Bright spatially coherent synchrotron X-rays from a table-top source

S. Kneip¹*, C. McGuffey², J. L. Martins³, S. F. Martins³, C. Bellei¹, V. Chvykov², F. Dollar², R. Fonseca^{3,4}, C. Huntington², G. Kalintchenko², A. Maksimchuk², S. P. D. Mangles¹, T. Matsuoka², S. R. Nagel¹, C. A. J. Palmer¹, J. Schreiber¹, K. Ta Phuoc⁵, A. G. R. Thomas², V. Yanovsky², L. O. Silva³, K. Krushelnick² and Z. Najmudin¹*

Each successive generation of X-ray machines has opened up new frontiers in science, such as the first radiographs and the determination of the structure of DNA. State-of-the-art X-ray sources can now produce coherent high-brightness Xrays of greater than kiloelectronvolt energy and promise a new revolution in imaging complex systems on nanometre and femtosecond scales. Despite the demand, only a few dedicated synchrotron facilities exist worldwide, in part because of the size and cost of conventional (accelerator) technology¹. Here we demonstrate the use of a new generation of laserdriven plasma accelerators², which accelerate high-charge electron beams to high energy in short distances³⁻⁵, to produce directional, spatially coherent, intrinsically ultrafast beams of hard X-rays. This reduces the size of the synchrotron source from the tens of metres to the centimetre scale, simultaneously accelerating and wiggling the electron beam. The resulting X-ray source is 1,000 times brighter than previously reported plasma wigglers^{6,7} and thus has the potential to facilitate a myriad of uses across the whole spectrum of light-source applications.

There are a number of proposals to use extreme nonlinear interactions of the latest generation of high-power ultrashort-pulse laser systems to produce beams of high-energy photons with high brightness and short pulse duration. For example, high-order harmonic generation promises trains of coherent pulselets⁸ and Compton scattering could extend energies into the γ -regime^{9,10}. An alternative proposal has been the use of compact laser–plasma accelerators to drive sources of undulating/wiggling radiation¹¹.

These accelerators use the plasma wakefield generated by the passage of an intense laser pulse through an underdense plasma¹². Such wakefields can have intrinsic fields of more than 1,000 times greater than the best achievable by conventional accelerator technology, and thus can accelerate particles to high energies in a fraction of the distance. Recently, it has been demonstrated that at high laser power, the wakefield can be driven to sufficient amplitude to be able to trap large numbers of particles (>100 pC) from the background plasma and accelerate them in a narrow energy spread beam^{3–5}, now producing beams of electrons of gigaelectronvolt-scale energy of the order of 1 cm (refs 13,14).

Such electron sources are of interest to replace the accelerators that drive current synchrotron sources, and typically use multiple periods of alternately poled magnets (undulators or wigglers) to reinforce the synchrotron emission over a length of a few metres. The first demonstrations of wakefield-driven radiation using external wigglers have also been reported, though still being limited to optical or near-optical wavelengths and modest peak brightness^{15,16}.

However, the particles being accelerated in the plasma accelerator also undergo transverse (betatron) oscillations when subject to the focusing fields of the plasma wave. These oscillations occur at the betatron frequency $\omega_{\beta} = \omega_{\rm p}/\sqrt{2\gamma}$, where $\omega_{\rm p}$ is the plasma frequency and γ is the Lorentz factor of the electron beam¹⁷. As this betatron wavelength $(\lambda_{\beta} \approx 2\pi c/\omega_{\beta})$ is much smaller than the period of comparable external wigglers, again as a result of the extremely large electric fields of the wakefield, the wavelength at which these particles radiate (of order $\lambda \approx \lambda_{\beta}/\gamma^2$) can be in the X-ray region for $\gamma \gtrsim 200$, that is, for an electron beam of energy W > 100 MeV. For small transverse amplitude and thus small strength parameter $K = \gamma r_{\beta} \omega_{\beta}/c \ll 1$ (that is, undulator limit), the radiation spectrum will be narrowly peaked about a fundamental energy. As $K \rightarrow 1$, radiation also appears at harmonics. For large $K \gg 1$ (that is, large-amplitude wiggler limit) a synchrotron spectrum with broad emission consisting of closely spaced harmonics is produced¹⁷.

$$\frac{\mathrm{d}^2 I}{\mathrm{d} E \mathrm{d} \Omega} \bigg|_{\theta=0} \simeq N_{\beta} \frac{3e^2}{2\pi^3 \hbar c \epsilon_0} \gamma^2 \bigg(\frac{E}{E_{\mathrm{crit}}}\bigg)^2 \cdot \mathcal{K}_{2/3}^2(E/E_{\mathrm{crit}})$$

Here, *I* is the radiated energy, N_{β} is the number of oscillations and $E_{\text{crit}} = 3\hbar K \gamma^2 \omega_{\beta}$ is twice the energy above and below which roughly half of the total energy is radiated. $\mathcal{K}_{2/3}$ is a modified Bessel function of the second kind. For $E > E_{\text{crit}}$, the spectrum of radiated energy decays exponentially. The radiation is confined to a cone with opening angle $\theta \approx K/\gamma$.

The first measurements of betatron radiation from a broad energy spread electron beam from a laser wakefield accelerator have demonstrated X-ray emission with energies up to $\simeq 1 \text{ keV}$ with a peak brightness up to 10^{19} photons per second per mrad² per mm² per 0.1% bandwidth⁶. Here we demonstrate that by entering the highly nonlinear, so-called bubble regime of laser wakefield acceleration², we always and immediately inject a much larger amount of charge (>100 pC), which is accelerated to narrow-energy-spread beams with much higher peak energy ($\simeq 230 \text{ MeV}$ over $\simeq 5 \text{ mm}$) than in earlier experiments. This results in a 1–100 keV X-ray source with 1,000 times greater

¹The Blackett Laboratory, Imperial College London, London, SW7 2BZ, UK, ²Center for Ultrafast Optical Science, University of Michigan, Ann Arbor, Michigan, 48109, USA, ³GoLP/Inst. Plasmas and Fusão Nuclear, Laboratório Associado, Instituto Superior Técnico, Lisbon, 1049-001, Portugal, ⁴DCTI, ISCTE, Lisbon University Institute, Lisbon, 1649-026, Portugal, ⁵Laboratoire d'Optique Appliquée, ENSTA, Ecole Polytechnique, Palaiseau, 91761, France. *e-mail: stefan.kneip@imperial.ac.uk; zn1@ic.ac.uk.

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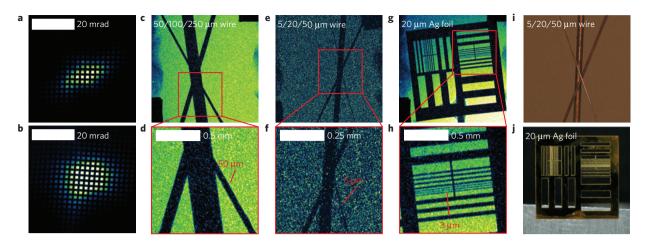


Figure 1 | **The quality of the X-ray beam is assessed by measuring its profile and imaging microscopic objects. a**,**b**, Single-shot X-ray beam profile (**a**) and sum of five consecutive shot profiles (**b**) show the imprint of a wire grid and a $4 \times 13 \text{ mrad}^2$ beam with 5 mrad pointing stability. **c-h**, X-ray radiographic images of wire triplets and a resolution test target. The objects contain features as small as $3 \mu m$, which are resolved on the radiographs, indicating a betatron source of size $\leq 3 \mu m$. **i**,**j**, Photographic images of the smallest wire triplet and resolution test target, from which the various feature sizes were obtained.

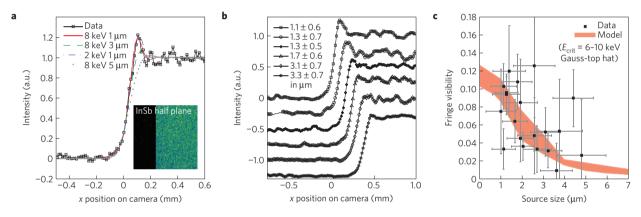


Figure 2 | The X-ray source casts a shadow of a half-plane on the detector. a, Close-up of measured intensity distribution (black squares) integrated along the edge of the half-shadow (inset) and exemplary intensity distributions using Fresnel diffraction modelling, for a source with Gaussian intensity distribution and synchrotron spectrum E_{crit}/w_x of 8 keV/1 µm (solid red), 8 keV/3 µm (dashed green), 2 keV/1 µm (dash-dotted blue) and 8 keV/6 µm (dotted grey). **b**, Series of measured intensity distributions corresponding to shots with decreasing source size w_x (from modelling) and fringe visibility. The numbers are the 95% confidence interval of the source size determined by a least-squares fit. **c**, The experimentally obtained fringe visibility is consistent with the fringe visibility obtained from the modelling, plotted for a 6-10 keV synchrotron spectrum with spatial profile of (super-) Gaussian to top-hat type. Horizontal error bars are given by the fitting procedure (typically ±30-40%) and vertical error bars are due to the noise in the experimental data (typically ±0.03-0.04).

peak brightness than previous laser-driven betatron sources^{6,7}. Furthermore, we demonstrate that the radiation has an appreciable degree of spatial coherence.

The experiment was carried out by focusing an intense shortpulse (\simeq 30 fs, \simeq 2 J) laser onto the front edge of 3, 5 and 10 mm helium gas jets (see Methods). A scintillating screen was placed in the electron beam to measure its beam profile, and a permanent magnet spectrometer to measure its energy spectrum. Electron beams with narrow energy spread features were observed from all nozzles, at electron densities of $4-22 \times 10^{18} \, \text{cm}^{-3}$. As there are consecutive phases of injection, the electron beam consists of multiple beamlets, which could be seen in both profile and spectral measurements¹⁴. For example, for an electron density of 8×10^{18} cm⁻³ on the 5 mm nozzle, electron beams of $W = (230 \pm$ 70) MeV with $\Delta W/W = (25 \pm 10)\%$ energy spread at full-width at half-maximum (FWHM) were observed with an average of 2.2 ± 0.4 beamlets per shot, with a root-mean-square (r.m.s.) divergence of $1.5 \times 1.8 \text{ mrad}^2$ and r.m.s. pointing fluctuation of $4.8 \times 4.7 \text{ mrad}^2$. We typically measure 100-300 pC of charge in the beam. The

average and maximum energy of the electron beam follow the typical wakefield electron density scaling law^{14,18}.

With the electron beam deflected away from laser axis by the spectrometer magnet, a bright (undeviated) beam of X-rays was also observed co-propagating along the laser axis. It was imperative to first prove that this X-ray source originates from the plasma itself. To do this a grid of silver wires (60 µm diameter, 310 µm separation) was placed a few centimetres from the target. X-rays originating from the interaction region project the outline of the mesh onto an imaging plate. A strongly directional beam of X-rays is evident in Fig. 1a. When either the laser power or electron plasma density was reduced to inhibit the electron beam, the X-ray beam also disappeared, showing that the generation of the X-rays is linked to the electron beam. The profile is elliptical, with a FWHM divergence of $\theta_x \times \theta_y = 4 \times 13 \text{ mrad}^2$, corresponding to a wiggler parameter $K = \theta \gamma$ of $K_x = 1.5$ and $K_y = 5$ for a simultaneously measured electron beam energy W = 200 MeV. The X-ray beam pointing is extremely stable, as can be deduced from Fig. 1b, which shows the sum of five consecutive shots. Their combined divergence is

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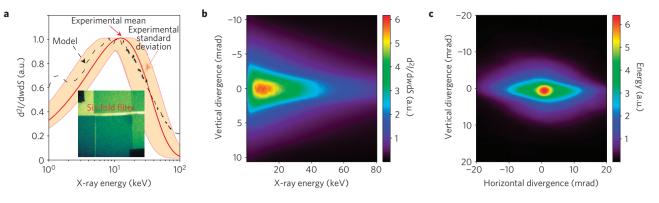


Figure 3 | Measurement and modelling of the X-ray spectrum and modelling of the X-ray profile. a, The modelled, normalized betatron spectrum agrees well with the experimentally measured spectrum obtained within one standard deviation. Inset: The experimental spectrum is inferred from measurements of the X-ray yield through various filters with known transmission. b, The angularly and spectrally resolved X-ray flux displays a peak on axis at 10 keV and a tail extending to ~100 keV. c, The spectrally integrated X-ray beam profile shows an elliptically elongated beam profile in the direction of laser polarization.

not significantly larger than that of a single-shot measurement, the r.m.s. pointing stability is 5 mrad in the horizontal and vertical direction, similar to the pointing stability of the electron beam.

To give an indication of the X-ray source size, microscopic objects were backlit with the X-ray beam. Figure 1c–h shows X-ray radiographic images of wire triplets of various sizes and of a resolution test target. Figure 1i,j gives photographic images of the smallest wire triplet and the test target. Even the smallest features of size 3 μ m are resolved, indicating that the betatron X-ray source is $\lesssim 3 \,\mu$ m, smaller than the size of the plasma wave in which the radiating electrons were trapped and oscillate ($\simeq 20 \,\mu$ m diameter).

To quantify the source size more precisely, a half-plane was backlit with the X-ray beam. A typical intensity distribution on the detector looks like a half-shadow (Fig. 2a inset), whose details convolve information about the X-ray source and half-plane. The half-plane was a 0.25-mm-thick cleaved InSb crystal (<6% transmission below 20 keV) and resembles an ideal step function.

To accurately model the shape of the intensity distribution, it is necessary to use Fresnel diffraction (see Methods), where the details of the diffraction pattern depend on the spatial and spectral intensity distribution of the source and the dimensions of the set-up. Figure 2a shows a close-up of a typical experimental and several modelled intensity distributions based on a Gaussian intensity profile and synchrotron spectrum. For the solid red curve in Fig. 2a, $E_{crit} = 8 \text{ keV}$ and a $1/e^2$ intensity radius $w_x = 1 \,\mu\text{m}$ was assumed, which best reproduced both the sharp rise and the amplitude and width of the first fringe. Changing w_x or E_{crit} under or overestimates the height and/or width of the overshoot and/or the rise, as shown by the other curves in Fig. 2.

To obtain a 95% confidence interval for the source size, a least-squares fit based on the Fresnel model was carried out with the measured intensity distributions (see Methods). Figure 2a shows exemplarily that the agreement between model and data depends only weakly on the critical energy. This was confirmed by independent measurements of the spectrum at identical experimental conditions, which found $E_{\rm crit} = 6-10$ keV, as discussed later. To accurately acknowledge the small uncertainty in critical energy and also possible deviations from a Gaussian source profile, the parameter space of the Fresnel model includes spectra with critical energies 6-10 keV and spatial profiles of (super-) Gaussian to top-hat type.

Figure 2b shows a series of experimental intensity distributions for which the source size was determined in this way. By changing the plasma density, we have some level of control over the source size. As the source size increases the visibility of the first Fresnel fringe decreases. This is in accordance with the modelling, from which it is expected that the fringes disappear for a source size $w_x \gtrsim 5 \,\mu\text{m}$. For the ringing to occur, the radiation needs to have an appreciable degree of spatial coherence at the place of the half-plane, which is merely u = 50 mm from the source. The van Cittert–Zernike theorem states that even the radiation of uncorrelated emitters with Gaussian intensity distribution can be spatially coherent¹⁹. The transverse coherence length is given by $L_{\text{trans}} = \lambda u/2\pi w_{x,y}$. In our case the relevant radiation wavelength is $\lambda \simeq 6 \times 10^{-10}$ m, which is the peak of the product of the synchrotron spectrum $E_{\text{crit}} \simeq 8 \text{ keV}$ and detector response. For a source size of $w_x = 2 \,\mu\text{m}$, $L_{\text{trans}} \simeq 3 \,\mu\text{m}$, which is enough to observe one Fresnel fringe. The spectral width of the source also reduces fringe visibility, as the temporal or longitudinal coherence length $L_{\text{long}} = \lambda^2/2\Delta\lambda \simeq \lambda$. We have assessed the amount of coherence by calculating the ex-

We have assessed the amount of coherence by calculating the experimental fringe contrast $(I_{\text{max}} - I_{\text{min}})/(I_{\text{max}} + I_{\text{min}})$ and comparing it with the prediction from the modelling of a (super-) Gaussian to top-hat source with $E_{\text{crit}} = 6-10$ keV, as shown in Fig. 2c. From this agreement we infer that the complex coherence factor μ of our beta-tron source must be close to the theoretical upper limit of $\mu = 0.88$, which would be achievable for a Gaussian intensity distribution¹⁹.

The spectral properties of the betatron radiation were determined by measuring the X-ray transmission through a set of filters (see the inset of Fig. 3a) with an X-ray CCD (charge-coupled device) detector⁷. Assuming the spectrum is synchrotron-like, the measured $E_{\rm crit}$ is (29 ± 13) keV for $n_e = (1.0\pm0.4) \times 10^{19}$ cm⁻³, and the 5 mm nozzle (see Fig. 3a). With twice the density and half the laser power, the measured $E_{\rm crit} = 6-10$ keV is found to be lower, consistent with what can be inferred from the fringe measurement ($E_{\rm crit} = 4-16$ keV) that was carried out at these conditions.

Numerical modelling was carried out using electron trajectories obtained from the fully relativistic particle-in-cell code OSIRIS. The simulations were run in the boosted frame, which, because of relativistic length contraction and time dilation, offers higher resolutions for shorter run times. The trajectories were postprocessed to yield the characteristics of the betatron radiation (see Methods). Figure 3b shows the modelled X-ray spectrum as a function of energy and angle. The X-ray flux peaks on-axis at 10 keV and extends to \simeq 100 keV. A lineout taken on-axis, as would be measured by our detector, is plotted in Fig. 3a, comparing well with the measured spectrum. A total of 10⁸ photons are predicted between 1 and 84 keV. We measure 106-108, depending on electron charge and (albeit small) pointing fluctuation. As a result of the difficulty in decoupling pointing and yield fluctuations, we take conservatively 5×10^7 photons, a $1/e^2$ source radius $w_{x,y} = 1 \,\mu\text{m}$, a divergence of $13 \times 4 \text{ mrad}^2$, a bandwidth of 100% and a pulse duration of 30 fs (refs 20,21). This gives a maximum achieved peak brightness of 1×10^{22} photons per second per mrad² per mm² per 0.1% bandwidth.

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The modelling reveals that the fastest electrons oscillate with the smallest amplitudes $r_{\beta} \lesssim 2 \,\mu$ m, consistent with the measured source sizes. *K*-parameters range from 1 to 30 with an average \simeq 7, close to the measured experimental value. Figure 3c shows the spectrally integrated X-ray beam profile obtained from the modelling for a 5 mm nozzle at $n_e = 8 \times 10^{18} \text{ cm}^{-3}$. The profile is elongated in the direction of laser polarization and the divergence $5 \times 16 \,\text{mrad}^2$ is in excellent quantitative agreement with the measurement (Fig. 1a).

The measured radiation is bright, spatially coherent and intrinsically ultrafast, opening up a multitude of advanced applications, such as phase contrast and lensless imaging, previously possible only with large conventional light sources^{22,23}. The laser plasma wiggler has the potential for making high-brightness radiation sources widespread, and thus impacting all areas of science and technology.

Methods

Laser. The experiments were carried out on the Hercules laser at the University of Michigan. Linearly polarized pulses with a central wavelength of $\lambda_0 = 800$ nm, a Gaussian FWHM pulse duration of $\tau = 32$ fs and a maximum energy of 2.3 J were focused to a peak intensity of $I = 4.7 \times 10^{19}$ W cm⁻² or normalized vector potential $a_0 = 4.7$ with an off-axis parabolic mirror of focal length f = 1 m and *F*-number of *F* = 10. The transverse intensity profile of the laser in vacuum yields a focal spot diameter of $d_{\rm FWHM} = 11.2 \,\mu\text{m}$. Typically 55% of the pulse energy is within $d_{\rm FWHM}$.

Fresnel diffraction. Fresnel diffraction occurs when a wave passes through an aperture and diffracts in the near field, causing any diffraction pattern observed to differ in size and shape, depending on the distance between the aperture and the projection *v*, the size of the aperture *A* and the wavelength of the wave λ . In the present case, $v \simeq 1,500 \text{ mm}$, $\lambda < 5 \times 10^{-7} \text{ mm} (>2 \text{ keV})$ and $a \simeq 125 \text{ µm}$ assuming an X-ray beam of 5 mrad divergence and a half-plane aperture at u = 50 mm from the source. Therefore the Fresnel number $F = A^2/(v\lambda) \simeq 16 > 1$, requiring a treatment in the Fresnel regime. *u* and *v* are the object and image distance respectively. The intensity distribution of a monochromatic point source on the projection is given by¹⁹

$$I(x) = E \cdot E^* = \frac{I_0}{2} \left\{ \left(\frac{1}{2} + \mathcal{C}(w(x)) \right)^2 + \left(\frac{1}{2} + \mathcal{S}(w(x)) \right)^2 \right\}$$

where *E* and *E*^{*} are the electric field and its complex conjugate, *C* and *S* are the Fresnel functions, $w(x) = x(u/(v+u)\sqrt{2(1/u+1/v)}/\lambda)$ and *x* is the position on the projection. If the source is not point-like, the electric field has to be convolved with the source function and spectral distribution $E(x) = \int_y \int_\lambda R(\lambda)g(y)E(\lambda, x - v/u \cdot y) \, dy \, d\lambda$ to compute I(x), where $R(\lambda)$ is the spectral response of the detector. The Fresnel intensity distribution can be modelled with a simple error function, which, for a given spectrum and profile, results in a relationship between the source size from the experimental data by applying a least-squares method.

Numerical modelling. Three-dimensional numerical simulations were carried out with the particle-in-cell code OSIRIS (ref. 24), in which a linearly polarized pulse with experimental parameters is focused 0.25 mm into the plasma. The electron plasma density profile increases linearly from zero to $n_e = 8 \times 10^{18}$ cm⁻³ in the first 0.5 mm, is then constant for 3.5 mm and finally falls linearly to zero in 0.5 mm. Simulations were carried out in a relativistically boosted frame ($\gamma = 5$), which allows for significant computational gains^{25,26}. The simulation box corresponds to $60 \times 94 \times 94 \,\mu\text{m}^3$ in the laboratory frame. A total of 9.2×10^8 particles (2 per cell) were pushed for 5×10^3 iterations. The resolution in the laser propagation direction z is $k_0 \Delta z = 0.11$, and $k_p \Delta x = k_p \Delta y = 0.16$ in the transverse directions. To obtain the radiation emission from the simulation, a post-processing diagnostic is used on a set of trajectories of injected electrons²⁷. The position and momentum history of the electrons is used to deposit the radiated fields on a virtual detector.

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Author contributions

The experiment and analysis was carried out in the main by S.K. and C.M. with contributions from C.B., F.D., C.H., S.P.D.M., T.M., S.R.N., and C.A.J.P. Simulations were carried out by J.L.M., S.F.M., R.F. and L.O.S., V.Y., G.K. and V.C. operated the laser, S.K., K.K., S.P.D.M., A.M., Z.N., K.T.P., J.S., A.G.R.T. contributed to planning, interpretation and manuscript preparation.

Additional information

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