



LAWRENCE
LIVERMORE
NATIONAL
LABORATORY

Femtoscscopy in Relativistic Heavy Ion Collisions

M. Lisa, S. Pratt, R. A. Soltz, U. Wiedemann

August 2, 2005

Annual Review of Nuclear and Particle Science

Disclaimer

This document was prepared as an account of work sponsored by an agency of the United States Government. Neither the United States Government nor the University of California nor any of their employees, makes any warranty, express or implied, or assumes any legal liability or responsibility for the accuracy, completeness, or usefulness of any information, apparatus, product, or process disclosed, or represents that its use would not infringe privately owned rights. Reference herein to any specific commercial product, process, or service by trade name, trademark, manufacturer, or otherwise, does not necessarily constitute or imply its endorsement, recommendation, or favoring by the United States Government or the University of California. The views and opinions of authors expressed herein do not necessarily state or reflect those of the United States Government or the University of California, and shall not be used for advertising or product endorsement purposes.

Femtoscopy in Relativistic Heavy Ion Collisions

MICHAEL ANNAN LISA

Department of Physics, The Ohio State University, Columbus, OH 43210, USA

SCOTT PRATT

Department of Physics and Astronomy, Michigan State University, East Lansing, MI 48824, USA

RON SOLTZ

N-Division, Livermore National Laboratory, 7000 East Avenue, Livermore, CA 94550, USA

URS WIEDEMANN

Theory Division, CERN, Geneva, SWITZERLAND

Key Words HBT, Intensity Interferometry, Heavy Ion Collisions, Femtoscopy

Abstract Analyses of two-particle correlations have provided the chief means for determining spatio-temporal characteristics of relativistic heavy ion collisions. We discuss the theoretical formalism behind these studies and the experimental methods used in carrying them out. Recent results from RHIC are put into context in a systematic review of correlation measurements performed over the past two decades. The current understanding of these results are discussed in terms of model comparisons and overall trends.

CONTENTS

Introduction	2
Theory and Phenomenology Basics	4
<i>Formalism</i>	4
<i>Identical-Particle Interference</i>	7
<i>Correlations from Coulomb and Strong Interactions</i>	7
<i>Coordinate Systems</i>	8
<i>Gaussian Parameterizations</i>	9
<i>The λ factor</i>	12
<i>Collective Flow and Blast-Wave Models</i>	13
<i>Generating Correlations Functions from Hydrodynamics and from Microscopic Simulations</i>	16
<i>Phase Space Density, Entropy and Coalescence</i>	17

arXiv:nucl-ex/0505014 v1 12 May 2005

Femtoscopic measurements	18
<i>Correlation Function Definition</i>	18
<i>Signal Construction</i>	19
<i>Background Construction</i>	20
<i>Corrections</i>	21
<i>Fitting</i>	22
Measured Femtoscopic Systematics	24
<i>System size: N_{part} and Multiplicity</i>	25
<i>Collision Orientation: Pair emission angle relative to \hat{b}</i>	26
<i>Boost invariance: $Y_{\pi\pi}$</i>	28
<i>Collective dynamics: k_T and particle mass</i>	29
<i>New possibilities: Non-identical particle correlations</i>	32
Interpretations of Experimental Results	34
<i>General Conclusions from Systematic Trends</i>	34
<i>Phase Space Density and Entropy</i>	35
<i>Dynamic Models and their Comparison with Data</i>	37
Summary	41

1 Introduction

The study of nucleus-nucleus collisions at ultra-relativistic energies aims at characterizing the dynamical processes via which matter at extreme densities is produced, and the fundamental properties which this matter exhibits. In nucleus-nucleus collisions, how do partonic equilibration processes proceed? For how long, over which spatial extension and at which density is a QCD equilibration state approached, and what are its properties? Particle densities attained during a heavy ion collision are expected to exceed significantly the inverse volume of a hadron. This implies that the high temperature phase of QCD, the Quark Gluon Plasma, comes within experimental reach. Chiral symmetry restoration and deconfinement phase transition are testable in heavy ion collisions. However, the experimental study of QCD at high temperatures and densities is complicated by the short lifetime and mesoscopic extension of the produced system. Femtoscopy, the spatio-temporal characterization of the collision region on the femtometer scale, is needed to frame any discussion of dynamical equilibration processes.

The Relativistic Heavy Ion Collider (RHIC) just finished the first part of a dedicated experimental heavy ion program. Center of mass energies ($\sqrt{s_{NN}} = 200$ GeV) exceeded those of previous fixed target experiments by a factor 10. The current discussion of RHIC data focuses mainly on several qualitatively novel phenomena which all support the picture that dense and rapidly equilibrating QCD matter is produced in the collision region (1; 2; 3; 4). In particular, identified single inclusive hadron spectra all appear to emerge from a common flow field whose size and dependence on transverse momentum and azimuth is consistent with expectations that the produced matter is a locally equilibrated, almost ideal fluid of very small viscosity. Moreover, high- p_T hadron spectra show a centrality dependent, strong suppression in Au+Au collisions, but not in a d+Au control experiment, indicating that even the hardest partons produced in the collision participate significantly in equilibration processes. In short, experiments at RHIC have demonstrated already that heavy ion collisions produce dense and equilibrating matter, and that controlled experimentation of this matter is pos-

sible, using a large variety of probes.

Despite these successes numerous questions remain concerning the state of the matter produced in these collisions. Most notably, the equation of state is far from being determined and issues concerning chiral symmetry restoration are largely unresolved. Addressing these fundamental questions about bulk matter requires a detailed understanding of the dynamics and chemistry of the collision which can only be acquired by thorough and coordinated analyses of data and theory. In particular, spatio-temporal aspects of the reaction need to be experimentally addressed. The small size, $\sim 10^{-14}$ m, and transient nature, $\sim 10^{-22}$ seconds, of the reactions preclude direct measurement of times or positions. Instead, femtoscopy must exploit measurements of asymptotic momenta. Correlations of two final-state particles at small relative momentum provide the most direct link to the size and lifetime of subatomic reactions (5; 6; 7; 8; 9; 10; 11; 12). Since correlations from either interactions or from identity interference are stronger for smaller separations in space-time, spatio-temporal information can be most easily extracted for small sources, opposite to the limitations of microscopes and telescopes.

The interference of two particles emitted from chaotic sources was first applied by Hanbury-Brown and Twiss (13; 14) where photons were exploited to determine source sizes for both laboratory and stellar sources in the 1950s and 1960s. Correlations of identical pions were shown to be sensitive to source dimensions in proton-antiproton collisions by Goldhaber-Goldhaber, Lee and Pais in 1960 (15). In the 1970s methods were refined by Kopylov (16; 17), Koonin, (18) and Gyulassy (19), and other classes of correlations were shown to be useful for source-size measurements, such as strong and Coulomb interactions. Bevalac analyses showed that interferometry was truly capable of quantitatively determining spatial and temporal source dimensions, (20), and of providing a stringent test of dynamical models (21). Throughout the last 25 years this phenomenology has developed into a precision tool for heavy ion collisions. Whereas hadronic sources are short lived and one measures correlations of the momentum of outgoing particles, stars are long-lived and require experimental filters to enforce the approximate simultaneity of the two photons. Despite the fact that the theory for these two classes of measurement are very different (22), the heavy-ion community often refers to any type of two-particles-at-small-relative-momentum analysis related to size and shape as HBT correlations, in reference to Hanbury-Brown and Twiss's original work with photons. To some, however, the term "HBT" refers only to identical-particle interferometry. Following Lednicky, we will employ the term "femtoscopy" (23; 24) to denote any measurement that provides spatio-temporal information, including coalescence analyses.

The theory and phenomenology of correlation femtoscopy are reviewed in the next section, with particular emphasis on describing the approximations used to derive the connection between spatio-temporal aspects of the emission function and correlations constructed from final-state momenta. Experimental methods and techniques are correspondingly reviewed in Sec. 3. Section 4 presents experimental results, with an emphasis on describing the systematics of source dimensions and lifetimes as a function of beam energy, system size, particle species and a particle's momentum. In addition to source dimensions, results for phase space density and entropy are presented. Comparisons of experimental results and transport models are presented in Sec. 5 with an emphasis on explaining the "HBT puzzle", i.e., the fact that dynamic descriptions that incorporate a phase

transition to a new state of matter with many degrees of freedom significantly over-predict observed source sizes. Section 6 summarizes the current state of the field and lists new directions and challenges for future theoretical and experimental analyses.

2 Theory and Phenomenology Basics

2.1 Formalism

Two-particle correlation functions are constructed as the ratio of the measured two-particle inclusive and single-particle inclusive spectra,

$$C^{ab}(\mathbf{P}, \mathbf{q}) = \frac{dN^{ab}/(d^3p_a d^3p_b)}{(dN^a/d^3p_a)(dN^b/d^3p_b)}, \quad (1)$$

$$P \equiv p_a + p_b, \quad q^\mu = (p_a - p_b)^\mu / 2 - \frac{(p_a - p_b) \cdot P}{2P^2} P^\mu.$$

The theoretical analysis of (1) aims at relating this experimentally measured correlation to the space-time structure of the particle emitting source (7; 6; 8; 11). Two forms are common for connecting the measured correlation function to the space-time emission function $s(p, x)$ through a convolution with the wave function ϕ . In the first form (25),

$$C^{ab}(\mathbf{P}, \mathbf{q}) = \frac{\int d^4x_a d^4x_b s_a(p_a, x_a) s_b(p_b, x_b) |\phi(\mathbf{q}', \mathbf{r}')|^2}{\int d^4x_a s_a(p_a, x_a) \int d^4x_b s_b(p_b, x_b)}. \quad (2)$$

In calculations of two particle correlation functions, the squared relative two-particle wave function $|\phi|^2$ serves generally as a weight, and the emission function $s(p, x)$ contains all space-time information about the source since it describes the probability of emitting a particle with momentum p from a space-time point x . Here, and throughout this section, primes denote quantities in the center-of-mass frame, i.e., the frame where $\mathbf{P} = 0$. The source function s_a is evaluated at the momentum $\bar{\mathbf{p}}_a = m_a \mathbf{P} / (m_a + m_b)$, $\bar{p}_a^0 = E_a(\bar{\mathbf{p}}_a)$.

The second form, which is equally justified as Equation 2 by the approximations described further below, is,

$$C^{ab}(\mathbf{P}, \mathbf{q}) - 1 = \int d^3r' \mathcal{S}_{\mathbf{P}}(\mathbf{r}') [|\phi(\mathbf{q}', \mathbf{r}')|^2 - 1], \quad (3)$$

$$\mathcal{S}_{\mathbf{P}}(\mathbf{r}') \equiv \frac{\int d^4x_a d^4x_b s_a(\bar{p}_a, x_a) s_b(\bar{p}_b, x_b) \delta(\mathbf{r}' - \mathbf{x}'_a + \mathbf{x}'_b)}{\int d^4x_a d^4x_b s_a(\bar{p}_a, x_a) s_b(\bar{p}_b, x_b)}.$$

This expression allows one to consider $|\phi|^2$ as a kernel with which one can transform from the coordinate-space basis to the relative-momentum basis. It also emphasizes the limitation of correlation functions, that they can provide, at best, the function $\mathcal{S}_{\mathbf{P}}(\mathbf{r}')$, the distribution of relative positions of particles with identical velocities and total momentum \mathbf{P} as they move in their asymptotic state. Thus, correlations do not measure the size of the entire source. Instead, they address the dimensions of the “region of homogeneity”, a term coined by Sinyukov (26), i.e., the size and shape of the phase space cloud of outgoing particles whose velocities have a specific magnitude and direction. If the collective expansion of the produced matter is strong, as is the case in central collisions, then the region

of homogeneity is significantly smaller than the entire source volume. In what follows, we discuss the assumptions on which Equations 2 and 3 are based.

We start from explicit expressions for the one- and two-particle spectra in terms of T -matrix elements. For one-particle emission,

$$E \frac{dN}{d^3p} = \int d^4x s(p, x) = \sum_{F'} \left| \int d^4x T_{F'}(x) e^{-ip \cdot x} \right|^2, \quad (4)$$

$$s(p, x) = \sum_{F'} \int d^4\delta x T_{F'}^*(x + \delta x/2) T_{F'}(x - \delta x/2) e^{ip \cdot \delta x}. \quad (5)$$

Here, F' refers to the state of all other particles in the system. All interactions with the residual system are incorporated into the T matrix. However, there is a choice as to whether mean-field interactions are included in the T matrix or are instead incorporated into the evolution matrix (27; 28; 29; 30; 31). For example, one can include the Coulomb interactions with the residual system by replacing the phase factor $e^{ip \cdot x}$ in Equation 4 with an outgoing Coulomb wave function. This can be quantitatively important, in particular for slow particles. It becomes difficult when the two particles interact with one another through the strong or Coulomb force, as this represents a quantum three-body problem.

Assumption 1: *Higher order (anti)symmetrization can be neglected.*

Equation 4 implies that all particles with asymptotic momentum p must have had their last interaction with the source at some point x . For distinguishable particles, this is indeed the case and Equation 4 does not represent an assumption. However, if there are $N_a > 1$ particles of the same type a , then one must consider $T_a(x_1 \cdots x_{N_a})$. The evolution matrix is then no longer a simple phase factor but includes $N_a!$ interference terms. The single-particle probability is then obtained by integrating over the other $N_a - 1$ momenta. This can be accomplished explicitly for simple source functions (32; 33; 34; 35; 36). The distortion to the single-particle spectra and to the two-particle correlation function were found to be important when the phase space density approached unity. Otherwise, Equation 4 is well justified.

Assumption 2: *The emission process is initially uncorrelated*

In writing Equation 4, one requires that two-particle matrix elements factorize, $T_{F''}(x_a, x_b) = T_{F'_a}(x_a) T_{F'_b}(x_b)$, i.e., that the emission is independent. If multi-particle symmetrization can be neglected, the two-particle evolution operator factorizes into a center-of-mass and a relative operator. One has $U(x_a, x_b; p_a, p_b) = u_{q'}(x'_a - x'_b) e^{iP \cdot (E'_a x_a / M_{inv} + E'_b x_b / M_{inv})}$ for non-identical particles, whereas for identical particles $U = e^{iP \cdot (x_a + x_b)/2} (u_{q'}(x'_a - x'_b) \pm (u_{q'}(x'_b - x'_a)) / \sqrt{2})$. This is illustrated in Figure 1. Then, the two-particle probability can be expressed in terms of one-particle source functions,

$$E_a E_b \frac{dN_{ab}}{d^3p_a d^3p_b} = \int d^4x_a d^4x_b d^4\tilde{q} s_a((E'_a / M_{inv})P + \tilde{q}, x_a) s_b((E'_b / M_{inv})P - \tilde{q}, x_b) \times d^4\delta r' e^{i\tilde{q} \cdot \delta r'} u_{q'}^*(x'_a - x'_b + \delta r'/2) u_{q'}(x'_a - x'_b - \delta r'/2). \quad (6)$$

Assumption 3: *Smoothness approximation*(37; 38; 39)

Equation 6 is difficult to evaluate as it requires knowledge of the source function evaluated off-shell. For the special case where the particles do not interact aside

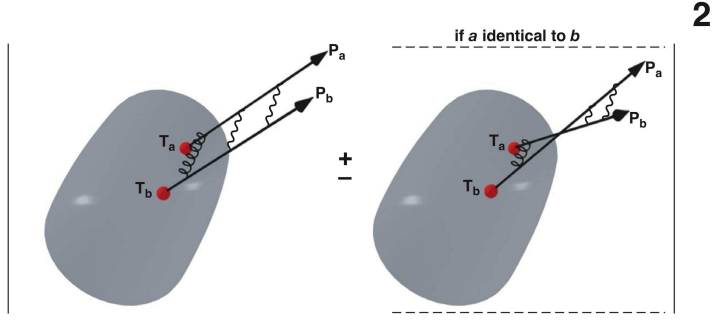


Figure 1: Schematic representation of the squared emission amplitude for two particles emitted independently from the grey-shaded source region and interacting with each other in the final states. For identical bosons (+) and fermions (-), correlation also involves interference between the paths.

from identical particle interference, $u_q(x_a - x_b) = [e^{iq \cdot (x_a - x_b)} \pm e^{iq \cdot (x_b - x_a)}] / \sqrt{2}$, the integrals over \tilde{q} and $\delta r'$ can be performed analytically,

$$E_a E_b \frac{dN_{ab}}{d^3 p_a d^3 p_b} = \int d^4 x_a d^4 x_b \{ s(p_a, x_a) s(p_b, x_b) \pm s(P/2, x_a) s(P/2, x_b) \cos[(p_a - p_b) \cdot (x_a - x_b)] \}. \quad (7)$$

The source functions in the interference term are evaluated off-shell for non-zero q , $P_0/2 \neq E(\mathbf{P}/2)$. The smoothness approximation replaces $s(P/2, x_a) s(P/2, x_b)$ with either $s(E(\mathbf{P}/2), \mathbf{P}/2, x_a) s(E(\mathbf{P}/2), \mathbf{P}/2, x_b)$ which leads to Equation 3, or with $s(p_a, x_a) s(p_b, x_b)$ which leads to Equation 2. If the first approximation is performed, one should also make the same approximation for the denominator. The smoothness approximation has been checked for expanding thermal sources, where it was found to be very reasonable for large (RHIC-like) sources, but quite questionable for smaller sources such as those found in pp or e^+e^- collisions (39).

Assumption 4: Equal time approximation

For the general case where the evolution operator incorporates Coulomb or strong interactions, deriving Equations 3 and 2 from Equation 6 is more complicated. First, the smoothness assumption amounts here to neglecting the \tilde{q} dependence in the product of the source functions in Equation 6. This assumption is somewhat more stringent in the presence of final state interactions, since the relevant range of \tilde{q} extends beyond q . With this assumption, one obtains a δ -function constraint for $\delta r'$, and the integrand of (6) is proportional to the squared evolution matrix $|u'_q(x'_a - x'_b)|^2$. This evolution matrix has non-zero time components, which must be neglected if one is to identify it with the relative wave function. Since the relative motion in the pair rest frame is small, one expects this approximation to be reasonable, but it has not yet been tested in model studies.

The above formalism is semi-classical in the sense that a quantum-mechanical particle emission probability, defined by the T -matrix elements, is usually approximated by classical source functions. As a consequence, quantum uncertainty limits the applicability of Equations 3 and 2. To illustrate this limitation, source functions have been evaluated by convoluting the emission function with wave packets(40; 41; 42; 43) of spatial width σ . This leads to a broadening of spatial distributions by $(\Delta R)^2 \sim 0.5 - 1.0 \text{ fm}^2$. Since quantum smearing may already be

incorporated into some of the semi-classical treatments through the choice of the initial density distribution, these calculations should be regarded as indicative of the theoretical uncertainty. Since this error affects the size in quadrature, it is negligible for large sources, but might be significant for sources near 1 fm in size with strong space-time correlations (42; 41; 44; 38; 39). In particular, analyses of $\pi\pi$ correlations from e^+e^- collisions, which result in source sizes of less than 1.0 fm(45; 46), are difficult to interpret in the above formalism.

2.2 Identical-Particle Interference

In the absence of strong and electromagnetic final state interactions, the wave function of an identical particle pair in Equation 3 becomes

$$|\phi(\mathbf{q}', \mathbf{r}')|^2 - 1 = \pm \cos(2\mathbf{q}' \cdot \mathbf{r}'). \quad (8)$$

The distribution of separations in coordinate space $\mathcal{S}_{\mathbf{P}}(\mathbf{r}')$ can then be determined by performing a three-dimensional Fourier transform of $C(\mathbf{q}') - 1$. It is instructive to consider the properties of this inversion in more detail. The curvature of $C(\mathbf{q})$ at $\mathbf{q} = 0$ can be related to the mean-square separation of the three-dimensional shape of $\mathcal{S}_{\mathbf{P}}(\mathbf{r})$ (we neglect the \mathbf{P} labels in C and \mathcal{S}),

$$-\frac{d^2 C(\mathbf{q})}{dq_i dq_j} = \langle r_i r_j \rangle = \int d^3 r \mathcal{S}_{\mathbf{P}}(\mathbf{r}) r_i r_j. \quad (9)$$

This relation has been useful to qualitatively illustrate the relation between specific space-time information and specific features of the correlator. However, applying the identity quantitatively requires careful consideration of pions from longer-lived resonances which can dominate the calculation of $\langle r^2 \rangle$ if not accounted for.

2.3 Correlations from Coulomb and Strong Interactions

Compared to the case for non-interacting identical particles, where the transformation between $\mathcal{S}_{\mathbf{P}}(\mathbf{r})$ and $C(\mathbf{P}, \mathbf{q})$ is a Fourier transform, analyzing the experimentally measured correlation function with Equation 3 to determine the source function is more complicated. Understanding the resolving power of the kernel $|\phi(\mathbf{q}, \mathbf{r})|^2$ requires a detailed understanding of the relative wave function. Once one averages over spins, the squared relative wave function is a function of q , r and $\cos \theta_{qr}$. In relativistic collisions, correlation analyses are usually confined to light singly charged hadrons. Coulomb-induced correlations are then weak and must be analyzed at small q where quantum effects become important ($qr/\hbar \sim 1$). The relative two-particle wave function in the presence of Coulomb interactions can then be written as a function of qr/\hbar , r/a_0 and $\cos \theta_{qr}$.

$$\phi = \Gamma(1 + i\eta) e^{-\pi\eta/2} e^{i\mathbf{q}\cdot\mathbf{r}} \left\{ 1 + \sum_{n=1}^{\infty} h_n \cdot (r/a_0)^n \right\}, \quad (10)$$

where a_0 is the Bohr radius, $h_1 = 1$ and $h_n = \frac{n-1-i\eta}{-in\eta} h_{n-1}$. Here, $\eta \equiv \mu e^2/\hbar q$ is independent of r , and for small r/a_0 the correlation function behaves as the Gamow factor, $G(\eta) \equiv e^{-\pi\eta} |\Gamma(1 + i\eta)|^2 = 2\pi\eta/(e^{2\pi\eta} - 1)$. Thus, the Coulomb

kernels have little resolving power for $\pi\pi$ correlations where the $a_0 = 387$ fm, but have greater resolving power for pK correlations where $a_0 = 84$ fm.

Coulomb interactions combine with identical-particle statistics. For $r \ll a_0$, the effects of Coulomb can be removed easily from the correlation function since η is independent of r (19),

$$|\phi|^2 \approx G(\eta)[1 + \cos(2qr \cos \theta_{qr})]. \quad (11)$$

For realistic source sizes, the order r/a_0 corrections are of the order of 10% for $\pi\pi$ correlations and are larger for heavier pairs (37; 47). Significant effort has been invested in “correcting” experimental correlation functions to remove Coulomb effects to all orders, but such “corrections” are model-dependent. The safest method for determining the source function is either inverting the full kernel (48; 49; 50; 51; 52; 53) or fitting $C(\mathbf{q})$ to some parameterized form for $\mathcal{S}(\mathbf{r})$ which is convoluted with the full kernel. Neither of these tactics are computationally prohibitive.

Strong interactions can also be exploited to provide size and shape information. If the size of the source is much larger than the range of the potential between the two particles, the kernel ($|\phi|^2 - 1$) can be determined entirely from knowledge of the phase shifts. Pairs that have a resonant interaction are especially useful, because the resonance will lead to a peak whose height is inversely proportional to the source volume, if $qR \gg \hbar$. At small q the kernel is determined by the scattering length, a , (54) and the height of the correlation at $q = 0$ becomes

$$C(\mathbf{q} = \mathbf{0}) - 1 = \left\langle -\frac{2a}{r} + \frac{a^2}{r^2} \right\rangle, \quad (12)$$

where the averaging is performed using $\mathcal{S}_P(\mathbf{r})$ as a weight. Of course, the effects of strong interactions, Coulomb and identical-particle interference can all combine as is the case for pp correlations. The pp kernel has been analyzed in depth by Brown and Danielewicz where the kernel was inverted and applied to experimental pp data. Evidence was seen for significant non-Gaussian behavior in the resulting source functions (52; 53). Strong-interactions also provide angular resolving power (55) which can be understood from the perspective of classical trajectories. Even s -wave scattering can be exploited to discern information about shape.

Strong and Coulomb induced correlations apply to both identical and non-identical particles. For non-identical particles the wave function is not symmetrized and $|\phi(\mathbf{q}, \mathbf{r})|^2 \neq |\phi(\mathbf{q}, -\mathbf{r})|^2$ which results in odd components of the correlations function, $C(\mathbf{q}) \neq C(-\mathbf{q})$ if there are odd components of $\mathcal{S}(\mathbf{r})$ as is the case for non-identical particle pairs. This asymmetry can be experimentally exploited to investigate the spatio-temporal differences between the emission functions of different particle species (56; 57); this requires, however, sufficient statistics to select on the angle between the total and relative momentum in the pair center of mass $\angle(\mathbf{q}, \mathbf{P})$ (58).

2.4 Coordinate Systems

Correlation functions depend on two three-dimensional momenta, \mathbf{P} and \mathbf{q} . For high-energy collisions, analyses are usually performed in the “longitudinally co-moving system” (LCMS), a rest frame moving along the longitudinal (beam)

direction such that $P_z = 0$. Axes are usually chosen according to the “out-side-long” prescription. The longitudinal axis is chosen parallel to the beam, while the outward axis points in the direction of \mathbf{P} , which is perpendicular to the beam axis. The sideward axis is chosen perpendicular to the other two. If the system is boost-invariant, observables expressed in the LCMS variables are independent of P_z and the source has a reflection symmetry about the $r_{\text{long}} = 0$ plane. If the collision is central, there is also a reflection symmetry about the $r_{\text{side}} = 0$ plane. Any four-vector V can be expressed in this coordinate system using the four-momentum P to project out the components(59; 60; 61),

$$\begin{aligned} V_{\text{long}} &= (P_0 V_z - P_z V_0)/M_T, \\ V_{\text{out}} &= (P_x V_x + P_y V_y)/P_T, \\ V_{\text{side}} &= (P_x V_y - P_y V_x)/P_T, \end{aligned} \quad (13)$$

where $M_T = P_0^2 - P_z^2$ and $P_T^2 = P_x^2 + P_y^2$. Dimensions of the source function are typically quoted in this coordinate system. One could also perform a second boost to the pair frame, in which the transverse components of the total momentum are zero. Then,

$$V'_{\text{out}} = \frac{M_{\text{inv}}}{M_T} \frac{(P_x V_x + P_y V_y)}{P_T} - \frac{P_T}{M_T M_{\text{inv}}} P \cdot V, \quad (14)$$

where $M_{\text{inv}}^2 = P^2$. Relative wave functions are more conveniently expressed in the frame of the pair. For instance, a sharp resonant peak is no longer sharp if the correlation is viewed away from the pair frame. For pairs where the correlation is influenced by Coulomb and strong interactions, most analyses are conducted in the pair frame.

For non-zero impact parameters, azimuthal symmetry is lost and source functions also depend on the azimuthal direction of the pair’s total momentum (62; 63; 64). Also, if boost-invariance is broken, the pair’s rapidity needs to be specified. In this more general case, reflection symmetries are broken and the choice of the coordinate axes becomes somewhat arbitrary. One could orient the axes according the event’s impact parameter, or one could rotate the coordinate system so that in the new frame cross terms such as $\langle xy \rangle$ vanish (illustrated in Figure 2). One would then specify the Euler angles as part of the description of the source function.

2.5 Gaussian Parameterizations

To gain a physical understanding of the three-dimensional spatio-temporal source distribution, it is useful to summarize its size and shape with a few parameters. This motivates the study of Gaussian parameterizations for the source $\mathcal{S}_{\mathbf{P}}(\mathbf{r}')$ and the two-particle correlator. Realistic sources deviate from Gaussians, e.g. by exponential tails due to resonance decay contributions. The extracted Gaussian source parameters may thus depend on details of the fitting procedures. These shortcomings can be overcome with imaging methods, or more complicated forms for the fitting. A more general three-dimensional analysis of correlations would involve decomposing both the correlation and source functions in either spherical or Cartesian harmonics (65; 66). Although the detailed non-Gaussian aspects of the correlation are important, the extra information can also cloud the main

trends in the data. In practice, Gaussian parameterizations provide the standard minimal description of experimental data.

2.5.1 THE GENERAL CASE The most general form for a Gaussian source is $\exp\{-A_{\alpha\beta}(x_\alpha - \bar{x}_\alpha)(x_\beta - \bar{x}_\beta)\}$, where x_α refers to the four dimensions x, y, z, t , and A is a 4×4 real symmetric matrix. The most general form has 14 parameters, 10 parameters for A and four more parameters for the offsets \bar{x}_α . Reflection symmetries can be used to eliminate certain cross terms and some of the offsets (67). Furthermore, if the particles are identical or have the same phase space distributions, all the offsets can be set to zero. Since the source function for the second species b might also have 14 parameters, there could be up to 28 Gaussian parameters in describing both s_a and s_b . However, since $\mathcal{S}(\mathbf{r})$ depends only on the distribution of relative spatial coordinates in the pair frame, only nine Gaussian parameters are required to describe the most general for \mathcal{S} for a given \mathbf{P} . Three of these nine parameters can be identified with Gaussian widths, three can be identified with offsets, and the last three can either be identified with cross terms, or with the three Euler angles describing the orientation of the three-dimensional ellipse.

Using the reflection symmetries for mid-rapidity sources in a symmetric central collision, a Gaussian parameterization of the emission function for particle species a in the out-side-long coordinate system reads

$$s_a(p, x) \sim \exp \left\{ -\frac{(x_{\text{out}} - \bar{x}_{a,\text{out}} - V_{s,a}(t - \bar{t}_a))^2}{2R_{a,\text{out}}^2} - \frac{(x_{\text{side}})^2}{2R_{a,\text{side}}^2} - \frac{(x_{\text{long}})^2}{2R_{a,\text{long}}^2} - \frac{(t - \bar{t}_a)^2}{2(\Delta\tau_a)^2} \right\}, \quad (15)$$

The symmetries forbid any cross terms in the exponential involving x_{side} or x_{long} , such as $x_{\text{side}}x_{\text{out}}$ or $x_{\text{long}}t$, but do not forbid a cross term between outward position x_{out} and time t . Here, this cross term is taken into account by allowing the source to move in the outwards direction with a velocity V_s . The symmetries also forbid offsets in the sideward or longitudinal direction. These other offsets have also been addressed for non-central collisions (68; 69).

The correlation function is determined by the phase space density of the final state, Equation 27. The resulting phase space density is

$$f_a(\mathbf{p}, \mathbf{r}, t) \sim \exp \left\{ -\frac{[x_{\text{out}} - \bar{X}_a(t)]^2}{2[R_{a,\text{out}}^2 + (V_{s,a} - V_\perp)^2(\Delta\tau_a)^2]} - \frac{x_{\text{side}}^2}{2R_{a,\text{side}}^2} - \frac{x_{\text{long}}^2}{2R_{a,\text{long}}^2} \right\},$$

$$\bar{X}_a(t) = \bar{x}_{a,\text{out}} + V_\perp(t - \bar{t}_a). \quad (16)$$

Here, V_\perp is the velocity of the pair in the LCMS frame.

The correlation function is determined by the relative distance function $\mathcal{S}_{\mathbf{P}}(\mathbf{r}')$,

see Equation 3,

$$\begin{aligned} \mathcal{S}_{\mathbf{P}}(\mathbf{r}') &\sim \exp \left\{ -\frac{[r'_{\text{out}} - \bar{X}_{\text{out}}]^2}{4\gamma_{\perp}^2 R_{\text{out}}^2} - \frac{r'_{\text{side}}{}^2}{4R_{\text{side}}^2} - \frac{r'_{\text{long}}{}^2}{4R_{\text{long}}^2} \right\} \quad (17) \\ R_{\text{out}}^2 &= \frac{1}{2} [R_{a,\text{out}}^2 + R_{b,\text{out}}^2 + (V_{s,a} - V_{\perp})^2 (\Delta\tau_a)^2 + (V_{s,b} - V_{\perp})^2 (\Delta\tau_b)^2], \\ R_{\text{side}}^2 &= \frac{1}{2} [R_{a,\text{side}}^2 + R_{b,\text{side}}^2], \quad R_{\text{long}}^2 = \frac{1}{2} [R_{a,\text{long}}^2 + R_{b,\text{long}}^2], \\ \bar{X}_{\text{out}} &= \bar{x}'_{a,\text{out}} - \bar{x}'_{b,\text{out}}, \end{aligned}$$

where $\gamma_{\perp} \equiv (1 - V_{\perp})^{-1/2}$. Thus, there are four measurable parameters, R_{out} , R_{side} , R_{long} and \bar{X}_{out} . In the absence of any symmetry, there are five more terms: three cross terms ($R_{\text{out,side}}^2$, $R_{\text{out,long}}^2$ and $R_{\text{side,long}}^2$), and two more offsets (X_{side} and X_{long}). For identical particles all the offsets are zero.

2.5.2 SENSITIVITY TO LIFETIME Given the symmetries used to derive Equation 17, the most experiment can provide is the determination of the four parameters, R_{out} , R_{side} , R_{long} and \bar{X}_{out} . The only way to independently determine the lifetime is to assume that the two transverse spatial sizes are approximately equal (70). After assuming $R_{a,\text{out}}^2 + R_{b,\text{out}}^2 \approx R_{b,\text{side}}^2 + R_{a,\text{side}}^2$, Equation 17 yields for identical particles ($V_{a,s} = V_{b,s} = V_s$, $\Delta\tau_a = \Delta\tau_b = \Delta\tau$),

$$(V_{\perp} - V_s)^2 (\Delta\tau)^2 \approx R_{\text{out}}^2 - R_{\text{side}}^2. \quad (18)$$

In general, however, the source velocity is not precisely known and outward and sideward spatial dimensions are not exactly equal which result in a significant systematic error when applying Equation 18, especially since temporal effects enter in quadrature. The $R_{\text{out}}/R_{\text{side}}$ ratio only provides a reliable estimate of the lifetime when $(V_{\perp} - V_s)\Delta\tau$ is much larger than the transverse size.

2.5.3 GAUSSIAN CROSS TERMS In addition to the three Gaussian parameters, R_{out} , R_{side} and R_{long} that describe the dimensions of a Gaussian source, one needs in the general non-symmetric case, three more parameters to describe the angular orientation of the principal axes. Figure 2 displays how the principal axes might differ from the out-side-long axes once the collisions are off center. These three Euler angles combined with three sizes can also be related to the parameters A_{ij} in the general form for a three-dimensional Gaussian, $\exp(-A_{ij}r_i r_j)$, where A is a symmetric matrix with six independent components (71; 67; 62; 64).

For pairs of identical particles, the correlation function for Gaussian sources is also Gaussian,

$$C(\mathbf{q}) = 1 + \exp(-(\mathcal{R}^2)_{ij} Q_i Q_j), \quad (19)$$

where $Q = 2q$. The six experimentally determined parameters, $(\mathcal{R}^2)_{ij}$, can be related to the moments of $\mathcal{S}(\mathbf{r}')$ (72),

$$\langle r_i r_j \rangle = (\mathcal{R}^2)_{ij}. \quad (20)$$

For central collisions, the source sizes $(R^2)_{ij}$ depend only on the longitudinal pair momentum P_L and on the modulus of the transverse pair momentum $|\vec{P}_T|$. This is different for non-central collisions for which the azimuthal direction $\phi_{\text{pair}} = \angle(\vec{P}_T, \hat{b})$ of the transverse pair momentum with respect to the impact parameter

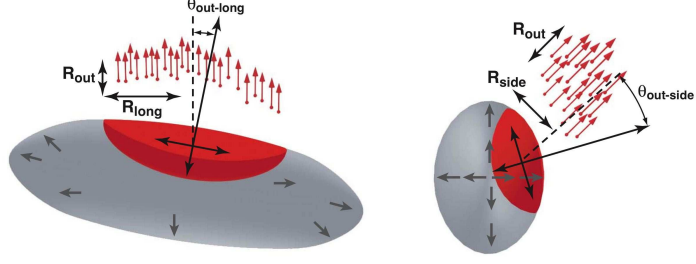


Figure 2: For non-central collisions the principal axes describing the orientation of the region of homogeneity can differ from the out-side-long axes. By viewing the source distribution from the perspective where the beam axis is oriented horizontally (left panel) and from peering down the beam pipe (right panel), the orientations leading to out-long and out-side cross terms are illustrated.

direction \hat{b} matters. The ϕ_{pair} -dependences of $(R^2)_{\text{out}}$, $(R^2)_{\text{side}}$ and $(R^2)_{\text{out-side}}$ then characterize the degree to which the initially out-of-plane-extended source geometry has expanded to the point where it becomes in-plane-extended (62; 63). The out-longitudinal and side-longitudinal cross terms contain information about the extent to which the main axis of the emission ellipsoid is *tilted* with respect to the beam axis(63). We return to this topic in Section 4.2.

The Yano-Koonin parameterization (73; 74; 67; 75) provides an alternative basis for describing the out-long cross term. The Yano-Koonin form is based on the assumption that one can boost along the beam axis to a source frame where the correlation function has a simple form.

$$C(\mathbf{P}, \mathbf{Q}) = 1 + \exp \left\{ -\tilde{Q}_{\perp}^2 R_{\perp}^2 - \tilde{Q}_{\parallel}^2 R_{\parallel}^2 - \tilde{Q}_0^2 R_0^2 \right\}, \quad (21)$$

where \tilde{Q} is the momentum difference defined in the source frame which has rapidity y_{YK} . In that frame R_0 is the Gaussian lifetime and R_{\perp} and R_{\parallel} are the dimensions of the source. This can be transposed to the out-long-side frame by boosting \tilde{Q} along the beam axis to the frame where $P_z = 0$, i.e. the LCMS frame, then using the fact that $Q_0 = Q_{\text{out}} P_{\perp} / P_0 \equiv Q_{\text{out}} V_{\perp}$ in the new frame. This yields $\tilde{Q}_0 = \cosh(y_{\pi\pi} - y_{YK}) Q_{\text{out}} V_{\perp} - \sinh(y_{\pi\pi} - y_{YK}) Q_{\text{long}}$, and $\tilde{Q}_{\parallel} = \cosh(y_{\pi\pi} - y_{YK}) Q_{\text{long}} - \sinh(y_{\pi\pi} - y_{YK}) Q_{\text{out}} V_{\perp}$, where $y_{\pi\pi}$ is the rapidity of the LCMS frame. Substituting these expressions into Equation 21 yields a cross term in the exponential equal to $Q_{\text{out}} Q_{\text{long}} (R_0^2 + R_{\parallel}^2) \sinh 2(y_{\pi\pi} - y_{YK})$, which disappears when $y_{YK} = y_{\pi\pi}$. By fitting y_{YK} as a function of the pair rapidity, aspects of boost invariance can be tested. Given that the distribution of source rapidities should fall off for large rapidities, one expects y_{YK} to lag $y_{\pi\pi}$ since pions of a given rapidity would more likely have been emitted from sources with smaller rapidities (74).

2.6 The λ factor

Many pions measured in experiment come from long-lived decays. Pions from weak decays may, or may not depending on the experiment, be identified and removed from the analysis since their decay vertices are typically a few centimeters from the reaction center. Decays from η or η' mesons occur a few thousand

fm away from the center of the collision. At these distances, they are effectively uncorrelated with other particles but can not be identified with experiment. If a fraction λ of the pairs originate from the spatio-temporal region relevant for correlations, the correlation is muted by the factor λ (19). If the source function is divided into two contributions, $\mathcal{S}(\mathbf{r}) = \lambda\mathcal{S}_{\text{local}} + (1 - \lambda)\mathcal{S}_{\infty}$, where both $\mathcal{S}_{\text{local}}$ and \mathcal{S}_{∞} integrate to unity, the resulting correlation is

$$C(\mathbf{q}) = (1 - \lambda) + \lambda \int d^3r' \mathcal{S}_{\text{local}}(\mathbf{r}') [|\phi(\mathbf{q}', \mathbf{r}')|^2 - 1]. \quad (22)$$

If the experimental sample includes a contamination from weak decays, η , or mis-identified particles of a fraction f , the lambda factor becomes $(1 - f)^2$. It is not uncommon for this contamination factor to be near 30% which results in λ near one half.

Certainly, this division is somewhat artificial as there are non-Gaussian tails, or halos (76), to $\mathcal{S}_{\text{local}}$ due to such causes as the exponential fall-off of the source function in the longitudinal direction or semi-long-lived resonances such as the ϕ whose lifetime is 40 fm/c. Non-Gaussian behavior is a subject for either imaging (48; 49; 50; 51; 52; 53), or for more complicated parameterizations.

The λ factor is often referred to as an incoherence factor, the name being motivated by the properties of a coherent state, $\exp\{i \int d\mathbf{p} \eta(\mathbf{p}) [a(\mathbf{p}) + a^\dagger(\mathbf{p})]\} |0\rangle$, which for identical particles leads to no correlation. Coherent states represent highly correlated emissions due to phase coherence and thus violate the assumption of incoherence or uncorrelated emission implied by Assumption 2 early in this section. Whether an observation of $\lambda \neq 1$ is due to coherence or due to contamination from particles from far outside the source volume can be tested by analyzing three-particle correlations (77; 78). Such analyses of data at both SPS and RHIC have been consistent with the incoherence conjecture (79; 80; 81). Microscopic model calculations at the AGS reproduce the excitation function of λ when resonances contributions are included (82).

2.7 Collective Flow and Blast-Wave Models

Both longitudinal and radial collective expansion reduce the size of the “region of homogeneity”, i.e., the relevant volume for particles of a given velocity. For an infinite volume, the size of this region is set by the length one must move before collective velocity overcomes the thermal velocity, $R \sim V_{\text{therm}}/(dv/dz)$.

2.7.1 LONGITUDINAL FLOW The Gaussian lifetime $\Delta\tau$ described in Equation 18 represents the spread of the emission times. A small value of $\Delta\tau$ does not imply that particles were emitted early, but that they were emitted suddenly. Inferring the mean time at which particles are emitted requires a different assumption. For instance, at RHIC, the initial nuclei are Lorentz contracted by a factor of 100, and if there were no subsequent expansion, R_{long} would be less than a Fermi. If one assumes that the system expands along the beam axis with no longitudinal acceleration, the collective velocity becomes

$$V_{\text{coll},z} = z/t. \quad (23)$$

If emission then comes from sources moving over a large range of rapidities (a boost-invariant expansion), the dimension along the beam axis for the source emitting zero-rapidity particles is determined by the distance one can move be-

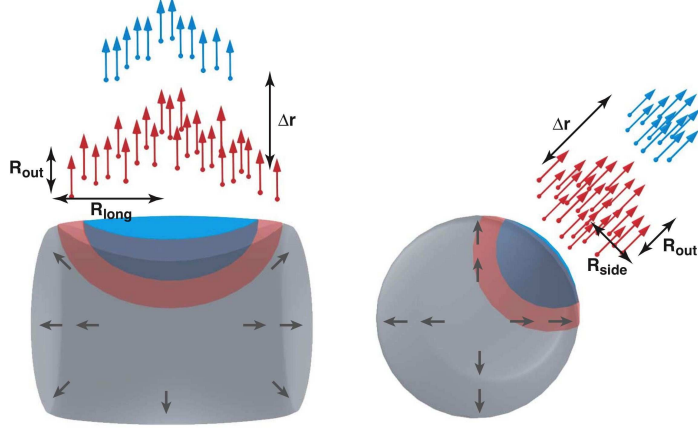


Figure 3: Since particles with heavier masses have smaller thermal velocities their source volumes are more strongly confined by collective flow. For longitudinal flow (left panel) this results in smaller values of R_{long} for particles with higher $m_T = \sqrt{m^2 + p_T^2}$. For radial flow (right panel) this confines heavier particles towards the surface which results in both a reduced volume and an offset Δr in the outward direction.

fore the collective velocity overwhelms the thermal velocity to force the emission function back to zero. The size can then be expressed as:

$$R_{\text{long}} \approx \frac{V_{\text{therm}}}{dv/dz} = V_{\text{therm}} \langle t \rangle. \quad (24)$$

Whereas $R_{\text{out}}/R_{\text{side}}$ gives information about the suddenness of emission, R_{long} provides insight into the mean time at which emission occurs given an estimate of the thermal velocity.

For a thermal source with relativistic motion, the thermal velocity along the beam axis is determined by the temperature and the transverse mass, $m_T = \sqrt{m^2 + p_T^2}$ (59). For large m_T the thermal velocity in the longitudinal direction becomes non-relativistic, $V_{\text{therm}} = \sqrt{T/m_T}$, and the source size falls as $1/\sqrt{m_T}$ which is referred to as m_T scaling (83). This is illustrated in Figure 3. However, this assumes all particles are emitted with the same Bjorken time τ_B and temperature, independent of the transverse mass. Since particles with high m_T are probably emitted at lower τ_B , and since the temperature roughly behaves at $\tau_B^{-4/3}$, the longitudinal size could fall even more quickly than $m_T^{-1/2}$.

In a boost invariant expansion, emission is a function of the Bjorken time $\tau_B = \sqrt{t^2 - z^2}$, not the time t , and since $t = \sqrt{\tau_B^2 + z^2}$, those particles emitted with small z have a head start. This is sometimes referred to as an inside-outside cascade. The transverse shape of $\mathcal{S}(\mathbf{r})$ is then affected non-trivially by the expansion along the beam. The resulting correlation function can be calculated analytically in the case of pure identical-particle correlations (84; 85).

As will be seen in the next subsection, boost invariance is also incorporated into blast-wave models with transverse expansion. Boost invariance is also assumed for many hydrodynamic models. The finite size of the system would alter the

results for two reasons. First, if the distribution of sources covers only a finite range in η , the tails of the distribution $\mathcal{S}(\mathbf{r})$ are cropped off. Assuming the distribution in η is Gaussian rather than uniform,

$$\frac{1}{R_{\text{long}}^2} \sim \frac{1}{V_{\text{therm}}^2 \tau_B^2} + \frac{1}{\eta_G^2 \tau_B^2}, \quad (25)$$

where η_G is the range of rapidities over which the sources are distributed. If η_G were 1.5 units of rapidity, the extracted values of τ_B from boost-invariant pictures would be underestimating τ_B by $\sim 10\%$.

A second shortcoming of boost-invariant models is that they ignore acceleration in the longitudinal direction. Accounting for this acceleration would alter the relation between the time and the velocity gradient. Neglecting this acceleration could also lead to a modest underestimate of τ_B .

2.7.2 TRANSVERSE COLLECTIVE FLOW Since transverse collective flow is intimately related to the pressure and viscosity, it is of central importance. Blast-wave models are based on pictures of thermal sources super-imposed onto the transverse and longitudinal collective velocity profiles. Simple forms are then chosen for the profiles. Only two parameters are important for analyzing spectra, the temperature and the transverse velocity. Since heavier particles are more sensitive to flow than are light particles, the two parameters can be adjusted to fit the spectra of several species.

In addition to the temperature and transverse velocity, correlation measurements are also sensitive to the space-time parameters of the blast wave. In a minimal parameterization this would include the lifetime τ_B and the transverse size R . More sophisticated models would also include a spread in lifetime $\Delta\tau$ and a surface diffuseness ΔR . Additional parameters ensue when one considers sensitivity to the reaction plane. Then, two parameters are needed to describe the transverse size, and two parameters are required to describe the transverse collective velocity. These parameters can then depend on the azimuthal direction of \mathbf{P} (86).

Choosing a blast-wave parameterization involves a number of choices with regards to the form of the parameterization (87). Chemical potentials and temperatures might be chosen to vary with the transverse position r (88; 89) or might be chosen to be uniform. A wide variety of parameterizations have been employed for the transverse velocity profile, which might choose linear profiles for either v , γv or the transverse rapidity $asinh(\gamma v)$. In some parameterizations, the velocity profile has been chosen to rise quadratically with r (90; 91). Although hydrodynamics has been invoked as justification for different parameterizations, profiles from hydrodynamics vary according to the equation of state.

For particles moving much faster than the surface velocity transverse flow manifests itself by constraining particles to an increasingly small fraction of the blast-wave volume for the same reason that R_{long} falls with m_T due to longitudinal expansion (92). For large m_T , this leads to both R_{side} and R_{out} falling as $1/\sqrt{m_T}$ (88; 89; 93). The fact that transverse dimensions fall with m_T might also result from the dynamics of cooling, superimposed with a growing fireball. This correlates high-energy particles with earlier times when the fireball was both smaller and at a higher temperature.

Non-identical particles are of special interest in a blast-wave. For particles moving faster than the surface of the blast wave, there is a stronger tendency

for heavier particles to be more confined to the region of the surface due to their slower thermal velocities (68). This results in heavy particles being ahead of lighter particles of the same asymptotic velocity, and leads to a non-zero $\Delta\mathbf{r}$ illustrated in Figure 3. As discussed in Sec. 2.3, these displacements are accessible through measuring odd components of the correlation function (56; 57; 58).

2.8 Generating Correlations Functions from Hydrodynamics and from Microscopic Simulations

Any model that predicts final-state space-time and momentum information of emitted particles can be used to predict correlation functions. This information may be extracted from both microscopic simulations or from hydrodynamic calculations.

Microscopic simulations model the collision by evolving particles along straight-line classical trajectories which are punctuated by collisions that are programmed to be consistent with free-space cross sections. When the modeling is done on a one-to-one basis, the simulations are referred to as cascades. Boltzmann simulations are similar but employ an over-sampling by a factor N_s accompanied by a scaling down of the cross sections by the same factor. These are then consistent with the Boltzmann equation and become local and relativistically covariant in the large N_s limit (94; 95). To generate correlation functions from either class of simulation, there are essentially two methods which are equally justified within the smoothness approximation. Method I is motivated by Equation 3. This involves first creating two lists, one for each species, of the space time coordinates x_a and x_b of all those particles that were emitted with momenta $m_a\mathbf{P}/(m_a + m_b)$ and $m_b\mathbf{P}/(m_a + m_b)$. From these lists, one generates $\mathcal{S}_{\mathbf{P}}(\mathbf{r})$ by sampling the distributions of $x_a - x_b$. This list is then convoluted with $|\phi(\mathbf{q}, \mathbf{r})|^2$ to generate $C(\mathbf{P}, \mathbf{q})$ for all \mathbf{q} . In Method II, one samples pairs randomly without regard to their momenta. The numerator of the correlation function is then calculated by generating pairs with the same weight as one expect to observe experimentally, and applying a weight given by the square of the relative wave function. The denominator would be calculated in a similar manner, but without the weight from the wave function. This method reflects the description of the correlation function in Equation 2. Acceptance effects or kinematic cuts can then be performed exactly as they would be performed for real particles. Method II has an advantage in that it is easier to accurately incorporate acceptance effects or tight kinematic cuts. Method I makes for a much quicker calculation since the procedure does not require sampling particles for irrelevant momenta (38).

Given the equation of state and the initial energy density, hydrodynamics provides the means for solving for the space-time development of the stress-energy tensor which can be used to make predictions for correlations (96; 97; 98). Viscous effects can also be incorporated and are non-negligible (99; 100). Generating source functions from the output of hydrodynamic calculations is not as straightforward as it might seem. The Cooper-Frye prescription (101) conserves energy and momentum if the equation of state is one of free particles, but suffers from the fact that the particles that cross backwards across the surface into the hydrodynamic volume enter the source function as a negative emission probability. If the relative velocity of the surface, as measured by an observer in the matter's rest frame, is not much faster than the thermal velocity, a different prescription is required. Numerous prescriptions have been proposed to address these issues

(102; 103; 104).

Hydrodynamic models, even those that incorporate viscosity, can not be justified once the system expands to the point that the mean free path is similar to the characteristic size of the system. However, Boltzmann descriptions or cascades are well justified at lower densities. Several efforts have thus focused on coupling the two approaches (105; 106; 107; 108). Since the final-state trajectories are established in the Boltzmann part of the prescription, one can apply either of the methods mentioned above.

2.9 Phase Space Density, Entropy and Coalescence

Since phase space density depends on both the momentum \mathbf{p} and the position \mathbf{r} , a measurement of the phase space density must specify the spatial region over which it is determined. In practice, spatial information from two-particle correlation functions is instrumental to this end. For identical particle pairs, the “area” under the correlation function determines the average phase space density (109). Substituting the final phase space density for the time-integrated source function,

$$\int^{t_f} \frac{dx_0}{E_p(2\pi)^3} s(p, x)|_{p_0=E_p} = f(\mathbf{p}, \mathbf{x}, t_f), \quad (26)$$

and inserting into Equation 3 with $Q = 2q$ leads to

$$\begin{aligned} \int d^3Q [C(\mathbf{Q}) - 1] &= (2\pi)^3 \frac{\int d^3r |f(\mathbf{P}/2, \mathbf{r}, t_f)|^2}{[\int d^3r f(\mathbf{P}/2, \mathbf{r}, t_f)]^2} = \frac{\bar{f}(\mathbf{P}/2)}{dN/d^3p}, \quad (27) \\ \bar{f}(\mathbf{p}) &= \frac{(2\pi)^3}{(2S+1)} \frac{E_p}{m} \frac{dN}{d^3p} \mathcal{S}_{\mathbf{P}=2\mathbf{p}}(\mathbf{r}' = 0). \end{aligned}$$

Equation 28 applies also for the case of non-identical particles of the same phase space density. We note that $\bar{f}(\mathbf{p})$ is the phase space density averaged over coordinate space for a specific momentum using the phase space density itself as the weight. Unless $f(\mathbf{p}, \mathbf{r}, t)$ is a constant within a fixed volume, the average phase space density will fall below the maximum phase space density (102). For instance, if $f(\mathbf{p}, \mathbf{r}, t)$ has a Gaussian profile in coordinate space, the average phase space density will be $2^{-3/2}$ of the maximum phase space density for that momentum. For a Gaussian source,

$$\bar{f}(\mathbf{p}) = \frac{\pi^{3/2}}{(2S+1)mR_{\text{inv}}^3} E \frac{dN}{d^3p}, \quad (28)$$

where $R_{\text{inv}}^3 = (E/m)R_{\text{out}}R_{\text{side}}R_{\text{long}}$ is the product of the three radii as measured in the frame of the pair. The phase space density is determined by combining a source size measurement with the spectra. Entropy can be related to the phase space density in the standard way (110)

$$S = (2S+1) \int \frac{d^3r d^3p}{(2\pi)^3} [-f \ln f \pm (1 \pm f) \ln(1 \pm f)], \quad (29)$$

$$dS/dy \approx \int d^2p_T E \frac{dN}{d^3p} \left[\frac{5}{2} - \ln(2^{3/2} \bar{f}(\mathbf{p})) \pm \bar{f}(\mathbf{p})/2 \right]. \quad (30)$$

Here, Equation 30 ignores higher powers of \bar{f} .

The average phase space density is also straight-forward to determine by constructing ratios of spectra with species that can either bind or form a resonance. If species a and b can bind to form species c , thermal arguments would state

$$f_c(\mathbf{r}, t) = f_a(\mathbf{P}m_a/m_c, \mathbf{r}, t)f_b(\mathbf{P}m_b/m_c, \mathbf{r}, t)e^{B/T}, \quad (31)$$

where B is the binding energy, or the excitation energy if the resonance is unstable. Coalescence arguments, which give the same expression but without the binding energy (111; 112), are identical if the binding energy is small compared to the temperature as is the case for nucleon coalescence. The average phase space density for a or b can be determined by inserting Equation 31 into Equation 28,

$$\bar{f}_{a,b}(\mathbf{p}) = e^{-B/T} \frac{m_{b,a}(2S_{b,a} + 1)}{m_c(2S_c + 1)} \frac{E_c dN_c/d^3P}{E_{b,a} dN_{b,a}/d^3p}. \quad (32)$$

Here, the binding energy needs to be expressed in the frame of the thermal bath. For the case where B is small the assumption of a thermal bath can be neglected. Two examples where particles with similar phase-space densities form low-energy resonances or bound states are $pn \rightarrow d$ and $\phi \rightarrow K^+K^-$.

By comparing the expression for \bar{f} from the ratio of spectra in Equation 32, with Equations 28 or 28, one can determine either $\mathcal{S}_{\mathbf{P}}(\mathbf{r}' = 0)$ or the Gaussian parameters from ratios of spectra.

$$\mathcal{S}_{\mathbf{P}}(\mathbf{r}' = 0) = e^{-B/T} \frac{m_a m_b (2S_a + 1)(2S_b + 1)}{(2\pi)^3 m_c (2S_c + 1)} \frac{E_c dN_c/d^3P}{E_a dN_a/d^3p_a \cdot dN/d^3p_b}, \quad (33)$$

$$R_{\text{inv}}^3(\mathbf{P}) = \frac{1}{(4\pi)^{3/2} \mathcal{S}_{\mathbf{P}}(\mathbf{r}' = 0)}. \quad (34)$$

Coalescence analyses can provide powerful measurements of volumes, but they only provide a single number, $\mathcal{S}(\mathbf{r}' = 0)$, and cannot provide any insight into either the shape or the r dependence.

3 Femtoscopic measurements

Experimental techniques have developed considerably in response to significant improvements in both the theory and the quantity and quality of experimental data. In this Section we discuss the general experimental approach for defining and analyzing femtoscopic correlations and their systematic dependence on global and kinematic quantities.

3.1 Correlation Function Definition

In practice, the formal definition of the correlation function in Equation 1 is seldom used in heavy ion physics. Instead the correlation of two particles, a and b for a given pair momentum \mathbf{P} and relative momentum \mathbf{q} is nominally given by,

$$C_{\mathbf{P}}^{ab}(\mathbf{q}) = \frac{A_{\mathbf{P}}^{ab}(\mathbf{q})}{B_{\mathbf{P}}^{ab}(\mathbf{q})} \cdot \xi_{\mathbf{P}}(\mathbf{q}). \quad (35)$$

where $A_{\mathbf{P}}^{ab}(\mathbf{q})$ is the signal distribution, $B_{\mathbf{P}}^{ab}(\mathbf{q})$ is the reference or background distribution which is ideally similar to A in all respects except for the presence of femtoscopic correlations, and $\xi_{\mathbf{P}}(\mathbf{q})$ is a correction factor introduced to compensate for non-femtoscopic correlations present in the signal that are not fully accounted for in the background as well as artifacts resulting e.g. from finite resolution and contamination.

3.2 Signal Construction

The signal $A_{\mathbf{P}}^{ab}(\mathbf{q})$ refers to the relative momentum distribution of particles a and b for a given range of pair momenta, \mathbf{P} , and a given set of event characterizations. Although not all analyses proceed in exactly the following fashion, the mechanics of constructing the signal and background are most easily understood if one considers as separate steps,

1. Event quality cuts and event-class binning,
2. Single-track (including particle identification) cuts and single-particle binning,
3. 2-particle pairing, 2-track cuts and pair momenta binning.

Here “event class” refers to both physics observables such as collision centrality and reaction plane orientation, and detector considerations such as event vertex position and the condition of the detector when the event was recorded, usually keyed by run number. The latter considerations are relevant to the proper construction of the background. For event-class, particle, and pair bin, the final signal is usually stored as a set of 3D-histograms in the canonical relative momentum variables.

Single-particle acceptances divide out with a properly constructed background, but 2-track acceptances can have a large effect on the correlation function. For this reason the analysis of 2-track cuts is the dominant consideration in the signal construction for most analyses. The cuts and terminology are different for TPC experiments with near continuous hit distributions and Drift Chamber experiments with projective geometry, but the goals are the same. Split-tracks (113; 114) and ghost-tracks both refer to single tracks which are incorrectly reconstructed as a pair of tracks with very low relative momenta. Even after the event-reconstruction algorithms (which generate a list of individual tracks) have been optimally tuned, small traces of these false pairs remain and must be removed from the analysis with identical *pairwise* cuts. Usually only a tiny fraction of tracks are split, and this effect may be ignored in essentially every physics *except* femtoscopic ones. Various methods have been developed for identifying likely split tracks, usually based on the number (114) or topology (113) of space-points associated with the track.

Pairwise effects usually also result in the *loss* of pairs at low relative momentum, since two tracks with very similar trajectories tend to be reconstructed as a single track. (Note that in tracking detectors, this is not a problem if one or both of the particles is a topologically-identified neutral particle. In that case, the decay daughters may be well-separated even if their parents have identical momentum.) Such merging issues are usually resolved by pairwise cuts which remove merged pairs. Developing efficient and appropriate cuts can be a subtle exercise, and requires good knowledge of the detector and event reconstruction software. In

most cases these cuts are supported by simulations, but final determinations are nearly always based upon data.

The question arises— what does it mean to cut out merged pairs? After all, if the tracks have merged, then the pair is lost anyway. The point is twofold. Firstly, the pair efficiency usually does not drop from 100% to 0% sharply as a function of any variable. Thus, the cuts are usually tuned to exclude all (82; 113) or most (115; 116; 117) of the inefficient region. If regions with less than perfect efficiency remain in the analysis, a two-track efficiency correction based on Monte Carlo simulations must be applied, typically leading to systematic uncertainties of a few percent. The second reason for the cut is that it is applied equally to the signal and to the background distribution $B_{\mathbf{P}}^{ab}(\mathbf{q})$ (118). Thus, if some fraction f of pairs is lost at some relative momentum \mathbf{q} in $A_{\mathbf{P}}^{ab}(\mathbf{q})$, the same fraction is lost in $B_{\mathbf{P}}^{ab}(\mathbf{q})$, and the ratio in Equation 35 is robust against the effect.

3.3 Background Construction

For reasons just described, *all* cut-imposed effects on the signal pair distribution A^{ab} must be applied to B^{ab} . This often means identifying which pairs “would have” been removed by merging, splitting, or other cuts, had the particles come from in the same event.

The ideal background should be identical to the signal in all respects except for the presence of femtoscopic correlations. Therefore, the global event characteristics, single particle distributions and acceptances should match those of the signal. A simple and straightforward way to construct such a background is to form pairs from different events within a single event class. This event-mixing technique (119) has gained wide acceptance in relativistic heavy-ion collisions where violations to energy-momentum conservation are negligible in the high multiplicity environment. This technique will be described in detail in what follows. However, other methods have also been used, especially if one considers femtoscopy in other systems.

For elementary-particle collisions or in low-multiplicity events, event mixing can violate total energy momentum conservation, especially when exclusive final states or jet-axes must be preserved; thus, the correlation function would reflect non-femtoscopic in addition to femtoscopic correlations. In these cases, the most common techniques form a background from unlike-signed pairs, with resonance regions excluded with cuts (120) or normalized with a correlation of like- to unlike-signed pairs from a Monte Carlo (121). Other experiments have constructed a background using only Monte Carlo generated pairs (122). A few experiments have investigated backgrounds formed by swapping (123) or reversing momentum components relative to a jet-axis (120), but these methods are not widely used. For detectors with symmetrical acceptance, such as the STAR TPC (124), momentum conservation effects may be eliminated by mixing pairs from the same event, with the lab momentum of one particle flipped (125). Backgrounds constructed from single-particle distributions as formally defined by Equation 1 have been used for heavy ion collisions at lower energies and shown to be consistent with the more commonly used event-mixing technique (126).

In order to avoid inducing artificial structure in the correlation function, the particles forming pairs in the background distribution should originate from “parent” events with the same event characteristics. The parent events should have similar vertex positions to within the experimental resolution. Because detector

acceptances can vary with time (e.g. components may fail for some runs) parent events should have been measured close in time to each other; this is usually easiest in any case, since event mixing is done “on the fly” as time-ordered data is read sequentially.

Parent events whose particles are mixed should also have the same single-particle momentum distributions. Thus, they should have similar centralities and orientations of the reaction plane. For example, mixing particles from events with very different p_T slopes or directions of preferred emission (elliptic flow) would produce differences between $A^{ab}(\mathbf{q})$ and $B^{ab}(\mathbf{q})$ even in the absence of physical correlations. Since almost all analyses to date have ignored these potential biases, it is comforting that accounting for them makes little difference in practice (113).

The list of event classes given here is by no means exhaustive. One can expect future analyses to incorporate the orientation of high p_T particles (jet axis) or any other event-related observable.

The procedure for deciding how many events to mix remains something of an art and involves optimizing over the range of data runs, bin width, and statistics. In order to minimize statistical errors, one typically forms approximately ten times the number of pairs in the background as in the signal. For the special case when all possible combinations are formed the variance of a particular relative momentum bin is proportional to $n^{\frac{3}{4}}$, where n is the number of entries in the bin (20). However, as the number of pairs formed is reduced, the variance per bin approaches the $n^{\frac{1}{2}}$ value expected for Poisson statistics (116). It is possible that non-Poisson fluctuations persist in the co-variance between different bins, but this has not yet been investigated.

Once the pairs have been mixed, the background must be subject to the same 2-track cuts that have been applied to the signal. For example, the exact same track merging cuts or minimum separation on a detector must be applied to both signal and background.

3.4 Corrections

Corrections to the correlation function fall into three categories: finite resolution effects, mis-identified particle contamination, and compensation for deficiencies in the background.

The first category concerns single-track momentum resolution, and reaction-plane resolution. We consider finite momentum resolution corrections first. Typically, momentum resolutions are on the order of 1%. One approach is to correct for momentum resolution by a double ratio of the ideal correlation function generated from a Monte Carlo simulation with perfect momentum resolution divided by a Monte Carlo correlation function with momentum resolution turned on. The femtoscopic weights are inserted into the simulations iteratively until the fitted radii converge (118; 114; 82; 113). A second approach is similar, but corrects only the Coulomb interaction term, which is most greatly affected by momentum resolution effects (117). In both cases, the corrections change the fitted radii by only $\lesssim 5\%$.

As discussed in Section 4.2, azimuthally-sensitive analyses (127; 128; 129) measure oscillations in correlations as a function of emission angle with respect to the reaction plane. Finite resolution effects in the reaction plane angle (130) artificially reduce the oscillation strengths. Methods have been developed to correct the distributions $A^{ab}(\mathbf{q})$ and $B^{ab}(\mathbf{q})$ for these effects (64; 131).

Another type of correction accounts for the inclusion of mis-identified and secondary-particle contamination. For example, electrons may be mistakenly identified as π^- mesons. It is usually assumed that the mis-identified or secondary particles are uncorrelated with other particles, so the net effect on the correlation function is to damp all structure uniformly in \mathbf{q} . For purely Gaussian correlations (c.f. Section 3.5), this effect is absorbed wholly into the λ factor discussed in Section 2.6; homogeneity lengths— derived from the width of the Gaussian correlation— are unaffected by the reduction in its strength. In many cases, however, the homogeneity length is extracted largely from the strength of the correlation, and so contamination effects must be removed. In the general case, for which the purity ρ depends on the relative momentum, the correlation function is be corrected according to $C^{\text{true}}(\mathbf{q}) = (C^{\text{raw}}(\mathbf{q}) - 1)/\rho(\mathbf{q}) + 1$ (58; 113).

It is more difficult to corrected for *correlated* contamination. For example, if cuts cannot completely distinguish primary protons from those coming from Λ decay, then measured $p - \Lambda$ correlations will contain contributions from $\Lambda - \Lambda$ correlations. Unlike the “white noise” contamination discussed above, this introduces structure into $C^{p\Lambda}(\mathbf{q})$ which would require detailed simulations to account for. Such corrections will become more important at RHIC due to copious resonance production, and especially for baryon correlation measurements, in which the heavy daughter carries most of the momentum of the parent resonance.

The last category of corrections are applied to fix deficiencies in the background distribution. This includes corrections is to account for two-particle inefficiencies, which have been discussed in the previous Section. A second correction of this type deals with the residual signal correlation that is present in *all* backgrounds derived from events that contribute to the signal. The “residual” correlation arises because femtoscopic correlations can modify the single particle distributions. This is especially true for small-aperture spectrometers. This effect can be removed with an iterative procedure (20; 132), however, for many large experiments the induced error is often 1% or less, and it is easier to fold this into the systematic errors (117).

3.5 Fitting

After the application of all cuts and corrections, the correlation is formed according to Equation 35 and then fit to determine spatial parameters. As described in Section 2, there are three approaches to fitting the correlation function: fitting to a simplified Gaussian form with strong and Coulomb interactions neglected or factored out, fitting to a convolution of the full kernel convoluted to a parameterization of $\mathcal{S}(\mathbf{r})$, and inverting the kernel to fit a source image. The simplified Gaussian fits based on Equation 19 are limited to correlations of identical pions, kaons, and photons, but it has been the most widely used method to date because of the computational demands of the other methods. We expect its use to continue for the large systematic studies in which binning in centrality, reaction plane, and k_T leads to fits of more than one hundred separate correlation functions for a single colliding system.

However, the functional form of the Gaussian parameterization used by experimentalists has evolved over the years. Before reviewing the most recent functional forms, it is necessary to review the treatment of the Coulomb interaction, and the fraction of pairs coming from the source that contribute to the femtoscopic correlations. Both were first introduced into the literature by Zajc (20) in the

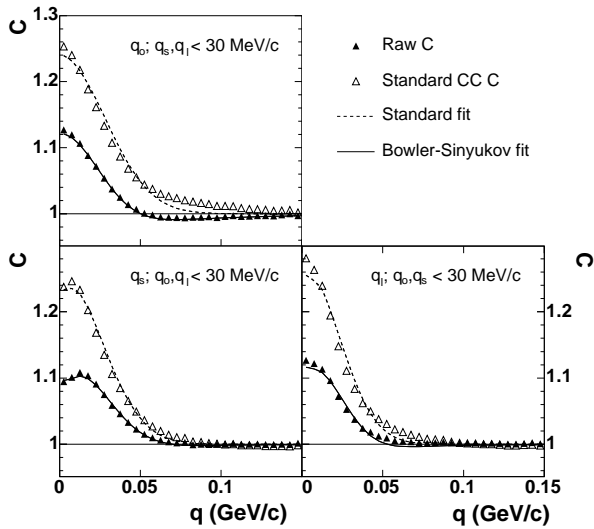


Figure 4: Projections of a 3-dimensional correlation function (integrated over 0-30 MeV/c in orthogonal components) for low- k_T π^- pairs for 200 GeV central Au+Au collisions (113) (filled symbols) with fit function. Open symbols include correction for Coulomb interaction among all pairs. Projections generated according to the prescription described in (135).

form of the Gamow factor given in Section 2 and an empirical parameter λ to account for the observation that not all pairs exhibit femtoscopic correlations. With steady improvements in data quality and CPU speed the Gamow factor has been replaced with a calculation of an squared unsymmetrized Coulomb wave for a finite Gaussian source. The improvements in data quality have also led to a self-consistent treatment of λ with respect to both Coulomb and Gaussian components of the fit function (133; 134). For this to be accurate we must assume that the source is fully chaotic, an assumption that has recently been verified with three-pion correlations (81; 77). The non-femtoscopic pairs consist of misidentified particles and particles that emanate from too far from the source for the correlation to be resolved experimentally. The region far from the source has been referred to the source halo, to differentiate it from the core. The correlation fit function is therefore given by Equation 36,

$$C(\mathbf{q}) = N [\lambda G(\mathbf{q})F(\mathbf{q}) + (1 - \lambda)], \quad (36)$$

where N is the overall normalization, F is the Coulomb component, and G is the Gaussian form for the un-damped correlation function, Equation 19 for out-long-side coordinates, or Equation 21 for Yano-Koonin variables.

Fig. 4 shows projections of a $\pi^- - \pi^-$ correlation function measured by the STAR collaboration (113). The filled symbols are the measured correlation function corrected for momentum resolution only and fit with Equation 36. The open symbols have been over-corrected by applying to all pairs the Coulomb correction for the fitted source dimensions. Depending on the shape of the correlation and degree of experimental contamination, extracted homogeneity lengths may vary by up to $\sim 15\%$ if the correlation function is over-corrected.

For proton-proton correlations and non-identical particle correlations direct fits are performed by convoluting the full kernel with a parameterized source. For these analyses, the paucity of statistics has been more of a limitation than the relatively modest demands in CPU power. The examples given in Section 4 are all for one-dimensional analyses, but recent data from RHIC will soon be analyzed

in multi-dimensions.

The ability to image the source by inverting the kernel is a relatively recent development, but one with very general applications. Because the source is parameterized by a series of B-splines, it is a very general form which is sensitive to non-Gaussian shapes. To date source imaging has been performed only with one-dimensional correlations, but like with the direct fits, a multi-dimensional kernel will soon be possible (136; 65).

Non-Gaussian effects were reported already in the first pion correlation measurement at RHIC (137). With higher statistics, the STAR Collaboration has studied the issue in greater detail (113), performing a functional expansion (the so-called Edgeworth expansion (138)) about a Gaussian shape. While significant non-Gaussian contributions were reported, the dominant length scales were already extracted in the purely Gaussian fits.

3.5.1 MINIMIZATION A simple chi-squared test is inappropriate for fitting correlation functions because the ratio of two Poisson distributions is not itself Poisson distributed, especially when taking the ratio of small numbers. For this reason a log-likelihood fit function of the form given in Equation 37 is preferred.

$$\chi^2_{PML} = -2 \left[A \ln \left(\frac{C(A+B)}{A(C+1)} \right) + B \ln \left(\frac{A+B}{B(C+1)} \right) \right]. \quad (37)$$

This equation derived from the principle of maximum likelihood assuming that both signal and background are Poisson distributed (116). The full derivation of Equation 37 and comparison to earlier log-likelihood functions is given in (116).

4 Measured Femtoscopic Systematics

The first systematic study to compare femtoscopic measurements across several systems and experiments was performed almost 20 years ago with data from intermediate-energy heavy ion-collisions at the Bevalac (139). The data, taken with from experiments with different acceptances, triggers, and analysis techniques, were sufficient to demonstrate a crude $A^{1/3}$ scaling of the 1-dimensional radii indicating that spatial scales were indeed being probed. The first femtoscopic measurements for relativistic heavy ion collisions were presented by the NA35 Collaboration at the Quark Matter meeting in Nordkirchen (140; 141). More detailed measurements followed with the availability of sulphur and silicon beams at the SPS (118) and AGS (142; 143).

Since then, increasingly sophisticated experiments at the AGS, SPS, and RHIC have performed femtoscopic measurements corresponding to a wide range of control parameters. The experimental community performing the measurements has reached critical mass and matured substantially; a common language and knowledge base has developed concerning sometimes subtle details in performing and interpreting femtoscopic measurements. The result of this effort is a striking degree of consistency across experiments in regions of phase space where acceptances overlap, and meaningful generation of systematics across experiments. Large-statistics data sets routinely allow three-dimensional correlation measurements with small statistical error bars. Systematic errors, which now dominate the experimental errors, have been reduced to the level of $\sim 5\%$, or ~ 0.25 fm for most measurements. It is no exaggerations to state that femtoscopic measurements have become a precision tool.

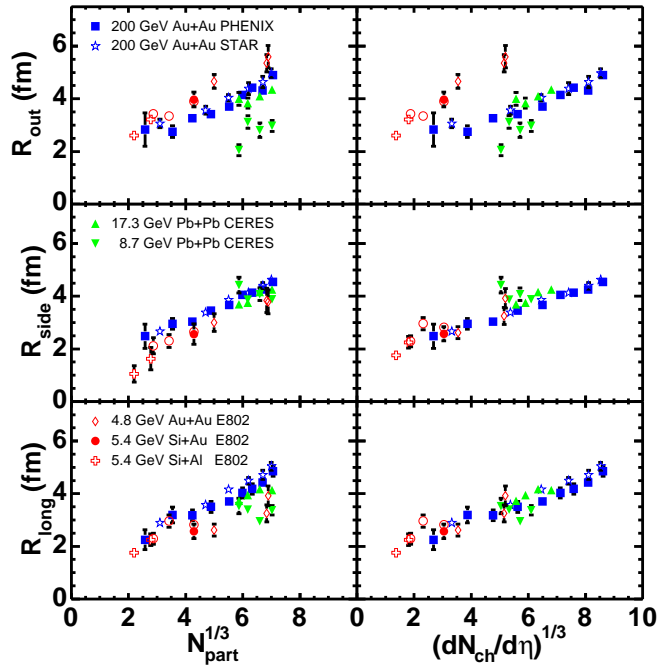


Figure 5: Pion source radius dependence on number of participants (left) and on charged particle multiplicity (right). Data are for Au+Au (Pb+Pb) collisions at several values of $\sqrt{s_{NN}}$, and also for Si+A collisions at the lowest energy. Average transverse momentum $\langle k_T \rangle \sim 450$ MeV/c for the RHIC data and ~ 390 MeV/c at the lower energies. Data from (116; 144; 117; 113).

Here, we cover the most important systematics of femtoscopic measurements from the AGS, SPS, and RHIC. We discuss only generally the physics probed by a given systematic, appealing to intuitive schematic models such as the blast wave (68). Full interpretations and comparisons to dynamic models are given in Section 5

4.1 System size: N_{part} and Multiplicity

As discussed earlier, femtoscopic radii probe homogeneity regions, and not the entire source (hereafter, the term “source” will be used to refer to the entire source of particle emission). Nevertheless, the claim that two-particle correlations probe spatial scales would be given little credence if the radii did not exhibit a strong, positive correlation with system size. Therefore, measuring the systematic variation of the radii vs. system composition and centrality represents the most basic test of both theoretical and experimental femtoscopic techniques.

Coalescence studies (145) and two-proton measurements at the AGS (146) and SPS (147) unambiguously demonstrate that nucleon homogeneity lengths increase with decreasing impact parameter and/or increasing projectile mass, continuing the trend mapped at lower energies (148; 149), where directional cuts have allowed measurement of the shape of the homogeneity region (150; 151; 152). More detailed information comes from pion correlations at relativistic energies, for which

three-dimensional analyses allow partial isolation of purely geometrical effects. The centrality dependence of Bertsch-Pratt source radii are shown in Figure 5 for a wide range of collision energies. The left panels show the dependence on the number of participating nucleons, N_{part} , a generalization of the $A^{1/3}$ linear scaling of nuclear radii used to approximate the initial overlap geometry. All of the radii exhibit a linear scaling in $N_{\text{part}}^{1/3}$, most with finite intercepts. It is possible to compare linear slopes for different energy ranges, however, the intercepts are difficult to interpret given the ambiguities involved in extrapolating to fewer than two participating nucleons. For R_{out} the linear slopes for all energies are approximately consistent, with the AGS data falling slightly above, and the lower energy SPS data falling well below the RHIC results. Only the slope of the R_{long} dependence shows a significant increase from the AGS to RHIC, consistent with a lifetime that increases with both centrality and $\sqrt{s_{\text{NN}}}$. The trend of increasing R_{long} with increasing $\sqrt{s_{\text{NN}}}$ is reversed for $\sqrt{s_{\text{NN}}} < 5$ GeV (82).

The right panels of Figure 5 show the same radii as a function of $(dN_{\text{ch}}/d\eta)^{1/3}$. The primary motivation for exploring the $(dN_{\text{ch}}/d\eta)^{1/3}$ dependence is its relation to the final state geometry through the density at freeze-out. However, the two scaling quantities are highly correlated. In fact, the values of $dN_{\text{ch}}/d\eta$ shown on the right side of Figure 5 were derived from N_{part} using the N_{part}^α parameterizations given in (153), and conversely, the N_{part} values are often calculated from multiplicity distributions using a Glauber model. Given this caveat, the R_{side} and R_{long} values exhibit a linear dependence on $(dN_{\text{ch}}/d\eta)^{1/3}$, again with finite intercepts. The strong uniformity from $\sqrt{s_{\text{NN}}}$ of 5 to 200 GeV leads one to believe that the N_{part} scaling (initial overlap geometry) is a result of the scaling with multiplicity (final freeze-out geometry) and not the other way around. The R_{out} differences between the AGS and RHIC apparent in the N_{part} scaling are magnified with $dN_{\text{ch}}/d\eta$ scaling, whereas the differences with the lower energy SPS data are compressed.

We note that the systematics in system size represent an initial “sanity check” for the femtoscopic technique. The obvious direct connection of the radii to the source geometry estimated in two ways refutes suggestions (154) that smaller, non-geometric length scales dominate experimentally extracted transverse radii.

4.2 Collision Orientation: Pair emission angle relative to \hat{b}

The variation of femtoscopic radii with the pair emission angle relative to \vec{b} (ϕ_{pair}) can be used to probe the 3-dimensional shape of the source (71; 155; 62; 156; 157; 63; 64; 68). The anisotropic shape transverse to the beam direction—the coordinate-space analog to the elliptic flow characterizing momentum-space—gives rise to $\cos(n\phi_{\text{pair}})$ (n even) oscillations in the squared transverse source radii R_{out}^2 , R_{side}^2 , $R_{\text{out,side}}^2$ (62; 64).

Just as one expects the source (and homogeneity regions) to be *larger* for decreasing $|\vec{b}|$, one also expects it to be *rounder*, reflected by small oscillations of the radii. Figure 6 for mid-rapidity pions from Au+Au collisions at RHIC confirms this expectation. As $|\vec{b}|$ increases, the oscillations indicate a transverse source increasingly elongated out of the reaction plane (129).¹

¹The out-of-plane nature of the elongation may be read directly from Figure 6. Ignoring collective flow or opacity effects (e.g. 63) an out-of-plane-extended Source would produce $R_{\text{side}}^2(\phi_{\text{pair}} = 0^\circ) > R_{\text{side}}^2(\phi_{\text{pair}} = 90^\circ)$, as seen in Figure 6. Collective flow effects complicate

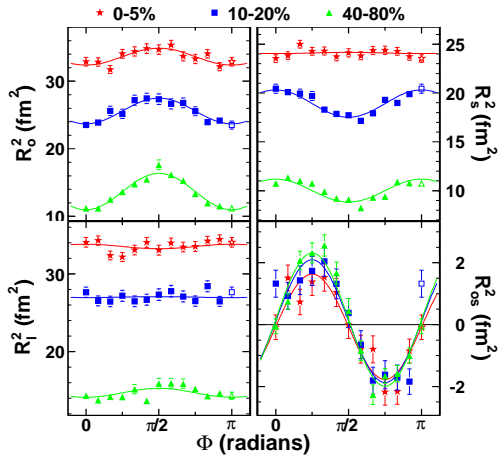


Figure 6: Squared source radii measured at mid-rapidity in Au+Au collisions at RHIC, as a function of pair emission angle relative to the reaction plane. Data for three centralities are shown. Figure from (129).

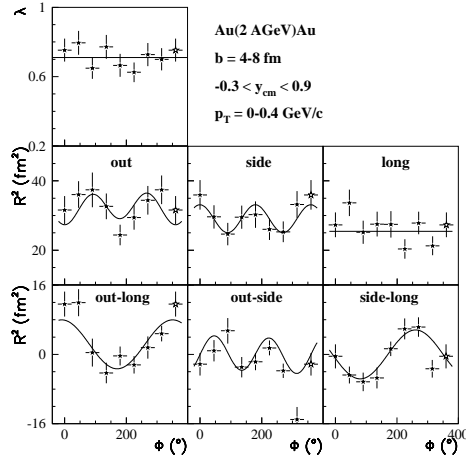


Figure 7: Squared source radii measured near mid-rapidity in mid-central Au+Au collisions at lowest-energy AGS settings, as a function of pair emission angle relative to the reaction plane. Figure from (127).

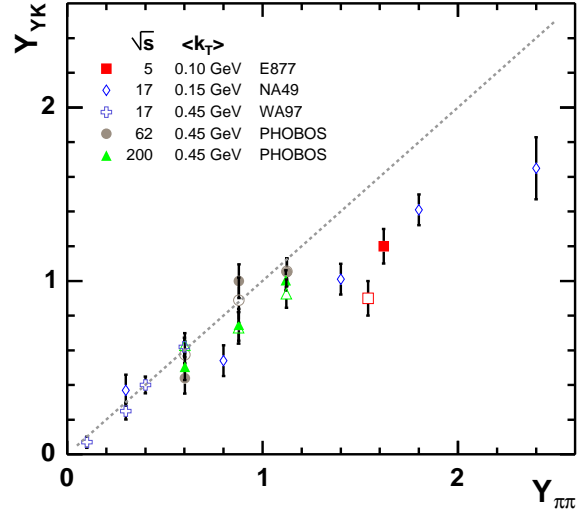
That the strong in-plane expansion (158) does not fully convert the initial out-of-plane (overlap) geometry into an in-plane-extended source at freeze-out, suggests a rather short evolution time; in essence, the system did not have time to reverse its deformation. However, this is only a hint, and a full dynamical transport calculation is required to extract physical timescales (86).

Whereas at the highest RHIC energy, the freeze-out anisotropy is $\sim \frac{1}{3}$ the of the initial (129), at low AGS energies, the final anisotropy is consistent with that of the initial overlap region (127), or perhaps slightly lower. Since elliptic flow “vanishes”—changes sign— at these energies (159), these trends make intuitive sense and suggest an underlying connection to the evolution dynamics. It would be desirable to map the source anisotropy at intermediate (AGS and SPS) energies, for which there may be interesting changes in the space-time systematics. At these energies, there have been intriguing hints of asymmetries in the homogeneity regions for pions (160; 161; 162; 163) and protons (164), and in the proton-pion separation (161; 69), although they have not been finalized.

If the impact parameter direction \hat{b} —not simply the 2nd-order event-plane angle (unambiguous only over a range $[0, \pi]$)—is known, then more detailed information may be obtained. In the left panel of Figure 2, the source is tilted with respect to the beam axis, towards \hat{b} . Just as anisotropic azimuthal geometry in the transverse plane is related to the structure of elliptic flow (86; 96), a tilted geometry can reveal important information on the underlying nature of directed flow (165; 166; 63; 127). The structure in $R_{o,l}^2$ and $R_{s,l}^2$ shown in Figure 7 is generated by this tilt. The spatial tilt has been measured only at low AGS energies (127); a measurement at RHIC might reveal exotic geometric configurations

this picture (62; 68), but the sign of the oscillations are determined by geometric, not dynamic, effects for realistic Sources at RHIC (86; 68).

Figure 8: The Yano-Koonin source rapidity is shown as a function of the pion pair rapidity for central Au(Pb)+Au(Pb) collisions over a broad range of energies (open symbols are for $\pi^-\pi^-$, closed symbols for $\pi^+\pi^+$). Both quantities are in the center-of-mass frame of the colliding system. Data from (171; 172; 173; 174).



generated by QGP formation (167) and would impact the important issue of boost-invariance at mid-rapidity.

4.3 Boost invariance : $Y_{\pi\pi}$

In high-energy hadronic collisions the initial parton distribution is expected to be approximately flat in rapidity. It has been conjectured that this momentum rapidity distribution corresponds to producing matter which initially exhibits a boost-invariant Hubble-type scaling correlation between longitudinal flow velocity and space-time points, $v_{L,flow} = z/t$ (168; 169). Relativistic hydrodynamics preserves boost-invariance of the initial conditions throughout its dynamical evolution (169). The combination of the above arguments underlies expectations that in ultra-relativistic heavy ion collisions, particle production emerges from boost-invariant longitudinal flow, and that dN/dy exhibits an approximately boost-invariant plateau around mid-rapidity. However, an extended plateau has never been observed from AGS through RHIC energies (3).

Since correlation measurements access spatio-temporal information, the question arises (74; 170) whether they allow us to test the relation $v_{L,flow} = z/t$ between the space-time rapidity and the momentum rapidity of the source. For boost-invariant sources, one can show that the pair momentum rapidity is equal to the Yano-Koonin source velocity which is directly obtained from the Gaussian radius parameters, $Y_{\pi\pi} \approx Y_{L,flow} \equiv \tanh^{-1} v_{L,flow}$. However, even if the source density distribution shows significant deviations from boost-invariance, this relation still holds approximately as long as the velocity profile is boost invariant, and k_T is sufficiently large (74).

Figure 8 reveals a roughly universal dependence of Y_{YK} on $Y_{\pi\pi}$ for pions from central collisions, depending weakly if at all on $\sqrt{s_{NN}}$ (c.f. Section 2.5.3). This trend is particularly striking given the very different center-of-mass projectile rapidities (~ 1.55 and 5.5 for $\sqrt{s_{NN}} = 5$ and 200 GeV, respectively) and corresponding widths of the pion distributions dN/dy .

By way of caution, we note that the results from RHIC are limited to the region $Y_{\pi\pi} < 1.2$, and that the deviations from boost-invariance are mostly in

the lower energy data. Extending the RHIC results to more forward rapidities would provide an important test for both the velocity scaling at RHIC and the energy-independence that is exhibited in Figure 8.

For central collisions the roughly universal behavior approximately obeys the boost-invariant consistency relationship discussed above. Moreover, $Y_{L,flow}$ shows a significant k_T dependence and falls below the linear relation in particular for small k_T (173). Qualitatively, this is consistent with blast-wave models in which a boost-invariant longitudinal flow is superimposed on a source density distribution of finite longitudinal width. However, a full dynamical understanding of the dependence is missing so far. However, the flat Y_{YK} dependence on $Y_{\pi\pi}$ measured at the SPS (173) for the most peripheral collisions is counter-intuitive, and requires further study.

The question of whether the source has boost-invariant space-time structure is an important one. There are many reports of very short evolution timescales (“lifetimes”) based on fits to the data with Equation 24, which is based upon an assumption of boost-invariance (83). Relaxation of that assumption might lead to considerably larger estimates (175).

4.4 Collective dynamics: k_T and particle mass

As discussed in Section 2.7.2, the dynamic substructure of the source is encoded in space-momentum ($\mathbf{x}\mathbf{p}$) correlations. Longitudinal $\mathbf{x}\mathbf{p}$ correlations, encoded in $R_{long}(k_T)$, are generally acknowledged (83; 68) to reflect longitudinal flow. Since all transverse correlations are generated in the collision itself, considerably more attention has generally been paid to the transverse substructure than to the longitudinal flow discussed in Section 4.3.

The most common explanation for transverse $\mathbf{x}\mathbf{p}$ correlations is collective transverse flow (92). These correlations have mostly been studied through pion correlations, but transverse flow implies also a systematic trend as the particle mass is varied.

4.4.1 k_T DEPENDENCE OF PION RADII Collective flow generates a characteristic fall-off of the pion source radii with k_T , which is ubiquitously observed in data. Final results for the k_T -dependence of Gaussian radii from central Au+Au (Pb+Pb) collisions exist at the AGS (82; 116), SPS (172; 132; 176; 173; 144), and RHIC (137; 177; 129; 113; 174; 117). As is clear from Figure 9, aside from a small variation in overall scale (discussed later), the k_T dependence is startlingly similar for all energies.

Figure 10 quantifies the evolution of the m_T -dependence of the pion source radii with $\sqrt{s_{NN}}$, using fits to $R_i(m_T) \sim m_T^{-\alpha_i}$. As discussed in Section 2.7.1, $\alpha_i = 0.5$ would represent expectations for instantaneous thermal emission for a three-dimensionally expanding fireball in the limit of large m_T .

The similarity persists as N_{part} is varied. In Au+Au collisions at RHIC, $R_i(m_T)$ ($i = out, side, long$) simply scale as $N_{part}^{1/3}$, with perhaps some flattening for $N_{part} < 100$ (113; 117). Very similar k_T dependence for different N_{part} is also observed in Pb+Pb collisions at SPS (144) and for Si+Au and Au+Au collisions at the AGS (116).

In a flow-dominated freeze-out scenario, the fall-off of transverse radii with m_T increases as flow increases and/or temperature decreases (e.g. 68). Blast-wave fits to spectra (178) indicate that freeze-out flow and temperature vary significantly

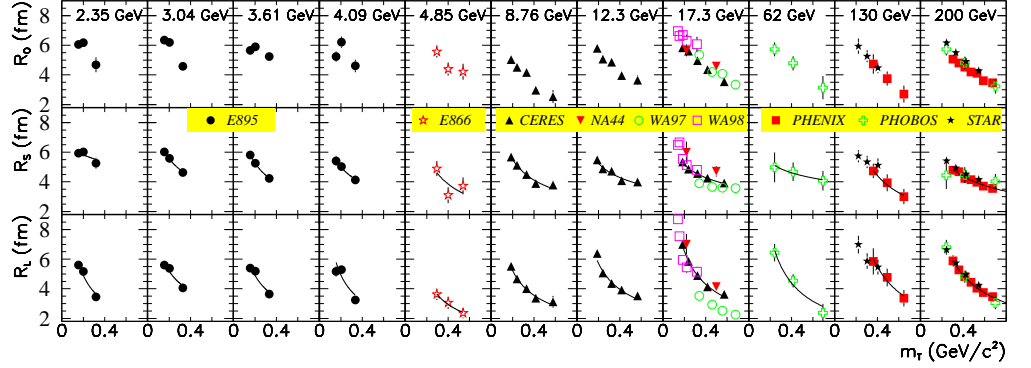


Figure 9: World data set of published m_T dependence of pion Bertsch-Pratt radii near mid-rapidity from Au+Au (Pb+Pb) collisions. Centrality selection is roughly top 10% of cross-section, but varies somewhat with experiment. Data from (82; 116; 144; 132; 173; 176; 177; 117; 174; 137; 113) Lines represent parameterized fits; see text for details.

with $\sqrt{s_{NN}}$ for $\sqrt{s_{NN}} \lesssim 10$ GeV. The overall approximate $\sqrt{s_{NN}}$ -independence of the α_i parameters may reflect the fact that significantly changing the slope of $R_i(m_T)$ requires very large changes in flow and temperature; on the other hand, it could be that the compensating effects of smaller (larger) homogeneity lengths generated by larger flow (temperature) cancel almost exactly in Nature. While $R_i(m_T)$ almost certainly reflects strong collective flow, the strength of that flow requires other information, such as particle spectra (90; 68).

Since the radii fall off roughly as $1/\sqrt{m_T}$ (c.f. Figure 10) and such a dependence has been discussed frequently in the literature (e.g. 179; 180), it is interesting to look at the overall scale parameter from a single-parameter fit to $R'_i/\sqrt{m_T}$. The $\sqrt{s_{NN}}$ -dependence of R'_{side} and R'_{long} are shown in Figure 11. The scale of the longitudinal homogeneity length grows significantly with $\sqrt{s_{NN}}$, consistent with an increase of the system evolution time. However, R' varies only very weakly with $\sqrt{s_{NN}}$.

4.4.2 SYSTEMATICS WITH PARTICLE MASS Systematic studies for different mass particles provide additional controls probing the space-time evolution of the source. In particular for kaons, the interpretation may be simplified due to reduced effects of resonance feed-down (181) and a reduced scattering cross-section for K^+ in nuclear matter, raising the possibility that kaon correlations could peer “farther back” to earlier stages of the collision (182). Indeed, the first kaon measurements (183; 184; 185; 171) reported smaller source radii for kaons. However, the observation that radii for K^+ and K^- were very similar (184) was an early experimental indication that different cross-sections were not the driving physics behind these smaller radii. This was supported by model calculations (186) which suggested that K^+ and K^- in fact scattered roughly equally in the dense medium created in heavy ion collisions. In this case, the smaller radii for kaons results from their increased mass in a flow field, not different cross-sections.

If indeed flow is generated in matter sufficiently dense that individual cross-sections are unimportant, then all particles participate equally in collective trans-

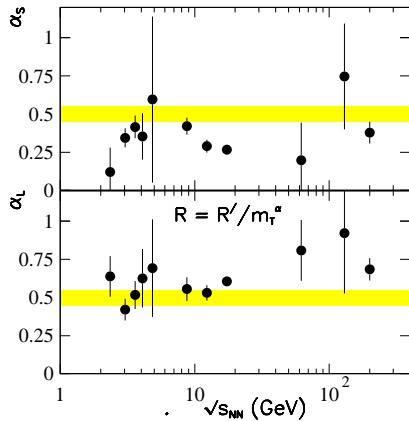


Figure 10: The excitation function of m_T power-law fall-off of pion source radii ($R_i(m_T) \sim m_T^{-\alpha_i}$). Shaded regions show $\alpha_i = 0.5 \pm 0.05$.

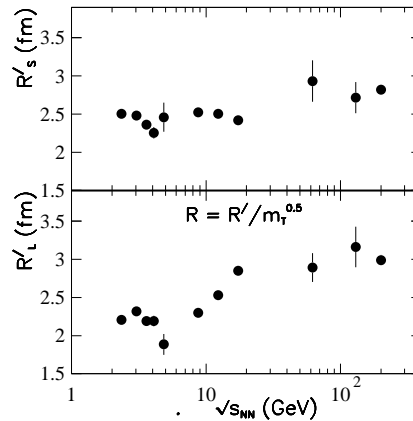


Figure 11: The excitation function of scale parameters R'_i from fits of the data in Figure 9 to $R_i(m_T) = R'_i/\sqrt{m_T}$. Note: this is a different functional form than that used in Figure 10.

verse flow. In this case, their source radii should approximately follow a common m_T scaling (26; 154; 187; 8). Within uncertainties, first results on kaon interferometry by NA44 at the SPS in S+Pb (179) and Pb+PB (188) collisions were consistent with a $1/\sqrt{m_T}$ scaling expected for isotropic “Hubble” flow (26; 154; 187). In some more recent analyses (189), the common m_T systematic for the transverse radii is less steep than for $R_{\text{long}}(m_T)$ (c.f. Figure 10), as might be expected for more boost-invariant (non-isotropic) flowing systems (8; 68).

Figure 12 collects the m_T dependence of homogeneity lengths for several energies. Left panels show results for Si+Au collisions at $\sqrt{s_{\text{NN}}}=5.4$ GeV, measured by E802 for pions (116) and kaons (183; 185). Femtoscopic radii for pions (132; 189; 144; 176) kaons (188), protons (190), and photons (191) measured in Pb+Pb collisions at the SPS are shown in the center panels. The right panel shows the one-dimensional radius parameter R_{inv} measured at RHIC for pions, charged kaons, and protons (192), neutral kaons (193), and with $\Lambda - p$ correlations (194).² To compare across energies, R_{inv} results are included for the AGS and also for the SPS, where the R_{inv} values were calculated from the 3D fit results by accounting for the boost along the outwards direction from the LCMS to the PCOM frame, $R_{\text{inv}}^2 = R_{\text{long}}^2 + R_{\text{side}}^2 + \gamma^2 R_{\text{out}}^2$, where γ is given by m_T/m of the pair. Note that for massless particles, such as photons, γ is given by k_T/Q_{inv} . For a given orientation of the photon pair momentum R_{inv} can be related to 3D radii in the LCMS frame through the following expression for the exponent of

²Protons and Λ baryons are not identical particles, of course. However, their masses are sufficiently close to try including $R_{\text{inv}}^{p-\Lambda}$ on a m_T -scaling plot. Adherence to the scaling is consistent with a flow-dominated scenario in which homogeneity lengths depend only on mass.

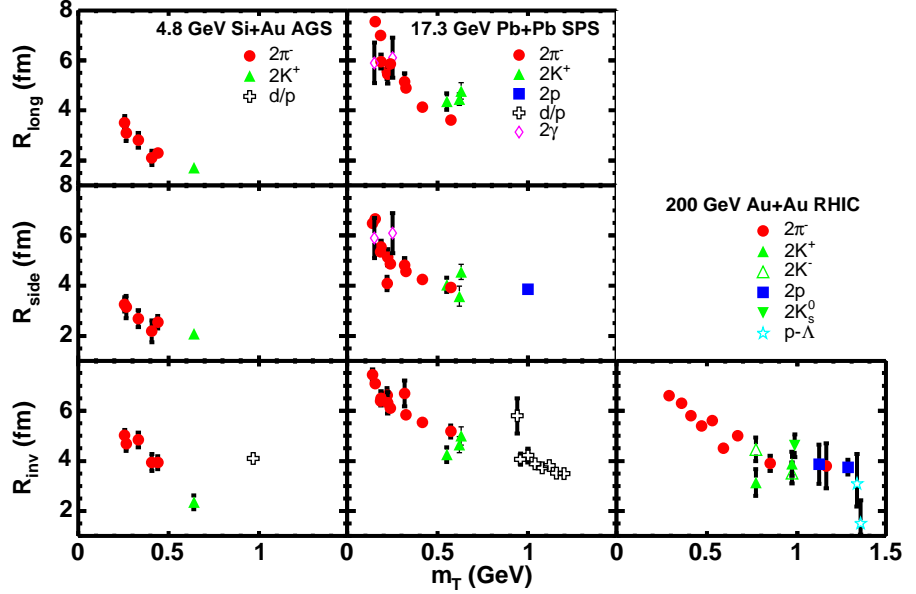


Figure 12: Transverse mass dependence of homogeneity lengths from correlations between particles of (nearly) identical mass.

the Gaussian source term,

$$\begin{aligned}
 Q_{\text{inv}}^2 R_{\text{inv}}^2 &= Q_{\text{inv}}^2 (R_{\text{long}}^2 \cos^2 \theta + R_{\text{side}}^2 \sin^2 \theta \sin^2 \phi + k_T^2 / Q_{\text{inv}}^2 R_{\text{out}}^2 \sin^2 \theta \cos^2 \phi) \\
 &= k_T^2 R_{\text{out}}^2 \sin^2 \theta \cos^2 \phi + Q_{\text{inv}}^2 (R_{\text{long}}^2 \cos^2 \theta + R_{\text{side}}^2 \sin^2 \theta \sin^2 \phi). \quad (38)
 \end{aligned}$$

Thus R_{out} drops completely out of the source term and enters into the λ coefficient (191), but as a function of k_T . For this reason the 1D Q_{inv} fits from WA98 are plotted alongside values of R_{side} and R_{long} instead.

This consistency between different particle types may carry an important message. It calls into question theoretical scenarios which appear to explain $R_i(m_T)$ for particular particle types only (29). Further, the consistency with emission from a *common* flow-dominated source may also support freeze-out scenarios in which the last scattering *in the dense phase* determines the homogeneity region, instead of milder rescatterings in the more dilute stage, which are dominated by particle species-dependent cross-sections (195; 196).

4.5 New possibilities: Non-identical particle correlations

Almost all femtoscopic measurements to date have been done through correlations of identical particles, usually charged pions. With the availability of high-statistics data sets and new theoretical ideas (197), however, experiments are beginning to make full multi-dimensional analyses of non-identical particle correlations. These correlations are being used to test and refine the treatment of Coulomb effects in identical-particle correlation analyses, to explore violations of flow-induced “universal” m_T scaling, and to extract qualitatively new information on the space-time substructure of the source.

4.5.1 OPPOSITE-SIGN PION CORRELATIONS The most common non-identical particle femtoscopy measurements have involved $\pi^+ - \pi^-$ correlations. Here, the natural assumption is that the homogeneity regions of these particles coincide. The primary interest is in whether the homogeneity volume is extracted from $\pi^\pm - \pi^\pm$ correlations—driven by quantum statistics and Coulomb FSI (with minimal strong FSI effects)—is consistent with that extracted from $\pi^+ - \pi^-$ correlations—determined only by Coulomb and strong FSI (113). This issue has obvious implications for the FSI weight $F(\mathbf{q})$ discussed in Section 3.5.

At the SPS, $\pi^+ - \pi^-$ correlations from reactions with sulphur (198) and lead (172) beams were consistent with emission from a homogeneity region with spatial scale roughly the same as the radii from identical-particle correlations. Preliminary analysis of opposite-sign pion correlations from Au+Au collisions at the AGS (171; 69) similarly find consistency with scales extracted from like-sign pion correlations. Recently, STAR reported (113) strong consistency at RHIC energies as well. Using large-statistics data sets, they further show that contributions to the $\pi^+ - \pi^-$ correlation function from strong-force interactions, though small, are nevertheless important in explaining the data in detail.

4.5.2 OTHER NON-IDENTICAL PARTICLE CORRELATIONS IN $|\vec{q}|$ For statistical reasons, most non-identical particle correlations are measured in one-dimensional $|\vec{q}|$ -space, and thus probe only the average spatial separation between the two particles at freeze-out in the pair center-of-momentum frame (c.f. Section 2.3). Because they are sensitive to the size of each particles emission source and this separation between them it is often necessary to rely on the identical particle correlations for interpretation of results.

Despite these ambiguities, one-dimensional non-identical particle correlations may be used to test existing systematics and expectations of the freeze-out scenario. Conversely, if the freeze-out geometry may be taken as given, non-identical particle correlations may place constraints on hadronic scattering parameters, e.g. to *measure* the squared relative wave-function of a pion and a Ξ baryon. First studies along these lines are underway at RHIC (199).

Wang and Pratt (54) suggested the measurement of $p - \Lambda$ correlations, which may be more sensitive to large structure than $p - p$ correlations, and have a higher 2-track reconstruction efficiency for some experiments. Published results from Au+Au collisions at the AGS (51) suggest a $p - \Lambda$ separation distribution with a width similar to the proton homogeneity length and roughly consistent with m_T scaling expectations (cf Section 4.4.2). Preliminary results at SPS (200) and RHIC (199; 201) give similar conclusions. Since the $p - \Lambda$ potential is not unambiguously known theoretically (54), it is unclear whether possible statistically marginal violations of m_T scaling (51) or differences between baryon and anti-baryon emission regions (199) are meaningful. To first order, existing $p - \Lambda$ correlations confirm existing systematics.

With the high-statistics and -quality data sets at the largest collision energies, truly exotic correlation studies are possible. Preliminary results (201) from RHIC on $\pi - \Xi$ correlations look particularly promising. Here, the well-measured pion emission distribution may be used to study the strange baryon freeze-out configuration. It may also provide information on the $\pi - \Xi$ final state interaction and scattering cross-section which in turn can be used to constrain our understanding of the sources of collective flow. Blast-wave calculations reproduce the preliminary $\pi - \Xi$ correlations, suggesting that the Ξ flow is determined by its

mass, not its quark content.

4.5.3 NEW INFORMATION– NON-IDENTICAL PARTICLE CORRELATIONS WITH DIRECTIONAL CUTS ON \vec{q} Non-identical particle correlation analyses as a function of– or with cuts on– the relative direction of \vec{q} and \vec{P} reveal qualitatively new information (56; 57; 197, and Section 2). In particular, in addition to the RMS width of the separation distribution, the direction and size of the *average* separation between the particles is probed, although offsets in time and space cannot be disentangled; this is shown as Δr in Figure 3. The correlation functions selected for $\vec{q} \parallel \vec{P}$ and $\vec{q} \not\parallel \vec{P}$ differ if $\Delta r \neq 0$. Furthermore, collective flow will induce *position*-momentum correlations detectable with directionally-selected non-identical particle correlations (57; 68).

These correlations are statistically challenging, and few results are available. At RHIC, STAR has reported (58) asymmetries in $K^\pm - \pi^\pm$ correlations measured in central Au+Au collisions. In Blast-wave calculations with transverse flow roughly adjusted to reproduce other observations at RHIC (68) describe the data semi-quantitatively. Preliminary studies of $p - \pi$ correlations at the SPS (200) and RHIC (201) exhibit very similar mass-ordered spatial asymmetries in the transverse plane. A preliminary study at the AGS (69) reported very large ($\Delta r_{\text{long}} \approx 10$ fm) average $p - \pi$ separations in the *longitudinal* (beam) and *impact parameter* direction for forward-moving particles, suggesting very strong longitudinal flow, however, this result was never confirmed.

We expect full three-dimensional analyses of a wide range of non-identical particle combinations to be available in the near future from RHIC experiments. Sophisticated analyses may probe non-trivial geometric substructure when selecting on reaction plane, including the *sideward* shift predicted by blast-wave calculations (68) when anisotropic flow structure is present.

5 Interpretations of Experimental Results

In this Section, we ask what we can learn from the spectrum of results just presented. Beginning with the broadest, least detailed observations, we move to two fundamental quantities which may be directly extracted from the data, and finally on to comparisons with specific models of heavy ion collisions.

5.1 General Conclusions from Systematic Trends

One of the first messages to take away from the discussion of Section 4 is that the results are stable across detector and method; experimental systematics and uncertainties are under control. Whatever difficulties we may have in interpreting measurements, we may be confident that they do not have their origin in experimental artifact.

The size and shape inferred from two-particle correlations tracked with collision geometry as anticipated. Kinematic and mass dependences of femtoscopic measurements showed the expected clear signatures of strong collective flow in the beam direction and perpendicular to it.

At a generic level (ignoring quantitative predictions), we are first given pause at the jejune nature of the $\sqrt{s_{NN}}$ -dependence of femtoscopic parameters. The most common example discussed is the pion source radii excitation function shown in Figure 13. In Section 4 we have explored the trends in considerably greater detail,

but the Figure conveys the right message: scanning $\sqrt{s_{NN}}$ through a range of two orders of magnitude, changes final-state geometry little.

Based on rather generic arguments of soft points in the Equation of State or entropy generation during a phase transition, there had been hopes for non-trivial structure in the excitation function, as the energy threshold for Quark Gluon Plasma (QGP) creation was crossed (59; 202; 203; 204). General expectations were for “long” (relative to the explosion of a purely hadronic system) system evolution time scales if QGP was formed. Blast-wave analyses of data (e.g. 68) appear to rule out systems with lifetimes in the neighborhood of 20 fm/c or higher. However, due to dynamic effects careful comparison with a dynamical model is required to extract detailed evolution information; this is discussed in Section 5.3.

Neglecting flow and other important details, and focusing solely on the excitation function in Figure 13, we note that its gradual but non-monotonic behavior has been successfully described in terms of a single parameter– a universal ($\sqrt{s_{NN}}$ -independent) pion mean-free-path at freezeout (205) of 1 fm, related to the pion number density and chemical composition of the system at freeze-out. It remains unclear how to interpret this intriguing simplification of the data, or whether it might be used to provide insight on the underlying physics of the collision.

5.2 Phase Space Density and Entropy

As shown in Section 2.9, average phase space densities, $\bar{f}(\mathbf{p})$, can be calculated by combining source-size measurements with spectra. Pionic phase space densities have been estimated for 130A GeV collisions at RHIC (110), at the top SPS energies, and for several AGS energies (208). Figure 14 shows results for all three regions. In each case, the phase space densities were calculated via Equation 28. For the SPS case, results were generated by applying Equation 28 to published spectra (209) and source-size measurements (210). We note that our calculations for the SPS are higher than previously published values at low p_t (211). This discrepancy is likely due to the fact that in Ref. (211) analytic parameterizations were used which significantly differ from published spectra at low p_t .

When applying Equation 28 an issue arises as to whether one should subtract the contribution from resonances to the spectra. Indeed, if the pions are created by decays so far outside the source volume that they do not contribute to the correlation function, they should not be considered as pions when calculating the phase space density. There are two strategies to correct for such pions. First, one could use spectra where such particles are subtracted and apply Equation 28 literally. As a second option, one could use the spectra without subtractions, but then multiply the expression for \bar{f} by $\sqrt{\lambda}$. Since most published spectra have been purged of the products of weak decays, the first method is usually applied. However, published spectra still include the contribution from η s which decay thousands of fm away from the source. The η contribution was accounted for in the SPS calculation for Figure 14 by reducing the spectra by 5%.

The phase-space densities in Figure 14 show a steady rise with beam energy that seems to plateau at SPS energies. Since the displayed phase space densities have been averaged over coordinate space, the peak values are higher (87; 102), by a factor of $2\sqrt{2}$ if the spatial profiles are Gaussian. For a breakup temperature of 110 MeV, this requires a rather high chemical potential, near or above 80 MeV

Figure 13: Panels (a-d): Excitation function of π^- source parameters at mid-rapidity and low k_T (~ 0.17 GeV/c) in central Au+Au(Pb+Pb) collisions. PHENIX data are for $k_T \sim 0.26$ GeV/c and so fall somewhat below the trend. Panel (e): Transverse freeze-out anisotropy parameter from non-central ($|\bar{b}| \sim 8$ fm) Au+Au collisions, estimated from the azimuthal dependence of source radii. Data are from (177; 117; 137; 129; 113; 116; 144; 176; 174; 132; 82; 127). Also shown are calculations (82; 127; 132; 206) at several energies with the RQMD model (207); hashed region at other values of $\sqrt{s_{NN}}$ interpolates between these calculations.

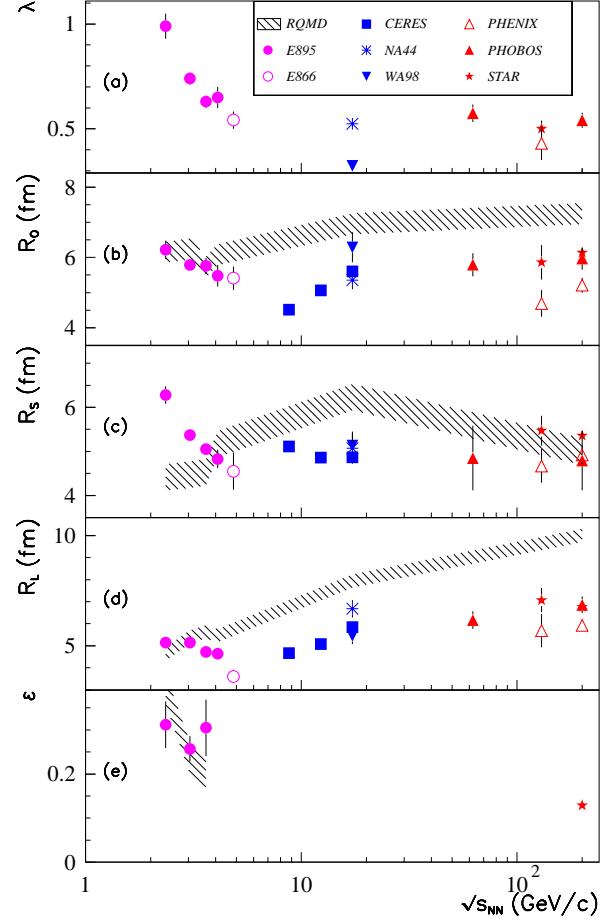
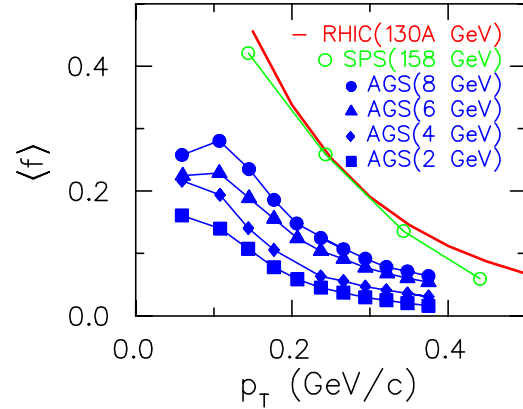


Figure 14: Average pionic phase space densities for central Au+Au and Pb+Pb collisions from the AGS to RHIC rise with beam energy but seem to plateau at SPS energies. Values were calculated with Equation 28.



(212; 213; 214).

Ratios of particle yields have been shown to be consistent with chemical freeze-out at temperatures near 170 MeV (215; 216; 217). This suggests an interpretation where at higher densities the system is so strongly interacting that yields equilibrate until the system reaches this temperature, and are then frozen during the subsequent freeze-out. At AGS energies, the increase in pion production

brought on by increased beam energy results in more pions being pushed into a given amount of phase space as realized in Figure 14 by the rising phase space density. However, one would expect that, as the excitation energy surpasses the threshold for a 170 MeV temperature, the phase space density would reach a limiting value. Since the average phase space density is preserved in an isentropic expansion with fixed particle number (not exactly true if several masses are mixed together), one would expect phase space densities to saturate once energy densities reached this value. Indeed, the behavior in Figure 14 is consistent with this scenario.

As shown in Section 2.9, entropy can be calculated from average phase space densities using Equation 30. Spectra and source-size measurements for baryons and mesons were used to estimate phase space densities and entropy for 130A GeV Au+Au collisions at RHIC. The total entropy in the central unit of rapidity was estimated at $dS/dy = 4450 \pm 10\%$ (110). In a hydrodynamic expansion entropy is conserved though viscosity and shock waves might result in a roughly 10% increase during the evolution. Thus, this measurement provides an upper bound for the entropy at $\tau \sim 1$ fm/c when thermalization first occurs. At $\tau = 1$ fm/c, the volume for particles in this rapidity slice is determined by the geometric cross-sectional areas of the over-lapping gold nuclei multiplied by $c\tau$, thus an upper bound for the entropy density can be determined $s \leq (dS/dy)/(\tau\pi R^2)$. Knowing the energy density at $\tau = 1$ fm/c would then provide a point in the equation of state, s vs. ϵ . The value of dS/dy estimated in (110) is consistent with lattice calculations if $\epsilon(\tau = 1 \text{ fm/c}) \sim 7 \text{ GeV/fm}^3$. Estimates of the original energy density from the final-state measurement of $dE_t/d\eta$ are in range of 4.5 GeV/fm^3 (218), but since these estimates neglect losses from longitudinal work or the energy from longitudinal thermal motion, the 7 GeV/fm^3 value seems reasonable.

5.3 Dynamic Models and their Comparison with Data

It is increasingly recognized that the comparison of dynamic models of heavy ion collisions to data is only insightful if it involves a sufficiently large variety of experimental data. A comparison of dynamic models to femtosopic measurements alone (or to any other class of measurements alone) is of limited value, simply because for realistic models, the number of possible model-dependent parameter choices then tends to exceed the number of experimental constraints. In fact, all the model results which we review in the current subsection remain unsatisfactory with this respect: they either deviate significantly from femtosopic data, or they reproduce these data at the price of missing other important experimental information. In particular, there is so far no dynamically consistent model which reproduces quantitatively both the systematic trends discussed in Section 4, and the corresponding single inclusive spectra. In this situation, the scope of this subsection is somewhat limited. We want to explain why a dynamical understanding of femtosopic measurements is important. We shall also discuss the key physics input which enters current attempts of dynamic modeling, and the uncertainties resulting from it. However, we shall try to bypass as far as possible model-dependent details, and rather focus on the question which qualitative changes in the underlying dynamics result in characteristic changes of femtosopic data.

Correlation measurements provide a “snap shot” of the geometrical distribu-

tion of particles at the time they decouple from the reaction. This geometrical distribution provides a unique test of the dynamical evolution of the produced matter at the late stage. Since the spatial extension, dynamical evolution and lifetime of the produced system determines phase space density and thus particle reaction rates, it is evident that any dynamic model for the latter has to be consistent with femtoscopic information. So far, correlation analyses have focused mainly on Boltzmann (or cascade) models, on hydrodynamic models, or on combinations of both (hybrid models). These model classes correspond to rather different equations of state.

The equations of state represented by cascade models tend to be stiff unless they incorporate large number of resonant scatterings. If particles collide via $2 \rightarrow 1 \rightarrow 2$ processes where the intermediate state has a finite lifetime, the equation of state can be softened (219; 220). A prominent example of a cascade model is RQMD (186), Relativistic Quantum Molecular Dynamics; it is the only one which has been compared to data at AGS, SPS and RHIC. Results for RQMD are shown in Figure 13. However, for RHIC energies, the dynamic consistency of RQMD is questionable since the model uses for most of its evolution hadronic degrees of freedom although RQMD simulations for RHIC yield an energy density which stays above that of normal nuclear matter for a significant duration (~ 5 fm/c). As can be seen in Figure 15 RQMD, which models the expansion as a hadronic cascade, over-predicts R_{out} and R_{long} at RHIC despite the fact that it under-predicts multiplicities. Another hadron cascade model, the ‘‘Hadronic Rescattering Model’’ (221; 222), based solely on hadronic rescattering with sudden collisions gives smaller sources, similar to those of AMPT which come closer to the data. Part of the difference between the two hadronic cascades may derive from RQMD’s treatment of scatterings as resonant interactions with finite lifetimes, which differs from the instantaneous collisions employed in the Hadronic Scattering Model. Figure 15 also displays results for MPC (223), which aims at providing a transparent partonic toy model, by modeling the collision of light partons which undergo a one-to-one hadronization to pions. MPC, which has only instantaneous two-to-two scatterings should have the stiffest effective equation of state and underestimates the source radii. AMPT (224; 225) aims at a realistic description of all aspects of the reaction dynamics and includes a partonic cascade coupled to a hadronic cascade employing a full list of resonant interactions. AMPT provides a good fit to experimental radii, though the k_t fall-off is stronger than that found in the data, which fall off $\sim m_T^{-1/2}$. The more rapid decrease of radius parameters with respect to k_T may be due to the continuous surface-like emission characteristic of microscopic models, which when combined with cooling, gives a higher relative weight for high-energy particles to have been emitted earlier in time, before the reaction volume has reached its full spatial extent. Any modification that would reduce this type of emission should improve agreement with data. Cautiously, one would conclude that the results in Figure 15 favor models with a stiff, but not too stiff, equation of state and no latent heat.

This conclusion is further supported if one compares the results of cascade models with those of hydrodynamic simulations or hybrid models. Figure 16 compares results from RHIC to a 3-dimensional hydrodynamic model of Hirano *et al.* (226) that investigates the effects of resonance decays on chemical composition, the 2-dimensional model of Heinz and Kolb (227), and a 2-dimensional chiral model by Zsische *et al.* (228) that performs calculations for both first order and

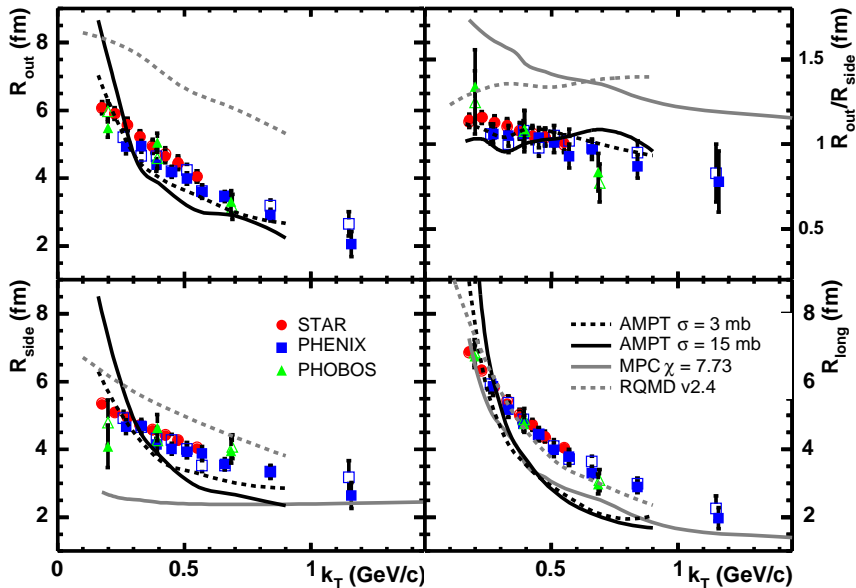


Figure 15: RHIC results are compared to three Boltzmann/cascade models, RQMD, MPC and AMPT with two different parton cross sections. Data are for $2\pi^-$ (open symbols) and $2\pi^+$ (closed symbols) source radii.

cross-over transitions. In all cases the more favorable calculations are compared with the data: partial chemical freeze-out for Hirano, and cross-over transition for Zschesche, but large discrepancies between the models and the data still remain.

In general, these models invoke equations of state which are typically softer than those used in cascades and Boltzmann calculations, and they often have latent heats to accompany the transition from the partonic phase. As a consequence, lifetime and emission duration of the produced matter are significantly larger than what one finds in cascade models, and such models often significantly over-predict R_{long} and $R_{\text{out}}/R_{\text{side}}$. The fact that R_{side} comes out smaller than the data in Figure 16 is mainly due to an attempt to compensate within the available model-parameter range the very large time-scales as much as possible. Purely hydrodynamic models can vary overall source volumes by adjusting the breakup criteria, but doing so can make it difficult to fit the three source dimensions, and their m_T dependence, as well as other observables. Several prescriptions have been applied to improve the modeling of the breakup in the late stage to depend on microscopic considerations determined by free-space cross sections without having an extra adjustable parameter, e.g. breakup density, (103; 104; 102). An alternative to improve the description of breakup is the use of hybrid models, in which hydrodynamic evolution in the early stage is combined with cascading in the late stage. In Figure 16, we compare results of one such hybrid model, URQMD (105; 106), to results of hydrodynamic calculations. Similar results were obtained by Teaney, et al., (107; 108). Compared to hydrodynamic simulations, hybrid descriptions do not seem to notably reduce the over-predicted lifetimes. They tend to emit most of the pions at a time near or above 15 fm/c and signif-

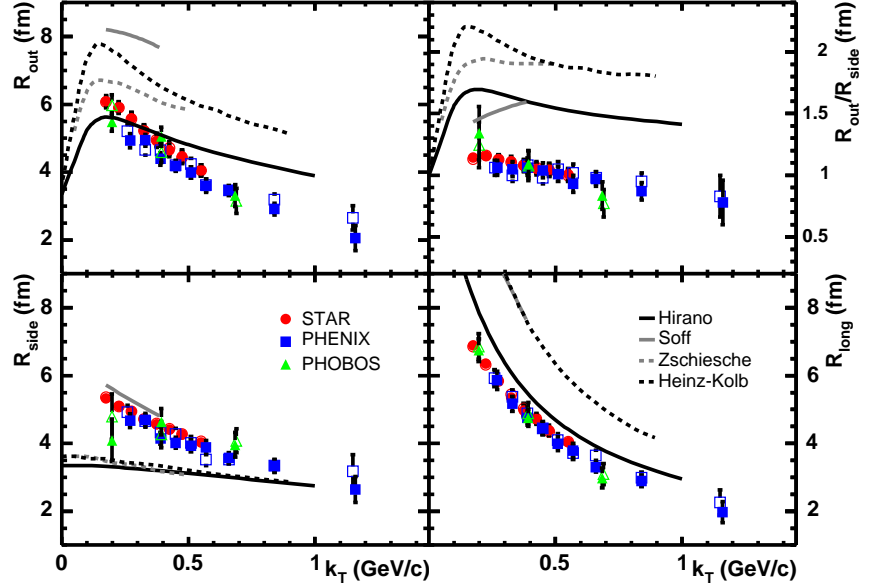


Figure 16: Hydrodynamic (Zschesche, Hirano and Kolb) and hybrid hydrodynamic/cascade (Soff) models calculations in comparison to RHIC data. Data are for $2\pi^-$ (open symbols) and $2\pi^+$ (closed symbols) source radii.

icantly over-predict R_{out}/R_{side} ratios, while blast-wave parameterizations favor breakup times near or slightly below 10 fm/c. However, firm conclusions that the relative failure of hybrid models derives from the chosen equations of state can not be made until comparisons are made between Boltzmann and hybrid calculations that use the same equation of state. Until such an analysis is performed, other issues will cloud the interpretation, such as whether viscous effects or details of the hydrodynamic/Boltzmann interface dominate results and might even invalidate the hydrodynamic approach.

Entropy and pressure are intimately related, in that knowing the entropy density as a function of the energy density determines the pressure as a function of energy density. Whereas one can intuitively understand why a lower pressure would lead to longer lifetimes and larger source dimensions, the manifestations of changing the entropy are less transparent. One way to understand the effects of changing the entropy is to associate higher entropy with a larger effective number of light degrees of freedom. For instance, if one were to hadronize via a one-to-one parton-to-pion scheme, the volume per identical particle would change by a factor of the number of degrees of freedom. If the r.m.s. momentum is the same before and after hadronization this would imply a change in entropy per particle equal to the logarithm of the ratio of effective degrees of freedom, and if ~ 40 light partonic degrees of freedom were immediately replaced by three pionic degrees of freedom, the system would lose $\ln(40/3)$ units of entropy per particle. To conserve entropy, a system must expand its volume by a similar ratio which implies an increase in radius parameters. Hydrodynamic models, which manifestly conserve entropy, use the energy stored in the latent heat to provide the heat necessary to

preserve entropy during hadronization. Entropy conservation is more difficult to enforce in microscopic approaches. This underlines the challenges involved in applying microscopic simulations in an environment of strongly interacting matter with ill-defined degrees of freedom, and re-emphasizes the importance of gaining a better understanding of the validity of hydrodynamic calculations in such manifestly finite situations. It is peculiar that the entropy extracted from source-size and spectral measurements is consistent with the lattice-inspired equation of state (107; 108), while the source sizes extracted from hybrid models incorporating similar equations of state significantly over-predict source sizes. Part of this contradiction can be explained by increases in the populations of baryons, which, since they inherently have more entropy per particle than do pions, can account for much of the missing entropy (110). Thus, the HBT puzzle does not necessarily imply an entropy puzzle.

The HBT puzzle is not so much that radius parameters can not be fit by models, but that our most sophisticated models, which incorporate a phase transition, fail to reproduce the data. The very gradual evolution of extracted source sizes as beam energies traverse a large range of energies is remarkable and puzzling in its own right. A simple explanation is that the equations of state do not dramatically change as the energy density changes from hadronic to super-hadronic densities, i.e., there is not even a hint of a latent heat. However, as emphasized above, a host of unresolved issues prevent more quantitative conclusions from being reached. These qualifiers can not be lifted until much more thorough analyses of models are performed, entailing a systematic exploration of the sensitivity of model predictions to both parameters and assumptions. This would necessitate a tremendous commitment from the community, but without it, many conclusions about the matter created in relativistic heavy ion collisions will remain vague.

6 Summary

Twenty five years ago, the goal of femtoscopy was to demonstrate that one could measure a hadronic length scale with correlations, and if a result was on the order of a few Fermi, the analysis was deemed a success. In contrast, with the improved accuracy of measurements, the enormous increase in statistics, and the simultaneous development of phenomenology and theory, femtoscopy is now considered a precision measurement. Ten percent deviations between theory and experiment are now taken seriously as evidence that the spatio-temporal description of a model is significantly flawed. At RHIC energies, all six dimensions of the correlation function have been exploited to provide truly three dimensional insight into the phase space cloud for particles of any momentum with any direction.

For relativistic heavy ion collisions, there were expectations that a transition from hadronic to partonic matter might be accompanied by a large latent-heat, which would bring about a dramatic change in the dynamics as beam energies traversed the range for exploring the mixed phase. During this range, it was expected that the latent heat and the associated softening of the equation of state would manifest itself by slowing the explosion with lifetimes approaching or exceeding 20 fm/c. The “signal” of the phase transition would have been an increase of the effective lifetime for a range of beam energies followed by a return to more explosive and shorter lived reactions at even higher energies.

Extended lifetimes were not observed. Increasing beam energies from AGS

to RHIC indeed causes larger energy densities and higher multiplicities which push towards increasing the source volumes. However, much of this increase in multiplicities is absorbed by higher phase space densities rather than larger source sizes. Combined with the increasing strength of radial flow, which provides smaller regions of homogeneity relative to the overall source volume, the result is that the effective dimensions change remarkably little over a wide range of beam energies. Furthermore, it appears that lifetimes of the reaction never leave the neighborhood of 10 fm/c. Not only does this represent a lack of evidence for a latent heat, it represents evidence that there is no such latent heat.

These conclusions remain only modestly guarded. Theory and phenomenology are progressing, but improvements in such aspects as mean-field effects or accounting for the smoothness approximation are not expected to change conclusions by more than 10%. Like the improvements in including Coulomb effects in Sec. 4, removing some of the distortions and aberrations from the analyses are likely to be significant for fine-tuning models, but should not alter the conclusion that there is no large latent heat associated with the reaction.

It is in the theory of modeling, i.e., generating the source functions, that there is the greatest need for progress. Since femtoscopic measurements are determined solely by the geometry of breakup, changes in chemical or kinetic evolution may have a significant impact on extracted source dimensions. The next generation of transport theories should be more flexible, and will probably incorporate numerous effects such as in-medium mass changes, in-medium reduction of scattering cross sections, viscous effects and dynamical solutions for chemical rates.

Despite the progress listed above, measurement has only begun to address the rich expanse of information available in correlations. Nearly all the three-dimensional analyses have been focused on identical-pion correlations. The huge data sets of the recent and upcoming runs at RHIC make it possible to analyze source functions for many pairs of particles in six full dimensions. In addition to providing important verification of identical-pion measurements, these analyses address other issues, such as whether all species flow and break up together.

This is not a field for the complacent. As emphasized above, efforts at RHIC are just beginning to explore wholly new classes of correlations. Energy densities at the LHC might surpass those at RHIC by the same factor that those at RHIC surpassed the AGS. Just as our visions of the future from twenty years ago proved largely naive, we should be prepared to be surprised with the femtoscopia of the next 25 years.

Acknowledgements

This work was supported by U.S. Department of Energy, Grants DE-FG02-03ER41259 and W-7405-ENG-48, and by U.S. National Science Foundation Grant PHY-0355007. The authors thank David Brown, Ulrich Heinz, Sergei Panitkin and Nu Xu for sharing insights. For providing and explaining data, we thank Harry Appelshäuser, Giuseppe Bruno, Mike Heffner, Mercedes López-Noriega, Michael Murray and the experimental collaborations who conscientiously post data in numerical form and respond to queries. We thank Mark Meamber of the Physics Art Team at Lawrence Livermore National Laboratory, for creating the illustrative figures.

This work was performed under the auspices of the U.S. Department of Energy by University of California, Lawrence Livermore National Laboratory under contract No. W-7405-Eng-48.

References

1. J. Adams et al. *nucl-ex/0501009*
2. K. Adcox et al. *nucl-ex/0410003*
3. B. B. Back et al. *nucl-ex/0410022*
4. I. Arsene et al. *nucl-ex/0410020*
5. D. H. Boal, C. K. Gelbke, and B. K. Jennings. *Rev. Mod. Phys.* 62:553–602, 1990
6. W. Bauer, C. K. Gelbke, and S. Pratt. *Ann. Rev. Nucl. Part. Sci.* 42:77–100, 1992
7. Ulrich W. Heinz and Barbara V. Jacak. *Ann. Rev. Nucl. Part. Sci.* 49:529–579, 1999
8. Urs Achim Wiedemann and Ulrich W. Heinz. *Phys. Rept.* 319:145–230, 1999
9. T. Csorgo. *Heavy Ion Phys.* 15:1–80, 2002
10. R. M. Weiner. *Phys. Rept.* 327:249–346, 2000
11. Boris Tomasik and Urs Achim Wiedemann. *hep-ph/0210250*
12. Gideon Alexander. *Rept. Prog. Phys.* 66:481–522, 2003
13. R. Hanbury-Brown and R.Q. Twiss. *Nature.* 178:1046, 1956
14. R. Hanbury-Brown and R.Q. Twiss. *Phil. Mag.* 45:663, 1954
15. Gerson Goldhaber, Sulamith Goldhaber, Won-Yong Lee, and Abraham Pais. *Phys. Rev.* 120:300–312, 1960
16. G. I. Kopylov and M. I. Podgoretsky. *Sov. J. Nucl. Phys.* 15:219–223, 1972
17. G. I. Kopylov, V. L. Lyuboshits, and M. I. Podgoretsky. JINR-P2-8069
18. S. E. Koonin. *Phys. Lett.* B70:43–47, 1977
19. M. Gyulassy, S. K. Kauffmann, and L. W. Wilson. *Phys. Rev.* C20:2267–2292, 1979
20. W. A. Zajc et al. *Phys. Rev.* C29:2173–2187, 1984
21. T. J. Humanic. *Phys. Rev.* C34:191–195, 1986
22. Gordon Baym. *Acta Phys. Polon.* B29:1839–1884, 1998
23. R. Lednicky and V. L. Lyuboshits. In *Nantes 1990, Proceedings, Particle correlations and interferometry in nuclear collisions* 42-54
24. R. Lednicky. *nucl-th/0212089*
25. R. Lednicky and V. L. Lyuboshits. *Sov. J. Nucl. Phys.* 35:770, 1982
26. S. V. Akkelin and Yu. M. Sinyukov. *Phys. Lett.* B356:525–530, 1995
27. H. W. Barz. *Phys. Rev.* C59:2214–2220, 1999
28. H. W. Barz. *Phys. Rev.* C53:2536–2538, 1996
29. John G. Cramer, Gerald A. Miller, Jackson M. S. Wu, and Jin-Hee Yoon. *nucl-th/0411031*
30. M. C. Chu, S. Gardner, T. Matsui, and R. Seki. *Phys. Rev.* C50:3079–3087, 1994
31. J. I. Kapusta and Y. Li. *nucl-th/0503075*
32. S. Pratt. *Phys. Lett.* B301:159–164, 1993
33. Ulrich W. Heinz, Pierre Scotto, and Qing-Hui Zhang. *Annals Phys.* 288:325–360, 2001
34. Q. H. Zhang, P. Scotto, and Ulrich W. Heinz. *Phys. Rev.* C58:3757–3760, 1998
35. T. Csorgo and J. Zimanyi. *Phys. Rev. Lett.* 80:916–918, 1998
36. J. Zimanyi and T. Csorgo. *Heavy Ion Phys.* 9:241, 1999
37. D. Anchishkin, Ulrich W. Heinz, and P. Renk. *Phys. Rev.* C57:1428–1439, 1998

38. Q. H. Zhang, U. A. Wiedemann, C. Slotta, and Ulrich W. Heinz. *Phys. Lett.* B407:33–38, 1997
39. S. Pratt. *Phys. Rev.* C56:1095–1098, 1997
40. J. Aichelin. *Nucl. Phys.* A617:510–525, 1997
41. U. A. Wiedemann et al. *Phys. Rev.* C56:614–618, 1997
42. Urs Achim Wiedemann, Daniel Ferenc, and Ulrich W. Heinz. *Phys. Lett.* B449:347–353, 1999
43. Sandra S. Padula and Cristiane G. Roldao. *Phys. Rev.* C58:2907–2913, 1998
44. M. Martin, H. Kalechofsky, P. Foka, and U. A. Wiedemann. *Eur. Phys. J.* C2:359–364, 1998
45. R. Barate et al. *Phys. Lett.* B478:50–64, 2000
46. P. Abreu et al. *Phys. Lett.* B471:460–470, 2000
47. S. Pratt. *Phys. Rev.* D33:72–79, 1986
48. D. A. Brown and P. Danielewicz. *Phys. Rev.* C64:014902, 2001
49. David A. Brown, Fu-qiang Wang, and Pawel Danielewicz. *Phys. Lett.* B470:33–38, 1999
50. David A. Brown and Pawel Danielewicz. *Phys. Lett.* B398:252–258, 1997
51. P. Chung et al. *Phys. Rev. Lett.* 91:162301, 2003
52. G. Verde et al. *Phys. Rev.* C65:054609, 2002
53. G. Verde et al. *Phys. Rev.* C67:034606, 2003
54. Fu-qiang Wang and Scott Pratt. *Phys. Rev. Lett.* 83:3138–3141, 1999
55. Scott Pratt and Silvio Petriconi. *Phys. Rev.* C68:054901, 2003
56. R. Lednicky, V. L. Lyuboshits, B. Erasmus, and D. Nouais. *Phys. Lett.* B373:30–34, 1996
57. S. Voloshin, R. Lednicky, S. Panitkin, and Nu Xu. *Phys. Rev. Lett.* 79:4766–4769, 1997
58. J. Adams et al. *Phys. Rev. Lett.* 91:262302, 2003
59. S. Pratt. *Phys. Rev.* D33:1314–1327, 1986
60. G. F. Bertsch. *Nucl. Phys.* A498:173c–180c, 1989
61. T. Csorgo, J. Zimanyi, J. Bondorf, H. Heiselberg, and S. Pratt. *Phys. Lett.* B241:301–307, 1990
62. Urs Achim Wiedemann. *Phys. Rev.* C57:266–279, 1998
63. Michael Annan Lisa, Ulrich W. Heinz, and Urs Achim Wiedemann. *Phys. Lett.* B489:287–292, 2000
64. U. W. Heinz, A. Hummel, M. A. Lisa, and U. A. Wiedemann. *Phys. Rev.* C66:044903, 2002
65. P. Danielewicz and S. Pratt. *nucl-th/0501003*
66. Z. Chajecski, T.D. Gutierrez, M.A. Lisa, and M. Lopez-Noriega. *nucl-ex/0505009*
67. Scott Chapman, J. Rayford Nix, and Ulrich W. Heinz. *Phys. Rev.* C52:2694–2703, 1995
68. Fabrice Retiere and Michael Annan Lisa. *Phys. Rev.* C70:044907, 2004
69. D. Miskowicz. *nucl-ex/9808003*
70. Ulrich W. Heinz, B. Tomasik, U. A. Wiedemann, and Y. F. Wu. *Phys. Lett.* B382:181–188, 1996
71. S. A. Voloshin and W. E. Cleland. *Phys. Rev.* C53:896–900, 1996
72. Urs Achim Wiedemann and Ulrich W. Heinz. *Phys. Rev.* C56:610–613, 1997

73. F. B. Yano and S. E. Koonin. *Phys. Lett.* B78:556–559, 1978
74. Y. F. Wu, Ulrich W. Heinz, B. Tomasik, and U. A. Wiedemann. *Eur. Phys. J.* C1:599–617, 1998
75. Scott Chapman, Pierre Scotto, and Ulrich W. Heinz. *Phys. Rev. Lett.* 74:4400–4403, 1995
76. T. Csorgo, B. Lorstad, and J. Zimanyi. *Z. Phys.* C71:491–497, 1996
77. Ulrich W. Heinz and Q. H. Zhang. *Phys. Rev.* C56:426–431, 1997
78. Ulrich W. Heinz and Alex Sugarbaker. *Phys. Rev.* C70:054908, 2004
79. H. Boggild et al. *Phys. Lett.* B455:77–83, 1999
80. I. G. Bearden et al. *Phys. Lett.* B517:25–31, 2001
81. J. Adams et al. *Phys. Rev. Lett.* 91:262301, 2003
82. M. A. Lisa et al. *Phys. Rev. Lett.* 84:2798–2802, 2000
83. A. N. Makhlin and Yu. M. Sinyukov. *Z. Phys.* C39:69, 1988
84. K. Kolehmainen. *Phys. Lett.* B180:203–206, 1986
85. Sandra S. Padula and Miklos Gyulassy. *Nucl. Phys.* B339:378–392, 1990
86. Ulrich W. Heinz and Peter F. Kolb. *Phys. Lett.* B542:216–222, 2002
87. Boris Tomasik and Ulrich W. Heinz. *Phys. Rev.* C65:031902, 2002
88. Mikolaj Chojnacki, Wojciech Florkowski, and Tamas Csorgo. *Phys. Rev.* C71:044902, 2005
89. T. Csorgo, S. V. Akkelin, Yojiro Hama, B. Lukacs, and Yu. M. Sinyukov. *Phys. Rev.* C67:034904, 2003
90. Kang Seog Lee, Ulrich W. Heinz, and Ekkard Schnedermann. *Z. Phys.* C48:525–541, 1990
91. Ekkard Schnedermann, Josef Sollfrank, and Ulrich W. Heinz. *Phys. Rev.* C48:2462–2475, 1993
92. S. Pratt. *Phys. Rev. Lett.* 53:1219–1221, 1984
93. Yu. M. Sinyukov, S. V. Akkelin, and Nu Xu. *Phys. Rev.* C59:3437–3440, 1999
94. Denes Molnar and Miklos Gyulassy. *Phys. Rev.* C62:054907, 2000
95. Sen Cheng et al. *Phys. Rev.* C65:024901, 2002
96. Peter F. Kolb and Ulrich Heinz. *nucl-th/0305084*
97. P. Huovinen, P. F. Kolb, Ulrich W. Heinz, P. V. Ruuskanen, and S. A. Voloshin. *Phys. Lett.* B503:58–64, 2001
98. Tetsufumi Hirano. *nucl-th/0410017*
99. Ulrich W. Heinz and Peter F. Kolb. *Nucl. Phys.* A702:269–280, 2002
100. Derek Teaney. *Phys. Rev.* C68:034913, 2003
101. Fred Cooper, Graham Frye, and Edmond Schonberg. *Phys. Rev.* D11:192, 1975
102. Boris Tomasik and Urs Achim Wiedemann. *Phys. Rev.* C68:034905, 2003
103. L. P. Csernai, V. K. Magas, E. Molnar, A. Nyiri, and K. Tamosiunas. *hep-ph/0406082*
104. Yu. M. Sinyukov, S. V. Akkelin, and Y. Hama. *Phys. Rev. Lett.* 89:052301, 2002
105. Sven Soff, Steffen A. Bass, and Adrian Dumitru. *Phys. Rev. Lett.* 86:3981–3984, 2001

106. S. A. Bass and A. Dumitru. *Phys. Rev. C* 61:064909, 2000
107. D. Teaney, J. Lauret, and Edward V. Shuryak. *Nucl. Phys. A* 698:479–482, 2002
108. D. Teaney, J. Lauret, and E. V. Shuryak. *nucl-th/0110037*
109. George F. Bertsch. *Phys. Rev. Lett.* 72:2349–2350, 1994
110. Subrata Pal and Scott Pratt. *Phys. Lett.* B578:310–317, 2004
111. P. Danielewicz and G. F. Bertsch. *Nucl. Phys. A* 533:712–748, 1991
112. W.J. Llope et al. *Phys. Rev. C* 52:2004, 1995
113. J. Adams et al. *Phys. Rev. C* 71:044906, 2005
114. M. A. Lisa et al. *nucl-ex/0503017*
115. H. Boggild et al. *Phys. Lett.* B349:386–392, 1995
116. L. Ahle et al. *Phys. Rev. C* 66:054906, 2002
117. S. S. Adler et al. *Phys. Rev. Lett.* 93:152302, 2004
118. H. Boggild et al. *Phys. Lett.* B302:510–516, 1993
119. G. I. Kopylov. *Phys. Lett.* B50:472–474, 1974
120. P. Abreu et al. *Phys. Lett.* B286:201–210, 1992
121. G. Abbiendi et al. *Eur. Phys. J. C* 16:423–433, 2000
122. Jorge Uribe Duque. *Phys. Rev. C* 49:4373, 2003. UMI-93-29679
123. P. Avery et al. *Phys. Rev. D* 32:2294–2302, 1985
124. M. Anderson et al. *Nucl. Instrum. Meth.* A499:659–678, 2003
125. A. V. Stavinskiy, K. R. Mikhailov, A. V. Vlassov, B. Erasmus, and G Renault. *Nukleonika.* 49:S23–S25, 2004
126. M. A. Lisa, W. G. Gong, C. K. Gelbke, and W. G. Lynch. *Phys. Rev. C* 44:2865–2868, 1991
127. M. A. Lisa et al. *Phys. Lett.* B496:1–8, 2000
128. Randall Wells. *Ph.D. thesis, Ohio State University.* 2002
129. J. Adams et al. *Phys. Rev. Lett.* 93:012301, 2004
130. A. M. Poskanzer and S. A. Voloshin. *Phys. Rev. C* 58:1671–1678, 1998
131. N. Borghini and J. Y. Ollitrault. *Phys. Rev. C* 70:064905, 2004
132. I. G. Bearden et al. *Phys. Rev. C* 58:1656–1665, 1998
133. M. G. Bowler. *Phys. Lett.* B270:69–74, 1991
134. Yu. Sinyukov, R. Lednický, S. V. Akkelin, J. Pluta, and B. Erasmus. *Phys. Lett.* B432:248–257, 1998
135. A. D. Chacon et al. *Phys. Rev. C* 43:2670–2688, 1991
136. David A. Brown, Pawel Danielewicz, Mike Heffner, and Ron Soltz. *nucl-th/0404067*
137. C. Adler et al. *Phys. Rev. Lett.* 87:082301, 2001
138. T. Csorgo, S. Hegyi, and W. A. Zajc. *Eur. Phys. J. C* 36:67–78, 2004
139. J. Bartke. *Phys. Lett.* B174:32–35, 1986
140. T. J. Humanic et al. *Z. Phys.* C38:79–84, 1988
141. A. Bamberger et al. *Phys. Lett.* B203:320–326, 1988
142. T. Abbott et al. *Phys. Rev. Lett.* 69:1030–1033, 1992
143. J. Barrette et al. *Phys. Lett.* B333:33–38, 1994
144. D. Adamova et al. *Nucl. Phys. A* 714:124–144, 2003
145. J. Barrette et al. *Phys. Rev. C* 50:1077–1084, 1994
146. J. Barrette et al. *Phys. Rev. C* 60:054905, 1999

147. H. Boggild et al. *Phys. Lett.* B458:181–189, 1999
148. M. A. Lisa et al. *Phys. Rev. Lett.* 70:3709–3712, 1993
149. R. Kotte et al. *Eur. J. Phys.* A23:271–278, 2005
150. M. A. Lisa et al. *Phys. Rev. Lett.* 71:2863–2866, 1993
151. M. A. Lisa et al. *Phys. Rev.* C49:2788, 1994
152. R. Kotte et al. *Z. Phys.* A359:47–54, 1997
153. S. S. Adler et al. *Phys. Rev.* C71:034908, 2005
154. T. Csorgo and B. Lorstad. *Phys. Rev.* C54:1390–1403, 1996
155. S. A. Voloshin and W. E. Cleland. *Phys. Rev.* C54:3212–3217, 1996
156. Henning Heiselberg. *Phys. Rev. Lett.* 82:2052–2055, 1999
157. Henning Heiselberg and Anne-Marie Levy. *Phys. Rev.* C59:2716–2727, 1999
158. K. H. Ackermann et al. *Phys. Rev. Lett.* 86:402–407, 2001
159. C. Pinkenburg et al. *Phys. Rev. Lett.* 83:1295–1298, 1999
160. D. Miskowiec. *Nucl. Phys.* A590:473c–476c, 1995
161. Kirill Filimonov et al. *Nucl. Phys.* A661:198–204, 1999
162. S. Nishimura et al. Talk given at 15th Winter Workshop on Nuclear Dynamics, Park City, UT, 9–16 Jan 1999
163. M. M. Aggarwal et al. *Nucl. Phys.* A663:729–732, 2000
164. Sergei Y. Panitkin. *nucl-ex/9905007*
165. L. P. Csernai and D. Rohrlich. *Phys. Lett.* B458:454, 1999
166. J. Brachmann et al. *Phys. Rev.* C61:024909, 2000
167. V. K. Magas, L. P. Csernai, and D. D. Strottman. *hep-ph/0101125*
168. Edward V. Shuryak. *Phys. Rept.* 61:71–158, 1980
169. J. D. Bjorken. *Phys. Rev.* D27:140–151, 1983
170. Ulrich W. Heinz. *Nucl. Phys.* A610:264c–277c, 1996
171. D. Miskowiec et al. *Nucl. Phys.* A610:227c–239c, 1996
172. H. Appelshauser et al. *Eur. Phys. J.* C2:661–670, 1998
173. F. Antinori et al. *J. Phys.* G27:2325–2344, 2001
174. B. B. Back et al. *nucl-ex/0409001*
175. Thorsten Renk. *Phys. Rev.* C69:044902, 2004
176. M. M. Aggarwal et al. *Phys. Rev.* C67:014906, 2003
177. K. Adcox et al. *Phys. Rev. Lett.* 88:192302, 2002
178. Nu Xu and Masashi Kaneta. *Nucl. Phys.* A698:306–313, 2002
179. H. Beker et al. *Phys. Rev. Lett.* 74:3340–3343, 1995
180. T. Csorgo and B. Lorstad. *Phys. Rev.* C54:1390–1403, 1996
181. Miklos Gyulassy and Sandra S. Padula. *Phys. Rev.* C41:21, 1990
182. S. Schnetzer et al. *Phys. Rev. Lett.* 49:989–992, 1982
183. Y. Akiba et al. *Phys. Rev. Lett.* 70:1057–1060, 1993
184. H. Beker et al. *Z. Phys.* C64:209–218, 1994
185. V. Cianciolo. *Nucl. Phys.* A590:459c–464c, 1995
186. H. Sorge, H. Stocker, and W. Greiner. *Nucl. Phys.* A498:567c–576c, 1989
187. Urs Achim Wiedemann, Pierre Scotto, and Ulrich W. Heinz. *Phys. Rev.* C53:918–931, 1996

188. I. G. Bearden et al. *Phys. Rev. Lett.* 87:112301, 2001
189. S. V. Afanasiev et al. *Phys. Lett.* B557:157–166, 2003
190. H. Appelshauser et al. *Phys. Lett.* B467:21–28, 1999
191. M. M. Aggarwal et al. *Phys. Rev. Lett.* 93:022301, 2004
192. M. Heffner. *J. Phys.* G30:S1043–S1047, 2004
193. S. Bekele. *J. Phys.* G30:S229–S234, 2004
194. G. Renault. *hep-ex/0404024*
195. Joseph I. Kapusta and Y. Li. *J. Phys.* G30:S1069–S1072, 2004
196. Cheuk-Yin Wong. *hep-ph/0403025*
197. R. Lednicky. *nucl-th/0112011*
198. T. Alber et al. *Z. Phys.* C73:443–448, 1997
199. G. Renault. *hep-ex/0406066*
200. C. Blume et al. *Nucl. Phys.* A715:55–64, 2003
201. Adam Kisiel. *J. Phys.* G30:S1059–S1064, 2004
202. G. Bertsch, M. Gong, and M. Tohyama. *Phys. Rev.* C37:1896–1900, 1988
203. Dirk H. Rischke and Miklos Gyulassy. *Nucl. Phys.* A608:479–512, 1996
204. John W. Harris and Berndt Muller. *Ann. Rev. Nucl. Part. Sci.* 46:71–107, 1996
205. D. Adamova et al. *Phys. Rev. Lett.* 90:022301, 2003
206. D. Hardtke and S. A. Voloshin. *Phys. Rev.* C61:024905, 2000
207. H. Sorge. *Phys. Rev.* C52:3291–3314, 1995
208. M. A. Lisa et al. *Nucl. Phys.* A698:185–192, 2002
209. S. V. Afanasiev et al. *Phys. Rev.* C66:054902, 2002
210. S. Kniege et al. *nucl-ex/0403034*
211. D. Ferenc, Ulrich W. Heinz, B. Tomasik, U. A. Wiedemann, and J. G. Cramer. *Phys. Lett.* B457:347–352, 1999
212. C. Greiner, C. Gong, and B. Muller. *Phys. Lett.* B316:226–230, 1993
213. S. Pratt and K. Haglin. *Phys. Rev.* C59:3304–3308, 1999
214. S. V. Akkelin and Yu. M. Sinyukov. *Phys. Rev.* C70:064901, 2004
215. Peter Braun-Munzinger, Krzysztof Redlich, and Johanna Stachel. *nucl-th/0304013*
216. P. Braun-Munzinger, D. Magestro, K. Redlich, and J. Stachel. *Phys. Lett.* B518:41–46, 2001
217. A. Andronic, P. Braun-Munzinger, K. Redlich, and J. Stachel. *Phys. Lett.* B571:36–44, 2003
218. K. Adcox et al. *Phys. Rev. Lett.* 87:052301, 2001
219. Pawel Danielewicz and Scott Pratt. *Phys. Rev.* C53:249–266, 1996
220. A. B. Larionov, M. Effenberger, S. Leupold, and U. Mosel. *Phys. Rev.* C66:054604, 2002
221. T. J. Humanic. *Nucl. Phys.* A715:641–644, 2003
222. T. J. Humanic. *nucl-th/0301055*
223. Denes Molnar and Miklos Gyulassy. *Phys. Rev. Lett.* 92:052301, 2004
224. Zi-wei Lin, C. M. Ko, and Subrata Pal. *Phys. Rev. Lett.* 89:152301, 2002
225. Zi-wei Lin and C. M. Ko. *J. Phys.* G30:S263–S270, 2004
226. Tetsufumi Hirano and Keiichi Tsuda. *Phys. Rev.* C66:054905, 2002

227. Ulrich W. Heinz and Peter F. Kolb. *hep-ph/0204061*
228. D. Zschesche, S. Schramm, H. Stocker, and W. Greiner. *Phys. Rev. C* 65:064902, 2002