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Impact of Microstructure on the Electron-hole Interaction in Metal Halide Perovskites

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25 Abstract

26 Despite the remarkable progress in the performance of devices based on the metal halide 27 perovskite semiconductor family, there is still a lack of consensus on their fundamental 28 photophysical properties. Here, using magneto-optical transmission spectroscopy we elucidate 29 the impact of the microstructure on the Coulomb interaction between photo-created electron-30 hole pairs in methylammonium lead triiodide (MAPbI₃) and the triple-cation lead mixed-31 halide composition, $C_{s_{0.05}}(MA_{0.17}FA_{0.83})_{0.95}Pb(I_{0.83}Br_{0.17})_{3}$ (Cs: Cesium, MA: 32 methylammonium, FA: formamidinium) by investigating thin films with a wide range of grain 33 sizes from tens of nanometers to microns. At low temperatures, in which thermal fluctuations 34 of the interactions are frozen and the rotational disorder of the organic cation is negligible, the 35 exciton binding energy and reduced effective mass of carriers remain effectively unchanged 36 with grain size. We conclude that the microstructure plays a negligible role in the Coulomb 37 interaction of the photo-created electron-hole pairs, in contrast to previous reports. This 38 renewed understanding of the relationship between these fundamental electronic properties 39 and the microstructure is critical for future fundamental studies and improving device design.

Key words: hybrid perovskites, microstructure, exciton binding energy, excitonic reduced
effective mass, magneto-optical spectroscopy

42 **Broader context:**

43 Metal halide perovskites such as methylammonium lead iodide (CH₃NH₃PbI₃) and alloyed 44 mixed-cation lead mixed-halide counterparts such as Cs_{0.05}(MA_{0.17}FA_{0.83})_{0.95}Pb(I_{0.83}Br_{0.17})₃ are 45 creating enormous excitement for their use in high-performance solar cells with potential for 46 exceptionally low capital-intensity for production. An important parameter that dictates the 47 design of a solar cell is the strength of the Coulomb interaction between the photo-generated 48 electron in the conduction band and the hole in the valence band (i.e. the exciton binding 49 energy). Recently, it has been suggested that the exciton binding energy of the perovskite material is influenced by the size of the grains of the perovskite film, and therefore different 50 51 device design will be required depending on the grain size. Here, we perform magneto-optic 52 measurements on films with different grain sizes to directly show that the microstructure has a 53 negligible influence on the excitonic properties in the samples, and that excitons do not play a 54 significant role in films with any of the polycrystalline morphologies studied. This means that excitons should not be factor for the design of optoelectronic devices based on polycrystalline 55 56 CH₃NH₃PbI₃ and similar materials with different grain sizes, making future device architecture optimisation less constrained. 57

58 Introduction

Solar cells based on lead halide perovskites (with general chemical formula of ABX₃; A = $CH_3NH_3^+$, $HC(NH_2)_2^+$ or Cs^+ ; B = Pb²⁺; X = I⁻, Br⁻ or Cl⁻) have achieved astonishingly high power conversion efficiencies exceeding 22% in only seven years since they were first used as sensitizers in photovoltaic solar cells ¹⁻³. Since then, the remarkable properties of these crystalline semiconductors, such as the tunability of their optoelectronic properties through fabrication of layers with different morphologies, compositions and crystal qualities ⁴⁻⁷, have been further exploited to develop a variety of promising optoelectronic devices ⁸⁻¹¹.

As for any emerging semiconductor material deployable in optoelectronic devices, the strength of the Coulomb interaction between the photo-generated electron in the conduction band and the hole in the valence band is of high importance. This was a focus of several studies on perovskites about two decades ago ¹²⁻¹⁴ and has recently regained attention ¹⁵⁻¹⁸ mainly due to the promising outlook of perovskites for solar photovoltaics ¹⁰, light-emitting diodes ¹¹ and lasing applications ¹⁹ with polycrystalline thin films and single crystals of superior quality than those synthesized two decades ago. 73 The latest studies have come to a consensus that the exciton binding energy – alternatively 74 known as effective Rydberg (Ry^*) in the Wannier-Mott exciton model – of well-studied CH₃NH₃PbI₃ and CH₃NH₃PbBr₃ polycrystalline thin films is ~5-12 meV ^{15, 17, 18} and ~25 meV 75 ^{16, 18} in the room temperature phase, respectively. These low values, with respect to thermal 76 energy at room temperature ($k_b \times T \approx 26$ meV; k_b is the Boltzmann constant and T is the 77 temperature in Kelvin), are due to generally weak many-body effects (i.e. Coulombic 78 interactions) in these semiconductor systems ²⁰. Nevertheless, there is ongoing debate 79 80 concerning the possible tunability of this parameter with variations of the microstructure and hence crystal quality of the perovskite layer ^{21, 22}, being speculated to influence the photo-81 82 excited exciton binding energies.

83 It has been proposed that disordered organic cation dipole domains in small and defective 84 crystals with a large portion of poly-crystallinity can create a large electrostatic potential fluctuation which causes a substantial screening of the photo-generated excitons ²². This could 85 lead to a very small value of Ry^* on the order of ~2-5 meV ^{23, 24} meaning that the material 86 behaves as if in the non-excitonic regime. On the other hand, large and defect-free crystals can 87 exhibit an order of magnitude greater values of $Ry^{*22, 25}$. As a result, one current hypothesis is 88 89 that it is not possible to assign a universal value to the exciton binding energy of the 90 commonly used CH₃NH₃PbI₃ hybrid lead halide perovskites.

A key consequence of this hypothesis could be a change in the working regime (i.e. photoexcited carrier dissociation and transport) of solar cells fabricated from active layers with diverse microstructures. Knowledge of the value of the exciton binding energy is essential because it determines the nature of the majority of the photo-generated species (i.e. free carriers versus excitons) in a semiconductor. For devices comprised of materials with binding energies greater than the probability of thermal dissociation at the working temperature, there would need to be an additional mechanism for exciton dissociation into free carriers, such as 98 an additional heterojunction, through which they can readily contribute to the photocurrent ²⁶.
99 In the case that the binding energies are so low that the fraction of photo-excited species that
100 are excitons is negligible, the n- and p-type contacts in the heterojunction structure of the solar
101 cell (cf. most perovskite solar cells) would simply act as charge-selective contacts for the
102 photo-generated electrons and holes, respectively, facilitating simpler device designs.
103 Altogether, these points highlight the need to fully understand the exciton-morphology
104 relationship.

105 Here, we elucidate the influence of the microstructure, manipulated through the perovskite fabrication method, on the exciton binding energy Ry^* and the excitonic reduced mass (μ) for 106 107 a variety of CH₃NH₃PbI₃ crystal sizes covering the size range reported in devices to date. The 108 impact of the local order and the degree of poly-crystallinity on electron-hole interaction is unraveled through a direct measurement of Ry^* and μ for thin films with different grain sizes. 109 110 These measurements are performed at low temperature (2 K) at which the impact of $CH_3NH_3^+$ 111 thermal orientational dynamic rearrangement is mitigated. We provide strong evidence that 112 the grain size has negligible influence on the exciton binding energy and effective mass in 113 CH₃NH₃PbI₃ and, thus, does not affect the dissociation probability of the bulk excitons and 114 their stability in the photo-excited charge carrier population.

115 **Results and Discussion**

116 Sample Morphology

117 A wide range of fabrication methods, from single-step 27 to sequential deposition 28 and 118 solution-processed 29 to vapor-assisted 30 techniques, have been employed to fabricate the 119 active layer of perovskite optoelectronic devices in the literature. Here, different deposition 120 methods, which have been shown to give corresponding high-performance solar cells with 121 efficiencies of >16 % $^{6, 31, 32}$, are used to controllably tune the perovskite grain size in the 122 films. These samples with CH₃NH₃PbI₃ chemical composition but fabricated via different 123 methods with various resulting morphologies simulate the range of different degrees of 124 structural disorder and poly-crystallinity in the CH₃NH₃PbI₃ layer in high-performance perovskite devices fabricated to date ²². The top-view scanning electron microscope (SEM) 125 126 images of the samples are presented in **Figure 1A-D**. The morphologies vary from a planar polycrystalline film with grain sizes of 772±227 nm ³² (A; large grain polycrystalline 127 annotated LPC) and polycrystalline film with grain sizes of 214±57 nm ³¹ (**B**; small grain 128 129 polycrystalline annotated SPC), to small crystals of 291±64 nm fabricated by a two-step dipping technique ⁶ (C; annotated SC) and also perovskite infiltrated into a $\sim 2 \mu m$ thick 130 131 mesoporous Aluminum oxide scaffold (mp-Al₂O₃) with grain sizes of <50 nm (**D**; annotated 132 MP). As can be seen in the top-view SEM image of the MP sample, there is no visible 133 capping layer formed on top of the mesoscopic phase and, therefore, a potential artefact from 134 the larger grain sizes of a capping layer does not influence the related data analysis.

135 Magneto-optical Transmission Measurements

136 Studying the optical behavior of excitons under external fields has proven to be a powerful 137 tool to reveal properties of these quasi-particles in semiconductors ³³. Here, we conducted 138 optical transmission measurements under high magnetic fields (B) up to 70T on the samples 139 presented in Figure 1. Typical transmission spectra for the large grain polycrystalline thin 140 film sample, showing a strong 1s excitonic transition, are given in Figure 2A. For 141 unambiguous identification of the weaker optical transitions, we show in Figure 2B the ratio 142 of the transmission spectrum at a specific high magnetic field over the transmission spectrum 143 taken at zero field, illustrating the free-carrier transitions. In Figure 2C we show transmission 144 measurements at several monochromatic probe energies under very high magnetic field 145 (<150T) achieved using a short pulse technique. The related spectra for other morphologies 146 are provided in the Supplementary Information, Figures S1-S3.

We note that a hydrogen-like excitonic behavior 34 can be assumed for CH₃NH₃PbI₃ in the low-temperature orthorhombic phase. The strong minimum in the transmission spectra (**Figure 2A**), attributed to the 1s state of the hydrogen-like free exciton ¹⁷, blue shifts with increase of the magnetic field. The second minimum, located at slightly higher energies than the 1s state, is attributed to the first excitonic excited state (labelled as 2s in **Figure 2A**) and also exhibits a blue-shift with magnetic field.

153 The Landau levels, the free carrier states with quantum numbers N=0,1,2,... formed by quantization of the particle's motion in the plane perpendicular to the magnetic field ³³, can be 154 observed at higher energies in the strong magnetic field regime in which $\hbar\omega_c >> Ry^*$ (155 $\hbar\omega_c = \hbar eB/\mu$), where \hbar and e are Planck's constant and the electron charge, respectively, ω_c is 156 the cyclotron frequency, and $\mu^{-1} = m_e^{-1} + m_h^{-1}$ is the reduced mass of the exciton (m_e and m_h are 157 158 electron and hole masses, respectively). The free carrier Landau level transitions are dominant at high magnetic fields with transition energy described by $E(B) = E_g + (N+1/2)\hbar\omega_{c-16, 17}$ 159 where E_{g} is the bandgap. We can then fit the measured optical transitions between Landau 160 161 levels to extract the reduced mass, μ , independent of the exciton binding energy. The value of 162 the reduced mass thus obtained is used as a fixed parameter in the fit of the hydrogen-like excitonic transitions to extract the exciton binding energy, based on a model which describes 163 the hydrogen atoms in high magnetic field $^{35}.$ In this model, a dimensionless parameter $^{\gamma}$ 164 which is defined as $\gamma = \hbar \omega_c / 2Ry^*$ is used. Makado and McGill calculated the transition 165 energies of excitonic states $E_n(\gamma)$ as a function of quantum number n up to 4 and y ranging 166 from 0 to 10 35 . Given values of the magnetic field *B* and the reduced mass μ , the binding 167

energy Ry^* is the only fitting parameter for the full set of hydrogenic transitions.

169 The complete set of these excitonic and free carrier transition energies for the LPC sample is 170 plotted in the fan diagram in Figure 2D. The observation of many Landau levels in the 171 magnetic field <150T allows precise determination of the reduced mass. Knowledge of the 172 reduced mass in turn imposes strong constraints on fitting the excitonic binding energy. An 173 additional influential constraint on the value of the binding energy is set if the higher 174 excitonic transition resonances such as the 2s state are resolved, as for the LPC morphology. 175 The observation of a well-resolved 2s state at intermediate magnetic field fixes the transition 176 energy of 2s at zero magnetic field using the model described earlier. The energy separation 177 of 1s-2s at zero magnetic field follows the series of 3D hydrogen-like energy states, $E_n = E_g - Ry^*/n^2$ where E_n is the excitonic transition energy at zero magnetic field. This 178 179 gives a strong constraint on the exciton binding energy.

180 The fitting to the hydrogen-like excitonic (dashed lines) and free carrier (solid lines) transitions, as explained in detail in refs. ^{16, 17, 35}, is presented in Figure 2D and the Ry^* and μ 181 values for the LPC sample are calculated to be 16 ± 1 meV and 0.102 ± 0.002 m₀, respectively. 182 183 For the morphologies other than LPC (i.e. SPC, SC and MP), the identification of free carrier 184 transitions at high energies allows us to obtain the values of the the reduced masses. The 185 extracted μ values are presented in **Table 1** and are very close to the value originally reported by Miyata et al. ($\mu = 0.104 \pm 0.003 \ m_0$)¹⁷. Importantly, we find that the microstructure has 186 187 negligible influence on the reduced mass. This indicates that the polaron coupling contant, 188 incorporating the effective longitudinal optical (LO) phonon energy and the optical and static 189 dielectric constants of the medium, which couples to the band-edge bare effective masses to form the polaronic masses 18, 36, is only minimally influenced by the variation in 190 191 microstructure.

192 In the absence of a clear observation of the 2s excitonic transitions for the SPC, SC and MP 193 samples, the global fit to the full excitonic and free-carrier transitions was employed and the 194 resulting unchanged Ry^* values are presented in **Table 1**. To visualize the negligible change in 195 the optical response under external magnetic fields of the CH₃NH₃PbI₃ films with different 196 morphology, the overlaid fan diagrams of the complete set of excitonic and free carrier optical 197 transitions after shifting by the corresponding bandgap offsets are presented in Figure 3(A). 198 Here we clearly show that these CH₃NH₃PbI₃ samples, although composed of dramatically 199 different grain sizes and morphologies, exhibit similar magneto-optical behaviour. 200 Furthermore, the XRD patterns (Figure S4) show that the CH₃NH₃PbI₃ SPC and SC thin 201 films have some PbI₂ impurities, indicating that the presence of small amounts of PbI₂ also 202 does not affect the excitonic properties of perovskites. We also note that, although the changes 203 in the grain size and morphology do not influence the low-temperature excitonic properties of 204 perovskites, they can affect the optoelectronic response of these materials including their PL 205 behaviour and the fraction of retained tetragonal phase within the orthorhombic phase upon lowering the temperature ³⁷⁻³⁹. 206

207 It was recently shown that when replacing CH₃NH₃⁺ by a larger monovalent organic cation, 208 formamidinium (HC(NH₂) $_2^+$), the excitonic characteristics of the corresponding perovskite, Ry^* and μ , decreased by less than 15 % ¹⁶. The larger steric size of formamidinium compared 209 210 to methylammonium results in a larger equatorial bond angle (lead-iodide-lead) and, thus, a 211 narrower optical bandgap 40. The variation of the crystal size would result in a modified 212 interplay between the organic and the inorganic moieties in the perovskite structure, which in 213 turn changes the strain induced on the crystal structure and thus, impacts the bandgap. The 214 negligible impact on the excitonic characteristics of the degree of dipolar cation order and 215 structural arrangement induced by the variation of perovskite microstructure might be 216 analogous to the small impact of cation substitution on the excitonic properties, with the leadhalogen sub-unit remaining unchanged in both cases. This indeed demonstrates the critical role of the electronic properties of the inorganic cage in the determination of the excitonic binding energy and reduced effective mass of perovskites. Moreover, this important observation agrees with the recent work by Perez-Osorio *et al.* in which the largest contribution to the static dielectric constant of $CH_3NH_3PbI_3$ was demonstrated to originate from the stretching and rocking vibrational modes of the lead-halogen system rather than the internal modes of the organic cation ⁴¹.

224 Using the values obtained for the effective dielectric constant ($\varepsilon = \sqrt{Ry^H \times \mu/Ry^* \times m_0}$

where $Ry^{H} = 13.6 \text{ eV}$ is the Rydberg constant of the hydrogen atom) and the reduced mass, the Bohr radius of the bound electron-hole system can be calculated. This is similar for all the CH₃NH₃PbI₃ perovskite morphologies and is about 5 nm in the low temperature phase. The Bohr radius is substantially smaller even than the perovskite crystal size of the **MP** sample and it is consistent with the absence of any signature of a quantum confinement effect on the optical transition of the corresponding excitons in the mesoscopic phase.

231 The correct interpretation of the optical absorption spectrum for the exciton binding energy 232 derivation in the room temperature tetragonal phase of CH₃NH₃PbI₃ (> 160 K) becomes more 233 difficult than in the orthorhombic phase as the large spectral broadening at higher 234 temperatures in the tetragonal phase can potentially mask critical optical transition features even under high magnetic fields ¹⁷. The increased broadening is most likely due to a 235 236 thermally-induced increase in the orientational disorder and rotational dynamics of the unlocked organic cation above the structural phase transition ⁴². The excitonic binding energy 237 has been reported in the literature ^{17, 18, 24, 43} to decrease upon orthorhombic-to-tetragonal 238 239 phase transition with a likely reason to be the slight increase in the dielectric constant of the 240 medium, which is in turn linked to the increase in the effective phonon energy coupling to the241 carriers.

242 We argue that the change in the binding energy at the transition to the tetragonal phase for all 243 the morphologies investigated in this study would be similar. This can be partly understood by 244 the dominant influence of thermal disorder on the electrostatic fluctuations at room 245 temperature over the degree of poly-crystallinity and molecular order which are dominant at 246 lower temperatures ²²; the latter properties we have already shown above to have insignificant 247 impact on the electron-hole interaction. This is further supported by a recent theoretical study 248 by Motta et al. in which the impact of the degree of molecular dipole orientational disorder -249 which could be induced by the grain size variation – on the binding energy in the high temperature phase of CH₃NH₃PbI₃ was found to be only $\sim 10 \%$ ⁴⁴. 250

251 To generalise our findings for CH₃NH₃PbI₃ to other perovskite compositions, we repeat the 252 state-of-the-art triple-cation lead mixed-halide measurements on a perovskite, $Cs_{0.05}(MA_{0.17}FA_{0.83})_{0.95}Pb(I_{0.83}Br_{0.17})_3^2$. We consider two significantly different morphologies 253 254 of this perovskite composition, namely a planar polycrystalline thin film and a film with the 255 perovskite infiltrated into a thick mp-Al₂O₃ scaffold (corresponding SEM images are shown 256 in Figure S5). We show in Figure S6 that both morphologies exhibit similar magneto-optical 257 responses and, thus, similar fundamental excitonic properties, with the binding energy and 258 reduced mass of 13 ± 2 meV and 0.096 ± 0.008 m₀, respectively (see **Table S1**). These results 259 allow us to extend our conclusions to the wider alloyed perovskite family, providing strong 260 evidence that the microstructure has a negligible impact on the excitonic properties in metal 261 halide perovskites.

Recently, Petrozza and co-workers used temperature-dependent time-resolved transient absorption (TA) spectroscopy to track the fate of excitons following photo-excitation and to

understand the relationship between the excitonic properties and film morphology ²². The 264 265 conclusion from this study was that the exciton binding energy and, therefore, exciton 266 stability can vary depending on the perovskite film microstructure in the tetragonal phase 267 which is most critical to solar cell operation. Their interpretation was based on the emergence 268 and evolution of the negative feature located at the high-energy end of the photobleach 269 signature in the transmission-difference spectrum, which was attributed to self-270 renormalization of the exciton energy due to a change in its concentration upon decrease in 271 temperature and/or grain size variation. We note that the negative feature developing in the 272 high-energy tail of the TA spectra over time has been proposed to be primarily a consequence of a bandgap renormalization effect (bandgap narrowing) and of a free carrier absorption, 273 among other possible mechanisms ^{45, 46}. The spectral TA analysis used in ref. ²², which is an 274 275 indirect technique for characterising excitonic properties, can also be influenced by 276 recombination mechanisms which are strongly influenced by the perovskite crystal morphology. Furthermore, we note from a previous report for CH₃NH₃PbBr₃ perovskite that 277 while Ry^* probably remains unchanged by increasing the temperature from 2 K to room 278 279 temperature, the increase in the broadening of the linear absorption spectra lessens the 1s transition strength ^{18, 47}. Although this spectral broadening increase is mainly attributed to 280 281 thermally-activated mechanisms, a change in the intrinsic broadening of perovskites with 282 different microstructures, due for instance to the difference in their inherent structural disorder, can similarly influence the spectral shape while Ry^* remains unchanged. This makes 283 284 unambiguous interpretation of the linear and transient absorption spectra difficult when relying on the detection and the strength of the excitonic feature – at the absorption edge – as 285 these techniques are indirect approaches for Ry^* determination compared to the direct 286 287 approach employed in this study (see Figure S7 for the zero-field transmission spectra of the 288 four different CH₃NH₃PbI₃ morphologies). We note that it is possible there are also transient changes in excitonic populations and properties (e.g. due to dynamic photo-excited changes indielectric constant) that will be averaged out in our steady-state measurements.

291 Conclusions

292 We investigated the relationship between the excitonic properties and microstructure of the 293 CH₃NH₃PbI₃ and Cs_{0.05}(MA_{0.17}FA_{0.83})_{0.95}Pb(I_{0.83}Br_{0.17})₃ perovskites using magneto-optical 294 spectroscopy measurements. We find that the microstructure does not influence the excitonic 295 binding energy of CH₃NH₃PbI₃ or the broader range of mixed-cation and mixed-halide 296 alloyed compositions. Thus, we conclude that excitons truly play a negligible role in the 297 operation of mixed organic-inorganic metal halide-based perovskites regardless of the thin 298 film deposition technique and final morphology, and that excitons need not be a factor for the 299 design of optoelectronic devices based on polycrystalline CH₃NH₃PbI₃ and similar materials 300 with different grain sizes. The universal values of ~16 meV and 0.102-0.109 m_0 for the 301 exciton binding energy and reduced effective mass in the low-temperature phase of 302 CH₃NH₃PbI₃ can be used as guidelines for interpretation of the spectroscopic data such as the 303 complex spectra of transient absorption measurements. It also implies that the electronic 304 structure of the inorganic cage $-(PbI_3)^-$ in the case of CH₃NH₃PbI₃ - is likely to have the 305 greatest contribution to the excitonic properties of organic-inorganic lead halide perovskites 306 rather than the degree of poly-crystallinity and the order of dipolar organic cation domains.

308 Figures





Figure 1 Morphology of CH₃NH₃PbI₃ samples. Top-view SEM images of CH₃NH₃PbI₃ layer with various morphologies fabricated on glass substrates. **A)** Polycrystalline thin film with grain sizes 772 \pm 227 nm (LPC). **B)** Polycrystalline thin film with grain sizes 214 \pm 57 nm (SPC). **C)** Small crystals of CH₃NH₃PbI₃ fabricated by the two-step solution-processed dipping technique, with resulting crystal sizes 291 \pm 64 nm (SC). **D)** CH₃NH₃PbI₃ infiltrated into a thick mesoporous-Al₂O₃ scaffold with grain sizes <50 nm (MP). Scale bars are 1 μ m.



316 Figure 2 Magneto-optical transmission spectrum and fan chart measured at 2 K. A) Sequence of typical 317 optical transmission spectra of the LPC sample measured at the indicated magnetic fields which show the 1s and 318 2s excitonic transitions with increased transmission. B) Sequences of the ratios of the transmission spectra in 319 magnetic field T(B) to that measured at zero field T(0). The resonant absorption features of the free carrier 320 Landau levels correspond to minima. C) Typical results of the low temperature monochromatic transmission as a 321 function of magnetic field obtained by the short pulse technique. **D**) Fan chart of $CH_3NH_3PbI_3$ for the large grain 322 polycrystalline sample. The circular data points are from long pulse field measurements and star symbol points 323 are collected by short pulse mega-gauss measurements. The solid lines and the dashed lines are the fits to the set 324 of Landau levels and the excitonic transitions, respectively.

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Figure 3 Comparison of the excitonic properties of the four different CH₃NH₃PbI₃ morphologies. A) Overlaid fan chart of the LPC, SPC, SC and MP samples. The solid lines and the dashed lines are the fits to the set of Landau levels and the excitonic transitions, respectively. B) The zoomed-in plot of the excitonic transitions region of the fan chart in panel (A). C) The bottom panel illustrates the change in the Ry^* and μ for the various morphologies.

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Table 1 Summary of the parameters of the fits to the full Landau fan chart for the four different morphologies of

337 CH₃NH₃PbI₃ in the low temperature (2 K), orthorhombic phases. The figures in the brackets are the error

338 estimates of the parameters.

	Size Distribution (nm)	E _g (meV)	μ (m ₀)	<i>Ry</i> [*] (meV)
Large Polycrystalline	772(227)	1642(2)	0.102(0.002)	16(1)
Small Polycrystalline	214(57)	1643(2)	0.105(0.002)	16(4)
Small Crystal	291(64)	1639(2)	0.109(0.002)	16(4)
Mesoporous	<50	1638(2)	0.107(0.003)	16(4)

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356 Author Contributions

AMS conceived the idea and designed the experiments with SDS and PP. TY, RB, AP and MA-J fabricated the samples. AMS, SDS and MA-J performed the SEM imaging and XRD measurements. The magneto-optic experiments were performed, under the supervision of PP, by ZY, AM, AS, KG, JU and NZ, and the data analyzed by ZY. AH-B and MAG supervised AMS, Y-BC supervised AP, OP supervised AM, SDS supervised MA-J, SDS and VB supervised RB. AMS prepared the first draft of the manuscript with technical and editorial inputs from SDS, PP and RJN. All authors contributed to the final version of the manuscript.

364 Methods

365 CH₃NH₃PbI₃ Perovskite Sample preparation

366 Large grain polycrystalline film (LPC)

367 LPC perovskite films were prepared using a methylammonium iodide (CH₃NH₃I) and lead acetate $Pb(Ac)_2 \cdot 3H_2O$ precursor mixture³². To generate the perovskite solution, CH_3NH_3I 368 369 (Dyesol) and Pb(Ac)₂ \cdot 3H₂O were dissolved in anhydrous N,N-dimethylformamide at a 3:1 370 molar ratio with final concentration of 30 wt%, and the stabilizer hypophosphorous acid 371 (HPA) was added at a molar ratio of 7.5% with respect to $Pb(Ac)2 \cdot 3H_2O$. $Pb(Ac)_2 \cdot 3H_2O$ 372 (316512) and HPA (214906) were purchased from Sigma Aldrich. Microscope slides and 373 coverslips were washed sequentially with soap (2% vol. Hellmanex in water), de-ionized 374 water, isopropanol, acetone and finally treated under oxygen plasma for 10 min. The 375 precursor solution was spin-coated at 2,000 r.p.m. for 45 s in a nitrogen-filled glovebox, and 376 the substrates were then dried at room temperature for 10 min before annealing at 100 °C for 5 377 min. The samples were then stored in a nitrogen-filled glovebox until used.

378 Small grain polycrystalline film (SPC)

Soda-lime glass substrates 1 mm thick were purchased from J. Melvin Freed. The substrates
were cleaned ultrasonically using soap solution (Helmanex), water, ethanol and iso-propanol
for 15 minutes, each.

382 PbI₂ (99.99%) was purchased from Alfa Aesar and used without further purification. MAI
383 was either synthesised in-house using a common method or else purchased from Dyesol. All
384 the film preparation work was carried out under nitrogen atmosphere.

385 CH₃NH₃PbI₃ films were prepared by the "gas-assisted" method of Huang *et al.* ³¹. A 25µL 45
386 wt% CH₃NH₃PbI₃ DMF solution, prepared from PbI₂ and CH₃NH₃I in a molar ratio of 1:1,

was spread on the substrate, on a spin-coater. An 80 L/min stream of nitrogen gas was blown
over the film during spinning at 6500 rpm, 2 s after the spinning commenced. The films were
annealed at 100°C on a hotplate for 10 min and then cooled to room temperature.

390 Small crystal sample (SC)

Soda-lime glass substrates treatment and precursor solution (PbI₂ and MAI) preparation were
similar to that of the small grain polycrystalline film.

393 CH₃NH₃PbI₃ films were prepared by the 2-step method of D'Innocenzo *et al.* ⁶. The clean 394 glass substrates were heated on a hotplate at 70° C. A PbI₂ solution in DMF (462 mg/mL) was 395 warmed to 70°C and then spin-coated at 2000 rpm for 60 s. The PbI₂ films were subsequently 396 annealed at 70°C for 30 min. CH₃NH₃PbI₃ films with small crystal size were prepared by 397 dipping the PbI₂-coated substrates in a 0.044 M CH₃NH₃I solution in isopropanol at 25°C for 398 2 minutes. Samples were finally rinsed in anhydrous IPA to remove unreacted CH₃NH₃I.

399 *Mesoporous Al₂O₃* (MP)

400 Unless otherwise specified, all materials were purchased from either Alfa Aesar or Sigma-401 Aldrich and used as received. The synthesis of CH_3NH_3I was performed using the same 402 previously reported method ³¹. The Al_2O_3 (Aldrich, 642991) nanoparticle paste used for 403 characterization measurements was formed by suspending ~ 45 nm particles in a solution of 404 ethanol, terpineol and ethyl cellulose.

Glass substrates were cleaned by sonication in ethanol for 10 mins. The mesoporous Al_2O_3 layer was formed by drop coating approximately 30 μ L/cm² of the nanoparticle solution on the glass substrates, which were then annealed at 500 °C for 30 mins. The perovskite precursor solution was formed by mixing stoichiometric amounts of CH₃NH₃I and PbI₂ in a combination of dimethyl sulfoxide (DMSO) and *N*-methyl-2-pyrrolidone (NMP) (7:3 vol) to form a (46 wt%) concentration solution. Approximately 30 μ L/cm² of the perovskite precursor solution was applied to the surface of the substrate prior to spin-coating. The films were spin-coated using a two-stage process: 1000 rpm for 5 s using an acceleration of 200 rpm/s, then 6000 rpm for 50 s using an acceleration of 6000 rpm/s. A nitrogen gas flow was introduced after 20 s of the second spin-coating step and sustained for a further 20 s. The perovskite films were then annealed on a hot-plate at 100 °C for a duration of 10 mins.

416 Mixed-cation Mixed-halide Perovskite Sample preparation

417 The organic cations were purchased from Dyesol; the lead compounds from TCI; 418 Dimethylformamide (DMF), Dimethyl Sulfoxide (DMSO) and cesium iodide from Sigma 419 Aldrich. The mixed cation lead mixed halide perovskite 420 (Cs_{0.05}(MA_{0.17}FA_{0.83})_{0.95}Pb(I_{0.83}Br_{0.17})₃) precursor solutions were prepared by dissolving 1 M 421 FAI, 1.1 M PbI₂, 0.2 M MABr, 0.2 M PbBr₂ in a 4:1 (v:v) mixture of anhydrous DMF:DMSO 422 and 1.5 M stock solution of CsI in DMSO was added to the above solution in a 5:95 423 (CsI:mixture) volume ratio. The perovskite solutions were then spin coated in a two-step 424 program at 1000 and 6000 rpm for 10 and 30 s respectively. During the second step, 50 ul of 425 chlorobenzene was dripped on the spinning substrate 10 s prior to the end. The perovskite film was annealed at 100° C for 30 min. 426

427 SEM imaging

The LPC samples were imaged using a Zeiss Merlin high-resolution scanning electron microscope with a beam energy of 3 kV. The SPC, SC and MP samples were imaged using a FEI Nova NanoSEM 450 scanning electron microscope unit with a beam energy of 15 kV.

431 Magneto-optical measurement

The magneto transmission measurements were performed combining long pulse magneticfield meausrements and 150 T short duration pulsed magnets.

434 For the long pulse measurements, the typical duration of the pulse is around 100 ms with the 435 magnetic field up to 66 T. For these measurements, the sample was placed in a liquid helium 436 cryostat. White light from a halogen lamp was used as the excitation source. The light emitted 437 from the lamp was coupled in a 200 µm-diameter multimode fiber, used to illuminate sample. 438 The transmitted light was collected by a 400 µm-diameter multimode fiber and guided to a 439 spectrometer equipped with a liquid nitrogen cooled CCD camera. The typical exposure time 440 was <3 ms, which ensured that the transmission spectra were acquired at an essentially 441 constant magnetic field value.

For very high magnetic field measurements (B<150 T), pulse having a typical duration < 10 μ s were generated by a single turn coil system with a bore diameter of 10 mm. A nonconducting helium-flow cryostat was located in the single turn coil. The sample was kept at a temperature of ~5 K. Magneto-transmission measurements were conducted by a using tunable Opical Parametric Oscillator pumped by a Ti:sapphire laser as the light source. A fast (100 MHz) silicon detector and a high speed digital oscilloscope was used for detection.

448 **Reference**

- D. Bi, W. Tress, M. I. Dar, P. Gao, J. Luo, C. Renevier, K. Schenk, A. Abate, F. Giordano, J. P. Correa Baena, J.-D. Decoppet, S. M. Zakeeruddin, M. K. Nazeeruddin, M. Grätzel and A.
 Hagfeldt, *Science Advances*, 2016, 2.
- 452 2. M. Saliba, T. Matsui, J.-Y. Seo, K. Domanski, J.-P. Correa-Baena, N. Mohammad K, S. M.
 453 Zakeeruddin, W. Tress, A. Abate, A. Hagfeldt and M. Gratzel, *Energy & Environmental* 454 *Science*, 2016, 9, 1989-1997.
- 455 3. A. Kojima, K. Teshima, Y. Shirai and T. Miyasaka, *Journal of the American Chemical* 456 *Society*, 2009, **131**, 6050-6051.
- 457 4. N. J. Jeon, J. H. Noh, W. S. Yang, Y. C. Kim, S. Ryu, J. Seo and S. I. Seok, *Nature*, 2015, 517, 476–480.
- 5. D. P. McMeekin, G. Sadoughi, W. Rehman, G. E. Eperon, M. Saliba, M. T. Hörantner, A. Haghighirad, N. Sakai, L. Korte, B. Rech, M. B. Johnston, L. M. Herz and H. J. Snaith, *Science*, 2016, **351**, 151-155.
- 462 6. V. D'Innocenzo, A. R. Srimath Kandada, M. De Bastiani, M. Gandini and A. Petrozza,
 463 *Journal of the American Chemical Society*, 2014, **136**, 17730–17733.
- 464 7. C. G. Bischak, E. M. Sanehira, J. T. Precht, J. M. Luther and N. S. Ginsberg, *Nano Letters*, 2015, 15, 4799-4807.
- 466 8. Z.-K. Tan, R. S. Moghaddam, M. L. Lai, P. Docampo, R. Higler, F. Deschler, M. Price, A.
 467 Sadhanala, L. M. Pazos, D. Credgington, F. Hanusch, T. Bein, H. J. Snaith and R. H. Friend,
 468 *Nat Nano*, 2014, 9, 687-692.
- M. Saliba, S. M. Wood, J. B. Patel, P. K. Nayak, J. Huang, J. A. Alexander-Webber, B.
 Wenger, S. D. Stranks, M. T. Hörantner, J. T.-W. Wang, R. J. Nicholas, L. M. Herz, M. B.
 Johnston, S. M. Morris, H. J. Snaith and M. K. Riede, *Advanced Materials*, 2016, 28, 923-929.
- 472 10. M. A. Green, A. Ho-Baillie and H. J. Snaith, *Nat Photon*, 2014, **8**, 506-514.
- 473 11. S. D. Stranks and H. J. Snaith, *Nature Nanotechnology*, 2015, **10**, 391-402.
- 474 12. T. Ishihara, Journal of Luminescence, 1994, 60-61, 269–274.
- 475 13. M. Hirasawa, T. Ishihara, T. Goto, K. Uchida and N. Miura, *Physica B: Condensed Matter*, 1994, 201, 427-430.
- 477 14. M. Hirasawa, T. Ishihara and T. Goto, *Journal of the Physical Society of Japan*, 1994, 63, 3870-3879.
- 479 15. M. E. Ziffer, J. C. Mohammed and D. S. Ginger, ACS Photonics, 2016, 3, 1060–1068.
- 480 16. K. Galkowski, A. Mitioglu, a. miyata, p. plochocka, o. Portugall, G. E. Eperon, J. T.-W.
 481 Wang, T. Stergiopoulos, S. D. Stranks, H. Snaith and R. J. Nicholas, *Energy & Environmental*482 *Science*, 2016, 9, 962-970.
- 483 17. A. Miyata, A. Mitioglu, P. Plochocka, O. Portugall, J. T.-W. Wang, S. D. Stranks, H. J. Snaith and R. J. Nicholas, *Nat Phys*, 2015, 11, 582–587.
- 485 18. A. M. Soufiani, F. Huang, P. Reece, R. Sheng, A. Ho-Baillie and M. A. Green, *Applied Physics Letters*, 2015, **107**, 231902.
- 487 19. Y. Fu, H. Zhu, C. C. Stoumpos, Q. Ding, J. Wang, M. G. Kanatzidis, X. Zhu and S. Jin, ACS
 488 Nano, 2016, 10, 7963-7972.
- 489 20. S. A. March, D. B. Riley, C. Clegg, D. Webber, X. Liu, M. Dobrowolska, J. K. Furdyna, I. G.
 490 Hill and K. C. Hall, *arXiv preprint arXiv:1602.05186*, 2016.
- 491 21. A. R. Srimath Kandada and A. Petrozza, *Accounts of Chemical Research*, 2016, **49**, 536–544.
- 492 22. G. Grancini, A. R. Srimath Kandada, J. M. Frost, A. J. Barker, M. De Bastiani, M. Gandini, S.
 493 Marras, G. Lanzani, A. Walsh and A. Petrozza, *Nat Photon*, 2015, 9, 695-701.
- 494 23. Q. Lin, A. Armin, R. C. R. Nagiri, P. L. Burn and P. Meredith, *Nature Photonics*, 2014, **9**, 106–112.
- 496 24. J. Even, L. Pedesseau and C. Katan, *The Journal of Physical Chemistry C*, 2014, **118**, 11566–
 497 11572.
- 498 25. V. D'Innocenzo, G. Grancini, M. J. P. Alcocer, A. R. S. Kandada, S. D. Stranks, M. M. Lee,
 499 G. Lanzani, H. J. Snaith and A. Petrozza, *Nat Commun*, 2014, 5.
- 500 26. L. M. Herz, Annual Review of Physical Chemistry, 2016, 67, 65-89.

- 501 27. J. Kim, J. S. Yun, X. Wen, A. M. Soufiani, C. F. J. Lau, B. Wilkinson, J. Seidel, M. A. Green,
 502 S. Huang and A. W. Y. Ho-Baillie, *The Journal of Physical Chemistry C*, 2016, **120**, 11262–
 503 11267.
- 504 28. J. Burschka, N. Pellet, S.-J. Moon, R. Humphry-Baker, P. Gao, M. K. Nazeeruddin and M.
 505 Gratzel, *Nature*, 2013, 499, 316-319.
- 506 29. W. S. Yang, J. H. Noh, N. J. Jeon, Y. C. Kim, S. Ryu, J. Seo and S. I. Seok, *Science*, 2015, 348, 1234-1237.
- 508 30. M. Liu, M. B. Johnston and H. J. Snaith, *Nature*, 2013, **501**, 395-398.
- 509 31. F. Huang, Y. Dkhissi, W. Huang, M. Xiao, I. Benesperi, S. Rubanov, Y. Zhu, X. Lin, L. Jiang
 510 and Y. Zhou, *Nano Energy*, 2014, **10**, 10-18.
- 32. W. Zhang, S. Pathak, N. Sakai, T. Stergiopoulos, P. K. Nayak, N. K. Noel, A. A. Haghighirad,
 V. M. Burlakov, D. W. deQuilettes, A. Sadhanala, W. Li, L. Wang, D. S. Ginger, R. H. Friend
 and H. J. Snaith, *Nat Commun*, 2015, 6.
- 514 33. C. F. Klingshirn, Semiconductor optics, Springer, 2007.
- 515 34. J. Tilchin, D. N. Dirin, G. I. Maikov, A. Sashchiuk, M. V. Kovalenko and E. Lifshitz, ACS
 516 Nano, 2016, 10, 6363-6371.
- 517 35. P. C. Makado and N. C. McGill, *Journal of Physics C: Solid State Physics*, 1986, **19**, 873.
- 518 36. G. Mahan, Proc. 1971 Antwerp Advanced Study Inst, 1972.
- 519 37. D. Li, G. Wang, H.-C. Cheng, C.-Y. Chen, H. Wu, Y. Liu, Y. Huang and X. Duan, *Nat* 520 *Commun*, 2016, **7**.
- 38. A. Osherov, E. M. Hutter, K. Galkowski, R. Brenes, D. K. Maude, R. J. Nicholas, P.
 522 Plochocka, V. Bulović, T. J. Savenije and S. D. Stranks, *Advanced Materials*, 2016, 28, 10757-10763.
- 524 39. K. Galkowski, A. A. Mitioglu, A. Surrente, Z. Yang, D. K. Maude, P. Kossacki, G. E. Eperon,
 525 J. T. W. Wang, H. J. Snaith, P. Plochocka and R. J. Nicholas, *Nanoscale*, 2017, 9, 3222.
- 526 40. M. R. Filip, G. E. Eperon, H. J. Snaith and F. Giustino, *Nature communications*, 2014, 5.
- M. A. Pérez-Osorio, R. L. Milot, M. R. Filip, J. B. Patel, L. M. Herz, M. B. Johnston and F. Giustino, *The Journal of Physical Chemistry C*, 2015, **119**, 25703-25718.
- A. M. A. Leguy, A. R. Goni, J. M. Frost, J. Skelton, F. Brivio, X. Rodriguez-Martinez, O. J.
 Weber, A. Pallipurath, M. I. Alonso, M. Campoy-Quiles, M. T. Weller, J. Nelson, A. Walsh and P. R. F. Barnes, *Physical Chemistry Chemical Physics*, 2016, 18, 27051-27066.
- 43. Y. Yamada, T. Nakamura, M. Endo, A. Wakamiya and Y. Kanemitsu, *IEEE Journal of Photovoltaics*.
- 534 44. C. Motta, P. Mandal and S. Sanvito, *Physical Review B*, 2016, **94**, 045202.
- M. B. Price, J. Butkus, T. C. Jellicoe, A. Sadhanala, A. Briane, J. E. Halpert, K. Broch, J. M.
 Hodgkiss, R. H. Friend and F. Deschler, *Nature Communications*, 2015, 6.
- 46. Y. Yang, D. P. Ostrowski, R. M. France, K. Zhu, J. van de Lagemaat, J. M. Luther and M. C.
 Beard, *Nat Photon*, 2015, **10**, 53–59.
- 539 47. N. Sestu, M. Cadelano, V. Sarritzu, F. Chen, D. Marongiu, R. Piras, M. Mainas, F. Quochi, M.
- Saba, A. Mura and G. Bongiovanni, *The Journal of Physical Chemistry Letters*, 2015, 6, 45664572.