

EFFECTS OF SOLAR RADIATION PRESSURE AND AERODYNAMIC FORCES ON SATELLITE— ATTITUDE DYNAMICS AND THEIR UTILIZATION FOR CONTROL: A SURVEY

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ABSTRACT

The environment exerts an important influence on the performance of space systems. A brief review of most of the studies, presented over the past eighteen years, relating to the influence and the possible utilization of the solar radiation pressure and aerodynamic forces, with particular reference to attitude dynamics and control of satellites is presented here. The semi-passive stabilizers employing these forces show promise of long life, low power and economic systems, which though slower in response, compare well with the active controllers. It is felt that much more attention is necessary to the actual implementation of these ideas and devices: some of which are quite ingenious and unique.

Key words: Solar radiation pressure, rarefied atmosphere, earth satellites, attitude dynamics, semi-passive control, gravity-gradient stabilization, flexibility, orbital perturbations.

INTRODUCTION

The application of space technology for development is now well recognized. The success of the missions, however, depends, to a great extent, on the stability of orbits and satellite orientations. The environmental forces present a major source of errors, especially for passive controllers like gravity-gradient systems which can provide only small correcting torques but are gaining increasing interest because of their simplicity and economy. The effects have been a subject of considerable investigation over the past eighteen

years with recent attempts at utilizing the perturbative forces to advantage. A brief account of some of the work relating to the attitude dynamics is included in the reviews by Roberson [1], Frye and Stearns [2], Ergin [3], Sabroff [4], and Shrivastava, Tscham and Modi [5]. A more complete, updated review of the investigations dealing with the environmental influence and their possible applications is presented here.

The orbital perturbation due to the environmental forces have been studied by a number of investigators [16-14 *et al.*]. Though very important this aspect falls beyond the scope of the present discussion. Here the attention is confined to attitude dynamics and control of satellites.

The relative magnitudes of torques due to various forces arising from gravity-gradient, solar radiations, earth and earth-reflected radiations atmospheric and magnetic forces, cosmic dust, etc., depend on the orbital elements and the satellite's shape, size, surface conditions, mass distribution and orientation. The actual systems are rather complex. For some simpler, shapes like plates, spheres and cylinders some of the torques have been mathematically modelled, through several assumptions, to a varying degree of accuracy by Roberson [15], Hall [16], Evans [17], Wiggins [18], Clancy and Mitchell [19], Hughes [20], Tidwell [21], Flanagan and Modi [22], and others. Figure 1, taken from Ref. [5], shows a comparison of maximum torques, at various altitudes, acting on a typical passive satellite (GEOS-A). It is apparent that for near-earth systems the atmospheric torques and for higher altitude those due to direct solar radiation pressure become comparable to the control moment provided by the gravity-gradient. For larger area/mass ratios, typical of the next generation of satellites, both these influences would tend to be much more significant. As such any analysis and design must include these considerations.

Noting the importance, which is further substantiated by the flight experience of Echo [23], Vanguard [6], Explorer XII [24], Aloutte [25, 26], GEOS [27, 28], Proton-2 [29], SKYNET [30] and several other satellites, a large number of investigations have been undertaken. These can be broadly classified as under:

1. Attitude perturbations due to
[a] Radiation pressure and [b] Aerodynamic forces.
2. Attitude stabilization and control using
[a] Solar radiation pressure, [b] Aerodynamic forces.

The studies have essentially proceeded along two lines: either detailed analytical and numerical studies of simple models, or analysis of complex

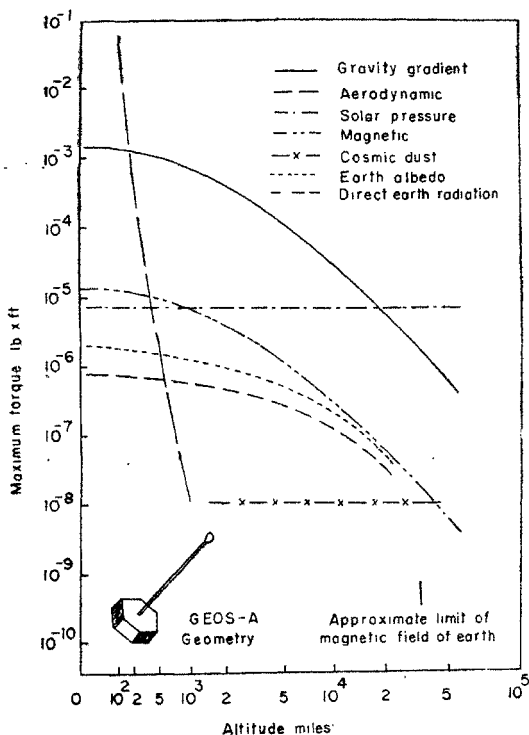


FIG. 1. Magnitude of environmental torques on a typical satellite⁵.

systems using simplified equations. On the experimental and hardware development aspects the work is still in the preliminary stages though a number of controllers have been proposed and the feasibility is established by the flight of COSMOS-149 [31], which employed atmospheric forces for its pitch control, and Mariner-IV [32], which demonstrated the usefulness of the solar radiation pressure. In the following sections the important aspects of most of the papers published on the subject are briefly reviewed.

1.1. Attitude Perturbations due to Radiation Pressure

According to the quantum theory of electromagnetic radiation, the pressure exerted by the radiation flux incident on a surface is the rate of change of momentum of the photons. A totally absorbent surface placed normal to the incident flux experiences a pressure: $P = m_p c'$ where m_p is the mass of photons incident per unit area and c' is the velocity of light. From Einstein's energy relation, the pressure in terms of the solar constant S is

$$P = S/c'$$

The pressure is a function of surface reflectivity, absorptivity and transmissibility, as well as the position relative to the emitting body. These properties themselves are functions of the wavelength and angle of incidence. The variations, however, are usually low and may be ignored [22].

The existence of the force has been known for a long time. Its magnitude being small [10^{-5} Newtons/m², for a perfectly reflecting surface] the designers of early satellites did not pay much attention. The attitude and orbital perturbations of Echo and Vanguard and spin-up of Explorer XII, which were later attributed to solar radiation pressure by Bryant [23], Evans [17], and Fedor [24], exhibited the necessity of a careful consideration of this effect in design and operation of even small satellites, especially those having only small control torques. Hall [16] derived an empirical force expression for direct solar radiation of the form $F = PAC_f$, where A is the surface area and C_f is the force coefficient which is found to be less than two for several convex shapes. McElvain [33] obtained an analytical expression for direct solar radiation torque and determined, for two geometries, the change in the satellite's angular momentum necessary to maintain a specified orientation. In a later study [34] ways to minimize the solar torque on a gravity-stabilized satellite, ATS, were suggested. The method involves optimal balance of surfaces and their characteristics. Karymov [35] derived equations of librational motion of a sun-orbiting satellite and analysed the stability of equilibrium along the local vertical. The derivations, however, are needlessly complicated. Clancy and Mitchell [19] undertook a more complete analysis accounting for three major sources of radiations, namely, the sun, earth and reflections from the earth and its atmosphere. The dynamical behaviour of the system is found as the motion of and about the angular momentum vector. In addition to the inherent limitations of the analysis, the resulting force expressions, given in an integral form, have to be evaluated numerically. The computation times involved appear to be very large and may render a comprehensive study of attitude dynamics virtually impossible [36].

A precise evaluation of these forces was undertaken by Flanagan and Modi [22] with particular reference to a flat plate model of a satellite, executing planar librations in an ecliptic orbit. The closed form character of the resulting expressions for the forces makes them ideally suited for satellite attitude dynamics studies. The system response under only direct solar radiations was analysed in detail using W.K.B.J. approach [37]. Later in conjunction with the concept of integral manifold in state-space, they were able to establish the altitude bounds where various environmental forces become significant [36, 38-40]. Figure 2, one of their results, shows the rapid reduction in the stability bounds and maximum allowable eccentricity for a relatively more stable gravity-gradient satellite. Modi and Kumar [41-43] concentrated upon direct solar radiation pressure, acting on a more realistic model—a cylindrical satellite. For the planar librations the system was analysed using digital as well as analog simulations [41]. The cross-plane motions in circular orbits were also included in a later study [42]. Here again the radiation reduces the stability region significantly. It may be mentioned here that a possibility of excitation of resonant motion across the orbital plane, due to radiations, was suggested for GEOS-II [27]. The observations confirmed it [28]. The importance of the influence on librational dynamics of a slowly spinning system is also established recently by Modi and Pande [46]. All these analyses, through extensive computations, clearly point out that the radiation pressure can play as important a role as the system inertia, orbital eccentricity, and spin, if any.

The designers, who often want a simple working model of a complex system, may find the study by Hodge [45], who, using the model given by Tidwell [21] for gravity, magnetic, aerodynamic, eddy-current and solar pressure torques on a rolling hexagonal satellite, pointed out the adequacy of a linearised analysis for short-term [1/3 orbit] predictions. This is mainly because the effects are small.

All the studies mentioned above assume the satellite to be rigid. The actual systems, especially those being designed now are far from this idealization. Their elasticity coupled with heating and presence of environmental forces can significantly affect the performance and success of the missions. Such a possibility, even for a small satellite containing elastic parts [e.g., Alouette and Explorer XX] was established by Etkins and Hughes [25]. A large number of studies that followed emphasized the behaviour further and resulted in many improvements in the components like booms and panels which may undergo very large deflections [46]. The combined effect of solar radiation pressure and differential heating influence the system response and stability rather adversely [47-48 *et al.*].

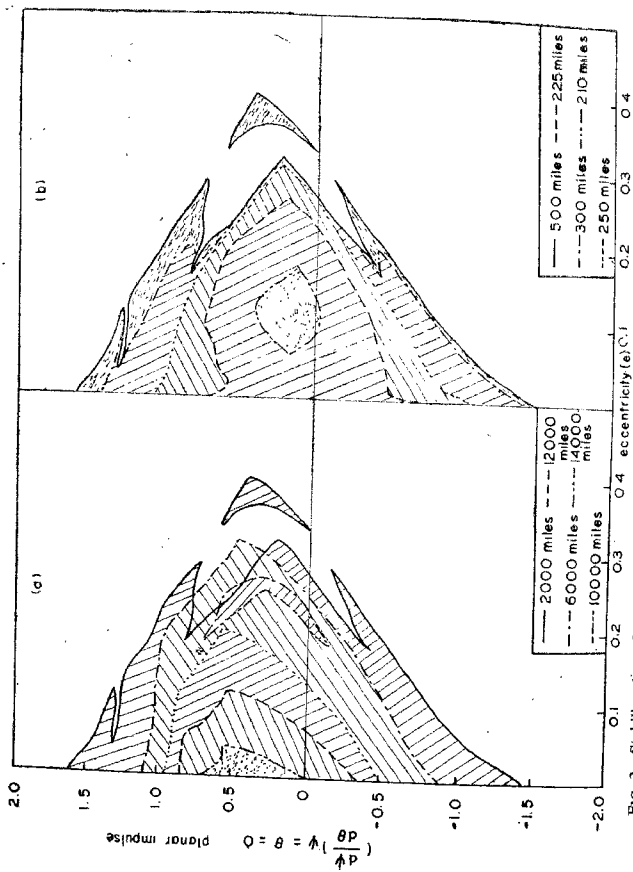


FIG. 2. Stability charts for planar librations of GEOS-A showing (a) Direct solar radiation at higher altitudes, (b) Aerodynamic effect at lower altitude.

The vast amount of literature, dealing with flexibility-control-interaction problems, thermoelastic behaviour of space structures, and analyses of flexible satellites subjected to environmental forces are well reviewed in a number of reports, *i.e.*, NASA special publication [49], which is briefly summarized by Noll *et al.* [50], and papers by Ashley [51], Likins and Bouver [52], Hughes [53] and most recently by Modi [54]. These should prove valuable to the designers.

1.2. Attitude Perturbations due to Aerodynamic Forces:

A number of studies indicate that the earth's atmosphere, though rarefied, plays a significant role in the satellite dynamics even at the altitudes of about 600-800 km. Due to large speeds of vehicles the drag becomes quite significant. The physics of the interaction with the free molecular flow is rather complex. The force depends on a number of surface and environmental parameters, including temperatures, reflectivity, Knudsen number, angle of incidence, etc. [55]. A simple model, extending relations in the continuum flow, can be made with the drag coefficient determined experimentally. In most cases the coefficient turns out to be about 2. With this approximation many studies relating to the equilibrium, response and stability of satellites of various shapes have been carried out.

Debra [56] discussed the variation of equilibrium for two gravity-stabilized satellites. Beletskii [57] treated the force as a small perturbative source acting on a rapidly spinning system. Through a linearized analysis Schrello [58] pointed out that the aerodynamic torques may exceed gravity gradient moment even at altitudes nearing 500 km. Evans [17] presented the disturbances in the fundamental form of pressure and shear stresses. The influence of a constant moment acting on a gravity stabilized system was determined through an infinitesimal analysis by Garber [60]. More directly Sarychev [60, 61] with a particular reference to the Russian Satellites, derived equations of motion and determined the necessary and adequate conditions for asymptotic stability of the eigen oscillations which are also caused by the rotating atmosphere. Meirovitch and Wallace [62] established the regions of guaranteed stability for a slowly spinning system with a constant aerodynamic force. For two configurations, the stability of equilibrium positions was established using Liapunov's direct method. Morozov [63, 64] included the magnetic forces and found conditions for stability of the steady state behaviour of a gyrostat.

Most of the studies discussed above consider a simpler case of axsymmetric satellites moving in circular orbits. A majority of the actual systems,

however, do possess inertial as well as geometric non-symmetries. In such cases, the interaction with the atmosphere may result in instabilities, as shown by Nurre [65]. The center of mass was still assumed to be in the plane of geometric symmetry. Frik [66], extending it to an arbitrarily shaped body, found that if the aerodynamic forces were conservative at least one stable equilibrium would exist. For non-conservative forces no stable position is possible.

The planar librations of gravity-stabilized satellite modelled like a plate moving in an elliptic orbit and subjected to various radiations and aerodynamic forces were investigated by Flanagan and Modi [40]. A more complete analysis, accounting for the transverse motion as well, for a cylindrical satellite moving in an arbitrary orbit and subjected to gravity gradient and aerodynamic torques is presented by Shrivastava and Modi [67-71]. After establishing the stability of equilibrium through infinitesimal as well as Liapunov's approach, detailed response and stability studies were carried out using the digital [67] and analog simulations [68]. The design plots of the allowable disturbances for non-tumbling motion indicate a rapid reduction in the stability region. The concept of integral manifold is successfully extended to the axisymmetric satellites in circular orbits. Three distinct types of motions corresponding to the fundamental and other periodic solutions are noted. The critical Hamiltonians representing the strength of the maximum permissible disturbance for many systems are established through Floquet's theory [70]. An approximate solution in terms of elliptic functions is also found using the constant first integral of the system [71]. Under the combined effect of eccentricity and atmosphere the equilibrium changes continuously and the response gets much more complicated. The stability bounds also shrink rapidly [72] [Fig. 3]. This puts a severe limitation on the usefulness of gravity gradient system for near-earth systems. Even for relatively stable dual-spin spacecraft of the type of SAS-A the effects of environment could not be ignored [73, 74]. Small attitude perturbations, mainly due to the atmosphere, were noted for this satellite which was put in a 500 km circular orbit.

The studies mentioned above are mostly deterministic in their approach and assume a rather simple model for a complex and uncertain phenomenon. Better results may be possible if one uses statistical methods. The study by Sheporaitis [76], who finds a parametric stability region using stochastic Liapunov function and gives probability estimates of the convergence of equilibrium position, should be of interest to the investigators,

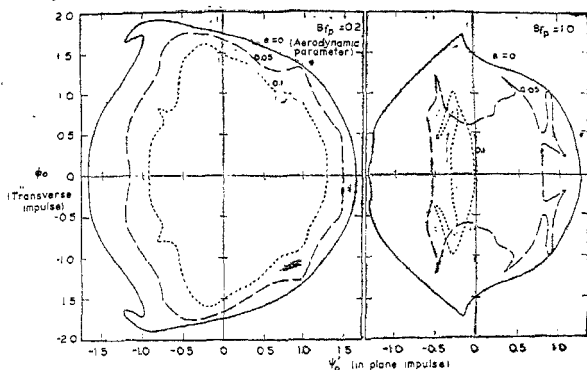


FIG. 3. Effect of aerodynamic torque and orbit eccentricity on the allowable impulsive disturbance for stable coupled motion of slender satellite⁷⁸.

All the results, particularly the latter few, indicate the significant role the forces due to the rarefied atmosphere can play, and emphasize their careful consideration in design.

2.1. Attitude Stabilization and Control Using Solar Radiation Pressure:

The solar radiation pressure is found to affect the performance of the spacecrafts adversely. A judicious design of the system can change its role-making it useful for stabilization and control. Such a possibility was first suggested by Garwin [76] who proposed "Solar-Sailing" for interplanetary missions. Though requiring a huge surface, it was shown to take less time than the chemical rockets for the distant planets. It was also thought to be simpler in implementation and operation than its competitor: ion propulsion. Sohn [77] suggested the use of a weathervane type solar attitude stabilizer. Frye and Stearns [2] thought of a trailing cone for the purpose. These were felt to be rather cumbersome appendages and Newton [78] preferred a satellite to be a big sphere having two types of coatings — one portion reflecting and the other absorbing. The local heating in such a system, however, could create problems. A possibility of increasing the effectiveness of the available force by focussing it through a set of reflecting and collecting mirrors was suggested by Hibbard [79]. Ule [80] designed

and built an array of wind-mill type mirrors with corner reflectors which could maintain the axis of rotation sun-oriented and could regulate the spin-rate. In absence of damping the effectiveness of all these systems would be limited. Accord and Nikalas [81] evolved a unique passive stabilizer, technically feasible, which could also provide damping, through a thermo-mechanical phase-lag component. No velocity sensor is necessary. The system, now patented [82], is said to be better than the conventional ones. Donlin and Randall's model [83] is essentially similar. Mer and Vigneron [84] thought of using the shape deformations of a balloon satellite of the type of Echo-II

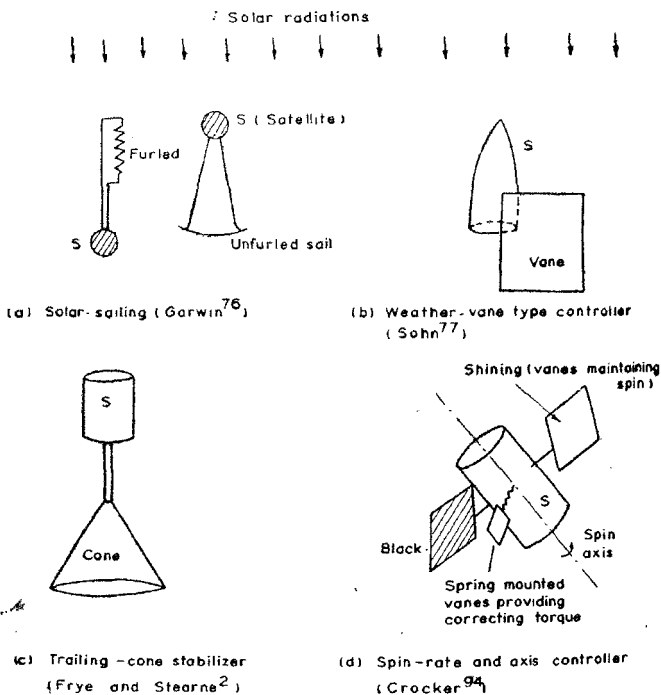


FIG. 4 Some concepts of solar controllers,

for spin-up. Galitskaya and Kislser [85] analysed a set of panels for three-axes stabilization and qualitatively established their optimum inclinations for the maximum utilization of the solar pressure. A number of theoretical studies [86-91] carried out at M.I.T. consider the dynamics, generally through linearization, of both spinning and non-spinning sun-orbiting spacecrafts stabilized by the solar radiation pressure. Aside from the conventional dampers a possibility of using the lag in thermal reradiation from the spacecraft's surface to produce non-conservative torques is also shown. The last of the above studies evolves a workable design of a three-axes controller for a small probe. Dzhumanoliev [92] analysed the small oscillations of a sun-pointing spacecraft with two coatings which result in a net control moment. A system of connected bodies to impart damping to a sun-pointing solar stabilizer is suggested by Merrick *et al.* [93]. To maintain spin and axis-orientation Crocker [94] proposed two sets of shining and black paddles. The spring-mounted set provides the desired change automatically. The preliminary analysis ignoring the gravity effects and assuming existence of a damper, shows good response. Figure 4 presents a few of the concepts developed above.

It is recognized that the gravity-oriented systems need damping to be effective. As discussed earlier they are also susceptible to even small disturbances. Mallach [95] using a phase-plane analysis and average torques considered the use of a solar pressure damper. Recently, Modi *et al.*, who had established the detrimental influence of radiations through the detailed nonlinear response and stability analyses, looked into the possibility of its utilization. Several semipassive systems involving a controlled change of area, moment arm or angle of incidence have been evolved [Fig. 5]. The first one was a simple velocity sensitive pitch damper. Its success both in circular and elliptic orbits under large external disturbances [96] lead to the development of a velocity- and position-sensitive controller which could stabilize the satellite about any desired in-plane orientation [97, 98]. The difficulties in making area changes through unfurlable material, etc., in space, were overcome by changing the moment arm [42]. This idea was later extended to coupled motion in circular orbit both in the ecliptic as well as other planes [99]. The response of several representative cases is analysed,

For dual-spin stabilized spacecrafts perhaps the only study on these lines is by Modi and Pande [100]. Nutation damping and attitude control is shown to be possible through a set of rotatable panels. The generalized analysis for vehicles in any orbit, with a reference to INTELSAT-IV and Anik-1 shows the transient time to be only 1/8 orbit with the steady-state errors

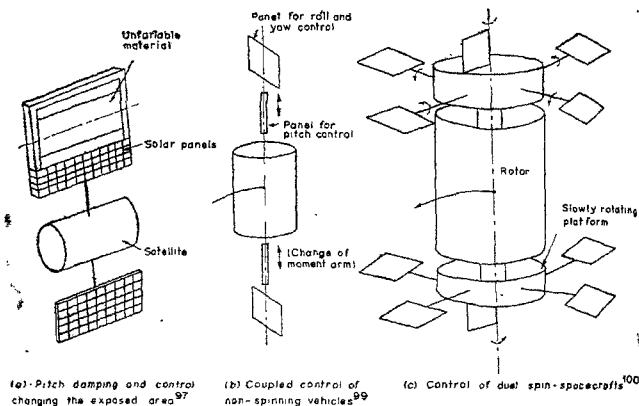


FIG. 5. Typical configuration of semi-passive solar controllers.

completely removed through a modified control function. The performance of this and other systems is further improved by numerical optimization [97, 101]. Though quite effective and promising these systems need attitude and angular-rate sensors and on-board computation. This may be difficult for smaller satellites.

It may be mentioned that a solar pressure controller has been successfully tested, despite some initial difficulties, on Mariner IV spacecraft [32, 81, 102]. This success may pave the way for implementation of the numerous ideas presented above.

2.2. Aerodynamic Controllers:

The aerodynamic forces have been shown to have a substantial influence on the attitude motion of the near-earth systems. Unlike solar radiation pressure, the literature on the advantageous applications of these forces is rather small. This may be due to the complex nature of the aerodynamic forces, rotation of atmosphere and strong dependence of density on height, season, sun-position and local change. [14]. The lack of a complete understanding of the atmospheric model also adds to the limited effort in their utilization. Yet, the few studies that have been made do show a feasibility of evolving a good aerodynamic attitude controller, which might also add

to the life of the near-earth system.. Such possibilities were indicated qualitatively in the early analyses by Debra and Stearns [103], Wall [104] and Schiello [105]. The success of aerodynamic pitch control, with the roll and yaw stabilized by gyros, was exhibited by COSMOS-149 [31]. In a model suggested by Hoffer [106] the gyros are replaced by a set of moving masses. Here the planar librations are reduced by a pair of drag flaps operating on an on-off control. This system may be difficult to implement because of the large changes in inertia. Ravindran and Hughes [107, 108] optimized, through linearization, a set of aerodynamic controller paddles for a satellite in a circular orbit. The system provides stabilization along the local horizontal. Using Liapunov's criteria the stability of such a controller is studied by Flanagan and Rangarajan [109].

For the general case of non-linear, coupled motion in an arbitrary orbit, Modi and Shrivastava [72] proposed several schemes of rotatable flaps which use both lift and drag. The velocity-sensitive controller results in an effective damping and stabilization of a gravity-gradient system. Later, an aerodynamic controller which stabilizes the system about any arbitrary spatial orientation is evolved using a velocity- and position-sensitive system. Through parametric optimization [110] the transient time is reduced to less than 1/3 orbit and the steady-state amplitude [in elliptic orbits] to smaller than 1°.

Noting the success of solar controllers and realizing the limitations of aerodynamic systems in elliptic orbits where the usable force may be significant only near the perigee, a simple hybrid control which employs atmospheric forces at lower altitudes and solar pressure at upper levels is developed [111]. The analysis is general enough to be applicable for aerodynamic, solar or a hybrid system. A modified control relation removes the steady-state errors even for large eccentricities. A similar system for a spinning satellite is also analysed [112, 113].

A schematic presentation of the above controllers is shown in Fig. 6.

The studies clearly indicate a need for further investigation of the possibilities of use of these semi-passive controllers because of their promise as light weight, low power, long life systems.

CONCLUDING REMARKS

The review of the literature suggests a considerable interest both in the influence of the environmental forces and their advantageous utilization for attitude control of satellites. A majority of the studies are, however

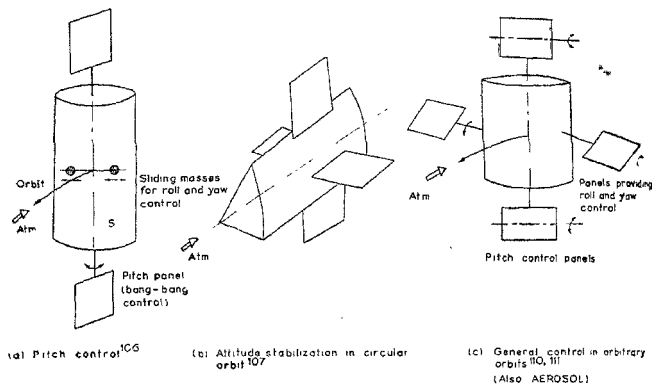


FIG. 6. Use of atmosphere for attitude stabilization and control.

theoretical or qualitative in nature. For the successful implementation of the promising ideas considerable work on hardware development and testing needs to be done.

It may be mentioned that this report does not cover the related important areas of orbital perturbations and corrections, radiation heating, and flexibility. In designing a system these too must be carefully considered.

An attempt has been made here to highlight the most important aspects of the studies available in the recent open literature. It is recognized that many important reports, particularly those in languages other than English may have been left out. It is felt, however, that this report may draw the attention of designers to the importance of the effects and may also stimulate thinking about some of the new concepts.

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