

1 Environmental controls on marine ecosystem recovery following mass extinctions, with an  
2 example from the Early Triassic

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## 23 **ABSTRACT**

24 The recovery of marine ecosystems following a mass extinction event involves an  
25 extended interval of increasing biotic diversity and ecosystem complexity. The pace of  
26 recovery may be controlled by intrinsic ecosystem or extrinsic environmental factors. Here,  
27 we present an analysis of changes in marine conditions following the end-Permian mass  
28 extinction with the objective of evaluating the role of environmental factors in the protracted

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29 (~5-Myr-long) recovery of marine ecosystems during the Early Triassic. Specifically, our  
30 study examines changes in weathering, productivity, and redox proxies in three sections in  
31 South China (Chaohu, Daxiakou, and Zuodeng) and one in northern India (Mud). Our results  
32 reveal: 1) recurrent environmental perturbations during the Early Triassic; 2) a general pattern  
33 of high terrestrial weathering rates and more intensely reducing marine redox conditions  
34 during the early Griesbachian, late Griesbachian, mid-Smithian, and (more weakly) the  
35 mid-Spathian; 3) increases in marine productivity during the aforementioned intervals except  
36 for the early Griesbachian; and 4) stronger and more temporally discrete intervals of  
37 environmental change in deepwater sections (Chaohu and Daxiakou) relative to shallow and  
38 intermediate sections (Zuodeng and Mud). Our analysis reveals a close relationship between  
39 episodes of marine environmental deterioration and a slowing or reversal of ecosystem  
40 recovery based on metrics of biodiversity, within-community (alpha) diversity, infaunal  
41 burrowing, and ecosystem tiering. We infer that the pattern and pace of marine ecosystem  
42 recovery was strongly modulated by recurrent environmental perturbations during the Early  
43 Triassic. These perturbations were associated with elevated weathering and productivity  
44 fluxes, implying that nutrient and energy flows were key influences on recovery. More regular  
45 secular variation in deepwater relative to shallow-water environmental conditions implies that  
46 perturbations originated at depth (i.e., within the oceanic thermocline) and influenced the  
47 ocean-surface layer irregularly. Finally, we compared patterns of environmental disturbance  
48 and ecosystem recovery following the other four “Big Five” Phanerozoic mass extinctions to  
49 evaluate whether commonalities exist. In general, the pace of ecosystem recovery depends on  
50 the degree of stability of the post-crisis marine environment.

51

52 *Keywords:* productivity; redox; anoxia; weathering; South China; India

53

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84 **1. Introduction**

85

86           Each major mass extinction event in the geologic record has been followed by an interval  
87 of restructuring of marine ecosystems, reflected in changes in clade dominance, ecological  
88 niche partitioning, and community organization (e.g., [Erwin, 1998](#)). Increased productivity  
89 among primary producers and consumers can generate ecological niches higher in the marine  
90 trophic system ([Kirchner and Weil, 2000](#)), allowing a progressive rebuilding of a stable,  
91 complex ecosystem structure ([Chen and Benton, 2012](#)). Although lacking a specific  
92 quantitative definition, “ecosystem recovery” is generally regarded as the reappearance of  
93 marine communities with a high biotic diversity and an integrated and complex structure that  
94 is stable at multimillion-year timescales ([Harries and Kauffman, 1990](#)). The progress of  
95 post-extinction recovery commonly has been evaluated using metrics related to overall  
96 biodiversity and/or species origination rates (e.g., [Jacobsen et al., 2011](#); [Payne et al., 2011](#)).  
97 However, “ecosystem recovery” is not simply a return to pre-extinction levels of biodiversity  
98 but, rather, the expansion and re-integration of entire marine ecosystems or communities  
99 ([Erwin et al., 2008](#); [Chen and Benton, 2012](#)) as reflected by metrics such as alpha diversity  
100 (i.e., within-community species richness; [Bambach, 1977](#); [Clapham et al., 2006](#)) and  
101 ecological tiering ([Twitchett, 1999](#); [Fraiser, 2011](#)).

102           In the case of the Permian-Triassic (P-Tr) boundary mass extinction, an initial, aborted  
103 recovery occurred soon after the end-Permian crisis, during the Induan stage of the Early  
104 Triassic ([Baud et al., 2008](#); [Brayard et al., 2009](#); [Stanley, 2009](#)), and a more sustained  
105 recovery took place during the late Olenekian stage (Spathian substage) ([Chen et al., 2011](#);  
106 [Song et al., 2011](#); [Payne et al., 2011](#)), but full ecosystem recovery probably did not occur until  
107 the Middle Triassic ([Erwin and Pan, 1996](#); [Bottjer et al., 2008](#); [Chen and Benton, 2012](#)). The  
108 recovery of marine invertebrate ecosystems following the end-Permian crisis was apparently  
109 the most protracted of any major mass extinction ([Bottjer et al., 2008](#)), i.e., the “Big Five”  
110 Phanerozoic mass extinctions of [Sepkoski \(1984, 1986\)](#). An important unresolved issue is  
111 what controlled the long duration of the post-extinction recovery interval during the Early  
112 Triassic. At least three hypotheses have been advanced, linking the protracted recovery to: (1)

113 the intensity of the mass extinction (Sepkoski, 1984; Solé et al., 2002), (2) the persistence of  
114 harsh environmental conditions (Hallam, 1991; Isozaki, 1997; Payne et al., 2004; Erwin,  
115 2007), and (3) episodic occurrence of strong environmental disturbances during the recovery  
116 interval (Algeo et al., 2007, 2008; Orchard, 2007; Retallack et al., 2011) (Fig. 1).

117 Examination of long-term records of Early Triassic marine environmental conditions has  
118 the potential to provide information relevant to these hypotheses. In this study, we (1) review  
119 existing literature on the recovery of marine ecosystems following the end-Permian mass  
120 extinction, (2) analyze changes in marine productivity and redox conditions at four locales in  
121 China and India from the latest Permian through the Spathian substage of the Early Triassic,  
122 (3) evaluate the importance of marine environmental changes during the Early Triassic as  
123 controls on the marine ecosystem recovery, and (4) compare the Early Triassic marine  
124 ecosystem recovery with those following other Phanerozoic mass extinctions. Our  
125 comparative analysis of recoveries following each of the ‘Big Five’ Phanerozoic mass  
126 extinctions is intended to identify general features or patterns of marine ecosystem recovery  
127 and their relationships to contemporaneous environmental changes.

128

## 129 **2. Background**

130

### 131 *2.1. The end-Permian biotic crisis*

132

133 The end-Permian mass extinction was the most severe biocrisis of the Phanerozoic (Fig.  
134 2; Erwin et al., 2002; Irmis and Whiteside, 2011). It killed ~80-96% of marine invertebrate  
135 species and ~70% of terrestrial vertebrate species (McKinney, 1995; Benton and Twitchett,  
136 2003). There appear to have been two pulses of marine extinction (Yin et al., 2012; Song-HJ  
137 et al., 2013) and environmental disturbance (Xie et al., 2005, 2007), rather than a single event  
138 during this biocrisis (Rampino and Adler, 1998; Jin et al., 2000; Shen-SZ et al., 2011). As an  
139 example, foram species in South China exhibit a ~57% extinction rate during the latest  
140 Permian pulse and a ~31% extinction rate during the earliest Triassic pulse (Song-HJ et al.,

141 [2013](#)). According to high-precision U-Pb dating in South China sections, the interval between  
142 these extinction pulses was  $60\pm 48$  kyr ([Burgess et al., 2014](#)). The end-Permian mass  
143 extinction coincided with eruption of the Siberian Traps Large Igneous Province ([Campbell et](#)  
144 [al., 1992](#); [Renne et al., 1995](#); [Reichow et al., 2009](#); [Sobolev et al., 2011](#)) as well as with major  
145 environmental changes including global sea-level rise ([Hallam and Wignall, 1999](#)), ocean  
146 anoxia ([Wignall and Twitchett, 1996](#); [Isozaki, 1997](#)), global warming ([Joachimski et al., 2012](#);  
147 [Sun et al., 2012](#); [Romano et al., 2013](#)), and, possibly, marine acidification ([Payne et al., 2010](#);  
148 [Hinojosa et al., 2012](#); [Kershaw et al., 2012](#)).

149

## 150 *2.2. The Early Triassic marine ecosystem recovery*

151

152 The recovery of marine ecosystems during the Early Triassic was a multi-step process.  
153 There were several phases of incomplete or aborted recovery during the Induan, and recovery  
154 from the P-Tr boundary mass extinction is generally regarded as not having been completed  
155 until the Middle Triassic, ~5 Myr after the end-Permian crisis ([Mundil et al., 2004](#);  
156 [Ovtcharova et al., 2006](#); [Lehrmann et al., 2006](#); [Shen-SZ et al., 2011](#)). Both benthic and  
157 planktonic cyanobacteria bloomed immediately after the end-Permian mass extinction ([Fig. 2](#);  
158 [Lehrmann, 1999](#); [Wang et al., 2005](#); [Xie et al., 2005](#); [Luo et al., 2011](#)). Cyanobacterial  
159 microbialites reappeared episodically in different regions throughout the Early Triassic but  
160 they largely disappeared by the early Middle Triassic ([Baud et al., 2007](#); [Xie et al., 2010](#)). An  
161 Early Triassic “chert gap” ([Beauchamp and Baud, 2002](#)) was caused by the loss of biosilica  
162 deposits from radiolarians and siliceous sponges, although occurrences of thin chert beds in  
163 the late Griesbachian and Dienerian ([Kakuwa, 1996](#); [Takemura et al., 2007](#); [Sano et al., 2010](#))  
164 document a temporary local early recovery of siliceous faunas.

165 Some secondary consumers such as conodonts and ammonoids rebounded rapidly from  
166 the end-Permian mass extinction ([Orchard, 2007](#); [Brayard et al., 2009](#); [Stanley, 2009](#)). Their  
167 rapid recovery may have been assisted by a microphagous habit ([Fischer and Bottjer, 1995](#)),  
168 allowing them to benefit directly from increased biomass among primary producers. These

169 clades subsequently declined during biocrises at the end of Griesbachian, Smithian, and  
170 Spathian substages of the Early Triassic, although they tended to rediversify rapidly during  
171 the intervening intervals (Fig. 2; Brayard et al., 2009; Stanley, 2009). However, conodonts  
172 display a strong Lilliput effect during the Smithian/Spathian boundary crisis (Chen et al.,  
173 2013). Compared to conodonts and ammonoids, recovery rates for benthic primary consumers  
174 such as foraminifers, gastropods, bivalves, brachiopods and ostracods were more gradual (Fig.  
175 2; Payne et al., 2011). Among foraminifers, a sustained diversity increase began in the early  
176 Smithian (early Olenekian) (Song et al., 2011) and accelerated during the Anisian (early  
177 Middle Triassic) (Payne et al., 2011). Similar recovery patterns are observed also among  
178 brachiopods (Chen et al., 2005) and ostracods (Crasquin-Soleau et al., 2007). The sizes of  
179 gastropod and bivalve shells were reduced across the P-Tr boundary and during the  
180 Griesbachian but returned to pre-extinction dimensions by the Anisian (Fig. 2; Fraiser and  
181 Bottjer, 2004; Payne, 2005; Twitchett, 2007). However, the high diversity, low dominance,  
182 and ecological complexity of mollusc fauna during the late Griesbachian and early Dienerian  
183 at Shanggan, South China (Hautmann et al., 2011) and on the Wasit Block in Oman (Krystyn  
184 et al., 2003; Twitchett et al., 2004) may represent an early recovery phase of these faunas.

185       The meso-consumer trace-makers and reef-builders can shed light on the recovery of  
186 benthic marine ecosystems. Generally, trace-makers decreased during the end-Permian  
187 biocrisis and recovered slowly in the Early Triassic (Fig. 2; Pruss and Bottjer, 2004; Chen et  
188 al., 2011). Locally, trace-fossil diversity shows occasional peaks during the Griesbachian to  
189 Smithian (Twitchett and Wignall, 1996; Twitchett, 1999; Zonneveld et al., 2010; Chen et al.,  
190 2011). However, small trace-fossil burrow size, low tiering levels, and low ichnofabric indices  
191 (bioturbation) generally persisted until the end of the Smithian substage, and the early  
192 Spathian is marked by a strong increase in trace-fossil diversity and complexity (Pruss and  
193 Bottjer, 2004; Chen et al., 2011). Nonetheless, Spathian ichnofaunas are less diverse than  
194 those of the Middle Triassic (Knaust, 2007). This pattern may suggest a stepwise recovery of  
195 trace-makers during Early to Middle Triassic (Twitchett and Barras, 2004). Furthermore, the  
196 recovery of trace-makers may have been diachronous, with a more rapid increase in

197 ichnodiversity at high northern paleolatitudes than in the equatorial region (Twitchett and  
198 Barras, 2004; Pruss and Bottjer, 2004). With regard to reef-builders, a new metazoan reef  
199 ecosystem formed by various sponges and serpulid worms associated with microbial  
200 carbonates and eukaryotic organisms developed in the early Smithian, latest Smithian, and  
201 early to middle Spathian on the eastern Panthalassic margin, in Utah and Nevada (Fig. 2;  
202 Brayard et al., 2011). These equatorial sponge-microbe reefs are found as early as 1.5 Myr  
203 after the P-Tr boundary and represent a temporary recovery at least regionally (Brayard et al.,  
204 2011, Chen and Benton, 2012). However, the “reef gap”, as represented by the absence of  
205 heavily calcified corals, persisted through the Early Triassic (Payne et al., 2006).

206 As for the top trophic level in the marine ecosystem, predatory fish and reptiles  
207 displayed different recovery trajectories. Fishes were rare in the Griesbachian-to-Smithian  
208 equatorial ocean (Fig. 2; Fraiser et al., 2005; Tong et al., 2006; Zhao and Lu, 2007; Sun et al.,  
209 2012) but more common in the middle to late Spathian (Goto, 1994; Wang et al., 2001;  
210 Benton et al., 2013). High-latitude regions had a more abundant and diverse fish fauna in the  
211 Early Triassic than the equatorial ocean (Scheaffer et al., 1976; Stemmerik et al., 2001; Mutter  
212 and Neuman, 2006; Romano and Brinkmann, 2010; Benton et al., 2013). Globally, fish  
213 diversity recovered by the Middle Triassic (Jin, 2006; Zhang et al., 2010; Hu et al., 2011).  
214 Marine reptiles first reappeared in the Smithian in high-latitude regions (Cox and Smith, 1973;  
215 Callaway and Brinkman, 1989) but later, in the Spathian, in equatorial regions (Li et al., 2002;  
216 Zhao et al., 2008). A high level of diversity among marine reptiles was achieved by the  
217 Middle to Late Triassic (Zhang et al., 2009).

218 To summarize, animals that were low in the marine trophic system tended to recover  
219 faster than those at higher trophic levels (Fig. 2; cf. Chen and Benton, 2012). Pelagic and  
220 nektonic faunas recovered faster than benthos as shown by rapid increases to multiple  
221 biodiversity peaks for ammonoids and conodonts during the Early Triassic, versus a slow  
222 return to pre-crisis diversity levels by the Middle Triassic for most bottom-dwellers. In  
223 Olenekian time, offshore benthos like calcareous algae and *Tubiphytes* recovered faster than  
224 those in nearshore environments in South China (Song et al., 2011). High-latitude biotas



225 recovered faster than equatorial marine biota (Pruss and Bottjer, 2004). These differentiated  
226 responses may suggest that the pattern and intensity of environmental changes during the  
227 Early Triassic had an important influence on the pathways and tempo of marine ecosystem  
228 recovery.

229

### 230 *2.3. Environmental change during the Early Triassic recovery*

231

232 During the recovery interval following the end-Permian mass extinction, major changes  
233 in the environment related to volcanism, sea level, and paleoceanographic conditions took  
234 place. The eruption of the Siberian Traps large igneous province (LIP), which had begun at  
235 ~252 Ma close to the P-Tr boundary, continued strongly for ~1.5 Myr and more weakly for  
236 several million years longer (Fig. 3). The eruption history of this LIP is delineated by U-Pb  
237 ages for gabbroic intrusive rocks of 252±4 Ma (Kuzmichev and Pease, 2007) and silicic tuff  
238 ages of 251.7±0.4 (Kamo et al., 2003), an Ar-Ar age of 250.3±1.1 Ma for the final stages of  
239 extrusive volcanism (Reichow et al., 2009), and younger Ar-Ar ages of 242.2±0.6 Ma for a  
240 basalt (Reichow et al., 2009). This range of dates documents activity of the Siberian Traps LIP  
241 from 252 Ma to 242 Ma with a main eruptive phase at ~252 to 250 Ma (Reichow et al., 2009).  
242 Altogether, the Siberian Traps degassed ~6300 to 7800 Gt sulfur, ~3400 to 8700 Gt chlorine,  
243 and ~7100 to 13600 Gt fluorine (Black et al., 2012). The high volatile contents increased the  
244 likelihood that volatiles reached the stratosphere and, thus, caused a drastic deterioration of  
245 global environments through direct toxicity and acid rainfall (Devine et al., 1984), ozone  
246 depletion (Johnston, 1980), and rapid climatic changes that may have included both global  
247 cooling (Sigurdsson et al., 1992; Wignall, 2001; Timmreck et al., 2010) and global warming  
248 (Ganino and Arndt, 2009). This interval coincided with a long-term eustatic rise from the Late  
249 Permian until the middle Late Triassic, with the most rapid rise during the Early Triassic (Fig.  
250 3; Haq et al., 1987; Haq and Schutter, 2008).

251 Major changes in tropical sea-surface temperatures accompanied the P-Tr boundary  
252 crisis. Temperatures increased gradually from ~60 kyr prior to the mass extinction event and

253 then spiked rapidly at the time of this event (Joachimski et al., 2012; Burgess et al., 2014).  
254 During the Early Triassic, temperatures reached a maximum in the mid- to late Griesbachian  
255 (~36-40°C), cooled slightly during the latest Griesbachian to the early Smithian, and then  
256 reached a second peak of extreme warmth in the late Smithian (Fig. 3; Sun et al., 2012,  
257 Romano et al., 2013). A pronounced retreat from peak temperatures occurred in the early  
258 Spathian, an event resulting in a major turnover and geographic displacement of marine  
259 invertebrate faunas (Galfetti et al., 2007a,b; Stanley, 2009). A weak warming episode in the  
260 mid-late Spathian was followed by a second large cooling step around the Early-Middle  
261 Triassic boundary, yielding distinctly more moderate temperatures during the Anisian  
262 although still warmer than in the pre-extinction Late Permian (Sun et al., 2012; Romano et al.,  
263 2013).

264 Ocean redox conditions exhibit pronounced geographic and secular variation during the  
265 latest Permian and Early Triassic. More reducing conditions developed widely at mid-water  
266 depths (i.e., with the oceanic thermocline) during the pre-extinction late Changhsingian  
267 (Algeo et al., 2012; Shen et al., 2013; Feng and Algeo, 2014). The end-Permian crisis was  
268 marked by a transient expansion of anoxia into shallow-marine settings, especially in the  
269 Tethyan Ocean (Fig. 3; Horacek et al., 2007; Grice et al., 2005; Algeo et al., 2007, 2008; Bond  
270 and Wignall, 2010; Brenneka et al., 2011; Shen-Y et al., 2011), although some places (e.g.,  
271 Oman, Iran) remained oxic (Krystyn et al., 2003; Richoz et al., 2010). Thereafter, the Early  
272 Triassic is characterized by a complex pattern of redox variation (Song et al., 2012; Grasby et  
273 al., 2013). The intensity of anoxia appears to have declined during the Spathian, and episodes  
274 of marine anoxia seem to have terminated around the Early-Middle Triassic boundary  
275 (Hermann et al., 2011; Song et al., 2012).

276 Marine productivity can vary greatly during major biocrises (Kump and Arthur, 1999).  
277 Several factors during the P-Tr boundary crisis might have led to higher productivity: 1)  
278 phosphate liberated from sediments under anoxic conditions can stimulate productivity (Ingall  
279 and van Cappellen, 1990), and 2) intensified subaerial weathering can increase the flux of  
280 river-borne P to the oceans (Algeo and Twitchett, 2010; Algeo et al., 2011b). Variations in

281 marine productivity can be reconstructed using carbon isotopes or elemental data (Kump and  
282 Arthur, 1999; Algeo et al., 2013; Schoepfer et al., 2014). The ‘biological pump’ removes  
283  $^{12}\text{C}$ -enriched carbon from the ocean-surface layer and transfers it to the ocean thermocline,  
284 producing a vertical gradient in the  $\delta^{13}\text{C}$  of dissolved inorganic carbon ( $\Delta^{13}\text{C}_{\text{DIC}}$ ). Changes in  
285  $\Delta^{13}\text{C}_{\text{DIC}}$  can thus provide information about the intensity of the organic carbon sinking flux  
286 and, indirectly, primary productivity (Hilting et al., 2008). A large vertical  $\delta^{13}\text{C}_{\text{DIC}}$  gradient in  
287 the Nanpanjiang Basin of South China was interpreted as evidence of elevated marine  
288 productivity during the Early Triassic (Fig. 3; Meyer et al., 2011), although this gradient has  
289 also been attributed to intensified water-column stratification (Song-HY et al., 2013; Luo et  
290 al., 2014). However, an analysis of marine productivity changes based on organic carbon  
291 burial fluxes suggested a productivity crash in Early Triassic seas of the South China craton  
292 (Algeo et al., 2013). The large carbon-isotope excursions of the Early Triassic (Payne et al.,  
293 2004; Tong et al., 2007; Clarkson et al., 2013) were hypothesized to have been due to marine  
294 productivity fluctuations (Algeo et al., 2011b), an inference supported by patterns of  
295  $\delta^{13}\text{C}$ - $\delta^{34}\text{S}$  covariation (Song et al., 2014). The ultimate control on these fluctuations appears to  
296 have been temperature, with warm intervals associated with reduced productivity (Song et al.,  
297 2014).

298 Seawater pH values may have fluctuated during the P-Tr boundary crisis, as shown by  
299 analysis of calcium isotopes (Payne et al., 2010; Hinojosa et al., 2012). Calcium isotopic  
300 fractionation caused by the precipitation of carbonate minerals results in  $^{40}\text{Ca}$ -rich marine  
301 sediments and  $^{44}\text{Ca}$ -rich in seawater (Skulan et al., 1997; De La Rocha and DePaolo, 2000;  
302 Fantle and DePaolo, 2005; Tang et al., 2008). Abrupt negative excursions of  $\delta^{44/40}\text{Ca}$  in both  
303 bulk carbonate and conodont apatite, representing a shift in seawater  $\delta^{44/40}\text{Ca}$ , occurred  
304 synchronously with the end-Permian biocrisis (Fig. 3; Payne et al., 2010; Hinojosa et al.,  
305 2012). The underlying cause of this change may have been eruption of the Siberian Traps,  
306 which injected a large amount of  $\text{CO}_2$  into the atmosphere-ocean system, causing seawater  
307 acidification and increased riverine  $^{40}\text{Ca}$ -rich calcium input owing to accelerated terrestrial

308 weathering of carbonates (Payne et al., 2010; Blätter et al., 2011). Ocean acidification during  
309 Permian-Triassic transition may lead to the preferential extinction of heavily calcified marine  
310 organisms (Knoll et al., 2007; Kiessling and Simpson, 2011; Clapham and Payne, 2011) and  
311 could explain the abrupt transition on carbonate platforms from skeletal to microbial and  
312 abiotic carbonate factories described by Kershaw et al. (2011).

313 To summarize, eruption of the Siberian Traps during the Late Permian to Early Triassic  
314 resulted in a major perturbation of the atmosphere-ocean system. Environmental changes  
315 linked to early phases of the eruption appear began slowly during an interval of at least ~60  
316 kyr preceding the main mass extinction, but accelerated sharply at the end of the Permian.  
317 Major environmental effects related to continuing eruption of Siberian Traps flood basalts  
318 persisted for ~1.5 to 2.0 million years during the Early Triassic, with some effects continuing  
319 until the Early-Middle Triassic boundary, nearly 5 million years after the end of the Permian.  
320 The main phase of the eruption, coinciding with the Induan Stage of the Early Triassic,  
321 coincided with highly disturbed marine ecosystems, sea-level rise, seawater acidification, and  
322 widespread oceanic anoxia. These relationships show that environmental instability coincided  
323 with, and probably caused or contributed to, the delayed recovery of marine ecosystems  
324 during the Early Triassic.

325

### 326 **3. Study sections**

327

328 Three of the sections chosen for this study are from the South China craton, which was  
329 located in the eastern Paleotethys Ocean during the Permian-Triassic transition. The Chaohu  
330 section was deposited in a deep ramp setting on the northeastern (paleo-northwestern) margin  
331 of this craton, Daxiakou on the mid-ramp of the same margin, and Zuodeng on a shallow  
332 carbonate platform within the Nanpanjiang Basin on the southwestern (paleo-southeastern)  
333 margin of this craton (Fig. 4A). These sections were widely separated, with a distance of ~650  
334 km between Chaohu and Daxiakou, and a distance of ~950 km between the latter and  
335 Zuodeng. The fourth study section is Mud, from the Spiti Valley of northern India, which was

336 located in the south-central Neotethys Ocean during the Permian-Triassic transition (Fig. 4B).  
337 We collected a total of 794 samples from 167 m of section at Chaohu, 302 samples from 71 m  
338 of section at Daxiakou, 351 samples from 109 m of section at Zuodeng, and 135 samples from  
339 26.5 m of section at Mud. Average sample spacing thus ranges from 20 to 31 cm for the four  
340 study sections, which equates to an average temporal interval of ~4 to 10 kyr between  
341 samples (see [Supplementary Table 1](#) for the geologic timescale used in this study, and the  
342 [Supplementary Information](#) for age-depth models of the study sections).

343

### 344 *3.1. Chaohu, Anhui Province, China*

345

346 The Chaohu section is located in proximity to Chaohu city in Anhui Province (Fig. 4A). It  
347 is a composite section comprising sections at West Majiashan, West Pingdingshan, and South  
348 Majiashan, all of which are located within a ~1-km<sup>2</sup> area (Tong et al., 2003). These sections  
349 contain, respectively, the narrow P-Tr boundary interval, the Griesbachian to Smithian, and  
350 the Spathian (Fig. 5), according to conodont biostratigraphic data (Zhao et al., 2007). The top  
351 of the South Majiashan section coincides approximately with the Spathian-Anisian  
352 (Early-Middle Triassic) boundary (Zhao et al., 2007). During the Early Triassic, the Chaohu  
353 area was on the deep lower margin of a ramp about 300 km to the north (paleo-west) of the  
354 Cathaysia Oldland (Fig. 4A; Tong et al., 2003). Estimated depositional water depths in the  
355 Chaohu area were ~300-500 m (Song-HY et al., 2013). However, relative sea-level elevations  
356 began to decrease during the Spathian (Tong et al., 2001, 2007b; Chen et al., 2011) as a  
357 consequence of a collision between the North China and South China blocks that culminated  
358 in the late Middle Triassic (Li, 2001).

359 This section has been subject to detailed analysis of conodont and ammonoid  
360 biostratigraphy (Zhao et al., 2007), sequence stratigraphy (Tong, 1997; Li et al., 2007), carbon  
361 isotopes (Tong et al., 2007a), and paleomagnetic polarity (Tong et al., 2003), permitting  
362 development of a high-resolution geochronological framework for this study. The West  
363 Pingdingshan section is a candidate for the Global Stratotype Section and Point (GSSP) of the

364 Induan-Olenekian boundary (Tong et al., 2003). Conodonts are found in abundance in the  
365 upper Griesbachian through Dienerian-Smithian boundary, the middle Smithian, and lower  
366 Spathian but are rarer in other stratigraphic intervals (Tong et al., 2003; Zhao et al., 2007).  
367 Foraminifers are found in the Induan stage (Song et al., 2011), ammonoids are particularly  
368 abundant around the Smithian-Spathian boundary, and some marine vertebrate fossils are  
369 found in the Olenekian (Tong et al., 2003).

370 The carbonate fraction of the sediment shows an increase upsection at Chaohu, from  
371 ~30% around the P-Tr boundary to ~40-70% in the Griesbachian and Dienerian, ~75% in the  
372 Smithian (except for a local decline to ~20% in the mid-Smithian), and ~87% in the Spathian  
373 (Fig. 5, Supplementary Table 2). Chert, which is probably mainly of biogenic origin,  
374 decreases upsection, from ~28% around the P-Tr boundary to ~5% in the Spathian.  
375 Clay-mineral content shows a similar upsection decrease, from ~50% around the P-Tr  
376 boundary to ~8% in the Spathian. These mineralogic changes are reflected in an upsection  
377 shift in lithology from cherty mudrock with minor limestone interbeds around the P-Tr  
378 boundary to thin-bedded marls with mudrock interbeds in the Griesbachian and Dienerian,  
379 dominant mudrock with marlstone interbeds in the Smithian, and thick-bedded limestone with  
380 marlstone interbeds in the Spathian (Fig. 5; Tong et al., 2003, 2007a; Guo et al., 2008).

381

### 382 3.2. Daxiakou, Hubei Province, China

383

384 The Daxiakou section is located in Xingshan county, Yichang city, in the Yangtze Gorge  
385 area of Hubei Province (Fig. 4A). During the Early Triassic, it was located in a deep-ramp  
386 setting on the northern margin of the South China Block (Tong and Yin, 2002; Zhao et al.,  
387 2005), ~850 km from the Kangdian Oldland (Fig. 4A). Conodont biostratigraphy shows that  
388 the section spans the early Changhsingian through mid-Smithian interval (Zhao et al., 2005).  
389 Fossils of ammonoids, conodonts, and bivalves, among others are found in particular  
390 abundance in upper Dienerian to lowermost Smithian strata (Li et al., 2009), implying  
391 relatively high primary productivity at that time (Tong, 1997). Estimated depositional water

392 depths in the Daxiakou area were ~200-300 m (Song-HY et al., 2013).

393 The carbonate component is high (>80%) throughout the section except for the P-Tr  
394 boundary interval and in Dienerian to lower Smithian strata (Fig. 5, Supplementary Table 3).  
395 In the P-Tr transition, average carbonate, chert, and clay-mineral contents are ~30%, ~20%,  
396 and 50%, respectively, and strata consist of thin-bedded, dark-gray to black cherty shales (cf.  
397 Wu et al., 2012). In the Dienerian to lower Smithian, average carbonate, chert, and  
398 clay-mineral contents are ~60%, ~10%, and ~30%, respectively, and strata consist of  
399 marlstones with mudrock intercalations.

400

### 401 3.3. Zuodeng, Guangxi Province, China

402

403 The Zuodeng section is located in Zuodeng county, Tiandong city, in Guangxi Province.  
404 During the Early Triassic, this section was located on a carbonate platform (the Debao  
405 Platform) within the Nanpanjiang Basin (Fig. 4A), a deep-marine embayment on the  
406 southwestern (paleo-southeastern) margin of the South China Block that existed from the Late  
407 Paleozoic to the Late Triassic (Enos et al., 1997). The Debao Platform was one of many  
408 isolated, shallow carbonate platforms within this basin, the largest being the Great Bank of  
409 Guizhou (Lehrmann et al., 2007). The Nanpanjiang Basin was adjacent to a subduction-zone  
410 volcanic arc along the South China-Indochina plate margin (Cai and Zhang, 2009), where  
411 volcanism was more intense than on the northern margin of the South China Block (e.g., Xie  
412 et al., 2010). This section ranges from the upper Changhsingian through the lower Spathian,  
413 as shown by conodont biostratigraphy (Yang et al., 1986; Tong et al., 2007a). Abundant  
414 gastropods and ostracods are found in the upper Griesbachian (Wang et al., 2001) and prolific  
415 ammonoids, conodonts, and fishes in the lowermost Smithian (Yang et al., 1986). Estimated  
416 depositional water depths in the Zuodeng area were ~30-50 m based on the energy subtidal  
417 feature of the Lower Triassic limestones for the Debao isolated platform (next to the Pingguo  
418 isolated platform, Lehrmann et al., 2007).

419 Carbonate content at Zuodeng is much higher than for the other study sections, averaging

420 ~95% in upper Changhsingian to Spathian strata with a small decrease to ~80% in the upper  
421 Smithian (Fig. 5, Supplementary Table 4). This section consists mainly of thin- to  
422 thick-bedded lime mudstone (cf. Wang et al., 2001). The lack of data around the P-Tr  
423 boundary is due to this interval being covered at the time of sample collection.

424

#### 425 3.4. Mud, Spiti Valley, India

426

427 The Mud section is located in the Spiti Valley, which is part of the district of Lahul  
428 and Spiti, a central area of the western Himalaya in northern India. Lower Triassic strata are  
429 well exposed in this area. During the Early Triassic, the study area was located at  
430 mid-southern latitudes (~30-35°S) on the northern Gondwanan margin (Fig. 4B; Krystyn et al.,  
431 2007). Middle Permian rifting (Stampfli et al., 1991; Garzanti et al., 1996) resulted in the  
432 formation of the Neo-Tethys Ocean (Stampfli et al., 1991; Garzanti et al., 1996), and the  
433 surface uplift of the rift shoulders resulted in widespread non-deposition, erosion and the  
434 unconformities in the stratigraphic record (Stampfli et al., 1991; Garzanti et al., 1996). The  
435 study section was deposited in a mid-shelf setting having a gentle slope, as implied by the  
436 modest water depths of deposition (~50-70 m) despite the distal location of the section  
437 (Krystyn et al., 2007).

438 The base of this section consists of Wuchiapingian to lower Changhsingian strata that are  
439 overlain by an unconformity (or highly condensed interval) spanning the upper  
440 Changhsingian and lower Griesbachian (Bhargava et al., 2004). The main part of the study  
441 section consists of a conformable succession of mid-Griesbachian to lowermost Spathian  
442 strata. Ammonoids are common in the upper Griesbachian to lower Smithian interval  
443 (Krystyn and Orchard, 1996; Krystyn et al., 2007). Average ammonoid shell size decreases  
444 from large in the lower Smithian to small in the middle Smithian (Krystyn et al., 2007),  
445 suggesting the development of more hostile environmental conditions at that time. Current  
446 bedforms are consistent with a well-oxygenated watermass during the earliest Smithian  
447 (Krystyn et al., 2007). The Mud section is a candidate GSSP for the Induan-Olenekian



448 boundary, which was formerly placed at the Bed 12/13 contact (Krystyn et al., 2007) but has  
449 been revised downward to approximately the Bed 9/10 contact (Brühwiler et al., 2010).

450 The Wuchiapingian to lower Changhsingian strata are composed of siliceous shale, with  
451 a ~25% chert fraction (Fig. 5, Supplementary Table 5). A sharp change in lithology occurs at  
452 the P-Tr unconformity, with Lower Triassic strata consisting dominantly of carbonates.  
453 However, low carbonate content is found in limited intervals of the Dienerian and lower  
454 Spathian, which consist mainly of marlstones (cf. Krystyn et al., 2007). For most of Lower  
455 Triassic strata, it consists of thin-bedded argillaceous limestone with shale intercalation (e.g.,  
456 Krystyn et al., 2007) and the average carbonate, shale and chert are ~78%, ~16% and ~6%,  
457 respectively (Fig. 5).

458

#### 459 **4. Results**

460

461 We report raw values for geochemical proxies for terrestrial chemical weathering, marine  
462 productivity, and marine redox conditions in Sections 4.1 to 4.3. We then used an  
463 age-thickness model for each study section (see Supplementary Information for details) in  
464 order to calculate fluxes for the same proxies (Sections 4.4 to 4.6). All raw chemostratigraphic  
465 data and calculated flux values for the four study sections are given in Supplementary Tables  
466 2 to 5.

467

##### 468 *4.1. Weathering proxies*

469

470 We used Al and Fe concentrations as well as the chemical index of alteration (CIA) to  
471 evaluate terrestrial weathering changes during the Early Triassic. In predominantly carbonate  
472 successions such as those of the present study, increases in Al and Fe (which are present  
473 mainly in clay minerals) can be due to climatically controlled fluctuations in subaerial  
474 weathering rates (cf. Sageman et al., 1997). CIA was calculated as  $Al_2O_3 / (Al_2O_3 + K_2O +$   
475  $Na_2O)$  (see Supplementary Information for details). It is a widely used proxy in reconstructing

476 paleoclimate since it is interpreted as a measure of the extent of conversion of feldspars  
477 related to the weathering (Young and Nesbitt, 1998; Price and Velbel, 2003). Note that the  
478 CIA results for each study section are described in conjunction with weathering fluxes  
479 (Section 4.4).

480 At Chaohu, Al ranges from <0.1 to 15.0%, with an average value of 4.6% (Fig. 6A). It  
481 shows generally high values from the P-Tr boundary to the Smithian, followed by generally  
482 lower values in the Spathian. Fe ranges from <0.1 to 14.9%, with an average value of 2.7%.  
483 Fe shows a similar pattern to Al throughout the study section.

484 At Daxiakou, Al ranges from <0.1 to 18.1%, with an average value of 2.3% (Fig. 6B). It  
485 shows lower values in the early Griesbachian and the early Smithian and higher values from  
486 the late Griesbachian through the Dienerian with a short interlude of relatively low values in  
487 the early Dienerian. Fe ranges from <0.1 to 9.0%, with an average value of 1.5%. Fe shows a  
488 similar pattern to Al within the Induan stage.

489 At Zuodeng, Al ranges from <0.1% to 14.3%, with an average value of 0.8% (Fig. 6C). It  
490 shows relatively higher values from the late Dienerian to the early Smithian, in the late  
491 Smithian, and in the middle Spathian but very low values in other intervals. Fe ranges from  
492 <0.1 to 5.0%, with an average value of 0.6%. Fe shows a similar pattern to Al throughout the  
493 study section except for the late Griesbachian, where the Al profile shows several peaks that  
494 the Fe profile does not.

495 At Mud, Al ranges from <0.1 to 11.4%, with an average value of 3.5% (Fig. 6D). It  
496 shows high values from the end of the Griesbachian to earliest Smithian and at the end of the  
497 Smithian but low values during most of the Smithian. Fe ranges from 0.4 to 20.7%, with an  
498 average value of 2.8%. Fe shows high values from the late Griesbachian to earliest Dienerian  
499 and in the early Smithian but low values during most of the Dienerian and Smithian. The Fe  
500 and Al profiles show rather different patterns in this section.

501

502 *4.2. Productivity proxies*

503

504 We used TOC, phosphorus (P), and excess barium ( $Ba_{xs}$ ) concentrations to evaluate  
505 marine productivity fluxes during the Early Triassic.  $Ba_{xs}$  was calculated as the amount of  
506 non-detrital barium (see Supplementary Information for details). These are widely used  
507 proxies for paleomarine productivity since their accumulation depends on organic matter  
508 abundance and preservation (Tribovillard et al., 2006; Calvert and Pedersen, 2007). A method  
509 to estimate actual paleomarine productivity values was developed by Schoepfer et al. (2014,  
510 [this issue](#)), who established regression equations to evaluate primary and export production as  
511 a function of TOC and P mass accumulation rates (MARs) using published data from  
512 Cenozoic sediment cores.

513 At Chaohu, TOC ranges from 0.02 to 5.17%, with an average value of 0.28% (Fig. 7A).  
514 It shows high values at the P-Tr boundary and in the mid Spathian, and moderately high  
515 values from the late Griesbachian to early Dienerian, the late Dienerian to earliest Smithian,  
516 and the mid to late Smithian. In contrast, it shows low values in the mid-Griesbachian, early  
517 Smithian, and most of the Spathian. P ranges from ~0 to 0.52%, with an average value of  
518 0.03%.  $Ba_{xs}$  ranges from 0.29 to 2992 ppm, with an average value of 176 ppm. Both P and  
519  $Ba_{xs}$  show patterns of variation that are similar to that of TOC, although  $Ba_{xs}$  exhibits  
520 relatively higher values in the late Griesbachian and mid to late Smithian.

521 At Daxiakou, TOC ranges from 0.06 to 4.65%, with an average value of 0.30% (Fig. 7B).  
522 It shows high values at the P-Tr boundary but relatively low values from the Griesbachian to  
523 early Smithian. P ranges from ~0 to 2.23%, with an average value of 0.03%. It exhibits a  
524 different pattern from TOC, showing relatively high values at the P-Tr boundary and in the  
525 late Griesbachian to early Smithian and low values in the mid-Griesbachian and early to  
526 mid-Smithian.  $Ba_{xs}$  ranges from 1.51 to 895 ppm, with an average value of 78 ppm. It shows a  
527 slightly different pattern, with high values at the P-Tr boundary and from the  
528 end-Griesbachian to the early Smithian and relatively lower values in the mid-Griesbachian  
529 and mid-Dienerian.

530 At Zuodeng, TOC ranges from 0.06 to 1.52%, with an average value of 0.15% (Fig. 7C).  
531 It shows generally low values for the entire Early Triassic, although with a small increase

532 during the late Griesbachian-early Dienerian and the early Smithian. P ranges from ~0 to  
533 0.30%, with an average value of 0.01%. It exhibits a different pattern from TOC, with  
534 generally high values during the late Smithian and Spathian (punctuated by peaks in the  
535 end-Smithian and mid-Spathian) and low values from the Griesbachian to mid-Smithian. Ba<sub>xs</sub>  
536 ranges from 0.39 to 920 ppm, with an average value of 40 ppm. It shows low values (<30  
537 ppm) through most of the Early Triassic but two peaks in the mid-Dienerian and  
538 mid-Spathian.

539 At Mud, TOC ranges from 0.07 to 2.71%, with an average value of 0.49% (Fig. 7D). It  
540 shows high values in the Dienerian and low values in the late Griesbachian, Smithian, and  
541 early Spathian. P ranges from 0.01 to 0.85%, with an average value of 0.07%. It shows a  
542 gradual upsection decrease. Ba<sub>xs</sub> ranges from 0.1 ppm to 1723 ppm, with an average value of  
543 166 ppm. It shows high values in the Dienerian and late Griesbachian, and low values in the  
544 Smithian and early Spathian.

545

#### 546 4.3. Redox proxies

547

548 We used Mo, U, and V concentrations to evaluate ocean redox changes during the Early  
549 Triassic. Redox-sensitive trace elements typically become enriched in marine sediments under  
550 reducing conditions (Algeo and Maynard, 2004; Algeo and Lyons, 2006; Tribovillard et al.,  
551 2006; Algeo and Tribovillard, 2009). Reducing conditions, characterized by low O<sub>2</sub> and/or  
552 high H<sub>2</sub>S concentrations in bottomwaters, are produced by some combination of decreased  
553 ventilation, commonly due to sluggish watermass circulation, and high respiratory oxygen  
554 demand, commonly due to a high sinking flux of organic matter (Pedersen and Calvert, 1990).

555 At Chaohu, Mo ranges from ~0 to 149 ppm, with an average of 3.4 ppm (Fig. 8A). U  
556 ranges from ~0 to 52 ppm, with an average of 3.5 ppm. V ranges from <1 to 2892 ppm, with  
557 an average of 110 ppm. For all three proxies, high values are observed at the P-Tr boundary  
558 and in the late Griesbachian, Dienerian, and late Smithian. In addition, the V profile exhibits  
559 enrichment in the mid-Smithian, and both the Mo and V profiles show a short episode of

560 somewhat higher values in the early Spathian.

561 At Daxiakou, Mo ranges from ~0 to 604 ppm, with an average of 7.3 ppm (Fig. 8B). U  
562 ranges from ~0 to 39 ppm, with an average of 2.9 ppm. V ranges from <1 to 1326 ppm, with  
563 an average of 53 ppm. For all three proxies, high values are observed at the P-Tr boundary  
564 and in the Dienerian, and V shows an additional peak in the late Griesbachian that is not seen  
565 in the Mo and U profiles.

566 At Zuodeng, Mo ranges from ~0 to 102 ppm, with an average of 3.8 ppm (Fig. 8C). U  
567 ranges from ~0 to 68 ppm, with an average of 6.8 ppm. V ranges from ~0 to 162 ppm, with an  
568 average of 14.0 ppm. For all three proxies, high values are observed in the late Dienerian and  
569 Smithian, with additional enrichment of V in the early Spathian.

570 At Mud, Mo ranges from <1 to 24 ppm, with an average of 2.4 ppm (Fig. 8D). U ranges  
571 from 0.3 to 7.3 ppm, with an average of 2.0 ppm. V ranges from 10 to 528 ppm, with an  
572 average of 132 ppm. For all three proxies, high values are observed in the Dienerian.

573

#### 574 4.4. Weathering fluxes

575

576 At Chaohu, the Al flux ranges from <0.1 to 36 g m<sup>-2</sup> y<sup>-1</sup>, with an average of 6.4 g m<sup>-2</sup> y<sup>-1</sup>  
577 (Fig. 9A). The Fe flux ranges from <0.1 to 25 g m<sup>-2</sup> y<sup>-1</sup>, with an average of 3.9 g m<sup>-2</sup> y<sup>-1</sup>. Both  
578 fluxes increase sharply at the P-Tr boundary and show peak values during the Griesbachian  
579 and Smithian, with a smaller increase in the early Spathian. The CIA ranges from 0.47 to 0.99,  
580 with an average of 0.75. High CIA values are found at the P-Tr boundary and in the  
581 Griesbachian and Smithian.

582 At Daxiakou, the Al flux ranges from <0.1 to 33 g m<sup>-2</sup> y<sup>-1</sup>, with an average of 3.8 g m<sup>-2</sup> y<sup>-1</sup>  
583 (Fig. 9B). The Fe flux ranges from <0.1 to 18 g m<sup>-2</sup> y<sup>-1</sup>, with an average of 2.7 g m<sup>-2</sup> y<sup>-1</sup>. Both  
584 fluxes increase sharply at the P-Tr boundary and show peak values during the Griesbachian  
585 and Smithian. The CIA ranges from 0.50 to 0.96, with an average of 0.80. High CIA values  
586 are found at the P-Tr boundary and in the late Griesbachian and Smithian, with significantly  
587 lower values in the mid-Griesbachian and Dienerian.

588 At Zuodeng, the Al flux ranges from  $<0.1$  to  $11.7 \text{ g m}^{-2} \text{ y}^{-1}$ , with an average of  $0.46 \text{ g m}^{-2}$   
589  $\text{y}^{-1}$  (Fig. 9C). The Fe flux ranges from  $<0.1$  to  $3.7 \text{ g m}^{-2} \text{ y}^{-1}$ , with an average of  $0.37 \text{ g m}^{-2} \text{ y}^{-1}$ .  
590 The Fe flux is relatively larger during the late Griesbachian, whereas the Al flux is greater  
591 during the late Dienerian; both fluxes exhibit higher values during the Smithian and early  
592 Spathian. The CIA ranges from 0.39 to 0.99, with an average of 0.81. Relatively higher CIA  
593 values are observed in the Griesbachian and Smithian.

594 At Mud, the Al flux ranges from  $<0.1$  to  $6.2 \text{ g m}^{-2} \text{ y}^{-1}$ , with an average of  $1.1 \text{ g m}^{-2} \text{ y}^{-1}$  (Fig.  
595 9D). The Fe flux ranges from  $<0.1$  to  $3.7 \text{ g m}^{-2} \text{ y}^{-1}$ , with an average of  $0.8 \text{ g m}^{-2} \text{ y}^{-1}$ . Both  
596 fluxes are large during the late Griesbachian and late Smithian, and Fe additionally shows a  
597 peak around the Dienerian-Smithian boundary. The CIA ranges from 0.40 to 0.98, with an  
598 average of 0.72. High CIA values are observed in the late Griesbachian and Smithian.

599 Summarizing patterns of variation in the weathering proxies, high Al and Fe  
600 concentrations are observed mainly at the P-Tr boundary and in the late Griesbachian,  
601 Dienerian, and mid to late Smithian (Fig. 6). With regard to fluxes, the main peaks in the Al  
602 and Fe profiles are at the P-Tr boundary and in the late Griesbachian and mid to late Smithian  
603 (Fig. 9). Thus, these intervals were probably associated with enhanced inputs of terrestrial  
604 detrital material to the marine study areas. The similar trends of these geochemical proxies  
605 despite differences in lithology among the four study sections suggest that lithologic variation  
606 did not exert a strong influence on these proxies. CIA values show essentially the same  
607 patterns of secular variation as the Al and Fe fluxes. This is a significant observation because  
608 CIA is independent of secular variation in bulk-sediment fluxes and, thus, serves to confirm  
609 patterns of secular variation in the other weathering proxies.

610

#### 611 4.5. Productivity fluxes

612

613 At Chaohu, the TOC flux ranges from 0.01 to  $3.9 \text{ g m}^{-2} \text{ y}^{-1}$ , with an average of  $0.36 \text{ g m}^{-2}$   
614  $\text{y}^{-1}$  (Fig. 10A). The P flux ranges from 0.1 to  $658 \text{ mg m}^{-2} \text{ y}^{-1}$ , with an average of  $48.5 \text{ mg m}^{-2}$   
615  $\text{y}^{-1}$ . The  $\text{Ba}_{\text{xs}}$  flux ranges from  $\sim 0$  to  $451 \text{ mg m}^{-2} \text{ y}^{-1}$ , with an average of  $27 \text{ mg m}^{-2} \text{ y}^{-1}$ . All three

616 proxies show a similar pattern of secular variation, with peak fluxes in the mid to late  
617 Griesbachian and the Smithian, and smaller increases around the P-Tr boundary and in the  
618 early Spathian.

619 At Daxiakou, the TOC flux ranges from  $<0.01$  to  $2.1 \text{ g m}^{-2} \text{ y}^{-1}$ , with an average of  $0.32 \text{ g}$   
620  $\text{m}^{-2} \text{ y}^{-1}$  (Fig. 10B). The P flux ranges from  $0.5$  to  $6395 \text{ mg m}^{-2} \text{ y}^{-1}$ , with an average of  $64.6 \text{ mg}$   
621  $\text{m}^{-2} \text{ y}^{-1}$ . The  $\text{Ba}_{\text{xs}}$  flux ranges from  $0$  to  $350 \text{ mg m}^{-2} \text{ y}^{-1}$ , with an average of  $10 \text{ mg m}^{-2} \text{ y}^{-1}$ . All  
622 three proxies show a similar pattern of secular variation, with peak fluxes in the Griesbachian  
623 and Smithian.

624 At Zuodeng, the TOC flux ranges from  $0.03$  to  $0.93 \text{ g m}^{-2} \text{ y}^{-1}$ , with an average of  $0.09 \text{ g}$   
625  $\text{m}^{-2} \text{ y}^{-1}$  (Fig. 10C). The P flux ranges from  $0.1$  to  $169.5 \text{ mg m}^{-2} \text{ y}^{-1}$ , with an average of  $5.7 \text{ mg}$   
626  $\text{m}^{-2} \text{ y}^{-1}$ . The  $\text{Ba}_{\text{xs}}$  flux ranges from  $\sim 0$  to  $61 \text{ mg m}^{-2} \text{ y}^{-1}$ , with an average of  $2 \text{ mg m}^{-2} \text{ y}^{-1}$ . The  
627 TOC and  $\text{Ba}_{\text{xs}}$  profiles show similar patterns of secular variation characterized by peak values  
628 in the late Griesbachian to early Dienerian, with low values through the remainder of the  
629 section. In contrast, the P profile shows peak values in the late Smithian to early Spathian,  
630 with low values through the remainder of the section.

631 At Mud, the TOC flux ranges from  $0.01$  to  $1.5 \text{ g m}^{-2} \text{ y}^{-1}$ , with an average of  $1.1 \text{ g m}^{-2} \text{ y}^{-1}$   
632 (Fig. 10D). The P flux ranges from  $1.9$  to  $183 \text{ mg m}^{-2} \text{ y}^{-1}$ , with an average of  $20.3 \text{ mg m}^{-2} \text{ y}^{-1}$ .  
633 The  $\text{Ba}_{\text{xs}}$  flux ranges from  $\sim 0$  to  $115 \text{ mg m}^{-2} \text{ y}^{-1}$ , with an average of  $6 \text{ mg m}^{-2} \text{ y}^{-1}$ . Patterns of  
634 secular variation differ among the three productivity-proxy fluxes. The P flux profile most  
635 closely matches secular variation in the South China sections, with high values in the late  
636 Griesbachian and Smithian, and low values in the Dienerian. In contrast, the TOC flux profile  
637 for Mud peaks in the Dienerian and shows low values in the Griesbachian and Smithian, and  
638 the  $\text{Ba}_{\text{xs}}$  flux profile peaks in the mid to late Smithian and shows low values through the  
639 remainder of the section.

640 Summarizing patterns of variation in the productivity proxies, high TOC, P and  $\text{Ba}_{\text{xs}}$   
641 concentrations are found mainly at the P-Tr boundary and in the late Griesbachian, Dienerian,  
642 and mid to late Smithian, with a smaller peak in the early Spathian (Fig. 7). With regard to  
643 fluxes, the main peaks in the TOC, P and  $\text{Ba}_{\text{xs}}$  profiles are in the late Griesbachian and mid to

644 late Smithian (Fig. 10). Thus, these intervals were probably associated with elevated rates of  
645 marine productivity relative to the remainder of the Early Triassic. In contrast to the  
646 weathering proxies, the productivity proxies exhibit low values around the P-Tr boundary,  
647 suggesting a decline in marine productivity during the end-Permian crisis interval.

648

#### 649 *4.6. Redox fluxes*

650

651 At Chaohu, the Mo flux ranges from 0.01 to 11 mg m<sup>-2</sup> y<sup>-1</sup>, with an average of 0.29 mg  
652 m<sup>-2</sup> y<sup>-1</sup> (Fig. 11A). The U flux ranges from 0.01 to 3.0 mg m<sup>-2</sup> y<sup>-1</sup>, with an average of 0.38 mg  
653 m<sup>-2</sup> y<sup>-1</sup>. The V flux ranges from the 0.1 to 174 mg m<sup>-2</sup> y<sup>-1</sup>, with an average of 11 mg m<sup>-2</sup> y<sup>-1</sup>. All  
654 three proxies show similar patterns of secular variation, with peak fluxes in the Griesbachian  
655 and Smithian. The Mo and U profiles also show a peak around the P-Tr boundary, and the Mo  
656 and V profiles show another peak in the early Spathian.

657 At Daxiakou, the Mo flux ranges from 0.01 to 2.1 mg m<sup>-2</sup> y<sup>-1</sup>, with an average of 0.19 mg  
658 m<sup>-2</sup> y<sup>-1</sup> (Fig. 11B). The U flux ranges from 0.01 to 2.6 mg m<sup>-2</sup> y<sup>-1</sup>, with an average of 0.27 mg  
659 m<sup>-2</sup> y<sup>-1</sup>. The V flux ranges from 0.01 to 73 mg m<sup>-2</sup> y<sup>-1</sup>, with an average of 5.9 mg m<sup>-2</sup> y<sup>-1</sup>. All  
660 three proxies show similar patterns of secular variation, with peak fluxes in the Griesbachian  
661 and Smithian. The Mo and U profiles also show a peak around the P-Tr boundary.

662 At Zuodeng, the Mo flux ranges from 0.01 to 7.4 mg m<sup>-2</sup> y<sup>-1</sup>, with an average of 0.24 mg  
663 m<sup>-2</sup> y<sup>-1</sup> (Fig. 11C). The U flux ranges from 0.01 to 3.7 mg m<sup>-2</sup> y<sup>-1</sup>, with an average of 0.41 mg  
664 m<sup>-2</sup> y<sup>-1</sup>. The V flux ranges from 0.01 to 6.4 mg m<sup>-2</sup> y<sup>-1</sup>, with an average of 0.83 mg m<sup>-2</sup> y<sup>-1</sup>. The  
665 three proxies show similar patterns of secular variation, although with minor differences. Peak  
666 values are in the late Dienerian and Smithian for the Mo flux profile, in the mid-Griesbachian,  
667 late Dienerian, and Smithian for the U flux profile, and in the Griesbachian, early Dienerian,  
668 late Dienerian, and Smithian for the V flux profile.

669 At Mud, the Mo flux ranges from 0.01 to 0.35 mg m<sup>-2</sup> y<sup>-1</sup>, with an average of 0.06 mg m<sup>-2</sup>  
670 y<sup>-1</sup> (Fig. 11D). The U flux ranges from 0.01 to 0.41 mg m<sup>-2</sup> y<sup>-1</sup>, with an average of 0.07 mg m<sup>-2</sup>  
671 y<sup>-1</sup>. The V flux ranges from 0.49 to 20 mg m<sup>-2</sup> y<sup>-1</sup>, with an average of 3.5 mg m<sup>-2</sup> y<sup>-1</sup>. The three



672 proxies show similar patterns of secular variation, with peak values in the late Smithian. The  
673 Mo and V profiles exhibit a second, but somewhat smaller, peak in the Dienerian.

674 Summarizing patterns of variation in the redox proxies, high Mo, U and V concentrations  
675 are found mainly at the P-Tr boundary and in the late Griesbachian, Dienerian, and mid to late  
676 Smithian (Fig. 8). With regard to fluxes, the main peaks are in the Griesbachian and Smithian,  
677 although modest increases are found also at the P-Tr boundary and in the early Spathian (Fig.  
678 11) and at Mud during the Dienerian. Thus, these intervals were probably associated with  
679 more reducing conditions in marine environments than the remainder of the Early Triassic.  
680 Secular variation in the redox proxies broadly mirrors that seen for the weathering and  
681 productivity proxies, suggesting close connections between all three environmental  
682 parameters.

683

## 684 **5. Discussion**

685

### 686 *5.1. Relationship of weathering, productivity, and redox variation to Early Triassic global* 687 *events*

688

689 The results above document major secular changes in weathering, productivity, and  
690 redox fluxes during the Early Triassic. In the following discussion, we consider relationships  
691 of these environmental proxies to coeval global events, in order to explore potential controls  
692 on the protracted recovery of Early Triassic marine ecosystems. Our analysis begins with the  
693 end-Permian mass extinction and proceeds through the Spathian, thus covering the full Early  
694 Triassic recovery interval.

695 The end-Permian crisis is generally regarded as having been triggered by the onset of  
696 massive eruptions of the Siberian Traps Large Igneous Province (Renne et al., 1995; Korte et  
697 al., 2010). It was marked by a general collapse of marine ecosystems, as reflected in  
698 biodiversity, trace fossil, and ecological tiering data (Erwin et al., 2002; Erwin, 2005; Fig. 12).  
699 This event was accompanied by an extreme climatic warming of >10°C (Joachimski et al.,

700 [2012; Sun et al., 2012](#)), a major expansion of oceanic anoxia globally ([Brennecke et al., 2011](#)),  
701 an abrupt incursion of sulfidic waters into the ocean-surface layer ([Grice et al., 2005; Algeo et](#)  
702 [al., 2007, 2008](#)), and large inputs of terrestrial material to shallow-marine areas ([Ward et al.,](#)  
703 [2000; Sephton et al., 2005; Xie et al., 2007; Algeo and Twitchett, 2010](#)), all of which are  
704 likely to have contributed to the biocrisis. Strong warming led to intensified stratification of  
705 the oceanic water-column, as reflected in a large vertical gradient of  $\delta^{13}\text{C}_{\text{DIC}}$  ([Song-HY et al.,](#)  
706 [2013](#)), and thus to a strongly reduced nutrient supply via upwelling, contributing to a sharp  
707 decline in marine productivity. The study sections exhibit only limited evidence for these  
708 major environmental changes, however, as the end-Permian and P-Tr boundary are  
709 characterized by, at most, a small increase in terrestrial weathering fluxes ([Fig. 9](#)) and a  
710 transient shift toward more reducing conditions ([Fig. 11](#); cf. [Grice et al., 2005; Cao et al.,](#)  
711 [2009](#)). The muted response of the terrestrial weathering and marine redox proxies in the study  
712 sections may be due to their distance from continental sources of siliciclastics and locations in  
713 areas with only limited local redox changes. Marine productivity exhibits a more visible  
714 change, declining sharply particularly across the South China craton ([Fig. 10](#)), a pattern  
715 possibly related to a productivity crash ([Algeo et al., 2013](#)) or to a shift in dominance from  
716 eukaryotic algae to bacterioplankton ([Luo et al., 2014](#)).

717       During the Griesbachian, the development of a hyper-greenhouse climate resulted in  
718 tropical sea-surface temperatures that were persistently  $>35^\circ\text{C}$  ([Fig. 12; Sun et al., 2012](#)). This  
719 warming contributed to expansion of marine anoxia ([Fig. 11](#)) through lowering of the  
720 solubility of dissolved oxygen in seawater and increasing the flux of river-borne nutrients to  
721 shallow-marine areas via enhanced chemical weathering ([Fig. 9; Algeo and Twitchett, 2010](#)).  
722 A consistently positive relationship is seen between redox conditions and marine productivity  
723 ([Fig. 12](#)), suggesting that organic carbon sinking fluxes controlled the expansion of oceanic  
724 oxygen-minimum zones ([Algeo et al., 2011a](#)). High seawater temperatures and widespread  
725 reducing conditions probably operated in tandem to keep benthic biotas under stress and to  
726 delay marine ecosystem recovery. As a result, benthic biotas were dominated by opportunistic

727 lineages of eurytopic bivalves, gastropods, and ostracods (Erwin, 1998). Relatively high  
728 productivity levels during the Griesbachian (Fig. 10) offered adequate food resources for  
729 nekton, resulting in a transient diversification among conodonts and ammonoids (Stanley,  
730 2009). High productivity may reflect dominance of bacterioplankton (Xie et al., 2010; Luo et  
731 al., 2014), which would have enhanced recycling of nutrients in the ocean-surface layer and  
732 reduced the organic carbon sinking flux (D'Hondt et al., 1998) and, thus, account for a  
733 decrease in the vertical gradient of  $\delta^{13}\text{C}_{\text{DIC}}$  (Song-HY et al., 2013). However, at the end of  
734 Griesbachian, extreme warmth (Sun et al., 2012) and more widespread oceanic anoxia (Fig.  
735 11) destroyed this surface-ocean ecosystem, resulting in a second-order mass extinction  
736 among conodonts and ammonoids (Brayard et al., 2006; Orchard, 2007; Stanley, 2009) and  
737 further depressing the benthic ecosystem. Expansion of the oceanic oxygen-minimum zone at  
738 this time would have resulted in a contraction of the ecospace available to planktic and nektic  
739 organisms (Fig. 13A).

740 The Dienerian was characterized by a warm climate, although one that was slightly  
741 cooler ( $\sim 32\text{-}35^\circ\text{C}$ ) than that of the late Griesbachian (Fig. 12; Sun et al., 2012). As a  
742 consequence of this relative cooling, terrestrial weathering fluxes were reduced (Fig. 9). In the  
743 marine environment, the Dienerian was characterized by lower marine productivity (Fig. 10)  
744 and a shift toward more oxidizing (or less reducing) conditions (Fig. 11). This substage was  
745 associated with a small negative excursion of  $\delta^{13}\text{C}_{\text{carb}}$  (Tong et al., 2007a) and intermediate  
746 and relatively stable vertical  $\delta^{13}\text{C}_{\text{DIC}}$  gradients (Song-HY et al., 2013), which are consistent  
747 with reduced marine productivity as well as a modest weakening of oceanic water-column  
748 stratification. With regard to marine biotas, the Dienerian exhibits increasing diversity among  
749 conodonts and ammonoids (Stanley, 2009) and other marine fauna (Tong et al., 2007a) and an  
750 increase in trace-fossil size (Twitchett, 1999; Chen et al., 2011; Fig. 2). Lower levels of  
751 oceanic productivity were probably associated with a greater proportion of eukaryotic  
752 plankton relative to bacterioplankton, which favored relatively greater export of organic  
753 carbon and nutrients from the ocean-surface layer (cf. D'Hondt et al., 1998). Comparatively

754 cooler climatic conditions and contraction of the oceanic oxygen-minimum zone would have  
755 resulted in an expansion of the ecospace available to conodonts and ammonoids in the surface  
756 ocean (Fig. 13B). Thus, somewhat less severe environmental conditions in the Dienerian  
757 (relative to the Griesbachian) triggered a limited marine ecosystem recovery, although the  
758 brief interval since the end-Permian mass extinction (~0.5 Myr) may have insufficient for a  
759 complete recovery of marine ecosystems (e.g., Kirchner and Weil, 2000).

760 The Dienerian-Smithian boundary was characterized by a transient temperature  
761 minimum (~30-32°; Sun et al., 2012) and a large positive excursion (ca. +6‰) of  $\delta^{13}\text{C}_{\text{carb}}$   
762 globally (Payne et al., 2004; Tong et al., 2007a; Fig. 12). Positive  $\delta^{13}\text{C}$  excursions are  
763 commonly associated with elevated marine productivity (Kump and Arthur, 1999). All four  
764 study sections show a substantial increase in terrestrial weathering fluxes at this time (Fig. 9),  
765 with two (Chaohu and Daxiakou) also showing evidence of increased marine productivity  
766 (Fig. 10). This pattern suggests that the increase in marine productivity may have been driven  
767 by enhanced riverine nutrient fluxes, possibly with an additional stimulus from upwelling of  
768 nutrient-rich deep waters owing to more vigorous thermohaline circulation as a consequence  
769 of climatic cooling and a steeper latitudinal temperature gradient. A concurrent shift toward  
770 somewhat more reducing conditions (Fig. 11) may have been driven by high  $\text{O}_2$  demand  
771 associated with an enhanced sinking flux of organic matter. With regard to marine biotas, this  
772 interval witnessed the maximum diversification of conodonts and ammonoids during the  
773 Early Triassic (Stanley, 2009; Fig. 12), and a limited increase in the diversity of echinoderms,  
774 brachiopods, and forams (Chen et al., 2005; Chen and McNamara, 2006; Song et al., 2011),  
775 suggesting improvements in both the ocean-surface and benthic ecosystems. Cooler  
776 temperatures and a contraction of oceanic oxygen-minimum zones resulted in an expansion of  
777 the ecospace available to marine faunas, and high productivity offered rich food resources for  
778 this ecosystem (Brayard et al., 2006; Orchard, 2007; Stanley, 2009). Enhanced oceanic  
779 overturning circulation generally results in improved ventilation of the global ocean (resulting  
780 in more ecospace availability), while simultaneously intensifying anoxia in limited areas of  
781 active upwelling (owing to greater nutrient fluxes to the ocean surface layer). The

782 Dienerian-Smithian boundary thus represents an episode of significantly ameliorated marine  
783 environmental conditions prior to the onset of the Smithian crisis, and it may have laid a  
784 foundation for more rapid ecosystem recovery at the beginning of the Spathian, ~0.5 Myr  
785 later (e.g., [Erwin, 2008](#); [Chen and Benton, 2012](#)).

786 The Smithian coincided with a major environmental and biotic crisis within the Early  
787 Triassic. It was characterized by development of a second hyper-greenhouse, with peak  
788 temperatures >38°C ([Sun et al., 2012](#); [Romano et al., 2013](#)), a large negative excursion of  
789  $\delta^{13}\text{C}_{\text{carb}}$  ([Payne et al., 2004](#); [Tong et al., 2007](#)), and a maximum vertical gradient in the  $\delta^{13}\text{C}$  of  
790 DIC ([Song-HY et al., 2013](#); [Fig. 12](#)). The negative shift in  $\delta^{13}\text{C}_{\text{carb}}$  is likely to reflect a strong  
791 decline in marine productivity, and the large vertical  $\delta^{13}\text{C}$  gradient an intensification of  
792 oceanic water-column stratification, both in response to extreme warming of the  
793 ocean-surface layer. In the study sections, the Smithian exhibits a large increase in terrestrial  
794 weathering fluxes ([Fig. 9](#)), reflecting stronger chemical weathering due to warming, and  
795 intensified marine anoxia ([Fig. 11](#)), due to a combination of riverine nutrient inputs and  
796 stronger water-column stratification. Productivity levels appear to have increased at this time  
797 in two of the study sections (Chaohu and Daxiakou; [Fig. 10](#)), although it is possible that these  
798 deep-water sections are recording enhanced organic matter preservation as a consequence of  
799 OMZ expansion rather than actual increases in surface-water productivity. Warming and other  
800 environmental stresses resulted in a major extinction event among conodonts and ammonoids  
801 at the end of the Smithian ([Brayard et al., 2006](#); [Orchard, 2007](#); [Stanley, 2009](#); [Fig. 12](#)).  
802 Expansion of the oceanic oxygen-minimum zone at this time would have resulted in a  
803 contraction of the ecospace available to planktic and nektic organisms ([Fig. 13C](#)).

804 The Spathian marks the onset of a sustained recovery of marine ecosystems that was  
805 completed in the Middle Triassic ([Bottjer et al., 2008](#); [Chen and Benton, 2012](#)). It was  
806 characterized by a pronounced climatic cooling from the hyper-greenhouse conditions of the  
807 preceding ~2 Myr, with tropical sea-surface temperatures falling to ~30-32°C ([Sun et al.,](#)  
808 [2012](#); [Romano et al., 2013](#); [Fig. 12](#)). In the study sections, it is marked by large declines in

809 terrestrial weathering fluxes, marine productivity, and the intensity of marine anoxia (Figs.  
810 9-11). The decline in productivity can be attributed to a reduced supply of nutrients from  
811 riverine sources following climatic cooling and stabilization of terrestrial landscapes (Looy et  
812 al., 1999, 2001; Hermann et al., 2011) and from upwelling sources following a flushing out of  
813 the deep-ocean nutrient inventory as a result of re-invigorated thermohaline circulation at the  
814 Smithian-Spathian boundary (Zhang et al., 2014). Improved ocean ventilation and reduced  
815 organic carbon sinking fluxes were responsible for a shift toward less reducing conditions  
816 globally (Fig. 12). These environmental changes coincided with a gradual rediversification of  
817 pelagic organisms, rapid rediversification among benthic organisms and trace-makers, and  
818 higher-level integration of marine trophic systems (Chen et al., 2005; Orchard, 2007; Stanley,  
819 2009; Song et al., 2011; Chen et al., 2011; Chen and Benton, 2012). The more sustained  
820 ecosystem recovery of the Spathian relative to the Dienerian can be attributed to several  
821 factors, including a longer time interval, a cooler climate, less widespread marine anoxia, and  
822 generally more stable environmental conditions (Fig. 13D).

823

## 824 *5.2. Spatial variation in Early Triassic marine environmental conditions*

825

826 Although the four study sections generally show similar patterns of secular variation in  
827 weathering, productivity, and redox proxies, some differences exist among the sections that  
828 are probably controlled by paleogeographic location, water depth, and local bathymetry. First,  
829 the deep-ramp sections (Chaohu and Daxiakou) show peak weathering fluxes that are ~3X  
830 greater than for the mid-shelf section (Mud) and ~10X greater than for the shallow-platform  
831 section (Zuodeng; Fig. 9). These differences reflect relative proximity to sources of detrital  
832 siliciclastics and local bathymetry (e.g., the relative isolation of the shallow-platform section  
833 from detrital influx). Second, average CIA values are somewhat higher in the peri-equatorial  
834 South China sections (~0.75-0.80) relative to the mid-latitude Mud section (0.72), a difference  
835 that is attributable to variations in weathering intensity as a function of climate. Third,  
836 productivity proxy fluxes show some variation among the study sections (Fig. 10). Fluxes are

837 similar for Chaohu, Daxiakou, and Mud but lower for Zuodeng, suggesting diminished  
838 marine productivity on shallow-platform tops relative to open-marine ramp and shelf settings.  
839 Fourth, redox proxy fluxes differ significantly among the study sections, with Chaohu  
840 exhibiting comparatively large fluxes, Daxiakou and Zuodeng intermediate fluxes, and Mud  
841 small fluxes (Fig. 11). These differences appear to be related to both water depth and  
842 paleogeographic location. Chaohu was the deepest section, with water depths of ~300-500 m  
843 putting it within the ocean thermocline and, thus, subject to influence by an expanding  
844 oxygen-minimum zone. However, the higher redox proxy fluxes for all South China sections  
845 relative to Mud suggest that the eastern Paleo-Tethys Ocean was subject to generally more  
846 strongly reducing conditions than the southern Neo-Tethys Ocean during the Early Triassic  
847 (Fig. 4B). Finally, all types of proxies exhibit a better-defined pattern of secular variation in  
848 the deep-ramp sections (Chaohu and Daxiakou) than elsewhere (Figs. 9-11). We attribute this  
849 relationship to differences in depositional water depth, which was >200 m for the deep-ramp  
850 sections but <100 m for the mid-shelf and shallow-platform sections (Section 3). With  
851 increasing water depths, sections were under greater influence by the oceanic  
852 oxygen-minimum zone, expansion of which occurred over discrete time intervals (Feng and  
853 Algeo, 2014). In contrast, the study sections located within the ocean-surface layer (<100 m)  
854 may have experienced more irregular secular variation in environmental conditions.

855

### 856 *5.3. Influences on weathering, productivity, and redox fluxes*

857

858 Modeling of geochemical proxy fluxes suggests a close relationship of changes in  
859 terrestrial weathering intensity, marine productivity rates, and ocean redox conditions  
860 throughout the Early Triassic (Figs. 9-11). Our interpretation, as presented above, is that this  
861 covariation reflects real relationships among these environmental parameters. Specifically,  
862 higher weathering intensities tend to result in increased riverine nutrient fluxes, leading to  
863 enhanced marine productivity (at least in coastal areas), and thus to intensified marine anoxia  
864 (again, possibly focused in coastal areas) (cf. Algeo et al., 1995, 2011a). These relationships

865 are natural consequences of strong climate warming, as occurred repeatedly during the Early  
866 Triassic (Fig. 12; Joachimski et al., 2012; Sun et al., 2012; Romano et al., 2013). We  
867 recognize, however, that the relationships among these environmental parameters may vary in  
868 detail and can have alternative linkages. For example, expansion of marine anoxia can  
869 potentially lead to enhanced organic carbon burial fluxes in the absence of any change in  
870 marine productivity. We also recognize that the patterns exhibited by the four study sections  
871 inherently represent local marine environmental conditions (Fig. 13) that may or may not  
872 mirror contemporaneous global oceanographic changes. However, the strong similarities  
873 among some of the study sections, particularly those deposited at deeper water depths (i.e.,  
874 Chaohu and Daxiakou), suggest that our results have probably captured some aspect of global  
875 marine environmental changes during the Early Triassic.

876 An additional important influence on the proxy fluxes is sediment bulk accumulation rate  
877 (BAR). These fluxes represent the multiplicative product of raw proxy concentrations and  
878 BAR (Supplementary Information), so an increase in either input variable can lead to higher  
879 calculated fluxes. The observation that, for each study section, the concentration profiles (Figs.  
880 6-8) and flux profiles (Figs. 9-11) tend to exhibit similar features reflects the influence of raw  
881 concentrations on calculated fluxes. However, a degree of auto-correlation among the various  
882 proxy fluxes results from the use of a common age-thickness model for each study section  
883 and is unavoidable in studies of this type. The analysis of four widely separated sections (note  
884 that the three Chinese sections cover >1500 km of the South China craton) helps to  
885 compensate for this situation and serves as a test of the validity of results because each study  
886 section makes use of an independent, site-specific age-thickness model (e.g., Section 5.2).  
887 Moreover, our use of CIA also provides a check on the degree of procedural auto-correlation  
888 of results because CIA is a proxy that is completely independent of BAR. CIA shows  
889 unambiguous positive covariation with BAR-based weathering proxies (i.e., Al and Fe) in the  
890 Chaohu, Daxiakou, and Mud sections (Fig. 9), in which CIA exhibits maxima at the PTB or  
891 early Griesbachian, the late Griesbachian, the Dienerian-Smithian boundary or early Smithian,  
892 and the Smithian-Spathian boundary. The weathering proxies in general do not show a



893 coherent pattern of secular variation in the third study section (Zuodeng), possibly because its  
894 paleodepositional setting (i.e., a shallow-marine carbonate platform) was not conducive to  
895 recording weathering fluxes.

896

#### 897 *5.4. Recovery patterns following other Phanerozoic mass extinctions*

898

899 The marine ecosystem recovery following the end-Permian mass extinction is regarded  
900 as having been longer than those following other major Phanerozoic biocrises ([Bottjer et al.,](#)  
901 [2008](#); [Chen and Benton, 2012](#)), although a detailed comparison with other recovery events  
902 has been lacking to date. In the following analysis, we examine patterns of marine ecosystem  
903 recovery following the other four “Big Five” Phanerozoic mass extinctions and consider their  
904 relationship to contemporaneous environmental conditions. This analysis reveals both  
905 commonalities and dissimilarities in the ecosystem recoveries following different mass  
906 extinction events.

907 The Cretaceous-Paleogene (K-Pg; formerly the Cretaceous-Tertiary, or K-T) boundary  
908 mass extinction at 66.0 Ma ([Renne et al., 2013](#)) is the most thoroughly investigated to date. It  
909 killed off ~50% of marine genera and ~70% of species ([Jablonski and Chaloner, 1994](#);  
910 [Sepkoski, 1998](#); [Alroy et al., 2008](#)), with high extinction rates among marine reptiles,  
911 ammonoids, rudist bivalves, planktonic foraminifera, and calcareous nanofossils ([Pospichal,](#)  
912 [1994](#); [Marshall and Ward, 1996](#); [Arenillas et al., 2000](#); [Bown, 2005](#); [Fastovsky and Sheehan,](#)  
913 [2005](#); [Fig. 14A](#)). The coincidence in timing between the Chicxulub bolide impact and the  
914 K-Pg boundary extinction suggests that the impactor was the major cause of this mass  
915 extinction ([Jolley et al., 2010](#); [Renne et al., 2013](#)), although the lethality of the impact may  
916 have been enhanced by long-term environmental stresses associated with the Deccan Traps  
917 eruptions ([Courtillot et al., 1986](#); [Courtillot et al., 1988](#); [White and Saunders, 2005](#); [Chenet et](#)  
918 [al., 2007](#)).

919 Algal primary productivity may have recovered very rapidly, in less than a century  
920 following the K-Pg boundary impact ([Fig. 14A](#), [Sepúlveda et al., 2009](#)). However,

921 redevelopment of an integrated marine ecosystem with extended trophic chains took longer  
922 and proceeded in two stages: a rapid initial phase and a delayed second phase (Coxall et al.,  
923 2006). The initial phase took about 1 Myr (Sepúlveda et al., 2009; Hull et al., 2011; Alegret,  
924 2012) and involved recovery of planktonic organisms such as dinoflagellates (Brinkhuis et al.,  
925 1998; Hildebrand-Habel and Streng, 2003), planktic foraminifera (Coccioni and Luciani, 2006;  
926 Hull et al., 2011), and radiolarians (Hollis et al., 2003), as well as coralline red algae (Aguirre  
927 et al., 2007). The delayed second phase of recovery lasted for ~4 Myr following the boundary  
928 crisis (Coxall et al., 2006) and culminated in a new diversity peak among planktic  
929 foraminifera (Olsson et al., 1999) and re-establishment of reef communities (Baceta et al.,  
930 2005).

931       The nutrient status of the early Danian ocean remains controversial, with some studies  
932 inferring high nutrient levels (Alegret and Thomas, 2009; Alegret et al., 2012) and others low  
933 nutrient levels (Coxall et al., 2006; Fuqua et al., 2008). Warmer conditions (as documented by  
934 a  $\delta^{18}\text{O}$  shift; Fig. 14A) caused the oceanic oxygen-minimum zone to expand, decreasing the  
935 ecospace available to pelagic organisms. These adverse conditions abated 3-4 Myr after the  
936 K-Pg boundary crisis, stimulating an increase in pelagic biodiversity and full redevelopment  
937 of the marine food web (Coxall et al., 2006; Yamamoto et al., 2010). The post-K-Pg-boundary  
938 recovery can be traced through carbon cycle changes in the early Paleocene (Fig. 14A). The  
939 vertical (shallow-to-deep)  $\delta^{13}\text{C}$  gradient in seawater DIC increased in two steps,  
940 corresponding to the two stages of ecosystem recovery (Coxall et al., 2006, Fig. 14A). The  
941 first step is marked by an increase in  $\Delta^{13}\text{C}_{\text{vert}}$  from 0 to ~1‰ within 1 Myr of the crisis, and  
942 the second step by an increase to ~2‰ within 3 million years of the crisis. The latter value is  
943 similar to the  $\Delta^{13}\text{C}_{\text{vert}}$  observed in modern marine systems and, thus, marks the  
944 re-establishment of an efficient biological pump by the late Danian, that is, less intense  
945 recycling of organic matter in the ocean-surface layer resulted in more export to the  
946 chemocline and deep ocean (D'Hondt et al., 1998; Coxall et al., 2006).

947       The Triassic-Jurassic (Tr-J) boundary mass extinction resulted in the demise of ~53% of

948 marine genera (Sepkoski, 1996) and ~80% of species (Sepkoski, 1994). It eliminated  
949 conodonts and severely affected brachiopods and gastropods (McRoberts et al., 1997;  
950 Tomašových and Siblík, 2007), as well as ammonoids, bivalves, corals and ostracods  
951 (McRoberts and Newton, 1995; Kiessling, 2001, 2005; van de Schootbrugge et al., 2007;  
952 Mander and Twitchett, 2008; Fig. 14B). Eruption of the Central Atlantic magmatic province  
953 (CAMP) and the resulting global warming were the major causes of this mass extinction  
954 (Marzoli et al., 1999; McElwain et al., 1999; Hesselbo et al., 2002; van de Schootbrugge et al.,  
955 2009; Schoene et al., 2010; Ruhl et al., 2011).

956 Marine ecosystem recovery began rapidly in the Early Jurassic, within 120 k.y. (Ruhl et  
957 al., 2010) to 290 k.y. of the extinction event (Bartolini et al., 2012). Recovery was marked by  
958 increases in diversity during the earliest Hettangian among pelagic carbonate producers  
959 (radiolarians, calcareous nanofossils) (Clémence et al., 2010), molluscs (McRoberts et al.,  
960 1997), ammonoids (Hesselbo et al., 2002; Guex et al., 2012), and brachiopods (Tomašových  
961 and Siblík, 2007). This initial recovery (Fig. 14B) was followed by a second extinction event  
962 among ammonoids during the early Hettangian (Guex et al., 2004, 2012). A second, longer  
963 (~2-3-Myr) recovery phase (Fig. 14B) is evidenced by more diverse radiolarian assemblages  
964 in the mid to late Hettangian (Longridge et al., 2007), rediversification of ammonoids in the  
965 late Hettangian (Guex et al., 2012), and increases in the diverse of gastropod and coral faunas  
966 in the early Sinemurian (Seuß et al., 2005).

967 A large (~5‰), rapid negative carbon isotope excursion occurred at the end of Triassic  
968 (Korte et al., 2009; Schoene et al., 2010; Bartolini et al., 2012), and was followed a positive  
969 excursion (~3‰) in the lower Hettangian (Williford et al., 2007; Korte et al., 2009; Schoene et  
970 al., 2010; Bartolini et al., 2012; Fig. 14B), suggesting an extreme carbon cycle disturbance at  
971 the T-J boundary. Paleobotanical data provide evidence of a rapid global warming at the T-J  
972 boundary (McElwain et al., 1999; Ruhl et al., 2011), and a subsequent cooling resulted from  
973 high marine productivity and enhanced organic matter burial (Korte et al., 2009). A long-term  
974 negative carbon isotope excursion during the early and middle Hettangian coincided with a  
975 greenhouse climate and widespread oceanic anoxia (Ruhl and Kürschner, 2011; Richoz et al.,

976 [2012; Bartolini et al., 2012](#)), hindering the recovery of marine ecosystems until the  
977 Hettangian-Sinemurian boundary ([Bartolini et al., 2012; van de Schootbrugge et al., 2013](#)),  
978 when the carbon cycle stabilized ([Bartolini et al., 2012; Guex et al., 2012](#)). The initial  
979 recovery of pelagic and benthic organisms during the early Hettangian was impeded first by  
980 large climate fluctuations and then by a long-term global warming that resulted in marine  
981 environmental stresses including warming, seawater acidification, and anoxia ([Richoz et al.,](#)  
982 [2012; van de Schootbrugge et al., 2013](#)). These factors contributed to an extinction event  
983 among ammonoids in the mid-Hettangian and a protracted recovery among many elements of  
984 the marine nekton and benthon during the Hettangian ([Guex et al., 2004, 2012](#)). Falling  
985 atmospheric  $p\text{CO}_2$  by the late Hettangian resulted in a cooler climate and ameliorated marine  
986 environmental conditions, leading to a second stage of recovery among plankton ([Bartolini et](#)  
987 [al., 2012](#)) and benthon ([Seuß et al., 2005](#)) during the Sinemurian stage. Thus, the Early  
988 Jurassic marine ecosystem recovery tracks contemporaneous environmental changes very  
989 well.

990       The Late Devonian mass extinction comprised a series of crises during an interval of ~20  
991 Myr, of which the largest were at the Givetian-Frasnian (G-F), Frasnian-Famennian (F-F), and  
992 Devonian-Carboniferous (D-C) boundaries ([Walliser, 1996, House, 2002; Morrow et al., 2011;](#)  
993 [Fig. 14C](#)). Collectively, these crises killed ~50-60% of marine genera and ~82% of species  
994 ([Jablonski, 1991; McGhee, 1996](#)). Many clades of marine invertebrates suffered multiple  
995 declines, including brachiopods, trilobites, corals, and stromatoporoids ([Copper, 1986; Stearn,](#)  
996 [1987](#)), and most colonial rugose corals went extinct at the F-F boundary ([Fig. 14C, Copper,](#)  
997 [2002; Shen and Webb, 2004](#)). The F-F mass extinction evidenced a collapse of the metazoan  
998 reef ecosystem after the mid-late Devonian acme of metazoan reefs and a replacement by  
999 microbial reefs ([Copper, 2002](#)).

1000       The immediate cause of the Late Devonian crisis appears to have been rapid changes in  
1001 seawater temperatures and redox conditions ([Joachimski et al., 2004; Chen et al., 2005](#)). For  
1002 example, the F-F cooling event severely affected the tropical-marine ecosystem, especially  
1003 reef metazoans and, thus, is the probable cause of this mass extinction ([Copper, 1986, 2002](#)).

1004 However, the ultimate cause of the Late Devonian crisis is likely to have been the spread of  
1005 higher land plants and consequent changes in nutrient cycling (Algeo et al., 1995, 2001;  
1006 Algeo and Scheckler, 1998). A progressive expansion of terrestrial floras during the Devonian  
1007 resulted in intensified chemical weathering of land areas, releasing more nutrients that  
1008 stimulated algal blooms and a consequent expansion of anoxia in epicontinental seas. These  
1009 paleobotanical developments resulted in a long-term decline in atmospheric  $p\text{CO}_2$  owing to an  
1010 increase in both organic carbon burial and silicate weathering, resulting in strong global  
1011 climatic cooling (Algeo et al., 1995). The Late Devonian was a time of transition from the  
1012 Middle Paleozoic greenhouse to the Late Paleozoic icehouse (Fig. 14C; Isaacson et al., 2008).  
1013 Each of the Late Devonian crises coincided with a major global cooling event, the episodes at  
1014 the F-F and D-C boundaries being particularly pronounced (Joachimski et al., 2004; Buggisch  
1015 and Joachimski, 2006; Kaiser et al., 2006, 2008). It is not certain whether the spread of higher  
1016 land plants was gradual and merely created background conditions for the development of  
1017 episodic marine biocrises, or whether it actively triggered each crisis through pulses of  
1018 expansion (Algeo and Scheckler, 2010).

1019 Because of the multi-episode nature of the crisis, there was at least a partial recovery of  
1020 marine ecosystems following each extinction event. For example, brachiopods and ostracods  
1021 underwent a modest recovery during early Famennian, following the F-F boundary event,  
1022 although they remained low in diversity (Casier and Lethiers, 1998; Baliński, 2002; Sokiran,  
1023 2002, Fig. 14C), and stromatoporoids began to recover during the early and middle  
1024 Famennian but went extinct at the D-C boundary (Metherell and Workman, 1969; Stearn,  
1025 1987; Webb, 1998). A permanent recovery did not begin until after the D-C boundary crisis.  
1026 Bryozoans recovered to their pre-extinction level of diversity during the Early Carboniferous  
1027 (Bigey, 1989). Some corals (e.g., *Pseudouralinia*, *Siphonophyllia*) and brachiopods (e.g.,  
1028 *Eochoristites*, *Martiniella*) recovered in the middle Tournaisian of the Early Carboniferous  
1029 (Liao, 2002). The extinction of colonial rugose corals at the F-F boundary resulted in  
1030 replacement of metazoan reefs by small microbial patch reefs (Pickett and Wu, 1990; Webb,  
1031 1998; Morrow et al., 2011), with regrowth of large barrier reefs delayed until the Viséan (Dix

1032 [and James, 1987; Webb, 1998, 1999; Wahlmann, 2002; Fig. 14C](#)). Loss of large metazoans  
1033 during this crisis permitted the establishment of novel ecologies dominated by microbial  
1034 communities ([Wood, 2004](#)). However, the final marine ecosystem recovery required the  
1035 high-trophic-level ecosystem community establishment (e.g., [Chen and Benthos, 2012](#)).

1036 The Late Ordovician (Hirnantian) mass extinction ([Fig. 14D](#)) eliminated ~24% of  
1037 families and 85% of species of marine invertebrates ([Jablonski, 1991; Sepkoski, 1996;](#)  
1038 [Brenchley et al., 2001; Sheehan, 2001](#)). It was particularly severe among trilobites,  
1039 brachiopods, molluscs ([Sepkoski, 1984; Adrain et al., 2000; Harper and Rong, 2008](#)), and  
1040 graptolites and conodonts ([Brenchley et al., 2001; Sheehan, 2001; Fan and Chen, 2007;](#)  
1041 [Rasmussen and Harper, 2011a,b](#)). The immediate cause of this extinction was the Hirnantian  
1042 glaciation ([Brenchley et al., 1995; Gibbs et al., 1997; Sheehan, 2001; Sutcliffe et al., 2006](#)). A  
1043 second extinction, ~1 Myr later, decimated the cool-adapted ‘Hirnantian fauna’ and was  
1044 caused by a rapid termination of glaciation ([Sheehan and Coorough, 1990](#)). The Early Silurian  
1045 (Llandovery epoch) was a transitional interval from the Late Ordovician icehouse to a middle  
1046 Paleozoic greenhouse ([Kaljo and Martma, 2000; Brand et al., 2006](#)). Warming conditions, the  
1047 killing factor in the second extinction, prevailed during the Llandovery, although interrupted  
1048 by two brief glaciations during the Aeronian stage ([Caputo, 1998; Azmy et al., 1998, 1999;](#)  
1049 [Delabroye et al., 2011; Finnegan et al., 2011; Fig. 14D](#)).

1050 Marine ecosystems began to recover following the end-Ordovician extinctions, although  
1051 climate fluctuations during the Aeronian stage complicated the recovery pattern. Microbialite  
1052 resurgence in the immediate aftermath of the Late Ordovician extinction coincided with an  
1053 interval of low-diversity megafaunal communities ([Sheehan and Harris, 2004](#)).

1054 Diversification of brachiopods and trilobites proceeded during the Rhuddanian, the first stage  
1055 of the Early Silurian ([Krug and Patzkowsky, 2004; Owen et al., 2008; Huang et al., 2012](#)),  
1056 representing the initial recovery of marine faunas ([Fig. 14D](#)). Crinoids and coral began to  
1057 diversify from the Rhuddanian in the Early Silurian ([Kaljo, 1996; Ausich and Deline, 2012](#)).

1058 Climate fluctuations during the Llandovery resulted in a delay in the recovery of marine  
1059 ecosystems ([Copper, 2001; Gouldey et al., 2010](#)). Full recovery of reef ecosystems took place

1060 by the mid-Aeronian stage, ~4 Myr after the end-Ordovician crisis ([Copper, 2001](#)).

1061

### 1062 *5.5. Evaluation of hypotheses regarding controls on marine ecosystem recovery*

1063

1064 Three hypotheses have been advanced for the apparent delay in recovery of marine  
1065 ecosystems following the end-Permian mass extinction, linking the duration of the recovery  
1066 interval to: (1) the intensity of the mass extinction ([Sepkoski, 1984](#); [Solé et al., 2002](#)), (2) the  
1067 persistence of harsh environmental conditions ([Hallam, 1991](#); [Isozaki, 1997](#); [Payne et al.,](#)  
1068 [2004](#); [Erwin, 2007](#)), and (3) episodic occurrence of strong environmental disturbances during  
1069 the recovery interval ([Algeo et al., 2007, 2008](#); [Orchard, 2007](#); [Retallack et al., 2011](#); [Fig. 1](#)).  
1070 Our analysis above of four Lower Triassic sections (Chaohu, Daxiakou, Zuodeng, and Mud)  
1071 demonstrates unambiguously that there were large fluctuations in marine environmental  
1072 conditions during the Early Triassic, and these disturbances were linked to transient  
1073 biodiversity crises among coeval marine faunas and relapses in marine ecosystem complexity  
1074 and integration. We conclude that episodic environmental disturbances were integral to the  
1075 pattern and pace of marine ecosystem recovery during the Early Triassic.

1076 Is it possible to draw general inferences about controls on marine ecosystem recovery  
1077 following mass extinctions? With regard to duration, there is considerable variation among  
1078 the recovery intervals following the “Big Five” Phanerozoic mass extinctions. If defined on  
1079 the basis of (1) re-attainment of biodiversity equal to or exceeding pre-crisis levels, and (2)  
1080 re-development of stable, well-integrated trophic systems, then the duration of the recovery  
1081 interval was ~4 Myr for the end-Ordovician crisis, ~10 Myr for the F-F crisis, ~5 Myr for the  
1082 end-Permian crisis, ~2.5 Myr for the end-Triassic crisis, and ~3 Myr for the end-Cretaceous  
1083 crisis ([Fig. 14](#)). These durations are closely linked to the interval of disturbed environmental  
1084 conditions that followed each extinction event. The shortest recovery interval, ~2.5 Myr after  
1085 the end-Triassic crisis, was associated with rapid cooling with minimal climate fluctuations  
1086 during the earliest Jurassic ([Korte et al., 2009](#)), suggesting that amelioration of marine  
1087 environmental conditions proceeded quickly following the CAMP eruptions ([Marzoli et al.,](#)

1088 [1999; McElwain et al., 1999; Hesselbo et al., 2002; Schoene et al., 2010; Ruhl et al., 2011](#)).  
1089 The ~3-Myr-long interval of recovery following the end-Cretaceous crisis was also associated  
1090 with comparatively stable environmental conditions during the early Paleogene ([Coxall et al.,](#)  
1091 [2006](#)). In contrast, the longest recovery interval, ~10 Myr after the F-F crisis, was interrupted  
1092 by two glaciation episodes ([Joachimski et al., 2004; Kaiser et al., 2006, 2008; Isaacson et al.,](#)  
1093 [2008](#)) during which environmental conditions deteriorated and marine ecosystem recovery  
1094 was halted or reversed ([Chen et al., 2005](#)). The second-longest recovery interval, ~5 Myr after  
1095 the end-Permian crisis, was also associated with repeated environmental disturbances ([Algeo](#)  
1096 [et al., 2007, 2008; Retallack et al., 2011](#)). Thus, unsettled environmental conditions following  
1097 the main extinction crisis appear to be a strong control on the pattern and pace of marine  
1098 ecosystem recovery.

1099       Various types of environmental perturbations can contribute to destabilization of  
1100 recovering marine ecosystems. First, temperature is clearly important, as most marine  
1101 creatures are adapted to live within a relatively narrow temperature range ([Brenchley and](#)  
1102 [Harper, 1998](#)). Extreme temperatures were a major factor in delayed ecosystem recovery  
1103 during the Early Triassic ([Sun et al., 2012; Romano et al., 2013](#)), and strong climate  
1104 fluctuations were important during other recovery intervals, e.g., the Late Devonian  
1105 ([Joachimski et al., 2004; Buggisch and Joachimski, 2006; Isaacson et al., 2008](#)) and the Early  
1106 Silurian ([Finney et al., 1999; Gouldey et al., 2010; Finnegan et al., 2011](#)). Second, nutrient  
1107 inventories and patterns of nutrient cycling can be important. Changes related to shifts from  
1108 eukaryotic to microbial primary production following the end-Cretaceous ([D'Hondt et al.,](#)  
1109 [1998](#)) and end-Permian mass extinctions ([Grice et al., 2005; Xie et al., 2010](#)) probably  
1110 influenced rebuilding of marine trophic systems (e.g., [Chen and Benton, 2012](#)). Third, ocean  
1111 redox conditions, which are linked to temperature and nutrient cycling, influence the  
1112 availability of ecospace for metazoans (e.g., [Fig. 13](#)). Development of reducing conditions  
1113 leads to hypercapnia and hypoxemia, which are lethal to most marine invertebrates ([Pörtner,](#)  
1114 [2001](#)). Fourth, ocean acidification, which is commonly linked to elevated atmospheric  $p\text{CO}_2$ ,  
1115 impedes the growth of calcifying organisms. A possible transient increase in seawater acidity



1116 during the Early Triassic (Payne et al., 2010; Hinojosa et al., 2012) and early Paleocene  
1117 (Alegret et al., 2012) are thought to have influenced the rate of recovery of some faunal  
1118 components of marine ecosystems.

1119 Other potential influences on rates of marine ecosystem recovery, i.e., persistently harsh  
1120 environmental conditions (Hallam, 1991; Isozaki, 1997; Payne et al., 2004; Erwin, 2007) or  
1121 magnitude of the extinction event (Sepkoski, 1984; Solé et al., 2002), may play a role as well.  
1122 Although environmental conditions exhibit a tendency to fluctuate strongly following a  
1123 biocrisis rather than remaining persistently harsh, at least some crises were followed by  
1124 protracted intervals of generally inhospitable conditions. The best-documented example is the  
1125 Early Triassic, during which tropical sea-surface temperatures remained persistently high  
1126 (>32°C) for at least 2 Myr following the end-Permian crisis (Sun et al., 2012; Romano et al.,  
1127 2013). The Late Devonian may provide another example, owing to the persistence of strongly  
1128 oxygen-depleted conditions in shallow-marine seas for intervals of millions of years  
1129 following the G-F and F-F crises (Algeo et al., 1995). In both cases, conditions fluctuated  
1130 during these extended intervals of environmental deterioration, yielding no distinct dividing  
1131 line between persistently harsh conditions and recurrent environmental disturbances. With  
1132 regard to the influence of magnitude of the extinction event, there appears to be no strong  
1133 correlation with the duration of the recovery interval (e.g., Kirchner and Weil, 2000; Erwin,  
1134 2001). There are intrinsic limits to how quickly ecosystems are capable of recovery that  
1135 depend on rates of biotic evolution and, thus, re-occupation of vacated ecological niches  
1136 (Sepkoski, 1998; Kirchner and Weil, 2000). However, it appears that such rates are at least an  
1137 order-of-magnitude faster than the durations of even the shorter marine ecosystem recoveries  
1138 (Hairston et al., 2008).

1139

## 1140 6. Conclusions

1141 The overriding control on the pattern and pace of marine ecosystem recovery following a  
1142 mass extinction event is environmental stability or lack thereof. An analysis of environmental  
1143 variation following the end-Permian mass extinction demonstrates that the protracted (~5-Myr)

1144 interval of recovery of Early Triassic marine ecosystems was due to recurrent environmental  
1145 perturbations. These perturbations were associated with high terrestrial weathering fluxes,  
1146 elevated marine productivity, and more intensely reducing oceanic redox conditions, and they  
1147 appear to have been triggered by episodes of strong climatic warming, possibly linked to  
1148 stages of increased magmatism in the Siberian Traps Large Igneous Province. The main  
1149 perturbations following the end-Permian extinction occurred during the early Griesbachian,  
1150 late Griesbachian, mid-Smithian, and (more weakly) the mid-Spathian. These episodes were  
1151 stronger and more temporally discrete in deepwater sections (Chaohu and Daxiakou) relative  
1152 to shallow and intermediate sections (Zuodeng and Mud), probably because warming and  
1153 attendant effects were felt most strongly in the oceanic thermocline region. The observed  
1154 relationships between weathering and productivity fluxes imply that nutrient and energy flows  
1155 were key influences on the pattern and pace of marine ecosystem recovery. Comparison with  
1156 recovery patterns following the other four “Big Five” Phanerozoic mass extinctions suggests  
1157 that marine ecosystem recovery in general depends on the stability of the post-crisis marine  
1158 environment. Persistent environmental stresses may also play a role in the pace of ecosystem  
1159 recovery, but there is no clear correlation to the magnitude of mass extinction event.

1160

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1162

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1167

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2092 **FIGURE CAPTIONS**

2093

2094 **Figure 1.** Three hypotheses to account for the protracted recovery of Early Triassic marine  
2095 ecosystems, linking it to (A) the intensity of the mass extinction (Solé et al., 2002); (B) the  
2096 persistence of harsh environmental conditions (Hallam, 1991; Isozaki, 1997; Payne et al.,  
2097 2004); and (C) the episodic recurrence of major environmental perturbations (Orchard, 2007;  
2098 Stanley, 2009; Algeo et al., 2011a; Retallack et al., 2011). The heavy solid line represents a  
2099 general biodiversity trend (cf. Tong et al., 2007b), and the shaded lines represent extinction  
2100 intensity (A) or environmental stresses (B and C). PTB: Permian-Triassic boundary. ET: Early  
2101 Triassic. MT: Middle Triassic.

2102

2103 **Figure 2.** General patterns of biodiversity and ecological change during the  
2104 Permian-Triassic transition and Early Triassic. Gr. = Griesbachian; Dien. = Dienerian; Sm. =  
2105 Smithian; SWI = sediment-water interface. For the tiering column, positive and negative  
2106 values are elevations in centimeters relative to the sediment-water interface (SWI).  
2107 Biodiversity data: conodont (Orchard, 2007; Stanley, 2009), ammonoid (Stanley, 2009; Yuri  
2108 and Abnavi, 2013), radiolarian (Racki and Cordey, 2000), foraminifera (Song et al., 2011;  
2109 Payne et al., 2011a), brachiopod (Chen et al., 2005; Yuri and Abnavi, 2013), and echinoderm  
2110 (Chen and McNamara, 2006). Trace fossils: diameter (Twitchett, 1999; Chen et al., 2011) and  
2111 ichnodiversity (Chen et al., 2011). Lilliput effect: maximum gastropod size (Payne, 2005) and  
2112 mean foraminifer size (Payne et al., 2011b; Rego et al., 2012). Tiering data (Twitchett, 1999)  
2113 and alpha diversity data (Hofmann et al., 2013, 2014). Recovery stages 1 and 2 are defined in  
2114 this study. The timescale is a modified version of that of Algeo et al. (2013) (see  
2115 Supplementary Table 1).

2116

2117 **Figure 3.** Volcanic and oceanic environmental changes during the Permian-Triassic  
2118 transition and Early Triassic. Abbreviations as in Figure 2. Timing of Siberian Traps eruptions

2119 is interpretative. Data sources:  $\delta^{13}\text{C}_{\text{carb}}$  (Payne et al., 2004), sea-level elevations (Haq et al.,  
2120 1987; Haq and Schutter, 2008), vertical  $\Delta^{13}\text{C}$  of DIC (Song-HY et al., 2013), bioapatite  $\delta^{18}\text{O}$   
2121 (Sun et al., 2012; Romano et al., 2013),  $\delta^{34}\text{S}_{\text{sulf}}$  (Song et al., 2014),  $\delta^{44/40}\text{Ca}$  (Payne et al., 2010;  
2122 Hinojosa et al., 2012), and ocean redox (Kakuwa, 2008; Wignall et al., 2010; Song et al., 2012;  
2123 Grasby et al., 2013).

2124

2125 **Figure 4.** Permian-Triassic paleogeography of (A) South China (modified from Tong et al.,  
2126 2007a), and (B) the world (modified from Algeo et al., 2013). Am = Amuria; Kz =  
2127 Kazakhstan; NC = North China; SC = South China; Tm = Tarim.

2128

2129 **Figure 5.** Stratigraphic variation in lithology of the four study sections. Lithologies  
2130 calculated per Eqs. 1-3 in Supplementary Information. The timescale at left is plotted relative  
2131 to thickness in the Chaohu section and is non-linear; note the different thickness scales for the  
2132 four sections.

2133

2134 **Figure 6.** Chemostratigraphic profiles of weathering proxies (Al and Fe) for the four study  
2135 sections. Vertical scales are identical to those in Figure 5.

2136

2137 **Figure 7.** Chemostratigraphic profiles of productivity proxies (TOC, P, and  $\text{Ba}_{\text{xs}}$ ) for the  
2138 four study sections. Vertical scales are identical to those in Figure 5.

2139

2140 **Figure 8.** Chemostratigraphic profiles of redox proxies (Mo, U, and V) for the four study  
2141 sections. Vertical scales are identical to those in Figure 5.

2142

2143 **Figure 9.** Profiles of weathering fluxes and CIA (chemical index of alteration) for the four  
2144 study sections.

2145

2146 **Figure 10.** Profiles of productivity proxy fluxes for the four study sections.

2147

2148 **Figure 11.** Profiles of redox proxy fluxes for the four study sections.

2149

2150 **Figure 12.** Generalized patterns of marine environmental, biodiversity, and ecosystem  
2151 change in the four study sections during the Early Triassic. The weathering, productivity, and  
2152 redox profiles are based on Figures 9-11. Data sources: sea-surface temperatures (SST) (Sun  
2153 et al., 2012);  $\delta^{13}\text{C}_{\text{carb}}$  and generic diversity (Tong et al., 2007a); fossil abundance (Yang et al.,  
2154 1986; Wang et al., 2001; Tong et al., 2003; Zhao et al., 2007; Krystyn et al., 2007; Li et al.,  
2155 2009; Song et al., 2011); trace fossil burrow size, ichnodiversity, and tiering (Chen et al.,  
2156 2011); and alpha diversity (Hofmann et al., 2013, 2014).

2157

2158 **Figure 13.** Integrated model showing relationships between environmental change,  
2159 shallow-marine ecospace, and marine ecosystem recovery during the Early Triassic. The  
2160 Griesbachian (A) and Smithian (C) are generally characterized by stronger volcanism,  
2161 enhanced weathering and riverine nutrient fluxes, an expanded OMZ, more intense  
2162 water-column stratification, weaker upwelling, and limited ecospace. In contrast, the  
2163 Dienerian (B) and Spathian (D) are generally characterized by weaker volcanism, decreased  
2164 weathering and riverine nutrient fluxes, a contracted OMZ, less intense water-column  
2165 stratification, stronger upwelling, and expanded ecospace.

2166

2167 **Figure 14.** Patterns of environmental change and marine ecosystem recoveries following  
2168 other major mass extinctions: (A) Cretaceous-Paleogene (K-Pg) boundary, (B)  
2169 Triassic-Jurassic (Tr-J) boundary, (C) Late Devonian, and (D) Ordovician-Silurian (O-S)  
2170 boundary. Data sources: (A) volcanism (Renne et al., 2013; Keller, 2003), impacts (Keller,  
2171 2003),  $\delta^{13}\text{C}_{\text{carb}}$  (D'Hondt et al., 1998; Coxall et al., 2006),  $\delta^{18}\text{O}$  (D'Hondt and Zachos, 1993;  
2172 Norris, 1996; Birch et al., 2012), planktic foraminifera diversity (Keller, 2003; Coxall et al.,  
2173 2006; Gallala et al., 2009), nanoplankton diversity (Hull et al., 2011), algal productivity

2174 (Sepúlveda et al., 2009; n.b., S/(S+H) = sterane/(sterane+hopane) ratio), nektonic carnivore  
2175 alpha diversity (Sessa et al., 2012), ecological tiering at the sediment-water interface (Sessa et  
2176 al., 2012); (B) CAMP volcanism (Olsen et al., 2002; Deenen et al., 2010; Blackburn et al.,  
2177 2013; Guex et al., 2012; Deenen et al., 2010; Ruhl et al., 2010),  $\delta^{13}\text{C}_{\text{org}}$  (Ruhl et al., 2011;  
2178 Bartolini et al., 2012; Williford et al., 2007),  $\delta^{18}\text{O}$  (Korte et al., 2009), warm and cool  
2179 intervals (Schoene et al., 2010), sea-level elevations (Haq et al., 1987), diversity of bivalves,  
2180 ammonites, brachiopods, and reefs (Hallam, 1996); (C) sea-level elevations (Johnson et al.,  
2181 1985; Algeo et al., 2007; Isaacson et al., 2008),  $\delta^{13}\text{C}_{\text{carb}}$  (Buggisch and Joachimski, 2006),  
2182 climatic oscillations (Isaacson et al., 2008), tropical sea surface temperatures (SSTs)  
2183 (Joachimski et al., 2004), R-CO<sub>2(atm)</sub> (Berner, 1994; n.b., RCO<sub>2</sub> = ratio of atmospheric CO<sub>2</sub> in  
2184 past to “modern” value of ~300 ppmv), stromatoporoid diversity (Stearn, 1987; Webb, 1998;  
2185 Morrow et al., 2011), reef-building coral diversity (Webb, 1998); and (D) sea-level elevations  
2186 (Johnson et al., 1991; Couto et al., 2013);  $\delta^{13}\text{C}_{\text{carb}}$  (Finney et al., 1999; Cramer et al., 2011;  
2187 Kaljo and Martma, 2000; Gouldey et al., 2010; Delabroye et al., 2011),  $^{87}\text{Sr}/^{86}\text{Sr}$  (Azmy et al.,  
2188 1999; Gouldey et al., 2010), tropical SSTs (Finnegan et al., 2011), glaciations (Caputo, 1998;  
2189 Azmy et al., 1998; Delabroye et al., 2011), crinoid diversity (Ausich and Deline, 2012),  
2190 graptolite diversity (Fan and Chen, 2007), coral diversity (Kaljo, 1996), brachiopod recovery  
2191 stage (Rong and Harper, 1999).  
2192

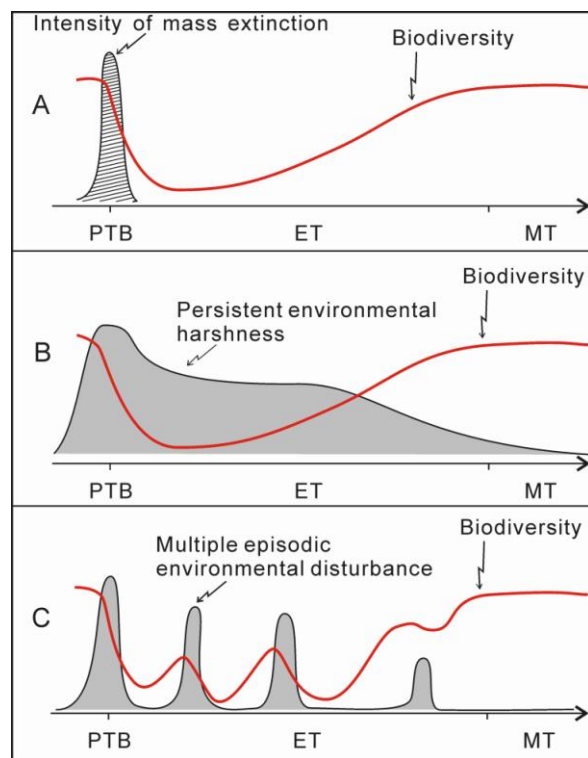


Figure 1.



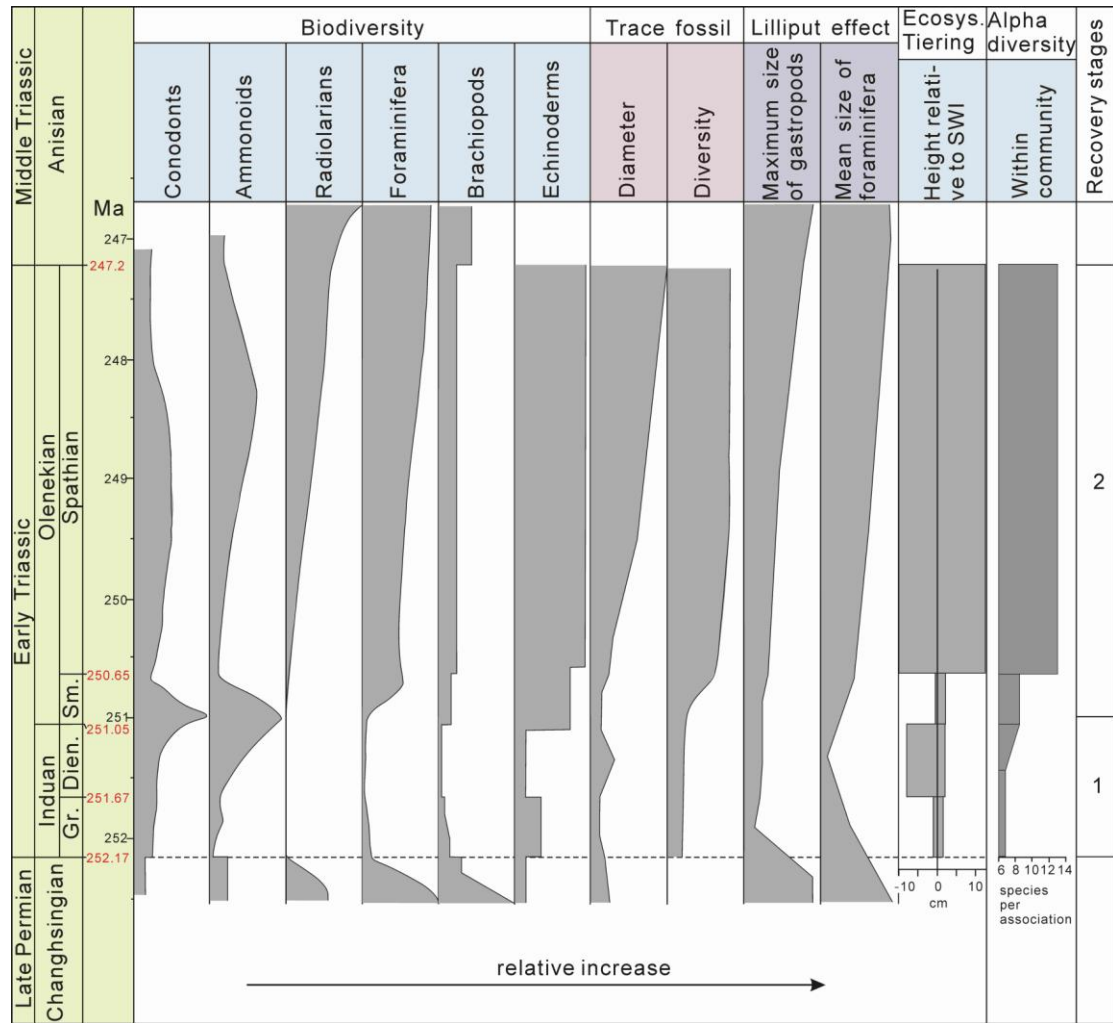


Figure 2.

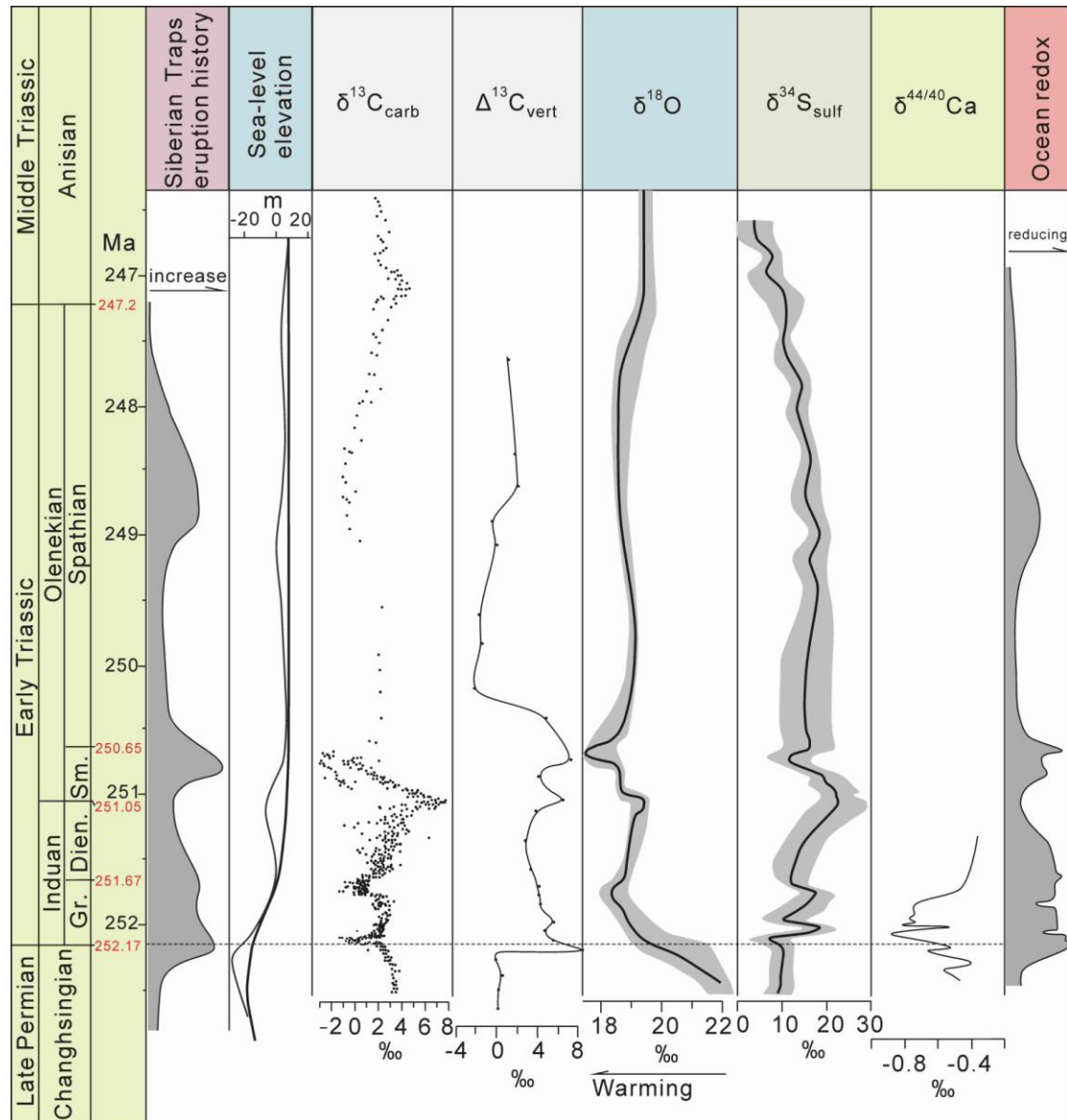


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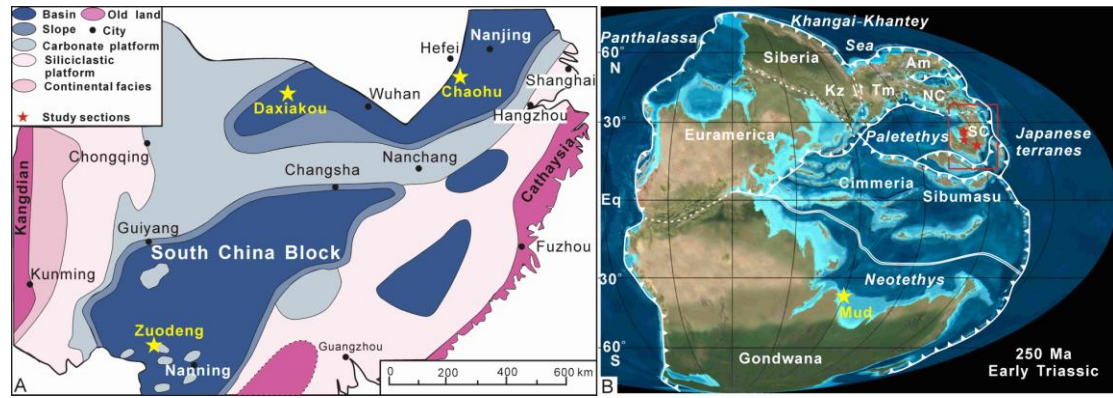


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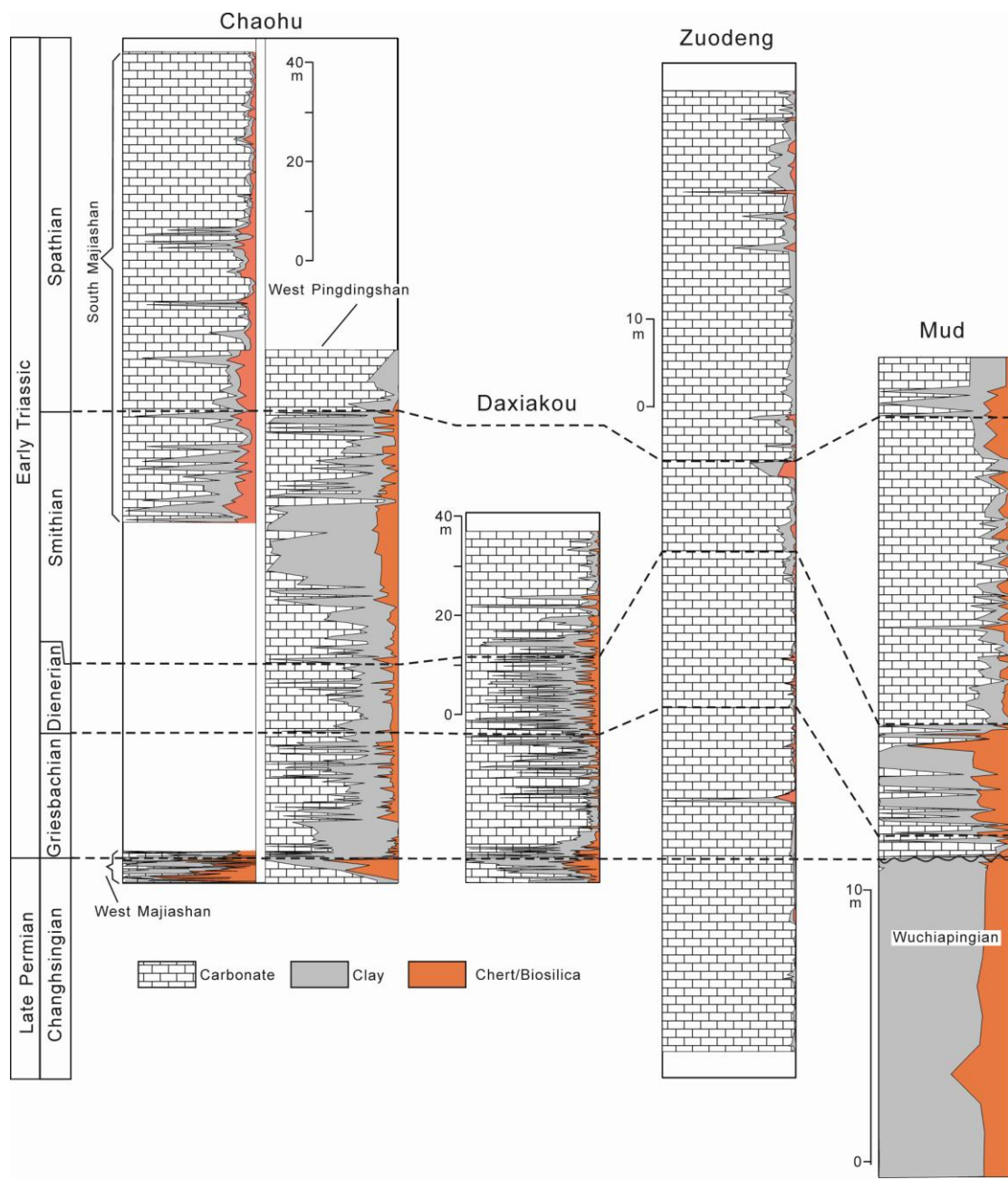


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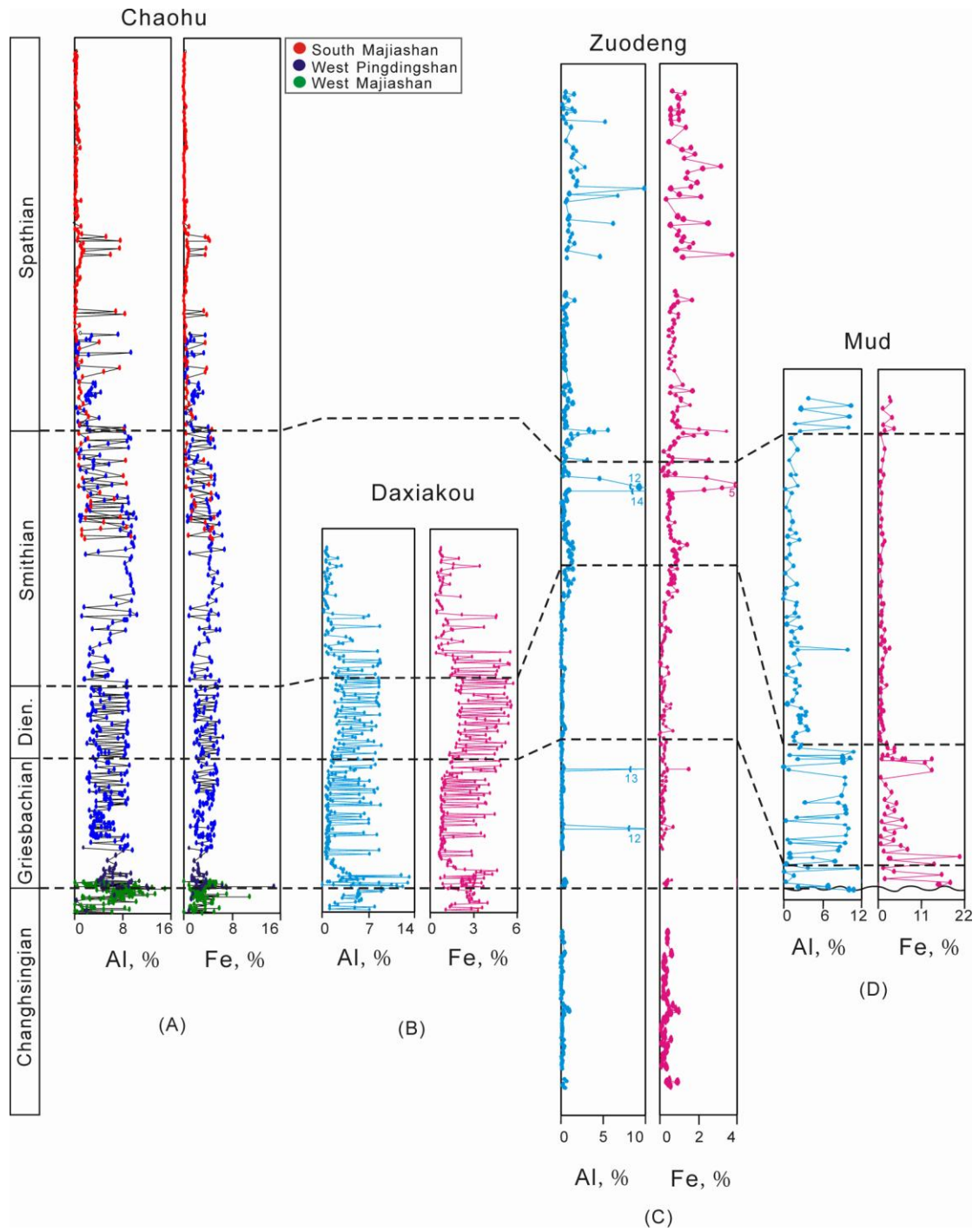


Figure 6.

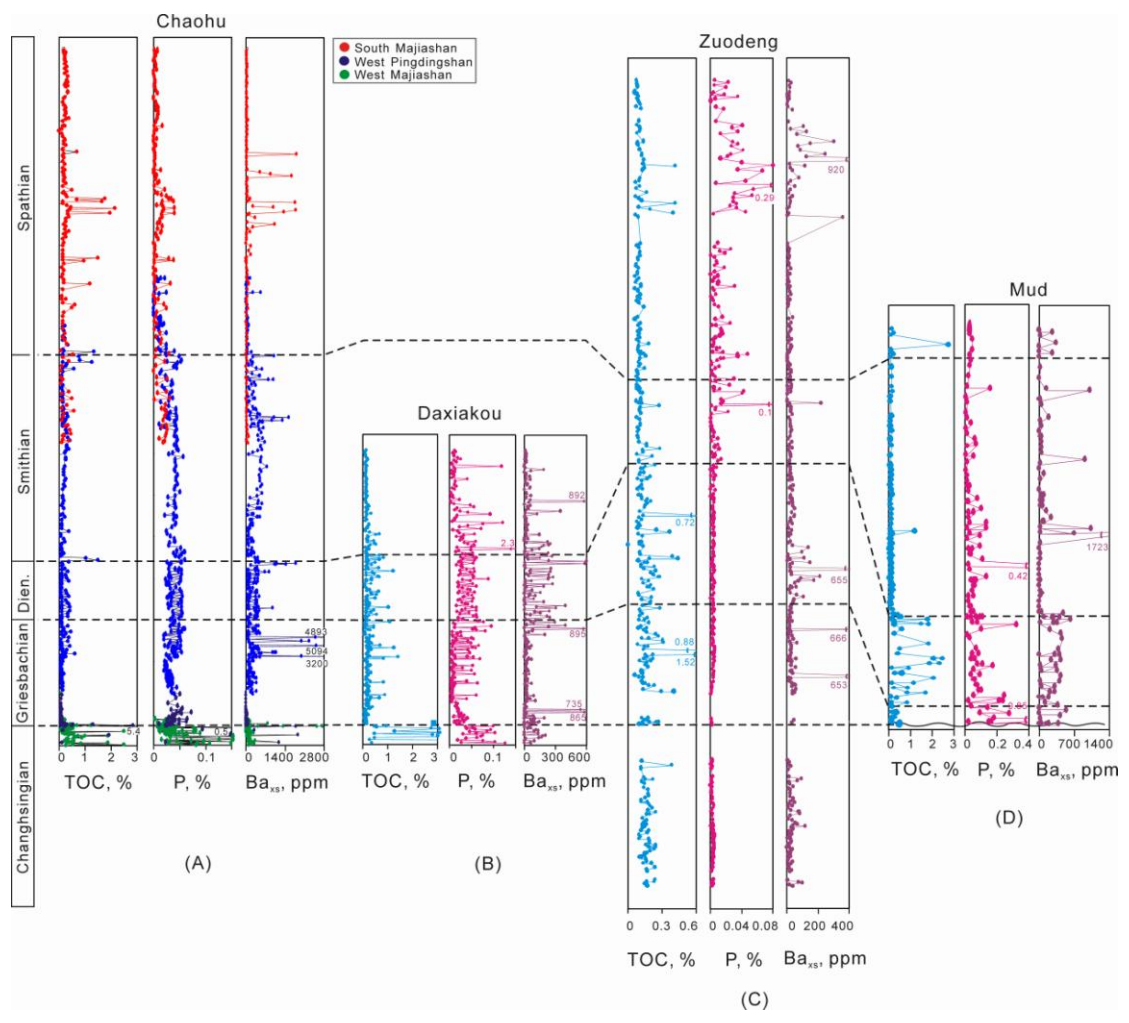


Figure 7.



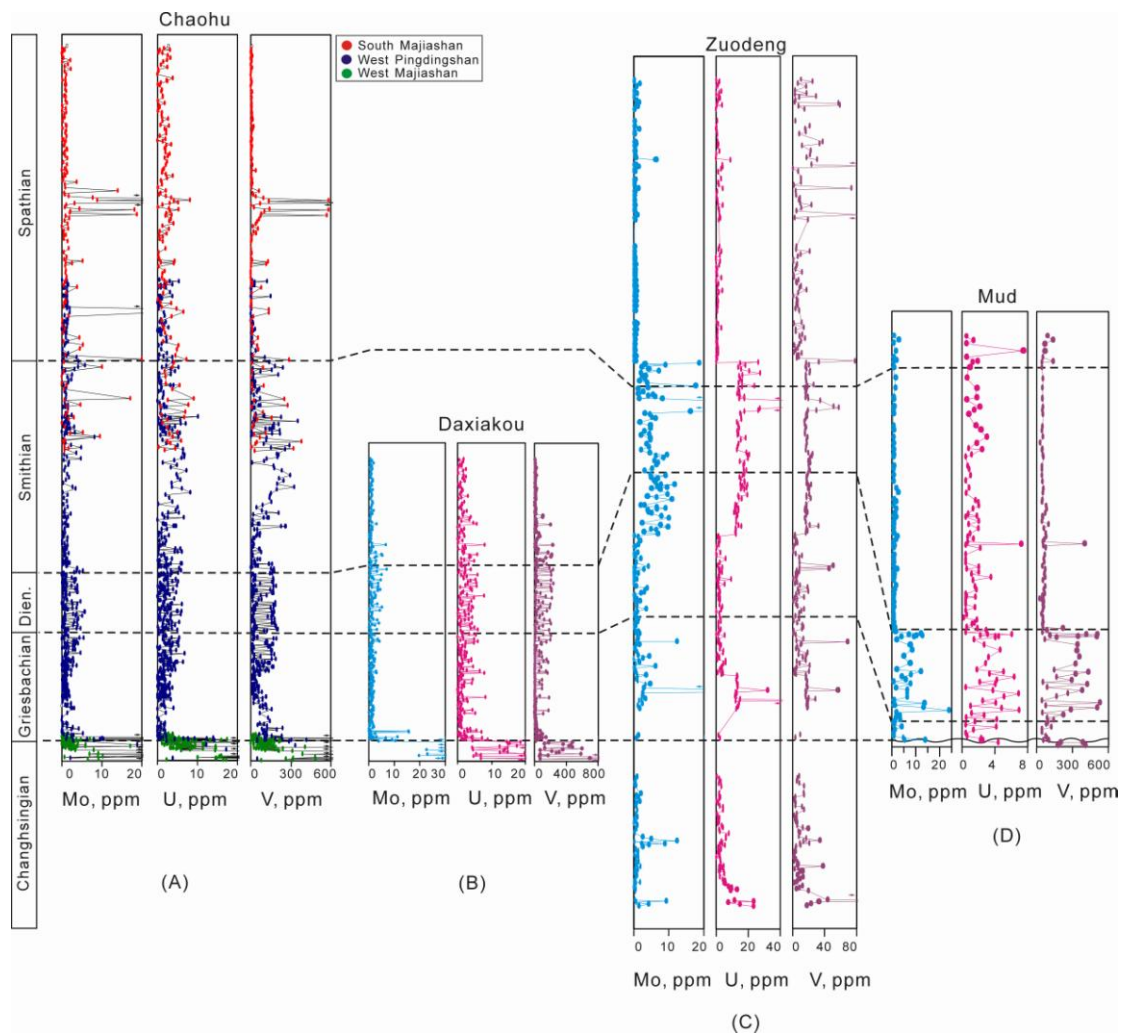


Figure 8.

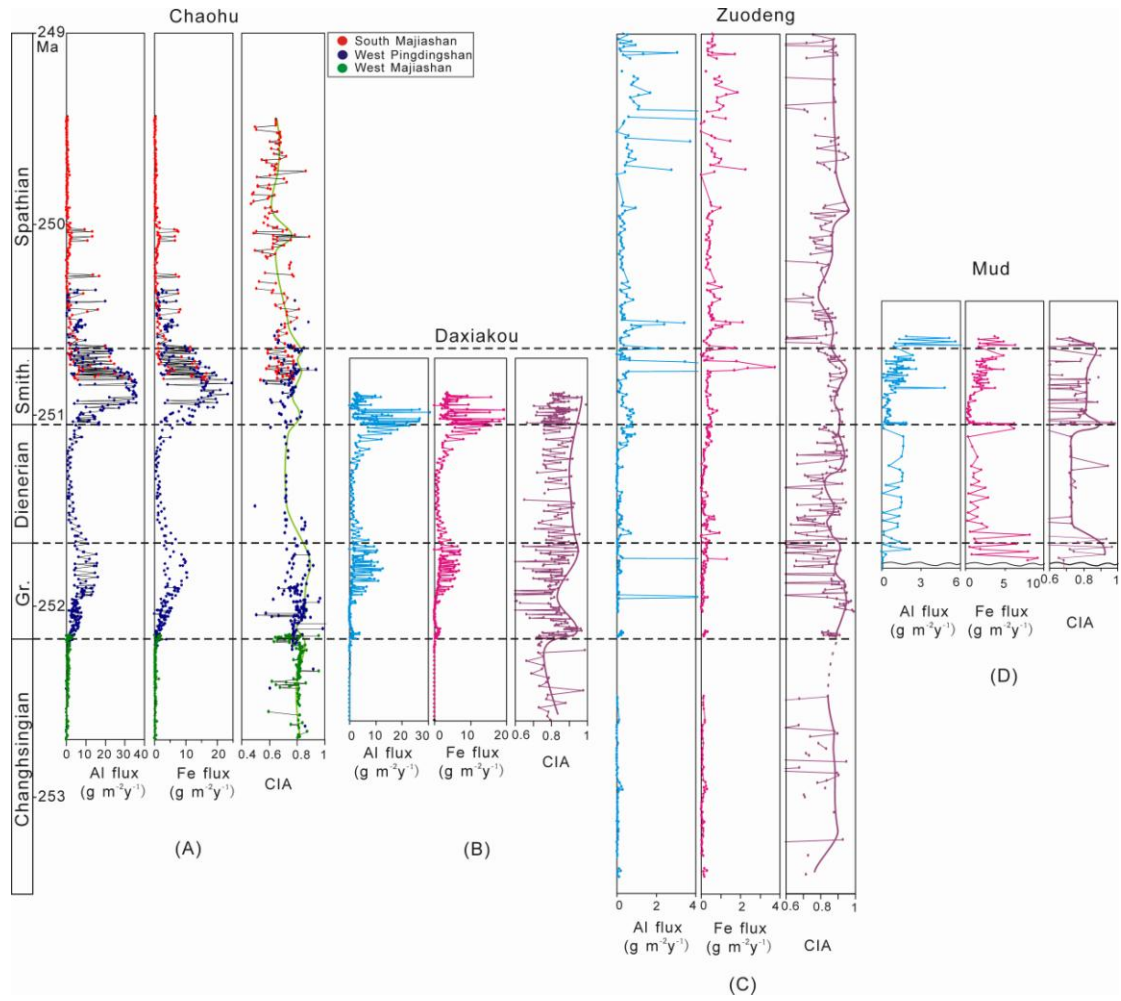


Figure 9.



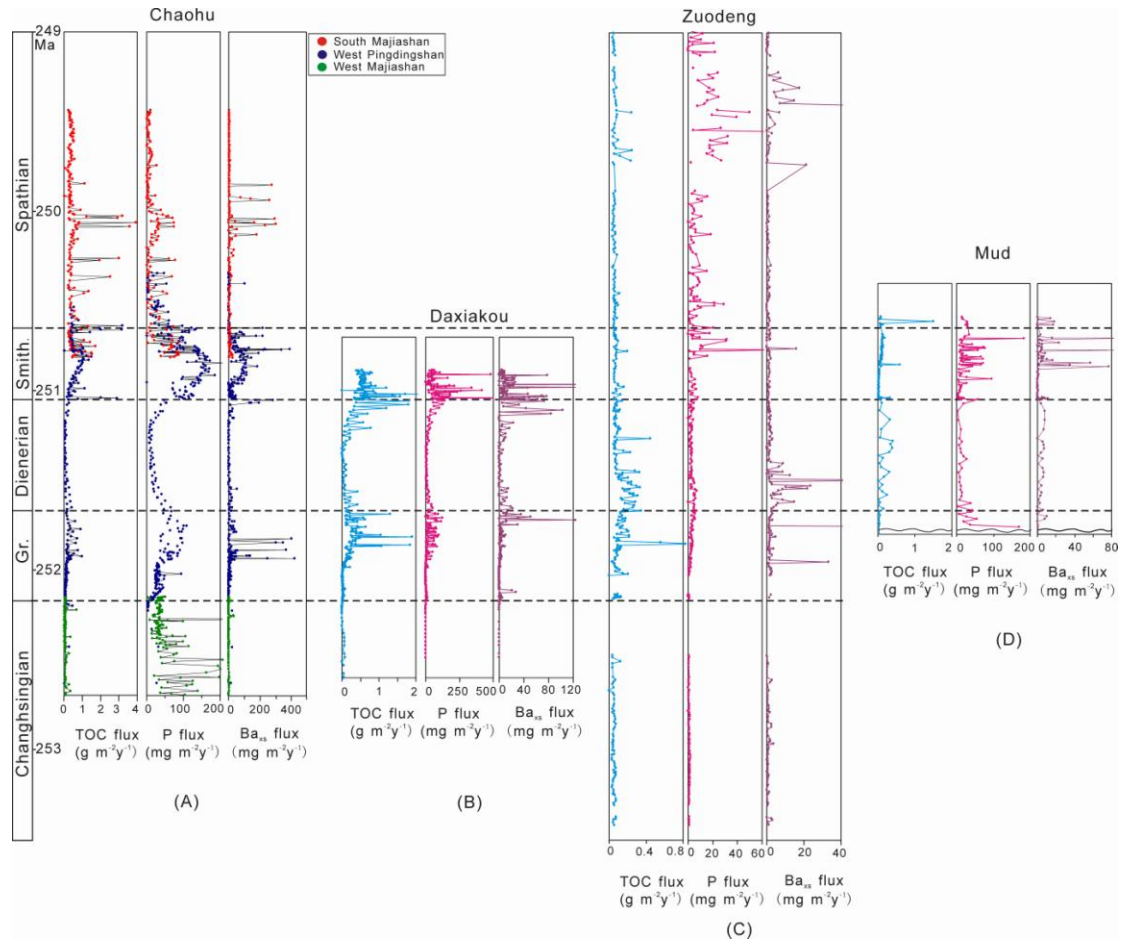


Figure 10.

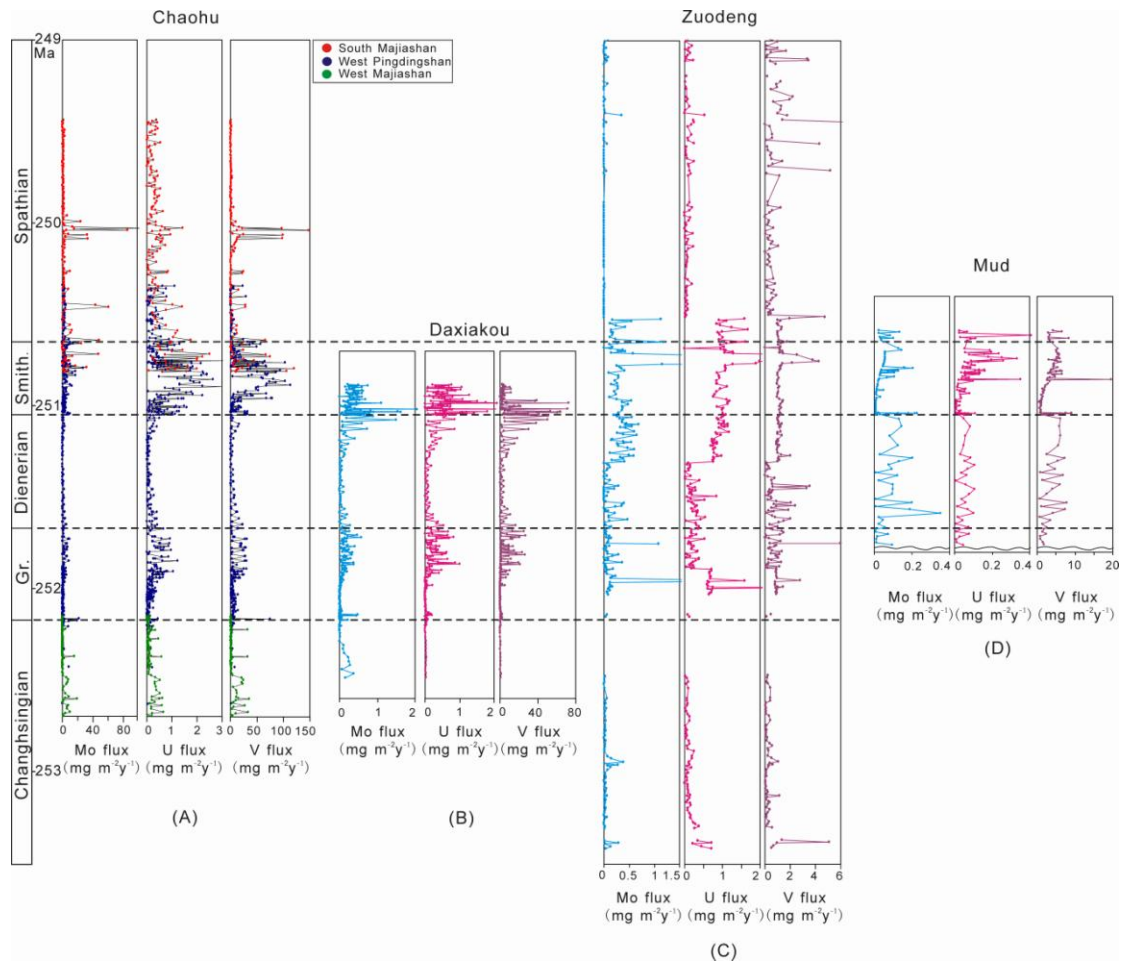


Figure 11.

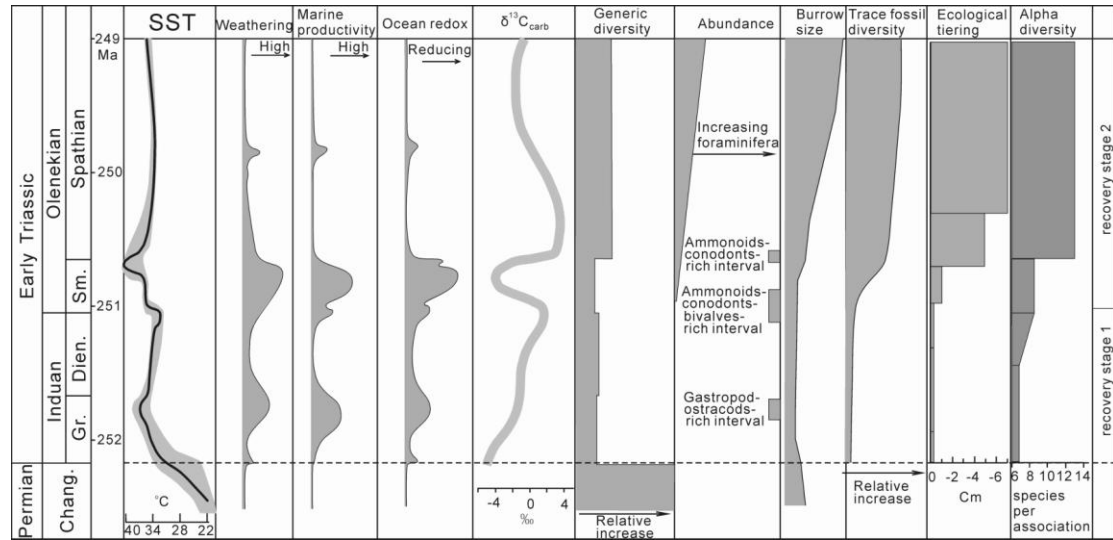


Figure 12.

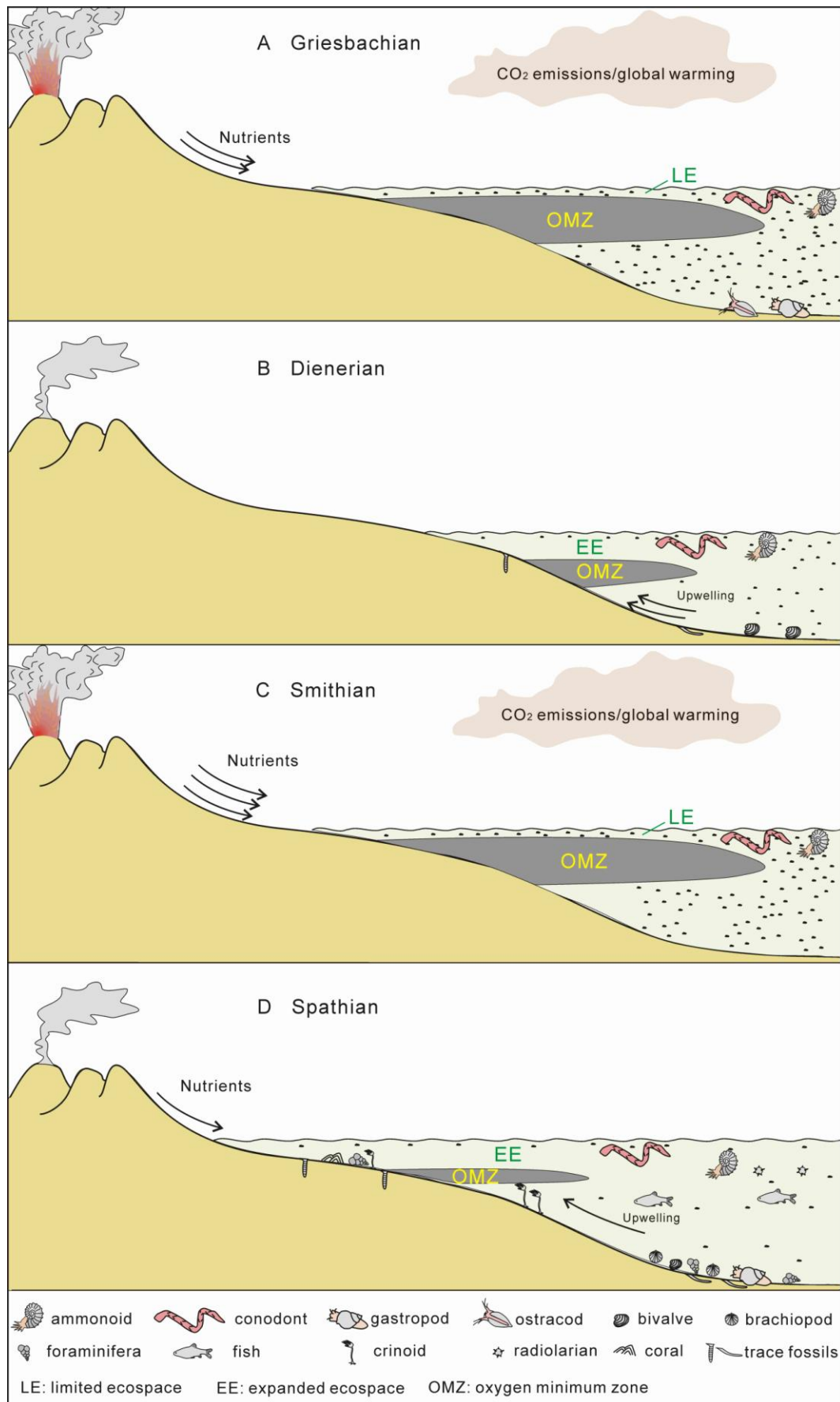


Figure 13.

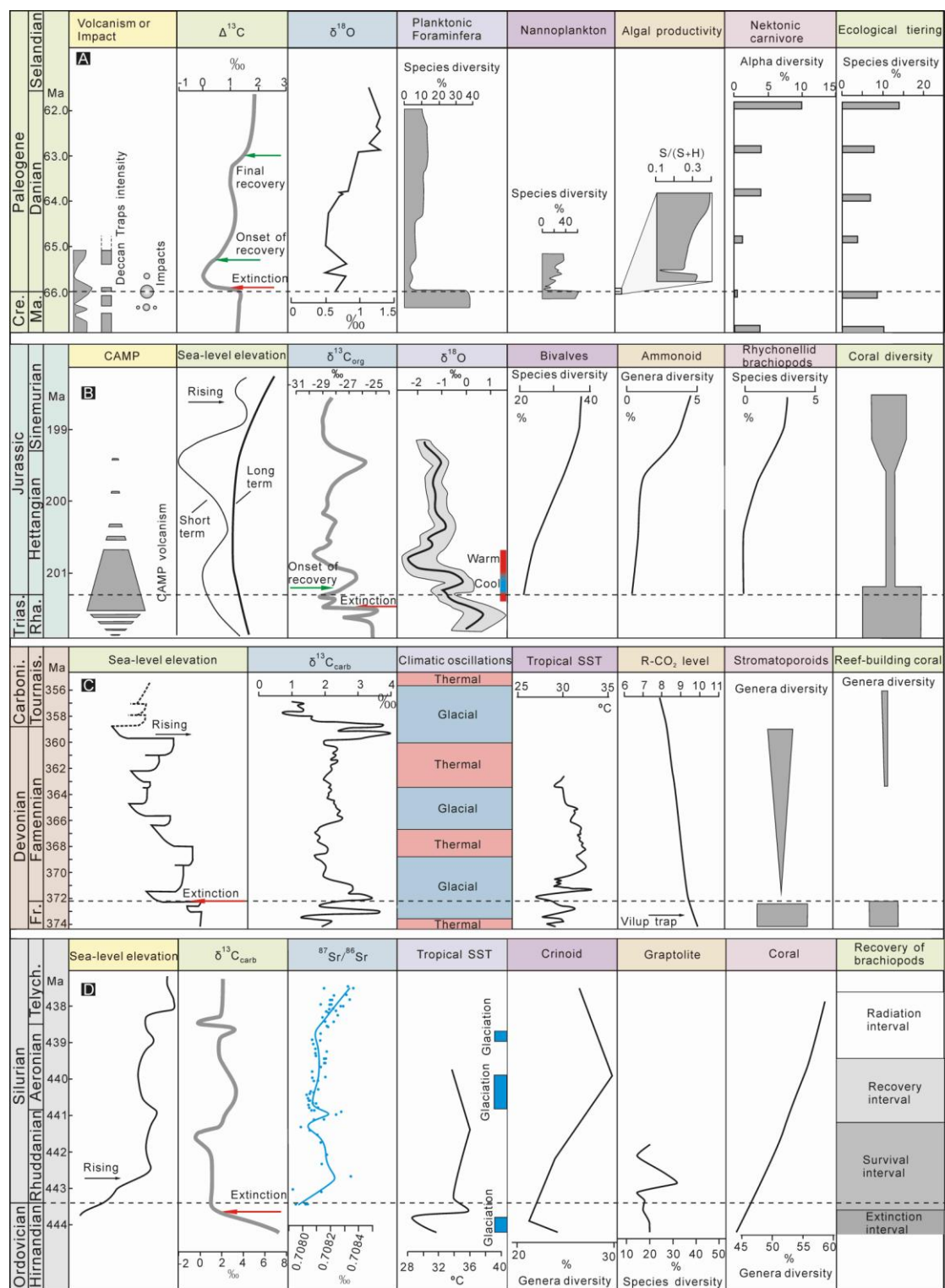


Figure 14.