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# Effect of calcination temperature on the properties of CZTS absorber layer prepared by RF sputtering for solar cell applications

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Abstract In present work, we report synthesis of nanocrystalline Kesterite copper zinc tin sulfide (CZTS) films by RF magnetron sputtering method. Influence of calcination temperature on structural, morphology, optical, and electrical properties has been investigated. Formation of CZTS has been confirmed by XPS, whereas formation of Kesterite-CZTS films has been confirmed by XRD, TEM, and Raman spectroscopy. It has been observed that crystallinity and average grain size increase with increase in calcination temperature and CZTS crystallites have preferred orientation in (112) direction. NC-AFM analysis revealed the formation of uniform, densely packed, and highly interconnected network of grains of CZTS over the large area. Furthermore, surface roughness of CZTS films increases with increase in calcination temperature. Optical bandgap estimated using UV-Visible spectroscopy decreases from 1.91 eV for as-deposited CZTS film to 1.59 eV for the film calcinated at 400 °C which is quite close to optimum value of bandgap for energy conversion in visible region. The photo response shows a significant improvement with increase in calcinations temperature. The employment these films in solar cells can improve the conversion efficiency by reducing recombination rate of photo-generated charge carriers due to larger grain size. However, further detail study is needed before its realization in the solar cells.

**Keywords** CZTS films · RF sputtering · Kesterite · Calcination · XRD, XPS, NC-AFM

# Introduction

As on today, silicon (Si) has the lion's share in the photovoltaic industry. The main reason behind it is the huge availability of Si on the earth and a developed and established industry for making high-quality Si solar cells. However, the cost of Si solar cells is still high due to the high production cost of device quality Si. The photovoltaic market nowadays is demanding low production cost of material and hence of solar cells [1]. Therefore, it is necessary to reduce the material cost of solar cells which effectively reduces the cost of solar cells [2]. Several other direct bandgap semiconductor materials, such as copper indium gallium sulfides (CIGS), cadmium telluride (CdTe), etc, have been tried for solar cell application. However, these materials have their own problems like Cd and Te which are toxic, while Ga and In are expensive, which restrict the future development of solar cells. Copper zinc tin sulfide (Cu<sub>2</sub>ZnSnS<sub>4</sub>) or simply CZTS is one of the promising absorber materials in thin-film solar cell because of its excellent material properties for obtaining high efficiency such as direct bandgap (1.45 eV) [3], which is very close to optimum bandgap for solar energy conversion, high absorption coefficients (> $10^4$  cm<sup>-1</sup>) [4], etc. In addition, CZTS does not contain any toxic and expensive element, resulting in realizing of solar cell with less environmentally damaging and low cost. It is composed of naturally abundant and nontoxic elements [5]. The maximum theoretical power conversion efficiency of CZTS solar cells reported was 29.4% [6]. Wang and his group fabricated laboratory scale CZTSSe solar cell having area



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 $0.42 \text{ cm}^2$  with efficiency 12.6%, which is the highest conversion efficiency achieved until today [7].

There are two methods used for the preparation of CZTS films, chemical methods, and physical/vacuum-based methods. The chemical methods include several techniques, such as chemical spray pyrolysis [8], photochemical depositions [9], sol-gel technique [10], spin coating [11], electrodeposition [12], electro-spinning [13], and successive ionic layer adsorption and reaction (SILAR) [14], etc. The physical or vacuum-based method includes atom beam sputtering [15], e-beam and thermal evaporation [16], pulsed laser deposition [17], etc. Each method has its own advantages and limitations. Among these methods, RF magnetron sputtering has received considerable attention in recent years owing to its capability to synthesize device quality CZTS films. It permits deposition at low substrate temperature, gives the good adhesion, possibility of large area deposition, maximum uniformity, controllable thickness, precise in chemical composition control, matching with tradition solar cell production line, as well as easy scale-up than other CZTS thin-film deposition methods [18].

Properties of CZTS thin films are greatly influenced by pre- and post-annealing or calcinations treatment in various gas atmospheres. Recently, small work has been done on effect of pre- and post-annealing or calcinations treatment in various gas atmospheres on structural, optical, and electrical properties of CZTS thin films deposited by various methods. Recently, Seboui et al. [19] investigated post growth effect on properties of CZTS thin films prepared by spray pyrolysis and reported that the post-annealing effect reduces the optical transmission and increases the bandgap of CZTS films. Secondary phases may remain in the film after heat treatment. Ericson et al. [20] obtained highly crystalline CZTS films after annealing in H<sub>2</sub>S atmosphere. Surgina et al. [21] have also studied the annealing effect on structural and optical properties of CZTS films grown by pulsed laser deposition in N2 atmosphere. Vanalakar et al. [22] explained the post-annealing effect on grain size and surface morphology of CZTS thin films in the different gas atmospheres. Recently, Liu et al. [23] reported preheating effect on CZTS film properties. Most of the authors reported the effect of pre- or post-annealing of CZTS films either at high temperature or in presence of toxic or hazardous gases. To best of our knowledge, low temperature calcination of CZTS in inert gas atmosphere is missing till date. With this motivation an attempt has been made to investigate low post calcination effect (>400 °C) in inert gas (Ar) atmosphere on structural, optical, morphology and electrical properties of CZTS thin films deposited by RF magnetron sputtering. It has been observed that by increasing calcination temperature in Ar atmosphere, it is possible to grow highly uniform, large area ( $\sim 4 \text{ cm}^2$ )



nanocrystalline kesterite-CZTS films with optimum bandgap ( $\sim 1.59 \text{ eV}$ ) which can be useful for enhancing the efficiency of CZTS solar cells.

# **Experimental details**

#### Film preparation and calcination

CZTS films were deposited on corning #7059 substrates using indigenously design and locally fabricated RF magnetron sputtering technique. It consists of a cylindrical stainless steel chamber (process chamber) coupled with a turbo molecular pump (TMP) followed by a roughing pump which yields a base pressure less than  $10^{-7}$  Torr. A target (Cu:Zn:Sn:S of 1.1:1.1:1.1:3) of 4 inch diameter (99.99%, RND-KOREA, Korea), 3 mm thick was used for the deposition of CZTS films and was kept facing the substrate holder  $\sim 9$  cm away. In order to get film uniformity substrates were kept rotating during the sputtering process using a stepper motor at a rate 12 rpm using speed controller. The substrate temperature was kept constant at 100 °C using the in-built thermocouple and temperature controller. The substrates can be clamped on the substrate holder which is heated by in-built heater using thermocouple and temperature controller. The pressure during deposition was kept constant by using automated throttle valve and measured with the capacitance manometer. For sputtering argon gas was introduced into the process chamber through a specially designed gas bank assembly which consists of mass flow controllers (MFCs) and gas mixing. The process parameters employed during the deposition of CZTS films are listed in Table 1.

Before each deposition the substrates were cleaned using a standard cleaning procedure using piranha solution. Prior to deposition, the substrate holder and deposition chamber were baked for two hours at 100 °C to remove any water vapor absorbed on the substrates and to reduce the oxygen contamination in the film. Sputter-etch of 10 min were used to remove the target surface contamination. As-deposited CZTS films then calcinated at different temperatures in argon atmosphere for 90 min in a cylindrical stainless steel chamber without air-break. During calcination the argon flow rate and pressure were kept constant at 50 sccm and 20 mTorr respectively. After calcination films were allowed to cool to room temperature in vacuum and then taken out for characterization.

# Film characterization

X-ray diffraction patterns were obtained by X-ray diffractometer (Bruker D8 Advance, Germany) using CuK $\alpha$  line ( $\lambda = 1.54$  Å) at a grazing angle of 1°. Raman spectra were

 Table 1 Process parameters employed during the deposition of CZTS films

Process parameter	Value
Deposition pressure	$6 \times 10^{-3}$ bar
Deposition time	60 min
RF power	100 W
Distance between substrate holder and target electrode	9 cm
Ar gas flow rate	60 sccm
Substrate temperature	100 °C

recorded in the range of  $100-600 \text{ cm}^{-1}$ . The spectrometer has the back-scattering geometry for detection of Raman spectrum with the resolution of  $1 \text{ cm}^{-1}$ . The excitation source was 532 nm line of He-Ne laser. The power of the Raman laser was kept less than 5 mW to avoid laser-induced crystallization of the film. The HR-TEM and SAED patterns were recorded using TECNAI G2-20-TWIN, transmission electron microscope operating at 200 kV. The optical bandgap of CZTS was deduced from absorbance spectra and was measured using a JASCO, V-670 UV-Visible spectrophotometer. The surface topology of the films was investigated NC-AFM (JEOL, JSPM-5200). The XPS spectra were recorded using a VSW ESCA instrument with a total energy resolution  $\sim 0.9$  eV fitted with an Al Ka source (soft X-ray source at 1486.6 eV) at base vacuum  $>10^{-9}$  Torr. The XPS signal was obtained after several scans in the acquisition process. The spectra were recorded for the specific elements (Cu, Zn, Sn, S). The photo response measurement of the CZ TS films was studied using a Keithley 2401 system. For light illumination, a PEC-L01 Portable Solar Simulator was used. Thickness of films was determined by profilometer (KLA Tencor, P-16+).

# **Results and discussion**

# X-ray diffraction (XRD) analysis

Figure 1 show the low angle-XRD pattern of as-deposited and calcinated CZTS films. As seen from the XRD pattern, as-deposited and calcinated CZTS film at 200 °C show only a diffraction peak at  $2\theta \sim 28.65^{\circ}$ . The CZTS film calcinated at 300 °C shows a tiny diffraction peak at  $2\theta \sim 31.83^{\circ}$  along with a diffraction peak at  $2\theta \sim 28.65^{\circ}$ . At 400 °C, the x-ray diffraction pattern shows three diffraction peaks at  $2\theta \sim 28.65^{\circ}$ ,  $31.83^{\circ}$  and  $58.98^{\circ}$  which are corresponding to (112), (200) and (224) planes of the kesterite-CZTS structure [JCPDS data card# 26-0575]. No peak of any other phase was found in XRD pattern. With increase in calcination temperature intensity of (112)

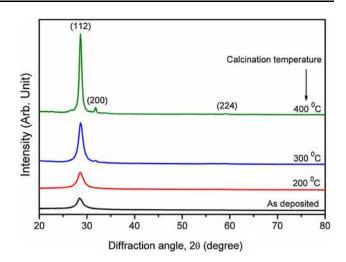


Fig. 1 XRD pattern of as-deposited and after calcination CZTS thin films at different temperatures for kesterite-CZTS crystal structure [JCPDS data card # 26-0575]

diffraction peaks increases whereas its sharpness gets reduced. The peak at  $2\theta \sim 28.65^{\circ}$  has highest intensity among all other peaks which indicate that CZTS crystallites have preferred orientations in (112) direction. The average grain (dx-ray) size has been estimated using the classical Scherer's formula [24],

$$d_{\rm x-ray} = \frac{0.9\,\lambda}{\beta\cos\theta_{\beta}} \tag{1}$$

where  $\lambda$  is the wavelength of the x-ray used,  $\beta$  is full-width at half-maximum (FWHM) and  $\theta$  is the Bragg diffraction angle. The average microstrain ( $\varepsilon$ ) developed in the asprepared and calcinated CZTS films was calculated by using the relation [25],

$$\varepsilon = \frac{\beta \, \cos \theta}{4} \tag{2}$$

The calculated structural parameters are presented in Table 2. As seen from Table 2, the average grain size of CZTS films increases with increasing calcination temperature. This may be due to coalescence and reorganization of grains with increase in calcination temperature. It is further supported by observed decrease in macrostrain with increase in calcination temperature. The coalescence and reorganization of grains fill voids due to calcination and the film become denser. The atomic force microscopy (AFM) analysis further supports this conjecture (discussed later).

# Raman spectroscopy analysis

Figure 2 shows Raman spectra of as-deposited and calcinated CZTS films in the range 100–600 cm<sup>-1</sup>. As seen for as-deposited film the dominant Raman peak is observed at  $\sim 338 \text{ cm}^{-1}$  which is in consistent with the peak of CZTS mono-grain powder reported in literature [26]. With



Calcination temperature (°C)	$d_{x-ray}$ (nm)	FWHM (°)	$d_{\rm hkl}$ (Å)	3
As deposited	7.66	1.07	3.20	$4.23 \times 10^{-3}$
200	7.67	1.07	3.12	$4.52 \times 10^{-3}$
300	10.52	0.78	3.11	$3.29 \times 10^{-3}$
400	16.83	0.49	3.11	$2.06 \times 10^{-3}$

**Table 2** Structural parameters, average grain size  $(d_{x-ray})$ , full-width half-maxima (FWHM), inter-planar spacing  $(d_{hkl})$ , and microstrain ( $\varepsilon$ ) for CZTS films

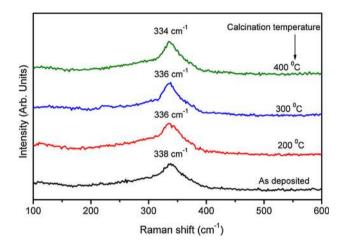


Fig. 2 Raman spectrum of kesterite CZTS of as deposited and after calcination CZTS thin films at different temperatures. The observed featured peaks are at 344, 336, and 338 cm<sup>-1</sup> and indicates that CZTS has synthesized without any impurity phases

an increase in calcination temperature the Raman peak shift towards lower wavenumber. Thus, CZTS film calcinated at 400 °C, show the Raman peak ~334 cm<sup>-1</sup>. The presence of internal compressive stress may responsible for shifting of Raman peak towards lower wavenumber. In addition, shrinking of substrate while cooling may also contribute to the internal strain in CZTS film [27]. The presence of the internal strain has been confirmed by low angle-XRD analysis (See Table 2). Normally, the main Raman peaks from the different phases potentially present in CZTS system are CuS ( $\sim 267 \text{ cm}^{-1}$ ), SnS ( $\sim 220 \text{ cm}^{-1}$ ), SnS<sub>2</sub>  $(\sim 202 \text{ cm}^{-1}), \text{ Sn}_2\text{S}_3 (\sim 234 \text{ cm}^{-1}), \text{ ZnS} (\sim 219 \text{ cm}^{-1}),$  $Cu_2SnS_3$  (~290 cm<sup>-1</sup>) and  $Cu_2ZnSnS_4$  (~337 cm<sup>-1</sup>) [28].However, in present study, we have observed only a Raman peak in the range 334–338 cm<sup>-1</sup> suggesting the existence of single phase CZTS films. Low angle-XRD results further support this.

# X-ray photoelectron spectroscopy (XPS) analysis

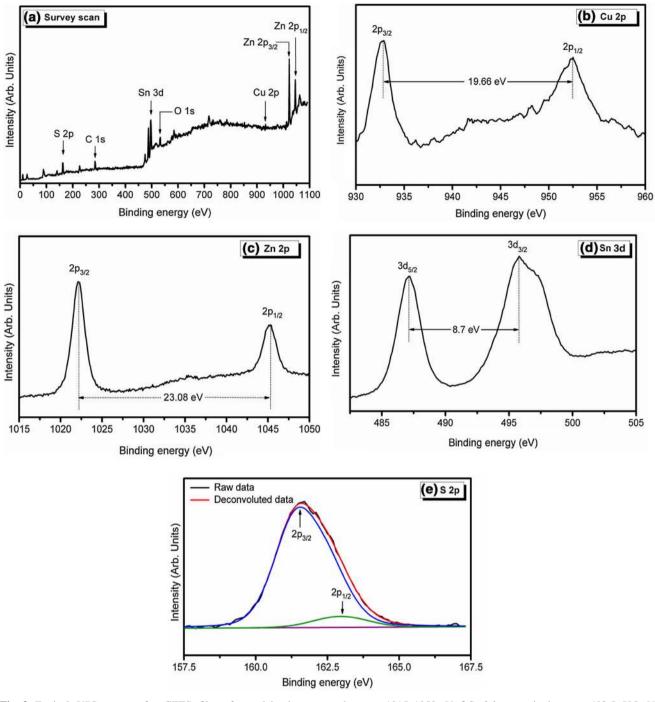
X-ray photoelectron spectroscopy (XPS) is very sensitive to the chemical composition and environment of the elements in a material. Figure 3a shows XPS survey spectra



(0-1100 eV) of CZTS film calcinated at 400 °C. The core level peaks corresponding to the elements copper (Cu 2p), Zinc (Zn 2p), Tin (Sn 3d) and sulfur (S 2p) can be visibly seen in the spectra. In addition, it also shows the presence of C peak as well as O peak at  $\sim 285$  and  $\sim 531$  eV respectively as impurities. These contaminants have also been identified in CZTS films by other workers [26]. The binding energy peak observed at ~932.80 eV corresponding to the Cu  $2p_{3/2}$  core level in CZTS film [29]. Figure 3b show high-resolution XPS spectrum of Cu 2p consisting two narrow and symmetric peaks at  $\sim 932.80$ and  $\sim$  952.46 eV, indicative of Cu(I) with a peak splitting of 19.66 eV. These results are well matched with the previously available data in the literature [30]. The core line of Zn exhibited a doublet at ~1022.17 eV and ~1045.25 eV corresponding to Zn  $3d_{5/2}$  and Zn  $3d_{3/2}$  peaks with spinorbit separation 23.08 eV (Fig. 3c) suggesting presences of Zn(II) in CZTS compound [31]. Figure 3d shows two peaks, one at ~495.80 eV corresponding to Sn  $3d_{3/2}$  core level in CZTS and other at  $\sim$  487.10 eV from Sn in SnS<sub>2</sub> phase [32] with spin-orbit separation 8.70 eV. Two S 2p peaks are located at ~161.5 and ~ 163.02 eV, respectively, showing a peak separation of 1.52 eV (Fig. 3e), which is also consistent with literature value for a metal sulfide [33]. Thus, XPS analysis confirms the formation of single phase CZTS films.

### Transmission electron microscopy (TEM) analysis

Figure 4 shows detailed TEM analysis of CZTS film calcinated at 400 °C. Figure 4a shows low resolution TEM images of CZTS film calcinated at 400 °C. As seen spherical CZTS nanoparticles having diameter in the range 5–25 nm and their agglomerates can also be clearly observed. The histogram of particle size distribution is plotted in Fig. 4b. The high-resolution TEM image shown in Fig. 4c indicates that CZTS nanoparticles are crystalline in nature. The inset figure shows the lattice fringes with the inter-planar distance of 0.31 nm which belong to (112) plane of kesterite-CZTS [34]. The electron diffraction pattern for the selected area (SAED) pattern shown in



**Fig. 3** Typical XPS spectra for CZTS film after calcination at 400 °C. **a** Survey scan in the range 0-1100 eV. **b** High-resolution XPS spectrum of Cu 2p in the range 930–960 eV. **c** Zn 2p spectra in

the range 1015–1050 eV. **d** Sn 3*d* spectra in the range 482.5–505 eV. **e** S 2p spectra in the range 157.5–167.5 eV

Fig. 4d is also in agreement with the low angle-XRD analysis, comprising of diffraction rings corresponding to the (112) and (200) planes of kesterite-CZTS. The diffraction ring corresponding to (224) plane is not clearly visible in SEAD pattern because of its weak intensity as seen in low angle-XRD pattern.

# Atomic force microscopy (AFM) analysis

Figure 5 shows surface topography of as-prepared and calcinated CZTS films investigated by non-contact atomic force microscopy (NC-AFM). The scan area for all AFM micrographs was 25  $\mu$ m<sup>2</sup> (5 × 5  $\mu$ m<sup>2</sup>). NC-AFM micrograph of



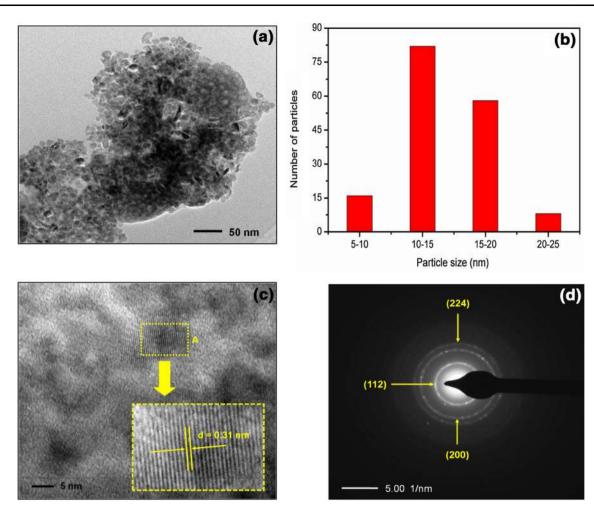


Fig. 4 a TEM image with low magnification, b particle size distribution histogram, representing particle size ranging from 5 to 25 nm, c HR-TEM, demonstrating nanoparticles have lattice fringes

as-prepared CZTS film revealed textured surface with tiny uniform island-like topography. It has been reported that such topography originates from the island growth of the Volmer-Weber mode and the kinetic energy at low temperature is not sufficient for the coalescence of island-like crystallites [35]. The root mean square (rms) surface roughness of as-prepared CZTS film was found  $\sim 0.32$  nm. The CZTS film calcinated at 200 °C (Fig. 5b) clearly indicate that these tinv textured island coalescence to form superstructure of smaller clusters with increased rms surface roughness to  $\sim 0.76$  nm. The coalescence of textured island may occur due to increase in the surface mobility with increase in calcination temperature. Further increase in calcination temperature to 300 °C and 400 °C (Fig. 5(c, d) one can have observed that these smaller clusters of CZTS are bound together and formed into non-uniform larger clusters with enhanced surface roughness. Thus, CZTS films calcinated at 300 °C and 400 °C the rms surface roughness was found  $\sim 0.60$  and 3.21 nm respectively. Therefore, from



with inter-planar distance of 0.31 nm, and **d** SAED pattern of concentric rings correspond to three major peaks in XRD of Kesterite-CZTS nanoparticles

AFM analysis it has been concluded that with increase in calcination temperature the particle size and surface roughness of CZTS film increases.

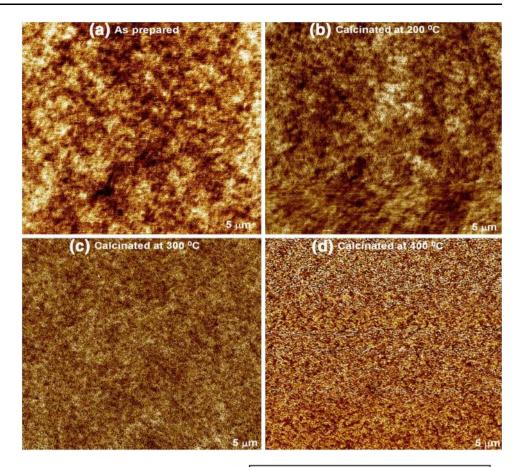
# UV-Visible spectroscopy analysis

The optical properties of as-deposited and calcinated CZTS films were investigated from UV–Visible spectroscopy. The optical absorption coefficient ( $\alpha$ ) can be calculated from the transmittance (*T*) and reflection (*R*) of the films with the formula [36],

$$\alpha = \frac{1}{d} \ln \left( \frac{T}{1 - R} \right) \tag{3}$$

where d is the thickness of the films. Figure 6a display the variation of absorbance for as-deposited and calcinated CZTS thin films. The optical absorption coefficient after calcinations CZTS films was found  $>10^4$  cm<sup>-1</sup> in the visible region indicating a direct bandgap characteristic of CZTS films.

Fig. 5 Two-dimensional (2D) AFM images of as-deposited and calcinated CZTS films. The scan area is 25  $\mu$ m<sup>2</sup> (5 × 5  $\mu$ m<sup>2</sup>)

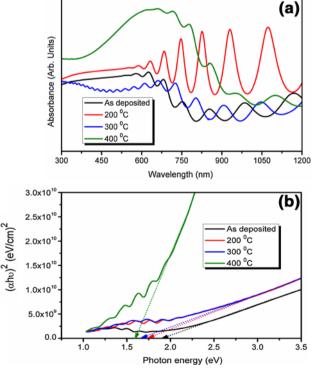


In the direct transition semiconductor, the optical energy bandgap ( $E_{opt}$ ) and the optical absorption coefficient ( $\alpha$ ) are related by [37],

$$(\alpha E)^{1/2} = B^{1/2} (E - E_{opt})$$
(4)

. ...

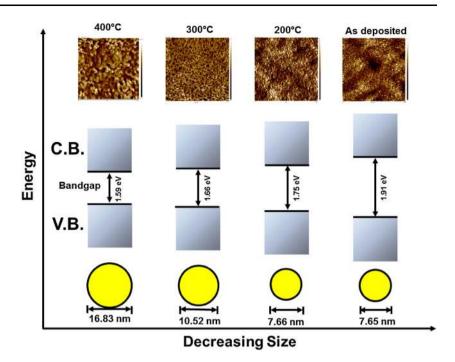
where  $\alpha$  is the absorption coefficient, B is the optical density of state and E is the photon energy. Therefore, optical bandgap can obtained by extrapolating the tangential line to the photon energy (E = hv) axis in the plot of  $(\alpha h \upsilon)^2$  versus photon energy (h $\upsilon$ ). Figure 6(b) shows plot of  $(\alpha h \upsilon)^2$  versus photon energy (h $\upsilon$ ) (Tauc plot) of asdeposited and calcinated CZTS films. As seen from the figure with increase in calcination temperature optical bandgap of CZTS films decreases from 1.91 to 1.59 eV. The obtained bandgap values are consistent with the bulk value of CZTS (1.45-1.90 eV) [38]. The main factor that affect the band gap of CZTS films are the average grain size [39] and presence of multiple phases of CZTS in the films [40]. As revealed from our low angle-XRD (Fig. 1) and Raman spectroscopy (Fig. 2) analysis the existence of multiple phases CZTS in the films are ruled out. Therefore, decrease in optical bandgap of CZTS films can attribute to increase in average grain size. Graphical presentation of dependence bandgap on average grain size is shown in Fig. 7. The optical bandgap of CZTS film calcinated at



**Fig. 6 a** UV–Visible NIR spectra recorded showing absorbance versus wavelength plot for CZTS thin films. **b** Plot of  $(\alpha h \upsilon)^2$  versus photon energy (h $\upsilon$ ) to determine optical bandgap for as-deposited and calcinated CZTS films



Fig. 7 Graphical presentation of dependence bandgap on average grain size of asdeposited and calcinated CZTS films



400 °C is  $\sim 1.59$  eV which is quite close to the optimum value bandgap for photovoltaic solar conversion in the visible region of solar spectrum.

#### Photo response measurement

Figure 8 shows the current versus time (I–t) plot as-deposited and calcinated CZTS films at constant 0.2 V bias voltage under dark and illumination conditions. For electrical properties measurement we have used samples of area  $0.5 \text{ cm}^2$ . As seen there is significant improvement in the current with increase in calcinations temperature. The improvement in electrical properties may attribute to improvement in crystalline nature, texture, grain size of CZTS films with increase in calcinations temperature. Such

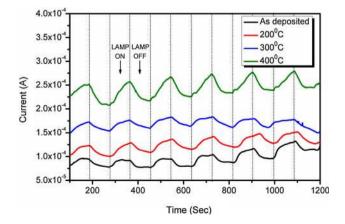


Fig. 8 Current versus time (I–t) plot of as-deposited and calcinated CZTS films at constant 0.2 V bias voltage for 60 s illuminations and for 60 s dark condition



larger grains CZTS thin films can be useful as an absorber layer for the improvement in photoelectric conversion efficiency because larger grain sized which can reduce the recombination rate of photo-generated charge carriers [41].

# Conclusion

In summary, nanocrystalline CZTS films have been prepared by home-made RF magnetron sputtering technique. Influence of calcination temperature in Ar atmosphere on structural, morphological, electrical and optical properties on CZTS films has been investigated. Formation of CZTS has been confirmed by x-ray photoelectron spectroscopy (XPS) whereas formation of Kesterite-CZTS films has been confirmed by X-ray diffraction (XRD), transmission electron microscopy (TEM) and Raman spectroscopy. We found that the calcination process has a great influence on growth and nucleation of grains. XRD analysis revealed that the crystallinity and average grain size increases with increase in calcination temperature. Raman spectroscopy analysis show shifting of Raman peak shift towards lower wavenumber with increase in calcination temperature. The presence of internal compressive stress and shrinking of substrate during cooling may responsible for shifting of Raman peak towards lower wavenumber. However, shrinking of substrate while cooling has not been verified experimentally. Detail surface study (morphology and topology) reveal that CZTS thin films have densely packed and a highly interconnected network of grains with large area (4 cm<sup>2</sup>). AFM show significant difference in surface

topography of CZTS films with change in calcination temperature. Increase in calcination temperature show increase in rms and average surface roughness of the CZTS films. UV-Visible spectroscopy analysis revealed that the absorption coefficient of as-deposited and calcinated CZTS films are in the range  $10^4 - 10^5$  cm<sup>-1</sup> in the visible region. The bandgap show decreasing trend with increase in calcination temperature (1.91–1.59 eV). The bandgap of CZTS film annealed at 400 °C was found  $\sim 1.59$  eV which is quite close to the optimum value for photovoltaic solar conversion in the visible region of solar spectrum. It is found that the photo response depends upon the grain size effect, whereas photo response increases with the increase of the grain size. Employment these films as an absorber layer in CZTS solar cells can improve the conversion efficiency by reducing recombination rate of photo-generated charge carriers due to increased grain size.

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