

of the radiation, that is, the relative magnitudes of the direct and diffuse components or of the ratio D_h/E_h . If it is assumed that for a given locality and season of the year the quality of the radiation is predominantly a function of the solar altitude a , then this latter factor is a convenient parameter on which to base a system of characteristic lines for a given value of θ_H . Typical characteristic lines are shown for $\theta_H = 90^\circ$ and $\theta_H = 40^\circ$ in Fig. 2. An analysis of the theory shows that if the diffuse component of radiation were uniformly distributed the slope M of the characteristic lines would be $(1 - D_h/E_h)$, but as the distribution is not uniform $(1 - D_h/E_h)$ is slightly smaller than M . The total radiation on a steeply inclined surface also includes some radiation from the heated surface of the ground, but this ceases to be appreciable when the angle of inclination θ_H is below 60° .

It is intended to publish charts or tabulated values of the constants C and M for various solar altitudes throughout the yearly cycle in the London area, which will enable the values of E_i/E_h to be calculated for any surface. The primary object of this communication, however, is to promote interest in the research in other countries and to stimulate the determination of similar data which will extend knowledge of the availability of solar radiation.

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Electron Temperature in a Laser-heated Plasma

WHEN a high-power laser beam is focused on to a material target, a small inertially confined plasma is generated near the target surface. Theoretical work by Basov and Krokhin^{1,2}, verified by Dawson³, suggests that very hot plasmas ($\sim 10^7$ °K) can ultimately be produced in this way. A spectroscopic investigation of plasmas generated by a small Q-switched ruby laser has already been reported⁴. Further experiments have now been carried out using a larger Q-switched neodymium glass laser, and an estimate of the electron temperature attained in a carbon plasma is reported in this communication. Other results will be described in detail elsewhere.

The peak power density at the focus of the neodymium glass laser beam was about 10^{11} W cm⁻². The output wave-length was 10,600 Å. Each laser pulse lasted about 1 μsec and consisted of three spikes. The energy in a pulse was about 1 joule. The focused beam broke down atmospheric air, so the target was placed in a vacuum. The plasma formed by the laser beam was imaged on the entrance slit of a Hilger medium spectrograph. When a graphite target was used, a single laser pulse was sufficient to produce a spectrogram on which lines of the CI, II, III, IV and V spectra were identified.

Kaufman and Williams^{5,6} have pointed out that under certain conditions the specific intensities of lines from atoms of the HeI isoelectronic sequence give a useful indication of the plasma electron temperature even when thermodynamic equilibrium is absent and the electron and ion densities are only very approximately known. A number of assumptions must be made whose validity will be discussed in a later paper. In particular, a Maxwellian distribution of electron velocities is assumed. The expression given by these authors for the specific intensity, I , divided by the product of the number densities of electrons, n_e , and of quadruply ionized carbon atoms in the ground state, n_4 , has been evaluated for the observed CV transition $2s^2S_{1-2}p^3P_2$ at 2270.91 Å and is plotted as a function of the electron temperature, T_e , in Fig. 1.

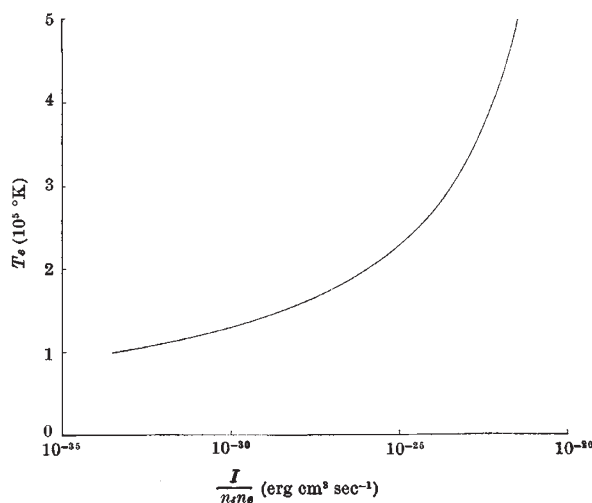


Fig. 1

Taking into account geometrical considerations, the observed blackening of the spectrographic plate indicates that the specific intensity of this CV line was at least 3×10^{11} erg cm⁻³ sec⁻¹. An extreme upper limit of 10^{21} cm⁻³ for n_4 is obtained from the volume and density of the graphite removed from the target by the laser pulse, and the volume of luminous plasma emitting CV light. The corresponding maximum electron density is 6×10^{21} cm⁻³, because the carbon plasma was very pure. Thus $I/(n_e n_4) \geq 5 \times 10^{-22}$ erg cm³ sec⁻¹. From Fig. 1, the lower limit for the peak electron temperature is therefore 1.1×10^6 °K. As may be seen from Fig. 1, at this temperature large errors in density or intensity estimates produce only small errors in the derived temperature. Nevertheless, as the true densities may well be orders of magnitude lower than the values used in the calculation the true electron temperature is probably somewhat higher.

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Saturation Characteristics of a Spherical Ionization Chamber and a Determination of Boag's Constant for Air

Thomson and Rutherford's¹ initial investigations of the saturation characteristics of a parallel plate ionization chamber were followed by theoretical treatments by Thomson², Mie³, Seeliger⁴ and others. More recently, Boag and Wilson⁵ have shown by dimensional analysis that the collection efficiency, f (that is, the ratio of the measured current, i , to the ideal saturation current, i_s), is a function of $d^2 q^2 V^{-1}$ where d is the distance between the plates, q is the rate of ionization and V is the applied voltage. Through the introduction of certain assumptions, they derived the following equation for the generalized saturation characteristic:

$$i_s = i \left[1 + \frac{i}{i_s} \cdot \frac{1}{6} \frac{\alpha/e}{k_1 k_2} \cdot \frac{d^4 q}{V^2} \right] \quad (1)$$