

Meshless Method of Lines for Numerical Solution of Kawahara Type Equations

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Abstract

In this work, an algorithm based on method of lines coupled with radial basis functions namely meshless method of lines (MMOL) is presented for the numerical solution of Kawahara, modified Kawahara and KdV Kawahara equations. The motion of a single solitary wave, interaction of two and three solitons and the phenomena of wave generation is discussed. The results are compared with the exact solution and with the results in the relevant literature to show the efficiency of the method.

Keywords: Kawahara Type Equations, Meshless Method of Lines (MMOL), Radial Basis Functions (RBFs)

1. Introduction

In this paper we study numerical solution of Kawahara equation, Modified Kawahara equation and KdV-Kawahara equation respectively given as:

 $u_t + u \ u_x + u_{xxx} - u_{xxxx} = 0, \quad a \le x \le b, \quad t \ge 0$ (1)

$$u_t + u^2 u_x + u_{xxx} - u_{xxxxx} = 0, \quad a \le x \le b, \quad t \ge 0$$
 (2)

$$u_t + u_x + u \ u_x + u_{xxx} - u_{xxxxx} = 0, \quad a \le x \le b, \quad t \ge 0$$
(3)

A variety of physical phenomena like, magneto acoustic waves in plasma [1], shallow water waves with surface tension [2] and capillary gravity water waves [3], are described by Kawahara Equation (1) and modified Kawahara Equation (2). KdV-Kawahara Equation (3) is used to describe the one dimensional evolution of small but finite amplitude long waves in various problems in fluid dynamics. This equation is a specific form of Benney-Lin equation [4,5].

Different analytic and numerical methods including the tanh-function method [6], Adomian decomposition method [7], sine-cosine method [8], variational iteration method, homotopy perturbation method [9], Crank-Nicolson Differential quadrature algorithms [10], Predictor corrector methods [11], Dual-Petrov Galerkin method [12] and RBF collocation method [13] have been proposed for solving the Kawahara type equations.

We shall use method of lines approach [14,15] using RBFs for numerical solution of the above problems using radial basis functions. The method of lines [16] is pow-

erful and comprehensive approach for solving time dependent partial differential equations (PDEs). This method involves two steps, first the approximation of spatial derivatives converting the PDEs to a system of ordinary differential equations (ODEs) and in the second step the resulting system of ODEs is integrated in time. One of the salient features of the MOL is the use of existing and well established numerical methods for ODEs. Coupling method of lines with radial basis functions makes it very simple by getting rid of mesh generation as compared to conventional mesh based methods. In 1990 Kansa [17, 18] for the first time used radial basis functions for numerical solution of PDEs. Other researchers like Fornberg et al. [19], Hon and Wu [20], Franke and Schaback [21], Wong et al. [22], Chen and Tanaka [23] also contributed a lot in this area. In order to solve the resultant collocation method Fasshauer modified the Kansa's method to a Hermite type collocation method [24]. For nonlinear time dependent PDEs see also [25,26]. In 1971 Hardy [27], developed MO to approximate geogra- phical surfaces. In Franke's [28] review paper, the MQ was ranked the best in some thirty scattered data interpolation methods. For solving PDEs the convergence proofs in applying the RBF's is given by Wu [29]. The accuracy of RBFs methods depends on the choice of a parameter called shape parameter involved in infinitely smooth RBFs like, Multiquadric (MQ), Gaussian (GA), Inverse multiquadric (IMQ) radial basis functions. Tarwater [30] found that the Root Mean Square of error decreases up to certain limit and then increases rapidly when $c \rightarrow \infty$.

Golberg, Chen, and Karur [31] and Hickernell and Hon [32] applied the technique of cross validation to obtain an optimal value of the shape parameter.

However the methods based on globally supported radial basis functions (GSRBFs) approach face the problem of ill-conditioning of the dense interpolation matrix. In order to overcome these difficulties several alternatives including domain decomposition [33], preconditioning [34] and use of compactly supported RBF [35] have been introduced. Another useful approach in this regard is locally supported RBFs instead of globally supported RBFs. The readers are recommended to visit Wu and Liu [36], C. K. Lie et al. [37], R. Vertnik and Bozidar Sarler [38]. It is worth mentioning that global, infinitely differentiable RBFs typically interpolate smooth data with spectral accuracy [39-42] and the shape parameter can be adjusted with the number of centers in order to produce a interpolation matrix which is well conditioned enough to be inverted in finite precision arithmetic [43]. The globally supported RBFs were used early on and for the problems whose size does not exceed 400-500 data points. These methods should still be the method of choice [44].

This paper is organized in three sections. In Section 2 the numerical scheme is explained and Section 3 contains the numerical examples for the justification of the method and we conclude in Section 4.

2. Numerical Scheme

For implementation of numerical method, we consider the Kawahara Equation (1) with the following initial and boundary conditions

$$u(x,0) = f(x), a \le x \le b,$$
(4)

$$u(a,t) = g_1(t), u(b,t) = g_2(t)$$
 (5)

To apply meshless method of lines, we first use radial basis functions to approximate space derivatives. The problem domain [a,b] is divided into nodes $x_i, i = 1, 2, \dots, N$. Out of these points $x_i, i = 2, \dots, N-1$ are interior points and $x_i, i = 1$ and N are the boundary points.

The approximate solution of u(x,t) is given by

$$u(x,t) = u^{n}(x,t) = \sum_{j=1}^{N} \lambda_{j}^{n}(t) \Psi(x_{j}) = \Psi^{T}(x) \lambda, \qquad (6)$$

where

$$\boldsymbol{\Psi}^{\mathrm{T}}(\boldsymbol{x}) = \left[\boldsymbol{\psi}_{1}(\boldsymbol{x}), \boldsymbol{\psi}_{2}(\boldsymbol{x}), \cdots, \boldsymbol{\psi}_{N}(\boldsymbol{x}) \right]^{\mathrm{T}}$$

We have used the following RBFs

$$\psi(x_j) = \sqrt{\left(x - x_j\right)^2 + c^2}, MQ$$
$$\psi(x_j) = \exp\left(-c\left(x - x_j\right)^2\right), GA$$

$$\psi(x_j) = \frac{1}{\sqrt{\left(x - x_j\right)^2 + c^2}} , IMQ$$

where $j = 1, 2, \dots, N$ and *c* being shape parameter. The Equation (6) in the matrix form is

 $\boldsymbol{u} = \boldsymbol{A}\boldsymbol{\lambda} \tag{7}$

where

$$\boldsymbol{u} = \left[u_{1}(t), u_{2}(t), \cdots, u_{N}(t)\right]^{\mathrm{T}}, \boldsymbol{\lambda} = \left[\lambda_{1}, \lambda_{2}, \cdots, \lambda_{N}\right]^{\mathrm{T}}$$

and

$$\boldsymbol{A} = \begin{bmatrix} \boldsymbol{\Psi}^{\mathrm{T}}(x_{1}) \\ \boldsymbol{\Psi}^{\mathrm{T}}(x_{2}) \\ \dots \\ \boldsymbol{\Psi}^{\mathrm{T}}(x_{N}) \end{bmatrix} = \begin{bmatrix} \psi_{1}(x_{1}) & \psi_{2}(x_{1}) & \dots & \psi_{N}(x_{1}) \\ \psi_{1}(x_{2}) & \psi_{2}(x_{2}) & \dots & \psi_{N}(x_{2}) \\ \vdots & \vdots & \ddots & \vdots \\ \vdots & \vdots & \ddots & \vdots \\ \psi_{1}(x_{N}) & \psi_{2}(x_{N}) & \dots & \psi_{N}(x_{N}) \end{bmatrix}$$

Using Equation (7) in Equation (6) it follows that

$$u(x,t) = \boldsymbol{\Psi}^{\mathrm{T}}(x)\boldsymbol{A}^{-1}\boldsymbol{u} = \boldsymbol{S}(x)\boldsymbol{u}, \qquad (8)$$

where $S(x) = \Psi^{T}(x)A^{-1}$

Applying Equation (8) to Equation (1), and collocating at each node x_i , we get system of first order ODEs

$$\frac{\mathrm{d}u_i}{\mathrm{d}t} + u_i \left(\boldsymbol{S}_x \left(x_i \right) \boldsymbol{u} \right) + \left(\boldsymbol{S}_{xxx} \left(x_i \right) \boldsymbol{u} \right) - \left(\boldsymbol{S}_{xxxxx} \left(x_i \right) \boldsymbol{u} \right) = 0, \quad (9)$$

$$i = 1, 2, \cdots, N$$

where $u_i(t) = u_i$

$$S_{x}(x_{i}) = [S_{1x}(x_{i}), S_{2x}(x_{i}), \dots, S_{Nx}(x_{i})], i = 1, 2, \dots, N$$
$$S_{jx}(x_{i}) = \frac{\partial}{\partial x}S_{j}(x_{i}), i, j = 1, 2, \dots, N$$
In the similar fashion

 $S_{xxx}(x_i) = \left[S_{1xxx}(x_i), S_{2xxx}(x_i), \cdots, S_{Nxxx}(x_i) \right],$ $i = 1, 2, \cdots, N$ $S_{jxxx}(x_i) = \frac{\partial^3}{\partial x^3} S_j(x_i), \quad i, j = 1, 2, \cdots, N$ $S_{xxxxx}(x_i) = \left[S_{1xxxxxx}(x_i), S_{2xxxxxx}(x_i), \cdots, S_{Nxxxxxxx}(x_i) \right],$ $i = 1, 2, \cdots, N$

$$S_{jxxxxx}\left(x_{i}\right) = \frac{\partial^{5}}{\partial x^{5}} S_{j}\left(x_{i}\right), \quad i, \quad j = 1, 2, \cdots, N$$

In order to write the above system of equations in terms of column vectors, let

$$\boldsymbol{U} = \begin{bmatrix} u_1 & u_2 & u_3 & \dots & u_{N-1} & u_N \end{bmatrix}^{\mathrm{T}},$$
$$\boldsymbol{S}_x = \begin{bmatrix} S_{jx} (x_i) \end{bmatrix}_{N \times N},$$

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$$\begin{split} \boldsymbol{S}_{xxx} &= \left\lfloor \boldsymbol{S}_{jxxx} \left(\boldsymbol{x}_{i} \right) \right\rfloor_{N \times N}, \\ \boldsymbol{S}_{xxxxxx} &= \left[\boldsymbol{S}_{jxxxxx} \left(\boldsymbol{x}_{i} \right) \right]_{N \times N}, \end{split}$$

Equation (9) then can be written as follows

$$\frac{\mathrm{d}\boldsymbol{U}}{\mathrm{d}t} + \boldsymbol{U} * (\boldsymbol{S}_{x}\boldsymbol{U}) + (\boldsymbol{S}_{xxx}\boldsymbol{U}) - (\boldsymbol{S}_{xxxx}\boldsymbol{U}) = \boldsymbol{0} \qquad (10)$$

Rewriting Equation (10) as

$$\frac{\mathrm{d}\boldsymbol{U}}{\mathrm{d}t} = G\left(\boldsymbol{U}\right) \tag{11}$$

where

 $G(U) = -U * (S_x U) - (S_{xxx} U) + (S_{xxxxx} U)$ and the symbol (*) denotes component by component multiplication of two vectors.

The initial condition is

$$U(t_{0}) = \left[u^{0}(x_{1}), u^{0}(x_{2}), \cdots, u^{0}(x_{N})\right]$$
(12)

From the boundary conditions described in (5) we get

$$u_1(t) = g_1(t) \text{ and } u_N(t) = g_2(t)$$
 (13)

Now we use classical fourth order Runge-Kutta scheme to solve Equations (11)-(13), namely

$$\begin{cases} \boldsymbol{U}^{n+1} = \boldsymbol{U}^{n} + \frac{\Delta t \left(K_{1} + 2(K_{2} + K_{3}) + K_{4} \right)}{6} \\ K_{1} = G \left(\boldsymbol{U}^{n} \right), K_{2} = G \left(\boldsymbol{U}^{n} + \frac{\Delta t}{2} K_{1} \right) \\ K_{3} = G \left(\boldsymbol{U}^{n} + \frac{\Delta t}{2} K_{2} \right), K_{4} = G \left(\boldsymbol{U}^{n} + \Delta t K_{3} \right) \end{cases}$$
(14)

The RK4 scheme does not face problem of stability as long as the time step Δt is chosen sufficiently small (see Collatz [45]) which is so in our case. The rule of thumb in selection of the time step Δt for stability is as follows [46]: "The method of lines is stable if the eigenvalues of the (linearized) spatial discretization operator, scaled by time step Δt , lie in the stability region of the timediscretization operator". So the method is stable if all the eigenvalues $(\alpha_i, j = 1, 2, \dots, N)$ of the Jacobian matrix in (11) lie inside the stability region R (*i.e.* $-2.78 < \Delta t \alpha_i < 0$) of RK4 scheme. For further details regarding stability of Runge-Kutta fourth order method, see Lambert [47] and Jain [48]. As far as the selection of the shape parameter cis concerned in this work, first we have to find an interval for c in which matrix A of radial basis functions is invertible and then select a value from that interval which gives us the most accurate results.

3. Numerical Application

In this section the numerical results for Kawahara, modified Kawahara and KdV Kawahara equations are presented. The root mean square norm and maximum error norms are calculated by the formulas

$$L_{2} = \left| u_{ex} - u_{num} \right|^{2} = \sum_{j=1}^{N} \Delta x \left| \left(u_{ex} \left(j \right) - u_{num} \left(j \right) \right)^{2} \right| \quad (15)$$

$$L_{\infty} = \left| u_{ex} - u_{num} \right|_{\infty} = \max_{j} \left| u_{ex} \left(j \right) - u_{num} \left(j \right) \right|$$
(16)

For Kawahara equation the lowest three conserved quantities, defined in [49] are also calculated using

$$I_{1} = \int_{-\infty}^{\infty} u dx, \quad I_{2} = \int_{-\infty}^{\infty} \frac{1}{2} u^{2} dx,$$
$$I_{3} = \int_{-\infty}^{\infty} \left(\frac{13}{8} \delta u_{x}^{2} - \frac{105}{96} \lambda^{2} u^{3} + \frac{1}{2} u_{xx}^{2}\right) dx$$

Example 3.1

Consider the Kawahara equation

 $u_t + uu_x + u_{xxx} - u_{xxxxx} = 0$

with the following initial and boundary conditions

$$u(x,0) = \frac{105}{169} \operatorname{sech}^{4} \left(k \left(x - x_{o} \right) \right)$$
(17)

$$u(a,t) = 0; u(b,t) = 0$$
 (18)

The exact solution given in [10] is,

$$u(x,t) = \frac{105}{169} \operatorname{sech}^{4} \left(k \left(x - \frac{36}{169} t - x_{o} \right) \right)$$
(19)

where $k = \frac{1}{2}\sqrt{\frac{1}{13}}$

For numerical computation we take [a,b] = [-20,30], $x_0 = 2, \Delta x = 1, N = 51$. The simulation is carried out up to time t = 25. L_2 and L_{∞} norms are calculated at t = 0, 5, 515 and 25, using MQ and GA radial basis functions with the value of shape parameter found to be in neighborhood of 4 as shown in Figure 1 and similarly for GA the value is found to be in neighborhood of 0.3. We have searched the optimal value of the shape parameter by plotting maximum error verses shape parameter with step 0.01. The three conserved quantities are also shown in the **Table 1**. The amplitudes and peak position of the solitary waves are also calculated. The results with present method using MQ are better than the polynomial based differential quadrature (PDQ) method [10] and are very close to cosine expansion based differential quadrature (CDQ) method [10]. While the results obtained by GA are better than both methods mentioned in [10]. In Figure 2 the forward motion of the solitary wave in comparison with exact solution (19) at different time levels is shown.

The Point wise rate of convergence in space is calculated using the following formula:

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Table 1. Results for Kawahara equation in comparison with [10].

Method	Time	$L_2 \times 10^3$	$L_{\infty} \times 10^3$	I_1	I ₂	I ₃	Height	Peak Position	CPU time(s)
MQ(c = 4.3)	0	0	0	5.97359	1.27250	-0.16458	0.62130	2	
	5	0.09468	0.04669	5.97348	1.27250	-0.16458	0.62119	3	0.094
	15	0.15362	0.05939	5.97343	1.27250	-0.16458	0.62038	5	0.188
	25	0.16818	0.04660	5.97355	1.27250	-0.16458	0.61880	7	0.328
	0	0	0	5.973599	1.272502	-0.16458	0.621301	2	
GA(c = 0.27)	5	0.10075	0.034297	5.973662	1.272502	-0.16458	0.621201	3	0.171
	15	0.10113	0.03830	5.973675	1.272502	-0.16458	0.620382	5	0.172
	25	0.13160	0.03990	5.973532	1.272502	-0.16458	0.618802	7	0.281
	0	0	0	5.97357	1.27250	-0.16458	0.62130	2	
$PDQ[10] (\Delta t = 0.1)$	5	1.986	0.921	5.97060	1.27250	-0.16458	0.62102	3	
	15	2.543	1.045	5.97014	1.27250	-0.16458	0.62047	5	
	25	2.851	0.863	5.97353	1.27250	-0.16457	0.61872	7	
CDO[10]	0	0	0	5.97357	1.27250	-0.16458	0.62130	2	
	5	0.151	0.043	5.97372	1.27250	-0.16458	0.62122	3	
	15	0.156	0.049	5.97364	1.27250	-0.16458	0.62037	5	
	25	0.159	0.076	5.97350	1.27250	-0.16458	0.61877	7	



Figure 1. Error verses shape parameter c for Example 3.1.



Figure 2. Travelling wave solution of Kawahara equation (solid lines showing numerical solution and dot (.) showing exact solution.

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$$\frac{\log_{10} \left(\left\| u - U_{h_i} \right\| / \left\| u - U_{h_{i+1}} \right\| \right)}{\log_{10} \left(h_i / h_{i+1} \right)}$$

where *u* and U_{h_i} represents the exact solution and the numerical solution respectively and h_i is spatial step size. We calculate spatial rate of convergence by keeping time step $\Delta t = 0.001$ fixed and varying the number of collocation points (F = 20, 40, 80). From the **Table 2** we can see that the order of convergence decreases with the smaller spatial step size. In all numerical examples we have used MQ and GA in order to calculate order of convergence.

Example 3.2

Consider the modified Kawahara equation

$$u_t + u^2 u_x + u_{xxx} - u_{xxxxx} = 0$$

with the following initial and boundary conditions

$$u(x,0) = D\operatorname{sech}^{2}(k(x))$$
(20)

$$u(a,t) = D \operatorname{sech}^2 \left(k \left(a - Bt \right) \right);$$

$$u(b,t) = D\operatorname{sech}^{2}(k(b-Bt))$$
⁽²¹⁾

The above conditions are extracted from the exact solution given in [50],

$$u(x,t) = D\operatorname{sech}^{2}(k(x-Bt))$$
(22)

where $D = -\frac{3}{\sqrt{10}}, k = \frac{1}{2}\sqrt{\frac{1}{5}}, B = \frac{4}{25}$

The calculation is carried out by taking [a,b] = [-30,30] with x = 1. We use MQ, GA and IMQ radial basis with shape parameter found to be in the neighborhood of 3 for MQ as shown in **Figure 3** and for GA and IMQ it is in the neighborhood of 0.5 and 0.0001. The L_2 and L_{∞} norms calculated at t = 0, 5, 15, 25 are shown in **Table 3**.

N	L_{∞}	order	L_2	Order
MQ				
20	2.81001×10^{-2}		1.01808×10^{-1}	
40	5.07036×10 ⁻⁵	9.1143	1.67092×10^{-4}	9.2509
80	3.56316×10 ⁻⁵	0.5089	1.28499×10 ⁻⁴	0.3788
GA				
20	1.49094×10^{-2}		5.38775×10^{-2}	
40	5.01268×10 ⁻⁵	8.2164	1.35443×10 ⁻⁴	8.6358
80	4.77379×10 ⁻⁵	0.0704	1.25773×10^{-4}	0.1068

Table 2 Degults for Madified Veryahara equation

Table 2. Spatial rate of convergence at for Example 3.1, t = 25.

Table 5. Results for Mounteu Rawanara equation.						
method	time	L_{∞}	L_2	I_1	I_2	CPU time(s)
MQ(c = 2.9)	0	0	0	-8.48525	2.68328	
	5	6.1995×10 ⁻⁵	1.7896×10^{-4}	-8.48524	2.683176	0.187
	15	1.0717×10^{-4}	2.7337×10^{-4}	-8.48487	2.68296	0.313
	25	1.2130×10^{-4}	3.4855×10^{-4}	-8.48464	2.68275	0.453
GA(c = 0.43)						
	0	0	0	-8.48525	2.68328	
	5	5.3996×10 ⁻⁵	1.0804×10^{-4}	-8.48510	2.68317	0.171
	15	8.6124×10 ⁻⁵	1.9575×10^{-4}	-8.48472	2.68296	0.313
IMQ(c = 0.0001)	25	7.8371×10 ⁻⁵	2.5257×10^{-4}	-8.48442	2.68275	0.543
	0	0	0	-8.48525	2.683281	
	5	1.80659×10^{-1}	5.2570×10^{-1}	-8.48525	2.683281	0.203
	15	5.18281×10^{-1}	5.2819×10^{-1}	-8.48134	2.683307	0.343
	25	7.62711×10^{-1}	2.2691×10^{0}	0 17121	2 602251	0.460





Figure 3. Error verses shape parameter c for Example 3.2.

GA and MQ are showing better accuracy than IMQ. The order of convergence in space decreases with increasing N as shown in **Table 4**. The solitary wave profile at different time levels in comparison with the exact solution is shown in Figure 4.

Example 3.3

Consider the KdV-Kawahara equation

$$u_t + uu_x + u_x + u_{xxx} - u_{xxxxx} = 0$$

with the following initial and boundary conditions

$$u(x,0) = \frac{105}{169} \operatorname{sech}^{4} \left(\frac{1}{2\sqrt{13}} (x - x_{o}) \right)$$
(23)

$$u(x,a) = \frac{105}{169} \operatorname{sech}^{4} \left(\frac{1}{2\sqrt{13}} \left(a - \frac{205}{169} t - x_{o} \right) \right)$$
(24)

$$u(x,b) = \frac{105}{169} \operatorname{sech}^{4} \left(\frac{1}{2\sqrt{13}} \left(b - \frac{205}{169} t - x_{o} \right) \right)$$
(25)

initial condition and boundary conditions are extracted from the exact solution given in [51].

$$u(x,t) = \frac{105}{169} \operatorname{sech}^{4} \left(\frac{1}{2\sqrt{13}} \left(x - \frac{205}{169} t - x_{o} \right) \right)$$
(26)

The calculation is carried out by taking [a,b] = [0,200]with $\Delta x = 1$. The discrete root mean square error norm L_2 and maximum error norm L_{∞} are calculated using MQ, GA and IMQ for time t = 1 up to 5. From the results shown in Table 5 we can see that the both MQ and GA are showing very good agreement with the exact solution. Shape parameter verses error plot for MQ is shown in Figure 5. The spatial rate of convergence is shown in Table 6. The order of convergence decreases by increasing collocation points for a fixed time step $\Delta t =$ 0.001. The forward movement of the solitary wave at different time levels in comparison with the exact solution (26) is shown in Figure 6, same as in [52].

Example 3.4

Considering Equation (1) for interaction of two positive solitary waves with the following initial condition

$$u(x,0) = \sum_{i=1}^{2} A_i^2 \operatorname{sech}^4\left(\frac{\sqrt{\sigma A_i}}{4}(x-x_i)\right).$$

We solve the problem by using MQ and GA RBFs taking [a,b] = [-50,100] with N = 201. The calculation is carried

N L_{∞} order L_2 order MQ 2.56860×10^{-1} 5.20872×10^{-1} 24 48 1.49043×10^{-3} 7.4291 2.65141×10⁻³ 7.6180 96 2.42285×10^{-5} 5.9428 5.67016×10⁻⁵ 5.5472

Table 4. Spatial rate of convergence at for Example 3.2, t = 5.

24 48 96	$\begin{array}{c} 3.47477 \times 10^{-2} \\ 5.00876 \times 10^{-4} \\ 2.17409 \times 10^{-5} \end{array}$	6.11 4.52	.63 259	9.95424> 9.19460> 5.46571>	<10 ⁻² <10 ⁻⁴ <10 ⁻⁵	6.7583 4.0723
	Table 5. Re	sults for KdV	⁷ Kawahai	ra equation	at $t = 5$.	
method	L_{∞}	L_2	I_1	I_2	Amplitude	CPU time(s)
MQ(c = 2.6)	b) 1.0977×10^{-4}	3.7697×10 ⁻⁴	5.97359	1.27250	0.621241	3.266
GA(c = 0.2)) 4.7924×10^{-6}	1.8465×10 ⁵	5.97368	1.27250	0.621200	3.390
IMQ(c = 0.00)	01) 5.0111×10^{-1}	1.5920×10^{0}	5.95790	1.27250	0.619041	3.641

Table 6. Spatial rate of convergence at for Example 3.3, t = 5.

Ν	L_{∞}	order	L_2	order
MQ				
40	3.12550×10^{-2}		1.33333×10^{-1}	
80	1.62584×10^{-3}	4.2648	4.60613×10 ⁻³	4.8553
160	1.56826×10^{-4}	2.1471	3.73104×10 ⁻⁴	2.7242
GA				
40	2.6983×10^{-3}		1.1620×10^{-2}	
80	1.4710×10^{-4}	4.1971	3.2768×10^{-3}	4.7639
160	1.2344×10 ⁻⁵	3.5749	2.4624×10 ⁻⁴	3.7341





out up to time t = 50 by taking time step $\Delta t = 0.001$. For our numerical calculations the values of the parameters involved in above equation are chosen as:

$$x_1 = 0, x_2 = 20, \sigma = \frac{4}{\sqrt{105}}, A_1 = \frac{10}{\sigma}, A_2 = \frac{8}{\sigma}$$

The two solitary waves propagate towards right as the time progresses. The process of interaction is shown in **Figure 7**. During this process the larger wave catches up the smaller one and then the both waves separate from each other maintaining their original shape. From **Table 7**, we can see that the invariants of motion remain almost conserved as time increases. The variation in the three



Figure 5. Error verses shape parameteric for Example 3.3.

conserved quantities is found to be in the range:

MQ 40.50925 $\leq I_1 \leq$ 40.48389, 45.83614 $\leq I_2 \leq$ 45.85093, $-32.37219 \leq I_3 \leq -32.15991$, GA 40.50925 $\leq I_1 \leq$ 40.41284, 45.83614 $\leq I_2 \leq$ 45.84364, $-32.37218 \leq I_3 \leq -32.14082$.

$$55014 \le 1_2 \le 45.04504, 52.57210 \le 1_3 \le$$

Example 3.5

For interaction of three solitary waves we consider Equation (1) with the initial condition

$$u(x,0) = \sum_{i=1}^{3} A_i^2 \operatorname{sech}^4 \left(\frac{\sqrt{\sigma A_i}}{4} (x - x_i) \right)$$



Figure 6. Solitary wave motion of KdV Kawahara equation (solid lines show numerical solution and dot (.) showing exact solution).

We use multiquadric and Gaussian radial basis functions in our numerical simulations to solve this problem. The spatial domain is selected as [a,b] = [-30,120] with $\Delta x = 0.75$. The calculation is carried out up to time t =50 taking time step $\Delta t = 0.001$. The values of the parameters used in above equation are selected as:

$$x_1 = -20, x_2 = 0, x_3 = 20, \ \sigma = \frac{4}{\sqrt{105}}, A_1 = \frac{10}{\sigma}, A_2 = \frac{8}{\sigma}, A_3 = \frac{6}{\sigma}$$

The three solitary waves propagate towards right. The process of interaction is shown in **Figure 8**. During this process the taller wave moves faster and catches up the smaller waves and then the three waves separate from each other. The shape of the three solitary waves after collision is maintained. From **Table 8** we can see that the invariants of motion remain almost conserved as time increases. The variation in the three conserved quantities



Figure 7. Interaction of two solitons for Kawahara equation Example 3.4.

Table 7. Invariants for interaction of two solitons for Example 3.4.

		MQ			GA	
time	I_1	I_2	I_3	I_1	I_2	I_3
0	40.509259	45.836141	-32.37219	40.509259	45.836141	-32.37218
10	40.507987	45.839240	-32.12856	40.492695	45.837775	-32.12820
20	40.499950	45.8427481	-32.73918	40.466653	45.839353	-32.73746
30	40.5207361	45.8359623	-32.58956	40.571742	45.840373	-32.59560
40	40.5552400	45.8443065	-32.10125	40.579624	45.841964	-32.12278
50	40.4838916	45.8509384	-32.15991	40.412842	45.843648	-32.14082

is found to be in the range:

$$\begin{split} \text{MQ } 51.47263 \leq I_1 \leq 51.56859, \\ 51.10049 \leq I_2 \leq 51.15838, \\ -33.63806 \leq I_3 \leq -33.06973, \end{split}$$

GA $51.47263 \le I_1 \le 51.47389, 51.10049 \le I_2 \le 51.10452$,

 $-33.63806 \le I_3 \le -33.39361.$

Example 3.6

In this example we will show the phenomena of wave generation for Equation (1). We consider the following initial condition

$$u(x,0) = \beta \operatorname{sech}^{4} \left(\frac{1}{2\sqrt{13}} (x - x_{o}) \right)$$

Computational domain [-40,130] with h = 0.625 is considered. The scheme is run up to time t = 18. We take

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 $\beta = 10$. The single solitary wave splits in to three solitary waves. The conserved quantities, height and position of three waves are calculated at various time levels as shown in **Table 9**. With passage of time the leading

wave owing to its faster velocity gets far from the other two waves as shown in **Figure 9**. Variation in three conserved quantities is found to be in the following range:

time		MQ			GA			
time	I_1	I_2	I_3	I_1	I_2	I_3		
0	51.472631	51.100493	-33.63806	51.472631	51.100493	-33.63806		
10	51.631982	51.094986	-33.40011	51.835962	51.098240	-33.39015		
20	51.688762	51.103862	-34.23846	51.848746	51.099504	-34.21111		
30	51.664151	51.127187	-35.19047	51.690220	51.101129	-35.15871		
40	51.601572	51.14098	-33.85613	51.493752	51.102653	-33.83892		
50	51.568591	51.15838	-33.06973	51.473898	51.104529	-33.39361		

 Table 8. Invariants for interaction of three solitons for Example 3.5.







Figure 9. Generation of waves for Example 3.6.

$$\begin{split} & \text{MQ } 96.14365 \leq I_1 \leq 96.14365, 329.65039 \leq I_2 \\ & \leq 329.65039, -875.41604 \leq I_3 \leq -701.42681, \\ & \text{GA } 96.14365 \leq I_1 \leq 96.14365, 329.65039 \leq I_2 \\ & \leq 329.65039, -875.41612 \leq I_3 \leq -706.85789. \end{split}$$

4. Conclusions

In this paper, we have used method of lines coupled with radial basis functions for numerical solution of Kawahara type equations. The numerical results describing motion of single solitary wave, interaction of two and three solitary waves and phenomena of wave generation have been discussed. The accuracy of the solution depends upon the choice of the shape parameter, which has been selected experimentally. The numerical results using MQ and GA for Kawahara equation are better than the Crank Nicolson differential quadrature algorithm [10]. The invariants of motion remained conserved during the process of computation for all cases. Two main advantages of this method are mesh less property and use of ODE solvers of high quality and their codes to approach the solution of PDEs. Also the presented method is simple, easy to implement because no mesh is required in the problem domain. Only radial distance between the nodes is used to approximate the solution.

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6. References

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