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## Crossover of angular dependent magnetoresistance with the metalinsulator transition in colossal magnetoresistive manganite films

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The temperature and magnetic field dependence of angular dependent magnetoresistance (AMR) along two orthogonal directions ([100] and [011]) was investigated in a charge-orbital-ordered  $Sm_{0.5}Ca_{0.5}MnO_3$  (SCMO) film grown on (011)-oriented  $SrTiO_3$  substrates. A dramatic decrease of AMR magnitude in both directions was observed with the appearance of magnetic-field-induced metal-insulator transition, which further led to a sign crossover in the AMR effect. The AMR crossover may give a direct evidence of the drastic modification of electronic structure or possible orbital reconstruction with the magnetic-destruction of charge/orbital ordering in SCMO films. © 2009 American Institute of Physics. [doi:10.1063/1.3240407]

The most striking aspect of physics for manganites is the occurrence of a magnetic-field-induced insulator-to-metal transition—colossal magnetoresistance (CMR) effect.<sup>1</sup> Besides magnetic fields, the metal-insulator transition (MIT) could be initiated by a number of other external stimuli, such as temperature, electric field, pressure, and irradiation with light, which exhibits intriguing applications potential in multifunctional oxide electronic devices.<sup>1</sup> The huge change in conductivity could be attributed to melting of long-range and/or short-range charge/orbital ordering (COO), a percolation transition, or a sliding charge-density wave.<sup>2</sup> However, no consensus in the microscopic mechanism of the MIT has been reached yet.

The competition between a ferromagnetic metallic state and a paramagnetic insulating phase underlies the CMR phenomenon. This phase competition appears due to the natural coupling between magnetism and orbital ordering. In fact, the spin-orbit coupling is also at the origin of the anisotropic magnetoresistance in ferromagnets.<sup>3,4</sup> Due to the strong coupling between the charge, orbital, spin, and lattice degrees of freedom, manganites present anomalous behavior in their anisotropic magnetoresistance or the angular dependent magnetoresistance (AMR) compared to the conventional ferromagnetic alloys.<sup>5-11</sup> For example, the AMR effect could persist up to the paramagnetic region. In particular, a peak in AMR is always observed near the Curie temperature  $T_C$  of ferromagnetic manganites,<sup>5-9</sup> where a spontaneous MIT together with a Jahn-Teller (JT) distortion occurs. Although it could not satisfactorily explained by the most traditional explanation of AMR as that in ferromagnetic metals appealing on the spin-orbit coupling and the anisotropic scattering rate of charge carriers, such AMR behavior seems to exhibit intimate correlation with JT distortion. Moreover, significant adjustments in the AMR, even a AMR sign change or crossover, have been observed in ultrathin Pr<sub>0.67</sub>Sr<sub>0.33</sub>MnO<sub>3</sub> (Ref. 10) and La<sub>0.65</sub>Ca<sub>0.35</sub>MnO<sub>3</sub> (Ref. 5) films, where the strainselected  $e_g$  orbital occupation prevails.<sup>12,13</sup> These results clearly indicate an important role of lattice and orbital effect in AMR of manganites. For charge-orbital-ordered manganites, the CMR effect is much larger and is always accompanied with a magnetic-field-induced MIT, during which a ferromagnetic metal generally takes place the insulating COO state and orbital reconstruction occurs. Thus, it should be rather interesting to investigate the AMR effect in COO manganites with the evolution of the magnetic-field-induced MIT. In this letter, a dramatic crossover in the AMR effect due to the collapse of COO is observed, which further indicates the close relation of the AMR abnormity and the MIT. The underlying mechanism is also discussed.

The SCMO films were grown on (011)-oriented SrTiO<sub>3</sub> substrates by pulsed laser deposition. The film thickness was  $\sim 60$  nm. The high quality of the films was confirmed by x-ray diffraction and reciprocal space maps measurements.<sup>14</sup> Magnetotransport properties along the [100] and  $[0\overline{1}1]$  directions,  $\rho_{100}$  and  $\rho_{011}$ , respectively, were measured using the four-terminal method in a bridge-shaped sample patterned by the conventional photolithography and chemical etching technique, which showed the dimensions of 0.1 mm in width and 0.4 mm in length. For AMR measurements, the magnetic field was rotated in the (011) plane to prevent the demagnetization effect and the Hall effect signals, which was performed in a Quantum Design physical property measurement system (PPMS-14h). The AMR effect was deduced from the total magnetoresistance data by subtracting the overwhelming CMR contributions.

No obvious AMR angular dependence  $[AMR(\theta)]$  is detected at a magnetic field  $\mu_0 H < 6$  T in the whole temperature range (2–300 K) for the nonferromagnetic SCMO films. Such situation persists even for the magnetic field up to 14 T for  $T > T_{CO}$  ( $T_{CO} \approx 235$  K, the charge ordering temperature). However, AMR( $\theta$ ) oscillations in both  $\rho_{100}$  and  $\rho_{0\bar{1}1}$  appear once the COO forms and the  $\mu_0 H$  is up to 6 T. Figures 1(a) and 1(b) show the linear (upper) and polar (lower) plots of the AMR( $\theta$ ) oscillations in  $\rho_{100}$  and  $\rho_{0\bar{1}1}$ , respectively, at 150 K under different magnetic field. Here, AMR( $\theta$ )=( $\rho_{\theta}$ 

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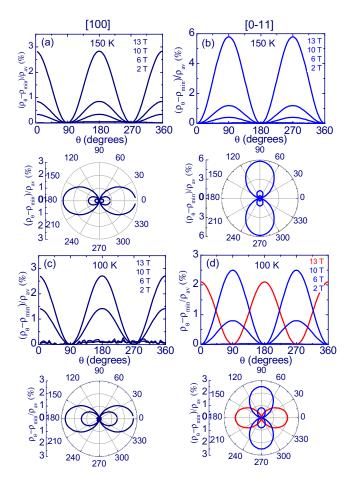


FIG. 1. (Color online) [(a)–(d)] Linear (upper) and polar (lower) plots of AMR oscillations,  $(\rho_{\theta} - \rho_{\min})/\rho_{av}$ , along both [100] and [011] directions at 100 K and 150 K under different magnetic field.

 $-\rho_{\rm min})/\rho_{\rm av}$ ,  $\rho_{\rm min}$ , and  $\rho_{\rm av}$  are the minimum and average resistivity, respectively, among the angular dependent resistivities. It is interesting to note that  $AMR(\theta)$  shows  $\cos^2 \theta$  dependence along [100] direction, while  $\sin^2 \theta$  dependence along  $[0\overline{1}1]$  direction, which is distinctly opposite to the AMR( $\theta$ ) observed in La<sub>0.95</sub>Ag<sub>0.05</sub>MnO<sub>3</sub> films with a ferromagnetic metallic state.<sup>6</sup> As also shown in the figures, further increasing magnetic field enhances the  $AMR(\theta)$  amplitude significantly. More interestingly, with decreasing temperature, a crossover of AMR( $\theta$ ) from sin<sup>2</sup>  $\theta$  to cos<sup>2</sup>  $\theta$  dependence is observed in  $\rho_{011}$  for  $\mu_0 H > 10$  T in 80 K < T <120 K, while the  $\cos^2 \theta$ -dependent AMR( $\theta$ ) remains for  $\rho_{100}$  as illustrated in Figs. 1(c) and 1(d) (T=100 K). When  $T \le 80$  K, another crossover of AMR( $\theta$ ) from  $\cos^2 \theta$  to  $\sin^2 \theta$  is also observed in  $\rho_{100}$ . As a result, the AMR( $\theta$ ) oscillations becomes similar to that of ferromagnetic metallic  $La_{0.95}Ag_{0.05}MnO_3$  films,<sup>6</sup> when T < 80 K and  $\mu_0 H > 10$  T.

Figure 2 displays a typical temperature dependent AMR( $\theta$ ) evolution in  $\rho_{100}$  and  $\rho_{0\bar{1}1}$  under 13 T during cooling. The corresponding temperature dependent resistivities and AMR values are shown in Figs. 3(a) and 3(b), respectively. Here, the AMR value is obtained from the form of  $(\rho_{\parallel}-\rho_{\perp})/\rho_{av}$ , where  $\rho_{\parallel}$  and  $\rho_{\perp}$  are the resistivities with the current parallel ( $\theta$ =0°) and perpendicular ( $\theta$ =90°) to the magnetization, respectively. As shown in Fig. 3(a), the SCMO film shows a quicker increase of the resistivity upon decreasing temperature to  $T_{CO}$  due to the formation of COO,

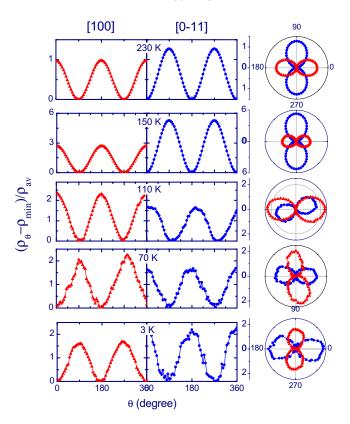


FIG. 2. (Color online) Linear (left) and polar (right) plots of temperature dependent evolution of AMR oscillations,  $(\rho_{\theta} - \rho_{\min})/\rho_{av}$ , along both [100] and  $[0\overline{1}1]$  directions under a field of 13 T.

where the collective JT distortion of the MnO<sub>6</sub> octahedron occurs and further evolves with decreasing temperature. In addition, the film exhibits a large transport anisotropy of  $\rho_{100}/\rho_{0\bar{1}1} \approx 10$  at room temperature due to the strain effect.<sup>14</sup> The anisotropic strain also leads to a significant different response of  $\rho_{100}$  and  $\rho_{0\bar{1}1}$  with respect to external magnetic fields, for example a giant CMR anisotropy of larger than 10<sup>3</sup>

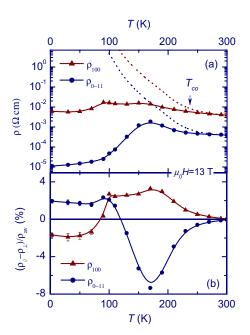


FIG. 3. (Color online) (a) The temperature dependent resistivities  $\rho_{100}$  and  $\rho_{0\bar{1}1}$  under 0 T (dotted line) and 13 T. (b) The temperature dependent AMR effect under 13 T along both [100] and  $[0\bar{1}1]$  directions.

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is obtained under a field of 13 T [Fig. 3(a)]. As shown in Fig. 3(b) and as mentioned previously, both AMR effects in  $\rho_{100}$ and  $\rho_{0\bar{1}1}$  appear at  $T=T_{CO}$ . Meanwhile, the AMR magnitudes increase fast with further decreasing temperature. However, AMR turns to decrease with decreasing temperature when T < 170 K, which is followed by the AMR( $\theta$ ) crossover at 120 K in  $\rho_{011}$  and near 80 K in  $\rho_{100}$ . Below 80 K, where "metallic" conducting behavior appears in both directions [Fig. 3(a)], the AMR exhibits little temperature dependence. Additionally, the value of  $\rho_{011}$  (with maximum of 7.3%) is essentially larger than that of  $\rho_{100}$  (with maximum of 3.3%) in the whole temperature range, which has also been observed in La<sub>2/3</sub>Ca<sub>1/3</sub>MnO<sub>3</sub> films.<sup>9</sup> The most striking feature in the figures is the occurrence of the distinct AMR decrease [Fig. 3(b)] once the magnetic-induced metallic conducting behavior begins to play a role [Fig. 3(a)]. The results clearly demonstrate that the magnetic-field-induced MIT also make a dramatic signature in AMR in COO manganites, besides the AMR peak in the vicinity of MIT for ferromagnetic manganites. In addition, the decrease of AMR magnitude and its sign change in the field-induced conducting region should indicate an opposite  $AMR(\theta)$  behavior between the COO state and the field-induced metallic one, while the latter exhibits a similar behavior as the ferromagnetic La<sub>0.95</sub>Ag<sub>0.05</sub>MnO<sub>3</sub>.<sup>6</sup>

Generally, the AMR mainly arises from the convoluted nature of the Fermi surface or the scattering anisotropies due to the spin-orbit interaction. Considering the disappearance of the  $H^2$ -dependent magnetoresistance and the small  $\omega_c \tau$ ( $\omega_c$  is the cyclotron frequency, and  $\tau$  the scatting time) in manganites,<sup>11</sup> the former contribution to the AMR should be excluded. Though more different from the conventional lowfield anisotropic magnetoresistance in ferromagnetic alloys, the AMR effect observed here resembles the ferromagnetic manganite AMR in that it also shows a close relation with JT distortion and MIT, which implies that the spin-orbital coupling may still play an important role behind it. Generally, the spin-orbital coupling is quantified by the Hamiltonian  $H_{SO} = \Delta_{SO} L \cdot S$ , where L and S are the orbital and magnetic moments, respectively, and  $\Delta_{SO}$  is the spin-orbit coupling parameter. For  $T > T_{CO}$ , the SCMO exhibits little JT distortion and the orbital angular momentum of  $t_{2g}^3$  and  $e_g^1$  configuration is completed quenched by the crystal field, i.e., L=0, thus no spin-orbit coupling.<sup>15</sup> The collective JT distortion of the MnO<sub>6</sub> octahedron accompanied with the COO formation leads to an unquenching of the 3d orbits, thus the appearance of spin-orbital coupling effects.<sup>13,15</sup> The evolution of collective JT distortion with decreasing temperatures could result in an increase of AMR below  $T_{\rm CO}$ . Meanwhile, the decrease of the AMR could be further ascribed to the suppression of JT distortion at field-induced MIT, which has also been observed in La<sub>0.69</sub>Ca<sub>0.31</sub>MnO<sub>3</sub> samples.<sup>8</sup> However, the crossover in AMR is a rather intriguing phenomenon and may be a much more complex process.

In the microscopic theory of AMR for ferromagnets,<sup>3,4</sup> the sign and the magnitude of AMR can be described by the formula:  $(\rho_{\parallel} - \rho_{\perp})/\rho_{av} = \gamma(\alpha - 1)$ , where  $\gamma$  is a spin-orbit coupling constant,  $\alpha$  is the ratio between spin  $\uparrow$  and spin  $\downarrow$ 

resistivities,  $\alpha = \rho_{\uparrow} / \rho_{\downarrow}$ . The positive AMR is generally regarded as a consequence of scattering of majority-spin (spin  $\uparrow$ ) electrons, while the minority-spin (spin  $\downarrow$ ) electrons are responsible for the negative AMR. Thus, the crossover in AMR may indicate that magnetic-destruction of the COO state results in severe changes in the electronic structure and a reversal in the spin polarization of dominating conductive electrons, footprint of which has also been found in La<sub>0.95</sub>Ag<sub>0.05</sub>MnO<sub>3</sub> films.<sup>6</sup> In addition, the essential larger AMR magnitude in  $\rho_{011}$  than that in  $\rho_{100}$  could indicate the stronger spin-orbit coupling along the  $[0\overline{1}1]$  direction than that along [100], which is consistent with the huger decrease in  $\rho_{0\bar{1}1}$  under magnetic field ( $\mu_0 H > 10$  T) and the results in La<sub>2/3</sub>Ca<sub>1/3</sub>MnO<sub>3</sub> films.<sup>9</sup> On the other hand, the AMR crossover could also be triggered by the reconstruction of orbital occupation accompanying the magnetic-field-induced MIT, which may be the case for the sign change in ultrathin  $Pr_{0.67}Sr_{0.33}MnO_3$  (Ref. 10) and  $La_{0.65}Ca_{0.35}MnO_3$  (Ref. 5) films. However, further investigations are needed to reveal the mechanism underlying the AMR crossover.

In summary, the AMR results exhibit a dramatic modification with temperature and magnetic field in charge-orbitalordered SCMO films. In particular, crossovers in AMR are observed due to the magnetic-destruction of charge/orbital ordering, which indicates a possible severe modification of the SCMO electronic structure or possible orbital reconstruction with the magnetic-field-induced MIT.

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- <sup>1</sup>E. Dagotto and Y. Tokura, MRS Bull. **33** 1037 (2008).
- <sup>2</sup>S. Cox, J. Singleton, R. D. McDonald, A. Migliori, and P. B. Littlewood, Nature Mater. 7, 25 (2008).
- <sup>3</sup>T. R. McGuire and R. I. Potter, IEEE Trans. Magn. 11, 1018 (1975).
- <sup>4</sup>I. A. Campbell and A. Fert, in *Ferromagnetic Materials*, edited by E. P. Wohlfarth (North-Holland, Amsterdam, 1982), Vol. 3, p. 747.
- <sup>5</sup>M. Egilmez, M. M. Saber, A. I. Mansour, R. Ma, K. H. Chow, and J. Jung, Appl. Phys. Lett. **93**, 182505 (2008).
- <sup>6</sup>I. C. Infante, V. Laukhin, F. Sanchez, J. Fontcuberta, O. Melnikov, O.Y. Gorbenko, and A. R. Kaul, J. Appl. Phys. **99**, 08C502 (2006).
- <sup>7</sup>J.-B. Yau, X. Hong, A. Posadas, C. H. Ahn, W. Gao, E. Altman, Y. Bason,
- L. Klein, M. Sidorov, and Z. Krivokapic, J. Appl. Phys. 102, 103901 (2007).
- <sup>8</sup>R. W. Li, H. B. Wang, X. W. Wang, X. Z. Yu, Y. Matsui, Z. H. Cheng, B. G. Shen, E. W. Plummer, and J. D. Zhang, Proc. Natl. Acad. Sci. U.S.A. **106**, 14224 (2009).
- <sup>9</sup>M. Bibes, V. Laukhin, S. Valencia, B. Martínez, J. Fontcuberta, O. Yu Gorbenko, A. R. Kaul, and J. L. Martínez, J. Phys.: Condens. Matter **17**, 2733 (2005).
- <sup>10</sup>Q. Li, H. S. Wang, Y. F. Hu, and E. Wertz, J. Appl. Phys. 87, 5573 (2000).
- <sup>11</sup>J. O'Donnell, J. N. Eckstein, and M. S. Rzchowski, Appl. Phys. Lett. 76, 218 (2000).
- <sup>12</sup>Z. Fang, I. V. Solovyev, and K. Terakura, Phys. Rev. Lett. **84**, 3169 (2000).
- <sup>13</sup>C. Aruta, G. Ghiringhelli, V. Bisogni, L. Braicovich, N. B. Brookes, A. Tebano, and G. Balestrino, Phys. Rev. B 80, 014431 (2009).
- <sup>14</sup>Y. Z. Chen, J. R. Sun, J. L. Zhao, J. Wang, B. G. Shen, and N. Pryds, J. Phys. Condens. Matter (to be published).
- <sup>15</sup>J. H. Song, J.-H. Park, J.-Y. Kim, B.-G. Park, Y. H. Jeong, H.-J. Noh, S.-J. Oh, H.-J. Lin, and C. T. Chen, Phys. Rev. B **72**, 060405(R) (2005).