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Multi-GNSS satellite clock estimation constrained with oscillator noise model in the existence of data discontinuity

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Abstract During the past years, Real-time Precise Point Positioning (PPP) has been proven to be an efficient tool in the applications of navigation, precise orbit determination of LEO as well as earthquake and tsunami early warning, etc. One of the most crucial issues of these applications is the high precision real-time GNSS satellite clock. Though the performance and character of the GNSS onboard atomic frequency standard have been widely studied, the white noise model is still the most popular hypothesis that employed in the real-time GNSS satellite clock estimation. However, concerning the real-time applications, significant data discontinuity may arise either due to the fact that only regional stations involved, or the failure in the stations, satellites and network connections. These data discontinuity would result in an arbitrary clock jump between adjacent arcs when the clock offsets are modelled as white noise. In addition, it is also expected that the detection and identification of outliers would be benefited from the constrains of the satellite oscillator noise model. Thus in this contribution, based on the statistic analysis of almost two years multi-GNSS precise clock products, we developed the oscillator noise model for the satellites of GPS, GLONASS, BDS and Galileo according to the oscillator type as well as the block type. Then, the efficiency of this oscillator noise model in multi-GNSS satellite clock estimation is demonstrated with two months data for both regional and global networks in simultaneous realtime mode. For the regional network, the results suggest that compared with the traditional solution based on white noise model, the improvement is 44.4% and

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12.1% on average for STD and RMS, respectively, and the improvement is mainly attributed to the efficiency of the oscillator noise model during the convergence period and the gross error resistance. Concerning the global experiment, since the stations guarantee the continuous tracking of the satellites with redundant observable, the improvement is not as evident as that of regional experiment for GPS, GLONASS and BDS. While, the STD of Galileo clock improves from 0.28ns to 0.19ns due to that, the satellites E14 and E18 still suffer significant data discontinuity during our experimental period.

Keywords Clock model, Data discontinuity, Multi-GNSS, Real-time

1 Introduction

The potential of Global Navigation Satellite Systems (GNSS) as an efficient tool in providing precise positioning has been widely recognized nowadays. Specifically, the so-called Precise Point Positioning (PPP) is one of the most promising techniques because of its cost-effectiveness, global coverage and high accuracy [Zumberge et al. (1997),Kouba and Héroux. (2001),Bisnath and Gao (2008)]. Moreover, in order to meet the growing requirements of real-time PPP applications, e.g. navigation, precise orbit determination of LEO, earthquake and tsunami early warning and GNSS-based weather forecasting, etc, the International GNSS Service (IGS) has issued a call for participation in IGS Real-time Pilot Project (IGS-RTPP) since 2007, and provided the precise orbit and clock correction for GPS and GLONASS officially in April 2013 [Weber et al. (2007),Shi et al. (2013), Zhang et al. (2015)].

Accuracy of orbit and clock products are fundamental to guaranty the performance of PPP. The predicted part of orbit provided by IGS ultra-rapid ephemeris in real-time has precision of centimeter level, which is suitable for PPP applications. The predicted clock of ultra-rapid product, however, deviates quickly due to the frequency variation originated from a variety of effects, e.g., ambient temperature, supply voltages, magnetic field, etc. Thus, the real-time estimation of satellite clocks is one of the key objectives of high precision GNSS service.

Obviously, a thorough understanding of the behavior of Atomic Frequency Standard (AFS) equipped on GNSS is useful for an optimal estimation of satellite clocks. And attribute to the successful deployment and wide-spread of GNSS, the stability of these onboard atomic oscillator has been analyzed in details since the early 1990s within the frequency transfer community [Petit and Thomas (1996)]. Based on the two-sample variance, referred as Allan variance, a number of efforts has been made to derive the noise spectral density of GPS Cesium oscillator, which is essential to characterize the frequency fluctuations of atomic clocks [Swift and Hermann (1988), Wübbena (1988)]. Along with the GPS program progress, increasing satellites are equipped with Rubidium clocks. As pointed out by [Hutsell (1995)], the Rubidium oscillator exhibits significant frequency drift (aging) and aging noise characteristics, which may leave the Allan variance analysis divergent. Thus, the three-sample variance, i.e., Hadamard variance, which removes the linear frequency aging automatically, presented an appropriate tool for characterizing these AFS noise [Hutsell (1995)]. By implementation of these two techniques, a wide range of valuable literatures has been published in parallel to the upgrading of the GPS satellites, based on a much wider

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scope of clock observables with an improved quality from IGS [Wright (2007), Senior et al. (2008), Huang et al. (2013A)]. More recently, along with the modernization of GPS and GLONASS, and the newly deployed BDS and Galileo, the oscillator stability of the new generation GNSS has received increasing interests again [Montenbruck et al. (2012), Steigenberger et al. (2013), Wang et al. (2017)]. It is concluded that the typical Allan variance of daily frequency stability is between 10^{-14} and 10^{-13} for the GNSS onboard oscillators [Steigenberger et al. (2013), Huang et al. (2013A)].

It is noted that most of the above studies are focused on the performance of onboard oscillator and frequency transfer based on the GNSS common-view technique. Concerning the satellite clock determination, however, the white noise model is still the most popular hypothesis that is underlining the parameter estimation. On the one hand, taking the high-resolution requirement of real-time applications into consideration, the corresponding clock estimation turns out to be too time-consuming to support PPP with a sampling rate up to 1 Hz. It is the reason why great efforts are underway in direction of an efficient satellite clock solution within the GNSS community. Among which, [Bock et al. (2009A)] proposed to perform low rate, i.e., 5 min, clock estimation and receiver clock synchronization using code measurements, and then to densify them with the clock increments derived from the epoch-differenced phase measurements. Similarly, the clock estimation model based on the un-differenced pseudo-range and epoch-differenced carrier-phase was also suggested by [Ge et al. (2011)], but in a real-time fashion. Obviously, the computation can be accelerated in the procedure of epoch-difference as the elimination of ambiguities. In contrast, the employment of the AFS noise model implies an extra computational burden in the satellite clock solution, since the oscillator noise model typically involves the simultaneous estimation of clock error, frequency offset, and aging. On the other hand, as evidenced by a wide range of publications, the performance of the clock corrections generated with the white noise model seems to work fairly well with an overall statistics of about 0.1 ns, which is sufficient for the centimeter level positioning with a stand-alone receiver [Bock et al. (2009B), Laurichesse et al. (2013), Zhang et al. (2015)].

However, as far as the real-time PPP applications are concerned, the problem is more complicated. Firstly, the assumption that the satellites can be continuously tracked relies on the distribution of the stations involved as well as the network condition between the stations and the real-time processing facility. However, there are numerous cases that only regional station's observables are available [Shi et al. (2017)]. In addition, a significant amount of data might be lost or arrive too late for processing due to network problems [Hauschild and Montenbruck (2009)]. This data discontinuity will result in an arbitrary clock jump between adjacent arcs when the clock offsets are modelled as white noise parameters, and consequently cause a degradation of the PPP performance. Secondly, the detection and identification of outlier satellites is more challenging under the practical environment, and it is expected that the filter robustness may benefit from the predicted satellite clock offsets based on the oscillator noise model. Last but not least, though the effect is potentially not harmful to positioning, it should be mentioned that a more stable satellite clock value constrained with an a priori oscillator noise model is always encouraged in the applications of GNSS based timing and synchronization. Unfortunately, to the best of our knowledge, only [Hauschild and Montenbruck (2009), Hauschild (2010)] have absorbed the oscillator noise model in real-time GPS clock estimation, and with an identical process noise setting for all the satellites. In addition, they have also failed to fully excavate the potential of oscillator noise model constrains in the resistance to the data discontinuity for clock estimation.

It is generally acknowledged that the satellite clock behavior depends on AFS oscillator type as well as block type [Senior et al. (2008)]. In addition, as evidenced by [Lou et al. (2015)], compared with the GPS only solution, better performance of PPP would be achieved by employing the multi-GNSS constellations. Furthermore, an efficient solution for real-time GNSS network processing has been promoted by [Gong et al. (2017)], in which the epoch multi-GNSS clock generation based on undifferenced pseudo-range and carrier-phase is around 0.3 s with over 80 globally distributed stations.

These deficiencies of the previous studies as well as the achievements in realtime multi-GNSS data processing recently as discussed above are the main drivers of this contribution, in which, we have demonstrated the efficiency of the AFS model in real-time multi-GNSS clock estimation in case of discontinued satellite tracking. This paper is organized as follows: after the introduction of the basic multi-GNSS data processing model with the emphasize on the oscillator model; the oscillator noise model for GPS, GLONASS, BDS and Galileo are studied and promoted; with the oscillator model, both regional and global satellite clock estimation are performed and analyzed in simulate real-time mode.

2 Method

2.1 Multi-GNSS Data Processing Model

In this paper, the satellite system identifiers $(G \ R \ C \ E)$ as suggested in RINEX 3.02 format are used to denote GPS, GLONASS, BDS and Galileo, respectively. The linear combination known as ionosphere-free (IF) that is typically involved in the satellite clock estimation for a specific GNSS system sys (sys $\in (G \ R \ C \ E)$) is expressed as

$$\left. \begin{array}{l}
P_r^s = p_r^s - \rho_r^s \\
= t_{r,sys} - t^s + \alpha_r^s \cdot T_r^z + e_{sys} \cdot b_r^k \\
\Phi_r^s = \phi_r^s - \rho_r^s \\
= t_{r,sys} - t^s + \alpha_r^s \cdot T_r^z + e_{sys} \cdot b_r^k - \lambda \cdot N_r^s + \varepsilon_\phi
\end{array} \right\}$$
(1)

in which, P_s^r , Φ_s^r are observation minus calculated (OMC) for pseudo-range and carrier-phase from receiver $r(r = 1, \dots, i)$ to satellite $s(s = 1, \dots, j)$ in length units, respectively; where *i* is the number of receivers involved, and *j* is the number of satellites being tracked; p_s^r , ϕ_s^r are pseudo-range and carrier-phase observable; ρ is the geometric distance, while the antenna phase center, tidal displacements, relativity, phase windup corrections should be applied to p, ϕ before ρ becomes unassociated with the frequency as well as the system sys; $t_{r,sys}$ is the clock offset corresponding to receiver r for the system sys; t^s denotes the clock offset of satellite s; T_r^z is the zenith tropospheric delay that can be converted to slant with the mapping function α ; b_r^k is the receiver code bias with the corresponding wavelength λ .

Note that the satellite code bias is not presented since the corresponding DCB products of IGS are held fixed in this study. Regarding the receiver code bias term b_r^k , the coefficient e_{sys} is related to the satellite system [Lou et al. (2015), Liu et al. (2017)]

$$e_{sys} = \begin{cases} 0, CDMA (G, C, E) \\ 1, FDMA (R) \end{cases}$$
(2)

since for the CDMA based systems, the receiver code bias is identical for different satellite, and thus it is absorbed in the receiver clock $t_{r,sys}$. While for the GLONASS with FDMA signal, the receiver code bias b_r^k is satellite frequency number k related, in which $k = (-7 - 4 - 3 \cdots 0 \cdots 6)$ (http://www.glonasscenter.ru).

Regarding the dynamic model for these parameters, it is typically assumed that the clocks for receiver and satellite are estimated as white noise, b_r and N_r are constant parameters, while the troposphere is regarded as random walk which satisfies

$$T_{r,n+1}^{z} = T_{r,n}^{z} + \omega_{t^{Z},n} \tag{3}$$

in which $\omega_{t^{Z},n}$ is a random process with

$$E\left(\omega_{t^{Z},n}\right) = 0 D\left(\omega_{t^{Z},m},\omega_{t^{Z},n}\right) = Q_{t^{Z}} \cdot \delta_{m,n}$$

$$(4)$$

and $\delta_{m,n}$ is the Kronecker delta function, which is zero except when m = n, then $\delta_{n,n} = 1$ and Q_{tz} is set as $15mm/\sqrt{h}$, typically.

2.2 Oscillator Noise Model

Since every oscillator, atomic or crystal, deviates from the ideal or "true" time in an abstract sense except as we may choose to define it [Allan (1987)]. The satellite clock bias $t^s(n)$ of a given epoch n emerged from the offset between the instantaneous reading of the clock $t_r^s(n)$ and the reference clock $t^0(n)$, can be expressed as

$$t^{s}(n) = t^{s}_{r}(n) - t^{0}(n) = (t^{s}_{0} + \int_{0}^{n} \frac{f^{s}(t)}{f^{0}} dt) - (t_{0} + \int_{0}^{n} \frac{f^{0}}{f^{0}} dt) = (t^{s}_{0} - t_{0}) + \int_{0}^{n} \frac{f^{s}(t) - f^{0}}{f^{0}} dt = a_{0} + \int_{0}^{n} y(t) dt$$
(5)

where t_0^s and t_0 is the clock reading of the oscillator and the "true" clock of a selected reference epoch; $f^s(t)$ is the instantaneous frequency of the oscillator; f^0 is the nominal frequency; a_0 is the clock bias at the reference epoch; y(t) is by definition the instantaneous frequency deviation.

Though the satellite clock term t^s is typically estimated as a white noise parameter, the sophisticated clock model is still a requirement for a more robust solution of real-time GNSS applications. Generally, the instantaneous oscillator frequency is assumed to be of the form

$$f^{s}(t) = f^{0} + df + \sum_{m=1}^{N-1} \frac{f^{(m)}}{m!} \cdot t^{m} + \varepsilon(t)$$
(6)

in which, df is the mean frequency bias from the nominal frequency f_0 ; $f^{(m)}$ is the *m*th-order frequency drift, that dominates the long-term instability; $\varepsilon(t)$ is the random frequency errors, that describe the short-term instability [Lindsey (1976)]. Thus from Eq. (5) and Eq. (6) the instantonous frequency domination reads

Thus from Eq. (5) and Eq. (6), the instantaneous frequency deviation reads

$$y(t) = \frac{df}{f^0} + \sum_{m=1}^{N-1} \frac{f^{(m)}}{f^0 \cdot m!} \cdot t^m + \frac{\varepsilon(t)}{f^0}$$
(7)

Substituting Eq. (7) into Eq. (5), we derive the general model of clock error

$$t^{s}(n) = a_{0} + \frac{df}{f^{0}} \cdot n + \sum_{m=2}^{N} \frac{f^{(m-1)}}{f^{0} \cdot m!} \cdot n^{m} + \int_{0}^{n} \frac{\varepsilon(t)}{f^{0}} dt$$
(8)

The rearranging of terms and a new notation yields the clock error

$$t^{s}(n) = a_{0} + a_{1} \cdot n + \frac{a_{2}}{2} \cdot n^{2} + \sum_{m=3}^{N} \frac{f^{(m-1)}}{f^{0} \cdot m!} \cdot n^{m} + r(n)$$
(9)

where a_0 is defined as Eq. (5); a_1 is the frequency bias; a_2 is the frequency drift; r(t) is the integrated random fractional frequency error [Wübbena (1988)].

By selection the proper degree, i.e., N, in the model (8) and (9), different dimensional algorithms may be derived. As demonstrated by [Shmaliy et al. (1999)], three-dimension ($a_0 \ a_1 \ a_2$) is the most efficient for the Rubidium oscillator, while the quadratic term a_2 brings less good results for the Cesium-beam standards [Allan (1987)], i.e.,

$$t^{s}(n) = c^{T} \cdot a$$

=
$$\begin{cases} \left(1 \ n \ \frac{n^{2}}{2} \ 1\right) \cdot \left(a_{0} \ a_{1} \ a_{2} \ r(n)\right)^{T}, Rb \\ \left(1 \ n \ 1\right) \cdot \left(a_{0} \ a_{1} \ r(n)\right)^{T}, Cs \end{cases}$$
(10)

In addition, the dynamic equation of a satellite clock is written as

$$a_{n+1} = \Phi_a(\tau) \cdot a_n + \omega_a \tag{11}$$

with τ denotes the time interval of epoch n and n+1, and

$$\boldsymbol{\Phi}_{\boldsymbol{a}}(\boldsymbol{\tau}) = \begin{cases} \begin{pmatrix} 1 & \tau & \frac{\tau^2}{2} & 0\\ 0 & 1 & \tau & 0\\ 0 & 0 & 1 & 0\\ 0 & 0 & 0 & 1 \end{pmatrix}, Rb \\ \begin{pmatrix} 1 & \tau & 0\\ 0 & 1 & 0\\ 0 & 0 & 1 \end{pmatrix}, Cs \end{cases}$$
(12)

$$\boldsymbol{\omega}_{\boldsymbol{a}} = \begin{cases} \begin{pmatrix} 0 & 0 & 0 & \omega_r \end{pmatrix}, Rb \\ \begin{pmatrix} 0 & 0 & \omega_r \end{pmatrix}, Cs \end{cases}$$
(13)

3 Satellite clock

3.1 Temporal Statistics

In section 2, both the observable model and the dynamic model for multi-GNSS satellite clock estimation have been presented. However, the temporal statistical property of the satellite clock terms a is still under developed. Thus, to get the quantitative measure of the correlation for these terms, the repeatability of $(a_0 a_1 a_2)$ and the variograms of r are derived on the sub-daily regime in this section.

Since May. 3rd, 2015, GFZ begins to provide satellite clocks for G, R, C, E with a resolution of up to 30 sec which can be obtained from

http://doi.org/10.5880/GFZ.1.1.2017.002. Thus, near 2 years final precise multi-GNSS clock products from May 3rd, 2015 to Feb. 25th, 2017 were collected in this statistical study based on model (10). It is noted that before this study, preprocessing should be carried out to get rid of the difficulties of large data gaps, outliers as well as the jump discontinuity across the processing boundary [Huang et al. (2012)].

The repeatability of $(a_0 a_1 a_2)$ is obtained in terms of STD by the comparison of the 12-hour prediction and the fitted value, while the variograms of r are obtained based on a finite sample with a robust method suggested by [Cressie and Hawkins (1980)] as:

$$2\hat{\gamma}(h) \equiv \left\{ \frac{1}{|N(h)|} \sum_{N(h)} |r(s_i) - r(s_j)|^{1/2} \right\}^4 \left/ \left(0.457 + \frac{0.494}{|N(h)|} \right) \right.$$
(14)

where $2\hat{\gamma}(h)$ is the estimate of variogram; r(s) is the satellite clock fitting residual of model (10) for epoch s; the sum is over $N(h) \equiv \{(i, j) : s_i - s_j = h\}$ and |N(h)|is the number of distinct elements of N(h). From definition, the variogram $2\gamma(h)$ relates to the covariance by the following equations

$$2\gamma(s_i, s_j) = 2(C(0) - C(s_i - s_j)) \tag{15}$$

with $C(\cdot)$ the covariagram which is defined by covariance directly

$$C(s_i - s_j) = cov(r(s_i), r(s_j))$$

$$(16)$$

Thus, the quantity variogram describes the temporal behavior of r exactly.

In the following paragraph, we present the results from our experiment directly. Shown in Tab. 1 to Tab. 4 are the optimal dimension and the STD of $(a_0 \ a_1 \ a_2)$ for G, R, C, E, respectively, and the oscillator informations are also presented. It should be noted, that the unit for a_1 and a_2 is ns/h and ns/h^2 , respectively.

Consistent with model (10), the optimal dimension for Rubidium and Cesium oscillator is 3 and 2, respectively regardless the satellite constellation involved. The best prediction precision for $(a_0 \ a_1 \ a_2)$ can be found in GPS, which is around 0.5ns, 0.06ns/h and $0.003ns/h^2$, respectively, while, the repeatability decreases to 0.6ns and 0.09ns/h for GLONASS satellites. Concerning Galileo satellites, the performance of rubidium and hydrogen oscillator are roughly the same with a precision of 0.96ns, 0.09ns/h and $0.006ns/h^2$ on average. Though the oscillator

information of BDS is not available publicly, the 3-dimension algorithm is suggested for all its satellites from our experiment, and the statistics also suggests that the terms $(a_0 \ a_1 \ a_2)$ of BDS is less stable compared with that of G, R and E. Moreover it seems that the behavior of C06 is abnormal with a much larger STD value.

12h prediction STD Sat. BLOCK Osc. Dim. a_0 a_2 a_1 ns/h^2 nsns/hG01IIF Rb 3 0.5160.047 0.002 G02IIR Rb 3 0.378 0.041 0.004 3 G03IIF Rb 0.4490.0430.002G05IIR-M \mathbf{Rb} 3 0.7570.069 0.004 G06IIF \mathbf{Rb} 3 0.3600.033 0.002IIR-M 3 G07 Rb 0.4630.0470.004G08 $\mathbf{2}$ 0.455IIF \mathbf{Cs} 0.114G09 \mathbf{IIF} \mathbf{Rb} 3 0.4040.040 0.002G10IIF \mathbf{Rb} 3 0.5150.047 0.002G11IIR Rb 3 0.938 0.088 0.004 G12IIR-M Rb 3 0.8010.068 0.003G13 IIR Rb 3 0.5470.0540.004G14IIR Rb 3 0.369 0.0400.004 G15IIR-M \mathbf{Rb} 3 0.4350.040 0.002 G16IIR Rb 3 0.7120.0630.003 G17IIR-M Rb 3 0.6730.060 0.004G18 HR. Rb 3 0.6300.0540.003G19 HR. Rb $\frac{3}{3}$ 0.3680.039 0.003G20IIR Rb 0.3900.041 0.003G21 IIR Rh $\frac{3}{3}$ 0.5690.0540.0030.003 G22 IIR Rh 0.4650.046G23IIR Rb 3 0.3200.0350.003 2 G24IIF 0.6310.158Cs 3 0.002 G25IIF Rb 0.5320.050HF 3 0.002G26 Rb 0.6830.045G27IIF 3 0.002 Rb 0.5570.0513 G28 IIR \mathbf{Rb} 0.838 0.095 0.009 G29IIR-M 0.5900.003Rb 3 0.056G30IIF \mathbf{Rb} 3 0.5270.039 0.0023 G31IIR-M \mathbf{Rb} 0.5910.0550.004 G32 \mathbf{IIF} Rb 3 0.494 0.037 0.0020.5470.056 0.003 Ave.

Table 1 Optimal algorithm dimension and the STD for a_0 , a_1 and a_2 of GPS

In addition, Fig. 1 to Fig. 4 present the quantitative measure of the temporal correlation of r for each satellite. The dashed line denotes the experimental variogram computed from the fitting residual. Overall, the variogram increases along with the distance: it approaches to zero as h approaches to zero and becomes a constant as h increases up to several hours. The faster increase and larger value of variogram implies that the temporal behavior of residual r is less stable. For GPS satellites, it is observed that the behavior of the random fractional frequency error r depends on not only the oscillator type, but also the block type, and the BLOCK IIF satellite equipped with rubidium oscillator gives the best performance. While,

~			12h prediction STD		
Sat.	Osc.	Dim.	a_0 ns	a_1 ns/h	
R01	\mathbf{Cs}	2	0.715	0.106	
R02	Cs	2	0.525	0.080	
R03	Cs	2	0.719	0.102	
R04	Cs	2	0.654	0.092	
R05	Cs	2	0.627	0.083	
R06	Cs	2	0.805	0.152	
R07	Cs	2	0.717	0.100	
R08	Cs	2	0.521	0.077	
R09	Cs	2	0.438	0.073	
R10	Cs	2	0.723	0.113	
R11	Cs	2	0.414	0.055	
R12	Cs	2	0.579	0.084	
R13	Cs	2	0.691	0.109	
R14	Cs	2	0.830	0.143	
R15	Cs	2	0.699	0.101	
R16	Cs	2	0.539	0.080	
R17	Cs	2	0.449	0.065	
R18	Cs	2	0.384	0.059	
R19	Cs	2	0.674	0.083	
R20	Cs	2	0.537	0.076	
R21	Cs	2	0.412	0.061	
R22	\mathbf{Cs}	2	0.571	0.081	
R23	\mathbf{Cs}	2	0.651	0.104	
R24	\mathbf{Cs}	2	0.644	0.092	
	Ave.		0.605	0.090	

Table 2 Optimal algorithm dimension and the STD for a_0 and a_1 of GLONASS

Table 3 Optimal algorithm dimension and the STD for a_0, a_1 and a_2 of BDS

		12h prediction STD		
Sat.	Dim.	$egin{array}{c} a_0 \\ ns \end{array}$	$egin{array}{c} egin{array}{c} egin{array}$	$egin{array}{c} a_2 \ ns/h^2 \end{array}$
C01	3	1.319	0.183	0.012
C02	3	1.146	0.145	0.010
C03	3	1.597	0.085	0.017
C04	3	1.455	0.153	0.022
C05	3	1.524	0.253	0.015
C06	3	5.362	0.761	0.062
C07	3	1.803	0.092	0.023
C08	3	1.823	0.172	0.017
C09	3	1.218	0.125	0.013
C10	3	1.633	0.123	0.019
C11	3	1.008	0.078	0.005
C12	3	0.835	0.067	0.005
C13 / C15	3	1.092	0.126	0.022
C14	3	0.686	0.055	0.004
Ave.		1.607	0.172	0.017

	Osc.	Dim.	12h prediction STD		
Sat.			$egin{array}{c} a_{0} \ ns \end{array}$	$egin{array}{c} a_{1} \ ns/h \end{array}$	$egin{array}{c} a_2 \ ns/h^2 \end{array}$
E01	Hy	3	0.351	0.0293	0.002
E02	Hy	3	0.319	0.0255	0.002
E08	Hy	3	0.784	0.0679	0.004
E09	Hy	3	0.861	0.0804	0.006
E11	Rb	3	0.744	0.0676	0.004
E12	Hy	3	0.857	0.0856	0.006
E14	Hy	3	1.086	0.1188	0.009
E18	Hy	3	1.694	0.1633	0.012
E19	Hy	3	0.972	0.0918	0.005
E22	Rb	3	0.987	0.0949	0.006
E24	Hy	3	2.178	0.2096	0.014
E26	Hy	3	0.805	0.0743	0.005
E30	Hy	3	0.859	0.0797	0.005
	Ave.		0.961	0.091	0.006

Table 4 Optimal algorithm dimension and the STD for a_0 , a_1 and a_2 of Galileo



Fig. 1 Variogram from satellite clock fitting residual of model (10) for GPS

the performance of Galileo satellite clock random error r is between that of GPS BLOCK IIF and BLOCK IIR satellite equipped with rubidium oscillator. The comparison between GLONASS and BDS suggests that BDS performs slightly better in terms of r with a longer correlation time. Moreover, from Tab. 3 and Fig. 3, the oscillator anomaly for C06 can be recognized.



Fig. 2 Variogram from satellite clock fitting residual of model (10) for GLONASS



Fig. 3 Variogram from satellite clock fitting residual of model (10) for BDS $\,$



Fig. 4 Variogram from satellite clock fitting residual of model (10) for GALILEO

From the above analysis, a spherical model is selected for the parametric variogram model which is summarized with Eq. (17) and Tab. 5:

$$\gamma(h) = \begin{cases} 0 & , \quad h = 0; \\ c_0 + c_s \left\{ (3/2)(h \mid a_s) - (1/2)(h \mid a_s)^3 \right\} & , \quad 0 < h < a_s; \\ c_0 + c_s & , \quad h \ge a_s. \end{cases}$$
(17)

As an illustration, the model variogram based on Eq. (17) and Tab. 5 is also plotted in the corresponding figure with solid black line. For more details concerning variogram, we refer to [Cressie (1993)] and [Shi et al. (2012)].

3.2 Discussion

Before the solution of the GNSS satellite clock estimation based on Eq. (1) and Eq. (11), there are actually more issues left to be solved, i.e., the rank deficiency

	Sys.	c_0	c_s	a_s
	IIF - Rb	0	0.005	15000
\mathbf{GPS}	IIR/IIR-M - Rb	0.0002	0.01	15000
	IIF - Cs	0.0002	0.03	4500
	GLONASS	0	0.032	10000
	BDS	0	0.023	15000
	Galileo	0	0.0075	15000

Table 5 Values of c_0 , c_s and a_s for G, R, C, E

raised from the linear dependency of clock and code bias terms. To reduce this datum deficiency problem, condition Eq. (18) is introduced in this study to make model (1) and (11) uniquely solvable.

$$0 = \boldsymbol{u} \cdot \boldsymbol{a}_{i}^{sys} \tag{18}$$

with $\boldsymbol{u} = (1 \ 1 \ \cdots \ 1)^T$. The subscript $i \in (0 \ 1 \ 2)$ and $sys \in (G \ R \ C \ E)$, thus \boldsymbol{a}_i^{sys} stands for the vector of a_i of different satellites for sys. As the rank deficiency problem is beyond the scope of this contribution, for more details we refer to the work of [Blaha (1982)] and [Gu et al. (2013)]. In the former work, the generally concept of free network solution is presented, while the application of this concept in the GNSS data processing is discussed in the later one.

Moreover, as indicated by Eq. (11) to Eq. (13), the terms $(a_0 a_1 a_2)$ are assumed as constant, while, the clock bias a_0 is indeed a constant by definition, the frequency bias and drift parameters are not strictly constant, but subject to random fluctuations. These variation effects on the frequency bias and drift are analyzed based on the Allan and Hadamard variance respectively. Once the Allan and Hadamard variance obtained, a different oscillator noise model may derived as presented by [Stein (1989)] and [Huang et al. (2013B)].

However in this contribution, we chose to follow the strategy of [Hauschild and Montenbruck (2009), Hauschild (2010)], and treat the frequency bias and drift as constant, and compensate the deviations of the system model from the truth with a process noise parameter r since that: this oscillator noise model is built for real time clock estimation with a regional network, thus the continuous data length of each satellite is typically less than 12 hours, in which case, the terms $(a_0 a_1 a_2)$ can be safely regarded as constant as demonstrated in the following sections.

4 Experiment

4.1 Software Development and Data Collection

To cope with the upcoming or available new and modernized GNSS signals, especially, the multi-frequency data processing, real-time applications, as well as the GNSS/INS integrated navigation, we launched the FUSING (FUSing IN Gnss) project. At present, the FUSING software is capable for the precise orbit determination (POD) of GNSS, high-frequency satellite clock estimation, ionosphere and troposphere modeling as well as multi-frequency precise positioning. In this study, we have developed a new module with the oscillator noise model described



Fig. 5 Distribution of the 18 regional GNSS stations; points in red represent stations capable of tracking G, R, C and one station capable of tracking G, R in black



Fig. 6 Distribution of the 80 global GNSS stations; points in red represent stations capable of tracking G, R, C, E and four stations capable of tracking G, R, E in black

in section 3 into the FUSING software. In the following demonstration, the new module is assessed with both regional and global networks for simulated real-time multi-GNSS G, R, C, E satellite clock estimation. And for comparison purpose, the traditional approach based on white noise model is also presented.

Shown in Fig. 5 and Fig. 6 are the distribution of 18 and 80 stations for a regional and global network, respectively. For the regional experiment, the data is collected by the crustal movement observation network of China (CMONOC) from DOY 153 to 182, 2016 with an interval of 30 s. It is a pity that the Galileo is not valid in the regional experiment, while, for the global experiment, the data is collected by the multi-GNSS experiment campaign (MGEX) from DOY 001 to 030 of 2017 with an interval of 30 s. Concerning the processing strategy of these experiments, please refer to Tab. 6. As indicated in Tab. 6, the satellite clock products estimated with no less than three stations are all included in the statistical analysis, i.e., the result is also affected by the convergence. In addition, we have ignored the effect of GLONASS inter-frequency code bias in this study.

4.2 Regional Experiment

4.2.1 Satellite discontinuity

To begin with the evaluation of the regional experiment, Tab. 7 illustrates the typical satellite tracking continuity averaged against the corresponding orbit period for a simulated network evenly distributed in China. Though the Galileo is not

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Item	Strategy
Observable	Ionosphere free combination Due he describe $C \to C$ is an expression static formation filter
Solution	Day by day with G, R, C, E in one square root information filter
Station coordinate	Fixed with static PPP solution averaged over one week using PANDA
Satellite orbit	Fixed with GFZ final multi-system orbit products
PCO/PCV	Corrected
Tides	Corrected
Phase windup	Corrected
Receiver clocks / ISBs	One clock for each system as white noise parameter, and no ISBs are involved
Satellite clocks ON	Estimated as Oscillator Noise parameters of model (11) for each satellite, and a_1 and a_2 are reset day by day
W N	Estimated as White Noise parameter for each satellite
Troposphere	Prior model with remaining estimated as a random walk process
Ambiguity	Constant for each continuous arc
Elevation angle cutoff	7°
Weighting	8 dm for pseudo-range and 8 mm for phase, low elevation observable and outliers are down-weighted
Quality control	Satellite clocks are generated only for those with at least three stations tracked simultaneously GFZ final multi-system clock products

Table 6 Details of the satellite clock estimation strategy

tracked by CMONOC up to now, the result is still presented in Tab . 7. It is concluded that the gap between the adjacent arcs is $6 \sim 9hours$ for the MEO satellites over China. Moreover, it should be noted that the actual tracking arc is subject to the status of satellites as well as stations. As an example, Fig. 7 presents the visibility of satellite G01 from different stations for DOY 153, 2016 with a cutoff angle of 7° from our experiment, while, the tracking arc for all satellites of this day is plotted in Fig. 8.

 ${\bf Table \ 7} \ {\bf The \ typical \ satellite \ tracking \ continuity \ averaged \ against \ the \ corresponding \ orbit \ period \ for \ a \ network \ evenly \ distributed \ in \ China$

Sys.		Period [Day]	Tracked [h]	Gap [h]
GPS		1	9.19	6.08
GLONASS		8	6.51	4.74
PDS IGSO		1	20.03	3.30
DDS	MEO	7	6.13	7.70
Galileo		10	6.66	9.01

4.2.2 Clock evaluation

Differenced with the GFZ final clock products, Fig. 9 gives the statistics of the regional satellite clock solution averaged over each constellation G, R, C from DOY 153 to 182, 2016. Since there may be more than one arc for each satellite during each day in the regional experiment, the STD in Fig. 9 is the mean value averaged over all STDs for different arcs which can be expressed as

$$STD = \frac{\sum_{i=1}^{m} STD_i}{m} \tag{19}$$

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Fig. 7 Visibility of satellite G01 from different stations for DOY 153, 2016 with a cutoff angle of 7°



Fig. 8 The typical satellite tracking arc of GPS, GLONASS and BDS with CMONOC for DOY 153, 2016

where $STD_i = \sqrt{\frac{\sum_{j=1}^n (dt_j - \bar{dt}_i)}{n-1}}$ and $\bar{dt}_i = \frac{\sum_{j=1}^n dt_j}{n}$ are the STD and mean clock bias for arc *i*, respectively. The RMS is derived over all the samples, i.e., regardless the arcs of the satellite for each day. Note that both the samples before and after convergence are included in the STD and RMS calculation, i.e., the convergence periods after the data discontinuity are also included in the evaluation of the clock solutions based on oscillator noise model and white noise model. It is obvious that when only regional stations involved, the performance of the satellite clock estimates can be improved significantly constrained with the oscillator noise model presented in section 3.1. Compared with the results based on white noise model,



Fig. 9 Statistics of the satellite clock estimation of GPS, GLONASS and BDS with CMONOC from DOY 153 to 182, 2016

the improvement is 44.4% and 12.1% on average for STD and RMS, respectively. Furthermore, the comparison among the G, R, C solutions suggests that the oscillator noise model is less efficient for the BDS solution. This is mainly due to the fact that the constellation of BDS mainly consists with the GEO and IGSO satellites, which suffer less data discontinuity. And for the same reason, the BDS satellite clock presents the best results with a STD of 0.34ns and RMS of 1.36ns.

Furthermore, Fig. 10 illustrates the 95% and 68% confidence level convergence series of a 6-hour pass during the initialization averaged over all the samples of the regional experiment. Together with Fig. 9, it is concluded that the improvement in the statistics, i.e., STD and RMS, are mainly attributed to the efficiency of the oscillator noise model during the first few hours for G, R. And not surprising, the improvement in convergence of BDS is not as evident as that of G, R. As presented in Tab. 8, the statistics based on samples after convergence, i.e., 3 hours later since the start of each satellite tracking arc, suggested that the precision of the solutions with different models are roughly the same after convergence. While, though the improvements after convergence are rather limited, it is still expected that the solution with oscillator noise model is more resistant to the gross error, especially during the end of each arc for satellite with a low elevation angle. This statement is confirmed with the clock difference series of G22, R12 and C12, respectively, for DOY 159, 2016 in Fig. 11.

Table 8 Statistics of the satellite clock estimation of GPS, GLONASS and BDS based on samples after convergence, i.e., 3 hours later since the start of each satellite tracking arc with from DOY 153 to 182

Sys.		GPS	GLONASS	BDS
STD [ns]	OM WN	$0.19 \\ 0.19$	$0.22 \\ 0.31$	$0.20 \\ 0.20$
RMS [ns]	OM WN	$1.36 \\ 1.35$	$3.33 \\ 3.54$	$1.23 \\ 1.29$



Fig. 10 The 95% and 68% confidence level convergence series of a 6-hour pass averaged over DOY 153 to 182, 2016 for GPS, GLONASS and BDS with CMONOC. While, the background scatter are the satellite clock convergence series for each individual arc for \mathbf{ON} (gray) and \mathbf{WN} (orange)



Fig. 11 Clock difference series of G22, R12 and C12, respectively, for DOY 159, 2016. The black dot denotes the solution with oscillator noise model, while, the red dot denotes the solution with white noise model

4.3 Global Experiment

Though this contribution aims at the satellite clock estimation with regional stations constrained with oscillator noise model, a global experiment was also carried out for comparison purpose.

In a similar manner of the regional experiment, presented in Fig. 12 are the STD and RMS averaged over DOY 001 to 030, 2017 for G, R, C, E with the global network. Generally speaking, the precision is at the same level regardless the clock model involved as demonstrated by Fig. 12. Apparent improvement can only be found in Galileo satellites with the STD decreased from 0.28ns to 0.19ns. A detailed analysis reveals that even with the stations in Fig. 6, the satellites E14 and E18 still suffer significant data discontinuity, i.e., 7.4h/day for our experimental



Fig. 12 Statistics of the satellite clock estimation of GPS, GLONASS, BDS and Galileo with MGEX from DOY 001 to 030, 2017

period. While the STD and RMS are improved to 0.19ns and 0.96ns, respectively, once we exclude E14 and E18 in the statistics. Though the root cause of these abnormal remains unknown, it is most likely due to the orbit recovery of these two satellites: E14 and E18 [Männel (2017)]. This result again confirms that the clock estimation in the existence of data discontinuity is expected to be benefited from the oscillator noise model promoted in this study.

Overall, the STD for the satellite clock with global network is about 0.2ns for G, C, E based on the whole daily session. While after convergence, the STD is improved to about 0.05ns as demonstrated in the previous studies of involving authors in [Gong et al. (2017)]. Compared with the results of G, C, E, the STD and RMS of GLONASS are much worse with a value of 0.36ns and 1.96ns, respectively. This is mainly due to the fact that the GLONASS inter-frequency code bias has not been taken into consideration in this study, while they have been estimated in the GFZ products. In addition, since the clock estimates depend on the type of observations and no official differential code biases (DCB) for Galileo are available, the solutions of this study and GFZ based on different stations may result in systemic bias [Gong et al. (2017)].

Furthermore, Fig. 13 presents the daily repeatability of frequency bias and frequency drift, i.e., a_1 and a_2 for each satellite of G, R, C, E with MGEX averaged over DOY 001 to 030, 2017. While the statistics of a_0 is not presented, the repeatability of a_0 is subjected to the daily boundary of the initial clock involved. It is concluded from Fig. 13, that the oscillator stability of Galileo is comparable with that of GPS, which is about 0.05ns/h for a_1 and $0.002ns/h^2$ for a_2 , while the STD of BDS is extremely larger than that of G, R, C. Besides the quality of the BDS oscillator itself, another possible reason is that the clock estimates may absorb the daily boundary of the BDS orbit. Additionally, it is noted that the frequency bias of Galileo is much larger than that of G, R, C.



Fig. 13 Daily repeatability of the oscillator noise model parameters a_1 (in black) and a_2 (in red) for GPS, GLONASS, BDS and Galileo with MGEX. The STD presented in this figure is the averaged value of the statistics for each satellite from DOY 001 to 030, 2017. Note that the satellite E22 has been removed in the statistic of Galileo as outlier



Fig. 14 Execution time per epoch and parameter number for DOY 001, 2017. The black line and dot denote the results of clock solution with oscillator noise model, while, the red line and dot denote the results of clock solution with white noise model

Besides the precision, the computational efficiency presents another key factor in satellite clock solution. Thus, Fig. 14 illustrates the processing efficiency and the parameter number of the satellite clock estimation with oscillator noise model and white noise model denoted in black and red, respectively. The analysis is conducted on a Centos 7 personal laptop with an Intel Core-3.3 GHz processor and 16 GB memory. Overall, more data processing effort is required with the new model due to the extra 154.2 parameters on average. Whereas, this extra time, i.e., 0.19s per epoch, is negligible and the total execute time per epoch, i.e., 1.61s, can fully satisfy the requirement of real-time precise satellite clock estimation. The processing efficiency here is comparable with our previous work for high efficient undifferenced GNSS data processing [Gong et al. (2017)].

5 Conclusion

The purpose of this contribution is to explore the potential of the oscillator noise model to constrain the real-time satellite clock estimation in the existence of the data discontinuity, i.e., regional network solution. Thus, the temporal behavior of the satellite oscillator was studied for GPS, GLONASS, BDS and Galileo based on the GFZ final multi-GNSS clock product from May. 3rd, 2015 to Feb. 25th, 2017. Different oscillator noise model were promoted and assigned to different satellites according to its oscillator type as well as the block type. In this oscillator noise models, the clock bias, frequency bias and frequency drift were estimated as constant, and the deviations of the model from the truth are further compensated with a random walk parameter.

By employing this model into the simulated real-time multi-GNSS clock solution, estimates of GPS, GLONASS and BDS satellite clock were derived with the CMONOC network with 18 stations evenly distributed over China from DOY 153 to 182, 2016. The comparison between the solution constrained with the oscillator noise model and the traditional solution with white noise demonstrates the efficiency of the proposed model with an improvement of 44.4% and 12.1% on average for STD and RMS, respectively. Furthermore, a global experiment was performed based on the MGEX network with 80 stations distributed worldwide from DOY 001 to 030, 2017. Though the precision of the clock estimates are comparable for these solutions with different models for GPS, GLONASS and BDS, the STD of Galileo clock improves from 0.28ns to 0.19ns due to that, the satellites E14 and E18 still suffer significant data discontinuity, i.e., 7.4h/day.

It should be mentioned, although rarely happened, that the satellite clock slip of BDS is detected from our experiment, and special attention should be given to this phenomenon when employing the oscillator noise model in the satellite clock estimation. Moreover, though the oscillator noise model seems redundant once the satellites are being tracked continuously, it still deserves to be studied: firstly, there are cases that the clock estimates should be predicted due to the connection problem of the real-time PPP users; secondly, the stability and smoothness of clock estimates are essential in the GNSS based timing and synchronization; and finally, with the development of the GNSS oscillators in the future, more precise constrains can be applied, which has the potential to further improve the performance.

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