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ITRF2008: an improved solution of the international terrestrial reference frame

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Abstract ITRF2008 is a refined version of the International Terrestrial Reference Frame based on reprocessed solutions of the four space geodetic techniques: VLBI, SLR, GPS and DORIS, spanning 29, 26, 12.5 and 16 years of observations, respectively. The input data used in its elaboration are time series (weekly from satellite techniques and 24-h sessionwise from VLBI) of station positions and daily Earth Orientation Parameters (EOPs). The ITRF2008 origin is defined in such a way that it has zero translations and translation rates with respect to the mean Earth center of mass, averaged by the SLR time series. Its scale is defined by nullifying the scale factor and its rate with respect to the mean of VLBI and SLR long-term solutions as obtained by stacking their respective time series. The scale agreement between these two technique solutions is estimated to be 1.05 ± 0.13 ppb at epoch 2005.0 and 0.049 ± 0.010 ppb/yr. The ITRF2008 orientation (at epoch 2005.0) and its rate are aligned to the ITRF2005 using 179 stations of high geodetic quality. An estimate of the origin components from ITRF2008 to ITRF2005 (both origins are defined by SLR) indicates differences at epoch 2005.0, namely: -0.5, -0.9 and -4.7 mm along X, Y and Z-axis, respectively. The translation rate differences between the two frames are zero for Y and Z, while we observe an X-translation rate of 0.3 mm/yr. The estimated formal errors of these parameters are 0.2 mm and 0.2 mm/yr, respectively. The high level of origin agreement between ITRF2008 and ITRF2005 is an indication of an imprecise ITRF2000 origin

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L. Métivier Institut de Physique du Globe de Paris, Universite Paris-Diderot, Batiment Lamarck, case 7011, 35 rue Helene Brion, 75013 Paris, France that exhibits a Z-translation drift of 1.8 mm/yr with respect to ITRF2005. An evaluation of the ITRF2008 origin accuracy based on the level of its agreement with ITRF2005 is believed to be at the level of 1 cm over the time-span of the SLR observations. Considering the level of scale consistency between VLBI and SLR, the ITRF2008 scale accuracy is evaluated to be at the level of 1.2 ppb (8 mm at the equator) over the common time-span of the observations of both techniques. Although the performance of the ITRF2008 is demonstrated to be higher than ITRF2005, future ITRF improvement resides in improving the consistency between local ties in co-location sites and space geodesy estimates.

Keywords Reference systems \cdot Reference frames \cdot ITRF \cdot Earth rotation

1 Introduction

Precisely determining satellite orbits or quantifying Earth rotation, tectonic plate motion or mean sea level rise and its variability in space and time fundamentally depend on the availability of a truly global Terrestrial Reference System (TRS) that only space geodesy is able to realize. Following the established terminology on reference systems (Kovalevsky et al. 1989), we distinguish between a TRS and its realization by a Terrestrial Reference Frame (TRF). The former is being ideal and conventional, with specified mathematical and physical properties, while the latter is constructed using space geodesy observations (hence with uncertainties) and is being accessible to the users through numerical values (e.g. positions as a function of time of a network of Earth crust-based points). The main physical and mathematical properties of a TRS (at the theoretical level) or of a TRF (at the realization level) are the origin, the scale, the orientation and their time evolution.



The International Terrestrial Reference System (ITRS), realized and maintained by the International Earth Rotation and Reference Systems Service (IERS) has been formally adopted and recommended for Earth science applications (IUGG 2007). For more details regarding the ITRS description and definition, the reader may refer to Chapter 4 of the IERS Conventions (Petit and Luzum 2010). The ITRS realization, through the International Terrestrial Reference Frame (ITRF) is regularly updated to take into account not only new accumulated data, but also improved analysis strategies applied by the analysis centers of the contributed techniques. Eleven ITRF versions were hence published, starting with the ITRF88 and ending with ITRF2008 described in this article. The space geodetic techniques that contribute to the ITRF construction are Very Long Baseline Interferometry (VLBI), Satellite Laser Ranging (SLR), Global Navigation Satellite Systems (GNSS) and Doppler Orbitography Radiopositioning Integrated by Satellite (DORIS). These techniques are organized as scientific services within the International Association of Geodesy (IAG) and known by the IERS as Technique Centers (TCs): the International VLBI Service (IVS), (Schlüter and Behrend 2007), the International Laser Ranging Service (ILRS), (Pearlman et al. 2002), the International GNSS Service, formerly the International GPS Service (IGS), (Dow et al. 2009) and the International DORIS Service (IDS), (Willis et al. 2010).

Although ITRF2005 (Altamimi et al. 2007) was an improved solution compared to past ITRF versions, in terms of internal consistency and robustness, it, however, shed light on some deficiencies of the four technique solutions as well as their co-locations. In preparation for the ITRF2008, the IAG technique services, together with their respective Analysis and Combination Centers invested considerable effort to improve their solutions. Without being exhaustive, we can cite that the reprocessed IGS solution involves new absolute phase center offsets and variations models for satellites and stations (Schmid et al. 2007) and a new tropospheric model (Boehm et al. 2006a, 2007); the reanalysis solution from IVS accounts for the mean pole tide correction following the IERS Conventions (Petit and Luzum 2010) and a more advanced troposphere modeling (Boehm et al. 2006b) and corrections for antenna thermal deformations (Nothnagel 2009); the improved ILRS solution takes into account new range bias values, a new tropospheric modeling (Mendes et al. 2002; Mendes and Pavlis 2004) and other station-dependent corrections, and for the first time the DORIS combined solution involves seven IDS Analysis Centers (Valette et al. 2010).

Each time an ITRF solution is prepared, one of the fundamental aspects of its formation is to ensure the optimal frame definition and its stability as a function of time. While the origin and the scale (having physical properties) are critical parameters of interest to Earth science applications, the orientation and its time variation (arbitrary and conventionally defined) are of importance to ensure the continuity of Earth rotation determination. Any bias or drift in these components will inevitably propagate into the geophysical results that depends on the usage of the ITRF, as for instance mean sea level assessment and its variability in space and time (Morel and Willis 2005; Beckley et al. 2007; Collilieux and Wöppelmann 2010). As it will be shown by the discussion of the ITRF2008 results, the current achievable accuracy of the ITRF scale is probably not better than 1 part-per-billion (ppb) and its temporal stability at the range of 0.05 ppb/yr. The total error budget includes systematic errors of not only VLBI and SLR, but also GPS that is crucially needed to tie the three other techniques together, as well as local tie errors.

The space geodetic technique solutions and local ties at co-location sites used to generate ITRF2008 are first presented in Sect. 2. Section 3 describes the combination strategy that has been applied to these data. The output results of this combination and ITRF2008 quality evaluation are described in Sect. 4. The main ITRF2008 results and performance are synthesized in the conclusion.

All the ITRF2008 files and results are available at the ITRF web site: http://itrf.ign.fr/ITRF_solutions/2008/

2 ITRF2008 input data

As for any ITRF solution, the ITRF2008 relies not only on space geodesy solutions, but also on local ties at co-location sites. In the following two sub-sections we describe the two sets of data used in the ITRF2008 construction.

2.1 Space geodesy solutions

We recall that the ITRF input time series solutions are provided on a weekly basis by the IAG International Services of satellite techniques: IGS, ILRS and IDS and on a daily (VLBI session-wise) basis by the IVS. Each per-technique time series is already a combination of the individual Analysis Center (AC) solutions of that technique. Table 1 summarizes the submitted solutions by the IAG services, specifying the time span, solution type and the constraints applied by the TCs as well as the estimated Earth Orientation Parameters (EOPs). For more details regarding the type of constraints applied by the techniques, and the minimum constraints concept in general, the reader may refer to Dermanis (2000, 2003); Sillard and Boucher (2001); Altamimi et al. (2002a, 2004) or to Chapt. 4 of the IERS Conventions (Petit and Luzum 2010). The submitted VLBI solution involves more than 4,000 session-wise SINEX files spanning the entire VLBI observation history (Bockmann et al. 2010). The SLR solution covers also its full observation history, and comprises fortnightly SINEX files, with polar motion and Length of Day (LOD) estimated each three days between 1983.0 and



Table 1 Summary of submitted solutions to ITRF2008

TC	Data-span	Solution type	Constraints	EOPs
IVS	1980.0–2009.0	Normal equation	None	Polar motion, rate, LOD, UT1-UTC
ILRS	1983.0-2009.0	Variance-covariance	Loose	Polar motion, LOD
IGS	1997.0-2009.5	Variance-covariance	Minimum	Polar motion, rate, LOD
IDS	1993.0-2009.0	Variance-covariance	Minimum	Polar motion, rate, LOD

1993.0 and weekly SINEX files with daily polar motion and LOD estimates afterwards (Pavlis et al. 2010). The GPS submitted solution represents a large part of the first reprocessed solution by the IGS and covers the time period 1997.0–2009.5 (Ferland 2010; Ferland and Piraszewski 2008). Note that a very small portion of GLONASS observations were used by some IGS ACs that contributed to the reprocessing effort. For the first time the DORIS contribution is a combined time series involving seven ACs and covers its full observation history, using data from all available satellites with onboard DORIS receiver, except Jason-2 (Valette et al. 2010).

The ITRF2008 network comprises 934 stations located at 580 sites, with 463 sites in the northern hemisphere and 117 in the southern hemisphere. The ITRF2008 combination involves 84 co-location sites where two or more technique instruments were or are currently operating and for which local ties are available. Figure 1 illustrates the full ITRF2008

network where we superimposed the VLBI, SLR and DORIS sites co-located with GPS. In fact all the 84 co-location sites comprise permanent GPS stations, except two sites: Dionysos (Greece) where DORIS and an old mobile SLR were co-located, and Richmond (Virginia, USA) where VLBI, SLR and DORIS systems were co-located.

2.2 Local ties in co-location sites

The local ties used in the ITRF2008 combination are provided in SINEX format with known measurement epochs, and 63% of them are available with full variance covariance information. Most of the local ties used in the ITRF2005 combination are used here with some updates, e.g., Tahiti (GPS, SLR, DORIS), Tsukuba (GPS, VLBI), Herstmonceux (GPS, SLR), Medicina and Noto (GPS, VLBI), Greenbelt (GPS, VLBI, SLR, DORIS), Maui/ Haleakala (GPS, SLR), San Fernando (GPS, SLR), Onsala (GPS, VLBI). Most of

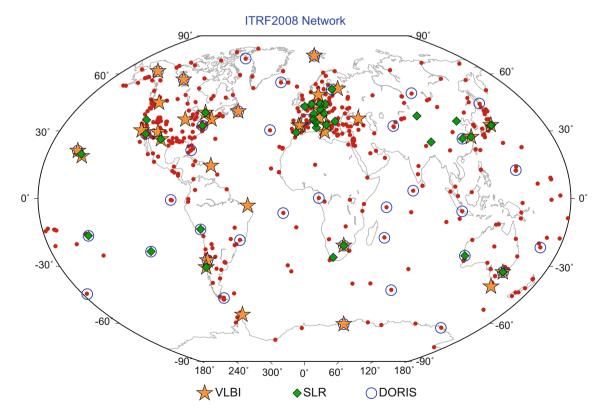


Fig. 1 ITRF2008 network highlighting VLBI, SLR and DORIS sites co-located with GPS



the local tie SINEX files were provided by the national agencies operating co-location sites, all sites in Australia (Johnston and Dawson 2004); three sites in Italy: Medicina, Noto (Sarti et al. 2004) and Matera (Bianco 2006); Wettzell and Concepcion (Schlüter et al. 2004a,b); Tsukuba (GSI 2010); Onsala (Haas 2010); and all the DORIS co-location sites were re-adjusted by the IGN survey department in order to generate full SINEX files, using the most recent surveys operated at these sites. All the local tie SINEX files used in the ITRF2008 combination are available at http://itrf.ign.fr/local_surveys.php.

Counting the number of co-locations between VLBI, SLR and DORIS, taken by pairs, we find eight VLBI-SLR, ten VLBI-DORIS and ten SLR-DORIS. These are very poor numbers of co-locations to allow a reliable combination of these three techniques alone. Therefore, the GPS is playing a major role in the ITRF combination by linking together the three other techniques (Altamimi and Collilieux 2009). We count in total 137 tie vectors between GPS and the three other techniques: 44 for VLBI, 48 for SLR and 45 for DORIS. 15 additional ties were also used between old and current DORIS beacon reference points in DORIS-only sites.

3 ITRF2008 data analysis

The CATREF combination model used for the ITRF combination is extensively described in previous publications and in particular (Altamimi et al. 2002b, 2007). For the ITRF2008 generation, we followed the same analysis strategy used for the ITRF2005 combination (Altamimi et al. 2007). We recall here that two steps are currently used for the ITRF construction: (1) stacking the individual time series to estimate a long-term solution per technique comprising station positions at a reference epoch, velocities and daily EOPs; and (2) combining the resulting long-term solutions of the four techniques together with the local ties in co-location sites. The main two equations of the combination model are given below for completeness. They involve a 14-parameter similarity transformation, station positions and velocities and EOPs and are written as

$$\begin{cases} X_{s}^{i} = X_{c}^{i} + (t_{s}^{i} - t_{0})\dot{X}_{c}^{i} \\ + T_{k} + D_{k}X_{c}^{i} + R_{k}X_{c}^{i} \\ + (t_{s}^{i} - t_{k})\left[\dot{T}_{k} + \dot{D}_{k}X_{c}^{i} + \dot{R}_{k}X_{c}^{i}\right] \\ \dot{X}_{s}^{i} = \dot{X}_{c}^{i} + \dot{T}_{k} + \dot{D}_{k}X_{c}^{i} + \dot{R}_{k}X_{c}^{i} \end{cases}$$

$$\begin{cases} x_{s}^{p} = x_{c}^{p} + R2_{k} \\ y_{s}^{p} = y_{c}^{p} + R1_{k} \\ UT_{s} = UT_{c} - \frac{1}{f}R3_{k} \\ \dot{x}_{s}^{p} = \dot{x}_{c}^{p} \\ \dot{y}_{s}^{p} = \dot{y}_{c}^{p} \end{cases}$$

$$(2)$$



where for each point i, X_s^i (at epoch t_s^i) and \dot{X}_s^i are positions and velocities of technique solution s and X_c^i (at epoch t_0) and \dot{X}_c^i are those of the combined solution c. For each individual frame k, as implicitly defined by solution s, D_k is the scale factor, T_k the translation vector and R_k the rotation matrix. The dotted parameters designate their derivatives with respect to time. The translation vector T_k is composed of three origin components, namely T_x , T_y , T_z and the rotation matrix of three small rotation parameters: R_x , R_y , R_z , following the three axes, respectively $X, Y, Z. t_k$ is a conventionally selected epoch of the seven transformation parameters. In addition to Eq. (1) involving station positions (and velocities), the EOPs are added by Eq. (2), following (Zhu and Mueller 1983), making use of pole coordinates x_s^p , y_s^p and universal time UT_s as well as their daily rates \dot{x}_s^p , \dot{y}_s^p and LOD_s . The link between the combined frame and the EOPs is ensured via the three rotation parameters appearing in the first three lines of Eq. (2).

Note that Eq. (1) uses the linearized form of the general similarity transformation formula, neglecting second and higher order terms. For more details regarding the mathematical derivation of Eq. (1), the reader may refer to Chapt. 4 of the IERS Conventions (Petit and Luzum 2010), and to Altamimi and Dermanis (2010).

The reader should note that in past publications, e.g., (Altamimi et al. 2007), we considered the three last lines of Eq. (2) as the time derivatives of the first three lines of the same equation, and so rotation rate terms have been added. However, the conventional EOP representation provided by the techniques is in the form of daily offset (at noon) and drift (or rate), while other types of EOP representation could also be considered, such as continuous piece-wise linear function where no EOP rates would appear. Therefore, the daily EOP rates are independent from the orientation time evolution of the reference frame, although they should be correlated with the daily EOP offset values. In case of time series stacking, the addition of such rotation rates has no impact on the results because they are not introduced in the normal equation. However, their addition in case of combining long-term solutions (which contain EOPs) yields biased frame rotation rates of the individual solutions with respect to (but without altering) the combined frame.

3.1 ITRF2008 time series analysis

The first line of Eq. 1 and the entire Eq. 2 are used to estimate long-term solutions for each technique, by accumulating (rigorously stacking) the individual technique time series of station positions and EOPs. In this process, the second line of Eq. 1 and the rates of the translation, scale and rotation parameters are not included because station velocities are not available/estimable at the weekly (daily) basis. Moreover, a precise definition of the reference frame associated with the

resulting long-term solution (comprising station positions at a reference epoch, station velocities and EOPs) has to be clearly specified. As transformation parameters are estimated between each weekly (or session-wise) frame and the long-term frame, it becomes necessary to define the longterm frame origin, scale, orientation and their time evolution and therefore to complete the rank deficiency of the constructed normal equation. It is essential that the longterm solutions be representative of the mean origin and mean scale information of the space geodesy techniques. The particular type of minimal constraints introduced in Altamimi et al. (2007) have been designed for such purpose and are consequently used here to preserve the long-term solution origin (for SLR and DORIS) and scale (for VLBI, SLR and DORIS). As the input GPS weekly frames have been explicitly expressed in the IGS05 (derived from ITRF2005), GPS long-term solution has been expressed with respect to ITRF2005 and was not used for the ITRF2008 origin and scale definition.

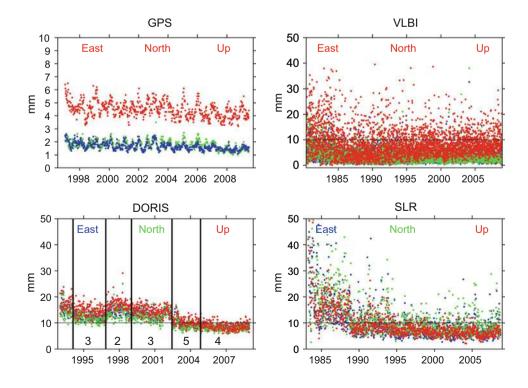
One of the main advantages of using time series as input data for the ITRF combination is the ability to identify station position discontinuities which are mostly due to earthquakes and equipment changes. There are in total 638 discontinuities; 550 of them are GPS-related. In the generation of the long-term technique solutions, we consider different station positions before and after each event. Nearby stations within the same site and multiple segments (in case of discontinuities) are constrained to have the same velocity, except

for sites where geophysical events such as large earthquakes occur.

As a by-product of the time series analysis we evaluate the internal precision (or repeatability) of each technique solution, through the computation of the Weighted Root Mean Scatter (WRMS) over the time series of weekly or daily residuals. Figure 2 displays the WRMS in East, North and Up components of the four technique solutions submitted to ITRF2008, with respect to the estimated long-term solutions. We computed the median values of these WRMS for the four technique solutions over the three components which are for GPS: 1.6, 1.7 and 4.4; for VLBI: 2.9, 3.0 and 7.4; for SLR: 7.9, 9.4, 7.9 and for DORIS: 11.9, 10.9 and 13.3 mm over the East, North and Up components, respectively. These median values are representative of the intrinsic precision (or repeatability) of each technique solution computed using the full time series, including the early less-precise observations, as well as all (less and well performing) stations. An inspection of the WRMS plotted in Fig. 2 indicates a precision improvement in the three components at recent dates for GPS, SLR and DORIS. It is interesting to note that for DORIS in particular the precision improves when observations from four to five satellites are used.

Seasonal signals can be noticed in the WRMS time series, and in particular for GPS. It has been shown that some of these variations, especially the annual, are related to loading effects (Altamimi and Collilieux 2009). Indeed, the Earth is continuously deforming due to non-tidal loading effects

Fig. 2 WRMS (internal precision) of weekly (or session-wise for VLBI) station position time series with respect to the long-term solution along the East, North and the Height components for GPS (top left); VLBI (top right); DORIS (bottom left) and SLR (bottom right). For DORIS, the number of satellites used is indicated on the plot





(Farrell 1972) but modeling of these phenomena is not yet integrated in the processing of geodetic observables for all the effects and all the techniques. It was decided not to incorporate the modeling of any non-tidal loading effects in the processing of geodetic data submitted by the TCs since not all available loading models have been fully evaluated. For instance, it was discovered recently that neglecting high-resolution topography changes induce errors in the atmospheric loading models (van Dam et al. 2010). However, it is worth noting that loading effects are not expected to affect ITRF frame parameters, especially the origin and scale (Collilieux et al. 2010), although the coordinates of the stations with few observations are impacted by this neglected effect. Therefore for the future, it becomes necessary that loading models, if they are accurate enough, be recommended by the IERS to be homogeneously and consistently used by the four techniques, since neglected effects, as the atmospheric tides, may be aliased into low frequencies signals (Tregoning and Watson 2009).

3.2 ITRF2008 multi-technique combination

The second step of the ITRF2008 analysis consists in combining the long-term solutions of the four technique solutions together with local ties at co-location sites. The ITRF2008 is specified by the following frame parameters:

- Origin: The ITRF2008 origin is defined in such a way that there are null translation parameters at epoch 2005.0 and null translation rates with respect to the ILRS SLR time series.
- Scale: The scale of the ITRF2008 is defined in such a
 way that there are null scale factor at epoch 2005.0 and
 null scale rate with respect to the mean scale and scale
 rate of VLBI and SLR time series.
- Orientation: The ITRF2008 orientation is defined in such a way that there are null rotation parameters at epoch 2005.0 and null rotation rates between ITRF2008 and ITRF2005. These two conditions are applied over a set of 179 reference stations located at 131 sites as illustrated by Fig. 3. The reference sites include 107 GPS, 27 VLBI, 15 SLR and 12 DORIS.

Special attention was devoted to the important issue of relative weighting of the space geodesy technique solutions and local ties. Several test combinations were performed by varying the weighting of all input data solutions in order to reach and adopt a statistically satisfactory ITRF2008 combination. The weighting of the individual solutions is operated by estimating an appropriate variance factor per data set and iterating as necessary. Because of the imbalance between the space geodesy technique solutions (which are global by nature)

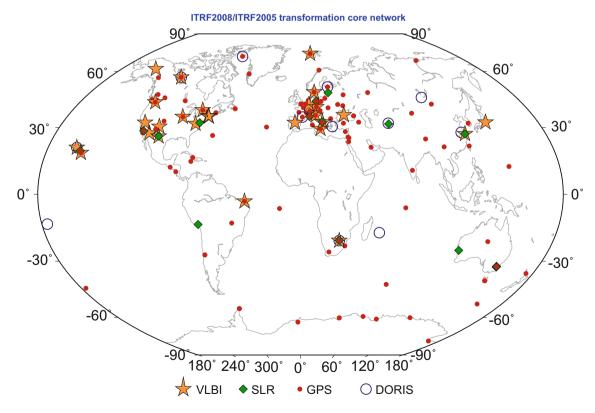


Fig. 3 Location of the core sites used in the estimation of the 14 transformation parameters between ITRF2008 and ITRF2005 and their orientation alignment



and the small, local-by-nature networks of co-location sites, it is very difficult to adequately use a mathematically or statistically prescribed method of variance component estimation as the degree of freedom or Helmert methods (Bähr et al. 2007). Therefore, the overall weighting process used in the ITRF2008 combination is a mixture of the degree of freedom method as described in (Altamimi et al. 2002b), Appendix A, Sect. A5, Eq. A16 and an empirical weighting. The long-term solutions of the four techniques are first properly weighted as a function of the variance factors (VF) obtained during the process of the individual stacking of the technique time series and computed as

$$VF = \frac{\sum_{s}^{S} v_{s}^{T} P_{s} v_{s}}{f} \tag{3}$$

where v_s is the vector of the station coordinates post fit residuals of solution s, P_s its weight matrix (the inverse of the variance-covariance matrix) and S is the total number of solutions included in the stacking. f is the number of degrees of freedom of the least squares adjustment, which is equal to the difference between the number of observations and the number of unknowns. Being a function of the residuals, normalized by the weight matrices, the VF is considered as a statistical indicator of the overall quality of the combination. Adequately weighting the individual solutions should lead to a variance factor close to unity. Therefore, an increase of the VF is a strong indication of the increase of the normalized residuals (raw residuals divided by their observation a priori errors), implying an inappropriate weighting of the individual solutions included in the combination or/and large outliers.

From this first step analysis, the following individual variance factors were obtained: 0.69 for GPS, 2.95 for SLR, 3.37 for VLBI and 2.20 for DORIS, respectively. As results of several ITRF2008 multi-technique test combinations, it was necessary to rescale the variance factors of VLBI and DORIS, by 1.2 and 2.5, respectively. In order to judge the appropriateness of this weighting scheme, we performed three main test combinations: a combination of the four technique velocity fields (downweighting station positions and local ties), and two combinations with and without EOPs. The obtained global variance factors of unit weight are in all cases close to unity, with variations less than 10%.

In the ITRF combination, local ties at co-location sites, provided in SINEX format, are used as observations with proper weighting as described in detail in Altamimi et al. (2002b). An empirical variance factor per local tie SINEX file is estimated during the combination process in such a way that (1) the normalized residual should not exceed a threshold of 3 and (2) the uncertainty per tie vector component should not be below 3 mm. The reasons for these two conditions are that (1) we believe a local tie between physically inaccessible instrumental measurement reference points is unlikely to be

precise to better than 3 mm, and (2) the agreement between local ties and space geodesy estimates are by far larger than 3 mm for most of the co-location sites and consequently (3) the local ties should be properly weighted in order to avoid contaminating the combined frame defining parameters by local tie and space geodesy discrepancies and at the same time to preserve consistency between individual technique solutions and ITRF2008. However, the discrepancies mean that either local ties or space geodesy estimates (or both) are imprecise or in error. One of the local survey difficulties is to precisely determine the eccentricity between the intersection of axes of VLBI or SLR telescopes and the ground physical markers. As an example, Sarti et al. (2009) estimated that the action of gravity on the structure of the Medicina (Italy) VLBI antenna biases the estimate of the instrument's reference point up to 1 cm.

A detailed analysis of the consistency between local ties and space geodesy estimates is discussed in Sect. 4.5.

4 ITRF2008 results

In the following subsections we discuss the main results of the ITRF2008 analysis as obtained from the two steps: time series stacking and multi-technique combination.

4.1 ITRF2008 origin

The submitted ILRS SLR solution was used to define the ITRF2008 origin, by fixing to zero (and consequently eliminating from the normal equation) the six parameters (translations and rates) of its corresponding long-term cumulative solution. Figure 4 illustrates the temporal behavior of the weekly SLR origin components with respect to ITRF2008, showing as expected (Dong et al. 1997) seasonal variations, but no shift or drift are visible. In order to evaluate the performance of the SLR origin determination, it is then important to compare the origin components between ITRF2008 and ITRF2005. An estimate of the translation components from ITRF2008 to ITRF2005 indicates differences at epoch 2005.0, namely -0.5, -0.9 and -4.7 mm along X, Y and Z-axes, respectively. The translation rate differences are zero for Y and Z, while we observe an X-translation rate of 0.3 mm/yr. These numbers suggest that the origin agreement between the two frames is at the level of or better than one centimeter over the entire time span of the SLR observations, and may be regarded as the level of the origin accuracy achievable today. We recall that there is a significant Z-translation rate of 1.8 mm/yr between ITRF2000 and ITRF2005 (Altamimi et al. 2007). A few studies (Greff-Lefftz 2000; Greff-Lefftz et al. 2010; Métivier et al. 2010) have shown that geophysical phenomena, such as present day ice melting or postglacial rebound, cannot lead to an



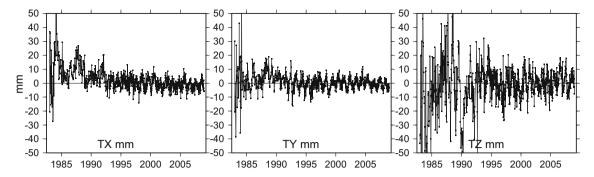


Fig. 4 Weekly translation components of the SLR ILRS solution with respect to ITRF2008, in millimeter along the X, Y and Z-axes: *left*, *middle* and *right*, respectively

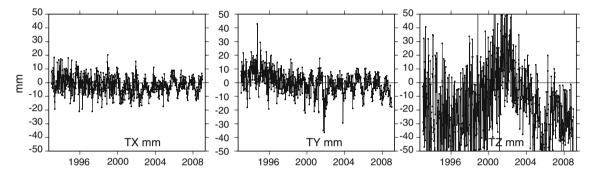


Fig. 5 Weekly translation components of the DORIS IDS-3 solution with respect to ITRF2008, in millimeter along the X, Y and Z-axes: left, middle and right, respectively

increase of the geocenter velocity along the Z-axis larger than 0.3 mm/yr in a decade. Given the fact that the ITRF2008 results show negligible translation rates with respect to ITRF2005, the Z-translation rate between ITRF2005 and ITRF2000 is most likely an indication of an imprecise origin determination of the ITRF2000 solution. This statement is also supported by the fact that both ITRF2005 and ITRF2008, being based on rigorous time series analysis, are by far more precise than ITRF2000.

While detailed analyses of the IDS submitted solution to ITRF2008 were published in (Altamimi and Collilieux 2010; Valette et al. 2010), for completeness Fig. 5 illustrates the temporal origin components of the DORIS IDS-3 solution with respect to ITRF2008. This plot shows a rather stable Tx, a slope change in Ty after 2002.0 and a very scattered Tz behavior.

4.2 ITRF2008 scale

The level of agreement of the scale and scale rate between VLBI and SLR solutions used in the ITRF2005 combination were 1.4 (± 0.11) ppb at epoch 2005.0 and 0.08 (± 0.01) ppb/yr, respectively. This low level of consistency was due to several factors, including uncorrected mean pole tide effect

of the VLBI and probably station range biases of the SLR, together with their poor co-locations and the degradation of their network with time.

As results from the final ITRF2008 combination we found that the scale and scale rate agreement between VLBI and SLR solutions are 1.05 (± 0.13) ppb at epoch 2005.0 and 0.049 (± 0.010) ppb/yr, respectively. This indicates an improvement compared to the past ITRF2005 results. This improvement is particularly due to the reprocessing effort of both IVS and ILRS. In addition, for some SLR co-located stations, we introduced breaks in the time series to account for significant discontinuities which improved the agreement of the estimated vertical velocities with GPS and consequently the scale rate agreement with VLBI. Propagating the scale discrepancy between SLR and VLBI at the start and end epochs of the time-span of VLBI data leads to a maximum discrepancy of 1.2 ppb (8 mm at the equator), which could be considered as the level of the scale accuracy achievable today. Therefore, defining the ITRF2008 scale to be in the middle of both technique solutions is the most appropriate choice that minimizes the scale impact for these two techniques when using the ITRF2008 products.

Figure 6 illustrates the daily scales of VLBI, and weekly scales of SLR and DORIS solutions with respect to ITRF2008 using the full time series. From that figure, we can distinguish



the level of agreement, scattering and temporal behavior of the scale factors of the three techniques. The estimated uncertainties (formal error at one sigma) of the scale and scale rates as results from the ITRF2008 combination are 0.10, 0.13, 0.19 ppb at epoch 2005.0 and 0.01, 0.01 and 0.03 ppb/yr for VLBI, SLR and DORIS, respectively.

The scale factor of the DORIS IDS-3 solution with respect to ITRF2008 is 0.70 ppb at epoch 2005.0, and so it is closer to VLBI than to SLR. As discussed in (Altamimi and Collilieux 2010), IDS-3 has a fairly stable scale (with negligible drift) during 1993–2002, and then displays uneven non-linear behavior afterward. This behavior is most likely to be related to the change of the number of satellites used.

4.3 Origin and scale seasonal variations

In order to evaluate the origin and scale seasonal variations visible in the time series plotted in Figs. 4, 5 and 6, we fitted annual amplitudes and phases (listed in Table 2) to the SLR and DORIS translation and scale components as well as to the scale of VLBI. For comparison, Table 2 reproduces also the values which were obtained using ITRF2005 data, published in (Altamimi et al. 2007). We restrict our discussion here to the quantification of the apparent geocenter motion and scale seasonal variation. The variations of the translation time series derived from the satellite techniques are expected to be close to the opposite of the non-linear part of the geocenter motion (Collilieux et al. 2009). The difference between the "true" geocenter motion and the translation

Fig. 6 Scale factors, in millimeters, of the VLBI, SLR and DORIS solutions with respect to ITRF2008

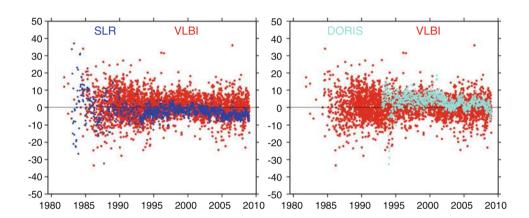


Table 2 Annual amplitude (A) and phase (ϕ) of the translation components and scale factor, according to the model $A\cos(\omega \cdot t - \phi)$ with t in decimal year

TC	TX		TY	TY		TZ		scale	
	A (mm)	ϕ (deg)	A (mm)	ϕ (deg)	A (mm)	ϕ (deg)	A (mm)	$\phi(\deg)$	
ITRF2008									
IVS	_	_	_	_	_	_	2.2	241	
\pm							0.1	3	
ILRS	2.6	222	3.1	135	5.5	202	0.6	255	
\pm	0.1	3	0.1	2	0.3	10	0.1	10	
IDS	3.9	327	4.6	160	4.4	115	0.3	206	
\pm	0.2	3	0.3	3	1.1	14	0.1	27	
ITRF2005 (A	Altamimi et al. 2	2007)							
IVS	_	_	_	_	_	_	2.7	220	
±							0.1	3	
ILRS	3.0	216	3.3	147	2.5	186	1.7	231	
±	0.2	4	0.2	3	0.4	11	0.2	7	
IDS-IGN	5.8	351	4.6	117	29.0	165	4.5	161	
±	0.5	5	0.5	6	2.5	5	0.3	4	
IDS-LCA	5.4	329	2.1	46	12.0	165	2.4	248	
±	0.5	5	0.5	13	1.8	9	0.4	9	



variations includes not only systematic errors related to orbit mis-modeling but also the aliasing of the loading effects (network effect) into the translation components (Collilieux et al. 2009; Lavallée et al. 2006). Scale annual variations also partly reflect this aliasing and can be interpreted as the mean vertical motion of the network (Collilieux et al. 2010). As all the geodetic networks are different, there is no reason to expect identical signals for all the techniques in the scale factor time series.

Comparing the annual amplitudes and phases of the SLR translations with those of ITRF2005 published in (Altamimi et al. 2007), we notice very similar estimates (i.e. within the uncertainties), except for the Z-translation amplitude which is two times larger here (5.5 mm). We checked that this Z-translation high amplitude is not significantly influenced by the pre 1993 SLR scattered data, being determined with only one satellite (Lageos 1). Contrariwise, while the SLR scale phase is statistically similar to ITRF2005 result, the amplitude here is smaller by 1 mm. Compared to ITRF2005 results, the VLBI scale factor annual phase is changed by about 20° and its annual amplitude is smaller by 0.5 mm, but this difference may be explained by the modeling of the antenna thermal deformation (Bockmann et al. 2010). Regarding DORIS, we note a more realistic annual amplitude in the Z component (4.4 mm), compared to the very large values exhibited by the two solutions used in the ITRF2005. This improvement is most likely to be related to the improvement of the solar radiation pressure treatment by some IDS analysis centers, leading to better estimate of the Z-geocenter (Gobindass et al. 2009). The DORIS X and Y amplitudes are higher by 1.5 mm than SLR, while only the phase in Y is closer to SLR. The DORIS scale annual amplitude is very small (0.3 mm), but the DORIS network has also the most homogeneous distribution over the Earth surface.

4.4 Impact of local ties on the ITRF2008 combination

Given the importance of local ties in the ITRF combination, we chose here to evaluate their impact on one of the most critical estimated parameters, namely the scale, being determined for the ITRF2008 by the mean of VLBI and SLR solutions. For this purpose, we performed different combination tests whose results are listed in Table 3. We evaluate the quality of these tests by discussing the obtained scale bias between VLBI and SLR solutions, as well as the variance factor of each adjustment. In reference to Eq. (3), an increase of the variance factor is the consequence of the increase of the normalized residuals (discrepancies between space geodesy estimates and local ties), implying an inappropriate relative weighting between space geodesy solutions and local ties.

Table 3 displays for each test, the scale factors (at epoch 2005.0) of SLR, GPS and DORIS with respect to VLBI solution, the global variance factor of each adjustment and the different cases of local ties handling. The first line of Table 3 lists the results of the final ITRF2008 combination where all available local tie SINEX files were properly weighted. Note that 63% of the local tie SINEX files are available with full variance-covariance information. The second combination is similar to the ITRF2008, but where the EOPs are not included and involves local ties vectors instead of SINEX files, weighted by their diagonal terms only. In addition, the epochs of these tie vectors were assumed to be all the same, namely 2005.0. This combination shows slight changes of the estimated scale factors of SLR, GPS and DORIS with an increase of their uncertainties as well as the variance factor. The next combination test (Case-1) involves a selection of good ties that agree with space geodesy estimates to better than 1 cm (for the worst cases), but with a uniform weighting of 1 mm over all ties and components. Although the increase of the variance factor to 4.0 indicating that the weight given to

Table 3 Scale factors (at epoch 2005.0) from VLBI long-term solution to SLR, DORIS and GPS using different weighting and selection of local ties at co-location sites

Case	SLR	DORIS	GPS	VF^a	Tie handling
ITRF2008	-1.05	0.18	0.67	0.90	All tie SINEX files, weighted
	± 0.13	± 0.20	± 0.10		
ITRF2008	-1.11	0.04	0.41	1.66	All tie vectors, weighted
without EOPs	± 0.26	± 0.31	± 0.19		
Case-1	-1.20	-0.31	0.49	4.00	Selection of local ties, $\sigma = 1 \text{ mm}$
	± 0.16	± 0.34	± 0.14		
Case-2	-1.02	-0.30	0.52	31.70	Selection of local ties, $\sigma = 0.1 \text{ mm}$
	± 0.31	± 0.84	± 0.31		
Case-3	0.0^{b}	0.48	1.28	4.28	Same as Case-1
		±0.32	± 0.08		

a Variance factor

^b parameter eliminated from the normal equation; assumed to be equal to VLBI scale



the local ties is too optimistic, this combination test increases the scale discrepancy between SLR and VLBI. The Case-2 test is the same as the previous one, but the local ties were given more weight by a factor of 10. It yields a very large variance factor and still shows a scale discrepancy of 1.02 ppb between SLR and VLBI solutions. The large variance factor obtained here is a clear indication that the heavy weight given to local ties increases considerably the normalized residuals, reflecting the discrepancies between space geodesy estimates and local ties that completely dominate this combination. The last test combination (Case-3) listed in Table 3 involves the same tie selection and weighting (1 mm) as Case-1, but where the SLR scale parameter was eliminated from the normal equation system (i.e. assuming zero scale difference with VLBI), this combination not only increases the variance factor, but it also insidiously transfers the SLR and VLBI scale discrepancy to GPS and DORIS. These test examples demonstrate that the adopted ITRF2008 final combination is the most statistically satisfactory combination. In addition, they show that the level of the scale agreement between VLBI and SLR solutions is not better than 1 ppb.

4.5 Consistency between local ties and space geodesy estimates

One of the most important by-products of the ITRF2008 combination is the assessment of the level of agreement between local ties and space geodesy estimates, through the availability of the post-fit residuals at co-location sites. In case of large discrepancies, discriminating between local ties and space geodesy estimates is a very delicate exercise, because the reasons for these discrepancies could be due to errors in local ties, in space geodesy estimates or in both. However, quantifying the level of agreement between the two ensembles is very critical for further investigation and hopefully for identifying the error sources. At co-location sites, not only station position residuals are computed, but also velocity residuals. Therefore, in order to take into account velocity disagreements between the technique solutions, it is more effective to compute the tie discrepancies at their measurement epochs. In order to identify the most performing co-location sites, we list in Table 4 the tie vectors where the discrepancies are less than 6 mm (corresponding to the level of scale consistency between VLBI and SLR solutions) in all three components: North, East and Up. As the GPS is playing the major role of connecting the three techniques together, the vectors listed in Table 4 are from GPS to other technique reference markers. With some exceptions, the geodetic instruments at the co-location sites listed in this table are still in operation in 2010. If we count the percentage of these sites listed in Table 4 over the total currently operating co-locations, we find approximately: 47, 43 and 34% for GPS-VLBI, GPS-SLR and GPS-DORIS, respectively.

However, using this sub-set of local ties only, would result in a non-optimal combination. Indeed, a test combination that was performed involving these co-location sites only yielded an increase of the uncertainties of the estimated parameters by a factor of 3.4, compared to the results of the ITRF2008 combination. In particular, the obtained scale factor between VLBI and SLR is 0.83 ppb, but with an increase of its uncertainty: ± 0.44 versus ± 0.13 ppb. In addition, we found the following percentages of co-location sites where tie discrepancies are larger than 10 mm: 29, 28 and 54% for GPS-VLBI, GPS-SLR and GPS-DORIS, respectively. Rejecting these co-location sites from the ITRF2008 combination increases the uncertainties of the estimated parameters by a factor of 2. Table 5 summarizes the tie discrepancy percentages following three categories: less than 6 mm, between 6 and 10 mm and larger than 10 mm.

The drawback of the GPS being the connecting technique is that any intrinsic GPS error would be transferred to the ITRF2008 estimated parameters. Therefore, properly weighting the local ties as a function of their agreement with space geodesy estimates is fundamental in order to ensure consistency between ITRF2008 results and individual technique solutions. Inspecting ITRF2008 local tie residuals, we found that some large discrepancies in the height component might be related to un-calibrated radomes covering the GPS antennas (Jim Ray and Ralf Schmid, personal communication). Examples of these cases include for GPS-VLBI co-locations: Onsala (Sweden), Tidbinbilla (Australia), Santiago (Chile), Fort Davis (USA), and for GPS-SLR co-locations: Greenbelt (USA), Yarragadee (Australia). Another patent example of the radome effect is the case of Forteleza (Brazil) GPS-VLBI co-location site. Using the original local tie values obtained from the adjustment of the survey conducted in 1994 exhibited a tie discrepancy of 19 mm in the height component. Correction was then applied for the apparent height shift due to the neglected effect of the conical radome covering the GPS antenna as reported by Ray et al. (2007) which reduced the discrepancy to 2.3 mm (see Table 4, line "FORT"). However, there are sites with GPS radomes where the discrepancies are within the tie uncertainties, such as CRO1 (Saint-Croix, Virgin Islands, USA) and NYA1 (Ny-Alesund, Norway) as listed in Table 4.

There most probably are other technique-specific errors related to the mis-modeling of the instrumental measurement reference points, not only for GPS (Schmid et al. 2007), but also for the other techniques. Sarti et al. (2010) for instance evaluated the impact of signal path variations caused by VLBI antenna gravitational deformations for Medicina and Noto (Italy) telescopes. They found that applying an elevation-dependent model of the gravitational deformations in VLBI data analysis for these two antennas shift the reference point positions upward by 8.9 and 6.7 mm, respectively. Applying this kind of elevation-dependent model to all VLBI



Table 4 ITRF2008 tie discrepancies less than 6 mm at tie epochs in (E)ast, (N)orth and (U)p: Residuals = Space geodesy - terrestrial tie (in mm)

GRAS 10002M006 7835 10002S001 1.3 -3.6 -5.6 99:284 SLR GRAS 10002M006 7845 10002S002 -1.1 -1.2 -0.6 99:284 SLR INSA 1003M009 TLSA 1000S001 -1.2 0.7 2.3 4.34 DORIS NYA1 10317M003 SPIB 10317S004 -1.2 5.7 2.2 3:228 DORIS NYA1 10317M003 SPIB 10317S005 -1.9 -0.8 -0.9 3:228 VLBI ONSA 10402M004 7231 10402S002 5.4 -0.6 -5.2 2:193 VLBI GRAZ 11001M002 7839 1101S002 1.4 -0.6 -5.9 92:319 VLBI GRAZ 1229M003 SAKA 12329S001 -0.4 -2.3 -2.9 4:231 DORIS YSSK 12329M003 SAKA 12329S001 -0.4 -2.3 -2.9 4:231 DORIS YSSK 12329M003 SAKA 12329S001 -0.4 -2.3 -2.9 4:231 DORIS YSSK 12329M003 SAKA 12329S001 -1.2 -3.	CODE DOMES #	CODE DOMES #	East	North	Up	Epoch	Technique
TLSE 10003M009 TLSA 10003S001 -1.2 0.7 2.3 4:34 DORIS NYA1 10317M003 SPIB 10317S004 -1.2 5.7 2.2 3:228 DORIS NYA1 10317M003 7331 10317S003 2.9 -2.7 1.2 3:228 VLBI ONSA 10402M004 7213 10402S002 5.4 -0.6 -5.2 2:193 VLBI GNZA 11001M002 7839 11001S002 1.4 -0.6 -5.2 2:193 VLBI BOR1 1220SM002 7811 1220SS001 0.5 1.7 -2.9 94:10 SLR YSSK 12329M003 SAKA 12329S001 -0.4 -2.3 -2.9 4:231 DORIS YSSK 12329M003 SAKA 12329S001 -0.4 -2.3 -2.9 4:231 DORIS YSSK 12329M003 SAKA 12329S001 -1.2 -3.1 -2.4 8:177 SLR HEES 13212M007 7840 13212S001 -1.2 -3.1 -2.4 8:177 SLR WTMM 14001M004 7810 14001S007 -1.9 -2.2 <td>GRAS 10002M006</td> <td>7835 10002S001</td> <td>1.3</td> <td>-3.6</td> <td>-5.6</td> <td>99:284</td> <td>SLR</td>	GRAS 10002M006	7835 10002S001	1.3	-3.6	-5.6	99:284	SLR
NYAI 10317M003 SPIB 10317S004 −1.2 5.7 2.2 3:228 DORIS NYAI 10317M003 SPIB 10317S003 2.9 −2.7 1.2 3:228 VLBI NYAI 10317M003 7331 10317S003 2.9 −2.7 1.2 3:228 VLBI ONSA 10402M004 7213 10402S002 5.4 −0.6 −5.9 92:319 SLR BOR1 1220SM002 7831 1020S001 0.5 1.7 −2.9 94:10 SLR YSSK 12329M003 SAKB 12329S001 −0.4 −2.3 −2.9 94:231 DORIS YSSK 12329M003 SAKB 12329S001 −0.4 −2.3 −2.9 4:231 DORIS YSSK 12329M003 SAKB 12329S001 −1.2 −3.1 −2.4 4:177 SLR YSSK 12329M003 SAKB 12329S001 −1.2 −3.1 −2.4 4:177 SLR ZIMM 14001M004 7810 14001S007 −1.9 −2.2 −5.2 96:95 SLR WTZZ 14212IM007 7490 1330303000 −1.2 −3.1	GRAS 10002M006	7845 10002S002	-1.1	-1.2	-0.6	99:284	SLR
NYAI 10317M003 SPJB 10317S005 −1.9 −0.8 −0.9 3:228 DORIS NYAI 10317M003 7331 10317S003 2.9 −2.7 1.2 3:228 VLBI ONSA 10402M004 7213 10402S002 5.4 −0.6 −5.2 2:193 VLBI GRAZ 11001M002 7839 11001S002 1.4 −0.6 −5.9 92:319 SLR BOR1 1220SM002 7811 1220SS001 −0.4 −2.3 −2.9 44:20 DORIS YSSK 12329M003 SAKB 12329S002 −2.3 −5.0 −2.3 42:31 DORIS YSSK 12329M003 SAKB 12329S002 −3.3 −4.7 −1.6 4:300 SLR MATE 12734M008 7941 12734S008 −3.7 −4.7 −1.6 4:300 SLR HERS 13212M007 7840 14201S001 −1.2 −3.1 −2.4 8:177 SLR WTZZ 14201M014 7224 14201S004 −1.4 −4.1 −0.3 ≥2:266 VLBI BIFS 21601M001 7249 21601S004 −0.6 <td< td=""><td>TLSE 10003M009</td><td>TLSA 10003S001</td><td>-1.2</td><td>0.7</td><td>2.3</td><td>4:34</td><td>DORIS</td></td<>	TLSE 10003M009	TLSA 10003S001	-1.2	0.7	2.3	4:34	DORIS
NYA1 10317M003 7331 10317S003 2.9 −2.7 1.2 3:228 VLBI ONSA 10402M004 7213 10402S002 5.4 −0.6 −5.2 2:193 VLBI GRAZ 11001M002 7839 1101S002 1.4 −0.6 −5.9 92:319 SLR BOR1 12205M002 7811 12205S001 0.5 1.7 −2.9 94:10 SLR YSSK 12329M003 SAKA 12329S001 −0.4 −2.3 −2.9 4:231 DORIS MATE 12734M008 7941 127345008 −3.7 −4.7 −1.6 4:300 SLR HERS 13212M007 7840 13212S001 −1.2 −3.1 −2.4 8:177 SLR YEZZ 14201M014 7224 14201S004 −1.4 −4.1 −0.3 −2.26 5.9 3:171 SLR WTZS 14201M014 7249 21601S004 0.6 −2.2 −5.9 96:95 SLR WTZS 14201M014 7249 12601S004 0.6 −2.2 −5.9 9:3171 SLR SISS 21730S007 −2.9 <	NYA1 10317M003	SPIB 10317S004	-1.2	5.7	2.2	3:228	DORIS
ONSA 10402M004 7213 10402S002 5.4 —0.6 —5.2 2:193 VLBI GRAZ 11001M002 7839 11001S002 1.4 —0.6 —5.9 92:319 SLR BOR1 12205M002 7811 12205S001 0.5 1.7 —2.9 94:10 SLR YSSK 12329M003 SAKA 12329S001 —0.4 —2.3 —2.9 4:231 DORIS YSSK 12329M003 SAKB 12329S002 2.3 —5.0 —2.3 4:231 DORIS MATE 12734M008 7941 12734S008 —3.7 —4.7 —1.6 4:300 SLR HERS 13212M007 7840 13212S001 —1.2 —3.1 —2.4 8:177 SLR ZIMM 14001M004 7810 14001S007 —1.9 —2.2 —5.2 96:95 SLR WTZZ 14201M014 7224 14201S004 —6.6 —2.2 —5.2 96:95 SLR BJFS 21601M001 7249 21601S004 —6.6 —2.2 —5.9 —3:171 SLR HRAO 30302M0404 7251 30302M003 —2.9 —0.8	NYA1 10317M003	SPJB 10317S005	-1.9	-0.8	-0.9	3:228	DORIS
GRAZ 11001M002 7839 11001S002 1.4 -0.6 -5.9 92:319 SLR BORI 1220SM002 7811 1220SS001 0.5 1.7 -2.9 94:10 SLR YSSK 12329M003 SAKA 12329S002 2.3 -2.3 -2.9 4:231 DORIS YSSK 12329M003 SAKB 12329S002 2.3 -5.0 -2.3 4:231 DORIS MATE 12734M008 7941 12734S008 -3.7 -4.7 -1.6 4:300 SLR HERS 13212M007 7840 13212S001 -1.2 -3.1 -2.4 8:177 SLR ZIMM 14001M004 7810 14001S007 -1.9 -2.2 -5.2 96:95 SLR WTZZ 14201M014 7224 14201S004 -1.4 -4.1 -0.3 2:266 VLBI BJFS 21601M001 7249 21601S004 0.6 2.2 5.9 3:171 SLR TSKB 21730S005 7345 21730S007 -2.9 -0.8 -3.0 8:16 VLBI HRAO 30302M004 7501 3030ZS006 5.4 -1.0	NYA1 10317M003	7331 10317S003	2.9	-2.7	1.2	3:228	VLBI
BORI 12205M002 7811 12205S001 0.5 1.7 -2.9 94:10 SLR YSSK 12329M003 SAKA 12329S001 -0.4 -2.3 -2.9 4:231 DORIS YSSK 12329M003 SAKB 12329S002 2.3 -5.0 -2.3 4:231 DORIS MATE 12734M008 7941 12734S008 -3.7 -4.7 -1.6 4:300 SLR HERS 13212M007 7840 13212S001 -1.2 -3.1 -2.4 8:177 SLR ZIMM 14001M004 7810 140018007 -1.9 -2.2 -5.2 96:95 SLR WTZZ 14201M010 7224 142018004 -1.4 -4.1 -0.3 2:266 VLBI BJFS 21601M001 7249 216018004 0.6 2.2 5.9 3:171 SLR TSKB 21730S005 7345 21730S007 -2.9 -0.8 -3.0 8:16 VLBI HRAO 30302M004 HBKB 30302S006 5.4 -1.0 -4.2 3:214 DORIS NKLG 32809M002 LIBB 32809S003 2.9 0.2	ONSA 10402M004	7213 10402S002	5.4	-0.6	-5.2	2:193	VLBI
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YSSK 12329M003 SAKB 12329S002 2.3 -5.0 -2.3 4:231 DORIS MATE 12734M008 7941 12734S008 -3.7 -4.7 -1.6 4:300 SLR HERS 13212M007 7840 13212S001 -1.2 -3.1 -2.4 8:177 SLR ZIMM 14001M004 7810 14001S007 -1.9 -2.2 -5.2 96:95 SLR WTZZ 14201M014 7224 14201S004 -1.4 -4.1 -0.3 2:266 VLBI BJES 21601M001 7249 21601S004 -0.6 2.2 5.9 3:171 SLR TSKB 21730S005 7345 21730S007 -2.9 -0.8 -3.0 8:16 VLBI HRAO 30302M004 7501 30302M003 -2.2 2.6 3.5 3:214 VLBI HRAO 30302M004 HBKB 30302S006 5.4 -1.0 -4.2 3:214 DORIS NKLG 32809M002 LIBB 32809S003 2.9 0.2 2.7 99:36 DORIS STIO 40101M001 7625 40127M001 -3.9 -3.9 <td>BOR1 12205M002</td> <td>7811 12205S001</td> <td>0.5</td> <td>1.7</td> <td>-2.9</td> <td>94:10</td> <td>SLR</td>	BOR1 12205M002	7811 12205S001	0.5	1.7	-2.9	94:10	SLR
MATE 12734M008 7941 12734S008 -3.7 -4.7 -1.6 4:300 SLR HERS 13212M007 7840 13212S001 -1.2 -3.1 -2.4 8:177 SLR ZIMM 14001M004 7810 14001S007 -1.9 -2.2 -5.2 96:95 SLR WTZZ 14201M014 7224 14201S004 -1.4 -4.1 -0.3 2:66 VLBI BJFS 21601M001 7249 21601S004 0.6 2.2 5.9 3:171 SLR TSKB 21730S005 7345 21730S007 -2.9 -0.8 -3.0 8:16 VLBI HRAO 30302M004 7232 30302S001 -1.6 3.2 2.0 3:214 VLBI HRAO 30302M004 HBKB 30302S006 5.4 -1.0 -4.2 3:214 DORIS NKLG 32809M002 LIBB 32809S003 2.9 0.2 2.7 99:36 DORIS YELL 40127M003 7255 40101M003 0.3 0.1 1.6 99:269 VLBI YELL 40127M003 7254 40456S001 -3.1 -2.8	YSSK 12329M003	SAKA 12329S001	-0.4	-2.3	-2.9	4:231	DORIS
HERS 13212M007	YSSK 12329M003	SAKB 12329S002	2.3	-5.0	-2.3	4:231	DORIS
ZIMM 14001M004 7810 14001S007 -1.9 -2.2 -5.2 96.95 SLR WTZZ 14201M014 7224 14201S004 -1.4 -4.1 -0.3 2:266 VLB1 BJFS 21601M001 7249 21601S004 0.6 2.2 5.9 3:171 SLR TSKB 21730S005 7345 21730S007 -2.9 -0.8 -3.0 8:16 VLB1 HRAO 30302M004 7233 30302S001 -1.6 3.2 2.0 3:214 VLB1 HRAO 30302M004 7501 30302M003 -2.2 2.6 3.5 3:214 DORIS HRAO 30302M004 HBKB 30302S006 5.4 -1.0 -4.2 3:214 DORIS NKLG 32809M002 LIBB 32809S003 2.9 0.2 2.7 99.36 DORIS STJO 40101M001 7625 40101M003 0.3 0.1 1.6 99:269 VLB1 YELL 40127M003 7285 40127M001 -3.9 -3.9 -1.0 1:285 VLB1 PIE1 40456M001 7234 40456S001 -3.1 -2.8 -0.3 92:336 VLB1 NLB 40465M001 7612 4046SS001 -3.4 -2.0 -5.2 93:64 VLB1 MKEA 40477M001 7617 40477S001 -3.7 -1.1 0.8 96:221 VLB1 MONP 40497M004 MONB 40497S008 3.1 -4.7 -0.8 5:335 DORIS MONP 40497M004 MONB 40497S008 3.1 -4.7 -0.8 5:335 DORIS MONP 40497M004 T1040497M001 3.2 -2.5 5.7 99:280 SLR FORT 41602M001 7297 41602S001 -0.6 -4.0 2.3 93:264 VLB1 GLPS 42005M002 SCRB 42005S001 -2.8 -1.0 -1.1 5:92 DORIS CRO1 43201M001 7615 43201S001 -1.5 2.2 1.6 94:16 VLB1 HOB2 50116M004 7242 50116S002 4.7 -3.0 -2.0 2:81 VLB1 HOB2 50116M004 7242 50116S002 4.7 -3.0 -2.0 2:81 VLB1 KERG 91201M002 KERB 91201S003 -3.2 -2.2 2.0 7:101 DORIS SYOG 66006S002 7342 66006S004 0.8 3.8 0.6 0:1 VLB1 ERRG 91201M000 ADFB 91501S003 -5.0 -3.0 3.3 8:39 DORIS DUM1 91501M001 ADFB 91501S003 -5.0 -3.0 3.3 8:39 DORIS THTI 92201M009 PAPB 92201S007 -1.3 1.7 4.4 7:278 DORIS	MATE 12734M008	7941 12734S008	-3.7	-4.7	-1.6	4:300	SLR
WTZZ 14201M014 7224 14201S004 -1.4 -4.1 -0.3 2:266 VLBI BJFS 21601M001 7249 21601S004 0.6 2.2 5.9 3:171 SLR TSKB 21730S005 7345 21730S007 -2.9 -0.8 -3.0 8:16 VLBI HRAO 30302M004 7232 30302S001 -1.6 3.2 2.0 3:214 VLBI HRAO 30302M004 7501 30302M003 -2.2 2.6 3.5 3:214 DORIS NKLG 32809M002 LIBB 32809S003 2.9 0.2 2.7 99:36 DORIS STJO 40101M001 7625 40101M003 0.3 0.1 1.6 99:269 VLBI YELL 40127M003 7285 40127M001 -3.9 -3.9 -1.0 1:285 VLBI YELL 40127M003 7285 40127M001 -3.1 -2.8 -0.3 92:336 VLBI MILB 40455M01 7612 40455S001 -3.1 -2.8 -0.3 92:336 VLBI MKEA 40477M001 7617 40477S001 -3.7 -1.1 <td>HERS 13212M007</td> <td>7840 13212S001</td> <td>-1.2</td> <td>-3.1</td> <td>-2.4</td> <td>8:177</td> <td>SLR</td>	HERS 13212M007	7840 13212S001	-1.2	-3.1	-2.4	8:177	SLR
BJFS 21601M001 7249 21601S004 0.6 2.2 5.9 3:171 SLR TSKB 21730S005 7345 21730S007 -2.9 -0.8 -3.0 8:16 VLBI HRAO 30302M004 7232 30302S001 -1.6 3.2 2.0 3:214 VLBI HRAO 30302M004 7501 30302M003 -2.2 2.6 3.5 3:214 DORIS NKLG 32809M002 LIBB 32809S003 2.9 0.2 2.7 99:36 DORIS STJO 40101M001 7625 40101M003 0.3 0.1 1.6 99:269 VLBI YELL 40127M003 7285 40127M001 -3.9 -3.9 -1.0 1:285 VLBI YELL 40127M003 7285 40127M001 -3.1 -2.8 -0.3 92:336 VLBI NLIB 40465M001 7612 40456S001 -3.1 -2.8 -0.3 92:336 VLBI MKEA 40477M001 7617 40477S001 -3.7 -1.1 0.8 96:221 VLBI MONP 40497M004 MONB 40497S008 3.1 -4.7 <td>ZIMM 14001M004</td> <td>7810 14001S007</td> <td>-1.9</td> <td>-2.2</td> <td>-5.2</td> <td>96:95</td> <td>SLR</td>	ZIMM 14001M004	7810 14001S007	-1.9	-2.2	-5.2	96:95	SLR
TSKB 21730S005 7345 21730S007 -2.9 -0.8 -3.0 8:16 VLBI HRAO 30302M004 7232 30302S001 -1.6 3.2 2.0 3:214 VLBI HRAO 30302M004 7501 30302M003 -2.2 2.6 3.5 3:214 SLR HRAO 30302M004 HBKB 30302S006 5.4 -1.0 -4.2 3:214 DORIS NKLG 32809M002 LIBB 32809S003 2.9 0.2 2.7 99:36 DORIS STJO 40101M001 7625 40101M003 0.3 0.1 1.6 99:269 VLBI YELL 40127M003 7285 40127M001 -3.9 -3.9 -1.0 1:285 VLBI PIE1 40456M001 7234 40456S001 -3.1 -2.8 -0.3 92:336 VLBI NLIB 40465M001 7612 40465S001 -3.4 -2.0 -5.2 93:64 VLBI MKEA 40477M001 7617 40477S001 -3.7 -1.1 0.8 96:221 VLBI MONP 40497M004 MONB 40497S008 3.1 -4.7 <td>WTZZ 14201M014</td> <td>7224 14201S004</td> <td>-1.4</td> <td>-4.1</td> <td>-0.3</td> <td>2:266</td> <td>VLBI</td>	WTZZ 14201M014	7224 14201S004	-1.4	-4.1	-0.3	2:266	VLBI
HRAO 30302M004 7232 30302S001 -1.6 3.2 2.0 3:214 VLBI HRAO 30302M004 7501 30302M003 -2.2 2.6 3.5 3:214 SLR HRAO 30302M004 HBKB 30302S006 5.4 -1.0 -4.2 3:214 DORIS NKLG 32809M002 LIBB 32809S003 2.9 0.2 2.7 99:36 DORIS STIO 40101M001 7625 40101M003 0.3 0.1 1.6 99:269 VLBI YELL 40127M003 7285 40127M001 -3.9 -3.9 -1.0 1:285 VLBI PIE1 40456M001 7234 40456S001 -3.1 -2.8 -0.3 92:336 VLBI NLIB 40465M001 7612 40465S001 -3.4 -2.0 -5.2 93:64 VLBI MKEA 40477M001 7617 40477S001 -3.7 -1.1 0.8 96:221 VLBI MONP 40497M004 MONB 40497S008 3.1 -4.7 -0.8 5:335 DORIS MONP 40497M004 7110 40497M001 3.2 -2.5 </td <td>BJFS 21601M001</td> <td>7249 21601S004</td> <td>0.6</td> <td>2.2</td> <td>5.9</td> <td>3:171</td> <td>SLR</td>	BJFS 21601M001	7249 21601S004	0.6	2.2	5.9	3:171	SLR
HRAO 30302M004 7501 30302M003 -2.2 2.6 3.5 3:214 SLR HRAO 30302M004 HBKB 30302S006 5.4 -1.0 -4.2 3:214 DORIS NKLG 32809M002 LIBB 32809S003 2.9 0.2 2.7 99:36 DORIS STIO 40101M001 7625 40101M003 0.3 0.1 1.6 99:269 VLBI YELL 40127M003 7285 40127M001 -3.9 -3.9 -1.0 1:285 VLBI PIE1 40456M001 7234 40456S001 -3.1 -2.8 -0.3 92:336 VLBI NLIB 40465M001 7612 40465S001 -3.4 -2.0 -5.2 93:64 VLBI MKEA 40477M001 7617 40477S001 -3.7 -1.1 0.8 96:221 VLBI MONP 40497M004 MONB 40497S008 3.1 -4.7 -0.8 5:335 DORIS MONP 40497M004 7110 40497M001 3.2 -2.5 5.7 99:280 SLR FORT 41602M001 7297 41602S001 -0.6 -4.0<	TSKB 21730S005	7345 21730S007	-2.9	-0.8	-3.0	8:16	VLBI
HRAO 30302M004 HBKB 30302S006 5.4 -1.0 -4.2 3:214 DORIS NKLG 32809M002 LIBB 32809S003 2.9 0.2 2.7 99:36 DORIS STJO 40101M001 7625 40101M003 0.3 0.1 1.6 99:269 VLBI YELL 40127M003 7285 40127M001 -3.9 -3.9 -1.0 1:285 VLBI PIE1 40456M001 7234 40456S001 -3.1 -2.8 -0.3 92:336 VLBI NLIB 40465M001 7612 40465S001 -3.4 -2.0 -5.2 93:64 VLBI MKEA 40477M001 7617 40477S001 -3.7 -1.1 0.8 96:221 VLBI MONP 40497M004 MONB 40497S008 3.1 -4.7 -0.8 5:335 DORIS MONP 40497M004 7110 40497M001 3.2 -2.5 5.7 99:280 SLR FORT 41602M001 7297 41602S001 -0.6 -4.0 2.3 93:264 VLBI GLPS 42005M002 SCRB 42005S001 -1.5 2.	HRAO 30302M004	7232 30302S001	-1.6	3.2	2.0	3:214	VLBI
NKLG 32809M002 LIBB 32809S003 2.9 0.2 2.7 99:36 DORIS STJO 40101M001 7625 40101M003 0.3 0.1 1.6 99:269 VLBI YELL 40127M003 7285 40127M001 -3.9 -3.9 -1.0 1:285 VLBI PIE1 40456M001 7234 40456S001 -3.1 -2.8 -0.3 92:336 VLBI NLIB 40465M001 7612 40465S001 -3.4 -2.0 -5.2 93:64 VLBI MKEA 40477M001 7617 40477S001 -3.7 -1.1 0.8 96:221 VLBI MONP 40497M004 MONB 40497S008 3.1 -4.7 -0.8 5:335 DORIS MONP 40497M004 7110 40497M001 3.2 -2.5 5.7 99:280 SLR FORT 41602M001 7297 41602S001 -0.6 -4.0 2.3 93:264 VLBI GLPS 42005M002 SCRB 42005S001 -2.8 -1.0 -1.1 5:92 DORIS CRO1 43201M001 7615 43201S001 -1.5 2.	HRAO 30302M004	7501 30302M003	-2.2	2.6	3.5	3:214	SLR
STIO 40101M001 7625 40101M003 0.3 0.1 1.6 99:269 VLBI YELL 40127M003 7285 40127M001 -3.9 -3.9 -1.0 1:285 VLBI PIE1 40456M001 7234 40456S001 -3.1 -2.8 -0.3 92:336 VLBI NLIB 40465M001 7612 40465S001 -3.4 -2.0 -5.2 93:64 VLBI MKEA 40477M001 7617 40477S001 -3.7 -1.1 0.8 96:221 VLBI MONP 40497M004 MONB 40497S008 3.1 -4.7 -0.8 5:335 DORIS MONP 40497M004 7110 40497M001 3.2 -2.5 5.7 99:280 SLR FORT 41602M001 7297 41602S001 -0.6 -4.0 2.3 93:264 VLBI GLPS 42005M002 SCRB 42005S001 -2.8 -1.0 -1.1 5:92 DORIS CRO1 43201M001 7615 43201S001 -1.5 2.2 1.6 94:16 VLBI STR1 50119M002 7849 50119S001 4.1 0.1	HRAO 30302M004	HBKB 30302S006	5.4	-1.0	-4.2	3:214	DORIS
YELL 40127M003 7285 40127M001 -3.9 -3.9 -1.0 1:285 VLBI PIE1 40456M001 7234 40456S001 -3.1 -2.8 -0.3 92:336 VLBI NLIB 40465M001 7612 40465S001 -3.4 -2.0 -5.2 93:64 VLBI MKEA 40477M001 7617 40477S001 -3.7 -1.1 0.8 96:221 VLBI MONP 40497M004 MONB 40497S008 3.1 -4.7 -0.8 5:335 DORIS MONP 40497M004 7110 40497M001 3.2 -2.5 5.7 99:280 SLR FORT 41602M001 7297 41602S001 -0.6 -4.0 2.3 93:264 VLBI GLPS 42005M002 SCRB 42005S001 -2.8 -1.0 -1.1 5:92 DORIS CRO1 43201M001 7615 43201S001 -1.5 2.2 1.6 94:16 VLBI STR1 50119M002 7849 50119S001 4.1 0.1 4.0 1:209 SLR CHAT 50207M001 CHAB 50207S001 0.1 0.6 </td <td>NKLG 32809M002</td> <td>LIBB 32809S003</td> <td>2.9</td> <td>0.2</td> <td>2.7</td> <td>99:36</td> <td>DORIS</td>	NKLG 32809M002	LIBB 32809S003	2.9	0.2	2.7	99:36	DORIS
PIE1 40456M001 7234 40456S001 -3.1 -2.8 -0.3 92:336 VLBI NLIB 40465M001 7612 40465S001 -3.4 -2.0 -5.2 93:64 VLBI MKEA 40477M001 7617 40477S001 -3.7 -1.1 0.8 96:221 VLBI MONP 40497M004 MONB 40497S008 3.1 -4.7 -0.8 5:335 DORIS MONP 40497M004 7110 40497M001 3.2 -2.5 5.7 99:280 SLR FORT 41602M001 7297 41602S001 -0.6 -4.0 2.3 93:264 VLBI GLPS 42005M002 SCRB 42005S001 -2.8 -1.0 -1.1 5:92 DORIS CRO1 43201M001 7615 43201S001 -1.5 2.2 1.6 94:16 VLBI HOB2 50116M004 7242 50116S002 4.7 -3.0 -2.0 2:81 VLBI STR1 50119M002 7849 50119S001 4.1 0.1 4.0 1:209 SLR SYOG 66006S002 7342 66006S004 0.8 3.8 <td>STJO 40101M001</td> <td>7625 40101M003</td> <td>0.3</td> <td>0.1</td> <td>1.6</td> <td>99:269</td> <td>VLBI</td>	STJO 40101M001	7625 40101M003	0.3	0.1	1.6	99:269	VLBI
NLIB 40465M001 7612 40465S001 -3.4 -2.0 -5.2 93:64 VLBI MKEA 40477M001 7617 40477S001 -3.7 -1.1 0.8 96:221 VLBI MONP 40497M004 MONB 40497S008 3.1 -4.7 -0.8 5:335 DORIS MONP 40497M004 7110 40497M001 3.2 -2.5 5.7 99:280 SLR FORT 41602M001 7297 41602S001 -0.6 -4.0 2.3 93:264 VLBI GLPS 42005M002 SCRB 42005S001 -2.8 -1.0 -1.1 5:92 DORIS CRO1 43201M001 7615 43201S001 -1.5 2.2 1.6 94:16 VLBI HOB2 50116M004 7242 50116S002 4.7 -3.0 -2.0 2:81 VLBI STR1 50119M002 7849 50119S001 4.1 0.1 4.0 1:209 SLR CHAT 50207M001 CHAB 50207S001 0.1 0.6 -3.3 99:56 DORIS SYOG 66006S002 7342 66006S004 0.8 3.8	YELL 40127M003	7285 40127M001	-3.9	-3.9	-1.0	1:285	VLBI
MKEA 40477M001 7617 40477S001 -3.7 -1.1 0.8 96:221 VLBI MONP 40497M004 MONB 40497S008 3.1 -4.7 -0.8 5:335 DORIS MONP 40497M004 7110 40497M001 3.2 -2.5 5.7 99:280 SLR FORT 41602M001 7297 41602S001 -0.6 -4.0 2.3 93:264 VLBI GLPS 42005M002 SCRB 42005S001 -2.8 -1.0 -1.1 5:92 DORIS CR01 43201M001 7615 43201S001 -1.5 2.2 1.6 94:16 VLBI HOB2 50116M004 7242 50116S002 4.7 -3.0 -2.0 2:81 VLBI STR1 50119M002 7849 50119S001 4.1 0.1 4.0 1:209 SLR CHAT 50207M001 CHAB 50207S001 0.1 0.6 -3.3 99:56 DORIS SYOG 66006S002 7342 66006S004 0.8 3.8 0.6 0:1 VLBI KERG 91201M002 KERB 91201S003 -3.2 -2.2	PIE1 40456M001	7234 40456S001	-3.1	-2.8	-0.3	92:336	VLBI
MONP 40497M004 MONB 40497S008 3.1 -4.7 -0.8 5:335 DORIS MONP 40497M004 7110 40497M001 3.2 -2.5 5.7 99:280 SLR FORT 41602M001 7297 41602S001 -0.6 -4.0 2.3 93:264 VLBI GLPS 42005M002 SCRB 42005S001 -2.8 -1.0 -1.1 5:92 DORIS CRO1 43201M001 7615 43201S001 -1.5 2.2 1.6 94:16 VLBI HOB2 50116M004 7242 50116S002 4.7 -3.0 -2.0 2:81 VLBI STR1 50119M002 7849 50119S001 4.1 0.1 4.0 1:209 SLR CHAT 50207M001 CHAB 50207S001 0.1 0.6 -3.3 99:56 DORIS SYOG 66006S002 7342 66006S004 0.8 3.8 0.6 0:1 VLBI KERG 91201M002 KERB 91201S003 -3.2 -2.2 2.0 7:101 DORIS DUM1 91501M001 ADEB 91501S003 -5.0 -3.0	NLIB 40465M001	7612 40465S001	-3.4	-2.0	-5.2	93:64	VLBI
MONP 40497M004 7110 40497M001 3.2 -2.5 5.7 99:280 SLR FORT 41602M001 7297 41602S001 -0.6 -4.0 2.3 93:264 VLBI GLPS 42005M002 SCRB 42005S001 -2.8 -1.0 -1.1 5:92 DORIS CRO1 43201M001 7615 43201S001 -1.5 2.2 1.6 94:16 VLBI HOB2 50116M004 7242 50116S002 4.7 -3.0 -2.0 2:81 VLBI STR1 50119M002 7849 50119S001 4.1 0.1 4.0 1:209 SLR CHAT 50207M001 CHAB 50207S001 0.1 0.6 -3.3 99:56 DORIS SYOG 66006S002 7342 66006S004 0.8 3.8 0.6 0:1 VLBI KERG 91201M002 KERB 91201S003 -3.2 -2.2 2.0 7:101 DORIS DUM1 91501M001 ADEB 91501S002 -0.2 -0.2 1.3 8:39 DORIS THTI 92201M009 PAPB 92201S007 -1.3 1.7	MKEA 40477M001	7617 40477S001	-3.7	-1.1	0.8	96:221	VLBI
FORT 41602M001 7297 41602S001 -0.6 -4.0 2.3 93:264 VLBI GLPS 42005M002 SCRB 42005S001 -2.8 -1.0 -1.1 5:92 DORIS CRO1 43201M001 7615 43201S001 -1.5 2.2 1.6 94:16 VLBI HOB2 50116M004 7242 50116S002 4.7 -3.0 -2.0 2:81 VLBI STR1 50119M002 7849 50119S001 4.1 0.1 4.0 1:209 SLR CHAT 50207M001 CHAB 50207S001 0.1 0.6 -3.3 99:56 DORIS SYOG 66006S002 7342 66006S004 0.8 3.8 0.6 0:1 VLBI KERG 91201M002 KERB 91201S003 -3.2 -2.2 2.0 7:101 DORIS DUM1 91501M001 ADEB 91501S002 -0.2 -0.2 1.3 8:39 DORIS THTI 92201M009 PAPB 92201S007 -1.3 1.7 4.4 7:278 DORIS NOUM 92701M003 NOUB 92701S002 0.5 5.8	MONP 40497M004	MONB 40497S008	3.1	-4.7	-0.8	5:335	DORIS
GLPS 42005M002 SCRB 42005S001 -2.8 -1.0 -1.1 5:92 DORIS CRO1 43201M001 7615 43201S001 -1.5 2.2 1.6 94:16 VLBI HOB2 50116M004 7242 50116S002 4.7 -3.0 -2.0 2:81 VLBI STR1 50119M002 7849 50119S001 4.1 0.1 4.0 1:209 SLR CHAT 50207M001 CHAB 50207S001 0.1 0.6 -3.3 99:56 DORIS SYOG 66006S002 7342 66006S004 0.8 3.8 0.6 0:1 VLBI KERG 91201M002 KERB 91201S003 -3.2 -2.2 2.0 7:101 DORIS DUM1 91501M001 ADEB 91501S002 -0.2 -0.2 1.3 8:39 DORIS THTI 92201M009 PAPB 92201S007 -1.3 1.7 4.4 7:278 DORIS NOUM 92701M003 NOUB 92701S002 0.5 5.8 2.2 5:236 DORIS	MONP 40497M004	7110 40497M001	3.2	-2.5	5.7	99:280	SLR
CRO1 43201M001 7615 43201S001 -1.5 2.2 1.6 94:16 VLBI HOB2 50116M004 7242 50116S002 4.7 -3.0 -2.0 2:81 VLBI STR1 50119M002 7849 50119S001 4.1 0.1 4.0 1:209 SLR CHAT 50207M001 CHAB 50207S001 0.1 0.6 -3.3 99:56 DORIS SYOG 66006S002 7342 66006S004 0.8 3.8 0.6 0:1 VLBI KERG 91201M002 KERB 91201S003 -3.2 -2.2 2.0 7:101 DORIS DUM1 91501M001 ADEB 91501S002 -0.2 -0.2 1.3 8:39 DORIS THTI 92201M009 PAPB 92201S007 -1.3 1.7 4.4 7:278 DORIS NOUM 92701M003 NOUB 92701S002 0.5 5.8 2.2 5:236 DORIS	FORT 41602M001	7297 41602S001	-0.6	-4.0	2.3	93:264	VLBI
HOB2 50116M004 7242 50116S002 4.7 -3.0 -2.0 2:81 VLBI STR1 50119M002 7849 50119S001 4.1 0.1 4.0 1:209 SLR CHAT 50207M001 CHAB 50207S001 0.1 0.6 -3.3 99:56 DORIS SYOG 66006S002 7342 66006S004 0.8 3.8 0.6 0:1 VLBI KERG 91201M002 KERB 91201S003 -3.2 -2.2 2.0 7:101 DORIS DUM1 91501M001 ADEB 91501S002 -0.2 -0.2 1.3 8:39 DORIS DUM1 91501M001 ADFB 91501S003 -5.0 -3.0 3.3 8:39 DORIS THTI 92201M009 PAPB 92201S007 -1.3 1.7 4.4 7:278 DORIS NOUM 92701M003 NOUB 92701S002 0.5 5.8 2.2 5:236 DORIS	GLPS 42005M002	SCRB 42005S001	-2.8	-1.0	-1.1	5:92	DORIS
STR1 50119M002 7849 50119S001 4.1 0.1 4.0 1:209 SLR CHAT 50207M001 CHAB 50207S001 0.1 0.6 -3.3 99:56 DORIS SYOG 66006S002 7342 66006S004 0.8 3.8 0.6 0:1 VLBI KERG 91201M002 KERB 91201S003 -3.2 -2.2 2.0 7:101 DORIS DUM1 91501M001 ADEB 91501S002 -0.2 -0.2 1.3 8:39 DORIS DUM1 91501M001 ADFB 91501S003 -5.0 -3.0 3.3 8:39 DORIS THTI 92201M009 PAPB 92201S007 -1.3 1.7 4.4 7:278 DORIS NOUM 92701M003 NOUB 92701S002 0.5 5.8 2.2 5:236 DORIS	CRO1 43201M001	7615 43201S001	-1.5	2.2	1.6	94:16	VLBI
CHAT 50207M001 CHAB 50207S001 0.1 0.6 -3.3 99:56 DORIS SYOG 66006S002 7342 66006S004 0.8 3.8 0.6 0:1 VLBI KERG 91201M002 KERB 91201S003 -3.2 -2.2 2.0 7:101 DORIS DUM1 91501M001 ADEB 91501S002 -0.2 -0.2 1.3 8:39 DORIS DUM1 91501M001 ADFB 91501S003 -5.0 -3.0 3.3 8:39 DORIS THTI 92201M009 PAPB 92201S007 -1.3 1.7 4.4 7:278 DORIS NOUM 92701M003 NOUB 92701S002 0.5 5.8 2.2 5:236 DORIS	HOB2 50116M004	7242 50116S002	4.7	-3.0	-2.0	2:81	VLBI
SYOG 66006S002 7342 66006S004 0.8 3.8 0.6 0:1 VLBI KERG 91201M002 KERB 91201S003 -3.2 -2.2 2.0 7:101 DORIS DUM1 91501M001 ADEB 91501S002 -0.2 -0.2 1.3 8:39 DORIS DUM1 91501M001 ADFB 91501S003 -5.0 -3.0 3.3 8:39 DORIS THTI 92201M009 PAPB 92201S007 -1.3 1.7 4.4 7:278 DORIS NOUM 92701M003 NOUB 92701S002 0.5 5.8 2.2 5:236 DORIS	STR1 50119M002	7849 50119S001	4.1	0.1	4.0	1:209	SLR
KERG 91201M002 KERB 91201S003 -3.2 -2.2 2.0 7:101 DORIS DUM1 91501M001 ADEB 91501S002 -0.2 -0.2 1.3 8:39 DORIS DUM1 91501M001 ADFB 91501S003 -5.0 -3.0 3.3 8:39 DORIS THTI 92201M009 PAPB 92201S007 -1.3 1.7 4.4 7:278 DORIS NOUM 92701M003 NOUB 92701S002 0.5 5.8 2.2 5:236 DORIS	CHAT 50207M001	CHAB 50207S001	0.1	0.6	-3.3	99:56	DORIS
DUM1 91501M001 ADEB 91501S002 -0.2 -0.2 1.3 8:39 DORIS DUM1 91501M001 ADFB 91501S003 -5.0 -3.0 3.3 8:39 DORIS THTI 92201M009 PAPB 92201S007 -1.3 1.7 4.4 7:278 DORIS NOUM 92701M003 NOUB 92701S002 0.5 5.8 2.2 5:236 DORIS	SYOG 66006S002	7342 66006S004	0.8	3.8	0.6	0:1	VLBI
DUM1 91501M001 ADFB 91501S003 -5.0 -3.0 3.3 8:39 DORIS THTI 92201M009 PAPB 92201S007 -1.3 1.7 4.4 7:278 DORIS NOUM 92701M003 NOUB 92701S002 0.5 5.8 2.2 5:236 DORIS	KERG 91201M002	KERB 91201S003	-3.2	-2.2	2.0	7:101	DORIS
THTI 92201M009 PAPB 92201S007 -1.3 1.7 4.4 7:278 DORIS NOUM 92701M003 NOUB 92701S002 0.5 5.8 2.2 5:236 DORIS	DUM1 91501M001	ADEB 91501S002	-0.2	-0.2	1.3	8:39	DORIS
NOUM 92701M003 NOUB 92701S002 0.5 5.8 2.2 5:236 DORIS	DUM1 91501M001	ADFB 91501S003	-5.0	-3.0	3.3	8:39	DORIS
	THTI 92201M009	PAPB 92201S007	-1.3	1.7	4.4	7:278	DORIS
REUN 97401M003 REUB 97401S002 0.4 2.5 -1.0 3:335 DORIS	NOUM 92701M003	NOUB 92701S002	0.5	5.8	2.2	5:236	DORIS
	REUN 97401M003	REUB 97401S002	0.4	2.5	-1.0	3:335	DORIS

The tie vectors are listed from GPS to other technique reference markers

antennas would probably increase its scale bias with respect to SLR. Appleby et al. (2009) reported that the introduction of a high-accuracy event timer at Herstmonceux (UK) SLR site in February 2007 has caused an apparent discontinuity

in the laser range data, suggesting that the data prior to this date of equipment change is biased by 12 mm. An appropriate correction was then applied by the ILRS for their ITRF2008 submission (Pavlis et al. 2010). As it can be seen in Table 4



Table 5 Tie discrepancy percentage

Discrepancy (mm)	GPS-VLBI	GPS-SLR	GPS-DORIS	
<6	47	43	34	
6–10	24	29	12	
>10	29	28	54	

(line HERS), the agreement between the local survey operated in June 2008 (IGN 2008) and space geodesy estimate is at the level of 3 mm.

4.6 ITRF2008 earth orientation parameters

Similar to the ITRF2005 experience, ITRF2008 provides consistent series of polar motion and its daily rates, universal time (UT1-UTC) and Length of Day (LOD), with the latter being determined by VLBI uniquely. The reason for using LOD values from VLBI only is to avoid contaminating the VLBI estimates by biased determinations from satellite techniques (Ray 1996, 2009). In order to evaluate the EOP consistency between the four technique solutions, Fig. 7 displays the post fit polar motion residuals as result from the ITRF2008 combination, where it can be seen that GPS is

dominating the three other techniques. It is worth noting that GPS polar motion helps tieing the three other technique solutions, via the two frame rotation parameters around the X and Y-axes. The WRMS values computed over the post fit residuals between the combined and the individual polar motion time series are (for the couple x and y components): (10,10), (142,120), (144,128), (239, 353) in micro-arc-seconds, for GPS, VLBI, SLR and DORIS, respectively.

4.7 ITRF2008 performance in station positions and velocities

In order to evaluate the ITRF2008 performance in station positions and velocities determination, we show the precision gain of these estimated parameters, compared to ITRF2005 results. Figure 8 illustrates for both frames the spherical

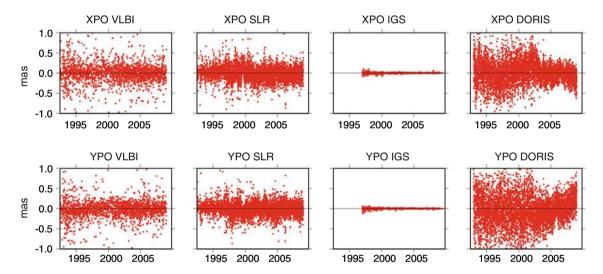


Fig. 7 ITRF2008 post fit residuals of polar motion in milli-arc-seconds

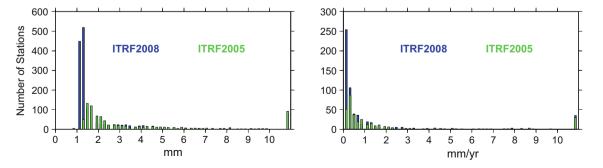


Fig. 8 Histogram of the ITRF2008 and ITRF2005 station position and velocity spherical errors



Fig. 9 ITRF2008 horizontal velocities with formal error less than 0.2 mm/yr. Major plate boundaries are shown according to Bird (2003)

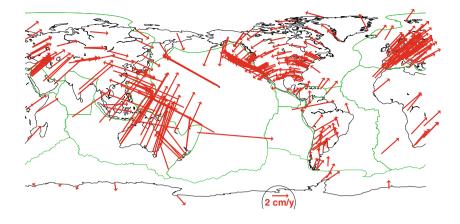
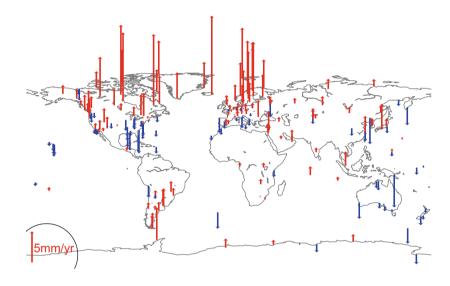


Fig. 10 ITRF2008 vertical velocities with formal error less than 0.2 mm/yr. Positive velocities are shown in *red* and negative in *blue*



errors (computed following Altamimi et al. (2002b), Appendix A7) in station positions and velocities, showing the improvement achieved by ITRF2008. The station position spherical errors are computed at the epochs of minimum variances, following Altamimi et al. (2002b), Appendix A6, Eq. A22. They correspond, for each point position and velocity, to the square root of the square sum of the formal errors along the three components. Although these errors are very often believed to be optimistic (Williams et al. 2004), the proportion of ITRF2008 sites having greater precision as indicted by this figure is relatively high to demonstrate the well performance of ITRF2008 determination. The velocity error histogram shows a number of stations with errors greater than 10 mm/yr which are due to short data spans for mostly old SLR and VLBI stations.

In addition, Figs. 9 and 10 show the ITRF2008 horizontal and vertical velocity fields, respectively, with formal error less than 0.2 mm/yr. At this high level of precision, ITRF2008 contains 406 sites in horizontal and 278 sites in vertical, versus 150 and 30 sites for ITRF2005.

4.8 Transformation parameters between ITRF2008 and ITRF2005

For many applications and in order to ensure the link between ITRF2008 and ITRF2005, it is essential to assess consistently the transformation parameters between the two frames. The same 179 stations (Fig. 3) that were used to ensure the alignment of the ITRF2008 orientation and its rate to the ITRF2005 were also used to estimate the transformation parameters between the two frames. The main criteria for selection of these 179 stations are (1) to have the best possible site distribution; (2) to involve as many as possible VLBI, SLR, GPS and DORIS stations and (3) to have the best agreement between the two frames in terms of postfit residuals of the 14-parameter transformation. Regarding this third criteria, the WRMS values of the 14-parameter similarity transformation fit are 2.4, 2.9 and 3.9 mm in position (at epoch 2005.0) and 0.4, 0.4, 0.7 mm/yr in velocity, in east, north and vertical components, respectively. Table 6 lists the transformation parameters from ITRF2008



	Tx (mm) $\dot{T}x \text{ (mm/yr)}$	Ty (mm) Ty (mm/yr)	Tz (mm) Tz (mm/yr)	<i>D</i> (ppb) <i>D</i> (ppb/yr)	Rx (mas) Rx (mas/yr)	Ry (mas) Ry (mas/yr)	Rz (mas) Rz (mas/yr)
±	-0.5	-0.9	-4.7	0.94	0.00	0.00	0.00
	0.2	0.2	0.2	0.03	0.08	0.08	0.08
\pm	0.3	0.0	0.0	0.00	0.00	0.00	0.00
	0.2	0.2	0.2	0.03	0.08	0.08	0.08

Table 6 Transformation Parameters at epoch 2005.0 and their rates from ITRF2008 to ITRF2005, to be used with Eq. 4

to ITRF2005, to be used with the transformation formula given by Eq. (4).

$$\begin{cases}
\begin{pmatrix} x \\ y \\ z \end{pmatrix}_{i05} = \begin{pmatrix} x \\ y \\ z \end{pmatrix}_{i08} + T + D \begin{pmatrix} x \\ y \\ z \end{pmatrix}_{i08} + R \begin{pmatrix} x \\ y \\ z \end{pmatrix}_{i08} \\
\begin{pmatrix} \dot{x} \\ \dot{y} \\ \dot{z} \end{pmatrix}_{i05} = \begin{pmatrix} \dot{x} \\ \dot{y} \\ \dot{z} \end{pmatrix}_{i08} + \dot{T} + \dot{D} \begin{pmatrix} x \\ y \\ z \end{pmatrix}_{i08} + \dot{R} \begin{pmatrix} x \\ y \\ z \end{pmatrix}_{i08}
\end{cases} (4)$$

where *i*05 designates ITRF2005 and *i*08 ITRF2008, T is the translation vector, $T = (T_x, T_y, T_z)^T$, D is the scale factor and R is the matrix containing the rotation angles, given by

$$R = \begin{pmatrix} 0 & -R_z & R_y \\ R_z & 0 & -R_x \\ -R_y & R_x & 0 \end{pmatrix}$$

The dotted parameters designate their time derivatives. The values of the 14 parameters are those listed in Table 6. Note that the inverse transformation from ITRF2005 to ITRF2008 follows by interchanging (i08) with (i05) and changing the sign of the transformation parameters.

5 Conclusion

The ITRF2008 follows the series of constantly improved ITRF solutions thanks to the continuous improvements of the input data provided by the technique services and the combination strategy. The results presented in this paper demonstrate the high performance of the ITRF2008, compared with past solutions including the ITRF2005, not only in terms of precision of station positions and velocities, but also the improvement of its defining parameters, specifically the origin and the scale. While the accuracy of the ITRF2008 origin is difficult to assess, being determined by SLR only, we believe that its agreement with the ITRF2005 origin at the level of or better than 1 cm (at any epoch covering the time-span of the SLR observations) could be regarded as the level of the origin accuracy attainable today. An assessment of the scale discrepancy between the two independent techniques (VLBI and SLR) solutions, and hence the current scale accuracy, is believed to be at the level of 1.2 ppb (8 mm) over the common time-span of their observations covering about 26 years.

Thorough analysis and several test combinations have been undertaken during the ITRF2008 project to evaluate the impact of the local ties and their usage on the results and in particular on the frame defining parameters. We showed that using the local ties at their measurement epochs as observations and properly weighted with full variance covariance information provides the most statistically satisfactory combination. Although newly reprocessed solutions of the four techniques were used for the ITRF2008 generation, discrepancies between local surveys and space geodesy estimates still exist: 50% of the available SLR and VLBI tie vectors to GPS exhibit residuals larger than 6 mm, and about 30% have residuals larger than 10 mm. We identified that some large tie discrepancies might be due to the effect of un-calibrated radomes covering the GPS antennas in some IGS sites. However, including all available local ties properly weighted and if possible with full variance covariance information has been shown to lead to the best inter-technique combination.

As results from the ITRF2008 combination, we evaluate the uncertainty of the transfer of SLR origin and SLR and VLBI mean scale to GPS frame to be at the level of 0.6 mm at epoch 2005.0. This uncertainty would certainly be improved if we had a better geometry of co-location network with better agreement between space geodesy and local tie estimates. Therefore, the prerequisite for improving the ITRF is to start by improving the geodetic infrastructure.

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