

# Gondwanide continental collision and the origin of Patagonia

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## Abstract

A review of the post-Cambrian igneous, structural and metamorphic history of Patagonia, largely revealed by a five-year programme of U–Pb zircon dating (32 samples), geochemical and isotope analysis, results in a new Late Palaeozoic collision model as the probable cause of the Gondwanide fold belts of South America and South Africa.

In the northeastern part of the North Patagonian Massif, Cambro-Ordovician metasediments with a Gondwana provenance are intruded by Mid Ordovician granites analogous to those of the Famatinian arc of NW Argentina; this area is interpreted as Gondwana continental crust at least from Devonian times, probably underlain by Neoproterozoic crystalline basement affected by both Pampean and Famatinian events, with a Cambrian rifting episode previously identified in the basement of the Sierra de la Ventana. In the Devonian (following collision of the Argentine Precordillera terrane to the north), the site of magmatism jumped to the western and southwestern margins of the North Patagonian Massif, although as yet the tectonics of this magmatic event are poorly constrained. This was followed by Early Carboniferous I-type granites representing a subduction-related magmatic arc and Mid Carboniferous S-type granites representing crustal anatexis. The disposition of these rocks implies that the North Patagonian Massif was in the upper plate, with northeasterly subduction beneath Gondwana prior to the collision of a southern landmass represented by the Deseado Massif and its probable extension in southeastern Patagonia. This ‘Deseado terrane’ may have originally rifted off from a similar position during the Cambrian episode. Intense metamorphism and granite emplacement in the upper plate continued into the Early Permian. Known aspects of Late Palaeozoic sedimentation, metamorphism, and deformation in the Sierra de la Ventana and adjacent Cape Fold Belt of South Africa are encompassed within this model. It is also compatible with modern geophysical and palaeomagnetic data that do not support previous hypotheses of southward-directed subduction and collision along the northern limit of Patagonia. Subsequent Permian break-off of the subducted plate, perhaps with delamination of the lower part of the upper plate, allowed access of heat to the overlying Gondwana margin and resulted in voluminous and widespread silicic plutonism and volcanism throughout Permian and into Triassic times. Thus the new model addresses and attempts to explain three long-standing geological enigmas—the origin of the Gondwanide fold belts, the origin of Patagonia, and the cause of widespread Permian silicic magmatism (Choiyoi province) in southern South America. Differing significantly from previous models, it has new implications for the crustal structure, mineral resources, and plant and animal distribution in this part of Gondwana, since the southern landmass would have had an independent evolution throughout the Early Palaeozoic.

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## 1. Introduction

Patagonia, conventionally considered as the continental region south of the Río Colorado (Fig. 1), has long been recognized as differing from the rest of South America in terms of its topography, environment, flora, fauna and palaeontological record. The idea that it could have had a separate geological history (Keidel, 1925) was stimulated by the recognition of terrane accretion processes in the early 1980s. Ramos (1984, 1986) proposed that an allochthonous (exotic) Patagonian terrane collided with cratonic South America (super-continental Gondwana) along the Río Colorado zone in Carboniferous times. In his model, the suture (observed

by much younger sediments) formed by closure of a previously intervening ocean, due to southwest-dipping subduction beneath the North Patagonian Massif. In more recent reviews of the tectonic evolution of Patagonia, Ramos (2002, 2004) has modified this idea to include a prior Early Palaeozoic collision within the Deseado Massif, also thought to result from southward-directed subduction, prior to Late Palaeozoic collision of the combined landmass so formed with cratonic South America, as before. In partial support of these models, Devonian–Carboniferous penetrative deformation, southward-verging folds and southward-directed thrusting of supracrustal rocks of the northeastern North Patagonian Massif was described by Chernikoff and

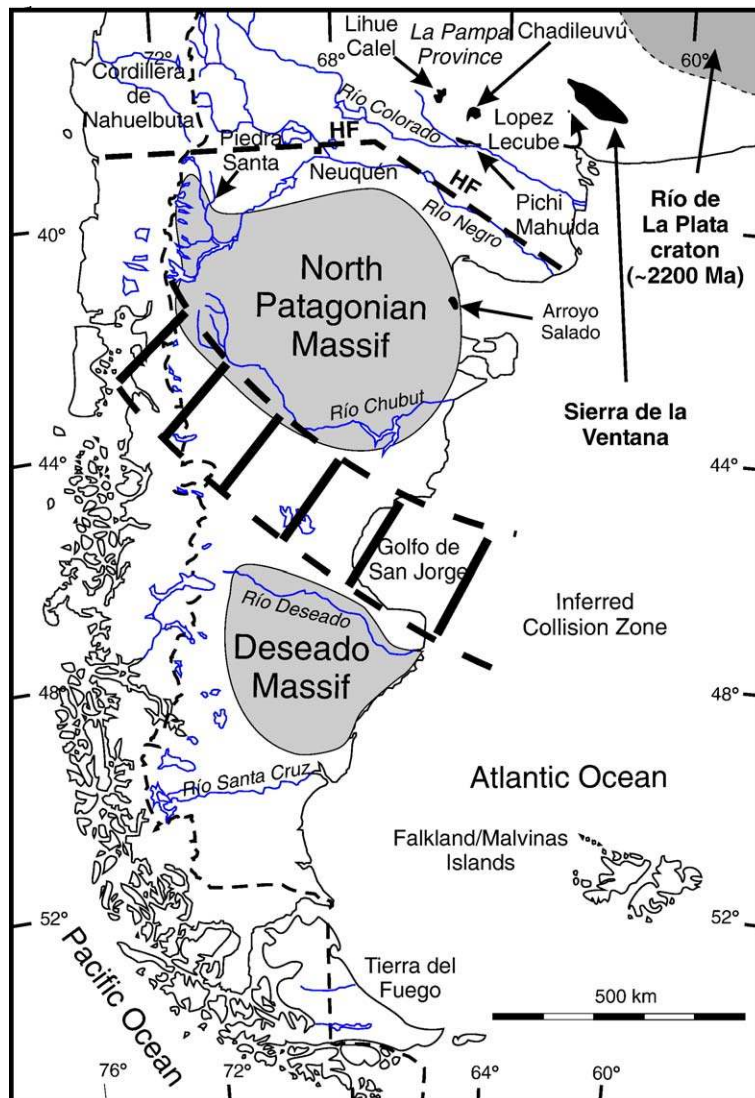


Fig. 1. Sketch map of Patagonia, showing the main pre-Jurassic tectonic elements, and the geographical relationship to the Sierra de la Ventana fold belt, as well as other significant basement rock exposures referred to in the text. HF=Huincul fault (Chernikoff and Zappettini, 2004).

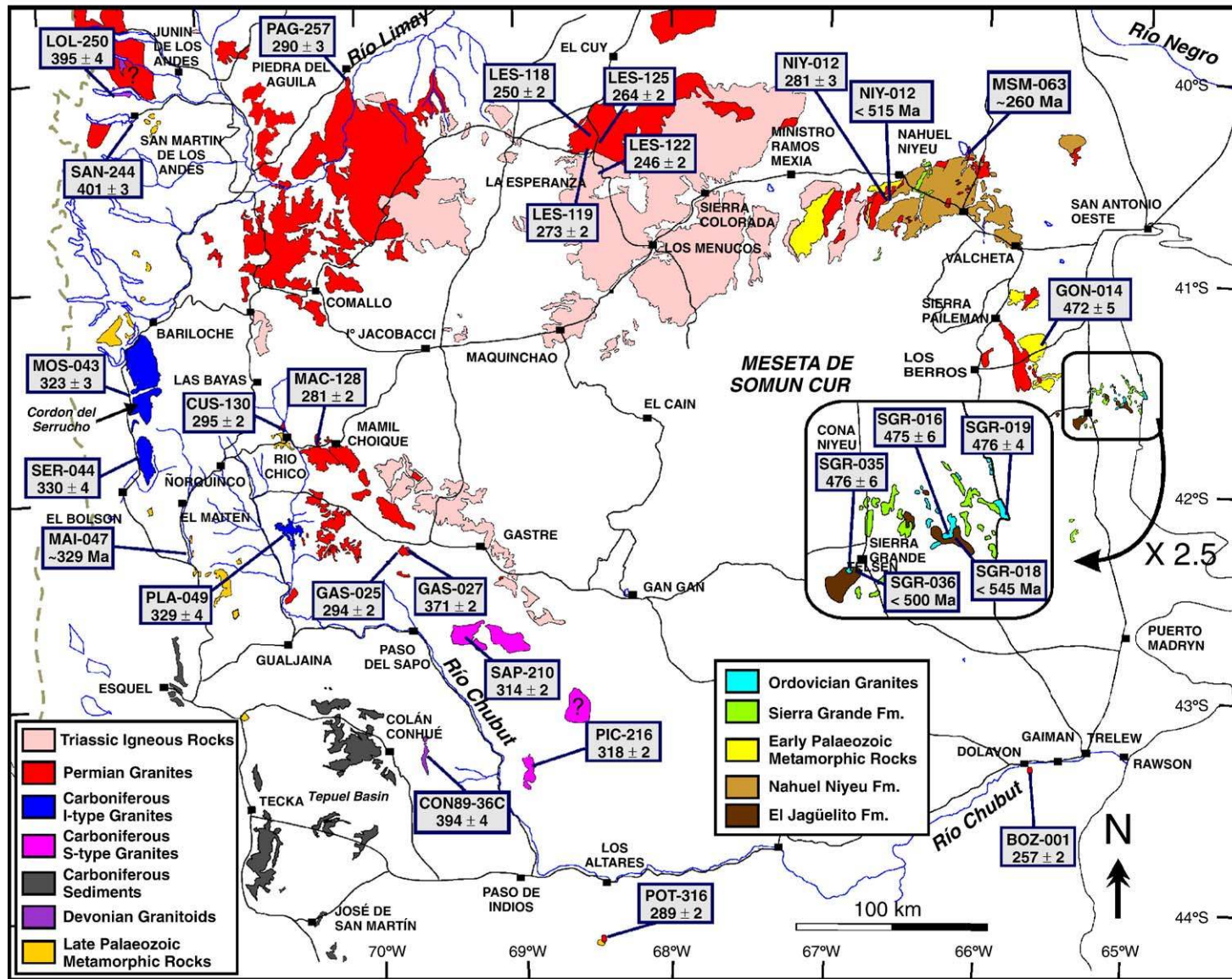


Fig. 2. Sketch map of the North Patagonian Massif, showing the pre-Cretaceous geology, the location of analysed samples, and the results of U–Pb zircon dating presented here (ages in Ma with 95% c. l. errors).

Table 1  
Geochemical data for Patagonian granites

	Subduction-related granites			Collision granites				Post-orogenic granites							
	MOS 43	SER 44	PLA 49	SAP 210	PIC 213	PIC 216	SAP 209	CUS 130	BOZ 1	GAS 25	NIY 010	MAC 128	MSM 063	POT 316	SLV 109
<i>Sample (wt.%)</i>															
SiO <sub>2</sub>	53.25	59.25	63.53	62.08	70.08	63.29	76.89	60.24	60.70	62.00	66.10	67.55	70.35	74.50	59.92
TiO <sub>2</sub>	0.93	0.81	0.44	0.72	0.39	0.81	0.23	0.90	0.68	0.91	0.33	0.48	0.30	0.01	0.85
Al <sub>2</sub> O <sub>3</sub>	18.59	17.31	17.24	17.29	15.64	17.55	11.33	17.93	15.36	16.99	16.50	16.07	14.18	14.36	14.40
Fe <sub>2</sub> O <sub>3t</sub>	8.23	6.82	4.26	5.44	2.52	5.15	1.70	5.61	4.93	5.99	3.28	3.26	2.44	1.13	5.86
MnO	0.17	0.13	0.07	0.07	0.02	0.03	0.02	0.06	0.12	0.08	0.06	0.06	0.05	0.02	0.10
MgO	4.34	3.06	2.47	2.03	0.92	2.22	0.48	2.20	3.84	2.57	1.35	1.12	0.88	0.04	2.62
CaO	7.05	6.55	5.05	5.04	3.64	4.74	1.47	6.36	4.27	5.61	2.90	3.51	1.45	1.21	3.94
Na <sub>2</sub> O	4.00	3.74	4.77	2.90	3.30	2.97	1.74	2.63	3.37	3.30	5.42	3.79	3.81	3.41	4.28
K <sub>2</sub> O	1.22	1.24	1.29	2.32	1.76	2.01	5.00	1.03	3.28	2.32	2.36	2.82	4.41	5.10	5.40
P <sub>2</sub> O <sub>5</sub>	0.14	0.23	0.10	0.22	0.13	0.33	0.04	0.20	0.17	0.24	0.14	0.17	0.13	0.04	0.46
LOI	2.10	0.89	0.74	0.98	0.80	0.85	0.85	2.10	3.00	0.40	1.12	0.44	0.82	0.07	0.78
Total	100.01	100.03	99.96	99.09	99.20	99.95	99.75	99.26	99.72	100.41	99.56	99.27	98.82	99.89	98.61
<i>Trace elements (ppm)</i>															
Cs	3.4	1.4	1	4.9	0.8	2.5	3	6	2.7	b.d.	b.d.	2.3	2.6	3.1	3.1
Rb	39	34	25	153	42	87	141	42	127	75	52	108	163	232	163
Sr	373	453	424	318	517	316	193	512	540	350	1030	449	596	286	1350
Ba	277	266	222	490	933	537	1153	415	1151	584	1050	599	1410	1157	2390
La	11.8	21.2	7.82	43	38.3	11.1	69.7	47.7	20	31.5	9.33	24.4	26.4	13.2	78.8
Ce	28.2	45	17.6	87.9	76.9	25	142	94.1	55.8	64.5	18.5	50.6	53.9	20.3	156
Pr	4.2	5.89	2.47	10.2	8.89	3.22	17.1	9.91	7.06	8.02	2.25	5.58	6.49	1.99	16.7
Nd	17.9	22.2	9.76	39.2	35.2	13.6	56	37.6	25.8	30.5	8.61	21.2	23.6	6.09	65
Sm	5.09	5.07	2.68	7.24	6.1	3.35	8.81	6.38	4.96	6.28	1.75	4.07	4.59	1.13	11.7
Eu	1.63	1.56	0.76	1.62	1.92	1.49	1.33	1.49	1.25	1.48	0.58	0.97	0.98	2.32	2.96
Gd	5.45	4.63	2.8	4.77	3.18	2.97	5.01	4.89	3.95	5	1.62	3.19	3.45	1.04	8.3
Tb	0.98	0.74	0.51	0.59	0.26	0.52	0.54	0.68	0.57	0.75	0.26	0.48	0.48	0.23	1.09
Dy	5.83	4.09	3.01	2.69	0.84	2.4	2.33	3.37	3.04	3.67	1.42	2.38	2.44	1.83	5.37
Ho	1.19	0.81	0.62	0.51	0.09	0.29	0.35	0.62	0.57	0.69	0.3	0.46	0.46	0.49	0.96
Er	3.51	2.38	1.87	1.64	0.2	0.62	0.92	1.64	1.65	1.87	0.9	1.29	1.32	2.04	2.82
Tm	0.54	0.36	0.28	0.24	b.d.	0.07	0.13	0.21	0.24	0.25	0.13	0.19	0.18	0.43	0.39
Yb	3.23	2.18	1.78	1.4	0.14	0.45	0.9	1.23	1.53	1.4	0.83	1.22	1.1	3.52	2.46
Lu	0.51	0.35	0.28	0.19	0.01	0.06	0.14	0.18	0.2	0.18	0.13	0.19	0.15	0.55	0.39
U	0.35	0.84	0.55	1.64	0.45	0.84	2.22	0.35	1.81	1.58	1.22	0.84	3.06	1.43	3.89
Th	0.53	4.07	1.32	10.7	9.32	0.46	25.5	8.11	6.6	6.72	3.94	7.95	11.7	0.46	18.7
Y	13.1	22.5	17.7	14.7	2.8	9.2	10.8	14.6	18.5	20.2	9.4	12.1	14.6	16.1	24.1
Nb	3.4	5.8	3.9	11	3.8	10.5	4.7	9.4	7.3	10.8	4.9	11.3	7.8	0.3	21.8
Zr	120	157	89	158	144	163	166	219	164	254	120	133	150	22	364
Hf	3.3	4.2	2.6	4.2	4.2	4.5	5	6.1	4.4	5.9	3	4	3.9	2	10
Ta	0.22	0.41	0.27	0.69	0.05	0.74	0.48	0.41	0.52	9.71	0.34	1.43	0.5	0.01	1.74
Sc	24	18	11	11	2	8	4	13	12	14	5	7	5	2	11
Ga	19	19	17	22	23	21	13	23	21	22	21	22	21	24	22

Analyses carried out by ACTLABS, Canada.

Major elements determined by ICP spectrometry.

Minor elements determined by ICP-MS spectrometry.

b.d.=below detection limit.

Camino (1996) and elaborated in a detailed structural study by von Gosen (2003), who argued for Permian rather than Carboniferous crustal shortening, and possibly a northeastward-directed accretionary process.

However, geophysical data (Chernikoff and Zappettini, 2004; Kostadinoff et al., 2005) fails to reveal

any significant crustal discontinuity beneath the Rio Colorado basin, and the physiographical boundary of Patagonia has been moved south to the line of the Huinacul fault, which in its eastern part follows the Rio Negro. Kostadinoff et al. (2005) interpret gravimetric anomalies as indicating a basement of

Table 2  
Nd Isotope data for samples from Patagonia

Sample	Age (Ma)	Sm ppm	Nd ppm	$^{147}\text{Sm}/$ $^{144}\text{Nd}$	$^{143}\text{Nd}/$ $^{144}\text{Nd}$	$^{143}\text{Nd}/$ $^{144}\text{Nd}$ initial	$\epsilon\text{Ndt}$	$T_{\text{DM}}$ (Ma)	Rb ppm	Sr ppm	$^{87}\text{Rb}/$ $^{86}\text{Sr}$	$^{87}\text{Sr}/^{86}\text{Sr}$	$^{87}\text{Sr}/^{86}\text{Sr}$ initial	
<i>Patagonia-metasedimentary</i>														
SGR-018	El Jaguelito Fm	530	5.24	25.91	0.1223	0.512190	0.511765	-3.7	1533	80	101	2.305	0.728736	0.711323
NIY-012	Nahuel Niyeu Fm	500	6.29	31.99	0.1189	0.512108	0.511718	-5.4	1636	66	89	2.147	0.728092	0.712798
GON-014	Mina Gonzalito gneiss	475	8.40	41.41	0.1226	0.512163	0.511782	-4.8	1574	121	163	2.161	0.725244	0.710619
MAI-047	El Maitén gneiss	330	5.40	28.15	0.1159	0.512109	0.511859	-6.9	1630	72	239	0.868	0.714467	0.710389
<i>Patagonia-igneous</i>														
SGR-016	Arroyo Salado granite	475	4.38	20.51	0.1290	0.512274	0.511873	-3.0	1440	168	172	2.829	0.727472	0.708327
SGR-035	Sierra Grande granite	475	7.12	35.27	0.1220	0.512258	0.511878	-2.9	1432	164	214	2.215	0.723166	0.708178
SGR-019	Punta Bahia granite	475	3.92	19.51	0.1214	0.512269	0.511891	-2.7	1412	129	126	2.960	0.727316	0.707281
PIM-113	Rio Colorado granodiorite	475	6.65	31.02	0.1297	0.512191	0.511787	-4.7	1565	153	160	2.785	0.732104	0.713259
PIM-115	Curaco granodiorite	475	3.93	15.83	0.1502	0.512221	0.511753	-5.3	1614	274	67	11.982	0.792570	0.711477
AB 157A	Lgo. Curruhue <sup>a</sup>	400	5.16	31.25	0.0998	0.512091	0.511830	-5.7	1592					
LCU-251	Lgo. Curruhue gneiss	400	3.36	17.77	0.1143	0.512139	0.511840	-5.5	1578	89	118	2.185	0.725960	0.713514
AB 13	San Martin tonalite <sup>a</sup>	400	6.43	31.63	0.1229	0.512220	0.511898	-4.4	1493	119	324	1.063	0.712180	0.706124
AB 152	San Martin tonalite <sup>a</sup>	400	6.65	32.44	0.1239	0.512286	0.511961	-3.2	1400	138	294	1.359	0.713970	0.706230
AB 154	C°Curruhuinca <sup>a</sup>	400	6.40	33.94	0.1140	0.512214	0.511915	-4.0	1468					
SAN-248	San Martin tonalite (enclave)	400	5.82	27.68	0.1272	0.512222	0.511889	-4.6	1507	182	227	2.321	0.720474	0.707256
LOL-250	Lago Lolog granite	400	3.35	16.32	0.1239	0.512360	0.512035	-1.7	1287	121	371	0.943	0.710517	0.705144
GAS-027	Caceres granite	372	7.62	34.32	0.1343	0.512229	0.511902	-5.0	1521	157	195	2.330	0.722228	0.709887
AB 165B	West of Sañicó <sup>a</sup>	360	6.56	29.89	0.1327	0.512030	0.511682	-8.6	1794					
AB 165A	West of Sañicó <sup>a</sup>	360	7.44	36.63	0.1228	0.512106	0.511784	-6.6	1655					
MOS-043	Canadon de la Mosca	323	5.26	19.97	0.1591	0.512564	0.512227	0.1	1084	39	168	0.306	0.705783	0.704374
AB 123	Cdón. Mosca <sup>a</sup>	323	4.72	18.47	0.1545	0.512519	0.512192	-0.6	1141					
SER-044	Serrucho granodiorite	330	4.96	22.19	0.1352	0.512566	0.512274	1.2	996	34	168	0.217	0.705646	0.704625
PLA-049	Platero granodiorite	330	2.57	10.77	0.1442	0.512667	0.512356	2.8	854	25	168	0.174	0.704189	0.703373
PIC-216	Pichinanes Bio-Grt granite	318	3.30	12.01	0.1659	0.512316	0.511970	-5.0	1485	108	168	0.956	0.712627	0.708301
SAP-209	Paso del Sapo mylonitised granite	314	5.60	35.91	0.0942	0.512161	0.511967	-5.2	1494	144	168	2.042	0.718937	0.709812
SAP-210	Paso del Sapo foliated granodiorite	314	4.33	21.49	0.1217	0.512204	0.511954	-5.5	1514	160	168	1.488	0.714411	0.707764
CUS-130	Rio Chico, 'Tunnel' tonalite	296	5.81	33.50	0.1048	0.512193	0.511990	-5.2	1483	40	168	0.217	0.708522	0.707608
91 RC26	Rio Chico, 'Tunnel' tonalite <sup>a</sup>	296	4.14	18.49	0.1354	0.512360	0.512098	-3.1	1324					
GAS-025	Laguna del Toro granodiorite	293	5.64	31.58	0.1080	0.512282	0.512075	-3.6	1362	83	168	0.664	0.709501	0.706734
PIC-213	Pichinanes granite	289	5.37	31.68	0.1024	0.512141	0.511928	-5.9	1546	44	168	0.241	0.710214	0.709124

(continued on next page)

Table 2 (continued)

Sample		Age (Ma)	Sm ppm	Nd ppm	<sup>147</sup> Sm/ <sup>144</sup> Nd	<sup>143</sup> Nd/ <sup>144</sup> Nd	<sup>143</sup> Nd/ <sup>144</sup> Nd initial	$\epsilon$ Ndt	$T_{DM}$ (Ma)	Rb ppm	Sr ppm	<sup>87</sup> Rb/ <sup>86</sup> Sr	<sup>87</sup> Sr/ <sup>86</sup> Sr	<sup>87</sup> Sr/ <sup>86</sup> Sr initial
POT-316	La Potranca leucogranite	289	25.96	112.82	0.1391	0.512055	0.511792	−9.3	1765	124	168	1.201	0.716822	0.711883
NIY-010	Navarette granodiorite	281	2.49	12.02	0.1253	0.512400	0.512170	−2.1	1233	59	168	0.166	0.705408	0.704745
NIY-011	Navarette quartz porphyry	281	2.76	12.96	0.1289	0.512363	0.512126	−2.9	1300	86	168	0.254	0.705685	0.704670
MAC-127	Leucogranite (La Pintada)	281	5.27	15.68	0.2031	0.512308	0.511934	−6.7	1579	210	168	12.211	0.752954	0.704134
MAC-128	Puesto Quintuplepu granodiorite	281	5.15	28.88	0.1078	0.512245	0.512047	−4.5	1418	118	168	0.774	0.709876	0.706782
AB 27C	Comallo <sup>a</sup>	280	5.77	30.88	0.1129	0.512125	0.511918	−7.0	1603					
LES-119	Prieto granodiorite	273	6.19	33.58	0.1113	0.512232	0.512033	−5.0	1448	161	168	1.065	0.710794	0.706659
AB 30C	Cerro Yuncón <sup>a</sup>	270	6.12	21.67	0.1707	0.512217	0.511915	−7.3	1618					
AB 30B	Cerro Yuncón <sup>a</sup>	270	3.80	10.10	0.2275	0.512492	0.512090	−3.9	1368					
AB56	Loma Miranda <sup>a</sup>	270	0.64	1.67	0.2317	0.512630	0.512221	−1.4	1168					
AB 120	Loma Miranda <sup>a</sup>	270	5.95	34.19	0.1052	0.512215	0.512029	−5.1	1457					
AB 121	Paso Flores <sup>a</sup>	270	10.27	34.96	0.1776	0.512241	0.511927	−7.1	1601					
LES-120	Donoso granite	270	2.98	17.59	0.1023	0.511968	0.511787	−9.8	1791	139	168	0.698	0.710168	0.707487
LES-125	La Esperanza rhyolite dome	264	3.85	23.65	0.0985	0.512163	0.511993	−6.0	1516	427	168	6.500	0.731292	0.706879
MSM-063	Mina San Martin granite	260	4.78	24.25	0.1192	0.512227	0.512024	−5.4	1476	153	168	0.757	0.710737	0.707938
SLV-109	Lopez Lecube syenite	258	11.38	64.87	0.1060	0.512200	0.512021	−5.6	1483	169	168	0.339	0.707617	0.706371
BOZ	Boca de la Zanja	257	5.47	27.55	0.1200	0.512253	0.512051	−5.0	1441	118	168	0.645	0.710755	0.708395
LES-118	Calvo granite	250	1.83	11.46	0.0966	0.512081	0.511923	−7.7	1630	377	168	29.096	0.807041	0.703565
International Standards														
JB-1			5.23	26.82	0.1178	0.512790								
G-2										487	0			
GSP-1										242	1			
NBS70a	(mean of 3)									66	8			

Isotope analyses carried out at NERC Isotope Geosciences Laboratory, BGS, Keyworth, Nottingham (Sm and Nd by TIMS-ID, Rb and by XRF). <sup>143</sup>Nd/<sup>144</sup>Nd relative to 0.511864 for La Jolla, <sup>87</sup>Sr/<sup>86</sup>Sr relative to 0.710235 for NBS987.

$T_{DM}$  = variable crust Sm/Nd multistage.

(DePaolo, D.J., Linn, A.M., Schubert, G., 1991. The continental crustal age distribution: methods of determining mantle separation ages from Sm–Nd isotopic data and application to the Southwestern United States. *J. Geophys. Res.* B96, 2071–2088).

<sup>a</sup> Data from Varela et al. (2005).

probable Early Palaeozoic age extending yet farther to the south.

Moreover, plate subduction and collision should result in a recognized pattern of related magmatism, metamorphism and crustal melting. The original allochthonous Patagonia hypothesis suffered a setback when it was shown that supposedly Carboniferous granites in the North Patagonian Massif were in fact Permian to Triassic in age (Pankhurst et al., 1992). All existing collision models lack supporting magmatic evidence for the postulated subduction and collision phases (i.e., subduction- and collision-related granite belts, and related metamorphism, of the right age and in the right relationship to the plate model).

We present the results of extensive study over the past 6 years of the ages and geochemistry of igneous and metamorphic rocks that form the pre-Mesozoic basement in Patagonia, and the provenance of some pre-Permian metasedimentary rocks. It must be emphasized that these outcrops are very small (often only a few hundred metres in plan) and sparsely distributed, due to the later extensive covering of Jurassic volcanic rocks, Cretaceous–Quaternary sedimentary basins, and Tertiary basaltic lavas. Nevertheless, the results, when combined with existing constraints, suggest that the North Patagonian Massif was already part of Gondwana in Ordovician times. In contrast, southern Patagonia seems to have belonged to an allochthonous entity that

collided with the North Patagonian Massif in Mid Carboniferous times as a result of north-easterly ocean-floor subduction. The implications and consequences of this model are reviewed, especially in relation to the role of continental collision in the formation of the Gondwana fold belts of South America, Africa and West Antarctica, as well as the distribution of Early Palaeozoic faunal provinces.

## 2. Analytical methods

Throughout this programme, we have dated the crystallization ages of zircon in igneous and metamorphic rocks by U–Pb geochronology using SHRIMP (Sensitive High-Resolution Ion Microprobe) technology at The Australian National University, Canberra (Williams, 1998). Some results previously published in conference proceedings have been recalculated using SQUID (Ludwig, 2000). The relevant data, together with precise sample localities, are presented as a Supplementary Appendix to this paper, and are summarised on a sketch map of northern Patagonian (Fig. 2). In addition to dating metamorphic events by analysis of metamorphic overgrowths, where possible, the detrital zircons in metasedimentary rocks have been dated to provide information on provenance of the protoliths. We have also undertaken whole-rock geochemistry (Table 1) and Sr–Nd isotope analysis (Table 2) in an attempt to constrain the tectonic environment of magmatism.

A parallel study using conventional U–Pb zircon dating has been published recently by Varela et al. (2005), in which earlier Rb–Sr and K–Ar are also reviewed (some of which must now be regarded as of dubious reliability in terms of crystallization age but, as pointed out by these authors, more probably relate to cooling and/or metamorphic effects). As indicated by Varela et al. (2005), the main events in the evolution of Patagonia are of Palaeozoic age, and metasedimentary rock units that can be ascribed to the latest Precambrian are mainly restricted to the northeastern North Patagonian Massif.

In view of the wide time-span covered by the Palaeozoic evolution of Patagonia, it is convenient to present the new results and discuss them for each of the tectonic stages that we now recognise. The time-scale used is that of Gradstein et al. (2004).

## 3. Cambrian rifting

The oldest stable area of continental basement in southern South America is the 2000–2200 Ma Río de la Plata craton (Santos et al., 2003). The Palaeozoic

sedimentary sequence of Sierra de la Ventana (also known as the Sierras Australes de Buenos Aires) crops out south of the proven extent of the craton, and rests on younger (Neoproterozoic) crystalline basement with crustal anatectic granite at  $607 \pm 5$  Ma and I/A-type granites at  $531 \pm 4$  and  $524 \pm 5$  Ma (Rapela et al., 2003). These igneous events correspond to the Brasiliano/Pampean orogenic cycles of South America farther north, and may now be thought of in terms of the final assembly of Gondwana (Veevers, 2005). The culminating phase of igneous activity here is represented by Late Cambrian ( $\sim 510$  Ma) peralkaline rhyolites that were derived from lithospheric mantle, and may be interpreted as recording extensional tectonics related to rifting of this part of the Gondwana margin. Rapela et al. (2003) suggested that this rifting episode resulted in the separation of a continental fragment but could not specifically identify it within the present-day Pacific margin of West Antarctica, which is a collage of small crustal blocks (see Rapela et al., 2003, Fig. 7). However, other authors agree that a Late Cambrian passive margin was established at the southern edge of Gondwana (as elsewhere, with reference to present-day geographical coordinates) along the span of the subsequent Gondwanide fold belt, e.g., Curtis (2001). Dalziel (1997) suggested that the rifted-away portion was a large plateau area attached to Laurentia and that this subsequently collided much farther to the north in the Ordovician, leaving behind a fragment represented by the Argentine Precordillera.

## 4. Ordovician magmatism

In northwest Argentina, the Early Cambrian Pampean orogenic belt lies to the west of the Río de la Plata craton, and was partially overprinted by the Early to Mid Ordovician Famatinian magmatic arc (Pankhurst et al., 1998, 2000), with intensive deformation in the Late Ordovician. The Famatinian belt has been considered as a continental marginal arc related to subduction during the approach and collision of the Precordillera terrane. However, recent geochronological and structural analysis suggests that Mesoproterozoic ‘Grenvillian’ rocks exposed in the Western Sierras Pampeanas may have become part of autochthonous Gondwana much earlier, being accreted to the Río de la Plata craton during the Pampean event (Escayola et al., 2005; Mulcahy et al., 2005; Rapela et al., 2005a). Such a continental basement, thinned during the Late Cambrian rifting, seems a likely candidate for the crust underlying the Sierra de la Ventana and, perhaps, the northwestern part of the North Patagonian Massif (see below).

Previous geochronological evidence has been advanced for extension of the Famatinian magmatic belt southeastwards through La Pampa province, as far as Chadileuvú just north of the Río Colorado and on to the Arroyo Salado area (Fig. 1) in the eastern part North Patagonian Massif (Tickyj et al., 1999; Varela et al., 1998). Our data confirm these results and establish the precise contemporaneity of magmatism in these areas, yielding indistinguishable Early Ordovician U–Pb SHRIMP ages of  $474 \pm 6$  Ma and  $475 \pm 5$  Ma for a (locally) early granodiorite and a late granite from Pichi Mahuida (Fig. 3a),  $475 \pm 6$  and  $476 \pm 4$  Ma for two granite bodies from the Arroyo Salado area, and  $476 \pm 6$  Ma for granite at Sierra Grande 25 km to the SW (Fig. 3b). Amphibolite-grade metamorphism associated with the Famatinian event in the North Patagonian Massif is recorded in quartzo-feldspathic gneiss from

Mina Gonzalito (Fig. 4a), where new zircon growth during amphibolite-grade metamorphism is dated at  $472 \pm 5$  Ma (Pankhurst et al., 2001, recalculated).

These Ordovician granites of the North Patagonian Massif do not show signs of the deformation characteristic of the contemporaneous Famatinian granites of northwest Argentina, but since the latter is thought to be a direct result of the collision of the Precordillera terrane, this would not be expected along the entire length of the magmatic belt, and there is no evidence for terrane collision at this time in the Patagonian region. Geochemical and isotopic data (Tables 1 and 2) show that these are equivalent to the igneous intrusions in the main part of the Famatinian belt, which are almost entirely intermediate to silicic, metaluminous, with evolved initial  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios (0.707–0.710) and  $\epsilon\text{Ndt}$  values (–2 to –6). It is possible that the magmatism

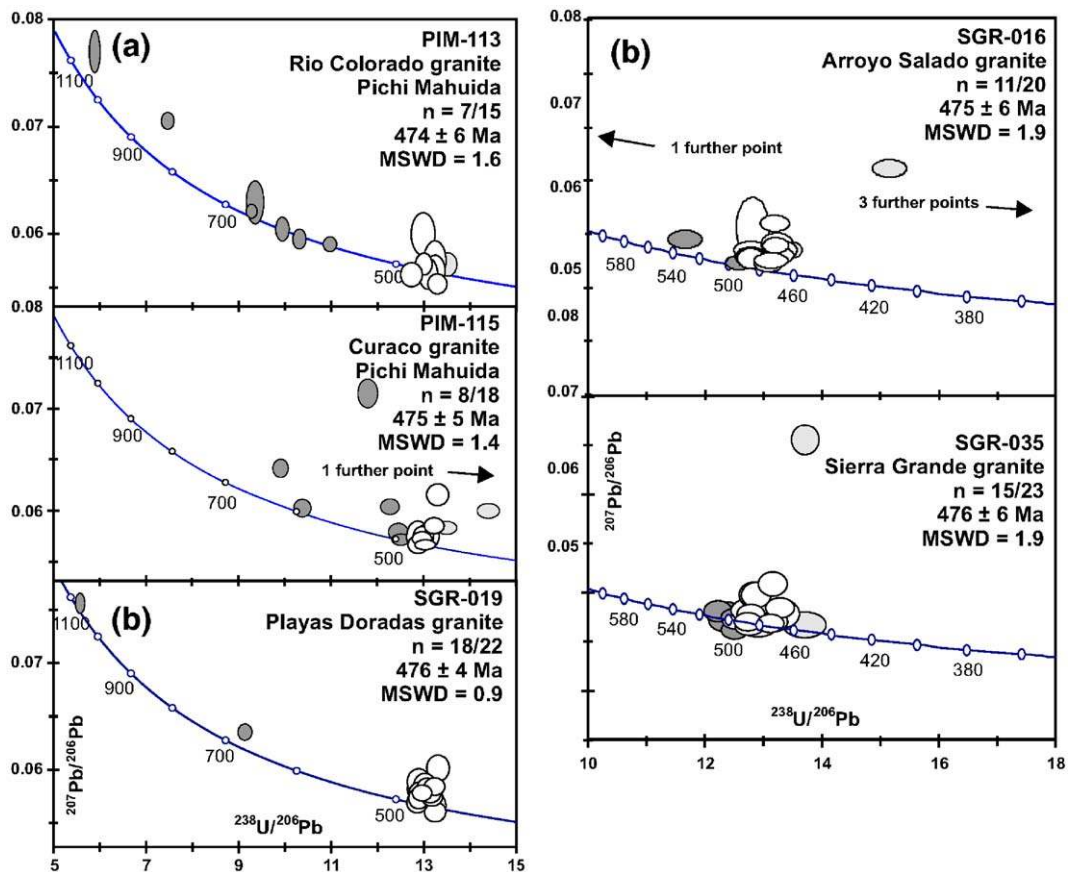


Fig. 3. New U–Pb SHRIMP zircon-dating results for Ordovician granitoids: (a) from Picchi Mahuida, Río Colorado (see Fig. 1) and (b) the northeastern North Patagonian Massif (see Fig. 2). The data (not corrected for common Pb content) are displayed in Tera-Wasserburg diagrams with 68.4% confidence limit error ellipses. The calculated age is the weighted mean of  $^{207}\text{Pb}$ -corrected  $^{238}\text{U}$ – $^{206}\text{Pb}$  ages for the white ellipses; light grey ellipses are for spots assumed to have suffered Pb-loss, and the dark grey ellipses are for those thought to contain a significant inherited component;  $n$  = number of points in age calculation/total number of areas analysed,  $\pm 2\sigma$  errors, MSWD = mean square of weighted deviates. The derived ages of granite intrusion in these two areas are indistinguishable, and fall within the range of Famatinian granitoids in northwestern Argentina.



throughout the belt was initiated by post-collisional lower crustal melting following the Pampean event, modified only in the northern areas by further collisional compression.

### 5. Pre-Carboniferous sedimentation

The country rocks of these granites in northeastern Patagonia consists of fine-grained meta-sandstones—the El Jagüelito and Nahuel Niyeu formations (see Fig. 2). González et al. (2002) argue for Cambrian deposition of the former on the basis of its relationship to the granites and its trace fossil content. Detrital zircons from the El Jagüelito Formation (Fig. 4a) have a youngest age peak at ~535 Ma consistent with this assignment: although this is strictly constraint on the maximum possible age, the dominance of this peak suggests erosion from a nearby active arc and in such cases it is commonly observed that the youngest detrital zircons effectively date deposition. That for the Nahuel Niyeu Formation is slightly younger at ~515 Ma (both have a very few grains with even younger ages that probably reflect partial Pb-loss during Ordovician granite emplacement or deformation). The older provenance of both samples is typically Gondwanan with ages of 550–750 Ma (Pampean and Brasiliano) and ~1000–1200 Ma ('Grenvillian'), as well as a small component of older ages, including some at ~2200 Ma. These patterns could relate to major provenance from the areas immediately to the north in Early to Mid Cambrian times, including the Río de la Plata craton (there is a very minor input of 2200–1900 Ma zircons), whereas the 'Grenvillian' component is consistent with a local basement similar to the Western Sierras Pampeanas as suggested above. The zircon cores from the Mina Gonzalito gneiss are similar, most notably in having their main provenance at 535–540 Ma (Fig. 4a), also suggesting Early Cambrian deposition and generation of the gneiss from the sandstone protolith by Famatinian-age metamorphism. Thus the metapelitic rocks of the northeastern North Patagonian Massif appear to have been originally deposited as sediments on a continental shelf at the southern margin of Gondwana.

Following intrusion of the Ordovician granites, the Sierra Grande sandstone–ironstone formation was deposited in Silurian or Early Devonian times (Limarino et al., 1999), one of the very few such deposits known after the Precambrian era. Its probable depositional environment was a quiescent platform on a stable passive margin, with no volcanic input (Spalletti, 1993). The invertebrate marine fauna is of Malvinokaffric type, similar to that found in the Falkland Islands. Palaeo-

magnetic data from this formation are consistent with Permian folding, but a prior pole position consistent with stable Gondwana was used to argue against it being part of a far-travelled exotic Patagonian terrane (Rapalini, 1998). A sample of ferruginous sandstone from just outside the Sierra Grande mine has a complex detrital zircon pattern (Fig. 4a) containing all the provenance elements seen in the El Jagüelito and Nahuel Niyeu formations, as well as a youngest peak at ~500 Ma. This suggests sediment derivation from the same source areas to the north, but after the eruption of the Late Cambrian rhyolites associated with rifting. One discordant grain has an apparent age of 470 Ma, but the general absence of Ordovician zircons suggests that the Famatinian granites were not exposed in the Silurian/Early Devonian source areas.

Thus one interpretation of the NE North Patagonian Massif is that it is underlain by basement rocks that were already part of the Gondwana continent by Cambrian times, perhaps thinned during the Late Cambrian rifting event, and affected by Famatinian plutonism in the Early Ordovician, then becoming a largely passive margin until post-Early Devonian times. Its essential integrity with the Gondwana margin in the Early Palaeozoic is consistent with the observation of Kostadinoff et al. (2005) that crustal magnetic signatures are continuous across the Huincul fault zone.

### 6. Devonian magmatism

After the Ordovician, active magmatism reappeared at the western margin of the North Patagonian Massif. Varela et al. (2005) have reported conventional  $^{238}\text{U}$ – $^{206}\text{Pb}$  zircon ages of  $419 \pm 27$  Ma (MSWD=43) and  $390.0 \pm 4.8$  Ma (MSWD=9) for tonalites near San Martín de los Andes (Fig. 2), and of  $348 \pm 11$  Ma (MSWD=63, possibly Carboniferous—see below) and  $386.6 \pm 5.4$  Ma (MSWD=9.3) for deformed leucogranites cutting schists about 50 km farther to the southwest. They ascribed these to a period of Devonian magmatism and migmatization associated with the Chanic orogenic event identified in the areas north of Patagonia (San Rafael, Precordillera and Sierras Pampeanas; Sims et al., 1998). We have confirmed and refined the age of this magmatism in NW Patagonia with U–Pb zircon SHRIMP ages of  $401 \pm 3$  Ma for the San Martín tonalite (Fig. 5a) and  $395 \pm 4$  Ma for an undeformed granite at Lago Lolog about 10 km farther north (Fig. 5b). Fig. 5a also illustrates ages of  $371 \pm 4$  Ma for megacrystic granite near Gastre in the southwestern North Patagonian Massif and  $394 \pm 4$  (recalculated from Pankhurst et al., 2001) for a similar megacrystic granite

at Colán Conhue 300 km southeast of Bariloche (Fig. 2) which, as suggested by Varela et al. (2005), may represent further extension of this belt. At this stage, the lack of geochemical data for this suite precludes assignment of the tectonic environment of the magmatism for the porphyritic granites, but the San Martín tonalite and Lago Lolog granite have initial  $\epsilon\text{Ndt}$  values of about  $-4$  and Sm–Nd model ages of  $\sim 1400$  Ma (in part recalculated from Varela et al., 2005, see Table 2), consistent with a Mesoproterozoic crustal component. If this belt, at least in its northwestern part, represents a subduction-related arc in the west at this time, it further reinforces the argument above that most of the North Patagonian Massif was autochthonous to Gondwana during the Early Palaeozoic. The jump in the magmatic arc position across the North Patagonian Massif after Mid Ordovician times was presumably related in some way to the change of the regional tectonic scheme following collision of the Precordillera to the north.

## 7. Carboniferous subduction

I-type granitoids of Carboniferous to Permian age form the core of the Coastal Cordillera of Chile from  $33^\circ$  to  $38^\circ\text{S}$  where they are seen in the Cordillera de Nahuelbuta (Fig. 1). Southeasterly extension of this belt and its metamorphosed envelope has been suggested, as far as the Piedra Santa complex near the northwestern boundary of the North Patagonian Massif (Franzese, 1995), but no supporting geochronology or geochemistry are available and until now it could not be traced any further. However, our results demonstrate the existence of a 120-km-long belt of Early Carboniferous I-type granodiorites in the eastern North Patagonian Massif, along the Cordón del Serrucho, between San Carlos de Bariloche and El Maitén ( $41$ – $42^\circ 15'\text{S}$ ) (Fig. 2). Two samples yield well-defined Early Carboniferous ( $\sim$ Visean) crystallization ages of  $323 \pm 3$  and  $330 \pm 4$  Ma; a sample of the El Platero tonalite in the Río Chico, about 50 km to the southeast, gave a comparable age of  $329 \pm 4$  Ma (Fig. 6a). A two-point conventional  $^{238}\text{U}$ – $^{206}\text{Pb}$  zircon age of  $321 \pm 2$  Ma has been independently reported for a sample from the same locality as MOS-043 (Varela et

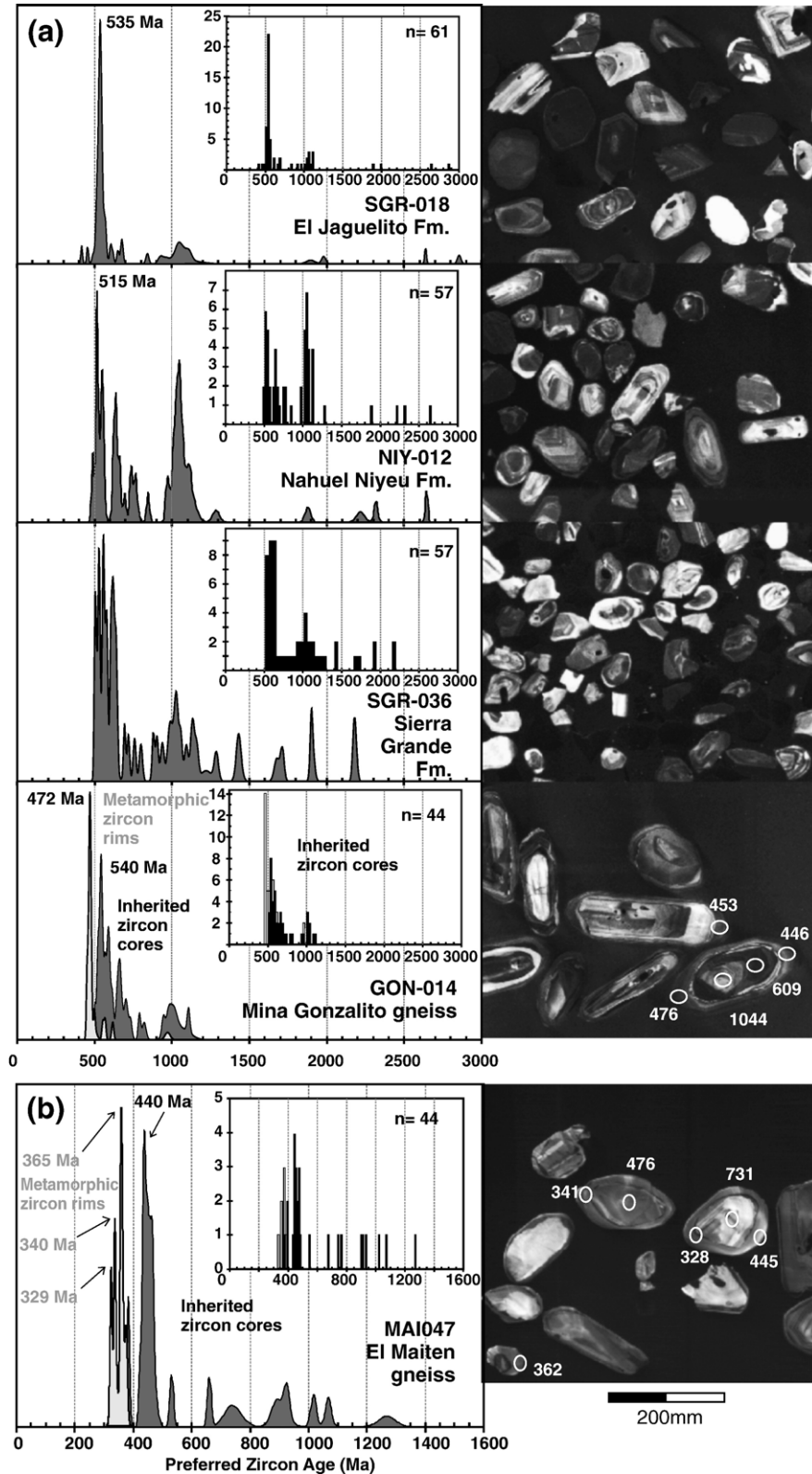
al., 2005). All three samples analysed, which are representative of continuous outcrop, are foliated metaluminous hornblende-biotite granitoids with low abundances of lithophile trace elements, rare earth element (REE) patterns typical of Andinotype calc-alkaline arc rocks, positive  $\epsilon\text{Ndt}$  values ( $+0.1$  to  $+2.8$ ), and low initial  $^{87}\text{Sr}/^{86}\text{Sr}$  ( $0.7034$ – $0.7046$ ) indicating a long-term light-REE depleted source such as the upper mantle (Table 1). Other authors have previously suggested Carboniferous magmatism in the western areas of the North Patagonian Massif, but mostly on the basis of imprecise Rb–Sr whole-rock errorochrons or K–Ar geochronological data (see Varela et al., 2005 for a summary).

## 8. Carboniferous collision-related magmatism

Two previously undated granite bodies in the southwestern North Patagonian Massif have yielded Mid Carboniferous (Serpuhkovian/Bashkirian) crystallization ages:  $314 \pm 2$  Ma from Paso del Sapo and  $318 \pm 2$  Ma from Sierra de Pichiñanes (Figs. 2 and 6b). These are both peraluminous S-type garnet-bearing leucogranites with high  $\text{SiO}_2$  contents, strongly depleted heavy-REE patterns, high  $\text{K}_2\text{O}/\text{Na}_2\text{O}$  ratios, unradiogenic  $\epsilon\text{Ndt}$  values ( $-5.0$  to  $-6.0$  on four samples) and relatively high initial  $^{87}\text{Sr}/^{86}\text{Sr}$  ( $0.7078$ – $0.7098$ ) indicative of generation by upper crustal melting (Table 2). They have multistage Sm–Nd model ages of  $\sim 1500$  Ma that could represent the age of their deep crustal source region.

These results show that in the southwestern part of the North Patagonian Massif a short-lived episode of subduction was followed by crustal anatexis during the Mid Carboniferous. The Cordón del Serrucho runs approximately N–S (Fig. 2), but the equivalence of the El Platero tonalite suggests that the arc was orientated more NW–SE, and the S-type granites occur somewhat to the northeast of this trend line. Although the wider distribution of I- and S-type granites in Patagonia is not clear, and the lack of outcrops may unfortunately prevent further elucidation, both are typically emplaced in the overriding plate above the subduction zone before and during

Fig. 4. (a) U–Pb zircon provenance age patterns for metasedimentary samples from NE Patagonia (see Fig. 2). The curves in the main diagrams are relative probability trends (Ludwig, 1999) based on the preferred age derived from individual measurements, which are also shown as inset histograms. For ages less than 1000 Ma, the  $^{238}\text{U}$ – $^{206}\text{Pb}$  age is used after correction for initial common Pb using the  $^{207}\text{Pb}$  measurement; for ages of 1000 Ma and more, the  $^{204}\text{Pb}$ -corrected  $^{207}\text{Pb}/^{206}\text{Pb}$  age is preferred. Representative portions of the cathodo-luminescence for each sample are shown on the right, with individual spot ages. The Mina Gonzalito gneiss zircons have Ordovician high-grade overgrowths on complex nuclei that have age patterns similar to the El Jagüelito Formation schist. (b) Similar plots and images for the El Maitén gneiss, which has zircon overgrowths formed during latest Devonian to Mid Carboniferous amphibolite-grade metamorphism.



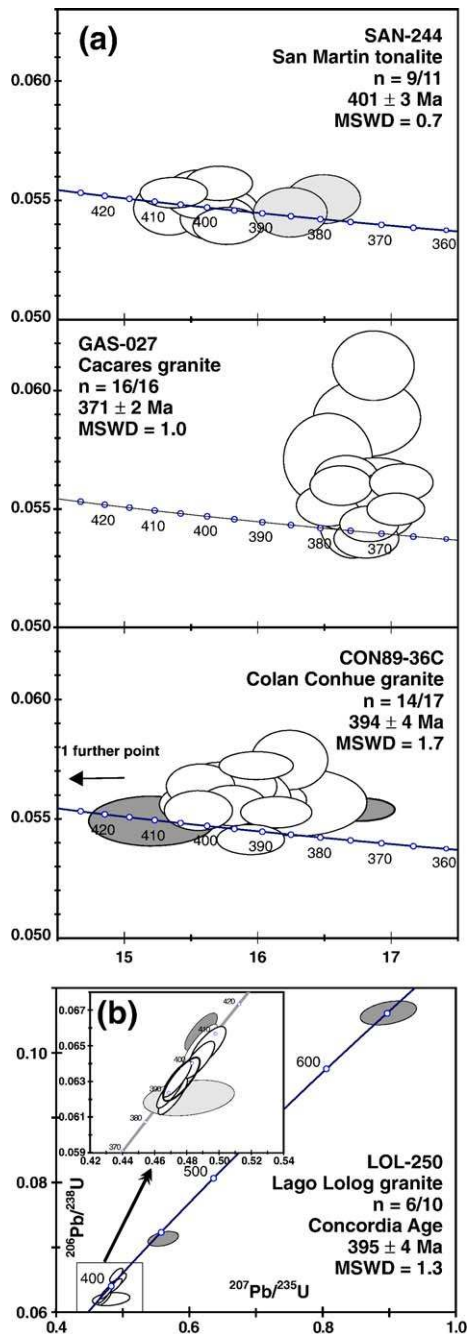
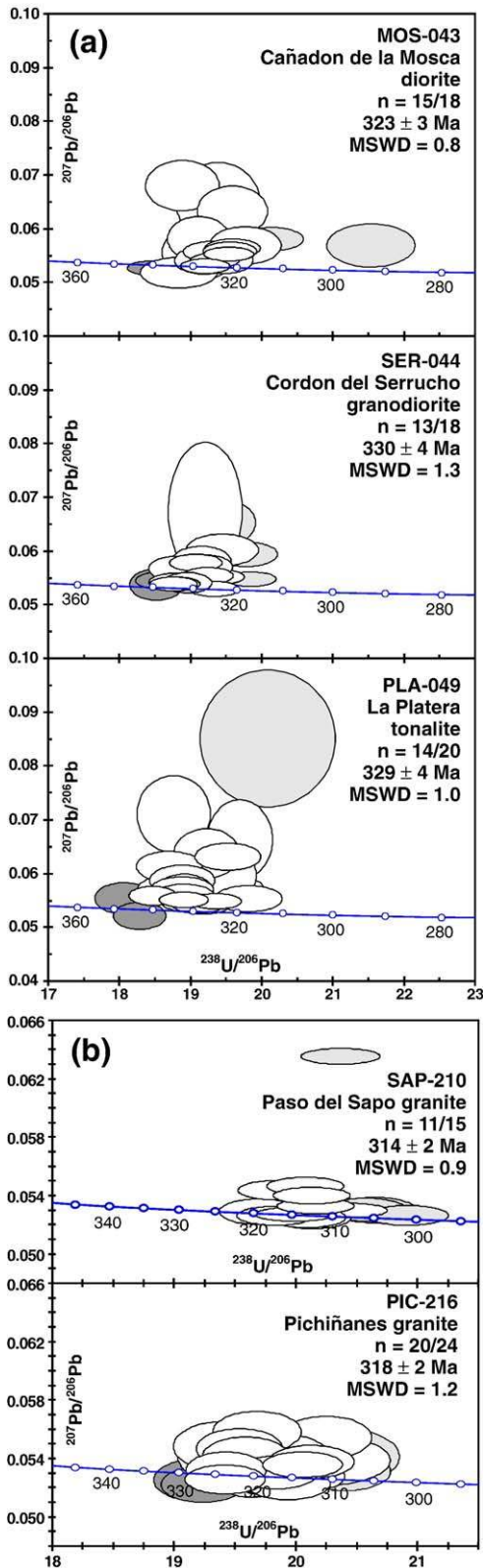


Fig. 5. New U–Pb SHRIMP zircon-dating results for Devonian granitoids in the western part of the North Patagonian Massif (Fig. 2). (a) Tera–Wasserburg diagrams for foliated tonalite from San Martín de los Andes, and porphyritic granites from Cáceres (southwest of Gastre) and Colán Conhue, details as in Fig. 3; (b)  $^{204}\text{Pb}$ -corrected data plotted in a Wetherill Concordia diagram for the Lago Lolog granite (since this sample contains zircons very rich in U, the preferred age calculation is based on six concordant data points in this diagram).

plate collision, respectively, and their disposition thus strongly suggests subduction of an oceanic plate from the southwest.

## 9. Deformation and metamorphism

The folded Lower Palaeozoic sequences of Sierra de la Ventana, the Cape region of South Africa, the Falkland Islands (in their pre-Jurassic position off the eastern Cape) and the Ellsworth Mountains of West Antarctica, consist of predominantly continental shelf deposits, commencing in the Mid Cambrian, with prominent Devonian quartzite horizons, and with Early Carboniferous diamictites often interpreted as of glacial origin. In the Sierra de la Ventana, deformation includes intense penetrative folding and cleavage and subsequent large-scale folding about approximately E–W axes, together top-to-the-southwest thrusting. Prior assumptions of a Late Permian age for deformation in both the Sierra de la Ventana and the Cape fold belt has been dependant on older publications of K–Ar data on illite and muscovite, together with the observation that the Permian strata at the top of the sequences have been deformed. However, there is notably less intense deformation in the Permian shales and sandstones at the top of the sequence, which were remagnetized within the fold and thrust belt, compared with the Lower Permian part of the sequence, which records stronger pre-folding magnetization northeastwards towards the foreland (Tomezzoli, 2001). Schematic cross-sections drawn by von Gosen et al. (1991) also illustrate the more intense thrust deformation in the Curamalal and Ventana groups below the level of an Early Permian unconformity. This was explained by proximity to the inferred collision zone, but could alternatively indicate that deformation began before this time but continued through the deposition of the Permian-to-Triassic Pillahuinco Group. The general context of deformation is compressive, although in the Cape Fold Belt dextral transpression is important. If resulting from a continental collision to the south of this belt, the main collision must have occurred in Mid Carboniferous times, as originally postulated by Ramos (1984, 1986), with less intense deformation continuing into the Permian. The timing of deformation of the Palaeozoic sedimentary series in the Cape Fold Belt, Falkland Islands and Ellsworth Mountains is mostly considered to be Permian in age, but this is also largely based on K–Ar dating of secondary mica in folded rocks, which can only be regarded as a minimum. Deformation certainly affects the Permian strata, but there is no time constraint on when it began.



Intense deformation following southward thrusting of the metasedimentary basement in the northwest part of the North Patagonian Massif has been ascribed to the Permian (von Gosen, 2003). However, the main constraint for the timing of deformation here is the age of the Navarrete granodiorite which post-dates the thrust tectonics but was affected by a late stage of NW–SE compression. We present a U–Pb zircon age of  $281 \pm 3$  Ma (15 points, MSWD=2.3) which falls well within the Early Permian (Artinskian) and requires thrust tectonics to have been completed prior to this time (Fig. 7a). In the same area, Basei et al. (2002) have reported a conventional U–Pb zircon age of  $300 \pm 6$  Ma for the Tardugno deformed granite, interpreted by von Gosen (2003) as a maximum age for thrust tectonics.

Carboniferous (?) tectono-metamorphism occurred in the southern and western parts of the North Patagonian Massif. Greenschist to lower amphibolite facies metamorphism was attained during peak deformation, forming extensive mylonite and ultramylonite ductile shear zones (Llambías et al., 2002 and references therein). As noted above, the igneous rocks of the Cordon del Serrucho are strongly foliated, and are associated with sheared and mylonitic schists (García-Sansegundo et al., 2005). A sample of amphibolite-grade paragneiss from south of El Maitén has zircon with metamorphic overgrowths recording events at 330, 340 and 365 Ma, and zircon cores with a major provenance at 440 Ma (Fig. 4b); it probably represents continental marginal sediments into which the Carboniferous arc was emplaced. Medium-pressure, garnet-bearing metapelitic schists, gneisses and migmatites of the Cushamen Formation at Río Chico (López de Luchi and Cerredo, 1997), of which this is probably an equivalent, are intruded by leucogranites with K–Ar muscovite ages of  $\sim 280$  Ma and have been proposed as higher-grade derivatives of low-grade schistose rocks at Esquel (Fig. 2) (Duhart et al., 2002). Other possibly pre-Permian gneisses and migmatites are found at Río Collon Cura, where Varela et al. (2005) recorded a U–Pb age of  $348 \pm 11$  Ma, see above, and at La Potranca, south of the Río Chubut, where altered orthopyroxene-garnet granulites host the S-type granite POT-316 ( $289 \pm 2$  Ma, Fig. 7c).

Fig. 6. New U–Pb SHRIMP zircon-dating results for Carboniferous granitoids in the southwestern part of the North Patagonian Massif (Fig. 2): (a) Tera-Wasserburg plots for three granodiorite samples representing the Early Carboniferous subduction-related arc, (b) Tera-Wasserburg diagrams for two anatectic S-type granites showing Mid Carboniferous crystallization ages. Details as for Fig. 3.

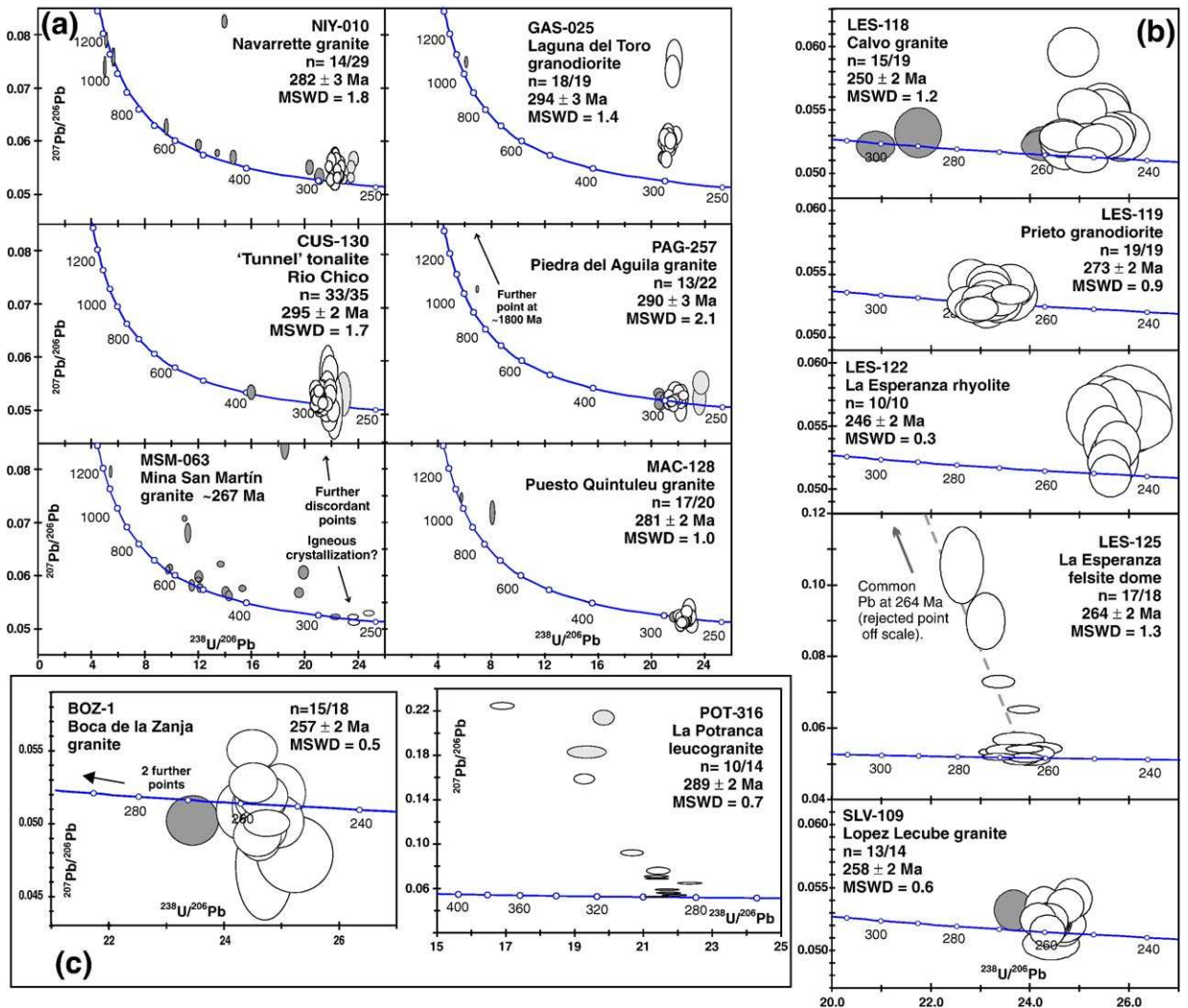


Fig. 7. New U–Pb SHRIMP zircon-dating results for samples representative of the Permian magmatism in the North Patagonian Massif: (a) granitoid rocks with significant inheritance from pre-existing Gondwana sources, including some at  $\sim 470$  Ma, 500–600 Ma and  $\sim 1000$  Ma; PAG-257 also has inherited Carboniferous zircons, (b) the granite-rhyolite complex of La Esperanza, central North Patagonian Massif, see Fig. 3 for full explanation of details, (c) two small bodies from the southern part of the North Patagonian Massif.

In contrast, the Tepuel basin in central-east Patagonia (Fig. 2, ca.  $43.5^\circ\text{S}$ ,  $70.5^\circ\text{W}$ ) has a lower section of highly deformed but low-grade Early to Mid Carboniferous metasediments, dominated by diamictites and siliciclastic rocks for which glacial, epi-glacial and marine environments are documented (Limarino et al., 1999). The upper section, located in the western sector of the basin, is also folded and comprises a deltaic sequence of conglomerates, sandstones and pelites carrying abundant Early Permian flora, and also includes also restricted levels with marine invertebrates (our unpublished U–Pb detrital zircon data are consistent with derivation predominantly from areas containing Silurian igneous rocks, such as the Deseado Massif).

## 10. Permian magmatism

Permian (and Triassic) I-type magmatism is known from the La Esperanza area in the northern part of the North Patagonian Massif on the basis of Rb–Sr whole-rock dating (Pankhurst et al., 1992; Rapela et al., 1996), and recently from the western North Patagonian Massif, where Varela et al. (2005) obtained  $^{238}\text{U}$ – $^{206}\text{Pb}$  zircon ages of  $272 \pm 2$  to  $286 \pm 13$  Ma (albeit mostly with MSWD values of 18–76 that must cast some doubt on the accuracy of their error estimates).

Our U–Pb SHRIMP results show that Early Permian granites are very extensive, in both time and space, occurring across the entire width of the massif (Figs. 2

and 7a). The oldest confirmed Permian intrusions, some of which show evidence of deformation, include the ‘Tunnel tonalite’ at Rio Chico ( $295 \pm 2$  Ma, dated at  $286 \pm 13$  Ma, MSWD=39 by Varela et al., 2005), the Laguna del Toro granodiorite near Gastre ( $294 \pm 3$  Ma, previously dated as Carboniferous by Rapela et al., 1992), a weathered two-mica granite from Piedra de Aguila ( $290 \pm 3$  Ma), and leucogranite associated with migmatite at La Potranca, south of the Río Chubut ( $289 \pm 2$  Ma). These may well have been generated during continued deformation. They are closely followed by the Navarrette granite ( $281 \pm 3$  Ma), and a muscovite migmatite west of Mamil Choique ( $281 \pm 2$  Ma). Many of these rocks, for which disparate age estimates have been previously published, belong to what has been denominated the Mamil Choique or Río Chico complex, once supposed to be Precambrian crystalline basement. Unlike the Carboniferous I-type granitoids, these characteristically contain zircon inherited from the older continental crust of Gondwana, as discrete grains or cores of polyphase crystals. We have not made a systematic study of the inherited components, but they typically have ages of  $\sim 320$  Ma,  $\sim 460$  Ma,  $500$ – $650$  Ma and  $\sim 1000$  Ma, representing the younger part of the Gondwana spectrum, with pre-Carboniferous components essentially the same as in the Cambrian metasediments of the northeastern North Patagonian Massif (Fig. 7a).

We have re-dated samples of the granites and associated volcanic rocks from La Esperanza in the northern North Patagonian Massif using the U–Pb SHRIMP zircon method (Fig. 7b). As is frequently observed, the U–Pb ages are slightly older, but consistent with the previously published Rb–Sr ages, with a total range of  $246$ – $275$  Ma. Finally, an isolated outcrop of Permian granitoid from the southeastern North Patagonian Massif, the Boca de la Zanja granodiorite from near Dolavon (Fig. 2), gave  $257 \pm 2$  Ma. The younger age range for these rocks (mostly Mid to Late Permian) than for the western and southern outcrops, suggest a northwesterly progression of magmatism during the Permian.

The composition of these Permian granitoids is quite variable (Figs. 8, 9), including rocks of both metaluminous I-type and peraluminous S-type affinities ( $60$ – $78\%$   $\text{SiO}_2$ ), and  $\epsilon\text{Nd}_t$  values range from  $-2$  (Navarrette) to  $-6$  (Mamil Choique), and even to  $-10$  (for the Donosa granite, La Esperanza, not dated in this study, but bracketed by the Prieto and Calvo granites according to field relationships), equivalent to source model ages  $1200$ – $1900$  Ma. Their initial  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios range from  $0.7036$  to  $0.7119$ . Trends in the Sr–Nd isotope diagram

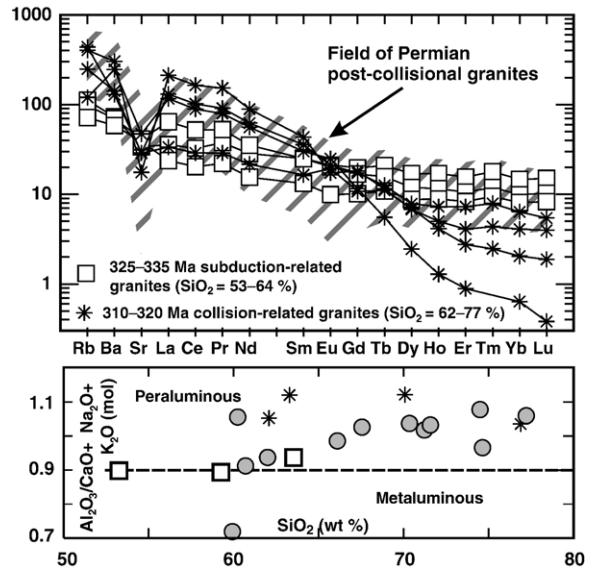


Fig. 8. Geochemical variation plots for the Carboniferous and Permian granitoids analysed in this study, showing clear distinctions in petrogenesis. The Early Carboniferous pre-collisional groups are I-type, with lithophile-element depleted signatures and primitive isotopic compositions, indicating a juvenile subduction-related source. The Mid Carboniferous collisional granites are lithophile-element rich and isotopically evolved (S-type), and have heavy REE depletion indicating anatexis at garnetiferous crustal depths. The widespread Permian, post-collisional granites (shaded circles in lower diagram) are intermediate in composition, and have patterns compatible with variable hybridization of crust- and mantle-derived magmas.

(Fig. 9) are distinct from that exhibited by the Ordovician granitoids and the Carboniferous S-type granites; whereas the latter show a shallow trend to higher initial  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios, which is characteristic of upper crustal involvement, the Permian data encompass a steeper decrease in  $\epsilon\text{Nd}_t$  values, especially in the Mid to Late Permian granites of La Esperanza, that is indicative of lower crustal components. The Permian intrusions thus appear to represent a major hybrid magmatic episode involving melting throughout the crustal section.

The North Patagonian Massif is thus a major site of Permian granitoid magmatism in southwest Gondwana. Emplacement began in earliest Permian times (ca.  $295$  Ma), gradually moved towards the northwest during the Mid Permian, and reached a climax after the main deformation of the Sierra de la Ventana fold and thrust belt. We attribute the more voluminous and widespread nature of this magmatism to major access of heat to the crust following break-off of the subducted slab after a continental collision that was initiated in Carboniferous times. In principle it extends as far north as the rhyolite volcanism of Lihue Calel in La Pampa province and

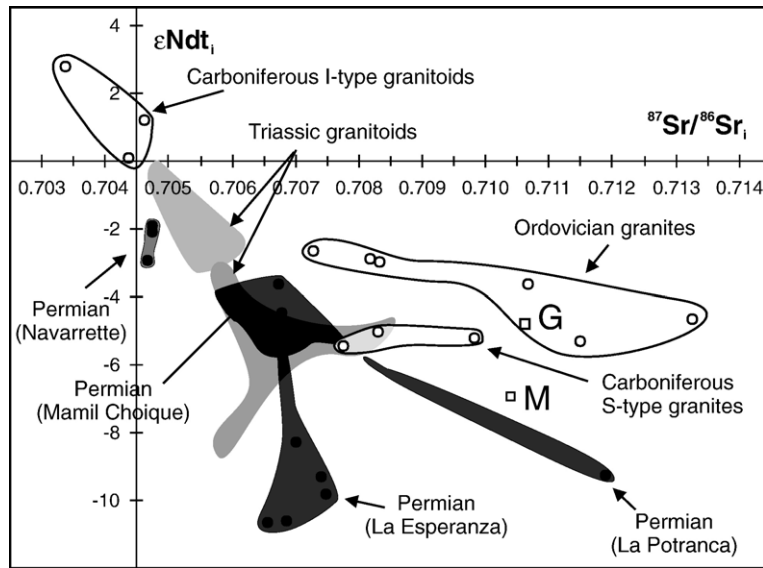


Fig. 9. Sr–Nd initial isotope composition plot for the Palaeozoic igneous rocks of Patagonia, based on the data in Table 2. The Early Carboniferous granitoids have primitive compositions largely in the long-term lithophile-element depleted quadrant, consistent with a dominant mantle-derived input. All the other granitoids have evolved initial isotope compositions indicating crustal contributions. Whereas the Ordovician and Carboniferous S-type granites follow a shallow trend suggesting young upper crustal sources with high Rb/Sr ratios, the Permian (and Triassic) granites are far more dispersed, with the steeper trend for La Esperanza granites in particular indicating older, low Rb/Sr sources associated with deeper continental crust. G=Mina Gonzalito gneiss, M=El Maitén gneiss.

Lopez Lecube syenite ( $258 \pm 2$  Ma, Fig. 7) near the Sierra de la Ventana (Fig. 1), where the alkaline character of the magmatism reflects its intraplate position, and continued into the Triassic (Rapela et al., 1996). These granites can be identified as the most important source so far recognised for the provenance of Permo-Triassic detritus in Late Palaeozoic sedimentary rocks along this part of the Pacific margin of Gondwana (Hervé et al., 2003).

## 11. Discussion: a new tectonic model

The latest information on the history of magmatism, including all new data obtained in this study, has been integrated with the sedimentary and tectonic record of the region extending from the Gondwana margin of the Sierra Ventana to southernmost Patagonia in Fig. 10. In the remainder of this section we develop our interpretation of these data in terms of the separate Early Palaeozoic evolution of southern Patagonia, largely represented by the Deseado Massif, and its collision with Gondwana in the Mid Carboniferous.

du Toit (1937) was among the first to apply the idea of continental drift to global tectonics. He recognized the essential continuity of stratigraphy and deformation in the Early Palaeozoic fold belts of Sierra de la Ventana in South America and the Cape region of South Africa,

defining his ‘Samfrau Geosyncline’. Prior to the dispersion of Gondwana as separate continental fragments, these belts would have been contiguous and collinear, and presumably extended to the Early Palaeozoic sequences of the Falkland Islands and the Ellsworth Mountains of West Antarctica. They share, at least in part, semi-continuous sedimentation from Mid Cambrian to Permian, and have several remarkable common features. With the arrival of plate tectonic theory, modern explanations of these sequences and their ‘Gondwanide’ deformation were sought, falling into two groups: those in which folding was ascribed to compression in a back-arc situation over a distant north-dipping subduction zone (Lock, 1980; Dalziel and Grunow, 1992; Trouw and de Wit, 1999; Dalziel et al., 2000), and those invoking continent–continent collision (Winter, 1984; Ramos, 1984, 1986). Cawood (2005) has proposed that this deformation represents the final stage of a long-lived ‘Terra Australis orogen’, that began in the Neoproterozoic, was developed along the entire Gondwana margin of South America, East Antarctica, and southeast Australia, and that was primarily accretional in nature, although this would allow for the accretion of small continental terrane fragments in certain places and at various times. Since the Cape Fold Belt faces the South Atlantic ocean, there is no remaining evidence of any possible colliding continental



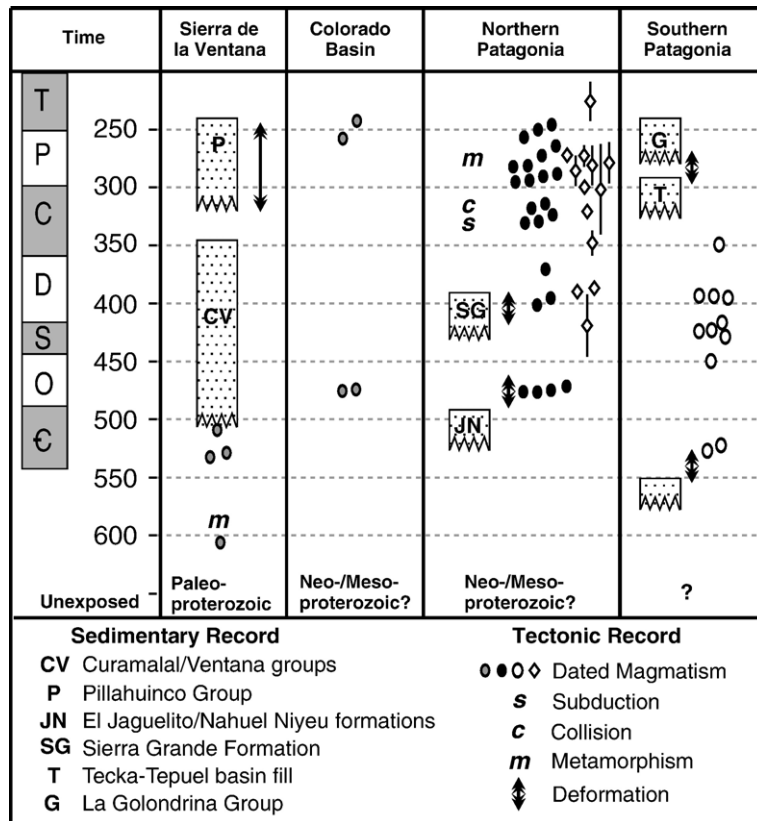


Fig. 10. Space–time diagram showing the pre-Mesozoic sedimentary, tectonic and magmatic history of the main regions of Patagonia and the adjacent Gondwana margin. Crystallization ages of granitoids represented by ellipses are from the present study (errors mostly included within the size of the symbols). Diamonds represent data from Basei et al. (2002) and Varela et al. (2005), with vertical lines indicating 2σ error bars where appropriate.

mass, but the landmass of Patagonia lies to the south of the Sierra de la Ventana.

In Fig. 11, we have attempted to demonstrate a genetic connection between the evidence presented above for Carboniferous–Early Permian collision in Patagonia with the deformation of the Gondwanide fold belts. Continental plate reconstruction of the SW margin of pre-Mesozoic Gondwana is complicated by major changes during break-up in Jurassic and Cretaceous times. These include rotation and translation of fragments such as the Falkland Plateau (Taylor and Shaw, 1989) and southern Patagonia (Vizán et al., 2005), movements on major dextral fault zones (Rapela and Pankhurst, 1992; Jacques, 2003), opening of the San Jorge and Magallanes sedimentary basins during the Cretaceous (Macdonald et al., 2003; Spalletti and Franzese, in press), growth of the westernmost areas by accretion of smaller terranes (Forsythe and Mpodozis, 1979), and crustal extension during the emplacement of granitic batholiths (Rapela et al., 2005b). Nevertheless, most recent models, e.g., Ghidella et al. (2002), concur with respect to placing the Deseado Massif much closer

to the southern tip of South Africa and the northern tip of the Antarctic Peninsula outboard of southernmost Patagonia. These factors have been incorporated as far as possible into the schematic of Fig. 11.

We suggest that the geological data summarised in this paper for the tectonic evolution Patagonia can be best explained by Mid Carboniferous collision on the southwest side of the North Patagonian Massif, resulting from ocean closure by subduction towards the northeast beneath an autochthonous Gondwana that included at least the greater part of the North Patagonian Massif. The only possible colliding crustal block is southern Patagonia, where the main exposure of Palaeozoic basement is the Deseado Massif (Fig. 11), which has a history of Neoproterozoic sedimentation and metamorphism followed by Silurian and Devonian granite magmatism (Pankhurst et al., 2003). The strongest doubt about such a collision being the main cause of deformation of the Gondwana fold belts is the relatively small size of the exposed massif, but its subsurface extension to the southeast is suggested by geophysical data showing the presence of an offshore basement high,

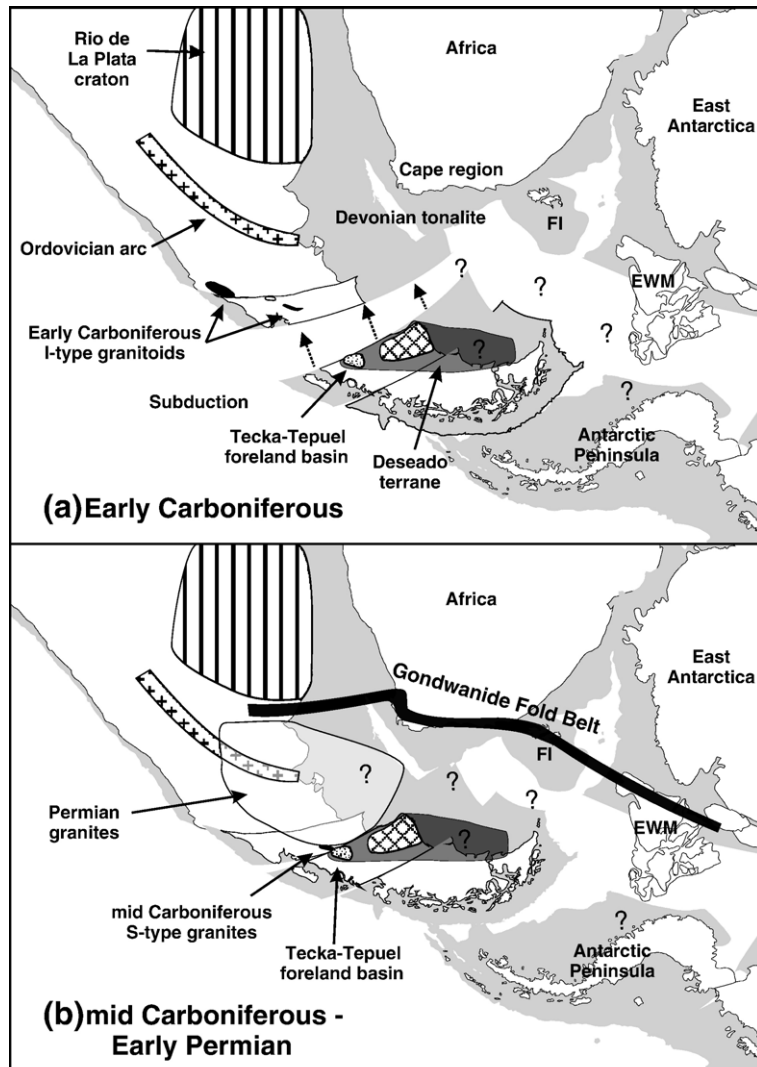


Fig. 11. Schematic reconstruction of SW Gondwana showing Late Palaeozoic plate configurations compatible with the data presented in this paper; (a) Early Carboniferous subduction stage, with the North Patagonian Massif forming part of the supercontinent since Ordovician (or at least Devonian) times, separated from a 'Deseado terrane' to the south, the true extent of which is unknown. Coastal areas of Chile and the Antarctic Peninsula consist of post-Carboniferous additions, but are shown as at the present day for the purpose of easy identification. (b) Mid Carboniferous collision stage, also showing the extent of subsequent deformation on the Gondwanide fold belts and Permian granitoid magmatism in the North Patagonian Massif. FI=Falkland Islands, EWM=Ellsworth-Whitmore mountains crustal block.

the Rio Chico-Dungeness Arch (Biddle et al., 1986). Even in the most recent model of accretionary rather collisional orogeny (Cawood, 2005), Patagonia is recognised as differing in nature from other accreted areas of the Gondwana margin, being classified as a 'peri-Gondwana continental assemblage'. Moreover, other continental basement areas in the West Antarctic collage could have been involved at this stage, viz. the Antarctic Peninsula, where the west coast shows evidence of Late Palaeozoic magmatic episodes superimposed on older metasedimentary rocks (Millar et al.,

2002). Ultimately, even if overall plate kinematic readjustments Cawood (2005) were preferred in order to explain Gondwana-wide folding, our evidence and arguments would still suggest that southern Patagonia, perhaps together with parts of West Antarctica, was continental crust to the south of the major plate boundary, so that its annexation was nevertheless an essential element in the Gondwana orogeny.

The suture zone would lie beneath the Mesozoic sediments of the San Jorge basin. The occurrence of a pre-collisional subduction-related arc and of post-

collisional anatectic granites in the southwestern part of the North Patagonian Massif also indicate that this was the upper plate and that ocean floor was subducted northeastwards beneath this active margin. The southern continental block is thus presumed to have had a passive margin, on which a foreland Mid Carboniferous–Early Permian basin represented by the Tepuel Group was formed before, during and after collision. Deformation was transmitted to the Palaeozoic sedimentary sequences of the Sierra de la Ventana, Cape region of South Africa, the Falkland Islands and the Ellsworth Mountains of West Antarctica as they were forced up against the back-stop of the Río de La Plata and Kaapvaal cratons, and continued into Mid Permian times. These relationships are illustrated in the schematic cross-sections of Fig. 12.

It is apparent that the colliding block could have consisted largely of continental crust that rifted off from the Gondwana margin in Cambrian times, and that was formerly part of the Neoproterozoic precursor of Gondwana—in tectonic terms it could be described as parautochthonous rather than allochthonous. This would be consistent with the occurrence of Cambrian magmatism similar to that of the Sierra de la Ventana in both the Cape fold belt basement and beneath the Jurassic volcanic cover of Tierra del Fuego (Söllner et al., 2000; Pankhurst et al., 2003). Separation may not have

been very great during the Cambrian–Carboniferous interval, which would explain the very short period of subduction (no more than 20 my) preceding collision.

Large-scale crustal melting as a result of slab break-off was also the mechanism invoked by Mpodozis and Kay (1990) to explain the Permian Choiyoi magmatism, which resulted in voluminous rhyolitic volcanic rocks and granites in the Frontal Cordillera and the San Rafael Massif, central western Argentina, 28–35°S. These extensive volcanic sequences unconformably overlie marine and continental Carboniferous and Early Permian deposits (Caminos and Azcuy, 1996). Mpodozis and Kay (1990) postulated collision of an unidentified exotic terrane in the west ('Terrane-X'), immediately prior to slab break-off. Geological evidence for this has never been established, although there is some evidence here also for Mid Carboniferous subduction-related magmatism, e.g., the Tabaquito pluton (Llambias and Sato, 1995) which is affected by pre-Permian fragile deformation. We suggest that an alternative to the Terrane-X hypothesis would be propagation of a progressive rupture in the slab ('slab tear-off') initially induced by the relatively small-scale Patagonian collision proposed in this paper. Detailed chronology of the magmatism of the Choiyoi volcanism that could provide support for this idea has yet to be carried out, but we note that the time-lag between the postulated collision at about

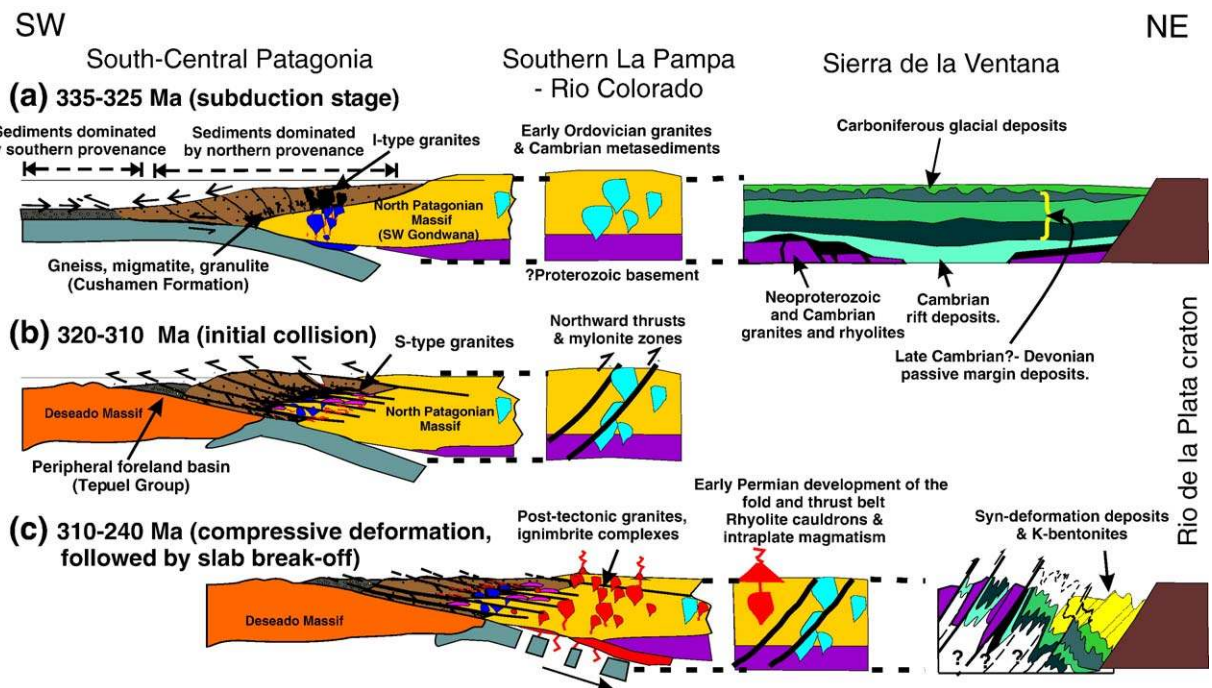


Fig. 12. Schematic cross-sections through northern Patagonia before, during and after Carboniferous continental collision according to the model advanced in Fig. 11.

320Ma and the start of Permian magmatism is comparable to that of 20–30my observed in the Himalaya after collision between India and Asia (Kohn and Parkinson, 2002).

## 12. Consequences of the new model

If the southern continental block represented by the Deseado Massif (etc.) was separated from SW Gondwana from Cambrian until Carboniferous times it could be expected to have a different Early/Mid Palaeozoic crustal history and basement evolution. There are already some notable distinctions (Pankhurst et al., 2003, and Fig. 10): i.e., it lacks Ordovician magmatism (although it was close to a source of Ordovician granite boulders by Permian times) and instead was the site of emplacement of Silurian S-type granites, which are also apparent in the detrital zircons of the Tepuel basin (our unpublished data).

It is quite possible that Cambrian rifting south of the North Patagonian Massif occurred along a pre-existing structural weakness, and thus the deep crustal structure of Patagonia south of the San Jorge basin could differ in age and origin from that to the north beneath the North Patagonian Massif. The recent discovery of Au mineralization associated with crust-derived Jurassic rhyolites in the Deseado Massif has not so far been repeated in similar rhyolites that cover the North Patagonian Massif, possibly as a result of their different deep geological composition. The flora and fauna developed during the Palaeozoic could also have followed significantly different evolutionary paths, depending on the geographical and climatic separation of the two continental areas. The Silurian–Devonian fauna of Sierra Grande in northwestern Patagonia is of Malvinokaffric type, similar to that in South Africa, the Falkland Islands and Sierra de la Ventana (Müller, 1964; Turner, 1980, Manceñido and Damborenea, 1984), which would be easily explained if this was part of the Gondwana margin at this time.

We conclude that Mid Carboniferous collision between continental areas represented by the Deseado and North Patagonian massifs was probably responsible for initial deformation of the Gondwanide fold belts, the effects of which lasted until the Mid Permian. Early Permian slab-break-off resulted in voluminous granite magmatism. Further tests for this model could come from examination of the few remaining unstudied pre-Mesozoic rocks, from deep seismic evidence for the nature of the deep crust of beneath the San Jorge basin, and further evaluation of the consequences of the model indicated above.

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## Appendix A. Supplementary data

Supplementary data associated with this article can be found, in the online version, at [doi:10.1016/j.earscirev.2006.02.001](https://doi.org/10.1016/j.earscirev.2006.02.001).

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## Gondwanide continental collision and the origin of Patagonia

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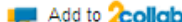



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### Abstract

A review of the post-Cambrian igneous, structural and metamorphic history of Patagonia, largely revealed by a five-year programme of U–Pb zircon dating (32 samples), geochemical and isotope analysis, results in a new Late Palaeozoic collision model as the probable cause of the Gondwanide fold belts of South America and South Africa.

In the northeastern part of the North Patagonian Massif, Cambro-Ordovician metasediments with a Gondwana provenance are intruded by Mid Ordovician granites analogous to those of the Famatinian arc of NW Argentina; this area is interpreted as Gondwana continental crust at least from Devonian times, probably underlain by Neoproterozoic crystalline basement affected by both Pampean and Famatinian events, with a Cambrian rifting episode previously identified in the basement of the Sierra de la Ventana. In the Devonian (following collision of the Argentine Precordillera terrane to the north), the site of magmatism jumped to the western and southwestern margins of the North Patagonian Massif, although as yet the tectonics of this magmatic event are poorly constrained. This was followed by Early Carboniferous I-type granites representing a subduction-related magmatic arc and Mid Carboniferous S-type granites representing crustal anatexis. The disposition of these rocks implies that the North Patagonian Massif was in the upper plate, with northeasterly subduction beneath Gondwana prior to the collision of a southern landmass represented by the Deseado Massif and its probable extension in southeastern Patagonia. This 'Deseado terrane' may have originally rifted off from a similar position during the Cambrian episode. Intense metamorphism and granite emplacement in the upper plate continued into the Early Permian. Known aspects of Late Palaeozoic sedimentation, metamorphism, and deformation in the Sierra de la Ventana and adjacent Cape Fold Belt of South Africa are encompassed within this model. It is also compatible with modern geophysical and palaeomagnetic data that do not support previous hypotheses of southward-directed subduction and collision along the northern limit of Patagonia. Subsequent Permian break-off of the subducted plate, perhaps with delamination of the lower part of the upper plate, allowed access of heat to the overlying Gondwana margin and resulted in voluminous and widespread silicic plutonism and volcanism throughout Permian and into Triassic times. Thus the new model addresses and attempts to explain three long-standing geological enigmas—the origin of the Gondwanide fold belts, the origin of Patagonia, and the cause of widespread Permian silicic magmatism (Choiyoi province) in southern South America. Differing significantly from previous models, it has new implications for the crustal structure, mineral resources, and plant and animal distribution in this part of Gondwana, since the southern landmass would have had an independent evolution throughout the Early Palaeozoic.

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**Keywords:** Patagonia; Gondwana; Sierras Australes; Cape fold belt; continental collision; U–Pb zircon

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## 1. Introduction

Patagonia, conventionally considered as the continental region south of the Río Colorado (Fig. 1), has long been recognized as differing from the rest of South America in terms of its topography, environment, flora, fauna and palaeontological record. The idea that it could have had a separate geological history (Keidel, 1925) was stimulated by the recognition of terrane accretion processes in the early 1980s. Ramos, 1984 and Ramos, 1986 proposed that an allochthonous (exotic) Patagonian terrane collided with cratonic South America (supercontinental Gondwana) along the Río Colorado zone in Carboniferous times. In his model, the suture (obscured by much younger sediments) formed by closure of a previously intervening ocean, due to southwest-dipping subduction beneath the North Patagonian Massif. In more recent reviews of the tectonic evolution of Patagonia, Ramos, 2002 and Ramos, 2004 has modified this idea to include a prior Early Palaeozoic collision within the Deseado Massif, also thought to result from southward-direct subduction, prior to Late Palaeozoic collision of the combined landmass so formed with cratonic South America, as before. In partial support of these models, Devonian–Carboniferous penetrative deformation, southward-verging folds and southward-directed thrusting of supracrustal rocks of the northeastern North Patagonian Massif was described by Chernikoff and Caminos (1996) and elaborated in a detailed structural study by von Gosen (2003), who argued for Permian rather than Carboniferous crustal shortening, and possibly a northeastward-directed accretionary process.



[Full-size image \(128K\)](#)

Fig. 1. Sketch map of Patagonia, showing the main pre-Jurassic tectonic elements, and the geographical relationship to the Sierra de la Ventana fold belt, as well as other significant basement rock exposures referred to in the text. HF = Huincul fault (Chernikoff and Zappettini, 2004).

However, geophysical data (Chernikoff and Zappettini, 2004 and Kostadinoff et al., 2005) fails to reveal any significant crustal discontinuity beneath the Río Colorado basin, and the physiographical boundary of Patagonia has been moved south to the line of the Huincul fault, which in its eastern part follows the Río Negro. Kostadinoff et al. (2005) interpret gravimetric anomalies as indicating a basement of probable Early Palaeozoic age extending yet farther to the south.

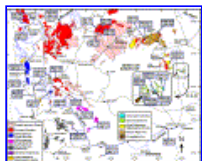
Moreover, plate subduction and collision should result in a recognized pattern of related magmatism, metamorphism and crustal melting. The original allochthonous Patagonia hypothesis suffered a set-back when it was shown that supposedly Carboniferous granites in the North Patagonian Massif were in fact Permian to Triassic in age (Pankhurst et al., 1992). All existing collision models lack supporting magmatic evidence for the postulated subduction and collision phases (i.e., subduction- and collision-related granite belts, and related metamorphism, of the right age and in the right relationship to the plate model).

We present the results of extensive study over the past 6 years of the ages and geochemistry of igneous and metamorphic rocks that form the pre-Mesozoic basement in Patagonia, and the provenance of some pre-Permian metasedimentary rocks. It must be emphasized that these outcrops are very small (often only a few hundred metres in plan) and sparsely distributed, due to the later extensive covering of Jurassic volcanic rocks, Cretaceous–Quaternary sedimentary basins, and Tertiary basaltic lavas. Nevertheless, the results, when combined with existing constraints, suggest that the North Patagonian Massif was already part of Gondwana in Ordovician times. In contrast, southern Patagonia seems to have belonged to an allochthonous entity that collided with the North Patagonian Massif in Mid Carboniferous times as a result of north-easterly ocean-floor subduction. The implications and consequences of this model are reviewed, especially in relation to the role of continental collision in the formation of the Gondwana fold belts of South America, Africa and West Antarctica, as well as the distribution of Early

Palaeozoic faunal provinces.

## 2. Analytical methods

Throughout this programme, we have dated the crystallization ages of zircon in igneous and metamorphic rocks by U–Pb geochronology using SHRIMP (Sensitive High-Resolution Ion Microprobe) technology at The Australian National University, Canberra (Williams, 1998). Some results previously published in conference proceedings have been recalculated using SQUID (Ludwig, 2000). The relevant data, together with precise sample localities, are presented as a Supplementary Appendix to this paper, and are summarised on a sketch map of northern Patagonian (Fig. 2). In addition to dating metamorphic events by analysis of metamorphic overgrowths, where possible, the detrital zircons in metasedimentary rocks have been dated to provide information on provenance of the protoliths. We have also undertaken whole-rock geochemistry (Table 1) and Sr–Nd isotope analysis (Table 2) in an attempt to constrain the tectonic environment of magmatism.



[Full-size image \(280K\)](#)

Fig. 2. Sketch map of the North Patagonian Massif, showing the pre-Cretaceous geology, the location of analysed samples, and the results of U–Pb zircon dating presented here (ages in Ma with 95% c.l. errors).

Table 1.

Geochemical data for Patagonian granites

	Subduction-related granites			Collision granites				Post-orogenic granites							
	MOS 43	SER 44	PLA 49	SAP 210	PIC 213	PIC 216	SAP 209	CUS 130	BOZ 1	GAS 25	NIY 010	MAC 128	MSM 063	POT 316	SLV 109
<i>Sample (wt.%)</i>															
SiO <sub>2</sub>	53.25	59.25	63.53	62.08	70.08	63.29	76.89	60.24	60.70	62.00	66.10	67.55	70.35	74.50	59.92
TiO <sub>2</sub>	0.93	0.81	0.44	0.72	0.39	0.81	0.23	0.90	0.68	0.91	0.33	0.48	0.30	0.01	0.85
Al <sub>2</sub> O <sub>3</sub>	18.59	17.31	17.24	17.29	15.64	17.55	11.33	17.93	15.36	16.99	16.50	16.07	14.18	14.36	14.40
Fe <sub>2</sub> O <sub>3t</sub>	8.23	6.82	4.26	5.44	2.52	5.15	1.70	5.61	4.93	5.99	3.28	3.26	2.44	1.13	5.86
MnO	0.17	0.13	0.07	0.07	0.02	0.03	0.02	0.06	0.12	0.08	0.06	0.06	0.05	0.02	0.10
MgO	4.34	3.06	2.47	2.03	0.92	2.22	0.48	2.20	3.84	2.57	1.35	1.12	0.88	0.04	2.62
CaO	7.05	6.55	5.05	5.04	3.64	4.74	1.47	6.36	4.27	5.61	2.90	3.51	1.45	1.21	3.94
Na <sub>2</sub> O	4.00	3.74	4.77	2.90	3.30	2.97	1.74	2.63	3.37	3.30	5.42	3.79	3.81	3.41	4.28
K <sub>2</sub> O	1.22	1.24	1.29	2.32	1.76	2.01	5.00	1.03	3.28	2.32	2.36	2.82	4.41	5.10	5.40
P <sub>2</sub> O <sub>5</sub>	0.14	0.23	0.10	0.22	0.13	0.33	0.04	0.20	0.17	0.24	0.14	0.17	0.13	0.04	0.46
LOI	2.10	0.89	0.74	0.98	0.80	0.85	0.85	2.10	3.00	0.40	1.12	0.44	0.82	0.07	0.78
Total	100.01	100.03	99.96	99.09	99.20	99.95	99.75	99.26	99.72	100.41	99.56	99.27	98.82	99.89	98.61
<i>Trace elements (ppm)</i>															

Cs	3.4	1.4	1	4.9	0.8	2.5	3	6	2.7	b.d.	b.d.	2.3	2.6	3.1	3.1
Rb	39	34	25	153	42	87	141	42	127	75	52	108	163	232	163
Sr	373	453	424	318	517	316	193	512	540	350	1030	449	596	286	1350
Ba	277	266	222	490	933	537	1153	415	1151	584	1050	599	1410	1157	2390
La	11.8	21.2	7.82	43	38.3	11.1	69.7	47.7	20	31.5	9.33	24.4	26.4	13.2	78.8
Ce	28.2	45	17.6	87.9	76.9	25	142	94.1	55.8	64.5	18.5	50.6	53.9	20.3	156
Pr	4.2	5.89	2.47	10.2	8.89	3.22	17.1	9.91	7.06	8.02	2.25	5.58	6.49	1.99	16.7
Nd	17.9	22.2	9.76	39.2	35.2	13.6	56	37.6	25.8	30.5	8.61	21.2	23.6	6.09	65
Sm	5.09	5.07	2.68	7.24	6.1	3.35	8.81	6.38	4.96	6.28	1.75	4.07	4.59	1.13	11.7
Eu	1.63	1.56	0.76	1.62	1.92	1.49	1.33	1.49	1.25	1.48	0.58	0.97	0.98	2.32	2.96
Gd	5.45	4.63	2.8	4.77	3.18	2.97	5.01	4.89	3.95	5	1.62	3.19	3.45	1.04	8.3
Tb	0.98	0.74	0.51	0.59	0.26	0.52	0.54	0.68	0.57	0.75	0.26	0.48	0.48	0.23	1.09
Dy	5.83	4.09	3.01	2.69	0.84	2.4	2.33	3.37	3.04	3.67	1.42	2.38	2.44	1.83	5.37
Ho	1.19	0.81	0.62	0.51	0.09	0.29	0.35	0.62	0.57	0.69	0.3	0.46	0.46	0.49	0.96
Er	3.51	2.38	1.87	1.64	0.2	0.62	0.92	1.64	1.65	1.87	0.9	1.29	1.32	2.04	2.82
Tm	0.54	0.36	0.28	0.24	b.d.	0.07	0.13	0.21	0.24	0.25	0.13	0.19	0.18	0.43	0.39
Yb	3.23	2.18	1.78	1.4	0.14	0.45	0.9	1.23	1.53	1.4	0.83	1.22	1.1	3.52	2.46
Lu	0.51	0.35	0.28	0.19	0.01	0.06	0.14	0.18	0.2	0.18	0.13	0.19	0.15	0.55	0.39
U	0.35	0.84	0.55	1.64	0.45	0.84	2.22	0.35	1.81	1.58	1.22	0.84	3.06	1.43	3.89
Th	0.53	4.07	1.32	10.7	9.32	0.46	25.5	8.11	6.6	6.72	3.94	7.95	11.7	0.46	18.7
Y	13.1	22.5	17.7	14.7	2.8	9.2	10.8	14.6	18.5	20.2	9.4	12.1	14.6	16.1	24.1
Nb	3.4	5.8	3.9	11	3.8	10.5	4.7	9.4	7.3	10.8	4.9	11.3	7.8	0.3	21.8
Zr	120	157	89	158	144	163	166	219	164	254	120	133	150	22	364
Hf	3.3	4.2	2.6	4.2	4.2	4.5	5	6.1	4.4	5.9	3	4	3.9	2	10
Ta	0.22	0.41	0.27	0.69	0.05	0.74	0.48	0.41	0.52	9.71	0.34	1.43	0.5	0.01	1.74
Sc	24	18	11	11	2	8	4	13	12	14	5	7	5	2	11
Ga	19	19	17	22	23	21	13	23	21	22	21	22	21	24	22

[Full-size table](#)

Analyses carried out by ACTLABS, Canada.

Major elements determined by ICP spectrometry.

Minor elements determined by ICP-MS spectrometry.

b.d. = below detection limit.

Table 2.

Nd Isotope data for samples from Patagonia

Sample		Age (Ma)	Sm ppm	Nd ppm	$^{147}\text{Sm}/^{144}\text{Nd}$	$^{143}\text{Nd}/^{144}\text{Nd}$	$^{143}\text{Nd}/^{144}\text{Nd}$ initial	$\epsilon_{\text{Nd}t}$	$T_{\text{DM}}$ (Ma)	Rb ppm	Sr ppm	$^{87}\text{Rb}/^{86}\text{Sr}$	$^{87}\text{Sr}/^{86}\text{Sr}$	$^{87}\text{Sr}/^{86}\text{Sr}$ initial
<i>Patagonia-metasedimentary</i>														
SGR-018	El Jaguelito Fm	530	5.24	25.91	0.1223	0.512190	0.511765	- 3.7	1533	80	101	2.305	0.728736	0.711323
NIY-012	Nahuel Niyeu Fm	500	6.29	31.99	0.1189	0.512108	0.511718	- 5.4	1636	66	89	2.147	0.728092	0.712798
GON-014	Mina Gonzalito gneiss	475	8.40	41.41	0.1226	0.512163	0.511782	- 4.8	1574	121	163	2.161	0.725244	0.710619
MAI-047	El Maitén gneiss	330	5.40	28.15	0.1159	0.512109	0.511859	- 6.9	1630	72	239	0.868	0.714467	0.710389
<i>Patagonia-igneous</i>														
SGR-016	Arroyo Salado granite	475	4.38	20.51	0.1290	0.512274	0.511873	- 3.0	1440	168	172	2.829	0.727472	0.708327
SGR-035	Sierra Grande granite	475	7.12	35.27	0.1220	0.512258	0.511878	- 2.9	1432	164	214	2.215	0.723166	0.708178
SGR-019	Punta Bahia granite	475	3.92	19.51	0.1214	0.512269	0.511891	- 2.7	1412	129	126	2.960	0.727316	0.707281
PIM-113	Río Colorado granodiorite	475	6.65	31.02	0.1297	0.512191	0.511787	- 4.7	1565	153	160	2.785	0.732104	0.713259
PIM-115	Curaco granodiorite	475	3.93	15.83	0.1502	0.512221	0.511753	- 5.3	1614	274	67	11.982	0.792570	0.711477
AB 157A	Lgo. Curruhue <sup>a</sup>	400	5.16	31.25	0.0998	0.512091	0.511830	- 5.7	1592					
LCU-251	Lgo. Curruhue gneiss	400	3.36	17.77	0.1143	0.512139	0.511840	- 5.5	1578	89	118	2.185	0.725960	0.713514
AB 13	San Martin tonalite <sup>a</sup>	400	6.43	31.63	0.1229	0.512220	0.511898	- 4.4	1493	119	324	1.063	0.712180	0.706124
AB 152	San Martin tonalite <sup>a</sup>	400	6.65	32.44	0.1239	0.512286	0.511961	- 3.2	1400	138	294	1.359	0.713970	0.706230
AB 154	C°Curruhuinca <sup>a</sup>	400	6.40	33.94	0.1140	0.512214	0.511915	- 4.0	1468					
SAN-248	San Martin tonalite (enclave)	400	5.82	27.68	0.1272	0.512222	0.511889	- 4.6	1507	182	227	2.321	0.720474	0.707256
LOL-250	Lago Lolog granite	400	3.35	16.32	0.1239	0.512360	0.512035	- 1.7	1287	121	371	0.943	0.710517	0.705144
GAS-027	Caceres granite	372	7.62	34.32	0.1343	0.512229	0.511902	- 5.0	1521	157	195	2.330	0.722228	0.709887
AB 165B	West of Sañicó <sup>a</sup>	360	6.56	29.89	0.1327	0.512030	0.511682	- 8.6	1794					
AB 165A	West of Sañicó <sup>a</sup>	360	7.44	36.63	0.1228	0.512106	0.511784	- 6.6	1655					
MOS-043	Canadon de la Mosca	323	5.26	19.97	0.1591	0.512564	0.512227	0.1	1084	39	168	0.306	0.705783	0.704374
AB 123	Cdón. Mosca <sup>a</sup>	323	4.72	18.47	0.1545	0.512519	0.512192	- 0.6	1141					
SER-044	Serrucho granodiorite	330	4.96	22.19	0.1352	0.512566	0.512274	1.2	996	34	168	0.217	0.705646	0.704625
PLA-049	Platero granodiorite	330	2.57	10.77	0.1442	0.512667	0.512356	2.8	854	25	168	0.174	0.704189	0.703373
PIC-216	Pichinanes Bio-Grt granite	318	3.30	12.01	0.1659	0.512316	0.511970	- 5.0	1485	108	168	0.956	0.712627	0.708301

SAP-209	Paso del Sapo mylonitised granite	314	5.60	35.91	0.0942	0.512161	0.511967	- 5.2	1494	144	168	2.042	0.718937	0.709812
SAP-210	Paso del Sapo foliated granodiorite	314	4.33	21.49	0.1217	0.512204	0.511954	- 5.5	1514	160	168	1.488	0.714411	0.707764
CUS-130	Rio Chico, 'Tunnel' tonalite	296	5.81	33.50	0.1048	0.512193	0.511990	- 5.2	1483	40	168	0.217	0.708522	0.707608
91 RC26	Rio Chico, 'Tunnel' tonalite <sup>a</sup>	296	4.14	18.49	0.1354	0.512360	0.512098	- 3.1	1324					
GAS-025	Laguna del Toro granodiorite	293	5.64	31.58	0.1080	0.512282	0.512075	- 3.6	1362	83	168	0.664	0.709501	0.706734
PIC-213	Pichinanes granite	289	5.37	31.68	0.1024	0.512141	0.511928	- 5.9	1546	44	168	0.241	0.710214	0.709124
POT-316	La Potranca leucogranite	289	25.96	112.82	0.1391	0.512055	0.511792	- 9.3	1765	124	168	1.201	0.716822	0.711883
NIY-010	Navarette granodiorite	281	2.49	12.02	0.1253	0.512400	0.512170	- 2.1	1233	59	168	0.166	0.705408	0.704745
NIY-011	Navarette quartz porphyry	281	2.76	12.96	0.1289	0.512363	0.512126	- 2.9	1300	86	168	0.254	0.705685	0.704670
MAC-127	Leucogranite ( La Pintada )	281	5.27	15.68	0.2031	0.512308	0.511934	- 6.7	1579	210	168	12.211	0.752954	0.704134
MAC-128	Puesto Quintulepu granodiorite	281	5.15	28.88	0.1078	0.512245	0.512047	- 4.5	1418	118	168	0.774	0.709876	0.706782
AB 27C	Comallo <sup>a</sup>	280	5.77	30.88	0.1129	0.512125	0.511918	- 7.0	1603					
LES-119	Prieto granodiorite	273	6.19	33.58	0.1113	0.512232	0.512033	- 5.0	1448	161	168	1.065	0.710794	0.706659
AB 30C	Cerro Yuncón <sup>a</sup>	270	6.12	21.67	0.1707	0.512217	0.511915	- 7.3	1618					
AB 30B	Cerro Yuncón <sup>a</sup>	270	3.80	10.10	0.2275	0.512492	0.512090	- 3.9	1368					
AB56	Loma Miranda <sup>a</sup>	270	0.64	1.67	0.2317	0.512630	0.512221	- - 1.4	1168					
AB 120	Loma Miranda <sup>a</sup>	270	5.95	34.19	0.1052	0.512215	0.512029	- 5.1	1457					
AB 121	Paso Flores <sup>a</sup>	270	10.27	34.96	0.1776	0.512241	0.511927	- 7.1	1601					
LES-120	Donoso granite	270	2.98	17.59	0.1023	0.511968	0.511787	- 9.8	1791	139	168	0.698	0.710168	0.707487
LES-125	La Esperanza rhyolite dome	264	3.85	23.65	0.0985	0.512163	0.511993	- 6.0	1516	427	168	6.500	0.731292	0.706879
MSM-063	Mina San Martin granite	260	4.78	24.25	0.1192	0.512227	0.512024	- 5.4	1476	153	168	0.757	0.710737	0.707938
SLV-109	Lopez Lecube syenite	258	11.38	64.87	0.1060	0.512200	0.512021	- 5.6	1483	169	168	0.339	0.707617	0.706371
BOZ	Boca de la Zanja	257	5.47	27.55	0.1200	0.512253	0.512051	- 5.0	1441	118	168	0.645	0.710755	0.708395
LES-118	Calvo granite	250	1.83	11.46	0.0966	0.512081	0.511923	- 7.7	1630	377	168	29.096	0.807041	0.703565
International Standards														
JB-1			5.23	26.82	0.1178	0.512790								
G-2									487	0				
GSP-1									242	1				
NBS70a	(mean of 3)								66	8				

[Full-size table](#)

Isotope analyses carried out at NERC Isotope Geosciences Laboratory, BGS, Keyworth, Nottingham (Sm and Nd by TIMS-ID, Rb and by XRF).

$^{143}\text{Nd}/^{144}\text{Nd}$  relative to 0.511864 for La Jolla,  $^{87}\text{Sr}/^{86}\text{Sr}$  relative to 0.710235 for NBS987.

$T_{\text{DM}}$  = variable crust Sm/Nd multistage.

(DePaolo, D.J., Linn, A.M., Schubert, G., 1991. The continental crustal age distribution: methods of determining mantle separation ages from Sm–Nd isotopic data and application to the Southwestern United States. *J. Geophys. Res.* B96, 2071–2088).

<sup>a</sup> Data from Varela et al. (2005).

A parallel study using conventional U–Pb zircon dating has been published recently by Varela et al. (2005), in which earlier Rb–Sr and K–Ar are also reviewed (some of which must now be regarded as of dubious reliability in terms of crystallization age but, as pointed out by these authors, more probably relate to cooling and/or metamorphic effects). As indicated by Varela et al. (2005), the main events in the evolution of Patagonia are of Palaeozoic age, and metasedimentary rock units that can be ascribed to the latest Precambrian are mainly restricted to the northeastern North Patagonian Massif.

In view of the wide time-span covered by the Palaeozoic evolution of Patagonia, it is convenient to present the new results and discuss them for each of the tectonic stages that we now recognise. The time-scale used is that of Gradstein et al. (2004).

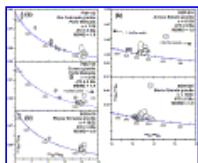
### 3. Cambrian rifting

The oldest stable area of continental basement in southern South America is the 2000–2200 Ma Río de la Plata craton (Santos et al., 2003). The Palaeozoic sedimentary sequence of Sierra de la Ventana (also known as the Sierras Australes de Buenos Aires) crops out south of the proven extent of the craton, and rests on younger (Neoproterozoic) crystalline basement with crustal anatectic granite at  $607 \pm 5$  Ma and I/A-type granites at  $531 \pm 4$  and  $524 \pm 5$  Ma (Rapela et al., 2003). These igneous events correspond to the Brasiliano/Pampean orogenic cycles of South America farther north, and may now be thought of in terms of the final assembly of Gondwana (Veevers, 2005). The culminating phase of igneous activity here is represented by Late Cambrian ( $\sim 510$  Ma) peralkaline rhyolites that were derived from lithospheric mantle, and may be interpreted as recording extensional tectonics related to rifting of this part of the Gondwana margin. Rapela et al. (2003) suggested that this rifting episode resulted in the separation of a continental fragment but could not specifically identify it within the present-day Pacific margin of West Antarctica, which is a collage of small crustal blocks (see Rapela et al., 2003, Fig. 7). However, other authors agree that a Late Cambrian passive margin was established at the southern edge of Gondwana (as elsewhere, with reference to present-day geographical coordinates) along the span of the subsequent Gondwanide fold belt, e.g., Curtis (2001). Dalziel (1997) suggested that the rifted-away portion was a large plateau area attached to Laurentia and that this subsequently collided much farther to the north in the Ordovician, leaving behind a fragment represented by the Argentine Precordillera.

### 4. Ordovician magmatism

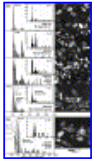
In northwest Argentina, the Early Cambrian Pampean orogenic belt lies to the west of the Río de la Plata craton, and was partially overprinted by the Early to Mid Ordovician Famatinian magmatic arc (Pankhurst et al., 1998 and Pankhurst et al., 2000), with intensive deformation in the Late Ordovician. The Famatinian belt has been considered as a continental marginal arc related to subduction during the approach and collision of the Precordillera terrane. However, recent geochronological and structural analysis suggests that Mesoproterozoic ‘Grenvillian’ rocks exposed in the Western Sierras Pampeanas may have become part of autochthonous Gondwana much earlier, being accreted to the Río de la Plata craton during the Pampean event (Escayola et al., 2005, Mulcahy et al., 2005 and Rapela et al., 2005a). Such a continental basement, thinned during the Late Cambrian rifting, seems a likely candidate for the crust underlying the Sierra de la Ventana and, perhaps, the northwestern part of the North Patagonian Massif (see below).

Previous geochronological evidence has been advanced for extension of the Famatinian magmatic belt southeastwards through La Pampa province, as far as Chadileuvú just north of the Río Colorado and on to the Arroyo Salado area (Fig. 1) in the eastern part North Patagonian Massif (Tickyj et al., 1999 and Varela et al., 1998). Our data confirm these results and establish the precise contemporaneity of magmatism in these areas, yielding indistinguishable Early Ordovician U–Pb SHRIMP ages of  $474 \pm 6$  Ma and  $475 \pm 5$  Ma for a (locally) early granodiorite and a late granite from Pichi Mahuida (Fig. 3a),  $475 \pm 6$  and  $476 \pm 4$  Ma for two granite bodies from the Arroyo Salado area, and  $476 \pm 6$  Ma for granite at Sierra Grande 25 km to the SW (Fig. 3b). Amphibolite-grade metamorphism associated with the Famatinian event in the North Patagonian Massif is recorded in quartzo-feldspathic gneiss from Mina Gonzalito (Fig. 4a), where new zircon growth during amphibolite-grade metamorphism is dated at  $472 \pm 5$  Ma (Pankhurst et al., 2001, recalculated).



Full-size image (103K)

Fig. 3. New U–Pb SHRIMP zircon-dating results for Ordovician granitoids: (a) from Pichi Mahuida, Río Colorado (see Fig. 1) and (b) the northeastern North Patagonian Massif (see Fig. 2). The data (not corrected for common Pb content) are displayed in Tera-Wasserburg diagrams with 68.4% confidence limit error ellipses. The calculated age is the weighted mean of  $^{207}\text{Pb}$ -corrected  $^{238}\text{U}$ – $^{206}\text{Pb}$  ages for the white ellipses; light grey ellipses are for spots assumed to have suffered Pb-loss, and the dark grey ellipses are for those thought to contain a significant inherited component;  $n$  = number of points in age calculation/total number of areas analysed,  $\pm 2\sigma$  errors, MSWD = mean square of weighted deviates. The derived ages of granite intrusion in these two areas are indistinguishable, and fall within the range of Famatinian granitoids in northwestern Argentina.



Full-size image (219K)

Fig. 4. (a) U–Pb zircon provenance age patterns for metasedimentary samples from NE Patagonia (see Fig. 2). The curves in the main diagrams are relative probability trends (Ludwig, 1999) based on the preferred age derived from individual measurements, which are also shown as inset histograms. For ages less than 1000 Ma, the  $^{238}\text{U}$ – $^{206}\text{Pb}$  age is used after correction for initial common Pb using the  $^{207}\text{Pb}$  measurement; for ages of 1000 Ma and more, the  $^{204}\text{Pb}$ -corrected  $^{207}\text{Pb}/^{206}\text{Pb}$  age is preferred. Representative portions of the cathodo-luminescence for each sample are shown on the right, with individual spot ages. The Mina Gonzalito gneiss zircons have Ordovician high-grade overgrowths on complex nuclei that have age patterns similar to the El Jagüelito Formation schist. (b) Similar plots and images for the El Maitén gneiss, which has zircon overgrowths formed during latest Devonian to Mid Carboniferous amphibolite-grade metamorphism.

These Ordovician granites of the North Patagonian Massif do not show signs of the deformation characteristic of the contemporaneous Famatinian granites of northwest Argentina, but since the latter is thought to be a direct result of the collision of the Precordillera terrane, this would not be expected along the entire length of the magmatic belt, and there is no evidence for terrane collision at this time in the Patagonian region. Geochemical and isotopic data (Table 1 and Table 2) show that these are equivalent to the igneous intrusions in the main part of the Famatinian belt, which are almost entirely intermediate to silicic, metaluminous, with evolved initial  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios (0.707–0.710) and  $\epsilon\text{Nd}$  values (– 2 to – 6). It is possible that the magmatism throughout the belt was initiated by post-collisional lower crustal melting following the Pampean event, modified only in the northern areas by further collisional compression.

## 5. Pre-Carboniferous sedimentation

The country rocks of these granites in northeastern Patagonia consists of fine-grained meta-sandstones—the El Jagüelito and Nahuel Niyeu formations (see Fig. 2). González et al. (2002) argue for Cambrian deposition of the former on the basis of its relationship to the granites and its trace fossil content. Detrital zircons from the El Jaguelito Formation (Fig. 4a) have a youngest age peak at  $\sim 535$  Ma consistent with this assignment: although this is strictly constraint on the maximum possible age, the dominance of this peak suggests erosion from a nearby active arc and in such cases it is commonly observed that the youngest detrital zircons effectively date deposition. That for the Nahuel Niyeu Formation is slightly younger at  $\sim 515$  Ma (both have a very few grains with even younger ages that probably reflect partial Pb-loss during Ordovician granite emplacement or deformation). The older provenance of both samples is typically Gondwanan with ages of 550–750 Ma (Pampean and Brasiliano) and  $\sim 1000$ – $1200$  Ma ('Grenvillian'), as well as a small component of older ages, including some at  $\sim 2200$  Ma. These patterns could relate to major provenance from the areas immediately to the north in Early to Mid Cambrian times, including the Río de la Plata craton (there is a very minor input of 2200–1900 Ma zircons), whereas the 'Grenvillian' component is consistent with a local basement similar to the Western Sierras Pampeanas as suggested above. The zircon cores from the Mina Gonzalito gneiss are similar, most notably in having their main provenance at 535–540 Ma (Fig. 4a), also suggesting Early Cambrian deposition and generation of the gneiss from the sandstone protolith by Famatinian-age metamorphism. Thus the metapelitic rocks of the northeastern North Patagonian Massif appear to have been originally deposited as sediments on a continental shelf at the southern margin of Gondwana.

Following intrusion of the Ordovician granites, the Sierra Grande sandstone–ironstone formation was deposited in Silurian or Early Devonian times (Limarino et al., 1999), one of the very few such deposits known after the Precambrian era. Its probable depositional environment was a quiescent platform on a stable passive margin, with no volcanic input (Spalletti, 1993). The invertebrate marine fauna is of Malvinokaffric type, similar to that found in the Falkland Islands. Palaeomagnetic data from this formation are consistent with Permian folding, but a prior pole position consistent with stable Gondwana was used to argue against it being part of a far-travelled exotic Patagonian terrane (Rapalini, 1998). A sample of ferruginous sandstone from just outside the Sierra Grande mine has a complex detrital zircon pattern (Fig. 4a) containing all the provenance elements seen in the El Jagüelito and Nahuel Niyeu formations, as well as a youngest peak at  $\sim 500$  Ma. This suggests sediment derivation from the same source areas to the north, but after the eruption of the Late Cambrian rhyolites associated with rifting. One discordant grain has an apparent age of 470 Ma, but the general absence of Ordovician zircons suggests that the Famatinian granites were not exposed in the Silurian/Early Devonian source areas.

Thus one interpretation of the NE North Patagonian Massif is that it is underlain by basement rocks that were already part of the Gondwana continent by Cambrian times, perhaps thinned during the Late Cambrian rifting event, and affected by Famatinian plutonism in the Early Ordovician, then becoming a largely passive margin until post-Early Devonian times. Its essential integrity with the Gondwana margin in the Early Palaeozoic is consistent with the observation of Kostadinoff et al. (2005) that crustal magnetic signatures are continuous across the Huincul fault zone.

## 6. Devonian magmatism

After the Ordovician, active magmatism reappeared at the western margin of the North Patagonian Massif. Varela et al. (2005) have reported conventional  $^{238}\text{U}$ – $^{206}\text{Pb}$  zircon ages of  $419 \pm 27$  Ma (MSWD = 43) and  $390.0 \pm 4.8$  Ma (MSWD = 9) for tonalites near San Martín de los Andes (Fig. 2), and of  $348 \pm 11$  Ma (MSWD = 63, possibly Carboniferous—see below) and  $386.6 \pm 5.4$  Ma (MSWD = 9.3) for deformed leucogranites cutting schists about 50 km farther to the southwest. They ascribed these to a period of Devonian magmatism and migmatization associated with the Chanic orogenic event identified in the areas north of Patagonia (San Rafael, Precordillera and Sierras Pampeanas; Sims et al., 1998). We have confirmed and refined the age of this magmatism in NW Patagonia with U–Pb zircon SHRIMP ages of  $401 \pm 3$  Ma for the San Martín tonalite (Fig. 5a) and  $395 \pm 4$  Ma for an undeformed granite at Lago Lolog about 10 km farther north (Fig. 5b). Fig. 5a also illustrates ages of  $371 \pm 4$  Ma for megacrystic granite near Gastre in the southwestern North Patagonian Massif and  $394 \pm 4$  (recalculated from Pankhurst et al., 2001) for a similar megacrystic granite at Colán Conhue 300 km southeast of Bariloche (Fig. 2) which, as suggested by Varela et al. (2005), may represent further extension of this belt. At this stage, the lack of geochemical data for this suite precludes assignment of the tectonic environment of the magmatism for the porphyritic granites, but the San Martín tonalite and Lago Lolog granite have initial  $\epsilon\text{Nd}$  values of about – 4 and Sm–Nd model ages of  $\sim 1400$  Ma (in part recalculated from Varela et al., 2005, see Table 2), consistent with a Mesoproterozoic crustal component. If this belt, at least in its northwestern part, represents a subduction-related arc in the west at this time, it further reinforces the argument above that most of the North Patagonian Massif was autochthonous to Gondwana during the Early Palaeozoic. The jump in the magmatic arc position across the North Patagonian Massif after Mid Ordovician times was presumably related in some way to the change of the regional tectonic scheme following collision of the Precordillera to the north.



[Full-size image](#) (63K)

Fig. 5. New U–Pb SHRIMP zircon-dating results for Devonian granitoids in the western part of the North Patagonian Massif (Fig. 2). (a) Tera-Wasserburg diagrams for foliated tonalite from San Martín de los Andes, and porphyritic granites from Cáceres (southwest of Gastre) and Colán Conhue, details as in Fig. 3; (b)  $^{204}\text{Pb}$ -corrected data plotted in a Wetherill Concordia diagram for the Lago Lolog granite (since this sample contains zircons very rich in U, the preferred age calculation is based on six concordant data points in this diagram).

## 7. Carboniferous subduction

I-type granitoids of Carboniferous to Permian age form the core of the Coastal Cordillera of Chile from  $33^\circ$  to  $38^\circ\text{S}$  where they are seen in the Cordillera de Nahuelbuta (Fig. 1). Southeasterly extension of this belt and its metamorphosed envelope has been suggested, as far as the Piedra Santa complex near the northwestern boundary of the North Patagonian Massif (Franzese, 1995), but no supporting geochronology or geochemistry are available and until now it could not be traced any further. However, our results demonstrate the existence of a 120-km-long belt of Early Carboniferous I-type granodiorites in the eastern North Patagonian Massif, along the Cordón del Serrucho, between San Carlos de Bariloche and El Maitén ( $41\text{--}42^\circ 15'\text{S}$ ) (Fig. 2). Two samples yield well-defined Early Carboniferous (~Visean) crystallization ages of  $323 \pm 3$  and  $330 \pm 4$  Ma; a sample of the El Platero tonalite in the Río Chico, about 50 km to the southeast, gave a comparable age of  $329 \pm 4$  Ma (Fig. 6a). A two-point conventional  $^{238}\text{U}\text{--}^{206}\text{Pb}$  zircon age of  $321 \pm 2$  Ma has been independently reported for a sample from the same locality as MOS-043 (Varela et al., 2005). All three samples analysed, which are representative of continuous outcrop, are foliated metaluminous hornblende-biotite granitoids with low abundances of lithophile trace elements, rare earth element (REE) patterns typical of Andinotype calc-alkaline arc rocks, positive  $\epsilon\text{Nd}_t$  values (+ 0.1 to + 2.8), and low initial  $^{87}\text{Sr}/^{86}\text{Sr}$  (0.7034–0.7046) indicating a long-term light-REE depleted source such as the upper mantle (Table 1). Other authors have previously suggested Carboniferous magmatism in the western areas of the North Patagonian Massif, but mostly on the basis of imprecise Rb–Sr whole-rock errorochrons or K–Ar geochronological data (see Varela et al., 2005 for a summary).



[Full-size image](#) (89K)

Fig. 6. New U–Pb SHRIMP zircon-dating results for Carboniferous granitoids in the southwestern part of the North Patagonian Massif (Fig. 2): (a) Tera-Wasserburg plots for three granodiorite samples representing the Early Carboniferous subduction-related arc, (b) Tera-Wasserburg diagrams for two anatectic S-type granites showing Mid Carboniferous crystallization ages. Details as for Fig. 3.

## 8. Carboniferous collision-related magmatism

Two previously undated granite bodies in the southwestern North Patagonian Massif have yielded Mid Carboniferous (Serpuhkovian/Bashkirian) crystallization ages:  $314 \pm 2$  Ma from Paso del Sapo and  $318 \pm 2$  Ma from Sierra de Pichiñanes (Fig. 2 and Fig. 6b). These are both peraluminous S-type garnet-bearing leucogranites with high  $\text{SiO}_2$  contents, strongly depleted heavy-REE patterns, high  $\text{K}_2\text{O}/\text{Na}_2\text{O}$  ratios, unradiogenic  $\epsilon\text{Nd}_t$  values (– 5.0 to – 6.0 on four samples) and relatively high initial  $^{87}\text{Sr}/^{86}\text{Sr}$  (0.7078–0.7098) indicative of generation by upper crustal melting (Table 2). They have multistage Sm–Nd model ages of  $\sim 1500$  Ma that could represent the age of their deep crustal source region.

These results show that in the southwestern part of the North Patagonian Massif a short-lived episode of subduction was followed by crustal anatexis during the Mid Carboniferous. The Cordón del Serrucho runs approximately N–S (Fig. 2), but the equivalence of the El Platero tonalite suggests that the arc was orientated more NW–SE, and the S-type granites occur somewhat to the northeast of this trend line. Although the wider distribution of I- and S-type granites in Patagonia is not clear, and the lack of outcrops may unfortunately prevent further elucidation, both are typically emplaced in the overriding plate above the subduction zone before and during plate collision, respectively, and their disposition thus strongly suggests subduction of an oceanic plate from the southwest.

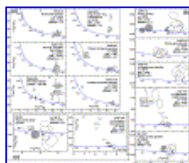
## 9. Deformation and metamorphism

The folded Lower Palaeozoic sequences of Sierra de la Ventana, the Cape region of South Africa, the Falkland Islands (in their pre-Jurassic position off the eastern Cape) and the Ellsworth Mountains of West Antarctica, consist of predominantly continental shelf deposits, commencing in the Mid Cambrian, with prominent Devonian quartzite horizons, and with Early Carboniferous diamictites often interpreted as of glacial origin. In the Sierra de la Ventana, deformation includes intense penetrative folding and cleavage and subsequent large-scale folding about approximately E–W axes, together top-to-the-southwest thrusting. Prior assumptions of a Late Permian age for deformation in both the Sierra de la Ventana and the Cape fold belt has been dependant on older publications of K–Ar data on illite and muscovite, together with the observation that the Permian strata at the top of the sequences have been deformed. However, there is notably less intense deformation in the Permian shales and sandstones at the top of the sequence, which were remagnetized within the fold and thrust belt, compared with the Lower Permian part of



the sequence, which records stronger pre-folding magnetization northeastwards towards the foreland (Tomezzoli, 2001). Schematic cross-sections drawn by von Gosen et al. (1991) also illustrate the more intense thrust deformation in the Curamalal and Ventana groups below the level of an Early Permian unconformity. This was explained by proximity to the inferred collision zone, but could alternatively indicate that deformation began before this time but continued through the deposition of the Permian-to-Triassic Pillahuinco Group. The general context of deformation is compressive, although in the Cape Fold Belt dextral transpression is important. If resulting from a continental collision to the south of this belt, the main collision must have occurred in Mid Carboniferous times, as originally postulated by Ramos, 1984 and Ramos, 1986, with less intense deformation continuing into the Permian. The timing of deformation of the Palaeozoic sedimentary series in the Cape Fold Belt, Falkland Islands and Ellsworth Mountains is mostly considered to be Permian in age, but this is also largely based on K–Ar dating of secondary mica in folded rocks, which can only be regarded as a minimum. Deformation certainly affects the Permian strata, but there is no time constraint on when it began.

Intense deformation following southward thrusting of the metasedimentary basement in the northwest part of the North Patagonian Massif has been ascribed to the Permian (von Gosen, 2003). However, the main constraint for the timing of deformation here is the age of the Navarrete granodiorite which post-dates the thrust tectonics but was affected by a late stage of NW–SE compression. We present a U–Pb zircon age of  $281 \pm 3$  Ma (15 points, MSWD = 2.3) which falls well within the Early Permian (Artinskian) and requires thrust tectonics to have been completed prior to this time (Fig. 7a). In the same area, Basei et al. (2002) have reported a conventional U–Pb zircon age of  $300 \pm 6$  Ma for the Tardugno deformed granite, interpreted by von Gosen (2003) as a maximum age for thrust tectonics.



Full-size image (156K)

Fig. 7. New U–Pb SHRIMP zircon-dating results for samples representative of the Permian magmatism in the North Patagonian Massif: (a) granitoid rocks with significant inheritance from pre-existing Gondwana sources, including some at  $\sim 470$  Ma, 500–600 Ma and  $\sim 1000$  Ma; PAG-257 also has inherited Carboniferous zircons, (b) the granite-rhyolite complex of La Esperanza, central North Patagonian Massif, see Fig. 3 for full explanation of details, (c) two small bodies from the southern part of the North Patagonian Massif.

Carboniferous (?) tectono-metamorphism occurred in the southern and western parts of the North Patagonian Massif. Greenschist to lower amphibolite facies metamorphism was attained during peak deformation, forming extensive mylonite and ultramylonite ductile shear zones (Lambias et al., 2002 and references therein). As noted above, the igneous rocks of the Cordon del Serrucho are strongly foliated, and are associated with sheared and mylonitic schists (García-Sansegundo et al., 2005). A sample of amphibolite-grade paragneiss from south of El Maitén has zircon with metamorphic overgrowths recording events at 330, 340 and 365 Ma, and zircon cores with a major provenance at 440 Ma (Fig. 4b); it probably represents continental marginal sediments into which the Carboniferous arc was emplaced. Medium-pressure, garnet-bearing metapelitic schists, gneisses and migmatites of the Cushamen Formation at Río Chico (López de Luchi and Cerredo, 1997), of which this is probably an equivalent, are intruded by leucogranites with K–Ar muscovite ages of  $\sim 280$  Ma and have been proposed as higher-grade derivatives of low-grade schistose rocks at Esquel (Fig. 2) (Duhart et al., 2002). Other possibly pre-Permian gneisses and migmatites are found at Río Collon Cura, where Varela et al. (2005) recorded a U–Pb age of  $348 \pm 11$  Ma, see above, and at La Potranca, south of the Río Chubut, where altered orthopyroxene-garnet granulites host the S-type granite POT-316 ( $289 \pm 2$  Ma, Fig. 7c).

In contrast, the Tepuel basin in central-east Patagonia (Fig. 2, ca.  $43.5^\circ\text{S}$ ,  $70.5^\circ\text{W}$ ) has a lower section of highly deformed but low-grade Early to Mid Carboniferous metasediments, dominated by diamictites and siliciclastic rocks for which glacial, epi-glacial and marine environments are documented (Limarino et al., 1999). The upper section, located in the western sector of the basin, is also folded and comprises a deltaic sequence of conglomerates, sandstones and pelites carrying abundant Early Permian flora, and also includes also restricted levels with marine invertebrates (our unpublished U–Pb detrital zircon data are consistent with derivation predominantly from areas containing Silurian igneous rocks, such as the Deseado Massif).

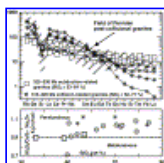
## 10. Permian magmatism

Permian (and Triassic) I-type magmatism is known from the La Esperanza area in the northern part of the North Patagonian Massif on the basis of Rb–Sr whole-rock dating (Pankhurst et al., 1992 and Rapela et al., 1996), and recently from the western North Patagonian Massif, where Varela et al. (2005) obtained  $^{238}\text{U}$ – $^{206}\text{Pb}$  zircon ages of  $272 \pm 2$  to  $286 \pm 13$  Ma (albeit mostly with MSWD values of 18–76 that must cast some doubt on the accuracy of their error estimates).

Our U–Pb SHRIMP results show that Early Permian granites are very extensive, in both time and space, occurring across the entire width of the massif (Fig. 2 and Fig. 7a). The oldest confirmed Permian intrusions, some of which show evidence of deformation, include the ‘Tunnel tonalite’ at Río Chico ( $295 \pm 2$  Ma, dated at  $286 \pm 13$  Ma, MSWD = 39 by Varela et al., 2005), the Laguna del Toro granodiorite near Gastre ( $294 \pm 3$  Ma, previously dated as Carboniferous by Rapela et al., 1992), a weathered two-mica granite from Piedra de Aguila ( $290 \pm 3$  Ma), and leucogranite associated with migmatite at La Potranca, south of the Río Chubut ( $289 \pm 2$  Ma). These may well have been generated during continued deformation. They are closely followed by the Navarrete granite ( $281 \pm 3$  Ma), and a muscovite migmatite west of Mamil Choique ( $281 \pm 2$  Ma). Many of these rocks, for which disparate age estimates have been previously published, belong to what has been denominated the Mamil Choique or Río Chico complex, once supposed to be Precambrian crystalline basement. Unlike the Carboniferous I-type granitoids, these characteristically contain zircon inherited from the older continental crust of Gondwana, as discrete grains or cores of polyphase crystals. We have not made a systematic study of the inherited components, but they typically have ages of  $\sim 320$  Ma,  $\sim 460$  Ma, 500–650 Ma and  $\sim 1000$  Ma, representing the younger part of the Gondwana spectrum, with pre-Carboniferous components essentially the same as in the Cambrian metasediments of the northeastern North Patagonian Massif (Fig. 7a).

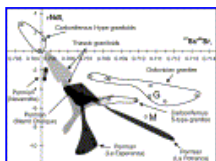
We have re-dated samples of the granites and associated volcanic rocks from La Esperanza in the northern North Patagonian Massif using the U–Pb SHRIMP zircon method (Fig. 7b). As is frequently observed, the U–Pb ages are slightly older, but consistent with the previously published Rb–Sr ages, with a total range of 246–275 Ma. Finally, an isolated outcrop of Permian granitoid from the southeastern North Patagonian Massif, the Boca de la Zanja granodiorite from near Dolavon (Fig. 2), gave  $257 \pm 2$  Ma. The younger age range for these rocks (mostly Mid to Late Permian) than for the western and southern outcrops, suggest a northwesterly progression of magmatism during the Permian.

The composition of these Permian granitoids is quite variable (Fig. 8 and Fig. 9), including rocks of both metaluminous I-type and peraluminous S-type affinities (60–78% SiO<sub>2</sub>), and  $\epsilon_{\text{Nd}}$  values range from – 2 (Navarrette) to – 6 (Mamil Choique), and even to – 10 (for the Donosa granite, La Esperanza, not dated in this study, but bracketed by the Prieto and Calvo granites according to field relationships), equivalent to source model ages 1200–1900 Ma. Their initial  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios range from 0.7036 to 0.7119. Trends in the Sr–Nd isotope diagram (Fig. 9) are distinct from that exhibited by the Ordovician granitoids and the Carboniferous S-type granites; whereas the latter show a shallow trend to higher initial  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios, which is characteristic of upper crustal involvement, the Permian data encompass a steeper decrease in  $\epsilon_{\text{Nd}}$  values, especially in the Mid to Late Permian granites of La Esperanza, that is indicative of lower crustal components. The Permian intrusions thus appear to represent a major hybrid magmatic episode involving melting throughout the crustal section.



Full-size image (47K)

Fig. 8. Geochemical variation plots for the Carboniferous and Permian granitoids analysed in this study, showing clear distinctions in petrogenesis. The Early Carboniferous pre-collisional groups are I-type, with lithophile-element depleted signatures and primitive isotopic compositions, indicating a juvenile subduction-related source. The Mid Carboniferous collisional granites are lithophile-element rich and isotopically evolved (S-type), and have heavy REE depletion indicating anatexis at garnetiferous crustal depths. The widespread Permian, post-collisional granites (shaded circles in lower diagram) are intermediate in composition, and have patterns compatible with variable hybridization of crust- and mantle-derived magmas.



Full-size image (46K)

Fig. 9. Sr–Nd initial isotope composition plot for the Palaeozoic igneous rocks of Patagonia, based on the data in Table 2. The Early Carboniferous granitoids have primitive compositions largely in the long-term lithophile-element depleted quadrant, consistent with a dominant mantle-derived input. All the other granitoids have evolved initial isotope compositions indicating crustal contributions. Whereas the Ordovician and Carboniferous S-type granites follow a shallow trend suggesting young upper crustal sources with high Rb/Sr ratios, the Permian (and Triassic) granites are far more dispersed, with the steeper trend for La Esperanza granites in particular indicating older, low Rb/Sr sources associated with deeper continental crust. G = Mina Gonzalito gneiss, M = El Maitén gneiss.

The North Patagonian Massif is thus a major site of Permian granitoid magmatism in southwest Gondwana. Emplacement began in earliest Permian times (ca. 295 Ma), gradually moved towards the northwest during the Mid Permian, and reached a climax after the main deformation of the Sierra de la Ventana fold and thrust belt. We attribute the more voluminous and widespread nature of this magmatism to major access of heat to the crust following break-off of the subducted slab after a continental collision that was initiated in Carboniferous times. In principle it extends as far north as the rhyolite volcanism of Lihue Calel in La Pampa province and Lopez Lecube syenite (258 ± 2 Ma, Fig. 7) near the Sierra de la Ventana (Fig. 1), where the alkaline character of the magmatism reflects its intraplate position, and continued into the Triassic (Rapela et al., 1996). These granites can be identified as the most important source so far recognised for the provenance of Permo-Triassic detritus in Late Palaeozoic sedimentary rocks along this part of the Pacific margin of Gondwana (Hervé et al., 2003).

## 11. Discussion: a new tectonic model

The latest information on the history of magmatism, including all new data obtained in this study, has been integrated with the sedimentary and tectonic record of the region extending from the Gondwana margin of the Sierra Ventana to southernmost Patagonia in Fig. 10. In the remainder of this section we develop our interpretation of these data in terms of the separate Early Palaeozoic evolution of southern Patagonia, largely represented by the Deseado Massif, and its collision with Gondwana in the Mid Carboniferous.



Full-size image (76K)

Fig. 10. Space–time diagram showing the pre-Mesozoic sedimentary, tectonic and magmatic history of the main regions of Patagonia and the adjacent Gondwana margin. Crystallization ages of granitoids represented by ellipses

are from the present study (errors mostly included within the size of the symbols). Diamonds represent data from [Basei et al. \(2002\)](#) and [Varela et al. \(2005\)](#), with vertical lines indicating  $2\sigma$  error bars where appropriate.

[du Toit \(1937\)](#) was among the first to apply the idea of continental drift to global tectonics. He recognized the essential continuity of stratigraphy and deformation in the Early Palaeozoic fold belts of Sierra de la Ventana in South America and the Cape region of South Africa, defining his 'Samfrau Geosyncline'. Prior to the dispersion of Gondwana as separate continental fragments, these belts would have been contiguous and collinear, and presumably extended to the Early Palaeozoic sequences of the Falkland Islands and the Ellsworth Mountains of West Antarctica. They share, at least in part, semi-continuous sedimentation from Mid Cambrian to Permian, and have several remarkable common features. With the arrival of plate tectonic theory, modern explanations of these sequences and their 'Gondwanide' deformation were sought, falling into two groups: those in which folding was ascribed to compression in a back-arc situation over a distant north-dipping subduction zone ([Lock, 1980](#), [Dalziel and Grunow, 1992](#), [Trouw and de Wit, 1999](#) and [Dalziel et al., 2000](#)), and those invoking continent–continent collision ([Winter, 1984](#), [Ramos, 1984](#) and [Ramos, 1986](#)). [Cawood \(2005\)](#) has proposed that this deformation represents the final stage of a long-lived 'Terra Australis orogen', that began in the Neoproterozoic, was developed along the entire Gondwana margin of South America, East Antarctica, and southeast Australia, and that was primarily accretional in nature, although this would allow for the accretion of small continental terrane fragments in certain places and at various times. Since the Cape Fold Belt faces the South Atlantic ocean, there is no remaining evidence of any possible colliding continental mass, but the landmass of Patagonia lies to the south of the Sierra de la Ventana.

In [Fig. 11](#), we have attempted to demonstrate a genetic connection between the evidence presented above for Carboniferous–Early Permian collision in Patagonia with the deformation of the Gondwanide fold belts. Continental plate reconstruction of the SW margin of pre-Mesozoic Gondwana is complicated by major changes during break-up in Jurassic and Cretaceous times. These include rotation and translation of fragments such as the Falkland Plateau ([Taylor and Shaw, 1989](#)) and southern Patagonia ([Vizán et al., 2005](#)), movements on major dextral fault zones ([Rapela and Pankhurst, 1992](#) and [Jacques, 2003](#)), opening of the San Jorge and Magellanes sedimentary basins during the Cretaceous ([Macdonald et al., 2003](#) and [Spalletti and Franzese, in press](#)), growth of the westernmost areas by accretion of smaller terranes ([Forsythe and Mpodozis, 1979](#)), and crustal extension during the emplacement of granitic batholiths ([Rapela et al., 2005b](#)). Nevertheless, most recent models, e.g., [Ghidella et al. \(2002\)](#), concur with respect to placing the Deseado Massif much closer to the southern tip of South Africa and the northern tip of the Antarctic Peninsula outboard of southernmost Patagonia. These factors have been incorporated as far as possible into the schematic of [Fig. 11](#).

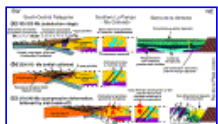


[Full-size image \(109K\)](#)

**Fig. 11.** Schematic reconstruction of SW Gondwana showing Late Palaeozoic plate configurations compatible with the data presented in this paper; (a) Early Carboniferous subduction stage, with the North Patagonian Massif forming part of the supercontinent since Ordovician (or at least Devonian) times, separated from a 'Deseado terrane' to the south, the true extent of which is unknown. Coastal areas of Chile and the Antarctic Peninsula consist of post-Carboniferous additions, but are shown as at the present day for the purpose of easy identification. (b) Mid Carboniferous collision stage, also showing the extent of subsequent deformation on the Gondwanide fold belts and Permian granitoid magmatism in the North Patagonian Massif. FI = Falkland Islands, EWM = Ellsworth-Whitmore mountains crustal block.

We suggest that the geological data summarised in this paper for the tectonic evolution Patagonia can be best explained by Mid Carboniferous collision on the southwest side of the North Patagonian Massif, resulting from ocean closure by subduction towards the northeast beneath an autochthonous Gondwana that included at least the greater part of the North Patagonian Massif. The only possible colliding crustal block is southern Patagonia, where the main exposure of Palaeozoic basement is the Deseado Massif ([Fig. 11](#)), which has a history of Neoproterozoic sedimentation and metamorphism followed by Silurian and Devonian granite magmatism ([Pankhurst et al., 2003](#)). The strongest doubt about such a collision being the main cause of deformation of the Gondwana fold belts is the relatively small size of the exposed massif, but its subsurface extension to the southeast is suggested by geophysical data showing the presence of an offshore basement high, the Rio Chico-Dungeness Arch ([Biddle et al., 1986](#)). Even in the most recent model of accretionary rather collisional orogeny ([Cawood, 2005](#)), Patagonia is recognised as differing in nature from other accreted areas of the Gondwana margin, being classified as a 'peri-Gondwana continental assemblage'. Moreover, other continental basement areas in the West Antarctic collage could have been involved at this stage, viz. the Antarctic Peninsula, where the west coast shows evidence of Late Palaeozoic magmatic episodes superimposed on older metasedimentary rocks ([Millar et al., 2002](#)). Ultimately, even if overall plate kinematic readjustments [Cawood \(2005\)](#) were preferred in order to explain Gondwana-wide folding, our evidence and arguments would still suggest that southern Patagonia, perhaps together with parts of West Antarctica, was continental crust to the south of the major plate boundary, so that its annexation was nevertheless an essential element in the Gondwana orogeny.

The suture zone would lie beneath the Mesozoic sediments of the San Jorge basin. The occurrence of a pre-collisional subduction-related arc and of post-collisional anatectic granites in the southwestern part of the North Patagonian Massif also indicate that this was the upper plate and that ocean floor was subducted northeastwards beneath this active margin. The southern continental block is thus presumed to have had a passive margin, on which a foreland Mid Carboniferous–Early Permian basin represented by the Tepuel Group was formed before, during and after collision. Deformation was transmitted to the Palaeozoic sedimentary sequences of the Sierra de la Ventana, Cape region of South Africa, the Falkland Islands and the Ellsworth Mountains of West Antarctica as they were forced up against the back-stop of the Río de La Plata and Kaapvaal cratons, and continued into Mid Permian times. These relationships are illustrated in the schematic cross-sections of [Fig. 12](#).



[Full-size image \(130K\)](#)

**Fig. 12.** Schematic cross-sections through northern Patagonia before, during and after Carboniferous continental collision according to the model advanced in [Fig. 11](#).

It is apparent that the colliding block could have consisted largely of continental crust that rifted off from the Gondwana margin in Cambrian times, and that was formerly part of the Neoproterozoic precursor of Gondwana—in tectonic terms it could be described as parautochthonous rather than allochthonous. This would be consistent with the occurrence of Cambrian magmatism similar to that of the Sierra de la Ventana in both the Cape fold belt basement and beneath the Jurassic volcanic cover of Tierra del Fuego (Söllner et al., 2000 and Pankhurst et al., 2003). Separation may not have been very great during the Cambrian–Carboniferous interval, which would explain the very short period of subduction (no more than 20 my) preceding collision.

Large-scale crustal melting as a result of slab break-off was also the mechanism invoked by Mpodozis and Kay (1990) to explain the Permian Choiyoi magmatism, which resulted in voluminous rhyolitic volcanic rocks and granites in the Frontal Cordillera and the San Rafael Massif, central western Argentina, 28–35°S. These extensive volcanic sequences unconformably overlie marine and continental Carboniferous and Early Permian deposits (Caminos and Azcuy, 1996). Mpodozis and Kay (1990) postulated collision of an unidentified exotic terrane in the west ('Terrane-X'), immediately prior to slab break-off. Geological evidence for this has never been established, although there is some evidence here also for Mid Carboniferous subduction-related magmatism, e.g., the Tabaquito pluton (Llambías and Sato, 1995) which is affected by pre-Permian fragile deformation. We suggest that an alternative to the Terrane-X hypothesis would be propagation of a progressive rupture in the slab ('slab tear-off') initially induced by the relatively small-scale Patagonian collision proposed in this paper. Detailed chronology of the magmatism of the Choiyoi volcanism that could provide support for this idea has yet to be carried out, but we note that the time-lag between the postulated collision at about 320 Ma and the start of Permian magmatism is comparable to that of 20–30 my observed in the Himalaya after collision between India and Asia (Kohn and Parkinson, 2002).

## 12. Consequences of the new model

If the southern continental block represented by the Deseado Massif (etc.) was separated from SW Gondwana from Cambrian until Carboniferous times it could be expected to have a different Early/Mid Palaeozoic crustal history and basement evolution. There are already some notable distinctions (Pankhurst et al., 2003, and Fig. 10): i.e., it lacks Ordovician magmatism (although it was close to a source of Ordovician granite boulders by Permian times) and instead was the site of emplacement of Silurian S-type granites, which are also apparent in the detrital zircons of the Tepuel basin (our unpublished data).

It is quite possible that Cambrian rifting south of the North Patagonian Massif occurred along a pre-existing structural weakness, and thus the deep crustal structure of Patagonia south of the San Jorge basin could differ in age and origin from that to the north beneath the North Patagonian Massif. The recent discovery of Au mineralization associated with crust-derived Jurassic rhyolites in the Deseado Massif has not so far been repeated in similar rhyolites that cover the North Patagonian Massif, possibly as a result of their different deep geological composition. The flora and fauna developed during the Palaeozoic could also have followed significantly different evolutionary paths, depending on the geographical and climatic separation of the two continental areas. The Silurian–Devonian fauna of Sierra Grande in northwestern Patagonia is of Malvinokaffric type, similar to that in South Africa, the Falkland Islands and Sierra de la Ventana (Müller, 1964, Turner, 1980 and Mancoff and Damborenea, 1984), which would be easily explained if this was part of the Gondwana margin at this time.

We conclude that Mid Carboniferous collision between continental areas represented by the Deseado and North Patagonian massifs was probably responsible for initial deformation of the Gondwanide fold belts, the effects of which lasted until the Mid Permian. Early Permian slab-break-off resulted in voluminous granite magmatism. Further tests for this model could come from examination of the few remaining unstudied pre-Mesozoic rocks, from deep seismic evidence for the nature of the deep crust of beneath the San Jorge basin, and further evaluation of the consequences of the model indicated above.

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## Appendix A. Supplementary data

Summary of U–Pb SHRIMP zircon analyses



(269 K)

[Supplementary\\_Table.xls](#)  [Help](#)

Microsoft Excel file 1.

Supplementary Table 1. U–Pb SHRIMP zircon data for igneous rock samples from Patagonia



(148 K)

[Supplementary\\_Table.xls](#)  [Help](#)

Microsoft Excel file 2.

Supplementary Table 2. U–Pb SHRIMP Data for metamorphic rock samples from Patagonia



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