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Nonequilibrium in a Low Power Arcjet Nozzle

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Abstract

Emission spectroscopy measurements were made of the plasma flow inside the nozzle of a 1 kW class arcjet thruster. The thruster propellant was a hydrogen-nitrogen mixture used to simulate fully decomposed hydrazine. 0.25 mm diameter holes were drilled into the diverging section of the tungsten thruster nozzle to provide optical access to the internal flow. Atomic electron excitation, vibrational, and rotational temperatures were determined for the expanding plasma using relative line intensity techniques. The atomic excitation temperature decreased from 18,000K at a location 3 mm downstream of the constrictor to 9,000K at a location 9 mm from the constrictor, while the molecular vibrational and rotational temperatures decreased from 6,500K to 2,500K and from 8,000K to 3,000K, respectively, between the same locations. The electron density, measured using hydrogen H_B line Stark broadening, decreased from ~10¹⁵ cm⁻³ to ~2 x 10¹⁴ cm⁻³ during the expansion. The results show that the plasma is highly nonequilibrium throughout the nozzle, with most relaxation times equal or exceeding the particle residence time.

Nomenclature

Cconstant for calculating vibrational relaxation
c _p specific heat capacity, J/kg/K
Eexcited state energy, J
E _H hydrogen ionization energy, 13.6 eV
foscillator strength
gstatistical weight
Gvibrational excitation energy level, J
GeffGaunt factor used for calculating electron densities
Hneutral atomic hydrogen
I intensities integrated over the complete spectral line or band

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k.....Boltzmann constant, 1.38 x 10⁻²³J/K K2.....constant for calculating vibrational relaxation λ.....wavelength, nm $ln(\Lambda)$coulomb logarithm m.....particle mass, kg M.....particle mass, amu NI, NII.....neutral and singly ionized atomic nitrogen n_e....electron density, cm⁻³ p.....pressure (eqn. 6), atm σ.....cross section, cm² τ.....relaxation times for temperature relaxation, sec T....electron temperature, K Tex....atomic electron excitation temperature, K T_{rot}.....rotational molecular temperature, K T_{vib}.....vibrational molecular temperature, K v.....particle velocity, m/s z.....degree of ionization +1

I. Introduction

Selection of 1 kW class arcjet thrusters for station keeping application on the Telstar IV series satellites has dramatically increased interest in arcjet technology. The ability to achieve specific impulses from 450 to 550 seconds with storable propellants, and the system simplicity makes arcjets attractive for application to geosynchronous satellite station keeping. Furthermore, modern satellite power busses have the capability to provide the required power/duty cycle combination for these thrusters. While arcjets have been accepted by the satellite community for geosynchronous stationkeeping, there is considerable room for improved thruster performance. Current arcjets convert only 25-35% of the input power into thrust power, with the dominant losses lying in unrecovered power deposited into the propellant. This paper describes an effort to empirically study the nozzle plasma in the expectation that a detailed understanding of the dominant physics will permit improvements of the arcjet.

In a typical arcjet, shown in Fig. 1, an arc is struck between the central, conical-tipped cathode and the coaxial, nozzle shaped anode. Propellant, injected with high swirl near the cathode tip, passes through the constrictor region (throat) and is ohmically heated by the arc. Curran¹, using an axially segmented arcjet anode, showed that the anode current attachment extends all the way down the length of the nozzle. This result demonstrated that energy is transferred to the gas throughout the entire flow passage, including subsonic, transition, and supersonic flow regimes. The energy transfer to the gas is thought to result predominantly from electron-ion or electron-neutral collisions as electrons are the dominant current carriers.

These collisions distribute the arc energy to several sinks, including random thermal

energy of the propellant, dissociation, ionization, and excitation of bound rotational, vibrational and electronic states. Energy deposited into these modes which is not converted to directed kinetic energy of the exhaust constitutes the major loss mechanism for arcjet thrusters. While an analytical understanding of these processes is clearly required, the complexity of the flow field and plasma physics presently precludes an ab initio approach. To provide improved insights into the internal arcjet flows, efforts were initiated to develop the diagnostic techniques required to directly measure several of the principal plasma parameters.

Most studies of arcjet plasmas have concentrated on the thruster plume. Studies utilizing Langmuir probes^{2,3} and emission spectroscopy⁴ have established electron density and temperature profiles and identified the dominant species in the plume downstream of the exit plane. Janson, et.al⁵ found differences between the plume spectrum of an arcjet tested with ammonia and one tested with a nitrogen-hydrogen mixture chosen to simulate ammonia. While this indicates that the mixture used in this study to simulate hydrazine may not exactly represent a hydrazine arcjet, it should be noted that in a flight-type system, hydrazine is decomposed in a catalyst bed before being injected into the arcjet, and performance comparisons⁶ indicate only small differences between simulated and actual hydrazine propellants. The only previous study of the internal arcjet plasma was performed on a low power helium arcjet with a 2mm diameter quartz constrictor⁷. However, due to the temperature restrictions of the quartz and the impact of the window on the energy transfer processes in the constrictor, it would be unrealistic to relate their results to flight-type arcjet thrusters.

This paper presents the results of an effort to non-intrusively measure the plasma properties in an arcjet thruster similar to those planned for flight application. Following a description of the experimental apparatus and thruster operating characteristics, the measurements and their analysis are presented, including identification of the plasma constituents, atomic and molecular temperatures, and the electron density. The results are then used to estimate various plasma relaxation rates. Finally the implications of the results for energy transfer processes and plasma models are discussed.

II. Experimental Apparatus

The thruster used in this experiment was a 1 kW class laboratory model device developed at the NASA Lewis Research Center⁸. The thruster was similar to those planned for application on geosynchronous communications satellites. The arcjet, shown in Fig. 1, consisted of a conical-tipped, 2% thoriated tungsten cathode, a 0.064 cm diameter by 0.064 cm length constrictor, and a conical nozzle/anode with an expansion angle of 20° and an area ratio of 225:1. Additional testing was performed with a nozzle expanding to an area ratio of 400:1 (same expansion angle). The anode material was 2% thoriated tungsten and all insulators were made of boron nitride.

The thruster was operated in a 0.64m diameter by 0.64m length vacuum chamber

evacuated using a mechanical roughing pump. The propellant flow consisted of a mixture of 59 mg/s hydrogen and 41.6 mg/s nitrogen (to simulate fully decomposed hydrazine) and was kept constant during all experiments. With this flow rate the pump maintained a back pressure of about 120 Pa. This back pressure, while affecting the plume, changed the thruster performance (thrust and voltage) by less than 5%. The power supply delivered a stable current that was varied between 9 and 12 A for the experiments. The thruster performance was measured on a thrust balance in a separate separate facility with higher pumping speed described in Ref. 1.

Two optical diagnostic techniques, described in Ref. 10, were developed to measure the emission spectra of the internal nozzle plasma. The more successful technique consisted of drilling 0.25 mm diameter holes radially into the nozzle wall at desired axial and radial positions to gain optical access to the plasma inside the nozzle. As described in the next section, these holes had no measurable impact on the arcjet performance and there was no evidence of preferential current attachment around the holes. The hole locations, shown in Fig. 2, were chosen to provide measurements at three area ratios in the 225:1 nozzle and to obtain coarse radial profiles in both the 225:1 and 400:1 nozzles. The radial profiles were obtained at only one axial location with the 225:1 nozzle and two axial locations in the 400:1 nozzle. While a simple Abel inversion could be applied to the radial measurements, the resolution was limited to 1/4 the nozzle radius by the small nozzle size.

The optics required to image the individual holes onto the spectrometer entrance slit are shown in Fig. 3. Two lenses were required to satisfy the spatial and magnification requirements and an iris aperture was used to eliminate stray light emitted from the glowing nozzle wall. In order to select an individual hole, the thruster was moved vertically inside the vacuum chamber on a stepper motor driven rail table until the desired hole was at the image location of the light path.

The 0.5 m Czerny-Turner spectrometer was equipped with a 2400 grooves/mm grating and an intensified 1025 pixel diode array detector to measure line intensities. The spectral resolution of the instrumentation was 0.016 to 0.019 nm per detector pixel, depending on the observed wavelength. Spectra were observed in real time to permit adjustment of the diode array integration time before storing them for later analyses.

III. Experimental Results

The current - voltage characteristic of the arcjet was obtained at the beginning of each measurement series to ensure the thruster was operating normally. Thrust measurements showed that the presence of the 0.25 mm diameter holes in the nozzle wall had no measurable effect (≤ 2%) on the thruster performance. This was verified by testing the thruster with the nozzle wall holes both sealed and open. These data, obtained with the 225:1 nozzle with axial holes, are summarized in Table 1. Excellent agreement was obtained with the lifetest results reported by Curran⁸.

In addition to the optical system complexity and small nozzle size, the measurements were

hindered by light reflection from the nozzle wall. This issue is distinct from that of the glowing nozzle wall contributing continuum radiation, and is due to plasma luminosity reflecting of the inner nozzle surface into the holes through which the measurements were made. This precluded a precise determination of the region where the light originated. The magnitude and origin of this light was determined by the nozzle geometry and surface reflectance. To establish the importance of this effect a small LED was slowly inserted up into the nozzle with the axial holes (geometry A in Fig. 2) and the light intensity emitting from the holes was measured as a function of the LED location. For each hole there were two axial positions at which the LED provided a measurable signal: the primary location when the LED was directly beneath the hole and the secondary corresponding to the location that caused direct reflection into the hole. The secondary peak was brightest for the hole closest to the exit plane, but never exceeded 30% of the intensity of the direct illumination. It is important to note that this experiment utilized a constant intensity source, whereas the plasma luminosity decreased rapidly in the downstream direction. While this intensity reduction mitigated the importance of reflected light, it was not possible to exactly quantify the impact of this problem.

Whenever possible the quantitative results were obtained from spectra separated by less than 40 nm to avoid the difficulties associated with intensity calibrations. For all measurements of steady-state conditions the thruster was turned on and allowed to stabilize for 15-20 minutes before a final alignment check was made to account for thermal expansion of the nozzle and arcjet components. Measurements were performed at least three times to ensure repeatability. The uncertainties of the results were calculated by propagating the errors of the individual parameters using standard algorithms¹¹.

Plasma Species

The plasma species identified from a spectrum recorded between 320 and 750 nm are listed in Table 2. Only species for which several lines were identified are reported. Observed wavelenghts were compared with tabulations in Ref. 12, 13 and 14. In addition to the molecular and atomic propellant species, tungsten impurity, NH, and NH₂ lines were observed. The high NH intensities observed at all axial positions indicates that recombination was taking place throughout the nozzle. Although the NH₂ and tungsten spectral lines were visible in most experiments, their intensities were too low to use them for more detailed investigations.

Plasma Properties

The spectral data were used to determine several plasma temperatures and the electron density. The temperatures included the atomic excitation temperature for NI, NII, and H, and molecular vibrational and rotational temperatures for N_2 . All temperatures were calculated assuming that the populations of the observed excited states followed a Boltzmann distribution at the local temperature. The validity of this assumption was checked a posteriori.

Atomic Excitation and Free Electron Temperatures

The atomic excitation temperature T_{ex} , was determined from emission line intensity ratios. This temperature can be calculated from two different lines (denoted 1 and 2) of the same

species¹⁵:

$$T_{ex} = \frac{E_1 - E_2}{k \ln[I_2 \lambda_2^3 g_1 f_1 / I_1 \lambda_1^3 g_2 f_2]}$$
(1)

The spectral lines used for this measurement are listed in Table 3, along with their oscillator strengths and the degeneracies of the upper excited states. A relative calibration was needed for the hydrogen measurements due to the wide wavelength separation of the observed transitions.

The axial excitation temperature distributions for NI, NII and H for thruster operation at currents of 9 and 12 amps with the 225:1 nozzle are shown in Figures 4a - c. While the absolute uncertainties are large, most of the error was due to the oscillator strengths, which does not affect the validity of the observed trends. The excitation temperatures for the three species were similar, dropping from ~ 18,000 K at an area ratio of 1:10 to ~ 10,000 K at an area ratio of 1:125. The excitation temperature increased slightly with arc current.

Radial electron excitation temperature profile measurements were hindered by the rapid decrease in line intensities away from the nozzle axis. For this reason, only the hydrogen excitation temperature could be measured. The measurement was made at four radial positions in both the 225:1 and 400:1 area ratio nozzles. The radial holes divided the nozzle section into four shells with an equal thickness of one fourth of the radius. The Abel inversion technique used assumed a uniform intensity in each of the individual ring shells, and the intensity in a radial ring was calculated from the geometric size of the observed shell portion¹⁶. This simple technique was used because the small nozzle dimensions precluded measurements at a large number of radial locations. The measured spectra were inverted and the plasma properties were determined from the inverted intensities. As shown in Fig. 5a and b the hydrogen excitation temperature did not vary within the inner 87% of the nozzle radius for either nozzle.

Background radiation from the glowing nozzle precluded direct measurement of the free electron temperature from continuum radiation. However, several arguments can be introduced to show that the electron temperature was near the atomic excitation temperature. First, there is the close agreement of excitation measurements for NI, NII, and H. Given the different cross-sections, radiative rates, and densities, it seems unlikely this agreement would occur unless the upper excited state populations were controlled by the same mechanisms of electronic collisional excitation and radiative decay. Second, the energy levels used for the measurements were close to the ionization potentials for the atoms. For example, for NI the measured states were less than 0.6 eV from the continuum. Third, the electron densities measured using Stark broadening were close to the threshold established by Griem¹⁵ for partial local thermal equilibrium, a condition for which $T_{\rm ex} = T_{\rm e}$. The latter result is further discussed in section IV.

Molecular Vibration Temperature

The nitrogen vibrational temperature was determined from the intensity ratio of two transitions. According to Herzberg¹² it can be calculated from two molecular bands (denoted 1

and $_2$):

$$T_{vib} = \frac{G_1 - G_2}{k \ln[I_2 \lambda_2^4 f_1 / I_1 \lambda_1^4 f_2]}$$
 (2)

Only two spectral bands, listed in Table 4 with the appropriate constants, were sufficiently intense and free from overlapping lines or bands to be used for the vibrational temperature determination. While the oscillator strengths for atomic transitions are well known for most species 13 , only a few have been measured for molecular transitions and the uncertainty is larger. Thus, the absolute uncertainty of the vibrational temperature determination is rather large (\pm 25%), though trends are accurate to within experimental error (\pm 10%).

Typical vibrational temperature measurements for the 225:1 nozzle at 9 and 12A are shown in Fig. 6, where it is seen that the temperature decreased from ~6500 K near the constrictor to ~2500 K at an area ratio of 126:1. While there was a slight dependence on arc current at the upstream end of the nozzle, this disappeared toward the exit plane. Radial profiles could not be measured due to the rapid decrease in spectral intensity at higher radii.

Rotational Temperature

The intensity ratio method could not be used to determine rotational temperatures because the rotational lines were too close together and overlapped, precluding the separation of individual lines within one band. To overcome this problem, the theoretical intensity distribution for a band was calculated with an assumed rotational temperature and compared to the measured spectrum. The rotational temperature was found iteratively by varying the temperature of the theoretical distribution, a method used by several other authors 4,17,18 . An examination of the observed spectrum led to selection of the N_2 $C^3\Pi_u$ $-B^3\Pi_g$ transition band at 375.5 nm. The equations governing the distributions the of P, Q and R rotational branches were taken from Herzberg 12 , and the appropriate wavelengths were obtained from Pearse and Gaydon 14 . Fifty transitions for each of the three branches were included for each vibrational band. A simple inverted parabola line shape was used for each transition to reduce computional time. This assumption had little effect where the lines were slightly separated, but introduced substantial errors near the band heads where the lines nearly completely overlap.

Figure 7 shows measured and calculated molecular bands for N_2 . It is clear that the distribution was well predicted away from the band head at 399.8 nm, but that the intensity of the band head itself was underpredicted. The agreement between the measured and calculated spectra, the latter based on a Boltzmann distribution, indicated that the rotational levels were not far from an equilibrium state. The discrepancy at the band head may have been due either to the assumed line profile or an overpopulation of the lower states. Figure 8 shows the rotational temperatures for the 225:1 nozzle at currents of 9 and 12A determined with this technique. The rotational temperature was found to be higher than the vibrational temperature, and the rotational temperature decreased more rapidly than the vibrational temperature. Note that for $T_{\rm rot}$ the increase in temperature with current level was evident throughout the nozzle.

The temporal behavior of the three plasma temperatures was measured in an attempt to

establish the importance of wall heat transfer to the plasma characteristics. Shown in Fig. 9 are the mean results from 100 spectra recorded every 0.5 to 3 seconds during the start-up phase of an arcjet test. The recording of the spectra was started before thruster ignition to study the start up phase. A curve of the form $T=T_0$ $(1-\exp(-t/\tau))$ was used to calculate the desired time constant, τ , for each plasma temperature. The time constants for the various plasma properties, listed in Table 5, were of the same order for all the plasma temperatures. The time constants varied with axial position in the nozzle, ranging from ~5 second near the constrictor, to ~8.5 seconds near the exit plane. The time constant for the nozzle temperature was approximately 130 seconds, and the thrust time constant was ~2 seconds. Given that the current and voltage rise times were <~2 seconds, this result implies that wall heat transfer played only a small role in plasma energy transfer for these operating conditions. It appears, however, that the observed equilibration times result from either thermal or mechanical relaxation of the arcjet components, since the rise times are much longer than acoustic or plasma relaxation rates.

Electron Density

The electron number density was determined from the Stark width of the hydrogen H_B line (486.1 nm)¹⁹. After the full width at half maximum (FWHM) for the line had been determined using a Lorentzian fit to the observed profile, the Doppler and Stark broadening contributions were deconvoluted using the results presented in Ref. 19. This process required knowledge of the hydrogen kinetic temperature, which was set equal to the electron temperature based on relaxation rate calculations presented in Section IV.

Typical results, shown in Fig.10, show that the density decreased from $\sim 10^{15}$ cm³ to $\sim 2\times 10^{14}$ cm⁻³ from axial locations at area ratios of 10:1 to 126:1. While it is clear that the upstream electron density increased with current level, the downstream density was not sensitive to the arc current. This may be a reflection of the nozzle current distribution.

The radial electron density distribution was measured at an area ratio of 126:1 for both the 225:1 and 400:1 area ratio nozzles. Measurements in the downstream holes for the longer nozzle were precluded by the low electron density and low spectrometer resolution. Line profiles were obtained at each radial station and the resulting profiles Abel inverted to obtain the FWHM as a function of radius. The resulting electron density profiles, shown in Fig. 11, are flat. This result, in combination with the excitation temperature measurements, shows that there was a wide central region with relatively uniform properties which was bounded by steep gradients near the wall. The increased density observed with the longer nozzle may be due to the presence of substantial current downstream of the measurement point. In other words, the arc current may extend over most of anode length anode for both the 400:1 and 225:1 nozzles; since these measurements were made at an area ratio of 126:1 with for both of the nozzles, a larger fraction of the current may extend beyond the measurement point with the longer nozzle than with the shorter one.

It is interesting that these measurements correlate well with measurements made in the thruster plume. The plume measurements shown in Fig. 12 were made using $H_{\rm B}$ Stark broadening (close to the nozzle) and Langmuir probes.

IV. Discussion

The large difference between the atomic excitation, vibrational and rotational temperatures clearly shows the plasma to be in a nonequilibrium state throughout the nozzle. The excitation temperature was more than twice as large as either molecular temperature throughout the expansion, and the rotational temperature was higher than the vibrational temperature. Three justifications were presented for equating the excitation temperature to the free electron temperature. First, T_{ex} was the same for the three species for which it was measured. Second, the excited states used in the measurement were energetically close to the continuum. Third, the states observed and the electron densities satisfied the partial local thermal equilibrium (PLTE) criterion developed by Griem¹⁵, and modified by Hey²⁰ for nonhydrogenic systems. This criterion is written as a lower electron density bound for which the excited state population is controlled by the free electron temperature:

$$n_e \ge \frac{2.55 \times 10^{17}}{G_{eff}} \left[\frac{\Delta E_{1-2}}{E_H} \right]^3 \sqrt{\frac{kT_e}{E_H}}$$
 (3)

where the effective Gaunt factors, $G_{\rm eff}$, were obtained from van Regemorter²¹. This criterion establishes a lower bound of $1x10^{15}{\rm cm}^{-3}$ for the 18,000K temperature near the constrictor and $8x10^{15}{\rm cm}^{-3}$ for the 7000K temperature near the exit plane. Thus, based on this PLTE criterion, equating the excitation temperature to the free electron temperature is acceptable near the constrictor but unjustified near the exit plane. However, Skorupski and Suckewer²², have shown that heavy particle elastic collisions can dramatically relieve the electron density requirements for PLTE. This coupled with the first two arguments above indicate that setting $T_{\rm ex} = T_{\rm e}$ was a reasonable approximation.

Several plasma relaxation rates were evaluated to establish which processes played a dominant role in the plasma expansion. These should be compared with the average particle residence time in the nozzle, which was estimated to be ~ 5 x 10⁻⁶ seconds by dividing the nozzle length by one-half the exhaust velocity. The latter was estimated from the measured specific impulses. The plasma relaxation times, summarized in Table 6, were calculated using the temperature and density measurements presented above and relations from Vincenti and Kruger²³ and Venugopalan²⁴. For example, the energy equilibration times between the observed plasma species were used to establish the atomic hydrogen kinetic temperature. This relaxation time, given by²⁴

$$\tau_{1-2} = 1.3 \times 10^{-5} \frac{(M_1 + M_2)^2}{n_2 M_1 M_2^{0.5} \langle \sigma_{1-2} (T_2) \rangle T_2}$$
 (4)

where the subscripts refer to the species in question, was used to compare the H atom - N_2 relaxation and the H atom - electron relaxation. In these estimates the N_2 kinetic temperature was assumed to be equal to the measured rotational temperature and the N_2 density was set to $\sim 10^{16}$ cm⁻³ based on the propellant flow rate, nozzle area and exhaust velocity. The momentum transfer cross-section for H - N_2 collisions is not known, so it was estimated from the geometric particle radii²⁵. The cross - section for electron - hydrogen collisions was obtained from Ref 26. The results, a relaxation time of 10^{-6} seconds for electron - hydrogen collisions, and 2 x 10^{-7}

seconds for N_2 - H were inconclusive. However, when the relaxation times were calculated for electron - proton and proton - atomic hydrogen collisions the results were more significant. The large Coulomb cross - section for the electron - proton interaction gives a relaxation time relation of 24

$$\tau_{eq}^{e-p} = 2.5 \times 10^2 \frac{T_e^{3/2}}{n_p \ln{(\Lambda)}}$$
 (5)

which, for the measured conditions, yields a relaxation time of 10^{-8} seconds. Equation 4 yields a time of 10^{-9} seconds for the H - P relaxation. Thus, the two-step process, consisting of an electron - proton collision followed by a proton - neutral hydrogen collision, is two orders of magnitude faster than the single step nitrogen molecule - hydrogen atom collision process. This led to the use of the electron excitation temperature as an estimate for the kinetic hydrogen temperature when the Doppler broadening was calculated in Section III. Similar two-step processes were investigated to check additional processes (electron - NII, NII - H and N_2 - N, N - H), but they were found to be slower than the electron - H⁺, H⁺ - H process described above. This is a result of the large mass difference between hydrogen atoms and nitrogen molecules (1:28) that slows this process and the fast relaxation between electrons and protons due to the large Coulomb cross-section.

While the measured rotational and vibrational temperatures were the same near the exit plane, the rotational temperature appeared to be slightly higher close to constrictor (Fig. 8). The observed difference in the behavior of the two temperatures during the expansion may result from the different relaxation times for the two processes. The rotational relaxation is somewhat slower than the translational relaxation²³. With equation 4 the translational relaxation was calculated to be 10⁻⁷s, so the rotational relaxation should have a time constant of 10⁻⁶ to 10⁻⁷s. The vibrational relaxation was calculated from²³

$$\tau = C \frac{e^{(K_2/T)^{1/3}}}{p}$$
 (6)

where the temperature T is in Kelvin and the pressure p in atm. The constants C and K_2 depend on the gas, for nitrogen $K_2 = 1.91 \times 10^6$ K and C = 7.12×10^9 s atm. The shortest possible time constant results from using the propellant feed pressure in eqn. 6, which results in a time constant of 10^{-6} seconds. Of course, during the expansion this value will increase by over three orders-of magnitude, so that the vibrational relaxation times are the longest time constants for the nozzle plasma. The difference in the time constants for rotation and vibration would explain the difference in the slope for these two temperatures.

V. Conclusion

The spectroscopic analysis of the light emitted from the plasma inside the nozzle of a 1 kW class arcjet thruster operated with a 2:1 hydrogen: nitrogen mixture showed the expected atomic and molecular nitrogen and hydrogen spectral lines. In addition, considerable NH and small traces of NH₂, and tungsten, eroded from the electrodes, were found. Measurements of

excitation, vibrational, and rotational temperatures showed the plasma to be in a nonequilibrium state throughout the nozzle. It appears that the upper atomic excited states can be described using a PLTE model, though this should be investigated carefully for the plasma near the exit plane. A collisional - radiative model, allowing for a minimum of two fluid temperatures, will be needed to accurately model the arcjet plasma expansion. The atomic excitation temperature was found to drop from 18,000 K to 9,000 K during the expansion, and the molecular temperatures were found to be lower than the atomic, ranging from 6,500 K to 3,000 K for the vibrational and 8,000 K to 3,000 K for the rotational temperature, respectively. The electron density dropped from ~1x10¹⁵ cm near the constrictor to 2x10¹⁴ cm⁻³ near the exit, and was sensitive to the arc current only near the constrictor. Radial profiles of the excitation temperature and electron density showed that both were flat over the observed 87% of the nozzle radius. This showed the presence of a broad central region with low gradients, surrounded by high gradients at the walls. The relaxation times calculated for the plasma conditions qualitatively explain the different temperature trends, and show that the free electron temperature should be close to the atomic hydrogen temperature. The latter could only be conclusively demonstrated using a two collision process, involving a hydrogen ion - electron Coulomb collision followed by a hydrogen ion - hydrogen neutral collision.

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Current Settings, A					
Parameter	9 A	10 A	11 A		
specific power (kJ/kg)					
sealed holes	21 510	23 360	25 160		
open holes	21 310	23 260	25 080		
thrust (mN)					
sealed holes	206	213	220		
open holes	204	212	219		
specific impulse (s)					
sealed holes	422	436	451		
open holes	419	435	449		
thrust efficiency (%)					
sealed holes	39.8	39.3	38.9		
open holes	39.2	39.1	38.7		

Table 1: Thruster performance with 225:1 expansion ratio nozzle with open and sealed holes

Identified Species	Identified Wavelengths, nm
H	656.3, 486.1, 434.1
H	588.8, 587.9, 580.6, 577.5, 573.7, 572.9, 571.3, 565.6, 550.5
NĪ	
NII	571.1,568.6, 566.7
N ₂	585.4, 581.5, 575,5 559.9 larger bands at: 405.9, 399.8, 394.0
4	385.8, 380.5, 375.6, 389.4, 388.9
N, ⁺	586.5, 585.3, 566.4, bands at: 391.4, 388.4, 385.8
	bands at 336.0, 337.1, 450.1
	587.0,570.7
	584.5, 579.6, 579.3,569.8, 566.1, 550.1, 412.5

Table 2: Species Identified within the Arcjet Nozzle

Species	λ _{ik} (nm)	f_{ik}	\mathbf{g}_{i}	
NI	583.0	0.00328	6	
	583.4	0.00390	2	
	585.0	0.00152	2	
NII	566.7	0.339	3	
	567.6	0.450	1	
	568.0	0.380	5	
Н	656.3	0.6407	2	
	486.1	0.1193	2	

Table 3: Transitions used to measure atomic excitation temperature.

Band Head λ (nm)	G (eV)	f	
380.5	7.24	0.76	
375.5	11.52	0.52	

Table 4: Bands used to measure N_2 vibrational temperature. Both are $C^3\Pi_u$ -B $^3\Pi_g$ transitions

Axial position from constrictor (expansion ratio)	atomic excitation	molecular vibration	molecular rotation
	temperature	temperature	temperature
3 mm (13.9:1)	5.0 - 5.4s	5.1 - 5.6s	4.9 - 5.6s
6 mm (56.1:1)	6.9 - 7.6s	7.0 - 7.8s	7.1 - 7.7s
9 mm (126:1)	8.1 - 8.5s	8.2 - 9.7s	8.3 - 8.6s

Table 5: Approximate time constants for plasma temperature to reach 63% of the final values after thruster ignition

time (s)	process
10 ⁻¹¹ -10 ⁻¹⁰	electron self relaxation, Ref. 18
	hydrogen-hydrogen (or proton) kinetic relaxation, Ref. 18
	electron - proton kinetic relaxation, Ref. 18
	direct electron - hydrogen relaxation, Ref. 18
	N ₂ rotational relaxation, Ref. 17
	hydrogen - N ₂ kinetic relaxation, Ref. 18
>10	average particle residence time in the nozzle
	N ₂ vibrational relaxation, Ref. 17

Table 6: Relaxation times in the arcjet nozzle plasma

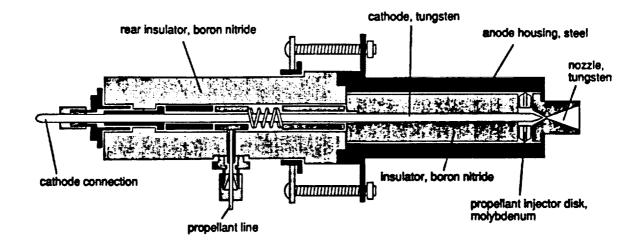


Figure 1. Schematic of arcjet used for spectroscopic measurements.

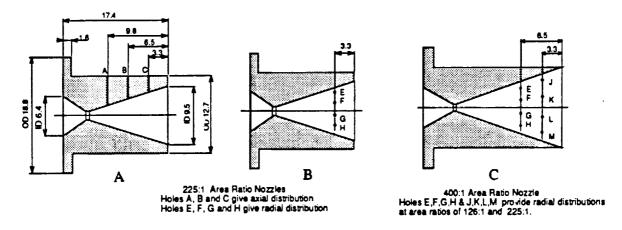


Figure 2. Schematics of three nozzles used in this study. Holes drilled in nozzle wall are all 0.25mm in diameter. Not to scale. All dimensions in millimeters.

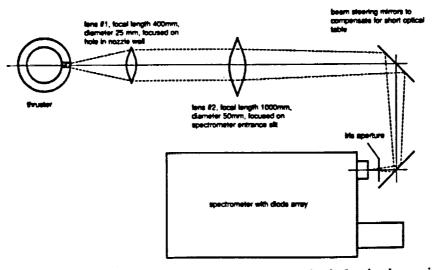


Figure 3. Optical arrangement used to image the holes in the arcjet nozzle walls onto the spectrometer entrance slit.

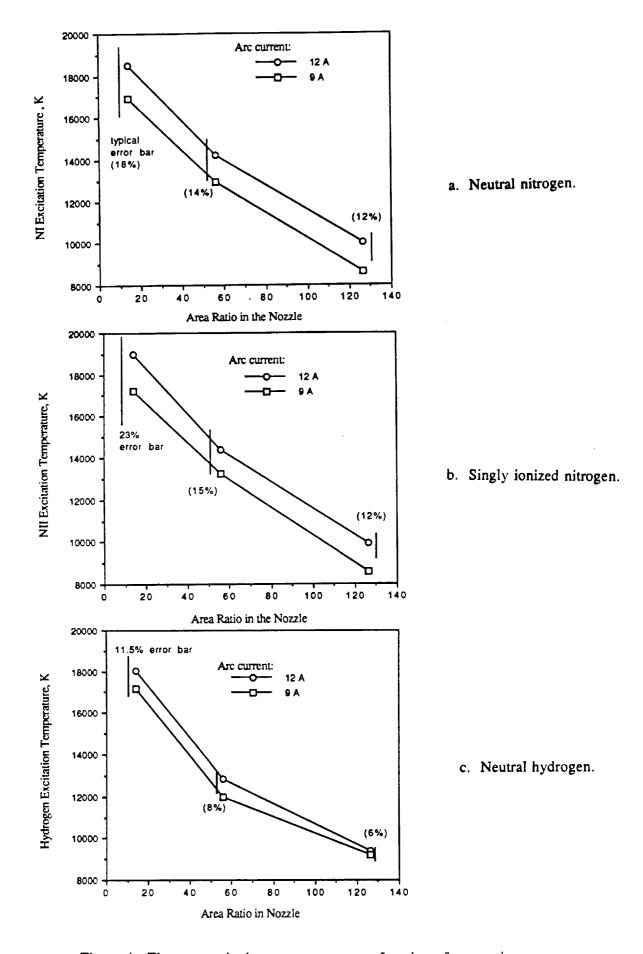
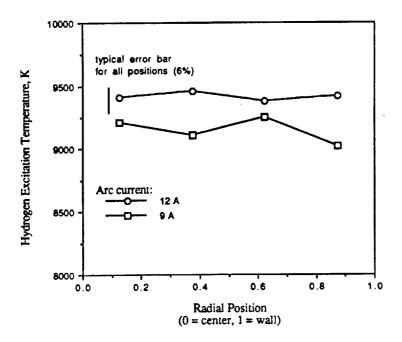
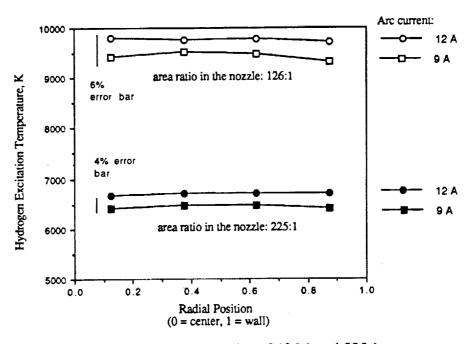


Figure 4. Electron excitation temperature as a function of area ratio in the nozzle for three plasma species.



a. 225:1 nozzle at area ratio of 126:1.



b. 400:1 nozzle at area ratios of 126:1 and 225:1.

Figure 5. Neutral hydrogen electron excitation temperature as a function of radial position in both the 225:1 and 400:1 nozzles.

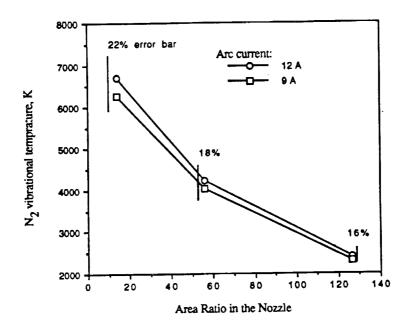


Figure 6. N_2 vibrational temperature as a function of area ratio in the 225:1 nozzle.

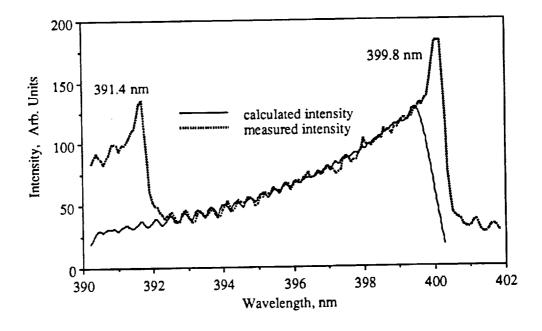


Figure 7. Typical experimental and calculated N_2 rotational band intensity distributions.

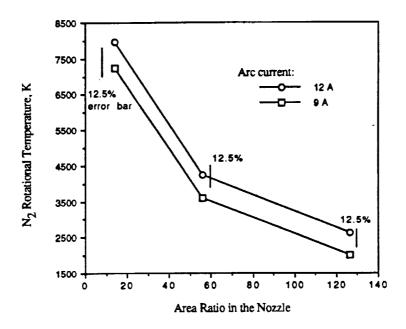


Figure 8. N_2 rotational temperature as a function of area ratio in the 225:1 nozzle.

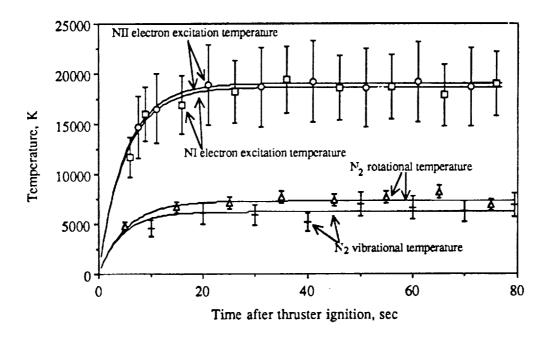


Figure 9. Time evolution of the plasma temperatures during arcjet start-up.

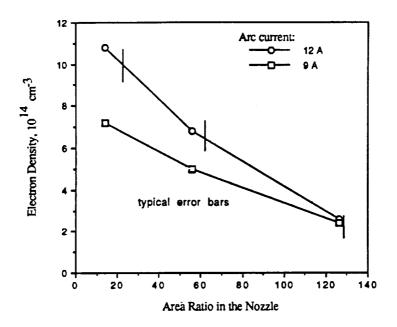


Figure 10. Electron density as a function of area ratio in the 225:1 nozzle.

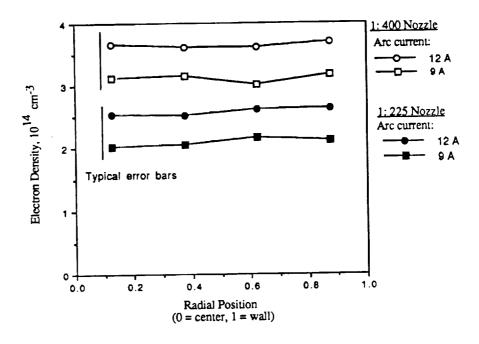


Figure 11. Electron density as a function of radial position for both the 225:1 and 400:1 nozzles at an area ratio of 126:1.

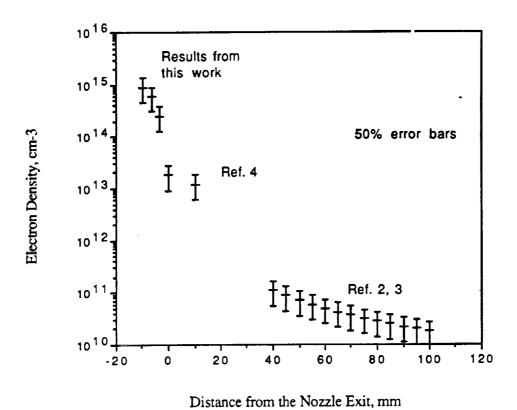


Figure 12. Comparison of electron density measurements in the arcjet plume with results from this work.

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the 27th Joint Propulsion Conference 27, 1991. Dieter M. Zube, Institute f Myers, Sverdrup Technology, Inc. R 16. Abstract Emission spectroscopy measurement The thruster propellant was a hydrog diameter holes were drilled into the dinternal flow. Atomic electron excita plasma using relative line intensity to 3 mm downstream of the constrictor and rotational temperatures decreased locations. The electron density, meas 10 ¹⁴ cm ⁻³ during the expansion. The most relaxation times equal or exceed	ir Raumfahrtsysteme, Usesponsible person, Rogers were made of the plassen-nitrogen mixture use liverging section of the stion, vibrational, and rotechniques. The atomic extensions of the section of the	na flow inside the nod to simulate fully do ungsten thruster noz ational temperatures citation temperatures and from 8000 to 30 g line Stark broadeniasma is highly nonec	ozzle of a 1 kW classecomposed hydrazin izle to provide optics were determined for edecreased from 18 rictor, while the mol 00 K, respectively, by decreased from	s arcjet thruster. se. 0.25 mm al access to the or the expanding 000 K at a location lecular vibrational between the same ~10 ¹⁵ cm ⁻³ to ~2 x
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