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# An interpretative view of open-vent volcanoes

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# 6 Abstract

7 Open-vent volcanoes are special systems where the dynamics of sustained magmatic processes can 8 be thoroughly investigated and where new monitoring tools can be tested and applied. However, 9 various aspects remain puzzling at open-vent volcanoes for which forecasting their behaviour can 10 be an important challenge. Recent papers highlight the very rapid improvements in spaceborn 11 instruments, data acquisition techniques, data treatment and modelling over the last decade, and 12 illustrate the fundamental contribution of long time-series data, either discontinuous or continuous, 13 and the development of multi-parameter studies.

Here we provide an interpretative overview of the main characteristics of open-vent volcanoes 14 15 on the basis of selected examples chosen to be representative of the diversity of their magma composition, their eruptive activity and their geodynamic context. We choose typical open-vent 16 17 volcanoes (Stromboli, Yasur and Erebus), some of them hosting a lava lake (Erta 'Ale, Nyiragongo, Villarrica, Ambrym and Masaya), to those with vigourous activity, which are associated to a long-18 19 lasting eruption (Arenal, Fuego, Popocatépetl, Santiaguito). We briefly review their surface activity and report the values of  $SO_2$  flux and the derived magma supply rate with emphasis on the key 20 21 results found on their behaviour. We show the key role of the magma viscosity and its implication 22 on the degassing. We present the current models used to explain how an open-vent volcano could be maintained, such as by the simultaneous rise and fall of a degassing and degassed magma (bi-23 directional flow models) and the few thermal models at lava lakes and in the conduit. Finally we 24 25 discuss the sulphur evolution for three nearby volcanoes at the triple junction in Central America 26 (Pacaya, Fuego, Santiaguito).

27

# 28 The diversity of geodynamic environments, geochemical compositions and 29 physical properties of magmas

Open-vent volcanoes are built up in various geodynamic contexts (Tables 1a and 1b). Their erupted magmas span wide ranges of chemical composition from basalts to nephelinite or trachyandesite and phonolite (Fig. 1), temperature and viscosity (see reviews of Rose et al. 2013 and Edmonds et al. 2022). Low to moderate viscosity systems are characterized by permanent plumes and gas emissions, eventually explosive activity at the craters and/or emissions of lava flows, whereas high
viscosity, silica-rich magmas able to retain gas bubbles promote lava dome growth and possibly
vulcanian eruptions.

These systems may remain continuously active over time periods spanning decades to millennia, 37 whereas changes in activity style and intensity occurs over the short (hours-days), medium (weeks-38 months) and long (years-decades) time scales. As an example Piton de la Fournaise (Réunion 39 Island) experienced an activity of basaltic lava lake, high lava fountaining and lava flowing, during 40 almost three centuries (1640-1848 CE; Michon et al. 2013). Arenal volcano in Costa Rica is not 41 presently active but erupted during nearly 42 years between July 1968 and 2010 (Parat et al. 2014), 42 with subplinian eruptions, pyroclastic flows and lava emissions (Cigolini et al. 1984; Wadge et al. 43 44 2006) and vulcanian and strombolian activity (Hagerty et al. 2000; Mora et al. 2022).

Short-lasting explosive and paroxysmal eruptions and hazardous fissure-type eruptions draining 45 46 the persistent lava lake at the craters represent ones of the major hazards at open-vent volcanoes. Destructive fissure eruptions are known at Erta 'Ale (Barnie et al. 2016), Kilauea (Halema'uma'u 47 48 crater; Patrick et al. 2019; Fig. 2a), Nyiragongo (Komorowski et al. 2004; Morrison et al. 2020; see Pouclet and Bram 2021 for a review; Fig. 2b), and Ambrym volcanoes (Németh and Cronin 2011; 49 50 Shreve et al. 2019). Short-term explosive eruptions, interrupting the persistent activity of the crater (mild-strombolian activity, passive degassing, lava lake) occur repeatedly at Villarrica in Chile in 51 2015 CE (Aiuppa et al. 2017; Johnson and Palma 2015; Johnson et al. 2018), at Stromboli in 52 Aeolian Islands in Italy (historic and present-day eruptions, Bertagnini et al. 2011; Rosi et al. 2013; 53 54 Métrich et al. 2021; Aiuppa et al. 2021; Pichavant et al. 2022; Fig. 2c and 2d) or as recently observed at Fuego (Naismith et al. 2019; Liu et al. 2020) and Santiaguito (Wallace et al. 2020) in 55 Guatemala. Erebus volcano (Antartica) also experienced an active explosive activity throughout the 56 Pleistocene and Holocene as testified by tephra layers found up to ~180 km away (Harpel et al. 57 58 2004).

59 Popocatépetl, with cyclic construction and destruction of lava domes and a large gas flux at the 60 surface may be considered part of the category of open-vent systems as do Arenal and Pacaya that 61 produced homogeneous magma for decades. The cyclic eruptions of high-level magnitude at Fuego 62 (Guatemala), as well as the long-term quasi-steady eruptive pattern at Santiaguito, and Pacaya are 63 other examples of open-vent systems, which are discussed here.

The timescales of magma storage, ascent and eruption are now studied through the measurement of the chemical zoning and diffusion patterns in crystals (Costa 2020). Focussing on open-vent volcanoes we note a remarkably high decompression rate of magma and a short time of magma crustal transfer leading to explosive eruptions as retrieved from the modelling of the diffusion

profiles of H<sub>2</sub>O, CO<sub>2</sub>, and S in glass embayments (open-pockets of magma) partly entrapped in 68 crystals. The decompression rate of the magma body that sustained the subplinian eruption (VEI 4) 69 on 17 October 1974 at Fuego reached 0.3-0.5 MPa/s, equivalent to 11-17 m/s ascent velocity 70 (Lloyd et al. 2014), when considering a single-stage decompression. Hence the magma rose 71 from~10 km depth - its storage region - within the extremely short span of time of 8 to 12 minutes. 72 The total ascent time is between 10 and 35 minutes for a two-stage decompression modelling taking 73 into account different depth ranges of CO<sub>2</sub> and H<sub>2</sub>O degassing. At Kilauea, Fergusson et al. (2016) 74 used the same method and report average decompression rates for the magmas that fed the 75 Hawaiian lava fountains to basaltic subplinian eruptions of ~0.05–0.45 MPa per second (Fig. 2a). 76 They calculate ascent timescales from  $\sim 5$  to  $\sim 36$  min for these magmas that are ascending from  $\sim 2$ 77 to ~4 km, respectively. The magma decompression rates of 0.1 to ~2 MPa/s during the 2017–2018 78 Plinian eruption of Aoba (Ambae; Vanuatu arc) suggest short travel time of less than 2 minutes 79 80 from  $\sim 0.5$  to 3 km depth (Moussallam et al. 2019).

Imbalance between the volume of magma that degassed and the volume of magma is the core of open-vent activity. First observed at basaltic volcanoes and named "endogenous growth" (Francis et al. 1993), it was widely used and debated in litterature (Shinohara 2008; Carn et al. 2017). The recent comparison between thermal and SO<sub>2</sub> fluxes suggests that the magma supply rate at Masaya and associated with SO<sub>2</sub> degassing may be 5.5 more than the amount of magma which reaches the surface from thermal fluxes estimated by satellites (Aiuppa et al. 2018).

Edmonds et al. (2022) provide a review of the degassing and volcanic activity at open-vent 87 88 volcanoes producing a wide range of magma compositions from basalts to dacites and rhyolites. 89 They propose that deep-derived volatile phase from magmas accumulating and degassing at crustal depths sustains the persistent degassing at these volcanic systems, which are gas-dominated. It 90 implies the differential transfer of the deep-derived gas phase whereas a major part of the volume of 91 degassing and crystallising magma would be accommodated at depth. As an example, at Masaya the 92 93 calculated supply rate of tholeiitic magma of 0.19±0.06 km<sup>3</sup> per year is suggested to be totally accommodated at depth by extensional tectonic (Zureck et al. 2019). These features are also in 94 95 agreement with the lack of seismicity other that associated with explosions at Erta 'Ale (Bouche et al. 2010). 96

97 Persistent degassing at open-vent volcanoes that may have an impact on the health of 98 inhabitants and on agriculture in downwind communities is also a chronic volcanic hazard (see 99 Edmonds et al. 2017 for a review). As an example, the two huge craters Benbow and Marum of 100 Ambrym volcano in Vanuatu, each of them hosting a lava lake until December 2018 (Moussallam 101 et al. 2021), deliver high fluxes of gas rich in trace elements, metals metalloids, halogens (Allard et

al. 2016a, 2016b). Human and environmental impacts of halogens concentrate in the downwind 102 western part of the island (Cronin and Sharp 2002; Bani et al. 2009; Allibone et al. 2012). Masaya 103 is one of the strongest degassing sources of SO<sub>2</sub> in the Central America Volcanic Arc (Aiuppa et al. 104 2018), and possibly maintained its steady state behaviour for 30 ka (Zureck et al. 2019). Masaya is 105 also recognised to emit such elements as Hg (Witt et al. 2008). Health and economic hazards 106 associated to Masaya persistent degassing have been long time identified (Delmelle et al. 2002; 107 Williams-Jones and Rymer 2015; Calkins and Delmelle 2021). The 2018 eruption along the lower 108 East Rift Zone (LERZ) at Kilauea produced large quantities of gas and aerosols (i.e., 7.1–13.6 Mt 109  $SO_2$  and 1.5–2.8 Mt  $CO_2$ ) in 3 months with significant impacts on air quality (Kern et al. 2020). 110 This aspect of the environmental impacts of the volcanic degassing will not be developed here. 111

In summary, the two main characteristics of open-vent volcanoes are the persistent and excess degassing on one side and on the other side their ability to prevent cooling either by reducing heat losses or by having additional sources of heat at a sufficient rate. We combine here geophysical modelling approach to magma geochemistry to pinpoint key aspects of a selection of well-known open-vent volcanoes. Our goal is to highlight the fundamental properties of these systems, and the rapid improvements of both techniques and modelling for better understanding of their volcanic activity and improvement of their eruption forecasting.

We focus on open-vent volcanoes, which cover a range of magma composition and rheology, 119 and show a diversity of activity from persistent degassing to strombolian and vulcanian 120 explosions, up to paroxysmal activity. These volcanoes are monitored and/or have been 121 122 extensively or more poorly studied, such as Erebus (Antartica), Stromboli (Italy) and Yasur (Vanuatu), respectively. They can host a lava lake (Masaya in Nicaragua, Villarica in Chile, 123 Ambrym in Vanuatu arc, Erta 'Ale and Nyiragongo-Nyamuragira system in Africa) or offer the 124 top of the magma column to air (Stromboli, Yasur, Erebus). In complement Popocatépetl, a dome-125 forming volcano is considered as one of the largest emitter of SO<sub>2</sub> and Arenal (Costa-Rica), a mid-126 127 point in term of magma composition between basalt and dacite, for its long-lasting eruption and the number of studies triggered by its eruption (Mora et al. 2022). Finally three nearby volcanoes 128 129 from Central American Volcanic Arc (Pacaya, Fuego and Santiaguito) are also selected to discuss the sulphur behaviour in arc-magmas. Kilauea volcano is mostly ignored, due to its enormous 130 number of studies, which are synthetised in a recent special issue of this journal (see Patrick et al. 131 2020 for an introduction). 132

We first discuss the bulk chemical homogeneity of the magma, before presenting basic knowledge necessary to understand the eruptive patterns of our selected volcanoes. We further

present the surface activity at lava lakes, including their degassing. We then focus on the 135 strombolian activity of Erebus, Stromboli and Yasur, firstly based on their petrology and magma 136 viscosity and secondly on their surface activity and degassing. We then turn to a summary on 137 vulcanian explosions before reviewing the thermal models, applicable to both lava lakes and 138 conduits. We present the various bi-directional flow models in the conduit and the key 139 informations that can be obtained from sound waves studies. Lastly, we discuss the possible 140 exsolution path of sulphur, firstly on the fate of the sulphur and secondly with the example of 141 Popocatépetl and thirdly on the three nearby volcanoes in Central America, Pacaya, Fuego and 142 143 Santiaguito.

144

# 145 A bulk chemical homogeneity of the magma on the long-term

A relatively homogenous bulk composition but a complex history of the crystal assemblage appear
to be shared characteristics of steady-state open-vent volcanoes that may testify to magma recharge
and mixing or convection motions without new magma injection.

Arenal volcano (Costa Rica) provides a good example of the long-term chemical homogeneity of 149 150 the bulk lava samples but a complex history recorded by the crystals, involving magma injection, mixing, mineral dissolution and crystallisation (Mora et al. 2022). The distribution of SiO<sub>2</sub> contents 151 152 of Arenal magma is unimodal and centered on the lavas at about 54 to 57 wt% SiO<sub>2</sub> (Bolge et al 2006; Fig. 1). Its mineralogy reveals a complex and diverse growth and mixing histories of the 153 crystal assemblage, although the basaltic-andesitic composition of the extruded magma did not 154 change since 1970 (Streck et al. 2002, Ryder et al. 2006; Parat et al. 2014). The basaltic andesites at 155 Arenal are experimentally reproduced after a minimum of 15% fractional crystallisation at 200 MPa 156 at 1000-1050°C from a basaltic melt (Parat et al. 2014). Streck et al. (2002) suggest, on the basis of 157 mineral zoning and compositions, that recurrent basaltic magma recharges occur with sub-decadal 158 frequency and may predate eruption by months or less. Stromboli is known for its persistent activity 159 at the craters since the eighth century CE (Bertagnini et al. 2011; Rosi et al. 2013). Both the major 160 and trace element of the magma vary in a very restricted range of basaltic-shoshonitic composition 161 162 (Fig. 1) over the last centuries (Pompilio et al. 2011) while the plagioclase crystals record successive dissolution-crystallisation events (Landi et al. 2006; 2009). Erebus crater hosts a long-163 lived exposed magma column, often referred to as a "lava lake" fed by a nearly anhydrous phonolite 164 (Giggenbach et al. 1973; Sims et al. 2021), while 1-cm-wide anorthoclase crystals register several 165 (from 1 to 3) complete cycles between the reservoir and the exposed magma column via the conduit 166 167 (Moussallam et al. 2015). These cycles also imply contrasting flow regimes in reservoir and

conduit, with convection in the former (with 1 to 3 cycles of 150 days at a speed of 0.5 mms<sup>-1</sup>) and
 more complex cycles of exchange flow and re-entrainment in the latter.

170 In summary, the petrological studies demonstrate the magma chemical homogeneity on 171 several-years time scales, despite the heterogeneous but steady pattern of minerals at the spatial 172 microscale scale.

## 173 Basic considerations on degassing and magma viscosity

Two main parameters, namely the degassing pattern and the magma viscosity, are playing a key role in understanding the eruptive regimes. To set the scene for open-vent volcanoes, we discuss here relevant basic informations that will be used later in the paper.

177

## 178 Degassing

Degassing at the surface can occur for rising bubbles with various geometries (small, spherical cap and Taylor bubbles (Fig. 3a, 3b and 3c). Laboratory experiments on bubble rise in inviscid to low viscosity fluids have shown that the geometry of rising bubbles depends only on a set of three dimensionless numbers (Clift et al. 1978). These are the Reynolds (*Re*), Eotvös (*Eo*) and Morton (*Mo*) numbers. The Reynolds number, *Re*, quantifies the respective role of viscous and inertia forces, the Eotvös number that of buoyancy versus surface tension forces and the Morton number (Mo) is the ratio between viscous and surface forces. They can be written as:

186  $Re = \rho dU/\mu$  (Eq.1)

187 Eo =  $\rho g D_{\Box}^2 / i \sigma$  (Eq. 2)

188 Mo =  $g\mu^4 / \rho \sigma^3$  (Eq. 3)

189 where  $\rho$  and  $\mu$  are the density (kg m<sup>-3</sup>) and dynamical viscosity (Pa.s) of the magma, respectively;

190 d is the bubble diameter (m);  $\sigma$  is the surface tension (kg s<sup>-2</sup>) and U the bubble upwards velocity

191 (m/s). The bubble velocity depends on the shape of the rising bubble for, small bubbles, i.e.

192 called the Stokes velocity (Clift et al. 1978; Eq. 16 in Vergniolle and Jaupart 1986), spherical cap

bubbles (Clift et al. 1978; Eq. 5 in Bouche et al. 2010) and Taylor bubbles (Clift et al. 1978; Eq.

12 and 17 in Vergniolle and Caplan-Auerbach 2006). Note that the Morton number only depends

195 on the material properties of the magma and the gas, while Reynolds and Eotvös numbers depend

196 on the bubble diameter. The regime diagram displaying the bubble shape during rise is classically

used by presenting the Reynolds number as a function of Eotvös number (Clift et al. 1978; Fig. 11

in Vergniolle and Jaupart 1986).

199 Bubbles can also be stationary (Fig. 3d), as those accumulated below the crust of a lava lake,

200 creating a foam. Local coalescence within this foam could lead to a long and thin bubble referred

to as a flat bubble (Fig. 3d). These flat bubbles have been shown to exist at the Erta 'Ale lava lake(Bouche et al. 2010).

203

#### 204 Key parameters on magma viscosity

The viscosity of a magma ( $\mu$ ), which is a three phase mixture of melt with dissolved water, crystals and bubbles (Le Losq et al. 2015), is calculated by summing terms as follows:

207  $\log \mu_{\text{magma}} = \log \mu_{\text{melt}} + X_{\text{water}} + Y_{\text{crystal}} + Z_{\text{bubble}}$  (Eq. 4)

where X is the weight fraction of water, and Y and Z the volume fraction of crystals and bubbles,respectively.

210 The respective effects of the temperature (Fig. S1a), the water concentration of the melt (Fig.

S1b), the fraction of crystals (Fig. S1c) and of the shape and size of bubbles (Fig. S1d) on the
magma viscosity are displayed in the supplementary material.

The review paper of Mader et al. (2013) provides an excellent summary of the current models 213 and of the physical processes, which play a role in estimating the magma viscosity at various 214 volcanoes. The first parameter is the crystal volume fraction. The linear approximation of Einstein, 215 who calculates magma viscosity by assuming that the crystal content is small (Mader et al. 2013), is 216 not valid above a crystal volume fraction of 10%. The lowest valid estimate for the viscosity is 217 218 obtained by Gluth and Gold, from the Taylor expansion of the linear Einstein's equation, made at the second order for the crystal volume fraction (Mader et al. 2013). Some models even predict an 219 220 exponential increase above a volume fraction exceeding 0.65 (Mader et al. 2013), a case that is never met by our chosen open-vent volcanoes (Tables 1a and 1b). 221

The second parameter to consider is the crystal size and shape and their distributions (Cimarelli et al. 2011; Mueller et al. 2011). The crystal shape is important because it exerts a strong control on the value of the maximum packing fraction, the value for which the viscosity becomes infinite (Mader et al. 2013). This effect is complicated to quantify due to the range of aspect ratios (Mader et al. 2013). The only currently available analysis has to be restricted to that of the average value (Mader et al. 2013) giving a large uncertainty on the final result.

The third parameter is the amount of dissolved water in the magma in the conduit, which when large decreases significantly the magma viscosity (Giordano et al. 2008; Fig.S1b). The fraction of dissolved water in magma is drastically changing as the result of gas exsolution when approaching the surface. An accurate prediction of magma viscosity in the conduit requires a good knowledge of the close link between water degassing and crystallisation (Pichavant and MacDonald 2007), both parameters increasing the viscosity close to surface (Fig. S1b and S1c). The depth at which the viscosity starts to increase due to simultaneous loss of dissolved water and increase in the

crystal volume fraction is a key ingredient to be known to have a correct view on the physical
processes associated to the gas-liquid flow in the conduit, assuming that the crystals passively
follow the magma global motion.

The fourth parameter is the gas volume fraction, which also increases due to gas exsolution 238 when approaching the surface. The changing gas volume fraction from 0, i.e. small, to 30 and 60% 239 increases the magma viscosity by a factor 2.5 and 10, respectively (Fig. S1c; Mader et al. 2013). 240 The exact calculation of the gas volume fraction is also complex because it depends on whether 241 the small bubbles present within the liquid can move in respect to the liquid, as for a separated gas 242 flow, or are stagnant, as for a homogeneous gas flow (Vergniolle and Jaupart 1986). The 243 consequence of having a separated gas flow is very important for the magma viscosity because the 244 245 gas volume fraction is significantly reduced when compared to that of a homogeneous gas flow (Fig 5b and 6 in Vergniolle and Jaupart 1986). The first estimates of the gas volume fraction had 246 247 assumed a homogeneous gas flow (Burton et al. 2007b; Allard et al. 2016b), whereas the framework to calculate the gas volume fraction for a separated gas flow had now been developed 248 249 (Pioli et al. 2012; Pansino et al. 2019). Three mechanisms can lead to a separated gas flow. The first one is related to the existence of small bubbles sufficiently large, prior to the depth where the 250 251 gas exsolution starts to be significant, to rise in respect to the melt. The second one is related to the existence of large bubbles and the third one to the existence of bubble clusters, i.e. an aggregate of 252 small bubbles able to rise on its own. Bubble clusters may form in the underlying conduit below 253 lava lakes, as shown to exist at Erta 'Ale (Bouche et al. 2010), and as discussed later in the lava 254 255 lake section.

The fifth parameter is the bubble shape and its effect is displayed, in steady shear, by the Capillary number (*Ca*) (Mader et al. 2013). This dimensionless number is the ratio between the viscous stress ( $\mu\dot{\gamma}$ ,where $\dot{\gamma}$  is the shear strain-rate) acting on a bubble of diameter d, and the restoring stress supplied by surface tension (2 $\sigma$ /d), giving:

260 
$$Ca = \frac{\mu d \dot{\gamma}}{2\sigma}$$
 (Eq. 5)

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The small and non-deformable bubbles increase the magma viscosity by adding some complexity and tortuosity into the path offered to the parcels of moving fluid, while the deformable bubbles decrease the magma viscosity by adding a free slip into the bubbly liquid (Mader et al 2013; Fig. S1d). In both cases, this effect starts to be noticeable for a gas volume fraction of 10 % and is more marked when the gas volume fraction increases. The viscosity of a bubbly magma containing a gas volume fraction of 35 % is either increased by a factor 1.4 for small bubbles or decreased by a factor 0.3 for deformable bubbles (Fig. S1d). The relevance of knowing the bubble

shape in the conduit, such as schematically drawn at Santiaguito (Bluth and Rose 2004), might
therefore be important. However, looking at the bubble shapes preserved in the ejecta might be
misleading due to additional effects associated to expulsion.

Another challenging phenomemum arises when a separated gas flow exists in the conduit. Laboratory experiments show that a global liquid motion can result from the bubble motion (Pioli et al. 2012; Pansino et al. 2019). When these conditions are met the small crystals will passively follow the magma motion, thereby decreasing the local crystal volume fraction.

In summary, the spatial organisation of the degassing gas bubbles in the conduit is complicated for a magma relatively rich in water, because of the competitive effects of the water content of magma and the gas bubble on the viscosity. The question of the progressive degassing as water is exsolving and how thick an overlying plug can be an interesting question that is addressed further because it requires the knowledge of the dynamics of the magma motion and of its bubbles.

280

# 281 Magma supply rate

282 Two simultaneous pioneer studies, applied to Stromboli (Allard et al. 1994) and Izu-Oshima (Japon; Kazahaya et al. 1994), combined records of SO<sub>2</sub> fluxes with sulphur content prior (melt 283 inclusions) and after (matrix glass) degassing at surface to assess the volume of the degassing 284 magma. An average magma supply rate can be calculated for a given period. This chemical budget 285 does not consider any specific physical process or depth. The only assumption is that the initial 286 sulphur content is quasi-fully lost between depth and surface. This framework had been largely 287 288 used to assess the magma supply rate at open-vent volcanoes (Tables 1a and 1b) using repetitive 289 measurements of SO<sub>2</sub> fluxes (Tables 2a and 2b).

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# 291 Long-term surface activity at lava lakes

### 292 Lava lakes: pressure gauges of the open-vent systems

Only a few open-vent volcanoes, although distributed all around the world in various geodynamic 293 294 environments (subduction, rift, and hot spot), host long-lasting lava lakes (6-7 presently known) and/or intermittently active lave lakes (Table 1a). Sensu stricto, a lava lake is a natural magma body 295 whose dimensions are large compared to the underlying conduit. In this respect, the so-called "lava 296 lake" at Erebus has a roughly conical shape (Oppenheimer et al. 2009), 40 m across for 30 m depth 297 with 4-10 m wide conduit opening at its base (Dibble 2008; Oppenheimer et al. 2009) and a 298 thickness of a few tens of meters (Jones et al. 2015). Its small diameter and thickness, compared to 299 that of the conduit, suggest it to be the top of a magma column rather than a proper lava lake. It 300

offers a direct view of the underlying conduit when looking from above. For simplicity, we shallrefer in the text of Erebus as an exposed magma column.

303 Now, the evidence emerges that most lava lakes have experienced sequences of draining-refilling along their life time that represents major hazards. At Ambrym volcano, several units derived from 304 ponded lava lake bodies were recognised along the Marum crater rim that testifies repeated 305 drainage, excavation or gentle subsidence (Németh and Cronin 2008). Pouclet and Bram (2021) 306 give a detailed review of the activity of the Nyiragongo lava lake, observed since 1948, to have 307 permanent to intermittent and episodic activity with long-term pauses. These authors discuss the 308 mechanisms at the origin of the periodic oscillations of a few tens of minutes for meter-scale level 309 variations. Short-term gas pistoning events and gas release for the rising and the falling phases 310 respectively, modelled by Witham and Llewellin (2006), is discarded because of the absence of 311 relationships between degassing process and oscillatory behaviour of the lava lake (Valade et al. 312 313 2018), as observed at Kilauea (Patrick et al. 2015). Sudden drops of several tens of meters with strong degassing are ascribed to temporary interruption of the magma convection in the conduit 314 315 with the lack of fresh magma ascent according to Bobrowski et al. (2017). A draining-refilling model was proposed for Nyamuragira, which involved a gravity-driven drainage of a shallow 316 317 magma body (or multiple interconnected magma pockets) induced by the eruption of deep undegassed magma along a lateral rift-parallel dike (Coppola et al. 2016). 318

The dynamics of the lava lakes has been intensively studied through the multiparametric 319 measurements of gas and heat fluxes, surface lake motions and flow velocities (Pering et al. 2019, 320 Lev et al. 2019). Measurements have been sporadic or inexistent in the past but they are becoming 321 continuous at a few well-monitored volcanoes. The level of the lava lakes is measured directly at 322 Kilauea (Tilling et al. 1987; Patrick et al. 2015; Fig. 4), Nyiragongo (Burgi et al. 2014; 2020; 323 Barrière et al. 2018; Pouclet and Bram 2021; Fig. 5), Erebus (Jones et al. 2015; Peters et al. 2014a; 324 2018; Fig. 6) or estimated undirectly from satellites (MODIS) at Erta 'Ale (Vergniolle and Bouche 325 326 2016; Fig. 7) or by measuring the shadow length cast by the crater rim on satellites imagery (Oppenheimer and Francis 1997; Barrière et al. 2018). Unfortunately, the use of satellites is 327 328 somehow restricted to lava lakes having a funnel shape, such as at Erta 'Ale (Vergniolle and Bouche 2016) and Nyiragongo (Burgi et al. 2014), a sufficient radiative surface and a reasonably unfrequent 329 cloud cover. 330

Lava lakes, by being in direct connection with a shallow magma reservoir, offer a unique opportunity to decipher the physical processes at depth. Their level can act as a pressure gauge, making its measurement a key parameter to follow. At Ambrym (Fig. 8), the lava lakes were visually observed to be at a high-level prior the effusive activity of 21 February 2015 (Radebaugh et

al. 2016; Hameling and Kilgour 2020; Shreve et al. 2021), as well as that of mid-December 2018 (Hameling et al. 2019; Shreve et al. 2019; Moussallam et al. 2021). This shows that the pressure was building up somewhere in the plumbing system prior to these major events. Similarly numerical simulations performed at Nyiragongo, and validated by 30 years of records, have also shown that (i) the level of lava lake rises prior to paroxysmal activity, but (ii) slows down in the preceding months/year (Burgi et al. 2020). Therefore, the fluctuations in the height of a lava lake can track changes in pressure and the dynamics of the magmatic systems.

At Masaya (Fig. 9), the lava lake has persisted with varying levels between 2015 to present (C. 342 Wauthier personal communication). Occasionally, these vents have enlarged in dimension, most 343 likely due to magma level increase in the conduit and crater floor collapse, ultimately leading to a 344 lava lake (Aiuppa et al. 2018). Its surface activity had been explained qualitatively, but not 345 quantitatively, by the behaviour of the stable foam accumulated at the top of the reservoir (Stix 346 347 2007). At Erta 'Ale, the lava lake level was also explained, and modelled, to be controlled by the thickness of the stable foam at the top of the shallow reservoir (Vergniolle and Bouche 2016). The 348 lava lake rises when the underlying gas flux below the foam exceeded that associated to the foam 349 spreading and it falls when the underlying gas flux had stopped (Fig. 10). The foam thickness was 350 351 modelled with these two different boundary conditions and compared to the level of the lava lake, 352 giving two independent constraints on the plumbing system, namely the area of the shallow magma reservoir and the underlying gas flux (Vergniolle and Bouche 2016). This analysis, done for several 353 years, showed that the gas flux at the top of a magma reservoir could, at time, be oscillatory with a 354 period of 10.8 days, continuous or nonexistent (Vergniolle and Bouche 2016). 355

The connection of the lava lake with a shallow plumbing system is in agreement with the good match observed between the pressure loss, the drop in lava lake level and the deformation pattern (Moore et al. 2019), as at Kilauea (Patrick et al. 2015; Poland and Carbone 2018), Ambrym (Németh and Cronin 2008), Nyiragongo (Pouclet and Bram 2021) and Erta 'Ale (Barnie et al. 2016) with a long-term influence of the regional tectonic activity.

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#### 362 Degassing and bubble regimes

We focus first on the present knowledge of lava lakes, which can often be seen from above, when they are not almost entirely covered by a solid crust and hidden as occurring, respectively in December 2004 at Erta 'Ale (Fig. 7c), and in January 2009 at Masaya (Fig. 9b).

At Erta 'Ale, the degassing occurs through the breaking of the spherical cap bubbles at a fixed location at the surface of the lava lake (insert in Fig. 7a) or through the flat bubbles (Fig. 3d) located on the incandescent zones, which sometimes separate the few cold plates of the crust

(Bouche et al. 2010; Fig. 7a). A spherical cap bubble rising in a conduit in such low viscosity 369 magma can grow a bubbly wake, which periodically detaches from its base upon reaching a 370 maximum volume proportional to that of the driving spherical cap bubble (Bouche et al. 2010). 371 This bubble cluster can then rise in the conduit on its own at a relatively large velocity. The 372 dismantling of these bubble clusters by the return flow of the lava lake is thought to be at the 373 origin of the small bubbles. These small bubbles accumulate below the cooling crust where they 374 can coalesce to produce the flat bubbles (Bouche et al. 2010). Spherical cap bubbles and flat 375 bubbles co-exist and correspond to two different types of degassing. Both types of degassing at 376 Erta 'Ale lead to a weak strombolian activity. However the driving gas overpressure leading to the 377 breaking of spherical cap bubbles is significantly larger than that driving the breaking of the flat 378 bubbles (Bouche et al. 2010). On some occasions, these bubbly clusters remain attached to their 379 driving spherical cap bubbles until very close to the surface of the lava lake, and burst at the 380 381 surface as small lava fountains (Bouche et al. 2010).

Is the rise of the large bubbles at other lava lakes sufficiently turbulent to grow a wake at the 382 383 surface or in the conduit? Can this bubbly wake detach and incorporate small bubbles as proposed by Bouche et al. (2010)? The key point of such a mechanism is that some magma is rapidly 384 385 brought to the surface, travelling at the velocity of the driving large bubble or that of a bubble cluster when within a detached bubbly wake. This depends on the value of the Reynolds number 386 (Eq. 1). For a value above 30, a wake, a vortex zone initially attached to the driving large bubble, 387 exists and can attract small bubbles in it (Fig. 15 in Vergniolle and Ripepe 2008; Fig. 27.11 in 388 389 Taddeucci et al 2015). Above a value of 70-100, the wake becomes open, with strong interactions 390 between the fluid around the bubble and the wake itself. This laminar wake is also unstable, hence can detach and rebuild. At Reynolds numbers exceeding 2000, the detachment of the wake, fully 391 turbulent, occurs periodically upon reaching a critical volume, proportional to the volume of the 392 driving bubble (Bouche et al. 2010). In this latter regime, the quantification of the volume of 393 394 magma within the detached bubbly wake is possible, providing a knowledge of the volume of the driving bubble and the gas volume fraction within the bubbly wake. The magma might be 395 396 sufficiently fluid at a depth where the amount of dissolved water is considered as partly or fully undegassed. This allows the large bubbles to rise at relatively high Reynolds number, despite their 397 compression in gas volume with depth. In this case, the degassing is organised at depth with large 398 driving bubbles and their detached bubbly wakes. Large bubbles and detached bubbly wakes can 399 400 then rise towards the surface and potentially have their Reynolds number decreasing, despite their increase in gas volume, due to the rapid increase in the gas volume fraction existing in the magma 401 402 containing stagnant bubbles. If the Reynolds number becomes below the critical value of 70-100,

the values estimated for the large bubbles at the surface do not reflect anymore the fact that the detached bubbly wakes were previously formed in the conduit and had already led to a proper separated gas flow at depth. If a separated gas flow is organised at depth, the newly gas exsolved above this depth is likely to mainly diffuse into the small bubbles already present in the separated gas flow, enhancing their differential velocity in respect to the magma, hence the vigour of the separated gas flow.

One should then consider the Reynolds number (Eq. 1) to decide on the existence of the 409 detached bubbly wakes somewhere in the conduit. For simplicity, we considered here the depth 410 where the magma is still rich in dissolved water to be 500 m below the surface for all the selected 411 lava lakes, based on the calculation of the gas volume fraction at Ambrym and its relatively high 412 413 dissolved water content (Allard et al. 2016b). Assuming that the magma is fully degassed, the Reynolds number calculated a depth of 100 m, is reduced by a multiplicative factor of 0.19 (Table 414 3). It would be reduced by a factor of 0.087 at 500 m due to gas compression when assuming, for 415 the sake of the argument, a fully degassed magma. 416

We are now going to discuss specific examples, using the Giordano's model (2008) unless specifically given elsewhere and the relevant parameters (Table 3). For each example, we discuss the final magma viscosity when adding crystals or small bubbles whenever possible.

The Reynolds number of the driving spherical caps at Erta 'Ale, is 1840 and 8000 at the surface 420 when using an updated value of the magma viscosity of 22-27 Pa.s (Table 3). It is still well above 421 the threshold value for which the detachment of the bubbly wake exists and is periodic. Again, this 422 example illustrates the importance of the viscosity on the bubble regime as the value of 10 Pa.s 423 used by Bouche et al. (2010) could have led to a different scenario. The formation of detached 424 bubbly wakes (Fig. 11a), which is a mechanism at work even during a period of very low activity 425 of the lava lake (Bouche et al. 2010; Vergniolle and Bouche 2016), is hence a persistent feature of 426 427 its degassing (Fig. 11).

Similarly, the large bubbles observed at Nyiragongo, 10 m in diameter (Tazieff 1996), led to relatively high values of the Reynolds number at the surface, between 3152 and 822 (Table 3), for a magma viscosity from 56 Pa.s to 230 Pa.s (Burgi et al. 2014; 2018; 2020), respectively, a value also sufficient to produce detached bubbly wakes. The Reynolds number at a depth of 500 m is also sufficiently high, 306 and 71 for a magma viscosity of 56-230 Pa.s, respectively, (Table 3) to lead to a fully separated gas flow in the conduit, organised with driving spherical cap bubbles and their bubbly wakes (Fig. 11a).

The Reynolds number of the spherical cap bubbles observed at Villarica is between 65 and 232 at the surface (Table 3), with gas volumes between 200 m<sup>3</sup> and 2500 m<sup>3</sup> (Gurioli et al. 2008),

respectively, and a magma viscosity of 1780 Pa.s, calculated by assuming a crystal volume 437 fraction of 38 vol% obtained with stagnant crystals (Witter et al. 2004). Such a value of the 438 Reynolds number is again mostly compatible with the formation of bubbly wakes and their 439 detachment close to the surface. The very high Reynolds number at a depth of 500 m, 20000-440 71500 and 556-1990, when assuming a negligible crystal volume fraction and a magma viscosity 441 of 0.5 Pa.s and 18 Pa.s, respectively, obtained for a dissolved water content of 4 wt% and 1 wt% 442 (Witter et al. 2004; Table 3), clearly show that the driving spherical cap bubbles are organised at 443 depth similarly to that of Erta 'Ale close to the surface (Fig. 11a), with a periodic detachment of 444 the bubbly wake onto reaching a maximum wake volume of 4-6 that of the driving spherical cap 445 (Bouche et al. 2010). 446

The bubbles observed at the surface of Ambrym lava lakes were very large, plurimetric (Allard 447 et al. 2016b) with a diameter of 5-10 m at Benbow on 5 October 2008 (P. Allard, pers. com. on 25 448 449 October 2021) and even larger at Marum in May 2014, 12 to 16 m in diameter (Radebaugh et al. 2016). The Reynolds number of these bubbles at the surface is between 318 and 1820, for a bubble 450 451 diameter of 5 and 16 m in diameter, respectively, and a magma viscosity between 170 and 250 Pa.s (Table 3). The conditions are therefore met to have the large bubbles rising in the lava lake 452 453 and at the origin of detached bubbly wakes at the surface. However, the above calculation of the magma viscosity does not consider the gas volume fraction of small bubbles potentially suspended 454 in the magma, a limit case for which most of the bubbles, even small, can move in respect to the 455 liquid, hence in a separated gas flow. In contrast, the gas volume fraction has been estimated by 456 457 combining records from open-path Fourier transform infrared spectroscopy of gas emissions at Ambrym, with the laws of the gas exsolution of  $H_2O$ ,  $CO_2$  and  $SO_2$  from the melt at equilibrium, 458 while all the newly produced gas is organised into small bubbles passively following the melt 459 phase (Allard et al. 2016b). In these conditions, the magma viscosity is increased by a factor 10 460 (Fig. S1c; Mader et al. 2013) as soon as it reaches the threshold value, for which the small bubbles 461 462 are touching one another, which occurs at a depth of 160 m (Allard et al. 2016b). In this case, the degassing at the surface of the lava lake at Ambrym does not present detached bubbly wakes for 463 464 the smallest observed bubbles, 5 m in diameter (Fig. 11b). However the detached bubbly wakes 465 can exist if the diameter of the driving large bubble exceeds 8.5 m (Fig. 11c), such as for the largest bubbles on 5 October 2008 (P. Allard, pers. com. 25 October 2021) and for the entire range 466 of bubble diameters seen in May 2014 (Radebaugh et al. 2016). 467

The bubble diameters are relatively small at Masaya, with a median value of 2.6 m and a range of 0.6-8.5 m, even during a period of vigorous activity (Pering et al. 2019). The Reynolds number of a bubble, whose diameter is 2.6 m, is 156 and 30 at the surface and a depth of 100 m,

471 respectively, (Table 3) for a magma viscosity between 140 and 180 Pa.s. The Reynolds number at the surface, calculated above with a negligible gas volume fraction, could be below the threshold 472 of 70-100, if the magma viscosity is multiplied by 2.5, such as for a gas volume fraction exceeding 473 a value of 30 % (Fig. S1c). The Reynolds number at a depth of 500 m, equal to 78 if we assume 474 that the magma still contains 1.45 wt% of dissolved water (Table 3), is just above the lowest value 475 for starting an open wake and initiating a separated gas flow. However, the magma at Masaya, 476 being richer in dissolved water, 1.45 wt%, than that at Ambrym, 1.2 wt%, may have partially lost 477 its initial dissolved water at a depth below than 500 m, up to a significant value. In this case, the 478 magma viscosity at this depth increases, thereby decreasing the Reynolds number. It may then 479 lead to a Reynolds number below the threshold value of 70-100, preventing both the formation of 480 open wakes and that of a separated gas flow at this depth. If we consider that the depth at which 481 the magma is still rich in dissolved water, 1.45 wt%, is 700 m rather than 500 m, the driving 482 spherical cap bubble at Masaya cannot grow an unstable wake at the lowest possible value of the 483 magma viscosity. This is even more the case if the depth, where part of the dissolved water is 484 485 significantly lost, is larger than 700 m, invaliding the possibility to have a separated gas flow by this mechanism at any depth below 500 m. 486

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#### 488 Small lava fountains: detached bubble wakes

489 The existence of small lava fountains, which have been interpreted at Erta 'Ale as the surface expression of newly detached bubbly wakes, could be used as an independent constrain in favour 490 491 of the existence of detached bubbly wakes at the surface (Bouche et al. 2010). Small lava fountains have also been observed at Nyiragongo (Burgi et al. 2014; 2018), Villarica (Palma et al. 492 493 2008) and Ambrym (Radebaugh et al. 2016). This suggests that a mechanism based on the 494 existence of detached bubbly wakes is possible at the surface of these lava lakes, in agreement 495 with relatively high values of the Reynolds number of their spherical cap bubbles close to the 496 surface (Table 3). Furthermore, the results from the sulphur isotopes, which have shown that the degassing at Erta 'Ale occur in disequilibria (De Moor et al. 2013), also reinforce our proposed 497 mechanism of spherical cap bubbles already existing at 500 m and sufficiently large to be able to 498 produce detached bubbly wakes even when the activity at the lava lake is very low (Bouche et al. 499 2010). 500

501 Small lava fountains have been observed at Ambrym at a time where the bubbles were 12-16 m 502 in diameter (Radebaugh et al. 2016). Their Reynolds numbers (Table 3) are either 1190-1820 or 503 119-182, when considering that the small bubbles suspended in the liquid phase move separately 504 and at a low gas volume fraction or together with the liquid at a gas volume fraction of 60 %, respectively. These Reynolds numbers, well above the value of 70-100, cannot be used to give anindication on the gas volume fraction at the surface.

The lack of small lava fountains at Masaya during all periods of observations (P. Delmelle, pers. 507 com., 2 November 2021) suggests that the Reynolds number at the surface is always below 70. 508 The Reynolds number of the typical bubble observed during a vigorous period, 2.6 m in diameter 509 (Pering et al. 2019), is equal to 156 when assuming a negligible gas volume fraction (Table 3). 510 This suggests that the gas volume fraction at the surface must always exceed 30 %. Such a value 511 can easily be explained by the richness in volatiles of the magma. The degassing at Masaya 512 (Martin et al. 2010) has been shown to be at equilibrium, when using the sulphur isotopes during a 513 period of very low activity (De Moor et al. 2013). The degassing at equilibrium is in agreement 514 with the lack of open wakes below a depth of 700 m and a large gas volume fraction close to the 515 surface. But we do not know whether the degassing at equilibrium is only occurring at periods of 516 517 low activity or is persistent at the lava lake.

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#### 519 Homogeneous versus separated gas flow?

We now discuss Ambrym and Masaya as typical examples. The records from open-path Fourier 520 521 transform infrared spectroscopy of gas emissions at Ambrym suggest that two types of degassing exist. The first one is associated to a separated gas flow, corresponding to the large bubbles 522 observed at the surface and shown to exist already at depth (Allard et al. 2016b). This is in 523 agreement with our proposed model of spherical cap bubbles and their detached bubbly wakes 524 525 already existing at a depth of 500 m (Fig. 11b and 11c). The second type of degassing at Ambrym is due to a coupled magma-gas ascent, leading to a very shallow degassing of puffs, hence a 526 separated gas flow, possibly occurring at depths above several tens of meters below the surface 527 (Allard et al. 2016b), i.e. within the lava lake. The coupled degassing, and its resulting separated 528 gas flow within puffs, can easily be produced if the gas volume fraction associated to the small 529 530 bubbles is large, thereby favouring a local bubble coalescence or their simultaneous breaking at the surface. The relatively frequent multimetric CO<sub>2</sub>-rich bubbles, 4 detected during 1h20min, on 531 5 October 2008 during a moderate period of activity (Allard et al. 2016b), suggests that the 532 separated gas flow may play a significant role in the degassing pattern, thereby reducing the gas 533 volume fraction in the conduit from that estimated by assuming gas exsolution within a fully 534 homogenous gas flow. 535

The Reynolds number of the bubbles, with a diameter of 5-10m at a depth of 100 m, 61-172 is just slightly below the value of 70-100 (Table 3), suggesting that most of the time an open wake can also exist at 100 m depth during this period. The existence of detached bubble wakes at 100 m

depth is even less questionable within the framework of the model proposed by Bouche et al. (2010) and the value used for the magma viscosity, when the bubbles are larger, 12-16 m in diameter (Radebaugh et al. 2016), due to their relatively large Reynolds number, 226-348, respectively (Table 3). We therefore propose that the degassing at Ambrym is mostly driven by a separated gas flow, induced below a depth of 500 m (Fig. 11b and 11c).

In the same line of reasoning, we could explore at other lava lakes the consequences of having 544 large bubbles at a depth of 500 m on the degassing pattern in the conduit. The calculation of the 545 Reynolds numbers is particularly revealing as it can show whether a wake can exist or no with a 546 threshold value of 70-100 and whether a wake is closed and laminar or open and turbulent (Fan 547 and Tsuchiya 1990). Among our choice of lava lakes, we note that the lava lake at Villarica 548 presents the largest values of the Reynolds numbers associated to driving large bubbles, even for 549 the largest value of the viscosity associated to the undegassed magma, 18 Pa.s (Table 3). The 550 551 Reynolds numbers of 556-1990 are slightly below the transition for a fully turbulent rise for the large bubble, 2000 (Bouche et al. 2010). It only requires a decrease in viscosity by a factor 2 to 552 553 place a typical driving large bubble, such as those observed (Gurioli et al. 2008), in the fully turbulent regime. This could perhaps be envisioned, as the maximum value of the initial dissolved 554 555 water is very high, 4 wt% (Witter et al 2004).

In summary, the main restriction is related to the assumption of equilibrium degassing, shown to be appropriate at Masaya and incorrect at Erta 'Ale (De Moor et al. 2013). This is perhaps a possibility to consider or discard in the future, given the suggestion that the mixing of fluids between the degassing and degassed magma is turbulent at Villarica in contrast to Erebus and Kilauea (Moussallam et al. 2016).

The calculation of the Reynolds numbers, associated to a given period during which the 561 measurements of the bubble volumes are performed, is dependent of the vigour of the lava lake, 562 which can anywhere between quiet or vigorous. This variation in activity emphasises the need for 563 564 continuous measurements. However, the technology to set continuous measurements even at lowdanger lava lakes is complicated and/or costly, as for the records from open-path Fourier 565 566 transform infrared spectroscopy, or the data analysis is very long to perform automatically and requires enormous amount of computer memory and storage, as for the video's measurements in 567 the visible, infra-red or ultra-violet range of recordings. 568

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## 570 Strombolian activity at Erebus, Stromboli and Yasur

571 Petrology and magma viscosity

572 Erebus, Yasur and Stromboli are textbook examples of open-vent volcanoes that maintained long-573 term steady-state behaviour (centuries). As the magma is continuously degassing and losing 574 temperature at the vents it crystallises and its viscosity significantly varies.

At Erebus, the mineral assemblage found in phonolitic lavas and tephra has not changed in 575 composition for at least ~17 ka (Kelly et al. 2008). At 950 ±25°C, the anorthoclase bearing 576 phonolite is almost dry (H<sub>2</sub>O <0.2 wt% versus 1-2 wt% in the parental basanite) (Moussallam et al. 577 2013; 2015; Sims et al. 2021), its viscosity is calculated to be 10<sup>6.3</sup> Pa.s (Le Losg et al. 2015) for 30 578 vol% anorthoclase, an amount of megacrystals that is kept constant through time (Kelly et al. 2008). 579 580 The temperature of the anorthoclase-bearing phonolite at Erebus is relatively high with respect to those of water-rich phonolites. Values of 760-780°C and 840-870°C are reported for the phonolitic 581 magma of Laacher See (Rout and Wörner 2020) having 4.0-3.6 wt% H<sub>2</sub>O (Harms and Schmincke 582 2000); 800-850°C for 3.0-2.5 wt% H<sub>2</sub>O at Tenerife, 935°C for 3 wt% H<sub>2</sub>O at Tambora; and 583 815±10°C for the 79AD phonolite having 6.2 wt% H<sub>2</sub>O at Vesuvius and even lower (785-800°C) 584 for the other Vesuvius phonolites for a range of H<sub>2</sub>O content between 7.5 and 3.5 wt% H<sub>2</sub>O 585 (Andujar and Scaillet 2012). Moussallam et al. (2015) propose that the CO<sub>2</sub> flux sustains the 586 587 convective motion observed at Erebus with recycling of degassed magma in accordance with a bidirectional flow model in the conduit (see details in separated section). This is in agreement with 588 589 the coexistence of two magma batches differing by their temperature and viscosity in a single conduit up to a depth of ~30 m, where the descending cold magma could be at least 100°C colder 590 591 than the ascending flow of bubbly magma coming from 2-3 km (Molina et al. 2015). A temperature range from 950°C to 1000-1050°C would signify a change in viscosity of at least one order of 592 593 magnitude  $(10^{6.3}-10^{5.3} \text{ Pa.s and possibly as high as } 10^7 \text{ Pa.s})$  (Le Losq et al. 2015.)

At Yasur, Firth et al. (2014) suggested that strombolian-style activity has persisted in its current 594 form for the last 630-850 years from their study of recently exposed tephra sequences. They show 595 that the emission rate is constant with  $\sim 5.7 \times 10^5$  kg per day expelled magma volume of basaltic 596 tranchyandesitic magma. Neither the crystallisation temperatures nor the major and trace element 597 composition of Yasur magma vary significantly over nearly 800 years activity (Firth et al. 2014). 598 The magma supply rate, calculated on the basis of the SO<sub>2</sub> flux and the sulphur content of olivine 599 600 hosted melt inclusions (Tables 1 and 2), achieves nearly 0.05 km<sup>3</sup> per year (Métrich et al. 2011; Woitischek et al. 2020). Woitischek et al. (2020) propose when assuming equilibrium conditions, 601 602 that the degassed shallow conduit at Yasur consists of a crystal-rich region with a thickness of at least 0.6 and up to 2 km, if extensive crystallisation is driven by H<sub>2</sub>O degassing at equilibrium 603 following Métrich et al. (2011). In these conditions, the proportion of crystals increases from 10 to 604 32 vol% on degassing. Calculations, using a software called MELTS (Ghiorso and Gualda 2015), 605

indicate that the bulk magma effective viscosity would vary from 5.8x10<sup>2</sup> Pa.s prior to degassing to 606 1.5x10<sup>5</sup> Pa.s after degassing and crystallisation, with H<sub>2</sub>O from 1.1 wt% to 0.2 wt% respectively 607 (Woitischek et al. 2020). While the trachyandesitic magma at Yasur is degassing and crystallising, 608 the SiO<sub>2</sub> concentration of the residual melt changes to reach 59 wt% (Métrich et al. 2011). 609 Simultaneously, its viscosity becomes 10<sup>3.4</sup> Pa.s for a water concentration of 0.3-0.2 wt% and using 610 Giordano et al.'s model (2008). A volume fraction of crystals of  $\sim 30$  % increases the viscosity by 611 0.5-1.0 log unit depending on their aspect ratio (Mader et al. 2013; Le Losq et al. 2015). It means 612 that the Yasur conduits may be capped by a relatively viscous magma in agreement with 613 equilibrium calculation of Woitischek et al. (2020). 614

At Stromboli, a magma supply rate of  $\sim 0.1 \text{ km}^3$  per year is needed to sustain the long-term 615 average of SO<sub>2</sub> flux (220-260 tons per day (Allard et al. 2008) (Tables 1 and 2). The major element 616 compositions of most degassed shoshonitic pyroclasts vary in a very restricted range (Pompilio et 617 618 al. 2011) as do the percentage of crystals (Landi et al. 2004; 2006; 2009). In contrast, a complex zoning of phenocrysts, particularly of plagioclase, has been reported for a long time, supporting 619 620 processes of magma mixing and up and down motion. As an example, the modal analysis of a series of 12 crystal-rich scoriae samples covering a period from 1986 to 2000 indicates 48 to 55 vol% 621 622 crystals made of plagioclase (27-38 vol%), clinopyroxene (16-10 vol%) and olivine (2-5 vol%) (Landi et al. 2004). Modal analysis performed on 9 lava samples of the 2002-2003 lava flow 623 provide similar results with 41 to 54 vol% (38 vol% in one sample) and similar proportions of 624 plagioclase (23-34 vol%), clinopyroxene (5-19 vol%), olivine (2-8 vol%) (Landi et al. 2006). In the 625 626 2007 lava samples (9 samples analysed) the proportion of phenocrysts varies from 45 to 51 vol% (plaglioclase 29-33 vol%; clinopyroxene 11-14 vol%; 4-5 vol%) (Landi et al. 2009). The melt 627 viscosity reaches 10<sup>2.7</sup> Pa.s at 1100°C, when using Giordano et al.'s model (2008) and a water 628 concentration of 0.3-0.2 wt%. 629

Erebus, Yasur and Stromboli conduits may be capped by degassed and relatively viscous 630 631 magmas (10<sup>4</sup> Pa.s) when assuming equilibrium conditions. They contain a significant proportion of crystals from nearly 30 vol% (Yasur, Erebus) up to 55 vol% (Stromboli), for temperature range 632 from 950°C (Erebus) to 1110-1120°C (Yasur, Stromboli). At Stromboli, a crystal content of 45 vol 633 % corresponds to an increase in viscosity by a factor of 0.6-0.9 log unit, while a crystal content of 634 55 vol% leads to an increase by 1.2 to nearly 1.8 orders of magnitude (Fig. S1c). In these 635 conditions, the magma existing during the classical strombolian activity has a viscosity of 500-820 636 637 Pa.s and 650-1000 Pa.s for 45 vol% and 55 vol% of crystals, respectively.

#### 639 Input from surface measurements

640 Erebus, Stromboli and Yasur are well-studied strombolian volcanoes.

The description and quantification of the surface activity at Erebus, which can safely be 641 observed from above (Fig. 6) are superbly done either using gas studies (Oppenheimer and Kyle 642 2008; Oppenheimer et al. 2009; 2011; Boichu et al. 2010; Ilanko et al. 2015a; 2015b) and/or 643 geophysical studies, which include videos, infrasound, seismicity and radars (Rowe et al. 2000; 644 Dibble et al. 2008; Johnson et al. 2008; Jones et al. 2008; Aster et al. 2003; 2008; Gerst et al. 645 2008; 2013). The degassing at Erebus (Dibble et al. 1988; Rowe et al. 2000; Dibble et al. 2008) is 646 also mainly two-fold. It takes the form of rare explosions (Johnson et al. 2008; Peters et al. 647 2014b), whose gas is initially strongly overpressurised, here called active degassing, and of more 648 frequent (each 11 min) and much smaller gas releases, associated to a very small gas overpressure 649 (blue cloud on photo small bubble breaking on Fig. 6; Oppenheimer et al. 2009; Peters et al. 650 651 2014b; Ilanko et al. 2015a and 2015b), hence here called passive degassing, and similar gas puffing. Gas studies have been restricted to short continuous periods, the longest lasting 10 hours 652 653 (Fig. 6d; Oppenheimer et al. 2009) and mostly mostly focused on the passive degassing (Oppenheimer and Kyle 2008; Peters et al. 2014b, 2018; Ilanko et al. 2015a) rather than on the 654 655 explosions, which only occurred on average 3-6 times a day (Jones et al. 2008; Oppenheimer et al. 2009; Oppenheimer et al. 2011; Ilanko et al. 2015b). In contrast, infrasonic, seismic and radar 656 measurements, which have been run continuously for several days, were able to constrain many 657 physical parameters, such as velocities and gas volumes, and mechanisms occurring at the vent 658 659 (Johnson et al. 2008; Jones et al. 2008; Aster et al. 2003; 2008; Gerst et al. 2008; 2013). Estimates of gas volume from explosions have been obtained from infrasonic measurements to vary between 660  $10^3$  and  $3 \times 10^4$  m<sup>3</sup> (Johnson et al. 2008; Table 3). Gas puffing, although markedly smaller than 661 explosions (Fig. 6d), could nevertheless significantly contribute to the gas fluxes due to their 662 occurrence (Oppenheimer and Kyle 2008; Peters et al. 2014b, 2018; Ilanko et al. 2015a). Aster et 663 al. (2003; 2008) showed that the Very Long Period of the seismic records (VLP) at Erebus could 664 be very well explained by the response of the magma column to the ascent of a Taylor bubble 665 666 (Fig. 3c) and its subsequent disequilibrium, in agreement with the laboratory experiments (Ripepe et al. 2001). The moment tensor inversion of the VLP events provided an estimate of the average 667 displaced mass of magma,  $8 \times 10^7$  kg for an acceleration of 1 m<sup>2</sup>/s (Aster et al. 2008). This mass 668 can be contained in a conduit radius of 5.6 m, when assuming a magma column height of 400 m 669 (Aster et al. 2008), in perfect agreement with observations of the conduit radius at the surface, 5-670 10 m (Dibble et al. 2008; Oppenheimer et al. 2009). The oscillation in volume of the rising Taylor 671 bubble (Fig. 3c), detected by using an array of 3 radars, showed that most of their gas overpressure 672

in the conduit is released by these precursory oscillations rather than during their bursting (Gerst et 673 al. 2013). They also show that the sound of the explosion is related to the bursting rather than by 674 the bubble volume oscillation, in contrast to what had been proposed at Stromboli (Vergniolle and 675 Brandeis 1994; 1996). The difference in the bursting mechanism of large bubbles between Erebus 676 and Stromboli is likely to result from the large differences in magma viscosity, which are 10<sup>63</sup> Pa.s 677 (Le Losq et al. 2015) and 10<sup>2.7</sup> Pa.s calculated using Giordano et al. (2008), respectively. The case 678 of Erebus illustrates the importance of having an accurate estimate of the magma viscosity. The 679 Taylor bubbles (Fig. 3c) associated to the explosions have been considered to be able to grow a 680 wake at their bottom, when using the viscosity previously estimated to be ~1600 Pa.s (Kyle et al. 681 1992; Calkins et al. 2008), a value now considered as too low (Le Losg et al. 2015). However, this 682 mechanism is now shown to be impossible when using the updated value of the magma viscosity 683 of 10<sup>6.3</sup> Pa.s (Le Losq et al. 2015), because their Reynolds number is very small (Table 3). This 684 drastic change in the interpretation of the degassing pattern, at the surface and in the conduit, 685 shows the key role of magma viscosity in understanding the eruptive behaviour. The result 686 687 associated to this updated magma viscosity is in perfect agreement with the finding of Oppenheimer et al. (2009) and Moussallam et al. (2016), who have suggested that the degassing at 688 689 Erebus is driven by a stable bi-directional flow, in which the new batch of magma regularly replaces the degassed sinking magma. 690

691 The degassing at Stromboli has been extensively described (Fig. 12) and quantified by gas and geophysical studies (Harris and Ripepe 2007; Ripepe et al. 2008; Ripepe et al. 2005; Del Donne 692 and Ripepe 2012; Pering et al. 2020; Thivet et al. 2021). The main techniques, which have been 693 used separately or simultaneously, consist into 1) videos in the visible and infrared range and 694 rarely high-speed cameras (Patrick et al. 2007; Harris et al. 2012; Taddeucci et al. 2012a, b; Harris 695 et al. 2013; Gaudin et al. 2014a; 2014b; Bombrun et al. 2015; Gaudin et al. 2017); 2) DOAS, 696 multi-gas, open-path Fourier transform infrared spectroscopy, ultraviolet-cameras for gas studies 697 698 (Burton et al. 2007; Aiuppa et al. 2010; Tamburello et al. 2012), 3) infrasound, seismicity and radar for geophysical studies (Weill et al. 1992; Vergniolle and Brandeis 1994; 1996; Chouet et al. 699 1999; 2003; Scharff et al. 2008; Genco and Ripepe 2010; Del Donne et al. 2016; Patanè et al. 700 701 2017) and 4) laboratory experiments and models (Lane et al. 2013; Del Bello et al. 2012; 2015). It is now well understood that the degassing at Stromboli is also mainly two-fold, with the weak 702 puffs, which occurs quasi-permanently at the central vents but can also exist at times at the 703 northeastern and southwestern vents (Harris and Ripepe 2007; Gaudin et al. 2017; Pering et al. 704 2020) and the explosions, whose physical parameters, such as ejecta and gas velocities or gas 705 706 volumes, are now well known (Harris and Ripepe 2007; Patrick et al. 2007; Harris et al. 2012;

Taddeucci et al. 2012a, b; Gaudin et al. 2014a; 2014b; Bombrun et al. 2015; Thivet et al. 2021). 707 Gas volumes from explosions have been estimated to be 2-100 m<sup>3</sup> (Vergniolle and Brandeis 1996), 708 20-35 m<sup>3</sup> (Ripepe and Marchetti 2002), 400 m<sup>3</sup> (Tamburello et al. 2012) whereas the gas volumes 709 in the puffs only reach 19-211 m<sup>3</sup> (Harris and Ripepe 2007), 1.4-21.5 m<sup>3</sup> (Tamburello et al. 2012), 710 1-100 m<sup>3</sup> (Gaudin et al. 2017). The occurrence of puffs, each 1 to 2 seconds, is responsible for the 711 major contribution of passive degassing to the total gas flux, > 95 %, (Ripepe and Marchetti 2002; 712 Ripepe et al. 2008; Gaudin et al. 2017). Explosions and puffs at Stromboli differ significantly in 713 their gas overpressure, large or small for explosions and puffs, respectively, and in a lesser extent 714 on their gas volumes, large or small for explosions and puffs, respectively. They also differ in their 715 depth of formation, 3 km for the deepest explosions and very shallow (11-150 m), for puffs 716 (Burton et al. 2007). Hence explosions correspond to active degassing while puffs belong to the 717 passive degassing. 718

To a lesser extent among the well-studied volcanoes, there is Yasur (Fig. 13a). Gas studies 719 (Bani and Lardy 2007; Bani et al. 2012; Oppenheimer et al. 2006; Woitischek et al. 2020), visible, 720 721 thermal and high speed videos (Bani et al. 2013; Gaudin et al. 2014a; 2014b; Gomez and Kennedy 2017) together with seismic, infrasonic and radar measurements (Nabyl et al. 1997; Battaglia et al. 722 723 2012; 2016a, b; Kremers et al. 2013; Marchetti et al. 2013; Spina et al. 2016; Meier et al. 2016; Jolly et al. 2017; Iezzi et al. 2019; Simmons et al. 2020; Perring et al. 2019; Fee et al. 2021) have 724 in the past 20 years help to constrain the eruptive pattern. Some of these studies have also shown 725 two classes of degassing (Oppenheimer et al. 2006; Bani et al. 2013; Meier et al. 2016; Kremers et 726 727 al. 2013; Marchetti et al. 2013; Spina et al. 2016; Simmons et al. 2020; Fee et al. 2021), although in practice a continuum of degassing events may exist (Gaudin et al. 2017). The frequent and 728 weakly overpressurised events on one side and the less frequent but strongly overpressurised 729 events on the other side are also examples of passive and active degassing, referred by Spina et al. 730 (2016) as minor and major events. Explosions at Yasur are very frequent, one per 1-2 min (Spina 731 732 et al 2016; Woitischek et al. 2020), while the minor events occurs as frequently as one event each 1-5 min (Bani et al. 2013; Fee et al. 2021) to 40 per min (Spina et al. 2016). The gas volumes of 733 the strong explosions have been estimated for at 5-80x10<sup>3</sup> m<sup>3</sup> from infrasound modelling during a 734 vigourous period (Iezzi et al. 2019) and on average at 14.25x10<sup>3</sup> m<sup>3</sup> from open-path Fourier 735 transform infrared spectroscopy combined with SO<sub>2</sub> flux measurements during an average period 736 (Woitischek et al. 2020). 737

These three laboratory open-vent volcanoes clearly show that degassing consists in both a strong and a weak degassing, here referred to as active and passive degassing, respectively. In this respect, the active degassing at Erta 'Ale (spherical cap bubble; Fig. 3a) shares a similar origin
than explosions at Stromboli and Erebus and to the major explosions at Yasur.

The strength of active degassing can be quantified by the maximum gas flux. Here we report 742 values that can be compared together because they all derive from modelling the acoustic waves. 743 The maximum gas fluxes reach 510 m<sup>3</sup>/s at Erta 'Ale (Bouche et al. 2010), 680 m<sup>3</sup>/s at Stromboli 744 (Vergniolle and Brandeis 1996),  $3x10^3$  -  $6x10^4$  m<sup>3</sup>/s at Yasur (Iezzi et al. 2019), and  $7x10^3$  -  $7x10^4$ 745 m<sup>3</sup>/s at Erebus (Johnson et al 2008). It is tempting to relate the difference in strength to the 746 respective viscosities of undegassed and degassed magma that range from 22-27 Pa.s (Erta 'Ale), 747 500 -  $10^3$  Pa.s (Stromboli), 5.8 ×  $10^2$  -  $1.5 × 10^5$  Pa.s (Yasur), up to  $10^{6.3}$  Pa.s (Erebus). Similarly, 748 the small gas release at Erebus (blue cloud; Fig. 6) should be compared to the puffs of Stromboli, 749 750 to the minor events at Yasur or to the flat bubbles at Erta 'Ale, although at Erta 'Ale additional 751 effects exist because of the cold crust.

This interpretation of the strombolian activity with a well identified two-fold degassing could be applied more broadly. Therefore time series need to be much larger than the typical return time of active degassing, if the goal is to assess its associated eruptive pattern. This may become a challenge at Erta 'Ale, from the relatively short time series possible in gas studies (Sawyer et al. 2008), as well as at Erebus, due to rarity of explosions, 4-6 per day (Rowe et al. 2000; Burgisser et al. 2012; Jones et al. 2015) and at Villarica (Sawyer et al. 2011; Moussallam et al. 2016).

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# 759 Vulcanian activity at open-vent volcanoes

Our selection of open-vent volcanoes, associated to intermediate to high-silica magma 760 compositions, may present short vulcanian explosions. It occurs at Arenal (Mora et al. 2022; Fig. 761 13b-d), Fuego (Morrissey et al. 2008; Marchetti et al. 2009; Diaz-Moreno et al. 2020), 762 Popocatépetl (Schaaf et al. 2005; Fig. 14a-b) and Santiaguito (Fig. 14c; Rose 1972; Rose et al. 763 1987; Harris et al. 2002; 2003; Bluth and Rose 2004; Johnson et al. 2004b; Sahetapy-Engel et al. 764 2008; Sahetapy-Engel and Harris 2009a, 2009b; Marchetti et al. 2009; Johnson et al. 2011; Scott 765 et al. 2013). A vulcanian explosion can be produced by the disruption of a solid cap suddenly 766 failing under a significant underlying gas overpressure. The overpressurised gas is suddenly 767 768 released and often leads to the formation of small eruptive columns of a few kilometers high or 769 above and sometimes with pyroclastic dense flows (Clarke et al. 2013). The gas ejection at the vent is short compared to the rising time of the eruptive columns, making the initial injection of 770 the gas in the atmosphere sustained by an initial pulse of overpressurised gas rather than by a 771 continuous gas flow feeding the eruptive column. 772

773 The pre-eruption pressure and ejection velocity are relatively high, <1-10 MPa and 400 m/s (Morrissey and Chouet 1997; Arciniega-Ceballos et al. 1999, 2012; Chouet et al. 2005; Clarke et 774 al. 2013). A shock wave, i.e. a pressure discontinuity, is produced in air upon disruption of the 775 plug and travels into the atmosphere at a velocity exceeding that of the sound, while a 776 decompression-induced fragmentation propagates down into the conduit (Kieffer 1981; Turcotte et 777 al. 1990; Woods 1995). This fragmentation, inducing the formation of ejecta, only occurs in the 778 shallowest portion of the conduit (Clarke et al. 2013). The magma left in the conduit cools until 779 being sufficiently cold and rigid to seal the top of the magma column. Meanwhile, some 780 vesiculation occurs below, inducing the building-up of pressure, ultimately responsible for the 781 next vulcanian explosion (Clarke et al. 2013). 782

Several volcanoes, such as Arenal (Mora et al. 2022), sometimes Fuego (Martin and Rose 783 1981; Morrissey and Mastin 2000), and perhaps Santiaguito with the small and strong ash-and-gas 784 785 explosions, are able to switch between periods of vulcanian explosions and periods of strombolian explosions. This implies that a drastic change of conditions, between a mechanically sealed vent 786 787 and an open vent, can occur at the vent, although perhaps one vent can be sealed while another one is simultaneously unsealed, as recently proposed for Arenal (Mora et al. 2022). The 788 789 conditions, which prevail at such transitions, have not yet been clearly identified. Alternatively, Clarke et al. (2013) had proposed that the difference between strombolian and vulcanian 790 791 explosions results in the bubble motion, largely separated or stagnant in the magma, respectively.

Popocatépetl is also remarkably surprising because the dome prior to a vulcanian explosion is 792 793 seen rising and growing into the conduit (R. Campion, pers. com., 7 July 2021; Fig. 14a and 14b), invalidating the conditions of a fully sealed vent. Furthermore the SO<sub>2</sub> flux is always quite high, 794 (Table 2b; Campion et al. 2018). This high  $SO_2$  flux exists not only during vulcanian explosions 795 but also between 2 successive vulcanian explosions. This observation also appears to be 796 797 incompatible with the condition of a fully mechanically sealed vent. The detection by a thermal 798 camera of one or two circular hot zones, of 20-50 m in diameter, is likely to be associated to the 799 gas percolation (R. Campion, pers. com., 7 July 2021).

In contrast, Santiaguito, which mostly display dacitic domes and sometimes, vulcanian explosions (Yamamoto et al. 2008; Scharff et al. 2012) as at Popocatépetl, has a surprisingly low  $SO_2$  flux (Table 2b). The classification of explosions at Santiaguito is still a matter to debate, from a closed vent prior to explosions to a fully open vent. Explosions are considered as weak vulcanian explosions (Johnson et al. 2011), or referred to as small to moderate gas-and-ash explosions (De Angelis et al. 2019; Wallace et al. 2020), rather than strombolian explosions (Bluth and Rose 2008), while the vent is sometimes considered as fully open for degassing (Holland et al. 2011).

The level of the dome at the conduit varies at different periods (Lamb et al. 2019; Gottschammer 807 et al. 2021). To our knowledge, no photographic evidence exists of a high-level dome being 808 systematically emplaced prior to a vulcanian explosion. But the dome extrusion, often cyclic, is a 809 frequent feature of its long-lasting eruption (Rose 1972; Rose et al. 1987; Harris et al. 2002; 810 2003). The initiation of a vulcanian explosion had been beautifully catched by a video taken from 811 above, showing that an explosion starts with a very weakly overpressurised degassing on an 812 annulus prior to the entire disruption of the solid cap in one unit from the center of this circle, as 813 initially quantified by Bluth and Rose, (2008). This may be indicative that the eruptive pattern is 814 similar at Popocatépetl and Santiaguito, although with two contrasting SO<sub>2</sub> fluxes (Table 2b) and 815 perhaps different amplitudes in the variations of level of magma in the conduit. The strength in the 816 817 explosions is also different at Popocatépetl and Santiaguito, typical and weak, respectively (Marchetti et al. 2009; Johnson et al. 2011; Scharff et al. 2012). 818

819 The differences between vulcanian and strombolian explosions have been tentatively quantified by looking at the simultaneous thermal and infrasonic records performed at Villarica, Stromboli, 820 821 Fuego and Santiaguito (Marchetti et al. 2009). Strombolian explosions at Stromboli and Villarrica present relatively low values of thermal radiation energy, consistent with the short explosive 822 events of these two volcanoes. In contrast, the longer-lasting thermal waveforms observed for 823 explosions at Santiaguito and Fuego result in the longer ash plume dynamics, which corresponds 824 to high values of thermal radiation energy. However, Santiaguito and Villarrica share similar 825 infrasonic energies, both being lower than the values produced by explosions at Stromboli and 826 827 Fuego (Marchetti et al. 2009). This may be associated to the very low level of activity both at Villarrica and Santiaguito during this period. The relationship between infrasonic energy, used as 828 a proxy of gas thrust, and the difference between infrasonic and thermal energies, used as a proxy 829 of buoyancy, shows that strombolian and vulcanian explosions are organised in two very different 830 clusters of points, suggesting a method to fully quantified these types of activity (Marchetti et al. 831 832 2009). Furthermore, combining thermal and infrasonic records on open-vent volcanoes opens an avenue for a quantified description of surface activity. These joined records could be used for 833 deciding whether the vent is properly sealed or open and also for understanding how the 834 transitions between regimes can occur at a single volcano. 835

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### 837 Thermal models

838 Open-vent volcanoes must release both gas and heat in order to maintain a persistent activity and 839 one key ingredient to maintain an open-vent volcano active is to reduce the heat loss. We discuss below three mechanisms by which heat transfer can occur, mainly conduction, thermal convectionand arrival of hot magma through bubbly wakes.

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## 843 Thermal models of lava lakes

The radiated heat output had been measured at several volcanoes, mostly at the surface of lava lakes in relatively close proximity, such as at Erta 'Ale (Burgi et al. 2002; Harris et al. 2005; Spampinato et al. 2008), Erebus (Calkins et al. 2008), Nyiragongo (Spampinato et al. 2013) and Kilauea (Spampinato et al. 2012). But thermal measurements have also been done by satellites (Ramsey and Harris 2013), and routinely such as by MODIS (Harris et al. 1999; Copola et al. 2016; Vergniolle and Bouche 2016; Wright 2016; Aiuppa et al. 2018; Naismith et al. 2019) or by SEVERI (Field et al. 2012).

Thermal models of lava lakes, initially performed without considering the convection (Peck et al. 851 852 1977), aimed at modelling the cooling of lava lakes, thereby estimating the thickness of the crust as a function of time and quantifying the role of small bubbles in stopping of convection (Worster et 853 854 al. 1993; Davaille 1993). The quantification of the heat budget of a lava lake or a magma reservoir requires the knowledge of the in and out heat fluxes. The heat transfer occurs by conduction for thin 855 856 lava lakes as well as by thermal convection, if its thickness exceeds a critical value (Davaille 1993). 857 The cooling is very slow for a sole thermal conduction, such as for the thinnest lava lakes at Kilauea (Davaille 1993) and at Erebus, whose magma depth in the crater area is only of a few tens of meters 858 (Jones et al. 2015). Once the thermal convection sets in, the heat losses are drastically enhanced. 859 Several additional factors can prevent thermal convection to exist, such as a gas volume fraction 860 exceeding 0.6 vol% (Davaille 1993), as for the cooling of the Makaopuhi lava lake (Hawaii), or the 861 settling of large crystals (Carrigan 1987) as for the Kilauea Iki lava lake (Helz, 1987; Davaille 862 1993; Jellinek and Kerr 2001). Additionally, the thickness of the solid crust growing at the surface 863 has been modelled under the assumption of a strong thermal convection within the lava lake 864 (Worster et al. 1993). 865

The temperature contrast between the magma within the lava lake and its surroundings rocks 866 867 should lead to its cooling if given a sufficient time. This occurs briefly at the lava lake of Erta 'Ale for several months around December 2004 (Fig. 7c). At this time, the degassing occurred at the site 868 of the 3 hornitos, located above the underlying conduits (Fig. 7c; Vergniolle and Bouche 2016). 869 Furthermore, the existence of hornitos, which could result from the local accumulation of ejecta 870 871 able to reach the surface of the partially encrusted lava lake rather than falling back within the lava lake, showed that the Strombolian explosions were still ongoing at this time. Therefore, lava lakes 872 873 never cooled entirely, hence could restart as a hot lava lake when the solid crust disappeared.

The lava lake at Villarrica, when encrusted, has small openings by which the degassing can occur (Gurioli et al. 2008; Goto and Johnson 2011; Johnson et al. 2018). The comparison between Villarrica and Erta 'Ale lava lakes suggests that the openings at Villarrica may be located above the underlying conduits. However the lack of a small structure around the openings at Villarrica suggests that the level of the underlying lava lake is further away from its solid crust than that at Erta 'Ale, as also suggested from acoustic records (Johnson et al. 2018).

The observed lack of cooling of lava lakes has been first explained by the arrival of hot magma 880 brought to the surface by convective processes, driven by a density difference between a volatile-881 rich magma and a degassed magma. This model, here referred to as a bi-directional flow model 882 (Fig. 15) considers that the hot and gas-rich magma rises while the degassed and cold magma falls 883 back in the conduit (Francis et al. 1993; Kazahaya et al. 1994; Stevenson and Blake 1998; 884 Oppenheimer and Francis 1998; Harris et al. 1999; Oppenheimer et al. 2004; Witter et al. 2004; 885 886 Harris et al. 2005; Jones et al. 2006; Huppert and Hallworth 2007; Harris 2008; Palma et al. 2011; Beckett et al. 2011; 2014). The bi-directional flow models of Stevenson and Blake, (1998; Fig. 887 888 15a) and Huppert and Hallworth, (2007; Fig. 15c) were used at Nyiragongo to estimate both the conduit diameter and the net magma flux necessary to maintain its giant lava lake (> $15 \times 10^6$  m<sup>3</sup>) in 889 a molten state (Burgi et al. 2014; 2020). 890

Alternatively the heat could be transferred from depth to the lava lake by the arrival of bubble 891 892 clusters, initially formed in the wake produced at the bottom of the driving spherical cap bubbles during their rise in the conduit (Bouche et al. 2010; Fig. 11). The heat brought to the lava lake is 893 provided by the interstitial magma initially present in the bubbly wake prior to its fragmentation in 894 the lava lake (Bouche et al. 2010) rather than that solely brought by the series of small hot 895 896 bubbles, as incorrectly deduced for Villarrica and Kilauea (Moussallam et al. 2016). The key point 897 of the qualitative thermal model of Bouche et al. (2010) is that the hot magma added to the lava 898 lake had been kept hot because it had travelled from depth to the surface at a relatively large velocity, that of the driving spherical cap bubble. The second point is that this mechanism for heat 899 transfer is only possible if the rise of the large bubbles is sufficiently vigorous to grow an instable 900 901 wake, laminar or a turbulent for Reynolds number of 70(100) - 2000 and >2000, respectively (Bouche et al. 2010). At Erebus, the large bubbles responsible for explosions cannot grow a wake 902 when using the updated value of the viscosity 10<sup>6.3</sup> Pa.s (Le Losq et al. 2015), in agreement with 903 904 the findings that the behaviour of the lava lake is governed by a magma exchange rather than by 905 an underlying gas flux from the reservoir (Peters et al. 2014b). The heat at Erebus is therefore provided either by magma upwards motion in a bi-directional flow in the conduit (Oppenheimer et 906 907 al. 2009, 2011; Moussallam et al. 2015) or by the enthalpy of the shallow reservoir (Aster et al.

2003), located either at 400 m (Aster et al. 2008) or between 80 m and 1300 m, when assuming that the volatile species are at equilibrium (Burgisser et al. 2012), or by a combination of the two. A source of heat, which combined the enthalpy of the shallow reservoir with a bi-directional flow in the conduit, is our preferred interpretation, given the comparatively large conduit diameter at Erebus, 4-10 m (Dibble et al. 2008; Oppenheimer et al. 2009), which exceeds the value of 1 m proposed to balance the heat loss from the conduit walls with the assumed magma supply rate in the conduit at Stromboli (Giberti et al. 1992).

It is particularly important to realise that lava lakes, which expose large hot areas to the surface, have a sole or an additional source of heat, which is provided by the existence of the detached bubbly wakes (Fig. 11), when compared to open-vent volcanoes only presenting tops of magma columns. The thermal effect at lava lakes, associated to the possible existence of detached bubbly wakes, is complicated to calculate in detail because it requires to know both the occurrence of the driving large bubbles on a long time scale, due to fluctuations in the level of activity, and the local gas volume fraction in the conduit.

922 Several main types of degassing at lava lakes can exist. On one side, the existence of detached bubbly wakes at depths of > 1 km and > 500 m-1 km at Erta 'Ale and Nyiragongo, respectively 923 (Table 3; Fig. 11a), suggests that their heat could be a significant source over the entire length of 924 the conduit up to 1 km. Hence it would promote efficient exchange of heat between this depth and 925 the surface during the rise of the driving large bubbles. On the other side, the detached bubbly 926 wakes may mostly exist at depths and be partially or fully suppressed in the conduit, such as at 927 928 Villarrica and Ambrym at times (Table 3; Fig. 11b and 11c), respectively, but still play a key role 929 in providing heat or in decreasing the gas volume fraction by favouring the separated gas flow.

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#### 931 Thermal models in conduit

The first model considered that the magma undergoes thermal convection in the conduit (Carrigan 932 933 1983). In this case, ascending hot (possibly gas-rich) magma mixes with the magma residing in the conduit and the cooling occurs at the expense of the reservoir enthalpy (Carrigan 1983). The second 934 935 model that was applied to Stromboli suggested that thermal convection could be ignored due to both 936 the small diameter of the conduit and the constant rise of large bubbles in the conduit (Giberti et al. 1992). They suggested that the only communication between magma reservoir and conduit occurs 937 after an explosion when the magma level in the conduit had returned to its equilibrium value, an 938 939 interpretation fully validated by laboratory experiments (Ripepe et al. 2001) and by the location of the tremor and Very Long Period events (Chouet et al. 1999; 2003). The numerical model of Giberti 940 941 et al. (1992), which considers a large reservoir located below a small conduit, showed that the

942 persistent state of Stromboli could be very well be explained by a shallow reservoir at a few 943 hundreds of meters below the vent, a result entirely validated by further independent studies 944 (Chouet et al. 1999; 2003; Harris and Ripepe 2007). The same model could be partly applicable to 945 Erebus, which presents a shallow reservoir at a depth of 400 m (Aster et al. 2003; 2008), and a 946 conduit slightly larger than twice that at Stromboli (Dibble et al. 2008; Oppenheimer et al. 2009).

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# 948 **Bi-directional flow models in conduit**

949 Another potential convective mechanism, that may occur in the conduit, is driven by the density difference between a hot and volatile-rich magma, light and rising, and a cold and volatile-poor 950 951 magma, hence dense, flowing in opposite direction (Fig. 15a). This mechanism is based on an exchange flow in the conduit (Kazahaya et al. 1994; Harris and Stevenson 1997; Stevenson and 952 Blake 1998). The excess degassing observed at the surface of Stromboli, estimated when comparing 953 the volumes expelled by gas and magma during strombolian explosions (Chouet et al 1974; Ripepe 954 et al. 1993; Patrick et al. 2007; Ripepe et al. 2008; Delle Donne and Ripepe 2012; Harris et al. 955 2012; Harris et al. 2013; Gaudin et al. 2014b), had been explained by the existence of such a flow 956 loop. This bi-directional flow is a potential mechanism for explaining both the persistent outgassing 957 958 and providing heat for the thermal budget of conduits and lave lakes.

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### 960 Existing models

Several models, now referred as bi-directional flow models (Fig. 15), were developed with various 961 962 assumptions detailed below. Several studies use this model to explain the behaviour of several open-vent volcanoes, without any clear indication on how such a flow loop could be initiated in the 963 964 conduit and maintained for very long periods. The rising magma is degassing on its way towards the surface due to gas exsolution, thereby increasing its positive buoyancy. Laboratory experiments 965 of exchange flows in vertical pipes, performed in conditions for which the net volume flux of 966 ascending magma is matched by the volume flux of the descending magma, have shown the key 967 role of the ratio of viscosities, between that of the more viscous fluid divided by that of the less 968 viscous fluid (Stevenson and Blake 1998; Huppert and Hallworth 2007; Beckett et al. 2011; Palma 969 970 et al. 2011). Models describing the spatial organisation of the upwards and downwards magma flows in the conduit had been called bi-directional flow models (Fig. 15). 971

The pioneer quantitative work of Stevenson and Blake (1998) used laboratory experiments to show that the bi-directional flow is a core-annular flow, in which the less viscous fluid flows in the center and the more viscous flows in an annulus close to conduit walls (Fig. 15a). They present a

theoretical model for the core-annular flow, which is used to give a reasonable estimate of conduit 975 radius at Stromboli (Stevenson and Blake 1998) and Villarrica (Witter et al. 2004). Burton et al. 976 (2007b) added to the core-annular flow (Fig. 15b) the degassing path at each depth along the 977 conduit, constrained by petrology data and SO<sub>2</sub> flux at Stromboli but ignore the crystallisation path. 978 979 The rising magma is modelled to be degassing at equilibrium due to the gas exsolution. The area available for the rise of the degassing magma is calculated by assuming that its velocity is 980 981 maximum (Burton et al. 2007b), an assumption later shown to be incorrect (Beckett et al. 2011; Suckale et al. 2018). Nevertheless, this early model gave an estimate of the radius and the velocity 982 of the rising degassing magma at each depth, as well as those of the falling degassed magma 983 (Burton et al. 2007b). They also provide an estimate of the conduit radius, in agreement with radius 984 observed at the surface. 985

This model was later refined by laboratory experiments (Fig. 15c), performed for both 986 987 vertical and inclined conduits, and a dimensionless analysis, which included lateral gradients in magma viscosity and density (Palma et al. 2011). But the properties of the rising and degassed 988 989 magma were taken to be constant during transport (Palma et al. 2011). This model was used to constrain the degassing and ascent rates of volatile-rich magma, which, when combined with 990 991 petrologic data on magmatic volatile content at Villarrica, led to an estimate of the conduit radius (Palma et al. 2011). Palma et al. (2011) only consider the flow in the lowest conduit, in which the 992 993 degassing occurs with stagnant CO<sub>2</sub>-bubbles and ignore the shallowest conduit for which the bubbles become sufficiently large to have a separated gas flow in respect to the magma, on which a 994 995 significant exsolution of water is added.

996 The model of Huppert and Hallworth, (2007), carefully and extensively performed to assess the effects of each physical parameter, showed that the core-annular flow is the most frequent regime 997 (Fig. 15d). They also noted that the fluid may be highly irregular and may present a varicose 998 999 instability, sometimes leading to the fragmentation of the degassing hot magma in many discrete 1000 blobs (Huppert and Hallworth 2007). When this situation occurs, the flow of the degassing magma is not organised in the conduit as a continuous stream but as a discontinuous series of vertically 1001 1002 spaced blobs of degassing magma. A coefficient flux, defined by a ratio of two dimensionless 1003 numbers, is used to analyse the flux of the exchange flow (Huppert and Hallworth 2007; Palma et al. 2011). This coefficient was experimentally estimated for vertical and inclined flows (Palma et al. 1004 2011). Note that these laboratory experiments were performed with fluid reservoirs at the end of the 1005 pipes to ensure a longer duration for the bi-directional flow in the conduit (Huppert and Hallworth 1006 2007; Beckett et al. 2011; Palma et al. 2011). 1007

When the ratio of viscosities exceeds 75, the bi-directional flow is organised with an 1009 axisymmetric core-annular flow (Fig. 15d), in which the less viscous fluid occupies a cylindrical 1010 core and the denser and more viscous fluid flows downwards in an annulus (Huppert and Hallworth 1011 2007; Beckett et al. 2011). The bi-directional flow is a side-by-side flow, for which both fluids are 1012 1013 in contact with the conduit walls and there is a single interface between them (Fig. 15 d), for a ratio of viscosities less than 117 (Beckett et al. 2011). The side-by-side flow is transitional (Fig. 15d) for 1014 a ratio of viscosities between 5 and 59 and presents simultaneously both a core-annular flow and a 1015 side-by-side flow at different depths (Beckett et al. 2011). In this case, the flow of the degassing 1016 magma is not a continuous stream from depth to the surface (Fig. 15d). 1017

Huppert and Hallworth (2007) use a dimensionless analysis for quantifying the flux of the dense fluid, referred with index u for upper fluid in equations, as a function of the ratio of the two viscosities,  $\gamma$ , and the Reynolds number of the upper fluid, Re<sub>U</sub>. The dimensionless number, called the transport number for exchange flow, *Te*, is defined by:

1022 
$$Te=\mu_U Q/(g\Delta\rho R^4)$$

1023where Q,  $\Delta \rho$  and g are the volume flux in either fluid in the tube of radius R, the difference in1024density between the denser and the lighter fluids and the acceleration of gravity, respectively. For1025low Reynolds number, two equations allow the calculation of the transport number *Te* as a function1026oftheviscosityratio

(Eq. 6)

1027 γ

1029 fluid divided by that of the less viscous fluid). The coefficient is estimated from laboratory 1030 experiment scaled by a dimensionless analysis.

1031 $Te=0.01\gamma \ (\gamma << 1)$ (Eq. 7a)1032 $Te=0.125 \ (\gamma >> 1),$ (Eq. 7b)

1033 whereas for large Reynolds number, the transport number, *Te*, depends on is the value of the 1034 Reynolds number of the upper fluid,  $Re_U$  as :

1035 
$$Te=0.556 \ \Re_U^{-1} \ (Re_U >> l)$$
 (Eq. 8)

1036 where  $Re_U$  is the Reynolds number of the upper fluid.

1037 Beckett et al. (2014) showed a similar relationship, for an axisymmetric core-annular flow, 1038 between the net magma flux of the ascending magma, Q, and the viscosity ratio between the 1039 degassed and the volatile-rich buoyant magma components,  $\beta$ :

1040  $Q = 0.059\beta^{-0.74}(g\Delta\rho R^4/2\mu_b)$  (Eq. 9)

1041 where  $\Delta \rho$  is their difference in density,  $\mu_b$  is the dynamic viscosity of the buoyant component, g is 1042 gravitational acceleration and R is the effective conduit radius.

1043

This model had been applied to Stromboli and combined to the degassing and crystallisation 1044 paths of the ascending and descending magmas, constrained by gas flux and melt inclusion data 1045 (Beckett et al. 2014). They propose, based on their choice of viscosity ratio, that the bi-directional 1046 flow below a depth of 3 km is a core-annular flow, in which the degassing magma rises in the center 1047 1048 while the degassed magma flows around it (Beckett et al. 2014). Above a depth of 3 km, they 1049 suggest that the bi-directional flow is a side-by-side flow, with both ascending and descending 1050 magma adjacent to a portion of the conduit walls (Beckett et al. 2014). In this model, the degassing 1051 due to gas exsolution is assumed to occur at equilibrium and with an implicit assumption that the 1052 small bubbles embodied in the magma are stagnant (Beckett et al. 2014). The implications of such a 1053 hypothesis, also used previously at Stromboli (Burton et al. 2007b) and later at Ambrym (Allard et al. 2016b), are strong, as previously discussed. No theoretical relationship between the exchange 1054 1055 flux and the pressure gradient so far exists (Beckett et al. 2014), limiting volcanological 1056 applications to the use of empirical laws, such as the coefficient flux (Huppert and Hallworth 2007; 1057 Palma et al. 2011) or that of Beckett et al. (2014) used at Ambrym (Allard et al. 2016b). The coreannular flow, the most common flow configuration, is also shown to be inherently bistable, i.e. able 1058 1059 to switch between two steady states, and the regime at work cannot be predicted solely from the material properties of the fluids and the tube geometry alone (Suckale et al. 2018). 1060

The spatial configuration of a bi-directional flow can only be organised as discussed above if no large bubble rises in the conduit and perturb the flow field. The existence of large bubbles rising in the conduit as spherical cap bubbles or Taylor bubbles (Fig. 3b and 3c), and the potential formation of detached bubbly wakes (Bouche et al. 2010; section lava lakes; Fig. 11) at several lava lakes, can erase and replace the above bi-directional flow models when the driving large bubbles are very frequent in the conduit. The above bi-directional flow models can be solely at work in between the rise of two successive large bubbles, 1 hr at lava lake of Erta 'Ale (Bouche et al. 2010).

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# 1069 Implications and examples of bi-directional flows

By definition, a bi-directional flow involves a rising degassing magma and the descending degassed magma, hence is a form of convection. Their relative velocity in the conduit cannot be estimated directly. In the past, it has been derived from the direct measurements of surface velocity at lava lakes, as summarised in Lev et al. (2019). A notable exception is the case of Kilauea, where ejecta were directly sampled during the enhanced activity of 9 April 2008 (Carey et al. 2013). They

allowed the first quantification of the convective velocity and the height of the convective cell in the
conduit (Fig. 15e) to be, 0.02-0.8 m/s and 100-300 m, respectively, from the dimensions of the halo
of re-dissolving bubbles within the downward flow of magma (Carey et al. 2013).

Modelling the bi-directional flow in a conduit implies to know the relative areas of the rising 1078 1079 degassing magma and the descending degassed magma. They vary with depth as a consequence of increasing the gas volume fraction within the degassing magma due to gas exsolution. The 1080 velocities of the rising and degassed magma were estimated at each depth based on the magma 1081 supply rate when assuming equilibrium conditions and full gas exsolution (Burton et al. 2007). The 1082 calculation of the gas volume fraction in the conduit, as estimated for Stromboli (Burton et al. 1083 2007b) and Ambrym (Allard et al. 2016b) is based on two implicit assumptions, one is that the 1084 1085 bubbles are small, hence cannot move in respect to the liquid, and the second is that gas and its 1086 surrounding magma are in chemical equilibrium for degassing.

1087 A certain amount of small bubbles may be produced in the rising magma mainly due to gas exsolution when approaching to the surface. Alternatively the excess in volatiles can migrate with 1088 1089 gas diffusion towards pre-existing bubbles. These bubbles, if small, can either be moving with the rising magma, such as in a homogeneous gas flow, or have their own differential velocity in respect 1090 1091 to the liquid, as in a separated gas flow. In the first case, the calculation of the magma viscosity, used by Beckett et al (2014), should consider the gas volume fraction in the magma as part of single 1092 1093 phase while in the second case, the magma viscosity should only consider the magma without the small bubbles, which are flowing on their own. The first consequence of a separated gas flow is to 1094 1095 reduce the gas volume fraction (Vergniolle and Jaupart 1986), by allowing a larger gas flux in the 1096 conduit. The second consequence is that a separated gas flow can enhance the disequilibrium between dissolved and exsolved gas, due to the relatively fast transport of exsolved gas towards the 1097 surface. In this respect, the sulphur isotopes could be used to assess the amount of chemical 1098 disequilibrium in volatiles, shown to exist at Erta 'Ale in contrast to Masaya (De Moor et al. 2013). 1099

1100 The bubbles present in the rising degassing magma may be sufficiently large to behave as a separated gas flow, or drift flow, in regards to that of the liquid (Vergniolle and Jaupart 1986). In 1101 1102 this case, the gas volume fraction should be calculated from the superficial velocity of the gas phase 1103 and the drift velocity of a bubble, which depends on both the bubble diameter and the liquid viscosity (Vergniolle and Jaupart 1986; Pioli et al. 2012; Pansino et al. 2019). These calculations 1104 rely on a good knowledge of the magma viscosity, estimated without the bubbles present in the 1105 separated gas flow, and of the conduit radius, a second key parameter (Pansino et al. 2019). 1106 Quantifying the gas volume fraction at the surface also requires the independent determination of 1107 1108 the flux of the rising magma and the total gas flux. The magma supply rate derives from the  $SO_2$ 

flux and the melt sulphur content, which implies full magma degassing from depth to surface. The 1109 total gas flux used by Pansino et al. (2019) combines the ratios of di-atomic gases such as  $CO_2/SO_2$ 1110 and H<sub>2</sub>O/ SO<sub>2</sub>, with the SO<sub>2</sub> flux at several volcanoes. However, the total gas flux, i.e. obtained by 1111 adding CO<sub>2</sub>, H<sub>2</sub>O, SO<sub>2</sub> fluxes, and the magma supply rate are not independent, invalidating the use 1112 1113 of the estimates of gas volume fractions obtained by Pansino et al. (2019). Nevertheless, a proper estimate of the gas volume fraction should consider whether the small bubbles are flowing in a 1114 homogenous flow, i.e. moving together with the liquid, or in a separated gas flow. Because of the 1115 key role of the viscosity in understanding the physical process at work in a given volcano, these two 1116 limit cases must be considered, unless being able to measure independently whether the gas flow of 1117 the small suspended bubbles, induced by gas exsolution, is homogeneous or separated. 1118

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## 1120 The key informations derived from sound waves studies

Studies on infrasonic (<20 Hz) and acoustic (>20 Hz) sound waves, applied for the first time to 1121 volcanoes in the 1990's (Vergniolle and Brandeis 1994; 1996; Vergniolle et al. 1996; Garcès et al. 1122 1996; Johnson et al., 1998), can also provide constraints on eruptive patterns. The use of broad-1123 band sensors such as microbarographs (a few tenths of Hz to 200 Hz) have been developed to run 1124 1125 continuously, in contrast to microphones (0.1 and 4 kHz) more complicated to protect outdoor. This technique is now expanding significantly for monitoring the volcanic activity, particularly at 1126 open-vent volcanoes with an "easy" access. The main interest of acoustic measurements is to 1127 provide a remote way to monitor the surface of the magma column and to locate the source of 1128 sound when using an array. The location of the source of the sound, often due to degassing at the 1129 vent (e.g., Johnson 2005; Johnson et al. 2011; Colo et al. 2010; Marchetti et al. 2013; Kondo et al. 1130 1131 2019), can also be produced by the front of a lahar, such as at Santiaguito (Johnson et al. 2011) or at Villarica, thereby providing an estimate of the flow front in time (Johnson and Palma 2015; 1132 1133 Johnson et al. 2015). The frequency of the infrasound had been used at Villarica as a mean of 1134 following the level of the lava lake (Goto and Johnson 2011; Richardson et al. 2014), which rises 2 days before the paroxysmal phase of 2015 (Johnson et al. 2018). However, this analysis was 1135 only possible because of a very specific geometry of the lava lake, mostly encrusted with a small 1136 opening in the crust (Johnson et al. 2018). The satellite SAR imagery at Nyiragongo, which can 1137 constrain the lava lake by measuring the length of the shadow cast by the rim of the pit crater 1138 1139 hosting the lava lake, can be joined to the analysis of a seismic and infrasonic records to deduce the lava lake level in the future from a single seismic and infrasonic station (Barrière et al. 2018). 1140 Combining infrasonic and thermal measurements allows the recognition of vulcanian and 1141

strombolian explosions (Marchetti et al. 2009), whereas it can be difficult from acoustic records 1142 alone (Mora et al. 2022). This approach could be particularly relevant in the future for open-vent 1143 volcanoes with a high-silica magma composition, if run continuously. These volcanoes can either 1144 produce vulcanian explosions when the vent is mechanically sealed, strombolian explosions when 1145 1146 the vent is mechanically open, and small to moderate gas-and-ash explosions as shown at Santiaguito (De Angelis et al. 2019; Wallace et al. 2020). But transitions exist at times between 1147 vulcanian and strombolian explosions, as at Arenal (Mora et al. 2022) and Fuego (Diaz-Moreno et 1148 al. 2020; Brill et al. 2022). This requires continuous measurements for a better modelling of this 1149 1150 transition.

1151 Various models exist to explain the sound produced by strombolian explosions, each being 1152 driven at the vent by the breaking of a Taylor bubble (Fig. 3c). These models are based on the: 1) 1153 oscillations in volume of a large bubble prior to its bursting (Vergniolle and Brandeis 1994; 1996); 1154 2) resonances of a rigid cavity opened by a small hole as for Helmholtz resonators (Vergniolle and Caplan-Auerbach 2004); 3) formation of a hole piercing the magma layer overlying a small 1155 1156 bubble, either with an extremely thin and inviscid liquid film, as for soap bubbles in water (Vidal et al. 2010; Sanchez et al. 2014) or in a viscous fluid (Koyabashi et al. 2010). The observations 1157 1158 that the source of the sound is mostly monopolar had led Johnson et al. (2008) and Gerst et al. (2013), to estimate the gas volume and the gas flux under this assumption for some strombolian 1159 explosions at Erebus, with further validation from comparison with video records, at least for the 1160 beginning of the explosion (Witsil and Johnson 2018). The robust assumption of a monopole 1161 1162 source is valid for any type of explosion (strombolian and vulcanian) providing that the sound propagation is linear, hence not associated to shock waves. The effect of the topography had also 1163 been included on the inversion of acoustic waveforms recorded with an array, giving estimates of 1164 gas flux and gas volume at Yasur (Iezzi et al. 2019, Fee et al. 2021). The use of a single acoustic 1165 sensor, set at a proximal distance from the vent and with an unobstructed line-of sight, provides 1166 1167 reasonable estimates of gas fluxes and gas volumes when using the geometrical spreading of a monopole source and ignoring the topography (Johnson and Miller 2014). Furthermore, the use of 1168 1169 a monopole approximation to recover the source parameters, such as maximum gas flux and gas 1170 volume, is more useful for monitoring purposes than just using the maximum radiated acoustic pressure, sometimes seen to be poorly correlated with the height of the eruptive column (Johnson 1171 and Miller 2014). Topography effects can sometimes affect the infrasonic waveforms, such as at 1172 Fuego (De Angelis 2019), but the perturbation can be much reduced by the use of an infrasonic 1173 array and inversions of waveforms (Iezzi et al. 2019; Diaz-Moreno et al. 2020). However, the use 1174 1175 of a single sensor, located at a proximal distance and with unobstructed line-of-sight to the vent,

had been proven to give a reasonable estimate of the source parameters also for vulcanian 1176 1177 explosions when using a simple monopole source (Johnson and Miller 2014). Inversions of infrasonic waveforms done for small and moderate gas-and-ash explosions at Santiaguito provides 1178 the time history of mass eruption rates in good agreement with thermal or visible imagery, 1179 1180 validating the use of acoustic measurements for estimates of gas volumes and gas fluxes (Johnson et al. 2010; De Angelis et al. 2016). The radiation pattern of the strombolian explosions at Yasur 1181 was also measured by using infrasound sensors aboard a tethered aerostat above the crater area, 1182 showing that both path effects from crater walls and source directionality could play a role in the 1183 infrasonic waveforms (Jolly et al. 2017). 1184

Non-linear effects in the sound propagation, such as existing during blast waves and shock 1185 1186 waves, have been reported for the most vigourous strombolian explosions, with a velocity just slightly above the sound wave at Yasur (Marchetti et al. 2013; Taddeucci et al. 2014; Maher et al. 1187 1188 2022) or well-marked during vulcanian explosions and during the 2018 paroxysm at Fuego (Morrissey et al. 2008; Diaz-Moreno et al. 2020). The modelling using strong shock waves could 1189 1190 be applicable to weak shock waves produced by volcanoes (Medici, et al. 2014). Blast waves can often be distinguished from linear sources of sound based on their specific shape and amplitudes 1191 1192 exceeding several hundreds of Pa within a few hundred metres from the source and also by their characteristic N-shape acoustic waveforms at a far distance (De Angelis et al. 2019). However, 1193 1194 specific acoustic waveforms, asymmetrical with the larger compressive peak in comparison with its following rarefaction peak, are sometimes associated to shock waves and supersonic fluid flow, 1195 1196 can be easily obtained for subsonic flows (Brogi et al. 2018), for which the monopole 1197 approximation holds. Futhermore, a recent study, also discussing supersonic jet flows, had shown that the sound associated to volcanic vortices and jet noise could be used to retrieve vent size and 1198 eruption parameters at open-vent volcanoes (Taddeucci et al. 2021). 1199

In addition, gas velocity and gas volume were also estimated by using the infrasonic power on 1200 1201 a subplinian basaltic eruptive column at Shishaldin (USA; Vergniolle and Caplan-Auerbach 2006) and during the 2006 eruption of Augustine (USA; Caplan-Auerbach et al. 2010). However care 1202 1203 should be used due to the multiphase nature (solid particles and liquid droplets carried by a hot 1204 gas) of the volcanic eruptive cloud when simplifying the acoustic source produced by a sustained eruptive column to a simple dipole or quadrupole (Matoza et al. 2009; 2013). The temporal 1205 evolution of the infrasonic power was also used to locate the depth of the reservoir at Shishaldin 1206 (Vergniolle and Caplan-Auerbach 2006), at 3.9 km in agreement with melt inclusions (> 3km) 1207 (Rasmussen et al., 2018), and the rise in the lava lake prior to the paroxysm at Villarica (Johnson 1208 1209 et al. 2018).

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The sole use of infrasonic records to estimate the gas volumes at strombolian volcanoes often 1210 leads to an underestimate on the total gas volume, as infrasonic records can only record the 1211 unsteady gas flow (Johnson 2003) or that of the overpressurised gas expanding at the surface 1212 (Vergniolle and Brandeis 1994, 1996). The difference in total gas volumes deduced from 1213 1214 infrasonic records with that obtained from records of videos in the visible or infrared or ultraviolet range, open-path Fourier transform infrared spectroscopy and radar is noticeable (see 1215 discussion in Vergniolle and Gaudemer 2015). The cause of this discrepancy is mainly a 1216 consequence of the significant amount of passive degassing associated to each explosion, keeping 1217 in mind that any technique, which cannot measure directly at the vent but above the vent, is also 1218 measuring air entrained within the volcanic cloud, such often the case with video, open-path 1219 Fourier transform infrared spectroscopy and radar techniques. The existence of passive degassing 1220 associated to explosions at Erebus is very clear from the comparison between videos and 1221 1222 infrasonic records, in full agreement in the early part of explosion (gas decompression at the vent) but not later (Witsil and Johnson 2018). Although passive degassing could sometimes be 1223 1224 associated to a bubbly wake attached to the bursting of the large bubble driving the explosion, the large viscosity of Erebus magma, 10<sup>6.3</sup> Pa.s (Le Losq et al. 2015), prevents such a mechanism to 1225 1226 be called for. It is then likely that the passive degassing expelled during an explosion at Erebus is induced by the bursting of numerous small bubbles present in the shallowest part of the magma 1227 column and made to burst by the rarefaction wave induced by the explosion and propagating down 1228 in the conduit. The active degassing, even if not major in term of gas volume at the vent in 1229 1230 comparison with passive degassing, needs to be estimated because is a key to unravel the eruptive pattern, and in particular estimate the gas flux at depth. The comparison between videos and 1231 infrasonic records also shows the strong potential of combining various techniques at open-vent 1232 volcanoes to assess the extent of both passive and active degassing, hence to obtain a refined 1233 1234 estimate of the gas flux.

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### 1236 Sulphur evolution path

### 1237 The fate of the suphur

New inventory and detection of new and unquantified sources of  $SO_2$  passive emissions in volcanic arcs (Aleutian, Papouasie New Guinea, Indonesia) (Carn et al. 2017), and a better evaluation of the budgets of global volcanic  $SO_2$  and  $CO_2$  emissions (Fisher et al. 2019) have benefited from the advances in space-borne instruments, data treatments, and the combination of satellite-derived measurements with ground-based techniques and networks. It is recognised that the Nyiragongo –

Nyamuragira system significantly contributes to the CO<sub>2</sub> volcanic flux of continental rift (Fisher et 1243 al. 2019), and to the global emissions of SO<sub>2</sub> (Carn et al. 2017), after Ambrym (Allard et al. 2016a, 1244 b). Since 2012, Nyamuragira strongly contributes to the global budgets of bromine (18–35% Br<sub>total</sub>) 1245 and iodine (8-18% I<sub>total</sub>) (Bobrowski et al. 2017). It is characterised by an "excess" of sulphur 1246 degassing during effusive activity (Bluth and Carn 2008). A comparison between the SO<sub>2</sub> emissions 1247 inferred from the melt inclusion data for the 1986 and 2006 eruptions with satellite based records is 1248 1249 compatible with pre eruption gas loss and gas accumulation together with degassing of unerupted shallow magma (Head et al. 2011). However, such an excess degassing is questioned and could not 1250 1251 be the rule at Nyamuragira (Coppola et al. 2013). Hence the sulphur-degassing budget remained difficult to be constrained because of its multi-valence states and dependence on such parameters as 1252 fO<sub>2</sub>, fS<sub>2</sub>, pressure and magma composition (e.g., Wallace and Edmonds 2011; Métrich 2021). 1253

Erebus is another case where the sulphur evolution path can be tricky. The phonolitic magma 1254 that fills the magma at the top of the column is derived from the extensive differentiation of a 1255 parental basanite, after the removal of 75 % of crystals (Kyle et al. 1992). The parental melt may 1256 contain up 0.25 wt% S versus 0.03-0.06 wt% S in the phonolite as measured in melt inclusions of 1257 1258 olivine and anorthoclase, respectively (Moussallam et al. 2014). A part of this sulphur is locked in the sulphide globules described in melt inclusions of their host crystals (Moussallam et al. 2016) as 1259 well as in the lava mineralogy (Kelly et al. 2008). A very rough estimation of the magma volume 1260 that is sustaining the low SO<sub>2</sub> flux at Erebus during passive degassing could be tentatively derived 1261 1262 from the dissolved S content (610 ppm the highest S content in anorthoclase) and the SO<sub>2</sub> flux (on average 52t/d; Oppenheimer et al. 2004; Sawyer et al. 2008; de Moor et al. 2013). The volume of 1263 1264 magma that sustained the passive SO<sub>2</sub> emissions at Erebus would be 0.25 m<sup>3</sup>/s (7.6  $10^{-3}$  km<sup>3</sup>/yr, 1265 Tables 1 and 2).

1266 New strategies now involve Unmanned Aerial Aircraft Systems (drone), in complement to satellite based-SO<sub>2</sub> flux measurements (Tropospheric Ozone Monitoring Instrument (TROPOMI); 1267 Ozone Mapping and Profiler Suite (OMPS)), to provide large scale and repeated measurements of 1268 1269 volcanic gas emissions, and they have been proved to be efficient for almost inaccessible volcanoes 1270 (Lui et al. 2020). Manam in Papua New Guinea whose vents have been active for several decades is an example of a newly identified contributor to the global volcanic fluxes of CO<sub>2</sub> and SO<sub>2</sub> (Lui et al. 1271 1272 2020; Edmonds et al. 2021). Observed correlations between the molar CO<sub>2</sub>/SO<sub>2</sub> ratios of surface gas emissions and incompatible trace element ratios in bulk rocks (i.e., Ba/La; Sr/Nd) of bulk rocks 1273 (Aiuppa et al. 2019) represent a major step forward towards a better quantification of the volcanic 1274 gas fluxes. Such an approach has been tested at Pacaya, regarded as the strongest degassing source 1275 in Central America (Battaglia et al. 2018), and used to predict the CO<sub>2</sub>/SO<sub>2</sub> emissions of so called 1276

<sup>1277</sup> "unmeasured" volcanoes, as Fuego and Santiaguito in Guatemala (Aiuppa et al. 2019). In Japan six <sup>1278</sup> persistently active volcanoes (Miyakejima, Sakurajima, Satsuma-Iwojima, Asama, Aso, <sup>1279</sup> Suwanosejima) dominate the annual  $SO_2$  flux (Shinohara 2013) and produce  $CO_2$ -poor gas <sup>1280</sup> emissions that mirror to the low amount of carbon in subducted sediments (Aiuppa et al. 2019).

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### 1282 The example of Popocatépetl volcano in Mexico

Popocatépetl volcano provides a good example of passive degassing (Delgado-Granados et al. 1283 2001) as a large part of the magma that supports the surface  $SO_2$  emissions is not erupting and 1284 continued to degas at depth while crystallising (Witter et al. 2005; Roberge et al. 2009). 1285 Popocatépetl is observed to be in a state of permanent open-vent activity since 1994 with cycles of 1286 building and destruction (associated to vulcanian explosions) of dacitic lava domes (Gómez-1287 Vazquez et al. 2016; Taquet et al. 2019) and strong gas emissions (5000 tons/d SO<sub>2</sub> as measured on 1288 28/2/2019) from the central part of the dome (Campion et al. 2018). The activity renewed after ~70 1289 years of repose time, the last eruption occurred in 1920 and lasted 7 years (Delgado-Granados 1290 2001). Campion et al (2018) report a 4-year average SO2 flux of 45 kg/s (3888 t/d against an 1291 average value of 1658±893 tons per day (1 $\sigma$ ) given by Carn (2017) from SO<sub>2</sub> camera measurements 1292 1293 between 2013 and 2016, during which degassing dominates (>95% of the time). They place 1294 Popocatépetl as one of the strongest permanent emitters of volcanic SO<sub>2</sub> over this period together with Ambrym (100 kg/s; Allard et al. 2016a, b), Kilauea (10-60 kg/s; Nadeau et al. 2015), and 1295 Nyamuragira (20–60 kg/s; Coppola et al. 2016). 1296

From March 1996 to 1<sup>st</sup> January 1998, the total SO<sub>2</sub> emissions achieved 9 Mt (9-13 kt/d) that is 4 1297 times higher than the SO<sub>2</sub> flux during the inter-eruption period (2-3 kt/d; Delgado-Granado et al. 1298 2001). Roberge et al (2009) conclude to a recharge of a minimum of 0.8 km<sup>3</sup> (equivalent of a 1299 magma supply rate of 0.285 kg per year over this period) of mafic magma (for a magma containing 1300 2000 ppm S) for sustaining this high discharge amount of SO<sub>2</sub>, rather than convection of magma in 1301 the uppermost part of the system. It would imply an initial CO<sub>2</sub> content of 1 wt% (Roberge et al. 1302 2009). They propose a model of decompression and degassing of magma (with up to 5 wt% H<sub>2</sub>O) at 1303 1304 150-350 MPa, with only a small portion of this magma reaching the surface, while being mixed with the silicic shallow magma. The effect of H<sub>2</sub>O loss promotes crystallisation and increases the 1305 mixture viscosity that may slow down the convection. They thus suggest a mechanism based on the 1306 1307 intrusion and degassing of a mafic magma batch at the origin of these emissions rather than one 1308 based of convection in the shallowest part of the subvolcanic conduits. On the basis of the presence 1309 of Mg-rich olivine together with textural and compositional evidence for mixing, Schaaf et al.

(2005) propose dominant recharge of olivine-bearing basaltic andesite magma with relatively highvolatile content.

Long-time series of the gas composition at Popocatépetl between 2012 and 2016, after the 2011 1312 magma reinjection, indicate variation of the SO<sub>2</sub> flux between 1990 to 3680 tons per day (Table 2b) 1313 1314 with the highest value corresponding active activity with magma extrusion and the lowest quiet period, and an average for the whole period of 2790 tons per day (Taquet et al. 2019). Following the 1315 approach developed by Roberge et al. (2009), we have tentatively calculated the magma supply rate 1316 for the 2012-2016 period using Taquet et al. (2019)  $SO_2$  flux of 2790 tons per day. The magma 1317 supply rate could achieve 0.1 km<sup>3</sup> per year while Campion et al. (2018) obtained a value of 0.16 1318 km<sup>3</sup> (Table 1b) from a compilation of the SO<sub>2</sub> flux since 1995, including periods of high and low 1319 1320 activity.

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# 1322 Persistent degassing and sulphur behaviour at open-vent volcanoes in Central American 1323 Volcanic Arc

Pacaya, Fuego and Santiaguito are three open-vent volcanoes in Guatemala and are located at the
triple junction of the North American, Cocos, and Caribbean tectonic plates that have sustained
persistent degassing activity during several decades (Battaglia et al. 2018).

The case of these three volcanoes provide the opportunity of discussing the sulphur behaviour in 1327 arc related basalts able to ascent before extensive crystal fractionation at Pacaya, to basalt and 1328 basaltic andesites at Fuego and dacitic magmas that are the final products of differentiation of such 1329 basaltic melts. Their magmas share comparable Ba/La ratios (on average 44-52) and CO<sub>2</sub>/S<sub>total</sub> ratio 1330 of 1-1.5 at Pacaya; 1.5±0.75 at Fuego, and 1.4±0.75 at Santiaguito (Aiuppa et al. 2017; Battaglia et 1331 al. 2019). According to these authors they thus share similar nature of slab component with 1332 1333 dominant aqueous fluids derived from the slab dehydration, and small amount of carbonate in subducted sediments. 1334

Pacaya volcano, resumed Strombolian activity in 1961, after 200 years of rest and produced lava 1335 flow and pyroclastic explosive eruptions over nearly 50 years of activity, with flank eruptions that 1336 1337 possibly emptying the upper most parts within the cone (Rose et al. 2013b). Violent explosive eruptions and lava fountains are reported in January and February 2000, and in 2010 (Rose et al. 1338 1339 2013b). Pacaya produced homogeneous crystal-rich high-Al basalt having between nearly 30 to 50 vol% phenocrysts with an average eruption rate for Pacaya of  $0.17 \pm 0.04$  m<sup>3</sup>/s (78 kg/s) from 1961 1340 1341 to 2001 (Rose et al. 2013b). The rapid ascent and crustal transfer of magma could have been 1342 facilitated by local extensional regime (Cameron et al. 2002; et al. 2018).

Time series of in-situ (Multi-GAS) and remote (ultra-violet camera) plume measurements carried 1343 out in January 2016, during a period of relatively mild-strombolian activity at Pacaya, provide a 1344 SO<sub>2</sub> flux of 885±550 t/d (Battaglia et al. 2019). In contrast, it has fluctuated since the onset of the 1345 ongoing eruptive period in 1965, from 260-290 tons/day in the period of 1972-1992 (Andres et al. 1346 1347 1993). It peaked at 1350 t/d within the 1999–2002 time period of higher level of explosive activity and reached 1570 tons per day while a lava lake is observed at the crater from August 2000 until 1348 May 2002 (Rodriguez et al. 2004). These observations led Rodríguez et al. (2004) to propose that 1349 Pacaya has a substantial convecting and circulating magma body near the surface that is degassing 1350 1351 high-Al basalt, only partially extruded.

A magma production rate of  $0.17 \pm 0.04 \text{ m}^3$ /s (equivalent to  $5 \times 10^{-3} \text{ km}^3$  per year) would imply an initial concentration of sulphur >> 1 wt% for a flux of varying from 260 to 1350 t/d that is unrealistic. The initial concentration of sulphur in parental basalt is not accurately known since only a few values are available ( $\leq 0.17 \text{ wt}\%$  S; Walker et al. 2003). Considering the highest value of 0.17 wt% S measured by Walker et al. (2003), that could be underestimated, would imply a volume ratio between the degassed and extruded magma of the order of 7, for an intermediate SO<sub>2</sub> flux of 885 t/d (Table 2b), with thus excess of sulphur degassing as concluded by Rose et al. (2013).

1359

Fuego is regarded as the most dangerous volcano in Central America known, since 1524, for its 1360 Subplinian eruptions (volcanic explosivity index 4), as occurred in October 1974 (Rose et al 1978; 1361 1362 Chesner and Rose 1984) after periods of repose times that lasted years or decades as reviewed by Naismith et al. (2019). The October 1974 Subplinian eruption produced hybrid magma from 1363 1364 basaltic (crystal-poor) to basaltic andesite (crystal-rich) magmas (Rose et al. 1978). It was sustained 1365 by an influx of ~0.1 km<sup>3</sup> H<sub>2</sub>O-rich high-alumina basaltic magma (Rose et al. 1978; 2008). Lowlevel Strombolian activity persisted until 1979 (Martin and Rose 1981) and from 1980 to 1999 1366 Fuego had irregularly spaced subplinian (volcanic explosivity index 1-2) events with periods of 1367 repose (Lyons et al. 2010). The magma erupted in 1999 and 2003, after two decades of quiescence 1368 1369 (Lyons et al. 2010), is partly residual from 1974 (Berlo et al. 2012). During 2 years of continuous 1370 observations (2005–2007) Fuego was persistently active, with a repeating cycle of activity: 1) passive lava effusion and minor strombolian explosion, followed by 2) paroxysmal eruptions, and 1371 finally 3) degassing explosions without lava effusion (Lyons et al. (2010). One of Fuego's largest 1372 eruptions in the period 1999-present occurred on 13th September 2012 whereas the most 1373 destructive eruption occurred on 3 June 2018 (Naismith et al. 2019; Liu et al. 2020). This explosive 1374 eruption was preceded by 5-month hiatus in activity that was proposed to be sufficient a time for 1375

1376 crystallisation of a shallow low permeability plug and by the flowing of hot and high-speed lava1377 flows (Liu et al. 2020).

It is found to be part of a 45-year cycle of high-energy eruption that produced a magma volume of  $0.04 \pm 0.01$  km<sup>3</sup> (dense rock equivallent) and 130 ktons of SO<sub>2</sub> (IASI satellite data) that could have involved the influx of basaltic magma (Pardini et al. 2019). This SO<sub>2</sub> flux would imply a concentration of sulphur of nearly 500 ppm a value by far lower than the S dissolved amount in the 1382 1974 olivine-hosted melt inclusions up to 2250 ppm (Lloyd et al. 2014) and ~2500 ppm (Berlo et al. 2012).

1384 The hybrid magma body erupted in between 14 and 23 October 1974 was proposed to reflect the 1385 contribution of several magma rather than a single magma composition, the basaltic magma (0.60 1386 wt% K<sub>2</sub>O) standing at >10 km depth and a second basaltic and site magma (0.90 wt% K<sub>2</sub>O) at 3–8 1387 km depth (Roggensack 2001). Accordingly, the observed variability is a reflection of magma 1388 hybridization immediately before the eruption and that the pre-mixing water contents of basaltic magma were likely 4-5 wt%. Lloyd et al. (2014) measured in the 1974 melt inclusions 4.5 wt% 1389 1390 H<sub>2</sub>O. Most data of 1974 olivine-hosted melt inclusions of Fuego can be reproduced by decompression-crystallisation trend between 1010°C and 1100°C in a vertically extended reservoir 1391 1392 (Cashman and Edmonds 2019). According to these authors, the high Al<sub>2</sub>O<sub>3</sub> content recorded in MIs (4.5 wt% H<sub>2</sub>O; Lloyd et al. 2014) prior to late-stage plagioclase crystallisation supports a high 1393 proportion (50%) of crystal fractionation (olivine and clinopyroxene) from a parental melt possibly 1394 having up to 7.6-9.6 wt%. Ion-probe analysis of S, H<sub>2</sub>O and CO<sub>2</sub> diffusion profile in open glassy 1395 gulfs (embayments) in olivine crystals still connected to the carrier magma indicate a 1396 decompression rate 0.24 to 0.60 MPa/s (~8-21 m/s) of the magma erupted on 17 October 1974 1397 1398 (Lloyd et al. 2014).

Liu et al. (2020) proposed a top-down decompression of magma in the conduit as the triggering mechanism of the eruption of 3 June 2018, shallow crystallised magma acting as a plug. Liu et al. (2020) also showed that the frequency of paroxysms at Fuego is broadly proportional to the gas supply rate.

Low sulphur concentration (~500 ppm) corresponds to the sulphide saturated melt, which has lost  $H_2O$  (Wade et al. 2006). Pyrrhotite is actually observed in association with titanomagnetite (Rose et al 1978, Berlo et al. 2012). Comparatively basaltic-andesite magma at Arenal contains similar S concentrations (2500 ppm), and becomes sulphide saturated as it is degassing and crystallising (Wade et al. 2006). Hence, low SO<sub>2</sub> flux reflects the degassing of crustal magma that is losing water and crystallising, a part of sulphur being locked in segregated sulphide globules.

Santiaguito complex consists of four domes, along an East-West trending fracture. The activity 1410 of dome-building began in 1922 within the crater formed during the cataclysmic 1902 eruption of 1411 Santa María (Rose 1972, 1987, 2013; Scott et al. 2012, 2013) and continues today. The longevity 1412 and continuous activity of Santiaguito are unusual among historic dome eruptions (Scott et al. 1413 2013). Magma extrusion rates have shown a cyclicity with 3-6-year-long periods of higher 1414 extrusion (0.6–2.1 m<sup>3</sup> s<sup>-1</sup>), which are separated by 4- to 11-year-long periods of lower (~0.2 m<sup>3</sup> s<sup>-1</sup>) 1415 extrusion rate (Rose 1987; Harris et al. 2003, 2004). The activity at Santiaguito since 1975 is 1416 characterised by effusion of blocky lava flows and small-to-moderate gas-rich explosions (Harris et 1417 al. 2003) reaching up to 1.5 km above the vent with explosions occurring at different intervals, from 1418 a few minutes up to ~6 days (Carter et al. 2020). Santiaguito dacitic rocks are typically porphyritic, 1419 with a phase assemblage dominated by plagioclase (~20 to ~30 modal%) (Scott et al. 2013). Dacitic 1420 magma is found to be derived from the andesite via ~20 to 25% fractional crystallisation (Scott et 1421 1422 al. 2013). Singer et al. (2011) propose that the dacite erupted in 1902 and during the early domeforming cycles of Santiaguito lavas resulted from ~25 ka of fractional crystallisation in a closed 1423 1424 system. The erupted magma has become progressively less evolved over time, with a decrease in whole rock SiO<sub>2</sub> from ~66 to ~62 wt% from 1922 to 2002 (Scott et al. 2013) with magma being 1425 1426 extracted progressively from a chemically-stratified lower crustal (~12 to ~25 km) storage zone thought to have formed during the  $\sim 25$  ka of quiescence that preceded the 1902 eruption. 1427 Santiaguito is by some aspect comparable to Colima for its cyclic emissions of SO<sub>2</sub>, with low 1428 background and high emissions of gas (1/3 of the emissions) leading to a range of SO<sub>2</sub> flux and 1429 1430 explosion phenomenology (Lamb et al. 2019).

Battaglia et al. (2018) report time-averaged SO<sub>2</sub> fluxes of  $127\pm58$  tons/day at Santiaguito (Table 2b), which reflects a very shallow magma degassing. Santiaguito parental melt have most likely similar parental basaltic magma, rich in water and sulphur as shown for Pacaya and Fuego. Whereas at Pacaya the SO<sub>2</sub> flux is sustained by high-alumina basalt degassing, dacitic magma evolution requires long-term fractional crystallisation (Singer et al. 2011), it thus implies titanomagnetite crystallisation, magma degassing and thus change in the redox conditions promoting melt sulphide saturation.

1438 In summary, these examples highlight the great need to combine time-series data, geochemical 1439 and petrological approaches as the necessity of having good geochemical modeling of depth-1440 dependent magma redox.

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## 1442 **Concluding points**

Combining time series infrasonic data and measurements of gas fluxes on well monitored open-vent 1443 volcanoes on the field is essential for the interpretation and modelling of the degassing pattern at the 1444 surface and in the conduit and understanding the eruptive activity. However, it cannot be 1445 1446 dissociated from the good knowledge of the texture of the erupted products, the depth-evolution of 1447 the dissolved amount of water in the melt and thus its crystallinity owing to the key role of magma viscosity in modelling the bubbles and magma motions. Key advances could be made by any 1448 techniques constraining the in-situ bubble diameters and the gas volume fraction of the small 1449 bubbles, as one way to view the different eruptive regimes at basaltic and explosive volcanoes is a 1450 consequence of the ability of their largest bubbles to rise with a differential velocity in respect to the 1451 liquid or be stagnant, respectively. An accurate modelling of the magma viscosity itself combining 1452 physical and thermodynamic parameters is thus a challenge. 1453

1454 A better modelling of the processes at the vent could be obtained by the joined and systematic use of thermal and infrasonic records. The inputs, which could be offered by imagery of the surface 1455 1456 activity, possibly at high-speed, such as in the visible, infra-red or UV range, are enormous for our 1457 understanding, as these open-vent volcanoes remain an open window onto a deeper plumbing 1458 system, in which the future activity at the surface is being prepared. Similarly, be able to measure 1459 systematically and continuously the level of the magma column in the conduit or in the lava lake could be used as a key to detect pressure changes at depth and/or be used as a precursory signal of 1460 an enhanced activity. The case of Masaya shed lights on the interest to record simultaneous thermal 1461 1462 and SO<sub>2</sub> fluxes to quantify the magma supply rate.

Another interesting avenue to explore is that provided by the analysis of long-time series as openvent volcanoes show sudden or progressive changes in behaviour. Several observatories across the world obtain these data but so far continuous long-time series are mainly restricted to geophysical methods.

Additionally, a key parameter to know for a better quantification of the  $SO_2$  flux in term of volume of degassed magma is the magma redox evolution during its crustal evolution and crystallisation. The fate of the degassed magma is also important, as it may remain at depth, be forcefully intruded into the edifice or be partly recycled by a convective loop in the conduit, i.e. the bi-directional flow at work at the less viscous open-vent volcanoes.

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- 133 134

#### 2244 **Figure captions**

Fig. 1 SiO<sub>2</sub> distribution in magmas of Erta 'Ale (Field et al. 2012); Masaya (Zureck et al. 2019);
Fuego (Berlo et al. 2012); Nyiragongo (Minissale et al. 2019; Demant and Lubala, 1994);
Nyamuragira (Head et al. 2011); Stromboli (Landi et al. 2006, 2009; Bertagnini et al. 2008;
Pompilio et al. 2011; Métrich et al. 2010, 2021); Ambrym (Allard et al. 2016a, b) ; Pacaya
(Cameron et al. 2002) ; Yasur (Métrich et al. 2011); Erebus (Kelly et al. 2008); Arenal (Bolge et al. 2006); Santiaguito (Scott et al. 2013); Popocatépetl (Schaaf et al. 2005).

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Fig. 2 Plumbing systems and associated mechanisms at: a) Kilauea (slightly adapted from Ferguson et al. 2016); b) Nyiragongo (Pouclet and Bram, 2020); c) Stromboli during the small-scale paroxysms (slightly adapted from Métrich et al. 2021); d) from seismic tomographic inversion from combining artificial and natural seismic sources, showing the magma pathway (red arrows) and the high velocity zone associated to the crystallised magma body (slightly adapted from Patanè et al. 2017). SdF, Stcc and iso, HVB stand for Sciara del Fuoco, Strombolicchio, isosurface of constant seismic velocity propagation (iso6 is for 6 km/s) and high-velocity body.

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Fig. 3 Potential bubble geometries for rising bubbles (a) and stationary bubbles (b): left: small;
middle: relatively large and defined as spherical cap bubble; right: very large and called Taylor
bubble; b) flat bubble, i.e. thin and very long, resulting from a local coalescence within a foam
accumulated below the crust at the top of a lava lake and defined in Bouche et al. 2010.

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Fig. 4 Lava lake of Kilauea at: a) Puu O'o on March 1988, showing a broken bubble of a few meters
in diameter (photo by S. Vergniolle); b) summit on 29 July 2016 (courtesy of M. Patrick); c)
summit on 6 May 2018, as the lake drained away during the 2018 lower East Rift Zone eruption,
with a crater of diameter of 150 m (courtesy of M. Patrick). Note the difference in the lava lake
level between 2b and 2c, analysed by Patrick et al, 2015.

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Fig. 5 Lava lake of Nyiragongo with a diameter of 250-260 m along the east-west axis and 200 m along the north-south axis (courtesy of B. Smets): a) Panoramic view of the lava lake on 12 September 2011 displaying a hot central zone and a partly solidified annulus. b) Many incandescent cracks can be seen in central area on 12 September 2011; c) overflows of the lava lake with the former lava lake appearing as an annular incandescent ring, on 12 September 2011;d) bubble bursting at the surface of the lava lake, in 2016.

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Fig. 6 Top of exposed magma column of Erebus (courtesy of C. Oppenheimer): a) with a diameter of 60 m, 7 December 2003 at 11h15 UTC; b) low-level lava lake, 30-40 m in diameter, 3 December 2012 at 9h54 UTC; c) fresh bombs can be seen on the side of the crater, 60 m in diameter, 15 December 2004 at 23h57 UTC; d) weak degassing (blue cloud), 17 December 2010 at 10:44:05.

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Fig. 7 Lava lake of Erta 'Ale: a) 29-30 January 2003 (courtesy of J-M Bardintzeff), the diameter is nearly 90 m; b) 22 March, 2003 (photo taken by T. Staudacher and slightly adapted from Vergniolle and Bouche, 2016); the diameter is 30 m; c) December 2004, the lava lake is entirely encrusted (diameter > 100 m; courtesy of J-M Bardintzeff); d) thermal image on 26 March 2003 (taken by T. Staudacher) with a lava lake of diameter equal 30 m, and with insert of visible image of a spherical cap bubble in March 2001 (adapted from Bouche et al. 2010 photo 4d).

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Fig. 8 Lava lakes of Ambrym: a) Benbow crater (courtesy of J-M Bardintzeff), with diameters of 2 2291 km for the caldera, 800 m for the largest crater in front, 200-300 m for the two small craters; b) 2292 Marum crater, 80 m in diameter, 13 August 2013 at 14h01 (courtesy of T. Boyer); c) Marum, 2293 bubble diameter of 8-10 m, photo taken at 80 m distance, October 2007 (courtesy of P. Allard); 2294 2295 d) Marum, bubble bursting on 22 June 2015 at 12h34 (courtesy of T. Boyer); e) Marum, diameter of 90 m, depth of 350m, 14 November 2016 at 11h20, the lava lake had a partial thin 2296 crust cover, with incandescent cracks (courtesy T. Boyer); f) lava flow on 9 March 2015, the 2297 cone in the back is Niri Maben Mbwelesu (courtesy T. Boyer). 2298

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Fig. 9 Lava lake at Masaya, a) 16 January 2009, degassing from hidden thermally radiating lava
lake (courtesy of P. Delmelle); b) in 2016 dimensions of 40x30 m, depth: 250 m (courtesy of JM Bardintzeff); c) 18 February 2017 (courtesy of P. Delmelle) approximatively 15 m of
diameter.

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Fig. 10 Sketch for the variations of the lava lake level at Erta 'Ale, explained to result from the
variations in the thickness of a foam accumulated at the top of the magma reservoir (slightly
adapted from Vergniolle and Bouche, 2016): a) lava lake is at a high level when the foam is
growing in the reservoir due to a non-zero underlying gas flux, as in 9-13 Dec 2002; b) low level

lava lake, when the underlying gas flux in the reservoir had stopped leading to the foam thinning 2309 due to its spreading in the conduit, as for the period December 2002-May 2003; c) changes of the 2310 lava lake volume, as deduced from MODIS records (diamonds) and compared with a model of 2311 foam thickness, used independently with two different boundary conditions, a non-zero 2312 2313 underlying gas flux (rise of lava lake) and a zero underlying gas flux (descent of the lava lake). The best fit between the MODIS records and the numerical solving gives two constrains, on the 2314 underlying gas flux and the area of the magma reservoir (Vergniolle and Bouche, 2016). Note 2315 that  $\mu_l$ ,  $\rho_l$ ,  $Q_{in}$  and  $Q_{out}$  stand for magma viscosity, magma density, underlying gas flux below the 2316 foam and the gas flux at the surface. The foam with a gas volume fraction,  $\varepsilon$ ,  $\Box \Box \Box \Box \Box$  towards 2317 the conduit at a velocity  $\square \square \square \square \square \square \square \square$  and has a thickness h(r,t), which varies with the radial 2318 coordinate, r, within the radius of the magma reservoir, R<sub>c</sub>, and the time t. t<sub>rise</sub> and t<sub>fall</sub> are the 2319 times at which the level of the lava lake rises and  $\Delta V_{\text{NTI}}$  is the relative volume of the lava lake in 2320 respect to that of March 2003 (see details in Vergniolle and Bouche, 2016). 2321

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Fig. 11 Models of degassing in the conduit, initially proposed for Erta 'Ale (Bouche et al. 2010) and 2323 2324 applied to: a) Erta 'Ale with an updated value of the magma viscosity; b-c) Ambrym at two different periods marked by different bubble diameters at the surface. The degassing at Erta Ale 2325 2326 is driven by the rise of large bubbles able to produce a bubbly wake, as deep as 1 km and up to 2327 the surface, while this type of degassing only exist from 500-1000m to the surface at Nyiragongo and at shallow depths at Villarica. The degassing at Ambrym is relatively similar to that of Erta 2328 Ale, when the bubbles are large, >8.5 m (see text for details). The drastic change in magma 2329 viscosity with the loss of dissolved water and the calculation of the gas volume fraction by 2330 assuming stagnant small bubbles and a fully degassing magma from gas exsolution had led us to 2331 use the viscosity of the undegassed magma below 500 m and that of the fully degassed magma 2332 above (see text and table 3 for details). Note that U<sub>b</sub> and U<sub>w</sub> are the vertical velocities of the 2333 driving large bubble and its detached bubbly wake (of length  $L_w$ ), respectively, while  $\alpha_l$  and  $\alpha_w$ 2334 are the gas volume fractions within the lava lake and the wake, respectively. 2335

2336

Fig. 12 Stromboli, a) a typical strombolian explosion in 1979, approximately of 150 m in height
(courtesy of J-M Bardintzeff); b) gas-rich explosion October 2005, height of ejecta between 3
and 5 m (photo taken by N. Métrich), 20 m in height (T Ricci); c) start of an ash-rich explosion
at vent N1, 29 September 2019 at 10h55 local time, approximately 30 m in height, (courtesy of
T. Ricci); d) southwestern vent in 1970, showing the top of the nearby magma column, an
exceptional photo due to danger, as well as on its information on Stromboli's eruptive pattern 40

years ago (courtesy of G. Sartoris), with an opening potentially of 1 m (order of magnitude); e)
lava pond due to bubble rise at vent N2, 16 November 2008 at 15hr46, length of major axis of
the red ellipse between 3 and 4 m, (courtesy of T. Ricci); f) a few seconds later at vent N2, 16
November 2008, 15hr46 (scale identical to e) (courtesy of T. Ricci).

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Fig. 13 a) an ash-rich explosion at Yasur, reaching a maximum height of 100m with ejecta falling
back in crater in a radius of 150m from the emission point, on 1 August 2008 at 17h18 (local
hour) (courtesy of A. Finizola); b-d) Arenal's edifice (courtesy of O. Aragon) with b) a typical
vulcanian explosion, leading to a small eruptive column of 400 m in height, in 1990, c) a typical
strombolian explosion with lava flows, in 1995 at 20h20, the red part with ballistic blocks is
approximately 500-600 m; d) close-up of a typical strombolian explosion with lava flows, in
1989 at 20h15, between 100 and 200 m, showing the two different vents existing in the C crater.

Fig. 14 a-b) Popocatépetl with a crater of 300 m in diameter (courtesy of R. Campion) with a) lava
dome at a high level, showing the passive degassing on an annular ring at the origin of the large
SO<sub>2</sub> flux, on 28 February 2019 and b) lava dome at a low level, on 18 February 2020, the rise of
the lava dome remaining puzzling for this dacitic magma composition, leading to vulcanian
explosions, hence initiated below a sealed vent; c) small ash-and-gas explosion at Santiaguito,
reaching a height of 500-1000 m, in 1978 (courtesy of J-M Bardintzeff).

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2363 Fig. 15 Bi-directional flow models for open-vent volcanoes, such as Stromboli, Yasur, Erebus and lava lakes, (see text for details), a) developed for explaining the excess degassing (slightly 2364 adapted from Stevenson and Blake, 1998); b) assuming that the velocity of the degassing 2365 magma, v<sub>s</sub>, is maximum and adding the degassing and ignore the crystallisation path at 2366 2367 Stromboli (slightly adapted from Burton et al. 2007b). Note that  $v_d$  is the velocity of the descending fluid of viscosity  $\eta$  and density  $\rho$ , while  $\phi$  is gas volume fraction in the ascending 2368 fluid and r<sub>s</sub> the radius of the ascending fluid; c) core-annular flow including lateral gradients in 2369 magma viscosity and density and solving for inclined and vertical conduit and combined with 2370 2371 petrologic data on magmatic volatile content at Villarica (slightly adapted from Palma et al. 2011). Note that  $\rho$  and  $\mu$  stand for magma density and viscosity for the ascending fluid (subscript 2372 c) and the descending fluid (subscript a). R is the conduit radius and r the radial coordinate; d) 3 2373 2374 regimes are possibly as a function of the viscosity ratio, although the core-annular flow is the most frequent (slightly adapted from Beckett et al. 2014). Note that the ascending fluid is marked 2375 2376 in grey; e) convective velocity and height of the convective cells, estimated from the re-

dissolution of the tiny bubbles within the downwards motion of the degassed magma (slightlyadapted from Carey et al. 2013).

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## 2382 Table captions:

Table 1 Main features of open-vent volcanoes (geodynamic context, activity period, magma composition, temperature, crystal volume fraction, magma supply rate) and associated references
for : a) persistent lava lakes (Nyiragongo, Erta Ale, Masaya, Villarica, Kilauea, Ambrym) ; b)
our selection of open-vent volcanoes without a lava lake (Stromboli, Yasur, Erebus, Popocatépetl, Arenal, Pacaya, Fuego, Santiaguito).

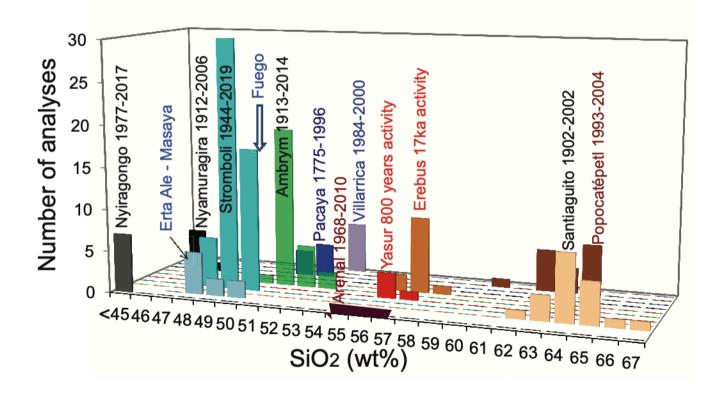
2388

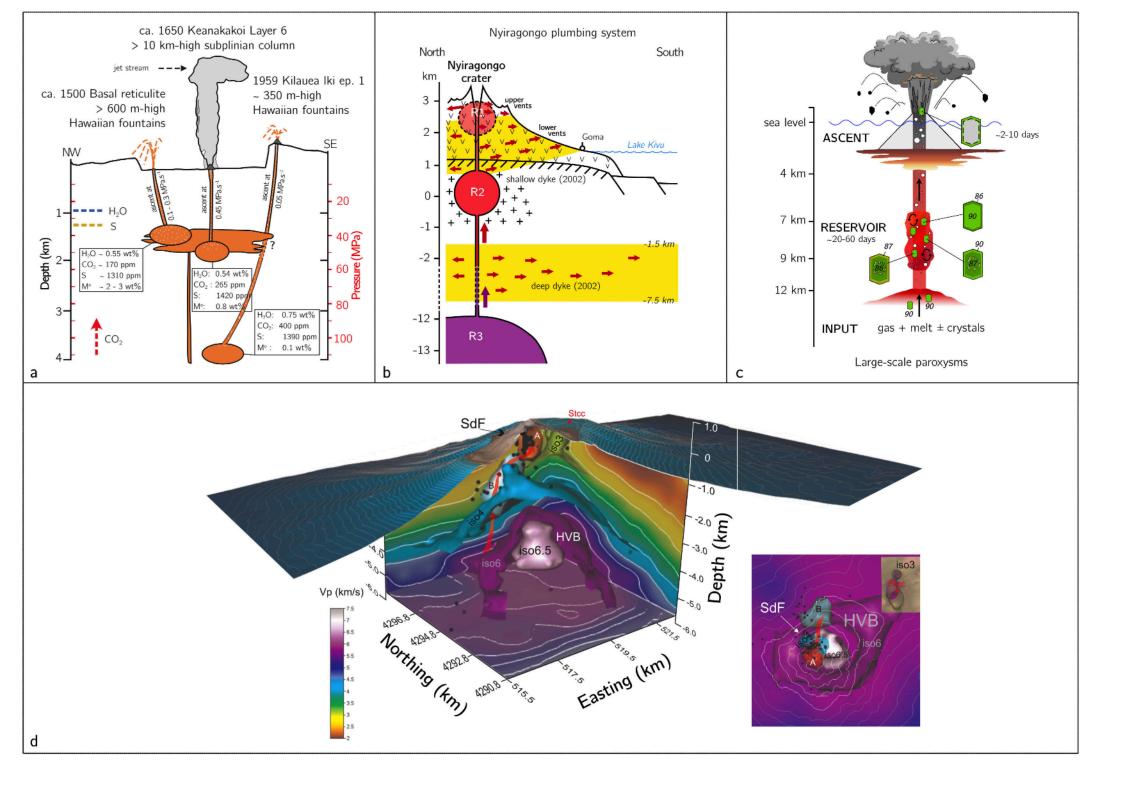
Table 2 Dissolved sulphur amounts in magmas (wt%) and SO2 fluxes for open-vent volcanic systems a) hosting a lava-lake (Nyiragongo, Erta Ale, Masaya, Villarica, Kilauea, Ambrym); b) without a lava lake (Stromboli, Yasur, Erebus, Popocatépetl, Arenal, Pacaya, Fuego, Santiaguito).

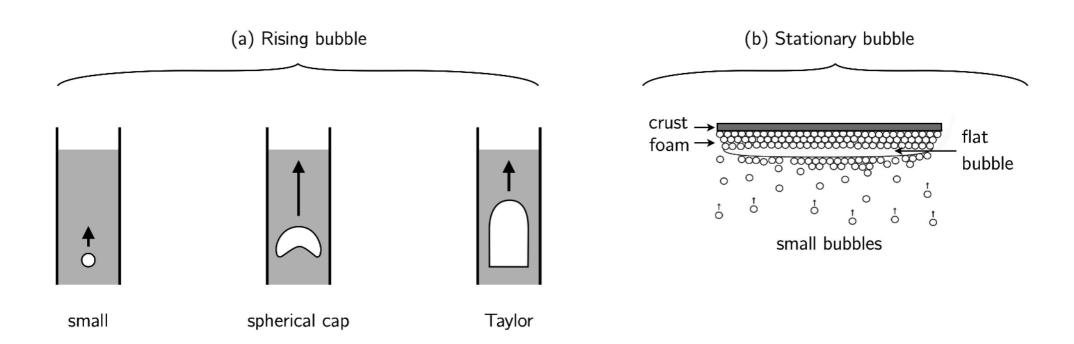
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Table 3 Parameters (silica content, temperature, dissolved water content, viscosity, bubble volume observed at the surface, Reynolds numbers) used for the calculation of the Reynolds number at lava lakes and at different depths (500 m, 100m, 0m), assuming undegassed magma below 500 m and fully degassed magma above (see text for details).

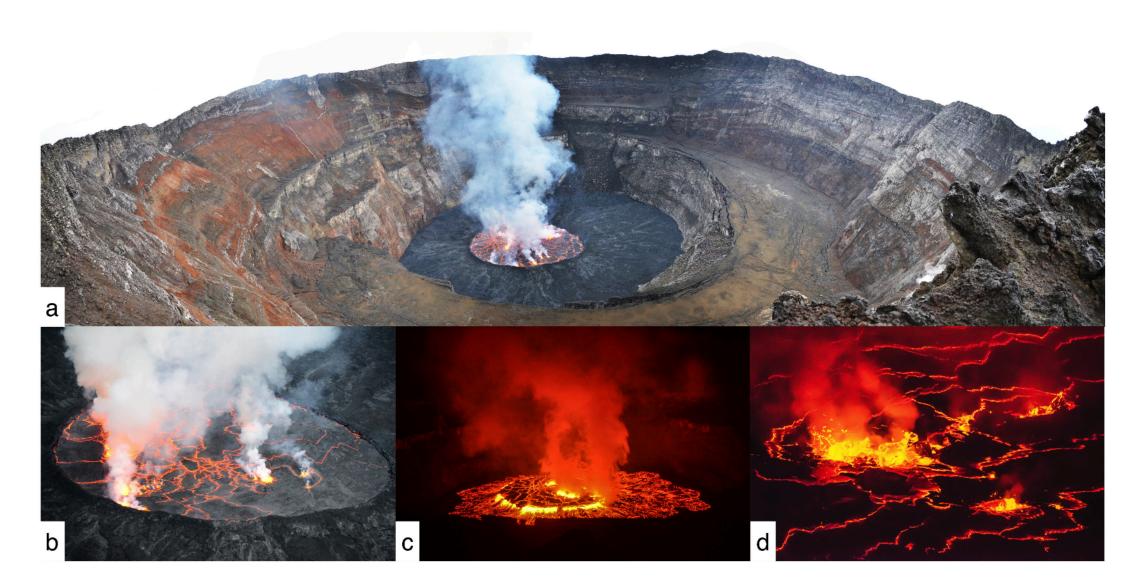
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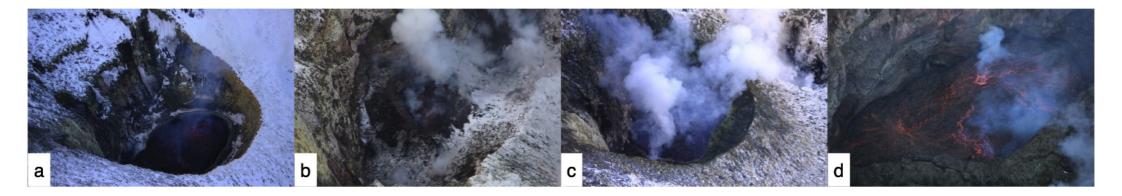


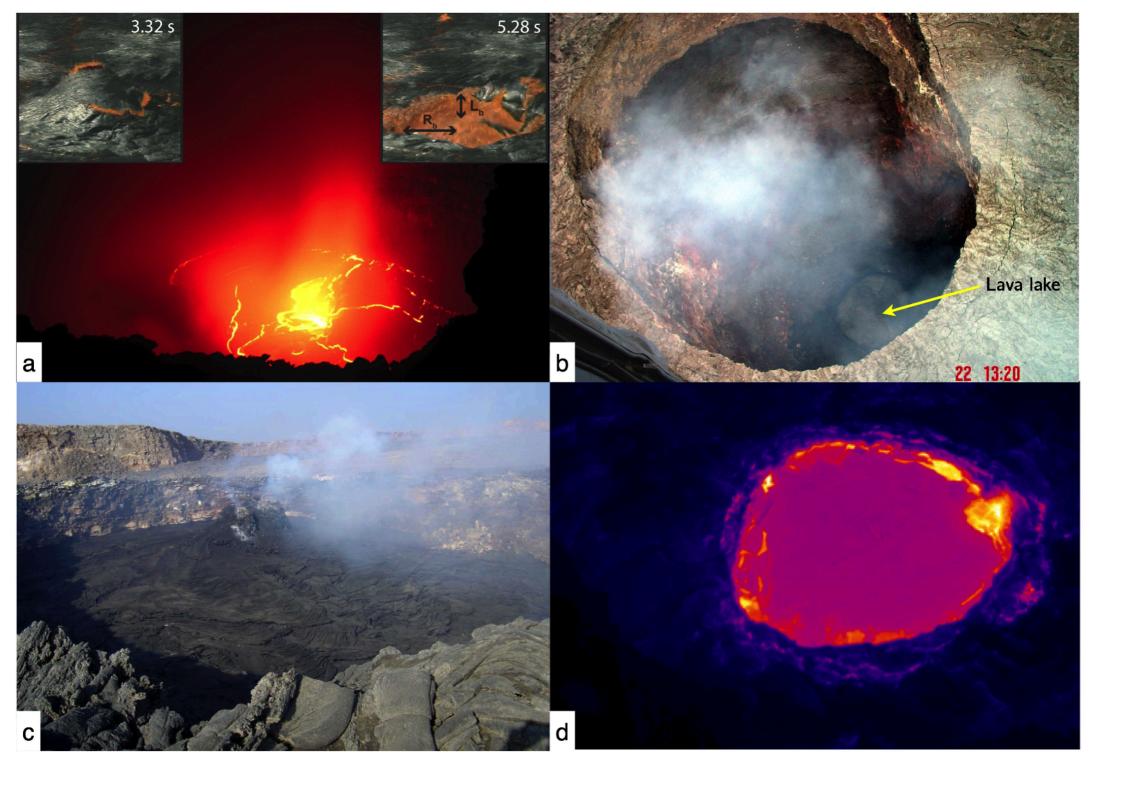






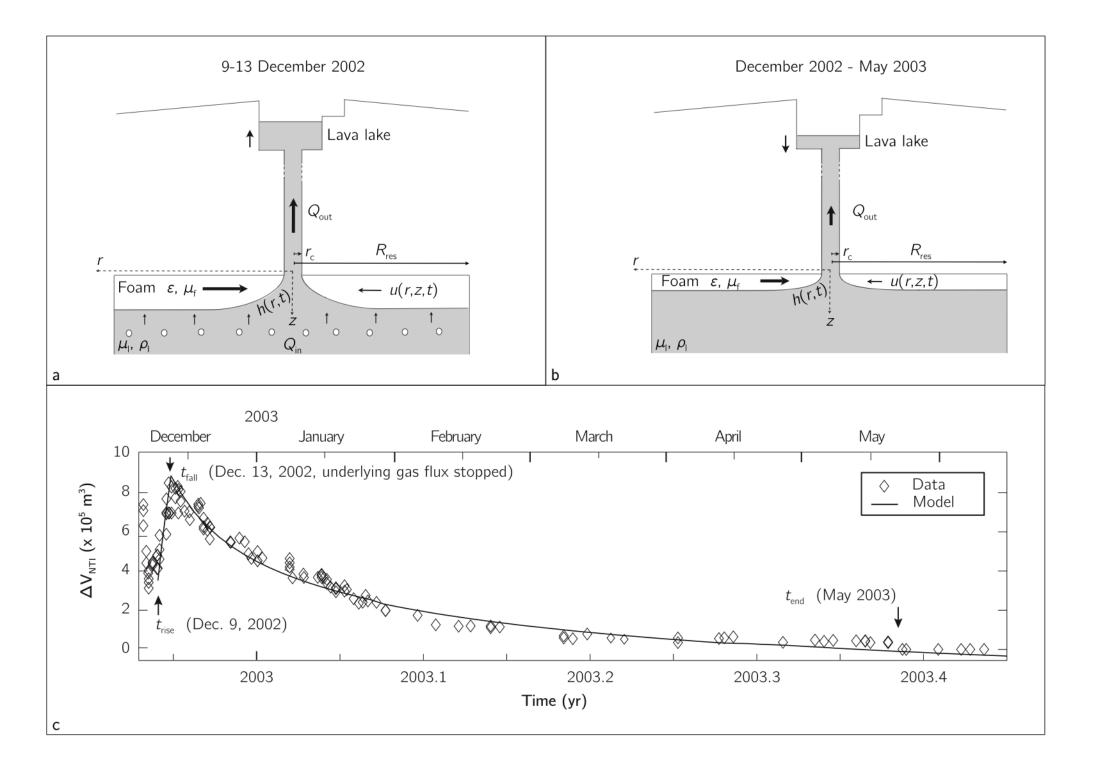


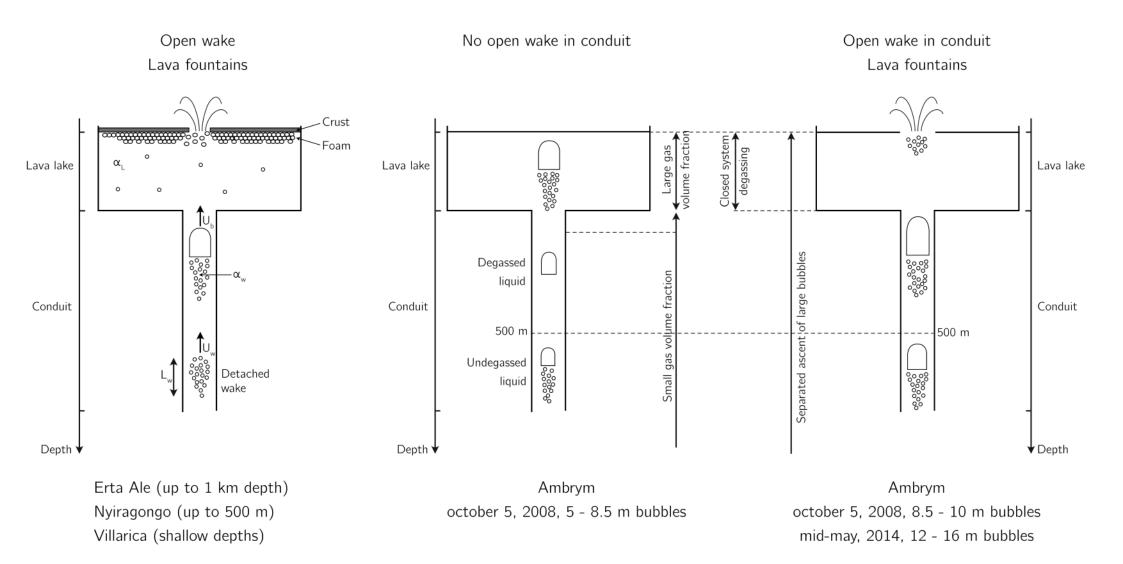


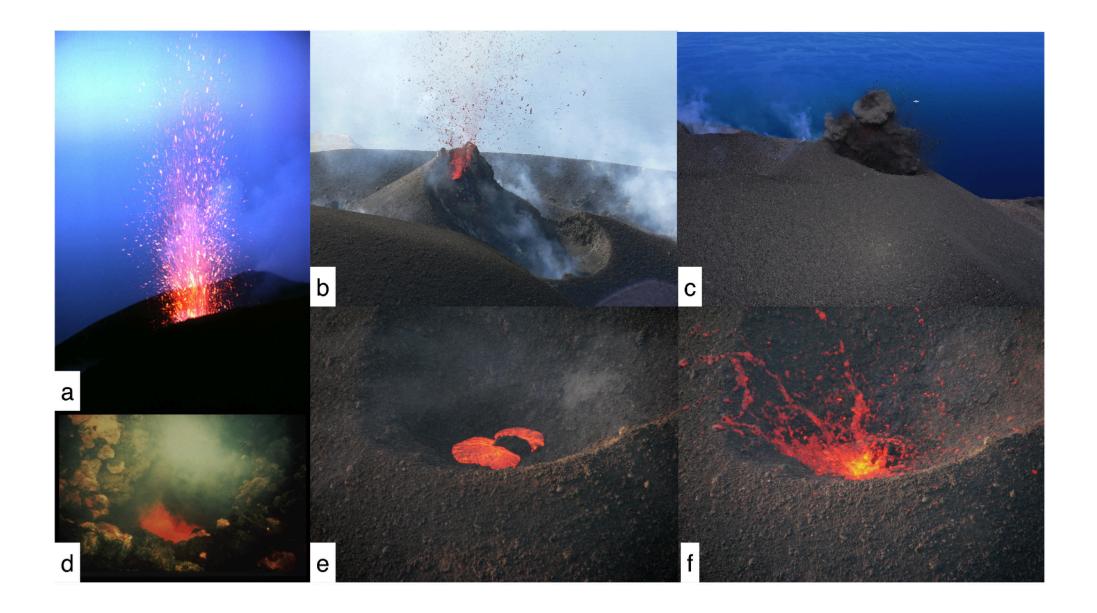


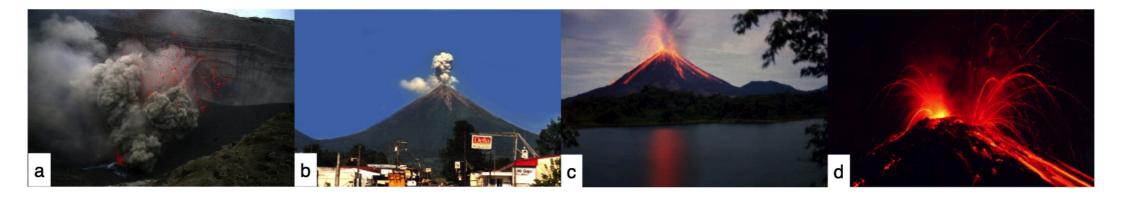




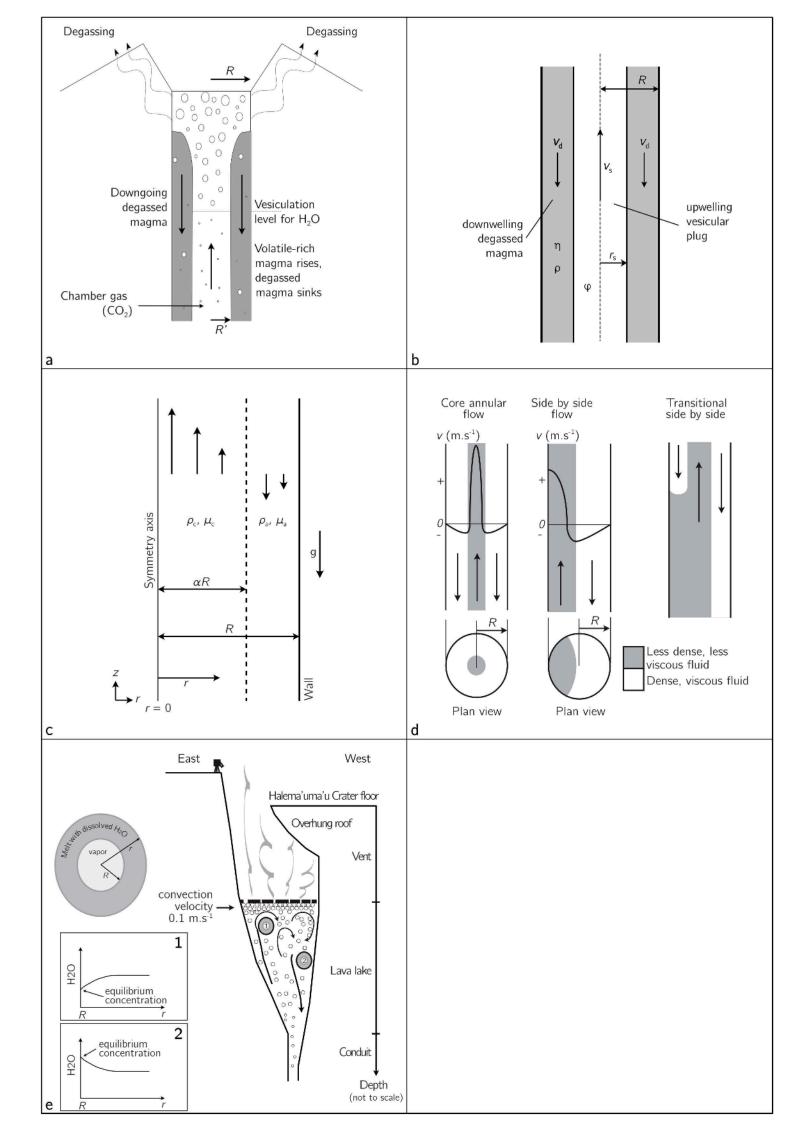












## Supplementary material by Vergniolle and Metrich

"An interpretative view of open-vent volcanoes" Figure caption:

Figure S1: Fig S1 (a) Viscosity (log Pa.s) as a function of inverse temperature (K<sup>-1</sup>) for anhydrous and crystal- and bubble-free phonolite melts of Erebus (Le Losq et al., 2015); for phonolite from Vesuvius (GP79) the experimental data are from Giordano et al., (2009), for rhyolite and andesite melts from Neuville et al. (1993), and for a basalt melt from Villeneuve et al. (2008). Figure redrawn from Le Losq et al., (2015). (b) Viscosity (log Pa.s) as a function the temperature and the H<sub>2</sub>O content of trachyandesitic melt of Yasur varying from 0 to 2.5 wt%, after Giordano et al., (2008). 52 wt% SiO<sub>2</sub> refers to Yasur present-day bulk rock and 56 wt% to the glassy matrix (Métrich et al., 2011). (c) Relative increase in magma viscosity as a function of the solid volume fraction, see text for details (slightly adapted from Mader et al., 2013). (d) Relative increase in magma viscosity as a function of the bubble shape, here quantified by the capillary number, see text for details, (slightly adapted from Mader et al., 2013).

