 9.3725 | 0.4 | 4.5490 | 1.4 | 0.3092 | 1.4 | 0.96 | 1743.7 | 7.6 | 99.6 |
| $\begin{aligned} & \text { OS081084A- } \\ & 29 \\ & \hline \end{aligned}$ | 157 | 86640 | 1.2 | 9.3931 | 0.6 | 4.5694 | 1.1 | 0.3113 | 0.9 | 0.86 | 1739.7 | 10.2 | 100.4 |
| $\begin{aligned} & \text { OS081084A- } \\ & 30 \end{aligned}$ | 247 | 47528 | 1.4 | 9.3907 | 0.3 | 4.6186 | 0.8 | 0.3146 | 0.7 | 0.92 | 1740.1 | 5.7 | 101.3 |
| $\begin{aligned} & \text { OS081084A- } \\ & 31 \\ & \hline \end{aligned}$ | 140 | 187648 | 1.5 | 9.3851 | 0.5 | 4.5498 | 0.9 | 0.3097 | 0.8 | 0.83 | 1741.2 | 9.4 | 99.9 |
| $\begin{aligned} & \text { OS081084A- } \\ & 32 \end{aligned}$ | 368 | 92979 | 1.4 | 9.3727 | 0.3 | 4.5781 | 0.7 | 0.3112 | 0.7 | 0.92 | 1743.7 | 5.2 | 100.2 |
| $\begin{aligned} & \text { OS081084A- } \\ & 33 \end{aligned}$ | 301 | 28489 | 2.4 | 8.8694 | 0.3 | 5.1836 | 1.1 | 0.3334 | 1.0 | 0.97 | 1844.2 | 5.1 | 100.6 |
| $\begin{aligned} & \text { OS081084A- } \\ & 35 \\ & \hline \end{aligned}$ | 221 | 3968 | 1.7 | 6.2203 | 1.5 | 9.5372 | 2.8 | 0.4303 | 2.4 | 0.84 | 2463.7 | 26.1 | 93.6 |
| $\begin{aligned} & \text { OS081084A- } \\ & 37 \\ & \hline \end{aligned}$ | 217 | 53461 | 0.8 | 7.8960 | 0.3 | 6.4130 | 1.0 | 0.3673 | 1.0 | 0.94 | 2051.9 | 6.1 | 98.3 |
| $\begin{aligned} & \text { OS081084A- } \\ & 38 \\ & \hline \end{aligned}$ | 123 | 67202 | 1.4 | 9.3877 | 1.0 | 4.5401 | 1.2 | 0.3091 | 0.7 | 0.54 | 1740.7 | 18.6 | 99.7 |
| $\begin{aligned} & \text { OS081084A- } \\ & 39 \\ & \hline \end{aligned}$ | 183 | 85874 | 2.1 | 9.1249 | 0.8 | 4.8650 | 1.3 | 0.3220 | 1.0 | 0.79 | 1792.6 | 14.6 | 100.4 |
| $\begin{aligned} & \text { OS081084A- } \\ & 40 \end{aligned}$ | 157 | 27788 | 1.4 | 6.0905 | 1.2 | 10.8122 | 8.9 | 0.4776 | 8.8 | 0.99 | 2499.3 | 20.3 | 100.7 |
| $\begin{aligned} & \text { OS081084A- } \\ & 41 \\ & \hline \end{aligned}$ | 127 | 8664 | 1.6 | 9.1150 | 1.2 | 4.5180 | 3.1 | 0.2987 | 2.9 | 0.92 | 1794.6 | 22.5 | 93.9 |
| $\begin{aligned} & \text { OS081084A- } \\ & 42 \end{aligned}$ | 115 | 111104 | 1.3 | 9.4312 | 1.1 | 4.3916 | 1.4 | 0.3004 | 0.8 | 0.60 | 1732.2 | 19.9 | 97.7 |
| $\begin{aligned} & \text { OS081084A- } \\ & 43 \\ & \hline \end{aligned}$ | 78 | 87863 | 0.5 | 5.8336 | 0.4 | 11.5480 | 0.8 | 0.4886 | 0.7 | 0.86 | 2571.6 | 7.1 | 99.7 |
| $\begin{aligned} & \hline \text { OS081084A- } \\ & 44 \\ & \hline \end{aligned}$ | 202 | 127828 | 1.9 | 6.1876 | 0.2 | 10.3607 | 0.7 | 0.4650 | 0.6 | 0.93 | 2472.6 | 4.1 | 99.5 |
| $\begin{aligned} & \text { OS081084A- } \\ & 45 \\ & \hline \end{aligned}$ | 77 | 30571 | 2.1 | 8.2306 | 0.7 | 5.9967 | 2.0 | 0.3580 | 1.9 | 0.94 | 1978.3 | 12.2 | 99.7 |
| $\begin{aligned} & \text { OS081084A- } \\ & 46 \end{aligned}$ | 313 | 51164 | 2.7 | 9.1050 | 0.4 | 4.7408 | 1.6 | 0.3131 | 1.5 | 0.97 | 1796.6 | 6.7 | 97.7 |
| $\begin{aligned} & \text { OS081084A- } \\ & 47 \\ & \hline \end{aligned}$ | 258 | 2346 | 2.9 | 9.3608 | 1.4 | 3.9489 | 3.0 | 0.2681 | 2.7 | 0.89 | 1746.0 | 25.1 | 87.7 |
| $\begin{aligned} & \text { OS081084A- } \\ & 48 \\ & \hline \end{aligned}$ | 487 | 43009 | 2.2 | 9.3240 | 0.8 | 4.6450 | 3.5 | 0.3141 | 3.4 | 0.97 | 1753.2 | 14.6 | 100.4 |
| $\begin{aligned} & \hline \text { OS081084A- } \\ & 49 \\ & \hline \end{aligned}$ | 174 | 20970 | 2.4 | 9.0127 | 0.4 | 4.8174 | 1.9 | 0.3149 | 1.8 | 0.97 | 1815.1 | 7.8 | 97.2 |
| $\begin{aligned} & \text { OS081084A- } \\ & 50 \\ & \hline \end{aligned}$ | 413 | 6301 | 1.6 | 8.9043 | 0.6 | 5.1005 | 1.9 | 0.3294 | 1.8 | 0.95 | 1837.0 | 10.6 | 99.9 |
| $\begin{aligned} & \text { OS081084A- } \\ & 51 \\ & \hline \end{aligned}$ | 275 | 265308 | 2.7 | 6.7867 | 0.2 | 8.7576 | 1.4 | 0.4311 | 1.4 | 0.99 | 2315.4 | 3.2 | 99.8 |
| $\begin{aligned} & \text { OS081084A- } \\ & 52 \\ & \hline \end{aligned}$ | 235 | 22017 | 2.0 | 5.1894 | 0.4 | 14.7560 | 8.0 | 0.5554 | 8.0 | 1.00 | 2765.3 | 5.9 | 103.0 |
| $\begin{aligned} & \text { OS081084A- } \\ & 53 \\ & \hline \end{aligned}$ | 243 | 19783 | 1.5 | 9.4385 | 0.9 | 4.6827 | 6.3 | 0.3206 | 6.3 | 0.99 | 1730.8 | 17.2 | 103.6 |
| $\begin{aligned} & \text { OS081084A- } \\ & 54 \end{aligned}$ | 475 | 31217 | 1.6 | 9.3914 | 0.3 | 4.4573 | 0.7 | 0.3036 | 0.6 | 0.91 | 1740.0 | 5.4 | 98.2 |
| $\begin{aligned} & \text { OS081084A- } \\ & 55 \end{aligned}$ | 118 | 28039 | 0.9 | 3.9003 | 0.4 | 22.1873 | 1.1 | 0.6276 | 1.0 | 0.94 | 3224.6 | 5.7 | 97.4 |


| OS081084A57 | 250 | 243530 | 4.6 | 9.0068 | 0.2 | 4.9847 | 0.6 | 0.3256 | 0.5 | 0.93 | 1816.3 | 4.0 | 100.0 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\begin{aligned} & \text { OS081084A- } \\ & 58 \end{aligned}$ | 205 | 10869 | 1.1 | 9.4491 | 0.4 | 4.1978 | 2.7 | 0.2877 | 2.7 | 0.99 | 1728.8 | 8.1 | 94.3 |
| $\begin{aligned} & \text { OS081084A- } \\ & 59 \end{aligned}$ | 214 | 38658 | 1.0 | 9.3396 | 1.3 | 4.7546 | 3.4 | 0.3221 | 3.1 | 0.93 | 1750.1 | 23.3 | 102.8 |
| $\begin{aligned} & \text { OS081084A- } \\ & 60 \\ & \hline \end{aligned}$ | 36 | 45771 | 2.1 | 5.8193 | 1.0 | 11.8006 | 2.0 | 0.4980 | 1.7 | 0.86 | 2575.7 | 16.9 | 101.2 |
| $\begin{aligned} & \text { OS081084A- } \\ & 62 \\ & \hline \end{aligned}$ | 274 | 256305 | 2.9 | 9.1619 | 0.2 | 4.7781 | 0.5 | 0.3175 | 0.5 | 0.92 | 1785.2 | 3.8 | 99.6 |
| $\begin{aligned} & \text { OS081084A- } \\ & 65 \end{aligned}$ | 241 | 21650 | 1.7 | 9.4060 | 0.5 | 4.5640 | 1.1 | 0.3113 | 0.9 | 0.87 | 1737.2 | 9.5 | 100.6 |
| $\begin{aligned} & \text { OS081084A- } \\ & 68 \end{aligned}$ | 114 | 67712 | 2.7 | 9.3177 | 1.2 | 4.6995 | 1.5 | 0.3176 | 0.8 | 0.57 | 1754.4 | 22.4 | 101.3 |
| $\begin{aligned} & \text { OS081084A- } \\ & 71 \\ & \hline \end{aligned}$ | 187 | 15202 | 0.7 | 9.4297 | 0.6 | 4.5665 | 2.5 | 0.3123 | 2.4 | 0.97 | 1732.5 | 11.4 | 101.1 |
| $\begin{aligned} & \text { OS081084A- } \\ & 72 \end{aligned}$ | 370 | 10148 | 2.1 | 8.8052 | 1.0 | 5.2982 | 5.6 | 0.3384 | 5.5 | 0.98 | 1857.3 | 18.9 | 101.2 |
| $\begin{aligned} & \text { OS081084A- } \\ & 74 \end{aligned}$ | 277 | 82513 | 2.2 | 6.1525 | 0.2 | 10.6335 | 1.1 | 0.4745 | 1.1 | 0.99 | 2482.2 | 2.7 | 100.8 |
| $\begin{aligned} & \text { OS081084A- } \\ & 75 \\ & \hline \end{aligned}$ | 179 | 332331 | 2.5 | 6.1642 | 0.2 | 10.3847 | 2.1 | 0.4643 | 2.1 | 1.00 | 2479.0 | 3.5 | 99.2 |
| $\begin{aligned} & \text { OS081084A- } \\ & 76 \\ & \hline \end{aligned}$ | 365 | 6562 | 1.7 | 9.2988 | 1.4 | 4.7600 | 2.8 | 0.3210 | 2.5 | 0.87 | 1758.1 | 25.7 | 102.1 |
| OS081084A- | 158 | 168317 | 1.1 | 9.3965 | 0.6 | 4.5518 | 0.9 | 0.3102 | 0.7 | 0.75 | 1739.0 | 10.7 | 100.2 |
| $\begin{aligned} & \text { OS081084A- } \\ & 80 \\ & \hline \end{aligned}$ | 365 | 32625 | 1.6 | 9.3441 | 0.3 | 4.5746 | 2.6 | 0.3100 | 2.5 | 0.99 | 1749.2 | 6.3 | 99.5 |
| $\begin{aligned} & \text { OS081084A- } \\ & 81 \\ & \hline \end{aligned}$ | 267 | 33389 | 3.1 | 9.1015 | 0.4 | 4.7792 | 1.2 | 0.3155 | 1.1 | 0.94 | 1797.3 | 7.3 | 98.3 |
| $\begin{aligned} & \text { OS081084A- } \\ & 82 \\ & \hline \end{aligned}$ | 206 | 39643 | 2.1 | 6.7622 | 0.3 | 8.7455 | 0.8 | 0.4289 | 0.8 | 0.94 | 2321.5 | 4.9 | 99.1 |
| $\begin{aligned} & \hline \text { OS081084A- } \\ & 83 \\ & \hline \end{aligned}$ | 222 | 7560 | 2.0 | 9.3087 | 0.4 | 4.4304 | 3.1 | 0.2991 | 3.1 | 0.99 | 1756.2 | 7.1 | 96.1 |
| $\begin{aligned} & \text { OS081084A- } \\ & 84 \end{aligned}$ | 204 | 100006 | 1.4 | 9.3630 | 0.5 | 4.5396 | 1.2 | 0.3083 | 1.1 | 0.89 | 1745.5 | 9.9 | 99.2 |
| $\begin{aligned} & \text { OS081084A- } \\ & 86 \end{aligned}$ | 300 | 63370 | 1.7 | 7.9590 | 0.3 | 6.5532 | 1.1 | 0.3783 | 1.0 | 0.97 | 2037.9 | 4.5 | 101.5 |
| $\begin{aligned} & \text { OS081084A- } \\ & 87 \\ & \hline \end{aligned}$ | 503 | 10371 | 2.6 | 8.1935 | 0.3 | 5.9866 | 1.8 | 0.3558 | 1.8 | 0.98 | 1986.4 | 6.0 | 98.8 |
| $\begin{aligned} & \text { OS081084A- } \\ & 88 \end{aligned}$ | 162 | 140138 | 1.8 | 9.3293 | 1.0 | 4.6720 | 1.5 | 0.3161 | 1.1 | 0.76 | 1752.2 | 17.4 | 101.1 |
| $\begin{aligned} & \text { OS081084A- } \\ & 89 \\ & \hline \end{aligned}$ | 431 | 21393 | 1.9 | 9.3533 | 0.3 | 4.5920 | 1.2 | 0.3115 | 1.2 | 0.96 | 1747.5 | 5.8 | 100.0 |
| $\begin{aligned} & \text { OS081084A- } \\ & 90 \end{aligned}$ | 531 | 10537 | 4.1 | 6.4613 | 0.7 | 8.7020 | 5.3 | 0.4078 | 5.2 | 0.99 | 2399.3 | 12.7 | 91.9 |
| $\begin{aligned} & \text { OS081084A- } \\ & 91 \\ & \hline \end{aligned}$ | 315 | 54495 | 2.7 | 9.0721 | 0.6 | 4.9656 | 1.1 | 0.3267 | 0.9 | 0.85 | 1803.2 | 10.2 | 101.1 |
| $\begin{aligned} & \text { OS081084A- } \\ & 92 \\ & \hline \end{aligned}$ | 326 | 5703 | 1.9 | 9.3805 | 0.9 | 3.9758 | 4.5 | 0.2705 | 4.4 | 0.98 | 1742.1 | 16.9 | 88.6 |
| $\begin{aligned} & \text { OS081084A- } \\ & 93 \end{aligned}$ | 525 | 134423 | 3.5 | 9.1451 | 0.3 | 4.7804 | 1.2 | 0.3171 | 1.2 | 0.98 | 1788.6 | 4.8 | 99.3 |
| $\begin{aligned} & \text { OS081084A- } \\ & 94 \\ & \hline \end{aligned}$ | 263 | 72629 | 1.0 | 9.3851 | 0.3 | 4.6370 | 1.2 | 0.3156 | 1.1 | 0.95 | 1741.2 | 6.4 | 101.6 |
| $\begin{aligned} & \text { OS081084A- } \\ & 95 \end{aligned}$ | 185 | 297147 | 1.2 | 9.4041 | 0.4 | 4.4688 | 1.3 | 0.3048 | 1.2 | 0.94 | 1737.5 | 8.2 | 98.7 |
| $\begin{aligned} & \text { OS081084A- } \\ & 96 \end{aligned}$ | 232 | 18951 | 1.7 | 9.3451 | 0.8 | 4.4869 | 1.2 | 0.3041 | 1.0 | 0.78 | 1749.1 | 13.8 | 97.9 |
| $\begin{aligned} & \text { OS081084A- } \\ & 97 \\ & \hline \end{aligned}$ | 295 | 30392 | 2.0 | 8.9764 | 0.3 | 5.1940 | 1.4 | 0.3381 | 1.3 | 0.97 | 1822.4 | 5.8 | 103.0 |
| $\begin{aligned} & \text { OS081084A- } \\ & 98 \\ & \hline \end{aligned}$ | 103 | 44189 | 0.8 | 9.1745 | 0.5 | 4.6945 | 0.8 | 0.3124 | 0.7 | 0.79 | 1782.7 | 9.5 | 98.3 |
| $\begin{aligned} & \text { OS081084A- } \\ & 99 \\ & \hline \end{aligned}$ | 93 | 119123 | 2.5 | 5.4863 | 0.5 | 13.3492 | 1.2 | 0.5312 | 1.1 | 0.93 | 2673.6 | 7.6 | 102.7 |
| $\begin{aligned} & \text { OS081084A- } \\ & 100 \\ & \hline \end{aligned}$ | 171 | 18365 | 1.6 | 9.0039 | 0.6 | 5.2613 | 4.2 | 0.3436 | 4.2 | 0.99 | 1816.9 | 11.3 | 104.8 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 1 \end{aligned}$ | 168 | 257185 | 2.0 | 5.9584 | 0.6 | 11.4868 | 2.0 | 0.4964 | 2.0 | 0.96 | 2536.1 | 9.3 | 102.5 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 2 \\ & \hline \end{aligned}$ | 399 | 67404 | 1.4 | 3.6427 | 0.2 | 24.1800 | 2.6 | 0.6388 | 2.6 | 1.00 | 3332.0 | 2.9 | 95.6 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 3 \end{aligned}$ | 438 | 238848 | 1.7 | 5.9819 | 0.2 | 11.4510 | 3.2 | 0.4968 | 3.2 | 1.00 | 2529.5 | 3.8 | 102.8 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 4 \\ & \hline \end{aligned}$ | 241 | 892575 | 2.3 | 8.8699 | 0.9 | 5.1960 | 2.6 | 0.3343 | 2.5 | 0.94 | 1844.0 | 16.1 | 100.8 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 6 \\ & \hline \end{aligned}$ | 194 | 1193113 | 2.4 | 6.0115 | 0.3 | 10.9488 | 2.6 | 0.4774 | 2.5 | 0.99 | 2521.2 | 5.0 | 99.8 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 7 \\ & \hline \end{aligned}$ | 262 | 711570 | 1.5 | 5.9991 | 0.7 | 11.2364 | 4.1 | 0.4889 | 4.0 | 0.98 | 2524.7 | 12.3 | 101.6 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 8 \end{aligned}$ | 257 | 338859 | 2.1 | 9.0977 | 1.0 | 5.0859 | 2.0 | 0.3356 | 1.7 | 0.86 | 1798.0 | 18.6 | 103.7 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 9 \\ & \hline \end{aligned}$ | 334 | 61241 | 2.2 | 5.5104 | 0.7 | 12.8636 | 2.2 | 0.5141 | 2.1 | 0.95 | 2666.4 | 10.9 | 100.3 |


| $\begin{aligned} & \text { K05110-108A- } \\ & 10 \end{aligned}$ | 387 | 31954 | 2.0 | 8.8455 | 0.5 | 5.0493 | 2.5 | 0.3239 | 2.5 | 0.98 | 1849.0 | 9.8 | 97.8 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\begin{aligned} & \text { K05110-108A- } \\ & 11 \end{aligned}$ | 587 | 284552 | 1.4 | 5.7491 | 0.3 | 11.9706 | 2.5 | 0.4991 | 2.5 | 0.99 | 2595.9 | 4.8 | 100.5 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 13 \end{aligned}$ | 29 | 108843 | 0.6 | 7.6644 | 3.3 | 7.2844 | 4.3 | 0.4049 | 2.8 | 0.64 | 2104.3 | 58.4 | 104.2 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 14 \end{aligned}$ | 84 | 662704 | 1.2 | 5.2498 | 0.6 | 14.3455 | 3.5 | 0.5462 | 3.5 | 0.99 | 2746.3 | 9.1 | 102.3 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 15 \end{aligned}$ | 208 | 801164 | 0.8 | 9.1931 | 0.7 | 4.7877 | 3.5 | 0.3192 | 3.4 | 0.98 | 1779.0 | 12.7 | 100.4 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 16 \end{aligned}$ | 163 | 1533661 | 1.1 | 5.4920 | 1.1 | 12.7298 | 2.0 | 0.5070 | 1.7 | 0.84 | 2671.9 | 18.2 | 99.0 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 18 \end{aligned}$ | 824 | 635018 | 1.6 | 6.0530 | 0.2 | 11.3382 | 4.1 | 0.4978 | 4.1 | 1.00 | 2509.6 | 3.0 | 103.8 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 19 \\ & \hline \end{aligned}$ | 1031 | 28086 | 4.8 | 8.9880 | 2.5 | 5.0196 | 2.8 | 0.3272 | 1.2 | 0.43 | 1820.1 | 45.8 | 100.3 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 20 \end{aligned}$ | 264 | 53197 | 2.2 | 8.7958 | 1.3 | 5.4121 | 9.7 | 0.3453 | 9.6 | 0.99 | 1859.2 | 23.7 | 102.8 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 21 \end{aligned}$ | 1195 | 81433 | 4.4 | 9.0182 | 0.3 | 4.8365 | 1.6 | 0.3163 | 1.6 | 0.98 | 1814.0 | 5.0 | 97.7 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 22 \end{aligned}$ | 392 | 1504985 | 7.0 | 9.0040 | 0.4 | 5.2685 | 1.8 | 0.3441 | 1.8 | 0.97 | 1816.9 | 7.6 | 104.9 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 23 \\ & \hline \end{aligned}$ | 485 | 475551 | 3.0 | 7.9718 | 0.3 | 6.6905 | 3.2 | 0.3868 | 3.2 | 0.99 | 2035.0 | 5.9 | 103.6 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 24 \end{aligned}$ | 550 | 71622 | 4.9 | 6.0990 | 0.5 | 10.6767 | 2.2 | 0.4723 | 2.2 | 0.97 | 2496.9 | 9.2 | 99.9 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 25 \end{aligned}$ | 281 | 826822 | 1.1 | 9.2011 | 0.6 | 4.6207 | 1.6 | 0.3084 | 1.5 | 0.92 | 1777.4 | 11.5 | 97.5 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 26 \end{aligned}$ | 318 | 988727 | 3.0 | 8.9762 | 0.8 | 5.2366 | 2.5 | 0.3409 | 2.4 | 0.95 | 1822.5 | 13.8 | 103.8 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 27 \end{aligned}$ | 786 | 276100 | 2.0 | 5.6072 | 0.2 | 12.3762 | 1.1 | 0.5033 | 1.1 | 0.98 | 2637.5 | 3.7 | 99.6 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 28 \\ & \hline \end{aligned}$ | 587 | 26839 | 1.8 | 8.8125 | 1.1 | 4.7879 | 2.8 | 0.3060 | 2.6 | 0.92 | 1855.8 | 19.1 | 92.7 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 29 \\ & \hline \end{aligned}$ | 457 | 87070 | 2.5 | 3.8490 | 0.2 | 24.2911 | 2.0 | 0.6781 | 2.0 | 1.00 | 3245.5 | 2.9 | 102.8 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 30 \\ & \hline \end{aligned}$ | 259 | 730485 | 2.3 | 6.0422 | 0.5 | 10.7038 | 3.9 | 0.4691 | 3.8 | 0.99 | 2512.6 | 8.5 | 98.7 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 31 \end{aligned}$ | 290 | 53440 | 1.9 | 5.1784 | 0.7 | 13.6381 | 2.3 | 0.5122 | 2.2 | 0.96 | 2768.8 | 10.9 | 96.3 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 32 \end{aligned}$ | 700 | 23777 | 3.1 | 7.8324 | 1.5 | 5.9974 | 3.2 | 0.3407 | 2.9 | 0.89 | 2066.2 | 26.1 | 91.5 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 34 \end{aligned}$ | 154 | 2326533 | 1.5 | 5.8391 | 0.6 | 11.2903 | 1.3 | 0.4781 | 1.1 | 0.87 | 2570.0 | 10.5 | 98.0 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 35 \end{aligned}$ | 759 | 25343 | 2.3 | 6.0588 | 0.9 | 9.4128 | 3.0 | 0.4136 | 2.8 | 0.95 | 2508.0 | 15.5 | 89.0 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 36 \\ & \hline \end{aligned}$ | 272 | 88091 | 2.0 | 6.0731 | 0.5 | 11.3153 | 1.2 | 0.4984 | 1.1 | 0.90 | 2504.1 | 8.9 | 104.1 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 37 \\ & \hline \end{aligned}$ | 126 | 511506 | 2.1 | 6.0607 | 0.9 | 10.8773 | 1.5 | 0.4781 | 1.3 | 0.83 | 2507.5 | 14.6 | 100.5 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 38 \end{aligned}$ | 147 | 1654619 | 3.3 | 8.6082 | 0.8 | 5.4850 | 1.5 | 0.3424 | 1.3 | 0.85 | 1898.1 | 14.3 | 100.0 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 39 \end{aligned}$ | 190 | 10289 | 2.6 | 5.9306 | 0.6 | 8.8708 | 3.3 | 0.3816 | 3.2 | 0.98 | 2543.9 | 10.8 | 81.9 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 40 \end{aligned}$ | 371 | 96622 | 1.8 | 5.7357 | 0.4 | 10.5061 | 3.0 | 0.4370 | 3.0 | 0.99 | 2599.8 | 6.3 | 89.9 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 41 \end{aligned}$ | 150 | 1326951 | 1.7 | 9.0954 | 0.9 | 4.8694 | 3.5 | 0.3212 | 3.4 | 0.97 | 1798.5 | 15.8 | 99.8 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 42 \\ & \hline \end{aligned}$ | 306 | 469995 | 1.3 | 9.1870 | 0.5 | 4.7814 | 2.4 | 0.3186 | 2.3 | 0.98 | 1780.2 | 9.4 | 100.1 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 44 \\ & \hline \end{aligned}$ | 299 | 76811 | 1.6 | 8.5867 | 0.5 | 5.0521 | 1.7 | 0.3146 | 1.7 | 0.95 | 1902.6 | 9.3 | 92.7 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 45 \end{aligned}$ | 198 | 634807 | 0.7 | 5.3595 | 0.4 | 13.5196 | 1.9 | 0.5255 | 1.9 | 0.97 | 2712.3 | 7.3 | 100.4 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 46 \end{aligned}$ | 162 | 360676 | 1.9 | 8.6830 | 1.3 | 5.4353 | 4.1 | 0.3423 | 3.9 | 0.95 | 1882.5 | 24.2 | 100.8 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 47 \end{aligned}$ | 216 | 802479 | 2.5 | 8.9695 | 1.1 | 4.9632 | 3.5 | 0.3229 | 3.3 | 0.94 | 1823.8 | 20.7 | 98.9 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 48 \\ & \hline \end{aligned}$ | 257 | 650903 | 1.9 | 5.6514 | 0.7 | 12.6758 | 3.7 | 0.5196 | 3.6 | 0.98 | 2624.4 | 10.8 | 102.8 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 49 \\ & \hline \end{aligned}$ | 215 | 231324 | 1.7 | 8.2562 | 2.1 | 5.9414 | 4.1 | 0.3558 | 3.5 | 0.86 | 1972.8 | 36.7 | 99.5 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 50 \\ & \hline \end{aligned}$ | 489 | 72799 | 1.4 | 8.9982 | 0.8 | 4.8188 | 3.1 | 0.3145 | 3.0 | 0.97 | 1818.0 | 13.7 | 97.0 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 51 \\ & \hline \end{aligned}$ | 726 | 48555 | 2.2 | 5.9707 | 1.0 | 11.1221 | 2.8 | 0.4816 | 2.6 | 0.93 | 2532.6 | 17.4 | 100.1 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 52 \end{aligned}$ | 314 | 20983 | 2.8 | 8.4661 | 0.8 | 4.9657 | 2.4 | 0.3049 | 2.3 | 0.95 | 1927.9 | 13.7 | 89.0 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 53 \end{aligned}$ | 984 | 59575 | 8.9 | 8.8744 | 0.3 | 5.0126 | 1.7 | 0.3226 | 1.7 | 0.98 | 1843.1 | 5.9 | 97.8 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 55 \\ & \hline \end{aligned}$ | 485 | 46637 | 1.8 | 8.9599 | 0.3 | 5.2535 | 1.3 | 0.3414 | 1.2 | 0.97 | 1825.7 | 5.8 | 103.7 |


| $\begin{aligned} & \text { K05110-108A- } \\ & 56 \\ & \hline \end{aligned}$ | 1104 | 20370 | 2.2 | 5.9407 | 0.8 | 10.6466 | 1.9 | 0.4587 | 1.7 | 0.90 | 2541.1 | 13.3 | 95.8 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K05110-108A57 | 242 | 1295032 | 2.1 | 6.0000 | 0.2 | 11.2338 | 2.7 | 0.4889 | 2.7 | 1.00 | 2524.4 | 3.6 | 101.6 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 58 \end{aligned}$ | 330 | 1753021 | 1.5 | 6.0043 | 0.4 | 11.1871 | 3.1 | 0.4872 | 3.1 | 0.99 | 2523.2 | 6.8 | 101.4 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 59 \end{aligned}$ | 387 | 1180681 | 2.2 | 8.2662 | 0.4 | 5.9075 | 2.1 | 0.3542 | 2.0 | 0.98 | 1970.6 | 7.7 | 99.2 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 60 \end{aligned}$ | 305 | 817129 | 1.6 | 6.0119 | 0.6 | 10.3524 | 3.8 | 0.4514 | 3.8 | 0.99 | 2521.1 | 10.8 | 95.3 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 61 \end{aligned}$ | 411 | 346525 | 1.2 | 9.0811 | 0.6 | 4.8019 | 3.8 | 0.3163 | 3.8 | 0.99 | 1801.4 | 10.1 | 98.3 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 62 \\ & \hline \end{aligned}$ | 492 | 108764 | 1.5 | 9.0666 | 0.3 | 4.8574 | 1.6 | 0.3194 | 1.6 | 0.98 | 1804.3 | 6.3 | 99.0 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 63 \\ & \hline \end{aligned}$ | 339 | 1080927 | 1.9 | 9.0902 | 0.3 | 4.9456 | 2.1 | 0.3261 | 2.1 | 0.99 | 1799.5 | 6.3 | 101.1 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 64 \end{aligned}$ | 354 | 675299 | 0.9 | 9.1067 | 0.7 | 4.8303 | 1.3 | 0.3190 | 1.1 | 0.83 | 1796.2 | 12.7 | 99.4 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 65 \end{aligned}$ | 471 | 38246 | 1.9 | 8.4533 | 0.4 | 5.6523 | 2.5 | 0.3465 | 2.5 | 0.99 | 1930.6 | 7.5 | 99.3 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 66 \end{aligned}$ | 548 | 76722 | 1.9 | 8.0247 | 0.5 | 6.1309 | 2.1 | 0.3568 | 2.0 | 0.97 | 2023.3 | 9.2 | 97.2 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 67 \\ & \hline \end{aligned}$ | 350 | 58720 | 2.6 | 6.2580 | 1.0 | 10.1722 | 3.4 | 0.4617 | 3.2 | 0.96 | 2453.5 | 16.9 | 99.7 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 68 \end{aligned}$ | 986 | 38550 | 1.8 | 5.9404 | 0.5 | 10.9015 | 1.5 | 0.4697 | 1.4 | 0.93 | 2541.2 | 8.9 | 97.7 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 69 \\ & \hline \end{aligned}$ | 432 | 578678 | 4.1 | 8.8422 | 0.9 | 5.1157 | 2.1 | 0.3281 | 1.9 | 0.91 | 1849.7 | 15.9 | 98.9 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 71 \end{aligned}$ | 230 | 107743 | 1.9 | 6.0345 | 0.6 | 10.6477 | 1.3 | 0.4660 | 1.2 | 0.90 | 2514.8 | 9.4 | 98.1 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 72 \end{aligned}$ | 486 | 271964 | 2.1 | 6.0363 | 0.4 | 11.0754 | 2.0 | 0.4849 | 2.0 | 0.98 | 2514.3 | 6.0 | 101.4 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 73 \end{aligned}$ | 183 | 168211 | 2.4 | 6.0867 | 1.4 | 10.8628 | 2.1 | 0.4795 | 1.5 | 0.72 | 2500.3 | 24.1 | 101.0 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 74 \\ & \hline \end{aligned}$ | 127 | 692000 | 2.5 | 6.0035 | 0.8 | 10.4094 | 2.2 | 0.4532 | 2.0 | 0.92 | 2523.5 | 14.3 | 95.5 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 75 \\ & \hline \end{aligned}$ | 409 | 55680 | 2.2 | 8.5104 | 0.5 | 5.2793 | 3.0 | 0.3259 | 3.0 | 0.99 | 1918.6 | 9.2 | 94.8 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 76 \end{aligned}$ | 355 | 2038199 | 1.6 | 6.0226 | 0.2 | 11.2478 | 2.0 | 0.4913 | 2.0 | 0.99 | 2518.1 | 3.4 | 102.3 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 77 \end{aligned}$ | 1537 | 712999 | 26.2 | 8.3047 | 0.5 | 5.7029 | 1.7 | 0.3435 | 1.6 | 0.95 | 1962.3 | 9.8 | 97.0 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 79 \end{aligned}$ | 358 | 70449 | 1.4 | 8.9584 | 1.8 | 4.9263 | 2.6 | 0.3201 | 1.9 | 0.74 | 1826.1 | 31.8 | 98.0 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 80 \\ & \hline \end{aligned}$ | 632 | 56696 | 1.7 | 5.4103 | 0.5 | 12.1708 | 2.8 | 0.4776 | 2.8 | 0.99 | 2696.7 | 8.0 | 93.3 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 81 \\ & \hline \end{aligned}$ | 1211 | 36642 | 2.5 | 8.6428 | 2.0 | 5.0105 | 3.5 | 0.3141 | 2.9 | 0.82 | 1890.8 | 36.8 | 93.1 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 82 \\ & \hline \end{aligned}$ | 257 | 19759 | 0.8 | 8.0816 | 1.7 | 5.7119 | 3.1 | 0.3348 | 2.5 | 0.82 | 2010.8 | 30.9 | 92.6 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 83 \\ & \hline \end{aligned}$ | 297 | 131164 | 1.7 | 2.6337 | 1.3 | 33.3601 | 1.9 | 0.6372 | 1.4 | 0.74 | 3830.6 | 19.5 | 83.0 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 84 \\ & \hline \end{aligned}$ | 500 | 173338 | 1.4 | 5.8614 | 1.2 | 11.0266 | 1.7 | 0.4688 | 1.3 | 0.73 | 2563.6 | 19.7 | 96.7 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 85 \end{aligned}$ | 330 | 3788622 | 1.6 | 5.3489 | 0.4 | 14.0220 | 2.6 | 0.5440 | 2.5 | 0.99 | 2715.5 | 7.3 | 103.1 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 86 \\ & \hline \end{aligned}$ | 620 | 115076 | 1.9 | 7.1644 | 2.2 | 7.6376 | 2.6 | 0.3969 | 1.5 | 0.56 | 2222.0 | 37.8 | 97.0 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 87 \\ & \hline \end{aligned}$ | 249 | 833652 | 2.2 | 8.9619 | 0.5 | 5.0703 | 2.7 | 0.3296 | 2.6 | 0.98 | 1825.3 | 8.6 | 100.6 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 88 \end{aligned}$ | 1587 | 32319 | 2.9 | 8.0452 | 0.5 | 6.3489 | 3.2 | 0.3705 | 3.2 | 0.99 | 2018.8 | 8.2 | 100.6 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 89 \end{aligned}$ | 536 | 51521 | 2.6 | 8.9762 | 0.6 | 4.7938 | 3.8 | 0.3121 | 3.8 | 0.99 | 1822.5 | 11.8 | 96.1 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 91 \end{aligned}$ | 389 | 617773 | 1.8 | 9.1807 | 0.3 | 4.9873 | 1.8 | 0.3321 | 1.8 | 0.98 | 1781.5 | 5.9 | 103.8 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 92 \end{aligned}$ | 546 | 90614 | 2.0 | 8.5073 | 0.6 | 5.3428 | 2.9 | 0.3297 | 2.9 | 0.98 | 1919.2 | 10.5 | 95.7 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 93 \\ & \hline \end{aligned}$ | 179 | 654416 | 0.6 | 5.7095 | 0.3 | 12.1507 | 2.8 | 0.5032 | 2.8 | 0.99 | 2607.4 | 4.8 | 100.8 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 94 \\ & \hline \end{aligned}$ | 312 | 111982 | 0.8 | 5.8422 | 0.7 | 11.8822 | 1.3 | 0.5035 | 1.1 | 0.86 | 2569.1 | 11.1 | 102.3 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 95 \\ & \hline \end{aligned}$ | 561 | 241146 | 2.3 | 8.9934 | 0.4 | 4.9038 | 3.0 | 0.3199 | 3.0 | 0.99 | 1819.0 | 6.5 | 98.4 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 96 \end{aligned}$ | 258 | 476812 | 1.3 | 8.8785 | 1.0 | 5.4137 | 2.0 | 0.3486 | 1.7 | 0.87 | 1842.3 | 17.4 | 104.6 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 97 \end{aligned}$ | 204 | 35676 | 2.2 | 8.7130 | 0.8 | 5.4415 | 1.4 | 0.3439 | 1.2 | 0.84 | 1876.3 | 14.0 | 101.5 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 99 \end{aligned}$ | 518 | 82585 | 0.6 | 5.7505 | 0.8 | 11.2853 | 2.1 | 0.4707 | 1.9 | 0.92 | 2595.5 | 13.9 | 95.8 |
| $\begin{aligned} & \text { K05110-108A- } \\ & 100 \\ & \hline \end{aligned}$ | 204 | 238620 | 2.7 | 7.1309 | 1.9 | 7.7744 | 2.4 | 0.4021 | 1.4 | 0.58 | 2230.1 | 33.8 | 97.7 |


| K061121A-2 | 119 | 65946 | 0.9 | 5.8761 | 0.2 | 11.1631 | 1.3 | 0.4757 | 1.3 | 0.99 | 2559.4 | 3.1 | 98.0 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K061121A-3 | 153 | 318351 | 1.7 | 5.7528 | 0.1 | 11.6724 | 1.7 | 0.4870 | 1.7 | 1.00 | 2594.8 | 2.2 | 98.6 |
| K061121A-4 | 192 | 31071 | 1.2 | 3.7737 | 0.4 | 20.1908 | 4.3 | 0.5526 | 4.2 | 1.00 | 3276.6 | 6.3 | 86.6 |
| K061121A-6 | 94 | 445799 | 1.7 | 5.5514 | 0.4 | 12.5710 | 1.2 | 0.5061 | 1.1 | 0.95 | 2654.1 | 6.1 | 99.5 |
| K061121A-8 | 82 | 32060 | 1.2 | 8.8998 | 0.4 | 4.6009 | 3.0 | 0.2970 | 3.0 | 0.99 | 1838.0 | 7.4 | 91.2 |
| K061121A-9 | 137 | 15048 | 1.7 | 6.3002 | 0.7 | 8.8815 | 5.5 | 0.4058 | 5.5 | 0.99 | 2442.1 | 11.3 | 89.9 |
| K061121A-10 | 170 | 18652 | 1.9 | 8.9818 | 0.2 | 4.7284 | 3.5 | 0.3080 | 3.5 | 1.00 | 1821.3 | 4.0 | 95.0 |
| K061121A-13 | 79 | 109057 | 0.6 | 9.2117 | 0.6 | 4.7450 | 1.3 | 0.3170 | 1.1 | 0.89 | 1775.3 | 10.5 | 100.0 |
| K061121A-15 | 190 | 9898 | 1.0 | 6.1047 | 0.7 | 9.7762 | 8.1 | 0.4328 | 8.1 | 1.00 | 2495.3 | 11.1 | 92.9 |
| K061121A-16 | 155 | 68242 | 0.8 | 5.8807 | 0.3 | 11.2486 | 0.9 | 0.4798 | 0.9 | 0.95 | 2558.1 | 5.1 | 98.8 |
| K061121A-17 | 166 | 98306 | 1.6 | 6.1634 | 0.2 | 10.4686 | 0.7 | 0.4680 | 0.6 | 0.93 | 2479.2 | 4.2 | 99.8 |
| K061121A-18 | 88 | 183949 | 1.0 | 5.4909 | 0.4 | 12.6604 | 0.6 | 0.5042 | 0.4 | 0.68 | 2672.2 | 7.2 | 98.5 |
| K061121A-19 | 138 | 287293 | 1.6 | 8.8409 | 0.5 | 5.2733 | 0.8 | 0.3381 | 0.6 | 0.74 | 1850.0 | 9.2 | 101.5 |
| K061121A-20 | 119 | 351988 | 2.0 | 6.1598 | 0.3 | 10.6880 | 0.5 | 0.4775 | 0.4 | 0.76 | 2480.2 | 5.4 | 101.5 |
| K061121A-21 | 146 | 273112 | 1.8 | 6.1608 | 0.3 | 10.6714 | 0.8 | 0.4768 | 0.7 | 0.93 | 2479.9 | 4.6 | 101.4 |
| K061121A-22 | 122 | 62098 | 2.2 | 8.7815 | 0.5 | 5.0495 | 1.2 | 0.3216 | 1.1 | 0.92 | 1862.2 | 8.2 | 96.5 |
| K06246-1A-1 | 275 | 119362 | 2.3 | 9.4230 | 0.3 | 4.5007 | 2.8 | 0.3076 | 2.7 | 0.99 | 1733.9 | 5.7 | 99.7 |
| K06246-1A-2 | 734 | 5358 | 7.2 | 4.2634 | 1.2 | 14.4105 | 2.9 | 0.4456 | 2.7 | 0.92 | 3083.4 | 18.5 | 77.0 |
| K06246-1A-3 | 656 | 3087 | 2.3 | 9.5061 | 1.8 | 3.9540 | 3.7 | 0.2726 | 3.3 | 0.88 | 1717.7 | 32.9 | 90.5 |
| K06246-1A-5 | 140 | 116347 | 2.5 | 6.1753 | 0.2 | 10.4418 | 0.6 | 0.4677 | 0.6 | 0.93 | 2475.9 | 3.9 | 99.9 |
| K06246-1A-6 | 250 | 44193 | 2.6 | 9.3995 | 0.4 | 4.5276 | 0.8 | 0.3087 | 0.6 | 0.83 | 1738.4 | 7.9 | 99.8 |
| K06246-1A-7 | 236 | 261923 | 2.9 | 9.4087 | 0.4 | 4.4931 | 1.1 | 0.3066 | 1.0 | 0.95 | 1736.6 | 6.5 | 99.3 |
| K06246-1A-8 | 274 | 1409383 | 1.1 | 3.8727 | 0.1 | 23.2149 | 0.9 | 0.6520 | 0.9 | 0.99 | 3235.9 | 2.2 | 100.0 |
| K06246-1A-9 | 401 | 6160 | 1.2 | 5.7906 | 4.1 | 10.3891 | 10.3 | 0.4363 | 9.4 | 0.92 | 2583.9 | 68.0 | 90.3 |
| K06246-1A-10 | 235 | 141645 | 2.9 | 9.4209 | 0.5 | 4.4994 | 0.8 | 0.3074 | 0.6 | 0.77 | 1734.3 | 9.6 | 99.6 |
| K06246-1A-11 | 241 | 254904 | 2.9 | 9.4203 | 0.4 | 4.5568 | 1.0 | 0.3113 | 0.9 | 0.93 | 1734.4 | 6.7 | 100.7 |
| K06246-1A-12 | 270 | 59398 | 2.7 | 9.3871 | 0.3 | 4.5520 | 1.1 | 0.3099 | 1.0 | 0.97 | 1740.8 | 5.0 | 100.0 |
| K06246-1A-13 | 266 | 113569 | 2.9 | 9.4011 | 0.4 | 4.6878 | 4.4 | 0.3196 | 4.4 | 1.00 | 1738.1 | 7.0 | 102.9 |
| K06246-1A-14 | 343 | 200038 | 2.5 | 9.4329 | 0.4 | 4.4205 | 4.9 | 0.3024 | 4.9 | 1.00 | 1731.9 | 6.5 | 98.3 |
| K06246-1A-15 | 347 | 9130 | 2.6 | 9.3956 | 0.3 | 4.3109 | 1.4 | 0.2938 | 1.3 | 0.97 | 1739.2 | 6.2 | 95.5 |
| K06246-1A-17 | 544 | 22230 | 1.1 | 5.5022 | 4.0 | 10.5337 | 7.0 | 0.4204 | 5.7 | 0.82 | 2668.9 | 66.3 | 84.8 |
| K06246-1A-18 | 226 | 45261 | 2.7 | 9.4198 | 0.4 | 4.5677 | 0.5 | 0.3121 | 0.4 | 0.72 | 1734.5 | 6.6 | 100.9 |
| K06246-1A-19 | 131 | 134109 | 3.5 | 6.2273 | 0.4 | 10.5243 | 1.0 | 0.4753 | 0.9 | 0.89 | 2461.8 | 7.3 | 101.8 |
| K06246-1A-20 | 368 | 32427 | 2.3 | 9.4064 | 0.1 | 4.4932 | 0.6 | 0.3065 | 0.6 | 0.97 | 1737.1 | 2.7 | 99.2 |
| K06246-1A-21 | 280 | 279398 | 2.6 | 5.4001 | 0.3 | 13.4078 | 1.2 | 0.5251 | 1.2 | 0.98 | 2699.8 | 4.4 | 100.8 |
| K06246-1A-22 | 210 | 481861 | 2.2 | 5.1164 | 0.2 | 14.6815 | 0.9 | 0.5448 | 0.9 | 0.98 | 2788.6 | 2.9 | 100.5 |
| K06246-1A-23 | 192 | 220996 | 2.7 | 6.1750 | 0.3 | 10.5045 | 0.5 | 0.4704 | 0.5 | 0.87 | 2476.0 | 4.3 | 100.4 |
| K06246-1A-24 | 143 | 114394 | 3.0 | 6.1895 | 0.2 | 10.5498 | 0.8 | 0.4736 | 0.8 | 0.98 | 2472.1 | 2.5 | 101.1 |
| K06246-1A-26 | 112 | 164263 | 3.3 | 6.2056 | 0.4 | 10.6589 | 1.4 | 0.4797 | 1.3 | 0.95 | 2467.7 | 7.2 | 102.4 |
| K06246-1A-27 | 188 | 49554 | 0.7 | 5.8788 | 0.3 | 11.6006 | 1.1 | 0.4946 | 1.1 | 0.97 | 2558.6 | 4.9 | 101.3 |
| K06246-1A-28 | 265 | 130083 | 2.6 | 9.0944 | 0.4 | 4.9591 | 0.8 | 0.3271 | 0.7 | 0.88 | 1798.7 | 6.6 | 101.4 |


| K06246-1A-29 | 958 | 4819 | 1.6 | 8.7869 | 0.3 | 4.3855 | 2.8 | 0.2795 | 2.7 | 0.99 | 1861.0 | 5.9 | 85.4 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K06246-1A-30 | 113 | 21152 | 1.5 | 8.9643 | 0.8 | 5.1216 | 1.5 | 0.3330 | 1.3 | 0.85 | 1824.9 | 14.0 | 101.5 |
| K06246-1A-31 | 466 | 6470 | 1.0 | 5.6217 | 0.5 | 11.8205 | 1.9 | 0.4819 | 1.8 | 0.97 | 2633.2 | 7.7 | 96.3 |
| K06246-1A-32 | 258 | 102984 | 2.6 | 9.4107 | 0.4 | 4.5576 | 0.8 | 0.3111 | 0.7 | 0.90 | 1736.2 | 6.5 | 100.6 |
| K06246-1A-33 | 279 | 29192 | 1.9 | 8.5900 | 0.4 | 5.5428 | 0.7 | 0.3453 | 0.6 | 0.87 | 1901.9 | 6.5 | 100.5 |
| K06246-1A-34 | 306 | 50100 | 2.6 | 9.4033 | 0.3 | 4.6688 | 0.7 | 0.3184 | 0.6 | 0.90 | 1737.7 | 5.4 | 102.5 |
| K06246-1A-35 | 208 | 61163 | 2.1 | 9.3828 | 0.4 | 4.5000 | 0.8 | 0.3062 | 0.7 | 0.86 | 1741.7 | 7.2 | 98.9 |
| K06246-1A-36 | 260 | 4337 | 3.0 | 8.9443 | 7.6 | 4.8755 | 10.1 | 0.3163 | 6.7 | 0.66 | 1828.9 | 137.3 | 96.9 |
| K06246-1A-37 | 134 | 81178 | 2.9 | 6.1882 | 0.6 | 10.5227 | 3.2 | 0.4723 | 3.2 | 0.98 | 2472.4 | 10.5 | 100.9 |
| K06246-1A-38 | 156 | 318372 | 2.3 | 6.1640 | 0.3 | 10.5057 | 1.1 | 0.4697 | 1.0 | 0.96 | 2479.0 | 5.2 | 100.1 |
| K06246-1A-39 | 263 | 69938 | 2.5 | 9.3834 | 0.2 | 4.5262 | 0.7 | 0.3080 | 0.7 | 0.96 | 1741.6 | 3.7 | 99.4 |
| K06246-1A-41 | 517 | 32318 | 4.5 | 9.5294 | 0.3 | 4.3428 | 2.0 | 0.3001 | 2.0 | 0.99 | 1713.2 | 5.8 | 98.8 |
| K06246-1A-42 | 260 | 167042 | 2.6 | 9.3893 | 0.4 | 4.5711 | 0.8 | 0.3113 | 0.7 | 0.88 | 1740.4 | 6.7 | 100.4 |
| K06246-1A-43 | 362 | 6899 | 0.7 | 8.8773 | 0.4 | 5.2079 | 2.0 | 0.3353 | 2.0 | 0.98 | 1842.5 | 7.8 | 101.2 |
| K06246-1A-44 | 453 | 13128 | 1.7 | 9.3854 | 0.4 | 4.4061 | 2.3 | 0.2999 | 2.3 | 0.98 | 1741.2 | 8.0 | 97.1 |
| K06246-1A-46 | 417 | 60996 | 1.8 | 9.4041 | 0.2 | 4.5466 | 1.3 | 0.3101 | 1.3 | 0.99 | 1737.5 | 3.0 | 100.2 |
| K06246-1A-47 | 399 | 4876 | 4.1 | 9.4071 | 0.7 | 4.1246 | 2.7 | 0.2814 | 2.6 | 0.96 | 1736.9 | 13.5 | 92.0 |
| K06246-1A-49 | 250 | 13048 | 2.6 | 9.3921 | 0.4 | 4.6460 | 1.3 | 0.3165 | 1.3 | 0.96 | 1739.9 | 6.6 | 101.9 |
| K06246-1A-50 | 344 | 5910 | 2.3 | 9.3665 | 0.6 | 4.5102 | 1.1 | 0.3064 | 0.9 | 0.82 | 1744.9 | 11.9 | 98.7 |
| K06246-1A-51 | 286 | 31916 | 2.4 | 9.4270 | 0.3 | 4.5991 | 0.6 | 0.3144 | 0.6 | 0.89 | 1733.1 | 5.3 | 101.7 |
| K06246-1A-52 | 296 | 35160 | 2.5 | 6.3544 | 0.2 | 9.6824 | 1.0 | 0.4462 | 1.0 | 0.98 | 2427.6 | 3.5 | 98.0 |
| K06246-1A-53 | 138 | 54169 | 0.6 | 5.5137 | 0.3 | 12.8861 | 2.7 | 0.5153 | 2.7 | 0.99 | 2665.4 | 4.5 | 100.5 |
| K06246-1A-54 | 361 | 7184 | 2.5 | 9.4517 | 0.5 | 4.3502 | 1.3 | 0.2982 | 1.2 | 0.92 | 1728.3 | 9.0 | 97.3 |
| K06246-1A-55 | 432 | 3846 | 3.7 | 9.4696 | 1.6 | 4.4097 | 2.9 | 0.3029 | 2.5 | 0.84 | 1724.8 | 29.0 | 98.9 |
| K06246-1A-56 | 398 | 4298 | 2.2 | 9.4120 | 0.6 | 4.2341 | 1.0 | 0.2890 | 0.8 | 0.82 | 1736.0 | 10.5 | 94.3 |
| K06246-1A-57 | 463 | 6672 | 2.2 | 9.4143 | 0.8 | 4.5491 | 5.2 | 0.3106 | 5.2 | 0.99 | 1735.5 | 15.4 | 100.5 |
| K06246-1A-59 | 257 | 13589 | 2.7 | 9.3844 | 0.4 | 4.6033 | 1.2 | 0.3133 | 1.1 | 0.93 | 1741.4 | 7.7 | 100.9 |
| K06246-1A-60 | 440 | 34335 | 2.1 | 9.3863 | 0.4 | 4.5240 | 2.2 | 0.3080 | 2.1 | 0.98 | 1741.0 | 7.4 | 99.4 |
| K06246-1A-61 | 536 | 5137 | 1.7 | 9.3885 | 0.7 | 4.5365 | 1.6 | 0.3089 | 1.5 | 0.91 | 1740.6 | 12.4 | 99.7 |
| K06246-1A-63 | 273 | 14812 | 1.3 | 5.8755 | 1.2 | 11.0348 | 3.4 | 0.4702 | 3.2 | 0.93 | 2559.6 | 20.1 | 97.1 |
| K06246-1A-64 | 418 | 14003 | 1.9 | 9.4074 | 0.4 | 4.3779 | 0.9 | 0.2987 | 0.8 | 0.91 | 1736.9 | 6.9 | 97.0 |
| K06246-1A-65 | 289 | 11794 | 2.6 | 9.3989 | 0.5 | 4.4842 | 1.6 | 0.3057 | 1.5 | 0.95 | 1738.5 | 9.5 | 98.9 |
| K06246-1A-66 | 259 | 71530 | 2.5 | 9.4184 | 0.4 | 4.5656 | 0.9 | 0.3119 | 0.9 | 0.91 | 1734.7 | 7.1 | 100.9 |
| K06246-1A-67 | 218 | 32459 | 2.8 | 9.4123 | 0.3 | 4.6215 | 1.3 | 0.3155 | 1.3 | 0.98 | 1735.9 | 5.2 | 101.8 |
| K06246-1A-69 | 286 | 168423 | 2.6 | 9.4324 | 0.3 | 4.4996 | 1.6 | 0.3078 | 1.5 | 0.98 | 1732.0 | 6.3 | 99.9 |
| K06246-1A-70 | 530 | 18323 | 1.6 | 9.4227 | 0.3 | 4.5031 | 1.7 | 0.3077 | 1.6 | 0.99 | 1733.9 | 4.9 | 99.7 |
| K06246-1A-71 | 284 | 2823 | 2.9 | 9.3712 | 0.2 | 4.0413 | 1.0 | 0.2747 | 1.0 | 0.97 | 1744.0 | 4.6 | 89.7 |
| K06246-1A-74 | 336 | 7328 | 1.3 | 9.4550 | 0.7 | 4.3185 | 1.7 | 0.2961 | 1.5 | 0.90 | 1727.6 | 13.2 | 96.8 |
| K06246-1A-75 | 248 | 31795 | 2.9 | 9.4096 | 0.5 | 4.5728 | 1.5 | 0.3121 | 1.4 | 0.94 | 1736.4 | 9.3 | 100.8 |
| K06246-1A-77 | 353 | 38988 | 2.2 | 9.4231 | 0.3 | 4.5343 | 1.1 | 0.3099 | 1.1 | 0.96 | 1733.8 | 5.6 | 100.4 |
| K06246-1A-78 | 242 | 90372 | 3.3 | 9.4217 | 0.5 | 4.5099 | 2.2 | 0.3082 | 2.2 | 0.97 | 1734.1 | 9.2 | 99.9 |


| K06246-1A-79 | 110 | 3159 | 1.7 | 6.1288 | 0.5 | 8.9566 | 1.2 | 0.3981 | 1.1 | 0.90 | 2488.7 | 9.2 | 86.8 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K06246-1A-81 | 253 | 46135 | 2.3 | 9.4291 | 0.4 | 4.5357 | 1.8 | 0.3102 | 1.7 | 0.98 | 1732.7 | 7.1 | 100.5 |
| K06246-1A-82 | 388 | 6648 | 2.3 | 9.3578 | 0.5 | 4.4239 | 1.2 | 0.3002 | 1.1 | 0.92 | 1746.6 | 8.5 | 96.9 |
| K06246-1A-83 | 323 | 19485 | 2.5 | 9.4036 | 0.6 | 4.5415 | 1.2 | 0.3097 | 1.0 | 0.85 | 1737.6 | 11.6 | 100.1 |
| K06246-1A-86 | 342 | 158585 | 2.4 | 9.3931 | 0.2 | 4.6764 | 0.9 | 0.3186 | 0.9 | 0.98 | 1739.7 | 3.7 | 102.5 |
| K06246-1A-88 | 365 | 70655 | 2.1 | 9.4020 | 0.2 | 4.5157 | 0.8 | 0.3079 | 0.7 | 0.96 | 1737.9 | 4.1 | 99.6 |
| K06246-1A-89 | 391 | 15711 | 2.5 | 9.4135 | 0.3 | 4.3955 | 1.0 | 0.3001 | 0.9 | 0.95 | 1735.7 | 5.9 | 97.5 |
| K06246-1A-90 | 323 | 4852 | 2.0 | 9.3729 | 0.8 | 4.1890 | 1.2 | 0.2848 | 0.9 | 0.77 | 1743.6 | 14.2 | 92.6 |
| K06246-1A-91 | 195 | 75545 | 2.5 | 9.3936 | 0.7 | 4.6267 | 1.0 | 0.3152 | 0.7 | 0.70 | 1739.6 | 12.8 | 101.5 |
| K06246-1A-92 | 330 | 184751 | 1.9 | 9.3925 | 0.4 | 4.6771 | 0.8 | 0.3186 | 0.7 | 0.84 | 1739.8 | 7.9 | 102.5 |
| K06246-1A-93 | 487 | 205451 | 2.0 | 9.3878 | 0.3 | 4.5375 | 1.0 | 0.3089 | 1.0 | 0.96 | 1740.7 | 5.0 | 99.7 |
| K06246-1A-94 | 246 | 121810 | 3.3 | 9.3935 | 0.5 | 4.5305 | 0.8 | 0.3087 | 0.7 | 0.82 | 1739.6 | 8.5 | 99.7 |
| K06246-1A-95 | 373 | 23183 | 2.6 | 8.7657 | 0.7 | 5.1765 | 1.2 | 0.3291 | 1.0 | 0.84 | 1865.4 | 12.2 | 98.3 |
| K06246-1A-96 | 330 | 13331 | 2.3 | 9.3982 | 0.5 | 4.4245 | 0.9 | 0.3016 | 0.7 | 0.86 | 1738.7 | 8.3 | 97.7 |
| K06246-1A-97 | 536 | 11859 | 1.0 | 5.2180 | 0.4 | 12.9523 | 2.3 | 0.4902 | 2.3 | 0.99 | 2756.3 | 6.1 | 93.3 |
| K06246-1A-98 | 245 | 499348 | 2.6 | 9.4188 | 0.2 | 4.4880 | 0.9 | 0.3066 | 0.9 | 0.97 | 1734.7 | 4.1 | 99.4 |
| $\begin{aligned} & \hline \text { K06246-1A- } \\ & 100 \end{aligned}$ | 294 | 5392 | 1.9 | 9.3902 | 0.7 | 4.6344 | 2.3 | 0.3156 | 2.1 | 0.95 | 1740.2 | 13.5 | 101.6 |

Notes:
Analyses with $>10 \%$ uncertainty (1-sigma) in $206 \mathrm{~Pb} / 238 \mathrm{U}$ age are not included.
Analyses with $>10 \%$ uncertainty (1-sigma) in $206 \mathrm{~Pb} / 207 \mathrm{~Pb}$ age are not included, unless $206 \mathrm{~Pb} / 238 \mathrm{U}$ age is $<500 \mathrm{Ma}$.
Best age is determined from 206Pb/238U age for analyses with 206Pb/238U age < 900 Ma and from 206Pb/207Pb age for analyses with 206Pb/238Uage > 900 Ma .
Concordance is based on 206Pb/238U age / 206Pb/207Pb age. Value is not reported for $206 \mathrm{~Pb} / 238 \mathrm{U}$ ages $<500 \mathrm{Ma}$ because of large uncertainty in 206Pb/207Pb age.
Analyses with $206 \mathrm{~Pb} / 238 \mathrm{U}$ age > 500 Ma and with >20\% discordance (<80\% concordance) are not included.
Analyses with 206Pb/238U age > 500 Ma and with >5\% reverse discordance (<105\% concordance) are not included.
All uncertainties are reported at the 1-sigma level, and include only measurement errors.
Systematic errors are shown as 206Pb/238U uncertainty, 206Pb/207Pb uncertainty to the right of each sample (at 2-sigma level).
$U$ concentration and U/Th are calibrated relative to Sri Lanka zircon and are accurate to $\sim 20 \%$. Common Pb correction is from 204 Pb , with composition interpreted from Stacey and Kramers (1975).

Uncertainties of 1.5 for $206 \mathrm{~Pb} / 204 \mathrm{~Pb}, 0.3$ for $207 \mathrm{~Pb} / 204 \mathrm{~Pb}$, and 2.0 for $208 \mathrm{~Pb} / 204 \mathrm{~Pb}$ are applied to common Pb composition.
$\mathrm{U} / \mathrm{Pb}$ and $206 \mathrm{~Pb} / 207 \mathrm{~Pb}$ fractionation is calibrated relative to fragments of a large Sri Lanka zircon of $563.5 \pm 3.2 \mathrm{Ma}$ (2-sigma).
$U$ decay constants and composition as follows: $238 \mathrm{U}=9.8485 \times 10-10,235 \mathrm{U}=1.55125 \times 10-10$, $238 \mathrm{U} / 235 \mathrm{U}=137.88$
Analytical methods as described by Gehrels et al. (2008).

Table A 2: Hf isotopic data for Vishnu Schist (ALC)

| Sample | ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}$ | $\pm$ (1s) | ${ }^{176} \mathrm{Lu}{ }^{177} \mathrm{Hf}$ | ${ }^{1 / 6} \mathrm{Hf} /^{1 / 7} \mathrm{Hf}$ <br> (T) | $\begin{gathered} \text { E-Hf } \\ (0) \\ \hline \end{gathered}$ | $\begin{gathered} \mathrm{E}-\mathrm{Hf}(0) \pm \\ (1 \mathrm{~s}) \end{gathered}$ | $\begin{gathered} \text { E-Hf } \\ (\mathrm{T}) \\ \hline \end{gathered}$ | $\begin{aligned} & \hline \mathrm{Age} \\ & \text { (Ma) } \\ & \hline \end{aligned}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| OS0884-2-1 | 0.281644 | 0.000035 | 0.001396 | 0.281594 | -40.4 | 1.2 | -0.3 | 1853 |
| OS0884-2-2 | 0.281638 | 0.000037 | 0.001754 | 0.281578 | -40.6 | 1.3 | -2.1 | 1800 |
| OS0884-2-3 | 0.280834 | 0.000034 | 0.000555 | 0.280806 | -69.0 | 1.2 | -9.5 | 2661 |
| OS0884-2-4 | 0.281022 | 0.000040 | 0.000912 | 0.280975 | -62.4 | 1.4 | -3.3 | 2671 |
| OS0884-2-6 | 0.281196 | 0.000041 | 0.000462 | 0.281173 | -56.2 | 1.5 | 2.2 | 2606 |
| OS0884-2-8 | 0.281217 | 0.000032 | 0.000690 | 0.281184 | -55.5 | 1.1 | -0.2 | 2484 |
| OS0884-2-9 | 0.281457 | 0.000039 | 0.000629 | 0.281436 | -47.0 | 1.4 | -7.5 | 1783 |
| OS0884-2-17 | 0.281145 | 0.000036 | 0.000649 | 0.281114 | -58.0 | 1.3 | -2.7 | 2485 |
| OS0884-2-18 | 0.281590 | 0.000028 | 0.000998 | 0.281554 | -42.3 | 1.0 | -1.5 | 1862 |
| OS0884-2-22 | 0.281581 | 0.000025 | 0.001013 | 0.281546 | -42.6 | 0.9 | -2.8 | 1817 |
| OS0884-2-27 | 0.281445 | 0.000036 | 0.001316 | 0.281399 | -47.4 | 1.3 | -7.4 | 1846 |
| OS0884-2-30 | 0.281632 | 0.000034 | 0.000931 | 0.281601 | -40.8 | 1.2 | -1.5 | 1790 |
| OS0884-2-32 | 0.281924 | 0.000050 | 0.001155 | 0.281885 | -30.4 | 1.8 | 8.2 | 1774 |
| OS0884-2-38 | 0.281275 | 0.000042 | 0.000919 | 0.281230 | -53.4 | 1.5 | 2.7 | 2541 |
| OS0884-2-40 | 0.281711 | 0.000033 | 0.001944 | 0.281645 | -38.0 | 1.2 | 0.0 | 1786 |
| OS0884-2-41 | 0.281318 | 0.000048 | 0.001904 | 0.281252 | -51.9 | 1.7 | -12.9 | 1831 |
| OS0884-2-50 | 0.281193 | 0.000050 | 0.000866 | 0.281151 | -56.3 | 1.8 | 0.5 | 2563 |
| OS0884-2-100 | 0.281648 | 0.000038 | 0.000832 | 0.281620 | -40.2 | 1.3 | -1.1 | 1777 |
| OS0884-2-53 | 0.281667 | 0.000044 | 0.000856 | 0.281636 | -39.5 | 1.5 | 1.5 | 1865 |
| OS0884-2-55 | 0.281189 | 0.000048 | 0.001586 | 0.281114 | -56.4 | 1.7 | -2.9 | 2478 |
| OS0884-2-56 | 0.281209 | 0.000038 | 0.000969 | 0.281160 | -55.7 | 1.3 | 2.2 | 2625 |
| OS0884-2-57 | 0.281688 | 0.000042 | 0.000911 | 0.281657 | -38.8 | 1.5 | 0.7 | 1801 |
| OS0884-2-58 | 0.281659 | 0.000041 | 0.001052 | 0.281623 | -39.8 | 1.5 | -0.7 | 1792 |
| OS0884-2-59 | 0.281229 | 0.000035 | 0.001033 | 0.281178 | -55.0 | 1.2 | 1.8 | 2577 |
| OS0884-2-60 | 0.281212 | 0.000033 | 0.000863 | 0.281171 | -55.6 | 1.2 | -0.9 | 2474 |
| OS0884-2-61 | 0.281122 | 0.000037 | 0.000694 | 0.281090 | -58.8 | 1.3 | -4.5 | 2445 |
| OS0884-2-62 | 0.281199 | 0.000035 | 0.000663 | 0.281167 | -56.1 | 1.2 | -0.8 | 2483 |
| OS0884-2-66 | 0.281223 | 0.000045 | 0.001134 | 0.281170 | -55.2 | 1.6 | -2.2 | 2419 |
| OS0884-2-67 | 0.281778 | 0.000043 | 0.001862 | 0.281716 | -35.6 | 1.5 | 2.0 | 1763 |
| OS0884-2-69 | 0.281746 | 0.000044 | 0.001003 | 0.281712 | -36.7 | 1.5 | 2.5 | 1794 |
| OS0884-2-70 | 0.281718 | 0.000040 | 0.000545 | 0.281698 | -37.7 | 1.4 | 3.5 | 1855 |
| OS0884-2-71 | 0.281747 | 0.000036 | 0.001348 | 0.281701 | -36.7 | 1.3 | 2.1 | 1790 |
| OS0884-2-72 | 0.281215 | 0.000044 | 0.001143 | 0.281161 | -55.5 | 1.5 | -1.0 | 2483 |
| OS0884-2-74 | 0.281267 | 0.000036 | 0.000660 | 0.281236 | -53.7 | 1.3 | 1.7 | 2485 |
| OS0884-2-75 | 0.281467 | 0.000039 | 0.001468 | 0.281403 | -46.6 | 1.4 | 3.5 | 2309 |
| OS0884-2-76 | 0.281581 | 0.000038 | 0.000555 | 0.281562 | -42.6 | 1.4 | -3.0 | 1784 |
| OS0884-2-78 | 0.281322 | 0.000030 | 0.000843 | 0.281282 | -51.7 | 1.1 | 2.7 | 2458 |
| OS0884-2-79 | 0.281912 | 0.000032 | 0.001573 | 0.281860 | -30.9 | 1.1 | 6.6 | 1744 |
| OS0884-2-80 | 0.281002 | 0.000035 | 0.001299 | 0.280935 | -63.1 | 1.3 | -4.1 | 2699 |
| OS0884-2-83 | 0.281705 | 0.000093 | 0.001011 | 0.281671 | -38.2 | 3.3 | 0.7 | 1778 |
| OS0884-2-88 | 0.281883 | 0.000032 | 0.001114 | 0.281846 | -31.9 | 1.1 | 6.3 | 1751 |
| OS0884-2-90 | 0.281736 | 0.000056 | 0.001254 | 0.281695 | -37.1 | 2.0 | 0.9 | 1748 |
| OS0884-2-92 | 0.281157 | 0.000045 | 0.001108 | 0.281101 | -57.6 | 1.6 | 0.3 | 2635 |
| OS0884-2-93 | 0.281787 | 0.000043 | 0.001753 | 0.281728 | -35.3 | 1.5 | 2.6 | 1770 |
| OS0884-2-95 | 0.281753 | 0.000061 | 0.002621 | 0.281665 | -36.5 | 2.2 | 0.4 | 1773 |
| OS0884-2-96 | 0.281571 | 0.000037 | 0.000672 | 0.281549 | -42.9 | 1.3 | -3.6 | 1779 |
| OS0884-2-97 | 0.281017 | 0.000050 | 0.000993 | 0.280964 | -62.5 | 1.8 | -1.4 | 2765 |
| OS0884-2-99 | 0.281191 | 0.000049 | 0.001772 | 0.281100 | -56.4 | 1.7 | 1.1 | 2668 |
| K0598.6-104-1 | 0.281637 | 0.000040 | 0.003100 | 0.281531 | -40.6 | 1.4 | -3.7 | 1802 |
| K0598.6-104-5 | 0.281757 | 0.000035 | 0.000834 | 0.281730 | -36.3 | 1.2 | 2.0 | 1742 |
| K0598.6-104-9 | 0.281704 | 0.000041 | 0.002097 | 0.281634 | -38.2 | 1.5 | -1.3 | 1747 |


| K0598.6-104-10 | 0.281722 | 0.000065 | 0.001356 | 0.281675 | -37.6 | 2.3 | 1.5 | 1805 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K0598.6-104-17 | 0.281815 | 0.000038 | 0.000770 | 0.281790 | -34.3 | 1.4 | 4.1 | 1744 |
| K0598.6-104-20 | 0.281821 | 0.000050 | 0.001304 | 0.281776 | -34.1 | 1.8 | 5.7 | 1832 |
| K0598.6-104-26 | 0.281657 | 0.000024 | 0.001065 | 0.281620 | -39.9 | 0.8 | -0.2 | 1819 |
| K0598.6-104-28 | 0.281749 | 0.000035 | 0.001700 | 0.281692 | -36.6 | 1.2 | 1.0 | 1759 |
| K0598.6-104-34 | 0.281779 | 0.000046 | 0.002534 | 0.281695 | -35.6 | 1.6 | 0.8 | 1744 |
| K0598.6-104-35 | 0.281744 | 0.000042 | 0.001256 | 0.281702 | -36.8 | 1.5 | 1.2 | 1750 |
| K0598.6-104-36 | 0.281813 | 0.000034 | 0.000764 | 0.281786 | -34.4 | 1.2 | 6.3 | 1842 |
| K0598.6-104-37 | 0.281912 | 0.000046 | 0.001247 | 0.281871 | -30.9 | 1.6 | 7.0 | 1744 |
| K0598.6-104-41 | 0.281808 | 0.000036 | 0.000984 | 0.281775 | -34.6 | 1.3 | 3.8 | 1754 |
| K0598.6-104-47 | 0.281753 | 0.000039 | 0.001459 | 0.281704 | -36.5 | 1.4 | 1.1 | 1743 |
| K0598.6-104-49 | 0.281558 | 0.000037 | 0.001014 | 0.281523 | -43.4 | 1.3 | -3.9 | 1806 |
| K0598.6-104-50 | 0.281773 | 0.000037 | 0.001261 | 0.281731 | -35.8 | 1.3 | 2.2 | 1750 |
| OS08108-4-2 | 0.281777 | 0.000049 | 0.001782 | 0.281718 | -35.6 | 1.7 | 1.6 | 1743 |
| OS08108-4-3 | 0.281626 | 0.000057 | 0.000798 | 0.281599 | -41.0 | 2.0 | -1.7 | 1783 |
| OS08108-4-4 | 0.281234 | 0.000048 | 0.000996 | 0.281184 | -54.8 | 1.7 | 3.2 | 2630 |
| OS08108-4-6 | 0.281808 | 0.000041 | 0.001627 | 0.281754 | -34.5 | 1.5 | 2.9 | 1744 |
| OS08108-4-7 | 0.281173 | 0.000045 | 0.000926 | 0.281130 | -57.0 | 1.6 | -2.3 | 2476 |
| OS08108-4-8 | 0.281789 | 0.000042 | 0.001315 | 0.281745 | -35.2 | 1.5 | 2.6 | 1744 |
| OS08108-4-11 | 0.281661 | 0.000048 | 0.001552 | 0.281609 | -39.8 | 1.7 | -2.2 | 1746 |
| OS08108-4-13 | 0.281702 | 0.000030 | 0.001009 | 0.281668 | -38.3 | 1.1 | 0.1 | 1757 |
| OS08108-4-17 | 0.281751 | 0.000033 | 0.001248 | 0.281709 | -36.6 | 1.2 | 1.9 | 1771 |
| OS08108-4-21 | 0.281740 | 0.000050 | 0.001247 | 0.281698 | -37.0 | 1.8 | 1.0 | 1750 |
| OS08108-4-27 | 0.281738 | 0.000037 | 0.001421 | 0.281691 | -37.0 | 1.3 | 0.6 | 1741 |
| OS08108-4-28 | 0.281803 | 0.000040 | 0.001957 | 0.281738 | -34.7 | 1.4 | 2.3 | 1744 |
| OS08108-4-30 | 0.281887 | 0.000036 | 0.001497 | 0.281837 | -31.8 | 1.3 | 5.8 | 1740 |
| OS08108-4-32 | 0.281736 | 0.000029 | 0.001010 | 0.281703 | -37.1 | 1.0 | 1.0 | 1744 |
| OS081084-1-39 | 0.281775 | 0.000041 | 0.001547 | 0.281722 | -35.7 | 1.4 | 2.9 | 1793 |
| OS081084-1-40 | 0.281437 | 0.000072 | 0.001242 | 0.281377 | -47.7 | 2.6 | 7.0 | 2499 |
| OS081084-1-43 | 0.281210 | 0.000041 | 0.000842 | 0.281168 | -55.7 | 1.4 | 1.3 | 2572 |
| OS081084-1-44 | 0.281308 | 0.000094 | 0.002268 | 0.281201 | -52.2 | 3.3 | 0.1 | 2473 |
| OS081084-1-47 | 0.281843 | 0.000044 | 0.001595 | 0.281790 | -33.3 | 1.6 | 4.2 | 1746 |
| OS081084-1-48 | 0.281756 | 0.000045 | 0.001497 | 0.281706 | -36.4 | 1.6 | 1.4 | 1753 |
| OS081084-1-57 | 0.281702 | 0.000034 | 0.002855 | 0.281603 | -38.3 | 1.2 | -0.8 | 1816 |
| K051108-108A-2 | 0.280736 | 0.000053 | 0.000992 | 0.280673 | -72.4 | 1.9 | 1.6 | 3332 |
| K051108-108A-3 | 0.280989 | 0.000058 | 0.001797 | 0.280903 | -63.5 | 2.0 | -9.2 | 2529 |
| K051108-108A-6 | 0.281140 | 0.000046 | 0.000575 | 0.281112 | -58.2 | 1.6 | -1.9 | 2521 |
| K051108-108A-7 | 0.281376 | 0.000016 | 0.001563 | 0.281300 | -49.8 | 0.6 | 4.9 | 2525 |
| K051108-108A-10 | 0.281930 | 0.000047 | 0.001207 | 0.281888 | -30.2 | 1.7 | 10.0 | 1849 |
| K051108-108A-11 | 0.280942 | 0.000056 | 0.000827 | 0.280901 | -65.2 | 2.0 | -7.7 | 2596 |
| K051108-108A-15 | 0.281929 | 0.000056 | 0.002623 | 0.281840 | -30.3 | 2.0 | 6.7 | 1779 |
| K051108-108A-21 | 0.281750 | 0.000043 | 0.001250 | 0.281707 | -36.6 | 1.5 | 2.8 | 1814 |
| K051108-108A-28 | 0.281665 | 0.000039 | 0.002695 | 0.281570 | -39.6 | 1.4 | -1.1 | 1856 |
| K051108-108A-29 | 0.280596 | 0.000036 | 0.001104 | 0.280527 | -77.4 | 1.3 | -5.7 | 3246 |
| K051108-108A-30 | 0.281386 | 0.000037 | 0.000900 | 0.281343 | -49.5 | 1.3 | 6.1 | 2513 |
| K051108-108A-32 | 0.281008 | 0.000054 | 0.000971 | 0.280970 | -62.8 | 1.9 | -17.5 | 2066 |
| K051108-108A-44 | 0.281476 | 0.000059 | 0.001567 | 0.281420 | -46.3 | 2.1 | -5.3 | 1903 |
| K051108-108A-46 | 0.281566 | 0.000041 | 0.000711 | 0.281540 | -43.1 | 1.4 | -1.5 | 1882 |
| K051108-108A-48 | 0.281201 | 0.000039 | 0.000656 | 0.281168 | -56.0 | 1.4 | 2.5 | 2624 |
| K051108-108A-50 | 0.281820 | 0.000045 | 0.001694 | 0.281761 | -34.1 | 1.6 | 4.8 | 1818 |
| K051108-108A-51 | 0.281152 | 0.000052 | 0.000737 | 0.281116 | -57.8 | 1.8 | -1.5 | 2533 |
| K051108-108A-52 | 0.281802 | 0.000044 | 0.001854 | 0.281734 | -34.8 | 1.6 | 6.4 | 1928 |
| K051108-108A-53 | 0.281694 | 0.000043 | 0.001210 | 0.281651 | -38.6 | 1.5 | 1.5 | 1843 |
| K051108-108A-57 | 0.281264 | 0.000040 | 0.000586 | 0.281236 | -53.8 | 1.4 | 2.6 | 2524 |
| K051108-108A-58 | 0.281051 | 0.000040 | 0.000732 | 0.281016 | -61.3 | 1.4 | -5.3 | 2523 |


| K051108-108A-59 | 0.281389 | 0.000053 | 0.000709 | 0.281362 | -49.4 | 1.9 | -5.8 | 1971 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K051108-108A-61 | 0.281641 | 0.000048 | 0.002329 | 0.281562 | -40.4 | 1.7 | -2.6 | 1801 |
| K051108-108A-62 | 0.281685 | 0.000056 | 0.001075 | 0.281648 | -38.9 | 2.0 | 0.5 | 1804 |
| K051108-108A-63 | 0.281646 | 0.000043 | 0.000880 | 0.281616 | -40.3 | 1.5 | -0.7 | 1800 |
| K051108-108A-64 | 0.281682 | 0.000039 | 0.001818 | 0.281620 | -39.0 | 1.4 | -0.7 | 1796 |
| K051108-108A-65 | 0.281624 | 0.000049 | 0.000228 | 0.281615 | -41.1 | 1.7 | 2.3 | 1931 |
| K051108-108A-66 | 0.281558 | 0.000033 | 0.001520 | 0.281500 | -43.4 | 1.2 | 0.3 | 2023 |
| K051108-108A-68 | 0.281117 | 0.000048 | 0.001394 | 0.281050 | -59.0 | 1.7 | -3.7 | 2541 |
| K051108-108A-69 | 0.281551 | 0.000041 | 0.001483 | 0.281499 | -43.6 | 1.5 | -3.7 | 1850 |
| K051108-108A-71 | 0.281298 | 0.000043 | 0.001171 | 0.281242 | -52.6 | 1.5 | 2.6 | 2515 |
| K051108-108A-72 | 0.281430 | 0.000045 | 0.001145 | 0.281375 | -47.9 | 1.6 | 7.3 | 2514 |
| K051108-108A-74 | 0.281135 | 0.000044 | 0.000455 | 0.281113 | -58.3 | 1.6 | -1.8 | 2523 |
| K051108-108A-75 | 0.281788 | 0.000031 | 0.000755 | 0.281760 | -35.3 | 1.1 | 7.1 | 1919 |
| K051108-108A-76 | 0.281236 | 0.000046 | 0.000973 | 0.281189 | -54.8 | 1.6 | 0.8 | 2518 |
| K051108-108A-77 | 0.281813 | 0.000045 | 0.000753 | 0.281785 | -34.4 | 1.6 | 9.0 | 1962 |
| K051108-108A-82 | 0.281466 | 0.000046 | 0.001174 | 0.281421 | -46.7 | 1.6 | -2.8 | 2011 |
| K051108-108A-84 | 0.281044 | 0.000041 | 0.000867 | 0.281001 | -61.6 | 1.4 | -4.9 | 2564 |
| K051108-108A-87 | 0.281783 | 0.000043 | 0.001332 | 0.281737 | -35.4 | 1.5 | 4.1 | 1825 |
| K051108-108A-89 | 0.281647 | 0.000036 | 0.001323 | 0.281601 | -40.2 | 1.3 | -0.7 | 1822 |
| K051108-108A-92 | 0.281478 | 0.000019 | 0.001907 | 0.281409 | -46.2 | 0.7 | -5.3 | 1919 |
| K051108-108A-93 | 0.281189 | 0.000035 | 0.001314 | 0.281123 | -56.4 | 1.2 | 0.5 | 2607 |
| K051108-108A-94 | 0.280728 | 0.000038 | 0.001725 | 0.280643 | -72.7 | 1.3 | -17.5 | 2569 |
| K051108-108A-99 | 0.281010 | 0.000041 | 0.000771 | 0.280972 | -62.8 | 1.4 | -5.2 | 2596 |
| K06-112-1-2 | 0.281092 | 0.000042 | 0.000914 | 0.281047 | -59.9 | 1.5 | -3.3 | 2559 |
| K06-112-1-3 | 0.281184 | 0.000040 | 0.001008 | 0.281134 | -56.6 | 1.4 | 0.6 | 2595 |
| K06-112-1-4 | 0.280690 | 0.000030 | 0.000460 | 0.280661 | -74.1 | 1.0 | -0.1 | 3277 |
| K06-112-1-6 | 0.281214 | 0.000031 | 0.000715 | 0.281177 | -55.6 | 1.1 | 3.5 | 2654 |
| K06-112-1-8 | 0.281581 | 0.000039 | 0.000436 | 0.281566 | -42.6 | 1.4 | -1.6 | 1838 |
| K06-112-1-9 | 0.281165 | 0.000040 | 0.001215 | 0.281109 | -57.3 | 1.4 | -3.9 | 2442 |
| K06-112-1-10 | 0.281722 | 0.000030 | 0.001049 | 0.281686 | -37.6 | 1.0 | 2.2 | 1821 |
| K06-112-1-13 | 0.281522 | 0.000054 | 0.000495 | 0.281505 | -44.7 | 1.9 | -5.2 | 1775 |
| K06-112-1-15 | 0.281110 | 0.000039 | 0.000993 | 0.281062 | -59.2 | 1.4 | -4.3 | 2495 |
| K06-112-1-16 | 0.281202 | 0.000034 | 0.001105 | 0.281148 | -56.0 | 1.2 | 0.2 | 2558 |
| K06-112-1-17 | 0.281179 | 0.000037 | 0.001039 | 0.281130 | -56.8 | 1.3 | -2.3 | 2479 |
| K06-112-1-18 | 0.281165 | 0.000030 | 0.000364 | 0.281146 | -57.3 | 1.1 | 2.9 | 2672 |
| K06-112-1-19 | 0.281777 | 0.000044 | 0.001392 | 0.281728 | -35.6 | 1.6 | 4.4 | 1850 |
| K06-112-1-20 | 0.281198 | 0.000044 | 0.000613 | 0.281169 | -56.1 | 1.5 | -0.8 | 2480 |
| K06-112-1-21 | 0.281109 | 0.000039 | 0.000794 | 0.281071 | -59.3 | 1.4 | -4.3 | 2480 |
| K06-112-1-22 | 0.281519 | 0.000049 | 0.000937 | 0.281486 | -44.8 | 1.7 | -3.9 | 1862 |
| K06-246-2-1 | 0.281806 | 0.000050 | 0.001527 | 0.281756 | -34.6 | 1.8 | 2.7 | 1734 |
| K06-246-2-5 | 0.281200 | 0.000035 | 0.000597 | 0.281172 | -56.0 | 1.2 | -0.8 | 2476 |
| K06-246-2-6 | 0.281846 | 0.000033 | 0.001337 | 0.281802 | -33.2 | 1.2 | 4.5 | 1738 |
| K06-246-2-7 | 0.281880 | 0.000036 | 0.000931 | 0.281849 | -32.0 | 1.3 | 6.1 | 1737 |
| K06-246-2-8 | 0.280557 | 0.000036 | 0.001872 | 0.280440 | -78.8 | 1.3 | -9.0 | 3236 |
| K06-246-2-12 | 0.281874 | 0.000041 | 0.001174 | 0.281835 | -32.2 | 1.4 | 5.7 | 1741 |
| K06-246-2-13 | 0.281970 | 0.000042 | 0.001898 | 0.281908 | -28.8 | 1.5 | 8.2 | 1738 |
| K06-246-2-14 | 0.281863 | 0.000046 | 0.001825 | 0.281803 | -32.6 | 1.6 | 4.3 | 1732 |
| K06-246-2-15 | 0.281837 | 0.000034 | 0.001743 | 0.281780 | -33.5 | 1.2 | 3.7 | 1739 |
| K06-246-2-17 | 0.280905 | 0.000034 | 0.001211 | 0.280843 | -66.5 | 1.2 | -8.0 | 2669 |
| K06-246-2-18 | 0.281806 | 0.000040 | 0.001224 | 0.281766 | -34.6 | 1.4 | 3.1 | 1735 |
| K06-246-2-20 | 0.281897 | 0.000040 | 0.001684 | 0.281842 | -31.4 | 1.4 | 5.8 | 1737 |
| K06-246-2-23 | 0.281127 | 0.000038 | 0.000544 | 0.281101 | -58.6 | 1.3 | -3.3 | 2476 |
| K06-246-2-24 | 0.281193 | 0.000047 | 0.000557 | 0.281167 | -56.3 | 1.6 | -1.1 | 2472 |
| K06-246-2-26 | 0.281227 | 0.000030 | 0.000498 | 0.281204 | -55.1 | 1.1 | 0.1 | 2468 |
| K06-246-2-28 | 0.281151 | 0.000053 | 0.001787 | 0.281090 | -57.8 | 1.9 | -19.4 | 1799 |


| K06-246-2-29 | 0.281626 | 0.000064 | 0.001227 | 0.281582 | -41.0 | 2.3 | -0.5 | 1861 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K06-246-2-31 | 0.281148 | 0.000047 | 0.001440 | 0.281075 | -57.9 | 1.7 | -0.6 | 2633 |
| K06-246-2-32 | 0.281881 | 0.000057 | 0.001181 | 0.281842 | -32.0 | 2.0 | 5.8 | 1736 |
| K06-246-2-33 | 0.281560 | 0.000041 | 0.000926 | 0.281527 | -43.3 | 1.5 | -1.6 | 1902 |
| K06-246-2-34 | 0.281886 | 0.000046 | 0.001347 | 0.281842 | -31.8 | 1.6 | 5.8 | 1738 |
| K06-246-2-42 | 0.281887 | 0.000043 | 0.001304 | 0.281844 | -31.8 | 1.5 | 6.0 | 1740 |
| K06-246-2-43 | 0.281833 | 0.000041 | 0.001657 | 0.281775 | -33.7 | 1.4 | 5.9 | 1843 |
| K06-246-2-46 | 0.281732 | 0.000080 | 0.002376 | 0.281654 | -37.2 | 2.8 | -0.8 | 1738 |
| K06-246-1-100 | 0.281949 | 0.000037 | 0.001830 | 0.281889 | -29.6 | 1.3 | 7.6 | 1740 |
| K06-246-1-53 | 0.281307 | 0.000097 | 0.002276 | 0.281191 | -52.3 | 3.4 | 4.3 | 2665 |
| K06-246-1-59 | 0.281918 | 0.000052 | 0.001461 | 0.281870 | -30.7 | 1.8 | 6.9 | 1741 |
| K06-246-1-60 | 0.281915 | 0.000061 | 0.002560 | 0.281831 | -30.7 | 2.2 | 5.5 | 1741 |
| K06-246-1-61 | 0.281958 | 0.000065 | 0.003116 | 0.281855 | -29.2 | 2.3 | 6.4 | 1741 |
| K06-246-1-63 | 0.281323 | 0.000044 | 0.000878 | 0.281280 | -51.7 | 1.6 | 5.0 | 2560 |
| K06-246-1-65 | 0.281920 | 0.000032 | 0.001402 | 0.281874 | -30.6 | 1.1 | 7.0 | 1739 |
| K06-246-1-71 | 0.281853 | 0.000033 | 0.001713 | 0.281797 | -32.9 | 1.2 | 4.4 | 1744 |
| K06-246-1-79 | 0.281362 | 0.000048 | 0.001759 | 0.281278 | -50.3 | 1.7 | 3.2 | 2489 |
| K06-246-1-82 | 0.281956 | 0.000057 | 0.001983 | 0.281890 | -29.3 | 2.0 | 7.8 | 1747 |
| K06-246-1-90 | 0.281876 | 0.000038 | 0.002300 | 0.281800 | -32.1 | 1.3 | 4.5 | 1744 |
| K06-246-1-91 | 0.281733 | 0.000045 | 0.001636 | 0.281679 | -37.2 | 1.6 | 0.1 | 1740 |
| K06-246-1-92 | 0.281881 | 0.000049 | 0.002339 | 0.281804 | -32.0 | 1.7 | 4.5 | 1740 |
| K06-246-1-93 | 0.281840 | 0.000050 | 0.002360 | 0.281762 | -33.4 | 1.8 | 3.1 | 1741 |
| K06-246-1-94 | 0.281980 | 0.000038 | 0.001262 | 0.281938 | -28.5 | 1.3 | 9.3 | 1740 |
| K06-246-1-95 | 0.281879 | 0.000039 | 0.001198 | 0.281837 | -32.0 | 1.4 | 8.6 | 1865 |
| K06-246-1-96 | 0.281963 | 0.000039 | 0.002109 | 0.281893 | -29.1 | 1.4 | 7.7 | 1739 |
| K06-246-1-97 | 0.280806 | 0.000033 | 0.001284 | 0.280738 | -70.0 | 1.2 | -9.7 | 2756 |

## Notes:

Hf fractionation is corrected by comparing measured $179 \mathrm{Hf} / 177 \mathrm{Hf}$ against known 179/177 (line by line). Beta Hf is applied as a power law. Yb fractionation is corrected by comparing measured $173 \mathrm{Yb} / 171 \mathrm{Yb}$ against known 173/171 (line by line) if 171 Yb intensity is more than $\sim 1$ mv . Beta Yb is applied as a power law.
Data are filtered by intensity of Hf (removed if below cutoff value determined by monitoring the average offset of the standards from their known values, which is set at the minimum offset)
Data are filtered by removing 1 max and 1 min value (out of 60 ).
Data are aflso filtered by $95 \%$ filter (rejected if outside of 2 -sigma std dev of full set)
Uncertainties are standard error of the mean, expressed at 1-sigma

Table A 3: U-Pb geochronologic analyses of Grand Canyon plutons (ALC)

|  |  |  |  |  |  | Isotope ratios |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Analysis | U | 206Pb | U/Th | 206Pb* | $\pm$ | 207Pb* | $\pm$ | 206Pb* | $\pm$ | error | $\begin{aligned} & \text { Best } \\ & \text { age } \end{aligned}$ | $\pm$ | Conc |
|  | (ppm) | 204 Pb |  | 207Pb* | (\%) | 235U* | (\%) | 238 U | (\%) | corr. | (Ma) | (Ma) | (\%) |
| K12-81L-3 | 380 | 13227 | 2.9 | 9.3282 | 0.3 | 4.3766 | 1.6 | 0.2961 | 1.5 | 0.98 | 1752.4 | 5.8 | 95.4 |
| K12-81L-4 | 326 | 31750 | 2.6 | 9.2709 | 0.4 | 4.6048 | 1.3 | 0.3096 | 1.2 | 0.94 | 1763.6 | 8.1 | 98.6 |
| K12-81L-5 | 461 | 13045 | 2.1 | 9.2852 | 0.5 | 4.7474 | 1.0 | 0.3197 | 0.8 | 0.86 | 1760.8 | 9.2 | 101.6 |
| K12-81L-7 | 613 | 32924 | 2.2 | 9.3032 | 0.3 | 4.7505 | 1.1 | 0.3205 | 1.0 | 0.96 | 1757.3 | 5.4 | 102.0 |
| K12-81L-8 | 535 | 717972 | 2.3 | 9.3100 | 0.1 | 4.7798 | 0.6 | 0.3227 | 0.6 | 0.98 | 1755.9 | 2.4 | 102.7 |
| K12-81L-11 | 331 | 25079 | 2.1 | 9.3239 | 0.4 | 4.5433 | 1.8 | 0.3072 | 1.8 | 0.98 | 1753.2 | 7.3 | 98.5 |
| K12-85-3L-1 | 431 | 799646 | 2.5 | 9.2997 | 0.2 | 4.7879 | 1.0 | 0.3229 | 1.0 | 0.99 | 1758.0 | 3.3 | 102.6 |
| K12-85-3L-2 | 355 | 463479 | 2.4 | 9.3229 | 0.2 | 4.6172 | 0.9 | 0.3122 | 0.9 | 0.98 | 1753.4 | 3.1 | 99.9 |
| K12-85-3L-3 | 337 | 203658 | 2.1 | 9.3185 | 0.2 | 4.6001 | 1.8 | 0.3109 | 1.8 | 1.00 | 1754.3 | 3.3 | 99.5 |
| K12-85-3L-4 | 207 | 130576 | 2.7 | 9.3132 | 0.3 | 4.5137 | 1.4 | 0.3049 | 1.4 | 0.98 | 1755.3 | 5.2 | 97.7 |
| K12-85-3L-6 | 439 | 506467 | 2.6 | 9.3141 | 0.1 | 4.7203 | 0.7 | 0.3189 | 0.7 | 0.99 | 1755.1 | 1.7 | 101.7 |
| K12-85-3L-7 | 420 | 64044 | 2.6 | 9.3325 | 0.2 | 4.6445 | 0.9 | 0.3144 | 0.8 | 0.98 | 1751.5 | 3.1 | 100.6 |
| K12-85-3L-8 | 584 | 315929 | 2.0 | 9.3141 | 0.2 | 4.6076 | 0.6 | 0.3113 | 0.6 | 0.93 | 1755.1 | 4.4 | 99.5 |
| K12-85-3L-8 | 400 | 426493 | 2.1 | 9.2974 | 0.1 | 4.6399 | 1.2 | 0.3129 | 1.2 | 0.99 | 1758.4 | 2.2 | 99.8 |
| K12-85-3L-9 | 429 | 98216 | 2.1 | 9.3035 | 0.1 | 4.6497 | 3.0 | 0.3137 | 3.0 | 1.00 | 1757.2 | 2.3 | 100.1 |
| K12-85-3L-10 | 548 | 75545 | 2.1 | 9.3054 | 0.2 | 4.5951 | 1.1 | 0.3101 | 1.0 | 0.98 | 1756.9 | 3.9 | 99.1 |
| K12-85-3L-11 | 377 | 527926 | 2.3 | 9.3240 | 0.3 | 4.5178 | 2.3 | 0.3055 | 2.3 | 0.99 | 1753.2 | 4.6 | 98.0 |
| K12-85-3L-12 | 482 | 140658 | 2.1 | 9.2977 | 0.2 | 4.7447 | 2.7 | 0.3199 | 2.7 | 1.00 | 1758.4 | 3.0 | 101.8 |
| K12-85-3L-13 | 246 | 351595 | 2.7 | 9.3070 | 0.3 | 4.8653 | 2.5 | 0.3284 | 2.5 | 0.99 | 1756.5 | 5.6 | 104.2 |
| K12-85-3L-14 | 291 | 709520 | 2.5 | 9.3189 | 0.2 | 4.7576 | 0.6 | 0.3215 | 0.6 | 0.92 | 1754.2 | 4.4 | 102.5 |
| K12-85-3L-15 | 254 | 208016 | 2.6 | 9.3676 | 0.4 | 4.6381 | 1.5 | 0.3151 | 1.4 | 0.97 | 1744.7 | 6.9 | 101.2 |
| K12-85-3L-16 | 594 | 20945 | 2.4 | 9.3229 | 0.3 | 4.4926 | 1.6 | 0.3038 | 1.6 | 0.98 | 1753.4 | 5.2 | 97.5 |
| K12-85-3L-17 | 405 | 33649 | 2.9 | 9.4042 | 0.3 | 4.3643 | 6.6 | 0.2977 | 6.6 | 1.00 | 1737.5 | 5.6 | 96.7 |
| K12-85-3L-18 | 417 | 824257 | 2.5 | 9.3182 | 0.2 | 4.8390 | 1.1 | 0.3270 | 1.1 | 0.98 | 1754.3 | 4.1 | 104.0 |
| K12-85-3L-19 | 150 | 135133 | 2.8 | 9.3566 | 0.5 | 4.1866 | 1.7 | 0.2841 | 1.6 | 0.95 | 1746.8 | 9.9 | 92.3 |
| K12-85-3L-20 | 464 | 381133 | 2.6 | 9.3217 | 0.1 | 4.6711 | 0.8 | 0.3158 | 0.8 | 0.99 | 1753.6 | 2.3 | 100.9 |
| K12-85-3L-23 | 681 | 15378 | 2.1 | 9.3178 | 0.5 | 3.8783 | 3.0 | 0.2621 | 2.9 | 0.99 | 1754.4 | 8.5 | 85.5 |
| K12-85-3L-24 | 467 | 7421 | 2.5 | 9.4001 | 0.4 | 3.9772 | 2.8 | 0.2712 | 2.7 | 0.99 | 1738.3 | 6.8 | 89.0 |
| K12-90-5R-1C | 115 | 33405 | 1.0 | 9.4239 | 1.0 | 4.4444 | 1.5 | 0.3038 | 1.2 | 0.76 | 1733.7 | 18.6 | 98.6 |
| K12-90-5R-1M | 119 | 37574 | 1.3 | 9.4634 | 0.8 | 4.4370 | 1.5 | 0.3045 | 1.3 | 0.87 | 1726.0 | 13.8 | 99.3 |
| K12-90-5R-2 | 251 | 211913 | 1.0 | 9.5060 | 0.3 | 4.4226 | 1.1 | 0.3049 | 1.0 | 0.95 | 1717.7 | 6.2 | 99.9 |
| K12-90-5R-3 | 141 | 15553 | 1.1 | 9.4735 | 0.8 | 4.4505 | 1.0 | 0.3058 | 0.6 | 0.64 | 1724.0 | 14.3 | 99.8 |
| K12-90-5R-4 | 136 | 32743 | 1.0 | 9.4269 | 0.8 | 4.4577 | 1.3 | 0.3048 | 1.0 | 0.80 | 1733.1 | 14.1 | 99.0 |
| K12-90-5R-5 | 207 | 141355 | 1.1 | 9.5034 | 0.4 | 4.4539 | 0.7 | 0.3070 | 0.6 | 0.86 | 1718.2 | 6.8 | 100.4 |
| K12-90-5R-6 | 208 | 70867 | 1.0 | 9.5112 | 0.6 | 4.4272 | 0.7 | 0.3054 | 0.4 | 0.57 | 1716.7 | 10.6 | 100.1 |
| K12-90-5R-7 | 189 | 178549 | 1.0 | 9.5204 | 0.4 | 4.4147 | 1.6 | 0.3048 | 1.6 | 0.97 | 1715.0 | 7.2 | 100.0 |
| K12-90-5R-8 | 181 | 91224 | 1.0 | 9.4794 | 0.7 | 4.4312 | 1.1 | 0.3046 | 0.8 | 0.73 | 1722.9 | 13.2 | 99.5 |
| K12-90-5R-9 | 441 | 378955 | 1.1 | 9.4358 | 0.9 | 4.6737 | 5.5 | 0.3198 | 5.4 | 0.99 | 1731.4 | 16.2 | 103.3 |
| K12-90-5R-10 | 154 | 99876 | 1.1 | 9.4618 | 0.5 | 4.4630 | 1.2 | 0.3063 | 1.1 | 0.92 | 1726.3 | 8.4 | 99.8 |
| K12-90-5R-11 | 167 | 117174 | 1.0 | 9.5222 | 0.5 | 4.4192 | 1.2 | 0.3052 | 1.1 | 0.92 | 1714.6 | 8.8 | 100.1 |
| K12-90-5R-12 | 117 | 19173 | 1.2 | 9.2963 | 3.2 | 4.5555 | 3.8 | 0.3071 | 2.0 | 0.52 | 1758.6 | 59.0 | 98.2 |
| K12-90-5R-13C | 138 | 25690 | 1.1 | 9.3503 | 2.9 | 4.5458 | 3.3 | 0.3083 | 1.7 | 0.51 | 1748.0 | 52.7 | 99.1 |
| K12-90-5R-13M | 110 | 121576 | 1.0 | 9.3826 | 1.3 | 4.4767 | 2.3 | 0.3046 | 1.9 | 0.83 | 1741.7 | 23.7 | 98.4 |
| K12-90-5R-14 | 193 | 92830 | 1.2 | 9.4921 | 0.7 | 4.4856 | 0.9 | 0.3088 | 0.5 | 0.61 | 1720.4 | 12.6 | 100.8 |
| K12-90-5R-15 | 169 | 117329 | 1.0 | 9.4638 | 0.5 | 4.4635 | 1.2 | 0.3064 | 1.1 | 0.89 | 1725.9 | 10.1 | 99.8 |
| K12-90-5R-16 | 324 | 30961 | 1.0 | 9.5005 | 0.5 | 4.4532 | 1.3 | 0.3068 | 1.2 | 0.94 | 1718.8 | 8.4 | 100.4 |
| K12-90-5R-17 | 515 | 326735 | 1.7 | 9.5112 | 0.1 | 4.4376 | 1.2 | 0.3061 | 1.2 | 1.00 | 1716.7 | 2.1 | 100.3 |
| K12-90-5R-18 | 258 | 343994 | 1.1 | 9.5194 | 0.4 | 4.4434 | 0.8 | 0.3068 | 0.7 | 0.88 | 1715.1 | 6.6 | 100.6 |
| K12-90-5R-19 | 141 | 156619 | 1.1 | 9.4506 | 0.4 | 4.4431 | 1.1 | 0.3045 | 1.0 | 0.93 | 1728.5 | 7.6 | 99.1 |


| K12-90-5R-20 | 180 | 48838 | 1.0 | 9.5083 | 0.4 | 4.4371 | 1.1 | 0.3060 | 1.0 | 0.93 | 1717.3 | 7.1 | 100.2 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K12-90-5R-21 | 196 | 147613 | 1.0 | 9.5292 | 0.4 | 4.4239 | 0.8 | 0.3057 | 0.7 | 0.84 | 1713.3 | 7.8 | 100.4 |
| K12-90-5R-22 | 219 | 37869 | 0.9 | 9.5326 | 0.4 | 4.4496 | 1.0 | 0.3076 | 0.9 | 0.89 | 1712.6 | 8.3 | 101.0 |
| K12-90-5R-23 | 170 | 83577 | 1.1 | 9.4930 | 0.5 | 4.4653 | 1.1 | 0.3074 | 1.0 | 0.90 | 1720.3 | 8.7 | 100.5 |
| K12-90-5R-24 | 99 | 57895 | 1.1 | 9.4984 | 0.9 | 4.4522 | 1.3 | 0.3067 | 1.0 | 0.74 | 1719.2 | 16.4 | 100.3 |
| K12-90-5R-25 | 74 | 89557 | 1.2 | 9.5656 | 1.3 | 4.4562 | 1.7 | 0.3092 | 1.0 | 0.61 | 1706.2 | 24.6 | 101.8 |
| K12-90-5R-26 | 273 | 96485 | 0.9 | 9.5024 | 0.2 | 4.4849 | 0.7 | 0.3091 | 0.7 | 0.97 | 1718.4 | 3.0 | 101.0 |
| K12-90-5R-27 | 154 | 81917 | 1.4 | 9.4674 | 0.9 | 4.4940 | 3.0 | 0.3086 | 2.9 | 0.95 | 1725.2 | 16.7 | 100.5 |
| K12-90-5R-28 | 78 | 83153 | 1.2 | 9.5267 | 0.6 | 4.4115 | 1.2 | 0.3048 | 1.0 | 0.86 | 1713.7 | 10.9 | 100.1 |
| K12-90-5R-29 | 212 | 262351 | 1.1 | 9.5087 | 0.3 | 4.4222 | 1.1 | 0.3050 | 1.1 | 0.96 | 1717.2 | 6.0 | 99.9 |
| K12-90-5R-30 | 143 | 138158 | 1.0 | 9.4820 | 0.5 | 4.3788 | 1.3 | 0.3011 | 1.2 | 0.93 | 1722.4 | 8.6 | 98.5 |
| K12-91-5R-1C | 96 | 77959 | 3.0 | 9.3640 | 0.6 | 4.6967 | 1.8 | 0.3190 | 1.6 | 0.93 | 1745.4 | 11.8 | 102.3 |
| K12-91-5R-1M | 138 | 8020 | 3.1 | 9.3817 | 0.6 | 4.4841 | 1.4 | 0.3051 | 1.3 | 0.90 | 1741.9 | 11.4 | 98.5 |
| K12-91-5R-3C | 72 | 84597 | 3.3 | 9.3512 | 0.8 | 4.6271 | 1.9 | 0.3138 | 1.7 | 0.90 | 1747.9 | 14.9 | 100.7 |
| K12-91-5R-4 | 138 | 110976 | 3.5 | 9.2780 | 0.4 | 4.6955 | 1.9 | 0.3160 | 1.8 | 0.98 | 1762.2 | 7.6 | 100.4 |
| K12-91-5R-5 | 77 | 92938 | 4.4 | 9.2800 | 0.8 | 4.7534 | 2.0 | 0.3199 | 1.9 | 0.93 | 1761.8 | 13.8 | 101.6 |
| K12-91-5R-6 | 55 | 9627 | 4.3 | 9.2835 | 1.0 | 4.6528 | 1.5 | 0.3133 | 1.2 | 0.75 | 1761.2 | 18.7 | 99.8 |
| K12-91-5R-7C | 99 | 13047 | 3.5 | 9.3867 | 0.6 | 4.7587 | 3.4 | 0.3240 | 3.4 | 0.98 | 1740.9 | 11.6 | 103.9 |
| K12-91-5R-7M | 73 | 91492 | 4.1 | 9.2886 | 0.6 | 4.7030 | 0.9 | 0.3168 | 0.7 | 0.76 | 1760.2 | 11.0 | 100.8 |
| K12-91-5R-8C | 114 | 94340 | 3.6 | 9.2979 | 0.6 | 4.8008 | 1.6 | 0.3237 | 1.4 | 0.93 | 1758.3 | 10.4 | 102.8 |
| K12-91-5R-9M | 150 | 8283 | 3.5 | 9.2875 | 0.7 | 4.6691 | 1.1 | 0.3145 | 0.8 | 0.75 | 1760.4 | 13.2 | 100.1 |
| K12-91-5R-11C | 68 | 68023 | 3.1 | 9.2391 | 0.8 | 4.6654 | 1.5 | 0.3126 | 1.2 | 0.83 | 1769.9 | 14.8 | 99.1 |
| K12-91-5R-11M | 183 | 7464 | 2.5 | 9.3039 | 0.6 | 4.5070 | 1.4 | 0.3041 | 1.3 | 0.91 | 1757.1 | 10.9 | 97.4 |
| K12-91-5R-12C | 251 | 4096 | 2.7 | 9.2960 | 1.6 | 4.0832 | 5.8 | 0.2753 | 5.6 | 0.96 | 1758.7 | 28.8 | 89.1 |
| K12-91-5R-12M | 173 | 1831 | 3.2 | 9.2921 | 1.5 | 4.4807 | 2.1 | 0.3020 | 1.4 | 0.68 | 1759.5 | 28.2 | 96.7 |
| K12-91-5R-13C | 164 | 7345 | 2.7 | 9.3155 | 0.8 | 4.6572 | 2.0 | 0.3147 | 1.8 | 0.91 | 1754.9 | 15.1 | 100.5 |
| K12-91-5R-14M | 105 | 68791 | 2.5 | 9.2795 | 0.8 | 4.6847 | 1.9 | 0.3153 | 1.7 | 0.91 | 1761.9 | 14.5 | 100.3 |
| K12-91-5R-15C | 232 | 118209 | 2.2 | 9.2958 | 0.3 | 4.5305 | 1.1 | 0.3054 | 1.1 | 0.97 | 1758.7 | 5.2 | 97.7 |
| K12-91-5R-15M | 182 | 4316 | 2.6 | 9.3209 | 0.7 | 4.1898 | 2.1 | 0.2832 | 2.0 | 0.94 | 1753.8 | 12.6 | 91.7 |
| K12-91-5R-16C | 88 | 23766 | 3.7 | 9.3794 | 0.8 | 4.6177 | 1.8 | 0.3141 | 1.6 | 0.90 | 1742.3 | 14.2 | 101.1 |
| K12-91-5R-16M | 128 | 3630 | 4.2 | 9.3646 | 0.7 | 4.4041 | 1.7 | 0.2991 | 1.6 | 0.92 | 1745.2 | 12.1 | 96.7 |
| K12-91-5R-17C | 160 | 242434 | 2.7 | 9.2920 | 0.6 | 4.7905 | 0.9 | 0.3228 | 0.7 | 0.77 | 1759.5 | 10.6 | 102.5 |
| K12-91-5R-17M | 95 | 97255 | 4.7 | 9.3484 | 0.7 | 4.7828 | 2.1 | 0.3243 | 2.0 | 0.94 | 1748.4 | 13.0 | 103.6 |
| K12-91-5R-18C | 174 | 49787 | 2.4 | 9.3289 | 0.3 | 4.8721 | 1.7 | 0.3296 | 1.6 | 0.98 | 1752.2 | 5.4 | 104.8 |
| K12-91-5R-18M | 73 | 85222 | 3.7 | 9.3621 | 0.7 | 4.6703 | 1.2 | 0.3171 | 0.9 | 0.80 | 1745.7 | 12.9 | 101.7 |
| K12-91-5R-19C | 130 | 86977 | 2.8 | 9.3324 | 0.6 | 4.5787 | 3.1 | 0.3099 | 3.0 | 0.98 | 1751.5 | 10.9 | 99.4 |
| K12-91-5R-19M | 156 | 14648 | 2.4 | 9.2814 | 0.6 | 4.4832 | 2.6 | 0.3018 | 2.5 | 0.97 | 1761.6 | 11.6 | 96.5 |
| K12-91-5R-20C | 216 | 176130 | 2.3 | 9.3241 | 0.4 | 4.7960 | 0.7 | 0.3243 | 0.5 | 0.75 | 1753.2 | 8.1 | 103.3 |
| K12-91-5R-20M | 324 | 5231 | 2.0 | 9.5152 | 1.2 | 3.7112 | 6.0 | 0.2561 | 5.9 | 0.98 | 1716.0 | 22.8 | 85.7 |
| K12-96-2L-1C | 299 | 433373 | 4.7 | 8.8942 | 0.3 | 5.3051 | 1.2 | 0.3422 | 1.1 | 0.96 | 1839.1 | 6.1 | 103.2 |
| K12-96-2L-1M | 159 | 12082 | 2.5 | 9.2863 | 2.7 | 4.2832 | 8.9 | 0.2885 | 8.5 | 0.95 | 1760.6 | 49.8 | 92.8 |
| K12-96-2L-3C | 137 | 66872 | 2.1 | 5.7429 | 0.2 | 11.8153 | 1.1 | 0.4921 | 1.0 | 0.98 | 2597.7 | 3.8 | 99.3 |
| K12-96-2L-3M | 296 | 2373 | 3.3 | 9.4565 | 0.9 | 4.0201 | 2.3 | 0.2757 | 2.1 | 0.92 | 1727.3 | 17.2 | 90.9 |
| K12-96-2L-4C | 109 | 38120 | 3.1 | 9.4614 | 2.0 | 4.2090 | 3.4 | 0.2888 | 2.8 | 0.81 | 1726.4 | 37.0 | 94.7 |
| K12-96-2L-4M | 353 | 6561 | 2.5 | 9.4284 | 1.0 | 3.9456 | 6.5 | 0.2698 | 6.4 | 0.99 | 1732.8 | 19.2 | 88.9 |
| K12-96-2L-6C | 143 | 7423 | 2.4 | 9.4053 | 0.8 | 4.2221 | 1.9 | 0.2880 | 1.7 | 0.89 | 1737.3 | 15.6 | 93.9 |
| K12-96-2L-8C | 247 | 7513 | 1.5 | 9.1461 | 5.0 | 4.6895 | 7.8 | 0.3111 | 6.1 | 0.77 | 1788.4 | 90.8 | 97.6 |
| K12-96-2L-8M | 393 | 12965 | 2.4 | 9.4255 | 0.4 | 4.4963 | 1.3 | 0.3074 | 1.3 | 0.96 | 1733.4 | 6.9 | 99.7 |
| K12-96-2L-10C | 245 | 21268 | 1.9 | 9.4165 | 0.6 | 4.5690 | 1.0 | 0.3120 | 0.7 | 0.78 | 1735.1 | 11.0 | 100.9 |
| K12-96-2L-10M | 303 | 29155 | 2.3 | 9.4425 | 0.3 | 4.4238 | 0.9 | 0.3030 | 0.9 | 0.96 | 1730.1 | 4.7 | 98.6 |
| K12-96-2L-11C | 370 | 11399 | 2.1 | 9.4134 | 1.0 | 4.4912 | 3.2 | 0.3066 | 3.0 | 0.95 | 1735.7 | 18.6 | 99.3 |
| K12-96-2L-13C | 190 | 20888 | 2.7 | 9.4546 | 0.8 | 4.6071 | 2.3 | 0.3159 | 2.2 | 0.95 | 1727.7 | 14.1 | 102.4 |
| K12-96-2L-15M | 289 | 12891 | 2.6 | 9.4467 | 0.9 | 4.0431 | 2.3 | 0.2770 | 2.1 | 0.92 | 1729.2 | 17.1 | 91.2 |
| K12-96-2L-16M | 292 | 5085 | 3.3 | 9.4271 | 1.1 | 4.3433 | 1.9 | 0.2970 | 1.5 | 0.79 | 1733.0 | 21.1 | 96.7 |
| K12-96-2L-17C | 119 | 128553 | 1.5 | 6.1481 | 0.4 | 10.3573 | 1.3 | 0.4618 | 1.2 | 0.95 | 2483.4 | 6.4 | 98.6 |


| K12-96-2L-20C | 270 | 8362 | 1.6 | 9.4455 | 0.9 | 4.5058 | 1.4 | 0.3087 | 1.1 | 0.76 | 1729.5 | 16.8 | 100.3 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K12-96-2L-20M | 281 | 8304 | 2.5 | 9.4712 | 0.7 | 4.3726 | 1.7 | 0.3004 | 1.6 | 0.92 | 1724.5 | 12.4 | 98.2 |
| K12-96-2L-21C | 288 | 4029 | 1.8 | 9.4417 | 1.0 | 3.9402 | 5.7 | 0.2698 | 5.6 | 0.99 | 1730.2 | 17.8 | 89.0 |
| K12-96-2L-21M | 289 | 15520 | 2.2 | 9.4660 | 0.4 | 4.3755 | 2.7 | 0.3004 | 2.7 | 0.99 | 1725.5 | 8.1 | 98.1 |
| K12-96-2L-22C | 385 | 12063 | 3.7 | 8.3873 | 0.4 | 5.5761 | 1.3 | 0.3392 | 1.2 | 0.94 | 1944.7 | 7.5 | 96.8 |
| K12-96-2L-23C | 170 | 40636 | 2.9 | 9.4925 | 0.5 | 3.9623 | 4.1 | 0.2728 | 4.1 | 0.99 | 1720.4 | 8.5 | 90.4 |
| K12-96-2L-23M | 305 | 4225 | 2.2 | 9.3194 | 2.1 | 4.2628 | 3.3 | 0.2881 | 2.6 | 0.78 | 1754.1 | 38.4 | 93.0 |
| K12-96-2L-24M | 235 | 2515 | 2.7 | 9.4278 | 0.7 | 4.0426 | 2.5 | 0.2764 | 2.4 | 0.96 | 1732.9 | 12.2 | 90.8 |
| K12-96-2L-25C | 156 | 4941 | 2.8 | 9.0883 | 1.6 | 4.8628 | 5.4 | 0.3205 | 5.2 | 0.96 | 1799.9 | 29.1 | 99.6 |
| 13H99R-1 | 182 | 26688 | 1.5 | 6.1151 | 0.3 | 9.9157 | 2.8 | 0.4398 | 2.8 | 0.99 | 2492.5 | 5.1 | 94.3 |
| 13H99R-2 | 594 | 3043 | 1.5 | 8.8354 | 0.9 | 4.3208 | 7.6 | 0.2769 | 7.5 | 0.99 | 1851.1 | 17.1 | 85.1 |
| 13H99R-3 | 284 | 7617 | 1.8 | 9.1764 | 0.8 | 4.6818 | 5.8 | 0.3116 | 5.8 | 0.99 | 1782.3 | 14.2 | 98.1 |
| 13H99R-5 | 236 | 8397 | 3.8 | 9.0816 | 0.5 | 4.5334 | 1.7 | 0.2986 | 1.6 | 0.95 | 1801.3 | 9.6 | 93.5 |
| 13H99R-6 | 488 | 31647 | 1.5 | 9.3239 | 0.5 | 4.3688 | 2.6 | 0.2954 | 2.5 | 0.98 | 1753.2 | 9.3 | 95.2 |
| 13H99R-7 | 210 | 2547 | 3.2 | 6.4635 | 1.1 | 7.4739 | 5.1 | 0.3504 | 4.9 | 0.98 | 2398.7 | 18.7 | 80.7 |
| 13H99R-8 | 104 | 8643 | 1.4 | 9.1875 | 0.9 | 4.5280 | 2.0 | 0.3017 | 1.8 | 0.91 | 1780.1 | 15.6 | 95.5 |
| 13H99R-9 | 273 | 376297 | 1.9 | 5.8047 | 0.7 | 11.0905 | 1.6 | 0.4669 | 1.5 | 0.89 | 2579.8 | 12.5 | 95.7 |
| 13H99R-10 | 87 | 52781 | 1.4 | 9.2677 | 1.0 | 4.6966 | 2.1 | 0.3157 | 1.8 | 0.87 | 1764.3 | 18.6 | 100.2 |
| 13H99R-11 | 171 | 27888 | 1.8 | 9.2666 | 0.4 | 4.5933 | 1.9 | 0.3087 | 1.8 | 0.97 | 1764.5 | 8.2 | 98.3 |
| 13H99R-14 | 292 | 21578 | 1.5 | 9.3921 | 0.3 | 4.4287 | 3.8 | 0.3017 | 3.8 | 1.00 | 1739.9 | 6.3 | 97.7 |
| 13H99R-16 | 96 | 112110 | 0.6 | 9.3901 | 1.1 | 4.6451 | 1.9 | 0.3164 | 1.5 | 0.81 | 1740.3 | 20.1 | 101.8 |
| 13H99R-17 | 311 | 13117 | 2.2 | 8.8841 | 0.4 | 4.8707 | 2.1 | 0.3138 | 2.1 | 0.98 | 1841.2 | 7.0 | 95.6 |
| 13H99R-18 | 393 | 502265 | 1.2 | 9.4576 | 0.3 | 4.4806 | 1.3 | 0.3073 | 1.2 | 0.98 | 1727.1 | 4.6 | 100.0 |
| 13H99R-21 | 229 | 5245 | 1.9 | 8.8350 | 0.5 | 4.4312 | 7.1 | 0.2839 | 7.1 | 1.00 | 1851.2 | 9.8 | 87.0 |
| 13H99R-22 | 275 | 125193 | 1.2 | 9.3497 | 0.5 | 4.6308 | 1.9 | 0.3140 | 1.8 | 0.96 | 1748.1 | 9.8 | 100.7 |
| 13H99R-25 | 120 | 1888 | 2.6 | 8.4678 | 2.2 | 5.4297 | 2.8 | 0.3335 | 1.7 | 0.62 | 1927.6 | 39.4 | 96.2 |
| 13H99R-26 | 251 | 6101 | 1.8 | 9.3748 | 0.4 | 4.1993 | 4.3 | 0.2855 | 4.2 | 1.00 | 1743.2 | 7.8 | 92.9 |
| 13H99R-27 | 297 | 17385 | 1.5 | 9.3579 | 1.1 | 4.2727 | 4.0 | 0.2900 | 3.8 | 0.96 | 1746.6 | 20.5 | 94.0 |
| 13H99R-29 | 404 | 2622 | 2.2 | 9.3227 | 1.5 | 3.6403 | 7.2 | 0.2461 | 7.0 | 0.98 | 1753.4 | 27.7 | 80.9 |
| 13H99R-31 | 88 | 44061 | 1.5 | 9.4100 | 1.1 | 4.6444 | 2.3 | 0.3170 | 2.0 | 0.88 | 1736.4 | 20.1 | 102.2 |
| 13H99R-32 | 78 | 41883 | 1.3 | 9.3974 | 1.0 | 4.5706 | 2.6 | 0.3115 | 2.4 | 0.93 | 1738.8 | 17.6 | 100.5 |
| 13H99R-34 | 457 | 1639 | 2.3 | 9.0079 | 1.4 | 4.0491 | 4.5 | 0.2645 | 4.3 | 0.95 | 1816.1 | 26.1 | 83.3 |
| 13H99R-35 | 497 | 14746 | 1.5 | 9.4478 | 0.4 | 4.5454 | 4.9 | 0.3115 | 4.9 | 1.00 | 1729.0 | 6.8 | 101.1 |
| 13H99R-36 | 221 | 2153 | 1.4 | 8.8360 | 3.4 | 4.9510 | 3.9 | 0.3173 | 1.9 | 0.48 | 1851.0 | 62.2 | 96.0 |
| 13H99R-37 | 430 | 5151 | 18.3 | 8.9898 | 1.4 | 4.6823 | 2.4 | 0.3053 | 1.9 | 0.81 | 1819.7 | 25.4 | 94.4 |
| 13H99R-38 | 87 | 11535 | 1.5 | 9.2538 | 1.0 | 4.4421 | 3.4 | 0.2981 | 3.3 | 0.96 | 1767.0 | 18.2 | 95.2 |
| 13H99R-40 | 163 | 1680 | 1.5 | 4.4711 | 1.9 | 14.8802 | 4.6 | 0.4825 | 4.2 | 0.91 | 3007.2 | 30.4 | 84.4 |
| 13H99R-41 | 509 | 17898 | 1.2 | 9.3356 | 0.6 | 4.7559 | 2.6 | 0.3220 | 2.5 | 0.97 | 1750.9 | 11.1 | 102.8 |
| 13H99R-44 | 149 | 633 | 1.3 | 8.7013 | 3.4 | 4.1948 | 8.8 | 0.2647 | 8.1 | 0.92 | 1878.7 | 60.9 | 80.6 |
| 13H99R-45 | 373 | 2183 | 4.3 | 8.4715 | 1.1 | 4.5690 | 3.9 | 0.2807 | 3.7 | 0.96 | 1926.8 | 19.7 | 82.8 |
| 13H99R-46 | 393 | 1727 | 1.7 | 9.1488 | 1.6 | 3.7464 | 7.3 | 0.2486 | 7.2 | 0.98 | 1787.8 | 28.7 | 80.1 |
| 13H99R-47 | 212 | 199649 | 1.4 | 6.1362 | 0.1 | 10.7699 | 1.6 | 0.4793 | 1.6 | 1.00 | 2486.7 | 2.5 | 101.5 |
| 13H99R-48 | 300 | 2111 | 1.7 | 9.3204 | 2.6 | 3.8341 | 3.4 | 0.2592 | 2.1 | 0.62 | 1753.9 | 47.9 | 84.7 |
| 13H99R-49 | 295 | 2911 | 3.7 | 8.5752 | 1.2 | 5.5863 | 2.1 | 0.3474 | 1.8 | 0.83 | 1905.0 | 21.5 | 100.9 |
| 13H99R-50 | 250 | 3809 | 3.5 | 8.7385 | 0.9 | 4.3227 | 2.3 | 0.2740 | 2.1 | 0.91 | 1871.0 | 17.1 | 83.4 |
| 13H99R-51 | 217 | 8454 | 1.4 | 9.3373 | 0.4 | 4.3790 | 3.1 | 0.2965 | 3.0 | 0.99 | 1750.6 | 8.1 | 95.6 |
| 13H99R-52 | 203 | 154052 | 1.4 | 5.1664 | 0.2 | 14.0706 | 2.3 | 0.5272 | 2.3 | 1.00 | 2772.6 | 3.5 | 98.5 |
| 13H99R-53 | 268 | 85058 | 1.4 | 9.4682 | 0.3 | 4.3289 | 1.1 | 0.2973 | 1.1 | 0.97 | 1725.1 | 4.8 | 97.3 |
| 13H99R-54 | 153 | 349886 | 1.6 | 9.0904 | 0.5 | 5.0225 | 2.2 | 0.3311 | 2.2 | 0.98 | 1799.5 | 8.3 | 102.5 |
| K12-115L-1 | 139 | 8526 | 5.2 | 8.8216 | 0.8 | 5.1334 | 2.1 | 0.3284 | 1.9 | 0.93 | 1853.9 | 14.3 | 98.8 |
| K12-115L-2 | 223 | 9227 | 3.5 | 8.7873 | 0.7 | 5.0700 | 1.9 | 0.3231 | 1.8 | 0.93 | 1860.9 | 12.4 | 97.0 |
| K12-115L-3 | 235 | 73680 | 4.2 | 8.8243 | 0.6 | 5.3889 | 5.9 | 0.3449 | 5.9 | 0.99 | 1853.4 | 10.8 | 103.1 |
| K12-115L-4C | 305 | 33153 | 3.2 | 8.6783 | 1.9 | 5.5604 | 3.4 | 0.3500 | 2.8 | 0.82 | 1883.5 | 34.5 | 102.7 |
| K12-115L-7 | 403 | 5072 | 2.4 | 8.8436 | 0.5 | 4.7426 | 2.0 | 0.3042 | 1.9 | 0.97 | 1849.4 | 8.5 | 92.6 |
| K12-115L-9 | 121 | 62444 | 4.2 | 8.8486 | 0.6 | 5.2156 | 0.9 | 0.3347 | 0.7 | 0.77 | 1848.4 | 11.0 | 100.7 |


| K12-115L-12 | 172 | 152547 | 5.4 | 8.7540 | 1.4 | 5.5466 | 4.0 | 0.3522 | 3.7 | 0.94 | 1867.8 | 24.5 | 104.1 |
| :--- | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: |
| K12-115L-15 | 446 | 27207 | 3.4 | 8.8187 | 0.5 | 5.3982 | 1.8 | 0.3453 | 1.7 | 0.96 | 1854.5 | 9.2 | 103.1 |
| K12-115L-18 | 583 | 8327 | 2.0 | 8.8279 | 0.7 | 5.0921 | 2.8 | 0.3260 | 2.7 | 0.97 | 1852.6 | 13.1 | 98.2 |
| K12-115L-19C | 393 | 8479 | 2.5 | 8.8105 | 0.5 | 5.2372 | 1.5 | 0.3347 | 1.4 | 0.93 | 1856.2 | 9.9 | 100.3 |
| K12-115L-22 | 335 | 15181 | 4.1 | 8.9769 | 0.4 | 4.8055 | 2.7 | 0.3129 | 2.7 | 0.99 | 1822.3 | 7.8 | 96.3 |
| K12-115L-25 | 279 | 12954 | 3.2 | 8.7714 | 0.8 | 5.2038 | 2.7 | 0.3310 | 2.6 | 0.96 | 1864.2 | 14.2 | 98.9 |
| K12-115L-26 | 335 | 11325 | 4.1 | 8.7958 | 0.7 | 5.2784 | 3.5 | 0.3367 | 3.4 | 0.98 | 1859.2 | 11.9 | 100.6 |
| K12-115L-27 | 441 | 47257 | 3.1 | 8.8096 | 0.6 | 5.2132 | 2.6 | 0.3331 | 2.5 | 0.97 | 1856.4 | 10.8 | 99.8 |

Notes:
Analyses with $>10 \%$ uncertainty (1-sigma) in $206 \mathrm{~Pb} / 238 \mathrm{U}$ age are not included.
Analyses with $>10 \%$ uncertainty (1-sigma) in $206 \mathrm{~Pb} / 207 \mathrm{~Pb}$ age are not included, unless $206 \mathrm{~Pb} / 238 \mathrm{U}$ age is $<500 \mathrm{Ma}$.
Best age is determined from 206Pb/238U age for analyses with 206Pb/238U age < 900 Ma and from 206Pb/207Pb age for analyses with 206Pb/238Uage > 900 Ma .
Concordance is based on $206 \mathrm{~Pb} / 238 \mathrm{U}$ age / 206Pb/207Pb age. Value is not reported for $206 \mathrm{~Pb} / 238 \mathrm{U}$ ages $<500 \mathrm{Ma}$ because of large uncertainty in 206Pb/207Pb age.
Analyses with $206 \mathrm{~Pb} / 238 \mathrm{U}$ age > 500 Ma and with >20\% discordance (<80\% concordance) are not included.
Analyses with 206Pb/238U age > 500 Ma and with >5\% reverse discordance (<105\% concordance) are not included.
All uncertainties are reported at the 1-sigma level, and include only measurement errors.
Systematic errors are shown as 206Pb/238U uncertainty, 206Pb/207Pb uncertainty to the right of each sample (at 2-sigma level).
$U$ concentration and U/Th are calibrated relative to Sri Lanka zircon and are accurate to ~20\%. Common Pb correction is from 204Pb, with composition interpreted from Stacey and Kramers (1975).

Uncertainties of 1.5 for $206 \mathrm{~Pb} / 204 \mathrm{~Pb}, 0.3$ for $207 \mathrm{~Pb} / 204 \mathrm{~Pb}$, and 2.0 for $208 \mathrm{~Pb} / 204 \mathrm{~Pb}$ are applied to common Pb composition.
$\mathrm{U} / \mathrm{Pb}$ and 206Pb/207Pb fractionation is calibrated relative to fragments of a large Sri Lanka zircon of $563.5 \pm 3.2 \mathrm{Ma}$ (2-sigma).
$U$ decay constants and composition as follows: $238 \mathrm{U}=9.8485 \times 10-10,235 \mathrm{U}=1.55125 \times 10-10$, $238 \mathrm{U} / 235 \mathrm{U}=137.88$
Analytical methods as described by Gehrels et al. (2008).

Table A 4: Hf isotopic data for Grand Canyon plutons (ALC)

| Sample | ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}$ | $\pm(1 \mathrm{~s})$ | ${ }^{176} \mathrm{Lu} /{ }^{177} \mathrm{Hf}$ | $\begin{gathered} { }^{1 / 6} \mathrm{Hf} /{ }^{1 / 7} \mathrm{Hf} \\ (\mathrm{~T}) \\ \hline \end{gathered}$ | $\begin{gathered} \mathrm{E}-\mathrm{Hf} \\ (0) \\ \hline \end{gathered}$ | $\begin{gathered} \mathrm{E}-\mathrm{Hf}(0) \pm \\ (1 \mathrm{~s}) \end{gathered}$ | E-Hf (T) | Age (Ma) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K12-81L-3 | 0.281931 | 0.000036 | 0.002576 | 0.281846 | -30.2 | 1.3 | 6.3 | 1752 |
| K12-81L-5 | 0.282024 | 0.000041 | 0.003186 | 0.281918 | -26.9 | 1.5 | 9.1 | 1761 |
| K12-81L-7 | 0.282124 | 0.000051 | 0.004981 | 0.281957 | -23.4 | 1.8 | 10.4 | 1757 |
| K12-81L-8 | 0.282152 | 0.000050 | 0.003547 | 0.282034 | -22.4 | 1.8 | 13.1 | 1756 |
| K12-81L-11 | 0.282006 | 0.000048 | 0.002820 | 0.281912 | -27.5 | 1.7 | 8.7 | 1753 |
| K12-85.3L-1 | 0.281997 | 0.000056 | 0.004510 | 0.281847 | -27.9 | 2.0 | 6.5 | 1758 |
| K12-85.3L-2 | 0.282141 | 0.000045 | 0.003789 | 0.282015 | -22.8 | 1.6 | 12.3 | 1753 |
| K12-85.3L-3 | 0.282081 | 0.000049 | 0.003088 | 0.281978 | -24.9 | 1.7 | 11.1 | 1754 |
| K12-85.3L-4 | 0.282002 | 0.000048 | 0.001816 | 0.281941 | -27.7 | 1.7 | 9.8 | 1755 |
| K12-85.3L-6 | 0.282064 | 0.000053 | 0.002334 | 0.281986 | -25.5 | 1.9 | 11.4 | 1755 |
| K12-85.3L-7 | 0.282039 | 0.000055 | 0.004704 | 0.281883 | -26.4 | 1.9 | 7.6 | 1751 |
| K12-85.3L-8 | 0.281937 | 0.000040 | 0.003132 | 0.281833 | -30.0 | 1.4 | 5.9 | 1755 |
| K12-85.3L-9 | 0.282112 | 0.000058 | 0.003778 | 0.281986 | -23.8 | 2.1 | 11.4 | 1757 |
| K12-85.3L-10 | 0.281998 | 0.000075 | 0.003973 | 0.281865 | -27.8 | 2.7 | 7.1 | 1757 |
| K12-85.3L-11 | 0.282058 | 0.000044 | 0.003757 | 0.281933 | -25.7 | 1.5 | 9.4 | 1753 |
| K12-85.3L-12 | 0.282036 | 0.000052 | 0.003929 | 0.281905 | -26.5 | 1.9 | 8.5 | 1758 |
| K12-85.3L-13 | 0.282065 | 0.000046 | 0.003384 | 0.281952 | -25.5 | 1.6 | 10.2 | 1756 |
| K12-85.3L-14 | 0.282003 | 0.000048 | 0.002453 | 0.281921 | -27.6 | 1.7 | 9.1 | 1754 |
| K12-85.3L-15 | 0.282042 | 0.000043 | 0.002178 | 0.281970 | -26.3 | 1.5 | 10.6 | 1745 |
| K12-85.3L-16 | 0.282187 | 0.000055 | 0.004845 | 0.282025 | -21.2 | 2.0 | 12.7 | 1753 |
| K12-85.3L-17 | 0.282003 | 0.000046 | 0.003463 | 0.281889 | -27.6 | 1.6 | 7.5 | 1737 |
| K12-85.3L-18 | 0.282062 | 0.000050 | 0.004324 | 0.281918 | -25.6 | 1.8 | 8.9 | 1754 |
| K12-85.3L-19 | 0.282019 | 0.000058 | 0.003019 | 0.281919 | -27.1 | 2.0 | 8.8 | 1747 |
| K12-85.3L-20 | 0.282035 | 0.000036 | 0.003011 | 0.281935 | -26.5 | 1.3 | 9.5 | 1754 |
| K12-85.3L-23 | 0.282058 | 0.000046 | 0.004162 | 0.281920 | -25.7 | 1.6 | 9.0 | 1754 |
| K12-85.3L-24 | 0.282087 | 0.000061 | 0.004303 | 0.281945 | -24.7 | 2.2 | 9.5 | 1738 |
| K12-90.5R-1C | 0.281923 | 0.000069 | 0.000422 | 0.281909 | -30.5 | 2.4 | 8.2 | 1734 |
| K12-90.5R-1M | 0.281952 | 0.000046 | 0.000308 | 0.281942 | -29.4 | 1.6 | 9.1 | 1726 |
| K12-90.5R-2 | 0.281900 | 0.000057 | 0.000946 | 0.281869 | -31.3 | 2.0 | 6.5 | 1724 |
| K12-90.5R-3 | 0.281941 | 0.000043 | 0.000455 | 0.281926 | -29.8 | 1.5 | 8.4 | 1718 |
| K12-90.5R-4 | 0.282016 | 0.000048 | 0.000672 | 0.281994 | -27.2 | 1.7 | 11.1 | 1733 |
| K12-90.5R-5 | 0.281906 | 0.000041 | 0.000639 | 0.281885 | -31.1 | 1.4 | 6.9 | 1718 |
| K12-90.5R-6 | 0.281894 | 0.000044 | 0.000909 | 0.281864 | -31.5 | 1.6 | 6.2 | 1717 |
| K12-90.5R-7 | 0.282050 | 0.000048 | 0.000899 | 0.282020 | -26.0 | 1.7 | 11.7 | 1715 |
| K12-90.5R-8 | 0.281903 | 0.000037 | 0.000762 | 0.281878 | -31.2 | 1.3 | 6.8 | 1723 |
| K12-90.5R-9 | 0.281860 | 0.000050 | 0.000737 | 0.281836 | -32.7 | 1.8 | 5.5 | 1731 |
| K12-90.5R-10 | 0.281895 | 0.000049 | 0.000675 | 0.281873 | -31.5 | 1.7 | 6.7 | 1726 |
| K12-90.5R-11 | 0.281931 | 0.000050 | 0.000819 | 0.281904 | -30.2 | 1.8 | 7.5 | 1715 |


| K12-90.5R-14 | 0.281995 | 0.000051 | 0.000765 | 0.281970 | -27.9 | 1.8 | 10.0 | 1720 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K12-90.5R-15 | 0.281862 | 0.000038 | 0.000659 | 0.281841 | -32.6 | 1.3 | 5.5 | 1726 |
| K12-90.5R-16 | 0.281895 | 0.000031 | 0.000968 | 0.281863 | -31.5 | 1.1 | 6.2 | 1719 |
| K12-90.5R-17 | 0.281988 | 0.000033 | 0.000990 | 0.281956 | -28.2 | 1.2 | 9.4 | 1717 |
| K12-90.5R-18 | 0.281923 | 0.000046 | 0.000635 | 0.281902 | -30.5 | 1.6 | 7.5 | 1715 |
| K12-90.5R-19 | 0.281938 | 0.000050 | 0.000618 | 0.281918 | -29.9 | 1.8 | 8.3 | 1728 |
| K12-90.5R-20 | 0.282068 | 0.000042 | 0.000871 | 0.282039 | -25.4 | 1.5 | 12.4 | 1717 |
| K12-90.5R-21 | 0.281981 | 0.000045 | 0.000878 | 0.281953 | -28.4 | 1.6 | 9.2 | 1713 |
| K12-90.5R-22 | 0.281937 | 0.000042 | 0.001057 | 0.281903 | -30.0 | 1.5 | 7.5 | 1713 |
| K12-90.5R-23 | 0.282026 | 0.000048 | 0.000726 | 0.282002 | -26.8 | 1.7 | 11.1 | 1720 |
| K12-90.5R-24 | 0.282021 | 0.000038 | 0.000789 | 0.281996 | -27.0 | 1.3 | 10.9 | 1719 |
| K12-90.5R-26 | 0.281942 | 0.000040 | 0.001108 | 0.281906 | -29.8 | 1.4 | 7.7 | 1718 |
| K12-90.5R-27 | 0.281900 | 0.000032 | 0.000372 | 0.281887 | -31.3 | 1.1 | 7.2 | 1725 |
| K12-90.5R-28 | 0.281957 | 0.000038 | 0.000527 | 0.281940 | -29.3 | 1.4 | 8.8 | 1714 |
| K12-90.5R-29 | 0.281923 | 0.000040 | 0.000762 | 0.281898 | -30.5 | 1.4 | 7.4 | 1717 |
| K12-90.5R-30 | 0.281958 | 0.000039 | 0.000605 | 0.281938 | -29.2 | 1.4 | 8.9 | 1722 |
| K12-91.5R-1C | 0.282031 | 0.000041 | 0.001308 | 0.281988 | -26.7 | 1.5 | 11.2 | 1745 |
| K12-91.5R-1M | 0.281959 | 0.000038 | 0.001341 | 0.281914 | -29.2 | 1.3 | 8.5 | 1742 |
| K12-91.5R-3C | 0.281997 | 0.000043 | 0.001336 | 0.281953 | -27.9 | 1.5 | 10.0 | 1748 |
| K12-91.5R-4 | 0.281937 | 0.000030 | 0.000943 | 0.281905 | -30.0 | 1.1 | 8.7 | 1762 |
| K12-91.5R-5 | 0.281978 | 0.000049 | 0.000894 | 0.281948 | -28.5 | 1.7 | 10.2 | 1762 |
| K12-91.5R-6 | 0.281988 | 0.000041 | 0.000778 | 0.281962 | -28.2 | 1.4 | 10.7 | 1761 |
| K12-91.5R-7C | 0.282034 | 0.000034 | 0.001137 | 0.281996 | -26.6 | 1.2 | 11.4 | 1741 |
| K12-91.5R-7M | 0.281859 | 0.000032 | 0.000753 | 0.281834 | -32.7 | 1.1 | 6.1 | 1760 |
| K12-91.5R-8C | 0.281981 | 0.000035 | 0.001137 | 0.281943 | -28.4 | 1.2 | 9.9 | 1758 |
| K12-91.5R-14M | 0.281941 | 0.000038 | 0.001543 | 0.281889 | -29.9 | 1.3 | 8.1 | 1762 |
| K12-91.5R-13C | 0.281980 | 0.000042 | 0.002134 | 0.281909 | -28.5 | 1.5 | 8.6 | 1755 |
| K12-91.5R-15C | 0.282030 | 0.000046 | 0.002696 | 0.281940 | -26.7 | 1.6 | 9.8 | 1759 |
| K12-91.5R-15M | 0.282010 | 0.000043 | 0.001286 | 0.281967 | -27.4 | 1.5 | 10.7 | 1754 |
| K12-91.5R-16C | 0.281934 | 0.000039 | 0.001204 | 0.281894 | -30.1 | 1.4 | 7.8 | 1742 |
| K12-91.5R-16M | 0.281961 | 0.000045 | 0.001336 | 0.281917 | -29.1 | 1.6 | 8.7 | 1745 |
| K12-91.5R-17C | 0.281969 | 0.000043 | 0.001268 | 0.281927 | -28.8 | 1.5 | 9.4 | 1759 |
| K12-91.5R-17M | 0.281962 | 0.000038 | 0.001021 | 0.281928 | -29.1 | 1.4 | 9.2 | 1748 |
| K12-91.5R-18M | 0.281958 | 0.000046 | 0.000686 | 0.281935 | -29.3 | 1.6 | 9.3 | 1746 |
| K12-91.5R-18C | 0.281890 | 0.000039 | 0.001869 | 0.281828 | -31.7 | 1.4 | 5.7 | 1752 |
| K12-91.5R-19C | 0.281904 | 0.000044 | 0.001600 | 0.281850 | -31.2 | 1.6 | 6.5 | 1751 |
| K12-91.5R-19M | 0.281892 | 0.000040 | 0.001373 | 0.281847 | -31.6 | 1.4 | 6.6 | 1762 |
| K12-91.5R-20C | 0.282012 | 0.000050 | 0.002182 | 0.281939 | -27.3 | 1.8 | 9.7 | 1753 |
| K12-91.5R-20M | 0.281989 | 0.000037 | 0.002581 | 0.281905 | -28.1 | 1.3 | 7.6 | 1716 |
| K12-96.2L-1C | 0.281889 | 0.000035 | 0.000929 | 0.281856 | -31.7 | 1.2 | 8.7 | 1839 |
| K12-96.2L-3M | 0.281986 | 0.000033 | 0.002053 | 0.281918 | -28.3 | 1.2 | 8.3 | 1727 |


| K12-96.2L-4M | 0.281995 | 0.000063 | 0.001817 | 0.281935 | -28.0 | 2.2 | 9.0 | 1733 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K12-96.2L-8M | 0.281910 | 0.000040 | 0.001212 | 0.281870 | -30.9 | 1.4 | 6.6 | 1726 |
| K12-96.2L-6C | 0.282045 | 0.000044 | 0.001055 | 0.282010 | -26.2 | 1.6 | 11.8 | 1737 |
| K12-96.2L-11C | 0.281910 | 0.000045 | 0.000925 | 0.281879 | -31.0 | 1.6 | 7.1 | 1736 |
| K12-96.2L-10C | 0.281974 | 0.000040 | 0.001080 | 0.281938 | -28.7 | 1.4 | 9.2 | 1735 |
| K12-96.2L-10M | 0.281927 | 0.000052 | 0.001242 | 0.281886 | -30.3 | 1.8 | 7.2 | 1730 |
| K12-96.2L-13C | 0.281884 | 0.000045 | 0.000972 | 0.281852 | -31.9 | 1.6 | 6.0 | 1728 |
| K12-96.2L-15M | 0.281927 | 0.000044 | 0.000795 | 0.281901 | -30.3 | 1.6 | 7.7 | 1729 |
| K12-96.2L-16M | 0.281877 | 0.000039 | 0.000706 | 0.281854 | -32.1 | 1.4 | 6.2 | 1733 |
| K12-96.2L-17C | 0.281156 | 0.000029 | 0.000960 | 0.281110 | -57.6 | 1.0 | -2.9 | 2483 |
| K12-96.2L-20C | 0.281947 | 0.000039 | 0.001250 | 0.281906 | -29.6 | 1.4 | 7.9 | 1729 |
| K12-96.2L-20M | 0.281901 | 0.000039 | 0.000892 | 0.281872 | -31.2 | 1.4 | 6.6 | 1724 |
| K12-96.2L-21M | 0.282034 | 0.000042 | 0.001848 | 0.281973 | -26.6 | 1.5 | 10.2 | 1725 |
| K12-96.2L-21C | 0.282024 | 0.000037 | 0.001022 | 0.281990 | -26.9 | 1.3 | 10.9 | 1730 |
| K12-96.2L-22C | 0.281412 | 0.000033 | 0.000360 | 0.281399 | -48.5 | 1.2 | -5.1 | 1945 |
| K12-96.2L-23C | 0.281933 | 0.000048 | 0.001024 | 0.281900 | -30.1 | 1.7 | 7.5 | 1720 |
| K12-96.2L-23M | 0.282071 | 0.000043 | 0.002679 | 0.281982 | -25.3 | 1.5 | 11.2 | 1754 |
| K12-96.2L-24M | 0.281872 | 0.000033 | 0.000939 | 0.281841 | -32.3 | 1.2 | 5.7 | 1733 |
| 13H-099R-1 | 0.281262 | 0.000050 | 0.001186 | 0.281205 | -53.9 | 1.8 | 0.7 | 2492 |
| 13H-099R-11 | 0.281907 | 0.000043 | 0.001571 | 0.281854 | -31.0 | 1.5 | 6.9 | 1764 |
| 13H-099R-14 | 0.281987 | 0.000052 | 0.001426 | 0.281940 | -28.2 | 1.8 | 9.4 | 1740 |
| 13H-099R-17 | 0.281918 | 0.000054 | 0.002224 | 0.281840 | -30.7 | 1.9 | 8.2 | 1841 |
| 13H-099R-18 | 0.282021 | 0.000040 | 0.000794 | 0.281995 | -27.0 | 1.4 | 11.1 | 1727 |
| 13H-099R-2 | 0.281524 | 0.000034 | 0.001427 | 0.281474 | -44.6 | 1.2 | -4.6 | 1851 |
| 13H-099R-21 | 0.281836 | 0.000039 | 0.001598 | 0.281780 | -33.5 | 1.4 | 6.3 | 1851 |
| 13H-099R-22 | 0.281787 | 0.000038 | 0.001631 | 0.281733 | -35.3 | 1.3 | 2.2 | 1748 |
| 13H-099R-26 | 0.281863 | 0.000046 | 0.000145 | 0.281858 | -32.6 | 1.6 | 6.6 | 1743 |
| 13H-099R-35 | 0.281913 | 0.000045 | 0.001775 | 0.281855 | -30.8 | 1.6 | 6.1 | 1729 |
| 13H-099R-47 | 0.281349 | 0.000050 | 0.001619 | 0.281272 | -50.8 | 1.8 | 3.0 | 2487 |
| 13H-099R-52 | 0.281087 | 0.000043 | 0.000528 | 0.281059 | -60.0 | 1.5 | 2.1 | 2773 |
| 13H-099R-6 | 0.281808 | 0.000041 | 0.002253 | 0.281733 | -34.6 | 1.5 | 2.3 | 1753 |
| 13H-099R-7 | 0.281319 | 0.000030 | 0.000543 | 0.281294 | -51.8 | 1.1 | 1.7 | 2399 |
| K12-115L-1 | 0.281891 | 0.000048 | 0.001511 | 0.281838 | -31.6 | 1.7 | 8.4 | 1854 |
| K12-115L-3 | 0.281896 | 0.000049 | 0.002146 | 0.281820 | -31.4 | 1.7 | 7.7 | 1853 |
| K12-115L-2 | 0.281993 | 0.000039 | 0.002361 | 0.281910 | -28.0 | 1.4 | 11.1 | 1861 |
| K12-115L-7 | 0.281954 | 0.000042 | 0.002559 | 0.281864 | -29.4 | 1.5 | 9.2 | 1849 |
| K12-115L-9 | 0.281921 | 0.000041 | 0.001254 | 0.281877 | -30.5 | 1.4 | 9.7 | 1848 |
| K12-115L-15 | 0.281916 | 0.000059 | 0.003148 | 0.281805 | -30.7 | 2.1 | 7.2 | 1854 |
| K12-115L-18 | 0.281920 | 0.000066 | 0.003738 | 0.281788 | -30.6 | 2.3 | 6.6 | 1853 |
| K12-115L-19C | 0.281941 | 0.000042 | 0.003042 | 0.281834 | -29.9 | 1.5 | 8.3 | 1856 |
| K12-115L-22 | 0.281943 | 0.000043 | 0.002153 | 0.281869 | -29.8 | 1.5 | 8.7 | 1822 |


| K12-115L-25 | 0.282010 | 0.000043 | 0.002696 | 0.281914 | -27.4 | 1.5 | 11.3 | 1864 |
| :--- | :--- | :--- | :--- | :--- | :---: | :---: | :---: | :---: |
| K12-115L-26 | 0.281908 | 0.000055 | 0.001978 | 0.281838 | -31.0 | 1.9 | 8.5 | 1859 |
| K12-115L-27 | 0.281882 | 0.000051 | 0.003121 | 0.281772 | -31.9 | 1.8 | 6.1 | 1856 |

Notes:
Hf fractionation is corrected by comparing measured 179Hf/177Hf against known 179/177 (line by line). Beta Hf is applied as a power law.
Yb fractionation is corrected by comparing measured 173Yb/171Yb against known 173/171 (line by line) if 171 Yb intensity is more than $\sim 1 \mathrm{mv}$. Beta Yb is applied as a power law.
Data are filtered by intensity of Hf (removed if below cutoff value determined by monitoring the average offset of the standards from their known values, which is set at the minimum offset) Data are filtered by removing 1 max and 1 min value (out of 60).
Data are aflso filtered by $95 \%$ filter (rejected if outside of 2 -sigma std dev of full set) Uncertainties are standard error of the mean, expressed at 1-sigma

Table A 5: U-Pb geochronologic data from Grand Canyon plutons (GEMOC)

|  |  |  |  | A G ES (common-Pb corrected, Ma) |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Analysis No. | Th | U | Th/U | $\begin{aligned} & \hline 207 \mathrm{~Pb} / \\ & { }^{206} \mathrm{~Pb} \end{aligned}$ | $\begin{gathered} \pm \\ 1 \sigma \end{gathered}$ | $\begin{gathered} { }^{207} \mathrm{~Pb} / \\ { }^{235} \mathrm{U} \end{gathered}$ | $\begin{gathered} \pm \\ 1 \sigma \end{gathered}$ | $\begin{aligned} & { }^{206} \mathrm{~Pb} / \\ & { }^{238} \mathrm{U} \end{aligned}$ | $\begin{gathered} \pm \\ 1 \sigma \end{gathered}$ | $\begin{aligned} & { }^{208} \mathrm{~Pb} / \\ & { }^{232} \mathrm{Th} \end{aligned}$ | $\begin{gathered} \pm \\ 1 \sigma \end{gathered}$ | Disc. |
|  | (ppm) | (ppm) |  |  |  |  |  |  |  |  |  | \% |
| K05-100.5-1Rim | 106 | 137 | 0.78 | 1739 | 8 | 1740 | 8 | 1740 | 16 | 1798 | 17 | 0.0 |
| K05-100.5-2Core | 130 | 129 | 1.01 | 2547 | 8 | 2547 | 9 | 2547 | 22 | 2729 | 25 | 0.0 |
| K05-100.5-3 | 68 | 149 | 0.46 | 2475 | 7 | 2474 | 9 | 2472 | 20 | 2588 | 23 | 0.1 |
| K05-100.5-4 | 91 | 124 | 0.74 | 2936 | 7 | 2936 | 9 | 2936 | 22 | 3047 | 26 | 0.0 |
| K05-100.5-5 | 86 | 146 | 0.59 | 2476 | 8 | 2476 | 9 | 2475 | 21 | 2596 | 23 | 0.0 |
| K05-100.5-6 | 67 | 123 | 0.55 | 1732 | 8 | 1716 | 8 | 1703 | 14 | 1770 | 17 | 1.9 |
| K05-100.5-7 | 199 | 286 | 0.69 | 1749 | 8 | 1747 | 8 | 1746 | 15 | 1825 | 16 | 0.2 |
| K05-100.5-8 | 129 | 218 | 0.59 | 2469 | 9 | 2465 | 11 | 2461 | 24 | 2693 | 30 | 0.4 |
| K05-100.5-9 | 106 | 209 | 0.51 | 2470 | 8 | 2471 | 9 | 2472 | 21 | 2575 | 23 | -0.1 |
| K05-100.5-12 | 90 | 166 | 0.54 | 2498 | 8 | 2497 | 9 | 2496 | 21 | 2829 | 26 | 0.1 |
| K05-100.5-19 | 78 | 555 | 0.14 | 2756 | 7 | 2754 | 9 | 2751 | 23 | 2872 | 26 | 0.2 |
| K05-100.5-39 | 175 | 194 | 0.90 | 1734 | 10 | 1731 | 10 | 1729 | 18 | 1870 | 21 | 0.3 |
| K05-100.5-54 | 103 | 205 | 0.51 | 1731 | 9 | 1731 | 8 | 1731 | 16 | 1817 | 18 | 0.0 |
| K05-100.5-57 | 220 | 497 | 0.44 | 2487 | 8 | 2487 | 10 | 2487 | 22 | 2785 | 26 | 0.0 |
| K05-100.5-58 | 280 | 1625 | 0.17 | 2397 | 9 | 2396 | 10 | 2394 | 23 | 3742 | 36 | 0.2 |
| K05-100.5-71 | 162 | 132 | 1.22 | 1778 | 8 | 1777 | 8 | 1776 | 15 | 1877 | 17 | 0.2 |
| K05-100.5-75 | 90 | 168 | 0.54 | 2504 | 8 | 2503 | 10 | 2502 | 22 | 2762 | 26 | 0.1 |
| K05-100.5-82 | 109 | 212 | 0.52 | 2468 | 7 | 2467 | 9 | 2465 | 20 | 2640 | 24 | 0.1 |
| K05-100.5-89 | 261 | 706 | 0.37 | 1758 | 34 | 1770 | 12 | 1780 | 18 | 1782 | 20 | -1.5 |
| K05-100.5-95 | 66 | 144 | 0.46 | 2417 | 30 | 2399 | 13 | 2378 | 21 | 2373 | 24 | 1.9 |
| K05-100.5-97 | 42 | 101 | 0.42 | 1731 | 8 | 1727 | 8 | 1723 | 15 | 2606 | 28 | 0.6 |
| K05-100.5-100 | 408 | 754 | 0.54 | 2473 | 8 | 2473 | 9 | 2472 | 20 | 2626 | 23 | 0.0 |
| K05-100.5-101 | 241 | 247 | 0.98 | 1853 | 8 | 1853 | 8 | 1853 | 16 | 1933 | 17 | 0.0 |
| K05-100.5-110 | 1403 | 1278 | 1.10 | 2373 | 30 | 1866 | 12 | 1446 | 13 | 1375 | 13 | 43.5 |
| K05-100.5-113 | 362 | 652 | 0.56 | 2472 | 35 | 2424 | 14 | 2367 | 26 | 2353 | 29 | 5.1 |
| K06-107-2 | 161 | 192 | 0.84 | 1721 | 23 | 1727 | 10 | 1732 | 16 | 1749 | 44 | -0.8 |
| K06-107-3 | 155 | 226 | 0.69 | 1723 | 20 | 1724 | 8 | 1725 | 14 | 1766 | 30 | -0.2 |
| K06-107-4 | 104 | 147 | 0.70 | 1739 | 31 | 1738 | 14 | 1737 | 19 | 1651 | 66 | 0.1 |
| K06-107-5 | 126 | 133 | 0.95 | 1720 | 20 | 1713 | 8 | 1709 | 14 | 1742 | 32 | 0.7 |
| K06-107-6 | 101 | 129 | 0.78 | 1741 | 24 | 1733 | 11 | 1726 | 17 | 1740 | 45 | 1 |
| K06-107-7 | 149 | 135 | 1.10 | 1722 | 29 | 1624 | 11 | 1550 | 12 | 1626 | 64 | 11.3 |
| K06-107-8 | 177 | 185 | 0.96 | 1735 | 23 | 1723 | 10 | 1714 | 15 | 1754 | 47 | 1.4 |
| K06-107-9 | 167 | 197 | 0.85 | 1724 | 23 | 1724 | 10 | 1723 | 16 | 1732 | 46 | 0.1 |
| K06-107-10 | 180 | 223 | 0.81 | 1725 | 25 | 1726 | 11 | 1727 | 17 | 1698 | 53 | -0.2 |


| K06-107-12 | 153 | 182 | 0.84 | 1721 | 21 | 1689 | 8 | 1663 | 14 | 1659 | 33 | 3.8 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K05-113-01 | 208 | 673 | 0.31 | 1843 | 9 | 1842 | 9 | 1841 | 17 | 1925 | 36 | 0.1 |
| K05-113-02 | 145 | 616 | 0.24 | 1853 | 10 | 1849 | 11 | 1846 | 18 | 1836 | 49 | 0.4 |
| K05-113-03 | 126 | 429 | 0.29 | 1847 | 12 | 1803 | 12 | 1766 | 16 | 1791 | 67 | 5.0 |
| K05-113-05 | 147 | 551 | 0.27 | 1811 | 14 | 1777 | 14 | 1748 | 20 | 1762 | 74 | 4.0 |
| K05-113-06 | 91 | 307 | 0.29 | 1834 | 9 | 1833 | 9 | 1832 | 16 | 1841 | 36 | 0.1 |
| K05-113-07 | 145 | 474 | 0.31 | 1842 | 10 | 1802 | 11 | 1768 | 18 | 1783 | 51 | 4.6 |
| K05-113-08 | 223 | 627 | 0.35 | 1847 | 11 | 1839 | 11 | 1832 | 18 | 1861 | 57 | 1.0 |
| K05-113-09 | 67 | 328 | 0.20 | 1832 | 9 | 1831 | 10 | 1831 | 15 | 1845 | 52 | 0.0 |
| K05-113-10 | 58 | 285 | 0.20 | 1791 | 26 | 1753 | 22 | 1722 | 23 | 1816 | 150 | 4.4 |
| K05-113-11 | 191 | 692 | 0.28 | 1836 | 11 | 1836 | 11 | 1836 | 17 | 1852 | 63 | 0.0 |
| K05-113-15 | 84 | 336 | 0.25 | 1828 | 14 | 1828 | 14 | 1829 | 20 | 1763 | 77 | 0.0 |
| K05-113-16 | 64 | 344 | 0.19 | 1821 | 39 | 1760 | 14 | 1709 | 18 | 1699 | 19 | 7.0 |
| K05-113-17 | 292 | 864 | 0.34 | 1845 | 12 | 1774 | 11 | 1715 | 16 | 1789 | 65 | 8.0 |
| K05-113-18 | 91 | 369 | 0.25 | 1839 | 27 | 1818 | 9 | 1799 | 15 | 1795 | 16 | 2.5 |
| K05-113-19 | 205 | 617 | 0.33 | 1847 | 20 | 1742 | 18 | 1656 | 21 | 1772 | 116 | 11.7 |
| K05-113-20 | 74 | 280 | 0.26 | 1840 | 9 | 1836 | 9 | 1833 | 16 | 1785 | 40 | 0.5 |
| K05-113-21 | 118 | 502 | 0.23 | 1866 | 14 | 1790 | 14 | 1726 | 19 | 1771 | 76 | 8.5 |
| K05-113-38 | 94 | 351 | 0.27 | 1867 | 18 | 1865 | 17 | 1863 | 22 | 1776 | 97 | 0.2 |
| K05-113-41 | 148 | 571 | 0.26 | 1841 | 10 | 1840 | 10 | 1839 | 17 | 1796 | 50 | 0.1 |
| K05-113-43 | 150 | 578 | 0.26 | 1848 | 9 | 1847 | 9 | 1847 | 15 | 1834 | 48 | 0.0 |
| K05-113-45 | 163 | 599 | 0.27 | 1846 | 10 | 1845 | 10 | 1844 | 17 | 1884 | 46 | 0.1 |
| K05-113-46 | 54 | 247 | 0.22 | 1850 | 11 | 1847 | 11 | 1844 | 18 | 2025 | 59 | 0.3 |
| K05-113-47 | 91 | 361 | 0.25 | 1840 | 10 | 1827 | 10 | 1815 | 16 | 1807 | 51 | 1.6 |
| K05-113-48 | 46 | 219 | 0.21 | 1853 | 17 | 1853 | 15 | 1853 | 19 | 2125 | 108 | 0.0 |
| K06-228.3-1-1 | 414 | 558 | 0.74 | 1737 | 21 | 1730 | 9 | 1724 | 15 | 1850 | 31 | 0.9 |
| K06-228.3-1-5 | 122 | 191 | 0.64 | 1726 | 21 | 1720 | 10 | 1715 | 16 | 1792 | 29 | 0.7 |
| K06-228.3-1-7 | 159 | 200 | 0.79 | 1757 | 21 | 1741 | 9 | 1728 | 16 | 1829 | 27 | 1.8 |
| K06-228.3-1-8 | 596 | 649 | 0.92 | 1733 | 20 | 1733 | 9 | 1732 | 15 | 1798 | 23 | 0.1 |
| K06-228.3-1-12 | 427 | 447 | 0.96 | 1820 | 25 | 1776 | 12 | 1738 | 19 | 1808 | 48 | 5.1 |
| K06-228.3-1-13 | 544 | 491 | 1.11 | 1740 | 20 | 1746 | 9 | 1751 | 15 | 1803 | 24 | -0.7 |
| K06-228.3-1-15 | 1023 | 883 | 1.16 | 1738 | 20 | 1742 | 9 | 1745 | 15 | 1792 | 24 | -0.4 |
| K06-228.3-1-16 | 1092 | 895 | 1.22 | 1735 | 21 | 1720 | 9 | 1707 | 16 | 1735 | 28 | 1.9 |
| K06-228.3-1-18 | 348 | 352 | 0.99 | 1742 | 25 | 1736 | 10 | 1732 | 15 | 1789 | 47 | 0.6 |
| K06-228.3-1-22 | 992 | 846 | 1.17 | 1738 | 21 | 1728 | 10 | 1720 | 16 | 1761 | 32 | 1.2 |
| K06-238-1 | 26 | 77 | 0.34 | 1724 | 23 | 1722 | 10 | 1719 | 16 | 1762 | 36 | 0.3 |
| K06-238-2 | 42 | 144 | 0.29 | 2275 | 45 | 2183 | 19 | 2086 | 24 | 2064 | 27 | 9.7 |
| K06-238-5 | 22 | 70 | 0.31 | 1737 | 25 | 1717 | 11 | 1701 | 16 | 1754 | 42 | 2.3 |
| K06-238-6 | 203 | 871 | 0.23 | 1721 | 23 | 1714 | 10 | 1709 | 15 | 1769 | 40 | 0.8 |


| K06-238-9 | 65 | 141 | 0.46 | 1739 | 22 | 1733 | 9 | 1729 | 14 | 1903 | 40 | 0.7 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K06-238-10 | 103 | 713 | 0.14 | 1730 | 28 | 1744 | 12 | 1755 | 17 | 1857 | 64 | -1.6 |
| K06-238-12 | 68 | 183 | 0.37 | 1728 | 23 | 1716 | 10 | 1705 | 15 | 1787 | 42 | 1.5 |
| K06-238-13Core | 51 | 253 | 0.20 | 1724 | 33 | 1722 | 12 | 1721 | 15 | 1720 | 18 | 0.2 |
| K06-238-13Rim | 85 | 729 | 0.12 | 1731 | 30 | 1581 | 13 | 1472 | 16 | 1474 | 60 | 16.7 |
| K06-238-16 | 141 | 278 | 0.51 | 1801 | 20 | 1802 | 9 | 1803 | 15 | 2033 | 34 | -0.1 |
| K06-238-17 | 131 | 301 | 0.43 | 1731 | 21 | 1752 | 9 | 1769 | 15 | 1859 | 33 | -2.5 |
| K06-238-18 | 111 | 240 | 0.46 | 1739 | 28 | 1740 | 13 | 1740 | 19 | 1818 | 57 | -0.1 |
| K06-238-19 | 233 | 419 | 0.56 | 1849 | 26 | 1872 | 12 | 1894 | 19 | 1941 | 58 | -2.8 |
| K06-238-22 | 53 | 179 | 0.30 | 1859 | 23 | 1848 | 11 | 1838 | 17 | 1861 | 44 | 1.3 |
| K06-238-24 | 453 | 730 | 0.62 | 1740 | 35 | 1746 | 16 | 1752 | 21 | 1937 | 93 | -0.8 |
| K06-238-26 | 260 | 450 | 0.58 | 1730 | 22 | 1733 | 10 | 1735 | 16 | 1835 | 36 | -0.3 |
| K06-238-30 | 170 | 444 | 0.38 | 2453 | 25 | 2472 | 13 | 2496 | 21 | 2618 | 83 | -2.1 |
| K06-238-31 | 83 | 268 | 0.31 | 1761 | 23 | 1765 | 10 | 1769 | 16 | 1848 | 43 | -0.6 |
| K06-238-32 | 195 | 433 | 0.45 | 1760 | 29 | 1776 | 13 | 1789 | 18 | 1871 | 68 | -1.9 |
| K06-238-35 | 208 | 463 | 0.45 | 1749 | 24 | 1753 | 11 | 1757 | 17 | 1925 | 49 | -0.5 |
| K06-238-36 | 156 | 413 | 0.38 | 1733 | 39 | 1734 | 18 | 1735 | 21 | 2044 | 115 | -0.1 |
| K06-245-2-3 | 100 | 421 | 0.24 | 2658 | 36 | 2537 | 16 | 2388 | 25 | 2353 | 30 | 12.1 |
| K06-245-2-4 | 103 | 383 | 0.27 | 1731 | 21 | 1721 | 8 | 1713 | 13 | 1771 | 33 | 1.2 |
| K06-245-2-6 | 202 | 531 | 0.38 | 1768 | 28 | 1580 | 11 | 1443 | 11 | 1513 | 51 | 20.5 |
| K06-245-2-7 | 269 | 862 | 0.31 | 1734 | 27 | 1689 | 12 | 1653 | 18 | 1705 | 53 | 5.3 |
| K06-245-2-11 | 241 | 822 | 0.29 | 1730 | 26 | 1628 | 11 | 1550 | 15 | 1617 | 47 | 11.7 |
| K06-245-2-13 | 403 | 876 | 0.46 | 1778 | 24 | 1738 | 11 | 1705 | 17 | 1780 | 43 | 4.6 |
| K06-245-2-14 | 237 | 644 | 0.37 | 1733 | 22 | 1596 | 8 | 1495 | 11 | 1477 | 32 | 15.4 |
| K06-245-2-15 | 391 | 1004 | 0.39 | 1737 | 27 | 1723 | 13 | 1712 | 19 | 1803 | 57 | 1.6 |
| K06-245-2-16 | 84 | 434 | 0.19 | 1744 | 31 | 1736 | 14 | 1729 | 20 | 1773 | 70 | 1 |
| K06-245-2-17 | 373 | 847 | 0.44 | 1748 | 23 | 1602 | 9 | 1494 | 13 | 1519 | 36 | 16.3 |
| K06-245-2-19 | 144 | 124 | 1.16 | 2645 | 20 | 2613 | 11 | 2571 | 21 | 3031 | 60 | 3.4 |
| K06-245-2-20 | 71 | 623 | 0.11 | 1680 | 33 | 1666 | 11 | 1655 | 16 | 1653 | 18 | 1.7 |
| K06-245-2-24 | 125 | 948 | 0.13 | 1741 | 21 | 1568 | 8 | 1443 | 12 | 1063 | 19 | 19.1 |
| K06-245-2-26 | 157 | 395 | 0.40 | 1738 | 27 | 1685 | 10 | 1642 | 12 | 1638 | 51 | 6.3 |
| K06-245-2-31 | 189 | 580 | 0.33 | 1747 | 22 | 1639 | 9 | 1556 | 13 | 1546 | 31 | 12.2 |
| K06-245-2-34 | 117 | 368 | 0.32 | 1731 | 29 | 1558 | 11 | 1433 | 12 | 1324 | 48 | 19.2 |
| K06-245-2-36 | 158 | 345 | 0.46 | 2454 | 21 | 2449 | 12 | 2443 | 24 | 2534 | 56 | 0.6 |
| K06-245-2-38 | 85 | 237 | 0.36 | 1750 | 27 | 1644 | 10 | 1563 | 12 | 1438 | 43 | 12 |

Table A 6: Hf isotopic data for Grand Canyon plutons (GEMOC)

| Analysis No. | ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}$ | 1 se | ${ }^{176} \mathrm{Lu} /{ }^{177} \mathrm{Hf}$ | ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}$ <br> ( T ) | E Hf (0) | Epsilon Hf <br> ( T ) | 1 sigma | Age |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K05-100.5-1Rim | 0.281768 | 0.000008 | 0.000519 | 0.281751 | -36.0 | 2.6 | 0.6 | 1739 |
| K05-100.5-2Core | 0.281134 | 0.000009 | 0.000679 | 0.281101 | -58.4 | -1.8 | 0.6 | 2547 |
| K05-100.5-3 | 0.281192 | 0.000009 | 0.000666 | 0.281161 | -56.3 | -1.3 | 0.6 | 2475 |
| K05-100.5-4 | 0.280895 | 0.000009 | 0.000843 | 0.280848 | -66.8 | -1.7 | 0.7 | 2936 |
| K05-100.5-5 | 0.281187 | 0.000011 | 0.000640 | 0.281157 | -56.5 | -1.4 | 0.8 | 2476 |
| K05-100.5-6 | 0.281806 | 0.000009 | 0.001013 | 0.281773 | -34.6 | 3.2 | 0.7 | 1732 |
| K05-100.5-7 | 0.281767 | 0.000008 | 0.000704 | 0.281744 | -36.0 | 2.6 | 0.6 | 1749 |
| K05-100.5-8 | 0.281196 | 0.000010 | 0.000563 | 0.281169 | -56.2 | -1.1 | 0.7 | 2469 |
| K05-100.5-9 | 0.281195 | 0.000009 | 0.000723 | 0.281161 | -56.2 | -1.4 | 0.7 | 2470 |
| K05-100.5-12 | 0.281168 | 0.000008 | 0.000576 | 0.281141 | -57.2 | -1.5 | 0.6 | 2498 |
| K05-100.5-19 | 0.280887 | 0.000009 | 0.000744 | 0.280848 | -67.1 | -5.9 | 0.6 | 2756 |
| K05-100.5-39 | 0.281802 | 0.000010 | 0.002025 | 0.281735 | -34.8 | 1.9 | 0.8 | 1734 |
| K05-100.5-54 | 0.281799 | 0.000006 | 0.001065 | 0.281764 | -34.9 | 2.9 | 0.5 | 1731 |
| K05-100.5-57 | 0.281209 | 0.000007 | 0.000493 | 0.281186 | -55.7 | -0.2 | 0.5 | 2487 |
| K05-100.5-58 | 0.281257 | 0.000010 | 0.001673 | 0.281180 | -54.0 | -2.4 | 0.8 | 2397 |
| K05-100.5-71 | 0.281601 | 0.000007 | 0.000834 | 0.281573 | -41.9 | -2.8 | 0.5 | 1778 |
| K05-100.5-75 | 0.281188 | 0.000011 | 0.000580 | 0.281160 | -56.5 | -0.7 | 0.8 | 2504 |
| K05-100.5-82 | 0.281191 | 0.000009 | 0.000698 | 0.281158 | -56.4 | -1.6 | 0.6 | 2468 |
| K05-100.5-89 | 0.281821 | 0.000009 | 0.001555 | 0.281769 | -34.1 | 3.7 | 0.7 | 1758 |
| K05-100.5-95 | 0.281383 | 0.000012 | 0.002979 | 0.281246 | -49.6 | 0.4 | 0.8 | 2417 |
| K05-100.5-97 | 0.281900 | 0.000010 | 0.002843 | 0.281807 | -31.3 | 4.4 | 0.7 | 1731 |
| K05-100.5-100 | 0.281225 | 0.000009 | 0.000711 | 0.281191 | -55.2 | -0.3 | 0.6 | 2473 |
| K05-100.5-101 | 0.281303 | 0.000008 | 0.000465 | 0.281287 | -52.4 | -11.3 | 0.6 | 1853 |
| K05-100.5-110 | 0.281275 | 0.000010 | 0.001983 | 0.281185 | -53.4 | -2.8 | 0.7 | 2373 |
| K05-100.5-113 | 0.281314 | 0.000009 | 0.002261 | 0.281207 | -52.0 | 0.3 | 0.7 | 2472 |
| K06-107-2 | 0.281942 | 0.000016 | 0.000547 | 0.281924 | -29.8 | 8.3 | 0.9 | 1721 |
| K06-107-3 | 0.281915 | 0.000011 | 0.000729 | 0.281891 | -30.8 | 7.2 | 0.6 | 1723 |
| K06-107-4 | 0.281969 | 0.000014 | 0.000443 | 0.281955 | -28.9 | 9.8 | 0.8 | 1739 |
| K06-107-5 | 0.281948 | 0.000010 | 0.000437 | 0.281934 | -29.6 | 8.7 | 0.6 | 1720 |
| K06-107-6 | 0.281941 | 0.000011 | 0.000482 | 0.281925 | -29.8 | 8.8 | 0.6 | 1741 |
| K06-107-7 | 0.281980 | 0.000014 | 0.000663 | 0.281958 | -28.5 | 9.6 | 0.8 | 1722 |
| K06-107-8 | 0.281966 | 0.000011 | 0.000554 | 0.281948 | -29.0 | 9.5 | 0.6 | 1735 |
| K06-107-9 | 0.281940 | 0.000013 | 0.000512 | 0.281923 | -29.9 | 8.4 | 0.7 | 1724 |
| K06-107-10 | 0.281958 | 0.000019 | 0.000567 | 0.281939 | -29.2 | 9.0 | 0.9 | 1725 |
| K06-107-12 | 0.281942 | 0.000011 | 0.000459 | 0.281927 | -29.8 | 8.4 | 0.6 | 1721 |
| K05-113-1 | 0.281954 | 0.000012 | 0.001725 | 0.281894 | -29.4 | 10.1 | 0.8 | 1843 |
| K05-113-2 | 0.281969 | 0.000008 | 0.002175 | 0.281893 | -28.9 | 10.3 | 0.5 | 1853 |
| K05-113-3 | 0.282057 | 0.000017 | 0.003080 | 0.281949 | -25.7 | 12.1 | 1.2 | 1847 |


| K05-113-5 | 0.281965 | 0.000015 | 0.002119 | 0.281892 | -29.0 | 9.3 | 1.3 | 1811 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K05-113-6 | 0.281967 | 0.000009 | 0.002032 | 0.281896 | -28.9 | 10.0 | 0.6 | 1834 |
| K05-113-7 | 0.282011 | 0.000008 | 0.003342 | 0.281894 | -27.4 | 10.1 | 0.6 | 1842 |
| K05-113-8 | 0.282113 | 0.000016 | 0.003846 | 0.281978 | -23.8 | 13.2 | 1.2 | 1847 |
| K05-113-9 | 0.281964 | 0.000009 | 0.002251 | 0.281886 | -29.0 | 9.5 | 0.6 | 1832 |
| K05-113-10 | 0.282005 | 0.000009 | 0.002145 | 0.281932 | -27.6 | 10.2 | 0.7 | 1791 |
| K05-113-11 | 0.282000 | 0.000010 | 0.002302 | 0.281920 | -27.8 | 10.8 | 0.7 | 1836 |
| K05-113-15 | 0.281982 | 0.000008 | 0.002253 | 0.281904 | -28.4 | 10.1 | 0.6 | 1828 |
| K05-113-16 | 0.281911 | 0.000012 | 0.001224 | 0.281869 | -30.9 | 8.7 | 1.0 | 1821 |
| K05-113-17 | 0.282067 | 0.000015 | 0.003010 | 0.281962 | -25.4 | 12.5 | 1.2 | 1845 |
| K05-113-18 | 0.281959 | 0.000011 | 0.002051 | 0.281887 | -29.2 | 9.8 | 0.7 | 1839 |
| K05-113-19 | 0.282080 | 0.000018 | 0.003978 | 0.281941 | -24.9 | 11.8 | 1.2 | 1847 |
| K05-113-20 | 0.281946 | 0.000012 | 0.001960 | 0.281878 | -29.7 | 9.4 | 0.7 | 1840 |
| K05-113-21 | 0.281947 | 0.000011 | 0.001724 | 0.281886 | -29.6 | 10.3 | 0.8 | 1866 |
| K05-113-38 | 0.282002 | 0.000012 | 0.003217 | 0.281888 | -27.7 | 10.4 | 0.9 | 1867 |
| K05-113-41 | 0.281969 | 0.000011 | 0.002407 | 0.281885 | -28.9 | 9.7 | 0.8 | 1841 |
| K05-113-43 | 0.282013 | 0.000011 | 0.003169 | 0.281902 | -27.3 | 10.5 | 0.7 | 1848 |
| K05-113-45 | 0.281992 | 0.000015 | 0.002161 | 0.281916 | -28.0 | 10.9 | 0.9 | 1846 |
| K05-113-46 | 0.281994 | 0.000016 | 0.002302 | 0.281913 | -28.0 | 10.9 | 1.1 | 1850 |
| K05-113-47 | 0.282033 | 0.000014 | 0.002679 | 0.281939 | -26.6 | 11.6 | 1.0 | 1840 |
| K05-113-48 | 0.282022 | 0.000017 | 0.003883 | 0.281885 | -27.0 | 10.0 | 1.2 | 1853 |
| K06-228.3-1-01 | 0.281968 | 0.000011 | 0.001576 | 0.281916 | -28.9 | 8.4 | 0.6 | 1737 |
| K06-228.3-1-05 | 0.281985 | 0.000020 | 0.001399 | 0.281939 | -28.3 | 9.0 | 1.0 | 1726 |
| K06-228.3-1-07 | 0.281958 | 0.000012 | 0.001299 | 0.281915 | -29.2 | 8.8 | 0.6 | 1757 |
| K06-228.3-1-08 | 0.281875 | 0.000011 | 0.001430 | 0.281828 | -32.2 | 5.2 | 0.6 | 1733 |
| K06-228.3-1-12 | 0.282013 | 0.000016 | 0.002845 | 0.281915 | -27.3 | 10.3 | 0.8 | 1820 |
| K06-228.3-1-13 | 0.282008 | 0.000015 | 0.003196 | 0.281902 | -27.5 | 8.0 | 0.8 | 1740 |
| K06-228.3-1-15 | 0.282035 | 0.000018 | 0.002188 | 0.281963 | -26.5 | 10.1 | 0.9 | 1738 |
| K06-228.3-1-16 | 0.281900 | 0.000013 | 0.001310 | 0.281857 | -31.3 | 6.3 | 0.7 | 1735 |
| K06-228.3-1-18 | 0.281869 | 0.000019 | 0.001574 | 0.281817 | -32.4 | 5.0 | 1.0 | 1742 |
| K06-228.3-1-22 | 0.282060 | 0.000014 | 0.003377 | 0.281948 | -25.6 | 9.6 | 0.8 | 1738 |
| K06-238-01 | 0.282010 | 0.000012 | 0.001099 | 0.281974 | -27.4 | 10.2 | 0.4 | 1724 |
| K06-238-02 | 0.281262 | 0.000010 | 0.000665 | 0.281233 | -53.9 | -3.4 | 0.4 | 2275 |
| K06-238-05 | 0.282020 | 0.000014 | 0.001346 | 0.281976 | -27.1 | 10.5 | 0.5 | 1737 |
| K06-238-06 | 0.281768 | 0.000013 | 0.000232 | 0.281760 | -36.0 | 2.5 | 0.5 | 1721 |
| K06-238-09 | 0.282022 | 0.000012 | 0.002446 | 0.281941 | -27.0 | 9.4 | 0.4 | 1739 |
| K06-238-10 | 0.281740 | 0.000011 | 0.000110 | 0.281736 | -37.0 | 1.9 | 0.4 | 1730 |
| K06-238-12 | 0.281961 | 0.000013 | 0.000929 | 0.281931 | -29.1 | 8.7 | 0.5 | 1728 |
| K06-238-13Core | 0.281898 | 0.000020 | 0.001382 | 0.281851 | -31.4 | 5.8 | 0.7 | 1724 |
| K06-238-13Rim | 0.281873 | 0.000020 | 0.001229 | 0.281833 | -32.3 | 5.3 | 0.7 | 1731 |
| K06-238-16 | 0.281979 | 0.000015 | 0.001634 | 0.281923 | -28.5 | 10.2 | 0.5 | 1801 |


| K06-238-17 | 0.281974 | 0.000011 | 0.001627 | 0.281921 | -28.7 | 8.5 | 0.4 | 1731 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K06-238-18 | 0.281925 | 0.000013 | 0.001209 | 0.281885 | -30.4 | 7.4 | 0.5 | 1739 |
| K06-238-19 | 0.281645 | 0.000010 | 0.000717 | 0.281620 | -40.3 | 0.5 | 0.4 | 1849 |
| K06-238-22 | 0.281571 | 0.000022 | 0.000269 | 0.281562 | -42.9 | -1.4 | 0.8 | 1859 |
| K06-238-24 | 0.282025 | 0.000022 | 0.002995 | 0.281925 | -26.9 | 8.8 | 0.8 | 1740 |
| K06-238-26 | 0.281916 | 0.000013 | 0.001002 | 0.281883 | -30.7 | 7.1 | 0.5 | 1730 |
| K06-238-30 | 0.281216 | 0.000011 | 0.000637 | 0.281186 | -55.5 | -0.9 | 0.4 | 2453 |
| K06-238-31 | 0.281671 | 0.000010 | 0.000977 | 0.281638 | -39.4 | -0.9 | 0.4 | 1761 |
| K06-238-32 | 0.281629 | 0.000015 | 0.001321 | 0.281584 | -40.9 | -2.8 | 0.5 | 1760 |
| K06-238-35 | 0.281890 | 0.000017 | 0.001538 | 0.281839 | -31.6 | 6.0 | 0.6 | 1749 |
| K06-238-36 | 0.281890 | 0.000017 | 0.001191 | 0.281851 | -31.6 | 6.0 | 0.6 | 1733 |
| K06-245-2-3 | 0.281106 | 0.000015 | 0.001992 | 0.281005 | -59.4 | -2.6 | 0.5 | 2658 |
| K06-245-2-4 | 0.281893 | 0.000013 | 0.000966 | 0.281861 | -31.5 | 6.3 | 0.5 | 1731 |
| K06-245-2-6 | 0.281890 | 0.000009 | 0.000871 | 0.281861 | -31.6 | 7.2 | 0.3 | 1768 |
| K06-245-2-7 | 0.281849 | 0.000011 | 0.000980 | 0.281817 | -33.1 | 4.8 | 0.4 | 1734 |
| K06-245-2-11 | 0.281907 | 0.000010 | 0.000690 | 0.281884 | -31.0 | 7.1 | 0.4 | 1730 |
| K06-245-2-13 | 0.281739 | 0.000013 | 0.000966 | 0.281707 | -37.0 | 2.0 | 0.5 | 1778 |
| K06-245-2-14 | 0.281864 | 0.000014 | 0.001410 | 0.281818 | -32.6 | 4.8 | 0.5 | 1733 |
| K06-245-2-15 | 0.281907 | 0.000014 | 0.000779 | 0.281881 | -31.0 | 7.2 | 0.5 | 1737 |
| K06-245-2-16 | 0.281845 | 0.000014 | 0.000466 | 0.281830 | -33.2 | 5.5 | 0.5 | 1744 |
| K06-245-2-17 | 0.281894 | 0.000012 | 0.001113 | 0.281857 | -31.5 | 6.6 | 0.4 | 1748 |
| K06-245-2-19 | 0.281368 | 0.000018 | 0.000429 | 0.281337 | -50.1 | 8.9 | 0.6 | 2645 |
| K06-245-2-20 | 0.281876 | 0.000011 | 0.000604 | 0.281850 | -32.1 | 4.8 | 0.4 | 1680 |
| K06-245-2-24 | 0.281824 | 0.000009 | 0.000390 | 0.281811 | -34.0 | 4.8 | 0.3 | 1741 |
| K06-245-2-26 | 0.281887 | 0.000016 | 0.000936 | 0.281856 | -31.8 | 6.3 | 0.6 | 1738 |
| K06-245-2-31 | 0.281891 | 0.000009 | 0.000852 | 0.281863 | -31.6 | 6.8 | 0.3 | 1747 |
| K06-245-2-34 | 0.281821 | 0.000014 | 0.000897 | 0.281791 | -34.1 | 3.9 | 0.5 | 1731 |
| K06-245-2-36 | 0.281257 | 0.000013 | 0.000859 | 0.281217 | -54.0 | 0.2 | 0.5 | 2454 |
| K06-245-2-38 | 0.281830 | 0.000011 | 0.000396 | 0.281817 | -33.8 | 5.2 | 0.4 | 1750 |

## APPENDIX B

## Geochemical Evolution and Metallogeny of Continents (GEMOC) Key Centre

## Methods

## U-Pb Geochronology

U-Pb geochronology at the GEMOC Key Centre was conducted in-situ using an HP 4500 inductively coupled plasma quadrupole mass spectrometer (ICP-MS) paired with a custom-made UV laser ablation microprobe (LAM) that incorporates a petrographic microscope for detailed sample scrutiny (Norman et al., 1996). Samples and standards were ablated also in a custom-made chamber and transported to the ICP-MS with He carrier gas in order to minimize $\mathrm{U} / \mathrm{Pb}$ fractionation. In addition, the laser was focused above the sample in order to further minimize fractionation effects; laser conditions were rigorously maintained throughout the duration of sample analysis.

Samples were compared to the zircon standard 02123, with four standard analyses completed before and after every 12 unknowns. The 02123 standard is a gem quality zircon from a Norwegian syenite that yields a perfectly concordant ID-TIMS age of 295 $\pm 1 \mathrm{Ma}$ (Ketchum et al., 2001). Isotope ratios for both standards and unknowns are determined from background-subtracted signals; the uncertainties in both the background and signal are added in quadrature.

Masses 206, 207, 208, 232, and 238 were measured, and all isotopic ratios were calculated using the in-house on-line data reduction software GLITTER. Mass 204 was not measured due to large isobaric interference from Hg . Common Pb correction was therefore conducted after Andersen (2002), using ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U},{ }^{207} \mathrm{~Pb} /{ }^{235} \mathrm{U}$, and ${ }^{208} \mathrm{~Pb} /{ }^{232} \mathrm{Th}$
ratios to solve mass-balance equations and correct the data in three-dimensional concordia space.

For additional details regarding analytical methods see Belousova et al. (2001), Griffin et al. (2004), and Jackson et al. (2004).

## Hf Isotopes

Hf-isotope analyses at GEMOC were carried out in-situ using a New Wave/Merchantek UP-213 laser-ablation microprobe, attached to a Nu Plasma multicollector ICPMS. The analyses were carried out with a beam diameter of ca $55 \mu \mathrm{~m}$ and a 5 Hz repetition rate. This resulted in total Hf signals of 1-6 $\times 10^{-11} \mathrm{~A}$, depending on conditions and the Hf contents. Typical ablation times were 100-120 seconds, resulting in pits $40-60 \mu \mathrm{~m}$ deep. He carrier gas transported the ablated sample from the laser-ablation cell via a mixing chamber to the ICPMS torch.

Interference of ${ }^{176} \mathrm{Lu}$ on ${ }^{176} \mathrm{Hf}$ is corrected by measuring the intensity of the interference-free ${ }^{175} \mathrm{Lu}$ isotope and using ${ }^{176} \mathrm{Lu} /{ }^{175} \mathrm{Lu}=0.02669$ (DeBievre \& Taylor 1993) to calculate ${ }^{176} \mathrm{Lu} /{ }^{177} \mathrm{Hf}$. Similarly, the interference of ${ }^{176} \mathrm{Yb}$ on ${ }^{176} \mathrm{Hf}$ has been corrected by measuring the interference-free ${ }^{172} \mathrm{Yb}$ isotope and using ${ }^{176} \mathrm{Yb} /{ }^{172} \mathrm{Yb}$ to calculate ${ }^{176} \mathrm{Yb} /{ }^{177} \mathrm{Hf}$. The appropriate value of ${ }^{176} \mathrm{Yb} /{ }^{172} \mathrm{Yb}$ was determined by spiking the JMC475 Hf standard with Yb , and finding the value of ${ }^{176} \mathrm{Yb} /{ }^{172} \mathrm{Yb}(0.58669)$ required to yield the value of $176 \mathrm{Hf} / 177 \mathrm{Hf}$ obtained on the pure Hf solution. Detailed discussions regarding the overlap corrections for ${ }^{176} \mathrm{Lu}$ and ${ }^{176} \mathrm{Yb}$ are provided in Pearson et al. (2008). Analyses of standard zircons (Griffin et al., 2000; Pearson et al., 2008) illustrate the precision and accuracy obtainable on the ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}$ ratio, despite the severe
corrections on ${ }^{176} \mathrm{Hf}$. The typical 2 SE precision on the ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}$ ratios presented here is $\pm 0.00002$, equivalent to $\pm 0.7 \varepsilon \mathrm{Eff}$ unit.

The Mud Tank and 91500 zircon standards, analyzed together with the samples, were used as independent control on reproducibility and instrument stability. Average ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}$ values obtained for the Mud Tank $(0.282523 \pm 0.000066)$ and 91500 $(0.282299 \pm 0.000042)$ during this study are similar to the long-term averages, which in turn are similar to the TIMS values (Griffin et al., 2006, 2007).

For the calculation of $\varepsilon H f$ values, we have adopted the chondritic values of Bouvier et al. (2008): ${ }^{176} \mathrm{Lu} /{ }^{177} \mathrm{Hf}\left(\mathrm{CHUR}\right.$, today) $=0.0336$ and ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}(\mathrm{CHUR}$, today $)=0.282785$. For the calculation of $\varepsilon H f$ values, we have adopted the decay constant $\left(1.867 \times 10-11 \mathrm{yr}^{-1}\right)$ for ${ }^{176} \mathrm{Lu}$ proposed by Scherer et al. (2001) because it gives the best fit for terrestrial rocks (Amelin and Davis, 2005; Albarède et al., 2006).

For additional details regarding analytical methods see Griffin et al. (2000, 2002, 2004).

## Arizona Laserchron Center (ALC) Methods

## U-Pb Geochronology

U-Pb geochronology of zircons is conducted in-situ by laser ablation multicollector inductively coupled plasma mass spectrometry (LA-MC-ICPMS). The analyses involve ablation of zircon with a New Wave UP193HE Excimer laser prior to May 2011, and afterwards a Photon Machines Analyte G2 excimer laser using a spot diameter of 30 microns. The ablated material is carried in helium into the plasma source of a Nu HR ICPMS, which is equipped with a flight tube of sufficient width that $\mathrm{U}, \mathrm{Th}$, and Pb isotopes are measured simultaneously. All measurements are made in static mode,
using Faraday detectors with $3 \times 10^{11}$ ohm resistors for ${ }^{238} \mathrm{U},{ }^{232} \mathrm{Th},{ }^{208} \mathrm{~Pb}-{ }^{206} \mathrm{~Pb}$, and discrete dynode ion counters for ${ }^{204} \mathrm{~Pb}$ and ${ }^{202} \mathrm{Hg}$. Ion yields are $\sim 0.8 \mathrm{mv}$ per ppm. Each analysis consists of one 15 -second integration on peaks with the laser off (for backgrounds), 15 one-second integrations with the laser firing, and a 30 second delay to purge the previous sample and prepare for the next analysis. The resulting ablation pit is $\sim 15$ microns in depth.

For each analysis, the errors in determining ${ }^{206} \mathrm{~Pb} / 238 \mathrm{U}$ and ${ }^{206} \mathrm{~Pb} /{ }^{204} \mathrm{~Pb}$ result in a measurement error of $\sim 1-2 \%$ (at 2-sigma level) in the ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ age. The errors in measurement of ${ }^{206} \mathrm{~Pb} /{ }^{207} \mathrm{~Pb}$ and ${ }^{206} \mathrm{~Pb} /{ }^{204} \mathrm{~Pb}$ also result in $\sim 1-2 \%$ (at 2-sigma level) uncertainty in age for grains that are $>1.0 \mathrm{Ga}$, but are substantially larger for younger grains due to low intensity of the ${ }^{207} \mathrm{~Pb}$ signal. For most analyses, the cross-over in precision of ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ and ${ }^{206} \mathrm{~Pb} /{ }^{207} \mathrm{~Pb}$ ages occurs at $\sim 1.0 \mathrm{Ga}$.
${ }^{204} \mathrm{Hg}$ interference with ${ }^{204} \mathrm{~Pb}$ is accounted for measurement of ${ }^{202} \mathrm{Hg}$ during laser ablation and subtraction of ${ }^{204} \mathrm{Hg}$ according to the natural ${ }^{202} \mathrm{Hg} /{ }^{204} \mathrm{Hg}$ of 4.35. This Hg is correction is not significant for most analyses because our Hg backgrounds are low (generally $\sim 150 \mathrm{cps}$ at mass 204).

Common Pb correction is accomplished by using the Hg -corrected ${ }^{204} \mathrm{~Pb}$ and assuming an initial Pb composition from Stacey and Kramers (1975). Uncertainties of 1.5 for ${ }^{206} \mathrm{~Pb} /{ }^{204} \mathrm{~Pb}$ and 0.3 for ${ }^{207} \mathrm{~Pb} /{ }^{204} \mathrm{~Pb}$ are applied to these compositional values based on the variation in Pb isotopic composition in modern crystal rocks.

Inter-element fractionation of $\mathrm{Pb} / \mathrm{U}$ is generally $\sim 5 \%$, whereas apparent fractionation of Pb isotopes is generally $<0.2 \%$. In-run analysis of fragments of a large zircon crystal (generally every fifth measurement) with known age of $563.5 \pm 3.2 \mathrm{Ma}$ (2-
sigma error) is used to correct for this fractionation. The uncertainty resulting from the calibration correction is generally $1-2 \%$ (2-sigma) for both ${ }^{206} \mathrm{~Pb} /{ }^{207} \mathrm{~Pb}$ and ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ ages.

For additional details regarding analytical methods see Gehrels et al. (2006, 2008), and Gehrels and Pecha (2014).

## Hf Isotopes

Hf isotope analyses are conducted with a Nu HR ICPMS connected to a New Wave UP193HE laser (2009-2010) or a Photon Machines Analyte G2 excimer laser (2011). Instrument settings are established first by analysis of 10 ppb solutions of JMC475 and a Spex Hf solution, and then by analysis of 10 ppb solutions containing Spex $\mathrm{Hf}, \mathrm{Yb}$, and Lu . The mixtures range in concentration of Yb and Lu , with ${ }^{176}(\mathrm{Yb}+\mathrm{Lu})$ up to $70 \%$ of the ${ }^{176} \mathrm{Hf}$. When all solutions yield ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}$ of $\sim 0.28216$, instrument settings are optimized for laser ablation analyses and seven different standard zircons (Mud Tank, 91500, Temora, R33, FC52, Plesovice, and Sri Lanka) are analyzed. These standards are included with unknowns on the same epoxy mounts. When precision and accuracy are acceptable, unknowns are analyzed using exactly the same acquisition parameters.

Laser ablation analyses are conducted with a laser beam diameter of 40 microns, with the ablation pits located on top of the $\mathrm{U}-\mathrm{Pb}$ analysis pits. CL images are used to ensure that the ablation pits do not overlap multiple age domains or inclusions. Each acquisition consists of one 40 -second integration on backgrounds (on peaks with no laser firing) followed by 60 one-second integrations with the laser firing. Using a typical laser
fluence of $\sim 5 \mathrm{~J} / \mathrm{cm} 2$ and pulse rate of 7 hz , the ablation rate is $\sim 0.8$ microns per second. Each standard is analyzed once for every $\sim 20$ unknowns.

Isotope fractionation is accounted for using the method of Woodhead et al. (2004): $\beta \mathrm{Hf}$ is determined from the measured ${ }^{179} \mathrm{Hf} /{ }^{177} \mathrm{Hf} ; \beta \mathrm{Yb}$ is determined from the measured ${ }^{173} \mathrm{Yb} /{ }^{171} \mathrm{Yb}$ (except for very low Yb signals); $\beta \mathrm{Lu}$ is assumed to be the same as $\beta \mathrm{Yb}$; and an exponential formula is used for fractionation correction. Yb and Lu interferences are corrected by measurement of ${ }^{176} \mathrm{Yb} /{ }^{171} \mathrm{Yb}$ and ${ }^{176} \mathrm{Lu} /{ }^{175} \mathrm{Lu}$ (respectively), as advocated by Woodhead et al. (2004). Critical isotope ratios are ${ }^{179} \mathrm{Hf} /{ }^{177} \mathrm{Hf}=0.73250$ (Patchett \& Tatsumoto, 1980); ${ }^{173} \mathrm{Yb} /{ }^{171} \mathrm{Yb}=1.132338$ (Vervoort et al. 2004); ${ }^{176} \mathrm{Yb} /{ }^{171} \mathrm{Yb}=0.901691$ (Vervoort et al., 2004; Amelin and Davis, 2005), ${ }^{176} \mathrm{Lu} /{ }^{175} \mathrm{Lu}=$ 0.02653 (Patchett, 1983). All corrections are done line-by-line. For very low Yb signals, $\beta \mathrm{Hf}$ is used for fractionation of Yb isotopes. The corrected ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}$ values are filtered for outliers (2-sigma filter), and the average and standard error are calculated from the resulting $\sim 58$ integrations. There is no capability to use only a portion of the acquired data.

All solutions, standards, and unknowns analyzed during a session are reduced together such that unknown values are calibrated based on the standards analyzed during the same session. The most weight is put on the following standards: Temore2, 91500, Mud Tank, FC1, and Plesovice, with less reliance placed on R33. The cutoff for using $\beta \mathrm{Hf}$ versus $\beta \mathrm{Yb}$ is determined by monitoring the average offset of the standards from their known values, and the cutoff is set at the minimum offset. For most data sets, this is achieved at $\sim 6 \mathrm{mv}$ of ${ }^{171} \mathrm{Yb}$. For sessions in which the standards yield ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}$ values that are shifted consistently from the know values, a correction factor is applied to the
${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}$ of all standards and unknowns. This correction factor is generally less than 1 epsilon unit. For example: all values were increased by 1.0 epsilon units for samples analyzed in August 2012, all values increased by 0.3 epsilon units for samples analyzed in May 2013, all values were decreased by 0.2 epsilon units for samples analyzed in May 2014.

The ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}$ at time of crystallization is calculated from measurement of present-day ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}$ and ${ }^{176} \mathrm{Lu} /{ }^{177} \mathrm{Hf}$, using the decay constant of ${ }^{176} \mathrm{Lu}(\lambda=1.867 \times$ $10^{-11}$ ) from Scherer et al. (2001) and Söderlund et al. (2004). No capability is provided for calculating Hf Depleted Mantle model ages because the ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}$ and ${ }^{176} \mathrm{Lu} /{ }^{177} \mathrm{Hf}$ of the source material(s) from which the zircon crystallized is not known.

For additional details regarding analytical methods see Cecil et al. (2011), and Gehrels and Pecha (2014).

## APPENDIX C

## Pluton Descriptions by Sample

## Grapevine Camp Pluton (K12-81L):

The $1737 \pm 1$ Ma Grapevine camp pluton is a medium grained, biotite granite exposed at river mile 81 as a narrow body aligned with $S_{2}$ foliation (Hawkins et al., 1996). Truncated on its western margin by the Vishnu fault zone, the Grapevine Camp pluton is marks the edge of the Mineral Canyon block. The Vishnu fault zone is the first of several thermal boundaries juxtaposing blocks with peak temperatures that vary up to $200{ }^{\circ} \mathrm{C}$, but remarkably constant pressures of $0.6-0.7 \mathrm{GPa}$ (Ilg et al., 1996; Dumond et al., 2007). The Mineral Canyon block is characterized by peak temperatures of $735 \pm 135$ ${ }^{\circ} \mathrm{C}$ (Dumond et al., 2007). The pluton preserves an annealed mylonitic fabric defined by ribbons of quartz and feldspar rich layers $\left(\mathrm{S}_{1}\right)$ that is cut by the Phanerozoic Vishnu fault (Huntoon et al., 1980), but the last ductile motion on the shear zone is inferred to be between $1.68-1.3 \mathrm{Ga}$ (Ilg et al., 1996). The eastern edge of the pluton displays clear intrusive relations to the Vishnu schist on its eastern side (Hawkins et al., 1996; Ilg et al., 1996). The age of the pluton and the presence of early gneissic layering suggests that the Grapevine Camp pluton is part of the $1.74-1.71 \mathrm{Ga}$ suite of arc-plutons, however it is compositionally similar to the younger granite-pegmatite anatectic suite.

Hawkins et al. (1996) described the external morphology an appearance of zircons separated from the Grapevine Camp pluton as slightly rounded and colorless, with pitted crystal faces and ubiquitous fractures and dusty inclusions. Our observations show similar features; zircons from the Grapevine Camp pluton range in size from long axes of approximately $70-200 \mu \mathrm{~m}$ with aspect ratios from $1: 1$ to $2: 1$. Grains are predominantly
rounded and subhedral. Internal textures revealed by CL imaging show many blotchy CL-dark regions disrupting, and in some cases completely obliterating, igneous growth zoning.

A total of 18 grains from the Grapevine Camp pluton were analyzed, with 21 total analyses conducted; 3 attempts were made to analyses distinct zircon domains, but those analyses did not pass the data reduction process. Of the 21 analyses, only 6 yielded satisfactory results. Most of the analyses were discarded due to high common Pb . Several grains yielded high U and Th concentrations ( $>1000 \mathrm{ppm}$ ), but all analyses yielded $\mathrm{U} / \mathrm{Th}$ ratios between $0.8-2.6$. No systematic variation was observed between age, U concentration, U/Th ratios, or concordance.

Hawkins et al. (1996) interpreted the external features of these zircons to mean that they were partially resorbed during magmatism, recrystallization, or deformation. The internal textures revealed by our CL provide additional evidence for the alteration of Grapevine Camp pluton zircons. Quenching of CL is an effect of radiation damage (Geisler et al., 2001; Nasdala et al., 2002, 2003), thus the morphology, internal texture, and U and Th concentrations of these grains are consistent with extensive metamictization. The lack of reliable ages obtained from these zircons and common Pb contamination is therefore not surprising. Nevertheless, 6 grains yielded $>95 \%$ concordant ages with a weighted mean of $1756 \pm 16 \mathrm{Ma}$ (MSWD $=0.3$ ). This age is in contrast with the age of $1737 \pm 1$ Ma reported by Hawkins et al. (1996), and the depositional constraints of $1750-1740$ Ma for the Grand Canyon Metamorphic Suite, but is within uncertainty of the latter. Given the high degree of metamictization of these grains, it is likely that some open-system behavior is influencing the age determination.

While the Hf data presented herein is linked to these ages, we defer to the $1737 \pm 1 \mathrm{Ma}$ reported by Hawkins et al. (1996) for the true crystallization age of the Grapevine Camp pluton.

Of the 6 Grapevine Camp pluton zircons that yielded reliable $\mathrm{U}-\mathrm{Pb}$ ages, we obtained Hf isotopic data from 4 of them. $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values ranged from 6.3 to 10.4 . The $\varepsilon \mathrm{Hf}$ of the depleted mantle at 1756 Ma is 10.1 . Three of the four analyses are indistinguishable from the depleted mantle $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$, and the fourth is still juvenile at 3.8 epsilon units below depleted mantle (Bahlburg et al., 2011).

## Zoroaster Pluton (K12-85.3L):

The $1740 \pm 2 \mathrm{Ma}$ Zoroaster pluton is a medium grained biotite granite to granodiorite orthogneiss (Hawkins et al., 1996). Well-documented intrusive relationships into the Grand Canyon Metamorphic Suite include screens of Vishnu schist that are entrained in the pluton (Figure 12). Localized mafic domains suggest some degree of mamga mixing (Figure 12). The compositional layering of the pluton and adjacent schists define a kilometer-scale $\mathrm{F}_{2}$ plunging $44^{\circ}$ towards $226^{\circ}$ (Lingley, 1973; Ilg et al., 1996). The Zoroaster pluton lies in the Clear Creek metamorphic block, which is characterized by temperatures from $518-552 \pm 110{ }^{\circ} \mathrm{C}$.

The morphology of Zoroaster pluton zircons as described by Hawkins et al. (1996) are doubly terminated and prismatic. Our observations consist of predominantly rounded and subhedral, with a few euhedral grains. Zircons range in size from long axes of approximately $50-250 \mu \mathrm{~m}$ with aspect ratios from 1:1 to $2: 1$. Internal textures revealed by CL imaging show many blotchy CL-dark regions disrupting, and in some
cases completely obliterating, igneous growth zoning. Some grains preserve pristine igneous growth zoning.

A total of 25 grains from the Zoroaster pluton were analyzed, with 22 grains passing the data reduction process. 2 analyses were discarded due to high ${ }^{204} \mathrm{~Pb}$ counts, and 1 was discarded due to reverse discordance. No attempts were made to analyze distinct crystal domains based on CL texture. The two grains excluded for high ${ }^{204} \mathrm{~Pb}$ yielded high U concentrations. Finally, 2 ages were manually removed from the population based on filtering of age, concordance, $U$ concentration, and U/Th ratios using the AgePick program (Gehrels, 2009). The Remaining 20 grains yielded a weighted mean age of $1755 \pm 14 \mathrm{Ma}(\mathrm{MSWD}=0.6)$. This is within uncertainty of the $1740 \pm 2 \mathrm{Ma}$ reported by Hawkins et al. (1996).

We obtained Hf isotopic data from 21 zircons separated from the Zoroaster pluton. $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values range from 5.9 to 12.7. The $\varepsilon \mathrm{Hf}$ value of the depleted mantle at 1755 Ma is 10.1 , and of the 21 Zoroaster pluton zircons, 19 of them yielded $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values that are indistinguishable from the depleted mantle at the $2 \sigma$ level. All analyses are juvenile after Bahlburg et al. (2011).

## Horn Creek Pluton (K12-90.5R):

The $1713 \pm 2 \mathrm{Ma}$ Horn pluton is a medium grained hornblende quartz diorite to tonalite (Hawkins et al., 1996). The pluton outcrops as a long sliver oriented parallel to the $S_{2}$ foliation, but displays intrusive relations with the Grand Canyon Metamorphic suite (Ilg et al., 1996). A strong northwest striking magmatic foliation $\left(S_{1}\right)$ is transposed by the $S_{2}$ foliation. Peak temperatures for the Trinity Creek metamorphic block, in which the Horn pluton lies, are $722 \pm 83^{\circ} \mathrm{C}$ (Dumond et al., 2007).

Hawkins et al. (1996) described the external morphology of zircons from the Horn pluton as clear grains ranging from prismatic to equant. Our observations show that zircons range in size from long axes of approximately $100-400 \mu \mathrm{~m}$ with aspect ratios from 1:1 to 2:1. Grains are predominantly subhedral, with many broken and some showing signs of resorption. Internal textures revealed by CL imaging show igneous growth zoning; most grains shows fairly pristine concentric zoning, but some show irregular sector zoning and some appear nearly homogenous. Blotchy, CL-dark sections are minimal any typically restricted to the edges of grains.

A total of 30 grains from the Horn Creek pluton were analyzed, and every analysis passed the data reduction process. Two attempts were made to date distinct crystal domains, however both attempted resulted in overlapping ages. Three analyses were excluded after manual filtering using the AgePick program revealed much higher U concentrations than all other grains. The remaining 29 analyses yielded a weighted mean age of $1719 \pm 14 \mathrm{Ma}(\mathrm{MSWD}=0.3)$, in good agreement with the $1713 \pm 1 \mathrm{Ma}$ age of Hawkins et al. (1996). Interestingly, Hawkins et al. (1996) reported a single grains that plotted to the right of their best chord. They interpreted this result as a mixture of a 1713 Ma overgrowth and an inherited core that they were unable to identify by transmitted light observation (Hanchar and Rudnick, 1995). However, our CL images show no evidence for inherited cores in Horn pluton zircons, and our attempts to identify distinct age domains failed.

We obtained 28 Hf isotopic analyses from the Horn pluton. $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values ranged from 5.5 to 12.4 . Of the 28 analyses, 18 were indistinguishable from the $\varepsilon H f$ value of the
depleted mantle at 1719 Ma , and all were within 5 epsilon units of the depleted mantle value of 10.2.

## Trinity Pluton (K12-91.5R):

The $1730 \pm 3$ Trinity pluton is a medium to coarse grained granodiorite to granitic biotite orthogneiss (Hawkins et al., 1996). In the past it has been speculated that the Trinity gneiss was basement to the supracrustal Grand Canyon Metamorphic Suite (Noble and Hunter, 1916), however its age and field relations clearly show that it is intrusive to the Brahma schist (Ilg et al., 1996). The pluton preserves a strong gneissic layering $\left(\mathrm{S}_{1}\right)$ and outcrops as an isoclinally folded $\left(\mathrm{F}_{2}\right)$ sheet (Ilg et al., 1996). Trinity pluton is naturally underlain by the Trinity Creek metamorphic block.

Hawkins et al. (1996) described the zircons of the Trinity pluton as clear, colorless, doubly terminated and euhedral. Our observations show that zircons from the Trinity pluton range in size from long axes of approximately $100-300 \mu \mathrm{~m}$ with aspect ratios from 1:1 to 3:1. Grains are predominantly elongate and subhedral to euhedral. Internal textures revealed by CL imaging show concentric igneous growth zoning and many grains with CL-dark rims that were avoided during analysis. Several apparently distinct cores were identified in CL images, and attempts were made to date discrete crystal domains (see below).

A total of 20 grains from the Trinity pluton were analyzed with 17 core-mantle pair analyses conducted yielding 37 total analyses. Of the 37, 28 analyses passed the data reduction process. 7 analyses were discarded due to high ${ }^{204} \mathrm{~Pb}$ counts, 1 due to unacceptable ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ age uncertainty, and 1 due to discordance. Many analyses
discarded for high ${ }^{204} \mathrm{~Pb}$ counts were conducted quite near to CL-dark rims and likely included some metamict material. Finally, 1 age was removed during manual filtering using AgePick.

Although there appeared to be many CL-distinct crystal domains, all analyses attempting to discern cores from mantles yielded overlapping ages. Furthermore, no systematic variation was observed between U concentration, $\mathrm{U} / \mathrm{Th}$, or concordance for cores and mantles. A weighted mean age of $1755 \pm 16 \mathrm{Ma}(\mathrm{MSWD}=0.4)$ was determined from 27 analyses. This age is in contrast with Hawkins et al. (1996) age of $1730 \pm 3 \mathrm{Ma}$, but within uncertainty of the depositional age of the Granite Gorge Metamorphic Suite into which it intrudes.

We obtained 23 Hf isotopic analyses from the Trinity pluton. $\varepsilon \mathrm{Hf}_{(t)}$ values range from 5.7 to 11.4 and of the 23 total grains 18 are indistinguishable from the $10.1 \varepsilon \mathrm{Hf}_{(\mathrm{t})}$ of the depleted mantle. All grains are within five epsilon units of the depleted mantle.

## Boucher Pluton (K12-96.2L):

The Boucher pluton is a granodiorite to tonalitic pluton that outcrops at river mile 96.2 and is bounded on its eastern side by the 96 Mile shear zone (Ilg et al., 1996). The pluton is weakly foliated with strain localization on its western margin related to motion on the 96 Mile shear zone (Hawkins et al., 199; Ilg et al., 1996). The 96 Mile shear zone marks the boundary between the Trinity Creek block and the Topaz Canyon block, which at greenshist to lower amphibolite grade preserves the lowest grade metamorphism in the Upper Granite Gorge (Ilg et al., 1996; Dumond et al., 1996). An identical age of $1714 \pm 1$ Ma was determined from both $\mathrm{Pb}-\mathrm{Pb}$ and U-Pb analysis of titanite (Hawkins, 1996),
indicating that peak temperatures in the Topaz Canyon block were below the closure temperature of titanite $\left(\sim 590{ }^{\circ} \mathrm{C}\right)$. In this paper, we present 25 new $\mathrm{U}-\mathrm{Pb}$ zircon ages from the Boucher pluton.

Zircons from the Boucher pluton range in size from long axes of approximately $100-300 \mu \mathrm{~m}$ with aspect ratios from 1:1 to $3: 1$. Grains are predominantly elongate and subhedral to euhedral. Internal textures revealed by CL imaging show concentric igneous growth zoning and many grains with CL-dark rims that were avoided during analysis. Several apparently distinct cores were identified in CL images, and attempts were made to date discrete crystal domains on every grain (see below).

A total of 25 grains from the Trinity pluton were analyzed with 25 core-mantle pair analyses conducted yielding 50 total analyses. Of the 50,25 analyses passed the data reduction process. 20 analyses were discarded due to high ${ }^{204} \mathrm{~Pb}$ counts, 1 due to unacceptable ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ age uncertainty, 2 due to reverse discordance, and 2 due to discordance.

In total, 8 core-mantle pair analyses passed the data reduction process, however only 2 core-mantle pairs yielded non-overlapping ages. 20 of the 25 analyses were used to calculate a weighted mean age of $1730 \pm 15 \mathrm{Ma}(\mathrm{MSWD}=0.5)$. This age includes 6 core-mantle pairs that yielded overlapping ages. This age is barely within uncertainty of the titanite ages of Hawkins (1996); however, with the higher closure temperature of zircon, it is likely that our new age of 1730 Ma is a more accurate crystallization age for the Boucher pluton.

No systematic variation was observed between U concentration, U/Th, or concordance in these core-mantle pairs. The remaining 4 analyses came from core analyses and yield a variety of ages. Grain K12-96.2L-1C yielded an age of $1839 \pm 19$ Ma, which corresponds to the $1840 \pm 1$ Ma age of the Elves Chasm gneiss (Hawkins et al., 1996). Grain K12-96.2L-22C yielded an age of $1945 \pm 22$ Ma. Grain K12-96.2L-17C yielded an age of $2483 \pm 24 \mathrm{Ma}$, corresponding quite closely to the 2481 Ma age peak defined in the Vishnu schist (Shufeldt et al., 2010). Finally, grain K12-96.2L-3C yielded an age of $2598 \pm 22 \mathrm{Ma}$. Grains K12-96.2L-1C and 22C did not yield mantle analyses that passed the data reduction process, however grains K12-96.2L-1C and 3C yielded mantle analyses that contributed to the 1730 Ma weighted mean age.

The morphology of these inherited cores varies, however all mantles display concentric igneous growth zoning. K12-96.2L-1C is a homogenous CL-dark care with an irregular shape. K12-96.2L-3C is slightly rounded with igneous growth zoning evident in CL-texture. Between the core and mantle is a thin CL-dark rim. K12-96.2L-17C is nearly euhedral with complex igneous zoning. K12-96.2L-22C is also nearly euhedral, with faint igneous growth zoning and a generally CL-dark texture.

We obtained 20 Hf isotopic analyses from Boucher pluton zircons. 17 Analyses from the 1730 Ma population yield $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values that range from 5.7 to 11.8 . All grains from this population yielded $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values within five epsilon units of the depleted mantle, and 15 analyses were indistinguishable from the depleted mantle $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ value of 10.1 at 1730 Ma . Grain K12-96.2L-1C yielded an epsilon Hf value of 8.7, which suggests the involvement of Elves Chasm gneiss like crust. Grain K12-96.2L-17C yielded an evolved $\varepsilon \mathrm{Hf}_{(t)}$ value of -2.9 , and grain $\mathrm{K} 12-96.2 \mathrm{~L}-17 \mathrm{C}$ yielded an intermediate value of -2.9 .

## Tuna Creek Pluton (13H-99R; K05-100.5):

Zircons from the Tuna Creek pluton range in size from long axes of approximately $50-200 \mu \mathrm{~m}$, though most grains are $\leq 100 \mu \mathrm{~m}$, with aspect ratios from 1:1 to 1:3. Grains are predominantly subhedral to euhedral. Internal textures revealed by CL imaging show concentric igneous growth zoning, and many have CL-dark rims. Distinct core and rim textures were identified in some cases, but were unable to be resolved by a $30 \mu \mathrm{~m}$ beam.

A total of 54 grains from the Tuna Creek pluton were analyzed at the ALC, with 40 grains passing the data reduction process. Manual filtering of ages via AgePick yielded interesting results; 34 grains show a near continuum of ages from 1725-1928 Ma. An additional 6 grains yield ages ranging from $2399-3007 \mathrm{Ma} .17$ grains of similar U/Th ratios, U concentration, and concordance yielded a weighted mean age of $1751 \pm 15$ Ma (MSWD = 1.1). Similarly, 25 grains analyzed at the GEMOC Key Centre yield ages ranging from $1732-2936 \mathrm{Ma}$, with 7 grains defining a weighted mean age of $1737 \pm 7$ $\mathrm{Ma}(\mathrm{MSWD}=0.7)$. In total, 65 ages from the Tuna Creek pluton range from $1725 \pm 10$ Ma to $3007 \pm 60 \mathrm{Ma} .24$ grains contribute to a peak age of $1740 \pm 14 \mathrm{Ma}(\mathrm{MSWD}=1.5)$ which we take to be the crystallization age of the Tuna Creek pluton. The inherited population defines an age peak of 2.48 Ga that is identical to that of the Vishnu Schist. The morphology and CL-texture of inherited grains is varied, but all together indistinguishable the younger population.

Inheritance was identified in the Tuna pluton previously, however those workers were unable to assign a reliable age to the Tuna pluton. Hawkins et al. (1996) reported that the age of the Tuna pluton ranged from $1750-1710 \mathrm{Ma}$ with $>2.0 \mathrm{Ga}$ inheritance.

Our age of $1740 \pm 14 \mathrm{Ma}$ is consistent with both previous geochronologic constraints and field relations; the Tuna pluton displays both $S_{1}$ and $S_{2}$ foliations, and is thus interpreted to have been emplaced either before or during $\mathrm{D}_{1}$ deformation (Karlstrom et al., 2003).

## Elves Chasm Gneiss (K12-115L; K06-113):

The Elves Chasm gneiss is a lineated hornblende-biotite tonalite to quartz diorite which at $1840 \pm 1 \mathrm{Ma}$ is presently the oldest plutonic rock in the southwestern United States.

Zircons from the Elves Chasm gneiss range in size from long axes of approximately $60-200 \mu \mathrm{~m}$ with aspect ratios from 1:1 to $2: 1$. Grains are predominantly subhedral. Internal textures revealed by CL imaging show concentric igneous growth zoning commonly with CL-dark rims.

A total of 27 grains from the Elves Chasm gneiss were analyzed at the ALC, with 4 core-mantle pair analyses conducted for a total of 31 analyses. No core-mantle pairs passed the data reduction process; however, our results show that there was no distinct age population of cores. In all, 14 analyses passed the data reduction process. 16 analyses were discarded due to high ${ }^{204} \mathrm{~Pb}$ counts, and 1 analysis due to low ${ }^{206} \mathrm{~Pb} /{ }^{204} \mathrm{~Pb}$ ratios. The remaining 14 grains yield a weighted mean age of $1850 \pm 18 \mathrm{Ma}(\mathrm{MSWD}=1.4)$. This age is within $2 \sigma$ error of the $1840 \pm 1 \mathrm{Ma}$ age of Hawkins et al. (1996).
$24 \mathrm{U}-\mathrm{Pb}$ ages were obtained from zircons analyzed at the GEMOC Key Centre. These ages yielded a weighted mean age of $1842 \pm 5 \mathrm{Ma}(\mathrm{MSWD}=0.9)$ in excellent agreement with the age of Hawkins et al. (1996).

## Ruby Pluton (K06-107):

The $1716.6 \pm 0.5$ Ma Ruby pluton, exposed from river mile $102-108$, comprises intermingled mafic to intermediate phases (Hawkins et al., 1996). The eastern margin of the pluton shows intrusive relations with supracrustal rocks. Meanwhile, the pluton is bounded in its western side by the Bass shear zone, and displays a weak magmatic foliation near the contact that is concordant with the $S_{1}$ fabric in the adjacent supracrustal rocks (Ilg et al., 1996). The Bass shear zone marks the western edge of the Tuna Creek metamorphic block, which is underlain mostly by the Ruby pluton and Tuna Creek plutons, and characterized by peak temperatures of $769 \pm 188^{\circ} \mathrm{C}$.

Hawkins et al. (1996) described the zircon population from the Ruby pluton as large, clear, equant, and inclusion free grains. BSE imaging of zircons analyzed at the GEMOC Key Centre shows rounded equant to subhedral grains with some inclusions and ubiquitous fractures. Back scattered electron images (BSE) are less effective at revealing internal textures than CL images, however internal oscillatory zoning is displayed in most grains. Some show irregular zoning and nearly homogenous textures. No inherited grains or xenocrystic cores were identified in the Ruby pluton.

A weighted mean of 10 analyses yields an age of $1726 \pm 14 \mathrm{Ma}(\mathrm{MSWD}=0.1)$, in good agreement with the previous age of Hawkins et al. (1996).
$\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values from the Ruby pluton yielded a tight cluster from 7.2 to 9.8 , suggesting that the Ruby pluton is derived from juvenile 1.75 Ga crust.

## Diamond Creek Pluton (K06-228.3):

The $1736 \pm 1$ Ma Diamond Creek pluton is exposed from river mile $212-228.5$ (Karlstrom et al., 2003). It displays abundant magma mingling textures including local ultramafic cumulate texture.

Zircons separated from the Diamond Creek pluton are rounded to anhedral with abundant inclusions and fractures. Internal textures revealed by BSE are irregular, although igneous zoning is locally preserved. A weighted mean of 10 analyses yields an age of $1738 \pm 14 \mathrm{Ma}(\mathrm{MSWD}=0.2)$, in excellent agreement with the previous age of Karlstrom et al. (2003). One grain yielded an age of $1820 \pm 50 \mathrm{Ma}$, which overlaps with the age of the Elves Chasm Gneiss, but no older xenocrystic grains were found.
$\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values range from 5.0 to 10.3 and most are indistinguishable from the depleted mantle curve. The presence of a juvenile 1820 Ma zircon may suggest that the Diamond Creek pluton was in part derived from, or interacted with Elves Chasm age crust during its formation.

## 238-Mile Pluton (K06-238-2):

The 238-Mile pluton is a granitic gneiss from within the Gneiss Canyon shear zone. The Gneiss Canyon shear zone extends from river mile $234-242$ and is a broad high strain zone with locally developed mylonite zones. A variety of supracrustal migmatitic gneisses are injected with granitic melts. The Gneiss Canyon shear zone represents the largest significant metamorphic discontinuity in the Grand Canyon region, where rocks to the west of the Gneiss Canyon shear zone record P-T conditions that correspond to $\sim 15-20 \mathrm{~km}$ depths are juxtaposed against $\sim 10 \mathrm{~km}$ rocks to the east.

Zircons separated from the 238-Mile pluton are primarily rounded with aspect ratios of $1: 1$ to $1: 2$. Some grains preserve pristine igneous zoning, and others yield more homogenous BSE textures.

The 238-Mile pluton has never been previously dated. A weighted mean of 21 analyses yields an age of $1731 \pm 14$. The 238-Mile pluton yielded 3 grains that overlap with the age of the Elves Chasm Gneiss, and an additional 2 ages of 2275 Ma and 2453 Ma. The latter corresponds well to the age peak found in the Vishnu Schist, and the age of 2275 is a globally rare age.
$\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values in the primary 1730 Ma population range from 10.5 to -2.8 ; values indistinguishable from the depleted mantle to more evolved values that reflect the involvement of older crustal material in the pluton's genesis. Of the 3 grains that overlap with the Elves Chasm Gneiss, only one grain yielded a juvenile $\varepsilon \mathrm{Hf}_{(t)}$ value. The oldest grains yield evolved $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values from - 0.9 to -3.4 , which is similar to the $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values of other xenocrystic grains found in the Tuna and 245-Mile plutons.

## 245-Mile Pluton (K06-245-2):

Zircons separated from the 245-Mile pluton are primarily subhedral and somewhat rounded, but two grains are extensively resorbed. Aspect ratios range from 1:1 to 1:3. Most grains contain inclusions and show some fracturing. Pristine igneous zoning is also displayed in all grains except the two that are extensively resorbed.

An age of $1720 \pm 5$ Ma was reported for the 245-Mile pluton (Karlstrom et al., 2003), which is in contrast to our weighted mean of 15 ages that yields $1741 \pm 13 \mathrm{Ma}$, however the well-developed $S_{1}$ foliation preserved in the pluton may suggest that the
older age is more accurate. Three older ages ranging from 2454 Ma to 2658 Ma were also obtained. One of these ages comes from a resorbed grain, however the other two are subhedral grains that display pristine igneous zoning indistinguishable from the younger population
$\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values in the young population range from 2.0 to 7.2 , which is somewhat lower than most of the other plutons in Grand Canyon. None of these primary grains yield $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values that are within error of the depeted mantle. In contrast, one of the older xenocrystic grains yielded a juvenile value of 8.9 at 2645 Ma . The other two grains yielded more evolved $\varepsilon \mathrm{Hf}_{(\mathrm{t})}$ values ranging from 0.2 to -2.6 .

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