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Analysis and Mapping of an Updated Terrestrial Heat Flow Data Set

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Francis Lucazeau

Institutions: Institut de Physique du Globe de Paris

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Analysis and mapping of an updated terrestrial heat flow dataset

F. Lucazeau¹

¹Université de Paris, Institut de physique du globe de Paris, CNRS, IGN, F-75005 Paris, France

Key Points:

- This compilation of terrestrial heat flow includes almost 70,000 measurements
- Global heat loss based on measurements only ranges between 40-42 TW, close to previous estimates using conductive cooling models (45-47 TW)
- A generalized similarity method is used to map global variations of surface heat flow on a 0.5° x 0.5° grid. The best prediction is obtained for 14 observables.

Corresponding author: F. Lucazeau, lucazeau@ipgp.fr

Abstract

The number of heat flow measurements at the Earth surface has significantly increased since the last global analysis (Pollack et al., 1993), and the most recent of them provide insights into key locations. This paper presents a new compilation, which includes approximately 70,000 measurements. Continental heat flow (67 mWm^{-2}) does not change significantly, but the differences are more important for oceanic heat flow. The divergence with conductive cooling models is reduced significantly for young ages of the seafloor, since the most recent measurements (92 mWm^{-2}) are significantly higher on average than the older ones (79 mWm^{-2}). This is related to a better quality and a better sampling of measurements in regions affected by hydrothermal circulation. The total Earth heat loss derived from these most recent measurements is estimated to $\sim 40\text{-}42 \text{ TW}$ and represents only 3-5 TW less than with a conductive cooling model ($\sim 45\text{-}47 \text{ TW}$). Hydrothermal heat loss in the oceanic domain is estimated with a new method based on the ruggedness of the seafloor, and represents $\sim 1.5 \text{ TW}$ more than previous estimates. The heat flow variability on continents is so large that defining a trend with stratigraphic or tectono-thermal age is difficult and makes extrapolation from age a poor predictor. On the other hand, additional geological and geophysical information can be combined with age for better predictions. A generalized similarity method was used here to predict heat flow on a global $0.5^\circ \times 0.5^\circ$ grid. The agreement with local measurements is generally good and increases with the number and quality of proxies.

1 Introduction

Surface heat flow, which represents the energy flux through a unit surface, is a fundamental expression of the Earth dynamics. The measurement technique, based on the Fourier's law, simply relates the surface heat flow to the product of the vertical thermal gradient and thermal conductivity. However, measuring thermal gradient is not practically easy and the measurement itself is complicated by several types of surface perturbations (climatic changes, topography, water circulation, erosion or sedimentation): this is why the number of values is so small compared to other geophysical observations. The first estimates of the Earth gradient were given in the first part of the 19th century (Fourier, 1827), but the first heat flow measurements were reported only one century later by Bullard (1939) and Benfield (1939) for continents, and by Revelle and Maxwell (1952) for oceans. Then, the number of heat flow measurements increased regularly, allowing statistical analysis, global mapping and heat budgets (Chapman & Pollack, 1975; Jessop et al., 1976; Sclater et al., 1980; Vitorello & Pollack, 1980; Sclater et al., 1981; Stein & Stein, 1992; Pollack et al., 1993). Up to the eighties, the heat flow community was active and well organized in the International Heat Flow Commission (IHFC) and produced regular compilations of data (Jessop et al., 1976; Hurtig et al., 1991; Cermak et al., 1992; Pollack et al., 1993) or guidelines to standardize and correct these data (Balling et al., 1981). But after the 1993 compilation (Pollack et al., 1993), the database was not maintained regularly by the IHFC, and only individual initiatives (Artemieva, 2006; Hasterok & Chapman, 2008; Lucazeau & Poort, 2015) tried to keep out an up-to-date record. One of these compilations (Hasterok & Chapman, 2008) with 35,523 terrestrial measurements and 23,013 marine measurements is available online (<http://www.heatflow.org/data>). I started myself to maintain a database named NGHF (New Global Heat Flow), which has now grown to almost 70,000 values, while the Pollack's database only included 24,774 observations. The quality of the new measurements is usually better: marine techniques have been considerably improved with micro-computerized devices, in-situ measurements of thermal conductivity and multi-penetrations measurements colocated along seismic profiles (Hyndman et al., 1979; Von Herzen et al., 1989; Davis et al., 1997). In addition, new techniques have extended the domains of measurements on

63 continental margins, with the use of BSR (Bottom Simulating Reflectors) observa-
 64 tions along high resolution seismic profiles (Yamano et al., 1982; Hyndman et al.,
 65 1992; Kaul et al., 2000; Lucazeau et al., 2004), and on mid-ocean ridges, with the use
 66 of short probes operated by submarine engines (H. P. Johnson et al., 1993; Becker et
 67 al., 1996) or thermal blankets (P. Johnson & Hutnak, 1997; H. P. Johnson et al., 2010;
 68 Salmi et al., 2014). In the 2000s, the depletion of conventional oil reserves has stimu-
 69 lated the research on deep continental margins, including acquisition of new heat flow
 70 measurements (White et al., 2003; Lucazeau et al., 2004, 2008; Calvès et al., 2010);
 71 processing of oil exploration data to derive heat flow has also been improved by new
 72 techniques for estimating thermal conductivity from geophysical logs (Hartmann et al.,
 73 2005; Goutorbe et al., 2006; Fuchs & Förster, 2013) or correcting bottom hole temper-
 74 atures (Goutorbe et al., 2007). On land, several teams have maintained measurements
 75 activity (Mareschal & Jaupart, 2004) and several regional compilations have identified
 76 heat flow data not included in Pollack’s compilation (Hu et al., 2001; Tanaka et al.,
 77 2004; Blackwell & Richards, 2004; Jiang et al., 2016, 2019).

78 Beside data collections, several attempts to extrapolate or predict heat flow from
 79 other observations have been developed. The most commonly used predictors are ge-
 80 ology (Lee, 1965) and age (Chapman & Pollack, 1975); recent analysis using them
 81 include Davies and Davies (2010) and Davies (2013), while Artemieva (2006) exclu-
 82 sively used them for continents. Elevation and age of the sea-floor constrain isostatic
 83 thermal models of the oceans (Sclater & Francheteau, 1970; Parsons & Sclater, 1977).
 84 Only few additional observations have been used as heat flow predictors: Shapiro and
 85 Ritzwoller (2004) extrapolated surface heat flow from upper mantle shear velocities, Li
 86 et al. (2017) used the Curie point depths based on EMAG2 and Hasterok et al. (2011)
 87 used distance to seamounts and sediment thickness as predictors for the efficiency
 88 of hydrothermal circulations. Goutorbe et al. (2011) extended the use of empirical
 89 predictors to an unlimited number of observations and tested different methods to
 90 combine them: the best results were obtained with a similarity method combining
 91 several proxies, which is also the most efficient in term of computational resources.

92 The objective of this paper is to present the state of the art for the Earth heat
 93 flow trends and suggest what could be further progress in the future. I analyze first the
 94 NGHF database and provide statistics that compare to previous studies. I reexamine
 95 also the heat flow age relationships that were debated for both oceans and continents.
 96 I finally discuss extrapolation of the surface heat flow to the entire Earth and propose
 97 a new global map at the resolution of $0.5^\circ \times 0.5^\circ$ based on a similarity method and
 98 suggest possible locations for future studies.

99 **2 The NGHF data set**

100 NGHF is a relational database powered by MariaDB (MYSQL fork), initially set
 101 up from the GLOBHEAT database (Pollack et al., 1993) and related to several other
 102 global databases (tectonic setting, age, basin types, crust or sediment thickness, etc...).

103 All the GLOBHEAT metadata were kept in NGHF, at least for comparisons with
 104 the initial data. While some of them duplicate information of the related databases
 105 (e.g. geology or geography), others remain useful (measurement techniques, quality)
 106 and were carefully reported, when available in publications. For instance, code 6 of
 107 GLOBHEAT reports the measurement quality, defined as the variation of heat flow
 108 with depth (**A** for a variation less than 10 %, **B** less than 20 %, **C** less than 30 %
 109 and **D** greater than 30 %). This information is often not available in publications, but
 110 in that case, I attempted to derive it from a combination of other criteria (depth and
 111 number of measurements for instance). However, there still remains a large number of
 112 records with no information on the quality.

Table 1. Heat flow statistics and comparison with Pollack et al. (1993). Heat loss includes volcanic and non volcanic contributions.

	N	HF	N	HF volcanic	N	HF Pollack	Area	Heat loss
	measurements	(mWm^{-2})	cells	(mWm^{-2})	cells	(mWm^{-2})	(10^6km^2)	(TW)
Oceans								
0.0- 1.6 Ma	1880	275.5 ± 337.6	94	315.2 ± 243.0	15	139.5 ± 93.4	4.2	1.16
1.6- 5.3 Ma	1615	171.0 ± 102.3	196	150.5 ± 69.3	12	109.1 ± 81.2	11.4	1.95
5.3- 23.7 Ma	2699	99.9 ± 48.7	917	127.3 ± 61.1	73	81.9 ± 55.5	56.5	5.68
23.7- 36.6 Ma	700	71.5 ± 31.0	391	86.7 ± 30.7	28	62.3 ± 39.9	36.6	2.62
36.6- 57.8 Ma	657	64.5 ± 22.1	490	63.7 ± 18.9	35	61.7 ± 29.3	46.6	3.01
57.8- 66.4 Ma	243	60.8 ± 19.6	171	61.0 ± 15.0	9	65.1 ± 34.3	16.5	1.00
66.4- 84.0 Ma	648	57.7 ± 16.1	321	59.8 ± 19.8	24	61.5 ± 31.5	31.2	1.80
84.0-119.0 Ma	1618	55.4 ± 13.6	669	52.3 ± 16.6	48	56.3 ± 21.8	50.3	2.78
119.0-144.0 Ma	987	51.8 ± 11.3	376	50.4 ± 13.4	55	53.0 ± 21.6	28.1	1.45
144.0-178.0 Ma	677	51.3 ± 11.4	286	52.0 ± 10.9	61	51.3 ± 16.9	17.7	0.91
≥ 178.0 Ma	118	38.0 ± 31.7	59		0		0.8	0.03
<i>average</i>	<i>11842</i>	<i>78.8 ± 76.0</i>	<i>3970</i>	<i>85.9 ± 83.2</i>	<i>360</i>	<i>66.2</i>	<i>299.9</i>	<i>23.78</i>
Continents								
Archean (>2500 Ma)	280	46.2 ± 20.9	98	85.2 ± 150.5	25	51.5 ± 25.6	4.7	0.23
Proterozoic (2500-570 Ma)	1737	59.2 ± 33.6	551	70.2 ± 55.4	76	58.3 ± 23.6	22.4	1.35
Paleozoic (570-245 Ma)	5743	59.3 ± 40.7	1605	83.1 ± 73.3	243	61.0 ± 30.2	29.2	1.81
Mesozoic (245-65 Ma)	7836	64.8 ± 40.4	1777	97.2 ± 133.3	151	63.7 ± 28.2	21.9	1.49
Cenozoic (65-2.55 Ma)	11937	62.8 ± 28.2	1577	107.7 ± 160.4	351	63.9 ± 27.5	26.1	1.78
Quaternary (2.5-0 Ma)	14294	69.8 ± 46.6	1648	142.9 ± 343.4	255		35.6	2.64
Continental margins	14098	74.0 ± 58.2	2989	147.8 ± 591.8	131	77.7 ± 53.6	64.9	5.02
<i>average</i>	<i>52641</i>	<i>67.1 ± 47.1</i>	<i>10146</i>	<i>111.1 ± 270.3</i>	<i>1230</i>	<i>65.0</i>	<i>211.3</i>	<i>14.90</i>
Basins								
Divergent margins	2327	69.4 ± 53.2	712	85.5 ± 51.2	48		16.0	1.13
Back-Arc basins	2237	84.3 ± 43.1	662	110.5 ± 41.8	7		7.5	0.64
Transform margins	306	75.6 ± 45.1	148	61.4 ± 26.8	3		3.4	0.25
Continental rifts	818	113.6 ± 93.8	108	302.4 ± 271.7	6		1.0	0.11
Cratonic basins	203	48.6 ± 17.1	22		0		2.4	0.11
Foreland basins	472	52.9 ± 21.3	98		0		1.5	0.08
Foldbelts	335	76.9 ± 37.1	92		0		1.5	0.14
Forearc basins	422	61.6 ± 35.1	159	43.6 ± 6.3	6		2.6	0.16
Subductions	909	72.2 ± 38.4	341	55.0 ± 15.6	9		5.0	0.36
<i>average</i>	<i>21600</i>	<i>69.5 ± 51.2</i>	<i>5061</i>	<i>119.6 ± 438.8</i>	<i>328</i>		<i>99.1</i>	<i>7.12</i>
Global average	64491	69.4 ± 52.3	14261	106.6 ± 438.8	1668		511.2	38.74

113 The new measurements added to NGHF are distributed all over the world and
 114 represent an enhancement of about 10% of the surface covered by data (more informa-
 115 tion on the distribution of the previous and new measurements is available in figures S1
 116 and S2 of the supporting information). The database was cross-checked with that of
 117 Hasterok and Chapman (2008) and with the original publications, in order to identify
 118 misplaced or forgotten data, but as this can be done only manually, it takes time and
 119 the process is not yet achieved. On the other hand, an automatic check to remove
 120 redundant data was used to clean up the database. There is now a total of 29323
 121 marine and 39901 terrestrial measurements, but 4702 very low quality measurements
 122 (**D**) were not used in this study.

123 As the distribution of data is quite uneven, two types of statistics were used, (1)
 124 arithmetic means for a selected province or a selected range of ages and (2) surface-
 125 weighting averages on a $0.5^\circ \times 0.5^\circ$ grid. Arithmetic average and associated standard
 126 deviation are reported only for estimating the quality of distributions, for which his-
 127 tograms are also provided in figure S3 of the supporting information. The second
 128 method involves a first step making arithmetic mean (or preferentially median) \bar{q}_i for
 129 each cell of the grid and then a second step making an average of heat flow weighted
 130 by the surface A_i of each cell, which can vary with latitude:

$$\bar{Q} = \frac{\sum_i A_i \bar{q}_i}{\sum_i A_i} \quad (1)$$

131 As mentioned in Pollack et al. (1993), biases related to hydrothermal circulation
 132 exist in both oceans and continents. I show later in this paper that these biases have
 133 been reduced considerably in oceans thanks to the number of dedicated high resolution
 134 studies. On continents, some sites are already tagged as geothermal sites, but as this
 135 information is not systematic, I considered that measurements above a threshold of
 136 250 mWm^{-2} are also belonging to geothermal areas, which appeared an empirically
 137 relevant limit. This does not solve however the problem of downward percolation
 138 associated with low heat flow.

139 In the following paragraphs of this section, I provide statistics that can be com-
 140 pared to Pollack et al. (1993), and discuss the differences and the similarities with
 141 their study. I used the same oceanic and continental classifications (table 1), but as
 142 the origin of geological information is more recent, this could also introduce some dif-
 143 ferences in the analysis that are difficult to estimate. In this study, the age of oceanic
 144 seafloor is derived from Müller et al. (2008) unless it was mentioned in the original
 145 publications and the surface geology results from a compilation of USGS (2000) and
 146 UNESCO CCGM/CGMW (2000) shapefiles (figure 1a). In order to be consistent with
 147 Pollack et al. (1993), only areal statistics are presented in table 1 and are used in the
 148 calculation of global heat loss. However, median values, arithmetic averages and stan-
 149 dard deviations are provided as quality estimators of the distributions. Statistics are
 150 also given both for normal lithosphere and lithosphere affected by intraplate volcanism
 151 (figure 1a).

152 In the oceanic domain, average values derived from NGHF are systematically
 153 higher than GLOBHEAT values for Cenozoic, but roughly similar for Mesozoic. At
 154 the locations where off-axis volcanism exists, there are also generally higher values for
 155 young ages of the seafloor that contribute to a significant increase of surface heat flow.
 156 The global heat loss of the oceanic domain can be extrapolated from the normal and
 157 volcanic areas of the Earth, and their respective ages. The resulting value is signif-
 158 icantly higher (23.8 TW) than for Pollack (19.7 TW). The quality of measurements
 159 has a significant effect on the statistics and oceanic heat loss. Quality A data (heat
 160 flow variation less than 10 % with depth) have an arithmetic mean of 126.3 ± 235.4
 161 mWm^{-2} , compared to $177.4 \pm 434.5 \text{ mWm}^{-2}$ for all measurements except the D cat-
 162 egory, but conversely, the surface-weighted average is higher for quality A ($87.3 \pm$

163 132.3 mWm^{-2} vs $78.8 \pm 76.0 \text{ mWm}^{-2}$), which increases the oceanic heat loss to 26.4
 164 TW. As the quality of measurements significantly increased when electronic devices
 165 were introduced in instruments (in the 90's), the date of publication can also be used
 166 to filter out the lesser quality data. Additionally, the improvement of seismic imag-
 167 ing techniques allowed a better positioning of heat flow measurements, where a thick
 168 sediment cover protects basement from hydrothermal circulation. For measurements
 169 after 1990, the surface weighted-average values is $84.2 \pm 114.1 \text{ mWm}^{-2}$ and $90.0 \pm$
 170 198.1 mWm^{-2} if combined with category A. On the other hand, if the Hasterok et
 171 al. (2011) criterion (sediment thickness greater than 400 m and distance to the closest
 172 seamount greater than 60 km) is used to exclude areas affected by hydrothermal cir-
 173 culation, the oceanic average heat flow decreases because a large number of high value
 174 near ridges axis are excluded. This can be overcome if the criterion is not applied
 175 to ages between 0 and 1.6 Ma: in that case, the surface-weighted average heat flow
 176 increases to $82.5 \pm 106.7 \text{ mWm}^{-2}$ and $88.1 \pm 181.9 \text{ mWm}^{-2}$ for category A only. A
 177 topographic ruggedness criterion was also proposed by Le Gal et al. (2018) to identify
 178 areas affected by hydrothermal circulation. This criterion is based on the variogram
 179 sill of topography (elevation difference at a given distance): the authors showed that
 180 for a sill greater than ~ 75 m, heat flow is largely affected by hydrothermal processes.
 181 However, the method is hardly suitable for a global analysis, and instead, I used a
 182 simpler ruggedness index (Wilson et al., 2007) with a threshold value of 50 m. In that
 183 case, oceanic heat flow average is $83.1 \pm 99.1 \text{ mWm}^{-2}$ and $89.1 \pm 173.0 \text{ mWm}^{-2}$ for
 184 category A only.

185 The best quality (category A, published after 1990, associated with a significant
 186 sediment thickness) oceanic measurements averages suggest therefore that the oceanic
 187 heat flow is at least 88-90 mWm^{-2} and the oceanic heat loss >27 TW.

188 In the continental domain, average values are relatively similar to those obtained
 189 by Pollack et al. (1993), but NGHF averages are again higher for Cenozoic, especially
 190 for regions affected by volcanism. A large number (22184) of data derived from oil
 191 exploration was included in the new database, but this did not introduce a significant
 192 bias as the weighted averages are relatively similar. The arithmetic mean is slightly
 193 lower for oil data ($62.4 \pm 21.3 \text{ mWm}^{-2}$) than for non oil data ($68.4 \pm 45.5 \text{ mWm}^{-2}$).
 194 The statistical effect related to the measurements quality is also limited: the arithmetic
 195 mean of quality A data (heat flow variation less than 10 % with depth) is 68.3 ± 37.9
 196 mWm^{-2} compared to $62.2 \pm 41.2 \text{ mWm}^{-2}$ for quality D data (heat flow variation
 197 higher than 30 %), which have not been included however in the present study. The
 198 global continental heat loss is slightly higher (14.9 TW) than the Pollack estimate
 199 (13.9 TW).

200 The total Earth heat loss based on the NGHF heat flow measurements is ~ 38.7
 201 TW (33.6 TW for Pollack), and up to 42.0 TW with the highest quality oceanic data.
 202 This is very similar to previous estimates using a conductive cooling model for the
 203 oceanic domain instead of heat flow data (Sclater et al., 1981; Pollack et al., 1993;
 204 Sclater et al., 2014; Jaupart et al., 2015), which would lead to 44-45 TW with the
 205 NGHF dataset.

206 In addition, I provide in table 1 statistics for sedimentary basins, where heat
 207 flow is often estimated from oil exploration wells. The average is comparable to the
 208 global average (table 1), which confirms that oil data do not add substantial statisti-
 209 cal bias, as the lower quality of temperature estimates is probably balanced by the
 210 greater depth of measurements. The variations of heat flow within sedimentary basin
 211 types are relatively straightforward, with the highest values for active extensional ar-
 212 eas (rifts, back-arc basins) and the lowest values for convergent areas (foreland and
 213 forearc basins) and cratonic basins.

214 3 Oceanic heat flow

215 The NGHF database contains 15333 oceanic data, ranging from -302 to 33448
 216 mWm^{-2} , with an arithmetic average $\text{HFa} = 205 \pm 672 \text{ mWm}^{-2}$, a median value
 217 $\text{HFm} = 78.2 \text{ mWm}^{-2}$ and a surface weighted average $\text{HFw} = 70.8 \text{ mWm}^{-2}$. The
 218 large range (including negative values) and the large standard deviation associated
 219 with the arithmetic average are mostly related to the non conductive processes near
 220 mid ocean ridges. When the lowest quality measurements (category **D**) are excluded,
 221 the number of values decreases to 11842, with an arithmetic average $\text{HFa} = 179 \pm$
 222 425 mWm^{-2} , a median value $\text{HFm} = 80.4 \text{ mWm}^{-2}$ and a surface weighted average
 223 $\text{HFw} = 79.3 \text{ mWm}^{-2}$. Histograms of the heat flow distribution are shown in figure
 224 S3 of the supporting information. These average values (table 1) are significantly
 225 higher than in Pollack et al. (1993), but this is mainly related to Cenozoic for which
 226 new measurements have a better quality and a better spatial distribution. However,
 227 the data coverage remains only 2.7% of the Earth surface (4.6 % of the global ocean
 228 surface) on a $0.5^\circ \times 0.5^\circ$ grid (see table S7 in supplementary material). Extrapolation
 229 of the average value HFw to the Earth ocean surface leads to an oceanic heat loss of
 230 23.8 TW compared to 18.5 TW for Pollack et al. (1993) and as discussed previously,
 231 rises up to 27 TW if one only considers the most recent data (publications starting in
 232 1990).

233 The decrease of heat flow with age and the subsidence of the sea floor are usually
 234 interpreted as a result of the passive cooling of oceanic lithosphere from the accretion
 235 ridge. Subsidence of the sea floor is therefore a major constraint on the oceanic ther-
 236 mal regime, because it is less affected by superficial processes than the surface heat
 237 flow: the most important of these effects is related to hydrothermal circulation, which
 238 has a dramatic mining effect where the impermeable sediment cover is not thick and
 239 continuous enough to prevent the recharge of sea water in a highly permeable upper
 240 crust (Lister, 1972; Davis et al., 1992; Villinger et al., 2002).

241 The first analysis of the ocean topography discussed whether the lithosphere was
 242 cooling continuously according to a boundary layer model (Davis & Lister, 1974) or
 243 reached a limit after a certain time like a plate model (Parsons & Sclater, 1977). The
 244 flattening of subsidence for ~ 80 Ma age (Parsons & Sclater, 1977) promoted the plate
 245 model rather than the boundary layer model in further analyses, but conversely the
 246 assumption of a constant temperature at the base of the lithosphere requires additional
 247 heat sources, such as the onset of small scale convection (Davaille & Jaupart, 1994) or
 248 lithosphere reheating (Menard & McNutt, 1982; Nagihara et al., 1996).

249 Different approaches were used to process heat flow and topographic data. While
 250 Parsons and Sclater (1977) selected only high quality data, Stein and Stein (1992) used
 251 a global statistical approach, which led to significantly different results. Both studies
 252 promoted the plate model, which explains the flattening of the topography for old
 253 ages of the sea floor, but the lithosphere was significantly hotter and thinner in the
 254 Stein and Stein (1992) analysis. Parsons and Sclater (1977) had a best fit using a 125
 255 km thick plate with 1333°C at the bottom, while Stein and Stein (1992) suggested
 256 95 km and 1450°C , which is too high for the average composition of MORB and
 257 the average thickness of oceanic crust (McKenzie et al., 2005; Jaupart et al., 2015).
 258 Hillier and Watts (2005) reexamined bathymetry of the North Pacific using a filtering
 259 method to remove components of the topography that are not related to thermal
 260 contraction. They found that the flattening occurs for older age (85 Ma) than Stein
 261 and Stein (1992), but still observed a deviation from the boundary layer model. The
 262 importance of the processing method was recently reopened by Adam and Vidal (2010)
 263 who analyzed bathymetry along trajectories representative of the underlying mantle
 264 flow and obtained conversely a better agreement with the boundary layer model. On
 265 the other hand, Goutorbe (2010) adding constraints provided by shear waves velocities
 266 in the upper mantle, preferred a constant heat flow at the base of the lithosphere

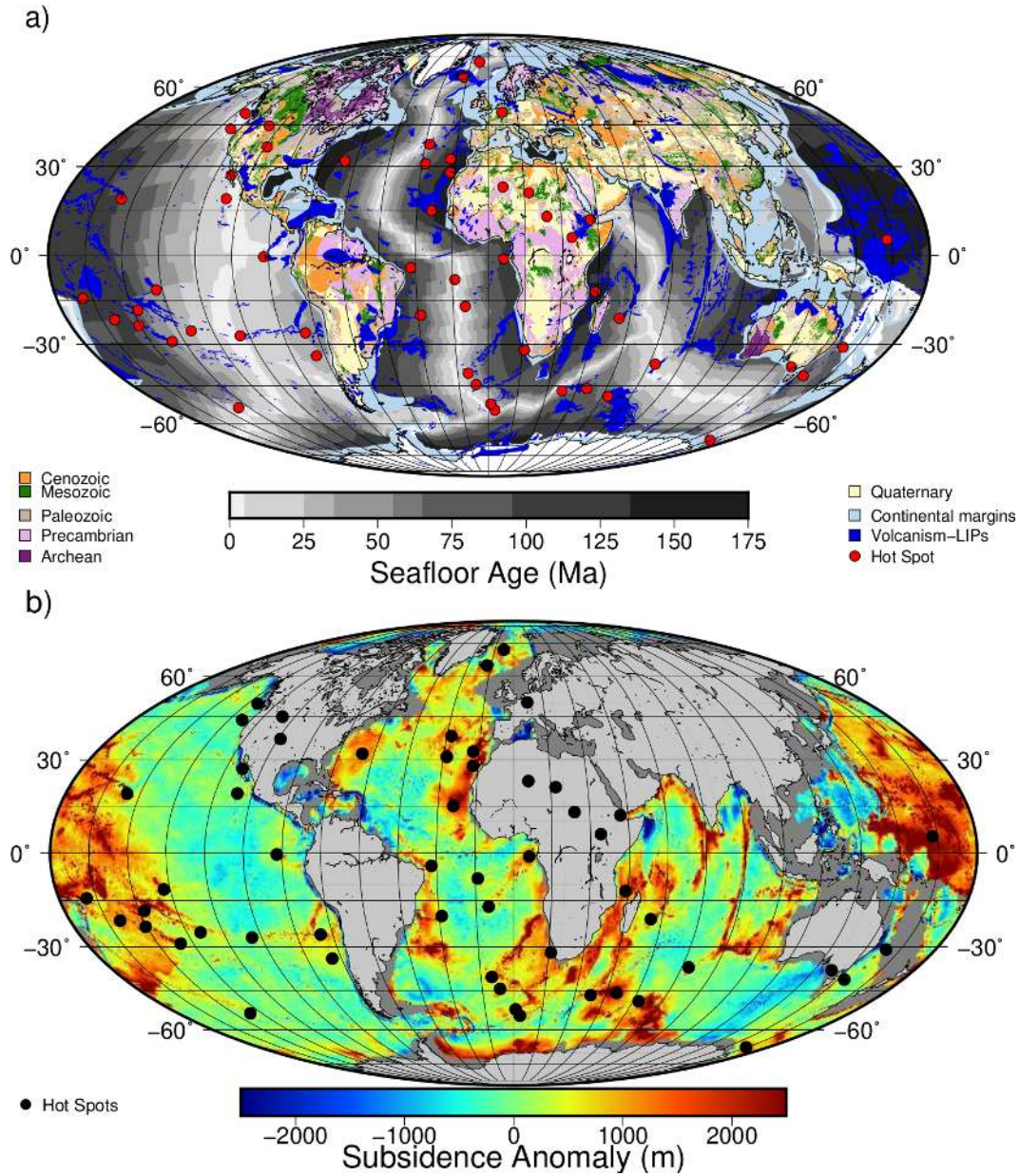


Figure 1. a) Age of continents (CCGM/CGMW, 2000) and oceans (Müller et al., 2008) used in the statistical analysis of this paper. The scale refers to the age of seafloor (in Ma). Red dot are the main hot spots, and blue color corresponds to LIPs or flood basalts (Johansson et al., 2018). b) Anomalies of subsidence (in meters) with respect to the elevation h given by the boundary layer model with Parsons and Sclater (1977) values: $h = 2500 + 350 \sqrt{\mathcal{T}}$ where \mathcal{T} is the age of the sea floor (Ma). Positive anomalies are mostly related to hot spots locations, hot spots tracks or large igneous provinces (see figure 1a for the location of LIPs).

(CHABLIS), initially proposed by Doin and Fleitout (1996). Other studies (Goutorbe & Hillier, 2013; Grose & Afonso, 2013; Hasterok, 2013) examined the effects of filtering global topography of the oceans or the effect of parameters values on the model results, but did not resolve the question of why there are so much difference between observed subsidence on selected oceanic domains as in Parsons and Sclater (1977) or Adam and Vidal (2010), and subsidence obtained from global topography of oceans.

In the present study, I reexamine this paradox and I started to reprocess topography at heat flow sites only. In most cases, I used topography as given at these sites in the publications, otherwise GTOPO30 was used. In order to have comparable subsidence estimates, topography was corrected for sediment loading (Le Douaran & Parsons, 1982) and for variations of the crustal thickness:

$$S = Z + e_{sed} \frac{\rho_m - \rho_s}{\rho_m - \rho_w} + (e_{crust} - e_{ref}) \frac{\rho_m - \rho_c}{\rho_m - \rho_w} \quad (2)$$

where S and Z are respectively the corrected and uncorrected topography, e_{sed} the sediment thickness, ρ_m the mantle density, ρ_s the sediment density, ρ_w the sea water density, ρ_c the oceanic crust density, e_{crust} the local crustal thickness and $e_{ref} = 6$ km the reference thickness of the crust. The sediment thickness was obtained from global maps (Divins, 2003; Whittaker et al., 2013) when no other information was available and the crustal thickness from CRUST1.0 (Laske et al., 2013).

Heat flow was binned for variable time intervals Δt , increasing with the age of the sea-floor, in order to have a better definition for young ages for which heat flow varies the most:

$$\Delta t = (2i - 1)\Delta t_0 \quad (3)$$

where i [1, 126] is the i^{th} age interval and $\Delta t_0 = 0.011236$ Ma (t varying from 0 to 178.4 Ma).

The resulting sea floor subsidence is significantly different from that predicted by GDH1 for age >65 Ma but in good agreement with PS model (figure 2) or the more recent analysis of Richards et al. (2018), which selected 2,028 reliable sites where sediment thickness and crustal thickness are well known. In order to check if these differences were related to methodological problems, the same analysis was performed using just the global topography (GTOPO30). In this case, the agreement with GDH1 is excellent (figure 2). Additionally, bathymetry at selected sites in the North Pacific and for ages of the sea floor >100 Ma (Parsons & Sclater, 1977) is in good agreement with the trend defined by bathymetry at heat flow sites. On the other hand, heat flow measurements are more consistent with the GDH1 model. I conclude therefore that (1) the heat flow measurements do not sample a representative surface of the oceanic domains, and (2) the existing models using global topography datasets do not provide a good fit for subsidence at heat flow sites.

Where the oceanic subsidence derived from GTOPO30 differs from the trend defined at heat flow sites only can be directly mapped as the difference between the Parsons and Sclater (1977) model and the observed topography (figure 1). The maximum deviations ($>+2000$ m) are all observed in the vicinity of hot spots, hot spots tracks, swells or large igneous provinces, like the Western Pacific around Ontong Java plateau, the African margins, the Crozet or Reunion swells and many other volcanic areas. In these anomalous zones, the average bathymetry is only 2500 m and the average age is 108 Ma. There are only a few heat flow measurements in these zones (41/15333 values and 34/5321 cells), which explains why subsidence at heat flow sites is so different from global data. However, the average heat flow (52.8 ± 16.4 mWm $^{-2}$) in these anomalous subsidence areas is not very different from that observed elsewhere on the old oceanic lithosphere.

314 The origin of the subsidence anomalies is obviously not related to crustal isostasy,
 315 since the variations of crustal thickness as well as the sediment load were corrected
 316 before. The coincidence with hot spots, LIPs or intraplate volcanic zones suggests
 317 that it is rather connected to residual thermal anomalies after an event that partially
 318 or totally reset the thermal state of the lithosphere (Menard & McNutt, 1982; Detrick
 319 & Crough, 1978) or to dynamic topography. The latter possibility does not seem
 320 however a likely explanation as the present day dynamic topography (Moucha et al.,
 321 2008) is totally uncorrelated with the anomalous subsidence. Nagihara et al. (1996)
 322 proposed a reheating process to explain heat flow and subsidence discrepancy with the
 323 PS model for an old oceanic lithosphere. As in the present study, they observed that
 324 heat flow is more consistent with GDH1 and subsidence with PS, but they interpreted
 325 the higher heat flow with respect to the PS model as a result of thermal reheating.
 326 In their interpretation, the surface heat flow would be perturbed by reheating of the
 327 lithosphere, but subsidence would follow the normal trend given by the PS model. My
 328 interpretation is different, and I rather postulate that reheating has affected specific
 329 zones (hot spots, LIPs), but that most of the data are characterized by a normal
 330 behavior that none of the existing models explains.

331 Therefore, as no existing model can match simultaneously heat flow and sub-
 332 sidence obtained at the same sites, I used a numerical model that provides more
 333 flexibility on parameters, and a Monte Carlo procedure to get their optimal values.
 334 The thermal model is a stationary and Eulerian 2D finite element model with a con-
 335 stant surface temperature, a temperature condition at the accretion axis, no heat flow
 336 through the other vertical boundary condition at a distance equivalent to 2 times the
 337 actual size of oceans and either a temperature or heat flow condition at the base of the
 338 box (figure 3). An advection term accounts for the thermal effect of oceanic accretion
 339 at a constant horizontal plate velocity V .

$$\frac{\partial}{\partial x} \left(\lambda(T, z) \frac{\partial T}{\partial x} \right) + \frac{\partial}{\partial z} \left(\lambda(T, z) \frac{\partial T}{\partial z} \right) - \rho C(z, T) V \frac{\partial T}{\partial x} = 0 \quad (4)$$

340 where T is the temperature, x the horizontal dimension, z the vertical dimension, ρ
 341 the density of the lithosphere, C the specific heat. The mesh is irregular and more
 342 refined near the ridge axis and near the surface (same progression as in equation 3).
 343 At the accretion axis, temperature of the mantle is a free parameter that decreases
 344 linearly in the upper 6 km of the lithosphere before reaching the surface temperature
 345 condition, for both numerical stability and to account for an accelerated cooling by
 346 hydrothermal processes in the crust like in Cochran and Buck (2001) or Spinelli and
 347 Harris (2011).

$$T(0, z) = T_L \quad \text{for } z_c \geq z \leq L \quad (5)$$

$$T(0, z) = T_L - \frac{T_L - T_s}{z_c} (z_c - z) \quad \text{for } z < z_c \quad (6)$$

349 where $x=0$ is the accretion axis, T_L the temperature of the mantle, T_s the surface
 350 temperature, z_c the thickness of the crust and L the depth of the lower boundary
 351 condition. The upper part of the lithospheric mesh is formed by a 6 km thick basaltic
 352 crust, which has a lower density ($\rho = 2850 \text{ Kg m}^{-3}$) and a lower thermal conductivity
 353 ($\lambda_c = 3.0 \text{ Wm}^{-1}\text{K}^{-1}$) than the lithospheric mantle. The importance of this insulating
 354 lid has been outlined by Grose and Afonso (2013), and leads to a significantly warmer
 355 lithosphere but lower heat flow for young ages of the seafloor (0-35 Ma).

356 As outlined by McKenzie et al. (2005), the thermal dependence of thermal con-
 357 ductivity including radiative transfer changes significantly the distribution of mantle
 358 temperatures and was included in the calculation. The radiative component differs
 359 significantly between two end-member models (Schatz & Simmons, 1972; Hofmeister,
 360 1999) for high temperatures (see figure S4 in the supporting information), but the
 361 differences for the lithospheric thermal regime remain negligible.

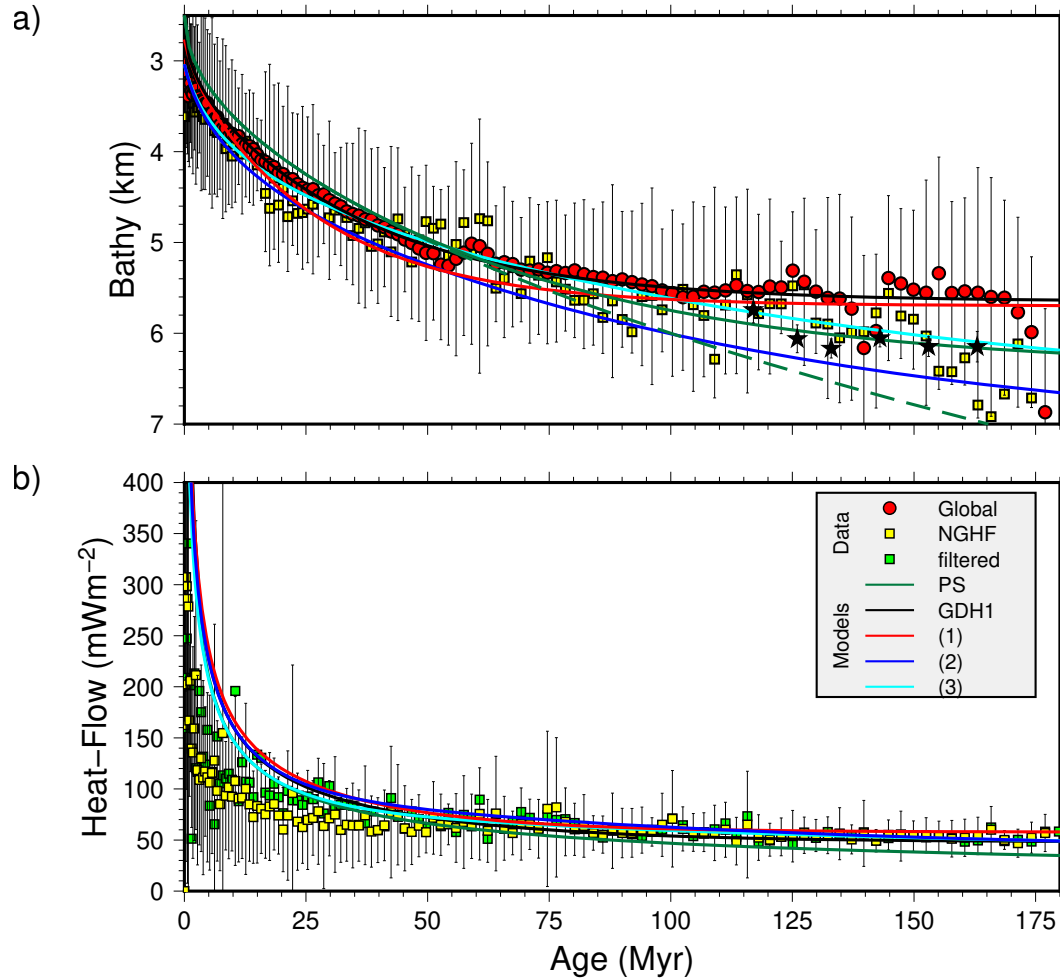


Figure 2. a) Elevation versus age of the ocean floor. Red and yellow symbols are respectively subsidence data from the global coverage and from a subset of locations where heat flow measurements exist. Both are corrected for sedimentation and crustal thickness variations and averaged for increasing time intervals. Stars are selected sites from North Pacific from the Parsons and Sclater (1977) paper. Lines correspond to models, black line to GDH1 (Stein & Stein, 1992), green line to PS (Parsons & Sclater, 1977), green dash line to boundary layer model and other color lines to numerical models (see text for details); b) heat flow versus age of the ocean floor; the green symbols correspond to filtered heat flow data in the same way as Hasterok et al. (2011), with sediment thickness and distance to the closest seamount larger than 400 m and 60 km respectively.

362 Because of the lateral heat conduction in the model, low plate velocities (<8
 363 km/Ma) result in higher surface heat flow near the ridge axis than in 1D models (see
 364 figure S5 in the supporting information. In this study, I used therefore a velocity (8
 365 km/Ma) that makes 2D results as close as possible with 1D models.

366 Subsidence of the sea floor results from thermal contraction when the lithosphere
 367 cools down and is obtained by integration of the density anomalies with respect to the
 368 ridge axis, using a local Airy assumption.

$$S_{calc} = \frac{\int_0^{Z_L} \rho(0, z, T) \partial z - \int_0^{Z_L} \rho(x, z, T) \partial z}{\rho_a - \rho_w} \quad (7)$$

369 with a temperature dependence of density:

$$\rho(T) = \rho_0(1 - \alpha T) \quad (8)$$

370 and with $\rho_a = \rho(T_L)$. The coefficient of thermal expansion α fitting lithosphere subsi-
 371 dence in previous studies ($\sim 3.1\text{-}3.2 \times 10^{-5} \text{ K}^{-1}$) is generally 15-30 % lower than de-
 372 termined experimentally (Korenaga, 2007; Goutorbe & Hillier, 2013; Grose & Afonso,
 373 2013). The origin of this reduction is not still well understood, and, here, I preferred
 374 to use it as a free parameter that makes comparison with conventional models easier.
 375 I let also specific heat constant, in order to minimize the number of free parameters.

376 Several other free parameters (table 2) were randomly selected using a Monte
 377 Carlo procedure in order to minimize the misfit with subsidence in the range 0-180
 378 Ma and the misfit with heat flow in the range 75-180 Ma. A unique estimator of the
 379 misfit was defined as:

$$RMS = \sqrt{\left(\frac{(S_{obs} - S_{calc})^2}{S_{ref}^2}\right)_{[0,180Ma]} + \left(\frac{(HF_{obs} - HF_{calc})^2}{HF_{ref}^2}\right)_{[75,180Ma]}} \quad (9)$$

380 where S_{obs} is the observed subsidence, S_{calc} the corresponding modeled subsidence,
 381 HF_{obs} the observed surface heat flow, HF_{calc} the modeled surface heat flow, and Z_{ref}
 382 and HF_{ref} are arbitrary values for normalization, respectively 5000 m and 40 mWm^{-2} .
 383 The choice of the type of lower boundary condition (temperature or heat flow) is also
 384 randomly chosen.

385 Three cases of subsidence curves are investigated here: (1) corresponds to the
 386 global set of topographic data, (2) to a subset of topographic data at heat flow sites
 387 only and (3) to the global set of topographic data, for which locations with subsidence
 388 anomalies $> +2000\text{m}$ were removed.

389 The optimal fit for case (1) corresponds to a fixed temperature at the base of
 390 the box (plate model) and is the closest of my results to the GDH1 model in term
 391 of subsidence flattening, but with a more pronounced curvature (figure 2). However,
 392 the optimal thickness (119 km) and mantle temperature (1300 °C) are different from
 393 those of GDH1 and more similar to the McKenzie et al. (2005) model (106 km and 1315
 394 °C), which also included the temperature dependence of thermal conductivity. Mantle
 395 thermal conductivity is high but fits the experimental results on sintered forsterite in
 396 Schatz and Simmons (1972) (see figure S4 in supporting information). The heat flow
 397 for old ages of the sea floor is also significantly higher than GDH1 (figure 2).

398 The optimal fit for case (2) corresponds to a constant heat flow 36.4 mWm^{-2} at
 399 the base of a 114 km thick lithosphere. Compared to the plate model for case (1), the
 400 temperature decreases faster near the axis and the slope of isotherm 600 °C is in very
 401 good agreement with the distribution of intraplate earthquakes (figure 3) published by
 402 McKenzie et al. (2005).

Table 2. Parameters of the numerical model, including the a priori ranges and the best a posteriori values. Mantle conductivity λ_L is the value at surface temperature; its dependence with in situ temperature includes a lattice component given by $\lambda_L(T) = \lambda_L(1 + \nu\Delta T)^{-1}$ where $\nu = 1.25 \cdot 10^{-3} K^{-1}$ and a radiative component. A map of all optimal values is also provided in supporting information. RMS is obtained according to equation 9. RMS values in parentheses correspond to subsidence only.

parameter	symbol	min	max	(1)	(2)	(3)	unit
Thermal expansion	α	$2.7 \cdot 10^{-5}$	$3.3 \cdot 10^{-5}$	$2.70 \cdot 10^{-5}$	$3.06 \cdot 10^{-5}$	$3.09 \cdot 10^{-5}$	K^{-1}
Boundary condition	L	90	120	119	114	118	km
Bottom temperature	T_L	1170	1430	1300	1235	1170	$^{\circ}C$
Bottom heat flow	q_L	27	33	0	36.4	37.1	mWm^{-2}
Ridge elevation	h_0	2.800	3.400	2.82	3.04	3.01	km
Mantle conductivity	λ_L	3.50	5.50	4.85	4.64	5.25	$Wm^{-1}K^{-1}$
RMS				0.88 (0.20)	0.89 (0.31)	0.75 (0.06)	
RMS GDH1				1.17 (0.41)	1.40 (0.88)	1.27 (0.63)	
RMS PS				2.46 (0.70)	2.49 (0.80)	2.41 (0.51)	
Heat loss				30.3	29.8	27.5	TW

403 The optimal fit for case (3) corresponds to a constant heat flow $37.1 mWm^{-2}$
 404 at the base of a 118 km thick lithosphere. As for case (1) and (2), mantle thermal
 405 conductivity is high but still in line with Schatz and Simmons (1972) results (see figure
 406 S4 in supporting information).

407 The RMS misfit only for heat flow / age (excluding subsidence) is less for cases
 408 (2) and (3) involving a basal heat flow boundary condition than for the plate model (1):
 409 these models can account for a slow decrease of heat flow for ages >65 Ma while the
 410 plate model predicts an almost constant value, that becomes higher than observations
 411 at oldest ages of the seafloor. Using the same RMS criteria, the GDH1 and PS models
 412 do not provide as good adjustment as numerical results, mostly because heat flow
 413 differs more significantly. As noticed before, the PS model heat flow is systematically
 414 lower than measurements, while GDH1 is close to case (2) for ages <40 Ma and for ages
 415 >150 Ma. Several other studies have explored different effects related to parameters
 416 values or different filters applied to topography of the seafloor. For instance, Grose
 417 and Afonso (2013) explored, through a sophisticated 2D numerical model, the effects
 418 of physical properties based on mineral physics and chemistry. They also accounted
 419 for axial hydrothermal circulation and for the effect of oceanic crust as a thermal
 420 insulator. Their models do not require an important reduction of the coefficient of
 421 thermal expansion as other studies (including this one), but they still need a reduction
 422 of 15 %. Although a direct comparison with their results is difficult, their conclusion
 423 on a cooler lithosphere than GDH1 model for seafloor younger than 35 Ma is similar to
 424 model (1), but their estimate of the oceanic heat loss (27 TW) is lower than model (1).
 425 On the other hand, as they do not consider models with a bottom heat flow condition,
 426 a comparison with model (2) and (3) is not possible. Compared to CHABLIS models,
 427 which imposes a constant heat flow at the base of an isotherm, bottom heat flow values
 428 in models (2) and (3) are similar to Doin and Fleitout (1996) and Goutorbe (2010).

429 The total oceanic heat loss corresponding to models (2) and (3) is 29.8 and 27.5
 430 TW. This does not include however a potential additional heat flow related to reheating

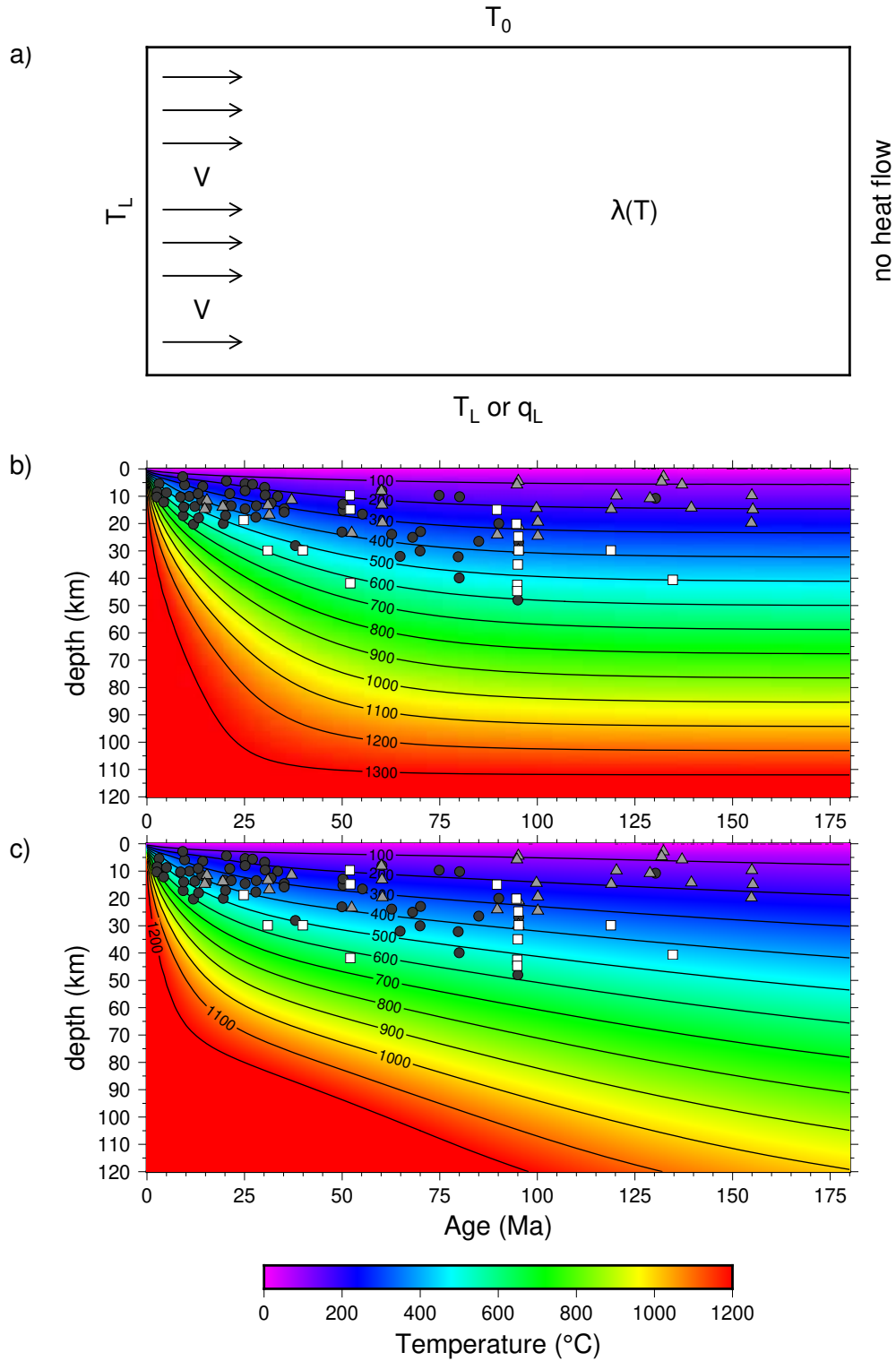


Figure 3. a) Finite element model sketch. The box width is $2 V t_{max} = 2 \times 8 \text{ km/Ma} \times 180 \text{ Ma} = 2880 \text{ km}$ and the box height is L . The lower boundary condition is either a fixed temperature T_L or heat flow q_L . Parameters values are given in table 2; b) lithospheric temperature field for model (1) with temperature as a lower boundary condition; c) lithospheric temperature field for model (2) with heat flow as a boundary condition. The dots (intraplate), triangles (outer rises) and squares (thrusts in outer rises) symbols show the depths of earthquakes within the oceanic lithosphere whose depths have been constrained by waveform modeling (McKenzie et al., 2005).

431 of the lithosphere, in regions where the subsidence anomaly exceeds +2000M (figure
432 1b).

433 Several studies estimated the hydrothermal heat flow as the difference between
434 conductive heat flow derived from thermal models and the observed heat flow. In
435 the Stein and Stein (1992) study, this difference is significant for ages < 65 Ma and
436 increases progressively with decreasing ages to reach 60 % of the conductive heat flow
437 near the ridge axis. The threshold at 65 Ma was interpreted as a possible permeability
438 closure in the uppermost lithosphere, which could limit hydrothermal circulation for
439 older ages of the sea floor, but several studies showed that such a permeability change
440 occurs more likely closer to the ridge axis (Fisher, 1998) and that the heat flow may
441 be perturbed by hydrothermal circulation for older ages of the sea floor (Von Herzen,
442 2004; Fisher & Von Herzen, 2005). The addition of new measurements or the filtering
443 of these measurements (Hasterok et al., 2011) can also change the resulting magnitude
444 of hydrothermal heat flow and/or the position of the permeability threshold (figure
445 2). I report in figure 4b the difference between a conductive model and the bins with
446 measurements obtained after 1990: the threshold would be near 20 Ma and the total
447 hydrothermal heat loss less than 2 TW! It seems therefore difficult to estimate the
448 oceanic hydrothermal heat loss as the difference between a conductive model and the
449 observed heat flow. On the other hand, the dispersion of heat flow values is clearly
450 related to the effect of hydrothermal circulation and can be used instead as a relative
451 indicator of its magnitude (figure 4a). For instance, the integration of one standard
452 deviation provides a first estimate (12.8 TW), which can be corrected for the natural
453 variability (15 mWm^{-2} in the oldest part of the oceanic domain, see figure 4a) and
454 give a cumulative estimate $\sim 9 \text{ TW}$ (figure 4b). An other and more quantitative way is
455 to extrapolate results of local studies (Villinger et al., 2002; Fisher, Davis, et al., 2003;
456 Fisher, Stein, et al., 2003; Lucazeau et al., 2006; Fisher & Wheat, 2010; Le Gal et al.,
457 2018), which show that the presence of permeable basalt exposures or alternatively of
458 a continuous impermeable sediment blanket are the main factors driving hydrothermal
459 circulation. Therefore, the local relief of the seafloor can show if the surface is smooth
460 and covered by impermeable sediment, or is rough and exposed to fluid circulation. Le
461 Gal et al. (2018) used variogram techniques to estimate the local relief and showed that
462 hydrothermal circulation effect reaches a maximum level of 60 % of the surface heat
463 flow for a local relief exceeding 150 m. Variogram is hard to use at global scale as it is
464 expensive in computer time, but other estimators like the ruggedness index (Riley et
465 al., 1999) can be used more easily. Although the local relief was not determined in the
466 same way, I used the relationship given by Le Gal et al. (2018) relating the fraction of
467 conductive heat flow observed at the surface to the ruggedness. This method provides
468 very similar results as the standard deviation method (figure 4b), showing a larger
469 range of ages (0-130 Ma) affected by hydrothermal circulation than proposed by Stein
470 and Stein (1992) and more consistent with observations of hydrothermal circulation on
471 old seafloor (Von Herzen, 2004). The previous estimates (Stein & Stein, 1994; Spinelli
472 & Harris, 2011) were not surprisingly far different from this one: they are mostly
473 affected by an offset of 1.7 TW that corresponds to the hydrothermal heat loss for age
474 older than 65 Ma (figure 4), which increases the hydrothermal heat loss to $\sim 10.5 \text{ TW}$.

475 3.1 Continental heat flow

476 The NGHF database contains 51621 continental data, more densely distributed
477 in North America (17885 from USA and 2985 from Canada). The range of values
478 varies from -52 to 15600 mWm^{-2} , with an arithmetic average $\text{HFa} = 68.1 \pm 51.1$
479 mWm^{-2} , a median $\text{HFm} = 62.8 \text{ mWm}^{-2}$ and a surface weighted average $\text{HFw} =$
480 68.0 mWm^{-2} . These statistics are not very different from those obtained by Pollack
481 et al. (1993) (table 1). 13139 of these measurements were obtained in vertical bore-
482 holes at thermal equilibrium, 18207 are derived from oil exploration wells and 14839

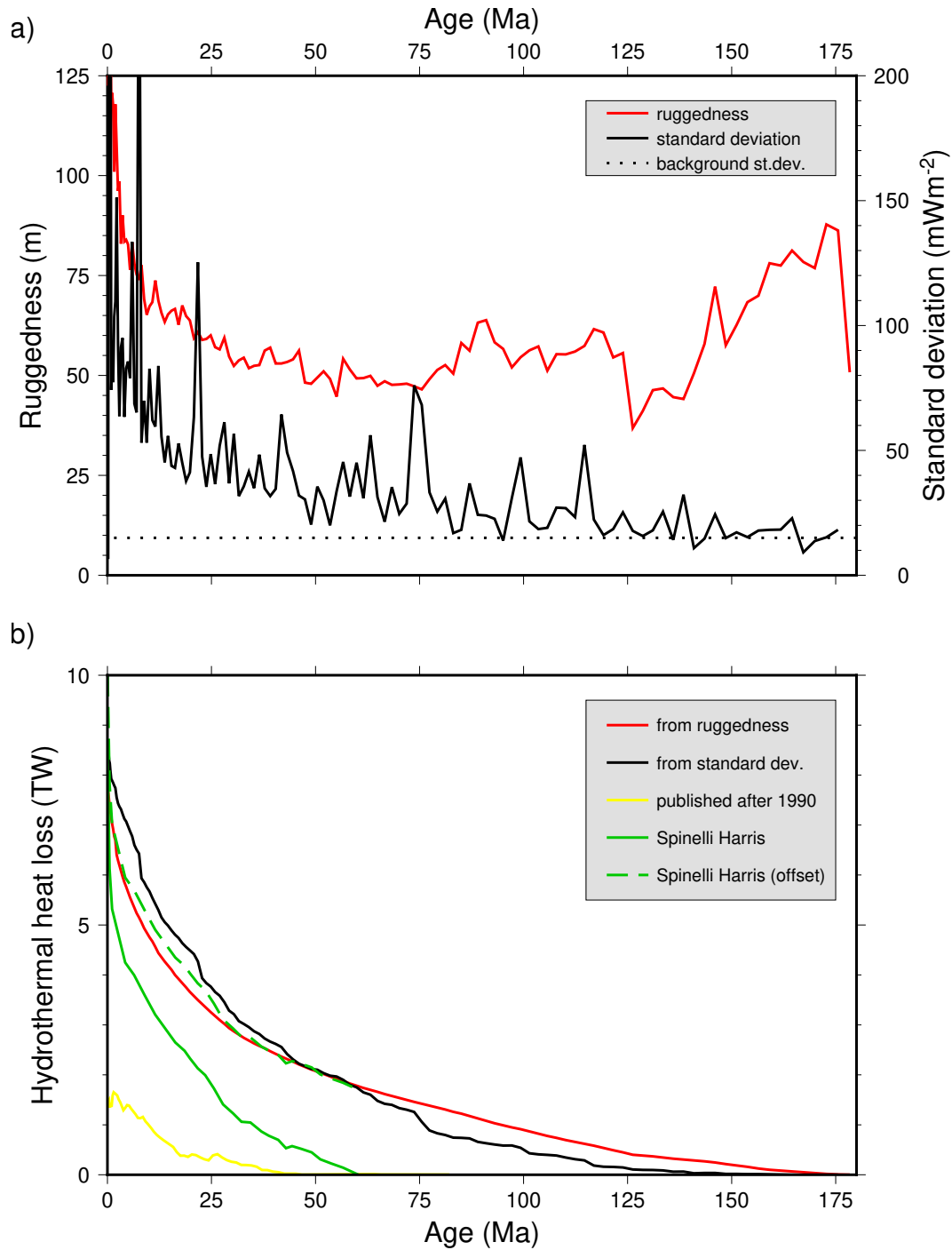


Figure 4. a) Average ruggedness and standard deviation (1σ) associated with observed heat flow measurements. The dotted line (15mWm^{-2}) is considered as the natural variability of heat flow in absence of hydrothermal circulation. b) Cumulative hydrothermal heat loss with respect to model 2 (see figure 2). The red line represents the cumulative perturbation according to an empirical relationship from local studies (Le Gal et al., 2018), which relates the fraction of conductive heat flow $\frac{Q}{Q_0}$ to the local relief S : $\frac{Q}{Q_0} = 3.8792S^{-0.4}$. The black line corresponds to an estimate of the hydrothermal heat flux based on the standard deviation. The yellow line is the difference between model 2 and measurements published after 1990. The green line is the Spinelli and Harris (2011) estimate and the green dash line is the same estimate offset by 1.7 TW, which corresponds to the hydrothermal heat loss for ages >65 Ma.

483 from marine techniques in lakes or continental margins. The origin of others measurements
 484 is not specified in the dataset. Average heat flow from oil wells measurements
 485 ($65 \pm 27 \text{ mWm}^{-2}$) is not statistically different from that obtained in boreholes at
 486 thermal equilibrium ($62 \pm 32 \text{ mWm}^{-2}$). There is also no important variation of the
 487 average heat flow between the main continental landmasses outside of the volcanic and
 488 geothermal areas, except Russia which has a lower value. Antarctica and Greenland
 489 are also different, but the differences are less significant because of the small number
 490 of measurements and the non representative areas covered by these measurements.

491 Several authors suggested that the continental heat flow could vary with age
 492 of the lithosphere (Polyak & Smirnov, 1968; Hamza & Verma, 1969; Vitorello et
 493 al., 1980; Sclater et al., 1981), in a similar way as oceanic lithosphere. The age of
 494 continents is more difficult to define, as a superposition of thermal processes (rifting
 495 or thickening of the crust, magmatic intrusions, erosion, etc.) can occur in space and
 496 in time. The last tectonic event having affected the thermal regime of the continental
 497 lithosphere is usually considered in the cited studies. In the first studies (Polyak &
 498 Smirnov, 1968; Hamza & Verma, 1969), a limited number of regional data was used,
 499 which provided relatively good relationships. In the later studies (Vitorello et al.,
 500 1980; Sclater et al., 1981), a larger dispersion of heat flow appeared, but still was
 501 interpreted by these authors as some thermal relaxation of the lithosphere with a
 502 time scale of ~ 0.5 Ga. Sclater et al. (1981) suggested a constant average heat flow
 503 of 46 mWm^{-2} for Precambrian, increasing to 77 mWm^{-2} at the end of Proterozoic.
 504 Vitorello et al. (1980) interpreted the heat flow / age relationship as the superposition
 505 of a background heat flow ($\sim 27 \text{ mWm}^{-2}$) at the base of an enriched radioactive upper
 506 crust, a radiogenic crustal component that would represent 40 % of the surface heat
 507 flow and a transient component with a time scale of ~ 500 Ma.

508 However, Rao et al. (1982) challenged the physical and statistical relevance of
 509 the heat flow / age relationship, arguing that, unlike oceanic lithosphere, no common
 510 process affects the thermal regime of the continental lithosphere and that other factors
 511 such as variations of the radiogenic content in the upper crust have much more effect
 512 than the age on the surface heat flow. It became evident later that the variations of
 513 surface heat flow in stable areas are mostly related to the variations of upper crustal
 514 heat production (Jaupart et al., 2014). As the number of continental measurements
 515 has significantly increased since these analysis, I examined again the continental heat
 516 flow / age relationship. Instead of using stratigraphy as in table 1, I defined the
 517 tectono thermal age from the USGS (1997) map, which distinguishes six categories:
 518 Archean (>2500 Ma), early Proterozoic (1600-2500 Ma), middle Proterozoic (900-1600
 519 Ma), late Proterozoic (570-900 Ma), Paleozoic (245-570 Ma) and Mesozoic-Cenozoic
 520 (0-245 Ma). I also subdivided continental surface into homogeneous units (e.g. Baltic
 521 shield, Williston basin, Sino-Korean craton...) and plotted them in several groups
 522 corresponding to the main landmasses (e.g. Europe, Africa, Central Asia...) in figure
 523 5, where some previous interpretations (Vitorello et al., 1980; Sclater et al., 1981) are
 524 also shown. It appears clearly that the dispersion of heat flow values is important
 525 whatever the age and is more important than previously estimated. This dispersion
 526 may be related to the inaccurate age domains of the USGS (1997) map (e.g. part of the
 527 Tertiary basin Aaiun-Tarfaya in Western Africa is located in Archean) or to the variable
 528 radiogenic contribution of the upper crust, such as the Precambrian regions of Aravalli
 529 and Singbhum in India (Roy & Rao, 2000), Slave in Canada (Perry et al., 2006) or
 530 South central Australia (Neumann et al., 2000). In order to avoid influence of uncertain
 531 age estimates, I cross checked the USGS (1997) tectono-thermal age map with the
 532 UNESCO stratigraphic map (CCGM/CGMW, 2000) and plot data belonging to the
 533 same era in both maps (figure 5): dispersion was significantly reduced for Archean,
 534 but remains important for Proterozoic and Paleozoic, as previously mentioned by
 535 Rao et al. (1982). Therefore, extrapolation of the continental heat flow based on the
 536 stratigraphic age (Davies & Davies, 2010) or the last tectono thermal age (Chapman

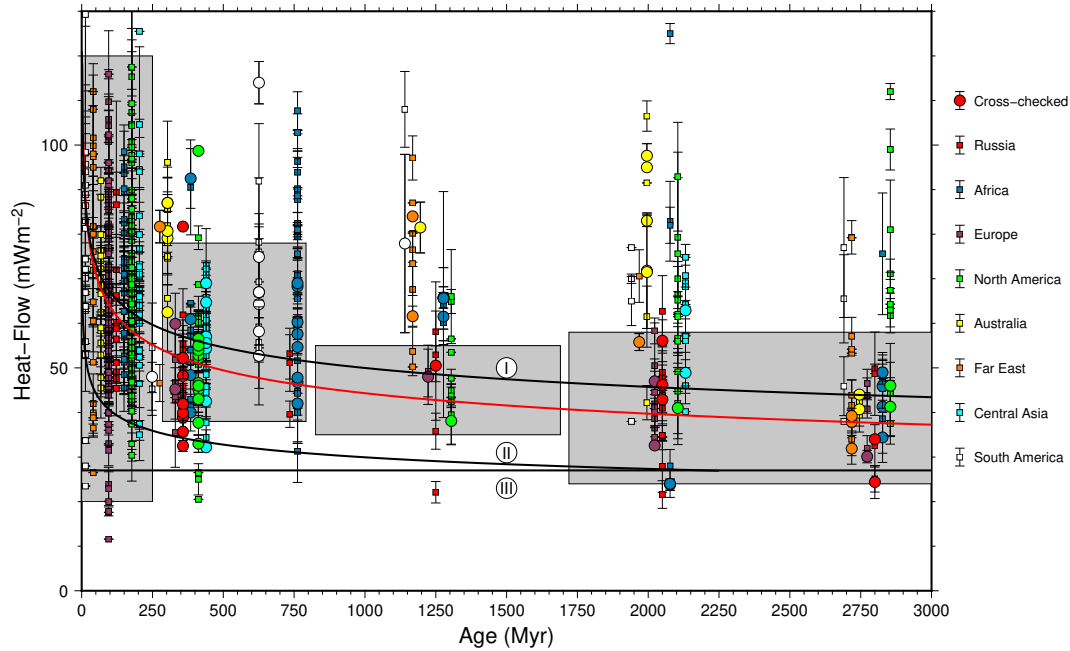


Figure 5. Continental heat flow vs last tectono thermal age. Heat flow is averaged for homogeneous geological units and according to the tectono thermal ages from a USGS map. Tectonothermal age is divided in six categories: Archean (>2500 Ma), early Proterozoic (1600-2500 Ma), middle Proterozoic (900-1600 Ma), late Proterozoic (570-900 Ma), Paleozoic (245-570 Ma) and Mesozoic-Cenozoic (0-245 Ma). For each of them, heat flow / age is displayed at the middle of the age interval with a color symbol corresponding to a specific landmass and eventually offset laterally for clarity (Russia is on the central position). The small square symbols are locations where the tectono-thermal age from USGS map is not consistent with the stratigraphic age (CCGM/CGMW, 2000). The larger circle symbols are conversely locations where they are similar. The red line corresponds to the best fit power law relationship : $HF = 132/Age^{0.158}$ for the Russian landmass. Grey boxes correspond to heat flow age domains defined by Slater et al. (1981). Black lines correspond to the trends proposed by (Vitorello et al., 1980): (I) is surface heat flow, (II) is the reduced heat flow (surface heat flow minus upper crust radiogenic contribution) that includes a transient component and (III) is background heat flow $\sim 27 \text{ mWm}^{-2}$ that includes $\sim 12 \text{ mWm}^{-2}$ mantle heat flow and $\sim 15 \text{ mWm}^{-2}$ lower crust radiogenic contribution.

537 & Pollack, 1975) may have limited utility as noted by the authors, and I will discuss
 538 a method to improve mapping in the next part of this paper.

539 In contrast to continents, divergent continental margins have a well defined thermal
 540 origin (rifting or break-up, magmatism,...) and a less important radiogenic contribution
 541 related to crustal thinning. Therefore, heat flow could be more likely linked to
 542 age in these specific domains. As the age of continental break-up is hard to obtain
 543 for all locations, the age of the closest oceanic domain was used systematically, except
 544 for Red Sea where it is not defined in the Müller et al. (2008) data.

545 Heat flow, corrected systematically for sedimentation effects with global datasets
 546 (Divins, 2003; Whittaker et al., 2013), was selected only in the deep continental domains
 547 of the margin including the Ocean Continent Transition (OCT), and was averaged by basin
 548 and country domains. The resulting relationship between heat flow and age (figure 6)
 549 is fairly compatible with an oceanic type conductive cooling. This suggests that continental
 550 margins evolve toward a different state after thermal relaxation than before rifting, when
 551 the mantle heat flow had a lower value. This could be an important aspect to consider
 552 when modelling petroleum systems.

553 Heat flow in other types of continental margins is not as simply related to the age.
 554 In transform margins, age is more difficult to define and the age of the closest oceanic
 555 domain is probably not relevant. Convergent margins are also more complex as the age
 556 cannot be simply defined as the age of the subducting plate. Consequently, the relation
 557 between heat flow and age cannot be established clearly for these types of continental
 558 margins.

559 4 Prediction of heat flow using a similarity method

560 Most of the previous studies (Chapman & Pollack, 1975; Sclater et al., 1980; Vitorello
 561 & Pollack, 1980; Sclater et al., 1981; Pollack et al., 1993; Davies & Davies, 2010;
 562 Davies, 2013; Jaupart et al., 2015) have used age or tectono-thermal age to extrapolate
 563 heat flow where no data exists. Other studies related surface heat flow to temperature
 564 proxies in the lithosphere, such as the upper mantle shear-wave velocities (Shapiro &
 565 Ritzwoller, 2004) or the Curie point depth (Artemieva, 2006; Li et al., 2017). As all
 566 of these approaches capture only part of the processes affecting surface heat flow,
 567 Goutorbe et al. (2011) suggested to stack these different proxies in order to get
 568 redundant information better constraining the heat flow predictions. They tested
 569 different methods of extrapolation and found that the similarity method was optimal,
 570 both in terms of prediction ability and computation time. The method is based on the
 571 evaluation, at each location of a grid ($0.5^\circ \times 0.5^\circ$ in the present study), of the number
 572 of similarities N_{sim} with several heat flow proxies (observables) at all other locations
 573 of the grid where surface heat flow is known. Two cells are similar if they belong
 574 to the same class (e.g. Paleozoic), the same range (e.g. elevation between 1000 and
 575 1200 m) or the same class of distance to a specific geological feature below a threshold
 576 value (e.g. less than 500 km to a subduction). Each additional similarity increases
 577 the weight w_i in the collection of heat flows by a factor K , which optimal value was
 578 estimated to $\simeq 10$ by Goutorbe et al. (2011):

$$w_i = K^{N_{sim}} \quad (10)$$

579 Classical statistics (mean \bar{Q} and standard deviation σ_Q) provide estimates of the
 580 average value \bar{Q} and its uncertainty σ_Q from the collection of heat flow:

$$\bar{Q} = \frac{\sum_i w_i Q_i}{\sum_i w_i} \quad (11)$$

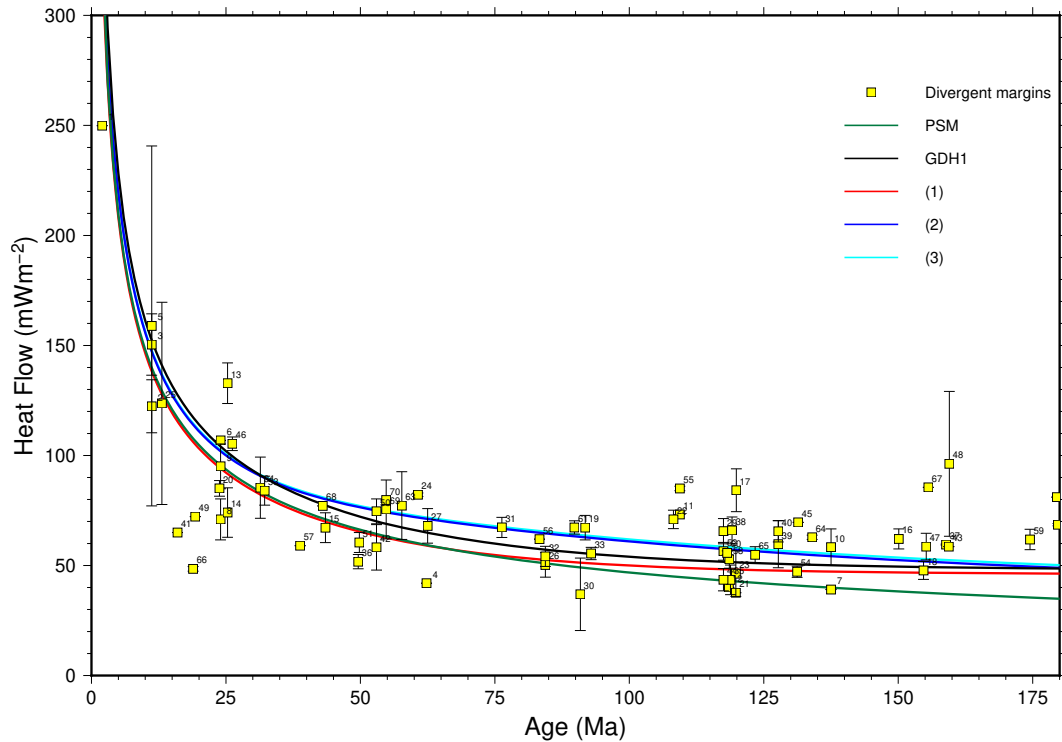


Figure 6. Heat flow vs age on divergent continental margins (red squares = active margins; blue squares = passive margins). Lines correspond to oceanic models described in figure 2. Heat flow is averaged for country subsets defined by numbers: 1- Red Sea Basin (Ethiopia), 2- Red Sea Basin (Eritrea), 3- Red Sea Basin (Egypt), 4- Burgas Basin (Bulgaria), 5- Central Red Sea Rift System (Egypt), 6- Djefara Basin (Libya), 7- Guadalquivir Basin (Portugal), 8- Gulf of Suez Basin (Egypt), 9- Pelagian Basin (Italy), 10- Rharb-Prerif Basin (Portugal), 11- Amazon Cone (Brazil), 12- Angola Basin (Angola), 13- Balingian Province (Malaysia), 14- Baram Delta (Brunei), 15- Bass Basin (Australia), 16- Browse Basin (Australia), 17- Campos Basin (Brazil), 18- Canning Basin (Australia), 19- Cantabrian Zone (Spain), 20- Catalano-Balearic Basin (Spain), 21- Cauvery Basin (India), 22- Celtic Sea Graben System (United Kingdom), 23- Congo Fan (Angola), 24- Djefara Basin (Tunisia), 25- East Greenland Basin (Greenland), 26- East Newfoundland Basin (Canada), 27- East Siberian Sea Basin (Russia), 28- Essaouira Basin (Morocco), 29- Gabon Coastal Basin (Gabon), 30- Galician Basin (Spain), 31- Gippsland Basin (Australia), 32- Grand Banks Basin (Canada), 33- Great Australian Bight (Australia), 34- Gulf of Aden (Oman), 35- Kwanza Basin (Angola), 36- Labrador Shelf (Canada), 37- Lamu Embayment (Somalia), 38- Lower Congo Basin (Gabon), 39- Lusitanian Basin (Portugal), 40- Mackenzie Delta (Canada), 41- Money Shoal Basin (Australia), 42- More Basin (Norway), 43- Mozambique Basin (Mozambique), 44- Namibe Basin (Angola), 45- Natal Trough (South Africa), 46- North Balearic-Ligurian Basin (Spain), 47- North Carnarvon Basin (Australia), 48- North Chukchi Basin (Russia), 49- Northwest Sabah Province (Philippines), 50- Norwegian-Danish Basin (Denmark), 51- Otway Basin (Australia), 52- Outeniqua Basin (South Africa), 53- Pearl River Mouth (Zhujiangkou) Basin (Taiwan), 54- Perth Basin (Australia), 55- Potiguar Basin (Brazil), 56- Rockall Trough (Ireland), 57- Sagaleh Basin (Somalia), 58- Santos Basin (Brazil), 59- Senegal Basin (Senegal), 60- Sergipe-Alagoas Basin (Brazil), 61- Sirte Basin (Libya), 62- Souss Trough (Morocco), 63- South Barents Sea Basin (Russia), 64- South Carnarvon Basin (Australia), 65- Southwest African Coastal Basin (Namibia), 66- Southwest Palawan Basin (Philippines), 67- Tanzania Coastal Basin (Tanzania), 68- Troms-Finmark Fault Complex (Norway), 69- Voring Basin (Norway), 70- Voring Plateau (Norway).

$$\sigma_Q = \sqrt{\frac{\sum w_i (Q_i - \bar{Q})^2}{\sum w_i}} \quad (12)$$

581 An example of collection of heat flow at some location of the grid is given in
 582 figure S9 of the supporting information. More details on the method can be found in
 583 Goutorbe et al. (2011). The method was also briefly discussed in Sclater et al. (2014).

584 The method was used here with NGHF, and I proceeded in several steps to
 585 explore how increasing number of observables changes the distribution of heat flow at
 586 the Earth surface. The predictability of the similarity method depends strongly on
 587 the number and on the quality of observables. The greater the number of observables,
 588 the smaller the misfit between measurements q_{obs} and predicted heat flow q_{pred} :

$$misfit = \sqrt{\frac{\sum (q_{obs} - q_{pred})^2}{N_{obs}}} \quad (13)$$

589 The misfit varies from $\simeq 90 \text{ mWm}^{-2}$ for 1 observable to $\simeq 4 \text{ mWm}^{-2}$ for 50 observ-
 590 ables. The combination of different observables affects the convergence of the process
 591 (Figure 7). In this figure, I show two examples that behave differently: the first set
 592 was optimized for a faster convergence, while the second set was chosen randomly. The
 593 first set evolves more or less as a function of $1/N$ (Figure 7), and after a rapid decrease
 594 of the RMS, there is no significant improvement when more observables are added.
 595 There is no systematic effect of one specific observable, as it can have a significant
 596 influence with some combination and no influence with another one. Therefore, the
 597 optimal combination was obtained by tries and errors.

598 A great number of observables can produce artifacts where no heat flow measure-
 599 ment exists, especially if the quality of these observables is not good. Therefore,
 600 there is a trade-off between an acceptable misfit and a more likely predictability. For
 601 the specific sets used in this study, the best models are obtained accordingly for 14
 602 observables (first set) and 19 observables (second set), assuming an acceptable RMS
 603 less than 6 mWm^{-2} (figure 7). The 14 observables used for the first set are displayed
 604 with their direct surface heat flow prediction in figures of the supporting information.

605 I also provide statistics (table 3) comparing usual extrapolation from heat flow
 606 measurements only, with values predicted by the similarity method. As shown pre-
 607 viously, oceanic heat flow measurements were improved in the 90s and therefore I
 608 selected only measurements published after 1990. I also included geothermal and vol-
 609 canic areas in the averaged values, which differs from the statistics in the first section
 610 of this paper. The differences between results of table 1 and 3 are therefore related to
 611 a slightly different processing of the data. Where several measurements exist within a
 612 specific cell of the grid, the arithmetic average was given to this cell: all cells with heat
 613 flow measurements are explored to determine if similarities exist with other cells with
 614 no measurements. The number of oceanic cells covered by measurements is 1672 (only
 615 data after 1990) and the number of continental cells 10914. Heat flow values for these
 616 cells range between -6 and 2777 mWm^{-2} , with an average of $89 \pm 128 \text{ mWm}^{-2}$ in the
 617 oceanic domain and $67 \pm 31 \text{ mWm}^{-2}$ for continents. The global heat loss estimated
 618 only from measurements is 40.8 TW (26.7 TW for oceans and 14.1 TW for continents).

619 Four predictions are discussed here (table 3 and figure 8).

620 The first prediction is based on 2 observables, the age of the oceanic seafloor
 621 (Müller et al., 2008) and 12 tomographic classes defined by Goutorbe et al. (2011)
 622 from the shear wave velocity model of Shapiro and Ritzwoller (2002). As the possible
 623 number of corresponding classes between 2 cells is limited, the heat flow map (figure

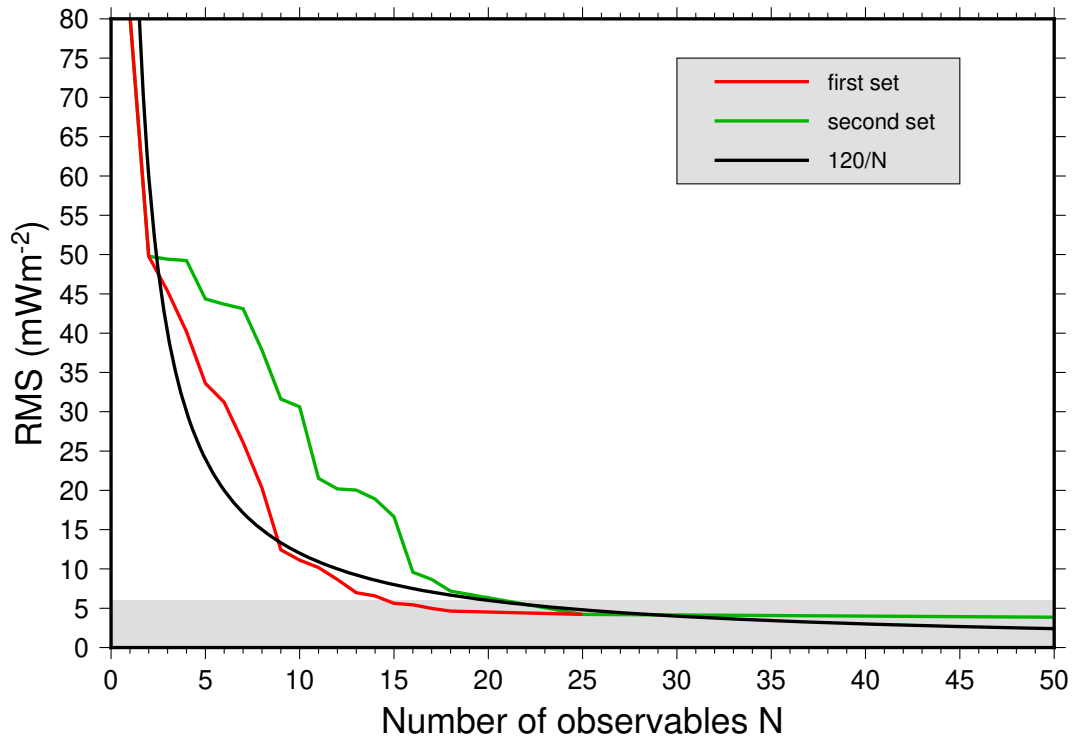


Figure 7. Evolution of the predictability (RMS) of the similarity method as a function of the number of observables. The grey area corresponds to RMS less than 6 %.

624 8a) shows smooth variations and the misfit between predicted and measured heat flow
 625 (50 mWm^{-2}) is the largest (table 3).

626 The second prediction is based on 14 observables, including age (Müller et al.,
 627 2008; CCGM/CGMW, 2000), seismic tomography (Shapiro & Ritzwoller, 2002), litho-
 628 spheric thickness (Conrad & Lithgow-Bertelloni, 2006), Curie point depth (Li et al.,
 629 2017), free air anomalies (Sandwell & Smith, 2009), topography (Kautz, 2017) and dis-
 630 tance to several tectonic features (rifts, orogens, active seismic zones, volcanic zones,
 631 Archean cores). All of them, and their potential effect on predicted heat flow, are dis-
 632 played in figure S10 to S23 of the supporting information. The heat flow map (figure
 633 8b) shows a better resolution than with 2 observables, revealing smaller scale anom-
 634 alies. The misfit between observed and predicted heat flow is noticeably reduced to 7
 635 mWm^{-2} (table 3).

636 Other predictions (25 and 50 observables) are not significantly improved (misfit
 637 is 7 and 4 mWm^{-2}), and increase the risk of generating noise elsewhere. For instance,
 638 the predicted heat flow map with 50 observables (figure 8c) shows a large variability
 639 in regions with no measurements such as off-axis oceanic domains. On the other hand,
 640 this could also suggest locations for new investigations.

641 A test of the prediction capability of the method was also made on Africa. All
 642 the terrestrial heat flow measurements in Africa were removed before the test and
 643 compared to the results obtained with these measurements included. In both cases,
 644 14 observables (preferred prediction) were used. The resulting statistical differences
 645 between the two prediction are insignificant ($64.6 \pm 17.2 \text{ mWm}^{-2}$ for the reference and
 646 $65.1 \pm 15.9 \text{ mWm}^{-2}$ for the test). The maps of African heat flow (figure 9) show some
 647 differences in a few specific regions, like the Saharian basins (Tindouf, Illizi, Ahnet, ...)

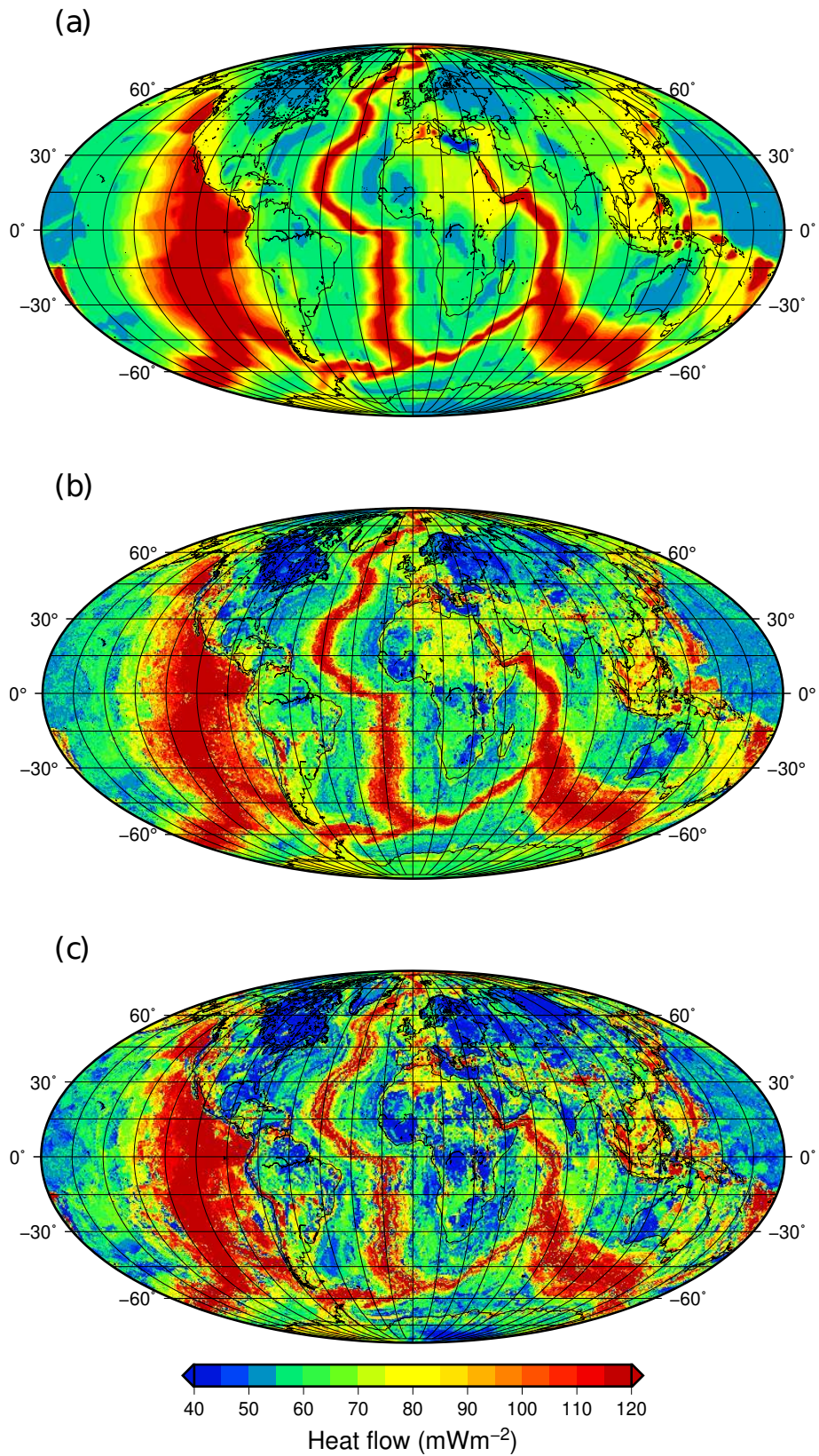


Figure 8. Global heat flow map based on similarities with (a) 2 observables; (b) 14 observables (preferred prediction); (c) 50 observables.

648 or the East African rift in Ethiopia where the high heat flow anomaly is not predicted
649 without the heat flow measurements constraints. However, most of the large scale
650 features of the reference are described satisfactorily in the test.

651 A comparison of statistical averages for stratigraphic or geographic subsets (table
652 3) shows that the average of predicted values is generally in agreement with the average
653 of observed values. There are a few exceptions, mostly for the young seafloor (0-5 Ma)
654 and some geographic subsets (Central America, Greenland, Antarctica) where there
655 exist few and non representative measurements. Large numbers of observables increase
656 the variability of oceanic heat flow near the ridge axis, which, for instance, becomes
657 larger than measurements in the range 1.6-5.3 Ma of seafloor ages from 1.6-5.3 Ma.

658 The predicted heat flow is stable with the number of observables for continents,
659 while it increases for the oceans, mostly in the range 0-1.6 Ma where the variability of
660 hydrothermal circulation is the most important. The total heat loss predicted with 14
661 observables is similar to that extrapolated from heat flow measurements, and increases
662 with a larger number of observables as a consequence of the oceanic heat flow increase
663 in the range 0-1.6 Ma.

664 In map view, the predicted heat flow (figure 8) outlines some remarkable trends.

665 In oceanic domains, map calculated with the largest numbers of observables
666 shows large heterogeneities. This is especially important for the Mid Atlantic Ridge
667 and the South East Indian ridge. This is of course not prescribed but related to the
668 dispersion of values caused by hydrothermal circulation and propagated to similar
669 locations.

670 The divergent continental margins are clearly marked by a higher heat flow than
671 their adjacent ocean and continent. This anomaly is particularly well developed on the
672 Atlantic margins (Brazil, Africa) and interestingly on the Antarctica margins (figure
673 8c) well constrained by measurements (Nagao et al., 2002; Morin et al., 2010). Several
674 studies have attributed the anomaly on mature continental margins to an edge driven
675 convection caused by lateral temperature and rheology contrast at the interface of
676 oceanic and continental lithospheres (King & Anderson, 1998; Lucazeau et al., 2004;
677 Hardebol et al., 2012).

678 In continental domains, North America, Europe and Russia are characterized by
679 large zones of low heat flow associated with shields, while in South America, Africa
680 and Australia, these zones are more fragmented. In Antarctica, a similar large domain
681 of low heat flow ($40\text{-}50\text{ mWm}^{-2}$) is predicted with 50 observables in the central eastern
682 part (figure 8c), as previously suggested by other authors (Shapiro & Ritzwoller, 2004;
683 Lieffering & Pattyn, 2013) but is weaker with 14 observables. In Greenland, some
684 new heat flow data and the occurrence of thermal springs were associated to a heat
685 flow anomaly the east-northeast area (Rogozhina et al., 2016; Rysgaard et al., 2018),
686 but the similarity method contains observables that hardly support the existence of
687 this anomaly, which appears only with 50 observables. In Central and South America,
688 the predicted heat flow is much higher than the extrapolation of heat flow measured in
689 this area. This is related to the presence of volcanic and back arcs, which correspond
690 generally to high heat flow like in the Andes or Cascade range in North America. Far
691 East margins are also associated with high heat flow of similar origin.

692 Conclusions

693 The objective of this paper was to reexamine the main trends of heat flow at the
694 Earth surface with a significantly upgraded database. An important result is that the
695 previous studies were statistically robust, and the previous estimates of the global heat
696 budget are not strongly modified. However, new oceanic heat flow measurements are

Table 3. Comparison of heat flow averages with predictions obtained from the similarity method with 2, 14, 25 and 50 proxies. The number of proxies clearly decreases the misfit with observed heat flow, while the different category averages do not vary significantly.

zone	area (10^6 km^2)	meas. Ncells	total Ncells	Heat Flow (mWm^{-2})			Heat Loss (TW)					
				2 obs	14 obs	25 obs	50 obs	2 obs	14 obs	25 obs	50 obs	
Oceans												
0.0-1.6 Ma	4.2	29	1819	249 ± 37	722 ± 524	810 ± 618	986 ± 876	2.5	1.0	3.0	3.4	4.1
1.6-5.3 Ma	11.5	116	4921	220 ± 34	214 ± 114	209 ± 125	227 ± 196	1.6	2.5	2.5	2.4	2.6
5.3-23.7 Ma	56.4	414	24663	129 ± 32	120 ± 28	125 ± 37	118 ± 45	6.4	7.3	6.8	7.1	6.7
23.7-36.6 Ma	36.5	129	16572	94 ± 6	95 ± 16	91 ± 17	94 ± 24	3.5	3.5	3.5	3.3	3.4
36.6-57.8 Ma	46.7	123	23552	73 ± 8	74 ± 15	74 ± 17	76 ± 19	3.2	3.4	3.4	3.5	3.5
57.8-66.4 Ma	16.5	51	8274	64 ± 4	65 ± 11	64 ± 10	65 ± 14	1.0	1.1	1.1	1.1	1.1
66.4-84.0 Ma	31.2	131	13717	59 ± 3	58 ± 8	59 ± 7	58 ± 11	1.8	1.8	1.8	1.8	1.8
84.0-119.0 Ma	50.3	308	20032	58 ± 2	58 ± 6	57 ± 7	58 ± 9	2.9	2.9	2.9	2.9	2.9
119.0-144.0 Ma	28.1	140	14701	53 ± 2	53 ± 5	53 ± 6	53 ± 8	1.5	1.5	1.5	1.5	1.5
144.0-178.0 Ma	17.7	165	6704	52 ± 2	53 ± 5	52 ± 5	53 ± 8	0.9	0.9	0.9	0.9	0.9
178.0-299.0 Ma	0.8	66	304	40 ± 10	42 ± 26	42 ± 27	41 ± 27	0.0	0.0	0.0	0.0	0.0
average	299.9	1672	135259	86 ± 45	91 ± 105	92 ± 120	95 ± 156	26.7	25.8	27.2	27.7	28.4
Continents												
Archean	4.7	139	2746	55 ± 7	46 ± 12	49 ± 13	42 ± 14	0.2	0.3	0.2	0.2	0.2
Proterozoic	22.4	643	9296	61 ± 8	58 ± 16	56 ± 17	53 ± 22	1.3	1.4	1.3	1.3	1.2
Paleozoic	29.3	1855	15338	62 ± 9	60 ± 18	60 ± 21	58 ± 25	1.8	1.8	1.7	1.8	1.7
Mesozoic	21.9	1897	10835	65 ± 10	64 ± 17	64 ± 20	61 ± 24	1.4	1.4	1.4	1.4	1.3
Cenozoic	26.1	1863	11441	66 ± 10	66 ± 18	65 ± 22	63 ± 27	1.8	1.7	1.7	1.7	1.6
Quaternary	35.6	1787	13813	67 ± 9	70 ± 19	69 ± 22	68 ± 27	2.5	2.4	2.5	2.5	2.4
contin. margins	63.3	2904	36065	69 ± 10	72 ± 16	71 ± 20	70 ± 28	4.5	4.4	4.5	4.5	4.5
average	211.3	10914	123941	65 ± 10	66 ± 17	66 ± 20	64 ± 25	14.1	13.7	14.0	14.0	13.5
Landmass												
Africa	39.2	1033	13485	66 ± 14	65 ± 17	63 ± 24	60 ± 27	2.7	2.6	2.5	2.5	2.4
Antarctica	22.6	137	34356	64 ± 14	69 ± 11	71 ± 14	66 ± 21	2.0	1.4	1.6	1.6	1.5
Australia	19.6	434	7580	67 ± 16	66 ± 17	68 ± 18	67 ± 22	1.3	1.3	1.3	1.3	1.3
Central America	6.7	304	2314	103 ± 48	112 ± 134	118 ± 153	123 ± 202	0.5	0.7	0.7	0.8	0.8
Central Asia	16.9	855	7148	68 ± 8	64 ± 18	63 ± 21	62 ± 25	1.1	1.1	1.1	1.1	1.0
Europe	13.3	1929	6954	66 ± 21	67 ± 37	67 ± 38	66 ± 41	0.9	0.9	0.9	0.9	0.9
Far East	17.3	813	6119	69 ± 14	70 ± 23	71 ± 25	71 ± 29	1.4	1.2	1.2	1.2	1.2
Greenland	4.4	44	5625	69 ± 19	71 ± 24	72 ± 47	69 ± 44	0.4	0.3	0.3	0.3	0.3
Middle East	7.9	261	2917	77 ± 22	76 ± 29	73 ± 38	72 ± 44	0.5	0.6	0.6	0.6	0.6
North America	33.7	3214	20626	63 ± 20	65 ± 73	64 ± 67	64 ± 88	2.3	2.1	2.2	2.2	2.2
Russia	24.5	1509	19587	64 ± 13	62 ± 18	62 ± 28	61 ± 35	1.4	1.6	1.5	1.5	1.5
South America	30.0	646	10788	76 ± 35	83 ± 78	84 ± 86	86 ± 122	2.3	2.3	2.5	2.5	2.6
<i>misfit</i> (mWm^{-2})				50	7	4	4					

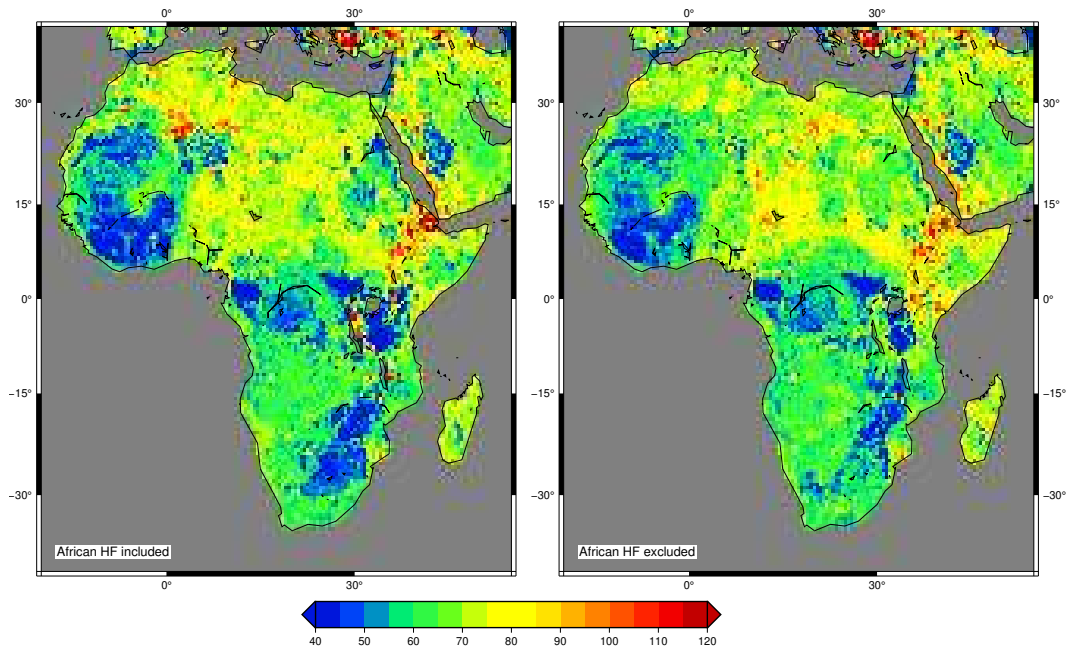


Figure 9. Comparison of heat flow predictions in Africa with (left part) or without (right part) heat flow measurements, and for 25 observables. The statistical differences for Africa are negligible ($64.6 \pm 17.2 \text{ mWm}^{-2}$ and $65.1 \pm 15.9 \text{ mWm}^{-2}$ with and without African HF data), but few differences in northern Africa (Sahara) and east Africa (Tanzania, Somalia) are noticeable on the maps.

697 significantly different from previous ones because they often involve multi-penetration
 698 techniques or new techniques that allowed investigations in virgin domains like mid
 699 ocean ridges axis. The consequence is that the statistical difference with conductive
 700 cooling models decreases in average for age younger than 65 Ma, but with a dispersion
 701 that is still important or even larger than in the Stein and Stein (1992) study. As this
 702 difference can no longer be used to estimate hydrothermal heat flow, I suggested two
 703 alternative methods that lead to a supplementary contribution of 1.5 TW accounting
 704 for the hydrothermal circulation effect in the old seafloor (>65 Ma).

705 The seafloor subsidence derived from the global binning of bathymetry is not con-
 706 sistent with that obtained at heat flow sites only. The origin of this difference is related
 707 to an anomalous subsidence in regions of hot spots and large igneous provinces, causing
 708 a flattening at global scale, and where heat flow measurements are few. Therefore, at
 709 heat flow sites, previous cooling models cannot simultaneously explain subsidence and
 710 heat flow for old oceans. This problem has already been raised by Nagihara et al.
 711 (1996) who studied subsidence and heat flow for a selection of nine locations with a
 712 basement older than 100 Ma and corrected for sediment and crustal thickness varia-
 713 tions. The observed heat flow at these locations appeared higher than predicted by the
 714 PS model, which conversely agreed very well with subsidence. The authors interpreted
 715 this paradox by reheating of the old lithosphere related to a more likely occurrence of
 716 thermal interactions with hot spots, but this should also affect subsidence. I suggest
 717 alternatively that this is the normal evolution of oceanic lithosphere for old age of the
 718 seafloor and explored if some model could explain the two sets of data simultaneously.
 719 The optimal model is obtained for a constant heat flow ($\sim 36.4 \text{ mWm}^{-2}$) at the base
 720 of a 114 km thick lithosphere and a moderate potential temperature of the mantle (\sim
 721 $1235 \text{ }^\circ\text{C}$).

In continents, the heat flow cannot be statistically related to the last tectono-thermal age or stratigraphic age as suggested frequently in past studies. There is a large variability that prevents any significant trend. Only Archean provinces (46 mWm^{-2}) are significantly lower and Quaternary domains significantly higher (70 mWm^{-2}) than other continental regions for ages spanning from early Proterozoic to Cenozoic ($60\text{--}65 \text{ mWm}^{-2}$). On the other hand, divergent continental margins have a heat flow dependence on age of the continental break-up and behave more like an oceanic domain.

The variability of heat flow independently of the age is not a real problem for global energy estimates, but makes a good mapping of surface heat flow difficult. I used a multi-proxies approach to derive a new global map at the scale of 0.5° square, which fairly reproduces the known local geological features and provides slightly modified statistics. For instance, Greenland and Antarctica would be notably cooler than existing heat flow measurements suggest. The number of observables decreases the misfit with observed heat flow, more or less as a function $1/N$ of the number of observables. The preferred prediction is obtained with 14 observables, for which RMS at heat flow locations is only 6 %. Heat flow anomalies are generated locally where no measurement exists. They could be real or artifacts, but could serve as potential targets for new studies. One especially interesting anomaly arising from this analysis is the quasi systematic higher heat flow surrounding continents that could attest of edge driven convection.

When observations only are considered, the global heat loss from this approach is 40.8 TW (14.1 TW for continents and 26.7 TW for oceans). By forcing oceanic heat flow with a conductive model fitting subsidence and heat flow, the total heat loss is 44 TW, similar to recent estimates (Davies & Davies, 2010; Sclater et al., 2014; Jaupart et al., 2015). Oceanic hydrothermal heat loss would represent 10.5 TW, i.e. one fourth of the total value.

Acronyms

BSR Bottom Simulating Reflector

CHABLIS Constant Heat flow At the Base of the LithoSphere

GDH1 Global Depth and Heat flow model (Stein & Stein, 1992)

GLOBHEAT Global Heat Flow dataset published by Pollack et al. (1993)

GTOPO30 Global topography/bathymetry grid from a wide variety of data sources

IHFC International Heat Flow Commission

NGHF New Global Heat Flow

PS Parsons and Sclater (1977) Model

Acknowledgments

All figures were drawn with GMT 5.4.5 (Wessel et al., 2013). The manuscript benefited from a first reading and discussions with Frédérique Rolandone and Valérie Vidal, and from constructive reviews by Derrick Hasterok, J. Huw Davies and John G. Sclater. Supporting information contains additional tables and figures, separate csv files including the NGHF dataset (NGHF.csv) in the same format as the previous GLOBHEAT file (Pollack et al., 1993) and the preferred heat flow prediction map (HFgrid14.csv). This is IGP contribution # 4051.

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