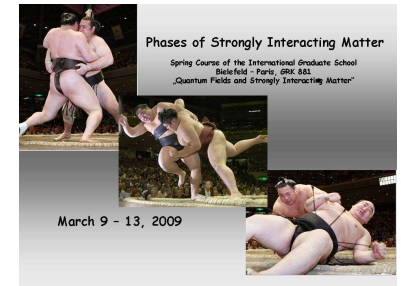


Andreas Schmitt

Institut für Theoretische Physik  
Technische Universität Wien  
1040 Vienna, Austria



## Color superconductivity in dense quark matter

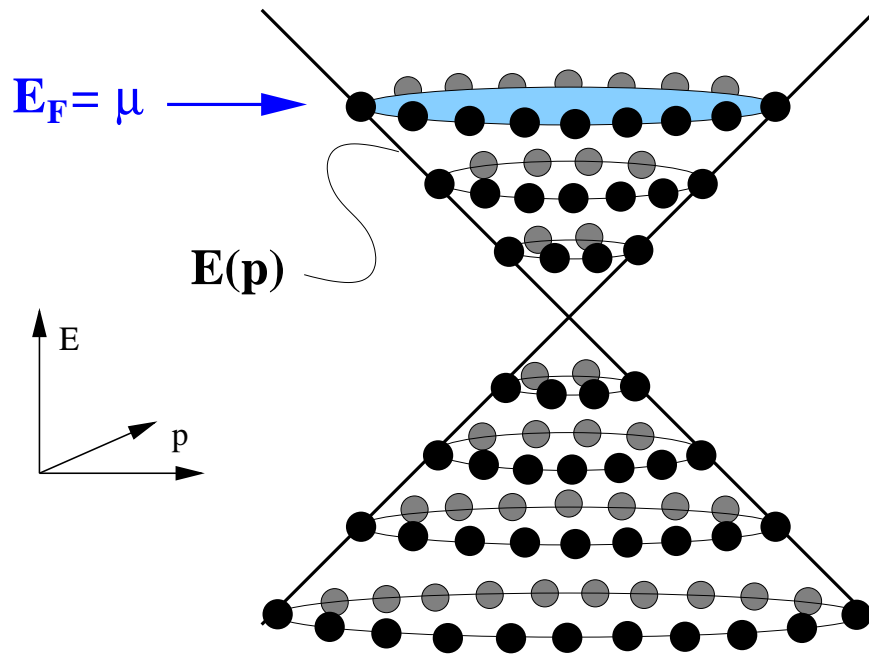
For a review, see

M. Alford, K. Rajagopal, T. Schäfer, A. Schmitt, *Rev. Mod. Phys.* 80, 1455 (2008)

- Basics of (color) superconductivity
- Color-flavor locking (CFL) – highest densities
- Stressed pairing – below CFL densities
- QCD calculations
- Effective theory of CFL – kaon condensation
- Transport properties – quark matter in compact stars

- **Part I**
  
- **Basics of (color) superconductivity**
- Color-flavor locking (CFL) – highest densities
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## • Basics of (color) superconductivity (page 1/5)



$$\text{free energy } \Omega = E - \mu N$$

- no interactions: add fermion at  $E = \mu$  without cost
- attractive interaction: add pair with gain
- pairs condense  
→ superconductivity

This Bardeen-Cooper-Schrieffer (BCS) argument holds for electrons in metal,  $^3\text{He}$  atoms, . . . , and quarks in quark matter

- Basics of (color) superconductivity (page 1/5)

## Fermi sphere



$$\text{free energy } \Omega = E - \mu N$$

- no interactions: add fermion at  $E = \mu$  without cost
- attractive interaction: add pair with gain
- pairs condense  
→ superconductivity

This Bardeen-Cooper-Schrieffer (BCS) argument holds for electrons in metal,  $^3\text{He}$  atoms, . . . , and quarks in quark matter

- Basics of (color) superconductivity (page 1/5)

## Cooper pairing



free energy  $\Omega = E - \mu N$

- no interactions: add fermion at  $E = \mu$  without cost
- attractive interaction: add pair with gain
- pairs condense  
→ superconductivity

This Bardeen-Cooper-Schrieffer (BCS) argument holds for electrons in metal,  $^3\text{He}$  atoms, . . . , and quarks in quark matter

- **Basics of (color) superconductivity (page 2/5)**  
**Electromagnetic vs. color superconductor**

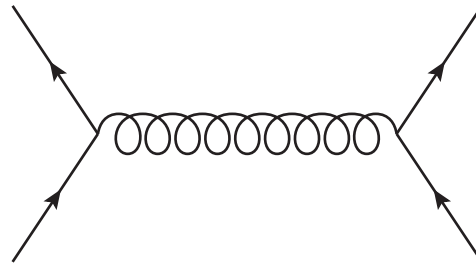
	<b>Where?</b>	<b>What?</b>	<b>Attractive force</b>	<b>Cooper pairs</b>	<b>Meissner effect</b>
“usual” superconductor	metals, alloys	ion lattice & electrons	phonons	electrons	photon
<b>color superconductor</b>	<b>neutron stars</b>	<b>quarks &amp; gluons</b>	<b>gluons</b>	<b>quarks</b>	<b>gluons (and photon)<sup>(*)</sup></b>

(\*) Most color superconductors are not electromagnetic superconductors (“rotated electromagnetism”)

Exception: **Spin-1 color superconductors**

A. Schmitt, Q. Wang, D. H. Rischke, PRL 91, 242301 (2003)

- Basics of (color) superconductivity (page 3/5): attractive quark-quark interaction
- one-gluon exchange



attractive in antisymmetric antitriplet channel  $[\bar{\mathbf{3}}]_c^a$

$$SU(3)_c : \quad [\mathbf{3}]_c \otimes [\mathbf{3}]_c = [\bar{\mathbf{3}}]_c^a \oplus [\mathbf{6}]_c^s$$

- flavor space

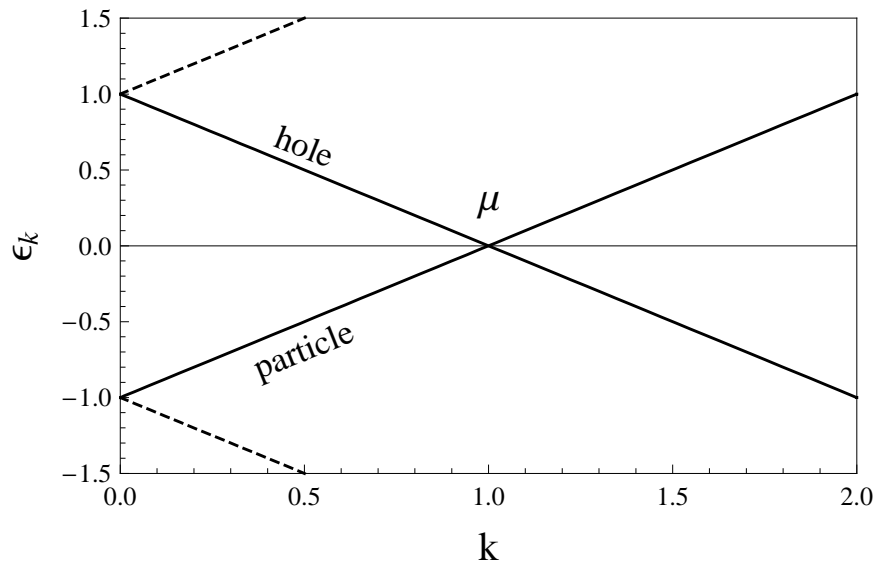
$$SU(3)_f : \quad [\mathbf{3}]_f \otimes [\mathbf{3}]_f = [\bar{\mathbf{3}}]_f^a \oplus [\mathbf{6}]_f^s$$

- order parameter (for spin-0 pairing):

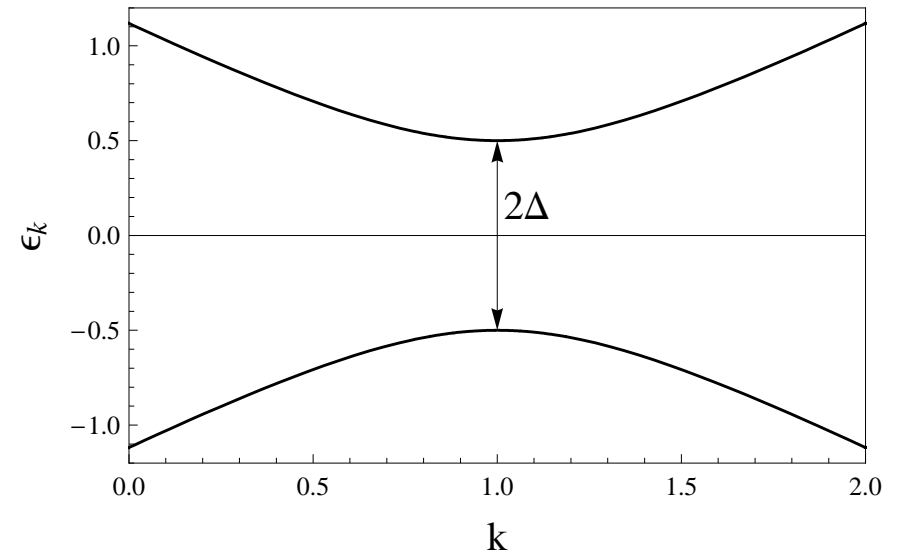
$$\langle \psi_i^\alpha C \gamma_5 \psi_j^\beta \rangle \propto \epsilon^{\alpha\beta A} \epsilon_{ijB} \phi_B^A \in [\bar{\mathbf{3}}]_c^a \otimes [\bar{\mathbf{3}}]_f^a$$

- Basics of (color) superconductivity (page 4/5): energy gap

## Fermion dispersions



$$\epsilon_k = \pm(k - \mu)$$



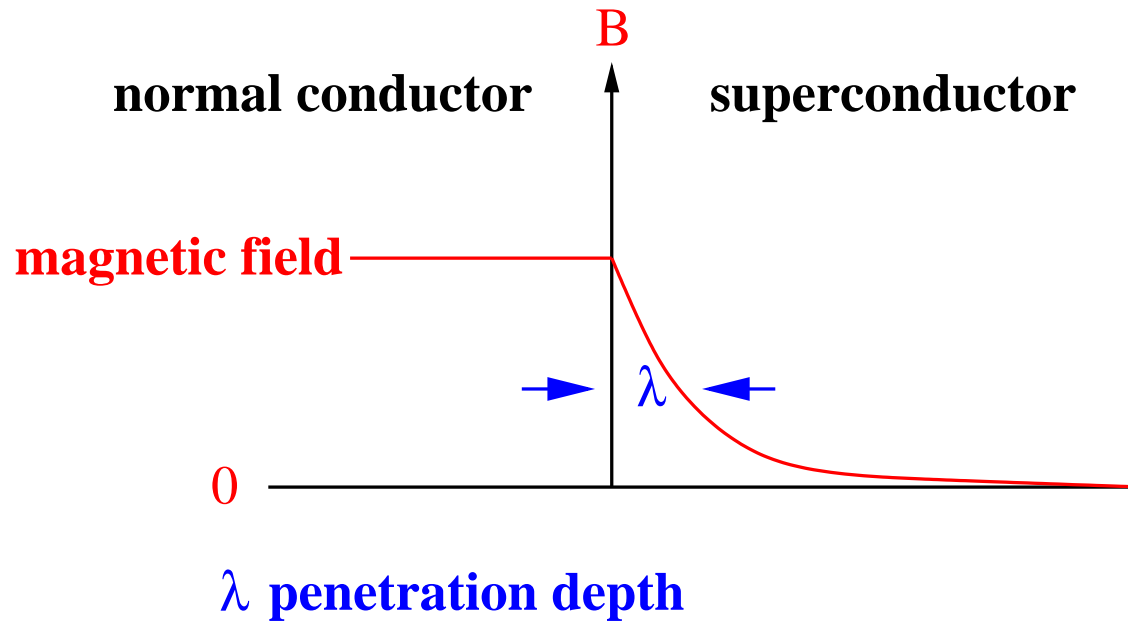
$$\epsilon_k = \pm\sqrt{(k - \mu)^2 + \Delta^2}$$

→ suppression of specific heat, viscosity, neutrino emissivity, etc.

**Critical temperature (in BCS theory):**  $T_c \simeq 0.57\Delta$



- **Basics of (color) superconductivity (page 5/5):  
Meissner effect**

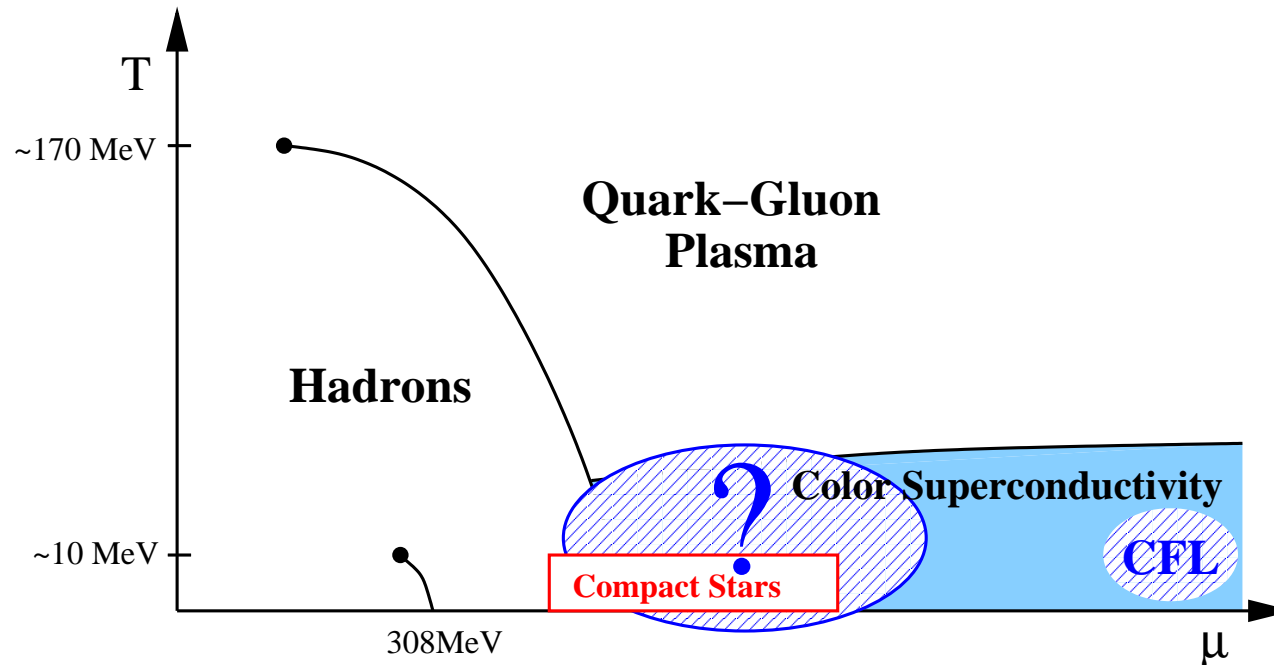


- spontaneous symmetry breaking:  $U(1)_{\text{em}} \rightarrow \mathbb{Z}_2$
- Anderson-Higgs mechanism: photon Meissner mass  $m_M$

$$m_M = \frac{1}{\lambda}, \quad B \propto e^{-m_M r}$$

- gluon Meissner masses in color superconductors

- **QCD phase diagram (page 1/3):  
Known and unknown territories**



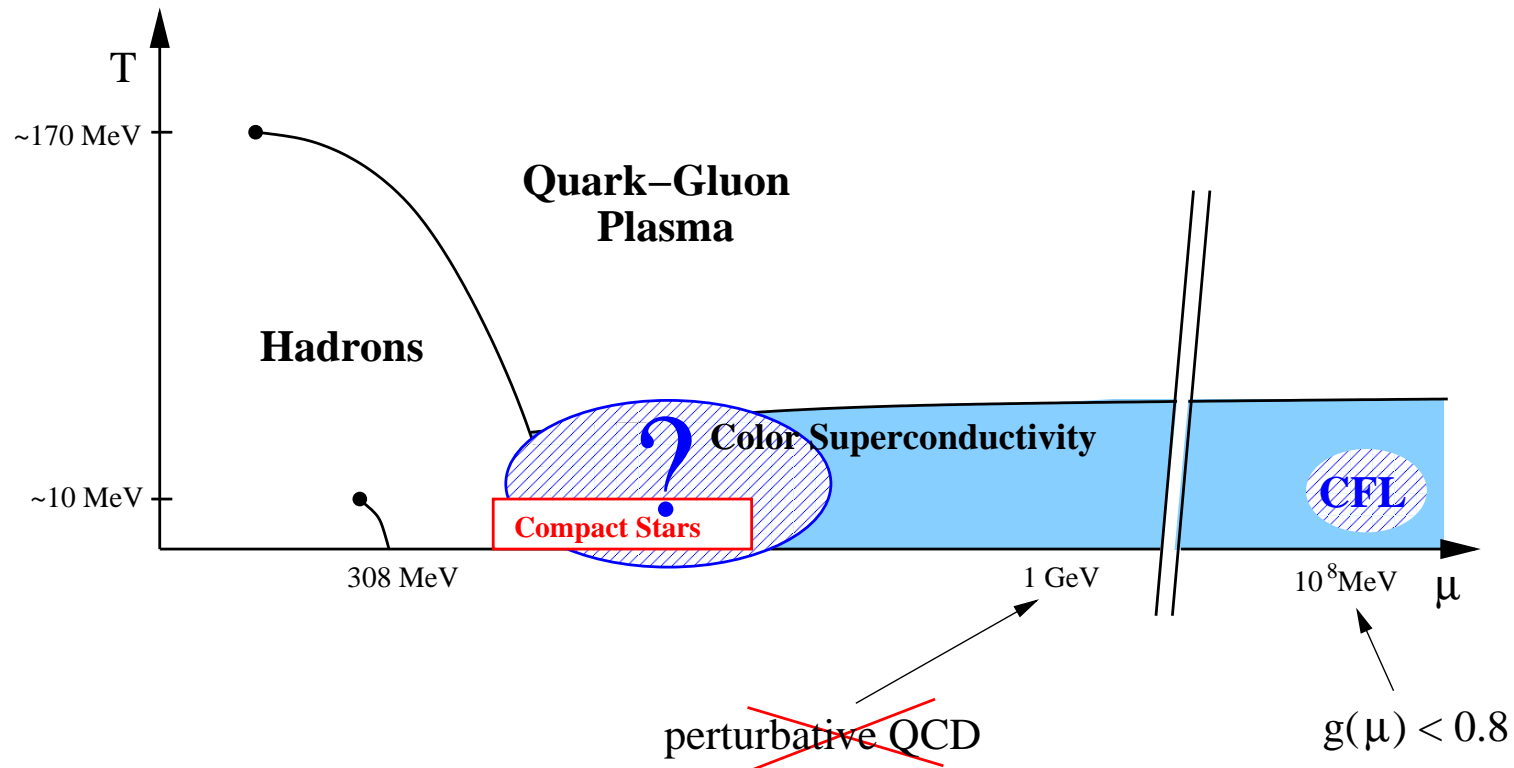
### High densities:

- rigorous theoretical control
- no nonperturbative gaps in our understanding

### Moderate densities:

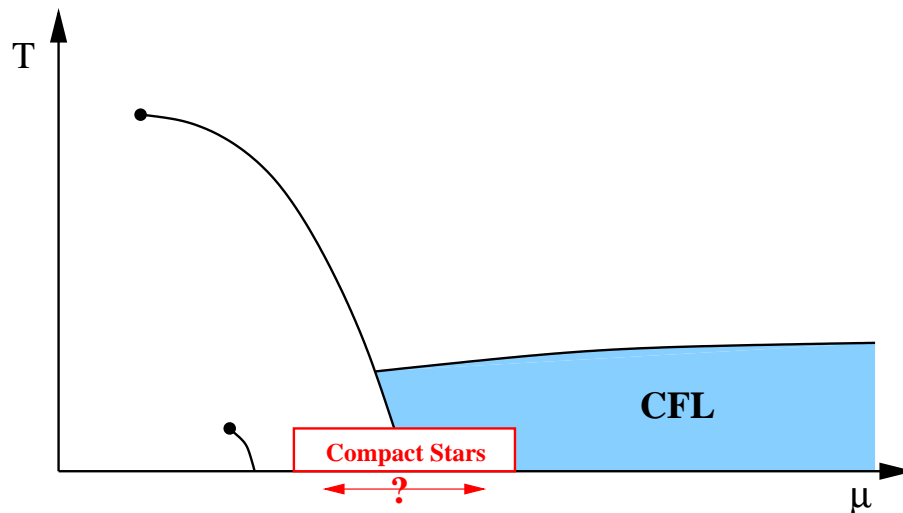
- perturbative QCD not valid
- strange mass & neutrality: stress on Cooper pairing

- QCD phase diagram (page 2/3):  
Validity of perturbative QCD



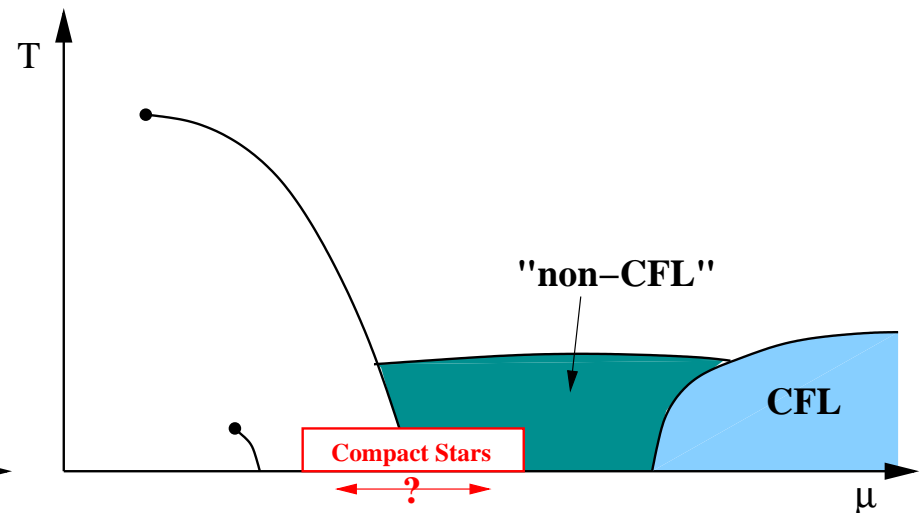
- here: always  $N_f = 3$  (ignore heavy quarks)
- $N_f > 3$ : T. Schäfer, NPB 575, 269 (2000)

- **QCD phase diagram (page 3/3): 2 Possible scenarios**



### CFL superseded by nuclear matter:

- effective theory for CFL in strongly-coupled regime
- CFL matter in the core of compact stars?



### CFL superseded by "non-CFL" matter:

- complicated phase structure?
- rely on Nambu-Jona-Lasinio-type models

## Question:

What is the ground state of deconfined quark matter at moderate densities (in the interior of compact stars)?

1. **Theoretical approach:** start from CFL and ask “what is next phase down in density?”  
(if not hadronic matter)
2. **Phenomenological approach:** “guess” possible phase, compute its properties and compare with astrophysical observations
3. (**Tabletop approach:** learn from parallels to cold fermionic atoms in magnetic trap)

- **Part II**

- Basics of (color) superconductivity
- **Color-flavor locking (CFL) – highest densities**
- Stressed pairing – below CFL densities
- QCD calculations
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- On safe grounds: Asymptotically large density

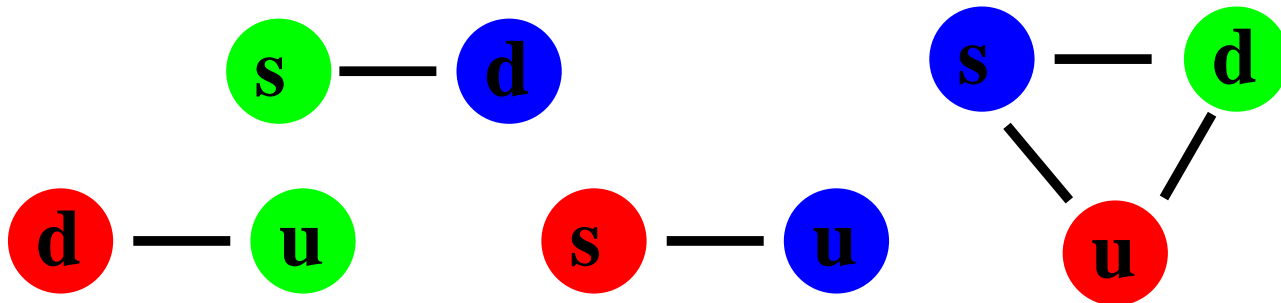
$$0 \simeq m_s \simeq m_u \simeq m_d \ll \mu \quad \text{all quark masses negligible}$$

“color-flavor locked phase (CFL)”

M. Alford, K. Rajagopal, F. Wilczek, NPB 537, 443 (1999)

$$\phi_B^A = \delta_B^A \Rightarrow \langle \psi_i^\alpha C \gamma_5 \psi_j^\beta \rangle \propto \epsilon^{\alpha\beta A} \epsilon_{ijA}$$

$$\Rightarrow SU(3)_c \times \underbrace{SU(3)_L \times SU(3)_R}_{\supset U(1)_Q} \times U(1)_B \rightarrow \underbrace{SU(3)_{c+L+R}}_{\supset U(1)_{\tilde{Q}}} \times \mathbb{Z}_2$$



- **Properties of CFL (page 1/3)**

**(1) chiral symmetry breaking**

- usual chiral symmetry breaking: LR pairing  $\langle \bar{\psi}_R \psi_L \rangle$
- here: LL, RR pairing  $\langle \psi_R \psi_R \rangle$ , however

$$SU(3)_c \times \underbrace{SU(3)_L \times SU(3)_R}_{\supset U(1)_Q} \times U(1)_B \rightarrow \underbrace{SU(3)_{c+L+R}}_{\supset U(1)_{\tilde{Q}}}$$

- chiral symmetry broken through “locking” to color
- octet of pseudo-Goldstone modes  $K^0, K^\pm, \pi^0, \dots$
- quark-hadron continuity? T. Schäfer, F. Wilczek, PRL 82, 3956 (1999)



- **Properties of CFL (page 2/3)**

- **(2) superfluidity**

$$SU(3)_c \times \underbrace{SU(3)_L \times SU(3)_R}_{\supset U(1)_Q} \times U(1)_B \rightarrow \underbrace{SU(3)_{c+L+R}}_{\supset U(1)_{\tilde{Q}}}$$

- *exactly massless* Goldstone mode  $\phi$
- vortices in rotating CFL [M. M. Forbes, A. R. Zhitnitsky, PRD 65, 085009 \(2002\)](#)

- **(3) rotated electromagnetism**

$$SU(3)_c \times \underbrace{SU(3)_L \times SU(3)_R}_{\supset U(1)_Q} \times U(1)_B \rightarrow \underbrace{SU(3)_{c+L+R}}_{\supset U(1)_{\tilde{Q}}}$$

- Cooper pairs neutral under  $\tilde{Q} = Q + \frac{2}{\sqrt{3}}T_8$
- photon-gluon mixing with (small) mixing angle

$$\cos^2 \theta = \frac{3g^2}{3g^2 + 4e^2} \simeq 1$$

- **Properties of CFL (page 3/3)**

**(3) rotated electromagnetism (continued)**

- Analogy to standard model

<b>Weinberg-Salam</b>	<b>CFL</b>
$SU(2)_I \times U(1)_Y$ isospin, hypercharge	$SU(3)_c \times U(1)_Q$ color, electromagnetism
$W_1, W_2, W_3, W_0$	$A_1, \dots, A_8, A$
$SU(2)_I \times U(1)_Y \rightarrow U(1)_Q$	$SU(3)_c \times U(1)_Q \rightarrow U(1)_{\tilde{Q}}$
$W^+, W^-$ $Z = \cos \theta_W W_3 + \sin \theta_W W_0$ $A = -\sin \theta_W W_3 + \cos \theta_W W_0$	$A_1, \dots, A_7$ $\tilde{A}_8 = \cos \theta A_8 + \sin \theta A$ $\tilde{A} = -\sin \theta A_8 + \cos \theta A$
$W^+, W^-, Z$ (massive)	$A_1, \dots, A_7, \tilde{A}_8$ (massive)
$A$ (massless)	$\tilde{A}$ (massless)

- **Why CFL is favored (page 1/2)**

- general order parameter

$$\langle \psi_i^\alpha C \gamma_5 \psi_j^\beta \rangle \propto \epsilon^{\alpha\beta A} \epsilon_{ijB} \phi_B^A$$

- complex  $3 \times 3$  matrix  $\phi_B^A \rightarrow$  *huge configuration space?!*

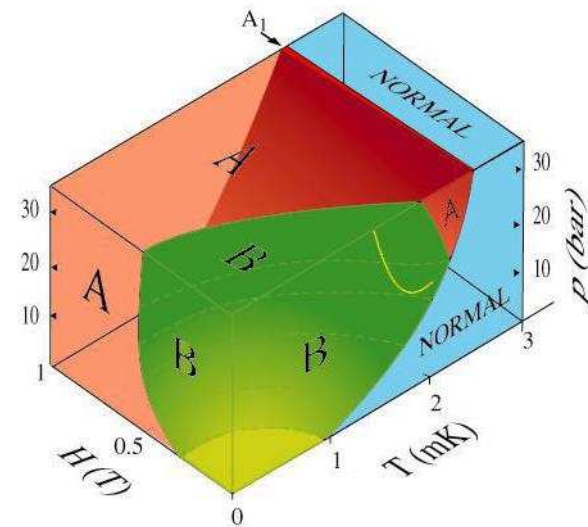
– cf. **superfluid  $^3\text{He}$**

$$SO(3)_L \times SU(2)_S \times U(1)$$

–  $3 \times 3$  order parameter in angular momentum  $L$ , spin  $S$

– **A phase:**  $\phi_B^A = \delta^{A3} (\delta_{B1} + i\delta_{B2})$

– **B phase:**  $\phi_B^A = \delta_B^A$



- **Why CFL is favored (page 2/2)**

- *simple at high densities!*  $\rightarrow$  full  $SU(3)_c \times SU(3)_f$  symmetry

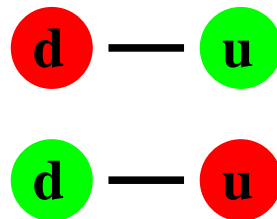
- can restrict to diagonal  $\phi_B^A$

$$\forall \phi \exists U \in SU(3)_c, V \in SU(3)_f : U^T \phi V \text{ diagonal}$$

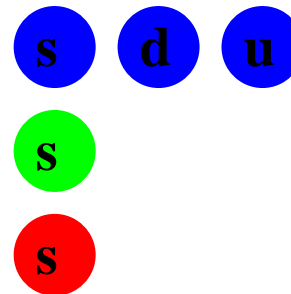
- then  $\phi_B^A = \delta_B^A$  is only phase where **all quarks pair**

- for instance **2SC phase**  $\phi_B^A = \delta^{A3} \delta_{B3}$

paired:

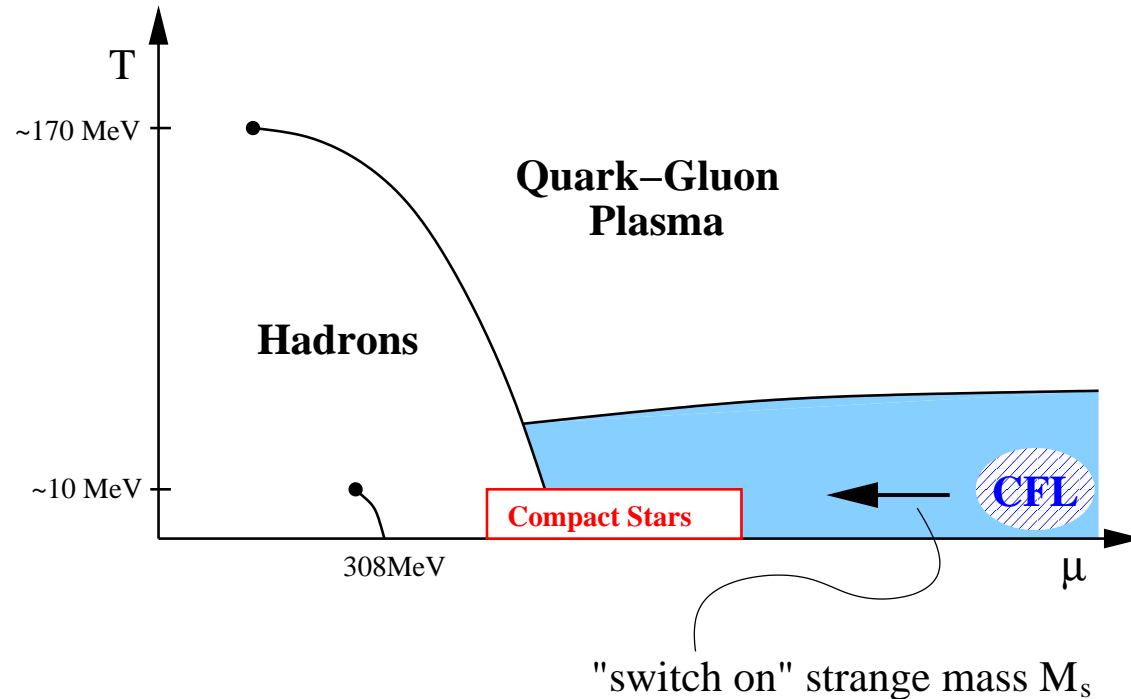


unpaired:



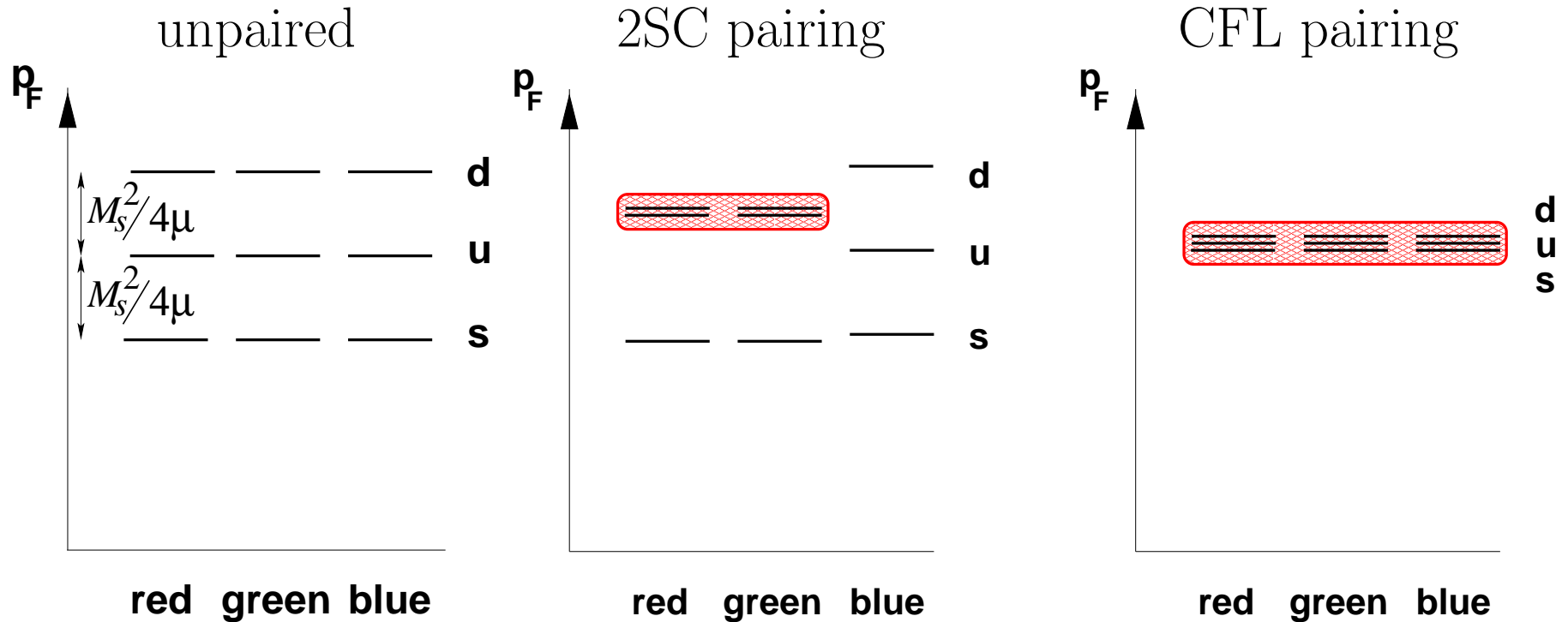
- **Part III**
- Basics of (color) superconductivity
- Color-flavor locking (CFL) – highest densities
- **Stressed pairing – below CFL densities**
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- **Going down in density:**  
**Large, but not asymptotically large, densities**
- strange mass  $M_s \simeq 120 \text{ MeV}$  no longer  $\ll \mu \simeq 400 \text{ MeV}$



- at given chemical potentials  $\mu, \mu_e$ :  $M_s \neq 0$  reduces  $p_F^s$
- electric neutrality: increase in  $p_F^d$  to compensate
- **Fermi momenta “try” to split apart**

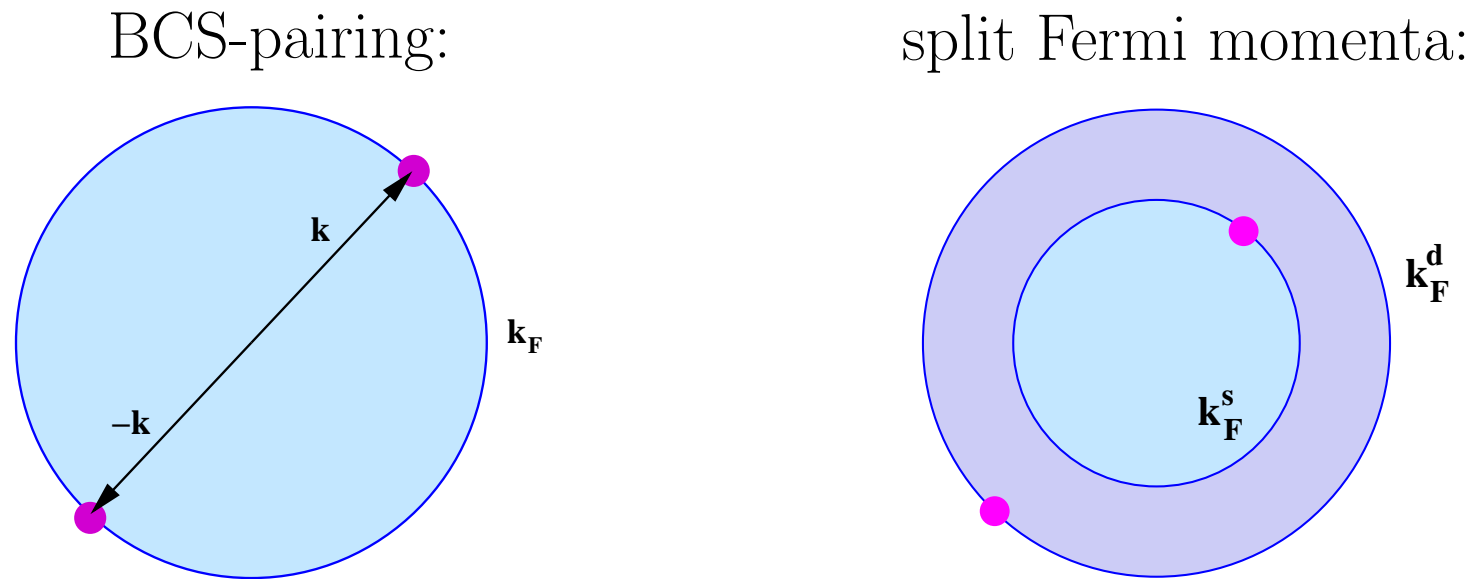
- $p_F$ 's splitting apart



- any pairing pattern most “comfortable” with  $M_s$  and neutrality?
- stressed pairing is **unavoidable!**

K. Rajagopal, A. Schmitt, PRD 73, 045003 (2006)

- **Stressed Cooper pairing:  
a general phenomenon (page 1/2)**



- for instance: two species (spin states) of cold fermionic atoms

- Experiments:

Y. Chin, M.W. Zwierlein, C.H. Schunck, A. Schirotzek, W. Ketterle, PRL 97, 030401 (2006)

G.B. Partridge, W. Li, R.I. Kamar, Y. Liao, R.G. Hulet, Science 311, 503 (2006)

- Theory (review):

D.E. Sheehy, L. Radzihovsky, Ann. Phys. 322, 1790 (2007)

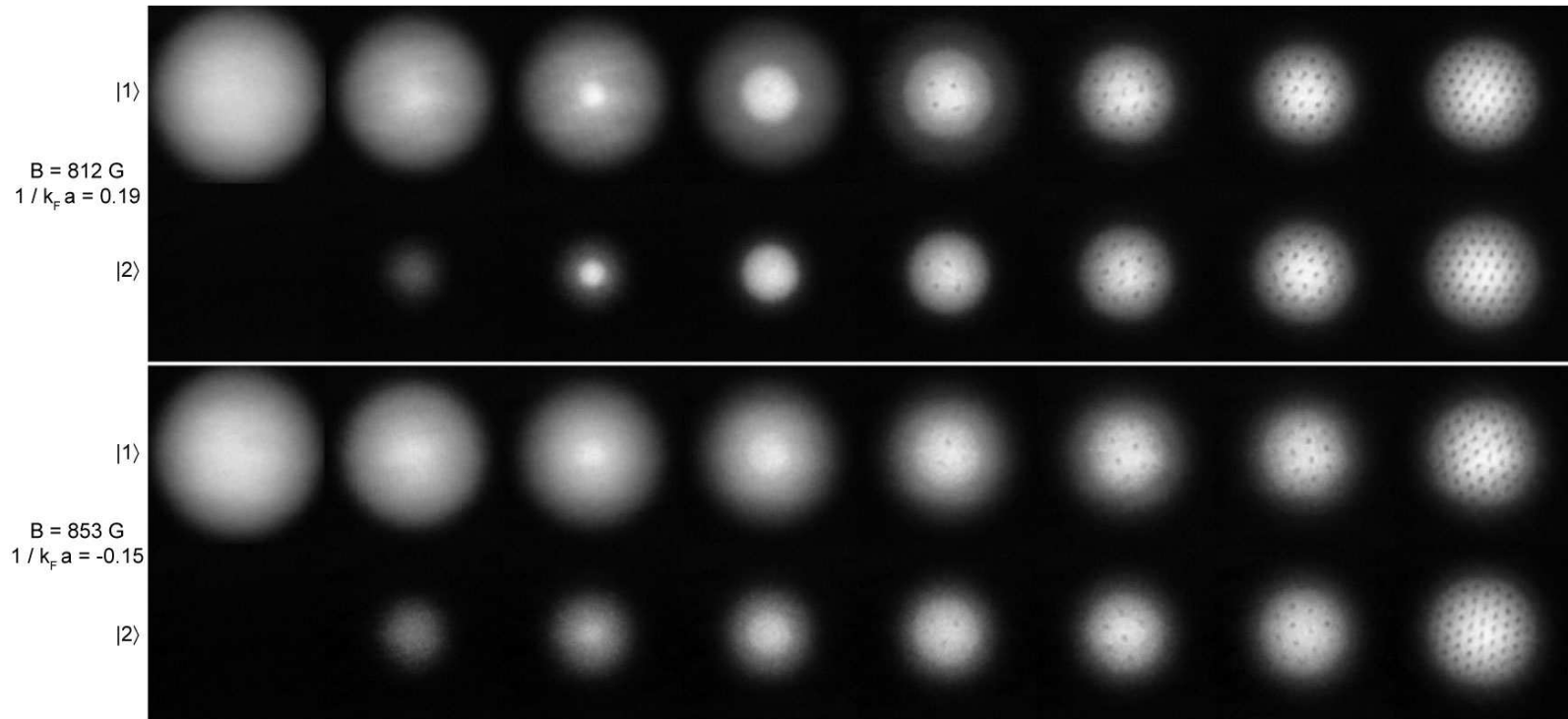


- **Stressed Cooper pairing:  
a general phenomenon (page 1/2)**

## Pairing with mismatch



- **Stressed Cooper pairing:  
a general phenomenon (page 2/2)**

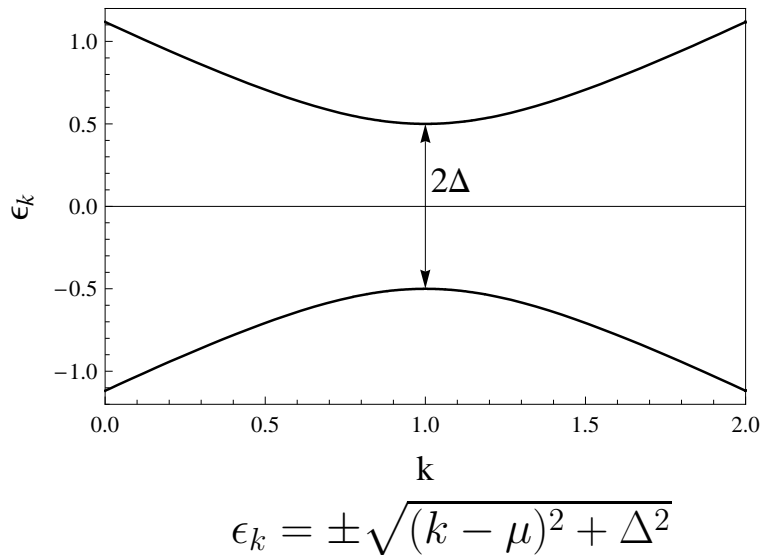
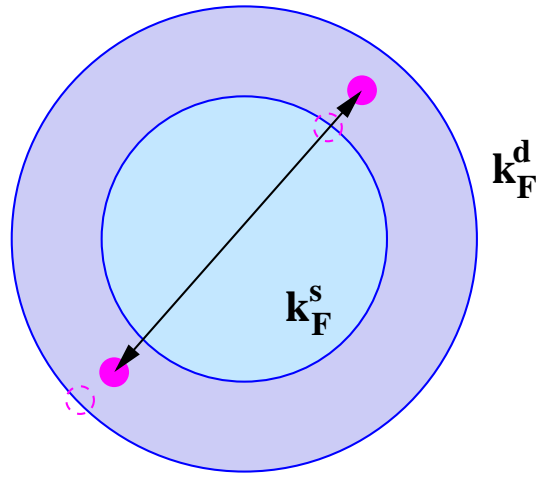


M.W. Zwierlein, A. Schirotzek, C.H. Schunck, W. Ketterle, *Science* 311, 492 (2006)

- **phase separation** of superfluid and normal components
- phase separation unlikely in quark matter (local color neutrality!)

M. Alford, C. Kouvaris, K. Rajagopal, *PRL* 92, 222001 (2004)

## • CFL pairing with small stress



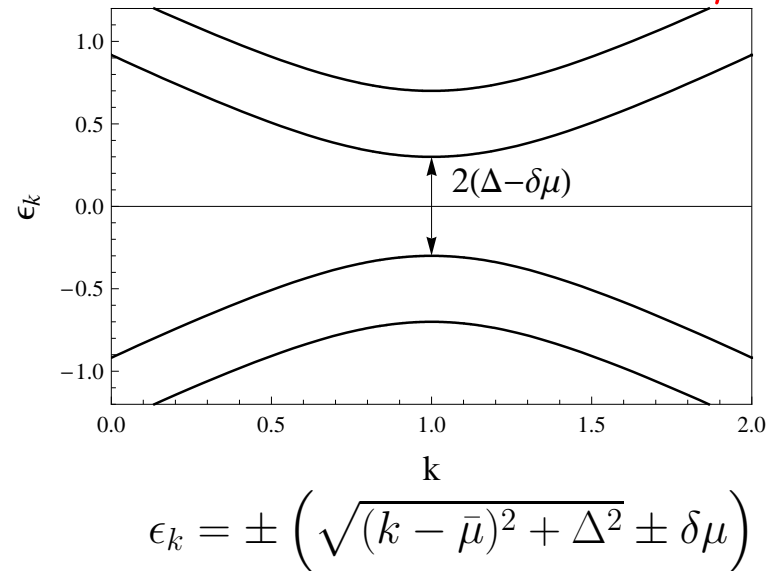
- create common Fermi surface:  
cost in free energy

$$\sim \delta p_F^2 \mu^2 \sim M_s^4$$

- form pairs:

$$\text{gain in free energy} \sim \Delta^2 \mu^2$$

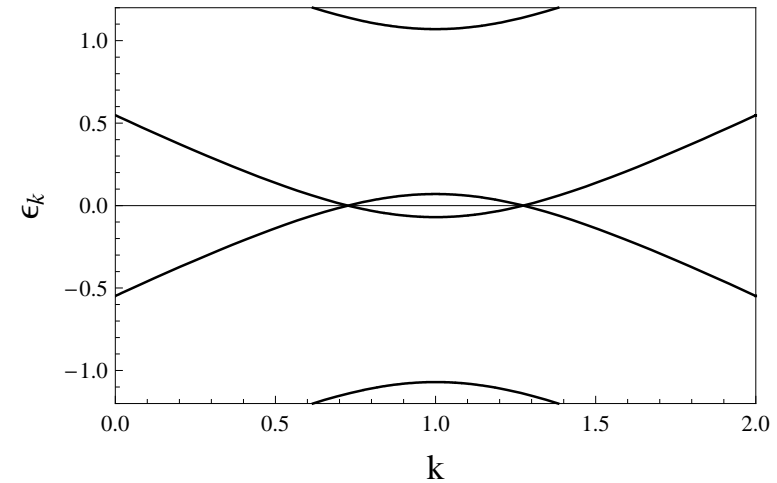
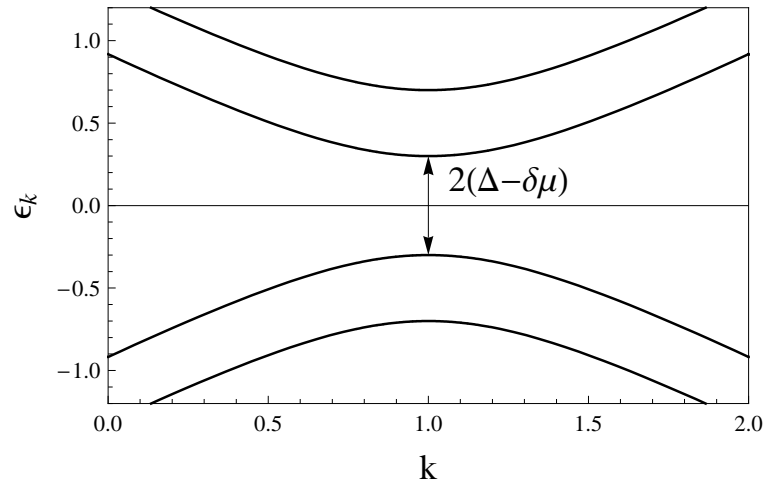
- CFL survives for  $\Delta \gtrsim \frac{M_s^2}{\mu}$



- **Large stress: gapless CFL?**

→ “gapless CFL” for  $\delta\mu > \Delta$

M. Alford, C. Kouvaris, K. Rajagopal, PRL 92, 222001 (2004)



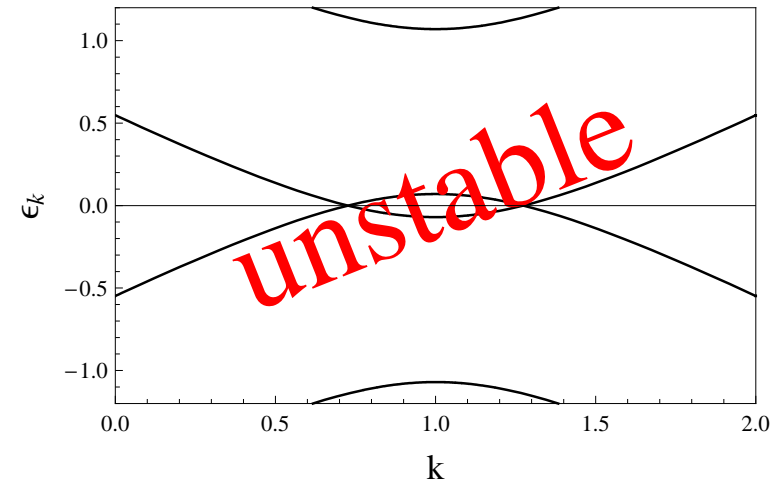
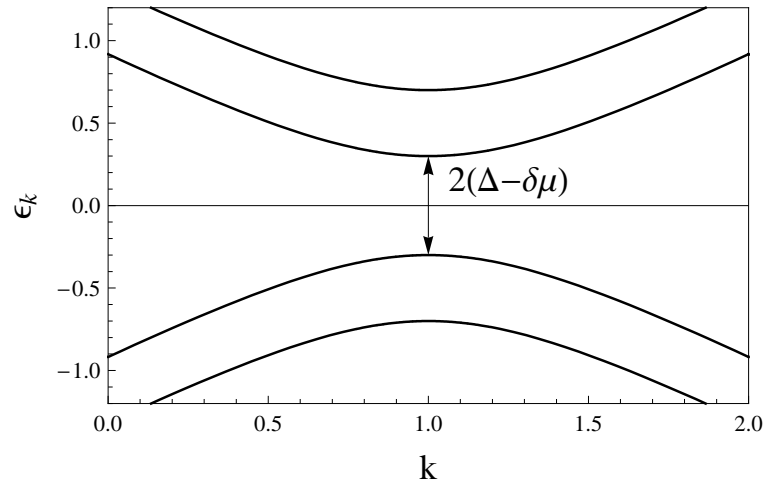
→ or: “breached” pairing

E. Gubankova, W. V. Liu and F. Wilczek, PRL 91, 032001 (2003)

- **Large stress: gapless CFL? No!**

→ “gapless CFL” for  $\delta\mu > \Delta$

M. Alford, C. Kouvaris, K. Rajagopal, PRL 92, 222001 (2004)



→ or: “breached” pairing

E. Gubankova, W. V. Liu and F. Wilczek, PRL 91, 032001 (2003)

→ **chromomagnetic instability** (at  $T = 0$ )

M. Huang, I. A. Shovkovy, PRD 70, 051501 (2004)

R. Casalbuoni, R. Gatto, M. Mannarelli, G. Nardulli, M. Ruggieri, PLB 605, 362 (2005)

→ ground state must be different → **currents**

- **Less (and less symmetric) pairing (page 1/4)**

## Kaon condensation “CFL- $K^0$ ”

P. F. Bedaque and T. Schäfer, NPA 697, 802 (2002)

- chiral field

$$\Sigma = \phi_L^\dagger \phi_R$$

- pure CFL:  $\Sigma = \mathbf{1}$

- kaon condensation  $\Rightarrow \Sigma = e^{i\varphi T_6}$   
(relative L/R rotation)

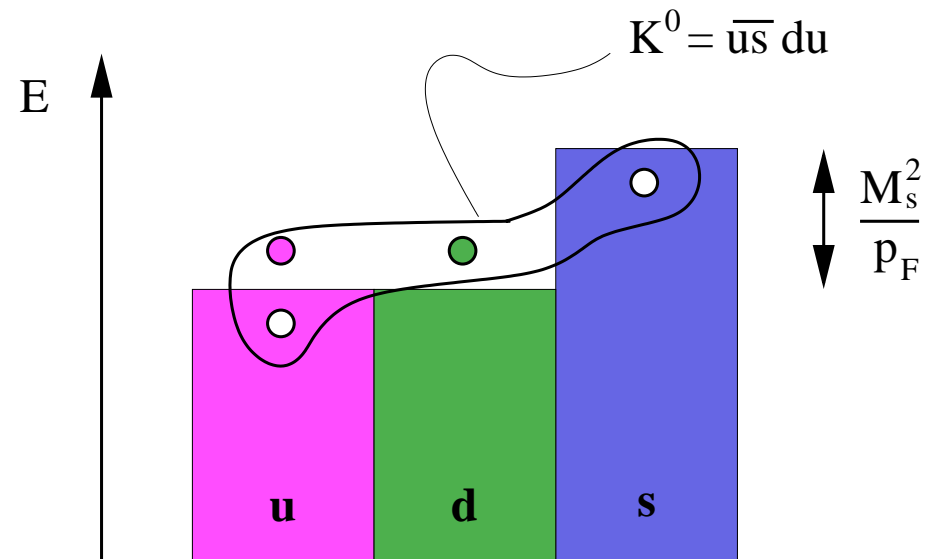
- **in other words:**

create kaon with mass

$$m_{K^0}^2 = a m_d (m_s + m_u) \ll \Delta^2$$

from  $0 \rightarrow \bar{s} + u + \bar{u} + d$

$$(a \sim \Delta^2 / \mu^2)$$



- Less (and less symmetric) pairing (page 2/4)

## (Super)currents in CFL: “curCFL- $K^0$ ”

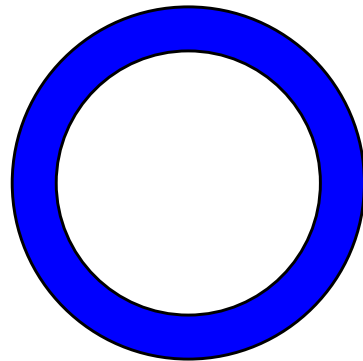
T. Schäfer, PRL 96, 012305 (2006)

A. Kryjevski, PRD 77, 014018 (2008)

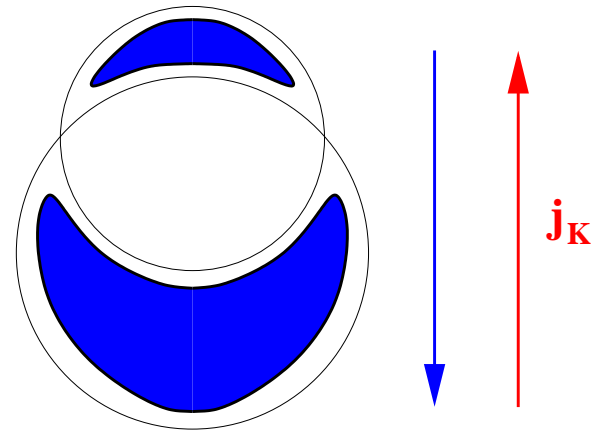
$$\phi_L(\mathbf{x}) = \Delta e^{i\mathbf{j}_K \cdot \mathbf{x}} T_8 e^{i(\varphi/2)T_6}$$

$$\phi_R(\mathbf{x}) = \Delta e^{i\mathbf{j}_K \cdot \mathbf{x}} T_8 e^{-i(\varphi/2)T_6}$$

- “anisotropic breach”
- stable and unstable Fermi surface topologies:



“breach” (unstable)



curCFL- $K^0$  (stable)

- Less (and less symmetric) pairing (page 3/4)

## More currents in CFL: crystalline structures (LOFF)

M. Alford, J. Bowers, K. Rajagopal, PRD 63, 074016 (2001)

M. Mannarelli, K. Rajagopal and R. Sharma, PRD 73, 114012 (2006)

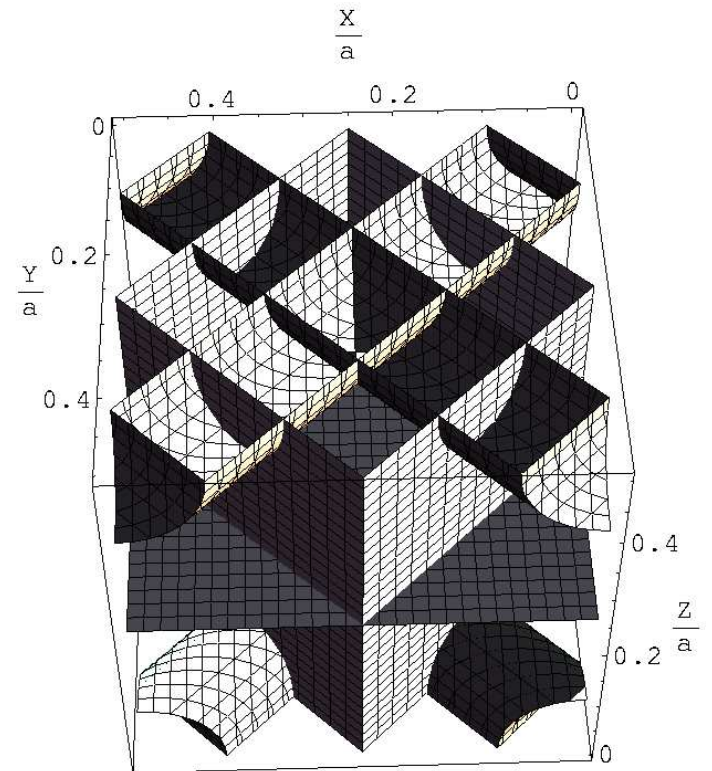
$$\langle ud \rangle \sim \Delta_3 \sum_a \exp(2i\mathbf{q}_3^a \cdot \mathbf{x})$$

$$\langle us \rangle \sim \Delta_2 \sum_a \exp(2i\mathbf{q}_2^a \cdot \mathbf{x})$$

$$\langle ds \rangle \simeq 0$$

- here: “CubeX”

- $\{\mathbf{q}_3\}$ ,  $\{\mathbf{q}_2\}$  each contain 4 vectors, together pointing to the corners of a cube



$\Delta_3(\mathbf{x}), \Delta_2(\mathbf{x})$



- **Less (and less symmetric) pairing (page 4/4)**

## Last resort: single flavor pairing

- need  $J = 1$  Cooper pairs

$$\phi \in [\bar{\mathbf{3}}]_c^a \otimes [\mathbf{3}]_J^s$$

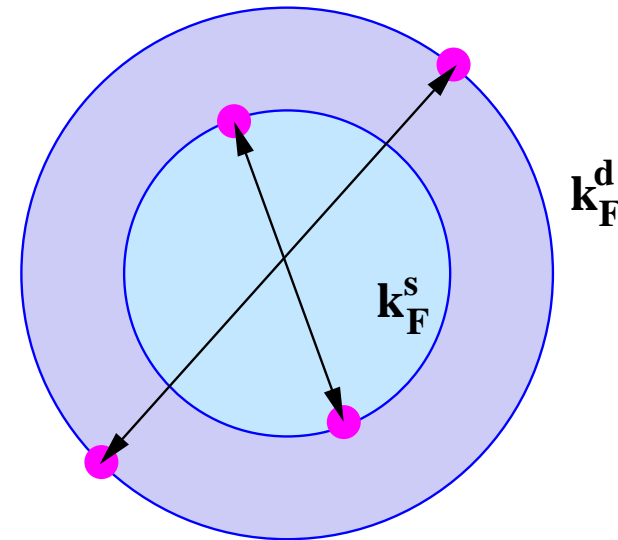
- different possible phases:  
Color-spin locking (CSL),  
A-phase, polar phase ...

T. Schäfer, PRD 62, 094007 (2000)

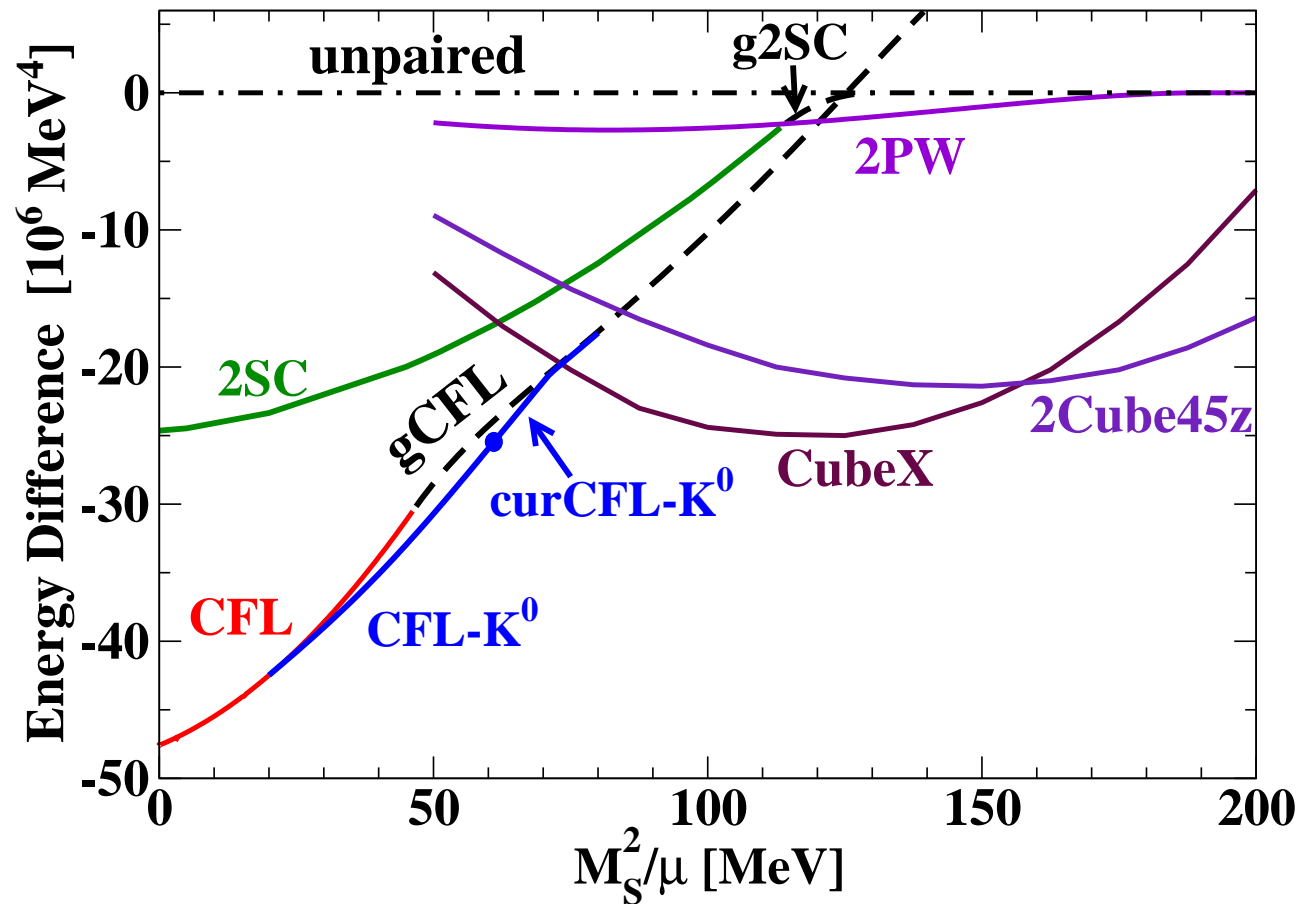
A. Schmitt, PRD 71, 054016 (2005)

- preferred phase at high densities: **CSL**
- gap much smaller than in spin-0 phases:

$$\Delta_{J=1} \lesssim 10^{-2} \Delta_{J=0}$$



- Stressed pairing: free energy comparison



here:  $\Delta_{\text{CFL}} = 25 \text{ MeV}$

(pert. QCD:  $\Delta_{\text{CFL}} \simeq 20 \text{ MeV}$ , NJL:  $\Delta_{\text{CFL}} \simeq (20 - 100) \text{ MeV}$ ).

- **Summary Parts I – III**
  
- **3-flavor quark matter** at asymptotically high densities is in the **Color-Flavor locked (CFL)** state
- **CFL** quark matter
  - breaks chiral symmetry
  - is a superfluid
  - is a transparent insulator
- phase(s) between CFL and hadronic matter (if there is/are any) is/are unknown
  - **large coupling**
  - **stressed pairing** renders ground state complicated

- **Part IV**

- Basics of (color) superconductivity
- Color-flavor locking (CFL) – highest densities
- Stressed pairing – below CFL densities
- **QCD calculations**
- Effective theory of CFL – kaon condensation
- Transport properties – quark matter in compact stars

- QCD Lagrangian and Nambu-Gorkov basis

$$\mathcal{L} = \bar{\psi}(i\gamma^\mu D_\mu + \hat{\mu}\gamma_0 - \hat{m})\psi - \frac{1}{4}G_a^{\mu\nu}G_{\mu\nu}^a$$

- high densities:  $m_u = m_d = m_s = 0$
- from weak interactions: only  $\mu, \mu_e$  (not  $\mu_u, \mu_d, \mu_s$ )
- introduce charge-conjugate spinors

$$\Psi = \begin{pmatrix} \psi \\ \psi_C \end{pmatrix}, \quad \bar{\Psi} = (\bar{\psi}, \bar{\psi}_C)$$

- gap matrix  $\Phi^\pm \propto \Delta \epsilon^{\alpha\beta A} \epsilon_{ijB} \phi_B^A$  in NG propagator

$$S^{-1} = \begin{pmatrix} [G_0^+]^{-1} + \Sigma^+ & \Phi^- \\ \Phi^+ & [G_0^-]^{-1} + \Sigma^- \end{pmatrix}$$

- **Gap equation**

- $72 \times 72$  ( $2 \cdot 4N_f N_c$ ) quark propagator

- $F^\pm$  “anomalous” propagators

$$S = \begin{pmatrix} G^+ & F^- \\ F^+ & G^- \end{pmatrix}$$

- one-loop self-energy

$$\Sigma = \begin{pmatrix} \text{diagram 1} & \text{diagram 2} \\ \text{diagram 3} & \text{diagram 4} \end{pmatrix}$$

- **gap equation** (Dyson-Schwinger equation)

$$\Phi^+(K) = g^2 \int_Q \gamma^\mu T_a^T F^+(Q) \gamma^\nu T_b D_{\mu\nu}^{ab}(K - Q)$$

- **Solution of the gap equation (page 1/3)**

- gap parameter  $\Delta$  at the Fermi surface at  $T = 0$

$$\Delta = 512\pi^4 \left( \frac{2}{N_f g^2} \right)^{5/2} \mu \exp\left(-\frac{\pi^2 + 4}{8}\right) e^{-d} e^{-\zeta} \exp\left(-\frac{3\pi^2}{\sqrt{2}g}\right) [1 + \mathcal{O}(g)]$$

D.T. Son, PRD 59, 094019 (1999)

T. Schäfer, F. Wilczek, PRD 60, 114033 (1999); R. D. Pisarski, D. H. Rischke, PRD 61, 051501 (2000); D.K. Hong, V.A. Miransky, I.A. Shovkovy, L.C.R. Wijewardhana, PRD 61, 056001 (2000)

W. E. Brown, J. T. Liu and H. c. Ren, PRD 61, 114012 (2000); Q. Wang, D. H. Rischke, PRD 65, 054005 (2002)

T. Schäfer, PRD 62, 094007 (2000); A. Schmitt, Q. Wang, D. H. Rischke, PRD 66, 114010 (2002); A. Schmitt, PRD 71, 054016 (2005)

- let's discuss the contributions separately ...

- **Solution of the gap equation (page 2/3)**

$$\Delta = 512\pi^4 \left( \frac{2}{N_f g^2} \right)^{5/2} \mu \exp\left(-\frac{\pi^2 + 4}{8}\right) e^{-d} e^{-\zeta} \exp\left(-\frac{3\pi^2}{\sqrt{2}g}\right) [1 + \mathcal{O}(g)]$$

- leading  $1/g$  term from **long-range magnetic gluons**

$$\ln\left(\frac{\Delta}{\mu}\right) = -\frac{b_{-1}}{g} - \bar{b}_0 \ln g - b_0 - \dots$$

- unlike **BCS (point-like interaction)**,  $\Delta \propto \exp(-1/g^2)$

$\Rightarrow$  color-superconducting gap *increases* at  $\mu \rightarrow \infty$ :

$$\frac{1}{g^2} \propto \ln \frac{\mu}{\Lambda} \quad \Rightarrow \quad \Delta \propto \mu e^{-const./g(\mu)} \text{ increases with } \mu$$

(of course  $\Delta/\mu$  decreases)



- Solution of the gap equation (page 2/3)

$$\Delta = 512\pi^4 \left( \frac{2}{N_f g^2} \right)^{5/2} \mu \exp\left(-\frac{\pi^2 + 4}{8}\right) e^{-d} e^{-\zeta} \exp\left(-\frac{3\pi^2}{\sqrt{2}g}\right) [1 + \mathcal{O}(g)]$$

- $\exp[-(\pi^2 + 4)/8] \simeq 0.2$  from “normal” quark self energy  $\Sigma^\pm$

$$S^{-1} = \begin{pmatrix} [G_0^+]^{-1} + \Sigma^+ & \Phi^- \\ \Phi^+ & [G_0^-]^{-1} + \Sigma^- \end{pmatrix} \quad \Sigma^\pm \simeq \gamma_0 \Lambda_{\mathbf{k}}^{(\pm)} \frac{g^2}{18\pi^2} k_0 \log \frac{48e^2 m_g^2}{\pi^2 k_0^2}$$

- smallness of spin-1 gaps

$$e^{-d} = \begin{cases} 1 & J = 0 \\ e^{-6} \simeq 2.5 \cdot 10^{-3} & J = 1; LL, RR \text{ pairing} \\ e^{-4.5} \simeq 0.01 & J = 1; LR, RL \text{ pairing} \end{cases}$$

- **Solution of the gap equation (page 3/3)**

$$\Delta = 512\pi^4 \left( \frac{2}{N_f g^2} \right)^{5/2} \mu \exp\left(-\frac{\pi^2 + 4}{8}\right) e^{-d} e^{-\zeta} \exp\left(-\frac{3\pi^2}{\sqrt{2}g}\right) [1 + \mathcal{O}(g)]$$

- $e^{-\zeta}$  depends on the **pairing pattern**  
(not on details of interaction)

- Define

$$\mathcal{M}_{ij}^{\alpha\beta} \equiv \epsilon^{\alpha\beta A} \epsilon_{ijB} \phi_B^A$$

- Denote

$\lambda_r$  ( $n_r$ -fold) eigenvalues of  $\mathcal{M}\mathcal{M}^\dagger$

- quasiparticle dispersions

$$\epsilon_{k,r} = \sqrt{(k - \mu)^2 + \lambda_r \Delta^2}$$

- condensation energy

$$\delta P = \frac{\mu^2}{4\pi^2} \sum_r n_r \lambda_r \Delta^2$$

- **Solution of the gap equation (page 4/4)**

- then

$$\zeta = \frac{1}{2} \frac{\langle n_1 \lambda_1 \log \lambda_1 + n_2 \lambda_2 \log \lambda_2 \rangle}{\langle n_1 \lambda_1 + n_2 \lambda_2 \rangle}$$

- $e^{-\zeta} \neq 1$  for **(i) two-gap structure...**

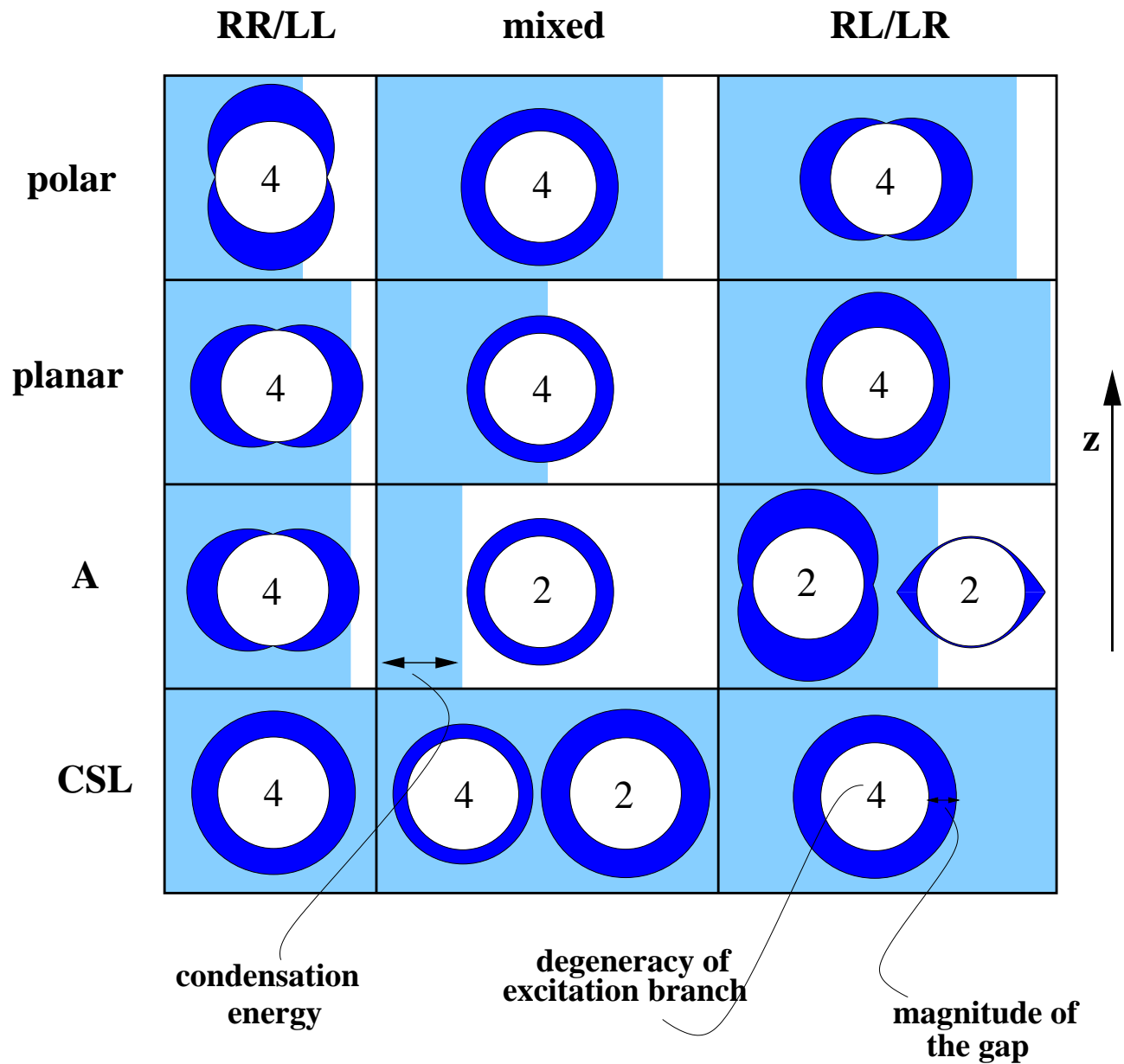
	CFL	2SC
$\lambda_1$ ( $n_1$ )	4 (1)	1 (4)
$\lambda_2$ ( $n_2$ )	1 (8)	0 (5)
$e^{-\zeta}$	$2^{-1/3}$	1

- consequently,

$$\Delta_{\text{CFL}} = 2^{-1/3} \Delta_{\text{2SC}}$$

- ...and/or **(ii) anisotropic gaps**

- **Anisotropic gaps** A. Schmitt, PRD 71, 054016 (2005)



- **Critical temperature**

- set  $\Delta = 0$  in gap equation to extract  $T_c$

$$T_c = e^\zeta \frac{e^\gamma}{\pi} \Delta [1 + \mathcal{O}(g)] \simeq 0.57 e^\zeta \Delta$$

- $e^\zeta$  violates BCS relation  $T_c \simeq 0.57 \Delta$

A. Schmitt, Q. Wang, D. H. Rischke, PRD 66, 114010 (2002)

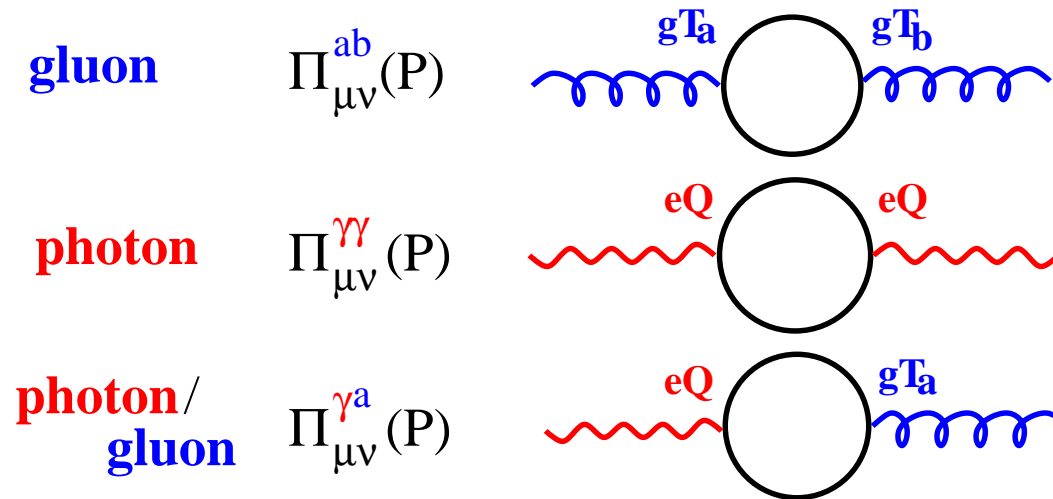
- include **gauge field fluctuations**:  
second order  $\rightarrow$  first order, and

$$\frac{T_c^* - T_c}{T_c} = \frac{\pi^2}{12\sqrt{2}} g$$

I. Giannakis, D. f. Hou, H. c. Ren, D. H. Rischke, PRL 93, 232301 (2004)

## • Meissner masses from polarization tensors

Meissner masses: 
$$m_{ab}^2 \equiv \frac{1}{2} \lim_{p \rightarrow 0} (\delta^{ij} - \hat{p}^i \hat{p}^j) \Pi_{ij}^{ab}(p_0 = 0, \mathbf{p})$$



Elimination of unphysical,  
“mixed” masses ...

$$m_{ab}^2 \rightarrow \tilde{m}_{ab}^2 \equiv \mathcal{O}_{ac} m_{cd}^2 \mathcal{O}_{db}^T$$

... goes along with rotation  
of gauge fields

$$A_a^\mu \rightarrow \tilde{A}_a^\mu \equiv \mathcal{O}_{ab} A_b^\mu$$



- **Part V**

- Basics of (color) superconductivity
- Color-flavor locking (CFL) – highest densities
- Stressed pairing – below CFL densities
- QCD calculations
- **Effective theory of CFL – kaon condensation**
- Transport properties – quark matter in compact stars



- **Low-energy degrees of freedom in CFL**

$$SU(3)_c \times \underbrace{SU(3)_L \times SU(3)_R}_{\supset U(1)_Q} \times U(1)_B \rightarrow \underbrace{SU(3)_{c+L+R}}_{\supset U(1)_{\tilde{Q}}}$$

→ pseudo-Goldstone octet  $(K^0, \bar{K}^0, K^\pm, \pi^0, \pi^\pm, \eta)$   
 + “superfluid mode”  $\varphi$

- **effective theory for mesons and superfluid mode**

P. F. Bedaque, T. Schäfer, NPA 697, 802 (2002)

D. T. Son, M. A. Stephanov, PRD 61, 074012 (2000)

$$\mathcal{L}_{\text{eff}}^\Sigma = \frac{f_\pi^2}{4} \text{Tr} [(\partial_0 \Sigma + i[A_0, \Sigma])(\partial_0 \Sigma^\dagger - i[A_0, \Sigma^\dagger]) - v_\pi^2 \partial_i \Sigma \partial_i \Sigma^\dagger] + \frac{a f_\pi^2}{2} \det \hat{M} \text{Tr}[\hat{M}^{-1}(\Sigma + \Sigma^\dagger)] + \dots$$

$$\mathcal{L}_{\text{eff}}^\phi = \frac{1}{2}(\partial_0 \phi)^2 - \frac{v_H^2}{2}(\nabla \phi)^2 - \frac{\pi}{9\mu^2} \partial_0 \phi \partial_\mu \phi \partial^\mu \phi + \frac{\pi^2}{108\mu^4} (\partial_\mu \phi \partial^\mu \phi)^2 + \dots$$

+ weak interactions between  $\Sigma$  and  $\phi$

- **Low-energy degrees of freedom in CFL**

- $\mathcal{L}_{\text{eff}}^{\Sigma} + \mathcal{L}_{\text{eff}}^{\phi}$  valid for  $T < \Delta \simeq 25 \text{ MeV}$
- form of  $\mathcal{L}_{\text{eff}}^{\Sigma} + \mathcal{L}_{\text{eff}}^{\phi}$  determined entirely by CFL symmetries

$$\mathcal{L}_{\text{eff}}^{\Sigma} = \frac{\mathbf{f}_{\pi}^2}{4} \text{Tr} [(\partial_0 \Sigma + i[A_0, \Sigma])(\partial_0 \Sigma^{\dagger} - i[A_0, \Sigma^{\dagger}]) - v_{\pi}^2 \partial_i \Sigma \partial_i \Sigma^{\dagger}] + \frac{\mathbf{a} \mathbf{f}_{\pi}^2}{2} \det \hat{M} \text{Tr}[\hat{M}^{-1}(\Sigma + \Sigma^{\dagger})] + \dots$$

$$\Sigma = e^{i\theta/f_{\pi}} = \phi_L \phi_R^{\dagger}, \quad \theta \in SU(3)$$

- parameters (reliably given only at high densities)

$$\mathbf{a} \equiv \frac{3\Delta^2}{\pi^2 \mathbf{f}_{\pi}^2}, \quad \mathbf{f}_{\pi} = \frac{21 - 8 \ln 2}{18} \frac{\mu^2}{2\pi^2}$$

$\Rightarrow$  lower densities: quantitative uncertainties, but theory applicable!  
 $\rightarrow$  e.g., compute transport properties from effective theory

- **Low-energy degrees of freedom in CFL**

$$\mathcal{L}_{\text{eff}} = \frac{f_\pi^2}{4} \text{Tr} [(\partial_0 \Sigma + i[\mathbf{A}_0, \Sigma])(\partial_0 \Sigma^\dagger - i[\mathbf{A}_0, \Sigma]^\dagger) - v_\pi^2 \partial_i \Sigma \partial_i \Sigma^\dagger] + \frac{a f_\pi^2}{2} \det \hat{M} \text{Tr}[\hat{M}^{-1}(\Sigma + \Sigma^\dagger)] + \dots$$

$$\Sigma = e^{i\theta/f_\pi} = \phi_L \phi_R^\dagger, \quad \theta \in SU(3)$$

- effective chemical potential  $\mathbf{A}_0 = \mu_Q Q - \hat{M}^2/2\mu$  and effective masses, e.g.,

$$E_{\pi^\pm} = \mu_{\pi^\pm} + [v_\pi^2 p^2 + a(m_u + m_d)m_s]^{1/2}$$

$$\mu_{\pi^\pm} = \mp \frac{m_d^2 - m_u^2}{2\mu}$$

$$E_{K^\pm} = \mu_{K^\pm} + [v_\pi^2 p^2 + a(m_u + m_s)m_d]^{1/2}$$

$$\mu_{K^\pm} = \mp \frac{m_s^2 - m_u^2}{2\mu}$$

$$E_{K^0, \bar{K}^0} = \mu_{K^0, \bar{K}^0} + [v_\pi^2 p^2 + a(m_d + m_s)m_u]^{1/2}$$

$$\mu_{K^0, \bar{K}^0} = \mp \frac{m_s^2 - m_d^2}{2\mu}$$

- **inverse mass ordering**  $m_K < m_\pi < \dots$

- **Inverse mass ordering**

- Remember:  $\phi_L, \phi_R$  are antitriplet diquarks

**usual pion**

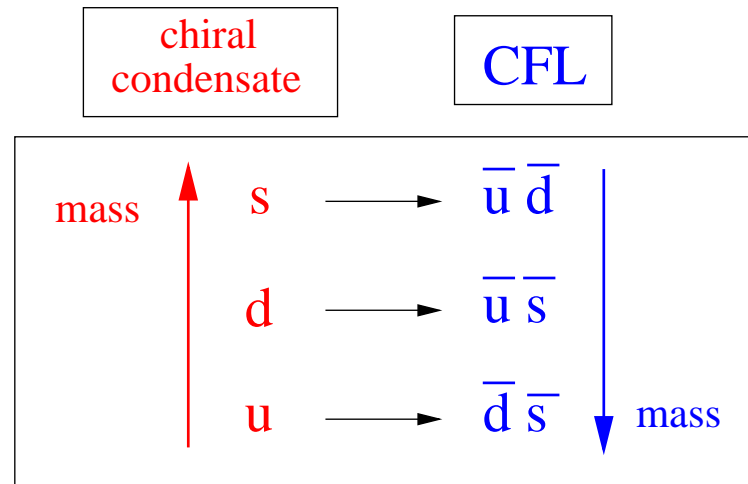
$$\pi^+ \sim \bar{d}u$$

$$m_{\pi^+} \propto m_u + m_d$$

**CFL pion:  $\bar{d} \rightarrow us, u \rightarrow \bar{d}\bar{s}$**

$$\pi^+ \sim us\bar{d}\bar{s}$$

$$m_{\pi^+}^2 \propto m_u m_s + m_d m_s$$



- **Inverse mass ordering**

- Remember:  $\phi_L, \phi_R$  are antitriplet diquarks

**usual pion**

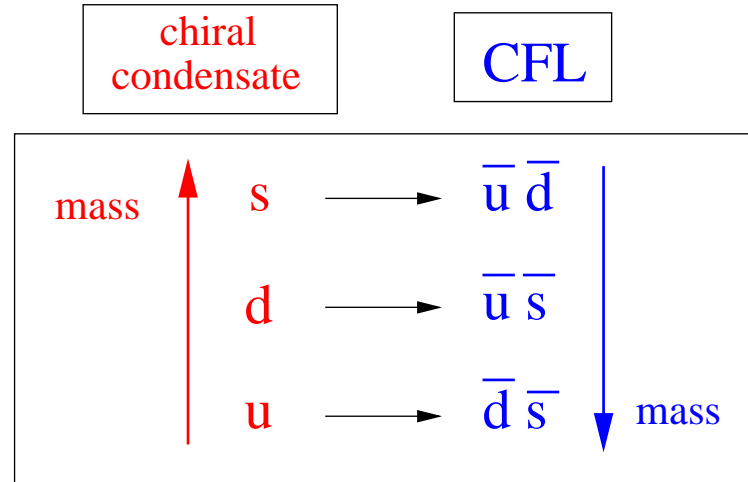
$$\pi^+ \sim \bar{d}u$$

$$m_{\pi^+} \propto m_u + m_d$$

**CFL pion:  $\bar{d} \rightarrow us, u \rightarrow \bar{d}\bar{s}$**

$$\pi^+ \sim us\bar{d}\bar{s}$$

$$m_{\pi^+}^2 \propto m_u m_s + m_d m_s$$

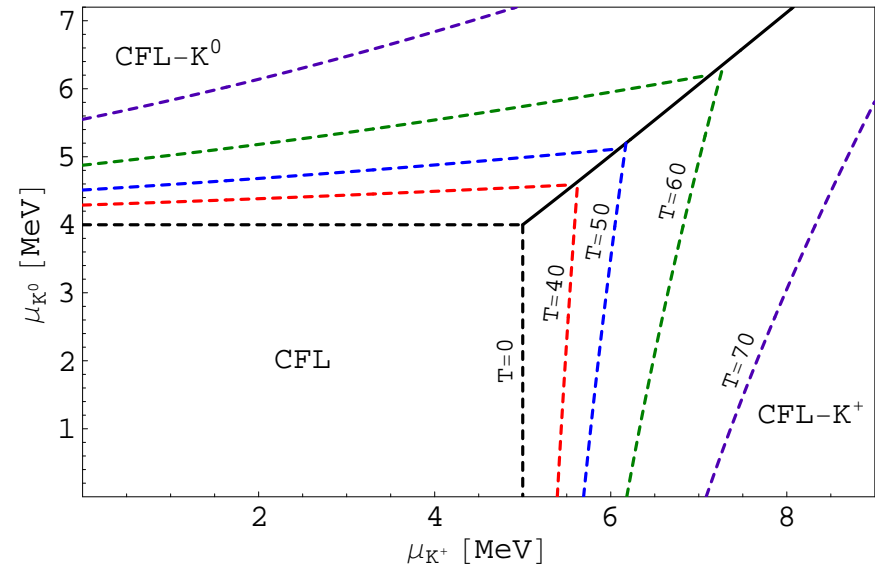


- **Kaon condensation (page 1/2)**

- kaon condensation at  $T = 0$  for  $\mu_{K^0} > m_{K^0}$

- $T \neq 0$ : 2PI formalism

M. G. Alford, M. Braby, A. Schmitt,  
J. Phys. G 35, 025002 (2008)



$\Rightarrow$  condensate (if present at  $T = 0$ ) very robust,  $T_c \sim T_c^{\text{CFL}}$

- Goldstone mode in CFL- $K^0$

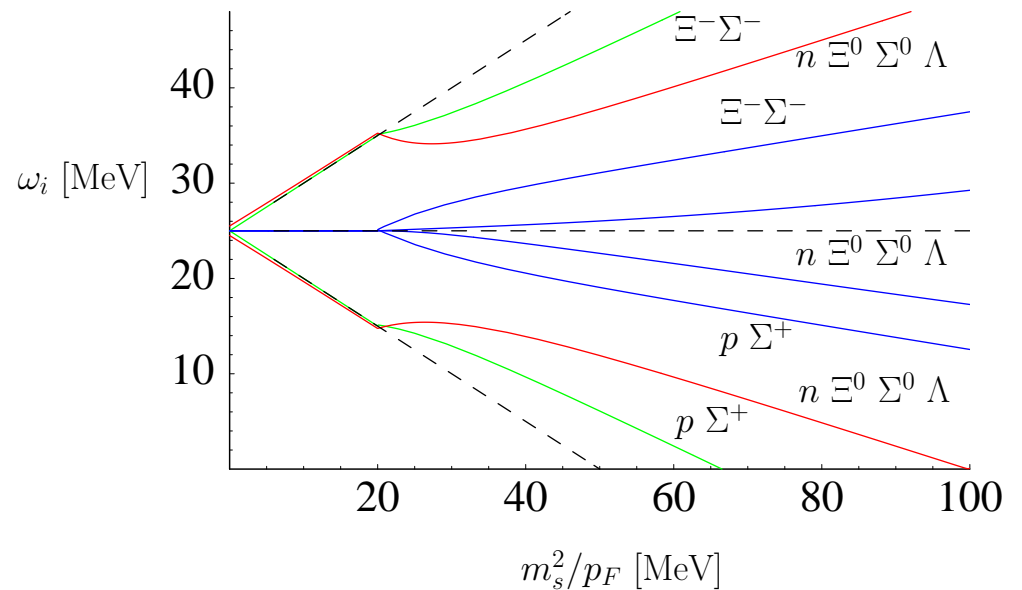
$$\epsilon_k = \begin{cases} \left\{ v_\pi^2 k^2 + M^2(T) + \mu_{K^0}^2 - \sqrt{4\mu_{K^0}^2 v_\pi^2 k^2 + [M^2(T) + \mu_{K^0}^2]^2} \right\}^{1/2} & \text{for } T < T_c \\ \sqrt{v_\pi^2 k^2 + M^2(T)} - \mu_{K^0} & \text{for } T > T_c \end{cases}$$

- **Kaon condensation (page 2/2)**

- **kaon condensate** shifts the onset of **gapless modes** (remember **chromomagnetic instability**)
- include fermions into effective theory as baryons  $N \sim \psi \langle \psi \psi \rangle$

A. Kryjevski, T. Schäfer, PLB 606, 52 (2005)

- fermion spectrum in CFL- $K^0$
- here:  $\Delta = 25$  MeV



Onset of **gapless modes**:  $\frac{M_s^2}{2p_F} = \Delta$  (CFL);  $\frac{M_s^2}{2p_F} = \frac{4\Delta}{3}$  (CFL- $K^0$ )

- **Part VI**
- Basics of (color) superconductivity
- Color-flavor locking (CFL) – highest densities
- Stressed pairing – below CFL densities
- QCD calculations
- Effective theory of CFL – kaon condensation
- **Transport properties – quark matter in compact stars**



# • Color superconductivity in compact stars (I)

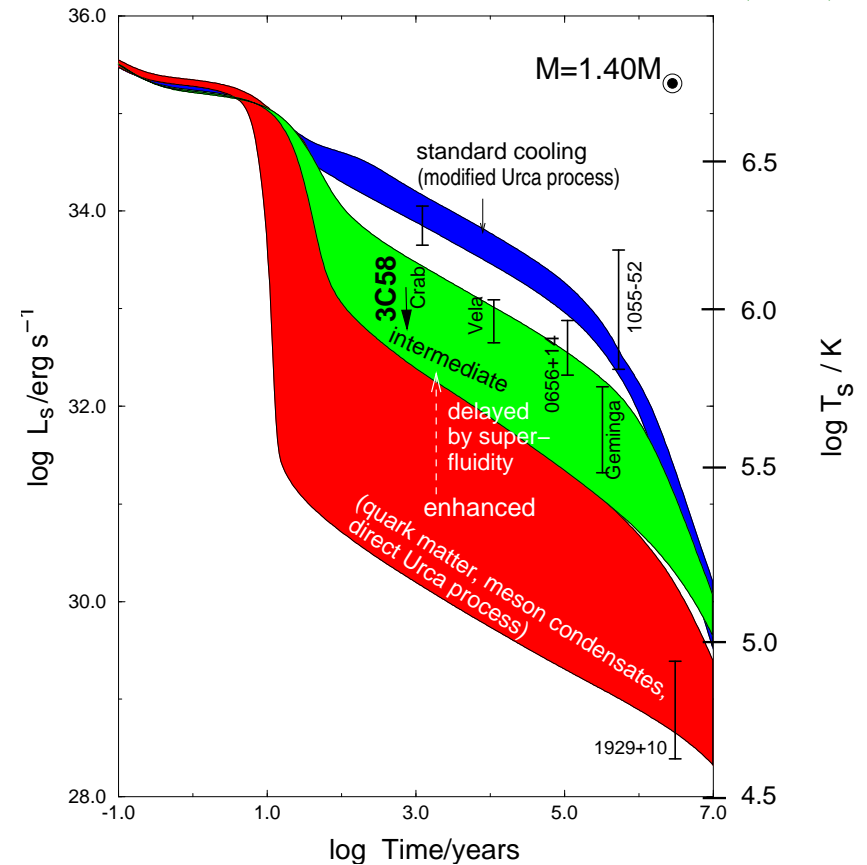
## Cooling of the star

### • neutrino emissivity

$$T \sim 10 \text{ MeV} \rightarrow 0.1 \text{ keV}$$

$$(t \sim 0 \rightarrow 10^6 \text{ yr})$$

F. Weber, Prog. Part. Nucl. Phys. 54, 193 (2005)



**unpaired quark matter** Iwamoto, PRL 44, 1637 (1980) **CFL** Jaikumar, Prakash, Schäfer, PRD 66, 063003 (2002) **2SC** Jaikumar, Roberts, Sedrakian, PRC 73, 042801 (2006) **spin-1** Schmitt, Shovkovy, Wang, PRD 73, 034012 (2006) **LOFF** Anglani, Nardulli, Ruggieri, Mannarelli, PRD 74, 074005 (2006)

- **Neutrino emissivity**

**CFL**

- (pseudo-)Goldstone processes

$$\pi^\pm, K^\pm \rightarrow e^\pm + \bar{\nu}_e$$

$$\pi^0 \rightarrow \nu_e + \bar{\nu}_e$$

$$\varphi + \varphi \rightarrow \varphi + \nu_e + \bar{\nu}_e$$

$\Rightarrow$  very small emissivity

$$\epsilon_\nu \sim \frac{G_F^2 T^{15}}{f^2 \mu^4}$$

**Non-CFL**

- **direct Urca process**

$$u + e \rightarrow d + \nu_e$$

$$d \rightarrow u + e + \bar{\nu}_e$$

$\Rightarrow$

$$\epsilon_\nu \simeq \frac{457}{630} \alpha_s G_F^2 T^6 \mu_e \mu_u \mu_d$$

(unpaired quark matter)

- non-CFL

(2SC, LOFF, spin-1, ...):

$\rightarrow$  **ungapped quarks**

- possible exception (?): CSL

- **Color superconductivity in compact stars (II)**

**magnetic fields**

- **spin-0 no Meissner effect**

M.G. Alford, J. Berges, K. Rajagopal,  
NPB 571, 269 (2000)

- **spin-1 Meissner effect**

A. Schmitt, Q. Wang, D.H. Rischke,  
PRL 91, 242301 (2003)

- ferromagnetism in the **curCFL- $K^0$**  phase

D. T. Son and M. A. Stephanov, PRD 77, 014021 (2008)

- “magnetic” CFL

E.J. Ferrer, V. de la Incera, C. Manuel,  
PRL 95, 152002 (2005)

- **de Haas-van Alphen oscillations**

K. Fukushima and H. J. Warringa,  
PRL 100, 032007 (2008)  
J. L. Noronha and I. A. Shovkovy,  
PRD 76, 105030 (2007)

- **Color superconductivity in compact stars (III)**

**glitches**

- **need vortex pinning at some crystal**
- maybe: neutron superfluidity + ion lattice in crust
- however: inconsistent with precession of star on  $\sim$  yr time scale

B. Link, arXiv:astro-ph/0608319

- better (?): crystalline color superconductivity
- **large shear modulus of CubeX** (and 2Cube45z)

M. Mannarelli, K. Rajagopal, R. Sharma, PRD 76, 074026 (2007)

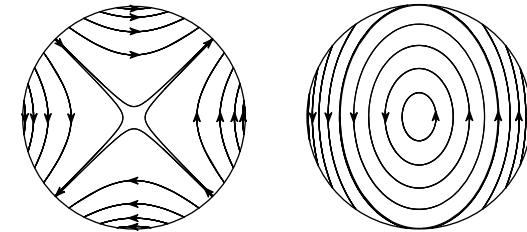
$$\nu = 3.96 \times 10^{33} \text{erg/cm}^3 \left( \frac{\Delta}{10 \text{ MeV}} \right)^2 \left( \frac{\mu}{400 \text{ MeV}} \right)^2$$

- 20 – 100 times more rigid than neutron star crust

## • Color superconductivity in compact stars (IV)

### r-mode instability

- **r-modes:**  
non-radial pulsation modes
- **grow unstable**  
in a **perfect-fluid** rotating star  
→ emission of gravitational waves
- fast rotating stars are observed!  
 $\omega \simeq 1 \text{ms}^{-1}$
- must be some damping mechanism → **bulk/shear viscosity**

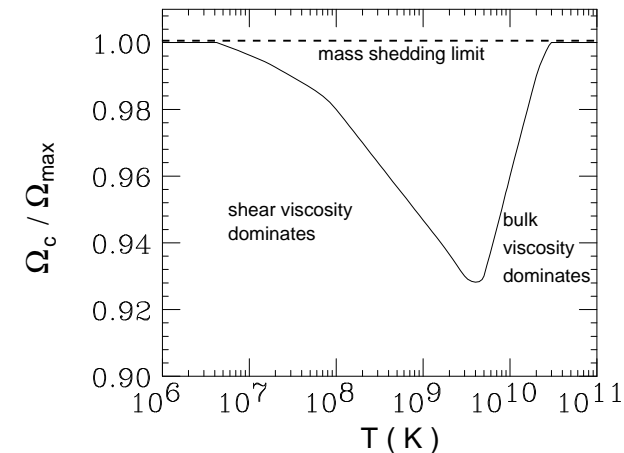


Polar View

Equatorial View

L. Lindblom, [arXiv:astro-ph/0101136](https://arxiv.org/abs/astro-ph/0101136)

- **spin down**  
the star drastically and quickly (within days)

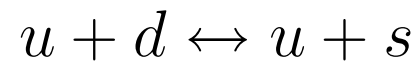


- **What is bulk viscosity?**

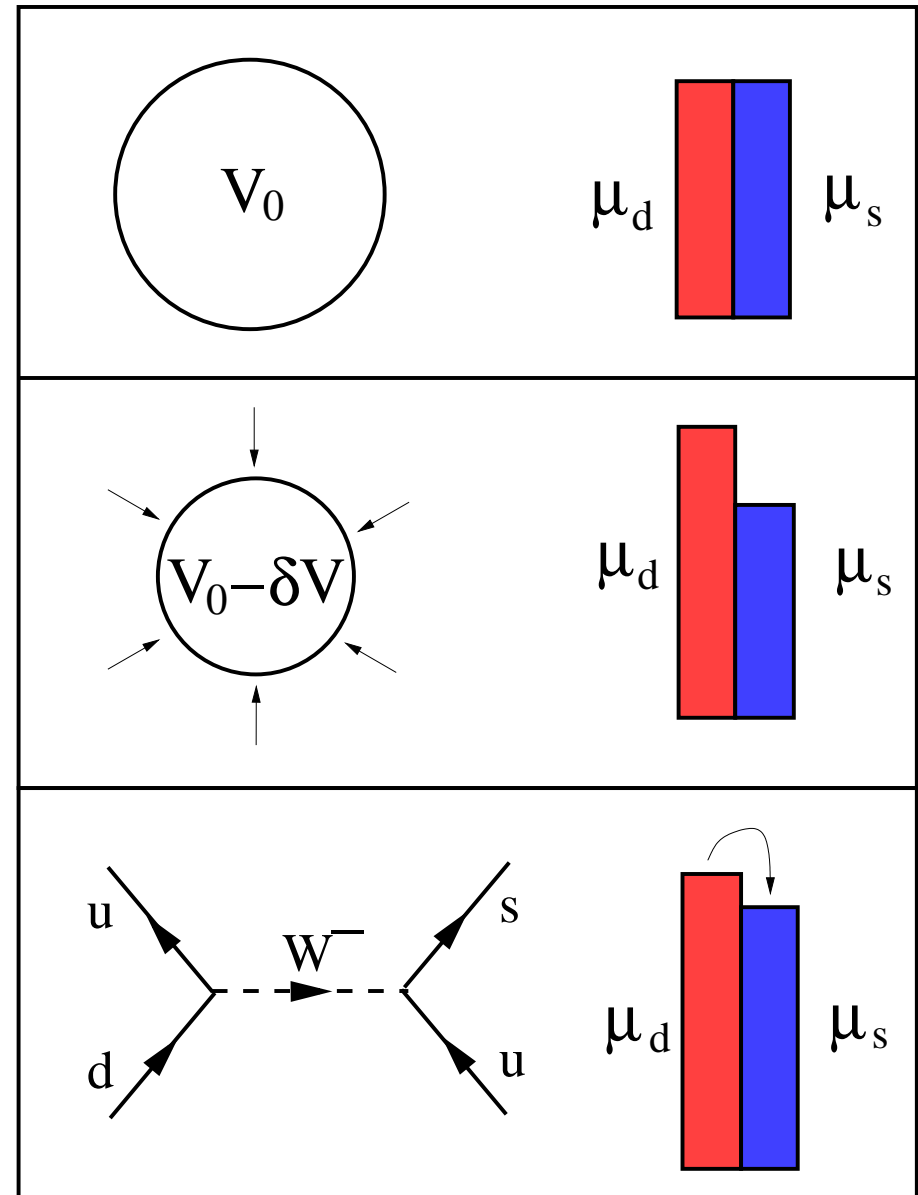
- volume oscillation  
→ chemical  
non-equilibrium

$$\mu_d - \mu_s \neq 0$$

- re-equilibration via

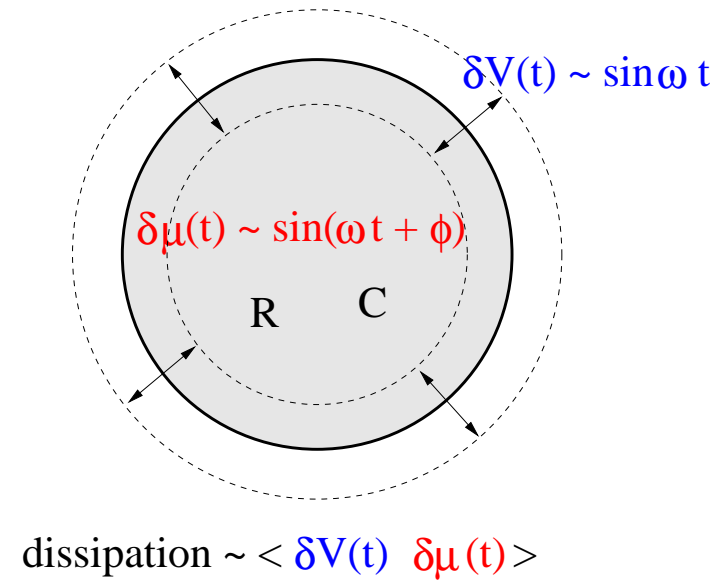
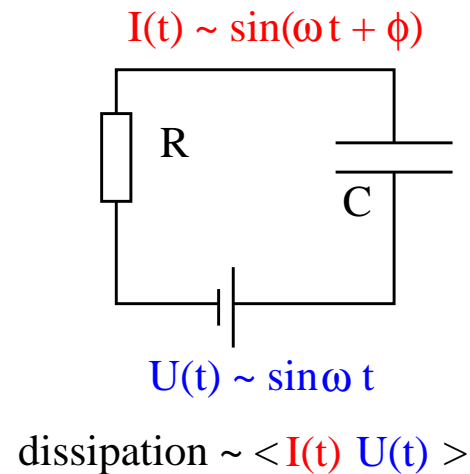


- **resonance phenomenon:**  
external oscillation  
vs. microscopic rate



- Bulk viscosity is a resonance phenomenon**

Just like an electric circuit!



“capacitance”  $C \leftrightarrow$  inverse microscopic rate  $\gamma^{-1}$   
(slow process  $\rightarrow$  store large chemical energy)

“resistance”  $R \leftrightarrow \left( n_u \frac{\partial \mu_d}{\partial n_u} + n_d \frac{\partial \mu_d}{\partial n_d} - n_s \frac{\partial \mu_s}{\partial n_s} \right)^{-1}$   
(same dispersion for  $d$  and  $s \rightarrow$  infinite “resistance”  $\rightarrow$  no dissipation)

**Bulk viscosity**

$$\zeta = \alpha \frac{\gamma}{\gamma^2 + \omega^2}$$

$$\alpha \equiv \frac{n_u \frac{\partial \mu_d}{\partial n_u} + n_d \frac{\partial \mu_d}{\partial n_d} - n_s \frac{\partial \mu_s}{\partial n_s}}{\frac{\partial \mu_d}{\partial n_d} + \frac{\partial \mu_s}{\partial n_s}}$$

• Compute rate for  $u + d \leftrightarrow u + s$  in 2SC

$$\Gamma_{2SC} = 4 \left[ \begin{array}{c} u \Delta \\ \Delta \\ d \Delta \end{array} \right] W^- \left[ \begin{array}{c} s \\ \Delta \\ u \Delta \end{array} \right] + 2 \left[ \begin{array}{c} u \Delta \\ \Delta \\ d \Delta \end{array} \right] W^- \left[ \begin{array}{c} s \\ u \\ u \end{array} \right] + 2 \left[ \begin{array}{c} u \\ \Delta \\ d \Delta \end{array} \right] W^- \left[ \begin{array}{c} s \\ \Delta \\ u \end{array} \right] + \left[ \begin{array}{c} u \\ \Delta \\ d \Delta \end{array} \right] W^- \left[ \begin{array}{c} s \\ u \\ u \end{array} \right]$$

$$+ 4 \left[ \begin{array}{c} u \Delta \\ \Delta \\ d \Delta \end{array} \right] W^- \left[ \begin{array}{c} s \\ \Delta \\ u \Delta \end{array} \right] + 2 \left[ \begin{array}{c} u \Delta \\ \Delta \\ d \Delta \end{array} \right] W^- \left[ \begin{array}{c} s \\ u \\ u \end{array} \right]$$

paired:



unpaired:



small temperatures,

$$T \ll T_c \simeq 30\text{MeV}$$

$$\Gamma_{2SC} = \frac{1}{9} \Gamma_{\text{unpaired}}$$

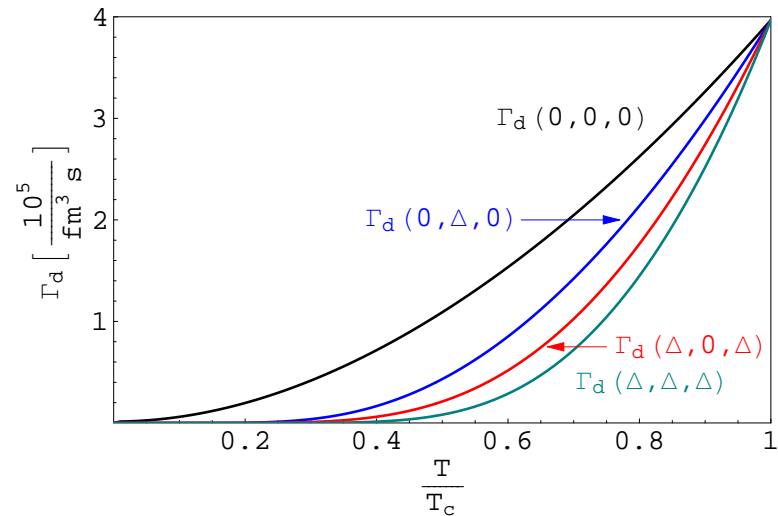
due to **exponential suppression**  
 $\exp(-\Delta/T)$  of gapped modes



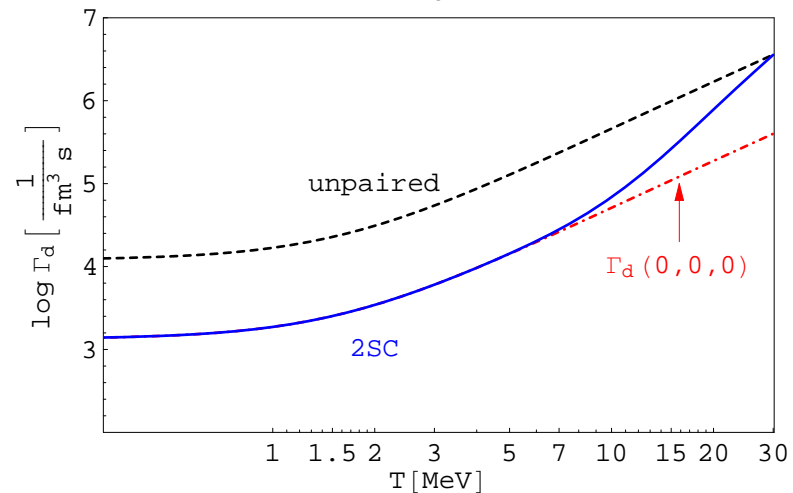
- **Results for all temperatures  $T < T_c$**

(i) fixed  $\delta\mu = \mu_s - \mu_d > 0 \rightarrow$  net production of  $d$  quarks,  $\Gamma_d > 0$

- **contributions of subprocesses**



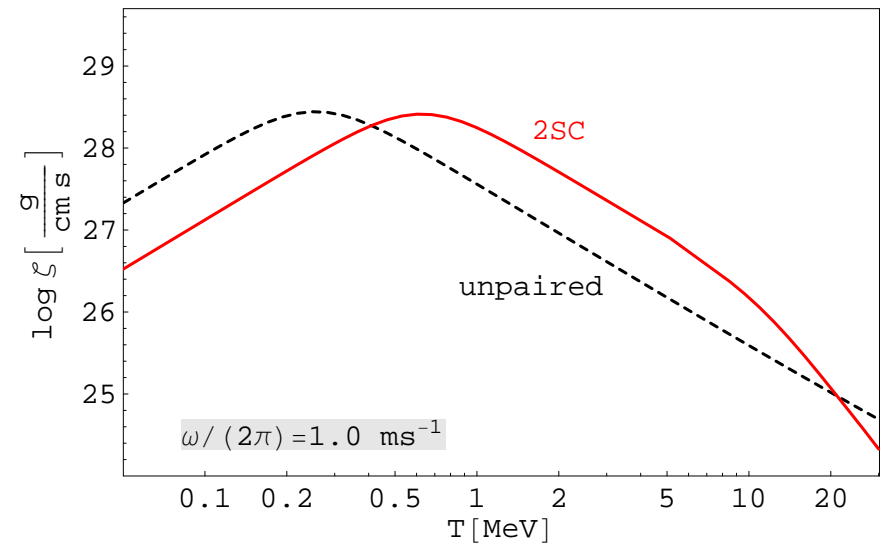
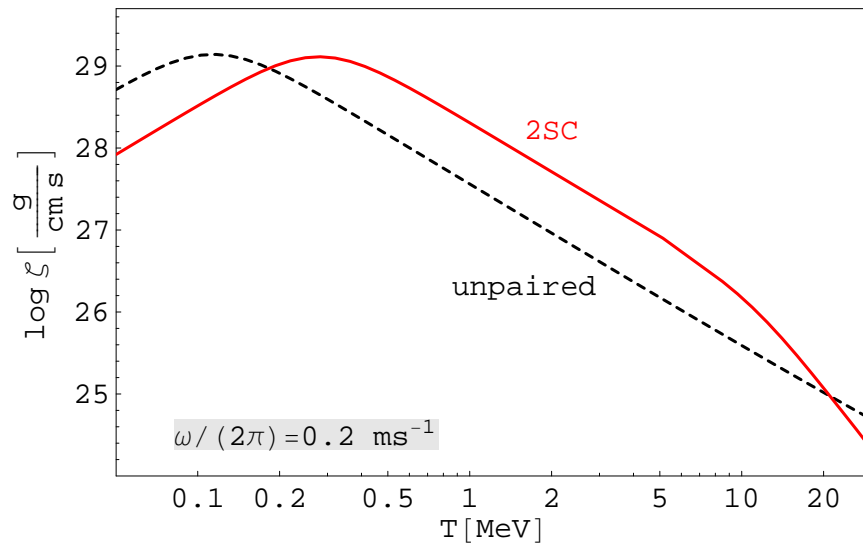
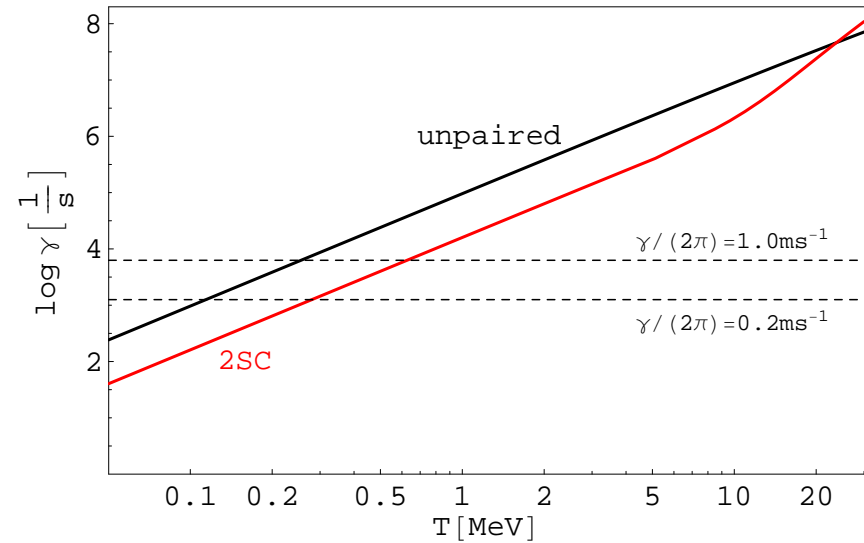
- **total rate compared to unpaired quark matter**



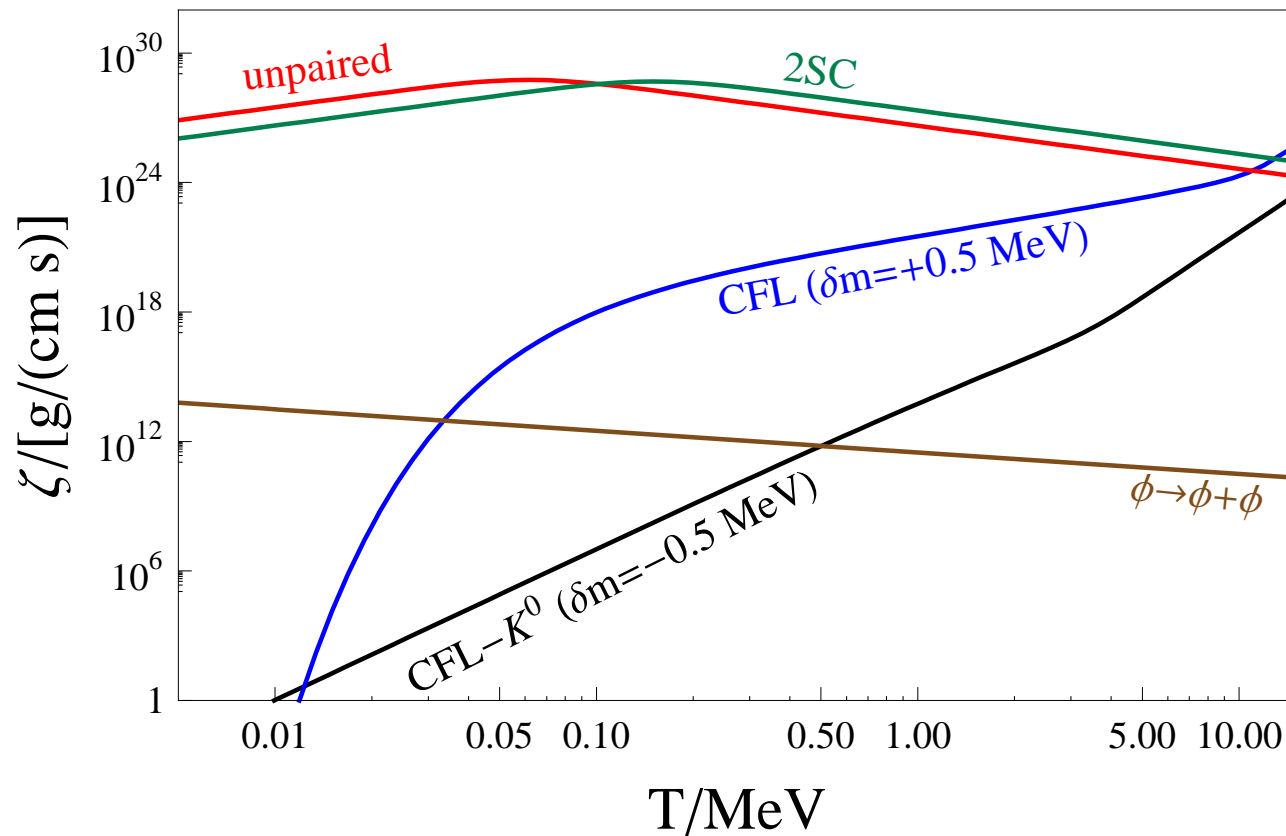
(ii) for bulk viscosity, consider  $\gamma = \frac{\partial \Gamma_d}{\partial \delta\mu}$

## • Results for bulk viscosity

$$\zeta = \alpha \frac{\gamma}{\gamma^2 + \omega^2}$$



## • Quark matter bulk viscosity: different phases



$$\omega/(2\pi) = 1 \text{ ms}^{-1}$$

$$\mu = 400 \text{ MeV}$$

$$\delta m \equiv m_{K^0} - \mu_{K^0}$$

**unpaired** J. Madsen, PRD 46, 3290 (1992)

**2SC** M.G. Alford, A. Schmitt, JPG 34, 67-101 (2007)

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- **Summary Parts IV – VI**
- **QCD** gap equation gives **parametrically different** result compared to BCS,  $\Delta \propto \exp(-\text{const}/g)$
- **effective theory** of CFL provides powerful tool to study **moderate densities** (all gluons & quarks are gapped)
- **transport properties** dominated by **Goldstone modes** in CFL ( $\rightarrow$  small  $\nu$ -emissivity, specific heat, bulk viscosity ...) and **ungapped quarks** in non-CFL ( $\rightarrow$  large emissivity ...)