

NUMERICAL MODELING OF TWO-PHASE FLOW WITH CONTACT DISCONTINUITIES

RONALD A. REMMERSWAAL* AND ARTHUR E.P. VELDMAN†

Johann Bernoulli Institute, University of Groningen
Nijenborgh 9, 9747 AG Groningen, The Netherlands
* r.a.remmerswaal@rug.nl † a.e.p.veldman@rug.nl

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Abstract. When simulating high Reynolds number two-phase flow, boundary layers develop at the interface, which are much thinner compared to the capillary length-scales that are of interest. Resolving such an interface layer is expensive and therefore it is often not resolved in a simulation.

Numerically such an underresolved interface layer results in a velocity discontinuity tangential to the interface. We propose to include such tangential velocity discontinuities in our numerical model. This results in a sharp two-fluid model for two-phase flow, where only the interface-normal component of the velocity field is smooth. This condition is implicitly enforced via a new jump condition on the pressure gradient, which we discretize using a multidimensional variant of the ghost fluid method [6].

Results are shown of breaking waves [2] as well as (breaking) waves impacting a solid wall [3] where we compare to state-of-the-art methods [3, 4]. We show that our proposed method is able to accurately simulate high Reynolds number two-phase flow without the need for resolving, or artificially thickening, of the interface layer.

1 INTRODUCTION

Numerical simulations can facilitate in the understanding of the physics underlying the observed variability in impact pressures during breaking wave impacts [3]. One of the mechanisms deemed responsible for pressure variability is the development of free surface instabilities at the wave crest. Numerical investigation of free surface instabilities on a breaking wave crest requires resolving the capillary length scales, amongst which the Kelvin-Helmholtz length scale¹ will be most restrictive. Due to viscous effects, an interface layer will form at the fluid interface, whose thickness² λ^{BL} is estimated to be significantly smaller than the capillary length scale that we are interested in resolving

$$\frac{\lambda^{\text{BL}}}{\lambda^{\text{KH}}} \approx \frac{\sqrt{|\llbracket u_\tau \rrbracket|}}{40},$$

where $\llbracket u_\tau \rrbracket$ is the tangential velocity difference across the interface. To that end we propose to model the interface layer by letting the velocity field be discontinuous in the interface tangential direction.

In section 2 we describe the continuous model which will be discretized in section 3. Results are shown

¹Here λ^{KH} denotes the fastest growing wavelength according to linearized potential flow theory.

²This length scale is estimated using the Blasius boundary layer approximation for a single phase flow.

in section 4 followed by concluding remarks in section 5.

2 MATHEMATICAL MODEL

We consider a d -dimensional domain $\Omega \subset \mathbb{R}^d$, which is separated by a time-dependent interface $I(t)$ resulting in a liquid domain $\Omega^l(t)$ and a gaseous domain $\Omega^g(t)$. Since the Navier-Stokes equations are equal for the liquid and gas phase, we state them for the π -phase, where we consider $\pi \in \{l, g\}$.

At the interface we define the jump of some function α in the following way

$$[[\alpha]](\mathbf{x}) = \alpha^g(\mathbf{x}) - \alpha^l(\mathbf{x}) = \lim_{s \rightarrow 0^+} \alpha(\mathbf{x} + s\boldsymbol{\eta}) - \lim_{s \rightarrow 0^-} \alpha(\mathbf{x} + s\boldsymbol{\eta}), \quad \mathbf{x} \in I,$$

where $\boldsymbol{\eta}$ denotes the interface normal pointing into the gas phase. Similarly we define the sum as

$$\{\{\alpha\}\} = \alpha^g + \alpha^l.$$

2.1 Two-fluid model

We model the two-phase flow using a two-fluid model, where each of the phases is modeled by the incompressible Navier-Stokes equations, which we write here in conservative form on some arbitrary and fixed control volume $\omega \subset \Omega$

$$\frac{d}{dt} \int_{\omega^\pi} \rho^\pi dV + \int_{\partial\omega^\pi \setminus I} \rho^\pi u_\eta^\pi dS = 0 \quad (1)$$

$$\frac{d}{dt} \int_{\omega^\pi} \rho^\pi \mathbf{u}^\pi dV + \int_{\partial\omega^\pi \setminus I} \rho^\pi \mathbf{u}^\pi u_\eta^\pi dS = \int_{\partial\omega^\pi} [(-p^\pi + \rho^\pi \mathbf{g} \cdot \mathbf{x})\mathbf{I} + \boldsymbol{\mu}^\pi \mathbf{S}^\pi] \cdot \boldsymbol{\eta} dS,$$

where $\omega^\pi(t) = \omega \cap \Omega^\pi(t)$, and $\rho, p, \mathbf{g}, \mathbf{u}, \mathbf{S}$ denote the density, pressure, gravitational acceleration, velocity and symmetric part of the velocity gradient respectively. The Navier-Stokes equations are supplemented with appropriate boundary conditions (usually slip) and an evolution equation for the interface

$$\frac{d}{dt} \mathbf{x} = \mathbf{u}^\pi(\mathbf{x}, t), \quad \forall \mathbf{x} \in I(t), \quad (2)$$

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where the interface motion may thus be modeled by either of the two velocities. The following velocity jump condition then uniquely determines the interface evolution

$$[[u_\eta]] = 0, \quad (3)$$

where $u_\eta = \mathbf{u} \cdot \boldsymbol{\eta}$. Furthermore we impose the Young-Laplace equation on the interface to model the effect of surface tension (we omit the jump in the diffusive stresses)

$$[[p]] = -\sigma\kappa, \quad (4)$$

where the interface mean curvature is denoted by κ and σ denotes the surface energy coefficient.

We choose to combine the two mass conservation equations (1) in the following way (making use of $|\omega| = |\omega^l| + |\omega^g|$)

$$\begin{aligned} \frac{d}{dt} \int_{\omega^l} dV + \int_{\partial\omega^l \setminus I} \mathbf{u}^l \cdot \boldsymbol{\eta} dS &= 0 \\ \int_{\partial\omega} \mathbf{u} \cdot \boldsymbol{\eta} dS &= 0, \end{aligned} \quad (5)$$

where the first equation is a volumetric constraint on the evolution equation of the interface, and the second equation states that the velocity field is divergence free in integral sense (note that this is not per phase).

2.2 Velocity discontinuities

Note that if $\mu^\pi = 0$ then the velocity field can develop tangential velocity discontinuities $[[u_\tau]] \neq 0$ due to surface tension and/or gravity, where $\tau \perp \eta$ is a vector tangent to the interface. If $\mu^\pi > 0$ then the Navier-Stokes solution will be smooth $[[u_\tau]] = 0$, numerically we however do not enforce this, as motivated in the introduction.

3 NUMERICAL MODEL

The numerical model is implemented in our in-house free surface Navier-Stokes solver ComFLOW which makes use of local and adaptive mesh refinement [11]. For simplicity in presentation however, we will assume the mesh to be 2D and rectilinear. The variables are arranged according to an Arakawa C-grid.

3.1 Notation

The notation used is similar to that of [5]. We denote the set of all cells by \mathcal{C} and the set of all faces by \mathcal{F} , where the faces of one particular cell c are denoted by $\mathcal{F}(c) \subset \mathcal{F}$ and the cells neighbouring one particular face f are denoted by $\mathcal{C}(f) \subset \mathcal{C}$.

Functions $p : c \mapsto p_c \approx p(\mathbf{x}_c)$ live in the function space denoted by \mathcal{C}^h , where \mathbf{x}_c is the centroid of the cell c . Similar definitions hold for \mathcal{F}^h , with $u : f \mapsto u_f \approx \mathbf{n}_f \cdot \mathbf{u}(\mathbf{x}_f)$, where \mathbf{n}_f is the face normal. The function $\alpha : \mathcal{C} \times \mathcal{F} \rightarrow \{-1, +1\}$ encodes the orientation of the face normal such that $\alpha_{c,f} \mathbf{n}_f$ is outward pointing relative to the cell c . The staggered control volume belonging to some face f is denoted by ω_f .

3.2 Interface model

The interface is tracked using a volume of fluid (VoF) method combined with a geometric PLIC reconstruction.

3.2.1 Representation

The interface is represented using the volume fraction field $\tilde{\chi}^l \in \mathcal{C}^h$, which is defined as $\tilde{\chi}_c^\pi = \frac{|c^\pi|}{|c|}$. During the interface advection step the volume fluxes are computed using a geometric PLIC reconstruction of the interface: per cell c we find a normal vector η_c and plane constant s_c such that

$$c^l = \{ \mathbf{x} \in c \mid \eta_c \cdot (\mathbf{x} - \mathbf{x}_c) \leq s_c \}, \quad c^g = c \setminus c^l.$$

Computation of the normal vector is done using either local height functions (LHFs) or the efficient least-squares VoF interface reconstruction algorithm (ELVIRA) [8] if insufficiently many LHFs are available. Using the interface reconstructions we then define the staggered volume fractions $\tilde{\chi}^\pi \in \mathcal{F}^h$ and the phase domains

$$\tilde{\chi}_f^\pi = \frac{|\omega_f^\pi|}{|\omega_f|}, \quad \omega_f^\pi = \omega \cap \Omega^\pi, \quad \Omega^\pi = \bigcup_{c \in \mathcal{C}} c^\pi. \quad (6)$$

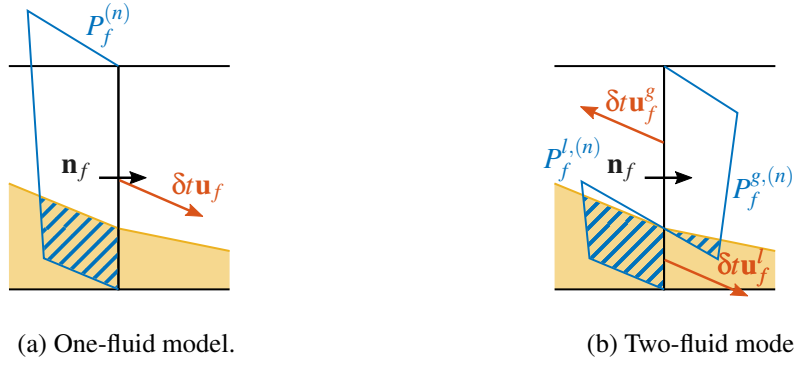


Figure 1: Illustration of the EMFPA flux polygons $P_f^{(n)}$ (blue). The blue hatched regions correspond to the liquid volume fluxes and the yellow regions correspond to the liquid domain $\Omega^{l,(n)}$.

The face apertures are given by $a_f^\pi = \frac{|f^\pi|}{|f|}$, where f^π is the part of the face f which contains the π -phase. Note that f^π is not uniquely defined since the PLIC reconstruction is not continuous across the faces, therefore we average the face aperture using the two neighbouring cells $c \in \mathcal{C}(f)$.

For the modeling of surface tension we compute the interface curvature $\kappa \in \mathcal{C}^h$ using the LHF method or generalized height function (GHF) method [9] if insufficiently many LHF's are available.

3.2.2 Advection

In case of a single fluid we summarize the interface advection in the following way

$$\bar{\chi}^{l,(n+1)} = \bar{\chi}^{l,(n)} - \delta t D m^{l,(n)}, \quad \delta t |f| m^{l,(n)} \approx \int_{t^{(n)}}^{t^{(n+1)}} \int_{f^l} u_\eta dS d\tau \quad (7)$$

where $D : \mathcal{F}^h \rightarrow \mathcal{C}^h$ is the discrete divergence operator (see section 3.3.5) and $m^{l,(n)}$ denotes the liquid volume flux which is approximated using the edge-matched flux polygon advection (EMFPA) method [7].

Here the Lagrangian backtracking of the face f , resulting in the flux polygon $P_f^{(n)} = P^{\text{EMFPA}}(f, -\delta t u^{(n)})$, is according to eq. (2), whereas the enforcement step tries to ensure that (here we consider the volume of the flux polygon to be signed)

$$|P_f^{(n)}| = \delta t |f| u^{(n)}.$$

The volume flux is then given by

$$\delta t |f| m_f^{l,(n)} = |P_f^{(n)} \cap \Omega^{l,(n)}|,$$

as shown in fig. 1a.

In case of a velocity discontinuity we modify the original EMFPA method by redefining the flux polygon as

$$P_f^{(n)} = P_f^{l,(n)} \cup P_f^{g,(n)}, \quad P_f^{\pi,(n)} = P^{\text{EMFPA}}(f^{\pi,(n)}, -\delta t u^{\pi,(n)}),$$

and the volume enforcement step is replaced by

$$|P_f^{\pi,(n)}| = \delta t |f| a_f^{\pi,(n)} u^{\pi,(n)}.$$

See also fig. 1b. It can be shown that this is a consistent approximation to the interface advection problem in the presence of a velocity discontinuity.

3.3 Momentum equations

3.3.1 Convection operator

The convection operator is very similar to that of [4]. This means that we use volume fluxes from the VoF method to transport momentum, per phase, along with the interface. Our approach differs from that of [4] in a few ways:

- We use the dimensionally unsplit EMFPA method as the underlying advection method. This is mainly because dimensional splitting does not combine well with the fact that our divergence operator depends on the interface configuration (which is updated at each dimensionally split step) via the face apertures (see section 3.3.5).
- In [4] the liquid and gas velocities are ‘merged’ into a single mass weighted velocity field \bar{u} after each momentum transport step

$$\bar{u} = \frac{\{\{\rho\bar{\chi}u\}\}}{\{\{\rho\bar{\chi}\}\}}.$$

We skip this step since we allow the velocity field to be discontinuous.

The resulting advection of some function $v^\pi \in \mathcal{F}^h$ using the velocity field u^π is denoted by

$$|\omega^{\pi,*}|_{v^{\pi,(n+1)}} = |\omega^{\pi,(n)}|_{v^{\pi,(n)}} - \delta t |\omega^{\pi,(n)}| \tilde{A}^\pi(u^{\pi,(n)})_{v^{\pi,(n)}},$$

note that we write $|\omega^{\pi,*}|$ and not $|\omega^{\pi,(n+1)}|$, this is because the latter follows from the ‘primary’ volume fractions $\bar{\chi}^{(n+1)} \in \mathcal{C}^h$ according to eq. (6), whereas the former follows from the staggered advection method. Due to this discrepancy the advection method is not exactly conservative.

3.3.2 Diffusion operator

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We use a second order accurate diffusion operator applied to \bar{u} with a dynamic viscosity which is based on weighted harmonic averaging

$$\bar{\mu} = \{\{\bar{\chi}\mu^{-1}\}\}^{-1}.$$

The reason we use the mass weighted velocity field \bar{u} for diffusion is that we want our solution with a velocity discontinuity to converge to the ordinary Navier-Stokes solution (hence without velocity discontinuity) as soon as the interface layer is resolved ($\delta x \rightarrow 0$ or $\text{Re} \rightarrow 0$).

3.3.3 Gravity

We model gravity according to the following finite volume discretisation

$$F = \{\{\rho a\}\} G(\mathbf{g} \cdot \mathbf{x}_i),$$

where \mathbf{x}_i denotes the interface centroid, and reduces to the cell centroid \mathbf{x}_c if the cell does not contain an interface. This model can be shown to sharply mimic the global conservation of energy and is well-balanced with the gradient operator described in section 3.3.6. That is, if the interface is globally linear,

then $\exists p \in \mathcal{C}^h$ for which

$$F = \{\{\rho\bar{\chi}\}\} g^\pi(p).$$

3.3.4 Time integration

The time integration is split in the following way. First we advect the momentum together with the interface eq. (7)

$$|\omega^{\pi,*}| \rho^\pi u^{\pi,*} = |\omega^{\pi,(n)}| \rho^\pi u^{\pi,(n)} - \delta t |\omega^{\pi,(n)}| \tilde{A}^\pi(u^{\pi,(n)}) \left(\rho^\pi u^{\pi,(n)} \right).$$

Then the diffusion operator and gravity force are integrated in time³

$$u^{\pi,**} = u^{\pi,*} + \frac{\delta t}{\{\{\rho\bar{\chi}\}\}} \left[\tilde{D}(2\tilde{\mu}\tilde{S}\tilde{u}^*) + F^{(n+1)} \right],$$

where \tilde{D}, \tilde{S} denote the divergence operator and symmetric part of the gradient operator respectively.

Finally the divergence constraint eq. (5) must be imposed, which is done via a pressure correction step

$$u^{\pi,(n+1)} = u^{\pi,**} - \delta t g^\pi(p), \quad g^\pi(p) \approx \frac{1}{\rho^\pi} \mathbf{n}_f \cdot (\nabla p^\pi)(\mathbf{x}_f),$$

for some pressure gradient $g^\pi(p)$. We can now also impose the jump condition eq. (3), since this condition can be written in terms of a jump condition on the pressure gradient

$$\eta_f \cdot (\mathbf{J}u^{(n+1)})_f = 0 \quad \Rightarrow \quad \eta_f \cdot (\mathbf{J}g^\pi(p))_f = \delta t^{-1} \eta_f \cdot (\mathbf{J}u^{**})_f, \quad (8)$$

where $(\mathbf{J}u)_f \approx \llbracket \mathbf{u} \rrbracket(\mathbf{x}_f)$ denotes the jump interpolant (see section 3.3.7) and η_f is an approximate interface normal located at \mathbf{x}_f . Hence the pressure gradient $g^\pi(p)$ should incorporate two jump conditions at the interface: the Young-Laplace equation eq. (4) as well as the newly introduced pressure gradient jump condition eq. (8).

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3.3.5 Divergence operator

The one-fluid divergence operator D is given by

$$|c|(Du)_c = \sum_{f \in \mathcal{F}(c)} |f| \alpha_{c,f} u_f \approx \int_{\partial c} \mathbf{u} \cdot \mathbf{n} \, dS.$$

At the interface two velocity fields exist, and we therefore modify the divergence operator using the cut-cell method (CCM) [10].

$$|c|(D\{\{au\}\})_c \approx \int_{\partial c^l \setminus I} \mathbf{u}^l \cdot \mathbf{n} \, dS + \int_{\partial c^g \setminus I} \mathbf{u}^g \cdot \mathbf{n} \, dS = \int_{\partial c} \mathbf{u} \cdot \mathbf{n} \, dS,$$

where $|f|\{\{au\}\}_f = |f^l|u_f^l + |f^g|u_f^g$ is the total flux through the face f , see also fig. 2.

³For low-Reynolds number flow we use the implicit midpoint rule for the time integration of diffusion.

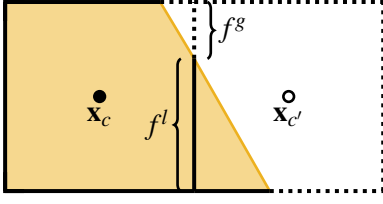


Figure 2: Illustration of the CCM. Each face is split into its liquid and gaseous part $f = f^l \cup f^g$. The yellow region corresponds to the liquid domain.

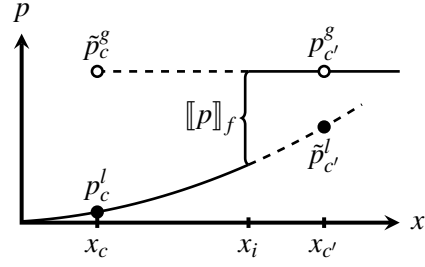


Figure 3: Illustration of the GFM. The pressure values denoted by \tilde{p} are ‘ghost’ pressures and not actually part of the solution. Here $x_i = x_c + r_f^l(x_{c'} - x_c)$.

3.3.6 Gradient operator

In the interior of each of the fluid domains the gradient operator G is simply the negative adjoint of the divergence operator

$$|\omega_f|(Gp)_f = -|\omega_f|(D^T p)_f = -|f| \sum_{c \in \mathcal{F}(f)} \alpha_{c,f} p_c \approx \mathbf{n}_f \cdot (\nabla p)(\mathbf{x}_f).$$

At the interface we need to modify the gradient to take the two jump conditions into account. Here we use the ghost fluid method (GFM) [6], which results in

$$g^\pi(p) = \frac{1}{\{\{\rho r\}\}} [Gp + \llbracket p \rrbracket G\chi + \llbracket g(p) \rrbracket \hat{\rho}^T], \quad (9)$$

where $\chi_c \in \{0, 1\}$ is the liquid indicator and $r^\pi \in [0, 1]$ is the fraction of the line connecting \mathbf{x}_c to $\mathbf{x}_{c'}$ which is covered in the π -phase, see also fig. 3. Moreover we define

$$\hat{\rho}^\pi = \begin{cases} -\rho^g r^g & \pi = l \\ \rho^l r^l & \pi = g \end{cases},$$

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such that $\llbracket \hat{\rho} \rrbracket = \{\{\rho r\}\}$ and $\{\{\rho r \hat{\rho}\}\} = 0$. We let $r = \bar{\chi}$ such that momentum is conserved if $\llbracket p \rrbracket = 0$

$$\sum_{f \in \mathcal{F}} |\omega_f| \mathbf{n}_f \{\{\rho \bar{\chi} g(p)\}\}_f = \sum_{f \in \mathcal{F}} |\omega_f| \mathbf{n}_f \left[(Gp)_f + \llbracket g(p) \rrbracket_f \frac{\{\{\rho \bar{\chi} \hat{\rho}\}\}_f}{\{\{\rho \bar{\chi}\}\}_f} \right] = \mathbf{0}.$$

3.3.7 The gradient jump

The GFM eq. (9) requires the face normal component of the gradient jump, $\llbracket g(p) \rrbracket_f$, to be known. However, the jump condition eq. (8) specifies the interface normal component of the gradient jump, rather than the face normal component of the gradient jump, which are in general unequal (unless the interface and face normals coincide: $\boldsymbol{\eta}_f = \pm \mathbf{n}_f$). Hence $\llbracket g(p) \rrbracket_f$ is not known a priori, and therefore we include it as an unknown in the pressure Poisson problem, and define it via the imposition of eq. (8).

For imposing eq. (8) we need the jump interpolant \mathbf{J} . We let the jump interpolant (in 2D on a rectilinear mesh) be given by

$$(\mathbf{J}u)_f = \mathbf{n}_f \llbracket u \rrbracket_f + \frac{1}{|\mathcal{F}_f^\perp(f)|} \sum_{\hat{f} \in \mathcal{F}_f^\perp(f)} \mathbf{n}_{\hat{f}} \llbracket u \rrbracket_{\hat{f}},$$

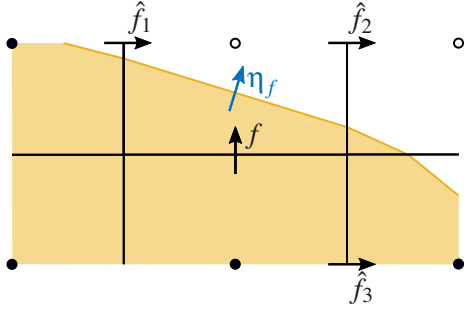


Figure 4: Example jump interpolant for a face f with $\left| \frac{\eta_f \cdot \mathbf{n}_{\hat{f}}}{\eta_f \cdot \mathbf{n}_f} \right| \leq 1$. Here $\mathcal{F}_I^\perp(f) = \{\hat{f}_1, \hat{f}_2, \hat{f}_3\}$.

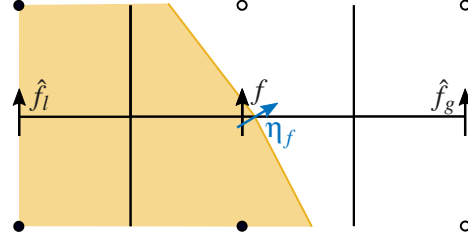


Figure 5: Example jump interpolant for a face f with $\left| \frac{\eta_f \cdot \mathbf{n}_{\hat{f}}}{\eta_f \cdot \mathbf{n}_f} \right| > 1$.

where $\mathcal{F}_I^\perp(f) \subset \mathcal{F}$ contains the faces \hat{f} neighbouring f whose normal $\mathbf{n}_{\hat{f}}$ is perpendicular to \mathbf{n}_f and whose corresponding staggered control volume $\omega_{\hat{f}}$ contains both phases, so $\tilde{\chi}_{\hat{f}}^\pi > 0$ for $\pi = \{l, g\}$. Hence the jump in the direction perpendicular to \mathbf{n}_f is defined as the average of the available jumps neighbouring the face f , see also fig. 4. We refer to the usage of GFM with our proposed multi dimensional jump interpolant \mathbf{J} as the multi dimensional GFM (MDGFM).

Once eq. (8) is imposed using the previously introduced jump interpolant, we can directly express $\llbracket g(p) \rrbracket_f$ in terms of the neighbouring jumps

$$\llbracket g(p) \rrbracket_f + \frac{1}{|\mathcal{F}_I^\perp(f)|} \sum_{\hat{f} \in \mathcal{F}_I^\perp(f)} \frac{\eta_f \cdot \mathbf{n}_{\hat{f}}}{\eta_f \cdot \mathbf{n}_f} \llbracket g(p) \rrbracket_{\hat{f}} = \delta t^{-1} \frac{\eta_f \cdot (\mathbf{J}u^{**})_f}{\eta_f \cdot \mathbf{n}_f}.$$

The ratio $\frac{\eta_f \cdot \mathbf{n}_{\hat{f}}}{\eta_f \cdot \mathbf{n}_f}$ is problematic as the denominator tends to zero when the interface normal becomes perpendicular to the face normal. If the ratio exceeds one in absolute sense then we do not impose eq. (8) but instead interpolate the jump $\llbracket g(p) \rrbracket$ from neighbouring pressure gradients

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$$\llbracket g(p) \rrbracket_f = g(p)_{\hat{f}_g} - g(p)_{\hat{f}_l},$$

where \hat{f}_π are neighbouring faces whose normal equals that of f , as shown in fig. 5.

The extended pressure Poisson problem can now be written as: find $p \in C^h$, $\llbracket g(p) \rrbracket \in \mathcal{F}_I^h$ such that

$$\begin{cases} D \llbracket ag(p) \rrbracket_c = \delta t^{-1} D \llbracket au^{**} \rrbracket_c & c \in \mathcal{C} \\ \eta_f \cdot (\mathbf{J}g(p))_f = \delta t^{-1} \eta_f \cdot (\mathbf{J}u^{**})_f & f \in \mathcal{F}_I \text{ for which } |\eta_f \cdot \mathbf{n}_f| \geq \sqrt{\frac{1}{2}}, \\ \llbracket g(p) \rrbracket_f = g(p)_{\hat{f}_g} - g(p)_{\hat{f}_l} & f \in \mathcal{F}_I \text{ for which } |\eta_f \cdot \mathbf{n}_f| < \sqrt{\frac{1}{2}} \end{cases},$$

where \mathcal{F}_I is the set of faces for which the corresponding staggered control volume contains both phases. Due to the use of the CCM for the divergence operator, it is guaranteed that the right-hand side of the pressure equation is in the image of the Laplace operator, hence a unique (up to a constant) pressure exists.

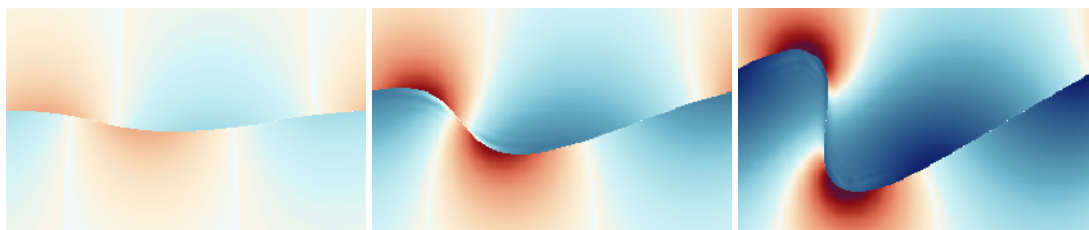


Figure 6: The velocity magnitude (clipped to $|\mathbf{u}|_2 \in [0, 1]$) at $t/T^\sigma = 0.3, 0.4, 0.5$ for $h/\lambda = 1/256$.

4 RESULTS

4.1 Inviscid Kelvin-Helmholtz instability

In order to isolate the treatment of the velocity discontinuity we consider the simulation of an inviscid Kelvin-Helmholtz instability. We let the initial interface profile be given by

$$\eta(x, 0) = \hat{\eta}(0) \cos(kx), \quad k = \frac{2\pi}{\lambda} = 1, \quad \hat{\eta}(0) = 10^{-2},$$

where the domain, which is periodic in the x -direction, is given by $\Omega = [0, \lambda] \times [-1.5\lambda, 1.5\lambda]$ with slip boundary conditions applied to the top and bottom walls. The initial velocity field is given by $u^s = \delta U/2$, $u^l = -\delta U/2$ and the corresponding Weber number is given by

$$We = \frac{\rho \delta U^2 \lambda}{\sigma} = 20.$$

According to linear potential flow theory the interface will be unstable if $We > 4\pi$ and therefore this initial condition will initially result in exponential growth of the interface amplitude $\hat{\eta}$.

In fig. 6 we show the evolution of the velocity magnitude where the time has been nondimensionalized using the capillary time scale $T^\sigma = \sqrt{\frac{2\rho\lambda^3}{\sigma}}$.

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Figures 7 and 8 show convergence to a reference solution which was computed using the author's implementation of the boundary integral method (BIM) described in [1] (therein referred to as Method III).

4.2 Third-order Stokes wave

We consider the simulation of a third-order Stokes wave, for details about the initial conditions and domain size, see [2]. Contrary to [2] we do not initialize the velocity field in the gas phase, hence $u^g = 0$ at $t = 0$. The density ratio is given by $\rho^g/\rho^l \approx 1.17 \cdot 10^{-3}$. In the notation of [2], we let the Reynolds and Bond number be given by $Re = 4 \cdot 10^4$ and $Bo = 10^3$ resulting in 'easy to resolve' interface layers and interface length scales. Hence this type of wave is appropriate for testing whether our model will converge to the desired Navier-Stokes solution for a sufficiently fine mesh. For a steepness of $\varepsilon = 0.55$ the initial condition results in a plunging breaker. In fig. 9 we show an example simulation using our proposed two-fluid model. Here time has been nondimensionalized in terms of the wave period $T^g = \frac{2\pi}{\sqrt{gk}}$.

In fig. 10 we show the convergence of the interface profile at three time instances as well as the convergence of the tangential velocity jump $[[u_\tau]]$ to zero. This shows that our two-fluid model essentially reduces to the one-fluid model once the interface layer is resolved.

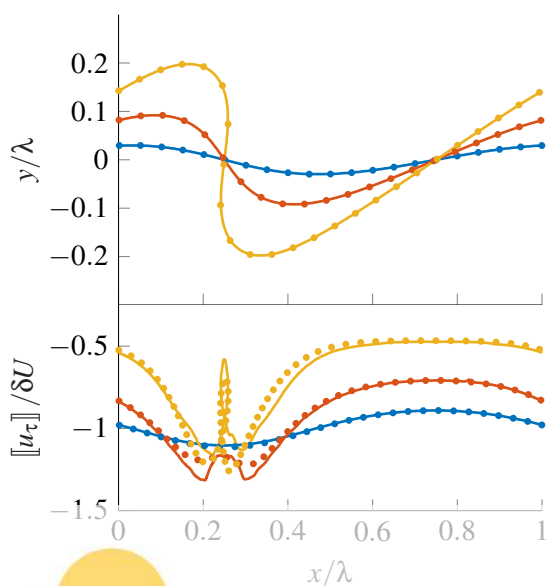


Figure 7: Example of the interface profile (top) and vortex sheet strength $[[u_\tau]]$ (bottom) for the Kelvin-Helmholtz problem using $h/\lambda = 1/512$. The solid lines correspond to the approximate solution using our proposed method whereas the markers correspond to the BIM solution.

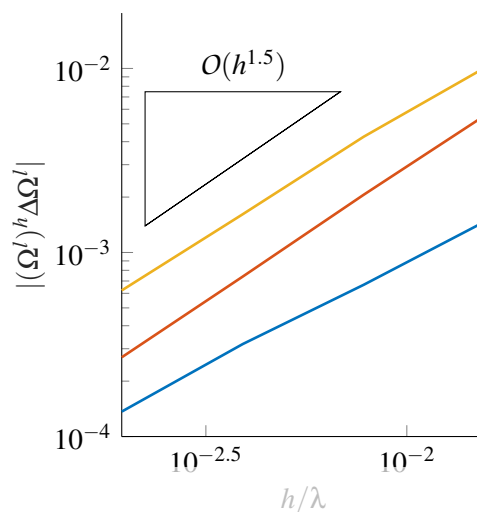


Figure 8: Area of the symmetric difference of the approximate liquid domain compared to the BI solution. The blue, red and yellow lines correspond to $t/T^g = 0.3, 0.4$ and 0.5 respectively.

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Figure 9: Vorticity (clipped to $\omega_y \in [-100, 100]$) and interface profile on time instances $t/T^g \approx 0.20, 0.36, \dots, 0.84$. Here $h/\lambda = 2^{-10}$ (using AMR).

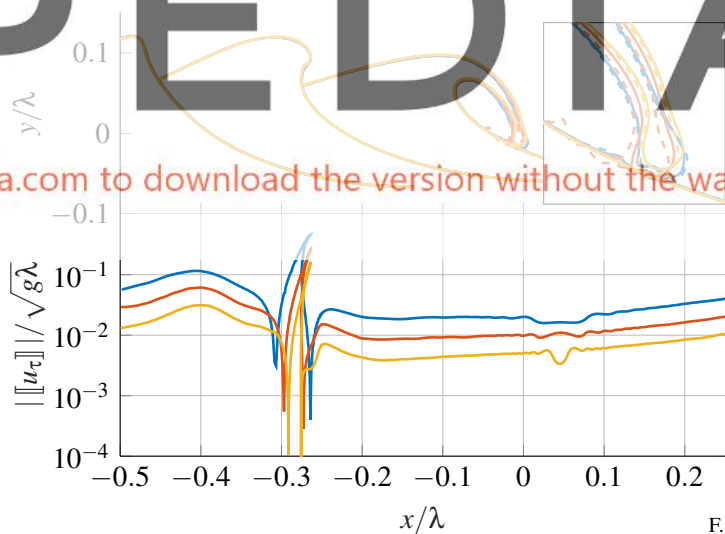


Figure 10: Convergence of the interface profile (at $t/T^g \approx 0.36, 0.52, 0.68$) and tangential velocity discontinuity (at $t/T^g \approx 0.52$) for $h/\lambda = 2^{-l}$, where $l = 9, 10, 11$ (using AMR) corresponding to blue, red and yellow respectively. The dashed and solid lines correspond to the one- and two-fluid model respectively.

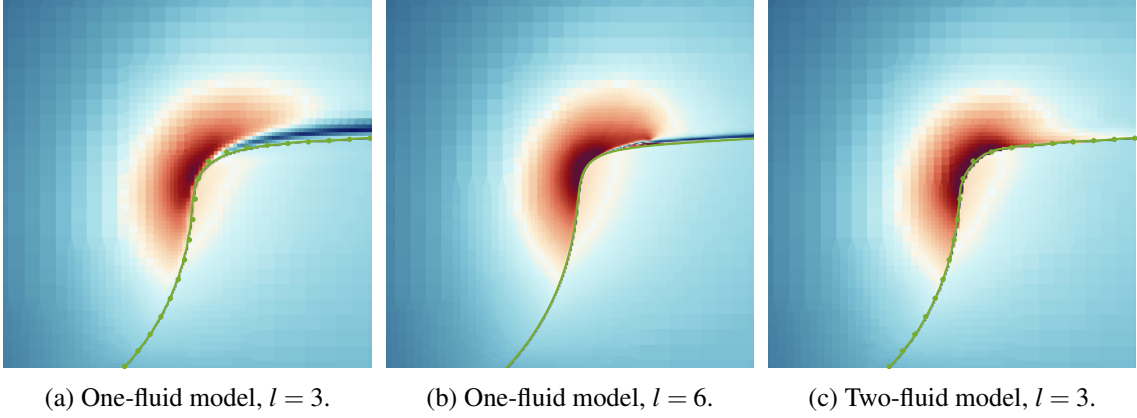


Figure 11: The velocity magnitude (clipped to $|\mathbf{u}| \in [0, 15]$ and zoomed in at $[3, 6] \times [5, 8]$) resulting from a large gas-pocket impact at $t = 1.62$. The refinement level l refers to an interface resolution of $h = 2^{-(2+l)}$. The markers correspond to the interface profile resulting from the one-fluid model with $l = 6$.

4.3 Large gas-pocket impact

Finally we consider a high Reynolds number test case for which resolving the interface layer is too expensive. We simulate a smoothed dam break which results in a large gas-pocket impact [3]. The fluids (water and air) are initially at rest with the following interface profile

$$y = 7.6 + 3.6 \tanh(0.44(x - 15.5)),$$

on a domain given by $\Omega = [0, 20] \times [0, 12] \setminus E$, where E is an ellipse centered at $\mathbf{0}$ with radii given by 18 and 2.8 respectively.

In fig. 11 we show close-ups of the velocity magnitude as well as the interface profile at $t = 1.62$ for the one- and two-fluid model⁴. Here we use a resolution of $h = 2^{-(2+l)}$ where $l = 3$ denotes the maximum level of refinements (we refine at the interface as well as in regions of high vorticity). For reference we also include a more accurate solution resulting from the one-fluid model with $l = 6$.

We find that both the one- and two-fluid model yield an accurate interface profile. The one-fluid model however results in an artificially thick interface layer which is not present in the two-fluid model. Further research into how this affects the development of free surface instabilities is however still needed.

5 DISCUSSION

We have proposed a sharp two-fluid model for two-phase flow in which underresolved interface layers are modeled by a tangential velocity discontinuity. The numerical model is based on novel generalizations of state-of-the-art numerical methods for the one-fluid model. We demonstrate that the resulting model is able to simulate inviscid (Kelvin-Helmholtz), viscous (Stokes-3 wave) as well as high-Reynolds number (large gas-pocket impact) two-phase flow problems at high accuracy.

⁴We picked an early time before impact such that it is still relatively easy to obtain a (nearly) resolved solution.

We will use the proposed model for the simulation of breaking wave impacts and in particular will study the development of free surface instabilities before impact. For future work we suggest the inclusion of compressibility as well as phase-change, the latter of which can sharply be taken into account by including a mass flux through the interface in the CCM divergence operator.

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