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Pattern formation in a suspension of swimming microorganisms: equations and stability theory

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A model for collective movement and pattern formation in layered suspensions of negatively geotactic micro-organisms is presented. The motility of the organism is described by an average upward swimming speed U and a diffusivity tensor **D**. It is shown that the equilibrium suspension is unstable to infinitesimal perturbations when either the layer depth or the mean concentration of the organisms exceeds a critical value. For deep layers the maximum growth rate determines a preferred pattern size explicitly in terms of U and **D**. The results are compared with observations of patterns formed by the ciliated protozoan *Tetrahymena pyriformis*.

1. Introduction

From time to time reports have appeared in the biological literature of observations of streaming patterns in liquid suspensions of swimming microorganisms. The phenomenon involves fluid dynamics in that the experiments strongly suggest that the visible patterns of high concentrations of the organisms, as well as the associated motion of the suspending fluid, arise from a process of 'bioconvection' (to use the term coined by Platt 1961), wherein natural dissipative losses (presumably due mainly to the viscosity of the fluid) are compensated for by the work done on the fluid by the micro-organisms. Robbins (1952), observing *Euglena gracilis*, and Loeffer & Mefferd (1952), observing cultures of the ciliated **pr**otozoan *Tetrahymena pyriformis*, found that patterns formed when the depth of the suspension or the organism number density exceeded critical values.[‡] (Figures 1 and 3*a*, plates 1 and 3, show the critical-depth effect.) The type of pattern apparently depends upon the suspension depth and upon the concentration and motility of the organisms. Wille & Ehret (1968), working with

[†] The critical values are given as 2 mm and 150 000 organisms/cm³. However, it is not clear whether or not these figures are intended to apply simultaneously at a critical state.

dense Tetrahymena cultures, observed two distinct steady-state patterns. The polka-dot pattern, in which spherical concentrations of the organism are arranged in a regular array, was formed in relatively shallow cultures (see figures 2 and 3b, plates 2 and 3). A reticulate pattern emerged in deeper cultures as at the deep end of figure 1. The latter is usually described as consisting of irregular cells ('regular cell' being used here to mean a unit repeating periodically along the layer of the suspension), with the organisms swimming predominantly upwards over the interior of the cell, being carried laterally near the top and bottom of the layer, and plummeting downwards in thin vertical columns or sheets which mark the main visible features of the pattern. The concentration within these descending regions can exceed the mean concentration by a factor of 10 or more, and the fluid speed there is typically of the order of 1 mm/s. This speed appreciably exceeds the mean vertical swimming speed of the organisms (typically 0.5 mm/s in *Tetrahymena*); thus the organisms in the sheets or columns are swept to the bottom of the layer, eventually to swim again to the top and repeat the process. The pattern formation time is typically 10-30 s, which is also roughly the cycle time of individual organisms within the pattern, and increases as the motility (in the present context, the average swimming speed) of the organisms decreases.

The dynamical explanation of bioconvection that has emerged in recent years (Platt 1961; Winet 1969; Winet & Jahn 1972; Plesset & Winet 1974) lies in the unstable stratification of the organism suspensions. The stratification is caused by the accumulation of organisms near the top of the layer and is a result of the swimming of organisms preferentially in the upward direction. We adopt the biological term and refer to this phenomenon as 'negative geotaxis'. The occurrence of negative geotaxis is now well established for *Tetrahymena*. Since the individual organisms are slightly denser than the ambient fluid, the process can produce an unstable subsurface layer of heavy material.

Plesset & Winet (1974; see also Plesset & Whipple 1974) liken the onset of bioconvection to Rayleigh-Taylor instability. They model the organism-rich sublayer as a layer of dense fluid overlying a less dense, deeper layer of fluid. They study the stability of this system under the influence of viscosity and find that the wavelength of the most rapidly growing mode agrees well with the observed scale of the reticular pattern. This result provides evidence that the onset of bioconvection is indeed due to a density inversion and that growing sedimentation fingers represent the formation of descending columns of organisms.

The model of Plesset & Winet does not directly consider the negative geotaxis of the organisms, nor even their motility, except implicitly as a means by which the unstable equilibrium could be set up. Although it is possible that the negative geotaxis might be inhibited by the formation of the subsurface layer, it is more likely that the mechanism which sets up the layering continues to act as the instability develops and influences the formation of a new stable equilibrium (should any exist), just as temperature differences remain the driving mechanism in nonlinear thermal convection. If the Plesset–Winet model were followed into the nonlinear range, the only possible new equilibrium would be one in which all or a major fraction of the subsurface fluid falls to the bottom, and which has a lighter and therefore stable subsurface layer.



FIGURE 1. Tilted layer of *Tetrchymena pyriformis*. The shallow end is at the top and the pattern ranges from polka-dot to reticulate as the depth increases. The diameter of the dish is 8 cm. Note also that the bottom of the dish in this and other figures is slightly raised in the centre.

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(Facing p. 592)



FIGURE 2. Pattern formation in a 3 mm layer of *C. cohnii* at a mean concentration of 10^6 organisms/cm³. Times: (a) 2 s, (b) 13 s, (c) 19 s, (d) 23 s, (e) 28 s, (f) 32 s, (g) 36 s, (h) 43 s, (i) 2 min. The suspension was swirled initially to render it homogeneous and the early patterns are forming during the decay of the swirl.

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(a)



(b)

FIGURE 3. Patterns in layers of C. cohnii. (a) Showing effect of slight variation in depth with shallow end at upper right. Filamentary structures are in the process of breaking up into clumps. (b) Close-up of a polka-dot pattern with some tori. Blurring towards edges is optical distortion.

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The purpose of the present paper is to develop a simple model for pattern formation by negatively geotactic micro-organisms, based upon the physical picture provided by Winet & Jahn (1972), but differing from the model of Plesset & Winet in this last respect. A system of equations is proposed in which the layering may be said to be internally generated as an equilibrium solution. Linear and nonlinear stability of this equilibrium solution can then be studied in the usual way. Guided by an analogy with solutes, we replace the discrete organism distribution by a continuous density and incorporate the negative geotaxis as a vertical drift of the organism 'stuff' relative to the suspending fluid. We also add an anisotropic diffusion. The motility of the organisms is thus parametrized by a vertical drift U and a diffusivity tensor **D**, both generally functions of vertical position and the local concentration of the organisms. The production of momentum associated with the locomotion of the organisms is averaged, so that the effect of the heavy stuff on the fluid motion is through a negative buoyancy term.

The equilibrium solution in this model describes the vertical stratification induced by the negative geotaxis, and in the present paper we investigate the linear stability of this stratified layer. In §2.4 we treat a special case satisfying the atypical condition that the layer depth is small compared with the virtual thickness of the subsurface layer. Some general results concerning the linear stability problem are given in §3. In §4 we treat the case where the layer depth greatly exceeds the sublayer thickness. A singular perturbation analysis allows us to study the transition from an exact mathematical analogy with Bénard convection under the condition of fixed heat flux to a convective instability with growth rates like those calculated by Plesset & Winet (1974), as the ratio of layer to sublayer thickness is increased. For deep layers we also derive an approximate expression in terms of U and D for the horizontal wavenumber for which growth is most rapid. Analytical details which are included here for completeness but which are not essential to the understanding of results are given in appendices. A discussion of the instability in physical terms is given by Levandowsky et al. (1975). Nonlinear aspects of the model and the construction of steady patterns will be taken up in a separate paper.

2. The model 2.1. Formulation

We consider a horizontally infinite, plane layer of homogeneous fluid bounded by the surfaces z = -H, 0 and containing in suspension a large number of impermeable micro-organisms. We denote the density of the fluid by ρ and the mean density of an organism by ρ_0 , and we consider the case $\rho_0 > \rho$. If the fractional volume occupied by organisms is $c(\mathbf{r}, t)$, the suspension has density $\rho_0 c + (1-c)\rho = \rho(1+\alpha c)$, where $\alpha = \rho_0/\rho - 1$. For *Tetrahymena* Winet & Jahn (1972) give $\alpha = 0.09$ while the values of c in the cultures discussed above are $\sim 10^{-2}$. This means that density fluctuations in the suspension are small and we describe its motion by the Boussinesq equations

$$\rho \, d\mathbf{u}/dt + \nabla p - \mu \nabla^2 \mathbf{u} = -g\rho(1 + \alpha c) \,\mathbf{k},\tag{2.1}$$

$$\nabla \cdot \mathbf{u} = 0, \qquad (2.2)$$

where **u** is the suspension velocity, p is the pressure, $g\mathbf{k}$ is the acceleration due to gravity and where in view of the smallness of c we have taken the viscosity μ to be constant.

To describe the evolution of c we write

$$\frac{dc}{dt} + \nabla \cdot \mathbf{J} = 0, \tag{2.3}$$

where \mathbf{J} is the flux of organisms through the fluid. We suppose that this flux consists of a part due to random motions, which is describable by diffusion, and a second part due to the negatively geotactic drift of organisms. Thus we write

$$\mathbf{J} = cU(c, z) \,\mathbf{k} - \mathbf{D} \cdot \nabla c,$$

$$\mathbf{D} = \kappa_1(c, z) \left(\mathbf{i} \cdot \mathbf{i} + \mathbf{j} \cdot \mathbf{j}\right) + \kappa(c, z) \,\mathbf{k} \cdot \mathbf{k},$$
 (2.4)

where $(\mathbf{i}, \mathbf{j}, \mathbf{k})$ are orthogonal unit vectors and U, κ and κ_1 are functions to be specified.

The boundary conditions to be adopted will depend upon the nature of the bounding planes. In all cases we require (with $\mathbf{u} = (u, v, w)$ and $\mathbf{J} = (J_1, J_2, J_3)$ and suppressing all independent variables but z)

$$w(0) = w(-H) = 0, \quad J_3(0) = J_3(-H) = 0,$$
 (2.5*a*, *b*)

which state that the vertical fluxes of fluid mass and of organisms vanish at both boundaries. We shall take the suspension boundary to coincide with the plane even when the boundary is free, and consider the cases ff (both free), fr (top free, lower rigid) and rr (both rigid), with conditions

$$\partial^2 w / \partial z^2 = 0$$
 on a free boundary, (2.5c)

$$\partial w/\partial z = 0$$
 on a rigid boundary. (2.5*d*)

A final special case will be that of an infinitely deep fluid; there we replace the lower conditions by the requirements that **u** and c vanish as $z \to -\infty$.

The central hypothesis of the model, that a continuous function $c(\mathbf{r}, t)$ may replace a complex distribution of self-propelled particles, is certainly a crude simplification of the phenomenon. In the densest parts of the pattern we are dealing with interparticle distances of the order of 10^{-2} to 10^{-3} cm, yet we wish to resolve pattern structure on the scale of the sublayer thickness, which in typical experiments is about 1 mm. It therefore seems likely that the averaging envisaged here is over only 10–100 organisms, and that the vertical resolution in determining sublayer distributions (see §2.2) will be a fraction of the sublayer thickness. Another difficulty is that the function $\mathbf{D}(c, z)$ is unknown, and for the most part we take \mathbf{D} to be constant. There are certainly substantial errors in such a description of the random component of the organisms' motion. Fortunately in the linear stability theory (§§3 and 4) we find that certain important results (e.g. pattern size) are insensitive to the sublayer concentration profile; also, it seems likely that simple diffusion gives a reasonably accurate description over the main body of the layer.

2.2. The equilibrium solution and definition of dimensionless variables

It is natural to take the basic pattern-free equilibrium to be a solution of (2.1)-(2.5) which is independent of x, y and t and has $\mathbf{u} = 0$. For such a solution (2.1) determines p for given c and (2.3) and (2.4) integrate (with condition (2.5b)) to give eU(a, z) = x(a, z) daldz = 0 (2.6)

$$cU(c,z) - \kappa(c,z) dc/dz = 0.$$

$$(2.6)$$

In this paper we shall refer to two special cases:

- (I) $U = U_0$, $\kappa = \kappa_0$, $\kappa_1 = \delta \kappa_0$; U_0 , κ_0 , δ = constants;
- (II) κ/U not explicitly dependent on z; κ_1 arbitrary.

Evidently, case II contains case I; we shall frequently specialize to case I for specific computations. In both of these cases we may integrate (2.6) in the form

$$z = \int_{K(0)}^{K} \frac{\kappa}{cU} dc$$
, $K(z)$ = equilibrium concentration profile,

which is an implicit definition of K as a function of z. Using the non-negativity of the integrand we may invert to obtain c = K(z) explicitly, where K is monotone increasing. The positive constant K(0) is arbitrary and is equal to the maximum concentration in the sublayer. Another quantity of interest is the mean concentration c_m . We define

$$c_0 = K(0), \quad c_m = \frac{1}{H} \int_{-H}^0 K dz.$$
 (2.7)

For example, in case I we have

$$K = c_0 \exp\left(U_0 z/\kappa_0\right), \quad c_0 = \frac{\lambda e^{\lambda}}{e^{\lambda} - 1} c_m, \quad \lambda = \frac{U_0 H}{\kappa_0}.$$
 (2.8)

To illustrate case II, consider the family of profiles generated by the choice

$$\kappa/U = z_0(k+1) (c/c_0)^k, \quad z_0 = \text{constant},$$
 (2.9)

where k is a positive number. We have

$$K(z) = \begin{cases} c_0 \left[\frac{k}{(k+1)} \frac{z}{z_0} + 1 \right]^{1/k}, & -z_0(k+1)/k \leq z \leq 0, \\ 0, & z < -z_0(k+1)/k, \\ c_m = c_0 z_0/H & \text{if } H > z_0(k+1)/k. \end{cases}$$
(2.10)

This family includes the exponential profile (2.8) as the limit for small k and a rectangular profile as the limit for large k, and describes tolerably well the observed profiles.

The equilibrium profiles provide convenient length scales for non-dimensionalization. Evidently, κ/U is a local scale height and its surface value is

$$h = \kappa_0 / U_0, \qquad (2.11 a)$$

where κ_0 and U_0 are the values of κ and U when z = 0 and $c = c_0$. An important dimensionless parameter is the ratio of H to h [cf. (2.8)], which we shall denote by

$$\lambda = H/h. \tag{2.11b}$$

However, in case II, where the profile may have a complicated form, it is sometimes useful to define an equivalent sublayer thickness h_e by

$$h_e = (c_m/c_0) H \quad (h_e \leq H),$$
 (2.12*a*)

and to define a corresponding ratio

$$\lambda_e = H/h_e \quad (\lambda_e \ge 1). \tag{2.12b}$$

If the layer is sufficiently deep, so that K vanishes on the lower boundary or is at least very small there, h_e will be equal to the thickness of a sublayer of constant concentration c_0 having the same total number of organisms per unit horizontal area as the entire layer.

For the illustration (2.9) and (2.10) of case II we see that $h_e = z_0$ if

$$H > z_0(k+1)/k.$$

The occasional advantage of (2.12) over (2.11) as a unit of length will be apparent in §4.4. We refer to a layer having small λ (equivalently λ_e close to 1) as shallow, and one for which $\lambda_e \simeq \lambda \gg 1$ as deep.

If U_0 is taken as a characteristic speed in the problem, a set of dimensionless variables appropriate to the analysis of the stability of the equilibrium K(z) can be obtained by introducing h as the unit of length:

$$r^{*} = h^{-1}r, \quad t^{*} = U_{0}h^{-1}t, \quad \mathbf{u}^{*} = U_{0}^{-1}\mathbf{u}, \quad p^{*} = (h/\mu U_{0})(p + \rho gz), \\ c^{*} = c_{0}^{-1}c, \quad \kappa^{*} = \kappa/\kappa_{0}, \quad \kappa_{1}^{*} = \kappa_{1}/\kappa_{0}, \quad U^{*} = U_{0}^{-1}U. \end{cases}$$
(2.13)

The scaling of c, which is already dimensionless, is included for convenience.

In the starred variables (2.13) the equations become

$$\sigma^{-1}d\mathbf{u}^*/dt^* + \nabla^* p^* - \nabla^{*2}\mathbf{u}^* = -Rc_*^*\mathbf{k}, \qquad (2.14a)$$

$$dc^*/dt^* + \nabla^* \cdot [c^*U^*\mathbf{k} - \mathbf{D}^* \cdot \nabla^* c^*] = 0$$
(2.14b)

and the flux condition is

$$c^*U^* - \kappa^* \partial c^* / \partial z^* = 0$$
 $(z^* = 0, -\lambda).$ (2.14c)

Here

$$\sigma = \nu/\kappa_0 \tag{2.15}$$

is a Schmidt number for vertical diffusion and

$$R = g\alpha c_0 h^3 / \nu \kappa_0 = g\alpha c_0 \kappa_0^2 / \nu U_0^3$$

$$\tag{2.16}$$

is a parameter which measures the magnitude of the Archimedean force. We shall refer to R as the 'Rayleigh number' although the parameter of the Bénard problem is more closely analogous to the parameter $\lambda^4 R$ [cf. (2.22)]. The experimental data discussed in §5 suggest that typical values of σ exceed 1 and that values of R are usually in the range 1–100.

2.3. Linear equations

We henceforth drop stars whenever the dimensionless variables (2.13) are used. The dynamical stability of the equilibrium will be studied in the usual way by linearizing (2.14) and separating out time and horizontal space variables. We set

 $\mathbf{596}$

 $c = K(z) + \phi(x, y, z, t)$, where ϕ is a perturbation, and eliminate the horizontal velocity components and pressure by applying the curl operator twice to the linearized momentum equation and taking the z component of the result. On substituting $(\phi, w) = e^{\gamma t} f(x, w) (\Phi(z), W(z)) = \nabla^2 f = -a^2 f$

$$(\phi, w) = e^{\gamma t} f(x, y) (\Phi(z), W(z)), \quad \nabla^2 f = -a^2 f,$$

we obtain the equations

$$(\gamma/\sigma) (D^2 - a^2) W - (D^2 - a^2)^2 W = a^2 R \Phi, \qquad (2.17)$$

$$\gamma \Phi + WDK + D\mathcal{F} + a^2 \kappa_1 \Phi = 0, \qquad (2.18)$$

where D = d/dz and

$$\mathscr{F} = \left[c \frac{\partial U}{\partial c} + U + \frac{\partial \kappa}{\partial z} \right]_{c=K} \Phi - D[\kappa(K, z) \Phi].$$

As is shown in appendix A, this last formula can be rewritten as

$$\mathscr{F} = \kappa(DK) D[\Phi/(DK)]. \tag{2.19}$$

The flux boundary conditions are

$$\mathscr{F}(0) = \mathscr{F}(-\lambda) = 0. \tag{2.20}$$

We show in appendix A that for cases I and II (see §2.2) γ is necessarily real for solutions of (2.17)–(2.20) with conditions from (2.5). The neutral-stability boundary will therefore be given by a function $R(\lambda, a)$ determined by the condition $\gamma = 0$. The critical values of R and a, denoted by R_c and a_c , are determined by minimizing R for fixed λ over all branches and all non-negative a.

2.4. The shallow-layer limit

It appears that bioconvection can only occur when a sublayer can in some sense be defined, so the limiting case considered in the present subsection is mainly of formal interest. The stability problem obtained for shallow layers may be used to illustrate the nature of the eigenvalue problem and the method of solution. To avoid inessential complications we consider only case I. If we rescale the (now dimensionless) vertical co-ordinate by λ , $z = \bar{z}\lambda$ (which in effect makes Hthe unit of length), the equation for K becomes

$$\overline{D}K = \lambda K, \quad \overline{D} = d/d\overline{z},$$
 (2.21)

and for small λ we have $K = 1 + \lambda \overline{z} + O(\lambda^2)$.

This switch in length scale from h to H suggests the change of variables

$$\lambda^2 \gamma = \overline{\gamma}, \quad \lambda^2 \overline{W} = W, \quad \lambda a = \overline{a},$$

in which case (2.17) indicates that the parameter

$$R_4 = \lambda^4 R = g \alpha c_0 U_0 H^4 / \nu \kappa_0^2 \tag{2.22}$$

will replace R. If we now combine (2.18) and (2.19) and allow λ to approach zero with barred quantities fixed, we obtain the following problem for shallow layers:

$$(\overline{\gamma}/\sigma)\,(\overline{D}{}^2 - \overline{a}{}^2)\,\overline{W} - (\overline{D}{}^2 - \overline{a}{}^2)^2\,\overline{W} = \overline{a}{}^2R_4\,\Phi,\tag{2.23}$$

$$\overline{\gamma}\Phi + \overline{W} - D^2\Phi + \overline{a}^2\delta\Phi = 0, \qquad (2.24)$$

$$\bar{D}\Phi = 0, \quad \bar{z} = -1, 0.$$
 (2.25)

If $\delta = 1$ equations (2.23)–(2.25) are identical to the linearized stability equations for a Boussinesq fluid at a Rayleigh number R_4 and Prandtl number σ , but with the boundary condition (2.25) on the heat flux replacing the conventional condition $\Phi = 0$ corresponding to fixed wall temperature. In this analogue $-\Phi$ replaces the temperature perturbation, κ_0 the thermal diffusivity, α the coefficient of thermal expansion and $c_0 U_0/\kappa_0$ the equilibrium temperature gradient.

Some calculations for this linear Bénard problem were included in a paper by Hurle, Jakeman & Pike (1967; see also Nield 1968), who showed that a treatment similar to the classical one is possible. As R_4 is increased, the mode which is first unstable is even in $\overline{z} + \frac{1}{2}$, and Hurle *et al.* found that R_{4c} is 120 for case *ff* and 720 for case *rr*, and that the critical wavenumber is zero in each case. Using estimates derived from two variational principles (Chandrasekhar 1961; see also §3) we find that case *fr* is similar, with $R_{4c} = 320$ and $\overline{a}_c = 0$.

The analysis of the onset of instability is thus simpler here in cases fr and rr than in the Bénard problem with conventional boundary conditions, since expansions with respect to \bar{a} may be used to study the critical point. To see how these expansions proceed, we first note that (2.24) may be integrated, using (2.25), to obtain $\int_{0}^{0} \int_{0}^{0} -$

$$\int_{-1}^{0} \Phi \, d\bar{z} (\bar{\gamma} + \delta \bar{a}^2) + \int_{-1}^{0} \overline{W} \, d\bar{z} = 0.$$
(2.26)

Since we are interested in solutions for small $\overline{\gamma}$ and \overline{a} , (2.26) suggests that both $\overline{\gamma}$ and \overline{W} should be taken to be $O(\overline{a}^2)$; (2.23) then suggests that R_4 is O(1). We thus try

$$(\overline{\gamma}, \overline{W}) = \sum_{n=0}^{\infty} \overline{a}^{2n+2}(\gamma_n, W_n), \quad \Phi = \sum_{n=0}^{\infty} \overline{a}^{2n} \Phi_n.$$
 (2.27)

Substitution gives a sequence of equations starting with

$$\overline{D}^2 \Phi_0 = 0 \qquad \qquad (2.28)$$

and continuing with systems of the form

$$\overline{D}^{4}W_{n} = -R_{4}\Phi_{n} + f_{n}(W_{n-1}, ..., W_{0}; \gamma_{n-1}, ..., \gamma_{0}), \qquad (2.29)$$

$$D^{2}\Phi_{n+1} = \gamma_{n}\Phi_{0} + g_{n}(W_{n}, ..., W_{0}; \Phi_{n-1}, ..., \Phi_{0}; \gamma_{n-1}, ..., \gamma_{0}), \qquad (2.30)$$

where $f_0 = 0$ and $g_0 = W_0 + \delta \Phi_0$. The boundary conditions follow from (2.5) and (2.25).

Because of the linearity and (2.25) we may take $\Phi_0 = 1$ and $\Phi_n(0) = 0$, n = 1, 2, ... The solutions are obtained at each stage by computing W_n , then γ_n and finally Φ_{n+1} . Since (2.30) with the boundary condition (2.25) is a self-adjoint problem, the solvability condition for (2.30) is

$$\gamma_n \int_{-1}^0 \Phi_0^2 d\bar{z} + \int_{-1}^0 g_n \Phi_0 d\bar{z} = \gamma_n + \int_{-1}^0 g_n d\bar{z} = 0.$$

The resulting expression for γ_n is identical to that obtained by substituting (2.27) into (2.26). The series formally determines $\overline{\gamma}$ as a function of R_4 and \overline{a} . The calculations are simple and we shall note only the expressions for γ_0 and γ_1 :

$$\begin{split} \gamma_0(R_4) &= \frac{1}{R_4} \int_{-1}^0 (\bar{D}^2 W_0)^2 d\bar{z} - \delta, \\ \gamma_1(R_4) &= \int_{-1}^0 \left[(\bar{D} \Phi_1)^2 - \left(\frac{\gamma_0}{\sigma} + 2\right) R_4^{-1} (\bar{D} W_0)^2 \right] d\bar{z}. \end{split}$$

Case	ſſ	fr	rr	
R_{4c}	120	320	720	
a_c	0	0	0	
$\gamma_1(R_{4c})$	-0.0301	~ 0.0340	-0.0368	
TABLE 1. Critical par	ameters for	the shallow l	ayer, $\delta = 1. \gamma_0(R_{4c}) =$	0.

The critical Rayleigh number R_{4c} is thus obtained by solving the simple boundary-value problem for W_0 , and γ_1 can be found once Φ_1 has been obtained. The results for $\delta = 1$ are given in table 1. Since $\gamma_1(R_{4c})$ is negative, the point $\bar{a} = 0$ is a relative minimum of the neutral-stability boundary, and when R_4 exceeds R_{4c} slightly the growth rate has a local maximum at the positive wavenumber $(-\frac{1}{2}\gamma_0/\gamma_1)^{\frac{1}{2}}$.

3. General results 3.1. Estimates of $R(\lambda, a)$

It is shown in appendix A that whenever U/κ has no z dependence (case II) the growth rate γ in (2.17) and (2.18) is necessarily real. The analogous statement for the shallow layer was proved by Hurle *et al.* (1967). Since γ then vanishes on the neutral curve, variational principles can be used to estimate $R(\lambda, a)$ (Chandrasekhar 1961); it is known from other similar problems that these approximations (which are upper bounds) can be accurate to within a few per cent for the simplest trial functions. We shall use this method to test the hypothesis that $a_c = 0$ for arbitrary λ in case I.

The two variational principles can be stated in terms of real functionals $I(\Phi)$ and Q(W) defined by

$$Q(W) = \int_{-\lambda}^{0} |D^2 W - a^2 W|^2 dz, \qquad (3.1)$$

$$I(\Phi) = \int_{-\lambda}^{0} (DK)^{-1} [\Gamma(z) |\Phi|^{2} + \kappa |D\Phi|^{2}] dz - [g |\Phi|^{2}]_{-\lambda}^{0}.$$
(3.2)

Here Q is a dimensionless viscous dissipation per unit horizontal area and the functions $\Gamma(z)$ and g(z) are defined in terms of K(z), κ and κ_1 in appendix A. Then R is given by either of

$$a^{2}R = \min_{\Phi} \left(I(\Phi)/Q(W) \right), \quad (D^{2} - a^{2})^{2} W = -\Phi, \quad W \text{ satisfies } (2.5 c, d), \quad (3.3)$$

$$a^{2}R = \min_{\mathbf{W}} (\tilde{Q}(W)/I(\Phi)), \quad \Phi \text{ satisfies (2.18)-(2.20);}$$
 (3.4)

the minimum being over all functions satisfying the boundary conditions, and the dependent function (W and Φ in (3.3) and (3.4) respectively) being defined for each trial function as indicated.

In table 2 we show results of calculations with constant $U, \kappa, \delta = 1$ using $\Phi = e^z$ in (3.3) and $W = \sin(\pi z/\lambda)$ in (3.4) respectively. Both calculations are for free boundaries. In all cases R had its minimum at a = 0 and for small λ the tabulated

λ.	0.5	1	2	0.2	1	2
= 0	2459	195-9	19.55	2465	197-0	19.9
0.1	2461	196-3	19.7	2467	197.3	20.1
0.2	2465	197.5	20.2	2470	198.5	20.5
0.3	2471	199.4	21.0	2476	200.5	21.3
0.4	2479	$202 \cdot 2$	$22 \cdot 1$	2485	$203 \cdot 2$	$22 \cdot 4$
0.5	2491	205.9	23.7	2496	206.7	$23 \cdot 9$
1.0	2585	$237 \cdot 1$	41.6	2588	$237 \cdot 4$	38· 2
$2 \cdot 0$	2981	385.3	128.2	2973	380.2	126.0

values bear out the very slow increase of R which is predicted by the values of γ_1 in table 1.

Estimates of $R(\lambda, a)$ when λ is large but λR is of order unity are given in §4. When $\lambda = \infty$, i.e. when the layer is semi-infinite, with free or rigid top surface, the eigenvalue problem may be solved exactly by the series method used in §3.3 below. The neutral-stability boundary is shown in table 3, along with estimates of R derived from (3.3) with $\Phi = e^z$. For each value of a in the table, wavenumbers smaller than a correspond to growing modes when R exceeds the tabulated values.

3.2. Approximate expressions for γ

The accuracy of (3.3) rests on the fact that the equation for W is solved exactly. A similar approach can be used to compute γ approximately. Given a trial function Φ we solve [cf. (2.17)]

$$(D^2 - a^2) (D^2 - a^2 q^2) W = -a^2 R \Phi, \quad q^2 = 1 + \gamma / a^2 \sigma. \tag{3.5}$$

As in the shallow-layer analysis an expression relating γ to Φ and W can be obtained by integrating (2.18) from $-\lambda$ to 0 and using (2.20) (see also §3.3 below):

$$\gamma \int_{-\lambda}^{0} \Phi \, dz + \int_{-\lambda}^{0} W DK \, dz + a^2 \int_{-\lambda}^{0} \kappa_1(K, z) \, \Phi \, dz = 0. \tag{3.6}$$

A simple choice for Φ in case II is

$$\Phi = DK \tag{3.7}$$

since \mathscr{F} then vanishes identically. In case I (with $\kappa_1 = \delta$ and $\Phi = e^z$) the computation is straightforward and gives, in the limit $\lambda \to \infty$, an implicit equation for γ :

$$\gamma = \frac{Ra}{2} \left[\frac{2 + a(1+q)}{(1+q)(1+a)^2(1+aq)^2} \right] - \delta a^2 \quad (\lambda = \infty, \text{ free surface}).$$
(3.8)

This equation was solved numerically and γ as a function of a always showed a unique maximum γ_m (at a wavenumber a_m). In figure 4 we show γ_m and a_m as functions of R, σ and δ , or more explicitly, γ_m/σ and a_m are shown as functions



FIGURE 4. The solid curves indicate (a) the predicted pattern wavenumber a_m and (b) the maximum growth rate divided by σ as functions of R/σ , for $\delta/\sigma = 0$ and 0.25, $\lambda = \infty$. The curves for the profile parameter k [cf. (2.10)] were obtained from (3.8) for k = 0 and from (3.11) for $k = \infty$. The plotted points are exact values for k = 0 computed as described in §3.3 for the following parameter values: \bigcirc , $\delta/\sigma = 0$, $\sigma = 1$; \bigcirc , $\delta/\sigma = 0$, $\sigma = 4$; \bigcirc , $\delta/\sigma = 0.25$, $\sigma = 4$.

of $R\sigma$ and δ/σ . Note that the estimates in table 3 for the free-surface case can be obtained by setting $\gamma = 0$ and $\delta = 1$ in (3.8).

If the horizontal diffusion term in (2.18) is neglected, the approximation (3.5)-(3.7) can be given a simple physical interpretation. In neglecting diffusion we nevertheless retain the sublayer thickness h as a parameter and regard K as an arbitrary prescribed concentration profile. Then Φ as defined by (3.7) determines an infinitesimal vertical shift of K(z) as the instability develops, \dagger and (3.6) is replaced by

$$\int_{-\lambda}^{0} (\gamma + W) DK dz = 0.$$
(3.9)

Viewed in this way γ is the growth rate for a prescribed concentration profile, the process by which the stratification was established having disappeared from the problem. This provides a way of looking at the model studied by Plesset & Winet (1974), who took

$$K(z) = \begin{cases} 1, & -1 \le z \le 0, \\ 0, & z < -1. \end{cases}$$
(3.10)

With (3.10) substituted in (3.9), $\gamma = -W(1)$ and (3.5) can be solved to obtain the following equation for infinite depth:

$$\gamma^2 + a^2 \sigma \gamma (1+q) = Ra\sigma/(1+\coth a)$$
 ($\lambda = 8$, free surface). (3.11)

This expression agrees with that given by Plesset & Winet if the q appearing on the left is replaced by 1. Curves based on (3.11) are shown in figure 4. Note that (3.11) is in fact an exact result when all diffusion and the negative geotaxis are neglected. This is because (3.9) with (3.10) is then equivalent to $(\gamma + W)DK = 0$.

† Physically, $\Phi = DK$ corresponds to a modulated vertical shift of the equilibrium concentration, and can be generated by $K(z+fe^{\gamma t})$ for small f(x, y).

	Rigid		Free		
a	Computed	$2(1+a)^4$	Computed	$2a(1+a)^{3}$	
0	2.00	2.00	0.00	0.00	
0.05	2.41	2.43	0·1 2	0·12	
0.10	2.83	2.93	0.26	0.27	
0.15	3.29	3.50	0.42	0.46	
0.20	3.80	4.15	0.62	0.69	
0.40	6.53	7.68	2.07	2·2 0	
0.60	10.83	13-11	4.60	4.92	
0.80	17.23	21.00	8.69	9.33	
1.00	26.26	32.00	14.85	16.00	

TABLE 3. Computed and estimated values of R for a fluid of infinite depth, case I, $\delta = 1$, for the upper surface rigid and free. The formulae follow from (3.3) with $\Phi = e^{z}$.

3.3. Solution for infinite depth, case I

To solve the eigenvalue problem exactly for $\lambda = \infty$, we write

$$W = A_1 W_1 + A_2 W_2 + A_3 W_3$$

where the W_i are solutions of

$$\begin{split} (\gamma + \delta a^2 + D - D^2) \left(D^2 - a^2 \right) \left(D^2 - a^2 q^2 \right) W &= a^2 R e^z W. \\ W_i &= \sum_{n=0}^{\infty} \exp\left[(p_i + n) z \right] c_n^{(i)}, \quad c_0^{(i)} = 1, \\ p_1 &= a, \quad p_2 = aq, \quad p_3 = \frac{1}{2} \{ 1 + [1 + 4(\delta a^2 + \gamma)]^{\frac{1}{2}} \}. \end{split}$$

We take

The boundary conditions are

$$W = D^2 W = (D-1)(D^2 - a^2)(D^2 - a^2q^2) W = 0, \quad z = 0.$$

Thus γ can be found by applying Newton's method to a 3×3 determinant. Numerical results show the expected dependence on σ when a_m and γ_m/σ are expressed as functions of R/σ and δ/σ . The exact values of a_m and γ_m/σ agree reasonably well with the approximate values calculated as in §3.2, the largest discrepancy in a_m occurring for $\delta = 0$, where the error is 7 % for $\sigma = 4$. Errors in γ_m/σ are less than 3 % for $\sigma = 0$, but for $\delta/\sigma = 0.25$ the exact value is larger than the approximate one by 15% at R = 8, $\sigma = 2$ (see figure 4). We recall that exponential and square profiles are two extreme cases within the family (2.10) of concentration profiles which might be considered representative in this problem.

3.4. Computation of $R_c(\lambda)$

We turn now to the λ dependence of the critical Rayleigh number. In view of the estimates in §3.1 the conjecture $a_c = 0$ appears to be reasonable in case I. We have not been able to show that this is true generally, but we shall show here that quite generally $R(\lambda, 0)$ is finite and positive. Assuming that the conjecture holds and $R_c(\lambda) = R(\lambda, 0)$, our sufficient condition for instability gives the critical Rayleigh number.

λ	ff	fr	rr	
0.2	2459.0	6213.0	14790.0	
1.0	$195 \cdot 9$	470·2	1185.0	
1.5	49.04	$122 \cdot 2$	300.0	
$2 \cdot 0$	19.55	$42 \cdot 82$	121.3	
$2 \cdot 5$	10.03	21.09	63.37	
3.0	6.017	12.19	38·8 2	
4 ·0	2.883	5.481	19.54	
5.0	1.734	3·132	12.44	
7.0	0·89 2	1.494	7.244	
10.0	0.492	0.771	4.808	
œ	0.000	0.000	2.000	

TABLE 4. $R_{c}(\lambda)/\delta$ for case I, with two free surfaces, one rigid and one free or both rigid.

The analysis follows the scheme outlined in §2.4. Expansions of the form

$$(\gamma, W) = \sum_{n=0}^{\infty} a^{2n+2}(\gamma_n, W_n), \quad \Phi = \sum_{n=0}^{\infty} a^{2n} \Phi_n$$

are substituted into (2.17)–(2.20), and $R(\lambda, 0)$ is obtained by setting $\gamma_0 = 0$. One analytical point requiring comment concerns the form of the solvability condition. The problem for Φ_n is now

$$D\mathscr{F}_n = -\gamma_n \Phi_0 + g_n, \quad \mathscr{F}_n = -\kappa(DK) D(\Phi/DK), \quad (3.12a, b)$$

$$= 0, \quad z = -\lambda, 0,$$
 (3.12*c*)

$$\Phi_0(0) = 1, \quad \Phi_n(0) = 0, \quad n \ge 1. \tag{3.13}$$

The boundary-value problem which is adjoint to the homogeneous version of (3.12) is $D_{1}(D_{1}K)$ $D_{2}(2.14)$

$$D[\kappa(DK)D\Psi] = 0, \qquad (3.14a)$$

$$D\Psi = 0, \quad z = -\lambda, 0. \tag{3.14b}$$

The solution for this problem is $\Psi = \text{constant}$. The solvability condition is therefore again obtained at each stage simply by integrating the equation for Φ_n [cf. (2.26) and (3.6)].

The remainder of the calculation is straightforward and there results

$$\gamma_0[1 - K(-\lambda)] - R \int_{-\lambda}^0 (D^2 W_0)^2 dz + \int_{-\lambda}^0 \kappa_1 DK dz = 0, \qquad (3.15a)$$

where $D^4 W_0 = DK$ with $W_0 = 0$, DW_0 or $D^2 W_0 = 0$ for $z = -\lambda, 0.$ (3.15b)

For case I the values for $R_c(\lambda)$ are given in table 4. We find

 \mathcal{F}_n

$$R_c \sim \delta \beta_1 / \lambda^4, \quad \lambda \to 0,$$

where $\beta_1 = (120, 320, 720)$ in cases (ff, fr, rr) in agreement with the shallow-layer theory. We also note that for a free upper surface

$$R_c \sim \delta \beta_2 / \lambda, \quad \lambda \to \infty,$$
 (3.16)

where $\beta_2 = (3, 4)$ for a (free, rigid) lower boundary. Thus a free sublayer on a sufficiently deep layer is always unstable. If R is fixed and the upper surface is

free, there is, moreover, a value of λ (and hence a value of H) below which the layer is stable. For fixed λ , there is correspondingly a value of R (hence c_0) below which the layer is stable. The last statement remains true if the reference concentration in (2.14) is taken to be the mean concentration c_m rather than c_0 ; this change in effect multiplies the tabulated values of R_c by $(1 - e^{-\lambda})/\lambda$.

4. Singular perturbation for deep layers

4.1. Remarks

The calculations given in §§ 3.2 and 3.3 describe the nature of the linear stability in deep layers when R is O(1) (that is, when R is significantly greater than R_c); there the maximum growth rate presumably gives an indication of pattern size. At the critical boundary, on the other hand, the onset of instability resembles a typical convective instability but with the unusual property that $a_c = 0$. Our aim in the present section is to study the transition between these two distinct regimes in a parameter range for which a perturbation analysis is feasible.

Such an analysis is motivated by observations of pattern formation in slightly tilted containers (see figure 1). Since the layer depth then varies linearly and rather slowly, one would expect to see rather large-scale structures near the critical depth (where a_m is very small), followed by a gradual reduction of pattern size until the value appropriate to infinite depth-is reached. What is in fact observed is an abrupt appearance of a well-defined horizontal scale (determined by the spacing of polka dots or columns) with no noticeable transition through small wavenumbers.

The key to a study of the transition region in deep layers is the introduction of the limit process

$$\lambda \to \infty$$
 with $\tilde{R} = \lambda R$, $\tilde{a} = \lambda a$, $\tilde{\gamma} = \lambda^2 \gamma$ fixed (4.1)

in place of the expansion for small a. In particular the horizontal wavelength will be of the order of the layer depth in the transition region. The justification for the choice (4.1) is the following: for $\lambda \ge 1$ the behaviour of R_c is given by (3.16), suggesting that \tilde{R} is the proper O(1) Rayleigh number in the transition region. Moreover, near a = 0 on the neutral-stability boundary we expect to have $R - R_c = O(a)$, and in order to be able to resolve the transition we shall therefore want $a \sim \lambda^{-1}$. Finally, the definition of $\tilde{\gamma}$ is suggested by (3.6) when tilde variables are substituted. (Physically we expect the maximum growth rate to be established by the effect of horizontal diffusion, which is $O(a^2)$ in (3.6).) If the quantities (4.1) are introduced into (3.8) we obtain in the limit (σ is fixed and O(1))

$$\tilde{\gamma} = \tilde{a}\tilde{R}/(1+q) - \delta\tilde{a}^2, \quad q^2 = 1 + \tilde{\gamma}/\sigma\tilde{a}^2. \tag{4.2}$$

We shall show that (4.2) is an asymptotic result provided that \tilde{R} and \tilde{a} are numerically *large*, with $\tilde{R}/\tilde{a} = O(1)$. Thus in a sense (4.2) is the asymptotic (for large \tilde{R}) form of a certain equation we wish to determine, the latter equation providing the dependence of $\tilde{\gamma}_m$ upon \tilde{R} right up to the critical value \tilde{R}_c (note that $\tilde{R}_c = 4$ for case fr).

4.2. Construction of matched expansions

The analysis is facilitated by the use of matched asymptotic expansions, since the sublayer (or inner) region requires special attention, while the flow in the main body of the layer (the outer region) has an especially simple structure. The correct choice of outer variable is suggested by the momentum equation (2.17) written in the variables (4.1):

$$D^{4}W = \lambda^{-2}\tilde{a}^{2}(1+q^{2}) D^{2}W - \lambda^{-3}\tilde{a}^{2}\tilde{R}\Phi - \lambda^{-4}\tilde{a}^{4}q^{2}W.$$
(4.3*a*)

The corresponding equation obtained from (2.18) is, in case I,

$$D^{2}\Phi - D\Phi = e^{z}W + \lambda^{-2}(\tilde{\gamma} + \delta\tilde{a}^{2})\Phi.$$
(4.3b)

For (4.3 *a*) there is a simple limit as $\lambda \rightarrow \infty$, namely

$$D^4W=0.$$

However, it is clear that this limit is not valid uniformly over the layer and we introduce an outer variable which makes D^2 formally of order λ^{-2} :

$$\tilde{z} = z/\lambda = O(1)$$
 (outer region). (4.4)

The fact that limits with \tilde{z} fixed are clearly non-trivial provides additional evidence that \tilde{a} is the correct O(1) wavenumber in the transition region.

If (4.3a) is written in terms of the outer variable we have

$$(\tilde{D}^2 - \tilde{a}^2) (\tilde{D}^2 - \tilde{a}^2 q^2) W = -\lambda \tilde{a}^2 \tilde{R} \Phi, \quad \tilde{D} = d/d\tilde{z}.$$
(4.5)

In view of (4.3*b*) it is reasonable to expect that, in case I, Φ is exponentially small in the outer region (a fact already suggested by the expansions in *a*), so that (4.5) reduces to $(\tilde{D}^2 - \tilde{a}^2)(\tilde{D}^2 - \tilde{a}^2a^2)W = 0$ (4.6)

$$(D^2 - \tilde{a}^2) (D^2 - \tilde{a}^2 q^2) W = \mathbf{0}.$$
(4.6)

We shall apply a particular ordering of W and Φ formally and show that the required matching can be carried out using standard methods (Cole 1968, chap. 1). We consider case I first.

Inner expansions having the forms

$$W = \lambda^{-2} W_0 + \lambda^{-3} W_1 + \lambda^{-4} W_2 + \dots, \qquad (4.7 a)$$

$$\Phi = \Phi_0 + \lambda^{-1} \Phi_1 + \lambda^{-2} \Phi_2 + \dots \tag{4.7b}$$

are introduced, and we consider an outer solution of (4.6):

$$\begin{split} \tilde{W} &= \lambda^{-1} [A(\lambda) \sinh \tilde{a}\tilde{z} + B(\lambda) \cosh \tilde{a}\tilde{z} + C(\lambda) \sinh \tilde{a}q\tilde{z} \\ &+ D(\lambda) \cosh \tilde{a}q\tilde{z}] \sim \lambda^{-1} \tilde{W}_0 + \lambda^{-2} \tilde{W}_1 + \dots \end{split}$$
(4.8)

The boundary conditions on the lower wall require

$$\tilde{W}_n(-1) = 0, \quad \tilde{D}\tilde{W}_n(-1) \quad \text{or} \quad \tilde{D}^2\tilde{W}_n(-1) = 0.$$
 (4.9)

Substitution of (4.7) into (4.3a, b) with the help of the conditions on z = 0 gives a series of easily solvable problems and we have for the first two terms

$$\Phi_0 = e^z, \quad W_0 = \alpha_0 z + \beta_0 z^3, \Phi_1 = 0, \quad W_1 = \tilde{a}^2 \tilde{R} (1 + \frac{1}{2} z^2 - e^z) + \alpha_1 z + \beta_1 z^3.$$
(4.10)

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Here α_n and β_n are arbitrary constants. To indicate how these constants are determined by matching, we consider the outer expansion of the inner partial sum of W up to terms of order λ^{-3} , obtained simply by discarding exponentially small terms:

$$\lambda^{-2}W_0 + \lambda^{-3}W_1 = \lambda\beta_0 \tilde{z}^3 + \beta_1 \tilde{z}^3 + \lambda^{-1} (\alpha_0 \tilde{z} + \frac{1}{2}\tilde{a}^2 \tilde{R} \tilde{z}^2) + O(\lambda^{-2}).$$
(4.11)

The addition of the term $\lambda^{-4}W_2$ to the left side would add a polynomial to the right side, but it is apparent that only multiples of \tilde{z}^5 , \tilde{z}^4 and \tilde{z}^3 respectively could thereby be introduced into terms of order λ , 1 and λ^{-1} , with even higher powers coming from subsequent terms. Thus the terms on the right side of (4.11) must coincide with terms of these orders in the power series for \tilde{W} . Now since (4.9) must be satisfied, the term of leading order in (4.11) cannot be a multiple of \tilde{z}^3 . Hence $\beta_0 = \beta_1 = 0$. The last term on the right side of (3.11) gives the three conditions (with $A = A_0 + \lambda^{-1}A_1 + \ldots$, etc.)

$$B_0 + D_0 = 0, \quad B_0 + q^2 D_0 = \tilde{R}, \tag{4.12a}$$

$$A_0 + qC_0 = \alpha_0/\tilde{a}. \tag{4.12b}$$

The first two of these, along with two conditions from (4.9), determine \tilde{W}_0 uniquely, while the last determines the remaining constant α_0 . To complete the calculation, the equation for $\tilde{\gamma}$, valid to order 1, follows from integration of (4.3 *a*) with $W_0 = \alpha_0 z$:

$$\tilde{\gamma} = \alpha_0 - \delta \tilde{a}^2.$$

Solving for $A_0, ..., D_0$ using (4.6) and (4.12*a*), and evaluating α_0 from (4.12*b*) we obtain

$$\tilde{\gamma} = \tilde{a}\tilde{R}G(\tilde{a},q) - \delta\tilde{a}^2, \qquad (4.13)$$

where, for case fr,

$$G(\tilde{a},q) = \frac{1}{q^2 - 1} \left[\frac{(1+q^2)\sinh\tilde{a}\sinh\tilde{a}q - 2q\cosh\tilde{a}\cosh\tilde{a}q + 2q}{q\sinh\tilde{a}\cosh\tilde{a}q - \cosh\tilde{a}\sinh\tilde{a}q} \right].$$
(4.14)

This is the desired equation for the growth rate, valid through the transition region as discussed in §4.1.

Before considering (4.13) in detail we indicate how the higher-order terms can be found. In order to determine the equation for $\tilde{\gamma}$ to terms of order λ^{-n} inclusive, inner terms up to W_{n+1} must be obtained. (Actually for the last term only D^2W_{n+1} need be known.) Each of these can be written as a finite sum of the form $p_0(z) + e^z p_1(z) + \ldots + e^{mz} p_m(z)$, where the p_i are polynomials. Only p_0 enters the matching. If we now suppose that $\alpha_0, \ldots, \alpha_{n-1}, \beta_0, \ldots, \beta_n$ are known, then the coefficients of z^0 and z^2 in the outer expansion of $\lambda^{-2}W_0 + \ldots + \lambda^{-(n+3)}W_{n+1}$ are known up to terms of order $\lambda^{-(n+1)}$. The corresponding matching conditions, along with the two conditions (4.9) on the lower wall, are then sufficient to determine \tilde{W}_n , and α_n can be found by matching the term in \tilde{z} . Finally β_{n+1} is obtained by matching with the \tilde{z}^3 term in \tilde{W}_{n-1} . We have carried out these calculations for n = 2 and record the results in appendix B.



FIGURE 5. Normalized pattern wavenumber and growth rate as functions of $\Delta = (R - R_c)/R_c$ for a deep layer. a_m^{∞} and γ_m^{∞} are the asymptotic values of a_m and γ_m for $\lambda \to \infty$, $R \leq 1$, and can be found from (4.2).

4.3. Discussion of (4.13)

We first note that, if \tilde{R} and \tilde{a} are both large with q and \tilde{R}/\tilde{a} both O(1), the expression (4.14) reduces to G = 1/(1+q) and (4.13) is then identical to (4.2). To determine $\tilde{\gamma}_m$ (the maximum of $\tilde{\gamma}$) and \tilde{a}_m when \tilde{R} is O(1) we first suppose that $\sigma \ge 1$ so that $q \simeq 1$. We then may use

$$\lim_{q \to 1} G(\tilde{a}, q) = \frac{1}{2} \left(\frac{\sinh^2 \tilde{a} - \tilde{a}^2}{\sinh \tilde{a} \cosh \tilde{a} - \tilde{a}} \right) \quad (\operatorname{case} fr).$$

With this simplification $\tilde{\gamma}_m$ and \tilde{a}_m are shown in figure 5 as functions of $\Delta = (R - R_c)/R_c$. It is found that once R exceeds R_c by 10% the pattern wavenumber remains within 25% of the asymptotic value

$$a_m^{\infty} = \frac{1}{4}R\delta \quad (R \ll 1, \, \lambda = \infty). \tag{4.15}$$

This is the value of a which maximizes γ in (4.2) and it is therefore appropriate to a layer of infinite depth when R is small compared with unity. Thus the transition from a 'stable' to a 'deep' layer is found to be remarkably abrupt for the weakly unstable situation considered here. We are led to conjecture that this abrupt transition will also occur when R has more realistic values in the range 1-10.

There remains the question of the possible effect of finite σ on these conclusions, (4.15) and figure 5 being formally valid only for large σ . An indication is given by the Prandtl-number dependence of the asymptotes $\tilde{\gamma}_m^{\infty}$ and \tilde{a}_m^{∞} . We write (4.2) in the form

$$\tilde{R}/\tilde{a}\delta = \theta = (q+1)[\eta(q^2 - 1) + 1] \quad (\eta = \sigma/\delta).$$
(4.16)

Differentiating (4.16) with respect to \tilde{a} and setting $d\tilde{\gamma}/d\tilde{a} = 0$ we obtain a second equation in q and θ : $q^2 + (1-\theta)q + \frac{1}{2}\theta = 0.$

Eliminating q from the last two equations produces a quadratic equation for $\eta(\theta)$:

$$9(\theta - \frac{1}{4}\theta^2)\eta^2 + (-2\theta^3 + 10\theta^2 - 10\theta - 1)\eta + 1 = 0.$$
(4.17)

Examination of the real roots of (4.17) shows that when $\eta > \frac{2}{3}$ (on the basis of estimates given in §5, $\frac{2}{3}$ is a satisfactory lower bound) the parameter θ varies between $2 + \sqrt{3}$ and 4, the latter corresponding to $\sigma = \infty$ [cf. (4.15)]. Thus a_m^{∞} is within 8% of the value at $\sigma = \infty$ when σ exceeds $\frac{2}{3}\delta$. The quantity $\gamma_m^{\infty 2}/R^2$ varies from 0.041 to 0.0625 as σ/δ increases from $\frac{2}{3}$ to ∞ . Consequently, in the limit considered here the effect of σ on the development of the instability is not very significant.

4.4. Case II

We now consider the effect of the choice of U and **D** upon the first-order expression (4.13) for the growth rate. We shall show that, if we define (still in the dimension-less variables (2.13))

$$\delta = \int_{-\lambda_s}^0 \kappa_1(K, z) \, DK \, dz$$

and use the equivalent sublayer thickness defined by (2.12a) in place of h in the definition (2.13) of the starred variables, then (4.13) holds also in case II.

The proof hinges on the fact that the exact form of K(z) enters into the computation of α_0 only through the term z^2 in the outer expansion of the inner W_1 ; this term contributes the only inhomogeneous matching condition at that order. Now in case II we have $\Phi_0 = DK$ and therefore

$$D^2 W_1 = -\tilde{a}^2 \tilde{R} \int_{-\lambda_s}^z K(z) \, dz + \tilde{a}^2 \tilde{R} \int_{-\lambda_s}^0 K(z) \, dz,$$

which immediately gives the coefficient of \tilde{z}^2 in the outer expansion of W_1 . If (4.13) is to be unchanged we must have

$$\int_{-\lambda_e}^{0} K(z) \, dz = 1. \tag{4.18}$$

But (4.18) follows from the choice of h_e as the unit of length, since (4.18) and (2.12*a*) are then equivalent. The above definition of δ follows easily from the integral of (2.18) with $W = W_0$ and $\Phi = \Phi_0 = DK$, with $\Phi_0(0)$ taken as unity.

5. Discussion

The measurements of Winet & Jahn (1972) and the data reported by Plesset & Winet (1974) (see table 5 below), both for *Tetrahymena pyriformis*, provide estimates of the values of the dimensionless parameters in typical *Tetrahymena* cultures. The variable c should be regarded as the volume density of only those organisms which actively participate in the pattern (here, the negatively geotactic organisms). A proper determination of c, U, κ and κ_1 would presumably

				Pattern wavenumber, a_m		
h (cm)	αgc_0	R	σ	Measured	$k = \infty$	k = 0
0.15	$1 \cdot 21 \times 10^{-4}$	5.9	$1 \cdot 5$	0.94	0.92	$0.73 \ (\delta = 0, \ \sigma = 1)$
0.05	1.1×10^{-3}	6.0	4.4	0.57	0.70	$0.59 \ (\delta = 0, \sigma = 4)$

TABLE 5. Data from Plesset & Winet (1974) compared with theoretical values of a_m obtained from the values shown in figure 4 (a). The theoretical values given by Plesset & Winet are 0.86 (h = 0.15) and 0.65 (h = 0.05), and were obtained from (3.11) with q = 1. The parameter k is defined in (2.9).

require a statistical average which accounts for the variability of organisms; as a rough approximation we shall simply regard c as the difference between the actual volume density and the (constant) value usually observed in deep cultures well outside the sublayer. We take $\alpha = 0.09$ and $U_0 = 0.045$ cm/s and assume throughout that $\nu = 0.01$ cm²/s. We are not aware of any direct measurement of effective diffusivities for the random component of the organisms' motion, and we shall therefore infer a value of κ_0 from the estimates of the sublayer thickness using $\kappa_0 = U_0 h$. For a typical organism density of 10^5 cm⁻³ and an organism volume of 1.8×10^{-8} cm³ we estimate c_0 as 1.8×10^{-3} . With h = 0.1 cm we obtain R = 3.5. The two measurements in table 5 are for R = 6 and σ in the range 1–5. We take R = 5, $\sigma = 2$ and h = 0.1 cm as typical of pattern-forming cultures of *Tetrahymena pyriformis*.

Wille & Ehret (1968) report a critical depth of 2 mm, and if we assume that this critical depth applies to our hypothetical culture with R = 5 we see from table 4 that R/δ would be about 45, implying that δ is about 0.1. Another estimate of δ can be obtained from the measurements of pattern size in table 5. These were calculated from the distribution of internodal distances when the pattern is reticulate. Although these data are not sufficient for detailed comparison of various choices of U, κ and κ_1 , the indication is that δ is small compared with unity and, if the functions (2.9) are adopted, that the profile parameter k should be positive. As a final check, we note that a simple random walk on a cubic lattice of side L would lead to isotropic diffusion with $\kappa_1 = \frac{1}{6}L^2/\Delta t = \frac{1}{6}U_0L$, and therefore $\delta = \kappa_1/\kappa_0 = \frac{1}{6}L/\hbar$. If the mean free path of horizontal diffusion is of the order of \hbar a value of δ of about 0.1 is not unreasonable.

The apparent difference between the effective vertical and horizontal diffusion prompts us to rewrite the diffusivity tensor as

$$\mathbf{D} = \kappa_g \mathbf{k} + \kappa_1 \mathbf{I} \tag{5.1}$$

and to conjecture that the parameter $\kappa_g = \kappa - \kappa_1$ vanishes when no gravity field is present. Thus κ_g represents the effect of the random vertical movements which are involved in the organisms' response to the gravitational field (and conceivably are involved in the mechanism by which the organisms sense the direction of the field). The isotropic term $\kappa_1 \mathbf{I}$ in (5.1) presumably accounts for the homogeneity of cultures which are not negatively geotactic. At present we do not know any way of obtaining κ_g independently so that a comparison with the observed sublayer thickness might be made. Roberts (1970) computes what is effectively

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a sublayer thickness for negative geotaxis of *Paramecium*, basing his calculation on the idea that the preferred orientation of the body is induced by its shape and density distribution. This is a restrictive assumption but Roberts' result is compatible with (5.1), in the sense that it shows one way of relating vertical diffusion directly to the response of the organism to the gravitational field. In any case, even for non-identical organisms, dispersion in the motility can be modelled in terms of these diffusion constants.

A related example of the combined effect of random and directed motion of organisms, leading to the formation of macroscopic structures, was studied in a continuum model by Keller & Segel (1970). In their work an equation similar to (2.4), involving a chemotactically induced flux of organisms, was coupled with diffusion-reaction equations. Their chemotactic flux, proportional to the gradient of concentration of a chemical, corresponds to our geotactic flux kcU. The analogy suggests that the latter term may also be best regarded as proportional to a gradient, in particular to the pressure gradient. In the dilute limit $c \ll 1$ this altered form would be equivalent to (2.4), but in general there would arise new physical effects, such as horizontally directed 'barotactic swimming'. Advantages of using the pressure-gradient form may appear, even in the dilute case, if centrifugal acceleration is important. (Some very preliminary observations of pattern formation on a rotating turntable indicate that patterns may be modified by rotation.) In addition, it is not impossible that actual chemotaxis is sometimes involved in pattern formation by swimming micro-organisms (Brinkman 1968).

Returning to the present results, we see on comparing table 4 and figure 5 that we can estimate the width of the transition from a marginally stable to a 'deep' culture as H changes slowly with x. We assume that R is constant and that the small R theory can be applied to the transition even though R_c is not small. For the culture with R = 5, a critical depth of 2h and $\delta = 0.1$, we see that once the depth has increased to 3h the value of R_c has dropped to about 1.2, making Δ in figure 4 about 1. This suggests that the critical depth is marked by a transition region in which the depth changes by no more than 0.5 mm. In some of our typical experiments with *Tetrahymena* we observed critical depths of about 4.5 mm, which, if δ is again assumed to be 0.1, would indicate an R of only about 0.4, but nevertheless a value $\Delta = 0.25$ was reached at a depth of 5 mm.

The linear stability problem considered in this paper is an idealization which cannot be completely realized in experiments, whether in tilted layers or in initially stirred, deep cultures. The sublayer forms in a time of the order of $H/U_0 \sim 30$ s, and if R is between 1 and 5 the equilibrium envisaged here presumably never has a chance to form completely. The transient nature of the convective instability is emphasized by the description of column formation given by Winet & Jahn (1972).

A related thermal instability has been analysed and observed by Foster (1965, 1969). He considered Bénard convection in a layer cooled from its top boundary. The formation of a cool zone at the top is clearly analogous to sublayer formation, and the onset of what Foster called "manifest convection", involving the abrupt falling away of thin sheets of cool fluid from the top layer, is similar to the rapid transition in a tilted layer envisaged in §4.3, the horizontal co-ordinate there

taking the place of the time variable in the thermal instability. In detail the two problems are different, but in a broad sense the physical description of the instability is the same.[†] Other dynamical analogies are suggested by observations of sedimentation of swarms of inert particles in a viscous fluid and by the behaviour of fluidized beds operating at small particle Reynolds number, but owing to differences in the boundary conditions neither of these analogies is exact.

The structure of bioconvection patterns following the initial instability considered above is likely to be a complex and nonlinear process since the Reynolds number based on U_0 and H is typically in the range 1–5 and any new steady equilibria are quite different from a simple vertically stratified layer. The application of the present model to this problem, and to the construction of steady solutions similar to the polka-dot patterns discussed in §1, will be taken up in a separate paper.

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Appendix A. Integral identities of linear theory

Multiply (2.17) by W^* (where * means 'complex conjugate' in this appendix) and integrate with respect to z from $-\lambda$ to 0 to obtain

$$\frac{\gamma}{\sigma} \int_{-\lambda}^{0} (|DW|^2 + a^2 |W|^2) dz + a^2 R \int_{-\lambda}^{0} \Phi^* W dz + \int_{-\lambda}^{0} |(D^2 - a^2) W|^2 dz = 0.$$
 (A 1)

Now differentiate condition (2.6) and evaluate the result for c = K(z):

$$\left[K\frac{\partial U}{\partial z} - \frac{\partial \kappa}{\partial z}DK\right]_{c=K} + \left[U + c\frac{\partial U}{\partial c} - \frac{\partial \kappa}{\partial c}DK\right]_{c=K}DK - \kappa(K,z)D^2K = 0.$$
(A 2)

Note that the first bracketed term in (A 2) vanishes for both case I and case II since $\Gamma_{AU} = UKar^{2}$

$$\left[K\frac{\partial U}{\partial z} - \frac{UK}{\kappa}\frac{\partial \kappa}{\partial z}\right]_{c=K} = K\kappa \left[\frac{\partial}{\partial z}\left(U/\kappa\right)\right]_{c=K} = 0.$$
 (A 3)

With the help of (A 2) and (A 3), note that (2.19) can be rewritten as

$$\mathscr{F} = -\kappa(DK) D(\Phi/DK).$$

Using this expression and (2.20), we multiply Φ/DK by the complex conjugate of (2.18) and integrate to obtain

$$\int_{-\lambda}^{0} (DK)^{-1} \left\{ \left[\gamma^* + a^2 \kappa_1 + \kappa \left(\frac{D^2 K}{DK} \right)^2 \right] |\Phi|^2 + \kappa |D\Phi|^2 \right\} dz + \int_{-\lambda}^{0} \frac{\kappa D^2 K}{(DK)^2} (\Phi D \Phi^* + \Phi^* D \Phi) dz + \int_{-\lambda}^{0} \Phi W^* dz = 0.$$
 (A 4)

† (Note added in proof.) Other methods of generating a relatively heavy sublayer as a model for the initial bioconvective instability were considered by H. Wager (1911). S. Childress, M. Levandowsky and E. A. Spiegel

If we subtract a^2R times (A 4) from (A 1) and take the imaginary part we find

$$\operatorname{Im}(\gamma)\left[a^{2}R\int_{-\lambda}^{0}(DK)^{-1}|\Phi|^{2}dz+\sigma^{-1}\int_{-\lambda}^{0}(|DW|^{2}+a^{2}|W|^{2})dz\right]=0,\quad (A\ 5)$$

which establishes that γ is real.

A partial integration of (A 4) yields

$$\gamma^* \int_{-\lambda}^0 (DK)^{-1} |\Phi|^2 dz + \int_{-\lambda}^0 \Phi W^* dz + \int_{-\lambda}^0 \kappa |D\Phi|^2 dz + I(\Phi) = 0, \quad (A 6)$$

where $I(\Phi)$ is given by (3.2) with

$$\Gamma(z) = \kappa \left(\frac{D^2 K}{D K}\right)^2 + D K D \left(\frac{\kappa D^2 K}{(D K)^2}\right) + a^2 \kappa_1, \quad g(z) = \kappa \frac{D^2 K}{(D K)^2}.$$
 (A 7)

The variational principles (3.3) and (3.4) follow from (A 1) and (A 6) with $\gamma = 0$ by computing the first variation and essentially reversing the steps in the above derivation.

Appendix B. Matched asymptotic expansions for a deep layer, case I, $ff, fr; a = \tilde{\lambda}a, \tilde{R} = \lambda R, \tilde{\gamma} = \lambda^2 \gamma, \lambda \to \infty$

$$\begin{split} Inner\ expansions \\ q^2 &= 1 + \tilde{\gamma}/\tilde{a}^2\sigma, \quad \Phi = \Phi_0 + \lambda^{-2}\Phi_2 + O(\lambda^{-3}), \\ W &= \lambda^{-2}(W_0 + \lambda^{-1}W_1 + \lambda^{-2}W_2 + \lambda^{-3}W_3) + O(\lambda^{-6}), \\ \Phi_0 &= e^z, \quad \Phi_2 = \frac{1}{2}\alpha_0 z^2 e^z, \\ W_0 &= \alpha_0 z, \quad W_1 = \tilde{a}^2 \tilde{R}[1 + \frac{1}{2}z^2 - e^z], \quad W_2 = \alpha_2 z + \beta_2 z^3, \\ W_3 &= \tilde{a}^4 \tilde{R}(1 + q^2) \left(\frac{1}{2}z^2 + \frac{1}{24}z^4 - e^z\right) - \alpha_0 \tilde{a}^2 \tilde{R} e^z (\frac{1}{2}z^2 - 4z + 10) \\ &+ 10\alpha_0 \tilde{a}^2 \tilde{R} + \frac{3}{2} \tilde{a}^2 \tilde{R} \alpha_0 z^2 + \alpha_3 z. \end{split}$$

$$\begin{split} &Outer\ expansions\\ &\tilde{z}=z/\lambda,\quad W=\lambda^{-1}[\tilde{W_0}+\lambda^{-2}\tilde{W_2}]+O(\lambda^{-3}),\\ &\tilde{W_n}=A_n\sinh\tilde{a}\tilde{z}+B_n\cosh\tilde{a}\tilde{z}+C_n\sinh\tilde{a}q\tilde{z}+D_n\cosh\tilde{a}q\tilde{z} \end{split}$$

The constants in the above are given by the following expressions:

$$\alpha_0 = \tilde{a}\tilde{R}G(\tilde{a},q), \quad \alpha_2 = \tilde{a}(A_2 + qC_2), \quad \beta_2 = \frac{1}{6}\tilde{a}^2(A_0 + q^3C_0).$$

 α_3 is determined by matching at the next stage. In the expression for α_0 the function G is given by (4.14) for case fr and by

$$G(\tilde{a},q) = \frac{1}{q^2 - 1} \left[\frac{q \sinh \tilde{a} \cosh \tilde{a} q - \cosh \tilde{a} \sinh \tilde{a} q}{\sinh \tilde{a} \sinh \tilde{a} q} \right]$$

for case ff. We give the coefficients in the outer expansion for case fr only:

$$\begin{split} \boldsymbol{A}_{0} &= [\Gamma R/(q^{2}-1)] [q \cosh \tilde{a} \cosh \tilde{a} q - \sinh \tilde{a} \sinh \tilde{a} q - q], \\ \Gamma &= [\cosh \tilde{a} \sinh \tilde{a} q - q \sinh \tilde{a} \cosh \tilde{a} q]^{-1}, \end{split}$$

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$$\begin{split} C_0 &= [\Gamma \tilde{R}/(q^2-1)] \left[\cosh \tilde{a} \cosh \tilde{a} q - q \sinh \tilde{a} \sinh \tilde{a} q - 1\right], \\ D_0 &= -B_0 = \tilde{R}/(q^2-1), \\ A_2 &= (3\alpha_0 + \tilde{a}^2 q^2) A_0 + \tilde{a}^2 \Gamma \tilde{R} [\sinh \tilde{a} \sinh \tilde{a} q - q \cosh \tilde{a} \cosh \tilde{a} q], \end{split}$$

 $C_2 = (3\alpha_0 + \tilde{a}^2 q^2) C_0 + \tilde{a} \Gamma \tilde{R}, \quad D_2 = (3\alpha_0 + \tilde{a}^2 q^2) D_0, \quad B_2 = \tilde{a}^2 \tilde{R} - D_1.$

The equation for $\tilde{\gamma}$ up to terms of order λ^{-2} is found to be

$$\begin{split} \tilde{\gamma} &= \tilde{a} \tilde{R} G(\tilde{a},q) - \delta \tilde{a}^2 - \lambda^{-1} (\frac{3}{2} \tilde{a}^2 \tilde{R}) \\ &+ \lambda^{-2} [(q^2 - \delta) \tilde{a}^3 \tilde{R} G(\tilde{a},q) + 3 \tilde{a}^2 \tilde{R}^2 G^2(\tilde{a},q) + \tilde{a}^3 \tilde{R} G_1(\tilde{a},q)] + O(\lambda^{-3}), \end{split}$$
e
$$G_1 &= \begin{cases} \coth(\tilde{a}), & \operatorname{case} ff, \\ -\Gamma(q^2 - 1) \sinh \tilde{a} \sinh \tilde{a} q + G(\tilde{a},q), & \operatorname{case} fr. \end{cases}$$

where

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