

RECEIVED: November 21, 2018

REVISED: March 26, 2019

ACCEPTED: June 30, 2019

PUBLISHED: July 11, 2019

Extremal and near-extremal black holes and near-CFT₁

Upamanyu Moitra, Sandip P. Trivedi and V. Vishal

*Department of Theoretical Physics, Tata Institute of Fundamental Research,
Colaba, Mumbai, 400005, India*

E-mail: upamanyu@theory.tifr.res.in, sandip@theory.tifr.res.in,
vishal@theory.tifr.res.in

ABSTRACT: We study the behaviour of extremal and near-extremal black holes at low energies and low temperatures and find that it can be understood from the near-horizon AdS_2 region. Our analysis includes charged matter and also goes beyond the S -wave approximation. We find that the leading behaviour at low energies arises from a mode linked to time reparametrisations and from phase modes arising from gauge fields. At somewhat higher energies, additional modes arising from higher partial waves can also be cumulatively significant. These results can be applied quite generally to cases where an $AdS_2 \times S^d$ near-horizon geometry arises, including black holes in asymptotically AdS and flat space-times.

KEYWORDS: 2D Gravity, AdS-CFT Correspondence, Black Holes, Conformal Field Theory

ARXIV EPRINT: [1808.08239](https://arxiv.org/abs/1808.08239)

Contents

1	Introduction	1
2	Spherically symmetric Reissner-Nordström black holes	3
3	Four point function with scalar exchange in the bulk	7
4	Charged matter coupled to a probe gauge field	10
4.1	The phase mode	12
4.2	Additional comments	13
5	Charged matter coupled to gravity and background gauge field	16
5.1	The two point function	16
5.2	The four point function	17
5.3	The action for time reparametrisation and phase modes	18
5.4	Additional comments	21
6	The near-AdS₂ near-CFT₁ correspondence	22
7	Conclusions	24
A	2-D action for charged matter coupled to a probe U(1) field	25
B	4-D action computation of 4 point function in background gauge field	28
C	Phase mode action	32

1 Introduction

The study of extremal and near-extremal black holes and black branes has proved to be a rich and rewarding area of investigation. The low energy and small temperature response in these systems is expected to arise from the near-horizon region where the gravitational red shift is big. For several brane configurations it is known that the near-horizon region contains an $AdS_n, n > 2$, factor and the resulting low energy dynamics is described by a conformal field theory (CFT) living on the boundary of this region. The CFT plays the role of an effective theory governing the interactions of the black brane with the outside at low energies. More remarkably, there is good evidence now that the CFT is often exactly equivalent to the gravity theory in the interior region described by the AdS_n geometry.

In contrast to black branes, which are extended, the situation vis-à-vis black holes has been less clear. The near-horizon region for extremal black holes often contain an AdS_2 factor. It is now known that the AdS_2 geometry can be used to correctly compute the

ground state degeneracy of extremal black holes to remarkable precision, [1, 2]. However, it has also been long suspected that for probing low energy excitations above extremality the AdS_2 geometry does not suffice and corrections would need to be incorporated. For example, it has been shown that the back reaction to a probe in AdS_2 would diverge rendering the AdS_2 unstable to any perturbation, [3]. The underlying physical reason is that all the directions transverse to the AdS_2 are finite for a black hole and this simply does not allow the perturbations enough room to dissipate.

What has been less clear, until recently, is how to incorporate the corrections to the AdS_2 geometry in a systematic way in analysing the low energy response. Interestingly, progress in answering this question has come from the study of quantum mechanical models in condensed matter physics, pioneered by Sachdev and Ye, [4], and Kitaev, [5], and extended in subsequent papers, [6–33] (we shall refer to the class of such models as SYK models below). It was found that an approximate conformal theory does arise in the IR in these models, but the breaking of conformal invariance cannot be neglected even at very low energies. This breaking is important in giving a finite action for a mode tied to the breaking of time reparametrisation invariance, and this mode is in fact responsible for the thermodynamics at low temperatures and dominates the low energy response of the system.

Remarkably, very similar behaviour was found to arise in a theory of two dimensional gravity coupled to a scalar called the dilaton — the Jackiw-Teitelboim (JT) model, [34, 35]. In this model the geometry is identically AdS_2 , with $SL(2, R)$ isometries, but the varying dilaton breaks the $SL(2, R)$ symmetry [36]. The fluctuations of the boundary of AdS_2 space then acquire a finite action which is of the same form as that in the SYK model, involving the Schwarzian derivative of the time reparametrisation of the boundary.

The study of SYK models was extended to models with complex fermions with conserved charge in [14, 16] It was found that in this case an additional mode arises corresponding to the phase degree of freedom conjugate to the charge. The breaking of conformal invariance is important for this mode as well, and its action is also suppressed by the scale at which the conformal invariance is broken.

The purpose of this paper is to investigate how general are the lessons learnt from the study of the SYK models for gravitational systems. The investigation we report on is a continuation of the analysis begun in [37] in which a simple system, the Reissner-Nordström black hole in AdS_4 , was analysed close to extremality in the S -wave sector and its dynamics at low-energies was found to agree with the JT model. More precisely, the low energy response of the extremal black hole or near-extremal black hole was found to arise from the near-horizon region and shown to be well approximated by the JT model. The role of the dilaton in this model is played by the radius of the S^2 transverse to the AdS_2 directions and it was found that the AdS_2 region can *in effect* be taken to have a boundary, located at a fixed value of this radius. Fluctuations of the boundary, due to fluctuations in the radius, then gives rise to the time reparametrisation mode mentioned above which dominates the low energy behaviour.

This paper extends the analysis in [37] in important ways. By considering charged matter we show that an additional phase mode is generically present with an action of the form given in [16] which is suppressed by the scale characterising departures from

the near horizon AdS_2 geometry. By studying the correlation functions of probe scalar fields we find that the exchange of this phase mode along with the time reparametrisation mode correctly gives rise to the low-energy response of the black hole. Some of the key formulas containing the time reparametrisation and phase modes along with their coupling to matter fields (eqs. (5.24), (5.25), (5.26)) and their finite temp counter-parts (eqs. (5.27), (5.28), (5.29), (5.30)) are contained in section 5.

We also go beyond the S -wave sector. The resulting dynamics can be conveniently analysed in the two dimensional theory obtained by carrying out a KK reduction over the S^2 . We show that additional phase modes arise from massless gauge fields which are tied to isometries of the internal space. In addition, massive fields also arise from the KK reduction. The effects of these massive states in correlation functions of probe fields can be calculated in AdS_2 space with the breaking of conformal invariance being neglected. While the S -wave sector and additional massless gauge fields dominate at low energies, the cumulative effect of the massive states, which arise from the higher partial waves, can also be significant at somewhat higher energies.

The lessons we learn in asymptotic AdS_4 space can be generalised to other situations, including higher dimensions and asymptotically flat space. This allows us to formulate a general method for calculating the low energy behaviour of extremal black holes whose near-horizon geometry is of the form $AdS_2 \times S^d$.

This paper is structured as follows. Having introduced the charged black hole system in AdS_4 and reviewed some background in section 2, we then analyse the effects of an intermediate scalar being exchanged in correlation functions in section 3. The effects of a probe gauge field, not excited in the background, are studied in section 4 and found to be governed, at leading order, by the dynamics of a phase mode. In section 5 we study a probe charged scalar field in the presence of the gravity and a background gauge field. These results are put together in section 5 and 6 which discuss how to calculate the response of an extremal black hole at low energies in general and extends this analysis to other situations, including higher dimensions, black holes in asymptotically flat space and multiply charged black holes.

Appendices A, B and C contain useful supplementary details.

It is worth drawing attention to some important references before we proceed further. See [36–70] for some recent discussion on AdS_2/CFT_1 correspondence and extremal black holes.

2 Spherically symmetric Reissner-Nordström black holes

Let us describe the geometry of a charged black hole in asymptotically AdS_4 space time which arises in a theory of Einstein gravity with a negative cosmological constant coupled to a Maxwell field. In this paper we take the black hole to be electrically charged. The action is given by

$$S = \frac{1}{16\pi G_N} \int d^4x \sqrt{-g} (R - 2\Lambda) - \frac{1}{4G_N} \int d^4x \sqrt{-g} F_{\mu\nu} F^{\mu\nu}. \quad (2.1)$$

The spherical Reissner-Nordström black hole solution is given by

$$ds^2 = -a(r)^2 dt^2 + a(r)^{-2} dr^2 + b(r)^2 (d\theta^2 + \sin^2\theta d\varphi^2),$$

$$a(r)^2 = 1 - \frac{2G_N M}{r} + \frac{4\pi Q^2}{r^2} + \frac{r^2}{L^2}, \quad b(r)^2 = r^2, \quad (2.2)$$

$$F_{rt} = \frac{Q}{r^2}. \quad (2.3)$$

Here M and Q are the mass and charge of the black hole, and L is the AdS₄ radius,

$$L = \sqrt{\frac{3}{|\Lambda|}}. \quad (2.4)$$

For an extremal black hole, the charge and mass get related as

$$Q^2 = \frac{1}{4\pi} \left(r_h^2 + \frac{3r_h^4}{L^2} \right), \quad (2.5)$$

$$M_{ext} = \frac{r_h}{G_N} \left(1 + \frac{2r_h^2}{L^2} \right). \quad (2.6)$$

With these values of M and Q , eq. (2.2) becomes

$$a^2(r) = \frac{(r - r_h)^2}{r^2 L^2} (L^2 + 3r_h^2 + 2rr_h + r^2). \quad (2.7)$$

In the first few sections of this paper we will primarily be interested in a big black hole whose horizon size is much bigger than the AdS radius,

$$r_h \gg L. \quad (2.8)$$

For such big black holes, the charge and mass at extremality can be approximated as

$$Q^2 \simeq \frac{1}{4\pi} \left(\frac{3r_h^4}{L^2} \right), \quad (2.9)$$

$$M_{ext} \simeq \left(\frac{2r_h^3}{G_N L^2} \right). \quad (2.10)$$

Upto $O(\frac{r-r_h}{r_h})$ corrections, the metric in the near-horizon region of the black hole is given by

$$ds^2 = \left[-\frac{(r - r_h)^2}{L_2^2} dt^2 + \frac{L_2^2}{(r - r_h)^2} dr^2 + r_h^2 (d\theta^2 + \sin^2\theta d\varphi^2) \right]. \quad (2.11)$$

We see that the near horizon metric is of the form $AdS_2 \times S^2$, where AdS_2 and S^2 radii are given by

$$R_{AdS_2} = L_2 \simeq \frac{L}{\sqrt{6}}, \quad (2.12)$$

$$R_{S^2} = r_h. \quad (2.13)$$

The metric eq. (2.11) is a good approximation when

$$\frac{r - r_h}{r_h} \ll 1. \quad (2.14)$$

The extremal black hole corresponds to a state at chemical potential μ in the boundary CFT_3 dual to the AdS_4 , with

$$\mu \sim \frac{r_h}{L^2}. \tag{2.15}$$

The entropy of the extremal black hole is

$$S_{ext} = \frac{\pi r_h^2}{G_N}. \tag{2.16}$$

A near-extremal black hole has a temperature $T \ll \mu$. Its entropy is given by,

$$S - S_{ext} = \frac{4\pi^2}{G_N} L_2^2 r_h T, \tag{2.17}$$

and the free energy is

$$F = -\frac{1}{\beta} \frac{\pi r_h^2}{G_N} - \frac{2\pi^2}{G_N} L_2^2 r_h T^2. \tag{2.18}$$

We see that the specific heat is linear near extremality.

It is known that the thermodynamical description breaks down very close to extremality; this has been often found to agree with the existence of a gap in the system. The break down occurs when the specific heat is order unity, leading to the condition,

$$T_{gap} \sim \frac{G_N}{L_2^2 r_h}. \tag{2.19}$$

We are interested in the low energy behaviour of the system which corresponds to being at a temperature $T \ll \mu$ and to studying the response at energies¹ $\omega \ll \mu$. Using eq. (2.15) these conditions become,

$$T \ll \frac{r_h}{L_2^2} \tag{2.20}$$

and

$$\omega \ll \frac{r_h}{L_2^2}. \tag{2.21}$$

We will also take $T \gg T_{gap}$.

Corrections to the $AdS_2 \times S^2$ geometry, especially the change in the radius of the S^2 , are important to keep in analysing the low energy behaviour, [36, 37]. The response of the system can be understood as arising from the region obtained by cutting off the near horizon geometry along a boundary which corresponds to a constant value of the dilaton. This boundary will lie in the asymptotic AdS_2 region. Working in Poincaré coordinates,

$$z = \frac{L_2^2}{(r - r_h)}, \tag{2.22}$$

the metric of AdS_2 can be written as

$$ds^2 = \frac{L_2^2}{z^2} (dt^2 + dz^2). \tag{2.23}$$

¹More generally, see eq. (3.2).

The boundary lies in the region where

$$\omega z \ll mL_2 \sim O(1). \tag{2.24}$$

From eq. (2.22) we see that this implies

$$\omega L_2^2 \ll (r - r_h) \tag{2.25}$$

which is consistent with eq. (2.14) since eq. (2.21) is being met.

It is also useful to briefly discuss the asymptotically flat case (which corresponds to setting $\Lambda = 0$ in eq. (2.1)). The behaviour of small black holes, with $r_h \ll L$ in AdS_4 is similar. The solution for the extremal RN black hole in asymptotically flat space is given by

$$ds^2 = \left(1 - \frac{\sqrt{4\pi} Q}{r}\right)^2 dt^2 + \left(1 - \frac{\sqrt{4\pi} Q}{r}\right)^{-2} dr^2 + r^2 d\Omega_2^2. \tag{2.26}$$

The radius of the S^2 at the horizon is $r_h = \sqrt{4\pi} Q$. The near horizon geometry is $AdS_2 \times S^2$ with both the radius of the S^2 and AdS_2 being r_h , i.e.,

$$L_2 = r_h. \tag{2.27}$$

This is in contrast with big black holes in AdS_4 ; these have a value of L_2 given in eq. (2.12) which does not depend on the charge Q .

From the field strength, working in the gauge $A_r = 0$, we get that

$$A_t = \frac{Q}{r_h} - \frac{Q}{r}. \tag{2.28}$$

The chemical potential of the black hole μ is the value of A_t at $r = \infty$; we see that $\mu = \frac{Q}{r_h}$ and does not grow with the charge Q since the radius of the black hole is also growing linearly with Q .² This is in contrast to big black holes in AdS_4 which we have discussed above for which $Q \sim \frac{r_h^2}{L}$ and $\mu \sim \frac{r_h}{L^2}$, so that the chemical potential does grow with the charge.

Despite these differences some of the key features of the analysis for big black holes in AdS_4 can be easily extended and apply to black holes in asymptotically flat space as well. Corrections to the near horizon geometry are small as long as eq. (2.14) is met. It can be easily seen that the condition for low energy and small temperature, so that the response arises from the near horizon region, still takes the form, eq. (2.20), eq. (2.21). From eq. (2.27) these become,

$$T \ll \frac{1}{r_h}, \tag{2.29}$$

$$\omega \ll \frac{1}{r_h}. \tag{2.30}$$

A similar analysis, for both black holes in asymptotically AdS and flat space, also applies in higher dimensions.

²Note from eq. (2.1) that in our units $F_{\mu\nu}$ is of dimension 1 so that the chemical potential is dimensionless.

3 Four point function with scalar exchange in the bulk

In this section we study a probe scalar ψ which is coupled to another scalar ϕ .

The exchange of this ϕ particle then gives rise to a four-point function for ψ . We will analyse the resulting four-point function in a partial wave expansion. Our conclusions will also generalise to contributions to the four-point function due to graviton or gauge field exchange.

We work in Euclidean space in this section. The matter action is

$$S_M = \frac{1}{2G_N} \left(\int d^4x \sqrt{g} (\partial_\mu \psi \partial^\mu \psi + m_\psi^2 \psi^2) + \int d^4x \sqrt{g} (\partial_\mu \phi \partial^\mu \phi + m_\phi^2 \phi^2) + \lambda \int d^4x \sqrt{g} \phi \psi \psi \right). \quad (3.1)$$

Here we have coupled the probe ψ by a cubic vertex to the scalar ϕ .

At low energies, eq. (2.21), the correlation functions for the operator dual to ψ arise from the near-horizon AdS_2 region, as discussed in section 3 of [37]. A sufficient condition is

$$\omega \ll m \frac{r_h}{L_2}. \quad (3.2)$$

Setting $m \sim O(\frac{1}{L_2})$ gives, eq. (2.21).

The metric of $AdS_2 \times S^2$ in Poincare coordinates is given by

$$ds^2 = \frac{L_2^2}{z^2} (dt^2 + dz^2) + r_h^2 d\Omega^2. \quad (3.3)$$

Expanding ψ in terms of the KK harmonics on the S^2 we get,

$$\psi = \sum_{lm} \psi_{lm}(z, t) Y_{lm}(\theta, \phi). \quad (3.4)$$

The fields $\psi_{lm}(z, t)$ are scalars in AdS_2 with mass

$$m_{\psi,l}^2 = m_\psi^2 + \frac{l(l+1)}{r_h^2}. \quad (3.5)$$

Using the standard dictionary it is easy to see that they correspond to operators with dimension

$$\Delta_+^{(l)} = \frac{1}{2} + \sqrt{\frac{1}{4} + m_{\psi,l}^2 L_2^2}. \quad (3.6)$$

Similarly for ϕ we get the KK modes $\phi_{lm}(z, t)$ with mass

$$m_{\phi,l}^2 = m_\phi^2 + \frac{l(l+1)}{r_h^2}. \quad (3.7)$$

Substituting eq. (3.4) in the action above eq. (3.1) gives,

$$S_{\text{matter}} = \frac{r_h^2}{G_N} \sum_{lm} \int d^2x \sqrt{g} \frac{1}{2} \left(|\partial \psi_{lm}|^2 + m_{\psi,l}^2 |\psi_{lm}|^2 + |\partial \phi_{lm}|^2 + m_{\phi,l}^2 |\phi_{lm}|^2 \right) + \frac{r_h^2}{G_N} \lambda \left(\sum_{l,l',l'';m,m',m''} C_{\{l,l',l''\}}^{\{m,m',m''\}} \int d^2x \sqrt{g} \phi_{l''m''} \psi_{l,m} \psi_{l'm'} \right), \quad (3.8)$$

where we have integrated over the S^2 and the metric in the expression is the 2-D metric. In the expression above,

$$C_{\{l,l',l''\}}^{\{m,m',m''\}} = \int d^2\Omega Y_{lm} Y_{l'm'} Y_{l''m''}. \quad (3.9)$$

Let us understand the contribution made by the various partial waves in more detail.

The value of ψ_{lm} in the AdS_2 bulk can be expressed in terms of the value it takes on the AdS_2 boundary located at

$$z = \delta \quad (3.10)$$

using the bulk to boundary propagator in AdS_2 ,

$$\psi_{lm}(z, t) = \int d\omega \hat{K}_{\nu_l}(\omega z) e^{-i\omega t} \psi_{lm}(\omega). \quad (3.11)$$

Here $\psi_{lm}(\omega)$, the boundary value for frequency ω , corresponds to the source of the dual operator turned on at the AdS_2 boundary. In eq. (3.11) \hat{K}_{ν} can be expressed in terms of the modified Bessel function $K_{\nu}(\omega z)$ as,

$$\hat{K}_{\nu_l}(\omega z) = \delta^{\Delta_{\pm}^{(l)}} \frac{\sqrt{\omega z} K_{\nu_l}(\omega z)}{\sqrt{\omega \delta} K_{\nu_l}(\omega \delta)} \quad (3.12)$$

with

$$\nu_l = \sqrt{\frac{1}{4} + m_{\psi,l}^2 L_2^2}, \quad (3.13)$$

and,

$$\Delta_{\pm}^{(l)} = \frac{1}{2} \pm \nu_l. \quad (3.14)$$

The bulk Green function for ϕ_{lm} satisfies the equation

$$[\nabla^2 - m_{\phi,l}^2] G_{lm} = \frac{\delta^2(x - x')}{\sqrt{g_2}}. \quad (3.15)$$

In frequency space, this gives

$$\left(\partial_z^2 - \omega^2 - \frac{m_{\phi,l}^2 L_2^2}{z^2} \right) G_{lm}(\omega, z, z') = \delta(z - z'). \quad (3.16)$$

The resulting on-shell action is then given by

$$S_{\text{onshell}} = -\frac{\lambda^2}{2} \frac{r_h^2}{G_N} \sum_{l_i, m_i} \left[\int \left(\prod_{i=1}^{i=4} \frac{d\omega_i}{2\pi} \psi_{l_i m_i}(\omega_i) \right) \delta\left(\sum \omega_i\right) \sum_{l, m} C_{\{l, l_1, l_2\}}^{\{m, m_1, m_2\}} C_{\{l, l_3, l_4\}}^{\{-m, m_3, m_4\}} \right. \\ \left. \int dz \frac{L_2^2}{z^2} \int dz' \frac{L_2^2}{z'^2} \hat{K}_{\nu_1}(\omega_1 z) \hat{K}_{\nu_2}(\omega_2 z) \hat{K}_{\nu_3}(\omega_3 z') \hat{K}_{\nu_4}(\omega_4 z') G_{lm}(\omega_1 + \omega_2, z, z') \right]. \quad (3.17)$$

This has four powers of the source $\psi_{lm}(\omega)$ and thus is a contribution to the 4-point function.

To understand this complicated looking result, it is useful to recall the scaling dimensions of various fields. The quadratic (free) part of the action gives the two-point function for ψ , which is given by

$$S_{2\text{pt}} = \frac{r_h^2}{G_N} C_1 \sum_{lm} \int d\omega \psi_{l,m} \psi_{l,-m} \omega^{2\Delta_l}, \quad (3.18)$$

where C_1 is a constant and $2\Delta_l = \Delta_+^{(l)} - \Delta_-^{(l)}$. From this it follows that under the scaling $\omega \rightarrow \lambda\omega$, $\psi_{lm}(\omega) \rightarrow \lambda^{-(\Delta_l + \frac{1}{2})}$, i.e. $\psi_{lm}(\omega)$ has dimensions $-(\Delta_l + \frac{1}{2})$. It also follows that the operator \mathcal{O}_{lm} for which ψ_{lm} is the source has dimension $(\Delta_l - \frac{1}{2})$ so that

$$S_{\text{coupling}} = \int d\omega \psi_{lm}(\omega) \mathcal{O}_{lm}(\omega) \quad (3.19)$$

is invariant.

Now consider a contribution to the four-point function of the form

$$S_{4\text{pt}} = \prod_i \int d\omega_i \psi_{l_i m_i}(\omega_i) \langle \mathcal{O}_{l_1 m_1} \mathcal{O}_{l_2 m_2} \mathcal{O}_{l_3 m_3} \mathcal{O}_{l_4 m_4} \rangle. \quad (3.20)$$

By the scaling symmetry of AdS_2 we see that

$$\langle \mathcal{O}_{l_1 m_1} \mathcal{O}_{l_2 m_2} \mathcal{O}_{l_3 m_3} \mathcal{O}_{l_4 m_4} \rangle \sim \omega^{\sum_i \Delta_i - 2} \quad (3.21)$$

where $\Delta_i = \Delta_{l_i}$. Just from the scaling properties it therefore follows that the operators with higher values of l , which have bigger anomalous dimensions, eq. (3.5), will make a smaller contribution at low frequencies. By comparison with eq. (3.17) we see that

$$\begin{aligned} \langle \mathcal{O}_{l_1 m_1} \mathcal{O}_{l_2 m_2} \mathcal{O}_{l_3 m_3} \mathcal{O}_{l_4 m_4} \rangle &= -\frac{\lambda^2}{2} \frac{r_h^2}{G_N} \delta\left(\sum_i \omega_i\right) \left[\sum_l C_{\{l, l_1 l_2\}}^{\{m, m_1, m_2\}} C_{\{l, l_3 l_4\}}^{\{-m, m_3, m_4\}} \right. \\ &\quad \left. \int dz \frac{L_2^2}{z^2} \int dz' \frac{L_2^2}{z'^2} \hat{K}_{\nu_1}(\omega_1 z) \hat{K}_{\nu_2}(\omega_2 z) \hat{K}_{\nu_3}(\omega_3 z') \hat{K}_{\nu_4}(\omega_4 z') G_{lm}(\omega_1 + \omega_2, z, z') \right] \end{aligned} \quad (3.22)$$

The sum over l arises from the contribution of the different intermediate partial waves of the ϕ field. If the $\psi_{l_i, m_i}(\omega)$ are comparable for the different l_i, m_i values, the S -wave in eq. (3.17) will dominate at sufficiently low energies and in this case the sum over l in eq. (3.22) will only get a contribution from the S -wave for the intermediate field ϕ . From the behaviour of the Bessel functions K_ν and the Green function G_{lm} which appear in eq. (3.22), it is easy to see directly that the four-point function scales with frequency as given in eq. (3.21).

An important fact to note from eq. (3.17), eq. (3.22) is that the 4-point function due to the intermediate scalar ϕ being exchanged is $O(r_h^2)$. In contrast, it was found in [37] that S wave contribution due to gravitational effects goes like $O(r_h^3)$ and is further enhanced by a factor of r_h . We will return to this point in section 5.

4 Charged matter coupled to a probe gauge field

Next, we consider the behaviour of a probe gauge field which is not turned on in the background, i.e. with respect to which the black hole is not charged. The action is

$$S_{gf} = -\frac{1}{G_N} \int d^4x \sqrt{g} \left(\frac{1}{4} F^2 + e A_\mu J^\mu \right), \quad (4.1)$$

where e is the charge of the scalar field. The current J^μ is quadratic in the charged matter field. For a charged scalar ψ ,

$$J_\mu = i(\psi \partial_\mu \psi^\dagger - \psi^\dagger \partial_\mu \psi). \quad (4.2)$$

For concreteness, we are interested in calculating the four-point function for ψ with an intermediate exchange of the probe gauge field. In particular, we are interested in studying the S -wave sector of this theory.

It was noted in [16] that an additional mode arises in the SYK model with charged matter. The departure from conformal invariance is important for the dynamics of this mode whose action is suppressed by the energy scale parametrising this departure. The behaviour of this mode is therefore analogous to the time reparametrisation modes which arise in the S -wave sector of gravity.

We will see below that in fact a similar mode does arise for a U(1) gauge field in the bulk in the S -wave sector. The exchange of this mode, which we will refer to as the “phase mode”, will dominate the low-energy contribution to the four-point function due to gauge field exchange. Let us note that such a mode was found earlier in AdS_3 gravity with a Chern-Simons action for the gauge field in [66].

The S wave sector can be analysed in 4 dimensions directly. We work in the gauge $A_t = 0$. The source J^μ is spherically symmetric and therefore only gives rise to a non-trivial value for A_r ; we can therefore also set $A_\theta, A_\phi = 0$. The equation of motion is

$$\partial_\beta (\sqrt{g} F^{\beta\alpha}) = e \sqrt{g} J^\alpha \quad (4.3)$$

with J^α being nonzero for $\alpha = r, t$. Here we are using the coordinates in eq. (2.2). Setting $\alpha = r$ gives,

$$F^{tr} = e \partial_t^{-1} J^r. \quad (4.4)$$

The reader should not be alarmed by the factor of ∂_t^{-1} . We are considering time dependent sources here and eq. (4.3) can be solved for each frequency, ω , separately with ∂_t^{-1} simply being short hand for a factor of $\frac{1}{\omega}$. Eq. (4.4) allows for an extra integration constant which is only a function of r , this has been set to vanish since we are computing the response to the time dependent source. Eq. (4.4) can be solved for A_r to give

$$A_r = e g_{tt} \partial_t^{-2} J_r. \quad (4.5)$$

Substituting in eq. (4.1) gives the on-shell action

$$S_{\text{onshell}} = -\frac{1}{G_N} 4\pi e^2 \int d^2x \sqrt{g_2} r^2 \left(\frac{1}{2} g^{rr} g_{tt} J_r \partial_t^{-2} J_r \right). \quad (4.6)$$

As was noted above, the current J_r is quadratic in the charged matter field. Thus eq. (4.6) is a contribution to the four-point function for the matter field. In eq. (4.6) we have already integrated over the S^2 and the remaining two dimensional integral left is over r, t , with the metric $g_{\alpha\beta}$ being the $r - t$ plane metric, and g_2 denoting its determinant.

At first sight it might seem that the resulting contribution to the four-point function is of $O(r_h^2)$ due to the factor of r^2 inside the integral in eq. (4.6) and not of $O(r_h^3)$ as in the S -wave sector of gravity. However, this conclusion is premature, as we now discuss.

The charged field ψ we are considering has the action

$$S_{\text{matter}} = \frac{1}{G_N} \int d^4x \sqrt{g} (|D\psi|^2 + m^2|\psi|^2), \quad (4.7)$$

$$D_\mu = \partial_\mu - i e A_\mu, \quad (4.8)$$

with mass m .

Now consider the asymptotic AdS_2 region. In this region ψ takes the form (see appendix A),

$$\psi(z, t) = \psi(t)z^{\Delta_-} + c(z^{\Delta_+} - z^{\Delta_-} \delta^{\Delta_+ - \Delta_-}) \int dt' \frac{\psi(t')}{|t - t'|^{2\Delta_+}}, \quad (4.9)$$

where c is a constant, eq. (A.5), and Δ_\pm are given by

$$\Delta_\pm = \frac{1}{2} \pm \sqrt{\frac{1}{4} + m^2 L_2^2}. \quad (4.10)$$

Working in the z coordinate, eq. (2.22), the contribution to the integral in eq. (4.6), is then given by,

$$\Delta S_{\text{onshell}} = -\frac{r_h^2}{G_N} 2\pi e^2 \int dt dz \sqrt{g} g_{tt} g^{zz} J_z(z, t) \partial_t^{-2} J_z(z, t). \quad (4.11)$$

Discarding contact terms it is easy to see from eq. (A.10) that

$$J_z(t, z) = i c (\Delta_+ - \Delta_-) \int dt' \left(\frac{\psi(t)\psi^\dagger(t') - \psi(t')\psi^\dagger(t)}{|t - t'|^{2\Delta_+}} \right). \quad (4.12)$$

In particular J_z is independent of z . Thus the z dependence in eq. (4.11) goes like

$$\Delta S_{\text{onshell}} \sim \int dz \frac{1}{z^2} \sim \frac{1}{z_c}, \quad (4.13)$$

where z_c is the value at the boundary of the asymptotic AdS_2 region.

One expects of course that this divergence is cut off once the corrections to the AdS_2 geometry become significant. Since this occurs at a scale,

$$z \sim \frac{1}{r_h}, \quad (4.14)$$

(see eqs. (2.14) and (2.22)) we get by setting $z_c \sim \frac{1}{r_h}$ in eq. (4.13) the resulting four-point function to be of order,

$$S_{4\text{pt}} \sim r_h^3 \int dt dt' dt'' \left(\frac{\psi(t)\psi^\dagger(t'') - \psi(t'')\psi^\dagger(t)}{|t - t''|^{2\Delta_+}} \right) \partial_t^{-2} \left(\frac{\psi(t)\psi^\dagger(t') - \psi(t')\psi^\dagger(t)}{|t - t'|^{2\Delta_+}} \right). \quad (4.15)$$

Note that due to the divergence the resulting 4 point function is in fact enhanced and of order r_h^3 , as expected in analogy with the SYK model. In appendix A, we calculate the coefficient in eq. (4.15) precisely. The resulting 4 point function is then given by

$$S_{4\text{pt}} = \left(\frac{r_h^2}{G_N}\right) r_h 2\pi c^2 e^2 (\Delta_+ - \Delta_-)^2 \times \int dt dt' dt'' \left(\frac{\psi(t)\psi^\dagger(t'') - \psi(t'')\psi^\dagger(t)}{|t - t''|^{2\Delta_+}}\right) \frac{1}{\partial_t^2} \left(\frac{\psi(t)\psi^\dagger(t') - \psi(t')\psi^\dagger(t)}{|t - t'|^{2\Delta_+}}\right). \quad (4.16)$$

We have focussed on the asymptotic AdS_2 region above. It is easy to see that the interior of the AdS_2 region gives a contribution to the 4 point function which is $O(r_h^2)$ and is therefore less important.

4.1 The phase mode

The 4-point function we have obtained above can be understood as arising due to the exchange of a mode with an action suppressed by $\frac{1}{r_h}$. Working in the $A_t = 0$ gauge we have chosen above, take A_r (in the (t, r) coordinate system) to be of the form³

$$A_r = \chi(r) \theta(t) \quad (4.17)$$

with

$$\chi(r) = \frac{r_h}{\sqrt{g} g^{tt} g^{rr}}. \quad (4.18)$$

Here the metric components refer to the extremal geometry given in eq. (2.2) and $\sqrt{g} = r^2$ is the determinant of the four dimensional metric.

With the form of A_r in eq. (4.18), the Maxwell action

$$-\frac{1}{4G_N} \int d^4x \sqrt{g} F^2 \quad (4.19)$$

gives

$$S_{\text{phase}} = -\left(\frac{r_h^2}{G_N}\right) 2\pi \int d^2x \frac{1}{\sqrt{g} g^{rr} g^{tt}} (\dot{\theta})^2, \quad (4.20)$$

where we have integrated over the S^2 . The upper limit of the r integral is the AdS_4 boundary. The lower limit of the integral can be taken to be the AdS_2 horizon (up to corrections subleading in the frequency ω , as discussed in the appendix A). This gives,

$$S_{\text{phase}} = -\left(\frac{r_h^2}{G_N}\right) 2\pi \frac{1}{r_h} \int dt (\dot{\theta})^2. \quad (4.21)$$

We will see below that $\theta(t)$ plays the role of the phase mode mentioned above. For now we note that its action is indeed suppressed by the scale r_h which is related to the energy scale $\mu \sim \frac{r_h}{L^2}$ at which departures from the AdS_2 geometry become significant.

Let us see how this mode enters in the thermodynamic properties. Consider a black hole in AdS_2 which now also carries charge q with respect to this probe field. As discussed

³We are grateful to Juan Maldacena for a discussion on this topic.

in appendix C, to calculate the partition function for the grand canonical ensemble, the action has an extra term compared to the canonical ensemble [71]. This results in an extra contribution in the free energy of the grand canonical ensemble:

$$\Delta F = -\frac{4\pi}{G_N} \frac{q^2}{r_h}. \quad (4.22)$$

This additional contribution is given by twice the contribution of the phase mode S_{phase} if we identify

$$q = r_h \dot{\theta}. \quad (4.23)$$

Note that eq. (4.23) solves the equation of motion for θ (we are considering the system without sources here):

$$\ddot{\theta} = 0. \quad (4.24)$$

Next let us see how the exchange of this mode gives rise to the $O(r_h^3)$ contribution to the 4 point function we obtained above. The charged matter field ψ , eq. (4.7), has a two-point function which arises from an action given by

$$S_{2\text{pt}} = -\left(\frac{r_h^2}{G_N}\right) 4\pi c(\Delta_+ - \Delta_-) \int dt dt' \frac{\psi^\dagger(t)\psi(t')}{|t-t'|^{2\Delta_+}}. \quad (4.25)$$

The action is evaluated at the boundary of the near AdS_2 region, with $\psi(t)$ being the boundary value for ψ . The field $\theta(t)$ is to be identified with the phase of ψ . This results in a coupling (at linear order in θ):

$$S_c = \left(\frac{r_h^2}{G_N}\right) 4\pi c(\Delta_+ - \Delta_-) e i \int dt dt' \frac{\psi^\dagger(t)\psi(t')}{|t-t'|^{2\Delta_+}} (\theta(t) - \theta(t')). \quad (4.26)$$

Integrating out θ using eq. (4.21) we now get a 4-pt function which agrees with eq. (A.18). Higher point functions, to leading order in $\frac{r_h^2}{G_N}$, arise from successive tree level exchanges of the phase mode, and can be calculated in a similar fashion.

4.2 Additional comments

It is worth noting that even though the radial integral involved in obtaining the full $O(r_h^3)$ term got its contribution from the region extending beyond the near-horizon geometry, the final result for the 4-point function could be expressed in terms of the value for ψ at the boundary of the near-horizon region, eq. (4.9). This result is expected on physical grounds since the low-energy response should arise from the near-horizon region.

It is also worth mentioning that the phase mode does not correspond to a gauge transformation in the AdS_2 region, since A_r in eq. (4.17) gives rise to a non-trivial field strength when θ is time dependent. What the mode does capture is both the charge dependence in the grand canonical ensemble and the Coulomb dressing effects for charged perturbations in the S -wave. Roughly, this mode can be thought of as a Wilson line (smeared on the S^2) in the radial direction extending from the horizon to the boundary of AdS_2 .

We have concentrated on the S -wave sector so far. The higher partial waves of the probe gauge field will give rise to massive gauge fields in the two dimensional theory obtained by dimensional reduction to AdS_2 . In addition, scalars will also arise from the KK reduction to two dimensions. These extra states will then also contribute as intermediate channels in the four-point function. To leading order their contribution can be calculated in the AdS_2 geometry itself, neglecting any departures, in a manner analogous to the effect of the intermediate scalar in section 3.

Expressing the 4 pt. function in Fourier space,

$$S_{4\text{pt}} = \left(\prod_{i=1}^4 \int \frac{d\omega_i}{2\pi} \right) \psi(\omega_1)\psi(\omega_2)\psi^\dagger(\omega_3)\psi^\dagger(\omega_4) \langle O(\omega_1)O(\omega_2)O(\omega_3)^\dagger O(\omega_4)^\dagger \rangle, \quad (4.27)$$

it is easy to see from eq. (4.16) that the S -wave contribution to $\langle O(\omega_1)O(\omega_2)O(\omega_3)^\dagger O(\omega_4)^\dagger \rangle$ will scale with frequency as⁴

$$\langle O(\omega_1)O(\omega_2)O(\omega_3)^\dagger O(\omega_4)^\dagger \rangle \sim e^2 \left(\frac{r_h^2}{G_N} \right) r_h \omega^{4\Delta-3}. \quad (4.28)$$

This might seem a bit surprising at first, since the anomalous dimension in AdS_2 for O would result in a scaling, see eq. (3.21), $\langle OOO^\dagger O^\dagger \rangle \sim \omega^{4\Delta-2}$.

The situation here is similar to what happens with the Schwarzian action for time reparametrisation modes. The action for the phase mode, eq. (4.21), breaks the scaling symmetry of AdS_2 under which $t \rightarrow \lambda t$ (with $\theta(t) \rightarrow \theta(\lambda t)$). The breaking is due to the factor of $\frac{1}{r_h}$ which suppresses the action; its presence we saw above is tied to the fact that the action arises not in the near AdS_2 region but in the region which interpolates from the asymptotic AdS_2 geometry to the AdS_4 boundary.

A convenient way of book keeping is to assign the factor of r_h which suppresses the action in eq. (4.21) a transformation law $r_h \rightarrow \frac{r_h}{\lambda}$ to keep track of this breaking. It is important to bear in mind that we do this while keeping the prefactor $\frac{r_h^2}{G_N}$ in eq. (4.21) invariant under the rescaling. The point being that $\frac{r_h^2}{G_N} \sim S_{ext}$ is the extremal entropy and therefore a measure of the degrees of freedom in the IR AdS_2 , while the extra factor of $\frac{1}{r_h}$ is due to the breaking of conformal invariance. It is easy to see that eq. (4.21) is invariant with this extra rescaling of r_h . It also then follows that the 4 pt. function scales as given in eq. (4.28) since it has an extra factor of r_h in front, eq. (4.16).

Consider now higher partial waves of the ψ field. For suitable values of their angular momentum the contribution of the S wave phase mode could still arise. For any mode ψ_{lm} with angular momentum l , the two-point function can be written in the form, eq. (4.25) with Δ_l being the anomalous dimension for the mode. This would then lead to a coupling with the phase mode analogous to eq. (4.26), resulting in a 4 pt function due to the exchange of the phase mode. In frequency space this would lead to

$$\langle O_l(\omega_1)O_l(\omega_2)O_{l'}(\omega_3)^\dagger O_{l'}(\omega_4)^\dagger \rangle \sim e^2 \left(\frac{r_h^2}{G_N} \right) r_h \omega^{2\Delta_l+2\Delta_{l'}-3}. \quad (4.29)$$

⁴The engineering dimensions are accounted for by suitable powers of L_2 .

As mentioned above, for higher partial waves of ψ , contributions due to the exchange of other intermediates would also arise. The breaking of conformal invariance can be neglected for these contributions leading to the behaviour

$$\langle O_{l_1}(\omega_1)O_{l_2}(\omega_2)O_{l_3}(\omega_3)^\dagger O_{l_4}(\omega_4)^\dagger \rangle \sim e^2 \left(\frac{r_h^2}{G_N} \right) \omega^{\sum_i \Delta_i - 2} \quad (4.30)$$

By comparing, eq. (4.29) and eq. (4.30) we see that S -wave exchange dominates, when it is allowed, at sufficiently low frequencies, meeting the condition,

$$\omega \ll \frac{r_h}{L_2^2}, \quad (4.31)$$

where we have inserted factors of L_2 as required by dimensional analysis.

This is the same condition as eq. (2.21).

In general, several higher partial waves can contribute in the intermediate channel. As a result, there can be an intermediate range of frequencies where the cumulative contribution of the higher partial wave exchange is significant. This happens if about⁵

$$n \sim \frac{r_h}{L_2^2 \omega} \quad (4.32)$$

number of partial waves become important.

Let us examine if eq. (4.32) can be met while keeping the mass of the KK mode intermediate states m_{KK} to be no bigger than of order unity in units of L_2 , and therefore their anomalous dimensions to be of order unity at most. The discussion in this section generalises in a straightforward way to the $AdS_2 \times S^d$ case, see section 6. We might as well analyse this more general case. The number of KK modes with $m_{\text{KK}} \leq O(\frac{1}{L_2})$ is

$$n_{\text{KK}} \sim \left(\frac{r_h}{L_2} \right)^d. \quad (4.33)$$

Thus eq. (4.32) is met if

$$\omega r_h \left(\frac{r_h}{L_2} \right)^{d-2} \sim O(1). \quad (4.34)$$

From eq. (2.21) and eq. (4.34) we learn that for ω lying in the range

$$\frac{1}{r_h} \left(\frac{L_2}{r_h} \right)^{d-2} \ll \omega \ll \frac{r_h}{L_2^2}, \quad (4.35)$$

the effect of the higher partial waves will be more significant than the S -wave sector while for smaller ω the S wave sector will dominate. For $d = 2$, we see that eq. (4.35) can be met for a big black hole with $\frac{r_h}{L_2} \gg 1$, while for $d > 2$ it becomes even easier to meet the condition for such black holes.

We will also discuss black holes in asymptotically flat space later in section 6; their behaviour is similar to small black holes in AdS space. The formulae above (expressed as

⁵The higher partial waves are further suppressed because they correspond to operators with higher anomalous dimensions, as we saw in section 3. We neglect this effect in the estimate below.

given in terms of the parameters, r_h, L_2) are applicable here as well, with $r_h = L_2$. The condition eq. (2.21) for the near AdS_2 analysis to apply is more appropriately stated as eq. (3.2) for a probe of mass m and eq. (4.35) then takes the form

$$\frac{1}{r_h} \left(\frac{L_2}{r_h} \right)^{d-2} \ll \omega \ll \frac{mr_h}{L_2}. \tag{4.36}$$

Setting $L_2 = r_h$ this becomes

$$\frac{1}{r_h} \ll \omega \ll m. \tag{4.37}$$

This condition can also be met for a sufficiently big black hole.

5 Charged matter coupled to gravity and background gauge field

Here we consider the effects of gravity and the coupled gauge field, which is turned on in the background extremal solution, for calculating the four-point function of a scalar field charged under this gauge field. The system is given by the Euclidean version of the action, eq. (2.1), coupled to the scalar,

$$S_{\text{matter}} = \frac{1}{G_N} \int d^4x \sqrt{g} (|D\psi|^2 + m^2|\psi|^2). \tag{5.1}$$

We are interested in the behaviour at low frequencies, eq. (2.21).

We begin the analysis by first considering the S -wave sector. We will see that this sector dominates the behaviour at sufficiently low frequencies. This is due to both contributions from the time reparametrisation mode and a phase mode analogous to the one described above. Correlation functions for a field uncharged under the background gauge field were already discussed in [37].

5.1 The two point function

Let us first examine the behaviour of the scalar field in the near-horizon region. To leading order, it satisfies the equation,

$$\bar{D}^\mu \bar{D}_\mu \psi - m^2 \psi = 0, \tag{5.2}$$

where $\bar{D}_\mu = \partial_\mu - ie \bar{A}_\mu$, with \bar{A}_μ being the background gauge field, e being the charge of the probe scalar⁶ and the indices in eq. (5.2) being raised by the background metric. We work in the gauge where only the \bar{A}_t component of the background gauge field is non-zero. In the AdS_2 region it is given by

$$\bar{A}_t = \frac{L_2^2}{r_h^2} \frac{Q}{z}. \tag{5.3}$$

As a result, it is easy to see that the “effective mass” of ψ is

$$m_{\text{eff}}^2 = m^2 + \frac{e^2 Q^2 L_2^2}{r_h^4}. \tag{5.4}$$

⁶In our conventions the gauge potential A_μ is dimensionless and e has dimensions of $(length)^{-1}$.

Let us take the boundary of the AdS_2 region to be at $z = \delta$, where z is defined in eq. (2.22). Then from eq. (5.2), eq. (5.4), we see that the asymptotic behaviour as $z \rightarrow \delta$ is

$$\psi(t, z) = z^{\tilde{\Delta}_-} \psi(t) + c(z^{\tilde{\Delta}_+} - z^{\tilde{\Delta}_-} \delta^{\tilde{\Delta}_+ - \tilde{\Delta}_-}) \int dt' \frac{\psi(t')}{|t - t'|^{2\tilde{\Delta}_+}}, \quad (5.5)$$

with

$$\tilde{\Delta}_\pm = \frac{1}{2} \pm \sqrt{\frac{1}{4} + m^2 L_2^2 + \frac{e^2 Q^2 L_2^4}{r_h^4}}. \quad (5.6)$$

$\psi(t)$ is the value for the scalar field at the AdS_2 boundary.

It then follows that the two-point function for the scalar at low energies is given by

$$S_{2pt} = -\left(\frac{r_h^2}{G_N}\right) 4\pi c (\tilde{\Delta}_+ - \tilde{\Delta}_-) \int dt dt' \frac{\psi^\dagger(t) \psi(t')}{|t - t'|^{2\tilde{\Delta}_+}}. \quad (5.7)$$

We see that the scalar corresponds to an operator of dimension $\tilde{\Delta}_+$.

5.2 The four point function

We turn to the four-point function next. As mentioned above we are considering the S -wave sector here. To begin, we work directly in the 4 dimensional theory.

The scalar, once turned on, gives rise to a stress tensor and charged current which in turn perturb the metric and the gauge field, giving rise to a four-point function. As discussed in appendix B, the resulting on-shell action to the required order for the four-point function is,

$$\begin{aligned} S_{OS} = & -32\pi^2 G_N \int dt dr \left(\frac{2a^2 b^3}{b'} T_{rr} \frac{1}{\partial_t} T_{tr}^\psi - a^2 b^2 \left(1 + \frac{2a'b}{b'a}\right) T_{tr} \frac{1}{\partial_t^2} T_{tr}^\psi \right) \\ & - \frac{2\pi e^2}{G_N} \int dt dr b^2 a^4 J_r \partial_t^{-2} J_r - \frac{2\pi e}{G_N} \int dt dr b^2 a^2 \delta A_r J_r. \end{aligned} \quad (5.8)$$

Here $T_{\alpha\beta}^\psi$ is the contribution to the stress tensor from the scalar ψ , eq. (B.19), and $T_{\alpha\beta}$ has contributions from both the scalar as well as the gauge field perturbation, expanded to the required order for the four-point correlator, see eq. (B.3). The current J_r is quadratic in ψ , see eq. (4.2), and the gauge field perturbation δA_r is given in terms of the metric perturbations in eq. (B.11).

We have seen, both in the uncharged case, and the probe gauge field case above, that an $O(r_h^3)$ contribution is expected to arise in the four-point function. For the terms in the first line of eq. (5.8) which are quadratic in the stress tensor, such contributions do arise and they get support from the near horizon region. The near horizon contribution can be expressed as an integral on the boundary of the near AdS_2 , as discussed in eq. (B.24) and takes the form,

$$S_{bdy}^\psi = -\frac{32\pi^2 G_N r_h^3}{L_2^2} \int dt (\tilde{T}_{tz}^\psi - z \partial_t \tilde{T}_{zz}^\psi) \frac{1}{\partial_t^4} (\tilde{T}_{tz}^\psi - z \partial_t \tilde{T}_{zz}^\psi). \quad (5.9)$$

Using the fact that

$$\tilde{T}_{\alpha\beta}^{\psi} = \frac{1}{2G_N} \left(\partial_{\alpha}\psi^{\dagger}\partial_{\beta}\psi + \partial_{\beta}\psi^{\dagger}\partial_{\alpha}\psi - \bar{g}_{\alpha\beta} (\bar{g}^{\mu\nu}\partial_{\mu}\psi^{\dagger}\partial_{\nu}\psi + m^2\psi\psi^{\dagger}) \right), \quad (5.10)$$

and the asymptotic form for ψ , eq. (5.5), the expression eq. (5.9) can be expressed in terms of the value ψ takes at the boundary of the near AdS_2 region to give,

$$S_{\text{bdy}}^{\psi} = -32\pi^2 c^2 (\tilde{\Delta}_+ - \tilde{\Delta}_-)^2 \tilde{\Delta}_+^2 \frac{r_h^2}{G_N} \frac{r_h}{L_2^2} \quad (5.11)$$

$$\int dt dt' dt'' \left(\frac{1}{4} F_1(t, t') \frac{1}{\partial_t^4} F_1(t, t'') - \tilde{\Delta}_- F_1(t, t') \frac{1}{\partial_t^4} F_2(t, t'') + \tilde{\Delta}_-^2 F_2(t, t') \frac{1}{\partial_t^4} F_2(t, t'') \right),$$

$$F_1(t, t') = \frac{\partial_t \psi^{\dagger}(t) \psi(t') + \partial_{t'} \psi(t) \psi^{\dagger}(t')}{|t-t'|^{2\tilde{\Delta}_+}},$$

$$F_2(t, t') = \frac{1}{2\tilde{\Delta}_+} \left(\psi^{\dagger}(t) \psi(t') + \psi(t) \psi^{\dagger}(t') \right) \partial_t |t-t'|^{-2\tilde{\Delta}_+}. \quad (5.12)$$

From the first term in the second line of eq. (5.8), which is quadratic in the current J_r , we again get a contribution of $O(r_h^3)$. This can be analysed in a manner completely analogous to the probe gauge field case in section 4. The resulting answer takes the form in eq. (4.16) with Δ_{\pm} replaced by $\tilde{\Delta}_{\pm}$ and is given by

$$S_J = \left(\frac{r_h^2}{G_N} \right) r_h 2\pi c^2 e^2 (\tilde{\Delta}_+ - \tilde{\Delta}_-)^2$$

$$\times \int dt dt' dt'' \left(\frac{\psi(t)\psi^{\dagger}(t'') - \psi(t'')\psi^{\dagger}(t)}{|t-t''|^{2\tilde{\Delta}_+}} \right) \frac{1}{\partial_t^2} \left(\frac{\psi(t)\psi^{\dagger}(t') - \psi(t')\psi^{\dagger}(t)}{|t-t'|^{2\tilde{\Delta}_+}} \right). \quad (5.13)$$

Finally, it can be shown that the last term in eq. (5.8) involving the gauge field perturbation δA_r cancels with a corresponding term in the first line of eq. (5.8), see appendix B for details. Thus $S_{\text{bdy}}^{\psi} + S_J$ is the total answer for the four-point function in the S -wave sector.

5.3 The action for time reparametrisation and phase modes

It is useful at this stage to consider the two dimensional theory obtained by S -wave reduction of the 4 dimensional theory, eqs. (2.1) and (5.1).

Expanding the dilaton field Φ (defined in eq. (C.2) of appendix) to linear order,

$$\Phi = r_h(1 + \phi) \quad (5.14)$$

we get, from an analysis similar to section 5 of [37] that in the near-horizon region the two dimensional theory has an action⁷ (see appendix C)

$$S = -\frac{r_h^2}{G_N} \left(\frac{1}{4} \int d^2x \sqrt{g} R + \frac{1}{2} \int_{\partial} \sqrt{\gamma} K \right) - \frac{r_h^2}{G_N} \frac{2\pi}{r_h} \int dt \dot{\theta}^2$$

$$- \frac{r_h^2}{G_N} \pi \int d^2x \sqrt{g} (F^2 - F_Q^2) - \frac{r_h^2}{G_N} 3\pi \int d^2x \sqrt{g} \phi (F^2 - F_Q^2)$$

$$- \frac{r_h^2}{G_N} \frac{1}{2} \int d^2x \sqrt{g} \phi (R - \Lambda_2) - \frac{r_h^2}{G_N} \int_{\partial} \sqrt{\gamma} \phi K + \frac{r_h^2}{G_N} \frac{1}{L_2} \int_{\partial} \sqrt{\gamma} \phi. \quad (5.15)$$

⁷We work in a gauge where the perturbation δA_{α} satisfies the boundary condition $F^{\alpha\beta} \delta A_{\alpha} n_{\beta} = 0$, with n_{β} being the normal to the boundary.

Here we have added boundary terms and an extra boundary counterterm, [37]. $F_{\alpha\beta}^Q$ above is given by

$$F_{rt}^Q = \frac{Q}{r_h^2 \sqrt{g} g^{rr} g^{tt}}, \tag{5.16}$$

and $\dot{\theta}$ is given by eq. (C.12).

We see that the resulting two dimensional theory is closely related to the JT theory of gravity in the presence of an extra gauge field. In the absence of any sources the theory admits a solution with $\phi = 0$, and with the metric being AdS_2 , and the gauge field taking the form,

$$F_{rt} = F_{rt}^Q. \tag{5.17}$$

It is easy to verify that the resulting stress tensor of the gauge field cancels the stress tensor which arises from the F_Q term.

We can now examine the dynamics of the near-horizon region in this two dimensional theory. The dynamics includes fluctuations of the boundary of this region corresponding to time reparametrisations⁸ and on general arguments in [36, 37] leads to the Schwarzian action,

$$S_{\text{sch}} = -\frac{r_h^2}{G_N} \frac{L_2^2}{r_h} \int_{\text{bdy}} Sch(f(t)), \tag{5.18}$$

where

$$Sch(f(t)) = -\frac{1}{2} \frac{(f'')^2}{(f')^2} + \left(\frac{f''}{f'} \right)'. \tag{5.19}$$

Writing

$$f(t) = t + \epsilon(t), \tag{5.20}$$

and expanding to quadratic order in ϵ , eq. (5.18) becomes

$$S_\epsilon = -\frac{r_h^2}{G_N} \frac{L_2^2}{r_h} \int dt \left(-\frac{3}{2} (\epsilon'')^2 - \epsilon' \epsilon''' \right). \tag{5.21}$$

As noted in [36, 37], the coefficient in eq. (5.21) leads to the correct near-extremal free energy in the Canonical Ensemble.

Similarly, a phase mode corresponding to the gauge field perturbation

$$\delta A_r = \frac{1}{\sqrt{g_4} g^{rr} g^{tt}} \delta\theta(t), \tag{5.22}$$

must arise to account for the Coulomb “dressing” of perturbations. Analogous to the phase mode in section 4 we get an action for this mode to be,

$$S_{\text{phase}} = -\frac{r_h^2}{G_N} 2\pi \frac{1}{r_h} \int dt \delta\dot{\theta}(t)^2. \tag{5.23}$$

⁸There are also additional effects since the geometry is no longer AdS_2 once F departs from the value F^Q .

The coefficient in eq. (5.23) is related, upto a factor of 2, to the coefficient of the term required to connect the canonical and grand canonical partition functions [71] as discussed in appendix C, eq. (C.11). Note that both terms in eq. (5.21) and eq. (5.23) are suppressed by a factor of $\frac{1}{r_h}$.

The coupling between ϵ and $\delta\theta$ can be obtained from symmetry considerations, [16] as explained in appendix C. This leads to a full action for ϵ and $\delta\theta$, which upto quadratic order in ϵ is given by

$$S_{\text{mode}} = -\frac{r_h^2}{G_N} 2\pi \frac{1}{r_h} \int dt \left(\delta\dot{\theta}(t) + \dot{\epsilon}(t)\dot{\theta} \right)^2 + \frac{r_h^2}{G_N} \frac{L_2^2}{2} \frac{1}{r_h} \int dt \epsilon(t) \epsilon^{(4)}(t). \quad (5.24)$$

We can now consider coupling these degrees of freedom to the source terms which arise from the scalar ψ . Carrying out a time reparametrisation on the two-point function, eq. (5.7), gives rise to a coupling to $\epsilon(t)$,

$$S_{\text{coup1}} = \frac{r_h^2}{G_N} 4\pi c (\tilde{\Delta}_- - \tilde{\Delta}_+) \tilde{\Delta}_+ \int dt dt' \left(\frac{\psi^\dagger(t)\psi(t') + \psi^\dagger(t')\psi(t)}{|t-t'|^{2\tilde{\Delta}_+}} \right) \left(\dot{\epsilon}(t) - \frac{2\epsilon(t)}{|t-t'|} \right). \quad (5.25)$$

Similarly a phase rotation and further time reparametrisation gives rise to the coupling

$$S_{\text{coup2}} = \frac{r_h^2}{G_N} 4\pi c e (\tilde{\Delta}_+ - \tilde{\Delta}_-) i \int dt dt' \left(\frac{\psi(t)\psi^\dagger(t') - \psi(t')\psi^\dagger(t)}{|t-t'|^{2\tilde{\Delta}_+}} \right) (\delta\theta(t) + \epsilon(t)\dot{\theta}). \quad (5.26)$$

The response to the sources can now be calculated and gives rise to a four-point function which is exactly the sum of the two terms, eq. (5.11) and (5.13) as discussed in appendix C.

We have focused on the zero temperature case above. At finite temperature the action for the time reparametrisation modes is given by

$$\begin{aligned} S_{\text{sch}} &= -\frac{r_h^2}{G_N} \frac{L_2^2}{r_h} \int_0^\beta d\tau \text{Sch} \left(\frac{\beta}{\pi} \tan \frac{\pi}{\beta} (\tau + \epsilon) \right) \\ &= -\frac{r_h^2}{G_N} \frac{L_2^2}{r_h} \int_0^\beta d\tau \left(-\frac{3}{2} (\epsilon'')^2 - \epsilon' \epsilon''' + \frac{2\pi^2}{\beta^2} (\epsilon')^2 \right) \end{aligned} \quad (5.27)$$

where $\tau \simeq \tau + \beta$ and $\epsilon' = \frac{d\epsilon}{d\tau}$. In going to the second line, we have expanded the action to quadratic order in ϵ . And for the phase mode, eq. (5.23) takes the form,

$$S_{\text{phase}} = -\frac{r_h^2}{G_N} 2\pi \frac{1}{r_h} \int_0^\beta d\tau \left(\frac{d\delta\theta}{d\tau} + \dot{\theta} \frac{d\epsilon}{d\tau} \right)^2, \quad (5.28)$$

where $\dot{\theta}$ continues to be given by eq. (C.12). The coupling between the time reparametrisation and phase modes to the field ψ at finite temperature can be obtained by doing a phase rotation and time reparametrisation $t = \frac{\beta}{\pi} \tan \frac{\pi}{\beta} (\tau + \epsilon)$ to the two point function. This gives the couplings for the time reparametrisation mode $\epsilon(\tau)$ and the phase mode

$\delta\theta(\tau)$ with matter to be

$$S_{\text{coup1}} = \frac{r_h^2}{G_N} 4\pi c (\tilde{\Delta}_- - \tilde{\Delta}_+) \tilde{\Delta}_+ \left(\frac{\pi}{\beta}\right)^{2\tilde{\Delta}_+} \int d\tau_1 d\tau_2 \left(\frac{\psi^\dagger(\tau_1)\psi(\tau_2) + \psi^\dagger(\tau_2)\psi(\tau_1)}{\sin^{2\tilde{\Delta}_+ \frac{\pi}{\beta}}(\tau_1 - \tau_2)} \right) \times \left(\epsilon'(\tau_1) - 2\epsilon(\tau_1) \cot \frac{\pi}{\beta}(\tau_1 - \tau_2) \right), \quad (5.29)$$

$$S_{\text{coup2}} = \frac{r_h^2}{G_N} 4\pi c e (\tilde{\Delta}_+ - \tilde{\Delta}_-) \left(\frac{\pi}{\beta}\right)^{2\tilde{\Delta}_+} i \int d\tau_1 d\tau_2 \left(\frac{\psi(\tau_1)\psi^\dagger(\tau_2) - \psi(\tau_2)\psi^\dagger(\tau_1)}{\sin^{2\tilde{\Delta}_+ \frac{\pi}{\beta}}(\tau_1 - \tau_2)} \right) \times \left(\frac{d\delta\theta}{d\tau_1} + \dot{\theta} \frac{d\epsilon}{d\tau_1} \right). \quad (5.30)$$

The terms eqs. (5.27), (5.28), (5.29) and (5.30) then give rise to the four-point function at finite temperature, and also the higher point functions which can be obtained through tree level exchange of these modes.

5.4 Additional comments

We have been considering the scalar field to be in the S -wave sector so far, and have analysed the effects of gravity and the gauge field which can be taken to be in the S -wave sector as well. Let us now consider higher partial waves for the scalar field. To analyse the resulting behaviour it is useful to carry out a KK reduction on the S^2 . From gravity we then get 3 additional massless gauge fields in the 2 dimensional theory which arise from the S^2 isometries. In addition, we get extra massive fields, which can be of spin 0, 1, 2, arising from the gravity and gauge perturbations which are coupled.

The contribution to the four-point function from the exchange of the massive degrees can be obtained by working in the AdS_2 geometry, and the breaking of the AdS_2 isometry due to the back reaction of the dilaton can be neglected in this sector. These contributions go like $O(r_h^2)$ and are analogous to the contributions studied in section 3 due to the exchange of an intermediate scalar.

The contribution due to the additional massless gauge fields will be analogous to the those studied in section 4 due to a probe gauge field. The higher partial waves of the scalar will carry charge under these gauge fields. The resulting contribution will be of $O(r_h^3)$ and of the form, eq. (4.29), with the anomalous dimensions Δ_i and charges being determined by the angular momentum of the scalar partial waves.

At low energies the $O(r_h^3)$ contributions, if present, will dominate in the four-point function. These will arise from both the S -wave sector and the additional KK massless gauge fields. The additional massive field discussed above will each contribute of $O(r_h^2)$, and will be suppressed at low energies. However, their cumulative contribution could be significant when the energy is not very small, see subsection 4.2 for related discussion.

We have focussed on the 4 point function above, higher point correlators can be obtained through successive tree level exchange of the time reparametrisation and phase modes.

6 The near- AdS_2 near- CFT_1 correspondence

The discussion above leads us to a formulation of the near AdS_2 near CFT_1 correspondence (henceforth referred to as the $NAdS_2/NCFT_1$ correspondence). We have been analysing a four dimensional theory above. However, we could start with a general higher dimensional situation as well.⁹ The near-horizon geometries to which our discussion applies have the form¹⁰ $AdS_2 \times S^d$. The resulting low energy dynamics can be analysed by constructing the two dimensional theory obtained by KK reduction on S^d . This will contain two dimensional gravity, coupled to the dilaton, which is the radius of the S^d , and the gauge fields turned on in the near-horizon region. In addition there will be fields which are not excited in the background. In general this will include massless gauge fields and additional massive degrees of freedom.¹¹

For concreteness let us consider the case of a single background gauge field in asymptotically AdS_{d+2} space. We have seen that the S -wave sector determines the low energy response of the system. In particular, working in the Canonical Ensemble at fixed charge, the Schwarzian action, eq. (5.21), gives rise to the free energy at small temperatures, $T \ll \mu$. In addition this action along with the phase mode, eq. (5.24), gives rise to the four and higher point correlators for probes at low energies, $\omega \ll \mu$. For the additional massless gauge fields in the two dimensional theory (including those arising from the isometries of S^d) there will be additional phase modes whose effects also need to be incorporated, as discussed above. The action for the time reparametrisation and phase modes will be suppressed by a power of $\frac{1}{r_h}$, where r_h is the scale characterising the departures from the AdS_2 geometry which is related to the chemical potential for the gauge field by $r_h \sim \mu L_{d+2}^2$. As a result, the four-point function for example, of a probe will be enhanced by an additional power of $\frac{r_h}{\omega L_{d+2}^2}$. The effects of massive fields, of various spins, can be incorporated in a straightforward manner as discussed above, and the departure from the AdS_2 geometry can be neglected in calculating these effects at low energies. The contributions due to the individual exchange of each massive mode will be suppressed, but cumulatively could be significant, see subsection 4.2.

One comment is worth making before we proceed. The same length scale r_h enters in the discussion of the infrared dynamics in two conceptually different ways in the example above of an extremal RN black hole in AdS_{d+2} . First, r_h is the radius of the horizon and therefore determines the ground state entropy. Second, it also characterises the scale at which corrections to the $AdS_2 \times S^d$ geometry become significant. It is helpful to keep this distinction in mind while analysing the low energy dynamics.

The essential points in the discussion above for the asymptotically AdS_{d+2} case are in fact valid more generally in any situation where the near-horizon geometry takes the form $AdS_2 \times S^d$. This includes extremal black holes in asymptotically flat space-time whose behaviour is similar to “small” black holes in AdS_{d+2} space.

It is worth commenting on this case in some more detail here. For the asymptotically flat case, the same scale r_h in fact plays three different roles. It is the radius of the near

⁹Asymptotic AdS_3 space can be different, we therefore restrict ourselves to four and higher dimensions.

¹⁰We have not considered rotating extremal black holes in this paper.

¹¹Massless scalars need to be analysed separately.

horizon AdS_2 , L_2 , as well as the radius of the S^d . In addition, it also characterises the scale at which corrections to the AdS_2 geometry become significant. Once one keeps track of this distinction carefully, all formulae for the asymptotically AdS_{d+2} case generalise and can be applied to this case as well. More explicitly, to apply the formulae obtained in the discussion of the asymptotically AdS_{d+2} case for the asymptotically flat case, one should express them first in terms of r_h (which enters in the AdS_{d+2} case both as the horizon value of the S^d , and the scale characterising the departure from the IR conformal symmetry) and L_2 , and thereafter set $L_2 = r_h$.

In fact, for the case of four dimensions, we have already been careful to present the formulae for the AdS_{d+2} case keeping these aspects in mind. The Schwarzian term and action for the phase modes as given by eq. (5.24) are also valid for the asymptotically flat case, with $L_2 = r_h$. Coupling them to the matter fields, as given in eqs. (5.25) and (5.26) then gives the correlation functions which characterise the response of the black hole in the four dimensional asymptotically flat case as well.

It is worth pointing out that the chemical potential of the extremal black hole in the asymptotically flat case is different from the AdS_{d+2} case and is of order unity (in our units) instead of being of order $\mu \sim \frac{r_h}{L_2^2}$ as noted in eq. (2.15). In addition, since L_2 is determined by the charge of the black hole in the asymptotically flat case it is more appropriate to think of the mass of the probe field as being an independent parameter. The condition for being at small enough energies, so that the resulting dynamics is described by the near- AdS_2 region, is then more appropriately stated as eq. (3.2), which becomes,

$$\omega \ll m. \tag{6.1}$$

The requirement of being at low temperatures is given by $T \ll \frac{r_h}{L_2^2} = \frac{1}{r_h}$ and the free energy continues to be given by eq. (2.18) and is correctly obtained from the Schwarzian term, eq. (5.21).

In this way one can apply the discussion developed above equally well to asymptotically flat case. In summary, the low energy behaviour arises from the near- AdS_2 dynamics, and is dominated at sufficiently low energies by the time reparametrisation and phase modes whose action is suppressed by the scale r_h which governs the corrections to the near-horizon geometry.

One expects that the near- AdS_2 dynamics which can be calculated as given above should correspond to the low energy behaviour of an approximate one dimensional CFT, in any consistent theory of gravity. Different black holes should correspond to different CFTs but regardless of microscopic difference their dynamics, at low energies, should be the same and determined by the time reparametrisation and phase modes whose action arises due to effects which break the conformal symmetry.

One can further extend these considerations to a general black hole carrying multiple charges. It is well known that extremal black holes in general exhibit the attractor mechanism, which is a general feature of supersymmetric and non-supersymmetric extremal black holes [72–88]. Moduli fields, which are generically present in string theory, are fixed at the horizon by the charges, and the entropy as a result is only a function of the charges and

not the asymptotic values the moduli take. On the other hand the deviation away from the AdS_2 region depends on the asymptotic value for these moduli, which determine how the attractor values are approached at the horizon. As a result, the ground state entropy is universal and determined by the charges alone, while we expect the Schwarzian and phase actions to have coefficients which are dependent on additional data pertaining to the asymptotic behaviour of the moduli which determine how the IR theory is approached from the UV. A more detailed analysis of the connections between the attractor phenomenon and the low energy dynamics of extremal black holes is left for the future.

7 Conclusions

We have analysed the behaviour of charged extremal and near-extremal black holes in this paper, continuing the analysis begun in [37]. In particular, we have considered charged matter fields and studied their correlation functions showing that at low energies the leading contribution arises due to the effects of the time reparametrisation mode and an additional phase mode. These effects are enhanced at low frequencies by a factor of $\frac{r_h}{L_2^2 \omega}$ where r_h is the scale characterising departures from the AdS_2 geometry and L_2 is the radius of the AdS_2 geometry. The action governing these modes at finite temperatures is given by eqs. (5.27) and (5.28). It is known that the Schwarzian term correctly captures the near-extremal specific heat in the canonical ensemble. We have also shown that the action eqs. (5.27) and (5.28) gives rise to the four-point function for the charged scalar field.

While our discussion has focussed on the four-point function, it can be easily extended to higher point correlators which can be calculated via successive tree level exchanges of the time reparametrisation and phase modes. Also, some of the formulae in the discussion on correlators above were written down for the extremal case, however they can easily be extended to the case of non-extremal black holes at low temperatures as discussed at the end of subsection 5.3.

It is also worth mentioning that the phase mode we find does not correspond to a gauge transformation in the AdS_2 space-time. For example, from eq. (4.17) we see that the field strength is non-zero since θ is time dependent. This is in contrast to the time reparametrisation mode which corresponds to asymptotic isometries in the AdS_2 space-time. Physically speaking, this difference arises because a spherically symmetric configuration of charged matter gives rise to an electric field, while a finite temperature black hole in AdS_2 is still described locally by the AdS_2 geometry.

We have also studied the behaviour beyond the S -wave sector and found that it can be analysed by carrying out a KK reduction down to the near AdS_2 theory. Additional massless gauge fields which arise from isometries in the KK reduction give rise to additional phase modes which give enhanced effects. Massive modes arising in the KK reduction give rise to effects which are not enhanced and which can be calculated in the AdS_2 geometry neglecting any corrections to it. While the effects of each massive mode is small, they can cumulatively have a significant effect.

We have also discussed how the analysis carried out for big black holes in asymptotically AdS space can be easily applied to more general situations, including asymptotically flat

black holes which give rise to a near-horizon geometry of the $AdS_2 \times S^d$ form. In the general case of multi-charge black holes, the attractor mechanism leads to the ground state entropy being universal, however, the departure from the AdS_2 region depends on the asymptotic values moduli take and this could affect the coefficient of the Schwarzian and the phase mode action.

In this way, our analysis allows us to formulate a general method for calculating the low energy behaviour of extremal black holes. This should correspond on general grounds to the low energy behaviour of microscopic dual theories which are approximate one-dimensional CFTs, with the breaking of conformal invariance being important in giving rise to the action for the time reparametrisation and phase modes.

We have not considered rotating Kerr black holes here. These are charged with respect to the KK gauge fields which arise from the isometries of S^d . We hope to return to this question in a subsequent paper, see [65] for some recent discussion.

There are many interesting microscopic models of extremal black holes in string theory, arising for example from D -brane constructions. It will be worthwhile applying the results of this paper to these models to see if more can be learnt about their dynamics.

Our results also provide good motivation for the study of the Jackiw-Teitelboim model of gravity, coupled to additional fields, in more detail, since we see that this model captures the low energy behaviour of extremal systems in many cases. We hope to report on some progress in this direction as well in the future.

Acknowledgments

We thank Gautam Mandal, Shiraz Minwalla, Subir Sachdev, Ashoke Sen, Ronak Soni and Spenta Wadia for insightful discussions and comments. We are especially grateful to Juan Maldacena for his generous insights and comments. We acknowledge the support by the International Centre for Theoretical Sciences (ICTS) during a visit for participating in the program “AdS/CFT at 20 and Beyond” where part of this work was done. We also thank the participants of the program for their questions and comments. UM and VV thank the “Infosys Foundation International Exchange Program of ICTS” for support in participating in Strings 2018 held in Okinawa, Japan. UM gratefully acknowledges support from the Simons Center for Geometry and Physics, Stony Brook University for the 2018 Simons Summer Workshop, while this work was in progress. We thank the DAE, Government of India, for support. SPT also acknowledges support from the J.C. Bose fellowship of the DST, Government of India. We also acknowledge support from the Infosys Endowment for Research on the Quantum Structure of Spacetime. Most of all, we thank the people of India for generously supporting research in String Theory.

A 2-D action for charged matter coupled to a probe U(1) field

In section 4 we considered a scalar field charged under a probe U(1) field in the background of a RN black hole in AdS_4 . Here we give some details of the calculation of the onshell action eq. (4.16).

We start by splitting the bulk integral in the 4D action eq. (4.1) into the near and far horizon regions,

$$S_{gf} = -\frac{1}{G_N} \int_{\text{NH}} d^4x \sqrt{g} \left(\frac{1}{4} F^2 + e A_\mu J^\mu \right) - \frac{1}{G_N} \int_{\text{FH}} d^4x \sqrt{g} \left(\frac{1}{4} F^2 + e A_\mu J^\mu \right). \quad (\text{A.1})$$

The near horizon integral spans from the horizon to the boundary in the asymptotic AdS_2 region and the far horizon region lies between the AdS_2 boundary and the AdS_4 boundary.

Now consider the far horizon bulk integral in eq. (A.1). We can use the equations of motion eq. (4.3) to simplify it to obtain the onshell expression

$$S_{\text{FH}} = -\frac{e}{2G_N} \int_{\text{FH}} d^4x \sqrt{g} A_\mu J^\mu + \frac{1}{2G_N} \int_{\partial AdS_2} d^3x \sqrt{\gamma} F^{\mu\nu} n_\mu A_\nu, \quad (\text{A.2})$$

where the boundary term is evaluated at the AdS_2 boundary. Here n_μ is the normal vector to the AdS_2 boundary. Note that we ignore the AdS_4 boundary term since it does not contribute to the on-shell action. Also we work in the gauge where the only nonzero component is A_r . We can take the boundary to be located at a constant value of r , to the required order, and hence we can ignore the boundary term in eq. (A.2) as well.

Evaluating the bulk action in the far horizon region eq. (A.2) using the solution, eqs. (4.4) and (4.5), we get

$$S_{\text{FH}} = -\frac{e}{2G_N} \int_{\text{FH}} d^4x \sqrt{g} A_r J^r = -\frac{e^2}{2G_N} \int_{\text{FH}} d^4x \sqrt{g} g_{tt} g^{rr} J_r \partial_t^{-2} J_r. \quad (\text{A.3})$$

We can obtain the solution for ψ in the far horizon region from its behaviour at the AdS_2 boundary. In terms of z coordinate eq. (2.22), at the AdS_2 boundary $z = \delta$ we have,

$$\psi(t, z)|_{\partial AdS_2} = z^{\Delta_-} \psi(t) + c(z^{\Delta_+} - z^{\Delta_-} \delta^{\Delta_+ - \Delta_-}) \int dt' \frac{\psi(t')}{|t - t'|^{2\Delta_+}}, \quad (\text{A.4})$$

where c is a constant,

$$c = 2^{-2\nu} \frac{\Gamma(-\nu)}{\Gamma(\nu)}, \quad \nu = \Delta_+ - \Delta_- \quad (\text{A.5})$$

and Δ_\pm are given in eq. (4.10). In terms of r coordinate, eq. (2.22), the expression eq. (A.4) becomes

$$\psi(t, r)|_{\partial AdS_2} = \left(\frac{L_2^2}{r - r_h} \right)^{\Delta_-} \psi(t) + c \left(\left(\frac{L_2^2}{r - r_h} \right)^{\Delta_+} - \left(\frac{L_2^2}{r - r_h} \right)^{\Delta_-} \delta^{\Delta_+ - \Delta_-} \right) \int dt' \frac{\psi(t')}{|t - t'|^{2\Delta_+}}. \quad (\text{A.6})$$

We can extend this solution beyond the AdS_2 boundary as follows. We are working with small values of ω which satisfy eq. (2.25). Therefore, the ω dependence or equivalently the time dependence of the solution for ψ is fixed to be that in eq. (A.6) for the region beyond the AdS_2 boundary where the condition eq. (2.25) is met. Therefore in this region, ψ takes the form,

$$\psi(t, r) = f_-(r) \psi(t) + c f_+(r) \int dt' \frac{\psi(t')}{|t - t'|^{2\Delta_+}}, \quad (\text{A.7})$$

where the functions $f_{\pm}(r)$ are solutions to the equation

$$\frac{1}{\sqrt{g}}\partial_r(\sqrt{g}g^{rr}\partial_rf_{\pm}) - m^2f_{\pm} = 0, \tag{A.8}$$

satisfying the condition that as $r \rightarrow r_c$

$$\begin{aligned} f_-|_{\partial AdS_2} &\rightarrow \left(\frac{L_2^2}{r-r_h}\right)^{\Delta_-}, \\ f_+|_{\partial AdS_2} &\rightarrow \left(\left(\frac{L_2^2}{r-r_h}\right)^{\Delta_+} - \left(\frac{L_2^2}{r-r_h}\right)^{\Delta_-}\delta^{\Delta_+-\Delta_-}\right). \end{aligned} \tag{A.9}$$

Plugging eq. (A.7) into the expression for the current eq. (4.2), we get,

$$J_r = -i c \int dt' \left(\frac{\psi(t)\psi^\dagger(t') - \psi(t')\psi^\dagger(t)}{|t-t'|^{2\Delta_+}}\right)(f_+\partial_rf_- - f_-\partial_rf_+). \tag{A.10}$$

The Wronskian in eq. (A.8), can be shown to be of the form

$$W[f_+, f_-] = f_+\partial_rf_- - f_-\partial_rf_+ = \frac{\tilde{c}}{\sqrt{g}g^{rr}}. \tag{A.11}$$

Here \tilde{c} is independent of r and can be obtained by examining the behaviour of $W[f_+, f_-]$ near the AdS_2 boundary. Using the boundary condition eq. (A.9) and the AdS_2 metric eq. (2.11),

$$\tilde{c} = W\sqrt{g}g^{rr} = r_h^2(\Delta_+ - \Delta_-). \tag{A.12}$$

Plugging in eqs. (A.10) and (A.11) into the action eq. (A.3) and simplifying, we obtain

$$S_{\text{FH}} = \frac{r_h^4}{2G_N}(\Delta_+ - \Delta_-)^2 c^2 e^2 \int_{\text{FH}} d^4x \frac{1}{\sqrt{g}} g_{tt} g_{rr} F(t)\partial_t^{-2}F(t), \tag{A.13}$$

where $F(t)$ is defined as

$$F(t) = \int dt' \left(\frac{\psi(t)\psi^\dagger(t') - \psi(t')\psi^\dagger(t)}{|t-t'|^{2\Delta_+}}\right). \tag{A.14}$$

Note that the metric is of the form

$$ds^2 = a(r)^2 dt^2 + a(r)^{-2} dr^2 + r^2 d\Omega_2^2, \tag{A.15}$$

where a^2 is given in eq. (2.2). Therefore the integral over r in eq. (A.13) becomes,

$$\int_{r_c}^{\infty} \frac{dr}{r^2} = \frac{1}{r_c}. \tag{A.16}$$

For small ω satisfying eq. (2.25), we can approximate r_c with r_h . This is because the integral over r in the far horizon region spans till the location r_c such that

$$\frac{r_c - r_h}{L_2} \sim \omega L_2 \Rightarrow r_c = r_h + O(\omega). \tag{A.17}$$

Therefore to leading order in ω after integrating over r and also over S^2 , the action in the far horizon region becomes

$$S_{\text{FH}} = \frac{r_h^2}{G_N} 2\pi r_h c^2 e^2 (\Delta_+ - \Delta_-)^2 \int dt F(t) \partial_t^{-2} F(t), \quad (\text{A.18})$$

where $F(t)$ is defined in eq. (A.14). We see that the region away from the horizon, near the AdS_2 boundary gives a contribution to the four-point function which is enhanced by a factor of r_h . It is straightforward to see that the near horizon region does not give such an enhanced contribution to the four-point function; it contributes at $O(r_h^2)$.

B 4-D action computation of 4 point function in background gauge field

In this appendix, we give details of the 4-point function computation of a scalar field charged under the gauge field which is excited in the background black hole solution. We work directly in 4 dimensions in the S -wave sector.

We start with the 4-D Euclidean action,

$$S = -\frac{1}{16\pi G_N} \int d^4x \sqrt{g} (R - 2\Lambda) - \frac{1}{4G_N} \int d^4x \sqrt{g} F^2 + \frac{1}{G_N} \int d^4x \sqrt{g} \left((D_\mu \psi)^\dagger D^\mu \psi + m^2 \psi^\dagger \psi \right) - \frac{1}{8\pi G_N} \int_\partial d^3x \sqrt{\gamma} K, \quad (\text{B.1})$$

where the Gibbons-Hawking term is evaluated at the AdS_4 boundary. This action admits the black hole solution eq. (2.2) and the scalar field ψ is charged under the background gauge field.

Varying the metric, we get the Einstein equations

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R + \Lambda g_{\mu\nu} = 16\pi G_N (T_{\mu\nu}^\psi + T_{\mu\nu}^F), \quad (\text{B.2})$$

where $T_{\mu\nu}^\psi$ and $T_{\mu\nu}^F$ are contributions to the stress tensor from the scalar field and gauge field respectively and are given by

$$T_{\mu\nu}^F = -\frac{1}{2G_N} \left(F_{\mu\rho} F_\nu^\rho - \frac{1}{4} g_{\mu\nu} F^2 \right), \quad (\text{B.3})$$

$$T_{\mu\nu}^\psi = \frac{1}{G_N} \left(\frac{1}{2} (D_\mu \psi)^\dagger D_\nu \psi + \frac{1}{2} (D_\nu \psi)^\dagger D_\mu \psi - \frac{1}{2} g_{\mu\nu} (g^{\alpha\beta} (D_\alpha \psi)^\dagger D_\beta \psi + m^2 \psi^\dagger \psi) \right). \quad (\text{B.4})$$

Note that when $\psi = 0$ and $J_r = 0$, we get the RN black hole solution given in eq. (2.2) which we will denote by $\bar{g}_{\mu\nu}$ below. Varying the gauge field gives the equation of motion,

$$\partial_\mu (\sqrt{g} g^{\mu\rho} g^{\nu\sigma} F_{\rho\sigma}) = e \sqrt{g} J^\nu, \quad (\text{B.5})$$

which can be solved to obtain

$$F_{tr} = -\frac{Q}{\sqrt{g} g^{rr} g^{tt}} + e g_{tt} \partial_t^{-1} J_r. \quad (\text{B.6})$$

Let us consider S -wave perturbations in the metric and work in the Fefferman-Graham gauge $\delta g_{\mu r} = 0$,

$$ds^2 = a^2 (1 + h_{tt}) dt^2 + a^{-2} dr^2 + b^2 (1 + h_{\theta\theta}) d\Omega_2^2, \quad (\text{B.7})$$

where a^2 and b^2 are given in eq. (2.2). These metric fluctuations induce perturbations in the gauge field. We have,

$$F_{tr} = \bar{F}_{tr} + F_{tr}^J + \delta F_{tr},$$

$$\bar{F}_{tr} = -\partial_r \bar{A}_t = -\frac{Q}{\sqrt{\bar{g}} \bar{g}^{rr} \bar{g}^{tt}}, \quad (\text{B.8})$$

$$F_{tr}^J = e \bar{g}_{tt} \partial_t^{-1} J_r, \quad (\text{B.9})$$

$$\delta F_{tr} = \partial_t \delta A_r, \quad (\text{B.10})$$

where \bar{F}_{tr} is the background field strength, F_{tr}^J is the response to the scalar current, upto the required order for the four-point function calculation, and δF_{tr} is the additional perturbation which arises due to metric fluctuations as we will see below. It is convenient to work in the gauge where \bar{A}_t is the nonzero component for the background gauge field and the perturbations in the gauge field have non-zero component only along the r direction. We denote the perturbation produced by the source J , eq. (B.9) as A^J , so that the full perturbation has a non-zero component $A_r^J + \delta A_r$.

δA_r can be obtained by linearising the first term in eq. (B.6),

$$-\frac{Q}{\sqrt{g} g^{rr} g^{tt}} = -\frac{Q}{\sqrt{\bar{g}} \bar{g}^{rr} \bar{g}^{tt}} \left(1 + \frac{1}{2} h_{tt} - h_{\theta\theta}\right) = \bar{F}_{tr} + \delta F_{tr}$$

$$\Rightarrow \delta A_r = -\bar{F}_{tr} \left(-\frac{1}{2} \partial_t^{-1} h_{tt} + \partial_t^{-1} h_{\theta\theta}\right). \quad (\text{B.11})$$

Expanding the action eq. (B.1) and using the equations of motion eqs. (B.2) and (B.5), we get the onshell action,

$$S_{\text{OS}} = \frac{1}{2} \int d^4x \sqrt{g} \delta g^{\mu\nu} T_{\mu\nu}^\psi - \frac{e}{2G_N} \int d^4x \sqrt{g} (A_r^J + \delta A_r) J^r$$

$$= -2\pi \int dt dr \left(\frac{b^2}{a^2} h_{tt} T_{tt}^\psi + 2h_{\theta\theta} T_{\theta\theta}^\psi \right) - \frac{2\pi e}{G_N} \int dt dr b^2 a^2 (A_r^J + \delta A_r) J_r. \quad (\text{B.12})$$

To get the four-point function for the field ψ , we express the perturbations in terms of the sources and compute the action to quartic order in ψ .

In the discussion below we will systematically examine the terms appearing in eq. (B.12). We define,

$$T_1 = -2\pi \int dt dr \left(\frac{b^2}{a^2} h_{tt} T_{tt}^\psi + 2h_{\theta\theta} T_{\theta\theta}^\psi \right), \quad (\text{B.13})$$

$$T_2 = -\frac{2\pi e}{G_N} \int dt dr b^2 a^2 A_r^J J_r, \quad (\text{B.14})$$

$$T_3 = -\frac{2\pi e}{G_N} \int dt dr b^2 a^2 \delta A_r J_r. \quad (\text{B.15})$$

Evaluating T_1

Consider first the term T_1 , eq. (B.13). This is infact similar to the expression one would obtain in the four-point function computation of an uncharged scalar field with a background magnetic charge, see section 3 of [37]. To simplify it, consider the tr and rr equations in eq. (B.2),

$$\left(\frac{a'}{a} - \frac{b'}{b}\right) \partial_t h_{\theta\theta} - \partial_t \partial_r h_{\theta\theta} = 16\pi G_N (T_{tr}^\psi + T_{tr}^F) = 16\pi G_N (T_{tr}^\psi + T_{tr}^J), \quad (\text{B.16})$$

$$\begin{aligned} \frac{1}{a^4} \partial_t^2 h_{\theta\theta} + \left(\frac{a'}{a} + \frac{b'}{b}\right) \partial_r h_{\theta\theta} + \frac{b'}{b} \partial_r h_{tt} + \frac{1}{a^2 b^2} h_{\theta\theta} &= 16\pi G_N (T_{rr}^\psi + T_{rr}^F) \\ \Rightarrow \frac{1}{a^4} \partial_t^2 h_{\theta\theta} + \left(\frac{a'}{a} + \frac{b'}{b}\right) \partial_r h_{\theta\theta} + \frac{b'}{b} \partial_r h_{tt} + \frac{1}{a^2 b^2} \left(1 - \frac{8\pi Q^2}{b^2}\right) h_{\theta\theta} &= 16\pi G_N (T_{rr}^\psi + T_{rr}^J), \end{aligned} \quad (\text{B.17})$$

where we have taken the terms involving the metric perturbation in $T_{\mu\nu}^F$ to the LHS. Here $T_{\mu\nu}^J$ is the $O(J)$ term in the stress tensor $T_{\mu\nu}^F$, eq. (B.3), and is defined as

$$T_{\mu\nu}^J = -\frac{1}{2G_N} \left(\bar{F}_{\mu\rho} (F^J)^\rho_\nu + F_{\mu\rho}^J \bar{F}_\nu^\rho - \frac{1}{2} \bar{g}_{\mu\nu} \bar{F}^{\alpha\beta} F_{\alpha\beta}^J \right), \quad (\text{B.18})$$

where $\bar{F}_{\mu\nu}$ and $F_{\mu\nu}^J$ are given in eq. (B.8) and eq. (B.9) respectively and the raising and lowering of indices are done with the background metric $\bar{g}_{\mu\nu}$. Also to the order we are interested in, we can write eq. (B.4) as

$$T_{\mu\nu}^\psi = \frac{1}{G_N} \left(\frac{1}{2} (\bar{D}_\mu \psi)^\dagger \bar{D}_\nu \psi + \frac{1}{2} (\bar{D}_\nu \psi)^\dagger \bar{D}_\mu \psi - \frac{1}{2} g_{\mu\nu} (g^{\alpha\beta} (\bar{D}_\alpha \psi)^\dagger \bar{D}_\beta \psi + m^2 \psi^\dagger \psi) \right), \quad (\text{B.19})$$

where the raising and lowering of indices are with the background metric $\bar{g}_{\mu\nu}$.

Note that the combined contribution to the stress tensor coming from ψ , eq. (B.19), and $F_{\mu\nu}^J$, eq. (B.18), is conserved with respect to the background metric i.e.

$$\nabla_{[\bar{g}]}^\mu (T_{\mu\nu}^\psi + T_{\mu\nu}^J) = 0, \quad (\text{B.20})$$

where $\nabla_{[\bar{g}]}$ denotes covariant derivative with respect to the background metric.

To simplify further, we write the expression eq. (B.13) as

$$\begin{aligned} T_1 &= -2\pi \int dt dr \left(\frac{b^2}{a^2} h_{tt} (T_{tt}^\psi + T_{tt}^J) + 2h_{\theta\theta} (T_{\theta\theta}^\psi + T_{\theta\theta}^J) \right) \\ &\quad + 2\pi \int dt dr \left(\frac{b^2}{a^2} h_{tt} T_{tt}^J + 2h_{\theta\theta} T_{\theta\theta}^J \right). \end{aligned} \quad (\text{B.21})$$

Let us denote the terms in the first and second lines of eq. (B.21) as T_{11} and T_{12} respectively. We can use the conservation equation eq. (B.20) and manipulate the equations, (B.16) and (B.17), analogous to [37] and show that T_{11} is given by

$$\begin{aligned} T_{11} &= -2\pi \int dt dr \left(\frac{b^2}{a^2} h_{tt} T_{tt} + 2h_{\theta\theta} T_{\theta\theta} \right) \\ &= -32\pi^2 G_N \int dt dr \left(\frac{2a^2 b^3}{b'} T_{rr} \frac{1}{\partial_t} T_{tr} - a^2 b^2 \left(1 + \frac{2a'b}{b'a} \right) T_{tr} \frac{1}{\partial_t^2} T_{tr} \right), \end{aligned} \quad (\text{B.22})$$

where

$$T_{\mu\nu} = T_{\mu\nu}^\psi + T_{\mu\nu}^J. \quad (\text{B.23})$$

Next, let us examine the contribution to the integral in T_{11} from the near-horizon region. As discussed in section 4 of [37], we can use the conservation equations for $T_{\mu\nu}$ to express eq. (B.22) as a boundary term at the AdS_2 boundary,

$$T_{11} = -\frac{32\pi^2 G_N r_h^3}{L_2^2} \int_{\text{bdy}} dt (T_{tz} - z\partial_t T_{zz}) \frac{1}{\partial_t^4} (T_{tz} - z\partial_t T_{zz}), \quad (\text{B.24})$$

where we have kept the leading behaviour which is $O(r_h^3)$. To simplify further, let us examine the behaviour of $T_{\mu\nu}^\psi$ and $T_{\mu\nu}^J$ at the AdS_2 boundary. From the form of ψ at the AdS_2 boundary, eq. (5.5), we get for the leading non-contact terms, $\psi^\dagger\psi = \psi^\dagger\partial_t\psi = \partial_t\psi^\dagger\partial_t\psi = 0$. Therefore, from eq. (B.19),

$$T_{tz}^\psi = \frac{1}{2G_N} \left(\partial_t\psi^\dagger\partial_z\psi + \partial_t\psi\partial_z\psi^\dagger \right) - \frac{e}{2G_N} \bar{A}_t J_z = \tilde{T}_{tz}^\psi - e \bar{A}_t J_z, \quad (\text{B.25})$$

$$T_{zz}^\psi = \frac{1}{G_N} \partial_z\psi\partial_z\psi^\dagger - \frac{1}{2G_N} \bar{g}_{zz} \left(\bar{g}^{\mu\nu} \partial_\mu\psi^\dagger\partial_\nu\psi + m^2\psi\psi^\dagger \right) = \tilde{T}_{zz}^\psi, \quad (\text{B.26})$$

where we $\tilde{T}_{tz}^\psi, \tilde{T}_{zz}^\psi$ are defined in eq. (5.10), J_z is defined in eq. (4.12) and \bar{A}_t is given in eq. (5.3).

Also, from eq. (B.18), we have

$$T_{tz}^J = 0, \quad (\text{B.27})$$

$$T_{zz}^J = -\frac{1}{2G_N} \frac{eQ}{r_h^2} \frac{L_2^2}{z^2} \partial_t^{-1} J_z, \quad (\text{B.28})$$

where we have used (see eqs. (B.8), (B.9))

$$\bar{F}_{tz} = \frac{Q}{r_h^2} \frac{L_2^2}{z^2}, \quad (\text{B.29})$$

$$F_{tz}^J = e \frac{L_2^2}{z^2} \partial_t^{-1} J_z. \quad (\text{B.30})$$

From eqs. (B.25), (B.26), (B.27), (B.28), it is straightforward to verify that

$$T_{tz} - z\partial_t T_{zz} = \tilde{T}_{tz}^\psi - z\partial_t \tilde{T}_{zz}^\psi. \quad (\text{B.31})$$

Therefore, eq. (B.24) becomes

$$T_{11} = -\frac{32\pi^2 G_N r_h^3}{L_2^2} \int_{\text{bdy}} dt (\tilde{T}_{tz}^\psi - z\partial_t \tilde{T}_{zz}^\psi) \frac{1}{\partial_t^4} (\tilde{T}_{tz}^\psi - z\partial_t \tilde{T}_{zz}^\psi). \quad (\text{B.32})$$

We see that the expression scales as $O(r_h^3)$ and agrees with eq. (5.9) of section 5. The contribution of the region away from the horizon to the integral eq. (B.22) gives an $O(r_h)^2$ contribution and is subleading.

Let us now examine the second line of eq. (B.21), T_{12} . From the expression for the stress tensor $T_{\mu\nu}^J$, eq. (B.18), we obtain

$$T_{tt}^J = -\frac{e}{2G_N} a^4 \bar{F}_{tr} \partial_t^{-1} J_r, \tag{B.33}$$

$$T_{\theta\theta}^J = \frac{e}{2G_N} b^2 a^2 \bar{F}_{tr} \partial_t^{-1} J_r. \tag{B.34}$$

Therefore, we have

$$\begin{aligned} T_{12} &= 2\pi \int dt dr \left(\frac{b^2}{a^2} h_{tt} T_{tt}^J + 2h_{\theta\theta} T_{\theta\theta}^J \right) \\ &= -\frac{2\pi e}{G_N} \int dt dr b^2 a^2 \bar{F}_{tr} \left(-\frac{1}{2} \partial_t^{-1} h_{tt} + \partial_t^{-1} h_{\theta\theta} \right) J_r. \end{aligned} \tag{B.35}$$

We will show that the expression eq. (B.35) cancels with the term T_3 , eq. (B.15).

Evaluating T_2 and T_3

Let us now examine the remaining terms in the action, eqs. (B.14) and (B.15). From eqs. (4.5) and (B.11), we get

$$T_2 = -\frac{2\pi e^2}{G_N} \int dt dr b^2 a^4 J_r \partial_t^{-2} J_r, \tag{B.36}$$

$$T_3 = \frac{2\pi e}{G_N} \int dt dr b^2 a^2 \bar{F}_{tr} \left(-\frac{1}{2} \partial_t^{-1} h_{tt} + \partial_t^{-1} h_{\theta\theta} \right) J_r. \tag{B.37}$$

The expression eq. (B.36) is same as what we obtained for the probe U(1) case as in section 4, see eq. (4.6). Therefore a similar calculation as done in appendix A follows resulting in an r_h^3 enhanced contribution to the four-point function which is given by eq. (5.13).

Moreover, the expression eq. (B.37) cancels with the term T_{12} , eq. (B.35).

Therefore, in the final expression for the onshell action eq. (B.12), the leading contribution which goes like $O(r_h^3)$ comes from the terms eq. (5.11) and eq. (5.13). These terms can also be obtained by the exchange of the time reparamerisation and phase modes as explained in section 5.

C Phase mode action

In appendix B, we derived the onshell 4-D action for a scalar charged under the background gauge field. Here we derive a 2-D action by dimensional reduction and going to the near horizon AdS_2 region. We start with the 4-D action

$$S = -\frac{1}{16\pi G_N} \int d^4x \sqrt{g} (R - 2\Lambda) - \frac{1}{4G_N} \int d^4x \sqrt{g} F^2 - \frac{1}{8\pi G_N} \int_{\partial} d^3x \sqrt{\gamma} K. \tag{C.1}$$

Let us now dimensionally reduce this action by taking the metric to be of the form

$$ds^2 = g_{\alpha\beta}(t, r) dx^\alpha dx^\beta + \Phi(t, r)^2 d\Omega_2^2, \tag{C.2}$$

where Φ is the dilaton field and $g_{\alpha\beta}$ is the 2-D metric in the $r - t$ plane. Since we are interested in the near horizon region and to make contact with the JT model [34, 35, 41], we do a Weyl rescaling of the metric,

$$g_{\alpha\beta} \rightarrow \frac{r_h}{\Phi} g_{\alpha\beta}. \tag{C.3}$$

The action eq. (C.1) becomes

$$S = -\frac{1}{4G_N} \int d^2x \sqrt{g} \left(\frac{2r_h}{\Phi} + \Phi^2 R - 2r_h \Phi \Lambda \right) - \frac{\pi}{G_N} \int d^2x \sqrt{g} \frac{\Phi^3}{r_h} F_{\alpha\beta} F_{\gamma\delta} g^{\alpha\gamma} g^{\beta\delta} - \frac{1}{2G_N} \int_{\text{bdy}} \sqrt{\gamma} \Phi^2 K. \tag{C.4}$$

We are interested in this action in the near- AdS_2 region. The boundary term in eq. (C.4) is then evaluated at the boundary of this region

Let us expand the dilaton in a perturbation in ϕ about its horizon value r_h ,

$$\Phi = r_h(1 + \phi). \tag{C.5}$$

We keep only terms upto $O(\phi)$ in the action; higher powers of ϕ do not contribute to the four-point function at $O(r_h^3)$ as argued in [37].

To $O(\phi^0)$, the action eq. (C.4) is

$$S_0 = -\frac{r_h^2}{4G_N} \left(\int d^2x \sqrt{g} R + 2 \int_{\text{bdy}} \sqrt{\gamma} K \right) - \frac{4\pi Q^2}{G_N r_h^2} \int d^2x \sqrt{g} - \frac{r_h^2}{G_N} \pi \int d^2x \sqrt{g} g^{\alpha\gamma} g^{\beta\delta} \left(F_{\alpha\beta} F_{\gamma\delta} - F_{\alpha\beta}^Q F_{\gamma\delta}^Q \right), \tag{C.6}$$

where $F_{\alpha\beta}^Q$ is given by eq. (5.16). The first two terms in eq. (C.6) are topological. The third term in the first line diverges but the divergence is cut off by corrections to the AdS_2 geometry. The correct expression for this term is obtained as follows. In the extremal or near extremal geometry, the action eq. (C.1) gives, on shell, the grand canonical partition function S_{GC} . This is related to the canonical ensemble partition function, obtained from S_{CE} as

$$S_{GC} = S_{CE} - \frac{1}{2G_N} \int d^4x \sqrt{g_4} F^2, \tag{C.7}$$

where g_4 corresponds to the 4-D metric. The canonical partition function S_{CE} is given by the topological piece at extremality. And the extra piece in eq. (C.7) is the more correct version of the second term, in eq. (C.6), accounting for corrections away from the near-horizon region. This gives

$$\frac{1}{2G_N} \int_{r_h}^{\infty} dr \int d^3x \sqrt{g_4} F^2 = \frac{4\pi Q^2}{G_N r_h} \int dt \tag{C.8}$$

where we used

$$F_{rt} = \frac{Q}{\sqrt{g_4} g^{rr} g^{tt}} = \frac{Q}{r^2}. \tag{C.9}$$

Therefore to $O(\phi^0)$, the action eq. (C.6) can be written as

$$S_0 = -\frac{r_h^2}{4G_N} \left(\int d^2x \sqrt{g} R + 2 \int_{\text{bdy}} \sqrt{\gamma} K \right) - \frac{4\pi Q^2}{G_N r_h} \int dt - \frac{r_h^2}{G_N} \pi \int d^2x \sqrt{g} g^{\alpha\gamma} g^{\beta\delta} \left(F_{\alpha\beta} F_{\gamma\delta} - F_{\alpha\beta}^Q F_{\gamma\delta}^Q \right). \quad (\text{C.10})$$

The second term can be written as,

$$-\frac{4\pi Q^2}{G_N r_h} \int dt = -\left(\frac{r_h^2}{G_N} \right) 4\pi \frac{1}{r_h} \int dt \dot{\theta}^2, \quad (\text{C.11})$$

where

$$\dot{\theta} = \frac{Q}{r_h}. \quad (\text{C.12})$$

Eq. (C.11) is of the same form, upto a factor of 2, as the action obtained for the phase mode in section 4, eq. (4.21).

Expanding eq. (C.4) to $O(\phi^1)$ gives

$$S_1 = -\frac{r_h^2}{2G_N} \int d^2x \sqrt{g} \phi (R - \Lambda_2) - \frac{r_h^2}{G_N} \int_{\text{bdy}} \phi \sqrt{\gamma} K - \frac{r_h^2}{G_N} 3\pi \int d^2x \sqrt{g} \phi g^{\alpha\gamma} g^{\beta\delta} \left(F_{\alpha\beta} F_{\gamma\delta} - F_{\alpha\beta}^Q F_{\gamma\delta}^Q \right). \quad (\text{C.13})$$

Here the terms in the first line correspond to the JT model, where Λ_2 is the two dimensional cosmological constant, $\Lambda_2 = -\frac{2}{L_2^2}$.

Combining eq. (C.6) and eq. (C.13), the complete action, including the boundary terms gives eq. (5.15). The terms in the last line of eq. (5.15) are the ones in JT action, including the last term which is a counter term needed for finiteness, see section 4 of [37]. Here F^Q is defined in eq. (5.16).

Let us make one comment before proceeding. We have not been very explicit about the boundary condition to be satisfied by the gauge field perturbation above and similarly not been explicit about a possible boundary term involving the perturbation which may be needed in the action, eq. (C.4). The explicit calculation we do carry out are done directly in 4-D as discussed in section 5, the 2-D action discussed here provides motivation for the action of the phase mode and time reparametrisation mode as we discuss further below.

Let us turn to the phase mode next and go back to the 4 dimensional action, eq. (C.1) for analysing it. Consider a gauge field configuration expanded about the background:

$$F_{rt} = \frac{Q}{\sqrt{g} g^{rr} g^{tt}} - \partial_t \delta A_r, \quad (\text{C.14})$$

where

$$\delta A_r = \frac{r_h}{\sqrt{g} g^{rr} g^{tt}} \delta \theta. \quad (\text{C.15})$$

Here the perturbation is analogous to that in the probe gauge field case, eq. (4.17).

Plugging eq. (C.14), eq. (C.15) into eq. (C.1) we get a term quadratic in $\delta\theta$, eq. (5.23). We see that $\delta\theta$ behaves like the phase mode. Note that the coefficient is suppressed by $\frac{1}{r_h}$ similar to the Schwarzian term above.

Next we analyse how this phase mode couples to time reparametrisations. In the asymptotically AdS_2 region, working in Poincare coordinates, eq. (2.22), we get from eq. (C.14)

$$F_{zt} = -\frac{Q}{r_h^2} \frac{L_2^2}{z^2} - \partial_t \delta A_z. \tag{C.16}$$

We work in the gauge where

$$A_\mu = \bar{A}_\mu + \delta A_\mu \tag{C.17}$$

where \bar{A}_t is given in eq. (5.3) and

$$\delta A_z = \frac{1}{r_h} \frac{L_2^2}{z^2} \delta\theta, \tag{C.18}$$

with the other components vanishing.

To find the coupling of this mode with the time reparametrisation mode it is convenient to change gauge. A gauge transformation,

$$\tilde{A}_\alpha = A_\alpha + \partial_\alpha \left(\frac{1}{r_h} \frac{L_2^2}{z} \delta\theta(t) \right), \tag{C.19}$$

allows us to go to the gauge in which the z component of the gauge field is zero. We have

$$\tilde{A}_z = 0, \quad \tilde{A}_t = \frac{1}{r_h} \frac{L_2^2}{z} \left(\frac{Q}{r_h} + \delta\dot{\theta} \right). \tag{C.20}$$

We drop the tildes henceforth.

We think of the boundary time reparamtrisation as being embedded in the asymptotic isometry of AdS_2 . Under the transformation,

$$\begin{aligned} t' &= t + \epsilon(t) \\ z' &= z(1 + \dot{\epsilon}), \end{aligned} \tag{C.21}$$

the gauge field transforms as

$$\begin{aligned} A'_z(t', z') &= \frac{\partial t}{\partial z'} A_t = 0, \\ A'_t(t', z') &= \frac{\partial t}{\partial t'} A_t = \frac{1}{r_h} \frac{L_2^2}{z} \frac{1}{1 + \dot{\epsilon}} \left(\frac{Q}{r_h} + \delta\dot{\theta} \right) \\ \Rightarrow \frac{1}{r_h} \frac{L_2^2}{z'} \partial_{t'} \theta(t') &= \frac{1}{r_h} \frac{L_2^2}{z} \frac{1}{1 + \dot{\epsilon}} \left(\frac{Q}{r_h} + \delta\dot{\theta} \right) \\ \Rightarrow \partial_t \theta' &= \left(\frac{Q}{r_h} + \delta\dot{\theta} \right) (1 + \dot{\epsilon}). \end{aligned} \tag{C.22}$$

Therefore under the transformation eq. (C.21), $\delta\theta$ transforms to linear order in perturbations as

$$\delta\dot{\theta} \rightarrow \delta\dot{\theta} + \dot{\theta} \dot{\epsilon}, \tag{C.23}$$

where $\dot{\theta}$ is given in eq. (C.12). This means that under the coordinate transformation eq. (C.21) the kinetic term for the phase mode transforms as

$$S_{\text{phase}} = -\frac{r_h^2}{G_N} 2\pi \frac{1}{r_h} \int d\tau \left(\frac{d\delta\theta}{d\tau} \right)^2 \rightarrow -\frac{r_h^2}{G_N} 2\pi \frac{1}{r_h} \int d\tau \left(\frac{d\delta\theta}{d\tau} + \frac{d\epsilon}{d\tau} \dot{\theta} \right)^2. \quad (\text{C.24})$$

To calculate the four-point function, we couple the action eq. (5.15) to sources. The coupling between ϵ and ψ comes from the kinetic part of the action for the field ψ , see appendix E of [37] and we obtain eq. (5.25). The coupling between $\delta\theta$ and ψ can be obtained in the same way as with the probe U(1) in 4.1 and is given in eq. (5.26). Combining all the contributions, we get the total 2-D action to be

$$S_{2D} = S_{\text{mode}} + S_{\text{coup1}} + S_{\text{coup2}}, \quad (\text{C.25})$$

where the S_{mode} , S_{coup1} , S_{coup2} are given in eqs. (5.24), (5.25), (5.26) respectively. Integrating out ϵ and $\delta\theta$, we get the final onshell action to be $S_{2D} = S_{\text{bdy}}^\psi + S_J$, see eqs. (5.11), (5.13).

Open Access. This article is distributed under the terms of the Creative Commons Attribution License ([CC-BY 4.0](https://creativecommons.org/licenses/by/4.0/)), which permits any use, distribution and reproduction in any medium, provided the original author(s) and source are credited.

References

- [1] A. Sen, *Entropy Function and AdS₂/CFT₁ Correspondence*, *JHEP* **11** (2008) 075 [[arXiv:0805.0095](https://arxiv.org/abs/0805.0095)] [[INSPIRE](#)].
- [2] A. Sen, *Quantum Entropy Function from AdS₂/CFT₁ Correspondence*, *Int. J. Mod. Phys. A* **24** (2009) 4225 [[arXiv:0809.3304](https://arxiv.org/abs/0809.3304)] [[INSPIRE](#)].
- [3] J.M. Maldacena, J. Michelson and A. Strominger, *Anti-de Sitter fragmentation*, *JHEP* **02** (1999) 011 [[hep-th/9812073](https://arxiv.org/abs/hep-th/9812073)] [[INSPIRE](#)].
- [4] S. Sachdev and J. Ye, *Gapless spin fluid ground state in a random, quantum Heisenberg magnet*, *Phys. Rev. Lett.* **70** (1993) 3339 [[cond-mat/9212030](https://arxiv.org/abs/cond-mat/9212030)] [[INSPIRE](#)].
- [5] A. Kitaev, *A simple model of quantum holography*, talk at KITP, 2015.
- [6] J. Maldacena and D. Stanford, *Remarks on the Sachdev-Ye-Kitaev model*, *Phys. Rev. D* **94** (2016) 106002 [[arXiv:1604.07818](https://arxiv.org/abs/1604.07818)] [[INSPIRE](#)].
- [7] E. Witten, *An SYK-Like Model Without Disorder*, [arXiv:1610.09758](https://arxiv.org/abs/1610.09758) [[INSPIRE](#)].
- [8] D.J. Gross and V. Rosenhaus, *A Generalization of Sachdev-Ye-Kitaev*, *JHEP* **02** (2017) 093 [[arXiv:1610.01569](https://arxiv.org/abs/1610.01569)] [[INSPIRE](#)].
- [9] R. Gurau, *The complete 1/N expansion of a SYK-like tensor model*, *Nucl. Phys. B* **916** (2017) 386 [[arXiv:1611.04032](https://arxiv.org/abs/1611.04032)] [[INSPIRE](#)].
- [10] I.R. Klebanov and G. Tarnopolsky, *Uncolored random tensors, melon diagrams and the Sachdev-Ye-Kitaev models*, *Phys. Rev. D* **95** (2017) 046004 [[arXiv:1611.08915](https://arxiv.org/abs/1611.08915)] [[INSPIRE](#)].
- [11] Y. Gu, X.-L. Qi and D. Stanford, *Local criticality, diffusion and chaos in generalized Sachdev-Ye-Kitaev models*, *JHEP* **05** (2017) 125 [[arXiv:1609.07832](https://arxiv.org/abs/1609.07832)] [[INSPIRE](#)].

- [12] M. Berkooz, P. Narayan, M. Rozali and J. Simón, *Higher Dimensional Generalizations of the SYK Model*, *JHEP* **01** (2017) 138 [[arXiv:1610.02422](#)] [[INSPIRE](#)].
- [13] A. Jevicki, K. Suzuki and J. Yoon, *Bi-Local Holography in the SYK Model*, *JHEP* **07** (2016) 007 [[arXiv:1603.06246](#)] [[INSPIRE](#)].
- [14] W. Fu, D. Gaiotto, J. Maldacena and S. Sachdev, *Supersymmetric Sachdev-Ye-Kitaev models*, *Phys. Rev. D* **95** (2017) 026009 [[arXiv:1610.08917](#)] [[INSPIRE](#)].
- [15] A.M. García-García and J.J.M. Verbaarschot, *Spectral and thermodynamic properties of the Sachdev-Ye-Kitaev model*, *Phys. Rev. D* **94** (2016) 126010 [[arXiv:1610.03816](#)] [[INSPIRE](#)].
- [16] R.A. Davison, W. Fu, A. Georges, Y. Gu, K. Jensen and S. Sachdev, *Thermoelectric transport in disordered metals without quasiparticles: The Sachdev-Ye-Kitaev models and holography*, *Phys. Rev. B* **95** (2017) 155131 [[arXiv:1612.00849](#)] [[INSPIRE](#)].
- [17] C. Krishnan, S. Sanyal and P.N. Bala Subramanian, *Quantum Chaos and Holographic Tensor Models*, *JHEP* **03** (2017) 056 [[arXiv:1612.06330](#)] [[INSPIRE](#)].
- [18] T. Li, J. Liu, Y. Xin and Y. Zhou, *Supersymmetric SYK model and random matrix theory*, *JHEP* **06** (2017) 111 [[arXiv:1702.01738](#)] [[INSPIRE](#)].
- [19] A. Eberlein, V. Kasper, S. Sachdev and J. Steinberg, *Quantum quench of the Sachdev-Ye-Kitaev Model*, *Phys. Rev. B* **96** (2017) 205123 [[arXiv:1706.07803](#)] [[INSPIRE](#)].
- [20] S. Choudhury, A. Dey, I. Halder, L. Janagal, S. Minwalla and R. Poojary, *Notes on melonic $O(N)^{q-1}$ tensor models*, *JHEP* **06** (2018) 094 [[arXiv:1707.09352](#)] [[INSPIRE](#)].
- [21] J. Sonner and M. Vielma, *Eigenstate thermalization in the Sachdev-Ye-Kitaev model*, *JHEP* **11** (2017) 149 [[arXiv:1707.08013](#)] [[INSPIRE](#)].
- [22] J. Yoon, *SYK Models and SYK-like Tensor Models with Global Symmetry*, *JHEP* **10** (2017) 183 [[arXiv:1707.01740](#)] [[INSPIRE](#)].
- [23] K. Bulycheva, *A note on the SYK model with complex fermions*, *JHEP* **12** (2017) 069 [[arXiv:1706.07411](#)] [[INSPIRE](#)].
- [24] D.J. Gross and V. Rosenhaus, *A line of CFTs: from generalized free fields to SYK*, *JHEP* **07** (2017) 086 [[arXiv:1706.07015](#)] [[INSPIRE](#)].
- [25] J. Murugan, D. Stanford and E. Witten, *More on Supersymmetric and 2d Analogs of the SYK Model*, *JHEP* **08** (2017) 146 [[arXiv:1706.05362](#)] [[INSPIRE](#)].
- [26] C. Krishnan, K.V. Pavan Kumar and D. Rosa, *Contrasting SYK-like Models*, *JHEP* **01** (2018) 064 [[arXiv:1709.06498](#)] [[INSPIRE](#)].
- [27] R. Bhattacharya, S. Chakrabarti, D.P. Jatkar and A. Kundu, *SYK Model, Chaos and Conserved Charge*, *JHEP* **11** (2017) 180 [[arXiv:1709.07613](#)] [[INSPIRE](#)].
- [28] S. Förste, J. Kames-King and M. Wiesner, *Towards the Holographic Dual of $N = 2$ SYK*, *JHEP* **03** (2018) 028 [[arXiv:1712.07398](#)] [[INSPIRE](#)].
- [29] A. Goel, H.T. Lam, G.J. Turiaci and H. Verlinde, *Expanding the Black Hole Interior: Partially Entangled Thermal States in SYK*, *JHEP* **02** (2019) 156 [[arXiv:1807.03916](#)] [[INSPIRE](#)].
- [30] C.-M. Chang, S. Colin-Ellerin and M. Rangamani, *On Melonic Supertensor Models*, *JHEP* **10** (2018) 157 [[arXiv:1806.09903](#)] [[INSPIRE](#)].

- [31] A. Blommaert, T.G. Mertens and H. Verschelde, *The Schwarzian Theory — A Wilson Line Perspective*, *JHEP* **12** (2018) 022 [[arXiv:1806.07765](#)] [[INSPIRE](#)].
- [32] T. Nosaka, D. Rosa and J. Yoon, *The Thouless time for mass-deformed SYK*, *JHEP* **09** (2018) 041 [[arXiv:1804.09934](#)] [[INSPIRE](#)].
- [33] D.V. Khveshchenko, *Thickening and sickening the SYK model*, *SciPost Phys.* **5** (2018) 012 [[arXiv:1705.03956](#)] [[INSPIRE](#)].
- [34] R. Jackiw, *Lower Dimensional Gravity*, *Nucl. Phys. B* **252** (1985) 343 [[INSPIRE](#)].
- [35] C. Teitelboim, *Gravitation and Hamiltonian Structure in Two Space-Time Dimensions*, *Phys. Lett.* **126B** (1983) 41 [[INSPIRE](#)].
- [36] J. Maldacena, D. Stanford and Z. Yang, *Conformal symmetry and its breaking in two dimensional Nearly Anti-de-Sitter space*, *PTEP* **2016** (2016) 12C104 [[arXiv:1606.01857](#)] [[INSPIRE](#)].
- [37] P. Nayak, A. Shukla, R.M. Soni, S.P. Trivedi and V. Vishal, *On the Dynamics of Near-Extremal Black Holes*, *JHEP* **09** (2018) 048 [[arXiv:1802.09547](#)] [[INSPIRE](#)].
- [38] A. Castro, D. Grumiller, F. Larsen and R. McNees, *Holographic Description of AdS_2 Black Holes*, *JHEP* **11** (2008) 052 [[arXiv:0809.4264](#)] [[INSPIRE](#)].
- [39] A. Castro and W. Song, *Comments on AdS_2 Gravity*, [arXiv:1411.1948](#) [[INSPIRE](#)].
- [40] K. Jensen, *Chaos in AdS_2 Holography*, *Phys. Rev. Lett.* **117** (2016) 111601 [[arXiv:1605.06098](#)] [[INSPIRE](#)].
- [41] A. Almheiri and J. Polchinski, *Models of AdS_2 backreaction and holography*, *JHEP* **11** (2015) 014 [[arXiv:1402.6334](#)] [[INSPIRE](#)].
- [42] J. Engelsöy, T.G. Mertens and H. Verlinde, *An investigation of AdS_2 backreaction and holography*, *JHEP* **07** (2016) 139 [[arXiv:1606.03438](#)] [[INSPIRE](#)].
- [43] A. Almheiri and B. Kang, *Conformal Symmetry Breaking and Thermodynamics of Near-Extremal Black Holes*, *JHEP* **10** (2016) 052 [[arXiv:1606.04108](#)] [[INSPIRE](#)].
- [44] M. Cvetič and I. Papadimitriou, *AdS_2 holographic dictionary*, *JHEP* **12** (2016) 008 [Erratum *ibid.* **01** (2017) 120] [[arXiv:1608.07018](#)] [[INSPIRE](#)].
- [45] G. Mandal, P. Nayak and S.R. Wadia, *Coadjoint orbit action of Virasoro group and two-dimensional quantum gravity dual to SYK/tensor models*, *JHEP* **11** (2017) 046 [[arXiv:1702.04266](#)] [[INSPIRE](#)].
- [46] D.J. Gross and V. Rosenhaus, *The Bulk Dual of SYK: Cubic Couplings*, *JHEP* **05** (2017) 092 [[arXiv:1702.08016](#)] [[INSPIRE](#)].
- [47] S. Förste and I. Golla, *Nearly AdS_2 SUGRA and the super-Schwarzian*, *Phys. Lett. B* **771** (2017) 157 [[arXiv:1703.10969](#)] [[INSPIRE](#)].
- [48] S.R. Das, A. Jevicki and K. Suzuki, *Three Dimensional View of the SYK/ AdS Duality*, *JHEP* **09** (2017) 017 [[arXiv:1704.07208](#)] [[INSPIRE](#)].
- [49] H. Kyono, S. Okumura and K. Yoshida, *Comments on 2D dilaton gravity system with a hyperbolic dilaton potential*, *Nucl. Phys. B* **923** (2017) 126 [[arXiv:1704.07410](#)] [[INSPIRE](#)].
- [50] C. Eling, *Holography and AdS_2 gravity with a dynamical aether*, *JHEP* **07** (2017) 147 [[arXiv:1705.04334](#)] [[INSPIRE](#)].

- [51] S. Dartois, H. Erbin and S. Mondal, *Conformality of $1/N$ corrections in SYK-like models*, [arXiv:1706.00412](#) [[INSPIRE](#)].
- [52] C. Krishnan and K.V.P. Kumar, *Towards a Finite- N Hologram*, *JHEP* **10** (2017) 099 [[arXiv:1706.05364](#)] [[INSPIRE](#)].
- [53] S. Dubovsky, V. Gorbenko and M. Mirbabayi, *Asymptotic fragility, near AdS_2 holography and $T\bar{T}$* , *JHEP* **09** (2017) 136 [[arXiv:1706.06604](#)] [[INSPIRE](#)].
- [54] M. Taylor, *Generalized conformal structure, dilaton gravity and SYK*, *JHEP* **01** (2018) 010 [[arXiv:1706.07812](#)] [[INSPIRE](#)].
- [55] D. Grumiller, R. McNees, J. Salzer, C. Valcárcel and D. Vassilevich, *Menagerie of AdS_2 boundary conditions*, *JHEP* **10** (2017) 203 [[arXiv:1708.08471](#)] [[INSPIRE](#)].
- [56] J. de Boer, E. Lladrés, J.F. Pedraza and D. Vegh, *Chaotic strings in AdS/CFT* , *Phys. Rev. Lett.* **120** (2018) 201604 [[arXiv:1709.01052](#)] [[INSPIRE](#)].
- [57] S.-K. Jian, Z.-Y. Xian and H. Yao, *Quantum criticality and duality in the Sachdev-Ye-Kitaev/ AdS_2 chain*, *Phys. Rev. B* **97** (2018) 205141 [[arXiv:1709.02810](#)] [[INSPIRE](#)].
- [58] R.-G. Cai, S.-M. Ruan, R.-Q. Yang and Y.-L. Zhang, *The String Worldsheet as the Holographic Dual of SYK State*, [arXiv:1709.06297](#) [[INSPIRE](#)].
- [59] N. Halmagyi and S. Mondal, *Tensor Models for Black Hole Probes*, [arXiv:1711.04385](#) [[INSPIRE](#)].
- [60] A. Kitaev and S.J. Suh, *The soft mode in the Sachdev-Ye-Kitaev model and its gravity dual*, *JHEP* **05** (2018) 183 [[arXiv:1711.08467](#)] [[INSPIRE](#)].
- [61] S.R. Das, A. Ghosh, A. Jevicki and K. Suzuki, *Three Dimensional View of Arbitrary q SYK models*, *JHEP* **02** (2018) 162 [[arXiv:1711.09839](#)] [[INSPIRE](#)].
- [62] S.R. Das, A. Ghosh, A. Jevicki and K. Suzuki, *Space-Time in the SYK Model*, *JHEP* **07** (2018) 184 [[arXiv:1712.02725](#)] [[INSPIRE](#)].
- [63] F.M. Haehl and M. Rozali, *Fine Grained Chaos in AdS_2 Gravity*, *Phys. Rev. Lett.* **120** (2018) 121601 [[arXiv:1712.04963](#)] [[INSPIRE](#)].
- [64] T.G. Mertens, *The Schwarzian theory — origins*, *JHEP* **05** (2018) 036 [[arXiv:1801.09605](#)] [[INSPIRE](#)].
- [65] A. Castro, F. Larsen and I. Papadimitriou, *5D rotating black holes and the $nAdS_2/nCFT_1$ correspondence*, *JHEP* **10** (2018) 042 [[arXiv:1807.06988](#)] [[INSPIRE](#)].
- [66] A. Gaikwad, L.K. Joshi, G. Mandal and S.R. Wadia, *Holographic dual to charged SYK from 3D Gravity and Chern-Simons*, [arXiv:1802.07746](#) [[INSPIRE](#)].
- [67] H.T. Lam, T.G. Mertens, G.J. Turiaci and H. Verlinde, *Shockwave S -matrix from Schwarzian Quantum Mechanics*, *JHEP* **11** (2018) 182 [[arXiv:1804.09834](#)] [[INSPIRE](#)].
- [68] Y.-Z. Li, S.-L. Li and H. Lü, *Exact Embeddings of JT Gravity in Strings and M-theory*, *Eur. Phys. J. C* **78** (2018) 791 [[arXiv:1804.09742](#)] [[INSPIRE](#)].
- [69] D. Harlow and D. Jafferis, *The Factorization Problem in Jackiw-Teitelboim Gravity*, [arXiv:1804.01081](#) [[INSPIRE](#)].
- [70] K.S. Kolekar and K. Narayan, *AdS_2 dilaton gravity from reductions of some nonrelativistic theories*, *Phys. Rev. D* **98** (2018) 046012 [[arXiv:1803.06827](#)] [[INSPIRE](#)].

- [71] S.W. Hawking and S.F. Ross, *Duality between electric and magnetic black holes*, *Phys. Rev. D* **52** (1995) 5865 [[hep-th/9504019](#)] [[INSPIRE](#)].
- [72] S. Ferrara, R. Kallosh and A. Strominger, *$N = 2$ extremal black holes*, *Phys. Rev. D* **52** (1995) R5412 [[hep-th/9508072](#)] [[INSPIRE](#)].
- [73] A. Strominger, *Macroscopic entropy of $N = 2$ extremal black holes*, *Phys. Lett. B* **383** (1996) 39 [[hep-th/9602111](#)] [[INSPIRE](#)].
- [74] S. Ferrara and R. Kallosh, *Supersymmetry and attractors*, *Phys. Rev. D* **54** (1996) 1514 [[hep-th/9602136](#)] [[INSPIRE](#)].
- [75] S. Ferrara and R. Kallosh, *Universality of supersymmetric attractors*, *Phys. Rev. D* **54** (1996) 1525 [[hep-th/9603090](#)] [[INSPIRE](#)].
- [76] G. Lopes Cardoso, B. de Wit and T. Mohaupt, *Corrections to macroscopic supersymmetric black hole entropy*, *Phys. Lett. B* **451** (1999) 309 [[hep-th/9812082](#)] [[INSPIRE](#)].
- [77] G. Lopes Cardoso, B. de Wit and T. Mohaupt, *Deviations from the area law for supersymmetric black holes*, *Fortsch. Phys.* **48** (2000) 49 [[hep-th/9904005](#)] [[INSPIRE](#)].
- [78] G. Lopes Cardoso, B. de Wit and T. Mohaupt, *Macroscopic entropy formulae and nonholomorphic corrections for supersymmetric black holes*, *Nucl. Phys. B* **567** (2000) 87 [[hep-th/9906094](#)] [[INSPIRE](#)].
- [79] G. Lopes Cardoso, B. de Wit, J. Kappeli and T. Mohaupt, *Stationary BPS solutions in $N = 2$ supergravity with R^2 interactions*, *JHEP* **12** (2000) 019 [[hep-th/0009234](#)] [[INSPIRE](#)].
- [80] H. Ooguri, A. Strominger and C. Vafa, *Black hole attractors and the topological string*, *Phys. Rev. D* **70** (2004) 106007 [[hep-th/0405146](#)] [[INSPIRE](#)].
- [81] A. Sen, *Black hole entropy function and the attractor mechanism in higher derivative gravity*, *JHEP* **09** (2005) 038 [[hep-th/0506177](#)] [[INSPIRE](#)].
- [82] K. Goldstein, N. Iizuka, R.P. Jena and S.P. Trivedi, *Non-supersymmetric attractors*, *Phys. Rev. D* **72** (2005) 124021 [[hep-th/0507096](#)] [[INSPIRE](#)].
- [83] R. Kallosh, *New attractors*, *JHEP* **12** (2005) 022 [[hep-th/0510024](#)] [[INSPIRE](#)].
- [84] D. Astefanesei, K. Goldstein and S. Mahapatra, *Moduli and (un)attractor black hole thermodynamics*, *Gen. Rel. Grav.* **40** (2008) 2069 [[hep-th/0611140](#)] [[INSPIRE](#)].
- [85] A. Dabholkar, A. Sen and S.P. Trivedi, *Black hole microstates and attractor without supersymmetry*, *JHEP* **01** (2007) 096 [[hep-th/0611143](#)] [[INSPIRE](#)].
- [86] S. Ferrara, G.W. Gibbons and R. Kallosh, *Black holes and critical points in moduli space*, *Nucl. Phys. B* **500** (1997) 75 [[hep-th/9702103](#)] [[INSPIRE](#)].
- [87] R. Kallosh, N. Sivanandam and M. Soroush, *The Non-BPS black hole attractor equation*, *JHEP* **03** (2006) 060 [[hep-th/0602005](#)] [[INSPIRE](#)].
- [88] M. Alishahiha and H. Ebrahim, *Non-supersymmetric attractors and entropy function*, *JHEP* **03** (2006) 003 [[hep-th/0601016](#)] [[INSPIRE](#)].