

Decoupling of fluids and fluid-mobile elements during shallow subduction: Evidence from halogen-rich andesite melt inclusions from the Izu arc volcanic front

Susanne M. Straub

GEOMAR Forschungszentrum an der Christian-Albrechts Universität zu Kiel, Wischhofstrasse 1–3, 24148 Kiel, Federal Republic of Germany

Now at Lamont Doherty Earth Observatory, Columbia University, 61 Route 9W, Palisades, New York 10964, USA. (smstraub@ldeo.columbia.edu)

Graham D. Layne

Woods Hole Oceanographic Institution, MS 23, Clark 114A, Woods Hole, Massachusetts 02543, USA

[1] Very rare, halogen-rich andesite melt inclusions (HRA) in bytownitic plagioclase phenocrysts (An₈₉₋₉₀) from tephra fallout of the Izu arc volcanic front (Izu VF) provide new insights into the processes of fluid release from slab trenchward to the volcanic front in a cool subduction zone. These HRA are markedly enriched in Cl, F and Li - by factors of up to 8 (Cl, F) and 1.5 (Li) - but indistinguishable with respect to the fluid-mobile large-ion lithophile elements (LILE; K, Sr, Rb, Cs, Ba, Pb, U), rare earths (REE) or high field strength elements (HFSE) from the low-K tholeiitic magmas of the Izu VF. We suggest that the chemical signature of the HRA reflects the presence of a fluid in the mantle source that originated from the serpentinized mantle peridotite above the metacrust. This "wedge serpentinite" presumably formed by fluid infiltration beneath the forearc and was subsequently down-dragged with the slab to arc front depths. The combined evidence from the Izu VF (~110 km above slab) and the outer forearc serpentinite seamounts (~25 to 30 km above slab) suggests that the slab flux of B and Cl is highest beneath the forearc, and decreases with increasing slab depths. In contrast, the slab flux of Li is minor beneath the forearc, but increases with depth. Fluorine may behave similarly to Li, whereas the fluid-mobile LILE appear to be largely retained in the slab trenchward from the Izu VF. Consequently, the chemical signatures of both Izu trench sediments and basaltic rocks appear preserved until arc front depths.

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1. Introduction

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[2] During shallow subduction (trenchward from the volcanic front), the descending lithosphere releases large amounts of water and carbon dioxide [e.g., Kerrick and Connolly, 1998; Wallmann, 2001]. Since water is a primary transport agent for many fluid-mobile elements (e.g., Li, B, Cl, Cs, Pb), these elements might be lost from slab during early subduction as well, implying that the composition of the subducted material changes significantly prior to reaching arc front depths [e.g., Bebout et al., 1999; Class et al., 2000]. Unfortunately, the critical region for early slab devolatilization (<110 km depth) cannot be directly observed at active convergent margins since there is no concomitant volcanism. To date, the most comprehensive information on the early chemical evolution of the subducting slab derives from studies of uplifted, metamorphosed slab fragments [e.g., Bebout et al., 1999; Scambelluri et al., 2001]. In active settings, however, information on the composition of the shallow fluids can only be obtained from serpentinite seamounts, located \sim 20–25 km above slab on the outer trench slope [Fryer et al., 1999; Benton et al., 2001], or from the volcanic arc, located ~ 110 km above slab [e.g., Taylor and Nesbitt, 1998; Hochstaedter et al., 2001].

[3] In this study, we present data on the Cl, F and Li contents of both matrix glasses and plagioclase melt inclusions from fallout tephra of the Izu arc volcanic front (Izu VF). These tephra glasses are an excellent proxy to Izu VF magma compositions (e.g., S. M. Straub, G. D. Layne, A. Schmidt and C. H. Langmuir, The recycling of fluids and sediment melts in volcanic arcs, manuscript submitted to Geochemisty Geophysics Geosystems, 2003, hereinafter referred to as Straub et al., submitted manuscript, 2003). Cl, F and Li are enriched in the subducting slab, owing to contributions from seawater, sediments, and basaltic oceanic crust [e.g., Ryan and Langmuir, 1987; Decitre et al., 2002]. Since Cl, F and Li are also highly mobile in hydrothermal fluids [e.g., Schilling et al., 1978; Ryan and Langmuir, 1987; Seyfried and Ding, 1995], they should become readily entrained in fluids during early subduction. The data show that Cl, F and Li display a much wider range in composition than the other recycled, fluid-mobile large-ion lithophile elements (LILE; K, Sr, Rb, Cs, Ba, Pb, U) in the Izu VF magmas. In combination with existing data from the outer forearc seamounts, our results provide new insights into the sequential release of fluid-mobile elements during shallow subduction.

2. Geological Setting

[4] The evolution and structure of the intraoceanic Izu Bonin Mariana arc/backarc system (IBM; NW Pacific) has been described in detail elsewhere [Taylor, 1992; Hochstaedter et al., 2001]. In summary, volcanism in the Izu and Mariana arcs began at ~49 Ma (middle Eocene), following the westerly subduction of the Mesozoic (\sim 130 Ma) Pacific plate beneath the Philippine plate. In the Oligocene (\sim 31 Ma) the Izu arc split along-strike. During the Miocene (\sim 24 to 15 Ma), the Shikoku backarc basin formed. Volcanism at the Izu VF waned during backarc spreading, but rejuvenated \sim 15 Ma ago, and has since been vigorous. A new period of intraarc rifting was initiated at ~2.8 Ma immediately west of the Izu VF (Figure 1). The Quaternary Izu VF is constructed on ~ 20 km thick crust [Suheviro et al., 1996], and located about $\sim 100-110$ km above a well-defined Wadati-Benioff zone [Katsumata and Sykes, 1969; Wicks and Richards, 1993]. The subducting crust is Cretaceous in age, with the basaltic basement being similar in composition to the mid-ocean ridge basalts of the East Pacific Rise [Fisk and Kelley, 2002]. The sediment cover consists of \sim 400 m thick Mesozoic and Cenozoic pelagic clays $(\sim 39\%)$, arc-derived ash $(\sim 5\%)$ and chert, nannofossil chalk and marls (\sim 56%) [Plank et al., 2000].

[5] The Izu VF is an informative setting to study, since it has a highly depleted subarc mantle wedge [*Taylor and Nesbitt*, 1998; *Langmuir et al.*, 2003] with no evidence of partial slab melts in the mantle source [*Taylor and Nesbitt*, 1998; *Hochstaedter et al.*, 2000, 2001]. Against this background, the chemistry of slab-derived fluids is clearly recog-

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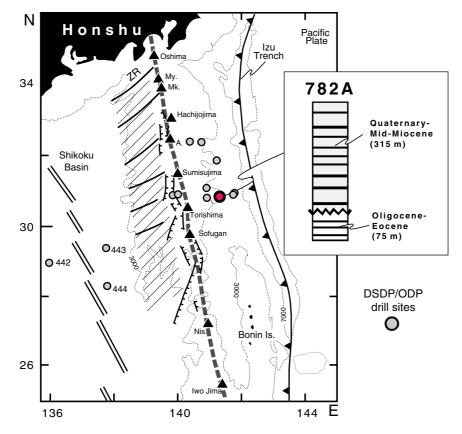


Figure 1. Geological setting of the Izu-Bonin arc/backarc system with DSDP and ODP drill sites. ODP Site 782A is indicated. From East to West: Pacific Plate, Izu Trench, Izu Forearc, Izu Volcanic Front (stippled line), rift (grabens), rear arc (hatched) and the inactive Shikoku Backarc Basin. Depth contours are in meters. ZR, Zenisu Ridge (subaerial volcanoes Nijiima and Kozushima are not shown); Subaerial volcanoes (black triangles) unless labeled are My, Miyakejima; Mk, Mikurajima; A, Aogashima; Nis, Nishinoshima. BI, Bonin islands (uplifted Eocene forearc basement). Diagonal solid lines in rear arc region denote rear-arc volcanic chains. Stratigraphy of ODP Site 782A after *Xu and Wise* [1992].

nizable. Moreover, the Izu Bonin is a classic example of a cool subduction zone, in which most of the fluid-mobile elements may be retained in the slab until arc front depths [e.g., Bebout et al., 1999]. Comprehensive studies of major and trace elements, and radiogenic isotopes (Sr, Nd, Pb), have shown that (1) the composition of the mantle and slab sources have been homogenous during the last 15 million years and that (2) the Neogene Izu VF magma source is infiltrated by fluid components from both the subducting sediment (\sim 3.5% of total fluid) and the subducting basaltic crust ($\sim 96.5\%$ of total fluid) [e.g., Taylor and Nesbitt, 1998; Hochstaedter et al., 2001; Straub et al., submitted manuscript, 2003]. Hydrated serpentinized peridotite, that provides additional fluid sources, is considered to be present either above [Straub and

Layne, 2002], or below, the slab [*Peacock*, 2001] (Figure 1). For these reasons, the Izu VF volcanics are ideally suited for an investigation of the provenance and pathways of slab-derived fluids.

3. Samples and Methods

[6] Halogen and Li abundances of Izu VF magmas were determined from matrix shards and melt inclusions from fallout tephra. The tephra was recovered from sediments at ODP site 782A, located \sim 120 km trenchward of the Quaternary Izu VF (Figure 1). Site 782A contains numerous fallout tephras, which provide a highly resolved temporal record (one event per 0.12 Ma on average) of Izu VF volcanism. Most of the fallout tephra investigated is of Neogene age (0.55–14.2 Ma); only a

single Oligocene fallout tephra (29.6 Ma) was included. A comprehensive data set exists for the 782A fallout tephra. This includes bulk analyses of Sr, Nd and Pb isotopes [Schmidt, 2001], as well as major and trace elements, and B isotopes, by microanalytical methods on single glass shards [Straub and Layne, 2002, 2003; Straub et al., submitted manuscript, 2003]. A representative sub-set of the glasses has been analyzed for trace elements by laser ablation-ICP-MS (Straub et al., submitted manuscript, 2003). The Cl, F and Li determinations were performed, by electron microprobe (Cl) and ion microprobe (F, Li), on the same suite of matrix glasses and melt inclusions previously analyzed for major elements (electron microprobe) and B isotopes (ion microprobe). The data relevant to this paper, and a short summary of the analytical procedures, is provided in Tables 1-4 for easy reference. A detailed description of the analytical procedures is given in Straub et al. (submitted manuscript, 2003) (major elements, B, Li), Straub and Layne [2002] (B isotopes) and Straub and Layne [2003] (Cl, F). None of these previous studies, however, addresses in detail the question of the origin of the halogen-rich andesite inclusions, and the potential implications for shallow slab devolatilization, which are the focus of this paper.

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[7] The glasses from the Site 782A fallout tephras are mostly low-K basalts to rhyolites that are very similar in major element, trace element and isotope composition to the Quaternary Izu VF lavas. The single Oligocene tephra studied (sample 108), which was heavily disturbed by the drilling, has a subordinate population of medium-K glasses, which is not further considered here. With the exception of a single high-MgO Izu lava from the island of Hachijojima (MgO = 8.23 wt%; Mg# = 62), the tephra glasses and Izu arc front lavas have similar maximum MgO (~6 wt%) and Mg# numbers [\sim 60; where Mg# = molar ratio Mg/(Mg + Fe^{2+}]. Basaltic through rhyolite lava and tephra series are indistinguishable in their Sr-, Nd-, Pb- and B isotope ratios, demonstrating that the series are co-genetic and originate from the same sources as the Izu VF lavas [Straub et al., submitted manuscript, 2002; Straub and Layne, 2002]. Further, no

evidence for post-eruptive alteration has been found. The glasses investigated are all clear, colorless to brown glass shards, or melt inclusions, without any telltale signs of birefringence under crossed polarizers or cloudiness indicative of incipient alteration. The consistency of glass and melt inclusion compositions, the apparently magmatic systematics of alteration-sensitive elements like Li and B (Figure 2), and the combination of high B $(\sim 10-40 \text{ ppm})$ with δ^{11} B of +5% to +12% in the Izu glasses, all militate against the presence of any syn- or post-eruptive seawater alteration [Straub and Layne, 2002, 2003]. Thus the high B contents and high- $\delta^{11}B$ values of the normal-group glasses and the HRA demonstrate the presence of a slab component in the Izu VF magma source and imply a slab origin for other fluid-mobile elements as well.

4. Results

4.1. Two Component Melts: Normal-Group Glasses and HRA

[8] On the basis of their Cl, F and Li contents, the Izu glasses can be divided into two groups (Figure 2). The majority of glasses and melt inclusions (termed "normal-group" glasses), contain 0.05-0.40 wt% Cl, 70-400 ppm F and 4-10 ppm Li. The normalgroup glasses comprise matrix glasses, and melt inclusions in plagioclase, that range from basalt to rhyolite in composition. In this group, Cl, F and Li behave similarly to the fluid-mobile LILE. This means that Cl, F and Li vary by about a factor of two at a given MgO, and increase by a factor of three with decreasing MgO (Figure 2; in the context of this paper the term "fluid-mobile LILE" refers to Cs, Rb, Ba, U, K, B, Sr and Pb). In contrast, the second group of glasses ("halogen-rich inclusions"; = HRA) consists of rare andesite melt inclusions in plagioclase (Figures 2-4). The HRA are selectively enriched in the halogens Cl (0.70–0.86 wt%), and F (700-900 ppm) by about a factor of 8. The abundance of Li (\sim 15 ppm) is increased by about a factor of 1.5, excepting a single HRA melt inclusion that has a 'normal' Li content (Figure 2). The HRA were found in only three out of forty of the 782A tephras investigated. In these three tephras, the HRA are readily identified by their characteristic Cl enrich-

Table 1 (Representative Sample). Composition of Matrix Glasses From 782A Fallout Tephra ^a [The full Table 1 is available in the HTML version of this article at http://www.g-cubed.org.]	presei v.g-cu	htative bed.or	e Sam g.]	ple).	Comp	ositio	n of N	latrix (Glasse	s Fron	1 782A	Fallou	ıt Teph	ura ^a [T.	he full	Table	1 is a	vailabl	e in th	e HTN	AL ve	rsion c	of this a	article
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Glass No. ^b	£3	a1	g2	al	fl	19-7	19 - 3	19-8	19-2	19 - 9	19-11	£	d5	b1	b2	d2	IJ	b1	с2	c1	34-2	34-3	34-3 34-10	34-7
Shard/Inc ^c	00	00	03	03	යය	00	00	83	යය	යය	ad	83	යය	00	00	00	00	00	යය	යය	80	00	යය	ad
SiO ₂ TiO ₂ Al2O3 FeO* FeO* MnO MnO MgO CaO Na ₂ O M ₂ O So ⁵ P ₂ O ₅ H ₂ O (SD) H ₂ O (SD)	54.96 0.96 14.58 11.39 0.23 3.260 0.10 0.10 0.10 0.10 0.10 0.10 0.10 0.10 0.10	54.96 56.35 0.96 1.01 14.58 14.85 11.39 11.36 0.23 0.23 3.55 3.55 3.60 3.55 2.96 2.81 0.36 0.39 0.10 0.11 97.4 98.6 2.8 1.4		57.30 58.69 67.31 1.00 0.95 0.58 14.87 14.53 13.14 14.87 14.53 13.14 10.93 10.78 7.04 0.22 0.23 0.23 3.50 2.76 1.15 7.92 7.24 4.68 7.93 0.47 0.72 0.33 0.47 0.72 0.30 0.47 0.72 0.10 0.11 0.14 0.90 98.7 98.5 1.0 1.3 1.5	67.31 0.58 13.14 7.04 1.15 3.74 0.72 0.14 98.5 1.5	$\begin{array}{c} 52.74\\ 1.11\\ 14.90\\ 12.30\\ 9.18\\ 9.18\\ 9.18\\ 0.21\\ 0.12\\ 0.12\\ 2.3\\ 2.3\\ 2.3\end{array}$	$\begin{array}{c} 52.80\\ 1.10\\ 15.07\\ 15.07\\ 15.07\\ 12.17\\ 0.23\\ 9.03\\ 2.59\\ 0.13\\ 97.7\\ 2.5\\ 2.5\\ 2.5\\ 2.5\\ 2.5\\ 2.5\\ 2.5\\ 2.5$	$\begin{array}{c} 53.06\\ 1.09\\ 15.11\\ 15.11\\ 15.11\\ 12.20\\ 9.16\\ 0.12\\ 0.12\\ 0.12\\ 0.12\\ 0.12\\ 0.12\\ 0.12\\ 0.12\\ 0.12\end{array}$	53.16 1.10 15.05 15.05 15.05 8.90 8.90 0.12 98.1 1.9 2.3 2.3	$\begin{array}{c} 53.5\\ 1.0\\ 15.0\\ 12.6\\ 8.8\\ 8.8\\ 8.8\\ 98.2\\ 1.8\\ 0.1\\ 1.8\\ 2.5\\ 2.5\\ 2.5\\ 2.5\\ 1.8\\ 1.8\\ 1.8\\ 1.8\\ 1.8\\ 1.8\\ 1.8\\ 1.8$	Abundances in wt% 7 1.11 1.00 0.5 7 1.11 1.00 0.5 7 1.11 1.00 0.5 7 1.11 1.00 0.5 7 1.11 1.00 0.5 10 15.11 15.24 15.1 11.60 10.91 10.0 10 12 11.66 10.91 10 3 3.76 3.82 3.2 4 8.46 8.28 8.2 5 0.26 0.38 0.1 1 0.12 0.13 0.13 0.1 3 98.6 97.5 97 97 3 98.6 97.5 97 97 3 1.4 2.6 2.6 2 5 2.4 2.7 2.7 2.7	es in wi 54.76 15.24 15.24 10.91 10.91 10.91 3.82 8.28 8.28 8.28 0.13 97.5 2.6 2.6	t% 5.518 5.518 15.28 10.75 10.75 3.74 8.36 2.74 0.30 0.16 97.6 2.4	$\begin{array}{c} 55.20\\ 0.96\\ 115.39\\ 111.11\\ 0.21\\ 0.37\\ 0.37\\ 0.14\\ 0.14\\ 1.8\end{array}$	$\begin{array}{c} 55.29\\ 0.94\\ 15.26\\ 10.86\\ 3.79\\ 3.79\\ 0.37\\ 0.37\\ 0.37\\ 2.2\\ 2.2\end{array}$	51.76 11.26 11.26 13.91 0.25 0.25 0.34 0.34 0.34 0.34 0.34 0.11 0.34 0.11 1.4 2.1	$\begin{array}{c} 552.27\\ 1.25\\ 14.69\\ 13.42\\ 0.19\\ 0.10\\ 0.39\\ 0.10\\ 0.39\\ 1.8\\ 1.8\\ 1.8\end{array}$	$\begin{array}{c} 55.28\\ 11.26\\ 12.63\\ 12.96\\ 0.19\\ 0.19\\ 0.11\\ 0.11\\ 0.11\\ 0.11\\ 2.3\\ 2.3\end{array}$	552.45 11.28 11.28 114.88 113.23 0.26 0.26 0.26 0.33 0.36 0.36 0.36 0.36 0.36 0.36 0.3	$\begin{array}{c} 552.49\\ 11.26\\ 11.26\\ 13.23\\ 0.22\\ 2.30\\ 0.40\\ 0.10\\ 0.10\\ 0.10\\ 1.3\end{array}$	$\begin{array}{c} 57.05\\ 11.00\\ 14.92\\ 10.65\\ 0.21\\ 3.45\\ 7.92\\ 2.86\\ 0.12\\ 98.4\\ 1.6\end{array}$	57.32 11.02 11.02 10.55 0.24 0.24 0.24 0.24 0.24 0.24 0.24 0.24	$\begin{array}{c} 57.42\\ 0.57.42\\ 14.87\\ 10.59\\ 10.59\\ 10.53\\ 3.22\\ 7.70\\ 0.34\\ 0.34\\ 0.12\\ 1.7\end{array}$	$\begin{array}{c} 57.52\\ 15.10\\ 15.10\\ 15.10\\ 10.89\\ 0.20\\ 3.51\\ 7.98\\ 0.38\\ 0.38\\ 0.12\\ 0.99.4\\ 0.6\end{array}$
CI LI BB 8 ¹¹ B	1760 164 6.90 0.41 18.7	1937 186 6.85 0.49 19.9 -	1907 196 7.51 0.42 20.7 -	2297 262 7.92 0.47 22.3	3310 287 8.42 0.57 36.9 -	1160 192 4.02 0.44 9.7 12.0	1190 211 4.01 0.50 9.5 10.4	1227 196 3.85 0.43 9.7	1108 233 3.91 0.57 9.6 11.0	Ab 1222 191 4.04 0.41 10.2 -	Abundances in ppm 2 1248 1657 11 2 236 203 19 4 4.35 6.94 7. 1 0.45 0.49 0. 2 11.3 17.2 17	es in p 1657 203 6.94 0.49 17.2 -	pm 1178 195 7.10 0.50 17.6 -	1563 200 6.61 0.53 17.0	1565 215 6.93 0.53 17.5	1093 149 7.29 0.52 23.9 7.9	1050 158 6.99 0.55 23.7	1077 151 6.65 0.48 7.0	1118 159 7.54 0.57 26.3	1047 157 7.00 0.52 2.4.0 7.6	1423 204 6.31 0.42 18.4 -	1235 189 6.13 0.43 17.8 -	1253 196 6.43 0.45 19.0	1338 233 6.08 0.48 17.9
^a Major element oxides, Cl and summation deficit (=100 minus sum of oxides; denoted by H ₂ O (SD)) by electron microprobe. H ₂ O (SIMS) and other trace element data by ion microprobe (SIMS)(see also <i>Straub and Layne</i> [2002] and Straub et al. (submitted manuscript, 2003) for detailed description of analytical methods. Major elements measured by electron microprobe Cameca SX50 at the Geoma and are normalized to the same chip of basalt glass JDF-D2. Precisions (one standard deviation) are typically (in wt%). SiO ₂ (0.26), TiO ₂ (0.02), Al ₂ O ₃ (0.11), P ₂ O ₅ (0.01). P ₂ O ₅ (0.01). The elements B, Li and F were measured by Cameca IMS 3f ion microprobe at the Woods Hole Oceanographic Institution. The reproducibility of B is better than 10% at B < 10 ppm and better than 5% at higher B contents. The reproducibility is $\pm 0.7\%$ at the given level of abundances. Column heads are as follows: sample codes, ODP Sample Cameca IMS 1270 ion microprobe at Woods Hole Oceanographic Institution. A typical reproducibility is $\pm 0.7\%$ at the given level of abundances. Column heads are as follows: sample codes, ODP Sample Number, and Age given in Ma. The "Sample code" (e.g. Sample 11) is as used in text. ^b 'Glass No.'' identifies individual glass or melt inclusions in polished section. ^c "ShardInc" specifies glass type analyses (g, matrix glass, cpx, opx, ph, melt inclusions in clinopyroxene, orthopyroxene and plagioclase). ^d Numeric ages based on clacreous biostratigraphy of <i>Xu and Wise</i> [1992].	nent oxi d Layne lized to 1 0.04), K < 10 pr < 10 pr < 270 ion Vge give " specif ges base	des, Cl [2002] the sam pm and micropt min Ma iffes inc iffes inc d on ca	and surr and Stra e chip o l]), P ₂ O ₅ better th obe at V a. The " irvidual is type a s type a	mation f basalt f basalt f basalt (0.01). nan 5% Noods F Voods F Voods F voods f ualyzed s biostra	deficit (l. (subm: glass JI The ele at hight Jole Occ code" t melt in 1 (g, ma atigraph	(=100 n itted ma DF-D2. ments E er B cor eanograp (e.g. Sa nclusion nclusion y of Xu	innus su muscrip Precisio 3, Li and tents. T phic Ins mple 1 is, cpx, is, cpx,	m of ox t, 2003) ns (one 1 F were the repri- titution. 1) is as ished s(opx, pl, se [199]	ides; de for dets standar e measu oducibil A typic used in used in melt ir melt ir	noted by ailed des d deviat red by C ity of L ity of L ity of L text. text.	y H ₂ O (S scription (control are t cameca II i is 4% (ducibility s in clino	D)) by e of analyt of analyt MS 3f ic relative) is ±0.7 pyroxer	electron ical met (in wt% n micro n micro % at the % at the te, ortho	micropr hods. [N] (a): SiO ₂ (b): SiO ₂ probe at produci i: given 1 pyroxer	obe. H_2^{t} fajor ele (0.26), t the Wo bility F evel of i evel of i	 SIM: SIM: ments n m	 and c neasured assured of relation ces. Co se). 	ther tradition of the t	ce eleme tron mi), FeO* c Institu on isoto ads are <i>i</i> ads are <i>i</i>	ant data croprob (0.11), 1 trion. Th ppe anal ppe anal as follow	by ion i e Came MnO (0 ne repro yses we vs: samj	micropr ca SX5(.02), M, ducibili pie code pie code	bbe (SIM at the C gO (0.06 y of B i rmed us s, ODP ((IS)(see icomar), CaO s better ing the Sample

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Shard/ Inc	ac	ac	ac	00	ad	cpx	cpx	cpx	xdo	pl	pl	pl	pl	pl	pl	pl	pl	pl	pl	pl	pl	pl	pl	pl	pl	pl	pl	pl
SiO_2	70.02	73.14	72.83	59.65	58.40	52.77	52.35	52.98	70.96	70.38	68.88	70.42	70.22 7	71.88 7	71.56 7	73.60 7	73.40	68.89	70.24	72.01	68.32	72.13	72.14 7	71.21 6	64.44 6	69.45 6	68.36	72.57
TiO_2	0.50			0.71	0.78	0.60		0.59	0.35		0.76	0.51																0.39
Al_2O_3	12.35		12.18	15.06		16.11			11.92	12.52	12.52	12.75	12.78 1		12.75 1		12.05	12.24		12.54				13.20 1		12.20 1	12.25	1.83
FeO*	3.73	3.36		8.58		9.18	9.19	9.11	3.14	4.11	5.38	3.75					2.66	3.51		3.37		2.65	3.13					2.61
MnO	0.11	0.12	0.12	0.17	0.19	0.20			0.13	0.12	0.18	0.14					0.08	0.15		0.14								0.10
MgO	0.52	0.38	0.32	2.79		5.97			0.28	0.62	0.95	0.53					0.44	0.49		0.51							0.66	0.32
CaO	2.74	2.54	2.22	6.68		11.25			2.21	2.84	3.13	2.87	2.92			2.15	2.30	2.56		2.53								2.13
Na_2O	4.01	3.95	4.06	3.57	3.41	2.06	2.11	2.06	3.92	4.45	4.50	3.61				3.71	3.84	3.57		3.76								4.12
K_2O	1.19	1.17	1.19	0.62	0.62	0.50	0.55	0.62	1.13	1.17	1.09	1.14				1.25	1.32	1.05		1.20								1.27
P_2O_5	0.07	0.09	0.06	0.14	0.15	0.15	0.15	0.12	0.0	0.13	0.14	0.12				0.07	0.08	0.10		0.07								0.01
Total	95.1	97.9	96.6	97.9	98.2	98.7	98.4	98.5	94.1	97.0	97.5	95.8		96.4	96.6	96.0	96.5	93.0	96.0	9.96	95.4			96.5	94.3	96.3	94.8	95.4
H_2O	4.9	2.1	3.4	2.1	1.8	1.3	1.6	1.5	5.9	3.0	2.5	4.2	3.9			4.0	3.5	7.0		3.4								4.6
(SD)																												
H_2O						1.6	1.6	1.4																				
(SIMS)																												
												A	Abundances in ppm	ices in	uidd													
CI	2532			1553	1517	1172	1192	1232	2143	3482	4933	3333	3518 3	3128		2350	_	2920	3289	2806	3926	2478		2479 5	5857 5			2942
Ч	445			299	272	214	214	189	433	607	712	572			511		388	589		487	666	460						413
Li	12.68		12.6	9.1	8.7	4.7	5.5	5.0	11.8	17.2	17.7	14.7		13.7	14.0	13.2	13.7	15.6	14.3	15.1	15.1	13.2	13.6	14.8	15.0	15.8	12.6	14.5
Be	0.61		0.62	0.54	0.53	0.47	0.45	0.46	0.68	0.63	0.57	0.56			0.62	0.59	0.64	0.63		0.59	0.57	0.62						0.59
В	62.7		64	35	31	17	18	18	62	83	48	61	58		74	99	62	62		70	50	60						65
$\delta^{11}B$	·	,	·	·	,	·	ı	,	·	,	·	ı	9.8	ı	11.2		ı	9.2		10.9	ı	9.5		9.2	ı	ı		
^a This table lists "mixed rhyolite" inclusions and basaltic inclusions only. Halogen- inclusions in single phenocrysts. For abbreviations and analytical methods see Table 1	^a This table lists "mixed rhyolite" inclusions and basalite inc lusions in single phenocrysts. For abbreviations and analytica	lists "r ngle pl	nixed 1 henocry	hyolite sts. Fo	" inclu r abbre	sviation	and ba:	saltic ii analvtic	nclusion cal met	ns only hods se	: Halog	cen-rich	lusions only. Halogen-rich andesite inclusions (HRA) from the same sample are listed in Table 3. Capital letters identify multiple melt I methods see Table 1.	te inch	usions	(HRA)	from	the san	ie sam	ple are	listed	in Tabl	e 3. C ²	pital le	tters ic	lentify	multip	le melt
					MOON I	IO INTL A		mm	10111 mo		1000 T 000	;																

Table 2. Composition of Matrix Shards and Melt Inclusions in Fallout Tephra Sample 3 (=782A-2H-4-113-114)^a

Geochemistry Geophysics Geosystems

0.11 0.09 0.14 0.14 97.61 98.93 96.02 96.17		04.16 00.16	94.92 97.33 97.46	95.18 94.92 97.33 97.46	93.48 95.18 94.92 97.33 97.46	95.36 93.48 95.18 94.92 97.33 97.46
	2.48 1.82 2.87 2.36	2.67 2.54 2.48 1.82 3.02 2.92 2.87 2.36	5.08 2.67 2.54 2.48 1.82 4.79 3.02 2.92 2.87 2.36	2.67 2.54 2.48 1.82 3.02 2.92 2.87 2.36	4.82 5.08 2.67 2.54 2.48 1.82 4.53 4.79 3.02 2.92 2.87 2.36	6.52 4.82 5.08 2.67 2.54 2.48 1.82 5.84 4.53 4.79 3.02 2.92 2.87 2.36
Abundances in ppm 933 838 1208 1412 82 100 152 163 5.5 6.0 6.7 6.5 0.44 0.49 0.48 0.42 13 9 18 16	528 708 84 81 4.9 4.4 0.48 0.43 9 7 5.5 70	503 608 528 708 102 89 84 81 7.7 7.6 4.9 4.4 0.45 0.37 0.48 0.43 11 11 9 7 6.4 - 5.5 7.0	6617 503 608 528 708 830 102 89 84 81 14.9 7.7 7.6 4.9 4.4 0.50 0.45 0.37 0.48 0.43 32 11 11 9 7 9.9 6.4 - 5.5 7.0	503 608 528 708 102 89 84 81 7.7 7.6 4.9 4.4 0.45 0.37 0.48 0.43 11 11 9 7 6.4 - 5.5 7.0	6687 6617 503 608 528 708 906 830 102 89 84 81 14.4 14.9 7.7 7.6 4.9 4.4 0.48 0.50 0.45 0.37 0.48 0.43 32 32 11 11 9 7 9.9 9.9 6.4 - 5.5 7.0	7292 6687 6617 503 608 528 708 779 906 830 102 89 84 81 9.9 14.4 14.9 7.7 7.6 4.9 4.4 0.51 0.48 0.50 0.45 0.37 0.48 0.43 40 32 32 11 11 9 7 - 9.9 9.9 6.4 - 5.5 7.0

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n^{2} 1 1 1 1 1 1 1 1 3 3 3 3 4 3 4<	2	Sample 3, /A, 0.55 Ma	Sample 3, 10A, 0.55 Ma	sample 3, 1015, 0.55 Ma	Sample 3, 11, 0.55 Ma	Sample 19, 2.84 Ma	Sample 29, 3.78 Ma	8.59 Ma	53mple 09, 10.3 Ma	Sample /1, 10.45 Ma	Sample 90, 13.35 Ma	sample 94, 14.2 Ma
inc inc inc inc inc inc inc inc 7.6 inc 1.07 7.6 $2.0.5$ 1.07 1.07 7.6 $2.0.5$ 1.07 1.07 0.29 0.73 0.46 0.45 0.23 0.54 0.45 0.45 0.23 0.54 0.45 0.45 0.211 0.45 0.45 0.211 1.05 0.45 0.211 1.05 0.45 0.211 0.932 0.45 0.211 0.932 0.45 0.211 0.932 0.45 0.212 0.46 1.97 0.35 0.46 1.97 0.35 0.90 1.95 0.90 1.97 0.599 $1.1.87$ 7.77 1.27 0.59 1.018 7.77 0.59 1.018 7.77 0.59 0.90 1.97 0.59 0.90 0.54 0.64 7.77 0.59 0.90 1.95 0.90 1.97 0.68 0.48 0.59 0.97 0.68 0.64 1.127 0.59 0.97 0.59 1.018 3.24 0.95 3.79 0.97 0.571 0.682 3.79 0.97 0.571 0.58 1.372 0.59 1.127 0.58 0.97 0.73 3.24 0.97 0.59 0.97 0.73 3.24 0.97 0.73 3.24 0.97 0.73 3.24 0.97 0.73 3.24 0.97 0.97 0.73 3.24 0.97 0.97 0.73 3.24 0.97 0.97 0.73 3.24 0.97 0.97 0.73 3.24 0.97 0.97 0.73 3.24 0.97 0.73 3.24 0.97 0.97 0.73 3.24 0.97 0.73 3.24 0.97 0.73 3.24 0.97 0.97 0.73 3.24 0.97 0.73 3.24 0.97 0.97 0.73 3.24 0.99 0.97 0.97 0.73 3.24 0.99 0.97 0.97 0.73 3.24 0.99 0.97 0.97 0.73 3.24 0.99 0.97 0.97 0.73 3.24 0.99 0.97 0.73 3.24 0.99 0.97 0.97 0.73 3.24 0.99 0.97 0.97 0.73 3.24 0.99 0.97 0.97 0.73 0.58 1.33 0.99 0.97 0.97 0.97 0.98 0.97 0.97 0.58 0.98 0.98 0.98 0.95 0.96 0.96 0.95 0.96 0.95 0.96 0.95 0.96 0.96 0.95 0.96 0.96 0.95 0.96 0.96 0.95 0.96 0.95 0.96 0.95 0.96 0.96 0.95 0.96 0.95 0.96	n ^b	1	1	1	1	3	3	3	3	4	3	4
0.78 1.62 1.07 7.6 20.5 15.3 97 20.5 15.3 97 0.29 0.73 0.46 0.29 0.73 0.45 0.45 0.29 0.73 0.46 0.45 0.23 0.54 0.45 0.45 0.23 0.54 0.607 1055 $0.43)^c$ 0.96 1.05 0.73 0.96 1.055 0.73 0.96 1.07 7.13 7.46 5.09 7.13 7.43 10932 7.13 13.43 11.87 7.13 7.46 5.09 1.30 1.95 1.07 7.77 13.43 11.87 7.77 13.43 11.87 7.77 13.43 11.97 3.09 1.95 0.64 1.777 12.22 10.94 3.09 0.59 0.66 0.68		inc	inc	inc	inc	60	00	60	ac	60	00	00
0.78 1.62 1.07 7.6 20.5 15.3 97 20.5 15.3 97 0.29 0.73 0.46 0.23 0.54 0.45 0.45 0.23 0.54 0.45 0.45 0.23 0.54 0.45 0.45 0.23 0.54 0.45 0.45 0.23 0.54 0.45 0.07 0.11 10353 10353 1.05 0.13 10353 10932 5.99 11.30 1.3.43 11.87 -0.07 7.13 13.43 11.87 -0.07 7.13 13.43 11.87 -0.07 3.80 1.95 90 -0.64 1.27 9.59 1.97 90 5.48 0.59 1.065 1.97 1.27 9.255 1.074 5.99 1.27 1.33.3 1.01.8 -1.12 3.09 1.255						Abundances i	in wt%					
7.6 20.5 15.3 97 202 187 0.29 0.73 0.46 0.23 0.54 0.45 0.23 0.54 0.45 0.23 0.54 0.45 0.23 0.54 0.45 0.23 0.54 0.45 0.23 0.54 0.45 0.11 <0.07 0.45 0.07 10353 0.06 1.0353 10353 1.093 5.40 1.05 1.093 7.13 13.43 11.87 7.13 13.43 11.87 7.13 1.95 1.97 90 3.80 1.95 127 95 90 548 0.59 10.18 7.77 1.222 10.18 7.77 1.2.22 10.18 3.09 1.38.3 2.57 0.6021 2.551 2.625 5.5.6 1.38.3 2.57 <td>C_{S}</td> <td>0.78</td> <td>1.62</td> <td>1.07</td> <td>1.90</td> <td>0.18</td> <td>0.51</td> <td>0.48</td> <td>0.60</td> <td>0.46</td> <td>1.45</td> <td>0.63</td>	C_{S}	0.78	1.62	1.07	1.90	0.18	0.51	0.48	0.60	0.46	1.45	0.63
97 202 187 0.29 0.73 0.46 0.23 0.54 0.45 0.23 0.54 0.45 0.23 0.54 0.45 0.143) ^c 0.96 1.05 0.143) ^c 0.96 1.05 0.11 < 0.11 < 0.07 $(0.43)^c$ 0.96 1.05 $(0.43)^c$ 0.96 1.05 $(0.07)^{-13}$ 10353 10932 2.81 13.43 11.87 7.13 13.43 11.87 7.13 13.43 11.87 7.13 13.43 11.87 7.13 1.3.43 11.87 7.77 9.5 90 5.48 0.195 1.97 127 9.59 10.18 7.77 1.2.22 10.18 7.77 1.2.22 10.18 3.09 1.3.8.3 1.01.18 3.09 0.95 0.662	Rb	7.6	20.5	15.3	20.1	2.1	4.9	4.7	8.0	5.7	16.5	7.8
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Ba	97	202	187	182	34	55	71	132	80	176	84
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Th	0.29	0.73	0.46	0.33	0.10	0.11	0.14	0.36	0.25	0.52	0.28
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	D	0.23	0.54	0.45	0.46	0.08	0.10	0.09	0.27	0.18	0.32	0.16
n.d. <0.11 <0.07 6007 10353 10932 2.81 5.40 4.61 7.13 13.43 11.87 3.80 7.46 5.99 1.30 1.95 1.97 3.80 7.46 5.99 1.27 95 90 548 314 338 0.48 0.59 10.94 7.77 12.22 10.18 7.77 12.22 10.18 7.77 12.22 10.94 3.09 4.56 2.72 5.66 138.3 100.4 5.77 0.68 0.73 3.79 0.97 3.24 0.68 0.97 2.72 3.79 4.56 2.72 3.79 4.67 3.24 0.68 0.97 0.73 3.79 0.97 3.24 0.61 0.97 0.73 3.79 0.97 0.73 <td>Nb</td> <td>$(0.43)^{c}$</td> <td>0.96</td> <td>1.05</td> <td>1.57</td> <td>0.25</td> <td>0.27</td> <td>0.31</td> <td>0.69</td> <td>0.49</td> <td>0.55</td> <td>0.30</td>	Nb	$(0.43)^{c}$	0.96	1.05	1.57	0.25	0.27	0.31	0.69	0.49	0.55	0.30
600710353109322.81 5.40 4.61 7.13 13.43 11.87 7.13 13.43 11.87 3.80 7.46 5.99 1.30 1.95 1.97 1.27 95 90 548 314 338 0.48 0.59 10.18 7.77 12.22 10.18 7.77 12.22 10.18 7.77 12.22 100.44 7.77 12.22 100.4 3.09 4.56 2.72 55.6 138.3 109.4 55.6 138.3 109.4 5.77 0.68 2.72 0.68 0.97 3.24 0.68 0.97 3.24 0.68 0.97 3.24 0.68 0.97 3.24 0.68 0.97 3.24 0.68 0.97 3.24 0.68 0.97 3.24 0.68 0.97 3.24 0.68 0.97 3.24 0.68 0.97 0.73 3.79 4.28 4.17 0.80 0.97 0.73 3.79 4.28 4.17 0.68 0.87 0.58 0.98 0.87 0.58 1.33 0.86 0.65 1.33 0.86 0.65 1.33 0.87 0.65 1.33 0.87 0.58 0.99 0.87 0.65 0.99 0.87 <t< td=""><td>Та</td><td>n.d.</td><td><0.11</td><td><0.07</td><td><0.15</td><td>0.02</td><td>0.02</td><td>0.02</td><td>0.06</td><td>0.05</td><td>0.04</td><td>0.03</td></t<>	Та	n.d.	<0.11	<0.07	<0.15	0.02	0.02	0.02	0.06	0.05	0.04	0.03
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	K	6007	10353	10932	9597	1769	3119	2799	5668	3399	9021	4465
7.13 13.43 11.87 3.80 7.46 5.99 1.30 1.95 90 548 314 338 648 0.59 1.97 7.77 95 90 548 314 338 601 7.77 12.22 10.18 7.77 12.22 10.18 3.36 7.77 12.22 10.18 3.24 7.77 12.22 10.94 2.72 55.6 138.3 109.4 2.72 55.6 138.3 109.4 2.72 6021 2.551 2.625 2.72 3.79 4.56 2.77 6.82 3.10 1.33 1.12 3.24 0.80 0.97 3.24 0.73 3.79 4.36 3.24 0.73 3.79 4.36 3.24 0.73 3.79 0.97 0.73 5.57 3.70 0.87 0.73 5.57 3.70 0.97 0.73 5.57	La	2.81	5.40	4.61	5.29	1.45	1.66	1.73	4.00	2.49	3.41	2.80
3.80 7.46 5.99 1.30 1.95 1.97 127 95 90 548 314 338 648 0.59 0.64 7.77 12.22 10.18 7.77 12.22 10.18 7.77 12.22 10.18 7.77 12.22 10.18 7.77 12.22 10.18 7.77 12.22 10.18 7.77 12.22 10.18 7.77 12.22 10.94 55.6 138.3 109.4 6021 2551 2625 3.79 4.67 3.24 0.68 0.97 0.73 3.79 4.36 3.24 0.80 0.97 0.73 3.79 4.36 3.24 0.80 0.97 0.73 3.79 4.36 3.24 0.80 0.97 0.73 3.79 4.36 3.24 0.80 0.97 0.73 3.70 4.36	Ce	7.13	13.43	11.87	13.12	4.52	5.14	4.96	12.11	7.90	9.09	6.90
1.30 1.95 1.97 127 95 90 548 314 338 548 314 338 0.48 0.59 0.64 7.77 12.22 10.18 7.77 12.22 10.18 3.09 4.56 2.72 55.6 138.3 109.4 55.6 138.3 109.4 6021 2.551 2.625 3.79 4.67 3.24 0.68 0.97 0.73 3.79 4.67 3.24 0.80 0.97 0.73 3.10 43.6 38.8 1.33 1.33 1.37 3.79 4.36 38.8 0.98 0.97 0.58 0.98 0.87 0.58 1.33 1.33 1.37 3.79 4.28 4.17 0.52 0.87 0.58 1.33 1.33 1.37 1.33 1.33 1.37 1.13 0.52 0	Pb	3.80	7.46	5.99	7.48	1.21	2.50	2.81	3.43	2.54	7.22	3.47
127 95 90 548 314 338 548 314 338 648 0.59 0.64 7.77 12.22 10.18 7.77 12.22 10.18 7.77 12.22 10.18 7.77 12.22 10.18 7.77 138.3 109.4 55.6 138.3 109.4 2.12 4.23 2.72 0.68 0.95 1.12 0.68 0.97 3.24 0.61 2551 2625 3.79 4.67 3.24 0.80 0.97 0.73 3.79 4.36 38.8 1.33 1.33 1.37 3.10 4.36 38.8 0.98 0.97 0.58 3.79 4.28 4.17 0.52 0.87 0.58 3.79 0.87 0.58 3.79 4.28 4.17 0.53 0.87 0.58 0.53 0.87	Pr	1.30	1.95	1.97	2.17	0.83	0.92	0.92	2.11	1.36	1.52	1.21
548 314 338 0.48 0.59 0.64 7.77 12.22 10.18 7.77 12.22 10.18 3.09 4.56 2.72 55.6 138.3 109.4 55.6 138.3 109.4 55.6 138.3 109.4 55.6 138.3 109.4 6021 2551 2.57 0.68 0.95 1.12 6021 2551 2625 3.79 4.67 3.24 0.80 0.97 0.73 3.79 4.67 3.24 0.80 0.97 3.24 0.80 0.97 3.24 0.80 0.97 3.24 0.80 0.97 3.24 3.1.0 43.6 3.8 3.1.0 43.6 3.8 3.79 4.28 4.17 0.53 0.87 0.58 3.79 4.28 4.17 0.53 0.54 0.58 3.70 0.87	Sr	127	95	90	107	170	174	179	161	206	177	223
0.48 0.59 0.64 7.77 12.22 10.18 7.77 12.22 10.18 3.09 4.56 2.72 55.6 138.3 109.4 55.6 138.3 109.4 2.12 4.23 2.57 0.68 0.95 1.12 6021 2551 2625 3.79 4.67 3.24 0.80 0.97 0.73 5.36 5.77 6.82 31.0 43.6 38.8 31.0 43.6 38.8 31.0 43.6 38.8 379 4.28 4.17 0.98 0.87 0.58 3.79 4.28 4.11 0.52 0.87 0.58 3.76 4.28 4.11 0.53 0.87 0.58 3.76 0.87 0.58 3.79 4.28 4.17 0.53 0.56 0.58 <td>Р</td> <td>548</td> <td>314</td> <td>338</td> <td>517</td> <td>531</td> <td>467</td> <td>445</td> <td>535</td> <td>737</td> <td>592</td> <td>661</td>	Р	548	314	338	517	531	467	445	535	737	592	661
7.77 12.22 10.18 3.09 4.56 2.72 55.6 138.3 109.4 55.6 138.3 109.4 2.12 4.23 2.57 0.68 0.95 1.12 0.68 0.95 1.12 0.6021 2551 2625 3.79 4.67 3.24 0.80 0.97 0.73 5.36 5.77 6.82 31.0 43.6 38.8 31.0 43.6 38.8 31.0 43.6 38.8 31.0 4.36 38.8 379 4.28 4.17 0.98 0.87 0.58 3.79 4.28 4.17 0.52 0.87 0.58 0.59 0.87 0.58 0.59 0.87 0.58 0.59 0.87 0.58 0.59 0.87 0.58 0.59 0.87 0.58 0.51 0.86 0.58 0.51 0.59	Be	0.48	0.59	0.64	0.61	0.49	0.51	0.40	0.67	0.57	0.55	0.44
3.09 4.56 2.72 55.6 138.3 109.4 55.6 138.3 109.4 2.12 4.23 2.57 0.68 0.95 1.12 6021 2551 2625 3.79 4.67 3.24 0.80 0.97 0.73 5.36 5.77 6.82 31.0 43.6 38.8 1.33 1.33 1.37 379 4.36 38.8 0.98 0.97 0.73 5.36 5.77 6.82 31.0 43.6 38.8 31.0 43.6 38.8 379 4.28 4.17 0.98 1.33 1.37 0.52 0.87 0.58 0.59 0.86 0.58 0.59 0.86 0.65 0.59 0.86 0.65 0.59 0.86 0.65 0.59 0.86 0.65 0.66 0.87 0.65 0.67 0.86 <t< td=""><td>Nd</td><td>7.77</td><td>12.22</td><td>10.18</td><td>11.49</td><td>5.07</td><td>5.52</td><td>5.45</td><td>12.06</td><td>7.77</td><td>8.53</td><td>7.03</td></t<>	Nd	7.77	12.22	10.18	11.49	5.07	5.52	5.45	12.06	7.77	8.53	7.03
55.6 138.3 109.4 2.12 4.23 2.57 0.68 0.95 1.12 6021 2551 2625 3.79 4.67 3.24 0.80 0.97 0.73 5.36 5.77 6.82 31.0 43.6 38.8 0.98 1.33 1.37 0.98 1.33 1.37 0.98 1.33 1.37 37.9 4.36 38.8 0.98 1.33 1.37 0.98 1.33 1.37 0.98 0.87 0.58 0.52 0.87 0.58 0.52 0.87 0.58 0.59 0.86 0.65 0.59 0.86 0.65 0.59 0.86 0.65 5.50 0.86 0.65 7.6 0.88 0.65 1.13 0.59 0.65 0.59 0.86 0.65 0.50 0.86 0.65 1.6 0.87	Sm	3.09	4.56	2.72	3.86	2.02	2.11	2.04	4.36	2.87	3.03	2.56
2.12 4.23 2.57 0.68 0.95 1.12 6021 2551 2625 3.79 4.67 3.24 0.80 0.97 0.73 5.36 5.77 6.82 31.0 43.6 38.8 0.98 1.33 1.37 0.98 1.33 1.37 0.98 1.33 1.37 0.98 1.33 1.37 0.98 1.33 1.37 0.98 1.33 1.37 0.52 0.87 0.58 0.52 0.87 0.58 0.59 0.86 0.65	Zr	55.6	138.3	109.4	112.7	33.2	32.4	34.7	92.5	57.1	64.6	42.0
0.68 0.95 1.12 6021 2551 2625 3.79 4.67 3.24 0.80 0.97 0.73 5.36 5.77 6.82 31.0 43.6 38.8 0.98 0.97 0.73 5.36 5.77 6.82 31.0 43.6 38.8 0.98 1.33 1.37 3.79 4.36 38.8 0.98 1.33 1.37 3.79 4.28 4.17 0.52 0.87 0.58 0.52 0.87 0.58 0.59 0.86 0.65 0.59 0.86 0.65 1 0.59 0.65 1 0.59 0.86 0.65 1 0.59 0.86 0.65 1 0.59 0.86 0.65	Hf	2.12	4.23	2.57	3.55	1.14	1.17	1.26	3.17	1.79	2.25	1.40
6021 2551 2625 3.79 4.67 3.24 0.80 0.97 0.73 5.36 5.77 6.82 31.0 43.6 38.8 0.98 0.97 0.73 5.36 5.77 6.82 31.0 43.6 38.8 0.98 1.33 1.37 3.79 4.28 4.17 0.52 0.87 0.58 3.76 4.92 4.11 0.57 0.86 0.58 0.59 0.86 0.58 0.59 0.86 0.65 5.76 0.86 0.65 5.76 0.86 0.65 7.6 0.86 0.65 1 0.59 0.86 0.65 1 0.86 0.86 0.65	Eu	0.68	0.95	1.12	1.03	0.82	0.83	0.79	1.35	1.04	0.94	0.87
3.79 4.67 3.24 0.80 0.97 0.73 5.36 5.77 6.82 31.0 43.6 38.8 0.98 1.33 1.37 0.98 1.33 1.37 0.98 1.33 1.37 3.79 4.28 4.17 0.52 0.87 0.58 3.76 4.92 4.11 0.59 0.86 0.65 6.59 0.87 0.65 76 4.92 0.11 0.59 0.86 0.65 for the purpose of comparison, laser ICP-MS analyses of norm 1.5 y el Table 1. For other abbreviations see Table 1. K, P and Ti by el 0.65 ndod PhysikCompex 110; W.F. McDonough and I. Horn, Harva 1.5 w el	Ξ	6021	2551	2625	3257	6543	7551	7312	3808	6107	4720	6442
0.80 0.97 0.73 5.36 5.77 6.82 31.0 43.6 5.88 31.0 43.6 5.88 31.0 43.6 5.88 31.0 43.6 5.88 379 4.28 4.17 0.52 0.87 0.58 3.76 4.92 4.11 0.59 0.86 0.65 6.59 0.86 0.65 1 0.59 0.86 0.65 1 0.59 0.86 0.65 1 0.59 0.86 0.65 1 0.59 0.86 0.65 1 0.59 0.86 0.65 1 0.59 0.86 0.65 1 0.59 0.86 0.65 1 0.59 0.86 0.65	Gd	3.79	4.67	3.24	5.51	2.95	3.07	3.19	5.66	4.05	4.24	3.58
5.36 5.77 6.82 31.0 43.6 5.88 31.0 43.6 38.8 0.98 1.33 1.37 3.79 4.28 4.17 0.52 0.87 0.58 3.76 4.92 4.11 0.59 0.86 0.65 750 0.86 0.65 1 0.59 0.86 0.65 1 0.59 0.86 0.65 1 0.59 0.86 0.65 1 0.59 0.86 0.65 1 0.59 0.86 0.65 1 0.59 0.86 0.65 1 0.59 0.86 0.65	Tb	0.80	0.97	0.73	0.76	0.51	0.55	0.54	1.02	0.71	0.75	0.61
31.0 43.6 38.8 0.98 1.33 1.37 3.79 4.28 4.17 3.79 4.28 4.17 0.52 0.87 0.58 3.76 4.92 4.11 0.59 0.86 0.65 For the purpose of comparison, laser ICP-MS analyses of norm Table 1. For other abbreviations see Table 1. K, P and Ti by el abda PhysikCompex 110; W.F. McDonough and I. Horn, Harva	Dy	5.36	5.77	6.82	6.81	3.84	4.08	3.92	7.54	4.96	5.47	4.41
0.98 1.33 1.37 3.79 4.28 4.17 3.76 4.92 4.11 0.52 0.87 0.58 3.76 4.92 4.11 0.59 0.86 0.65 76 4.92 1.15 76 4.92 0.65 77 0.86 0.65 78 0.86 0.65 79 0.86 0.65 70 0.86 0.65 70 0.86 0.65 70 0.86 0.65 70 0.86 0.65 70 0.86 0.65 70 0.86 0.65 70 0.86 0.86 70 0.86 0.96 70 0.86 0.96 70 0.86 0.65	Y	31.0	43.6	38.8	39.5	22.0	22.4	23.0	43.8	29.6	32.5	26.9
Er 3.79 4.28 4.17 4.44 2.48 2.62 2.56 4.99 3.40 3.63 2.8 Tm 0.52 0.87 0.58 0.76 0.37 0.38 0.39 0.77 0.50 0.56 0.4 Yb 3.76 4.92 4.11 4.16 2.57 2.65 2.60 5.36 3.41 3.95 3.0 Lu 0.59 0.86 0.65 0.73 0.40 0.40 0.84 0.53 0.61 0.4 ^a For the purpose of comparison, laser ICP-MS analyses of normal group glasses are also shown (samples 19, 29, 59, 69, 71, 90, 94; data from <i>Straub et al.</i> [2003]). For full identification of sample-se Table 1. For other abbreviations see Table 1. K. P and Ti by electron microprobe, Be by ion probe. All other trace elements by laser ablation ICP-MS (VG Elemental Plasma Quad with Excimpte Lambda PhysikCompex 110; W.F. McDonough and I. Horn, Harvard University). Errors (1s) are within 10% (REE 5%), except for Cs (25%), Rb (12%) see <i>Straub et al.</i> [2003] for anal [2003] for anal	Но	0.98	1.33	1.37	1.21	0.83	0.89	0.85	1.64	1.15	1.20	0.98
Tm 0.52 0.87 0.58 0.76 0.37 0.38 0.39 0.77 0.50 0.56 0.4 Yb 3.76 4.92 4.11 4.16 2.57 2.65 2.60 5.36 3.41 3.95 3.0 Lu 0.59 0.86 0.65 0.73 0.40 0.41 0.40 0.61 0.61 0.61 0.61 0.61 0.61 0.61 0.61 0.61 0.41 0.40 0.61 0.41 0.40 0.61 0.41 0.40 0.61 0.61 0.41 0.41 0.40 0.61 0.41	Er	3.79	4.28	4.17	4.44	2.48	2.62	2.56	4.99	3.40	3.63	2.84
Yb 3.76 4.92 4.11 4.16 2.57 2.65 2.60 5.36 3.41 3.95 3.0 Lu 0.59 0.86 0.65 0.73 0.40 0.41 0.40 0.84 0.53 0.61 0.4 ^a For the purpose of comparison, laser ICP-MS analyses of normal group glasses are also shown (samples 19, 29, 59, 69, 71, 90, 94; data from <i>Straub et al.</i> [2003]). For full identification of sample-see Table 1. For other abbreviations see Table 1. K. P and Ti by electron microprobe, Be by ion probe. All other trace elements by laser ablation ICP-MS (VG Elemental Plasma Quad with Excinent Lambda PhysikCompex 110; W.F. McDonough and I. Horn, Harvard University). Errors (1s) are within 10% (REE 5%), except for Cs (25%), Rb (12%) and Pb (16%) (see <i>Straub et al.</i> [2003] for anal	Tm	0.52	0.87	0.58	0.76	0.37	0.38	0.39	0.77	0.50	0.56	0.44
Lu 0.59 0.86 0.65 0.73 0.40 0.41 0.40 0.84 0.53 0.61 0.4 ^a For the purpose of comparison, laser ICP-MS analyses of normal group glasses are also shown (samples 19, 29, 59, 69, 71, 90, 94; data from <i>Straub et al.</i> [2003]). For full identification of sample-see Table 1. For other abbreviations see Table 1. K. P and Ti by electron microprobe, Be by ion probe. All other trace elements by laser ablation ICP-MS (VG Elemental Plasma Quad with Excime Lambda PhysikCompex 110; W.F. McDonough and I. Horn, Harvard University). Errors (1s) are within 10% (REE 5%), except for Cs (25%), Rb (12%) and Pb (16%) (see <i>Straub et al.</i> [2003] for anal 0.41	Yb	3.76	4.92	4.11	4.16	2.57	2.65	2.60	5.36	3.41	3.95	3.02
^a For the purpose of comparison, laser ICP-MS analyses of normal group glasses are also shown (samples 19, 29, 59, 69, 71, 90, 94; data from <i>Straub et al.</i> [2003]). For full identification of sample see Table 1. For other abbreviations see Table 1. K, P and Ti by electron microprobe, Be by ion probe. All other trace elements by laser ablation ICP-MS (VG Elemental Plasma Quad with Excimer Lambda PhysikCompex 110; W.F. McDonough and I. Horn, Harvard University). Errors (1s) are within 10% (REE 5%), except for Cs (25%), Rb (12%) and Pb (16%) (see <i>Straub et al.</i> [2003] for anal	Lu	0.59	0.86	0.65	0.73	0.40	0.41	0.40	0.84	0.53	0.61	0.46
defailed	^a For th see Table Lambda Pl	le purpose of con 1. For other abb hysikCompex 1	mparison, laser ICP-M: previations see Table 1. 10; W.F. McDonough 6	S analyses of normal { . K , P and Ti by elect and I. Horn, Harvard I	group glasses are al: tron microprobe, B¢ University). Errors (so shown (sample s by ion probe. A (1s) are within 10	es 19, 29, 59, 69, Jl other trace ele % (REE 5%), ex	, 71, 90, 94; data ments by laser at cept for Cs (25%	from <i>Straub et a</i> , olation ICP-MS (), Rb (12%) and	<i>l</i> . [2003]). For fu (VG Elemental P Pb (16%) (see <i>S</i> i	ll identification o lasma Quad with traub et al. [2003	of sample code: 1 Excimer lase 3] for analytica

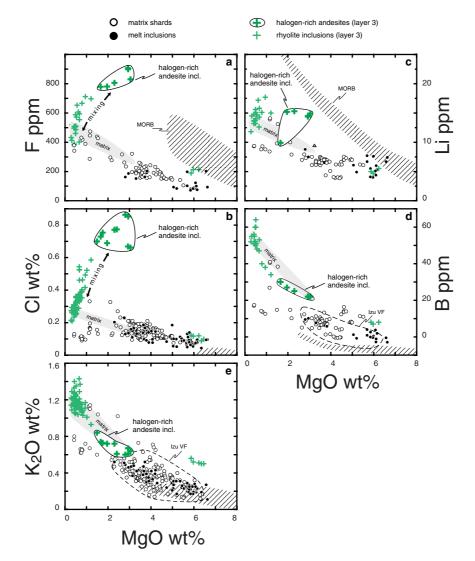


Figure 2. Variations of F, Cl, Li, K₂O and B versus MgO in the Izu VF glasses. The normal-group glasses (melt inclusions are filled circles; matrix glasses are open circles; grey field denoted "matrix" indicated matrix shards in sample 3) display a general enrichment in these elements with decreasing MgO. The halogen-rich andesite (HRA) inclusions (bold green crosses) are selectively enriched in Cl, F and Li. Mixed rhyolite inclusions (green crosses) lie along apparent mixing trends between the HRA and normal-group rhyolite glasses. The fields for N-MORB are from *Ryan and Langmuir* [1987] and *Ryan and Langmuir* [1993] (B, Li), and *Michael and Cornell* [1998] (Cl, K). The Cl reference field is for depleted MORB with Cl/K < 0.08. Izu VF after *Ryan and Langmuir* [1993], *Taylor and Nesbitt* [1998] and *Langmuir et al.* [2003].

ments and high Cl/K₂O ratios. However, only one fallout tephra (layer 3 at 0.55 Ma) contained melt inclusions that are large enough (up to 100 micrometer diameter) for the analysis of trace elements and B isotopes by laser ablation-ICPMS and ion microprobe, respectively. These latter analyses show that the normal-group glasses and the HRA are indistinguishable in all other incompatible trace elements, and in the B isotopes [*Straub and Layne*, 2002, 2003]. In each fallout

tephra layer, the HRA always constitute a minor portion (<1 vol.%) of the entire melt inclusion population in the plagioclases. The HRA are always included within bytownitic (An_{89-90}) plagioclase cores, the most calcic plagioclase compositions present in any given fallout tephra (Figures 4a and 4b). The bytownitic cores are always mantled by labradoritic ($\sim An_{50-60}$) rims. These rims commonly contain "mixed" dacitic and rhyolitic melt inclusions, that lie on apparent mixing

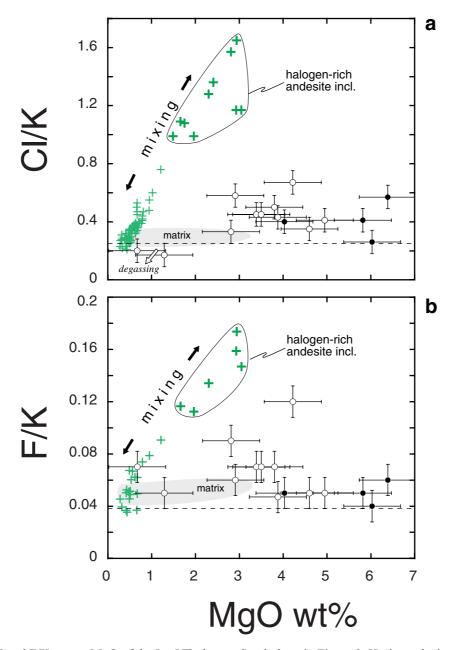


Figure 3. Cl/K and F/K versus MgO of the Izu VF glasses. Symbols as in Figure 2. Horizontal stippled lines are the minimum Cl/K and F/K ratios of the normal-group glasses. No decrease in Cl/K and F/K with decreasing MgO is apparent, implying that only a single, Cl- and F-poor, aqueous vapor phase was exsolved, resulting in no significant loss of halogens. Only the most siliceous normal-group glasses fall below this perceived minimum for Cl/K (white arrow), which is attributed to the incipient exsolution of hypersaline brine, producing substantial depletion of melt Cl.

trends between the lowest MgO normal-group rhyolite glasses and the HRA (Figures 4a and 4b).

4.2. The Effects of Fractionation

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[9] In the low-K Izu VF magmas, Li, Be, B, F, Cl, and K all have bulk partition coefficients <1 with

respect to the observed mineral assemblage, which is dominated by plagioclase and pyroxene. The D's are, in decreasing order; $D_F \sim 0.8$, $D_{Cl} \sim 0.54$, $D_{Be} \sim 0.27$, $D_B \sim 0.09$ and $D_K \sim 0.09$, based on a compilation of the most recent partitioning data from *Bindeman et al.* [1998], *Brenan et al.* [1998] and *Dunn and Sen* [1994] (see *Straub and Layne*



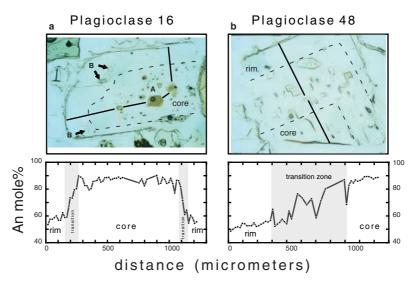


Figure 4. Photomicrographs of plagioclase phenocrysts with HRA and mixed rhyolite inclusions in fallout tephra 125-782A-2H-4-113-114 (= tephra sample 3 at 0.55 Ma). Each fragment is approximately one millimeter across. (a) A broad core ($\sim An_{89-90}$) contains several brown, translucent HRA melt inclusions. Bubbles indicate fluid exsolution. The rim is a sodium-rich labradorite of $\sim An_{55-60}$, containing several colorless inclusions of mixed rhyolite (B with arrows). Note abrupt transition from core to rim with a drop of ~ 30 mole% An. (b) A broad, mesh-textured transition zone with mixed rhyolite inclusions separates a typically calcic core (An_{88-89} but without HRA) from a labradoritic rim ($\sim An_{55}$).

[2003] for more details). Thus these elements are not perceptibly fractionated from each other during melt crystallization as the degree of crystallization at the Izu VF is likely less than 70% [e.g., Langmuir et al., 2003; Straub and Layne, 2003]. Olivine, which is present in the Izu VF lavas, but not in the tephras, does not appreciably fractionate these elements either. Apatite can take up Cl and F, but it occurs in small quantities (<1 vol.%) in highsilica glasses only, and has no perceptible effect on the halogen content of the melt [Straub and Layne, 2003]. Further, the halogens are not fractionated by degassing, although the Izu melts do lose up to several wt.% magmatic H₂O at upper crustal levels [Straub and Layne, 2003]. This can be shown by ratioing Cl and F to K, an incompatible, nonvolatile trace element (Figure 3). Note that the vesicular, dacitic (<2 wt% MgO) matrix shards of the normal-group have the same minimum Cl/K, F/K and Li/K (not shown) ratios as undegassed basaltic melt inclusions, which effectively precludes substantiative halogen loss by degassing. Consequently, the wide range of Cl/K, F/K and Li/K observed in the glasses should reflect mantle source heterogeneity.

[10] It is not clear whether the parental melts of the HRA are basaltic or andesitic as both are primary melts in subduction zones [e.g., *Hirose*, 1997; *Carmichael*, 2002; *Tamura and Tatsumi*, 2002]. As to the Cl/K, F/K and Li/K ratios considered, this question is not relevant, since the effects of fractional crystallization on these ratios are negligible at any rate. We note, however, that *Kohut et al.* [2002] report high Cl contents (~0.8 wt%) in basaltic melt inclusions in olivine phenocrysts from the submarine Southern Mariana arc, demonstrating that such Cl enrichments can exist in basaltic melts as well.

4.3. Sources of Cl, F and Li at the Izu VF

[11] In Figure 5, Cl, F and Li are compared to the other incompatible elements. As the recycled fluid-mobile LILE (Cs, Ba, Rb, B, U, Th, K, Pb, Sr), the halogens, and Li, are substantially enriched relative to non-fluid-mobile incompatible elements (Nb, Ta, Zr, Hf, middle and heavy REE). Given the high degree of melting beneath the Izu VF ($\sim 20-25\%$ [*Plank and Langmuir*, 1988; *Hochstaedter et al.*, 2000], significant fractionation of the incompatible elements during mantle melting can be



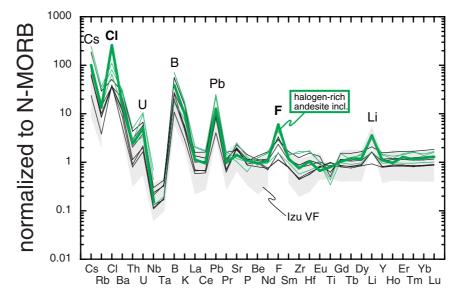


Figure 5. Incompatible elements in Izu VF glasses measured by LA-ICPMS. Thick green line denotes halogen-rich andesite inclusions (Ta below detection limit), thin lines are normal-group glasses (black) and mixed rhyolites (green), respectively. Izu VF after *Taylor and Nesbitt* [1998] and *Langmuir et al.* [2003].

excluded. Therefore assuming that the excess enrichments relative to similarly incompatible, but non-recycled elements (e.g., Nb, Ta, middle and heavy REE) are due to slab additions, the percentages of Cl, F and Li added from slab can be calculated using a simple mass balance model:

$$c_{VF} = c_{BM} * 0.975 + c_{SF} * 0.025 \tag{1}$$

whereby c_{VF} = concentration of an element in Izu VF mantle source, c_{BM} = concentration in the Izu background mantle, c_{SF} = concentration in the slab fluid (Straub et al., submitted manuscript, 2003). In this simplified model, the slab fluid is considered as single component, although it might gain elements from several slab sources (e.g., metasediment, metabasalt, wedge and serpentinite). The "Izu VF mantle" source is the subarc mantle that is metasomatized by the fluid from the slab. The "Izu background mantle" is the Izu VF source minus the additions from the slab. The mixing ratio of background wedge ($\sim 97.5\%$) and slab fluid $(\sim 2.5\%)$ has been inferred previously from Pb and Pb isotope modeling (Straub et al., submitted manuscript, 2003). The details of the calculations, including the determination of Izu VF source and background mantle compositions, can be found in Straub et al. (submitted manuscript, 2003), and are not reiterated here. The end-member compositions needed, however, are listed in Table 5 for easy reference. In summary, these calculations show that >99% of Cl, ~60% of F and ~34% of Li in the normal-group glasses are derived from slab; the remainders derive from the subarc mantle wedge. In the HRA component melt, the percentages of slab-derived Cl, F and Li are higher, being ~99.9% Cl, ~92% F and ~76% Li, respectively.

4.4. Mixing Systematics of Normal-Group Glasses and HRA

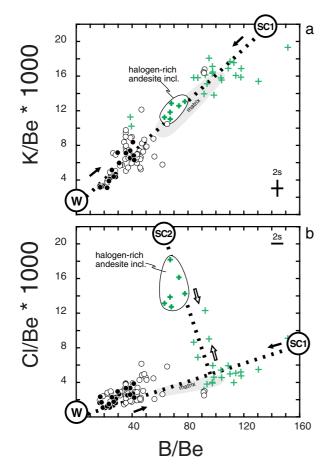
[12] Ratioing Cl, F and Li to similarly incompatible elements, that derive either from the slab or the subarc wedge, highlights source characteristics (Figures 6 and 7). The most important information gained from these diagrams, is that - in terms of Cl, F and Li - a compositional spectrum of fluids exists that is derived from at least two subarc mantle sources (Figure 6). The element Be was chosen as incompatible wedge-derived element, since it has been analyzed in all glasses investigated by ion microprobe, in contrast to the other wedge-derived elements (e.g., Nb, middle and heavy REE). Two observations show that Be has no significant slab addition at the Izu VF: (1) the Be content of Izu VF glasses and lavas are both significantly lower than

	1	1a	2	2a	3	4	5	5a	9	6a
	Avg Izu VF normal group glasses	Avg Izu VF HRA	Avg Izu VF source normal group glasses	Avg Izu VF source <i>HRA</i>	Izu background mantle	bulk partition coeff.	Avg comp. fluid normal group glasses	Avg comp. fluid <i>HRA</i>	Perc% Added normal group glasses	Perc% Added HRA
Cs Rb	0.32 3.47		0.0811 0.89		$0.0002 \\ 0.016$	0.0081 0.0081	3.24 35		99.7 98.0	
C	925	6000	236	1528	0.83	0.0062	9390	61087	9.66	6.66
Ba	45.1		11.4		0.174	0.0043	450		98.4	
Πh	0.11		0.029		0.0033	0.0039	1.02		87.5	
D	0.09		0.022		0.0013	0.0036	0.81		93.8	
qN	0.24		0.060		0.0531	0.0034	0		0	
Та	0.020		0.005		0.0048	0.0030	0		0	
Х	2148		541		34	0.0025	20272		93.0	
La	1.42		0.37		0.204	0.0154	6.7		43.7	
Ce	4.08		1.09		0.786	0.0228	12.2		27.6	
Pb	1.79		0.49		0.036	0.0315	18		91.6	
Sr	152		42		10.9	0.0315	1224		73.4	
ΡN	4.45		1.25		1.012	0.0403	9.4		18.8	
ц	150	680	41	194	16	0.0508	066	7122	59.6	92
Sm	1.71		0.51		0.476	0.0613	1.2		5.9	
Li	4.64	12	1.69	4.4	1.06	0.153	7.3	133	33.8	76
^a Se respec recons is the t for Cs, wedge	e Straub et al. [21 tively. All element tructed from averag aackground mantle Th, U, Nb, Ta, Pb mixing proportion	003] and <i>Straub</i> a values in ppm. 1 a ge Izu VF (column wedge. 4 is the bu and Sr are interpol of 0.025 : 0.975	^a See <i>Straub et al.</i> [2003] and <i>Straub and Layne</i> [2003] for details of calculation. Trace elements other than Cl, <i>F</i> and Li are identical in components based on normal group glasses and HRA, respectively. All element values in ppm. 1 and 1a and the average of primitive magmas from the Izu VF, based on fractionation-corrected composition of Izu glasses. 2 and 2a are the Izu VF mantle source, reconstructed from average Izu VF (column 1), assuming 25% degree of partial melting (batch melting), a depleted peridotite (OLIV:OPX:CPX = 60:25:15) and the bulk partition coefficients of column 4. 3 is the background mantle wedge. 4 is the bulk partition coefficients (DS) used in melting equation, compiled from <i>Ewart and Hawkesworth</i> [1987] (Rb, Ba, K) and <i>Miller et al.</i> [1992] (La, Ce, Nd, Sm). D's for Cs, Th, U, Nb, Ta, Pb and Sr are interpolated using the sequence of incompatibility sequence in <i>Sun and McDonough</i> [1989]. 5 and 5a are the average of composite fluid, calculated using on a slab fluid: wedge mixing proportion of 0.025 : 0.975 [<i>Straub et al.</i> , 2003]. 6 and 6a are the percentage of slab-derived elements in Izu VF magmas.	ails of calculation. primitive magmas 4 of partial melting (5s) used in melting f incompatibility se nd 6a are the perc	details of calculation. Trace elements other than Cl, <i>F</i> and Li are identical in components based on normal group glasses and HRA. of primitive magmas from the Izu VF, based on fractionation-corrected composition of Izu glasses. 2 and 2a are the Izu VF mantle source, gree of partial melting (batch melting), a depleted peridotite (OLIV:OPX:CPX = $60:25:15$) and the bulk partition coefficients of column 4. 3 s (Ds) used in melting equation, compiled from <i>Ewart and Hawkesworth</i> [1987] (Rb, Ba, K) and <i>Miller et al.</i> [1992] (La, Ce, Nd, Sm). D's ce of incompatibility sequence in <i>Sun and McDonough</i> [1989]. 5 and 5a are the average of composite fluid, calculated using on a slab fluid. 6 and 6a are the percentage of slab-derived elements in Izu VF magmas.	r than Cl, F and L d on fractionation-co leted peridoite (OL om Ewart and Haw cDonough [1989]. 5 d elements in Izu V	i are identical in cc orrected compositior IV:OPX:CPX = 60:27 <i>kesworth</i> [1987] (Rt 6 and 5a are the avera F magmas.	in of Izu glasses. 2 55:15) and the bu 5, Ba, K) and Miu 9, Ba, Composite	on normal group g and 2a are the Izu ' Ik partition coefficie <i>lev et al.</i> [1992] (La, fluid, calculated usii	Jasses and HRA, VF mantle source, nts of column 4.3 Ce, Nd, Sm). D's g on a slab fluid :

Table 5. Trace Element Composition of Slab and Mantle Wedge End-Members^a

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Figure 6. Cl/Be and K/Be versus B/Be of Izu VF glasses to demonstrate the selective enrichment of the halogens as exemplified by Cl (F and Li are not shown), and normalizing abundance to a wedge-derived element (Be). W, Wedge, SC1 and SC2 are the two different slab components identified. Error for Cl/Be is within symbol size. For other symbols see Figure 2.

MORB, a characteristic of depleted island arc lavas [*Ryan and Langmuir*, 1987], and (2) the Izu glasses have Be/Nd (0.07 ± 0.02 ; n = 28) within the range of the global average of N-MORB (0.05).

[13] Normalizing Cl, F and Li to Be identifies two slab components: Slab component 1 (SC1) that is present in the source of all normal-group glasses, and slab component 2 (SC2), an additional fluid component in the source that is selectively enriched in Cl, F and Li (Figure 6b). Interestingly, in Figure 6b, the linear mixing trend from SC2 (represented by the HRA) does not trend toward the depleted mantle wedge (W), but toward the rhyolite melts of the normal-group glasses. This argues against the two fluids infiltrating separate subarc mantle domains at different times. Rather, the two fluids must infiltrate the Izu VF more or less simultaneously becoming mixed until at crustal levels the melts are nearly homogenized. This is consistent with the petrographic features, which strongly suggest that the HRA are relicts of the pre-aggregated andesitic component melts, that subsequently mixed with the far more numerous normal-group melt batches. In most cases the

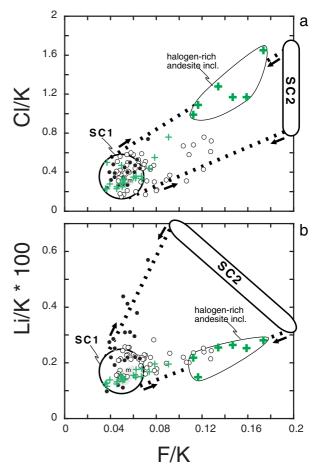


Figure 7. Cl/K and Li/K versus F/K of Izu VF glasses to illustrate the compositional variability of the two slab components. The slab component infiltrating the mantle source of the normal-group glasses (SC1) has an inherent variability (large circle); m, matrix shards. For other symbols see Figure 2. Note the clearly manifested mixing line of HRA (bold green crosses) and mixed rhyolite inclusion (thin green crosses). However, not all trends emanating from this end-member trend toward the HRA, indicating that the selectively enriched second slab component (SC2) is much more variable than indicated by the HRA of fallout tephra 3.



subordinate HRA component melt apparently did not survive the mixing process, and faded into the normal-group melt. If so, the significant range in Cl/Be (Figure 6) (as well as in F/Be and Li/Be, not shown) of the normal-group glasses can be interpreted as a "remnant signal," produced by the preeruptive mixing of these multiple component melts.

[14] Normalizing Cl, F and Li to another slabderived element (e.g., K) further highlights the compositional variability of these slab fluids. In the F/K versus Cl/K diagram (Figure 7a), the data plot in a broad band, extending from the normalgroup glasses (SC1 with low Cl/K and F/K) toward SC2 (HRA), which has high Cl/K and F/K (Figure 7). The array is roughly linear, confirming the existence of two, partially mixed, endmember components. However, having a rather limited range, the mixed rhyolites and HRA appear to represent an example of this array only, implying that the Cl/F of SC2 (HRA) could be even lower. The heterogeneity of the HRA fluid (SC2) is more evident in F/K versus Li/K space (Figure 7b). Apparently, halogen-rich fluids can have very variable Li contents, showing that the Li enrichment is not strictly coupled to the halogen flux. Because of its high Cl and F contents, SC2 (HRA) must affect the halogen contents of the entire spectrum of fluids, even if present in much smaller quantity than the SC1 component. This would also explain why, in the normal-group Izu glasses, the paired correlations of Cl and F with the other fluidmobile LILE are substantially poorer than the correlations of the fluid-mobile LILE with each other.

5. Discussion

5.1. Origin by Assimilation at Upper Crustal Levels?

[15] In the preceding section, we implied that the Cl, F and Li variations in the Izu VF tephras reflect mantle source signatures. However, similarly elevated Cl contents are found in basaltic glasses and melt inclusions from intraplate and backarc environments [*Michael and Schilling*, 1989; *Kent et al.*, 2002; *Lassiter et al.*, 2002], where they have been

attributed to the assimilation of seawater, or seawater-altered crust, at upper crustal levels. Nonetheless, the more extensive data set collected for the Izu glasses, including B, F and Li concentrations, does not support an assimilation model.

[16] Assimilation of a seawater-derived brine must be ruled out, since the chemical signatures of the Izu glasses (both normal-group glasses and HRA) differ so strongly from seawater. Seawater has high Cl (19500 ppm), versus relatively low abundances of F (1.3 ppm), Li (0.18 ppm), and B (4.4 ppm). In contrast, the F (\sim 100–800 ppm) and Li (\sim 5– 15 ppm) abundances of the Izu glasses are much higher relative to their Cl contents (~1000-8000 ppm). Current data indicate that seawater processing within the oceanic crust does not produce fluids with the appropriate chemical signature either. For example, Cl-rich hydrothermal fluids become progressively enriched in Li (as well as the other alkalis, like Rb and Cs), but their B and F abundances remain close to those of seawater [e.g., Von Damm, 1995]. Thus if such a fluid impregnated the crust, from which HRA component melt was subsequently produced by partial fusion, considerable fractionation during melting would be required, the details of which are not clear [e.g., Lassiter et al., 2002].

[17] Another apparent problem concerns the amount of Cl, F, Li and B that can be assimilated. The normal-group glasses have Cl, F, Li and B abundances that are common in arc magmas worldwide [Ryan and Langmuir, 1993; Sisson and Lavne, 1993; Cervantes and Wallace, 2003]. Assimilation of several hundred ppm Cl, or several tens of ppm F, as reported in several studies [Sigvaldason and Oskarsson, 1986; Kent et al., 1999; Nielsen et al., 2000; Hauri, 2002], might add to the variability, but are insufficient to create the enrichment observed in the Izu HRA. Other studies, however, suggest an increase in Cl by 1 to 2 wt.%; i.e., the amount of Cl presumed to be assimilated is higher than the total Cl of the HRA (~0.8 wt%) [Michael and Schilling, 1989; Kent et al., 2002; Lassiter et al., 2002]. If this amount of Cl was added to the normal-group glasses melt by means of a partial crustal melt, then the absence of any variation in the alkaline and alkaline earth

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elements (e.g., Na, K, Rb) that are also enriched in seawater-derived fluids, is suspicious. Unfortunately, existing models of seawater assimilation in the intraplate and backarc basin environments are based primarily on Cl variability. The other light elements (F, Li, B) are not considered. Without this additional information, a direct comparison to the genetic processes in the Izu arc volcanic front remains unsatisfactory. We note, however, that in a single case of contaminated intraplate basalts, where F was analyzed [Hauri, 2002], it appears that the Cl and F enrichments of the glasses are unrelated. Thus if assimilation creates diversity, how could it then produce the systematic, longlasting enrichments of Cl, F and Li observed in the Izu VF glasses? Furthermore, the subducting slab is a known source of the elements Cl, Li and B in arc magmas [Ito et al., 1983; Ryan and Langmuir, 1993; Cervantes and Wallace, 2003], and slab processing at the relevant temperatures and pressures should be capable of providing the systematic excess Cl, F and Li observed both in the normalgroup glasses and the HRA.

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5.2. Mantle Sources of the Halogens and Li

[18] According to current theory, several fluid sources are present in shallow subduction zones (Figure 8).

[19] 1. The "classic" slab fluid source is the metacrust, which consists of the metamorphosed subducting sediment (metasediment) and basaltic crust (metabasalt). Metacrust fluids are generated by the decay of hydrous phases (e.g., chlorite, serpentine, amphibole) during progressive slab metamorphism [e.g., Tatsumi and Eggins, 1995; Schmidt and Poli, 1998; Bebout et al., 1999]. The capacity of these fluids to extract nominally "fluidmobile" elements from the metacrust, however, is contested. Even if extraction appears to be efficient for some elements (e.g., B [Ryan and Langmuir, 1993; Ryan et al., 1995, 1996; Benton et al., 2001]), there is increasing evidence, from experiments and from field studies in metamorphic terrains, that larger amounts of both subducted H₂O and potentially fluid-mobile LILE are buried in the slab itself by incorporation into the hydrous phases that appear at higher pressures, e.g., phen-

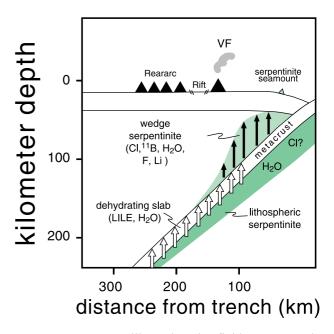


Figure 8. Cartoon illustrating the fluid sources and fluid distribution in shallow subduction zones, as discussed in the text. The hydrated wedge peridotite (light green) apparently forms beneath the forearc, and becomes down-dragged with the slab to arc front depth [e.g., *Straub and Layne*, 2002]. The serpentinized lithosphere (estimated thickness of 20 km; darker green) forms outboard of the arc, in the oceanic basin, prior to subduction into the trench [*Ranero et al.*, 2001].

gite, paragonite, lawsonite and epidote-group minerals [*Schmidt*, 1996; *Schmidt and Poli*, 1998; *Bebout et al.*, 1999; *Becker et al.*, 1999; *Hermann and Green*, 2001]. Only beyond arc front depths will the gradual destruction of these phases slowly release the elements that then control the chemistry of the slab components [*Schmidt*, 1996; *Melzer and Wunder*, 2000; *Hermann and Green*, 2001].

[20] 2. An alternative fluid source is the hydrated peridotite mantle above the metacrust (*Tatsumi and Eggins* [1995] and "wedge serpentinite" in Figure 8). The peridotite is hydrated trenchward from the VF by infiltration of slab fluids, and subsequently becomes down-dragged together with the metacrust. Tatsumi and coworkers [e.g., *Tatsumi and Eggins*, 1995; *Kogiso et al.*, 1997; *Tatsumi and Kosigo*, 1997] suggest that the decay of amphibole at \sim 75 km depth triggers substantial loss of fluid-mobile LILE. The LILE are then trapped in new phases in the hydrated mantle above (e.g., phlogopite, pargasitic amphibole), and transported to arc

front depths and beyond. This redistribution of the LILE, coupled with fractionation, is considered to control the trace element chemistry of the slab fluids [e.g., *Kogiso et al.*, 1997]. A variant of this model suggests that the shallow slab fluids carry only a few, extremely fluid-mobile elements, such as B or Cl [*Ryan et al.*, 2001; *Straub and Layne*, 2002], with the fluid-mobile LILE being largely retained in the slab. The extensively serpentinized mantle wedge was still down dragged with the metacrust to arc front depths, becoming the prime source of the isotopically heavy B that is observed in many arc lavas [*Straub and Layne*, 2002].

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[21] 3. Irrespective of the presence of these first two fluid sources, the largest fluid volume might actually originate from the serpentinized lithospheric mantle of the subducting slab, below the metacrust (Ranero et al. [2001], Rüpke et al. [2002], "lithospheric serpentinite" in Figure 8). This serpentinized lithosphere, which may reach 20 km or more in thickness, forms prior to subduction by (1) the infiltration of seawater during deep hydrothermal circulation, at spreading centers and in the oceanic basins [e.g., Seyfried et al., 1984; Agrinier et al., 1988; Nielsen et al., 2000], or (2) along deep faults cutting through the oceanic crust during flexural up bending as the crust approaches the trenches [Peacock, 2001; Ranero et al., 2001]. Potentially, this process is very efficient in recycling elements that are readily dissolved in hydrothermal fluids and seawater (e.g., Cl, B, rare gases) back into the mantle [*Rüpke* et al., 2002; Scambelluri et al., 2001].

[22] In short, the fluid sources present beneath the Izu VF fall into two categories: (1) the metacrust, and (2) serpentinized (hydrated) upper mantle, located either above or below the metacrust (Figure 8). In contrast to the metacrust, which is highly enriched in all LILE, the serpentinized mantle is highly depleted, except for those elements that are added by the fluids. Apparently, the metacrust is then the most likely source of the LILE-rich fluids that infiltrated the mantle source of the normal-group glasses. However, the serpentinized mantle could produce the additional fluid component whose signal is preserved in the HRA. To date, there is no evidence to suggest wholesale transport of LILE from the oceanic crust or seawater into the mantle peridotite, which could in turn serve as the major source for the LILE-rich fluids. Field studies from the Mariana forearc serpentinites, from exhumed segments of the metacrust, and from the serpentinized upper mantle, all indicate that these contain only minor quantities of the fluid-mobile LILE relative to the metacrust [Ryan et al., 2001; Scambelluri et al., 2001]. For example, hydrated upper mantle exposed in the Alps usually has <10 ppm Sr (and a maximum of 34 ppm Sr) [Scambelluri et al., 2001], well below the average of the metacrust (~ 100 ppm Sr). On the other hand, the serpentinized peridotite provides an excellent sink for a few highly fluid-mobile elements, such as Cl, F and Li [Rucklidge and Patterson, 1977; Agrinier et al., 1988; Puga et al., 1999; Anselmi et al., 2000]. The Cl and F replace the (OH⁻)-groups in hydrous Mg-silicates (e.g., serpentine and chlorite, or amphibole and phlogopite at greater depths), whereas the Li ion (Li⁺) substitutes for Mg²⁺. Studies show that modern and ancient oceanic serpentinites can contain several 1000 ppm of Cl and F, with maxima of 9000 ppm Cl and 3000 ppm F, respectively [Rucklidge and Patterson, 1977; Agrinier et al., 1988; Puga et al., 1999; Anselmi et al., 2000]. In contrast, the average metacrust contains only \sim 113 ppm Cl and \sim 241 ppm F [e.g., Straub and Layne, 2003]. The serpentinite of the outer Mariana forearc contains 2-19ppm Li [Benton et al., 1999] compared to an average upper mantle value of ~ 1 ppm Li (Table 5).

5.2.1. Lithospheric or Wedge Serpentinite as Fluid Source?

[23] Of the two serpentinite reservoirs (lithospheric serpentinite and wedge serpentinite), the wedge serpentinite appears to be the most likely source of the HRA fluid component, for several reasons.

[24] First, the volumetric dominance of the normalgroup glasses relative to the HRA suggests that most of the total fluid volume that infiltrates the Izu VF mantle originates from the metacrust. The serpentinite must be a subsidiary fluid source. This characteristic appears to be more consistent with the wedge serpentinite, as the lithospheric serpentinite may produce an even larger fluid volume

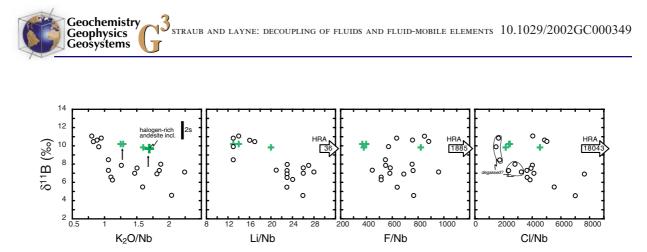


Figure 9. K₂O/Nb, Li/Nb, F/Nb and Cl/Nb versus δ^{11} B of Izu glasses. For symbols see Figure 2. K₂O/Nb exemplifies the trends of LILE/Nb and B/Nb in normal-group glasses (open circles; see *Straub and Layne* [2002, Figure 5]). HRA (bold green cross) and the mixed rhyolite inclusions (smaller green crosses) have slightly higher δ^{11} B values (indicated by black arrows) than the normal-group glasses (note that the K₂O contents of HRA and normal-group glasses are identical; see Figure 2), consistent with their origin from a high- δ^{11} B source (i.e., wedge serpentinite with δ^{11} B ~+15‰ [*Straub and Layne*, 2002]). Numbers within arrows give abundances (in ppm) of Li, F and Cl in HRA, which plot outside the diagram. Note linear mixing trends defined by the HRA and the mixed rhyolites.

than the metacrust [Ranero et al., 2001; Rüpke et al., 2002].

[25] Second, any fluid from the lithospheric serpentinite must pass through the ~ 6.4 km thick metacrust, from which it is very likely to scavenge additional elements [e.g., Rüpke et al., 2002]. Therefore if the lithospheric serpentinite was a significant fluid source, the fluid-mobile LILE and the halogens should be correlated, rather than the observed decoupling. If the lithospheric serpentinite was the HRA fluid source, then the decoupling is particularly difficult to explain for Li, an element that is more abundant in the metacrust than in the serpentinite [e.g., Ryan and Langmuir, 1987]. A scenario in which the fluid from the lithospheric serpentinite passed through the metacrust along continuous cracks, or even through a slab window, is not consistent with the co-existence of the HRA and the normal-group glasses in crustal chambers, an observation that implies that the two fluid components must infiltrate the mantle source more or less simultaneously.

[26] Third, the normal-group Izu glasses display inverse correlations of δ^{11} B and LILE/Nb ratios [*Straub and Layne*, 2002, Figure 5]. These inverse correlations provide a strong argument against both heavy B and the LILE being derived from the same fluid source. If the metacrust (or perhaps the lithospheric serpentinite) was the major source of heavy B, as suggested by several workers [e.g., *Ishikawa et al.*, 2001; *Tonarini et al.*, 2001; *Rüpke* *et al.*, 2002], then δ^{11} B and LILE/Nb should be positively correlated. Moreover, at arc front depths, the metacrust apparently has δ^{11} B < 0, after having lost heavy B during shallow devolatilization [*Peacock and Hervig*, 1990; *Straub and Layne*, 2002]. If this is the case, then it appears likely that the lithospheric serpentinite is similarly depleted in heavy B at arc front depths. Therefore the most likely source of heavy B is, in fact, the wedge serpentinite, which became impregnated with ¹¹Brich fluids trenchward of the Izu VF during shallow dewatering of the slab.

[27] In summary, the wedge serpentinite is the most likely source of the HRA fluid. Although this model does not preclude some lesser contribution of fluids from the lithospheric serpentinite, these "lithospheric" fluids are likely to completely mix with the metacrust fluids, forming a single, composite fluid of minor compositional variability.

5.2.2. Relative Contributions of Cl, F and Li From the Two Fluid Sources

[28] Correlating Cl, F and Li with δ^{11} B allows a first order estimate of the relative importance of the wedge serpentinite versus metacrust fluid sources for these elements at the Izu VF. This can be done because the Izu VF B originates in approximately equal amounts from the serpentinite wedge (~53% of B) and the metacrust (~47% of B) that have very different δ^{11} B values (+14‰ and +1‰, respectively [*Straub and Layne*, 2002]). Figure 9

shows that the normal-group glasses are inversely correlated only in the Li/Nb versus δ^{11} B space, suggesting that the fluid-transported Li - as B - was contributed both from the metacrust and the wedge serpentinite in about equal amounts. In contrast, Cl/Nb and F/Nb are uncorrelated with δ^{11} B, even if some loss of Cl by degassing is taken into account (Figure 9). This lack of clear correlation with $\delta^{11}B$ suggests that Cl and F are supplied predominantly from only one of the two fluid sources, which should be the wedge serpentinite according to our model. This interpretation is supported by the elevated $\delta^{11}B$ values the HRA and the mixed rhyolite inclusions display relative to the normalgroup glasses, given the fact that both groups have similar K₂O at a given MgO (Figure 9).

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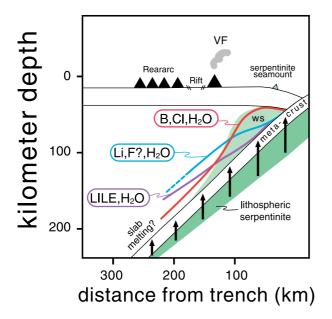
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[29] Therefore if the wedge serpentinite is the dominant source of the fluid-transported Cl and F at the Izu VF, the contributions of B and Li from this reservoir must be less. A weaker relative fluid flux of Li from the wedge serpentinite is supported by the more moderate Li enrichment of the HRA compared to the halogens, and the lack of selective Li enrichment in a single HRA inclusion (Figure 2c). Further, the lack of a selective B enrichment in the HRA implies that the relative B flux from the wedge serpentinite is even somewhat less than the relative flux of Li. Thus the B contribution from the serpentinite wedge is entirely concealed in the abundance data (but not in the B isotopes). Since \sim 53% of the total B at the Izu VF is wedge-derived [Straub and Lavne, 2002], more than 53% of the Cl, F and Li must originate from the wedge serpentinite at the Izu VF, with only the remainder being contributed from the metacrust fluids.

5.2.3. Formation of the Wedge Serpentinite

[30] The Cl-, F- and Li-bearing wedge serpentinite is likely to form by fluid infiltration trenchward to the arc volcanic front. However, the joint infiltration of Cl, F, Li and heavy B, beneath the outer forearc is not supported by the data from the Mariana serpentinite seamounts. These data show that only B and possibly Cl are removed substantially at these depths ($\sim 20-25$ km). Note that despite being less saline than seawater (19500 ppm Cl) - the fluids upwelling beneath the outer Mariana forearc are still rich in Cl, with abundances ranging from 9900 to 18550 ppm Cl [Fryer et al., 1999; Benton et al., 2001], which could be due to Cl addition from slab. On the other hand, and despite the elevated Li of the Mariana serpentinites $(\sim 2-19 \text{ ppm } [Benton \ et \ al., 1999])$, the Mariana forearc fluids carry only minor Li ($\sim 1-2$ ppm [Ryan et al., 2001]) relative to the amounts present in the metacrust (several tens of ppm [Rvan and Langmuir, 1987]). This observation is consistent with the lower mobility of Li relative to B in hydrothermal experiments at 150° to 375°C [Sevfried at al., 1984]. Nothing yet is known about the F contents of these shallow fluids, but F is expected in slab fluids, since it becomes highly fluid-mobile at temperatures >250°C [Seyfried and Ding, 1995]. A simple explanation to reconcile these disparities was that the mobilities of Cl, F and Li in fluids increase with increasing slab depth, in order to account for their abundance in the wedge peridotite. In this context, the joint enrichment of Cl, F and Li might be significant, given the fact that these elements are all hosted in amphibole, which is the major metamorphic phase in the subducting crust [e.g., Tatsumi and Eggins, 1995]. In the metacrust, amphibole becomes unstable at \sim 70 to 80 km [Schmidt and Poli, 1998], still trenchward to the arc volcanic front. Hence amphibole should be an excellent source of Cl, Fand Li in slab fluids. However, the other phases, such as phengite, might release one or more of these elements as well, as they gradually release elements with depth [e.g., Bebout et al., 1999].

[31] Figure 10 attempts to summarize the above discussion in terms of "fluid release curves." Note that the lithospheric serpentinite is shown as an additional source of fluids to the wedge. We suggest that the release curves of B and Cl peak early, but then decline with increasing depth toward the rear arc. This is in agreement with models that propose that B and Cl become rapidly scavenged from the subducting slab during early subduction, which in turn prevents these elements from being recycled back into mantle [*Ito et al.*, 1983; *Ryan and Langmuir*, 1993; *Ryan et al.*, 1995, 1996; *Bebout et al.*, 1999; *Straub and Layne*, 2003]. In contrast to B and Cl, the fluid release



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Figure 10. Cartoon to illustrate the 'fluid release curves' discussed in the text. WS = wedge serpentinite. The lithospheric serpentinite is shown as an additional fluid source next to the metacrust. Three successive fluid release curves are suggested; for B and Cl (red), Li and possibly F (blue), and the fluid-mobile LILE (lilac). Water is implied to be released continuously from slab.

curves of Li and F, the latter being very speculative, are shown to increase with increasing slab depth, and, at least in the case of Li, may extend beyond arc front depth [Paquin and Altherr, 2002]. Despite the evidence for fluid removal of Li and F from the slab, however, it is important to keep in mind that the relative amount removed from the slab is probably minor (<10% of slab Li and F), as indicated by mass balance models [Ryan et al., 2001; Straub and Layne, 2003], and the study of metamorphic complexes [Bebout et al., 1999]. This might reflect the fact that most of F and Li are stored in anhydrous phases, replacing $\mathrm{Mg}^{\scriptscriptstyle +}$ and O²⁻ [e.g., Paquin and Altherr, 2002]. Their incorporation into the mineral structure should protect these elements from entrainment into fluids, and they may only become substantially mobilized in partial slab melts that form well beyond arc front depths [e.g., Bebout et al., 1999; Plank and Kelley, 2001]. Finally, the release of the fluid-mobile LILE, the curve is drawn as single line for simplicity, appears to be negligible trenchward from the Izu VF as well, and increases only significantly at arc front depths.

6. Some Implications of the Model Proposed

[32] The fluid release model proposed here has several implications. Importantly, it supports the studies of Bebout et al. [1999] and Ryan et al. [2001] in proposing that the loss of many fluidmobile LILE is insignificant during early slab devolatilization. In the case of the "cool" Izu Bonin subduction zone, this apparently includes Cs as well, whereby Cs might be mobilized to a larger extent in subduction zones with higher thermal gradients [Bebout et al., 1999]. The fluid-mobile LILE are largely sequestered in the hydrous minerals, such as phengite, lawsonite, epidote, clinozoisite and zoisite, that sequentially stabilize with progressive metamorphism [Schmidt, 1996; Becker et al., 1999; Melzer and Wunder, 2000]. Consequently, the elemental and isotopic signatures of even highly fluid-mobile elements like Cs, Pb or Sr should be largely preserved in the subducted materials until arc front depths. This validates many earlier studies on elemental recycling in arcs, in which this assumption is implicit [e.g., Elliott et al., 1997; Patino et al., 2000; Hochstaedter et al., 2001].

[33] Another aspect concerns the experimental studies of Keppler and Wyllie [1990] and Keppler [1996] that predict a critical role for the Cl and F of slab fluids. These studies suggest that, owing to their capacity for complexing, Cl- and F-bearing fluids are capable of fractionating the fluid-mobile LILE from Nb and Ta. They may also fractionate U from Th, generating a typical "arc signature" [Keppler and Wyllie, 1990; Keppler and Wyllie, 1991; Keppler, 1996]. At first glance, this hypothesis appears to be supported since the Cl- and Frich slab fluids beneath the Izu VF have exactly the chemical characteristics (elevated LILE/Nb and U/Th) predicted by this model [Hochstaedter et al., 2001; Straub et al., submitted manuscript, 2003]. However, the selective mobilization of halogens and Li documented by the HRA clearly does not support a key role for the halogens in



elemental fractionation. Rather, it suggests that slab devolatilization was a multistep process, during which fluid release was controlled by the fluid-mobility of the individual elements, resulting in the depth-dependent, en echelon fluid release curves depicted in Figure 10, and in the fractionation of Cl, F, Li and B from the other fluidmobile LILE. If the Cl flux indeed decreases, as indicated in Figure 10, the role of Cl in controlling the fractionation of the elements mobilized in the deeper fluids is even more doubtful. Thus without ruling out halogen-complexing completely, it appears that the sequential formation and decay of metamorphic mineral phases exerts a stronger role as to fractionating elements during slab devolatilization.

7. Conclusions

[34] Four primary conclusions may be drawn from this study.

[35] 1. The Izu VF volcanics contain a subordinate component melt that is selectively enriched in Cl, F and Li. Despite being preserved only in very rare melt inclusions this halogen-rich and esite ("HRA") component melt appears to be a common constituent of the Izu VF magmas.

[36] 2. The HRA component melt is interpreted to reflect the presence of a subordinate fluid component in the Izu VF magma source that is derived, at arc front depths, from the "wedge serpentinite" above the metacrust. The major portion of this fluid, which transports the fluid-mobile LILE (K, Sr, Rb, Cs, Ba, Pb, U), originates from the metacrust, with a possible contribution from the lithospheric serpentinite below the metacrust.

[37] 3. This interpretation implies that slab fluids entrain only a few elements (Cl, B, and minor amounts of F and Li) during shallow subduction, trenchward from the VF of the "cool" Izu subduction zone. Most of the fluid-mobile LILE (K, Sr, Rb, Cs, Ba, Pb, U) appear to be retained in the slab until arc front depths.

[38] 4. Elemental fractionation by halogen-complexing appears to play a minor role compared to the sequential formation and decay of metamorphic mineral phases during progressive slab metamorphism.

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