## The Flatness Problem and the Pulsating Universe.

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(Nuovo Cimento B, 108 (1993) 1253)

PACS 04.20.Cv - Fundamental problems and general formalism.

PACS 98.80.Dr - Theoretical cosmology.

PACS 99.10 - Errata.

On p. 1255, second paragraph, line 2, for eq. (3), read eq. (2); on line 3, for eq. (4), read eq. (3); on the last line of the page, for eq. (3), read eq. (2).

On p. 1263, following eq. (57) for  $\xi = (1 - \cos \phi)$  read  $\xi = (1 - \varepsilon \cos \phi)$ , and in eq. (58), after the integral sign, for  $(1 - \cos \phi)$ , read  $(1 - \varepsilon \cos \phi)$ .

## Light Travel Times around a Closed Universe.

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On p. 932, eq. (1), for  $ds^2 = c^2 dt^2 - a_e^2 (d\theta_1^2 + \sin^2 \theta_1)(d\theta_2^2 + \sin^2 \theta_2 d\theta_3^2)$ , read  $ds^2 = c^2 dt^2 - a_e^2 [d\theta_1^2 + \sin^2 \theta_1 (d\theta_2^2 + \sin^2 \theta_2 d\theta_3^2)]$ .

Also, in both these papers, as well as in the one that preceded them (Nuovo Cimento B, 108 (1993) 911), the energy-stress tensor  $T^{\mu}_{\nu}(A^2)$  was interpreted as being associated with a gas of «free particles» in negative-energy states. I no longer believe this interpretation to be valid, and that rather the tensor is most likely associated with a band of filled negative-energy states at the top of the Dirac sea which are not shielded for their gravitational effects in the field equations. A paper discussing this revised interpretation is in preparation. Also, the discussion of the solution for the expansion parameter for the case  $A^2 = 0$ ,  $B^2 > 0$  exhibits heretofore unnoticed anomalous behavior in the neighborhood of t = 0, and a detailed analysis is needed and will be given in a forthcoming paper.