2014

New Perspectives on the Bishop Tuff from Zircon Textures, Ages and Trace Elements

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RECEIVED MAY 23, 2013; ACCEPTED NOVEMBER 11, 2013

We present zircon textural, trace element and U-Pb age data obtained by secondary ion mass spectrometry (SIMS) (SHRIMP-RG: sensitive high-resolution ion microprobe, reverse geometry) from 15 stratigraphically controlled Bishop Tuff samples and two Glass Mountain (GM) lava samples (domes OD and YA). Bishop zircon textures divide into four suites: (a) dominant sector-zoned grains, with (b) subordinate grains showing bright rims [lower U, Th, rare earth elements (REE)] in CL imaging, (c) sparse GM-type grains (texturally similar to zircons from GM dome YA) and (d) sparse Mesozoic xenocrysts from Sierran granitoid country rocks. All Bishop zircons from suites (a)-(c)combined have a weighted mean age of 777.9 $\pm\,2.2\,ka$ (95% confidence) and a tail back to \sim 845 ka. Our eruption age estimate using the weighted mean of 166 rim ages of $766.6 \pm 3.1 \text{ ka}$ (95%) confidence) is identical within uncertainty to published estimates from isotope-dilution thermal ionization mass spectrometry (ID-TIMS) (767.1 $\pm 0.9 \, ka, 2\sigma$) and ${}^{40}Ar / {}^{39}Ar$ (767.4 $\pm 2.2 \, ka,$ 2σ) techniques, the latter using the 28.172 Ma age for the Fish Canyon sanidine standard. We estimate also an eruption age for GM dome YA of $862 \pm 23 \, ka$ (95% confidence), significantly older than the currently accepted 790 ± 20 ka K-Ar age. The oldest zircon cores from late-erupted Bishop material (including those with GM-type textures) have a weighted mean age of 838.5 ± 8.8 ka (95% confidence), implying that the Bishop Tuff system was active for only ~ 80 kyr, and had effectively no temporal overlap with the GM system. Trace element variations in Bishop zircons are influenced strongly for many elements by sector zoning, producing up to $3 \times$ concentration differences between sides and tips within the same growth zone. Contrasting trends in molar $(Sc + \Upsilon + REE^{3+})/P$ ratios between sides and tips indicate contrasting mechanisms of substitution in different sectors of the same crystal. Concentrations of Ti in tips are double those in the sides of crystals, hindering applicability of the Ti-in-zircon thermometer, in addition to variations inherent to the 0.15-0.67 range in values proposed for aTiO₂. The bright-rim portions of grains are inferred to have crystallized from the same magma as that which generated the bright rims seen under cathodoluminescence or back-scattered electron imaging on quartz and feldspar, respectively. This less evolved, slightly hotter magma invaded the deeper parts of the chamber represented in the late-erupted northern units possibly up to ~ 10 kyr prior to eruption, but invaded shallower levels only very shortly before eruption, as shown by our textural information and previously proposed from the sharp delineation of quartz bright rims. By obtaining a large number of analyses from zircon separates that systematically cover the entire Bishop Tuff eruption sequence we can produce an eruption age estimate using SIMS to the same precision and accuracy as ID-TIMS and ⁴⁰Ar/³⁹Ar techniques.

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KEY WORDS: zircon; U–Pb dating; Bishop Tuff; supervolcano; sector zoning; trace elements

INTRODUCTION

The Bishop Tuff has become an iconic representative of the largest group of felsic explosive volcanic eruptions ('supereruptions') and a central focus of debates on the origin, assembly and compositional zonation of large silicic magma bodies. In particular, following the seminal studies of Hildreth (1977, 1979) and their setting into regional and global perspectives (Hildreth, 1981, 2004), several key areas have remained controversial. The central inference of Hildreth (1977, 1979) of a pre-eruptive, vertically and laterally continuous >600 km³ body of moderate- to highsilica rhyolite, vertically zoned with respect to composition (particularly trace elements), crystal and volatile contents, and temperature has remained (Dunbar & Hervig, 1992; Wallace et al., 1999; Anderson et al., 2000; Bindeman & Valley, 2002; Hildreth & Wilson, 2007; Roberge et al., 2013, but see also Gualda et al., 2012a).

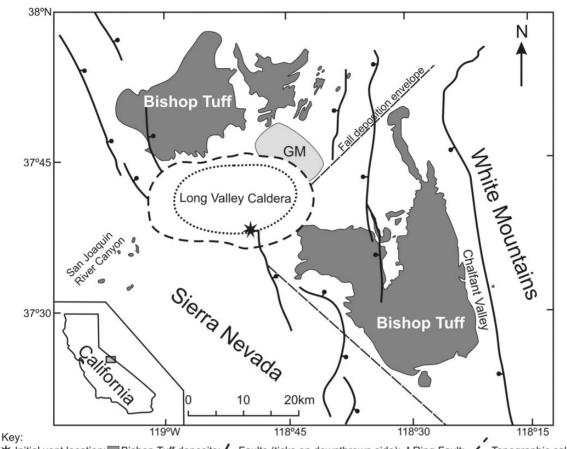
There remains vigorous debate, however, about two aspects of this picture, which we consider here from the zircon perspective. The first concerns inconsistencies between estimates of the eruption age derived from ${}^{40}\text{Ar}/{}^{39}\text{Ar}$ experiments on sanidine, and between those estimates and values derived from U-Pb dating of zircons using highprecision isotope-dilution thermal ionization mass spectrometry (ID-TIMS). The second concerns the longevity of the Bishop magma chamber, the timing of its assembly and the magmatic regime, derived from various lines of evidence and modelling. To address these issues, we present a comprehensive suite of zircon age determinations and associated trace element data obtained by secondary ion mass spectrometry (SIMS). We have analyzed zircons from 15 stratigraphic horizons in the Bishop Tuff to trace the geochronological and trace-element development of the magma chamber as recorded in discrete zones of single crystals. Although single SIMS age determinations are less precise than those derived from single-zircon ID-TIMS techniques, the ability to create large datasets from discrete zones in single zircons from different pumice types and different eruptive units within the Bishop Tuff places powerful constraints on the pre-eruptive history of the magma body. Additionally, collection of trace-element data and the linkage to zircon textures as shown in cathodoluminescence (CL) imaging allow the overall petrogenetic context and broad timings of interaction of the zircon crystals with different magma compositions to be assessed.

THE BISHOP TUFF Geological background

The $>600 \text{ km}^3$ Bishop Tuff was erupted in association with the formation of Long Valley caldera in eastern California

(Fig. 1; Bailey et al., 1976; Hildreth, 1979, 2004; Wilson & Hildreth, 1997; Hildreth & Wilson, 2007). The eruption represented the culmination of ~4.5 Myr of magmatism and the rhyolite erupted in the Bishop Tuff event is greater in volume than all other felsic eruptions combined in the Long Valley area, before or since (Hildreth, 2004). The Bishop Tuff in its proximal outcrop area between the Sierra Nevada and the White Mountains (Fig. 1) consists of a sequence of plinian pumice fall deposits generated before, and coevally with, lobes of ignimbrite, together with voluminous co-eruptive caldera fill (Wilson & Hildreth, 1997). In these proximal areas where we (and all other workers concerned with pumice in the Bishop Tuff) have sampled, the fall deposits are found only in an $\sim 90^{\circ}$ sector east of the source; however, the overall blanket of fall deposits (plinian and co-ignimbrite) covers much of the western USA to >2000 km from the source (Izett et al., 1988). Marker horizons in the proximal fall deposits were used to divide them into nine units (F1-F9). In turn, interbedding relationships and recognition of other marker horizons were used to correlate between the fall and flow deposits and to estimate that the eruption occurred (with only one possible very short hiatus between F8 and F9) over a period of the order of 6 days (Wilson & Hildreth, 1997). Lithic assemblages in the fall deposits and ignimbrite serve to locate the initial vent in the southeastern sector of what became the caldera and then follow the subsequent concurrent development of venting along the northern and eastern portions of the ring fracture (Figs 1 and 2; Hildreth & Mahood, 1986; Wilson & Hildreth, 1997).

Prior to the Bishop Tuff eruption the Glass Mountain rhyolitic system was active from 2.2 Ma to 790 ka, producing more than 60 effusive and explosive eruptions, and building a lava dome complex with a surrounding fan of primary and reworked volcaniclastic material (Metz & Mahood, 1985, 1991; Metz & Bailey, 1993; Hildreth, 2004). This system has two subgroups: a 2.2-1.3 Ma sequence of chemically variable, exceedingly evolved, high-silica rhyolites, and a younger group (1.2 Ma to 790 ka) of chemically more uniform, highly evolved high-silica rhyolites, the latter similar in bulk composition to the early Bishop eruption products (Metz & Mahood, 1991; Hildreth, 2004). Glass Mountain rhyolites are dominantly crystal poor (<8% crystals), with similar mineral assemblages to the Bishop Tuff (Metz & Mahood, 1991). Earlier rhyolitic volcanism is represented by plinian pumice fall deposits, which are best exposed in the Blind Spring Valley area east of Glass Mountain but which also extend for hundreds of kilometres farther east (Sarna-Wojcicki et al., 2005). Younger parts of this sequence have been correlated with Glass Mountain activity, but older lapilli-bearing (hence probably within tens of kilometres of the vent) fall deposits have returned ⁴⁰Ar/³⁹Ar single-crystal sanidine ages of



🗱 Initial vent location; 🔳 Bishop Tuff deposits; 🖌 Faults (ticks on downthrown side); 🕻 Ring Fault; 🖊 Topographic caldera outline

Fig. 1. Map of the Long Valley area, eastern California, USA, after Hildreth & Wilson (2007). The topographic outline of the caldera is shown as a dashed line, with the ring-fault zone shown as a dotted ellipse inside the topographic margin. Regions of Bishop pyroclastic-flow deposition are shaded grey. The envelope enclosing precursory Glass Mountain (GM) lava domes is highlighted in light grey. The line marked 'Fall deposition envelope' marks the westerly limits of where Bishop fall deposits are found in this area.

between c. 2.2 and 2.8 Ma (Sarna-Wojcicki *et al.*, 2005). As such, these deposits apparently pre-date activity in the Glass Mountain system and came from sources as yet undefined within the Long Valley area.

Eruption age estimates

The age of the Bishop Tuff has long been of interest globally because of its proximity in time to the Matuyama Reversed to Brunhes Normal (M–B) paleomagnetic transition (e.g. Sarna-Wojcicki *et al.*, 2000; Mark *et al.*, 2012; Singer *et al.*, 2012). Two contrasting approaches have been used to assess the Bishop Tuff eruption age: U–Pb geochronology, both by ID-TIMS and SIMS techniques, and Ar–Ar geochronology.

Crowley *et al.* (2007) reported high-precision ID-TIMS U–Pb age determinations on a suite of zircons from a densely welded zone of the Bishop Tuff in the Owens River gorge. From their description we infer their sample to be from the dense-welded portion of IglEb (welding zone c of Wilson & Hildreth (2003); Fig. 2b). The zircons illustrated by Crowley *et al.* (2007) are dominated by a sectorzoned structure, often with an apparent central domain, and U concentrations ranging from under 1000 ppm to 2000 ppm in the central domain to over 5000 ppm in the pyramidal tips. These zircons yielded a weighted mean age estimate of $767 \cdot 1 \pm 0.9$ ka [2 σ , mean square of weighted deviates (MSWD) = 1.3] from 17 of 19 grains analysed.

U-Pb age determinations by SIMS have been reported by Reid & Coath (2000), Simon & Reid (2005) and Reid & Schmitt (2012). The ion microprobe uses a focused ion beam to sputter the target and has the advantage of allowing selection of particular domains in the zircon. However, the fine sampling scale of the ion microprobe greatly diminishes the sample size and there is a commensurate loss of precision. However, the issue of ion microprobe analysis is not simply the final determination of a bulk age, but rather the association of that age with

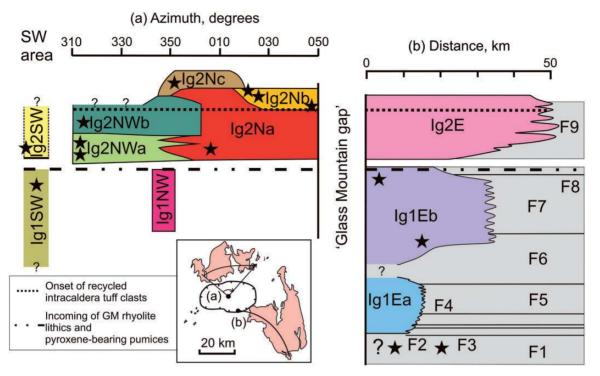


Fig. 2. Summary stratigraphy of the Bishop Tuff, after Hildreth & Wilson (2007). Locations for samples in this study are starred (see Table 1 for sample numbers and sites). (a) Hypothetical cross-section from NW to NE across the northern ignimbrite lobes as viewed from the centre of Long Valley caldera. (b) Schematic proximal to distal cross-section approximately along the line of Owens River Gorge, where the horizontal axis represents distance from the initial vent site of Hildreth & Mahood (1986).

discrete structures in the zircon. Reid & Coath (2000) and Simon & Reid (2005) determined that zones in the zircons showed a significant pre-eruptive history extending back hundreds of thousands of years. Although such prehistory should also have been apparent in the bulk ID-TIMS analyses, the samples analyzed were not the same and the possibility arises that there could be stratigraphic variations in the zircon age characteristics. Simon & Reid (2005) reported weighted mean crystallization ages (uncertainties reported as $1\sigma_{\rm m}$) of 811 ± 7 ka (MSWD = 3.4, 20 analyses on 14 grains) for interiors of grains from Ig2NW[b], 841 ±8 ka (MSWD = 0.9, n = 29) from rims of grains from IglEb, and 823 ± 14 ka (MSWD = 0.8, 22 analyses on 20 grains) from IglEb [also reported by Reid & Coath (2000)]. Pooled minimum ages (i.e. first-order estimates of eruption age) from the first two of these datasets are 753 ± 11 ka (n = 7) and 767 ± 30 ka (n = 4), respectively.

Closely similar eruption age estimates to those from U–Pb dating were obtained from 20–25 grain multi-crystal total fusion ${}^{40}\text{Ar}/{}^{39}\text{Ar}$ experiments on Bishop sanidines by Rivera *et al.* (2011) from the same sample (79G94) as used by Sarna-Wojcicki *et al.* (2000). From their descriptions, we infer this sample to be from ignimbrite packages Ig2Na or Ig2Nb of Wilson & Hildreth (1997). Rivera *et al.* (2011) derived a weighted mean age for the Bishop sanidines of 767.4 \pm 2.2 ka (2 σ , external errors), along with an

equivalent astronomically intercalibrated $^{40}\mathrm{Ar}/^{39}\mathrm{Ar}$ age for the Fish Canyon Tuff (FCT) sanidine standard (FCT) of 28:172 ± 0.028 Ma (2 σ , external errors).

The apparently excellent agreement between the U–Pb age determinations and the 40 Ar/ 39 Ar ages conflicts, however, with proposals for contrasting values for the 40 K decay constant and consequent FCT standard age values that would have the Bishop Tuff eruption age being significantly older (Renne *et al.*, 2010, 2011; Mark *et al.*, 2012). The eruption age estimate derived from the experiments of Sarna-Wojcicki *et al.* (2000) would shift to $778 \cdot 4 \pm 3 \cdot 7$ ka [2 σ , using the 28:305 \pm 0:036 Ma FCT age from Renne *et al.* (2010)] or $778 \cdot 1 \pm 3 \cdot 7$ ka [2 σ , after Renne *et al.* (2011): FCT age of $28 \cdot 294 \pm 0 \cdot 036$ Ma]. Further age determinations by Mark *et al.* (2012) on the same feldspar sample as used by Sarna-Wojcicki *et al.* (2000) and Rivera *et al.* (2011) were reported as $776 \cdot 4 \pm 0 \cdot 9$ ka (inferred by us to be 1σ) on the basis of the Renne *et al.* (2011) FCT value.

Magma accumulation timescales

A range of methods has been used to investigate timescales of magma accumulation and residence in silicic systems [for reviews see Hawkesworth *et al.* (2000, 2004)]. It should be noted, however, that the term 'residence time' may be interpreted in different ways. A distinction is drawn by us between the lifetimes of crystals (reflecting

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complex, possibly prolonged patterns of growth, stasis or dissolution) and the lifetime of the melt-dominant body in which they were suspended when quenched on eruption. Hildreth (1979) proposed that the crystal population in the Bishop pumices grew in situ and that the species present reflected the overall zonation in bulk-compositional characteristics. On this basis, the age range of the crystals and the lifetime of the melt-dominant body should be effectively the same. In contrast, studies of comparably sized systems (Bindeman et al., 2001; Charlier et al., 2005, 2008, 2010) show that crystal populations in deposits pre- and post-dating the climactic events are influenced by inheritance and recycling of antecrysts (from earlier crystallization events in the system) and xenocrysts (from local basement rocks). The age spectra of zircons within the Oruanui (New Zealand) pumices demonstrably bear no resemblance to the much shorter timescale for growth of the final melt-dominant body that was evacuated in the Oruanui eruption (Wilson & Charlier, 2009; Allan et al., 2013).

For the Bishop Tuff, at one extreme, timescales of the order of hundreds of thousands to a million years have been proposed for development of the parental magma body. Glass-feldspar Rb/Sr isochrons from Glass Mountain lavas (particularly those from ~1·1 Ma onwards), and pumices from the Bishop Tuff itself, yield model ages 300-1300 kyr older than notional eruption ages (Halliday et al., 1989; Christensen & DePaolo, 1993; Christensen & Halliday, 1996; Davies & Halliday, 1998). This protracted timescale was apparently supported by ⁴⁰Ar/³⁹Ar ages from melt inclusions in quartz (2.0-2.3 Ma; van den Bogaard & Schirnick, 1995), but these ages were subsequently attributed to excess Ar (Winick et al., 2001). Theoretical models of felsic magma genesis have also been used to estimate a million year timescale for development of the Bishop magma body (Fowler & Spera, 2010), but this approach has been suggested to be unrealistic on the grounds of thermal constraints (Fowler & Spera, 2011; Gualda & Ghiorso, 2011).

Studies focusing on the chronological information from zircons have, however, yielded shorter and contrasting timescales. Zircon in the Bishop Tuff is considered to be one of the earliest crystallizing phases (Hildreth, 1979) on the basis of its presence as inclusions within all other crystal species. It would thus be expected that the zircon age spectra should date the timescales of crystallization during the generation of the Bishop magma in the overall crystal-dominated roots of the magmatic system (the mush zone: Bachmann & Bergantz, 2004; Hildreth, 2004). There is, however, a distinct contrast between zircon mean ages of 811-841 ka, plus the ~160 kyr range of preeruptive ages from SIMS (Reid & Coath, 2000; Simon & Reid, 2005; Reid *et al.*, 2011), and the ~12 kyr maximum range proposed from ID-TIMS (Crowley *et al.*, 2007). This

contrast has been further highlighted by proposals (Gualda *et al.*, 2012*a*, 2012*b*; Pamukcu *et al.*, 2012) that the crystallization histories of the major mineral phases in the Bishop magma occurred over a much smaller range of magmatic temperatures (757–762°C: Gualda *et al.*, 2012*a*) than those inferred from Fe–Ti-oxide equilibria (714–818°C: Hildreth & Wilson, 2007) and over associated shorter timescales, more consistent with those implied by the results of Crowley *et al.* (2007).

Trace element patterns

The great variations in trace-element abundances within the Bishop pumices were first highlighted by Hildreth (1977, 1979) and attributed to liquid-state fractionation (Soret diffusion). This specific mechanism has since been demonstrated to be inappropriate (Lesher, 1986), and the variations are now considered to result from crystal fractionation processes (Michael, 1983; Cameron, 1984) in the mush zone, with extraction of variably fractionated liquids to form the melt-dominant body (Hildreth, 2004; Hildreth & Wilson, 2007), although debate still remains. The crystal-specific record shows, however, further complexities. The outer zones of crystals (together with their melt inclusions) from pumices in ignimbrite lobes north and northwest of the caldera rim show contrasting textures that correlate with trace-element patterns. Many quartz, sanidine and zircon crystals from these inferred latererupted materials have outer zones that are brighter under CL (quartz, zircon) or back-scattered electron (BSE) imaging (sanidine), and have volatile abundances and trace-element patterns indicative of crystallization from a less-evolved, inferred hotter melt (Dunbar & Hervig, 1992; Wallace et al., 1999; Anderson et al., 2000; Peppard et al., 2001; Wark et al., 2007; Reid et al., 2011). Originally interpreted to reflect sinking of the crystals from the 'normal' Bishop magma (represented by the earlier-erupted plinian fall deposits and eastern-side ignimbrite lobes) into the contrasting magma (Wallace et al., 1999; Anderson et al., 2000; Peppard et al., 2001), these features are now interpreted to reflect ingress into the Bishop chamber of the contrasting magma type (e.g. Wark et al., 2007; Reid et al., 2011). The boundaries between the 'normal' (i.e. crystals typical of earlier-erupted plinian fall deposits and eastern-side ignimbrite) cores and contrasting rims to the crystals are often sharp to <10 µm. Such abrupt changes have been linked to introduction of a contrasting magma that may have triggered the onset of the eruption itself (e.g. Wark et al., 2007; Reid et al., 2011), and/ or accompanied rapid rise of the host magma to shallow depths (Thomas et al., 2010; cf. Wilson et al., 2012).

Trace-element patterns within Bishop zircons from 'early' versus 'late'-erupted material show wide variations (Hildreth, 1979; Mahood & Hildreth, 1983, from bulk crystal aliquots; Reid *et al.*, 2011 from single crystals) that match closely the variations seen in the quartz and

feldspar. As shown by Reid *et al.* (2011) and amplified in this study, however, most of the trace element contrasts arise from variations within the single crystals (Fig. 3). Outer rims, brighter under CL imaging, occur on virtually no grains in the early erupted material, but are frequent and up to tens of microns thick on grains from the lateerupted material. Melt compositions indicated by the bright-rim material or represented by melt inclusions in these rims were inferred or demonstrated to represent less-evolved and/or hotter melts (Wallace *et al.*, 1999; Anderson *et al.* 2000; Wark *et al.*, 2007). However, our work shows that this picture is neither as simple nor as uniform as implied by these earlier studies.

Sample suites for this study

Fifteen samples of the Bishop Tuff were selected for zircon extraction (Table 1; Fig. 2), plus two samples from Glass Mountain domes OD (K–Ar age of 1.33 ± 0.01 Ma) and

YA $(0.79 \pm 0.01 \text{ Ma})$ (ages from Metz & Mahood, 1985). The tuff was systematically sampled by us from the earliest fall unit (Fl) to the latest ignimbrite unit (Ig2Nc) to track any vertical and lateral variations present within the single zoned magma body. A range of pumice types was selected, from crystal-poor to crystal-rich in the 'normal' pumice spectrum, as well as some of the 'variant' very crystal-poor pumices (Hildreth & Wilson, 2007; see Table 1 for brief sample descriptions). Fourteen samples are from single pumice clasts (Bishop) or obsidian dome carapace material (Glass Mountain). For Bishop IglSW and Ig2SW samples, owing to a lack of pumices of adequate size, zircons were extracted from multiple lapilli-sized pumices. For the samples from fall Fl, bulk samples were taken, sieved (to exclude free crystals), and a pumice fraction was separated by water panning and hand-picking (to remove lithic fragments) before crushing for zircon extraction.

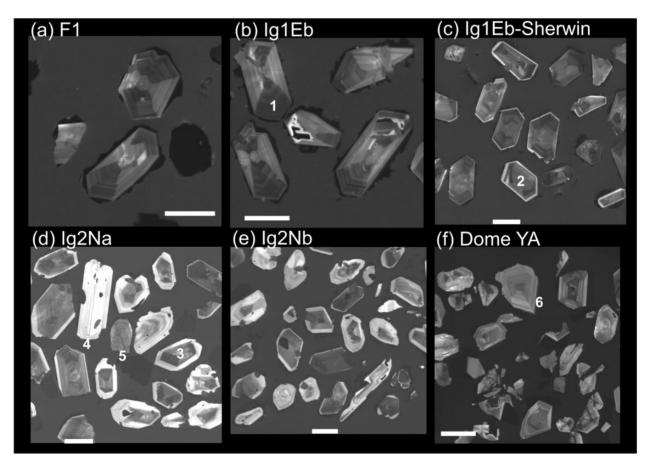


Fig. 3. Zircon cathodoluminescence (CL) textures in samples from the Bishop Tuff (a–e) compared with those of the youngest Glass Mountain dome [YA, (f); see text for details]. (a) BPI78-FI; (b) BP087-IglEb; (c) BP040-IglEb-Sherwin; (d) BP059-Ig2Na; (e) BP170-Ig2Nb (See Table 1 for sample descriptions). Types of zircons described in the text are as follows: 1, dominantly sector-zoned 'normal Bishop' zircons showing the characteristic oscillatory zoning; 2, 'normal Bishop'-type zircon with a narrow ($<10 \,\mu$ m) bright rim; 3, a wide bright rim overgrowing a 'normal Bishop'-type zircon; 4, a completely CL-bright zircon; 5, Triassic zircon showing oscillatory zoning with no sector zoning or bright rim; the crystal is fractured but still glass coated; 6, Glass Mountain-type oscillatory zoned crystal. Scale bar represents 100 μ m in each image.

Table 1: Samples studied

Sample no.	Unit	Туре	Location	Grid ref. (UTM)	Brief description
BP109	F1	multi	Owens River Gorge	03594, 41564	Bottom 12 cm bulk sample of F1 (pumices sieved to >1 mm)
BP178	F1	multi	Owens River Gorge	03550, 41607	Bulk pumice sample of 2.8-8 mm fraction from F1
BP087	lg1Eb	хр	Owens River Gorge	03594, 41564	Very fresh, large (26 cm) pumice crystal-poor to -moderate
BP098	lg1Eb	vxp-sw	Owens River Gorge	03594, 41564	Moderately vesicular vxp pumice with a sugary texture-swirly type
BP130	lg1SW	multi	Sotcher Lake	03174, 41661	Bulk sample of the largest single unoxidized pumices
BP040	lg1Eb	xr	Sherwin Grade	03483, 41578	Oxidized, crystal-rich, 'Adobe' type single pumice
	(Sherwin)				
BP216	lg2SW	multi	Pumice Butte	03169, 41572	Bulk sample of largest single pumice clasts from the upper parts or Ig2SW
BP118	lg2NWa	xr	Crestview	03241, 41812	Moderately vesicular crystal-rich pumice with a uniaxial fibrous textur
BP115	lg2NWa	vxr	Crestview	03241, 41812	Moderately vesicular very crystal-rich pumice with a planar fabric
BP059	lg2Na	xr	Bald Mountain	03381, 41879	Fresh, moderately vesicular, very crystal-rich single pumice
			Spring		
BP124	lg2NWb	xr	Aeolian Buttes	03175, 41925	Moderately vesicular crystal-rich single clast with a fibrous texture
BP170	lg2Nb	xr	McGee Lobe	03477, 41871	Core of a very fresh crystal-rich pumice with a crimped-fibrous textur
BP172	lg2Nb	xp-sw	McGee Lobe	03477, 41871	Very crystal-poor, moderately vesicular light grey swirly pumice
BP232	lg2Nb	xp-sw	McGee Lobe	03476, 41872	Pale grey, finely vesicular pumice with a foamy texture-swirly
BP206	lg2Nc	xr	Granite Mountain	03422, 41971	Crystal-rich single clast with an Adobe-like texture
BR228	Dome YA	xp lava	Dome YA, Glass	03409, 41846	Very crystal-poor obsidian
			Mountain		
BR226	Dome OD	xm	Dome OD, Glass	03449, 41877	Crystal-moderate to -rich lava
			Mountain		

Erupted units and types are from Hildreth & Wilson (2007).

ANALYTICAL TECHNIQUES

Zircon crystals for analysis were handpicked from a 60-250 µm sieve fraction pumice crush, which had been concentrated using standard density and magnetic separation techniques. Age determinations were made by SIMS techniques on the sensitive high-resolution ion microprobe, reverse geometry (SHRIMP-RG) instruments at the joint USGS-Stanford University facility (SUMAC) and at the Research School of Earth Sciences, Australian National University (ANU). Zircons were mounted in epoxy resin, polished to expose the cores of the grains, and imaged in reflected light, transmitted light and by CL on a scanning electron microprobe (SEM) at SUMAC and ANU. To minimize contamination by common Pb, the mounts were cleaned in detergent, petroleum spirits or dilute EDTA, and 1 M HCl, with intervening rinses in distilled water, then gold coated.

During analysis, the primary beam was rastered for 180 s over an area of $35 \,\mu\text{m} \times 45 \,\mu\text{m}$ prior to data acquisition to remove the gold coat and any possible surface contamination. Secondary ions were then sputtered with a 5–6 nA primary O_2^- beam focused to an ~25 $\mu\text{m} \times 35 \,\mu\text{m}$ spot.

The mass spectrometer was cycled through peaks corresponding to (bracketed species run at SUMAC only) $^{(172}$ Yb 16 O), 90 Zr $_{2}^{16}$ O, (180 Hf 16 O), 204 Pb, background, 206 Pb, 207 Pb, 208 Pb, (232 Th), 238 U, 232 Th 16 O and 238 U 16 O, with a total analysis time of 750-900s. Extended count times were used for ²⁰⁶Pb (20 or 30 s) and ²⁰⁷Pb (20 s) and six scans were run through the mass sequence. The concentration standard was MAD (Madagascar green: Barth & Wooden, 2010; 4196 ppm U) at SUMAC and SL13 (238 ppm U) at ANU. The U-Pb age standard throughout was R33 (420 Ma; ID-TIMS age from: http:// earth.boisestate.edu/isotope/analytical-capabilities/id-timsu-pb). Data reduction used SQUID 2 (Version 1.51; Ludwig, 2009), and data were plotted using Isoplot 3.76 (Ludwig, 2008). For single analyses uncertainties given by SQUID2 are reported here at 2σ , whereas for our grouped datasets uncertainties are given at the 95% confidence interval as generated in Isoplot. In these grouped datasets weighted means are calculated using the 'reject OK?' option engaged in Isoplot.

The ²³⁰Th correction for initial disequilibrium was calculated using the measured Th and U concentrations in the zircons from the SIMS measurements, together with melt Th/U values derived from laser ICP-MS data from glass separates for the Bishop Tuff samples (K. J. Chamberlain, unpublished data) and published instrumental neutron activation analysis (INAA) data for the two Glass Mountain samples (Metz & Mahood, 1991). For comparison, Crowley et al. (2007) used a single value of 2.81 from the mean of 12 melt inclusion analyses published by Anderson et al. (2000). A correction factor was then applied using $f = (Th/U_{zir})/(Th/U_{wr})$, following Schärer (1984). The presence of common Pb was evaluated by monitoring ²⁰⁴Pb and ²⁰⁷Pb/²⁰⁶Pb, and a correction was then applied using the recorded ²⁰⁷Pb/²⁰⁶Pb values and a common-Pb isotopic composition for the sample age from the average crust model of Stacey & Kramers (1975). This approach was tested using a regression of ²³⁸U/²⁰⁶Pb vs ²⁰⁷Pb/²⁰⁶Pb for 20 spots with >20% of the ²⁰⁶Pb attributable to common-Pb. This plot yielded a ²⁰⁷Pb/²⁰⁶Pb intercept of 0.82 ± 0.03 , consistent with use of the Stacey & Kramers (1975) value for the sample age (see Supplementary Data: Electronic Appendix 1). Analyses of grains with >10%common Pb at U >500 ppm were excluded from the dataset. However, some zircons from later-erupted pumices showed CL-bright zones that had <500 ppm U and accompanying higher proportions of common-Pb. For these 31 analyses a cut-off of 30% common-Pb was used and these data were handled separately.

SIMS analyses for trace elements were performed on the SHRIMP-RG at SUMAC using a 15-3 nA primary beam current and a focused $15 \,\mu\text{m} \times 12 \,\mu\text{m}$ spot. Analytical sites were selected to avoid inclusions visible in reflected light images and to representatively sample different CL zones, as well as crystal cores and rims. Trace element analytical spots were not superimposed on the U–Pb age dating spots (and vice versa), but care was taken to locate trace element analyses within the same growth zones of the zircon crystals as those that were age dated. The MAD zircon standard was used to calibrate trace element concentrations (Electronic Appendix 1).

RESULTS

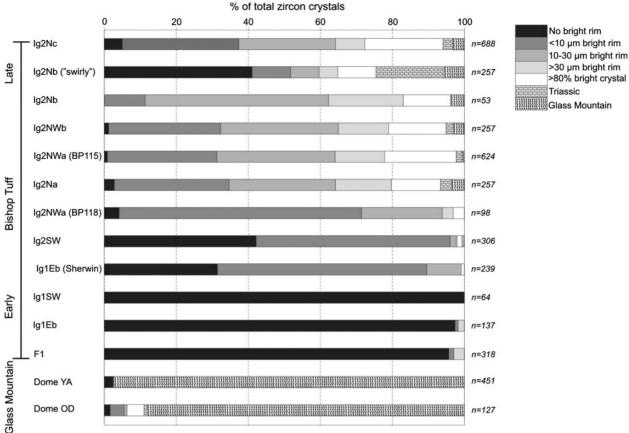
Our data suite consists of three quantifiable aspects of zircons: the textures as shown under CL imaging, the age data, and the trace-element data, which we first consider as separate entities.

Zircon textures

A total of \sim 3300 zircon crystals from the Bishop Tuff and Glass Mountain samples were characterized for their respective CL zoning patterns. When mounting the crystals, care was taken to lay prismatic crystals flat so that when they were ground down and polished, a representative cross-section was made of textures in the plane of the *a*- and *c*- or *b*- and *c*-axes. The patterns observed are grouped into four main suites (Figs 3 and 4; Table 2), as follows.

- (a) The 'typical' Bishop suite (referred to hereafter as 'normal Bishop') of euhedral prismatic grains, almost always with strongly developed sector zoning but lacking bright rims (Fig. 3a-c). This type overwhelmingly dominates the zircon populations from the fall deposit and easterly distributed ignimbrite, and is present in all samples from the northern ignimbrite lobes. All the grains illustrated by Crowley *et al.* (2007) belong to this type.
- (b) A subordinate suite having some proportion of the crystal that displays distinctly bright CL shades when compared with the normal Bishop type, and labelled 'bright-rim' type (Fig. 3c-e). The bright material most often forms discrete, sharply defined rims on grains, the cores of which have morphologies and CL patterns similar to those of the normal Bishop grains (above), but also comprises whole grains or continuously CL-zoned grains in samples from some of the northern lobes. The zircons with bright rims were further separated into four categories: <10 µm rim; $10-30 \,\mu\text{m}$ rim; rim >30 μm wide overgrowing a core of contrasting composition; >80% of the crystal showing bright CL shades and inferred to have largely or wholly crystallized from the 'bright-rim magma'.
- (c) A sparsely developed suite of oscillatory zoned, poorly shaped crystals with little or no visible indication of sector zoning (Fig. 3f). This type comprises virtually all of the grains from the YA dome sample (Fig. 4) and is thus termed 'Glass Mountain type'. Grains with similar characteristics are a sparse (0–6%) component of the zircon suites in pumices from northern and northwestern Bishop ignimbrite lobes.
- (d) A sparsely developed suite of concentrically zoned, irregularly shaped to euhedral grains that are commonly fractured, oscillatory zoned, and darker in CL than the Glass Mountain type (Fig. 3d, labelled as grain 5). Trace-element analyses, incorporating short counts for ²⁰⁶Pb (10 s) and ²³⁸U¹⁶O (5 s) for Pb and U concentrations, respectively, showed that such grains are xenocrystic (although glass coated, showing that they are not accidental), returning Mesozoic, overwhelmingly Triassic, ages inferred to reflect a source in the local Sierra Nevada granites (Barth *et al.*, 2011). These grains are labelled as 'basement type' hereafter.

Systematic changes in zircon textures are observed in pumices collected through the Bishop Tuff eruptive sequence (Fig. 4; Electronic Appendix 2).



C Fig. 4. Zircon CL zoning patterns (Fig. 3) correlated to stratigraphic position (from Wilson & Hildreth, 1997) within the deposits studied here. (Note the systematic changes through the Bishop eruption sequence from a dominance of 'normal Bishop'-type zircons with no bright rim to a dominance of zircons that have various degrees of development of the CL-bright rims.)

Sector zoning

Sector zoning occurs both in non-rimmed grains and the cores of many bright-rimmed grains. Variable grey-scale intensities in the sector zoning dominate the CL images, whether oscillatory zoning is present or not (Fig. 3). Zircons from the fall units and early eastern ignimbrite lobes are dominantly (~85%) sector zoned, whereas the abundance of sector zoning diminishes to 24% in the latest unit (Ig2Nc). This diminution is gradual, with no sharp definable change corresponding to changes in eruption vent location (Hildreth & Mahood, 1986; Wilson & Hildreth, 1997). Sector zoning is also common in zircons from the IgISW and Ig2SW units, which have previously been suggested to be from a source similar to (but not necessarily co-vented from) the early erupted fall deposits and ignimbrite (Hildreth & Wilson, 2007).

Bright overgrowths

Bright overgrowths on zircons become the dominant feature in northern-erupted units, where <5% of the crystals show no bright overgrowth, and $\sim65\%$ of all crystals have >10 µm wide bright rims. The core-rim contact shows no evidence for dissolution of the crystal core (Fig. 3; Electronic Appendix 2), implying that ingress of the contrasting magma composition or any changes in physical conditions did not cause significant dissolution of pre-existing grains. The width of the bright rim generally increases in zircons from pumices later in the eruption sequence, with some crystals showing purely bright-CL characteristics (Fig. 4). Sector zoning is greatly subdued in the bright zones. The bright CL rims in zircons are inferred to be the counterparts to the bright rims reported from CL and BSE images on quartz and sanidine crystals, respectively (Peppard et al., 2001; Morgan & Blake, 2006; Wark et al., 2007; Pamukcu et al., 2012). Bright rims on quartz and sanidine are generally reported only in the later units erupted from the northern caldera rim, although Gualda (2007) reported sparse bright-rimmed quartz and feldspar grains from the early erupted sequences. Our observation that <10 µm bright rims are present in late IglEb (Sherwin subunit of Hildreth & Wilson, 2007) and Ig2SW is in agreement with glass and melt

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Sample	Eruptive unit	Number	Sector	No bright	$< 10\mu m$	10-30 μm	$>$ 30 μ m	All	Mesozoic (%)	Glass
		of grains	zoned (%)	rim (%)	bright	bright	bright	bright (%)		Mtn (%)
					rim (%)	rim (%)	rim (%)			
BP206	lg2Nc	688	24	5	32	27	8	22	3	3
BP124	lg2NWb	257	23	1	31	33	14	16	2	3
BP172	lg2Nb	67	42	43	9	9	6	15	18	0
BP232	lg2Nb	190	44	41	13	7	5	7	21	6
BP170	lg2Nb	53	15	0	11	51	21	13	0	4
BP059	lg2Na	257	61	3	32	30	16	14	3	4
BP115	lg2NWa	624	32	1	30	33	14	20	2	1
BP118	lg2NWa	98	89	4	67	22	3	3	0	0
BP216	lg2SW	306	67	42	54	2	0	1	1	0
BP040	Ig1Eb (Sherwin)	239	86	31	58	10	0	1	0	0
BP130	lg1SW	64	66	100	0	0	0	0	0	0
BP087	lg1Eb	79	99	100	0	0	0	0	0	0
BP098	lg1Eb	58	79	97	2	0	0	0	2	0
BP178	F1	284	83	100	0	0	0	0	0	0
BP109	F1	34	82	94	3	0	0	0	3	0
BR228	DomeYA	451	3	2	0	0	0	0	0	97
BR226	Dome OD	127	7	2	4	1	0	0	1	88

Table 2: Summary of textural features observed in zircons from Bishop Tuff samples examined in this study

Eruptive units are from Hildreth & Wilson (2007).

inclusion chemistry (Roberge *et al.*, 2013) showing some degree of mixing in 'middle-erupted' units of the Bishop Tuff. Bright rims are not found in zircons from Glass Mountain dome YA, yet are sparsely present on crystals from dome OD. We infer that the magma from which the bright zones crystallized in zircon (and other minerals) is not simply a short-lived feature associated only with the Bishop Tuff magma body, but represents a recurrent feature.

Zircons from swirly pumices

Two pumices sampled were of the crystal-poor, highly vesicular swirly type (Hildreth & Wilson, 2007), collected from Ig2Nb. Particularly high proportions (18 and 21%) of basement-derived zircons occur in these clasts; the grains wholly lack any overgrowths ('normal' or 'bright rim'), and are commonly fractured but always glass coated.

U-Pb age determinations

A total of 616 acceptable U–Pb ages were obtained on 12 samples of the Bishop Tuff, as well as 38 ages from Glass Mountain dome YA and 18 from dome OD. Acceptable' for the purpose of this study means analyses with <10% of the ²⁰⁶Pb attributable to common Pb, regardless of the U concentration at the spot point. Single age determinations

on grains with U>500 ppm have 2σ uncertainties between 14 and 179 kyr, with an average of ± 48 kyr; however, when the analyses are combined in weighted mean estimates the uncertainties are greatly reduced.

Bishop Tuff: general age data

Age determinations, together with values for ²⁰⁶Pb/²³⁸U ratios, concentrations of U and Th, and the proportions of ²⁰⁶Pb attributable to common Pb for all zircons analysed are given in Electronic Appendix 1. These data are annotated also for cores versus rims, sides versus tips of sectorzoned grains, and binned for the bright-rim widths. Zircons with a wide range of U concentrations were analysed (Fig. 5) although at lower U concentrations the analytical uncertainties increase and the proportions of 206 Pb attributable to common Pb often are beyond our 10% cutoff for acceptable analyses (see Methods section). Age determinations do not, within analytical uncertainties, increase at high uranium concentrations (Fig. 5) as has been reported elsewhere and attributed to the effects of metamictization or high trace element concentrations resulting in analyses of unknowns no longer being matrix-matched by the zircon standard (e.g. Gagnevin et al., 2011; White & Ireland, 2012). Variations in U concentrations are not systematic with sample or eruptive timing (Electronic

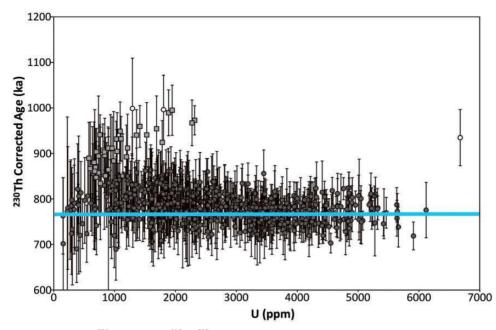


Fig. 5. U concentrations (ppm) vs 230 Th-corrected 206 Pb/ 208 U ages (ka) for all Bishop Tuff (circles) and Glass Mountain dome YA (squares) data points. The three open circles represent data points that fall outside of the bulk range of Bishop Tuff data and may be erroneous. The thick line indicates eruption age estimates from 40 Ar/ 39 Ar experiments on feldspar and ID-TIMS experiments on zircon (767 ka; see text for details and references). Error bars represent 2σ uncertainties.

Appendix l). Our age determinations for all Bishop eruptive units and all locations within crystals range from \sim 700 ka to \sim 880 ka. A probability density function (PDF) curve generated in Isoplot (Ludwig, 2008) for all acceptable Bishop age determinations shows a unimodal peak (Fig. 6), with an associated weighted mean value of 777.9 ± 2.2 ka (95% confidence limits), where 26 of 616 analyses were rejected (MSWD = 2.1, probability = 0.00).

Bishop Tuff: details in age data

Sector-zoned ages. Given the dominance of sector zoning in zircons from early erupted samples, we investigated whether the different sectors yielded the same ages for the same growth zones, as indicated by the 'growth rings' of oscillatory zoning. When sides and tips were compared there was no systematic age difference between the sectors (Fig. 7): the CL darker tips yield a weighted mean of 773.5 ± 2.8 ka (nine of 267 rejected, MSWD = 1.8) whereas the lighter sides have a weighted mean of $776 \cdot 1 \pm 7 \cdot 2$ ka (three of 60 rejected, MSWD=1.7). Side and tip ages within single crystals from samples BP098 and BP178, differ between the two sectors by <20 kyr for 13 of 22 crystals, but by 20-70 kyr for the others (Fig. 7; Electronic Appendix 1). When viewed together as part of a large sample suite it is apparent that there are no resolvable age differences between sides and tips.

Early versus late-erupted units. Table 3 summarizes weighted means for each of the 12 samples analysed for age. Four

samples from Fl and lowest IglEb representing 'early erupted' material have age spectra with weighted-mean values ranging between $752 \cdot 3 \pm 5 \cdot 7$ ka and $776 \cdot 1 \pm 5 \cdot 1$ ka. In later-erupted samples the weighted means are somewhat older, ranging from $790 \cdot 8 \pm 8 \cdot 5$ ka to 825 ± 13 ka. When plotting single sample PDF curves through the eruption sequence, older ages are represented more in the later erupted material (see Electronic Appendix 1). However, some bias has been introduced to analyses of zircons in the later-erupted samples owing to our specific targeting of cores within the bright-rimmed grains to ascertain whether they showed a longer crystallization history (Simon & Reid, 2005). By separating core and rim analyses prior to calculating weighted means, we account for this bias.

Cores versus rims. In the four 'early erupted' samples there is no significant difference in core versus rim ages, both yielding age-equivalent unimodal peaks (Fig. 8). Cores give a weighted mean of 774 ± 10 ka (one of 22 rejected, MSWD=16) and rims a weighted mean of 766.6 ± 3.1 ka (six of 172 rejected, MSWD=13). Three samples from the late-erupted Ig2Nb and Ig2Nc have rims that yield a slightly older weighted mean of 787.3 ± 5.2 ka (two of 62 rejected, MSWD=17). However, the cores of these lateerupted samples yield a bimodal PDF curve with a weighted mean of 807.5 ± 7.7 ka (three of 55 rejected, MSWD=14), which can be split into two age groups (<808 ka and >808 ka). The younger age group gives a weighted mean of

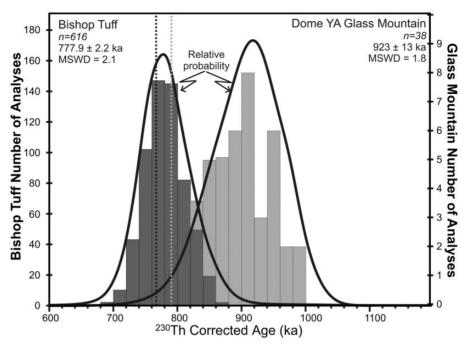


Fig. 6. Zircon ²³⁰Th-corrected ²⁰⁶Pb/²⁰⁸U age data for the entire Bishop Tuff (dark grey), and for dome YA (Glass Mountain, light grey). The dashed lines are published eruption ages of the Bishop Tuff (767 ka; see text) and dome YA (790 \pm 20 ka: Metz & Mahood, 1985). The probability density function (PDF) curve (from Isoplot; Ludwig, 2008) and histograms demonstrate the clear temporal separation between the youngest precursory Glass Mountain eruption and the Bishop Tuff. Values for the two datasets represent the number of spots used, and the weighted means, uncertainties at 95% confidence and mean square of weighted deviates (MSWD) values for each population. (For full dataset see Electronic Appendix 1.)

784·1 ±8·7 ka (none of 27 rejected, MSWD = 0·42), whereas as the older age group yields a weighted mean of 838·5 ±8·7 ka (none of 28 rejected, MSWD = 0·75). When the 'unmix' option in Isoplot is applied to the datasets, two weighted means are obtained of 790.6 ± 10.0 ka and 836.9 ± 11.0 ka. Both of these approaches imply that there are two populations of grains, both distinctly older than those in the early erupted samples, but which are indistinguishable on a textural basis.

Bright-rimmed grains. Age determinations from the bright rims (Table 2) have greater uncertainties commensurate with the lower U and Th concentrations (sometimes <200 ppm; see Electronic Appendix 2). When plotted on a Tera–Wasserburg diagram, the collection of all low-U analyses with <30% of the ²⁰⁶Pb attributable to common Pb yield a lower age intercept with a disequilibrium-corrected concordia value of 772 ±28 ka (Fig. 9), matching the data from the higher-U grains or domains. Ages from the higher-U cores, but from spots sited just inside narrow (<l0 µm) bright rims have a weighted mean of 777.5 ± 5.4 ka (five of 90 rejected, MSWD = 2.3).

Glass Mountain samples

Eighteen ages for dome OD have a weighted mean of 1.750 ± 0.021 Ma (one of 18 rejected, MSWD = 2.8: Table 3),

similar to the 'minimum age' average of 1.728 ± 0.043 Ma (1 σ , n = 4) from Simon & Reid (2005). Both figures are significantly older than the K–Ar eruption age of 1.35 ± 0.10 Ma (Metz & Mahood, 1985), but consistent with the 40 Ar/ 39 Ar age determination from step-heating experiments on sanidine of 1.702 ± 0.011 Ma (1 σ) by Davies *et al.* (1994). [Note that Davies *et al.* (1994) reported a value of 1.686 ± 0.011 Ma relative to a value of 1071 Ma for standard Hb3gr, for which Renne *et al.* (2011) gave an age of 1081 ± 1.2 Ma.]

Zircons from dome YA yield 38 acceptable ages, one of which $(1.682 \pm 0.079 \text{ Ma})$ is inferred to be xenocrystic, despite having identical textural characteristics to the others. The 37 remaining ages range from 1000 ka to 820 ka and yield a unimodal PDF peak with a weighted mean of 923 ± 13 ka (none rejected; MSWD = 1.8), somewhat younger than the average reported by Simon & Reid (2005) of 995 ± 19 ka (n = 9, MSWD = 2.2). A positive correlation of age versus U concentration is observed in the dome YA data (Fig. 5), in contrast to Bishop zircons, but we attribute this correlation to secular changes in the uptake of U into the grains, not to metamictization or excessive trace element concentrations. The weighted mean for the youngest 30% of our age data yields a value of 862 ± 23 ka (n = 12, MSWD = 0.096). We infer, based on previous work with New Zealand examples where eruption

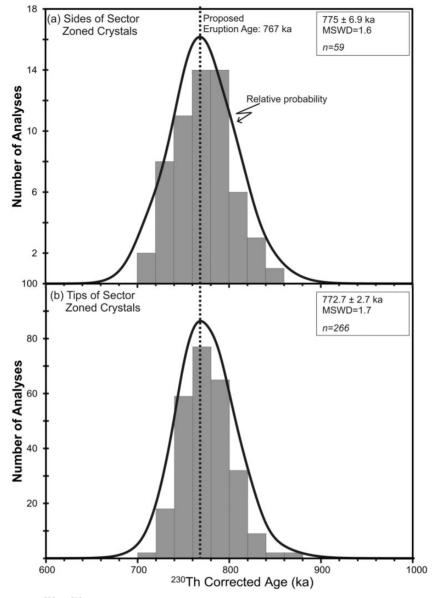


Fig. 7. Zircon 230 Th-corrected 206 Pb/ 208 U age data for all sector-zoned zircons, to compare age determinations from the tips vs sides of the same crystals. Dashed line represents the proposed eruption age from 40 Ar/ 39 Ar experiments on feldspar and ID-TIMS experiments on zircon (767 ka; see text for details and references). Probability density function (PDF) curves from Isoplot. Values for the datasets represent the number of spots used (*n*), and the weighted means, uncertainties at 95% confidence and MSWD values for the respective populations. (For full dataset see Electronic Appendix 1.)

ages have been independently determined (e.g. Milicich *et al.*, 2013), that this value represents an accurate estimate for the eruption age of dome YA. This age is in contrast to the 790 \pm 20 ka K–Ar determination by Metz & Mahood (1985).

Trace elements in Bishop zircons

Trace element analyses of zircons from seven Bishop Tuff samples were undertaken to characterize the trace element variation across the range of ages and CL characteristics, taking account of variations between sectors within single crystals as well as those between normal cores and the bright rims. Additionally, we collected trace element data for the Mesozoic grains in sample BP232 (Ig2Nb: Table 1) to see if the compositional patterns differed significantly from those for the young grains. As in all microanalytical approaches to mineral chemistry, there is the risk of analyses reflecting contamination by melt or crystal inclusions. Although analytical spots were placed away from any visible inclusions, whether in CL images or in reflected light, JOURNAL OF PETROLOGY VOLUME 55 NUMBER 2 FEBRUARY 2014

Table 3:	U - Pb	ages	measured	using	SIMS

Sample	Unit		ratio from LA-ICPMS sis of glass shards	U (ppm)	Th (ppm)	²³² Th/ ²³⁸ U	Total ²⁰⁶ Pb/ ²³⁸ U	²³⁰ Th corrected 206/238 age (ka)	Error; 2σ (ka)
BP109	F1	2·5	average	3393	2134	0.6178	0.00011	770	46
			average 'normal' core	3549	2487	0.6538	0.00011	781	45
			average 'normal' rim	3301	1980	0.6027	0.00011	767	46
BP178	F1	2·5	average	3233	1986	0.5808	0.00011	773	40
			average 'normal' core	3298	2181	0.6066	0.00011	790	42
			average 'normal' rim	3357	1996	0·5752	0.00011	764	37
BP087	lg1Eb	2·7	average	3685	2481	0.6697	0.00011	760	44
			average 'normal' rim	3714	2517	0.6745	0.00011	760	43
0000	L.:15h	07	-						
BP098	lg1Eb	2.7	average	3440	2039	0.5779	0.00011	777	42 42
			average 'normal' rim	3416	2019	0.5759	0.00011	776	42
BP040	Ig1Eb-Sherwin	2.8	average	2845	1749	0.5918	0.00011	767	41
			average 'normal' core	2429	1312	0.5191	0.00011	761	52
			average 'normal' rim	3271	2145	0.6445	0.00011	758	41
BP059	lg2Na	3.3	average	2404	1476	0.6030	0.00011	796	50
			average 'normal' rim	2913	1838	0.6151	0.00011	790	45
			average 'normal' core	1654	877	0.5736	0.00012	816	60
			average 'bright'	457	228	0.5278	0.00011	778	73
BP170	lg2Nb	4·0	average	1414	733	0.5360	0.00011	809	72
	-		average 'normal' rim	1808	889	0.4971	0.00011	820	65
			average 'normal' core	1534	738	0.4782	0.00011	814	69
			average 'bright' rim	383	289	0.7005	0.00011	775	114
			average 'bright' core	677	423	0.6486	0.00011	786	82
BP206	lg2Nc	3.7	average	2073	1248	0.5739	0.00011	793	46
	.9	•	average 'normal' core	1864	1047	0.5665	0.00011	800	58
			average 'bright' core	942	436	0.4741	0.00011	809	49
			average rim	2979	1962	0.6315	0.00011	773	37
BP118	lg2NWa	2.6	average	3003	2094	0.6746	0.00011	773	50
DELIO	Igzivva	2.0	average 'normal' rim	3080	2094	0·6740 0·6791	0.00011	771	49
			average 'normal' core	3375	2727	0.7710	0.00011	775	45 50
			average 'bright'	2095	1300	0.6369	0.00011	783	71
00445									
BP115	lg2NWa	3·1	average	1590	905	0.5770	0.00011	802	54
			average 'normal' rim	2600	1640	0.6383	0.00011	791	37
			average 'normal' core	1548	783	0.5097	0.00011	817	42
			average 'bright'	451	297	0.6526	0.00011	780	107
BP124	lg2NWb	3.7	average	2410	1489	0.6225	0.00011	793	51
			average 'normal' rim	3042	2031	0.6686	0.00011	794	45
			average 'normal' core	2073	1190	0.5852	0.00011	794	57
			average 'bright'	747	544	0.7627	0.00010	731	87
BP232	lg2Nb (sw)	3.0	average	2930	2002	0.6785	0.00011	796	38
			average 'normal' rim	3631	2484	0.6855	0.00011	790	31
			average 'normal' core	2343	1550	0.6522	0.00012	816	41
			average 'bright'	387	201	0.5384	0.00012	780	59
BR228	Dome YA	3·2	average	1099	512	0.4665	0.00013	928	71
			-						
BR226	Dome OD	2.3	average	8129	4569	0.5583	0.00026	1759	58

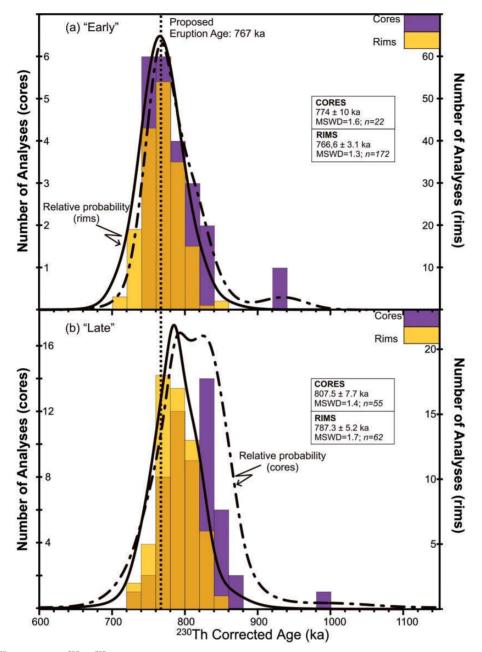


Fig. 8. Zircon ²³⁰Th-corrected ²⁰⁶Pb/²⁰⁸U age data compared for spots in cores and rims from typical early and late-erupted samples. Dashed line represents the proposed eruption age from ⁴⁰Ar/³⁹Ar experiments on feldspar and ID-TIMS experiments on zircon (767 ka; see text for details and references). (a) Early erupted samples (FI-BPI09 and BPI78; IglEb-BP087; see Table 1) have cores with a mean age that is closely similar to that of the rims from these samples. Both are within uncertainty of the independently determined eruption age estimates (see text). (b) Late-erupted rims from Ig2Nb/c (BPI70, BP206 and BP232; see Table 1) samples are slightly older than eruption age, whereas cores show a marked tail-off to older ages. Probability density function (PDF) curve from Isoplot. Values for the datasets represent the number of spots used (*n*), and the weighted means, uncertainties at 95% confidence and MSWD values for the respective populations. (For full dataset see Electronic Appendix 1.)

in some cases it became apparent that inclusions were intersected. We used common elements that are normally found at very low concentrations in zircon compared with the contaminant phase to monitor such contamination: Ca, Al, Na and K for feldspar and glass; Ca, P and F for apatite; Fe for Fe–Ti oxides; and Ca and Fe for allanite. Analyses showing significant degrees of contamination were culled accordingly. The full data are given in Electronic Appendix 3; spots with severe contamination are omitted, whereas others with minor contamination are highlighted

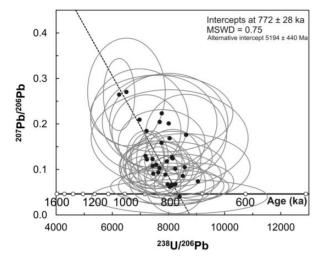


Fig. 9. Inverse isochron (Tera–Wasserburg) diagram generated in Isoplot of all analyses from grains with U <500 ppm and up to 30% of the ²⁰⁶Pb attributable to common Pb, using the common-Pb composition for this age from the crustal evolution model of Stacey & Kramers (1975) (see text for discussion). The best-fit line is not pinned to any intercept value. Ellipses represent 2σ uncertainty on single analyses. (For dataset see Electronic Appendix 1.)

and have suspect concentrations of affected elements struck through.

Bishop Tuff: general trace element data

When plotted all together, the trace element data form two contrasting end-member patterns (Fig. 10), between a single coherent array where intra-grain variations (see section below on sector zoning) fall within the overall trend of the array, or two separated but subparallel arrays. The overall trends in Fig. 10 reflect fractionation-driven changes in the melt from which the zircons were growing (Reid et al., 2011; Fig. 10a and b). It is typical for Ti and Th/U and Nd/Yb ratios to decrease and Yb/Gd to increase with increasing melt fractionation as measured by higher Hf concentrations or lower Eu/Eu* ratios. U typically covaries with Yb and molar $(Y + Sc + REE^{3+})/P$ (Belousova et al., 2002; Claiborne et al., 2006, 2010; Barth & Wooden, 2010; Barth et al., 2012). Zircons from the earlier-erupted samples tend to have Eu/Eu* ratios around or lower than 0.01, whereas those from later-erupted samples have higher values owing to crystallization from a (still very evolved) melt that has experienced slightly less feldspar fractionation (Reid et al., 2011). The xenocrystic Triassicage zircons follow the same general trends, but differ in values of Eu/Eu*, Yb/Gd and Th/U (Fig. 10e and f).

In contrast to the general fractionation-driven trends in Fig. 10, some elements or elemental ratios clearly plot as two parallel arrays. When subdivided into distinct groups the data reflect the overarching influence of sector zoning (Fig. 10). Two subparallel trends are apparent when Yb or another heavy REE (HREE), or molar $(Y + Sc + REE^{3+})/P$ (Fig. 10c and d), is plotted against U; these elements require careful consideration if magmatic trends are to be interpreted from these parameters.

Sector-zoning contrasts

Whenever present, sector zoning strongly influences the distributions and abundances of many trace elements (Table 4; Figs 10 and 11). We have thus differentiated the data into those from crystal sides (growth along the aand *b*-axes; lighter shades under CL) and those from tips (growth along the c-axis; darker under CL) (Fig. 11). CL emission in zircons is primarily thought to be generated by Dy, Sm, Eu, Tb, Gd, Er and Nb, and can be suppressed by U, Y, Ca, P and Fe (Nasdala et al., 2003). For the Bishop zircons that exhibit sector zoning, plots of many trace elements or their ratios against an element or ratio that can be used as an index of fractionation (e.g. Hf, Yb/Gd; Grimes et al., 2009; Claiborne et al., 2010; Carley et al., 2011) define two arrays that might, in the absence of textural context, be interpreted as reflecting separate compositional suites. When compared with the lighter sides, the darker tips are enriched in Ti by $\sim 45\%$, in the $(Sc + Y + REE^{3+})/P$ molar ratio by ~100% (Fig. 11a and b), in U and Th by up to about three times, and in Nb, U, Sm, Nd, Gd and Dy by 1:3-2 times (see Table 5 and Electronic Appendix 3). The lighter sides more closely follow the 'xenotime substitution' mechanism, where the molar ratio $(Y + Sc + REE^{3+})/P$ is unity (Hoskin & Schaltegger, 2003; Reid et al., 2011), whereas the darker tips show values of this ratio between 2.1 and 3.0, increasing with U, Th or Yb. This contrast between the different sectors appears to mostly reflect a relative depletion of P in the darker tips (Table 5, Fig. 11d). Lighter sides have lower concentrations of trace elements in general, when compared with the dark tips. However, the separate data arrays for the two sectors are roughly subparallel in elongation and show similar evolutionary trends. Some elements, such as the HREE and Y (Fig. 1le and f), are apparently less influenced by the sector zoning.

Core-bright-rim contrasts

In zircons that lack a bright rim, sector zoning dominates the variations in trace-element patterns. However, where bright rims were wide enough to analyse [three samples: IglEb (BP040), Ig2NWa (BP118) and Ig2Nb (BP170)] it is clear that the normal, sector-zoned Bishop cores (as defined on textural grounds) are enriched (on average) by $\sim 50\%$ in middle REE (MREE), Th and U relative to the bright rims (Fig. 12a and b). Although the absolute enrichment differs when bright rims are compared with dark tips or light sides, the enrichment is nonetheless present, and therefore we use an average of the sides and tips to calculate our enrichment factors. The bright rims crystallized from less evolved melts, displaying higher Eu/Eu*

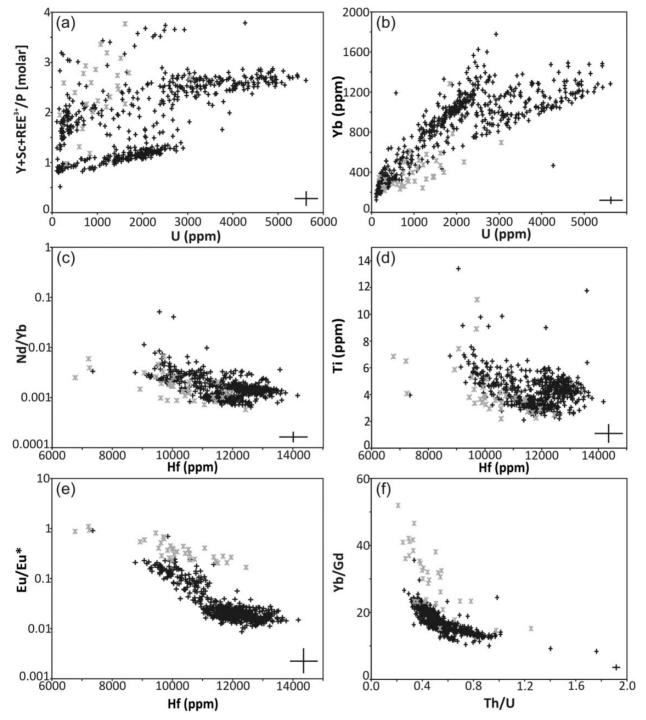


Fig. 10. Summary plots of zircon trace element data from samples from the full range of the Bishop eruptive sequence. With certain trace elements or elemental ratios, sector zoning is clearly represented by the presence of split arrays (a, b), whereas its effect is subdued in other cases (c-f). Cross in bottom right of each panel represents 2σ uncertainties based on repeat analyses of the MAD green zircon (Barth & Wooden, 2010). Triassic-age grains are marked in grey; all Bishop data are in black. (For full dataset see Electronic Appendix 3.)

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471 97 369 2026 32 0003 499 102 300 003 499 102 300 3037 795 6625 1197 1037 <i>11E Sharwin, lightar sides</i> 1077 156 299 2964 31 0012 401 095 516 0125 647 1046 1046 1046 1688 183 360 3555 52 0001 252 049 550 144 1893 3696 1416 1646 1564 1416 1563 168 144 449 357 66 0112 50 054 159 1563 1416 1563 171 144 449 357 66 0112 50 054 59 3696 1646 1416 1563 1563 1563 1563 1563 1563 1563 1563 1563 1563 1563 1563 1563 1563 1564 <	471 97 3.69 2026 32 0.003 49.9 $1Eb$ (Sherwin), lighter sides 1077 156 2.99 2964 31 0.012 40.1 1077 156 2.99 2964 31 0.012 40.1 1077 156 2.99 2964 31 0.012 40.1 1688 183 3.60 3535 52 0.029 52.5 448 124 2.125 225 0.030 114.2 90.03 721 144 4.49 3577 66 0.030 114.2 90.030 721 144 4.49 3577 66 0.030 114.2 90.030 114.2 721 101 3.617 857 96 0.030 114.2 90.030 114.2 721 101 3.61 2.51 39.4 114.2 90.030 114.2 721 101 3.61 2.92 90.030 $2.$	117·3 2·00	0·209	` ຕ	37·8	447·8	829.2	179·2	1491·1	262.6	13239	0.534	5434	4806
IED (Sherwin), lighter sides 10 2 4 10 541 10 641 10 </td <td>IEb [Sherwin], lighter sides 1077 156 2:99 2964 31 0:012 40:1 1688 183 3:60 3535 5:2 0:029 5:2:5 1688 183 3:60 3535 5:2 0:029 5:2:5 1688 124 2:125 22 0:011 25:2 17Eb (Sherwin), darker tips 721 144 4:49 3577 66 0:012 90:0 721 144 4:49 3:577 66 0:012 90:0 114:2 721 144 4:49 3:577 66 0:012 90:0 114:2 721 144 4:49 3:577 66 0:012 90:0 52:2 2Na, lighter sides 1227 109 3:00 2827 18 0:045 39:4 1227 109 3:00 2827 18 0:0127 52:1 1227 109 3:00 285 0:020 29:2 0:020 29:4 1463 170 3:48 3676</td> <td>49-9 1-02</td> <td>0.083</td> <td>œ</td> <td>16.3</td> <td>195.6</td> <td>373-7</td> <td>79-5</td> <td>662·5</td> <td>119-7</td> <td>12087</td> <td>0.237</td> <td>2144</td> <td>957</td>	IEb [Sherwin], lighter sides 1077 156 2:99 2964 31 0:012 40:1 1688 183 3:60 3535 5:2 0:029 5:2:5 1688 183 3:60 3535 5:2 0:029 5:2:5 1688 124 2:125 22 0:011 25:2 17Eb (Sherwin), darker tips 721 144 4:49 3577 66 0:012 90:0 721 144 4:49 3:577 66 0:012 90:0 114:2 721 144 4:49 3:577 66 0:012 90:0 114:2 721 144 4:49 3:577 66 0:012 90:0 52:2 2Na, lighter sides 1227 109 3:00 2827 18 0:045 39:4 1227 109 3:00 2827 18 0:0127 52:1 1227 109 3:00 285 0:020 29:2 0:020 29:4 1463 170 3:48 3676	49-9 1-02	0.083	œ	16.3	195.6	373-7	79-5	662·5	119-7	12087	0.237	2144	957
10715629629643100124010954070086517208566167611119610652134813461681883603556520029525129516012061124932001392639213961748216412554481242442442493501441883396939693969149615631416716532449357660101252056014838335951445646317482164125837211444493576601029001556630114196833145517451661303721144449357680030114219683335744951865836117451961303721101361257380025520830144498307449517691374196132620141361371383369369636963696369636963696369636962014146317036137137336953696369636963696369636962014146317036839136936963696369636963696369636962014 </td <td></td>														
183 360 355 52 0029 57.6 1/29 54.7 139.2 639.2 139.9 1174.8 2154.2 1552 174 244 215 2 0001 25.2 055 2.67 0.048 350 144 1893 3959 896 7563 1416 11591 nvmi/ darker tips 1 1 3 </td <td>1688 183 3-60 3535 52 0-029 52-5 448 124 2-44 2125 22 0-001 25-2 721 144 4-49 3577 66 0-012 90-0 721 144 4-49 3577 66 0-012 90-0 838 212 5-32 4522 96 0-030 114-2 571 101 3-61 2257 38 0-002 52.2 2Na, lighter sides 1101 3-61 2257 38 0-012 90-0 1227 109 3-00 2827 18 0-012 52-2 1463 170 3-48 3676 40 0-137 52-1 1021 46 2-49 1995 2-5 0-020 29-2 2Na, darker tips 626 98 4.35 3147 34 0-038 2758 165 5-50 4104 87 0-030 121-8</td> <td>40.1 0.95</td> <td>0·089</td> <td></td> <td>266.6</td> <td>116·7</td> <td>541·1</td> <td>119-6</td> <td>1005·2</td> <td>184·8</td> <td>11945</td> <td>0·183</td> <td>1970</td> <td>880</td>	1688 183 3-60 3535 52 0-029 52-5 448 124 2-44 2125 22 0-001 25-2 721 144 4-49 3577 66 0-012 90-0 721 144 4-49 3577 66 0-012 90-0 838 212 5-32 4522 96 0-030 114-2 571 101 3-61 2257 38 0-002 52.2 2Na, lighter sides 1101 3-61 2257 38 0-012 90-0 1227 109 3-00 2827 18 0-012 52-2 1463 170 3-48 3676 40 0-137 52-1 1021 46 2-49 1995 2-5 0-020 29-2 2Na, darker tips 626 98 4.35 3147 34 0-038 2758 165 5-50 4104 87 0-030 121-8	40.1 0.95	0·089		266.6	116·7	541·1	119-6	1005·2	184·8	11945	0·183	1970	880
48 124 244 214 215 2 0 1 2 1 <th< td=""><td>448 124 2.44 2125 22 0.001 25.2 1fEb (Sherwin), darker tips 721 144 4.49 3577 66 0.012 90.0 721 144 4.49 3577 66 0.012 90.0 838 212 5.32 4522 96 0.030 1142 521 101 3.61 2257 38 0.002 52.2 2Na, lighter sides 3.00 2827 18 0.045 39.4 1227 103 3.00 2827 18 0.045 39.4 1227 103 3.00 2827 18 0.045 39.4 1227 103 3.00 2827 18 0.045 39.4 1227 103 3.00 2827 18 0.045 39.4 1021 46 2.49 1995 2.5 0.020 29.2 2Na, darker tips 626 98 4.35 3147 34 0.038 556 4104 87 0.030 121.8 758 165 5.50 4104 87 0.030</td><td>52.5 1.29</td><td>0·122</td><td></td><td>320-0</td><td>139·2</td><td>639·2</td><td>139-9</td><td>1174·8</td><td>215-4</td><td>12552</td><td>0·274</td><td>2354</td><td>1130</td></th<>	448 124 2.44 2125 22 0.001 25.2 1fEb (Sherwin), darker tips 721 144 4.49 3577 66 0.012 90.0 721 144 4.49 3577 66 0.012 90.0 838 212 5.32 4522 96 0.030 1142 521 101 3.61 2257 38 0.002 52.2 2Na, lighter sides 3.00 2827 18 0.045 39.4 1227 103 3.00 2827 18 0.045 39.4 1227 103 3.00 2827 18 0.045 39.4 1227 103 3.00 2827 18 0.045 39.4 1227 103 3.00 2827 18 0.045 39.4 1021 46 2.49 1995 2.5 0.020 29.2 2Na, darker tips 626 98 4.35 3147 34 0.038 556 4104 87 0.030 121.8 758 165 5.50 4104 87 0.030	52.5 1.29	0·122		320-0	139·2	639·2	139-9	1174·8	215-4	12552	0·274	2354	1130
(1E) (Sherwin), darker tips 721 144 449 3571 66 0.012 900 155 66.3 0.115 791 369.5 144.5 646.3 135.9 1111.7 196.0 12823 721 144 449 3577 66 0.012 900 155 67.3 0.146 186.5 836.1 178.5 1475.2 266.8 13094 838 212 552 452 96 0.030 142 196 0.495 496 885.1 1785 487.8 136.7 148.9 136.9 1111.7 196.0 12500 204, johter sides 200 282 0.409 0.984 498 200 241.9 93.9 147.9 143.2 143.2 12680 204, johter sides 300 2827 49.9 0.993 249.9 196.0 143.6 143.6 143.9 143.9 143.9 143.9 143.9 143.9 143.9 143.9 <	(1Eb (Sherwin), darker tips 721 144 4.49 3577 66 0-012 90-0 838 212 5.32 4522 96 0-030 1142 521 101 3.61 2257 38 0-002 52.2 2Na, lighter sides 101 3.61 2257 38 0-045 39-4 1227 109 3.00 2827 18 0-045 39-4 1463 170 3.48 3676 40 0-137 52-1 1021 46 2.49 1995 2-5 0-020 29-2 2Na, darker tips 66 98 4.35 3147 34 0-038 85-4 756 98 4.35 3147 34 0-030 121-8 7 755 166 5-0 4104 87 0-030 121-8 7	25·2 0·55	0.048		188·3	83·8	396.9	89.6	756.3	141 <i>·</i> 6	11591	0·122	1418	507
721144449357766 0.012 900155 6.63 0.115 79.1 359.5 144.5 646.3 135.9 1111.7 196.0 12823 838212 5.32 4522 96 0.030 114.2 1.96 8.39 0.148 983 37.6 4496 1865 836.1 1785 1475 2668 1304 521101 3.61 2257 38 0.002 52.2 0.83 4.09 0.948 498 201.7 95.9 4489 93.9 807.2 1432 1280 $2Na, lighter sides$ 103 3.61 2251 0.83 4.09 0.934 498 211.7 95.9 4489 93.9 807.2 1432 1280 $2Na, lighter sides$ 103 300 2827 18 0.025 52.4 0.93 394 111.6 201.7 2662 5448 1209 1324 1224 1453 340 340 111.6 0.93 481 111.6 207 2662 5448 1209 1324 1224 1463 340 2827 1450 2662 1489 239.2 165 2419 12806 11719 $2Na, lighter sides170348111.6201226226481209129212419146313612962129212921262129212921292119<$	721 144 4.49 3577 66 0-012 90-0 838 212 5.32 4522 96 0-030 114.2 521 101 3-61 2257 38 0-002 52.2 2Na, lighter sides 1227 109 3-00 2827 18 0-045 39-4 1227 109 3-00 2827 18 0-045 39-4 1463 170 3-48 3576 40 0-137 52-1 1021 46 2-49 1995 2-5 0-020 29-2 2Na, darker tips 626 98 4.35 3147 34 0-038 85-4 758 165 5-50 4104 87 0-030 121-8													
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521 101 361 225 38 0.002 522 0.83 4.09 0.084 4.98 200 241.7 959 4.90 307.2 143.2 1232 1232 1332 1332 1332 1232 1232 1237 1232 1234 1024	521 101 3.61 2257 38 0.002 52.2 2Na, lighter sides	114·2 1·96	0·148		449-6	186-5	836.1	178-5	1475-2	266·8	13094	0.492	5256	4534
2Na, lighter sides 2Na, li	2Na, lighter sides 1227 109 3.00 2827 18 0.045 39.4 1463 170 3.48 3676 40 0.137 52.1 1021 46 2.49 1995 2.5 0.020 29.2 2Na, darker tips 626 98 4.35 3147 34 0.038 85.4 758 165 5.50 4104 87 0.090 121.8	52·2 0·83	0.084		241-7	95.9	448·9	93·9	807·2	143·2	12580	0·308	2976	1982
127 109 3.00 2827 18 0.045 394 0.955 3 48.1 11.6 20.7 266.2 54.48 120.9 1033.4 188.0 1224 1433 170 3.48 3676 40 0.137 52.1 1.60 6.37 0.130 65.7 1450 280 370.2 157.9 1350.5 241.9 1205 1021 46 2.49 1956 2.5 0.020 292 0.62 2.63 0.068 32.4 76.9 137.7 178.6 71.9 84.6 735.8 137.5 11719 2Na, larker tips 1 2.4 76.9 137.7 178.6 371.4 84.6 735.8 137.5 11719 2Na, larker tips 1 3 4.35 1.57 0.78 377.4 84.6 735.8 137.5 11719 256 9 1 7.79 177.9 1277.5 265.9 247.1 1377.4 <	1227 109 3.00 2827 18 0.045 39.4 1463 170 3.48 3676 40 0.137 52.1 1021 46 2.49 1995 2.5 0.020 29.2 2Na, darker tips 626 98 4.35 3147 34 0.038 85.4 758 165 5.50 4104 87 0.090 121.8													
145 170 348 3676 40 0·137 52.1 160 6·37 0·130 65.7 145.0 28.0 359.0 750.2 157.9 1350.5 241.9 12805 1021 46 2.49 1995 2.5 0·020 29.2 0·62 2.63 0·088 32.4 76.9 137 1786 371.4 84.6 735.8 1375 1719 2No, darker tips 0·130 29.2 0·137 1786 371.4 84.6 735.8 1375 1719 2No, darker tips 0·131 0·152 0·130 67'2 127'5 26'9 327'9 55'6 135'7 1719 2116 758 165 5·50 4104 87 0·030 121'8 2·14 7.79 0·170 92'5 188'4 36'8 141'1 137'4 137'14 758 3·54 2·16 <t< td=""><td>1463 170 3.48 3676 40 0.137 52.1 1021 46 2.49 1995 2.5 0.020 29.2 2Na, darker tips 626 98 4.35 3147 34 0.038 85.4 758 165 5.50 4104 87 0.090 121.8</td><td>39.4 0.95</td><td>0.093</td><td></td><td>20.7</td><td>266·2</td><td>544·8</td><td>120-9</td><td>1033·4</td><td>188·0</td><td>12241</td><td>0·188</td><td>1859</td><td>797</td></t<>	1463 170 3.48 3676 40 0.137 52.1 1021 46 2.49 1995 2.5 0.020 29.2 2Na, darker tips 626 98 4.35 3147 34 0.038 85.4 758 165 5.50 4104 87 0.090 121.8	39.4 0.95	0.093		20.7	266·2	544·8	120-9	1033·4	188·0	12241	0·188	1859	797
1021 46 2-49 1995 2-5 0-020 29-2 0-62 2-63 0-068 32-4 76-9 13-7 178-6 371-4 84-6 735-8 137-5 11719 <i>2Nu-darket tips</i> 2 6 0-020 29-2 0-62 2-63 0-068 32-4 76-9 13-7 178-6 371-4 84-6 735-8 137-5 11719 <i>2Nu-darket tips</i> 3 3147 34 0-038 85-4 1-52 5-71 0-130 67-2 127-5 26-9 327-9 556-0 127-5 1659-8 185-6 13150 758 165 5-50 4104 87 0-090 121-8 2-14 779 0-170 92-5 168-4 36-8 444-4 778-2 167-5 1392-8 241-1 13774 48 35 3-64 2152 2-6 10-17 3-74 0-074 42-6 86-9 17-3 143-7 1332-8 241-1 13774 48 35 3-18 242-6 17-7 3-17 <td< td=""><td>1021 46 2.49 1995 2.5 0.020 29.2 2 ¿2Na, darker tips </td><td>52·1 1·60</td><td>0·130</td><td></td><td>28·0</td><td>359-0</td><td>720·2</td><td>157-9</td><td>1350-5</td><td>241-9</td><td>12805</td><td>0·304</td><td>2562</td><td>1214</td></td<>	1021 46 2.49 1995 2.5 0.020 29.2 2 ¿2Na, darker tips	52·1 1·60	0·130		28·0	359-0	720·2	157-9	1350-5	241-9	12805	0·304	2562	1214
2 <i>Na, darker tips</i> 626 98 4:35 3147 34 0·038 85-4 1·52 5·71 0·130 67·2 127·5 26·9 327·9 595·0 127·5 1059·8 185·6 13150 758 165 5·50 4104 87 0·090 121·8 2·14 7·79 0·170 92·5 168·4 36·8 444.4 778·2 167·5 1392·8 241·1 13774 448 35 3·64 2152 2·6 0·016 57·6 1·17 3·47 0·074 42·6 86·9 17·3 218·0 420·8 93·2 783·1 143·7 12563	2Na, darker tips 626 98 4.35 3147 34 0.038 85-4 758 165 5-50 4104 87 0.090 121-8 700 000 121-8	29·2 0·62	0·068		13·7	178·6	371-4	84·6	735·8	137·5	11719	0·089	1141	422
626 98 4.35 3147 34 0.038 85.4 1.52 5.71 0.130 67.2 127.5 26.9 327.9 595.0 127.5 1059.8 185.6 13150 758 165 5.50 4104 87 0.090 121.8 2.14 7.79 0.170 92.5 168.4 36.8 444.4 778.2 167.5 1392.8 241.1 13774 448 35 3.64 2162 1.17 3.47 0.074 42.6 86.9 17.3 218.0 423.7 1363.2 783.1 143.7 12563	626 98 4.35 3147 34 0.038 85.4 758 165 5.50 4104 87 0.090 121.8													
758 165 5-50 4104 87 0-090 121:8 2:14 7.79 0:170 92:5 168:4 36:8 444:4 778:2 167:5 1392:8 241:1 13774 448 35 3:64 21:62 2:6 0:016 57:6 1:17 3:47 0:074 42:6 86:9 17:3 218:0 420:8 93:2 783:1 143:7 12563	758 165 5-50 4104 87 0-090 121-8 27 0.01 0.110	85·4 1·52	0.130		26.9	327-9	595.0	127-5	1059-8	185.6	13150	0·374	3463	2370
448 35 3.64 2152 2.6 0.016 57.6 1.17 3.47 0.074 42.6 86.9 17.3 218.0 420.8 93.2 783.1 143.7 12563		121·8 2·14	0·170		36·8	444-4	778·2	167-5	1392·8	241·1	13774	0.538	4883	4235
	448 35 3·64 2152 2·6 0·016 57·6	57.6 1.17	0.074	9	17-3	218·0	420·8	93·2	783·1	143·7	12563	0·150	1448	582

Table 4: Average and range in trace element concentrations for all samples analysed in the Bishop Tuff

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	Ч	Sc	Ξ	٨	Nb	La	Ce	Nd	Sm	Eu	Gd	Тb	Dy	Но	Er	Tm	Чb	Lu	Hf	Pb	D	Th
BP170, Ig2Nb, cores	Nb, core.	s																				
average	714	114	3.94	2377	23	0.026	58.1	1-46	4.72	0·135	50.2	94-4	19.1	228-4	432·2	93·3	773-0	140-4	11739	0·160	1536	972
maximum	1245	184	60.9	4613	88	0·118	134-9	3·31	9-49	0·284	108·2	180·9	41·2	475-9	822·1	173·1	1384·9	238·8	12302	0.530	5148	4988
minimum	359	49	2.71	1400	2.4	0.005	20.0	0.51	1 [.] 82	0.020	23·1	52·3	9.3	119.6	251·5	57.1	487·1	90·1	10428	0.041	655	249
BP170, Ig2Nb, bright rims	Nb, brigł	it rims																				
average	303	51	4.99	819	4·0	0.023	34-7	0·74	1·91	0·187	17-5	32.7	6.5	77-4	156-1	35·2	304-9	58·8	10309	0.053	362	184
maximum	684	67	6.80	1488	5.8	0·173	57-7	1.39	3·67	0.581	34.0	65 - 4	13·8	165-5	320-6	70·7	598·1	110-0	12228	0·182	1185	638
minimum	146	28	3·13	327	2.1	0·004	20.3	0·27	0.79	0.050	7.0	12·6	2.3	30.0	61·5	13.7	123·6	25-1	9296	0·008	110	49
BP118, Ig2NWa, lighter sides	NWa, lig.	hter sidé	Sé																			
average	1331	148	3·24	2936	26	0.014	42.1	0.95	3.96	0.092	49.4	110.1	20.1	254.6	517.6	114·3	956-3	175-5	11588	0·182	1757	859
maximum	1688	197	3·73	4401	42	0.040	53·7	1·25	5.21	0·128	68·6	155 · 0	28.4	357·8	723·3	160.0	1312·3	236·0	12284	0·288	2747	1400
minimum	1075	107	2.70	2126	17	0.003	33.9	0.67	3·03	0.073	39.6	82·6	15.3	189·8	379-9	82.9	712·7	133·3	10771	0·109	1234	585
BP118, Ig2NWa, darker tips	NWa, da	rker tips	(*																			
average	686	142	4.75	3273	57	0.015	91.6	1.57	6·12	0·132	72·3	132 <i>·</i> 4	27-7	329.8	594.4	125.9	1027·1	181·0	12429	0.351	3500	2612
maximum	872	186	5.37	4555	103	0.052	125.2	1·99	8·53	0·234	104·2	185 · 0	40.1	471·2	842·0	176-7	1462·7	252.5	12908	0.503	5423	4479
minimum	489	95	4·10	1910	30	0.001	50.5	0·89	3·52	0.085	41·0	77-4	15·8	188·1	352-9	76·2	636.9	115·8	11438	0·213	1918	1060
BP118, Ig2NWa, cores	NWa, cor	es.																				
average	1043	140	4·16	2773	30	0.031	60.7	1·21	4·73	0·122	54·2	108·9	21.3	259-7	501·6	109-1	912·1	166-0	11621	0·190	2029	1235
maximum	1567	178	5.79	4049	76	0·153	108·1	1·80	8·07	0·172	91.0	164 <i>·</i> 0	34-7	405·8	735.6	155-4	1289-0	225-9	12239	0·306	3789	2697
minimum	573	<u> 8</u> 6	2.76	1871	10	0·004	28-9	0-57	2.51	0.076	33·1	71·2	13·4	165-7	326-6	72.5	606·0	114·5	10939	0·101	838	386
BP118, Ig2NWa, bright rims	NWa, bri	ight rims	5																			
average	459	64	5.20	1093	8·0	0·304	38·6	1·25	2.46	0·161	22·1	43·0	8.4	101·3	202-9	45·2	385·7	72.5	10608	0.069	522	262
maximum	1092	136	7.56	2230	17	1·896	53·8	3·17	3.95	0·324	40.7	86.4	15.9	197-4	400-4	88·4	756-5	139.2	11179	0.105	945	489
minimum	188	42	3.17	570	2.9	0.007	21·3	0.53	1.35	0.065	12.7	22·2	4.6	52.5	105-3	24·1	210·7	39.6	9514	0.037	196	111
BP232, Ig2Nb, cores	Nb, core.	S																				
average	1037	149	4·17	3369	61	0.037	70·2	1·52	5.47	0·149	66·6	131 <i>·</i> 2	25.9	316·7	601·8	128·3	1077-4	190·2	11876	0·314	2716	1739
maximum	1583	296	5.59	4415	300	0·106	121-4	2·52	7·93	0.544	98·5	171-7	37.5	438·1	770-9	165-0	1349·7	237·1	13035	0.653	5320	5015
minimum	368	30	2.90	1199	2.1	0·00	34.5	0.92	3·18	0-071	32.9	47·6	10.6	120·3	225·2	48·3	422·5	78·4	9651	0.045	265	172
BP232, Ig2Nb, bright rims	Nb, brigł	it rims																				
average	795	113	4.43	2559	29	0.055	48·2	1.51	4·63	0·197	51·1	100·9	19·8	242.0	471·0	101·6	863.5	154-9	11309	0·203	1751	915
maximum	1618	179	7·84	4488	71	0·297	84-4	4·08	10·98	1·068	110-0	182·1	40·5	467-4	836-6	172.7	1406·6	241·2	13605	0.490	4113	2900
minimum	110	32	2.67	636	2.2	0.013	23·7	0·62	1·64	0.063	14·8	25·0	5.1	62·4	119·8	26·7	241-4	45.0	7354	0·018	156	100

413

Table 4: Continued

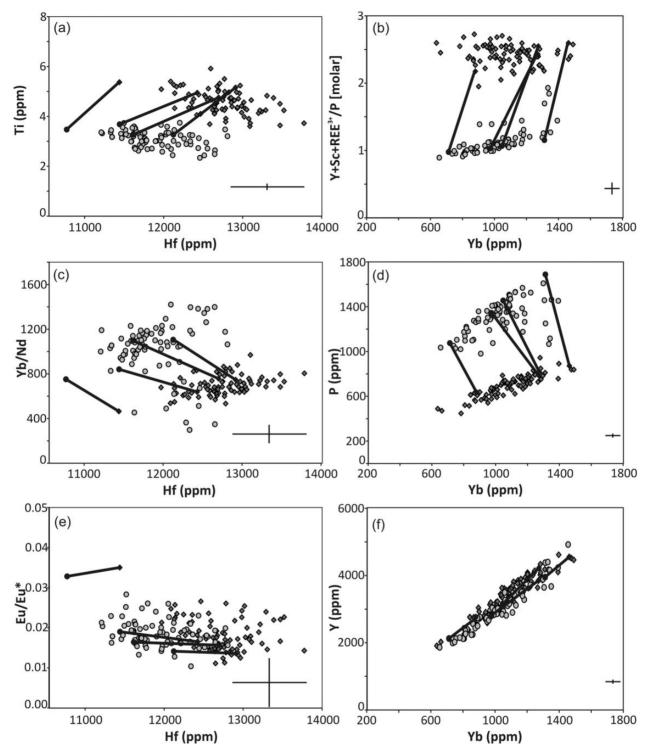


Fig. 11. Compositional variations within zircon crystals with textures dominated by sector zoning: dark grey diamonds represent darker tips; light grey circles represent lighter sides. Representative tie-lines connect sides and tips of single crystals. Cross in bottom right of each panel represents 2σ uncertainties. (For full dataset see Electronic Appendix 3.)

Table 5: Enrichment factors for the darktips in sector-zoned zircon crystals whencompared with the lighter sides

Element	Dark tip enrichment factor
Th	2.61
Nb	2.00
U	1.75
Nd	1.56
Sm	1.42
Ti	1.37
Gd	1.37
Eu	1.32
Dy	1.22
Tb	1.21
Но	1.35
Y	1.00
Yb	1.00
Tm	1.00
Sc	1.00
Lu	1.00
Hf	1.00
Er	1.00
Р	0.75
$(Y + Sc + REE^{3+})/P$	1.88

An enrichment factor of 1.00 shows no enrichment. Factor <1 indicates depletion in the dark tips; a factor >1 shows enrichment.

ratios (0.05-0.2) than the normal sector-zoned cores (0.01-0.04; Fig. 12c), indicative of a melt that had experienced less feldspar fractionation than the cores that the rims overgrew. However, although the absolute concentrations show variability, several element ratios show no variation between cores and bright rims; for example, Yb/Gd vs Th/U (Fig. 12d). The bright-rim analyses do not display as significant an influence of sector zoning on trace-element distributions; only a single trend in trace element patterns is defined, unlike the subparallel trends seen in 'normal' Bishop grains (see previous section and Fig. 11).

DISCUSSION

As emphasized by Hildreth (1977, 1979), zircon was among the first crystallizing mineral phases in the Bishop magma. Integration of the information available from single zircons is thus relevant to the origins and evolution of the Bishop system across the entirety of its lifespan as a felsic, zircon-saturated entity.

Zircon chronological information in the Bishop Tuff

Eruption age of the Bishop Tuff

Recent studies on the Bishop Tuff have raised issues regarding the eruption age, primarily from comparisons between ⁴⁰Ar/³⁹Ar age determinations on feldspar and ID-TIMS determinations on zircon (e.g. Sarna-Wojcicki et al., 2000; Crowley et al., 2007; Rivera et al., 2011; Singer et al., 2012). This apparent conflict arises from two independent aspects: (1) controversy over the appropriate values for the ⁴⁰K decay constant and consequent ages for the Fish Canyon (FCT) sanidine standard that is a benchmark for most young ⁴⁰Ar/³⁹Ar age experiments (e.g. Kuiper *et al.*, 2008; Renne et al., 2010, 2011; Rivera et al. 2011); (2) debate over the comparisons between single-crystal ID-TIMS age determinations and those obtained by spot analyses using SIMS techniques (Simon & Reid, 2005; Crowley et al., 2007). This second debate in turn focuses on three issues: (1) the contrasting precisions obtainable by each technique; (2) the possibility of inheritance of grains with significant residence histories, skewing age determinations and populations towards older values; (3) the three orders of magnitude contrasts in sample volume between the half- or whole-crystals analysed by ID-TIMS [(~l-3) $\times 10^{6} \mu m^{3}$] versus the small volumes [(~1.5- $2 \times 10^3 \mu m^3$ excavated by SIMS analysis (see Crowley et al., 2007; Simon et al., 2008; Reid et al., 2011; Reid & Schmitt, 2012).

We obtain a weighted mean age of 766.6 ± 3.1 ka (95%) confidence) for grains that we have analysed in cross-section as close to the rims as feasible. These grains, which lack CL-bright overgrowths, come from samples in early erupted material and hence represent the closest approximation we can make to measuring immediately pre-eruptive ages. Our mean value is the same, within uncertainty, as the weighted mean of $767 \cdot 1 \pm 0.9$ ka (2σ) from ID-TIMS work on single grains (Crowley et al., 2007), and the published ${}^{40}\text{Ar}/{}^{39}\text{Ar}$ age of 767.4 ± 2.2 ka (2σ) of Rivera et al. (2011), the latter based on a value for the FCT standard of 28.172 ± 0.028 Ma modified from Kuiper *et al.* (2008). Our data are, in turn, not in agreement with the older ages suggested for the Bishop eruption by Renne et al. [(2010, 2011; 778·1 \pm 7·4 ka (2 σ), based on recalculation of the original data of Sarna-Wojcicki et al. (2000)] and Mark et al. (2012; 776.4 \pm 1.8 ka, 2 σ). These differences are not bridged in our ²⁰⁶Pb/²³⁸U age determinations by the Th/U melt values adopted (Table 1) and its influence on the initial ²³⁰Th disequilibrium correction. Comparisons between the range of melt Th/U ratios we use versus the uniform 2.81 value used by Crowley et al. (2007) indicate that ²³⁰Th-corrected age values are 2-3 kyr younger for

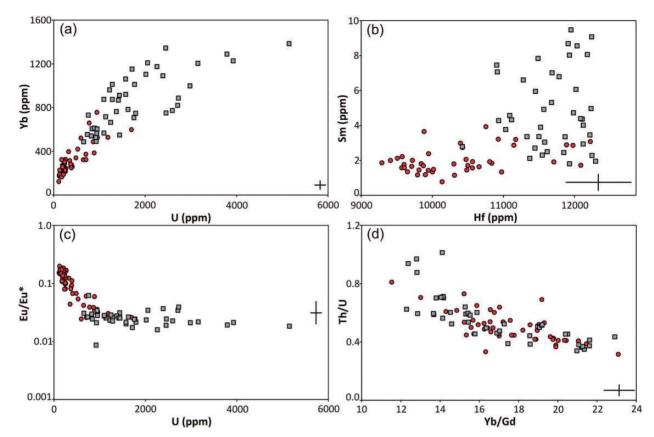


Fig. 12. Representative trace element compositional variations in CL-darker cores vs CL-brighter rims within Bishop zircon crystals. Cores are grey squares; rims are filled circles. Cross in bottom right of each panel represents 2σ uncertainties. (For full dataset see Electronic Appendix 3.)

Th/U = 2.5 (typical of early eruptive rocks) and 6-8 kyr older for Th/U = 4.0 (the maximum for later eruptive rocks: see Electronic Appendix 1).

Hence, despite the markedly different sampling strategies employed by SIMS and ID-TIMS, the results are identical within uncertainties. However, we note that there are some discrepancies in the datasets that could have produced systematic bias. The SIMS data clearly resolve the appropriate common Pb composition because there is no blank Pb to consider. As such, we clearly establish that the common Pb in the zircons (at the attogramfemtogram level) is magmatic and not surface related. The common Pb value that we obtain from the inverse concordia plot of our common-Pb-rich data points is consistent with the value from Stacey & Kramers (1975) used in SQUID 2, and also within uncertainty the same as the published bulk-rock value of 207 Pb/ 206 Pb = 0.818 derived by Halliday et al. (1989) and used by Simon & Reid (2005). Conversely, the ID-TIMS data rely on the common Pb being from the laboratory blank at the appropriate composition (see Crowley et al., 2007, their supplementary data table 1). In addition, our SIMS data utilize the Th/U compositions of the glass phase from the respective samples to make sample-specific Th/U fractionation corrections to the raw $^{238}\text{U}-^{206}\text{Pb}$ ages (see Table 3 and Electronic Appendix 1, which show the differences in age determinations between sample-specific Th/U ratios and a constant Th/U ratio of 2·81). On the other hand, the ID-TIMS data use an averaged Th/U value from the data of Anderson *et al.* (2000) with which to calculate the age correction for initial 230 Th disequilibrium. The coincidence of the final ages from the two techniques indicates that common Pb and Th/U fractionation is treated appropriately in both datasets.

Although the precisions we obtain are an order of magnitude poorer than those from single-crystal ID-TIMS work, the utility of SIMS instruments to undertake orders of magnitude more analyses in the same time span and to give spatially resolved ages within single crystals offers a complementary perspective that is, in this case, in agreement. As highlighted by Reid & Schmitt (2012), bulk ID-TIMS analysis is a whole-grain average, weighted by the U (and Pb) contents of each of the single zones that make up the crystal. Volumetrically small core zones could therefore be completely dominated by the large outer growth zones (Charlier & Zellmer, 2000). Another aspect that arises is comparisons between surface profiling techniques (e.g. Reid et al., 2011; Storm et al., 2011; Reid & Schmitt, 2012) and the cross-sectional techniques we have employed, with the outermost rims being argued to yield a 'true' U-Pb eruption age owing to there being no influence from older cores. Reid & Schmitt (2012) reported an average of SIMS surface profiling ages from Bishop zircons of 781 ± 22 ka (uncertainty inferred by us to be 2σ). With cross-sectional analysis we can thus replicate as good an approximation to eruption age as the uncertainties on the measurements allow, despite not analysing the outermost parts of any crystal. With both approaches (e.g. Charlier et al., 2005; Storm et al., 2011), however, there are commonly grains within samples that have inherited ages, whether through their recent (re-)introduction into the magma as antecrysts from the underlying mush pile or as xenocrysts from country rock. In either case, the scatter of ages obtained may reflect subtle issues of inheritance, not just position within the crystal or the uncertainties associated with single age determinations.

Temporal relationships between Glass Mountain and the Bishop Tuff

A previous tenet of the magmatic history at Long Valley is the temporal continuity and close compositional relationships between the younger series of eruptions from Glass Mountain and the Bishop Tuff. The YA dome K–Ar age of 790 ± 20 ka (Metz & Mahood, 1985) has been accepted by all workers and has been used to consider that there was temporal overlap of the Bishop and Glass Mountain systems. Our data suggest, in contrast, that the two systems are separate in time and distinct, with our estimate of 862 ± 23 ka for the eruption age of dome YA.

Other data also support an older age for the cessation of Glass Mountain activity (Sarna-Wojcicki et al., 2005). An averaged single-crystal sanidine 40 Ar/39 Ar age of 890 ± 31 ka [adjusted relative to a Taylor Creek sanidine standard age of 28.485 ± 0.033 Ma to match Rivera *et al.* (2011)] was reported by Sarna-Wojcicki et al. for their bed BT-16, underlying the Bishop Tuff at Chalk Bluffs. This sample was reported by them as yielding the youngest reliable (i.e. uncontaminated) age for pre-Bishop fall deposits at this location. The older eruption ages suggested by our zircon ages for domes YA and OD [the latter matched by the ${}^{40}\text{Ar}/{}^{39}\text{Ar}$ age determination by Davies *et al.* (1994)] imply that the entire Glass Mountain system may be significantly older than previously reported. A revision to the age spread for the Glass Mountain complex therefore could also suggest that the older ⁴⁰Ar/³⁹Ar dated tephras of Sarna-Wojcicki et al. (2005) in the Blind Spring Hill area were early products of the Glass Mountain system, rather than from an as-yet unknown older system in Long Valley.

In comparison, a weighted mean estimate for the oldest 10% of our acceptable Bishop zircon ages (all of which

are core analyses) is 844 ± 6 ka (n = 62, MSWD = 0.78, probability of 0.89; data points shown in Fig. 5). This age is indistinguishable from the older population of cores from the northern units derived by unmixing $(838.5 \pm 8.7;$ Fig. 7). We thus infer on the basis of age data, coupled with qualitative observations of the zircon textures, that the Glass Mountain and Bishop systems were not only temporally but also physically separate entities. Such a view is not in conflict with abundant evidence for their chemical continuity, on the basis of compositional and isotopic data for the younger Glass Mountain activity, from ~1·1 Ma onwards from published K-Ar ages (e.g. Halliday et al., 1984, 1989; Metz & Mahood, 1991; Bindeman & Valley, 2002; Simon et al., 2007). We would contend, however, that growth of the Bishop zircon population and assembly of the Bishop melt-dominant body post-dated and were separate from the youngest activity from the Glass Mountain system.

Assembly of the Bishop magma body

We interpret our age data to show that crystallization of the Bishop magma began from 844 ± 6 ka (as above) and continued up to the time of eruption, at 766.6 ± 3.1 ka. Previous workers have reported age ranges consistent with a prolonged crystallization history of zircons in the Bishop Tuff of ~160 kyr (Reid & Coath, 2000; Simon & Reid, 2005). Peak PDF estimates $(2\sigma \text{ uncertainties})$ of 823 ± 28 ka (early Bishop Tuff), and 811 ± 14 ka (late Bishop Tuff) can be calculated from the data of Reid & Coath (2000) and 841 ± 16 ka (early Bishop Tuff) from Simon & Reid (2005). It has been inferred that both the early and late-erupted Bishop units had a history of zircon crystallization extending as far back as 900 ka, with average maxima of 917 ± 27 ka for the early Bishop Tuff and 901 \pm 17 ka for the late Bishop Tuff (Simon & Reid, 2005), and a value of 892 ± 26 ka reported by Reid & Schmitt (2012). In contrast, we can see evidence for only ~80 kyr of crystallization history in the overwhelming majority of zircons within the tuff.

The older suite of ages we obtain (Fig. 8 and Electronic Appendix 1) from zircon cores, particularly from pumices erupted later and from the northern vents, are inferred to represent the onset of zircon crystallization within the Bishop mush zone; that is, the onset of crystallization for most of the mineral species represented in the Bishop pumices (Hildreth, 1979). What cannot be discerned from our data alone is whether this onset of zircon growth accompanied or pre-dated growth of the melt-dominant, stratified body. Overall growth rates for Bishop zircons implied by our age data are similar to the generalized 10^{-18} - $10^{-19}\,\mathrm{m\,s^{-1}}$ value for granitic melts reported by Watson & Liang (1995), and grains from the 844 ka average-age suite show no differences in texture from those zircons yielding ages within uncertainty of the eruption age. Where possible, zircons with a narrow (<10 µm) bright rim were dated at spots within the 'normal' sector-zoned Bishop interior, directly adjacent to the bright rim. These ages give the oldest age for the bright rim (in those thinly rimmed crystals), and thus the difference in the weighted mean of these ages (777.5 ka) from the proposed eruption age (766.6 ka) can be used as the maximum time for growth of a $10 \,\mu m$ (or less) rim. This age difference yields a maximum growth rate of $\sim 3 \times 10^{-19} \,\mathrm{m \, s^{-1}}$ (up to $1.3 \times 10^{-18} \,\mathrm{m \, s^{-1}}$, and as slow as $1.6 \times 10^{-19} \,\mathrm{m \, s^{-1}}$, with uncertainties), assuming that no dissolution or cessation of zircon growth occurred during that time. If the zircons across the age range from 845 to 766 ka grew in contrasting physical environments (i.e. a crystal-dominated mush versus a melt-dominant magma body) then these contrasts have left no distinctive textural or compositional signals.

Zircon compositional information in the Bishop Tuff

Implications of sector-zoning variations

The ubiquitous sector zoning observed in zircons from early erupted samples and in the cores of many crystals in samples from the northern ignimbrite lobes is in contrast to the more homogeneous CL textures generally documented in dated zircons (e.g. Corfu et al., 2003; Munoz et al., 2012). Experimental studies that investigated sector zoning and trace-element uptake in pyroxene, titanite and zircon have found that these minerals have apparent partition coefficients that vary depending on the crystal face (Paterson & Stephens, 1992; Watson & Liang, 1995; Lofgren et al., 2006; Schwandt & McKay, 2006). It has been proposed that if the crystal growth rate is fast enough to 'trap' trace elements at the crystal-melt interface, and the diffusivity of elements within the crystal lattice is not rapid enough to eradicate differences in concentration, sector zoning is generated (Watson & Liang, 1995; Watson, 1996). However, such conditions are almost always met in volcanic zircons regardless of whether or not sector zoning is visible in CL imagery. Sector zoning is not as obvious in the Glass Mountain or the bright-CL Bishop zircons, suggesting that there were differences in the crystal growth regimes when compared with the normal Bishop sector-zoned grains.

The differences in trace elements between sectors in the normal Bishop zircons raise questions about the relevant substitution mechanisms. Although the lighter sides have molar $(Y + Sc + REE^{3+})/P$ values of ~ 1 [the 'xenotime substitution' of Hoskin & Schaltegger (2003)], the darker tips have considerably higher values (Figs 1la and 13a), in large part owing to a depletion in P. Although it is beyond the scope of this paper to suggest an alternative substitution mechanism, it appears that the xenotime substitution may be operable only on certain crystal faces along the *a*- and *b*-axes (100 and 010 in the tetragonal

grains we are dealing with here) (Hoskin & Schaltegger, 2003). It is therefore possible that P fits better, and Y and the larger REE somewhat less well, into the lattice configuration along the 100 and 010 faces than the faces intersecting with the *c*-axis. An alternative charge-balancing scheme for crystallization in the dark sectors may need to be considered.

Ti concentrations and Ti-in-zircon thermometry

Our data show that the Ti concentrations in the same growth zone (as indicated by oscillatory zonation patterns) in sector-zoned zircons vary significantly between the sides and tips. This has strong implications for the applicability of Ti-in-zircon thermometry to the Bishop Tuff. As currently formulated, a single value of the activity of TiO_2 in the magma ($dTiO_2$) is adopted when applying the Ti-in-zircon thermometer (Watson & Harrison, 2005; Watson et al., 2006; Ferry & Watson, 2007; Fu et al., 2008). Five estimates for the value(s) of $dTiO_2$ have been proposed for the Bishop magma: 0.6, estimated by Watson et al. (2006) and Wark et al. (2007); 0.53 as calculated from Fe-Ti-oxide equilibrium modelling (Thomas et al., 2010; Reid et al., 2011); 0.15 using MELTS and 0.23 using rhyolite-MELTS (Thomas & Watson, 2012); and 0.45-0.67 calculated from Fe-Ti-oxide compositions by Ghiorso & Gualda (2013). An additional universal assumption is that the value of $aSiO_2$ is 10, based on the abundance of crystallizing quartz in the Bishop Tuff.

Using the calibration of Ferry & Watson (2007), we calculated apparent temperatures (on 101 side-tip pairs) from measured Ti concentrations in the same growth zone. These calculations yield average apparent temperatures for the darker tips between $712^{\circ}C$ ($dTiO_2 = 0.67$) and $863^{\circ}C$ (*a*TiO₂ = 0.15) depending upon which value of aTiO₂ is used. In contrast, the lighter sides yield average apparent temperatures of 686° C (*d*TiO₂ = 0.67) to 829° C $(aTiO_2 = 0.15)$ (Electronic Appendix 3). This gives an average apparent temperature difference of 20-50°C for different sectors within the same crystal growth zone (Fig. 13a). If a constant T of 750°C is assumed the difference in calculated $aTiO_2$ between the darker tips ($aTiO_2 = 0.44$) and the lighter sides $(aTiO_2 = 0.34)$ is on average ~0.1. This difference in $dTiO_2$ increases at cooler temperatures, but is reduced at temperatures $>750^{\circ}$ C.

Ti concentrations also vary between bright overgrowths and their respective cores (neither of which are visibly sector zoned in order for these variations to be estimated). The bright rims produce average apparent temperatures between 718°C ($d\text{TiO}_2 = 0.67$) and 871°C ($d\text{TiO}_2 = 0.15$) with an average range in apparent temperatures of 87°C across all bright rims. Cores yield apparent temperatures between 697°C ($d\text{TiO}_2 = 0.67$) and 805°C ($d\text{TiO}_2 = 0.15$) with an average range in apparent temperatures of 77°C across the cores. The trends in Ti–U space of cores and rims are very different, both apparently evolving to

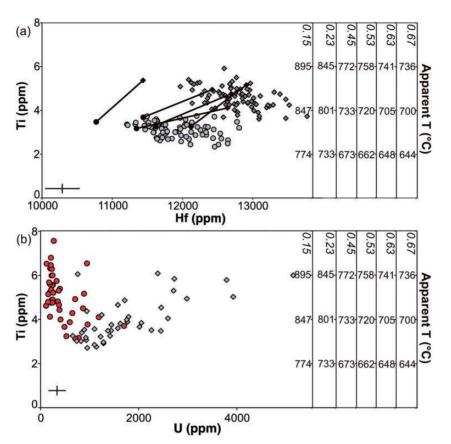


Fig. 13. Summary of results from Ti-in-zircon thermometry from Bishop zircons. (a) Comparisons of model temperatures obtained from Ti concentrations between different sectors of the same crystals: light grey circles are lighter sides; dark grey diamonds are darker tips. (b) Differences between CL-darker cores and CL-brighter rims of the same crystals: filled circles represent rims; grey diamonds represent cores. The *j*-axis is labelled with the measured Ti concentrations (ppm) on the left, and the resulting model temperature scales on the right based on the multiple published estimates of dTiO₂ (given in italics) proposed for the Bishop magma (see text for discussion). Cross represents 2σ uncertainty. (For full dataset see Electronic Appendix 2.)

(or from) a similar point (Fig. 13b). These calculated apparent temperatures for sectors and core–rim pairs are distinctly lower than those calculated using Fe–Ti-oxide thermometry (Hildreth, 1979; Hildreth & Wilson, 2007), two-pyroxene thermometry (Hildreth & Wilson, 2007) and Ti-in-quartz thermometry, when the same values of dTiO₂ are used (Wark *et al.*, 2007; Thomas *et al.*, 2010).

Considering Ti concentration alone, the Bishop zircons display a relatively limited range in Ti concentrations (2–8 ppm) when compared with other large silicic eruptions, such as the Oruanui (2–14 ppm), Ongatiti (3–18 ppm) and Kidnappers (2–18 ppm) examples in New Zealand (C. J. N. Wilson *et al.*, unpublished data). The nano-scale Ti variability reported by Hofmann *et al.* (2013), if generally present, would raise further issues over the measurement of Ti concentrations in zircon and interpretation of these values in terms of entry of Ti into the zircon structure (e.g. as nano-inclusions versus as a structural component) and consequent temperature estimates. However, the great contrast in the sampling volumes of the NanoSIMS versus ion probe means that the two suites of data are not yet comparable.

Contributing components to the zircon population in the Bishop Tuff magma

By combining the textural, age and trace element datasets it is clear that there are at least four components that have contributed to the zircon assemblages in the Bishop Tuff: (1) the magma that grew the sector-zoned 'normal Bishop' crystals, which are the dominant population in most of the Bishop samples examined (Table 2; Fig. 4); (2) the magma that caused the growth of bright rims (and sparse whole crystals) in the zircon population; (3) the xenocrystic crustal component, mostly from Triassic granitoids, and present at up to 20% of the zircon population in two samples (Table 2; Fig. 4); (4) the Glass Mountain component, which dominates the YA dome sample but forms only a minor component (6% maximum) of grains in the Bishop samples.

The top of the magma chamber (sampled by Fl and IglEb) is dominated by the normal Bishop zircons, rapidly grown (reflecting the sector zoning) and with no bright overgrowths. This situation highlights a lack of mixing within the upper parts of the magma chamber, consistent with the limited ranges in corresponding Bishop Tuff phenocryst and whole-rock compositions (Hildreth, 1979; Hildreth & Wilson, 2007). These early erupted zircons have trace element variations dominated by sector zoning and ages close to the eruption age (cores of Fl and IglEb vary dominantly from 760 to 820 ka). However, it is apparent from their relatively restricted Th/U ratios, only mildly affected by sector zoning (Fig. 10), that the source for the 'normal Bishop' zircons involved melting of an arcsourced lithology (Miller & Wooden, 2004; Claiborne et al., 2010). The trends evident in U vs Yb (Fig. 10) and most other element-element and element-ratio plots, in which element concentrations are increasing with fractionation, are likely to reflect continuing high-percentage fractionation of guartz and feldspar (up to 90% of the erupted crystal assemblage; Hildreth, 1979), which makes most of the trace elements common to zircon behave incompatibly in the melt. We cannot find evidence for allanite fractionation influencing zircon trace element patterns, as there is no decrease in U and Th concentrations with increasing Hf (Fig. 10). Allanite fractionation has been proposed to explain a coupled decrease in light REE (LREE) and Zr/Hf in zircon (Reid et al., 2011). Allanite is not present, however, in the late Bishop Tuff and the extremely low abundances of LREE in any zircons make monitoring LREE changes challenging.

With progression through the eruption sequence (i.e. deeper in the magma chamber) the zircon textures become more variable, with the magma responsible for the bright zircon rims first becoming present in significant amounts in the latest IglEb deposits (Sherwin subunit). It had previously been thought that the bright rims on quartz and sanidine crystals (Hildreth, 1979; Anderson et al., 2000; Peppard et al., 2001; Wark et al., 2007) were restricted to the northern units. Recent investigations of quartz-hosted melt inclusions has, however, revealed a more complex history (Roberge et al., 2013). This study documented changes in melt inclusion compositions, from less evolved than the host melt, to more evolved than the host melt, taken as evidence of melt mixing occurring at depth within the magma chamber. This pattern of less evolved glass coupled with more evolved melt inclusions is not, however, restricted to the northern-erupted Bishop Tuff, but occurs also in the mid- to late-erupted eastern units (Ig2Ea, Ig2Eb). When coupled with the presence of the bright rims on zircons in late IglEb, these observations imply that the introduction of a new magmatic component into the chamber caused growth of the zircon bright rims.

Samples from the later-erupted Bishop units (Ig2N, Ig2NW) contain a smaller proportion of sector-zoned zircons, with sparse crystals of Glass Mountain type (Table 2; Fig. 4). These contrasts suggest that zircon suites in the later-erupted units of the Bishop Tuff, inferred to come from deeper in the chamber (Wallace *et al.*, 1999), incorporate some influence from a longer-lived part of the magma system. Some minor inheritance from the Glass Mountain system is also inferred to be present on the basis of textural observations, although this is not reflected in the age spectra. If such grains are from a source common to the Glass Mountain system, then magmatism continued after eruption of the youngest Glass Mountain domes.

The later-erupted deeper-sourced magma, which contains zircons that began to crystallize at ~845 ka, produced grains that do not show obvious sector zoning in CL, but are similar in trace element composition to those found in the early deposits (including some sectoral zoning in trace elements such as P and Ti). Our data indicate that the melts from which the sector-zoned and non-sector-zoned zircons grew were broadly similar in composition and had experienced comparable degrees of fractionation. This cogenetic nature of the early versus late-erupted zircons is consistent with the unitary nature of the Bishop Tuff magma body (Hildreth, 1979; Hildreth & Wilson, 2007) and counts against any notion of two separate magma bodies (Gualda *et al.*, 2012*a*).

The trace to subordinate component of Triassic zircons, most abundant in the swirly pumice type (Hildreth & Wilson, 2007), indicates a late-stage incorporation of country-rock components into the magma (see also Hildreth, 1985). The Triassic zircons have a lower range in Th/U compared with the normal Bishop-type zircons (Fig. 10), indicating that melts of Triassic crust have not been a source for the main Bishop Tuff magma body. The lack of overgrowths on any of the zircon crystals implies that xenocrystic incorporation was very late stage, most probably syn-eruptive.

Origin of the 'bright-rim' magma

Development of the CL-bright zircon rims is linked to changes in melt inclusion chemistry (Wallace *et al.*, 1999; Anderson *et al.*, 2000; Roberge *et al.*, 2013), and bright overgrowths found on quartz (CL imagery) and sanidine (BSE imagery) in the northern-erupted units (Anderson *et al.*, 2000; Peppard *et al.*, 2001; Morgan & Blake, 2006; Wark *et al.*, 2007; Gualda *et al.*, 2012*b*). It has been suggested that varying growth conditions (i.e. changes in *T*, *P*, fO_2) could have caused crystallization of these bright rims on quartz, sanidine and zircon (Anderson *et al.*, 2000; Peppard *et al.*, 2001; Wark *et al.*, 2007; Thomas *et al.*, 2010). However, our trace-element data from bright rims show low concentrations of elements such as the HREE, U and Th compared with those of the sector-zoned zircons, indicating that these species were at lower concentrations within the parental melt. Using the partition coefficients of Sano et al. (2002) on the rims and cores of the same sectors indicates that the melt that crystallized the bright zircon was higher in Eu, but lower in the MREE and HREE than the melt associated with the sector-zoned zircons, assuming constant partition coefficients for both growth regimes. The difference between the bright rims and normal cores is more variable in Ig2Nb when compared with earlier-erupted units. This 'bright-rim' melt had Th/ U ratios of 0.3-0.6, similar to the normal Bishop sectorzoned cores, which indicates that both the 'bright-rim' melt and the 'normal Bishop' melt involved the melting of a similar, arc-lithology source, consistent with the low U/ Yb values of less than 10 (Fig. 10; Electronic Appendix 3). The lower values of HREE agree with previously published melt-inclusion and glass data (e.g. Anderson et al., 2000; Roberge et al., 2013), showing that these changes in the zircon trace elements match the evolution of host melt compositions. Increases in Ti concentration in the bright overgrowths yield $\sim 20^{\circ}$ C higher apparent temperatures over the cores (independent of the value of $aTiO_2$ used), possibly showing that the magma was hotter. This is in

agreement with Fe–Ti-oxide thermometry, which shows that the eruption temperatures of the melts are higher in the later units (~815°C versus ~715°C; Hildreth & Wilson, 2007). However, owing to the highly incompatible nature of many key trace elements in zircon and the obvious overprinting from the sector zoning, more detail about the nature of this 'bright-rim' melt cannot be derived from zircon chemistry alone. Further investigations are currently under way into the nature of the 'bright-rim' magma and its relationship to the minor less-evolved components in the Bishop Tuff (Hildreth, 1979; Hildreth & Wilson, 2007).

Introduction of the bright-rim forming magma has previously been considered to be on such a short timescale (<100 years) that it was a possible trigger for the Bishop Tuff eruption (Wark *et al.*, 2007). By calculating the weighted means of our age determinations from directly inside the bright rim it is apparent that these zones are systematically older than the true rim analyses (whether they are 'bright rim' or 'normal'; Fig. 14). However, there is no linear relationship between the bright-rim width and age inside the rim, so we cannot determine absolutely

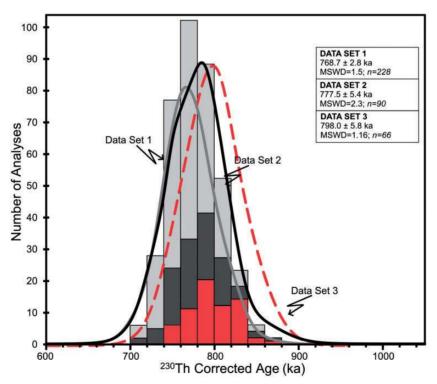


Fig. 14. Variation in ²³⁰Th-corrected ²⁰⁶Pb/²⁰⁸U ages measured in 'normal Bishop' zircons (as defined from CL imagery; see text and Fig. 3). Dataset 1 (light grey) combines all age determinations (this study) on rims of 'normal Bishop' zircons where there is no development of any CL-bright rim overgrowth. Dataset 2 (dark grey) represents measurements on Bishop zircons that have narrow ($<10 \mu$ m) bright rims, where the spots were sited in the darker CL cores just inside the brighter rim. Dataset 3 (intermediate shading, dashed line) represents measurements on Bishop zircons that have wide (>10 µm) bright rims, where the spots were sited in the darker CL cores just inside the brighter rim. Probability density function (PDF) curves from Isoplot. Values for the datasets represent the number of spots used (*n*), and the weighted means, uncertainties at 95% confidence and MSWD values for the respective populations. (For full dataset see Electronic Appendix 1.)

when the bright-rim magma started to invade the magma chamber, apart from that it must be within 10 kyr of the eruption age, based on ages within the cores of crystals with wide bright rims (Fig. 14). There are therefore two possibilities for the timing of interaction with this 'bright-rim' melt. The first is that the systematic differences in rim width are due to a constant zircon growth rate, and therefore the increase in mean ages with bright-rim width (Fig. 14) shows that the 'bright-rim' magma was interacting with the roots of the Bishop system under the northern side of what became Long Valley caldera for a longer period of time. However, this is inconsistent with quartz diffusion timescales that indicate <100 years of interaction with the 'bright-rim' magma (Wark et al., 2007; Gualda et al., 2012b). The second possible interpretation of the differences in timescales with rim width is that it reflects a cessation of zircon growth (but no dissolution, hence the euhedral core morphologies). Therefore, in the deepest parts of the magma chamber (where rims are typically thicker; see Fig. 4) the magma was close to (or just above) zircon saturation, and zircons had ceased growing until interaction with the 'bright-rim' magma triggered renewed growth. In the upper parts of the magma system, where bright rims on zircon are rare and thin, the magma remained zircon saturated (and therefore zircon continued to grow) until the shortly pre-eruptive interaction with the 'bright-rim' magma (Wark et al., 2007).

To assess if introduction of the hotter 'bright-rim' magma would allow crystallization of the observed bright rims, we calculated the state of zirconium saturation in the melt, using the Watson & Harrison (1983) calibration. In the samples dominated by 'normal' Bishop zircons the Zr saturation temperatures are within uncertainty of the published Fe–Ti-oxide and $\Delta^{18}O$ (Qz–Mt) temperatures, and therefore the melt was zirconium saturated, at least immediately prior to eruption (Hildreth, 1979; Bindeman & Valley, 2002). However, in the three samples that display high proportions of bright rims the Zr saturation temperatures are up to 50°C lower than the Fe-Ti-oxide thermometry for those units, implying that the 'bright-rim' melt was slightly Zr undersaturated. This result would seem to indicate that the 'bright rims' must have grown before those modelled in guartz. However, Harrison et al. (2007) proposed that zircon can crystallize at up to 80°C higher than the zircon saturation temperature calculated from Watson & Harrison (1983). In addition, it is unlikely that the zircon bright rims are unrelated to the bright rims observed in quartz and sanidine, given the similar associated trace element changes (Hildreth, 1979; Anderson et al., 2000; Wark et al., 2007; Simakin & Bindeman, 2008).

CONCLUSIONS

This study demonstrates some key findings that are not only relevant to the development and eruption of the Bishop Tuff magma body, but that also raise important considerations for zircon studies on any silicic system.

- (1) A comprehensive suite of U-Pb ages from the earliest to the latest units of the Bishop Tuff reveals no evidence for a long 'tail' to the precursory Glass Mountain system, and we cannot replicate published peaks in age spectra greater than ~845 ka. Our nearest-rim analyses in samples from early erupted material yield a weighted-mean estimate for the eruption age of 766.6 ± 3.1 ka (95% confidence), fully consistent with the ID-TIMS estimate of Crowley et al. (2007) and ⁴⁰Ar/³⁹Ar age determinations of Rivera et al. (2011). Our data are not consistent with proposals of older ages for the Fish Canyon standard that would have the Bishop eruption age in the 776-778 ka range (Renne et al., 2010, 2011; Mark et al., 2012). SIMS chronologies on zircons, at least in this key example, can match in overall accuracy and precision those obtained by ID-TIMS, with the advantage of being able to link analytical spot positions to textural features within grains.
- (2) Zircon age data from the youngest Glass Mountain dome (YA) are consistent with an eruption age of 862±23 ka, in contrast to the universally used K-Ar value of 790±20 ka (Metz & Mahood, 1985). Our data, coupled with zircon ages and ⁴⁰Ar/³⁹Ar age determinations from Glass Mountain dome OD (Davies *et al.*, 1994; Simon & Reid, 2005), imply that the K-Ar chronological framework of Metz & Mahood (1985) may be significantly inaccurate.
- (3) Our age data show that the Bishop Tuff magma and mush systems were physically separated from the Glass Mountain system and must have evolved in separate fashion. Textural data imply that there was very little (<10%) zircon crystal inheritance from the Glass Mountain roots even in the deepest parts of the magma chamber. Our ages imply that the Bishop magma system began its operations with zircon-saturated compositions, only to begin crystallizing zircon ~80 kyr prior to eruption. This is not necessarily the assembly time for the melt-dominant magma body, but the lack of systematic textural contrasts between zircons spanning this ~80 kyr range suggests that this was the case.
- (4) The Bishop zircons can be classified into four categories: (a) a numerically dominant suite of sector-zoned grains; (b) a subordinate suite of grains with rims of CL-bright material (and cores that may or may not show sector zoning); (c) a sparse suite of grains with oscillatory zoning and shapes that match the dominant population in Glass Mountain dome YA; (d) oscillatory zoned Mesozoic crustal xenocrysts. We demonstrate that a major control on trace-element variations in the Bishop zircon population is imposed

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by sector zoning. This sector zoning is most dominant in crystals from the early Bishop Tuff units, which are representative of the archetypical and numerically dominant 'normal Bishop' type of zircon. Application of Ti-in-zircon model temperatures in particular is hindered both by the sector zoning (differences of $20-50^{\circ}$ C between sides and tips in the same growth zone in the same crystals) and by contrasting values proposed for the value of $d\text{TiO}_2$ (0·15–0·67).

(5) Despite the strong influence of sector zoning, key magmatic features can still be identified. By combining textural observations from CL imagery and trace-element analyses, two major end-member magmas for the Bishop Tuff are identified, one responsible for the 'normal Bishop' sector-zoned zircons and the other responsible for the CL-bright components. The 'bright-rim' component was a less evolved latestage contributor to the Bishop system, and it also influenced growth of bright rims on sanidine and quartz. However, 'rejuvenation' of the Bishop magma chamber is observed to reach levels as shallow as those tapped by the latest IglEb unit, earlier than has been previously identified. The sources for the Bishop Tuff magmas (whether 'normal' or that responsible for the bright rims) appear to involve an evolved arc source, which subsequently experienced significant amounts of quartz and feldspar fractionation.

ACKNOWLEDGEMENTS

We thank George Cooper, Wes Hildreth, Dan Morgan, Joel Baker, Joe Hiess and Millie Wilkinson for valuable discussions and advice at various stages of this work, and Matt Coble, John Foster, Peter Holden, Brad Ito and Jorge Vazquez for their help in the ion probe laboratories at Stanford University and the Australian National University. K.J.C. would like to thank Diane Seward, Nick Burrows and Simon Barker for their help during zircon separations. Ilya Bindeman, John Wolff and an anonymous reviewer are thanked for their constructive reviews, and Richard Price is thanked for editorial handling.

FUNDING

K.J.C. was supported by an NZ International Doctoral Research Scholarship administered by Education New Zealand, and acknowledges with thanks a Jack Kleinman Award administered by the US Geological Survey and two Victoria University of Wellington Faculty Strategic Research Grants. C.J.N.W. acknowledges support from the Royal Society of New Zealand through Marsden Fund grant VUW0813 and award of a James Cook Fellowship.

SUPPLEMENTARY DATA

Supplementary data for this paper are available at *Journal* of *Petrology* online.

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