

1 **Geant4 Monte Carlo simulation of absorbed dose and radiolysis yields enhancement from**
2 **a gold nanoparticle under MeV proton irradiation**

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36 **Abstract**

37 Gold nanoparticles have been reported as a possible radio-sensitizer agent in radiation therapy
38 due to their ability to increase energy deposition and subsequent direct damage to cells and
39 DNA within their local vicinity. Moreover, this increase in energy deposition also results in
40 high yields of chemical species which have been previously shown to significantly contribute
41 to cellular damage. In this work we present, for the first time, an in-silico investigation utilising
42 the general purpose Monte Carlo simulation toolkit Geant4 into energy deposition and radical
43 species production around a spherical gold nanoparticle 50 nm in diameter via proton
44 irradiation. Simulation were performed for incident proton energies ranging from 2 to 170
45 MeV, which are of interest for clinical proton therapy.

46 **Keywords**

47 Geant4-DNA, Gold nanoparticle, radiolysis, proton beam, radiation therapy, radiotherapy,
48 Monte Carlo simulation, radiation chemistry;

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57 **1. Introduction**

58 In order to identify potential new approaches for improving clinical outcomes in radiation
59 therapy, and reduce collateral effects, further study of the use of ionizing radiation in
60 healthcare is needed.

61 One of the most promising new methods in clinical radiation therapy is the delivery of
62 chemical compounds made of high-Z materials to the site of cancerous cells for which ionizing
63 radiation interaction cross sections are significantly higher than in liquid water, the main
64 compound of biological media. As a result, the absorbed dose from irradiation may be
65 enhanced in tumor tissues. For instance, the use of high-Z materials as radio-sensitizers agents
66 was described by Matsudaira *et al.* in 1980 [1], who measured a radio-enhancing effect of
67 iodine on cultured cells irradiated with X-rays but not with gamma rays. The absorbed dose
68 from low energy photon irradiation (X-rays) was increased due to a larger photoelectric cross
69 section in such materials compared to water. This work was then followed by the studies of
70 refs [2] to [6] which utilised iodine and barium containing contrast agents in diagnostic
71 imaging as radio-sensitizers in order to demonstrate their therapeutic advantage in tumors
72 irradiated with low energy photons.

73 Due to their small size, gold nanoparticles (GNP) have been found to easily penetrate cells [7].
74 Exposure of human cells to such nanoparticles (NP) does not cause cytotoxicity [8]. In the
75 initial studies of gold nanoparticles using X-rays, Regulla *et al.* [9] found a factor of 100 of
76 dose enhancement in tissue-equivalent polymethylmethacrate close to the surface of a thin
77 metallic gold foil. In experimental studies on mice, Herold *et al.* [10] and Hainfeld *et al.* [11]
78 have found an increased biologically effective dose thanks to the use of gold nanoparticles in

79 X-ray therapy. For nearly a decade, one has observed *in vivo* the increase of DNA double-
80 strand breaks in irradiated cell populations loaded with gold nanoparticles [12]. The *in vivo*
81 study of Liu *et al.* [13] demonstrated a ~2-45% decrease of cell survival after irradiation using
82 polyethylene-glycol-Au nanoparticles. Recently, Kim *et al.* [14] underlined the role of
83 secondary electrons and characteristic X-rays emitted from metallic nanoparticles irradiated by
84 protons and observed *in vivo* complete tumor regression increase with dose in mice tumors, as
85 well as *in vitro* intracellular reactive oxygen species level increase with dose.

86 In order to understand the possible mechanisms involved in the observed radio-sensitizing
87 effect induced by GNPs under ionizing irradiation, various studies based on Monte Carlo
88 simulations have been undertaken. Such simulations have the potential to fully describe
89 interaction processes of ionizing radiation in biological materials [15]. These investigations
90 focus on the estimation, at the nano-scale, of energy deposition in the vicinity of a single GNP
91 [16] [17] [18]. For instance, Lechtman *et al.* [16] compared Monte Carlo simulations at the
92 nano-scale with experimental macroscopic dose enhancement predictions to demonstrate the
93 influence of gold nanoparticle intracellular localization. A good agreement was shown between
94 experimental survival and the Monte Carlo-based AuNP radiosensitization predictive model
95 (ARP), which takes into account the detailed energy deposition at the nano-scale. Based on
96 Monte Carlo simulations using the BEAMnrc/DOSXYZnrc codes and the interaction cross
97 sections of the PEGS4 and EGS4 codes, Cho [19] has shown a dependence of the dose
98 enhancement factor (DEF) with the concentration of GNPs within the tumor. In the approach
99 of Jones *et al.* [20], the Monte Carlo code EGSnrc was used to quantify the dose distribution
100 from secondary electron dose point kernels within a GNP-loaded tumor; they demonstrated that
101 the microscopic dose around a GNP is enhanced by factors up to more than 100. More recently,
102 Lin *et al.* [21] performed a comparison of dose in the vicinity of a single GNP and a water
103 nanoparticle (WNP) for various energies of incident photons and protons in connection to
104 radiotherapy. They predicted a dose enhancement up to 14, independently of the proton energy.

105 Ionizing particles that pass through biological materials can create reactive oxygen species
106 which can participate in unwanted reactions resulting in cell damage, during the so-called
107 “chemical” and “biological” stages lasting from a few picoseconds to many months after
108 irradiation. This was reported *in vivo* in the experiment of Baluchamy *et al.* [22] which showed
109 that proton irradiation alters oxidant and antioxidant levels in cells.

110 The description of the physical interactions, such as electronic excitations and ionizations,
111 between the ionizing particles and the medium, as well as the rearrangement of the targeted
112 molecules and the subsequent chemistry, can enable us to estimate numerically the amount and
113 types of the generated radiolytic species. Among these species, two, namely the hydroxyl
114 radical and the solvated electron, are created in a significant amount and are particularly
115 reactive. Karamitros *et al.* [23], in the framework of the Geant4-DNA project [24], reported the
116 development of a mechanistic model of radiation chemistry in Geant4, in particular for the
117 evaluation of the yields of hydroxyl radicals and solvated electrons. The potential of
118 enhancement of reactive species production due to an increase of the relative dose distribution
119 around GNPs in an aqueous environment has been recently mentioned in the work of Kim *et*
120 *al.* [14]. In view of the findings of Kim *et al.* [14], regarding the enhancement of reactive
121 species production around GNP, and in order to further understand the role of GNP, we used in
122 this work for the first time the Geant4 general purpose Monte Carlo simulation toolkit [25] and
123 its very low energy “Geant4-DNA” extension to simulate absorbed dose distributions and
124 production of chemical species generated in the vicinity of a single GNP irradiated by a proton.
125 These quantities are estimated as a function of the distance from the NP. The results are
126 obtained for incident mono-energetic protons ranging from 2 MeV to 170 MeV impacting the
127 GNP and are compared to the results obtained when the GNP is replaced by liquid water (water
128 nanoparticle, noted as WNP).

129 Section 2 presents a brief description of the Geant4 simulation toolkit and outlines our
130 implementation with specific emphasis on the selected physics processes. Section 3 then

131 presents and discusses the results of our simulations focusing on the impact of the GNP in
132 liquid water on the observed absorbed dose (Section 3.1) and radiolysis enhancement (Section
133 3.2). Finally, an overall conclusion follows in Section 4.

134 2. Materials and methods

135 2.1 The simulation toolkit

136 Geant4 is a general purpose and open source Monte Carlo toolkit allowing the simulation of
137 particle-matter interactions. It is nowadays utilized in a variety of application domains, from
138 high energy physics, to aerospace and medical physics [25] [26]. Users can simulate physical
139 interactions by specifying to Geant4 the list of physics “processes” and “models” for all
140 particles involved in the simulation. A physics “process” describes a physical interaction (such
141 as ionization, multiple scattering...) and can evoke several physics “models” (which can be
142 fully theoretical, semi-empirical...). Such models compute the physical interaction total cross
143 section and fully describe the final state of the colliding system, including the production of
144 secondary particles, energy loss and emission angles.

145 Geant4 has been recently extended to microdosimetry and nanodosimetry applications in liquid
146 water at very low energies and submicrometer scale [27]-[28] in the framework of the Geant4-
147 DNA project. As a component of Geant4 electromagnetic physics, Geant4-DNA currently
148 simulates the dominant physical interactions of electrons, hydrogen and helium atoms with
149 charge states (H^0 , H^+ , He^0 , He^+ , He^{2+}) in liquid water down to very low energies [24], [29]. The
150 physico-chemical and chemical module in Geant4-DNA [30] was released for the first time in
151 Geant4 10.1 (December 2014). This module is intended for the simulation of radiolytic species
152 production, their diffusion, and their mutual interactions in liquid water following the modeling
153 of physical interactions [23]. In this study, we use Geant4-DNA to calculate the energy
154 deposition and the distribution of chemical species in liquid water surrounding the
155 nanoparticles. In parallel, Geant4 electromagnetic physics **are** used for the simulation of
156 physical interactions in GNP and WNP, as ad-hoc low energy physics models are not currently

157 available in Geant4-DNA for materials other than liquid water. The list of Geant4
 158 electromagnetic processes and models used in this study is given in Table 1 for protons,
 159 electrons and photons. This combination is possible since Geant4-DNA physics and Geant4
 160 electromagnetic physics use the same software design. The “Livermore” set [31] of physics
 161 models in Geant4 low energy electromagnetic physics was selected for the simulation of
 162 electron interactions in the NP. A step size limit of 1 nm and secondary particle production
 163 threshold of 13.6 eV (the first ionization level of H) were selected. Atomic de-excitation was
 164 activated in order to simulate fluorescence, Auger electron production and particle induced X-
 165 ray emission (PIXE). **No production threshold has been applied on atomic de-excitation**
 166 **products.** The interactions of secondary photons are described by the “Livermore” set of
 167 physics models.

Physical interaction	Geant4 process class	Geant4 model class	Energy range
Electron			
Multiple Scattering	G4eMultipleScattering	G4UrbanMscModel	< 100 MeV
		G4WentzelVIModel	
Coulomb Scattering	G4eCoulombScattering	G4eCoulombScatteringModel	> 100 MeV
Ionization	G4eIonisation	G4LivermoreIonisationModel	< 0.1 MeV
		G4MollerBhabhaModel	> 0.1 MeV
Bremsstrahlung	G4eBremsstrahlung	G4LivermoreBremsstrahlungModel	
Proton			
Multiple Scattering	G4hMultipleScattering	G4UrbanMscModel	
Ionization	G4hIonisation	G4BraggModel	< 2 MeV

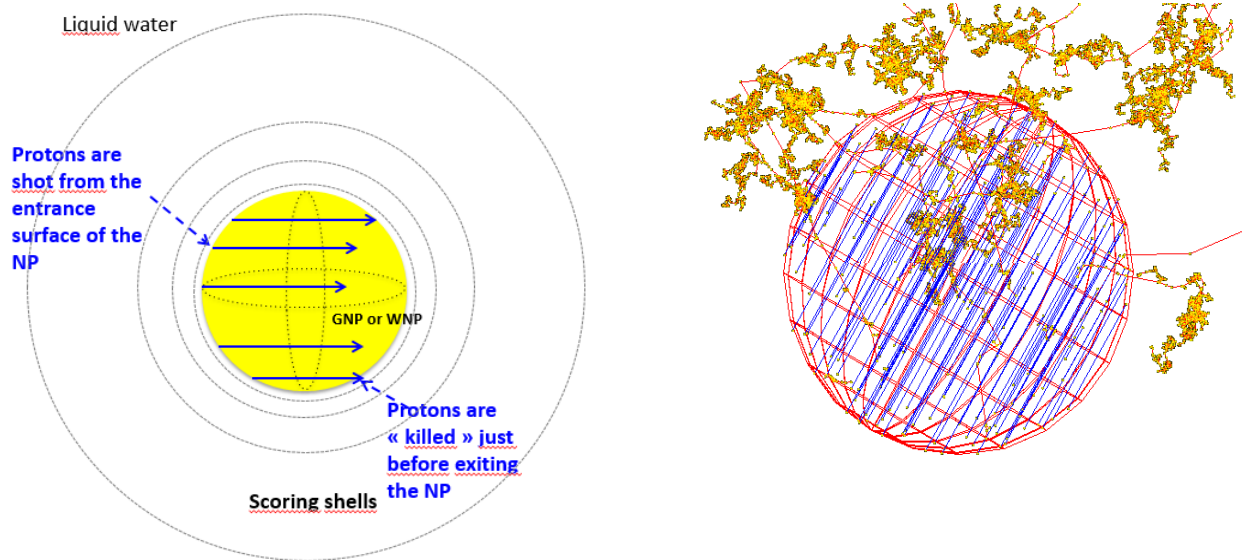
	G4BetheBlochModel	> 2 MeV
Photon		
Photoelectric effect	G4PhotonElectricEffect	G4LivermorePhotoElectricModel
Compton scattering	G4ComptonScattering	G4LivermoreComptonModel
Rayleigh scattering	G4RayleighScattering	G4LivermoreRayleighModel
Gamma conversion	G4GammaConversion	G4LivermoreGammaConversionModel

168
169 *Table 1: Geant4 physics **processes and models** used for the simulation of electron, proton and*
170 *photon interactions in the GNP and WNP. Process and model classes are indicated, as well as*
171 *the energy range of applicability of each model.*

172 *2.2 Irradiation conditions*

173 **The simulation geometry was comprised of a single NP immersed in liquid water irradiated**
174 **with mono-energetic proton beam.** We chose a NP with a diameter of 50 nm as suggested by
175 Chithrani *et al.* [12] and Lin *et al.* [21]. The transverse size of the proton beam was limited
176 within a circle of a 50 nm diameter in front of the surface of the NP. **The propagation axis of**
177 **the proton beam was set parallel to the z-axis (see Figure 1) and nine different energies were**
178 **simulated: 2, 3, 5, 10, 30, 50, 75, 100, and 170 MeV.** In order to score energy depositions, we
179 divided the region around the NP in 170 spherical shells of log-scale thickness, from the NP's

180 surface to 10^7 nm. All energy depositions are recorded in these spherical shells. The incident
181 protons are shot from the surface of the NP as a parallel beam and they are stopped when the
182 protons exit the NP in order to avoid the influence of energy deposition by protons in the
183 scoring shells, as shown in Figure 1.



184 *Figure 1: Left: schematic diagram of the simulated geometry showing the incident parallel*
185 *protons (full line) inside the NP (yellow sphere) and the scoring concentric spherical shells.*
186 *Right: example of visualization obtained with Geant4 when the NP (in red) is irradiated with a*
187 *parallel proton beam (blue tracks), showing emitted secondary electron interactions (red*
188 *tracks and yellow vertices).*

189 2.3 Analysis

190 The number of selected incident protons were fixed to 10^7 projectiles for GNP and to 10^8
191 projectiles for WNP to ensure good statistics and reasonable computing times for the radiolysis
192 simulation. Deposited energy was accumulated for all secondary interacting particles.
193 Normalized radial absorbed dose distribution per proton was obtained via the division of the
194 total energy deposited in each spherical shell by the mass of the corresponding shell and
195 number of incident protons. To evaluate the effectiveness of GNP, we estimated the absorbed
196 dose enhancement factor (DEF) by secondary particles outside the NP via the calculation of the

197 ratio of the dose absorbed outside the GNP with respect to dose absorbed outside the WNP
198 [21]. Only the NP particle material – gold or liquid water – was changed between these two
199 simulations (geometry and physics were not modified).

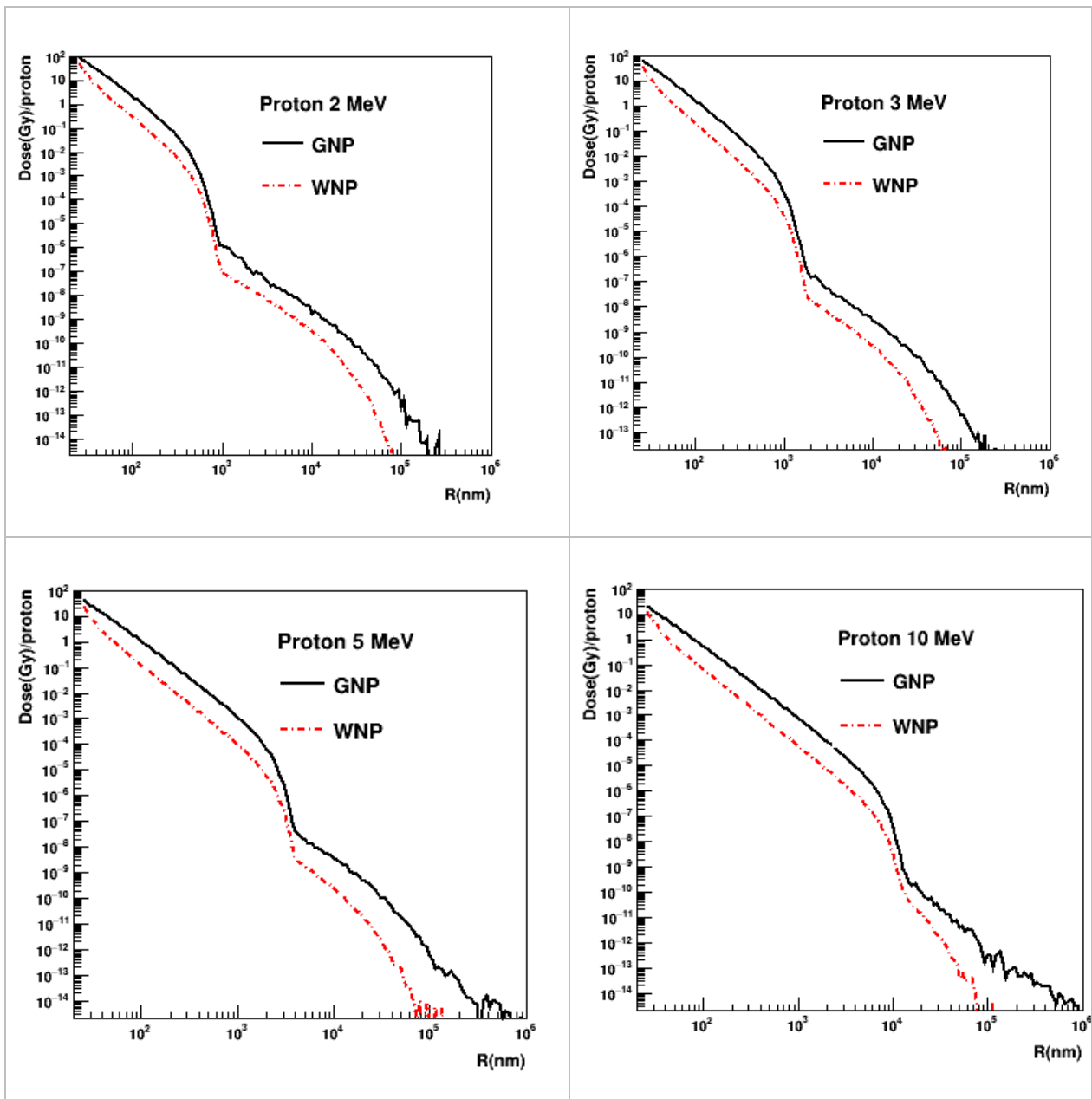
200 **3. Results and discussion**

201 *3.1 Absorbed dose enhancement*

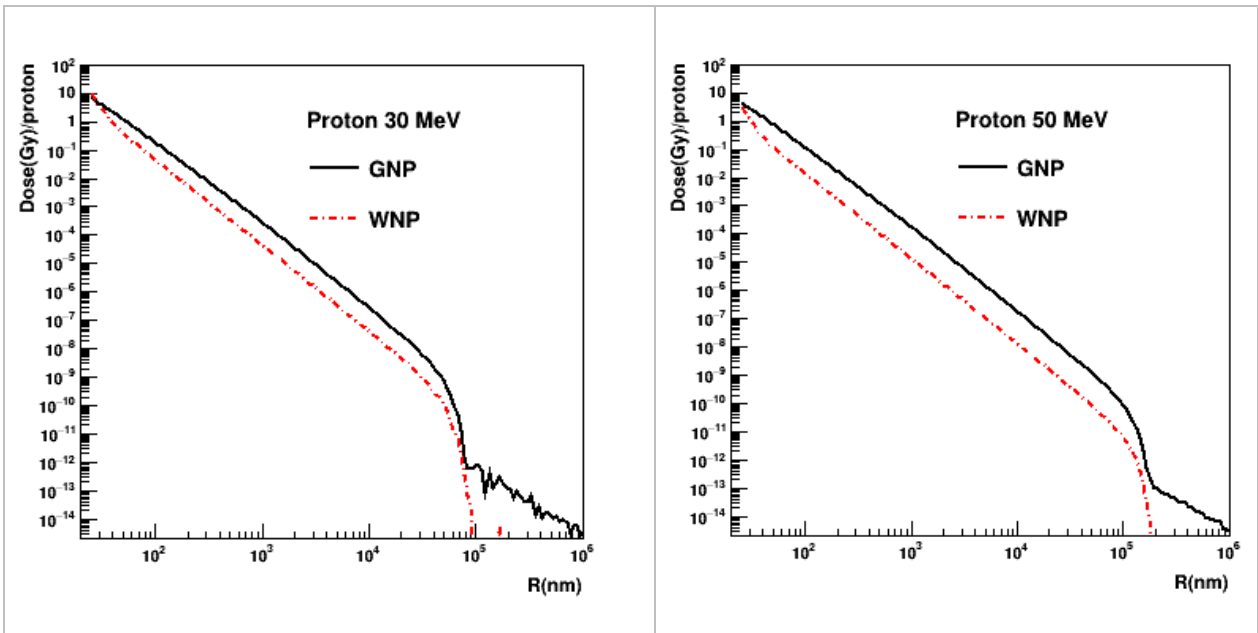
202 **Figures 2, 3 and 4** show the comparison of simulated radial absorbed dose distributions in
203 liquid water outside the NP, in the case of a GNP or a WNP, as a function of radial distance
204 from the NP. Results are presented for **2, 3, 5, 10, 30, 50, 75, 100 and 170 MeV** incident
205 protons. The energy deposited in liquid water outside the NP is primarily **due to the** secondary
206 electrons that are generated within and then exit the NP. In all cases, the absorbed dose rapidly
207 decreases as a function of radial distance. In addition, the dose absorbed in liquid water due to
208 secondary electrons emitted from the GNP is always larger than the absorbed dose due to
209 electrons emitted by the WNP, for a given incident proton energy. **This is expected as protons**
210 **interacting with gold will generate significantly more secondary electrons through ionization in**
211 **the NP than in the water case. Therefore, since the number of secondary electrons are more**
212 **numerous per interaction and gold is denser than water, the secondary electrons that exit the**
213 **GNP will deposit more energy outside the NP.**

214

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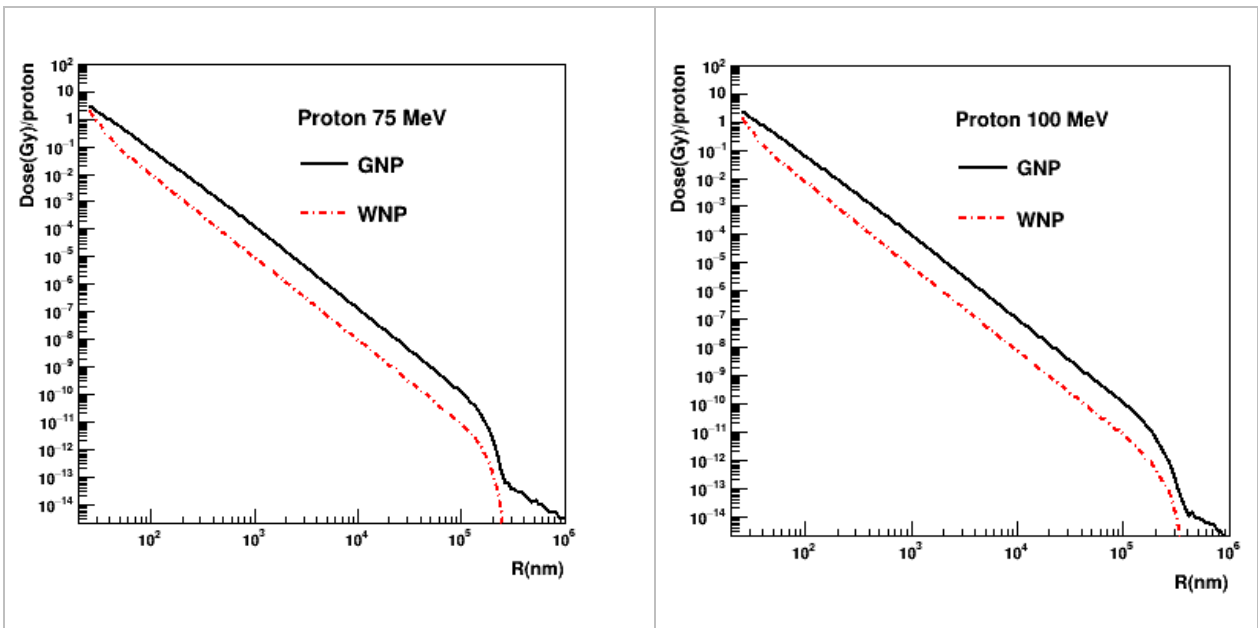


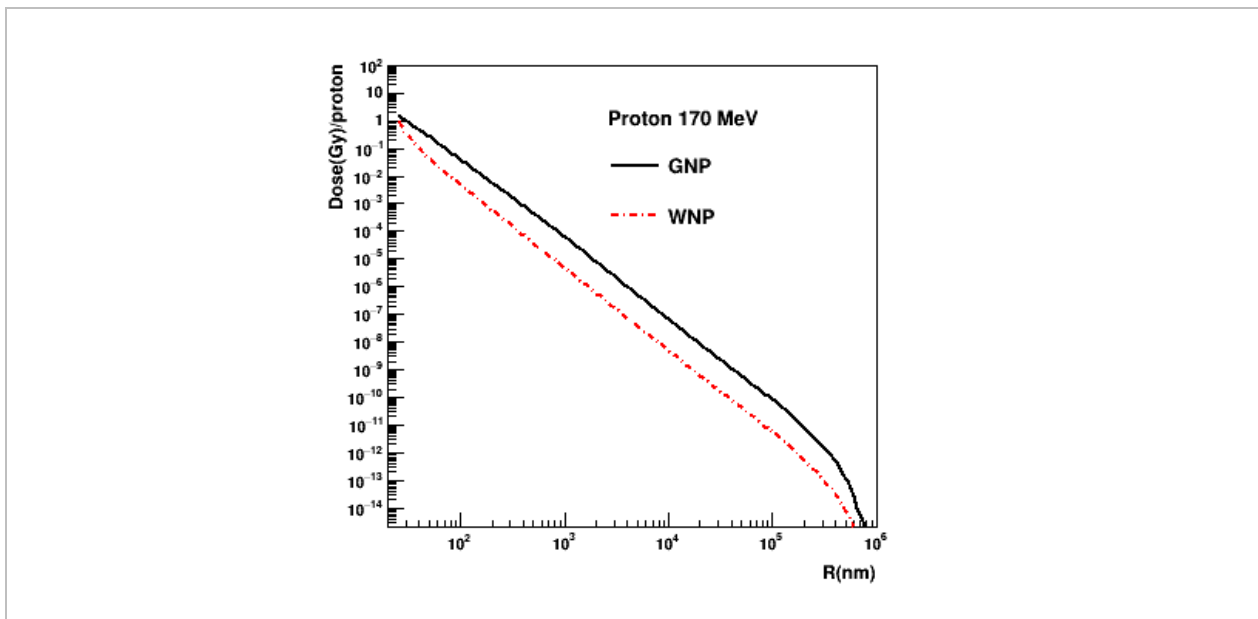
216 *Figure 2: Comparison of radial absorbed dose distribution for GNP (full black lines) and*
 217 *WNP (red dot-dashed lines) for 2 MeV, 3 MeV, 5 MeV and 10 MeV incident protons.*



218 *Figure 3: Comparison of radial absorbed dose distribution for GNP (full black lines) and*
 219 *WNP (red dot-dashed lines) for 30 MeV and 50 MeV incident protons.*

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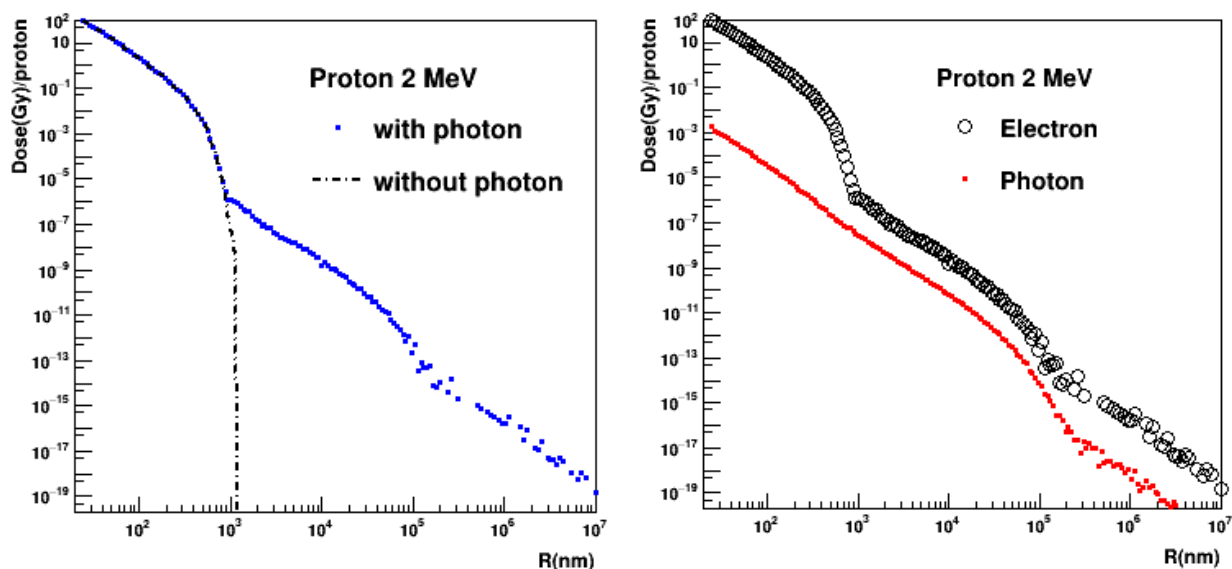


221 *Figure 4: Comparison of radial absorbed dose distribution for GNP (full black lines) and*
 222 *WNP (red dot-dashed lines) for 75 MeV, 100 MeV and 170 MeV incident protons.*

223
 224 The influence of secondary photons on the radial absorbed dose distribution can be observed in
 225 the case of 2 MeV incident protons at large distances ($r > 1 \mu\text{m}$) from the NP surface for both
 226 of GNP and WNP (and at larger distances for the three other incident energies as seen in Figure
 227 2). The maximum energy transferred by a 2 MeV proton to a free and at-rest electron is
 228 approximately 4.4 keV. This electron has a range in water of about $1 \mu\text{m}$, according to our
 229 estimations, which is consistent with the results shown in Figure 2.

230 Thus the observed gradient continuation beyond $1 \mu\text{m}$ in Figure 2 can be attributed to the
 231 interaction of secondary photons with the surrounding liquid water medium. These photons are
 232 generated through bremsstrahlung and atomic de-excitation in the NP. They can exit the NP
 233 and interact by photoelectric effect or Compton scattering in the water medium generating
 234 photoelectrons at large distances, which cannot be reached by secondary electrons produced by
 235 proton impact. The left panel in Figure 5 shows an example of this effect for 2 MeV incident
 236 protons through the comparison of radial absorbed dose surrounding the GNP with (blue
 237 points) and without (full black dotted line) the inclusion of energy depositions due to secondary

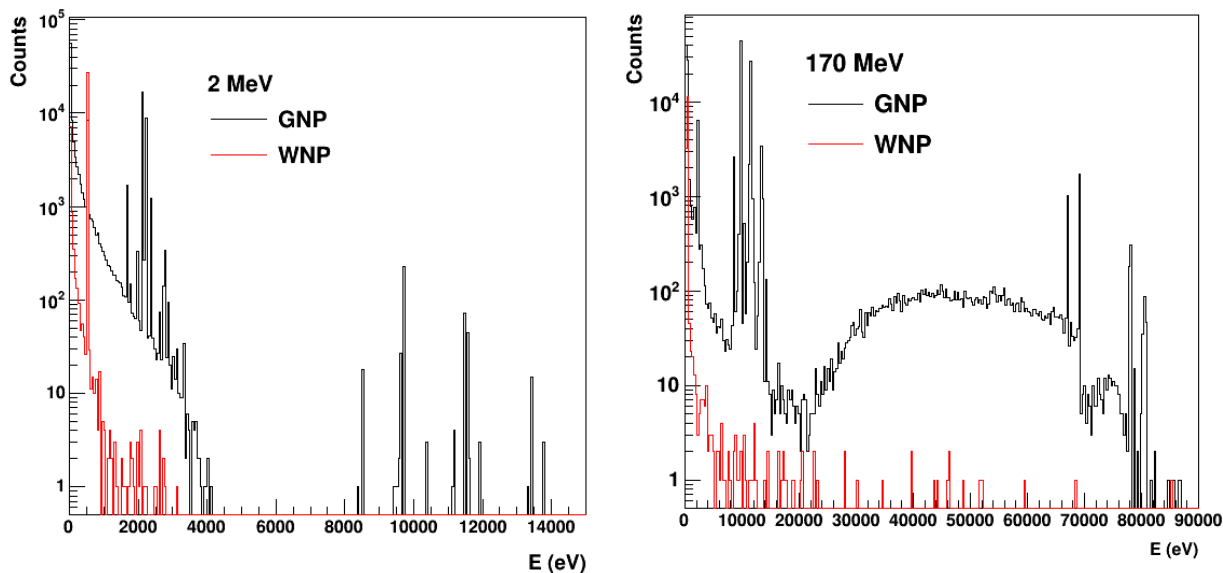
238 photons. This figure clearly illustrates the contribution of secondary photons to radial absorbed
 239 dose. Further evidence of their contribution can be observed in the radial absorbed dose
 240 distributions for electrons and photons at 2 MeV shown in the right panel of Figure 5. Here it
 241 can be seen that the contribution from photons to the total absorbed radial dose is negligible in
 242 comparison to electrons at distances below 1 μm .



243
 244 *Figure 5: (Left): Comparison of radial absorbed dose distribution around a GNP irradiated by*
 245 *2 MeV protons, when photoelectron production is deactivated (dot-dash black line) and when*
 246 *it is activated (dot blue line). (Right): Comparison of electron (open circles) and photon (red*
 247 *square) contributions to the radial absorbed dose.*

248 Figure 6 shows the comparison of the energy **spectra** of secondary photons for the GNP and
 249 WNP irradiated by 10^8 protons, for the 2 MeV and 170 MeV cases. For this specific
 250 configuration, which does not require the simulation of the chemical stage but only physical
 251 interactions, we used identical statistics for GNP and WNP. The number of photons generated
 252 by the GNP is larger than the number of photons created by the WNP. In the case of the GNP,
 253 the dominant deexcitation lines are observed: $M_{\alpha 1}$ (2 keV), $L_{\alpha 1}$, $L_{\alpha 2}$, $L_{\beta 1}$, $L_{\beta 2}$ and L_{γ} (from 9
 254 keV to 13 keV) for 2 MeV protons. For 170 MeV protons, dominant $K_{\alpha 1}$, $K_{\alpha 2}$ and $K_{\beta 1}$ lines are

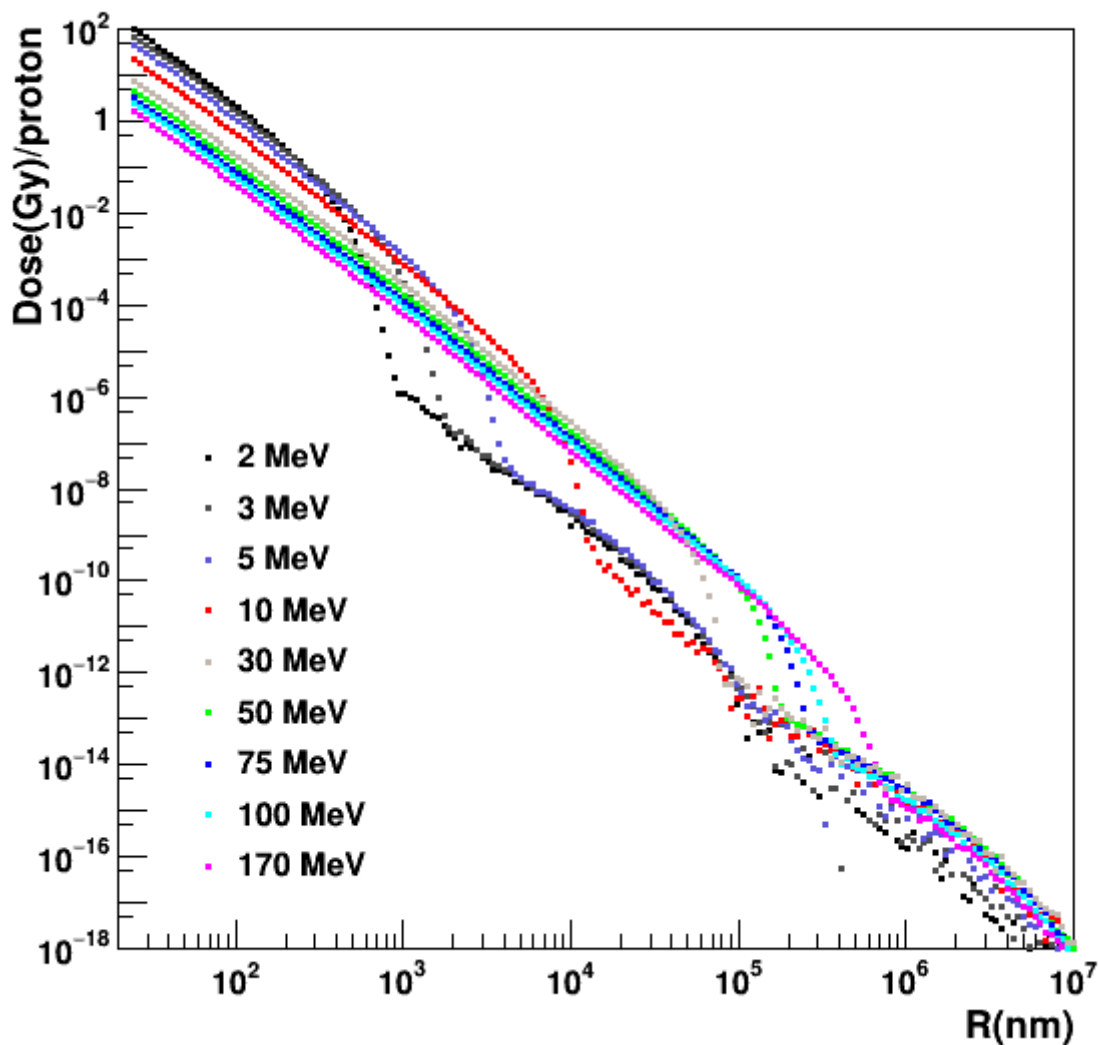
255 observed (68-78 keV, see details on de-excitation lines in [32]). K line of Oxygen at 524 eV is
 256 also observed for the WNP. The bremsstrahlung background is also observed in both cases.
 257



258
 259 *Figure 6: Comparison of the energy spectrum of secondary photons for the GNP (black line)*
 260 *and the WNP (red line) irradiated by 10^8 protons, for the 2 MeV (left) and 170 MeV (right)*
 261 *cases. .*

262 Figure 7 shows the comparison of absorbed dose as a function of radial distance from the GNP
 263 for nine incident proton energies (2, 3, 5, 10, 30, 50, 75, 100 and 170 MeV). Up to
 264 approximately 1 cm from the GNP all curves have the same monotonic dependence with radial
 265 dose; i.e. the larger the kinetic energy of the protons, the smaller the absorbed radial doses.
 266 This is also expected since for the selected incident energy range (2 MeV – 170 MeV) the
 267 linear energy transfer of protons in the GNP decreases with increasing kinetic energy, leading
 268 to a decrease of the production of secondary electrons in the GNP and thus a decreasing
 269 absorbed dose in liquid water outside the NP at a given radial distance. The range of secondary
 270 electrons produced by the projectile impact increases as the incident proton kinetic energy
 271 increases; this is explained by the fact that the maximum energy transferred from a proton to a

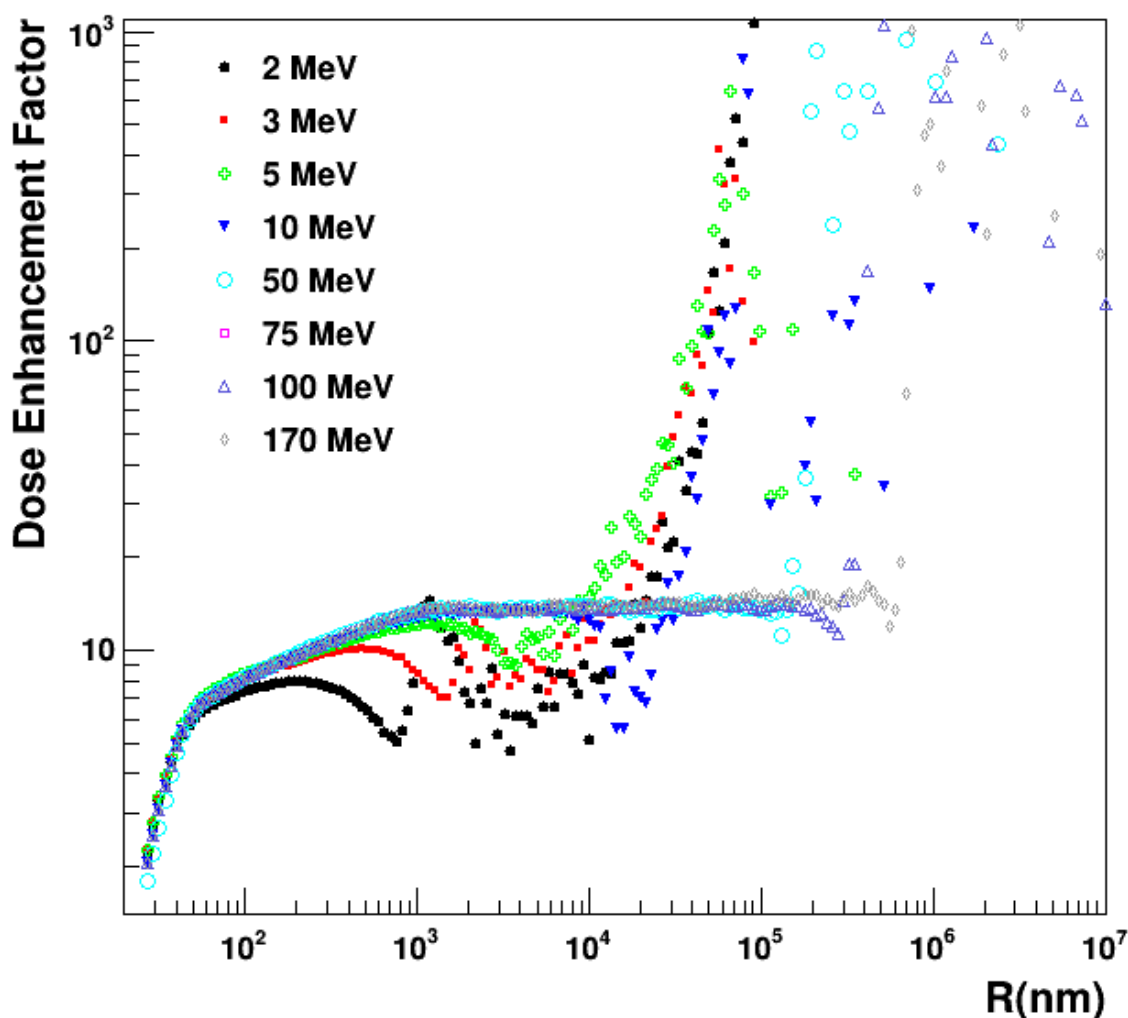
272 free and at-rest electron (as those in valence shells or quasi-free electrons in metals) increases
273 linearly with the proton energy. These electrons determine the gradient-discontinuity region
274 observed in the radial dose curve.
275



276
277 *Figure 7: Absorbed radial dose distribution for GNP for 2, 3, 5, 10, 30, 50, 75, 100 and 170*
278 *MeV protons.*
279

280 The contribution of photoelectrons to the absorbed radial dose distribution is observed for
281 incident proton energies (2 MeV-170 MeV) in Figure 7 for distances larger than that defined
282 by the gradient-discontinuity region (see also Figure 5).

283 Finally, Figure 8 presents the absorbed dose enhancement factor (DEF) between GNP and
284 WNP obtained for proton beams of 2, 3, 5, 10, 50, 75, 100 and 170 MeV as a function of radial
285 distance from the NP. The DEF is always larger than 1, demonstrating the absorbed dose
286 increase induced by the GNP. For **higher** incident energies (> 10 MeV), the DEF clearly
287 increases with radial distance up to **the micron-scale** and becomes nearly constant beyond. At
288 larger distances (millimeter-scale), we observe a fast enhancement which can reach a DEF of a
289 few hundred for low energy values, such as 2, 3, 5 and 10 MeV.



290

291 *Figure 8: Absorbed Dose Enhancement Factor (DEF) obtained by mono-energetic proton of 2,*
292 *3, 5, 10, 50, 75, 100 and 170 MeV.*

293 The behavior of the DEF for protons with incident kinetic energies of a few MeV (2, 3 and 5
294 MeV) for radial distances of less than 10 μm from the NP is complex. These profiles, within
295 this confined region, undergo an initial increase, followed by a decrease, and finally an
296 increase up to large values beyond 10 μm . For example, in the 2 MeV case, the DEF first
297 increases up to 7.5 at a distance of about 200 nm, then decreases when approaching 1 μm and
298 increases again for larger radial distances. The stages of this behavior can be explained as
299 follows:

- 300 • For small radial distances, the DEF increases with radial distance, up to a distance
301 (about 200 nm from the NP) corresponding to a kinetic energy equal to the mean
302 ionization potential of gold (790 eV). This increase is due to a large number of
303 secondary ionization electrons created in the GNP by delta electrons with kinetic
304 energies above 790 eV and the incident proton itself. In the case of water, this limit is
305 equal to 78 eV and the lower density of the material drastically reduces the number of
306 generated secondary electrons. Thus, the observed increase must be considered as an
307 approximation only as the Geant4 ionization models for incident protons (the
308 G4BraggModel and the G4BetheBlochModel) have a low energy applicability limit
309 equal to the mean ionization potential of the medium and they can not generate delta
310 electrons below these values (note that this is not the case for Livermore models which
311 are usable for electrons down to a few 10 eVs). In order to further investigate this DEF
312 increase at small radial distances, we plan in the near future to include discrete inelastic
313 models applicable to gold and other materials into Geant4-DNA.
- 314 • At larger radial distances, the DEF decreases. This decrease corresponds to the fall-off
315 observed in Figure 2 and Figure 5 in the GNP and WNP: at a given incident proton
316 energy, most energetic electrons generated by the GNP and the WNP behave similarly.

317 Their ranges are similar and absorbed doses for both the GNP and the WNP decrease
318 rapidly at the end of their range. Figure 8 shows that the DEF decreases near the
319 transition region defined by these ranges.

320 • Finally, beyond the range of most energetic secondary electrons (the fall-off),
321 photoelectron interactions dominate at very large distances. Due to the larger number of
322 secondary electrons generated by the GNP than by the WNP, more photoelectrons are
323 generated by the GNP than by the WNP and this leads to the observed large increase of
324 the DEF. It should be noted that high Z targets preferentially decay by fluorescence
325 emission after inner shell ionizations (K, L, M shells), unlike low Z targets that decay
326 mainly through Auger emission. A high-energy fluorescence X-ray produced in gold
327 may remove electrons in water by Compton scattering or photo-electric effect. In
328 addition, an enhancement of Bremsstrahlung production can be observed (see Figure 6)
329 for gold when compared to that of water. These two facts lead to an increase of the
330 secondary photon production and thus of the secondary electrons in water out of the
331 GNP. Note that the peak between 1 and 2 μm in Figure 8 is caused by the transition
332 between secondary electrons and photoelectrons. The observed decrease in DEF just
333 after the peak can be attributed to the fluorescence K line photons of Oxygen in the
334 WNP. Thus, the large increase of the DEF beyond 10 μm comes from the absence of
335 high energy fluorescence photons generated by the WNP compared to the GNP case.

336 3.2. Radiolysis enhancement

337 Thanks to the recent addition of the modeling capabilities of physico-chemical and chemical
338 processes in Geant4-DNA [23], it is now possible to simulate the production of radiolytic
339 species during the so-called “physico-chemical” and “chemical” stages of water radiolysis, up
340 to 1 μs after irradiation. After the so-called “physical” stage, excited and ionized water
341 molecules may dissociate into new chemical species (such as e^-_{aq} , H_2 , $\text{H}\cdot$, $\cdot\text{OH}$, H_3O^+), which
342 can diffuse and interact mutually producing other species (such as OH , H_2O_2). The

343 dissociation scheme of excited and ionized water molecules, diffusion coefficients of diffusing
344 species, and list of chemical reactions and reaction rates are all taken from our previous
345 publication[30]. We performed the simulation of chemical stage up to 1 μ s in order to
346 investigate the effect of the GNP on the production of chemical species, as a function of radial
347 dose from the GNP, time and incident proton energy. For consistency, the chemical species
348 were scored using the same spherical shell approach employed in Section 3.1 (see Section 2.2
349 for more details).

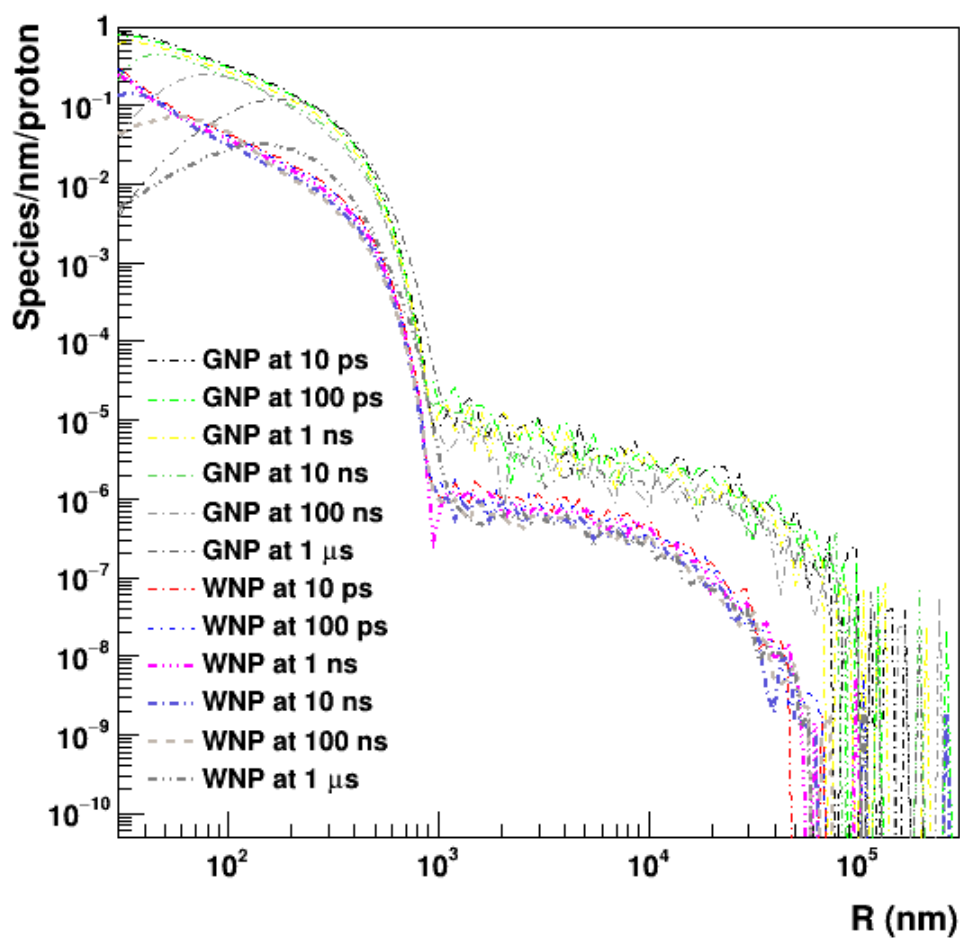
350 - Time-dependence of chemical species production

351 Figure 9 and Figure 10 show the time evolution of the distribution of all chemical species
352 produced by 2 MeV and 50 MeV incident protons respectively, as a function of radial distance,
353 from either the GNP or the WNP. Results are shown at six different times after irradiation: 10
354 ps, 100 ps, 1 ns, 10 ns, 100 ns and 1 μ s. At very short times (10 ps, 100 ps and 1 ns), radial
355 distributions on the total number of chemical species are very similar, with the same order of
356 magnitude, and do not show any significant time dependence. From 10 ns to 1 μ s, radial
357 distributions decrease faster with time in the closest vicinity of the NP. Indeed, in this dense
358 area, the radiolytic species recombine to water molecules.

359 Figure 11 and Figure 12 show the Radiolysis Enhancement Factor (REF) for 2 MeV protons
360 and 50 MeV protons respectively, calculated in a given shell, as the ratio of the number of
361 chemical species generated in the presence of the GNP, to the number of chemical species
362 generated in the presence of the WNP. At 10 ps, 100 ps and 1 ns, REFs increase and have the
363 same dependence on radial distance. For the 2 MeV case, the enhancement increases up to 7.5
364 at a distance of about 200 nm, then decreases until 1000 nm. For the 50 MeV case, REFs are
365 larger than 10 beyond 1 μ m away from the NP, illustrating again the strong influence of the
366 GNP on the production of radiolytic species compared to the water case. For both cases, from
367 10 ns to 1 μ s, REFs increase slower as a function of time from the surface of the NP.

368 The observed distributions and the time independence of the REF in Figures 11 and 12 during
369 the first few ns of the chemical phase are not unexpected.

370 These observed trends are due to the combination of three factors; 1) the increase in the
371 production of chemical species is full driven by the physical stage and, as such, more species
372 are generated in the case of the GNP than the WNP, 2) at short time scales, most of the
373 recombination will only happen around the deposited energy points where chemical kinetics
374 are not yet dependent on the structure of the secondary electrons , and 3) at times greater than a
375 few ns, the radiolytic species start to diffuse away from the deposited energy positions and then
376 become dependent on the dynamics of the track structures. The time evolution of the chemical
377 stage is experimentally investigated through the evolution of radiochemical yields defined as
378 the ratio of the number of a given radiolytic species to the deposited energy. The time
379 evolution of this ratio may depend on the incident radiation LET (see for eg. [23]), underlining
380 the strong influence of the physical stage on yields.

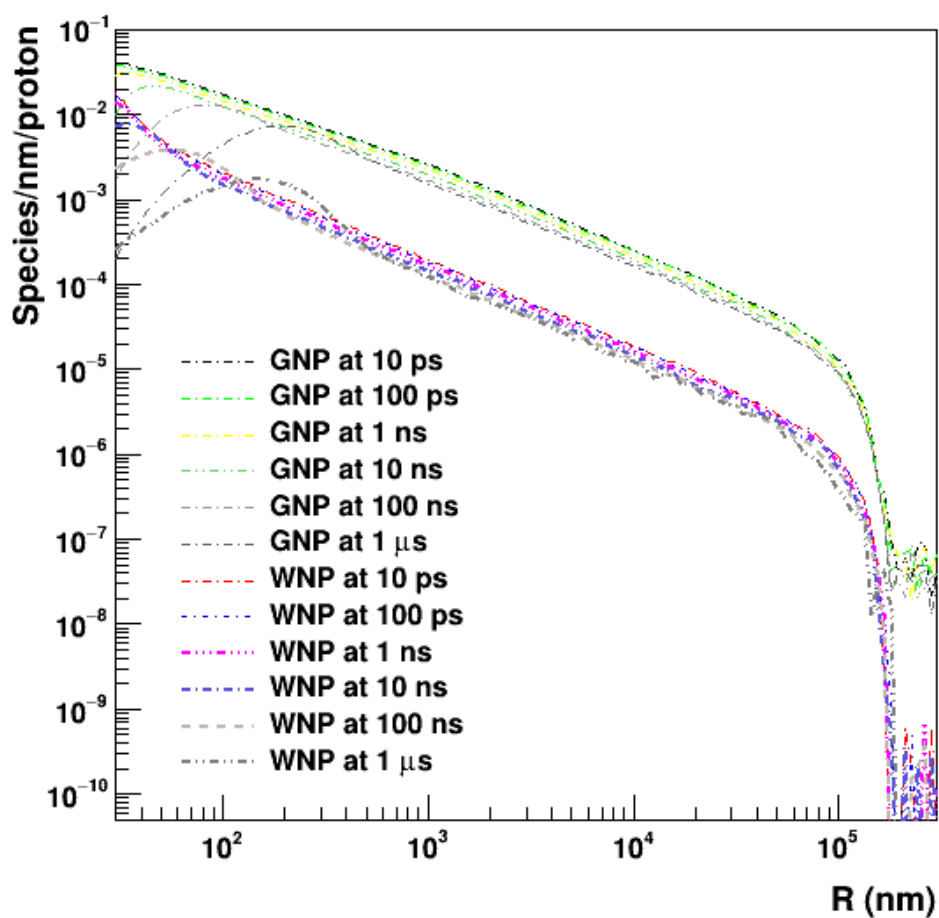


381

382 *Figure 9: Time evolution of the distribution of all radiolytic species produced by incident 2*
 383 *MeV protons as a function of radial distance, either for the GNP or WNP. Results are shown at*
 384 *six different times after irradiation (10 ps, 100 ps, 1 ns, 10 ns, 100 ns and 1 μs).*

385

386



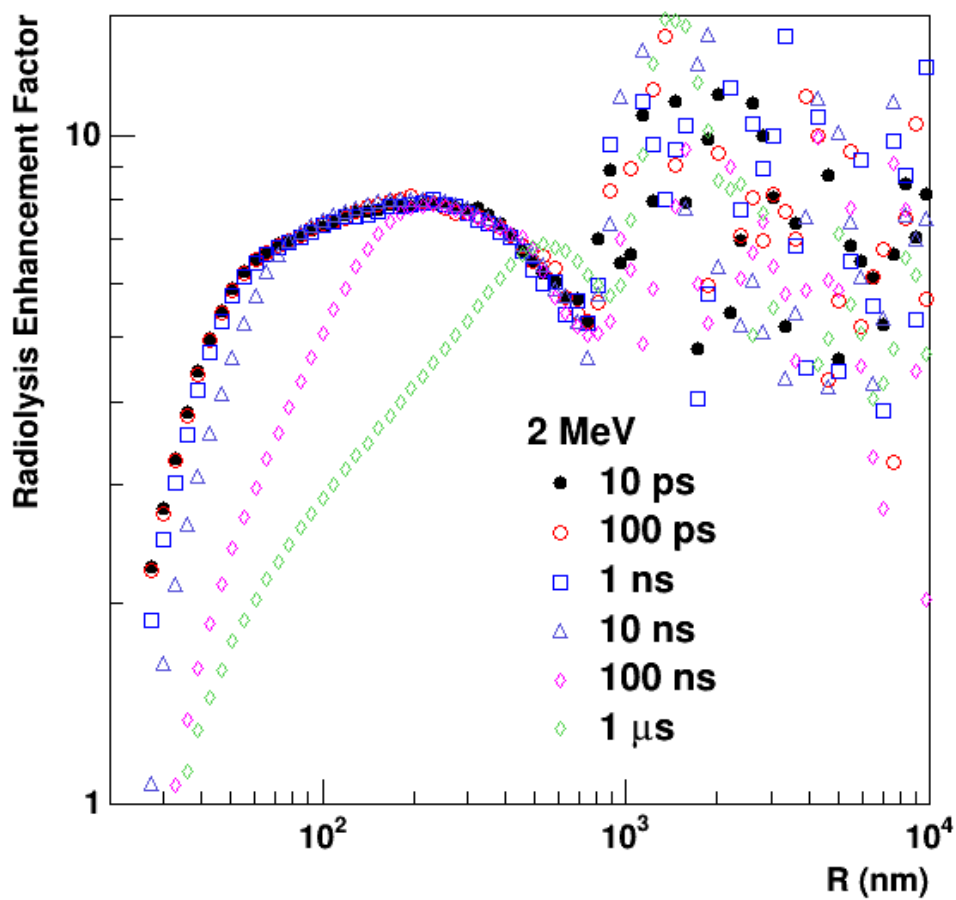
387

388 *Figure 10: Time evolution of the distribution of all chemical species produced by incident 50*
 389 *MeV protons as a function of radial distance, either for the GNP or WNP. Results are shown at*
 390 *six different times after irradiation (10 ps, 100 ps, 1 ns, 10 ns, 100 ns and 1 μ s).*

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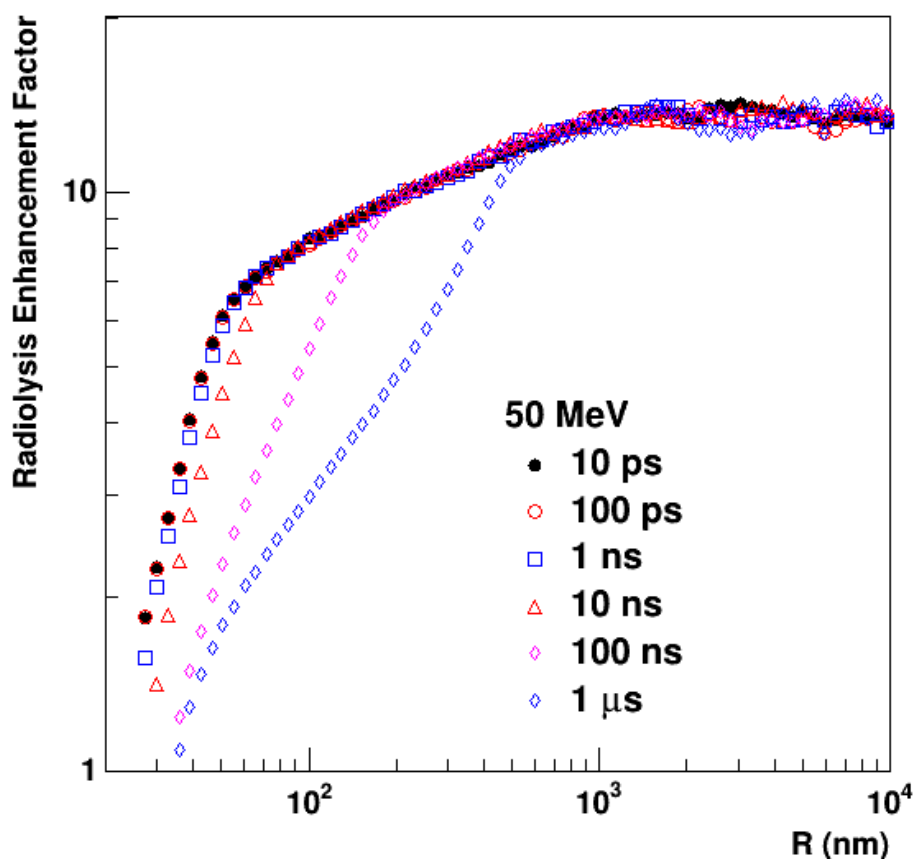
395 *Figure 11: Radial Radiolysis Enhancement Factor (REF) for 2 MeV protons for six time*
 396 *intervals after irradiation (10 ps, 100 ps, 1 ns, 10 ns, 100 ns and 1 us)*

397

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399

400



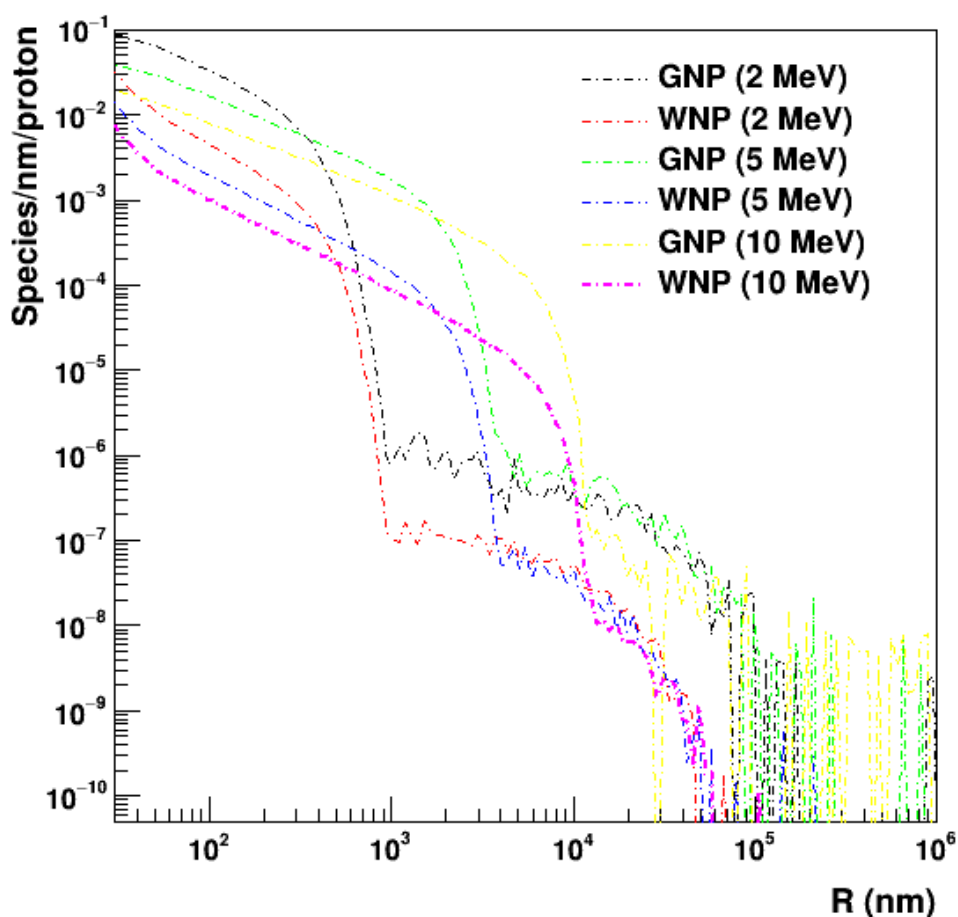
401

402 *Figure 12: Radial Radiolysis Enhancement Factor (REF) for 50 MeV protons for six time*
 403 *intervals after irradiation (10 ps, 100 ps, 1 ns, 10 ns, 100 ns and 1 us).*

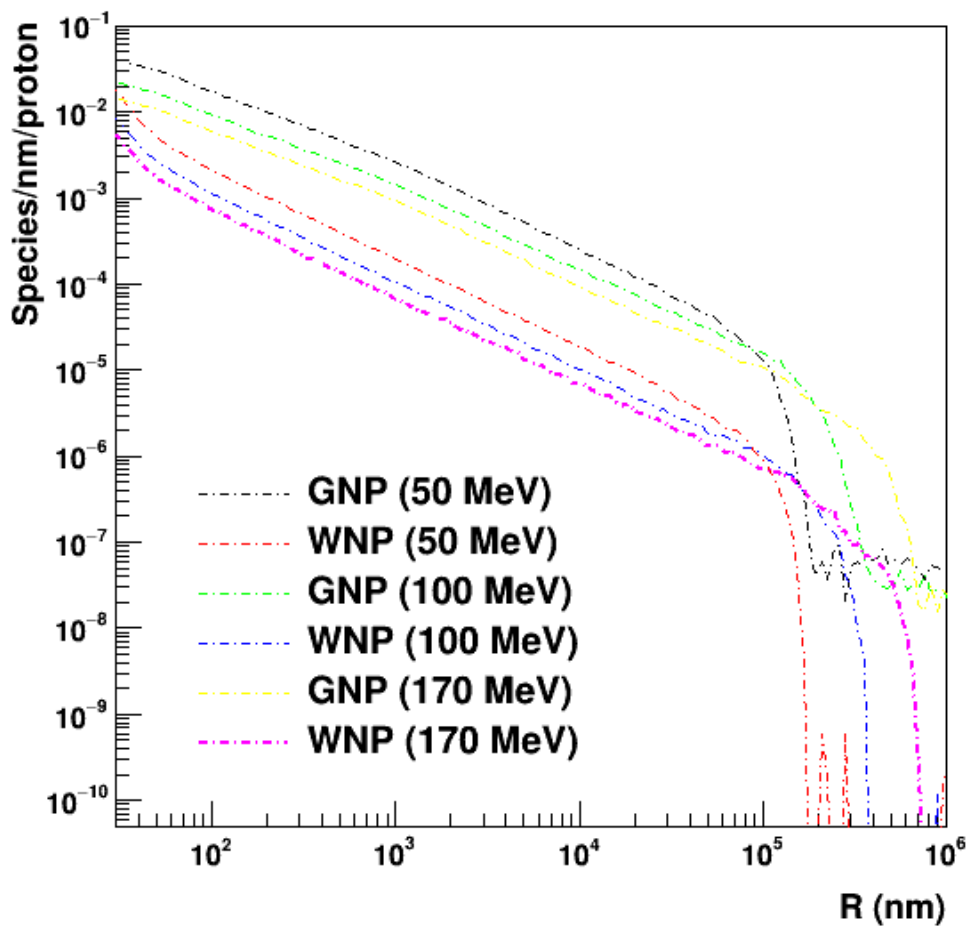
404 - Incident energy dependence of radiolytic species production

405 The energy dependence on the production of chemical species at 10 ps after irradiation is
 406 presented in Figure 13 and Figure 14. The distribution of species as a function of radial dose is
 407 presented for six incident proton energies: 2, 5, 10, 50, 100 and 170 MeV. For these six
 408 energies, species generated around the GNP are always in larger number than for the WNP
 409 case. This is related to the larger absorbed dose observed around the GNP compared to the
 410 WNP, as we previously discussed in Section 3.1. Further inspection of Figures 13 and 14 also
 411 shows that there is an inverse relationship between the radial distribution of radiolytic species
 412 and incident proton energy. This can be attributed to the radial absorbed dose dependence with
 413 LET on the incident proton energy (discussed in Section 3.1). Further evidence of this

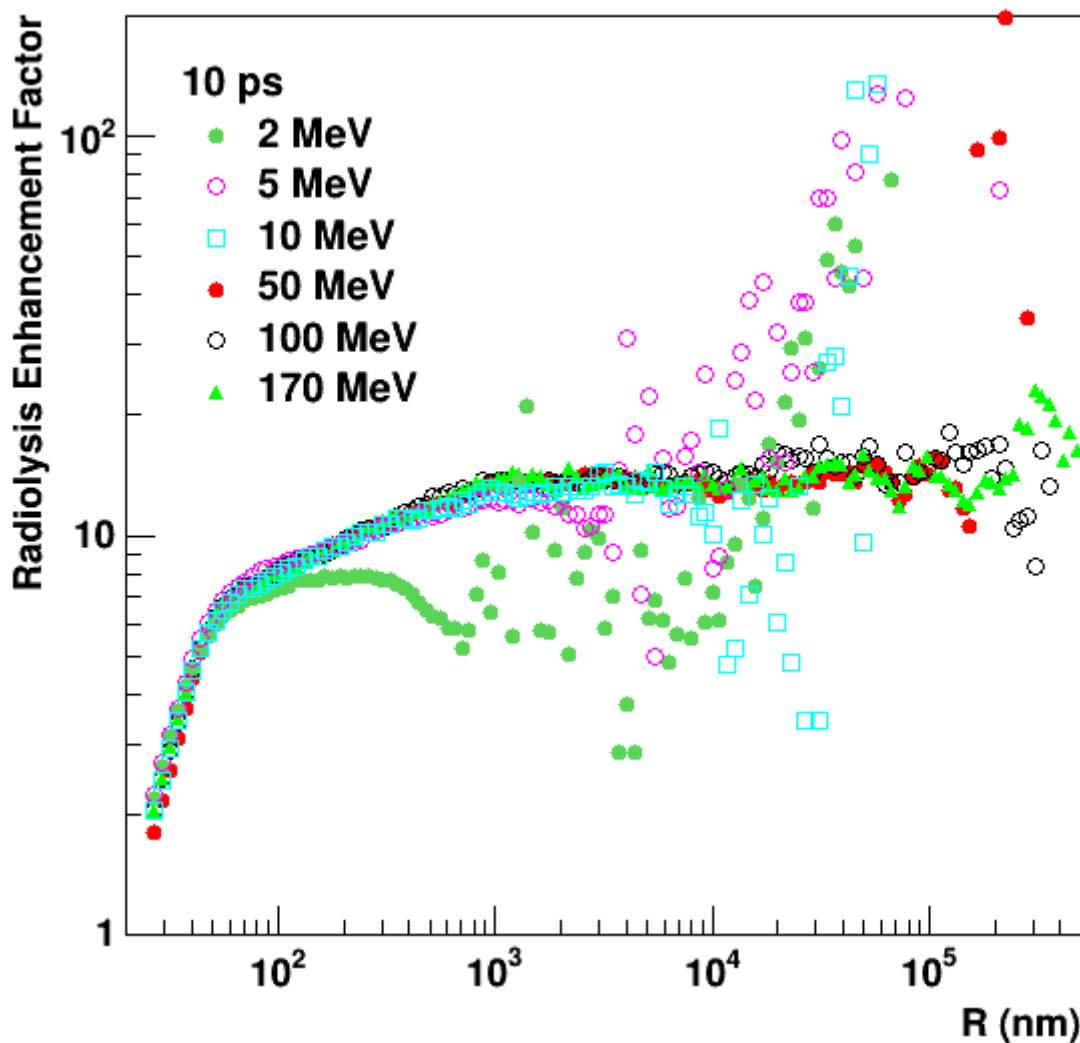
414 relationship between radiolytic species production and incident proton energy around the GNP
415 can be seen in Figure 15. This figure presents the REF radial profiles for each of the six tested
416 incident proton energies (2, 5, 10, 50, 100 and 170 MeV) at 10 ps after irradiation. It can be
417 seen that the maximum REF in each radial profile before the fall-off moves further out from
418 the edge of the GNP surface with increasing proton energy. However a non-trivial relationship
419 between the maximum achieved REF and incident proton energy is also present, e.g. at 2 MeV
420 a maximum REF of 7.5 can be seen at 2 MeV but it then plateaus at around 10 for the other
421 five investigated energies.



422
423 *Figure 13: Distribution of radiolytic species as a function of radial distance for three incident*
424 *proton energies (2, 5, 10 MeV) for the GNP and WNP, at 10 ps after irradiation.*



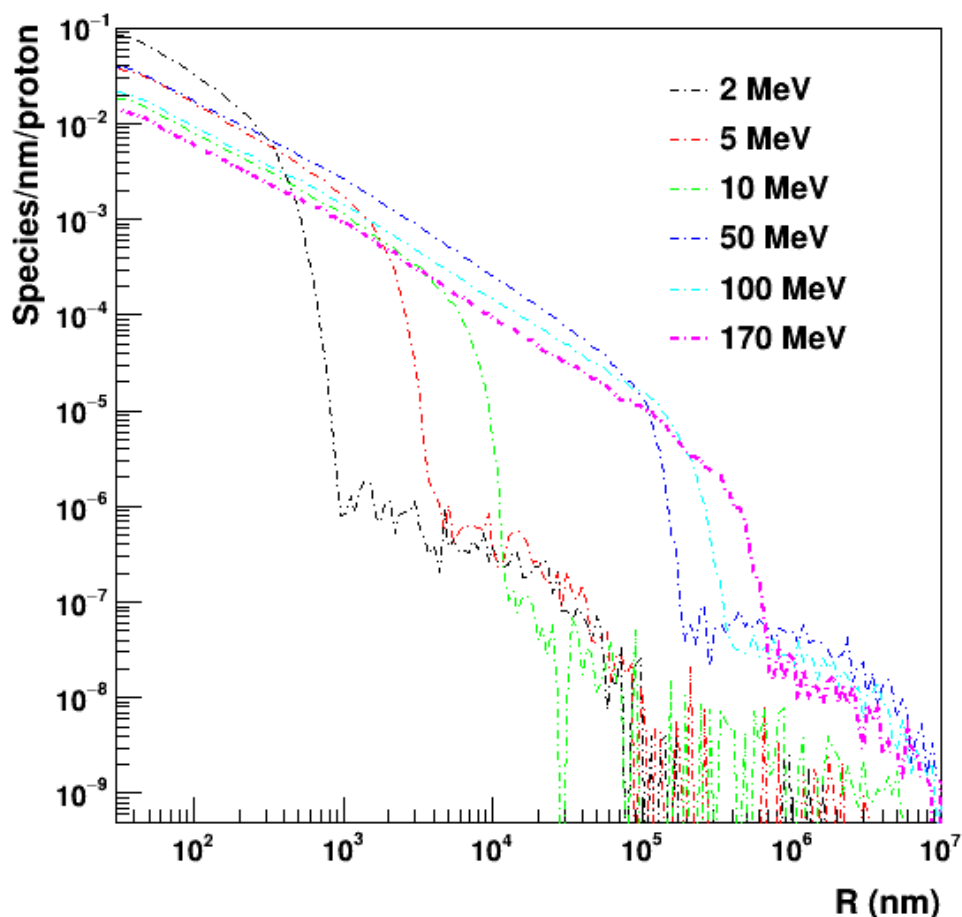
425
 426 *Figure 14: Distribution of radiolytic species as a function of radial distance for three incident*
 427 *proton energies (50, 100 and 170 MeV) for the GNP and WNP, at 10 ps after irradiation.*



428
 429 *Figure 15: REF dependence with radial distance at 10 ps for six incident proton energies (2, 5,*
 430 *10, 50, 100 and 170 MeV).*

431
 432 **Figure 16 shows the distribution of radiolytic species as a function of radial distance from the**
 433 **surface of the GNP for all six investigated proton energies (2, 5, 10, 50, 100 and 170 MeV).**
 434 **For radial distances of up to about 10 μm from the GNP, all distributions possess the same**
 435 **monotonic dependence on radius and the fall-off at larger distances. The distance at which the**
 436 **fall-off in REF occurs is directly proportional to incident proton energy. As the trends observed**
 437 **in Figure 16 match the absorbed dose profiles for the GNP shown in Figure 7, we can clearly**

438 illustrate that the generation of chemical species is directly dependent on absorbed dose, i.e.
439 any increase in absorbed dose outside the NP will result in an increase in chemical species
440 production. Thus we can directly link the generated chemical species at radial distances after
441 each profiles fall-off to the photons created during the interaction of the proton with the GNP.

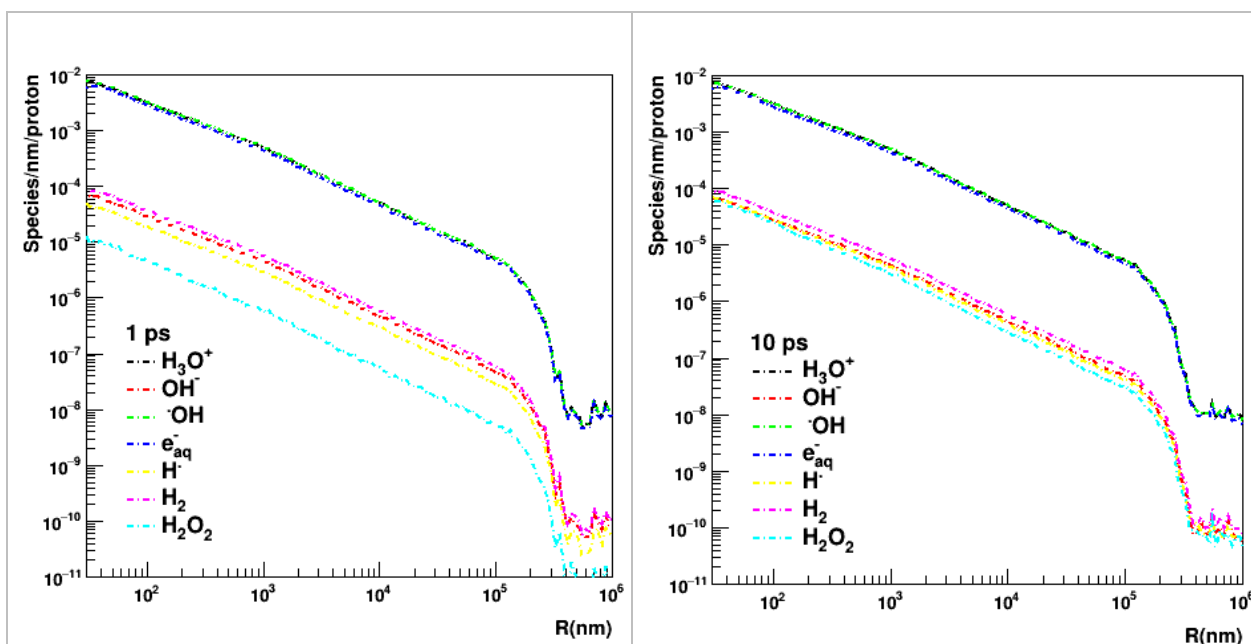


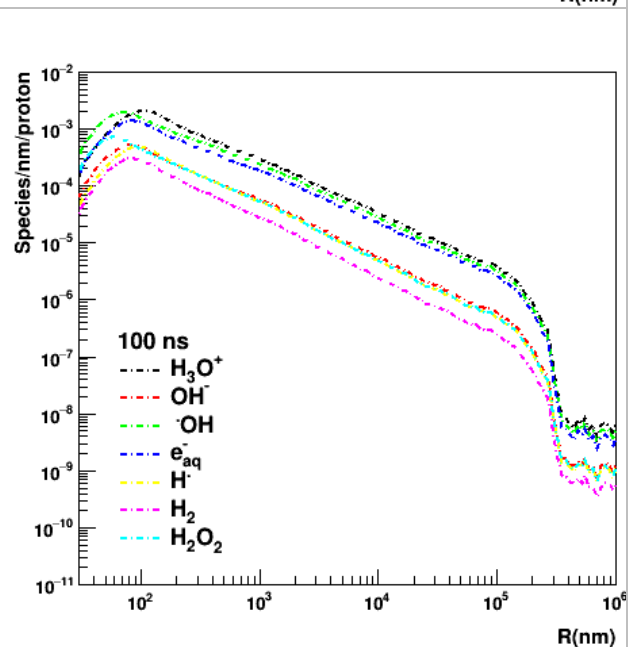
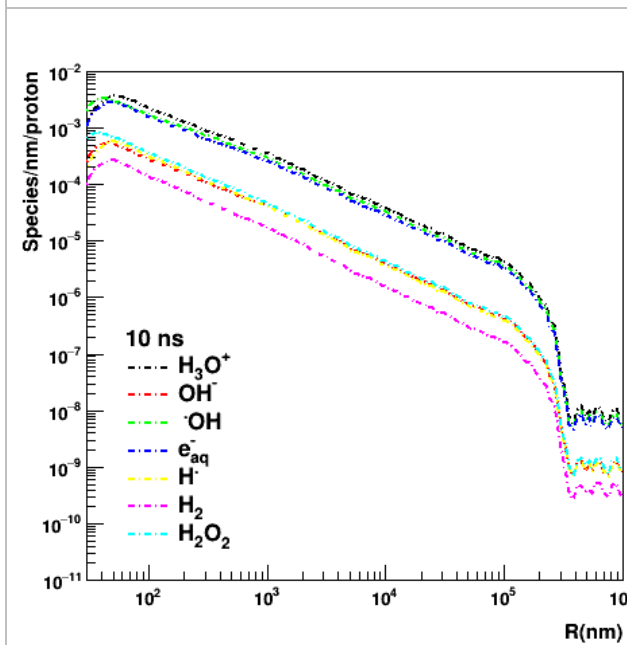
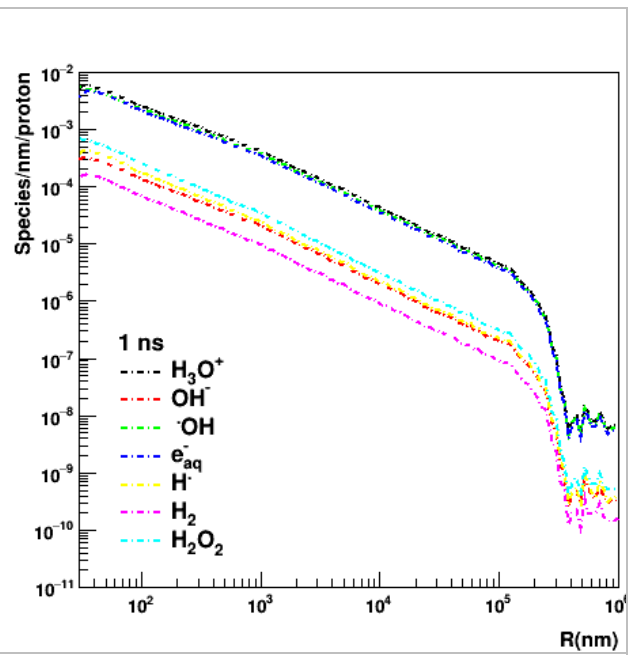
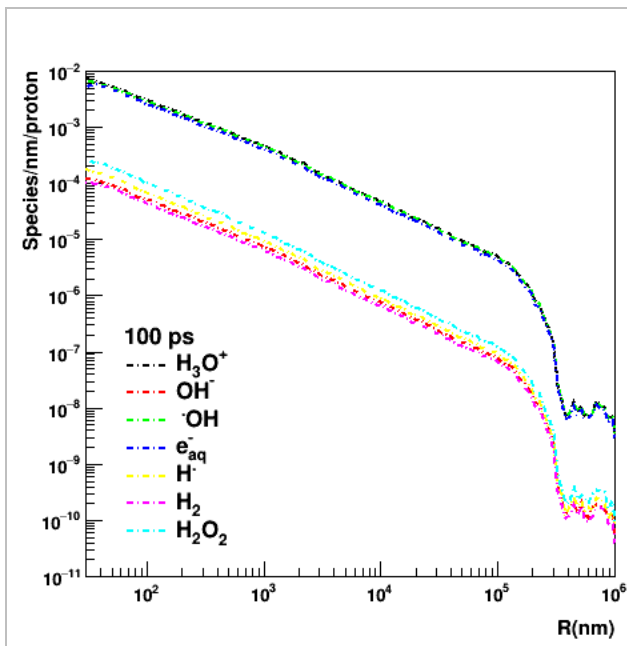
442
443 *Figure 16: Radial distribution of chemical species 10 ps after irradiation for all six*
444 *investigated proton energies (2, 5, 10, 50, 100 and 170 MeV) incident on the GNP.*

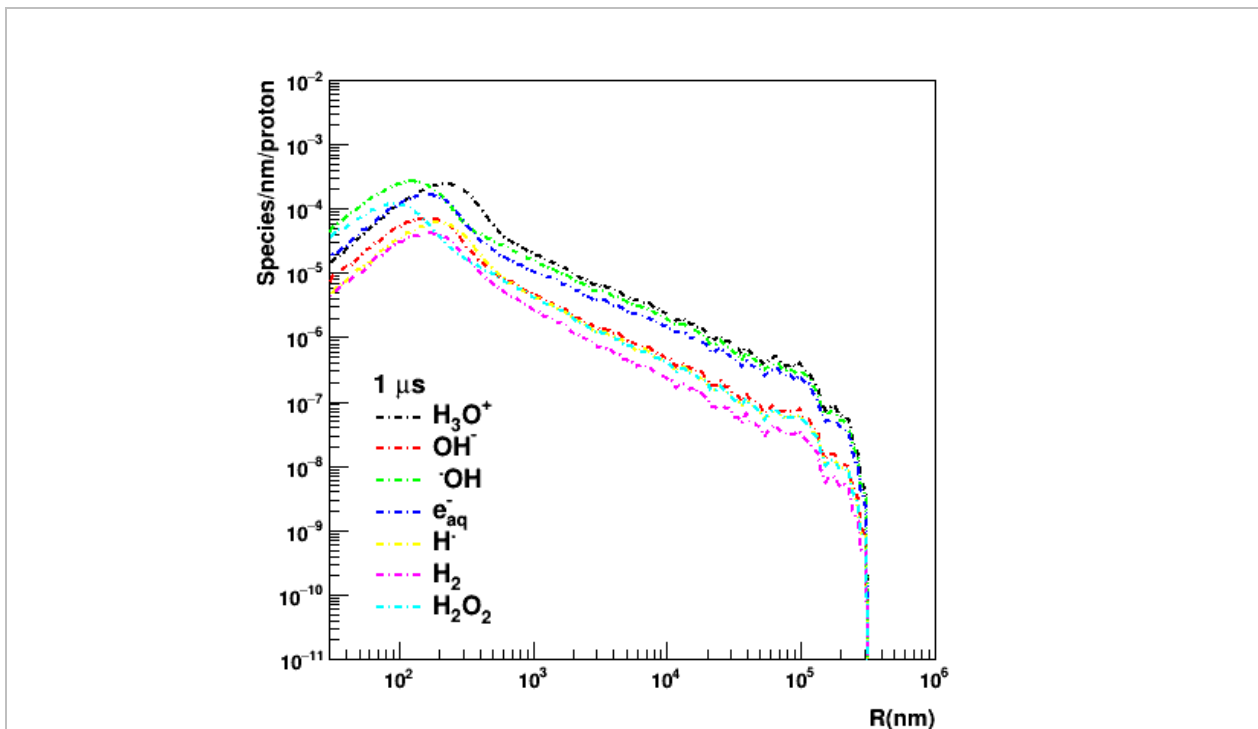
445
446
447 Figure 17 shows the relative distributions of chemical species (H_3O^+ , OH^- , $\bullet\text{OH}$, e_{aq}^- , $\text{H}\bullet$, H_2 ,
448 H_2O_2) produced by incident 100 MeV protons and their time evolution from 1 ps to 1 μs , as a
449 function of radial distance from the GNP. All distributions of chemical species have the same

450 dependence with radial distance. It can also be observed that the relative distributions evolve
451 differently with time: we observe an increase for OH^- , H_2O_2 and H^\bullet , a decrease for $\bullet\text{OH}$, H_3O^+
452 and e_{aq}^- , and a slow change for H_2 . This observation is in agreement with the previously
453 published time-dependent radiolytic yields in [23].

454 At short times after irradiation (1, 10, 100 ps and 1 ns), the distributions of radiolytic species
455 from the NP are similar with the absorbed dose distribution. But from 10 ns to 1 μs , the
456 deviation of the radiolytic species distribution compared to the absorbed dose distribution can
457 be observed more easily as the peak of radiolytic species distribution moves in time from the
458 close vicinity of the NP to about 200 nm (Figure 17). It should be recalled, in this case, that the
459 incident protons were stopped when they exit the NP.





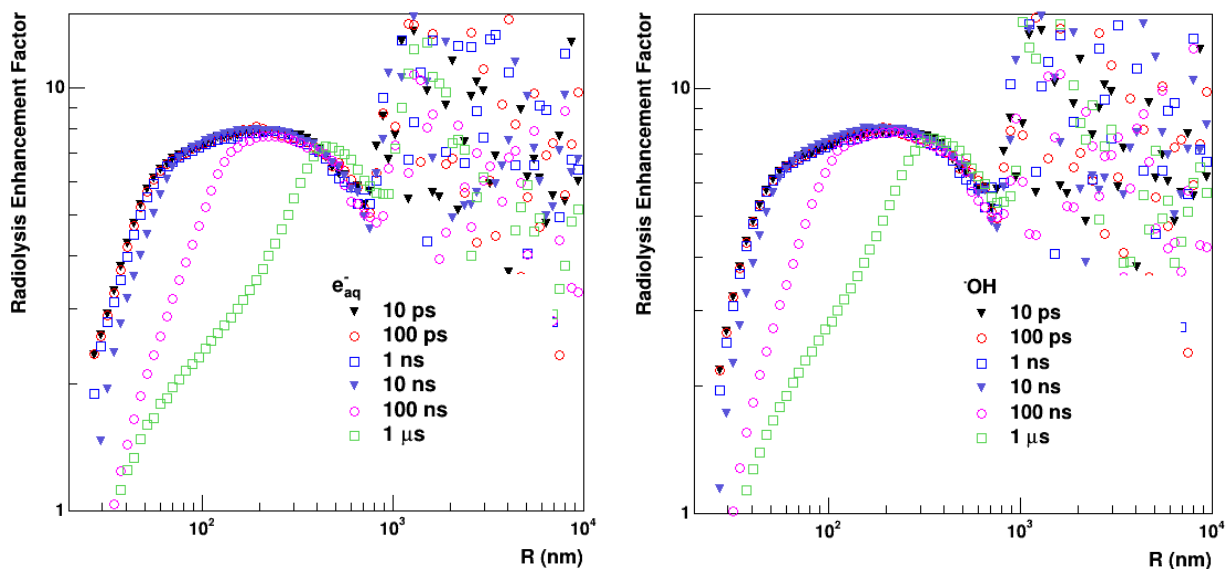


460
 461 *Figure 17: Time evolution of the distribution of chemical species (H_3O^+ , OH^- , $\bullet OH$, e_{aq}^- , $H\bullet$,*
 462 *H_2 , H_2O_2) produced by incident 100 MeV protons as a function of radial distance from the*
 463 *GNP. Results are shown at seven different times after irradiation (1 ps, 10 ps, 100 ps, 1 ns, 10*
 464 *ns, 100 ns, 1 μ s).*

465 Figure 18 and **Figure 19** show the REF of the distribution of chemical species (e_{aq}^- , $\bullet OH$)
 466 produced by incident 2 MeV and 100 MeV protons as a function of radial distance from the
 467 GNP. Results at six different times after irradiation (10 ps, 100 ps, 1 ns, 10 ns, 100 ns, 1 μ s)
 468 show a similar distribution for the two species.

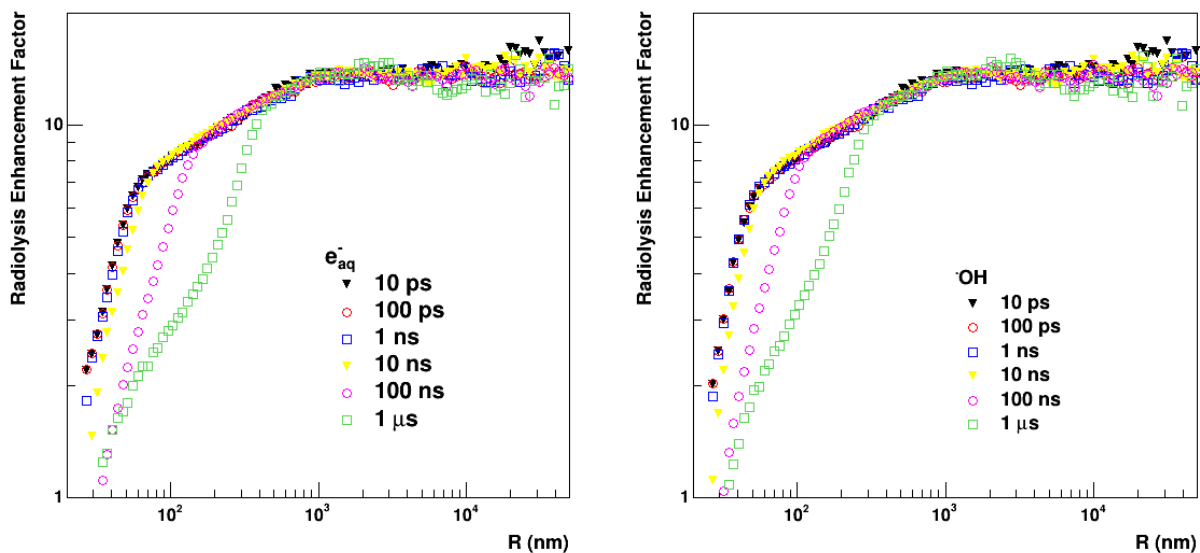
469

470



471 *Figure 18: Radiolysis Enhancement Factor (REF) of the distribution of chemical species (e_{aq}^- ,*
 472 *$\bullet OH$) produced by incident 2 MeV protons as a function of radial distance from the NP. Results*
 473 *are shown at six different times after irradiation (10 ps, 100 ps, 1 ns, 10 ns, 100 ns, 1 μs).*

474



475 *Figure 19: Radiolysis Enhancement Factor (REF) of the distribution of chemical species (e_{aq}^- ,*
 476 *$\bullet OH$) produced by incident 100 MeV protons as a function of radial distance from the NP.*

477 Results are shown at six different times after irradiation (10 ps, 100 ps, 1 ns, 10 ns, 100 ns, 1
 478 μ s).

479 The average total number of each chemical species per incident proton as a function of both
 480 evolution time and incident proton energy is presented in Table 2. Inspection of this data
 481 enables us to draw three main conclusions:

- 482
- 483 • For a given incident energy, there is a decrease with increasing time of the average
 484 number of chemical species, both for the GNP and WNP cases. This is due to the
 485 recombination of the species to liquid water.
 - 486 • For a given time, there is a decrease with increasing incident energy of the average
 487 number of chemical species, both for the GNP and WNP cases. This is related to the
 488 decrease of the incident particle LET in the NP with its incident kinetic energy, in the
 489 studied energy range.
 - 490 • For a given (time, energy) combination, the average number of chemical species
 491 generated around the GNP is always significantly larger than the average number of
 492 chemical species generated around the WNP. This is a direct consequence of the larger
 493 generation of secondary electrons in gold than in liquid water.
- 494

Time	2 MeV		5 MeV		10 MeV		50 MeV		100 MeV		170 MeV	
	GNP	WNP	GNP	WNP	GNP	WNP	GNP	WNP	GNP	WNP	GNP	WNP
10 ps	80.3179	14.0348	69.9185	8.5858	49.4015	5.47051	18.0477	1.71277	11.3257	0.99165	7.92319	0.693841
100 ps	75.7778	13.2372	66.1415	8.14597	47.0519	5.20458	17.2413	1.63309	10.6674	0.965804	7.62902	0.671854
1 ns	64.7229	11.4046	58.1127	7.08668	41.7588	4.5966	15.4223	1.44684	9.66332	0.861639	6.77494	0.588084
10 ns	51.979	9.2844	48.5478	5.87829	35.576	3.78756	13.3285	1.23121	8.42813	0.736346	5.76033	0.489763
100 ns	39.0212	8.16022	39.3362	5.14844	29.541	3.33195	11.632	1.0945	7.35567	0.660627	4.94079	0.424286
1 μ s	31.9774	7.68632	33.4445	4.75699	25.8717	3.11938	10.4302	1.01443	6.58677	0.635982	4.3565	0.405011

495 Table 2: Average numbers of radiolytic species generated per incident proton around the GNP
 496 and WNP at different times after the physical stage.

497

498 **4. Conclusion**

499 For the first time, we have presented in this work the combination of Geant4 physics modeling
500 capabilities with Geant4-DNA physics, physico-chemical and chemical modeling capabilities,
501 for the study of the irradiation of a gold **spherical** nanoparticle immersed in liquid water by an
502 incident proton beam. This combination allowed us studying the absorbed dose enhancement
503 and the production of chemical species around the nanoparticle for different incident energies
504 (2 MeV – 170 MeV) and at different times after irradiation up to 1 μ s. The reader should
505 **however** keep in mind that the physico-chemical and chemical module of Geant4-DNA is still
506 in a prototype state. The presented results may be affected by the evolution of the description
507 of the processes, in particular in the physico-chemical stage, as for instance the electron-**cation**
508 recombination or the way the radiolytic products are placed after water dissociation, which is a
509 source of uncertainty for this type of modeling. **Moreover, Geant4 existing ionization models**
510 **for incident protons (the G4BraggModel and the G4BetheBlochModel) cannot simulate the**
511 **production of delta electrons below the mean ionization potential of the medium, leading to an**
512 **underestimation of the number of delta electrons at very low energies. This may directly**
513 **influence the DEF in close vicinity of the NP as well as the associated production of chemical**
514 **species.**

515 The simulation of Dose Enhancement Factors (DEF) and Radiolysis Enhancement Factors
516 (REF) between GNP and WNP as a function of radial distance show similar trends for the
517 simulated incident proton energy range. While lower incident proton energies show higher
518 absorbed dose distribution near the NP, higher DEF and REF are found with higher incident
519 proton energies: there is thus a competition near the NP between DEF (and REF) on one side,
520 and absorbed dose distribution on the other side, as a function of incident proton energy.
521 Further mechanistic investigation at larger scale (for example at the cellular or tissue scale),
522 taking into account the full transport of protons in the water medium, and using a larger

523 number of NPs (including a study of coating influence) is still required before being able to
524 draw any conclusion related to possible benefits of using gold NPs in proton irradiation.
525 However, this work illustrates the current technical capabilities of the Geant4-DNA extension
526 for further study of involved physical, physico-chemical and chemical mechanisms.

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533

534

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