This paper was presented at a colloquium entitled "Carbon Dioxide and Climate Change," organized by Charles D. Keeling, held November 13–15, 1995, at the National Academy of Sciences, Irvine, CA.

## Gases in ice cores

MICHAEL BENDER\*, TODD SOWERS<sup>†</sup>, AND EDWARD BROOK\*<sup>‡</sup>

\*Graduate School of Oceanography, University of Rhode Island, Kingston, RI 02881; <sup>†</sup>Department of Geosciences, 447 Deike Building, Pennsylvania State University, University Park, PA 16802; and <sup>‡</sup>Departments of Geology and Environmental Science, Washington State University, Vancouver, WA 98686

ABSTRACT Air trapped in glacial ice offers a means of reconstructing variations in the concentrations of atmospheric gases over time scales ranging from anthropogenic (last 200 yr) to glacial/interglacial (hundreds of thousands of years). In this paper, we review the glaciological processes by which air is trapped in the ice and discuss processes that fractionate gases in ice cores relative to the contemporaneous atmosphere. We then summarize concentration–time records for  $CO_2$  and  $CH_4$  over the last 200 yr. Finally, we summarize concentration–time records for  $CO_2$  and  $CH_4$  during the last two glacial–interglacial cycles, and their relation to records of global climate change.

Ice crystals near the surface of a glacier are compressed by the continual addition of snow at the surface. As the ice crystals travel down into the glacier, they grow and reorient themselves into a closer packing. Density rises, open porosity decreases, and by some depth between 40 and 120 m, the crystals are sintered together into an impermeable mass ("glacial ice") in which about 10% of the volume is composed of isolated bubbles. The gas trapped in these bubbles is close in composition to contemporaneous air and allows us to reconstruct changes in atmospheric chemistry on three important time scales: the last 200 years (relating to anthropogenic change), the Holocene (last 10,000 years), and the last several hundred thousand years (relating to glacial-interglacial cycles). In this paper we discuss the gas trapping process and summarize data on the changes in the greenhouse gas concentrations of air over the three time scales of interest.

## **Physics of Gases in Glaciers**

The density of ice at the surface of an ice sheet is typically 0.3-0.35 g cm<sup>-3</sup>; the corresponding porosity is 62-67%. Settling and packing cause the density to rise rapidly to about 0.55 g cm<sup>-3</sup> by a depth of 10-30 m. Below, recrystallization and other processes drive a somewhat slower increase in density, which continues until individual crystals are fused together into an impermeable mass of glacial ice (1). At the "bubble closeoff depth," about 10-15% of the volume is air, and the density is about 0.81-0.84 g cm<sup>-3</sup> (2, 3). The "firn" is the zone of porous snow and ice above the closeoff depth, and the depth interval in which bubbles close is termed the "firn–ice transition." Below the transition, densification continues by the compression of bubbles due to hydrostatic pressure.

When the snow accumulation rate at the surface of an ice sheet is greater than about 4 cm yr<sup>-1</sup> (expressed as the ice-equivalent thickness of annual layers), discrete seasonal layers of snow are preserved that have characteristic physical and chemical properties. Wintertime layers initially have higher densities than summertime layers. This density contrast

@ 1997 by The National Academy of Sciences 0027-8424/97/948343-72.00/0 PNAS is available online at http://www.pnas.org.

is maintained during the densification process (Fig. 1). The firn-ice transition occurs at the same density for wintertime and summertime layers, but wintertime layers attain the closeoff density at a shallower depth. Consequently, there is an interval of about 10 m in which wintertime layers are more extensively sealed than summertime layers. In this interval, permeable and impermeable layers alternate in the ice sheet (3).

There are three regimes of gas transport in the firn (4, 5). The uppermost layer, which appears to extend down to about 10-m depth at some sites where firn air has been sampled and analyzed, is affected by convective mixing driven by surface wind stress. Underlying the convective zone is the "stagnant air column," in which transport is by molecular and atomic diffusion only. Diffusivities of gases are typically about 1 m<sup>2</sup> day<sup>-1</sup> at 10-m depth. Below, they decrease with increasing density (5), due to a combination of lower porosity and higher tortuosity (the latter factor accounts for the extra distances gas atoms and molecules must travel as they wind their way through the ice crystals to move from one depth to another).

The diffusivity of the firn is such that air at the base of the stagnant column today has a "CO<sub>2</sub> age" ranging from about 6 yr for the GISP2 core (central Greenland) to about 40 years at Vostok (East Antarctica). In point of fact, however, air in firn at a given depth is not of a single age. The composition of firn air is convoluted by a number of processes. Air in the convective zone responds instantaneously to changes in atmospheric chemistry. These changes then propagate down into the stagnant column by molecular or atomic diffusion. Schwander *et al.* (6) calculated that the age of maximum abundance at the base of the firn is about  $0.65 \cdot z_t^2/D$ , where  $z_t$  is the height of the stagnant column and *D* is the free air diffusivity of the gas at the ambient temperature and pressure of the site. A small fraction of the gas is younger, and there is a long tail to older ages.

The diffusivity of an element or compound decreases with increasing mass and increasing atomic or molecular diameter. Thus each element or compound diffuses at a different rate, and each isotope of a compound diffuses at a different rate. In consequence, the covariation between the composition of one gas and another (e.g., CO<sub>2</sub> and CH<sub>4</sub>) in firn is different from their historical covariation in air. The isotopic composition of a gas (e.g.,  $CO_2$ ) in firn air also varies with the concentration of that gas in a way that is different from the historical relationship. The concentrations of gases and isotopes that diffuse most rapidly will be closest to their current atmospheric concentrations. Because light isotopes diffuse more rapidly, the concentration of a gas in firn air will be more depleted in heavy isotopes than was the atmosphere at the time it had the same concentration as a firn air sample. Differential diffusivity is a first-order effect that must be taken into account when interpreting data on the concentration and isotopic composition of gases in firn air and ice cores (7).

Abbreviation: kyr, thousands of years.

<sup>8343</sup> 



FIG. 1. (From Martinerie *et al.*, ref. 3). (*Upper*) Closed porosity vs. density in the firn at Summit, Greenland. Note the monotonic relationship between the two properties. When density = 0.83, open porosity = 0. The decrease in closed porosity at higher densities is due to compression of bubbles. (*Lower*) Closed porosity vs. depth for maximum-density layers and minimum-density layers. Note how maximum-density layers are completely closed by 75-m depth, and minimum-density layers by 83-m depth. In the intervening depth interval, parts of the firn remain permeable but gases cannot migrate vertically.

Several additional factors influence the composition of air in the firn. Gravitational fractionation is one (8-10). The pressure of gases increases with depth below the surface of the firn according to the barometric equation:

$$P/P_z = \exp(mgz/RT)$$

 $m = \text{mass in units of g mol}^{-1}$ , g is the gravitational acceleration constant, z is depth, R is the ideal gas constant, and T is Kelvin temperature.

As Craig *et al.* (8) and Schwander (10) recognized, this equation applies not only to bulk air but to each individual constituent of air in that (dominant) depth interval of the firm where transport is essentially entirely by diffusion (the stagnant air column). The rate at which the enrichment-per-mass unit increases with depth, expressed in the  $\delta$  notation, is  $(\Delta mg/RT - 1)\cdot 1,000$ , or about 0.005%/amu per meter at typical firm air temperatures. The relative enrichment with depth for different species is directly proportional to the mass difference. The firm air data for the GISP2 site, central Greenland, demonstrate the expected enrichment for the  $\delta^{15}N$  of N<sub>2</sub> (Fig. 2). The enrichment or depletion is significant for nearly all species, corresponding to 3 ppmv of CO<sub>2</sub> at the base of deep firm profiles, for example.

Seasonal changes in the concentrations of gases in air cause seasonal variations in firn air chemistry. The magnitude of these variations relative to their secular trends depends on location and property. The effect is perhaps largest for O<sub>2</sub>, CO<sub>2</sub>, and  $\delta^{13}$ C of CO<sub>2</sub> in Greenland. Seasonal variations are damped out with depth and become very small below 30–50 m.

Thermal fractionation also affects the isotopic and elemental composition of firn air. Severinghaus (11) and Severinghaus et al. (12) first recognized the importance of thermal fractionation in porous environmental media in their studies of the composition of air in sand dunes. Temperature gradients cause fractionation, with heavier gases or isotopes being enriched in colder regions. For <sup>15</sup>N, the fractionation is about 0.025%/°C. Thermal fractionation is large in firn because gases diffuse faster than heat. In consequence, steep seasonal temperature gradients occur in the upper  $\approx 5$  m of the firn and gases nearly equilibrate with these temperature gradients. This effect produces large seasonal variations in isotopic compositions and in the  $O_2/N_2$  ratio in the top few meters of the firn. The seasonal anomalies decrease with depth, and for most species are insignificant below 30 m. O2 is an exception; the concentration of this gas in air is changing so slowly (on a percentage basis) that seasonal thermal gradients are significant down to 60-m depth.

These processes combine to influence the composition of gas throughout the firn, and at its base where gases are trapped as bubbles in impermeable ice. Here, two modes of trapping are possible. First, seasonal layering may be absent and air may be trapped throughout the bubble closeoff zone. In this case, the composition of the bulk trapped gases in ice cores will be further convoluted because of the finite closeoff interval. At Vostok, for example, the bubble closeoff zone is about 8 m



FIG. 2.  $\delta^{15}N$  (*Left*) and CH<sub>4</sub> (*Right*) vs. depth in firn air from the GISP2 site at Summit, Greenland. The subsurface maximum in  $\delta^{15}N$  is due to thermal fractionation (<sup>15</sup>N is enriched at 5- and 10-m depth because the firn at these depths, which remains at the mean annual temperature, is colder than air at the surface during summertime, when sampling was done). The increase below 20-m depth is due to gravitational fractionation. CH<sub>4</sub> decreases very slowly to the top of the bubble closeoff zone at 70-m depth. Below it decreases very rapidly because gases cannot migrate vertically and the age of the gas in the firn increases as rapidly as the age of the ice (about 4 yr/m).

thick. A single layer of ice traps bubbles throughout the  $\approx 300$ yr it moves through this zone. This process accounts for the largest share of the dispersion of gas ages in a single sample of ice. At Summit, Greenland, on the other hand, high-density layers are completely sealed as they pass through the top of the bubble closeoff zone. Sealing forms vertically impermeable layers that prevent additional diffusive mixing and "locks in" the composition of gas present in the open, intervening, low-density layers. In such a case, individual ice samples can resolve time periods as short as a decade. The influence of lock-in at Summit can be seen from the firn air data at GISP2 (Fig. 2).  $\delta^{15}$ N ceases to rise below 70 m depth. CH<sub>4</sub> and CO<sub>2</sub> concentrations fall rapidly below this depth; CO<sub>2</sub> and CH<sub>4</sub> "ages" of firn air increase about as rapidly as ice ages below 70 m (4 yr/m). These firn air results are similar to those of Schwander et al. (4), who developed the method for firn air sampling and studied the composition of air in the firn at the nearby GRIP site.

The final process influencing the composition of gas in polar firm and ice cores is effusion. Craig *et al.* (8) suggested this process to account for the depletion of  $O_2$  and Ar relative to  $N_2$  in polar ice samples, as measured first by Raynaud and Delmas (13) and later by Craig *et al.* (8) and Sowers *et al.* (9). Craig *et al.* (8) pointed out that  $O_2$  and Ar, with diameters of about 3 Å, were smaller than  $N_2$  (diameter 3.3 Å). They suggested that cracks and imperfections in the ice of about 3 Å spacing would allow the effusive loss of  $O_2$  and Ar while selectively retaining nitrogen. It is not clear whether effusion takes place *in situ* during bubble closeoff or after ice cores are retrieved. Some indications for the former possibility are that deep firm air samples are enriched in  $O_2$  and Ar relative to  $N_2$ , and that the  $O_2$  and Ar depletions of ice core samples do not increase after cores have been on the surface for a few days.

The processes affecting gases in ice cores need to be taken into account in reconstructions of the composition of the past atmosphere. First, measured concentrations of gases in ice cores and firn air need to be corrected for effects of gravitational fractionation and, where appropriate, thermal fractionation. Second, gas records are useful only when dated absolutely or on a time scale common to other records. Because bubbles close at depths of 40-120 m, gases are younger than the ice enclosing them. The gas age-ice age difference ( $\Delta$ age) is as great as 7 kyr in glacial ice from Vostok; it is as low as 30 yr in the rapidly accumulating Antarctic core DE 08. There are substantial uncertainties associated with  $\Delta$ age, limiting our ability to interpret some records. This is not a problem when reconstructing the anthropogenic transient from ice core studies, because one can align the recent part of ice core records with direct observations and assume that  $\Delta$ age is constant below the interval of overlap.

Once trapped in bubbles, air in ice cores is subject to two additional processes. First, bubbles are compressed under hydrostatic pressure (14). Second, gases eventually begin to dissolve in the ice as air hydrates (14, 15). Nucleation is kinetically limited, and at ambient temperatures and pressures occurs over order  $10^4$  yr (15). As a result, air hydrates form at depths of about 400–1,500 m; cold temperatures and slow accumulation favor formation at shallower depths.

Different gases form clathrates at different pressures (16). In the long zone over which air hydrates and bubbles coexist in ice cores, the composition of gases in bubbles must be different from the composition in bulk ice. When gas samples are extracted after crushing of ice for a sufficiently long period of time, as is done for CO<sub>2</sub> analysis, individual compounds are apparently not fractionated despite the fact that overall extraction efficiency is <100%. The evidence for this statement is that coherent records of CO<sub>2</sub> are obtained by analyzing different ice cores, despite the fact that the dissolution of gases occurs in samples of different ages.

## Reconstructions of the Anthropogenic Transient from Ice Core and Firn Air Chemistry

**CO**<sub>2</sub>. The most extensive study of the preindustrial CO<sub>2</sub> concentration of air and its anthropogenic rise is that of Etheridge *et al.* (17). Their results are based largely on studies of the DE 08 ice core, from Law Dome, Antarctica (66° 43′ S, 113° 12′ E; elevation 1,250 m). The high accumulation rate, about 1.2 m/yr, and warm annual temperature ( $-19^{\circ}$ C) at the site of this core (which causes the closeoff depth to be relatively shallow) allow time to be resolved exceptionally well. Etheridge *et al.* (17) estimate the gas age–ice age difference to be only 30 yr and the duration of the bubble closeoff process to be 8 yr. They supplement their results from DE-08 with the nearby DSS core and compare their CO<sub>2</sub>–time curve with that from Siple Dome, analyzed by Neftel *et al.* (18).

Results are summarized in Fig. 3. The CO<sub>2</sub> concentration of air varies in the range  $275 \pm 5$  ppm between 1350 and 1800. The anthropogenic rise begins between 1780 and 1870, the later age corresponding to the time when CO<sub>2</sub> clearly begins rising above the highest level of the previous 500 yr. As inferred from the study of the DE 08 core, CO<sub>2</sub> generally rises at a rate that tends to increase with time. The most dramatic exception is the interval of constant concentration inferred for the period between 1935 and 1945. The DE 08 record is in good agreement with the DSS and Siple data. The exception is the period between 1935 and 1945, when CO<sub>2</sub> concentrations from the latter two cores fall below those of DE 08 by up to 5 ppmv.

Box-diffusion model deconvolutions, which assume  $CO_2$  concentrations to be in good agreement with the Siple results, invoke a net biosphere release of about 0.1–0.5 Gt C/yr during this period (20, 21). Etheridge *et al.* (17) suggest that one can balance the carbon cycle between 1935 and 1945 with a combination of



FIG. 3. (*Upper*)  $CO_2$  vs. time before present, as inferred by Etheridge *et al.* (17) from ice core studies. (*Lower*)  $CH_4$  vs. time before present, as inferred by Etheridge *et al.* (19) from ice core studies.



FIG. 4. Selected ice core records for the last 350 kyr. Data from refs. 34-37.

lower respiration, increased growth due to  $CO_2$  fertilization and changes in temperature and precipitation, and suppressed  $CO_2$ input to the atmosphere as a result of exceptional El Nino events. Otherwise the new data of Etheridge *et al.* (17) are consistent with the Siple  $CO_2$  data and estimates of carbon fluxes derived therefrom. These invoke rising  $CO_2$  beginning around 1780, with the biosphere as the major source. Fossil fuel becomes the dominant source around 1910. The biosphere continues to be a source until 1940 in the deconvolution of Siegenthaler and Oeschger (20), when it becomes neutral, and 1970 in the deconvolution of Keeling *et al.* (21), when it becomes a sink. **CH4.** Two recent records (19, 22) supplement earlier work of Craig and Chou (23), Etheridge *et al.* (24), Khalil and Rasmussen (25), and Stauffer *et al.* (26) to provide a detailed curve of the anthropogenic increase in atmospheric CH4. The results of Blunier *et al.* (22) suggest a stable preindustrial value of about 720 ppbv over Antarctica between about 1000 A.D. and 1750 A.D. The Antarctic concentration rise after 1840, as measured in the DE 08 core by Etheridge *et al.* (24), is shown in Fig. 3.

The factors maintaining the preanthropogenic CH<sub>4</sub> background, and those responsible for the anthropogenic increase,



FIG. 5. Selected ice core records for the last 100 kyr. Data from refs. 30, 34, 35, and 38.

have been discussed in a number of papers (refs. 27 and 28 and citations therein). According to Chappellaz *et al.* (29, 30), preindustrial CH<sub>4</sub> production was dominated by tropical wetlands, with grassland and temperate to boreal forests also making significant contributions. They argue that the natural source has diminished by 20% since preindustrial times. An-thropogenic sources are, of course, far greater. Chappellaz *et al.* (29) estimate that an anthropogenic decrease of about 20% in the atmospheric OH concentration has caused the atmospheric lifetime of CH<sub>4</sub> to rise and thereby contributed to the anthropogenic increase.

**Other Gases.** Studies of trapped gases in ice cores have revealed preanthropogenic levels of two other properties. The preanthropogenic N<sub>2</sub>O concentration is estimated to have been about 275–285 ppbv, compared with a contemporary value of about 311 ppbv (24, 31, 32). Because of its high solubility, the precise measurement of this gas in ice cores is very difficult. The  $\delta^{13}$ C of CO<sub>2</sub> over Antarctica decreased from a preindustrial level of about -6.5% to approximately -8.0% today (33). The decrease roughly mirrors the CO<sub>2</sub> increase.

Firn Air Studies. Sampling and analysis of gas in the firn is an emerging approach for improving existing records of anthropogenic transients (4, 7). Firn air sampling allows one to collect 1,000 liters or more of isotopically unfractionated air as old as 90 yr. These samples can be used to measure concentrations of trace organic compounds,  $O_2/N_2$  ratios, and the isotopic composition of greenhouse gases, all measurements that are far more difficult to make on ice core samples.

## Glacial-Interglacial Changes in Atmospheric Chemistry Recorded in Ice Cores

Selected climate records are summarized in Figs. 4 and 5, covering the periods from 0 to 350 kyr and from 0 to 100 kyr before the present (B.P.), respectively. The  $\delta^{18}$ O of calcitic foraminifera from deep sea sediments is a proxy indicator for ice volume. The  $\delta$ D or  $\delta^{18}$ O of ice from ice cores is a proxy indicator of temperature in the area of the ice core. Recent inversions of borehole temperature data conclude that the glacial–interglacial temperature change in Greenland was about 20°C (39, 40). The dust content of ice cores is a proxy for dryness of the source areas. High dust contents are generally attributed to dry source areas from which dust is readily suspended into the atmosphere. CO<sub>2</sub> and CH<sub>4</sub> concentrations are measured in bubbles of polar ice as described above. June solar insolation at 65° N latitude is plotted at the bottom of the figure.

On a broad scale, the atmospheric  $CO_2$  concentration is clearly linked to climate change. Lorius *et al.* (41) argued that the glacial–interglacial  $CO_2$  change (with a small contribution from  $CH_4$ ) was responsible for a 2°C temperature change when all feedbacks are considered. Obviously this number is highly uncertain but it is very likely that increasing  $CO_2$  concentrations in air contributed to global warming during glacial terminations 1 and 2. The link between  $CO_2$  and higherfrequency climate changes is weak. For example, there is no minimum in  $CO_2$  corresponding to the cold period (Glacial Stage 5d) at about 110 kyr B.P. Similarly,  $CO_2$  does not rise during the interstadial events of the past 35 kyr (ref. 42; Fig. 5).

Many factors have been invoked to explain the glacialinterglacial change in the CO<sub>2</sub> concentration of air. The cause of this change was vigorously debated for a decade beginning in 1982. Recently attention to this subject has diminished, more because the protagonists are exhausted than because the issue is resolved. It is widely accepted that the  $pCO_2$  of the atmosphere is regulated by the  $pCO_2$  of surface seawater, because about 99% of the  $CO_2$  in the ocean/atmosphere system resides in the oceans. In his classic paper initiating the debate, Broecker (43) noted that lower glacial temperatures would cause the  $\dot{CO}_2$  concentration of air to fall, whereas higher salinities would cause  $CO_2$  to rise. A glacial ocean temperature decrease of 1.5°C would cause CO<sub>2</sub> to fall by 20 ppmv. This change would be roughly offset by a salinity increase of 1%. Recent controversial results showing that equatorial temperatures may have increased by up to 5°C suggest that the glacial-interglacial change in sea surface temperatures may have been much larger than Broecker initially estimated. If this is correct, the temperature change may have contributed 20-30 ppm of the glacial-interglacial change in atmospheric pCO2 after correcting for the compensating effect of salinity (the temperature dependence of pCO2 on sea surface temperature is about 13 ppmv/°C; ref. 43).

Clearly other factors must contribute to the glacialinterglacial difference in atmospheric  $pCO_2$ . These factors must act either by decreasing the total  $CO_2$  concentration of surface seawater or by increasing the surface alkalinity. Changes in ocean circulation, the nutrient/carbon ratio of organic matter, and varying rates of nutrient utilization in Antarctic waters have been invoked to account for the changing surface water total  $CO_2$  concentrations. Removal of CaCO<sub>3</sub> during sea level rise, changing ratios of CaCO<sub>3</sub> to organic carbon in biogenic debris falling out of the surface ocean, and changes in the calcium carbonate compensation depth in the deep ocean have been invoked to change the alkalinity of surface waters.

There have been large changes in the CH<sub>4</sub> concentration of air during the last  $\approx 200$  kyr (Figs. 4 and 5). CH<sub>4</sub> changes are very closely linked with changes in climate (29, 44). CH<sub>4</sub> differs from CO<sub>2</sub> in two important ways. First, CO<sub>2</sub> changes are about five times more important than those of CH<sub>4</sub> in driving glacial-interglacial temperature changes (41). Second, CH<sub>4</sub> changes, which are believed to be caused mainly by changes in emission rates, are somewhat better understood and serve much better as a climate proxy. Chappellaz *et al.* (27) generally interpreted changes in atmospheric CH<sub>4</sub> as reflecting increases in the extent of low-latitude wetlands due to increased precipitation and perhaps temperatures.

Recent records of atmospheric CH<sub>4</sub> variations over the last 200 kyr (Figs. 4 and 5) reveal several very interesting features. As noted above, there is a large glacial-interglacial change, with Antarctic CH<sub>4</sub> concentrations varying between glacial levels of 350 ppbv and interglacial levels of 700 ppbv. There are changes of somewhat smaller magnitude associated with climate changes linked to the 20-kyr climate cycles associated with precession. Examples are CH<sub>4</sub> maxima associated with warm periods at 80 kyr and 100 kyr B.P. Third, there are large CH4 variations associated with interstadial events of the last 40 kyr. These events, recorded most dramatically in Greenland ice cores (e.g., ref. 30; Fig. 5), reflect rapid warming over Greenland, slow cooling, and rapid cooling back to baseline glacial values. During the last 40 kyr at least, interstadial warmings are accompanied by increases of about 150 ppbv in the CH<sub>4</sub> concentration of air. Chappellaz et al. (30) invoked wetter and warmer conditions in the tropics to explain the CH4 increases. Support for this idea comes from the fact that the longer interstadial events, at least, are recorded in Antarctica, illustrating their global nature (45), and are sometimes coincident with deposition of cave deposits in Botswanaland, which Holmgren et al. (46) linked to increased tropical precipitation. Finally, the CH<sub>4</sub> concentration of the atmosphere was surprisingly variable during the Holocene. Methane was at a maximum early in the Holocene, fell to a broad minimum about 5 kyr B.P., and has risen during the last 3 kyr (22). Only the abrupt minimum at about 8.2 kyr B.P. has been directly linked to a climate event, namely an abrupt cooling that occurred at that time. The causes of the mid-Holocene minimum and subsequent slow rise are topics of current investigation.

This research was supported by grants from the New England Regional Center of the National Institute of Global Environmental Change, Department of Energy (Subagreement 901214-HAR), and the Office of Polar Programs, National Science Foundation.

- 1. Herron, M. M. & Langway, C. C., Jr. (1980) J. Glaciol. 25, 373–385.
- Raynaud, D. & Lorius, C. (1977) in *Proceedings of the Grenoble* Symposium, August/September 1975 (International Association of Hydrological Sciences, United Kingdom), Vol. 118, pp. 326–335.
- Martinerie, P., Raynaud, D. Etheridge, D. M., Barnola, J.-M. & Mazaudier, D. (1992) *Earth Planet. Sci. Lett.* 112, 1–13.
- Schwander, J., Barnola, J.-M., Andrie, C., Leuenberger, M., Ludin, A., Raynaud, D. & Stauffer, B. (1993) J. Geophys. Res. 98, 2831–2838.
- Sowers, T., Bender, M., Raynaud, D. & Korotkevich, Y. S. (1992) J. Geophys. Res. 97, 15683–15697.
- Schwander, J., Stauffer, B. & Sigg, A. (1988) Annu. Glaciol. 10, 141–145.
- Battle, M., Bender, M., Sowers, T., Tans, P., Butler, J., Elkins, J., Ellis, J., Conway, T., Zhang, N., Lang, P. & Clarke, A. (1996) *Nature (London)* 383, 231–235.
- 8. Craig, H., Horibe, Y. & Sowers, T. (1988) Science 242, 1675–1678.
- Sowers, T., Bender, M. & Raynaud, D. (1989) J. Geophys. Res. 94, 5137–5150.

- Schwander, J. (1989) in *The Environmental Record in Glaciers and Ice Sheets*, eds. Oeschger, H. & Langway, C. (Wiley, New York).
- 11. Severinghaus, J. P. (1995) Ph.D. dissertation (Columbia University, New York).
- Severinghaus, J. P., Bender, M. L., Keeling, R. F. & Broecker, W. A. (1996) *Geochim. Cosmochim. Acta* **60**, 1005–1018.
- Raynaud, D. & Delmas, R. (1977) in Proceedings of the Grenoble Symposium, August/September 1975 (International Association of Hydrological Sciences, United Kingdom), Vol. 118, pp. 377–381.
- 14. Gow, A. J. & Williamson, T. (1975) J. Geophys. Res. 80, 5101– 5108.
- 15. Price, P. B. (1975) Science 267, 1802–1804.
- 16. Miller, S. L. (1969) Science 165, 489-490.
- Etheridge, D. M., Steele, L. P., Langenfelds, R. L., Francey, R. J., Barnola, J.-M. & Morgan, V. I. (1995) *J. Geophys. Res.* 101, 4115–4128.
- 18. Neftel, A., Moor, E., Oeschger, H. & Stauffer, B. (1985) *Nature* (*London*) **315**, 45–47.
- Etheridge, D. M., Pearman, G. I. & Fraser, P. J. (1992) *Tellus* 44B, 282–294.
- 20. Siegenthaler, U. & Oeschger, H. (1987) Tellus 39B, 140-154.
- Keeling, C. D., Bacastow, R. B., Carter, A. F., Piper, S. C., Whorf, T. P., Heimann, M., Mook, W. G. & Roeloffzen, H. (1989) in *Aspects of Climate Variability in the Pacific and Western Americas*, Geophysical Monograph, ed. Peterson, J. H. (Amer. Geophysical Union, Washington, DC), Vol. 55, pp. 165–236.
- 22. Blunier, T., Chappellaz, J., Schwander, J., Stauffer, B. & Raynaud, D. (1995) Nature (London) 374, 46-49.
- 23. Craig, H. & Chou, C. C. (1982) Geophys. Res. Lett. 9, 1221-1224.
- Etheridge, D. M., Pearman, G. I. & de Silva, F. (1988) Annu. Glaciol. 10, 28–33.
- 25. Khalil, M. A. K. & Rasmussen, R. A. (1987) *Atmos. Environ.* 21, 2445–2452.
- Stauffer, B., Fischer, G., Neftel, A. & Oeschger, H. (1985) Science 229, 1386–1388.
- 27. Aselmann, I. & Crutzen, P. J. (1989) J. Atmos. Chem. 8, 307-358.
- 28. Fung, I., John, J., Lerner, J., Matthews, E., Prather, M., Steele,
- L. P. & Fraser, P. J. (1991) J. Geophys. Res. 96, 13033-13065.
- 29. Chappellaz, J., Fung, I. & Thompson, A. (1993) Tellus 45B, 228-241.

- Chappellaz, J., Blunler, T., Raynaud, D., Barnola, J. M., Schwander, J. & Stauffer, B. (1993) *Nature (London)* 366, 443–445.
- 31. Khalil, M. A. K. & Rasmussen, R. (1988) Annu. Glaciol. 10, 73–80.
- 32. Machida, T., Nakazawa, T., Fujii, Y., Aoki, S. & Watanabe, O. (1995) *Geophys. Res. Lett.* **22**, 2921–2924.
- Friedli, H., Lotscher, H., Oeschger, H., Siegenthaler, U. & Stauffer, B. (1986) *Nature (London)* 324, 237–238.
- Shackleton, J. H. & Pisias, N. G. (1985) in *The Carbon Cycle and Atmospheric CO<sub>2</sub>: Natural Variations Archean to Present*, Geophysical Monograph Series, eds. Sundquist, E. T. & Broecker, W. S. (Amer. Geophysical Union, Washington, DC), Vol. 32, pp. 303–317.
- Jouzel, J., Barkov, N. I., Barnola, J. M., Bender, M. L., Chappellaz, J., Genthon, C., Kotlyakov, V. M., Lipenkov, V., Lorius, C., Petit, J. R., Raynaud, D., Raisbeck, G., Ritz, C., Sowers, T., Stievenard, M., Yiou, F. & Yiou, P. (1993) *Nature (London)* 364, 407–412.
- Barnola, J. M., Pimienta, P., Raynaud, D. & Korotkevich, Y. S. (1991) *Tellus* 43B, 83–90.
- Petit, J. R., Mounier, L., Jouzel, J., Korotkevich, Y., Kotlyakov, V. & Lorius, C. (1990) Nature (London) 343, 56–58.
- Grootes, P. M., Stuvier, M., White, J. W. C., Johnsen, S. & Jouzel, J. (1993) Nature (London) 366, 552–555.
- Cuffey, K., Alley, R., Grootes. P., Bolzan, J. & Anandakrishnan, S. (1994) J. Glaciol. 40, 341–349.
- Johnsen, S., Dahl-Jensen, D., Dansgaard, W. & Gundestrup, N. (1995) *Tellus* 47B, 624–629.
- 41. Lorius, C., Jouzel, J., Raynaud, D., Hansen, J. & Le Treut, H. (1990) *Nature (London)* **347**, 139–145.
- 42. Sowers, T. & Bender, M. (1995) Science 269, 210-214.
- 43. Broecker, W. (1982) Geochim. Cosmochim. Acta 46, 1689-1705.
- Chappellaz, J., Barnola, J. M., Raynaud, D., Korotkevich, Y. S. & Lorius, C. (1990) *Nature (London)* 345, 127–131.
- Bender, M., Sowers, T., Dickson, M. L., Orchardo, J., Grootes, P., Mayewski, P. A. & Meese, D. A. (1994) *Nature (London)* 372, 663–666.
- Holmgren, K., Karlen, W. & Shaw, P. A. (1995) Quat. Res. 43, 320–328.