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Large-scale Mass Wasting on the Miocene Continental Margin of Western India

BULLETIN

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Abstract:

A giant mass transport complex was recently discovered in the eastern Arabian Sea, exceeding in volume all but one other known complex on passive margins worldwide. The complex, named the Nataraja Slide, was drilled by International Ocean Discovery Program (IODP) Expedition 355 in two locations where it is \sim 300 m (Site U1456) and \sim 200 m thick (Site U1457). The top of this mass transport complex is defined by the presence of both reworked microfossil assemblages and deformation structures, such as folding and faulting. The deposit consists of two main phases of mass wasting, each which consists of smaller pulses, with generally fining-upward cycles, all emplaced just prior to 10.8 Ma. The base of the deposit at each site is composed largely of matrix-supported carbonate breccia that is interpreted as the product of debris flows. In the first phase, these breccias alternate with well-sorted calcarenites deposited from a high energy current, coherent limestone blocks that are derived directly from the Indian continental margin, and a few clastic mudstone beds. In the second phase, at the top of the deposit, muddy turbidites dominate and become increasingly more siliciclastic. At Site U1456, where both phases are seen, a 20 m section of hemipelagic mudstone is present, overlain by a ~40 m thick section of calcarenite and slumped interbedded mud and siltstone. Bulk sediment geochemistry, heavy- mineral analysis, clay mineralogy, isotope geochemistry, and detrital zircon U-Pb ages constrain the provenance of the clastic, muddy material to being reworked Indus-derived sediment, with input from western Indian rivers (e.g., Narmada and Tapti Rivers), and some material from the Deccan Traps. The carbonate blocks found within the breccias are shallow-water limestones from the outer western Indian continental shelf that was oversteepened from enhanced clastic sediment delivery during the mid-Miocene. The final emplacement of the material was likely related to seismicity as there are modern analogues for intraplate earthquakes close to the source of the slide. Although we hypothesize this area is at low risk for future mass wasting events, it should be noted that other oversteepened continental margins around the world could be at risk for mass failure as large as the Nataraja Slide.

1	Large-scale Mass Wasting on the Miocene Continental Margin of Western
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69 Abstract

70 A giant mass-transport complex was recently discovered in the eastern Arabian Sea, 71 exceeding in volume all but one other known complex on passive margins worldwide. The 72 complex, named the Nataraja Slide, was drilled by International Ocean Discovery Program 73 (IODP) Expedition 355 in two locations where it is ~300 m (Site U1456) and ~200 m thick (Site 74 U1457). The top of this mass transport complex is defined by the presence of both reworked 75 microfossil assemblages and deformation structures, such as folding and faulting. The deposit consists of two main phases of mass wasting, each which consists of smaller pulses, with 76 77 generally fining-upward cycles, all emplaced just prior to 10.8 Ma. The base of the deposit at 78 each site is composed largely of matrix-supported carbonate breccia that is interpreted as the product of debris flows. The first phase, these breccias alternate with well-sorted calcarenites 79 80 deposited from a high energy current, coherent limestone blocks that are derived directly from 81 the Indian continental margin, and a few clastic mudstone beds. In the second phase, at the top of 82 the deposit, muddy turbidites dominate and become increasingly more siliciclastic. At Site 83 U1456, where both phases are seen, a 20-m section of hemipelagic mudstone is present, overlain 84 by a ~ 40 m thick section of calcarenite and slumped interbedded mud and siltstone. Bulk 85 sediment geochemistry, heavy--mineral analysis, clay mineralogy, isotope geochemistry, and 86 detrital zircon U-Pb ages constrain the provenance of the clastic, muddy material to being 87 reworked Indus-derived sediment, with input from western Indian rivers (e.g., Narmada and 88 Tapti Rivers), and some material from the Deccan Traps. The carbonate blocks found within the 89 breccias are shallow-water limestones from the outer western Indian continental shelf-that was oversteepened from enhanced clastic sediment delivery during the mid-Miocene. The final 90 \bigcirc 91 emplacement of the material was likely related to seismicity as there are modern analogues for

92 intraplate earthquakes close to the source of the slide. Although we hypothesize this area is at
93 low risk for future mass wasting events, it should be noted that other oversteepened continental
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95

96 INTRODUCTION

97 Large-scale mass wasting of continental margins is an important process in controlling the geomorphology of continental slopes fringingall ocean basins (Coleman and Prior, 1988). 98 99 The scale of large mass transport complexes (MTCs) makes them significant as geohazards, 100 directly through mass wasting (Dan et al., 2007; Yamada et al., 2012), by generating tsunamis 101 (Tappin et al., 2001), as well as posing risks for seafloor infrastructure such as oil and gas 102 platforms, pipelines (Bea et al., 1983), and communication cables (Hsu et al., 2008). Moreover, 103 the emplacement of MTCs can have significant influence on the stratigraphy of deep ocean 104 basins, as well as for the continental margin from which it was derived. 105 Although the largest mass transport deposits are associated with active margins (Burg et 106 al., 2008), where earthquakes are more common and can act as triggers for emplacement, passive 107 margins are also recognized to host some of the largest gravitational collapses in the modern 108 oceans (Embley and Jacobi, 1977). Seismic surveying in the eastern Arabian Sea offshore 109 western India has identified one of the largest such complexes, totaling around 19,000 km³ 110 (Calvès et al., 2015). Mapping of the deposit by seismic methods suggests that it may be up to 111 800 m thick in places (Calvès et al., 2015). In 2015 this deposit was drilled by International 112 Ocean Discovery Program (IODP) during Expedition 355. During the expedition, the MTC was 113 sampled on its southern edge, where the thicknesses were considerably thinner (Pandey et al., 114 2016c)(Fig. 1). The deposit, named the Nataraja Slide, shows substantial run out from its inferred

\bigcirc

115 source regions offshore Saurashtra (Fig. 1), being emplaced ~ 500 km into the Indian Ocean. In 116 this study, we examine the sedimentary rocks recovered by IODP in order to infer the 117 depositional mechanisms active during emplacement. We further make inferences about what processes triggered its formation, which is dated as being just before 10.8 Ma (Pandey et al., 118 2016a). Are MTCs of this magnitude formed by the same processes that we see at much smaller 119 120 scales, or are these mega-scale complexes unique in their modes of emplacement and triggers? 121 Given the profound potential geohazards for human settlements in coastal regions, understanding 122 the origins and impacts of the Nataraja Slide MTC are of both great scientific and societal 123 significance.

124

125 GEOLOGY OF LARGE MTCS

126 Mass transport complexes are an extreme form of gravity induced sediment transport 127 (Hampton et al., 1996). Most submarine gravity driven sediment transport involves redeposition 128 of individual sediment particles suspended in water (e.g., in a turbidity current) or as a fluidized 129 sediment suspension (e.g., a debris flow or mud flow)(Pickering et al., 1986; Talling et al., 130 2012). Sediment may also be mobilized when the proportion of water is very low, such as a 131 slow--moving sediment grain flow or creep (Carter, 1975; Lowe, 1976). However, large volumes 132 of material can also be transported rapidly (hours to days) in the form of slope failures where 133 coherent masses of material can be transported by sliding, rolling, falling, and/or slumping 134 (Coleman and Prior, 1988). Slumps involve displacement of a stratigraphic package above a 135 concave-upward detachment surface and can leave the slumped material in a relatively 136 undisturbed state after removal from an area that then shows an arcuate scar (Hampton et al., 137 1996; Moore, 1961). Slumps differ from slides in that motion is along a pre-existing weakness,

138 such as a bedding plane or joint surface, but the displaced package can move as a coherent mass, 139 or can be become disaggregated depending on the length and speed of transport. Significant 140 progress has been made in understanding mass transport through outcrop studies, such as the 141 Carboniferous (Pennsylvanian) Ross Slide of Ireland (Martinsen and Bakken, 1990; Strachan, 142 2002), the Eocene of the Pyrenean foreland basin (Farrell, 1984), and the Pliocene of Sicily 143 (Trincardi and Argnani, 1990). In all examples, each MTC was emplaced over a sharply defined 144 basal décollement once the deposit reached the lower slope after erosive mass wasting of the 145 steeper upper slope.

146 The geometry and internal structure of any gravitationally driven slump, slide or debris 147 flow reflect the mechanism of failure and the morphology of the slope where the transport occurs 148 (Lucente and Pini, 2003). The style of deformation and the mode of transport are controlled by 149 sediment and rock rheology that in turn are dependent on the lithology and strain rate. For this 150 reason, the largest MTCs are different from shallow debris flows and slumps because they 151 incorporate both lithified and unconsolidated materials. There are few exposures of very large 152 MTCs and those in the oceans are hard to access, especially through drilling. MTCs are often 153 seismically homogeneous (Vardy et al., 2010) but can show important changes in sediment 154 facies with depth and with distance from their source. For example, swath bathymetric mapping 155 of the Ebro margin in the western Mediterranean featuring the pre-11 ka BIG'95 Slide shows 156 that only finer sediments have reached the most distal areas, yet coherent rafts of continental 157 margin sedimentary rock are seen at the base of the slope (Lastras et al., 2004). Analysis of the geometry and distribution of sedimentary facies and structures can be used to reconstruct the 158 159 evolving sedimentary and deformational strain history of any individual MTC. By doing so, it is 160 possible to derive a kinematic model of emplacement that can be compared with other examples.

161 The Storegga Slide in offshore Norway is one of the best studied large-volume mass 162 transport complex. This MTC is entirely siliciclastic and its generation has been linked to sliding on ourite sand and silts that became overpressured as a result of rapid burial by glacial 163 164 maximum aged debris-flow sediments (Bryn et al., 2005). However, rapid sedimentation on any 165 clastic margin receiving sediment from the continent would provide weak layers on which 166 sliding could occur. Overpressuring has also been linked to growth and migration of silica 167 diagenetic fronts (Davies and Clark, 2006). Slope oversteepening increases the chances of mass 168 wasting simply by the consequence of rapid sediment delivery, although the tendency may be 169 heightened by the pre-existing basement structure of the continental margin (Lastras et al., 2004). 170 Slope oversteepening by itself cannot explain large-scale mass wasting because giant MTCs on 171 European continental margins are mostly associated with low gradient glacial margins. In 172 contrast, turbidity currents appear to dominate on steeper non-glacial margins which might 173 otherwise be expected to suffer mass wasting due to their gradient (Leynaud et al., 2009). In 174 these cases, differences in the sediment types and the timing of sediment delivery favor 175 gravitational instabilities at different times, with non-glaciated margins tending to mass waste 176 more during sealevel lowstands, where the opposite more often occurs on glaciated margins. 177 Modelling indicates that continental margins with more cohesive clay-rich sediments tend to 178 experience coherent sliding more frequently than sand-rich margins whose gravitational slides 179 tend to disintegrate in turbidity currents (Elverhoi et al., 2010). 180 The triggering of MTC emplacement can be attributed to a number of potential processes, 181 including seismicity (Moernaut et al., 2007; Piper et al., 1985), volcanic eruptions (Carracedo,

182 1999) and meteorite impacts (Klaus et al., 2000; Parnell, 2008). Dissociation of gas hydrates

183 during times of warming seawater could have aided liquefaction in the case of Storegga Slide

(Mienert et al., 2005), with seismicity possibly related to post-glacial isostatic rebound providing
the final impetus for redeposition (Evans et al., 2002). In the eastern Mediterranean Sea, MTC
emplacement has also be linked to biogenic gas and slope oversteepening acting individually or
in tandem with one another (Frey Martinez et al., 2005).

188 Mechanisms for MTC emplacement differ between clastic and carbonate margins. This is 189 because carbonate sediment production occurs *in situ* and can result in steep platform margins, 190 sometimes almost vertically where reef complexes develop in outer shelf areas. Carbonate 191 production is strongly linked to sealevel and was fastest when sealevel was high after the onset 192 of Northern Hemispheric Glaciation (NHG, ~2.4 Ma)(Schlager et al., 1994). Many carbonate 193 MTCs are linked to platform margin collapse and result in deposits with numerous coherent 194 blocks suspended within a more fluidized matrix. Seismic mapping around the Great Bahama 195 Bank has identified coherent Plio-Pleistocene sedimentary rock rafts 0.5–2.0 km in length, 0.3– 196 1.5 km in width, and 50 m in thickness (Principaud et al., 2015). Adjacent deposits have also 197 been observed on the Florida margin (Mullins et al., 1986), as well as offshore Nicaragua (Hine, 198 1992), all with a similar Plio-Pleistocene age. Plio-Pleistocene MTCs are larger than most known 199 older examples because the rapidly changing sealevel since the start of the NHG enhanced 200 carbonate production and induced gravitational instability as sealevel rose and fell (Schlager et 201 al., 1994). Among these older deposits, only the Cretaceous Ayabacas MTC of Peru is 202 noteworthy for its large volume, long run out and presence of slide blocks measuring kilometers 203 in length (Callot et al., 2008).

204

205 GEOLOGICAL SETTING

206 The Nataraja Slide lies within the Laxmi Basin offshore the western continental margin 207 of India (Fig. 1A and B). The Laxmi Basin is separated from the main Arabian Basin by the 208 Laxmi Ridge (Fig. 1). The Laxmi Basin is a rift basin that formed between India and the Laxmi 209 Ridge prior to the opening of the main Arabian Sea in the early Paleoeene (Bhattacharya et al., 210 1994), where the ridge is generally interpreted to be a rifted fragment of Indian continental crust 211 (Pandey et al., 1995). The age of rifting is somewhat controversial, but likely just predates the 212 emplacement of the Deccan Traps flood basalts in the latest Cretaceous, based on analysis of 213 magnetic anomalies (Bhattacharya et al., 1994) and the geochemistry of the basalts sampled at 214 IODP Site U1457 (Pandey et al., 2016b). The sediments in the Laxmi Basin can be divided into 215 three major units described below. The oldest, dated as Lower Paleocene, largely comprises red-216 brown mudstones eroded from peninsular India and sampled at IODP Site U1457 (Pandey et al., 217 2016b). These deposits are overlain by the Nataraja Slide and by younger distal turbidite 218 sandstones and siltstones, as well as hemipelagic mudstones that form the Indus submarine fan. 219 These latter sediments were supplied through the Indus River via erosion from the western 220 Himalaya and Karakoram (Pandey et al., 2016c). The age of the Indus Fan in the Laxmi Basin is 221 not well defined, although within the main Arabian basin the fan is typically considered to date 222 from at least 45 Ma, continuing to the present time (Clift et al., 2001). It is within these deposits 223 that the Nataraja Slide (MTC) was emplaced just before $\overline{10.8}$ Ma.

Towards the east, the Laxmi Basin is bounded by the rifted passive margin of India, which has been supplied by sediment from the erosion of the peninsula via a number of significant rivers that drain towards the west (e.g., Mahi, Tapti, and Narmada). Oil exploration drilling has furthermore identified significant repeated buildups of carbonate on the shelf, especially towards the shelf edge where the supply of clastic material was more limited (Rao and

229 Talukdar, 1980; Wandrey, 2004). It is generally presumed that extensional deformation in the 230 area ceased after the rifting that formed the Laxmi Basin. The area has been largely seismically 231 inactive except towards the north where the Rann of Kutch forms an active structure within the 232 Indian Craton. This structure is linked to flexure of the plate as a result of the collision between 233 India and Asia (Bilham et al., 2003; Biswas, 2005), presumed to have started in the Eocene 234 (Najman et al., 2010) or even earlier (DeCelles et al., 2014). Towards the north, the Indian 235 peninsula is cut by the NE-SW-trending Cambay Basin which formed as an initial early 236 Cretaceous rift that was then reactivated in the Cenozoic and experienced significant inversion in 237 the early Miocene (Chowdhary, 2004).

The MTC run-out distance is estimated to be about 550 km, with a rength of 338 km and 238 a maximum width of 193 km (Calvès et al., 2015). Prior work on the Nataraja Slide found this 239 240 MTC to be acoustically homogenous in seismic lines, with few identified rafts preser and to 241 have a flat, rather than significantly angular erosive base over older deposits (Fig. 2)(Calvès et 242 al., 2015; Pandey et al., 2016c). However, closer inspection in the vicinity of the drilling sites 243 finds this is not always the case. In the case of IODP Site U1456 where the slide is somewhat 244 thicker, there is a significant missing section of submarine fan turbidites from ~15.6 to 10.8 Ma 245 (Pandey et al., 2016a). In that area the upper part of the deposit appears to be more acoustically washed out and homogenous, but the lower regions are marked by strong reflections that show 246 247 limited lateral continuity suggestive of some internal structure within the deposit. This raises the 248 possibility that this is not simply a single depositional package (Fig. 2). Such strong reflections 249 are reminiscent of coherent slide blocks seen in seismic images of other MTCs (Gamboa et al., 250 2012; Krastel et al., 2012; Principaud et al., 2015). The same is not true at the more distal Site

U1457 location where the MTC onlaps the Laxmi Ridge and its acoustic character is moreuniform.

253

254 METHODS

Sedimentary cores were collected and initially described during IODP Expedition 355 255 256 but several cores are re-examined in order to obtain more detailed descriptions of critical 257 sedimentary structures and facies. In addition to preparing sedimentary logs designed to 258 highlight the contrasting sedimentary facies, samples for sediment petrography were examined to 259 allow investigation into the different sediment types at both the macro and microscopic scale. 260 These methods allowed us to better define the depositional processes that operated during 261 Nataraja Slide emplacement and to provide constraints on the origin(s) of the MTC. Geochemical without were employed in order to further constrain the provenance of the 262

Reterials, and in particular, to verify the proposed western Indian continental margin source for 263 much of the MTC argued by Calvès et al. (2015). This approach is predicated on the fact that 264 source rocks of MTC deposits have different bulk geochemical compositions and that Himalayan 265 266 sources can be effectively discriminated from peninsular sources when considering provenance 267 due to different bedrock source compositions and contrasting chemical weathering histories. 268 Forty-four samples were selected for determination of major element composition, 269 together with select trace elements (Ni, Ba, V, Zr, Sc, Y, Sr). These were determined by 270 inductively coupled plasma emission spectrometry (ICP-ES) at Boston University, with precision

quantified to be better than 2% of the measured value for all elements. Accuracy was constrained

272 by analysis of certified Standard Reference Materials (BHVO-2) and results were accurate within

273 precision. Table 1 provides analyses of samples as well as repeated analyses of the standard.

274 The neodymium (Nd) isotope compositions of sediments are generally considered to be 275 minimally affected by chemical weathering, such that source terranes faithfully translate their 276 isotopic signature to eroded sediments (i.e., Goldstein et al. (1984)) and can be utilized for 277 sedimentary provenance studies. Strontium (Sr) isotopes are additionally considered, while 278 recognizing that Sr isotope compositions may be affected by chemical alteration largely during 279 transport across flood plains (Derry and France-Lanord, 1996). Together these isotopic systems 280 have a record of being powerful provenance proxies in the Arabian Sea (e.g., (Clift and 281 Blusztajn, 2005; Clift et al., 2008a). Care was taken to decarbonate samples prior to analysis 282 with 20% acetic acid because Sr isotope compositions are strongly controlled by carbonate 283 compositions and this study targets the siliciclastic sediment compositions only. Decarbonation 284 lasted for six days until no further fizzing was observed when samples were exposed to 285 unreacted acid. Samples were washed by deionized water before being ground into powders. 286 Twenty-five samples were selected throughout the Nataraja Slide/MTC at Sites U1456 and U1457 for the determination of ¹⁴³Nd/¹⁴⁴Nd and ⁸⁷Sr/⁸⁶Sr values. Isotopic compositions were 287 288 determined by Finnigan Neptune multi-collector inductively coupled plasma mass spectrometer 289 (MC-ICP-MS) at the Woods Hole Oceanographic Institute for both Nd and Sr isotopes. Nd and Sr isotope analyses were corrected against La Jolla Nd standard ¹⁴³Nd/¹⁴⁴Nd=0.511847 and 290 291 NBS987 standard ⁸⁷Sr/⁸⁶Sr=0.710240. Procedural blanks were 20–25 pg for Sr and 50–70 pg for Nd. We calculate the parameter ϵ_{Nd} after (DePaolo and Wasserburg, 1976) using a ¹⁴³Nd/¹⁴⁴Nd 292 293 value of 0.512638 for the Chondritic Uniform Reservoir (CHUR) (Hamilton et al., 1983). Results are presented in Table 2. 294 295 Heavy-mineral analysis was applied to study the mineralogy of the MTC deposits in

296 order to further constrain the source of the materials and to estimate the potential impact of

297 diagenetic dissolution. Sediment left after thin section preparation was gently crushed in water 298 with mortar and pestle and wet-sieved using a standard 500 µm steel sieve and a special 299 handmade 15 µm tissue-net sieve. A wide size window (15–500 µm) was chosen to include a 300 large range of the size distribution (Garzanti et al., 2009). A gravimetric separation of dense 301 grains was achieved with a centrifuge using Na-polytungstate (density 2.90 g/cm³), and heavy 302 minerals recovered by partial freezing in liquid nitrogen. An appropriate amount of the dense 303 fraction thus obtained was split with a micro-riffle box and mounted with Canada balsam. Heavy 304 minerals were counted under a polarizing microscope with the area method (Mange and Maurer, 305 1992). Grains of uncertain character were systematically checked and identified by an inViaTM 306 Renishaw Raman spectrometer, equipped with a 532 nm laser and a 50x LWD objective (Andò 307 and Garzanti, 2014). Heavy-mineral and transparent-heavy-mineral concentrations (HMC and 308 tHMC indices of Garzanti and Andò (2007), representing fundamental parameters for 309 unravelling provenance and detecting hydraulic-sorting effects and diagenesis, allow us to 310 distinguish poor (tHMC < 1), and very rich (tHMC > 10) transparent-heavy-mineral suites. The 311 resulting assemblages were compared with those of modern sediments of the Tapti River 312 (sampled at 21°08'40.7" N, 72°44'08.1"E) and Indus River. Results are presented in Table 3. 313 U-Pb dating of detrital zircon has been widely used for provenance analysis in siliciclastic 314 systems because zircon is a common mineral in continental rocks of many compositions and is 315 chemically and mechanically resistant to weathering during transport (Carter and Bristow, 2003). Furthermore, zircon has a closure temperature of ~750°C for the U/Pb isotope system (Hodges 316 317 2003), making it very robust and unsusceptible to change during multiple stages of recycling. 318 Mineral separation and grain mounting were performed at GeoSep Services (GSS) Laboratory, 319 Moscow, ID. Only one sample was analyzed for zircon U-Pb dating because much of the core

320 lacked suitable layers for this method. Zircowere separated via hand picking and used for age 321 dating as described by Donelick et al. (2005). This process enhances the recovery of all possible 322 grain sizes while minimizing the potential loss of smaller grains within a sample by the use of 323 water-table devices. The method used by Donelick et al. (2005) further ensures the preservation 324 of complete grains by minimizing grain breakage and/or fracturing that can be associated with 325 traditional procedures of isolating individual grains from whole rock samples. Recovered zircon 326 were mostly medium silt to fine sand-sized grains. Epoxy wafers containing zircon grains for 327 laser ablation inductively coupled plasma mass spectrometry (LA-ICP-MS) were polished 328 manually using 3.0 µm and 0.3 µm Al₂O₃ slurries to expose internal zircon grain surfaces. The 329 polished grain surfaces were washed in 5.5 M HNO₃ for 20 sec. at 21°C in order to clean the 330 surfaces prior to introduction into the laser system sample cell.

331 A total of 51 individual zircon grains were targeted for data collection using a New Wave 332 YP213 213 nm solid state laser ablation system with a 20 µm diameter laser spot size, 5 Hz laser 333 firing rate, and ultra-high purity He as the carrier gas. Isotopic analyses of the ablated zircons 334 were performed using a ThermoScientific Element 2 magnetic sector mass spectrometer using high purity Ar as the plasma gas. Ages from the ratios ²⁰⁷Pb/²³⁵U, ²⁰⁶Pb/²³⁸U, and ²⁰⁷Pb/²⁰⁶Pb 335 336 were calculated for each data scan and checked for concordance. Concordance was defined as 337 overlap of all three ages at the 1σ level. If the number of concordant data scans for a spot was greater than zero, the more precise age from the concordant-scan-weighted ratio ²⁰⁷Pb/²³⁵U, 338 ²⁰⁶Pb/²³⁸U, or ²⁰⁷Pb/²⁰⁶Pb was chosen as the preferred age, and whichever exhibited the lower 339 340 relative error. If zero concordant data scans were observed, the common Pb-corrected age based 341 on isotopic sums of all acceptable scans was chosen as the preferred age. Results of zircon U-Pb 342 dating are shown in Table 4.

Clay mineralogy was examined for provenance purposes based on the concept that different environmental conditions and source terranes can produce characteristic assemblages. This allows us to separate material derived from the Indus River from material more closely linked to peninsular India. Although there may have been some change in mineralogy during initial diagenesis, the relatively shallow burial depths of these cores means that there is no significant thermal diagenesis and we can consider the observed mineralogy to be largely representative of that at the time of sedimentation.

350 Clay mineralogy was determined by using X-Ray Powder Diffraction (XRD) at Louisiana 351 State University using a Panalytical Empyrean X-Ray Diffractometer. Forty selected samples 352 within the MTC were soaked in water until there was no flocculation, with Na₃PO₄ added to de-353 flocculate when necessary. Samples were centrifuged for separation of the <2 µm material. Four 354 XRD patterns were generated from each oriented sample smear. The first pattern was collected 355 from the sample in air-dried conditions. The second XRD pattern was generated from a 356 glycolated sample after the slide was then placed in a desiccator with ethylene glycol for a 357 minimum of 8 h at 25°C. t. The third and fourth XRD datasets were collected after the sample 358 was subjected to heat treatments of 300°C for 1 h, and then 550°C for 1 h, respectively. XRD 359 analysis began immediately after glycolation, and immediately after the first heat treatment. In 360 this study we use the semi-quantitative method of Biscaye (1965) to estimate the clay 361 assemblage, which is based on peak-intensity factors determined from calculated XRD patterns 362 as measured by MACDIFF software. For clay minerals present in amounts >10 wt% uncertainty 363 is estimated as better than ± 5 wt% at the 95% confidence level. Uncertainty of peak area 364 measurement based on repeated measurements is typically <5%. Data are presented as relative 365 concentrations of the total clay assemblage in Table 5.

366

367 **DEFINING THE TOP AND BASE**

368 Microfossil assemblages within the sediments provide constraints on the age of emplacement. The oldest sediment overlying the MTC was det around 10.8 Ma based on 369 370 nannofossil assemblages and paleomagnetic stratigraphy (Pandey et al., 2016c). In Hole U1456D 371 the first appearance of Discoaster hamatus (10.55 Ma) marks the top of Zone NN8 (Pandey et 372 al., 2016a), while in Hole U1457C the interval 859.49–995.93 mbsf contains Catinaster coalitus, 373 which has a total age range of 9.69–10.89 Ma (Pandey et al., 2016b). The presence of *Discoaster* 374 *bellus* (first appearance at 10.40 Ma) within this interval also constrains the age to between 9.69 375 and 10.40 Ma. Much of the interval from 1009.21 to 1054.34 mbsf at Site U1457 contains a 376 mixture of different nannofossil species.

377 Above the MTC there is a coherent assemblage of nannofossils suggestive of hemipelagic 378 sedimentation and not the mixed assemblage of early Neogene and Paleogene forms found 379 within the MTC, as might be associated with a reworked deposit. We use this noticeable change 380 in nannofossil assemblage as a criteria for defining the top of the MTC. In this study we define 381 both a sedimentary and biostratigraphic top from the core, as well as the top inferred from the 382 strong reflector in the seismic image, typically associated with massive carbonate beds. The 383 sedimentary top of the deposit marks the transition from sediment that is clearly slumped or 384 tilted in the core and appears to have been affected by syn-sedimentary deformation (Figs. 3 and 385 4) while the biostratigraphic top represents the transition from reworked into pristine nannofossil 386 assemblages. The difference in depth, ~35 m, is significant and could represent continued 387 slumping and reworking of young sediments after the initial emplacement of the main MTC 388 bodies.

389 The base of the complex is easily established in both drilling sites, being marked by the 390 presence of carbonate breccias immediately overlying fine-grained sediments (Figs. 3 and 5). 391 The depth of this contact is 1101.65 and 1054.1 mbsf at Sites U1456 and U1457, respectively. A 392 key observation is that in the thicker Site U1456 section there is a 20-m-thick interval in which 393 normal hemipelagic sedimentation was briefly reestablished, based on the lack of reworking in 394 the nannofossil assemblages. This spans from around 956 to 935 mbsf (Figs. 3 and 6). This 395 shows that the MTC must have been emplaced in at least two phases separated by a pause, 396 despite the fact that this is not apparent in the seismic image. What is surprising is that the top of 397 this hemipelagic hiatus in mass wasting is not marked by a fresh influx of clearly reworked 398 brecciated carbonate material. Much of the hemipelagic interval comprises massive or parallel-399 laminated mudstones with a couple of medium-bedded to massive sandstones representing less 400 than 10% of the section (Fig. 6A). This is only moderately different from the material which lies 401 above the hemipelagic layer that is characterized by mudstones interbedded with thin beds of 402 siltstone. Above the hemipelagic layer, however, there is clear evidence for slump folding, tilted 403 bedding and microfaulting, which testifies to the redeposited character of these sequences, as 404 well as the mixed nannofossil assemblage. It is only in the somewhat shallower part of the 405 section at Site U1456 there is evidence for a fresh influx of very coarse redeposited carbonate 406 debris flow material, above 874.2 mbsf (Fig. 3A).

407 At both sites, the topmost part of the deposit largely comprises fine-grained, bioturbated 408 claystones and clay-rich siltstones that are otherwise hard to distinguish from the background 409 deposits of the Indus submarine fan, especially when they are not deformed. Tilted bedding is 410 suggestive of deformation but might be interpreted as being coring related. The presence of

411	slump folds close to the sedimentary top of each drilled section is, however, more conclusive in
412	demonstrating continued mass wasting above the coarser grained basal units.
413	
414	SEDIMENTARY FACIES
415	The sedimentary facies within the MTC were determined on the basis of core
416	descriptions and, in particular, the analysis of sedimentary structures that give clues to the
417	depositional processes that were operating during emplacement. We here describe the major
418	sediment types and provide interpretations of the depositional mechanisms. These are
419	summarized in Figure 3.
420	

421 Limestones

422 Short intervals of the MTC comprise coherent sections of fine-grained limestone that show little evidence for the action of high energy reworking depositional processes. Limestones 423 424 are found at Site U1456 within the lower part of the section around 1050 mbsf depth (Fig. 3). 425 The limestones are typically massive and generally fine-grained micrite with moderate amounts 426 of clay that give them an off-white color. Heavily bioturbated sediment with vertical Zoophycos 427 trace fossils are typical of sedimentation in moderately deep water, often close to the shelf edge 428 (Fig. 7A)(Ekdale et al., 1984; Seilacher, 1967). Figure 7B shows a massive micritic limestone 429 with some evidence of bioturbation, but which indicates minor recrystallization along stylolites, 430 highlighted by thin clay-rich partings. Neither deposit contains indication of strong current 431 activity, such as ripples or laminations, or even a well sorted granular texture, but rather 432 sedimentation in a low energy carbonate-rich environment probably below storm-wave base 433 (<40 m)(Peters and Loss, 2012). Short intervals of limestone are also found at Site U1457 very

close to the base of the MTC ~1050 mbsf. These are granular and porous and may be the product
of higher energy sedimentation in relatively shallow water depths (<30 m). Again, the limestones
are tan-colored rather than being pure white that is indicative of a modest clay content. Given the
modern significant water depth (3523 m at Site U1457) we propose that these limestones
represent coherent blocks of relatively shallow water material that were emplaced as part of the
brecciated units near the base of the MTC.

440

441 Carbonate Breccia Debrites

442 The vast majority of the carbonate sediment in the MTC are breccia clasts found mostly in the bottom part of the deposit at Site U1456 (970–1101 mbsf), with further yet more limited 443 444 clasts in the upper part of the MTC at the same site. They are also found immediately above the 445 base of the MTC at Site U1457 (Fig. 3). These breccias are thick-bedded, ranging close to 20 m 446 thick for individual beds separated by finer grained units. At Site U1456 there are multiple such 447 breccia units, stacked on top of each other, that are preferentially developed towards the base of 448 the sequence. The breccias are sometimes overlain by calcarenites (described below) or by 449 mudstones with a sharp boundary between the two lithologies. The breccias are extremely 450 poorly-sorted and the individual clasts are angular to sub-angular. Clast size ranges up to and 451 greater than the width of the core (>10 cm). There is usually no trend towards fining or 452 coarsening upwards within individual units, although one coarsening upwards sequence is seen 453 in Section U1456D-43R-1 (860 mbsf). The fabric of the sediment is rarely clast-supported (Fig. 454 8A) but is normally suspended in a dark muddy matrix (Fig. 8B). 455 The limestone clasts are pale tan to bright white with the interior showing a very fine-

456 grained or slightly granular sediment classified as micrite or more rarely packstone and

wackestone (Dunham, 1962). In the part of the section densest in limestone clasts (~1036 mbsf at
Site U1457), clasts are seen to indent one another both in core surfaces (Fig. 8A), as well as in
microscope thin sections (Fig. 9D). We interpret this as a result of dissolution during diagenesis
and burial.

461 The vast majority of the carbonate rocks redeposited in the debris flows appear to have 462 been lithified prior to their resedimentation. In combination with the observation of angular 463 clasts, we see coherent rafts of sediment (>10 cm width) floating within finer grained material 464 (Fig. 8B). There is some evidence that some of the carbonate sediment was not lithified during 465 emplacement because soft sediment folding of the deposits, such as seen in muddy limestones 466 (Fig. 10A) can be observed. However, these deformed deposits only represent a relatively small 467 part of the total sequence. It is clear that brittle deformation is important locally, especially 468 between and within the more coherent carbonate blocks. Slickensides especially testify to rapid 469 brittle deformation of the carbonate rocks during their emplacement (Fig. 8C). Most of the debris 470 flow units are extremely poorly-sorted but sometimes are represented by coarse sandstones 471 devoid of larger clasts (Fig. 8D). In these, larger granular clasts are supported in a muddy 472 sandstone matrix with no clear grading within the unit.

Although limestone fragments dominate the debris flows, it is noteworthy that in places
there is evidence for reworking of volcanic rocks into the flows (Fig. 10B). These clasts are
weathered red-brown and are sub-rounded. The largest single clast was found at 879 mbsf at Site
U1456 within a poorly indurated conglomeratic part of the debris flow sequence. The clast is an
8-cm-wide fragment of vesicular aphyric basalt that is presumed to be derived by erosion from
the Deccan Plateau volcanic sequences exposed across peninsular India. The clasts were likely

479 eroded on to and then reworked across the continental shelf because being redeposited in the480 MTC.

481 The limestone, from which the carbonate clasts were derived, formed as a typical 482 shallow-water deposit in a biologically productive zone mostly starved of clastic sediment input. 483 Original water depths were within the photic zone on the continental shelf or within a back-reef 484 setting (<50 m), with only moderate amounts of current activity, since we see no evidence for 485 strong sorting or high energy deposits such as oolites or grainstones (Dunham, 1962). These 486 original rocks have mostly been broken and reworked as debris flow deposits during the 487 emplacement of the MTC. The muddy matrix has a separate provenance, either from the deep-488 water slope of peninsular India or from the Indus Fan itself, as discussed below.

489

490 Calcarenites

491 Calcarenite is present in each carbonate section, in the form of massive, well-sorted units 492 suggestive of high energy current transport. Beds of calcarenite are several meters thick and 493 generally massive and structureless, although they can develop a sub-horizontal fabric suggestive 494 of current flow. Where the deposits are finer (Fig. 10D), there is a shear-type fabric developed 495 within the calcareous siltstones. In the coarser grained units (Fig. 10C) there is some evidence 496 for internal soft sediment deformation, although generally the units are homogenous and 497 comprise uniform, gray, coarse-grained sandstone. They are well-sorted and clast-supported, 498 with very little muddy matrix, suggestive of a high energy depositional regime. The majority of 499 the clasts are carbonate, although there are a significant number of dark grains of organic carbon 500 origin. These calcarenites often have sharp tops that are interpreted to reflect erosion of the 501 deposit prior to the emplacement of overlying units. Figure 10D shows a calcareous siltstone

sharply overlain by conglomeratic sandstones deposited as debris flows. Very few sedimentary structures are seen within these deposits, so that we infer sedimentation in an upper flow regime resulting in relatively laminar deposits without any current ripples or finer interbeds. Sediment concentrations are inferred to have been very high during deposition, which terminated rapidly.

200

507 Turbidites and Hemipelagic Mudstones

508 Apart from the carbonate-dominated debris flows, minor turbidite sandstones and 509 dominant siltstones and mudstones make up the largest part of the MTC. These are also 510 interbedded with associated hemipelagic mudstones. In the coarsest sandstones, each turbidite 511 shows a classic fining upward sequence (Fig. 11A), with largest carbonate fragments suspended 512 in a dark clastic mud matrix. Locally, there are sub-horizontal lamination although sedimentary 513 structures are poorly developed, with up-section fining dominating characteristic of these 514 deposits. In the upper parts of the MTC at both sites, muds show lamination and interbedding of 515 modest amounts of muddy silt (Fig. 11B). Elsewhere, the deposits are massive, dark gray 516 mudstones with few sedimentary structures. These contrast with the draping mudstones that 517 overlie the catastrophically emplaced MTC where typical deep-water trace fossil assemblages 518 (i.e., Zoophycos; (Fig. 11C) characterize the hemipelagic sedimentation and eliminate the 519 possibility of large-scale mass wasting. This is in contrast to the muddy upper sections of the 520 MTC itself, where there is evidence for laminar current flow that follows the initial emplacement 521 of the carbonate debris flow deposits at the base of each cycle. In general, the grain sizes are 522 relatively limited, with only few a thin-bedded sandstones and occasional siltstones developed 523 within what is otherwise a dominantly (95%) muddy sequence. Distinguishing muddy sediment

with the MTC from the hemipelagic interval within Site U1456 is difficult without the help ofmicropaleontology evidence.

Syn-sedimentary deformation within the muddy turbidities include folds, micro-faults, 526 527 and tilted bedding (Fig. 11D) and are particularly easy to see in well-laminated sequences. Dip of 528 lamina can be high $(>50^\circ)$, indicating significant deformation of the muddy units after 529 sedimentation. In addition to ductile structures, there is evidence for compressional reverse 530 faulting. Significant dips and deformation are evidence for incorporation as part of the MTC 531 rather than the subsequent hemipelagic sedimentation of the Indus Fan, which is only gently 532 inclined like the seafloor or the top of the MTC ($\sim 1.2^{\circ}$ according to Calvès et al. (2015)). 533 534 **Micro-Facies** 535 Petrographic analysis can be used to help interpret paleoenvironment and depositional 536 mechanisms from facies identified in the cores. Figure 9A shows a silty laminated mudstone 537 from the upper part of the MTC at Site U1457 that is interpreted here as a turbidite deposit. The 538 massive calcarenite beds that overlie debris flow conglomerates are seen to be relatively poorly 539 sorted and matrix supported, at least in places, in thin section (Fig. 9B). Clasts are rarely 540 composed of calcite crystals but are dominated by a variety of finer limestone facies, especially 541 micrite. Aggregates of dolomite crystals are observed (Fig. 9C) and interpreted to represent 542 diagenetic alteration of original calcite via interaction with magnesium-rich waters prior to 543 resedimentation. Their presence is suggestive of redeposition from shallow water areas where 544 this mineral generally forms. 545 There are large numbers of microfossils and their fragments within the breccia limestone

546 clasts. Foraminifers are abundant (Figs. 12A, 12B, 12F). In addition, we also confirm the

547	presence of crinoid fragments (Fig. 12D), bryozoans, and rare radiolarians (Fig. 12E). The
548	skeletal assemblage of most limestone clasts is dominated by calcareous red algae and benthic
549	foraminifera (including both miliolids and large rotaliids; Fig. 12C). Rare echinoderms, mollusks
550	and hermatypic coral fragments are also present. Some skeletal grains, originating from a
551	shallow-water environment (coralline algae, large echinoid spines, large benthic foraminifera),
552	also occur within the matrix (Figs. 12H, 12I). The occurrence of what is likely to be Lockhartia,
553	together with the peyssoneliacean red-alga Polystrata alba, suggests that at least part of the
554	eroded limestone was of Paleogene age (Fig. 12C)(Bassi and Nebelsick, 2000; BouDagher-
555	Fadel, 2018). The matrix is largely dominated by planktonic foraminifera with minor
556	contribution from small rotaliids (Figs. 12G).
557	These characteristics suggest that the MTC involved both lithified inner platform deposits

(the source of limestone fragments) and outer platform deposits still composed of loose grains
(the source of the muddy matrix with planktonic foraminifera).

560

561 **DEPOSITIONAL MECHANISMS**

562 Most sediment within the MTC are either debris flow deposits, well-sorted calcarenites, 563 or dominantly clastic turbiditic siltstones and mudstones. Both phases of the MTC at Site U1456 564 (Fig. 3) show large-scale fining upwards cycles, with a dominance of carbonate debris flows 565 towards the base grading into more siliciclastic turbidite sedimentation towards the top. Smaller, 566 shorter phases of fining upwards cycles are further observed within the two overall fining 567 upwards cycles at Site U1456. For example, the upper part of Phase 1 (Fig. 3), comprises a basal 568 unit from between 999.2 and 984.0 mbsf that is dominated by rafted carbonate sheets and 569 carbonate debris flow material (Figs. 3 and 6B). This interval is likely a second pulse after the

570 initial Phase 1 event. Above 984.0 mbsf there is a transition to massive thick-bedded calcarenite

571 with slump folds, although this is truncated sharply at 973 mbsf by mudstones that rapidly

572 transition into the hemipelagic sediment described above (Fig. 6B). This implies that the basal

573 Phase 1 unit, especially at Site U1456 comprises a series of pulses rather than one single gigantic

by the seismic data alone (c.f.(Calvès et al., 2015)(Fig. 2).

575 The base of Phase 1 at both sites is characterized by a thick-bedded sequence of debris 576 flow calcareous breccias and rafts of undeformed shallow water carbonate (Fig. 5). These are not 577 surprisingly the thickest such deposits within the entire drilled section. Although Site U1456 is in 578 a more central location within the basin, the oldest debris flow breccia at the base of Phase 1 is 579 thinner in this location than at Site U1457 and transitions more rapidly up into thick-bedded 580 breccia and interbedded calcarenites. Both sections, however, do show an overall fining upward 581 between the base and overlying mudstone units. The initial debris flow sedimentation appears to 582 be ~94 m thick at Site U1456 (1101.6-1007.2 mbsf) and ~48 m thick at Site U1457 (1006.4- \mathcal{O} 583 1054.3 mbsf; Figs. 3 and 5).

584 In general, calcarenites alternate with debris flow conglomerates (Fig. 5A) indicating 585 alternating depositional mechanisms within a single emplacement episode. Individual debris 586 flow events are followed by high energy upper flow regime periods of sedimentation where 587 massive well-sorted calcarenites were deposited before being followed by another debris flow 588 unit. However, presumably all this material was emplaced over a relatively short period of time. 589 The carbonate-dominated debris flows form the initial erosive base of the MTC, followed by 590 mud-dominated turbidite sedimentation and hemipelagic fallout representing the tail of the MTC. 591 At Site U1456 this sequence is then repeated after the hemipelagic break. Soft sediment 592 deformation is commonly seen in the more laminated sections indicative of slumping after

sedimentation. It seems unlikely that poorly consolidated mudstones and siltstones could have
been emplaced hundreds of kilometers in a semi-coherent form, unlike the well-lithified
limestone clasts seen close to the base of each section.

596

597 **GEOCHEMISTRY**

598 Bulk Geochemistry

599 We use a CN-A-K ternary diagram to illustrate major element geochemistry of MTC 600 samples compared to sediments from the Indus Canyon and delta. The sediment from the MTC 601 largely plots within the range of the Indus Canyon and trends towards higher values of Al₂O₃ 602 (Fig. 13A). MTC samples appear to have higher values that trend towards the illite end-members 603 and may be more depleted in biotite and feldspars compared to the delta. This is likely a result of 604 sediment transport, similar to what has been observed in the Indus Canyon (Li et al., 2018). 605 Sediments in the muddy upper part of the MTC at Site U1457 largely plot with low 606 Chemical Index of Alteration (CIA), which is a proxy of the state of weathering of a sediment 607 compared to pristine bedrock (Nesbitt et al., 1980). The muddy upper MTC samples trend more 608 towards the Quaternary Indus Delta field compared to the lower parts of both Phase 1 and Phase 609 2, which show more overlap with western Indian Shelf sediments, largely derived from rivers 610 draining the Deccan Plateau (Kurian et al., 2013). This plot implies that the upper muddy 611 sediments at Site U1457 had a dominant source from the Indus River/Fan and little inputs from 612 western peninsular India.

613 The sediment in the MTC can also be characterized using other major element 614 discrimination diagrams. Figure 13B shows the scheme of Herron (1988) in which the Phase 1 615 and Phase 2 samples largely plot within the Fe shale field, with a few slightly depleted in Fe and

616 plotting as shales. Again, we plot these samples along with the western Indian Shelf, Indus 617 Canyon and delta sediments. Samples from the upper muddy top to Phase 1 at Site U11457 form 618 a cluster within the range of the Indus Canyon sediments, suggesting a dominant provenance of 619 reworked Indus material. Comparison with sediment from the western Indian shelf shows a 620 significant difference, with the shelf sediment typically plotting with much higher Fe contents, 621 similar to the lower Phase 1 and 2 sediments. We infer that the bulk of the sediment in the lower 622 MTC comprises mostly Indian margin sediment with muddy top dominated by sediment eroded 623 and redeposited from the Indus Fan.

624

625 Nd and Sr Isotopes

626 We use Sr and Nd isotope values to constrain the provenance of siliciclastic sediment in 627 the MTC. By cross-plotting Nd and Sr isotopic compositions from source regions such as the 628 Deccan Traps, peninsular Indian rivers, Transhimalaya, Karakoram, Greater Himalaya, Kirthar 629 and Sulaiman Ranges, and modern/Quaternary Indus-derived sediment allows the origin of the 630 sediment to be further constrained (Fig. 14). This diagram shows that the MTC samples form a 631 relatively discrete cluster with one exception that has especially positive ε_{Nd} values that fall 632 within the Deccan and Transhimalayan arrays. When we compare these data with potential 633 sources, it is clear that the bulk of the sediments lie within the isotopic range defined by the 634 Indus submarine fan sediments at the same drilling sites (Clift et al., 2018). This is consistent 635 with the argument that much of this material may be reworked Indus-derived sediment. 636 However, we note that it is impossible to exclude mixing of sediment from the peninsular Tapti 637 or Narmada Rivers. The isotope compositions by themselves do not allow us to quantify the 638 degree of reworking from these sources as they are similar to the Indus. Although the MTC

639 samples plot with higher ε_{Nd} values compared to the Quaternary Indus Canyon, as well as the 640 Kirthar and Sulaiman ranges, such a composition could largely be explained through temporal 641 variation in the Indus River itself (Clift and Blusztajn, 2005; Clift et al., 2018). The one very 642 positive ε_{Nd} sample is anomalous and plots with even more positive values than the Tapti River. 643 This is strongly suggestive of erosion from peninsular India and is corroborated by the presence 644 of vesicular Deccan Plateau basalt fragments as previously noted.

We can look at the stratigraphic variation in isotopic compositions through time at both sites (Fig. 15). In both cases, Nd isotope compositions plot within error of the Quaternary Indus or with slightly more positive ε_{Nd} values. We note that the most positive ε_{Nd} values in each borehole are found within the debris flow conglomerate units bearing basaltic clasts at the base of the lower part of the MTC. This is especially true at Site U1456 (Fig. 15A). Variations in ⁸⁷Sr/⁸⁶Sr also mirror this general evolution.

651 The provenance of the coarse-grained carbonate debris flow deposits is different from 652 those of the finer grained sediments overlying them. The fine-grained sediments may represent 653 recycling of pre-existing fan sediments into the top of the MTC, while the debris flow deposits 654 are more closely associated with mass wasting from the western Indian continental margin. It is 655 possible that some Indus River sediment could have been transported east along the shelf, carried 656 by longshore currents from the river mouth, and deposited offshore Saurashtra before being 657 redeposited as part of the MTC. However, there is no evidence that significant Indus sediment 658 travels farther east than the Rann of Kutch (Khonde et al., 2017; Kurian et al., 2013). The 659 simplest interpretation is that the upper muddy layers of the MTC represent entrained and 660 reworked Indus Fan material.

661

662

Heavy Mineral Analysis

663 The heavy-mineral assemblages help to constrain the source area of the MTC. The 664 concentration of heavy minerals in all samples is very low suggesting a strong depletion due to 665 intrastratal dissolution of unstable silicates (Garzanti, 2017). Consequently, a relative enrichment 666 of ultrastable minerals is observed (ZTR index of Hubert (1962)). The two samples (U1456E-667 15R-1W, 61-63 cm and U146E-17R-4W, 131-133 cm), analyzed from the carbonate breccia 668 present extremely low HMC (0.04-0.05%) with common augitic clinopyroxene ($\sim 6\%$) and rare 669 spinel (2–3%). The minerals also show corroded surficial textures, indicating a strong diagenetic 670 overprinting (Ando et al., 2012). A similar fingerprint is detected in Sample U1456E-7R-1, 80-671 82 cm where green and brown augite are abundant (48%). In all these samples, there are 672 common garnets associated either with apatite, titanite, epidote, zircon, tourmaline, and 673 metamorphic Ca-amphiboles, potentially derived from recycled sediments from the Himalaya-674 derived Indus Fan turbidites eroded by the MTC. Notwithstanding diagenetic dissolution, the 675 highly unstable augitic clinopyroxene (volcanic origin) always dominates over metamorphic 676 amphiboles, suggesting a sizable contribution to the MTC from the Indian passive margin, and 677 especially from Deccan Plateau basaltic lavas. Sample U1456E-4R-1W, 110-111 cm is a 678 calcarenite within which hydraulic sorting and high-energy currents preferentially selected the 679 available heavy minerals suite derived from the MTC, concentrating platy heavy minerals such 680 as chloritoid, Ca-amphiboles and tourmaline (lighter). The sample is partially depleted in denser 681 garnet. This assemblage is completed with the presence of abundant apatite, common titanite, epidote and spinel with trace of kyanite, and alusite and staurolite. 682 683 Sample U1457C-88R-4W, 58-60 cm was deposited far from the Indian Passive margin

684 and the mineralogy reflects a dominant contribution from recycled minerals derived from the

\bigcirc

685	erosion and re-deposition of the Indus Fan turbidites. The tHMC is very low (0.08%), and
686	mineralogy is dominated by abundant epidote and garnet with common apatite and titanite. Ca-
687	amphiboles dominate over clinopyroxenes, with a ratio 8:1, pointing to a major contribution
688	from the Indus River and the Himalaya in this sample. The assemblage also includes tourmaline,
689	zircon, chloritoid, Cr-spinel and trace of and kyanite, staurolite and andalusite.
690	The modern Tapti River was analyzed close to its mouth. The sample contains a very rich
691	assemblage of heavy minerals (tHMC 17%) with dominant augitic clinopyroxenes (92%) and
692	subordinate amount of metamorphic heavy-mineral, Ca-amphiboles, epidote, garnet and
693	sillimanite. This mineralogical signature differs from the observed suite of orogenic heavy
694	minerals observed in the modern Indus River and his delta (Garzanti et al., 2005).
695	The heavy mineral assemblage in the MTC and the very low concentration of heavy
696	minerals points to different sources for the siliciclastic sediments, i.e., partially derived axially
697	from the Himalayas via the Indus River (especially at Site U1457C) and partially derived
698	transversally from the Indian peninsula (especially at Site U1456).
699	
700	Zircon U-Pb Ages
701	To further constrain provenance, we compare detrital zircon U-Pb ages with existing data
702	from the Indus river mouth (Clift et al., 2004), Indus Fan turbidites above and below the MTC
703	(Clift et al., 2018), and with bedrock data from potential sources in the river catchment (Fig.
704	16)(DeCelles et al., 2000; Gehrels et al., 2011). Although the zircon ages from source bedrock
705	overlap with each other, each source regions demonstrates strong preferential age spectra that
706	can be used to discriminate between them. Zircons from Nanga Parbat, Kohistan, the
707	Transhimalaya, and the Karakoram generally have younger ages (<300 Ma) than those from the

Himalayan ranges (Alizai et al., 2011)(Fig. 16). Both the Greater and Tethyan Himalaya have UPb age peaks at 300–750 Ma and 750–1250 Ma, with older ages at ~1850 Ma characterizing the
Lesser Himalaya

711 The volume of sample available for U-Pb dating from Core U1457C-7R (the only suitable 712 sediment seen in the MTC) was extreme mited such that only 51 grains yielded concordant 713 ages, which is somewhat lower than the 113 minima suggested by Vermeesch (2004) for a 714 sample with complex provenance. Nonetheless, some inferences concerning provenance can be 715 made. What is clear is that young ages dominate with 17 grains dated at less than 100 Ma (Fig. 716 16). The age spectrum bears most similarity with Indus Fan turbidites dated at 7.8, 8.3, and 15.6 717 Ma, but all are in contrast with the ages from the modern river. The match between these young 718 grains and sources in the Karakoram and Kohistan argue for the sand to be an Indus-derived 719 sediment and not from sediment transported from the Indian peninsula where zircon ages are 720 Paleozoic or typically much older. This conclusion is consistent with the Nd and Sr isotope data 721 from the upper parts of the MTC. The analyzed sandstone was sampled below the 722 sediment/structurally defined top of the MTC but above the carbonate-dominated debris flow 723 facies at the base of the complex, i.e., within the muddy but slumped top of the MTC. This 724 implies that the upper parts of the MTC are Indus Fan sediments entrained in the tail of the MTC 725 during emplacement.

726

727 Clay Mineralogy

The clay mineral assemblages within the MTC can be used to assess provenance by semiquantitative analysis and comparison with existing data from other sources. When plotted on the ternary diagram of (illite+chlorite), kaolinite, and smectite (Fig. 17) there is significant

731 overlap between the new MTC data and other Arabian Sea sediments (Rao and Rao, 1995). In 732 general, the MTC clays are low in kaolinite and form an array between the smectite and (illite+chlorite) end members. In this respect, they show a similar character to sediments from the 733 734 Indus fan and have significant overlap with Quaternary clays from the Indus Canyon (Li, 2018). 735 Samples from Phase 1 of the MTC have very high smectite contents, similar to the Paleocene 736 sediments overlying basement at Site U1457, suggestive of a volcanic source. They are close to 737 sediments recovered from the inner shelf offshore Saurashtra and from the Gulf of Cambay. 738 Phase 2 sediments and the hemipelagic layer are slightly less smectite rich but overlap with the 739 Holocene Indus Shelf, as well as some modern Indian Shelf sediments. We note that the bulk of 740 the muddy upper Phase 1 sediments plot with higher (illite+chlorite) values and they also tend to 741 have slightly higher kaolinite compared with analyses of sediments from the Indus floodplains 742 (Alizai et al., 2012). These sediments are similar to the assemblage recognized from the outer 743 Saurashtra margin (Rao and Rao, 1995) and are similar to many clay assemblages within Indus 744 Fan turbidite sequences. Overall, the MTC deposits have lower kaolinite compared with most 745 Western Indian shelf deposits but some samples plot closely to the shelf. It is also noteworthy 746 that the MTC assemblages generally show lower (illite+chlorite) compared with many of the 747 Miocene-Recent Indus submarine fan deposits, which likely indicates a mixed provenance of 748 Indus and Indian peninsular sediment. However, because illite and chlorite are the product of 749 physical weathering rather than chemical weathering their relatively high contribution to the 750 MTC could also indicate reduced chemical weathering of fan sources since MTC emplacement. 751 These data are consistent with a dominant recycling of Indus Fan deposits in the upper muddy 752 parts of the MTC, but with greater involvement of clays derived from the Western Indian margin

in the lower part, especially in Phase 1. The similarity with modern nearshore sediments offshoreSaurashtra and Cambay is consistent with an origin in this part of the margin.

755 Clay mineralogy shows significant variation with depth (Fig. 15). At Site U1456 the 756 carbonate-rich part of the section shows particularly high smectite contents and relatively low 757 (illite+chlorite) values. Smectite only becomes less abundant than these two physically 758 weathered clays above the upper Phase 2 carbonate debris flow unit. At Site U1457 the 759 carbonate-rich part of the section similarly is smectite-rich, but immediately above this level the 760 sediments become dominated by an (illite+chlorite) assemblage similar to the Indus Fan. It is 761 noteworthy that the Paleocene sediments beneath the MTC at Site U1457 are ~100% smectite, 762 possibly reflecting chemical weathering of the underlying basaltic basement. Clay mineralogy 763 supports the Nd and Sr isotope compositions in showing a characteristic difference between the 764 carbonate-dominated sections that indicate similarity to the western Indian margin, whereas the 765 mudstone dominated sequences further upsection in the MTC are most similar to compositions 766 associated with the Indus Fan.

767

768 SEDIMENT BUDGET

To assess the potential of sediment delivery rates and margin oversteepening as triggering mechanisms of the MTC, a sediment budget from the western Indian margin was generated using standard two-dimensional backstripping methods from seismic profile data (Clift, 2006; Kusznir et al., 1995). This was to primarily test the hypothesis that the rapid accumulation of sediment on the continental margin resulted in an unstable stratigraphy that was then more liable to mass wasting events like the Nataraja MTC. There is strong evidence that the Western Indian continental margin is gravitationally unstable as a result of the large-scale compressional thrusts 776 seen in seismic profiles towards the base of the continental slope seen between the Saurashtra 777 shelf and Bombay High (Fig. 1)(Calvès et al., 2015; Nair and Pandey, 2018). These features are 778 often associated with slopes prone to gravitational collapse, which in this region, has yet to 779 manifest in the dramatic fashion of the Nataraja MTC. In order to estimate the mass flux of the 780 margin, we use the cross-margin seismic reflection profile of Nair and Pandey (2018)(Figs. 1 and 781 18). Their northernmost profile lies immediately south of the scarp region identified by Calvès et 782 al. (2015) and which we consider to be potentially representative of the sedimentation in the 783 source regions of the MTC prior to its redeposition. For the purpose of this study, we use the age 784 control provided by Nair and Pandey (2018), at least for the continental shelf and slope areas 785 (Fig. 18A). West of the toe of the slope sedimentation is linked to the Indus Fan and may not be 786 representative of the mass flux to the Saurashtra Shelf. Figure 18A shows the interpretation of 787 Nair and Pandey (2018) with a conversion from their seismic travel time scale to depth made on 788 the basis of multichannel seismic stacking velocities derived from the Indus shelf, as used by 789 Clift et al. (2002)(Table 6). We do this because of the absence of such data from the Saurashtra 790 region itself. We prefer to use velocity data from the Indus continental shelf rather than from the 791 deep basin because as the sediment thicknesses are much greater under the continental shelf, they 792 are more comparable to those seen offshore the Indus River mouth. Based on the lateral 793 variability in velocities seen on the Indus Shelf, we estimate that this conversion may introduce 794 uncertainties as high as $\pm 20\%$ (Clift, 2006). Stratigraphic ages are then assigned numerical ages 795 based on the timescale of Gradstein et al. (2012).

The depth-converted line was then backstripped using standard decompaction methods (Kusznir et al., 1995; Sclater and Christie, 1980). This was done to restore each dated sediment layer to its original thickness prior to burial. Knowledge of the sediment type is important to this
799 calculation because shales experience much greater loss of porosity during burial than do 800 sandstones (Sclater and Christie, 1980), and in this case, we used lithological data from Wandrey 801 (2004) and Rao and Talukdar (1980). These studies show a mixed Cenozoic sequence dominated 802 by silty muds and carbonates offshore Saurashtra. The decompaction process involves 803 accounting for the loss of porosity of the sediment during burial, which would otherwise result in 804 an underestimation of deposited volumes for the older, deeper buried sediment packages. After 805 the original, uncompacted volume of sediment in each dated interval has been determined, the 806 mass of rock delivered during that time period can be calculated. Errors in lithology and 807 compaction history are much smaller than the time-depth conversion and rarely exceed 5%. 808 In this study two-dimensional decompaction was calculated using the program Flex-809 DecompTM (Kusznir et al., 1995). It must be assumed that the analyzed profile is representative 810 of the total mass flux to the margin since rifting of the Arabian Sea ~66 Ma (Bhattacharya et al., 811 1994). Because we only have one profile close to the area of mass wasting, and no estimate of 812 the total sediment mass offshore Saurashtra, it is not possible to make a volume calculation. 813 However, the two-dimensional budget does at least allow us to estimate the volumes of sediment 814 delivered per kilometer of margin close to the source of the MTC. Our results show a clear trend 815 to increasing mass flux after 26 Ma (Fig. 18B), with a peak between 16 and 11 Ma. Because the 816 resolution of the budget is constrained by the presence of the dated horizons, it is not possible to 817 accurately say when the peak sediment flux was achieved, but this analysis confirms that the 818 Middle Miocene was a time of rapid sedimentation offshore Saurashtra, a pattern that it shares 819 with many other Asian delta systems. As a result, it seems likely that the pulse was caused by 820 faster erosion driven by heavy summer monsoon rains (Clift, 2006). We suggest that much of the 821 gravitational instability on the western Indian margin was caused by rapid sedimentation in the

Middle Miocene causing oversteepening of the shelf edge, comprising large thicknesses of
sediment liable to incomplete dewatering during burial. The reducing sedimentation rates after
11 Ma may explain why a second such slide has not been emplaced in this part of the margin.

826 SEISMICITY

827 As well as an over-steepened continental margin caused by increased sediment flux, we 828 investigate the possible triggering of the MTC as a result of seismic activities that are often 829 implicated in the emplacement of large mass wasting complexes (Kastens, 1984). Figure 1 shows 830 the location of earthquakes greater than 4.5 magnitude since 1960 in the vicinity of the source 831 region for the MTC. There is some seismicity related to the plate boundary west of the Indus 832 delta and there are small amounts of activity in the Saurashtra Peninsula itself, immediately 833 opposite the scar in the continental shelf. It is apparent that the greatest concentration of seismic 834 activity is however around the Rann of Kutch, where historic intraplate events up to 7.7 835 magnitude have been recorded (Bilham, 1999). This activity reflects reactivation of earlier rift-836 related faults due to compression linked to the India-Eurasia collision (Bilham et al., 2003; 837 Biswas, 2005). This part of the Indian plate is a weak zone and may well have been active as a 838 seismic hotspot for significant periods of time. We suggest that it is the relative proximity of the 839 Saurashtra margin to this tectonic feature (<300 km) which may have initiated the mass wasting 840 in that region, rather than further south along the margin where sediment flux was also high. 841

842 SYNTHESIS AND CONCLUSIONS

843 This study, made possible through drilling, reveals for the first time the internal structure 844 and origin of the Nataraja MTC, and extends our understanding based on the earlier seismic

845 surveying of the deposit. At Site U1456, there is clear evidence that the MTC was emplaced in 846 two major phases separated by a significant break (Fig. 19). Even the larger, earlier Phase 1 can 847 be broken down into at least two stages, indicative of pulsed emplacement. The basal part of 848 each drilled section of the complex comprises debris flow carbonate breccias and larger rafts of 849 shallow water limestone, which can be traced back to collapse of the carbonate edge of the 850 continental shelf offshore Saurashtra. The MTC is emplaced as a number of fining upward 851 sequences with debris flow breccias, overlain by well sorted, coarse calcarenite deposited by 852 high velocity currents following in the wake of the initial mass wasting landslide. These are 853 overlain by muddy and turbiditic deposits, which are increasingly siliciclastic in character. At 854 Site U1457, only a thinner section of the earlier Phase 1 appears to be preserved, but a second 855 Phase 2 is apparent at Site U1456. Again, there was an emplacement of carbonate-rich debris 856 flows, although these were preceded and followed by muddy turbidite deposits, largely reworked 857 from pre-existing sediments of the Indus Fan. The top of each drilled sequence shows a 858 separation between sediment where the biostratigraphy is mixed and where slumping continues 859 to occur in the aftermath of the original depositional event.

860 Nd and Sr isotopic data, together with heavy-mineral assemblages, show that the 861 siliciclastic fraction of the deposit is associated with the western Indian continental margin, at 862 least in the debris flow part of the deposits although the overlying muddy turbidite units share the 863 same characteristics as the Indus submarine fan and suggest entrainment of sediment already 864 deposited in Laxmi Basin in the wake of the carbonate-rich debris flows that formed the MTC in 865 the first place. Limited zircon data at Site U1457 also show the clear signature of the Indus 866 River, although this applies only to the muddy units overlying the carbonate debris flows. We 867 envisage that enhanced sediment delivery to the western Indian continental margin driven by

868 strong monsoon during the middle Miocene resulted in an oversteepened continental margin that 869 was in a gravitationally unstable state. Exactly what triggered the collapse is not clear, but may 870 well be related to seismic activity in the nearby Rann of Kutch where large earthquakes continue 871 to the present day. Compressional deformation structures in the western Indian continental 872 margin south of Saurashtra suggest that this region too is in a compressional and potentially 873 unstable situation. However, decreasing sediment flux to the continental margin since the middle 874 Miocene has lessened the instability of the continental slope and reduced the chance of mass 875 wasting, especially further south away from potential seismic triggers. The western Indian 876 margin, however, has also experienced the increasing sedimentation rates linked to the onset of 877 northern hemisphere glaciation and so the potential for significant geohazard still exists. 878 Nonetheless, the fact that there has been no similar large event since 10.8 Ma does argue for this 879 being relatively low risk at the present time.

880

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885

886

887 Figure Captions

888 Figure 1. A) Shaded topographic and bathymetric map of the Arabian Sea showing the location 889 of the core sites discussed in this study (yellow dots). Base map from GeoMapApp. Dashed 890 yellow lines show proposed continent-ocean boundaries. Dashed white lines show oceanic 891 transform faults. Numbered red circles indicate existing scientific boreholes from Deep Sea Drilling Project (DSDP) and Ocean Drilling Program (ODP). Pink squares show pajor cities. 892 893 Magnetic anomalies (thin gray numbered lines) are from Miles et al. (1993). Green-filled circles 894 show earthquakes >4.5 magnitude since 1960 recorded by US Geological Survey. B) Close-up 895 map of Laxmi Basin showing the precise location of the drill sites. A pink dashed line shows the 896 extent of the Nataraja MTC (Calvès et al., 2015). Light blue lines show locations of seismic 897 profiles shown in Figure 2. 898 899 Figure 2. Seismic profiles of the core sites (left) with interpretation (right) showing the mass-900 transport complex in the immediate vicinity of (A) IODP Site U1456 and (B) IODP Site U1457. 901 Modified from Pandey et al. (2016c). See Figure 1 for locations of lines. 902 903 Figure 3. Summary stratigraphic columns showing the lithologies and interpreted facies of the 904 mass-transport complex at (A) IODP Site U1456 and (B) IODP Site U1457. Black shading in 905 second column indicates recovery, with white showing lost section. mbsf = meters below 906 seafloor. 907 908 Figure 4. (A) Sedimentary log showing the top of the deposit, U1456D-33R to U1456D-42R; (B)

909 Sedimentary log showing the top of the deposit, U1457C-69R to U1457C-78R. Black shading in

910	second column indicates recovery, with white showing lost section. mbsf = meters below
911	seafloor.

912

913	Figure 5. (A)	Sedimentary	log showing	the bottom	of the MTC,	U1456E-16R to	0 U1456E-19R;
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914 (B) Sedimentary log showing bottom of the MTC, U1457C-86R to U1457C-92R. Lithological

915 patterns and sedimentary structures same as Figure 4. Black shading in second column indicates

916 recovery, with white showing lost section. mbsf = meters below seafloor.

917

918 Figure 6. (A) Sedimentary log showing the deposit above and within the hemipelagic layer,

919 U1456D-50R to U1456D- 53R. As shown, soft sediment deformation occurs until pelagic layer

920 begins; (B) Sedimentary log showing the second pulse of carbonate debris flow material,

921 U1456D-56R to U1456D-61R. Lithological patterns and sedimentary structures same as Figure

922 4. Black shading in second column indicates recovery, with white showing lost section. mbsf =
923 meters below seafloor.

924

925 Figure 7. (A) Limestone with burrows (20 cm long), U1456E-12R-1, 42-47 cm (1045 mbsf); (B)

Stylolite in limestone, U1456E-10R-3, 30-40 cm (1030 mbsf). Vertical scale is in cm below the
section top. See locations on Figure 3.

928

929 Figure 8. (A) Coarse carbonate breccia with mudstone matrix, U1457C-90R-2, 75-83 cm (1036

930 mbsf); (B) Debris flow conglomerate with faulted mudstone raft (larger faults shown with white

931 lines), U1456E-9R-4, 78-88 cm (1021 mbsf); (C) Core photograph of slickensides on a fault

932 within silty claystone, U1456E-9R-4, 37-51 cm (1021 mbsf), (D) Coarse sandy, calcarenite,

933 U1457C-88R-5, 38-48 cm (1022 mbsf). Vertical scale is in cm below the section top. See
934 location on Figure 3.

935

- 936 Figure 9. Thin section plane polarized photomicrographs of (A) Laminated sandy siltstone with
- 937 quartz grains, U1457C-85R-1, 22-26 cm (997 mbsf). Note the finer muddy center of the image
- and the poorly sorted silt interbeds on either side with dominant quartz clasts; (B) Calcarenite,
- 939 U1456D- 60R-1, 13-17 cm (1006 mbsf); (C) Euhedral calcite/dolomite crystals within larger

940 grain, U1456E-15R-1, 12-16 cm (1073 mbsf); (D) Suture grain contact of carbonate clasts in

- 941 breccia, U1456D-45R-4-52-57 cm (870 mbsf). See location on Figure 3.
- 942

943 Figure 10. (A) Slump folded calcareous siltstone, U1456D-58R-2, 43-53 cm (989 mbsf); (B)

944 Deccan vesicular basalt clast, U1456D-46R-1, 16-25 cm (879 mbsf); (C) Massive calcarenite

945 with ductile folded layer U1456D-41R-3A, 114-124 cm (841 mbsf); (D) Sharp contact between

calcarenite and calcareous siltstone, U1457C-88R-7, 61-70 cm (1025 mbsf). Vertical scale is incm below the section top. See location on Figure 3.

948

- 949 Figure 11. (A) Sandy siltstone showing gradual normal grading, U1457C-71R-3, 101-115 cm
- 950 (865 mbsf) (B) Tilted, laminated turbidite deposit U1457C- 73R-2, 140-148 cm (881 mbsf); (C)
- 951 Mudstone with Zoophycos burrows (one outlined for clarity), U1457C-68R-1, 42-52 cm (832
- mbsf); (D) Steeply dipping laminated mudstone showing reverse faulting, U1456D-46R-3A,
- 953 139-148 cm (883 mbsf). Vertical scale is in cm below the section top. See location on Figure 3

Figure 12. Thin section plane polarized photomicrographs of (A) Uniserial benthic foraminifer in
breccia clast, U1456E-15R-1, 12-16 cm (1072 mbsf); (B) Siltstone with planktonic foraminifers,
U1456D-58R-2, 40-44 cm (989 mbsf); (C) Limestone clast with a specimen of *Lockhartia*,
U1456E-17R-4, 131-133 cm (1086 mbsf); (D) Echinoderm spine in carbonate clast, U1456D61R-2, 44-48 cm (1017 mbsf); (E) Foraminifer fragments in siltstone, U1456D-58R-2, 40-44 cm
(989 mbsf); (F) Planktic foraminifers and bioclasts in carbonate breccia, U1457C-90R-1-6-10 cm
(1034 mbsf). G) Planktonic foraminifer, U456E-7R-1, 80-82 cm (999 mbsf); H) Fragments of

962 coralline algae (white arrows) included in the planktonic-foraminifer-dominated matrix; Plk =

963 planktonic foraminifer, U1457C-88R-4, 58-60 cm (1021 mbsf); I) Orthophragminid fragment

964 (white arrow) included in the planktonic-foraminifer-dominated-matrix; Dl = dolomite crystal,

965 U1457C-88R-4, 58-60 cm (1021 mbsf). See locations on Figure 3.

966

967 Figure 13. (A) Geochemical signature of the analyzed samples illustrated by a CN-A-K ternary 968 diagram (Fedo et al., 1995). CN denotes the mole weight of Na₂O and CaO* (CaO* represent the 969 CaO associated with silicate, excluding all the carbonate (Singh et al., 2005)). A and K indicate 970 the content of Al₂O₃ and K₂O respectively. CIA values are calculated and shown on the left side, 971 with values correlated with on the CN-A-K ternary. Samples from the delta have the lowest CIA 972 values and indicate high contents of CaO and Na₂O and plagioclase. Abbreviations: sm 973 (smectite), pl (plagioclase), ksp (K-feldspar), il (illite), m (muscovite). B) Geochemical 974 classification of sediments from the Indus delta (Clift et al., 2010), Indus Canyon (Li et al., 2018) and western Indian Peninsular shelf north of Goa (Kurian et al., 2013) following the scheme of 975 976 Herron (1988). Phase 1 and Phase 2 sediments, together with the hemipelagic drape are the 977 materials of the MTC.

978

979	Figure 14. Cross plot of Sr versus Nd isotope data from the MTC, adjacent drill sites, major
980	source regions onshore, and modern Mahi, Tapti, and Narmada River sediments (Goswami et al.,
981	2012). Kirthar and Sulaiman data is from Zhuang et al. (2015). Deccan Plateau data are from
982	GEOROC compilation (http://georoc.mpch-mainz.gwdg.de/georoc/). Transhimalaya data are
983	from Rolland et al. (2002), Singh et al. (2002), and Khan et al. (1997). Greater Himalayan data
984	are from Ahmad et al.(2000), Deniel et al. (1987), Inger et al. (1993) and Parrish and Hodges
985	(1996). Karakoram data are from Crawford and Searle (1992) and Schärer et al. (1990),
986	
987	Figure 15. Downhole plots of Nd and Sr isotope compositions and clay mineralogy of
988	siliciclastic sediments from IODP sites (A) U1456 and (B) U1457. Gray vertical bar shows
989	compositional range of Quaternary sediments in the Indus Delta (Clift et al., 2010; Clift et al.,
990	2008b), as well as modern Tapti and Narmada River sediments (Goswami et al., 2012). Deccan
991	Plateau volcanic rocks plot outside this range at more positive ϵ_{Nd} values and lower ${}^{87}Sr/{}^{86}Sr$
992	values. Nd and Sr isotope analyses include errors recently suggested by Jonell et al. (2018) for
993	bulk sediment compositions. Error bars encompass the total expected error for any bulk sample
994	as a result of variable grain size and mineralogy, and analytical error contributed during sample
995	preparation, homogenization, and analysis.
996	
997	Figure 16. Kernel density estimate (KDE) plots for detrital zircon U-Pb ages for the Nataraja

998 MTC compared to major source terrains in the western Himalayas, from the compilation of

Alizai et al. (2011), as well as a modern sand from the river mouth (Clift et al., 2004) and select

1000 Indus Fan turbidites also from IODP Sites U1456 and U1457 (Clift et al., 2018). Deccan Zincons

1001 at ~65 Ma would plot within the Karakoram-Kohistan range but the inset box at the top shows 1002 that grains <100 Ma from the MTC do not cluster at this age and are better match to sources in 1003 the Indus suture zone. Data from the Tethyan, Greater and Lesser Himalaya are compiled from 1004 DeCelles et al. (2004). Karakoram data is from Le Fort et al. (1983), Parrish and Tirrul (1989), 1005 Schärer et al. (1990), Fraser et al. (2001) and Ravikant et al. (2009). Nanga Parbat data is from 1006 Zeitler and Chamberlain (1991) and Zeitler et al. (1993), Transhimalayan data is from Honegger 1007 et al. (1982), Schärer et al. (1984), Krol et al. (1996), Weinberg and Dunlap (2000), Zeilinger et 1008 al. (2001), Dunlap and Wysoczanski (2002), (Singh et al., 2007), and Ravikant et al. (2009). 1009

Figure 17. Ternary diagram of clay minerals from IODP Site U1456 and U1457 indicates a clay
mineral assemblage consisting mostly of smectite, chlorite and illite. Clay mineral data from
source regions are plotted to compare their clay assemblages. Data from western Indian shelf
modern sediments are from Rao and Rao (1995), Indus canyon data is from Li et al. (2018),
Indus flood plain and delta data is from Alizai et al. (2012), and Indus Fan data is from Peng
Zhou (unpublished).

1016

Figure 18. (A) Interpretation of the depth-converted seismic section of the western Indian
continental shelf immediately to the south of the source region for the Nataraja Slide based on
the seismic profile of Nair and Pandey (2018) and using the seismic velocities shown in Table 5;
and (B) A calculated sediment budget for the Indian shelf derived from two-dimensional
sediment backstripping of this profile derived from FlexDecomp[™] software.

1023	Figure 19. Schematic cartoon illustrating the over-steepened Indian margin (A), the first phase
1024	of emplacement of the Nataraja MTC (B) separated by a short time of quiescence with
1025	hemipelagic sedimentation (C) from the second smaller phase of emplacement (D).
1026	
1027	Table 1. Bulk sediment geochemistry analyzed by ICP-ES.
1028	
1029	Table 2. Neodymium and strontium isotope data.
1030	
1031	Table 3. Heavy mineral data. HM = heavy minerals; tHM = transparent heavy minerals. The
1032	ZTR index is the sum of zircon, tourmaline and rutile over total transparent heavy minerals
1033	(Hubert, 1962) and is classically used to estimate the mineralogical durability of the assemblage
1034	(i.e., the extent of recycling and/or intrastratal dissolution).
1035	
1036	Table 4. U-Pb zircon data for sample U1456C-71R-1, 110 cm.
1037	
1038	Table 5. Quantitative estimates of major clay mineral assemblages.
1039	
1040	Table 6. Seismic interval velocities for the main stratigraphic units interpreted by Nair and
1041	Pandey (2018) used to depth convert the seismic profile before backstripping.
1042 1043 1044	

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Figure 1 Dailey et al.

(A) Site U1456



(B) Site U1457









Figure 4 Dailey et al.



Figure 5 Dailey et al.





Figure 7 Dailey et al.



Figure 8 Dailey et al.



Figure 9 Dailey et al.



Figure 10 Dailey et al.



Figure 11 Dailey et al.



Figure 12 Dailey et al.



Figure 13 Dailey et al.





Figure 15 Dailey et al.



Figure 16 Dailey et al.
Figure 17





Figure 18



Figure 18 Dailey et al

Figure 19



Figure 19 Dailey et al

Sample (mbsf) (%) ((ppm) 403 54
IODP U1456D 35R-4, 107-122 cm 784.47 0.18 48.71 0.09 8.98 3.75 15.80 0.98 0.32 0.15 2.87 162.01 119.59 4 38R-1, 22-24 cm 808.00 0.12 52.99 0.04 7.93 3.33 17.12 1.12 0.19 0.40 2.98 152.43 96.48 4	403 54
35R-4, 107-122 cm 784.47 0.18 48.71 0.09 8.98 3.75 15.80 0.98 0.32 0.15 2.87 162.01 119.59 4 38R-1, 22-24 cm 808.00 0.12 52.99 0.04 7.93 3.33 17.12 1.12 0.19 0.40 2.98 152.43 96.48 4	403 54
38R-1, 22-24 cm 808.00 0.12 52.99 0.04 7.93 3.33 17.12 1.12 0.19 0.40 2.98 152.43 96.48 4	405.54
	431.06
42R-6, 40-42 cm 854.30 0.13 53.39 0.05 8.56 4.15 16.21 0.97 0.18 0.69 3.34 162.15 82.60 3	368.68
46R-4, 8-10 cm 883.40 0.12 53.59 0.04 8.33 3.17 16.77 1.44 0.27 0.64 2.90 199.20 89.80 4	404.49
49R-1, 50-52 cm 908.00 0.12 49.57 0.06 9.50 3.01 17.18 1.51 0.54 0.38 2.72 178.69 125.28 4	414.05
52R-5, 65-67 cm 943.10 0.10 56.31 0.04 6.78 4.13 15.81 1.00 0.20 0.77 3.22 204.61 80.95 3	350.21
54R-1, 5-7 cm 956.50 0.12 53.22 0.04 8.28 3.18 17.09 1.29 0.23 0.48 3.09 166.90 90.29 3	366.60
59R-5, 10-12 cm 1002.60 0.09 54.77 0.05 7.50 3.17 11.51 0.71 1.31 0.55 2.30 171.54 738.40 20	2088.35
60R-1, 109-111 cm 1006.50 0.26 53.12 0.10 6.96 3.37 11.39 0.74 1.48 0.39 1.98 82.44 308.40 18	1806.13
IODP U1456E	
9R-4, 64-66 cm 1021.00 0.04 51.60 0.07 8.54 3.43 12.40 0.89 1.63 0.15 1.56 159.79 282.36 18	1890.01
12R-1, 112-114 cm 1044.70 0.07 51.45 0.03 8.85 3.25 12.89 0.89 1.25 0.28 1.84 181.84 108.74 7	736.09
17R-7, 42-44 cm 1090.10 0.06 49.46 0.05 9.64 3.26 14.60 1.05 1.21 0.20 2.17 159.60 108.94 12	1250.60
IODP 1457C	
68R-7, 128-130 cm 842.50 0.11 56.57 0.04 7.16 3.42 16.51 1.05 0.20 0.74 2.89 168.76 82.06 4	409.00
69R-4, 104-107 cm 846.53 0.15 53.50 0.17 7.60 4.92 15.50 0.91 1.30 0.65 2.84 176.97 100.61 3	346.80
69R-7, 13-16 cm 850.38 0.12 51.61 0.08 8.59 3.55 16.02 1.02 0.65 0.50 3.09 161.96 104.08 3	385.13
70R-1, 6-8 cm 851.70 0.12 54.35 0.04 8.05 3.18 17.25 1.22 0.43 0.56 3.20 154.32 85.26 4	416.50
70R-5, 95-98 cm 858.16 0.34 53.30 0.29 7.90 4.85 14.87 0.86 1.05 0.67 2.84 169.51 105.17 3	344.01
71R-1, 7-9 cm 861.50 0.12 52.49 0.04 8.26 3.36 17.39 1.20 0.30 0.37 3.12 143.35 73.07 3	337.56
71R-2, 109-111 cm 863.49 0.17 50.69 0.11 6.35 3.18 13.48 0.97 7.89 0.66 2.39 191.88 317.03 3	305.86
72R-1, 107-110 cm 871.49 0.14 51.68 0.10 8.23 3.53 16.34 1.04 2.21 0.46 2.94 173.34 142.98 3	346.71
73R-1, 10-12 cm 857.78 0.11 53.52 0.03 8.18 3.28 17.40 1.12 0.09 0.38 3.39 156.13 69.86 3	357.88
74R-2, 19-21 cm 891.75 0.12 62.72 0.03 5.73 2.76 14.14 0.96 0.50 1.00 2.67 216.07 93.56 3	370.40
75R-1, 36-40 cm 900.06 0.15 50.12 0.20 8.51 4.00 16.01 0.98 2.34 0.28 2.91 162.80 155.71 3	325.69
76R-1, 90-92 cm 911.89 0.15 58.24 0.04 7.05 3.10 16.64 1.01 0.42 0.60 2.98 183.48 83.48 3	333.03
76R-3, 44-47 cm 912.57 0.16 49.16 0.11 9.53 3.67 16.17 1.21 1.95 0.43 2.69 176.56 136.57 3	344.65
77R-2, 3-5 cm 923.50 0.13 53.97 0.04 8.05 3.30 17.02 1.21 0.28 0.43 2.93 171.88 78.31 3	359.82
77R-5, 26-29 cm 924.94 0.11 50.07 0.09 8.35 3.60 16.23 1.00 3.02 0.38 2.88 162.60 188.13 3	320.07
78R-4, 25-28 cm 933.28 0.14 49.34 0.13 9.62 3.79 16.33 1.17 1.61 0.39 2.80 178.48 132.21 3	305.89
78R, 5, 50-52 cm 933.78 0.12 56.60 0.03 6.28 3.10 15.49 1.00 0.30 0.68 2.95 182.32 81.19 3	325.06
79R-5, 120-122 945.74 0.15 58.44 0.04 7.06 3.16 16.02 1.02 0.40 0.66 3.11 215.88 81.25 2	298.55
81R-1, 30-32 cm 958.20 0.15 53.22 0.12 7.45 3.46 14.95 0.90 3.36 0.45 2.77 179.71 157.01 2	270.30
81R-1, 42-44 cm 958.32 0.15 57.32 0.05 7.19 3.34 16.27 0.97 0.42 0.62 3.15 176.67 81.33 2	279.34
82R-3, 6-9 cm 970.66 0.15 58.78 0.13 6.10 2.96 11.45 0.73 4.59 0.93 2.13 246.76 168.04 2	245.10
82R-3, 89-91 cm 971.55 0.17 62.47 0.04 6.40 2.99 14.53 0.96 0.62 0.95 2.88 222.25 94.57 3	306.87
83R-1, 61-63 cm 978.00 0.10 52.09 0.03 7.98 3.13 17.49 1.25 0.16 0.35 3.06 168.56 67.93 2	282.34
84R-6, 101-103 cm 991.90 0.14 56.77 0.06 7.12 3.88 15.83 0.90 0.28 0.69 3.31 185.01 78.23 2	266.53
85R-1, 60-62 cm 998.40 0.12 57.92 0.05 7.09 3.98 16.02 0.92 0.29 0.71 3.36 184.78 82.51 2	274.00
85R-3, 46-49 cm 999.74 0.16 53.58 0.15 7.86 4.65 15.32 0.89 0.99 0.72 3.21 185.22 116.33 2	261.00
86R-2, 44-47 cm 1008.04 0.11 18.69 0.07 2.73 1.75 6.14 0.31 31.32 0.07 0.95 57.03 720.11 6	611.45
87R-1, 14-18 cm 1011.22 0.10 15.76 0.07 2.11 1.45 5.20 0.24 32.94 0.08 0.71 47.83 649.62 13	1357.92
89R-2, 57-59 cm 1027.62 0.08 52.64 0.06 8.63 3.41 11.45 0.76 1.93 0.26 1.63 94.10 222.78 13	1315.27
89R-3, 119-121 cm 1029.30 0.05 52.31 0.03 7.46 4.03 13.01 0.77 0.90 0.43 2.41 138.94 82.01 3	308.97
96R-1, 62-66 cm 1090.74 0.09 47.67 1.10 10.46 3.37 12.29 0.95 0.92 0.49 2.22 139.86 113.58 1	102.65

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Sample	⁸⁷ Sr/ ⁸⁶ Sr	¹⁴³ Nd/ ¹⁴⁴ Nd	Epsilon Nd
IODP U1456D			
35R-4, 107-122 cm	0.718548	0.512186	-8.8
38R-1, 22-24 cm	0.719026	0.512219	-8.2
42R-6, 40-42 cm	0.716960	0.512233	-7.9
46R-4, 8-10 cm	0.717475	0.512210	-8.3
49R-1, 50-52 cm	0.718123	0.512200	-8.5
52R-5, 65-67 cm	0.717166	0.512191	-8.7
54R-1, 5-7 cm	0.720084	0.512137	-9.8
59R-5, 10-12 cm	0.708516	0.512560	-1.5
IODP U1456F			
12R-1 112-114 cm	0 713705	0 512251	-75
17R-7 42-44 cm	0.715006	0.512184	-8.9
19R-CC, 17-22 cm	0.725510	0.512101	-8.7
IODB 1457C			
68P 7 128 130 cm	0 717787	0 512160	0.1
60R 1 100 104 cm	0.717/07	0.512109	-9.1
$69R_{-5}$ 136-148 cm	0.713033	0.512187	-8.6
$70R_{-4}$ 137_152 cm	0.717610	0.512190	-8.6
$70R_{-5}$, 95_{-97} cm	0.717017	0.512305	-6.5
$71R_{-6}$ 18-28 cm	0.719332	0.512303	-9.6
74R-2 19-21 cm	0.719552	0.512147	-10.1
75R-1 36-40 cm	0.712895	0.512207	-8.4
78R, 5, 50-52 cm	0.719460	0.512089	-10.7
81R-1, 30-32 cm	0.713224	0.512169	-9.1
83R-1, 61-63 cm	0.721653	0.512128	-9.9
87R-1, 14-18 cm	0.708510	0.512357	-5.5
89R-3, 119-121 cm	0.716107	0.512198	-8.6
96R-1, 62-66 cm	0.709144	0.512348	-5.7

Sample	Facies	Grain size (µm)	% fine tail	% class analyzed	% coarse tail	n° transparent HM	tot. grains counted	HMC %weight	tHM% weight	zircon	tourmaline	rutile	anatase/brookite	apatite	titanite
Tapti River	Sand bar	15-500	15%	78%	7%	215	309	24	17	0	0	0	0	0	1
1456E-4R-1, 110 cm 1456E-7R-1, 80 cm 1456E-15R-1, 61 cm 1456E-17R-4, 131 cm 1457C-88R-4, 58 cm	Packstone Packstone Breccia Breccia Packstone	15-500 15-500 15-500 15-500 15-500	63% 46% 52% 50% 42%	37% 54% 33% 24% 53%	0.1% 0.2% 15% 26% 5%	209 21 68 64 216	3336 1449 2557 1144 819	0.08 0.06 0.04 0.05 0.14	0.02 0.00 0.00 0.01 0.08	5 5 10 6 4	13 5 4 6 6	2 0 1 3 1	1 0 0 1	24 10 16 22 11	10 5 13 8 9
	andalusite	kyanite	sillimanite	amphibole	green augite	Cr-spinel	Total	ZTR	% transparent	% opaque	% Fe oxide	% Ti oxide	% HM turbid	% rock fragments	% soils & turbid
Tapti River	0	0	1	2	92	0.0	4.7	0	70	24	4	0	0	1	0
1456E-4R-1, 110 cm	0.5	2	0	2	1	4	90.4	20	6	8	7	1	0	0	0
1456E-7R-1, 80 cm	0	0	0	10	48	0	42.9	10	1	61	0	0	0	0	0
1456E-15R-1, 61 cm	3	1	0	0	6	3	86.8	16	3	19	5	0	0	0	0
1456E-17R-4, 131 cm	0	0	0	3	6	2	89.1	16	6	10	7	0	0	0	0
1457C-88R-4, 58 cm	0.5	1	0	8	1	2	87.5	11	26	11	5	1	0	0	0

barite	diaspore	epidote	garnet	chloritoid	staurolite	
0	0	2	1	0	0	
3	0	9	15	8	1	
5	0	0	14	0	0	
4	0	9	21	3	4	
5	0	9	25	3	2	
3	0.5	30	18	4	0.5	
% phosphate	% chlorite	% biotite	% carbonates	% light minerals	Total	
0	0	0	0	1	100	
35	6	6	30	1	100	
17	0	1	19	0	100	
57	1	1	13	1	100	
59	1	3	10	3	100	
19	6	12	15	5	100	

									(Concordant S	Scans: Age	\$			
Preferred			Concorda	nt						²⁰⁶ Pb/ ²³⁸ U			²⁰⁷ Pb/ ²⁰⁶ Pb		
Age (Ma)	2 -sigma	2 +sigma	Scans	²⁰⁷ Pb/ ²³⁵ U	2 sigma	²⁰⁶ Pb/ ²³⁸ U	2 sigma	²⁰⁷ Pb/ ²⁰⁶ Pb	2 sigma	Age (Ma)	2 -sigma	2 +sigma	Age (Ma)	2 -sigma	2 +sigma
47.01	1.37	1.37	64	0.05045	0.00660	0.00732	0.00021	0.04999	0.00657	47.01	1.37	1.37	194.64	320.86	292.11
47.33	1.26	1.26	64	0.05143	0.00657	0.00737	0.00020	0.05062	0.00651	47.33	1.26	1.26	223.55	311.81	284.54
47.79	1.85	1.85	46	0.04993	0.00845	0.00744	0.00029	0.04866	0.00834	47.79	1.85	1.85	131.60	263.20	379.94
48.02	1.73	1.73	46	0.05000	0.00839	0.00748	0.00027	0.04850	0.00826	48.02	1.73	1.73	123.77	247.55	378.11
48.53	1 99	1 99	47	0.05389	0.00852	0.00756	0.00031	0.05172	0.00821	48.53	1 99	1.99	273.19	385.93	344.89
48.99	3 14	3 14	20	0.05734	0.01738	0.00763	0.00049	0.05452	0.01668	48.99	3 14	3.14	392.40	771 22	621.88
49.14	1.90	1.90	47	0.05466	0.00857	0.00765	0.00030	0.05181	0.00818	40.77	1.90	1.90	277.18	383.24	342 72
49.14	1.90	1.90	48	0.05400	0.00847	0.00765	0.00020	0.05124	0.00809	49.14	1.90	1.90	251.45	38/ 05	344.15
49.17	3.20	3 20	10	0.05756	0.00047	0.00769	0.00029	0.05432	0.01715	49.17	3.20	3.20	384.50	769.00	640.63
50.14	2.00	2.00	27	0.05046	0.01166	0.00701	0.00030	0.05432	0.01/15	50.14	2.00	2.00	421.89	472.00	412.19
50.14	2.00	2.00	24	0.05940	0.01100	0.00781	0.00031	0.05324	0.01089	50.29	2.00	2.00	421.88	472.90	412.10
70.08	2.50	2.50	20	0.03839	0.01191	0.00783	0.00030	0.03427	0.00220	70.08	2.50	2.50	167.27	490.10	429.90
70.08	2.50	2.50	20	0.07447	0.01329	0.01093	0.00039	0.04941	0.00869	70.08	2.50	2.50	107.57	220.02	293.24
70.13	2.20	2.20	39	0.07401	0.01292	0.01094	0.00036	0.04946	0.00803	70.15	2.20	2.20	109.31	269.05	266.09
70.16	2.20	2.20	43	0.07510	0.01242	0.01094	0.00035	0.04977	0.00830	70.16	2.20	2.20	184.37	308.75	300.98
72.23	2.74	2.74	37	0.08171	0.01524	0.01127	0.00043	0.05259	0.00993	72.23	2.74	2.74	311.22	460.83	403.35
/6.83	3.05	3.05	68	0.08023	0.00532	0.01199	0.00048	0.04853	0.00298	/6.83	3.05	3.05	125.30	148.01	141.61
77.15	3.05	3.05	67	0.08014	0.00533	0.01204	0.00048	0.04828	0.00299	77.15	3.05	3.05	112.82	149.29	142.79
100.16	3.15	3.15	59	0.10617	0.00460	0.01566	0.00050	0.04917	0.00188	100.16	3.15	3.15	156.18	90.86	88.40
100.99	2.82	2.82	62	0.10697	0.00434	0.01579	0.00044	0.04914	0.00181	100.99	2.82	2.82	154.32	87.35	85.07
118.08	5.47	5.47	58	0.12988	0.01607	0.01849	0.00086	0.05096	0.00637	118.08	5.47	5.47	238.93	301.60	276.00
120.22	5.26	5.25	58	0.13169	0.01594	0.01882	0.00083	0.05074	0.00622	120.22	5.26	5.25	228.91	296.48	271.72
484.75	13.50	13.48	120	0.60703	0.01998	0.07810	0.00226	0.05637	0.00139	484.75	13.50	13.48	467.11	55.05	54.12
519.35	18.13	18.10	111	0.67939	0.02913	0.08390	0.00305	0.05873	0.00208	519.35	18.13	18.10	557.10	78.23	76.36
547.29	23.73	23.68	71	0.73343	0.05847	0.08861	0.00400	0.06003	0.00498	547.29	23.73	23.68	604.79	184.88	174.70
565.78	24.78	24.73	72	0.75974	0.06307	0.09173	0.00419	0.06007	0.00510	565.78	24.78	24.73	606.01	189.07	178.44
691.81	24.28	24.23	17	1.05756	0.04642	0.11329	0.00419	0.06771	0.00244	691.81	24.28	24.23	859.54	75.81	74.02
742.77	112.79	111.81	5	1.28928	0.25300	0.12212	0.01955	0.07657	0.01460	742.77	112.79	111.81	1109.98	406.32	358.96
754.15	21.48	21.45	92	1.08175	0.06391	0.12410	0.00374	0.06322	0.00367	754.15	21.48	21.45	715.56	125.85	121.02
762.15	21.70	21.67	99	1.13684	0.04166	0.12550	0.00379	0.06570	0.00233	762.15	21.70	21.67	796.76	75.25	73.49
764.53	19.53	19.50	84	1.09582	0.06500	0.12592	0.00341	0.06312	0.00375	764.53	19.53	19.50	712.23	128.77	123.72
766.14	25.46	25.41	41	1.19877	0.06376	0.12620	0.00444	0.06889	0.00368	766.14	25.46	25.41	895.57	112.18	108.28
937.35	22.90	22.86	98	1.55283	0.04656	0.15651	0.00410	0.07196	0.00189	937.35	22.90	22.86	984.77	53.80	52.88
997.31	29.09	29.03	97	1.73089	0.05731	0.16732	0.00526	0.07503	0.00227	997.31	29.09	29.03	1069.26	61.51	60.31
1002.01	40.31	40.18	126	1.68389	0.07518	0.16817	0.00729	0.07262	0.00249	1002.01	40.31	40.18	1003.43	70.40	68.84
1017.09	21.65	21.62	104	1.79155	0.04889	0.17090	0.00393	0.07603	0.00212	1017.09	21.65	21.62	1095.83	56.39	55.38
1064.57	23.22	23.18	79	1.90134	0.06993	0.17956	0.00424	0.07680	0.00272	1064.57	23.22	23.18	1115.94	71.52	69.90
1173.37	25.27	25.22	68	2.30640	0.06114	0.19964	0.00470	0.08379	0.00200	1173.37	25.27	25.22	1287.68	46.92	46.21
1703.26	43.19	42.57	107	4.30372	0.10657	0.29906	0.00565	0.10437	0.00243	1686.66	28.06	28.00	1703.26	43.19	42.57
1719.35	37.26	36.80	121	4.36640	0.09580	0.30078	0.00576	0.10529	0.00212	1695.16	28.56	28.50	1719.35	37.26	36.80
1734.05	45.62	44.93	35	4.14546	0.12716	0.28328	0.00809	0.10613	0.00262	1607.86	40.70	40.57	1734.05	45.62	44.93
1751.80	44.50	43.84	92	4.44995	0.11308	0.30116	0.00596	0.10717	0.00259	1697.04	29.55	29.48	1751.80	44.50	43.84
1809.44	50.74	49.89	114	4.95575	0.14377	0.32495	0.00818	0.11061	0.00306	1813.86	39.84	39.72	1809.44	50.74	49.89
1824.80	35.56	35.14	85	4.65207	0.13339	0.30247	0.00829	0.11155	0.00217	1703.54	41.07	40.94	1824.80	35.56	35.14
2126.41	196.81	184.41	11	6.01281	0.90134	0.33006	0.05028	0.13213	0.01436	1838.66	246.02	241.41	2126.41	196.81	184.41
2404.03	46.88	46.13	115	9.69673	0.27654	0.45314	0.01182	0.15520	0.00425	2409.19	52.52	52.31	2404.03	46.88	46.13
2417.39	46.37	45,64	103	9,10358	0.25984	0.42209	0.01083	0.15643	0.00424	2269.95	49.21	49.02	2417.39	46.37	45.64
2504.10	52.23	51.30	45	9,98345	0.42634	0.43973	0.01793	0.16466	0.00507	2349.43	80.51	80.01	2504.10	52.23	51.30
2584.49	44.66	43.98	101	11.36799	0.39304	0.47726	0.01606	0.17275	0.00459	2515.34	70.28	69.90	2584.49	44.66	43.98
2594.62	47 11	46 35	96	11 43402	0.43489	0.47713	0.01803	0 17380	0.00487	2514 75	78.95	78 47	2594.62	47.11	46 35
2636.60	60.15	58.92	103	11 99958	0 44152	0.48876	0.01801	0 17824	0.00437	2563.14	78.26	77 79	2636.60	60.15	58.92
2649.00	61 30	60.11	102	12 03610	0.46870	0.48610	0.01015	0 17958	0.00057	2553.78	83 33	82.80	2649.00	61 39	60.11
2047.00	01.57	00.11	102	12.05010	0.40070	0.10010	0.01/15	0.1720	0.00057	2000.10	05.55	02.00	2017.00	01.57	00.11

	Smectite	Chlorite	Illite	Kaolinite	Palygorskite
Sample	(%)	(%)	(%)	(%)	(%)
IODP 111456D					
35R-4 107-122 cm	37.6	18.7	33.0	9.6	0.0
39R-1, 12-14 cm	40.9	13.8	36.8	8.2	0.0
40R-1 60-62 cm	43.1	13.9	8.0	8.1	34.1
43R-7 2-4 cm	62.2	12.9	19.9	4.4	0.0
47R-4 58-60 cm	65.2	13.8	13.4	7.3	0.0
50R-1 31-33 cm	53.8	15.5	21.9	8.5	0.0
51R-6 20-22 cm	61.4	11.9	15.8	10.0	0.0
53R-1 5-7 cm	49.1	14.6	27.0	8.7	0.0
57R-7 75-77 cm	53.1	8.4	26.5	0.0	10.6
58R-2 2-4 cm	69.9	3.8	17.5	0.0	67
61R-1 40-42 cm	91.9	0.8	4.8	0.0	2.5
01IC-1, 40-42 cm	<i>J</i> 1. <i>J</i>	0.0	1.0	0.0	2.5
IODP U1456E					
5R-2, 25-27 cm	100.0	0.0	0.0	0.0	0.0
14R-1, 75-77 cm	96.9	0.0	0.0	0.0	0.0
16R-2, 5-7 cm	54.8	4.6	22.5	0.0	18.0
19R-CC, 17-22 cm	40.0	25.0	28.3	5.1	0.0
IODP 01457C	17.0	27.2	22.5	10.0	0.0
69R-4, 104-106 cm	17.2	21.2	33.5 25.6	10.9	9.9
69R-6, 13-15 cm	23.2	26.7	35.0	8.2	4.8
69R-7, 112-114 cm	29.0	22.0	37.4	9.7	0.0
/0R-5, 95-9/ cm	32.3 29.5	18.8	30.6	9.6	7.6
71R-2, 109-111 cm	38.3 20.2	17.9	33.Z	9./	0.0
72R-1, 107-109 cm	30.3	23.2	34.2	10.8	0.0
74R-2, 25-27 cm	31.8	23.1	33.2	10.3	0.0
75R-1, 36-40 cm	30.2	21.0	33.0 25.9	8.4 11.2	0.0
76R-3, 44-46 cm	44.5	1/.9	25.8	11.2	0.0
7/R-5, 26-28 cm	41.3	19.2	28.5	10.1	0.0
/8R-4, 25-27 cm	39.1	19.9	28.1	11.9	0.0
/9R-5, 129-131 cm	26.2	23.1	37.9	11./	0.0
81R-1, 30-32 cm	27.1	22.2	39.3 42.6	10.4	0.0
82R-3, 6-8 cm	23.3	20.9	42.0	10.2	0.0
83R-2, 18-20 cm	$\frac{3}{.}$	10.8	34.1 20.0	10.4	0.0
84R-3, 143-145 cm	23.1	25.0	39.9	10.4	0.0
85R-3, 46-48 cm	31.1 20.6	13.8	30.2 25.2	ð.1 0.1	8.1
86R-1,6-8 cm	29.0 75.5	10.2 5 1	55.5 14.4	9.1	6.9 1.5
80K-2, 44-46 cm	13.3	J.1	14.4	3.1 2.6	1.5
δ/K-1, 14-18 cm	//.0	4.3	13.0	2.0	1.5
95K-1, 50-52 cm	100.0	0.0	0.0	0.0	0.0
95K-3, 50-52 cm	100.0	0.0	0.0	0.0	0.0
94K-2, 55-5 / cm	100.0	0.0	0.0	0.0	0.0
95К-1, 12-14 cm	100.0	0.0	0.0	0.0	0.0

Table 5

96R-1, 62-66 cm	100.0	0.0	0.0	0.0	0.0
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Dailey et al.

Table 6

Stratigraphic	Age at base of	Interval velocity
interval	interval	(km/s)
H7	Top Miocene	1.70
H6	Top M. Miocene	2.20
Н5	Top L. Miocene	2.40
H4	Top. L. Oligocene	2.45
H3	Top. L. Eocene	2.50
H2	Top U. Paleocene	2.60
H1	Basement (66 Ma)	2.65