

Gravitational couplings of intrinsic spin

This article has been downloaded from IOPscience. Please scroll down to see the full text article.

2000 Class. Quantum Grav. 17 2399

(<http://iopscience.iop.org/0264-9381/17/12/312>)

View [the table of contents for this issue](#), or go to the [journal homepage](#) for more

Download details:

IP Address: 128.206.162.204

The article was downloaded on 27/09/2010 at 14:42

Please note that [terms and conditions apply](#).

Gravitational couplings of intrinsic spin

Bahram Mashhoon

Department of Physics and Astronomy, University of Missouri-Columbia, Columbia, MO 65211, USA

Received 27 March 2000

Abstract. The gravitational couplings of intrinsic spin are briefly reviewed. A consequence of the Dirac equation in the exterior gravitational field of a rotating mass is considered in detail, namely, the difference in the energy of a spin- $\frac{1}{2}$ particle polarized vertically up and down near the surface of a rotating body is $\hbar\Omega \sin \theta$. Here θ is the latitude and $\Omega = 2GJ/(c^2R^3)$, where J and R are, respectively, the angular momentum and radius of the body. It seems that this relativistic quantum gravitational effect could be measurable in the foreseeable future.

PACS numbers: 0420C, 0365, 0480C

1. Introduction

About 40 years ago, Kobzarev and Okun [1] considered the theoretical possibility that a nuclear particle may possess a gravitoelectric dipole moment. This would lead to a violation of the equivalence principle through an interaction of the form $H_{\text{int}} = A\boldsymbol{\sigma} \cdot \mathbf{g}$, where A is an amplitude, $\boldsymbol{\sigma}$ is the particle spin and \mathbf{g} is the gravitational acceleration due to a massive body such as the Earth. Similar spin-gravitoelectric couplings of the form $f(r)\boldsymbol{\sigma} \cdot \hat{\mathbf{r}}$ have been considered by a number of authors in connection with the possible breakdown of parity and time-reversal invariance in gravitation [2]. Leitner and Okubo used the hyperfine splitting of the ground state of hydrogen to set an upper limit on the strength of such an interaction [2]. Meanwhile, Dabbs *et al* [3] studied the freefall of neutrons polarized vertically up and down in the gravitational field of the Earth and found no splitting in the gravitational acceleration greater than a few per cent of g . However, a few years later observational evidence was reported for the gravitoelectric dipole moment of the proton [4]. This was soon shown to be spurious by the experiments of Vasil'ev [5] and Young [6]. In particular, Young [6] placed an upper limit of 0.3 Hz on the gravity shift of the proton Larmor frequency in 1969. Finally, a significant upper limit of 10^{-4} Hz was placed on a possible shift of the deuteron Larmor frequency due to the Earth's gravitational field by Wineland and Ramsey [7] in 1972.

In 1989, the observation of an anomalous difference in the weight of mechanical gyroscopes rotating vertically upward and downward was reported [8]. Again, the existence of such a rotor weight change was soon contradicted by subsequent experiments [9].

The observational search for the role of intrinsic spin in the gravitational interaction as well as the spacetime torsion has continued and many significant experiments have been performed [10–16]. These experiments have also explored finite-range axion-like interactions, which could be of the $\hat{\mathbf{r}} \cdot \boldsymbol{\sigma}$ ('monopole–dipole') form as well as a linear combination of $\boldsymbol{\sigma}_A \cdot \boldsymbol{\sigma}_B$ and $\hat{\mathbf{r}} \cdot \boldsymbol{\sigma}_A \hat{\mathbf{r}} \cdot \boldsymbol{\sigma}_B$ ('dipole–dipole') form, and have placed useful restrictions on the parameters of

such interactions. Indeed, the past few decades have witnessed the emergence of extremely precise measurement techniques [17] that make it possible to detect frequency shifts of the order of 10^{-9} Hz, an improvement of five orders of magnitude over what was possible three decades ago [7].

The aim of the present paper is to discuss the gravitomagnetic coupling of intrinsic spin due to the fact that according to the standard theory a spinning particle possesses a gravitomagnetic dipole moment. This moment couples to the gravitomagnetic field of a rotating mass (such as the Earth) in complete analogy with the $-\boldsymbol{\mu} \cdot \boldsymbol{B}$ interaction in electrodynamics. Instead of treating the Dirac equation in the exterior gravitational field of a rotating mass, a heuristic derivation of this general interaction is given in sections 2 and 3 on the basis of the gravitational Larmor theorem. For a spin- $\frac{1}{2}$ particle near the surface of the Earth, the effect involves a frequency shift of the order of 10^{-14} Hz. Section 4 contains a brief discussion of the prospects for the measurement of this relativistic quantum gravitational effect.

2. Inertia of intrinsic spin

Imagine an observer in a laboratory on the Earth using Earth-based coordinate axes to describe the results of measurements. The particles involved in the experiments on the rotating Earth are waves propagating in inertial spacetime and it is natural to assume that they would keep their polarization aspects fixed in the underlying inertial frame. As measured by the observer, however, such intrinsic spin must ‘precess’ in a sense opposite to the sense of rotation of the Earth. The Hamiltonian associated with such motion would be of the form $H = -\boldsymbol{\sigma} \cdot \boldsymbol{\Omega}$, where $\boldsymbol{\Omega}$ is the frequency of rotation of the laboratory frame. The existence of such a Hamiltonian would show that intrinsic spin has rotational inertia. In quantum mechanics, mass and spin characterize the irreducible unitary representations of the inhomogeneous Lorentz group. The inertial properties of mass are well known in classical mechanics through various translational and rotational acceleration effects. In quantum mechanics, the inertial properties of mass have been investigated experimentally by a number of authors [18]. It is therefore interesting to consider the inertial properties of spin [19].

The coupling of intrinsic spin with rotation indicated above may be illustrated by a simple example. Imagine an observer rotating counterclockwise with uniform frequency Ω about the direction of propagation of a plane linearly polarized monochromatic electromagnetic wave of frequency $\omega \gg \Omega$. For instance, the observer could be in an Earth-based laboratory and Ω would then be the frequency of the proper rotation of the Earth. We neglect gravitational effects in this section and consider all phenomena in a global inertial frame in Minkowski spacetime. Let the observer move on a circle of radius r with speed $c\beta = r\Omega$ in the (x, y) -plane of the inertial frame and let the electric field of the wave be given by the real part of

$$\boldsymbol{E} = E_0 \hat{\boldsymbol{x}} e^{-i\omega t + ikz}, \quad (1)$$

where E_0 is a constant amplitude, $\boldsymbol{k} = k\hat{\boldsymbol{z}}$ is the wavevector and $\omega = ck$. From the viewpoint of the rotating observer, the direction of linear polarization that is fixed in the inertial frame must drift in a clockwise sense about the direction of propagation, i.e.

$$\boldsymbol{E} = E_0 (\cos \Omega t \hat{\boldsymbol{x}}' - \sin \Omega t \hat{\boldsymbol{y}}') e^{-i\omega t + ikz}, \quad (2)$$

where $\hat{\boldsymbol{x}}' = \hat{\boldsymbol{x}} \cos \Omega t + \hat{\boldsymbol{y}} \sin \Omega t$, $\hat{\boldsymbol{y}}' = -\hat{\boldsymbol{x}} \sin \Omega t + \hat{\boldsymbol{y}} \cos \Omega t$ and $\hat{\boldsymbol{z}}' = \hat{\boldsymbol{z}}$ denote the Cartesian coordinate axes in the rotating frame of the observer. Specifically, the two coordinate systems are related by a simple rotation such that

$$\hat{\boldsymbol{x}} + i\hat{\boldsymbol{y}} = e^{\pm i\Omega t} (\hat{\boldsymbol{x}}' \pm i\hat{\boldsymbol{y}}'). \quad (3)$$

The linearly polarized wave (1) is a coherent superposition of a right circularly polarized (RCP) wave and a left circularly polarized (LCP) wave, i.e.

$$\mathbf{E} = \frac{1}{2}E_0 (\hat{\mathbf{x}} + i\hat{\mathbf{y}}) e^{-i\omega t + ikz} + \frac{1}{2}E_0 (\hat{\mathbf{x}} - i\hat{\mathbf{y}}) e^{-i\omega t + ikz}. \quad (4)$$

From the viewpoint of the rotating observer, these eigenstates of the radiation field remain invariant,

$$\mathbf{E} = \frac{1}{2}E_0 (\hat{\mathbf{x}}' + i\hat{\mathbf{y}}') e^{-i(\omega - \Omega)t + ikz} + \frac{1}{2}E_0 (\hat{\mathbf{x}}' - i\hat{\mathbf{y}}') e^{-i(\omega + \Omega)t + ikz}, \quad (5)$$

except that the frequency of the RCP component is perceived to be $\omega - \Omega$, while that of the LCP wave is perceived to be $\omega + \Omega$ with respect to inertial time t . The proper time of the observer is, however, $\tau = t/\gamma$, where $\gamma = (1 - \beta^2)^{-1/2}$. Thus we find that the proper frequencies measured by the observer are

$$\omega' = \gamma(\omega \mp \Omega). \quad (6)$$

Here the Lorentz factor accounts for time dilation, which is all that should happen according to the transverse Doppler effect. Instead, we have in (6) the additional ‘angular Doppler terms’ $\mp\Omega$ that have the following physical origin: in an RCP (LCP) wave, the electric and magnetic fields rotate in the positive (negative) sense about the direction of propagation with frequency ω . Since the observer rotates in the positive sense with frequency Ω , it perceives the effective frequency of the RCP (LCP) wave to be $\omega - \Omega$ ($\omega + \Omega$). In the JWKB limit, $\omega \rightarrow \infty$ and the ‘angular Doppler terms’ disappear since $\mp\Omega/\omega \rightarrow 0$. Our heuristic treatment ignores certain relativistic corrections that are not essential for the purposes of this discussion. Writing equation (6) in terms of the photon energy as $E' = \gamma(E \mp \hbar\Omega)$, we see that the deviation from the simple transverse Doppler effect is due to the coupling of the spin of a circularly polarized photon to the rotation of the observer, since a RCP (LCP) photon carries an intrinsic spin of \hbar ($-\hbar$) along its direction of propagation [20]. These elementary considerations already contain the basic aspects of the phenomenon of spin–rotation coupling, as can be seen from the following discussion based on the theory of relativity [21].

The special theory of relativity consists of two main elements: the principle of relativity (i.e. Lorentz invariance) and the hypothesis of locality. The latter specifies what an accelerated observer measures by establishing a connection between the accelerated observer and an inertial observer. Indeed, it requires that an accelerated observer be at each instant locally equivalent to a momentarily comoving inertial observer. This is a non-trivial axiom since there exist definite acceleration scales of time and length that are associated with an accelerated observer. In the case under consideration, for example, the acceleration length of the rotating observer is $\mathcal{L} = c/\Omega$ and the corresponding temporal scale is $\mathcal{L}/c = \Omega^{-1}$. Moreover, an elementary application of the hypothesis of locality would imply that $\omega' = \gamma\omega$ by the transverse Doppler effect, since the connection between the instantaneous inertial frame of the accelerated observer and our global inertial frame simply results in the standard Doppler and aberration formulae with a time-dependent velocity $c\beta(t)$. On the other hand, it should be clear that to measure wave characteristics such as the frequency, one must observe at least a few periods of the oscillations of the wave before a determination of the frequency becomes even possible. In this way, the curvature of the observer’s worldline would have to be taken into consideration, and hence the standard Doppler and aberration formulae of relativity theory are valid only to the extent that the period of the wave $T = 2\pi/\omega$ is negligible compared to Ω^{-1} , i.e. $\Omega T = 2\pi\Omega/\omega \rightarrow 0$. In view of these remarks, it is therefore natural to apply the locality axiom only to the electromagnetic field; then, the measured field could be Fourier analysed, which is a non-local operation, to obtain the frequency and wavevector content of the field in the accelerated frame. This is indeed the physical basis for the extension of relativistic

wave equations to accelerated systems; in fact, this extended hypothesis of locality for wave phenomena is equivalent to the assumption of minimal coupling. Using this approach, one finds that for $\omega \gg \Omega$, the standard Doppler and aberration formulae should be modified to

$$\omega' = \gamma(\omega - c\boldsymbol{\beta} \cdot \mathbf{k}) - \gamma \hat{\mathbf{H}} \cdot \boldsymbol{\Omega}, \quad (7)$$

$$\mathbf{k}' = \mathbf{k} + (\gamma - 1)(\hat{\boldsymbol{\beta}} \cdot \mathbf{k})\hat{\boldsymbol{\beta}} - \frac{1}{c}\gamma\omega\boldsymbol{\beta} + \frac{1}{c}\gamma(\hat{\mathbf{H}} \cdot \boldsymbol{\Omega})\boldsymbol{\beta}, \quad (8)$$

where $\hat{\mathbf{H}} = \pm ck/\omega$ is the unit helicity vector. One can then consider optical interferometry in a rotating frame that would be based on the spin of the photon in contrast to the Sagnac effect that is connected to its orbital angular momentum [21].

The general expression for spin–rotation coupling can be written as

$$E' = \gamma(E - \hbar M \Omega), \quad (9)$$

where M is the total (orbital plus spin) ‘magnetic’ quantum number along the axis of rotation; that is, $M = 0, \pm 1, \pm 2, \dots$ for a scalar or a vector field, while $M \mp \frac{1}{2} = 0, \pm 1, \pm 2, \dots$ for a Dirac field. In the JWKB approximation, equation (9) can be written as $E' = \gamma(E - \boldsymbol{\Omega} \cdot \mathbf{J})$, where $\mathbf{J} = \mathbf{L} + \mathbf{S} = \mathbf{r} \times \mathbf{P} + \mathbf{S}$. Thus $E' = \gamma(E - \mathbf{v} \cdot \mathbf{P}) - \gamma \mathbf{S} \cdot \boldsymbol{\Omega}$, so that in the absence of intrinsic spin we recover the classical expression for the energy of a particle as measured in the rotating frame with $\mathbf{v} = \boldsymbol{\Omega} \times \mathbf{r}$. The energy and momentum of a spinning particle as measured by an accelerated observer are then

$$E' = \gamma(E - \mathbf{v} \cdot \mathbf{P} - \mathbf{S} \cdot \boldsymbol{\Omega}), \quad (10)$$

$$\mathbf{P}' = \mathbf{P} + (\gamma - 1)(\mathbf{P} \cdot \hat{\boldsymbol{\beta}})\hat{\boldsymbol{\beta}} - \frac{1}{c}\gamma E \boldsymbol{\beta} + \frac{1}{c}\gamma(\mathbf{S} \cdot \boldsymbol{\Omega})\boldsymbol{\beta}, \quad (11)$$

using the same JWKB approach as in the derivation of equations (7) and (8). It follows that

$$E'^2 - c^2 P'^2 = m^2 c^4 - 2E(\mathbf{S} \cdot \boldsymbol{\Omega}) + (\mathbf{S} \cdot \boldsymbol{\Omega})^2. \quad (12)$$

These results reduce to the equations appropriate for light once we set $E = \hbar\omega$, $\mathbf{P} = \hbar\mathbf{k}$, $\mathbf{S} = \hbar\hat{\mathbf{H}}$ and $m = 0$.

Experimental evidence for helicity–rotation coupling exists in the microwave and optical regimes via the phenomenon of frequency shift of polarized radiation [19]. Moreover, there is observational evidence for the coupling of spin- $\frac{1}{2}$ particles with the rotation of the Earth [19]. The analogous gravitational coupling of intrinsic spin is considered in the next section.

3. Spin–gravitomagnetic coupling

To extend the physics of spin–rotation coupling to the gravitational field, one must resort to Einstein’s heuristic principle of equivalence. It is possible to interpret this principle in the post-Newtonian approximation via the gravitational Larmor theorem [22]. Newton’s law of gravitation is formally analogous to Coulomb’s law of electricity; therefore, one may describe Newtonian gravitational effects in terms of a gravitoelectric field. The classical tests of general relativity are all due to post-Newtonian gravitoelectric corrections. However, any consistent framework that brings Newtonian gravitation and Lorentz invariance together must of necessity contain a gravitomagnetic field that would be due to a mass current. A direct measurement of the gravitomagnetic field of the Earth via the precession of superconducting gyroscopes in a polar orbit about the Earth is one of the goals of the Stanford gyroscope experiment (GP-B) planned for 2001.

In the linear approximation of general relativity, where gravitational effects are treated as linear perturbations in a global inertial frame in Minkowski spacetime, one can express the gravitational field equations as Maxwell's equations for the gravitoelectric field \mathbf{E}_g and the gravitomagnetic field \mathbf{B}_g once $O(c^{-4})$ terms are neglected in the post-Newtonian metric perturbations [22]. Specifically, we let $g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$, where $\eta_{\mu\nu}$ is the Minkowski metric and for the linear perturbation $h_{\mu\nu}$ we define $\bar{h}_{\mu\nu} = h_{\mu\nu} - \frac{1}{2}\eta_{\mu\nu}h^\alpha_\alpha$. Then $\bar{h}^{00} = 4\phi_g/c^2$, $\bar{h}^{0i} = 2A_g^i/c^2$ and $\bar{h}^{ij} = O(c^{-4})$. Here $\phi_g(t, \mathbf{x})$ is the gravitoelectric potential and $\mathbf{A}_g(t, \mathbf{x})$ is the gravitomagnetic vector potential. That is, of the ten effective gravitational potentials $\bar{h}_{\mu\nu}$ in general relativity, we neglect the six spatial potentials \bar{h}_{ij} as these are of $O(c^{-4})$ for non-relativistic (astronomical) sources and from the remaining four potentials one can construct a consistent theory of gravitoelectromagnetism (GEM) in this approximation scheme. Let us note that $\bar{h}^{0\mu} = 2c^{-2}(2\phi_g, \mathbf{A}_g)$, so that the Lorentz gauge condition $\bar{h}^{\mu\nu}{}_{,\nu} = 0$ reduces in this case to

$$\frac{2}{c} \frac{\partial \phi_g}{\partial t} + \nabla \cdot \mathbf{A}_g = 0. \tag{13}$$

Thus $A^\mu = (2\phi_g, \mathbf{A}_g)$ is the effective GEM potential and the spacetime metric is given by

$$ds^2 = -c^2 \left(1 - \frac{2\phi_g}{c^2}\right) dt^2 - \frac{4}{c} (\mathbf{A}_g \cdot d\mathbf{x}) dt + \left(1 + \frac{2\phi_g}{c^2}\right) \delta_{ij} dx^i dx^j. \tag{14}$$

The analogy with electrodynamics turns out to be exact, except for the fact that the ratio of the gravitomagnetic charge to the gravitoelectric charge is two, $q_B/q_E = 2$; that is, linear gravity is a spin-2 field in contrast to the spin-1 character of the electromagnetic field that implies $q_B/q_E = 1$ for the Maxwell theory.

In electrodynamics, Larmor established a theorem regarding the local equivalence of magnetism and rotation for all charged particles with the same charge-to-mass ratio q/m . In fact, the electromagnetic field can be locally replaced by an accelerated frame with translational acceleration $\mathbf{a}_L = -(q/m)\mathbf{E}$ and rotational (Larmor) frequency $\omega_L = q\mathbf{B}/(2mc)$. In electrodynamics, q/m can be positive, zero or negative; however, the gravitational charge-to-mass ratio is universal due to the experimentally well tested equivalence of gravitational and inertial masses. This leads directly to Einstein's principle of equivalence and hence a geometric theory of gravitation. Einstein's heuristic principle of equivalence traditionally involves the local equivalence of the gravitoelectric field with the translational acceleration of the 'Einstein elevator' in Minkowski spacetime. The interpretation of Einstein's principle in terms of the gravitational Larmor theorem would then involve, in addition, the local equivalence of the gravitomagnetic field with the Larmor rotation of the elevator as well.

Let us consider the exterior field of an almost spherical rotating astronomical body (such as the Earth) with GEM potentials

$$\phi_g \simeq \frac{GM}{r}, \quad \mathbf{A}_g \simeq \frac{G}{c} \frac{\mathbf{J} \times \mathbf{r}}{r^3}, \tag{15}$$

where M is the mass and \mathbf{J} is the angular momentum of the source. These potentials can be obtained from the electromagnetic analogy by assuming that the source has positive gravitoelectric charge $Q_E = M$ and gravitomagnetic charge $Q_B = 2M$. The GEM fields are then

$$\mathbf{E}_g = -\nabla\phi_g - \frac{1}{2c} \frac{\partial}{\partial t} \mathbf{A}_g, \quad \mathbf{B}_g = \nabla \times \mathbf{A}_g. \tag{16}$$

The motion of test particles in the gravitational field of a rotating mass can be obtained from the Lorentz force law if we assume that for a test particle of inertial mass m the gravitational

charges are negative, i.e. $q_E = -m$ and $q_B = -2m$, in order to take due account of the dominant gravitational attraction between the test particle and the source. It turns out that an ideal test gyroscope at rest outside the rotating source undergoes gravitomagnetic precession

$$\frac{d\mathbf{S}}{dt} = \boldsymbol{\Omega}_P \times \mathbf{S} \quad (17)$$

with frequency

$$\boldsymbol{\Omega}_P = \frac{1}{c} \mathbf{B}_g = \frac{GJ}{c^2 r^3} [3(\hat{\mathbf{r}} \cdot \hat{\mathbf{J}})\hat{\mathbf{r}} - \hat{\mathbf{J}}]. \quad (18)$$

Imagine now that we replace the gravitomagnetic field by a rotating frame in the neighbourhood of the gyroscope. As referred to observers at rest in the rotating frame, the motion of the gyroscope would be the same as before if the observers rotate with Larmor frequency $\boldsymbol{\omega}_L = -\boldsymbol{\Omega}_P$. This relation is consistent with the Larmor formula $\boldsymbol{\omega}_L = q\mathbf{B}/(2mc)$ once we set $q_B = -2m$ and $\mathbf{B}_g = c\boldsymbol{\Omega}_P$ as in equation (18). Thus a consistent and complete gravitoelectromagnetic formalism can be developed along these lines [22].

In particular, the spin-rotation coupling can be extended to gravitomagnetism via the Larmor theorem with $\boldsymbol{\omega}_L = -\boldsymbol{\Omega}_P$. That is, the interaction of intrinsic spin with the gravitomagnetic field is given by the Hamiltonian $H = \boldsymbol{\sigma} \cdot \boldsymbol{\Omega}_P$, since this interaction in the Larmor frame would be $H = -\boldsymbol{\sigma} \cdot \boldsymbol{\omega}_L$ as described in section 2. Moreover, the Heisenberg equations of motion for the spin-gravity interaction $H = \boldsymbol{\sigma} \cdot \boldsymbol{\Omega}_P$ are formally the same as equations (17) and (18) for the precession of an ideal test gyroscope.

In classical electrodynamics, the magnetic dipole moment for a particle of mass m and charge q is given by $\boldsymbol{\mu} = q\mathbf{S}/(2mc)$, where \mathbf{S} is its orbital angular momentum. The energy associated with the interaction of this magnetic moment with a magnetic field \mathbf{B} is $-\boldsymbol{\mu} \cdot \mathbf{B}$. Extending these notions to GEM with $q_B = -2m$, we find that a gravitomagnetic dipole moment for a gyroscope of spin \mathbf{S} is $\boldsymbol{\mu}_g = -\mathbf{S}/c$ and the energy of interaction with a gravitomagnetic field is $-\boldsymbol{\mu}_g \cdot \mathbf{B}_g = \mathbf{S} \cdot \boldsymbol{\Omega}_P$. A further extension of this result to the intrinsic spin of particles naturally leads to the interaction Hamiltonian $H = \boldsymbol{\sigma} \cdot \boldsymbol{\Omega}_P$. The gravitoelectric analogue of this interaction has already been discussed in section 1; that is, $H_{\text{int}} = -\mathbf{d}_g \cdot \mathbf{E}_g$, where $\mathbf{d}_g = A\boldsymbol{\sigma}$ would be the hypothetical gravitoelectric dipole moment of a particle [1, 2] and $\mathbf{E}_g = -\mathbf{g}$ from (15) and (16).

Let us imagine an experiment in a laboratory near the surface of an astronomical body (such as the Earth) involving the difference in the energy of a particle of spin $\sigma = s\hbar$ polarized vertically up and down (i.e. perpendicular to the surface). According to the spin-gravitomagnetic coupling, the result is

$$E_+ - E_- = 2s\hbar\Omega \sin\theta. \quad (19)$$

Here θ is the geographic latitude (i.e. $E_+ = E_-$ at the equator) and Ω is the effective frequency associated with the gravitomagnetic field

$$\Omega = \frac{2GJ}{c^2 R^3}, \quad (20)$$

where R is the mean radius of the body. Equation (19) expresses a relativistic quantum gravitational effect; indeed, one can write $\hbar\Omega = (2cJ/R^3)L_P^2$, where $L_P = (\hbar G/c^3)^{1/2}$ is the Planck length. Let us note that for the Earth $\hbar\Omega_E \simeq 2 \times 10^{-29}$ eV, while near the surface of Jupiter $\hbar\Omega_J \simeq 10^{-27}$ eV; similarly, for the Sun $\hbar\Omega_S \sim 10^{-27}$ eV, but for a neutron star $\hbar\Omega_{NS} \sim 10^{-14}$ eV.

It is important to point out that the spin-rotation-gravity coupling has appeared in the work of many authors who have studied wave equations in accelerated systems and gravitational

fields [23, 24]. In particular, the $\boldsymbol{\sigma} \cdot \boldsymbol{\Omega}_P$ interaction under scrutiny in this work first appeared in the work of de Oliveira and Tiomno [23]. The observation of wave phenomena associated with such couplings was first investigated independently in connection with possible limitations of the general theory of relativity in [25]. Dynamics in electromagnetic fields can be generated by the transformation of the momentum via $p_\mu \rightarrow p_\mu - (q/c)A_\mu$, where $A_\mu = (-\phi, \mathbf{A})$ is the electromagnetic (EM) potential. The same holds in the GEM case, except that the analogue of A_μ is $(-2\phi_g, \mathbf{A}_g)$. Let us consider, for instance, the motion of electromagnetic waves in the exterior field of a rotating mass. The effective gravitational charge in this case should be determined based on the fact that a photon of energy $\hbar\omega$ in ‘cyclotron’ motion has an effective inertial mass of $\hbar\omega/c^2$ and hence the effective GEM charges are $q_E = -\hbar\omega/c^2$ and $q_B = -2\hbar\omega/c^2$. The eigenvalue problem in gravitomagnetic fields leads to discreteness properties for the modes reminiscent of the Fock–Darwin–Landau levels in a magnetic field. Imagine, for instance, the motion of electromagnetic waves in a gravitomagnetic field characterized by the magnitude of the effective ‘cyclotron’ frequency $\Omega_c \simeq 2GJ/(c^2r^3)$. It follows from the explicit solution of Maxwell’s equations in this background [25] that the wavefunctions are proportional to Hermite polynomials. These polynomials vary over a harmonic characteristic length scale l_g that is given by

$$l_g = \frac{c}{(\omega\Omega_c)^{1/2}} \quad (21)$$

for an electromagnetic mode of frequency ω . If in this equation we set $\hbar\omega = mc^2$ and Ω_c as the cyclotron frequency in a magnetic field, we recover the *magnetic length* that is well known in the discussion of the motion of a charged particle of mass m in a magnetic field. It is interesting to note that the gravitomagnetic acceleration length is given by $\mathcal{L}_g = c/\Omega_c$, so that the *gravitomagnetic length* (21) is the geometric mean of the reduced wavelength of radiation λ and \mathcal{L}_g . The gravitomagnetic length l_g is essentially the same as the radius of the ‘cyclotron’ orbit for a mode with frequency equal to the ‘cyclotron’ frequency ($\omega = \Omega_c$). The eigenvalue spectrum clearly shows the existence of a gravitomagnetic coupling between the photon spin and the rotation of the source [25]. One can show that in the eikonal approximation the gravitational helicity–rotation coupling leads to a differential deflection of polarized radiation, thus violating the universality of freefall in a gravitational field beyond the geometric optics limit [22, 25]. That the spin–gravity interaction violates the universality of freefall is already apparent from $H = \boldsymbol{\sigma} \cdot \boldsymbol{\Omega}_P$, since this Hamiltonian depends only on the spin of the particle and is independent of its mass.

Imagine, for instance, the scattering of electromagnetic radiation by a black hole (i.e. pure geometry free of matter). For a Schwarzschild black hole, the scattering amplitude is independent of the polarization of the incident radiation, hence the polarization properties of the radiation are preserved in the scattering process. For a Kerr black hole, however, the scattering amplitude is dependent upon the polarization of the incident radiation. It is possible to give only rough and partial estimates for the motion of *wavepackets* in a gravitomagnetic field [25, 26]. The influence of helicity–rotation coupling on the gravitational deflection of electromagnetic radiation is rather weak and far below the existing observational upper limits [27], but could become important in future microlensing experiments with polarized radiation. To provide useful astrophysical estimates of the resulting polarization-dependent deflection of radiation, an *eikonal* approach has been developed for the motion of *rays* based on equations (7) and (8), i.e. $\omega(\mathbf{r}, \mathbf{k}) = ck \pm \hat{\mathbf{k}} \cdot \boldsymbol{\Omega}_P(\mathbf{r})$, so that the Einstein deflection is ignored for the sake of simplicity and only the helicity–rotation coupling is taken into account [22]. In this treatment, the total differential deflection of positive and negative helicity rays approaching the source together from asymptotic infinity and travelling to infinity after deflection vanishes in contrast

to what is expected from the wave treatment; however, it is possible to obtain useful estimates for radiation originating near the source. For instance, consider radiation originating over a pole and propagating normal to the rotation axis with an impact parameter D ; then, RCP and LCP waves separate by a total angle of $\delta \approx 4\lambda GJ/(c^3 D^3)$ about the average Einstein deflection angle. A *qualitative* description of this effect is given in [28]. The gravitomagnetic splitting δ is small; it amounts to about 1 marcsec for GHz radio waves passing over a pole of a neutron star. In addition to this splitting, one expects a wavelength-independent gravitomagnetic rotation of the plane of polarization along a ray, i.e. the Skrotskii effect that is the gravitational analogue of the Faraday effect [25, 29]. Moreover, the difference in the arrival times of positive and negative helicity radiation originating near a rotating mass and propagating freely outward to a distant point is estimated to be $T_+ - T_- = -2\lambda G\mathbf{J} \cdot \mathbf{r}/(c^4 r^3)$, where \mathbf{r} is the position vector of the point of origin of the radiation relative to the centre of the rotating source. This differential time delay due to the different phase speeds of RCP and LCP waves is too small to be measurable at present [22].

The violation of the universality of freefall is a wave effect, so that it vanishes in the $\lambda/\mathcal{L}_g \rightarrow 0$ limit. Consider, for instance, a spinning particle in a gravitomagnetic field with the interaction Hamiltonian $H = \boldsymbol{\sigma} \cdot \boldsymbol{\Omega}_p$. This potential energy is position dependent; therefore, there exists a gravitomagnetic Stern–Gerlach force $\mathbf{F} = -\nabla H$ acting on the particle that is independent of mass and hence violates the universality of the gravitational acceleration. Specifically,

$$\mathbf{F} = \frac{3GJ}{c^2 r^4} \{ [5(\boldsymbol{\sigma} \cdot \hat{\mathbf{r}})(\hat{\mathbf{J}} \cdot \hat{\mathbf{r}}) - \boldsymbol{\sigma} \cdot \hat{\mathbf{J}}] \hat{\mathbf{r}} - (\boldsymbol{\sigma} \cdot \hat{\mathbf{r}}) \hat{\mathbf{J}} - (\hat{\mathbf{J}} \cdot \hat{\mathbf{r}}) \boldsymbol{\sigma} \}, \quad (22)$$

so that the weight operator for the particle $W = mg - \mathbf{F} \cdot \hat{\mathbf{r}}$ is given by $W = mg - 3H/r$. If the spin is polarized vertically up or down in a laboratory near the Earth,

$$W_{\pm} = mg \mp \frac{3s}{R} \hbar \Omega \sin \theta, \quad (23)$$

so that $W_{\pm} = mg(1 \mp \epsilon)$, where ϵ can be expressed as

$$\epsilon = 6s \left(\frac{I}{MR^2} \right) \left(\frac{\hbar \omega}{mc^2} \right) \sin \theta. \quad (24)$$

Here $J = I\omega$, I is the moment of inertia and ω is the proper rotation frequency of the Earth. For a neutron near the Earth's surface, $\hbar\omega/(m_n c^2) \simeq 5 \times 10^{-29}$; hence, ϵ is too small to be measurable in the foreseeable future. It follows that for polarized materials the relevant ϵ is expected to be even smaller. Let us note that ϵ is directly proportional to $\hbar\omega/(mc^2)$, which can be expressed as the ratio of the Compton wavelength of the particle (\hbar/mc) to the rotational acceleration length of the observer (c/ω). Indeed, the extended nature of the particle makes it possible for its intrinsic spin to couple to the spacetime curvature, resulting in a force \mathbf{F} that has an exact analogue in the classical Mathisson–Papapetrou spin-curvature force [22, 28].

Thus far we have discussed the gravitomagnetic spin–rotation coupling in terms of a single rotating source such as the Earth. However, the universality of the gravitational interaction implies that the whole mass–energy content of the universe is involved in every physical experiment via the gravitational interaction. In classical physics, the gravitational force of the rest of the universe enters only through its gradients, which turn out to be rather small for experiments in the solar system. The situation is in general different in quantum physics, however. For instance, in the calculation of the spin–gravity coupling, the gravitomagnetic field generated by the total mass–energy current must be taken into account. This is a

difficult problem; however, to obtain some idea of what is involved here we may use the linear approximation to write the interaction Hamiltonian as

$$H = \frac{G}{c^2} \sum_a \frac{3(\mathbf{r}_a \cdot \mathbf{J}_a)(\mathbf{r}_a \cdot \boldsymbol{\sigma}) - r_a^2(\mathbf{J}_a \cdot \boldsymbol{\sigma})}{r_a^5}, \quad (25)$$

where the sum is over all astronomical sources and $\mathbf{r}_a = \mathbf{x}_a - \mathbf{x}$ is the vector of relative separation between the particle of spin $\boldsymbol{\sigma}$ at \mathbf{x} and the centre of mass of the source a . Equation (25) can be expressed as $H = c(\boldsymbol{\sigma} \cdot \partial_{\mathbf{x}})\Phi_g$, where

$$\Phi_g = \frac{G}{c^3} \sum_a \frac{\mathbf{J}_a \cdot \mathbf{r}_a}{r_a^3} \quad (26)$$

is the net dimensionless scalar gravitomagnetic potential defined by $\mathbf{B}_g = c^2 \nabla \Phi_g$. For a laboratory experiment near the Earth, it is simple to show that the net contribution due to the Sun, the Moon and the other planets is negligible. Therefore, to compute Φ_g one must investigate the cosmic mass–current distribution. This is a difficult observational problem and much remains unknown regarding the distribution of angular momentum in the universe. It is likely that over the largest scales no preferred sense of rotation would be discernible. These considerations lead one to surmise that near the Earth (or Jupiter) the main contribution to the Hamiltonian is simply due to the Earth (or Jupiter), though a completely satisfactory resolution is not available. Conversely, observational data regarding the gravitomagnetic spin–rotation coupling could in principle set limits on the cosmic mass–current distribution.

4. Discussion

The gravitational coupling of intrinsic spin with rotation has been described in this work and the consequences of the gravitomagnetic interaction $H = \boldsymbol{\sigma} \cdot \boldsymbol{\Omega}_p$ have been pointed out. In particular, the gravitomagnetic shift in the Larmor frequency of a nuclear particle has been estimated. Efforts are underway to improve the sensitivity of the measurement of such frequency shifts by several orders of magnitude. This could potentially make the effect measurable near the surface of Jupiter [30]. Let us recall that for Jupiter $\hbar\Omega_J \simeq 10^{-27}$ eV, corresponding to a gravitomagnetic Larmor shift of about 3×10^{-13} Hz. In view of the current interest in planetary exploration, it appears that the gravitomagnetic coupling of intrinsic spin with rotation could be measurable in the foreseeable future.

Acknowledgments

I am grateful to Friedrich Hehl and Michael Romalis for helpful discussions and correspondence.

References

- [1] Kobzarev I Yu and Okun L B 1963 *Sov. Phys.–JETP* **16** 1343
- [2] Leitner J and Okubo S 1964 *Phys. Rev. B* **136** 1542
Hari Dass N D 1976 *Phys. Rev. Lett.* **36** 393
Peres A 1978 *Phys. Rev. D* **18** 2739
Morgan T A and Peres A 1962 *Phys. Rev. Lett.* **9** 79
- [3] Dabbs J W T, Harvey J A, Paya D and Horstmann H 1965 *Phys. Rev. B* **139** 756
- [4] Velyukhov G E 1968 *JETP Lett.* **8** 229
- [5] Vasil'ev B V 1969 *JETP Lett.* **9** 175

- [6] Young B A 1969 *Phys. Rev. Lett.* **22** 1445
- [7] Wineland D J and Ramsey N F 1972 *Phys. Rev. A* **5** 821
- [8] Hayasaka H and Takeuchi S 1989 *Phys. Rev. Lett.* **63** 2701
- [9] Faller J E, Hollander W J, Nelson P G and McHugh M P 1990 *Phys. Rev. Lett.* **64** 825
Quinn T J and Picard A 1990 *Nature* **343** 732
Nitschke J M and Wilmarth P A 1990 *Phys. Rev. Lett.* **64** 2115
- [10] Adelberger E G, Heckel B R, Stubbs C W and Rogers W F 1991 *Ann. Rev. Nucl. Part. Sci.* **41** 269
- [11] Wineland D J *et al* 1991 *Phys. Rev. Lett.* **67** 1735
- [12] Venema B J *et al* 1992 *Phys. Rev. Lett.* **68** 135
- [13] Ritter R C, Winkler L I and Gillies G T 1993 *Phys. Rev. Lett.* **70** 701
- [14] Berglund C J *et al* 1995 *Phys. Rev. Lett.* **75** 1879
Youdin A N *et al* 1996 *Phys. Rev. Lett.* **77** 2170
- [15] Ni W-T *et al* 1999 *Phys. Rev. Lett.* **82** 2439
Chui T C P and Ni W-T 1993 *Phys. Rev. Lett.* **71** 3247
- [16] Vorobyov P V and Gitarts Ya I 1988 *Phys. Lett. B* **208** 146
Bobrakov V F *et al* 1991 *JETP Lett.* **53** 294
Vorob'ev P V 1994 *JETP Lett.* **59** 510
- [17] Jacobs J P *et al* 1995 *Phys. Rev. A* **52** 3521
- [18] Werner S A, Staudenmann J-L and Colella R 1979 *Phys. Rev. Lett.* **42** 1103
Rauch H and Werner S A 2000 *Neutron Interferometry* (Oxford: Clarendon)
Atwood D K, Horne M A, Shull C G and Arthur J 1984 *Phys. Rev. Lett.* **52** 1673
Bonse U and Wroblewski T 1983 *Phys. Rev. Lett.* **51** 1401
Hasselbach F and Nicklaus M 1993 *Phys. Rev. A* **48** 143
Moorhead G F and Opat G I 1996 *Class. Quantum Grav.* **13** 3129
Schwab K, Bruckner N and Packard R E 1997 *Nature* **386** 585
Gustavson T L, Bouyer P and Kasevich M A 1997 *Phys. Rev. Lett.* **78** 2046
- [19] Mashhoon B 1995 *Phys. Lett. A* **198** 9
Mashhoon B, Neutze R, Hannam M and Stedman G E 1998 *Phys. Lett. A* **249** 161
- [20] Beth R A 1936 *Phys. Rev.* **50** 115
- [21] Mashhoon B 1988 *Phys. Rev. Lett.* **61** 2639
Mashhoon B 1992 *Phys. Rev. Lett.* **68** 3812
Mashhoon B 1989 *Phys. Lett. A* **139** 103
Mashhoon B 1990 *Phys. Lett. A* **143** 176
Mashhoon B 1990 *Phys. Lett. A* **145** 147
Mashhoon B 1993 *Phys. Rev. A* **47** 4498
- [22] Mashhoon B 1993 *Phys. Lett. A* **173** 347
Mashhoon B, Gronwald F and Theiss D S 1999 *Ann. Phys., Lpz.* **8** 135
Mashhoon B, Gronwald F and Lichtenegger H I M 2000 *Testing Relativistic Gravity in Space* ed C Lämmerzahl, C W F Everitt and F W Hehl (Berlin: Springer)
(Mashhoon B, Gronwald F and Lichtenegger H I M 1999 *Preprint gr-qc/9912027*)
- [23] de Oliveira C G and Tiomno J 1962 *Nuovo Cimento* **24** 672
Mitskievich N V 1969 *Physical Fields in General Relativity Theory* (Moscow: Nauka) (in Russian)
Schmutzer E 1973 *Ann. Phys., Lpz.* **29** 75
Barker B M and O'Connell R F 1975 *Phys. Rev. D* **12** 329
Schmutzer E and Plebański J 1977 *Fortschr. Phys.* **25** 37
- [24] Hehl F W and Ni W-T 1990 *Phys. Rev. D* **42** 2045
Cai Y Q and Papini G 1991 *Phys. Rev. Lett.* **66** 1259
Cai Y Q and Papini G 1992 *Phys. Rev. Lett.* **68** 3811
Anandan J 1992 *Phys. Rev. Lett.* **68** 3809
Andretsch J and Lämmerzahl C 1992 *Appl. Phys. B* **54** 351
Silverman M P 1992 *Nuovo Cimento D* **14** 857
Huang J 1994 *Ann. Phys., Lpz.* **3** 53
Soares I D and Tiomno J 1996 *Phys. Rev. D* **54** 2808
Ryder L H 1999 *Gen. Rel. Grav.* **31** 775
- [25] Mashhoon B 1974 *Nature* **250** 316
Mashhoon B 1974 *Phys. Rev. D* **10** 1059
Mashhoon B 1975 *Phys. Rev. D* **11** 2679
- [26] Damour T and Ruffini R 1974 *C.R. Acad. Sci. A* **279** 971

- de Logi W K and Kovacs S J Jr 1977 *Phys. Rev. D* **16** 237
Leahy D A 1982 *Int. J. Theor. Phys.* **21** 703
Mashhoon B 1987 *Phys. Lett. A* **122** 299
Futterman J A H, Handler F A and Matzner R A 1988 *Scattering from Black Holes* (Cambridge: Cambridge University Press)
Feng L L and Lu T 1991 *Class. Quantum Grav.* **8** 851
Carini P, Feng L L, Li M and Ruffini R 1992 *Phys. Rev. D* **46** 5407
[27] Harwit M *et al* 1974 *Nature* **249** 230
Dennison B, Dickey J and Jauncey D 1976 *Nature* **263** 666
Dennison B *et al* 1978 *Nature* **273** 33
[28] Mashhoon B 1999 *Gen. Rel. Grav.* **31** 681
[29] Kopeikin S and Mashhoon B 2000 *Preprint*
[30] Romalis M 2000 Private communication