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# CFL3D User's Manual (Version 5.0)

Sherrie L. Krist BaNANePOS, Inc., Hampton, Virginia

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# Nomenclature

Symbol:	Definition:
<b>A</b> , <b>B</b> , <b>C</b>	flux Jacobians
a	speed of sound
b	part of viscous term in energy equation; also, compression parameter in minmod flux limiter
С	constant
$C_{\mu}$	modeling variable in turbulence model equations
D	diffusion term in turbulence model equations
$D_{1}, D_{2}$	damping terms in Baldwin-Barth turbulence model
d	distance to nearest wall; also, directed distance to the wall
Ε	total energy
e	total energy per unit volume
F	function
f	damping function; also, frequency
<b>F</b> , <b>G</b> , <b>H</b>	inviscid fluxes
$\mathbf{F}_{v}, \mathbf{G}_{v}, \mathbf{H}_{v}$	viscous fluxes
$G_1$	modeling variable (similar to $-C_{\mu}$ term) in turbulence model equations
Н	total enthalpy
Ι	restriction operator
Ι	identity matrix
i, j, k	grid indices
J	transformation Jacobian, $J = \partial(\xi, \eta, \zeta) / \partial(x, y, z)$
k <sub>r</sub>	reduced frequency
k <sub>g</sub>	growth rate

N

Symbol:	Definition:
k	kinetic energy in turbulence model equations
$ ilde{L}$	characteristic length
L <sub>ref</sub>	length in grid corresponding to $\tilde{L}$
$\tilde{L}_R$	reference length used by code $\tilde{L}_R = \tilde{L}/L_{ref}$
l	reference length; also, length scale in Baldwin-Lomax turbulence model
М	Mach number, $M =  \tilde{\mathbf{V}}  / \tilde{a}$
Μ	transformation matrix from conserved variables to primitive variables, $\partial Q/\partial q$
т	sub-iteration counter
Ν	spatially-factored implicit matrix term
n	direction normal to the wall
n	current iteration
Р	production term in turbulence model equations
Pr	Prandtl number
р	static pressure
Q	conserved variables
q	primitive variables
ġ	heat flux terms
R	residual vector
R	turbulent Reynolds number term in Baldwin-Barth turbulence model; also, residual term
$Re_{\tilde{L}}$	Reynolds number, $Re_{\tilde{L}} = \tilde{\rho}_{\infty}  \tilde{\mathbf{V}} _{\infty} \tilde{L} / \tilde{\mu}_{\infty}$
S	mean rate-of-strain tensor
S <sub>P</sub>	production source term in turbulence model equations
S <sub>D</sub>	destruction source term in turbulence model equations
S	entropy; also, parameter used in smooth flux limiter
Т	matrix of eigenvectors
t	time; also, parameter used in smooth flux limiter
U, V, W	contravariant velocities

Symbol:	Definition:
<i>u</i> , <i>v</i> , <i>w</i>	Cartesian velocities in $x, y, z$ directions
<i>u</i> <sup>+</sup>	law-of-the-wall variable, $u^+ = \tilde{u} \sqrt{\tilde{\rho} / \tilde{\tau}_w}$
V	corrections on coarser meshes, used to update finer mesh in the multigrid algorithm
V	velocity vector, $(u, v, w)$
$ \mathbf{V} $	total velocity
W	mean vorticity tensor
Х	represents either k, $\omega$ , or $\epsilon$ in general turbulence model equations
<i>x</i> , <i>y</i> , <i>z</i>	Cartesian coordinates
<i>y</i> <sup>+</sup>	law-of-the-wall variable, $y^+ = \sqrt{\tilde{\rho}\tilde{\tau}_w}\tilde{y}/\tilde{\mu}$
α	angle of attack; also, used as constant in turbulence model equations
β	side-slip angle; also, used as constant in turbulence model equations
γ	ratio of specific heats, $\gamma = 1.4$ ; also, variable in turbulence model equations
$\Delta$	incremental quantity; also, forward difference operator
δ	difference operator
8	dissipation term in turbulence model equations; also, small constant used in flux limiters
к	spatial differencing parameter; also, von Karman constant used in turbu- lence model equations
Λ	matrix of eigenvalues
λ	bulk viscosity coefficient
μ	molecular viscosity coefficient
ν	kinematic viscosity
Ŷ	field equation variable in Spalart-Allmaras turbulence model
ξ, η, ζ	general curvilinear coordinates; also, $\eta$ and $\zeta$ used as variables in EASM turbulence model equations
ρ	density
τ	shear stress tensor; also, relative truncation error
φ	parameter governing the temporal order of accuracy of the scheme
Ω	magnitude of vorticity

#### Nomenclature

Symbol:	Definition:
ω	rotational velocity; also, variable in turbulence model equations,
	$\omega = \varepsilon/k$
$\nabla$	gradient operator $\nabla X = \partial X / \partial x_i$

# Subscripts

Symbol:	Definition:
е	denotes estimated value
i, j, k	denote grid indices; also, summation convention where specified
inv	denotes inviscid part
k	denotes k turbulence model quantity
L	denotes left-hand state; also, denotes laminar quantity where specified
l	denotes summation convention
R	denotes right-hand state
Т	denotes turbulent quantity
t	denotes total quantity; e.g. $p_t \Rightarrow$ total pressure; also, denotes differentia-
	tion with respect to time
v	denotes viscous term
W	denotes wall condition
<i>x</i> , <i>y</i> , <i>z</i>	denote differentiation with respect to $x, y, z$ ; $x$ also denotes tensor notation where specified
ε	denotes $\varepsilon$ turbulence model quantity
ω	denotes $\omega$ turbulence model quantity
$\infty$	denotes reference conditions, typically free-stream conditions
+, -	denotes forward or backward difference operator
ξ, η, ζ	denotes reference to a particular coordinate direction

# Superscripts

Symbol:	Definition:
b	denotes biased gradient
С	denotes correction term

## **Superscripts**

Symbol:	Definition:
r	denotes residual smoothing term
٨	denotes quantities in generalized coordinates
~	denotes dimensional value; also, denotes Roe-averaged variable where specified
,	denotes partial derivative with respect to $\mathbf{q}$ and intermediate values in the time advancement scheme
$\rightarrow$	denotes a vector quantity
+, -	denotes forward or backward difference operator

### Abbreviations

CFD	Computational Fluid Dynamics
CFL	Courant number
CFL3D	Computational Fluids Laboratory 3-Dimensional (flow solver)
CPU	Central Processing Unit
EASM	Explicit Algebraic Stress Model
FAST	Flow Analysis Software Toolkit <sup>44</sup>
LRR	Launder-Reese-Rodi
GRIDGEN	GRID GENeration <sup>36</sup>
MaGGiE	MultiGeometry Grid Embedder <sup>11</sup>
NACA	National Advisory Committee for Aeronautics
PLOT3D	PLOT 3-Dimensional <sup>43</sup>
RAE	Royal Aircraft Establishment
SSG	Speziale-Sarkar-Gatski
SST	Shear Stress Transport
TLNS3D	Thin-Layer Navier-Stokes 3-Dimensional (flow solver) <sup>42</sup>



# CHAPTER 1 Introduction

Welcome to CFL3D (Version 5.0), a Reynolds-Averaged thin-layer Navier-Stokes flow solver for structured grids. The original version of CFL3D was developed in the early 1980's in the Computational Fluids Laboratory at NASA Langley Research Center; hence the name of the code, which is an acronym for the Computational Fluids Laboratory <u>3</u>-Dimensional flow solver. As the number of people who have utilized CFL3D has increased over time, so have the demands on the code. Consequently, it is constantly "under construction" with numerous researchers having contributed to code upgrades. Currently, the primary developers/supporters of CFL3D are Dr. Christopher L. Rumsey and Dr. Robert T. Biedron of the Aerodynamics and Acoustics Branch at NASA Langley. An overview of the history of the code can be found in Rumsey, Biedron, and Thomas.<sup>32</sup>

CFL3D solves the time-dependent conservation law form of the Reynolds-averaged Navier-Stokes equations. The spatial discretization involves a semi-discrete finite-volume approach. Upwind-biasing is used for the convective and pressure terms, while central differencing is used for the shear stress and heat transfer terms. Time advancement is implicit with the ability to solve steady or unsteady flows. Multigrid and mesh sequencing are available for convergence acceleration. Numerous turbulence models are provided, including 0-equation, 1-equation, and 2-equation models. Multiple-block topologies are possible with the use of 1-1 blocking, patching, overlapping, and embedding. CFL3D does not contain any grid generation software. Grids must be supplied extraneously.

Version 5.0 of CFL3D has several additional utilities over earlier versions of the code. Most notably, Version 5.0 has the capability to employ sliding patched-zone interfaces, such as might be required to perform rotor-stator computations, for example. However, it is stressed here that CFL3D has been developed primarily as a tool for external aerodynamics analysis. Its use for internal turbomachinery applications has been only as a basic research code thus far; other Navier-Stokes codes specifically designed for turbomachinery applications, such as ADPAC<sup>22</sup>, UncleTurbo<sup>23</sup>, or ROTOR<sup>29</sup> may be better suited to the analysis of such flows.

The purpose of this user's manual is to describe the code and provide instruction for its usage. Chapter 2 explains the set up and running of the code. Various files available with CFL3D are described and directions for their usage are provided. Step by step instructions are listed for running CFL3D, as well as the pre-processing codes needed for grid-overlapping and patching.

Chapter 3 provides a line-by-line description of all the input parameters utilized in CFL3D. This will probably be the chapter referred to the most. It will aid users in understanding the sample problems as well as in setting up their own problems. While some parameters are strictly case dependent, some are fairly general and recommended values are usually indicated. Certain problems may, of course, require altering the recommended values of these parameters as well.

To avoid the use of feet, meters, pounds, grams, etc. in coding the equations, CFL3D solves the Navier-Stokes equations in "nondimensional" form. Each flow-field parameter is nondimensionalized by reference values. For example, all points on a wing may be non-dimensionalized by the chord length of the wing. The nondimensionalizations used in CFL3D are described in Chapter 4.

Chapter 5 explains the file formats. The most important of these is the grid file format, since users will need to translate their grids into an appropriate CFL3D format. The restart file format may be needed for user-designed post-processing programs. This chapter also describes the various output files.

Besides providing general information about boundary condition usage in CFL3D, Chapter 6 discusses the various physical boundary conditions available. It also provides descriptions of the multiple block capabilities in CFL3D, mainly, 1-1 blocking, patching, grid overlapping, and grid embedding. Input examples for the various techniques are provided.

The advantages of using multigrid for convergence acceleration are explained in Chapter 7. The theory behind multigrid and its usage are discussed. The use of multigrid with embedded grids is described. Mesh sequencing is also defined and discussed. Input examples are provided there as well.

Chapter 8 discusses time accurate simulation of unsteady flows using sub-iteration schemes. Two types of sub-iterations are currently implemented in CFL3D. The effects of the different types of sub-iterations, as well as the strategy for pursuing time-accurate computations in general, are described.

Chapter 9 contains several test cases for the user to practice with when learning how to run CFL3D. Two-dimensional sample cases include a single block RAE airfoil case, a grid-overlapping NACA 0012 airfoil case, a patched grid NACA 0012 airfoil case, and a grid-overlapping multielement airfoil case. Also described are examples for a fixed flat plate, a vibrating plate, a multistream nozzle, and a rotor-stator. The three-dimensional test cases include an axisymmetric bump, a single block F-5 wing, a single block ONERA M6 wing, and a single block delta wing.

Finally, Chapter 10 is dedicated to troubleshooting. It is basically a compilation of suggestions for common problems encountered by users over the years. It is recommended that this chapter be read at the onset of any difficulties which occur while using CFL3D. Precious time may be saved by learning from the experiences (and, yes, the mistakes!) of

others. It is prudent at this point to suggest to the user that any problems encountered while using CFL3D be reported to the support personnel so that others may benefit from hitherto unknown experiences that may occur.

There are several appendices at the end of this manual which provide more detailed information about the code. Appendix A describes the governing equations. Appendix B explains the time advancement procedure. The spatial discretizations of the inviscid and viscous fluxes are described in Appendix C. The multigrid algorithm is described in Appendix D. The matrix needed to convert the equations from conservative variables to primitive variables is derived in Appendix E. A step-by-step explanation of how to convert the governing equations from Cartesian coordinates to generalized coordinates is provided in Appendix F. Appendix G shows how the forces and moments are calculated. Appendix H describes the turbulence models available in CFL3D. Appendix I illustrates the angle ( $\alpha$  and  $\beta$ ) definitions used in the code. Appendix J contains a list of all the subroutines in CFL3D and their purposes. Finally, Appendix K explains the differences between Versions 4.1 and 5.0 of CFL3D.

The last two sections of the manual are source listings. Beginning on page 337, there is a partial list of papers that have been published over the years as CFL3D has developed. The last section of the manual is a list of the references used for this document.

It is appropriate at this point to discuss some of the terminology used in the CFL3D *User's Manual.* Terms, such as grid, zone, mesh, and block, are often used interchangeably in CFD discussions and literature. For clarity in this manual, the following distinctions are defined. The *grid* refers to those sets of points that define the flow field and are generated by the user to be read in by CFL3D. A grid may be one entity or it may be composed of several component grids. These component grids may also be called *grid zones*. The grid zones may communicate with one another through 1-1 blocking, patching, overlapping, or embedding. The term *block* is used for bookkeeping purposes. All component grids are blocks and are assigned a block number. The grid or the set of component grids is often referred to as the *global* grid, encompassing the entire flow field. When the component grids are coarsened for such options as multigrid and mesh sequencing, the coarser levels are labelled as blocks and are also assigned block numbers. *Embedded* grids are finer than the component grid in which they reside, but do not encompass the entire zone. Instead, they are placed in regions known to have high gradients to further resolve those areas.

Another terminology definition that should be clarified is the phrase *free stream*. As used throughout this manual, the words *free stream* imply a reference state. CFL3D was developed primarily as an external flow solver, in which case the appropriate reference state *is* free stream. However, for purely internal flows, the concept of free stream has no meaning and the more general concept of *reference state* should be used.

Further distinctions between terms are required in discussing the methods of communication between blocks. When two blocks share a face or a portion of a face and the grid points correspond point to point, the boundary condition communication set up between the two blocks is called 1-1 blocking. The blocks in this case are often referred to as being  $C^0$  continuous. Grid patching, on the other hand, refers to the boundary condition interpolations set up between blocks that share a common face or portion of a face, but which do not match point to point. For example, a fine grid face of one block could pass flow information to a coarse grid face of another block as long as the two blocks shared the face. Grid overlapping has neither the restriction of point to point connectivity or a common face between blocks. For example, a cylindrical "inner" grid could communicate with a cubic "outer" grid even though the grid topologies are completely different. The terms overset grids and chimera grids are synonymous with overlapped grids.

A distinction should also be made for the methods of time advancement. The term *time accurate* means that the flow is actually tracked in time. For example, if the flow is allowed to advance for ten seconds, the flow conditions at a certain point in the flow field might have ten different values, one at each second. For a *steady state* problem, after convergence is reached, the flow conditions will not change with time.

Here is one final tip for reading this manual. In discussions and examples, input variables are typed in **bold** print. Such items as variable names, array names, lines of actual code, input samples, etc. are printed in the Courier font. Note that these conventions do not necessarily convey in the World-Wide-Web version of this manual. The CFL3D internet address is: http://fmad-www.larc.nasa.gov/~rumsey/CFL3D/cfl3d.html .

Hopefully, this manual will answer most questions encountered when running CFL3D. If problems arise that are not addressed in the sample inputs or in Chapter 10, please contact

Dr. Christopher L. Rumsey (757)864-2165 c.l.rumsey@larc.nasa.gov

or

Dr. Robert T. Biedron (757)864-2156 r.t.biedron@larc.nasa.gov

NASA Langley Research Center Mail Stop 128 Hampton, VA 23681-0001

Please also let these representatives know if any "bugs" or other useful insights are discovered during the usage of CFL3D.

Every effort has been made to insure that this document represents the correct usage and theoretical underpinnings of CFL3D. The intention has been to make this manual as complete and as self-consistent as possible. However, no guarantee can be given that this manual covers all possible aspects of the code and its usage or that this manual is without error.



#### **CHAPTER 2**

While the ultimate goal in utilizing CFL3D will be to solve problems pertinent to the user's work, it is highly recommended for new users to begin by running some of the test cases described in Chapter 9. Experience gained while running these will help the user when running a "real" problem. The following sections provide general instructions for compiling and running CFL3D.

## 2.1 Acquiring the Code and Example Files

The files needed to run CFL3D are located on Vonneumann and Eagle, two Cray supercomputers located at NASA Ames Research Center. The files are in

```
~rumsey/Cfl3dv5/cfl3dv5.ascii
```

The files are tarred, compressed, encrypted, and uuencoded. A keyword (obtained from the Aerodynamic and Acoustic Methods Branch at NASA Langley – see below) must be used to decode the file cfl3dv5.ascii. The steps for this procedure are as follows:

<u>Step 1</u>

```
uudecode cfl3dv5.ascii
```

# Step 2

crypt keyword < cfl3dv5.crypt > cfl3dv5.tar.Z

# <u>Step 3</u>

uncompress cfl3dv5.tar.Z

# <u>Step 4</u>

tar xvf cfl3dv5.tar

This step should create a directory named Cfl3dv5 with the appropriate CFL3D files in it.

### Step 5

rm cfl3dv5.ascii cfl3dv5.tar cfl3dv5.crypt

If the test cases are not required, a more convenient option might be to obtain the file cfl3dv5.codeonly.ascii. This file is also tarred, compressed, encrypted, and uuen-coded, but it is much smaller than cfl3dv5.ascii.

If access to Vonneumann or Eagle is not possible, the encrypted codes are also available on the Langley CFD anonymous ftp site. To obtain the code in this manner, take the following steps:

#### <u>Step 1</u>

```
ftp tabdemo.larc.nasa.gov
```

#### or

```
ftp 128.155.24.42
```

### Step 2

When prompted for the username, type

anonymous

#### Step 3

Type the user's e-mail address as the password.

#### Step 4

cd incoming

#### Step 5

cd Rumsey

#### <u>Step 6</u>

get cfl3dv5.ascii

#### or

get cfl3dv5.codeonly.ascii

To obtain the current keyword needed to decode the source files, send mail or e-mail to Dr. Christopher L. Rumsey (see page 4) requesting the code, along with an explanation describing the planned usage of the code. Include a FAX number with the request. All users will be asked to fill out and sign a CFL3D Usage Agreement form. After the completed form is received, the keyword will be provided.

After obtaining CFL3D, it is required that the user remember not to distribute any part of the code to others outside of his or her own working group without prior permission from NASA Langley. The CFL3D code is currently restricted to use within the United States only.

#### 2.1.1 The Code and Supplementary Files

The items listed below are made available to the user when CFL3D is obtained. The names in **bold** face indicate a directory.

2dtestcases	isrcheq_wkstn.f	makeprecfl3d_hp_sngl
3dtestcases	lbcx.f	makeprecfl3d_rs6000
Advice	makecfl3d_cray	makeprecfl3d_sgi
Maggie	makecfl3d_decalpha	makeprecfl3d_sgi_R10000
Multitask_cray	makecfl3d_hp_dbl	makeprecfl3d_sgi_R8000
README	makecfl3d_hp_sngl	makeprecfl3d_sun
Ronnie	makecfl3d_rs6000	plot3d_hp_dbl.f
Tools	makecfl3d_sgi	precfl.h
af3f.f	makecfl3d_sgi_R10000	precfl3d.f
bc.f	makecfl3d_sgi_R8000	rhs.f
cbsem.f	makecfl3d_sun	second_hp_dbl.f
cvmgt_wkstn.f	makeprecfl3d_cray	second_rs6000.f
dynptch.f	makeprecfl3d_decalpha	second_wkstn.f
input.doc	<pre>makeprecfl3d_hp_dbl</pre>	turbs.f

The CFL3D code consists of one main driver subroutine package, cbsem.f, and six sets of library subroutines, bc.f, rhs.f, af3f.f, lbcx.f, turbs.f, and dynptch.f. Cbsem.f contains the main subroutine (mgblk) of the program as well as many of the routines needed for the input and output of information. The file bc.f contains the boundary condition subroutines, including physical boundary conditions, 1-1 blocking, and grid embedding. It also contains the necessary routines for processing supplemental overlapped-grid and static patched-grid interpolation information. Rhs.f contains the subroutines needed for solving the right-hand side of the governing equations, while af3f.f contains the subroutines needed for calculating the left-hand side of the governing equations. (See Appendix A.) Lbcx.f contains a variety of subroutines, including those needed for multigrid and mesh sequencing, those providing the metrics and other grid information, those involved with computing the forces, and several others. The file turbs.f contains the subroutines needed for dynamic grid patching.

Some supplementary files are also needed for CFL3D. A "makefile" is used to compile and link subroutines and to create an executable for the code. A preprocessing program called precf13d.f is used to determine the memory and array dimensions needed by CFL3D for a particular case. For precf13d.f, use makeprecf13d\_machinename. Here, and in the discussions that follow, the terms in italics pertain to user-specific items. For instance, *machinename* might be cray, sgi, or sun for the Cray supercomputer, SGI, and Sun workstations, respectively. Precfl.h is the parameter file required by this makefile. The array dimensions set in precfl.h should be large enough for nearly any problem and should rarely need modification; if a larger value of a particular parameter is required, precfl3d will stop and request the user to increase the relevant parameter.

A makefile called makecf13d\_machinename is also needed to compile and link the CFL3D subroutines and create the executable, cf13d. Several files are needed to compile the code on workstations. These files contain Fortran source code for the corresponding Cray intrinsic functions. For instance, when utilizing makecf13d\_rs6000, make sure that second\_rs6000.f is available. Likewise, when using makecf13d\_hp\_db1, make sure that plot3d\_hp\_db1.f and second\_hp\_db1.f are in the working directory. Note that for double precision on the HP, the user must actually *remove* the subroutines plot3d, plot3c, and plot3t from cbsem.f prior to compiling with makecf13d\_hp\_db1. For all other makecf13d makefiles, cvmgt\_wkstn.f, isrcheq\_wkstn.f, and second\_wkstn.f are utilized and therefore must be available. (Note that wkstn is short for "workstation".)

The file input.doc contains the information found in Chapter 3. It is a listing and description of the input parameters used in CFL3D.

The files in directory 2dtestcases are:

0012_patch	Flatplate	Rae10
0012_xmera	Multielem	Rotorstator
README.2d	Multistream	Vibrate

The files in directory **3dtestcases** are:

README.3d	Delta	Oneram6
Axibump	F5wing	

Directory 2dtestcases contains various two-dimensional test cases. A patching example involving the NACA 0012 airfoil is located in 0012\_patch. In 0012\_xmera, a case utilizing grid overlapping for the same airfoil is found. A three-element airfoil case involving grid overlapping is set up in Multielem. Directory Rae10 contains a single block case for the RAE 2822 airfoil. Flatplate contains the files needed to run the flat plate test case. A multistream nozzle case is in Multistream. In vibrate are the files for a vibrating plate case. A rotor-stator case is set up in Rotor.

The three-dimensional test cases are located in the directory **3dtestcases**. A case solving the flow over an axisymmetric bump is in **Axibump**. Other cases currently available are for a delta wing (in **Delta**), an F-5 wing (in **F5wing**), and an Onera M-6 wing (in **Oneram6**). The steps for running these test cases are described in Chapter 9.

The files in directory Advice are:

cfl3dadvice.give	cfl3d.references	v4_to_v5.inputdif
cfl3dadvice.grid	cfl3d.turb.references	

The first four files contain the same information given in Chapter 10 and in "CFL3D Papers" on page 337. The v4\_to\_v5.inputdif file describes the differences between the Version 4.1 and Version 5.0 input files. (This information is also given in Appendix K.)

The files in directory Maggie are:

README	maggie.doc	makemaggie_sgi
ismax_wkstn.f	maggie.f	makemaggie_sun
ismin_wkstn.f	makemaggie_cray	second_wkstn.f
mag1.h	makemaggie_hp_sngl	

The files in the Maggie directory are needed for cases involving grid overlapping. The chimera scheme code is called maggie.f. Maggie.doc explains the input parameters for maggie.f. Makefiles for various machines are available and require magl.h, which contains the appropriate dimensions for the problem at hand. Make sure ismax\_wkstn.f, ismin\_wkstn.f, and second\_wkstn.f are available before running MaGGiE on a workstation.

The files in Multitask\_cray are:

README	makecfl3d_cray_multi	xlim_multi.f
fhat_multi.f	tinvr_multi.f	

These files should be used in conjunction with the standard CFL3D files when Cray multitasking is desired.

The files in directory Ronnie are:

README	makeronnie_sgi	ronnie.f
makeronnie_cray	makeronnie_sun	<pre>second_wkstn.f</pre>
makeronnie_decalpha	ron1.h	
makeronnie_hp_sngl	ronnie.doc	

When utilizing grid patching, the files in the **Ronnie** directory are needed. The code that sets up the patching interpolation stencils is called ronnie.f. Ronnie.doc explains the input parameters for ronnie.f. Makefiles for various machines are available and utilize ron1.h in the compilation. The user must set the parameters in ron1.h for each particular case. Also, make sure second\_wkstn.f is available before running ronnie on a workstation.

The files in the **Tools** directory are:

README	makeltol.f	reverseijk.f
everyother.f	<pre>make_p3dtotec3d_cray</pre>	v4tov5_input.f
everyotherp3d.f	p3dtotec3d.f	v4tov5_restart.f
historytec3d.f	p3dtotec3d_pre.c	v5inpdoubhalf.f

These programs are useful for converting files from one format to another. The code everyother.f takes an existing CFL3D-type grid and coarsens it by taking every other grid point. The everyotherp3d.f program performs the same task for PLOT3D-type grids. The file historytec3d.f reads in the residual history file and creates a corresponding Tecplot<sup>4</sup> file. It will also convert iterations to work units if desired. The make1to1.f program assists in the creation of the boundary condition and one-to-one blocking sections of the input file. The p3dtotec3d.f tool (along with p3dtotec3d\_pre.c and make\_p3dtotec3d\_cray) creates a Tecplot file from PLOT3D grid and solution files. The reverseijk.f code reads a CFL3D-type grid file, swaps indices (as desired), and writes out a new CFL3D-type grid file. The v4tov5\_input.f tool takes a Version 4.1 input file and converts it to Version 5.0 form. The v4tov5\_restart.f tool converts a Version 4.1 restart file to Version 5.0 form. Finally, v5inpdoubhalf.f takes an existing input file and creates a new input file with either double or half the grid points of the original.

#### 2.2 Obtaining a Grid File

The first step in solving a CFD problem is obtaining a grid. A pre-packaged tool such as GRIDGEN<sup>36</sup> could be used to create a grid, or perhaps a user-written program could be used for relatively simple configurations. How ever the grid is obtained, an essential step toward success with CFL3D will be having the grid file written in CFL3D or PLOT3D/TLNS3D format. See "Grid File" on page 65 for the appropriate formats for the file. Time can also be saved at this point by considering what options will be desirable later. For example, since the use of multigrid is highly recommended, choose grid dimensions wisely. See "A Word About Grid Dimensions" on page 127. If viscous cases will be run, make sure the grid spacing normal to the wall in the boundary layer is fine enough. At least 20 grid points are recommended in a *laminar* boundary layer. At least 3 points are

recommended in the laminar sub-layer of a *turbulent* boundary layer  $(y^+)$  of first grid point off the wall should be O(1). Basically, consider how the problem will be solved and choose dimensions, blocking strategies, grid-line stretching, etc., accordingly.

#### 2.3 Creating the Input File

Chapter 3 lists and describes all the parameters in the input file. The easiest way to create an input file for a particular case is to copy the input file from the test case (see Chapter 9) most resembling the case at hand. Then change those parameters that pertain to the current case. The other parameters should already be set to the recommended values. Besides the test case sample inputs, some input examples for 1-1 blocking and grid embedding are provided in Chapter 6 and several input examples for multigrid, mesh sequencing, and grid embedding are provided in Chapter 7. CFL3D provides additional help in setting up the input file with an extensive set of diagnostics which halt execution at detectable errors and provide the user with a message indicating the problem.

#### 2.4 Utilizing Patched and/or Overlapped Grids

When using patched and/or overlapped grids, the files containing the appropriate interpolation coefficients must be obtained prior to running CFL3D. The preprocessors ronnie and MaGGiE<sup>11</sup> will generate the needed coefficients for patched grids and overlapped grids, respectively, and will output the interpolation information in the format needed by CFL3D. Information on using ronnie and MaGGiE can be found in the ronnie.doc and maggie.doc files, respectively. The preprocessor precfl3d, which sets all the array size parameters needed by CFL3D, will need the appropriate patch and/or overlap data files in order to properly size the arrays needed for these options. Therefore, be sure to run ronnie and/or MaGGiE *before* running precfl3d.

#### 2.4.1 Running ronnie

When utilizing the grid patching option of CFL3D, the preprocessor ronnie must be run first. The basic steps for running this code are as follows.

#### <u>Step 1</u>

Prepare an input deck, typically called ronnie.inp, for the case. Also, modify the header file ron1.h as needed.

#### Step 2

Use the makefile to compile, link, and create the executable for the ronnie code (be sure ron1.h is in the current directory):

make -f makeronnie\_machinename

#### <u>Step 3</u>

Run the ronnie code:

ronnie < ronnie.inp</pre>

If successful, ronnie will create a file containing patch interpolation data. The name of the patch file is specified by the user in the input file; it is typically called patch.bin.

#### 2.4.2 Running MaGGiE

If grid overlapping is used in some portion of the grid configuration, then the preprocessor MaGGiE must be utilized. The basic steps for running this code are as follows.

#### <u>Step 1</u>

Prepare an input deck, typically called maggie.inp, for the case. Also, modify the header file mag1.h as needed.

#### Step 2

Use the makefile to compile, link, and create the executable for the MaGGiE code (be sure mag1.h is in the current directory):

```
make -f makemaggie_machinename
```

#### Step 3

Run the MaGGiE code:

maggie < maggie.inp</pre>

If successful, MaGGiE will create a file containing overlap interpolation data. The name of the overlap file is specified by the user in the input file; it is typically called ovrlp.bin.

#### 2.5 Running CFL3D

The basic steps for running CFL3D are as follows. Remember, if using grid patching and/or grid overlapping, run ronnie and/or MaGGiE before proceeding with the following steps.

#### Step 1

Compile, link, and create the executable for the precfl3d.f code. Make sure the parameter file precfl.h is available.

make -f makeprecfl3d\_machinename

The resulting executable will be called precf13d.

#### <u>Step 2</u>

Run precfl3d.

precfl3d < inputfilename</pre>

The *inputfilename* here is the CFL3D input file name. Detailed diagnostic information is printed out in the file precfl3d.out. Most items in the input deck will be checked for

consistency. However, since precf13d does not read in the grid file, it cannot detect grid problems and, most importantly, it cannot detect ordering errors in the 1-1 block interfaces. Check the output from CFL3D itself to verify that 1-1 block interfaces are set correctly. If precf13d has not run successfully, then the most likely cause is an error in the input file; precf13d echoes the input file as it goes along, so check the bottom of precf13d.out to determine the approximate location of the input error. A successful execution of precf13d will have one of the following lines at the end of the output file and print to the screen:

you \*MUST\* recompile cfl3d

or

you do not need to recompile cfl3d

When the code has run successfully, there will be no further need for this program until a new input file is developed.

Running precf13d will either create or modify the parameter files, cf11.h, cf12.h, cf13.h, cf14.h, and cf15.h, which are used when compiling CFL3D. Remember that these files *are* case dependent. Therefore, another case will generally require that precf13d be run again and CFL3D recompiled.

#### Step 3

Compile, link, and create the executable for CFL3D.

make -f makecfl3d\_machinename

The executable will be called cfl3d.

#### <u>Step 4</u>

Run cfl3d. This step can be performed interactively or a submittal file can be utilized to send the case to a particular queue of the machine being used.

```
cfl3d < inputfilename
```

Users have been known to submit cases for large numbers of iterations on the first run only to discover that an input parameter was set incorrectly. It is wise to begin running CFL3D with only a few iterations and then check the output. Look over the main output file. Make sure there are no warning messages and that, in general, it "looks right" for the particular case being run. Take the PLOT3D files to a Silicon Graphics (or compatible) workstation and check any block boundaries to see if the flow is passing from block to block as expected. *Then* submit the case for a more extensive computation.



#### **CHAPTER 3**

Descriptions of each of the input line types are provided in this chapter where the section numbers denote the line type (LT) number referred to throughout the manual. An example input file is provided below for a flat plate case. Note that not all of the line types of information will be required for every application.

<pre>I/O FILES grdflat5.bi plot3dg.bin plot3dq.bin cfl3d.out cfl3d.res cfl3d.turre cfl3d.bloma cfl3d.out15 cfl3d.prout cfl3d.out20 ovrlp.bin patch.bin restart.bin</pre>	s x						
turbule XMACH	nt flat pl ALPHA	ate (plate. BETA	from j=17 REUE,MIL	-65, prior TINF,DR	to 17 is IALPH	symmetry) IHSTRY	
0.2000	00.000	0.0	06.000	460.0	IALPH 0	0	
SREF	CREF	BREF	XMC	YMC	ZMC	-	
1.00000	1.00000	1.0000	0.00000	0.00	0.00		
DT	IREST	IFLAGTS	FMAX	IUNST	CFLTAU		
-5.000 NGRID	0 NPLOT3D	000 NPRINT	05.0000 NWREST	0 ICHK	10.0 I2D	NTSTEP	ITA
NGRID 1	NFL013D	NFRINI 0	1200	0	120	NISIEF 1	1
NCG	IEM	IADVANCE	IFORCE	IVISC(I)	IVISC(J)	IVISC(K)	±
2	0	0	001	0	0	7	
IDIM		DIM					
02	65	97					
ILAMLO 1	ILAMHI 2	JLAMLO 1	JLAMHI 17	KLAMLO 1	KLAMHI 97		
INEWG	IGRIDC	IS	JS	KS	IE	JE	KE
0	0	0	0	0	0	0	0
IDIAG(I)	IDIAG(J)	IDIAG(K)	IFLIM(I)	IFLIM(J)	IFLIM(K)		
1	1	1	0	0	0		
IFDS(I)	IFDS(J)	IFDS(K)	RKAPO(I)	RKAPO(J)	RKAPO(K)		
1 GRID	1 NBCIO	1 NBCIDIM	0.3333 NBCJ0	0.3333 NBCJDIM	0.3333 NBCK0	NBCKDIM	IOVRLP
GRID 1	1	NBCIDIM 1	NBC00	NBCODIM 1	2	NBCRDIM 1	0
IO: GRID	SEGMENT	BCTYPE	JSTA	JEND	KSTA	KEND	NDATA
1	1	1001	0	0	0	0	0
IDIM: GRID	SEGMENT	BCTYPE	JSTA	JEND	KSTA	KEND	NDATA
1 JO: GRID	1 SEGMENT	1002 BCTYPE	0 ISTA	0 IEND	0 KSTA	0 KEND	0 NDATA
JU: GRID	SEGMENI 1	1008	151A 0	1END 0	KSIA 0	CEND 0	NDATA 0
JDIM: GRID	SEGMENT	BCTYPE	ISTA	IEND	KSTĂ	KEND	NDATA
1	1	1002	0	0	0	0	0
K0: GRID	SEGMENT	BCTYPE	ISTA	IEND	JSTA	JEND	NDATA
1	1	1001	0	0	1	17	0
1	2 TWTYPE	2004 CQ	0	0	17	65	2
	IWIIPE 0.	0.					
KDIM: GRID	SEGMENT	BCTYPE	ISTA	IEND	JSTA	JEND	NDATA
1	1	1003	0	0	0	0	0

1	MGFLAG 1 EPSSSC(1)	0		0	NGAM 02	FDCCCI	> (1) 도		FDCCCD(3)
1550	0.3	0.3							0.3
-	MGLEVG				0		0.5	0.5	0.5
	03	00		000					
MIT1	MIT2	MIT3		MIT4	MIT5	N	4IT6	MIT7	MIT8
01	01	01		01	01		1	1	1
1-1 BLO	CKING DATA:	:							
NBLI									
0									
NUMBER (		ISTA		-		JEND			ISVA2
		ISTA	JSTA	KSTA	IEND	JEND	KEN	D ISVA1	ISVA2
NINTER	RFACE DATA								
NINIER									
PLOT3D O	ITPIIT •								
	TYPE ISTAR	[ IEND	IINC	JSTART	JEND	JINC	KSTAR'	r kend	KINC
	0 (								0
IMOVIE									
0									
PRINT OUT	Γ:								
	TYPE ISTAR	[ IEND	IINC	JSTART	JEND	JINC	KSTAR	I KEND	KINC
CONTROL S	SURFACE:								
NCS									
	TART IENI		7778		יעד דירוג	END IV			
GRID IS.	IAKI IENI	JJJARI	JEI	ND KSI	AKI KI	עד מאיז	VALL	INORM	

# 3.1 LT1 - Input/Output File Names

Input/Output file names, up to 60 characters each, starting in column 1. Specify the user file names in the following order (one name per line), using a "dummy" file name if a file is not actually needed for the current problem:

Input/Output File File Type
grid (unit 1)
PLOT3D grid (unit 3) binary <sup>[2],[3]</sup> output
PLOT3D solution (unit 4) binary <sup>[2],[3]</sup> output
primary case information (unit 11) ascii output
residual and force coefficient history (unit 12) ascii output
turbulence model residual history (not Baldwin-Lomax) (unit 13) ascii output
Baldwin-Lomax (unit 14) ascii output
secondary case information (unit 15) ascii output

<sup>[1]</sup> Defaulted in the code to standard binary for the particular machine on which the code is executed. On Cray computers the default can be changed to IEEE Big unformatted by uncommenting "call asnfile" one line prior to the open statement for unit 1 in the main routine.

<sup>[2]</sup> Defaulted in the code to standard binary for the particular machine on which the code is executed. On Cray computers the default can be changed to IEEE Big unformatted by uncommenting the "call asnfile" lines one line prior to the open statements for both unit 3 and unit 4 in the main routine.

<sup>[3]</sup> Can be changed to formatted by changing the open statements for unit 3 and unit 4 to "formatted" and changing ibin from 1 to 0 in subroutine qout.

flow-field variables printout (unit 17) ...... ascii output unsteady pressures from forced oscillations (**iunst** > 0) (unit 20) .... ascii output interpolation coefficients for chimera grids (unit 21) ......binary input interpolation coefficients for patched grids (unit 22) ..... binary input restart (unit 2) .......binary input/output

3.2 LT2 - Case Title

title describing case

low Conditions
– free-stream <sup>[4]</sup> Mach number
– angle of attack (See Appendix I.)
– side-slip angle (See Appendix I.)
- free-stream <sup>[4]</sup> Reynolds number per unit grid length (millions) <u>How to Set reue</u> Suppose it is desired to simulate a flow in which the Reynolds number based on some characteristic length (e.g. the chord) is $Re$ . If $L$ is the cor- responding length in the grid (ignore the dimension, if any, of $L$ ), then set <b>reue</b> = $Re/L$ . For example, if unit "1" in the grid corresponds to "1 inch" in the experiment, then give <b>reue</b> as the Reynolds number per inch. If unit "1" in the grid corresponds to "16.7845 furlongs", then give <b>reue</b> as the Reynolds number per 16.7845 furlongs. (Also remember that in the input file, <b>reue</b> /1×10 <sup>6</sup> is the actual input value, so put a "20" for "20 million", etc.) (See "Reynolds Number Examples" on page 55.)
- free-stream <sup>[4]</sup> temperature (degrees Rankine)
- indicator for determining how angle of attack is measured in PLOT3D- type grids ( <b>ngrid</b> < 0) <u>If</u> <u>Then</u> <b>ialph</b> = 0 <b>alpha</b> is measured in the $x - z$ plane (with z "up"), i.e. <b>alpha</b> = 90° would give a free-stream velocity vector in the positive z direction.

<sup>[4]</sup> See the note on page 3 about the usage of the phrase *free stream*.

ialph > 0 alpha is measured in the x - y plane (with y "up"), i.e. alpha = 90° would give a free-stream velocity vector in the positive y direction.

<u>Note</u>

(1) For CFL3D-type grids, alpha is always measured in the x-z plane, with z "up". (If it is desired to have y "up" with CFL3D-type grids, then set ialph > 0 and comment out the following line in subroutine global:

- determines which variables are to be tracked for a convergence history

if (ngrid.gt.0) ialph=0

ihstry

- (i.e., which variables are output to file cfl3d.res (unit 12) and file cfl3d.subit.res (unit 23))
  If Then
  - **ihstry** = 0 standard convergence history: residual,  $C_l$ ,  $C_d$ ,  $C_y$  or  $C_z$ depending on whether z or y is "up", respectively, and  $C_{my}$  or  $C_{mz}$  (for z or y "up", respectively)
  - **ihstry** > 0 control surface history: residual and mass flow, pressure force, viscous force, thrust (momentum) force (forces are resultant forces, i.e. if  $f_x$ ,  $f_y$ , and  $f_z$  are the force components in the x, y, and z directions, then the resultant force is  $\sqrt{f_x^2 + f_y^2 + f_z^2}$ ; must have **ncs** > 0 and specify which

control surfaces are to be tracked - only those surfaces with **inorm** 0 are included in the sum)

<u>Note</u>

(1) "LT3 - Flow Conditions" of CFL3D Versions 4.1 and earlier had additional parameters isnd and c2spe to govern wall temperature and heat transfer. These parameters are no longer used, but their functions are implemented in a more general fashion in boundary condition type 2004. (See notes under "LT14 - I0 Boundary Condition Specification" on page 32). The position formerly occupied by isnd is now occupied by ialph, so that the value of ialph for PLOT3D-type grids is no longer hard-wired in subroutine rp3d (as in earlier versions of the code).

#### 3.4 LT4 - Reference Quantities

- sref reference area used to compute non-dimensional forces and moments;
   sref is generally taken to be the plan-form area (*not* the wetted area) of a wing, for example. It should be given in grid dimensions. (You must also account for the grid-distance between 2-D planes when i2d=1!)
- **cref** reference length used to compute non-dimensional moments

bref- reference span used to compute non-dimensional momentsxmc- moment center in x directionymc- moment center in y directionzmc- moment center in z direction

# 3.5 LT5 - Time Step Parameters

dt

– time step <i><u>If</u></i>	<u>Then</u>		
$\mathbf{dt} < 0$			CFL number = $ dt $ ; Use <b>ncyc</b> to con- ycles. Sub-iterations are not used.
<b>dt</b> > 0	constant time step equal to <b>dt</b> ; Use <b>ntstep</b> to control the number of time steps and <b>ncyc</b> to control the number of sub-iterations. Two different sub-iteration strategies can be employed: $t$ -TS and $\tau$ -TS. The $t$ -TS method can have unacceptable time-step limitations depending on the case. The $\tau$ -TS method allows for much larger time steps, although larger time steps require more sub-iterations. Multigrid may be used to enhance convergence of sub-iterations and is recommended in general.		
	<u>If</u>	<u>Then</u>	
	t-TS	ntstep	= number of time steps
		ita	= +1  or  +2 (+2  recommended)
		mgflag	= 0  or  1 (1  recommended)
		mglevel	= number of multigrid levels
		ncyc	= number of sub-iterations + 1 (typically from 2 to 15) ( <b>ncyc</b> = 1 will yield standard time-accurate method with <i>no</i> sub-iterations.)
	τ-TS	ntstep	= number of time steps
		ita	= -1 or -2 (-2 recommended) (See "LT6 - Options and Specifica- tions" on page 23)
		mgflag	= 0 or 1 (1 recommended)
		mglevel	= number of multigrid levels
		ncyc	= number of sub-iterations + 1 (typically from 5 to 25)

## Notes

	<ul> <li>the physical larger time ber of sub-increase.</li> <li>(2) It is <i>stro</i> monitored cient sub-should dr values due point, it is</li> <li>(3) The τ-TS "pseudo" number as</li> </ul>	ber of sub-iterations required depends partly on the size of cal time step used. More sub-iterations are required for e steps. The use of multigrid will generally reduce the num- iterations required, although the cost per sub-iteration will <i>ngly</i> recommended that the sub-iteration convergence be in the output file cfl3d.subit_res to determine if suffi- iterations are being used for a particular case. The residual op and the force coefficients should tend toward constant ring each sub-iteration. However, from an efficiency stand- not generally desirable to fully converge each time step. S scheme utilizes, in addition to the physical time step, a time step based on local time stepping. Thus, there is a CFL esociated with $\tau$ -TS (see <b>cfl_tau</b> on page 23). e and Time Step" on page 56 for the nondimensionalization e step.
		second order accuracy in time ( $ ita  = 2$ ), sub-iterations ) must be used.
irest	- restart flag <u>If</u> <b>irest</b> = 0 <b>irest</b> > 0 <b>irest</b> < 0	<u>Then</u> no restart restart from previous solution, using restart file (unit 2) restart from previous solution, but do not save previous history information
iflagts	<pre>- time step ram If iflagts = 0 iflagts &gt; 0</pre>	ping flag <u>Then</u> constant <b>dt</b> <b>dt</b> ramped over <b>iflagts</b> steps to <b>dt</b> × <b>fmax</b>
fmax	<u>Notes</u> (1) The ramp occurring (2) When the ber to set	crease in dt; $dt_{final} = fmax \times dt_{initial}$ sing of the time step/CFL number to $dt_{final}$ is non-linear, slowly at first and then increasing in rate. time step has successfully been ramped to $dt_{final}$ , remem- the input value of dt to $dt_{final}$ before restarting with the solution. Also, remember to change <b>iflagts</b> to zero and/or .0.
iunst	- unsteady mes $\underline{lf}$ <b>iunst</b> = 0 <b>iunst</b> = 1	

#### Notes

 $\overline{(1) \text{ If }} dt < 0$ , **iunst** is automatically set to 0.

- (2) See "LT33 Number of Translated Grids" on page 43 and beyond for specification of dynamic grid input.
- cfl\_tau- CFL number for  $\tau$ -TS scheme; cfl\_tau is always > 0 and it is not used for<br/>the t-TS scheme (ita > 0) or non-time-accurate runs (dt < 0). A value in<br/>the range from 5 to 10 is recommended; however, a smaller value may be<br/>required if the sub-iterations do not converge.

## 3.6 LT6 - Options and Specifications

ngrid	- number of grids input = abs(ngrid) <u>If</u> <u>Then</u> ngrid > 0 CFL3D grid format ngrid < 0 PLOT3D (or TLNS3D) grid format (See "Grid File" on page 65.)
nplot3d	<ul> <li>number of flow-field data sets to be output in PLOT3D format <u>Notes</u> <ol> <li>If nplot3d &lt; 0, then the PLOT3D files are automatically set to include all solid surfaces (no field points) for 3-d cases or all field points for 2-d cases.</li> </ol> </li> <li>(2) If nplot3d &lt; 0, <i>do not</i> input any PLOT3D data lines in the PLOT3D input section below ("LT28 - PLOT3D Output Specifications"), that is, if nplot3d &lt; 0, treat the PLOT3D input section as if nplot3d = 0.</li> </ul>
nprint	<ul> <li>number of data sets to be sent to an output file <u>Notes</u> (1) If nprint &lt; 0, then the printout file is automatically set to include all solid surfaces (no field points) for 3-d cases or all field points for 2-d cases. (2) If nprint &lt; 0, <i>do not</i> input any print out data lines in the print out input section below ("LT30 - Print Out Specifications"), that is, if nprint &lt; 0, treat the print out input section as if nprint = 0.</li></ul>
nwrest	- number of iterations between updates of the binary restart file (Note that the binary restart file is also updated at the last iteration of the run regard- less of the value of <b>nwrest</b> .) Must be > 0. Used to save an earlier restart file in case the full run does not finish. Generally, set <b>nwrest</b> > <b>ncyc</b> (or > <b>ntstep</b> for time-accurate) unless it is suspected that the run may not fin- ish, or for a margin of safety. Restart file is overwritten each time it is updated.

ichk	- checks for negative densities and/or pressures in subroutines gfluxr, hfluxr, ffluxr, and conu if set equal to 1.		
		for negative densities and/or pressures requires additional s, so utilize <i>only</i> to diagnose problem cases.	
i2d	<ul> <li>(2) For i2d ≠</li> <li>(3) For i2d =</li> </ul>	2-d case	
ntstep	<u>Notes</u> (1) For time-a ations. (2) For steady	cles for time-accurate computations ( <b>dt</b> > 0) accurate computations, <b>ncyc</b> controls the number of sub-iter- y-state computations ( <b>dt</b> < 0), <b>ntstep</b> defaults to 1 and <b>ncyc</b> he number of cycles.	
ita	$   \underbrace{If}_{ita = +1} \\   ita = +2 \\   ita = -1 \\   ita = -2 \\   \underbrace{Notes}_{(1) \text{ The approxime introversion is duced. Sun nalization time) solution time) solution time)} $	-accuracy/two-time scheme flag (used only for $dt > 0$ ) <u>Then</u> first order in time; physical time term only ( <i>t</i> -TS method) second order in time; physical time term only ( <i>t</i> -TS method) first order in time; physical <i>and</i> pseudo time terms ( $\tau$ -TS method) second order in time; physical <i>and</i> pseudo time terms ( $\tau$ -TS method) second order in time; physical <i>and</i> pseudo time terms ( $\tau$ -TS method) (recommended) eximate factorization scheme used to advance the solution in duces first order errors in time. Furthermore, if the diagonal utilized ( <b>idiag</b> = 1), additional errors of order $\Delta t$ are intro- ub-iterations can be used to drive these factorization/diago- errors to zero. Therefore, if a formally second-order (in tion is desired, sub-iterations <i>must</i> be used; the run will ter- <b>ita</b> > 1 and <b>ncyc</b> = 1.	

(2) The overhead for second order verses first order is relatively small when compared to the gain in accuracy and is therefore recommended. (3) The inclusion of a pseudo time term increases (often dramatically) the maximum allowable time step one can take for a particular problem. However, sub-iterations (ncyc > 1) are therefore mandatory and multigrid is recommended. Also, note that larger time steps imply greater temporal error; so again, second order is recommended.

# 3.7 LT7 - Grid, Force, and Viscous Options

(Data for Line Type Seven should be repeated **ngrid** times.)

ncg	<ul> <li>number of coarser grids to construct for multigrid and/or mesh sequencing (Set ncg = 0 for an embedded mesh.)</li> <li><u>Note</u> <ul> <li>(1) With the exception of embedded grids, all grids should have the same value of ncg.</li> </ul> </li> </ul>		
iem	- embedded me <u>If</u> <b>iem</b> = 0 <b>iem</b> = <b>l</b>	esh flag <u>Then</u> global grid level number ( <b>l</b> ) of this embedded grid above the global grid level	
iadvance	<u><i>If</i></u> iadvance ≥ 0	e residual/update calculations <u>Then</u> evaluate the residual and update the solution in the current block skip the residual/update calculations for the current block (rarely used)	
iforce	iforce is a 3-c $Where$ $IntI = 0noI = 1calI = 2calI = 3calJ = 0noJ = 1calJ = 2calJ = 3calMK = 0K = 0no$	the calculation of forces on block faces with solid surfaces; digit number of the form IJK: <i>tiates</i> force calculations on the <i>i</i> faces leulation of the force contribution from the $i = 1$ face leulation of the force contributions from both the $i = 1$ face leulation of the force contributions from both the $i = 1$ face d the $i = idim$ face force calculations on the <i>j</i> faces leulation of the force contribution from the $j = 1$ face leulation of the force contribution from the $j = 1$ face leulation of the force contributions from both the $j = 1$ face leulation of the force contributions from both the $j = 1$ face leulation of the force contributions from both the $k = 1$ face d the $j = jdim$ face force calculations on the <i>k</i> faces leulation of the force contribution from the $k = 1$ face	

- K = 2 calculation of the force contribution from the k = kdim face
- K = 3 calculation of the force contributions from both the k = 1 face and the k =kdim face

Notes

- Only solid surfaces contribute to the force computations; i.e. only those boundaries with boundary condition types 1005 and 2004 can contribute to the force totals. Thus, wakes (with boundary condition type 0) are not computed in the force total, regardless of whether or not they are on a boundary that has been flagged with **iforce** > 0.
- (2) The iforce parameter indicates the calculations of force contributions from entire faces; i.e. **iforce**  $\neq 0$  will cause *all* segments on that particular face to contribute to the force total (assuming those segments have solid-surface boundary conditions, as discussed previously). To eliminate from the force computation a segment that would otherwise contribute to the force total, set segment < 0 in "LT14 - I0 Boundary Condition Specification" on page 32 through "LT19 - KDIM Boundary Condition Specification" on page 35. This will eliminate that particular segment from the total. This feature is useful, for example, when the fuselage and sting are both on the k = 1 boundary of grid 1. The force contribution from the fuselage is desired, but not the contribution from the sting. In this case, let the k = 1 boundary be defined by two segments (the first for the fuselage and the second for the sting). Set **iforce** = 001 for grid 1 and then set **segment** = -2 for the second segment in "LT18 - K0 Boundary Condition Specification", e.g.:

K0: GRID	SEGMENT	BCTYPE	ISTA	IEND	JSTA	JEND	NDATA
1	1	2004	1	97	0	0	2
1	-2	2004	97	129	0	0	2

- viscous/inviscid surface flag with  $m = 1, 2, 3 \Rightarrow m = i, j, k$ 

ivisc(m)

<u>If</u>	Then
ivisc = 0	inviscid
<b>ivisc</b> = 1	laminar
ivisc = 2	turbulent – Baldwin-Lomax model
ivisc = 3	turbulent – Baldwin-Lomax with Degani-Schiff modification
ivisc = 4	turbulent – Baldwin-Barth model
ivisc = 5	turbulent – Spalart-Allmaras model
<b>ivisc</b> = 6	turbulent – Wilcox $k - \omega$ model
ivisc = 7	turbulent – Menter's $k - \omega$ SST model
<b>ivisc</b> = 8	turbulent $-k - \omega$ Explicit Algebraic Stress Model Gatski-
	Speziale (EASM Gatski-Speziale) in eddy-viscos-
	ity formulation
ivisc = 9	turbulent $-k - \varepsilon$ EASM Girimaji in eddy-viscosity formula-
	tion
<b>ivisc</b> = 10	turbulent – Abid $k - \varepsilon$ model

- **ivisc** = 11 turbulent  $-k \varepsilon$  EASM Gatski–Speziale *Nonlinear*
- **ivisc** = 12 turbulent  $-k \omega$  EASM Gatski–Speziale *Nonlinear*
- **ivisc** = 13 turbulent  $-k \varepsilon$  EASM Girimaji *Nonlinear*

```
Notes
```

(1) If ivisc(m) < 0 on input, a wall function is employed. This option should only (officially) be used for attached flow, when the first grid

point off the wall is not in the sub-layer  $(y^+ > 10 \text{ or more})$ . The wall function with the Baldwin-Lomax model (**ivisc** = -2 or -3) can be particularly non-robust and, although it may sometimes work well, it is not recommended in general.

- (2) The thin-layer viscous terms (laminar or turbulent) can be included in the *j*, *k*, or *i* directions, either separately or combined. Cross-derivative terms are not included. The exception is for turbulent flow using the Baldwin-Lomax model; terms can be included simultaneously in, at most, two directions, either j k or i k, for any particular grid.
- (3) Unlike previous versions of the code, the Baldwin-Lomax model can now be applied at any boundary or boundaries; i.e., the walls may now be at *j* / *k* / *i* = 1 and/or *j* / *k* / *i* = jdim/kdim/idim.
- (4) For multiple-zone applications, be aware that the Baldwin-Lomax model is *not* a field-equation model and it relies on distances to j/k/i = 1 and/or j/k/i = jdim/kdim/idim walls *in a given zone*. If a particular zone does not contain walls but it is still desired that **ivisc** = 2 or 3 there, the model will default to using a length scale associated with a distance to the nearest j/k/i = 1 face *in that zone*. Another option is to set **ivisc** = 0 or 1 in the zones with no walls, but this may not be viable if the zone in question is very near to a wall(s) in other zones.
- (5) For the reasons listed above, the field-equation modes 4 and up, which use the very general minimum distance function (which finds the minimum distance to the nearest wall, regardless of what zone it is in) are generally preferred, especially in multi-zone applications. However, the Baldwin-Lomax model (with the Degani-Schiff option) still seems to be the best model for vortical flows.
- (6) It is preferable to let k be the primary viscous direction and i be the secondary viscous direction. See discussion in Section 5.1.1. If i is used as a primary viscous direction, it may be necessary to switch the order of inversions in subroutine af3f (e.g. from j,k,i to j,i,k). Keep in mind that, unlike i and k, altering the location of the j inversion in the approximate factorization sequence requires substantial recoding.

(7) The minimum distance function smin is computed from viscous walls only. If a wall is inviscid, then as far as smin is concerned, it is invisible. This is important to remember when viscous boundary conditions are turned on after running a case inviscidly for some time since smin may never have been computed!

#### <u>Caution</u>

The EASM models are currently very preliminary and should be considered as "researcher-oriented" as opposed to "production-oriented". Due to the sensitive nature of the variable coefficient cmu, the models are generally less robust than other models (particularly for 3-d cases and particularly for the nonlinear versions). It is recommended that EASM solutions be restarted from previously converged  $k - \omega$  or  $k - \varepsilon$  solutions for this reason. Even then, certain cases may *still* experience trouble! However, preliminary tests indicate that the Girimaji EASM  $k - \varepsilon$  model (**ivisc** = 13) tends to be the most robust of the EASM models.

### 3.8 LT8 - Grid Dimensions

(Data for Line Type Eight should be repeated **ngrid** times.)

idim	- number of grid points in the <i>i</i> direction (must be 2 for $ i2d  = 1$ )
jdim	- number of grid points in the <i>j</i> direction
kdim	- number of grid points in the k direction

## 3.9 LT9 - Laminar Region Specification

(Data for Line Type Nine should be repeated **ngrid** times.)

ilamlo	- lower <i>i</i> grid point index defining the laminar region of the current block
ilamhi	– upper $i$ grid point index defining the laminar region of the current block
jlamlo	- lower $j$ grid point index defining the laminar region of the current block
jlamhi	– upper $j$ grid point index defining the laminar region of the current block
klamlo	– lower $k$ grid point index defining the laminar region of the current block
klamhi	– upper $k$ grid point index defining the laminar region of the current block
	<ul> <li><u>Notes</u> <ul> <li>(1) These parameters are used for simulating transition only if ivisc &gt; 1.</li> <li>Set ilamlo, jlamlo, klamlo = 0 for fully turbulent flow. Currently, only one transition region per grid is allowed.</li> </ul> </li> </ul>

(2) When **ilamlo**, etc. are used, they define a region inside which the production terms in the turbulence model equations (for **ivisc** > 3) are turned "off" (for the Baldwin-Lomax model, the eddy viscosity is merely zeroed out). The flow usually remains laminar inside that region as a result. In contrast, "fully turbulent" (**ilamlo**, etc. = 0) means that the production terms in the turbulence model equations are "on" everywhere. However, in practice, one might see behavior which indicates that the model is still "transitioning" on its own. In other words, one might see a low "laminar"  $C_f$  near the leading edge of an

airfoil, transitioning fairly rapidly to a higher "turbulent"  $C_f$ , for example. Note that turning off the production term does *not* keep turbulence from convecting into a region, if turbulence exists upstream.

(3) Currently for multigrid, the turbulent viscosity is only computed on the finest level grid, then it is restricted to coarser levels. Therefore, the **ilamlo**, etc. values need not be "good" multigrid numbers when used.

#### 3.10 LT10 - Embedded Mesh Specifications

(Data for Line Type Ten should be repeated **ngrid** times.)

	v 1	
inewg	– restart flag fo <u>If</u>	or grid (not needed if <b>irest</b> = 0) <u>Then</u>
	$\vec{inewg} = 0$	
	inewg = $1$	
	Notes	
	<ul> <li><u>Notes</u> <ul> <li>(1) The purpose of <b>inewg</b> is to allow embedded grids to be added <i>after</i> solution process has begun; that is, if a solution has already be computed, but additional resolution in certain parts of the flow field desired. In such cases, embedded grids may be appended to the file and the case may be restarted. Setting <b>inewg</b> = 1 initializes solution on the new embedded level by interpolating from a solu on a coarser level, thereby providing a reasonable starting solution the new embedded level (the solution for the new embedded grin not contained in the restart file at this point).</li> <li>(2) <i>Important:</i> <b>inewg</b> should be set to 1 <i>only for the first restart</i> after new embedded level has been added. On subsequent restarts, <b>inew</b> 0 should be used, since the embedded solution has become part of</li> </ul> </li> </ul>	
	restart file	e and should not be re-initialized.
igridc	<u>If</u>	ag for embedded meshes <u>Then</u> the current grid is a global mesh ( <b>iem</b> = 0)

	igridc > 0 igridc is the grid number to which the embedded mesh (iem > 0) connects
is, js, ks	- starting indices in the connecting grid for placement of an embedded mesh (set $is = 0$ , $js = 0$ , $ks = 0$ for global meshes)
ie, je, ke	<ul> <li>ending indices in the connecting grid for placement of an embedded mesh (set ie = 0, je = 0, ke = 0 for global meshes)</li> </ul>
	$\frac{Note}{(1)}$ The embedded meshes must be a regular refinement in all directions of the grid to which it connects. The exception is in the <i>i</i> direction for which the embedded mesh <i>may</i> have the same grid spacing as the grid to which it connects if desired (this ensures that <b>idim</b> = 2 for all grid levels in 2-d cases, including embedded grid levels).

# 3.11 LT11 - Matrix Inversion and Flux Limiter

(Data for Line Type Eleven should be repeated **ngrid** times.)

idiag(m)	- matrix inversion flag with $m = 1, 2, 3 \Rightarrow m = i, j, k$				
	<u>If</u>	<u>Then</u>			
	idiag = 0	$5 \times 5$ block tridiagonal inversions			
	<b>idiag</b> = 1	scalar tridiagonal inversions			
	Notes				
	$\overline{(1)}$ The scal	ar tridiagonal inversions are recommended for steady-state			
	computat	tions. They may or may not be appropriate for time-accurate			
	computat	tions. For time-accurate computations with scalar tridiagonal			
		s, it is suggested that sub-iterations be used to reduce the			
	•	zation errors.			
	(2) When <b>id</b>	iag = 0, viscous terms are included on the left-hand side of			
	the impli	cit time advancement scheme in the $k$ direction <i>only</i> .			
iflim( <i>m</i> )	– flux limiter f	lag with $m = 1, 2, 3 \Rightarrow m = i, j, k$			
	<u>If</u>	<u>Then</u>			
	<b>iflim</b> = $0$	unlimited			
	<b>iflim</b> = 1	smooth limiter			
	<b>iflim</b> = 2	min-mod limiter (recommended if limiter is needed and $rkap0 \neq 1/3$ )			
	<b>iflim</b> = 3	smooth limiter tuned to $\kappa = 1/3$ with a cut-off to eliminate			
		limiting in regions of small gradients (recommended if			
		limiter is needed and $rkap0 = 1/3$ )			
	Notes				
	$\overline{(1)}$ Many purely subsonic flows do not require a limiter.				
	(2) Use of iffim = 3 will override the $rkap0$ value input and force				
	upwind-t	biased third order.			

## 3.12 LT12 - Spatial Differencing

(Data for Line Type Twelve should be repeated **ngrid** times.)

- spatial differencing parameter for Euler fluxes with **ifds**(*m*)  $m = 1, 2, 3 \Rightarrow m = i, j, k$ <u>If</u> <u>Then</u> ifds = 0flux-vector splitting **ifds** = 1 flux-difference splitting (Roe's scheme) (recommended) rkap0(m)- spatial differencing parameter for Euler fluxes with  $m = 1, 2, 3 \Rightarrow m = i, j, k$ <u>If</u> <u>Then</u> rkap0 = -1fully upwind **rkap0** = 0 Frommes' scheme **rkap0** = 1 central **rkap0** = 1/3upwind-biased third order (recommended)

## 3.13 LT13 - Boundary Condition Segments

(Data for Line Type Thirteen should be repeated **ngrid** times.) boundary condition segment flags:

grid	<ul> <li>– current grid number (must be in order)</li> </ul>		
nbci0	- number of segments on $i = 0$ boundary		
nbcidim	– number of segments on $i = idim$ boundary		
nbcj0	- number of segments on $j = 0$ boundary		
nbcjdim	– number of segments on $j = \mathbf{jdim}$ boundary		
nbck0	- number of segments on $k = 0$ boundary		
nbckdim	– number of segments on $k = kdim$ boundary		
iovrlp	- grid-overlapping flagIfTheniovrlp = 0no overlapping occurs for the current gridiovrlp = 1the current grid receives data from an overlapped grid(s) $\underline{Note}$ (1) The minimum number of segments on any boundary is 1.		

# 3.14 LT14 - IO Boundary Condition Specification

(Data for Line Type Fourteen should be repeated  $\sum_{n=1}^{ngrid} nbci0(n)$  times.)

I0 Boundary:

grid	– current grid number (must be in order)	
segment	– current segment (must be in order and the total must equal <b>nbci0</b> )	
bctype	- boundary condition type for the current segment	
jsta	– <i>j</i> starting gri	d point location for the current segment
jend	– <i>j</i> ending grid	l point location for the current segment
ksta	– k starting gri	d point location for the current segment
kend	-k ending grid	d point location for the current segment
ndata	- number of additional data values required for this boundary condition	
	$\frac{Notes}{(1) \text{ Setting jsta}} = \text{jend} = 0 \text{ and/or } \text{ksta} = \text{kend} = 0 \text{ is a shorthand way of specifying the entire range of } j \text{ and/or } k \text{ for this segment.}$	
	The following notes pertain to "LT14 - I0 Boundary Condition Specific tion" on page 32 through "LT19 - KDIM Boundary Condition Specific tion" on page 35:	
	(2) If ndata > 0, then, following the current line, a header line must appear, followed by a single line with the ndata values for this bound ary condition.	
	(3) When multigrid is being used, it is best if all <i>segment</i> lengths are a multigridable. This is not a requirement. However, if one-to-one connections are involved, it is possible that the use of non-multigrida segment lengths will result in geometric mismatches on coarser g levels, causing a "mismatch" message to be written to unit 11 a most likely, eventual disruption of the execution.	
	<ul> <li>(4) If a particular segment should not be counted in the force total, s segment to be negative. (See <i>Note</i> (2) under iforce in Section 3.7.)</li> <li>(5) Input values for bctype (i.e. boundary condition types currently supported) as follows:</li> </ul>	
	bctype	boundary condition
	1000	free stream
	1000	general symmetry plane
	1001	extrapolation
	1002	inflow/outflow
	1003	(no longer evoilable, use 2004 instead)

1004 (no longer available, use 2004 instead)

1008constant enthalpy and entropy inflow1011singular axis – half-plane symmetry1012singular - full plane1013singular axis – partial plane2002specified pressure ratio2003inflow with specified total conditions2004no-slip wall2005periodic in space2006set pressure to satisfy the radial equilibrium equation2007set all primitive variables2102pressure ratio specified as a sinusoidal function of time	1005	inviscid surface
1012singular – full plane1013singular axis – partial plane2002specified pressure ratio2003inflow with specified total conditions2004no-slip wall2005periodic in space2006set pressure to satisfy the radial equilibrium equation2007set all primitive variables	1008	constant enthalpy and entropy inflow
1013singular axis – partial plane2002specified pressure ratio2003inflow with specified total conditions2004no-slip wall2005periodic in space2006set pressure to satisfy the radial equilibrium equation2007set all primitive variables	1011	singular axis – half-plane symmetry
2002specified pressure ratio2003inflow with specified total conditions2004no-slip wall2005periodic in space2006set pressure to satisfy the radial equilibrium equation2007set all primitive variables	1012	singular – full plane
2003inflow with specified total conditions2004no-slip wall2005periodic in space2006set pressure to satisfy the radial equilibrium equation2007set all primitive variables	1013	singular axis – partial plane
2004no-slip wall2005periodic in space2006set pressure to satisfy the radial equilibrium equation2007set all primitive variables	2002	specified pressure ratio
<ul> <li>2005 periodic in space</li> <li>2006 set pressure to satisfy the radial equilibrium equation</li> <li>2007 set all primitive variables</li> </ul>	2003	inflow with specified total conditions
2006set pressure to satisfy the radial equilibrium equation2007set all primitive variables	2004	no-slip wall
2007 set all primitive variables	2005	periodic in space
1	2006	set pressure to satisfy the radial equilibrium equation
2102 pressure ratio specified as a sinusoidal function of time	2007	set all primitive variables
	2102	pressure ratio specified as a sinusoidal function of time

Descriptions of the boundary conditions can be found in Chapter 6.

3.15 LT15 - IDIM Boundary Condition Specification

(Data for Line Type Fifteen should be repeated  $\sum_{n=1}^{ngrid} nbcidim(n)$  times.)

IDIM Boundary:

All entries in this line type are the same as in "LT14 - I0 Boundary Condition Specification" except:

**segment** – current segment (must be in order and the total must equal **nbcidim**)

## 3.16 LT16 - J0 Boundary Condition Specification

(Data for Line Type Sixteen should be repeated  $\sum_{n=1}^{ngrid} nbcj0(n)$  times.)

J0 Boundary:

grid	<ul> <li>– current grid number (must be in order)</li> </ul>	
segment	– current segment (must be in order and the total must equal <b>nbcj0</b> )	
bctype	- boundary condition type for the current segment	
ista	-i starting grid point location for the current segment	
iend	-i ending grid point location for the current segment	
ksta	-k starting grid point location for the current segment	
kend	-k ending grid point location for the current segment	

ndata – number of additional data values required for this boundary condition

<u>Notes</u>

- (1) Setting ista = iend = 0 and/or ksta = kend = 0 is a shorthand way of specifying the entire range of i and/or k for this segment.
- (2) See general notes in "LT14 IO Boundary Condition Specification".

## 3.17 LT17 - JDIM Boundary Condition Specification

(Data for Line Type Seventeen should be repeated  $\sum_{n=1}^{ngrid} nbcjdim(n)$  times.)

JDIM Boundary:

All entries in this line type are the same as in "LT16 - J0 Boundary Condition Specification" except:

segment – current segment (must be in order and the total must equal **nbcjdim**)

## 3.18 LT18 - K0 Boundary Condition Specification

(Data for Line Type Eighteen should be repeated  $\sum_{n=1}^{ngrid} nbck0(n)$  times.)

K0 Boundary:

grid	- current grid number (must be in order)		
segment	– current segment (must be in order and the total must equal <b>nbck0</b> )		
bctype	- boundary condition type for the current segment		
ista	-i starting grid point location for the current segment		
iend	-i ending grid point location for the current segment		
jsta	-j starting grid point location for the current segment		
jend	-j ending grid point location for the current segment		
ndata	- number of additional data values required for this boundary condition		
	<ul> <li><u>Notes</u> <ul> <li>(1) Setting ista = iend = 0 and/or jsta = jend = 0 is a shorthand way of specifying the entire range of <i>i</i> and/or <i>j</i> for this segment.</li> <li>(2) See general notes in "LT14 - I0 Boundary Condition Specification".</li> </ul> </li> </ul>		

### 3.19 LT19 - KDIM Boundary Condition Specification

(Data for Line Type Nineteen should be repeated  $\sum_{n=1}^{ngrid} nbckdim(n)$  times.)

### **KDIM Boundary:**

All entries in this line type are the same as in "LT18 - K0 Boundary Condition Specification" except:

**segment** – current segment (must be in order and the total must equal **nbckdim**)

# 3.20 LT20 - Mesh Sequencing and Multigrid

mseq	<pre>- mesh sequen for mesh sequen If mseq = 1 mseq = 2</pre>	cing flag for global grids (maximum number of grid levels uencing is 5) <u>Then</u> single solution on the finest grid level solution on the second finest grid advanced $ncyc(1)$ cycles, followed by $ncyc(2)$ cycles on the finest grid. The initial solution on the finest grid is obtained by interpolation from the coarser grid solution. If $ncyc(2) = 0$ , the compu- tations are terminated on the second finest grid after ncyc(1) cycles and the restart file is written for the second finest grid. sequencing from coarsest to finest mesh as above
mgflag	- multigrid flag If mgflag = 0 mgflag = 1 mgflag = 2	<sup>g</sup> <u>Then</u> no multigrid multigrid on coarser global meshes multigrid on coarser global meshes and on embedded meshes
iconsf	- conservation $\underline{If}$ iconsf = 0 iconsf = 1	<u>Then</u> nonconservative flux treatment for embedded grids
mtt	– flag for addit <u>If</u> mtt = 0 mtt > 0	ional iterations on the "up" portion of the multigrid cycle <u>Then</u> no additional iterations (recommended) <b>mtt</b> additional iterations
ngam	- multigrid cyc <i>If</i> <b>ngam =</b> 1 <b>ngam =</b> 2	<u>Then</u>

 $\frac{Note}{(1) \text{ If } mglevg} = 2, \text{ ngam } defaults to 1.$ 

3.21 LT21	- Smoothing	
issc	$- \text{ correction s}$ $\frac{If}{\mathbf{issc}} = 0$	moothing flag <u>Then</u> no correction smoothing (usually recommended, but see note 18 in Chapter 10)
	$\mathbf{issc} = 1$	correction smoothing
epsssc(m)	- correction smoothing coefficient with $m = 1, 2, 3 \Rightarrow m = i, j, k$ ; typical values: 0.3, 0.3, 0.3	
issr	- residual sm <i>If</i> <b>issr</b> = 0 <b>issr</b> = 1	<u>Then</u> no residual smoothing (usually recommended, but see note 18 in Chapter 10)
epsssr(m)	<ul> <li>residual sm values: 0.3,</li> </ul>	noothing coefficient with $m = 1, 2, 3 \Rightarrow m = i, j, k$ ; typical 0.3, 0.3

### 3.22 LT22 - Iterations and Multigrid Levels

(Data for Line Type Twenty-Two should be repeated for each sequence 1 through **mseq** (from coarsest to finest).)

ncyc - number of cycles for steady-state computations (dt < 0); number of subiterations + 1 for time-accurate computations (dt > 0) <u>Notes</u>

- Set ncyc ≥ 1, unless using mesh sequencing and terminating execution on a coarser grid level (see description of mseq = 2 for "LT20 - Mesh Sequencing and Multigrid" on page 35, as well as "Mesh Sequencing" on page 134).
- (2) For time-accurate computations (dt > 0), when ncyc = 1, the code is iterating on each time step *once*, which means *no* sub-iterations are used. Using ncyc = 1 corresponds to the standard non-sub-iterative time-accurate scheme that was in all previous versions of CFL3D. Similarly, ncyc = 2 means that the code is iterating on each time step *twice*, which means sub-iterations are used, since it is *one extra* sub-iteration over the standard non-sub-iterative time-accurate scheme.
- (3) For time-accurate computations (dt > 0) with multigrid (**mgflag** > 0), at least one sub-iteration must be used (**ncyc**  $\ge 2$ ).

mglevg	<ul> <li>number of grids to use in multigrid cycling for the globa</li> <li><u>If</u> <u>Then</u></li> <li>mglevg = 1 single grid level</li> <li>mglevg = 2 two grid levels</li> <li>mglevg = m m grid levels</li> </ul>	al meshes
nemgl	- number of embedded grid levels above the finest global <u>If</u> <u>Then</u> <b>nemgl</b> = 0 no embedded grids <b>nemgl</b> = 1 one embedded grid <b>nemgl</b> = m m embedded grids <u>Note</u> (1) Set <b>nemgl</b> = 0 for global grids coarser than the finest	
nitfo1	– number of first order iterations ( <b>nitfo1</b> = 0 recommende	d)

### 3.23 LT23 - Coarse Grid Iterations

(Data for Line Type Twenty-Three should be repeated for each sequence 1 through **mseq** (from coarsest to finest).)

mitL – iterations on level L for each level L from coarsest to finest (mitL = 1 recommended, in general, although for some cases 3 2 1 or 2 2 1 can yield better multigrid convergence than 1 1 1 when three levels are used, for example)

# 3.24 LT 24 - Number of 1-1 Interfaces

**nbli** – number of 1-1 grid-point-connecting block interfaces

## 3.25 LT25 - 1-1 Blocking Connections

(Data for Line Type Twenty-Five should be repeated nbli times.)

number	<ul> <li>identifying number to aid the user with "bookkeeping" the 1-1 interfaces in the input file (not used by CFL3D)</li> </ul>		
grid	<ul> <li>grid/block identifier indicating which grid in the common interface is being described in this line of input</li> </ul>		
ista	– starting <i>i</i> limit for 1-1 block interface		
jsta	– starting <i>j</i> limit for 1-1 block interface		
ksta	– starting $k$ limit for 1-1 block interface		

iend	– ending $i$ limit for 1-1 block interface		
jend	– ending $j$ limit for 1-1 block interface		
kend	– ending $k$ limit for 1-1 block interface		
isva1	<ul> <li>varying index on 1-1 block interface (isva1 in Line Type Twenty-Five varies with isva1 in Line Type Twenty-Six)</li> </ul>		
isva2	<ul> <li>varying index on 1-1 block interface (isva2 in Line Type Twenty-Five varies with isva2 in Line Type Twenty-Six)</li> </ul>		
	<ul> <li>Notes <ul> <li>(1) The code will check the input connection data by computing the geometric mismatch between both sides of the interface, based on the specified ranges. A true 1-1 interface will have zero (or machine zero) mismatch. Any mismatches larger than ε (where ε is the larger of 10<sup>-9</sup> and 10 times machine zero) will cause a warning message to be printed out in the main output file (look for the words "geometric mismatch". Always check the main output file (unit 11) for these words, since the code will only print out the message; it will not stop. A good practice is to make one run with only one iteration, then check the output for "geometric mismatch". Small mismatches may be acceptable, but large mismatches [O(1)] most likely indicate that one of the segments on the block boundary has been specified incorrectly.</li> <li>(2) Any block segments that have 1-1 connectivity must also be input in the boundary condition section ("LT13 - Boundary Condition Segments" on page 31 through "LT19 - KDIM Boundary Condition Specification" on page 35), setting bctype = 0.</li> </ul></li></ul>		

# 3.26 LT26 - 1-1 Blocking Connections

(Data for Line Type Twenty-Six should be repeated **nbli** times.)

Line Type Twenty-Six has the same parameter types as Line Type Twenty-Five. It gives the grid/block 1-1 connection information on the opposite side of the interface.

# 3.27 LT27 - Number of Patched-Grid Interfaces

ninter	- patched-grid flag (zonal interface along a common surface where points		
	need not match)		
	<u>If</u>	<u>Then</u>	
	<b>ninter</b> = 0	no patched-grid data is needed (no patched interfaces)	
	<b>ninter</b> < 0	patched-grid data supplied in a separate file	

#### <u>Notes</u>

- (1) In Version 5.0 of CFL3D, patched-grid interpolation data for static patched interfaces (interfaces that do not change with time) must be obtained as a preprocessing step (as for chimera overlapped grids). The code ronnie is designed for this task. Dynamic patched interfaces (such as occur when blocks slide past one another) are computed internally to CFL3D. The input for dynamic patched interfaces is described in "LT41 Number of Dynamic Patched-Grid Interfaces" on page 49 through "LT45 Displacement and Rotation Amounts" on page 52.
- (2) Any block segments that are patched must also be input in the boundary condition section ("LT13 - Boundary Condition Segments" on page 31 through "LT19 - KDIM Boundary Condition Specification" on page 35), setting **bctype** = 0.

#### 3.28 LT28 - PLOT3D Output Specifications

(Data for Line Type Twenty-Eight should be repeated nplot3d times.)

- designated grid number for output grid iptype - type of PLOT3D file output If Then iptype = 0grid point type – grid file and  $\mathbf{Q}$  file output iptype = 1cell center type – grid file and  $\mathbf{Q}$  file output iptype = 2cell center type – grid file and turbulence file output (ivisc > 1 only) (The current defaults for output are the production term  $\tilde{P} \cdot \tilde{L}_R / (\tilde{u}_{\infty})^3$  and the Reynolds stress components  $\overline{u'w'}/(\tilde{u}_{\infty})^2$ ,  $\overline{u'u'}/(\tilde{u}_{\infty})^2$ , and  $\overline{w'w'}/(\tilde{u}_{\infty})^2$  for 2-d; and the same plus  $Sk/\epsilon$  for 3-d. (See Ristorcelli<sup>30</sup> for a description of these and other available variables.) These may be modified to output other variables if the user desires. If no turbulence model is being used or if the turbulence model cannot obtain a specific parameter being asked for, a value of zero is output. The "hard wire" occurs in subroutine plot 3t. Note that the u', v', and w' quantities are aligned with the x, y, and z axes of the grid, respectively. cell center type - grid file and PLOT3D function file outiptype > 2 put Specific functions currently available: iptype = 3minimum distance to nearest viscous wall or directed distance (**ivisc** > 1 only) iptype = 4eddy viscosity (**ivisc** > 1 only)

istart	- starting location in <i>i</i> direction		
iend	– ending location in <i>i</i> direction		
iinc	– increment factor in $i$ direction		
jstart	– starting location in <i>j</i> direction		
jend	– ending location in <i>j</i> direction		
jinc	- increment factor in $j$ direction		
kstart	- starting location in $k$ direction		
kend	- ending location in $k$ direction		
kinc	- increment factor in $k$ direction		
<ul> <li>Notes <ul> <li>(1) Setting istart = iend = iinc = 0, jstart = jend = jinc = 0, and/or kst = kend = kinc = 0 is a shorthand way of specifying the entire range <i>i</i>, <i>j</i> and/or <i>k</i>, respectively.</li> <li>(2) Grid files are written with a blank array.</li> <li>(3) All files are in multi-block format (even if there is a single zone).</li> <li>(4) Files are written as 2-d if i2d ≠ 0.</li> <li>(5) Function files contain only one variable.</li> <li>(6) Files are written out in single precision.</li> <li>(7) On Cray computers, "uncommenting" the lines <ul> <li>c call asnfile(plt3dg, '-F f77 -N ieee', IEI</li> <li>and</li> <li>c call asnfile(plt3dg, '-F f77 -N ieee', IEI</li> <li>in module cbsem.f will cause the files to be written in IEEE bind.</li> <li>The files may then be transferred directly to a Silicon Graphics workstation and read into FAST or PLOT3D using the unformatted opting the set of t</li></ul></li></ul></li></ul>			

(8) This output will appear in the PLOT3D files (units 3 and 4) named near the top of the input deck.

# 3.29 LT29 - Movie Option

movie - flag to append periodically to the PLOT3D output files (grid and solution) as the solution progresses. Only one grid file and one solution file are generated. If

<u>Then</u>

default (binary) option.

- **movie** = 0no output of intermediate solutions (if **|nplot3d**| > 0, then a single solution is written at the end of the run)
- movie > 0output of additional solutions every movie iterations (the contents of the augmented PLOT3D files are governed by the parameters under "LT28 - PLOT3D Output Specifications" on page 39)
- **movie** < 0output of the initial flow field at the beginning of the run and output of additional solutions every |movie| iterations

**Caution** 

Use with care: PLOT3D files will get very large very quickly! If the complete flow field is output, this is really a viable option only in 2-d. In 3-d, it is generally a viable option only if subsets of the flow field are output (i.e. the surface data). Also, set nprint = 0 when movie > 0 or large print files will result as well.

### 3.30 LT30 - Print Out Specifications

(Data for Line Type Thirty should be repeated **nprint** times.)

grid	– designated grid number for output	
iptype	If $Ihen$ $If$ $Ihen$ $iptype = 0$ grid point type $iptype = 1$ cell center type	
istart	– starting location in <i>i</i> direction	
iend	- ending location in <i>i</i> direction	
iinc	- increment factor in <i>i</i> direction	
jstart	- starting location in $j$ direction	
jend	- ending location in $j$ direction	
jinc	– increment factor in <i>j</i> direction	
kstart	- starting location in $k$ direction	
kend	- ending location in k direction	
kinc	- increment factor in $k$ direction	
	$\frac{Notes}{(1) \text{ Setting istart}} = \text{iend} = \text{iinc} = 0, \text{ jstart} = \text{jend} = \text{jinc} = 0, \text{ and/or kstart}$ $= \text{kend} = \text{kinc} = 0 \text{ is a shorthand way of specifying the entire range of}$ $i, j \text{ and/or } k.$	

(2) This output will appear in the printout file (unit 17) named near the top of the input deck.

Section 5.2.2 gives a detailed sample of the contents of the printout file.

# 3.31 LT31 - Number of Control Surfaces

ncs – the number of control surfaces; control surfaces are user-specified grid surfaces through which mass flow, thrust (momentum) force, pressure force, and viscous force may be monitored. Control surfaces *must* be specified if **ihstry** > 0. (See "LT3 - Flow Conditions" on page 19.)

# 3.32 LT32 - Control Surface Specifications

block	- block number of the control surface	
istart	– starting location in <i>i</i> direction	
iend	– ending location in <i>i</i> direction	
jstart	- starting location in <i>j</i> direction	
jend	- ending location in $j$ direction	
kstart	- starting location in $k$ direction	
kend	- ending location in $k$ direction	
iwall	If $Then$ $iwall = 0$ for a flow surface $iwall = 1$ for a solid wall	
inorm	- defines the direction of the control surface outward normal relative to the grid surface normal. This is only required in the calculation of the total forces for all control surfaces to obtain the correct signs. If <u>Then</u> inorm = 1 the control volume outward normal is in the same direction as the grid surface normal inorm = -1 the control volume outward normal is in the opposite direction to the grid surface normal inorm = 0 do not add this surface into the summation of the total forces <u>Notes</u>	

- (1) The starting and ending indices in one of the directions must be identical to define the surface. If they are not, no calculation is performed. In the remaining two directions, setting both the starting and ending indices to zero is a shorthand way of specifying the entire index range.
- (2) If all the surfaces have **inorm** = 0, the totals are not calculated.
- (3) This output will appear in the printout file named near the top of the input deck; the control surface data appears after the data generated by setting **nprint** > 0.

### THE INPUT DECK TERMINATES HERE FOR STEADY-STATE CASES OR TIME-DEPENDENT CASES IN WHICH THE GRID IS STATIONARY (**iunst** = 0)

Line Types Thirty-Three through Forty-Five are required only if **iunst** > 0 (dynamic grid).

## 3.33 LT33 - Number of Translated Grids

**ntrans** – number of grids undergoing translation motion <u>Notes</u>

- (1) If **ntrans** = 0, only the headers for "LT34 Translational Reference Length" on page 43 through "LT36 - Maximum Translational Displacements" on page 45 are required (i.e. no numerical values required).
- (2) If the moment center also needs to be translated, set **ntrans** = "number of grids undergoing translation + 1". If the moment center is not translated, moments are computed relative to a fixed point in space.

# 3.34 LT34 - Translational Reference Length

(Data for Line Type Thirty-Four should be repeated **ntrans** times.)

Iref – the "grid equivalent" of the dimensional reference length used to define either the reduced frequency needed for itrans = 2 or the decay rate needed for itrans = 3 (See "LT35 - Translational Information and Velocities" on page 44 below.) For itrans ≤ 1, any value may be input (it is reset to 1.0 internally).

### Note

(1) For example, if the input value of reduced frequency for a sinusoidally plunging wing was defined based on the dimensional chord of the wing and, in the grid, the chord of the wing is 2.0, then set **lref** = 2.0.

3.35 LT35 - Translational Information and Velocities (Data for Line Type Thirty-Five should be repeated ntrans times.) grid - designated number of grid to be translated Note (1) If grid = 0, then the following input governs the motion of the moment center; on output, the value of 0 will be replaced by the letters "MC" as a reminder. - type of translation itrans If Then itrans = 0no translation itrans = 1translation with constant speed itrans = 2sinusoidal variation of displacement itrans = 3smooth increase in displacement, asymptotically reaching a maximum displacement rfreq - reduced frequency when **itrans** = 2; growth rate to maximum displacement when **itrans** = 3 (set **rfreq** = 0 for **itrans** = 0 or **itrans** = 1) - translation velocity in x direction when **itrans** = 1; maximum displaceutrans ment in x direction when itrans > 1 (set utrans = 0 if there is no x displacement) - translation velocity in y direction when **itrans** = 1; maximum displacevtrans ment in y direction when **itrans** > 1 (set **vtrans** = 0 if there is no y displacement) wtrans - translation velocity in z direction when **itrans** = 1; maximum displacement in z direction when **itrans** > 1 (set wtrans = 0 if there is no z displacement) Notes (1) Depending on the type of motion, the nondimensional input parameters listed above have different definitions and it is important to understand the distinction. These multiple definitions allow a range of grid motions with a minimal number of input parameters. (2) The sign of utrans/vtrans/wtrans governs the direction of the motion: "+" for motion in the increasing coordinate direction, "-" for motion in the decreasing coordinate direction. For itrans = 2 (sinusoidal), the sign dictates the "initial" direction of motion. Caution

If PLOT3D-type grids, where y is "up", are used, it is important to note that the grid is internally treated as if z is "up" and that **utrans**, **vtrans**, and **wtrans** refer to the internal velocities. Thus, if y is "up" in the input grid and translation in the "up" direction is desired, then **wtrans** must be used to set the speed. It is advised that a visualization package such as FAST or PLOT3D be used in the early stages of the computation to verify that the grid is moving in the desired direction.

### 3.36 LT36 - Maximum Translational Displacements

(Data for Line Type Thirty-Six should be repeated ntrans times.)

grid	<ul> <li>designated number of grid to be translated</li> <li><u>Note</u> <ul> <li>(1) If grid = 0, then the following input governs the motion of the moment center; on output, the value of 0 will be replaced by the letters "MC" as a reminder.</li> </ul> </li> </ul>
dxmax	- maximum (absolute) translational displacement in the x direction to be allowed for this grid, measured from the $t = 0$ position; set <b>dxmax</b> = 0 if no restriction is required
dymax	- maximum (absolute) translational displacement in the y direction to be allowed for this grid, measured from the $t = 0$ position; set <b>dymax</b> = 0 if no restriction is required
dzmax	<ul> <li>maximum (absolute) translational displacement in the <i>z</i> direction to be allowed for this grid, measured from the <i>t</i> = 0 position; set dzmax = 0 if no restriction is required</li> <li><i>Note</i></li> </ul>

(1) Setting dxmax, dymax, and/or dzmax > 0 will have the following effect: if the grid moves further than |dxmax|, |dymax|, and/or |dzmax| from its position at t = 0, the grid will be reset backward the appropriate amount, from where the specified grid motion will start again. The grid will be reset every time it reaches the specified limit. Though each block may have an independent limit, care should be taken to insure that any blocks that must be reset concurrently have the *same* values of dxmax, dymax, and/or dzmax. This resetting feature is used primarily in 2-d turbomachinery applications. For example, in 2-d rotor-stator interaction problems where only a few of an infinite number of rotor and stator blades are actually modeled in the grid system, resetting the grids periodically allows the simulation to be continued for an arbitrarily long period of time with only a limited number of zones. This resetting option is only applicable for constant translational speed (itrans = 1).

## 3.37 LT37 - Number of Rotational Grids

- **nrotat** number of grids undergoing rotational motion *Notes* 
  - (1) If **nrotat** = 0, only the headers for "LT38 Rotational Reference Length" on page 46 through "LT40 - Maximum Rotational Displacements" on page 48 are required (i.e. no numerical values required).
  - (2) If the moment center also needs to be rotated, set **nrotat** = "number of grids undergoing rotation" + 1.

# 3.38 LT38 - Rotational Reference Length

(Data for Line Type Thirty-Eight should be repeated **nrotat** times.)

lref

the "grid equivalent" of the dimensional reference length used to define either the nondimensional rotation rate needed for irotat = 1, the reduced frequency needed for irotat = 2, or the growth rate needed for irotat = 3 (See "LT39 - Rotational Information and Velocities" below.) For itrans ≤ 1, any value may be input (it is reset to 1.0 internally). *Note*

(1) For example, if the input value of reduced frequency for a sinusoidally pitching wing was defined based on the dimensional chord of the wing and, in the grid, the chord of the wing = 2.0, then set lref = 2.0.

	3.39 LT39 - Rotational Information and Ve	
3.39 LT39	9 - Rotational Information and Velocities	
(Data for Line	e Type Thirty-Nine should be repeated <b>nrotat</b> times.)	
grid	- designated number of grid to be rotated $\frac{Note}{(1)}$ If grid = 0, then the following input governs the motion of the moment center; on output, the value of 0 will be replaced by the let- ters "MC" as a reminder.	
irotat	<ul> <li>type of rotation</li> <li>If <u>Then</u></li> <li>irotat = 0 no rotation</li> <li>irotat = 1 rotation with constant angular speed</li> <li>irotat = 2 sinusoidal variation of angular displacement</li> <li>irotat = 3 smooth increase in displacement, asymptotically reaching a maximum angular displacement</li> </ul>	
rfreq	- reduced frequency when irotat = 2; growth rate to maximum angular dis- placement when irotat = 3 (set rfreq = 0 for irotat = 0 or irotat = 1)	
omegax	<ul> <li>- x component of rotational velocity when irotat = 1; maximum angular displacement about the x axis when irotat &gt; 1; set omegax = 0 if there is no x rotation</li> </ul>	
omegay	- y component of rotational velocity when irotat = 1; maximum angular displacement about the y axis when irotat > 1; set omegay = 0 if there is no y rotation	
omegaz	- z component of rotational velocity when <b>irotat</b> = 1; maximum angular displacement about the z axis when <b>irotat</b> > 1; set <b>omegaz</b> = 0 if there is no z rotation	
xorig	-x coordinate of origin of the rotation axis	
yorig	– y coordinate of origin of the rotation axis	
zorig	-z coordinate of origin of the rotation axis	
	<u>Notes</u> (1) Depending on the type of motion, the nondimensional input parame- ters listed above have different definitions and it is important to understand the distinction. These multiple definitions allow a range of grid motions with a minimal number of input parameters. (See	

Section 4.5.2 in the Nondimensionalization chapter for details.)

(2) The sign of **omegax**, **omegay**, and **omegaz** governs the direction of the rotation: "+" for motion in the positive direction, "-" for motion in the negative direction where "positive" and "negative" are dictated by the right-hand rule in which the thumb points in the positive direction of the axis of rotation and the fingers curl in the positive rotational direction. (See "The Right-Hand Rule" on page 67.)

# 3.40 LT40 - Maximum Rotational Displacements

(Data for Line Type Forty should be repeated **nrotat** times.)

grid	<ul> <li>designated number of grid to be rotated Note</li> </ul>	
	$\overline{(1)}$ If grid = 0, then the following input governs the motion of the moment center.	
dthxmx	- maximum (absolute) rotational displacement about the x axis to be allowed for this grid (set $dthxmx = 0$ if no restriction is required)	
dthymx	<ul> <li>maximum (absolute) rotational displacement about the y axis to be allowed for this grid (set <b>dthymx</b> = 0 if no restriction is required)</li> </ul>	
dthzmx	- maximum (absolute) rotational displacement about the z axis to be allowed for this grid (set <b>dthzmx</b> = 0 if no restriction is required)	
	<u>Notes</u> (1) Setting <b>dthxmx</b> , <b>dthymx</b> , and/or <b>dthzmx</b> > 0 will have the following effect: if the grid rotates more than   <b>dthxmx</b>  ,   <b>dthymx</b>  , and/or   <b>dthzmx</b>   from it's position at $t = 0$ , the grid will be reset backward the appropriate amount, from where the specified grid motion will start again. The grid will be reset every time it reaches the specified limit. Though each block may have an independent limit, care should be taken to insure that any blocks that must be reset concurrently have the <i>same</i> values of <b>dthxmx</b> , <b>dthymx</b> , and/or <b>dthzmx</b> . This resetting feature is used primarily in 3-d turbomachinery applications. For example, in 3-d rotor-stator interaction problems where only a few of the large number of rotor and stator blades are actually modeled in the grid system, resetting the grids periodically allows the simulation to be continued for an arbitrarily long period of time with only a limited number of zones. This resetting option is only applicable for constant rotational speed ( <b>irotat</b> = 1).	

(2) dthxmx, dthymx, and dthzmx are output in degrees.

### 3.41 LT41 - Number of Dynamic Patched-Grid Interfaces

**ninter2** – dynamic patched-grid flag (zonal interface along a common interface, with grids in relative motion on either side of the interface)

<u>If</u> <u>Then</u>

**ninter2** = 0 no patched-grid data is needed (no patched interfaces)

**ninter2** > 0 patched-grid data calculated at every time step

- Notes
- (1) This type of patched-grid data cannot be read in from a file since it must be recalculated for every time step. If there are interfaces for which static patch data is required, use the standard patch input method (precalculate with ronnie and set **ninter** < 0 in "LT27 - Number of Patched-Grid Interfaces" on page 38).
- (2) Any block segments that are dynamically patched must also be input in the boundary condition section ("LT13 - Boundary Condition Segments" on page 31 through "LT19 - KDIM Boundary Condition Specification" on page 35), using bctype = 0.
- (3) Further information on patching can be found in the input.doc file for the ronnie preprocessor. It is recommended that some experience be gained with static patching (via ronnie) before attempting dynamic patching with moving zones.
- (4) While most of the input for static and dynamic patching is the same, dynamic patching requires a few extra pieces of information, described below.

### 3.42 LT42 - Dynamic Patched-Grid Specifications

(Data for Line Type Forty-Two should be repeated **ninter2** times.)

int	– interpolati	on number
ifit	– type of fit	
	If	<u>Then</u>
	<b>ifit</b> = 1	bi-linear fit
	<b>ifit</b> = 2	serendipity (degenerate) bi-quadratic fit
	<b>ifit</b> = 3	quadratic fit in $\xi$ ; linear fit in $\eta$
	ifit = 4	linear fit in $\xi$ ; quadratic fit in $\eta$
limit	– maximum 1 recomm	step size (number of cells) to use in the search routine ( <b>limit</b> = ended)
itmax		number of iterations allowed to find each "to" cell ( <b>itmax</b> = grid dimension recommended)

mcxie	<ul> <li>flag to indicate whether the ξ = 0 boundaries on either side of the patch interface are to be rendered coincident (generally required if ξ = 0 corresponds to a viscous surface) </li> <li>If <ul> <li>Then</li> <li>mcxie = 0</li> <li>do not render the ξ = 0 boundaries coincident</li> <li>mcxie &gt; 100</li> <li>render the ξ = 0 boundaries coincident</li> <li>mcxie = 1 to 100</li> <li>The code will try to decide whether or not the boundaries should be rendered coincident. The larger the value, the greater the number differences between the two boundaries can be before deciding they should not be rendered coincident.</li> </ul> </li> <li>Mote <ul> <li>(1) For multiple "from" blocks, in which multiple ξ, η coordinate systems are defined, the proper ξ = 0 boundary to consider is the one closest to the ξ = 0 boundary on the "to" side.</li> </ul> </li> </ul>
mceta	- flag to indicate whether the $\eta = 0$ boundaries on either side of the patch interface are to be rendered coincident (generally required if $\eta = 0$ corresponds to a viscous surface)
ic0	-flag for C-0 continuous boundaries ( <i>not applicable to dynamically patched interfaces; set</i> <b>ic0</b> = 0)
iorph	<ul> <li>flag to allow points which cannot be located by the search routine to be tagged as "orphan" points and to be marked as being "interpolated" from block 0</li> <li><i>If</i> <u>Then</u></li> <li>iorph = 0 do not allow orphan points</li> <li>iorph = 1 allow orphan points (not recommended for dynamic patch interfaces at this time)</li> </ul>
itoss	<ul> <li>flag to allow faster patching for cases in which the interface lies along a constant (or very nearly so) coordinate surface </li> <li><i>If Then</i> itoss = 0 interface is non-planar </li> <li>itoss = 1 interface is an <i>x</i> = constant plane</li> <li>itoss = 2 interface is an <i>y</i> = constant plane</li> <li>itoss = 3 interface is an <i>z</i> = constant plane</li> <li><i>Notes</i> (1) While itoss = 0 can always be used, it will slow down the dynamic patching considerably. Thus, if the interface is on a constant coordinate surface, the appropriate value of itoss is strongly recommended. </li> <li>(2) All of the above parameters except itoss are also used as input to ronnie.</li> </ul>

# 3.43 LT43 - Dynamic Patching "To" Grid Specifications

(Data for Line Type Forty-Three should be repeated **ninter2** times.)

– interpolation number
<ul> <li>a three or four digit number indicating the block number and face of the "to" surface ("to" refers to the block being interpolated)</li> <li>to = Nmn where</li> </ul>
N is the block number (N may be up to two digits) m is the coordinate direction; $1 = i$ , $2 = j$ , $3 = k$ n is the minimum or maximum face; $1 = minimum$ , $2 = maximum$
– starting index in $\xi$ for which interpolation coefficients will be found on the "to" side of the interface
– ending index in $\xi$ for which interpolation coefficients will be found on the "to" side of the interface
– starting index in $\eta$ for which interpolation coefficients will be found on the "to" side of the interface
– ending index in $\eta$ for which interpolation coefficients will be found on the "to" side of the interface
<ul> <li>number of block boundaries which make up the "from" side of the patch surface ("from" refers to the block(s) from which the interpolations are made)</li> </ul>
<ul> <li><u>Notes</u> <ul> <li>(1) If the entire range is desired, a shortcut is to set xie1 = xie2 = eta1 = eta2 = 0.</li> </ul> </li> <li>(2) <i>Important</i>: For dynamic patching, additional "ghost" block boundaries are often required. The baseline value of nfb is set from the zonal setup at t = 0 (as in static patching). However, as soon as the zones start to move past one another, portions of some block faces may not have an opposing block to interpolate from. To provide the patching algorithm with a block to interpolate from, one or more of the "from" blocks that match to the "to" side at t = 0 are duplicated and then translated or rotated by an appropriate amount to provide coverage of the "to" side as the zones move past one another. The baseline value of ntb must then be increased accordingly.</li> </ul>

# 3.44 LT44 - Dynamic Patching "From" Grid Specifications

(Data for Line Type Forty-Four should be repeated **nfb** times for *each* Line Type Forty-Three.)

from	<ul> <li>a three or four digit number indicating the block number and face of the "from" surface ("from" refers to the block(s) from which the interpolations are made)</li> <li>from = Nmn where</li> <li>N is the block number (N may be up to two digits)</li> <li>m is the coordinate direction; 1 = i, 2 = j, 3 = k</li> <li>n is the minimum or maximum face; 1 = minimum, 2 = maximum</li> </ul>
xie1	– starting index in $\xi$ for which interpolation coefficients will be found on "from" side of interface
xie2	– ending index in $\xi$ for which interpolation coefficients will be found on "from" side of interface
eta1	– starting index in $\eta$ for which interpolation coefficients will be found on "from" side of interface
eta2	– ending index in $\eta$ for which interpolation coefficients will be found on "from" side of interface
factj	- expansion factor in the $\xi$ direction (not implemented at this time (set <b>factj</b> = 0))
factk	- expansion factor in the $\eta$ direction (not implemented at this time (set <b>factk</b> = 0))
	$\frac{Note}{(1)}$ If the entire range is desired, a shortcut is to input zeros for the beginning and ending indices.

# 3.45 LT45 - Displacement and Rotation Amounts

(Data for Line Type Forty-Five should be repeated **nfb** times for *each* Line Type Forty-Three.)

dx	– amount (in grid units) to displace the "from" block in the $x$ direction
dy	– amount (in grid units) to displace the "from" block in the y direction
dz	– amount (in grid units) to displace the "from" block in the $z$ direction
dthetx	– amount (in degrees) to rotate the "from" block in the $x$ direction
dthety	– amount (in degrees) to rotate the "from" block in the y direction
dthetz	– amount (in degrees) to rotate the "from" block in the $z$ direction



#### **CHAPTER 4**

# Nondimensionalization

The goal of many CFD problems is to solve the flow over configurations that have been tested in wind tunnels or flight tests. In experimental tests, the model being studied is of known dimensions. Since there is no standard system of measurement for such testing, some models are measured in feet, some in meters, some in inches, etc. Similarly, wind speeds might be measured in feet/second or meters/second or some other way. Nondimensionalizing the flow-field parameters in a CFD code removes the necessity of converting from one system to another within the code.

Throughout this chapter, the term *free stream* is used. Please see the note on page 3 concerning its usage.

# 4.1 Dimensional Parameter Definitions

The following parameters are used to nondimensionalize the flow-field variables:

 $\tilde{L}$  = characteristic length, feet (e.g. chord)

 $L_{ref}$  = corresponding length in the grid (nondimensional)

 $\tilde{L}_R = \tilde{L}/L_{ref}$  = reference length used by code, feet (dimensional)

 $\tilde{\rho}_{\infty}$  = free-stream density, slug/feet^3

 $\tilde{a}_{\infty}$  = free-stream speed of sound, feet/second

 $|\tilde{\mathbf{V}}|_{\infty}$  = free-stream velocity, feet/second

 $\tilde{\mu}_{\infty}$  = free-stream molecular viscosity, slug/feet-second

 $\tilde{t} = \text{time}, \text{seconds}$ 

The ~ here and elsewhere indicates a dimensional quantity. Note that meters or inches, etc. could be substituted for feet in the above definitions, as long as it is done consistently.

# 4.2 General Flow-Field Variables

In CFL3D, the nondimensionalizations are performed as follows:

$$\rho = \frac{\rho}{\tilde{\rho}_{\infty}} \qquad \rho_{\infty} = 1$$

$$u = \frac{\tilde{u}}{\tilde{a}_{\infty}} \qquad u_{\infty} = M_{\infty} \cos \alpha \cos \beta$$

$$v = \frac{\tilde{v}}{\tilde{a}_{\infty}} \qquad v_{\infty} = -M_{\infty} \sin \beta \qquad (4-1)$$

$$w = \frac{\tilde{w}}{\tilde{a}_{\infty}} \qquad w_{\infty} = M_{\infty} \sin \alpha \cos \beta$$

$$p = \frac{\tilde{p}}{\tilde{\rho}_{\infty} (\tilde{a}_{\infty})^{2}} \qquad p_{\infty} = \frac{1}{\gamma}$$

Also,

$$e = \frac{\tilde{e}}{\tilde{\rho}_{\infty}(\tilde{a}_{\infty})^{2}} \qquad e_{\infty} = \frac{1}{\gamma(\gamma - 1)} + \frac{(M_{\infty})^{2}}{2}$$

$$a = \frac{\tilde{a}}{\tilde{a}_{\infty}} \qquad a_{\infty} = 1 \qquad (4-2)$$

$$T = \frac{\tilde{T}}{\tilde{T}_{\infty}} = \frac{\gamma p}{\rho} = a^{2} \qquad T_{\infty} = 1$$

$$x = \frac{\tilde{x}}{\tilde{L}_{R}} \qquad y = \frac{\tilde{y}}{\tilde{L}_{R}}$$

$$z = \frac{\tilde{z}}{\tilde{L}_{R}} \qquad t = \frac{\tilde{t}\tilde{a}_{\infty}}{\tilde{L}_{R}} \qquad (4-3)$$

Also, Reynolds number based on the characteristic length is

4.3 Reynolds Number Examples

$$Re_{\tilde{L}} = \frac{\tilde{\rho}_{\infty} |\mathbf{V}|_{\infty} \tilde{L}}{\tilde{\mu}_{\infty}}$$
(4-4)

Note that  $Re_{\tilde{L}}$  is *not* input into the code, but rather the Reynolds number per unit grid length, in millions:

$$\mathbf{reue} = \frac{Re_{\tilde{L}}}{L_{ref}} \times 10^{-6} = \frac{\tilde{\rho}_{\infty} |\tilde{\mathbf{V}}|_{\infty} \tilde{L}_R}{\tilde{\mu}_{\infty}} \times 10^{-6} = Re_{\tilde{L}_R} \times 10^{-6}$$
(4-5)

The free-stream Mach number is

$$M_{\infty} = \frac{|\mathbf{\tilde{V}}|_{\infty}}{\tilde{a}_{\infty}}$$
(4-6)

where the free-stream velocity magnitude is

$$|\tilde{\mathbf{V}}|_{\infty} = \sqrt{(\tilde{u}_{\infty})^2 + (\tilde{v}_{\infty})^2 + (\tilde{w}_{\infty})^2}$$
(4-7)

The molecular viscosity is defined from Sutherland's<sup>45</sup> law as follows:

$$\mu = \frac{\tilde{\mu}}{\tilde{\mu}_{\infty}} = \left(\frac{\tilde{T}}{\tilde{T}_{\infty}}\right)^{\frac{3}{2}} \left[\frac{\tilde{T}_{\infty} + \tilde{c}}{\tilde{T} + \tilde{c}}\right]$$

$$= (T)^{\frac{3}{2}} \left[\frac{1 + \frac{\tilde{c}}{\tilde{T}_{\infty}}}{T + \frac{\tilde{c}}{\tilde{T}_{\infty}}}\right]$$
(4-8)

where  $\tilde{c} = 198.6^{\circ}R = 110.4^{\circ}K$  is Sutherland's constant.

## 4.3 Reynolds Number Examples

Experience has shown that some confusion arises when it comes to setting the parameter **reue** in CFL3D. Some examples may help.

Suppose the configuration to be analyzed has a body length of 48 inches and it is desired to run a simulation in which the Reynolds number based on the body length is  $6 \times 10^6$ . Thus,  $\tilde{L} = 48$  inches and  $Re_{\tilde{L}} = 6 \times 10^6$ .

To carry out the simulation, a grid is required. Suppose three people each create a grid. In the first person's grid, the length of the body is 48, i.e. the configuration is full size in the grid coordinate system. Thus,  $L_{ref}$  in grid 1 is 48. Therefore, by Equation (4-5),

$$\mathbf{reue} = \frac{6 \times 10^6}{48} \times 10^{-6} = 0.125$$
 (4-9)

The second person decides to normalize the grid such that the body has a length of 1. Thus,  $L_{ref}$  in grid 2 is 1 and by Equation (4-5),

$$\mathbf{reue} = \frac{6 \times 10^6}{1} \times 10^{-6} = 6.0 \tag{4-10}$$

The third person was given the geometry definition in centimeters. Grid 3 is therefore generated with a body length of 121.9 (48 inches  $\approx$  121.9 centimeters), so  $L_{ref} = 121.9$  and

$$\mathbf{reue} = \frac{6 \times 10^6}{121.9} \times 10^{-6} = 0.04922$$
(4-11)

### 4.4 Time and Time Step

Time is nondimensionalized as follows:

$$t = \frac{\tilde{t}\tilde{a}_{\infty}}{\tilde{L}_R}$$
(4-12)

If a particular dimensional time step (**dt** in "LT5 - Time Step Parameters" on page 21) is required, then this relation may be used to determine the corresponding input value of  $\Delta t$ .

For the time step,

 $\Delta t$  = desired time step, seconds

Therefore,

$$\Delta t = \frac{\Delta t \tilde{a}_{\infty}}{\tilde{L}_R}$$
(4-13)

## 4.5 Moving Mesh

These additional values may be needed for the nondimensionalization of the movingmesh parameters:

 $\tilde{f}$  = frequency, cycles/second

- $k_r$  = reduced frequency,  $k_r = \tilde{f}\tilde{L}/\tilde{a}_{\infty}$  Note: This definition of reduced frequency differs from the "standard" definition  $k_r = \tilde{f}\tilde{L}/|\tilde{\mathbf{V}}|_{\infty}$ . Also, note that the quantities  $\tilde{L}$  and  $L_{ref}$  used to nondimensionalize the moving mesh data need not be the *same* as those used to nondimensionalize x, y, z, and t. That is why  $L_{ref}$  is input explicitly in the moving mesh section of the input file.
- $k_g = \tilde{L}/(2\pi \tilde{a}_{\infty} \tilde{t}_0)$  = rate of growth of displacement, where  $\tilde{t}_0$  = the time to reach 63.2% of the maximum displacement, seconds

#### 4.5.1 Translational Motion

These dimensional parameters are also needed for the translational nondimensionalizations:

- $\tilde{u}_{trans} = x$  component of translational velocity, feet/second (needed for **itrans** = 1)
- $\tilde{v}_{trans} = y$  component of translational velocity, feet/second (needed for **itrans** = 1)

 $\tilde{w}_{trans} = z$  component of translational velocity, feet/second (needed for **itrans** = 1)

 $\tilde{x}_{max}$  = maximum x displacement, feet (needed for itrans > 1)

 $\tilde{y}_{max}$  = maximum y displacement, feet (needed for itrans > 1)

 $\tilde{z}_{max}$  = maximum z displacement, feet (needed for itrans > 1)

Again note that meters or inches, etc. could be substituted for feet in the above definitions, as long as it is done consistently. Also, not all of the above dimensional parameters are relevant to a given problem. For example, in sinusoidal plunging (**itrans** = 2),  $k_g$ ,  $\tilde{u}_{trans}$ ,  $\tilde{v}_{trans}$ , and  $\tilde{w}_{trans}$  are irrelevant.

4.5.1.1 Translation With Constant Velocity

For **itrans** = 1, input the following nondimensional parameters:

rfreq = 0.0

**utrans** =  $\tilde{u}_{trans} / \tilde{a}_{\infty}$  (Mach number of x translation)

**vtrans** =  $\tilde{v}_{trans} / \tilde{a}_{\infty}$ 

wtrans =  $\tilde{w}_{trans}/\tilde{a}_{\infty}$ 

The nondimensional displacement in the code is then governed by

$$x = u_{trans}t$$

$$y = v_{trans}t$$

$$z = w_{trans}t$$
(4-14)

# 4.5.1.2 Sinusoidal Plunging

For **itrans** = 2, input the following nondimensional parameters:

 $rfreq = k_r$ 

utrans = 
$$\tilde{x}_{max}/L_R$$

**vtrans** = 
$$\tilde{y}_{max}/L_R$$

wtrans = 
$$\tilde{z}_{max}/\tilde{L}_R$$

Note that,  $\tilde{x}_{max}/\tilde{L}_R$ , for example, is simply the maximum x displacement in "grid units". The nondimensional displacement in the code is then governed by

$$x = u_{trans} \sin\left(2\pi \frac{k_r t}{L_{ref}}\right)$$
  

$$y = v_{trans} \sin\left(2\pi \frac{k_r t}{L_{ref}}\right)$$
  

$$z = w_{trans} \sin\left(2\pi \frac{k_r t}{L_{ref}}\right)$$
  
(4-15)

If the time step is to be chosen so that there are N steps per cycle, the required input value of **dt** (see "LT5 - Time Step Parameters" on page 21) may be determined from

$$\mathbf{dt} = \frac{L_{ref}}{k_r \cdot N} \tag{4-16}$$

The corresponding dimensional time step is

$$\tilde{\Delta t} = \frac{1}{N \cdot \tilde{f}} \tag{4-17}$$

#### 4.5.1.3 Acceleration to Constant Displacement

For **itrans** = 3, input the following nondimensional parameters:

**rfreq** =  $k_g$ **utrans** =  $\tilde{x}_{max}/\tilde{L}_R$ 

**vtrans** = 
$$\tilde{y}_{max} / \tilde{L}_{F}$$

wtrans =  $\tilde{z}_{max}/\tilde{L}_R$ 

Note that,  $\tilde{x}_{max}/\tilde{L}_R$ , for example, is simply the maximum x displacement in "grid units". The nondimensional displacement in the code is then governed by

$$x = u_{trans} \left[ 1.0 - \exp\left(-2\pi \frac{k_r t}{L_{ref}}\right) \right]$$
  

$$y = v_{trans} \left[ 1.0 - \exp\left(-2\pi \frac{k_r t}{L_{ref}}\right) \right]$$
  

$$z = w_{trans} \left[ 1.0 - \exp\left(-2\pi \frac{k_r t}{L_{ref}}\right) \right]$$
  
(4-18)

Thus, with **utrans** =  $x_{max}$ , x will attain 63.2% of it's maximum value, i.e.

$$\frac{x}{x_{max}} = 1 - \frac{1}{e}$$
 (4-19)

at a nondimensional time of

$$t_0 = \frac{L_{ref}}{2\pi k_r} \tag{4-20}$$

rfreq for itrans = 3 may therefore be defined as

$$\mathbf{rfreq} = \frac{L_{ref}}{2\pi t_0} = \frac{\tilde{L}}{2\pi \tilde{a}_{\infty} \tilde{t}_0}$$
(4-21)

where  $\tilde{t}_0$  is the time (seconds) to reach 63.2% of the maximum displacement. Increasing **rfreq** has the effect of decreasing the time to attain a specified percentage of the asymptotic maximum displacement.

#### 4.5.2 Rotational Motion

These additional dimensional parameters are needed for the rotational nondimensionalizations:

 $\tilde{\omega}_x = x$  component of rotational velocity, revolutions/second (needed for **irotat** = 1)

$$\tilde{\omega}_{v} = y$$
 component of rotational velocity, revolutions/second (needed for **irotat** = 1)

 $\tilde{\omega}_z = z$  component of rotational velocity, revolutions/second (needed for **irotat** = 1)

$$\theta_{x, max}$$
 = maximum angular displacement about x axis, degrees (needed for irotat >1)

 $\tilde{\theta}_{y, max}$  = maximum angular displacement about y axis, degrees (needed for irotat >1)

 $\tilde{\theta}_{z, max}$  = maximum angular displacement about z axis, degrees (needed for irotat >1)

Again note that meters or inches, etc. could be substituted for feet in the above definitions, as long as it is done consistently. Not all of the above dimensional parameters are relevant to a given problem. For example, in sinusoidal pitching (**irotat** = 2)  $k_g$ ,  $\tilde{\omega}_x$ ,  $\tilde{\omega}_y$ , and  $\tilde{\omega}_z$  are irrelevant.

Using these dimensional values, the nondimensionalizations are performed as follows:

#### 4.5.2.1 Rotation With Constant Angular Velocity

For **irotat** = 1, input the following nondimensional parameters:

rfreq = 0.

omegax =  $\tilde{\omega}_x \tilde{L} / \tilde{a}_\infty$ 

omegay =  $\tilde{\omega}_{y}\tilde{L}/\tilde{a}_{\infty}$ 

omegaz =  $\tilde{\omega}_{z}\tilde{L}/\tilde{a}_{\infty}$ 

For turbomachinery type applications,  $\tilde{L}$  is typically chosen as the prop/blade diameter, so that  $\tilde{\omega}_x$ ,  $\tilde{\omega}_y$ ,  $\tilde{\omega}_z = M_{\infty}$  / (advance ratio). The angular displacement (radians) in the code is then governed by

$$\theta_{x} = 2\pi \frac{\omega_{x}t}{L_{ref}}$$

$$\theta_{y} = 2\pi \frac{\omega_{y}t}{L_{ref}}$$

$$\theta_{z} = 2\pi \frac{\omega_{z}t}{L_{ref}}$$
(4-22)

#### 4.5.2.2 Sinusoidal Variation of Angular Displacement

For **irotat** = 2, input the following parameters:

**rfreq** = 
$$k_r$$

**omegax** =  $\tilde{\theta}_{x, max}$ , degrees

**omegay** =  $\tilde{\theta}_{y, max}$ , degrees

**omegaz** =  $\tilde{\theta}_{z, max}$ , degrees

The rotational displacement (radians) in the code is then governed by

$$\begin{aligned} \theta_x &= \omega_x \frac{\pi}{180} \sin\left(2\pi k_r \frac{t}{L_{ref}}\right) \\ \theta_y &= \omega_y \frac{\pi}{180} \sin\left(2\pi k_r \frac{t}{L_{ref}}\right) \\ \theta_z &= \omega_z \frac{\pi}{180} \sin\left(2\pi k_r \frac{t}{L_{ref}}\right) \end{aligned}$$
(4-23)

If the time step is to be chosen so that there are N steps per cycle, the required input value of **dt** (see "LT5 - Time Step Parameters" on page 21) may be determined from

$$\Delta t = \frac{L_{ref}}{k_r N} \tag{4-24}$$

The corresponding dimensional time step is:

$$\tilde{\Delta t} = \frac{1}{N\tilde{f}} \tag{4-25}$$

# 4.5.2.3 Acceleration to a Constant Rotational Displacement

For **irotat** = 3, input the following nondimensional parameters:

rfreq = 
$$k_g$$
  
omegax =  $\tilde{\theta}_{x, max}$ , degrees  
omegay =  $\tilde{\theta}_{y, max}$ , degrees  
omegaz =  $\tilde{\theta}_{z, max}$ , degrees

•

The rotational displacement (radians) in the code is then governed by

$$\theta_{x} = \frac{\pi \omega_{x}}{180} \left[ 1 - \exp\left(-2\pi k_{r} \frac{t}{L_{ref}}\right) \right]$$

$$\theta_{y} = \frac{\pi \omega_{y}}{180} \left[ 1 - \exp\left(-2\pi k_{r} \frac{t}{L_{ref}}\right) \right]$$

$$\theta_{z} = \frac{\pi \omega_{z}}{180} \left[ 1 - \exp\left(-2\pi k_{r} \frac{t}{L_{ref}}\right) \right]$$
(4-26)

Thus, with  $\omega_x = \tilde{\theta}_{x, max}$ ,  $\theta_x$  will attain 63.2% of it's maximum value, i.e.

$$\theta_x / \theta_{x, max} = 1 - \frac{1}{e}$$
(4-27)

at a nondimensional time of

$$t_0 = \frac{L_{ref}}{2\pi k_r} \tag{4-28}$$

Increasing  $k_r$  has the effect of decreasing the time to attain a specified percentage of the asymptotic maximum displacement. **rfreq** for **irotat** = 3 may therefore be defined as

$$\mathbf{rfreq} = \frac{L_{ref}}{2\pi t_0} = \frac{\tilde{L}}{2\pi \tilde{a}_{\infty} \tilde{t}_0}$$
(4-29)

where  $\tilde{t}_0$  is the time (seconds) to reach 63.2% of maximum displacement.



# **CHAPTER 5**

This chapter describes the input and output files listed in Section 3.1. Typical names are given below; however, these names can be changed by the user to pertain to certain cases.

I/O FILES
grid.bin
plot3dg.bin
plot3dq.bin
cfl3d.out
cfl3d.res
cfl3d.turres
cfl3d.blomax
cfl3d.out15
cfl3d.prout
cfl3d.out20
ovrlp.bin
patch.bin
restart.bin

# 5.1 Input Files

There are two input files that *must* be available to CFL3D whenever a case is submitted. These files are the grid file (unit 1) and the input file (unit 5) defining the case. The latter file is described in detail in Chapter 3. A restart file (unit 2) *must* be available for a restart case; however, for a start from "scratch", it is not needed. In addition, there are two files that *may* be needed at submission depending on the problem being solved. When overlapped grids are used, the interpolation and boundary data file (unit 21) must be available. When patching is utilized, the patched boundary data file (unit 22) must be available.

# 5.1.1 Grid File

# grid.bin (unit 1)

The grid must be generated using grid-generation software (CFL3D does not generate grids). Take care when generating a grid. Grid spacings should vary smoothly; the grid should "look nice". (See the troubleshooting notes about grids in Chapter 10). Unless there is a good reason for not doing so, the grids, as well as all segments, should be multi-gridable (see Section 7.1.2 on page 127). Grid stretching should not be too severe if it can be avoided since it can degrade the accuracy and convergence of the code.

The grid file is in binary format. In the input file, the parameter **ngrid** is set as the total number of grids to be read in by CFL3D. The *sign* of **ngrid** indicates whether the grid will be read in CFL3D format or in PLOT3D format. (Note that TLNS3D<sup>42</sup> uses PLOT3D format.) CFL3D-type grids always have **alpha** measured in the x - z plane, with z as the "up" direction (i.e. **alpha** = 90° would give a free-stream velocity vector in the positive z direction). PLOT3D-type grids may be utilized with either z being "up" (**alpha** measured in the x - z plane) or y being "up" (**alpha** measured in the x - y plane) depending on the value of **ialph** (see "LT3 - Flow Conditions" on page 19). See Appendix I for an illustrated definition of  $\alpha$ . When PLOT3D-type grids are input, they are immediately converted internal to the code to CFL3D format. Thus the actual computations are *always* carried out internally as if z is "up".

If **ngrid** > 0, the following CFL3D grid format is used to read in the grid(s):

If **ngrid** < 0, the following PLOT3D grid format is used to read in the grid(s):

It does not matter which grid format is used. What *does* matter is the particular orientation of *i*, *j*, and *k* relative to the body and flow field, since the matrix inversions are done in a particular order. Experience has indicated that it is usually best (from speed of convergence perspective) to have the *k* direction be the primary viscous direction (i.e., the primary direction of the grid lines normal to the body). If bodies exist with more than one grid line orientation, then choose one of them to serve as the indicator of the "primary" viscous direction. Also, it is usually best if the *j* direction is taken as the streamwise direction and the *i* direction as the spanwise direction. Hence, for an existing grid with *i* is in the streamwise direction, *j* in the normal direction, and *k* in the spanwise direction, it might be a good idea to first swap the indices:  $i \rightarrow j$ ,  $j \rightarrow k$ , and  $k \rightarrow i$ , to put them in CFL3D's "preferred" orientation before running. The new grid created can be either CFL3D-type or PLOT3D-type. Finally, note that it is not always necessary to follow this guideline; performance is case dependent. Just be sure to satisfy the right-hand rule at all times, as discussed in the next section.

For 2-D grids (i2d=1), idim must be 2 (i.e., the grid must be described by 2 repeated planes). We recommend making the grid-distance between the planes = 1, although any distance is okay provided that it is accounted for in the **sref** term in the input file.

# 5.1.2 The Right-Hand Rule

The right-hand rule *must* be obeyed by the grid coordinates (x, y, z) and the grid indices (i, j, k). Error messages and an immediate halt to execution will result if it is not. As a reminder of this important rule, consider the system of axes in Figure 5-1. Imagine pointing the right-hand index finger along the positive x axis. Then curl the rest of the right-hand fingers toward the positive y axis. The thumb, pointing straight out, is now in the positive z direction.

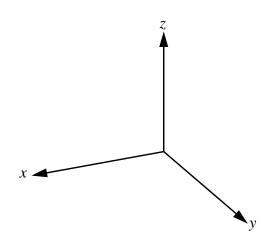


Figure 5-1. Three-dimensional system of axes showing positive x, y, and z directions.



Figure 5-2. Index direction examples.

Keep in mind that the same rule applies to the indices i, j, k. Suppose the positive i, j, and k directions run along the positive x, y, and z axes, respectively, as drawn in Figure 5-1. This is the simplest case and is illustrated in Figure 5-2(a). Suppose it is necessary, however, for the positive i index direction to run in the opposite direction of positive x as drawn in Figure 5-2(b). (This means, for example, that if x varies from 0.0 to 1.0 and **idim** = 5, then at i = 1, x = 1.0 and at i = 5, x = 0.0.) One of the other index directions

must also be changed to maintain the right-hand rule. In Figure 5-2(b), the j index direction has been changed.

Also, *i*, *j*, and *k* do not *have* to be in the *x*, *y*, and *z* directions, respectively. Any permutation is valid as long as the right-hand rule is upheld. (However, keep in mind the discussion above in Section 5.1.1 as well as *Note* (6) describing the **ivisc**(m) input on page 27 when setting the index directions.)

5.1.3 Restart File

#### restart.bin (unit 2)

The restart file is in binary format. It must be available at the beginning of the program execution for a restart of a previous solution (**irest** > 0). *The restart file is overwritten during program execution; therefore, before resubmitting the case, be sure to save any restart files that may be needed later!* The binary restart file has the following format. (Note that this Version 5.0 format is different from the formats in Version 4.1 and earlier versions.)

General case information, the q array, and the turbulence quantities are written as follows (nblk is the total number of blocks):

```
do nb=1, nblk
      write(3) titlw, xmachw, jdim, kdim, idim, alphw, reuew, ntr, time
С
      if (nb.eq.1) then
                              n=1, ntr), (clw(n),
         write(3) (rms(n),
                                                       n=1,ntr),
                   (cdw(n),
                              n=1, ntr), (cdpw(n),
                                                       n=1,ntr),
                   (cdvw(n),
                               n=1,ntr),(cxw(n),
                                                       n=1, ntr),
                   (cyw(n),
                               n=1,ntr),(czw(n),
                                                       n=1, ntr),
                   (Cmxw(n),
                               n=1, ntr), (cmyw(n),
                                                      n=1, ntr),
                              n=1, ntr), (fmdotw(n),
                                                      n=1,ntr),
                   (cmzw(n),
                   (cftmomw(n), n=1,ntr), (cftpw(n),
                                                      n=1,ntr),
                   (cftvw(n), n=1,ntr), (cfttotw(n), n=1,ntr)
      end if
С
      write(3) ((((q(j,k,i,l,nb),j=1,jdim-1),k=1,kdim-1),
                      i=1,idim-1),l=1,5)
С
      write(3) iv1, iv2, iv3
      if (nb.eq.1) then
         write(3) (rmstr1(n), n=1, ntr), (rmstr2(n), n=1, ntr),
                   (nneg1(n), n=1,ntr), (nneg2(n), n=1,ntr)
      end if
С
      if (iv1.ge.2 .or. iv2.ge.2 .or. iv3.ge.2) then
        write(3) (((vist3d(j,k,i,nb),j=1,jdim-1),k=1,kdim-1),i=1,idim-1)
      end if
      if (iv1.ge.4 .or. iv2.ge.4 .or. iv3.ge.4) then
         write(3) ((((tursav(j,k,i,m,nb),j=1,jdim-1),k=1,kdim-1),
                       i=1,idim-1),m=1,2)
         write(3) (((smin(j,k,i,nb),j=1,jdim-1),k=1,kdim-1),i=1,idim-1)
         if (iv1.eq.4 .or. iv2.eq.4 .or. iv3.eq.4) then
         write(3)(((xjb(j,k,i,nb),j=1,jdim-1),k=1,kdim-1),i=1,idim-1)
         write(3)(((xkb(j,k,i,nb),j=1,jdim-1),k=1,kdim-1),i=1,idim-1)
         write(3)(((blnum(j,k,i,nb),j=1,jdim-1),k=1,kdim-1),i=1,idim-1)
```

```
end if
          if (iv1.eq.9 .or. iv2.eq.9 .or. iv3.eq.9
                                                        .or.
              iv1.eq.13.or. iv2.eq.13.or. iv3.eq.13) then
          write(3) (((cmuv(j,k,i,nb),j=1,jdim-1),k=1,kdim-1),i=1,idim-1)
          end if
      end if
      end do
С
      if (dt.gt.0) then
         write(3) iflagg
         if (iflagg.eg.1 .or. iflagg.eg.3) then
           do nb=1, nblk
              write(3) jdim,kdim,idim
              write(3) ((((qc0(j,k,i,l,nb),j=1,jdim-1),k=1,kdim-1),
                             i=1,idim-1),l=1,5)
           end do
         end if
С
         if (iflagg.eq.2 .or. iflagg .eq.3) then
           do nb=1, nblk
              write(3) iunst
              if (iunst .ne. 0) then
                write(3) jdim,kdim,idim
                write(3) itrans,rfreqt,xorig,yorig,zorig,xorig0,yorig0,
                           zorig0, utrans, vtrans, wtrans, dxmx, dymx, dzmx,
                           itransmc, rfreqtmc, xorigmc, yorigmc, zorigmc,
                           xorig0mc, yorig0mc, zorig0mc, utransmc, vtransmc,
                           wtransmc, xmc, ymc, zmc, dxmxmc, dymxmc, dzmxmc,
                           irotat, rfreqr, thetax, thetay, thetaz,
                           omegax, omegay, omegaz, dthxmx, dthymx, dthzmx,
                           irotatmc, rfreqrmc, thetaxmc, thetaymc, thetazmc,
                           omegaxmc, omegaymc, omegazmc, dthxmxmc, dthymxmc,
                           dthzmxmc,time2,time2mc,dt
              end if
           end do
         end if
      end if
Where
                 Indicates
                 a real array of length 20 describing the case
titlw
                 Mach number
xmachw
jdim, kdim, idim
                 grid dimensions
                 angle of attack
alphw
                 Reynolds number
reuew
                 number of iterations
ntr
                 time (for time-accurate computations)
time
                 residual
rms
                 lift coefficient
clw
                 drag coefficient
cdw
                 force coefficient in y direction
суw
                 moment coefficient about y
cmyw
                 pressure drag coefficient
cdpw
                 viscous drag coefficient
cdvw
                 force coefficient in x direction
CXW
                 force coefficient in z direction
CZW
                 moment coefficient about x
CMXW
```

Cmzw	moment coefficient about z					
fmdotw	mass flow through control surface					
cftmomw	magnitude of resultant momentum force on control surface					
cftpw	magnitude of resultant pressure force on control surface					
cftvw	magnitude of resultant viscous force on control surface					
cfttotw	magnitude of resultant (pressure + momentum + viscous) force on					
	control surface					
q	nondimensional primitive variables; $\mathbf{q} = [\rho, u, v, w, p]$					
iv1,iv2,iv3	ivisc(1), ivisc(2), and ivisc(3)					
rmstr1	residual for first turbulence model equation (if used)					
rmstr2	residual for second turbulence model equation (if used)					
nneg1	number of points at which first turbulence model variable needed to be					
	limited					
nneg2	number of points at which second turbulence model variable needed to					
	be limited					
vist3d	nondimensional turbulent eddy viscosity $\mu_T$					
tursav	nondimensional turbulence quantities for field-equation turbulence					
	models (varies with model); tursav(2) is also used to store nearest					
	wall <i>i</i> location for the Baldwin-Barth turbulence model					
smin	minimum distance function for field-equation turbulence models					
xjb	nearest wall j location for Baldwin-Barth turbulence model					
xkb	nearest wall k location for Baldwin-Barth turbulence model					
blnum	nearest wall block number for Baldwin-Barth turbulence model					
cmuv	variable $C_{\rm u}$ coefficient for Girimaji turbulence model					
qc0	$\mathbf{Q}$ values at previous time step					
iflagg	parameter describing temporal order and whether grid is moving or					
TTTAGG	stationary					
	= 0 not second order, stationary grid					
	= 1 second order, stationary grid					
	= 2 not second order, moving grid					
	= 3 second order, moving grid					
	× 00					

The dynamic grid parameters in the restart file are not described here.

# 5.1.4 Grid-Overlapping and Patching Data Files

# ovrlp.bin (unit 21)

The ovrlp.bin file is a binary file, which is only used for grid-overlapping cases. It is always listed at the top of the input deck. The file contains the interpolation stencils for overlapped grids and is created by running MaGGiE.

#### patch.bin (unit 22)

The patch.bin file is a binary file, which is only used for grid-patching cases. It is also always listed at the top of the input deck. The file contains the interpolation stencils for patched grids and is created by running ronnie.

# 5.2 Output Files

CFL3D produces several output files. The following sections describe the contents and formats of these files.

#### 5.2.1 PLOT3D Files

The PLOT3D files are binary files (or ascii, with the modification to the code as described in the footnote at the beginning of Chapter 3) which are input to flow visualization programs like PLOT3D or FAST. Partial or entire flow-field regions may be specified by the user in the input file. The parameter **nplot3d** specifies the number of regions to be written to the files and "LT28 - PLOT3D Output Specifications" on page 39 indicates the desired index ranges. Data may also be extracted from these files for post-processing programs written by the user. In order to read these files for such a purpose, use the formats below. The number of grids to be read is indicated by nplots and id, jd, kd are the number of points to be read on each grid. Note that these files can be specified as grid-point data files (**iptype =** 0) or cell-center data files (**iptype =** 1).

#### plot3dg.bin (unit 3)

For 3-d grids,

For 2-d grids,

#### plot3dq.bin (unit 4)

The flow-field data file has the following formats. For 3-d grids,

```
write(4) nplots
write(4) (id(n), jd(n), kd(n), n=1, nplots)
do n=1, nplots
write(4) xmach, alpha, reue, time
write(4) ((((q(j,k,i,m,n), i=1, id(n)), j=1, jd(n)), k=1, kd(n)), m=1, 5)
end do
```

For 2-d flow-field data files,

```
write(4) nplots
write(4) (jd(n),kd(n),n=1,nplots)
do n=1,nplots
write(4) xmach,alpha,reue,time
write(4) (((q(j,k,m,n),j=1,jd(n)),k=1,kd(n)),m=1,4)
end do
```

Note that five flow-field variables are output for 3-d grids, whereas four variables are output for 2-d grids. Also, note that q here is the array of conserved variables **Q**.

When using **iptype** = 0, be aware that the solution on block edges and corners may be inaccurate. This is because CFL3D solves for cell-center values. Values at grid points must be reconstructed from cell-center data and the reconstruction process at edges/corners is not unique.

When iptype = 2, CFL3D outputs turbulence quantity information rather than the conserved variables **Q**. See the note in Section 3.28 on page 39.

When **iptype** > 2, a PLOT3D function file is output in place of the **Q** file. The function file format for 3-d grids is

```
write(4) nplots
write(4) (id(n), jd(n), kd(n), 1, n=1, nplots)
do n=1, nplots
write(4) (((f(j,k,i,n),i=1,id(n)), j=1, jd(n)), k=1, kd(n))
end do
```

For 2-d function files,

```
write(4) nplots
write(4) (jd(n),kd(n),1,n=1,nplots)
do n=1,nplots
write(4) ((f(j,k,n),j=1,jd(n)),k=1,kd(n))
end do
```

Note that function files are currently written for a single function only. See Section 3.28 on page 39 for information on what functions are currently available.

# 5.2.2 Ascii Format Files

# cfl3d.out (unit 11)

The cfl3d.out file is the primary output file. It echoes the input file with the appropriate bookkeeping changes, such as updates in block numbers when coarser grids are generated internally. The file then gives a "blow-by-blow" message about what steps have been performed: the grid is read in, the metrics are computed, etc. until finally the other output files are written. Another important function of this file is to print any warnings and error messages. For example, it will reveal any mismatches in one-to-one block connections, signaling a problem with the grid or input file. It is a good idea to run a new job for a few iterations only and then to study this file to make sure everything looks as expected.

# cfl3d.res (unit 12)

When **ihstry** = 0, the cfl3d.res file contains the residual, lift, drag, side force, and moment histories. When **ihstry** > 0, the residual, mass flow, pressure force, viscous force, and momentum force are given instead. This type of file has been used for all of the "iteration history" plots presented in this manual. The following sample format shows a portion of the cfl3d.res file (when **ihstry** = 0) for the flat plate test case discussed in Section 9.1.5 on page 173. When sub-iterations are used in a time-accurate computation, only the final residuals and forces from the last sub-iteration of each time step are output to unit 12. (For the sub-iteration residual history file, see page 76.)

```
turbulent flat plate (plate from j=17-65, prior to 17 is symmetry)
Mach= 0.2000E+00, alpha= 0.0000E+00, ReUe=
                                            0.6000E+07
Final res= 0.3623E-09
Final cl,cd,cy,cmy= -0.1869E-02 0.2827E-02 0.0000E+00 0.3174E-03
  800 it
          log(res)
                       cl
                                cd
                                                     CV
                                                                   cmy
         -0.15700E+02
                       0.25376E-12
                                     0.67093E+00
                                                   0.00000E+00 -0.12688E-12
      1
      2
         -0.59009E+01 0.13542E-01
                                     0.40383E-01
                                                  0.00000E+00 -0.72299E-02
         -0.94402E+01 -0.18662E-02
    799
                                     0.28275E-02
                                                   0.00000E+00
                                                                0.31613E-03
         -0.94409E+01 -0.18695E-02
                                     0.28273E-02
                                                   0.00000E+00
                                                                0.31742E-03
    800
```

Here, res is the L-2 norm of the residual for the equation for density.

# cfl3d.turres (unit 13)

The cfl3d.turres file contains the residual history for the one or two-equation turbulence models (**ivisc** > 3). The sample below shows part of the cfl3d.turres file for the flat plate test case discussed in Section 9.1.5 on page 173.The log(turres1) quantity is the log of the residual for all one-equation models or the log of the residual for the  $\omega$  or  $\varepsilon$ equation for two-equation models. Log(turres2) is the log of the residual for the k equation for two-equation models and is not used for one-equation models. Nneg1 and nneg2 indicate how many field points require limiting (to avoid going negative) during that iteration. They should be zero or, at most, a small fraction of the total number of points for a converged solution. If they are large, they indicate a problem with convergence. (They may be large during start-up, as in this example, but should always go to zero or small values as the solution progresses.) When sub-iterations are used in a time-accurate computation, only the final residuals from the *last* sub-iteration of each time step are output.

#### cfl3d.blomax (unit 14)

The cfl3d.blomax file provides information about the Baldwin-Lomax turbulence model (ivisc = 2).

#### cfl3d.out15 (unit 15)

The cfl3d.out15 file is a secondary output file. It provides more in depth information about the progression of a case. It is primarily used by the programmer for debugging purposes.

#### cfl3d.prout (unit 17)

The cfl3d.prout file contains the output as specified in the input file with the **nprint** parameter and "LT30 - Print Out Specifications" on page 41. Note that this file can be written for values at the grid points (**iptype** = 0) or cell-centers (**iptype** = 1). Keep in mind that this file can grow quite large rather quickly, so limit the regions specified for output. The cfl3d.prout file will also contain the control surface data if **ncs** > 0 (see Section 3.31 on page 42). (The sample below is from the flat plate test case discussed in Section 9.1.5 on page 173.) Note that dn, Cf, Ch, and yplus are only output when **iptype** = 0.

```
turbulent flat plate (plate from j=17-65, prior to 17 is symmetry)
     Mach
             alpha
                         beta ReUe Tinf,dR
                                                      time
                    0.00000 0.600E+07 460.00000
  0.20000 0.00000
                                                  0.00000
BLOCK 1 IDIM, JDIM, KDIM=
                                        97
                             2
                                   65
                            Z
  І Ј К
                    Y
            X
                                    U/Uinf
                                               V/Vinf
                                                         W/Winf
                                                                   P/Pinf
                       cp
       MACH cp tur. vis.
1 1 -.3333E+00 0.0000E+00 0.0000E+00 0.9987E+00 0.0000E+00 -.7200E-07
T/Tinf
   1 1
0.1000E+01 0.1000E+01 0.1997E+00 0.5992E-02 0.2054E-02
   1 2
          1 -.3125E+00 0.0000E+00 0.0000E+00 0.9989E+00 0.0000E+00 0.1323E-21
0.1000E+01 0.1000E+01 0.1998E+00 0.6012E-02 0.1379E-02
   1 64
          1 0.9792E+00 0.0000E+00 0.0000E+00 0.0000E+00 0.0000E+00 0.0000E+00
0.9999E+00 0.1008E+01 0.0000E+00 -.3123E-02 0.1966E-09
   1 65 1 0.1000E+01 0.0000E+00 0.0000E+00 0.3675E-02 0.0000E+00 0.5962E-08
0.9999E+00 0.1008E+01 0.7320E-03 -.3607E-02 0.1963E-09
```

IJK х Υ 7. dn P/Pinf T/Tinf Cf Ch yplus 2 1 1 -0.31250E+00 0.00000E+00 0.00000E+00 0.10000E-05 0.10002E+01 0.10000E+01 0.00000E+00 0.00000E+00 0.00000E+00 1 -0.29167E+00 0.00000E+00 0.00000E+00 0.10000E-05 0.10002E+01 1 3 0.10000E+01 0.00000E+00 0.00000E+00 0.00000E+00 1 63 1 0.95833E+00 0.00000E+00 0.00000E+00 0.10000E-05 0.99994E+00 0.10080E+01 0.24677E-02 0.33657E-02 0.20858E+00 1 0.97916E+00 0.00000E+00 0.00000E+00 0.10000E-05 0.99991E+00 1 64 0.10080E+01 0.24660E-02 0.24780E-02 0.20851E+00 BLOCK 1 IDIM, JDIM, KDIM= 2 65 97 тлк Х Y Ζ U/Uinf V/Vinf W/Winf P/Pinf T/Tinf MACH ср tur. vis. 1 49 1 0.6667E+00 0.0000E+00 0.0000E+00 0.0000E+00 0.0000E+00 0.0000E+00 0.1000E+01 0.1008E+01 0.0000E+00 0.1029E-02 0.2188E-09 1 49 2 0.6667E+00 0.0000E+00 0.1000E-05 0.8330E-02 0.0000E+00 -.1105E-06 0.1000E+01 0.1008E+01 0.1659E-02 0.1029E-02 0.2112E-07 1 49 96 0.6667E+00 0.0000E+00 0.9594E+00 0.9987E+00 0.0000E+00 0.2794E-03 0.1000E+01 0.1000E+01 0.1997E+00 0.2587E-02 0.7274E-02 1 49 97 0.6667E+00 0.0000E+00 0.9837E+00 0.9987E+00 0.0000E+00 0.2672E-03 0.1000E+01 0.1000E+01 0.1997E+00 0.2609E-02 0.7274E-02 I J K Х Υ 7. dn P/Pinf T/Tinf yplus Cf Ch 1 49 1 0.66667E+00 0.00000E+00 0.00000E+00 0.10000E-05 0.10000E+01 0.10080E+01 0.25533E-02 0.16414E-02 0.21218E+00 1 49 97 0.66667E+00 0.00000E+00 0.98367E+00 0.24255E-01 0.10001E+01 0.10000E+01 0.23455E-13 0.60677E-09 0.15760E-01

A sample of the type of control surface information that can be expected when  $\mathbf{ncs} > 0$  is shown here (this is no longer for the flat plate test case):

Control Surface 1

Grid 1 (Block 1) i = 25, 25 j = 1, 33 k = 1, 25 iwall = 0 Normal = 1 x-area = 0.2687E+02 y-area = 0.1842E+00 z-area = -0.1842E+00 total-area = 0.2687E+02 Mass averaged properties = P/Pinf 0.9446E+00 Pt/Pinf = 0.1022E+01 T/Tinf = 0.9840E+00 Tt/Tinf = 0.1007E+01 Mach number = 0.3367E+00 Mass flow / (rhoinf\*vinf) = 0.42300E+02 x-force y-force z-force resultant-force lift-force drag-force -0.52122E+02 -0.23196E+01 0.23205E+01 0.52226E+02 0.23205E+01 Pressure force -0.52122E+02 Thrust force 0.12872E+03 0.23886E+02 -0.23848E+02 0.13307E+03 -0.23848E+02 0.12872E+03 Total force 0.76598E+02 0.21566E+02 -0.21527E+02 0.82437E+02 -0.21527E+02 0.76598E+02

This information is printed at the end of the cfl3d.prout file.

#### cfl3d.out20 (unit 20)

The cfl3d.out20 file is an auxiliary output file providing information when **iunst** > 0. It provides the unsteady pressures from forced oscillations. This file is primarily used by the programmer for debugging purposes. To output this file, hard-wire irite = 1 in subroutine resp in rhs.f.

#### cfl3d.subit\_res (unit 23)

The cfl3d.subit\_res file is similar to the cfl3d.res file, except that it outputs residuals for all sub-iterations when time-accurate computations with sub-iterations are performed. This file is hard-wired with this name and is *not* in the file list at the top of the

input file. However, whereas unit 12 outputs only  $R(\mathbf{q}^m)$  for the equation for density (in Equation (B-12) or Equation (B-22) of Appendix B), unit 23 outputs the entire right-hand side of Equation (B-12) or Equation (B-22) for the equation for density. Therefore, the residuals in the two files do not match.

#### cfl3d.subit\_turres (unit 24)

The cfl3d.subit\_turres file is similar to the cfl3d.turres file, except that it outputs field-equation turbulence model residuals for all sub-iterations when time-accurate cases with sub-iterations are run. This file is hard-wired with this name and is *not* in the file list at the top of the input file.

#### cfl3d.dynamic\_patch (unit 25)

The cfl3d.dynamic\_patch file is output when using the moving grid option with dynamic patching. This file is hard-wired with this name and is *not* in the file list at the top of the input file.

#### 5.2.3 PLOT3D Function Files

If desired, CFL3D can output PLOT3D function files (for 3-d cases) with surface grid,

 $y^{+}$ , eddy viscosity, and normal spacing at the first grid point above all viscous walls. These files are:

#### surf\_xyz.fmt (unit 28)

surf\_y+.fmt (unit 29)

surf\_vist.fmt (unit 30)

#### surf\_dn.fmt (unit 31)

For 2-d cases, a single file is output:

# surf\_y+\_2d.fmt (unit 28)

This file is in column format, containing all of the above information, with headers between different segments. By default, this option to write these files is currently turned *off.* To activate this option, the user must change ifunc from 0 to 1 in subroutine yplusout in lbcx.f.

# **CHAPTER 6**

# **Boundary Conditions**

The equations of motion require boundary conditions on all sides of the domain in which the solution is to be obtained, as well as on all surfaces of any objects which lie within the domain. For CFD methods, this implies that boundary conditions must be applied at each face of each computational block. Here, the term boundary conditions is used somewhat loosely as it refers to both physical conditions (e.g. solid wall, symmetry plane, etc.) and interfaces between adjacent or overlapped blocks.

In CFL3D, the boundary condition data are located in the arrays qi0(jdim,kdim,5,4), qj0(kdim,idim-1,5,4), and qk0(jdim,idim-1,5,4) for the *i*, *j*, and *k* directions, respectively. The third index in the arrays indicates the position for the five flow-field primitive variables (Equation (E-3)). The fourth dimension indicates the storage location for the boundary value positions; indices 1 and 2 are the positions for the boundary data at the left face and indices 3 and 4 are the positions for the boundary data at the right face. "Left face" here refers to the *i* = 1, *j* = 1, and *k* = 1 faces. "Right face" refers to the *i* = idim, *j* = jdim, and *k* = kdim faces.

Note that the locations defining the boundary condition regions in "LT14 - I0 Boundary Condition Specification" through "LT19 - KDIM Boundary Condition Specification" are *grid points*, but the actual boundary conditions are applied at *cell-face centers*.

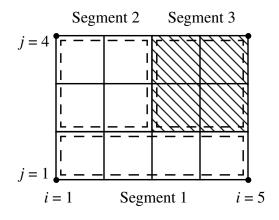


Figure 6-1. Boundary condition segments.

For example, say that Figure 6-1 represents the surface of a wing. At the unshaded cells, it is desired to apply a heated wall boundary condition, while at the shaded cells it is

desired to apply an adiabatic wall boundary condition. One way to accomplish this objective would be to divide the  $\kappa 0$  boundary condition into three segments, as indicated by the dashed lines in the figure. The CFL3D input file would look like this:

#### Line Type

18	K0: GRID	SEGMENT	BCTYPE	ISTA	IEND	JSTA	JEND	NDATA
	1	1	2004	1	5	1	2	2
	twtype	Cq						
	1.6	0.0						
	1	2	2004	1	3	2	4	2
	twtype	Cq						
	1.6	0.0						
	1	3	2004	3	5	2	4	2
	twtype	Cq						
	0.0	0.0						

Note that, for example, the *grid points* i = 1 to 5, j = 1 to 2 define the boundary of the *cells* at which the boundary condition type for segment 1 is applied.

Two types of boundary condition representations are employed in CFL3D, namely, cell-center and cell-face. For cell-center type boundary conditions, the flow-field variables are specified at "ghost" points corresponding to two cell-center locations analytically extended outside the grid as illustrated in Figure 6-2. For example, qi0(j,k,l,1) contains the boundary data at the ghost point locations right next to the i = 1 cell-center data and qi0(j,k,l,2) contains the boundary data in the second set of ghost point locations at the left boundary. Likewise, qi0(j,k,l,3) contains the boundary data at the ghost point locations at the boundary data at the ghost point locations right next to the i = idim cell-center data and qi0(j,k,l,4) contains the boundary data in the second set of ghost point locations at the second set of ghost point locations at the right boundary. The same boundary data location definitions apply for the qj0 and qk0 arrays.

For cell-face type boundary conditions, the flow-field variables and their gradients are specified at the cell-face boundary as illustrated in Figure 6-3. For cell-face type boundary conditions, qi0(j,k,l,1) contains the flow-field variable data at the i = 1 grid cell-face and qi0(j,k,l,2) contains the flow-field variable gradients at the i = 1 grid cell-face. Likewise, qi0(j,k,l,3) contains the flow-field variable data at the i = idim grid cell-face and qi0(j,k,l,4) contains the flow-field variable gradients at the i = idim grid cell-face. The same boundary data location definitions apply for the qj0 and qk0 arrays.

Note that when running in 2-d mode (i2d = 1), it *does not matter* what boundary conditions are used on the i = 1 and i = idim faces (provided that the boundary condition itself does not require more than one interior cell). Bctype is generally set to 1002 on these faces.

Boundary conditions for the eddy viscosity are stored in vi0, vj0, and vk0, while boundary conditions for the field equation turbulence models are stored in ti0, tj0, and tk0. Both of these are cell-center type, although only the first ghost points (1 and 3) are used. The field equation turbulence model boundary conditions are described in more detail in Section H.7 on page 296.

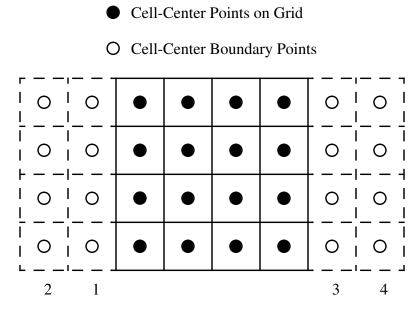


Figure 6-2. Cell-center type boundary locations.

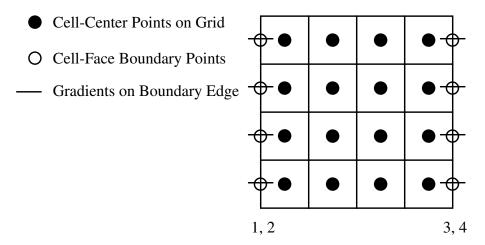


Figure 6-3. Cell-face type boundary conditions.

# 6.1 Physical Boundary Conditions

The 1000 series of boundary condition types are used for physical boundary conditions and are specified at any of a block's six faces in "LT14 - I0 Boundary Condition Specification" on page 32 through "LT19 - KDIM Boundary Condition Specification" on page 35. In each case, **ndata** = 0, since no additional information is required for implementation of the condition. The 1000 series boundary condition types currently supported are as follows:

bctype_	boundary condition
1000	free stream
1001	general symmetry plane
1002	extrapolation
1003	inflow/outflow
1004	(no longer available, use 2004 instead)
1005	inviscid surface
1008	tunnel inflow
1011	singular axis – half-plane symmetry
1012	singular axis – full plane
1013	singular axis – partial plane

# 6.1.1 Free Stream

#### bctype 1000

The free stream boundary conditions are cell-center type boundary conditions. The five flow-field variables for both sets of ghost points are set equal to the initial values, which are:

$$\rho_{initial} = 1.0$$

$$u_{initial} = M_{\infty} \cos \alpha \cos \beta$$

$$v_{initial} = -M_{\infty} \sin \beta$$

$$w_{initial} = M_{\infty} \sin \alpha \cos \beta$$

$$p_{initial} = \rho_{initial} (a_{initial})^2 / \gamma$$
(6-1)

where  $a_{initial} = 1.0$ .

6.1.2 General Symmetry Plane

bctype 1001

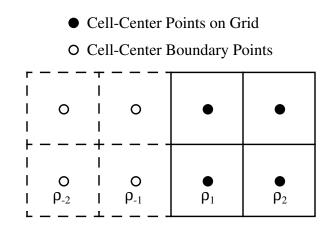


Figure 6-4. Symmetry boundary conditions (bctype 1001) for density.

The symmetry plane boundary conditions are cell-center type boundary conditions. As the name implies, symmetry is assumed across an axis. The ghost point density values are set equal to their "mirror image" counterparts. With the points defined as in Figure 6-4,

$$\begin{aligned} \rho_{-1} &= \rho_1 \\ \rho_{-2} &= \rho_2 \end{aligned} \tag{6-2}$$

(Note that  $\rho_{-1} = qi0(j,k,1,1)$  at the i = 1 face, for example.) The pressure values are assigned in the same way.

The velocity components at the ghost cells are obtained as follows. Consider ghost cells at an i = 1 face. Note that the normalized contravariant velocity  $\overline{U}$  is normal to an i = constant face. Let  $\overline{U}_1$  be the normalized contravariant velocity at cell center 1 in Figure 6-4. For an i symmetry plane,  $\overline{U}$  must have opposite signs on each side of the plane. Thus

$$u_{-1} = u_1 - 2\hat{\xi}_x \overline{U}_1$$

$$v_{-1} = v_1 - 2\hat{\xi}_y \overline{U}_1$$

$$w_{-1} = w_1 - 2\hat{\xi}_z \overline{U}_1$$
(6-3)

where  $\hat{\xi}_x$ ,  $\hat{\xi}_y$ , and  $\hat{\xi}_z$  are the metrics (unit normals) at the i = 1 face.

Since  $(\hat{\xi}_x)^2 + (\hat{\xi}_y)^2 + (\hat{\xi}_z)^2 = 1$ ,  $\overline{U}_{-1} = \overline{U}_1$ . The velocity components at ghost cell center –2 are set in a similar manner using the data at cell center 2. (Note that this is now a "general" symmetry condition, applicable to any plane, not just the x - z plane as in earlier versions of the code. Consequently, there is no longer a **bctype** 1006.)

#### 6.1.3 Extrapolation

#### bctype 1002

The extrapolation boundary conditions are cell-center type boundary conditions. The ghost points are extrapolated from the computational domain. Based on the locations of  $\rho_1$ ,  $\rho_{-1}$  and  $\rho_{-2}$  depicted in Figure 6-4, the extrapolated values would be

$$\rho_{-1} = \rho_1$$

$$\rho_{-2} = \rho_1$$
(6-4)

The same zeroth order extrapolation is used for the boundary values of the other four flowfield variables.

#### 6.1.4 Inflow/Outflow

#### bctype 1003

The inflow/outflow boundary conditions are cell-face type boundary conditions. In the far field, the velocity normal to the far boundary (pointing *out* of the grid) and the speed of sound are obtained from two locally 1-d Riemann invariants:

$$\mathbf{R} \equiv \bar{u} \pm \frac{2a}{\gamma - 1} \tag{6-5}$$

where the boundary is considered a surface of constant  $\xi$  (in this example) with decreasing  $\xi$  corresponding to the interior of the domain and

$$\bar{u} = \overline{U} - \frac{\xi_t}{|\nabla \xi|}$$
(6-6)

$$\overline{U} = \frac{\xi_x}{|\nabla\xi|} u + \frac{\xi_y}{|\nabla\xi|} v + \frac{\xi_z}{|\nabla\xi|} w + \frac{\xi_t}{|\nabla\xi|}$$
(6-7)

 $R^-$  can be evaluated locally from conditions outside the computational domain and  $R^+$  can be determined locally from inside the domain. The normal velocity and speed of sound are determined from

$$\bar{u}_{face} = \frac{1}{2} (R^{+} + R^{-})$$

$$a_{face} = \frac{\gamma - 1}{4} (R^{+} - R^{-})$$
(6-8)

The Cartesian velocities are determined by decomposing the normal and tangential velocity vectors:

$$u_{face} = u_{ref} + \frac{\xi_x}{|\nabla \xi|} (\bar{u}_{face} - \bar{u}_{ref})$$

$$v_{face} = v_{ref} + \frac{\xi_y}{|\nabla \xi|} (\bar{u}_{face} - \bar{u}_{ref})$$

$$w_{face} = w_{ref} + \frac{\xi_z}{|\nabla \xi|} (\bar{u}_{face} - \bar{u}_{ref})$$
(6-9)

For inflow,  $ref \Rightarrow \infty$ . For outflow, *ref* represents the values from the cell inside the domain adjacent to the boundary.

The sign of the normal velocity  $\overline{U}_{face} = \overline{u}_{face} + \xi_t / |\nabla \xi|$  determines whether the condition is at inflow ( $\overline{U}_{face} < 0$ ) or outflow ( $\overline{U}_{face} > 0$ ). The entropy  $p/\rho^{\gamma}$  is determined using the value from outside the domain for inflow and from inside the domain for outflow. The entropy and speed of sound are used to determine the density and pressure on the boundary:

$$\rho_{face} = \left[\frac{\left(a_{face}\right)^2}{\gamma s_{face}}\right]^{\frac{1}{\gamma - 1}}$$

$$p_{face} = \frac{\rho_{face}\left(a_{face}\right)^2}{\gamma}$$
(6-10)

Note that when i2d = -1, the 1003 boundary condition is augmented by the far-field point-vortex correction<sup>38</sup> (for 2-d, x - z plane only).

# 6.1.5 Inviscid Surface

bctype 1005

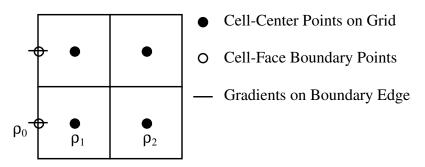


Figure 6-5. Inviscid surface boundary condition (bctype 1005) for density.

The inviscid surface boundary conditions are cell-face type boundary conditions. The cell-face boundary values for density are approximated with the density values of the near-est cell-center points on the grid. With the points defined as in as Figure 6-5,

$$\rho_0 = \rho_1$$
(6-11)

The boundary values for pressure are obtained in the same manner.

The velocity component normal to the surface is set to zero in the following manner. Assume that the surface in Figure 6-5 is a k = constant surface, for which the metrics (unit normals) are  $\hat{\zeta}_x$ ,  $\hat{\zeta}_y$ , and  $\hat{\zeta}_z$ . Then the normalized contravariant velocity (normal to the k direction) at cell center 1 is

$$\overline{W}_1 = u_1 \hat{\zeta}_x + v_1 \hat{\zeta}_y + w_1 \hat{\zeta}_z + \hat{\zeta}_t$$
(6-12)

The velocity components at the wall are then calculated using

$$u_0 = u_1 - \hat{\zeta}_x \overline{W}_1$$

$$v_0 = v_1 - \hat{\zeta}_y \overline{W}_1$$

$$w_0 = w_1 - \hat{\zeta}_z \overline{W}_1$$
(6-13)

Using  $(\hat{\zeta}_x)^2 + (\hat{\zeta}_y)^2 + (\hat{\zeta}_z)^2 = 1$ , it may be seen that  $\overline{W}_0 = 0$  so that the normal velocity at the wall is zero.

The gradient values for all five flow-field variables needed at the cell-face are obtained using two-point differences. For example, for density,

(6-14)

$$\nabla \rho = 2(\rho_1 - \rho_0)$$

# 6.1.6 Constant Enthalpy and Entropy Inflow

bctype 1008

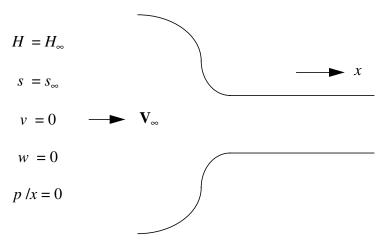


Figure 6-6. Constant enthalpy and entropy inflow boundary conditions (bctype 1008).

The constant enthalpy and entropy boundary condition (sometimes referred to as the wind tunnel inflow boundary condition) is a cell-center type condition. A grid set up to use the tunnel inflow condition should always have the x coordinate running along the length of the tunnel as shown in Figure 6-6. It is assumed that the entropy and the enthalpy are at the free-stream conditions. The v and w components of velocity are set to zero and the pressure gradient at the boundary is also zero. The density at the boundary is obtained from

$$s_{\infty} = \frac{p}{\rho^{\gamma}}$$
(6-15)

The u component of velocity is obtained from

$$H_{\infty} = \frac{a^2}{\gamma - 1} + \frac{u^2}{2}$$
(6-16)

The zero pressure gradient condition is used to set the pressure in the ghost cell equal to the nearest interior value.

For running internal (wind tunnel type) flows, **bctype** 2003 is usually preferred as the inflow boundary condition over 1008, although both may work. For the corresponding wind tunnel outflow boundary condition, **bctype** 2002 should generally be used.

6.1.7 Singular Axis Using Half-plane Symmetry

#### bctype 1011

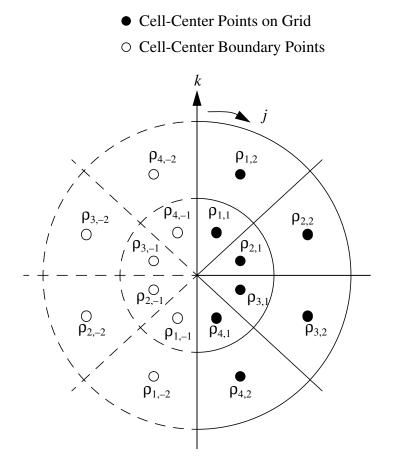


Figure 6-7. Singular axis with half-plane symmetry boundary condition (bctype 1011).

The singular axis using half-plane symmetry boundary conditions are cell-center type boundary conditions. As an example of this boundary condition, assume the singular axis occurs at k = 1. If **jdim** = 5, an i = constant plane might look like the one drawn in Figure 6-7. In the figure, the subscripts for  $\rho$  are j and then k (i.e.  $\rho_{j,k}$ ). The boundary values for density are assigned as

$$\begin{aligned}
\rho_{1,-1} &= \rho_{4,1} & \rho_{1,-2} &= \rho_{4,2} \\
\rho_{2,-1} &= \rho_{3,1} & \rho_{2,-2} &= \rho_{3,2} \\
\rho_{3,-1} &= \rho_{2,1} & \rho_{3,-2} &= \rho_{2,2} \\
\rho_{4,-1} &= \rho_{1,1} & \rho_{4,-2} &= \rho_{1,2}
\end{aligned}$$
(6-17)

The other four flow-field variables are assigned in a similar fashion; however, the normalized contravariant velocities and metrics are used to determine the correct signs for the velocity components in a manner similar to boundary condition type 1001. First note that by the assumption of half-plane symmetry, the direction cosines on j = 1 are the same as j = jdim apart from the sign; let these metrics be denoted  $\hat{\eta}_x$ ,  $\hat{\eta}_y$ , and  $\hat{\eta}_z$ . Then, for example,

$$\overline{V}_{4,1} = u_{4,1}\hat{\eta}_x + v_{4,1}\hat{\eta}_y + w_{4,1}\hat{\eta}_z + \hat{\eta}_t$$

$$u_{1,-1} = u_{4,1} - 2\hat{\eta}_x \overline{V}_{4,1}$$

$$v_{1,-1} = v_{4,1} - 2\hat{\eta}_y \overline{V}_{4,1}$$

$$w_{1,-1} = w_{4,1} - 2\hat{\eta}_z \overline{V}_{4,1}$$
(6-19)

# 6.1.8 Singular Axis Using Full Plane, Flux Specified

#### bctype 1012

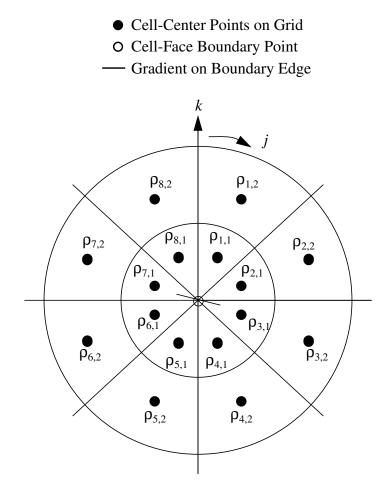


Figure 6-8. Singular axis using full plane boundary condition (bctype 1012).

The singular axis using full plane boundary conditions are cell-face type boundary conditions. As an example of this boundary condition, assume the singular axis occurs at k = 1. If **jdim** = 9, an i = constant plane might look like the one drawn in Figure 6-8. In the figure, the subscripts for  $\rho$  are j and then k (i.e.  $\rho_{i,k}$ ).

For the cell-face boundary point value of density ( $\rho_0$ ), an average value of all the values at k = 1 is obtained, i.e.

$$\rho_{j,0} = \frac{\sum_{j=1}^{j\dim -1} \rho_{j,1}}{j\dim -1}$$
(6-20)

The flux value is obtained with a two point extrapolation using  $\rho_1$  and  $\rho_0$ :

$$(\nabla \rho)_{i} = 2(\rho_{i,1} - \rho_{i,0})$$
 (6-21)

The boundary values for all five flow-field variables are assigned as described for density.

A known problem exists when using this boundary condition with the Baldwin-Lomax turbulence model. In such cases, the code employs the "wall" Baldwin-Lomax equation option rather than the "wake" Baldwin-Lomax equation option at the 1012 boundary.

#### 6.1.9 Singular Axis Using Extrapolation (Partial Plane)

#### bctype 1013

The singular axis using extrapolation boundary conditions are cell-center type boundary conditions. The ghost points are extrapolated from the computational domain. This boundary condition is used with singular axes for which neither boundary condition type 1011 or 1012 is appropriate (quarter planes, for instance). For example, the density boundary values would be approximated as

$$\begin{split} \rho_{-1} &= \rho_1 \\ \rho_{-2} &= \rho_1 \end{split} \tag{6-22}$$

The same first order extrapolation is used for the boundary values of the other flow-field variables.

#### 6.1.10 A Word About Singular Metrics

Version 5.0 will automatically detect collapsed grid lines (e.g. singular metrics, cell faces with zero area) on block boundaries. (Collapsed grid lines in the "interior" of a block

are not allowed.) The detection of singular metrics is no longer keyed to the specification of either **bctype** 1011, 1012, or 1013. While these boundary conditions are still applicable, the new singular metric detection allows any other standard boundary condition that does not rely on the grid metrics at the boundary to be used as well. Note that the following boundary conditions rely on the grid metrics at the boundary and so should not be used on a singular face/face segment: 1001 (symmetry), 1005 (inviscid), 1003 (inflow/outflow), 2003 (inflow with specified total conditions) and 2006 (radial equilibrium).

When CFL3D detects singular metrics on a particular block, a message is written to the main output file (unit 11) indicating the location. It is always a good idea to verify that any singular block faces/face segments are in fact correctly detected. Metrics are treated as singular if the total area of a face/face segment is less than a parameter atol (set near the top of subroutine metric, in module lbcx.f). In certain cases, atol may need to be changed (increased in magnitude if faces that are known to be non-singular are flagged as singular and decreased if faces that are known to be singular are not flagged as such, in which case the code will give floating point overflow in subroutine metric).

# 6.2 Physical Boundary Conditions with Auxiliary Data

For the "2000" series of boundary conditions, auxiliary data is required. Therefore,  $ndata \neq 0$ . The following sections describe the boundary condition types for constant input data (ndata > 0). For these conditions, the information in "LT14 - I0 Boundary Condition Specification" on page 32 through "LT19 - KDIM Boundary Condition Specification" on page 35 is immediately followed by a header line, then a single data line with the ndata values appropriate for the particular boundary condition. Section 6.2.8 describes how to use the same boundary condition types with variable input data.

Input values for **bctype** (i.e. boundary condition types currently supported) as follows:

bctype_	boundary condition
2002	specified pressure ratio
2003	inflow with specified total conditions
2004	no-slip wall
2005	periodic in space
2006	set pressure to satisfy the radial equilibrium equation
2007	set all primitive variables
2102	pressure ratio specified as a sinusoidal function of time

# 6.2.1 Specified Pressure Ratio

# bctype 2002

The specified pressure ratio boundary condition, generally used as the outflow boundary condition for internal flows, is a cell-center type boundary condition. A single pressure ratio,  $\tilde{p}/\tilde{p}_{\infty}$ , is specified on input. The parameter **ndata** must be set to 1 for the input pressure ratio value to be read. This pressure ratio is used to set both cell-center pressure boundary values ( $p_{-1}$  and  $p_{-2}$ ). Extrapolation from inside the computational domain is used to set the boundary values for  $\rho$ , u, v, and w. See "Extrapolation" on page 84. An example of the input lines is:

p/pinf 0.910

When using **bctype** 2002 as the outflow condition for internal (wind tunnel type) flows, the tunnel Mach number and/or mass flow should be monitored and **p/pinf** adjusted accordingly to obtain the correct conditions. This is usually an iterative process.

# 6.2.2 Inflow With Specified Total Conditions

# bctype 2003

The inflow with specified total conditions boundary condition (sometimes referred to as an "engine inflow" boundary condition because it is often used to specify the conditions at an inflow face where an engine exhaust is located) is a cell-center type boundary condition. The following five pieces of information are provided on input (**ndata** = 5): an estimated inflow Mach number ( $M_e$ ), the total pressure ratio ( $\tilde{p}_t/\tilde{p}_{\infty}$ ), the total temperature ratio ( $\tilde{T}_t/\tilde{T}_{\infty}$ ), and the flow directions ( $\alpha$  and  $\beta$ ) in degrees. These values are used as the external state in a 1-d characteristic boundary condition. An example of the input lines is:

 Mach
 Pt/Pinf
 Tt/Tinf
 Alphae
 Betae

 0.300
 4.000
 1.17555
 0.0
 0.0

Boundary condition type 2003 is very similar to boundary condition type 1003 in that either interior or exterior values are chosen, depending on the sign of the characteristics at the boundary. One difference between them is that, while 1003 uses far-field reference-state levels for the exterior values, 2003 uses the total temperature, pressure, Mach number, and flow angle that are input to determine the reference exterior values. Another difference is the way the density and pressure boundary condition values are determined. Boundary condition type 1003 calculates these values as shown in Equation (6-10). Boundary condition type 2003 first determines a Mach number at the boundary using the velocities and speed of sound at the boundary (which were calculated through the characteristic method):

$$M^{2} = \frac{u^{2} + v^{2} + w^{2}}{a^{2}}$$
(6-23)

It should be noted that the inflow Mach number may end up slightly different from the input  $M_e$  for a converged solution. The pressure and density boundary conditions are then determined with

$$p = \frac{p_t/p_{\infty}}{\gamma \left(1 + \frac{\gamma - 1}{2}M^2\right)^{\gamma/(\gamma - 1)}}$$

$$\rho = \frac{\gamma p}{a^2}$$
(6-24)

The boundary condition velocity components are determined the same way as for boundary condition type 1003.

## 6.2.3 Viscous Surface

#### bctype 2004

The viscous surface boundary conditions are cell-face type boundary conditions. The no-slip condition ( $\mathbf{V} = 0$ ) is applied at the surface. Two pieces of auxiliary information are supplied on input (**ndata** = 2): the wall temperature ( $\tilde{T}_w/\tilde{T}_\infty$ ) and the mass flow ( $C_q$ ), where  $C_q = (\rho u_{normal})/(\rho u)_\infty$ . ( $C_q$  is zero if there is no flow through the wall.) An example of the input lines is:

where

**Twtype** > 0sets  $\tilde{T}_w/\tilde{T}_\infty =$ **TwtypeTwtype** = 0sets adiabatic wall**Twtype** < 0</th>sets  $\tilde{T}_w$  at the stagnation temperature**Cq** < 0</th>results in mass flow in the direction of "decreasing" computational coordinate (i.e.  $C_q < 0$  results in suction on minimum wall, blowing on maximum wall)**Cq** > 0results in mass flow in the direction of "increasing" computational coordinate (i.e.  $C_q > 0$  results in blowing on minimum wall, suction on maximum wall)**Cq** = 0results in no flow through the wall (same as old **bctype** 1004)

Note that for a dynamic grid, **bctype** 2004 gives no-slip relative to the moving wall. To set the wall velocity to zero relative to a non-moving reference frame when the mesh is moving, use –2004 instead of 2004; this option makes sense only if the grid motion is tangential to the surface.

Also note that boundary condition type 2004 supersedes boundary condition types 1004 *and* 2004 in previous versions of CFL3D. The main reason for the replacement is that 1004 relied on global parameters **isnd** and **c2spe** to determine whether adiabatic wall or constant-temperature wall conditions were invoked. As a consequence, all 1004 segments had to be the same. Now, with 2004 (along with its additional data field **Twtype**), every no-slip wall segment can be set with different wall temperature conditions, if desired. Boundary condition type 2004 also allows for mass flow through the wall (suction or blowing) through the second additional data field **Cq**.

The no-slip wall boundary condition is implemented as follows:

• The nondimensional pressure on the body  $p_b$  is determined through linear extrapolation:

$$p_b = p_1 - (p_2 - p_1)/2 \tag{6-25}$$

But if  $p_b < 0$ , then  $p_b = p_1$ . Here the indices 1 and 2 indicate the first and second cellcenter values away from the wall, respectively.

• The nondimensional square of the speed of sound, variable c2, at the wall is next determined. (Note that c2 =  $(\tilde{a}_w/\tilde{a}_\infty)^2 = \tilde{T}_w/\tilde{T}_\infty$ .)

If **Twtype** > 0, c2 =**Twtype**.

If **Twtype** < 0, c2 = 
$$1 + \frac{\gamma - 1}{2} (M_{\infty})^2$$
.

If **Twtype** = 0, c2 = 
$$\left(\frac{a_1}{a_{\infty}}\right)^2 \left[1 + \frac{\gamma - 1}{2}(M_1)^2\right]$$
.

• The surface velocities are then determined as

$$u_{w} = M_{\infty}C_{q} \frac{\xi_{x}}{|\nabla\xi|} \frac{(c2)}{\gamma p_{b}} + u_{mesh}$$

$$v_{w} = M_{\infty}C_{q} \frac{\xi_{y}}{|\nabla\xi|} \frac{(c2)}{\gamma p_{b}} + v_{mesh}$$

$$w_{w} = M_{\infty}C_{q} \frac{\xi_{z}}{|\nabla\xi|} \frac{(c2)}{\gamma p_{b}} + w_{mesh}$$
(6-26)

where  $u_{mesh}$ ,  $v_{mesh}$ , and  $w_{mesh}$  are the velocity components of the mesh (they equal 0 if the grid/body is not in motion).

• Since 2004 is a cell-face boundary condition type (i.e. the boundary conditions and their gradient are applied at the cell face rather than in ghost cells),

$$\rho_{\rm w} = \frac{\gamma p_b}{c^2}$$

$$p_{\rm w} = p_b$$
(6-27)

The gradients at the wall for all the primitive variables are determined via  $\nabla \rho = 2\rho_1 - 2\rho_w$ , etc.

Note that smin is computed from viscous walls *only*. If a wall is inviscid, then as far as smin is concerned, it is invisible. This is important to remember when viscous boundary conditions are turned on after running a case inviscidly for some time since smin may never have been computed!

# 6.2.4 Periodic (In Space)

#### bctype 2005

The periodic boundary conditions are cell-center type boundary conditions. Four pieces of additional information must be input (**ndata** = 4). The number of the grid with which the current grid is periodic and the rotation angle  $(d\theta_x, d\theta_y, d\theta_z)$  about one of the coordinate axes (x, y, z) are needed. Only *one* of the three angles can be used at a time and the other two angles *must* be identically zero. The angles should be determined from the right-hand rule. For example, if rotation is desired about the N axis (where N is either x, y or z, point the right-hand thumb in the direction of the +N axis. The fingers will curl in the direction of the positive angle. When setting the angle for a particular face (i.e. the i = 1 face), set the angle of rotation equal to the angle that the *periodic face* would have to move through to get to this face.

For a sample input, assume the current face on which the boundary condition is being set is j = 1. The the following input will cause the current face to be periodic with the **jdim** face in grid 2, where a rotation of +45 degrees about the x axis would map the **jdim** face of grid 2 onto the j = 1 face of the current grid:

ngridp dthx dthy dthz 2 45.0 0.0 0.0

At present, it is assumed that the current block and the block with which it is periodic match 1-1 at their corresponding faces after the rotation. Note that there is *not* a check for this! Also, the two blocks are assumed to be aligned similarly. That is i, j, and k on the first block must be defined *exactly* the same on the second block. For example, if the

k = kdim face of the current block has the periodic boundary condition applied to it, it is implicitly assumed that it is periodic with the k = 1 face of the specified second block *and i* and *j* run in the exact same directions on both blocks. Note that this means that two of the dimensions (**idim** and **jdim** in the example) must be the same on both blocks.

The periodic boundary condition also works for a grid with one cell (two grid planes) in the periodic direction if the grid is set to be periodic with itself. Note, however, that if a particular block is periodic with a *different* block, then neither block should be only one cell wide in the periodic direction.

In addition, the periodic boundary condition may be used for linear displacement, provided the rotation angles are set to zero. Alternatively, the 1-1 block connection can be used (linear displacement only!). The 1-1 internal check will flag geometric mismatches, which should be equal to the desired linear periodic displacement. This is a good check of the input which is not available with boundary condition type 2005.

Boundary condition type 2005 is a limited-use periodic (in space) boundary condition. For example, if one is solving for flow through a duct and it is known that the solution is periodic over 90 degrees (i.e. the solution is identical in each of the four quadrants of the duct), then a solution need only be obtained on a grid spanning 90 degrees, with periodic boundary conditions applied at the two edges.

For an illustration of how this boundary condition works, consider Figure 6-9. This example is for flow through a 90 degree wedge (the flow direction is in the third dimension, out of the page). The flow is periodic over 90 degrees, so periodic boundary conditions are applied at the j = 0 and j = jdim faces.

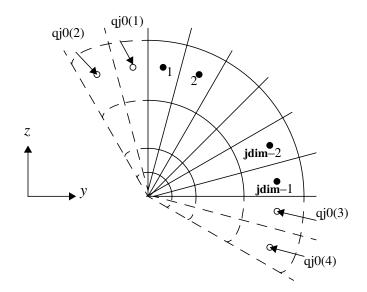


Figure 6-9. Periodic boundary condition (bctype 2005) example.

For density and pressure, boundary condition type 2005 sets:

$$qj0(1) = q(jdim-1)$$
  
 $qj0(2) = q(jdim-2)$   
 $qj0(3) = q(1)$   
 $qj0(4) = q(2)$   
(6-28)

The boundary conditions for the velocity components u, v, and w are assigned similarly, except that they are first rotated through the periodic angle, in this case 90°. For example,

$$u_{1,BC} = qj0(k, i, 2, 1) = q(j, k, i, 2) \quad (\text{unchanged})$$

$$v_{1,BC} = qj0(k, i, 3, 1) = q(j, k, i, 3)\cos\theta_{p} - q(j, k, i, 4)\sin\theta_{p} \quad (6-29)$$

$$w_{1,BC} = qj0(k, i, 4, 1) = q(j, k, i, 3)\sin\theta_{p} + q(j, k, i, 4)\cos\theta_{p}$$

For an example of how this boundary condition can be used for linear displacement, see the 2-d vibrating plate sample test case in Section 9.1.6 on page 177.

## 6.2.5 Radial Pressure Equilibrium

#### bctype 2006

Boundary condition type 2006 is a cell-center type boundary condition. Radial pressure equilibrium is used as a downstream boundary condition when it is desired to specify a pressure, but a large swirling component is present in the flow. Typically, this boundary condition would be used in turbomachinery applications in which a swirling flow is established by a rotor that is not corrected by the presence of stators or exit guide vanes.

Four additional pieces of information are needed for this boundary condition type (**ndata** = 4). The grid number (**ngridc**) of the grid *from* which the integration of pressure is to be continued is specified. If the integration in this grid is not continued from another grid, input 0. A value for  $\tilde{p}/\tilde{p}_{\infty}$  is input. If **ngridc** = 0, this value will be the starting value for the integration. If **ngridc**  $\neq$  0, then the pressure value from the connecting point in grid **ngridc** is used as the starting value for the integration in the current block. The direction in which integration is to proceed is specified with the parameter **intdir**. It may have an absolute value of 1, 2, or 3 for integration in the *i*, *j*, or *k* directions, respectively. The sign of **intdir** indicates whether the integration proceeds in the increasing (positive) or decreasing (negative) coordinate direction. The **intdir** direction *must* be the radial direction. The fourth piece of information needed is the physical direction along which the radial axis lies, denoted with the *x*, *y*, or *z* axes. The input lines for this boundary condition type would look like:

ngridc P/Pinf intdir axcoord

0 0.9 +3 1

The radial equilibrium condition requires that the pressure satisfy

$$\frac{dp}{dr} = \rho \frac{\left(V_{\theta}\right)^2}{r}$$
(6-30)

where  $V_{\theta}$  is the circumferential velocity component and *r* is the radius. The pressure ratio,  $\tilde{p}/\tilde{p}_{\infty}$ , is set at either the bottom or the top of the block face and then the radial equilibrium equation is integrated in either the increasing or decreasing radial direction to give the pressure at all other radii. The trapezoidal rule is used to perform the integration. The other flow-field variables ( $\rho$ , *u*, *v*, *w*) are extrapolated from inside the computational domain. See "Extrapolation" on page 84.

This boundary condition assumes that one coordinate direction on the block face is essentially radial and the other is essentially circumferential and that the integration is being carried out in the radial direction. Since there is no way to verify this in the code, care *must* be exercised when using this boundary condition to insure that this and the following restrictions are met. Continuation of the radial equilibrium condition through block boundaries is restricted to blocks that have the *same* orientation. For example, if the equilibrium condition is to be continued through a *k* boundary from an adjacent block, then *i* and *j* must run in the same direction in both blocks. This also implies that the *k* boundary must be k = 1 in one block and k = kdim in the second (i.e. the boundary can not be *k* = 1 in both blocks). Also, the segment must run the entire length of the block face in the direction in which the integration is being carried out. For example, if the integration is being carried out in the *k* direction, then **ksta** must be set to 1 and **kend** must be set to **kdim**. This restriction applies only if the integration is being continued from another block.

Consider a case in which boundary condition 2006 is to be applied at an i = idim face, with j = constant radial lines (i.e. lines of constant angle  $\theta$ ) and k = constant circumferential lines (i.e. lines of constant radius). Assume the block face lies in an x = constant plane, as shown in Figure 6-10.

Assuming integration in the +k direction (intdir = 3), the pressure (1 = 5) in the first plane of ghost cells is obtained from

$$qi0(j,1,5,3) = \left(\frac{\tilde{p}}{\tilde{p}_{\infty}}\right) p_{\infty}$$

$$qi0(j,k,5,3) = p_{k,j} = p_{k-1,j} + \frac{\rho_{k,j}[(V_{\theta})^2]_{k,j}}{\bar{r}_{k,j}} \Delta r_{k,j} \qquad k = 2, \text{ kdim} - 1$$
(6-31)

where

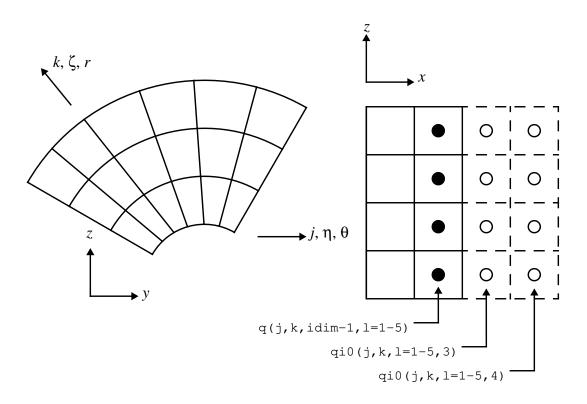
$$[(V_{\theta})^{2}]_{k, j} = \mathbf{V}_{k, j} \cdot \overline{\nabla \eta_{j}}$$

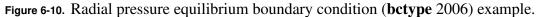
$$\overline{\nabla \eta_{j}} = \frac{1}{2} [\nabla \eta_{k, j+1/2} + \nabla \eta_{k, j-1/2}]_{\mathbf{idim}-1}$$
(6-32)

 $\mathbf{V}_{k, j}$  is the velocity at the cell center  $(j, k, \mathbf{idim} - 1)$  and  $\bar{r}_{k, j}$  is the average radius in cell (k, j):

$$[r_{k-1/2, i} + r_{k+1/2, i}]/2$$
(6-33)

(k, j) denotes a cell-center value.  $(k \pm 1/2, j \pm 1/2)$  denotes a face-center value.





The velocity and density in the first layer of ghost cells is determined by zeroth order extrapolation from the adjacent interior points. For the case shown in Figure 6-10, the density (1 = 1) is obtained from

$$qi0(j,k,1,3) = q(j,k,idim-1,1)$$
 (6-34)

All flow quantities in the second layer of ghost cells are obtained by zeroth order extrapolation from the first layer of ghost cells:

$$qi0(j,k,l=1-5,4) = qi0(j,k,l=1-5,3)$$
 (6-35)

6.2.6 Specify All Primitive Variables

bctype 2007

Boundary condition type 2007 is a cell-center type boundary condition. It involves setting the boundary conditions with the five (**ndata** = 5) primitive variables, using standard CFL3D normalizations:  $\tilde{\rho}/\tilde{\rho}_{\infty}$ ,  $\tilde{u}/\tilde{a}_{\infty}$ ,  $\tilde{v}/\tilde{a}_{\infty}$ ,  $\tilde{w}/\tilde{a}_{\infty}$ , and  $\tilde{p}/[\tilde{\rho}_{\infty}(\tilde{a}_{\infty})^2]$ . An example of the input lines is:

rho/rho\_inf u/a\_inf v/a\_inf w/a\_inf p/(rho\_inf\*a\_inf\*\*2) 1.0 0.3 0.0 0.0 0.0

Note that the input pressure is *not*  $\tilde{p}/\tilde{p}_{\infty}$  but rather  $\tilde{p}/(\gamma \tilde{p}_{\infty})$ . Also note that the turbulence quantities are *not* currently allowed to be specified in the same way as the **q**'s. Instead, they are extrapolated from the interior when **bctype** 2007 is employed.

## 6.2.7 Specified Pressure Ratio As Sinusoidal Time Function

#### bctype 2102

The specified pressure ratio as a sinusoidal function of time boundary conditions are cell-center type boundary conditions. The pressure ratio,  $\tilde{p}/\tilde{p}_{\infty}$ , is specified as a sinusoidal function of time. The other flow-field variables ( $\rho$ , u, v, w) are extrapolated from inside the computational domain. Three pieces of additional information are specified on input (**ndata** = 4): the desired baseline (steady) pressure ratio ( $\tilde{p}/\tilde{p}_{\infty}$ ), the amplitude of the pressure oscillation ( $\Delta \tilde{p}/\tilde{p}_{\infty}$ ), the reduced frequency of the pressure oscillation ( $k_r$ ), and the "grid equivalent" of the dimensional reference length used to define the reduced frequency ( $L_{ref}$ ). The pressure will vary as

$$\gamma p = \frac{\tilde{p}}{\tilde{p}_{\infty}} + \frac{\Delta \tilde{p}}{\tilde{p}_{\infty}} \sin(2\pi k_r t)$$
(6-36)

The reduced frequency is non-dimensionalized by

$$k_r = \frac{\tilde{f}\tilde{L}}{\tilde{a}_{\infty}}$$
(6-37)

where  $\tilde{f}$  is the frequency in cycles per second,  $\tilde{L}$  is a characteristic length and  $\tilde{a}_{\infty}$  is the free-stream speed of sound. (Note that this definition of reduced frequency differs from

the "standard" definition  $k_r = \tilde{f}\tilde{L}/|\tilde{\mathbf{V}}|_{\infty}$ .) An example of the lines in the input file for this boundary condition is:

p/pinf	deltap/pinf	rfreqp	lref
0.910	0.001	175.93	1.0

## 6.2.8 Variable Data for the 2000 Series

To use this option, set ndata = -(the ndata used for the constant data value(s) as described above). Then, instead of the line containing the constant data value(s), substitute a line with the name (up to 60 characters) of a formatted file that has the appropriate array of data values. The data file will then be read with the following format:

The roles of m, n vary depending on which face the segment is located:

m, $n \rightarrow j, k$	on the $i = 1$ and the $i = idim$ faces where
$m_{J} m \rightarrow J, \kappa$	mdim = jend-jsta and ndim = kend-ksta
m, $n \rightarrow k, i$	on the $j = 1$ and the $j = jdim$ faces where
$m_{\ell} n \rightarrow \kappa, \ell$	mdim = kend-ksta and ndim = iend-ista
m n \ii	on the $k = 1$ and the $k = kdim$ faces where
m, n $\rightarrow j, i$	<pre>mdim = jend-jsta and ndim = iend-ista</pre>

Note that zeroes are *not* acceptable for mdim or ndim. Use the actual values of **jdim**-1, **kdim**-1, and/or **idim**-1 for the full face. Only boundary condition type 2007 can make use of two planes of data. For all other boundary conditions, set np = 1.

# 6.3 Block Interface Boundary Conditions

For all the types of block interface boundary conditions, set **bctype** = 0. If **bctype**  $\neq$  0, but block interface boundary conditions are set, the block interface boundary conditions will supersede. All block interface boundary conditions (one-to-one, patched, and overset) are cell-center type boundary conditions.

# 6.3.1 One-to-One Blocking

## (Input Line Types Twenty-Four through Twenty-Six)

One-to-one (1-1) blocking, sometimes called  $C^0$  continuous, means that the faces shared by two grids are exactly (to machine zero) the same. Several examples showing a variety of blocking strategies are discussed below. In the examples, the variable, **l**, represents the five flow-field variables. The sample inputs illustrate how the index ranges (**ista** to **iend**, **jsta** to **jend**, **ksta** to **kend**) in "LT25 - 1-1 Blocking Connections" on page 37 are assigned so that the correct communication between blocks is established. A good initial check to determine if the one-to-one blocking input is set up correctly is to compare the quantity of points in the range. For example, if two grids share a common portion of a j = constant face, the following must be true:

$$[\mathbf{jend} - \mathbf{jsta} + 1]_{\text{grid } 1} = [\mathbf{jend} - \mathbf{jsta} + 1]_{\text{grid } 2}$$
(6-38)

Keep in mind that j = constant faces can communicate with i = constant and/or k = constant faces as well (and vice versa), in which case the check will be

$$[\mathbf{jend} - \mathbf{jsta} + 1]_{\text{grid } 1} = [\mathbf{iend} - \mathbf{ista} + 1]_{\text{grid } 2}$$
(6-39)

or

$$[\mathbf{jend} - \mathbf{jsta} + 1]_{\text{grid } 1} = [\mathbf{kend} - \mathbf{ksta} + 1]_{\text{grid } 2}$$
(6-40)

respectively. Note that Equation (6-38) through Equation (6-40) are necessary for a oneto-one interface to be specified correctly, but they are not sufficient; the direction in which the indices are input must be correct as well. (See Example 3 on page 106.)

**Example 1** 

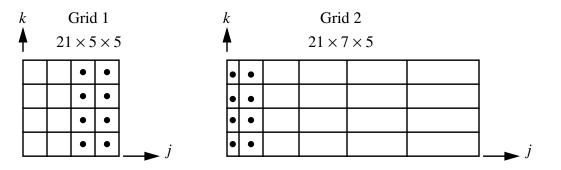


Figure 6-11. One-to-one blocking example 1.

As a simple illustration of 1-1 blocking, consider Figure 6-11. The figure shows an i = constant (not necessarily the *same* constant) face of two grids. Suppose communica-

Line	<u>Type</u>							
6	NGRID	NPLOT3D	NPRINT	NWREST	ICHK	I2D	NTSTEP	ITA
	2	1	0	100	0	0	1	1
8	IDIM	JDIM	KDIM					
Ũ	21	5	5					
	21	7	5					
13	GRID	NBCI0	NBCIDIM	NBCJ0	NBCJDIM	NBCK0	NBCKDIM	IOVRLP
10	1	1	1	1	1	1	1	0
	2	1	1	1	1	1	1	0
16	J0: GRID	SEGMENT	BCTYPE	ISTA	IEND	KSTA	KEND	NDATA
	1	1	1003	1	21	1	5	0
	2	1	0	1	21	1	5	0
17	JDIM: GRID	SEGMENT	BCTYPE	ISTA	IEND	KSTA	KEND	NDATA
	1	1 1	0	1	21	1	5 5	0
	Z	1	1003	1	21	1	5	0
	1-1 BI	OCKING DA	TA:					
24	NBLI							
~ -	1							
25	NUMBER			JSTA KS			KEND ISVA1	ISVA2
26	1	1	1	5	1 21	5	5 1	3
26	NUMBER		ISTA	JSTA KS			KEND ISVA1	ISVA2
	1	2	1	1	1 21	1	5 1	3

tion is desired between the j = jdim face of grid 1 and the j = 1 face of grid 2. The pertinent input would look something like:

Note that in the sample input, both grids have the same value for **idim**. This does not *have* to be true. One grid can share only a portion of a face with another grid.

The boundary conditions at the j = jdim face on any *i* plane, denoted i1, of grid 1 would be set as:

<u>Grid 1</u>	<u>Grid 2</u>
qj0(1,i1, <b>l</b> ,3) =	q(1,1,i2, <b>l</b> )
qj0(2,i1, <b>l</b> ,3) =	q(1,2,i2, <b>l</b> )
qj0(3,i1, <b>l</b> ,3) =	q(1,3,i2, <b>l</b> )
qj0(4,i1, <b>l</b> ,3) =	q(1,4,i2, <b>l</b> )
qj0(1,i1, <b>l</b> ,4) =	q(2,1,i2, <b>l</b> )
qj0(2,i1, <b>l</b> ,4) =	q(2,2,i2, <b>l</b> )
qj0(3,i1, <b>l</b> ,4) =	q(2,3,i2, <b>l</b> )
qj0(4,i1, <b>l</b> ,4) =	q(2,4,i2,l)

The boundary conditions at the j = 1 face on any *i* plane, denoted i2, of grid 2 would be set as:

<u>Grid 2</u>	<u>Grid 1</u>
qj0(1,i2, <b>l</b> ,1) =	q(4,1,i1, <b>l</b> )
qj0(2,i2, <b>l</b> ,1) =	q(4,2,i1, <b>l</b> )
qj0(3,i2, <b>l</b> ,1) =	q(4,3,i1, <b>l</b> )
qj0(4,i2, <b>l</b> ,1) =	q(4,4,i1, <b>l</b> )
gj0(1,i2, <b>l</b> ,2) =	q(3,1,i1, <b>l</b> )
qj0(2,i2, <b>l</b> ,2) =	q(3,2,i1, <b>l</b> )
gj0(3,i2, <b>l</b> ,2)=	q(3,3,i1, <b>l</b> )
qj0(4,i2, <b>l</b> ,2) =	q(3,4,i1, <b>l</b> )

#### Example 2

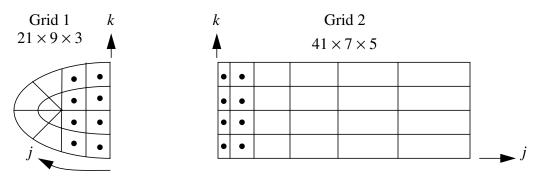


Figure 6-12. One-to-one blocking example 2.

A slightly more complicated example of 1-1 blocking is shown in Figure 6-12. Again, the figure shows an i = constant face (not necessarily the *same* constant) of two grids. In grid 1, j is now a circumferential direction. Communication is desired between the j = 1 and the  $j = \mathbf{jdim}$  faces of grid 1 and the j = 1 face of grid 2. (In this example, it is assumed that the k = 1 boundary of grid 1 is a solid wall flat plate, so there is no 1-1 connectivity there.) The pertinent input would look something like:

<u>Type</u>							
NGRID	NPLOT3D	NPRINT	NWREST	ICHK	I2D	NTSTEP	ITA
2	1	0	100	0	0	1	1
IDIM	JDIM	KDIM					
21	9	3					
41	7	5					
CDID	NDCTO	NDCIDIM	NDC TO	NDCIDIM	NDCKO	NDCVDTM	
GRID 1	NBCIU 1	NBCIDIM 1	NBCJU 1	NBCJDIM 1	NBCRU 1	NBCADIM 1	IOVRLP
1	1	1	1	1	1	1	0
2	T	Ţ	Z	T	T	T	0
J0: GRID	SEGMENT	BCTYPE	ISTA	IEND	KSTA	KEND	NDATA
1	1	0	1	21	1	3	0
2	1	0	1	21	1	5	0
2	2	1001	21	41	1	5	0
JDIM: GRID	SEGMENT	BCTYPE	ISTA	IEND	KSTA	KEND	NDATA
	NGRID 2 IDIM 21 41 GRID 1 2 J0: GRID 1 2 2	NGRID         NPLOT3D           2         1           IDIM         JDIM           21         9           41         7           GRID         NBCI0           1         1           2         1           J0:         GRID           SEGMENT         1           2         1           2         2	NGRID         NPLOT3D         NPRINT           2         1         0           IDIM         JDIM         KDIM           21         9         3           41         7         5           GRID         NBCI0         NBCIDIM           1         1         1           2         1         1           30:         GRID         SEGMENT           J0:         GRID         SEGMENT           2         1         0           2         1         0           2         1         0           2         1         0	NGRID         NPLOT3D         NPRINT         NWREST           2         1         0         100           IDIM         JDIM         KDIM         100           21         9         3         3           41         7         5         5           GRID         NBCI0         NBCIDIM         NBCJ0           1         1         1         1           2         1         1         2           J0:         GRID         SEGMENT         BCTYPE         ISTA           1         1         0         1         1           2         1         0         1         2           J0:         GRID         SEGMENT         BCTYPE         ISTA           1         1         0         1         1           2         1         0         1         2           2         2         1001         21         2	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	NGRID         NPLOT3D         NPRINT         NWREST         ICHK         I2D         NTSTEP           2         1         0         100         0         0         1           IDIM         JDIM         KDIM

	1 2	1 1	0 1003		1 1	21 41		1 1	3 5	(	) )
24	1-1 BLOC NBLI 2	KING DAT.	A:								
25	NUMBER GR 1 2	ID : 1 1	ISTA 1 1	JSTA 1 9	KSTA 1 1	IEND 21 21	JEND 1 9	KEND 3 3	ISVA1 1 1	ISVA2 3 3	
26	NUMBER GR 1 2	ID : 2 2	ISTA 1 1	JSTA 1 1	KSTA 3 3	IEND 21 21	JEND 1 1	KEND 1 5	1 1 1	ISVA2 3 3	

Note how k ranges on grid 2 to coincide with the appropriate k points on the j = 1 and j = jdim faces of grid 1. This sample input also shows how grid 1 can share only a portion of the j = 1 face of grid 2 in the *i* direction.

The boundary conditions at the j = 1 face on any i plane, denoted i1, of grid 1 would be set as:

<u>Grid 1</u>	<u>Grid 2</u>
qj0(1,i1, <b>l</b> ,1) =	q(1,2,i2, <b>l</b> )
qj0(2,i1, <b>l</b> ,1) =	q(1,1,i2, <b>l</b> )
qj0(1,i1, <b>l</b> ,2) =	q(2,2,i2, <b>l</b> )
qj0(2,i1, <b>l</b> ,2) =	q(2,1,i2, <b>l</b> )

The boundary conditions at the j = jdim face for the i1 plane of grid 1 would be set as:

<u>Grid 1</u>	Grid 2
qj0(1,i1, <b>l</b> ,3) =	q(1,3,i2, <b>l</b> )
qj0(2,i1, <b>l</b> ,3) =	q(1,4,i2, <b>l</b> )
qj0(1,i1, <b>l</b> ,4) =	q(2,3,i2, <b>l</b> )
qj0(2,i1, <b>l</b> ,4) =	q(2,4,i2, <b>l</b> )

The boundary conditions at the j = 1 face on any i plane, denoted i2, of grid 2 would be set as:

<u>Grid 2</u>	<u>Grid 1</u>
qj0(1,i2, <b>l</b> ,1) =	q(1,2,i1, <b>l</b> )
qj0(2,i2, <b>l</b> ,1) =	q(1,1,i1, <b>l</b> )
qj0(3,i2, <b>l</b> ,1) =	q(8,1,i1, <b>l</b> )
qj0(4,i2, <b>l</b> ,1) =	q(8,2,i1, <b>l</b> )
qj0(1,i2, <b>l</b> ,2) =	q(2,2,i1, <b>l</b> )
qj0(2,i2, <b>l</b> ,2) =	q(2,1,i1, <b>l</b> )
qj0(3,i2, <b>l</b> ,2) =	q(7,1,i1, <b>l</b> )
qj0(4,i2, <b>l</b> ,2) =	q(7,2,i1, <b>l</b> )

6.3.1 One-to-One Blocking

#### Example 3

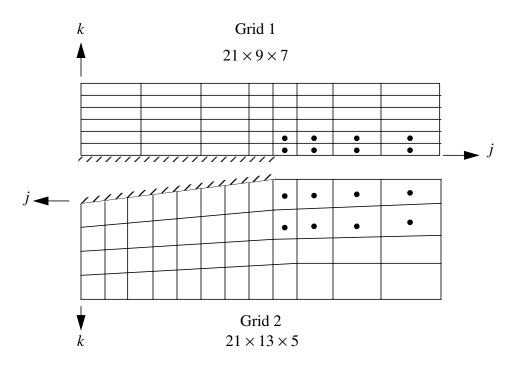


Figure 6-13. One-to-one blocking example 3.

The previous two examples show 1-1 blocking over the entire k range of two grids. In some cases, only segments of faces will utilize 1-1 communication between grids. In Figure 6-13, two grids share a portion of a face beyond boundaries defined as inviscid surfaces. Again, the figure shows an i = constant face (not necessarily the *same* constant) of two grids. Communication is desired between **jsta** = 5 and **jend** = 9 on the k = 1 face of grid 1 and **jsta** = 1 and **jend** = 5 on the k = 1 face of grid 2. The pertinent input would look something like:

Line '	<u>Type</u>									
6	NGRID	NPLOT3D	NPRINT	NWREST	ICHK	I2D	NTSTEP	ITA		
	2	1	0	100	0	0	1	1		
8	IDIM	JDIM	KDIM							
	21	9	7							
	21	13	5							
13	GRID	NBCI0	NBCIDIM	NBCJ0	NBCJDIM	NBCK0	NBCKDIM	IOVRLP		
15	GRID	NDCIU 1	NDCIDIM 1	NDC00	NDCODIM 1	NDCR0	NDCKD1M 1	TOAKTE		
	1	1	1	1	1	2	1	0		
	2	T	T	T	Ţ	Z	T	0		
18	K0: GRID	SEGMENT	BCTYPE	ISTA	IEND	JSTA	JEND	NDATA		
	1	1	1005	1	21	1	5	0		
	1	2	0	1	21	5	9	0		
	2	1	0	1	21	1	5	0		
	2	2	1005	1	21	5	13	0		
	1-1 BLOCKING DATA:									

24	NBLI									
	1									
25	NUMBER GRID	:	ISTA	JSTA	KSTA	IEND	JEND	KEND	ISVA1	ISVA2
	1 1		1	5	1	21	9	1	1	2
26	NUMBER GRID	:	ISTA	JSTA	KSTA	IEND	JEND	KEND	ISVA1	ISVA2
	1 2		1	5	1	21	1	1	1	2

Note that, since the j index runs in opposite directions on the two grids, one range is increasing and one range is decreasing in the 1-1 blocking input. It does not matter which grid's range increases and which one decreases as long as they map the points in the correct order.

The boundary conditions at the k = 1 face from **jsta** to **jend**-1 on any *i* plane, denoted i1, of grid 1 would be set as:

<u>Grid 1</u>	<u>Grid 2</u>
qk0(5,i1,l,1) =	q(4,1,i2, <b>l</b> )
qk0(6,i1,l,1) =	q(3,1,i2, <b>l</b> )
qk0(7,i1,l,1) =	q(2,1,i2, <b>l</b> )
qk0(8,i1,l,1) =	q(1,1,i2, <b>l</b> )
qk0(5,i1,l,2) =	q(4,2,i2, <b>l</b> )
qk0(6,i1,l,2) =	q(3,2,i2, <b>l</b> )
qk0(7,i1,l,2) =	q(2,2,i2, <b>l</b> )
qk0(8,i1,l,2) =	q(1,2,i2, <b>l</b> )

The boundary conditions at the k = 1 face from **jsta** to **jend**-1 on any *i* plane, denoted i2, of grid 2 would be set as:

Grid 2	<u>Grid 1</u>
qk0(1,i2,l,1) =	q(8,1,i1, <b>l</b> )
qk0(2,i2,l,1) =	q(7,1,i1, <b>l</b> )
qk0(3,i2,l,1) =	q(6,1,i1, <b>l</b> )
qk0(4,i2,l,1)=	q(5,1,i1, <b>l</b> )
qk0(1,i2,l,2) =	q(8,2,i1, <b>l</b> )
qk0(2,i2,l,2)=	q(7,2,i1, <b>l</b> )
qk0(3,i2,l,2) =	q(6,2,i1, <b>l</b> )
qk0(3,i2,l,2) =	q(5,2,i1, <b>l</b> )

#### **Example 4**

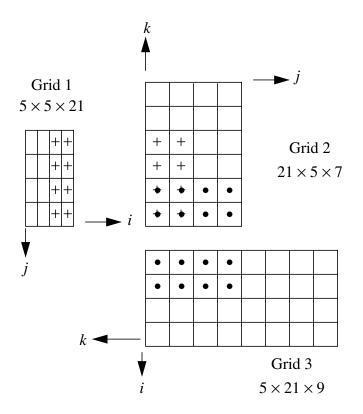


Figure 6-14. One-to-one blocking example 4.

The last example of 1-1 blocking involves three grids shown in Figure 6-14. In the figure, a k = constant (denoted k1) plane of grid 1, an i = constant (denoted i2) plane of grid 2, and a j = constant (denoted j3) plane of grid 3 are shown. Communication between grids 1 and 2 is desired between the i = idim face of grid 1 and the j = 1 face of grid 2. Communication between grids 2 and 3 is desired between the k = 1 face of grid 2 and the i = 1 face of grid 3. The pertinent input would look something like:

Line T	<u>[ype</u>							
6	NGRID	NPLOT3D	NPRINT	NWREST	ICHK	I2D	NTSTEP	ITA
	3	1	0	100	0	0	1	1
8	IDIM	JDIM	KDIM					
0	5	5	21					
	21	5						
	5	21	9					
13	GRID	NBCI0	NBCIDIM	NBCJ0	NBCJDIM	NBCK0	NBCKDIM	IOVRLP
-	1	1	1	1	1	1	1	0
	2	1	1	2	1	1	1	0
	3	2	1	1	1	1	1	0
14	IO: GRID	SEGMENT	BCTYPE	JSTA	JEND	KSTA	KEND	NDATA
	1	1	1003	1	5	1	21	0
	2	1	1003	1	5	1	7	0
	3	1	1005	1	21	1	5	0

	3	2	0		1	21		5	9	0
1511	DIM: GRIDSE	GMENTBCTYPE	e jsta	J	END	KSTA	KEN	ID N	IDATA	
	1 1	0	1		5	1	2	1	0	
	2 1	1001	1		5	1		7	0	
	3 1	1000	1		21	1		9	0	
16	J0: GRID	SEGMENT	BCTYPE	I	STA	IEND	KSI	'A	KEND	NDATA
	1				1	5		1	21	0
	2	1	0		1	21		1	5	0
	2	2	1000		1			5	7	0
	3	1	1003		1	5		1	9	0
10										
18	K0: GRID	SEGMENT	BCTYPE	I	STA	IEND	JST	A	JEND	NDATA
	1	1			1	5		1	5	0
	2	1			1	21		1	5	0
	3	1	1003		1	5		1	21	0
	1_1 1	LOCKING DAT	· 7. •							
24		DOCKING DAI	<b>A</b> •							
24	NBLI 2									
25										
25		GRID :								
	1	_	5	1			5		2	3
24	2	2	1	1		21		1	1	2
26	NUMBER	•••••	ISTA		KSTA				ISVA1	
	1		21	1	5	1	1	1	3	1
	2	3	1	1	9	1	21	5	2	3

Notice that the i index range for grid 2 runs in the opposite direction of the k index range for grid 3 due to the right-hand rule.

The boundary conditions at the i = idim face of grid 1 would be set as:

<u>Grid 1</u>	Grid 2
qi0(1,k1, <b>l</b> ,3) =	q(1,4,i2, <b>l</b> )
qi0(2,k1, <b>l</b> ,3) =	q(1,3,i2, <b>l</b> )
qi0(3,k1, <b>l</b> ,3) =	q(1,2,i2, <b>l</b> )
qi0(4,k1, <b>l</b> ,3) =	q(1,1,i2, <b>l</b> )
qi0(1,k1, <b>l</b> ,4) =	q(2,4,i2, <b>l</b> )
qi0(2,k1, <b>l</b> ,4) =	q(2,3,i2, <b>l</b> )
qi0(3,k1, <b>l</b> ,4) =	q(2,2,i2, <b>l</b> )
qi0(4,k1, <b>l</b> ,4) =	q(2,1,i2, <b>l</b> )

The boundary conditions from **ksta** to **kend**-1 on the j = 1 face of grid 2 would be set as:

Grid 2	<u>Grid 1</u>
qj0(1,i2, <b>l</b> ,1) =	q(4,k1,4, <b>l</b> )
qj0(2,i2, <b>l</b> ,1) =	q(3,k1,4, <b>l</b> )
qj0(3,i2, <b>l</b> ,1) =	q(2,k1,4, <b>l</b> )
qj0(4,i2, <b>l</b> ,1) =	q(1,k1,4, <b>l</b> )
qj0(1,i2, <b>l</b> ,2) =	q(4,k1,3, <b>l</b> )
qj0(2,i2, <b>l</b> ,2) =	q(3,k1,3, <b>l</b> )
qj0(3,i2, <b>l</b> ,2) =	q(2,k1,3, <b>l</b> )
qj0(4,i2, <b>l</b> ,2) =	q(1,k1,3, <b>l</b> )

The boundary conditions at the k = 1 face of grid 2 would be set as:

<u>Grid 2</u>	Grid 3
qk0(1,i2,l,1) =	q(j3,8,1, <b>l</b> )
qk0(2,i2,l,1) =	q(j3,7,1, <b>l</b> )
qk0(3,i2, <b>l</b> ,1) =	q(j3,6,1, <b>l</b> )
qk0(4,i2, <b>l</b> ,1) =	q(j3,5,1, <b>l</b> )
qk0(1,i2, <b>l</b> ,2) =	q(j3,8,2, <b>l</b> )
qk0(2,i2,l,2) =	q(j3,7,2, <b>l</b> )
qk0(3,i2,l,2) =	q(j3,6,2, <b>l</b> )
qk0(4,i2,l,2) =	q(j3,5,2, <b>l</b> )

The boundary conditions from **ksta** to **kend**-1 on the i = 1 face of grid 3 would be set as:

Grid 3	Grid 2
qi0(j3,5, <b>l</b> ,1) =	q(4,1,i2, <b>l</b> )
qi0(j3,6,l,1) =	q(3,1,i2, <b>l</b> )
qi0(j3,7, <b>l</b> ,1) =	q(2,1,i2, <b>l</b> )
qi0(j3,8,1,1) =	q(1,1,i2, <b>l</b> )
qi0(j3,5,l,2) =	q(4,2,i2, <b>l</b> )
qi0(j3,6, <b>l</b> ,2) =	q(3,2,i2, <b>l</b> )
qi0(j3,7,l,2) =	q(2,2,i2, <b>l</b> )
qi0(j3,8,l,2) =	q(1,2,i2, <b>l</b> )

6.3.2 Patched-Grid Interpolation

#### (Input Line Type Twenty-Seven)

6.3.2.1 General Information

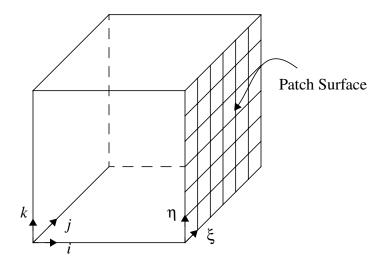


Figure 6-15. Patched-grid surface.

This section describes the "basics" of patching in CFL3D and then shows a simple static patching example to illustrate the basic concepts. Next, an example of dynamic patching is given. Patched-grid interpolation is designed for communication between grids that share a common face, but are *not* C<sup>0</sup> continuous. For example, a global (3-d) grid is represented by the "box" in Figure 6-15. The grid must be right-handed as shown with the *i*, *j*, and *k* axes drawn in the figure. Patching may occur on an *i* = constant, *j* = constant, or *k* = constant surface. In the figure, an *i* = constant surface is indicated as the patched surface. Note that the adjacent grid involved in the patch is not shown in the figure. Patching works best when the spacing of the adjacent grid in the normal direction to the patch is the same as that in the other grid. While the patch surface is indicated by the  $\xi$  and  $\eta$  axes.

The global and local indices correspond as follows (this is important when determining the input parameters for dynamic patching in Section 3.42 on page 49):

If the patch surface is on an i = constant surface (as shown in Figure 6-15), then

$$k \Leftrightarrow \eta$$

$$j \Leftrightarrow \xi$$
(6-41)

If the patch surface is on a j = constant surface, then

$$i \Leftrightarrow \eta$$
  
 $k \Leftrightarrow \xi$ 
(6-42)

If the patch surface is on a k = constant surface, then

$$i \Leftrightarrow \eta$$

$$j \Leftrightarrow \xi$$
(6-43)

One grid may have multiple grids patched to it. For example, the flow over a nose cone illustrated in Figure 6-16 shows block 1 patched to both block 2 and block 3. Note that the orientation of grid lines in block 3 is different from that in blocks 1 and 2 to illustrate the features of the patching algorithm.

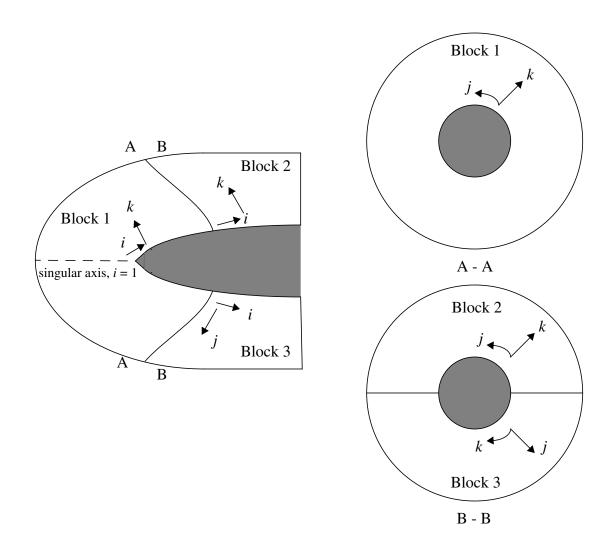


Figure 6-16. Patched-grid example.

To provide two-way communication between all blocks, a total of seven interpolations are required for this case, as follows:

Interpolation:	Description:
1	To $i = idim$ of block 1 from $i = 1$ of blocks 2 and 3 (two "from" blocks)
2	To $i = 1$ of block 2 from $i = idim$ of block 1
3	To $i = 1$ of block 3 from $i = idim$ of block 1
4	To $j = 1$ of block 2 from $k = 1$ of block 3
5	<i>To</i> $j = jdim$ of block 2 <i>from</i> $k = kdim$ of block 3
6	To $k = 1$ of block 3 from $j = 1$ of block 2
7	To $k = kdim$ of block 3 from $j = jdim$ of block 2

The interpolations may be input in any order (but must be consistent once the order is chosen).

## 6.3.2.2 Description/Discussion of Input Parameters

Basically, the patch algorithm works as follows. The interpolations are cycled through, one at a time, so that at any given time there is one surface being interpolated to. Note, however, that there may be more than one surface being interpolated from. In order to interpolate from one (or more) block(s) to another, interpolation coefficients are required. These are found by expressing the cell-center coordinates  $x_c$ ,  $y_c$ ,  $z_c$  on the "to" side of the patch surface in terms of a nonlinear polynomial in  $\xi$ ,  $\eta$ , where  $\xi$ ,  $\eta$  are the local coordinates on the "from" side of the patch surface. Newton iteration is used to invert the polynomial to find the local coordinates of the cell center,  $\xi_c$ ,  $\eta_c$ .

Because of the nonlinearity of the equation, problems may arise in the iteration process. The parameters **ifit**, **limit**, and **itmax** may be adjusted to try to overcome any convergence difficulties. There must be **ninter2** values for each of these three parameters. **Limit** is the maximum step size in  $\xi$  or  $\eta$  allowed during the search procedure. The value 1 seems to be a good general choice. **Itmax** is the maximum number of search steps allowed per grid point. A rule of thumb is **limit** × **itmax** ≈ the maximum dimension of *i*, *j*, or *k* on a patch surface.

If it controls the order of the polynomial fit used to relate x, y, or z to  $\xi$  and  $\eta$ . If t = 1 for linear in both  $\xi$  and  $\eta$  (bilinear). If t = 2 for quadratic in both  $\xi$  and  $\eta$  (degenerate biquadratic). If t = 3 for quadratic in  $\xi$ , linear in  $\eta$ . If t = 4 for linear in  $\xi$ , quadratic in  $\eta$ .

Some tips for choosing ifit are:

- A grid with highly curved grid lines in one of the *local* directions will need a quadratic polynomial in that direction.
- Refer to Section 6.3.2.1 and the user's own knowledge of the grid to decide which (if any) of the local directions  $\xi$  or  $\eta$  are highly curved.
- Use the lowest order fit which seems reasonable in each direction.

• One-to-one patching can always be done with a bilinear fit, regardless of curvature.

To and from are numbers containing attributes of the "to" side of the patch surface and the "from" side of the patch surface. For each of the **ninter2** interpolations there is only one value of to, but there may be more than one value of from which is set by the parameter **nfb**. (**Nfb** is the number of blocks on the "from" side of the patch surface.) The values of to and from are of the form:

## to/from = Nmn

where "N" indicates the block number; "m" indicates the coordinate which is constant on the patch surface (may differ on either side of the patch):

 $m = 1 \Rightarrow i = \text{constant}$   $m = 2 \Rightarrow j = \text{constant}$  $m = 3 \Rightarrow k = \text{constant}$ 

and "n" indicates on which of the two possible m = constant surfaces the patch surface occurs:

n = 1 for patch on m = 1 surface n = 2 for patch on m = mdim surface

The following inputs pertain to the nose cone case in Figure 6-16:

Interpolation #	<u>to</u>	<u>nfb</u>	<u>from</u>
1	112	2	211, 311
2	211	1	112
3	311	1	112
4	221	1	331

Example

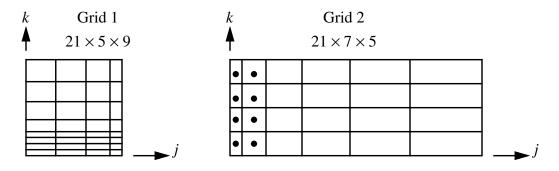


Figure 6-17. Patched-grid example.

Suppose the *k* dimension of grid 1 in the first example of 1-1 blocking (see Figure 6-11 on page 102) was actually **kdim** = 9. Grid 1 might look something like that drawn in Figure 6-17. As in Figure 6-11, Figure 6-17 shows an i = constant face (not necessarily the *same* constant) of two grids. Notice that the spacing in the direction normal to the interface (i.e. *j*) is approximately constant across the interface.

In Version 5.0 of CFL3D, patched-grid interpolation data for *static* patched interfaces (interfaces that do not change with time) must be obtained as a preprocessing step. The code ronnie is designed for this task. *Dynamic* patched interfaces, such as occur when grids slide past one another, are computed internally in CFL3D.

Note that, if two grids with very different sizes are patched together, it may be necessary to use a limiter on the gradients there. To do this, replace (hard-wire) the calls to int2 with calls to int3 instead.

# 6.3.3 Chimera Grid Interpolation

## (Input Line Type Ten)

Grid overlapping, also known as the overset-grid method or chimera technique, requires neither 1-1 connectivity nor a shared interface to pass flow information from one grid to another. With this method, a variety of grid topologies can be used together. The chimera implementation used in CFL3D is based on the method of Benek et al.<sup>12</sup> As a simplified example of grid overlapping, consider a cross section of a polar grid and a Cartesian mesh as shown Figure 6-18. Suppose the grids overlap as in Figure 6-19. Since both grids cover the same area, computations on both grids in this area would be redundant. Therefore, certain points on the Cartesian grid will be eliminated from the computation. The polar grid is used to carve a "hole" in the Cartesian grid. In this example, the hole is defined as any cell-center point of the Cartesian grid interior to the k = 4 grid line of the polar mesh. The Cartesian mesh with the hole carved out is illustrated in Figure 6-20. Any cell center point of the Cartesian grid located within this hole is designated a "hole point". The first two "nonhole" cell-center points of the Cartesian grid that border a hole point both vertically and horizontally are labelled "fringe points". The remaining grid points that have not been designated as either hole or fringe points are called "field" points. Figure 6-21 depicts the hole, fringe, and boundary points for this example. Each fringe point of the Cartesian grid falls within a "target cell" of the polar grid.

A searching algorithm is used to identify the particular eight points that define the hexahedral target cell. The search begins with an initial guess for the target cell. Next, the current target cell is isoparametrically mapped into a unit cube in computational space. The same transformation into the mapped coordinate system is then applied to the fringe point; if the mapped fringe point lies in the same unit cube as the current target cell, then that target cell in fact encases the fringe point. If the mapped fringe point lies outside the unit cube, then the current target cell is not the correct choice. However, the magnitude and direction of the mapped fringe point relative to the current target cell may be used to choose a new guess for the target cell. The mapping process is repeated until the correct target cell is identified. With the correct target cell identified, the data are transferred from the target cell to the fringe point with trilinear interpolation in computational space. Outer boundary values of the polar grid are determined in a similar manner. The MultiGeometry Grid Embedder (MaGGiE) code, written specifically for CFL3D by Baysal et al.<sup>11</sup>, is used to determine the interpolation information between grids.

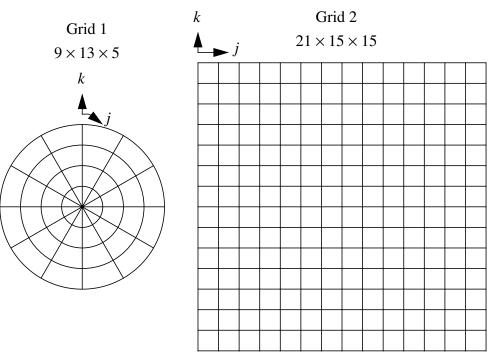


Figure 6-18. Grid-overlapping grid examples.

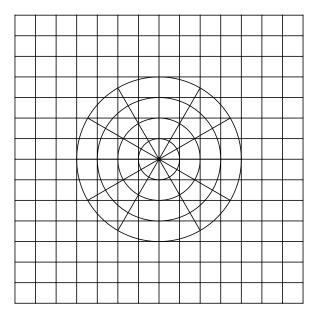


Figure 6-19. Overlapped grids.

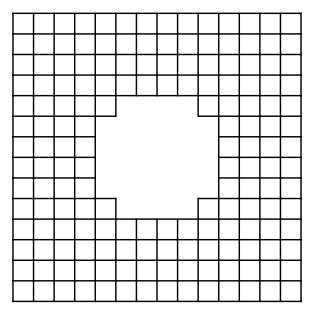


Figure 6-20. Hole in Cartesian grid in region of polar grid.

- Hole Point for Cartesian Grid
- + Fringe Point for Cartesian Grid
- Boundary Point for Polar Grid

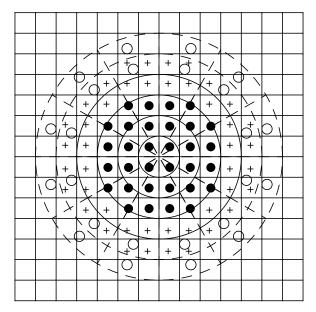


Figure 6-21. Hole, fringe, and boundary points for overlapped grid example.

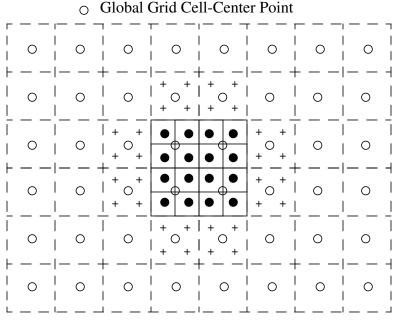
# 6.3.4 Embedded Mesh

#### **Regular refinement of coarse mesh**

Embedded grids are useful when high gradient areas are limited to an identifiable region (see reference 25). An embedded grid can be placed in that region to resolve the flow field without refining the entire mesh. As an example, an embedded grid scheme is shown in Figure 6-22 for two dimensions. The diagram represents full refinement in both directions. The solid lines define a finer mesh embedded completely within a coarser mesh depicted by the dashed lines. In the figure, a portion of the flow field is covered by both the embedded mesh and a portion of the coarser grid. The grids are coupled together during the solution process. The cell-center variables on a coarser grid cell which underlies a finer embedded grid cell are replaced with a volume-weighted restriction of variables from the four (2-d) or eight (3-d) finer grid cells, similar to the restriction operators used in a global multigrid scheme.

## Embedded Grid Cell-Center Point

+ Embedded Grid Boundary Point



#### Figure 6-22. Embedded grid example.

For the embedded grid, the computational boundaries occur either at a physical boundary, such as a wing, a symmetry plane, an inflow/outflow plane along a zonal interface (one-to-one, patched, or overset), or along an interior computational surface of a coarser grid. Along an interior surface, two additional lines of data corresponding to an analytical continuation of the finer grid cell centers are constructed from linear interpolation of the coarser grid state variables. An enlarged view of the lower-left corner of the embedded mesh of Figure 6-22 is shown in Figure 6-23. Arrows indicate which coarse grid points are used in the linear interpolation scheme to obtain a couple of the boundary points drawn. If the input parameter **iconsf** = 1 ("LT20 - Mesh Sequencing and Multigrid"), then global conservation is enforced by replacing the coarse grid flux at an embedded grid boundary with the sum of the finer grid fluxes which share the common interface. If **iconsf** = 0, the coarse grid flux is computed using the volume-weighted restriction from the fine grid, and conservation is not insured.

- Embedded Grid Cell-Center Point
- + Embedded Grid Boundary Point
- O Global Grid Cell-Center Point

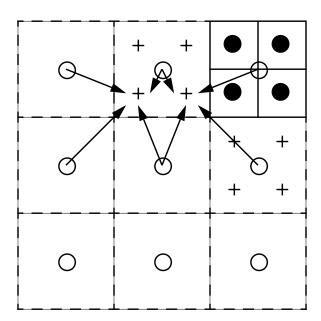


Figure 6-23. Enlarged portion of embedded mesh.

When designing an embedded mesh, it is important to remember that at least two layers of coarse grid cells should surround the embedded boundaries unless the embedded boundary is set with a physical boundary condition (e.g. solid wall, plane of symmetry, or far-field) or a zonal boundary such as a one-to-one, patched, or overset type. In the example of Figure 6-22, there are two layers of coarse grid cells above and below the embedded mesh and three layers of coarse grid cells on the left and right sides of the embedded mesh.

Suppose the grids in Figure 6-22 are in an i = constant plane. Also, **jdim** = 9 and **kdim** = 7 for the coarse grid and **jdim** = 5 and **kdim** = 5 for the embedded mesh. The pertinent input would look something like:

Line Type								
6	NGRID	NPLOT3D	NPRINT	NWREST	ICHK	I2D	NTSTEP	ITA
	2	1	0	100	0	0	1	1
7	NCG	IEM	IADVANCE	IFORCE	IVISC(I)	IVISC(J)	IVISC(K)	
	0	0	0	0	0	0	0	
	0	1	0	0	0	0	0	
8	IDIM	JDIM	KDIM					
	21	9	7					
	41	5	5					
10	INEWG	IGRIDC	IS	JS	KS	IE	JE	KE
	0	0	0	0	0	0	0	0
	0	1	1	4	3	21	6	5
20	MSEO	MGFLAG	ICONSF	MTT	NGAM			
20	1	0	1001001	0	0			
	1	0	T	0	0			
22	NCYC	MGLEVG	NEMBL	NITFO				
	1	1	1	0				

In this example, the embedded mesh extends the entire length of the global grid in the i direction. The dimensions of the embedded mesh must satisfy a particular relationship to the data in "LT10 - Embedded Mesh Specifications". The input parameters **is**, **ie**, **js**, **je**, **ks**, and **ke** are indices in the global grid to which the embedded mesh is connected. Thus, the following must be true:

$$\mathbf{idim}_{embedded} = 2(\mathbf{ie} - \mathbf{is})_{global} + 1$$
(6-44)

Analogous relationships hold for the *j* and *k* directions. Note that the only exception to Equation (6-44) occurs in the *i* direction for 2-d cases, where  $\mathbf{is} = 2$ ,  $\mathbf{ie} = 1$  and  $\mathbf{idim}_{embedded} = \mathbf{idim}_{global} = 2$ .

There is also an option called "semi-coarsening" in which the number of points in the *i* direction are the same for the embedded grid and the global grid in the embedded region. This is helpful if there are sufficient points in the *i* direction, but grid enrichment in j - k planes is desirable. There is an internal check for this option and the only change in the input sample above is to set **idim** = 21 for the embedded grid as well. Note that semi-coarsening can only be used in the *i* direction.

Another nice option with grid embedding is to add an embedded grid to a previouslyrun coarse grid problem. For instance, suppose a converged coarse grid solution is obtained and the flow field looks well resolved everywhere except in a vortex region. An embedded grid for that region can be tacked on to the original grid file and the case restarted. On the *first* run of the restart, set **inewg** = 1 for the embedded grid. This lets the code know that a solution not on the current restart file is beginning. On the next run, the embedded grid solution *will* be on the restart file, so set **inewg** = 0.



# **CHAPTER 7**

# Convergence Acceleration

In CFL3D, the convergence rate at which many problems are solved can be accelerated with the use of multigrid, mesh sequencing, or a combination of the two. The following sections describe the two techniques. The use of multigrid with grid overlapping and embedded grids is also discussed.

Multigrid is a *highly* recommended option available in CFL3D. The improvement in convergence acceleration afforded with multigrid makes it very worthwhile to learn about and utilize. Table 7–1 illustrates this fact for two cases, a 2-d airfoil and a 3-d forebody. A work unit is defined here to be the "equivalent" fine grid iteration. For example, on the finest grid level, 1 iteration = 1 work unit; on the next-to-finest grid level, 1 iteration = 1/8 (1/4 in 2-d) of a work unit; on the next coarser level, 1 iteration = 1/64 (1/16 in 2-d) of a work unit; and so forth. The residual and lift coefficient convergence histories for these two cases are shown in Figure 7-1 and Figure 7-2. The term "full multigrid" implies that mesh sequencing is used in conjunction with multigrid. See "Mesh Sequencing" on page 134. Note that there is an additional CPU penalty for the overhead associated with the multigrid scheme that is not reflected in the work unit measure, however, for most problems, the overhead is small.

	Approximate Number of Work Units to Achieve the Final Lift						
Case	Without Multigrid	With Multigrid	With Full-Multigrid				
2-d NACA 4412 Airfoil	10,000	1000	750				
3-d F-18 Forebody	>> 1500*	500	500				

Table 7–1. Convergence improvements with multigrid.

\*This case was stopped after 1500 iterations when it became clear that the lift coefficient was far from converged and to continue running the case would have been a waste of computer resources.

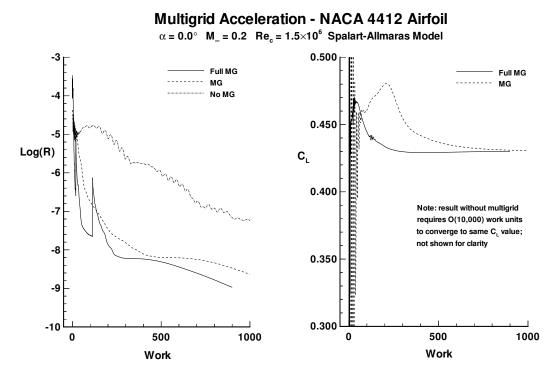


Figure 7-1. Convergence acceleration for 2-d NACA 4412 airfoil case.

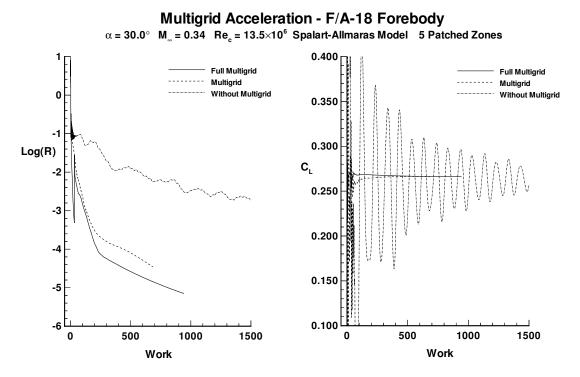


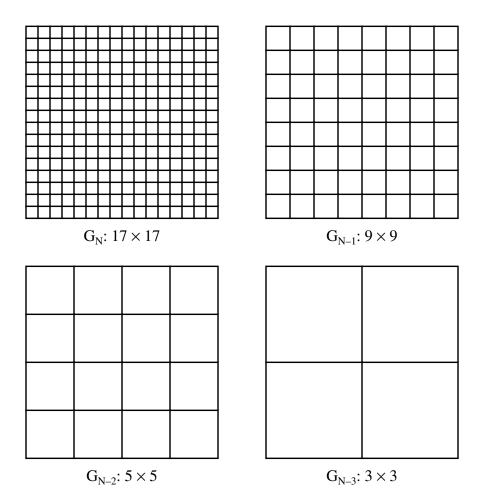
Figure 7-2. Convergence acceleration for 3-d F-18 Forebody case.

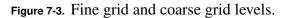
# 7.1 Multigrid

# 7.1.1 Global Grids

In a single-grid-level algorithm, the flow equations are solved only on the finest grid. The convergence rate for this algorithm is typically good during the initial stages of the computation but soon degrades significantly due to poor damping of low-frequency errors.<sup>13</sup> Multigrid methods, in which a sequence of coarser grids are used to eliminate the low-frequency errors, can significantly improve the convergence rate of the algorithm. The full-approximation-scheme multigrid algorithm, as developed and applied to inviscid three-dimensional flows in references 5, 6 and 7, is available in CFL3D to accelerate convergence of a solution to a steady state. A sequence of grids  $G_0, G_1, ..., G_N$  is defined, where  $G_N$  denotes the finest grid and coarser grids are formed by successively deleting every other grid line in all three coordinate directions (except in 2-d, where the *i* direction is not coarsened). The fine grid serves to damp the high-frequency errors; the coarser grids damp the low-frequency errors. The coarse grids are solved with a forcing function on the right-hand side arising from restricting the residual from the finer meshes. The forcing function is the relative truncation error between the grids such that the solutions on the coarser meshes are driven by the fine grid residual. A correction is calculated on the coarser meshes and passed up to the next finest mesh using trilinear interpolation in the prolongation step.

As an example, consider a 2-d grid of dimensions  $17 \times 17$  as shown in Figure 7-3. Suppose it is desired to use four levels of multigrid, i.e. the fine grid level and three coarser grid levels. The three coarser grid levels, along with their dimensions, are also shown in Figure 7-3. Figure 7-4 illustrates a typical W-cycle multigrid scheme for four grid levels. In the figure, *n* refers to the current iteration. The multigrid cycle begins with a Navier-Stokes calculation performed at the finest grid level specified. The fine grid residual and flow-field variables are restricted to the next grid level where they are utilized to perform a Navier-Stokes calculation at this level. The process continues until the coarsest grid level is reached. After a Navier-Stokes calculation is performed at this level, the solution is prolongated back "up" the cycle. With the W-cycle, additional calculations and the restrictions and prolongation steps are repeated at the coarser grid levels giving the "W" pattern illustrated in Figure 7-4. The other option available is a V-cycle. An example of a V-cycle for the four grid level case is illustrated in Figure 7-5. Generally, a W-cycle is more effective since additional work is done at the coarse grid level. However, this may be case dependent. Note that, if only two grid levels are specified, a V-cycle will always be performed even if a W-cycle is specified in the input file.





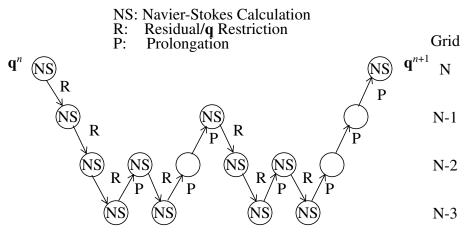


Figure 7-4. Multigrid W-cycle.

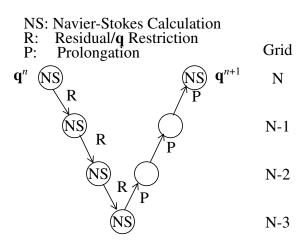


Figure 7-5. Multigrid V-cycle.

For a multigrid W-cycle, the pertinent lines of input for the example in Figure 7-3 would look something like:

Line Typ	<u>be</u>							
5	DT	IREST	IFLAGTS	FMAX	IUNST	CFLTAU		
	-1.0	0	0	1.0	0	10.0		
6	NGRID	NPLOT3D	NPRINT	NWREST	ICHK	I2D	NTSTEP	ITA
	1	0	0	0	0	1	1	1
7	NCG	IEM	IADVANCE	IFORCE	IVISC(I)	IVISC(J)	IVISC(K)	
	3	0	0	1	0	0	0	
8	IDIM	JDIM	KDIM					
	2	17	17					
20	MSEQ	MGFLAG	ICONSF	MTT	NGAM			
	1	1	0	0	2			
22	NCYC	MGLEVG	NEMGL	NITFO				
	200	4	0	0				
23	MIT1	MIT2	MIT3	MIT4	MIT5			
	1	1	1	1	1			

# 7.1.2 A Word About Grid Dimensions

When determining how many multigrid levels are available for a particular grid, use the following formula:

$$m_c = \frac{m_f - 1}{2} + 1 = \frac{m_f + 1}{2}$$
(7-1)

where  $m_f$  is a fine-grid dimension (**idim**, **jdim**, or **kdim**) and  $m_c$  is the corresponding coarse-grid dimension. Then rename  $m_c$  as  $m_f$  and compute Equation (7-1) again to determine an even coarser grid dimension. When  $m_c$  is an even number, that is as coarse as that grid dimension can go.

For example, suppose a grid has idim = 37, jdim = 65, kdim = 65. The coarse grid below that would have idim = 19, jdim = 33, kdim = 33. The next coarser grid would have idim = 10, jdim = 17, kdim = 17. Since idim is an even number, that is as coarse as the grid can go; even though jdim and kdim can be reduced to coarser numbers, the entire grid is restricted by the idim = 10 value.

If the use of multigrid is desirable, it is important to plan ahead in the grid generation step of the computational problem. Choose grid dimensions that are "good" multigrid numbers, such as 129 (65, 33, 17, 9, 5, 3), 73 (37, 19), and 49 (25, 13, 7). Generally, two or three coarser grid levels are satisfactory.

It is best if all grid *segments* are also multigridable (see "LT14 - I0 Boundary Condition Specification" on page 32 through "LT19 - KDIM Boundary Condition Specification" on page 35). For example, a face with **jdim** = 65 might have a portion from j = 1 to 17 as a one-to-one interface, a portion from j = 17 to 41 as a viscous wall, and a portion from j = 41 to 65 as a patched interface. Each of these segment lengths (7, 25, and 25) is multigridable down three levels. Note that CFL3D will sometimes work fine even if some grid segments are not multigridable: the code can assign indices on the coarser levels that denote different physical locations than the indices on the fine grid and still converge on the finest level. However, this does *not* work all the time. For example, if a C-mesh has a wrap-around dimension of **jdim** = 257 and the wake extends from j = 1 to 40 and j = 218to 257 (with one-to-one point matching), the code will create a coarser level with the wake from i = 1 to 20 and i = 109 to 129 (using Equation (7-1)). On this level, i = 20 is a phys*ically different point* from j = 109 (it is not even the actual trailing edge), and will yield a boundary condition error when the code is run, because these points are expected to match in a one-to-one fashion. It is currently not necessary to specify laminar regions with multigridable numbers. See Note (3) on page 29 in the LT9 - Laminar Region Specification description.

It is also important to consider the grid *geometry* when defining the grid dimensions. If multigrid is used, be sure to have any important geometric features (such as corners) located at multigridable points. Otherwise, on coarser levels, the geometry may change significantly, resulting in poor multigrid performance.

Table 7-2. Grid sizes multigridable to three additional level.													
Grid:	Coa	arser Lev	els:		Grid:	Coarser Levels:				Grid:	Coarser Levels:		els:
9	5	3	2		345	173	87	44		673	337	169	85
17	9	5	3		353	177	89	45		681	341	171	86
25	13	7	4		361	181	91	46		689	345	173	87
33	17	9	5		369	185	93	47		697	349	175	88
41	21	11	6		377	189	95	48		705	353	177	89
49	25	13	7		385	193	97	49		713	357	179	90
57	29	15	8		393	197	99	50		721	361	181	91
65	33	17	9		401	201	101	51		729	365	183	92
73	37	19	10		409	205	103	52		737	369	185	93
81	41	21	11		417	209	105	53		745	373	187	94
89	45	23	12		425	213	107	54		753	377	189	95
97	49	25	13		433	217	109	55		761	381	191	96
105	53	27	14		441	221	111	56		769	385	193	97
113	57	29	15		449	225	113	57		777	389	195	98
121	61	31	16		457	229	115	58		785	393	197	99
129	65	33	17		465	233	117	59		793	397	199	100
137	69	35	18		473	237	119	60		801	401	201	101
145	73	37	19		481	241	121	61		809	405	203	102
153	77	39	20		489	245	123	62		817	409	205	103
161	81	41	21		497	249	125	63		825	413	207	104
169	85	43	22		505	253	127	64		833	417	209	105
177	89	45	23		513	257	129	65		841	421	211	106
185	93	47	24		521	261	131	66		849	425	213	107
193	97	49	25		529	265	133	67		857	429	215	108
201	101	51	26		537	269	135	68		865	433	217	109
209	105	53	27		545	273	137	69		873	437	219	110
217	109	55	28		553	277	139	70		881	441	221	111
225	113	57	29		561	281	141	71		889	445	223	112
233	117	59	30		569	285	143	72		897	449	225	113
241	121	61	31		577	289	145	73		905	453	227	114
249	125	63	32		585	293	147	74		913	457	229	115
257	129	65	33		593	297	149	75		921	461	231	116
265	133	67	34		601	301	151	76		929	465	233	117
273	137	69	35		609	305	153	77		937	469	235	118
281	141	71	36		617	309	155	78		945	473	237	119
289	145	73	37		625	313	157	79		953	477	239	120
297	149	75	38		633	317	159	80		961	481	241	121
305	153	77	39		641	321	161	81		969	485	243	122
313	157	79	40		649	325	163	82		977	489	245	123
321	161	81	41		657	329	165	83		985	493	247	124
329	165	83	42		665	333	167	84		993	497	249	125
337	169	85	43										

For a handy reference, the following table lists the grid sizes (< 1000) that are multigridable to three additional levels:

### 7.1.3 Overlapped Grids

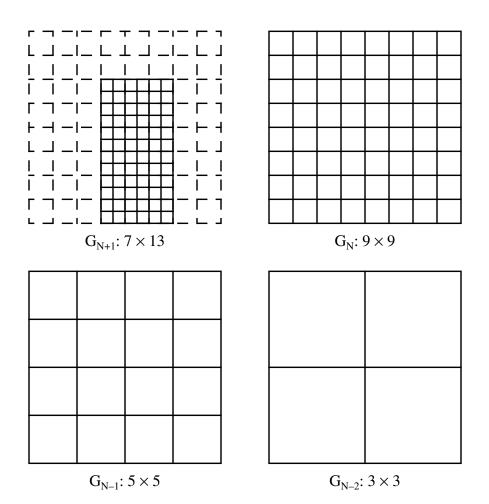
With the grid-overlapping scheme, interpolation stencils are determined at the finest grid level only. Therefore, information would be missing at the coarser grid levels of the multigrid scheme. To avoid the need to obtain interpolation stencils at the coarser grid levels prior to the computation and the corresponding problems that most likely will arise, the multigrid scheme has been modified to accommodate grid-overlapping cases.<sup>24</sup>

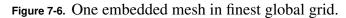
First, if a fine grid boundary is supplied overlap information as the boundary condition, extrapolation is used for that boundary condition on all coarser grids. Second, for the restriction step of the multigrid scheme, all points are interpolated to coarser grids regardless of whether or not they are in a hole at the fine grid level. The flow variables in the hole are simply at free-stream conditions. Finally, for the prolongation step, only corrections on field points are used to update the finest mesh solution. The hole points are not updated since they will be overwritten with free-stream values when a Navier-Stokes calculation is made at each fine grid level. Likewise, the fringe points are not corrected since they will be updated with flow information from other grids at the same fine grid level.

Note that in Chapter 3 under "LT20 - Mesh Sequencing and Multigrid" on page 35, it is stated that the W-cycle is *not* recommended with overlapped grids. Therefore, set ngam = 1.

### 7.1.4 Embedded Grids

For time advancement, embedded grids basically become an extension of the multigrid scheme. Consider the grid system shown in Figure 7-6. The finest global grid has dimensions  $9 \times 9$  and there are two coarser global grids for multigrid purposes. In addition, there is a  $7 \times 13$  grid embedded in the finest global grid as shown in the figure. The multigrid (W-cycle) scheme would follow the course drawn in Figure 7-7.





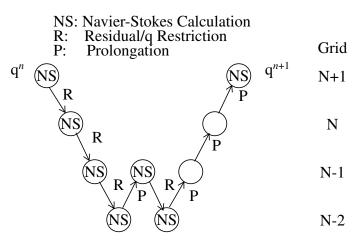


Figure 7-7. Multigrid W-cycle with one embedded grid level.

For a multigrid W-cycle, the pertinent lines of input for the example in Figure 7-6 would look something like:

Line	<u>Type</u>							
5	DT	IREST	IFLAGTS	FMAX	IUNST	CFLTAU		
	-1.0	0	0	1.0	0	10.0		
6	NGRID	NPLOT3D	NPRINT	NWREST	ICHK	I2D	NTSTEP	ITA
	2	0	0	0	0	1	1	1
7	NCG	IEM	IADVANCE	IFORCE	IVISC(I)	IVISC(J)	IVISC(K)	
	2	0	0	1	0	0	0	
0	0	1	0	0	0	0	0	
8	IDIM	JDIM	KDIM					
	2	9	9					
	2	7	13					
10								
10	INEWG	IGRIDC	IS	JS	KS	IE	JE	KE
	0	0	0	0	0	0	0	0 7
	0	T	T	4	T	2	/	1
20	MSEO	MGELAG	ICONSF	MTT	NGAM			
20	1	1	1	0	2			
	1	Ŧ	-	0	-			
22	NCYC	MGLEVG	NEMGL	NITFO				
	200	3	1	0				
23	MIT1	MIT2	MIT3	MIT4	MIT5			
	1	1	1	1	1			

For the boundary conditions on the embedded grid, use the appropriate physical or connectivity **bctype** on any face that lies along an edge of its parent grid and **bctype** = 0 for any face that lies wholly within its parent grid. Suppose another embedded grid level is added to the grid system as in Figure 7-8. In the figure two grids are embedded in the first embedded grid. Note that only one embedded grid *level* is added by their addition.

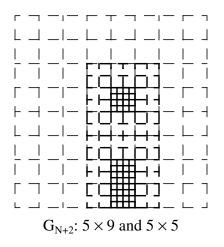


Figure 7-8. Two embedded grids within an embedded grid.

Line 7	Гуре							
5	DT	IREST	IFLAGTS	FMAX	IUNST	CFLTAU		
	-1.0	1	0	1.0	0	10.0		
6	NGRID	NPLOT3D	NPRINT	NWREST	ICHK	I2D	NTSTEP	ITA
	4	0	0	0	0	1	1	1
7	NCG	IEM	IADVANCE	IFORCE	IVISC(I)	IVISC(J)	IVISC(K)	
	2	0	0	1	0	0	0	
	0	1	0	0	0	0	0	
	0	2	0	0	0	0	0	
0	0	2	0	0	0	0	0	
8	IDIM	JDIM	KDIM					
	2	9	9					
	2	7	13					
	2	5 5	9 5					
	Z	5	5					
10	INEWG	IGRIDC	IS	JS	KS	IE	JE	KE
	0	0	0	0	0	0	0	0
	Ő	1	1	4	1	2	7	7
	1	2	1	3	1	2	5	5
	1	2	1	3	9	2	5	11
20								
20	~	MGFLAG		MTT	NGAM			
	1	1	1	0	2			
22								
LL	NCYC		NEMGL	NITFO				
22	200	3		0	_			
23	MIT1		MIT3	MIT4	MIT5			
	1	1	1	1	1			

The corresponding input would resemble:

Note that **inewg** should be set to 0 for grids 3 and 4 if this case is restarted.

Multigrid can also be performed at the embedded grid levels themselves by setting mgflag = 2. The embedded grid uses the coarser grid levels in which it is embedded, so ncg for the embedded grid itself should remain 0. The input for this case would resemble:

Line T	ype							
5	DT	IREST	IFLAGTS	FMAX	IUNST	CFLTAU		
	-1.0	0	0	1.0	0	10.0		
6	NGRID	NPLOT3D	NPRINT	NWREST	ICHK	I2D	NTSTEP	ITA
_	2	0	0	0	0	1	1	1
7	NCG	IEM	IADVANCE	IFORCE	IVISC(I)	IVISC(J)	IVISC(K)	
	2	0	0	1	0	0	0	
0	0	1	0	0	0	0	0	
8	IDIM	JDIM	KDIM					
	2	9	9					
	2	7	13					
10	INEWG	IGRIDC	IS	JS	KS	IE	JE	KE
10	0	0	0	0	0	0	0	0
	0	1	1	4	1	2	7	7
	0	-	-	-	-	-		
20	MSEQ	MGFLAG	ICONSF	MTT	NGAM			
	1	2	1	0	2			
22								
22	NCYC	MGLEVG	NEMGL	NITFO				
• •	200	3	1	0				
23	MIT1	MIT2	MIT3	MIT4	MIT5			
	1	1	1	1	1			

### 7.2 Mesh Sequencing

When setting up a CFD problem, initial conditions are set at every point on a grid. Usually, the closer the guess is to the final solution the quicker the case will converge. In CFL3D, the initial conditions for a problem are set at free-stream conditions for a single-grid-level case. However, if mesh sequencing is utilized, a better "guess" can be made for the initial conditions on the finer grid where the computations are most expensive.

Suppose a solution is desired on the  $9 \times 9$  grid depicted in Figure 7-9. Instead of starting the solution with free-stream conditions, a solution could first be obtained on the  $5 \times 5$ grid shown in Figure 7-10. While the solution on the  $9 \times 9$  grid would be more accurate than that on the  $5 \times 5$  grid, the coarser grid's solution would be closer to the fine grid solution than free-stream conditions. Since the coarse grid solution can be computed more quickly than the fine grid solution, it is usually beneficial to use the coarse grid solution as the initial condition for the fine grid.

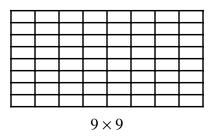


Figure 7-9. Simple grid example.

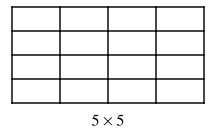


Figure 7-10. Mesh sequencing sample grid.

There are two ways to implement mesh sequencing in CFL3D. The first way is to completely converge the solution on the coarse grid before mapping it up to the fine grid. For the sample grids, the pertinent input would look like:

Line Type

5	DT	IREST	IFLAGTS	FMAX	IUNST	CFLTAU		
	-1.0	0	0	1.0	0	10.0		
6	NGRID	NPLOT3D	NPRINT	NWREST	ICHK	I2D	NTSTEP	ITA

1

		_					
_	1	0	0	0	0	1	1
7	NCG	IEM	IADVANCE	IFORCE	IVISC(I)	IVISC(J)	IVISC(K)
	2	0	0	1	0	0	0
8	IDIM	JDIM	KDIM				
	2	9	9				
20	MSEQ	MGFLAG	ICONSF	MTT	NGAM		
	2	1	0	0	1		
22	NCYC	MGLEVG	NEMGL	NITFO			
	200	2	0	0			
	0	3	0	0			
23	MIT1	MIT2	MIT3	MIT4	MIT5		
	1	1	1	1	1		
	1	1	1	1	1		

Note that multigrid is also employed. The coarse grid solution can be examined and, if the solution on the coarse grid is not converged, the case can be restarted on that grid. After the coarse grid solution is converged, the solution is mapped to the fine grid and the calculations continue at the fine grid level. The pertinent input for this step would look something like:

Line Typ	e							
5	DT	IREST	IFLAGTS	FMAX	IUNST	CFLTAU		
	-1.0	1	0	1.0	0	10.0		
6	NGRID	NPLOT3D	NPRINT	NWREST	ICHK	I2D	NTSTEP	ITA
	1	0	0	0	0	1	1	1
7	NCG	IEM	IADVANCE	IFORCE	IVISC(I)	IVISC(J)	IVISC(K)	
	2	0	0	1	0	0	0	
8	IDIM	JDIM	KDIM					
	2	9	9					
20	MSEO	MGFLAG	ICONSF	MTT	NGAM			
20	2	1 Inor	0	0	1			
	2	Ŧ	0	0	T			
22	NCYC	MGLEVG	NEMGL	NITFO				
	1	2	0	0				
	200	2	0	0				
23	MIT1	MIT2	MIT3	MIT4	MIT5			
	1	1	1	1	1			
	1	1	1	1	1			

Here, one iteration is performed on the coarse grid and two-hundred iterations are performed on the fine grid. At the end of this run, only the fine grid solution is available for examination. (So be sure to save the coarse grid solution *before* this step if it will be needed later for grid refinement studies, etc.) The case can be resubmitted for additional computations on the fine grid with the following input:

Line T	<u>ype</u>							
5	DT	IREST	IFLAGTS	FMAX	IUNST	CFLTAU		
	-1.0	1	0	1.0	0	10.0		
6	NGRID	NPLOT3D	NPRINT	NWREST	ICHK	I2D	NTSTEP	ITA
	1	0	0	0	0	1	1	1
7	NCG	IEM	IADVANCE	IFORCE	IVISC(I)	IVISC(J)	IVISC(K)	
	2	0	0	1	0	0	0	
8	IDIM	JDIM	KDIM					
	2	9	9					
20	MSEQ	MGFLAG	ICONSF	MTT	NGAM			

	1	1	0	0	1
22	NCYC	MGLEVG	NEMGL	NITFO	
23	200 MIT1	3 MIT2	0 MIT3	0 MIT4	MIT5
	1	1	1	1	1

This input could be used until satisfactory convergence has been achieved.

Even a partially-converged coarse grid solution is typically a better initial guess than free-stream conditions. Therefore, the second way to set up the mesh sequencing assumes that a set number of iterations at the coarse grid level will be adequate to improve the convergence at the fine grid level. If this is the case, the coarse grid solution is mapped to the fine grid on the first submittal. For example:

#### Line Type

5	• •	DT	IREST	IFLAGTS	FMAX	IUNST	CFLTAU		
-		-1.0	0	0	1.0	0	10.0		
6		NGRID	NPLOT3D	NPRINT	NWREST	ICHK	I2D	NTSTEP	ITA
		1	0	0	0	0	1	1	1
7		NCG	IEM	IADVANCE	IFORCE	IVISC(I)	IVISC(J)	IVISC(K)	
		2	0	0	1	0	0	0	
8		IDIM	JDIM	KDIM					
		2	9	9					
•									
20		MSEQ	MGFLAG	ICONSF	MTT	NGAM			
		2	1	0	0	1			
22				1000					
LL		NCYC	MGLEVG	NEMGL	NITFO				
		200	2	0	0				
22		200	3	0	0				
23		MIT1	MIT2	MIT3	MIT4	MIT5			
		1	1	1	1	1			
		1	1	1	1	1			

When an input like the one above is used, the coarse grid solution is not available for examination before proceeding to the fine grid. In addition, the restart file will contain the solution for the fine grid only. This is important to remember when a grid refinement study is being conducted (in which case the first approach to mesh sequencing would be beneficial). To restart from a run with the above input, use something like:

Line 7	Гуре							
5	DT	IREST	IFLAGTS	FMAX	IUNST	CFLTAU		
	-1.0	0	0	1.0	0	10.0		
6	NGRID	NPLOT3D	NPRINT	NWREST	ICHK	I2D	NTSTEP	ITA
	1	0	0	0	0	1	1	1
7	NCG	IEM	IADVANCE	IFORCE	IVISC(I)	IVISC(J)	IVISC(K)	
	2	0	0	1	0	0	0	
8	IDIM	JDIM	KDIM					
	2	9	9					
20	MSEQ	MGFLAG	ICONSF	MTT	NGAM			
	1	1	0	0	1			
22	NCYC	MGLEVG	NEMGL	NITFO				
	200	3	0	0				
23	MIT1	MIT2	MIT3	MIT4	MIT5			
	1	1	1	1	1			

One last note about mesh sequencing is to emphasize, as with multigrid, the advantage of "good" grid dimensions. See "A Word About Grid Dimensions" on page 127. Planning for mesh sequencing should be made at the grid generation step of the CFD problem.



Although most computations being performed today are for steady-state cases, it appears as though CFD will be used more and more for unsteady, time-accurate cases in the future. Therefore, since this realm is still somewhat in its infancy, this chapter has been written in order to explore in detail what is currently known about the subject, relative to CFL3D's capabilities.

CFL3D has been used extensively for time-accurate computations (dt > 0). See, for example, reference 33. Two types of sub-iterations, called "*t*-TS" and " $\tau$ -TS" are currently implemented in the code and are also described in this reference. This chapter describes the effects of the different types of sub-iterations, as well as the strategy for pursuing time-accurate computations in general.

When performing steady-state computations, the primary numerical accuracy issue about which the user needs to be concerned is that of spatial accuracy. Generally, the user runs a problem on a series of successively finer grids. As long as the solutions are fully converged on each grid, the user can get a clear picture of the accuracy of a solution on a given grid. In fact, the user can determine the numerical global order of accuracy by using a series of at least three grids *from the same family* and plotting some global quantity of interest as a function of a measure of the average grid spacing, such as  $1/(\sqrt{number of grid points})$  for 2-d or  $1/(\sqrt[3]{number of grid points})$  for 3-d, on a log-log plot. The slope of the plotted line represents the spatial order of accuracy of the scheme. When the standard  $\kappa = -1/3$  scheme is employed, CFL3D has been demonstrated in the past to be globally approximately second-order accurate for most grids. However, this is problem-dependent and the accuracy can degrade somewhat on grids that are too coarse or on grids with extremely severe stretching.

For time-accurate problems, temporal accuracy becomes an additional numerical accuracy issue of concern for the user. Now, not only does the effect of the grid need to be assessed for each problem, but the effect of the time step as well. Additionally, because CFL3D is an implicit code and employs approximate factorization, linearization and factorization errors are introduced during each time step, which can degrade the accuracy of the time-accurate simulation. (In fact, if no sub-iterations are employed, the best that can possibly be hoped for is first-order temporal accuracy.) This is why sub-iterations are generally recommended for time-accurate computations. Sub-iterations "iterate away" the linearization and factorization errors. The more sub-iterations performed, the more accurate the simulation. But how many sub-iterations are enough? And which type of sub-iteration scheme works the best? Hopefully, this chapter will help to answer these questions.

#### 8.1 General Effects of Numerical Parameters

A graphic representing the general effects of sub-iterations, time step, and grid is shown in Figure 8-1. Here, some "quantity of interest" is shown as a function of time step. For example, the quantity could represent Strouhal number for an unsteady circular-cylinder case. Say that the user ran a series of computations on a given grid, using three subiterations. Each successive computation used a smaller and smaller time step and each was completely converged to periodic quasi-steady-state. If the user plotted a global quantity of interest as a function of time step, a curve like the lower-most curve in the figure might be obtained. (Note that the figure shows the quantity of interest increasing with decreasing time step, but the trend could also be in the opposite direction, depending on the case and the quantity of interest chosen.)

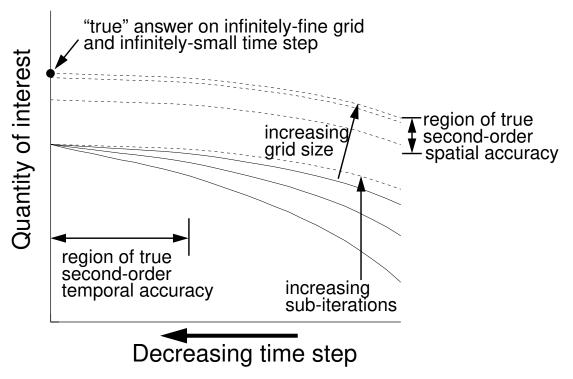


Figure 8-1. Time accuracy trends.

Now say that the user repeated the entire series of computations, except this time using six sub-iterations instead of three. A curve similar to the second curve from the bottom of the figure might be obtained. Finally, with an infinite number of sub-iterations per time step, the user would obtain the dashed curve, which represents the best possible solution *on that grid*. Note that, using a given time step, an increasing number of sub-iterations yields an increasingly better answer, but even the best answer (with infinite sub-iterations) is still in error from the answer using an infinitely-small time step.

The dashed line should behave either first or second order accurate in time (depending on the accuracy input by the user in the input file), but only for sufficiently small time steps. This is what is graphically represented by the left horizontal arrow in the figure. If the time step is too large, the solution may not exhibit the expected temporal accuracy.

It is evident from this figure that, if an extremely small time step is taken, then subiterations are not as beneficial as when larger time steps are taken. The solution is already pretty good. Hence the user needs to make a trade-off between accuracy and efficiency. An extremely small time step can be taken, but at a greater cost, or a larger time step can be taken with some degradation in accuracy. And, sub-iterations may or may not contribute much toward improving accuracy, depending on the time step. Also, at one time step a certain number of sub-iterations may be enough, but at a different time step that same number may be either insufficient or overkill.

The effect of the grid size is also represented in Figure 8-1. On finer and finer grids, the location of the dashed curve will change, approaching the "true" answer on an infinitely-fine grid. (Note that the figure shows the quantity of interest increasing with finer grids, but it could also go the opposite direction, depending on the case and the quantity of interest chosen.) The standard  $\kappa = -1/3$  CFL3D scheme should generally behave spatially second-order accurate on sufficiently fine grids. The filled-in circle in the figure represents the "true" answer on an infinitely fine grid with an infinitely small time step. Obviously, the user would generally like to get as close to this answer as possible, but with a reasonable expenditure of resources.

## 8.2 Effect of Sub-iterations With Time Step and Grid Size

The effect of sub-iterations with time step and grid size is explored in this section for a sample test case of laminar flow over a circular cylinder at Reynolds number 1200 and  $M_{\infty} = 0.2$ . The input file, for a relatively fine O-grid with *t*-TS multigrid sub-iterations, is given here:

cylnew.bin plot3dg.bin cfl3d.out cfl3d.res cfl3d.turres cfl3d.blomax cfl3d.out15 cfl3d.prout cfl3d.out20 ovrlp.bin patch.bin restart.bin							
circ cyl	inder						
XMACH	ALPHA	BETA	REUE,MIL	TINF,DR	IALPH	IHIST	
0.2000	00.000	0.0	0.0012	460.0	0	0	
SREF	CREF	BREF	XMC	YMC	ZMC		
1.00000	1.00000	1.0000	0.00000	0.00	0.00		
DT	IREST	IFLAGTS	FMAX	IUNST	CFLTAU		
+0.1000	0	000	05.0000	0	5.		
NGRID	NPLOT3D	NPRINT	NWREST	ICHK	I2D	NTSTEP	ITA
1	1	1	6100	0	1	0100	+2
NCG	IEM	IADVANCE	IFORCE	IVISC(I)	IVISC(J)	IVISC(K)	

2		0	1	0	1	1	
IDIM 2		KDIM 97					
ILAMLC 0		JLAMLO 0	JLAMHI 0	KLAMLO 0	KLAMHI 0		
INEWG		IS	JS	KS	IE	JE	KE
O IDIAG(I)	IDIAG(J)	0 IDIAG(K)	0 IFLIM(I)	0 IFLIM(J)	0 IFLIM(K)	0	0
1 IFDS(I) 1	IFDS(J)	1 IFDS (K) 1	0 RKAP0(I) 0.3333	0 RKAP0(J) 0.3333	0 RKAP0(K) 0.3333		
GRID		NBCIDIM	NBCJ0	NBCJDIM	NBCK0	NBCKDIM	IOVRLP
1 10: GRID 1	SEGMENT	1 BCTYPE 1001	1 JSTA 0	1 JEND 0	1 KSTA 0	1 KEND 0	0 NDATA 0
IDIM: GRID	SEGMENT	BCTYPE	JSTA	JEND	KSTA	KEND	NDATA
1 J0: GRID 1	SEGMENT	1002 BCTYPE 0	0 ISTA 0	0 IEND 0	0 KSTA 0	0 KEND 0	0 NDATA 0
JDIM: GRID	SEGMENT	BCTYPE	ISTA	IEND	KSTA	KEND	NDATA
1 K0: GRID 1	SEGMENT	0 BCTYPE 2004	0 ISTA 0	0 IEND 0	0 JSTA 0	0 JEND 0	0 NDATA 2
-	TWTYPE	CQ	0	0	Ŭ	0	2
KDIM: GRID		0. BCTYPE	ISTA	IEND	JSTA	JEND	NDATA
1 MSEQ		1003 ICONSF	0 MTT	0 NGAM	0	0	0
1 ISSC		0 EPSSSC(2)	0 EPSSSC(3)	01 ISSR	EPSSSR(1)	EPSSSR(2)	EPSSSR(3)
ISSC 0	EPSSSC(1) 0.3	EPSSSC(2) 0.3	EPSSSC(3) 0.3		EPSSSR(1) 0.3	EPSSSR(2) 0.3	EPSSSR(3) 0.3
ISSC O NCYC 10	EPSSSC(1) 0.3 MGLEVG 03	EPSSSC(2) 0.3 NEMGL 00	EPSSSC(3) 0.3 NITFO 000	ISSR 0	0.3	0.3	0.3
ISSC 0 NCYC	EPSSSC(1) 0.3 MGLEVG 03 MIT2	EPSSSC(2) 0.3 NEMGL	EPSSSC(3) 0.3 NITFO	ISSR			
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The initial study is to investigate the effect of number and type of sub-iterations on the sub-iteration convergence of residual and drag. These levels are printed out automatically to the file cfl3d.subit\_res (unit 23). (See Section 5.2.2.)

<u>Grid</u>	$\Delta t$	Sub-iteration Type
$97 \times 49$	0.02	<i>t</i> -TS, multigrid
$97 \times 49$	0.02	<i>t</i> -TS, no multigrid
$97 \times 49$	0.02	$\tau$ -TS, multigrid
97 × 49	0.02	$\tau$ -TS, no multigrid
97 × 49	0.10	<i>t</i> -TS, multigrid
97 × 49	0.10	<i>t</i> -TS, no multigrid
97 × 49	0.10	$\tau$ -TS, multigrid
97 × 49	0.10	$\tau$ -TS, no multigrid
97 × 49	0.50	<i>t</i> -TS, multigrid
97 × 49	0.50	<i>t</i> -TS, no multigrid
97 × 49	0.50	$\tau$ -TS, multigrid
97 × 49	0.50	$\tau$ -TS, no multigrid
193 × 97	0.10	<i>t</i> -TS, multigrid
193 × 97	0.10	<i>t</i> -TS, no multigrid
193 × 97	0.10	$\tau$ -TS, multigrid
193 × 97	0.10	$\tau$ -TS, no multigrid

For this study, the following runs were performed (the coarser grid consists of every other point from the fine grid):

When multigrid was employed, a 3-level V-cycle was used.

Figure 8-2 shows the residual for density and the drag coefficient as a function of **ncyc** (number of sub-iterations +1) at a time step of 0.02. This is a fairly fine time step, yielding over 1000 steps per period. At this time step, both the *t*-TS and  $\tau$ -TS with multigrid converge the drag with **ncyc** = 4 (3 sub-iterations) and *t*-TS and  $\tau$ -TS without multigrid require about **ncyc** = 6 (5 sub-iterations). The residual also converges quicker with multigrid, as expected. Note that the *t*-TS with multigrid has a slightly better residual convergence rate than  $\tau$ -TS with multigrid at this time step.

At a higher time step of dt = 0.10, the trends in Figure 8-3 are similar: both *t*-TS and  $\tau$ -TS with multigrid converge the quickest, requiring about **ncyc** = 5, while the non-multigrid methods require about **ncyc** = 14-16. At this time step, the residual for  $\tau$ -TS with multigrid converges slightly better than *t*-TS with multigrid. This time step corresponds to a little over 200 steps per period.

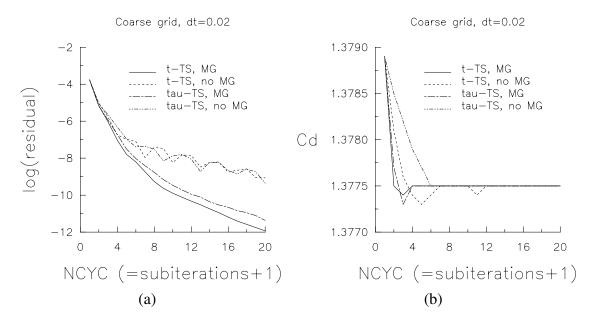


Figure 8-2. Coarse grid residual and drag coefficient histories for a single time step of  $\Delta t = 0.02$ .

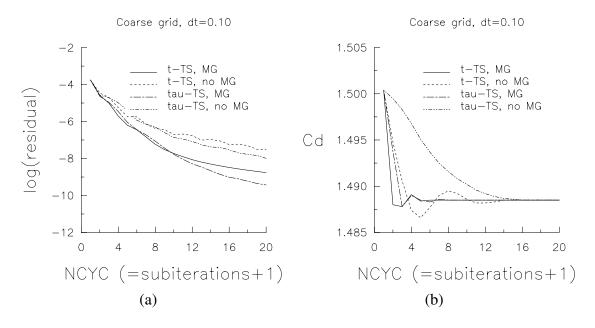


Figure 8-3. Coarse grid residual and drag coefficient histories for a single time step of  $\Delta t = 0.10$ .

At the highest time step of dt = 0.5, Figure 8-4, the multigrid sub-iterations require about ncyc = 12, while the non-multigrid sub-iterations are not fully converged even after ncyc = 40. Also note from Figure 8-4(a), at this time step the residual for *t*-TS with multigrid method now does not converge as well as either  $\tau$ -TS method. This time step corresponds with less than 50 steps per period.

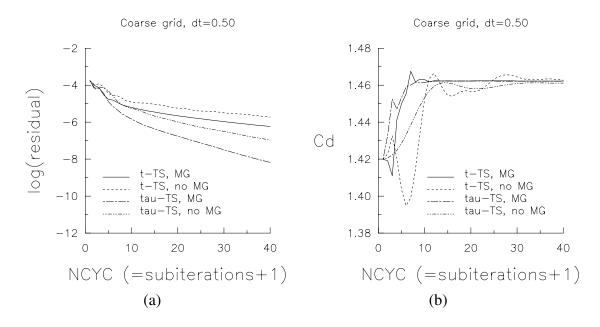


Figure 8-4. Coarse grid residual and drag coefficient histories for a single time step of  $\Delta t = 0.50$ .

Results for the fine  $193 \times 97$  grid with dt = 0.1 are shown in Figure 8-5. Results are qualitatively similar to those in Figure 8-3. However, on this finer grid, more sub-iterations are required to converge the drag: 6-8 iterations for multigrid and well over 20 iterations for non-multigrid.

With a 3-level V-cycle, the multigrid method costs roughly 1.5 times as much as the non-multigrid method. (The  $\tau$ -TS is only marginally more expensive than *t*-TS, so they may be considered essentially equivalent.) Hence, at the smallest time step of **dt** = 0.02, the user is roughly at a break-even point in terms of whether it is more efficient to use multigrid or no multigrid. However, at the larger time steps, using multigrid is clearly beneficial: for example, at **dt** = 0.1 (around 200 steps per cycle), multigrid converges the drag in approximately 0.36 the number of sub-iterations on the coarse grid at 1.4 times the cost; this means a savings of almost 50%! The savings is even greater on the fine grid. Due to this substantial savings, the remaining results will include only cases utilizing multigrid.

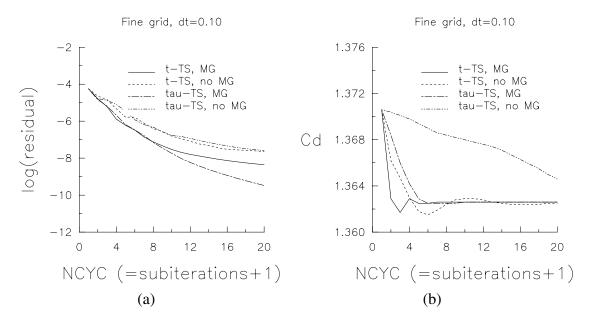


Figure 8-5. Fine grid residual and drag coefficient histories for a single time step of  $\Delta t = 0.10$ .

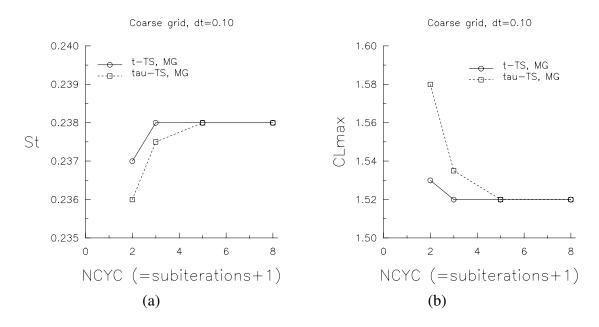


Figure 8-6. Coarse grid Strouhal number and lift coefficient histories for fully periodic solutions using  $\Delta t = 0.10$ .

Figure 8-2 through Figure 8-5 give a feel for the effect of sub-iterations on the convergence to the next physical time step *for one iteration*, but what is the effect of the number of sub-iterations on *global* quantities, over a long period of time? Figure 8-6 shows Strouhal number and maximum lift coefficient (in absolute value) as a function of **ncyc** for **dt** = 0.10 on the coarse grid, where the solution is obtained using the given value of **ncyc** over a long time (until periodic quasi-steady-state is reached). If at least 4 sub-iterations are run for this case (**ncyc** = 5), both *t*-TS and  $\tau$ -TS with multigrid converge to the same result. This is consistent with the results for maximum lift coefficient shown in Figure 8-3(b). However, if less than this number of sub-iterations is run, then the *t*-TS method appears to give the better result.

For the higher time step of dt = 0.5, results are shown in Figure 8-7. If **ncyc** is less than about 7 for this time step, both *t*-TS and  $\tau$ -TS sub-iterations with multigrid yield non-physical solutions (not shown in the figures). For example, *t*-TS yields a highly non-regular lift cycle, while  $\tau$ -TS yields a regular lift cycle with non-zero mean. At least **ncyc** = 15-20 is required to converge the sub-iterative schemes sufficiently at this time step. This is roughly consistent with the results in Figure 8-4(b). If less than this number of sub-iterations is used, then the  $\tau$ -TS method gives the better result. This is the opposite result from that given above for **dt** = 0.1, but it is consistent with the trend seen in the residual plots of Figure 8-2(a), Figure 8-3(a), and Figure 8-4(a). In other words, it appears that the *t*-TS method may require less sub-iterations at low time steps, while  $\tau$ -TS requires less sub-iterations at higher time steps.

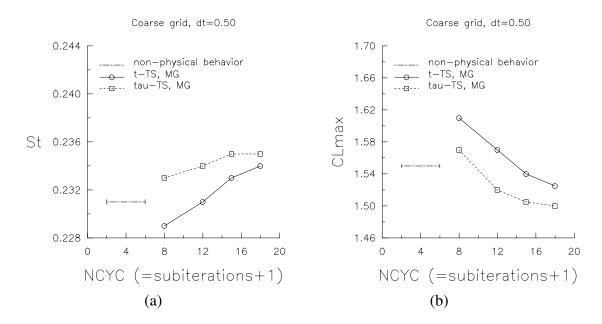


Figure 8-7. Coarse grid Strouhal number and lift coefficient histories for fully periodic solutions using  $\Delta t = 0.50$ .

Other time-accurate circular cylinder cases with turbulence model(s) employed (not shown), have revealed similar trends to the study shown here. One point worth mentioning in connection with these other cases is that, in some instances, the *t*-TS method with multigrid has been seen to either diverge or else give nonphysical answers, *regardless of the number of sub-iterations taken*, when the time step is too large. Unfortunately, determining what time step is too large remains elusive at this point. Trial and error seems to be the only way to determine it for such cases. Therefore, since the  $\tau$ -TS method has not exhibited such errant behavior, it appears that, as a general rule, the safest bet is to go with the  $\tau$ -TS method rather than *t*-TS.

From this study, combined with experience running the CFL3D code for time-accurate cases, the following conclusions are made:

- 1. In general, it is recommended that the user employ multigrid when using sub-iterations.
- 2. The larger the time step (the less steps per period), the more sub-iterations are required to converge the sub-iterative scheme.
- 3. The larger the grid, the more sub-iterations are required to converge the sub-iterative scheme.
- 4. *t*-TS and  $\tau$ -TS are roughly equivalent in their ability to converge a quantity like "drag". However, for lowering residual, *t*-TS is slightly more efficient than  $\tau$ -TS at small time steps, while the reverse is true at higher time steps.
- 5. *t*-TS appears to require slightly less sub-iterations at low time steps, while  $\tau$ -TS appears to require slightly less sub-iterations at higher time steps.
- 6. Since it is not possible to know in advance what time step is "low" and what is "high" (in connection with conclusions 4. and 5.) and since *t*-TS has been known to *not converge* regardless of the number of sub-iterations for some cases when the time step is too high, it is recommended that  $\tau$ -TS (with multigrid) be used in practice as a general rule.

Also, the following recommendation is made. When performing time-accurate computations, *always* monitor the cfl3d.subit\_res file (unit 23) in order to insure that the subiterative scheme is converging sufficiently. Additionally, it is recommended that the user perform a time step study (vary **dt**), along with the usual grid density study, to determine the solution's sensitivity to time step.

## 8.3 Convergence Criterion for Sub-iterations

The sub-iteration equation (Equation (B-7), with  $\phi' = 0$ ) is

$$\begin{bmatrix} \left(\frac{1}{J\Delta\tau} + \frac{1+\phi}{J\Delta t}\right)I + \delta_{\xi}\mathbf{A} + \delta_{\eta}\mathbf{B} + \delta_{\zeta}\mathbf{C}\end{bmatrix}\Delta\mathbf{Q}^{m} = \frac{\phi\Delta\mathbf{Q}^{n-1}}{J\Delta t} - \frac{(1+\phi)(\mathbf{Q}^{m} - \mathbf{Q}^{n})}{J\Delta t} + R(\mathbf{Q}^{m})$$
(8-1)

for  $m \to \infty$ ,  $\mathbf{Q}^m \to \mathbf{Q}^{n+1}$ , and the left-hand side, which is the sub-iteration residual,  $R_{subit}(\mathbf{Q}^m)$ , approaches 0. Write the sub-iteration residual as

$$R_{subit}(\mathbf{Q}^m) = \frac{\phi \Delta \mathbf{Q}^{n-1}}{J\Delta t} - \frac{(1+\phi)(\mathbf{Q}^m - \mathbf{Q}^n)}{J\Delta t} + R(\mathbf{Q}^m)$$
(8-2)

For  $m \to \infty$ ,

$$R_{subit}(\mathbf{Q}^m) \sim \frac{\phi \Delta \mathbf{Q}^{n-1}}{J\Delta t} - \frac{(1+\phi)(\mathbf{Q}^{n+1} - \mathbf{Q}^n)}{J\Delta t} + R(\mathbf{Q}^m)$$
(8-3)

where  $\frac{\phi \Delta \mathbf{Q}^{n-1}}{J\Delta t} - \frac{(1+\phi)(\mathbf{Q}^{n+1}-\mathbf{Q}^n)}{J\Delta t} \Rightarrow \text{discrete} - \frac{\partial \mathbf{Q}}{\partial t}.$ 

Then

$$\frac{\partial \mathbf{Q}^{n}}{\partial t} = R(\mathbf{Q}^{m}) - R_{subit}(\mathbf{Q}^{m})$$
(8-4)

where  $\frac{\partial \mathbf{Q}^n}{\partial t} = R(\mathbf{Q}^m) \Rightarrow$  discrete Navier-Stokes equations. Therefore, to insure that the discrete Navier-Stokes equations are solved accurately,  $R_{subit}(\mathbf{Q}^m) \ll R(\mathbf{Q}^m)$  are needed. These values are calculated in the code.  $R_{subit}(\mathbf{Q}^m)$  (for  $\mathbf{Q}$  = density) is output to cfl3d\_subit.res, while  $R(\mathbf{Q}^m)$  (for  $\mathbf{Q}$  = density) is output to the user-specified file on unit 12 (typically called cfl3d.res in the sample input files).



#### **CHAPTER 9**

Test Cases

CFL3D provides multiple options for solving CFD problems. A variety of 0-equation, 1-equation, and 2-equation turbulence models are available. The code has static or dynamic mesh capabilities. If the grid has multiple zones, there are several choices for communication between the zones which can be used independently or in conjunction with one another. For convergence acceleration, multigrid and mesh sequencing are available.

The test cases described in this chapter provide a sampling of CFL3D's capabilities. After studying the test cases, the user will hopefully be able to choose the best strategy for his or her particular applications. For information on how to obtain the files needed for the test cases see "Acquiring the Code and Example Files" on page 7.

Several two-dimensional test cases discussed in this chapter involve airfoils and flat plates. The use of a single block is exemplified with a RAE 2822 airfoil case. A NACA 0012 airfoil case is used as an example for both grid patching and grid overlapping. Also included is a multielement airfoil case, involving grid overlapping. The flat plate examples include a turbulent flat plate case and a vibrating flat plate case which illustrates the dynamic mesh capabilities of CFL3D. Also included are a multistream nozzle case and a rotor-stator case.

One three-dimensional example is for an axisymmetric bump. By taking advantage of periodicity, it is solved on a grid with only two planes in the circumferential direction. Three of the three-dimensional examples are for wing topologies. A single block case is set up for an F-5 wing. A case solving for the viscous flow over the Onera M-6 wing is also set up using a single block. A delta wing case with laminar flow is also available. Keep in mind that, in order to have cases that are "quick" to run, the three-dimensional grids used in some of these examples are relatively coarse compared to what one *should* use to adequately resolve the flow.

Note: you may see slight differences in your results, due to errors in CFL3D that have been corrected since the plots in this chapter were generated.

# 9.1 Two-dimensional Test Cases

CFL3D solves for the primitive variables at the *cell centers* of a grid. Therefore, for two-dimensional cases, *two* grid planes are needed for one plane of cell-center points to exist. The "2-d direction" is the *i* direction designated by setting **idim** = 2 (and **i2d** = 1). Typically, after a 2-d grid is generated, it is simply duplicated such that identical planes

exist at i = 1 and i = 2 with a constant value in the third direction. For example, if x is the third direction, x is typically set to 0.0 at i = 1 and x might equal 1.0 or -1.0 at i = 2. When setting up the third direction by duplicating the grid plane, keep in mind that the right-hand rule *must* be satisfied. See "The Right-Hand Rule" on page 67. Also note that, while this step *is* doubling the number of grid points, only *one* plane of data is actually computed. Therefore, the number of points in *one* plane of the grid should be used when estimating the time required to run the code.

9.1.1 RAE 2822 Airfoil

This test case solves for the viscous flow over the RAE 2822 airfoil at  $\alpha = 2.72^{\circ}$  with  $M_{\infty} = 0.75$ . These are corrected conditions from Case 10 of Cook et al.<sup>15</sup> The grid consists of a single zone with 24929 points in one plane. Menter's  $k - \omega$  SST model is used to solve the turbulent flow with a Reynolds number of 6.2 million. The memory requirement for this example is 5.7 million words. A typical timing for this case is 382 CPU seconds on a CRAY YMP (NASA LaRC's Sabre as of June 1996). A close-up of the grid near the airfoil is shown in Figure 9-1.

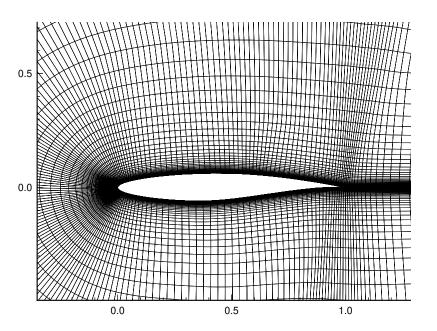


Figure 9-1. Single zone RAE 2822 airfoil grid.

Besides the CFL3D code, the following files are needed to run this test case:

<u>File</u>	Description
rae10.inp	input for CFL3D
rae10.grd	formatted single plane grid
grid2dto3d.f	converter for creating 2 grid planes

The steps for running this case on the YMP are as follows:

<u>Step 1</u>

Compile the grid converter code:

```
cft77 grid2dto3d.f
```

### <u>Step 2</u>

Link the grid converter object file:

segldr -o grid2dto3d grid2dto3d.o

### <u>Step 3</u>

Run the grid converter program (the binary file rae10.bin will be output):

grid2dto3d

In answer to the questions, type:

```
rae10.grd
rae10.bin
2
0
```

### <u>Step 4</u>

Use the makefile to compile, link, and create the executable for the precfl3d code (be sure precfl.h is in the current directory):

```
make -f makeprecfl3d_cray
```

#### <u>Step 5</u>

Run the precf13d code (the cflx.h files will be output):

precfl3d < rae10.inp</pre>

### <u>Step 6</u>

Use the makefile to compile, link, and create the executable for the CFL3D code:

make -f makecfl3d\_cray

## <u>Step 7</u>

Run the CFL3D code:

cfl3d < rae10.inp

The input file for this case is:

FILES: rae10.bin plot3dg.bi cf13d.out cf13d.out cf13d.turr cf13d.blom cf13d.out1 cf13d.out1 cf13d.out2 ovrlp.bin patch.bin restart.bi	n es ax 5 t 0 0						
XMACH	ALPHA	n SST model BETA	REUE,MIL	TINF,DR	IALPH	IHSTRY	
0.7500	02.720	0.0	6.2000	460.0	0	0	
SREF	CREF	BREF	XMC	YMC	ZMC		
1.00000 DT	1.00000 IREST	1.0000 IFLAGTS	0.00000 FMAX	0.00 IUNST	0.00 CFLTAU		
-5.0000	0	000	05.0000	0	10.0		
NGRID	NPLOT3D	NPRINT	NWREST	ICHK	I2D	NTSTEP	ITA
1	1	1	6100	0	1	0001	1
NCG 2	IEM 0	IADVANCE 0	IFORCE 1	IVISC(I) 0	IVISC(J) 0	IVISC(K) 7	
IDIM	JDIM	KDIM	T	0	0	/	
2	257	97					
ILAMLO	ILAMHI	JLAMLO	JLAMHI	KLAMLO	KLAMHI		
1	2	88	159	1	97		
INEWG 0	IGRIDC 0	IS 0	JS 0	KS 0	IE O	JE 0	KE 0
IDIAG(I)	IDIAG(J)	IDIAG (K)	IFLIM(I)	IFLIM(J)	IFLIM(K)	0	0
1	1	1	3	3	3		
IFDS(I)	IFDS(J)	IFDS(K)	RKAPO(I)	RKAP0(J)	RKAPO(K)		
1 GRID	1 NBCIO	1 NBCIDIM	0.3333 NBCJ0	0.3333 NBCJDIM	0.3333 NBCK0	NBCKDIM	IOVRLP
GRID 1	NBC10 1			NECODIM			
IO: GRID		1	1	1	3	1	0
TO: OUTD	SEGMENT	1 BCTYPE	1 JSTA	1 JEND	3 KSTA	1 KEND	0 NDATA
1	SEGMENT 1	BCTYPE 1001	JSTA 0	JEND 0	ksta 0	kend 0	NDATA 0
1 IDIM: GRID	SEGMENT 1 SEGMENT	BCTYPE 1001 BCTYPE	JSTA 0 JSTA	JEND 0 JEND	KSTA 0 KSTA	KEND 0 KEND	NDATA 0 NDATA
1 IDIM: GRID 1	SEGMENT 1 SEGMENT 1	BCTYPE 1001 BCTYPE 1002	JSTA 0 JSTA 0	JEND 0 JEND 0	KSTA 0 KSTA 0	KEND 0 KEND 0	NDATA 0 NDATA 0
1 IDIM: GRID	SEGMENT 1 SEGMENT	BCTYPE 1001 BCTYPE	JSTA 0 JSTA	JEND 0 JEND	KSTA 0 KSTA	KEND 0 KEND	NDATA 0 NDATA
1 IDIM: GRID 1 J0: GRID	SEGMENT 1 SEGMENT 1 SEGMENT	BCTYPE 1001 BCTYPE 1002 BCTYPE	JSTA 0 JSTA 0 ISTA	JEND 0 JEND 0 IEND	KSTA 0 KSTA 0 KSTA	KEND 0 KEND 0 KEND	NDATA 0 NDATA 0 NDATA
1 IDIM: GRID 1 JO: GRID 1 JDIM: GRID 1	SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1	BCTYPE 1001 BCTYPE 1002 BCTYPE 1002 BCTYPE 1002	JSTA 0 JSTA 0 ISTA 0 ISTA 0	JEND 0 JEND 0 IEND 0 IEND 0	KSTA 0 KSTA 0 KSTA 0 KSTA 0	KEND 0 KEND 0 KEND 0 KEND 0	NDATA 0 NDATA 0 NDATA 0 NDATA 0
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1 IDIM: GRID J0: GRID JDIM: GRID 1 K0: GRID 1	SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1	BCTYPE 1001 BCTYPE 1002 BCTYPE 1002 BCTYPE 1002 BCTYPE 0	JSTA 0 JSTA 0 ISTA 0 ISTA 0 ISTA 0	JEND 0 JEND 0 IEND 0 IEND 0 IEND 0 IEND	KSTA 0 KSTA 0 KSTA 0 KSTA 0 JSTA 1	KEND 0 KEND 0 KEND 0 KEND 0 JEND 41	NDATA 0 NDATA 0 NDATA 0 NDATA 0 NDATA 0
1 IDIM: GRID 1 JO: GRID 1 JDIM: GRID 1 KO: GRID	SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1	BCTYPE 1001 BCTYPE 1002 BCTYPE 1002 BCTYPE 1002 BCTYPE 0 2004	JSTA O JSTA O ISTA O ISTA	JEND 0 JEND 0 IEND 0 IEND 0 IEND	KSTA 0 KSTA 0 KSTA 0 KSTA 0 JSTA	KEND 0 KEND 0 KEND 0 KEND 0 JEND	NDATA 0 NDATA 0 NDATA 0 NDATA 0 NDATA
1 IDIM: GRID J0: GRID JDIM: GRID 1 K0: GRID 1	SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 2	BCTYPE 1001 BCTYPE 1002 BCTYPE 1002 BCTYPE 0 2004 CQ 0.	JSTA 0 JSTA 0 ISTA 0 ISTA 0 ISTA 0	JEND 0 JEND 0 IEND 0 IEND 0 IEND 0 0	KSTA 0 KSTA 0 KSTA 0 KSTA 1 1 41	KEND 0 KEND 0 KEND 0 KEND 0 JEND 41 217	NDATA 0 NDATA 0 NDATA 0 NDATA 0 2
1 IDIM: GRID 1 J0: GRID 1 JDIM: GRID 1 K0: GRID 1 1	SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 2 TWTYPE 0. 3	BCTYPE 1001 BCTYPE 1002 BCTYPE 1002 BCTYPE 0 2004 CQ 0. 0	JSTA O JSTA O ISTA O ISTA O O O	JEND 0 JEND 0 IEND 0 IEND 0 IEND 0 0	KSTA 0 KSTA 0 KSTA 0 JSTA 1 41 217	KEND 0 KEND 0 KEND 0 JEND 41 217	NDATA 0 NDATA 0 NDATA 0 NDATA 0 2
1 IDIM: GRID 1 JO: GRID 1 JDIM: GRID 1 KO: GRID 1 KDIM: GRID	SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 2 TWTYPE 0. 3 SEGMENT	BCTYPE 1001 BCTYPE 1002 BCTYPE 1002 BCTYPE 0 2004 CQ 0. 0 BCTYPE	JSTA 0 JSTA 0 ISTA 0 ISTA 0 0 ISTA	JEND 0 JEND 0 IEND 0 IEND 0 IEND 0 0 IEND	KSTA 0 KSTA 0 KSTA 0 KSTA 1 41 217 JSTA	KEND 0 KEND 0 KEND 0 JEND 41 217 257 JEND	NDATA 0 NDATA 0 NDATA 0 NDATA 0 2 0 NDATA
1 IDIM: GRID J0: GRID JDIM: GRID K0: GRID 1 KDIM: GRID 1	SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 2 TWTYPE 0. 3 SEGMENT 1	BCTYPE 1001 BCTYPE 1002 BCTYPE 1002 BCTYPE 1002 BCTYPE 0 2004 CQ 0. 0 BCTYPE 1003	JSTA 0 JSTA 0 ISTA 0 ISTA 0 0 ISTA 0 0 ISTA	JEND 0 JEND 0 IEND 0 IEND 0 IEND 0 IEND	KSTA 0 KSTA 0 KSTA 0 JSTA 1 41 217 JSTA 0	KEND 0 KEND 0 KEND 0 JEND 41 217 257 JEND	NDATA 0 NDATA 0 NDATA 0 NDATA 0 2 0 NDATA
1 IDIM: GRID J0: GRID JDIM: GRID K0: GRID 1 KDIM: GRID 1	SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 2 TWTYPE 0. 3 SEGMENT 1 MGFLAG	BCTYPE 1001 BCTYPE 1002 BCTYPE 1002 BCTYPE 0 2004 CQ 0. 0 BCTYPE 1003 ICONSF	JSTA 0 JSTA 0 ISTA 0 ISTA 0 0 ISTA 0 NTT	JEND 0 JEND 0 IEND 0 IEND 0 IEND 0 IEND	KSTA 0 KSTA 0 KSTA 0 JSTA 1 41 217 JSTA 0	KEND 0 KEND 0 KEND 0 JEND 41 217 257 JEND	NDATA 0 NDATA 0 NDATA 0 NDATA 0 2 0 NDATA
1 IDIM: GRID J0: GRID JDIM: GRID K0: GRID 1 KDIM: GRID KDIM: GRID MSEQ 1	SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 2 TWTYPE 0. 3 SEGMENT 1 MGFLAG 1	BCTYPE 1001 BCTYPE 1002 BCTYPE 1002 BCTYPE 1002 BCTYPE 0 2004 CQ 0. 0 BCTYPE 1003 ICONSF 0 EPSSSC(2)	JSTA 0 JSTA 0 ISTA 0 ISTA 0 0 ISTA 0 0 ISTA 0 0 EPSSSC(3)	JEND 0 JEND 0 IEND 0 IEND 0 IEND 0 IEND 0 NGAM 02 ISSR	KSTA 0 KSTA 0 KSTA 0 JSTA 1 41 217 JSTA 0 EPSSSR(1)	KEND 0 KEND 0 KEND 0 JEND 41 217 257 JEND	NDATA 0 NDATA 0 NDATA 0 NDATA 0 2 0 NDATA 0 0 NDATA 0
1 IDIM: GRID J0: GRID JDIM: GRID K0: GRID 1 KDIM: GRID KDIM: GRID MSEQ 1	SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 MGFLAG 1 EPSSSC(1) 0.3	BCTYPE 1001 BCTYPE 1002 BCTYPE 1002 BCTYPE 1002 BCTYPE 0 2004 CQ 0. 0 BCTYPE 1003 ICONSF 0 EPSSSC(2) 0.3	JSTA 0 JSTA 0 ISTA 0 ISTA 0 0 ISTA 0 0 EPSSSC(3) 0.3	JEND 0 JEND 0 IEND 0 IEND 0 IEND 0 IEND 0 NGAM 02	KSTA 0 KSTA 0 KSTA 0 KSTA 0 JSTA 1 41 217 JSTA 0 EPSSSR(1)	KEND 0 KEND 0 KEND 0 JEND 41 217 257 JEND 0	NDATA 0 NDATA 0 NDATA 0 NDATA 0 2 0 NDATA 0 2 0 NDATA 0 2 0 NDATA 0 2 0 NDATA 0 0 NDATA

500 MIT1 01 1-1 BLOCK NBLI 1	03 MIT2 01 ING DATA:	00 MIT3 01		000 MIT4 01	MIT5 01	Ν	4IT6 1	MIT7 1	MIT8 1
NUMBER GR	•	ISTA	JSTA	KSTA	IEND	JEND	KEND	ISVA1	ISVA2
1	-	1	1	1	2	41	1	1	2
NUMBER GR	ID : 1	ISTA 1	JSTA	KSTA 1	IEND 2	JEND 217	KEND 1	ISVA1 1	ISVA2 2
PATCH SURF NINTER 0	-	Ť	237	Ĩ	2	217	Ť	Ţ	2
PLOT3D OUT	PUT:								
GRID IPTY	PE ISTART	IEND		JSTART		0 1110	KSTART	112112	KINC
1	0 1	01	1	01	999	1	1	999	1
MOVIE 0 PRINT OUT:									
GRID IPTY	PE ISTART	IEND	IINC	JSTART	JEND	JINC	KSTART	KEND	KINC
1	0 1	01	1	41	217	1	1	1	1
CONTROL SU NCS 0	RFACE:								
GRID ISTA	RT IEND	JSTART	JEI	ND KSI	TART KI	END IV	VALL IN	ORM	

After running this test case a result such as that shown in Figure 9-2 should be obtained. In the figure, surface pressure coefficients are plotted along with experimental data for this case. The computational surface pressures can be obtained from file cfl3d.prout. Experimental surface pressure coefficients from Cook et. al<sup>15</sup> are included with this test case for comparison purposes. The file is called rae10.cpexp. The residual plots shown in Figure 9-3 should also be duplicated. These convergence histories can be found in cfl3d.res.

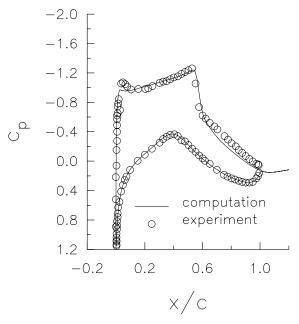
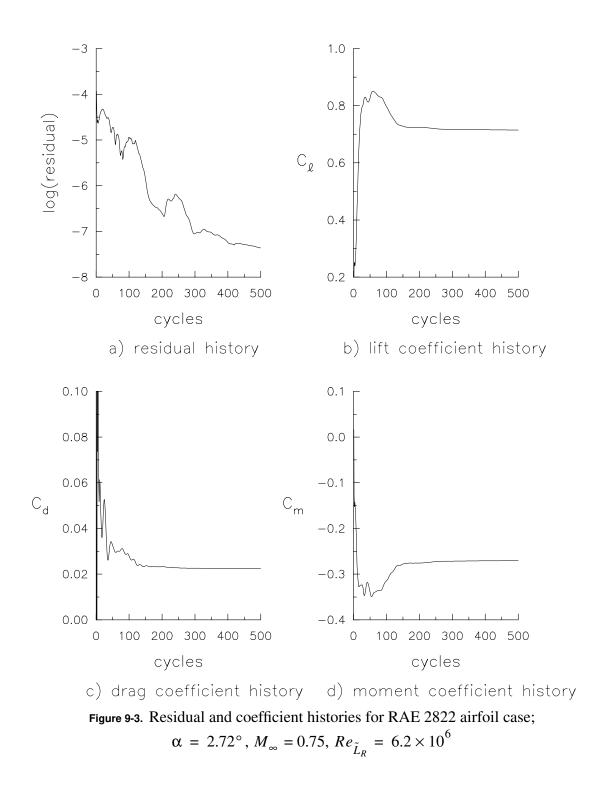


Figure 9-2. Surface pressure coefficients for RAE 2822 airfoil;  $\alpha = 2.72^{\circ}, M_{\infty} = 0.75, Re_{\tilde{L}_R} = 6.2 \times 10^6.$ 



### 9.1.2 NACA 0012 Airfoil with Overlapped Grids

This test case solves for the inviscid flow over the NACA 0012 airfoil at  $\alpha = 5^{\circ}$  with  $M_{\infty} = 0.2$ . The grid has a total of 4850 points on two grid zones which communicate with one another through overlapped grid stencils. Therefore, the MaGGiE code is used in addition to CFL3D. The memory requirement for this case is 1.8 million words. A typical timing for this case is 43 CPU seconds on a CRAY YMP (NASA LaRC's Sabre as of June 1996). A close-up of the grid near the airfoil is shown in Figure 9-4.

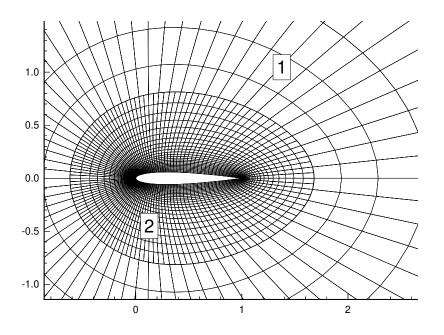


Figure 9-4. Two-zone overlapped grid system for NACA 0012 airfoil.

Besides the CFL3D and MaGGiE codes the following files are needed to run this test case:

<u>File</u>	Description
0012x.inp	input for CFL3D
0012x.fmt	formatted grid in PLOT3D format
fmttobin_p3d.f	grid converter
mag1.h	parameters for MaGGiE makefile
maggie.inp	input for MaGGiE

The steps for running this case on the YMP are as follows:

## <u>Step 1</u>

Compile the grid converter code:

cft77 fmttobin\_p3d.f

## Step 2

Link the grid converter object file:

```
segldr -o fmttobin_p3d fmttobin_p3d.o
```

## <u>Step 3</u>

Run the grid converter program (the binary file 0012x.bin will be output):

```
fmttobin_p3d
```

## Step 4

Use the makefile to compile, link, and create the executable for the MaGGiE code (be sure mag1.h is in the current directory):

make -f makemaggie\_cray

# <u>Step 5</u>

Run the MaGGiE code (the file ovrlp.bin will be output):

```
maggie < maggie.inp</pre>
```

## Step 6

Use the makefile to compile, link, and create the executable for the precfl3d code (be sure precfl.h is in the current directory):

```
make -f makeprecfl3d_cray
```

# <u>Step 7</u>

Run the precfl3d code (the cflx.h files will be output):

```
precfl3d < 0012x.inp</pre>
```

## <u>Step 8</u>

Use the makefile to compile, link, and create the executable for the CFL3D code:

```
make -f makecfl3d_cray
```

## <u>Step 9</u>

Run the CFL3D code:

### cfl3d < 0012x.inp

# The input file for this case is:

<pre>I/O FILES 0012x.bin plot3dg.bir plot3dq.bir cf13d.out cf13d.res cf13d.turre cf13d.bloma cf13d.out15 cf13d.out20 ovrlp.bin patch.bin restart.bir</pre>	n es ax 5 2						
2-block ( XMACH	012 airfoi) ALPHA	l as simpl BETA	e chimera REUE,MIL	test TINF,DR	IALPH	IHSTRY	
.200	5.000	0.0	0.0	520.0	1	0	
SREF	CREF	BREF	XMC 0.25000	YMC	ZMC		
1.00000 DT	1.00000 IREST	1.0000 IFLAGTS	0.25000 FMAX	0.00 IUNST	0.00 CFLTAU		
-5.00	0	000	1.00	0	10.		
NGRID -2	NPLOT3D 2	NPRINT 0	NWREST 100	ICHK 0	I2D 1	NTSTEP 1	ITA 1
NCG	IEM	IADVANCE	IFORCE	IVISC(I)	IVISC(J)	IVISC(K)	
2 2	0 0	0 0	0 1	0	0	0	
IDIM	JDIM	KDIM	±	Ũ	Ũ	Ũ	
002 002	65 129	25 25					
ILAMLO	ILAMHI	JLAMLO	JLAMHI	KLAMLO	KLAMHI		
00	00	000	000	0	0000		
00 INEWG	00 IGRIDC	000 IS	000 JS	0 KS	0000 IE	JE	KE
0	0	0	0	0	0	0	0
0 IDIAG(I)	0 IDIAG(J)	0 IDIAG(K)	0 IFLIM(I)	0 IFLIM(J)	0 IFLIM(K)	0	0
1DIAG(1)	1DIAG(0)	1D1AG(R)	0	0	0		
1	1	1	0	0	0		
				DUUDO(T)			
IFDS(I) 1	IFDS(J)	IFDS (K) 1	RKAP0(I)	RKAP0(J)	RKAP0(K)		
1	IFDS(J) 1 1	IFDS(K) 1 1	.3333	.3333	.3333		
1 1 GRID	IFDS(J) 1 1 NBCIO	1 1 NBCIDIM	.3333 .3333 NBCJ0	.3333 .3333 NBCJDIM	.3333 .3333 NBCK0	NBCKDIM	IOVRLP
1	IFDS(J) 1 1	1	.3333	.3333	.3333	NBCKDIM 1 1	IOVRLP 1 1
1 GRID 1 2 IO: GRID	IFDS(J) 1 NBCIO 1 SEGMENT	1 NBCIDIM 1 BCTYPE	.3333 .3333 NBCJ0 1 1 JSTA	.3333 .3333 NBCJDIM 1 1 JEND	.3333 .3333 NBCK0 1 1 KSTA	1 1 KEND	1 1 NDATA
1 GRID 1 2 I0: GRID 1	IFDS(J) 1 NBCI0 1 SEGMENT 1	1 NBCIDIM 1 BCTYPE 1002	.3333 .3333 NBCJ0 1 1 JSTA 0	.3333 .3333 NBCJDIM 1 1 JEND 0	.3333 .3333 NBCK0 1 1	1 1	1 1 NDATA 0
1 GRID 1 10: GRID 1 2 IDIM: GRID	IFDS(J) 1 NBCIO 1 SEGMENT 1 SEGMENT	1 1 NBCIDIM 1 BCTYPE 1002 1002 BCTYPE	.3333 .3333 NBCJ0 1 JSTA 0 0 JSTA	.3333 .3333 NBCJDIM 1 JEND 0 0 JEND	.3333 .3333 NBCK0 1 1 KSTA 0 0 KSTA	1 KEND 0 0 KEND	1 NDATA 0 0 NDATA
1 GRID 1 2 IO: GRID 1 IDIM: GRID 1	IFDS(J) 1 NBCIO 1 SEGMENT 1 SEGMENT 1	1 1 NBCIDIM 1 BCTYPE 1002 BCTYPE 1002	.3333 .3333 NBCJ0 1 JSTA 0 JSTA 0	.3333 .3333 NBCJDIM 1 JEND 0 0 JEND 0	.3333 .3333 NBCK0 1 KSTA 0 KSTA 0	1 1 KEND 0 0 KEND 0	1 1 NDATA 0 0 NDATA 0
1 GRID 1 2 IO: GRID 1 IDIM: GRID 1	IFDS(J) 1 NBCIO 1 SEGMENT 1 SEGMENT 1 SEGMENT	1 1 NBCIDIM 1 BCTYPE 1002 BCTYPE 1002 1002 BCTYPE	.3333 .3333 NBCJ0 1 JSTA 0 0 JSTA 0 0 ISTA	.3333 .3333 NBCJDIM 1 JEND 0 0 JEND 0 0 JEND	.3333 .3333 NBCK0 1 1 KSTA 0 0 KSTA 0 0 KSTA	1 1 KEND 0 KEND 0 KEND	1 1 NDATA 0 0 NDATA 0 0 NDATA
1 GRID 1 2 IO: GRID 1 2 IDIM: GRID 1 2 JO: GRID 1	IFDS(J) 1 NBCIO 1 SEGMENT 1 SEGMENT 1 SEGMENT	1 1 NBCIDIM 1 BCTYPE 1002 BCTYPE 1002 1002 BCTYPE	.3333 .3333 NBCJ0 1 JSTA 0 0 JSTA 0 0 ISTA	.3333 .3333 NBCJDIM 1 JEND 0 0 JEND 0 0 JEND	.3333 .3333 NBCK0 1 1 KSTA 0 0 KSTA 0 0 KSTA	1 1 KEND 0 KEND 0 KEND	1 1 NDATA 0 NDATA 0 0 NDATA
1 GRID 1 2 I0: GRID 1 IDIM: GRID 1 J0: GRID 1 2 J0: GRID	IFDS(J) 1 NBCIO 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT	1 1 NBCIDIM 1 BCTYPE 1002 1002 BCTYPE 0 0 BCTYPE 0 0 BCTYPE	.3333 .3333 NBCJ0 1 1 JSTA 0 0 JSTA 0 0 ISTA 0 0 ISTA	.3333 .3333 NBCJDIM 1 JEND 0 0 JEND 0 0 IEND 0 0 IEND	.3333 .3333 NBCK0 1 1 KSTA 0 0 KSTA 0 0 KSTA 0 0 KSTA	1 1 KEND 0 KEND 0 KEND 0 0 KEND	1 1 NDATA 0 0 NDATA 0 0 NDATA 0 0 NDATA
1 GRID 1 2 I0: GRID 1 IDIM: GRID 1 J0: GRID 1 2 J0: GRID	IFDS(J) 1 NBCIO 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT	1 1 NBCIDIM 1 BCTYPE 1002 1002 BCTYPE 0 0 BCTYPE 0 0 BCTYPE	.3333 .3333 NBCJ0 1 1 JSTA 0 0 JSTA 0 0 ISTA 0 0 ISTA	.3333 .3333 NBCJDIM 1 JEND 0 0 JEND 0 0 IEND 0 0 IEND	.3333 .3333 NBCK0 1 1 KSTA 0 0 KSTA 0 0 KSTA 0 0 KSTA	1 1 KEND 0 KEND 0 KEND 0 0 KEND	1 1 NDATA 0 0 NDATA 0 0 NDATA 0 0 NDATA
1 GRID 1 2 I0: GRID 1 IDIM: GRID 1 J0: GRID 1 2 J0: GRID	IFDS(J) 1 NBCIO 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT	1 1 NBCIDIM 1 BCTYPE 1002 1002 BCTYPE 0 0 BCTYPE 0 0 BCTYPE	.3333 .3333 NBCJ0 1 1 JSTA 0 0 JSTA 0 0 ISTA 0 0 ISTA	.3333 .3333 NBCJDIM 1 JEND 0 0 JEND 0 0 IEND 0 0 IEND	.3333 .3333 NBCK0 1 1 KSTA 0 0 KSTA 0 0 KSTA 0 0 KSTA	1 1 KEND 0 KEND 0 KEND 0 0 KEND	1 1 NDATA 0 0 NDATA 0 0 NDATA 0 0 NDATA
1 GRID 1 2 I0: GRID 1 IDIM: GRID 1 J0: GRID 1 2 J0: GRID	IFDS(J) 1 NBCIO 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT	1 1 NBCIDIM 1 BCTYPE 1002 1002 BCTYPE 0 0 BCTYPE 0 0 BCTYPE	.3333 .3333 NBCJ0 1 1 JSTA 0 0 JSTA 0 0 ISTA 0 0 ISTA	.3333 .3333 NBCJDIM 1 JEND 0 0 JEND 0 0 IEND 0 0 IEND	.3333 .3333 NBCK0 1 1 KSTA 0 0 KSTA 0 0 KSTA 0 0 KSTA	1 1 KEND 0 KEND 0 KEND 0 0 KEND	1 1 NDATA 0 0 NDATA 0 0 NDATA 0 0 NDATA
1 GRID 1 2 IO: GRID 1 2 IDIM: GRID 1 JO: GRID 1 2 JDIM: GRID 1 2 KO: GRID 1 2	IFDS (J) 1 NBCI0 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 1 1 1 1 1 1 1 1 1 1 1 1	1 1 1 1 1 1 1 1 1 1 1 1 1 1	.3333 .3333 NBCJ0 1 1 JSTA 0 0 JSTA 0 0 ISTA 0 0 ISTA 0 0 ISTA 0 0	.3333 .3333 NBCJDIM 1 JEND 0 0 JEND 0 1END 0 1END 0 1END 0 1END 0 0	.3333 .3333 NBCK0 1 1 KSTA 0 0 KSTA 0 0 KSTA 0 0 KSTA 0 0 KSTA 0 0 XSTA 0 0 0 XSTA 0 0	1 1 KEND 0 KEND 0 KEND 0 KEND 0 0 KEND 0 0 JEND 0 0	1 1 NDATA 0 0 NDATA 0 0 NDATA 0 0 NDATA 0 0 NDATA 0 0
1 GRID 1 2 IO: GRID 1 2 IDIM: GRID 1 JO: GRID 1 2 JDIM: GRID 1 2 KO: GRID 1 2	IFDS (J) 1 NBCI0 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 1 1 1 1 1 1 1 1 1 1 1 1	1 1 1 1 1 1 1 1 1 1 1 1 1 1	.3333 .3333 NBCJ0 1 1 JSTA 0 0 JSTA 0 0 ISTA 0 0 ISTA 0 0 ISTA 0 0	.3333 .3333 NBCJDIM 1 JEND 0 0 JEND 0 1END 0 1END 0 1END 0 1END 0 0	.3333 .3333 NBCK0 1 1 KSTA 0 0 KSTA 0 0 KSTA 0 0 KSTA 0 0 KSTA 0 0 XSTA 0 0 0 XSTA 0 0	1 1 KEND 0 KEND 0 KEND 0 KEND 0 0 KEND 0 0 JEND 0 0	1 1 NDATA 0 0 NDATA 0 0 NDATA 0 0 NDATA 0 0 NDATA 0 0
1 GRID 1 2 IO: GRID 1 2 IDIM: GRID 1 JO: GRID 1 2 JDIM: GRID 1 2 KO: GRID 1 2	IFDS (J) 1 NBCI0 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 1 1 1 1 1 1 1 1 1 1 1 1	1 1 1 1 1 1 1 1 1 1 1 1 1 1	.3333 .3333 NBCJ0 1 1 JSTA 0 0 JSTA 0 0 ISTA 0 0 ISTA 0 0 ISTA 0 0	.3333 .3333 NBCJDIM 1 JEND 0 0 JEND 0 1END 0 1END 0 1END 0 1END 0 0	.3333 .3333 NBCK0 1 1 KSTA 0 0 KSTA 0 0 KSTA 0 0 KSTA 0 0 KSTA 0 0 XSTA 0 0 0 XSTA 0 0	1 1 KEND 0 KEND 0 KEND 0 KEND 0 0 KEND 0 0 JEND 0 0	1 1 NDATA 0 0 NDATA 0 0 NDATA 0 0 NDATA 0 0 NDATA 0 0
1 GRID 1 2 IO: GRID 1 2 IDIM: GRID 1 JO: GRID 1 2 JDIM: GRID 1 2 KO: GRID 1 2	IFDS (J) 1 NBCI0 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 1 1 1 1 1 1 1 1 1 1 1 1	1 1 1 1 1 1 1 1 1 1 1 1 1 1	.3333 .3333 NBCJ0 1 1 JSTA 0 0 JSTA 0 0 ISTA 0 0 ISTA 0 0 ISTA 0 0	.3333 .3333 NBCJDIM 1 JEND 0 0 JEND 0 1END 0 1END 0 1END 0 1END 0 0	.3333 .3333 NBCK0 1 1 KSTA 0 0 KSTA 0 0 KSTA 0 0 KSTA 0 0 KSTA 0 0 XSTA 0 0 0 XSTA 0 0	1 1 KEND 0 KEND 0 KEND 0 KEND 0 0 KEND 0 0 JEND 0 0	1 1 NDATA 0 0 NDATA 0 0 NDATA 0 0 NDATA 0 0 NDATA 0 0
1 GRID 1 2 I0: GRID 1 1 DIM: GRID 1 J0: GRID 1 JDIM: GRID 1 KDIM: GRID 1 KEQ	IFDS (J) 1 NBCI0 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 SEGMENT 1 1 1 1 1 1 1 1 1 1 1 1 1	1 NBCIDIM 1 BCTYPE 1002 1002 BCTYPE 0 0 BCTYPE 0 0 BCTYPE 0 0 BCTYPE 0 1005 BCTYPE 1003 0 ICONSF 1 EPSSC (2)	. 3333 . 3333 NBCJO 1 1 JSTA 0 0 JSTA 0 0 ISTA 0 0 ISTA 0 0 ISTA 0 0 ISTA 0 0 ISTA 0 0 EPSSC (3)	.3333 .3333 NBCJDIM 1 1 JEND 0 0 JEND 0 1END 0 1END 0 1END 0 1END 0 1END 0 0 1END 0 0 1END 0 0 1END 0 0 1END 0 0 1END 0 0 1END 0 0 1END 0 0 1END 0 0 1END 0 0 1END 0 0 1END 0 0 1END 0 0 1END 0 0 1END 0 0 1END 0 0 1END 0 0 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	. 3333 . 3333 NBCK0 1 1 KSTA 0 0 KSTA 0 0 KSTA 0 0 USTA 0 0 JSTA 0 0 JSTA 0 0 USTA	1 1 KEND 0 KEND 0 KEND 0 KEND 0 0 KEND 0 0 JEND 0 0	1 1 NDATA 0 0 NDATA 0 0 NDATA 0 0 NDATA 0 0 NDATA 0 0 NDATA 0 0 NDATA 0 0 NDATA 0 0 NDATA

NCYC 500 MIT1 01 1-1 BLOCI NBLI 2	Ν	LEVG 03 MIT2 01 DATA:	NEMGL 00 MIT3 01	1	11TFO 000 MIT4 01	MIT5 01				
NUMBER GI	RID	:	ISTA	JSTA	KSTA	IEND	JEND	KEND	ISVA1	ISVA2
1	1		1	1	1	2	1	25	1	3
2	2		1	1	1	2	1	25	1	3
	RID	:	ISTA	JSTA	KSTA	IEND	JEND	KEND	ISVA1	ISVA2
1	1			65	1	2	65	25	1	3
2	2		1	129	1	2	129	25	1	3
PATCH SUR	FACE I	DATA:								
NINTER										
0 PLOT3D OU										
GRID IPT			TEND	TINO	JSTART	TEND	TTNO	KSTART	VEND	KINC
GRID IPI. 1	1PE 13 0	1 1	IEND 001	1 INC	051AR1 01	JEND 999	JINC 1	LSIARI 1	KEND 999	1
1 2	0	1	001	1	01	999	1	1	999	1
MOVIE	0	T	001	Ŧ	01	222	T	T	555	T
0										
PRINT OUT	•									
GRID IPT	•	START	IEND	IINC	JSTART	JEND	JINC	KSTART	KEND	KINC
CONTROL SI										
NCS										
0										
GRID IST	ART	IEND	JSTART	JEI	ND KSI	CART KE	ND IV	VALL IN	ORM	

The residual and coefficient histories for this case are plotted in Figure 9-5.

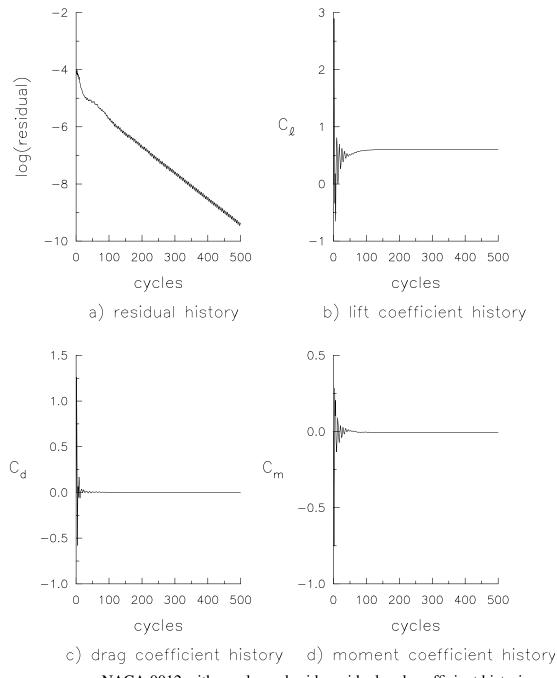


Figure 9-5. NACA 0012 with overlapped grids residual and coefficient histories;  $\alpha = 5^{\circ}, M_{\infty} = 0.2.$ 

#### 9.1.3 NACA 0012 Airfoil with Patched Grids

This test case solves for the inviscid flow over the NACA 0012 airfoil at  $\alpha = 1.25^{\circ}$  with  $M_{\infty} = 0.8$ . The grid has a total of 4949 points in seven zones which communicate with one another utilizing the patching option. Therefore, the ronnie code is used in addition to CFL3D. An advantage of using patched grids is that finer grids can be placed in high gradient regions while relatively coarser grids can be placed elsewhere thus reducing the CPU time and memory needed. In this case, the finest grids are located in the regions where the upper and lower shocks are expected to occur in order to better resolve these flow phenomena. The memory requirement for this example is 1.9 million words. A typical timing is 87 CPU seconds on a CRAY YMP (NASA LaRC's Sabre as of June 1996). A close-up of the grid near the airfoil is shown in Figure 9-6. In the figure, the grids are labelled one through seven and this is the grid order in which the information is set up in the input file.

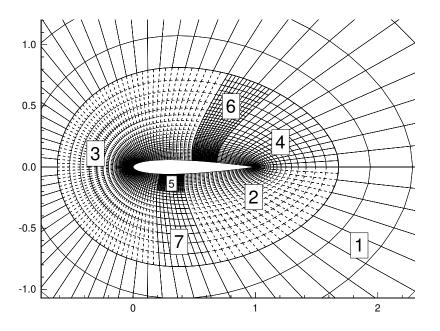


Figure 9-6. Seven-zone patched grid system for NACA 0012 airfoil.

Besides the CFL3D and ronnie codes the following files are needed to run this test case:

<u>File</u>	<b>Description</b>
0012.inp	input for CFL3D
0012.fmt	formatted grid
fmttobin.f	grid converter

<u>File</u>	Description
ron1.h	parameters for ronnie makefile
ronnie.inp	input for ronnie

The steps for running this case on the YMP are as follows:

<u>Step 1</u>

Compile the grid converter code:

cft77 fmttobin.f

#### <u>Step 2</u>

Link the grid converter object file:

segldr -o fmttobin fmttobin.o

#### <u>Step 3</u>

Run the grid converter program (the binary file 0012.bin will be output):

fmttobin

#### <u>Step 4</u>

Use the makefile to compile, link, and create the executable for the ronnie code (be sure ron1.h is in the current directory):

make -f makeronnie\_cray

#### <u>Step 5</u>

Run the ronnie code (the file patch.bin\_0012 will be output):

ronnie < ronnie.inp</pre>

#### <u>Step 6</u>

Use the makefile to compile, link, and create the executable for the precfl3d code (be sure precfl.h is in the current directory):

```
make -f makeprecfl3d_cray
```

### Step 7

Run the precf13d code (the cflx.h files will be output):

precfl3d < 0012.inp</pre>

# <u>Step 8</u>

Use the makefile to compile, link, and create the executable for the CFL3D code:

```
make -f makecfl3d_cray
```

## <u>Step 9</u>

Run the CFL3D code:

cfl3d < 0012.inp

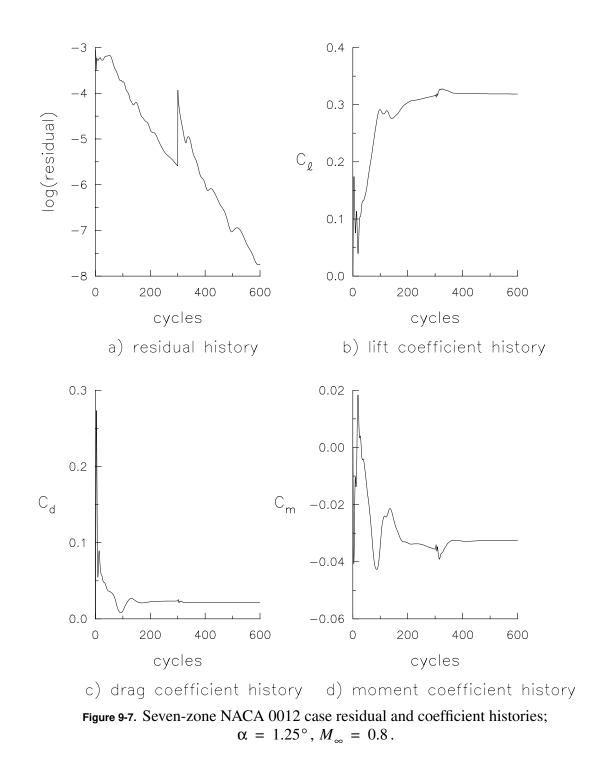
The input file for this case is:

<pre>I/O FILES 0012.bin plot3dg.bin plot3dq.bin cf13d.out cf13d.res cf13d.turre cf13d.bloma cf13d.out15 cf13d.out20 ovrlp.bin patch.bin_0 restart.bin</pre>	s x 012						
input for 7 XMACH	ALPHA	BETA	REUE, MIL	TINF,DR	IALPH	IHSTRY	
0.80 SREF	1.25 CREF	0.0 BREF	0.000 XMC	122.0 YMC	0 ZMC	0	
1.00000	1.00000	1.0000	0.25000	0.00	0.00		
DT	IREST	IFLAGTS	FMAX	IUNST	CFLTAU		
-5.00 NGRID	0 NPLOT3D	000 NPRINT	1.00 NWREST	0 ICHK	10.0 I2D	NTSTEP	ITA
NGRID 7	NPLOISD 7	0	100	0	120	1	1
NCG	IEM	IADVANCE	IFORCE	IVISC(I)	IVISC(J)	IVISC(K)	
1	0 0	0	0 1	0	0 0	0	
1	0	0	1	0	0	0	
1	0	0	1	0	0	0	
1	0	0	1	0	0	0	
1	0	0	1	0	0	0	
IDIM	JDIM	KDIM	0	0	0	0	
2	65	13					
2	23 79	25 25					
2	15	25					
2	17	25					
2	13	49					
2 ILAMLO	9 ILAMHI	13 JLAMLO	JLAMHI	KLAMLO	KLAMHI		
00	00	000	000	0 CLAMLO	0000		
00	00	000	000	0	0000		
00	00	000	000	0	0000		
00000	00 00	000 000	000 000	0	0000 0000		
00	00	000	000	0	0000		
00	00	000	000	0	0000		
INEWG	IGRIDC	IS	JS	KS	IE	JE	KE
0	0 0	0	0 0	0	0 0	0	0 0
0	0	0	0	0	0	0	0
0	0	0	0	0	0	0	0
0	0	0	0	0	0	0	0

0	0	0	0	0	0	0	0
0 IDIAG(I)	0 IDIAG(J)	0 IDIAG(K)	0 IFLIM(I)	0 IFLIM(J)	0 IFLIM(K)	0	0
0 0	0 0	0 0	3 3	3 3	3 3		
0	0	0	3	3	3		
0	0 0	0 0	3 3	3 3	3 3		
0	0	0 0	3 3	3 3	3 3		
IFDS(I)	IFDS(J)	IFDS(K)	RKAP0(I)	RKAP0(J)	RKAP0(K)		
0	0	0 0	.3333 .3333	.3333 .3333	.3333 .3333		
0	0	0	.3333	.3333	.3333		
0 0	0 0	0 0	.3333	.3333 .3333	.3333 .3333		
0	0	0 0	.3333 .3333	.3333 .3333	.3333 .3333		
GRID	NBCIO	NBCIDIM	NBCJ0	NBCJDIM	NBCK0	NBCKDIM	IOVRLP
1 2	1 1	1 1	1 1	1 1	1 1	1 1	0 0
3 4	1	1 1	1 1	1 1	1 1	1 1	0 0
5	1	1	1	1	1	1	0
6 7	1	1 1	1 1	1 1	1 1	1 1	0 0
IO: GRID	SEGMENT	BCTYPE	JSTA	JEND	KSTA	KEND	NDATA
1 2	1 1	1002 1002	0 0	0 0	0 0	0 0	0 0
3 4	1 1	1002 1002	0	0 0	0 0	0 0	0 0
5	1	1002	0	0	0	0	0
6 7	1 1	1002 1002	0 0	0 0	0 0	0 0	0 0
IDIM: GRID 1	SEGMENT 1	BCTYPE 1002	JSTA 0	JEND 0	KSTA 0	KEND 0	NDATA 0
2	1	1002	0	0	0	0	0
3 4	1 1	1002 1002	0 0	0 0	0 0	0 0	0 0
5 6	1 1	1002	0	0	0	0	0
6 7	1	1002 1002	0	0	0	0	0
JO: GRID 1	SEGMENT 1	BCTYPE 0	ISTA 0	IEND 0	ksta 0	KEND 0	NDATA 0
2	1	0	0	0	0	0	0
3 4	1 1	0 0	0 0	0 0	0 0	0 0	0 0
5 6	1	0 0	0 0	0 0	0 0	0 0	0 0
7	1	0	0	0	0	0	0
JDIM: GRID 1	SEGMENT 1	BCTYPE 0	ISTA 0	IEND 0	ksta 0	KEND 0	NDATA 0
2 3	1 1	0 0	0 0	0 0	0 0	0 0	0 0
4	1	0	0	0	0	0	0
5 6	1	0 0	0	0 0	0 0	0 0	0 0
7	1	0	0	0	0	0	0
KO: GRID 1	SEGMENT 1	BCTYPE 0	ISTA 0	IEND 0	JSTA 0	JEND 0	NDATA 0
2 3	1 1	1005 1005	0	0 0	0 0	0 0	0 0
4	1	1005	0	0	0	0	0
5 6	1 1	1005 1005	0 0	0 0	0 0	0 0	0 0
7 KDIM: GRID	1 SEGMENT	0 BCTYPE	0 ISTA	0 IEND	0 JSTA	0 JEND	0 NDATA
1	1	1003	0	0	0	0	0
2 3	1 1	0 0	0 0	0 0	0 0	0 0	0 0
4	- 1 1	0	0	0	0	0	0
5	Ţ	U	U	U	U	U	U

	6	1	0		0	0		0	0		0
	7	1	0		0	0		0	0		0
MS	EQ	MGFLAG	ICONSF		MTT	NGAM					
то	2 ISC EF	1 SSC(1)	1 EPSSC(2)	EDC	0 SC(3)	02 ISSR	EPSSI	D(1) T		EDCCD (	21
15	0	.3	LPSSC(2)	EPS:	.3	1558	EF 221	.3	EPSSR(2)	EPSSR (	.3
NC	YC	MGLEVG	.J NEMGL	r	.J NITFO	0		• 5			• 5
	00	01	00	1	000						
	00	02	00		000						
	T1	MIT2	MIT3		MIT4	MIT5					
	01	01	01		01	01					
	01	01	01		01	01					
1-1 B	LOCKIN	IG DATA:									
NB	LI										
	1										
NUMBER	GRII		ISTA	JSTA	-	IEND	JEND	KENI	-	ISVA2	
1	1	-	1	1	1	2	1	13		3	
NUMBER	GRII		ISTA	JSTA	-	IEND	JEND	KENI	-	ISVA2	
1 DATCI	1 CUDEAC	E DATA:	1	65	1	2	65	13	3 1	3	
NINT		E DAIA:									
	-1										
PLOT3D	-	лт:									
			IEND	IINC	JSTART	JEND	JINC	KSTARI	r kend	KINC	
1	(			1	01	999	1	- 1	L 999	1	
2	(	) 1	001	1	01	999	1	1	L 999	1	
3	(	) 1	001	1	01	999	1	1	L 999	1	
4	(			1	01	999	1	-	L 999	1	
5	(			1	01	999	1	-	L 999	1	
6	(			1	01	999	1	-	L 999	1	
7	(	) 1	001	1	01	999	1	1	L 999	1	
MOVIE											
0	0110										
PRINT		. ISTART	IEND	TINC	JSTART	JEND	TINC	KSTARI	r kend	KINC	
CONTRO		-	TEND	TINC	JSIARI	JEND	JINC	KSIAKI	L KEND	KINC	
NCS	L SURE	ACE:									
0											
-	ISTARI	IEND	JSTART	JEI	ND KST	ART KE	ND I	WALL ]	INORM		

The residual and force coefficient history plots for this case is shown in Figure 9-7. The sharp spike in the residual history plot depicts the iteration at which the grid levels changed for mesh sequencing. These convergence histories can be found in cfl3d.res.



### 9.1.4 Multielement Airfoil with Overlapped Grids

This test case solves for the viscous, turbulent flow over a three-element airfoil with  $M_{\infty} = 0.2$ ,  $\alpha = 8.109^{\circ}$ , and a Reynolds number of 9 million. The Spalart-Allmaras turbulence model is used. The grid, with a total of 59051 points, consists of three zones, one for each element. The grid zones communicate with one another utilizing the grid overlapping option. Therefore, the MaGGiE code is used in addition to CFL3D. The memory requirement for this case is 9.5 million words. A typical timing is 2849 CPU seconds on a CRAY YMP (NASA LaRC's Sabre as of June 1996). A close-up of the grid near the airfoil is shown in Figure 9-8. In the figure, the grids are labelled one through three and this is the grid order in which the information is set up in the input file.

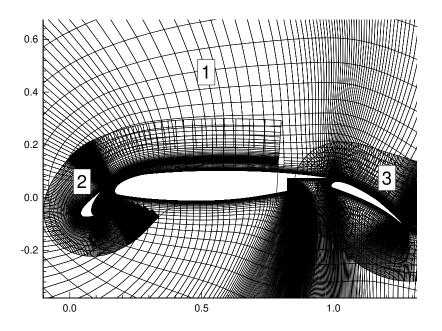


Figure 9-8. Three-zone overlapped grid system for a three-element airfoil.

Besides the CFL3D and MaGGiE codes the following files are needed to run this test case:

<u>File</u>	Description
multi.inp	input for CFL3D
grid.fmt	formatted grid
fmttobin.f	grid converter
mag1.h	parameters for MaGGiE makefile
mag.inp_multi	input for MaGGiE

The steps for running this case on the YMP are as follows:

### <u>Step 1</u>

Compile the grid converter code:

cft77 fmttobin.f

### <u>Step 2</u>

Link the grid converter object file:

segldr -o fmttobin fmttobin.o

### <u>Step 3</u>

Run the grid converter program (the binary file multi.bin will be output):

fmttobin

### <u>Step 4</u>

Use the makefile to compile, link, and create the executable for the MaGGiE code (be sure mag1.h is in the current directory):

make -f makemaggie\_cray

#### <u>Step 5</u>

Run the MaGGiE code (the file ovrlp.bin will be output):

```
maggie < mag.inp_multi</pre>
```

### <u>Step 6</u>

Use the makefile to compile, link, and create the executable for the precfl3d code (be sure precfl.h is in the current directory):

make -f makeprecfl3d\_cray

## <u>Step 7</u>

Run the precf13d code (the cflx.h files will be output):

```
precfl3d < multi.inp</pre>
```

### <u>Step 8</u>

Use the makefile to compile, link, and create the executable for the CFL3D code:

```
make -f makecfl3d_cray
```

# <u>Step 9</u>

Run the CFL3D code:

cfl3d < multi.inp

# The input file for this case is:

multi plot3 plot3 cf13d cf13d cf13d cf13d cf13d cf13d cf13d cf13d	dg.bin dq.bin l.out l.res l.turre l.bloma l.out15 l.prout l.out20 o.bin	s x						
3 е	element XMACH	airfoil - ALPHA	chimera-t BETA	ype grids REUE,MIL	- Spalart- TINF,DR	Allmaras t IALPH	urb model IHSTRY	
	.2000	8.109	0.0	9.0	520.0	0	0	
1	SREF 00000	CREF 1.00000	BREF 1.0000	XMC 0.25000	YMC 0.00	ZMC 0.00		
1.	DT	IREST	IFLAGTS	FMAX	IUNST	CFLTAU		
	-5.00	0	000	1.00	0	10.0		
	NGRID 3	NPLOT3D 3	NPRINT 0	NWREST 100	ICHK 0	12D 1	NTSTEP 1	1 ITA
	NCG	IEM	IADVANCE	IFORCE	IVISC(I)	IVISC(J)	IVISC(K)	T
	2	0	0	001	0	0	5	
	2 2	0 0	0	001 001	0	0	5 5	
	IDIM	JDIM	KDIM	001	0	0	J	
	002	361	65					
	002	369	57					
т	002 LAMLO	297 ILAMHI	49 JLAMLO	JLAMHI	KLAMLO	KLAMHI		
-								
	00	00	000	000	0	0000		
	00	0 0 0 0	000 000	000 000	0 0	0000 0000		
	0 0 0 0	0 0 0 0 0 0	000 000 000	000 000 000	0 0 0	0000 0000 0000	TE	VE
	00	0 0 0 0	000 000	000 000	0 0	0000 0000	JE 0	KE 0
	00 00 INEWG 0 0	00 00 IGRIDC 0 0	000 000 000 IS	000 000 000 JS	0 0 0 KS	0000 0000 0000 IE	0 0	0 0
TDT	00 00 INEWG 0 0	00 00 IGRIDC 0 0 0	000 000 IS 0 0	000 000 JS 0 0	0 0 KS 0 0	0000 0000 IE 0 0	0	0
IDI	00 00 INEWG 0 0 0 CAG(I)	00 00 IGRIDC 0 0 0 IDIAG(J)	000 000 IS 0 0 0 IDIAG(K)	000 000 JS 0 0 0 1FLIM(I)	0 0 KS 0 0 0 1FLIM(J)	0000 0000 IE 0 0 0 1FLIM(K)	0 0	0 0
IDI	00 00 INEWG 0 0	00 00 IGRIDC 0 0 0	000 000 IS 0 0	000 000 JS 0 0	0 0 KS 0 0	0000 0000 IE 0 0	0 0	0 0
	00 00 INEWG 0 0 CAG(I) 1 1	00 00 IGRIDC 0 0 IDIAG(J) 1 1 1	000 000 IS 0 0 IDIAG(K) 1 1	000 000 JS 0 0 IFLIM(I) 3 3 3	0 0 KS 0 0 1FLIM(J) 3 3 3 3	0000 0000 IE 0 0 IFLIM(K) 3 3 3	0 0	0 0
	00 00 INEWG 0 0 CAG(I) 1 1 1 CDS(I)	00 00 IGRIDC 0 0 IDIAG(J) 1 1 1 1 IFDS(J)	000 000 IS 0 0 IDIAG(K) 1 1 1 IFDS(K)	000 000 JS 0 0 1FLIM(1) 3 3 3 RKAP0(1)	0 0 KS 0 0 1FLIM(J) 3 3 3 RKAP0(J)	0000 0000 IE 0 0 IFLIM(K) 3 3 3 RKAP0(K)	0 0	0 0
	00 00 INEWG 0 0 CAG(I) 1 1	00 00 IGRIDC 0 0 IDIAG(J) 1 1 1	000 000 IS 0 0 IDIAG(K) 1 1	000 000 JS 0 0 IFLIM(I) 3 3 RKAP0(I) .3333	0 0 KS 0 0 IFLIM(J) 3 3 3 RKAP0(J) .3333	0000 0000 IE 0 0 0 IFLIM(K) 3 3 RKAP0(K) .3333	0 0	0 0
	00 00 INEWG 0 0 CAG(I) 1 1 TDS(I) 1	00 00 IGRIDC 0 0 IDIAG(J) 1 1 1 IFDS(J) 1	000 000 IS 0 0 IDIAG(K) 1 IFDS(K) 1	000 000 JS 0 0 1FLIM(1) 3 3 3 RKAP0(1)	0 0 KS 0 0 1FLIM(J) 3 3 3 RKAP0(J)	0000 0000 IE 0 0 IFLIM(K) 3 3 3 RKAP0(K)	0 0	0 0
	00 00 INEWG 0 0 CAG(I) 1 1 TDS(I) 1 1 GRID	00 00 IGRIDC 0 0 IDIAG(J) 1 1 IFDS(J) 1 1 NBCI0	000 000 IS 0 0 1DIAG(K) 1 1 IFDS(K) 1 1 NBCIDIM	000 000 JS 0 0 1FLIM(I) 3 3 RKAP0(I) .3333 .3333 .3333 NBCJ0	0 0 KS 0 0 1FLIM(J) 3 3 3 RKAP0(J) .3333 .3333 .3333 .3333 NBCJDIM	0000 0000 IE 0 0 1FLIM(K) 3 3 RKAP0(K) .3333 .3333 .3333 NBCK0	0 0 0 NBCKDIM	0 0 0 IOVRLP
	00 00 INEWG 0 0 CAG(I) 1 1 TDS(I) 1 1 GRID 1	00 00 IGRIDC 0 0 1DIAG(J) 1 1 1 IFDS(J) 1 1 NBCI0 1	000 000 IS 0 0 IDIAG(K) 1 1 IFDS(K) 1 1 NBCIDIM 1	000 000 JS 0 0 0 1FLIM(I) 3 3 3 RKAP0(I) .3333 .3333 .3333 .3333 NBCJ0	0 0 KS 0 0 1FLIM(J) 3 3 3 RKAP0(J) .3333 .3333 .3333 NBCJDIM	0000 0000 IE 0 0 1FLIM(K) 3 3 3 RKAP0(K) .3333 .3333 .3333 .3333 NBCK0 3	0 0 0 NBCKDIM	0 0 0 IOVRLP
	00 00 INEWG 0 0 CAG(I) 1 1 TDS(I) 1 1 GRID	00 00 IGRIDC 0 0 IDIAG(J) 1 1 IFDS(J) 1 1 NBCI0	000 000 IS 0 0 1DIAG(K) 1 1 1 IFDS(K) 1 1 NBCIDIM 1 1	000 000 JS 0 0 0 1FLIM(I) 3 3 3 RKAP0(I) .3333 .3333 .3333 .3333 NBCJ0	0 0 KS 0 0 1FLIM(J) 3 3 3 RKAP0(J) .3333 .3333 .3333 NBCJDIM	0000 0000 IE 0 0 1FLIM(K) 3 3 RKAP0(K) .3333 .3333 .3333 NBCK0	0 0 0 NBCKDIM	0 0 0 IOVRLP
IF	00 00 INEWG 0 0 CAG(I) 1 1 TDS(I) 1 GRID 1 2 3 GRID	00 00 IGRIDC 0 0 1DIAG(J) 1 1 1 IFDS(J) 1 1 NBCI0 1 1 SEGMENT	000 000 IS 0 0 1DIAG(K) 1 1 IFDS(K) 1 NBCIDIM 1 NBCIDIM 1 1 SCTYPE	000 000 JS 0 0 0 1FLIM(I) 3 3 3 RKAP0(I) .3333 .3333 .3333 .3333 NBCJ0 1 1 1 1 JSTA	0 0 KS 0 0 0 1FLIM(J) 3 3 3 RKAP0(J) .3333 .3333 .3333 .3333 NBCJDIM 1 1 1 1 JEND	0000 0000 IE 0 0 0 0 1FLIM(K) .33333 .3333 .3333 .3333 .3333 .3333 .33333 .33333 .33333 .3333 .333333	0 0 0 NBCKDIM 1 1 1 1 KEND	0 0 0 IOVRLP 1 1 1 NDATA
IF	00 00 INEWG 0 0 CAG(I) 1 1 TDS(I) 1 GRID 1 2 3 GRID 1	00 00 IGRIDC 0 0 1DIAG (J) 1 1 1 IFDS (J) 1 1 NBCI0 1 1 SEGMENT 1	000 000 IS 0 0 1DIAG(K) 1 1 IFDS(K) 1 NBCIDIM 1 NBCIDIM 1 1 SCTYPE	000 000 JS 0 0 0 1FLIM(I) 3 3 3 RKAP0(I) .3333 .3333 .3333 .3333 NBCJ0 1 1 1 1 JSTA	0 0 KS 0 0 0 1FLIM(J) 3 3 3 RKAP0(J) .3333 .3333 .3333 .3333 NBCJDIM 1 1 1 1 JEND	0000 0000 IE 0 0 0 0 1FLIM(K) .33333 .3333 .3333 .3333 .3333 .3333 .33333 .33333 .33333 .3333 .333333	0 0 0 NBCKDIM 1 1 1 1 KEND	0 0 0 1 1 1 NDATA 0
IF	00 00 INEWG 0 0 CAG(I) 1 1 TDS(I) 1 GRID 1 2 3 GRID	00 00 IGRIDC 0 0 1DIAG (J) 1 1 1 IFDS (J) 1 1 NBCI0 1 1 SEGMENT 1	000 000 IS 0 0 1DIAG(K) 1 1 IFDS(K) 1 NBCIDIM 1 NBCIDIM 1 1 SCTYPE	000 000 JS 0 0 0 1FLIM(I) 3 3 3 RKAP0(I) .3333 .3333 .3333 .3333 NBCJ0 1 1 1 1 JSTA	0 0 KS 0 0 0 1FLIM(J) 3 3 3 RKAP0(J) .3333 .3333 .3333 .3333 NBCJDIM 1 1 1 1 JEND	0000 0000 IE 0 0 0 0 1FLIM(K) .33333 .3333 .3333 .3333 .3333 .3333 .33333 .33333 .33333 .3333 .333333	0 0 0 NBCKDIM 1 1 1 1 KEND	0 0 0 IOVRLP 1 1 1 NDATA
IF IO:	00 00 INEWG 0 0 CAG(I) 1 1 CDS(I) 1 1 GRID 1 2 3 GRID 1 2 3 GRID 1 2 3 GRID	00 00 IGRIDC 0 0 1DIAG (J) 1 1 1 1FDS (J) 1 1 1 NBCI0 1 1 SEGMENT 1 1 SEGMENT	000 000 IS 0 0 1DIAG(K) 1 1 1 IFDS(K) 1 1 1 NBCIDIM 1 1 NBCIDIM 1 1 BCTYPE 1002 1002 1002 1002 BCTYPE	000 000 JS 0 0 0 1FLIM(I) 3 3 3 RKAP0(I) .3333 .3333 .3333 NBCJO 1 1 1 1 JSTA 0 0 0 JSTA	0 0 KS 0 0 0 1FLIM(J) 3 3 3 RKAP0(J) .3333 .3333 NBCJDIM 1 1 1 1 JEND 0 0 0 0 JEND	0000 0000 IE 0 0 0 1FLIM(K) 3 3 3 RKAP0(K) .3333 .3333 .3333 NBCK0 3 3 3 KSTA 0 0 0 KSTA	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	0 0 0 1 1 1 NDATA 0 0 0 0 NDATA
IF IO:	00 00 INEWG 0 0 CAG(I) 1 1 CDS(I) 1 1 GRID 1 2 3 GRID 1 2 3 GRID 1 2 3 GRID 1	00 00 IGRIDC 0 0 1DIAG (J) 1 1 1 1 1 1 1 1 1 1 1 1 1 1 5 2 5 2 5 2	000 000 IS 0 0 1DIAG(K) 1 1 1 IFDS(K) 1 1 1 NBCIDIM 1 1 NBCIDIM 1 1 BCTYPE 1002 1002 1002 1002 BCTYPE	000 000 JS 0 0 0 1FLIM(I) 3 3 3 RKAP0(I) .3333 .3333 .3333 NBCJO 1 1 1 1 JSTA 0 0 0 JSTA	0 0 KS 0 0 0 1FLIM(J) 3 3 3 RKAP0(J) .3333 .3333 NBCJDIM 1 1 1 1 JEND 0 0 0 0 JEND	0000 0000 IE 0 0 0 1FLIM(K) 3 3 3 RKAP0(K) .3333 .3333 .3333 NBCK0 3 3 3 KSTA 0 0 0 KSTA	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	0 0 0 1 1 1 NDATA 0 0 0 NDATA 0
IF IO:	00 00 INEWG 0 0 CAG(I) 1 1 TDS(I) 1 1 GRID 1 2 3 GRID 1 2 3 GRID 1 2 3 GRID 1 2 3 3 GRID 1 2 3 3 1 2 3 3 1 2 3 3 1 2 3 3 3 1 2 3 3 3 1 2 1 2	00 00 IGRIDC 0 0 1DIAG (J) 1 1 1 1 1 1 1 1 1 1 1 1 1 1 5 2 5 2 5 2	000 000 IS 0 0 1DIAG(K) 1 1 1 IFDS(K) 1 1 1 NBCIDIM 1 1 NBCIDIM 1 1 BCTYPE 1002 1002 1002 1002 BCTYPE	000 000 JS 0 0 0 1FLIM(I) 3 3 3 RKAP0(I) .3333 .3333 .3333 NBCJO 1 1 1 1 JSTA 0 0 0 JSTA	0 0 KS 0 0 0 1FLIM(J) 3 3 3 RKAP0(J) .3333 .3333 NBCJDIM 1 1 1 1 JEND 0 0 0 0 JEND	0000 0000 IE 0 0 0 1FLIM(K) 3 3 3 RKAP0(K) .3333 .3333 .3333 NBCK0 3 3 3 KSTA 0 0 0 KSTA	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	0 0 0 1 1 1 1 1 1 1 0 0 0 0 0 0 0 0 0 0
IF IO: IDIM:	00 00 INEWG 0 CAG(I) 1 1 CDS(I) 1 1 GRID 1 2 3 GRID 1 2 3 GRID 1 2 3 GRID 1	00 00 IGRIDC 0 0 1DIAG (J) 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 5 EGMENT 1 1 1 5 EGMENT 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	000 000 IS 0 0 1DIAG(K) 1 1 1 IFDS(K) 1 1 1 NBCIDIM 1 1 NBCIDIM 1 1 BCTYPE 1002 1002 1002 1002 BCTYPE 1002	000 000 JS 0 0 0 0 1FLIM(I) .3333 .3333 .3333 .3333 .3333 NBCJO 1 1 1 1 1 JJSTA 0 0 0 JJSTA 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	0 0 KS 0 0 0 0 1FLIM(J) 3 3 3 3 RKAP0(J) .3333 .3333 .3333 .3333 NBCJDIM 1 1 1 1 1 1 1 0 0 0 0 JEND 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	0000 0000 IE 0 0 0 0 1FLIM(K) 3 3 3 RKAP0(K) .3333 .3333 .3333 .3333 NBCK0 3 3 3 S KSTA 0 0 0 KSTA 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	0 0 0 1 1 1 1 1 1 NDATA 0 0 0 NDATA 0 0 0 0 0 0 0 0 0 0 0

1	0	0	0	0	0	0
1	0	0	0	0	0	0
SEGMENT	BCTYPE	ISTA	IEND	KSTA	KEND	NDATA
1	1003	0	0	0	0	0
1	0	0	0	0	0	0
1	0	0	0	0	0	0
SEGMENT	BCTYPE	ISTA	IEND	JSTA	JEND	NDATA
1	0	0	0	1	73	0
2	2004	0	0	73	289	2
TWTYPE	CQ					
Ο.	0.					
3	0	0	0	289	361	0
1	0	0	0	1	65	0
2	2004	0	0	65	305	2
TWTYPE	CQ					
0.	0.					
3	0	0	0	305	369	0
1	0	0	0	1	49	0
2	2004	0	0	49	249	2
TWTYPE	CQ					

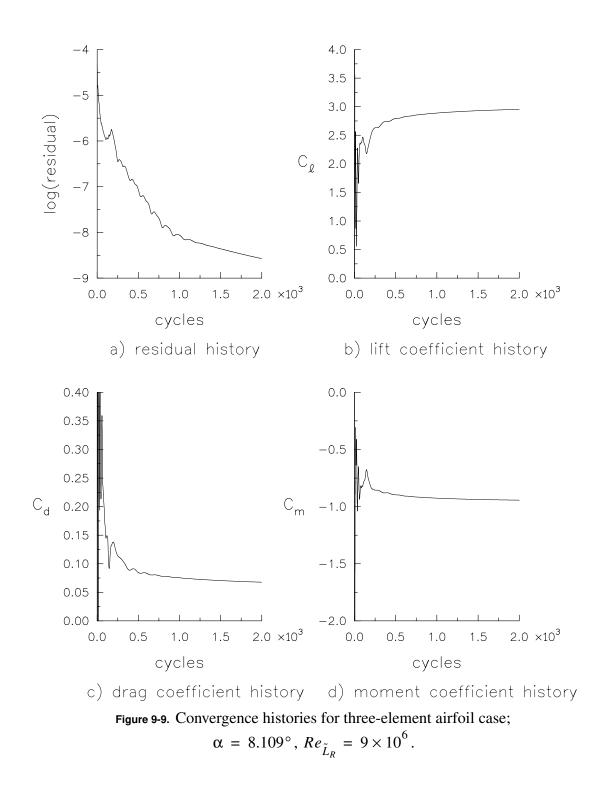
GRID

JDIM: GRID

K0:

	TWTYPE	CQ								
	0.	Ο.								
2		0		0	0		305	369	0	
3	8 1	0		0	0		1	49	0	
3	3 2	2004		0	0		49	249	2	
	TWTYPE	CQ								
	0.	0.								
3	3	0		0	0		249	297	0	
KDIM: GRID	SEGMENT	BCTYPE		ISTA	IEND	, L	JSTA	JEND	NDATA	
1	. 1	1003		0	0		0	0	0	
2	. 1	0		0	0		0	0	0	
3	3 1	0		0	0		0	0	0	
MSEC	) MGFLAG	ICONSF		MTT	NGAM					
1	1	1		0	01					
ISSC	EPSSC(1)	EPSSC(2)	EPSS	SC (3)	ISSR	EPSSE	R(1) 1	EPSSR(2)	EPSSR(3)	
C	.3	.3		.3	0		.3	.3	.3	
NCYC	MGLEVG	NEMGL	1	VITEO						
2000	03	00		000						
MIT1	MIT2	MIT3		MIT4	MIT5					
01	. 01	01		01	01					
1-1 BLC	CKING DATA:									
NBLI										
3	3									
NUMBER	GRID :	ISTA	JSTA	KSTA	IEND	JEND	KENI	D ISVA1	ISVA2	
1	1	1	1	1	2	73		1 1	2	
2	2	1	1	1	2	65		1 1	2	
3	3	1	1	1	2	49		1 1	2	
NUMBER	GRID :	ISTA	JSTA	KSTA	IEND	JEND	KENI	D ISVA1	ISVA2	
1	1	1	361	1	2	289		1 1	2	
2	2	1	369	1	2	305		1 1	2	
3	3	1	297	1	2	249		1 1	2	
PATCH SU	IRFACE DATA:									
NINTER	ξ									
C										
PLOT3D C	UTPUT:									
GRID IP	TYPE ISTART	IEND	IINC	JSTART	JEND	JINC	KSTAR'	T KEND	KINC	
1	1 1	001	1	01	999	1		1 999	1	
2	1 1	001	1	01	999	1		1 999	1	
3	1 1	001	1	01	999	1		1 999	1	
MOVIE										
0										
PRINT OU	JT:									
GRID IP	TYPE ISTART	IEND	IINC	JSTART	JEND	JINC	KSTAR'	T KEND	KINC	
	SURFACE:									
NCS										
0										
GRID IS	TART IEND	JSTART	JEI	ND KST	ART KE	ND IV	VALL	INORM		

After running this test case, the residual and force coefficient convergence histories should look like those in Figure 9-9. These convergence histories can be found in file cfl3d.res. Note the unusually high number of multigrid cycles required to converge this case. While quite large, this is the behavior typically seen (with CFL3D) for multielement airfoil cases, even when one-to-one blocking is employed.



## 9.1.5 Flat Plate

The viscous, turbulent flow with a Reynolds number of 6 million over a flat plate is solved in this test case. The grid consists of a single grid zone with 6305 points. Menter's  $k - \omega$  SST turbulence model is utilized in this example. The memory requirement is 2.3 million words. A typical timing for this case is 157 CPU seconds on a CRAY YMP (NASA LaRC's Sabre as of June 1996). The entire flat plate grid is illustrated in Figure 9-10.

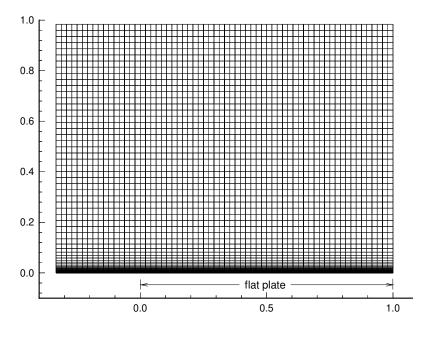


Figure 9-10. Single zone flat plate grid.

Besides the CFL3D code the following files are needed to run this test case:

<u>File</u>	<b>Description</b>
grdflat5.inp	input for CFL3D
grdflat5.grd	formatted grid
grid2dto3d.f	grid converter

The steps for running this case on the YMP are as follows:

#### <u>Step 1</u>

Compile the grid converter code:

```
cft77 grid2dto3d.f
```

# <u>Step 2</u>

Link the grid converter object file:

segldr -o grid2dto3d grid2dto3d.o

## Step 3

Run the grid converter program (the binary file grdflat5.bin will be output):

grid2dto3d

In answer to the questions, type:

```
grdflat5.grd
grdflat5.bin
2
0
```

### Step 4

Use the makefile to compile, link, and create the executable for the precfl3d code (be sure precfl.h is in the current directory):

```
make -f makeprecfl3d_cray
```

### <u>Step 5</u>

Run the precf13d code (the cflx.h files will be output):

```
precfl3d < grdflat5.inp</pre>
```

## Step 6

Use the makefile to compile, link, and create the executable for the CFL3D code:

```
make -f makecfl3d_cray
```

Step 7

Run the CFL3D code:

cfl3d < grdflat5.inp

The input file for this case is:

I/O FILES grdflat5.bin plot3dg.bin cfl3d.out cfl3d.res cfl3d.turres cfl3d.blomax cfl3d.out15 cfl3d.prout

cfl3d.out20 ovrlp.bin patch.bin							
restart.bir turbule	ent flat pl	ate (plate	e from j=17			symmetry)	
XMACH 0.2000	ALPHA 00.000	BETA 0.0	REUE,MIL 06.000	TINF,DR 460.0	IALPH 0	IHSTRY 0	
SREF 1.00000	CREF 1.00000	BREF	XMC 0.00000	YMC 0.00	ZMC 0.00		
DT	IREST	1.0000 IFLAGTS	FMAX	IUNST	CFLTAU		
-5.000 NGRID	1 NPLOT3D	000 NPRINT	05.0000 NWREST	0 ICHK	10. I2D	NTSTEP	ITA
1 NCG	1 IEM	2 IADVANCE	1200 IFORCE	0 IVISC(I)	1 IVISC(J)	1 IVISC(K)	1
2	0	0	001	0	0	12	
IDIM 02	JDIM F 65	DIM 97					
ILAMLO 1	ILAMHI 2	JLAMLO 1	JLAMHI 17	KLAMLO 1	KLAMHI 97		
INEWG 0	IGRIDC 0	IS 0	JS 0	KS 0	IE O	JE 0	KE 0
IDIAG(I)	IDIAG(J)	IDIAG(K)	IFLIM(I)	IFLIM(J)	IFLIM(K)	0	0
1 IFDS(I)	1 IFDS(J)	1 IFDS (K)	0 RKAPO(I)	0 RKAP0(J)	0 RKAP0(K)		
1 GRID	1 NBCIO	1 NBCIDIM	0.3333 NBCJ0	0.3333 NBCJDIM	0.3333 NBCK0	NBCKDIM	IOVRLP
1 IO: GRID	1 SEGMENT	1 BCTYPE	1 JSTA	1 JEND	2 KSTA	1 KEND	0 NDATA
1	1	1001	0	0	0	0	0
IDIM: GRID 1	SEGMENT 1	BCTYPE 1002	JSTA 0	JEND 0	ksta 0	kend 0	NDATA 0
JO: GRID 1	SEGMENT 1	BCTYPE 1008	ISTA 0	IEND 0	KSTA 0	KEND 0	NDATA 0
JDIM: GRID 1	SEGMENT 1	BCTYPE 1002	ISTA 0	IEND 0	KSTA 0	KEND 0	NDATA 0
K0: GRID	SEGMENT	BCTYPE	ISTA	IEND	JSTA	JEND	NDATA
1 1	1 2	1001 2004	0 0	0 0	1 17	17 65	0 2
	TWTYPE 0.	CQ 0.					
KDIM: GRID 1	SEGMENT 1	BCTYPE 1003	ISTA 0	IEND 0	JSTA 0	JEND 0	NDATA 0
MSEQ	MGFLAG	ICONSF	MTT	NGAM	0	0	0
1 ISSC		0 EPSSSC(2)	0 EPSSSC(3)		EPSSSR(1)		
0 NCYC	0.3 MGLEVG	0.3 NEMGL	0.3 NITFO	0	0.3	0.3	0.3
0500 MIT1	03 MIT2	00 MIT3	000 MIT4	MIT5	MIT6	MIT7	MIT8
01	01	01		01	1		1
I-I BLOC NBLI	CKING DATA:						
0 NUMBER 0	GRID :	ISTA	JSTA KST	A IEND	JEND KH	END ISVA1	ISVA2
NUMBER C	GRID : RFACE DATA:	ISTA	JSTA KST	A IEND	JEND KI	END ISVA1	ISVA2
NINTER 0							
PLOT3D OU							
GRID IP1 1	CYPE ISTARI 0 C	IEND 0	IINC JSTAR	T JEND 0 0	JINC KSTA 0	ART KEND 0 0	KINC 0
IMOVIE 0							
PRINT OUT			TING TOWARD				WING
GRID IPI 1	0 C	) <u> </u>	IINC JSTAR 0 0 4	0 0	OINC KSIA O	ARI KEND 1 1	1
1 CONTROL S		) 0	0 4	9 49	1	0 0	0
NCS 0							
-	TART IEND	) JSTART	JEND K	START KI	END IWALL	INORM	

After this test case is run, the residual history, found in file cfl3d.res, should look like that plotted in Figure 9-11. In Figure 9-12, values of  $u^+$  verses  $y^+$  are plotted at two cross sections of the flat plate and compared with theoretical values. The  $u^+$  and  $y^+$  values were extracted from the PLOT3D grid and q files using a postprocessor currently not available for general use.

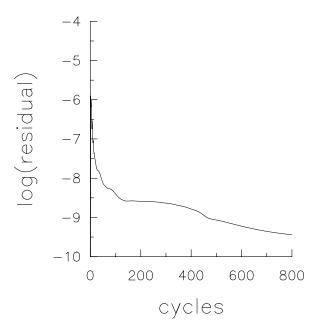


Figure 9-11. Residual history for single grid flat plate case.

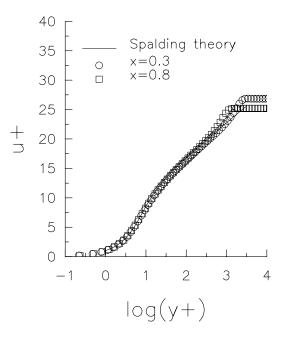


Figure 9-12. Flat plate calculation compared with experiment;  $Re_{\tilde{L}_R} = 6 \times 10^6$ .

Turbulence Model	ivisc	Approximate CPU seconds per cycle	Percent increase over Baldwin- Lomax per cycle
Baldwin-Lomax	2	0.153	0
Baldwin-Lomax with Degani-Schiff	3	0.153	0
Baldwin-Barth	4	0.177	16
Spalart-Allmaras	5	0.182	19
Wilcox $k - \omega$	6	0.186	22
SST $k - \omega$	7	0.196	28
$k - \omega$ EASM Gatski-Speziale (Linear)	8	0.218	43
$k - \varepsilon$ EASM Girimaji (Linear)	9	0.236	54
$k - \varepsilon$ (Abid version)	10	0.197	29
$k - \varepsilon$ EASM Gatski-Speziale (Nonlinear)	11	0.236	54
$k - \omega$ EASM Gatski-Speziale (Nonlinear)	12	0.236	54
$k - \varepsilon$ EASM Girimaji (Nonlinear)	13	0.254	66

This flat plate case has been studied with *all* the turbulence models currently available in CFL3D and a summary of the timings is tabulated below:

This case requires between 800 to 1800 cycles to converge, depending on the turbulence model employed and the convergence criterion chosen. Generally,  $k - \varepsilon$  models take longer than  $k - \omega$  models and two-equation models tend to take longer than one-equation models. (See Appendix H.) Memory requirements also depend on which turbulence model is being used, varying from 2.2 to 2.4 million words.

# 9.1.6 Vibrating Flat Plates

This test case solves for the unsteady, time-accurate inviscid flow through an "infinite" row of vibrating flat plates. The plates, located from x = 0.0 to x = 1.0, are vibrating up and down with a sinusoidal motion. The maximum displacement is h = 0.001 and the nominal distance between the plates is 1.0. The reduced frequency, defined by  $k_r = \tilde{\omega}\tilde{c}/2|\tilde{V}|_{\infty}$ , where  $\omega$  is in radians/second, is 4.0. (Note that the input to CFL3D defines  $\tilde{\omega}$  in cycles/second; therefore, **rfreq** = 0.63662.) The plate vibration generates acoustic waves which propagate both downstream and upstream. The solution invokes periodicity at both the upper and lower boundaries (except between x = 0.0 and x = 1.0), thus mimicking an infinite row of flat plates.

This example also employs a second block downstream of x = 2.0. This block is a "sliding block" that is used to test the effect of a sliding block interface (such as might be used in a rotor-stator computation) on the transmission of acoustic waves. Currently, it is set to translate "up" with speed  $w/a_{\infty} = 1/\pi = 0.31831$ .

This test problem is set up to run for 961 time steps, using three multigrid sub-iterations per time step. The time step is  $\pi/320 = 0.0098175$ . At this time step, 160 time steps yield one complete cycle of plate oscillation, so 961 steps yield six complete cycles of plate oscillations. The code, taking advantage of the periodicity of the solution, "resets" the sliding zone (zone 2) whenever its motion exceeds **dzmax** = 1.0. At the end of time step 961, zone 2 is again physically aligned with zone 1. If the number of steps taken is such that **ntstep**-1 is not evenly divisible by 160, then zone 2 will appear displaced some distance "up" from zone 1 when looking at the solution. While this is not a problem since the solution is periodic, it is easier to visualize the whole flow field when the two zones are aligned.

The grid consists of two grid zones with a total of 5502 points. The memory requirement for this case is 2.1 million words. A typical timing for six cycles of plate oscillation is 422 CPU seconds on a CRAY YMP (NASA LaRC's Sabre as of July 1996).

Besides the CFL3D code the following files are needed to run this test case:

<u>File</u>	<b>Description</b>
vibrate.inp	input for CFL3D
cartesian.f	grid generator

The steps for running this case on the YMP are as follows:

### <u>Step 1</u>

Compile the grid generator code:

```
cft77 cartesian.f
```

# Step 2

Link the grid converter object file:

segldr -o cartesian cartesian.o

# <u>Step 3</u>

Run the grid generator program (the binary file grid.bin will be output):

cartesian

In answer to the questions, type

2 -6,2,0,1 .05,.05 0 2,7,0,1 .05,.05 0

## <u>Step 4</u>

Use the makefile to compile, link, and create the executable for the precfl3d code (be sure precfl.h is in the current directory):

```
make -f makeprecfl3d_cray
```

## <u>Step 5</u>

Run the precf13d code (the cflx.h files will be output):

precfl3d < vibrate.inp</pre>

### <u>Step 6</u>

Use the makefile to compile, link, and create the executable for the CFL3D code:

make -f makecfl3d\_cray

### Step 7

Run the CFL3D code:

cfl3d < vibrate.inp

The input file for this case is:

```
I/O FILES
grid.bin
plot3dg.bin
plot3dq.bin
cfl3d.out
cfl3d.res
cfl3d.turres
cfl3d.blomax
cfl3d.out15
cfl3d.prout
cfl3d.out20
ovrlp.bin
patch.bin
restart.bin
  Flat plate vibrating cascade with sliding interface downstream
     XMACH
               ALPHA
                          BETA REUE,MIL
                                            TINF,DR
                                                        IALPH
                                                                  IHSTRY
      .500
                0.00
                           0.0
                                     1.07
                                              520.0
                                                            0
                                                                       0
                CREF
                          BREF
                                      XMC
                                                          ZMC
      SREF
                                                YMC
   1.00000
             1.00000
                        1.0000
                                  0.25000
                                               0.00
                                                          0.00
                       IFLAGTS
        DT
               IREST
                                     FMAX
                                              IUNST
                                                       CFLTAU
  .0098175
                   0
                            000
                                     1.00
                                                  1
                                                            5.
```

NGRID 2	NPLOT3D 2	NPRINT 2	NWREST 2000	ICHK 0	I2D 1	NTSTEP 961	ITA -2
NCG	IEM	IADVANCE	2000 IFORCE	IVISC(I)	IVISC(J)	IVISC(K)	
2	0	0	JLAMHI	0	0	0	
IDIM	JDIM	KDIM					
2	101	21 21					
ILAMLO	ILAMHI	JLAMLO	JLAMHI	KLAMLO	KLAMHI		
00	00	000	000	0	0000		
INEWG	IGRIDC	IS	000 000 JS 0 0	KS	IE	JE	KE
0	0	0	0	0	0	0	KE 0 0
IDIAG(I)	IDIAG (J)	IDIAG(K)		TF.TTW(1)	TE, TTW (K)		
1	1	1	0 0 RKAP0(I)	0	0		
IFDS(I)	IFDS(J)	IFDS(K)	RKAP0(I)	RKAP0(J)	RKAP0(K)		
1	1	1	.3333	.3333	.3333		
GRID	NBCIO	NBCIDIM	NBCJ0	NBCJDIM	NBCK0	NBCKDIM	IOVRLP
1	1	1	1	1	3	3	0
IO: GRID	SEGMENT	BCTYPE	JSTA	JEND	KSTA	KEND	NDATA
1	1	1001	0	0	0	0	0
IDIM: GRID	SEGMENT	BCTYPE	JSTA	JEND	KSTA	KEND	NDATA
1	1	1002	0	0	0	0	0
J0: GRID	SEGMENT	BCTYPE	ISTA	IEND	KSTA	KEND	NDATA
1	1	1003	0	0	0	0	0
JDIM: GRID	SEGMENT	BCTYPE	ISTA	IEND	KSTA	KEND	NDATA
1	1	0	0	0	0	0	0
K0: GRID	SEGMENT	BCTYPE	ISTA	IEND	JSTA	JEND	NDATA
1	1	2005	1	2	1	121	4
	NBLP 1	0.0	DIHIY 0.	DIHIZ 0.			
1 IFDS(I) 1 GRID 2 I0: GRID 1 2 IDIM: GRID 1 J0: GRID 1 2 J0: GRID 1 2 K0: GRID 1 2 KDIM: GRID 1 1 2 KDIM: GRID 1 1 2 KDIM: GRID 1 1 1 2 KDIM: GRID 1 1 1 1 1 1 1 1 1 1 1 1 1	2	1005	1	2	121	141	0
Ţ	3 NBLP	DTHTX	L DTHTY	2 DTHTZ	141	101	4
0	1	0.0	0.	0.	0	0	4
2	l NBLP	DTHTX	U DTHTY	0 DTHTZ	0	0	4
	2	0.0	0.	0.			
KDIM: GRID 1	SEGMENT 1	BCTYPE 2005	ISTA 1	lend 2	JSTA 1	JEND 121	NDATA 4
	NBLP	DTHTX	DTHTY	DTHTZ			
1	1 2					141	
1	3	2005		2			4
	NBLP 1						
2	1	2005		0	0	0	4
	NBLP 2	DTHTX 0.0		DTHTZ 0.			
MSEQ	MGFLAG	ICONSF	MTT				
1 TSSC			0 EPSSC(3)		EPSSR(1)	EPSSR(2)	EPSSR(3)
0	.3	.3	.3	0		.3	
NCYC 3	MGLEVG 03						
MIT1	MIT2	MTT3	MIT4 01	MIT5			
01 1-1 BLOC	01 KING DATA:	01	01	01			
NBLI							
0 NUMBER G	RTD .	TSTA	JSTA KST	A TEND	JEND KE	ND ISVA1	TSVA2
NUMBER G	RID :	ISTA	JSTA KST			ND ISVA1	
PATCH SUR NINTER	FACE DATA:						

0 PLOT3D OUTPUT: GRID IPTYPE ISTART IEND IINC JSTART JEND JINC KSTART KEND KINC 0 999 01 999 999 1 1 1 1 1 1 999 2 0 1 999 1 01 1 1 999 1 MOVIE Ω PRINT OUT: GRID IPTYPE ISTART IEND IINC JSTART JEND JINC KSTART KEND KINC 1 1 2 1 999 999 1 1 999 1 1 2 2 1 1 1 1 1 1 1 999 1 CONTROL SURFACE: NCS 0 GRID ISTART IEND JSTART JEND KSTART KEND IWALL INORM MOVING GRID DATA - TRANSLATION NTRANS 2 LREF 0.5 ITRANS RFREO XMAG YMAG ZMAG GRID 0. 0.00100 1 2 .31831 Ο. 1 Ο. 0. 0.31831 2 Ο. GRID DXMAX DYMAX DZMAX Ο. Ο. 0.000 1 2 Ο. Ο. 1.000 MOVING GRID DATA - ROTATION NROTAT 0 LREF GRID IROTAT RFREQ THXMAG THYMAG THZMAG XORTG YORTG ZORIG GRID THXMAX THYMAX THZMAX DYNAMIC PATCH INPUT DATA NINTER 2 LIMIT MCETA C - 0TNT IFIT ITMAX MCXIE TORPH TTOSS 1 -1 5 20 0 0 0 0 1 -1 2 5 20 0 0 0 0 1 ТО XIE1 XIE2 ETA1 ETA2 NFB INT 1 122 0 0 0 0 2 FACTK FROM XIE1 XIE2 ETA1 ETA2 FACTJ 221 0 0 0 0 Ο. 0. DX DY DTHETX DTHETY DZDTHET7 Ο. Ο. Ο. Ο. Ο. 0. FROM XIE2 ETA1 XIE1 ETA2 FACTJ FACTK 221 Ο Ο 0 0 0. 0. DX DY DΖ DTHETX DTHETY DTHETZ -1. 0. 0. 0. 0. 0. ETA1 INT ТО XIE1 XIE2 ETA2 NFB 2 221 0 0 0 0 2 FACTJ ETA1 FACTK FROM XIE1 XIE2 ETA2 0 0 0 0 Ο. Ο. 122 DTHETX DTHETY DTHETZ DX DY DZ Ο. Ο. Ο. Ο. Ο. 0. FROM XIE1 XIE2 ETA1 ETA2 FACTJ FACTK 122 0 0 0 0 Ο. Ο. DX DY DΖ DTHETX DTHETY DTHETZ 0. 0. +1. 0. 0. 0.

The resulting residual and lift coefficient histories for this case are shown in Figure 9-13 and Figure 9-14, respectively. The oscillatory nature of the flow is clearly evident. Figure 9-15 shows a profile of the flow as defined by pressure contours.

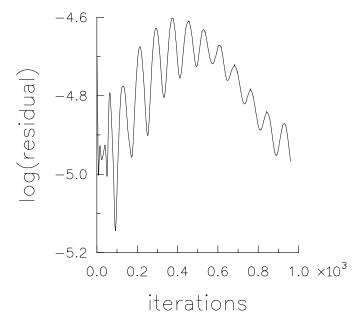


Figure 9-13. Residual history for inviscid flow through vibrating flat plates.

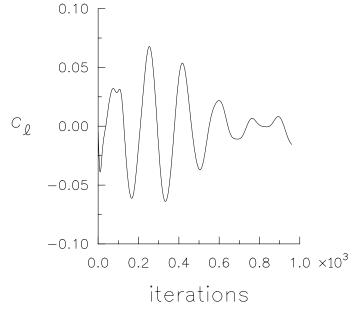


Figure 9-14. Lift coefficient history for inviscid flow through vibrating flat plates;  $M_{\infty} = 0.5$ .

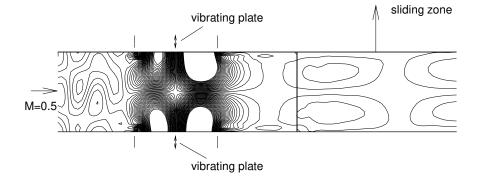


Figure 9-15. Pressure contours for inviscid flow through vibrating flat plates;  $M_{\infty} = 0.5$ .

### 9.1.7 Multistream Nozzle

This case simulates, in two dimensions, the flow though a converging/diverging nozzle with multiple streams. The case is meant to model the exhaust from an engine (with a hot core and cooler outer flow modeled as a "top hat" temperature profile) entering an s-shaped converging/diverging nozzle. Two additional streams are injected downstream of the throat to provide additional cooling of the exhaust.

The grid consists of thirteen patched zones with a total of 15897 points in one plane. The memory requirement for this example is 3.2 million words. A typical timing for this case is 1550 CPU seconds on a CRAY YMP (NASA LaRC's Sabre as of October 1996). A cross-section of the grid is shown in Figure 9-16.

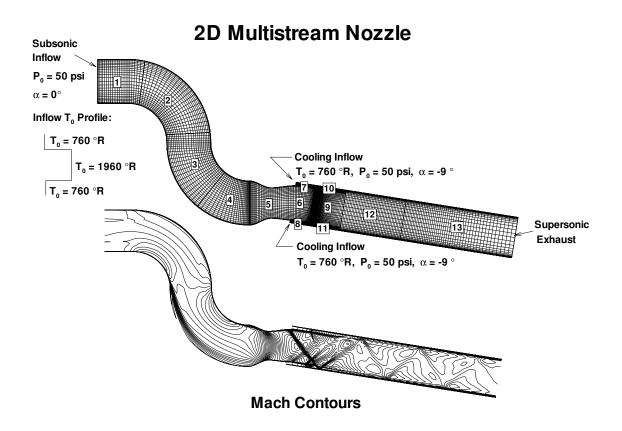


Figure 9-16. Cross-section of multistream nozzle zonal patching and Mach contours.

Boundary conditions with user-defined input, including constant data supplied via the main input deck and variable (point-to-point) data supplied via an auxiliary boundary condition data file are exemplified. Control surfaces are used to monitor mass-flow convergence (here, two control surfaces are used to measure the "difference" between mass in

and mass out). In addition, the case illustrates how one can define reference conditions for CFL3D in the case of a purely internal flow, as discussed below.

#### 9.1.7.1 Nondimensionalization

The conditions provided for this case are that the total pressure of the primary inflow is 50 psi, with a total temperature of 1960°R in the core and 760°R in the outer region. The flow enters the s-duct with zero angle. The secondary cooling flows also have a total pressure of 50 psi and a uniform total temperature of  $760^{\circ}$ R. The cooling flow enters at –9 degrees relative to horizontal. The exhaust from the nozzle system is supersonic. The throat height is 1 foot. The primary inflow height is 1.449 feet; the core flow region (i.e. where the total temperature is  $1960^{\circ}$ R) spans the central 0.769 feet.

Although inflow stagnation conditions would be a natural reference state for this problem, the associated Mach number is zero. Since the viscous terms are scaled with the reference Mach number, another reference state is needed. A second natural reference state for nozzle flows is the sonic point. From isentropic relations, the sonic conditions can be obtained once the inflow total conditions are known. However, for this problem there are two total conditions owing to the "top hat" temperature profile. Thus, to have just one reference state, the total temperature of the inflow is area-averaged and the resulting average total temperature, together with the given total pressure, is used to determine the sonic conditions at the throat.

In what follows, stagnation conditions are denoted by 0, sonic conditions by \*, and dimensional quantities by ~. First, determine the average inflow total temperature:

$$\tilde{T}_0 = \frac{0.769}{1.449(1960)} + \frac{(1.449 - 0.769)}{1.449(760)} = 1397^{\circ} R$$
(9-1)

Next, determine the stagnation density and speed of sound:

$$\tilde{\rho}_0 = \frac{p_0}{\tilde{R}\tilde{T}_0} = \frac{50(144)}{1716(1397)} = 0.00300 \text{ slugs/feet}^3$$
(9-2)

$$\tilde{a}_0 = \sqrt{\gamma \tilde{R} \tilde{T}_0} = \sqrt{1.4(1716)(1397)} = 1832 \text{ feet/second}$$
 (9-3)

where  $\hat{R} = 1716 \text{ feet}^2/(\text{second}^2 - {}^\circ R)$ . From the isentropic relations,

$$\frac{p}{p_0} = 0.528, \frac{T}{T_0} = 0.833, \frac{\rho}{\rho_0} = 0.634, \frac{a}{a_0} = 0.913$$
(9-4)

Thus, the desired reference pressure, temperature, speed of sound and density are

$$\tilde{p} = 50(0.528) = 26.4 \text{ psi}$$
  
 $\tilde{T} = 1397(0.833) = 1164^{\circ}\text{R}$   
 $\tilde{p} = 0.003(0.634) = 0.0019 \text{ slug/feet}^{3}$   
 $\tilde{a} = 1832(0.913) = 1673 \text{ feet/second}$ 
(9-5)

Assuming a molecular viscosity coefficient of  $3.7 \times 10^{-7}$  slugs/(feet-seconds) for a temperature of 520°R, then the power law  $\mu_2/\mu_1 = (T_2/T_1)^{0.76}$  gives

$$\tilde{\mu}^* = 3.7 \times 10^{-7} \left(\frac{1164}{520}\right)^{0.76} = 6.82 \times 10^{-7} \text{ slugs/(feet-seconds)}$$
 (9-6)

The reference Reynolds number based on throat height and the reference sonic values is

$$Re^{*} = \frac{\tilde{\rho}^{*}\tilde{a}^{*}(\text{throat height})}{\tilde{\mu}^{*}} = 0.0019(1673)\frac{1}{6.82 \times 10^{-7}} = 4.66 \times 10^{6}$$
(9-7)

In the grid, the throat height is 12 inches, so the input parameter reue is

**reue** = 
$$\frac{Re^* \times 10^{-6}}{12} = \frac{4.66}{12} = 0.388$$
 (9-8)

Finally, the nondimensional input values for boundary condition type 2003 are determined from the reference sonic conditions (note that in CFL3D parlance, in this problem the \* conditions are the "infinity" conditions):

$$p_{t} = 50 \text{ psi} \rightarrow \frac{p_{t}}{p^{*}} = \frac{p_{t}}{p_{\infty}} = \frac{50}{26.4} = 1.894$$

$$T_{t} = 760^{\circ} \text{R} \rightarrow \frac{T_{t}}{T^{*}} = \frac{T_{t}}{T_{\infty}} = \frac{760}{1164} = 0.653$$

$$T_{t} = 1960^{\circ} \text{R} \rightarrow \frac{T_{t}}{T^{*}} = \frac{T_{t}}{T_{\infty}} = \frac{1960}{1164} = 1.684$$

Boundary condition type 2003 also needs an estimate of the local Mach number. For the primary inflow, the inlet height (area) to throat height (area) is 1.449/1. The isentropic relations give a corresponding Mach number of approximately 0.45. The local Mach number for the cooling inflow cannot be determined a priori; 1.0 is used in the boundary condition. The computations give the cooling inflow Mach number as approximately 0.85; 1.0 is deemed as a sufficiently close estimate since the solution does not change perceptibly if 0.85 is used instead of 1.0.

The auxiliary boundary condition data file provided for this example (inflow.data) contains the data for the primary inflow:

$$M = 0.45$$

$$\frac{P_t}{p_{\infty}} = 1.894$$
(9-10)
$$\frac{T_t}{T_{\infty}} = \frac{0.653}{1.684}$$

in top hat distribution,  $\alpha = 0, \beta = 0$ . The cooling inflow:

$$M = 1.0$$

$$\frac{p_t}{p_{\infty}} = 1.894$$

$$\frac{T_t}{T_{\infty}} = 0.653$$

$$\alpha = -9$$

$$\beta = 0$$
(9-11)

is specified explicitly in the main data file.

### 9.1.7.2 Running CFL3D

Besides the CFL3D and ronnie codes the following files are needed to run this test case:

Description
input for CFL3D
formatted grid
grid converter
auxiliary boundary condition data
parameters for ronnie makefile
input for ronnie

The steps for running this case on the YMP are as follows:

## <u>Step 1</u>

Compile the grid converter code:

```
cft77 formtobin.f
```

## <u>Step 2</u>

Link the grid converter object file:

segldr -o formtobin formtobin.o

### <u>Step 3</u>

Run the grid generator program (the binary file grid\_multistream.bin will be output):

formtobin

### <u>Step 4</u>

Use the makefile to compile, link, and create the executable for the ronnie code (be sure ron1.h is in the current directory):

```
make -f makeronnie_cray
```

### <u>Step 5</u>

Run the ronnie code (the file patch\_multistream.bin will be output):

ronnie < ronnie.inp</pre>

### <u>Step 6</u>

Use the makefile to compile, link, and create the executable for the precfl3d code (be sure precfl.h is in the current directory):

```
make -f makeprecfl3d_cray
```

### <u>Step 7</u>

Run the precf13d code (the cflx.h files will be output):

```
precfl3d < multistream.inp</pre>
```

### Step 8

Use the makefile to compile, link, and create the executable for the CFL3D code:

```
make -f makecfl3d_cray
```

#### <u>Step 9</u>

Run the CFL3D code (be sure the inflow.data file is available and correct for this case):

cfl3d < multistream.inp

The input file for this case is:

```
I/O FILES
grid_multistream.p3d
plot3dg.bin
```

plot3dq.bin cf13d.out cf13d.res cf13d.turres cf13d.blomax cf13d.out15 cf13d.out20 ovrlp.bin patch_multis restart.bin Multistream XMACH 1.0000 DT -1.0000 NGRID -13 NCG 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	tream.bir ALPHA 0.00 CREF 1.00000 IREST 0 NPLOT3D 13 IEM 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	(sonic cond BETA 0.0 BREF 1.0000 IFLAGTS 000 NPRINT 0 IADVANCE 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	itions as REUE,MIL 0.388 XMC 0.25000 FMAX 1.0 NWREST 500 IFORCE 000 000 000 000 000 000 000 000 000 0	reference TINF,DR 1163.0 YMC 0.00 IUNST +1 ICHK 0 IVISC(I) 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	state) IALPH 1 ZMC 0.00 CFL_TAU 5. I2D 1 IVISC(J) 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	IHIST 1 NTSTEP 2 IVISC (K) +7 +7 +7 +7 +7 +7 +7 +7 +7 +7 +7 +7 +7 +	ITA -2
2 2 2 1LAMLO 00 00 00 00 00 00 00 00 00 00 00 00 00	49 49 49 41 49 1LAMHI 00 00 00 00 00 00 00 00 00 00 00 00 00	41 21 61 57 JLAMLO 000 000 000 000 000 000 000 000 000 0	JLAMHI 000 000 000 000 000 000 000 000 000 0	KLAMLO 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	KLAMHI 0000 0000 0000 0000 0000 0000 0000 0	JE 0 0 0 0 0 0 0 0 0 0 0 0 0	KE 0 0 0 0 0 0 0 0 0 0 0 0 0 0

0 0 0 0	0 0 0	0 0 0	0 0 0	0 0 0	0 0 0	0 0 0 0	0 0 0 0
IDIAG(I)	IDIAG(J)	IDIAG(K)	IFLIM(I)	IFLIM(J)	IFLIM(K)		
1	1 1	1 1	3 3	3 3	3 3		
1	1 1	1 1	3 3 3 3	3	3		
1	1	1	3	3	3		
1	1 1	1 1	3 3	3	3		
1	1	1	3	3	3		
1	1 1	1 1	3	3	3 3 3 3 3 3 3 3 3 3 3 3 3		
1	1	1	3 3	3 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3	3		
1	1 1	1 1	3 3	3	3 3		
IFDS(I) 1	IFDS(J) 1	IFDS(K) 1	RKAP0(I) .3333	RKAP0(J) .3333	RKAP0(K) .3333		
1	1	1	.3333	.3333	.3333		
1	1 1	1 1	.3333 .3333	.3333 .3333	.3333 .3333		
1	1	1	.3333	.3333	.3333		
1	1 1	1 1	.3333 .3333	.3333 .3333	.3333 .3333		
1	1	1	.3333	.3333	.3333		
1 1	1 1	1 1	.3333 .3333	.3333 .3333	.3333 .3333		
1	1 1	1 1	.3333 .3333	.3333 .3333	.3333 .3333		
1	1	1	.3333	.3333	.3333		
GRID 1	NBCIO 1	NBCIDIM 1	NBCJ0 1	NBCJDIM 1	NBCKO 1	NBCKDIM 1	IOVRLP 0
2	1	1	1	1	1	1	0
3 4	1 1	1 1	1 1	1 1	1 1	1 1	0 0
5 6	1 1	1 1	1 1	1 1	1 1	1 1	0 0
7	1	1	1	1	1	1	0
8 9	1 1	1 1	1 1	1	1 1	1 1	0 0
10	1	1	1	1	1	1	0
11 12	1 1	1 1	1 1	1 1	1 1	1 1	0 0
13 IO: GRID	1 SEGMENT	1 BCTYPE	1 JSTA	1 JEND	1 KSTA	1 KEND	0 NDATA
1	1	1002	0	0	0	0	0
2 3	1 1	1002 1002	0 0	0 0	0 0	0 0	0 0
4	1	1002	0	0	0	0	0
5 6	1 1	1002 1002	0 0	0 0	0 0	0 0	0 0
7 8	1 1	1002 1002	0 0	0 0	0 0	0 0	0 0
9	1	1002	0	0	0	0	0
10 11	1 1	1002 1002	0	0	0 0	0 0	0 0
12	1	1002	0	0	0	0	0
13 IDIM: GRID	1 SEGMENT	1002 BCTYPE	0 JSTA	0 JEND	0 KSTA	0 KEND	0 NDATA
1 2	1 1	1002 1002	0 0	0 0	0 0	0 0	0 0
3	1	1002	0	0	0	0	0
4 5	1 1	1002 1002	0 0	0 0	0 0	0 0	0 0
6	1	1002	0	0	0	0	0
7 8	1 1	1002 1002	0 0	0 0	0 0	0 0	0 0
9 10	1	1002	0	0	0	0	0
10	1 1	1002 1002	0 0	0 0	0	0 0	0 0

J0:	12 13 GRID	1 1 SEGMENT	1002 1002 BCTYPE	0 0 ISTA	0 0 IEND	0 0 KSTA	0 0 KEND	0 0 NDATA
	1 Pt/Pi ow.data		2003 alpha k	0 Deta	0	0	0	-5
111111	2 3	1 1	0	0 0	0 0	0 0	0 0	0 0
	4	1	0	0	0	0	0	0
	5	1	0	0	0	0	0	0
	6 7	1 1	0 2003	0	0	0	0	0 5
Mac 1.0	ch Pt/2 00 1.	Pinf Tt/Tin 894 0.653	nf alpha 3 -9.	beta 0.	-	-	-	
Mac 1.(	8 ch Pt/1 00 1.		2003 nf alpha 3 -9.	0 beta 0.	0	0	0	5
	9	1	0	0	0	0	0	0
	10	1	0	0	0	0	0	0
	11 12	1 1	0	0 0	0 0	0 0	0 0	0 0
	13	1	Ő	0	0	Ő	Ő	Ő
JDIM	: GRID	SEGMENT	BCTYPE	ISTA	IEND	KSTA	KEND	NDATA
	1 2	1 1	0 0	0 0	0 0	0 0	0 0	0 0
	2 3	1	0	0	0	0	0	0
	4	1	Ő	Ő	Ő	Ő	Ő	Ő
	5	1	0	0	0	0	0	0
	6 7	1 1	0	0 0	0	0 0	0	0
	8	1	0	0	0	0	0	0
	9	1	0	0	0	0	0	0
	10	1 1	0	0 0	0	0 0	0 0	0
	11 12	1	0	0	0	0	0	0
	13	1	1002	Ő	Ő	Ő	Ő	Ő
к0:	GRID	SEGMENT	BCTYPE	ISTA	IEND	JSTA	JEND	NDATA
Τw	1 Cq	1	2004	0	0	0	0	2
0.	0.							
	2	1	2004	0	0	0	0	2
Tw 0.	Cq 0.							
0.	3	1	2004	0	0	0	0	2
Τw	Cq							
0.	0. 4	1	2004	0	0	0	0	2
Τw	4 Cq	1	2004	0	0	0	0	Z
0.	0. 5							
	5	1	2004	0	0	0	0	2
Τw 0.	Cq 0.							
	6	1	2004	0	0	0	0	2
Τw	Cq							
0.	0. 7	1	2004	0	0	0	0	2
Τw	, Cq	T	2004	0	0	0	0	2
0.	Ο.							
	8	1	2004	0	0	0	0	2
Τw Ο.	Cq 0.							
	9	1	0	0	0	0	0	0
	10	1	2004	0	0	0	0	2
Tw	Cq							
0.	0. 11	1	0	0	0	0	0	0
	12	1	2004	Ő	Ő	0	0	2
Tw	Cq							
0.	0. 13	1	2004	0	0	0	0	2
Τw	Cq	+	2001	0	0	0	U U	2
0.	0.							

KDIM:	: GRID 1	SEGMENT 1	BCTYPE 2004	ISTA 0	IEND 0	JSTA 0	JEND 0	NDATA 2
Tw O.	Cq 0.							
Tw O.	2 Cq 0.	1	2004	0	0	0	0	2
Tw 0.	3 Cq 0.	1	2004	0	0	0	0	2
Tw	4 Cq	1	2004	0	0	0	0	2
0. Tw	0. 5 Cq	1	2004	0	0	0	0	2
0. Tw	0. 6 Cq	1	2004	0	0	0	0	2
0.	0. 7	1	2004	0	0	0	0	2
Τw 0.	Cq 0. 8	1	2004	0	0	0	0	2
Tw 0.	Cq 0. 9	1	0	0	0	0	0	0
<b>T</b>	10 11	1 1	0	0	0	0	0	0 2
Τw 0.	Cq 0. 12	1	2004	0	0	0	0	2
Tw 0.	Cq 0. 13	1	2004	0	0	0	0	2
Tw O.	Cq 0. MSEQ	MGFLAG	ICONSF	MTT	NGAM			
	1 ISSC	1 EPSSC(1)	1 EPSSC(2)	0 EPSSC(3)	02 ISSR	EPSSR(1)	EPSSR(2)	EPSSR(3)
	0 NCYC 3000	.3 MGLEVG 02	.3 NEMGL 00	.3 NITFO 0	0	.3	.3	.3
1-	MIT1 01 -1 BLOC	MIT2 01 KING DATA	MIT3 01	MIT4 01	MIT5 01			
	NBLI 0			70 <b>0</b> 3 <b>1</b> /0 <b>0</b> 3	TEND			
NU PAT	CCH SUF	GRID : GRID : RFACE DATA	ISTA	JSTA KSTA JSTA KSTA	IEND IEND	JEND KENI JEND KENI		ISVA2 ISVA2
	-1 C3D OUI		iend iinc	jsta jend	jing k	ksta kend	kinc	
911	1 2	0 0 0 0		0 0		0 0 0 0	0	
	3 4	0 0 0 0	0 0 0	0 0	0	0 0 0 0	0	
	5	0 0 0 0	0 0 0	0 0	0	0 0 0 0	0	
	7 8	0 0 0 0	0 0 0 0	0 0	0 0	0 0 0 0	0 0	
	9 LO	0 0 0 0	0 0 0 0	0 0	0 0	0 0 0 0	0 0	
1	L1 L2	0 0 0 0	0 0 0 0	0 0	0 0	0 0 0 0	0 0	
	L3 /IE 0	0 0	0 0	0 0	0	0 0	0	
GF	INT OUT RID IPT		T IEND	IINC JSTART	JEND	JINC KSTA	RT KEND	KINC

NCS 4										
GRID	ISTA	IEND	JSTA	JEND	KSTA	KEND	IWALL	INORM		
1	0	0	1	1	0	0	0	-1		
7	0	0	1	1	0	0	0	-1		
8	0	0	1	1	0	0	0	-1		
13	0	0	49	49	0	0	0	1		
The	inflo	w.data	file is	:						
auxil 40, 2 5	-	c data	, j-fa	ce of	block	1, mul	ltistre	eam nozzl	le	
								16*0.6530 )., 40*0		0001,

After running this test case, the residual history and mass flow convergence history shown in Figure 9-17 results. Also, a plot of Mach contours should have the flow features of those plotted in Figure 9-16.

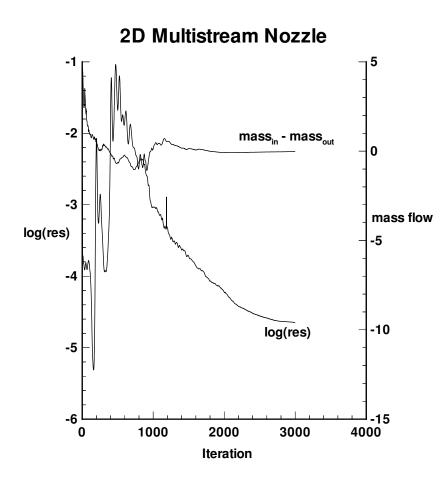


Figure 9-17. Multistream nozzle case residual and mass flow convergence history.

### 9.1.8 Rotor Stator

This case simulates, in two dimensions, the unsteady flow through a single stage turbine in which the ratio of stator to rotor blades is 3:4. The axial gap between the two blades is 50% of the blade chord. The case exercises a number of capabilities of CFL3D including unsteady flow, moving (translating) zones, dynamic patching between zones in relative motion, grid overlapping, and boundary conditions with user-defined input.

The original grid for this case was provided by D. J. Dorney<sup>17</sup> of Western Michigan University, although the grid given out for the test case contains only half the number of points of the original grid. The grid consists of fourteen zones with a total of 18374 points in one plane. A close-up of the grid near the airfoil is shown in Figure 9-18. The grid zones communicate with one another through both patching and overlapping. At a time step of 1.0, it takes 270 time steps for the eight rotor zones (containing four blades) to completely traverse the six stator zones (containing three blades). The rotor zones are reset after each complete traverse. The input file is set for 1500 time steps (using five multigrid sub-iterations per time step), which is sufficient to establish a time-periodic solution. The memory requirement for this example is 4.0 million words. A typical timing for this case (1500 time steps) is 4205 CPU seconds on a Cray YMP (NASA LaRC's Sabre as of October 1996). On a DEC Alpha workstation, the timing is 18303 CPU seconds, using single precision (as of June 1996).

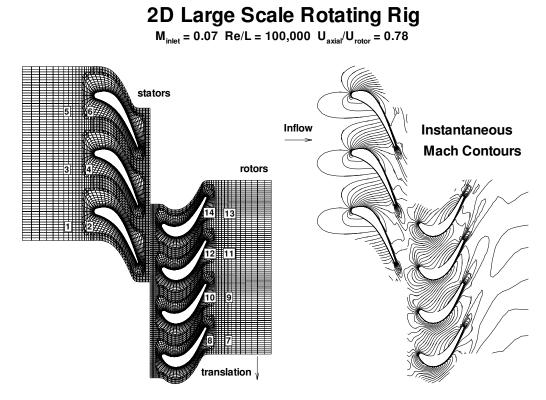


Figure 9-18. Fourteen zone rotor-stator grid system and Mach contours.

#### 9.1.8.1 Experimental Conditions

The experimental blade count was 22 stator blades and 28 rotor blades. In order to run an exact simulation, a minimum of 11 stator blades and 14 rotor blades would be required (Dring et.al<sup>18</sup>). To reduce the problem size in the computation, the ratio of stator to rotor blades was reduced to 3:4 (equivalent to 21:28), and the stators were scaled by a factor of 22/21 to maintain the same pitch-to-chord ratio as in the experiment. The experimental set-up was a three-dimensional configuration; the corresponding 2-d simulation was set up from conditions at the mid-span radius of 27 inches, for a rotor speed of 410 rpm, with a nominal axial velocity of 75 feet/second. The inlet Mach number in the experiment was approximately 0.07, and the Reynolds number/inch was approximately 100,000.

#### 9.1.8.2 Input Setup

Inlet conditions are used as the reference conditions, so **xmach** = 0.07. The grid is full scale, with dimensions in inches. Therefore, **reue** = 0.1. The inlet temperature was assumed to be  $60^{\circ}F$ , so **tinf** =  $520^{\circ}R$ .

Boundary condition type 2003 is used to specify total pressure and total temperature at the inlet. From isentropic flow relations or tables, for an inlet flow Mach number of 0.07,

$$M_{inlet} = 0.07$$

$$\frac{p_{t,inlet}}{p_{\infty}} = 1.0035$$

$$\frac{T_{t,inlet}}{T_{\infty}} = 1.0010$$
(9-12)

Also, alphae = betae = 0 (purely axial flow is assumed).

Boundary condition type 2002 is used to specify an exit pressure. Dorney gives a ratio of static pressure to inlet total pressure = 0.963 at the rotor trailing edge plane. Assuming this value to hold at the exit as well gives

$$\frac{p_{exit}}{p_{\infty}} = \frac{p_{exit}}{p_{t, inlet}} \frac{p_{t, inlet}}{p_{\infty}} = 0.963 \times 1.0035 = 0.967$$
(9-13)

Note that the inflow Mach number used in boundary condition type 2003 is an estimate; if the exit pressure were not set correctly, the computed inflow Mach number would not be close to the specified inflow value (when a time-periodic state is reached or at convergence in steady state). By specifying control surfaces at the inflow plane, the user is able to verify after the computation is complete that the average inflow Mach number is approximately 0.071; this was deemed to be close enough to the desired value. If desired, the exit pressure could be adjusted (raised in this case) and the solution re-run until a new time-periodic solution (and a new inlet Mach number) is established.

It should be noted that the input grid is in PLOT3D format, with y as the "up" direction (**ialph** = 1; z is the spanwise, 2-d direction). However, the grid motion parameters *must* be set as if z is the up direction. Recall that if the input grid has y as the up direction, CFL3D will internally swap y and z so that the code always computes on a grid in which z is up. (See the caution in "LT35 - Translational Information and Velocities" on page 44.)

Given the rotor speed and mid-span radius, the translational velocity for a 2-d simulation corresponding to the mid-span radius is

$$\tilde{w}_{trans} = \omega r = (410/60 \times 2\pi)(27/12) = 96.6$$
 feet/second (9-14)

This gives

$$\frac{\tilde{u}_{axial}}{\tilde{w}_{trans}} = \frac{75}{96.6} = 0.78$$
(9-15)

The input value **wtrans** is  $\tilde{w}_{trans}/\tilde{a}_{\infty}$ , so with the reference Mach number 0.07, **wtrans** = 0.07/0.78 = (-) 0.0897 (the negative gives a downward rotor motion).

In order to be able to run an arbitrarily long simulation, the grid resetting option was employed. The top-to-bottom length of the grid is 24.23514 inches and the rotor and stator zones start out in alignment, so **dzmax** = 24.23514. Thus the rotor zones are reset whenever the displacement exceeds 24.23514 inches.

#### 9.1.8.3 Running CFL3D

Besides the CFL3D code, the following files are needed to run this test case:

<u>File</u>	Description
lsrr.inp	input for CFL3D
lsrr_coarse.p3d_fmt	formatted single plane grid
fmttobin_p3d.f	converter for creating 2 grid planes
mag1.h	parameters for MaGGiE makefile
maggie.inp	input for MaGGiE

The steps for running this case on the DEC are as follows:

## <u>Step 1</u>

Compile the grid converter code:

```
cft77 fmttobin_p3d.f
```

## <u>Step 2</u>

Link the grid converter object file:

```
segldr -o fmttobin_p3d fmttobin_p3d.o
```

# <u>Step 3</u>

Run the grid converter program (the binary file lsrr\_coarse.p3d will be output):

```
fmttobin_p3d
```

# <u>Step 4</u>

Use the makefile to compile, link, and create the executable for the MaGGiE code (be sure mag1.h is in the current directory):

```
make -f makemaggie_cray
```

# Step 5

Run the MaGGiE code (the file ovrlp.bin will be output):

```
maggie < maggie.inp</pre>
```

# <u>Step 6</u>

Use the makefile to compile, link, and create the executable for the precfl3d code (be sure precfl.h is in the current directory):

```
make -f makeprecfl3d_cray
```

# <u>Step 7</u>

Run the precfl3d code (the cflx.h files will be output):

```
precfl3d < lsrr.inp</pre>
```

# <u>Step 8</u>

Use the makefile to compile, link, and create the executable for the CFL3D code:

```
make -f makecfl3d_cray
```

# <u>Step 9</u>

Run the CFL3D code:

cfl3d < lsrr.inp

The input file for this case is:

<pre>I/O FILES lsrr_coarse.] plot3dg.bin plot3dq.bin cfl3d.out cfl3d.res cfl3d.turres cfl3d.blomax cfl3d.out15 cfl3d.out20 ovrlp.bin patch.bin restart.bin</pre>	p3d						
XMACH 0.070 SREF 1.00000 DT +1.0000	ALPHA 0.000 CREF 1.00000 IREST 0	BETA 0.0 BREF 1.0000 IFLAGTS 000	REUE,MIL 0.1 XMC 0.00000 FMAX 1.00	TINF,DR 520.0 YMC 0.0000 IUNST 1	IALPH 1 ZMC 0.0000 CFL_TAU 5.	- 50% Axial IHSTRY 0	
NGRID -14 NCG	NPLOT3D 14 TEM	NPRINT 0 TADVANCE	NWREST 100 IFORCE	ICHK 0 IVISC(I)	I2D 1 IVISC (J)	NTSTEP 1500 IVISC (K)	ITA -2
NCG 1 1 1 1 1 1 1 1 1 1 1 1 1	IEM 0 0 0 0 0 0 0 0 0 0 0 0 0	IADVANCE 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	IFORCE 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	IVISC(I) 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	IVISC (J) 0 0 0 0 0 0 0 0 0 0 0 0 0 0	IVISC (K) 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5	
I LAMLO 00 00 00 00 00 00 00 00 00 00 00 00 00	ILAMHI 00 00 00 00 00 00 00 00 00 00 00 00 00	JLAMLO 000 000 000 000 000 000 000 000 000 0	JLAMHI 000 000 000 000 000 000 000 000 000 0	KLAMLO 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	KLAMHI 0000 0000 0000 0000 0000 0000 0000 0		

00 00 INEWG 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	00 00 IGRIDC 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	000 000 IS 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	000 JS 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	0 KS 0 0 0 0 0 0 0 0 0 0 0 0 0 0	0000 0000 IE 0 0 0 0 0 0 0 0 0 0 0 0 0 0	JE 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	KE 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0
IDIAG(I) 1 1 1 1 1 1 1 1 1 1 1 1 1	IDIAG (J) 1 1 1 1 1 1 1 1 1 1 1 1 1	IDIAG (K) 1 1 1 1 1 1 1 1 1 1 1 1 1	IFLIM(I) 0 0 0 0 0 0 0 0 0 0 0 0 0	IFLIM(J) 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	IFLIM(K) 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0		
1 1 1 1 1 1 1 1 1 GRID 1 2 3	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	. 3333 . 11 . 11	. 3333 . 11 . 11	. 3333 . 11 . 1	NBCKDIM 1 1	IOVRLP 1 1 1
4 5 6 7 8 9 10 11 12 13 14 10: GRID 1 2 3 4 5 6 7 8 9	1 1 1 1 1 1 1 1 SEGMENT 1 1 1 1 1 1 1 1 1 1 1 1	1 1 1 1 1 1 1 1 1 1 1 1 1 1	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1

	10 1 11 1 12 1 13 1	1002 1002 1002 1002	0 0 0	0 0 0	0 0 0	0 0 0	0 0 0
IDIM:	14 1 GRID SEGMENT 1 1	1002 BCTYPE 1002	0 JSTA 0	0 JEND 0	0 KSTA 0	0 KEND 0	0 NDATA 0
	2 1 3 1 4 1	1002 1002 1002	0 0 0	0 0 0	0 0 0	0 0 0	0 0 0
	$egin{array}{cccc} 5 & 1 \ 6 & 1 \ 7 & 1 \ 8 & 1 \end{array}$	1002 1002 1002 1002	0 0 0	0 0 0 0	0 0 0 0	0 0 0 0	0 0 0 0
	9 1 10 1 11 1	1002 1002 1002 1002	0 0 0	0 0 0	0 0 0	0 0 0	0 0 0
	12 1 13 1 14 1	1002 1002 1002	0 0 0	0 0 0		0 0 0	0 0 0
J0: Mach	GRID SEGMENT 1 1 Pt/Pinf Tt/Tinf	BCTYPE 2003	ISTA 0 ta	IEND 0	KSTA 0	KEND 0	ndata 5
0.07	1.0035 1.0010 2 1 3 1		0. 0 0	0 0	0 0	0 0	0 5
Mach 0.07	Pt/Pinf Tt/Tinf 1.0035 1.0010 4 1	0.	0.	0	0	0	0
Mach 0.07	5 1 Pt/Pinf Tt/Tinf 1.0035 1.0010	- O.	0.	0	0	0	5
	$egin{array}{cccc} 6 & 1 \ 7 & 1 \ 8 & 1 \ 9 & 1 \end{array}$	0 0 0 0	0 0 0 0	0 0 0 0	0 0 0 0	0 0 0 0	0 0 0 0
	10 1 11 1 12 1	0 0 0	0 0 0	0 0 0	0 0 0	0 0 0	0 0 0
JDIM:	13 1 14 1 GRID SEGMENT 1 1	0 0 BCTYPE 0	0 0 ISTA 0	0 0 IEND 0	0 0 KSTA 0	0 0 KEND 0	0 0 NDATA 0
	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	0 0 0	0 0 0	0 0 0	0 0 0	0 0 0	0 0 0
	5 1 6 1 7 1	0 0 2002	0 0 0	0 0 0	0 0 0	0 0 0	0 0 1
pexit 0.967	/pinf 8 1	0	0	0	0	0	0
pexit 0.967		2002	0	0	0	0	1
	10 1 11 1 2/pinf	0 2002	0 0	0 0	0 0	0 0	0 1
0.967	12 1 13 1	0 2002	0 0	0 0	0 0	0 0	0 1
0.967 K0:	/pinf 14 1 GRID SEGMENT	0 BCTYPE	0 ISTA	0 IEND	0 JSTA	0 JEND	0 NDATA
Tw/Ti	$\begin{array}{ccc} 1 & 1 \\ 2 & 1 \end{array}$	0 2004	0 0	0 0	0 0 0	0 0 0	0 2
0.	$\begin{array}{ccc} 0. \\ 3 & 1 \\ 4 & 1 \end{array}$	0 2004	0 0	0 0	0 0	0 0	0 2
Tw/Ti	.nf C_q						

0. 0 5 6		0 2004	0 0	0	0 0	0	0 2
Tw/Tinf C	C_q ).	0	0	0	0	0	
8 Tw/Tinf C	1 2_q	2004	0	0	0	0	0 2
9 10	1	0 2004	0 0	0 0	0 0	0 0	0 2
		0	0	0	0	0	0
12 Tw/Tinf C 0. 0	1 C_q ).	2004	0	0	0	0	2
13 14 Tw/Tinf C		0 2004	0 0	0 0	0 0	0 0	0 2
0. 0	).						
KDIM: GRID	SEGMENT	BCTYPE	ISTA	IEND	JSTA	JEND	NDATA
1 2	1	0	0	0 0	0 0	0	0 0
3	1	0	Ő	0	Ő	Ő	0
4	1	0	0	0	0	0	0
5 6	1 1	0	0 0	0 0	0 0	0	0 0
7	1	0	0	0	0	0	0
8	1	0	0	0	0	0	0
9	1	0	0	0	0	0	0
10 11	1 1	0	0 0	0 0	0 0	0 0	0 0
12	1	0	0	0	0	0	0
13	1	0	0	0	0	0	0
14	1	0	0	0	0	0	0
MSEQ 1	MGFLAG 1	ICONSF 1	MTT 0	NGAM 01			
ISSC 0	EPSSC(1)	EPSSC(2)	EPSSC(3)	ISSR 0	EPSSR(1) .3	EPSSR(2)	EPSSR(3) .3
NCYC 5 MIT1	MGLEVG 02 MIT2	NEMGL 00 MIT3	NITFO 000 MIT4	MIT5			
01	01	01	01	01			
1-1 BLOC	CKING DATA:						
NBLI							
14 NUMBER	GRID :	ISTA J	ISTA KSTA	IEND	JEND KENI	D ISVA1	ISVA2
1	2	1	1 1	2	1 2		3
2	4	1	1 1	2	1 2		3
3 4	6 8	1 1	$\begin{array}{ccc} 1 & 1 \\ 1 & 1 \end{array}$	2 2	1 2 1 2		3 3 3 3 3 3 2
5	10	1	1 1	2	1 2		3
5 6	12	1	1 1	2	1 21	1 1	3
7							3
8 9	14	1	1 1	2	1 21		-
	14 1	1	1 1	2 2	55	1 1	2
10	14 1 7	1 1	$\begin{array}{ccc} 1 & 1 \\ 1 & 1 \end{array}$	2 2 2	55 5 61	$egin{array}{ccc} 1 & 1 \ 1 & 1 \end{array}$	2 2 2
10 11	14 1	1	1 1	2 2 2 2 2	55 5 61 55 23	1 1 1 1 3 1	2 2 2
10 11 12	14 1 7 1 3 7	1 1 1 1 1	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	2 2 2 2 2 2 2	55 61 55 2 55 2 61 2	1 1 1 1 3 1 3 1 3 1 3 1	2 2 2 2
10 11 12 13	14 1 7 1 3 7 9	1 1 1 1 1	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	2 2 2 2 2 2 2 2 2	55 61 55 2 55 2 61 2 61 2	1 1 1 1 3 1 3 1 3 1 3 1 3 1	2 2 2 2 2 2
10 11 12 13 14	14 1 7 1 3 7 9 11	1 1 1 1 1 1	$\begin{array}{cccc} 1 & 1 \\ 1 & 1 \\ 1 & 23 \\ 1 & 23 \\ 1 & 23 \\ 1 & 23 \\ 1 & 23 \\ 1 & 23 \end{array}$	2 2 2 2 2 2 2 2 2	55 61 55 2 55 2 61 2 61 2 61 2 61 2	1 1 1 1 3 1 3 1 3 1 3 1 3 1 3 1 3 1	2 2 2 2 2 2 2
10 11 12 13 14 NUMBER 1	14 1 7 3 7 9 11 GRID :	1 1 1 1 1 1	1 1 1 23 1 23 1 23 1 23 1 23 1 23 1 23 1 23 VSTA KSTA	2 2 2 2 2 2 2 2 1 END 2	55 61 55 55 61 2 61 2 61 2 JEND KEN	1 1 1 1 3 1 3 1 3 1 3 1 3 1 3 1 3 1 5 ISVA1	2 2 2 2 2 1SVA2
10 11 12 13 14 NUMBER 1 2	14 1 7 9 11 GRID : 2 4	1 1 1 1 1 1 ISTA 3 1 1	1 1 1 23 1 57A KSTA 61 1 61 1	2 2 2 2 2 2 2 1 END 2 2 2	55 2 61 2 55 2 61 2 61 2 JEND KENI 61 2 61 2 61 2	1 1 1 1 3 1 3 1 3 1 3 1 3 1 3 1	2 2 2 2 2 1SVA2
10 11 12 13 14 NUMBER 1 2 3	14 1 7 1 3 7 9 11 GRID : 2 4 6	1 1 1 1 1 1 ISTA 3 1 1	1 1 1 23 1 57A KSTA 61 1 61 1 61 1	2 2 2 2 2 2 2 2 1END 2 2 2 2	55 61 55 61 2 61 2 61 2 JEND KENI 61 2 61 2 61 2 61 2	1 1 1 1 3 1 3 1 3 1 3 1 3 1 3 1	2 2 2 2 2 1SVA2
10 11 12 13 14 NUMBER 1 2 3 4	14 1 7 1 3 7 9 11 GRID : 2 4 6 8	1 1 1 1 1 1 ISTA 3 1 1 1 1	1 1 1 23 1 61 1 61 1 61 1 61 1	2 2 2 2 2 2 2 2 1END 2 2 2 2 2 2 2	55 2 61 2 55 2 61 2 61 2 JEND KEN 61 2 61 2 61 2 61 2 61 2 61 2	1 1 1 1 3 1 3 1 3 1 3 1 3 1 3 1	2 2 2 2 2 1SVA2
10 11 12 13 14 NUMBER 1 2 3 4 5	14 1 7 1 3 7 9 11 GRID : 2 4 6 8 10	1 1 1 1 1 1 ISTA 3 1 1 1 1 1	1 1 1 23 1 23 1 23 1 23 1 23 1 23 1 23 ISTA KSTA 61 1 61 1 61 1 61 1 61 1	2 2 2 2 2 2 2 2 1END 2 2 2 2 2 2 2 2 2	55 2 61 2 55 2 61 2 61 2 JEND KENN 61 2 61 2 61 2 61 2 61 2 61 2 61 2 61 2 61 2 61 2	1       1         1       1         3       1         3       1         3       1         3       1         3       1         1       1         1       1         1       1         1       1         1       1         1       1         1       1         1       1	2 2 2 2 2 1SVA2
10 11 12 13 14 NUMBER 1 2 3 4	14 1 7 9 11 GRID : 2 4 6 8 10 12 14	1 1 1 1 1 1 ISTA 3 1 1 1 1	1       1         1       23         1       23         1       23         1       23         1       23         1       23         1       23         1       23         1       23         1       23         1       23         1       23         1       23         1       23         1       23         1       23         1       1         61       1         61       1         61       1         61       1         61       1         61       1         61       1         61       1         61       1         61       1         61       1	2 2 2 2 2 2 2 2 2 1END 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	55 2 61 2 55 2 61 2 61 2 61 2 JEND KENN 61 2 61 2 6	1 1 1 1 3 1 3 1 3 1 3 1 3 1 3 1	2 2 2 2 2 1SVA2 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3
10 11 12 13 14 NUMBER 1 2 3 4 5 6 7 8	14 1 7 9 11 GRID : 2 4 6 8 10 12 14 5	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	1       1         1       1         1       23         1       23         1       23         1       23         1       23         1       23         1       23         1       23         1       23         1       23         1       1         61       1<	2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	55 2 61 2 55 2 61 2 61 2 61 2 JEND KENI 61 2 61 2 6	1       1         1       1         3       1         3       1         3       1         3       1         3       1         3       1         1       1         1       1         1       1         1       1         1       1         1       1         1       1         1       1         1       1         1       1         3       1	2 2 2 2 2 1SVA2 3 3 3 3 3 3 3 2
10 11 12 13 14 NUMBER 1 2 3 4 5 6 7	14 1 7 9 11 GRID : 2 4 6 8 10 12 14	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	1       1         1       23         1       23         1       23         1       23         1       23         1       23         1       23         1       23         1       23         1       23         1       23         1       1         61       1         61       1         61       1         61       1         61       1         61       1         61       1         61       1         61       1         61       1         61       1         61       1         61       1         61       1         61       1         61       1	2 2 2 2 2 2 2 2 2 1END 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	55 2 61 2 55 2 61 2 61 2 61 2 JEND KENN 61 2 61 2 6	1       1         1       1         3       1         3       1         3       1         3       1         3       1         3       1         1       1         1       1         1       1         1       1         1       1         1       1         1       1         1       1         1       1         1       1         3       1	2 2 2 2 2 1SVA2 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3

#### CHAPTER 9 Test Cases

1 1 1 PATCH NINT	1 5 2 9 3 11 4 13 SURFACE D	PATA:	1 1 1 1	1 1 1 1	1 1 1 1	2 2 2 2 2	55 55 61 61 61	1 1 1 1	1 1 1 1	2 2 2 2 2 2
	OUTPUT: IPTYPE IS 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	TART 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	IEND 001 001 001 001 001 001 001 001 001 00	IINC 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	JSTART 01 01 01 01 01 01 01 01 01 01 01 01	JEND 999 999 999 999 999 999 999 999 999 9	JINC 1 1 1 1 1 1 1 1 1 1 1 1 1 1	KSTART 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	KEND 999 999 999 999 999 999 999 999 999 9	KINC 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1
PRINT BLOCK CONTROL NCS	OUT: IPTYPE IS SURFACES		IEND	IINC	JSTART	JEND	JINC	KSTART	KEND	KINC
NTRANS	1 1 1 1 1 GRID DAT	2 2	9 999 9 999 9 999 9 999 1 1 1 1 1 1	9 9 9 9 1 1 1	A         KENE           0         0           0         0           0         0           0         0           0         0           0         0           0         0           0         0           0         0		INORM 1 1 1 0 0 0			
NROTAT 0	ITRANS 1 1 1 1 1 1 1 1 1 1 1 1 1	0. 0. 0. 0. 0. 0. 0.	UTRAN DZMZ -24.235 -24.235 -24.235 -24.235 -24.235 -24.235 -24.235 -24.235 -24.235 -24.235 -24.235 -24.235 -24.235	0. 0. 0. 0. 0. 0. 0. 0. 0. 0.	0. 0. 0. 0. 0. 0.	WTRANS -0.0897 -0.0897 -0.0897 -0.0897 -0.0897 -0.0897 -0.0897 -0.0897				
GRID	IROTAT THXMAX C PATCH I	THYM2	AX TH	AX O IZMAX	MEGAY	OMEGAZ	XORI	G YOR:	IG ZOH	RIG

	7							
INT		LIMIT	ITMAX	MCXIE	MCETA	C-0	IORPH	ITOSS
1	1		30 30	0	0 0	0	0	1
2 3	1 1			0 0	0	0	0	1 1
4	1	1	30	0	0 0	0		1
5	1	1	30	0	0 0	0		1
6 7	1 1	1 1	30 30	0 0	0	0	0	1
INT	TO	XIE1	XIE2	ETA1	ETA2	NFB	0	T
1	122	0	0	0	0	6		
	FROM 721	XIE1	XIE2	ETA1	ETA2	FACTJ 0.	FACTK 0.	
	DX	DY	0 DZ 0.	DTHETX	DTHETY	DTHETZ		
	0.	0.	0. XIE2	0.	0.	0.		
	FROM 921	XIE1 0	XIE2	ETA1	ETA2	FACTJ	FACTK 0.	
	DX	DY	DZ	DTHETX	0 DTHETY 0.	DTHETZ	0.	
	0.		0. XIE2	0.	0.	0.		
	FROM 1121	XIE1 0	XIE2	ETA1	ETA2	FACTJ	FACTK 0.	
	DX	DY	DZ	DTHETX	0 DTHETY 0.	DTHETZ	0.	
	0.	0.	0. XIE2	0.	0.	0.		
	FROM 1321	XIE1 0	XIE2	ETA1	ETA2	FACTJ	FACTK 0.	
	DX	DY	DZ	DTHETX	0 DTHETY 0.	DTHETZ	0.	
	0.		0. XIE2	0.	0.	0.		
	FROM 721	XIE1	XIE2	ETA1	ETA2	FACTJ	FACTK 0.	
	DX	DY	0 DZ	DTHETX	DTHETY	DTHETZ		
	0.	0.	24.23514 XIE2	0.	0.	0.		
	FROM 921	XIE1	XIE2	ETA1	ETA2	FACTJ 0.	FACTK 0.	
	DX	DY	0 DZ	DTHETX	DTHETY	DTHETZ	0.	
	0.	0.	24.23514 XIE2	0.	0.	0.		
INT 2	TO 322	XIE1	XIE2	ETA1	ETA2	NFB 6		
Z	FROM	XIE1	0 XIE2	ETA1	ETA2	FACTJ	FACTK	
	921	0	0 DZ 0.	0	0	0.	0.	
	DX 0.	DY	DZ	DTHETX	DTHETY	DTHETZ		
	FROM	XIE1	XIE2	ETA1	ETA2	FACTJ	FACTK	
	1121	0	0 DZ 0.	0	0	0.	0.	
	DX 0.	DY 0	DZ O	DTHETX	DTHETY	DTHETZ		
	FROM	XIE1	XIE2	ETA1	ETA2	FACTJ	FACTK	
	1321	0	0 DZ	0	0	0.	0.	
	DX 0.	DY 0.	DZ 0.	DTHETX 0.	DTHETY 0.	DTHETZ 0.		
	FROM	XIE1	XIE2	ETA1	ETA2	FACTJ	FACTK	
	721	0	0	0	0		0.	
	DX 0.	DY 0.		DTHETX 0.	DTHETY 0.	DTHETZ 0.		
	FROM	XIE1	XIE2	ETA1	ETA2	FACTJ	FACTK	
	921	0	0	0	0		0.	
	DX 0.	DY 0.		DTHETX 0.	DTHETY 0.	DTHETZ 0.		
	FROM	XIE1	XIE2	ETA1	ETA2	FACTJ	FACTK	
	1121	0	0	0 DTHETX	0 DTHETY		0.	
	DX 0.	DY 0.			DIHEIY 0.	DTHETZ 0.		
INT	TO	XIE1	XIE2	ETA1	ETA2	NFB		
3	522	0 XIE1	0 XIE2	0	0	6		
	FROM 1121	XIEI 0	XIEZ 0	ETA1 0	ETA2 0		FACTK 0.	
	DX	DY	DZ	DTHETX	DTHETY			
	0. FROM	0. XIE1	0. XIE2	0. ETA1	0. ETA 2	0. FACTJ	FACTK	
	FROM 1321	XIEI 0	0	0	0		FACIK 0.	
	DX	DY	DZ	DTHETX	DTHETY	DTHETZ		
	0. FROM	0. XIE1	0. XIE2	0. ETA1	0. ETA2	0. FACTJ	FACTV	
	LINOM	VIET	ALEZ	DIAL	IAZ	TACIU	TACIN	

	721	0	0	0	0	0.	0.
	DX 0.	DY	DZ 24.23514	DTHETY	DTHETY	DTHETZ	
	FROM 921	XIE1 0	XIE2 0	ETA1 0	ETA2 0	FACTJ	FACTK 0.
	DX	DY	DZ 24.23514	DTHETX	DTHETY	DTHETZ	0.
	0. FROM	XIE1	XIE2	ETA1	ETA2 0	FACTJ	FACTK
	1121 DX	0 DY	D7.	DTHETX	DTHETY	DTHETZ	0.
	0. FROM	0. XIE1	24.23514 XIE2 0	0. ETA1	0. ETA2	FACTJ	FACTK
	1321 DX	DY	D7.	DTHETY	DTHETY	0. DTHETZ	0.
INT	0. TO	Ο.	24.23514	Ο.	0.	0.	
4	721		XIE2 0 XIE2		ETA2 0	NFB 4 EACTI	EACTE
	FROM 122		XIE2 0		0	0.	FACTK 0.
	DX 0.	DY 0.	D2 0.	DTHETX 0.	DTHETY 0.	DTHETZ 0.	
	FROM 522	XIE1 0	XIE2 0	ETA1 0	ETA2 0	FACTJ 0.	FACTK 0.
	DX 0.	DY		DTHETX	DTHETY	DTHETZ	
	FROM 322	XIE1	XIE2 0	ETA1	ETA2	FACTJ	FACTK 0.
	DX	DY	D7	DTHETX	DTHETY	DTHETZ	0.
	0. FROM	U. XIE1	-24.23514 XIE2 0	U. ETA1	0. ETA2	0. FACTJ	FACTK
	122 DX	DY	D7	DTHETX	DTHETY	DTHETZ	0.
INT	0. TO	0. XIE1	-24.23514	0.	0.	Ο.	
5	921 FROM	0 XTE1		0 ETA1	ETA2 0 ETA2	5 FACTJ	FACTK
	122 DX	0 DY	XIE2 0			0.	0.
	0.	0.	0.	0.	DTHETY 0.		
	FROM 322	XIE1 0	0	0	ETA2 0	0.	FACTK 0.
	DX 0.	DY 0.	DZ 0.	DTHETX 0.	DTHETY 0.	DTHETZ 0.	
	FROM 522			ETA1 0	ETA2 0	FACTJ 0.	FACTK
	DX 0.	DΥ		DTHETY	DTHETY	DTHETZ	
	FROM	XIE1	XIE2	ETA1	ETA2	FACTJ	
	DX	DY	0 DZ	DTHETX	DTHETY	DTHETZ	0.
	0. FROM		-24.23514 XIE2		0. ETA2		FACTK
	122 DX	0 DY		0	0	0. DTHETZ	0.
INT	0. TO	0. XIE1	-24.23514		0. ETA2	0. NFB	
6	1121 FROM	0	0 XIE2	0	0	5 FACTJ	FACTK
	322	0	0	0	0	0.	0.
	DX 0.	DY 0.	DZ O.	0.	0.	0.	
	FROM 522	XIE1 0	XIE2 0	ETA1 0	ETA2 0	0.	FACTK 0.
	DX 0.	DY 0.	DZ O.	DTHETX 0.	DTHETY 0.	DTHETZ 0.	
	FROM 122	XIE1 0	XIE2 0	ETA1 0	ETA2 0	FACTJ	FACTK 0.
	DX	DY	DZ	DTHETX	DTHETY	DTHETZ	0.
	0. FROM	0. XIE1	0. XIE2	0. ETA1	0. ETA2	FACTJ	FACTK
	522 DX	0 DY	0 DZ	ETA1 0 DTHETX	0 DTHETY		0.

	0.	0.	-24.23514	0.	0.	0.	
	FROM	XIE1	XIE2	ETA1	ETA2	FACTJ	FACTK
	322	0	0	0	0	0.	0.
	DX	DY	DZ	DTHETX	DTHETY	DTHETZ	
	Ο.	Ο.	-24.23514	Ο.	0.	0.	
INT	TO	XIE1	XIE2	ETA1	ETA2	NFB	
7	1321	0	0	0	0	4	
	FROM	XIE1	XIE2	ETA1	ETA2	FACTJ	FACTK
	522	0	0	0	0	0.	Ο.
	DX	DY	DZ	DTHETX	DTHETY	DTHETZ	
	Ο.	Ο.	0.	Ο.	0.	0.	
	FROM	XIE1	XIE2	ETA1	ETA2	FACTJ	FACTK
	322	0	0	0	0	0.	Ο.
	DX	DY	DZ	DTHETX	DTHETY	DTHETZ	
	Ο.	Ο.	Ο.	Ο.	0.	Ο.	
	FROM	XIE1	XIE2	ETA1		FACTJ	FACTK
	122	0	0	0	0	Ο.	Ο.
	DX	DY	DZ	DTHETX	DTHETY	DTHETZ	
	Ο.	Ο.	Ο.	Ο.	Ο.	0.	
	FROM	XIE1	XIE2	ETA1	ETA2	FACTJ	FACTK
	522	0	0	0	0	0.	0.
	DX	DY	DZ	DTHETX	DTHETY	DTHETZ	
	0.	0.	-24.23514	0.	0.	0.	

The convergence histories for residual, mass flow, and rotor lift coefficient as shown in Figure 9-19 should be obtained.

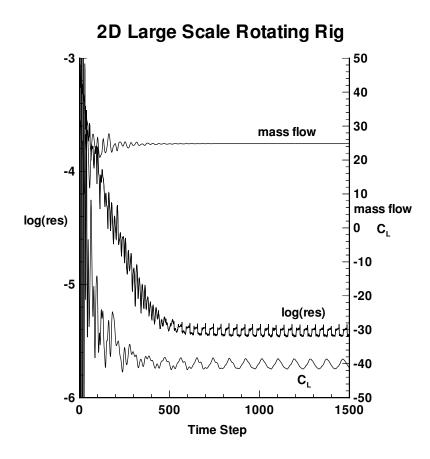


Figure 9-19. Convergence histories for rotor-stator case.

# 9.2 Three-dimensional Test Cases

#### 9.2.1 Axisymmetric Bump Flow

This test case solves for the turbulent flow over an axisymmetric bump. The flow is modeled in 3-d using two computational planes (separated by an angle of 1 degree), with periodic boundary conditions; hence **bctype** is 2005 and **dthtx** is –1.0 and 1.0 on the I0 and IDIM boundaries, respectively. The grid consists of a single zone with a total of 36562 points. The memory requirement for this example is 4.9 million words. A typical timing for this case is 1026 CPU seconds on a CRAY YMP (NASA LaRC's Sabre as of October 1996). A close-up of the grid near the bump is shown in Figure 9-20.

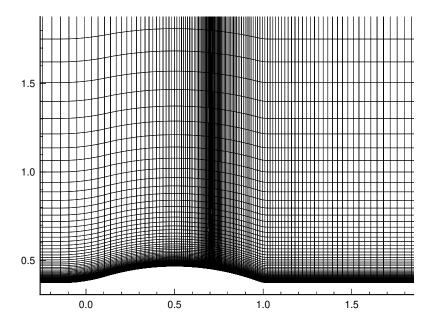


Figure 9-20. Axisymmetric bump grid.

Besides the CFL3D code, the following files are needed to run this test case:

<u>File</u>	Description
bumpv5periodic.inp	input for CFL3D
bump.grd	formatted single plane grid
gridaxi.f	converter for creating 2 grid planes

The steps for running this case on the YMP are as follows:

# <u>Step 1</u>

Compile the grid converter code:

```
cft77 gridaxi.f
```

# Step 2

Link the grid converter object file:

segldr -o gridaxi gridaxi.o

# <u>Step 3</u>

Run the grid converter program (the binary file bumpgrd.bin will be output):

gridaxi

In answer to the question, type:

bump.grd

#### <u>Step 4</u>

Use the makefile to compile, link, and create the executable for the precfl3d code (be sure precfl.h is in the current directory):

```
make -f makeprecfl3d_cray
```

# <u>Step 5</u>

Run the precf13d code (the cflx.h files will be output):

```
precfl3d < bumpv5periodic.inp</pre>
```

# <u>Step 6</u>

Use the makefile to compile, link, and create the executable for the CFL3D code:

```
make -f makecfl3d_cray
```

#### <u>Step 7</u>

Run the CFL3D code:

cfl3d < bumpv5periodic.inp

The input file for this case is:

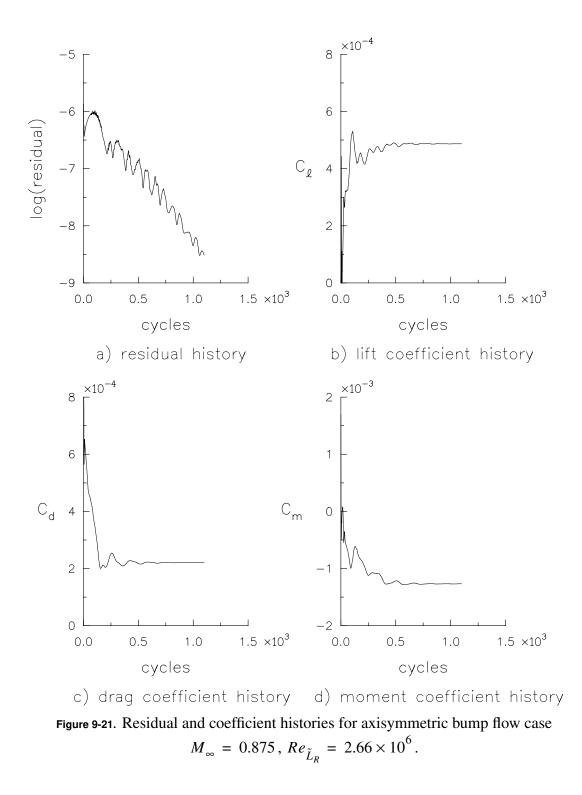
I/O FILES bumpgrd.bin plot3dg.bin cfl3d.out cfl3d.res cfl3d.turres cfl3d.blomax

cfl3d.out15 cfl3d.prout							
cfl3d.out20							
ovrlp.bin patch.bin							
restart.bin		flow 2-c	a uning	2 planes and	1 poriodia	PCa	
XMACH	ALPHA		REUE,MI		I periodic IALPH	IHIST	
0.8750 SREF	00.000 CREF	0.0 BREF	02.60 XN		0 ZMC	0	
1.00000	1.00000	1.0000	0.0000		0.00		
DT -05.000	IREST 0	IFLAGTS 000	FMA 5.000		CFLTAU 10.		
NGRID	NPLOT3D	NPRINT	NWRES		I2D	NTSTEP	ITA
1 NCG 2	1 IEM 0	1 IADVANCE 0	120 IFORC 00	CE IVISC(I)	0 IVISC(J) 0	1 IVISC(K) 07	1
IDIM	JDIM K	DIM	00	0	0	07	
02 ILAMLO	181 ILAMHI	101 JLAMLO	JLAMH	HI KLAMLO	KLAMHI		
0	0	0		0 0	0		
INEWG 0	IGRIDC 0	IS 0	, L	JS KS 0 0	IE O	JE O	KE 0
IDIAG(I)	IDIAG(J)	IDIAG(K)	IFLIM(]	I) IFLIM(J)	IFLIM(K)	Ű	Ũ
1 IFDS(I)	1 IFDS (J)	1 IFDS (K)	RKAP0(1	3 3 I) RKAPO(J)	3 RKAP0(K)		
1	1	1	0.333	0.3333	0.3333		
GRID 1	NBCIO 1	NBCIDIM 1	NBCJ	JONBCJDIM11	NBCKO 1	NBCKDIM 1	IOVRLP 0
IO: GRID	SEGMENT	BCTYPE	JSI	TA JEND	KSTA	KEND	NDATA
1	1 NBLP	2005 DTHTX	DTHI	0 0 TY DTHTZ	0	0	4
TRIM ORID	1	-1.0		0.	KOMA	VEND	
IDIM: GRID 1	SEGMENT 1	BCTYPE 2005	JSI	TA JEND 0 0	KSTA 0	KEND 0	NDATA 4
	NBLP	DTHTX	DTHI				
J0: GRID	1 SEGMENT	+1.0 BCTYPE	ISI	). O. TA IEND	KSTA	KEND	NDATA
1 IDIM: CDID	1 CECMENT	1003	тел	0 0	0 КСТА	0 KEND	0
JDIM: GRID 1	SEGMENT 1	BCTYPE 1003	IST	TA IEND 0 0	ksta 0	KEND 0	NDATA 0
KO: GRID 1	SEGMENT 1	BCTYPE 2004	IST	TA IEND 0 0	JSTA 0	JEND 0	NDATA 2
T	TWTYPE	2004 CQ		0 0	0	U	2
KDIM: GRID	0. SEGMENT	0. BCTYPE	IST	TA IEND	JSTA	JEND	NDATA
RDIM: GRID	J	1003	191	0 0	0	0 END	NDATA 0
MSEQ 1	MGFLAG 1	ICONSF 0	MI	0 NGAM 02			
		EPSSSC(2)	EPSSSC (3			EPSSSR(2)	
0 NCYC	0.3 MGLEVG	0.3 NEMGL	0. NITE		0.3	0.3	0.3
1100	03	00	00	00			
MIT1 01	MIT2 01			T4 MIT5 D1 01			MIT8 1
1-1 BLOC	KING DATA:				-	-	1
NBLI 0							
NUMBER G						END ISVA1	
NUMBER G PATCH SUR	RID : FACE DATA:		JSTA P	KSTA IEND	JEND KI	END ISVA1	ISVA2
NINTER							
0 PLOT3D OU	TPUT:						
BLOCK IPT	YPE ISTARI				JINC KSTA		KINC
IMOVIE	0 C	0	0	0 0	0	0 0	0
0							
PRINT OUT BLOCK IPT	: YPE ISTARI	IEND	IINC JST	TART JEND	JINC KSTA	ART KEND	KINC
1 CONTROL S	0 1	. 1	1	0 0	0	1 1	1
CONTROL 2	UNFACE:						

9.2.1 Axisymmetric Bump Flow

NCS 0 GRID ISTART IEND JSTART JEND KSTART KEND IWALL INORM

After running this test case, the convergence history plots shown in Figure 9-21 should be duplicated.



Also, a result such as that shown in Figure 9-22 should be obtained. In the figure, surface pressure coefficients are plotted along with experimental data for this case. The computational surface pressures can be obtained from file cfl3d.prout. Experimental surface pressure coefficients from Bachalo et. al<sup>8</sup> are included with this test case for comparison purposes. The file is called bumpcpdata.dat.

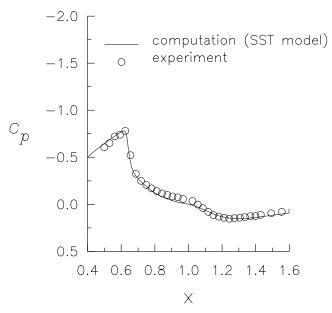


Figure 9-22. Pressure coefficients for axisymmetric bump case

$$M_{\infty} = 0.875, Re_{\tilde{L}_R} = 2.66 \times 10^6.$$

# 9.2.2 F-5 Wing

The inviscid flow over an F-5 wing is solved in this test case. The grid consists of a single grid zone with a C-H mesh topology and is composed of 210,177 points. The memory requirement for this example is 10.5 million words. A typical timing for this case is 984 CPU seconds on a CRAY YMP (NASA LaRC's Sabre as of September 1996). The wing surface grid and wake, as well as the plane of symmetry grid are illustrated in Figure 9-23.

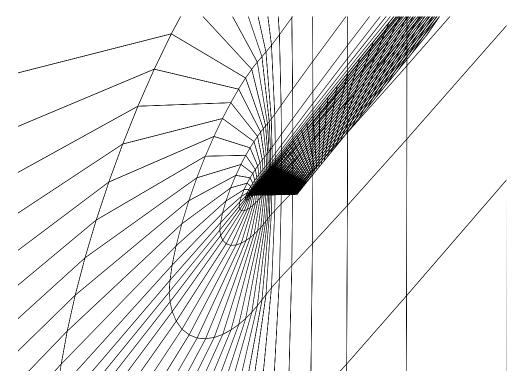


Figure 9-23. Single zone F-5 wing surface grid and plane of symmetry grid.

Besides the CFL3D code the following files are needed to run this test case:

File	Description
f5wing.inp	input for CFL3D
f5grid.dat	formatted wing section grid
f5wing_grid.f	grid converter

The steps for running this case on the YMP are as follows:

# <u>Step 1</u>

Compile the grid converter code:

cft77 f5wing\_grid.f

#### Step 2

Link the grid converter object file:

```
segldr -o f5wing_grid f5wing_grid.o
```

#### <u>Step 3</u>

Run the grid converter program to generate the 3-d volume grid (the binary file f5wing.grd will be output):

f5wing\_grid

#### <u>Step 4</u>

Use the makefile to compile, link, and create the executable for the precfl3d code (be sure precfl.h is in the current directory):

```
make -f makeprecfl3d_cray
```

#### <u>Step 5</u>

Run the precf13d code (the cf1x.h files will be output):

precfl3d < f5wing.inp</pre>

#### <u>Step 6</u>

Use the makefile to compile, link, and create the executable for the CFL3D code:

```
make -f makecfl3d_cray
```

#### <u>Step 7</u>

Run the CFL3D code:

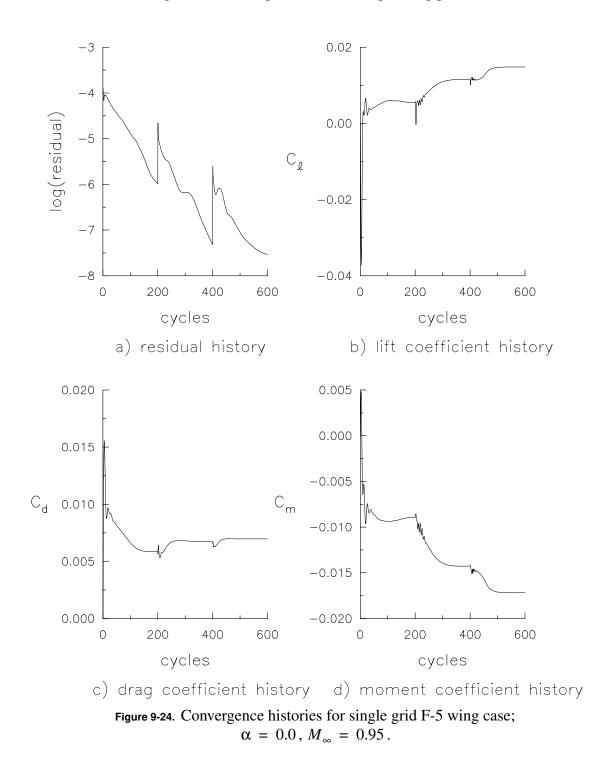
cfl3d < f5wing.inp

The input file for this case is:

```
I/O FILES:
f5wing.grd
plot3dg.bin
plot3dq.bin
cfl3d.out
cfl3d.res
cfl3d.turres
cfl3d.blomax
cfl3d.out15
cfl3d.prout
cfl3d.out20
ovrlp.bin
patch.bin
restart.bin
     F5 Wing, cfl3d type grid
     XMACH
               ALPHA
                          BETA REUE, MIL
                                            TINF,DR
                                                        IALPH
                                                                   IHIST
```

0.950 SREF	00.000 CREF	0.0 BREF		460.0 YMC	0 ZMC	0	
1.00000		1.0000		0.00	0.00		
DT -5.000	IREST 0	IFLAGTS 000	FMAX 1.000	IUNST 0	CFLTAU 10.		
NGRID	NPLOT3D 0	NPRINT 0	NWREST 100	ICHK 0	IO. I2D	NTSTEP 1	ITA 1
NCG	IEM	IADVANCE	IFORCE	IVISC(I)	IVISC(J)	IVISC(K)	1
2 IDIM	0 JDIM	0 KDIM	1	0	0	0	
33 ILAMLO	193 ILAMHI	33 JLAMLO	JLAMHI	KLAMLO	KLAMHI		
00 INEWG	00 IGRIDC	000 IS	000 JS	0 KS	0000 IE	JE	KE
INEWG 0	IGRIDC 0	13	0	0	16	0	0
IDIAG(I) 1	IDIAG(J) 1	IDIAG(K) 1	IFLIM(I) 3	IFLIM(J) 3	IFLIM(K) 3		
IFDS(I) 1	IFDS(J) 1	IFDS(K) 1	RKAP0(I) .3333	RKAP0(J) .3333	RKAP0(K) .3333		
GRID 1	NBCIO 1	NBCIDIM 1	NBCJ0	NBCJDIM 1	NBCK0	NBCKDIM 1	
IO: GRID	SEGMENT	BCTYPE	JSTA	JEND	KSTA	KEND	NDATA
1 IDIM: GRID	1 SEGMENT	1001 BCTYPE	JSTA	0 JEND	0 KSTA	0 KEND	0 NDATA
1 J0: GRID 1	1 SEGMENT 1	1002 BCTYPE 1003	0 ISTA 0	0 IEND 0	0 KSTA 0	0 KEND 0	0 NDATA 0
JDIM: GRID	SEGMENT	BCTYPE	ISTA	IEND	KSTA	KEND	NDATA
1 KO: GRID	1 SEGMENT	1003 BCTYPE	0 ISTA	0 IEND	0 JSTA	0 JEND	0 NDATA
1	1	0 1005	1	33 21	1 41	41 153	0
1	3	0	21	33	41	153	0
1 KDIM: GRID	4 SEGMENT	0 BCTYPE	1 ISTA	33 IEND	153 JSTA	193 JEND	0 NDATA
1	1	1003	0	0	0	0	0
MSEQ 3	MGFLAG 1	ICONSF 0	MTT 0	NGAM 02			
ISSC	EPSSSC(1)	EPSSSC(2)	EPSSSC(3)	ISSR		EPSSSR(2)	
0 NCYC	0.3 MGLEVG	0.3 NEMGL	0.3 NITFO	0	0.3	0.3	0.3
200	01	00	000				
200 200	02 03	00 00	000 000				
MIT1	MIT2	MIT3	MIT4	MIT5			
01	01	01	01	01			
01 01	01 01	01 01	01 01	01 01			
	CKING DATA:	:					
NBLI 2							
NUMBER (	RTD .	ISTA 1	JSTA KST 1	A IEND 1 33	JEND KI 41	END ISVA1 1 1	ISVA2 2
2 NUMBER (	1 	21 TSTA	1 41 JSTA KST	1 33 A TEND	97 JEND KI	1 1	2 TSV22
1	1	1	193 153	1 33	153		2
2 PATCH SUF	I RFACE DATA:	21	153	1 33	97	T T	Z
NINTER 0							
PLOT3D OU	JTPUT:						
GRID IPI MOVIE	IYPE ISTAR	[ IEND	IINC JSTAR	T JEND	JINC KST	ART KEND	KINC
0 PRINT OUT	г.						
GRID IPT	TYPE ISTAR	r iend	IINC JSTAR	T JEND	JINC KST	ART KEND	KINC
CONTROL S NCS	SURFACE:						
0	יייתיד דיסגיו	חינייטד (		פייז פייז	דיייזיד רואק	TNODM	
GKID ISI	iaki ieni	JSIART	JEND K	SIAKI KI	ТМАГГ ЛИЗ	INORM	

After this test case is run, the convergence history, found in file cfl3d.res, should look like that plotted in Figure 9-24. The two sharp spikes in the residual history are at the iterations at which the grid levels change in the mesh sequencing procedure.



#### 9.2.3 Onera M-6 Wing

In this case, a turbulent Navier-Stokes computation is performed over the Onera M-6 wing, on a coarse grid, using a grid in PLOT3D-type format is performed. The grid consists of a single grid zone with a C-O mesh topology and is composed of 41,225 points. (Keep in mind that one needs a grid at least double this size in each direction, e.g.  $193 \times 49 \times 33$  or larger, to actually resolve the flow. A coarser grid is used here to shorten the test run.) The wing surface grid and wake, as well as the plane of symmetry grid are illustrated in Figure 9-25. The memory requirement for this example is 3.4 million words. A typical timing for this case is 453 CPU seconds on a CRAY YMP (NASA LaRC's Sabre as of September 1996).

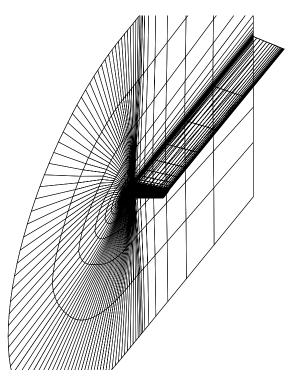


Figure 9-25. Single zone Onera wing surface grid and plane of symmetry grid.

The viscous direction in this PLOT3D-formatted grid is taken as the j direction rather than the k direction as generally recommended. (Due to the order in which CFL3D approximately factors the three index directions, the CFL3D code is usually most efficient when the primary viscous direction is taken as the k direction.) In this case, the convergence is not hurt by the altered directionality. (In some cases, however, it can be!) Note, however, that this case *is* more efficient (CPU timewise) when run on a vector machine with k as the viscous direction, due to the distribution of individual i, j, k index lengths and the way the code is vectorized. For this case, the difference on Sabre is a factor of 17% (with a CFL3D-type k viscous grid, the code runs in 374 seconds as opposed to 453 seconds). It is possible to duplicate this result by changing the hard-wired parameter iplot3d to 0 in the form2bin.f file and using the input file oneram6.inp\_cfl3d instead of oneram6.inp\_p3d. This exercise will demonstrate the differences between PLOT3D-type and CFL3D-type grids, as well as the corresponding differences in the input files.

Besides the CFL3D code the following files are needed to run this test case:

File	<b>Description</b>
oneram6.inp_p3d	input for CFL3D
m6_i97.fmt_p3d	formatted grid
form2bin.f	grid converter

The steps for running this case on the YMP are as follows:

#### <u>Step 1</u>

Compile the grid converter code:

cft77 form2bin.f

# <u>Step 2</u>

Link the grid converter object file:

segldr -o form2bin form2bin.o

#### <u>Step 3</u>

Run the grid converter program to generate the 3-d volume grid (the binary file m6\_i97.grd\_p3d will be output):

form2bin

# <u>Step 4</u>

Use the makefile to compile, link, and create the executable for the precfl3d code (be sure precfl.h is in the current directory):

```
make -f makeprecfl3d_cray
```

#### <u>Step 5</u>

Run the precfl3d code (the cflx.h files will be output):

precfl3d < oneram6.inp\_p3d</pre>

# <u>Step 6</u>

Use the makefile to compile, link, and create the executable for the CFL3D code:

make -f makecfl3d\_cray

# <u>Step 7</u>

Run the CFL3D code:

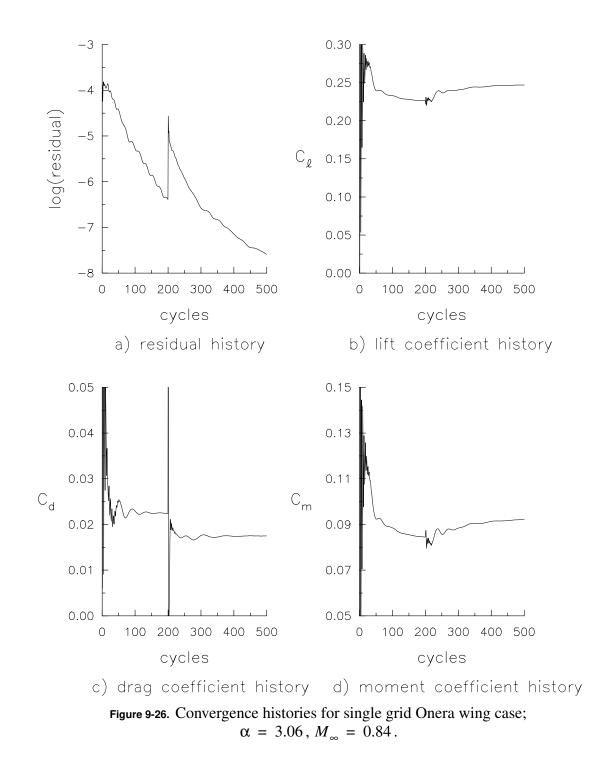
cfl3d < oneram6.inp\_p3d

# The input file for this case is:

,								
	FILES:							
m6_i9								
plot3								
cfl3d		1						
cfl3d								
	.turre	25						
	.bloma							
cfl3d	.out15	5						
cfl3d	.prout	5						
cfl3d	.out20	)						
ovrlp								
patch								
resta								
				e grid, coa			TUTOR	
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			0	0	0	0	0	0
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IF IO: IDIM: JO: JDIM: KO:	AG(I) 1 DS(I) 1 GRID 1 GRID 1 GRID 1 GRID 1 GRID 1 GRID 1 GRID 2	IDIAG(J) 1 IFDS(J) 1 NBCIO 1 SEGMENT	IDIAG(K) 1 IFDS(K) 1 NBCIDIM 1 BCTYPE 1003 BCTYPE 1003 BCTYPE 1003 BCTYPE 1003 BCTYPE 1003 BCTYPE 1001 BCTYPE 0 ICONSF 0	IFLIM(I) 3 RKAP0(I) .3333 NBCJO 3 JSTA 0 JSTA 0 ISTA 1 13 85 ISTA 0 0 ISTA 0 0 0 ISTA 0 0 0 0 I I I I I I I I I I I I I	IFLIM(J) 3 RKAP0(J) .3333 NBCJDIM 1 JEND 0 JEND 0 IEND 0 IEND 0 IEND 0 IEND 0 IEND 0 IEND 0 0 0 0 0 0 0 0 0 0 0 0 0	IFLIM(K) 3 RKAPO(K) .3333 NBCKO 1 KSTA 0 KSTA 0 KSTA 0 KSTA 0 JSTA 0 JSTA 0	NBCKDIM 1 KEND 0 KEND 0 17 0 KEND 0 JEND 0 JEND 0 0 JEND 0 0 JEND 0 0 0 0 0 0 0 0 0 0 0 0 0	0 NDATA 0 NDATA 0 NDATA 0 NDATA 0 NDATA 0 NDATA 0 NDATA 0 NDATA

(	)0 71 )1 )1 LOCKING	02 02 MIT2 01 01 G DATA:	00 00 MIT3 01 01		000 000 MIT4 01 01	MIT5 01 01				
NUMBER	GRID	:	ISTA	JSTA	KSTA	IEND	JEND	KEND	ISVA1	ISVA2
1	1		1	1	1	13	1	17	1	3
2	1		1	1	17	49	25	17	1	2
NUMBER	GRID	:	ISTA	JSTA	KSTA	IEND	JEND	KEND	ISVA1	ISVA2
1	1		97	1		85	1	17	1	3
2	1		97	1	17	49	25	17	1	2
PATCH S NINTE		DATA:								
PLOT3D	OUTPUT	' <b>:</b>								
GRID 1	IPTYPE	ISTART	IEND	IINC	JSTART	JEND	JINC	KSTART	KEND	KINC
1	0	0	0	0	0	0	0	0	0	0
MOVIE 0										
PRINT C										
GRID I CONTROI NCS		-	IEND	IINC	JSTART	JEND	JINC	KSTART	KEND	KINC
0										
GRID 1	ISTART	IEND	JSTART	JEI	ND KST	ART KE	ND IV	VALL IN	ORM	

After this test case is run, the convergence histories, found in file cfl3d.res, should look like those plotted in Figure 9-26. The sharp spikes in the plots indicate the iteration at which the grid level changes in the mesh sequencing process.



# 9.2.4 Delta Wing

The laminar flow over a 75° swept delta wing is solved in this test case. The grid consists of a single grid zone with 156,325 points. (Note that this grid is coarser than what one would normally use to resolve this flow.) The memory requirement for this example is 8.0 million words. A typical timing for this case is 2236 CPU seconds on a CRAY YMP (NASA LaRC's Sabre as of September 1996). The surface grid (k = 1) and a trailing edge grid plane are shown in Figure 9-27.

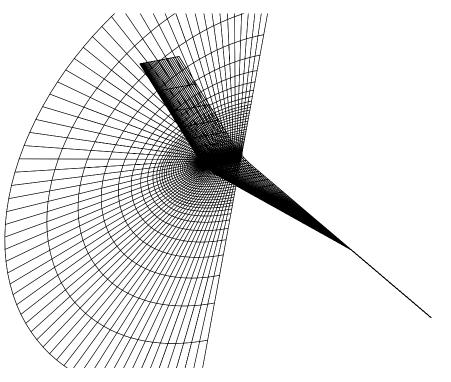


Figure 9-27. Single zone delta wing surface grid and trailing edge plane grid.

Besides the CFL3D code the following files are needed to run this test case:

File	<b>Description</b>
delta.inp	input for CFL3D
delta.fmt	formatted grid
form2bin.f	grid converter

The steps for running this case on the YMP are as follows:

# <u>Step 1</u>

Compile the grid converter code:

cft77 form2bin.f

# Step 2

Link the grid converter object file:

```
segldr -o form2bin form2bin.o
```

#### <u>Step 3</u>

Run the grid converter program to generate the 3-d volume grid (the binary file delta.bin will be output):

form2bin

# <u>Step 4</u>

Use the makefile to compile, link, and create the executable for the precfl3d code (be sure precfl.h is in the current directory):

```
make -f makeprecfl3d_cray
```

#### <u>Step 5</u>

Run the precf13d code (the cflx.h files will be output):

```
precfl3d < delta.inp</pre>
```

#### <u>Step 6</u>

Use the makefile to compile, link, and create the executable for the CFL3D code:

```
make -f makecfl3d_cray
```

#### <u>Step 7</u>

Run the CFL3D code:

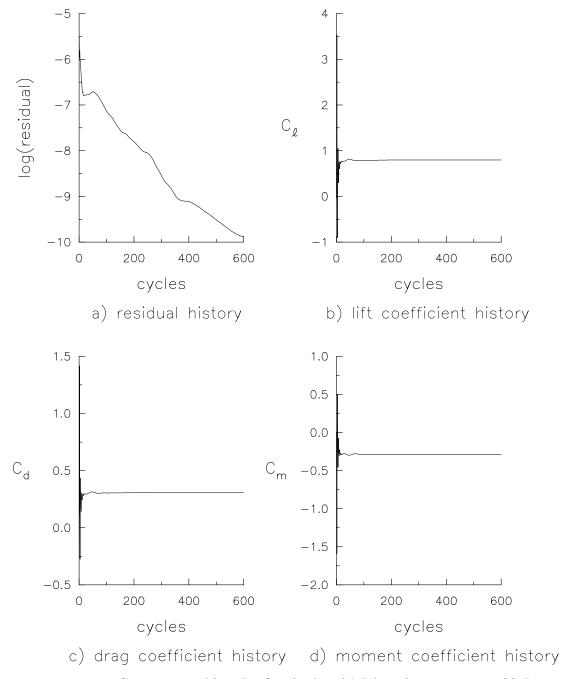
cfl3d < delta.inp

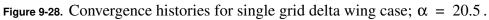
The input file for this case is:

```
I/O FILES
delta.bin
plot3dg.bin
plot3dq.bin
cfl3d.out
cfl3d.res
cfl3d.turres
cfl3d.blomax
cfl3d.out15
cfl3d.prout
cfl3d.out20
ovrlp.bin
patch.bin
restart.bin
 75 Degree Swept Delta Wing - 37x65x65 - Laminar
     XMACH
               ALPHA
                          BETA REUE, MIL
                                            TINF,DR
                                                        IALPH
                                                                   IHIST
```

		0.0	0.500			0		
SREF .13398			XMC 0.25000		ZMC 0.			
DT	IREST	IFLAGTS	FMAX	IUNST	CFLTAU			
-10.0	0	0	1.0		10.			
NGRID 1	NPLOT3D 1	NPRINT 0	NWREST 0100	ICHK 0	I2D 0	NTSTEP 1	ITA 1	
NCG	IEM	IADVANCE			-		±	
2	0	0	IFORCE 001	0	0	1		
IDIM 37	JDIM K 65	KDIM 65						
ILAMLO			JLAMHI	KLAMLO	KLAMHI			
00	00	000	000	0	0000			
INEWG 0	IGRIDC 0	IS 0	JS 0	KS 0	IE O	JE O	KE O	
IDIAG(I)	U IDIAG(J)			IFLIM(J)	IFLIM(K)	0	0	
1	1	1	4	4	4			
IFDS(I) 1	IFDS(J) 1		RKAP0(I) .3333					
GRID	NBCIO	NBCIDIM		NBCJDIM	.3333 NBCK0	NBCKDIM	IOVRLP	
1	1	1	1	1	3	1		
IO: GRID	SEGMENT	BCTYPE		JEND	KSTA 1	KEND	NDATA	
1 IDIM: GRID	1 SEGMENT	1003 BCTYPE	1 JSTA	65 JEND	L KSTA	65 KEND	0 NDATA	
1	1	1003		65	1	65	0	
JO: GRID	SEGMENT	BCTYPE	ISTA	IEND	KSTA	KEND	NDATA	
1 JDIM: GRID	1 SEGMENT	1001 BCTYPE	1 ISTA	37 IEND	1 KSTA	65 KEND	0 NDATA	
1	1	1001		37	1	65	0	
KO: GRID	SEGMENT	BCTYPE	ISTA	IEND	JSTA	JEND	NDATA	
1	1 2		1 9	9 25	1		0 2	
Ŧ	TWTYPE	CQ	2	20	Ŧ	00	2	
	-1.	0.						
1 KDIM: GRID	3 SEGMENT	0 BCTYPE	25 ISTA	37 IEND	1 JSTA		0 NDATA	
1	JEGMENT 1	1003	131A	37	1	65	NDATA 0	
MSEQ	MGFLAG	ICONSF	MTT	NGAM				
1	1 EDSSSC(1)	1			EDCCCD(1)	EDCCCD(2)	EDCCCD(2)	
1350	0.3	0.3	EPSSSC(3) 0.3	1556	LFSSSR(1) 0.3	LF355K(2) 0.3	LF355K(3) 0.3	
NCYC	MGLEVG	NEMGL	NITFO					
600 MTT1				МТТБ	MTTC	мтт7	MTTO	
MIT1 01	01	MIT3 01	MIT4 01	M115 01	MIT6 1	MII / 1	MIT8 1	
	CKING DATA:		• -		_	_	_	
NBLI								
1 NUMBER (	GRID :	ISTA	JSTA KSTA	A IEND	JEND KE	END ISVA1	ISVA2	
1	1	25	1 1	L 37	33	1 1	2	
NUMBER G	GRID :	ISTA	JSTA KSTA 65	A IEND	JEND KE	END ISVA1	ISVA2	
PATCH SUF	FACE DATA:	25	60 -	L 37	33	1 1	Z	
NINTER								
0								
PLOT3D OU GRID IPI		TEND	IINC JSTAR	L TEND	JINC KSTA	ART KEND	KINC	
1	0 0	0	0 (	0 0	0	0 0	0	
MOVIE								
0 PRINT OUT:								
		IEND	IINC JSTAR	[ JEND	JINC KSTA	ART KEND	KINC	
CONTROL S	SURFACE:							
NCS 0								
	TART IEND	) JSTART	JEND KS	START KI	END IWALL	INORM		

After this test case is run, the convergence histories, found in file cfl3d.res, should look like those plotted in Figure 9-28.







# **CHAPTER 10**

Mistakes made in setting up a problem for CFL3D are (hopefully) diagnosed by the code's error checks. The error messages, which are printed to unit 11 should be clear enough so that the user can easily identify and remedy the problem. This chapter contains a few general tips based on common user errors.

Unfortunately, it is sometimes the case with CFL3D (as with any CFD code) that the code blows up, even though the user has apparently done everything right. For example, if the code tries to calculate a negative square root, it will bomb with a floating point error. What to do in such cases is never clear and is often more of an art form than a science. Listed below are some things to try and/or look for when experiencing difficulty with CFL3D (when it blows up for no easily-apparent reason). These limited suggestions for what to try when problems occur are based in part on the code-developers' experiences and in part on what is heard from the user community:

- 1. When the code gives the message "negative volume" or blows up (end-of-file) when trying to read the grid, check the following:
  - (a) Be sure the grid is written correctly, using the right-hand rule (see "The Right-Hand Rule" on page 67) for both x, y, z and i, j, k. Either CFL3D format or PLOT3D/TLNS3D format may be used. See "Grid File" on page 65 for a description of the two formats. Remember to set **ngrid** > 0 for CFL3D format and **ngrid** < 0 for PLOT3D/TLNS3D format.
  - (b) Check the grid near the point where the negative volume is indicated. Make sure there are no grid lines that cross or are positioned incorrectly.
  - (c) When running on a workstation, be sure to use the necessary precision for arrays (many turbulent grids are too fine to use single precision).
  - (d) Be consistent with the precision between the grid and the code (i.e., the grid must be created in double precision if the CFL3D code is "made" that way).
- 2. Check the grid. It seems that 95% of the time, the problem is with the grid. Too much stretching too quickly, poor "quality", badly-skewed cells, etc. can all cause problems. Also be sure that the grid *and all its coarser levels* are of a reasonable size (when using multigrid, a coarser grid level should not be *too* coarse).
- 3. Be sure that the correct direction (either z or y) is "up" according to the type of grid being used. See "Grid File" on page 65.

4. If problems occur when starting a run for a very large grid, debugging on a coarser grid is recommended. Unfortunately, running on a coarser grid level using the mseq flag still requires a large percentage of the full amount of memory. Because this can be frustratingly slow for very big grids, the use of the following tools, found in the Tools subdirectory (see "The Code and Supplementary Files" on page 9), is recommended:

everyotherp3d.f - reads a PLOT3D grid and writes out a coarser (every other grid point) grid

v5inpdoubhalf.f - reads a CFL3D input file and creates a new input file applicable for either half or double the number of grid points

After employing these tools to create a smaller grid and the corresponding input file, rerun precf13d on the new input file and then recompile CFL3D and debug the problem on the coarser grid. The required memory will be substantially smaller than that necessary for the original grid.

- 5. Be sure to save a copy of the original code *before* making any modifications. If an error *is* made, the revision can be compared with the original version to aid with debugging. If additional subroutines are added, be sure they are added correctly to the makefile. If they are merely "tacked on" to the beginning of line 2 in the makefile, then they will not necessarily be recompiled correctly when cfll.h, etc. are changed.
- 6. Peruse the output files early in the computation to make sure that they make sense. For example, if part of the grid is in motion, is it moving at the correct speed and in the expected direction? Do the results "look" as expected? Valuable CPU time can be saved by not running erroneous calculations for hundreds of iterations.
- 7. Check to make sure the boundary conditions are implemented correctly. This includes all 1-to-1 and patched interfaces. For 1-to-1 boundaries, look for "mismatch" in the output files for the "geometric mismatches" that may be a source of error. These numbers should all be close to zero. Small mismatches may be acceptable, but large mismatches (O(1)) most likely indicate that one of the segments on the block boundary has been specified incorrectly, or possibly backwards. For patched or overlapped interfaces, it is quite a bit more difficult to assure valid communication stencils. Check all outputs from ronnie for patched grids and from MaGGiE for overlapped grids.
- 8. Try lowering the CFL number significantly initially and allowing it to "creep" up as the solution progresses. Unless the grid is *really* bad, the CFL number should not have to be set below 0.1 (dt = -0.1). Sometimes it may be necessary to *remain* at low CFL numbers for particularly difficult problems. The optimum CFL number to run at is generally around 5 (dt = -5.0), for most problems on decent grids.
- 9. Make sure the executable is appropriate for the input file, particularly if any changes have been made to it *or* to the grid since the previous run. Run precf13d to make sure the dimensioning is correct in the cfl\*.h files. If precf13d changes them, CFL3D must be recompiled.
- Try employing mesh sequencing. Many people have found this to be beneficial for tough problems! For a description of mesh sequencing, see "Mesh Sequencing" on page 134.

- 11. Peruse all output files (there are a lot of them!) for any clues as to what might be wrong.
- 12. Try restarting from a less difficult, converged case on the same grid. For example, run and converge the configuration at alpha = 0 first; then restart at the higher, more difficult angle of attack (or Mach number or whatever).
- 13. When *really* desperate, try running with first order in space for awhile (**nitfo = ncyc**), then restart with third order from that. Or, try using second order (**rkap0 =** -1) rather than third order (**rkap0 =** +0.333333).
- 14. If multigrid is being used, try turning multigrid off for a short time. If multigrid is *not* being used, turn it *on*! In general, for steady-state problems, multigrid should definitely be used. See "Multigrid" on page 125 for instructions.
- 15. Be sure to use the latest version of the code (as of November 1996, Version 5.0).
- 16. There only seems to be sporadic success when varying **idiag**, **ifds**, or **iflim**. In general, if trouble arises when the default values are used for these variables, then trouble will occur even if they are altered. There are of course exceptions. For example, often hypersonic flow cases run more successfully with flux-vector splitting (**idiag** = 0). Also, sometimes, if a particular configuration is marginally stable (i.e., on the brink of going unsteady), then flux-vector splitting, which has more inherent dissipation than flux-difference splitting, may yield a steady solution while flux-difference splitting may go unsteady. What this means, however, is unclear, since it may be that the real physics of the flow *should* go unsteady. Presumably in such a case, if the grid is refined extensively, even flux-vector splitting should probably go unsteady.
- 17. There is little recent experience with this, but experimentation can be made with **ngam** (type of multigrid cycle), **mitL** (number of iterations on each multigrid level), and **mglevg** (the number of multigrid levels).
- 18. Another feature that is rarely used, but could be experimented with is residual correction and/or smoothing with multigrid. Set **issc** and/or **issr** to 1. (Zero is the default.) These parameters have been known, in isolated instances (particularly for hypersonic cases), to cause a "bombing" solution to work. In fact, some users turn residual correction and smoothing *on* as default. This is *not* recommended however because, for the majority of cases the code developers have seen, these parameters seem to hurt more than help. It probably depends a lot on the type of configuration or case being run.
- 19. Try running a different turbulence model that is more robust (Baldwin-Lomax and Spalart-Allmaras are probably the most robust of the models). Then restart the case from this converged solution using the desired model.
- 20. Monitor cfl3d.turres. Make sure there are either no or relatively few nneg values. Make sure the residual of the turbulent equations is not going *up* gradually over time rather than down. If there are problems, try setting factor lower in the appropriate turbulent model subroutine (such as subroutine spalart when using the Spalart-Allmaras turbulence model).
- 21. Sometimes the order of the indices in the grid can make a difference, since the code performs the approximate factorization (AF) in a particular order. Experience has shown that it is usually best to let the *primary* viscous direction (if there is one), such

as the direction normal to the wing surface, be the k direction. Eventually, at steadystate, the answer should be the same regardless of the index directions, but how *well* the code converges to the steady state can unfortunately sometimes be influenced by the choice.

- 22. When restarting and switching turbulence models, it is sometimes necessary, at least temporarily, to lower the CFL number or time step to get the new solution going. After a time, the CFL number or time step may then be "bumped up" to the desired level.
- 23. Hypersonic cases can often be difficult to start. Some suggestions are:
  - (a) Start with a very low CFL number (on the order of 0.1) and run for a while, then later ramp it up.
  - (b) Use mesh sequencing.
  - (c) Use flux-vector splitting (ifds = 0), at least in the beginning.
  - (d) Restart from a previous similar solution.
  - (e) Experiment with varying mitL (number of iterations on each multigrid level).
- 24. If the residuals near patched boundaries are high (particularly when the grid sizes are very different near the patch), try replacing calls to int2 with calls to int3 instead (int3 employs gradient limiting on the patch stencil).



#### APPENDIX A

# Governing Equations

The computational algorithm employed in CFL3D for the three-dimensional Navier-Stokes code CFL3D is described in Thomas et al.<sup>37</sup> The governing equations, which are the thin-layer approximations to the three-dimensional time-dependent compressible Navier-Stokes equations, can be written in terms of generalized coordinates as

$$\frac{\partial \hat{\mathbf{Q}}}{\partial t} + \frac{\partial (\hat{\mathbf{F}} - \hat{\mathbf{F}}_{v})}{\partial \xi} + \frac{\partial (\hat{\mathbf{G}} - \hat{\mathbf{G}}_{v})}{\partial \eta} + \frac{\partial (\hat{\mathbf{H}} - \hat{\mathbf{H}}_{v})}{\partial \zeta} = 0$$
 (A-1)

A general, three-dimensional transformation between the Cartesian variables (x, y, z) and the generalized coordinated  $(\xi, \eta, \zeta)$  is implied. (See Appendix F for details.) The variable *J* represents the Jacobian of the transformation:

$$J = \frac{\partial(\xi, \eta, \zeta, t)}{\partial(x, y, z, t)}$$
(A-2)

In Equation (A-1),  $\mathbf{Q}$  is the vector of conserved variables, density, momentum, and total energy per unit volume, such that

$$\hat{\mathbf{Q}} = \frac{\mathbf{Q}}{J} = \frac{1}{J} \begin{bmatrix} \rho \\ \rho u \\ \rho v \\ \rho w \\ e \end{bmatrix}$$
(A-3)

The inviscid flux terms are

$$\hat{\mathbf{F}} = \frac{\mathbf{F}}{J} = \frac{1}{J} \begin{bmatrix} \rho U \\ \rho Uu + \xi_x p \\ \rho Uv + \xi_y p \\ \rho Uv + \xi_z p \\ (e+p)U - \xi_t p \end{bmatrix}$$
(A-4)

$$\hat{\mathbf{G}} = \frac{\mathbf{G}}{J} = \frac{1}{J} \begin{bmatrix} \rho V \\ \rho V u + \eta_x p \\ \rho V v + \eta_y p \\ \rho V w + \eta_z p \\ (e+p)V - \eta_t p \end{bmatrix}$$
(A-5)  
$$\hat{\mathbf{H}} = \frac{\mathbf{H}}{J} = \frac{1}{J} \begin{bmatrix} \rho W \\ \rho W u + \zeta_x p \\ \rho W v + \zeta_y p \\ \rho W v + \zeta_z p \\ (e+p)W - \zeta_t p \end{bmatrix}$$
(A-6)

The contravariant velocities are given by

$$U = \xi_x u + \xi_y v + \xi_z w + \xi_t$$
  

$$V = \eta_x u + \eta_y v + \eta_z w + \eta_t$$
  

$$W = \zeta_x u + \zeta_y v + \zeta_z w + \zeta_t$$
  
(A-7)

The viscous flux terms are

$$\hat{\mathbf{F}}_{v} = \frac{\mathbf{F}_{v}}{J} = \frac{1}{J} \begin{bmatrix} 0 \\ \xi_{x}\tau_{xx} + \xi_{y}\tau_{xy} + \xi_{z}\tau_{xz} \\ \xi_{x}\tau_{xy} + \xi_{y}\tau_{yy} + \xi_{z}\tau_{yz} \\ \xi_{x}\tau_{xz} + \xi_{y}\tau_{yz} + \xi_{z}\tau_{zz} \\ \xi_{x}b_{x} + \xi_{y}b_{y} + \xi_{z}b_{z} \end{bmatrix}$$
(A-8)  
$$\hat{\mathbf{G}}_{v} = \frac{\mathbf{G}_{v}}{J} = \frac{1}{J} \begin{bmatrix} 0 \\ \eta_{x}\tau_{xx} + \eta_{y}\tau_{xy} + \eta_{z}\tau_{xz} \\ \eta_{x}\tau_{xy} + \eta_{y}\tau_{yy} + \eta_{z}\tau_{yz} \\ \eta_{x}\tau_{xz} + \eta_{y}\tau_{yz} + \eta_{z}\tau_{zz} \\ \eta_{x}b_{x} + \eta_{y}b_{y} + \eta_{z}b_{z} \end{bmatrix}$$
(A-9)

$$\hat{\mathbf{H}}_{v} = \frac{\mathbf{H}_{v}}{J} = \frac{1}{J} \begin{vmatrix} 0\\ \zeta_{x}\tau_{xx} + \zeta_{y}\tau_{xy} + \zeta_{z}\tau_{xz}\\ \zeta_{x}\tau_{xy} + \zeta_{y}\tau_{yy} + \zeta_{z}\tau_{yz}\\ \zeta_{x}\tau_{xz} + \zeta_{y}\tau_{yz} + \zeta_{z}\tau_{zz}\\ \zeta_{x}b_{x} + \zeta_{y}b_{y} + \zeta_{z}b_{z} \end{vmatrix}$$
(A-10)

The shear stress and hear flux terms are defined in tensor notations (summation convention implied) as

$$\tau_{x_i x_j} = \frac{M_{\infty}}{R e_{\tilde{L}_R}} \left[ \mu \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) + \lambda \frac{\partial u_k}{\partial x_k} \delta_{ij} \right]$$
(A-11)

$$b_{x_i} = u_j \tau_{x_i x_j} - \dot{q}_{x_i}$$
 (A-12)

$$\dot{q}_{x_i} = -\left[\frac{M_{\infty}\mu}{Re_{\tilde{L}_R}Pr(\gamma-1)}\right]\frac{\partial a^2}{\partial x_i}$$
(A-13)

The pressure is obtained by the equation of state for a perfect gas

$$p = (\gamma - 1) \left[ e - \frac{\rho}{2} (u^2 + v^2 + w^2) \right]$$
 (A-14)

The above equations have been nondimensionalized in terms of the free-stream<sup>[1]</sup> density,  $\tilde{\rho}_{\infty}$ , the free-stream speed of sound,  $\tilde{a}_{\infty}$ , and the free-stream molecular viscosity,  $\tilde{\mu}_{\infty}$ . (See Chapter 4.) The chain rule is used to evaluate derivatives with respect to (x, y, z) in terms of  $(\xi, \eta, \zeta)$ . Consistent with the thin-layer assumption, only those derivatives in the direction normal to the wall ( $\zeta$ ) are retained in the shear stress and heat flux terms. Equation (A-1) is closed by the Stokes hypothesis for bulk viscosity ( $\lambda + 2\mu/3=0$ ) and Sutherland's law for molecular viscosity.<sup>45</sup>

The CFL3D code also has the capability to solve the Euler equations, which are obtained when the  $\hat{\mathbf{F}}_{v}$ ,  $\hat{\mathbf{G}}_{v}$ , and  $\hat{\mathbf{H}}_{v}$  terms are omitted from Equation (A-1).

The numerical algorithm uses a semi-discrete finite-volume formulation, resulting in a consistent approximation to the conservation laws in integral form

$$\frac{\partial}{\partial t} \iiint_{V} \mathbf{Q} dV + \iint_{S} \overset{\flat}{\mathbf{f}} \cdot \overset{\flat}{\mathbf{n}} dS = 0$$
(A-15)

<sup>[1]</sup> See the note on page 3 about the usage of the phrase *free stream*.

where  $\dot{f}$  denotes the net flux through a surface *S* with unit normal  $\dot{n}$  containing the (time-invariant) volume *V*. Integration of Equation (A-15) over a control volume bounded by lines of constant  $\xi$ ,  $\eta$ , and  $\zeta$  gives the semi-discrete form

$$\left( \frac{\partial \hat{\mathbf{Q}}}{\partial t} \right)_{i, j, k} + (\hat{\mathbf{F}} - \hat{\mathbf{F}}_{v})_{i+1/2, j, k} - (\hat{\mathbf{F}} - \hat{\mathbf{F}}_{v})_{i-1/2, j, k} + (\hat{\mathbf{G}} - \hat{\mathbf{G}}_{v})_{i, j+1/2, k} - (\hat{\mathbf{G}} - \hat{\mathbf{G}}_{v})_{i, j-1/2, k} + (\hat{\mathbf{H}} - \hat{\mathbf{H}}_{v})_{i, j, k+1/2} - (\hat{\mathbf{H}} - \hat{\mathbf{H}}_{v})_{i, j, k-1/2} = 0$$
(A-16)

where, for convenience,

$$\Delta \xi = \xi_{i+1/2, j, k} - \xi_{i-1/2, j, k} = 1$$
  

$$\Delta \eta = \eta_{i, j+1/2, k} - \eta_{i, j-1/2, k} = 1$$
  

$$\Delta \zeta = \zeta_{i, j, k+1/2} - \zeta_{i, j, k-1/2} = 1$$
  
(A-17)

The discrete values  $\hat{\mathbf{Q}}_{i, j, k}$  are regarded as average values taken over a unit computational cell; similarly, discrete values of  $\hat{\mathbf{F}}$ ,  $\hat{\mathbf{G}}$ , and  $\hat{\mathbf{H}}$  are regarded as face-average values. The convective and pressure terms are differenced using either the upwind flux-difference-splitting technique of Roe<sup>31</sup> or the flux-vector-splitting technique of van Leer.<sup>39</sup> The MUSCL (Monotone Upstream-centered Scheme for Conservation Laws) approach of van Leer<sup>40</sup> is used to determine state-variable interpolations at the cell interfaces. The shear stress and heat transfer terms are centrally differenced.



# Time Advancement

For a nondeforming mesh, Equation (A-1) can be written as

$$\frac{1}{J}\frac{\partial \mathbf{Q}}{\partial t} = R(\mathbf{Q}) \tag{B-1}$$

where

APPENDIX B

$$R = -\left[\frac{\partial(\hat{\mathbf{F}} - \hat{\mathbf{F}}_{v})}{\partial\xi} + \frac{\partial(\hat{\mathbf{G}} - \hat{\mathbf{G}}_{v})}{\partial\eta} + \frac{\partial(\hat{\mathbf{H}} - \hat{\mathbf{H}}_{v})}{\partial\zeta}\right]$$
(B-2)

The time term can be discretized with backward differencing:

$$\frac{(1+\phi)(\mathbf{Q}^{n+1}-\mathbf{Q}^n)-\phi(\mathbf{Q}^n-\mathbf{Q}^{n-1})}{J\Delta t} = R(\mathbf{Q}^{n+1})$$
(B-3)

where the superscripts indicate time level. When  $\phi = 0$  the method is first-order temporally accurate; when  $\phi = 1/2$  the method is second-order accurate. This equation is implicit because the right-hand side is a function of the unknown flow variables at time level n + 1.

The CFL3D code is advanced in time with an implicit approximate-factorization method. The implicit derivatives are written as spatially first-order accurate, which results in block-tridiagonal inversions for each sweep. However, for solutions that utilize FDS the block-tridiagonal inversions are usually further simplified with a diagonal algorithm (with a spectral radius scaling of the viscous terms).

Because of the method which the left-hand side is treated for computational efficiency in steady-state simulations (approximate factorization, first-order accuracy), second-order temporal accuracy is forfeited for unsteady computations. One method for recovering the desired accuracy is through the use of sub-iterations. Two different sub-iteration strategies have been implemented in CFL3D. The first method is termed "pseudo time sub-iteration ( $\tau$ -TS)". The method is also often referred to as the "dual time stepping" method. The other method, termed "physical time sub-iteration (t-TS)," follows Pulliam.<sup>28</sup>

For the  $\tau$ -TS method, a pseudo time term is added to the time-accurate Navier-Stokes equations.

APPENDIX B Time Advancement

$$\frac{1}{J}\frac{\partial \mathbf{Q}}{\partial \tau} + \frac{(1+\phi)(\mathbf{Q}^{n+1} - \mathbf{Q}^n) - \phi(\mathbf{Q}^n - \mathbf{Q}^{n-1})}{J\Delta t} = R(\mathbf{Q}^{n+1})$$
(B-4)

This equation is then discretized and iterated in m, where m is the sub-iteration counter.

$$\frac{(1+\phi')(\mathbf{Q}^{m+1}-\mathbf{Q}^{m})-\phi'(\mathbf{Q}^{m}-\mathbf{Q}^{m-1})}{J\Delta\tau} + (\mathbf{B}-5)$$

$$\frac{(1+\phi)(\mathbf{Q}^{m+1}-\mathbf{Q}^{n})-\phi(\mathbf{Q}^{n}-\mathbf{Q}^{n-1})}{J\Delta t} = R(\mathbf{Q}^{m+1})$$

In Equation (B-4),  $\phi$  and  $\phi'$  govern the order of accuracy of the physical and pseudo time terms, respectively. In practice, the pseudo time term is treated as first order (i.e.,  $\phi'=0$ ), but the general form is shown here for completeness. As  $m \to \infty$ , the pseudo time term vanishes if the sub-iterations converge and  $\mathbf{Q}^{m+1} \to \mathbf{Q}^{n+1}$ . If *R* is linearized with

$$R(\mathbf{Q}^{m+1}) \cong R(\mathbf{Q}^{m}) + \frac{\partial R}{\partial \mathbf{Q}} \Delta \mathbf{Q}^{m}$$
(B-6)

and the quantity  $-(1 + \phi)\mathbf{Q}^m/(J\Delta t)$  is added to both sides of Equation (B-4)), then Equation (B-4) becomes

$$\begin{bmatrix} \left(\frac{1+\phi'}{J\Delta\tau} + \frac{1+\phi}{J\Delta t}\right)I + \delta_{\xi}\mathbf{A} + \delta_{\eta}\mathbf{B} + \delta_{\zeta}\mathbf{C}\end{bmatrix}\Delta\mathbf{Q}^{m} =$$

$$\frac{\phi'\Delta\mathbf{Q}^{m-1}}{J\Delta\tau} + \frac{\phi\Delta\mathbf{Q}^{n-1}}{J\Delta t} - \frac{(1+\phi)(\mathbf{Q}^{m} - \mathbf{Q}^{n})}{J\Delta t} + R(\mathbf{Q}^{m})$$
(B-7)

where

$$\Delta \mathbf{Q}^{m} = \mathbf{Q}^{m+1} - \mathbf{Q}^{m}$$
(B-8)

$$\mathbf{A} = \frac{\partial (\hat{\mathbf{F}} - \hat{\mathbf{F}}_{v})}{\partial \mathbf{Q}}$$
(B-9)

$$\mathbf{B} = \frac{\partial (\hat{\mathbf{G}} - \hat{\mathbf{G}}_{v})}{\partial \mathbf{Q}}$$
(B-10)

$$\mathbf{C} = \frac{\partial (\hat{\mathbf{H}} - \hat{\mathbf{H}}_{v})}{\partial \mathbf{Q}}$$
(B-11)

Equation (B-7) is approximately factored and written in primitive variable form; it is solved as a series of sweeps in each coordinate direction as

$$\begin{bmatrix} \left(\frac{(1+\phi')\mathbf{M}}{J\Delta\tau} + \frac{(1+\phi)\mathbf{M}}{J\Delta t}\right) + \delta_{\xi}\mathbf{A}^{*} \end{bmatrix} \Delta \mathbf{q}' = \frac{\phi'\mathbf{M}\Delta\mathbf{q}^{m-1}}{J\Delta\tau} + \frac{\phi\mathbf{M}\Delta\mathbf{q}^{n-1}}{J\Delta\tau} - \qquad (B-12)$$
$$\frac{(1+\phi)\mathbf{M}(\mathbf{q}^{m}-\mathbf{q}^{n})}{J\Delta t} + R(\mathbf{q}^{m})$$
$$\begin{bmatrix} \left(\frac{(1+\phi')\mathbf{M}}{J\Delta\tau} + \frac{(1+\phi)\mathbf{M}}{J\Delta\tau}\right) + \delta_{\eta}\mathbf{B}^{*} \end{bmatrix} \Delta \mathbf{q}'' = \left(\frac{(1+\phi')\mathbf{M}}{J\Delta\tau} + \frac{(1+\phi)\mathbf{M}}{J\Delta\tau}\right) \Delta \mathbf{q}' \qquad (B-13)$$
$$\begin{bmatrix} \left(\frac{(1+\phi')\mathbf{M}}{J\Delta\tau} + \frac{(1+\phi)\mathbf{M}}{J\Delta\tau}\right) + \delta_{\zeta}\mathbf{C}^{*} \end{bmatrix} \Delta \mathbf{q}^{m} = \left(\frac{(1+\phi')\mathbf{M}}{J\Delta\tau} + \frac{(1+\phi)\mathbf{M}}{J\Delta\tau}\right) \Delta \mathbf{q}'' \qquad (B-14)$$
$$\mathbf{q}^{m+1} = \mathbf{q}^{m} + \Delta \mathbf{q}^{m} \qquad (B-15)$$

where the primitive variables are

$$\mathbf{q} = \begin{bmatrix} \rho \\ u \\ v \\ w \\ p \end{bmatrix}$$
(B-16)

$$\mathbf{M} = \frac{\partial \mathbf{Q}}{\partial \mathbf{q}} \tag{B-17}$$

$$\mathbf{A}^* = \frac{\partial (\hat{\mathbf{F}} - \hat{\mathbf{F}}_v)}{\partial \mathbf{q}}$$
(B-18)

$$\mathbf{B}^* = \frac{\partial (\hat{\mathbf{G}} - \hat{\mathbf{G}}_v)}{\partial \mathbf{q}}$$
(B-19)

$$\mathbf{C}^* = \frac{\partial (\mathbf{\hat{H}} - \mathbf{\hat{H}}_v)}{\partial \mathbf{q}}$$
(B-20)

The quantity  $\Delta \tau$  is based on a constant CFL number set by the input parameter **cfl\_tau** (See "LT5 - Time Step Parameters" on page 21). Multigrid is used to drive  $\Delta \mathbf{q}^m$  to zero in a reasonable number of sub-iterations.

In the t-TS method, Equation (B-3) is merely iterated in m, where m is the sub-iteration counter:

$$\frac{(1+\phi)(\mathbf{Q}^{m+1}-\mathbf{Q}^n)-\phi(\mathbf{Q}^n-\mathbf{Q}^{n-1})}{J\Delta t} = R(\mathbf{Q}^{m+1})$$
(B-21)

The quantity  $-(1 + \phi)\mathbf{Q}^m/(J\Delta t)$  is added to both sides, the residual is linearized, and the equation is approximately factored and written in primitive variable form as

$$\left[\frac{(1+\phi)\mathbf{M}}{J\Delta t} + \delta_{\xi}\mathbf{A}^{*}\right]\Delta\mathbf{q}' = \frac{\phi\mathbf{M}\Delta\mathbf{q}^{n-1}}{J\Delta t} - \frac{(1+\phi)\mathbf{M}(\mathbf{q}^{m}-\mathbf{q}^{n})}{J\Delta t} + R(\mathbf{q}^{m})$$
(B-22)

$$\left[\frac{(1+\phi)\mathbf{M}}{J\Delta t} + \delta_{\eta}\mathbf{B}^{*}\right]\Delta\mathbf{q}^{\prime\prime} = \frac{(1+\phi)\mathbf{M}}{J\Delta t}\Delta\mathbf{q}^{\prime}$$
(B-23)

$$\left[\frac{(1+\phi)\mathbf{M}}{J\Delta t} + \delta_{\zeta}\mathbf{C}^*\right]\Delta\mathbf{q}^m = \frac{(1+\phi)\mathbf{M}}{J\Delta t}\Delta\mathbf{q}^{\prime\prime}$$
(B-24)

$$\mathbf{q}^{m+1} = \mathbf{q}^m + \Delta \mathbf{q}^m \tag{B-25}$$

As  $m \to \infty$ ,  $\mathbf{q}^{m+1} \to \mathbf{q}^{n+1}$ . When only one series of sweeps is performed,  $\mathbf{q}^m = \mathbf{q}^n$  and the standard time-accurate CFL3D scheme is recovered (i.e., no sub-iterations). Unlike the  $\tau$ -TS method, this sub-iteration procedure (Equation (B-22) through Equation (B-25)) utilizes only one time step: the physical time step  $\Delta t$  (= constant).

Prior to the execution of Equation (B-25) in the code, the corrections are constrained in order to maintain the positivity of the thermodynamic variables  $\rho$  and p. For example, the update to pressure is taken as

$$p^{n+1} = p^{n} + \Delta p \left[ 1 + \phi_c \left( \alpha_c + \left| \frac{\Delta p}{p^n} \right| \right) \right]^{-1}$$
(B-26)

whenever  $\Delta p / p^n \le \alpha_c$ . Currently,  $\alpha_c = -0.2$  and  $\phi_c = 2.0$ .

In the limit of  $\Delta p/p^n \to -\infty$ ,  $p^{n+1} \to p^n/2$ . This modification improves the robustness of the method by allowing it to proceed through local transients encountered during the convergence process which would otherwise terminate the calculation.

When running steady-state computations (dt < 0), the time step advanced locally in each cell is related to the input CFL number by

$$\Delta t = \frac{\text{CFL}}{|\nabla\xi|t_1 + |\nabla\eta|t_2 + |\nabla\zeta|t_3}$$
(B-27)

where

$$t_{1} = |\overline{U}| + a + 2|\nabla\xi|(\mu + \mu_{T})\max\left(\frac{4}{3}, \frac{\gamma}{Pr}\right)\frac{M}{Re_{\tilde{L}_{R}}}\frac{1}{\rho}$$

$$t_{2} = |\overline{V}| + a + 2|\nabla\eta|(\mu + \mu_{T})\max\left(\frac{4}{3}, \frac{\gamma}{Pr}\right)\frac{M}{Re_{\tilde{L}_{R}}}\frac{1}{\rho}$$

$$t_{3} = |\overline{W}| + a + 2|\nabla\zeta|(\mu + \mu_{T})\max\left(\frac{4}{3}, \frac{\gamma}{Pr}\right)\frac{M}{Re_{\tilde{L}_{R}}}\frac{1}{\rho}$$
(B-28)

where  $\overline{U} = U/|\nabla\xi|$ ,  $\overline{V} = V/|\nabla\eta|$ ,  $\overline{W} = W/|\nabla\zeta|$  and U, V, and W are defined in Equation (A-7) in Appendix A. The viscous scaling terms (the last term in each equation of Equation (B-28)) are only used when the solution includes viscous terms. They arise from a spectral radius scaling (see Coakley<sup>14</sup>).



#### APPENDIX C

# Spatial Discretization

### C.1 Inviscid Fluxes

The spatial derivatives of the convective and pressure terms are written conservatively as a flux balance across a cell as, for example,

$$(\delta_{\xi} \hat{\mathbf{F}})_{i} = \hat{\mathbf{F}}_{i+\frac{1}{2}} - \hat{\mathbf{F}}_{i-\frac{1}{2}}$$
 (C-1)

where the *i* index denotes a cell-center location and  $i \pm 1/2$  corresponds to a cell-interface location. The interface flux is determined from a state-variable interpolation and a locally one-dimensional flux model. For flux-vector splitting (FVS), Equation (C-1) is split into forward and backward-moving pieces as:

$$(\delta_{\xi}\hat{\mathbf{F}})_{i} = (\delta_{\xi}^{-}\hat{\mathbf{F}}^{+} + \delta_{\xi}^{+}\hat{\mathbf{F}}^{-})_{i} = [\hat{\mathbf{F}}^{+}(\mathbf{q}_{L}) + \hat{\mathbf{F}}^{-}(\mathbf{q}_{R})]_{i+\frac{1}{2}} - [\hat{\mathbf{F}}^{+}(\mathbf{q}_{L}) + \hat{\mathbf{F}}^{-}(\mathbf{q}_{R})]_{i-\frac{1}{2}}$$
(C-2)

For flux-difference splitting (FDS), the interface flux is written as an exact solution to an approximate Riemann problem as:

$$(\boldsymbol{\delta}_{\xi} \hat{\mathbf{F}})_{i} = \frac{1}{2} [\hat{\mathbf{F}}(\mathbf{q}_{L}) + \hat{\mathbf{F}}(\mathbf{q}_{R}) - |\tilde{\mathbf{A}}_{inv}| (\mathbf{q}_{R} - \mathbf{q}_{L})]_{i+\frac{1}{2}}$$

$$- \frac{1}{2} [\hat{\mathbf{F}}(\mathbf{q}_{L}) + \hat{\mathbf{F}}(\mathbf{q}_{R}) - |\tilde{\mathbf{A}}_{inv}| (\mathbf{q}_{R} - \mathbf{q}_{L})]_{i-\frac{1}{2}}$$
(C-3)

Further details on the FVS and FDS methods are given below. In both cases, interpolated values  $\mathbf{q}_L$  and  $\mathbf{q}_R$  at each interface are required. The state variable interpolations determine the resulting accuracy of the scheme. They are constructed from interpolation of the primitive variables. For first-order fully-upwind differencing:

Higher order accuracy is given by the family of interpolations:

$$(\mathbf{q}_{L})_{i+\frac{1}{2}} = \mathbf{q}_{i} + \frac{1}{4} [(1-\kappa)\Delta_{-} + (1+\kappa)\Delta_{+}]_{i}$$

$$(\mathbf{q}_{R})_{i+\frac{1}{2}} = \mathbf{q}_{i+1} - \frac{1}{4} [(1-\kappa)\Delta_{+} + (1+\kappa)\Delta_{-}]_{i+1}$$
(C-5)

where

$$\Delta_{+} \equiv \mathbf{q}_{i+1} - \mathbf{q}_{i}$$

$$\Delta_{-} \equiv \mathbf{q}_{i} - \mathbf{q}_{i-1}$$
(C-6)

The parameter  $\kappa \in [-1,1]$  forms a family of difference schemes.  $\kappa = -1$  corresponds to second-order fully-upwind differencing,  $\kappa = 1/3$  to third-order upwind-biased differencing, and  $\kappa = 1$  to central differencing. Note, however, that setting  $\kappa = 1$  in CFL3D will result in odd-even point decoupling, since there is no artificial dissipation in the code. Also note that "third order" for  $\kappa = 1/3$  is only obtained for one-dimensional flows. For two and three dimensions, the formal order of accuracy is second order.

In practice, the gradients of density and pressure are biased by an average value in order to improve the robustness of the calculation at higher Mach numbers and in the early transient stages of a solution.

$$\Delta^{\rm b}_{+} = \frac{\Delta_{+}}{\mathbf{q}_{i} + \frac{1}{2}\Delta_{+}}$$

$$\Delta^{\rm b}_{-} = \frac{\Delta_{-}}{\mathbf{q}_{i} - \frac{1}{2}\Delta_{-}}$$
(C-7)

And,

$$(\mathbf{q}_{L})_{i+\frac{1}{2}} = \mathbf{q}_{i} + \frac{\mathbf{q}_{i}}{4} [(1-\kappa)\Delta_{-}^{b} + (1+\kappa)\Delta_{+}^{b}]_{i}$$

$$(\mathbf{q}_{R})_{i+\frac{1}{2}} = \mathbf{q}_{i+1} - \frac{\mathbf{q}_{i+1}}{4} [(1-\kappa)\Delta_{+}^{b} + (1+\kappa)\Delta_{-}^{b}]_{i+1}$$
(C-8)

Note that Equation (C-8) is not performed for the gradients of velocity; this is essentially an extrapolation of the log of the density and pressure (first-order expansion).

### C.1.1 Flux Limiting

For solutions with discontinuities (such as shock waves), schemes of order higher than one generally require a flux limiter to avoid numerical oscillations in the solution. CFL3D has several limiter options.

The smooth limiter (iflim = 1) is implemented via

$$(\mathbf{q}_{L})_{i+\frac{1}{2}} = \mathbf{q}_{i} + \left\{ \frac{s}{4} [(1-\kappa s)\Delta_{-} + (1+\kappa s)\Delta_{+}] \right\}_{i}$$

$$(\mathbf{q}_{R})_{i+\frac{1}{2}} = \mathbf{q}_{i+1} - \left\{ \frac{s}{4} [(1-\kappa s)\Delta_{+} + (1+\kappa s)\Delta_{-}] \right\}_{i+1}$$

$$(C-9)$$

where

$$s = \frac{2\Delta_{+}\Delta_{-} + \varepsilon}{\left(\Delta_{+}\right)^{2} + \left(\Delta_{-}\right)^{2} + \varepsilon}$$
(C-10)

and  $\varepsilon$  is a small number ( $\varepsilon = 1 \times 10^{-6}$ ) preventing division by zero in regions of null gradient.

The min-mod limiter (iflim = 2) is implemented via

$$(\mathbf{q}_{L})_{i+\frac{1}{2}} = \mathbf{q}_{i} + \frac{1}{4} [(1-\kappa)\overline{\Delta}_{-} + (1+\kappa)\overline{\Delta}_{+}]_{i}$$

$$(\mathbf{q}_{R})_{i+\frac{1}{2}} = \mathbf{q}_{i+1} - \frac{1}{4} [(1-\kappa)\overline{\Delta}_{+} + (1+\kappa)\overline{\Delta}_{-}]_{i+1}$$

$$(C-11)$$

where

$$\Delta_{-} = \operatorname{minmod}(\Delta_{-}, b\Delta_{+})$$
  
$$\overline{\Delta}_{+} = \operatorname{minmod}(\Delta_{+}, b\Delta_{-})$$
  
(C-12)

$$minmod(x, y) = max\{0, min[x sign(y), by sign(x)]\}sign(x)$$
(C-13)

The parameter b is a compression parameter,  $b = (3 - \kappa)/(1 - \kappa)$ .

The smooth limiter tuned to  $\kappa = 1/3$  (iffim = 3) is implemented as follows.

$$(\mathbf{q}_{L})_{i+\frac{1}{2}} = \mathbf{q}_{i} + \frac{1}{2} (\delta_{L} \mathbf{q})_{i}$$

$$(\mathbf{q}_{R})_{i+\frac{1}{2}} = \mathbf{q}_{i+1} - \frac{1}{2} (\delta_{R} \mathbf{q})_{i+1}$$
(C-14)

where

$$(\delta_L \mathbf{q})_i = \mathbf{I}(\mathbf{q}_{i+1} - \mathbf{q}_i, \mathbf{q}_i - \mathbf{q}_{i-1})$$
  

$$(\delta_R \mathbf{q})_i = \mathbf{I}(\mathbf{q}_i - \mathbf{q}_{i-1}, \mathbf{q}_{i+1} - \mathbf{q}_i)$$
(C-15)

$$I(x, y) = \frac{x(y^2 + 2\varepsilon^2) + y(2x^2 + \varepsilon^2)}{2x^2 - xy + 2y^2 + 3\varepsilon^2}$$
(C-16)

and I is designed to recover the state variable to third-order accuracy in the one-dimensional case in smooth regions of the flow and interpolate without oscillations near discontinuities. The parameter  $\varepsilon^2$  is a small constant of order  $\Delta x^3$  which is used to improve the accuracy near smooth extremum and reduce the nonlinearity of the interpolation in regions of small gradient.

**T** /

#### C.1.2 Flux-Vector Splitting

The flux-vector splitting (FVS) method of van Leer<sup>39</sup> is implemented as follows. The generalized inviscid fluxes  $\hat{\mathbf{F}}$ ,  $\hat{\mathbf{G}}$ , and  $\hat{\mathbf{H}}$ , representing the pressure and convection terms, are upwind differenced by splitting into forward and backward contributions and differencing accordingly. For example, the flux difference in the  $\xi$  direction is

$$\delta_{\xi} \hat{\mathbf{F}} = \delta_{\xi}^{-} \hat{\mathbf{F}}^{+} + \delta_{\xi}^{+} \hat{\mathbf{F}}^{-}$$
(C-17)

where  $\delta_{\xi}^{-}$  and  $\delta_{\xi}^{+}$  denote general backward and forward divided difference operators, respectively. For flux-vector splitting,  $\hat{\mathbf{F}}$  is split according to the contravariant Mach number in the  $\xi$  direction, defined as

$$M_{\xi} = \frac{\overline{U}}{a} \tag{C-18}$$

where

$$\overline{U} = \frac{U}{|\nabla\xi|}$$
(C-19)

and

$$U = \xi_x u + \xi_y v + \xi_z w + \xi_t$$
(C-20)

For locally supersonic flow, where  $|M_{\xi}| \ge 1$ ,

$$\hat{\mathbf{F}}^{+} = \hat{\mathbf{F}}$$
  $\hat{\mathbf{F}}^{-} = 0$   $M_{\xi} \ge +1$   
 $\hat{\mathbf{F}}^{-} = \hat{\mathbf{F}}$   $\hat{\mathbf{F}}^{+} = 0$   $M_{\xi} \le -1$  (C-21)

and for locally subsonic flow, where  $|M_{\xi}| < 1$ ,

$$\hat{\mathbf{F}}^{\pm} = \frac{|\nabla\xi|}{J} \begin{bmatrix} \mathbf{F}_{1}^{\pm} \\ \mathbf{F}_{2}^{\pm} \\ \mathbf{F}_{3}^{\pm} \end{bmatrix} = \frac{|\nabla\xi|}{J} \begin{bmatrix} f_{\text{mass}}^{\pm} \\ f_{\text{mass}}^{\pm} [\hat{\xi}_{x}(-\overline{U} \pm 2a)/\gamma + u] \\ f_{\text{mass}}^{\pm} [\hat{\xi}_{y}(-\overline{U} \pm 2a)/\gamma + v] \\ f_{\text{mass}}^{\pm} [\hat{\xi}_{z}(-\overline{U} \pm 2a)/\gamma + w] \end{bmatrix}$$
(C-22)

where the direction cosines of the cell interfaces are

$$\hat{\xi}_{x} = \frac{\xi_{x}}{|\nabla\xi|} \qquad \hat{\xi}_{y} = \frac{\xi_{y}}{|\nabla\xi|} \qquad \hat{\xi}_{z} = \frac{\xi_{z}}{|\nabla\xi|}$$
(C-23)

and

$$f_{\rm mass}^{\pm} = \pm \frac{\rho a}{4} (M_{\xi} \pm 1)^2$$
 (C-24)

$$f_{\text{energy}}^{\pm} = f_{\text{mass}}^{\pm} \left[ \frac{(1-\gamma)\overline{U}^2 \pm 2(\gamma-1)\overline{U}a + 2a^2}{(\gamma^2 - 1)} + \frac{(u^2 + v^2 + w^2)}{2} - \frac{\hat{\xi}_t}{\gamma} (-\overline{U} \pm 2a) \right]$$
(C-25)

and

$$\hat{\xi}_t = \frac{\xi_t}{|\nabla \xi|} \tag{C-26}$$

For an implicit scheme, the Jacobians of the right-hand-side flux terms must be determined (see Appendix B). They appear as terms in the left-hand-side matrix multiplying  $\Delta q$ , and the resulting system of equations is solved with a full 5 by 5 block-tridiagonal inversion procedure. For flux-vector splitting, the split flux Jacobians (with respect to the

primitive variables) for the  $\xi$  direction, i.e.  $\frac{\partial \hat{F}^{\pm}}{\partial q}$  , are

$$\frac{\partial F_{1}^{\pm}}{\partial \rho} = \frac{1}{4} (\overline{U} \pm a) \left[ -1 \pm \frac{3}{2a} (\overline{U} \pm a) \right]$$

$$\frac{\partial F_{1}^{\pm}}{\partial u} = \pm \frac{\rho}{2a} (\overline{U} \pm a) \hat{\xi}_{x}$$

$$\frac{\partial F_{1}^{\pm}}{\partial v} = \pm \frac{\rho}{2a} (\overline{U} \pm a) \hat{\xi}_{y}$$

$$\frac{\partial F_{1}^{\pm}}{\partial w} = \pm \frac{\rho}{2a} (\overline{U} \pm a) \hat{\xi}_{z}$$

$$\frac{\partial F_{1}^{\pm}}{\partial p} = \pm \frac{\rho}{4a^{2}} (\overline{U} \pm a) \left[ -1 \pm \frac{1}{2a} (\overline{U} \pm a) \right]$$
(C-27)

$$\frac{\partial F_{2}^{\pm}}{\partial \rho} = \pm \frac{1}{4} (\overline{U} \pm a) \left[ \frac{3\overline{U}}{2a\gamma} (-\overline{U} \pm a) \hat{\xi}_{x} + u \left( \frac{3}{2a} (\overline{U} \pm a) \mp 1 \right) \right]$$

$$\frac{\partial F_{2}^{\pm}}{\partial u} = \pm \frac{\rho}{4a} (\overline{U} \pm a) \left[ \hat{\xi}_{x} \left( \frac{3\hat{\xi}_{x}}{\gamma} (-\overline{U} \pm a) + 2u \right) + (\overline{U} \pm a) \right]$$

$$\frac{\partial F_{2}^{\pm}}{\partial v} = \pm \frac{\rho}{4a} (\overline{U} \pm a) \hat{\xi}_{y} \left[ \frac{3\hat{\xi}_{x}}{\gamma} (-\overline{U} \pm a) + 2u \right]$$

$$\frac{\partial F_{2}^{\pm}}{\partial w} = \pm \frac{\rho}{4a} (\overline{U} \pm a) \hat{\xi}_{z} \left[ \frac{3\hat{\xi}_{x}}{\gamma} (-\overline{U} \pm a) + 2u \right]$$

$$\frac{\partial F_{2}^{\pm}}{\partial w} = \pm \frac{\rho}{4a} (\overline{U} \pm a) \hat{\xi}_{z} \left[ \frac{3\hat{\xi}_{x}}{\gamma} (-\overline{U} \pm a) + 2u \right]$$

$$\frac{\partial F_{2}^{\pm}}{\partial p} = \pm \frac{1}{4a^{2}} (\overline{U} \pm a) \left\{ \hat{\xi}_{x} \left[ 2a - \frac{\overline{U}}{2a} (-\overline{U} \pm a) \pm \gamma u \left( 1 \mp \frac{1}{2a} (\overline{U} \pm a) \right) \right] \right\}$$

$$\begin{split} \frac{\partial F_{3}^{\pm}}{\partial \rho} &= \pm \frac{1}{4} (\overline{U} \pm a) \Big[ \frac{3\overline{U}}{2a\gamma} (-\overline{U} \pm a) \hat{\xi}_{y} + v \Big( \frac{3}{2a} (\overline{U} \pm a) \mp 1 \Big) \Big] \\ \frac{\partial F_{3}^{\pm}}{\partial u} &= \pm \frac{\rho}{4a} (\overline{U} \pm a) \hat{\xi}_{x} \Big[ \frac{3\hat{\xi}_{y}}{\gamma} (-\overline{U} \pm a) + 2v \Big] \\ \frac{\partial F_{3}^{\pm}}{\partial v} &= \pm \frac{\rho}{4a} (\overline{U} \pm a) \Big[ \hat{\xi}_{y} \Big( \frac{3\hat{\xi}_{y}}{\gamma} (-\overline{U} \pm a) + 2v \Big) + (\overline{U} \pm a) \Big] \\ \frac{\partial F_{3}^{\pm}}{\partial v} &= \pm \frac{\rho}{4a} (\overline{U} \pm a) \hat{\xi}_{z} \Big[ \frac{3\hat{\xi}_{y}}{\gamma} (-\overline{U} \pm a) + 2v \Big] \\ \frac{\partial F_{3}^{\pm}}{\partial \rho} &= \pm \frac{1}{4a^{2}} (\overline{U} \pm a) \hat{\xi}_{z} \Big[ \frac{3\hat{\xi}_{y}}{\gamma} (-\overline{U} \pm a) \pm 2v \Big] \\ \frac{\partial F_{4}^{\pm}}{\partial \rho} &= \pm \frac{1}{4a^{2}} (\overline{U} \pm a) \hat{\xi}_{z} \Big[ \frac{3\hat{\xi}_{y}}{\gamma} (-\overline{U} \pm a) \pm 2v \Big] \\ \frac{\partial F_{4}^{\pm}}{\partial \rho} &= \pm \frac{1}{4a^{2}} (\overline{U} \pm a) \Big[ \hat{\xi}_{y} \Big[ 2a - \frac{\overline{U}}{2a} (-\overline{U} \pm a) \pm \gamma v \Big( 1 \mp \frac{1}{2a} (\overline{U} \pm a) \Big) \Big] \Big] \\ \frac{\partial F_{4}^{\pm}}{\partial \mu} &= \pm \frac{\rho}{4a} (\overline{U} \pm a) \hat{\xi}_{x} \Big[ \frac{3\hat{\xi}_{z}}{\gamma} (-\overline{U} \pm a) \pm 2w \Big] \\ \frac{\partial F_{4}^{\pm}}{\partial u} &= \pm \frac{\rho}{4a} (\overline{U} \pm a) \hat{\xi}_{y} \Big[ \frac{3\hat{\xi}_{z}}{\gamma} (-\overline{U} \pm a) + 2w \Big] \\ \frac{\partial F_{4}^{\pm}}{\partial v} &= \pm \frac{\rho}{4a} (\overline{U} \pm a) \hat{\xi}_{y} \Big[ \frac{3\hat{\xi}_{z}}{\gamma} (-\overline{U} \pm a) + 2w \Big] \\ \frac{\partial F_{4}^{\pm}}{\partial v} &= \pm \frac{\rho}{4a} (\overline{U} \pm a) \hat{\xi}_{z} \Big[ \frac{3\hat{\xi}_{z}}{\gamma} (-\overline{U} \pm a) + 2w \Big] \\ \frac{\partial F_{4}^{\pm}}{\partial v} &= \pm \frac{\rho}{4a} (\overline{U} \pm a) \hat{\xi}_{z} \Big[ 2a - \frac{\overline{U}}{2a} (-\overline{U} \pm a) + 2w \Big] \\ \frac{\partial F_{4}^{\pm}}{\partial v} &= \pm \frac{\rho}{4a} (\overline{U} \pm a) \Big[ \hat{\xi}_{z} \Big[ 2a - \frac{\overline{U}}{2a} (-\overline{U} \pm a) \pm 2w \Big] + (\overline{U} \pm a) \Big] \Big] \\ \frac{\partial F_{4}^{\pm}}}{\partial p} &= \pm \frac{1}{4a^{2}} (\overline{U} \pm a) \Big\{ \hat{\xi}_{z} \Big[ 2a - \frac{\overline{U}}{2a} (-\overline{U} \pm a) \pm 2w \Big] + (\overline{U} \pm a) \Big] \Big\} \\ \frac{\partial F_{5}^{\pm}}}{\partial \rho} &= \pm \frac{1}{4a^{2}} (\overline{U} \pm a) \Big\{ \hat{\xi}_{z} \Big[ 2a - \frac{\overline{U}}{2a} (-\overline{U} \pm a) \pm \gamma w \Big( 1 \mp \frac{1}{2a} (\overline{U} \pm a) \Big] \Big\} \Big\} \\ \frac{\partial F_{5}^{\pm}}}{\partial \mu} &= \pm \frac{1}{4} (\overline{U} \pm a) \Big\{ \hat{\xi}_{x} \Big[ \frac{2a^{2}}{\gamma - 1} + \Big( \frac{4\overline{U}}{\gamma + 1} - \frac{3\hat{\xi}_{y}}{\gamma} \Big) - \overline{U} \pm a + u^{2} + v^{2} + w^{2} \Big] + u (\overline{U} \pm a) \Big\} \Big\} \\ \frac{\partial F_{5}^{\pm}}}{\partial v} &= \pm \frac{\rho}{4a} (\overline{U} \pm a) \Big\{ \hat{\xi}_{z} \Big[ \frac{2a^{2}}{\gamma - 1} + \Big( \frac{4\overline{U}}{\gamma + 1} - \frac{3\hat{\xi}_{y}}{\gamma} \Big) - \overline{U} \pm a + u^{2} + v^{2} + w^{2} \Big\} + v (\overline{U} \pm a) \Big\}$$

 $\frac{\partial F_5^\pm}{\partial \rho}$ 

 $\frac{\partial \mathbf{F}_5^{\pm}}{\partial u}$ 

 $\frac{\partial \mathbf{F}_5^{\pm}}{\partial v}$ 

$$\frac{\partial F_5^{\pm}}{\partial p} = \pm \frac{\gamma}{4a} (\overline{U} \pm a) \left\{ \left[ \left( \frac{\overline{U}}{\gamma + 1} - \frac{\hat{\xi}_l}{\gamma} \right) \left( 2a - \frac{\overline{U}}{2a} (-\overline{U} \pm a) \right) + \frac{a}{\gamma^2 - 1} (\overline{U} \pm 3a) + \frac{u^2 + v^2 + w^2}{2} \left( -\frac{1}{2a} (\overline{U} \pm a) \pm 1 \right) \right] \right\}$$

#### C.1.3 Flux-Difference Splitting

In the flux-difference splitting (FDS) method of  $\text{Roe}^{31}$ , the interface flux in the  $\xi$  direction is written

$$\hat{\mathbf{F}}_{i+\frac{1}{2}} = \frac{1}{2} [\hat{\mathbf{F}}(\mathbf{q}_L) + \hat{\mathbf{F}}(\mathbf{q}_R) - |\tilde{\mathbf{A}}_{inv}| (\mathbf{q}_R - \mathbf{q}_L)]_{i+\frac{1}{2}}$$
(C-32)

where  $\tilde{A}_{inv}$  is  $A_{inv}$  evaluated with Roe-averaged variables defined below. ( $A_{inv}$  is the inviscid part of the A matrix defined in Appendix B.) Hence,

$$|\tilde{\mathbf{A}}_{inv}| = |\mathbf{A}_{inv}(\tilde{\mathbf{q}})|$$
 (C-33)

and

$$\tilde{\mathbf{q}} = \tilde{\mathbf{q}}(\mathbf{q}_L, \mathbf{q}_R)$$

$$\mathbf{A}_{inv} = \frac{\partial \hat{\mathbf{F}}}{\partial \mathbf{Q}} = \mathbf{T} \Lambda \mathbf{T}^{-1} = \mathbf{T} (\Lambda^+ + \Lambda^-) \mathbf{T}^{-1}$$

$$\Lambda^{\pm} = \frac{\Lambda \pm |\Lambda|}{2}$$

$$|\mathbf{A}_{inv}| = \mathbf{T} |\Lambda| \mathbf{T}^{-1}$$
(C-34)

The diagonal matrix  $\Lambda$  is the matrix of eigenvalues of  $A_{inv}$ , T is the matrix of right eigenvectors as columns, and  $T^{-1}$  is the matrix of left eigenvectors as rows. They are all evaluated at Roe-averaged values such that

$$\hat{\mathbf{F}}(\mathbf{Q}_R) - \hat{\mathbf{F}}(\mathbf{Q}_L) = \tilde{\mathbf{A}}_{inv}(\mathbf{Q}_R - \mathbf{Q}_L)$$
 (C-35)

is satisfied exactly. The term  $|\tilde{\mathbf{A}}_{inv}|(\mathbf{Q}_R - \mathbf{Q}_L)$  can be written

$$\begin{split} \left|\tilde{\mathbf{A}}_{\mathrm{inv}}\right|(\mathbf{Q}_{R}-\mathbf{Q}_{L}) &\equiv \left|\tilde{\mathbf{A}}_{\mathrm{inv}}\right| \Delta \mathbf{Q} = \begin{bmatrix} \alpha_{4} \\ \tilde{u}\alpha_{4} + \hat{\xi}_{x}\alpha_{5} + \alpha_{6} \\ \tilde{v}\alpha_{4} + \hat{\xi}_{y}\alpha_{5} + \alpha_{7} \\ \tilde{w}\alpha_{4} + \hat{\xi}_{z}\alpha_{5} + \alpha_{8} \end{bmatrix} \text{ (C-36)} \\ \tilde{w}\alpha_{4} + \hat{\xi}_{z}\alpha_{5} + \alpha_{8} \\ \tilde{H}\alpha_{4} + (\tilde{\overline{U}} - \hat{\xi}_{t})\alpha_{5} + \tilde{u}\alpha_{6} + \tilde{v}\alpha_{7} + \tilde{w}\alpha_{8} - \frac{\tilde{a}^{2}\alpha_{1}}{\gamma - 1} \end{bmatrix} \end{split}$$

where

$$\begin{aligned} \alpha_{1} &= \left| \frac{\nabla \xi}{J} \right| \left| \overline{\overline{U}} \right| \left( \Delta \rho - \frac{\Delta p}{\tilde{a}^{2}} \right) \\ \alpha_{2} &= \frac{1}{2\tilde{a}^{2}} \left| \frac{\nabla \xi}{J} \right| \left| \overline{\overline{U}} + \tilde{a} \right| (\Delta p + \tilde{\rho} \tilde{a} \Delta \overline{U}) \\ \alpha_{3} &= \frac{1}{2\tilde{a}^{2}} \left| \frac{\nabla \xi}{J} \right| \left| \overline{\overline{U}} - \tilde{a} \right| (\Delta p - \tilde{\rho} \tilde{a} \Delta \overline{U}) \\ \alpha_{4} &= \alpha_{1} + \alpha_{2} + \alpha_{3} \\ \alpha_{5} &= \tilde{a} (\alpha_{2} - \alpha_{3}) \\ \alpha_{6} &= \left| \frac{\nabla \xi}{J} \right| \left| \overline{\overline{U}} \right| (\tilde{\rho} \Delta u - \hat{\xi}_{x} \tilde{\rho} \Delta \overline{U}) \\ \alpha_{7} &= \left| \frac{\nabla \xi}{J} \right| \left| \overline{\overline{U}} \right| (\tilde{\rho} \Delta v - \hat{\xi}_{y} \tilde{\rho} \Delta \overline{U}) \\ \alpha_{8} &= \left| \frac{\nabla \xi}{J} \right| \left| \overline{\overline{U}} \right| (\tilde{\rho} \Delta w - \hat{\xi}_{z} \tilde{\rho} \Delta \overline{U}) \end{aligned}$$
(C-37)

Here, the notation ~ denotes the following Roe-averaged evaluations:

$$\tilde{\rho} = \sqrt{\rho_L \rho_R}$$

$$\tilde{u} = \frac{u_L + u_R \sqrt{\rho_R / \rho_L}}{1 + \sqrt{\rho_R / \rho_L}}$$

$$\tilde{v} = \frac{v_L + v_R \sqrt{\rho_R / \rho_L}}{1 + \sqrt{\rho_R / \rho_L}}$$

$$\tilde{w} = \frac{w_L + w_R \sqrt{\rho_R / \rho_L}}{1 + \sqrt{\rho_R / \rho_L}}$$

$$\tilde{H} = \frac{H_L + H_R \sqrt{\rho_R / \rho_L}}{1 + \sqrt{\rho_R / \rho_L}}$$

$$\tilde{a}^2 = (\gamma - 1)\tilde{H} - \frac{\tilde{u}^2 + \tilde{v}^2 + \tilde{w}^2}{2}$$
(C-38)

The terms  $\hat{\xi}_x$ ,  $\hat{\xi}_y$ ,  $\hat{\xi}_z$ ,  $\hat{\xi}_t$  and  $\overline{U}$  are defined in Section C.1.2 and

$$\tilde{\overline{U}} = \frac{1}{|\nabla\xi|} (\xi_x \tilde{u} + \xi_y \tilde{v} + \xi_z \tilde{w} + \xi_t)$$
(C-39)

For flux-difference splitting, the true Jacobians are expensive to compute. Two alternatives are employed in CFL3D. When the system of equations is solved with a full 5 by 5 block-tridiagonal inversion procedure (**idiag** = 0), the following approximations for the left-hand-side flux-difference splitting Jacobians are employed:

$$(\mathbf{A}_{inv}^{+})_{i+\frac{1}{2}} = \frac{1}{2} \left[ (\mathbf{A}_{inv})_{i} + \left| \tilde{\mathbf{A}}_{inv} \right|_{i+\frac{1}{2}} \right]$$

$$(\mathbf{A}_{inv}^{-})_{i+\frac{1}{2}} = \frac{1}{2} \left[ (\mathbf{A}_{inv})_{i+1} - \left| \tilde{\mathbf{A}}_{inv} \right|_{i+\frac{1}{2}} \right]$$

$$(C-40)$$

Flux-difference splitting also has an option to employ a diagonal approximation for the left-hand side (idiag = 1). In this method, each of the spatial factors is approximated with a diagonal inversion as

$$\left[\frac{\mathbf{I}}{J\Delta t} + \delta_{\xi} \frac{\partial \mathbf{F}}{\partial \mathbf{Q}}\right] \Delta \mathbf{Q} \approx \mathbf{T} \left[\frac{\mathbf{I}}{J\Delta t} + \delta_{\xi}^{-} \Lambda^{+} + \delta_{\xi}^{+} \Lambda^{-}\right] \mathbf{T}^{-1} \Delta \mathbf{Q}$$
 (C-41)

Because of the repeated eigenvalues of  $\Lambda$ , only three scalar tridiagonal LU decompositions are needed in each direction, resulting in a significant savings in run time.

#### C.2 Viscous Fluxes

The viscous terms, which represent shear stress and heat transfer effects, are discretized with second-order central differences. The second derivatives are treated as differences across cell interfaces of first-derivative terms. Hence, in the  $\xi$  direction for example, the viscous terms are discretized as

$$\delta_{\xi}(\hat{\mathbf{F}}_{v})_{i} = (\hat{\mathbf{F}}_{v})_{i+\frac{1}{2}} - (\hat{\mathbf{F}}_{v})_{i-\frac{1}{2}}$$
(C-42)

The term  $\hat{\mathbf{F}}_{v}$  is given in Equation (A-8) of Appendix A. Using the thin-layer approximation, it can be written as:

$$\hat{\mathbf{F}}_{v} = \frac{M_{\infty}\mu}{Re_{\tilde{L}_{R}}J} \begin{bmatrix} 0 \\ \phi_{1}\frac{\partial u}{\partial\xi} - \xi_{x}\phi_{2} \\ \phi_{1}\frac{\partial v}{\partial\xi} - \xi_{y}\phi_{2} \\ \phi_{1}\frac{\partial w}{\partial\xi} - \xi_{z}\phi_{2} \\ \phi_{1}\left[\delta_{\xi}\left(\frac{|\mathbf{V}|^{2}}{2}\right) + \frac{1}{Pr(\gamma-1)}\delta_{\xi}(a^{2})\right] + (U - \xi_{t})\phi_{2} \end{bmatrix}$$
(C-43)

where

$$\phi_1 = \xi_x^2 + \xi_y^2 + \xi_z^2 \qquad \phi_2 = \left(\xi_x \frac{\partial u}{\partial \xi} + \xi_y \frac{\partial v}{\partial \xi} + \xi_z \frac{\partial w}{\partial \xi}\right)/3 \qquad (C-44)$$

Implemented in a finite-volume approach, Equation (C-43) requires an approximation to the volume at the cell interface  $(1/J)_{i+1/2}$ , which is calculated by averaging the neighboring values.

When the system of equations is solved with a full 5 by 5 block-tridiagonal inversion procedure (**idiag** = 0), the viscous Jacobians (left-hand-side implicit terms) are employed only for the viscous terms in the  $\zeta$  direction (corresponding with the *k* index of the grid). This is a hold-over from the early days of the CFL3D code, when viscous terms were employed in only one direction. Hence, keep in mind that if the **idiag** = 0 option is used when viscous terms are included in either the *j* or *i* directions, the convergence may suffer due to the lack of the appropriate left-hand-side terms. The missing left-hand-side terms have no effect on *converged* solutions, however.

However, almost all runs with CFL3D employ flux-difference splitting with the diagonal left-hand-side option (**idiag** = 1), because flux-difference splitting is generally more accurate for Navier-Stokes computations (see van Leer et al<sup>41</sup>), and **idiag** = 1 is significantly less expensive than **idiag** = 0. For the diagonal method, a spectral radius scaling for the viscous Jacobian matrices is used, similar to that developed by Coakley.<sup>14</sup> In this reference, the true left-hand-side viscous matrix term is replaced by the matrix  $\tilde{v}I$ , where I is the identity matrix,  $\tilde{v} = \tilde{\mu}_{max}/\tilde{\rho}$ , and  $\tilde{\mu}_{max}$  is the largest eigenvalue of the one-dimensional Navier-Stokes equation,  $\tilde{\mu}_{max} = \max(4\tilde{\mu}/3, \gamma\tilde{\mu}/Pr)$ . In CFL3D, nondimensional  $\mu_{max}$  is taken as  $2(1 + \mu_t)$ , where  $\mu_t$  is the nondimensional turbulent eddy viscosity.



#### APPENDIX D

# Multigrid Algorithm

The full-approximation storage (FAS) multigrid algorithm is used to accelerate convergence to steady state (or to accelerate convergence of sub-iterations during a time-accurate computation). A sequence of grids  $G_0, G_1, ..., G_N$  is defined, where  $G_N$  denotes the finest grid, and coarser grids are formed by successively deleting every other grid line in all three coordinate directions. The fine grid serves to damp the high-frequency errors; the coarser grids damp the low-frequency errors. The coarse grids are solved with a forcing function on the right-hand side, arising from restricting the residual from the finer meshes. The forcing function is the relative truncation error between the grids, such that the solution on the coarser meshes are driven by the fine grid residual. The resulting scheme on mesh  $G_i$  is given as

$$N_i \Delta \mathbf{q}_i^c = -[L_i(\mathbf{q}_i^c) - \tau_i] \equiv -R_i$$
(D-1)

where  $\mathbf{q}_i^c$  is the current approximation to the solution on mesh  $G_i$ ,  $N_i$  is the spatiallyfactored implicit matrix, and  $\tau_i$  is the relative truncation error (where  $\tau_N$  is defined to be zero). The relative truncation error is calculated as

$$\tau_i = L_i (I_{i+1}^i \mathbf{q}_{i+1}^c) - \hat{I}_{i+1}^i R_{i+1}$$
(D-2)

The operator  $I_{i+1}^{i}$  is a volume-weighted restriction operator that transfers values on the finer grid to the coarser grid

$$I_{i+1}^{i}\mathbf{q}_{i+1} = \frac{\sum_{i+1}^{\mathbf{q}_{i+1}}}{\sum_{i+1}^{\mathbf{1}}}$$
(D-3)

where the summation is taken over the eight finer grid cells that make up the coarser grid cell. The restriction operator  $\hat{I}_{i+1}^{i}$  represents a summation over the finer grid cells which make up the coarser cell

$$\hat{I}_{i+1}^{i}R_{i+1} = \sum R_{i+1}$$
 (D-4)

The corrections  $V_i$  on the coarser meshes are used to update the finer mesh

$$\begin{aligned} \mathbf{V}_{i} &= \mathbf{q}_{i}^{c} - I_{i+1}^{i} \mathbf{q}_{i+1}^{c} \\ \mathbf{q}_{i}^{c} &\leftarrow \mathbf{q}_{i}^{c} + I_{i-1}^{i} \mathbf{V}_{i-1} \end{aligned} \tag{D-5}$$

where the prolongation operator corresponds to trilinear interpolation.

When correction smoothing (**issc** = 1) is employed, the corrections  $V_i$  are smoothed on  $G_i$ , before prolongation, with a Laplacian-type operator factored into three sweeps. Equation (D-5) is replaced with

$$V_{i} = \mathbf{q}_{i}^{c} - I_{i+1}^{i} \mathbf{q}_{i+1}^{c}$$

$$(I - \varepsilon_{\xi}^{c} \delta_{\xi\xi})(I - \varepsilon_{\eta}^{c} \delta_{\eta\eta})(I - \varepsilon_{\zeta}^{c} \delta_{\zeta\zeta})\tilde{V}_{i} = V_{i}$$

$$\mathbf{q}_{i}^{c} \leftarrow \mathbf{q}_{i}^{c} + I_{i-1}^{i}\tilde{V}_{i-1}$$
(D-6)

The  $\varepsilon_{\xi}^{c}$ ,  $\varepsilon_{\eta}^{c}$ , and  $\varepsilon_{\zeta}^{c}$  coefficients are user-input values. When used, typical values are 0.3. The correction smoothing overcomes an odd-even decoupling sometimes encountered with the FAS algorithm on highly-stretched grids.

When residual smoothing (**issr** = 1) is employed, the same Laplacian-type operator is used to smooth the  $\Delta \mathbf{q}_i^c$  values, just prior to updating the primitive variables on a given grid level.

$$(I - \varepsilon_{\xi}^{r} \delta_{\xi\xi})(I - \varepsilon_{\eta}^{r} \delta_{\eta\eta})(I - \varepsilon_{\zeta}^{r} \delta_{\zeta\zeta}) \tilde{\mathbf{q}}_{i}^{c} = \Delta \mathbf{q}_{i}^{c}$$
(D-7)

The  $\varepsilon_{\xi}^{r}$ ,  $\varepsilon_{\eta}^{r}$ , and  $\varepsilon_{\zeta}^{r}$  coefficients are again user-input values. When used, typical values are 0.3.



#### APPENDIX E

# Primitive Variables

While CFL3D solves the governing equations in conservation law form, only the primitive variables are stored in permanent memory. This necessitates the use of the transformation matrix:

$$\mathbf{M} = \frac{\partial \mathbf{Q}}{\partial \mathbf{q}} \tag{E-1}$$

where

$$\mathbf{Q} = \begin{bmatrix} \rho & \rho u & \rho v & \rho w & e \end{bmatrix}^{\mathrm{T}}$$
(E-2)

$$\mathbf{q} = \begin{bmatrix} \rho & u & v & w & p \end{bmatrix}^{\mathrm{T}}$$
(E-3)

From Equation (A-14),

$$e = \frac{p}{\gamma - 1} + \frac{\rho}{2}(u^2 + v^2 + w^2)$$
(E-4)

Therefore,

$$\mathbf{M} = \begin{bmatrix} \frac{\partial \rho}{\partial \rho} & \frac{\partial \rho}{\partial u} & \frac{\partial \rho}{\partial v} & \frac{\partial \rho}{\partial w} & \frac{\partial \rho}{\partial \rho} \\ \frac{\partial \rho u}{\partial \rho} & \frac{\partial \rho u}{\partial u} & \frac{\partial \rho u}{\partial v} & \frac{\partial \rho u}{\partial w} & \frac{\partial \rho u}{\partial \rho} \\ \frac{\partial \rho v}{\partial \rho} & \frac{\partial \rho v}{\partial u} & \frac{\partial \rho v}{\partial v} & \frac{\partial \rho v}{\partial \rho} \\ \frac{\partial \rho w}{\partial \rho} & \frac{\partial \rho w}{\partial u} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial \rho} \\ \frac{\partial \rho w}{\partial \rho} & \frac{\partial \rho w}{\partial u} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial \rho} \\ \frac{\partial \rho w}{\partial \rho} & \frac{\partial \rho w}{\partial u} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial \rho} \\ \frac{\partial \rho w}{\partial \rho} & \frac{\partial \rho w}{\partial u} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial \rho} \\ \frac{\partial \rho w}{\partial \rho} & \frac{\partial \rho w}{\partial u} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial \rho} \\ \frac{\partial \rho w}{\partial \rho} & \frac{\partial \rho w}{\partial u} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial \rho} \\ \frac{\partial \rho w}{\partial \rho} & \frac{\partial \rho w}{\partial u} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial \rho} \\ \frac{\partial \rho w}{\partial \rho} & \frac{\partial \rho w}{\partial u} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial \rho} \\ \frac{\partial \rho w}{\partial \rho} & \frac{\partial \rho w}{\partial u} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial \rho} \\ \frac{\partial \rho w}{\partial \rho} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial \rho} \\ \frac{\partial \rho w}{\partial \rho} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial \rho} \\ \frac{\partial \rho w}{\partial \rho} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial \rho} \\ \frac{\partial \rho w}{\partial \rho} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial \rho} \\ \frac{\partial \rho w}{\partial \rho} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial \rho} \\ \frac{\partial \rho w}{\partial \rho} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial \rho} \\ \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial \rho} \\ \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial \rho} \\ \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial \rho} \\ \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} \\ \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} \\ \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} \\ \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} \\ \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} \\ \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} \\ \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} \\ \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} \\ \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} \\ \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} \\ \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} & \frac{\partial \rho w}{\partial v} \\ \frac{\partial \rho w}{\partial v}$$

The inverse of M is

$$\mathbf{M}^{-1} = \begin{bmatrix} 1 & 0 & 0 & 0 & 0 \\ -\frac{u}{\rho} & \frac{1}{\rho} & 0 & 0 & 0 \\ -\frac{v}{\rho} & 0 & \frac{1}{\rho} & 0 & 0 \\ -\frac{w}{\rho} & 0 & 0 & \frac{1}{\rho} & 0 \\ \frac{\gamma - 1}{2}(u^2 + v^2 + w^2) - u(\gamma - 1) - v(\gamma - 1) - w(\gamma - 1) & 1 \end{bmatrix}$$

(E-6)



# APPENDIX F Generalized Coordinates

F.1 Navier-Stokes Equations in Cartesian Coordinates

The compressible three-dimensional Navier-Stokes equations, excluding body forces and external heat sources, in Cartesian coordinates are

$$\frac{\partial \mathbf{Q}}{\partial t} + \frac{\partial \mathbf{f}}{\partial x} + \frac{\partial \mathbf{g}}{\partial y} + \frac{\partial \mathbf{h}}{\partial z} = 0$$
 (F-1)

where

$$\mathbf{Q} = \begin{bmatrix} \boldsymbol{\rho} \\ \boldsymbol{\rho} \boldsymbol{u} \\ \boldsymbol{\rho} \boldsymbol{v} \\ \boldsymbol{\rho} \boldsymbol{w} \\ \boldsymbol{e} \end{bmatrix}$$
(F-2)

$$\mathbf{f} = \begin{bmatrix} \rho u \\ \rho u^{2} + p - \tau_{xx} \\ \rho uv - \tau_{xy} \\ \rho uw - \tau_{xz} \\ (e + p)u - u\tau_{xx} - v\tau_{xy} - w\tau_{xz} + q_{x} \end{bmatrix}$$
(F-3)

$$\mathbf{g} = \begin{bmatrix} \rho u v - \tau_{xy} \\ \rho v^2 + p - \tau_{yy} \\ \rho v w - \tau_{yz} \\ (e + p) v - u \tau_{xy} - v \tau_{yy} - w \tau_{yz} + q_y \end{bmatrix}$$
(F-4)

$$\mathbf{h} = \begin{bmatrix} \rho w \\ \rho u w - \tau_{xz} \\ \rho v w - \tau_{yz} \\ \rho w^{2} + p - \tau_{zz} \\ (e + p)w - u\tau_{xz} - v\tau_{yz} - w\tau_{zz} + q_{z} \end{bmatrix}$$
(F-5)

and

$$\tau_{xx} = \frac{2}{3}\mu \left( 2\frac{\partial u}{\partial x} - \frac{\partial v}{\partial y} - \frac{\partial w}{\partial z} \right)$$
(F-6)

$$\tau_{yy} = \frac{2}{3}\mu \left( 2\frac{\partial v}{\partial y} - \frac{\partial u}{\partial x} - \frac{\partial w}{\partial z} \right)$$
(F-7)

$$\tau_{zz} = \frac{2}{3}\mu \left( 2\frac{\partial w}{\partial z} - \frac{\partial u}{\partial x} - \frac{\partial v}{\partial y} \right)$$
(F-8)

$$\tau_{xy} = \mu \left( \frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right) = \tau_{yx}$$
 (F-9)

$$\tau_{xz} = \mu \left( \frac{\partial w}{\partial x} + \frac{\partial u}{\partial z} \right) = \tau_{zx}$$
 (F-10)

$$\tau_{yz} = \mu \left( \frac{\partial v}{\partial z} + \frac{\partial w}{\partial y} \right) = \tau_{zy}$$
 (F-11)

The pressure is defined by the equation of state for an ideal gas:

$$p = (\gamma - 1) \left[ e - \frac{\rho}{2} (u^2 + v^2 + w^2) \right]$$
 (F-12)

# F.2 Transformation to Generalized Coordinates

Now apply the generalized coordinate transformation

$$\begin{aligned} \xi &= \xi(x, y, z, t) \\ \eta &= \eta(x, y, z, t) \\ \zeta &= \zeta(x, y, z, t) \end{aligned} \tag{F-13}$$

From the chain-rule for a function of multiple variables,

$$\frac{\partial}{\partial x} = \xi_x \frac{\partial}{\partial \xi} + \eta_x \frac{\partial}{\partial \eta} + \zeta_x \frac{\partial}{\partial \zeta} + t_x \frac{\partial}{\partial t}$$

$$\frac{\partial}{\partial y} = \xi_y \frac{\partial}{\partial \xi} + \eta_y \frac{\partial}{\partial \eta} + \zeta_y \frac{\partial}{\partial \zeta} + t_y \frac{\partial}{\partial t}$$

$$\frac{\partial}{\partial z} = \xi_z \frac{\partial}{\partial \xi} + \eta_z \frac{\partial}{\partial \eta} + \zeta_z \frac{\partial}{\partial \zeta} + t_z \frac{\partial}{\partial t}$$

$$\frac{\partial}{\partial t} = \xi_t \frac{\partial}{\partial \xi} + \eta_t \frac{\partial}{\partial \eta} + \zeta_t \frac{\partial}{\partial \zeta} + t_t \frac{\partial}{\partial t}$$
(F-14)

where  $\xi_x$ ,  $\eta_x$ ,  $\zeta_x$ ,  $\xi_y$ ,  $\eta_y$ ,  $\zeta_y$ ,  $\xi_z$ ,  $\eta_z$ ,  $\zeta_z$ ,  $\xi_t$ ,  $\eta_t$ ,  $\zeta_t$  are the metrics.

#### F.2.1 Obtaining the Metrics

The metrics are determined as follows. (For a description of the geometrical evaluation of the metrics, see "Geometrical Evaluation of the Metrics" on page 263.) The derivatives of the generalized coordinates can be written

$$d\xi = \xi_x dx + \xi_y dy + \xi_z dz + \xi_t dt$$
  

$$d\eta = \eta_x dx + \eta_y dy + \eta_z dz + \eta_t dt$$
  

$$d\zeta = \zeta_x dx + \zeta_y dy + \zeta_z dz + \zeta_t dt$$
  

$$dt = t_x dx + t_y dy + t_z dz + t_t dt$$
  
(F-15)

or, in matrix form (noting that  $t_x = t_y = t_z = 0$  and  $t_t = 1$ ),

$$\begin{bmatrix} d\xi \\ d\eta \\ d\zeta \\ dt \end{bmatrix} = \begin{bmatrix} \xi_x & \xi_y & \xi_z & \xi_t \\ \eta_x & \eta_y & \eta_z & \eta_t \\ \zeta_x & \zeta_y & \zeta_z & \zeta_t \\ 0 & 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} dx \\ dy \\ dz \\ dt \end{bmatrix}$$
(F-16)

Similarly, the derivatives of the Cartesian coordinates can be written

$$\begin{bmatrix} dx \\ dy \\ dz \\ dt \end{bmatrix} = \begin{bmatrix} x_{\xi} x_{\eta} x_{\zeta} x_{t} \\ y_{\xi} y_{\eta} y_{\zeta} y_{t} \\ z_{\xi} z_{\eta} z_{\zeta} z_{t} \\ 0 & 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} d\xi \\ d\eta \\ d\zeta \\ dt \end{bmatrix}$$
(F-17)

Therefore, it is evident that

-

$$\begin{bmatrix} \xi_x & \xi_y & \xi_z & \xi_t \\ \eta_x & \eta_y & \eta_z & \eta_t \\ \zeta_x & \zeta_y & \zeta_z & \zeta_t \\ 0 & 0 & 0 & 1 \end{bmatrix} = \begin{bmatrix} x_\xi & x_\eta & x_\zeta & x_t \\ y_\xi & y_\eta & y_\zeta & y_t \\ z_\xi & z_\eta & z_\zeta & z_t \\ 0 & 0 & 0 & 1 \end{bmatrix}^{-1}$$
(F-18)

To determine the inverse matrix represented on the right-hand side of Equation (F-18), first the cofactor matrix is found to be:

$$\begin{bmatrix} (y_{\eta}z_{\zeta} - y_{\zeta}z_{\eta}) & -(y_{\xi}z_{\zeta} - y_{\zeta}z_{\xi}) & (y_{\xi}z_{\eta} - y_{\eta}z_{\xi}) & 0\\ -(x_{\eta}z_{\zeta} - x_{\zeta}z_{\eta}) & (x_{\xi}z_{\zeta} - x_{\zeta}z_{\xi}) & -(x_{\xi}z_{\eta} - x_{\eta}z_{\xi}) & 0\\ (x_{\eta}y_{\zeta} - x_{\zeta}y_{\eta}) & -(x_{\xi}y_{\zeta} - x_{\zeta}y_{\xi}) & (x_{\xi}y_{\eta} - x_{\eta}y_{\xi}) & 0\\ CF_{41} & CF_{42} & CF_{43} & CF_{44} \end{bmatrix}$$
(F-19)

where

$$CF_{41} = -x_{t}(y_{\eta}z_{\zeta} - y_{\zeta}z_{\eta}) - y_{t}(x_{\zeta}z_{\eta} - x_{\eta}z_{\zeta}) - z_{t}(x_{\eta}y_{\zeta} - x_{\zeta}y_{\eta})$$

$$CF_{42} = x_{t}(y_{\xi}z_{\zeta} - y_{\zeta}z_{\xi}) + y_{t}(x_{\zeta}z_{\xi} - x_{\xi}z_{\zeta}) + z_{t}(x_{\xi}y_{\zeta} - x_{\zeta}y_{\xi})$$

$$CF_{43} = -x_{t}(y_{\xi}z_{\eta} - y_{\eta}z_{\xi}) - y_{t}(x_{\eta}z_{\xi} - x_{\xi}z_{\eta}) - z_{t}(x_{\xi}y_{\eta} - x_{\eta}y_{\xi})$$

$$CF_{44} = x_{\xi}(y_{\eta}z_{\zeta} - y_{\zeta}z_{\eta}) + x_{\eta}(y_{\zeta}z_{\xi} - y_{\xi}z_{\zeta}) + x_{\zeta}(y_{\xi}z_{\eta} - y_{\eta}z_{\xi})$$
(F-20)

The transpose of the cofactor matrix (i.e. the adjoint matrix) is

$$\begin{bmatrix} (y_{\eta}z_{\zeta} - y_{\zeta}z_{\eta}) & -(x_{\eta}z_{\zeta} - x_{\zeta}z_{\eta}) & (x_{\eta}y_{\zeta} - x_{\zeta}y_{\eta}) & CF_{41} \\ -(y_{\xi}z_{\zeta} - y_{\zeta}z_{\xi}) & (x_{\xi}z_{\zeta} - x_{\zeta}z_{\xi}) & -(x_{\xi}y_{\zeta} - x_{\zeta}y_{\xi}) & CF_{42} \\ (y_{\xi}z_{\eta} - y_{\eta}z_{\xi}) & -(x_{\xi}z_{\eta} - x_{\eta}z_{\xi}) & (x_{\xi}y_{\eta} - x_{\eta}y_{\xi}) & CF_{43} \\ 0 & 0 & 0 & CF_{44} \end{bmatrix}$$
(F-21)

The determinant of the inverse matrix represented on the right-hand side of Equation (F-18) is

$$\begin{vmatrix} x_{\xi} & x_{\eta} & x_{\zeta} & x_{t} \\ y_{\xi} & y_{\eta} & y_{\zeta} & y_{t} \\ z_{\xi} & z_{\eta} & z_{\zeta} & z_{t} \\ 0 & 0 & 0 & 1 \end{vmatrix} = x_{\xi}(y_{\eta}z_{\zeta} - y_{\zeta}z_{\eta}) - x_{\eta}(y_{\xi}z_{\zeta} - y_{\zeta}z_{\xi}) + x_{\zeta}(y_{\xi}z_{\eta} - y_{\eta}z_{\xi})$$
(F-22)

The Jacobian of the transformation is defined to be

$$J = \frac{\partial(\xi, \eta, \zeta, t)}{\partial(x, y, z, t)} = \begin{vmatrix} \xi_x & \xi_y & \xi_z & \xi_t \\ \eta_x & \eta_y & \eta_z & \eta_t \\ \zeta_x & \zeta_y & \zeta_z & \zeta_t \\ 0 & 0 & 0 & 1 \end{vmatrix}$$
(F-23)

Therefore,

$$\frac{1}{J} = \frac{\partial(x, y, z, t)}{\partial(\xi, \eta, \zeta, t)} = \begin{vmatrix} x_{\xi} & x_{\eta} & x_{\zeta} & x_{t} \\ y_{\xi} & y_{\eta} & y_{\zeta} & y_{t} \\ z_{\xi} & z_{\eta} & z_{\zeta} & z_{t} \\ 0 & 0 & 0 & 1 \end{vmatrix}$$
(F-24)

Note that  $CF_{44} = 1/J$ . Now, from the property of an invertible matrix **D**,

$$\mathbf{D}^{-1} = \frac{1}{det(\mathbf{D})} a dj(\mathbf{D})$$
(F-25)

let

$$\mathbf{D} = \begin{bmatrix} x_{\xi} \ x_{\eta} \ x_{\zeta} \ x_{t} \\ y_{\xi} \ y_{\eta} \ y_{\zeta} \ y_{t} \\ z_{\xi} \ z_{\eta} \ z_{\zeta} \ z_{t} \\ 0 \ 0 \ 1 \end{bmatrix}$$
(F-26)

It follows from Equation (F-18) that

$$\mathbf{D}^{-1} = \begin{bmatrix} \xi_{x} & \xi_{y} & \xi_{z} & \xi_{t} \\ \eta_{x} & \eta_{y} & \eta_{z} & \eta_{t} \\ \zeta_{x} & \zeta_{y} & \zeta_{z} & \zeta_{t} \\ 0 & 0 & 0 & 1 \end{bmatrix} = J \begin{bmatrix} (y_{\eta}z_{\zeta} - y_{\zeta}z_{\eta}) & -(x_{\eta}z_{\zeta} - x_{\zeta}z_{\eta}) & (x_{\eta}y_{\zeta} - x_{\zeta}y_{\eta}) & \mathrm{CF}_{41} \\ -(y_{\xi}z_{\zeta} - y_{\zeta}z_{\xi}) & (x_{\xi}z_{\zeta} - x_{\zeta}z_{\xi}) & -(x_{\xi}y_{\zeta} - x_{\zeta}y_{\xi}) & \mathrm{CF}_{42} \\ (y_{\xi}z_{\eta} - y_{\eta}z_{\xi}) & -(x_{\xi}z_{\eta} - x_{\eta}z_{\xi}) & (x_{\xi}y_{\eta} - x_{\eta}y_{\xi}) & \mathrm{CF}_{43} \\ 0 & 0 & 0 & \frac{1}{J} \end{bmatrix}$$
(F-27)

Therefore, the metrics are

$$\begin{aligned} \xi_x &= J(y_{\eta}z_{\zeta} - y_{\zeta}z_{\eta}) \\ \xi_y &= J(x_{\zeta}z_{\eta} - x_{\eta}z_{\zeta}) \\ \xi_z &= J(x_{\eta}y_{\zeta} - x_{\zeta}y_{\eta}) \\ \xi_t &= -x_t\xi_x - y_t\xi_y - z_t\xi_z \\ \eta_x &= J(y_{\zeta}z_{\xi} - y_{\xi}z_{\zeta}) \\ \eta_y &= J(x_{\xi}z_{\zeta} - x_{\zeta}z_{\xi}) \\ \eta_z &= J(x_{\zeta}y_{\xi} - x_{\xi}y_{\zeta}) \\ \eta_t &= -x_t\eta_x - y_t\eta_y - z_t\eta_z \\ \zeta_x &= J(y_{\xi}z_{\eta} - y_{\eta}z_{\xi}) \\ \zeta_y &= J(x_{\eta}z_{\xi} - x_{\xi}z_{\eta}) \\ \zeta_z &= J(x_{\xi}y_{\eta} - x_{\eta}y_{\xi}) \\ \zeta_t &= -x_t\zeta_x - y_t\zeta_y - z_t\zeta_z \end{aligned}$$
(F-28)

#### F.2.2 Applying the Transformation

Now, to apply the transformation to the Navier-Stokes equations represented in Equation (F-1), substitute Equation (F-14) into Equation (F-1) and multiply by 1/J:

$$\frac{1}{J}\frac{\partial \mathbf{Q}}{\partial t} + \frac{1}{J}\frac{\partial \mathbf{Q}}{\partial \xi}\xi_t + \frac{1}{J}\frac{\partial \mathbf{Q}}{\partial \eta}\eta_t + \frac{1}{J}\frac{\partial \mathbf{Q}}{\partial \zeta}\zeta_t +$$
(F-29)  
$$\frac{1}{J}\frac{\partial \mathbf{f}}{\partial t}t_x + \frac{1}{J}\frac{\partial \mathbf{f}}{\partial \xi}\xi_x + \frac{1}{J}\frac{\partial \mathbf{f}}{\partial \eta}\eta_x + \frac{1}{J}\frac{\partial \mathbf{f}}{\partial \zeta}\zeta_x +$$
$$\frac{1}{J}\frac{\partial \mathbf{g}}{\partial \xi}t_y + \frac{1}{J}\frac{\partial \mathbf{g}}{\partial \xi}\xi_y + \frac{1}{J}\frac{\partial \mathbf{g}}{\partial \eta}\eta_y + \frac{1}{J}\frac{\partial \mathbf{g}}{\partial \zeta}\zeta_y +$$
$$\frac{1}{J}\frac{\partial \mathbf{h}}{\partial t}t_z + \frac{1}{J}\frac{\partial \mathbf{h}}{\partial \xi}\xi_z + \frac{1}{J}\frac{\partial \mathbf{h}}{\partial \eta}\eta_z + \frac{1}{J}\frac{\partial \mathbf{h}}{\partial \zeta}\zeta_z = 0$$

(Remember, however, that  $t_x = t_y = t_z = 0$ ). Since

$$\frac{\partial}{\partial \xi} \left( \mathbf{f} \frac{\xi_x}{J} \right) = \mathbf{f} \frac{\partial}{\partial \xi} \left( \frac{\xi_x}{J} \right) + \frac{\xi_x}{J} \frac{\partial \mathbf{f}}{\partial \xi}$$
(F-30)

Then,

$$\frac{\xi_x}{J}\frac{\partial \mathbf{f}}{\partial \xi} = \frac{\partial}{\partial \xi} \left(\mathbf{f}\frac{\xi_x}{J}\right) - \mathbf{f}\frac{\partial}{\partial \xi} \left(\frac{\xi_x}{J}\right)$$
(F-31)

So, Equation (F-29) becomes

(F-32)

$$\begin{aligned} \frac{\partial}{\partial t} \left( \frac{\mathbf{Q}}{J} \right) &- \mathbf{Q} \frac{\partial}{\partial t} \left( \frac{1}{J} \right) \\ &+ \frac{\partial}{\partial \xi} \left( \frac{\mathbf{Q}}{J} \xi_t \right) - \mathbf{Q} \frac{\partial}{\partial \xi} \left( \frac{\xi_t}{J} \right) + \frac{\partial}{\partial \eta} \left( \frac{\mathbf{Q}}{J} \eta_t \right) - \mathbf{Q} \frac{\partial}{\partial \eta} \left( \frac{\eta_t}{J} \right) + \frac{\partial}{\partial \zeta} \left( \frac{\mathbf{Q}}{J} \zeta_t \right) - \mathbf{Q} \frac{\partial}{\partial \zeta} \left( \frac{\xi_t}{J} \right) \\ &+ \frac{\partial}{\partial \xi} \left( \frac{\mathbf{f}}{J} \xi_x \right) - \mathbf{f} \frac{\partial}{\partial \xi} \left( \frac{\xi_x}{J} \right) + \frac{\partial}{\partial \eta} \left( \frac{\mathbf{f}}{J} \eta_x \right) - \mathbf{f} \frac{\partial}{\partial \eta} \left( \frac{\eta_x}{J} \right) + \frac{\partial}{\partial \zeta} \left( \frac{\mathbf{f}}{J} \zeta_x \right) - \mathbf{f} \frac{\partial}{\partial \zeta} \left( \frac{\xi_x}{J} \right) \\ &+ \frac{\partial}{\partial \xi} \left( \frac{\mathbf{g}}{J} \xi_y \right) - \mathbf{g} \frac{\partial}{\partial \xi} \left( \frac{\xi_y}{J} \right) + \frac{\partial}{\partial \eta} \left( \frac{\mathbf{g}}{J} \eta_y \right) - \mathbf{g} \frac{\partial}{\partial \eta} \left( \frac{\eta_y}{J} \right) + \frac{\partial}{\partial \zeta} \left( \frac{\mathbf{g}}{J} \zeta_y \right) - \mathbf{g} \frac{\partial}{\partial \zeta} \left( \frac{\xi_y}{J} \right) \\ &+ \frac{\partial}{\partial \xi} \left( \frac{\mathbf{h}}{J} \xi_z \right) - \mathbf{h} \frac{\partial}{\partial \xi} \left( \frac{\xi_z}{J} \right) + \frac{\partial}{\partial \eta} \left( \frac{\mathbf{h}}{J} \eta_z \right) - \mathbf{h} \frac{\partial}{\partial \eta} \left( \frac{\eta_z}{J} \right) + \frac{\partial}{\partial \zeta} \left( \frac{\mathbf{h}}{J} \zeta_z \right) - \mathbf{h} \frac{\partial}{\partial \zeta} \left( \frac{\xi_z}{J} \right) = 0 \end{aligned}$$

Regrouping,

$$\frac{\partial}{\partial t} \left( \frac{\mathbf{Q}}{J} \right) + \frac{\partial}{\partial \xi} \left[ \frac{1}{J} (\mathbf{f} \boldsymbol{\xi}_{x} + \mathbf{g} \boldsymbol{\xi}_{y} + \mathbf{h} \boldsymbol{\xi}_{z} + \mathbf{Q} \boldsymbol{\xi}_{t}) \right]$$

$$+ \frac{\partial}{\partial \eta} \left[ \frac{1}{J} (\mathbf{f} \boldsymbol{\eta}_{x} + \mathbf{g} \boldsymbol{\eta}_{y} + \mathbf{h} \boldsymbol{\eta}_{z} + \mathbf{Q} \boldsymbol{\eta}_{t}) \right]$$

$$+ \frac{\partial}{\partial \zeta} \left[ \frac{1}{J} (\mathbf{f} \boldsymbol{\zeta}_{x} + \mathbf{g} \boldsymbol{\zeta}_{y} + \mathbf{h} \boldsymbol{\zeta}_{z} + \mathbf{Q} \boldsymbol{\zeta}_{t}) \right]$$

$$- \mathbf{f} \left[ \frac{\partial}{\partial \xi} \left( \frac{\boldsymbol{\xi}_{x}}{J} \right) + \frac{\partial}{\partial \eta} \left( \frac{\boldsymbol{\eta}_{x}}{J} \right) + \frac{\partial}{\partial \zeta} \left( \frac{\boldsymbol{\zeta}_{x}}{J} \right) \right]$$

$$- \mathbf{g} \left[ \frac{\partial}{\partial \xi} \left( \frac{\boldsymbol{\xi}_{y}}{J} \right) + \frac{\partial}{\partial \eta} \left( \frac{\boldsymbol{\eta}_{y}}{J} \right) + \frac{\partial}{\partial \zeta} \left( \frac{\boldsymbol{\zeta}_{y}}{J} \right) \right]$$

$$- \mathbf{h} \left[ \frac{\partial}{\partial \xi} \left( \frac{\boldsymbol{\xi}_{z}}{J} \right) + \frac{\partial}{\partial \eta} \left( \frac{\boldsymbol{\eta}_{z}}{J} \right) + \frac{\partial}{\partial \zeta} \left( \frac{\boldsymbol{\xi}_{z}}{J} \right) \right]$$

$$- \mathbf{Q} \left[ \frac{\partial}{\partial t} \left( \frac{1}{J} \right) + \frac{\partial}{\partial \xi} \left( \frac{\boldsymbol{\xi}_{t}}{J} \right) + \frac{\partial}{\partial \eta} \left( \frac{\boldsymbol{\eta}_{z}}{J} \right) + \frac{\partial}{\partial \eta} \left( \frac{\boldsymbol{\eta}_{z}}{J} \right) = 0$$

Consider the fifth major term in Equation (F-33):

$$\frac{\partial}{\partial \xi} \left(\frac{\xi_x}{J}\right) = \frac{\partial}{\partial \xi} (y_{\eta} z_{\zeta} - y_{\zeta} z_{\eta}) = y_{\eta} z_{\xi\zeta} + y_{\xi\eta} z_{\zeta} - y_{\zeta} z_{\xi\eta} - y_{\xi\zeta} z_{\eta}$$

$$\frac{\partial}{\partial \eta} \left(\frac{\eta_x}{J}\right) = \frac{\partial}{\partial \eta} (y_{\zeta} z_{\xi} - y_{\xi} z_{\zeta}) = y_{\zeta} z_{\xi\eta} + y_{\eta\zeta} z_{\xi} - y_{\xiz\eta\zeta} - y_{\xi\eta} z_{\zeta}$$

$$\frac{\partial}{\partial \zeta} \left(\frac{\zeta_x}{J}\right) = \frac{\partial}{\partial \zeta} (y_{\xi} z_{\eta} + y_{\eta} z_{\xi}) = y_{\xi} z_{\eta\zeta} + y_{\xi\zeta} z_{\eta} - y_{\eta} z_{\xi\zeta} - y_{\eta\zeta} z_{\xi}$$
(F-34)

The summation of the terms in Equation (F-34) is zero. The same can be shown for the sixth, seventh and eighth major terms in Equation (F-33). Therefore, Equation (F-33) can be written

$$\frac{\partial}{\partial t} \left( \frac{\mathbf{Q}}{J} \right) + \frac{\partial}{\partial \xi} \left[ \frac{1}{J} (\mathbf{f} \boldsymbol{\xi}_x + \mathbf{g} \boldsymbol{\xi}_y + \mathbf{h} \boldsymbol{\xi}_z + \mathbf{Q} \boldsymbol{\xi}_t) \right]$$

$$+ \frac{\partial}{\partial \eta} \left[ \frac{1}{J} (\mathbf{f} \boldsymbol{\eta}_x + \mathbf{g} \boldsymbol{\eta}_y + \mathbf{h} \boldsymbol{\eta}_z + \mathbf{Q} \boldsymbol{\eta}_t) \right]$$

$$+ \frac{\partial}{\partial \zeta} \left[ \frac{1}{J} (\mathbf{f} \boldsymbol{\zeta}_x + \mathbf{g} \boldsymbol{\zeta}_y + \mathbf{h} \boldsymbol{\zeta}_z + \mathbf{Q} \boldsymbol{\zeta}_t) \right] = 0$$
(F-35)

When summed, the expression in the second major term in Equation (F-35) can be written

$$\mathbf{f}\boldsymbol{\xi}_{x} + \mathbf{g}\boldsymbol{\xi}_{y} + \mathbf{h}\boldsymbol{\xi}_{z} + \mathbf{Q}\boldsymbol{\xi}_{t} = \begin{bmatrix} \rho U \\ \rho Uu + p\boldsymbol{\xi}_{x} \\ \rho Uv + p\boldsymbol{\xi}_{y} \\ \rho Uw + p\boldsymbol{\xi}_{z} \\ (e+p)U - \boldsymbol{\xi}_{t}p \end{bmatrix} - \begin{bmatrix} 0 \\ \boldsymbol{\xi}_{x}\boldsymbol{\tau}_{xx} + \boldsymbol{\xi}_{y}\boldsymbol{\tau}_{xy} + \boldsymbol{\xi}_{z}\boldsymbol{\tau}_{xz} \\ \boldsymbol{\xi}_{x}\boldsymbol{\tau}_{xy} + \boldsymbol{\xi}_{y}\boldsymbol{\tau}_{yy} + \boldsymbol{\xi}_{z}\boldsymbol{\tau}_{yz} \\ \boldsymbol{\xi}_{x}\boldsymbol{\tau}_{xz} + \boldsymbol{\xi}_{y}\boldsymbol{\tau}_{yz} + \boldsymbol{\xi}_{z}\boldsymbol{\tau}_{zz} \\ \boldsymbol{\xi}_{x}\boldsymbol{k}_{x} + \boldsymbol{\xi}_{y}\boldsymbol{k}_{y} + \boldsymbol{\xi}_{z}\boldsymbol{k}_{z} \end{bmatrix}$$
(F-36)

where

$$U = \xi_x u + \xi_y v + \xi_z w + \xi_t$$

$$b_x = u\tau_{xx} + v\tau_{xy} + w\tau_{xz} + \dot{q}_x$$

$$b_y = u\tau_{xy} + v\tau_{yy} + w\tau_{yz} + \dot{q}_y$$

$$b_z = u\tau_{xz} + v\tau_{yz} + w\tau_{zz} + \dot{q}_z$$
(F-38)

Equation (F-38) can be written compactly using indicial notation as

$$b_{x_i} = u_j \tau_{x_i x_j} - \dot{q}_{x_i}$$
(F-39)

with i = 1, 2, 3 and j = 1, 2, 3 where 1 indicates the x direction, 2 indicates the y direction, and 3 indicates the z direction.

F.3 Navier-Stokes Equation in Generalized Coordinates

Now, let

$$\mathbf{F} = \begin{bmatrix} \rho U \\ \rho Uu + p\xi_x \\ \rho Uv + p\xi_y \\ \rho Uw + p\xi_z \\ (e+p)U - \xi_t p \end{bmatrix} \qquad \mathbf{F}_v = \begin{bmatrix} 0 \\ \xi_x \tau_{xx} + \xi_y \tau_{xy} + \xi_z \tau_{xz} \\ \xi_x \tau_{xy} + \xi_y \tau_{yy} + \xi_z \tau_{yz} \\ \xi_x \tau_{xz} + \xi_y \tau_{yz} + \xi_z \tau_{zz} \\ \xi_x b_x + \xi_y b_y + \xi_z b_z \end{bmatrix}$$
(F-40)

Combining the other terms in a similar manner and letting  $\hat{\mathbf{Q}} = \mathbf{Q}/J$ ,  $\hat{\mathbf{F}} = \mathbf{F}/J$ ,  $\hat{\mathbf{F}}_{\nu} = \mathbf{F}_{\nu}/J$ ,  $\hat{\mathbf{G}} = \mathbf{G}/J$ ,  $\hat{\mathbf{G}}_{\nu} = \mathbf{G}_{\nu}/J$ ,  $\hat{\mathbf{H}} = \mathbf{H}/J$ ,  $\hat{\mathbf{H}}_{\nu} = \mathbf{H}_{\nu}/J$ , Equation (F-35) can be written

$$\frac{\partial \hat{\mathbf{Q}}}{\partial t} + \frac{\partial (\hat{\mathbf{F}} - \hat{\mathbf{F}}_{\nu})}{\partial \xi} + \frac{\partial (\hat{\mathbf{G}} - \hat{\mathbf{G}}_{\nu})}{\partial \eta} + \frac{\partial (\hat{\mathbf{H}} - \hat{\mathbf{H}}_{\nu})}{\partial \zeta} = 0$$
 (F-41)

The terms are as shown in Appendix A, Equations (A-3) through (A-14).

#### F.4 Geometrical Evaluation of the Metrics

By analogy with the integral form of the equations, a geometrical interpretation of the metric terms can be made. The vector  $\nabla k/J$  is the directed area of the cell interface normal to a k = constant coordinate direction (k =  $\xi$ ,  $\eta$ , or  $\zeta$ ). (See Figure F-1.) The unit vector (k<sub>x</sub>, k<sub>y</sub>, k<sub>z</sub>)/ $|\nabla k|$  is composed of the direction cosines of the cell interface and  $|\nabla k|/J$  is the area of the cell interface. Note that

$$\nabla \mathbf{k} = \mathbf{k}_x \hat{i} + \mathbf{k}_y \hat{j} + \mathbf{k}_z \hat{k}$$
  
$$|\nabla \mathbf{k}| = \sqrt{\mathbf{k}_x^2 + \mathbf{k}_y^2 + \mathbf{k}_z^2}$$
 (F-42)

The directed areas are calculated as one-half the vector product of the two diagonal vectors connecting opposite vertex points of a cell face, taken such that the directed area is parallel to the direction of increasing k.

Likewise, the normalized contravariant velocity,  $\overline{U} = U/|\nabla\xi|$ , for example, represents the relative velocity normal to a  $\xi$  interface, where  $-\hat{\xi}_t = -\xi_t/|\nabla\xi|$  is the grid speed normal to the interface. The volume of the cell is 1/J and is determined by sum-

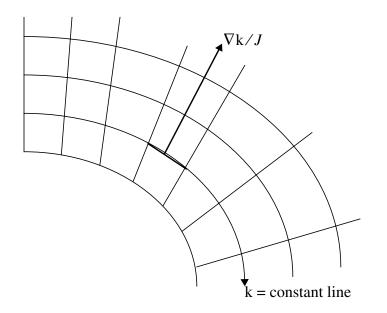


Figure F-1. Directed area schematic.

ming the volumes of the six pentahedra forming the hexagonal cell. Each pentahedron is defined by one of the six cell faces and the average point in the volume. The net effect is that the difference equations are satisfied identically when evaluated at free-stream conditions on arbitrary meshes.

In the code, the metric arrays are set up as follows using the  $\xi$  direction as an example.

$\operatorname{si}(j,k,i,1) = \xi_x /  \nabla \xi  \equiv \hat{\xi}_x$	= $x$ component of unit normal to $i$ face
$si(j,k,i,2) = \xi_y /  \nabla \xi  \equiv \hat{\xi}_y$	= $y$ component of unit normal to $i$ face
si(j,k,i,3) = $\xi_z /  \nabla \xi  \equiv \hat{\xi}_z$	= z component of unit normal to <i>i</i> face
si(j,k,i,4) = $ \nabla \xi /J$	= i face area
$si(j,k,i,5) = -\xi_t /  \nabla \xi $	= i face normal velocity



#### APPENDIX G

# Force and Moment Calculations

# G.1 Forces

The forces are calculated in CFL3D as follows. Let  $\vec{F}_l$  be the total dimensional force acting on the surface element l, with area  $\tilde{s}_l$ , normalized by  $\tilde{q}_{\infty}\tilde{s}_{ref}$ , where

$$\tilde{q}_{\infty} = \frac{1}{2} \tilde{\rho}_{\infty} |\tilde{\mathbf{V}}|_{\infty}^{2}$$
(G-1)

and  $\tilde{s}_{ref}$  is the reference area. In what follows,  $\tilde{s}_{ref}$  and  $\tilde{s}_l$  are taken in terms of grid dimensions. Then the dimensionless force (force coefficient) acting on element l is

$$\vec{F}_{l} = \frac{\vec{F}_{l}}{\tilde{q}_{\infty}\tilde{s}_{ref}}$$
(G-2)

 $\vec{F}_l$  is composed of pressure and viscous components,  $\vec{F}_l^p$  and  $\vec{F}_l^v$ , respectively. The total force coefficient is computed by summing the contributions from all specified surface elements:

$$\vec{F} = \sum_{l} \vec{F}_{l}$$
 (G-3)

### G.1.1 Pressure Component

 $\dot{F}_{l}^{p}$  is normal to the surface.

$$\tilde{s}_{ref}\vec{F}_{l}^{p} = \frac{\tilde{p} - \tilde{p}_{\infty}}{\frac{1}{2}\tilde{\rho}_{\infty}|\tilde{\mathbf{V}}|_{\infty}^{2}}\tilde{s}_{l}\vec{n} = 2\frac{\tilde{p}/(\tilde{\rho}_{\infty}\tilde{a}_{\infty}^{2}) - \tilde{p}_{\infty}/(\tilde{\rho}_{\infty}\tilde{a}_{\infty}^{2})}{|\tilde{\mathbf{V}}|_{\infty}^{2}/\tilde{a}_{\infty}^{2}}\tilde{s}_{l}\vec{n} = 2\frac{p - \frac{1}{\gamma}}{M_{\infty}^{2}}\tilde{s}_{l}\vec{n}$$
(G-4)

Therefore

$$\tilde{s}_{ref}\vec{F}_l^p = \frac{2}{\gamma M_{\infty}^2}(\gamma p - 1)\tilde{s}_l\vec{n}$$
(G-5)

The x, y, and z components of  $\vec{F}_l^p$  are obtained by multiplying Equation (G-5) by the appropriate direction cosine. For example, if element *l* lies on an *i* = constant surface,

$$\vec{\hat{n}} = \nabla \hat{\xi}$$
(G-6)
$$\tilde{s}_{ref}(\vec{\hat{F}}_l^p)_x = \frac{2}{\gamma M_{\infty}^2} (\gamma p - 1) \tilde{s}_l \hat{\xi}_x$$

$$\tilde{s}_{ref}(\vec{\hat{F}}_l^p)_y = \frac{2}{\gamma M_{\infty}^2} (\gamma p - 1) \tilde{s}_l \hat{\xi}_y$$
(G-7)
$$\tilde{s}_{ref}(\vec{\hat{F}}_l^p)_z = \frac{2}{\gamma M_{\infty}^2} (\gamma p - 1) \tilde{s}_l \hat{\xi}_z$$

where  $\hat{\xi}_x = \xi_x / |\nabla \xi|$ ,  $\hat{\xi}_y = \xi_y / |\nabla \xi|$ , and  $\hat{\xi}_z = \xi_z / |\nabla \xi|$ .

### G.1.2 Viscous Component

 $\dot{\mathbf{F}}_{l}^{v}$  is tangential to the surface.

$$\tilde{s}_{ref}\vec{F}_{l}^{v} = \frac{\tilde{\overleftarrow{\tau}}}{\frac{1}{2}\tilde{\rho}_{\infty}|\tilde{\mathbf{V}}|_{\infty}^{2}}\tilde{s}_{l} = 2\frac{\tilde{\overleftarrow{\tau}}/(\tilde{\rho}_{\infty}\tilde{a}_{\infty}^{2})}{|\tilde{\mathbf{V}}|_{\infty}^{2}/\tilde{a}_{\infty}^{2}}\tilde{s}_{l} = \frac{2}{M_{\infty}^{2}}\vec{\overleftarrow{\tau}}\tilde{s}_{l}$$
(G-8)

Consider the flow near a surface element l of an i = constant surface. (See Figure G-1.)

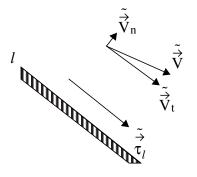


Figure G-1. Viscous force component example.

In the figure,

$$\vec{\tilde{V}} = \text{velocity vector}$$
  
$$\vec{\tilde{V}}_n = \text{normal velocity vector}$$
  
$$\vec{\tilde{V}}_t = \text{tangential velocity vector (0 on the surface)}$$
  
(G-9)

and

$$\vec{\widetilde{V}}_t = \vec{\widetilde{V}} - \vec{\widetilde{V}}_n$$
 (G-10)

$$\tilde{\vec{\tau}}_{l} = \tilde{\mu} \frac{\partial \tilde{\vec{V}}_{t}}{\partial \tilde{n}} = \tilde{\mu} \frac{\partial \left(\vec{V} - \vec{V}_{n}\right)}{\partial \tilde{n}}$$
(G-11)

where

$$\vec{\vec{V}}_n = \left(\vec{\vec{V}} \cdot \vec{\vec{n}}\right)\vec{\vec{n}}$$
(G-12)

 $(\vec{\vec{V}} \cdot \vec{\vec{n}})$  is the normalized contravariant velocity and  $\vec{\vec{n}}$  is the unit surface normal.) The x component of  $\vec{\vec{\tau}}_l$  is  $\vec{\vec{\tau}}_l \cdot \vec{\vec{i}}$ :

$$(\tilde{\vec{\tau}}_l)_x = \tilde{\mu} \frac{\partial \left(\tilde{\vec{\nabla}} \cdot \tilde{\vec{i}} - \tilde{\vec{\nabla}}_n \cdot \tilde{\vec{i}}\right)}{\partial \tilde{n}} = \tilde{\mu} \frac{\partial \left[\tilde{u} - \left(\tilde{\vec{\nabla}} \cdot \tilde{\vec{n}}\right)\hat{\xi}_x\right]}{\partial \tilde{n}}$$
(G-13)

Similarly,

$$(\tilde{\vec{t}}_l)_y = \tilde{\mu} \frac{\partial \left[\tilde{v} - \left(\tilde{\vec{V}} \cdot \tilde{\vec{n}}\right)\hat{\xi}_y\right]}{\partial \tilde{n}}$$
 (G-14)

$$(\tilde{\vec{\tau}}_l)_z = \tilde{\mu} \frac{\partial \left[\tilde{w} - (\vec{V} \cdot \vec{n})\hat{\xi}_z\right]}{\partial \tilde{n}}$$
(G-15)

Consider the nondimensionalization of the x component of  $\dot{\tau}_l$ :

$$(\dot{\tilde{\tau}}_l)_x = \frac{(\tilde{\tilde{\tau}}_l)_x}{\tilde{\rho}_{\infty}\tilde{a}_{\infty}^2} = \frac{1}{\tilde{\rho}_{\infty}\tilde{a}_{\infty}^2} \cdot \tilde{\mu} \frac{\partial \left[\tilde{u} - \left(\vec{\tilde{V}} \cdot \vec{\tilde{n}}\right)\hat{\xi}_x\right]}{\partial \tilde{n}}$$
(G-16)

(See Equation (G-8).) Then

$$(\tilde{\tau}_{l})_{x} = \frac{\tilde{\mu}_{\infty}}{\tilde{\rho}_{\infty} |\tilde{\mathbf{V}}|_{\infty} \tilde{L}_{R}} \frac{|\tilde{\mathbf{V}}|_{\infty}}{\tilde{a}_{\infty}} \frac{\tilde{\mu}}{\tilde{\mu}_{\infty}} \frac{\partial \left\{ \left[ \tilde{u} - \left( \vec{\tilde{\mathbf{V}}} \cdot \vec{\tilde{n}} \right) \hat{\boldsymbol{\xi}}_{x} \right] / \tilde{a}_{\infty} \right\}}{\partial (\tilde{n} / \tilde{L}_{R})}$$
(G-17)

Therefore,

$$(\vec{\tau}_l)_x = \frac{M_{\infty}}{Re_{\tilde{L}_R}} \mu \frac{\partial [u - (\vec{\nabla} \cdot \vec{n})\hat{\xi}_x]}{\partial n}$$
(G-18)

with similar expressions for  $(\dot{\tau}_l)_y$  and  $(\dot{\tau}_l)_z$ . The derivative is evaluated using the cell-center and wall values of a cell volume like that shown in Figure G-2.

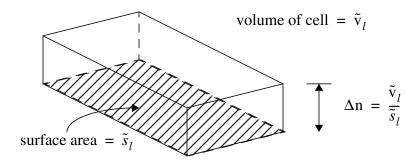


Figure G-2. Cell volume example.

That is, for the *x* component,

$$\frac{\partial \left[\tilde{u} - \left(\tilde{\vec{\nabla}} \cdot \vec{n}\right)\hat{\xi}_x\right]}{\partial n} = \frac{\left[\tilde{u} - \left(\tilde{\vec{\nabla}} \cdot \vec{n}\right)\hat{\xi}_x\right]_c - 0}{\frac{1}{2}\Delta n}$$
(G-19)

where the subscript *c* denotes the cell-center value and  $\left[\tilde{u} - (\tilde{\vec{V}} \cdot \vec{n})\hat{\xi}_x\right] \equiv 0$  on a solid wall with the no-slip assumption. So

$$(\vec{\tau}_l)_x = \frac{M_{\infty}}{Re_{\tilde{L}_R}} \mu \frac{\left[\tilde{u} - \left(\vec{\nabla} \cdot \vec{n}\right) \hat{\xi}_x\right]_c}{\frac{1}{2}\Delta n} = \frac{2M_{\infty}}{Re_{\tilde{L}_R}} \mu \left(\frac{\tilde{s}_l}{\tilde{v}_l}\right) \left[\tilde{u} - \left(\vec{\nabla} \cdot \vec{n}\right) \hat{\xi}_x\right]_c$$
(G-20)

Thus, for the x component of the viscous force,

$$\tilde{s}_{ref}(\vec{F}_l)_x = \frac{2}{M_{\infty}^2} (\vec{\tau}_l)_x \tilde{s}_l = \frac{4}{M_{\infty} R e_{\tilde{L}_R}} \mu [u - (\vec{\nabla} \cdot \vec{n}) \hat{\xi}_x]_c \frac{\tilde{s}_l^2}{\tilde{v}_l}$$
(G-21)

Similarly,

$$\tilde{s}_{ref}(\vec{F}_l^v)_y = \frac{4}{M_{\infty}Re_{\tilde{L}_R}}\mu[v - (\vec{V} \cdot \vec{n})\hat{\xi}_y]_c \frac{\tilde{s}_l^2}{\tilde{v}_l}$$
(G-22)

$$\tilde{s}_{ref}(\vec{F}_l^v)_z = \frac{4}{M_{\infty}Re_{\tilde{L}_R}}\mu[w - (\vec{V} \cdot \vec{n})\hat{\xi}_z]_c \frac{\tilde{s}_l^2}{\tilde{v}_l}$$
(G-23)

## G.2 Moments

The moments due to the forces acting on an element are determined as follows. Let

$$(\vec{F}_l)_x = (\vec{F}_l^p + \vec{F}_l^v)_x$$

$$(\vec{F}_l)_y = (\vec{F}_l^p + \vec{F}_l^v)_y$$

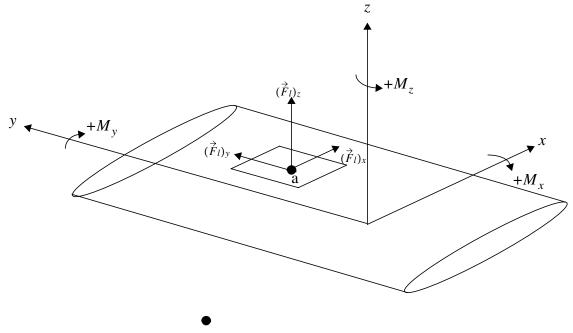
$$(\vec{F}_l)_z = (\vec{F}_l^p + \vec{F}_l^v)_z$$
(G-24)

Figure G-2 illustrates the directions of the moments. All moments are positive if counterclockwise when viewed from the positive axis. The conventions (assuming x points downstream and z points up) are

 $M_x$ : rolling moment; positive for counter-clockwise roll when viewed from downstream.

 $M_{v}$ : pitching moment; positive for pitch up.

 $M_z$ : yawing moment; positive for counter-clockwise yaw when viewed from above. Therefore,



mc (moment center)

Figure G-3. Moment directions.

$$(\vec{M}_l)_x = [+(\vec{F}_l)_x(\tilde{y}_a - \tilde{y}_{mc}) - (\vec{F}_l)_y(\tilde{z}_a - \tilde{z}_{mc})]/\tilde{b}_{ref}$$

$$(\vec{M}_l)_y = [-(\vec{F}_l)_z(\tilde{x}_a - \tilde{x}_{mc}) + (\vec{F}_l)_x(\tilde{z}_a - \tilde{z}_{mc})]/\tilde{c}_{ref}$$

$$(\vec{M}_l)_z = [+(\vec{F}_l)_y(\tilde{x}_a - \tilde{x}_{mc}) - (\vec{F}_l)_x(\tilde{y}_a - \tilde{y}_{mc})]/\tilde{b}_{ref}$$
(G-25)

Note that the reference length used to nondimensionalize  $(\vec{M}_l)_y$  is  $\tilde{c}_{ref}$ , while  $(\vec{M}_l)_z$  is made dimensionless with  $\tilde{b}_{ref}$ . This is the default for CFL3D-type grids or PLOT3D-type grids with **ialph** = 0 (see "LT3 - Flow Conditions" on page 19). However, if a PLOT3Dtype grid is used with **ialph** = 1, then  $(\vec{M}_l)_y$  is made nondimensional with  $\tilde{b}_{ref}$  and  $(\vec{M}_l)_z$ is made dimensionless with  $\tilde{c}_{ref}$ .  $(\vec{M}_l)_x$  is always made dimensionless with  $\tilde{b}_{ref}$ . By switching the reference lengths based on **ialph**, the moment coefficient that is normally associated with the pitching moment is always based on  $\tilde{c}_{ref}$ , while the moment coefficient that is normally associated with the yawing moment is always based on  $\tilde{b}_{ref}$ . Because  $(\vec{M}_l)_x$  always uses  $\tilde{b}_{ref}$ , the moment coefficient associated with the rolling moment is based on  $\tilde{b}_{ref}$ .



#### **APPENDIX H**

CFL3D has several turbulence model capabilities. Appendix H provides the derivation for the advanced turbulence model equations. Be aware that while some variables in this appendix are consistent with the rest of the manual (and are listed in the Nomenclature section), many are introduced, defined, and used only in these sections and may or may not appear in the Nomenclature listing. Also, for simplicity's sake in this Appendix,  $Re = Re_{\tilde{L}_{\nu}}$ .

### H.1 Equations of Motion

Following Wilcox<sup>46</sup>, Favre<sup>19</sup> averaging can be used with the Navier-Stokes equations to account for turbulent fluctuations. The resulting equations of motion can be written using the summation convention as follows. The full Navier-Stokes equations are shown here, but in CFL3D, they are solved as the thin-layer approximation in pre-selected coordinate direction(s). The ~ indicates a dimensional quantity.

$$\frac{\partial \tilde{\rho}}{\partial \tilde{t}} + \frac{\partial}{\partial \tilde{x}_{j}} (\tilde{\rho} \tilde{u}_{j}) = 0$$
(H-1)

$$\frac{\partial}{\partial \tilde{t}} (\tilde{\rho} \tilde{u}_i) + \frac{\partial}{\partial \tilde{x}_j} (\tilde{\rho} \tilde{u}_j \tilde{u}_i) = -\frac{\partial \tilde{p}}{\partial \tilde{x}_i} + \frac{\partial \tilde{\tau}_{ji}}{\partial \tilde{x}_j}$$
(H-2)

$$\frac{\partial}{\partial \tilde{t}} (\tilde{\rho} \tilde{\mathbf{E}}) + \frac{\partial}{\partial \tilde{x}_{j}} (\tilde{\rho} \tilde{u}_{j} \tilde{\mathbf{H}}) = \frac{\partial}{\partial \tilde{x}_{j}} [\tilde{u}_{i} \tilde{\tau}_{ij} - \tilde{\dot{q}}_{j} + \tilde{\psi}_{j}]$$
(H-3)

where

$$\tilde{p} = (\gamma - 1) \left[ \tilde{\rho} \tilde{E} - \frac{1}{2} \tilde{\rho} (\tilde{u}^2 + \tilde{v}^2 + \tilde{w}^2) \right]$$
(H-4)

$$\tilde{E} = \tilde{e} + \frac{1}{2}(\tilde{u}^2 + \tilde{v}^2 + \tilde{w}^2)$$
(H-5)

$$\tilde{H} = \tilde{E} + \frac{\tilde{p}}{\tilde{\rho}}$$
(H-6)

$$\tilde{\dot{q}}_{j} = -\frac{1}{\gamma - 1} \left( \frac{\tilde{\mu}}{\Pr} + \frac{\tilde{\mu}_{T}}{\Pr_{T}} \right) \frac{\partial \tilde{a}^{2}}{\partial \tilde{x}_{j}}$$
(H-7)

$$\tilde{a}^2 = \frac{\gamma \tilde{p}}{\tilde{\rho}} \tag{H-8}$$

Note that the kinetic energy of the fluctuating turbulent field  $\tilde{k}$  is ignored in the definition of  $\tilde{E}$  in CFL3D (it is assumed that  $\tilde{k} \le \tilde{e} + \frac{1}{2}(\tilde{u}^2 + \tilde{v}^2 + \tilde{w}^2)$ ). Define

$$\begin{split} \tilde{\mathbf{S}}_{ij} &= \frac{1}{2} \left( \frac{\partial \tilde{u}_i}{\partial \tilde{x}_j} + \frac{\partial \tilde{u}_j}{\partial \tilde{x}_i} \right) \\ \tilde{\mathbf{W}}_{ij} &= \frac{1}{2} \left( \frac{\partial \tilde{u}_i}{\partial \tilde{x}_j} - \frac{\partial \tilde{u}_j}{\partial \tilde{x}_i} \right) \end{split}$$
(H-9)

Also, the magnitude of vorticity is

$$\tilde{\Omega} = \sqrt{2\tilde{W}_{ij}\tilde{W}_{ij}}$$
 (H-10)

The shear stress terms  $\tilde{\tau}_{ij}$  is composed of a laminar and a turbulent component as

$$\tilde{\tau}_{ij} = \tilde{\tau}_{ij}^{\rm L} + \tilde{\tau}_{ij}^{\rm T}$$
(H-11)

where

$$\tilde{\tau}_{ij}^{L} = 2\tilde{\mu} \left( \tilde{S}_{ij} - \frac{1}{3} \frac{\partial \tilde{u}_{k}}{\partial \tilde{x}_{k}} \delta_{ij} \right)$$
(H-12)

For all *eddy-viscosity* models in CFL3D, the following approximations are made:

$$\tilde{\tau}_{ij}^{\mathrm{T}} = 2\tilde{\mu}_{\mathrm{T}} \left( \tilde{\mathbf{S}}_{ij} - \frac{1}{3} \frac{\partial \tilde{u}_{k}}{\partial \tilde{x}_{k}} \delta_{ij} \right)$$
(H-13)

$$\tilde{\Psi}_j = 0$$
 (H-14)

Currently, for the nonlinear models in CFL3D,

$$\begin{split} \tilde{\tau}_{ij}^{\mathrm{T}} &= 2\tilde{\mu}_{\mathrm{T}} \left( \tilde{\mathrm{S}}_{ij} - \frac{1}{3} \frac{\partial \tilde{u}_{k}}{\partial \tilde{x}_{k}} \delta_{ij} \right) - \frac{2}{3} \tilde{\rho} \tilde{\mathrm{k}} \delta_{ij} + \\ & 2\tilde{\mu}_{\mathrm{T}} K_{1} \frac{\tilde{\mathrm{k}}}{\tilde{\epsilon}} (\tilde{\mathrm{S}}_{ik} \tilde{\mathrm{W}}_{kj} - \tilde{\mathrm{W}}_{ik} \tilde{\mathrm{S}}_{kj}) - \\ & 2\tilde{\mu}_{\mathrm{T}} K_{2} \frac{\tilde{\mathrm{k}}}{\tilde{\epsilon}} \left( \tilde{\mathrm{S}}_{ik} \tilde{\mathrm{S}}_{kj} - \frac{1}{3} \tilde{\mathrm{S}}_{kl} \tilde{\mathrm{S}}_{lk} \delta_{ij} \right) \\ & \tilde{\psi}_{j} &= \left( \tilde{\mu} + \frac{\tilde{\mu}_{\mathrm{T}}}{\sigma_{\mathrm{k}}} \right) \frac{\partial \tilde{\mathrm{k}}}{\partial \tilde{x}_{j}} \end{split}$$
(H-15)

where  $\tilde{k}/\tilde{\epsilon}$  is replaced by  $1/\tilde{\omega}$  when the  $k - \omega$  equations rather than the  $k - \epsilon$  equations are employed.

The Navier-Stokes equations are nondimensionalized and written in generalized coordinates, as described Appendix F. For *eddy-viscosity* models, the end result is that the turbulent Navier-Stokes equations are identical to the laminar equations with the exception that

$$\tilde{\mu}$$
 is replaced by  $\tilde{\mu} + \tilde{\mu}_T$ 

and 
$$\frac{\tilde{\mu}}{Pr}$$
 is replaced by  $\frac{\tilde{\mu}}{Pr} + \frac{\mu_{\rm T}}{Pr_{\rm T}}$ 

where  $\tilde{\mu}_{T}$  is the eddy viscosity value obtained by whatever turbulence model is used, and Pr = 0.72,  $Pr_{T} = 0.9$ . For the *nonlinear* models, both the above substitutions must be made *and* the following additions as well. The term

$$-\frac{2}{3}\tilde{\rho}\tilde{k}\delta_{ij}+2\tilde{\mu}_{\mathrm{T}}K_{1}\frac{\tilde{k}}{\tilde{\epsilon}}(\tilde{S}_{ik}\tilde{W}_{kj}-\tilde{W}_{ik}\tilde{S}_{kj})-2\tilde{\mu}_{\mathrm{T}}K_{2}\frac{\tilde{k}}{\tilde{\epsilon}}\left(\tilde{S}_{ik}\tilde{S}_{kj}-\frac{1}{3}\tilde{S}_{kl}\tilde{S}_{lk}\delta_{ij}\right)$$

is added to  $\tilde{\tau}_{ij}$  in the momentum and energy equations and

$$\frac{\partial}{\partial \tilde{x}_j} \left[ \left( \tilde{\mu} + \frac{\tilde{\mu}_{\rm T}}{\sigma_{\rm k}} \right) \frac{\partial \tilde{k}}{\partial \tilde{x}_j} \right]$$

is added to the energy equation. These additions are made in subroutines gfluxv, hfluxv, and ffluxv.

### H.2 Nondimensionalizations

The turbulence equations are nondimensionalized by the same reference quantities as the Navier-Stokes equations:  $\tilde{a}_{\infty}$ ,  $\tilde{\rho}_{\infty}$ ,  $\tilde{\mu}_{\infty}$ , and  $\tilde{L}_R = \tilde{L}/L_{ref}$ . The nondimensionalized variables used with the turbulence models are:

$$\begin{aligned} \mathbf{k} &= \frac{\tilde{k}}{\tilde{a}_{\infty}^{2}} \qquad \omega = \frac{\tilde{\mu}_{\infty}\tilde{\omega}}{\tilde{\rho}_{\infty}\tilde{a}_{\infty}^{2}} \qquad \mathbf{P}_{\mathbf{k}} = \frac{\tilde{\mathbf{P}}_{\mathbf{k}}\tilde{L}_{R}^{2}}{\tilde{\mu}_{\infty}\tilde{a}_{\infty}^{2}} \\ \boldsymbol{\varepsilon} &= \frac{\tilde{\mu}_{\infty}\tilde{\varepsilon}}{\tilde{\rho}_{\infty}\tilde{a}_{\infty}^{4}} \qquad \boldsymbol{\rho} = \frac{\tilde{\rho}}{\tilde{\rho}_{\infty}} \qquad \mathbf{P}_{\omega} = \frac{\tilde{\mathbf{P}}_{\omega}\tilde{L}_{R}^{2}}{\tilde{\rho}_{\infty}\tilde{a}_{\infty}^{2}} \\ \boldsymbol{u} &= \frac{\tilde{u}}{\tilde{a}_{\infty}} \qquad \boldsymbol{p} = \frac{\tilde{p}}{\tilde{\rho}_{\infty}\tilde{a}_{\infty}^{2}} \qquad \mathbf{P}_{\omega} = \frac{\tilde{\mathbf{P}}_{\omega}\tilde{L}_{R}^{2}}{\tilde{\rho}_{\infty}\tilde{a}_{\infty}^{2}} \end{aligned} \tag{H-17}$$
$$\boldsymbol{x} &= \frac{\tilde{x}}{\tilde{L}_{R}} \qquad \boldsymbol{\mu} = \frac{\tilde{\mu}}{\tilde{\mu}_{\infty}} \qquad \mathbf{P}_{\varepsilon} = \frac{\tilde{\mathbf{P}}_{\varepsilon}\tilde{L}_{R}^{2}}{\tilde{\rho}_{\infty}\tilde{a}_{\infty}^{4}} \\ \boldsymbol{\tau}_{ij} &= \frac{\tilde{\tau}_{ij}\tilde{L}_{R}}{\tilde{\mu}_{\infty}\tilde{a}_{\infty}} \qquad \boldsymbol{t} = \frac{\tilde{t}\tilde{a}_{\infty}}{\tilde{L}_{R}} \qquad \boldsymbol{\Omega} = \frac{\tilde{\Omega}\tilde{L}_{R}}{\tilde{a}_{\infty}} \end{aligned}$$

### H.3 Zero-equation Models

#### H.3.1 Baldwin-Lomax Model

(ivisc = 2)

The Baldwin-Lomax<sup>10</sup> model is not a field-equation model; it is an algebraic model. Because it is the original model employed in CFL3D, its implementation differs in many respects from the more advanced models. First of all, it does not use the *minimum distance function* as its length scale, like the other models. Instead, it uses a *directed distance* from the i = 1, j = 1, k = 1, i = idim, j = jdim, or k = kdim point along a constant index line. For example, if the "body" is at k = 1, then the directed distance to a given point in the field is the directed (normal) distance from  $k_{given}$  to k = 1, keeping *i* and *j* fixed. See Figure H-1(a). The directed distance is

$$d = \overrightarrow{r_k} \cdot \overrightarrow{n} - \overrightarrow{r_0} \cdot \overrightarrow{n}$$
(H-18)

where  $\vec{r}_k$  is the vector from the origin to  $k_{given}$  and  $\vec{r}_0$  is the vector from the origin to the k = 1 point. Note that if grid lines curve significantly d can become negative as in Figure H-1(b). If the grid lines do this, CFL3D currently does not allow the distance to go

negative. Instead, it computes d out to its maximum, then sets all distances thereafter to that maximum value. However, if the grid behaves like the grid in Figure H-1(b), the Baldwin-Lomax model, which is inherently dependent on the grid structure, is probably not a good choice anyway. The use of any of the other field-equation models is recommended instead.

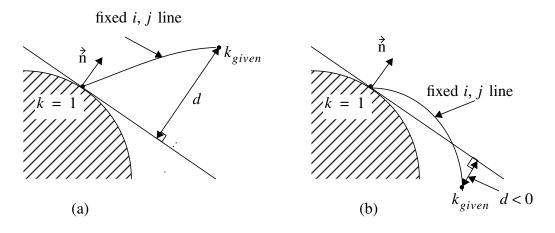


Figure H-1. Directed distance schematic.

In the current implementation in CFL3D, if the grid contains multiple blocks, Baldwin-Lomax should only be implemented on blocks that contain bodies. This restriction applies to the Baldwin-Lomax model only. It does *not* apply to the more general fieldequation models.

For the Baldwin-Lomax model,

$$\mu_{T} = \mu_{T, inner} \qquad y \le y_{crossover}$$

$$\mu_{T} = \mu_{T, outer} \qquad y > y_{crossover}$$
(H-19)

where  $y_{crossover}$  is the location along the constant index line where  $\mu_{T, inner}$  exceeds  $\mu_{T, outer}$ , marching away from the wall. The inner eddy viscosity is

$$\mu_{\mathrm{T, inner}} = \rho l^2 \Omega \left( \frac{Re}{M_{\infty}} \right) \tag{H-20}$$

where  $\Omega$  is the magnitude of the vorticity and

$$l = y[1 - \exp(-y^{+}/26)]$$
(H-21)

Since it now appears to be fairly widely accepted that the definition of  $y^+$  for the damping term needs to be modified in Baldwin-Lomax to give better answers when there are large temperature gradients near the wall,

$$y^{+} = \frac{\sqrt{\tilde{\rho}\tilde{\tau}_{w}}}{\tilde{\mu}}\tilde{y}$$
(H-22)

is used rather than

$$y^{+} = \frac{\sqrt{\tilde{\rho}_{w}\tilde{\tau}_{w}}}{\tilde{\mu}_{w}}\tilde{y}$$
(H-23)

Nondimensionally, this becomes

$$y^{+} = \frac{\sqrt{\rho\tau_{w}}}{\mu} y \left(\frac{Re}{M_{\infty}}\right)^{\frac{1}{2}}$$
(H-24)

The outer eddy viscosity is

$$\mu_{\mathrm{T, outer}} = 0.0168(1.6)\rho F_{wake} F_{kleb} \left(\frac{Re}{M}\right) \tag{H-25}$$

where

$$F_{wake} = \min[y_{max}F_{max}, 1.0y_{max}u_{dif}^2/F_{max}]$$

$$F(y) = y\Omega[1 - \exp(-y^*/26)]$$
(H-26)

In wakes,  $\exp(-y^{+}/26)$  is set to zero.  $F_{max}$  is the maximum value of F(y) that occurs in a profile and  $y_{max}$  is the value at which  $F_{max}$  occurs. Also,

$$F_{kleb} = \left[1 + 5.5 \left(\frac{0.3y}{y_{max}}\right)^{6}\right]^{-1}$$

$$u_{dif} = (\sqrt{u^{2} + v^{2} + w^{2}})_{max} - (\sqrt{u^{2} + v^{2} + w^{2}})_{min}$$
(H-27)

The second term in  $u_{dif}$  is taken to be zero, except in wakes. The region for the search for  $F_{max}$  is currently bound by

$$0.2\eta_{max} + 1 < \eta < 0.8\eta_{max} + 1 \tag{H-28}$$

where  $\eta$ , in this case, is the index direction away from the body.

## *H.3.2 Baldwin-Lomax with Degani-Schiff Modification* (**ivisc** = 3)

The Degani-Schiff<sup>16</sup> modification to the Baldwin-Lomax model is an algorithmic change which attempts to select the *first* occurrence of  $F_{max}$  in a search from the wall outwards. This can be important when there is a vortex somewhere above the body surface. If the code is not forced to select the  $F_{max}$  in the boundary layer, it may choose a length scale corresponding to the distance to the vortex, since often F can be larger in the vortex.

The test in CFL3D is very simple-minded and can often fail. However, it is quite difficult to find a test that works for all circumstances; so, for lack of anything better, the following is used. Marching outward away from the body,  $F_{max}$  is updated index by index. Then, if

$$F < 0.9F_{max},$$
 (H-29)

the code stops searching.

## H.4 One- and Two-equation Field-equation Models

All of the one- and two-equation model equations can be written in the general form

$$\frac{\partial}{\partial t}(\mathbf{X}) + u_j \frac{\partial}{\partial x_j}(\mathbf{X}) = \mathbf{S}_{\mathbf{P}} + \mathbf{S}_{\mathbf{D}} + D$$
(H-30)

where  $S_P$  is a "production" source term(s),  $S_D$  is a "destruction" source term(s), and *D* represents diffusion terms of the form  $\frac{\partial}{\partial x_j} \left[ ( \ ) \cdot \frac{\partial X}{\partial x_j} \right]$ . Note that  $S_P$  and  $S_D$  have been grouped together rather loosely. In Spalart's model in CFL3D, for example, part of Spalart's production term is grouped in  $S_D$  for convenience. In Menter's model in CFL3D, the cross-derivative term is grouped in  $S_D$ . In CFL3D, the  $S_P$  terms are treated explicitly while the  $S_D$  terms are linearized and treated implicitly. This is Spalart's "third strategy."<sup>35</sup>

All of the field-equation models are solved *uncoupled* from the Navier-Stokes equations. All of the models are solved in essentially the same fashion. Details are given in "Solution Method" on page 299 in the form of an example.

Also, all of the one- and two-equation models are based on incompressible turbulence equations. No compressibility corrections have been added. Hence, for problems where the turbulent Mach number,  $M_{\rm T} = \sqrt{2\tilde{k}/\tilde{a}^2}$ , is high, these turbulence models may not be applicable. Note that for most subsonic through low supersonic aerodynamic applications, the incompressible forms of the turbulence models are generally expected to be valid.

All of the field equation models except for Wilcox  $k - \omega$  make use of the *minimum distance function*, or the distance to the nearest wall, smin. This distance differs from the *directed distance* used by the Baldwin-Lomax model in that it does *not* follow grid lines and can be computed across zone boundaries. Hence, it is more easily applicable to multiple zone applications. The minimum distance function represents the distance to the nearest viscous wall and takes into account grid skewness when computing the nearest wallpoint location (in subroutine findmin\_new). The exception is the Baldwin-Barth model which uses a minimum distance algorithm (subroutine findmin) that does not take grid skewness into account. (The reason for this exception is that the Baldwin-Barth model requires other variables not currently implemented in subroutine findmin\_new.)

#### H.4.1 Baldwin-Barth Model

(ivisc = 4)

The Baldwin-Barth<sup>9</sup> model solves a single field equation for a turbulent Reynolds number term R:

$$\frac{\partial R}{\partial t} + u_j \frac{\partial R}{\partial x_j} = (C_{\epsilon_2} f_2 - C_{\epsilon_1}) \sqrt{RP} + \frac{M_{\infty}}{Re} \left( v + \frac{v_T}{\sigma_{\epsilon}} \right) \frac{\partial^2 R}{\partial x_j^2} - \frac{M_{\infty}}{Re} \frac{1}{\sigma_{\epsilon}} \frac{\partial}{\partial x_j} \left( v_T \frac{\partial R}{\partial x_j} \right)$$
(H-31)

where

$$\frac{1}{\sigma_{\varepsilon}} = (C_{\varepsilon_2} - C_{\varepsilon_1}) \frac{\sqrt{C_{\mu}}}{\kappa^2}$$
(H-32)

$$v_{\rm T} = C_{\mu} R D_1 D_2 \tag{H-33}$$

$$D_1 = 1 - \exp\left(-\frac{y^+}{A^+}\right)$$
 (H-34)

$$D_2 = 1 - \exp\left(-\frac{y^+}{A_2^+}\right)$$
 (H-35)

$$f_{2} = \frac{C_{\varepsilon_{1}}}{C_{\varepsilon_{2}}} + \left(1 - \frac{C_{\varepsilon_{1}}}{C_{\varepsilon_{2}}}\right) \left(\frac{1}{\kappa y^{+}} + D_{1}D_{2}\right) \left\{ \sqrt{D_{1}D_{2}} + \frac{y^{+}}{\sqrt{D_{1}D_{2}}} \left[\frac{1}{A^{+}} \exp\left(-\frac{y^{+}}{A^{+}}\right) D_{2} + \frac{1}{A_{2}^{+}} \exp\left(-\frac{y^{+}}{A_{2}^{+}}\right) D_{1}\right] \right\}$$

$$\mu_{T} = \rho \nu_{T}$$
(H-37)

$$\begin{aligned} \kappa &= 0.41 & C_{\varepsilon_1} = 1.2 & C_{\varepsilon_2} = 2.0 \\ C_{\mu} &= 0.09 & A^+ = 26 & A_2^+ = 10 \end{aligned} \tag{H-38}$$

The production term P is given by

$$P = 2v_{\rm T} S_{ij} S_{ij} - \frac{2}{3} v_{\rm T} \left(\frac{\partial u_k}{\partial x_k}\right)^2$$
(H-39)

but is approximated as

$$P \cong v_{\rm T} \Omega^2 \tag{H-40}$$

Hence, using the general form in Equation (H-30) with X = R,

$$\mathbf{S}_{\mathbf{P}} \cong (C_{\varepsilon_2} f_2 - C_{\varepsilon_1}) \sqrt{C_{\mu} D_1 D_2} \Omega R \tag{H-41}$$

$$D = \frac{M_{\infty}}{Re} \left( v + \frac{v_{\rm T}}{\sigma_{\varepsilon}} \right) \frac{\partial^2 R}{\partial x_j^2} - \frac{M_{\infty}}{Re} \frac{1}{\sigma_{\varepsilon}} \frac{\partial}{\partial x_j} \left( v_{\rm T} \frac{\partial R}{\partial x_j} \right)$$
(H-42)

Since the Baldwin-Barth model requires  $y^+$  (rather than just the minimum distance to the wall), additional information about the nearest wall point needs to be stored as well as the minimum distance function d. For simplicity, it is currently assumed that in the regions where  $y^+$  has an effect, these additional wall values are in the same grid zone as the point in question.

#### H.4.2 Spalart-Allmaras Model

(ivisc = 5)

The Spalart-Allmaras  $^{34}$  model solves a single field equation for a variable  $\hat{\nu}$  related to the eddy viscosity through

$$\mu_{\rm T} = \rho \hat{\nu} f_{\nu_1} \tag{H-43}$$

where

$$f_{\nu_1} = \frac{\chi^3}{\chi^3 + C_{\nu_1}^3}$$
(H-44)

$$\chi \equiv \frac{\hat{\nu}}{\nu} \tag{H-45}$$

The equation is

$$\begin{split} \frac{\partial \hat{\mathbf{v}}}{\partial t} + u_j \frac{\partial \hat{\mathbf{v}}}{\partial x_j} &= C_{b_1} [1 - f_{t_2}] \Omega \hat{\mathbf{v}} \qquad (\text{H-46}) \\ &+ \frac{M_{\infty}}{Re} \bigg\{ C_{b_1} [(1 - f_{t_2}) f_{v_2} + f_{t_2}] \frac{1}{\kappa^2} - C_{w_1} f_w \bigg\} \left( \frac{\hat{\mathbf{v}}}{d} \right)^2 \\ &- \frac{M_{\infty}}{Re} \frac{C_{b_2}}{\sigma} \hat{\mathbf{v}} \frac{\partial^2 \hat{\mathbf{v}}}{\partial x_j^2} \\ &+ \frac{M_{\infty}}{Re} \frac{1}{\sigma} \frac{\partial}{\partial x_j} \bigg[ (\mathbf{v} + (1 + C_{b_2}) \hat{\mathbf{v}}) \frac{\partial \hat{\mathbf{v}}}{\partial x_j} \bigg] \end{split}$$

(Note that Spalart's trip function is not used.) Also,

$$f_{t_2} = C_{t_3} \exp(-C_{t_4} \chi^2)$$
 (H-47)

d = distance to the closest wall = minimum distance function (H-48)

$$f_{w} = g \left[ \frac{1 + C_{w_{3}}^{6}}{g^{6} + C_{w_{3}}^{6}} \right]^{\frac{1}{6}} = \left[ \frac{g^{-6} + C_{w_{3}}^{-6}}{1 + C_{w_{3}}^{-6}} \right]^{\frac{1}{6}}$$
(H-49)

$$g = r + C_{w_2}(r^6 - r)$$
(H-50)

$$r = \frac{\hat{v}}{\hat{S}\left(\frac{Re}{M_{\infty}}\right)\kappa^2 d^2}$$
(H-51)

where

$$\hat{S} = \Omega + \frac{\hat{v}f_{v_2}}{\left(\frac{Re}{M_{\infty}}\right)\kappa^2 d^2}$$
(H-52)

$$f_{v_2}$$
 = no longer used (H-53)

$$f_{\nu_2} = 1 - \frac{\chi}{1 + \chi f_{\nu_1}} \tag{H-54}$$

CFL3D currently uses Spalart's Version Ia<sup>34</sup>. The fv3 term was employed as a smooth fix to prevent  $\hat{S}$  from going negative prior to 12/97, but was removed after an error was discovered ( $\hat{S}$  and fv2 were also different). The constants are

$$\begin{split} C_{b_1} &= 0.1355 \qquad \sigma = \frac{2}{3} \qquad C_{b_2} = 0.622 \qquad \kappa = 0.41 \\ C_{w_3} &= 2.0 \qquad C_{v_1} = 7.1 \qquad C_{t_3} = 1.2 \qquad C_{t_4} = 0.5 \qquad C_{w_2} = 0.3 \\ C_{w_1} &= \frac{C_{b_1}}{\kappa^2} + \frac{(1+C_{b_2})}{\sigma} \qquad \text{(note typo in ref. 34)} \end{split}$$

For the general form in Equation (H-30):

$$X = \hat{v} \tag{H-56}$$

$$S_{P} = C_{b_{1}} [1 - f_{t_{2}}] \Omega \hat{v}$$
 (H-57)

$$S_{\rm D} = \frac{M_{\infty}}{Re} \left\{ C_{b_1} [(1 - f_{t_2})f_{v_2} + f_{t_2}] \frac{1}{\kappa^2} - C_{w_1} f_w \right\} \left( \frac{\hat{v}}{d} \right)^2$$
(H-58)

$$D = -\frac{M_{\infty}C_{b_2}}{Re}\hat{\sigma}\hat{v}\frac{\partial^2\hat{v}}{\partial x_j^2} + \frac{M_{\infty}}{Re}\frac{1}{\sigma\partial x_j}\left[(v + (1 + C_{b_2})\hat{v})\frac{\partial\hat{v}}{\partial x_j}\right]$$
(H-59)

Note that in CFL3D, the S<sub>P</sub> and S<sub>D</sub> terms are grouped differently than Spalart's. Part of Spalart's production term is grouped into S<sub>D</sub> because it has the common factor  $M_{\infty}/Re$ , like the other destruction terms.

## H.4.3 Wilcox k-Omega Model (ivisc = 6)

For Wilcox's<sup>46</sup> model,

$$\frac{\partial \mathbf{k}}{\partial t} + u_j \frac{\partial \mathbf{k}}{\partial x_j} = \frac{1}{\rho} \mathbf{P}_{\mathbf{k}} \left( \frac{M_{\infty}}{Re} \right) - \beta' \mathbf{k} \omega \left( \frac{Re}{M_{\infty}} \right) + \frac{1}{\rho} \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_T}{\sigma_k} \right) \frac{\partial \mathbf{k}}{\partial x_j} \right] \left( \frac{M_{\infty}}{Re} \right)$$
(H-60)

$$\frac{\partial \omega}{\partial t} + u_j \frac{\partial \omega}{\partial x_j} = \frac{1}{\rho} P_\omega \left( \frac{M_{\infty}}{Re} \right) - \beta \omega^2 \left( \frac{Re}{M_{\infty}} \right) + \frac{1}{\rho} \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_T}{\sigma_\omega} \right) \frac{\partial \omega}{\partial x_j} \right] \left( \frac{M_{\infty}}{Re} \right)$$
(H-61)

In this model,  $C_{\mu}$  is incorporated into the definition of  $\omega,$  so

$$\mu_{\rm T} = \frac{\rho k}{\omega} \tag{H-62}$$

The production terms are approximated by

$$P_{k} = \mu_{T} \Omega^{2}$$
(H-63)

$$P_{\omega} = \gamma \rho \Omega^2$$
 (H-64)

The constants are

$$\gamma = \frac{\beta}{C_{\mu}} - \frac{\kappa^2}{\sigma_{\omega}\sqrt{C_{\mu}}} \qquad \beta' = C_{\mu} = 0.09$$
  

$$\sigma_k = 1/0.5 \qquad \beta = 0.075 \qquad (H-65)$$
  

$$\sigma_{\omega} = 1/0.5 \qquad \kappa = 0.41$$

The minimum distance function d is not required by this model. However, at the present time it is still computed and stored by CFL3D. For the general form in Equation (H-30),

$$X_{k} = k \tag{H-66}$$

$$S_{P,k} = \frac{1}{\rho} \mu_T \Omega^2 \left( \frac{M_{\infty}}{Re} \right)$$
(H-67)

$$S_{D, k} = -\beta' k \omega \left(\frac{Re}{M_{\infty}}\right)$$
(H-68)

$$D_{k} = \frac{1}{\rho} \frac{\partial}{\partial x_{j}} \left[ \left( \mu + \frac{\mu_{T}}{\sigma_{k}} \right) \frac{\partial k}{\partial x_{j}} \right] \left( \frac{M_{\infty}}{Re} \right)$$
(H-69)

and

$$X_{\omega} = \omega \tag{H-70}$$

$$S_{P,\omega} = \gamma \Omega^2 \left(\frac{M_{\infty}}{Re}\right)$$
(H-71)

$$S_{D,\omega} = -\beta \omega^2 \left(\frac{Re}{M_{\infty}}\right)$$
(H-72)

$$D_{\omega} = \frac{1}{\rho} \frac{\partial}{\partial x_{j}} \left[ \left( \mu + \frac{\mu_{\mathrm{T}}}{\sigma_{\omega}} \right) \frac{\partial \omega}{\partial x_{j}} \right] \left( \frac{M_{\infty}}{Re} \right)$$
(H-73)

#### H.4.4 Menter's k-Omega SST Model

(ivisc = 7)

Menter's k- $\omega$  model is assessed in reference 27. Note that there are two corrections to this reference. Equation (1) should be

$$\sigma \mu_t = \mu + \sigma \mu_t$$

$$\sigma^* \mu_t = \mu + \sigma^* \mu_t$$
(H-74)

and Equation (17) should be

$$\Gamma = \max(2\Gamma_3, \Gamma_1) \tag{H-75}$$

The two equations for this model are

$$\frac{\partial \mathbf{k}}{\partial t} + u_j \frac{\partial \mathbf{k}}{\partial x_j} = \frac{1}{\rho} \mathbf{P}_{\mathbf{k}} \left( \frac{M_{\infty}}{Re} \right) - \beta' \mathbf{k} \omega \left( \frac{Re}{M_{\infty}} \right) + \frac{1}{\rho} \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_T}{\sigma_k} \right) \frac{\partial \mathbf{k}}{\partial x_j} \right] \left( \frac{M_{\infty}}{Re} \right)$$
(H-76)

$$\frac{\partial \omega}{\partial t} + u_j \frac{\partial \omega}{\partial x_j} = \frac{1}{\rho} P_{\omega} \left( \frac{M_{\infty}}{Re} \right) - \beta \omega^2 \left( \frac{Re}{M_{\infty}} \right) + \frac{1}{\rho} \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_{\rm T}}{\sigma_{\omega}} \right) \frac{\partial \omega}{\partial x_j} \right] \left( \frac{M_{\infty}}{Re} \right) + 2(1 - F_1) \sigma_{\omega_2} \frac{1}{\omega} \frac{\partial k}{\partial x_j} \frac{\partial \omega}{\partial x_j} \left( \frac{M_{\infty}}{Re} \right)$$
(H-77)

In this model,  $C_{\mu}$  is incorporated into the definition of  $\omega.$  The eddy viscosity is given by

$$\mu_{\rm T} = \min\left[\frac{\rho k}{\omega}, \frac{a_1 \rho k}{\Omega F_2} \left(\frac{Re}{M}\right)\right] \tag{H-78}$$

The production terms are approximated by

$$P_k = \mu_T \Omega^2 \tag{H-79}$$

$$P_{\omega} = \gamma \rho \Omega^2$$
 (H-80)

The constants are calculated from  $\phi = F_1\phi_1 + (1 - F_1)\phi_2$ , where the  $\phi$ 's are the constants:

$$\begin{split} \gamma_{1} &= \frac{\beta_{1}}{C_{\mu}} - \frac{\kappa^{2}}{\sigma_{\omega_{1}}\sqrt{C_{\mu}}} & \gamma_{2} &= \frac{\beta_{2}}{C_{\mu}} - \frac{\kappa^{2}}{\sigma_{\omega_{2}}\sqrt{C_{\mu}}} \\ \sigma_{k_{1}} &= 1/0.85 & \sigma_{k_{2}} &= 1.0 \\ \sigma_{\omega_{1}} &= 1/0.5 & \sigma_{\omega_{2}} &= 1/0.856 \\ \beta_{1} &= 0.075 & \beta_{2} &= 0.0828 \\ \kappa &= 0.41 & a_{1} &= 0.31 \\ \beta' &= C_{\mu} &= 0.09 \end{split}$$
(H-81)

$$F_{1} = \tanh(\Gamma^{4})$$

$$\Gamma = \min[\max(\Gamma_{1}, \Gamma_{3}), \Gamma_{2}]$$

$$\Gamma_{1} = \frac{500\nu}{d^{2}\Omega} \left(\frac{M_{\infty}}{Re}\right)^{2} \qquad \Gamma_{2} = \frac{4\rho\sigma_{2}k}{d^{2}(CD_{k-\omega})} \qquad \Gamma_{3} = \frac{\sqrt{k}}{C_{\mu}\omega d} \left(\frac{M_{\infty}}{Re}\right)$$

$$CD_{k-\omega} = \max\left(\rho\frac{2\sigma_{2}}{\omega}\frac{\partial k}{\partial x_{j}}\frac{\partial \omega}{\partial x_{j}}, 1 \times 10^{-20}\right)$$

$$F_{2} = \tanh\Pi$$

$$\Pi = \max(2\Gamma_{3}, \Gamma_{1})$$
(H-82)

For the general form in Equation (H-30),

X<sub>k</sub> = k (H-83)

$$S_{P,k} = \frac{1}{\rho} \mu_T \Omega^2 \left( \frac{M_{\infty}}{Re} \right)$$
(H-84)

$$S_{D,k} = -\beta' k \omega \left(\frac{Re}{M_{\infty}}\right)$$
(H-85)

$$D_{k} = \frac{1}{\rho} \frac{\partial}{\partial x_{j}} \left[ \left( \mu + \frac{\mu_{T}}{\sigma_{k}} \right) \frac{\partial k}{\partial x_{j}} \right] \left( \frac{M_{\infty}}{Re} \right)$$
(H-86)

and

$$X_{\omega} = \omega \tag{H-87}$$

$$S_{P,\omega} = \gamma \Omega^2 \left(\frac{M_{\infty}}{Re}\right)$$
(H-88)

$$S_{D,\omega} = -\beta\omega^2 \left(\frac{Re}{M_{\infty}}\right) + 2(1 - F_1)\sigma_{\omega_2} \frac{1}{\omega} \frac{\partial k}{\partial x_j} \frac{\partial \omega}{\partial x_j} \left(\frac{M_{\infty}}{Re}\right)$$
(H-89)

$$D_{\omega} = \frac{1}{\rho} \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_{\rm T}}{\sigma_{\omega}} \right) \frac{\partial \omega}{\partial x_j} \right] \left( \frac{M_{\infty}}{Re} \right)$$
(H-90)

H.4.5 Abid k-Epsilon Model

(**ivisc** = 10)

For the Abid<sup>2</sup> k- $\varepsilon$  model, the equations are

$$\frac{\partial \mathbf{k}}{\partial t} + u_j \frac{\partial \mathbf{k}}{\partial x_j} = \frac{1}{\rho} \mathbf{P}_{\mathbf{k}} \left( \frac{M_{\infty}}{Re} \right) - \varepsilon \left( \frac{Re}{M_{\infty}} \right) + \frac{1}{\rho} \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_{\mathrm{T}}}{\sigma_{\mathrm{k}}} \right) \frac{\partial \mathbf{k}}{\partial x_j} \right] \left( \frac{M_{\infty}}{Re} \right)$$
(H-91)

$$\frac{\partial \varepsilon}{\partial t} + u_j \frac{\partial \varepsilon}{\partial x_j} = \frac{1}{\rho} P_{\varepsilon} \left( \frac{M_{\infty}}{Re} \right) - C_{\varepsilon_2} \frac{\varepsilon^2}{k} f_2 \left( \frac{Re}{M_{\infty}} \right) + \frac{1}{\rho} \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_T}{\sigma_{\varepsilon}} \right) \frac{\partial \varepsilon}{\partial x_j} \right] \left( \frac{M_{\infty}}{Re} \right)$$
(H-92)

$$\mu_{\rm T} = C_{\mu} f_{\mu} \frac{\rho k^2}{\epsilon}$$
(H-93)

The production terms are approximated by

$$P_{k} = \mu_{T} \Omega^{2}$$

$$P_{\epsilon} = C_{\epsilon_{1} k} \frac{\epsilon}{\mu_{T}} \Omega^{2}$$
(H-94)

The constants are

$$C_{\varepsilon_1} = 1.45$$
  $\sigma_k = 1.0$   
 $C_{\varepsilon_2} = 1.83$   $\sigma_{\varepsilon} = 1.4$  (H-95)  
 $C_{\mu} = 0.09$ 

The damping functions are given by

$$f_{\mu} = [1 + 4(Re_{\tau}^{-0.75})] \tanh(0.008Re_{k})$$
  
$$f_{2} = \left[1 - \exp\left(-\frac{Re_{k}}{12}\right)\right]$$
 (H-96)

where

$$Re_{\tau} = \frac{\rho k^{2}}{\mu \varepsilon}$$

$$Re_{k} = \frac{\rho \sqrt{k} d}{\mu} \left(\frac{Re}{M}\right)$$
(H-97)

where d is the distance to the nearest wall. For the general form in Equation (H-30),

$$X_{k} = k \tag{H-98}$$

$$S_{P,k} = \frac{1}{\rho} \mu_T \Omega^2 \left( \frac{M_{\infty}}{Re} \right)$$
(H-99)

$$S_{D, k} = -\varepsilon \left(\frac{Re}{M_{\infty}}\right)$$
(H-100)

$$D_{k} = \frac{1}{\rho} \frac{\partial}{\partial x_{j}} \left[ \left( \mu + \frac{\mu_{T}}{\sigma_{k}} \right) \frac{\partial k}{\partial x_{j}} \right] \left( \frac{M_{\infty}}{Re} \right)$$
(H-101)

and

$$X_{\epsilon} = \epsilon$$
 (H-102)

$$S_{P, \epsilon} = \frac{1}{\rho} C_{\epsilon_1} \frac{\epsilon}{k} \mu_T \Omega^2 \left(\frac{M_{\infty}}{Re}\right)$$
(H-103)

$$S_{D, \epsilon} = -C_{\epsilon_2} \frac{\epsilon^2}{k} f_2 \left(\frac{Re}{M_{\infty}}\right)$$
(H-104)

$$D_{\varepsilon} = \frac{1}{\rho} \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_{\rm T}}{\sigma_{\varepsilon}} \right) \frac{\partial \varepsilon}{\partial x_j} \right] \left( \frac{M_{\infty}}{Re} \right)$$
(H-105)

## H.4.6 EASM Gatski-Speziale k-Omega Model

(**ivisc** = 8 and 12)

EASM stands for Explicit Algebraic Stress Model.<sup>3</sup> **Ivisc** = 8 is the "linear"  $k - \omega$  version, treated as an eddy viscosity model, whereas **ivisc** = 12 is the fully nonlinear  $k - \omega$  version. The two versions obtain  $\mu_T$  with identical methods. However, the nonlinear model includes nonlinear terms added to the Navier-Stokes equations, whereas the linear model does not. (Its effect is felt only through the  $\mu_T$  term.)

The k –  $\omega$  equations are of the same form as in the Wilcox k –  $\omega$  model (Equation (H-60) and Equation (H-61)). However,  $C_{\mu}$  is not incorporated into the definition of  $\omega$  (it is now a variable coefficient), so,

$$\mu_{\rm T} = C_{\mu} \rho \frac{k}{\omega} \tag{H-106}$$

The constants are different as well:

$$\sigma_{k} = 1.4$$

$$\sigma_{\omega} = 2.2$$

$$\beta' = 1.0$$

$$\beta = 0.83$$

$$\kappa = 0.41$$

$$(H-107)$$

The production terms are

$$P_{k} = \tau_{ij}^{T} \frac{\partial u_{i}}{\partial x_{j}}$$

$$P_{\omega} = \gamma \frac{\omega}{k} \tau_{ij}^{T} \frac{\partial u_{i}}{\partial x_{j}}$$
(H-108)

where  $\tau_{ij}^{T}$  is given (dimensionally) in Equation (H-15). The variable  $C_{\mu}$  and the coefficients  $K_1$  and  $K_2$  in Equation (H-15) are determined as follows:

$$C_{\mu} = \frac{3(1+\eta^2) + 0.2(\eta^6 + \zeta^6)}{3+\eta^2 + 6\zeta^2\eta^2 + 6\zeta^2 + \eta^6 + \zeta^6} \alpha_1$$
(H-109)

$$\eta = \frac{\alpha_2}{\omega} (S_{ij} S_{ij})^{\frac{1}{2}} \left( \frac{M_{\infty}}{Re} \right)$$
(H-110)

$$\zeta = \frac{\alpha_3}{\omega} (W_{ij} W_{ij})^{\frac{1}{2}} \left( \frac{M_{\infty}}{Re} \right)$$
(H-111)

$$K_1 = \alpha_3 \tag{H-112}$$
$$K_2 = 2\alpha_2$$

where

$$\alpha_{1} = \left(\frac{4}{3} - C_{2}\right)^{g}_{2}$$

$$\alpha_{2} = (2 - C_{3})^{g}_{2}$$

$$\alpha_{3} = (2 - C_{4})^{g}_{2}$$

$$g = \left(\frac{C_{1}}{2} + C_{5} - 1\right)^{-1}$$
(H-113)

(Note that  $\eta$ ,  $\zeta$ , and  $\alpha$  are just used here to denote terms in the equations and do *not* reflect the definitions in the Nomenclature list.) Currently, the pressure-strain correlation is modeled with the Speziale-Sarkar-Gatski (SSG) correlation, which uses:

$$C_1 = 6.8$$
  $C_3 = 1.25$   $C_5 = 1.88$   
 $C_2 = 0.36$   $C_4 = 0.4$   $C_{\mu}^* = 0.081$  (H-114)

To improve convergence, the  $\mu_{T}$  terms in  $\tau_{ij}^{T}$  multiplying the nonlinear terms

$$S_{ik}W_{kj} - W_{ik}S_{kj}$$
 and  $S_{ik}S_{kj} - \frac{1}{3}S_{kl}S_{lk}\delta_{ij}$ 

are replaced by

$$\mu_{\rm T}' = C_{\mu}' \rho \frac{\rm k}{\omega} \tag{H-115}$$

where

$$C_{\mu}' = \frac{3(1+\eta^2)}{3+\eta^2+6\zeta^2\eta^2+6\zeta^2+\eta^6+\zeta^6}\alpha_1$$
 (H-116)

Also, the  $\mu_T$  terms in the  $D_k$  and  $D_{\omega}$  diffusion terms are replaced by

$$\mu_{\rm T}^* = C_{\mu}^* \rho \frac{\rm k}{\omega} \tag{H-117}$$

The minimum distance function d is not required by this model. However, at the present time it is still computed and stored by CFL3D. For the general form in Equation (H-30),

$$X_{k} = k \tag{H-118}$$

$$S_{P,k} = \frac{1}{\rho} \tau_{ij}^{T} \frac{\partial u_i}{\partial x_j} \left( \frac{M_{\infty}}{Re} \right)$$
(H-119)

$$S_{D, k} = -\beta' k \omega \left(\frac{Re}{M_{\infty}}\right)$$
(H-120)

$$D_{k} = \frac{1}{\rho} \frac{\partial}{\partial x_{j}} \left[ \left( \mu + \frac{\mu_{T}^{*}}{\sigma_{k}} \right) \frac{\partial k}{\partial x_{j}} \right] \left( \frac{M_{\infty}}{Re} \right)$$
(H-121)

and

$$X_{\omega} = \omega$$
 (H-122)

$$S_{P,\omega} = \frac{1}{\rho} \gamma \frac{\omega}{k} \tau_{ij}^{T} \frac{\partial u_i}{\partial x_j} \left( \frac{M_{\infty}}{Re} \right)$$
(H-123)

$$S_{D,\omega} = -\beta \omega^2 \left(\frac{Re}{M_{\infty}}\right)$$
(H-124)

$$D_{\omega} = \frac{1}{\rho} \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_T^*}{\sigma_{\omega}} \right) \frac{\partial \omega}{\partial x_j} \right] \left( \frac{M_{\infty}}{Re} \right)$$
(H-125)

## H.4.7 EASM Gatski-Speziale k-Epsilon Model

(ivisc = 11)

The  $k - \varepsilon$  version of the Gatski-Speziale EASM model<sup>3</sup> is only implemented as a nonlinear model. Its equations are identical to those of the Abid  $k - \varepsilon$  model (Equation (H-91) and Equation (H-92)), with

$$\mu_{\rm T} = C_{\mu} \frac{\rho k^2}{\epsilon} \tag{H-126}$$

(There is no  $f_{\mu}$  term.) The constants are

$$C_{\varepsilon_{2}} = 1.83$$

$$C_{\varepsilon_{1}} = C_{\varepsilon_{2}} - \frac{\kappa^{2}}{\sigma_{\varepsilon}\sqrt{C_{\mu}}^{*}} \qquad \sigma_{\varepsilon} = 1.3$$

$$\kappa = 0.41$$

$$(H-127)$$

The production terms are

$$P_{k} = \tau_{ij}^{T} \frac{\partial u_{i}}{\partial x_{j}}$$

$$P_{\varepsilon} = C_{\varepsilon_{1}} \frac{\varepsilon}{k} \tau_{ij}^{T} \frac{\partial u_{i}}{\partial x_{j}}$$
(H-128)

The variable  $C_{\mu}$ , and all constants are identical to those given for the EASM k –  $\omega$  model in "EASM Gatski-Speziale k-Omega Model" on page 286, except that  $\omega$  is replaced by  $\epsilon/k$ . The damping functions are given by

$$f_2 = \left[1 - \exp\left(-\frac{Re_k}{12}\right)\right] \tag{H-129}$$

where

$$Re_{\rm k} = \frac{\rho\sqrt{\rm k}d}{\mu} \left(\frac{Re}{M_{\infty}}\right) \tag{H-130}$$

The  $\mu_T$  terms in  $\tau_{ij}^T$  multiplying the nonlinear terms are replaced in exactly the same way as for the EASM k –  $\omega$  model. Also, the  $\mu_T$  terms in the  $D_k$  and  $D_{\varepsilon}$  diffusion terms are replaced by

$$\mu_{\rm T}^* = C_{\mu}^* \frac{\rho k^2}{\epsilon} \tag{H-131}$$

For the general form in Equation (H-30),

$$X_{k} = k \tag{H-132}$$

$$S_{P,k} = \frac{1}{\rho} \tau_{ij}^{T} \frac{\partial u_i}{\partial x_j} \left( \frac{M_{\infty}}{Re} \right)$$
(H-133)

$$S_{D, k} = -\varepsilon \left(\frac{Re}{M_{\infty}}\right) \tag{H-134}$$

$$D_{k} = \frac{1}{\rho} \frac{\partial}{\partial x_{j}} \left[ \left( \mu + \frac{\mu_{T}^{*}}{\sigma_{k}} \right) \frac{\partial k}{\partial x_{j}} \right] \left( \frac{M_{\infty}}{Re} \right)$$
(H-135)

and

$$X_{\epsilon} = \epsilon$$
 (H-136)

$$S_{P,\epsilon} = \frac{1}{\rho} C_{\epsilon_1} \frac{\epsilon}{k} \tau_{ij}^T \frac{\partial u_i}{\partial x_j} \left( \frac{M_{\infty}}{Re} \right)$$
(H-137)

$$S_{D, \epsilon} = -C_{\epsilon_2} \frac{\epsilon^2}{k} f_2 \left(\frac{Re}{M_{\infty}}\right)$$
(H-138)

$$D_{\varepsilon} = \frac{1}{\rho} \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_{\rm T}^*}{\sigma_{\varepsilon}} \right) \frac{\partial \varepsilon}{\partial x_j} \right] \left( \frac{M_{\infty}}{Re} \right)$$
(H-139)

H.4.8 EASM Girimaji k-Epsilon Model

(**ivisc** = 9 and 13)

The Girimaji<sup>21</sup> version of the EASM  $k - \varepsilon$  model is implemented both as a "linear" (eddy-viscosity) (**ivisc** = 9) and "nonlinear" (**ivisc** = 13) version. The  $k - \varepsilon$  equations are

$$\frac{\partial \mathbf{k}}{\partial t} + u_j \frac{\partial \mathbf{k}}{\partial x_j} = \frac{1}{\rho} \mathbf{P}_{\mathbf{k}} \left( \frac{M_{\infty}}{Re} \right) - \varepsilon \left( \frac{Re}{M_{\infty}} \right) + \frac{1}{\rho} \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_T}{\sigma_k} \right) \frac{\partial \mathbf{k}}{\partial x_j} \right] \left( \frac{M_{\infty}}{Re} \right)$$
(H-140)

$$\frac{\partial \varepsilon}{\partial t} + u_j \frac{\partial \varepsilon}{\partial x_j} = \frac{1}{\rho} P_{\varepsilon} \left( \frac{M_{\infty}}{Re} \right) - C_{\varepsilon_2} \frac{\varepsilon^2}{k} \left( \frac{Re}{M_{\infty}} \right) + \frac{1}{\rho} \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_T}{\sigma_{\varepsilon}} \right) \frac{\partial \varepsilon}{\partial x_j} \right] \left( \frac{M_{\infty}}{Re} \right)$$

$$+ C_{\varepsilon_2} \frac{\varepsilon}{k} \frac{2\mu}{\rho} \left( \frac{\partial\sqrt{k}}{\partial x_j} \right)^2 \left( \frac{M_{\infty}}{Re} \right)$$
(H-141)

$$\mu_{\rm T} = -G_1 f_{\mu} \frac{\rho k^2}{\epsilon} \tag{H-142}$$

The extra  $(\partial \sqrt{k}/\partial x_j)^2$  term in the  $\varepsilon$  equation was added to replace the  $f_2$  damping function term on the  $\varepsilon^2/k$  term, which is needed because  $\varepsilon^2/k \to \infty$  at the wall. The constants are

$$C_{\varepsilon_1} = 1.44$$
  $\sigma_k = 1.0$   
 $C_{\varepsilon_2} = 1.83$   $\sigma_{\varepsilon} = 1.3$  (H-143)

The production terms are approximated by

$$P_{k} = \tau_{ij}^{T} \frac{\partial u_{i}}{\partial x_{j}}$$

$$P_{\epsilon} = C_{\epsilon_{1}} \frac{\epsilon}{k} \tau_{ij}^{T} \frac{\partial u_{i}}{\partial x_{j}}$$
(H-144)

The damping functions  $f_{\mu}$  is

$$f_{\mu} = \tanh\left[0.015 \frac{\sqrt{k}d\rho}{\mu} \left(\frac{Re}{M_{\infty}}\right)\right]$$
(H-145)

 $\tau_{ij}^{T}$  is given (dimensionally) in Equation (H-15). However,  $K_1$  and  $K_2$  in that equation are now given by

$$K_{1} = \frac{G_{2}}{G_{1}}$$

$$K_{2} = -\frac{G_{3}}{G_{1}}$$
(H-146)

and

for 
$$\eta_1 = 0$$
  $G_1 = \frac{L_1^0 L_2}{(L_1^0)^2 + 2\eta_2 (L_4)^2}$   
for  $L_1^1 = 0$   $G_1 = \frac{L_1^0 L_2}{(L_1^0)^2 + \frac{2}{3}\eta_1 (L_3)^2 + 2\eta_2 (L_4)^2}$   
for  $D > 0$   $G_1 = -\frac{p}{3} + \left(-\frac{b}{2} + \sqrt{D}\right)^{\frac{1}{3}} + \left(-\frac{b}{2} - \sqrt{D}\right)^{\frac{1}{3}}$  (H-147)  
for  $D < 0, b < 0$   $G_1 = -\frac{p}{3} + 2\sqrt{-\frac{a}{3}}\cos\left(\frac{\theta}{3}\right)$   
for  $D < 0, b > 0$   $G_1 = -\frac{p}{3} + 2\sqrt{-\frac{a}{3}}\cos\left(\frac{\theta}{3}\right)$ 

where

$$\eta_{1} = \left(\frac{k}{\varepsilon}\right)^{2} S_{ij} S_{ij} \left(\frac{M_{\infty}}{Re}\right)^{2}$$

$$\eta_{2} = \left(\frac{k}{\varepsilon}\right)^{2} W_{ij} W_{ij} \left(\frac{M_{\infty}}{Re}\right)^{2}$$
(H-148)

$$p = -\frac{2L_1^0}{\eta_1 L_1^1} \qquad r = -\frac{L_1^0 L_2}{(\eta_1 L_1^1)^2}$$
(H-149)

$$q = \frac{1}{(\eta_1 L_1^1)^2} \left[ (L_1^0)^2 + \eta_1 L_1^1 L_2 - \frac{2}{3} \eta_1 (L_3)^2 + 2\eta_2 (L_4)^2 \right]$$
(H-150)

$$a = \left(q - \frac{p^2}{3}\right) \qquad b = \frac{1}{27}(2p^3 - 9pq + 27r)$$

$$D = \frac{b^2}{4} + \frac{a^3}{27} \qquad \cos\theta = \frac{-b/2}{\sqrt{-a^3/27}}$$
(H-151)

$$G_2 = \frac{-L_4 G_1}{L_0^1 - \eta_1 L_1^1 G_1} \qquad G_3 = \frac{+2L_3 G_1}{L_0^1 - \eta_1 L_1^1 G_1}$$
(H-152)

C

$c^0$	$L_2 = \frac{C_2}{2} - \frac{2}{3}$	
$L_0^1 = \frac{C_1^0}{2} - 1$	$L_3 = \frac{C_3}{2} - 1$	(H-153)
$L_1^1 = C_1^1 + 2$	$L_4 = \frac{C_4}{2} - 1$	

For the general form in Equation (H-30),

$$X_{k} = k \tag{H-155}$$

$$S_{P,k} = \frac{1}{\rho} \tau_{ij}^{T} \frac{\partial u_i}{\partial x_j} \left( \frac{M_{\infty}}{Re} \right)$$
(H-156)

$$S_{D, k} = -\varepsilon \left(\frac{Re}{M_{\infty}}\right)$$
(H-157)

$$D_{k} = \frac{1}{\rho} \frac{\partial}{\partial x_{j}} \left[ \left( \mu + \frac{\mu_{T}}{\sigma_{k}} \right) \frac{\partial k}{\partial x_{j}} \right] \left( \frac{M_{\infty}}{Re} \right)$$
(H-158)

and

$$X_{\epsilon} = \epsilon$$
 (H-159)

$$S_{P,\varepsilon} = \frac{1}{\rho} C_{\varepsilon_1} \frac{\varepsilon}{k} \tau_{ij}^T \frac{\partial u_i}{\partial x_j} \left( \frac{M_{\infty}}{Re} \right) + C_{\varepsilon_2} \frac{\varepsilon}{k} \frac{2\mu}{\rho} \left( \frac{\partial \sqrt{k}}{\partial x_j} \right)^2 \left( \frac{M_{\infty}}{Re} \right)$$
(H-160)

$$S_{D, \epsilon} = -C_{\epsilon_2} \frac{\epsilon^2}{k} \left(\frac{Re}{M_{\infty}}\right)$$
(H-161)

$$D_{\varepsilon} = \frac{1}{\rho} \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_{\rm T}}{\sigma_{\varepsilon}} \right) \frac{\partial \varepsilon}{\partial x_j} \right] \left( \frac{M_{\infty}}{Re} \right)$$
(H-162)

## H.5 Generalized Coordinate Form

For the transformation from Cartesian coordinates to generalized coordinates, for example,

$$u_j \frac{\partial \mathbf{k}}{\partial x_j} = U \frac{\partial \mathbf{k}}{\partial \xi} + V \frac{\partial \mathbf{k}}{\partial \eta} + W \frac{\partial \mathbf{k}}{\partial \zeta}$$
(H-163)

where

$$U = \xi_x u + \xi_y v + \xi_z w + \xi_t$$
  

$$V = \eta_x u + \eta_y v + \eta_z w + \eta_t$$
  

$$W = \zeta_x u + \zeta_y v + \zeta_z w + \zeta_t$$
  
(H-164)

In the diffusion-type terms, cross-derivative terms are neglected. For example,

$$\begin{aligned} \frac{\partial}{\partial x_{j}} \left( \mu \frac{\partial k}{\partial x_{j}} \right) &\cong \xi_{x} \frac{\partial}{\partial \xi} \left( \xi_{x} \mu \frac{\partial k}{\partial \xi} \right) + \xi_{y} \frac{\partial}{\partial \xi} \left( \xi_{y} \mu \frac{\partial k}{\partial \xi} \right) + \xi_{z} \frac{\partial}{\partial \xi} \left( \xi_{z} \mu \frac{\partial k}{\partial \xi} \right) \\ &+ \eta_{x} \frac{\partial}{\partial \eta} \left( \eta_{x} \mu \frac{\partial k}{\partial \eta} \right) + \eta_{y} \frac{\partial}{\partial \eta} \left( \eta_{y} \mu \frac{\partial k}{\partial \eta} \right) + \eta_{z} \frac{\partial}{\partial \eta} \left( \eta_{z} \mu \frac{\partial k}{\partial \eta} \right) \\ &+ \zeta_{x} \frac{\partial}{\partial \zeta} \left( \zeta_{x} \mu \frac{\partial k}{\partial \zeta} \right) + \zeta_{y} \frac{\partial}{\partial \zeta} \left( \zeta_{y} \mu \frac{\partial k}{\partial \zeta} \right) + \zeta_{z} \frac{\partial}{\partial \zeta} \left( \zeta_{z} \mu \frac{\partial k}{\partial \zeta} \right) \end{aligned}$$
(H-165)

An example of the discretization is given for the first term on the right-hand side of Equation (H-165):

$$\xi_{x}\frac{\partial}{\partial\xi}\left(\xi_{x}\mu\frac{\partial k}{\partial\xi}\right) \rightarrow \xi_{x_{i}}\xi_{x_{i+\frac{1}{2}}}\mu_{i+\frac{1}{2}}\left(\frac{\partial k}{\partial\xi}\right)_{i+\frac{1}{2}} - \xi_{x_{i}}\xi_{x_{i-\frac{1}{2}}}\mu_{i-\frac{1}{2}}\left(\frac{\partial k}{\partial\xi}\right)_{i-\frac{1}{2}}$$
(H-166)

where

$$\left(\frac{\partial k}{\partial \xi}\right)_{i+\frac{1}{2}} = k_{i+1} - k_i \qquad \left(\frac{\partial k}{\partial \xi}\right)_{i-\frac{1}{2}} = k_i - k_{i-1}$$
(H-167)

The locations of the indices are shown in Figure H-2.

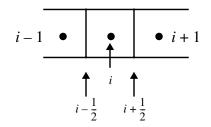


Figure H-2. Definition of left and right states for generalized coordinate transformation.

## H.6 Initial Conditions

For the Baldwin-Barth and Spalart-Allmaras models, the levels of the turbulence quantities *R* and  $\hat{v}$  are initialized at their free-stream levels everywhere in the flow field. For the two-equation models, it helps the solution start-up phase to initialize the k and  $\omega$  (or k and  $\varepsilon$ ) values to crude profiles that simulate typical profiles in a boundary layer. Currently, for k –  $\omega$ :

$$k_{IC} = \max(k_{\infty}, -C_1 d^2 + C_2 d)$$

$$\omega_{IC} = \max\left(-12444d + 0.54, \frac{C_{\mu}k_{IC}}{v_3 d}\right)$$
(H-168)

where d is the distance to the nearest wall and, for EASM models,  $C_{\mu} = 0.09$ ; otherwise,  $C_{\mu} = 1.0$ . Also,

$$v3d = \frac{100k_{IC}}{t_{max}}$$

$$t_{max} = -C_1 S_{max}^2 + C_2 S_{max}$$

$$S_{max} = \frac{C_2}{2C_1}$$

$$C_1 = 45.8 \qquad C_2 = 1.68$$
(H-169)

For  $k - \epsilon$ :

$$k_{IC} = \min\{zk4, \max[zk1, \min(zk2, zk3)]\}$$
(H-170)  

$$zk1 = k_{\infty}$$

$$zk2 = 10^{-471d + 0.47}$$

$$zk3 = 10^{-37.5d - 3.7}$$

$$zk4 = 6.7d$$
(H-171)  

$$\epsilon_{IC} = \min\{ep4, \max[ep1, \min(ep2, ep3)]\}$$
(H-172)  

$$ep1 = \epsilon_{\infty}$$

$$ep2 = 10^{-555d - 6}$$

$$ep3 = 10^{-280d - 9.2}$$

$$ep4 = \min(1 \times 10^{20}, 10^{13333d - 9.8})$$

## H.7 Boundary Conditions

## H.7.1 Free-stream Levels

For the Baldwin-Barth model,

H.7.2 Boundary Conditions at Solid Walls

$$R_{\infty} = 0.1$$
 (H-174)

For the Spalart-Allmaras model,

$$\hat{v}_{\infty} = 1.341946$$
 (H-175)

For the standard k –  $\omega$  or Menter's SST models (C<sub>µ</sub> is incorporated into the definition of  $\omega$ ),

$$k_{\infty} = 9 \times 10^{-9}$$

$$\omega_{\infty} = 1 \times 10^{-6}$$
(H-176)

For the EASM  $(k - \omega)$  models,

$$k_{\infty} = 9 \times 10^{-9}$$

$$\omega_{\infty} = 9 \times 10^{-8}$$
(H-177)

For the  $k - \varepsilon$  models,

$$k_{\infty} = 1 \times 10^{-9}$$

$$\epsilon_{\infty} = 1 \times 10^{-17}$$
(H-178)

The end result for all models is

$$\mu_{\rm T, \infty} \cong 0.009$$
 (H-179)

in the free stream. Note that  $\omega_{\infty} = 9 \times 10^{-8}$  and  $\varepsilon_{\infty} = 1 \times 10^{-17}$  for the EASM models are based on  $C_{\mu} = 0.09$ , which is only approximate since  $C_{\mu}$  is variable.

#### H.7.2 Boundary Conditions at Solid Walls

With the parameters known at the cell-centers near the solid wall boundary (see the wall region schematic in Figure H-3), the following boundary conditions are applied.

For Baldwin-Barth,

$$R_{\rm w} = 0 \tag{H-180}$$

For Spalart-Allmaras,

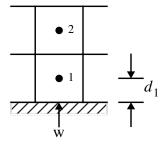


Figure H-3. Parameter locations at wall boundary.

$$\hat{\mathbf{v}} = \mathbf{0} \tag{H-181}$$

For the  $k - \omega$  models,

$$\mathbf{k}_{\mathrm{w}} = \mathbf{0} \tag{H-182}$$

$$\omega_{\rm w} = \frac{60\mu_1}{\rho_1 \beta(d_1)^2} \left(\frac{M_{\infty}}{Re}\right)^2$$
(H-183)

where

$$\beta = 0.83$$
 for the EASM k –  $\omega$  models and  
 $\beta = 0.075$  otherwise.

For the k-  $\epsilon$  models,

$$k_{w} = 0$$
 (H-184)

$$\varepsilon_{\rm w} = \frac{2\mu_1}{\rho_1} \left(\frac{\partial\sqrt{k}}{\partial n}\right)_{\rm w}^2 \left(\frac{M_{\infty}}{Re}\right)^2 \tag{H-185}$$

where n is the direction normal to the wall.

Note that the actual wall boundary conditions for the turbulence quantities (unlike the primitive variables) are applied at ghost cells. Hence,

$$\chi_{\rm BC} = 2\chi_{\rm w} - \chi_1 \tag{H-186}$$

where  $\chi$  represents R,  $\hat{v}$ , k,  $\omega$ , or  $\epsilon$ .

Version 4.1 of CFL3D applied the above boundary condition for  $\omega$  (Equation (H-183)) at the ghost cell center rather than at the wall. This is not correct, but supposedly any wall boundary condition for  $\omega$  that is bigger by some factor than the "analytical" value  $6\mu_1/(\rho\beta y^2)$  gives basically the same result. (See reference 26.) The current version of CFL3D (Version 5.0) obtains the ghost cell value using linear interpolation from  $\omega_w$  and  $\omega_1$  (or  $\varepsilon_w$  and  $\varepsilon_1$ ).

## H.8 Solution Method

The one or two equations are solved, decoupled, using implicit approximate factorization (AF). The  $S_P$  terms are treated explicitly, lagged in time while the  $S_D$  and D terms are treated implicitly (they are linearized and a term is brought to the left-hand-side of the equations). This procedure is described by Spalart and Allmaras<sup>35</sup> (strategy #3). The advective terms are discretized using first-order upwinding. Treating the destruction terms implicitly helps increase the diagonal dominance of the left-hand-side matrix. Most of the  $S_D$  terms are linearized in a simplified fashion by assuming no coupling between the variables. For example,

$$\beta' k \omega^{(n+1)} \cong \beta' k \omega^{(n)} + \frac{\partial}{\partial k} (\beta' k \omega) \Delta k$$

$$\cong \beta' k \omega^{(n)} + \beta' \omega \Delta k$$
(H-187)

For the S<sub>D</sub> terms in the k –  $\varepsilon$  equations, however, a more sophisticated approach was taken. When linearizing, the k and  $\varepsilon$  variables are assumed to be coupled, with the eddy viscosity  $\mu_{\rm T}$  assumed to be fixed. For example, in the k equation, S<sub>D</sub> =  $-\varepsilon(Re/M_{\infty})$  is linearized as follows:

$$\varepsilon^{(n+1)} \cong \varepsilon^{(n)} + \frac{\partial}{\partial k} (\varepsilon) \Delta k \tag{H-188}$$
$$\cong \varepsilon^{(n)} + \frac{\partial}{\partial k} \left( \frac{C_{\mu} f_{\mu} \rho k^{2}}{\mu_{T}} \right) \Delta k$$
$$\cong \varepsilon^{(n)} + 2 \frac{C_{\mu} f_{\mu} \rho k}{\mu_{T}} \Delta k$$
$$\cong \varepsilon^{(n)} + 2 \frac{\varepsilon}{k} \Delta k$$

The S<sub>D</sub> =  $-C_{\varepsilon_2} \frac{\varepsilon^2}{k} f_2(Re/M_{\infty})$  term in the  $\varepsilon$  equation is treated similarly, for consistency:

$$\frac{\varepsilon^{2(n+1)}}{k} \cong \frac{\varepsilon^{2(n)}}{k} + \frac{\partial}{\partial \varepsilon} \left( \frac{\varepsilon^2}{k} \right) \Delta \varepsilon$$

$$\cong \frac{\varepsilon^{2(n)}}{k} + \frac{\partial}{\partial \varepsilon} \left[ \left( \frac{C_{\mu} f_{\mu} \rho}{\mu_{\rm T}} \right)^{\frac{1}{2}} \varepsilon^{\frac{3}{2}} \right] \Delta \varepsilon$$

$$\cong \frac{\varepsilon^{2(n)}}{k} + \frac{3}{2} \left( \frac{C_{\mu} f_{\mu} \rho \varepsilon}{\mu_{\rm T}} \right)^{\frac{1}{2}} \Delta \varepsilon$$

$$\cong \frac{\varepsilon^{2(n)}}{k} + \frac{3}{2} \frac{\varepsilon}{k} \Delta \varepsilon$$

$$(H-189)$$

Menter's SST k –  $\omega$  model has an additional cross-diffusion term in the  $\omega$  equation. This term is treated implicitly (See reference 26).

$$C_{\omega}^{(n+1)} \cong C_{\omega}^{(n)} + \frac{\partial}{\partial \omega} C_{\omega} \Delta \omega$$

$$\cong C_{\omega}^{(n)} - \frac{|C_{\omega}|}{\omega} \Delta \omega$$
(H-190)

(The negative sign insures that, when taken to the left-hand side, this term increases the diagonal dominance of the implicit matrix.)

#### H.8.1 Example

This example shows the solution method for the k equation in Menter's SST model. All other equations are solved in a similar fashion.

$$\frac{\partial \mathbf{k}}{\partial t} + u_j \frac{\partial \mathbf{k}}{\partial x_j} = \frac{1}{\rho} \mathbf{P}_{\mathbf{k}} \left( \frac{M_{\infty}}{Re} \right) - \beta' \mathbf{k} \omega \left( \frac{Re}{M_{\infty}} \right) + \frac{1}{\rho} \frac{\partial}{\partial x_j} \left[ \left( \mu + \frac{\mu_{\mathrm{T}}}{\sigma_{\mathrm{k}}} \right) \frac{\partial \mathbf{k}}{\partial x_j} \right] \left( \frac{M_{\infty}}{Re} \right) \tag{H-191}$$

Written in generalized coordinates, with the production term treated explicitly (lagged in time), this equation becomes:

$$\begin{split} \frac{\partial \mathbf{k}^{(n+1)}}{\partial t} &= -U \frac{\partial \mathbf{k}^{(n+1)}}{\partial \xi} - V \frac{\partial \mathbf{k}^{(n+1)}}{\partial \eta} - W \frac{\partial \mathbf{k}^{(n+1)}}{\partial \zeta} \end{split} \tag{H-192} \\ &+ \frac{1}{\rho} P_{\mathbf{k}} \left( \frac{M_{\infty}}{Re} \right)^{(n)} - \beta' \mathbf{k} \omega \left( \frac{Re}{M_{\infty}} \right)^{(n+1)} \\ &+ \frac{1}{\rho} \left[ \frac{\partial}{\partial \xi} \left\{ (\xi_x^2 + \xi_y^2 + \xi_z^2) \left( \mu + \frac{\mu_T}{\sigma_k} \right) \frac{\partial \mathbf{k}}{\partial \xi} \right\} \right] \\ &+ \frac{\partial}{\partial \eta} \left\{ (\eta_x^2 + \eta_y^2 + \eta_z^2) \left( \mu + \frac{\mu_T}{\sigma_k} \right) \frac{\partial \mathbf{k}}{\partial \eta} \right\} \\ &+ \frac{\partial}{\partial \zeta} \left\{ (\zeta_x^2 + \zeta_y^2 + \zeta_z^2) \left( \mu + \frac{\mu_T}{\sigma_k} \right) \frac{\partial \mathbf{k}}{\partial \zeta} \right\} \right] \left( \frac{M_{\infty}}{Re} \right)^{(n+1)} \\ &= \mathrm{RHS}_k \end{split}$$

where the superscripts denote the time level. Note that the terms  $(\xi_x^2 + \xi_y^2 + \xi_z^2)$  etc. are not strictly correct as written. This is really short-hand notation. The correct way to expand the diffusion terms is given in "Generalized Coordinate Form" on page 294. This section also indicates how these metric terms are discretized.

$$U = \xi_x u + \xi_y v + \xi_z w + \xi_t$$
  

$$V = \eta_x u + \eta_y v + \eta_z w + \eta_t$$
  

$$W = \zeta_x u + \zeta_y v + \zeta_z w + \zeta_t$$
  
(H-193)

Define

$$\Delta k = k^{(n+1)} - k^{(n)}$$
(H-194)

Then

$$\frac{\Delta k}{\Delta t} = RHS_k \tag{H-195}$$

(H-196)

Linearize the right-hand-side terms that are taken at time level (n + 1):

$$-U\frac{\partial \mathbf{k}^{(n+1)}}{\partial \xi} \cong -U\frac{\partial \mathbf{k}^{(n)}}{\partial \xi} - U\delta_{\xi}^{upwind}(\Delta \mathbf{k})$$
(H-197)

$$-\left(\frac{Re}{M_{\infty}}\right)\beta'k\omega^{(n+1)} \cong -\left(\frac{Re}{M_{\infty}}\right)\beta'k\omega^{(n)} - \left(\frac{Re}{M_{\infty}}\right)\beta'\omega(\Delta k)$$
(H-198)

$$\left(\frac{M_{\infty}}{Re}\right) \frac{1}{\rho} \frac{\partial}{\partial \xi} \left\{ \chi_{\xi} \frac{\partial \mathbf{k}}{\partial \xi} \right\}^{(n+1)} \cong \left(\frac{M_{\infty}}{Re}\right) \frac{1}{\rho} \frac{\partial}{\partial \xi} \left\{ \chi_{\xi} \frac{\partial \mathbf{k}}{\partial \xi} \right\}^{(n)} + \left(\frac{M_{\infty}}{Re}\right) \frac{1}{\rho} \frac{\partial}{\partial \xi} \left\{ \chi_{\xi} \delta_{\xi}(\Delta \mathbf{k}) \right\}$$
(H-199)

where  $\chi_{\xi}$  represents  $(\xi_x^2 + \xi_y^2 + \xi_z^2) \left(\mu + \frac{\mu_T}{\sigma_k}\right)$ .

The equation is approximately factored and the contribution of the  $S_D$  term to the lefthand side is added in the first factor only.

$$\begin{bmatrix} \mathbf{I} + \Delta t U \delta_{\xi}^{upwind} + \Delta t \left(\frac{Re}{M_{\infty}}\right) \beta' \omega - \Delta t \left(\frac{M_{\infty}}{Re}\right) \frac{1}{\rho} \frac{\partial}{\partial \xi} (\chi_{\xi} \delta_{\xi}) \end{bmatrix}$$
(H-200)  
$$\begin{bmatrix} \mathbf{I} + \Delta t V \delta_{\eta}^{upwind} - \Delta t \left(\frac{M_{\infty}}{Re}\right) \frac{1}{\rho} \frac{\partial}{\partial \eta} (\chi_{\eta} \delta_{\eta}) \end{bmatrix}$$
$$\begin{bmatrix} \mathbf{I} + \Delta t W \delta_{\zeta}^{upwind} - \Delta t \left(\frac{M_{\infty}}{Re}\right) \frac{1}{\rho} \frac{\partial}{\partial \zeta} (\chi_{\zeta} \delta_{\zeta}) \end{bmatrix} \Delta \mathbf{k} = \Delta t \mathbf{R} \mathbf{H} \mathbf{S}_{\mathbf{k}}$$

This equation is solved in a series of three sweeps in each of the three coordinate directions:

$$\begin{bmatrix} I + \Delta t U \delta_{\xi}^{upwind} + \Delta t \left(\frac{Re}{M_{\infty}}\right) \beta' \omega - \Delta t \left(\frac{M_{\infty}}{Re}\right) \frac{1}{\rho} \frac{\partial}{\partial \xi} (\chi_{\xi} \delta_{\xi}) \end{bmatrix} \Delta k^{*} = \Delta t R H S_{k}$$

$$\begin{bmatrix} I + \Delta t V \delta_{\eta}^{upwind} - \Delta t \left(\frac{M_{\infty}}{Re}\right) \frac{1}{\rho} \frac{\partial}{\partial \eta} (\chi_{\eta} \delta_{\eta}) \end{bmatrix} \Delta k^{**} = \Delta k^{*}$$

$$\begin{bmatrix} I + \Delta t W \delta_{\zeta}^{upwind} - \Delta t \left(\frac{M_{\infty}}{Re}\right) \frac{1}{\rho} \frac{\partial}{\partial \zeta} (\chi_{\zeta} \delta_{\zeta}) \end{bmatrix} \Delta k = \Delta k^{**}$$

$$k^{(n+1)} = k^{(n)} + \Delta k$$
(H-201)

Each sweep requires the solution of a scalar tridiagonal matrix.

## H.9 Limiting

As a precaution, many of the turbulence quantities are limited in practice. Whether all of these conditions are necessary or not for the models to work and/or converge well is unknown. Some are probably included only because they were tried at some point and never removed.

- •In all models, the linearized S<sub>D</sub> term is added to the left-hand-side matrix only if its contribution is positive.
- In all models, the terms D,  $\hat{v}$ , k,  $\omega$ , and  $\varepsilon$  are not allowed to become negative. If the equation "wants" to yield a negative value, it is instead set to some very small positive number. (A counter also sums the number of times this happen and it is output in the \*.turres output file.)
- •In all the  $k \omega$  models as well as the  $k \varepsilon$  models for Abid and Gatski-Speziale, the production term in the k equation (i.e.

$$\frac{1}{\rho} \tau_{ij}^{\mathrm{T}} \frac{\partial u_i}{\partial x_i} \left( \frac{M_{\infty}}{Re} \right)$$

or an approximation thereof) is limited to be less than or equal to 20 times the destruction term in that equation.

- •In all EASM models, the  $\tau_{11}$ ,  $\tau_{22}$ , and  $\tau_{33}$  terms are limited to be positive (realiz-ability constraint).
- In all EASM models, the production term  $\tau_{ij}^{T} \frac{\partial u_i}{\partial x_j}$  is limited to be positive.
- In the Girimaji EASM models, the production term

$$\frac{1}{\rho} \tau_{ij}^{\mathrm{T}} \frac{\partial u_i}{\partial x_i} \left( \frac{M_{\infty}}{Re} \right)$$

is limited (in both the k and  $\varepsilon$  equations) to be less than or equal to  $40\varepsilon \left(\frac{Re}{M_{\infty}}\right)$ .

• In the Girimaji EASM models, the term

$$\varepsilon \left(\frac{Re}{M_{\infty}}\right) - 2\frac{\mu}{\rho} \left(\frac{\partial \sqrt{k}}{\partial x_j}\right)^2 \left(\frac{M_{\infty}}{Re}\right)^2$$

is limited to be positive.

•In all models, the nondimensional eddy viscosity  $\mu_T$  is limited to be less than or equal to 100,000.

- In the Gatski-Speziale EASM models,  $\eta^2$  and  $\zeta^2$  are each limited to be less than or equal to 10.
- •In the Gatski-Speziale EASM models,  $C_{\mu}$  is limited to be between 0.005 and 0.187.
- In the  $k \varepsilon$  models,  $f_{\mu}$  is limited to be less than or equal to 1.
- In the Girimaji EASM models,  $\eta_1$  and  $\eta_2$  are each limited to be less than or equal to 1200.
- In the Girimaji EASM model,  $G_1$  is limited to be between -0.005 and -0.2.
- In the Gatski-Speziale EASM  $k \omega$  models, " $\omega$ " in the denominator of the nonlinear terms in  $\tau_{ij}^{T}$  is limited to be greater than or equal to the free-stream value of  $\omega$ .
- •In the EASM  $k \varepsilon$  models, " $\varepsilon$ "in the denominator of the nonlinear terms in  $\tau_{ij}^{T}$  is limited to be greater than or equal to the free-stream value of  $\varepsilon$  only in terms that appear in the Navier-Stokes equations. This limiting is not done for the  $\tau_{ij}$  terms that appear in the S<sub>P</sub> terms in the k and  $\varepsilon$  equations.
- In the one- and two-equation models, the diffusion term in one coordinate direction, for example, can be written as

$$\frac{\partial}{\partial \xi} \left[ B \frac{\partial \chi}{\partial \xi} \right] = B_{i+\frac{1}{2}} \chi_{i+1} - \left( B_{i+\frac{1}{2}} + B_{i-\frac{1}{2}} \right) \chi_i + B_{i-\frac{1}{2}} \chi_{i-1}$$

In Baldwin-Barth and Spalart-Allmaras, the terms are limited as follows:

$$B_{i+\frac{1}{2}} = \max\left(B_{i+\frac{1}{2}}, 0\right)$$
$$B_{i-\frac{1}{2}} = \max\left(B_{i-\frac{1}{2}}, 0\right)$$

#### H.10 Wall Function

A simple approach for wall functions is taken from Abdol-Hamid, Lakshmanan, and Carlson<sup>1</sup>. This approach modifies the eddy-viscosity value  $\mu_{T, w}$  at the wall only. The wall shear stress is estimated from law-of-the-wall values at the first cell center off the wall and is used to determine an "equivalent" eddy viscosity at the wall.

Wall functions can be employed when the grid spacing is too coarse in the direction normal to the wall; typically  $y^+$  lies in the range  $30 < y^+ < 200$ . (The  $y^+$  value of the first grid point off the wall *should* be O(1) to ensure decent turbulent computations when no wall functions are used. See Section 2.2.) Wall functions are invoked in CFL3D by setting the value of **ivisc**(m) to be negative.

A caution: While wall functions can work very well, they can also sometimes cause problems. They are not strictly valid for separated flows, although many people use them anyway with reasonable success. It is recommended that wall functions *not* be used with the Baldwin-Lomax model in CFL3D. With other turbulence models, it is recommended that they be used sparingly, if at all. In our opinion, it is better to make a turbulent grid  $(y^+ = O(1))$  whenever possible and *avoid the use of wall functions*.

The wall function replaces the eddy viscosity at the ghost cell center at walls as follows:

$$\mu_{\text{T, ghost}} = 2\mu_{\text{T, w}} - \mu_{\text{T, 1}}$$

$$\mu_{\text{T, w}} = \mu_1 \left[ \frac{n^{+2}\mu_1}{\rho_1 \left(\frac{\partial u}{\partial y}\right)_1 d^2} \left(\frac{M_{\infty}}{Re}\right) - 1 \right]$$

$$\left(\frac{\partial u}{\partial y}\right)_1 = \frac{\sqrt{(u_1 - u_w)^2 + (v_1 - v_w)^2 + (w_1 - w_w)^2}}{d_1}$$
(H-202)

where  $d_1$  is the distance to the nearest wall from the first cell center off the wall and

for 
$$R_c \le 20.24$$
  $n^+ = \sqrt{R_c}$   
for  $20.24 < R_c \le 435$   $n^+ = a_0(1) + \sum_{n=2}^7 a_0(n)R_c^{n-1}$   
for  $435 < R_c \le 4000$   $n^+ = a_1(1) + \sum_{n=2}^5 a_1(n)R_c^{n-1}$  (H-203)  
for  $R_c > 4000$   $n^+ = a_2(1) + \sum_{n=2}^3 a_2(n)R_c^{n-1}$   
 $R_c = \frac{\rho_1 \sqrt{(u_1 - u_w)^2 + (v_1 - v_w)^2 + (w_1 - w_w)^2} d_1(\frac{Re}{M_{\infty}})$  (H-204)  
 $a_1(1) = 5.777191$   
 $a_0(1) = 2.354039$   $a_1(2) = 6.8756983 \times 10^{-2}$   
 $a_0(2) = 0.117984$   $a_1(3) = -7.1582745 \times 10^{-6}$   
 $a_0(4) = 2.0404148 \times 10^{-6}$   $a_1(4) = 1.5594904 \times 10^{-9}$   
 $a_0(4) = 2.0404148 \times 10^{-6}$   $a_1(5) = -1.4865778 \times 10^{-13}$  (H-205)

$$a_1(5) = -1.4865778 \times 10^{-13}$$
 (H-205)

$$a_0(5) = -5.1775775 \times 10^{-9}$$

$$a_0(6) = 6.2687308 \times 10^{-12}$$
  $a_2(1) =$   
 $a_0(7) = -2.916958 \times 10^{-15}$   $a_2(2) =$ 

$$a_2(1) = 31.08654$$
  
 $a_2(2) = 5.0429072 \times 10^{-2}$   
 $a_2(3) = -2.0072314 \times 10^{-8}$ 

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## APPENDIX I

# Angle Definitions

The definitions of  $\alpha$  and  $\beta$  are the same as that given in reference 20 with the exception that the direction of positive  $\beta$  is reversed. The figures below illustrate the difference.

## I.1 Standard-type Grid

(z is up and y is out the span)

This is the default for CFL3D-format grids (also applies when **ialph** = 0 for PLOT3D-format grids).

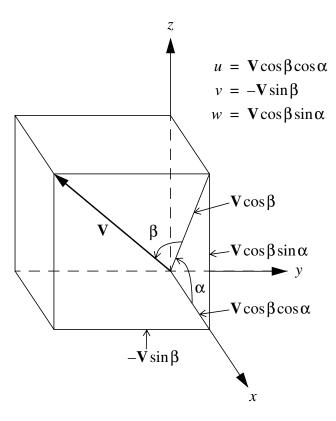


Figure I-1. Grid coordinate system for standard-type grids.

Drag = component of forces parallel to  $\mathbf{V}$  direction

Lift = component of forces perpendicular to  $V \cos \beta$  direction in the x - z plane

I.2 Non-standard-type Grid

(y is up and z is out the span)

(**ialph** > 0 for PLOT3D-format grids)

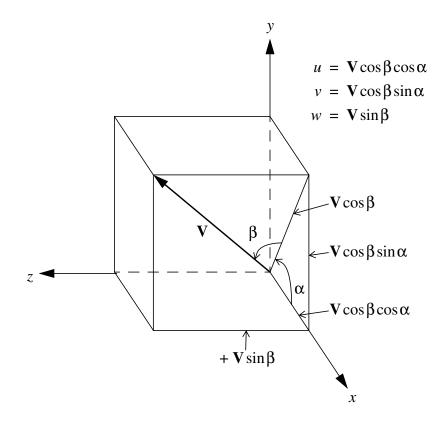


Figure I-2. Grid coordinate system for non-standard-type grids.

Drag = component of forces parallel to V direction

Lift = component of forces perpendicular to  $V \cos \beta$  direction in the x - y plane



#### APPENDIX J

## Subroutine Listing

CFL3D contains all the subroutines listed below:

CBSEM mgblk setup rp3d setblk global lead qinter qout qface plot3c plot3d plot3t update updateg resetg xtbatb grdmove resetwk rpatch setqc0 setdqc0 resadd readdat getdhdr	fluxp fluxm gfluxv hfluxv ffluxv dird wmag delv prntcp	dabcjz	<b>BC</b> bc bc1000 bc1001 bc1002 bc1003 bc1005 bc1008 bc1011 bc1012 bc1013 bc2002 bc2003 bc2004 bc2005 bc2006 bc2007 bc2102 chksym chkrap chkrot rield rielde blockk blockj blocki cblkk cblkj cblki int2 int3 rotateq rotateq0 rotateqb bcchk xupdt getibk intrbc ccf	LBCX tau tau2x colldat collx collv collq coll2q collqc0 collxt collxtb addx add2x add2x add2xv init initvist rrest wrest wrest wrestg metric tmetric cellvol ctime vlutr vlutrp bsub bsubp q8sdot q8smax q8smin q8ssum q8vrev l2norm l2norm2 force csurf csout rsmooth xmukin findmin	LBCX (continued) findmin_new bbdist bbdst1 calc_dist collect_surf distcc distcg initf initi ifree ffree iialloc ifalloc makebb calcbb getvrt shells spltbb push pop sort_x move_real move_integer triang heap_sort cntsurf celcen yplusout calyplus forceout histout	dynptch global2 patcher loadgr collapse rechk expand invert shear arc diagnos direct project extra extrae topol topol2 dsmin xe xe2 newfit trace	TURBS blomax barth3d spalart twoeqn triv
--	--	--------	---	--	---	---	---

#### J.1 CBSEM Routines

#### mgblk

Advances the solution in time using multigrid acceleration. At each level in the multigrid procedure, all the blocks are advanced before moving to the next level.

#### setup

Reads in the grid and restart data and calculates some preliminary information, such as the metrics, for subsequent use.

#### rp3d

Reads in the grids in PLOT3D format.

#### setblk

Initializes the blank array.

#### global

Reads in the case input file.

#### lead

Installs the attributes of a particular block into the common blocks.

#### qinter

Interpolates the solution from a coarser mesh to a finer mesh. The finer mesh can be either a global mesh or an embedded mesh. Also updates grid position of finer mesh if meshes are in motion.

#### qout

Outputs data for plotting or printing.

#### qface

Determines the primitive variables at the edges of the grid and installs them in the qj0/qk0/qi0 arrays for use in output routines.

#### plot3c

Writes the output file for the cell-centered points in PLOT3D format. (Outputs the grid and/or solution in single precision for use with FAST and/or PLOT3D.) Also prints solution data to a printout file for a specified range of points.

#### plot3d

Writes the output file at the grid points in PLOT3D format. (Outputs the grid and/or solution in single precision for use with FAST and/or PLOT3D.) Also prints solution data to a printout file for a specified range of points.

#### plot3t

Writes turbulent information for the cell-centered points in PLOT3D format.

#### update

Updates the solution in time.

#### updateg

Updates the grid to a new position and obtains corresponding grid-boundary velocities for use in the boundary conditions. Also collocates new grid position to coarser levels and obtains grid-boundary velocities on coarser levels. Also updates moment center.

#### resetg

Checks to see if any blocks in the grid have been translated and/or rotated out of the bounds set in the input deck. If so, resets these blocks so that they are at or near the initial positions. If the rotational displacement of a block is reset, the solution must also be rotated to correspond to the reset position. (Resetting is allowed only for constant translational speed, **itrans** = 1, or constant rotational speed, **irotat** = 1.)

#### xtbatb

Stores grid speeds and accelerations on the boundaries for use in setting no-slip and wall pressure boundary conditions for dynamic meshes.

#### grdmove

Moves the grid from one position to another.

#### resetwk

Replaces all locations of the work array which were filled with integer values in subroutine plot3d with real values. (Otherwise, problems may arise later when a real array attempts to access the integer value located in memory.)

#### rpatch

Reads in the generalized-coordinate interpolation data for grid patching from a file.

#### setqc0

Stores conservative variables for use in second-order temporal differencing and subiteration.

#### setdqc0

Stores conservative variables  $(\mathbf{Q}_n - \mathbf{Q}_{n-1})$  for use in second-order temporal differencing.

#### resadd

Adds additional terms to the right-hand side for sub-iteration and second-order temporal accuracy.

#### readdat

Reads in auxiliary data arrays for the "2000 series" of boundary conditions.

#### getdhdr

Sets the character data for the main output file headers when the "2000" series of boundary conditions are used.

#### J.2 RHS Routines

#### resid

Computes the residual contributions to the right-hand side.

#### resp

Computes and prints residuals and sums the forces.

#### fa

Accumulates fluxes to insure conservation with grid embedding.

#### i2x

Interpolates the primitive variables from coarser meshes onto twice finer meshes for grid embedding.

#### i2xs

Interpolates the primitive variables from coarser meshes onto twice finer meshes for grid embedding.

#### fa2xj

Accumulates fluxes in the j direction for use on a twice coarser mesh to insure conservation with grid embedding.

#### fa2xk

Accumulates fluxes in the k direction for use on a twice coarser mesh to insure conservation with grid embedding.

#### fa2xi

Accumulates fluxes in the i direction for use on a twice coarser mesh to insure conservation with grid embedding.

#### gfluxr

Computes residual contributions for the right-hand side in the j direction from the inviscid terms.

#### hfluxr

Computes residual contributions for the right-hand side in the k direction from the inviscid terms.

#### ffluxr

Computes residual contributions for the right-hand side in the i direction from the inviscid terms.

#### xlim

Performs monotone interpolations to the interfaces of the cells.

#### fhat

Computes Roe's<sup>31</sup> generalized flux at the interface given the left and right states at the interface.

#### fill

Fills the edges of the q array for safety using a multi-plane vectorization technique.

#### fluxp

Computes the "positive" parts of the fluxes using the flux-vector-splitting method of van Leer.<sup>39</sup>

#### fluxm

Computes the "negative" parts of the fluxes using the flux-vector-splitting method of van Leer.<sup>39</sup>

#### gfluxv

Calculates the right-hand-side residual contributions in the j direction due to the viscous terms.

#### hfluxv

Calculates the right-hand-side residual contributions in the k direction due to the viscous terms when idf = 0. Calculates the implicit matrix terms when idf > 0.

#### ffluxv

Calculates the right-hand-side residual contributions in the i direction due to the viscous terms.

#### dird

Evaluates directed distance from the k = 0 wall and the i = 0/j = 0 wall for use in evaluating the Baldwin-Lomax<sup>10</sup> turbulence model.

#### wmag

Evaluates the vorticity magnitude for use in determining the turbulent eddy viscosity.

#### delv

Evaluates the velocity derivatives at cell centers.

#### prntcp

Writes the pressures on the body (actually the cell centers closest to the body) to an output file for unsteady flow.

#### trans

Determines the increment to grid position due to a grid translation.

#### transmc

Determines the increment to moment center due to a grid translation.

#### rotate

Determines the increment to grid position due to a grid rotation.

#### rotatmc

Determines the increment to moment center due to a grid rotation.

#### hole

Zeroes out the right-hand-side residuals for the blanked points.

#### dthole

Updates the  $\Delta t$  values for the hole and fringe cells. The values will be replaced with  $\Delta t_{\min}$ .

#### blkmax

Determines the location of the maximum residual.

#### J.3 AF3F Routines

#### af3f

Advances the solution in time using a 3-factor spatially-split approximate factorization algorithm.

#### amafj

Formulates the implicit matrices in the j direction for the 3-factor algorithm.

#### swafj

Solves the block  $5 \times 5$  tridiagonal equations for the 3-factor spatially-split algorithm in the *j* direction.

#### amafk

Formulates the implicit matrices in the k direction for the 3-factor algorithm.

#### swafk

Solves the block  $5 \times 5$  tridiagonal equations for the 3-factor spatially-split algorithm in the *k* direction.

#### amafi

Formulates the implicit matrices in the i direction for the 3-factor algorithm.

#### swafi

Solves the block  $5 \times 5$  tridiagonal equations for the 3-factor spatially-split algorithm in the *i* direction.

#### diagj

Solves the scalar tridiagonal equations to approximate the spatially-split factor in the j direction of the 3-D spatially-split algorithm.

#### diagk

Solves the scalar tridiagonal equations to approximate the spatially-split factor in the k direction of the 3-D spatially-split algorithm.

#### diagi

Solves the scalar tridiagonal equations to approximate the spatially-split factor in the *i* direction of the 3-D spatially-split algorithm.

#### tinvr

Multiplies the inverse of the diagonalizing matrix **T** times the residual contribution  $(\mathbf{T}^{-1}R)$ .

#### tdq

Multiplies the inverse of the diagonalizing matrix **T** times the change in characteristic combination of variables  $(\mathbf{T}^{-1}\Delta \mathbf{q})$ .

#### dlutr

Performs the scalar tridiagonal (LU) decomposition.

#### dfbtr

Performs the back substitution for a scalar tridiagonal system of equation.

#### dlutrp

Performs the LU decomposition for a periodic scalar tridiagonal system of equations.

#### dfbtrp

Performs the back substitution for a periodic scalar tridiagonal system of equations.

#### dfhat

Computes a Jacobian matrix with respect to the primitive variables at the cell interface. The Jacobian evaluation is approximate, being taken as either  $A^+$  or  $A^-(T\Lambda T^{-1})$ , and is computed with metric terms from the interface and dependent variables fro the cell centers.

#### dfluxpm

Computes "positive" or "negative" parts of the flux Jacobians using the flux-vector-splitting method of van Leer.<sup>39</sup>

#### gfluxl

Computes the left-hand flux contributions due to the inviscid terms for the j direction.

#### hfluxl

Computes the left-hand flux contributions due to the inviscid terms for the k directions.

#### ffluxl

Computes the left-hand flux contributions due to the inviscid terms for the i direction.

#### abciz

Zeroes out the left-hand-side matrix element with the help of the blank array. For a point with blank = 0, all elements of matrices a and c become zero. Only diagonal elements matrix b is changed to 1.0 for *i* implicit; *j* sweep.

#### abcjz

Zeroes out the left-hand-side matrix element with the help of the blank array. For a point with blank = 0, all elements of matrices a and c become zero. Only diagonal elements matrix b is changed to 1.0 for *j* implicit; *k* sweep.

#### abckz

Zeroes out the left-hand-side matrix element with the help of the blank array. For a point with blank = 0, all elements of matrices a and c become zero. Only diagonal elements matrix b is changed to 1.0 for k implicit; j sweep.

#### dabciz

Uses the blank values to modify the coefficient matrices, a, b, c, for the diagonal inversion in the i direction.

#### dabcjz

Uses the blank values to modify the coefficient matrices, a, b, c, for the diagonal inversion in the j direction.

#### dabckz

Uses the blank values to modify the coefficient matrices, a, b, c, for the diagonal inversion in the k direction.

#### J.4 BC Routines

#### bc

Determines boundary data/conditions at edges of grids.

#### bc1000

Sets free-stream boundary conditions.

#### bc1001

Sets symmetry plane boundary conditions.

#### bc1002

Sets extrapolation boundary conditions.

#### bc1003

Sets characteristic inflow/outflow boundary conditions.

#### bc1005

Sets inviscid surface boundary conditions.

#### bc1008

Sets tunnel inflow boundary conditions.

#### bc1011

Sets singular axis (half plane) boundary conditions.

#### bc1012

Sets singular axis (full plane) boundary conditions.

#### bc1013

Sets extrapolation boundary conditions for a singular axis.

#### bc2002

Sets pressure ratio; extrapolates other quantities.

#### bc2003

Sets characteristic inlet boundary conditions at engine inlet given (estimated) inlet Mach number, total pressure ratio, total temperature ratio, and flow angle.

#### bc2004

Sets solid wall (viscous wall) boundary conditions.

#### bc2005

Sets periodic boundary conditions given angular rotation angle to the periodic face and its block number.

#### bc2006

Sets pressure via radial equilibrium condition; extrapolates other quantities.

#### bc2007

Sets all the primitive variables with standard CFL3D normalization;  $\tilde{\rho}/\tilde{\rho}_{\infty}$ ,  $\tilde{u}/\tilde{a}_{\infty}$ ,

$$\tilde{v}/\tilde{a}_{\infty}, \tilde{w}/\tilde{a}_{\infty}, \tilde{p}/(\tilde{\rho}_{\infty}\tilde{a}_{\infty}^2).$$

#### bc2102

Sets the pressure ratio as a function of time; extrapolates the other flow-field quantities.

#### chksym

Checks for symmetry boundary conditions in j, k, or i directions in order to apply boundary condition type 1011 (singular axis with half-plane symmetry).

#### chkrap

Checks for wrap-around in j, k, or i directions in order to apply boundary condition type 1012 (singular axis with full plane).

#### chkrot

Checks to make sure that the proper rotation angle for periodic boundary conditions has been set.

#### rield

Determines far-field boundary data using quasi-1-D characteristic relations for boundary condition type 1003.

#### rie1de

Determines far-field boundary data using quasi-1-D characteristic relations for boundary condition type 2003.

#### blockk

Transfers information from the block designated ir to the qk0 array of the block designated it.

#### blockj

Transfers information from the block designated ir to the qj0 array of the block designated it.

#### blocki

Transfers information from the block designated ir to the qi0 array of the block designated it.

#### cblkk

Checks information transferred from the block designated ir to the qk0 array of the block designated it.

#### cblkj

Checks information transferred from the block designated ir to the qj0 array of the block designated it.

#### cblki

Checks information transferred from the block designated ir to the qi0 array of the block designated it.

#### int2

Linearly interpolates q from one grid to ghost cells of another grid using generalized coordinates without using a limiter on the gradients.

#### int3

Linearly interpolates q from one grid to ghost cells of another grid using generalized coordinates using a limiter on the gradients.

#### rotateq

Rotates the solution contained in array q through a specified angle and stores the rotated solution in grot.

#### rotateq0

Rotates the solution at ghost points contained in array q0 (either qi0, qj0, or qk0 through the angle  $\Delta \theta_x / \Delta \theta_y / \Delta \theta_z$  and stores the rotated solution in q0rot.

#### rotateqb

Rotates the solution in the qb array through the angle  $\Delta \theta_x / \Delta \theta_y / \Delta \theta_z$  for the chimera scheme with rotating grids.

#### bcchk

Determines if the boundary conditions have been set and fills the end-points for safety.

#### xupdt

Updates the fringe points of overlapped grids with boundary values which have been interpolated from other grids to provide the mechanism for coupling the various grids.

#### getibk

Reads the output from MaGGiE (but not the grids).

#### intrbc

Interpolates the corrections for boundary values for all grids overlapped in the current mesh.

#### ccf

Modifies  $u_f$ ,  $w_f$ ,  $a_f$  (velocities and speed of sound at the far field) based on point vortex correction (used when **i2d** = -1 and the far-field boundary condition type is 1003).

#### J.5 LBCX Routines

#### tau

Computes a residual correction and stores the values of q for later use in determining corrections to finer grids in the multigrid iteration scheme.

#### tau2x

Puts the restricted residual from a finer embedded mesh into a coarser mesh.

#### colldat

Restricts auxiliary boundary condition data arrays to coarser meshes.

#### collx

Restrict x, y, and z values to coarser meshes.

#### collv

Restricts volumes to coarser meshes.

#### collq

Restricts q (the primitive variables) with a volume-weighted interpolation and residuals to coarser meshes. Also restricts turbulent eddy viscosity in the case of turbulent flows to coarser meshes.

#### coll2q

Restricts q (the primitive variables) with a volume-weighted interpolation and residuals from finer embedded meshes to coarser meshes.

#### collqc0

Restricts conservative variables  $\mathbf{Q}_n$  and  $\mathbf{Q}_n - \mathbf{Q}_{n-1}$  to coarser meshes via summation over fine-grid cells for use in time-accurate multigrid.

#### collxt

Restricts xt (array containing grid speeds) to coarser meshes.

#### collxtb

Restricts xtb and atb (arrays containing grid boundary velocity and acceleration, respectively) to coarser meshes.

#### addx

Interpolates the solution or the correction from a coarser mesh to a finer mesh.

#### addxv

Interpolates the turbulence quantities from a coarser mesh to a finer mesh during mesh sequencing.

#### add2x

Interpolates the solution or the correction from a coarser mesh to a finer embedded mesh.

#### add2xv

Interpolates the turbulence quantities from a global mesh to an embedded mesh during mesh sequencing.

#### init

Sets the initial conditions on a mesh to be free stream.

#### initvist

Sets the turbulent initial conditions on a mesh.

#### rrest

Reads the restart file for a block.

#### rrestg

Reads the restart file to get the latest position for a dynamic mesh, along with corresponding metrics and grid-boundary speeds. Also reads qc0 for a second-order accurate in time restart.

#### wrest

Writes the restart file for a block.

#### wrestg

Appends the latest position of a dynamic mesh to the end of the restart file. Also writes qc0 for a second-order accurate in time restart.

#### metric

Calculates the cell-interface directed areas.

#### tmetric

Calculates the time-metric terms for a grid in motion.

#### cellvol

Calculates the cell volumes.

#### ctime

Calculates the time step for an input fixed Courant number or calculates the Courant number based on an input value of  $\Delta t$ .

#### vlutr

Performs the LU decomposition for a block  $5 \times 5$  tridiagonal system of equations.

#### vlutrp

Performs the LU decomposition for a block  $5 \times 5$  tridiagonal system of equations which is periodic.

#### bsub

Performs the back substitution for a block  $5 \times 5$  tridiagonal matrix equation solution.

#### bsubp

Performs the back substitution for a block  $5 \times 5$  tridiagonal matrix equation solution which is periodic.

q8sdot (function)

Computes the dot product between two vectors.

q8smax (function)

Finds the maximum value in an array.

```
q8smin (function)
```

Finds the minimum value in an array.

```
q8ssum (function)
```

Sums the elements of a vector

#### q8vrev

Reverses the elements in an array.

#### 12norm

Computes the L2-norm of the residuals or the change in primitive variables from one time to the next.

#### 12norm2

Computes the L2-norm of the residuals, after subtracting out the contribution of the unsteady terms that were added in subroutine resadd.

#### force

Integrates the forces on the body.

#### csurf

Integrates control surface mass flow and momentum/forces.

#### csout

Writes control surface mass flow and momentum/forces to an output file.

#### rsmooth

Performs implicit residual smoothing (constant coefficient).

#### xmukin

Computes Sutherland's formula with linear law at low temperatures.

#### findmin

Finds minimum distances from field points to viscous wall(s).

#### findmin\_new

Serves as a driver routine for computing distances to the closest viscous surface.

#### bbdist

Serves as a driver routine for determining the nearest bounding box for each field point.

#### bbdst1

Identifies the nearest bounding box for each field point.

#### calc\_dist

Finds the minimum distance to field points using the recursive-box algorithm.

#### collect\_surf

Stores coordinates of surface points and identifies the neighbors of each surface point.

#### distcc

Averages the distance function based at grid points to one based at cell centers.

#### distcg

Collocates the fine-grid minimum distance function on a fine grid to a coarser grid.

#### initf

Initializes pointers to floating-point variables in recursive-box algorithm.

#### initi

Initializes pointers to integer variables in recursive-box algorithm.

#### ifree

"Frees up" pointers to integer variables in recursive-box algorithm.

#### ffree

"Frees up" pointers to floating-point variables in recursive-box algorithm.

#### iialloc (function)

Increments pointers to integer variables in recursive-box algorithm.

#### ifalloc (function)

Increments pointers to floating-point variables in recursive-box algorithm.

#### makebb

Serves as a driver routine for generating the bounding boxes for the recursive-box algorithm.

#### calcbb

Calculates bounding boxes.

#### getvrt

Searches for all points that fall within a bounding box.

#### shells

Performs a shell sort.

#### spltbb

Subdivides bounding boxes.

#### push

Places an item into the stack and adjusts the pointer accordingly.

#### рор

Removes an item from the stack and adjusts the pointer accordingly.

#### sort\_x

Sorts surface points with respect to *x*-coordinate.

move\_real

Rearranges items in the real array x based on the pointer iperm.

#### move\_integer

Rearranges items in the integer array x based on the pointer iperm.

#### triang

Finds the closest distance from a field point to the actual surface (i.e. not simply the closest discrete surface point), using local triangulation of the surface.

```
heap_sort
```

Sorts a list of points.

#### cntsurf

Counts the number of viscous surface points.

#### celcen

Finds cell centers of field points.

#### yplusout

Calls the routines necessary for calculating yplus at the first grid point above solid walls.

#### calyplus

Calculates  $y^+$  in turbulent flows at solid surfaces in a block and calculates statistics for the  $y^+$  distribution (average  $y^+$ , maximum  $y^+$  and its location, standard deviation).

#### forceout

Writes the forces and moments on individual blocks, as well as a global force/moment summary, to an output file.

#### histout

Writes the convergence history for mean-flow equations and turbulence equations to an output file.

#### J.6 DYNPTCH Routines

#### dynptch

Establishes zone-to-zone communication for block interfaces that move relative to one another, using a patched-grid technique (nonconservative).

#### global2

Reads the dynamic patch input parameters.

#### patcher

Calculates patched-grid interpolation coefficients.

#### loadgr

Loads the proper grid from a 1-D storage array to a 2-D work array.

#### collapse

Checks for collapsed points in the grid and expands any collapsed lines detected.

#### rechk

Checks for branch cuts.

#### expand

Expands the grid at boundaries.

#### invert

Determines generalized coordinates of cell centers of the "to" grid in terms of the generalized coordinate system(s) defined on the "from" grid(s).

#### shear

Determines the generalized coordinates of cell edge midpoints on  $\xi = 0$  and  $\eta = 0$  boundaries and determines the requisite shearing correction to the generalized coordinates near  $\xi = 0$  and/or  $\eta = 0$  boundaries.

#### arc

Performs arc-length correction to the generalized coordinates near a boundary if required when shearing correction has failed.

#### diagnos

Performs diagnostic checks on interpolation coefficients (generalized coordinates) found via search and inversion routines.

#### direct

Computes normalized directed area components, or equivalently, components of the unit normal to a cell face.

project

Projects a point onto a plane.

#### extra

Computes extra mid-cell points in  $\xi$  direction.

#### extrae

Computes extra mid-cell points in  $\boldsymbol{\eta}$  direction.

#### topol

Searches appropriate "from" blocks for current "to" cell center. Driver routine for determining  $\xi$  and  $\eta$  of the cell center.

#### topol2

Searches appropriate "from" blocks in one direction only for current "to" cell center

#### dsmin

Finds the closest point in a grid to a specified point.

#### xe

Selects proper coordinates to use for inversion.

#### xe2

Sets up the coefficients for locally fitting a polynomial variation in the  $\xi$  and  $\eta$  directions.

#### newfit

Determines a new polynomial fit for cells with stubborn convergence.

#### trace

Writes the search routine history for the current "to" cell to unit 7.

#### transp

Translates the "from" block to provide complete coverage for interpolation for cases in which the complete physical domain is not modeled.

#### rotatp

Rotates "from" block to provide complete coverage for interpolation for cases in which the complete physical domain is not modeled.

#### J.7 TURBS Routines

#### blomax

Computes the turbulent viscosity distributions using the Baldwin-Lomax<sup>10</sup> two-layer eddy-viscosity model.

#### barth3d

Computes the turbulent viscosity distributions using the one-equation Baldwin-Barth<sup>9</sup> turbulence model.

#### spalart

Computes the turbulent viscosity distributions using the one-equation Spalart<sup>35</sup> turbulence model.

#### twoeqn

Computes the turbulent viscosity distributions using the two-equation turbulence models.

#### triv

Solves a scalar tridiagonal system of equation



#### APPENDIX K

## Version Differences

### K.1 Significant Differences Between Versions 5.0 and 4.1

- 1. The most significant difference between the two versions is the generalization of moving grid capability and the added capability to handle sliding patched interfaces.
- 2. The viscous wall boundary condition now sets wall temperature locally rather than globally.
- 3. The capability to output control surface information and turbulence quantity information has been added.
- 4. There are slight differences in how the  $k \omega$  wall boundary conditions are set internal to the code (the new way is more "correct"). This change may result in very small differences in convergence, but the end result should be nearly identical to the old way.
- 5. The Baldwin-Lomax model can be applied on **jdim/kdim/idim** bodies now.
- 6. In 2-d, the "far-field point-vortex correction" boundary condition can now be applied. Set i2d = -1, in combination with 1003.
- 7. A few more boundary conditions have been added.
- 8. Several more turbulence models have been added. Also, the "SSTZ" version of the SST  $k \omega$  model has been removed.
- 9. A crude wall function capability has been added.
- 10. Some very minor differences, which may result in nearly insignificant differences between Version 4.1 and Version 5.0 results, are:
  - (a) For the Wilcox  $k \omega$  model, Version 5.0 limits the production term in the k equation to be less than 20 times the destruction term. Version 4.1 does not.
  - (b) In Version 4.1, the input alpha and beta angles (in degrees) are transformed to radians by dividing by 57.2958. In Version 5.0, the transformation (using parameter radtodeg) is more precise.
- 11. Both the input file and the restart file are different. See the following section for details on the differences in the input file. Any existing Version 4.1 input file or restart file can be automatically transformed to a Version 5.0 file by using v4tov5\_input.f or v4tov5\_restart.f, located in the **Tools** directory. (See "The Code and Supplementary Files" on page 9.)

#### K.2 Summary of Changes to the Input File

Version 5.0 of CFL3D originated when the generalization of moving grid capability (including sliding patched interfaces) was added to Version 4.1. As a consequence, the method for controlling the movement of any dynamic grids, patched or not, is now controlled from a new "section" at the end of the input file. Other than this, most changes to the input file, for either improved clarity or added capability, are relatively minor. After obtaining Version 5.0 of CFL3D, veteran users may notice some minor differences from Version 4.1. For those changes involving new input parameters, be sure to read about them in Chapter 3. Any existing Version 4.1 input file can be automatically transformed to a Version 5.0 input file by using v4tov5\_input.f, located in the **Tools** directory. (See "The Code and Supplementary Files" on page 9.)

In the following, note that sample numerical values, which of course are case dependent, are included.

1. Line Type Three changed from

	<b>V</b> I	U					
	XMACH	ALPHA	BETA	REUE,MIL	TINF,DR	ISND	C2SPE
	0.8750	00.000	0.0	02.660	460.0	0	0.0
to							
	XMACH	ALPHA	BETA	REUE,MIL	TINF,DR	IALPH	IHIST
	0.8750	00.000	0.0	02.660	460.0	0	0

(Note: **isnd** and **c2spe** now handled by **Twtype** in the boundary condition section.)

2. Line Type Five changed from

1

1

		-02.000	0	000	05.0000	0	10.	
		DT	IREST	IFLAGTS	FMAX	IUNST	CFLTAU	
	to							
0	.50000		0	000	03.0000	0	0.39000	0.22200
С	LOC	-02.000	0	000	05.0000	0	0.39600	0.22200
		DT	IREST	IFLAGTS	FMAX	IUNST	RFREQ	ALPHAU

(Note: Unsteady grid motion is now handled in Line Types Thirty-Three through Fourth-Five)

3. Any boundary condition (Line Types Fourteen through Nineteen) with viscous wall (**bctype** 1004) changed from

1	1004	0	0	0	0	0
1 TWTYPE 0	2004 CQ 0	0	0	0	0	2

4. The 1-1 blocking section of the input file has been modified so that the parameter names are more pertinent and so that the bookkeeping of 1-1 interfaces is simplified. Line Types Twenty-Four through Twenty-Seven have been changed from

```
1-1 BLOCKING DATA:
          NBLI
         NBLON
                 NBLK(1)
                            NBLK(2)
             0
                        1
             0
                        3
                                    4
    LIMBLK(1) LIMBLK(2) LIMBLK(3) LIMBLK(4) LIMBLK(5) LIMBLK(6) ISVA(1,1)
ISVA(1,2)
             1
                       1
                                  1
                                             2
                                                      41
                                                                  1
                                                                            1
                                                                                       2
```

	1 LIMBLK(2)	1 LIMBLK(3)	2 LIMBLK(4)	63 LIMBLK(5)	1 LIMBLK (	6) ISN	1 VA(2,1)	2
ISVA(2,2) 1 1	257 1	1 51	2 2	217 63	1 51		1 2	2 3
to								
1-1 BLOCKING DATA: NBLI 2								
NUMBER 1 2 NUMBER 1	GRID : 1 3 GRID : 1	1 1	JSTA KS 1 JSTA KS 257	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	JEND 41 63 JEND 217	KEND 1 KEND 1	ISVA1 1 ISVA1 1	ISVA2 2 2 ISVA2 2
2	4	1	1	51 2	63	51	2	3

The parameter **nblon** is no longer used (it is always 0). The values for nblk(1) and nblk(2) are now assigned under **grid** in Line Types Twenty-Five and Twenty-Six, respectively. The parameter **number** is intended to help the user keep track of interface numbers. This is particularly useful when there are a very large number of 1-1 interfaces. This parameter is not used internal to the code.

5. After **nprint** section, the following Line Types (Thirty-One and Thirty-Two) must be added:

```
CONTROL SURFACE:
NCS
0
```

GRID ISTART IEND JSTART JEND KSTART KEND IWALL INORM

These lines (in combination with **ihstry**) control output of control surface information, such as mass flow. (Note: When using a control surface for print out, then **ncs** 0 and there will be at least one additional line under **grid**...)

6. The above items 1 - 5 are all that are *necessary* to transform an existing Version 4.1 input file (with non-moving grid) to a Version 5.0 input file. For a moving grid, Version 5.0 now has the additional section Line Types Thirty-Three through Forty-Five, required if and only if **iunst** > 0.

# P

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Equation (1) should be:  

$$\sigma \mu_t = \mu + \sigma \mu_t$$

$$\sigma^* \mu_t = \mu + \sigma^* \mu_t$$

Equation (17) should be:  $\Gamma = \max(2\Gamma_3, \Gamma_1)$ 

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