

Observation of a vector charmoniumlike state in $e^+e^- \rightarrow D_s^+ D_{s1}^-(2536)^- + c.c.$

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Using a data sample of 921.9 fb^{-1} collected with the Belle detector, we study the process of $e^+e^- \rightarrow D_s^+ D_{s1}(2536)^- + \text{c.c.}$ via initial-state radiation. We report the first observation of a vector charmoniumlike state decaying to $D_s^+ D_{s1}(2536)^- + \text{c.c.}$ with a significance of 5.9σ , including systematic uncertainties. The measured mass and width are $(4625.9_{-6.0}^{+6.2}(\text{stat}) \pm 0.4(\text{syst})) \text{ MeV}/c^2$ and $(49.8_{-11.5}^{+13.9}(\text{stat}) \pm 4.0(\text{syst})) \text{ MeV}$, respectively. The product of the $e^+e^- \rightarrow D_s^+ D_{s1}(2536)^- + \text{c.c.}$ cross section and the branching fraction of $D_{s1}(2536)^- \rightarrow \bar{D}^{*0} K^-$ is measured from the $D_s \bar{D}_{s1}(2536)$ threshold to 5.59 GeV .

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In the past decade, measurements of the exclusive cross sections for e^+e^- annihilation into charmed or charmed-strange meson pairs above the open-charm threshold have attracted much attention [1–10]. These open-charm final states are dominantly produced from the Okubo-Zweig-Iizuka (OZI)-allowed strong decays of excited vector charmonium states (ψ states). A comprehensive study of the exclusive e^+e^- cross sections to various open-charm final states could help one to understand the couplings of these ψ states, and extract their resonant parameters.

Many additional Y states with $J^{PC} = 1^{--}$ with masses above the open-charm threshold have been discovered in the last 14 years [11–19]. It has been noticed that the Y states above the open-charm threshold do not appear explicitly as peaks either in the total hadronic cross section or in the exclusive e^+e^- cross sections to open-charm final states [20] [the only vector charmoniumlike states which reveal themselves as peaks at threshold are the $Y(4630)$ and $Y(4220)$ observed in the $\Lambda_c^+ \bar{\Lambda}_c^-$ and $\pi^+ D^0 D^{*-}$ final states, respectively [21,22]]. In $e^+e^- \rightarrow Y \rightarrow \pi^+ \pi^- J/\psi$ and $\pi^+ \pi^- \psi(2S)$ [$Y = Y(4260)$, $Y(4660)$] processes, events in $\pi^+ \pi^-$ mass spectra tend to accumulate at the $f_0(980)$ nominal mass, which has an $s\bar{s}$ component. Thus, it is natural to search for Y states with a $(c\bar{s})(\bar{c}s)$ quark component. As mentioned in Ref. [23], bound states of $D_s \bar{D}_s$ mesons, e.g., $D_s \bar{D}_{s1}(2536)$, can appear as a result of $f_0(980)$ exchange. Unfortunately, open-charmed-strange production associated with these Y states has not yet been observed.

In this Letter, we perform a measurement of the exclusive cross section for $e^+e^- \rightarrow D_s^+ D_{s1}(2536)^- (D_{s1}(2536)^- \rightarrow \bar{D}^{*0} K^- / D^{*-} K_S^0)$ as a function of center-of-mass (C.M.) energy from the $D_s^+ D_{s1}(2536)^-$ mass threshold to 5.59 GeV

via initial-state radiation (ISR) [24]. In this process, a charmoniumlike state decaying to $D_s^+ D_{s1}(2536)^-$ is observed for the first time. The data used in this analysis correspond to 921.9 fb^{-1} of integrated luminosity at C.M. energies of 10.52 , 10.58 , and 10.867 GeV collected by the Belle detector [25] at the KEKB asymmetric-energy e^+e^- collider [26,27].

We use PHOKHARA [28] to generate signal Monte Carlo (MC) events, determine the detector efficiency, and optimize selection criteria for signal events. Generic MC samples of $\Upsilon(4S) \rightarrow B^+ B^- / B^0 \bar{B}^0$, $\Upsilon(5S) \rightarrow B_s^{(*)} \bar{B}_s^{(*)}$, and $e^+e^- \rightarrow q\bar{q}$ ($q = u, d, s, c$) at $\sqrt{s} = 10.52$, 10.58 , and 10.867 GeV with four times the luminosity of data are used to study possible backgrounds.

We fully reconstruct the ISR photon γ_{ISR} , D_s^+ , and K^- / K_S^0 , but do not reconstruct the \bar{D}^{*0} / D^{*-} . Since the \bar{D}^{*0} / D^{*-} decays are not reconstructed, the detection efficiency for the $e^+e^- \rightarrow D_s^+ D_{s1}(2536)^- (\rightarrow \bar{D}^{*0} K^- / D^{*-} K_S^0)$ process is greatly improved. For the measurement of the $e^+e^- \rightarrow D_s^+ D_{s1}(2536)^-$ cross section, we determine the invariant mass spectrum of $D_s^+ D_{s1}(2536)^-$ ($M(D_s^+ D_{s1}(2536)^-)$), which is equivalent to the mass recoiling against γ_{ISR} ($M_{\text{rec}}(\gamma_{\text{ISR}})$). Here, $M_{\text{rec}}(\gamma_{\text{ISR}})$ is calculated using $M_{\text{rec}}(\gamma_{\text{ISR}}) = \sqrt{(P_{\text{C.M.}} - P_{\gamma_{\text{ISR}}})^2}$, where $P_{\text{C.M.}}$ and $P_{\gamma_{\text{ISR}}}$ are the four-momenta of the initial e^+e^- system and the ISR photon, respectively. However, the energy resolution of γ_{ISR} is very poor due to its high energy. We constrain the recoil mass of the $\gamma_{\text{ISR}} D_s^+ K^- / \gamma_{\text{ISR}} D_s^+ K_S^0$ to the nominal mass of the \bar{D}^{*0} / D^{*-} meson [29] to improve the resolution for the ISR photon for events within the \bar{D}^{*0} / D^{*-} signal region. Before applying the mass constraint, the mass resolution of the $M(D_s^+ D_{s1}(2536)^-)$ system is about $180 \text{ MeV}/c^2$. As a result of the constraint, the mass resolution is significantly improved, to about $5 \text{ MeV}/c^2$.

The D_s^+ candidates are reconstructed using eight decay modes: $\phi\pi^+$, $\bar{K}^*(892)^0 K^+$, $K_S^0 K^+$, $K^+ K^- \pi^+ \pi^0$, $K_S^0 \pi^0 K^+$, $K^*(892)^+ K_S^0$, $\eta\pi^+$, and $\eta'\pi^+$. We use the techniques of Ref. [30] to reconstruct particles such as photons, charged

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pions and kaons, and K_S^0 . The ϕ , $\bar{K}^*(892)^0$, and $K^*(892)^+$ candidates are reconstructed in the K^+K^- , $K^-\pi^+$, and $K_S^0\pi^+$ decay modes. The invariant masses of the K_S^0 , ϕ , $\bar{K}^*(892)^0$, and $K^*(892)^+$ candidates are required to be within 10, 10, 50, and 50 MeV/c^2 of the corresponding nominal masses [29] (>95% signal events are retained), respectively.

The most energetic ISR photon is required to have energy greater than 3 GeV in the e^+e^- C.M. frame. Pairs of photons are combined to form π^0 candidates. The energies of the photons from π^0 are required to be greater than 50 MeV in the calorimeter barrel and 100 MeV in the calorimeter end caps [31] in the laboratory frame. The η candidates are reconstructed via $\gamma\gamma$ and $\pi^+\pi^-\pi^0$ decay modes. Photon candidates from $\eta \rightarrow \gamma\gamma$ are required to have energies greater than 100 MeV in the laboratory frame. The reconstructed η candidates are then combined with $\pi^+\pi^-$ pairs to form η' candidates. The mass windows applied for π^0 , $\eta \rightarrow \gamma\gamma$, $\eta \rightarrow \pi^+\pi^-\pi^0$, and η' candidates are ± 12 , ± 20 , ± 10 , and ± 10 MeV/c^2 , which are within approximately 2.5σ of the corresponding meson nominal masses [29]. After applying the mass window requirements, mass-constrained fits are applied to the π^0 , η , and η' candidates to improve their momentum resolutions.

Before calculation of the D_s^+ candidate mass, a fit to a common vertex is performed for charged tracks in the D_s^+ candidate. After the application of the above requirements, D_s^+ signals are clearly observed. We define the D_s^+ signal region as $|M(D_s^+) - m_{D_s^+}| < 12$ MeV/c^2 ($\sim 2.0\sigma$). Here and throughout the text, m_i represents the nominal mass of particle i [29]. To improve the momentum resolution of the D_s^+ meson candidate, a mass-constrained fit to the D_s^+ nominal mass [29] is performed. The D_s^+ mass sideband regions are defined as $1912.34 < M(D_s^+) < 1936.34$ MeV/c^2 and $2000.34 < M(D_s^+) < 2024.34$ MeV/c^2 , which are twice as wide as the signal region. The D_s^+ candidates from the sidebands are also constrained to the central mass values in the defined D_s^+ sideband regions. The D_s^+ candidate with the smallest χ^2 from the D_s^+ mass fit is kept. Besides the selected ISR photon and D_s^+ , we require at least one additional K^- or K_S^0 candidate in the event, and retain all the combinations (the fraction of events with multiple candidates is 1.7%).

Figure 1(a) shows the sum of the recoil mass spectra against the $\gamma_{\text{ISR}}D_s^+K^-$ and $\gamma_{\text{ISR}}D_s^+K_S^0$ systems after requiring the events be within the $D_{s1}(2536)^-$ signal region (see below) in data. Due to the poor recoil mass resolution, the \bar{D}^{*0}/D^{*-} signal is very wide. The \bar{D}^{*0}/D^{*-} signal component is modeled using a Gaussian function convolved with a Novosibirsk function [32] derived from the signal MC samples, while the combinatorial backgrounds are described by a second-order polynomial. The solid curve is the total fit; the \bar{D}^{*0}/D^{*-} signal yield is 275 ± 32 . We define an asymmetric requirement of $-200 < M_{\text{rec}}(\gamma_{\text{ISR}}D_s^+K^-/K_S^0) - m_{\bar{D}^{*0}/D^{*-}} < 400$ MeV/c^2 for the \bar{D}^{*0}/D^{*-} signal region.

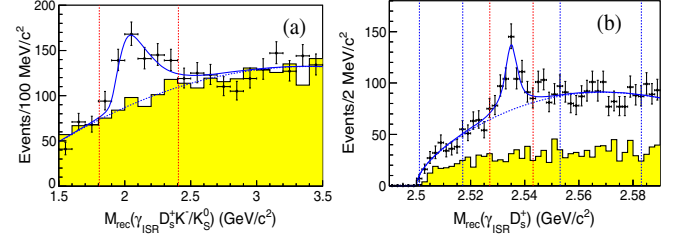


FIG. 1. (a) The recoil mass spectrum against the $\gamma_{\text{ISR}}D_s^+K^-/K_S^0$ system before applying the \bar{D}^{*0}/D^{*-} mass constraint. The yellow histogram shows the normalized $D_{s1}(2536)^-$ mass sidebands (see below). The red dashed lines show the required \bar{D}^{*0}/D^{*-} signal region. (b) The recoil mass spectrum against the $\gamma_{\text{ISR}}D_s^+$ system in data. The yellow histogram shows the normalized $D_{s1}(2536)^-$ mass sidebands. The red dashed lines show the required $D_{s1}(2536)^-$ signal region, and the blue dashed lines show the $D_{s1}(2536)^-$ mass sidebands.

Hereinafter the \bar{D}^{*0}/D^{*-} mass constraint is applied for events in the \bar{D}^{*0}/D^{*-} signal region to improve mass resolution.

The recoil mass spectrum against the $\gamma_{\text{ISR}}D_s^+$ system after requiring the events within \bar{D}^{*0}/D^{*-} signal region is shown in Fig. 1(b). A clear $D_{s1}(2536)^-$ signal is observed. The signal shape is described by a double Gaussian function (all the parameters are fixed to those from a fit to the MC simulated distribution), and a threshold function is used for the backgrounds. The threshold function is $(M_{\text{rec}} - x_{\text{thr}})^\alpha e^{\beta_1(M_{\text{rec}} - x_{\text{thr}}) + \beta_2(M_{\text{rec}} - x_{\text{thr}})^2}$, where M_{rec} is the recoil mass of the $\gamma_{\text{ISR}}D_s^+$; the parameters α , β_1 , and β_2 are free; the threshold parameter x_{thr} is fixed from generic MC simulations. The fit yields 254 ± 36 $D_{s1}(2536)^-$ signal events as shown in Fig. 1(b) [33]. We define the $D_{s1}(2536)^-$ signal region as $|M_{\text{rec}}(\gamma_{\text{ISR}}D_s^+) - m_{D_{s1}(2536)^-}| < 8$ MeV/c^2 ($\sim 2.5\sigma$), and sideband regions as shown by blue dashed lines, which are three times as wide as the signal region. To estimate the signal significance of the $D_{s1}(2536)^-$, we compute $\sqrt{-2\ln(\mathcal{L}_0/\mathcal{L}_{\text{max}})}$ [34], where \mathcal{L}_0 and \mathcal{L}_{max} are the maximized likelihoods without and with the $D_{s1}(2536)^-$ signal, respectively. The statistical significance of the $D_{s1}(2536)^-$ signal is 8σ .

The $D_s^+D_{s1}(2536)^-$ invariant mass distribution is shown in Fig. 2(a). There is a significant peak around 4626 MeV/c^2 , while no structure is seen in the normalized $D_{s1}(2536)^-$ mass sidebands shown as the yellow histogram. In addition, no peaking background is found in the $D_s^+D_{s1}(2536)^-$ mass distribution from generic MC samples. We therefore interpret the peak in the data as evidence for an exotic charmoniumlike state [35] decaying into $D_s^+D_{s1}(2536)^-$, called $Y(4626)$ hereafter.

One possible background, which is not included in the $D_{s1}(2536)^-$ mass sidebands, is from $e^+e^- \rightarrow D_s^{*+}(\rightarrow D_s^+\gamma)D_{s1}(2536)^-$, where the photon from the D_s^{*+} remains undetected. To estimate such a background contribution, we measure this process with the data

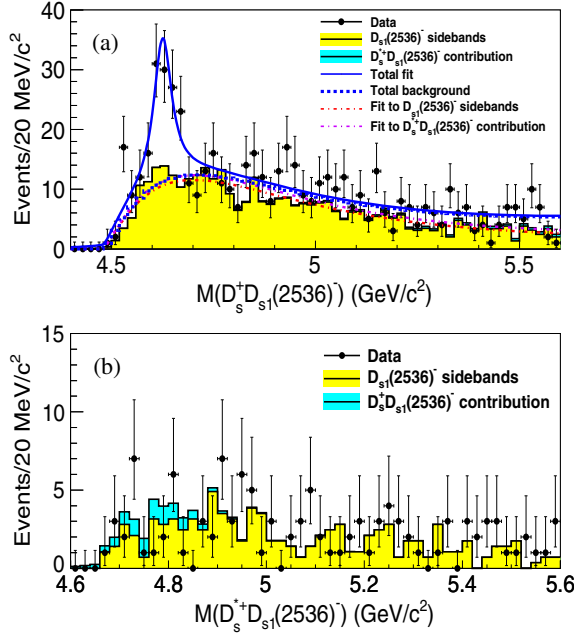


FIG. 2. (a) The $D_s^+ D_{s1}^{2536-}$ invariant mass spectrum for $e^+e^- \rightarrow D_s^+ D_{s1}^{2536-}$. (b) The $D_s^{*+} D_{s1}^{2536-}$ invariant mass spectrum for $e^+e^- \rightarrow D_s^{*+} D_{s1}^{2536-}$. All the components including those from the fit to the $D_s^+ D_{s1}^{2536-}$ invariant mass spectrum are indicated in the labels and described in the text. Note that the cyan shaded histograms in the top/bottom show the $D_s^+ D_{s1}^{2536-}/D_s^{*+} D_{s1}^{2536-}$ invariant mass spectrum from $D_s^{*+} D_{s1}^{2536-}/D_s^+ D_{s1}^{2536-}$ background contribution after applying the requirements to reconstruct $e^+e^- \rightarrow D_s^+ D_{s1}^{2536-}/e^+e^- \rightarrow D_s^{*+} D_{s1}^{2536-}$.

following the same procedure as used for the signal process. We require an extra photon with $E_\gamma > 50$ MeV in the barrel or $E_\gamma > 100$ MeV in the end caps to combine with the D_s^+ to form the D_s^{*+} candidate. The mass and vertex fits are applied to the D_s^{*+} candidates to improve their momentum resolution. In events with multiple candidates, the best candidate is chosen using the lowest χ^2 value from the mass-constrained fit. The same \bar{D}^{*0}/D^{*-} signal region requirement on $M_{\text{rec}}(\gamma_{\text{ISR}} D_s^{*+} K^-/K_S^0)$ and the \bar{D}^{*0}/D^{*-} mass constraint are applied as before in $e^+e^- \rightarrow D_s^+ D_{s1}^{2536-}$. In the recoil mass spectrum of the $\gamma_{\text{ISR}} D_s^{*+}$ an excess of 28 ± 13 D_{s1}^{2536-} signal events with a statistical significance of 2.4σ is observed in the D_{s1}^{2536-} signal region.

After requiring the $D_s^+ K^-/K_S^0$ mass to be within the D_{s1}^{2536-} signal region, the $D_s^{*+} D_{s1}^{2536-}$ invariant mass distribution is shown in Fig. 2(b). Note that the process $e^+e^- \rightarrow D_s^+ D_{s1}^{2536-}$ is a source of backgrounds for the $e^+e^- \rightarrow D_s^{*+} D_{s1}^{2536-}$ when the D_s^+ candidates are combined with low energy photons to form D_s^{*+} candidates. From Fig. 2(b), no obvious structure is observed. The normalized contribution from $e^+e^- \rightarrow D_s^{*+} D_{s1}^{2536-}$ to $e^+e^- \rightarrow D_s^+ D_{s1}^{2536-}$ is the cyan shaded histogram which is shown in

Fig. 2(a), and which is normalized to correspond to $N_{D_s^{*+} D_{s1}^{2536-}}^{\text{obs}} \epsilon_{D_s^{*+} D_{s1}^{2536-}} / \epsilon_{D_s^+ D_{s1}^{2536-}}$ events. Here $N_{D_s^{*+} D_{s1}^{2536-}}^{\text{obs}}$ is the yield of $e^+e^- \rightarrow D_s^{*+} D_{s1}^{2536-}$ signal events in each $M(D_s^{*+} D_{s1}^{2536-})$ bin in data after subtracting the normalized D_{s1}^{2536-} sidebands and the $e^+e^- \rightarrow D_s^+ D_{s1}^{2536-}$ background contribution, and $\epsilon_{D_s^+ D_{s1}^{2536-}}$ and $\epsilon_{D_s^{*+} D_{s1}^{2536-}}$ are the reconstruction efficiencies for $e^+e^- \rightarrow D_s^+ D_{s1}^{2536-}$ and $e^+e^- \rightarrow D_s^{*+} D_{s1}^{2536-}$, respectively, where the ratio of efficiencies is (1.00 ± 0.02) . The yield of $D_s^{*+} D_{s1}^{2536-}$ after background subtraction for the entire region in Fig. 2(b) is (11.6 ± 3.6) . A similar method is applied to estimate the background contribution from $e^+e^- \rightarrow D_s^+ D_{s1}^{2536-}$ to $e^+e^- \rightarrow D_s^{*+} D_{s1}^{2536-}$.

We perform an unbinned likelihood fit simultaneously to the $M(D_s^+ D_{s1}^{2536-})$ distributions of all selected D_{s1}^{2536-} signal candidates, the normalized D_{s1}^{2536-} mass sidebands, and the $e^+e^- \rightarrow D_s^{*+} D_{s1}^{2536-}$ contribution. The yields in the normalized D_{s1}^{2536-} mass sidebands and the $e^+e^- \rightarrow D_s^{*+} D_{s1}^{2536-}$ contribution are fixed in the fit. The following components are included in the fit to the $M(D_s^+ D_{s1}^{2536-})$ distribution: a resonance signal, a nonresonant contribution, the D_{s1}^{2536-} mass sidebands, and an $e^+e^- \rightarrow D_s^{*+} D_{s1}^{2536-}$ contribution. A Breit-Wigner (BW) function convolved with a Gaussian function (with its width fixed at 5.0 MeV/c^2 according to the MC simulation), multiplied by an efficiency function that has a linear dependence on $M(D_s^+ D_{s1}^{2536-})$ and the differential ISR effective luminosity [36], is taken as the signal shape. Here the BW formula used has the form [37]

$$BW(\sqrt{s}) = \frac{\sqrt{12\pi\Gamma_{ee}\mathcal{B}_f\Gamma}}{s - M^2 + iM\Gamma} \sqrt{\frac{\Phi_2(\sqrt{s})}{\Phi_2(M)}}, \quad (1)$$

where M is the mass of the resonance, Γ and Γ_{ee} are the total width and partial width to e^+e^- , $\mathcal{B}_f = \mathcal{B}(Y(4626) \rightarrow D_s^+ D_{s1}^{2536-}) \times \mathcal{B}(D_{s1}^{2536-} \rightarrow \bar{D}^{*0} K^-)$ is the product branching fraction of the $Y(4626)$ into the final state, and Φ_2 is the two-body decay phase space factor that increases smoothly from the mass threshold with \sqrt{s} , respectively. A two-body phase space form is also taken into account for the nonresonant contribution. The D_{s1}^{2536-} mass sidebands and the $e^+e^- \rightarrow D_s^{*+} D_{s1}^{2536-}$ contribution are parametrized with threshold functions.

The fit results are shown in Fig. 2(a), where the solid blue curve is the best fit, the blue dotted curve is the sum of the backgrounds, the red dot-dashed curve is the fitted result to the normalized D_{s1}^{2536-} mass sidebands, and the violet dot-dashed curve is for the $e^+e^- \rightarrow D_s^{*+} D_{s1}^{2536-}$ contribution. The yield of the $Y(4626)$ signal is 89_{-16}^{+17} . The statistical significance of the $Y(4626)$

signal is 6.5σ , calculated from the difference of the logarithmic likelihoods [34], $-2 \ln(\mathcal{L}_0/\mathcal{L}_{\max}) = 50.4$, where \mathcal{L}_0 and \mathcal{L}_{\max} are the maximized likelihoods without and with a signal component, respectively, taking into account the difference in the number of degrees of freedom ($\Delta\text{ndf} = 3$). The parametrization of the nonresonant contribution is the dominant systematic uncertainty for the estimate of the signal significance. Changing the two-body phase space form to a threshold function parametrized by $\sqrt{M - x_{\text{thr}}}$ or a two-body phase space form plus a threshold function, the $Y(4626)$ signal significance is reduced to 5.9σ . We take this value as the signal significance with systematic uncertainties included. The fitted mass and width for the $Y(4626)$ are $(4625.9^{+6.2}_{-6.0}(\text{stat}) \pm 0.4(\text{syst})) \text{ MeV}/c^2$ and $(49.8^{+13.9}_{-11.5}(\text{stat}) \pm 4.0(\text{syst})) \text{ MeV}$, respectively. The value of $\Gamma_{ee} \times \mathcal{B}(Y(4626) \rightarrow D_s^+ D_{s1}(2536)^-) \times \mathcal{B}(D_{s1}(2536)^- \rightarrow \bar{D}^{*0} K^-)$ is obtained to be $(14.3^{+2.8}_{-2.6}(\text{stat}) \pm 1.5(\text{syst})) \text{ eV}$. The systematic uncertainties are discussed below.

The $e^+e^- \rightarrow D_s^+ D_{s1}(2536)^-$ cross section is extracted from the recoil mass spectrum against the $\gamma_{\text{ISR}} D_s^+$ system. The product of the $e^+e^- \rightarrow D_s^+ D_{s1}(2536)^-$ dressed cross section (σ) [38] and the decay branching fraction $\mathcal{B}(D_{s1}(2536)^- \rightarrow \bar{D}^{*0} K^-)$ for each $D_s^+ D_{s1}(2536)^-$ mass bin from threshold to $5.59 \text{ GeV}/c^2$ in steps of $20 \text{ MeV}/c^2$ is computed as

$$\sigma(e^+e^- \rightarrow D_s^+ D_{s1}(2536)^-) \mathcal{B}(D_{s1}(2536)^- \rightarrow \bar{D}^{*0} K^-) = \frac{N_{\text{fit}}^{D_{s1}(2536)^-}}{d\mathcal{L} \times [\sum_i (\varepsilon_i^{\bar{D}^{*0} K^-} \times \mathcal{B}_i) + R_{\bar{D}^{*0} K^-}^{D^* K_S^0} \times \sum_i (\varepsilon_i^{D^* K_S^0} \times \mathcal{B}_i)]}, \quad (2)$$

where $N_{\text{fit}}^{D_{s1}(2536)^-}$ is the yield of fitted $D_{s1}(2536)^-$ signal events after subtracting the $e^+e^- \rightarrow D_s^{*+} D_{s1}(2536)^-$ background contribution in data, $d\mathcal{L}$ is the effective luminosity [36], $\sum_i (\varepsilon_i^{\bar{D}^{*0} K^-} \times \mathcal{B}_i)$ and $\sum_i (\varepsilon_i^{D^* K_S^0} \times \mathcal{B}_i)$ are the sums of the product of the reconstruction efficiency and branching fraction for each D_s^+ decay mode (i) in $D_{s1}(2536)^- \rightarrow \bar{D}^{*0} K^-$ and $D_{s1}(2536)^- \rightarrow D^* K_S^0$, and $R_{\bar{D}^{*0} K^-}^{D^* K_S^0} = \mathcal{B}(D_{s1}(2536)^- \rightarrow D^* K_S^0) / \mathcal{B}(D_{s1}(2536)^- \rightarrow \bar{D}^{*0} K^-) = 0.425 \pm 0.06$ taken from Ref. [29]. The values used to calculate $\sigma(e^+e^- \rightarrow D_s^+ D_{s1}(2536)^-) \times \mathcal{B}(D_{s1}(2536)^- \rightarrow \bar{D}^{*0} K^-)$ are summarized in Supplemental Material [33]. In the fit to the recoil mass spectrum of $\gamma_{\text{ISR}} D_s^+$ combinations in each $D_s^+ D_{s1}(2536)^-$ mass bin, the $D_{s1}(2536)^-$ signal shape is fixed to that from the overall fit, as shown by the blue solid curve in Fig. 1, and a threshold function is used for the backgrounds. The resulting $\sigma(e^+e^- \rightarrow D_s^+ D_{s1}(2536)^-) \times \mathcal{B}(D_{s1}(2536)^- \rightarrow \bar{D}^{*0} K^-)$ value as a function of $M(D_s^+ D_{s1}(2536)^-)$ is shown in Fig. 3 with the statistical and systematic uncertainties discussed below summed in quadrature.

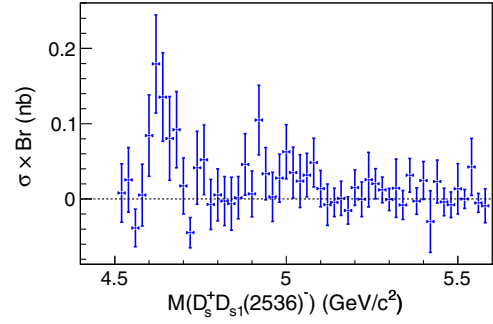


FIG. 3. The product of the $e^+e^- \rightarrow D_s^+ D_{s1}(2536)^-$ cross section and branching fraction $\mathcal{B}(D_{s1}(2536)^- \rightarrow \bar{D}^{*0} K^-)$ as a function of $M(D_s^+ D_{s1}(2536)^-)$; here the statistical and systematic uncertainties are summed in quadrature. Here, the correlated systematic uncertainties from the detection efficiency, branching fraction, and luminosity are not included.

The sources of systematic uncertainties for the cross section measurement include detection-efficiency-related uncertainties, branching fractions of the intermediate states, fit uncertainty, resonance parameters, the MC event generator, $e^+e^- \rightarrow D_s^{*+} D_{s1}(2536)^-$ background contribution, mass resolution as well as the integrated luminosity. The detection-efficiency-related uncertainties include those for tracking efficiency (0.35%/track), particle identification efficiency (1.1%/kaon and 0.9%/pion), K_S^0 selection efficiency (1.4%) [40], π^0 reconstruction efficiency (2.25%/ π^0) and photon reconstruction efficiency (2.0%/photon). The above individual uncertainties from different D_s^+ decay channels are added linearly, weighted by the product of the detection efficiency and D_s^+ partial decay width. These uncertainties are summed in quadrature to obtain the final uncertainty related to the reconstruction efficiency. Uncertainties for D_s^+ decay branching fractions and $R_{\bar{D}^{*0} K^-}^{D^* K_S^0}$ are taken from Ref. [29]; the final uncertainties on the D_s^+ partial decay widths are summed in quadrature over the eight D_s^+ decay modes weighted by the product of the efficiency and the D_s^+ partial decay width. Systematic uncertainties associated with the fitting procedure are estimated by changing the order of the background polynomial and the range of the fit. The deviations from nominal fit results are taken as systematic uncertainties. Changing the values of mass and width of $D_{s1}(2536)^-$ by 1σ [29] in each $M(D_s^+ D_{s1}(2536)^-)$ bin has no effect on the fits. Thus, the uncertainty from the resonance parameters can be neglected. The PHOKHARA generator calculates the ISR-photon radiator function with 0.1% accuracy [28]. The uncertainty attributed to the generator can also be neglected.

By fitting the $D_{s1}(2536)^-$ mass spectrum in each $M(D_s^+ D_{s1}(2536)^-)$ bin for $e^+e^- \rightarrow D_s^{*+} D_{s1}(2536)^-$, we find the signal yields are less than 1. In addition, the $D_{s1}(2536)^-$ signal from the $e^+e^- \rightarrow D_s^{*+} D_{s1}(2536)^-$ contribution has a much poorer mass resolution according to

TABLE I. Summary of the absolute systematic uncertainties (σ_{sys}) on the product of $e^+e^- \rightarrow D_s^+ D_{s1}(2536)^-$ cross section and the decay branching fraction $\mathcal{B}(D_{s1}(2536)^- \rightarrow \bar{D}^{*0} K^-)$ for different $M(D_s^+ D_{s1}(2536)^-)$ bins.

Source	σ_{sys} (pb)
Detection efficiency	0.0–8.1
Branching fractions	0.0–7.4
Fit uncertainty	0.5–36.2
Mass resolution	0.1–7.7
Luminosity	0.0–2.5
Quadratic sum	0.6–36.3

MC simulation. Therefore, the systematic uncertainty associated with the $e^+e^- \rightarrow D_s^{*+} D_{s1}(2536)^-$ contribution is neglected. The MC simulation is known to reproduce the resolution of mass peaks within 10% over a large number of different systems. The systematic uncertainty in the mass resolution is estimated by comparing the yields when the mass resolution is changed by 10%. The total luminosity is determined to 1.4% precision using wide-angle Bhabha scattering events. All the uncertainties are summarized in Table I. Assuming all the sources are independent, we sum them in quadrature to obtain the total systematic uncertainties.

The following systematic uncertainties on the measured mass and width for the $Y(4626)$, and the $\Gamma_{ee} \times \mathcal{B}(Y(4626) \rightarrow D_s^+ D_{s1}(2536)^-) \times \mathcal{B}(D_{s1}(2536)^- \rightarrow \bar{D}^{*0} K^-)$ are considered. The resultant systematic uncertainties attributed to the mass resolution in the width and $\Gamma_{ee} \times \mathcal{B}(Y(4626) \rightarrow D_s^+ D_{s1}(2536)^-) \times \mathcal{B}(D_{s1}(2536)^- \rightarrow \bar{D}^{*0} K^-)$ are 0.3 MeV and 0.1 eV. By changing the nonresonant background shape to a threshold function or to the sum of a two-body phase space form and a threshold function, the differences of 0.3 MeV/ c^2 and 3.9 MeV in the measured mass and width, and 1.3 eV for the $\Gamma_{ee} \times \mathcal{B}(Y(4626) \rightarrow D_s^+ D_{s1}(2536)^-) \times \mathcal{B}(D_{s1}(2536)^- \rightarrow \bar{D}^{*0} K^-)$, respectively, are taken as systematic uncertainties. The uncertainty in the efficiency correction from detection efficiency, branching fractions of the intermediate states, and integrated luminosity is 4.9%. Changing the efficiency function by 4.9% gives a 0.1 MeV/ c^2 change on the mass, 0.2 MeV on the width, and 0.7 eV on the product $\Gamma_{ee} \times \mathcal{B}(Y(4626) \rightarrow D_s^+ D_{s1}(2536)^-) \times \mathcal{B}(D_{s1}(2536)^- \rightarrow \bar{D}^{*0} K^-)$. Finally, the total systematic uncertainties on the $Y(4626)$ mass, width, and $\Gamma_{ee} \times \mathcal{B}(Y(4626) \rightarrow D_s^+ D_{s1}(2536)^-) \times \mathcal{B}(D_{s1}(2536)^- \rightarrow \bar{D}^{*0} K^-)$ are 0.4 MeV/ c^2 , 4.0 MeV, and 1.5 eV, respectively.

In summary, the product of the $e^+e^- \rightarrow D_s^+ D_{s1}(2536)^-$ cross section and the decay branching fraction $\mathcal{B}(D_{s1}(2536)^- \rightarrow \bar{D}^{*0} K^-)$ is measured over the C.M. energy range from the $D_s^+ D_{s1}(2536)^-$ mass threshold to 5.59 GeV for the first time. We observe the first

vector charmoniumlike state decaying to a charmed-antistrange and anticharmed-strange meson pair $D_s^+ D_{s1}(2536)^-$ with a signal significance of 5.9σ with systematic uncertainties included. The measured mass and width are $(4625.9_{-6.0}^{+6.2}(\text{stat}) \pm 0.4(\text{syst})) \text{ MeV}/c^2$ and $(49.8_{-11.5}^{+13.9}(\text{stat}) \pm 4.0(\text{syst})) \text{ MeV}$, respectively, which are consistent with the $Y(4660)$ mass of $(4643 \pm 9) \text{ MeV}/c^2$ and width of $(72 \pm 11) \text{ MeV}$ [29] within uncertainties. The $\Gamma_{ee} \times \mathcal{B}(Y(4626) \rightarrow D_s^+ D_{s1}(2536)^-) \times \mathcal{B}(D_{s1}(2536)^- \rightarrow \bar{D}^{*0} K^-)$ is obtained to be $(14.3_{-2.6}^{+2.8}(\text{stat}) \pm 1.5(\text{syst})) \text{ eV}$.

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- [1] G. Pakhlova *et al.* (Belle Collaboration), *Phys. Rev. D* **77**, 011103 (2008).
- [2] G. Pakhlova *et al.* (Belle Collaboration), *Phys. Rev. Lett.* **98**, 092001 (2007).
- [3] G. Pakhlova *et al.* (Belle Collaboration), *Phys. Rev. Lett.* **100**, 062001 (2008).
- [4] G. Pakhlova *et al.* (Belle Collaboration), *Phys. Rev. D* **80**, 091101 (2009).
- [5] G. Pakhlova *et al.* (Belle Collaboration), *Phys. Rev. D* **83**, 011101(R) (2011).
- [6] B. Aubert *et al.* (BABAR Collaboration), *Phys. Rev. D* **76**, 111105 (2007).
- [7] B. Aubert *et al.* (BABAR Collaboration), *Phys. Rev. D* **79**, 092001 (2009).
- [8] P. del Amo Sanchez *et al.* (BABAR Collaboration), *Phys. Rev. D* **82**, 052004 (2010).
- [9] D. Cronin-Hennessy *et al.* (CLEO Collaboration), *Phys. Rev. D* **80**, 072001 (2009).
- [10] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Rev. Lett.* **120**, 132001 (2018).
- [11] B. Aubert *et al.* (BABAR Collaboration), *Phys. Rev. Lett.* **95**, 142001 (2005).
- [12] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Rev. Lett.* **100**, 102003 (2008).
- [13] T. E. Coan *et al.* (CLEO Collaboration), *Phys. Rev. Lett.* **96**, 162003 (2006).
- [14] C. Z. Yuan *et al.* (Belle Collaboration), *Phys. Rev. Lett.* **99**, 182004 (2007).
- [15] Z. Q. Liu *et al.* (Belle Collaboration), *Phys. Rev. Lett.* **110**, 252002 (2013).
- [16] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Rev. Lett.* **118**, 092001 (2017).
- [17] B. Aubert *et al.* (BABAR Collaboration), *Phys. Rev. Lett.* **98**, 212001 (2007).
- [18] X. L. Wang *et al.* (Belle Collaboration), *Phys. Rev. Lett.* **99**, 142002 (2007).
- [19] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Rev. D* **96**, 032004 (2017).
- [20] T. V. Uglov, Y. S. Kalashnikova, A. V. Nefediev, G. V. Pakhlova, and P. N. Pakhlov, *JETP Lett.* **105**, 1 (2017).
- [21] G. Pakhlova *et al.* (Belle Collaboration), *Phys. Rev. Lett.* **101**, 172001 (2008).
- [22] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Rev. Lett.* **122**, 102002 (2019).
- [23] M. Karliner and J. L. Rosner, *Nucl. Phys.* **A954**, 365 (2016).
- [24] Throughout this analysis, for any given mode, the corresponding charge-conjugate mode is implied.
- [25] A. Abashian *et al.* (Belle Collaboration), *Nucl. Instrum. Methods Phys. Res., Sect. A* **479**, 117 (2002); also see Section II in J. Brodzicka *et al.*, *Prog. Theor. Exp. Phys.* **2012**, 04D001 (2012).
- [26] S. Kurokawa and E. Kikutani, *Nucl. Instrum. Methods Phys. Res., Sect. A* **499**, 1 (2003), and other papers included in this Volume.
- [27] T. Abe *et al.*, *Prog. Theor. Exp. Phys.* **2013**, 03A001 (2013) and references therein.
- [28] G. Rodrigo, H. Czyż, J. H. Kühn, and M. Szopa, *Eur. Phys. J. C* **24**, 71 (2002).
- [29] M. Tanabashi *et al.* (Particle Data Group), *Phys. Rev. D* **98**, 030001 (2018).
- [30] S. Jia *et al.* (Belle Collaboration), *Phys. Rev. D* **98**, 092015 (2018).
- [31] H. Ikeda *et al.* (Belle Collaboration), *Nucl. Instrum. Methods Phys. Res., Sect. A* **441**, 401 (2000).
- [32] The Novosibirsk function is defined as $f(x) = \exp[-\frac{1}{2}(\ln^2(1+\Lambda(x-x_0))/\tau^2 + \tau^2)]$ with $\Lambda = \sinh(\tau\sqrt{\ln 4})/(\sigma\sqrt{\ln 4})$. The parameters represent the mean (x_0), the width (σ) and the tail asymmetry (τ).
- [33] See Supplemental Material at <http://link.aps.org/supplemental/10.1103/PhysRevD.100.111103>, for the recoil mass spectra against the $\gamma_{\text{ISR}}D_s^+$ system in $e^+e^- \rightarrow D_s^+D_{s1}(2536)^- (\rightarrow \bar{D}^{*0}K^-)$ and $e^+e^- \rightarrow D_s^+D_{s1}(2536)^- (\rightarrow D^{*-}K_S^0)$ and for a summary of the values for the effective luminosity, the total reconstruction efficiency, the number of fitted signal events, and the product of the dressed cross section and the decay branching fraction $\sigma(e^+e^- \rightarrow D_s^+D_{s1}(2536)^-) \times \mathcal{B}(D_{s1}(2536)^- \rightarrow \bar{D}^{*0}K^-)$ in each $D_s^+D_{s1}(2536)^-$ mass bin.
- [34] S. S. Wilks, *Ann. Math. Stat.* **9**, 60 (1938).
- [35] S. Godfrey and N. Isgur, *Phys. Rev. D* **32**, 189 (1985).
- [36] E. A. Kuraev and V. S. Fadin, *Yad. Fiz.* **41**, 733 (1985) [*Sov. J. Nucl. Phys.* **41**, 466 (1985)].
- [37] X. Y. Gao, C. P. Shen, and C. Z. Yuan, *Phys. Rev. D* **95**, 092007 (2017).
- [38] The dressed cross section is $\sigma = \sigma^B/|1 - \Pi|^2$, where σ^B is the Born cross section, and $|1 - \Pi|^2 = 0.94$ is the vacuum polarization factor [39].
- [39] S. Actis *et al.*, *Eur. Phys. J. C* **66**, 585 (2010).
- [40] N. Dash *et al.* (Belle Collaboration), *Phys. Rev. Lett.* **119**, 171801 (2017).