## ON SOME 3-DIMENSIONAL COMPLETE RIEMANNIAN MANIFOLDS SATISFYING $R(X, Y) \cdot R = 0$

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1. Introduction. Let (M, g) be a Riemannian manifold. By R we denote the Riemannian curvature tensor. By  $T_x(M)$  and  $\operatorname{Exp}_x$  we denote the tangent space to M at x and the exponential mapping of (M, g) at x. For  $X, Y \in T_x(M), R(X, Y)$  operates on the tensor algebra as a derivation at each point  $x \in M$ . In a locally symmetric space  $(\nabla R = 0)$ , we have

(\*)  $R(X, Y) \cdot R = 0$  for any point  $x \in M$  and X,  $Y \in T_x(M)$ .

We consider the converse under some additional conditions.

THEOREM A (S. Tanno [8]). Let (M, g) be a complete and irreducible 3-dimensional Riemannian manifold. If (M, g) satisfies (\*) and the scalar curvature S is positive and bounded away from 0 on M, then (M, g) is of constant curvature.

Other results concerning this problem may be found in references. In this paper, we shall prove

THEOREM B. Let (M, g) be a complete and irreducible 3-dimensional Riemannian manifold satisfying (\*). If the volume of (M, g) is finite, then (M, g) is of constant curvature, and hence,  $\nabla R = 0$ .

COROLLARY B. Let (M, g) be a compact and irreducible 3-dimensional Riemannian manifold satisfying (\*). Then (M, g) is of constant curvature.

It may be noticed that (\*) implies in particular

$$(**) R(X, Y) \cdot R_1 = 0,$$

where  $R_1$  denotes the Ricci tensor of (M, g).

In this paper, (M, g) is assumed to be connected, complete and of class  $C^{\infty}$  unless otherwise specified.

2. Preliminaries. Let (M, g) be a 3-dimensional Riemannian manifold. Assume (\*). dim M = 3 implies

$$(2.1) R(X, Y) = R^{\scriptscriptstyle 1} X \wedge Y + X \wedge R^{\scriptscriptstyle 1} Y - (S/2) X \wedge Y,$$

where

$$g(R^{i}X, Y) = R_{i}(X, Y)$$
 and  $(X \wedge Y)Z = g(Y, Z)X - g(X, Z)Y$ .

Let  $(K_1, K_2, K_3)$  be eigenvalues of the Ricci transformation  $R^1$  at a point x. Then (\*) is equivalent to

$$(2.2) (K_i - K_j)(2(K_i + K_j) - S) = 0.$$

Therefore we may have only three cases:

(K, K, K), (K, K, 0), (0, 0, 0) at each point.

First, if (K, K, K),  $K \neq 0$ , holds at some point x, then it holds on some open neighborhood U of x. Hence U is an Einstein space, and Kis constant on U and on M. Therefore (M, g) is of constant curvature (cf. H. Takagi and K. Sekigawa [6]). From now, we assume that rank  $R^1 \leq 2$  on M. Let  $W = \{x \in M; \text{rank } R^1 = 2 \text{ at } x\}$ . By  $W_0$  we denote one component of W. On  $W_0$ , we have two  $C^{\infty}$ -distributions  $T_1$  and  $T_0$  such that

$$T_1 = \{X; R^1X = KX\},$$
  
 $T_0 = \{Z; R^1Z = 0\}.$   
For X,  $Y \in T_1$  and  $Z \in T_0$ , by (2.1), we have  
(2.3)  $R(X, Y) = KX \wedge Y,$   
 $R(X, Z) = 0.$ 

This shows that  $T_0$  is the nullity distribution. Since the index of nullity at each point of M is 1 or 3, the nullity index of (M, g) is 1. Thus integral curves of  $T_0$  are geodesics (and complete if (M, g) is complete) (cf. Y. H. Clifton and R. Maltz [2], etc.). Let  $(E_1, E_2, E_3) = (E)$  be a local field of orthonormal frame such that  $E_3 \in T_0$  (consequently,  $E_1, E_2 \in T_1$ ) and

$$abla_{\scriptscriptstyle E_i} E_i = 0$$
 ,  $i = 1, 2, 3$  .

We call this (E) an adapted frame field. If we put  $\nabla_{E_i} E_j = \sum_{k=1}^3 B_{ijk} E_k$ , then we get  $B_{ijk} = -B_{ikj}$  and

$$(2.4) B_{3\,ij} = 0, i, j = 1, 2, 3.$$

The second Bianchi identity and (2.3) give

(2.5) 
$$E_3K + K(B_{1\,31} + B_{2\,32}) = 0$$
, or  
div  $E_3 = -E_3K/K$ .

By (2.4) and  $R(E_i, E_3)E_3 = \nabla_{E_i}\nabla_{E_3}E_3 - \nabla_{E_3}\nabla_{E_i}E_3 - \nabla_{[E_i, E_3]}E_3 = 0$ , we get

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$$(2.6) E_3 B_{131} + (B_{131})^2 + B_{132} B_{231} = 0,$$
  

$$E_3 B_{132} + B_{131} B_{132} + B_{132} B_{232} = 0,$$
  

$$E_3 B_{231} + B_{231} B_{131} + B_{232} B_{231} = 0,$$
  

$$E_3 B_{232} + (B_{232})^2 + B_{231} B_{132} = 0.$$

(2.5) and  $(2.6)_2$ , (2.5) and  $(2.6)_3$ , (2.5) and  $(2.6)_{1.4}$  imply

$$(2.7) B_{1\,32} = C_1(E)K, B_{2\,31} = C_2(E)K,$$

$$(2.8) B_{131} - B_{232} = D(E)K,$$

where  $C_1(E)$ ,  $C_2(E)$  and D(E) are functions defined on the same domain as (E) such that  $E_3C_1(E) = E_3C_2(E) = E_3D(E) = 0$ .

By (2.5) and (2.8), we get

(2.9) 
$$2B_{131} = D(E)K - E_{3}K/K.$$

Now, let  $\gamma_x(s)$  be an integral curve of  $T_0$  through  $x = \gamma_x(0) \in W_0$  with arclength parameter s, i.e.,  $\gamma_x(s) = \operatorname{Exp}_x s(E_3)_x$ . Then (2.6)<sub>1</sub>, (2.7) and (2.9) give

(2.10) 
$$\frac{1}{2} \frac{d}{ds} \left( \frac{1}{K} \frac{dK}{ds} \right) = HK^2 + \frac{1}{4} \left( \frac{1}{K} \frac{dK}{ds} \right)^2, \text{ along } \gamma_x(s),$$

where

$$H = H(E) = D(E)^2/4 + C_1(E)C_2(E)$$
.

(2.10) implies that H is independent of the choice of the adapted frame fields (E). Solving (2.10), we get

$$(2.11) K = \gamma , (for H = 0) , or$$

(2.12) 
$$K = \pm 1/((\alpha s - \beta)^2 - H/\alpha^2)$$
, (for  $H \neq 0$ ),

where  $\alpha$ ,  $\beta$  and  $\gamma$  are constant along  $\gamma_{x}(s)$ ,  $\alpha \neq 0$ .

With respect to our arguments, without loss of essentiality, we may assume that M is orientable. Let (E) be any adapted frame field which is compatible with the orientation. We call it an oriented adapted frame field. Then we see that  $f = (C_1(E) - C_2(E))K$  is independent of the choice of oriented adapted frame fields, and hence f is a function of class  $C^{\infty}$ on  $W_0$ . f = 0 holds on an open set  $U \subset W_0$ , if and only if  $T_1$  is integrable on U. This is a geometric meaning of f. In the sequel, we assume that the volume of (M, g) is finite. We can see that  $H = H(E) = D(E)^2/4 +$  $C_1(E)C_2(E)$  is a function of class  $C^{\infty}$  on  $W_0$ . Let  $W(H) = \{x \in W_0; H \neq 0$ at  $x\}$ . We assume that  $W(H) \neq \emptyset$ . Let  $W(H)_0$  be one component of W(H). By (2.12) and completeness of (M, g), H must be negative on  $W(H)_0$ . For each point  $x \in W(H)_0$ , consider  $\gamma_x(s)$ . Then  $\gamma_x(s) \in W(H)_0$ ,

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for all s. Let  $x_0 = \gamma_x(\beta/\alpha)$ . For  $(E_1)_{x_0}$ ,  $(E_2)_{x_0} \in T_1(x_0)$ , there exists a 2dimensional submanifold,  $\{\varphi(u_1, u_2) \in W(H)_0; (u_1, u_2) \in (-\varepsilon, \varepsilon)^2, \varepsilon > 0\}$ , such that  $\varphi(0, 0) = x_0$  and  $(\partial \varphi/\partial u_1)(0, 0) = (E_1)_{x_0}, (\partial \varphi/\partial u_2)(0, 0) = (E_2)_{x_0}$ . Now, we define a mapping

$$\Phi: (-\varepsilon, \varepsilon)^2 \times (-\delta, \delta) \rightarrow W(H)_0$$
 by

(2.13)  $\Phi(u_1, u_2, w_3) = \operatorname{Exp}_{\varphi(u_1, u_2)} w_3 E_3$ , for some  $\delta > 0$ .

Then  $\Phi$  is of class  $C^{\infty}$  and furthermore, for small  $\varepsilon$ ,  $\delta$ ,  $V(\varepsilon, \delta) = {\Phi(u_1, u_2, w_3) \in W(H)_0; (u_1, u_2, w_3) \in (-\varepsilon, \varepsilon)^2 \times (-\delta, \delta)}$  is a local coordinate neighborhood with origin at  $x_0$ . In  $V(\varepsilon, \delta)$ , by (2.12), we get

$$(2.14) K = \pm 1/((Aw_3 - B)^2 - H/A^2) ,$$

where A and B are functions of class  $C^{\infty}$  on  $V(\varepsilon, \delta)$  such that  $\partial A/\partial w_3 = \partial B/\partial w_3 = 0$  on  $V(\varepsilon, \delta)$  and  $A = \alpha$ , B = 0 at  $x_0$ .

By continuity of A and B in (2.14), there is  $\varepsilon_0$ ,  $0 < \varepsilon_0 < \varepsilon$  such that  $-\delta/4 < B/A < \delta/4$ , for  $(u_1, u_2) \in (-\varepsilon_0, \varepsilon_0)^2$ .

Now, we define a mapping  $\psi: (-\varepsilon_0, \varepsilon_0)^2 \to V(\varepsilon, \delta)$  by

$$(2.15) \qquad \qquad \psi(u_1, u_2) = \operatorname{Exp}_{\varphi(u_1, u_2)}(B(u_1, u_2)/A(u_1, u_2))E_3.$$

And furthermore, we define a mapping  $\Psi: (-\varepsilon_0, \varepsilon_0)^2 \times (-\delta_0, \delta_0) \rightarrow V(\varepsilon, \delta)$ by

(2.16) 
$$\Psi(u_1, u_2, u_3) = \operatorname{Exp}_{\psi(u_1, u_2)} u_3 E_3$$
,  $\delta_0 = \delta/4$ .

Then  $\Psi$  is of class  $C^{\infty}$  and

$$U(\varepsilon_0, \ \delta_0) = \{ \Psi(u_1, \ u_2, \ u_3) \in \ V(\varepsilon, \ \delta); \ (u_1, \ u_2, \ u_3) \in (-\varepsilon_0, \ \varepsilon_0)^2 \times \ (-\delta_0, \ \delta_0) \}$$

is a local coordinate neighborhood with origin at  $x_0$ .

Between  $w_3$  in  $V(\varepsilon, \delta)$  and  $u_3$  in  $U(\varepsilon_0, \delta_0)$ , the following relation holds:

$$(2.17)$$
  $w_{\scriptscriptstyle 3} = u_{\scriptscriptstyle 3} + B/A$  , in  $U(arepsilon_{\scriptscriptstyle 0},\,\delta_{\scriptscriptstyle 0})$  .

Thus (2.14) and (2.17) imply

(2.18) 
$$K = \pm 1/((Au_3)^2 - H/A^2)$$
, on  $U(\varepsilon_0, \delta_0)$ .

Let  $\gamma(u_1, u_2)$  be the integral curve of  $T_0$  starting from  $\psi(u_1, u_2)$ ,  $(u_1, u_2) \in (-\varepsilon_0, \varepsilon_0)^2$ , i.e.,  $\gamma(u_1, u_2)(s) = \operatorname{Exp}_{\psi(u_1, u_2)} sE_3$ . Then, in  $U(\varepsilon_0, \delta_0)$ ,  $u_3$  can be considered as the arc-length parameter of  $\gamma(u_1, u_2)$ . We put  $L(u_1, u_2) = \{\gamma(u_1, u_2)(s) \in M; -\infty < s < \infty\}$ . Since dim  $T_0 = 1$ , taking account of (2.12) and (2.18), we can see that  $\gamma(u_1, u_2)(s_1) \neq \gamma(u_1, u_2)(s_2)$  for  $s_1 \neq s_2$ . From (2.12) and (2.18), dK/ds = 0 for s = 0 and otherwise  $dK/ds \neq 0$  along  $L(u_1, u_2)$ , for any  $(u_1, u_2) \in (-\varepsilon_0, \varepsilon_0)^2$ . Thus, we can see that if  $(u_1, u_2) \neq (v_1, v_2)$ ,  $(u_1, u_2)$ ,  $(v_1, v_2) \in (-\varepsilon_0, \varepsilon_0)^2$ , then  $L(u_1, u_2) \cap L(v_1, v_2) = \emptyset$ . Now, we put

 $U(arepsilon_0)=\{\widehat{\varPsi}(u_1,\,u_2,\,u_3)\in M;\,(u_1,\,u_2)\in (-arepsilon_0,\,arepsilon_0)^2,\,-\infty\,< u_3<\infty\}$  ,

where  $\hat{\varPsi}$  denotes an extension of  $\varPsi$  defined by

 $\hat{\mathscr{\Psi}}(u_1, u_2, u_3) = \operatorname{Exp}_{\psi(u_1, u_2)} u_3 E_3$ , on  $(-\varepsilon_0, \varepsilon_0)^2 \times (-\infty, \infty)$ .

Then, from the above arguments, we have the following

LEMMA 2.1.  $U(\varepsilon_0)$  is a local coordinate neighborhood with origin at  $x_0$ .

For any G > 0, we put

$$V_{G} = \{\widehat{\mathscr{V}}(u_{1}, u_{2}, u_{3}) \in U(\varepsilon_{0}); (u_{1}, u_{2}) \in (-\varepsilon_{0}/2, \varepsilon_{0}/2)^{2}, 0 < u_{3} < G\}$$
.

Then  $\overline{V}_{g} \subset U(\varepsilon_{0})$ . Let vol (M, g) and vol  $(V_{g})$  denote the volumes of (M, g) and the open subspace  $V_{g}$  of (M, g), respectively. Then, by the assumption, we have

(2.19) 
$$\operatorname{vol}(V_g) < \operatorname{vol}(M, g) < \infty$$
, for any  $G > 0$ .

On the other hand, since  $E_3 = \partial/\partial u_3$  on  $U(\varepsilon_0)$ , we have

$${
m div}\ E_{\scriptscriptstyle 3} = (1/\sqrt{g_{\scriptscriptstyle 0}}) (\partial \sqrt{g_{\scriptscriptstyle 0}}/\partial u_{\scriptscriptstyle 3})$$
 on  $U(arepsilon_{\scriptscriptstyle 0})$  ,

where

$$g_{_0}=\det{(g_{_{ij}})}$$
 ,  $g_{_{ij}}=g(\partial/\partial u_i,\,\partial/\partial u_j)$  ,  $i,\,j=1,\,2,\,3$  .

Thus, by (2.5), we get

$$(2.20) \qquad (1/\sqrt{g_0})(\partial\sqrt{g_0}/\partial u_3) + (1/K)(\partial K/\partial u_3) = 0 \quad \text{on} \quad U(\varepsilon_0) .$$

Solving (2.20), we get

$$(2.21) \sqrt{g_{\scriptscriptstyle 0}} = C/K \,,$$

where  $C = C(u_1, u_2)$  is a function of class  $C^{\infty}$  on  $U(\varepsilon_0)$ . Thus, from (2.18) and (2.21), we get

$$ext{vol}\left(V_{G}
ight) = \int_{r_{G}} dM = \int_{-\epsilon_{0}/2}^{\epsilon_{0}/2} \int_{-\epsilon_{0}/2}^{\sigma} \int_{0}^{G} (C/K) du_{1} du_{2} du_{3}$$
  
 $\geq a(\epsilon_{0})^{2}G, \quad ext{for any} \quad G > 0,$ 

where

 $a = \min_{\substack{-\varepsilon_0/2 \leq u_1, u_2 \leq \varepsilon_0/2 \\ u_3 = 0}} C/K > 0.$ 

But, this contradicts (2.19). Thus we have the following

LEMMA 2.2. If vol (M, g) is finite, then, for each point  $x \in W$ , S = 2K is constant along  $\gamma_x(s)$ ,  $-\infty < s < \infty$ . 3. Proof of Theorem B. In the sequel, we shall assume that  $\operatorname{vol}(M, g)$  is finite and rank  $R^1$  is at most 2 on M and rank  $R^1 = 2$  at some point of M. From Lemma 2.2, H = 0 on  $W_0$ . Let  $V = \{x \in W_0; f(x) \neq 0\}$ . Now, we assume that  $V \neq \emptyset$ . Let  $V_0$  be one component of V. H = H(E) = 0 implies  $D(E)^2 = -4C_1(E)C_2(E)$ . Put  $\cos 2\theta(E) = K(C_1(E) + C_2(E))/f$  and  $\sin 2\theta(E) = KD(E)/f$ . Define  $(E^*)$  by  $E_3^* = E_3$  and

$$egin{array}{lll} E_{_1}^{\,*} &= \cos \, heta(E) E_{_1} - \sin \, heta(E) E_{_2} \;, \ E_{_2}^{\,*} &= \sin \, heta(E) E_{_1} + \cos \, heta(E) E_{_2} \;. \end{array}$$

Then we have  $D(E^*) = 0$ . Furthermore, for (E) and (E'), we have  $E_1^*(E) = \pm E_1^*(E')$  and  $E_2^*(E) = \pm E_2^*(E')$ . H = 0 and  $D(E^*) = 0$  imply  $C_1(E^*)C_2(E^*) = 0$ . So we can assume that  $C_2(E^*) = 0$  (otherwise, change  $(E_1^*, E_2^*, E_3^*) \rightarrow (E_2^*, -E_1^*, E_3^*)$ ). Then we get

$$(3.1) B_{1\ 32}^{*} \neq 0 , B_{1\ 31}^{*} = B_{2\ 31}^{*} = B_{2\ 32}^{*} = 0 ,$$

where

$$abla_{E_{i}} * E_{j}^{\, st} = \sum\limits_{k=1}^{3} B_{i\,\, j\, k}^{\, st} \, E_{k}^{\, st} \; .$$

 $R(E_1^*, E_3^*)E_2^* = 0$  implies

 $(3.2) E_3^* B_{1_{21}}^* = 0.$ 

 $R(E_1^*, E_2^*)E_3^* = 0$  implies  $B_{221}^* = 0$  and

$$(3.3) E_2^* B_{1\ 32}^* + B_{1\ 21}^* B_{1\ 32}^* = 0$$

 $R(E_1^*, E_2^*)E_1^* = -KE_2^*$  implies

$$(3.4) E_2^* B_{1\,21}^* + (B_{1\,21}^*)^2 = -K \, .$$

By  $B_{2ij}^* = 0$ , each trajectory of  $E_2^*$  is a geodesic. Put  $h = B_{121}^*$  and  $F = (E_1^* f)^2$ . Then F is a function of class  $C^{\infty}$  on  $V_0$ . From Lemma 2.2, and (3.3), we get

(3.5)  
$$E_{3}^{*}(E_{1}^{*}f) = E_{1}^{*}(E_{3}^{*}f) + [E_{3}^{*}, E_{1}^{*}]f$$
$$= -B_{1}^{*}{}_{32}(E_{2}^{*}f) = f^{2}h, \quad \text{i.e.},$$
$$d(E_{1}^{*}f)/ds = f^{2}h, \quad \text{along} \quad \gamma_{x}(s), \quad x \in V_{0}.$$

From Lemma 2.2, for each point  $x \in V_0$ ,  $\gamma_x(s) \in V_0$ ,  $-\infty < s < \infty$ . Taking account of (3.2) and solving (3.5), we get

$$(3.6) \quad \cdot \qquad F = (f(x)^2 h(x)s + c)^2 , \quad \text{along} \quad \gamma_x(s) , \quad -\infty < s < \infty ,$$

where c is constant along  $\gamma_x(s)$ .

Let  $V^* = \{x \in V_0; h(x) \neq 0\}$ . From (3.4), we see that  $V^* \neq \emptyset$ . Let  $V_0^*$  be one component of  $V^*$ . Then, by (3.2), we see that, for each point  $x \in V_0^*$ ,  $\gamma_x(s) \in V_0^*$ ,  $-\infty < s < \infty$ . For each point  $x \in V_0^*$ , consider  $\gamma_x(s)$ . Let  $x_0 = \gamma_x(-c/f(x)^2h(x))$  in (3.6).

Then we have

$$(3.7) extsf{F} = ((f^2h)w_3 + k)^2 extsf{, on } V(arepsilon, \delta) \cap V_0^* extsf{,}$$

where  $k = k(u_1, u_2)$  is a function of class  $C^{\infty}$  on  $V(\varepsilon, \delta) \cap V_0^*$  such that k(0, 0) = 0, and  $V(\varepsilon, \delta)$  is a local coordinate neighborhood with origin at  $x_0$  constructed by the similar fashion as in §2. From (3.7), by applying the similar arguments as in the proof of Lemma 2.1, to the function F instead of K, we can construct a local coordinate neighborhood

$$U(arepsilon^*) = \{ arPsilon^*(u_1,\, u_2,\, u_3) \in \, V_0^*;\, (u_1,\, u_2) \in (-arepsilon^*,\, arepsilon^*)^2,\, -\infty \, < \, u_3 < \, \infty \}$$
 ,

with origin at  $x_0$  such that  $F = ((f^2h)u_3)^2$  on  $U(\varepsilon^*)$ , where  $\varepsilon^* > 0$ , and  $\Psi^*$ is a mapping of class  $C^{\infty}$  defined by the similar way as  $\hat{\Psi}$  in § 2. For any G>0, let  $V_G^* = \{\Psi^*(u_1, u_2, u_3) \in U(\varepsilon^*); (u_1, u_2) \in (-\varepsilon^*/2, \varepsilon^*/2)^2, 0 < u_3 < G\}$ . From Lemma 2.2, and (2.5), we have div  $E_3^* = 0$ . Thus, we can see that if  $G \to \infty$ , then vol  $(V_G^*) \to \infty$ . But, this is a contradiction. Thus, we can conclude that f = 0 on  $W_0$  and hence  $T_1$  is integrable on  $W_0$ . Thus,  $T_1$  and  $T_0$  are parallel on  $W_0(cf.$  S. Tanno [7]). If W is dense in M, the restricted homogeneous holonomy group of (M, g) is reducible. If W is not dense in M, then the interior of the complement of W in M is flat. Hence, also in this case, the restricted homogeneous holonomy group of (M, g) is reducible. Lastly, if rank  $R^1 = 0$  on M, then (M, g) is flat. Therefore, this completes a proof of Theorem B.

From our arguments in this paper, we can also show the following

THEOREM C. Let (M, g) be a complete and simply connected 3-dimensional Riemannian manifold satisfying (\*). If the volume of (M, g) is finite, then (M, g) is isometric to a 3-dimensional sphere.

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