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DEIMOS: A beamline dedicated to dichroism measurements in the 350–2500 eV energy range

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The DEIMOS (Dichroism Experimental Installation for Magneto-Optical Spectroscopy) beamline was part of the second phase of the beamline development at French Synchrotron SOLEIL (Source Optimisée de Lumière à Energie Intermédiaire du LURE) and opened to users in March 2011. It delivers polarized soft x-rays to perform x-ray absorption spectroscopy, x-ray magnetic circular dichroism, and x-ray linear dichroism in the energy range 350–2500 eV. The beamline has been optimized for stability and reproducibility in terms of photon flux and photon energy. The main end-station consists in a cryo-magnet with 2 split coils providing a 7 T magnetic field along the beam or 2 T perpendicular to the beam with a controllable temperature on the sample from 370 K down to 1.5 K. © 2014 AIP Publishing LLC. [<http://dx.doi.org/10.1063/1.4861191>]

I. INTRODUCTION

The DEIMOS (Dichroism Experimental Installation for Magneto-Optical Spectroscopy) beamline is the beamline dedicated to soft x-ray dichroic measurements at SOLEIL (Source Optimisée de Lumière à Energie Intermédiaire du LURE). Such measurements are obtained by comparing two x-ray absorption (XA) spectra recorded for different sample states and/or different light polarizations. The photon source is constituted by two undulators installed on the I07-m medium straight section of the SOLEIL storage ring: a 1.65 m long Apple-II helical undulator¹ with a period of 52.4 mm, providing circularly polarized light (left and right) over the full energy range 260–2500 eV (with the first harmonic covering the range 260–1300 eV) and horizontal and vertical linearly polarized light from 350 eV to 2500 eV. The second undulator is a hybrid electromagnet/permanent magnet helical undulator (EMPHU)² with a period of 65.0 mm optimized for a fast switching (5 Hz) of the circular polarized light, with no access to the pure linear horizontal polarization. The 350–2500 eV energy range was chosen to cover the principal chemical elements linked to materials showing magnetic properties: the first (*3d*) and second (*4d*) row transition metal *L*-edges, the rare earth element *M*-edges and carbon, nitrogen, oxygen, and sulphur K-edges which are particularly interesting in the field of molecular magnetism.³ This beamline is devoted primarily to the studies of magnetic and electronic

states of nanoscale structures. XMCD (X-ray Magnetic Circular Dichroism) and XMLD (X-ray Magnetic Linear Dichroism) as well as XNCD (X-ray Natural Circular Dichroism) and XNLD (X-ray Natural Linear Dichroism) are indeed very sensitive techniques allowing to probe minute amounts of matter^{4,5} with the unique capability to provide information with chemical and orbital selectivity by measuring the appropriate absorption edge. Moreover, for some edges, XMCD can also be used to obtain quantitative information on the spin and orbital magnetic moments using specific sum rules.⁶ As the magnetic properties, particularly in nanoscale objects, are exacerbated by large magnetic field and low temperature, the main end-station is a superconducting cryo-magnet producing 7 T along the beam or 2 T perpendicular to the beam; the temperature on the sample can be set from 1.5 K to 370 K.⁷

In this paper, we will first present the optical layout of the beamline along with some results from the commissioning of the monochromator. Then we will concentrate on describing the technical developments specifically made to address our beamline demanding characteristics. These developments are (i) the control-command with the implantation of a fast XA spectra data acquisition; (ii) the hybrid undulator (EMPHU), and (iii) the use of a multilayer grating designed to reach the upper energy range (up to 2500 eV).

II. BEAMLINE IMPLEMENTATION

The schematic layout of the beamline, from the source to the experimental end-station, is shown in Figure 1; the main parameters are summarized in Table I (sources) and Table II (optics).

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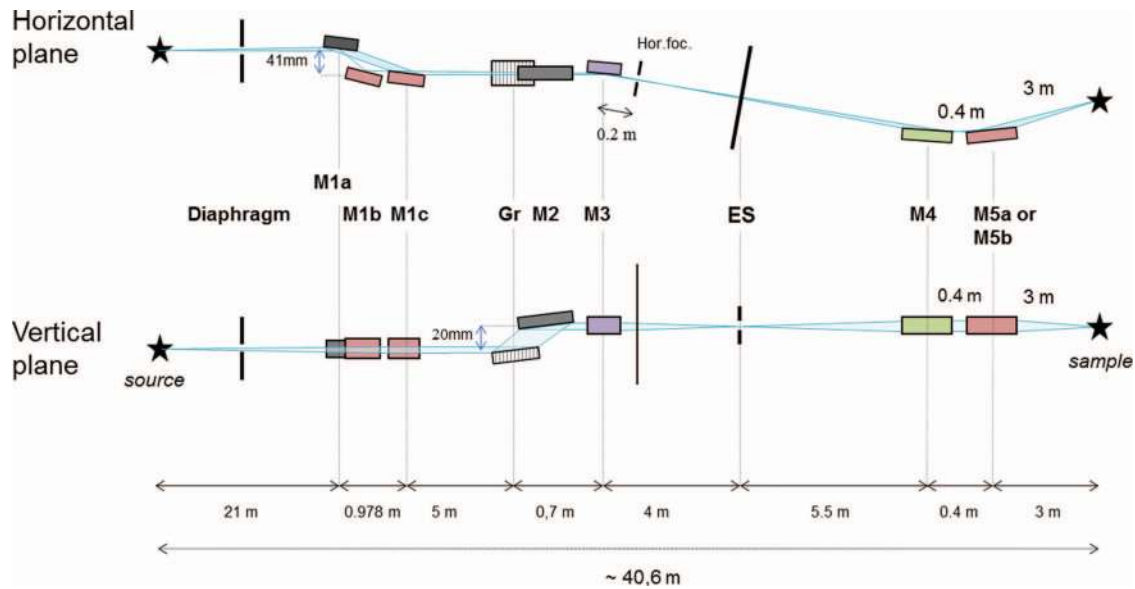


FIG. 1. Horizontal and vertical view of the schematic layout of the DEIMOS beamline from the source (on the left) to the sample (on the right).

A. Sources

The DEIMOS straight section (I07-m) is equipped with two undulators: an Apple-II HU-52 undulator and an EMPHU-65 (ElectroMagnet/Permanent magnet Helical Undulator) undulator.⁸

1. HU-52

The 52.4 mm period of the Apple-II undulator⁹ has been chosen in order to reach all polarizations (linear from horizontal to vertical and circular left and right) from 350 eV up to 2500 eV. In practice with the lower gap allowed (15.5 mm) it is possible to reach ≈ 320 eV in the linear vertical geometry and even smaller energy values with linear horizontal and circular polarizations. However, in order to minimize the deformations on the water cooled SiC flat mirror, the radiated power should be kept below ~ 160 W on the first optic. Therefore, only circular polarizations are allowed in the range 260–350 eV. A variable diaphragm is located at 13.09 m from the middle of the straight section and allows selecting the useful part of the polarized light.¹⁰ With an optimized diaphragm opening, on the first harmonic the circular polarization rate is 99% while on the second and third harmonic it is around 98%.¹⁰ The pink beam emitted from

the undulators can be visualized and characterized using the so-called DiagOn device.¹¹ This equipment, developed at SOLEIL, consists of a multilayer and a Si(111) single crystal, both reflecting upward the beam in Bragg diffraction on a scintillator screen (YAG:Ce). The multilayer and the crystal have been optimized to work for different ranges of undulator gaps according to their diffracted energy. The multilayer is made of 50 pairs of Cr(1 nm)/Sc(1.4 nm) layers and reflects the beam at 367.5 eV according to its fixed Bragg angle of 45° ; it is optimized to study the first two harmonics of the undulator set for linear horizontal (LH) polarized light. The Si(111) crystal, which reflects the beam at 2796 eV, is used to study the third to the seventeenth undulator harmonics (still with LH polarization). This makes the DiagOn an ideal tool to analyse the HU-52 undulator properties over a wide harmonic range.¹²

2. EMPHU-65

The purpose of the EMPHU-65 undulator is to supplement the HU-52 undulator by enhancing sensitivity.¹³ Indeed, the use of an electromagnet to generate the vertical magnetic field allows to switch rapidly (5 Hz repetition rate) from right to left circular polarization. The drawback of using coils is that emission of linear horizontal polarization is made impossible. Nevertheless, its ability to flip rapidly between circular polarizations makes it an ideal source to measure very small XMCD and XNCD signals. An alternative way of switching rapidly the polarization has been developed on other beamlines. It consists in the use of two identical undulators set at two different polarizations; the two sources are switched at each energy point of the spectrum.^{14,15} This approach implies that the beam follows different optical paths for each polarization, thus potentially leading to nonlinear effects which may jeopardize the measurements, in particular when looking for very small XMCD signals. By using the same optical path, the use of an EMPHU ensures a very high

TABLE I. Undulators parameters.

	HU-52	EMPHU-65
Device type	Apple-II	EMPHU
Technology	NdFeB magnets	NdFeB magnets + electromagnet
Max. magnetic field	$B_x = 0.53$ T $B_z = 0.76$ T	$B_x = 0.28$ T $B_z = 0.24$ T
Period length	52.4 mm	65.0 mm
Number of periods	30+2	24
Minimum gap	15.5 mm	14.7 mm
Phase-shift range	± 26.2 mm	± 350 A

TABLE II. Optical parameters.

Parameter	M1a	M1b	M1c	Gra	Grb	M2a	M2b	M3	M4	M5a	M5b
Shape	Flat	Toroid	Toroid	Flat	Flat	Flat	Flat	Cylinder	Sphere	Toroid	Toroid
Dist. from the straight section center (m)	21	21.470	21.978	27.448	27.448	~27.7 ^b	~27.7 ^b	28.148	37.648	38.048	38.048
Max. absorbed power (W)	~160	~15	~15	~1	~1	-	-	-	-
Dimension (mm)											
Length	220	130	230	70	50	290	200	40	340	320	320
Width	20	25	25	40	20	30	20	20	35	35	35
Thick	25	60	60	60	60	60	60	20	40	40	40
Beam footprint (mm ²)	184 × 4	92 × 4	184 × 4	60 × 1	35 × 1	30 × 1	30 × 1	20 × 10	320 × 3	300 × 3	300 × 3
RMS roughness (nm) ^a	0.32	0.14	0.27	0.6	0.25	0.10	0.23	0.15	0.10	0.20	0.16
RMS slope error tan. (μrad) ^a	0.8	2.6	2.8	0.23	0.6	0.4	0.6	1.0	1.1	1.3	0.7
RMS slope error sag. (μrad) ^a	0.4	0.7	1.91	...	0.7	0.8	1.1	2.2	3.0	4	2.5
Bulk material	SiC-CVD	Si	Si	Si	Si	Si	Si	Si	SiO ₂	SiO ₂	SiO ₂
Coating	Pt	Pt	Rh	Pt	[B ₄ C/Mo ₂ C] 2d = 5.36nm	Pt	[B ₄ C/Mo] 2d = 5.35nm	Pt	Pt	Pt	Pt
Total deflection angle (°)	5.06° or 2.4°	5.06°	2.4°	Variable	Variable	Variable	Variable	2.4°	2.4°	2.4°	2.4°
Tangential radius ^a	-19.5 km ^c	233.03 m	455.5 m	-3.7 km ^c	6.7 km	360 km	42.3 km	592.8 m	296.2 m	955 m	2.9 km
Sagittal radius ^a	3.39 km	1894 mm	924 mm	...	2.8 km	1.7 km	1.8 km	167.8 mm	188.1 m	83.6 mm	150.3 mm
Line density (l/mm)	1600	2400	-	-	-	-
Groove depth (nm) ^a	3.6 to 15.4 on 28 mm	≈2.48	-	-	-	-
Groove width to period ratio	0.43-0.56	0.46	-	-	-	-

^aMeasured values.^bVaries with the energy and the factor *c*.^cA minus sign denotes a convex surface.

reproducibility and stability of the beam and consequently an improved sensitivity in the dichroic measurements (see Ref. 16 for some examples of dichroism in absorption experiments performed at the very edge of this technique using an EMPHU undulator).

B. M1 mirrors chamber

The first optical chamber allows choosing between two optical paths, M1a-M1b or M1a-M1c, by translating in and out of the optical path the M1b mirror. The mirrors M1b and M1c differ by their coatings, platinum, and rhodium, respectively. The M1a-M1b path is optimized for high energy harmonics rejection below 1300 eV due to a working angle of 2.53°, whereas M1a-M1c is used for higher energies (working angle on the Rh coated M1c mirror: 1.20°) (see the cross curve on Figure 4(a)). For the maximum angular acceptance of 0.2×0.2 mrad² defined by the variable diaphragm, the footprints on M1a are 92×4 mm² and 184×4 mm² for respectively the M1a-M1b and M1a-M1c trajectories. The power coming on the first optics (maximum of 160 W for an angular acceptance of 0.2×0.2 mrad² according to the calculation using SRW¹⁷) is dissipated using a water cooled SiC-CVD mirror (SiC-CVD from *Rohm and Haas*

having a thermal conductivity of 300 W/mK). The deformation due to the thermal heat load has been evaluated by finite element analysis and input to the optical ray tracing.¹⁸ The comparison of these calculations to the perfectly flat mirror case show that, at worse, the resolving power is reduced by ~6% and the horizontal beam size is increased by a factor 1.3 ± 0.1 .

The horizontal deflecting toroidal mirror (M1b or M1c) focus this light in the horizontal plane to a point located 200 mm after the M3 mirror (inside the monochromator chamber) while in the vertical direction the beam is collimated in an almost parallel beam.

C. Plane grating monochromator (PGM)

The monochromator is equipped with two plane gratings, Gra and Grb, which are respectively a variable groove depth (VGD)¹⁹ grating with 1600 lines/mm, and a Mo₂C/B₄C alternate multilayer (AML) grating with 2400 lines/mm.²⁰⁻²² The VGD grating is optimized for the energy range 240-1500 eV while the AML grating is optimized for the energy domain 1000-2500 eV. The efficiency of both gratings is detailed in Sec. III C. The dispersed light is vertically deflected onto the associated M2 plane mirror (M2a or M2b in Table II); note

that in the case of the AML grating, M2b is a multilayer mirror with a matched period. The beam is then deviated in the horizontal plane by a cylinder mirror M3 which focuses the beam on the exit slit (see Fig. 1).

With the VGD grating the monochromator is usually operated in the Petersen geometry with a fixed c factor.²³ This c factor is defined as: $c = \sin \beta / \sin \alpha$, where α is the grazing incidence angle of the light on the grating and β the grazing diffraction angle. The monochromator is working in the first outside diffraction order, with the grating as the first optics; hence the c values are below 1. Usable c factor values are comprised between 0.05 and 0.4; the smaller c values correspond to the higher dispersion, i.e., to the higher resolving power for a given exit slit (and consequently a lower transmitted flux). The high groove density of the VGD grating (1600 lines/mm) has been chosen to improve the energy stability. As a matter of fact, achieving a given energy resolution with a large grating dispersion and a moderately closed exit slit reduces the influence of the mechanical angular accuracy (of both the grating and the mirror) onto the energy accuracy. Nevertheless, it is also possible to achieve large resolving power if one decides to work with a closed exit slit, and the ability of changing the c factor gives further flexibility.

In order to determine the resolving power of the beamline, resolution measurements have been performed using a gas cell located between the exit slit and the Wolter chamber²⁴ (mirror pair M4-M5). In this cell a stable pressure (around 1×10^{-3} mbar) is established with N_2 or Ne gas. An electrode inside this chamber collects the electrons emitted during the photon absorption; the current recorded gives a good measurement of the x-ray absorption spectroscopy (XAS). On Figure 2, we plot a typical K -edge XA spectrum for N_2 .

From such XA spectrum, using a Lorentzian function decomposition, we can determine the variations of the full width at half maximum as a function of the c factor and the opening of the exit slits. The use of the K -edge of N_2 or Ne allows characterizing this dependence at respectively 410 and 870 eV. Using this method, and knowing the natural width of

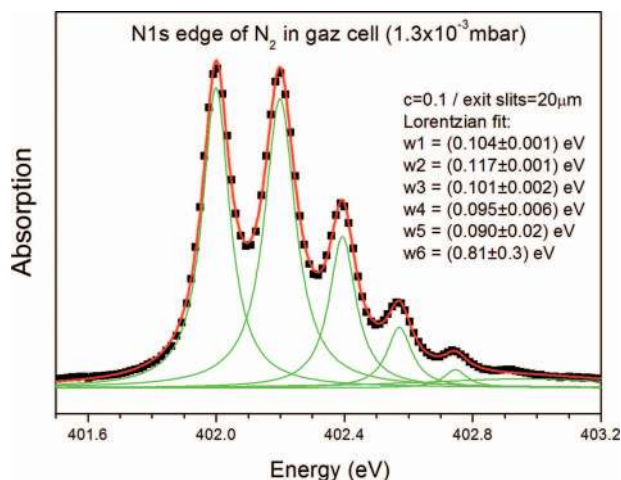


FIG. 2. K -edge XA spectrum for N_2 with c set to 0.1 and the exit slits open at $20 \mu\text{m}$; dotted curve: experimental data, red curve: total spectrum of 6 Lorentzians, and green curves: individual Lorentzian contributions.

the K -edges (i.e., 0.123 eV and 0.230 eV for respectively N_2 and Ne^{25}), we verified that the resolving power of the VGD grating is within 8% of the theoretical value, i.e., an ultimate resolving power between 30 000 and 35 000 for the energy range 350–1500 eV and for a factor c of 0.2.

The energy stability was measured using a reference Fe_3O_4 single crystal sample, which is chemically very stable under x-rays and provides very sharp $L_{2,3}$ edges mandatory for such measurement. The principle consists in recording over several hours the same spectrum in exactly the same conditions (light polarization and sample state). Then, spectra taken at different moments (separated by 1–2 h) are subtracted to each other and the resulting signal (ideally 0) is compared to the mathematical derivative of the average of all measurements. Using this method, we could measure a stability in energy of ≈ 1 meV/h, which is a very competitive value. This value is comparable to the drift in energy (read from the encoders' rotations of the monochromator) due to the variations in temperature produced by the day-night cycles, although the monochromator has its own hutch with a dedicated air-conditioner which keeps the temperature variations below $\pm 0.5^\circ\text{C}$. This means that the energy stability is mainly limited, in our case, by the thermal stability of the monochromator. It should nevertheless be noted that the monochromator does not have an entrance slit to define the source, but rather uses the image of the undulator source reflected by the M1 mirror pair. This means that the energy stability can also be affected by temperature variation of the first optics hutch and machine stability.

D. Wolter mirror pair

The last optical stage is a simplified Wolter mirror combination associating a spherical mirror M4 and a toroidal mirror M5 as approximations of the ellipsoidal shapes of the canonical Wolter combination.²⁴ The two successive reflections in the horizontal plane are used to cancel the coma aberration in this plane, and hence accommodate the large image aperture which results from the large horizontal demagnification imposed by SOLEIL horizontal source size. Two toroidal mirrors (M5a and M5b) are available and can be switched under vacuum in order to produce two different beam sizes in the end-station located 3 m downstream. While M5a exactly focuses on the experimental point, M5b is focusing somewhat beyond in order to increase by a factor of ~ 10 the spot size on the sample, both horizontally and vertically. This divides by ~ 100 the photon density on the sample going from a spot size of $\sim 80 \times 80 \mu\text{m}^2$ for M5a to $\sim 800 \times 800 \mu\text{m}^2$ for M5b. Additionally, due to its angle tolerant design, the Wolter permits to move the beam on the sample in a range of a few millimeters.

E. End-station

The end-station consists of a superconducting magnet providing ± 7 T along the x-rays beam or ± 2 T perpendicular to the beam. The variable temperature insert (VTI) has been developed for temperatures in the 1.5–370 K range.⁷ The absorption signal can be recorded in 3 different ways: (i) in

transmission mode (for semi-transparent samples) with a removable diode²⁶ located at the back of the end-station chamber, (ii) in total electron yield, and (iii) in fluorescence yield. For all these modes, the currents are measured with *Novelec* electrometers connected to a counting card NI-6602 from *National Instruments*.²⁷

The end-station is connected to a glove-box and to a set of ultra-high vacuum (UHV, $<10^{-9}$ mbar) chambers dedicated to *in situ* sample conditioning, including sample preparation and characterization under inert atmosphere (standard wet chemistry techniques) and under UHV conditions (evaporation, Ar⁺ sputtering, annealing, Auger electron spectroscopy, low energy electron diffraction, and scanning tunneling microscopy).

III. SPECIFIC DEVELOPMENTS

A. Turboscan (on-the-fly energy scan)

As mentioned above, the DEIMOS beamline has been optimized for dichroic measurements obtained by comparison of XA spectra recorded with different polarizations and/or different magnetic/electronic states. This statement implies that for such measurements a very good stability and fast acquisition process are necessary in order to minimize the impact of the residual instabilities due to the experimental environment. This fast data acquisition is a common issue on beamlines performing scans in energy. In a conventional step-by-step approach, the monochromator and the insertion device are sequentially driven to the desired photon energies. The limiting factor of this approach is the time needed to accurately position all the motors: gratings and mirrors cradles in the monochromator as well as gap and phase motors of the insertion device, as their acceleration and deceleration are time consuming.

An alternative approach is to use an on-the-fly energy scanning mode. However, rigorous approach requires sophisticated multi-axis motion control with nonlinear motion trajectories; this is still part of an on-going development at SOLEIL. In the meantime, we have thus chosen to rely on a software based solution²⁸ using the available Tango tools (the so-called Tango is an object oriented distributed control system used at SOLEIL²⁹). It consists in a software synchronization of the monochromator and the undulator motors by determining proper initial position, final position, and speed for each motor.

From the user point of view, the parameters are the start (E_S) and final (E_F) energies, the total time (t) and the integration time (dt). Using these values (E_S , E_F , and t), the trajectories (initial, final positions, and speed) of the 2 rotations of the monochromator and the gap of the undulator are determined. When the on-the-fly scan is launched all these motor positions are recorded by a counting card in a synchronous fashion with the integration time dt . The trajectory in energy is then reconstructed from these known motor positions. To improve the synchronization between the energy of the monochromator and the energy of the undulator, a master-slave dynamical coupling is established between the two devices. In practice, as the photon energy on the sample is always given by the

monochromator, the latter is given the role of “master.” During the scan, every 500 ms, the position in energy of the insertion device is compared to the position of the monochromator and its speed is adjusted in order to minimize this offset.

Compared to a step-by-step XA spectrum, this on-the-fly spectrum allows a gain of ≈ 10 in time with similar statistics and resolution in energy. In addition, we have checked several important points of these on-the-fly scans: (i) *c factor*: as explained previously, the monochromator is normally operated with a fixed c factor. But during the on-the-fly scan, the coupling between grating and M2 mirror rotations is not ensured. Nevertheless, we can calculate at the end of a scan the actual c factor from the α and β angle values recorded at each measurement point, and it appears that it never deviates by more than 0.5% from the desired value. This deviation is low enough to claim that acquisition conditions are similar to a step-by-step scan. (ii) *Monochromator-undulator synchronization*: the feedback during the scan allows reducing the offset between the monochromator and undulator to values as low as 3 eV. Compared to the typical width of the undulator peak (15–40 eV), we can conclude that the synchronization of the undulator and the monochromator is very good, even for a wide scan of several hundreds of eV. (iii) *Polarization*: the phase is not adjusted during the on-the-fly scan. In the case of circular polarized light, a constant phase over an energy range of ~ 100 eV means a loss in the polarization rate limited to few percents. Such deviation is well below the typical errors bars on a dichroic measurement.

In conclusion, the on-the-fly scan in energy is a powerful tool to perform XA spectra in a fast and accurate way. The typical time to acquire a scan of ~ 100 eV width is 2–3 min.

B. EMPHU-65

Another elegant way to perform dichroic measurements in absorption is to record at each point in energy the dichroic signal. In the case of XMCD, this is obtained by flipping the circularly polarized light direction at each point in energy. To perform such measurements, we have chosen a second undulator which uses a combination of permanent magnets and electromagnets. The electromagnets allow flipping the circular polarization at a rate of 5 Hz. Practically, this means that at each energy point of an absorption spectrum the signal is measured for the two circular polarizations which are switched within 0.2 s. This very short delay between the two measurements minimizes the impact of all the measurement instabilities and offers the unique capability to measure very weak dichroic signals.

While tuning such a device the beam steering in the undulator is a critical parameter, and should be identical when switching from one polarization state to the other. This has been monitored using the diaphragm as a slit to image the beam intensity profile, and it allowed to define the fine parameters of the insertion device (correction tables). As a final check, the beam on the sample ($\sim 80 \times 80 \mu\text{m}^2$) has been imaged on a YAG:Ce screen and no change in position has been observed when flipping the polarization, within the resolution of our measurement (10 μm).

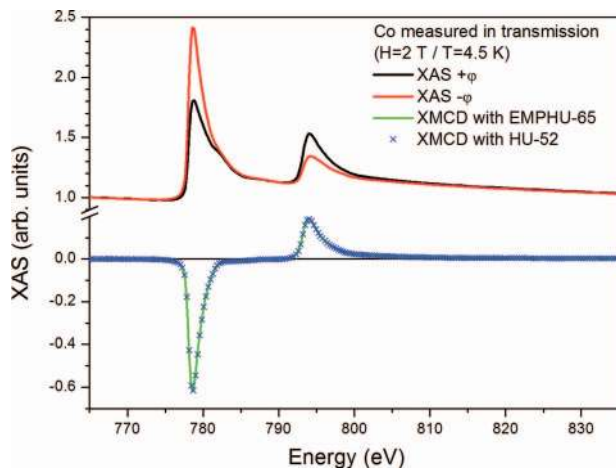


FIG. 3. XAS and XMCD at the $L_{2,3}$ edges of Co. The solid lines correspond to the measurements obtained using the EMPHU-65 while the dotted XMCD spectrum is given as a reference and has been obtained using the HU-52 in the same conditions.

A second important issue is the polarization symmetry between circular left and right. The calculations using the magnetic measurements predicted polarization rates of 99.99% and 99.82% for respectively the right and left circularly polarized light (in the case of a gap of 27.5 mm corresponding to ~ 800 eV). The first XMCD spectra obtained on a Co sample confirm this statement: the deduced circular polarization rate indicates a circular polarization rate above 99% (Figure 3). Indeed, the XMCD signal has exactly the same amplitude as the one obtained with the HU-52, leading to the magnetic moment expected from the sum rules. These results confirm the good performance of the EMPHU-65.

C. Multilayer grating

The AML grating has been developed to improve significantly the efficiency of the monochromator at high energies. Lamellar gratings have rather low efficiencies at energies above ~ 1500 eV and DEIMOS VGD grating is not optimized to work at energies higher than ~ 1200 eV. Blazed gratings with triangular line profiles are a known way of overcoming this limitation:³⁰ they offer high efficiency in the 2 keV energy range when the blaze angle (angle of the triangular facet with the grating plane) is in the order of $1-2^\circ$, and they can be coupled with a grazing incidence mirror to offer a good throughput. Nevertheless, the shallow angle of the profile and the manufacturing precision it requires make the realization of such gratings difficult to master on a large optical surface. Moreover, they are not suitable for holographic fabrication and generally rely on a mechanical ruling with all associated regularity issues. For these reasons SOLEIL, together with Horiba Scientific (formerly Horiba Jobin-Yvon SAS), have been working in the last years on an alternative technology named AML grating.³¹ The basic idea is to deposit alternate layers of two materials on a lamellar grating substrate in order to synthesize a 2D periodic structure in the plane perpendicular to the grating lines. The AML grating has properties similar to crystals used in high energy double crystal monochromator, namely only waves satisfying a Bragg condition are

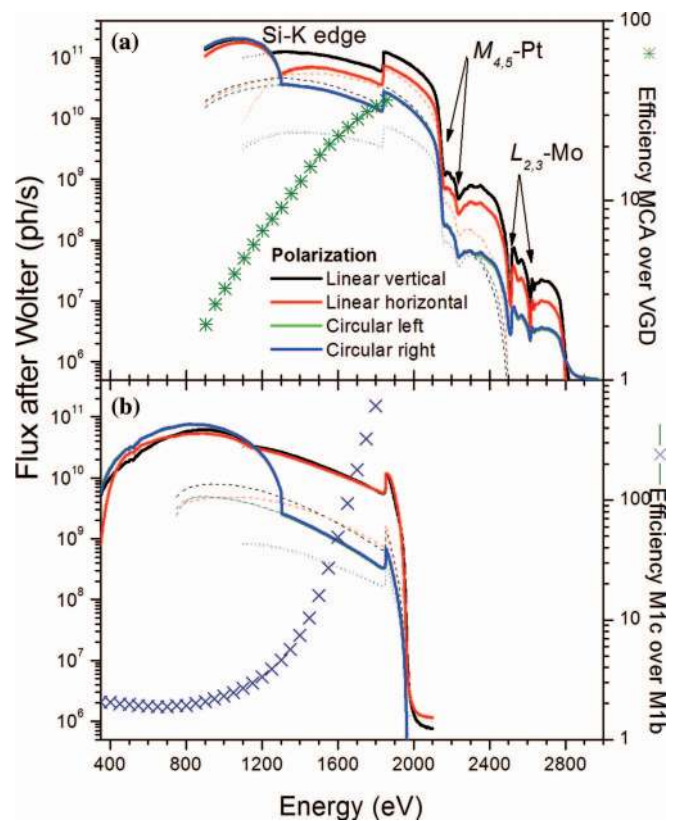


FIG. 4. Flux measured after the Wolter chamber for different light polarizations and a ring current of 430 mA. (a) Using the AML grating (exit slit set to $200 \mu\text{m}$) and (b) the VGD grating (exit slit set to $100 \mu\text{m}$). The 1st harmonic goes up to ~ 1300 eV and the 2nd (dashed lines) and 3rd harmonics (dotted lines) are used to go to higher energies value. The symbols -x- in blue (b) represent the efficiency of the M1a-M1c optical path over M1a-M1b while the green *- (a) illustrate the efficiency of the AML grating over the VGD. The Si-K edge is due to the material of the diode while the Mo- $L_{2,3}$ and Pt- $M_{4,5}$ edges correspond to the optical coating materials.

reflected. This enhances the diffraction efficiency in the allowed orders. In a monochromator, the AML grating should be paired with a mirror coated with a similar multilayer which compensates the grating deviation.²⁰ DEIMOS AML grating and multilayer mirror pair is the first to be fully commissioned for operation at SOLEIL.

DEIMOS AML grating Grb was fabricated by deposition of 30 periods of a $[\text{Mo}_2\text{C}/\text{B}_4\text{C}]$ multilayer coating, on a silicon lamellar grating substrate etched with a line density of 2400 mm^{-1} . The materials were chosen for allowing high reflectivity around 2 keV and their ability to form clean interfaces, almost without any interdiffusion. The layer period, 5.36 nm, was selected to match the grating groove depth and results in a Bragg angle of about 0.69° . This grating/mirror pair has an optimal efficiency before Mo- L edges in the 1000–2450 eV range. The details of the fabrication and the choice of materials for the AML, as well as the first characterization, are given elsewhere (see Refs. 20 and 21) and we will concentrate in this paper on the gain in terms of photon flux.

The flux was measured, using a calibrated AXUV-100 photodiode,³² for the 2 optical pathways M1a-M1b and M1a-M1c in the case of the classical VGD grating and with the AML grating. In both cases, the aperture was 0.1×0.1 mrad

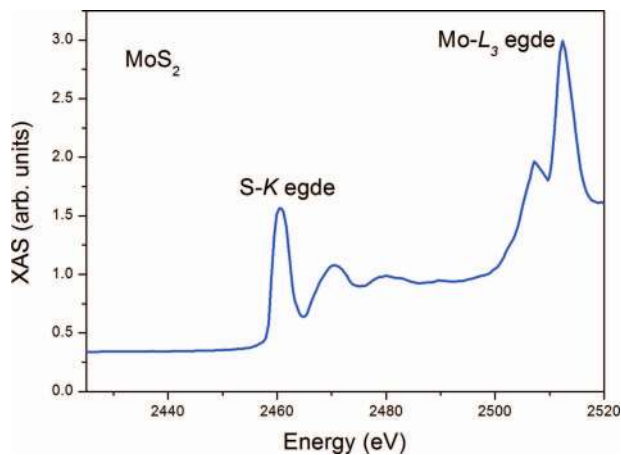


FIG. 5. X-ray absorption spectrum of S-K and Mo-L₃ edges for MoS₂.

and only the settings of the monochromator were different: for the VGD the *c* factor was set to 0.2 and the exit slit to 100 μm, to be close to the usual settings, while for the AML grating the exit slit was set to 200 μm to improve the flux at the cost of the ultimate resolution. Flux measurements have been done for all 4 standard light polarizations produced by the HU-52, using the first three harmonics to cover an energy range from 350 to 3000 eV (see Figure 4).

On the same figure, we also represent different efficiencies (plot with the large symbols): the efficiency of the M1a-M1c optical path over M1a-M1b and more interestingly the AML grating efficiency over the VGD. From this curve, we see that the AML grating offers a good efficiency from ~1000 eV and that above 2000 eV it is the only way to get some flux on the sample. Although there is a strong flux drop above 2000 eV, due to the Pt and Mo mirrors coatings, the AML grating still provides a reasonable flux and we are able to perform absorption measurements at this energy as can be seen from the x-ray absorption spectrum of MoS₂ recorded at S-K and Mo-L₃ edges displayed in Figure 5.

IV. CONCLUSION AND PROSPECTS

The DEIMOS beamline has been optimized for XAS and all types of dichroic measurements in the 350–2500 eV energy range. In order to reach such high energy on a soft X-ray beamline, we developed AML gratings. The capability to record very small XMCD signal has been improved by the implantation of the EMPHU insertion device, first hybrid electromagnet/permanent magnet undulator build to work in the soft energy range. Other instrumental developments are in progress: (i) additionally to the standard VTI we are developing an insert with 12 current leads to the sample. These current leads can have several purposes, such as, for instance, the application of an electrical potential difference across the sample, or the possibility to perform 4-points resistance measurements simultaneously to XMCD measurements. (ii) We have started the development of a third dedicated VTI for measurements from 200 to 1000 mK using a ³He-⁴He dilution refrigerator. (iii) We will install by the end of the year a new end-station consisting of an electromagnet with a VTI dedicated to measure in the high temperature range from 20

to 1000 K. (iv) With the growing interest of time resolved magnetic studies we are now considering the possible development of a second branch line dedicated to time resolved measurements in the 0.1 ps scale using the sliced beam.³³ This branch would be optimized for providing the maximum flux on the sample knowing that the slicing scheme means the source flux will be reduced by 6 orders of magnitude.

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