

Evidence for enhanced ^{10}Be deposition in Mediterranean sediments 35 Kyr BP

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Abstract. In this letter, we report the profiles (~160 samples) of ^{10}Be , ^9Be , Zn in a Mediterranean sediment core spanning the last 60 kyr. We show the existence of a ^{10}Be peak, whose absolute age is estimated to be 34 ± 3 kyr BP, based on its stratigraphic position between two tephra layers originating from two volcanic eruptions (Campanian Ignimbrite and Citara), dated by K-Ar, ^{40}Ar - ^{39}Ar and ^{14}C methods. This peak is similar to the increase in the ^{10}Be concentration observed in Vostok and Dome C ice cores at the same age. The increase with approximately similar amplitude in both ice and sediment cores at different latitudes and hemispheres cannot be explained by changes in the archives. The present result supports the view of a well defined global enhancement of the ^{10}Be flux, related to an increase of the ^{10}Be production rate and not to a redistribution of the ^{10}Be fallout between different latitudes.

Introduction

The ^{10}Be concentration in Vostok and Dome C ice cores, Antarctica, displays an increase by a factor of two at about 35 kyr BP which cannot be explained by changes in the accumulation rate [Yiou et al., 1985; Raisbeck et al., 1987]. Since its discovery in the south polar regions, the ^{10}Be peak has been confirmed in another Antarctic record: in the Byrd core [Beer et al., 1992]. In the northern hemisphere, the evidence for its existence was up to now confined to a limited number of data points in the Camp Century ice core. ^{10}Be anomalies in a sediment core of the Gulf of California were also found around this age [Mc Hargue et al., 1993]. Here we report on a successful search for this peak in a Mediterranean sediment core, at middle northern latitudes. These results from a different archive support the view of a well defined global enhancement of the ^{10}Be flux.

Experimental procedure

The Mediterranean sediment core CT85-5 considered here was taken in the Tyrrhenian Sea ($40^{\circ}19'02''$ N, $11^{\circ}15'42''$ E) at a water depth of 2833 m by the Italian C.N.R. Vessel "Bannock" in 1985. The core has a diameter of 9 cm and a

length of 616 cm. It consists mainly of deep sea clays and carbonates (25 to 50%). We performed ^{10}Be , ^9Be and Zn measurements in about 160 samples of 1 g each (taken from 1 cm intervals) at different depths. ^9Be and Zn were measured in order to estimate biogeochemical and sedimentological processes capable of influencing the ^{10}Be concentration.

The dating of the core CT 85-5 is based on tephrochronology [Cini Castagnoli et al., 1992a, 1993]. The procedure consists of two steps, namely: a) to recognize the different layers of volcanic origin interbedded in the sediment and b) to establish their terrestrial equivalents, which are more suitable for dating.

In the upper 20 cm of the core, historically dated eruptions of Vesuvius of the last 2 Millennia are recognized on the basis of the frequency distribution of volcanic pyroxene grains, characterized by skeletal morphology and sector zoning. The two major pyroxene peaks at 17 cm and 20 cm correspond to the volcanic events of Pollena (472 AD) and Pompei (79AD) leading to a sedimentation rate of ~10 cm/kyr. At greater depths in the lower half of the core two main tephra layers (vitric tuffs) can be used as time markers. The chemical composition of the glass shards was determined by SEM-EDS analysis. The terrestrial equivalents of the marine tephra were recognized [Cini Castagnoli et al., 1993], by comparison of the respective chemical analysis, according to the method of Paterne et al., 1986.

Evidence for an average constant sedimentation rate along the core and the absence of any significant postdepositional process comes also from the spectral analysis of the CaCO_3 content, which reveals very similar spectra in the upper and the lower part of the record [Cini Castagnoli et al., 1992a], and from the good agreement between the $\delta^{18}\text{O}$ curves of CT85-5 [Cini Castagnoli et al., 1992b] and KET 8003 [Paterne et al., 1986] cores for the uppermost 180 cm.

A thick tephra layer composed of trachytic glass (Table 1), admixed with carbonate mud and a few grains of phenocrysts of sanidine, plagioclase, biotite and Na-pyroxene is found between 334 and 290 cm. Extension and chemical analysis make the Campanian Ignimbrite the most probable candidate for the origin of this important marine tephra layer. It represents volcanic products from a major event in the Phleagrean Fields, which caused collapse of a 12 km wide caldera and a flow with an estimated volume of 80 km^3 , covering the entire Campanian plane between Roccamonfina and Salerno [Rosi et al., 1987]. The chemical composition of the land products indicates features similar to those determined for the marine shards (Table 1). Both are trachytic in composition with a $\text{K}_2\text{O}/\text{Na}_2\text{O}$ ratio of ~1.6 and similar Ca, Mg and Fe concentrations. The lower content of total alkalis in the marine glass shards is attributed to slight leaching effects in the marine environment. The thickness of the core section characterized by the Campanian volcanic debris, the sedimentological features and the existence of shallow water

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Table 1. Chemical analysis of the glassy material from the CT85_5 Thyrrhenian core section 290-333 and of the associated land products of the Campanian Ignimbrite (average of 27 samples [Rosi et al., 1987]).

Sample number depth(cm)	SiO ₂	TiO ₂	Al ₂ O ₃	FeO	MgO	CaO	Na ₂ O	K ₂ O	$\frac{\text{K}_2\text{O}}{\text{Na}_2\text{O}}$
290-291	62.3±0.6	0.3±0.2	18.7±0.3	3.5±0.3	0.6±0.2	2.4±0.3	4.4±0.4	7.7±0.7	1.75
333-334	63.0±0.6	0.4±0.2	19.2±0.6	3.2±0.2	0.4±0.2	1.8±0.2	4.7±0.3	7.3±0.3	1.55
Land 33.5±1.5 kyr	61.3±0.7	0.40±0.05	18.5±0.4	3.1±0.2	0.5±0.2	2.1±0.5	5.4±1.4	7.8±0.8	1.6

gastropodes (Niessen, personal communication) indicates that part of this section may be of turbiditic origin.

The Campanian Ignimbrite volcanic rocks have scattered ^{14}C ages (28-38 Kyr B.P.) [Alessio et al., 1973, 1974] and a K/Ar age of 33.5 ± 1.5 kyr [Cassinol et al., 1982]. Recent measurements by ^{14}C and $^{40}\text{Ar}/^{39}\text{Ar}$ methods give respectively (34.6 ± 1) kyr and (37.1 ± 0.4) kyr [Deino et al., 1994]. In the section 339-334 cm, adjacent to the vitric layer, we have found biotite and pyroxene grains, remnants of the starting phase of the eruption. We adopt for this section an age of 34 ± 3 kyr. Following the same procedure of chemical analysis as outlined for the Campanian Ignimbrite, we have identified the glass material of the section 426-402 cm as remnants of the Citara eruption on the island of Ischia 40 ± 2 kyr ago [Gillot, 1984]. The chemical analysis of the Citara marine and land products [Vezzoli, 1988] (Tab.2) display a very similar trachytic composition.

The ^{10}Be peak, which according to the Vostok time scale was expected at ~ 35 kyr BP, is thus flanked in the Tyrrhenian core by the volcanic ash layers of the Citara eruption and interrupted by the Campanian Ignimbrite event.

Analytical methods and results

For the ^{10}Be , ^9Be and Zn analysis the samples were prepared using a 6M HCl leaching technique [Shen et al., 1992]. A set of samples were also prepared with 1M HCl attack; this test procedure showed that 1M HCl randomly leaches only 40 to 90% compared to 6M HCl; moreover it has been shown by Shen et al., 1992, that leaching with 6M HCl removes effectively all Be adsorbed to particle surfaces, which includes Be scavenged from ocean water (including cosmogenic ^{10}Be). ^{10}Be was measured using the AMS facility of ETH/PSI in Zürich. ^9Be and Zn were determined by ICP/AES at EAWAG. The results are displayed as a function of depth for ^{10}Be in Fig.1a, for ^9Be in Fig.1b and for Zn in Fig.1c. In these figures the shadowed areas correspond to tephra layers. We notice that: i) the ^{10}Be profile has a pronounced peak of ~ 2.5 times the average value, interrupted by the Campanian Ignimbrite tephra layer in which it is strongly diluted; ii) ^9Be concentration is highly enhanced in the tephra layers; iii) Zn is relatively constant at all depths.

In detail the profiles of ^{10}Be and ^9Be show three principal regimes, which are characterized by: (I) ordinary ^{10}Be and ^9Be

concentrations: the average values in this section are respectively (6 ± 1) 10^8 at/g and (0.9 ± 0.2) ppm; (II) increasing ^{10}Be up to $15 \cdot 10^8$ at/g and ordinary ^9Be ; (III) volcanic sections: high ^9Be (> 2 ppm); low ^{10}Be ($< 4 \cdot 10^8$ at/g) in the thick Ignimbrite event (C.I. in Fig.1), while ordinary in the less pronounced eruptions (e.g. Citara, C. in Fig.1).

In section II, the ^{10}Be concentration slowly increases with respect to the ordinary background level in an interval spanning 46 cm (384-339 cm) and increases as a spike between 288-273 cm. The ^{10}Be peak is in fact interrupted by the Campanian Ignimbrite event. The maximum value of ^{10}Be ($15.1 \cdot 10^8$ at/g) is reached at 280 cm near the top of the tephra. We note in the same section the ^9Be flatness. The different shapes of the Be isotope profiles are crucial to investigate the origin of the ^{10}Be peak.

In section III, ^9Be is clearly increased by the eruptions, showing a peak in correspondence to each volcanic event. In particular, ^9Be increases in the starting phase of the Campanian Ignimbrite eruption (thin double-shaded area in Fig.1) recorded between 339-334 cm by the presence of biotite and pyroxene grains. It extends with a stronger intensity into the vitric tuffs. Here the ^{10}Be concentration is low due to a dilution in the bulk of tephra matrix.

The Be built into silicate structures will not dissolve until these structures are broken, which requires a whole rock digestion procedure. For this reason, samples were also treated using an HF, HNO₃, HClO₄ bomb digestion [Albrecht et al., 1992]. The ^9Be results averaged over the 3 sections are displayed in Tab.3. We notice that in sections I and II, 6M HCl and total Be are similar, indicating no significant portion of silicate bound Be, while in section III, silicate bound Be increases significantly due to tephra input (here the difference between average leached (2.7 ppm) and whole rock Be (5 ppm) is most striking). At the boundary to the volcanic layer (339-334 cm), characterized by silicate bound Be, ^{10}Be is still elevated, indicating a partial overlap of the ^{10}Be peak with volcanic activity. In the same Tab.3, the results obtained for the 1M leachates show the low (and unreliable) efficiency of this attack. The volcanic influence of ^9Be concentration can be observed both in 6 and 1M leaches. These results clearly show that the volcanic influence on ^9Be can not be avoided regardless of leaching method chosen.

Table 2. Chemical analysis of the glassy material from the CT85_5 Thyrrhenian core 424-425 cm and of the land products of Citara [Vezzoli, 1988].

Sample number depth(cm)	SiO ₂	TiO ₂	Al ₂ O ₃	FeO	MnO	MgO	CaO	Na ₂ O	K ₂ O	$\frac{\text{K}_2\text{O}}{\text{Na}_2\text{O}}$
424-425	64±2	0.50	18.5±0.5	2.8±0.1	0.2±0.2	0.4±0.1	1.3±0.2	5.5±0.3	5.5±0.2	1.0
Land 40±2 kyr	63.0±1	0.5±0.1	18.8±0.4	2.5±0.1	0.1±0.1	0.5±0.1	1.1±0.2	6.2±0.4	6.7±0.1	1.1

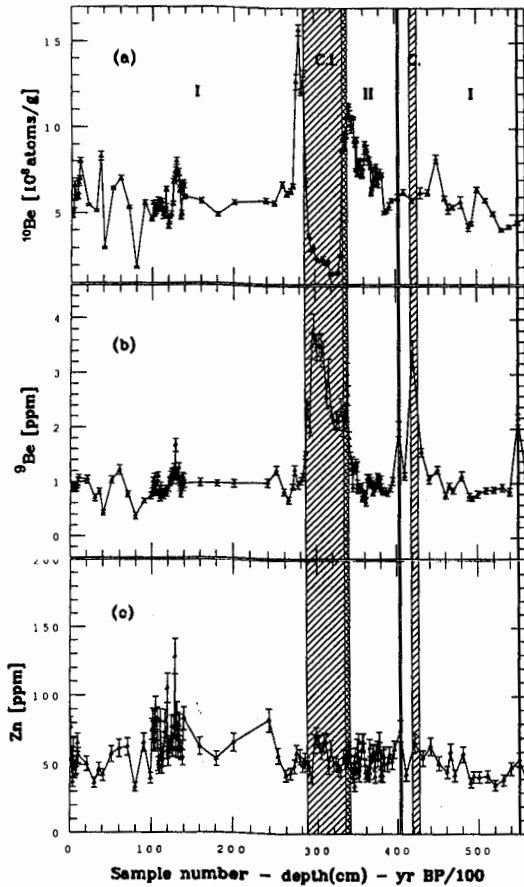


Figure 1. (a) ¹⁰Be, (b) ⁹Be and (c) Zn concentrations as a function of the core depth. Region (I) represents sections of ordinary ¹⁰Be and ⁹Be concentrations; region (II) contains the ¹⁰Be peak. Shaded sections represent tephra. The double shaded area is the starting phase of the Campanian Ignimbrite (C.I.) eruption.

Another trace metal whose profile may through light on significant variations of metal transport mechanisms to the sediment is Zn. It may not have a similar biogeochemical behaviour compared to Be, but it is not influenced by volcanic ashes. In fact the Zn concentration values obtained with 6M HCl leaching, shows a rather constant concentration of ~60 ppm over the last 60 kyr (Fig.1c), with the only significant exception being the deglaciation event around 115 cm and the anthropogenically influenced last 100 yrs (top first cm).

Discussion

¹⁰Be, like other metals in the ocean, is scavenged from the water column by sinking suspended particles [Turekian, 1977]. An increase in the metal concentration found in a sediment core can be explained by an increased atmospheric flux or by an

Table 3. Average ⁹Be concentrations in the CT85-5 Thyrrenian core. Errors are 1 σ standard deviations.

leaching	units	section I	section II	section III
6 M HCl	ppm	0.9±0.2	1.0±0.2	2.6±0.6
1 M HCl	ppm	0.6±0.1	0.7±0.3	1.7±0.7
tot digest	ppm	1.3±0.7	1.2±0.3	5±3

enhanced particle flux and/or lateral transport of dissolved metals, known as sediment focusing or enhanced scavenging [Anderson et al., 1990], here called enhanced sediment pumping. If two metals have similar biogeochemical behaviour and a constant flux, the effect of enhanced sediment pumping should be alike and normalization should obliterate this effect. The most obvious choice to investigate this question is a comparison of ⁹Be with ¹⁰Be. Their sources are different, but transport and transfer pathways are to some extent similar, i.e. particulate dissolved riverine and particulate atmospheric. The most significant difference is the input of ⁹Be by volcanic particles. The question regarding the similar biogeochemical behaviour of both Be isotopes can be rephrased as to whether a large portion of both was in solution a sufficient amount of time for equilibrium to be established.

Concentrations of Be in filtered and unfiltered ocean waters were essentially identical [Measures and Edmond, 1993], indicating that the most significant portion of Be in sea water is in solution. Even eolian dust is partially dissolved [Brown et al., 1992] and Be removed from the trophic zone in particulate form is returned to solution at depth [Peng et al., 1990]. Thus Be is mostly in solution and we can conclude that the isotopes ¹⁰Be and ⁹Be ought to be equilibrated. Any biogeochemical enrichment process that might have affected one should have affected the other as well. We therefore normalize ¹⁰Be to ⁹Be [following the suggestion by Bourles et al., 1989]. The results are displayed in Fig.2a. The position of the peak does not change. The distinct double peak caused by a dilution effect on ¹⁰Be and an increasing ⁹Be are both related to the high volcanic particle flux. The normalization of ¹⁰Be to Zn, the other trace metal whose profile has been measured (Fig.1c), does not change the amplitude and the position of the peak, as shown in Fig.2b. The fact that the normalizations to ⁹Be and Zn changes the ¹⁰Be concentration peak only slightly is a strong argument against any significant change of the sediment pumping during this time.

The mean ¹⁰Be concentration of (6±1) 10⁸ atoms/g corresponds to a flux of (4.3±0.8) 10⁶ atoms/cm²/yr, based on a density of 0.8 g/cm³ and a sedimentation rate of 0.01 cm/yr, which is comparable to that measured in Lake Zürich (5.10⁶

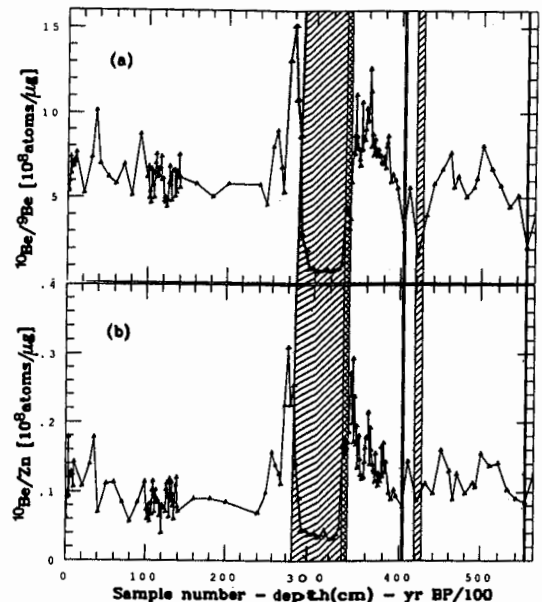


Figure 2.(a) ¹⁰Be normalized to ⁹Be and (b) normalized to Zn.

at/cm²/yr, [Schuler et al., 1991]) and to that of the precipitation in Berne (3.5 10⁶ at/cm²/yr, Luder and Beer, unpublished results). An additional source of ¹⁰Be is related to the input of aeolian dust and riverine particles. Based on ¹⁰Be determined by us in eolian dust samples ((4.4±1.3)10⁸at/g) and suspended Rhone particles (2.5 10⁸at/g) we estimate this additional flux to ~1.2 10⁶ at/cm²/yr. An increased flux of eolian or river borne particles would be reflected in both higher ¹⁰Be and ⁹Be fluxes, confirming in part the findings by Brown et al., 1992. The ¹⁰Be/⁹Be ratio would not change significantly. Our limited results on Saharan dust and Rhone samples give an indication that ¹⁰Be/⁹Be in dust and sediments are comparable.

Finally one might ask which geologic period recorded in our core would be the best candidate to show changes in the sediment pumping, which may alter ¹⁰Be concentration. One candidate is certainly the end of the last glaciation 10 kyr ago (~100 cm in our core). In Fig. 1a, 1b we can see that this period of significant climatic and hydrologic change caused only minor fluctuations in the ¹⁰Be and in ⁹Be concentration. The largest variation in the Zn profile appears indeed around this epoch (Fig. 1c).

These are strong arguments in favour of the hypothesis that the increase in the ¹⁰Be concentration up to a factor ~2.5, lasting a few kyr, is the consequence of an increased production rate and not of a change in physical processes with constant atmospheric ¹⁰Be flux. Further evidence that the ¹⁰Be peak is caused by an increase of the global atmospheric inventory derives from the fact that it has a similar amplitude in the sediment and in the ice cores.

Changes of the production rate can be related to changes of the cosmic ray flux due to helio- and geomagnetic modulation effects or due to fluctuations in the primary cosmic rays flux, e.g. a nearby supernova explosion [Sonett et al., 1987; Kocharov et al., 1992]. A contribution from cosmic dust can be excluded based on two ²⁶Al measurements performed on our core in the peak region. If the observed ¹⁰Be peak is indeed the result of an increase of the production rate, then this has some important implications:

1. All cosmogenic radioisotopes with appropriate half-lives and production rates (e.g. ³⁶Cl and ¹⁴C) should show a peak in the same region.
2. All archives recording with a sufficient time resolution the atmospheric fall-out should contain this signal. Therefore this ¹⁰Be peak provides a unique relative time marker to synchronize ice cores and sedimentary archives; absolute dating of the peak in one specific archive allows to improve the time scales of all other archives.

The fact that the peak is stratigraphically situated between two dated tephra layers at 33.5 and 40 kyr BP allows to estimate its age to 34±3 kyr. The tephra layers offer the opportunity for additional dating leading possibly to a more accurate age of the ¹⁰Be peak.

It will be the subject of future investigations to figure out the precise cause of the observed peak and the environmental implications of an increased cosmic ray flux to the Earth.

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References

Albrecht, A., G.S. Hall and G.F. Herzog, Determination of trace element concentrations in meteorites by inductively coupled plasma-mass spectrometry, *J. Radioanal. Nucl. Chem. Lett.*, 164(1), 13, 1992

Alessio, M., F. Bella, S. Improta, G. Belluomini, G. Calderoni, C. Cortesi and F. Turi, University of Rome C-14 dates X, *Radiocarbon*, 15/1, 165, 1973

Alessio, M., F. Bella, S. Improta, G. Belluomini, G. Calderoni, C. Cortesi and F. Turi, University of Rome C-14 dates XII, *Radiocarbon*, 16/3, 358, 1974

Anderson, R.F., Y. Lao, W.S. Broecker, S.E. Trumbore, H.J. Hoffmann and W. Wolfli, Boundary scavenging in the Pacific Ocean: a comparison of ¹⁰Be and ²³⁴Pa, *Earth and Planet. Sc. Lett.*, 96, 287, 1990

Beer, J., S.J. Johnsen, G. Bonani, R.C. Finkel, C.C. Langway, H. Oeschger, B. Stauffer, M. Suter and W. Wolfli, in *The Last Deglaciation: Absolute and Radiocarbon Chronologies* (ed. Bard, E. and Broecker, W.S.) NATO ASI Series, 12, 141, Springer-Verlag, Berlin Heidelberg, 1992

Bourles, D., G.M. Raisbeck, and F. Yiou, ¹⁰Be and ⁹Be in Marine Sediments and Their Potential for Dating, *Geochim. Cosmochim. Acta*, 53, 443, 1989

Brown, E.T., C.I. Measures, J.M. Edmond, D.L. Bourles, G.M. Raisbeck and F. Yiou, Continental inputs of beryllium to the ocean, *Earth and Planet. Sc. Lett.*, 114, 101, 1992

Cassignol C. and P.Y. Gillot, Range and effectiveness of unspiked potassium-argon dating: experimental groundwork and applications, in *Numerical dating in stratigraphy* (ed. G.S. Odin) 159, Wiley, New York, 1982

Cini Castagnoli, G., G. Bonino, A. Provenzale, M. Scario and E. Callegari, The CaCO₃ profiles of deep and shallow Mediterranean Sea cores as indicators of past Solar-Terrestrial relationships *Nuovo Cimento*, 15C, 547, 1992a

Cini Castagnoli, G., G. Bonino, M. Scario, G.M. Zhu, A.M. Borsetti, L. Capotondi and C. Vergnaud Grazzini, Biostatigraphy and carbonates in the Tyrrhenian abyssal plane as indicators of past climatic variations, *Annales Geophysicae*, 10, suppl. III, 419, 1992b

Cini Castagnoli, G., G. Bonino, E. Callegari, C. Taricco and G.M. Zhu, Thermoluminescence in sea sediments during the cosmogenic isotopes enhancement 35000 yr BP, *Nuovo Cimento*, 16C, 807, 1993

Deino, A.L., J. Southon, G. Orsi, ¹⁴C and ⁴⁰Ar/³⁹Ar dating of the Campanian Ignimbrite, Phlegrean Fields, Italy, Proc. VIII Int. Conf. on Geo-Cosmochronology, Berkeley, June 1994

Gillot, P.Y., *Datations par la methode du potassium argon des roches volcaniques recentes (Pleistocene et Holocene)*. These Dr. Sci. Paris-Sud, 249 pp., 1984

Kocharov, G.E., A.N. Konstantinov, V.A. Levchenko, A.E. Amosov, E.G. Berezhko and G.F. Krymsky, Cosmic rays from supernova explosion, *Rare Nuclear processes*, ed. P. Povinec, World Scientific, Singapore, 203, 1992

Mellargue, L., P. Damon and D. Donahue, ¹⁰Be anomalies, geomagnetic excursions and supernovae, Proc. 23rd Int. Cosmic Ray Conf., Calgary, 3, 854, 1993

Measures, C.I. and J.M. Edmond, The geochemistry of Beryllium-9: A reconnaissance, *Earth Planet. Sci. Lett.*, 66, 101, 1983

Paterna, M., F. Guichard, J. Labeyrie, P.Y. Gillot and J.C. Duplessy, Tyrrhenian sea tephrochronology of the oxygen isotope record for the past 60000 years, *Mar. Geol.*, 72, 259, 1986

Peng, T.-H., T.-L. Ku, J. Southon, C.I. Measures and W.S. Broecker (Ed.), *Factors controlling the distribution of ¹⁰Be and ⁹Be in the ocean*, Bangalore, Indian Academy of Sciences, 1990

Raisbeck, G.M., F. Yiou, D. Bourles, C. Lorius, J. Jouzel and N.I. Barkov, Evidence for two intervals of enhanced ¹⁰Be deposition in Antarctic ice during the last glacial period, *Nature*, 326, 273, 1987

Rosi, M., A. Sbrana, Phlegrean Fields, *C.N.R. Quaderni de "La Ricerca Scientifica"*, 114, 144, 1987

Shen, Ch., J. Beer, L. Tungsheng, H. Oeschger, G. Bonani, M. Suter, W. Wolfli, ¹⁰Be in Chinese loess, *Earth Planet. Sci. Lett.*, 109, 169, 1992

Schuler, C., E. Wieland, P.H. Santschi, M. Sturm, A. Lueck, S. Bollhalder, J. Beer, G. Bonani, H.J. Hoffmann, M. Suter and W. Wolfli, A Multitracer Study of Radionuclides in Lake Zurich, Switzerland, Comparison of Atmospheric and Sedimentary Fluxes of ⁷Be, ¹⁰Be, ²¹⁰Pb, ²¹⁰Po and ¹³⁷Cs, *J. Geophys. Res.* 96, 17,051-17,065 1991

Sonett, C.P., G.E. Morfill and J.R. Jokipii, Interstellar shock waves and ¹⁰Be from ice cores, *Nature*, 330, 458, 1987

Turekian, K.K., The fate of metals in the oceans, *Geochim. et Cosmochim. Acta*, 41, 1139, 1977

Yiou, F., G.M. Raisbeck, D. Bourles, C. Lorius and N.I. Barkov, ¹⁰Be in ice at Vostok Antarctica during the last climatic cycle, *Nature*, 316, 616, 1985

Vezzoli, L., Island of Ischia, *CNR Quaderni de "La Ricerca Scientifica"*, 114, 106, 1988.

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