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1 **ASUCA: The JMA Operational**
2 **Non-hydrostatic Model**

3 **ISHIDA Junichi**

4 *Numerical Prediction Division, Japan Meteorological Agency,*
5 *Tokyo, Japan*

6 **ARANAMI Kohei**

7 *Office of Marine Prediction, Japan Meteorological Agency,*
8 *Tokyo, Japan*

9 **KAWANO Kohei**

10 *Numerical Prediction Development Center, Japan*
11 *Meteorological Agency, Tsukuba, Japan*

12 **MATSUBAYASHI Kengo**

13 *Numerical Prediction Development Center, Japan*
14 *Meteorological Agency, Tsukuba, Japan*

15

KITAMURA Yuji

16

Numerical Prediction Development Center, Japan

17

Meteorological Agency, Tsukuba, Japan

18

MUROI Chiashi

19

Planning Division, Japan Meteorological Agency, Tokyo,

20

Japan

21

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Corresponding author: ISHIDA Junichi, Numerical Prediction Division, Japan
Meteorological Agency, 3-6-9 Toranomom, Minato City, Tokyo, 105-8431, Japan.
E-mail: j-ishida@met.kishou.go.jp

Abstract

22
23 The non-hydrostatic numerical weather prediction (NWP) model ASUCA
24 developed by the Japan Meteorological Agency (JMA) was launched into
25 operation as 2 and 5 km-resolution regional models in 2015 and 2017, re-
26 spectively. This paper outlines specifications of ASUCA with focus on the
27 dynamical core and its configuration/accuracy as an operational model.
28 ASUCA is designed for high computational stability and efficiency, mass
29 conservation and forecast accuracy. High computational stability is achieved
30 via a time-split integration scheme to compute acoustic terms and an advec-
31 tion scheme with a flux-limiter function to avoid numerical oscillation. In
32 addition, vertical advection and sedimentation are calculated together with
33 another exclusive time-splitting technique. ASUCA adopts hybrid paral-
34 lelization using Message Passing Interface (MPI) and Open Multi Process-
35 ing (OpenMP) for high computational efficiency on massive parallel scalar
36 computers. The three-dimensional arrays are allocated such that the ver-
37 tical direction is the stride-one innermost dimension to make effective use
38 of cache and multi-thread parallelization. This is particularly advantageous
39 for physical processes evaluated in a vertical column. To ensure mass con-
40 servation, density rather than pressure is integrated as a prognostic variable
41 in flux-form fully compressible governing equations. ASUCA exhibited bet-
42 ter performance than the previous operational model in idealized and NWP

43 tests.

44 1. Introduction

45 Numerical weather prediction (NWP) models form the technical foun-
46 dation of weather forecasting by the Japan Meteorological Agency (JMA);
47 their precision directly affects the accuracy of weather advisories/warnings
48 and various other types of weather information. As weather-related disas-
49 ters in Japan have become more intensified in recent times, optimization of
50 prediction accuracy is an important factor in disaster mitigation. Against
51 such a background, stable operation and sustainable development of JMA's
52 operational NWP model are vital.

53 JMA has operated regional mesoscale NWP models with 2 and 5 km hor-
54 izontal resolutions since 2012 and 2006, respectively, for purposes including
55 mitigation of disasters caused by torrential rain. The Agency began re-
56 gional mesoscale NWP model operation in 2001 with a hydrostatic model
57 featuring a horizontal resolution of 10 km. This was replaced in 2004 by
58 the non-hydrostatic model JMA-NHM (Saito et al. 2001, 2006), which was
59 initially developed for research and subsequently adopted for operational
60 use.

61 Though the JMA-NHM had been utilized extensively in research fields,
62 its operation highlighted various problems, such as difficulties in achieving
63 numerical stability, high-performance computing and sophistication of an
64 NWP system including data assimilation.

65 As the reliance on NWP products increases in weather forecasting, higher
66 numerical stability is required. To improve the computational stability, var-
67 ious methods (including artificial horizontal diffusion) have been introduced
68 to the JMA-NHM. However, the JMA-NHM has occasionally caused compu-
69 tational instability or produced artificial noise for various complex reasons.
70 The strength of the artificial diffusions applied to avoid numerical instabil-
71 ity have been set empirically due to a lack of any scientific basis for such
72 determination. Furthermore, sedimentation and vertical advection for wa-
73 ter substances are treated independently of each other in the JMA-NHM.
74 This treatment often affects the vertical distribution of the water substances
75 and numerical stability. Accordingly, the application of artificial numerical
76 diffusions in the JMA-NHM does not solve this problem; an overall recon-
77 struction of the dynamical core is required.

78 Another significant issue relates to the rapid progress of high-performance
79 computing. JMA upgrades its supercomputer system every five or six years
80 with an increased number of CPUs. The sixth-generation system (1996 –
81 2001) had only 4 CPU cores, while the seventh (2001 – 2006) and eighth
82 (2006 - 2012) had 640 and 2,560 cores respectively. The number of CPUs in
83 supercomputer systems is expected to maintain exponential growth, giving
84 rise to an urgent need for higher parallel computation efficiency.

85 However, the various efforts implemented to solve these problems com-

86 plicated the source code of the JMA-NHM, and eventually hindered further
87 development toward higher forecast accuracy. To promote ongoing NWP
88 model development, sophisticated code management was needed. Against
89 this background, JMA began development of the new non-hydrostatic dy-
90 namical core ASUCA, which is a recursive acronym of “ASUCA is a System
91 based on a Unified Concept for Atmosphere”, in 2007.

92 For high computational stability, the monotonicity-preserving advection
93 scheme proposed by Koren(1993) is utilized to avoid numerical oscillation,
94 and the third-order Runge-Kutta method (Wicker and Skamarock 2002) is
95 employed. Time splitting is also applied for vertical advection and falling
96 water substances to satisfy the Courant-Friedrichs-Lewy (CFL) condition.
97 These enable the exclusion of additional terms for computational stability
98 such as numerical diffusion and divergence damping (Skamarock and Klemp
99 1992). The terms for vertical advection and falling water substances are
100 calculated together, as independent treatment may cause unrealistic vertical
101 separation of water substances.

102 To ensure accurate mass conservation, density rather than pressure is
103 integrated as a prognostic variable in flux-form fully compressible equations
104 with the finite volume method. Horizontally split-explicit and vertically im-
105 plicit time integration method based on the conservative Split-Explicit Time
106 Integration Method (Klemp et al. 2007) is employed to control acoustic and

107 gravity waves.

108 Hybrid parallelization using Message Passing Interface (MPI) and Open
109 Multi Processing (OpenMP) is adopted for high computational efficiency on
110 massive parallel scalar computers. Computation, communication and disk-
111 I/O are overlapped as much as possible, and three-dimensional arrays are al-
112 located such that the vertical is the stride-one innermost dimension to make
113 effective use of cache, multi-thread parallelization and Single-Instruction
114 Multi-Data (SIMD) instructions. This design enables ASUCA to achieve
115 high scalability on current parallel supercomputer systems.

116 A modern software management system including source code review,
117 documentation, version control and project management tools is used to
118 improve code quality and ensure a scientific research basis. To promote
119 the development of physical process schemes which play key roles on NWP
120 performance, physical process schemes are implemented via an independent
121 library of the ASUCA dynamical core. Here, physical process schemes are
122 designed as vertical one-dimensional models with unified coding and inter-
123 face rules to support development using single-column models. The data
124 assimilation system (Ikuta et al. 2021) and the forecast model are man-
125 aged with a unified source code tree to maintain consistency between the
126 4D-Var assimilation system and the forecast model. This system facilitates
127 the development and maintenance of source code quality. Intensive testing

128 and checking were performed in the development of the operational model
129 to avoid unexpected effects such as downgraded forecast accuracy.

130 ASUCA was found to perform better than the JMA-NHM, and replaced
131 it as the Local Forecast Model (LFM) with 2 km resolution in 2015 and as
132 the Meso Scale Model (MSM) with 5 km resolution in 2017. The ASUCA-
133 based Mesoscale Ensemble Prediction System (MEPS; Ono et al. 2021) and
134 the 4D-Var assimilation system (Ikuta et al. 2021) have been operated to
135 provide uncertainty information and initial fields for the MSM since 2019
136 and 2020, respectively.

137 This paper outlines specifications of ASUCA with focus on the dynam-
138 ical core and its configuration/accuracy as an operational model. Section
139 2 details the governing equations, and Section 3 describes discretization
140 including the treatment of advection and time integration along with the
141 derivation of the split-explicit method. Parallelization, which is essential
142 for future high performance computing (HPC), is detailed in Section 4, and
143 Section 5 presents simple specifications of physical process schemes used in
144 the operational system and their coupling with dynamics. Section 6 com-
145 pares the performance of ASUCA to that of the JMA-NHM in idealized and
146 realistic simulations, while Section 7 provides a summary and outlines the
147 future development plan of ASUCA.

2. Governing Equations

ASUCA involves the use of non-hydrostatic fully compressible governing equations written in flux form for mass conservation. Spherical curvilinear orthogonal and hybrid terrain-following coordinates with a shallow assumption are employed. To enable application of Lambert conformal map projection (as supported by the JMA-NHM and used in operational regional NWP), ASUCA employs generalized coordinates for flexible three-dimensional transformation. Map factors for conformal projection are incorporated into the transformation metric tensor. Derivation is described in JMA (2014).

ASUCA employs a total mass density ρ and a modified moist potential temperature θ_m , defined as

$$\begin{aligned}\rho &\equiv \rho_d + \rho_v + \rho_c + \rho_r + \rho_i + \rho_s + \rho_g, \\ \theta_m &\equiv \theta \left\{ 1 + \left(\frac{1 - \epsilon}{\epsilon} \right) q_v - q_c - q_r - q_i - q_s - q_g \right\}.\end{aligned}\tag{1}$$

The definition of θ_m is identical to that of Klemp et al.(2007) except in the incorporation of water substances. The subscripts d, v, c, r, i, s and g represent dry air, water vapor, cloud water, rain, cloud ice, snow and graupel, respectively. q_α is the ratio of the density of water substances α to the total mass density ($\alpha = v, c, r, i, s, g$). ϵ is the ratio of the gas constant for dry air R_d to the gas constant for water vapor R_v , and θ is the potential

166 temperature defined as

$$\theta \equiv \frac{T}{\pi}. \quad (2)$$

167 π is the Exner function defined as

$$\pi = \left(\frac{p}{p_0} \right)^{\frac{R_d}{C_p}}, \quad (3)$$

168 where p is the total pressure of moist air, p_0 is the reference pressure (typi-
169 cally 10^5 Pascals), and C_p is approximated by the specific heat capacity of
170 dry air at constant pressure.

171 The density ρ_b , which is the sum of dry air and water substances whose
172 terminal fall velocity is assumed to be zero, is described as

$$\rho_b = \sum_{\alpha \neq \text{sed}} \rho_\alpha = \rho \left(1 - \sum_{\alpha = \text{sed}} q_\alpha \right), \quad (4)$$

173 where “sed” denote the collection of water substances which are assumed
174 to have non-zero terminal velocity.

175 The Jacobian of transformation from the Cartesian coordinates (x, y, z)
176 to generalized coordinates (ξ, η, ζ) is defined as

$$J \equiv \begin{vmatrix} \xi_x & \xi_y & \xi_z \\ \eta_x & \eta_y & \eta_z \\ \zeta_x & \zeta_y & \zeta_z \end{vmatrix}, \quad (5)$$

177 where, $\xi_x \equiv \frac{\partial \xi}{\partial x} \Big|_{y,z}$ and the same description is applied to other metrics.
 178 A restriction is imposed on vertical coordinates to satisfy $\xi_z = \eta_z = 0$ as
 179 required for application of the split-explicit time integration scheme as seen
 180 in section 3.4.

181 Velocity components in generalized coordinates (U, V, W) are defined as

$$\begin{aligned} U &= \xi_x u + \xi_y v + \xi_z w \\ V &= \eta_x u + \eta_y v + \eta_z w \\ W &= \zeta_x u + \zeta_y v + \zeta_z w. \end{aligned} \quad (6)$$

Here, (u, v, w) represent velocity components in Cartesian coordinates. The terminal fall velocity of water substances α in generalized coordinates is defined as

$$W_{t\alpha} = \zeta_z w_{t\alpha}, \quad (7)$$

182 where $w_{t\alpha}$ is the terminal fall velocity in Cartesian coordinates.

183 The variables ρ , $\rho\theta_m$ and π are decomposed into the basic state and the
 184 related deviation as

$$\rho = \bar{\rho} + \rho', \quad \rho\theta_m = \overline{\rho\theta_m} + (\rho\theta_m)', \quad \pi = \bar{\pi} + \pi', \quad (8)$$

185 where the basic state is time-independent and satisfies the hydrostatic equi-
 186 librium

$$\gamma R_d \bar{\pi} \zeta_z \frac{\partial}{\partial \zeta} (\overline{\rho\theta_m}) + \bar{\rho} \mathbf{g} = 0. \quad (9)$$

187 $\gamma = C_p/C_v$ is the ratio of the heat capacities, where C_v is the specific heat
 188 capacity of dry air at constant volume, and \mathbf{g} is gravity acceleration. The
 189 definitions of all variables are summarized in Appendix A.

190 2.1 Momentum equations

191 As described above, the basic equations are transformed to generalized
 192 coordinates using map projection with spherical curvilinear orthogonal co-
 193 ordinates based on a shallow assumption. The transformed equations of
 194 motion are described as

$$\begin{aligned}
& \frac{\partial}{\partial t} \left(\frac{1}{J} \rho u \right) + \frac{\partial}{\partial \xi} \left(\frac{1}{J} \rho u U \right) + \frac{\partial}{\partial \eta} \left(\frac{1}{J} \rho u V \right) + \frac{\partial}{\partial \zeta} \left(\frac{1}{J} \rho_b u W \right) \\
& + \gamma R_d \pi \left\{ \frac{1}{J} \xi_x \frac{\partial}{\partial \xi} (\rho \theta_m)' + \frac{1}{J} \eta_x \frac{\partial}{\partial \eta} (\rho \theta_m)' + \frac{1}{J} \zeta_x \frac{\partial}{\partial \zeta} (\rho \theta_m)' \right\} \quad (10) \\
& = - \sum_{\alpha=\text{sed}} \frac{\partial}{\partial \zeta} \left\{ \frac{1}{J} \rho q_{\alpha} u (W + W_{t\alpha}) \right\} - \frac{1}{J} \rho v \Gamma - \frac{1}{J} \rho v f + \frac{1}{J} F_{\rho u},
\end{aligned}$$

$$\begin{aligned}
& \frac{\partial}{\partial t} \left(\frac{1}{J} \rho v \right) + \frac{\partial}{\partial \xi} \left(\frac{1}{J} \rho v U \right) + \frac{\partial}{\partial \eta} \left(\frac{1}{J} \rho v V \right) + \frac{\partial}{\partial \zeta} \left(\frac{1}{J} \rho_b v W \right) \\
& + \gamma R_d \pi \left\{ \frac{1}{J} \xi_y \frac{\partial}{\partial \xi} (\rho \theta_m)' + \frac{1}{J} \eta_y \frac{\partial}{\partial \eta} (\rho \theta_m)' + \frac{1}{J} \zeta_y \frac{\partial}{\partial \zeta} (\rho \theta_m)' \right\} \quad (11) \\
& = - \sum_{\alpha=\text{sed}} \frac{\partial}{\partial \zeta} \left\{ \frac{1}{J} \rho q_{\alpha} v (W + W_{t\alpha}) \right\} + \frac{1}{J} \rho u \Gamma + \frac{1}{J} \rho u f + \frac{1}{J} F_{\rho v},
\end{aligned}$$

$$\begin{aligned}
& \frac{\partial}{\partial t} \left(\frac{1}{J} \rho w \right) + \frac{\partial}{\partial \xi} \left(\frac{1}{J} \rho w U \right) + \frac{\partial}{\partial \eta} \left(\frac{1}{J} \rho w V \right) + \frac{\partial}{\partial \zeta} \left(\frac{1}{J} \rho_b w W \right) \\
& + \gamma R_d \pi \left\{ \frac{1}{J} \zeta_z \frac{\partial}{\partial \zeta} (\rho \theta_m)' \right\} + \left(\frac{\rho'}{J} - \frac{\pi'}{\bar{\pi}} \frac{\bar{\rho}}{J} \right) \mathfrak{g} \quad (12) \\
& = - \sum_{\alpha=\text{sed}} \frac{\partial}{\partial \zeta} \left\{ \frac{1}{J} \rho q_{\alpha} w (W + W_{t\alpha}) \right\} + \frac{1}{J} F_{\rho w},
\end{aligned}$$

195 where

$$\Gamma = u \frac{m_2}{m_1} \frac{\partial m_1}{\partial \eta} - v \frac{m_1}{m_2} \frac{\partial m_2}{\partial \xi}. \quad (13)$$

196 The variables m_1 and m_2 are map factors relating to map projections
197 (Saito et al. 2001). f is the Coriolis parameter. $F_{\rho u}$, $F_{\rho v}$ and $F_{\rho w}$ are the
198 source and sink terms of momentum based on physical processes for the x -,
199 y -, and z -directions, respectively.

200 The governing equations are solved on the coordinate system based on
 201 hybrid terrain-following vertical coordinates and Lambert conformal projec-
 202 tion in regional configurations, in which the metric tensor is determined by
 203 the vertical coordinate transformation factor and the map factors. Details
 204 of the map projection and vertical coordinates employed in the operational
 205 regional NWP are given in Appendix B.

206 *2.2 Equation for mass conservation*

207 The equation for mass conservation is

$$\frac{\partial}{\partial t} \left(\frac{1}{J} \rho' \right) + \frac{\partial}{\partial \xi} \left(\frac{1}{J} \rho U \right) + \frac{\partial}{\partial \eta} \left(\frac{1}{J} \rho V \right) + \frac{\partial}{\partial \zeta} \left(\frac{1}{J} \rho_b W \right) = - \sum_{\alpha=\text{sed}} \frac{\partial}{\partial \zeta} \left(\frac{1}{J} \rho q_\alpha (W + W_{t\alpha}) \right) + \frac{1}{J} F_\rho, \quad (14)$$

208 where F_ρ is the source, sink and sub-grid transport term of total mass
 209 density.

210 *2.3 Prognostic equation for potential temperature*

211 The thermodynamic equation is

$$\frac{\partial}{\partial t} \left(\frac{1}{J} (\rho \theta_m)' \right) + \frac{\partial}{\partial \xi} \left(\frac{1}{J} \rho \theta_m U \right) + \frac{\partial}{\partial \eta} \left(\frac{1}{J} \rho \theta_m V \right) + \frac{\partial}{\partial \zeta} \left(\frac{1}{J} \rho \theta_m W \right) = \frac{1}{J} \left(\rho_d + \frac{\rho_v}{\epsilon} \right) Q_\theta, \quad (15)$$

212 where Q_θ is diabatic heating.

213 2.4 Prognostic equation for water substances

214 The prognostic equation for the density of water substances is

$$\frac{\partial}{\partial t} \left(\frac{1}{J} \rho q_\alpha \right) + \frac{\partial}{\partial \xi} \left(\frac{1}{J} \rho q_\alpha U \right) + \frac{\partial}{\partial \eta} \left(\frac{1}{J} \rho q_\alpha V \right) + \frac{\partial}{\partial \zeta} \left(\frac{1}{J} \rho q_\alpha (W + W_{t_\alpha}) \right) = \frac{1}{J} F_{\rho\alpha},$$

(16)

215 where $F_{\rho\alpha}$ is the source and sink term for the density of water substances

216 α based on physical processes.

217 2.5 State equation

218 Using ρ and θ_m , the state equation for the ideal gas can be written in

219 the same manner as that for dry conditions:

$$p = R_d \pi \rho \theta_m. \tag{17}$$

220 3. Discretization

221 3.1 Spatial discretization

222 The grid structures of the model are the Arakawa-C type (Arakawa and

223 Lamb 1977) in the horizontal direction and the Lorenz type (Lorenz 1960) in

224 the vertical direction. The equations are spatially discretized using the finite
 225 volume method to conserve total mass across the entire domain; this mass
 226 is controlled by inflow and outflow from the lower and lateral boundaries.

227 3.2 *Advection scheme*

228 Numerical oscillation should be avoided because it can cause spurious
 229 negative values for positive definite prognostic variables (e.g. density) and
 230 computational instability. However, higher-order linear advection schemes
 231 except the first-order scheme are non-monotone (Godunov 1959). Accord-
 232 ingly, a flux limiter function combining the solutions of the higher-order
 233 scheme and the first-order scheme is used to achieve higher-order accuracy
 234 without spurious oscillations (Durrant 2010). Here, let us consider a simple
 235 one-dimensional transport problem with a scalar variable ϕ ,

$$\frac{\partial \phi}{\partial t} + \frac{\partial (u\phi)}{\partial x} = 0, \quad (18)$$

236 and discretize the advection term as

$$\begin{aligned} \mathcal{F}_{i+\frac{1}{2}} &= (u\phi)_{i+\frac{1}{2}} = \left[\phi_i + \frac{1}{2}\Phi(s_{i+\frac{1}{2}})(\phi_i - \phi_{i-1}) \right] u_{i+\frac{1}{2}}, \\ s_{i+\frac{1}{2}} &= \frac{\phi_{i+1} - \phi_i}{\phi_i - \phi_{i-1}}, \end{aligned} \quad (19)$$

237 where $\mathcal{F}_{i+\frac{1}{2}}$, $u_{i+\frac{1}{2}}$, $s_{i+\frac{1}{2}}$ and $\Phi(s_{i+\frac{1}{2}})$ are the flux and wind speed at the edge
 238 of the i -th cell, the smoothness parameter and the flux limiter function,
 239 respectively.

240 The model employs the flux limiter function proposed by Koren(1993)
 241 which combines the third- and first-order upwind schemes using the smooth-
 242 ness parameter s :

$$\Phi(s) = \max \left[0, \min \left\{ 2s, \frac{1}{3} + \frac{2}{3}s, 2 \right\} \right]. \quad (20)$$

243 Figure 1 shows the Sweby diagram (Sweby 1984) for the flux limiter
 244 function in Eq. (20). The striped area indicates the region in which $\Phi(s)$
 245 must lie to preserve monotonicity. When the distribution of ϕ_i is smooth,
 246 the parameter s is close to unity and $\Phi(s) = \frac{1}{3} + \frac{2}{3}s$. Equation (19) provides
 247 the third-order upwind scheme. However, when ϕ_i has a local minimum or
 248 maximum, s is negative and $\Phi(s)$ is zero. The equation then gives the first-
 249 order scheme. Thus, the third-order upwind scheme, which provides high
 250 accuracy, and the first-order upwind scheme, which preserves monotonicity,
 251 are smoothly connected. Based on Eq. (20), Eq. (19) can be written as

$$\mathcal{F}_{i+\frac{1}{2}} = \begin{cases} [\phi_i + \max\{0, \min(x_i, y_i, z_i)\}] u_{i+\frac{1}{2}} & (x_i \geq 0) \\ [\phi_i + \min\{0, \max(x_i, y_i, z_i)\}] u_{i+\frac{1}{2}} & (x_i < 0) \end{cases} \quad (21)$$

$$x_i = \phi_i - \phi_{i-1}, \quad y_i = \phi_{i+1} - \phi_i, \quad z_i = \frac{x_i}{6} + \frac{y_i}{3}.$$

252 Note that division by zero disappears in Eq. (21) in contrast to Eq. (20).

253 Figure 2 shows the results of a one-dimensional (1D) transport problem
 254 in comparison with the first- and third-order advection schemes. A rectan-
 255 gular pulse is advected 2500 time steps in a 200 grid periodic domain using
 256 Courant number of 0.16. In the test, the time integration scheme in Section
 257 3.3 is used. It can be seen that Koren’s flux limiter suppresses overshoot
 258 and undershoot without numerical diffusions.

Fig. 1

Fig. 2

259 3.3 Time integration

260 The three-stage third-order Runge-Kutta (RK3) scheme proposed by
 261 Wicker and Skamarock(2002) is adopted for time integration. In this scheme,
 262 the differential equation

$$\frac{d\phi}{dt} = F(\phi), \quad (22)$$

263 is integrated from $\phi(t)$ to $\phi(t + \Delta t)$ as

$$\begin{aligned}
\phi^* &= \phi(t) + F(\phi(t)) \cdot \frac{1}{3}\Delta t, \\
\phi^{**} &= \phi(t) + F(\phi^*) \cdot \frac{1}{2}\Delta t, \\
\phi(t + \Delta t) &= \phi(t) + F(\phi^{**}) \cdot \Delta t.
\end{aligned}$$

264 This helps to reduce memory consumption because the updated value in
265 each stage can be calculated from the value in the previous stage and $\phi(t)$.

266 3.4 Horizontally explicit and vertically implicit (HE-VI) scheme

267 Terms related to sound waves and gravity waves are evaluated on a
268 short time step using a horizontally explicit and vertically implicit (HE-VI)
269 scheme (Klemp et al. 2007) and the RK3 scheme is also applied on the
270 short time step. Forward time integrations with the short time step $\Delta\tau$ are
271 used for the horizontal momentum equations:

$$\left(\frac{1}{J}\rho u\right)^{\tau+\Delta\tau} = \left(\frac{1}{J}\rho u\right)^{\tau} - \gamma R_d \pi^t \left\{ \frac{1}{J}\xi_x \frac{\partial}{\partial\xi}(\rho\theta_m)'_{\tau} + \frac{1}{J}\eta_x \frac{\partial}{\partial\eta}(\rho\theta_m)'_{\tau} + \frac{1}{J}\zeta_x \frac{\partial}{\partial\zeta}(\rho\theta_m)'_{\tau} \right\} \Delta\tau + \mathfrak{R}_u^t \Delta\tau,$$

(23)

$$\left(\frac{1}{J}\rho v\right)^{\tau+\Delta\tau} = \left(\frac{1}{J}\rho v\right)^{\tau} - \gamma R_d \pi^t \left\{ \frac{1}{J}\xi_y \frac{\partial}{\partial\xi}(\rho\theta_m)'_{\tau} + \frac{1}{J}\eta_y \frac{\partial}{\partial\eta}(\rho\theta_m)'_{\tau} + \frac{1}{J}\zeta_y \frac{\partial}{\partial\zeta}(\rho\theta_m)'_{\tau} \right\} \Delta\tau + \mathfrak{R}_v^t \Delta\tau,$$

(24)

272 where

$$\begin{aligned}\mathfrak{R}_u &= -\frac{\partial}{\partial\xi}\left(\frac{1}{J}\rho u U\right) - \frac{\partial}{\partial\eta}\left(\frac{1}{J}\rho u V\right) - \frac{\partial}{\partial\zeta}\left(\frac{1}{J}\rho u W\right) \\ &\quad - \sum_{\alpha}\frac{\partial}{\partial\zeta}\left(\frac{1}{J}\rho u q_{\alpha} W_{t_{\alpha}}\right) - \frac{1}{J}\rho v \Gamma - \frac{1}{J}\rho v f + \frac{1}{J}F_{\rho u},\end{aligned}\quad (25)$$

$$\begin{aligned}\mathfrak{R}_v &= -\frac{\partial}{\partial\xi}\left(\frac{1}{J}\rho v U\right) - \frac{\partial}{\partial\eta}\left(\frac{1}{J}\rho v V\right) - \frac{\partial}{\partial\zeta}\left(\frac{1}{J}\rho v W\right) \\ &\quad - \sum_{\alpha}\frac{\partial}{\partial\zeta}\left(\frac{1}{J}\rho v q_{\alpha} W_{t_{\alpha}}\right) + \frac{1}{J}\rho u \Gamma + \frac{1}{J}\rho u f + \frac{1}{J}F_{\rho v}.\end{aligned}\quad (26)$$

273 The pressure gradient / buoyancy terms in the vertical momentum equa-
 274 tion and the vertical advection terms in the potential temperature / density
 275 equations are implicitly evaluated to ensure computational stability as

$$\left(\frac{1}{J}\rho w\right)^{\tau+\Delta\tau} = \left(\frac{1}{J}\rho w\right)^{\tau} - \left\{ \gamma R_d \pi^t \frac{1}{J} \zeta_z \frac{\partial}{\partial\zeta} (\rho \theta_m)^{\tau+\Delta\tau} + \frac{\rho'^{\tau+\Delta\tau}}{J} \mathbf{g} - \frac{\pi^t \bar{\rho}}{\bar{\pi}} \frac{1}{J} \mathbf{g} \right\} \Delta\tau + \mathfrak{R}_w^t \Delta\tau, \quad (27)$$

$$\left(\frac{1}{J}(\rho \theta_m)'\right)^{\tau+\Delta\tau} = \left(\frac{1}{J}(\rho \theta_m)'\right)^{\tau} - \left\{ \frac{\partial}{\partial\zeta} \left(\frac{1}{J} \zeta_z \theta_m^{\tau} (\rho w)^{\tau+\Delta\tau} \right) \right\} \Delta\tau + \mathfrak{R}_{\theta_m}^t \Delta\tau, \quad (28)$$

$$\left(\frac{1}{J}\rho'\right)^{\tau+\Delta\tau} = \left(\frac{1}{J}\rho'\right)^{\tau} - \left\{ \frac{\partial}{\partial\zeta} \left(\frac{1}{J} \zeta_z (\rho w)^{\tau+\Delta\tau} \right) \right\} \Delta\tau + \mathfrak{R}_{\rho}^t \Delta\tau, \quad (29)$$

276 where

$$\begin{aligned}
\mathfrak{R}_w &= -\frac{\partial}{\partial\xi}\left(\frac{1}{J}\rho wU\right) - \frac{\partial}{\partial\eta}\left(\frac{1}{J}\rho wV\right) - \frac{\partial}{\partial\zeta}\left(\frac{1}{J}\rho wW\right) \\
&\quad - \sum_{\alpha} \frac{\partial}{\partial\zeta}\left(\frac{1}{J}\rho wq_{\alpha}W_{t_{\alpha}}\right) + \frac{1}{J}F_{\rho w}, \\
\mathfrak{R}_{\theta_m} &= -\left\{\frac{\partial}{\partial\xi}\left(\frac{1}{J}\theta_m(\widetilde{\rho U})\right) + \frac{\partial}{\partial\eta}\left(\frac{1}{J}\theta_m(\widetilde{\rho V})\right) + \frac{\partial}{\partial\zeta}\left(\frac{1}{J}\theta_m(\widetilde{\rho W})\right)\right\} \\
&\quad + \frac{1}{J}\left(\rho_d + \frac{\rho_v}{\epsilon}\right)Q_{\theta}, \\
\mathfrak{R}_{\rho} &= -\left\{\frac{\partial}{\partial\xi}\left(\frac{1}{J}(\widetilde{\rho U})\right) + \frac{\partial}{\partial\eta}\left(\frac{1}{J}(\widetilde{\rho V})\right) + \frac{\partial}{\partial\zeta}\left(\frac{1}{J}(\widetilde{\rho W})\right)\right\} \\
&\quad - \sum_{\alpha} \frac{\partial}{\partial\zeta}\left(\frac{1}{J}\rho q_{\alpha}W_{t_{\alpha}}\right) + \frac{1}{J}F_{\rho},
\end{aligned} \tag{30}$$

277 and

$$(\widetilde{\rho U}) = \xi_x(\rho u)^{\tau+\Delta\tau} + \xi_y(\rho v)^{\tau+\Delta\tau}, \tag{31}$$

$$(\widetilde{\rho V}) = \eta_x(\rho u)^{\tau+\Delta\tau} + \eta_y(\rho v)^{\tau+\Delta\tau}, \tag{32}$$

$$(\widetilde{\rho W}) = \zeta_x(\rho u)^{\tau+\Delta\tau} + \zeta_y(\rho v)^{\tau+\Delta\tau}. \tag{33}$$

It should be noted that terms including $(\rho w)^{\tau+\Delta\tau}$ in Eqs (31) – (33) are omitted under the assumption that the vertical coordinate is restricted to satisfy $\xi_z = \eta_z = 0$ as outlined in Section 2. This restriction is necessary for the vertical implicit treatment of Eqs (27) – (29). Eliminating $\left(\frac{1}{J}(\rho\theta_m)'\right)^{\tau+\Delta\tau}$ and $\left(\frac{1}{J}\rho'\right)^{\tau+\Delta\tau}$ from Eq. (27) using Eqs (28) and (29), the one dimensional Helmholtz type equation for $\omega \equiv \left(\frac{1}{J}\rho w\right)^{\tau+\Delta\tau}$ is de-

terminated as

$$-\Delta\tau^2\gamma R_d\pi^t\frac{1}{J}\zeta_z\frac{\partial}{\partial\zeta}\left(J\frac{\partial}{\partial\zeta}(\zeta_z\theta_m^\tau\omega)\right)-\Delta\tau^2\mathbf{g}\frac{\partial}{\partial\zeta}(\zeta_z\omega)+\omega=\mathfrak{R}, \quad (34)$$

where

$$\begin{aligned} \mathfrak{R} &= \left(\frac{1}{J}\rho w\right)^\tau - \gamma R_d\pi^t\Delta\tau\frac{1}{J}\zeta_z\frac{\partial}{\partial\zeta}\{(\rho\theta_m)^\tau + J\mathfrak{R}_{\theta_m}\Delta\tau\} \\ &\quad - \Delta\tau\mathbf{g}\left(\frac{1}{J}\rho'^\tau + \mathfrak{R}_\rho\Delta\tau\right) + \left\{-\left(1-\frac{\pi^t}{\bar{\pi}}\right)\frac{\bar{\rho}}{J}\mathbf{g} + \mathfrak{R}_w\right\}\Delta\tau. \end{aligned} \quad (35)$$

278 $\omega = 0$ is imposed at the top and bottom boundaries upon resolution of the
279 Helmholtz equation. This is derived from $W = 0$ at these boundaries.

280 3.5 Time splitting

281 a. Time splitting for vertical advection

282 For real-case simulations including physical processes, a strong vertical
283 velocity that does not satisfy the CFL condition is often computed. To
284 ensure computational stability, ASUCA employs a time-splitting scheme
285 for evaluation of vertical advection on the basis of the three-dimensional
286 CFL condition.

287 The stability condition of three-dimensional advection depends on the
288 advection scheme as well as the time integration scheme. The CFL condition
289 for the advection and time integration schemes used in ASUCA is given as,

$$C_1 \leq C_{\text{crit}}, \quad (36)$$

$$C_{\text{crit}} \simeq 1.25,$$

290 where, C_1 is the Courant number for 1D advection and C_{crit} is the criti-
 291 cal value for satisfying the CFL condition as is shown in Appendix C. As
 292 described in Section 3.3, the RK3 scheme is employed in ASUCA with par-
 293 allel splitting (Dubal et al. 2004) of advection in each direction. The CFL
 294 condition for these specifications is

$$C_\xi + C_\eta + C_\zeta \leq C_{\text{crit}}, \quad (37)$$

295 where C_ξ , C_η and C_ζ are the Courant numbers in the ξ , η and ζ directions,
 296 respectively. As this condition can be easily violated in typhoons with
 297 stormy horizontal winds and strong updrafts, time splitting for vertical ad-
 298 vection is adopted. When a time step in evaluation of vertical advection
 299 (in the ζ direction) is divided into N substeps, the condition of Eq. (37) is
 300 modified to

$$C_\xi + C_\eta + C_\zeta/N \leq C_{\text{crit}}. \quad (38)$$

301 In the model, time splitting is applied to columns where

$$C_\xi + C_\eta + C_\zeta \geq \lambda C_{\text{crit}}. \quad (39)$$

302 Here, λ is a safety coefficient set as 0.95 in ASUCA.

303 When time splitting is invoked, RK3 for the short time step $\Delta\tau$ is nested
 304 in the original RK3 time integration with the time step Δt (Fig. 3). Note
 305 that $\Delta\tau$ is given independent of the short time step for the HE-VI scheme
 306 described in Section 3.4. This involves greater computational cost, but
 307 produces the desired higher stability. Divergence damping can be excluded
 308 using RK3 for short time steps. This is desirable for accurate dispersion
 309 relations in compressible equations (e.g. Gassmann and Herzog 2007).

Fig. 3

310 In time splitting, horizontal and vertical advection are evaluated sequen-
 311 tially (Dubal et al. 2004). The prognostic variables are updated using the
 312 horizontal flux \mathcal{F}_ξ and \mathcal{F}_η , and vertical flux \mathcal{F}_ζ is then evaluated with the
 313 updated variables as

$$\phi^{H*} = \phi^n - \left(\frac{\partial}{\partial \xi} \mathcal{F}_\xi^n + \frac{\partial}{\partial \eta} \mathcal{F}_\eta^n \right) \Delta\tau, \quad (40)$$

$$\phi^{n+1} = \phi^{H*} - \left(\frac{\partial}{\partial \zeta} \mathcal{F}_\zeta^{H*} \right) \Delta\tau. \quad (41)$$

314 Sequential time splitting is advantageous for its higher computational
 315 stability as compared to parallel splitting. However, this approach pro-
 316 duces directional distortion because the updated value ϕ^{n+1} depends on the

317 evaluation sequence. Accordingly, sequential splitting is used only when the
 318 condition of Eq. (37) cannot be satisfied in order to minimize errors.

319 *b. Time splitting for sedimentation of precipitable water substances*

320 As sedimentation of precipitable water substances (e.g., rain, snow and
 321 graupel) with high terminal velocity can cause computational instability,
 322 a time-splitting method is employed. The vertical velocities of such sub-
 323 stances are defined as the sum of the vertical velocity of air W and the
 324 terminal velocities $W_{t\alpha}$ as determined from cloud microphysics (e.g. Gunn
 325 and Kinzer 1949).

326 The time-split interval $\Delta\tau_{\text{sed}}$ for sedimentation is dynamically deter-
 327 mined for each column depending on the Courant number C_{sed} for sedimen-
 328 tation. This number for the first time-split step at the vertical level k of
 329 the column is defined as

$$C_{\text{sed},k}^1 = \frac{(W_k^1 + W_{t\alpha,k}^1)\Delta t}{\Delta\zeta_k}, \quad (42)$$

330 where the overscript 1 indicates the first time-level index of the time-split
 331 step.

332 The first time-split step interval $\Delta\tau_{\text{sed}}^1$ for the column is then determined
 333 as

$$\Delta\tau_{\text{sed}}^1 = \begin{cases} \Delta t & (\max(C_{\text{sed},k}^1) \leq 1) \\ \beta \frac{\Delta t}{\max(C_{\text{sed},k}^1)} & (\max(C_{\text{sed},k}^1) > 1). \end{cases} \quad (43)$$

334 Here, $\max(C_{\text{sed},k}^1)$ is the maximum Courant number for the column and β
 335 is a parameter for determining time-split step set as 0.9 in ASUCA.

336 After time integration with $\Delta\tau_{\text{sed}}^1$, the residual time step is $\Delta t' = \Delta t - \Delta\tau_{\text{sed}}^1$.

337 The next time-split step interval $\Delta\tau_{\text{sed}}^2$ is determined from the Courant
 338 number $C_{\text{sed},k}^2 = (W_k^2 + W_{t\alpha,k}^2)\Delta t'/\Delta\zeta_k$ using the updated terminal veloci-
 339 ties $W_{t\alpha,k}^2$, and time integration with $\Delta\tau_{\text{sed}}^2$ is calculated. This procedure is
 340 repeated until no residual time step is left.

341 3.6 Boundary conditions

342 a. Rayleigh damping

343 To prevent wave reflection at the lateral and upper boundaries, the
 344 Rayleigh damping term

$$\frac{\partial\phi}{\partial t} = -m(x, y, z) [\phi(t) - \phi_{\text{ext}}(t)], \quad (44)$$

345 is added to the time tendencies of the prognostic variables ρ' , ρu , ρv ,
 346 ρw , $(\rho\theta_m)'$ and ρq_α near the boundaries. In Eq. (44), ϕ denotes the prog-
 347 nostic variable at the first state of each time step (i.e., Eq. (44) is solved

348 explicitly), and ϕ_{ext} is the value of the parent model providing the lat-
 349 eral and upper boundaries in regional configurations. As the parent model
 350 provides the boundary data with a much larger time interval and coarser
 351 resolution than that of the model, ϕ_{ext} is interpolated in time and space
 352 from the provided data. It should be noted that $\rho w_{\text{ext}} = 0$.

353 The location-based function $m(x, y, z)$ is used to determine the 1/e-
 354 folding time for the boundaries. The function has a maximum at the bound-
 355 ary and decreases with subsequent grid point distance defined as

$$m(x, y, z) = \max(m_x, m_y, m_z), \quad (45)$$

356 where

$$m_x = \begin{cases} \gamma_h \sin^2 \left[\frac{\pi}{2} \left(1 - \frac{d_x}{d_h} \right) \right] & (d_x < d_h) \\ 0 & (d_x \geq d_h), \end{cases} \quad (46)$$

$$m_y = \begin{cases} \gamma_h \sin^2 \left[\frac{\pi}{2} \left(1 - \frac{d_y}{d_h} \right) \right] & (d_y < d_h) \\ 0 & (d_y \geq d_h), \end{cases} \quad (47)$$

$$m_z = \begin{cases} \gamma_v \sin^2 \left[\frac{\pi}{2} \left(1 - \frac{d_z}{d_v} \right) \right] & (d_z < d_v) \\ 0 & (d_z \geq d_v). \end{cases} \quad (48)$$

357 Here, d_x, d_y and d_z are the distances from the boundaries, d_h and d_v
358 are the distances from the lateral and upper boundaries where Rayleigh
359 damping is observed, and γ_h and γ_v are parameters determining damping
360 strength, respectively. d_h, d_v, γ_h and γ_v are empirically determined.

361 *b. Lateral flux adjustment*

362 In regional models, changes in total mass depend on i) the source and
363 sink terms at the surface (i.e. evaporation and precipitation), ii) the density
364 change due to Rayleigh damping, and iii) inflow and outflow at the upper
365 and lateral boundaries as determined from the parent model covering the
366 regional model's domain. As the orders of magnitude of i), ii) and iii) are
367 10^{11} , 10^{11} and 10^{13} [kg] respectively in the operational LFM, we assume i)
368 and ii) are negligible in relation to iii). Then, the change in total mass can
369 be approximated as

$$\frac{\partial M(t)}{\partial t} = F(t), \quad (49)$$

370 where $M(t)$ and $F(t)$ are the total mass in the model domain and the sum
371 of mass flux at the boundaries, respectively. As the parent model provides
372 $M(t)$ and $F(t)$ with a time interval much larger than that of the model, re-
373 gional models must compute $F(t)$ at each time step by interpolating bound-
374 ary data temporally. However, this produces an overall mass error because

375 the interpolated mass flux $F_g(t)$ differs from the parent model prediction,
 376 i.e.:

$$\begin{aligned}
 M^p(t_{n+1}) - M^p(t_n) &= \int_{t_n}^{t_{n+1}} F^p(t) dt \\
 &\neq \sum_{t=t_n}^{t_{n+1}} F_g(t) \Delta t.
 \end{aligned}
 \tag{50}$$

377 Here, t_n and t_{n+1} are the times at which lateral boundary data are given,
 378 and the superscript p indicates variables predicted by the parent model.
 379 Note that $F_g(t)$ at each time step is computed by interpolating $F^p(t_n)$ and
 380 $F^p(t_{n+1})$ temporally and spatially. If the total mass flux predicted by the
 381 parent model $F^p(t)$ exceeds the interpolated mass flux $F_g(t)$, the regional
 382 model will predict a total mass smaller than this and, consequently, a lower
 383 pressure field. To reduce this error, correction for mass flux at boundaries
 384 is required in the regional model.

385 To ensure overall mass consistency with the parent model, regional con-
 386 figurations of ASUCA employ flux adjustment with the value $F_{adj}(t)$ modi-
 387 fying lateral inflow and outflow:

$$M^p(t_{n+1}) - M^p(t_n) = \sum_{t=t_n}^{t_{n+1}} [F_g(t) + F_{adj}(t)] \Delta t.
 \tag{51}$$

388 The adjustment does not correct mass flux at the upper boundary because
 389 vertical velocity at the model top is assumed to be zero (Section 3.4).

390 As there is no established approach for determining $F_g(t)$ and $F_{adj}(t)$
 391 which satisfy Eq. (51), ASUCA employs a method that produces smooth
 392 $M(t)$ values with regard to time, and calculates $F_{adj}(t)$ consequently. $M(t)$
 393 is approximated as a sequence of polynomials $M_n(t)$ via third-order spline
 394 interpolation as

$$M_n(t) \simeq a_n + b_n(t - t_n) + c_n(t - t_n)^2 + d_n(t - t_n)^3, \quad (52)$$

395 where $M_n(t_n) = M^p(t_n)$ and $M_n(t_{n+1}) = M^p(t_{n+1})$. a_n , b_n , c_n and d_n are
 396 coefficients in the interval $t_n \leq t \leq t_{n+1}$ and determined via spline interpola-
 397 tion (i.e., first and second derivatives of $M_n(t_n)$ is identical to $M_{n+1}(t_{n+1})$).
 398 Equations. (49) and (52) give $F(t)$ as

$$\begin{aligned} F(t) &= \frac{\partial M(t)}{\partial t}, \\ &= b_n + 2c_n(t - t_n) + 3d_n(t - t_n)^2. \end{aligned} \quad (53)$$

399 $F_g(t)$ is linearly interpolated using $F(t_n)$ and $F(t_{n+1})$, and $F_{adj}(t)$ is deter-
 400 mined as

$$F_{adj}(t) = b_n + 2c_n(t - t_n) + 3d_n(t - t_n)^2 - F_g(t). \quad (54)$$

401 As ASUCA employs momentums for prognostic variables, horizontal mo-
 402 mentum is adjusted to ensure mass consistency. To mitigate adjustment-

403 related shock, horizontal momentum adjustment is applied over the whole
 404 domain rather than only at boundaries. The value is linearly reduced de-
 405 pending on distance from lateral boundaries as

$$\begin{aligned} \frac{1}{J}\rho\widehat{U}(t, \xi, \eta, \zeta) &= \frac{1}{J}\rho U(t, \xi, \eta, \zeta) + \left(1 - 2\frac{dx(\xi)}{Dx}\right) A(z, t) \\ \frac{1}{J}\rho\widehat{V}(t, \xi, \eta, \zeta) &= \frac{1}{J}\rho V(t, \xi, \eta, \zeta) + \left(1 - 2\frac{dy(\eta)}{Dy}\right) A(z, t), \end{aligned} \quad (55)$$

406 where \widehat{U} and \widehat{V} are adjusted velocities. Dx and Dy are the sizes of the
 407 computational domain in the x - and y - directions, respectively, while $dx(\xi)$
 408 and $dy(\eta)$ are the distances from the western and southern boundaries.
 409 $A(z, t)$ is the horizontal momentum adjustment defined as

$$A(z, t) = \frac{F_{adj}(t)}{2(S_{\eta\zeta} + S_{\xi\zeta})} \frac{\bar{\rho}(z)\Delta z}{\int \bar{\rho}(z)dz}, \quad (56)$$

410 where $S_{\eta\zeta}$ and $S_{\xi\zeta}$ are the areas of the sides of full model domain, $\bar{\rho}(z)$ is the
 411 basic state density, and $\frac{\bar{\rho}(z)\Delta z}{\int \bar{\rho}(z)dz}$ is the mass-fraction of the discretized grid in
 412 the vertical column. $A(z, t)$ is formulated so that total inflow at boundaries
 413 is equivalent to $F_{adj}(t)$ and adjustment horizontal velocity is approximately
 414 uniform with height.

415 This lateral flux adjustment enables ASUCA to predict total mass and
 416 pressure field values consistent with those of the parent model. Related
 417 performance is described in Section 6.2.

418 4. Parallel Computing

419 Parallel computing plays a crucial role in NWP modeling due to the
420 current trend of supercomputer architecture toward massively parallel pro-
421 cessing. To achieve high computational efficiency on massive parallel scalar
422 computers, ASUCA employs hybrid parallelization using the OpenMP inter-
423 face for shared memory parallelization and the MPI for distributed memory
424 parallelization.

425 The three-dimensional arrays are allocated so that the vertical z is the
426 stride-one innermost dimension to make the z - loop contiguous in the mem-
427 ory address, enabling ASUCA to make effective use of the CPU cache. This
428 is also beneficial for code management, as physics schemes, which are gen-
429 erally designed as single-column models (Moncrieff et al. 1997), can be
430 easily implemented in the model. Calculations of physical process schemes
431 at different columns are essentially independent, meaning that OpenMP
432 parallelization can be applied for horizontal loops.

433 The model domain is split into horizontally two-dimensional sub-domains,
434 and each decomposed sub-domain is assigned to one of the MPI processes
435 (Aranami and Ishida 2004). The OpenMP interface is used for paralleliza-
436 tion inside the sub-domains. OpenMP threads are applied to loops for the
437 y -direction, and in some horizontal loops, the x - and y -loops are fused to
438 increase loop length such that load imbalance between threads is minimized.

439 The sub-domains have halo regions that are exchanged with immedi-
440 ately adjacent MPI processes. As MPI communication and file I/O are
441 time-consuming operations with the current supercomputer architecture,
442 two types of overlapping are used in the model to significantly improve
443 computational efficiency. One is overlapping of halo exchanges with com-
444 putation (Cats et al. 2008) to minimize the overhead of communication
445 between MPI processes. The OpenMP interface is also used for this oper-
446 ation; while one thread calls a MPI function to exchange data in the halo
447 region with another MPI process, the other threads simultaneously exe-
448 cute computation in the inner region. The other technique involves an I/O
449 server approach to overlap file I/O with computation. Figure 4 illustrates
450 a schematic diagram of the I/O server approach. In this method, certain
451 MPI processes are dedicated to file I/O. While computation continues, ded-
452 icated I/O processes read data from disks and send them to the relevant
453 computational processes. When output is required, the processes save the
454 data in a dedicated buffer to invoke send operation and immediately con-
455 tinue computation. I/O processes receive the data and output them to the
456 disk. In this approach, computation and disk I/O are asynchronously pro-
457 cessed. It is advantageous for hiding disk I/O latency because disk I/O is
458 a time consuming process. The optimum number of I/O ranks depends on
459 calculation amount, frequency of disk I/O, and computer architecture.

Fig. 4

460 Single-Instruction Multi-Data (SIMD) vectorization is applied to the in-
461 nermost z - direction. However, this is not applicable for z -loops that have
462 loop-carried dependency, such as the ordinary tri-diagonal matrix solver
463 used in the vertical implicit solver for HE-VI (section 3.4), due to verti-
464 cal dependency. To remedy the issue, the Ends Toward Center scheme
465 (Samukawa 2001) is employed for better use of SIMD in the tri-diagonal
466 matrix solver. This contributes to optimization of the model because this
467 calculation is required at every short time step in the HE-VI scheme.

468 The parallel efficiency of ASUCA is shown in Fig.5 for the configuration
469 of the operational LFM. In this experiment, a 10-hour forecast with $1,581 \times$
470 $1,301$ grid points in the horizontal and 58 layers in the vertical was produced
471 with total input/output data sizes of 97 and 17 GB, respectively. Figure
472 5 shows the ideal and measured acceleration ratios from JMA's current
473 supercomputer system Cray XC50 on which each of nodes is equipped with
474 two Intel Xeon Platium 8160 processors with a clock frequency of 2.1GHz
475 and 24 cores per processors (JMA 2019a). The horizontal axis represents
476 the number of CPU cores. There are 8 OpenMP threads up to 14,400 cores
477 while 12, 16 and 24 threads are used for 19,200 and 28,800, 24,000, and
478 38,400 and 48,000 cores, respectively. The model demonstrates more than
479 50% of ideal scalability up to 24,000 cores even though full-size output to
480 the disk in operational LFM configuration is included in the elapsed time.

Fig. 5

481 5. Physical Processes

482 Physical process schemes are implemented via the independent Physics
483 Library of the ASUCA dynamical core (Hara et al. 2012). The library is a
484 group of various subroutines related to physical process schemes, and pro-
485 vides a common interface based on the unified coding rules. Physical process
486 schemes in the library are designed as the vertical one-dimensional models
487 independently of the horizontal grids. This enables the constitution of a
488 single-column model for unit testing and comparison of parameterization
489 schemes. The coding policy is also suitable for modern scalar computers
490 because an improved cache hit rate is crucial for processing speed.

491 The Physics Library is utilized in the procedures described here. Model
492 variables are converted to variables required as arguments for subroutines
493 implemented in the library. For instance, if the velocity u is required in the
494 library, u is calculated from ρ and ρu which are the prognostic variables of
495 ASUCA. The subroutine in the library calculates and returns the tendency
496 of u , which is then converted to the tendency of the model variable ρu .
497 Subroutines implemented in the library are not used to directly update
498 prognostic variables. Hence, an NWP model can call a subroutine in a
499 common style regardless of how it is implemented. These rules contribute
500 to efficient updating of the physical process schemes applied to ASUCA.

Table 1

501 The physical process schemes are regularly updated in operational use.

502 Those used in the LFM since March 2021 are summarized in Table 1, and
503 the schemes are detailed in JMA (2019b). The surface scheme employs
504 a tiled approach in which area fractions of different surface types such as
505 land and sea are given in the same grid. Turbulent fluxes for all tiles are
506 calculated, and a grid point value of these fluxes is evaluated as the weighted
507 average in proportion to the area fraction of each tile in the grid.

508 Computational stability is essential for NWP model operation while a
509 sufficiently small time step could not be adopted for evaluating physical
510 processes, because calculation must finish within a certain time. To sat-
511 isfy these contradictory requirements, some physical process schemes (e.g.,
512 cloud microphysics, surface flux and boundary layer) are implicitly calcu-
513 lated. In cloud microphysics, processes in which change rate of a variable
514 is proportional to the amount of the variable itself (e.g., accretion of cloud
515 ice by snow; Lin et al. 1983) are solved implicitly. It should be noted that
516 vertical flux in the boundary layer scheme is evaluated independently of
517 the surface flux scheme while both must be coupled for implicit evaluation.
518 The Physics Library provides an implicit solver to enable coupling of the
519 boundary layer and the surface schemes with these schemes implemented
520 as separated packages.

521 In ASUCA, radiation, boundary layer and surface, and convection schemes
522 are computed using the first state of time-steps independently of each other

523 (i.e., parallel splitting; Dubal et al. 2004). Microphysics is computed at
524 the end of time-steps sequentially to guarantee non-negative hydrometeors
525 while saturation adjustment is computed at every RK3 steps.

526 **6. Validation Tests for Operational Use**

527 *6.1 Ideal experiments*

528 Various ideal experiments were conducted to validate basic ASUCA dy-
529 namics performance. Ishida et al.(2010) reported the results of experiments
530 for non-hydrostatic inertia gravity waves as originally proposed by Ska-
531 marock and Klemp(1994), and for non-linear density current with the re-
532 sults obtained by Straka et al.(1993) referenced as a benchmark. The results
533 of ideal mountain wave and rising thermal simulation tests are presented
534 below in comparison with the JMA-NHM outcomes.

535 *a. Mountain wave tests*

536 ASUCA simulation provided better results than the JMA-NHM in a two-
537 dimensional linear non-hydrostatic mountain wave test as per the “Standard
538 Test Set for Non-hydrostatic Dynamical Cores of NWP Models” (Skamarock
539 et al. 2004), which enables evaluation of simulated non-hydrostatic topo-
540 graphic flows based on comparison to the analytic solution.

541 Uniform flow with a constant wind speed of 10 ms^{-1} and a Brunt-Väisälä

542 frequency of 0.01 s^{-1} over mountainous terrain were considered. The moun-
543 tain profile, $h(x)$, was given as

$$h(x) = h/(1 + (x/a)^2), \quad (57)$$

544 where $a = 2 \text{ km}$ and $h = 1 \text{ m}$. The computational domain size was 144 km
545 horizontally and 30 km vertically, with the grid spacings of 400 and 250 m ,
546 respectively. The mountain is located at the center of the horizontal domain.
547 Cyclic boundary conditions for the lateral boundaries were assumed, and
548 Rayleigh damping was used for the top 6 km layers to relax the state back to
549 the initial field to reduce the artificial reflection of gravity-wave. Time-step
550 intervals of 3 and 1 s were used for ASUCA and the JMA-NHM, respectively.

551 Figure 6 shows the analytic solution and the simulation results from
552 ASUCA and the JMA-NHM. Note that the effect of Rayleigh damping does
553 not appear in Fig. 6 as the displayed domain is lower 12 km of the com-
554 putational domain. The difference between the mountain wave simulated
555 by ASUCA and the related analytical solution is smaller than that of the
556 JMA-NHM. The normalized L2 norm of the error in the vertical velocity
557 from the analytic solution for ASUCA and JMA-NHM results are 0.192 and
558 0.477 , respectively. Note that the error in ASUCA is smaller than that in
559 the JMA-NHM even though the time-step interval used in ASUCA is three
560 times larger in this experiment.

Fig. 6

561 *b. Rising thermal simulation*

562 Numerical simulation for a rising thermal in a uniform horizontal flow in
563 a two-dimensional adiabatic atmosphere, based on Wicker and Skamarock(1998),
564 was carried out to evaluate basic performance for idealized convection and
565 advection. The grid spacing is 125 m in both the x - and z -directions, and the
566 computational domain is 20 km wide and 10 km deep. The initial thermal
567 (diameter: 4 km) is placed at a height of 2 km with a potential temperature
568 of 2 K higher than the surrounding environment and neutral stratification.
569 The test imposes a uniform horizontal flow of 20 ms^{-1} and integrated for
570 1000 s, so that the thermal is laterally advected in a horizontally periodic
571 domain and should be located in the center of the horizontal domain. A
572 time-step interval of 2 s is used for ASUCA, while a 1 s time interval is used
573 for the JMA-NHM because serious deterioration in simulation is produced
574 with a 2 s interval.

Fig. 7

575 Figure 7 shows the results for ASUCA and the JMA-NHM. The po-
576 tential temperature and vertical velocity fields with ASUCA are symmet-
577 rical, while the potential temperature field with the JMA-NHM is much
578 less symmetrical and the vertical velocity field is distorted and dispersive.
579 The asymmetric result produced by JMA-NHM is mainly due to its fourth-
580 order advection and leapfrog time integration scheme, as shown by Wicker
581 and Skamarock(1998). The normalized L2 norm of the error in the vertical

582 velocity, using the results from each model simulation with no horizontal
583 flow as the benchmark, are 0.189 for ASUCA and 0.236 for JMA-NHM,
584 respectively.

585 *6.2 Performance as an operational NWP*

586 This section compares the performance of the ASUCA-LFM (the 2 km-
587 resolution operational regional model) to that of the NHM-LFM (Aranami
588 et al. 2015).

589 A simulation involving Karman vortex streets, which often form down-
590 wind of islands during the cold-air outbreaks in winter, is presented here
591 as an example of favorable representation using ASUCA dynamics. In Fig.
592 8, the ASUCA-LFM appears to reproduce the phenomenon better than the
593 NHM-LFM. The JMA-NHM with numerical diffusion coefficients weaker
594 than those of operational configuration could reproduce the Karman vor-
595 texes streets. However, these weakened values frequently cause computa-
596 tional instability.

Fig. 8

597 In regional models, the predicted synoptic-scale pressure field should be
598 consistent with that prescribed as the lateral boundary condition. Figure
599 9 shows differences in sea level pressure forecasts between the LFM and
600 the external model (MSM) providing the boundary condition for 1200 UTC
601 on 25 Dec. 2012. The ASUCA-LFM (left) follows MSM prediction at

602 the synoptic-scale, while the local pattern differs due to differences in their
603 prediction properties. However, the synoptic-scale pressure field determined
604 from the NHM-LFM (right) deviates from the prediction of the MSM. This
605 superior consistency mainly comes from improved total mass conservation
606 in the ASUCA-LFM. The lateral boundaries control the net mass flux of
607 the model domain, and consequently control the synoptic-scale pressure
608 field. ASUCA explicitly calculates mass inflow and outflow because density
609 is directly integrated as a prognostic variable. Accordingly, mass change
610 across the entire domain coincides exactly with the total mass change due
611 to inflow and outflow imposed at the region boundaries, source and sink
612 at the surface, and density change by Rayleigh damping. However, the
613 NHM-LFM, in which pressure is a prognostic variable, does not readily
614 conserve total mass because errors are inevitable in evaluating density from
615 the equation of state.

Fig. 9

616 The ASUCA-LFM also achieved higher NWP performance including
617 quantitative precipitation forecast (QPF) accuracy than NHM-LFM (Fig.
618 10). It should be noted that the QPF change is attributed to the updates
619 of physical process schemes as well as the dynamical core itself. As it was
620 also confirmed that ASUCA-LFM had better performance in terms of con-
621 sistency with ground-based and radiosonde observations, ASUCA replaced
622 JMA-NHM as the LFM in 2015, and the MSM in 2017 subsequently.

Fig. 10

7. Future Plans

As described in Section 1, development of the new non-hydrostatic numerical model ASUCA was begun in 2007. The model is designed to ensure accuracy for NWP model, maintainability of components including physical processes and the data assimilation system and code structure suitable for future supercomputer architectures toward the establishment of a long-term operational forecasting infrastructure with a new-generation NWP model. ASUCA replaced the previous regional NWP model JMA-NHM (Saito et al. 2006), in 2015 as the LFM with 2 km resolution and in 2017 as the MSM with 5 km resolution. The ASUCA-based Mesoscale Ensemble Prediction System (Ono et al. 2021) has been in operation since 2019 and the ASUCA-based 4D-Var system (Ikuta et al. 2021), which is an outcome of relational developments, has been used to provide initial MSM fields since 2020. We close this paper with some ongoing development of ASUCA.

Increases of computational power in the future enable us to operate a regional NWP with higher model resolution and a wider forecast region, which could contribute to improving forecasts of severe weather. However, this may also give rise to new issues to be addressed in the improvement of high-resolution model accuracy.

While numerical models start to partly resolve cumulus convection with increased resolution, unresolved motions still remain to be parameterized

644 due to the incompleteness of the motions resolved in the model. As assump-
645 tions made in conventional parameterization schemes are also often violated
646 in such a regime, new parameterization schemes suitable for partially unre-
647 solved processes are required. This is known as the *gray zone* problem, and
648 has recently drawn attention in the fields of research on cumulus convec-
649 tion (e.g., Arakawa and Wu 2013) and boundary layer turbulence (Honnert
650 et al. 2020). It should be emphasized that the improvement of dynamical
651 processes is even more important than ever in the gray zone as resolved part,
652 which is represented by dynamical processes, increases at higher resolution.

653 The utilization of Koren’s flux limiter enables the elimination of an ex-
654 plicit numerical filter and only advection scheme involves diffusion which
655 highly depends on wind speed and direction (i.e. acting only in the wind-
656 ward direction). This results in overly frequent prediction of intense vertical
657 velocity because diffusion relating to horizontal advection is relatively small
658 in such situations. Accordingly, parameterization of horizontal diffusion
659 with physical consideration may be necessary.

660 Some recent convection schemes (e.g., Kuell et al. 2007; Malardel and
661 Bechtold 2019) relax the conventional assumption that the change of den-
662 sity by convection is negligible (i.e., environmental subsidence cancels the
663 convective mass fluxes). The coupling of such physics schemes with the cur-
664 rent dynamical core is a significant challenge because terms evaluated by

665 those schemes relate to sound waves and their implementation may require
666 the modification of the HE-VI scheme.

667 A numerical model with higher spatial resolution can also resolve smaller
668 scales of topography, whose favorable representation is known to help im-
669 prove model performance (Kanehama et al. 2019; Sandu et al. 2019).
670 For instance, local circulation and precipitation processes derived from the
671 topography can be more accurate. However, finer topographical represen-
672 tations incorporating steep slopes can significantly distort the shape of the
673 control volume and affect numerical stability because the vertical axis in
674 the terrain-following coordinate is restricted to the direction of gravity in
675 the current configuration.

676 Steeper terrain can also cause significant pressure gradient force errors.
677 In slope-containing grids, evaluation of horizontal pressure gradient force
678 requires consideration of the vertical pressure gradient in generalized coor-
679 dinates as per Eqs. (10) and (11). As the centered difference is used for
680 the vertical pressure gradient, pressure is linearly interpolated using sur-
681 rounding pressure values. Linear vertical interpolation of pressure creates
682 larger discretization errors for steeper slopes because pressure changes with
683 height are almost exponential. Modification of pressure gradient force com-
684 putation via the interpolation of pressure to a constant height for adjacent
685 columns (e.g., Klemp 2011) may reduce such errors. However, this requires

686 further consideration in future work.

687 The currently operational non-hydrostatic model ASUCA has improved
688 NWP accuracy for heavy rain and typhoons as well as contributing to a
689 better understanding of extreme weather conditions around Japan and Asia.
690 Ongoing development of the model is expected to improve NWP accuracy
691 even more.

692 **A. List of symbols**

693 Symbols used in this paper are listed below in alphabetical order. The
694 subscript α refers to dry air, water vapor, cloud water, cloud ice, rain, snow,
695 and graupel for d, v, c, i, r, s and g respectively. Cartesian coordinates and
696 generalized coordinates are referred to as (x, y, z) and (ξ, η, ζ) , respectively.

A	momentum flux adjustment
β	parameter for determining sedimentation short time step
C	Courant number
C_{sed}	Courant number for sedimentation
C_p	specific heat capacity of dry air at constant pressure
C_v	specific heat capacity of dry air at constant volume
d_h	arbitrary parameter for determination of damping zone width from lateral boundaries
d_v	arbitrary parameter for determination of damping layer thickness from upper boundary
d_x	x -direction distance from lateral boundaries

d_y	y -direction distance from lateral boundaries
d_z	z -direction distance from upper boundary
D_x	x -direction size of computational domain
D_y	y -direction size of computational domain
ϵ	ratio of gas constants for dry air and water vapor (R_d/R_v)
f	Coriolis parameter
\mathcal{F}	fluxes of prognostic variables
F_M	mass flux from lateral boundaries
$F_{M,g}$	mass flux interpolated temporally from lateral boundaries
F_ρ	source, sink, and sub-grid transport term of total mass density
$F_{\rho\alpha}$	source, sink, and sub-grid transport term of density of α
$F_{\rho\theta_m}$	source, sink, and sub-grid transport term of $\rho\theta_m$
$F_{\rho u}$	source, sink, and sub-grid transport term of momentum for x -direction
$F_{\rho v}$	source, sink, and sub-grid transport term of momentum for y -direction
$F_{\rho w}$	source, sink, and sub-grid transport term of momentum for z -direction
\mathbf{g}	gravity acceleration
γ	ratio of specific heat capacities for dry air at constant pressure and volume (C_p/C_v)
γ_h	arbitrary parameter for determination of damping strength at lateral boundaries
γ_v	arbitrary parameter for determination of damping strength at upper boundary
Γ	curvature of map factors
J	Jacobian of coordinate transformation from Cartesian to generalized coordinates

m	location-based function for determination of damping strength at lateral and/or upper boundaries
m_x	damping strength at x -direction lateral boundaries
m_y	damping strength at y -direction lateral boundaries
m_z	damping strength at upper boundary
m_1	map factor for x -direction
m_2	map factor for y -direction
nx	grid points number in x -direction
ny	grid points number in y -direction
ω	w -direction momentum at short time step
p	pressure
p_0	reference pressure
ϕ	scalar variable
Φ	flux limiter function
π	Exner function
q_α	density ratio of water substance α for total mass density
Q_θ	diabatic heating by physical processes
ρ	total mass density
ρ_α	density of category α
ρ_b	total density of dry air and water substances assumed to have zero terminal velocity
R_d	gas constant for dry air
R_v	gas constant for water vapor

\mathfrak{R}	term defined in Eq. (35)
\mathfrak{R}_{θ_m}	term defined in Eq. (30)
\mathfrak{R}_ρ	term defined in Eq. (30)
\mathfrak{R}_u	x -direction momentum tendency terms solved with long time step
\mathfrak{R}_v	y -direction momentum tendency terms solved with long time step
\mathfrak{R}_w	w -direction momentum tendency terms solved with long time step
s	smoothness parameter for determination of flux limiter value
t	time
Δt	full model time step
τ	time at Runge-Kutta step
$\Delta\tau$	short time step for determination of acoustic and gravity wave modes
$\Delta\tau_{\text{sed}}$	time-split step for sedimentation
θ	potential temperature
θ_m	modified moist potential temperature
u	velocity component in x -direction
U	velocity component in ξ -direction
v	velocity component in y -direction
V	velocity component in η -direction
w	velocity component in z -direction
W	velocity component in ζ -direction
W_{t_α}	terminal fall velocity of water substance α

z_s surface height
 z_T model top height

697 Subscripts and superscripts

$()_x$ partial differential of $()$ with respect to x
 $()_y$ partial differential of $()$ with respect to y
 $()_z$ partial differential of $()$ with respect to z
 $\overline{()}$ basic state satisfying hydrostatic equilibrium
 $()'$ deviation from basic state
 $()^t$ value at time step
 $()^\tau$ values at Runge-Kutta steps
 $()^{\tau+\Delta\tau}$ values at HE-VI short time steps
 $\tilde{()}$ generalized coordinate momentum in HE-VI short time steps

698 **B. Map projection and vertical coordinates in the op-** 699 **erational models**

700 The details of map projection and the vertical coordinates employed in
701 the LFM and MSM, which are operational regional NWP with 2 km and
702 5 km resolutions respectively, are presented here.

703 ASUCA employs the Lambert conformal map projection. The map fac-

704 tors m_1 and m_2 (for the x and y directions) here are given by

$$m_1 = m_2 = m = \left(\frac{\cos \varphi}{\cos \varphi_1} \right)^{a-1} \left(\frac{1 + \sin \varphi_1}{1 + \sin \varphi} \right)^a, \quad (58)$$

$$a = \ln \left(\frac{\cos \varphi_1}{\cos \varphi_2} \right) / \ln \left\{ \frac{\tan \left(\frac{\pi}{4} - \frac{\varphi_1}{2} \right)}{\tan \left(\frac{\pi}{4} - \frac{\varphi_2}{2} \right)} \right\}. \quad (59)$$

705 where φ is latitude, and $\varphi_1 = 30^\circ$ and $\varphi_2 = 60^\circ$ are used as standard
706 parallels in the operational models.

707 The hybrid terrain-following vertical coordinate (Ishida 2007) is adopted
708 to reduce the influences of topography with height. Since the horizontal
709 pressure gradient term and the horizontal advection term are split into hor-
710 izontal and vertical derivatives with non-flat coordinates, and the vertical
711 grid spacing of NWP models is generally larger in the upper atmosphere,
712 reduction of errors associated with related difference calculation is advan-
713 tageous. The vertical coordinate ζ is transformed using

$$z = \zeta + z_s h(\zeta), \quad (60)$$

714 where z is height above sea level and z_s is ground height. The function $h(\zeta)$
715 is given by

$$h(\zeta) = \frac{b \left\{ 1 - \left(\frac{\zeta}{z_T} \right)^n \right\}}{b + \left(\frac{\zeta}{z_T} \right)^n}, \quad b = \frac{\left(\frac{z_c}{z_T} \right)^n}{1 - 2 \left(\frac{z_c}{z_T} \right)^n}, \quad (61)$$

716 where z_T is the model top. z_c and n are parameters characterizing the
 717 influence of terrain; z_c is the height at which the center of the transition,
 718 between the terrain-following coordinate and the flat coordinate, is located,
 719 and n determines the varying rate of the transition. $z_c = 7000\text{m}$ and $n = 3$
 720 are employed in the LFM and MSM. Figure 11 shows the model levels over
 721 idealized mountain in the hybrid terrain-following coordinate in contrast to
 722 the basic terrain-following coordinate (Gal-Chen and Somerville 1975). The
 723 hybrid terrain-following coordinate is identical to the terrain-following and
 724 flat coordinate at $z = z_s$ and $z = z_T$, respectively, and the two coordinates
 725 are smoothly connected.

Fig. 11

726 C. CFL condition for the advection with the Koren 727 flux limiter and RK3 scheme

728 The one-dimensional advection equation with a uniform velocity $u(> 0)$
 729 is considered as follows:

$$\frac{\partial \tilde{f}}{\partial t} = -u \frac{\partial \tilde{f}}{\partial x}, \quad x \in [0, 2\pi]. \quad (62)$$

730 The spatial direction is discretized with the grid spacing $\Delta x = 2\pi/N$ (N :
731 the number of grid cells). We assume the following solution in Eq. (62):

$$\tilde{f}(x, t) = f(t)e^{ikx}, \quad k = 0, 1, \dots, N/2. \quad (63)$$

732 The advection term is approximated by the first- or third-order upwind
733 difference in the Koren flux limiter. The spatial derivative at the j -th grid
734 using Eq. (63) is represented as

$$\left. \frac{\partial \tilde{f}}{\partial x} \right|_j \simeq \frac{\tilde{f}_j - \tilde{f}_{j-1}}{\Delta x} = \frac{1 - e^{-ik\Delta x}}{\Delta x} \tilde{f}_j, \quad (64)$$

735 for the first-order upwind difference. Substituting Eq. (64) into (62) yields

$$\frac{df}{dt} = \frac{u}{\Delta x} (e^{-ik\Delta x} - 1)f. \quad (65)$$

736 In the third-order Runge-Kutta time integration, Eq. (65) is stable if

$$|f^{n+1}/f^n| = |1 + z + z^2/2 + z^3/6| \leq 1, \quad z = C_1(e^{-ik\Delta x} - 1), \quad (66)$$

737 is satisfied for $0 \leq k\Delta x \leq \pi$. Here, the superscript n denotes the n -th
738 timestep and $C_1 = u\Delta t/\Delta x$ is the Courant number. Solving Eq. (66)
739 numerically, we can obtain the CFL condition $C_1 \leq 1.25$ for the first-order
740 upwind difference.

741 The spatial derivative for the third-order upwind difference is written as

$$\left. \frac{\partial \tilde{f}}{\partial x} \right|_j \simeq \frac{2\tilde{f}_{j+1} + 3\tilde{f}_j - 6\tilde{f}_{j-1} + \tilde{f}_{j-2}}{6\Delta x} = \frac{2e^{ik\Delta x} + 3 - 6e^{-ik\Delta x} + e^{-2ik\Delta x}}{6\Delta x} \tilde{f}_j. \quad (67)$$

742 The CFL condition $C_1 \leq 1.61$ for the third-order upwind difference can be
 743 obtained by the similar procedure. This value coincides with that by Wicker
 744 and Skamarock(2002). Because the critical value of C_1 for the first-order
 745 is smaller than that for the third-order, $C_1 \leq 1.25$ should be chosen as the
 746 CFL condition for the Koren flux limiter.

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751 Data Availability Statement

752 Some of the datasets and program codes used in this study are not
 753 publicly available due to the management policy of the Japan Meteorological
 754 Agency, but may be available from the relevant authors for reasonable usage
 755 upon request. All rights remain with JMA.

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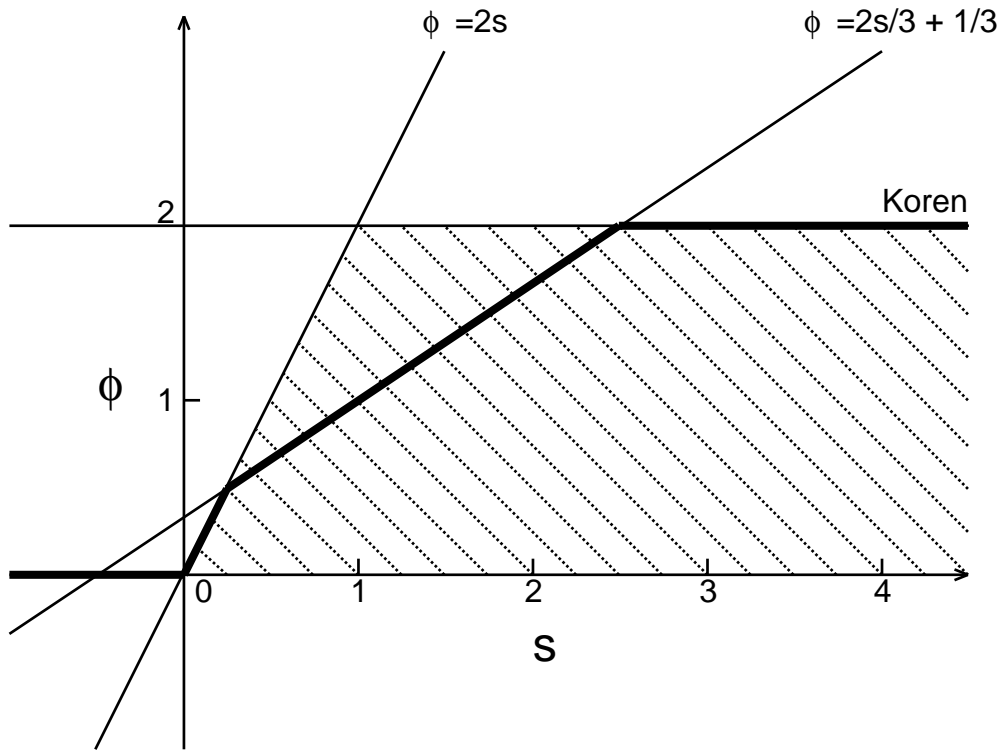


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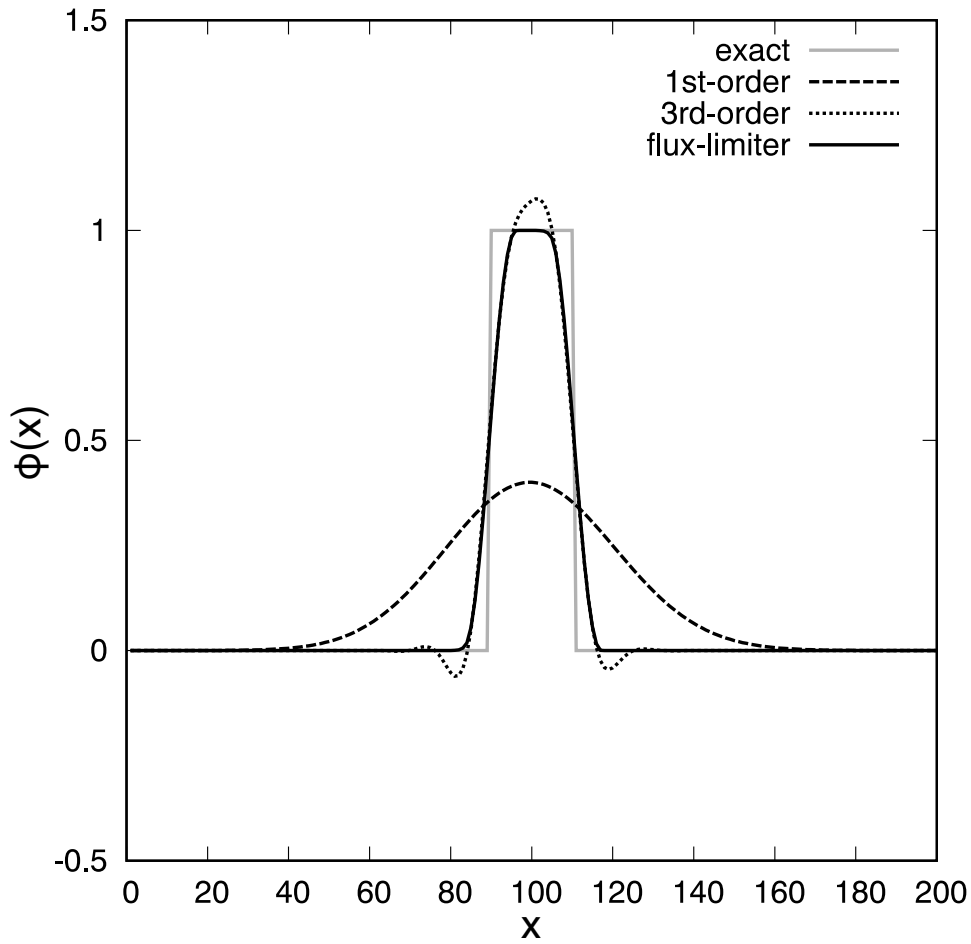


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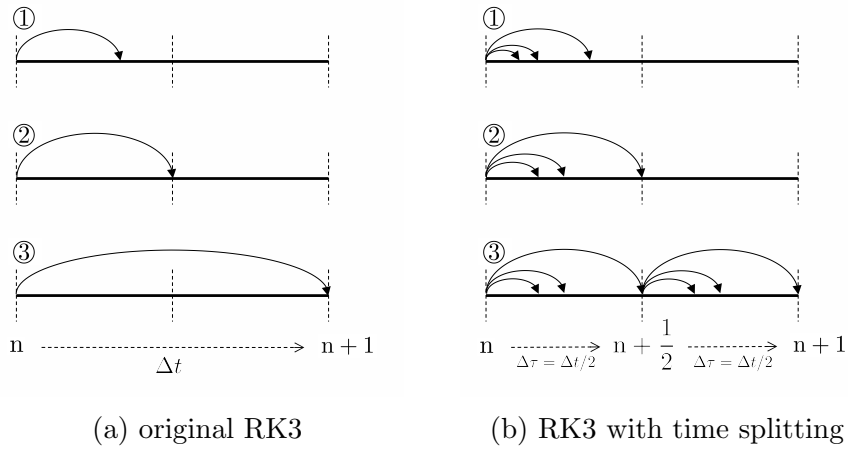


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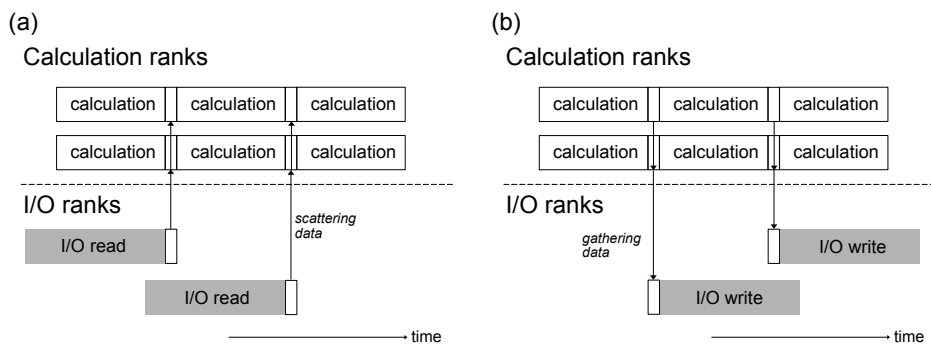


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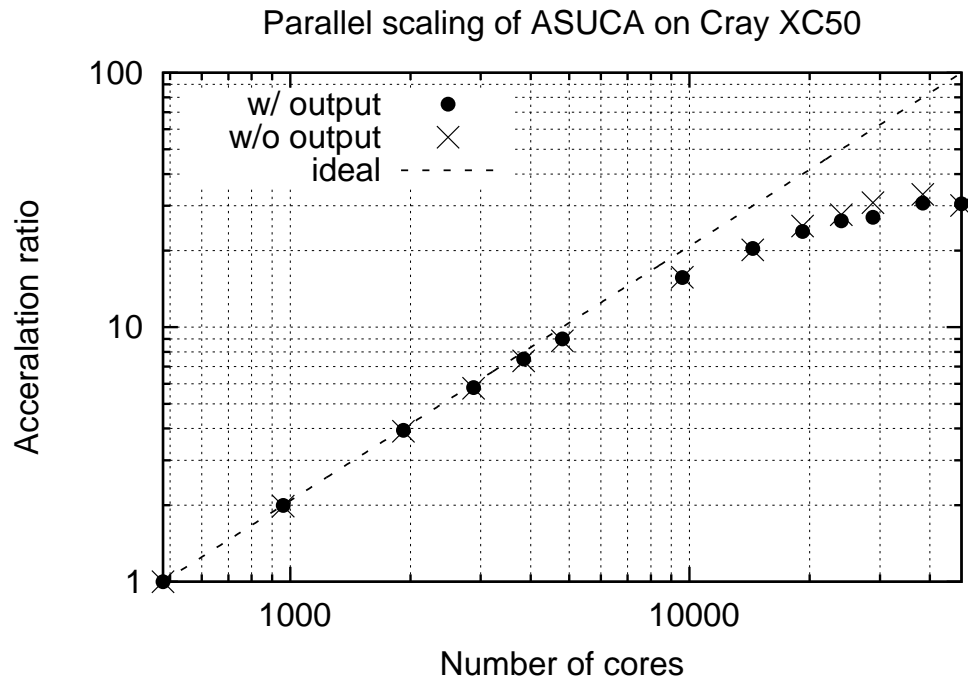


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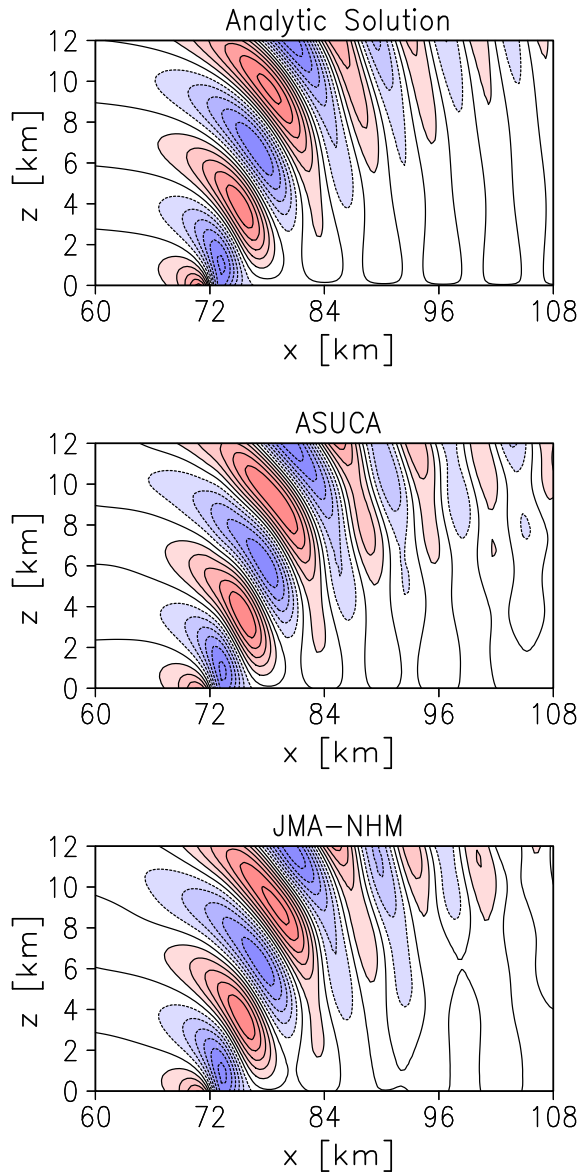


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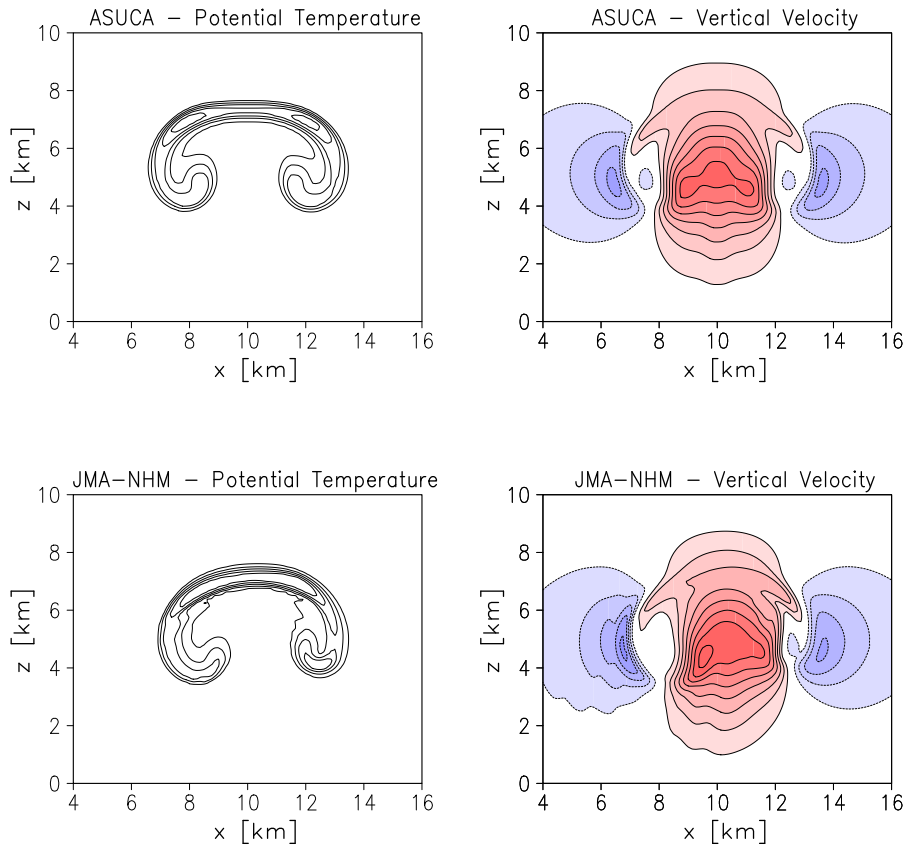


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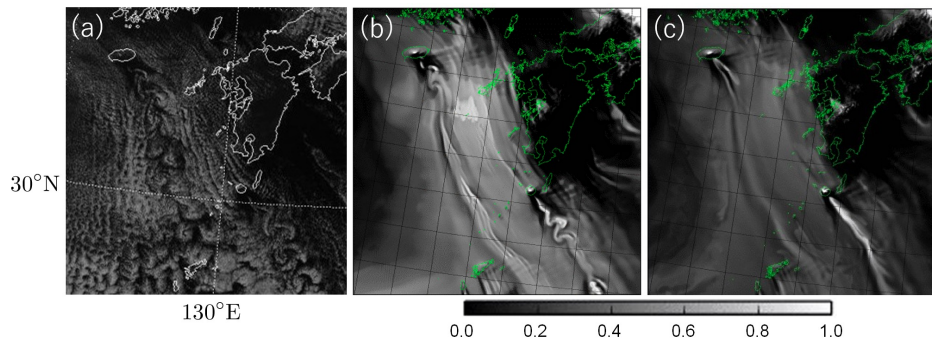


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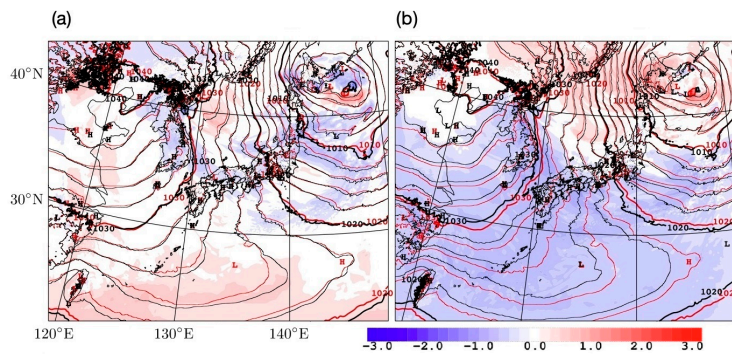


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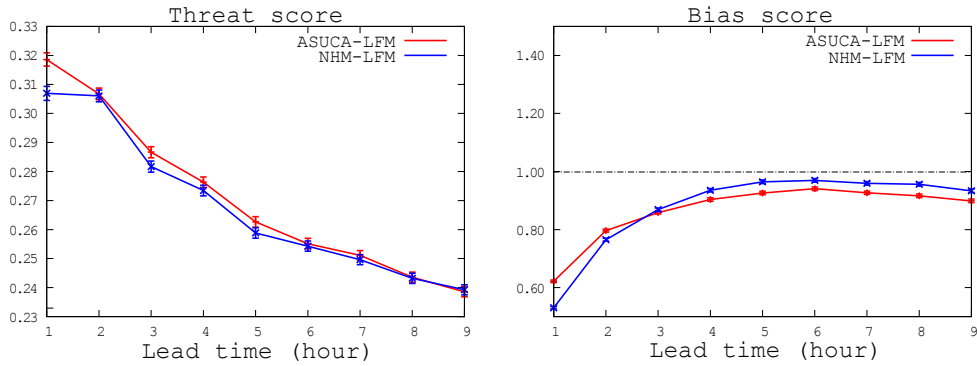


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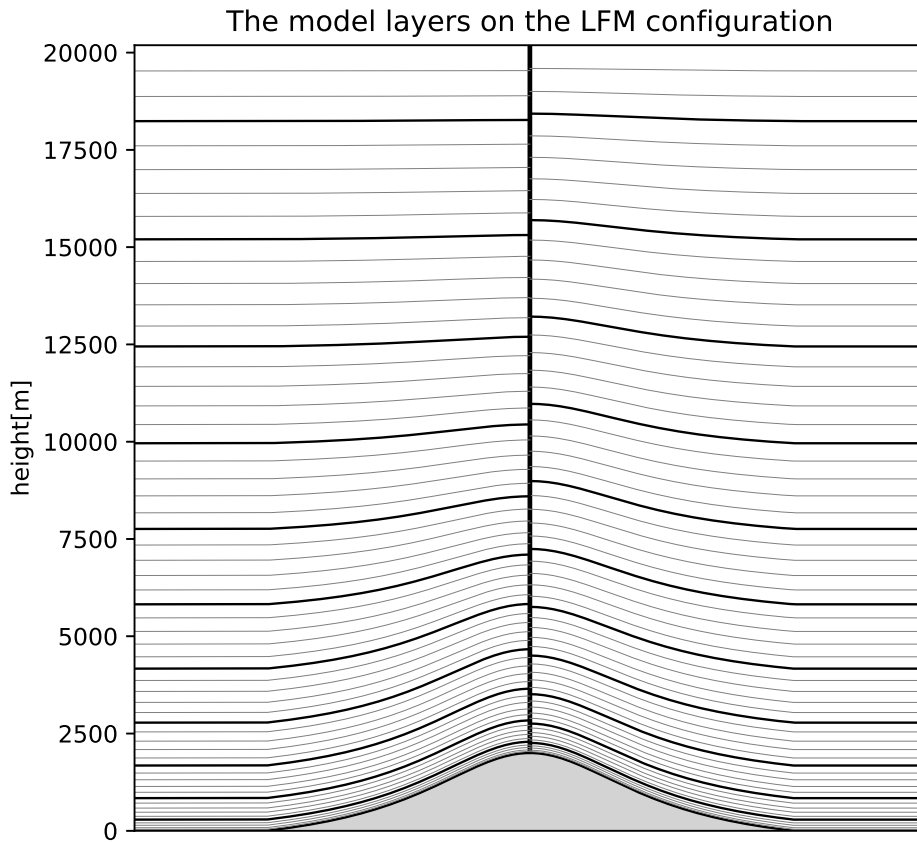


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978		March 2021.	76

Table 1: Physical process schemes used in the LFM operated since March 2021.

Process	Scheme
Radiation	Short wave: two-stream with delta-Eddington approximation (evaluated every 15 minutes) (Joseph et al. 1976; Coakley et al. 1983) Long wave: two-stream absorption approximation (evaluated every 15 minutes) (Yabu 2013)
Boundary layer	Mellor-Yamada-Nakanishi-Niino Level-3 scheme (Nakanishi and Niino 2009)
Surface flux	Monin-Obukhov similarity theory with stability function (Beljaars and Holtslag 1991; Gryanik et al. 2020)
Soil	Ground temperature prediction using an eight-layer ground model (Noilhan and Planton 1989)
Convection	Kain-Fritsch convection scheme (Kain 2004; Kain and Fritsch 1990)
Cloud microphysics	Single moment, three-ice bulk method (Lin et al. 1983; Ikawa and Saito 1991)