



# Lower oceanic $\delta^{13}\text{C}$ during the last interglacial period compared to the Holocene

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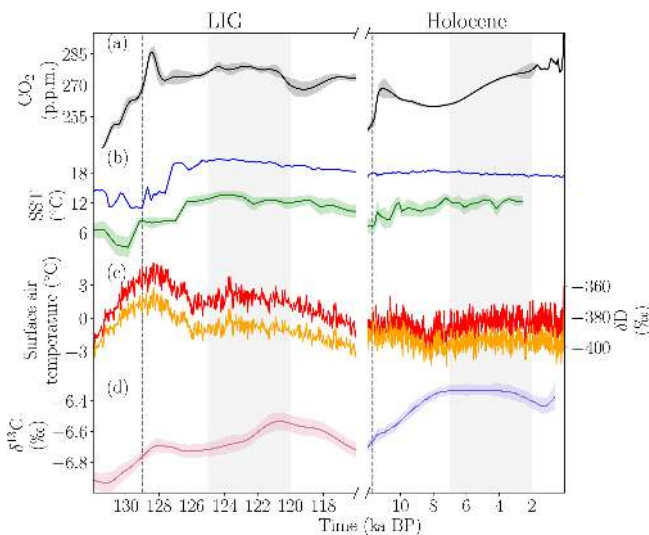
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**Abstract.** The last time in Earth's history when high latitudes were warmer than during pre-industrial times was the last interglacial period (LIG, 129–116 ka BP). Since the LIG is the most recent and best documented interglacial, it can provide insights into climate processes in a warmer world. However, some key features of the LIG are not well constrained, notably the oceanic circulation and the global carbon cycle. Here, we use a new database of LIG benthic  $\delta^{13}\text{C}$  to investigate these two aspects. We find that the oceanic mean  $\delta^{13}\text{C}$  was  $\sim 0.2\%$  lower during the LIG (here defined as 125–120 ka BP) when compared to the Holocene (7–2 ka BP). A lower terrestrial carbon content at the LIG than during the Holocene could have led to both lower oceanic  $\delta^{13}\text{C}$  and atmospheric  $\delta^{13}\text{CO}_2$  as observed in paleo-records. However, given the multi-millennial timescale, the lower oceanic  $\delta^{13}\text{C}$  most likely reflects a long-term imbalance between weathering and burial of carbon. The  $\delta^{13}\text{C}$  distribution in the Atlantic Ocean suggests no significant difference in the latitudinal and depth extent of North Atlantic Deep Water (NADW) between the LIG and the Holocene. Furthermore, the data suggest that the multi-millennial mean NADW transport was similar between these two time periods.

## 1 Introduction

The most recent and well documented warm time period is the last interglacial period (LIG), which is roughly equivalent to Marine Isotope Stage (MIS) 5e (Past Interglacials Working Group of PAGES, 2016; Shackleton, 1969). The LIG began at the end of the penultimate deglaciation and ended with the last glacial inception ( $\sim 129$ – $116$  ka BP; Dutton and Lambeck, 2012; Govin et al., 2015; Masson-Delmotte et al., 2013; Menviel et al., 2019). The LIG was globally warmer than the pre-industrial period (PI,  $\sim 1850$ – $1900$ ; IPCC, 2013; Shackleton et al., 2020), with PI estimated to be  $\sim 0.4^\circ\text{C}$  cooler than the peak of the Holocene (10–5 ka BP) (Marcott et al., 2013). Though not an exact analogue for future warming, the LIG may still help shed light on future climates. In particular, we seek to constrain the mean LIG ocean circulation and estimate the global oceanic mean  $\delta^{13}\text{C}$ .

As greenhouse gas concentrations were comparable to the Holocene, the LIG was most likely relatively warm because of the high boreal summer insolation (Laskar et al., 2004). During the LIG, the atmospheric  $\text{CO}_2$  concentration was relatively stable around  $\sim 280$  ppm (Bereiter et al., 2015; Lüthi et al., 2008), while during the Holocene  $\text{CO}_2$  first decreased by about 8 ppm starting at 11.7 ka BP before increasing by  $\sim 17$  to 277 ppm at  $\sim 2$  ka BP (Fig. 1a) (Köhler et al., 2017).  $\text{CH}_4$  and  $\text{N}_2\text{O}$  peaked at  $\sim 700$  and  $\sim 267$  ppb, respectively, during both the LIG and the Holocene (Flückiger et al., 2002;



**Figure 1.** LIG and Holocene time series of (a)  $\text{CO}_2$  stack smoothed with a spline based on the age model AICC2012 (Köhler et al., 2017), (b) sea surface temperatures (SSTs) determined from alkenones and aligned with oxygen isotopes from the Iberian Margin (MD01-2444, blue, Martrat et al., 2007b) and the North Atlantic (GIK23414-6, green, Candy and Alonso-Garcia, 2018), (c) EPICA Dome C ice core (EDC96) deuterium measurements (orange) and estimated surface air temperature anomaly relative to the mean of the last 1 kyr (red, Bazin et al., 2013; Jouzel et al., 2007) on the AICC2012 timescale, and (d) spline of atmospheric  $\delta^{13}\text{C}$  from EPICA Dome C and the Talos Dome ice cores (Holocene, Eggleston et al., 2016) and Monte Carlo average of three Antarctic ice cores atmospheric  $\delta^{13}\text{C}$  (LIG, Schneider et al., 2013) both based on the age model AICC2012. Shading around the lines indicates  $1\sigma$ . Vertical grey shading indicates the periods of analysis in this paper. Vertical dotted grey lines indicate the commencement of the LIG and Holocene.

Petit et al., 1999; Spahni et al., 2005). Global sea level was 6–9 m higher at the LIG compared to PI (Dutton et al., 2015; Kopp et al., 2009), thus indicating significant ice mass loss from both Antarctica and Greenland.

Strong polar warming is supported by terrestrial and marine temperature reconstructions. A global analysis of sea surface temperature (SST) records suggests that the mean surface ocean was  $0.5 \pm 0.3^\circ\text{C}$  warmer during the LIG compared to 1870–1889 (Hoffman et al., 2017), similar to another global estimate, which suggests SSTs were  $0.7 \pm 0.6^\circ\text{C}$  higher during the LIG compared to the late Holocene (McKay et al., 2011). However, there were differences in the timing of these SST peaks in different regions compared to the 1870–1889 mean: North Atlantic SSTs peaked at  $+0.6 \pm 0.5^\circ\text{C}$  at 125 ka BP (e.g. Fig. 1b) and Southern Hemisphere extratropical SSTs peaked at  $+1.1 \pm 0.5^\circ\text{C}$  at 129 ka BP (Hoffman et al., 2017). On land, proxy records from mid-latitudes to high latitudes indicate higher temperatures during the LIG compared to PI, particularly in North America (Anderson et al., 2014; Axford et al.,

2011; Montero-Serrano et al., 2011). Similarly, the EPICA DOME C record suggests that the highest Antarctic temperatures from the last 800 kyr occurred during the LIG (Masson-Delmotte et al., 2010) (Fig. 1c).

Polar warming was also associated with significant changes in vegetation. Pollen records suggest a contraction of tundra and an expansion of boreal forests across the Arctic (CAPE, 2006), in Russia (Tarasov et al., 2005), and in North America (Govin et al., 2015; Muhs et al., 2001; de Vernal and Hillaire-Marcel, 2008). The few Saharan records suggest a green Sahara period during the LIG (Drake et al., 2011; Larasoña et al., 2013), consistent with a stronger West African monsoon (Otto-Bliesner et al., 2021). Although these reconstructions indicate changes in vegetation distribution during the LIG, the total amount of carbon stored on land remains poorly constrained.

Recent numerical experiments of the LIG as part of the Paleomodel Intercomparison Project Phase 4 (PMIP4) simulate significant warming over Alaska and Siberia in boreal summer, with mean annual temperature anomalies of close to zero, which is in good agreement with the proxy record (Otto-Bliesner et al., 2021). Despite this and other recent data compilations and modelling efforts (including Bakker et al., 2013), there are many open questions remaining about the LIG. In particular, stronger constraints are needed on the extent of Greenland and Antarctic ice sheets; on ocean circulation and the global carbon cycle, including  $\text{CaCO}_3$  accumulation in shallow waters; and peat and permafrost carbon storage (Brovkin et al., 2016).

It is important to constrain the state of the Atlantic Meridional Overturning Circulation (AMOC) at the LIG given its significant role in modulating climate. Seven coupled climate models integrated with transient 130–115 ka BP boundary conditions simulate different AMOC trends, with some models producing a strengthening of the AMOC, while others simulate a weakening during the LIG (Bakker et al., 2013). Paleoproxy records suggest equally strong and deep North Atlantic Deep Water (NADW) during the LIG and the Holocene (e.g. Böhm et al., 2015; Lototskaya and Ganssen, 1999), with a possible southward expansion of the Arctic front related to changes in the strength of the subpolar gyre (Mokeddem et al., 2014), and AMOC weakening during a few multi-centennial-scale events between 127 and 115 ka BP (e.g. Galaasen et al., 2014b; Helmens et al., 2015; Lehman et al., 2002; Mokeddem et al., 2014; Oppo et al., 2006; Rowe et al., 2019; Tzedakis et al., 2018).

Stable carbon isotopes are a powerful tool for investigating ocean circulation (e.g. Curry and Oppo, 2005; Eide et al., 2017) and the global carbon cycle (e.g. Menviel et al., 2017; Peterson et al., 2014). Since the largest carbon isotope fractionation occurs during photosynthesis, organic matter is enriched in  $^{12}\text{C}$  (low  $\delta^{13}\text{C}$ ), while atmospheric  $\text{CO}_2$  and surface water dissolved inorganic carbon (DIC) become enriched in  $^{13}\text{C}$  (high  $\delta^{13}\text{C}$ ). Organic matter on land includes the terrestrial biosphere, as well as carbon stored in soils,

such as in peats and permafrost. Different photosynthetic pathways (which differentiate  $\text{C}_3$  and  $\text{C}_4$  plants) fractionate carbon differently, producing typical signatures of about  $-37\text{‰}$  to  $-20\text{‰}$  for  $\text{C}_3$  plants (Kohn, 2010) and around  $-13\text{‰}$  for  $\text{C}_4$  plants (Basu et al., 2015), though these values vary with a number of factors, including precipitation, atmospheric  $\text{CO}_2$  concentration and  $\delta^{13}\text{C}$ , light, nutrient availability, and plant species (Cernusak et al., 2013; Diefendorf et al., 2010; Diefendorf and Freimuth, 2017; Farquhar, 1983; Farquhar et al., 1989; Keller et al., 2017; Leavitt, 1992; Schubert and Jahren, 2012). In the ocean, phytoplankton using the  $\text{C}_3$  photosynthetic pathway are found to have fractionation during photosynthesis that depends on the concentration of dissolved  $\text{CO}_2$ . Thus, atmospheric  $\delta^{13}\text{CO}_2$  during the LIG (Fig. 1d) is influenced by the cycling of organic carbon within the ocean, changes in the amount of carbon stored in vegetation and soils, temperature-dependent air–sea flux fractionation (Lynch-Stieglitz et al., 1995; Zhang et al., 1995), and on longer timescales by interactions with the lithosphere (Tschumi et al., 2011). The mean surface DIC is enriched by  $\sim 8.5\text{‰}$  compared to the atmosphere due to fractionation during air–sea gas exchange (Menviel et al., 2015; Schmittner et al., 2013).

NADW is characterised by low nutrients and high  $\delta^{13}\text{C}$  as a result of a high nutrient and carbon utilisation by marine biota and fractionation during air–sea gas exchange in the northern North Atlantic. Along its path through the Atlantic basin interior, organic matter remineralisation and mixing with southern source waters lowers  $\delta^{13}\text{C}$ , with  $\delta^{13}\text{C}$  values of  $\sim 0.5\text{‰}$  in the deep Southern Ocean.

The tight relationship between the water masses' apparent oxygen utilisation, nutrient content and  $\delta^{13}\text{C}$  allows  $\delta^{13}\text{C}$  to be used as a water mass ventilation tracer (e.g. Boyle and Keigwin, 1987; Curry and Oppo, 2005; Duplessy et al., 1988; Eide et al., 2017). The  $\delta^{13}\text{C}$  of benthic foraminifera shells, particularly of the species *Cibicides wuellerstorfi*, has been found to reliably represent the  $\delta^{13}\text{C}$  signature of DIC (Belanger et al., 1981; Duplessy et al., 1984; Zahn et al., 1986) and has therefore been used to better constrain the extent of different water masses. Mass balances of  $\delta^{13}\text{C}$  between the atmosphere, ocean, and land have been previously used to constrain changes in terrestrial carbon between the last glacial maximum ( $\sim 20$  ka BP) and Holocene (e.g. Peterson et al., 2014). However, on longer timescales, exchanges with the lithosphere including volcanic outgassing (Hasencllever et al., 2017; Huybers and Langmuir, 2009),  $\text{CaCO}_3$  burial in sediments and weathering, release of carbon from methane clathrates, and the net burial of organic carbon also influence the global mean  $\delta^{13}\text{C}$ . It has been estimated that the amount of carbon both entering and exiting the lithosphere due to weathering and burial of organic carbon fluxes could be from  $0.274$  to  $0.344$   $\text{Gt C yr}^{-1}$  (Schneider et al., 2013), though these vary through time (Hoogakker et al., 2006). Over timescales greater than 10 kyr, the influence of weathering and burial of carbon might dominate the  $\delta^{13}\text{C}$  signal

(Jeltsch-Thömmes et al., 2019; Jeltsch-Thömmes and Joos, 2020), and thus a mass balance cannot be accurately applied to evaluate terrestrial carbon changes between the LIG and Holocene.

Here, we present a new compilation of benthic  $\delta^{13}\text{C}$  from *Cibicides wuellerstorfi* spanning the 130–118 ka BP time period. We use this data to compare the  $\delta^{13}\text{C}$  signal of the LIG with that of the Holocene and to determine the difference in average ocean  $\delta^{13}\text{C}$  between the two time periods. We then investigate the AMOC during the LIG with our new benthic  $\delta^{13}\text{C}$  database. Finally, we qualitatively explore the role of the various processes affecting the  $\delta^{13}\text{C}$  difference between the LIG and the Holocene.

## 2 Database and methods

### 2.1 Database

We present a new compilation of benthic  $\delta^{13}\text{C}$  covering the periods 130–118 and 8–2 ka BP. From these two sets of data, we select data pertaining to the LIG and compare it to data from the Holocene. Our database only includes measurements on *Cibicides wuellerstorfi* as no significant fractionation between the calcite shells and the surrounding DIC has been measured in this species (Belanger et al., 1981; Duplessy et al., 1984; Zahn et al., 1986).

Our compilation is predominantly based on Lisiecki and Stern (2016) (53 cores) but includes 14 cores described in Oliver et al. (2010), as well as a few other records (CH69-K09, Labeyrie et al., 2017; MD03-2664, Galaasen et al., 2014a; MD95-2042, Martrat et al., 2007a; ODP 1063, Deaney et al., 2017; and U1304, Hodell and Channell, 2016). The full core lists are provided in Tables 1 and 2 for the LIG and the Holocene, respectively.

### 2.2 Age models

Due to the lack of absolute age markers, such as tephra layers, the LIG age models mostly rely on alignment strategies that tie each record to a well-dated reference record. The age model tie points used in this study are taken from the original age model publications. The reference records (LS16; Lisiecki and Stern, 2016) consist of eight regional stacks (one for the intermediate and one for the deep ocean each for the North Atlantic, South Atlantic, Pacific, and Indian oceans) of benthic  $\delta^{18}\text{O}$  that were dated through alignment with other climatic archives such as ice-rafted debris records, synthetic ice core records, and speleothems. The use of regional stacks, rather than a single global stack, improved stratigraphic alignment targets and provided more robust age models. The estimated age model uncertainty ( $2\sigma$ ) for this group of cores is 2 kyr. Please refer to Lisiecki and Stern (2016) for further details. Oliver et al. (2010) defined their age tie points assuming that sea level minima and benthic  $\delta^{18}\text{O}$  maxima are synchronous. The benthic  $\delta^{18}\text{O}$  records

**Table 1.** List of cores for the last interglacial period (LIG). Provided is the core name (“Core”), latitude (“Lat”, °), longitude (“Long”, °), depth (“Dep”, m), the region, and the reference. Regions are abbreviated as follows: NEA: northeastern Atlantic; SWA: southwestern Atlantic; SEA: southeastern Atlantic; SA: southern Atlantic; NP: northern Pacific; SP: southern Pacific; I: Indian. Reference abbreviations are as follows: BW96: Bickert and Wefer (1996); CL82: Curry and Lohmann (1982); dA03: de Abreu et al. (2003); KJ8994: Keigwin and Jones (1989, 1994); KS02: Keigwin and Schlegel (2002); L99: Labeyrie et al. (1999); MB99: Mackensen and Bickert (1999); OH00: Oppo and Horowitz (2000); SH84: Shackleton and Hall (1984); SS0405: Skinner and Shackleton (2004, 2005); VH02: Venz and Hodell (2002); V99: Venz et al. (1999); ZM1011: Zarriss and Mackensen (2010, 2011).

Core	Lat	Long	Dep (m)	Region	Reference	Core	Lat	Long	Dep (m)	Region	Reference
ODP758	5.38	90.36	2935	I	Chen et al. (1995)	SU90-39	52.5	-22	3955	NEA	Cortijo (2003)
RC12-339	9.13	90.03	3010	I	Members (2006)	ODP983	60.4	-23.64	1984	NEA	McInyre et al. (1999)
GEOB3004-1	14.61	52.92	1803	I	Schmiedl and Mackensen (2006)	SU90-03	40.05	-32	2475	NEA	Chapman and Shackleton (1999)
MD01-2378	-13.08	121.79	1783	I	Holbourn et al. (2005)	UI308	49.88	-24.24	3883	NEA	Hodell et al. (2008)
Y69-71	0.1	-95.65	2740	NP	Lyle et al. (2002)	ODP980	55.49	-14.7	2168	NEA	McManus et al. (1999), Oppo et al. (1998)
ODP677	1.2	-83.73	3450	NP	SH84 Shackleton et al. (1990)	ODP982	57.51	-15.85	1134	NEA	Jansen et al. (1996), V99, VH02
ODP849	0.18	-110.52	3839	NP	Shackleton et al. (1990)	EW9209-IJPC	5.91	-44.2	4056	NWA	Curry and Oppo (1997)
V24-109	0.43	158.8	2367	NP	Duplessy et al. (1984)	GEOB4403-2	6.13	-43.44	4503	NWA	Bickert and Mackensen (2003)
Y69-106	2.98	-86.55	2870	NP	Lyle et al. (2002), Pistas and Mix (1997)	ODP1063	33.68	-57.62	4584	NWA	Deaney et al. (2017)
ODR807A	3.61	156.63	2804	NP	Zhang et al. (2007)	CH69-K9	41.75	-47.35	4100	NWA	L99, Waelbroeck et al. (2001)
GK17961-2	8.51	112.33	1795	NP	Wang et al. (1999)	SU90-11	44.07	-40.02	3645	NWA	Jullien et al. (2006), Labeyrie et al. (1995)
MD97-2151	8.73	109.87	1598	NP	Lee et al. (1999), Wei et al. (2006)	UI304	53.06	-33.53	3065	NWA	Hodell and Channell (2016)
ODP143	9.36	113.29	2772	NP	Cheng et al. (2004)	V27-20	54.0	-46.2	3510	NWA	Ruddiman and Members (1982)
V28-304	28.53	134.13	2942	NP	Duplessy et al. (1984)	MD03_2664	57.44	-48.61	3442	NWA	Galahsen et al. (2014a)
V32-128	36.47	177.17	2923	NP	Duplessy et al. (1984)	ODP925	4.2	-43.49	3040	NWA	Bickert et al. (1997)
PS2495	-41.28	-14.49	3134	SA	Mackensen et al. (2001)	ODP926	3.72	-42.91	3598	NWA	Curry et al. (1995)
ODP1089	-40.94	9.89	4621	SA	Hodell et al. (2001)	ODP928	5.46	-43.75	4012	NWA	Bickert et al. (1997)
PS2082	-43.22	11.74	4610	SA	McCorkle and Holder (2001)	V28-127	11.65	-80.13	3237	NWA	Oppo and Fairbanks (1987)
MD06-3018	-23	166.15	2470	SP	Russon et al. (2009)	KNR140-57JPC	31.41	-75.26	3000	NWA	Curry and Oppo (2005), KS02
RC13-110	-0.1	-95.65	3231	SP	Mix et al. (1991)	GEOB3801-6	-29.51	-8.31	4546	SEA	Bickert and Mackensen (2003)
ODR846	-3.1	-90.82	3296	SP	Shackleton et al. (1995)	GEOB1214	-24.69	7.24	3210	SEA	BW96
V19-27	-0.47	-82.07	1373	SP	Mix et al. (1991)	GEOB1710	-23.43	11.7	2987	SEA	BW96
GEOB1101	1.66	-10.98	4588	NEA	Zahn et al. (1986)	GEOB1034	-21.74	5.42	3772	SEA	Schmiedl and Mackensen (1997)
GK16402	14.42	-20.54	4202	NEA	Sarnthein et al. (1994)	GEOB1035	-21.59	5.03	4453	SEA	BW96
GK12392-1	25.17	-16.85	2573	NEA	Shackleton (1977), Zahn et al. (1986)	V22-174	-10.07	-12.82	2630	SEA	Bickert and Mackensen (2003)
GK16004	29.98	-10.65	1512	NEA	Sarnthein et al. (1994)	GEOB1112	-5.78	-10.75	3125	SEA	BW96, MB99
GEOB4216	30.63	-12.4	2324	NEA	Freudenthal et al. (2002)	GEOB1115	-3.56	-12.56	2945	SEA	BW96, MB99
GK15669	34.89	-7.82	2022	NEA	Sarnthein et al. (1994)	GEOB1041	-3.48	-7.6	4033	SEA	BW96, MB99
GK15612-2	44.36	-26.54	3050	NEA	Sarnthein et al. (1994)	GK16867	-2.2	5.1	3891	SEA	Sarnthein et al. (1994)
NO79-28	45.63	-22.75	3625	NEA	Duplessy (1996)	GEOB1105	-1.67	-12.43	3225	SEA	BW96, MB99
GK23416-4	51.57	-20.0	3616	NEA	Sarnthein et al. (1994)	GK16772-1	-1.34	-11.97	3911	SEA	Sarnthein et al. (1994)
NEAP18K	52.77	-30.35	3275	NEA	Chapman and Shackleton (1999)	V29-135	-19.7	8.88	2675	SEA	Bickert and Mackensen (2003)
GK23415-9	53.18	-19.15	2472	NEA	CL82, Sarnthein et al. (1994)	RC13-228	-22.33	11.2	3204	SEA	Lynch-Stiegitz et al. (2006)
GK23414-9	53.54	-20.29	2196	NEA	Sarnthein et al. (1994)	ODP1087	-31.46	15.31	1372	SEA	Ryan et al. (1999)
CH73-139	54.63	-16.35	2209	NEA	Curry et al. (1988), Sarnthein et al. (1994)	MD96-2080	-36.27	19.48	2488	SEA	Vidal et al. (2002)
GK17049-6	55.26	-26.73	3331	NEA	Sarnthein et al. (1982), Sarnthein et al. (1994)	GEOB2109-1	-27.91	-45.88	2504	SWA	Ruddiman and Members (1982)
V28-56	68.03	-6.12	2941	NEA	Ruddiman and Members (1982)	V22-38	-9.55	-34.25	3797	SWA	BW96, MB99
ODP984	61	-24	1650	NEA	Raymo et al. (2004)	GEOB1117	-3.82	-14.9	3984	SWA	BW96, MB99
V29-202	61	-21	2658	NEA	Oppe and Lehman (1995)	GEOB1118	-3.56	-16.43	4675	SWA	BW96, MB99
ODP664	0.11	-23.23	3806	NEA	Raymo et al. (1997)						

**Table 2.** List of cores for the Holocene. Provided is the core name (“Core”), latitude (“Lat”, °), longitude (“Long”, °), depth (“Dep”, m) the region, and the reference. Regions are abbreviated as follows: NEA: northeastern Atlantic; NWA: northwestern Atlantic; SWA: southwestern Atlantic; SEA: southeastern Atlantic; SA: southern Atlantic; NP: northern Pacific; SP: southern Pacific; I: Indian. Reference abbreviations are as follows: BW96: Bickert and Wefer (1996); CL82: Curry and Lohmann (1982); dA03: de Abreu et al. (2003); KJ8994: Keigwin and Jones (1989, 1994); KS02: Keigwin and Schlegel (2002); L99: Labeyrie et al. (1999); MB99: Mackensen and Bickert (1999); OH00: Oppo and Horowitz (2000); SH84: Shackleton and Hall (1984); SS0405: Skinner and Shackleton (2004, 2005); VH02: Venz and Hodell (2002); V99: Venz et al. (1999); ZM1011: Zarriss and Mackensen (2010, 2011).

Core	Lat	Long	Dep (m)	Region	Reference	Core	Lat	Long	Dep (m)	Region	Reference
ODP758	5.38	90.36	2935	I	Chen et al. (1995)	GIK23419	54.96	-19.76	1487	NEA	Sarnthein et al. (1994)
GEOB3004-1	14.61	52.92	1803	I	Schmiedl and Mackensen (2006)	GIK17049-6	55.26	-26.73	3331	NEA	Sarnthein et al. (1994)
M5-3A-422	24.39	58.04	2732	I	Sirocko et al. (2000)	DSDP552	56.04	-23.22	2311	NEA	SH84
MD01-2378	-13.08	121.79	1783	I	Holbourn et al. (2005)	GIK17051	56.16	-31.99	2295	NEA	Sarnthein et al. (1994)
MD79-254	-17.53	38.4	1934	I	Curry et al. (1988)	GIK23519	64.8	-29.6	1893	NEA	Millo et al. (2006)
RC11-120	-43.52	79.87	3193	I	CL82	ODP984	61	-24	1650	NEA	Raymo et al. (2004)
MD88-770	-46.02	96.46	3290	I	Labeyrie et al. (1996), Sowers et al. (1993)	V29-202	61	-21	2658	NEA	Oppo and Lehman (1995)
V35-5	7.2	112.08	1953	NP	Wang et al. (1999)	MD95-2042	37.8	-10.17	3146	NEA	Martrat et al. (2007a)
V24-109	0.43	158.8	2367	NP	Duplessy et al. (1984)	SU90-39	52.5	-22	3955	NEA	Cortijo (2003)
Y69-106	2.98	-86.55	2870	NP	Lyle et al. (2002), Pistias and Mix (1997)	ODP983	60.4	-23.64	1984	NEA	McIntyre et al. (1999)
ODP807A	3.61	156.63	2804	NP	Zhang et al. (2007)	V22-197	14.17	-18.58	3167	NEA	Sarnthein et al. (1994)
GIK17964-2	6.16	112.21	1556	NP	Wang et al. (1999)	ODP659	18.08	-21.03	3082	NEA	Sarnthein et al. (1994)
GIK17961-2	8.51	112.33	1795	NP	Wang et al. (1999)	V30-49	18.43	-21.08	3093	NEA	Mix and Fairbanks (1985)
MD97-2151	8.73	109.87	1598	NP	Lee et al. (1999), Wei et al. (2006)	MD03-2698	38.24	-10.39	4602	NEA	Lebreiro et al. (2009)
GIK17940-2	20.12	117.38	1727	NP	Wang et al. (1999)	SU90-03	40.05	-32	2475	NEA	Chapman and Shackleton (1999)
V28-304	28.53	134.13	2942	NP	Duplessy et al. (1984)	V23-81	54.25	-16.83	2393	NEA	Sarnthein et al. (1994)
EW9504-05	32.48	-118.13	1818	NP	Stott et al. (2000)	NA87-22	55.48	-14.68	2161	NEA	Sarnthein et al. (1994)
MD02-2489	54.39	-148.92	3640	NP	Gebhardt et al. (2008)	ODP980	55.49	-14.7	2168	NEA	Sarnthein et al. (1994)
ODP1090	-42.91	8.9	3702	SA	Hodell et al. (2000), Hodell et al. (2003)	ODP982	57.51	-15.85	1134	NEA	Jansen et al. (1996), V99, VH02
ODP1089	-40.94	9.89	4621	SA	Hodell et al. (2001)	V28-14	64.78	-29.57	1855	NEA	Duplessy et al. (1984)
PS2082	-43.22	11.74	4610	SA	McCorle and Holder (2001)	KNR110-50	4.87	-43.21	3995	NWA	Curry et al. (1988)
MD07-3076	-44.07	-14.21	3770	SA	Waelbroeck et al. (2011)	KNR110-55	4.95	-42.89	4556	NWA	Curry et al. (1988)
MD06-3018	-23	166.15	2470	SP	Russon et al. (2009)	EW9209-1JPC	5.91	-44.2	4056	NWA	Curry and Oppo (1997)
RC13-110	-0.1	-95.65	3231	SP	Mix et al. (1991)	GEOB4403-2	6.13	-43.44	4503	NWA	Bickert and Mackensen (2003)
ODP846	-3.1	-90.82	3296	SP	Shackleton et al. (1995)	KNR31-GPC5	33.69	-57.63	4583	NWA	KJ8994, Keigwin et al. (1991)
V19-27	-0.47	-82.07	1373	SP	Mix et al. (1991)	CH69-K9	41.75	-47.35	4100	NWA	L99, Waelbroeck et al. (2001)
H214	-36.92	177.43	2045	SP	Sikes et al. (2009)	U1304	53.06	-33.53	3065	NWA	Hodell and Channell (2016)
RS147-07	-45.15	146.28	3300	SP	Pahnke and Zahn (2005)	ODP925	4.2	-43.49	3040	NWA	Bickert et al. (1997)
MD97-2120	-45.53	174.93	1210	SP	Pahnke and Zahn (2005)	V25-59	1.37	-33.48	3824	NWA	Mix and Fairbanks (1985)
GEOB1101	1.66	-10.98	4588	NEA	BW96	ODP926	3.72	-42.91	3598	NWA	Curry et al. (1995)
EN066-29	2.46	-19.76	5105	NEA	Sarnthein et al. (1994)	KNR110-75	4.34	-43.41	3063	NWA	Curry et al. (1988)
EN066-32	2.47	-19.73	4998	NEA	Sarnthein et al. (1994)	KNR110-82	4.34	-43.49	2816	NWA	Curry et al. (1988)
EN066-26	3.09	-20.02	4745	NEA	Sarnthein et al. (1994)	KNR110-71	4.36	-43.7	3164	NWA	Curry et al. (1988)
EN066-21	4.23	-20.63	3792	NEA	Sarnthein et al. (1994)	KNR110-66	4.56	-43.38	3547	NWA	CL82, Curry et al. (1988)
EN066-36	4.31	-20.21	4095	NEA	Boyle (1992)	KNR110-91	4.76	-43.31	3810	NWA	Curry et al. (1988)
EN066-38	4.92	-20.5	2937	NEA	Sarnthein et al. (1994)	KNR110-58	4.79	-43.04	4341	NWA	Curry et al. (1988)
EN066-44	5.26	-21.71	3423	NEA	Sarnthein et al. (1994)	ODP927	5.46	-44.48	3326	NWA	Bickert et al. (1997)
EN066-16	5.45	-21.14	3160	NEA	Boyle (1992)	ODP928	5.46	-43.75	4012	NWA	Bickert et al. (1997)
GIK13519-1	5.67	-19.85	2862	NEA	Zahn et al. (1986)	V28-127	5.98	-43.74	4369	NWA	Bickert et al. (1997)
EN066-10	6.64	-21.9	3527	NEA	Sarnthein et al. (1994)	ODP929	5.46	-43.75	4012	NWA	Bickert et al. (1997)
GEOB9526	12.44	-18.06	3223	NEA	ZM1011, Zarriss et al. (2011)	M35003-4	11.65	-80.13	3237	NWA	Oppo and Fairbanks (1987)
GIK16402	14.42	-20.54	4202	NEA	Sarnthein et al. (1994)	KNR140-37JPC	12.09	-61.24	1299	NWA	Hüls (1999), Zahn and Stüber (2002)
GEOB9508-5	14.5	-17.95	2384	NEA	Multiza et al. (2008)	V26-176	31.41	-75.26	3000	NWA	Curry and Oppo (2005), KS02
							36.05	-72.38	3942	NWA	Curry et al. (1988)

Table 2. Continued.

Core	Lat	Long	Dep (m)	Region	Reference	Core	Lat	Long	Dep (m)	Region	Reference
GIK12347-2	15.83	-17.86	2576	NEA	Sarnthein et al. (1994)	GEOB1214	-24.69	7.24	3210	SEA	BW96
GEOB7920-2	20.75	-18.58	2278	NEA	Collins et al. (2011), Tjallingii et al. (2008)	GEOB1211	-24.48	7.53	4084	SEA	BW96
GIK12328-5	21.15	-18.57	2778	NEA	Sarnthein et al. (1994)	GEOB1710	-23.43	11.7	2987	SEA	Schmiedl and Mackensen (1997)
GIK16030	21.24	-18.06	1516	NEA	Sarnthein et al. (1994)	GEOB1032	-22.92	6.04	2505	SEA	BW96, Bickert et al. (2003)
GIK12379-3	23.14	-17.75	2136	NEA	Sarnthein et al. (1994)	GEOB1034	-21.74	5.42	3772	SEA	BW96
GIK12392-1	25.17	-16.85	2573	NEA	Shackleton (1977), Zahn et al. (1986)	GEOB1035	-21.59	5.03	4453	SEA	BW96
GEOB4240	28.89	-13.23	1358	NEA	Freudenthal et al. (2002)	GEOB1028-5	-20.1	9.19	2209	SEA	Bickert and Mackensen (2003)
GIK16004	29.98	-12.4	1512	NEA	Sarnthein et al. (1994)	GEOB1112	-5.78	10.75	3125	SEA	BW96, MB99
GEOB4216	30.63	-10.65	2324	NEA	Freudenthal et al. (2002)	GEOB1115	-4.0	10.0	1000	SEA	Sarnthein et al. (1994)
GIK15672	34.86	-8.13	2460	NEA	CL82, Sarnthein et al. (1994)	BT4	-3.56	-12.56	2945	SEA	BW96, MB99
GIK15669	34.89	-7.82	2022	NEA	Sarnthein et al. (1994)	GEOB1041	-3.48	-7.6	4033	SEA	BW96, MB99
GIK11944-2	35.65	-8.06	1765	NEA	Sarnthein et al. (1994)	GIK16867	-2.2	5.1	3891	SEA	Sarnthein et al. (1994)
KP13	37.58	-31.84	2690	NEA	Curry et al. (1988)	GEOB1105	-1.67	-12.43	3225	SEA	BW96, MB99
MD99-2334	37.8	-10.17	3146	NEA	Skinner et al. (2003), SS0405	V29-135	-19.7	8.88	2675	SEA	Sarnthein et al. (1994)
MD95-2040	40.58	-9.86	2465	NEA	DA03, Schönfeld et al. (2003)	RC13-229	-25.5	11.3	4194	SEA	Sarnthein et al. (1994)
CHN82-24	41.72	-32.85	3427	NEA	Boyle and Keigwin (1985)	RC13-228	-22.33	11.2	3204	SEA	Bickert and Mackensen (2003)
GIK15612-2	44.36	-26.54	3050	NEA	Sarnthein et al. (1994)	ODP1087	-31.46	15.31	1372	SEA	Lynch-Stiegler et al. (2006)
NO79-28	45.63	-22.75	3625	NEA	Dupleissy (1996)	MD96-2080	-36.27	19.48	2488	SEA	Rau et al. (2002)
GIK17055-1	48.22	-27.06	2558	NEA	Sarnthein et al. (1994)	GEOB2109-1	-27.91	-45.88	2504	SWA	Vidal et al. (1999)
UI308	49.88	-24.24	3883	NEA	Hodell et al. (2008)	KNR159-36	-27.51	-46.47	1268	SWA	Came et al. (2003), OH00
GIK23417-1	50.67	-19.43	3850	NEA	Sarnthein et al. (1994)	GEOB1117	-3.82	-14.9	3984	SWA	BW96, MB99
GIK23416-4	51.57	-20.0	3616	NEA	Sarnthein et al. (1994)	GEOB1118	-3.56	-16.43	4675	SWA	BW96, MB99
GIK23418-8	52.55	-20.33	2841	NEA	Sarnthein et al. (1994)	RC16-84	-26.7	-43.33	2438	SWA	OH00
GIK23415-9	53.18	-19.15	2472	NEA	CL82, Sarnthein et al. (1994)	RC16-119	-27.7	-46.52	1567	SWA	OH00
GIK23414-9	53.54	-20.29	2196	NEA	Sarnthein et al. (1994)	V24-253	-26.95	-44.67	2069	SWA	OH00

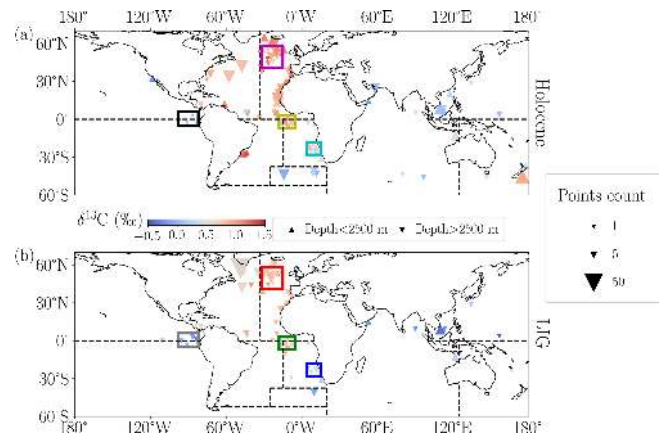
were aligned with each other and then tied to the Dome Fuji chronology (based on  $\text{O}_2/\text{N}_2$ ) (Kawamura et al., 2007). Please refer to Shackleton et al. (2000) and Oliver et al. (2010) for an extensive method description. The age model uncertainty on this group of cores is estimated to range from 1 to 2.5 kyr.

The published age models for the additional cores were determined using similar alignment techniques: SSTs were correlated to the NGRIP Greenland ice core for CH69-K09 and MD95-2042 (Govin et al., 2012). The age model for MD03-2664 was determined by correlating MD03-2664  $\delta^{18}\text{O}$  with previously dated MD95-2042  $\delta^{18}\text{O}$  (Galaasen et al., 2014b). ODP 1063 and U1304  $\delta^{18}\text{O}$  were originally aligned to the LR04 stack (Lisiecki and Raymo, 2005). In order to align all of the records, adjustments to the age models of cores from Oliver et al. (2010) and the five additional cores (CH69-K09, MD95-2042, MD03-2664, ODP 1063, and U1304) were made by aligning the  $\delta^{18}\text{O}$  minima during the LIG to the corresponding  $\delta^{18}\text{O}$  minima of the nearest LS16 stack. The  $\delta^{18}\text{O}$  data before and after the alignment is given in Fig. S1 in the Supplement.

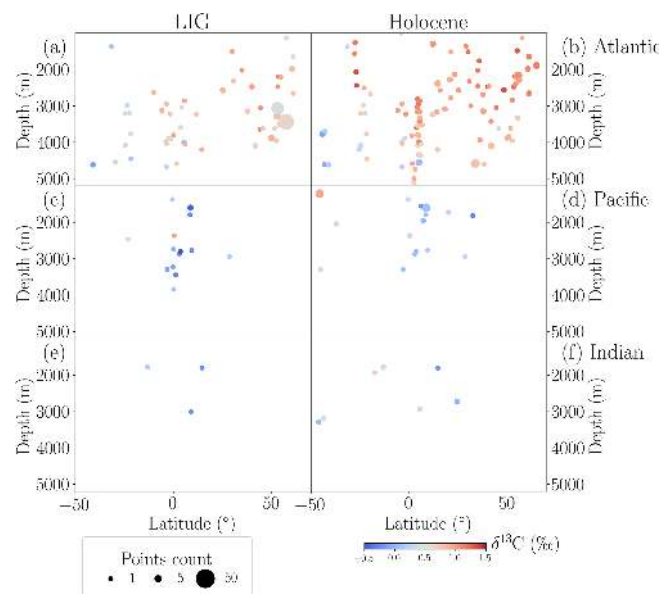
The Holocene age models are based on planktonic foraminifera radiocarbon dates (Stern and Lisiecki, 2014; Waelbroeck et al., 2001) that have been converted into calendar ages using IntCal13 and using reservoir ages based on modern observations (Key et al., 2004), which are assumed to have remained fairly stable across the Holocene. The age uncertainty associated with these Holocene radiocarbon-based age models is generally less than 0.5 kyr. However, it is important to note that Holocene age models from Oliver et al. (2010) were derived using the same method as their LIG age models, leading to larger age uncertainties of about 1–2.5 kyr for this set of Holocene records (four cores). The tie points were used to derive a full age–depth model assuming a constant sedimentation rate between tie points (i.e. linear interpolation).

### 2.3 Spatial coverage

The spatial distribution of the database for the Holocene and the LIG is shown in Fig. 2, and the depth distribution in each ocean basin is shown in Fig. 3. There are more data in the Atlantic Ocean (65 LIG, 118 Holocene) than in the Pacific (15 LIG, 19 Holocene) and Indian (3 LIG, 7 Holocene) oceans. We used this database to determine (1) if there is a significant difference in the average ocean  $\delta^{13}\text{C}$  signal at the LIG compared to the Holocene and (2) if ocean circulation patterns were comparable. Due to the sparsity of data in the Indian and Pacific oceans, our investigation is primarily focused on the Atlantic. Additionally, the temporal uncertainties ( $\sim 2$  kyr) do not permit an investigation of centennial-scale events, and therefore we restrict our analysis to mean LIG and Holocene conditions.



**Figure 2.** Global distribution of benthic foraminifera  $\delta^{13}\text{C}$  covering the periods studied here: the Holocene (7–2 ka BP) (a) and LIG (125–120 ka BP) (b). Symbol size indicates the number of values per core, colour indicates average  $\delta^{13}\text{C}$ , and the triangle direction indicates the proxy depth (upward-pointing triangle: between 1000 and 2500 m depth; downward-pointing triangle: deeper than 2500 m). Four specific regions used in Sect. 3.1 are outlined: eastern equatorial Pacific (black, grey), equatorial Atlantic (yellow, green), southeastern Atlantic (cyan, blue), and northeastern Atlantic (magenta, red). Regional boundaries used to calculate the global volume-weighted mean  $\delta^{13}\text{C}$  (Sect. 3.2) are indicated by dotted black lines as defined in Peterson et al. (2014).



**Figure 3.** Zonal distribution of benthic foraminifera  $\delta^{13}\text{C}$  (‰) during the LIG (125–120 ka BP; a, c, e) and the Holocene (7–2 ka BP; b, d, f) in the Atlantic Ocean (a, b), Pacific Ocean (c, d), and Indian Ocean (e, f). Symbol size indicates the number of measurements per core and colour indicates average  $\delta^{13}\text{C}$ .

### 3 Results

The  $\delta^{13}\text{C}$  signal varies significantly regionally and with depth. The highest average  $\delta^{13}\text{C}$  values are associated with NADW and are generally found at depths between  $\sim 1500$  and  $3000$  m in the North Atlantic, with organic matter remineralisation and mixing with southern source waters leading to a  $\delta^{13}\text{C}$  decrease along the NADW path. The lowest  $\delta^{13}\text{C}$  values are in the deep South Atlantic ( $> 4000$  m) because the Antarctic Bottom Water (AABW) end member is much lower than its NADW counterpart. Since the Indian and Pacific oceans are mostly ventilated from southern-sourced water masses,  $\delta^{13}\text{C}$  generally decreases northward in these two basins.

Since the number of cores is not consistent across the two time periods and given the high regional variability observed in  $\delta^{13}\text{C}$ , it is not possible to simply average all available data to determine the global mean  $\delta^{13}\text{C}$ . Furthermore, the spatial heterogeneity of the data density adds to the complexity of the problem. To address these points, we first analyse differences between the LIG and Holocene records for pre-defined small regions with high data density. We then calculate regional volume-weighted  $\delta^{13}\text{C}$  means for larger regions from which we estimate the global LIG–Holocene anomaly.

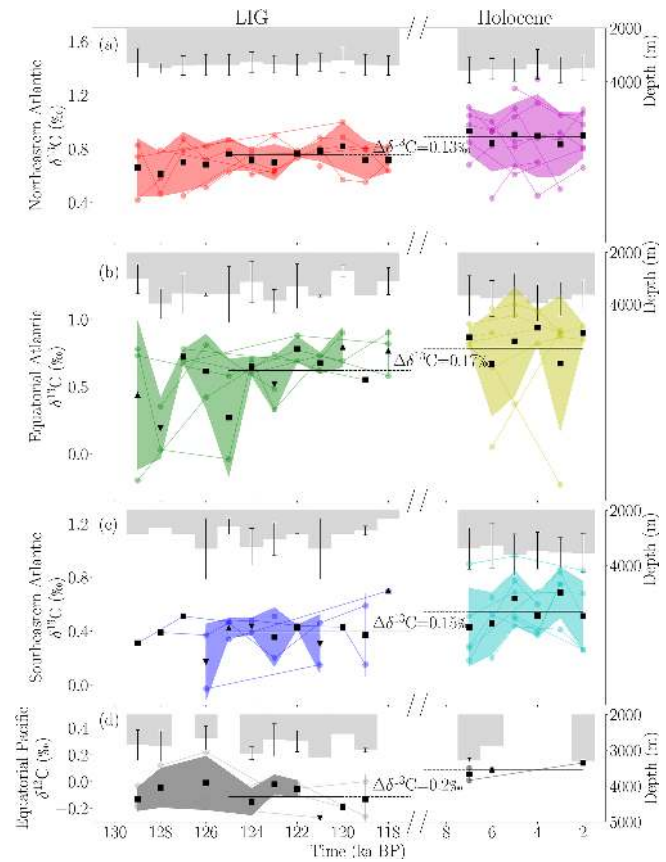
#### 3.1 Regional reconstruction of $\delta^{13}\text{C}$

We define regions with high densities of cores to reconstruct regional mean  $\delta^{13}\text{C}$  (Fig. 2). These regions need to be small enough to assume reasonably small spatial variability in the  $\delta^{13}\text{C}$  signal and yet still have enough data to establish a reliable statistical difference between the two time periods.

Based on these requirements, four regions are selected: the northeastern Atlantic, the equatorial Atlantic, a region off the Namibian Coast (southeastern Atlantic), and a region around the Galapagos Islands (eastern equatorial Pacific). The boundaries of each region are defined in Table 3.

We then define the time periods within the LIG and the Holocene to perform our analyses. For the Holocene, as most of the available data are dated prior to 2 ka BP, we define the end of our Holocene time period as 2 ka BP. To capture as much of the Holocene data as possible, we include data back to 7 ka BP, ensuring that we do not include instability associated with the 8.2 ka BP event (Alley and Ágústsdóttir, 2005; Thomas et al., 2007). This provides a time span of 5 kyr of data that we will consider for our analysis of the Holocene.

For the LIG, we seek to avoid data associated with the end of the penultimate deglaciation, which is characterised by a benthic  $\delta^{13}\text{C}$  increase in the Atlantic until  $\sim 128$  ka BP (Govin et al., 2015; Menviel et al., 2019; Oliver et al., 2010, Fig. 4). In addition, a millennial-scale event has been identified in the North Atlantic between  $\sim 127$  and 126 ka BP (Galaasen et al., 2014b; Tzedakis et al., 2018). Considering the typical dating uncertainties associated with the LIG data (2 kyr), we thus decide to start our LIG time period at



**Figure 4.** Benthic foraminifera  $\delta^{13}\text{C}$  (left y axis, ‰) during the LIG (left) and Holocene (right) for four defined regions: northeastern Atlantic (a), equatorial Atlantic (b), southeastern Atlantic (c), and eastern equatorial Pacific (d). Data are presented in discrete time slices spanning 1 kyr. Only cores deeper than 2500 m are shown. Circular, coloured points connected by lines show each average  $\delta^{13}\text{C}$  value per core per time slice. Black symbols represent  $\delta^{13}\text{C}$  averages per slice. Each slice has a corresponding averaged depth (right y axis, m), with 1 standard deviation on either side shown on the bars. Slices with an average depth within  $\pm 300$  m of the mean core depth of all slices are represented with a square point. Slices with an average depth 300 m shallower than the mean are shown with an upward triangle, and slices that are than 300 m deeper than the mean are shown with a downward triangle. Shading shows 1 standard deviation on either side of the mean for slices where more than one point exists.

125 ka BP. To ensure that the two time periods are of the same length (5 kyr), we define the LIG period for our analysis to be 125–120 ka BP. We note that our definition should also avoid data associated with the glacial inception (Govin et al., 2015; Past Interglacials Working Group of PAGES, 2016). We verify that the LIG time period has sufficient data across the selected four regions, noting that the highest density of data falls within the 125–120 ka BP time period – particularly in the equatorial Atlantic and southeastern Atlantic (Fig. 4b, c).



**Table 3.** Regional summary of  $\delta^{13}\text{C}$  below 2500 m depth for the LIG (125–120 ka BP) and Holocene (7–2 ka BP) using a single value per core for each time slice. Shown are the non-volume-weighted means ( $\delta^{13}\text{C}$ , ‰), standard deviations ( $\sigma$ , ‰), and counts ( $N$ ) for both time periods, along with the time period regional anomalies ( $\Delta\delta^{13}\text{C}$ , ‰), propagated standard deviations for the anomaly ( $\sigma$ , ‰), and  $p$  values from two-sample  $t$  tests between the two time periods.

Region	Latitude	Longitude	Holocene			LIG			LIG-Holocene		
			$\delta^{13}\text{C}$ (‰)	$\sigma$ (‰)	$N$	$\delta^{13}\text{C}$ (‰)	$\sigma$ (‰)	$N$	$\Delta\delta^{13}\text{C}$ (‰)	$\sigma$ (‰)	$P$ value
Northeastern Atlantic	41–58° N	32–15° E	0.89	0.21	34	0.76	0.11	23	−0.13	0.12	0.0096
Equatorial Atlantic	7° S–3° N	18–5° E	0.79	0.32	22	0.62	0.23	14	−0.17	0.20	0.1110
Southeastern Atlantic	28–18° S	4–15° W	0.55	0.22	27	0.40	0.11	12	−0.15	0.12	0.0361
Equatorial Pacific	5° S–6° N	98–82° E	0.09	0.05	4	−0.11	0.10	8	−0.20	0.06	0.0056

We round the data to the nearest 1 kyr, find an average per 1 kyr, and refer to this as a time slice. We consider the LIG–Holocene anomaly across these two time periods for the four regions selected, and qualitatively consider the influence of changes in the average depth in which the proxies were recorded, as indicated by the direction of the black triangles in Fig. 4.

The average  $\delta^{13}\text{C}$  anomaly between the LIG and Holocene periods for cores deeper than 2500 m is consistent across the different regions despite their geographic separation, suggesting a significantly lower  $\delta^{13}\text{C}$  during the LIG than the Holocene, with differences ranging from  $-0.13$  ‰ in the northeastern Atlantic to  $-0.20$  ‰ in the equatorial Pacific (Table 3). The statistical significance between the two time periods is established using a two-tailed  $t$  test on data of one mean value per core and spans all time slices (125–120 and 7–2 ka BP). The  $t$  test shows that there is a statistically significant difference everywhere except in the equatorial Atlantic, with confidence intervals varying from 0.13 in the equatorial Atlantic to 0.04 in the northeastern Atlantic. When using a single tail  $t$  test instead, the difference becomes significant in the equatorial Atlantic, giving a new  $p$  value of 0.02. Figure 4 suggests that the difficulty in determining significance in this region for cores deeper than 2500 m might be due to a singular outlier measurement in the equatorial Atlantic: a value of  $-0.23$  ‰ from GeoB1118 at  $\sim 3.5$  ka BP. If this value is excluded, then an anomaly of  $-0.22$  with a  $p$  value less than 0.005 is observed.

We also compare the distribution of  $\delta^{13}\text{C}$  for cores deeper than 2500 m for three overlapping periods within the LIG (early LIG: 128–123 ka BP; LIG: 125–120 ka BP; late LIG: 123–118 ka BP). The results for the four regions are shown in Fig. 5. The statistical characteristics do not show much variation between the LIG and late LIG  $\delta^{13}\text{C}$  distributions. In the equatorial Pacific, the difference between the early LIG and the Holocene is smaller than between the LIG and Holocene, but this is countered with a larger difference in the equatorial Atlantic between the early LIG and Holocene. The spread in the data is generally larger during the Holocene than during the other time periods, which might be due to the

greater number of measurements during the Holocene. The spread of data during the early LIG is slightly larger than during the LIG and late LIG in the equatorial and southeastern Atlantic. The equatorial Atlantic is the only region that displays significantly more points with lower  $\delta^{13}\text{C}$  during the early LIG. Overall, these distributions do not suggest that the LIG–Holocene anomalies that we have determined would be significantly impacted by slight variations in the selected time window. We perform an analysis of variance (ANOVA) on each region and post hoc tests on the data. We find that the Holocene data are significantly different from the three LIG periods in the northeastern Atlantic, the southeastern Atlantic, and the equatorial Pacific, while the three periods within the LIG are not significantly different from each other for any of the regions.

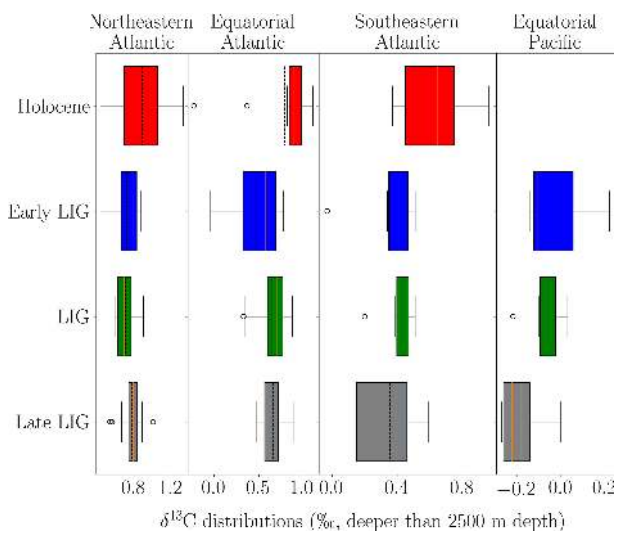
### 3.2 Volume-weighted regional $\delta^{13}\text{C}$

The second approach we use to further constrain the LIG–Holocene  $\delta^{13}\text{C}$  anomaly is to estimate the volume-weighted regional  $\delta^{13}\text{C}$ . We define our regional boundaries based on the regions described in Peterson et al. (2014); however, we only include the regions where there are enough data to justify an analysis. For all the data in each of these regions, we calculate a mean value by taking the direct averages of all data. We divide the ocean basins into eight regions (Table 4, shown in Fig. 2) and calculate the volume-weighted averages  $\delta^{13}\text{C}$  for each of these regions. Since the Atlantic and Pacific oceans have more data than the Indian Ocean, there is greater confidence in the  $\delta^{13}\text{C}$  estimates for these regions. These regional averages are then used to calculate a global volume-weighted  $\delta^{13}\text{C}$ .

Results for the Atlantic and Pacific oceans are given in Fig. 6 and show a mean LIG–Holocene anomaly of  $-0.21$  ‰ and  $-0.27$  ‰, respectively, slightly higher than the range of estimates for the four regions selected in Sect. 3.1. There is a higher offset estimated in this definition of the southwestern Atlantic ( $-0.45$  ‰) than in Sect. 3.1; however, there are only four cores available in this region during the Holocene and two during the LIG.

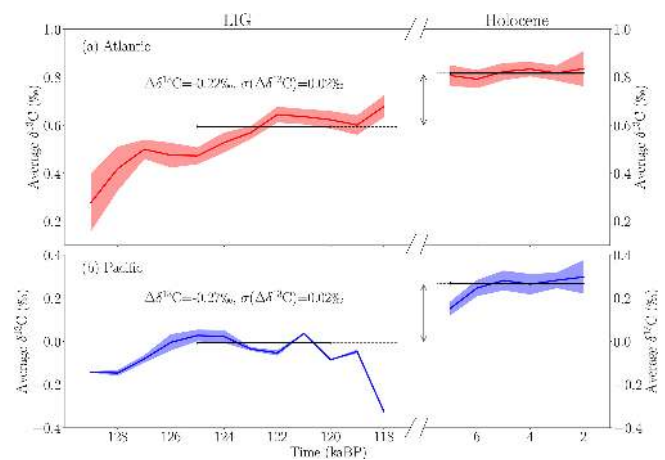
**Table 4.** Regional breakdown of  $\delta^{13}\text{C}$  data for all depths during the Holocene (7–2 ka BP) and LIG (125–120 ka BP) averaged across the 1 kyr time slices. For each region: the average number of data points (labelled as “Points”) and cores per time slice (labelled as “Cores”), the average standard deviation of  $\delta^{13}\text{C}$  per time slice (‰), the mean depth (m) across time slices, and the standard deviation of depth (m) between time slices ( $\sigma_{\text{depth}}$ ). Regions are abbreviated as follows: NEA: northeastern Atlantic; NWA: northwestern Atlantic; SA: southern Atlantic; SEA: southeastern Atlantic; SWA: southwestern Atlantic; I: Indian, NP: northern Pacific; SP: southern Pacific.

Area	Holocene						LIG					
	$\delta^{13}\text{C}$ (‰)	Points	Cores	$\sigma_{\delta^{13}\text{C}}$ (‰)	Mean depth (m)	$\sigma_{\text{depth}}$ (m)	$\delta^{13}\text{C}$ (‰)	Points	Cores	$\sigma_{\delta^{13}\text{C}}$ (‰)	Mean depth (m)	$\sigma_{\text{depth}}$ (m)
NEA	0.94	73	47	0.19	2853	944	0.76	32	22	0.21	2746	789
NWA	0.81	28	13	0.27	3698	867	0.64	41	9	0.27	3679	455
SA	0.08	6	4	0.18	4103	429	−0.14	3	2	0.11	4533	120
SEA	0.63	13	12	0.26	3306	787	0.55	14	14	0.23	3163	799
SWA	0.96	6	4	0.32	2302	929	0.51	2	2	0.12	4156	172
I	0.23	6	4	0.26	2287	529	0.06	4	4	0.19	2347	581
NP	0.03	14	7	0.20	2015	448	−0.10	9	8	0.24	2815	673
SP	0.45	12	5	0.30	2285	924	0.06	3	3	0.15	2724	709



**Figure 5.** Distributions of  $\delta^{13}\text{C}$  for all core measurements deeper than 2500 m during the Holocene (7–2 ka BP, red), the early LIG (128–123 ka BP), the LIG (125–120 ka BP), and the late LIG (123–118 ka BP) across four regions (equatorial Pacific, equatorial Atlantic, southeastern Atlantic, northeastern Atlantic). The lower end of the box indicates quartile 1 (Q1), and the upper end indicates quartile 3 (Q3). Orange vertical lines show the median, and dotted vertical lines show the mean. The whiskers indicate the lower and upper fences of the data calculated as  $Q1 - 1.5 \times (Q3 - Q1)$  and  $Q3 + 1.5 \times (Q3 - Q1)$ , respectively, and the clear circles are outliers.

The estimated LIG–Holocene anomaly in the southern Pacific is relatively high at  $-0.39\text{‰}$ , giving a relatively large Pacific anomaly estimate of  $-0.27\text{‰}$ . This could be due in part to the deeper location of the LIG cores compared to the mean of the Holocene cores (439 m difference, Table 4). There is less confidence in the estimate of the Pa-



**Figure 6.** Comparison of volume-weighted  $\delta^{13}\text{C}$  for the Atlantic (red) and Pacific (blue) for the LIG and Holocene, calculated using the regions from Peterson et al. (2014) from data covering all depths. Solid coloured lines indicate the mean volume-weighted  $\delta^{13}\text{C}$ , and the shading indicates the volume-weighted sum of square deviations from the mean. The horizontal bars indicate the mean of the stable period determined from the regional analysis as defined in Sect. 3.1 (LIG: 125–120 ka BP; Holocene: 7–2 ka BP), with the  $\Delta\delta^{13}\text{C}$  indicating the mean anomaly between these two averages and the standard deviation ( $\sigma(\delta^{13}\text{C})$ , ‰).

cific volume-weighted mean since the proxy data are sparse, and the majority of cores are from the eastern equatorial Pacific as shown in Fig. 2. We also note that the average depths of cores from the Pacific Ocean (LIG: 2711 m; Holocene: 2131 m) and Indian Ocean (LIG: 2383 m; Holocene: 2303 m) are shallower than that of the Atlantic Ocean (LIG: 3531 m; Holocene: 3157 m; Fig. 3). However, as the vertical gradient below 2000 m depth in the Pacific Ocean is small (e.g. Eide et al., 2017), this might not significantly impact our results.

There is a small positive trend in the average Atlantic  $\delta^{13}\text{C}$  from 125 ka BP, reaching a maximum value at 118 ka BP (Fig. 6). The average core depth over the 125–120 ka BP time period does not suggest that a change in the mean depth could explain this variation. Fitting a linear regression over this period indicates an increase in  $\delta^{13}\text{C}$  of  $0.03\text{‰ kyr}^{-1}$  in the Atlantic, with a  $p$  value of 0.01 and an  $R^2$  of 0.85 (Fig. 4a). For the Pacific, there is a  $\sim 0.13\text{‰}$  increase in  $\delta^{13}\text{C}$  between 7 and 5 ka BP, which could be associated with the early Holocene terrestrial regrowth (Menviel and Joos, 2012).

For the Indian Ocean, we only include four cores, as these are the only ones spanning both the LIG and Holocene. An LIG anomaly of  $-0.13\text{‰}$  in the Indian Ocean compared to the Holocene is therefore associated with higher uncertainties. The whole ocean mean LIG  $\delta^{13}\text{C}$  anomaly is  $-0.25\text{‰}$ , but it is associated with higher uncertainties than each region anomaly.

Both the regional analysis of our new database and our volume-weighted estimate indicate that the global mean  $\delta^{13}\text{C}$  was about  $0.2\text{‰}$  lower during the LIG than during the Holocene. We further test the robustness of this result in the next section.

### 3.3 Reconstruction of the LIG Atlantic Meridional Overturning Circulation

In this section, we analyse the spatial  $\delta^{13}\text{C}$  distribution in the Atlantic Ocean to assess potential changes in the penetration depth and southward expansion of NADW during the LIG, defined here as 125–120 ka BP, with respect to the Holocene. A change in NADW might influence our estimate of the mean  $\delta^{13}\text{C}$ , given that most of the available data are localised in the Atlantic Ocean.

We use simple statistical regression models to reconstruct NADW and AABW separately with a quadratic-with-depth and linear-with-latitude equation following the method of Bengtson et al. (2019). For consistency, the regression algorithm only includes records from cores that span both the LIG and Holocene and uses a weighted least-squares approach, where the weighting equals the number of samples per core. The modelled region is defined between  $40^\circ\text{S}$  and  $60^\circ\text{N}$  as this is the region where we can expect to find both the NADW and AABW  $\delta^{13}\text{C}$  signals.

The results are shown in Fig. 7. We test the robustness of our statistical model using the jackknifing technique. We systematically exclude each individual core from the database one at a time, fit the parameters using this modified database, and compare the model prediction against the core, which was excluded. This produces small variations in the average mean response of the statistical models (the standard deviations were  $0.01\text{‰}$  for both the LIG and Holocene, respectively).

We calculate end member values based on proxies located near the water mass sources. These are taken as  $0.79\text{‰}$  and  $1.02\text{‰}$  for NADW for the LIG and Holocene, respec-

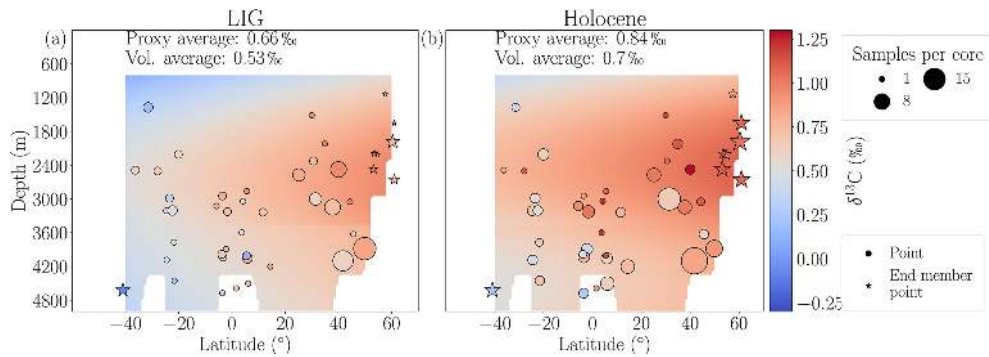
tively, and  $-0.09\text{‰}$  and  $0.23\text{‰}$  for AABW for the LIG and Holocene, respectively. The end member values are calculated as the average of cores shallower than 3000 m but deeper than 1000 m and located between  $50$  and  $70^\circ\text{N}$  for NADW. The NADW end member cores have an average depth of 2043 m and a standard deviation of 478 m during the LIG. For the AABW end member, the only eligible core is ODP1089, which is at  $\sim 41^\circ\text{S}$  and 4621 m.

The mean volume-weighted  $\delta^{13}\text{C}$  for the Atlantic Ocean between  $40^\circ\text{S}$  and  $60^\circ\text{N}$  based on this interpolation is  $0.53\text{‰}$  for the LIG and  $0.70\text{‰}$  for the Holocene (Fig. 7). This suggests a  $0.17\text{‰}$  lower Atlantic  $\delta^{13}\text{C}$  at the LIG than the Holocene. Our statistical reconstruction points to a very similar NADW depth ( $\sim 2600\text{ m}$ ) for both time periods (Fig. 7). The NADW depth is defined here as the depth of maximum  $\delta^{13}\text{C}$  in the North Atlantic.

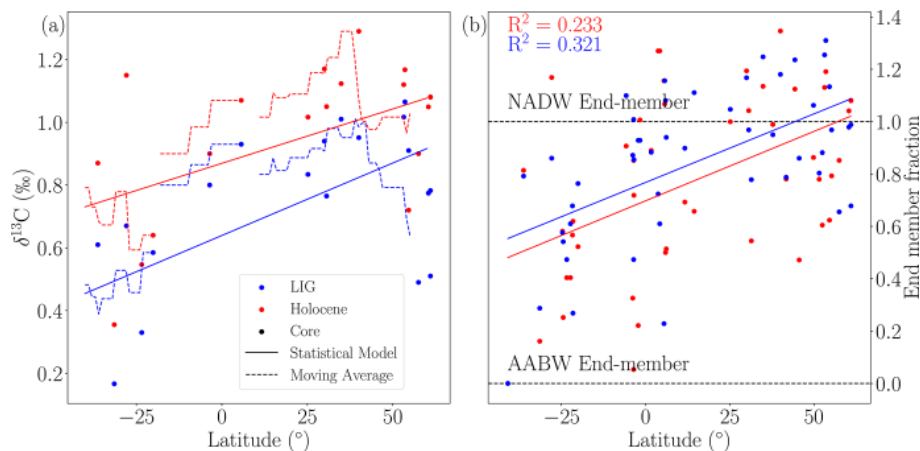
We also investigate the meridional gradient in  $\delta^{13}\text{C}$  in the Atlantic Ocean to determine whether the NADW southward penetration, transport, and remineralisation rates were significantly different during the LIG compared to the Holocene. We only consider cores that are located between depths of 1000 and 3000 m in order to stay within the main pathway of NADW (Fig. 8a). Though there is significant scatter, in accordance with our previous findings, a moving average through the Holocene and the LIG data shows that LIG  $\delta^{13}\text{C}$  is typically lower than the Holocene counterparts. However, the meridional  $\delta^{13}\text{C}$  statistical model gradients are not very different for the LIG ( $0.0036\text{‰ per degree latitude}$ ) and the Holocene ( $0.0030\text{‰ per degree latitude}$ ) (Fig. 8a), suggesting a similar southward penetration of NADW.

Using  $\delta^{13}\text{C}$  of the end members for NADW and AABW, we use a simple binary mixing model for all cores deeper than 1000 m to estimate changes in NADW penetration (Fig. 8b). The LIG and Holocene  $\delta^{13}\text{C}$  slopes in the Atlantic are similar, indicating similar southward penetration of NADW during both time periods. This suggests that the differences in  $\delta^{13}\text{C}$  between the two time periods is most likely due to change in end member values, while the mean Atlantic oceanic circulation was likely similar.

Based on our analysis, there appears to be no significant difference in the mean time-averaged AMOC between the LIG and the Holocene. Negative LIG–Holocene anomalies are found for each of the smaller regions selected (northeastern Atlantic, equatorial Atlantic, southeastern Atlantic, and eastern equatorial Pacific), with statistical significance seen in all regions except the equatorial Atlantic, where an unusual low  $\delta^{13}\text{C}$  value in one core is responsible for narrowing the difference between the two period means. Additionally, our volume-weighted mean  $\delta^{13}\text{C}$  estimates have similar anomalies in the Atlantic and Pacific oceans ( $-0.21\text{‰}$  and  $-0.27\text{‰}$ , respectively, Fig. 6).



**Figure 7.** Reconstructed Atlantic  $\delta^{13}\text{C}$  (‰) meridional section during the LIG (125–120 ka BP) and Holocene (7–2 ka BP). The circular points represent the proxy data, showing the average  $\delta^{13}\text{C}$  with colour and the number of points per core with size. The stars represent the proxy data that make up the end members. Background shading shows the reconstructed  $\delta^{13}\text{C}$  using a quadratic statistical regression of the proxy data following the method described in Bengtson et al. (2019).



**Figure 8.** The meridional gradient of the Atlantic Ocean benthic  $\delta^{13}\text{C}$  (‰). **(a)** Holocene (red) and LIG (blue)  $\delta^{13}\text{C}$  for each core (points) between 1000 and 3000 m. Dotted lines are the moving averages of the cores. Solid lines indicate the results of our statistical model at 2000 m. **(b)** Average  $\delta^{13}\text{C}$  for each record deeper than 1000 m as a proportion of the end members. A value of one indicates the NADW end member, and a value of zero indicates the AABW end member. Solid lines show the linear regressions of the records.

#### 4 Discussion

One of the goals of our study is to assess the mean change in oceanic  $\delta^{13}\text{C}$  between the LIG and the Holocene. Given the uncertainties in the chronologies, avoiding data that pertains to deglaciation, and capturing the same length of time during the LIG and the Holocene, we chose the periods 125 to 120 ka BP for the LIG and 7 to 2 ka BP for the Holocene. Using a similar geographical distribution of data points for both periods, we find that the oceanic  $\delta^{13}\text{C}$  was  $\sim 0.2\text{‰}$  lower during the LIG than the Holocene.

Our analysis of the  $\delta^{13}\text{C}$  signal suggests consistent LIG–Holocene  $\delta^{13}\text{C}$  anomalies in different regions of the Atlantic basins, as well as in the Pacific and Indian oceans, even if there are significant uncertainties with the later due to fewer available records. The  $\delta^{13}\text{C}$  distribution in the Atlantic Ocean suggests that there was no significant mean change in the southward penetration or depth of NADW during the LIG

(125–120 ka BP) compared to the Holocene (7–2 ka BP). A statistical reconstruction of the early LIG (128–123 ka BP)  $\delta^{13}\text{C}$  compared to our 125–120 ka BP reconstruction does not reveal a significant difference in either the NADW core depth or NADW extent, as indicated by the meridional  $\delta^{13}\text{C}$  gradients (Fig. S2). The volume-weighted average  $\delta^{13}\text{C}$  during the early LIG is  $0.06\text{‰}$  lighter than during the LIG period considered here (125–120 ka BP). Since both time slices (128–123 and 125–120 ka BP) are 5 kyr averages and include dating uncertainties of  $\sim 2$  kyr, it is not possible to resolve potential centennial-scale oceanic circulation changes (e.g. Galaasen et al., 2014b; Tzedakis et al., 2018).

Explanations for the  $0.2\text{‰}$  lower  $\delta^{13}\text{C}$  anomaly in the ocean may include a redistribution between the ocean–atmosphere system. Such a redistribution can result from a change in end member values (Fig. 8). As fractionation during air–sea gas exchange is temperature dependent, globally

higher SSTs at the LIG could lead to a lower oceanic  $\delta^{13}\text{C}$ . However, the effect of this is likely small (Brovkin et al., 2002) and would also lead to a higher atmospheric  $\delta^{13}\text{CO}_2$  at the LIG, which is inconsistent with Antarctic ice core measurements that suggest an anomaly of  $-0.3\text{‰}$  (Schneider et al., 2013). Lower nutrient utilisation in the North Atlantic would decrease surface ocean  $\delta^{13}\text{C}$  and thus the  $\delta^{13}\text{C}$  end members. However, this would also imply that less organic carbon would be remineralised at depth. Therefore, it is unlikely that the lower average oceanic mean  $\delta^{13}\text{C}$  results from a change in end members through lower surface ocean nutrient utilisation. Currently, there is still a lack of constraints on nutrient utilisation in these end member regions during the LIG compared to the Holocene. Therefore, the lower  $\delta^{13}\text{C}$  in the ocean–atmosphere system cannot be explained by a simple redistribution of  $\delta^{13}\text{C}$  between the atmosphere and the ocean.

An alternative explanation for the anomaly is a change in the terrestrial carbon storage, which has a typical signature of approximately  $-37\text{‰}$  to  $-20\text{‰}$  for  $\text{C}_3$ -derived plant material (Kohn, 2010) and  $-13\text{‰}$  for  $\text{C}_4$ -derived plant material (Basu et al., 2015). The total land carbon content at the LIG is poorly constrained. Proxies generally suggest extensive vegetation during the LIG compared to the Holocene (CAPE, 2006; Govin et al., 2015; Larrasoana et al., 2013; Muhs et al., 2001; Tarasov et al., 2005; de Vernal and Hillaire-Marcel, 2008), which would imply a greater land carbon store. However, other terrestrial carbon stores including peatlands and permafrost may also have differed during the LIG compared to the Holocene. With an estimated  $\sim 550\text{ GtC}$  stored in peats today (mean  $\delta^{13}\text{C} \sim -28\text{‰}$ ; Dioumaeva et al., 2002; Novák et al., 1999) and  $\sim 1000\text{ GtC}$  in the active layer in permafrost, which may have been partially thawed during the LIG (Reyes et al., 2010; Schuur et al., 2015; Stapel et al., 2018), less carbon stored in peat and permafrost at the LIG could have led to a lower total land carbon store compared to the Holocene. However, it is not possible to infer this total land carbon change from the oceanic and atmospheric  $\delta^{13}\text{C}$  anomalies because it cannot be assumed that the mass of carbon and  $^{13}\text{C}$  is preserved within the ocean–atmosphere–land biosphere system on glacial–interglacial timescales. There is indeed continuous exchange of carbon and  $^{13}\text{C}$  between the lithosphere and the coupled ocean, atmosphere, and land biosphere carbon reservoirs. Isotopic perturbations associated with changes in the terrestrial biosphere are communicated to the burial fluxes of organic carbon and  $\text{CaCO}_3$  and are therefore attenuated on multi-millennial time scales (Jeltsch-Thömmes et al., 2019; Jeltsch-Thömmes and Joos, 2020). Nevertheless, when hypothetically neglecting any exchange with the lithosphere, we find that the change in terrestrial carbon needed to explain the difference in  $\delta^{13}\text{C}$  would be on the order of  $310 \pm 44\text{ GtC}$  less during the LIG than the Holocene (Text S1).

In addition, due to the warmer conditions at the LIG than during the Holocene, there could have been a release

of methane clathrates which would have added isotopically light carbon ( $\delta^{13}\text{C}$ :  $\sim -47\text{‰}$ ) to the ocean–atmosphere system. However, available evidence suggests that geological  $\text{CH}_4$  sources are rather small (Bock et al., 2017; Dyonisius et al., 2020; Hmiel et al., 2020; Saunio et al., 2020) making this explanation unlikely, although we cannot completely exclude the possibility that the geological  $\text{CH}_4$  source was larger at the LIG than the Holocene. Similarly, since  $\delta^{13}\text{CO}_2$  from volcanic outgassing has a similar value to atmospheric  $\delta^{13}\text{CO}_2$  (Brovkin et al., 2016) and modelling suggests volcanic outgassing likely only had a minor impact on  $\delta^{13}\text{CO}_2$  (Roth and Joos, 2012), it is unlikely that volcanic outgassing of  $\text{CO}_2$  played a significant role in influencing the mean oceanic  $\delta^{13}\text{C}$ .

While changes in terrestrial carbon could have impacted the oceanic  $\delta^{13}\text{C}$  at the LIG, the LIG–Holocene differences in the isotopic signal of both the atmosphere and ocean were most likely due to a long-term imbalance between the isotopic fluxes to and from the lithosphere, including the net burial (or redissolution) of organic carbon and  $\text{CaCO}_3$  in deep-sea sediments, and changes in shallow water sedimentation and coral reef formation (Jeltsch-Thömmes and Joos, 2020).

## 5 Conclusions

We present a new compilation of benthic  $\delta^{13}\text{C}$  from 130 to 115 ka BP covering the LIG. Over this time period, benthic  $\delta^{13}\text{C}$  generally display a maximum value at  $\sim 121\text{ ka BP}$  ( $\pm 2\text{ kyr}$ ), in phase with the maximum atmospheric  $\delta^{13}\text{CO}_2$  (LIG value of  $-6.5\text{‰}$  at  $\sim 120\text{ ka BP}$ ). As there are significant chronological uncertainties associated with LIG records, we analyse data between 125 and 120 ka BP to avoid data associated with millennial-scale events and the deglaciation. We compare this LIG benthic  $\delta^{13}\text{C}$  data to a similar database covering the Holocene (7–2 ka BP). We find that during these specific time periods, LIG oceanic  $\delta^{13}\text{C}$  was about  $0.2\text{‰}$  lower than during the Holocene. This anomaly is consistent across different regions in the Atlantic Ocean. Even though there are fewer records available, benthic  $\delta^{13}\text{C}$  data from the Pacific Ocean also support an anomaly of about  $0.2\text{‰}$ .

An analysis of  $\delta^{13}\text{C}$  gradients across the Atlantic Ocean suggests that there were no significant changes in mean long-term ocean circulation between the two intervals. While reduced high northern latitude peat and permafrost caused by higher temperatures at the LIG than during the Holocene (Otto-Bliesner et al., 2021) could have led to a lower atmospheric and oceanic  $\delta^{13}\text{C}$ , the most likely explanation for the lower LIG oceanic  $\delta^{13}\text{C}$  is a long-term imbalance in the weathering and burial of carbon. Additional studies are required to further constrain the LIG carbon balance.

**Data availability.** The data are published on Research Data Australia at DOI <https://doi.org/10.26190/5efe841541f3b> (Bengtson et al., 2020).

**Supplement.** The supplement related to this article is available online at: <https://doi.org/10.5194/cp-17-507-2021-supplement>.

**Author contributions.** SAB, LCM, and KJM designed the research. CDP and LEL provided significant portions of the  $\delta^{13}\text{C}$  data. SAB, LCM, KJM, and LM analysed the data and developed the methodology. FJ assisted in the interpretation of the results. SAB prepared the manuscript with contributions from all co-authors.

**Competing interests.** The authors declare that they have no conflict of interest.

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