## COMPUTATION OF FOKKER-PLANCK EQUATION

By

## STEPHEN S.-T. YAU

Department of Mathematics, Statistics and Computer Science, M/C 249, 851 South Morgan Street, Chicago, Illinois 60607-7045

**Abstract.** In plasma physics, the interaction of radio-frequency waves with a plasma is described by a Fokker-Planck equation with an added quasilinear term. In nonlinear filtering with conditional probability density of the state  $x_t$  given the observation  $\{y(s): 0 \le s \le t\}$  is also described by a Fokker-Planck equation with an added first order term. Method for solving Fokker-Planck equation by means of ordinary differential equations is discussed.

1. Introduction. Fokker-Planck models are most useful for the study of rf-driven currents [Fi] or neutral beam heating in tokamaks on time scales longer than the collisional time  $(\tau_{\text{coll}})$ . In [Ka], Karney took the plasma to be azimuthally symmetric about the magnetic field and homogeneous (representative of the central portion of a tokamak plasma). He presented some numerical methods to solve the Fokker-Planck equation in time and two velocity (or momentum) dimensions only. However, a complete Fokker-Planck treatment of rf or neutral beam heating in tokamaks generally requires solution of an equation which is at least two dimensions in momentum/velocity space and two dimensions in configuration space. A reduction in dimensionality occurs in cases where the bounce/transit time of the particles,  $\tau_b$ , is shorter when compared to the collision time, i.e.  $\tau_b \ll \tau_{\rm coll}$ . The present generation of larger tokamak experimental devices often operate with most of the plasma in this low-collisionality "banana" regime. Moreover, it is usually the case that the non-Maxwellian particles generated by auxiliary heating and current drive are in the low-collisionality regime. In such cases, a "bounce-average" over the bounce or toroidal transit motion of the particle is appropriate, reducing the Fokker-Planck equation to be essentially three-dimensional since the particle distributions as a function of poloidal angle become constant when expressed as a function of the collisionless constants of motion. In [Ke-Mc], Kerbel and McCoy developed a numerical solution scheme for the 3-dimensional Fokker-Planck equation. Therefore in plasma

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 $E ext{-}mail\ address: yauQuic.edu}$ 

physics, it remains very desirable to solve the Fokker-Planck equation for dimensions larger than 3.

On the other hand, nonlinear filtering is concerned with making estimates of quantities associated with a stochastic process  $\{x_t\}$  on the basis of information gleaned from a related process  $\{y_t\}$ . The process  $\{x_t\}$  is called the signal or state process and  $\{y_t\}$  is called the observation process. The goal is the computation, for each t, of least square estimates of functions of the signal  $\{x_t\}$  given the observation history  $\{y_s: 0 \le s \le t\}$ , i.e., the computation of conditional expectations of the form  $E[\phi(x_t)|y_s: 0 \le s \le t]$ , or perhaps even the computation of the entire conditional distribution  $\rho(t,x)$  of  $x_t$ . It is well known that  $\rho(t,x)$  is given by normalizing a function  $\sigma(t,x)$  which satisfies a Duncan-Mortensen-Zakai (DMZ) equation. By gauge transformation, one can show that it is sufficient to solve the robust DMZ equation, which is essentially the Fokker-Planck equation with an added first order term. In [Ya-Ya], it is shown that in order to get a numerical solution of the robust DMZ equation, it is sufficient to find numerical solution of the Fokker-Planck equation. Hence from the point of view of nonlinear filtering, it is also very desirable to solve the Fokker-Planck equation for dimensions larger than 3.

The purpose of this paper is to present an ODE method to solve the Fokker-Planck equation. These ODEs are of first order. The total number of ODEs is  $n + n^2$ , where n is the state space dimension of the Fokker-Planck equation. Since there are many well-known ODE solvers, the Fokker-Planck equation can be solved very efficiently, even for very large n.

In Sec. 2, we recall the background of the Fokker-Planck equation from the viewpoints of plasma physics and nonlinear filtering. In Sec. 3, we show that certain kinds of Fokker-Planck equations can be solved easily by means of solution of system of first order ordinary differential equations.

## 2. Preliminaries.

2.1. The Fokker-Planck equation in plasma physics. We first recall the Fokker-Planck equation from the point of view of plasma physics. The Fokker-Planck equation for the electron e can be written as

(2.1.1) 
$$\frac{\partial f_e}{\partial t} - \sum_s C(f_e, f_s) + \nabla \cdot \overrightarrow{S}_w + \frac{q_e \overrightarrow{E}}{m_e} \cdot \nabla f_e = 0,$$

where  $q_s$  and  $m_s$  are the charge and mass of species s,  $C(f_a, f_b)$  is the collision term for species a colliding off species b, the sum extends over all the species of the plasma (typically electrons and ions),  $\vec{S}_w$  is the wave (w)-induced quasilinear flux, and  $\vec{E} = E \vec{V}_{\parallel}$  is the electric field (assumed to be parallel to the magnetic field). The quantity  $q_s$  carries the sign of the charge, thus  $q_e = -e$ . The subscript  $\parallel$  refers to the direction parallel to the magnetic field. The  $\nabla = \partial/\partial \vec{V}$  operator operates in velocity space.

Because collisions in a plasma are primarily due to small-angle scattering, the collision term can be written as the divergence of a flux

(2.1.2) 
$$C(f_a, f_b) = -\nabla \cdot S_c^{a/b}$$

in which Eq. (2.1.1) can be expressed as

(2.1.3) 
$$\frac{\partial f_e}{\partial t} + \nabla \cdot \overrightarrow{S} = 0,$$

where

$$\vec{S} = \vec{S}_c + \vec{S}_w + \vec{S}_e$$

is the total flux in velocity space, and

$$(2.1.4) \overrightarrow{S}_c = \sum_{c} S_c^{e/s}$$

$$(2.1.5) \vec{S_e} = \frac{q_e \vec{E}}{m_e} f_e$$

are the collisional (c)- and electric-field (e) - induced electron fluxes.

Typically, two types of terms appear in  $\overline{S}$ : a diffusion term and a friction term

$$(2.1.6) \qquad \overrightarrow{S} = -D \cdot \nabla f_e + \overrightarrow{F} f_e.$$

Combining (2.1.3) and (2.1.6), we see that the Fokker-Planck equation in plasma physics looks like:

(2.1.7) 
$$\frac{\partial f_e}{\partial t} = \nabla (D \cdot \nabla f_e) + \nabla \cdot (\overrightarrow{F} f_e).$$

2.2. Fokker-Planck equation in nonlinear filtering. The filtering problem considered here is based on the following signal observation model:

(2.2.1) 
$$\begin{cases} dx(t) = f(x(t))dt + g(x(t))dv(t) & x(0) = x_0 \\ dy(t) = h(x(t))dt + dw(t) & y(0) = 0 \end{cases}$$

in which x, v, y, and w are respectively  $\mathbb{R}^n$ ,  $\mathbb{R}^p$ ,  $\mathbb{R}^m$ , and  $\mathbb{R}^m$  valued processes and v and w have components that are independent, standard Brownian processes. We further assume that n = p; f, g, and h are vector-valued, matrix-valued, and vector-valued  $C^{\infty}$  smooth functions. We shall refer to x(t) as the state of this system at time t and to y(t) as the observation at time t.

Let  $\rho(t,x)$  denote the conditional probability density of the state given the observation  $\{y(s): 0 \leq s \leq t\}$ . It is well known that  $\rho(t,x)$  is given by normalizing a function  $\sigma(t,x)$  that satisfies the following Duncan-Mortensen-Zakai equation:

(2.2.2) 
$$\begin{cases} d\sigma(t,x) = L_0\sigma(t,x)dt + \sum_{i=1}^m L_i\sigma(t,x)dy_i(t) \\ \sigma(0,x) = \sigma_0(x), \end{cases}$$

where

$$L_0 = \frac{1}{2} \sum_{i,j=1}^n \frac{\partial^2}{\partial x_i \partial x_j} \left[ g(x) g^T(x) \right]_{ij} - \sum_{i=1}^n f_i \frac{\partial}{\partial x_i} - \sum_{i=1}^n \frac{\partial f_i}{\partial x_i} - \frac{1}{2} \sum_{i=1}^m h_i^2$$

for  $i = 1, ..., m, L_i$  is the zero degree differential operator of multiplication by  $h_i$ , and  $\sigma_0$  is the probability density of the initial point  $x_0$ . In most of the applications,  $[g(x)g^T(x)]_{ij}$ 

are assumed to be constants  $G_{ij}$ ,  $1 \le i, j \le n$ . Note that  $G_{ij} = G_{ji}$ . Then

(2.2.3) 
$$L_{0} = \frac{1}{2} \sum_{i,j=1}^{n} G_{ij} \frac{\partial^{2}}{\partial x_{i} \partial x_{j}} - \sum_{i=1}^{n} f_{i} \frac{\partial}{\partial x_{i}} - \sum_{i=1}^{n} \frac{\partial f_{i}}{\partial x_{i}} - \frac{1}{2} \sum_{i=1}^{m} h_{i}^{2}.$$

Equation (2.2.2) is a stochastic partial differential equation in the sense of Stratonovich. Define a new unnormalized density

(2.2.4) 
$$u(t,x) = \exp\left(-\sum_{i=1}^{m} h_i(x)y_i(t)\right)\sigma(t,x).$$

Then we can reduce (2.2.2) to the following time varying partial differential equation

(2.2.5) 
$$\begin{cases} \frac{\partial u}{\partial t}(t,x) &= L_0 u(t,x) + \sum_{i=1}^m y_i(t) [L_0, L_i] u(t,x) \\ &+ \frac{1}{2} \sum_{i,j=1}^m y_i(t) y_j(t) [[L_0, L_i], L_j] u(t,x) \\ u(0,x) &= \sigma_0(x), \end{cases}$$

where  $[L_0, L_i]$  denotes the commutator of the differential operators.

Using the similar technique developed in [Ya-Ya], one can show that it is enough to solve the Fokker-Planck equation of the following form:

$$(2.2.6) \quad \frac{\partial u}{\partial t}(t,x) = \frac{1}{2} \sum_{i,j=1}^{m} G_{ij} \frac{\partial^{2} u}{\partial x_{i} \partial x_{j}}(t,x) - \sum_{i=1}^{n} f_{i}(x) \frac{\partial u}{\partial x_{i}}(t,x) - \left(\sum_{i=1}^{n} \frac{\partial f_{i}}{\partial x_{i}}(x) + \frac{1}{2} \sum_{i=1}^{m} h_{i}^{2}(x)\right) u(t,x).$$

2.3. Fokker-Planck equation in general form. In view of (2.1.7) and (2.2.6), we shall consider the following general form of the Fokker-Planck equation on  $\mathbb{R}^n$ :

(2.3.1)

$$\begin{cases} \frac{\partial u}{\partial t}(t,x) &= \sum_{i,j=1}^{n} G_{ij} \frac{\partial^{2} u}{\partial x_{i} \partial x_{j}}(t,x) - \sum_{i=1}^{n} f_{i}(x) \frac{\partial u}{\partial x_{i}}(t,x) + q(x)u(t,x) \\ u(0,x) &= u_{0}(x) \end{cases}$$

where  $G_{ij} = G_{ji}$  are constants.

3. Explicit solution of the Fokker-Planck equation in terms of solutions of ODEs. We shall solve the Fokker-Planck equation (2.3.1) where  $f_i(x)$  are degree one polynomials and Q(x) is a degree 2 polynomial:

(3.1) 
$$f_i(x) = \ell_i(x) = \sum_{j=1}^n d_{ij}x_j + d_i, \quad 1 \le i \le n$$

$$(3.2) q(x) = x^T Q x + p^T x + r.$$

where  $Q = (q_{ij})$  is an  $n \times n$  symmetric matrix,  $p^T = (p_1, \dots, p_n)$  and  $x^T = (x_1, \dots, x_n)$  are  $1 \times n$  matrices, and r is a scalar.

It is well known that any distribution is well approximated by a finite linear combination of Gaussians of the form  $\alpha_1 G_1 + \cdots + \alpha_k G_k$ , where  $\alpha_i$ 's are real numbers and  $G_i$ 's

are Gaussian distributions. Let  $u_i$  be the solution of (2.3.1) with initial distribution  $G_i$ . Since (2.3.1) is a linear partial differential equation, it follows that the solution of (2.3.1) is of the form  $\alpha_1 u_1 + \cdots + \alpha_k u_k$ . Therefore it remains to solve (2.3.1) with Gaussian initial distribution. The following theorem gives an explicit solution of (2.3.1) with Gaussian initial distribution in terms of solutions of ODEs.

**Theorem 3.1.** Consider the Fokker-Planck equation on  $\mathbb{R}^n$  with Gaussian initial distribution

(3.3) 
$$\begin{cases} \frac{\partial u}{\partial t}(t,x) = \sum_{i,j=1}^{n} G_{ij} \frac{\partial^{2} u}{\partial x_{i} \partial x_{j}}(t,x) - \sum_{i=1}^{n} \ell_{i}(x) \frac{\partial u}{\partial x_{i}}(t,x) + q(x)u(t,x) \\ u(0,x) = e^{x^{T} A(0)x + B^{T}(0)x + C(0)}, \end{cases}$$

where  $G_{ij} = G_{ji}$  are constants,  $A(0) = (A_{ij}(0))$  is an  $n \times n$  matrix,  $B^T(0) = (B_1(0), \ldots, B_n(0))$ , and  $x^T = (x_1, \ldots, x_n)$  are  $1 \times n$  matrices, and C(0) is a scalar. Suppose conditions (3.1) and (3.2) hold. Then the solution of (3.3) is of the following form:

(3.4) 
$$u(t,x) = e^{x^T A(t)x + B^T(t)x + C},$$

where A(t) is an  $n \times n$  symmetric matrix valued function of t,  $B^{T}(t) = (B_{1}(t), \dots, B_{n}(t))$  is a  $1 \times n$  matrix valued function of t, and C(t) is a scalar function of t. Moreover, A(t),  $B^{T}(t)$ , and C(t) satisfy the following system of ODEs:

(3.5) 
$$\frac{dA}{dt} = A^{T}GA + 2AGA + AG^{T}A^{T} - (A^{T} + A)D + Q$$
(3.6) 
$$\frac{dB^{T}}{dt} = B^{T}GA + 2B^{T}G^{T}A^{T} + B^{T}G^{T}A - B^{T}D - d^{T}A - d^{T}A^{T} + p^{T}$$
(3.7) 
$$\frac{dC}{dt} = 2tr(GA) + B^{T}GB - d^{T}B + r.$$

PROOF: Differentiating (3.4) with respect to t and  $x_i$ , we get the following equations:

(3.8) 
$$\frac{\partial u}{\partial t} = \left(x^T \frac{dA}{dt} x + \frac{dB^T}{dt} x + \frac{dC}{dt}\right) u$$

$$\frac{\partial u}{\partial x_j} = \left[\sum_{i,k=1}^n A_{ik} \left(\frac{\partial x_i}{\partial x_j} x_k + x_i \frac{\partial x_k}{\partial x_j}\right) + B_j\right] u$$

$$= \left(\sum_{k=1}^n A_{jk} x_k + \sum_{i=1}^n A_{ij} x_i + B_j\right) u$$

$$\nabla u^T = \left[(Ax)^T + x^T A + B^T\right] u = (x^T A^T + x^T A + B^T) u$$

$$\begin{split} \frac{\partial^2 u}{\partial x_i \partial x_j} &= \left[ A_{ji} + A_{ij} + \left( \sum_{k=1}^n A_{jk} x_k + \sum_{k=1}^n A_{kj} x_k + B_j \right) \left( \sum_{k=1}^n A_{ik} x_k \right. \right. \\ &+ \sum_{k=1}^n A_{ki} x_k + B_i \right] u \\ &= \left[ A_{ji} + A_{ij} + \sum_{k,\ell=1}^n A_{jk} A_{i\ell} x_k x_\ell + \sum_{k,\ell=1}^n A_{jk} A_{\ell i} x_k x_\ell + B_i \sum_{k=1}^n A_{jk} x_k \right. \\ &+ \sum_{k,\ell=1}^n A_{kj} A_{i\ell} x_k x_\ell + \sum_{k,\ell=1}^n A_{kj} A_{\ell i} x_k x_\ell + B_i \sum_{k=1}^n A_{kj} x_k \\ &+ B_j \sum_{k=1}^n A_{ik} x_k + B_j \sum_{k=1}^n A_{kj} x_k + B_j B_i \right] u \\ &\sum_{i,j=1}^n G_{ij} \frac{\partial^2 u}{\partial x_i \partial x_j} &= \left[ \sum_{i,j=1}^n G_{ij} A_{ji} + \sum_{i,j=1}^n G_{ij} A_{ij} + \sum_{i,j,k,\ell=1}^n G_{ij} A_{jk} A_{i\ell} x_k x_\ell \right. \\ &+ \sum_{i,j,k=1}^n G_{ij} B_{ij} A_{jk} x_k + \sum_{i,j,k,\ell=1}^n G_{ij} A_{kj} A_{i\ell} x_k x_\ell + \sum_{i,j,k,\ell=1}^n G_{ij} A_{kj} A_{\ell i} x_k x_\ell \\ &+ \sum_{i,j,k=1}^n G_{ij} B_i A_{jk} x_k + \sum_{i,j,k=1}^n G_{ij} B_j A_{ik} x_k + \sum_{i,j,k=1}^n G_{ij} B_j A_{ki} x_k \\ &+ \sum_{i,j}^n G_{ij} B_j B_i \right] u \\ &= \left[ \sum_{i,j=1}^n G_{ij} A_{ji} + \sum_{i,j=1}^n G_{ji} A_{ij} + \sum_{i=1}^n \sum_{j=1}^n A_{i\ell} x_\ell \right) \left( \sum_{j,k=1}^n G_{ij} A_{jk} x_k \right. \\ &+ \sum_{i=1}^n \sum_{\ell=1}^n A_{\ell i} x_\ell \right) \left( \sum_{j,k=1}^n G_{ij} A_{jk} x_k \right. \\ &+ \sum_{i=1}^n \sum_{k=1}^n X_k A_{kj} \right) \left( \sum_{i,\ell=1}^n A_{i\ell} A_{ij} \right) \\ &+ \sum_{i=1}^n \sum_{k=1}^n X_k A_{kj} \right) \left( \sum_{i,\ell=1}^n A_{i\ell} A_{ij} \right) \\ &+ \sum_{i=1}^n \sum_{k=1}^n X_k A_{kj} \right) \left( \sum_{i,\ell=1}^n A_{i\ell} A_{ij} \right) \\ &+ \sum_{i=1}^n \sum_{k=1}^n X_k A_{kj} \right) \left( \sum_{i,\ell=1}^n A_{i\ell} A_{ij} \right) \\ &+ \sum_{i=1}^n \sum_{k=1}^n X_k A_{kj} \right) \left( \sum_{i,\ell=1}^n A_{i\ell} A_{ij} \right) \\ &+ \sum_{i=1}^n \sum_{k=1}^n X_k A_{kj} \right) \left( \sum_{i,\ell=1}^n A_{i\ell} A_{ij} \right) \sum_{k=1}^n G_{ij} B_j \right) \left( \sum_{k=1}^n A_{ik} A_{kj} \right) \\ &+ \sum_{i=1}^n \sum_{k=1}^n X_k A_{kj} \right) \left( \sum_{i,\ell=1}^n A_{i\ell} A_{ij} \right) \sum_{k=1}^n G_{ij} B_j \right) \left( \sum_{k=1}^n A_{ik} A_{kj} \right) \\ &+ \sum_{i=1}^n \sum_{k=1}^n X_k A_{kj} \right) \left( \sum_{i,\ell=1}^n A_{i\ell} A_{ij} \right) \sum_{k=1}^n G_{ij} B_j \right) \left( \sum_{k=1}^n A_{ik} A_{kj} \right) \\ &+ \sum_{i=1}^n \sum_{k=1}^n X_k A_{kj} \right) \left( \sum_{i,\ell=1}^n A_{i\ell} A_{ij} \right) \sum_{k=1}^n A_{ik} A_{kj}$$

$$\begin{split} &= [2tr(GA) + (Ax)^T(GAx) + (x^TA)(GAx) + B^TGAx + x^TAGAx \\ &+ x^TA(x^TAG)^T + (x^TA)GB + (GB)^TAx + (x^TA)(GB) + B^TGB]u \\ &= [x^T(A^TGA + AGA + AGA + AG^TA^T)x + (B^TGA + B^TG^TA^T \\ &+ B^TG^TA + B^TG^TA^T)x + 2tr(GA) + B^TGB]u \\ &= [x^T(A^TGA + 2AGA + AG^TA^T) + (B^TGA + 2B^TG^TA^T + B^TG^TA)x \\ &+ 2tr(GA) + B^TGB]u. \end{split}$$

Let  $D = (d_{ij})$  be an  $n \times n$  matrix and  $d^T = (d_1, \dots, d_n)$  be a  $1 \times n$  matrix. Then

$$\sum_{i=1}^{n} \ell_i(x) \frac{\partial u}{\partial x_i} = \sum_{i=1}^{n} \sum_{j=1}^{n} d_{ij} x_j \frac{\partial u}{\partial x_i} + \sum_{i=1}^{n} d_i \frac{\partial u}{\partial x_i}$$

$$= \nabla u^T D x + d^T \nabla u$$

$$= (x^T A^T + x^T A + B^T) u D x + d^T (A x + A^T x + B) u$$

$$= [x^T (A^T + A) D x + (B^T D + d^T A + d^T A^T) x + d^T B] u.$$

Thus the L.H.S. of (3.3) is given by

$$(3.9) \quad \sum_{i,j=1}^{n} G_{ij} \frac{\partial^{2} u}{\partial x_{i} \partial x_{j}}(t,x) - \sum_{i=1}^{n} \ell_{i}(x) \frac{\partial u}{\partial x_{i}}(t,x) + q(x)u(t,x)$$

$$= [x^{T} (A^{T}GA + 2AGA + AG^{T}A^{T} - (A^{T} + A)D + Q)x + (B^{T}GA + 2B^{T}G^{T}A^{T} + B^{T}G^{T}A - B^{T}D - d^{T}A - d^{T}A^{T} + p^{T})x + 2tr(GA) + B^{T}GB - d^{T}B + r]u.$$

Equating (3.8) and (3.9) and comparing terms, we get equations (3.5), (3.6), and (3.7). Q.E.D.

**Theorem 3.2.** The solution of the following Fokker-Planck equation on  $\mathbb{R}^n$  can be found by the method described above.

(3.10)

$$\frac{\partial u}{\partial t}(t,x) = \sum_{i,j=1}^{n} G_{ij} \frac{\partial^{2} u}{\partial x_{i} \partial x_{j}}(t,x) - \sum_{i=1}^{n} \left( \sum_{j=1}^{n} G_{ij} \frac{\partial F}{\partial x_{j}}(x) + \ell_{i}(x) \right) \frac{\partial u}{\partial x_{i}}(t,x) + q(x)u(t,x)$$

where  $G_{ij} = G_{ji}$  are constants;  $\ell_i(x)$  are degree one polynomials; F(x) and q(x) are  $C^{\infty}$  functions on  $\mathbb{R}^n$  such that

$$(3.11) \quad \widetilde{q}(x) := \frac{1}{2} \sum_{i,j=1}^{n} G_{ij} \frac{\partial^{2} F}{\partial x_{i} \partial x_{j}}(x) + \frac{1}{4} \sum_{i,j=1}^{n} \frac{\partial F}{\partial x_{i}}(x) \frac{\partial F}{\partial x_{j}}(x) - \frac{1}{2} \sum_{i=1}^{n} \frac{\partial F}{\partial x_{i}}(x) \left( \sum_{j=1}^{n} G_{ij} \frac{\partial F}{\partial x_{j}}(x) + \ell_{i}(x) \right) + q(x)$$

is a polynomial of degree two. In fact, let

(3.12) 
$$u(t,x) = e^{\frac{1}{2}F(x)}\widetilde{u}(t,x).$$

Then (3.10) can be reduced to the following equation:

(3.13)

$$\frac{\partial \widetilde{u}}{\partial t}(t,x) = \sum_{i,j} G_{ij} \frac{\partial^2 \widetilde{u}}{\partial x_i \partial x_j}(t,x) - \sum_{i=1}^n \ell_i(x) \frac{\partial \widetilde{u}}{\partial x_i}(t,x) + \widetilde{q}(x) \widetilde{u}(t,x).$$

PROOF: By putting (3.12) into (3.10), one deduces that (3.13) holds. Q.E.D.

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