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Large-Eddy Simulation of the Temporal Mixing Layer Using the Clark Model

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Abstract. The Clark model for the turbulent stress tensor in large-eddy simulation is investigated from a theoretical and computational point of view. In order to be applicable to compressible turbulent flows, the Clark model has been reformulated. Actual large-eddy simulation of a weakly compressible, turbulent, temporal mixing layer shows that the eddy-viscosity part of the original Clark model gives rise to an excessive dissipation of energy in the transitional regime. On the other hand, the model gives rise to instabilities if the eddy-viscosity part is omitted and only the "gradient" part is retained. A linear stability analysis of the Burgers equation supplemented with the Clark model is performed in order to clarify the nature of the instability. It is shown that the growth-rate of the instability is infinite in the inviscid limit and that sufficient (eddy-)viscosity can stabilize the model. A model which avoids both the excessive dissipation of the original Clark model as well as the instability of the "gradient" part, is obtained when the dynamic procedure is applied to the Clark model. Large-eddy simulation using this new dynamic Clark model is found to yield satisfactory results when compared with a filtered direct numerical simulation. Compared with the standard dynamic eddy-viscosity model, the dynamic Clark model the new model provides equal accuracy at a lower computational effort.

1. Introduction

Large-eddy simulation (LES) is a promising technique to simulate turbulent flows (Rogallo and Moin, 1984). In LES only the large scales are explicitly solved and for this purpose the Navier–Stokes equations are filtered. The filtering introduces new terms, the so-called subgrid terms representing the effect of the small scales, which have to be modeled in order to close the filtered equations. A number of subgrid models for the most important subgrid term, the turbulent stress tensor, exist. Using an eddy-viscosity hypothesis is the most common procedure to model this tensor and the Smagorinsky model is the best-known example (Smagorinsky, 1963). Recently, a dynamic formulation of this model has been proposed (Germano, 1992), in which the model coefficient is dynamically adjusted to the local structure of the flow. This procedure considerably improves the performance of the eddy-viscosity model in transitional and inhomogeneous flows (Germano *et al.*, 1991). However, the eddy-viscosity concept is not the only strategy which has been followed to model the turbulent stress tensor. Alternative formulations are provided by, e.g., the similarity model (Bardina *et al.*, 1984) and the Clark model (Clark *et al.*, 1979). An increasing amount of research is directed toward the use of the similarity idea in LES. The validity of this idea has been investigated using experimental data (Liu *et al.*, 1994). Furthermore, LES has been conducted using the mixed model, in which the similarity model is supplemented with the Smagorinsky eddy-viscosity (Erlebacher *et al.*, 1992). The

dynamical formulation of the mixed model has also been adopted (Zang *et al.*, 1993; Vreman *et al.*, 1994b) and improvement over the standard dynamic eddy-viscosity model was observed. Compared with other subgrid models the dynamic mixed model produces good results, but a simulation using this model is quite expensive, because many explicit filtering operations are needed in each time step.

The model developed by Clark et al. (1979), which is the subject of this paper, has received less attention than the similarity model, although a priori tests of this model yield correlations which are at least as good as those of the similarity model (Clark et al., 1979; Vreman et al., 1993). The Clark model consists of two parts; the first part, which is based on Taylor expansions of the velocity field, contains inner products of velocity gradients and is called the "gradient" model, and the other part is the Smagorinsky eddy-viscosity model. Furthermore, the model satisfies several basic requirements, such as Galilean invariance (Speziale, 1985) and realizability (Vreman et al., 1994a). However, actual simulations using the Clark model are rarely found in the literature. One of the examples is the work of Love (1980) on the one-dimensional Burgers equation. In the present paper we present results of LES with the Clark model for the temporal mixing layer at low Mach number. First, in Section 2 we demonstrate that the derivation of the Clark model for the turbulent stress tensor in incompressible flow can be extended to compressible flow. Next, in Section 3 results of LES with the original Clark model show that the model is too dissipative in the transitional regime. However, if the eddy-viscosity part of the model, which is responsible for the excessive dissipation, is omitted, the simulation becomes unstable. An analysis of the nature of this instability for the onedimensional Burgers equation is presented in Section 4, which shows that the instability can be removed if a sufficient amount of dissipation is added. For this reason the dynamic Clark model is formulated in Section 5, which overcomes the problems of both excessive dissipation and instability. Finally, in Section 6 we summarize our conclusions.

2. Closure of the Filtered Equations with the Clark Model

We first present the filtered Navier–Stokes equations for a compressible flow and then demonstrate how the Clark model can be derived for the turbulent stress tensor using Taylor expansions.

The partial differential equations which govern a compressible flow are the Navier–Stokes equations, representing conservation of mass, momentum, and energy. In the LES of turbulent flows these equations are filtered in order to reduce the amount of scales to be solved. The filter operation extracts the large scale part \overline{f} from a flow variable f as follows:

$$\overline{f}(\mathbf{x}) = \int_{\Omega} G_{\Delta}(\mathbf{x} - \mathbf{y}) f(\mathbf{y}) \, d\mathbf{y}.$$
(1)

Here Ω is the flow domain and Δ denotes the "filter width" associated with the kernel G_{Δ} . We adopt a filter width which does not depend on the postion vector **x** and, consequently, the filter operation is a convolution and commutes with spatial derivatives (Rogallo and Moin, 1984). For compressible flows, Favre (1983) has introduced a related filter operation,

$$\tilde{f} = \frac{\overline{\rho f}}{\bar{\rho}},\tag{2}$$

where ρ denotes the fluid density. The filtered Navier–Stokes equations can be written in the following form (Vreman *et al.*, 1994c):

$$\partial_t \bar{\rho} + \partial_i (\bar{\rho} \tilde{u}_i) = 0, \tag{3}$$

$$\partial_i(\bar{\rho}\tilde{u}_i) + \partial_j(\bar{\rho}\tilde{u}_i\tilde{u}_j) + \partial_i\bar{p} - \partial_j\check{\sigma}_{ij} = -\partial_j(\bar{\rho}\tau_{ij}),\tag{4}$$

$$\partial_t \check{e} + \partial_i ((\check{e} + \bar{p})\tilde{u}_i) - \partial_i (\check{\sigma}_{ij}\tilde{u}_i) + \partial_i \check{q}_j = R,$$
(5)

where the symbols ∂_t and ∂_j denote the partial differential operators $\partial/\partial t$ and $\partial/\partial x_j$, respectively. Furthermore, the summation convention for repeated indices is used. The independent variables t and x_j represent time and the spatial coordinates, respectively. Concerning the flow variables, the Favre-filtered velocity is denoted by $\tilde{\mathbf{u}}$, while $\bar{\rho}$ is the filtered density and \bar{p} is the filtered pressure. Moreover, \check{e} is the total energy density of the filtered variables

$$\check{e} = \frac{\bar{p}}{\gamma - 1} + \frac{1}{2} \bar{\rho} \tilde{u}_i \tilde{u}_i.$$
(6)

The viscous stress tensor based on filtered variables, $\check{\sigma}_{ij}$, is defined as

$$\check{\sigma}_{ij} = \frac{\mu(T)}{Re} S_{ij}(\tilde{\mathbf{u}}) \qquad \text{with} \quad S_{ij}(\tilde{\mathbf{u}}) = (\partial_j \tilde{u}_i + \partial_i \tilde{u}_j - \frac{2}{3} \delta_{ij} \partial_k \tilde{u}_k), \tag{7}$$

where δ_{ij} is the Kronecker delta and $\mu(\tilde{T})$ is the dynamic viscosity, expressed by Sutherland's law for air,

$$\mu(\tilde{T}) = \tilde{T}^{3/2} \frac{1+C}{\tilde{T}+C}$$
(8)

with C = 0.4. In addition **ğ** represents the heat-flux vector based on filtered variables, given by

$$\check{q}_{j} = -\frac{\mu(\tilde{T})}{(\gamma - 1)Re Pr M_{R}^{2}} \,\partial_{j}\tilde{T}.$$
(9)

The Favre-filtered temperature \tilde{T} is related to the filtered density and the filtered pressure by the ideal gas law

$$\tilde{T} = \gamma M_{\mathbf{R}}^2 \frac{\bar{p}}{\bar{\rho}}.$$
(10)

These equations have been made dimensionless by introducing a reference length $L_{\rm R}$, velocity $u_{\rm R}$, density $\rho_{\rm R}$, temperature $T_{\rm R}$, and viscosity $\mu_{\rm R}$. In addition γ , the ratio of the specific heats C_P and C_V , and the Prandtl number Pr are given the values $\gamma = 1.4$ and Pr = 1. The values of the Reynolds number $Re = \rho_{\rm R}u_{\rm R}L_{\rm R}/\mu_{\rm R}$ and the reference Mach number $M_{\rm R} = u_{\rm R}/\sqrt{\gamma R_{\rm g} T_{\rm R}}$, where $R_{\rm g}$ is the universal gas constant, are given below.

In this description the left-hand sides of (3)–(5) are expressed in the filtered variables $\bar{\rho}$, \tilde{u}_j , and \bar{p} . The right-hand sides of the filtered equations contain the so-called subgrid terms. The subgrid term in the momentum equation (4) contains the turbulent stress tensor, defined as

$$\bar{\rho}\tau_{ij} = \overline{\rho u_i u_j} - \frac{\overline{\rho u_i \rho u_j}}{\bar{\rho}} = \bar{\rho}(\widetilde{u_i u_j} - \widetilde{u}_i \widetilde{u}_j).$$
(11)

The effects of the subgrid terms in the energy equation (5) are negligible for the weakly compressible mixing layer considered in this paper (Vreman *et al.*, 1994c). Terms created by nonlinearities in the viscous stress tensor can also be neglected.

In the following we derive the Clark model for $\rho \tau_{ij}$ using Taylor expansions of the filtered velocity. Our procedure is slightly different from the procedure followed by Clark *et al.* (1979), since we do not apply the expansions to the Leonard, cross, and Reynolds components separately, but to the turbulent stress itself. Furthermore, we present the formulation for the compressible turbulent stress tensor, thus generalizing the derivation by Clark *et al.* The same technique can also be applied in the construction of models for other subgrid terms, for example, the pressure-velocity or pressure dilation subgrid term (Vreman *et al.*, 1994c). The filter that we use is the top-hat filter, which is defined by its kernel:

$$G_{\Delta}(\mathbf{y}) = \begin{cases} \frac{1}{\Delta_1 \Delta_2 \Delta_3} & \text{if } |y_i| < \frac{\Delta_i}{2} \quad (i = 1, 2, 3), \\ 0 & \text{otherwise,} \end{cases}$$
(12)

where Δ_i is the filter width in the *i*-direction and $\Delta = (\Delta_1 \Delta_2 \Delta_3)^{1/3}$. Filters commonly used in LES are the top-hat, the Gaussian, and the spectral cut-off filter. Vreman *et al.* (1994a) argue that from the viewpoint of realizability the Clark model should be used in conjunction with positive filters, e.g., the top-hat or Gaussian filter. In this work we have selected the top-hat filter, since it is more attractive than the Gaussian

filter with respect to computational cost. For the top-hat filter \overline{f} is defined as

$$\overline{f}(\mathbf{x}) = \frac{1}{\Delta_1 \Delta_2 \Delta_3} \int_{-\Delta_3/2}^{\Delta_3/2} \int_{-\Delta_2/2}^{\Delta_2/2} \int_{-\Delta_1/2}^{\Delta_1/2} f(\mathbf{x} + \mathbf{y}) \, d\mathbf{y}.$$
(13)

The function $f(\mathbf{x} + \mathbf{y})$ is expanded as a Taylor series around \mathbf{x} , and after evaluation of the integral we obtain

$$\overline{f} = f + \frac{1}{24} \Delta_k^2 \partial_k^2 f + O(\Delta^4).$$
(14)

The same formula holds for Gaussian filters (Love, 1980). We use this formula to rewrite the turbulent stress tensor:

$$\bar{\rho}\tau_{ij} = \overline{\rho u_i u_j} - \frac{\overline{\rho u_i \rho u_j}}{\bar{\rho}}$$

$$= \rho u_i u_j + \frac{1}{24} \Delta_k^2 \partial_k^2 (\rho u_i u_j) - \frac{(\rho u_i + \frac{1}{24} \Delta_k^2 \partial_k^2 (\rho u_i))(\rho u_j + \frac{1}{24} \Delta_k^2 \partial_k^2 (\rho u_j))}{\rho + \frac{1}{24} \Delta_k^2 \partial_k^2 \rho} + O(\Delta^4)$$

$$= \frac{1}{12} \Delta_k^2 \rho (\partial_k u_i) (\partial_k u_j) + O(\Delta^4), \qquad (15)$$

where we used the relation

$$\frac{1}{\rho + \frac{1}{24}\Delta_k^2 \partial_k^2 \rho} = \frac{1}{\rho} - \frac{1}{24\rho^2} \Delta_k^2 \partial_k^2 \rho + O(\Delta^4).$$
(16)

The next step is to express (15) into filtered variables. Using (16), the Favre-filtered velocity can be written as

$$\begin{split} \tilde{u}_{i} &= \frac{\overline{\rho u_{i}}}{\overline{\rho}} \\ &= \frac{\rho u_{i} + \frac{1}{24} \Delta_{k}^{2} \partial_{k}^{2} (\rho u_{i})}{\rho + \frac{1}{24} \Delta_{k}^{2} \partial_{k}^{2} \rho} + O(\Delta^{4}) \\ &= u_{i} + \frac{1}{24} \Delta_{k}^{2} \partial_{k}^{2} u_{i} + \frac{1}{12\rho} \Delta_{k}^{2} (\partial_{k} \rho) (\partial_{k} u_{i}) + O(\Delta^{4}). \end{split}$$
(17)

We observe that for both the bar and Favre filter, unfiltered and filtered variables differ by a term of the order $O(\Delta^2)$:

$$\rho = \bar{\rho} + O(\Delta^2),\tag{18}$$

$$u_i = \tilde{u}_i + O(\Delta^2). \tag{19}$$

Substituting expressions (18) and (19) into (15) yields

$$\bar{\rho}\tau_{ii} = \frac{1}{12}\Delta_k^2 \bar{\rho}(\partial_k \tilde{u}_i)(\partial_k \tilde{u}_i) + O(\Delta^4).$$
⁽²⁰⁾

The first term on the right-hand side is referred to as the "gradient" model. Observe that this expansion is mathematically correct provided the variables can be differentiated sufficiently often. However, for rapidly fluctuating variables, the $O(\Delta^4)$ term may not be small and hence simply omitting the rest-term can be inaccurate. Clark *et al.* have inserted the Smagorinsky eddy-viscosity model for this rest-term, which finally yields the original Clark model:

$$\bar{\rho}\tau_{ij} = \frac{1}{12}\Delta_k^2 \bar{\rho}(\partial_k \tilde{u}_i)(\partial_k \tilde{u}_j) - C_s^2 \Delta^2 \bar{\rho} |S(\tilde{\mathbf{u}})| S_{ij}(\tilde{\mathbf{u}}),$$
(21)

where $|S(\tilde{\mathbf{u}})|^2 = \frac{1}{2}S_{ij}^2(\tilde{\mathbf{u}})$ and the model coefficient $C_s = 0.17$, a value suggested by Clark *et al.* (1979) and Schumann (1991). Equations (18) and (19) show that the Smagorinsky model is $O(\Delta^2)$, whereas according to (20) a model for the rest-term is preferred to be $O(\Delta^4)$. At present, however, no satisfactory $O(\Delta^4)$ model is available, therefore we consider the Smagorinsky model, which is of the order $O(\Delta^2)$, for the rest-term.

The Clark *et al.* model has been called "mathematically indefensible" by Love (1980) (compare the work by Rogallo and Moin (1984)), since Clark *et al.* used Taylor expansions of the subgrid-scale velocity field $u'_i = u_i - \bar{u}_i$, which is not smoothly varying over lengths of $O(\Delta)$. The derivation in this paper is different from

the derivation by Clark *et al.*, since we do not use Taylor expansions of the subgrid-scale velocity field u'_i , but Taylor expansions of the *unfiltered* variables ρ and u_i . However, like the subgrid-scale quantities the unfiltered variables are not smoothly varying over lengths of $O(\Delta)$ either. This point does not imply that the expansion is mathematically incorrect (assuming that derivatives exist), although the higher-order terms may not be small compared with the terms retained. An alternative procedure is to use Taylor expansions of *filtered* variables only, which vary more smoothly over lengths of $O(\Delta)$. This can be achieved by considering not the turbulent stress itself, but the similarity model for the turbulent stress (Vreman *et al.*, 1994d):

$$\bar{\rho}\tau_{ij}^{\rm sim} = \overline{\bar{\rho}\tilde{u}_i\tilde{u}_j} - \frac{\bar{\bar{\rho}}\tilde{u}_i\bar{\rho}\tilde{u}_j}{\bar{\bar{\rho}}},\tag{22}$$

which is the definition of the turbulent stress in compressible flow applied to *filtered* variables. For incompressible flows this model reduces to the sum of the Leonard term and Bardina's similarity model for the cross-term (Bardina *et al.*, 1984). By analogy to (15) the tensor given in (22) can be rewritten as

$$\bar{\rho}\tau_{ij}^{\rm sim} = \frac{1}{12}\Delta_k^2\bar{\rho}(\partial_k\tilde{u}_i)(\partial_k\tilde{u}_j) + O(\Delta^4).$$
⁽²³⁾

In this derivation only Taylor expansions of the *filtered* quantities $\bar{\rho}$ and \tilde{u}_i are employed, whereas in (15) Taylor expansions of *unfiltered* variables are used. The expressions in (20) and (23) are identical, but the $O(\Delta^4)$ term in the latter case will in general be smaller. From this point of view the Smagorinsky eddy-viscosity in the Clark model accounts not only for the relatively small $O(\Delta^4)$ term in (23), but also for the discrepancy between the similarity model $\bar{\rho}\tau_{ii}^{sim}$ and the turbulent stress $\bar{\rho}\tau_{ij}$.

3. Shortcomings of the Clark Model in Actual LES

In this section we present results of LES of the temporal mixing layer using the Clark subgrid model. The results are compared with the results of a filtered direct numerical simulation (DNS). We first discuss the DNS, in which the unfiltered Navier–Stokes equations are solved.

In order to simulate a three-dimensional temporal mixing layer we solve the flow equations in a cubic geometry $[0, L] \times [-\frac{1}{2}L, \frac{1}{2}L] \times [0, L]$. Periodic boundary conditions are imposed in the x_1 - and x_3 -direction, while in the x_2 -direction the boundaries are free slip walls, i.e., the normal velocity and the normal derivatives of the density, pressure, and tangential velocities are zero. The basic initial velocity profile is the hyperbolic tangent profile $u = \tanh x_2$. The initial temperature profile is obtained from the Busemann-Crocco law (Ragab and Wu, 1989) and the initial pressure distribution is uniform. The reference length L_R is half the initial vorticity thickness, while the reference density, velocity, temperature, and dynamic viscosity are the initial upper stream values. As in other simulations of the three-dimensional mixing layer (Moser and Rogers, 1993; Comte *et al.*, 1992), the length L of the domain is set to four times the wavelength of the most unstable mode according to linear stability theory. We superimpose a three-dimensional large amplitude perturbation on the mean profile, as described by Sandham and Reynolds (1991). The amplitudes of the disturbances are 0.05 for the two-dimensional and 0.15 for the oblique instability waves. In this paper we do not study compressibility effects, so a low convective Mach number $M_R = 0.2$ is used (Vreman *et al.*, 1994c), while Re = 50, following Comte *et al.* (1992).

The DNS database has been obtained by solving the Navier–Stokes equations, without any subgrid model, on a fine grid with 192³ cells. The numerical method employs a second-order accurate four-stage Runge–Kutta method for marching in time (Vreman *et al.*, 1992), whereas the spatial derivatives are discretized using a fourth-order scheme for the convective and a second-order scheme for the viscous terms (Vreman *et al.*, 1994c). A higher-order scheme for the marching in time is not required, since the trunctation errors introduced by the discretization of the time derivatives are considerably smaller than the truncation errors introduced by the spatial discretizations. Visualization of the DNS demonstrates the roll-up of the fundamental instability and successive pairings. Four rollers with mainly negative spanwise voriticity are observed at t = 20. After the first pairing (t = 40) the flow has become highly three-dimensional. Another pairing (t = 80) yields a single roller in which the flow exhibits a complex structure, with many regions of positive spanwise vorticity (Vreman *et al.*, 1993). This structure is an effect of the transition to turbulence which has been triggered by the pairing process at t = 40 (Moser and Rogers, 1993). The simulations are stopped at t = 100, since the single roller at t = 80 cannot undergo another pairing. The accuracy of the



Figure 1. Results of LES with the Clark model (solid), the filtered DNS (circles), and LES without a subgrid-model (dashed-dotted). The evolution of the total kinetic energy (a) and the evolution of the momentum thickness (b) are shown.

simulation was found to be satisfactory. First, the linear growth rates of the dominant instability modes were captured within 1%. Furthermore, in a simulation on a coarser grid (128³ cells) the results for the kinetic energy and momentum thickness were identical within plotting accuracy, whereas deviations in local quantities (e.g., slices through the vorticity field in the turbulent regime) were small.

The LES presented in this section have been performed on a grid with 32^3 cells, which is considerably coarser than the DNS grid, and a filter width $\Delta = L/16$ is used. Thus the filter width Δ equals twice the grid spacing of the coarse grid, which is an optimal choice (Vreman *et al.*, 1994d). The simulation is performed with the same numerical method as the DNS. In order to compare the LES and DNS results the DNS results are filtered on the fine grid. Next, the filtered data is easily obtained on the coarse grid through restriction of the filtered fine-grid data. This procedure in particular provides the initial conditions of the LES.

Figure 1 shows the total kinetic energy E and the momentum thickness d calculated from the LES with the Clark model. The quantities E and d are defined as

$$E = \int_{\Omega} \frac{1}{2} \bar{\rho} \tilde{u}_i \tilde{u}_i \, d\mathbf{x},\tag{24}$$

$$d = \int_{-L/2}^{L/2} \langle \bar{\rho} \rangle (1 - \langle \tilde{u}_1 \rangle) (1 + \langle \tilde{u}_1 \rangle) dx_2, \tag{25}$$

where $\langle \cdot \rangle$ denotes averaging over the homogeneous x_1 - and x_3 -directions. The curves corresponding to the filtered fine-grid DNS and LES without incorporating a subgrid model ("coarse-grid DNS") have been included in Figure 1. Comparison with the fine-grid DNS results indicates that the behavior of the subgrid model is poor; LES without a subgrid model is better than LES with the Clark model. More specifically, the subgrid model is too dissipative: the decay of the total kinetic energy and the growth of the momentum thickness in the initial stages are too large. This behavior is caused by the Smagorinsky part of the Clark model. Indeed, the pure Smagorinsky model also exhibits an excessive dissipation in the transitional regime (Piomelli *et al.*, 1990: Vreman *et al.*, 1994d).

Since the Clark model is too dissipative in the transitional regime, we also perform a simulation using the gradient model, i.e., omitting the eddy-viscosity part in the Clark model. Figure 2 shows the evolution of the total kinetic energy of an LES with this subgrid model. The behavior in the transitional regime is more



Figure 2. The evolution of the total kinetic energy for LES with the gradient model (solid), the filtered DNS (circles), and LES without a subgrid-model (dashed-dotted).

accurate than LES without a subgrid model. However, the simulation becomes unstable, which is illustrated by the strong increase of the total kinetic energy after t = 70. The reason for this instability cannot simply be that the grid is too coarse, because the coarse-grid DNS which employs the same grid remains stable. Furthermore, similar to findings for the similarity model (22) (Vreman *et al.*, 1994d), the gradient model can dissipate energy, since until t = 70 the decay of energy is stronger than in the coarse-grid DNS. In *a priori* tests the gradient model gives correlations which are as high or even higher than those of the similarity model (Vreman *et al.*, 1993). However, *a posteriori* tests show that although both models are able to dissipate energy, the gradient model causes stability problems, whereas the similarity model gives a stable simulation with acceptable results (Vreman *et al.*, 1994d).

Summarizing we have the following shortcomings of the Clark model: if the eddy-viscosity in incorporated, it is excessively dissipative in the transitional regime, whereas without the eddy-viscosity the LES is unstable. In the next section we analyze the nature of this instability for a one-dimensional model problem. Then in Section 5 a new version of the Clark model is presented, resulting in a simulation which is both stable and considerably more accurate than the original Clark model.

4. Instabilities in the Burgers Equation Using the Gradient Subgrid Model

It is well known that the Burgers equation is a simple model that describes flow phenomena which are in several respects qualitatively similar to Navier–Stokes flows. Since the Burgers equation is a onedimensional scalar equation, mathematical analysis is often possible. The equation has been the subject of several studies of LES (Love, 1980; Humi, 1990). In this section we examine the Burgers equation supplemented with the one-dimensional version of the gradient part of the Clark model. The linear stability of a sinusoidal profile is investigated, in order to gain some understanding of the instability of the gradient model encountered in the previous section. The connection between linear and nonlinear stability is known for Navier–Stokes flows and has been formulated in the following way (Henningson and Reddy, 1994). If a flow is linearly unstable, then it is nonlinearly unstable to arbitrarily small initial disturbances. On the other hand, if a flow is linearly stable, then it is nonlinearly stable, provided the initial disturbance is sufficiently small. Linear analysis thus provides information on the nonlinear equation.

The Burgers equation with the gradient subgrid model is written as

$$\partial_t u + \frac{1}{2} \partial_x (u^2) - v \partial_x^2 u = -\frac{1}{2} \eta \partial_x (\partial_x u)^2 + f(x), \tag{26}$$

where ∂_t and ∂_x denote the time and spatial derivative, respectively, u is the one-dimensional velocity, and

v is the viscosity. The left-hand side of this equation contains all terms in the standard Burgers equation. The right-hand side represents the gradient model with parameter $\eta = \frac{1}{2}\Delta^2$ plus a forcing term f.

The following analysis shows that smooth solutions of (26) can be extremely sensitive to small perturbations, leading to severe instabilities. In particular, we consider the linear stability of a 2π -periodic stationary solution, $U(x, t) = \sin(x)$, on the domain $[0, 2\pi]$ with periodic boundary conditions. The forcing function f is determined by the requirment that U is a solution of (26). We notice that no forcing is needed to ensure that U is an exact solution for the inviscid case with $\eta = 1$. We substitute a superposition of U and a perturbation v,

$$u(x,t) = U(x) + v(x,t),$$
(27)

into (26) and linearize around U, omitting higher-order terms in v:

$$\partial_t v + (1 - \eta) \sin(x) \partial_x v + (v + \eta \partial_x^2 v) \cos(x) = v \partial_x^2 v.$$
(28)

We use the following Fourier expansion for v:

$$v = \sum_{k=-\infty}^{\infty} \alpha_k(t) e^{ikx},$$
(29)

where *i* is the imaginary unit with $i^2 = -1$. After substitution of this series into (28) and ordering of terms, we obtain an infinitie system of ordinary differential equations for the Fourier coefficients α_k :

$$\dot{\alpha}_{k} = \frac{1}{2}k(\eta k - \eta - 1)\alpha_{k-1} - k^{2}\nu\alpha_{k} + \frac{1}{2}k(\eta k + \eta + 1)\alpha_{k+1},$$
(30)

where k is a number in \mathbb{Z} .

The instability encountered in the previous section concerns the gradient model, in which no eddyviscosity is adopted, while the molecular viscosity is relatively small. Therefore, to understand the nature of this instability for the Burgers equation, we first analyse system (30) assuming v = 0. Instead of the infinite system, we consider a sequence of finite-dimensional systems,

$$\dot{\mathbf{z}}_n = M_n \mathbf{z}_n,\tag{31}$$

where z_n is a vector containing the 2n + 1 Fourier coefficients $a_{-n} \cdots a_n$ and M_n is a $(2n + 1) \times (2n + 1)$ tridiagonal matrix:

$$\mathbf{z}_{n} = \begin{bmatrix} \alpha_{-n} \\ \vdots \\ \alpha_{-1} \\ \alpha_{0} \\ \alpha_{1} \\ \vdots \\ \alpha_{n} \end{bmatrix}, \qquad M_{n} = \begin{bmatrix} 0 & l_{n} \\ r_{n} & \ddots & \ddots & \\ & \ddots & \ddots & l_{2} \\ & & r_{2} & 0 & l_{1} \\ & & & 0 & 0 & 0 \\ & & & l_{1} & 0 & r_{2} \\ & & & & l_{2} & \ddots & \ddots \\ & & & & & l_{2} & \ddots \\ & & & & & l_{2} & \ddots & \ddots \\ & & & & & & l_{n} & 0 \end{bmatrix},$$
(32)

with

$$l_k = \frac{1}{2}k(\eta k - \eta - 1), \tag{33}$$

$$r_k = \frac{1}{2}(k-1)(\eta k+1). \tag{34}$$

The eigenvalues of A_n determine the stability of the problem. The system is unstable if the maximum of the real parts of the eigenvalues is positive.

In the following we consider the asymptotic behavior for large values of n of the maximum of the real parts of the eigenvalues of M_n . We denote the eigenvalues of M_n by λ_i , whereas λ_{max} represents an eigenvalue with

$$|\lambda_{\max}| = \max|\lambda_i|. \tag{35}$$

The proof of the following three properties is given in the Appendix:

1. If
$$\lambda$$
 is an eigenvalues, then $-\lambda$ is an eigenvalue. (36)

$$2. |\lambda_{\max}| \sim \eta n^2. \tag{37}$$

3.
$$|\operatorname{Im}(\lambda_{\max})| \le n - 1. \tag{38}$$

The first point implies that λ_{max} can be chosen such that $\text{Re}(\lambda_{max}) \ge 0$. Hence, the combination of these three properties yields the asymptotic behavior of the maximum of the real parts of the eigenvalues:

$$\operatorname{Re}(\lambda_{\max}) \sim \eta n^2. \tag{39}$$

Thus we have shown that the inviscid system is linearly unstable and that the largest real part of the eigenvalues is asymptotically proportional to n^2 , where n is the number of Fourier modes taken into account.

It should be observed that the instability is severe, since the system is not only unstable, but the growth rate of the instability is infinitely large as $n \to \infty$. The instability is fully due to the incorporation of the gradient model, since all eigenvalues of the matrix M_n are purely imaginary in case the inviscid Burgers equation without subgrid model is considered ($\eta = 0$). In numerical simulations the instability will grow with a finite speed, since then the number of Fourier modes is limited by the finite grid. Moreover, expression (39) illustrates that grid refinement (with η kept constant), which corresponds to a larger n, will not stabilize the system, but only enhance the instability. In order to investigate whether three-dimensional simulations show features similar to this one-dimensional behavior, we have refined the grid of the unstable 32^3 LES in the previous section to 64^3 cells, using the same value for Δ . Hence, whereas $\Delta = 2h$ in the simulation on the 32³ grid, $\Delta = 4h$ in the 64³ LES, where h is the grid spacing. The 64³ LES with the gradient model collapsed at t = 21, which is about a quarter of the collapse time of the corresponding 32^3 -LES model. However, if the filter width Δ is reduced simultaneously when the grid is refined, i.e., Δ/h is kept constant, the instability is not necessarily enhanced. This is illustrated by a simulation on the 64^3 grid using the gradient model with $\Delta = 2h$. This simulation collapsed at approximately the same time as the 32³ LES using $\Delta = 2h$. In fact the growth rate of the instability of the one-dimensional problem can be expressed in Δ and the grid spacing h: $\eta n^2 \sim (\Delta/h)^2$. Consequently, the instability is not enhanced if the ratio between Δ and h is kept constant.

Finally, we consider the more complicated case $v \neq 0$. The linear system in (30) now gives rise to matrices M_n which have a negative principal diagonal. It is known that for every fixed value of *n* there exists an eigenvalue arbitrarily close to the eigenvalue of the inviscid system (λ_{max}) if *v* is sufficiently small (Chatelin, 1993 (Lemma 4.3.1)). Hence for small values of *v* the viscous system for finite *n* is still linearly unstable. The matrix M_n is strictly diagonally dominant if

$$\nu > \eta + 1, \tag{40}$$

while all rows except n and n + 2 are already diagonally dominant if $v > \eta$. If the matrix is diagonally dominant, the real parts of all eigenvalues are negative and, consequently, the system is stable. This indicates that stability can be achieved by a sufficiently large viscosity, which does not depend on n, but only on η . Thus, we conclude that if the gradient model is supplemented with an adequate eddy-viscosity the instability will be removed. The original Clark model, which indeed contains enough eddy-viscosity to avoid instabilities, is too dissipative. In the next section an adequate eddy-viscosity formulation which stabilizes the gradient model, while not being too dissipative, is presented for three-dimensional LES.

5. The Dynamic Clark Model

In the previous sections we have encountered two problems caused by the Clark model. If the Smagorinsky eddy-viscosity is incorporated, the model is too dissipative. However, if the eddy-viscosity part is omitted, the simulation becomes unstable. The analysis in the previous section indicated that this instability is caused by the model, not by the numerical method, and can be overcome by sufficient dissipation. In this section it is shown that the dynamic procedure is able to solve both problems.

In order to overcome the excessive dissipation of the Smagorinsky model in laminar regimes the dynamic eddy-viscosity model has been developed by Germano (Germano, 1992) and has subsequently been applied to a number of flows (Moin and Jimenez, 1993). This model adopts Smagorinsky's eddy-viscosity formulation, but the square of the Smagorinsky constant is replaced by a coefficient which is dynamically obtained and depends on the local structure of the flow. Apart from the grid-filter level ("*F*-level"), denoted by the bar filter, Germano introduced a test filter (at the "G-level"), which is denoted by the hat (\hat{r}) and corresponds to a filter width 2Δ . The consecutive application of these two filters, resulting in, e.g., $\hat{\rho}$, defines a filter on the "*FG*-level," with which a filter width $\kappa\Delta$ can be associated. For top-hat filters the optimum value for κ equals $\sqrt{5}$ (Vreman *et al.*, 1994b). Next we consider Germano's identity, which reads

$$\hat{\bar{\rho}}T_{ij} - \bar{\bar{\rho}\tau_{ij}} = L_{ij}.$$
(41)

The right-hand side of (41) can be explicitly calculated from the variables on the F-level:

$$L_{ij} = \left(\frac{\overline{\rho u_i \rho u_j}}{\overline{\rho}}\right)^2 - \frac{\overline{\rho u_i \rho u_j}}{\overline{\rho}},\tag{42}$$

where $(\cdot)^{\hat{}}$ indicates that the hat filter is applied to the expression between the brackets. This notation is used in conjunction with the identically defined notation $(\hat{})$ for convenience in the exposure. The terms on the left-hand side of the Germano identity are the turbulent stress on the FG-level,

$$\hat{\rho}T_{ij} = \widehat{\overline{\rho u_i u_j}} - \frac{\overline{\widehat{\rho u_i \rho u_j}}}{\widehat{\hat{\rho}}}, \tag{43}$$

and the turbulent stress on the *F*-level filtered with the test filter, respectively. The terms on the left-hand side cannot be calculated from the variables on the *F*-level.

In order to obtain a dynamic eddy-viscosity model, Germano substituted the Smagorinsky eddyviscosity model into identity (41). Similarly, a dynamic mixed model has been formulated, which employs the mixed model by Bardina *et al.* as the base model (Zang *et al.*, 1993; Vreman *et al.*, 1994b). In the following we formulate the dynamic Clark model,

$$\bar{\rho}\tau_{ij} = \frac{1}{12}\Delta_k^2 \bar{\rho}\partial_k \tilde{u}_i \partial_k \tilde{u}_j - C_{\rm C}\Delta^2 \bar{\rho} |S(\tilde{\mathbf{u}})| S_{ij}(\tilde{\mathbf{u}}), \tag{44}$$

which resembles (21), with the model constant C_s^2 replaced by a model coefficient C_c , which is allowed to be a function of space and time. In order to obtain an expression for C_c , the subgrid model (equation (44)) is substituted into the Germano identity, which means that expressions for T_{ij} and τ_{ij} are obtained by formulating the subgrid model in FG-filtered quantities and F-filtered quantities, respectively. Denoting the Favre-filtered velocity on the FG-level by

$$v_i = \frac{\widehat{\overline{\rho u}_i}}{\widehat{\rho}},\tag{45}$$

the Clark model on the FG-level reads

$$\hat{\bar{\rho}}T_{ij} = \frac{1}{12}(\kappa\Delta_k)^2 \hat{\bar{\rho}}\partial_k v_i \partial_k v_j - C_{\rm C}(\kappa\Delta)^2 \hat{\bar{\rho}} |S(\mathbf{v})| S_{ij}(\mathbf{v}), \tag{46}$$

where $|S(\mathbf{v})|^2 = \frac{1}{2}S_{ij}^2(\mathbf{v})$. Substitution of (44) and (46) into the Germano identity (41) yields

$$H_{ij} + C_{\rm C} M_{ij} = L_{ij}, \tag{47}$$

where H_{ij} and M_{ij} are terms related to the gradient and eddy-viscosity part of the Clark model, respectively:

$$H_{ij} = \frac{1}{12} (\kappa \Delta_k)^2 \,\hat{\bar{\rho}} \partial_k v_i \partial_k v_j - \left(\frac{1}{12} \Delta_k^2 \bar{\rho} \partial_k \tilde{u}_i \partial_k \tilde{u}_j\right),\tag{48}$$

$$M_{ij} = -\hat{\bar{\rho}}(\kappa\Delta)^2 |S(\mathbf{v})| S_{ij}(\mathbf{v}) + (\bar{\rho}\Delta^2 |S(\tilde{\mathbf{u}})| S_{ij}(\tilde{\mathbf{u}}))^2.$$
(49)

To obtain the expression for M_{ij} we have neglected variations of $C_{\rm C}$ on the scale of the test-filter width. Since, (47) represents a system of five equations for the single unknown $C_{\rm C}$, a least-squares approach (Lilly, Large-Eddy Simulation of the Temporal Mixing Layer Using the Clark Model

1992) is followed to calculate this coefficient,

$$C_{\rm c} = \frac{\langle M_{ij}(L_{ij} - H_{ij}) \rangle}{\langle M_{ij} M_{ij} \rangle}.$$
(50)

In order to prevent numerical instability, caused by negative values of $C_{\rm C}$, the numerator and denominator in (50) are averaged over the homogeneous directions, which is expressed by the symbol $\langle \cdot \rangle$. In the simulations reported here the coefficient $C_{\rm C}$ in the dynamic Clark model is artificially set to zero at locations where the right-hand side of (50) returns negative values. It has, however, been checked that simulations in which negative values of $C_{\rm C}$ are allowed give rise to negligible differences.

This subgrid model satisfies the requirements following from the analysis in the previous section, if the coefficient $C_{\rm C}$ is sufficiently large. If on a given grid the model is stable, it is expected to remain stable after grid refinement. Indeed, the magnitude of the eddy-viscosity part of the dynamic Clark model, $C_{\rm C}\Delta^2 \bar{\rho} |S(\tilde{\mathbf{u}})|$, varies only slightly with the mesh size when keeping Δ constant. Hence, this model enables a separation of numerical errors from modeling errors.

We compare results of the dynamic Clark model with the dynamic eddy-viscosity model (Germano, 1992) and the dynamic mixed model (Zang *et al.*, 1993; Vreman *et al.*, 1994b). The dynamic eddy-viscosity model employs the Smagorinsky model as the base model, and the dynamic model coefficient is obtained using (47) with the tensor H_{ij} equal to zero. The dynamic mixed model employs the mixed model by Bardina *et al.* as the base model; in that case the tensor H_{ij} in (47) is replaced by a tensor using the similarity model on different levels (Vreman *et al.*, 1994b). With respect to computational efficiency the dynamic Clark model does not require much more work than the dynamic eddy-viscosity model, unlike the dynamic mixed model. The derivatives of the filtered velocity needed for the gradient part are already calculated in order to obtain the strain rate at different filter levels. The extra work needed is mainly due to the six filtering operations in (48) indicated by $(\cdot)^{\circ}$. The calculation of H_{ij} for the dynamic mixed model, however, contains considerably more filtering operations, required for the evaluation of the similarity model at different filtering levels. Consequently, the computational cost for the dynamic mixed model is considerably higher than for the dynamic eddy-viscosity and dynamic Clark model.

Figure 3 shows results of LES for the three dynamic models, using the same grid and filter as described in Section 3. Three quantities are compared: the evolution of the total kinetic energy E and the momentum thickness d, and the profile of the turbulent kinetic energy k, which is defined as

$$k(x_2) = \frac{1}{2} \langle (\tilde{u}_i - \langle \tilde{u}_i \rangle)^2 \rangle.$$
⁽⁵¹⁾

The k-profile is calculated at t = 70, which is well beyond the starting point of the mixing transition process, but just before the final pairing has been accomplished. The time t = 70 is rather arbitrary; it could equally well be t = 60 or t = 80, without altering the conclusions. All three models shows an improvement over the coarse-grid simulation without any subgrid model, indicating that in all three cases subgrid modeling makes sense. Furthermore, the results for the dynamic Clark and dynamic mixed model are much closer to the filtered DNS results than the results of the dynamic eddy-viscosity model, in particular for the evolution of the total kinetic energy and the momentum thickness.

Figure 4 shows the evolutions of the value of the dynamic model coefficient for the three dynamic models. As can be inferred from (50) the model coefficient is a function of time and the normal direction x_2 only. The results in Figure 4 show C_c averaged over the x_2 -direction. The dynamic eddy viscosity model gives rise to higher values for the model coefficient than the other two models. The reason is that the nonedddy viscosity parts of the dynamic Clark and dynamic mixed model already provide some dissipation and thus reduce the role of the eddy-viscosity part. Furthermore, it is observed that the similarity part in the dynamic mixed model is more dissipative than the gradient part in the dynamic Clark model, since the model coefficient is lower in the first case.

Hence, the dynamic Clark and dynamic mixed model lead to comparable simulation results, which are considerably more accurate than can be obtained with the dynamic eddy-viscosity model. Moreover, the dynamic procedure applied to the Clark model solves the problems encountered in Section 3. Firstly, the dynamic Clark model provides proper dissipation, thus avoiding the excessive dissipation of the original Clark model. On the other hand, the dissipation in the dynamic Clark model is sufficient to overcome the stability problem of the gradient model. We have verified that LES with the dynamic Clark model is also



Figure 3. Results of LES with the dynamic Clark model (solid), the dynamic eddy-viscosity model (dashed), the dynamic mixed model (dotted), the filtered DNS (circles), and LES without a subgrid-model (dashed-dotted). The evolution of the total kinetic energy (a), the evolution of the momentum thickness (b), and the profile of the turbulent kinetic energy at t = 70 (c) are shown.

stable at other grids, both for $\Delta = 4h$ keeping Δ constant, and for both Δ and h taken smaller and retaining $\Delta = 2h$. These findings are in agreement with the analysis in Section 4.

6. Conclusions

In this paper we have investigated the Clark subgrid-scale model, which is the sum of the gradient model, based on Taylor expansions, and the Smagorinsky eddy-viscosity model. We have extended the formal derivation of the incompressible Clark model to be able to use the model in compressible flow simulations as well. In LES of the weakly compressible mixing layer, the Clark model is too dissipative when the



Figure 4. The averaged model coefficient for LES with the dynamic Clark model (solid), the dynamic eddy-viscosity model (dashed), and the dynamic mixed model (dotted).

eddy-viscosity part is incorporated. However, the model destabilizes the simulation if the eddy-viscosity part is omitted. In order to analyse the gradient part of the Clark model and to clarify the nature of the instability, the one-dimensional Burgers equation supplemented with the gradient model has been considered. A linear stability analysis of a sinusoidal profile of the modified Burgers equation has been performed. In the limit of vanishing viscosity, the maximum of the positive real parts of the eigenvalues has been shown to be proportional to the square of the number of modes, if a finite number of modes is taken into account. This shows that the instability becomes more severe, if the resolution is increased. For an infinite number of modes, the growth-rate of the instability is infinitely large. These analytical results for the one-dimensional case are similar to the behavior of the instability observed in three-dimensional simulations with the gradient model. Furthermore, the one-dimensional analysis indicates that a sufficient amount of viscosity is able to solve the stability problem. For these reasons we have introduced the dynamic Clark model, which is the sum of the gradient model and a dynamic eddy-viscosity model. Simulations using this new model show that the dissipation is sufficient to prevent the instability encountered in simulations employing the pure gradient model. On the other hand the dynamic Clark model avoids the excessive dissipation in the transitional regime observed in simulations using the original Clark model.

The results obtained with the dynamic Clark model are reasonably accurate as comparison with filtered DNS results indicates. Furthermore, the dynamic Clark model has been compared with the standard dynamic eddy-viscosity model and the recently developed dynamic mixed model. Compared with the dynamic eddy-viscosity model, the dynamic Clark model gives more accurate predictions. Compared with the recently developed dynamic mixed model, the results are of comparable accuracy. With respect to computational costs the dynamic Clark model is more efficient than the dynamic mixed model, since considerably less extra filtering operations are required in order to evaluate the model. We conclude that the dynamic Clark model has been found to give good results for the weakly compressible temporal mixing layer. Understanding this model will further increase if it is also investigated in LES of different flows.

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Appendix

In this appendix the proof of the three statements in (36)–(38) concerning the eigenvalues of the matrix M_n is given. The structure of the matrix M_n is such that an eigenvalue is equal to zero or an eigenvalue of the following matrix:

$$A_{n} = \begin{bmatrix} 0 & r_{2} & & & \\ l_{2} & \ddots & \ddots & & \\ & \ddots & \ddots & r_{n-1} & \\ & & l_{n-1} & 0 & r_{n} \\ & & & & l_{n} & 0 \end{bmatrix}.$$
 (A.1)

Consequently, to consider the eigenvalues of A_n is sufficient. These eigenvalues are roots of the characteristic polynomial $P_n(\lambda)$ of the matrix A_n :

$$P_n(\lambda) = \det(\lambda I_n - A_n), \tag{A.2}$$

where I_n is the $n \times n$ unity matrix. If we decompose this determinant with respect to the last column, we obtain

$$P_1(\lambda) = \lambda, \tag{A.3}$$

$$P_2(\lambda) = \lambda^2 - r_2 l_2, \tag{A.4}$$

$$P_n(\lambda) = \lambda P_{n-1}(\lambda) - r_n l_n P_{n-2}(\lambda), \qquad n > 2.$$
(A.5)

This recursive relation demonstrates that $P_n(\lambda) = 0$ implies $P_n(-\lambda) = 0$. Thus if λ is an eigenvalue, $-\lambda$ is also an eigenvalue and thus property (36) has been established.

Next, we consider the statement about the asymptotic behavior of $|\lambda_{max}|$. In order to prove this statement we formulate a lower and upper bound for $|\lambda_{max}|$ which are both asymptotically proportional to ηn^2 . An upper bound is obtained when the Gershgorin theory is applied to A_n :

$$|\lambda_{\max}| \le l_{n-1} + r_n = \eta (n-1)^2 \sim \eta n^2.$$
(A.6)

Since $|\lambda_{\max}| > |\lambda_i|$ for i = 1, ..., n and λ_i represents one of the *n* roots of the polynomial $P_n(\lambda)$, the following relation provides a lower bond for $|\lambda_{\max}|$:

$$|\lambda_{\max}| \ge |\lambda_1 \lambda_2 \cdots \lambda_n|^{1/n} = |P_n(0)|^{1/n}.$$
(A.7)

The recursive relation (A.5) provides

$$P_n(0) = -r_n l_n P_{n-2}(0), \tag{A.8}$$

with

$$r_n l_n = \frac{1}{4} (n-1)n(\eta^2 (n^2 - n) - \eta - 1) \sim \frac{1}{4} \eta^2 (n-1)^2 n^2.$$
(A.9)

If n is even and the coefficient l_k is nonzero for even k (r_k is always nonzero), we obtain the following estimate, using Stirling's formula:

$$|\lambda_{\max}| \ge |P_n(0)|^{1/n} \sim ((\frac{1}{4}\eta^2)^{n/2} (n!)^2)^{1/n} \ge \frac{1}{2}\eta \left(\frac{n}{e}\right)^2 \sim \eta n^2.$$
(A.10)

Equations (A.6) and (A.10) together yield

$$|\lambda_{\max}| \sim \eta n^2. \tag{A.11}$$

The case for odd *n* is more complicated, but does not need to be considered, since the behavior for even *n* already provides sufficient information about the system. Furthermore, the above argument alters when $l_k = 0$ for a certain (even) value of k. If $l_k = 0$, the characteristic polynomial can be written as

$$P_n(\lambda) = P_{k+1}(\lambda)Q_{n-k+1}(\lambda), \tag{A.12}$$

where $Q_{n-k+1}(\lambda)$ represents the characteristic polynomial of the tridiagonal matrix with lower diagonal $l_{k+1} \cdots l_n$ and upper diagonal $r_{k+1} \cdots r$. The maximum root of Q_{n-k+1} can be estimated like in (A.10) and its asymptotic behavior is also proportional to ηn^2 .

Finally, we derive an upper bound for the imaginary part of an eigenvalue λ of A_n . Denoting the corresponding eigenvector with y we have

$$A_n \mathbf{y} = \lambda \mathbf{y}.\tag{A.13}$$

The matrix A_n can be split into a symmetric matrix B_n and an antisymmetric matrix C_n :

$$A_n = B_n + C_n \tag{A.14}$$

with

$$B_{n} = \begin{bmatrix} 0 & b_{2} & & \\ b_{2} & \ddots & \ddots & \\ & \ddots & \ddots & b_{n} \\ & & b_{n} & 0 \end{bmatrix}, \qquad Cn = \begin{bmatrix} 0 & c_{2} & & \\ -c_{2} & \ddots & \ddots & \\ & \ddots & \ddots & c_{n} \\ & & -c_{n} & 0 \end{bmatrix},$$
(A.15)

and

$$b_k = \frac{1}{2}(r_k + l_k) = \frac{1}{2}\eta k^2 - \frac{1}{2}\eta k + \frac{1}{4},$$
(A.16)

$$c_k = \frac{1}{2}(r_k - l_k) = \frac{1}{2}k - \frac{1}{4}.$$
(A.17)

In the following we denote the Euclidean inner product in \mathbb{C}^n by (\cdot, \cdot) and the complex conjugate by the superscript *. Next, we derive the following relation between λ and λ^* :

$$\lambda(\mathbf{y}, \mathbf{y}) = (A_n \mathbf{y}, \mathbf{y})$$

$$= (B_n \mathbf{y}, \mathbf{y}) + (\mathbf{C}_n \mathbf{y}, \mathbf{y})$$

$$= (\mathbf{y}, B_n^T \mathbf{y}) + (\mathbf{y}, C_n^T \mathbf{y})$$

$$= (\mathbf{y}, B_n \mathbf{y}) - (\mathbf{y}, C_n \mathbf{y})$$

$$= (\mathbf{y}, A_n \mathbf{y}) - 2(\mathbf{y}, C_n \mathbf{y})$$

$$= \lambda^* (\mathbf{y}, \mathbf{y}) - 2(\mathbf{y}, C_n \mathbf{y}).$$
(A.18)

Now we can express an upper bound for the imaginary part of λ as

$$|\operatorname{Im}(\lambda) = \frac{1}{2}|\lambda - \lambda^{*}| = \left|\frac{(\mathbf{y}, C_{n}\mathbf{y})}{(\mathbf{y}, \mathbf{y})}\right| \le \frac{\|\mathbf{y}\| \|C_{n}\mathbf{y}\|}{\|\mathbf{y}\|^{2}} \le \|C\|_{\mathrm{sp}},\tag{A.19}$$

since $||C_n \mathbf{y}|| \le ||C_n||_{sp} ||\mathbf{y}||$, where the subscript sp denotes the spectral matrix norm. The spectral matrix norm $||C||_{sp}$ is defined as the square root of the maximum of the absolute eigenvalues of the matrix $C_n^T C_n$, which is equal to the maximum absolute eigenvalue of C_n , since C_n is antisymmetric. Finally, the following result is obtained using the Gershgorin theory:

$$|\operatorname{Im}(\lambda)| \le |c_{n-1}| + |c_n| = n - 1, \tag{A.20}$$

which expresses an upper bound for the imaginary parts of the eigenvalues.

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