

# Development, characterisation, and deployment of the SNO+ liquid scintillator



## SNO+ collaboration

M. R. Anderson<sup>a</sup> S. Andringa<sup>b</sup> L. Anselmo<sup>c</sup> E. Arushanova<sup>d</sup> S. Asahi<sup>a</sup> M. Askins<sup>e,f,g</sup>  
 D. J. Auty<sup>h</sup> A. R. Back<sup>d,i</sup> Z. Barnard<sup>j</sup> N. Barros<sup>b,k,l,m</sup> D. Bartlett<sup>a</sup> F. Barão<sup>b,n</sup> R. Bayes<sup>j</sup>  
 E. W. Beier<sup>l</sup> A. Bialek<sup>c,j,h</sup> S. D. Biller<sup>o</sup> E. Blucher<sup>p</sup> R. Bonventre<sup>e,f,l</sup> M. Boulay<sup>a</sup> D. Braid<sup>j</sup>  
 E. Caden<sup>c,j,a</sup> E. J. Callaghan<sup>e,f</sup> J. Caravaca<sup>e,f</sup> J. Carvalho<sup>q</sup> L. Cavalli<sup>o</sup> D. Chauhan<sup>c,j,b,a</sup>  
 M. Chen<sup>a</sup> O. Chkvorets<sup>j</sup> K. J. Clark<sup>o,a,i</sup> B. Cleveland<sup>c,j</sup> D. Cookman<sup>o</sup> C. Connors<sup>j</sup>  
 I. T. Coulter<sup>o,l</sup> M. A. Cox<sup>r,b</sup> D. Cressy<sup>j</sup> X. Dai<sup>a</sup> C. Darrach<sup>j</sup> B. Davis-Purcell<sup>s</sup> C. Deluce<sup>c,j</sup>  
 M. M. Depatie<sup>j</sup> F. Descamps<sup>e,f</sup> J. Dittmer<sup>m</sup> F. Di Lodovico<sup>d,t</sup> N. Duhaime<sup>j</sup> F. Duncan<sup>c,j,†</sup>  
 J. Dunger<sup>o</sup> A. D. Earle<sup>i,†</sup> D. Fabris<sup>c</sup> E. Falk<sup>i</sup> A. Farrugia<sup>j</sup> N. Fatemighomi<sup>c,a</sup> V. Fischer<sup>g</sup>  
 E. Fletcher<sup>a</sup> R. Ford<sup>c,j</sup> K. Frankiewicz<sup>u</sup> N. Gagnon<sup>c</sup> A. Gaur<sup>h</sup> K. Gilje<sup>h</sup>  
 O. I. González-Reina<sup>ab</sup> D. Gooding<sup>u</sup> P. Gorel<sup>h</sup> K. Graham<sup>a</sup> C. Grant<sup>u,g</sup> J. Grove<sup>j</sup> S. Grullon<sup>l</sup>  
 E. Guillian<sup>a</sup> S. Hall<sup>c</sup> A. L. Hallin<sup>h</sup> D. Hallman<sup>j</sup> S. Hans<sup>v</sup> J. Hartnell<sup>i</sup> P. Harvey<sup>a</sup>  
 M. Hedayatipour<sup>h</sup> W. J. Heintzelman<sup>l</sup> J. Heise<sup>a</sup> R. L. Helmer<sup>s</sup> D. Horne<sup>a</sup> B. Hreljac<sup>a,j</sup> J. Hu<sup>h</sup>  
 A. S. M. Hussain<sup>j</sup> T. Iida<sup>a</sup> A. S. Inácio<sup>b,k</sup> C. M. Jackson<sup>e,f</sup> N. A. Jelley<sup>o</sup> C. J. Jillings<sup>n,j</sup>  
 C. Jones<sup>o</sup> P. G. Jones<sup>o,d</sup> K. Kamdin<sup>e,f</sup> T. Kaptanoglu<sup>e,f,l</sup> J. Kaspar<sup>w</sup> K. Keeter<sup>x</sup>  
 C. Kefelian<sup>e,f</sup> P. Khaghani<sup>j</sup> L. Kippenbrock<sup>w</sup> J. R. Klein<sup>l</sup> R. Knapik<sup>y,l</sup> J. Kofron<sup>w</sup>  
 L. L. Kormos<sup>z</sup> S. Korte<sup>j</sup> B. Krar<sup>a</sup> C. Kraus<sup>j,a</sup> C. B. Krauss<sup>h</sup> T. Kroupova<sup>o</sup> K. Labe<sup>p</sup>  
 F. Laffleur<sup>c</sup> I. Lam<sup>a</sup> C. Lan<sup>a</sup> B. J. Land<sup>l,e,f</sup> R. Lane<sup>d</sup> S. Langrock<sup>d</sup> A. LaTorre<sup>p</sup> I. Lawson<sup>c,j</sup>  
 L. Lebanowski<sup>l</sup> G. M. Lefevre<sup>i</sup> E. J. Leming<sup>o,i</sup> A. Li<sup>u</sup> J. Lidgard<sup>o</sup> B. Liggins<sup>d</sup> Y. H. Lin<sup>c,j</sup>  
 X. Liu<sup>a</sup> Y. Liu<sup>a</sup> V. Lozza<sup>b,k,m</sup> M. Luo<sup>l</sup> S. Maguire<sup>v</sup> A. Maio<sup>b,k</sup> K. Majumdar<sup>o</sup> S. Manecki<sup>c,a,j</sup>  
 J. Maneira<sup>b,k</sup> R. D. Martin<sup>a</sup> E. Marzec<sup>l</sup> A. Mastbaum<sup>p,l</sup> J. Mauel<sup>a</sup> N. McCauley<sup>r</sup>  
 A. B. McDonald<sup>a</sup> P. Mekarski<sup>h</sup> M. Meyer<sup>m</sup> C. Miller<sup>a</sup> C. Mills<sup>i</sup> M. Mlejnek<sup>i</sup> E. Mony<sup>a</sup>  
 I. Morton-Blake<sup>o</sup> M. J. Mottram<sup>d,i</sup> S. Nae<sup>b,k</sup> M. Nirkko<sup>i</sup> L. J. Nolan<sup>d</sup> V. M. Novikov<sup>a</sup>  
 H. M. O’Keeffe<sup>z,a</sup> E. O’Sullivan<sup>a</sup> G. D. Orebi Gann<sup>e,f,l</sup> M. J. Parnell<sup>z</sup> J. Paton<sup>o</sup>  
 S. J. M. Peeters<sup>i</sup> T. Pershing<sup>g</sup> Z. Petriw<sup>h</sup> J. Petzoldt<sup>m</sup> L. Pickard<sup>g,1</sup> D. Pracsovics<sup>j</sup> G. Prior<sup>b</sup>  
 J. C. Prouty<sup>e,f</sup> S. Quirk<sup>a</sup> A. Reichold<sup>o</sup> S. Ricetto<sup>a</sup> R. Richardson<sup>j</sup> M. Rigan<sup>i</sup> A. Robertson<sup>r</sup>  
 J. Rose<sup>r</sup> R. Rosero<sup>v</sup> P. M. Rost<sup>j</sup> J. Rumleskie<sup>j</sup> M. A. Schumaker<sup>j</sup> M. H. Schwendener<sup>j</sup>  
 D. Scislawski<sup>w</sup> J. Secrest<sup>aa,l</sup> M. Seddighin<sup>a</sup> L. Segui<sup>o</sup> S. Seibert<sup>l</sup> I. Semeneč<sup>a,j</sup> F. Shaker<sup>h</sup>  
 T. Shantz<sup>j</sup> M. K. Sharma<sup>h</sup> T. M. Shokair<sup>l</sup> L. Sibley<sup>h</sup> J. R. Sinclair<sup>i</sup> K. Singh<sup>h</sup> P. Skensved<sup>a</sup>

**M. Smiley<sup>e,f</sup> T. Sonley<sup>a</sup> R. Stainforth<sup>r</sup> M. Strait<sup>p</sup> M. I. Stringer<sup>d,i</sup> R. Svoboda<sup>g</sup> A. Sørensen<sup>m</sup>  
B. Tam<sup>a,1</sup> J. Tatar<sup>w</sup> L. Tian<sup>a</sup> N. Tolich<sup>w</sup> J. Tseng<sup>o</sup> H. W. C. Tseung<sup>w</sup> E. Turner<sup>o</sup> R. Van Berg<sup>l</sup>  
J. G. C. Veinot<sup>h</sup> C. J. Virtue<sup>j</sup> B. von Krosigk<sup>m</sup> E. Vázquez-Jáuregui<sup>ab,j</sup> J. M. G. Walker<sup>r</sup>  
M. Walker<sup>a</sup> S. C. Walton<sup>j</sup> J. Wang<sup>o</sup> M. Ward<sup>a</sup> O. Wasalski<sup>s</sup> J. Waterfield<sup>i</sup> J. J. Weigand<sup>m</sup>  
R. F. White<sup>i</sup> J. R. Wilson<sup>d,t</sup> T. J. Winchester<sup>w</sup> P. Woosaree<sup>j</sup> A. Wright<sup>a</sup> J. P. Yanez<sup>h</sup> M. Yeh<sup>v</sup>  
T. Zhang<sup>g</sup> Y. Zhang<sup>h</sup> T. Zhao<sup>a</sup> K. Zuber<sup>m,ac</sup> A. Zummo<sup>l</sup>**

<sup>a</sup>Queen's University, Department of Physics, Engineering Physics & Astronomy, Kingston, ON K7L 3N6, Canada

<sup>b</sup>Laboratório de Instrumentação e Física Experimental de Partículas (b), Av. Prof. Gama Pinto, 2, 1649-003, Lisboa, Portugal

<sup>c</sup>SNOLAB, Creighton Mine #9, 1039 Regional Road 24, Sudbury, ON P3Y 1N2, Canada

<sup>d</sup>Queen Mary, University of London, School of Physics and Astronomy, 327 Mile End Road, London, E1 4NS, UK

<sup>e</sup>University of California, Berkeley, Department of Physics, CA 94720, Berkeley, USA

<sup>f</sup>Lawrence Berkeley National Laboratory, 1 Cyclotron Road, Berkeley, CA 94720-8153, USA

<sup>g</sup>University of California, Davis, 1 Shields Avenue, Davis, CA 95616, USA

<sup>h</sup>University of Alberta, Department of Physics, 4-181 CCIS, Edmonton, AB T6G 2E1, Canada

<sup>i</sup>University of Sussex, Physics & Astronomy, Pevensey II, Falmer, Brighton, BN1 9QH, UK

<sup>j</sup>Laurentian University, Department of Physics, 935 Ramsey Lake Road, Sudbury, ON P3E 2C6, Canada

<sup>k</sup>Universidade de Lisboa, Faculdade de Ciências (FCUL), Departamento de Física, Campo Grande, Edifício C8, 1749-016 Lisboa, Portugal

<sup>l</sup>University of Pennsylvania, Department of Physics & Astronomy, 209 South 33rd Street, Philadelphia, PA 19104-6396, USA

<sup>m</sup>Technische Universität Dresden, Institut für Kern und Teilchenphysik, Zellescher Weg 19, Dresden, 01069, Germany

<sup>n</sup>Universidade de Lisboa, Instituto Superior Técnico (IST), Departamento de Física, Av. Rovisco Pais, 1049-001 Lisboa, Portugal

<sup>o</sup>University of Oxford, The Denys Wilkinson Building, Keble Road, Oxford, OX1 3RH, UK

<sup>p</sup>The Enrico Fermi Institute and Department of Physics, The University of Chicago, Chicago, IL 60637, USA

<sup>q</sup>Universidade de Coimbra, Departamento de Física and Laboratório de Instrumentação e Física Experimental de Partículas (LIP), 3004-516, Coimbra, Portugal

<sup>r</sup>University of Liverpool, Department of Physics, Liverpool, L69 3BX, UK

<sup>s</sup>TRIUMF, 4004 Wesbrook Mall, Vancouver, BC V6T 2A3, Canada

<sup>t</sup>King's College London, Department of Physics, Strand Building, Strand, London, WC2R 2LS, UK

<sup>u</sup>Boston University, Department of Physics, 590 Commonwealth Avenue, Boston, MA 02215, USA

<sup>v</sup>Brookhaven National Laboratory, Chemistry Department, Building 555, P.O. Box 5000, Upton, NY 11973-500, USA

---

<sup>1</sup>Corresponding Author

<sup>w</sup> *University of Washington, Center for Experimental Nuclear Physics and Astrophysics, and Department of Physics, Seattle, WA 98195, USA*

<sup>x</sup> *Idaho State University, 921 S. 8th Ave, Mail Stop 8106, Pocatello, ID 83209-8106*

<sup>y</sup> *Norwich University, 158 Harmon Drive, Northfield, VT 05663, USA*

<sup>z</sup> *Lancaster University, Physics Department, Lancaster, LA1 4YB, UK*

<sup>aa</sup> *Armstrong Atlantic State University, 11935 Abercorn Street, Savannah, GA 31419, USA*

<sup>ab</sup> *Universidad Nacional Autónoma de México (UNAM), Instituto de Física, Apartado Postal 20-364, México D.F., 01000, México*

<sup>ac</sup> *MTA Atomki, 4001 Debrecen, Hungary*

<sup>†</sup> *Deceased*

*E-mail: [benjamin.tam@queensu.ca](mailto:benjamin.tam@queensu.ca), [ljpickard@ucdavis.edu](mailto:ljpickard@ucdavis.edu)*

**ABSTRACT:** A liquid scintillator consisting of linear alkylbenzene as the solvent and 2,5-diphenyloxazole as the fluor was developed for the SNO+ experiment. This mixture was chosen as it is compatible with acrylic and has a competitive light yield to pre-existing liquid scintillators while conferring other advantages including longer attenuation lengths, superior safety characteristics, chemical simplicity, ease of handling, and logistical availability. Its properties have been extensively characterized and are presented here. This liquid scintillator is now used in several neutrino physics experiments in addition to SNO+.

**KEYWORDS:** Double-beta decay detectors; Neutrino detectors; Scintillators, scintillation and light emission processes (solid, gas and liquid scintillators)

---

## Contents

<b>1</b>	<b>Introduction</b>	<b>1</b>
1.1	The SNO+ Experiment	1
<b>2</b>	<b>Scintillator Development</b>	<b>2</b>
2.1	Solvent	2
2.2	Fluor	4
2.3	Future development	4
<b>3</b>	<b>The SNO+ Liquid Scintillator</b>	<b>4</b>
<b>4</b>	<b>Scintillator Properties</b>	<b>6</b>
4.1	Light Yield	6
4.1.1	Quenching	7
4.2	Scattering	8
4.3	Emission and Absorption	8
4.4	Transfer efficiencies	10
4.5	Re-emission	11
4.6	Timing Measurements	11
4.7	Refractive Index	13
4.8	Density	13
<b>5</b>	<b>Scintillator Cocktail Deployment</b>	<b>13</b>

---

# 1 Introduction

Scintillation counters continue to be among the most common particle detectors. Scintillators are materials that emit light following excitation by ionising radiation [1]; in particle detectors, this light is typically observed with photomultiplier tubes (PMTs).

Organic scintillators — the favoured scintillation medium in large liquid particle detectors — are aromatic solvents with an emission mechanism that depends on the excitation and subsequent de-excitation of benzene rings in various molecules [2]. Liquid scintillators are typically “cocktails” that include fluors and wavelength shifters in order to maximize light yield and transparency, optimize emission times, minimize self-absorption, and tune the emission spectra to match the quantum efficiency curves of the observing PMTs [3]. Recent neutrino experiments that have successfully deployed large volumes of organic liquid scintillator include Borexino, which used 1,2,4-trimethylbenzene (i.e. pseudocumene, PC) [4] and KamLAND, which used PC diluted in dodecane [5]. Both Borexino and KamLAND used 2,5-diphenylloxazole (PPO) as the fluor.

## 1.1 The SNO+ Experiment

SNO+ is a large-scale, multipurpose neutrino physics experiment. The detector is situated 2 km underground at the SNOLAB facility located in the Creighton mine near Sudbury, Canada. The basic infrastructure of the SNO+ detector was inherited from the SNO experiment, including the 6-m radius spherical acrylic vessel (AV) that serves to contain the active target [6]. A detailed description of the SNO+ detector is given in [7].

Acrylic is not compatible with existing widely-used liquid scintillators such as PC. This incompatibility motivated the SNO+ collaboration to search for a new option that would not only be compatible with acrylic, but have a competitive light yield while being safer to handle. This search culminated in the identification of a cocktail consisting of linear alkylbenzene (LAB) as the solvent and PPO as the fluor. Since its development and adoption by the SNO+ collaboration, LAB-based liquid scintillators have been successfully deployed in a number of neutrino detectors such as Daya Bay [8] and RENO [9], and will be used in future neutrino experiments including JUNO [10]. LAB-based scintillators are also used in veto detectors for current dark matter experiments such as COSINE-100 [11], and are considered for use in future dark matter experiments such as SABRE [12].

After an extensive physics campaign operating as an ultrapure water Cherenkov detector [13–15], liquid scintillator has now been deployed in the SNO+ detector. The physics program in this phase of the experiment requires precise characterization of the scintillator to facilitate modelling of the detector response. The physics program also required the scintillator to be of ultra high purity in order to limit radiological backgrounds and minimize light attenuation [7]; a purification plant was built to purify the liquid scintillator prior to deployment into the SNO+ detector.

This paper discusses the SNO+ liquid scintillator: its conception and development, the characterization of its properties, and the purification and deployment techniques used.

## 2 Scintillator Development

The incompatibility of acrylic with commonly used scintillating solvents motivated the SNO+ collaboration to search for a new cocktail. Diluted and mixed solvents, such as the 80% dodecane + 20% PC + PPO (1.52 g/L) mixture used in the KamLAND experiment, have been demonstrated to be sufficiently compatible with acrylic [5]. However, there was a desire to avoid using PC in SNO+ due to safety concerns — PC is highly volatile, toxic, and has a low flash point. This prompted the development of a new liquid scintillator with a light yield competitive with PC, while prioritising acrylic compatibility and superior safety handling considerations. As the AV is submerged in a water-filled cavity that acts as external shielding, a liquid scintillator density close to 1 g/cm<sup>3</sup> was also desirable.

### 2.1 Solvent

There were seven solvents identified as potential candidates for the SNO+ liquid scintillator. These are listed in Table 1.

**Table 1:** Solvent candidates investigated for the SNO+ liquid scintillator.

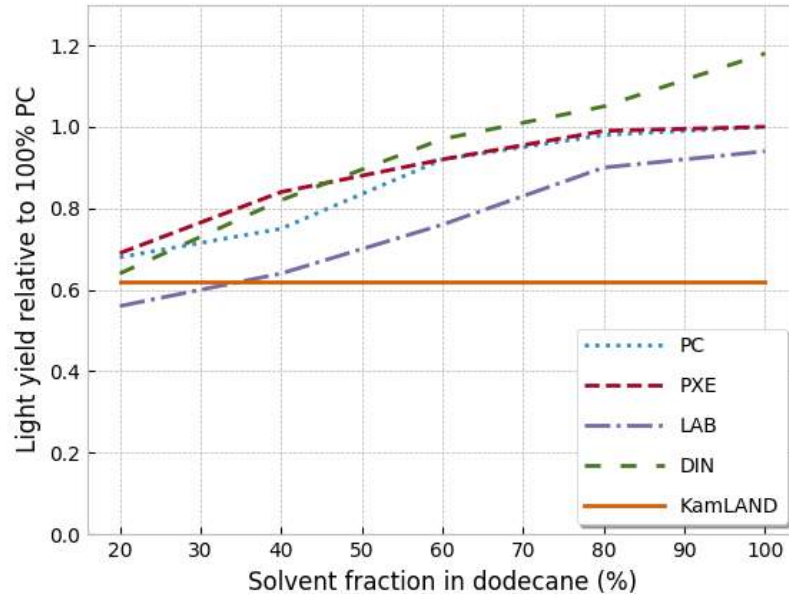
Solvent	Density (g cm <sup>-3</sup> )	Flash point (°C)
Pseudocumene (PC)	0.889	44
Phenylcyclohexane (PCH)	0.950	98
Linear alkylbenzene (LAB)	0.856	143
Di-isopropylnaphthalene (DIN)	0.960	140
Phenyl-o-xylylene (PXE)	0.985	167
1-Methylnaphthalene (1-MN)	1.020	82
Dodecylbenzene (DCB)	0.870	140

Acrylic compatibility, the chief factor in the initial evaluation of the various solvents, was determined by submerging 80 cast acrylic cubes (6.35 mm × 6.35 mm × 6.35 mm) in 175 mL of each candidate solvent prior to stirring with a magnetic teflon rod. A Photon Technology International (PTI) QuantaMaster<sup>TM</sup> fluorescence spectrometer subsequently measured Mie scattering caused by any acrylic particulates suspended in each of the acrylic-exposed solvents; such particulates are indicative of chemical degradation of the cubes. Poor compatibility excluded DCB and PCH as candidates. As expected, PC was also not compatible with acrylic. The remaining solvents exhibited varying — but tolerable — degrees of degradation, and were used in further selection testing.

Light yield was another major factor during selection. To determine the light yields of the remaining solvents, each were doped with PPO (2 g/L) and excited with a 10 mCi <sup>125</sup>I source. The resulting emission spectra were measured using a PTI QuantaMaster<sup>TM</sup> fluorescence spectrometer; the intensities between 350–425 nm were integrated and normalized to measurements using PC to give a relative pulse height (RPH). 1-MN was measured to have a poor light yield, though it is known that the light yield of 1-MN is highly dependent on purity. Furthermore, 1-MN was previously

found to need careful treatment due to issues of robustness [16]. Since 1-MN exhibited no obvious advantages, these factors excluded this solvent from further selection testing.

As other experiments have deployed solvents diluted in dodecane to improve acrylic compatibility [17], this test was repeated for each of the remaining candidates after dilution with dodecane at various concentrations. The measurements were compared to a sample of KamLAND scintillator, as shown in Figure 1.



**Figure 1:** The effect of dodecane dilution on the light yield of the remaining solvent candidates. The solid line is a measurement made of the KamLAND scintillator.

DIN appeared to be a viable choice due to its competitive acrylic compatibility, high intrinsic light yield, and comparable density to water. However, compared to LAB, DIN has a shorter attenuation length of 3.3–4.4 m at 430 nm [18] which is comparable to the scale of the SNO+ detector.<sup>1</sup> Mixtures which minimized acrylic damage while maintaining competitive light yields were also considered — particularly those involving PXE. However, PXE is incompatible with acrylic when undiluted and has been demonstrated to soften acrylic above a solvent fraction of 20% in dodecane [19]. Compared to LAB, no combination of solvents and/or dilutions conferred sufficient advantages to justify the increased cocktail complexity.

LAB was determined to be the superior compromise between acrylic compatibility, light yield, cocktail simplicity, and safety. Furthermore, LAB is a single-component solvent, has a near-identical index of refraction to acrylic, and is easy to handle due to its non-toxicity and mild odour. Commercial LAB has minimal Rayleigh scattering and high transparency, with attenuation lengths

<sup>1</sup>Preliminary measurements of LAB demonstrated mean attenuation lengths of over 20 m. More extensive measurements were later performed on LAB, and are discussed in Sections 4.2 and 4.3.

at relevant wavelengths of over 10 m even prior to purification. LAB is also a logistically favorable choice due to the relative proximity of SNOLAB to the CEPSA Química Bécancour (formerly known as Petresa Canada) linear alkylbenzene facility in Bécancour, Quebec.

Once LAB was identified as the premier solvent, an extensive investigation was initiated to confirm its acrylic compatibility, following ASTM D543 “Standard Practices for Evaluating the Resistance of Plastics to Chemical Reagents”. After 100 months of testing at elevated stresses and strains, it was found that LAB has a negligible chemical effect on the acrylic used in the SNO+ experiment. Details of the long-term acrylic compatibility tests and the experimental design are given in [20] with the most recent results given in [21].

## 2.2 Fluor

PPO has been the dominant choice of fluor in liquid scintillators for decades [22]. Two additional solutes, 2-(4-Biphenyl)-5-phenyloxazole (BPO) and 2-(1-Naphthyl)-5-phenyloxazole (NPO), were also identified as potential options. Furthermore, phenyls such as p-terphenyl (another common fluor), along with p-quinquephenyl and p-sexiphenyl (both less common), were briefly considered but excluded due to their low solubility in LAB. Although BPO and NPO have slightly higher intrinsic light yields, PPO was ultimately chosen as the fluor due to its low cost, wide availability, and the extensive existing literature as a result of its widespread use.

## 2.3 Future development

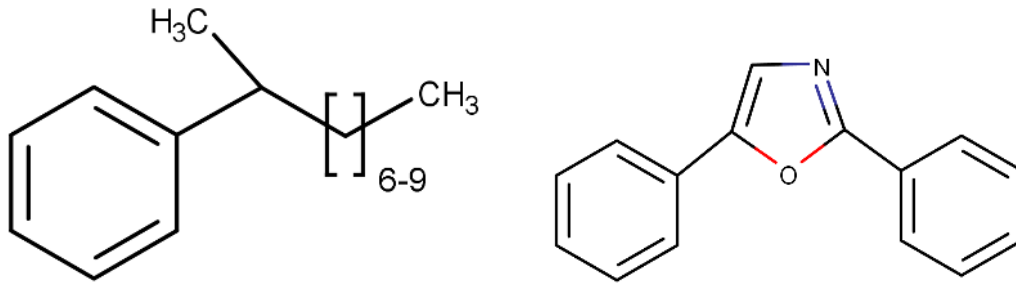
In the initial scintillator development phases, metal-loading capabilities were an independent consideration and did not impact the selection process. A technique was ultimately established to load natural Te into the scintillator in order to allow for the study of neutrinoless double beta decay with  $^{130}\text{Te}$ . To mitigate absorption by Te, 1,4-bis(2-methylstyryl)benzene (bis-MSB) can be added to the scintillator as an additional secondary wavelength shifter. Bis-MSB is known to perform well with PPO as a fluor-shifter mixture. The Te loading process will be discussed in a forthcoming paper.

## 3 The SNO+ Liquid Scintillator

The final SNO+ liquid scintillator (“the scintillator cocktail”) was determined to be 2 g PPO per 1 L LAB. Increasing the concentration of PPO in LAB increases both the primary light yield and self-absorption. Therefore, at distances relevant to the SNO+ detector, the increase in light yield due to the increase in PPO starts to plateau at concentrations higher than 2 g/L.

LAB acts as both the solvent and primary absorber. The carbon chain attached to the phenyl group varies from 9–14 carbons in length, as seen in Figure 2 (left); each chain length has subtle emission spectra differences. As discussed in Section 4.3, the presence of a fluor mitigates self-absorption of LAB by transferring energy from LAB to PPO (primarily via non-radiative Förster resonant energy transfer), resulting in light being emitted at longer wavelengths. The structural formula for PPO can be seen in Figure 2 (right).

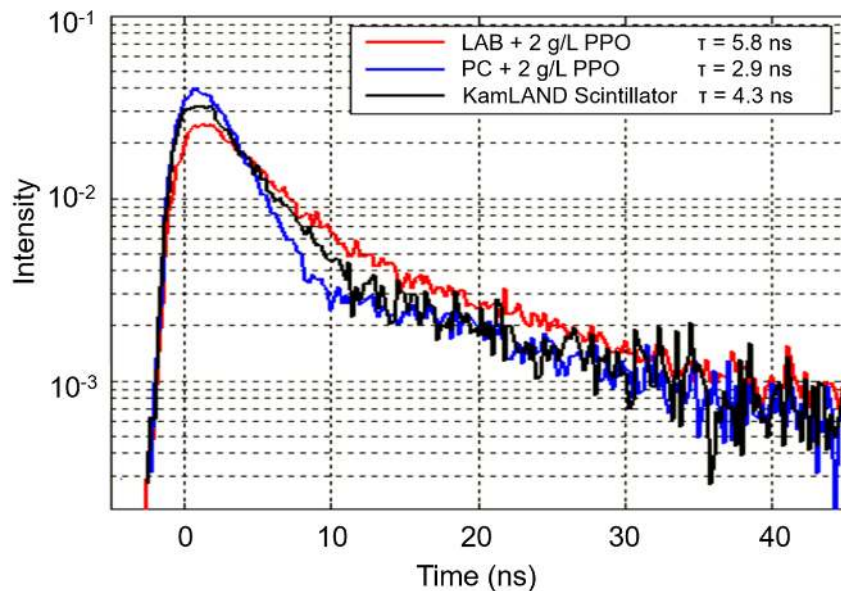




**Figure 2:** Structural formulae of the two primary components of the SNO+ scintillator cocktail. Left: LAB is a phenyl group attached to a carbon chain varying between 9-14 carbons in length. Right: PPO is the fluorophore of the liquid scintillator cocktail.

The scintillator cocktail exhibits acrylic compatibility and competitive light yields with respect to existing liquid scintillators, while maintaining long attenuation lengths, improved safety characteristics, chemical simplicity, ease of handling, and logistical availability to the SNO+ site.

As part of the final selection, a preliminary comparison of the decay time was performed between the unpurified scintillator cocktail, scintillator used in the KamLAND experiment, and PC + 2g/L PPO (the scintillator formula used in the Borexino experiment). These results are presented in Figure 3. The scintillation times were deduced using a single exponential fit after UV excitation. A more thorough analysis of the scintillator cocktail timing was subsequently undertaken and is presented in Section 4.6.



**Figure 3:** Comparison of scintillation decay times for the KamLAND scintillator, PC + 2 g/L PPO, and LAB + 2 g/L. The time constant  $\tau$  for each was determined using a single exponential fit.

## 4 Scintillator Properties

The properties of the scintillator cocktail have been extensively characterized and subsequently incorporated into the SNO+ Monte Carlo framework to ensure that the detector is accurately and precisely modelled [6].

### 4.1 Light Yield

As with all scintillation detectors, characterization of the light yield was required to understand the energy resolution of the SNO+ experiment. Two different methods to determine the light yield were carried out: irradiation using an  $^{125}\text{I}$  source with the light output measured using a PMT-based spectrometer, and irradiation using a  $^{90}\text{Sr}$  source with the light output measured using a PMT.

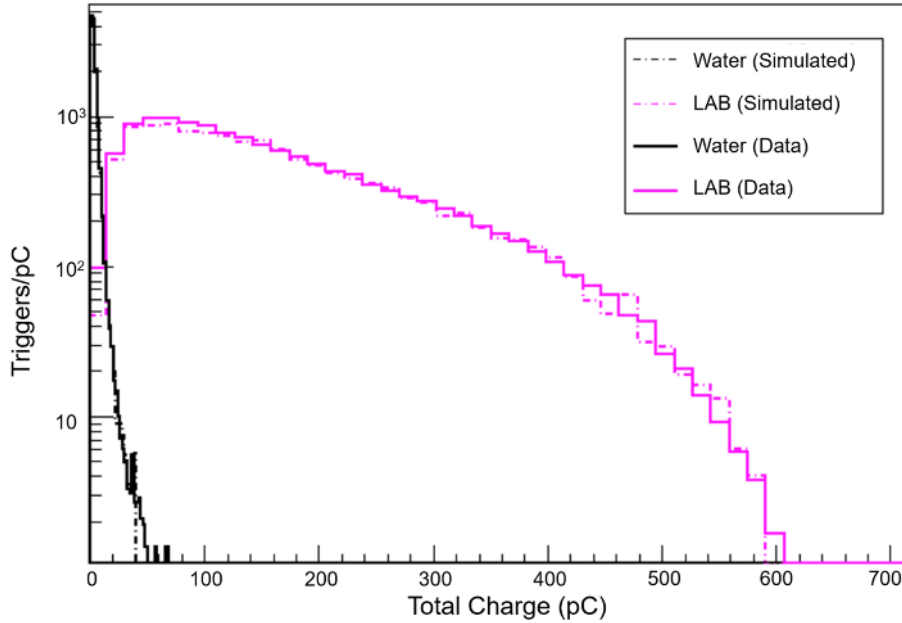
**$^{125}\text{I}$  irradiation** The light yield of the scintillator cocktail was quantified through comparison to PC + 2 g/L PPO. Each were poured into a  $1 \times 1$  cm quartz cuvette located in a dark box, and subsequently irradiated with 35 keV X-rays from an  $^{125}\text{I}$  source. The emission spectra were measured using a PMT-based spectrometer. The relative light yield was determined by integrating their spectra between 350–425 nm. After accounting for the  $\sim 5$  cm attenuation lengths and the density differences between the solvents, LAB was found to have a relative light yield of 96% compared to PC with equivalent PPO loading.

The light yield of PC + PPO (at 1.5 g/L) was determined by the Borexino collaboration to be  $11500 \pm 1000$  photons/MeV [4]. To obtain a 2.0 g/L equivalent loading, this light yield was scaled by 1.075 based on measurements that were performed. Accounting for the 96% relative measurement, the absolute light yield of the SNO+ scintillator cocktail was determined to be  $11900 \pm 1100$  photons/MeV.

**$^{90}\text{Sr}$  irradiation** Irradiation with a  $0.1 \mu\text{Ci}$   $^{90}\text{Sr}$  source provided a complementary light yield measurement of the scintillator cocktail. The source was embedded in a  $3.5 \text{ cm} \times 3.5 \text{ cm} \times 3.5 \text{ cm}$  acrylic cube with a hollowed out, 2 cm diameter,  $\sim 11$  mL cylindrical sample cavity. The size of the cube was optimized to reduce self-absorption.  $^{90}\text{Sr}$  decays to  $^{90}\text{Y}$ , emitting a 0.546 MeV  $\beta^-$  that is fully absorbed by the acrylic. However,  $^{90}\text{Y}$  decays to  $^{90}\text{Zr}$ , emitting a 2.28 MeV  $\beta^-$  which penetrates through the acrylic and ionizes the sample. A 1" Hamamatsu R7600-200 HQE PMT was used to trigger an acquisition, while three optically isolated 12" Hamamatsu R11780 HQE PMT's were used to collect the scintillation light. This setup was located in a  $2 \text{ m} \times 2 \text{ m} \times 0.9 \text{ m}$  dark box surrounded by a Helmholtz coil and FINEMET<sup>®</sup> shielding to reduce the effect of the Earth's magnetic field.

For each triggered event, the total charge measured by all three 12" PMTs was used to build a charge spectrum. The experiment and simulation was also performed for Cherenkov light by placing water in the sample container, which was used to normalize the integrated light collection as shown in Figure 4. The accumulated spectrum was compared to simulation, and the light yield of the scintillator cocktail was determined to be  $10830 \pm 570$  photons/MeV. As oxygen is known to be responsible for quenching effects in scintillators [23], this method was repeated after the sample was de-oxygenated through 30 minutes of sparging with  $\text{N}_2$ . The light yield of the scintillator cocktail

after de-oxygenation was increased to  $11920 \pm 630$  photons/MeV. This is in good agreement with the  $^{125}\text{I}$  X-ray assay. The systematic uncertainties for this test are presented in Table 2.



**Figure 4:** The measured charge spectra for the scintillator cocktail compared to a simulated best fit. The integrated light collection was normalized to Cherenkov light emitted by a water sample placed in the same sample container.

**Table 2:** A breakdown of systematic uncertainties for the  $^{90}\text{Sr}$ -irradiated light yield measurement.

Systematic Uncertainty	Value
PMT distance to source	$\pm 0.5$ cm
Acrylic window thickness	$\pm 0.5$ mm
Source disc thickness	$\pm 0.25$ mm
Birks' constant	$\pm 4$ $\mu\text{m}/\text{MeV}$
PMT quantum efficiency	$\pm 3.5\%$
Overall systematic uncertainty	$\pm 5.3\%$

#### 4.1.1 Quenching

There are a number of non-radiative de-excitations or transitions that do not result in light generation. Such quenching effects reduce the light yield of the scintillator and are modeled through the Birks' equation:

$$\frac{dL}{dx} = S \frac{\frac{dE}{dx}}{1 + k_B \frac{dE}{dx}} \quad (4.1)$$

where  $\frac{dL}{dx}$  is the fluorescent energy emitted per unit path length,  $\frac{dE}{dx}$  is the specific energy loss of the charged particle, and  $S$  is the scintillation efficiency which provides a normalization.  $k_B$  is a material-specific Birks' constant.

Various quenching effects and the Birks' constant of the scintillator cocktail has been extensively studied by the SNO+ collaboration. Recent measurements of the Birks' constant are in the range of  $0.0066 \pm 0.0016$  to  $0.0076 \pm 0.0003$  cm/MeV for  $\alpha$ -particles [24],  $0.0094 \pm 0.0002$  to  $0.0098 \pm 0.0003$  cm/MeV for protons [25], and  $\sim 0.0074$  cm/MeV for electrons [26]. Additionally, the quenching of the scintillator cocktail has a temperature dependence that varies based on the particle type [27]. These measurements have been corroborated through multiple other studies [28, 29].

## 4.2 Scattering

The attenuation of photons via scattering was measured in order to accurately model SNO+ in simulations. Of the three main scattering mechanisms — Rayleigh, Mie, and Raman — Rayleigh scattering (in which the photon wavelength is much larger than the scattering molecule) dominates due to the size of the LAB and PPO molecules.<sup>2</sup> The Rayleigh ratio and isotropic nature of the scattered light was assessed to understand this effect [30].

The Rayleigh ratio is the relationship between the intensities of incident and scattered light, and can be used to determine mean attenuation lengths. The Rayleigh ratio for LAB was determined by comparing its scattered light intensity with those of ultrapure water, acetone, cyclohexane, and toluene. This measurement was performed using a PTI QuantaMaster<sup>TM</sup> fluorescence spectrometer after the solvents were placed in a  $1 \times 1$  cm quartz cuvette within a dark box and illuminated with a Xe lamp monochromatized to 546 nm. The Rayleigh ratio of LAB corresponding to each solvent was determined using

$$R_{LAB} = R_s \frac{I_{LAB}}{I_s} \frac{n_{LAB}}{n_s} A, \quad (4.2)$$

where  $R_{LAB}$  and  $R_s$  are the Rayleigh ratio of the LAB and solvent,  $I_{LAB}$  and  $I_s$  are the measured intensities of the LAB and solvent, and  $n_{LAB}$  and  $n_s$  are the refractive index of LAB and solvent.  $A$  is a correction factor implemented to account for the differences in refractive indices between the solvent and the quartz cuvette.

As seen in Table 3, the intensity of light scattered off of LAB was compared to that of each solvent to determine the Rayleigh ratio, with an average of  $(16.60 \pm 3.14) \times 10^{-6} \text{ cm}^{-1}$  between the four solvents. Calculated from the Rayleigh scattering cross section, the resulting mean scattering length was determined to be  $71.90 \pm 13.6$  m at 546 nm.

## 4.3 Emission and Absorption

The emission and absorption spectra of the scintillator cocktail were characterized in order to determine the Stokes' shift and understand the effect of self-absorption.

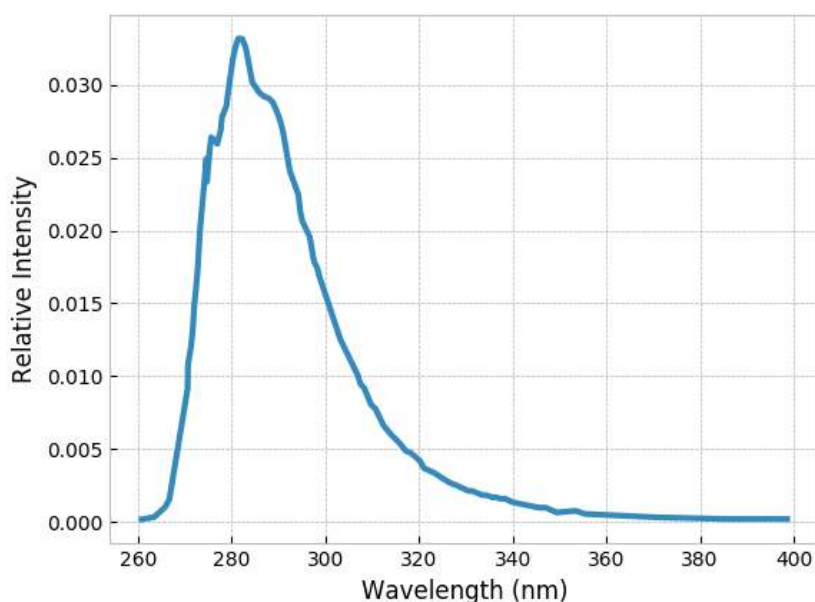
To determine the emission spectrum, LAB was first diluted in spectrophotometric-grade cyclohexane (2.5 mL LAB per 1 L cyclohexane) to minimize self-absorption. A PTI QuantaMaster<sup>TM</sup> fluorescence spectrometer with an excitation wavelength of 250 nm was then used to measure the emission spectrum for the diluted LAB, as is presented in Figure 5. As can be seen, LAB was measured to emit in the 275–350 nm range.

---

<sup>2</sup>Rayleigh scattering is the dominant mechanism only if the scintillator is free of impurities with large molecular sizes, as is the case for the scintillator cocktail after purification.

**Table 3:** Using the Rayleigh ratio of known solvents, the Rayleigh ratio of LAB was determined by comparing the intensity of scattered light between LAB and each solvent. The correction factor accounts for discrepancies in refractive indices between the solvent and quartz cuvette used for the measurements.

Solvent	Rayleigh ratio (cm <sup>-1</sup> )	Correction Factor	Corresponding Rayleigh ratio of LAB (cm <sup>-1</sup> )
Ultrapure Water	1.05	21	$(16.63 \pm 2.99) \times 10^{-6}$
Acetone	4.47	5	$(16.79 \pm 3.14) \times 10^{-6}$
Cyclohexane	4.68	4	$(16.46 \pm 3.00) \times 10^{-6}$
Toluene	20.5	1	$(16.61 \pm 3.45) \times 10^{-6}$

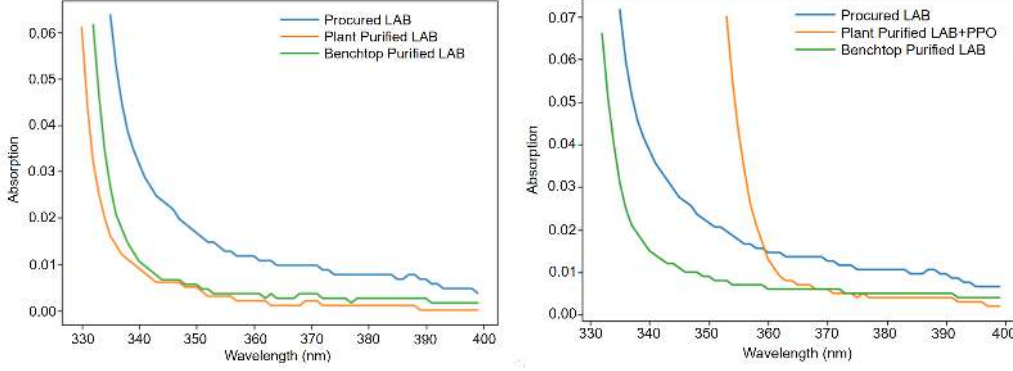


**Figure 5:** Emission spectrum of LAB, diluted to 2.5 mL LAB per 1 L cyclohexane to minimize self-absorption and excited using a 250 nm fluorescence spectrometer.

As mentioned in Section 1.1 and more thoroughly discussed in Section 5, LAB was distilled in a purpose-built scintillator purification plant prior to deployment. The absorption spectrum of purified LAB was determined by placing the solution in a 1×1 cm quartz cuvette, which was measured between 190–1100 nm using an Orion Aquamate 8000 spectrophotometer. The purified LAB was compared to the procured LAB after 0.2 μm filtration but before distillation, as well as the best possible “benchtop” laboratory-purified LAB without using the purification plant. The LAB purified by the plant had consistently superior optical clarity in the 330–400 nm region of interest, as presented in Figure 6 (left).

The absorption and emission spectra of PPO have been extensively studied, and have previously

been presented in compilations such as [31]. The SNO+ collaboration also measured the absorption and emission spectra of PPO, obtaining results consistent with these previous publications. The absorption spectrum of scintillator cocktail deployed within the SNO+ detector is shown in Figure 6 (right). As in Figure 6 (left), this spectrum is also compared to both undistilled and “benchtop” distilled LAB. At wavelengths above which PPO is highly absorbing ( $\sim 365$  nm), the superior optical clarity of the liquid scintillator is maintained after addition of the fluor.



**Figure 6:** Absorption spectra of pure LAB and the scintillator cocktail (2 g PPO per 1 L LAB), produced by the SNO+ scintillator purification plant (orange), procured pure LAB after  $0.2 \mu\text{m}$  filtration but before distillation (blue) and best possible “benchtop” laboratory-purified pure LAB (green). Left: Pure LAB has superior optical clarity after distillation. Right: The scintillator cocktail has superior optical clarity after distillation at all wavelengths above the region where PPO is highly absorbing ( $\sim 365$  nm).

#### 4.4 Transfer efficiencies

In order to determine the efficiency of the non-radiative transfer from LAB to PPO, the scintillator cocktail was optically excited in order to fit the decay time to the fluorescence response function for fluor ‘y’, given a solvent ‘x’,

$$n_y(t) = \frac{n_{0x} \frac{a_{xy}}{\tau_{txy}}}{\frac{1}{\tau_{0x}} - \frac{1}{(\tau_{0y})_0}} \left( e^{-\frac{t}{(\tau_{0x})_0}} - e^{-\frac{t}{\tau_{0x}}} \right). \quad (4.3)$$

Here,  $n_y$  is the number of excited fluor molecules,  $n_{0x}$  is the number of excited solvent molecules at  $t = 0$ ,  $(\tau_{0y})_0$  is the molecular fluorescence time of the fluor (1.66 ns for PPO [31]),  $\tau_{0x}$  is the decay time for excitations of the solvent,  $a_{xy}$  is the radiative transfer efficiency, and  $\tau_{txy}$  is the non-radiative transfer decay time [1].

Fitting Equation 4.3 to extract  $\tau_{0x}$ , the non-radiative transfer efficiency can then be deduced using

$$f_{xy} = 1 - \frac{\tau_{0x}}{(\tau_{0x})_0}, \quad (4.4)$$

where  $(\tau_{0x})_0$  is the pure solvent’s decay constant (measured using a time-based laser-excited fluorimeter to be 22.7 ns for LAB). The measurements of non-radiative transfer efficiency for LAB loaded with different concentrations of PPO are presented in Table 4.

**Table 4:** The non-radiative transfer efficiencies for various PPO concentrations in LAB.

PPO Concentration (g PPO per 1 L LAB)	Non-Radiative Transfer Efficiency %
4	$86.0 \pm 0.8$
2	$78.2 \pm 1.5$
1	$67.7 \pm 2.3$
0.5	$59.3 \pm 3.2$
0.25	$48.7 \pm 5.0$

#### 4.5 Re-emission

Photons absorbed by the scintillator cocktail have a non-zero probability of being re-emitted. Measurements of the quantum yield (the emission probability of a fluorophore) typically involve comparing the fluorescence intensity of the sample against a known reference, such as quinine sulfate ( $C_{20}H_{24}O_2N_2$ ). These measurements are prone to error, and literature values vary over a large range. This is due to a variety of parameters including oxygen quenching, wavelength dependencies, geometry, and dependence on the presence of the solvent matrix (in which numerous corrections must be included to extract accurate values) [32].

**Absorption by PPO** The quantum yield of PPO has been extensively measured with values ranging from 0.71–1.00. SNO+ adopted a quantum yield of 0.82, as measured by the Borexino collaboration [33]. A recent paper reports a value of  $0.842 \pm 0.042$  for the quantum yield of PPO [32]. When modelling re-emission time, light absorbed by PPO molecules is assumed to fluoresce with the intrinsic PPO exponential decay time constant of 1.6 ns.

**Absorption by LAB** In cases where the photon is absorbed by a LAB molecule, the quantum yield is assumed to be  $0.20 \pm 0.02$ , as measured by [32]. However, as discussed in Section 4.4, non-radiative energy transfer to a PPO molecule can occur with a probability of  $(78.2 \pm 1.5)\%$ ; a conservative probability of 72% can be safely assumed. Taking the product of the conservative non-radiative transfer efficiency and the quantum yield of the PPO gives a calculated re-emission probability of 0.59 for a photon absorbed by an LAB molecule. When modelling re-emission, the dominant PPO emission spectrum is assumed. The re-emission time, following light absorption by LAB molecules and the subsequent re-emission by PPO after non-radiative transfer, was measured using UV fluorescence excitation to be  $5.1 \pm 0.1$  ns.

#### 4.6 Timing Measurements

The ability to discriminate between alpha particles ( $\alpha$ ) and electrons ( $\beta$ ) is an important capability for liquid scintillator detectors. Consequently, significant efforts were undertaken to parameterize the timing profile of the scintillator cocktail; these characteristics were also required for accurate modelling in Monte Carlo simulations.

When LAB is excited after interaction with an ionising particle, the excitation energy can be transferred to PPO through Förster resonance energy transfer, a dipole-dipole non-radiative interaction.



PPO then undergoes radiative decay, and de-excites from singlet or triplet states. Singlet states will decay directly to the ground state resulting in the prompt emission of scintillation light. De-excitation from the triplet to the singlet state is forbidden if intersystem crossing occurred, resulting in a delayed emission of scintillation light [1]. The ratio between singlet and triplet states depend on the ionization density of the incident particle. Particles with high ionization densities, such as  $\alpha$ -particles, will undergo quenching and produce a higher proportion of triplet states when compared to  $\beta$ -particles. Therefore,  $\alpha/\beta$  discrimination can be achieved by using the liquid scintillator timing profile — the intensity of scintillation light as a function of time.

Measurements have previously been made to determine the scintillation time profile of oxygenated and de-oxygenated LAB-based scintillator cocktails under  $\alpha$  and  $\beta$  excitation, with results published in [23]. The importance of de-oxygenation in the ability to perform  $\alpha/\beta$  discrimination was demonstrated. However, subsequent measurements made by [34, 35] were unable to replicate the very effective separation observed between  $\alpha$  and  $\beta$  events. As the setup and methodology were well understood and verified, it remains unclear why this particular result was seemingly better than other observations.

Due to this tension, two further complementary measurements were carried out by the SNO+ collaboration to characterize the timing profile of the scintillator cocktail. Both measurements used the single-photon sampling technique described in [36] with slightly different methodologies. In each, a minimum of two PMTs were used; one to trigger scintillation events (“trigger PMT”) and another placed such that the average number of photons detected per trigger was less than one (“timing PMT”). The time difference between the trigger and timing PMTs yielded the emission time of the detected photon. The overall timing profile can then be built by accumulating these measurements over many events.

**Method 1** Measurements were taken by placing the liquid scintillator in a  $3.5\text{ cm} \times 3.5\text{ cm} \times 3.5\text{ cm}$  acrylic cube with a hollowed out, 2 cm diameter,  $\sim 11$  ml cylindrical sample cavity. The acrylic cube was optically coupled to a 1" trigger PMT, and a 30 mm  $\times$  30 mm timing PMT was placed at an offset of  $\sim 30$  cm.  $\beta$  events were produced using a  $^{90}\text{Sr}$  source while  $\alpha$  events were produced using a  $^{210}\text{Po}$  source. Both sources were placed above the acrylic cube and separated from the cocktail by a  $\sim 1$  mm air gap.

**Method 2** Measurements used three 2" cylindrical PMTs, and were taken by placing the liquid scintillator into an ultraviolet-transparent acrylic (UVT) sample holder. The sample holder was optically coupled to a trigger PMT. A secondary PMT was also coupled to the sample holder, which was deployed to understand the scintillator rise time. The timing PMT was masked and placed at a distance where it would only observe  $< 1\%$  of triggered events.  $\beta$  events were produced using a  $^{137}\text{Cs}$  source placed beneath the sample holder, while  $\alpha$  events were produced using a  $^{244}\text{Cm}$  source immersed within the scintillator cocktail.

**Results** The timing profiles of Method 1 were fit using a  $\chi^2$ -minimization approach while accounting for the dark rate and after-pulsing; the data was well described with a summation of four exponentials convolved with Gaussians. Conversely, the timing profiles of Method 2 were parameterized using a maximum likelihood fit in which the dark current, single photon transit time



of the PMTs, and trigger distribution were accounted for; the data was also well described using summation of four exponentials. The fit parameters and resulting timing profiles between the two methods were in good agreement, and are presented in Table 5.

**Table 5:** The fit parameters for the  $\alpha$  and  $\beta$  timing profiles of the scintillator cocktail.  $\tau_i$  are the decay constants and  $A_i$  are the scaling factors for each exponential.

	Method 1		Method 2	
Fit	$\alpha$	$\beta^-$	$\alpha$	$\beta^-$
$\tau_1$ (ns)	4.66±0.17	5.1±0.2	4.79±0.18	4.88±0.08
$\tau_2$ (ns)	14.2±2.0	17.6±2.0	18.4±0.9	15.4±0.7
$\tau_3$ (ns)	64.3±6.1	45.3±5.1	92.0±5.0	66.0±4.0
$\tau_4$ (ns)	578±313	498±50	900±110	400±40
$A_1$	0.44±0.02	0.66±0.02	0.427±0.01	0.665±0.008
$A_2$	0.31±0.01	0.20±0.01	0.313±0.01	0.218±0.009
$A_3$	0.16±0.01	0.08±0.01	0.157±0.06	0.083±0.004
$A_4$	0.09±0.02	0.06±0.02	0.102±0.003	0.035±0.002

#### 4.7 Refractive Index

The relationship between refractive index and wavelength in LAB+PPO was first published in [37] using 3.0 g PPO per 1 L LAB. A Woollam M2000 ellipsometer was used to scan wavelengths of this higher concentration scintillator cocktail between 210 - 1000 nm at 5 incident angles (55°, 60°, 65°, 70° and 75°). The average of all 5 incident angles was used to determine the refractive index at each wavelength. As shown in Figure 7, these results were compared to other measurements made by SNO+ and RENO [38]. More details about this measurement can be found in [37].

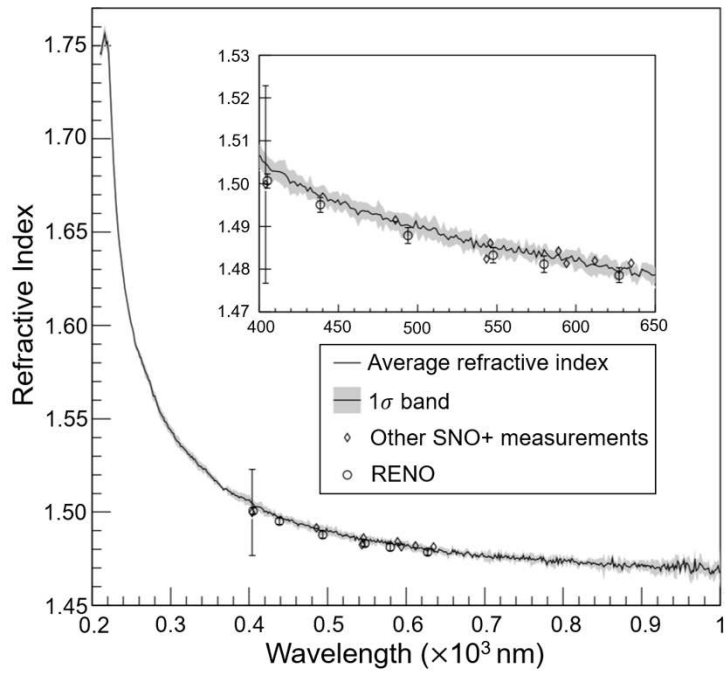
#### 4.8 Density

The density of the scintillator cocktail depends on temperature. To characterize this, samples of both pure LAB and the scintillator cocktail were either placed in glass syringes and cooled in a refrigeration unit, or placed in a glass beaker and heated on a hot plate. The samples were then measured using a PAAR DMA 35 densitometer. As shown in Figure 8, the density of pure LAB and the scintillator cocktail was assessed between 10–26°C. These measurements are consistent with those published in [39].

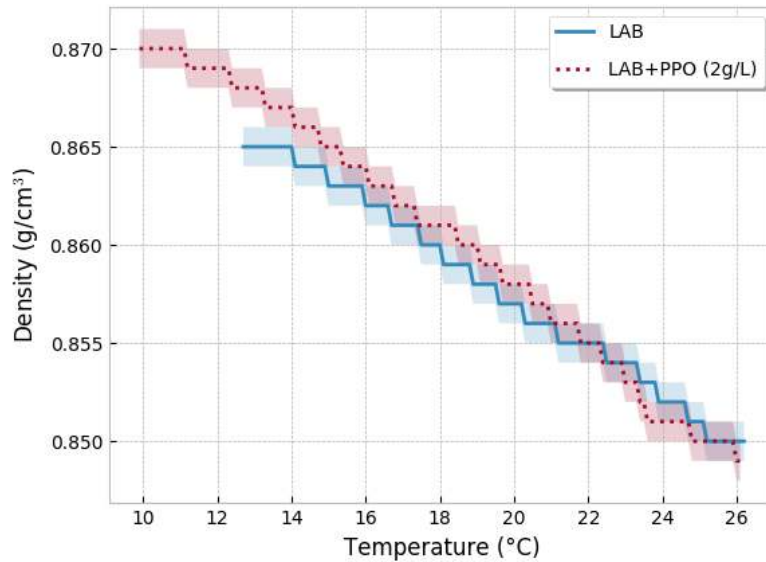
### 5 Scintillator Cocktail Deployment

The SNO+ scintillator production and purification methodology was previously discussed in [40]. This section will provide an overview of the production, purification and quality assurance processes used during the deployment of the scintillator cocktail.

In order to meet the strict cleanliness and purity requirements of the experiment [7], P 500-Q LAB was purchased from CEPESA Química Bécancour Inc. Upon receiving shipments, the ultraviolet-visible (UV-Vis) and Fourier-transform infrared (FTIR) spectra, density, and turbidity were measured. If the delivered LAB had sufficient quality as deemed by these measurements, it was stored



**Figure 7:** The wavelength dependent refractive index of the scintillator cocktail with a higher PPO loading of 3.0 g/L. The shaded region represents  $1\sigma$  deviations. These results are compared to previous refractometry measurements by SNO+ and RENO [38]. Figure taken from [37].



**Figure 8:** Comparison of density variation of the scintillator cocktail between 10–26°C. The shaded regions represent instrumental uncertainties from the densitometer.

in a stainless steel holding tank located in the SNOLAB surface facility. This tank was previously passivated with citric acid and cleaned, and the LAB stored within was kept under a blanket of high purity nitrogen gas. The LAB was then loaded onto stainless steel railcars, each cleaned in the same manner as the surface storage container. These railcars were subsequently transported underground, where the LAB was unloaded into two storage tanks, each capable of holding 50 tonnes of LAB. These storage tanks have a polypropylene liner, and were cleaned in accordance with MIL-STD-1247C level 50 using both a 1% Alconox detergent solution and ultrapure water.<sup>3</sup>

Once underground, the LAB was purified using the SNO+ scintillator purification plant. The first step of this purification was a continuous multi-stage distillation. This system heats the LAB to 220°C at an absolute pressure of 40 Torr. The distillation column was refluxed such that a distillate flow rate of up to 750 kg/hr was achieved. This process removed lower volatility impurities including heavy metals (Bi, K, Pb, Po, Ra and Th) and oxidised organics (carboxyl groups and 1,4-benzoquinone). The distilled LAB then underwent N<sub>2</sub> stripping at 100°C and an absolute pressure of 150 Torr. This process removed dissolved gases (Ar, Kr, O<sub>2</sub> and Rn) and volatile liquids (such as residual water). The capability to perform steam stripping on LAB was developed, and may be used to polish the deployed scintillator cocktail in the future.

The distilled and N<sub>2</sub>-stripped LAB was systematically monitored throughout all purification processes. Improvements to the solvent were assessed using UV-Vis spectroscopy, densitometry, and nephelometry, which served as proxies for radiopurity. Tested every hour, only LAB that passed this strict quality assurance regime were utilised in further stages. An example of the UV-Vis spectra performed during this verification process is shown in Figure 6.

The PPO used in the deployed scintillator cocktail was purchased as “neutrino grade” from PerkinElmer, delivered to SNOLAB in plastic drums, and shipped directly to the underground laboratory. The PPO was mixed with purified LAB to a target concentration of ~120 g PPO per 1 L LAB to produce a high-concentration “master solution”. This master solution was then sparged three times with high purity N<sub>2</sub> gas using a vacuum pump/purge system to remove dissolved gases such as O<sub>2</sub> and Rn. Three solvent-solvent water extractions were then performed on the master solution, in which the raffinate was mixed with ultrapure water and allowed to re-separate. This removed U, Th, Ra, K, and Pb, as well as other charged particulates and ionic impurities. The master solution was subsequently filtered and distilled in a single-stage kettle at 230°C and an absolute pressure of 20 Torr. At this temperature and pressure, the LAB flashed through the system while the PPO was distilled, and the two components were re-mixed in a condenser.

The distilled master solution was mixed in-line with purified LAB to attain the desired scintillator cocktail concentration of 2 g PPO per 1 L of LAB. This was then passed through 0.05 μm filters before being systematically tested every hour using UV-Vis spectroscopy, densitometry and nephelometry. Further tests were also regularly performed; these included light yield, FTIR, and neutron activation analysis. Once the purity of the scintillator cocktail was assured through a multi-stage

---

<sup>3</sup>MIL-STD-1247C is the United States Military Standard on Product Cleanliness Levels and Contamination Control Program. [www.everyspec.com/MIL-STD/MIL-STD-1100-1299/MIL\\_STD\\_1246C\\_131/](http://www.everyspec.com/MIL-STD/MIL-STD-1100-1299/MIL_STD_1246C_131/).

approval process, the cocktail was cooled to a target temperature of 12 °C and sent to the SNO+ AV.

Following deployment of the scintillator cocktail within the SNO+ detector, the purification plant is designed to maintain the optical clarity and radiopurity of the scintillator through solvent-solvent extraction with ultrapure water, steam stripping, and the use of metal scavengers. These processes are designed to operate at a flow rate of ~130 litres per minute, allowing for a recirculation of the entire 780 tonne detector volume in ~100 hours, comparable to the half-life of  $^{222}\text{Rn}$ .

## **6 Summary**

Due to the incompatibility between existing widely-used liquid scintillators and the acrylic used to construct the SNO+ detector, the SNO+ collaboration developed a new liquid scintillator. The resulting scintillator cocktail of 2 g PPO per 1 L LAB is not only compatible with acrylic, but exhibits a competitive light yield to existing liquid scintillators while maintaining other advantages including longer attenuation lengths, superior safety characteristics, chemical simplicity, ease of handling, and logistical availability to the SNO+ site.

In preparation for the SNO+ scintillator phase physics program, properties of the scintillator cocktail were extensively characterized to ensure that the detector is accurately modelled in Monte Carlo simulations. The LAB-based liquid scintillator developed by the SNO+ collaboration is now used in multiple large-scale experiments, and has been deployed in the SNO+ experiment.

## Acknowledgments

Capital construction funds for the SNO+ experiment were provided by the Canada Foundation for Innovation (CFI) and matching partners. This research was supported by: **Canada:** Natural Sciences and Engineering Research Council, the Canadian Institute for Advanced Research (CIFAR), Queen's University at Kingston, Ontario Ministry of Research, Innovation and Science, Alberta Science and Research Investments Program, National Research Council, Federal Economic Development Initiative for Northern Ontario (FedNor), Northern Ontario Heritage Fund Corporation, Ontario Early Researcher Awards, Arthur B. McDonald Canadian Astroparticle Physics Research Institute; **Germany:** the Deutsche Forschungsgemeinschaft (DFG); **Mexico:** DGAPA-UNAM and Consejo Nacional de Ciencia y Tecnología; **Portugal:** Fundação para a Ciência e a Tecnologia (FCT-Portugal); **UK:** Science and Technology Facilities Council (STFC), the European Union's Seventh Framework Programme under the European Research Council (ERC) grant agreement, the Marie Curie grant agreement; **US:** Department of Energy Office of Nuclear Physics, National Science Foundation, the University of California, Berkeley, Department of Energy National Nuclear Security Administration through the Nuclear Science and Security Consortium.

We would like to thank SNOLAB and its staff for support through underground space, logistical, and technical services. SNOLAB operations are supported by the Canada Foundation for Innovation and the Province of Ontario, with underground access provided by Vale Canada Limited at the Creighton mine site.

## References

- [1] J. B. Birks, *The Theory and Practice of Scintillation Counting*. Pergamon Press, 1964.
- [2] F. Brooks, A scintillation counter with neutron and gamma-ray discriminators, *Nucl. Instrum. Meth. A* **4** (1959) 151–163.
- [3] F. Brooks, *Development of Organic Scintillators*, *Nucl. Instrum. Meth.* **162** (1979) 477–505.
- [4] F. Elisei et al., *Measurements of liquid scintillator properties for the Borexino detector*, *Nucl. Instrum. Meth. A* **400** (1997) 53–68.
- [5] S. Yoshida et al., *Light output response of KamLAND liquid scintillator for protons and C-12 nuclei*, *Nucl. Instrum. Meth. A* **622** (2010) 574–582.
- [6] SNO+ collaboration, S. Andringa et al., *Current Status and Future Prospects of the SNO+ Experiment*, *Adv. High Energy Phys.* **2016** (2016) 6194250, [1508.05759].
- [7] SNO+ collaboration, M. Anderson et al., *The sno+ experiment*, *JINST* (forthcoming).
- [8] DAYA BAY collaboration, F. An et al., *The Detector System of The Daya Bay Reactor Neutrino Experiment*, *Nucl. Instrum. Meth. A* **811** (2016) 133–161, [1508.03943].
- [9] J. Park et al., *Production and optical properties of Gd-loaded liquid scintillator for the RENO neutrino detector*, *Nucl. Instrum. Meth. A* **707** (2013) 45–53.
- [10] D. Cao et al., *Light Absorption Properties of the High Quality Linear Alkylbenzene for the JUNO Experiment*, *Nucl. Instrum. Meth. A* **927** (2019) 230–235, [1801.08363].
- [11] G. Adhikari et al., *Initial Performance of the COSINE-100 Experiment*, *Eur. Phys. J. C* **78** (2018) 107, [1710.05299].
- [12] SABRE collaboration, M. Antonello et al., *The SABRE project and the SABRE Proof-of-Principle*, *Eur. Phys. J. C* **79** (2019) 363, [1806.09340].
- [13] SNO+ collaboration, M. Anderson et al., *Measurement of the  $^8\text{B}$  solar neutrino flux in SNO+ with very low backgrounds*, *Phys. Rev. D* **99** (2019) 012012, [1812.03355].
- [14] SNO+ collaboration, M. Anderson et al., *Search for invisible modes of nucleon decay in water with the SNO+ detector*, *Phys. Rev. D* **99** (2019) 032008, [1812.05552].
- [15] SNO+ collaboration, M. Anderson et al., *Measurement of neutron-proton capture in the SNO+ water phase*, *Phys. Rev. C* **102** (2020) 014002, [2002.10351].
- [16] G. Goldstein and W. Lyon, *Liquid scintillators using 1-methylnaphthalene*, *The International Journal of Applied Radiation and Isotopes* **15** (1964) 133–137.
- [17] C. Aberle, C. Buck, F. Hartmann, S. Schonert and S. Wagner, *Light output of Double Chooz scintillators for low energy electrons*, *JINST* **6** (2011) P11006.
- [18] S. Song et al., *Feasibility study of a gadolinium-loaded DIN-based liquid scintillator*, *Journal of the Korean Physical Society* **63** (2013) 970–974.
- [19] C. Aberle, C. Buck, B. Gramlich, F. Hartmann, M. Lindner, S. Schonert et al., *Large scale Gd-beta-diketonate based organic liquid scintillator production for antineutrino detection*, *JINST* **7** (2012) P06008, [1112.5941].
- [20] A. J. Wright, *Robust Signal Extraction Methods and Monte Carlo Sensitivity Studies for the Sudbury Neutrino Observatory and SNO+ Experiments*. PhD thesis, Queen’s University, 2009.

- [21] B. Daniel, *Quality assurance testing, and sensitivity studies for the SNO+ experiment*, Master's thesis, Queen's University, 2018.
- [22] S. Wunderly and J. Kauffman, *New quench-resistant fluors for liquid scintillation counting*, *Nucl. Instrum. Meth. A* **41** (1990) 809–815.
- [23] H. O'Keeffe, E. O'Sullivan and M. Chen, *Scintillation decay time and pulse shape discrimination in oxygenated and deoxygenated solutions of linear alkylbenzene for the SNO+ experiment*, *Nucl. Instrum. Meth. A* **640** (2011) 119–122, [[1102.0797](#)].
- [24] B. von Krosigk et al., *Measurement of  $\alpha$ -particle quenching in LAB based scintillator in independent small-scale experiments*, *Eur. Phys. J. C* **76** (2016) 109, [[1510.00458](#)].
- [25] B. von Krosigk, L. Neumann, R. Nolte, S. Röttger and K. Zuber, *Measurement of the proton light response of various LAB based scintillators and its implication for supernova neutrino detection via neutrino-proton scattering*, *Eur. Phys. J. C* **73** (2013) 2390, [[1301.6403](#)].
- [26] H. Wan Chan Tseung, J. Kaspar and N. Tolich, *Measurement of the dependence of the light yields of linear alkylbenzene-based and EJ-301 scintillators on electron energy*, *Nucl. Instrum. Meth. A* **654** (2011) 318–323, [[1105.2100](#)].
- [27] A. Sørensen, S. Hans, A. Junghans, B. Krosigk, T. Kögler, V. Lozza et al., *Temperature quenching in LAB based liquid scintillator*, *Eur. Phys. J. C* **78** (2018) 9.
- [28] B. Von Krosigk, *Measurement of proton and  $\alpha$  particle quenching in LAB based scintillators and determination of spectral sensitivities to supernova neutrinos in the SNO+ detector*. PhD thesis, Dresden, Tech. U., Dept. Math., 2015.
- [29] A. Sørensen, *Temperature quenching in LAB based liquid scintillator and muon-induced backgrounds in the SNO+ experiment*. PhD thesis, Dresden, Tech. U., 2016.
- [30] S. Asahi, *The Rayleigh ratio measurement for SNO+ liquid scintillator and background studies of the internal ropes*, Master's thesis, Queen's University, 2012.
- [31] I. Berlman, *Handbook of fluorescence spectra of aromatic molecules*. Academic Press, New York and London, 1971.
- [32] C. Buck, B. Gramlich and S. Wagner, *Light propagation and fluorescence quantum yields in liquid scintillators*, *JINST* **10** (2015) P09007, [[1509.02327](#)].
- [33] J. Maniera, *Calibration and Monitoring for the Borexino Solar Neutrino Experiment*. PhD thesis, Lisboa, U., 2001.
- [34] P. Lombardi, F. Ortica, G. Ranucci and A. Romani, *Decay time and pulse shape discrimination of liquid scintillators based on novel solvents*, *Nucl. Instrum. Meth. A* **701** (2013) 133–144.
- [35] X.-B. Li et al., *Timing properties and pulse shape discrimination of LAB-based liquid scintillator*, *Chinese Physics C* **35** (2011) 1026–1032.
- [36] L. Bollinger and G. Thomas, *Measurement of the Time Dependence of Scintillation Intensity by a Delayed-Coincidence Method*, *Review of Scientific Instruments* **32** (1961) 1044–1050.
- [37] H. C. Tseung and N. Tolich, *Ellipsometric measurements of the refractive indices of linear alkylbenzene and EJ-301 scintillators from 210 to 1000 nm*, *Phys. Scripta* **03** (2011) 035701, [[1105.2101](#)].
- [38] I. Yeo et al., *Measurement of the refractive index of the LAB-based liquid scintillator and acrylic at RENO*, *Phys. Scripta* **82** (2010) 065706.

- [39] X. Zhou, Q. Zhang, Q. Liu, Z. Zhang, Y. Ding, L. Zhou et al., *Densities, isobaric thermal expansion coefficients and isothermal compressibilities of linear alkylbenzene*, *Phys. Scripta* **90** (2015) 055701, [[1408.0877](#)].
- [40] SNO+ collaboration, R. J. Ford, *A Scintillator Purification Plant and Fluid Handling System for SNO+*, *AIP Conf. Proc.* **1672** (2015) 080003, [[1506.08746](#)].