Petrogenesis and Stratigraphy of the High-Ti/Y Urubici Magma Type in the Paraná Flood Basalt Province and Implications for the Nature of 'Dupal'-Type Mantle in the South Atlantic Region

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The high-Ti/Y Urubici (or Khumib) magma type of the Paraná-Etendeka large igneous province has a restricted spatial extent, near the southeast Brazilian coast and in the northern Etendeka (Namibia). Urubici flows are interbedded with low-Ti/Y Gramado flows. Flow correlations indicate that local topographic relief was important in controlling emplacement of flows, and that lavas near the coast have undergone up to 1 km of post-magmatic uplift relative to inland areas. Urubici magmas have undergone extensive fractional crystallization (MgO <5.5 wt %). Stratigraphic variations highlight complexities of mixing and minor crustal assimilation indicative of open-system magmatic plumbing. The least contaminated samples have high La/Nb (~1.5) and $(Tb/Yb)_{N}$ (~2.5), Sr-Nd isotopes close to Bulk Earth (⁸⁷Sr/⁸⁶Sr_i ~0.7050; ε_{Nd_i} –2.7), and Dupal Pb isotopes with unradiogenic ${}^{206}Pb/{}^{204}Pb$ (~17.6). These features are similar to those of the Walvis Ridge DSDP (Deep Sea Drilling Project) Site 525A basalts that define the EM1 oceanic mantle component, and many are also shared with local Cretaceous alkalic magmas that are inferred to be lithospheric mantle melts. Low ²⁰⁶Pb/²⁰⁴Pb material found in the Urubici and Site 525A basalts is not seen as a mixing end-member within the modern Tristan plume system or in South Atlantic mid-ocean ridge basalt. An origin from lithospheric mantle material, delaminated and dispersed within the asthenosphere following continental break-up, is preferred. Thus the South Atlantic Dupal mantle anomaly cannot be considered as a single entity: Urubici flood basalts and Walvis Ridge Site 525A basalts have a relatively shallow origin within originally lithospheric mantle, whereas the Tristan plume is a deep mantle upwelling.

KEY WORDS: Paraná flood basalts; lava stratigraphy; Dupal mantle anomaly; lithospheric mantle; crustal assimilation

INTRODUCTION

Eruption of the Paraná–Etendeka flood basalt province was associated with the initial opening of the South Atlantic Ocean during the Early Cretaceous. The magmatic history of the Tristan mantle plume can be traced back from its recent activity on the islands of the Tristan group and Gough, along the Rio Grande Rise and the

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Fig. 1. Map of the South Atlantic region. Shallow features, between 4 km water depth and sea level, are highlighted in white. The magmatic effects of the Tristan mantle plume, whose present location is marked by the Tristan da Cunha island group and Gough island, can be traced back along the Rio Grande Rise and the Walvis Ridge, to the flood basalt exposures of the Paraná and the Etendeka, respectively. The Tristan da Cunha island group also includes Inaccessible and Night-ingale islands, which lie 40 km to the WSW and SSW, respectively, of Tristan da Cunha.

Walvis Ridge, to the flood basalt exposures (Fig. 1). The presence of the Tristan plume was clearly an important factor in the generation of the flood basalts, although its exact role is controversial (Morgan, 1981; Fodor, 1987; Hawkesworth et al., 1988; Piccirillo & Melfi, 1988; White & McKenzie, 1989; Peate et al., 1990; Gallagher & Hawkesworth, 1992; Saunders et al., 1992; Arndt et al., 1993; Gibson *et al.*, 1995*a*; Peate & Hawkesworth, 1996; Turner et al., 1996). The issue is whether the basalts were generated by decompressional melting of Tristan plume mantle and/or mid-ocean ridge basalt (MORB)-source asthenospheric mantle, accompanied by some lithospheric contamination, or whether the plume had a more passive role in supplying the heat required to produce the basalts by melting of lithospheric mantle. In a wider context, this issue links to discussions about the nature and location of the 'enriched' mantle components observed in some ocean-island basalts (OIB) and other mantle-derived magmas (e.g. Zindler & Hart, 1986; Hofmann, 1997).

Hawkesworth *et al.* (1986) showed that Paraná flood basalts have Dupal isotope features (high ²⁰⁷Pb/²⁰⁴Pb and ²⁰⁸Pb/²⁰⁴Pb for a given ²⁰⁶Pb/²⁰⁴Pb relative to MORB and OIB from the Northern Hemisphere; Hart, 1984). Most MORB and OIB from the South Atlantic Ocean region between the Paraná and Etendeka flood basalt exposures also show Dupal isotope characteristics. The origin and original location of this isotopically anomalous Dupal mantle is not resolved, with opinion divided between a relatively shallow source in delaminated continental lithospheric mantle material (e.g. Hawkesworth *et al.*, 1986) or a deep source in the lower mantle (e.g. Castillo, 1988). In this paper, we focus on the high-Ti/Y Urubici lavas of southern Brazil because of their compositional similarities (Hawkesworth *et al.*, 1986) to

the Deep Sea Drilling Project (DSDP) Site 525A Walvis Ridge samples that Zindler & Hart (1986) used to define their hypothesized EM1 mantle component. Is this an intrinsic component of the Tristan plume or is it lithospheric mantle material that was first sampled during the flood basalt event and then sampled 60 my later along part of the Walvis Ridge as a result of being detached and entrained in the asthenosphere (Richardson *et al.*, 1982; Hawkesworth *et al.*, 1986; Carlson *et al.*, 1996; Milner & le Roex, 1996)?

Extensive flow-by-flow sampling allowed us to determine the lava stratigraphy in detail within a small region on the coastal escarpment (the Serra Geral) of southern Brazil. This provides information about the progressive emplacement of the lava pile, and can help resolve the relative contributions of pre-existing topography and post-magmatic uplift in controlling the present-day configuration of lavas along the coastal escarpment. Temporal variations in composition can provide important clues to the petrogenetic evolution of the Urubici magmas and allow us to estimate better the compositional characteristics of unmodified mantle melts. From these results, we can discuss the mantle origins of the Paraná basalts and the implications for the nature of Dupal mantle in the South Atlantic region.

THE URUBICI MAGMA TYPE AND SAMPLE DETAILS

The Paraná basalts are divided into six distinct compositional magma types (Peate et al., 1992). These are grouped into high-Ti/Y (>310: Urubici, Pitanga, Paranapanema, Ribeira) and low-Ti/Y (<310: Gramado, Esmeralda). The Urubici magma type is volumetrically minor, making up <5% of the total preserved volume of the Paraná lavas (compared with ~50% for the high-Ti/ Y Paranapanema and Pitanga magma types: Fig. 2). Its surface extent is restricted to a relatively small strip (~100 km \times 350 km) along the north-east flank of the lava field in southern Brazil (Peate et al., 1992). Duncan (1987) found high-Ti/Y flows (Khumib magma type) in the northern part of the formerly adjacent Etendeka lavas of Namibia (Fig. 2), which are compositionally equivalent to the Urubici magma type (e.g. Fig. 3). Despite its small volume, the Urubici magma type is of interest because its Sr-Nd-Pb isotopic composition lies closest, of the high-Ti/Y magma types, to that of the DSDP Site 525A Walvis Ridge basalts that define the EM1 mantle component in the nomenclature of Zindler & Hart (1986).

New compositional data are presented in Tables 1–3 for Urubici samples collected from road profiles traversing the Serra Geral escarpment near São Joaquim (Fig. 4). The Urubici flows are interbedded with low-Ti/Y flows of the Gramado magma type, which have been discussed



Fig. 2. Pre-drift reconstruction showing the Paraná–Etendeka flood basalt province in relation to the proto-Atlantic rift (Peate *et al.*, 1990). The outlined extents of the high-Ti/Y magma types highlight the difference in erupted volume between the Urubici magma type (including the equivalent Khumib lavas in the Etendeka: Duncan, 1987) and the more extensive Pitanga and Paranapanema lavas (Peate *et al.*, 1992). The box marks the São Joaquim field area, shown in detail in Fig. 4. ES, GO, RO and AV are boreholes where single Urubici samples were recovered at depth (Table 1; Peate *et al.*, 1992). RA is a road profile which contains the southernmost known Urubici flow (Bellieni *et al.*, 1984). FD is the coast-parallel Florianopolis dyke swarm.

by Peate & Hawkesworth (1996). The number of Urubici flows within a given profile decreases southwards: none are found south of the RA profile (Fig. 4: Bellieni *et al.*, 1984). Coast-parallel dykes (striking ~NE–SW) of Urubici composition were sampled near the coast at Florianopolis, ~175 km NE of São Joaquim (Fig. 2), and several Urubici dykes of similar orientation were found in the São Joaquim area (Fig. 4).

Urubici samples are subaphyric to weakly porphyritic (<5% phenocrysts) with phenocrysts predominantly of plagioclase and augite. The lavas have a restricted compositional range (MgO = $3 \cdot 7 - 5 \cdot 4$ wt %; SiO₂ = 50 - 53 wt %), with the exception of two andesitic–dacitic flows (DSM-24, 55 wt % SiO₂; DSM-04/08, 60 wt % SiO₂) and one coastal dyke (ST93-5, 55 wt % SiO₂). Urubici and Pitanga basalts have the highest TiO₂ (>3 wt %)



Fig. 3. Fe₂O₃(t) vs Sr diagram. Urubici magma type is defined as samples with TiO₂ >2 wt % that have Sr >550 ppm and Fe₂O₃(t) <14.5 wt %. Values of high-Ti/Y Khumib samples from the Etendeka, Namibia (A. R. Duncan, unpublished data, 1992) overlap completely with the Urubici field. Data for other Paraná basalt magma types from Peate & Hawkesworth (1996) and Peate (1990). Two Pitanga samples with high Sr >550 ppm also have anomalously high Al₂O₃ and have probably accumulated plagioclase.

and highest incompatible trace element abundances (except for Rb, Ba, Th, K) of the Paraná basalt magma types. Urubici lavas have higher SiO₂, TiO₂, K₂O and lower Fe₂O₃ than Pitanga magmas, at similar MgO contents. Sr and Fe are the most diagnostic elements to distinguish between these two magma types: Urubici magmas have Sr >550 ppm and Fe₂O₃(t) <14.5 wt % (Peate *et al.*, 1992; Fig. 3). Data on coexisting groundmass Ti-magnetite–ilmenite pairs indicate that Urubici magmas equilibrated at lower $f(O_2)$ conditions [close to the QFM (quartz–fayalite–magnetite) buffer] than Pitanga and Paranapanema magmas (~0.5 log units above QFM) (Bellieni *et al.*, 1984).

Urubici magmas have less radiogenic present-day Pb isotope compositions (²⁰⁶Pb/²⁰⁴Pb = 17·4–18·3), lower ⁸⁷Sr/⁸⁶Sr_i, and a more restricted range in ⁸⁷Sr/⁸⁶Sr_i and ε_{Nd_i} (0·7047–0·7065, and –1·9 to –5·7) than the low-Ti/ Y Gramado and Esmeralda magmas (Fig. 5; Peate & Hawkesworth, 1996). Urubici lavas, in common with other Paraná lavas (Hawkesworth *et al.*, 1986; Peate, 1990; Peate & Hawkesworth, 1996) have Dupal Pb isotope compositions, with $\Delta 7/4 > 7$ and $\Delta 8/4 > 85$ ($\Delta 7/4$ 4 and $\Delta 8/4$ are measures of the vertical deviations in ²⁰⁷Pb/²⁰⁴Pb and ²⁰⁸Pb/²⁰⁴Pb above the Northern Hemisphere Reference Line, a best-fit line to data for MORB and OIB from the Northern Hemisphere; Hart, 1984).

Sample	Road	Alt.	Flow	SiO ₂	TiO ₂	AI_2O_3	Fe ₂ O ₃	MnO	MgO	CaO	Na₂O	K₂O	P_2O_5	LOI	V	Co	Ni	Cu	Zn	Ga	Rb	Sr	Y	Zr	Nb	Ba
		(m)																								
DGB-09	GB	905	1	52.55	3.59	13.74	12.48	0.15	4.22	7.73	2.78	2.24	0.52	0.56	336	31	52	147	109	25	57	649	37	305	28.0	573
DGB-39	GB	1000	2	51.88	3.79	13.00	13.28	0.17	4.50	8.57	2.52	1.77	0.52	0.25	341	41	65	n.a.	122	23	26	765	35	324	30.0	575
DGB-36	GB	1030	2	52·01	3.67	13.48	13.14	0.17	4.49	8.30	2.49	1.76	0.47	0.93	n.a.	n.a.	48	136	112	25	34	782	36	302	28.7	n.a.
DUP-01	PE	1185	3	51.37	3.89	13.14	13.33	0.18	4.80	9.15	2.50	0.93	0.54	1.39	354	38	57	171	117	26	11	894	39	313	28.4	592
DUP-02	PE	1220	3	50.53	3.89	13.48	13.60	0.18	4.87	8.90	2.85	1.19	0.52	1.98	371	42	60	185	119	25	27	977	43	325	28.6	732
DUP-03	PE	1220	3	50.29	3.95	13.20	14.03	0.18	4.99	8.86	2.60	1.39	0.51	1.80	370	40	59	175	112	25	26	830	39	314	27.3	562
DUP-07	PE	1240	3	49.73	3.74	13.92	13.45	0.16	4.93	8.93	2.69	1.95	0.50	1.65	354	40	56	84	112	32	30	811	40	307	28.8	592
DGB-31	GB	1105	3	50.07	4.00	14.12	14.00	0.16	4.61	9.25	2.20	1.03	0.56	2.31	n.a.	n.a.	40	154	112	27	38	718	43	303	27.8	n.a.
DUP-05	PE	1235	4	52·20	3.55	13.39	13.09	0.18	4.35	8.05	2.68	1.92	0.60	0.75	323	32	36	118	116	24	32	832	38	318	31.3	746
DAB-08	AB	1130	4	51.70	3.60	13.41	12.64	0.16	4.63	8.37	3.19	1.78	0.52	0.08	327	40	91	n.a.	131	25	38	724	44	300	26.0	698
DAB-06	AB	1160	4	51.67	3.63	13.73	13.20	0.17	4.39	7.87	2.82	1.98	0.54	0.15	n.a.	n.a.	39	145	119	24	42	691	41	311	28.9	n.a.
DAB-15	AB	1210	4	51.65	3.61	13.72	13.21	0.17	4.38	7.83	2.96	1.93	0.55	-0.01	n.a.	n.a.	40	154	112	27	38	718	43	303	27.8	n.a.
DUP-04	PE	1230	5	53.00	3.51	13.38	12.76	0.16	4.01	7.39	3.06	2.13	0.60	0.87	323	38	38	134	121	25	34	697	40	326	32.7	686
DUP-06	PE	1235	5	53·52	3.56	13.26	12.87	0.17	4.06	7.45	2.49	2.06	0.54	0.86	332	35	38	125	118	27	34	661	39	322	31.9	690
DAB-10	AB	1190	5	52.83	3.60	13.24	13.02	0.16	4.00	6.91	3.61	2.00	0.63	0.89	295	44	95	n.a.	142	26	43	670	44	335	28.0	817
DGB-25	GB	1175	5	53.74	3.68	12.94	12.94	0.18	3.73	7.26	2.66	2.31	0.56	0.48	344	35	47	163	123	27	47	734	39	351	34.0	816
DUP-46	UR	1190	6	51.40	3.46	12.93	12.86	0.16	4.41	7.77	3.64	1.83	0.51	0.82	330	33	47	152	96	22	43	619	39	301	26.9	457
DSM-37	SM	1235	6	52.53	3.52	13.49	12.68	0.16	4.53	7.55	3.08	1.90	0.57	0.26	331	36	50	132	103	25	44	637	40	310	27.9	785
DSM-36	SM	1260	6	52.48	3.52	13.42	12.81	0.17	4.51	7.69	2.99	1.88	0.53	0.06	344	41	49	126	117	23	42	613	40	313	28.0	701
DUP-44	UR	1220	7	51.51	3.36	13.19	12.93	0.19	5.30	8.49	2.89	1.68	0.45	0.91	339	41	92	124	99	22	40	652	39	269	23.3	794
DUP-43	UR	1225	7	50.99	3.34	13.30	12.89	0.17	5.33	9.29	2.78	1.43	0.47	0.56	365	34	78	148	104	23	23	711	35	269	24.6	437
DSM-35	SM	1275	7	51.61	3.39	13.46	12.67	0.17	5.38	8.53	2.74	1.56	0.49	0.70	343	42	99	151	104	26	31	674	38	268	23.7	503
DUP-42	UR	1270	8	51.54	3.35	13.66	12.48	0.16	4.88	8.73	2.83	1.92	0.46	0.27	336	45	66	142	108	23	38	664	34	271	25.7	525
DSM-34	SM	1300	8	51.48	3.40	13.75	13.25	0.19	4.67	8.32	3.08	1.38	0.47	0.85	314	37	54	166	102	23	29	820	34	272	26.8	709
DSM-33	SM	1305	8	50.44	3.44	14.13	12.97	0.18	4.86	8.86	2.91	1.20	0.48	1.06	365	43	60	159	109	26	19	863	37	277	28.2	736
DUP-40	UR	1290	9	52.04	3.50	13.16	12.80	0.18	5.05	8.85	2.50	1.46	0.47	1.08	358	39	67	158	103	25	21	809	34	275	26.1	572
DSM-31	SM	1325	9	51.84	3.60	13.43	12.61	0.19	4.96	8.32	2.89	1.64	0.53	1.10	347	36	60	136	113	22	31	706	39	278	24.0	743
DAB-12	AB	1250	9/10	51.62	3.51	13.35	12.79	0.18	4.92	8.58	2.78	1.76	0.51	0.08	354	49	81	n.a.	126	23	38	701	39	305	28.0	645
DW-21	CO	1275	9/10	52.00	3.62	13.11	12.86	0.16	4.83	8.15	3.10	1.69	0.48	0.49	344	33	63	135	96	23	36	602	35	279	23.8	502
DRR-18	RR	1245	9/10	51.84	3.61	13.31	13.02	0.17	4.88	8.00	2.68	2.02	0.49	0.13	n.a.	n.a.	69	187	115	23	36	643	38	279	25.4	n.a.
DUP-37	UR	1345	10	52.93	3.55	13.02	12.50	0.16	4.81	8.20	2.63	1.70	0.50	1.00	349	35	52	145	109	28	47	745	36	284	25.1	574
DSM-30	SM	1350	10	51.74	3.61	13.22	12.87	0.17	5.01	8.28	2.89	1.71	0.49	0.26	360	38	62	102	104	27	36	657	37	279	25.0	551

Table 1: XRF major and trace element data for basaltic lavas of the high-Ti/Y Urubici magma type, and selected low-Ti Palmas rhyolites

Bindia Air Flow Sloy TOJ Air C N S N S N S N																																	
Simple Flow Slo, TIO, AlO, FaO, Maro Foo Al C Al C Al C Al C Al	Ba	550	600	600	576	682	609	608	519	670	615	776	673	684	628	753	718	715	732	615	727	n.a.	707	738	691	674	630	592	595	570	563	1137	1115
Bindia Rundia Flow S10, TiO, Ai,O, Flow S10, S11, S2 S3 S3 <td>ЧN</td> <td>23.8</td> <td>27.5</td> <td>26.8</td> <td>26.6</td> <td>28.5</td> <td>28.0</td> <td>30.4</td> <td>25.1</td> <td>27.0</td> <td>28.2</td> <td>38.4</td> <td>32.9</td> <td>32.3</td> <td>33.2</td> <td>32.3</td> <td>35.7</td> <td>32.3</td> <td>34-5</td> <td>32.9</td> <td>33.1</td> <td>32.8</td> <td>31.5</td> <td>31-4</td> <td>32.4</td> <td>27.7</td> <td>27.2</td> <td>28.5</td> <td>23-9</td> <td>24-5</td> <td>23.7</td> <td>46.8</td> <td>47.9</td>	ЧN	23.8	27.5	26.8	26.6	28.5	28.0	30.4	25.1	27.0	28.2	38.4	32.9	32.3	33.2	32.3	35.7	32.3	34-5	32.9	33.1	32.8	31.5	31-4	32.4	27.7	27.2	28.5	23-9	24-5	23.7	46.8	47.9
Bindia Rand Hi, Flow St0, TU, Ai,0, Field, Min C Ni C Ni C Ni Cir Zir G Ni SMM 3 Min Flow SU, TU,0, Ai,0, Field, Min Yir Co Ni Zir Zi	Zr	282	300	307	305	309	312	311	300	307	298	391	326	328	328	326	339	330	344	336	339	334	343	351	346	332	312	306	272	282	259	482	499
Bindia Flow Sloi TiO2 Ajo2 FayO MarO MarO FO Ni Co Ni Ni <th>≻</th> <th>36</th> <th>36</th> <th>39</th> <th>40</th> <th>38</th> <th>41</th> <th>38</th> <th>37</th> <th>44</th> <th>38</th> <th>48</th> <th>41</th> <th>42</th> <th>40</th> <th>46</th> <th>40</th> <th>38</th> <th>42</th> <th>39</th> <th>41</th> <th>41</th> <th>40</th> <th>41</th> <th>39</th> <th>39</th> <th>38</th> <th>36</th> <th>37</th> <th>39</th> <th>36</th> <th>48</th> <th>50</th>	≻	36	36	39	40	38	41	38	37	44	38	48	41	42	40	46	40	38	42	39	41	41	40	41	39	39	38	36	37	39	36	48	50
ample Found Flow Slo1 Tuo, Slo1 V Co<	Sr	669	764	764	673	882	829	706	951	1050	832	552	737	884	840	892	930	875	833	707	608	938	789	879	662	812	869	826	701	690	372	687	600
Binnel Rund Ninol Ninol <th< th=""><th>Rb</th><th>33</th><th>36</th><th>30</th><th>39</th><th>47</th><th>20</th><th>41</th><th>38</th><th>45</th><th>24</th><th>72</th><th>44</th><th>21</th><th>21</th><th>39</th><th>46</th><th>41</th><th>43</th><th>47</th><th>59</th><th>19</th><th>43</th><th>31</th><th>41</th><th>45</th><th>37</th><th>58</th><th>45</th><th>47</th><th>41</th><th>113</th><th>109</th></th<>	Rb	33	36	30	39	47	20	41	38	45	24	72	44	21	21	39	46	41	43	47	59	19	43	31	41	45	37	58	45	47	41	113	109
Bannola Aludi Flow Silo Aludi Flow Silo Aludi Col Ni Col<	Ga	26	21	26	30	29	26	25	25	26	25	26	27	26	29	32	22	30	20	27	24	27	24	24	25	n.a.	26	26	27	26	27	27	27
Samuple Foury Kit. Flow Sloi Tup. Al. Mark Kit. Flow Sloi Tup. Mark	Zn	111	118	107	105	112	114	125	117	109	119	137	121	119	109	114	110	110	129	125	130	125	114	111	133	n.a.	109	122	105	124	113	109	114
Samueli Raad Ait. Flow SiO_1 TO_2 Ai_2 FaO_1 Na_1 Na_1 <	Cu	152	136	267	134	182	182	177	185	179	140	72	160	196	201	170	191	185	192	188	190	196	169	181	165	n.a.	165	152	188	200	165	40	44
Samuple Read Alt. Flow Slo_1 To_2 Al_0 Fe_0 MnO MgO Cal Ne_0 Slo_1 V Cal J(m) (m) Slo_1 To_2 Al_0 Fe_0 MnO MgO Cal Ne_0 Slo_1 V Col V Col Slo	ï	54	60	58	54	61	56	42	64	52	61	25	65	65	51	49	59	49	59	54	55	48	51	56	45	57	62	54	37	34	39	4	ო
Sample Road Alt. Flow SiO ₂ TO ₂ Al ₁ O ₁ Fe ₀ O ₁ MnO MgO CaO Na ₂ O P_0 LOI V (m) (m) SiO 3137 10 5141 357 1343 1263 017 514 327 1343 1263 017 616 326 323 3223 1343 1266 140 57 324 279 177 0.66 0.66 326 323 2007-33 UR 1406 11 5526 326 1341 140 11 6163 326 1341 140 143 126 140 146 32 324 323 323 323 323 323 323 323 323 323 323 324 324 327 337 324 343 325 344 324 324 325 337 324 323 337 337 337 337 337	Co	43	35	36	36	35	36	40	n.a.	40	42	n.a.	45	38	39	40	39	36	45	38	39	n.a.	39	39	n.a.	n.a.	38	n.a.	32	n.a.	37	n.a.	17
Sample Road Alt. Flow Silo, TIO, Al,O, Fe,O, MnO MgO CaO Na/O Fo,O Pa/O Pa/O (m) (m) (m) Silo TIO, Silo, TIO, Silo, TIO, Filo, Fil	>	355	357	343	324	326	358	379	337	342	358	257	398	371	364	341	361	362	375	393	379	n.a.	370	358	353	349	349	378	387	373	407	119	122
Sample Road Al. Flow SiO ₂ TO ₂ Al ₂ O ₁ Fe ₂ O ₁ MnO GaO Na ₂ O K ₂ O P ₁ O ₆ Im) Im) Im) SiO ₂ TO ₂ TO ₂ Fe ₂ O ₁ MnO GaO Na ₂ O K ₂ O P ₁ O ₆ SSM-40 SM 1375 10 51-46 3-57 13-43 12-63 0-11 4-73 2-70 1-10 0-56 SSM-41 SM 1405 11 52-28 3-37 13-16 1-10 4-71 8-42 2-39 1-10 0-56 DUP-35 UR 1435 11 52-36 3-37 13-16 1-10 4-76 8-32 2-37 1-10 0-56 DUP-35 UR 1430 11 52-36 3-36 1-3-36 1-11 2-36 1-30 0-50 DUP-35 UR 1430 11 52-32 3-36 1-30 1-31 1-30 1-31 1-30 <th>LOI</th> <th>96-0</th> <th>0-69</th> <th>0.26</th> <th>0.57</th> <th>1.05</th> <th>1.06</th> <th>0.94</th> <th>1.21</th> <th>1.54</th> <th>0.63</th> <th>0.50</th> <th>0.10</th> <th>1.10</th> <th>1.94</th> <th>1.26</th> <th>0.96</th> <th>0.72</th> <th>0.33</th> <th>0.83</th> <th>1.76</th> <th>0.72</th> <th>0.91</th> <th>0.91</th> <th>0.49</th> <th>1.12</th> <th>0.62</th> <th>0.73</th> <th>0.75</th> <th>1-65</th> <th>0.86</th> <th>0-87</th> <th>0.88</th>	LOI	96-0	0-69	0.26	0.57	1.05	1.06	0.94	1.21	1.54	0.63	0.50	0.10	1.10	1.94	1.26	0.96	0.72	0.33	0.83	1.76	0.72	0.91	0.91	0.49	1.12	0.62	0.73	0.75	1-65	0.86	0-87	0.88
Sample Road Att. Flow Slo ₃ TuO ₃ A ₁ O ₃ Fe ₃ O ₃ MnO MgO CaO Na ₃ O K ₃ O SSM-20 SM 1375 10 51-48 327 13-43 12-80 13-7 15-1 SSM-20 SM 1475 10 51-48 327 13-12 13-17 14-7 17-7 17-1 SSM-20 SM 1405 11 52.27 33-1 13-12 13-10 14-3 17-1 17-1 SUP-36 UR 1405 11 52.28 33-1 13-12 13-0 14-4 17-1 17-1 SUP-31 UR 1400 111 52.48 37-3 13-48 14-40 17-466 8-32 24-1 17-7 SUM-41 SM 1400 111 51-48 37-3 13-64 14-60 17-7 4-66 8-32 24-1 17-7 SUM-41 SM 1400 11	P_2O_5	0.55	0.49	0.58	0.51	0.56	0.59	0.59	0.52	0-53	0.49	0.70	0.58	0.55	0.58	0.56	0.60	0.59	0.58	0.59	0.60	0.58	0.61	0.62	0.60	0.60	0.57	0.55	0.51	0.53	0.50	0.70	0.69
Sample Road Air. Flow SiO ₂ TO ₂ Ai,O ₃ Fa ₂ O ₃ MnO MgO CaO Ma ₂ O SSM-29 SM 1375 10 51-41 357 13-43 12-63 0.17 6-14 8-30 3-17 SSM-40 SM 1406 10 51-66 3-62 13-21 13-11 0.17 4-68 8-33 2-67 SSM-40 SM 1406 11 52-28 3-30 13-12 13-01 0.17 4-68 8-33 2-67 SUP-33 UR 1436 11 52-28 3-76 13-40 0.17 4-68 8-33 2-67 SUP-33 UR 1430 11 52-28 3-76 13-40 0.17 4-68 8-33 2-67 SM-41 SM 1430 11 52-18 3-76 13-41 0.17 4-68 8-33 2-61 SM-41 SM 13-69 13-6 13-61<	K ₂ 0	1.57	1.72	1.70	1.80	1.11	1.35	1.83	1.72	1.77	1.68	2.86	2.22	1.08	1.00	1.21	1.31	1.47	1.88	2.03	2.29	1.44	2.11	1.84	2.36	1.36	1.64	1.54	1.80	1-87	1.71	4.12	4-01
Bample Road Att. Flow SiO ₂ TO ₂ A ₃ O ₃ Fe ₂ O ₃ MnO MgO CaO Im) m) m) mod siO ₂ TO ₂ A ₃ O ₃ Fe ₂ O ₃ MnO MgO CaO DSM-20 SM 1375 10 51-66 362 13-26 10 4-34 8-36 DUP-35 UR 1376 11 52-13 376 13-16 13-16 14-36 8-33 DUP-36 UR 1405 11 52-13 376 13-16 13-16 4-36 8-36 DUP-31 UR 1436 11 52-13 374 13-44 1-66 8-32 DUP-31 UR 1436 11 52-13 374 13-44 1-66 8-33 DUP-31 UR 1446 11 51-13 1-16 1-46 8-33 8-44 DUP-31 UR 1446 11 51-13 1-16	Na ₂ O	3.17	2.79	2.57	2.66	2.70	2.68	2.56	2.23	2.41	2.61	3.29	2.67	2.62	2.30	3.13	3.25	2.98	2.72	2.89	3.05	2.55	2.91	2.50	2.82	3.07	3.17	2.95	3.01	2.72	2.94	3.84	3.89
Sample Road Att. Flow SiO ₂ TIO ₂ A ₃ O ₃ Fe ₂ O ₃ MnO MgO DSM-20 SM 1375 10 51-43 357 13-43 12-63 0.71 4.71 DVP-36 UR 1305 11 52-01 375 13-263 0.71 4.71 DVP-36 UR 1405 11 52-27 375 13-16 13-10 14-65 DVP-36 UR 1405 11 52-27 375 13-16 13-10 14-65 DVP-33 UR 1435 11 52-43 376 12-91 13-06 14-66 DVP-31 UR 1430 11 52-44 376 13-16 14-06 14-66 DVP-31 UR 1446 11 52-44 374 13-46 14-66 14-66 DVP-31 UR 1446 13 51-34 13-46 14-66 14-66 DVP-31 UR </th <th>CaO</th> <th>8.38</th> <th>8.42</th> <th>8.30</th> <th>7.93</th> <th>8.31</th> <th>8.38</th> <th>8.23</th> <th>8.48</th> <th>8.33</th> <th>8.02</th> <th>6.44</th> <th>7.68</th> <th>8.67</th> <th>8.99</th> <th>8.37</th> <th>8.24</th> <th>8.33</th> <th>8-07</th> <th>7.87</th> <th>7.27</th> <th>8.32</th> <th>7.62</th> <th>7.85</th> <th>7.27</th> <th>8.21</th> <th>8.18</th> <th>8.40</th> <th>8-01</th> <th>7.93</th> <th>8.04</th> <th>4.11</th> <th>4.11</th>	CaO	8.38	8.42	8.30	7.93	8.31	8.38	8.23	8.48	8.33	8.02	6.44	7.68	8.67	8.99	8.37	8.24	8.33	8-07	7.87	7.27	8.32	7.62	7.85	7.27	8.21	8.18	8.40	8-01	7.93	8.04	4.11	4.11
Sample Road Alt. Flow SiO ₂ TiO ₂ Al ₂ O ₃ Fe ₂ O ₃ MnO SSM-40 SM 1375 10 51-41 3:57 13:43 12:68 0:17 SSM-40 SM 1405 10 51-66 3:62 13:27 13:13 0:17 SSM-40 SM 1405 11 53:01 3:76 12:48 0:17 SUP-35 UR 1405 11 52:27 3:75 13:40 0:17 SUP-31 UR 1435 11 52:28 3:79 13:30 0:17 SUP-33 UR 1435 11 52:85 3:79 13:30 0:17 SUM-21 SM 1400 11 51:8 3:73 14:0 0:17 SUM-21 SM 1400 11 51:8 3:73 13:40 0:17 SUM-21 SM 1400 11 51:8 3:74 13:40 0:17	MgO	5.14	4.71	4.34	4.65	4.36	4-50	4.65	4-50	4-60	4.48	3.18	4.40	4-53	4-50	4.35	4.29	4.42	4.39	4.35	4.32	4.42	4.12	4.06	4-03	4-07	4.45	4.56	4-30	4.21	4.37	2.05	1.92
Sample Road Alt. Flow SiO2 TIO2 AlgO3 Fe2O3 Im) <	MnO	.17	.17	0.19	.17	0.19	0.16	0.17	0.15	0.16	0.19	.17	0.19).18	.17	0.19	0.19	.17	.17	0.16	.17	.17	0.16	.17	0.16	0.16	.17	.17	0.19	.17	0.17	0.14	0.15
Sample Road Alt. Flow SiO2 TiO2 Al ₂ O3 F (m) (m) (m) 51-41 3-57 13-43 1 DSM-29 SM 1375 10 51-41 3-57 13-43 1 DUP-35 UR 1365 11 53-01 3-76 12-86 1 DUP-35 UR 1365 11 53-01 3-76 12-91 1 DUP-34 UR 1405 11 52-27 3-76 12-91 1 DUP-33 UR 1435 11 52-36 3-90 13-12 1 DUP-33 UR 1435 11 52-36 3-76 12-91 1 DUP-31 PE 1390 11 51-30 3-76 12-91 1 13-94 1 DUP-31 PE 1390 11 51-30 3-76 12-91 1 12-91 1 12-91 13-91 13-91	e2O3 1	2.63 (3.13	2.68 (3.10 (3.01 (3.25 (4-04 (3.09 (2.98 (3.45 (1.53 (3.83 (4-02 (4.16 (3.94 (3.57 (3.49 (3.50 (3.40 (3.43 (3.92 (3.39 (3.36 (3.00	3.70 (3.36 (3.20 (3.87 (4.15 (4.05 (9.20 (9.38
Sample Road Alt. Flow SiO2 TIO2 A Im) im) im) im) im) im) in) in) SSM-29 SM 1375 10 51.41 3.57 i i SSM-40 SM 1375 10 51.41 3.57 i i DUP-35 UR 1365 11 53.01 3.76 i <t< td=""><td>l₂O₃ F</td><td>3.43 1</td><td>3.27 1</td><td>2.86 1</td><td>3.16 1</td><td>3.12 1</td><td>2.96 1</td><td>3.13 1</td><td>2.91 1</td><td>3.48 1</td><td>3.54 1</td><td>3.94 1</td><td>2.87 1</td><td>2.96 1</td><td>2.98 1</td><td>2.71 1</td><td>2.91 1</td><td>2.92 1</td><td>3.13 1</td><td>3.14 1</td><td>3.17 1</td><td>3.27 1</td><td>3.07 1</td><td>3.10 1</td><td>2.64 1</td><td>3.11 1</td><td>3.33 1</td><td>3.71 1</td><td>3.24 1</td><td>3.32 1</td><td>3.32 1</td><td>4-06</td><td>3.86</td></t<>	l₂O₃ F	3.43 1	3.27 1	2.86 1	3.16 1	3.12 1	2.96 1	3.13 1	2.91 1	3.48 1	3.54 1	3.94 1	2.87 1	2.96 1	2.98 1	2.71 1	2.91 1	2.92 1	3.13 1	3.14 1	3.17 1	3.27 1	3.07 1	3.10 1	2.64 1	3.11 1	3.33 1	3.71 1	3.24 1	3.32 1	3.32 1	4-06	3.86
Sample Road Alt. Flow SiO2 T Im) (m) (m) 510.4 3 DSM-29 SM 1375 10 51.41 3 DSM-29 SM 1375 10 51.41 3 DUP-35 UR 1365 11 53.01 3 DUP-35 UR 1365 11 52.27 3 DUP-33 UR 1405 11 52.28 3 DUP-33 UR 1435 11 51.28 3 3 DUP-33 UR 1435 11 51.64 3 3 3 DUP-33 UR 1435 11 51.64 3 3 3 3 DUP-31 PE 1336 11 51.63 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3	02 A	57 1:	62 1:	76 1:	75 1:	80 1:	79 1:	72 1:	76 1:	88	74 1:	95 1:	21 1:	15 12	13 1	16 1	20 1:	20 1:	33 1:	23 1:	21 1:	20 1:	16 1	15 13	04 1:	11	88	66 1;	51 1:	48 1:	49 1:	01 1/	03 1:
Sample Road Alt. Flow S (m) (m) (m) 5 (m) 5 DSN-29 SM 1375 10 5 5 SSW-40 SM 1375 10 5 DUP-35 UR 1365 11 5 DUP-35 UR 1365 11 5 DUP-31 UR 1405 11 5 DUP-33 UR 1435 11 5 DUP-33 UR 1435 11 5 DUP-33 UR 1420 11 5 DUP-31 PR 1420 11 5 DUP-31 PR 1420 13 5 DUP-31 PR 1450 13 5 DUP-31 PR 1445 13 5 DUP-31 PR 1445 14 5 DUP-32 SM 1445 14 5 <td< th=""><th>i0₂ T</th><th>1.41 3.</th><th>1-66 3.</th><th>3.01 3.</th><th>2.27 3.</th><th>2.85 3.</th><th>2.35 3.</th><th>1.08 3.</th><th>2.64 3.</th><th>1-87 3.</th><th>1.79 3.</th><th>4-93 2.</th><th>1.35 4.</th><th>1.23 4.</th><th>1.20 4.</th><th>1.38 4.</th><th>1-44 4.</th><th>1.42 4.</th><th>1.23 4.</th><th>1.34 4.</th><th>1.49 4.</th><th>1.13 4.</th><th>1-83 4</th><th>2.35 4.</th><th>3-09 4-</th><th>1-60 4-</th><th>1.28 3.</th><th>1.25 3.</th><th>1-56 3.</th><th>1-61 3.</th><th>1.43 3.</th><th>9.78 2.</th><th>9.96 2.</th></td<>	i0 ₂ T	1.41 3.	1-66 3.	3.01 3.	2.27 3.	2.85 3.	2.35 3.	1.08 3.	2.64 3.	1-87 3.	1.79 3.	4-93 2.	1.35 4.	1.23 4.	1.20 4.	1.38 4.	1-44 4.	1.42 4.	1.23 4.	1.34 4.	1.49 4.	1.13 4.	1-83 4	2.35 4.	3-09 4-	1-60 4-	1.28 3.	1.25 3.	1-56 3.	1-61 3.	1.43 3.	9.78 2.	9.96 2.
Sample Road Alt. Fld Imple Road Alt. Fld DSN-29 SM 1375 10 DSN-29 SM 1375 10 DUP-35 UR 1365 11 DUP-35 UR 1365 11 DUP-35 UR 1405 11 DUP-33 UR 14365 11 DUP-33 UR 1420 11 DUP-33 UR 14365 11 DUP-33 UR 1430 11 DUP-12 PE 13365 11 DUP-31 UR 1445 12 DUP-13 PE 1420 13 DUP-31 PE 1445 14 DUP-28 UR 1450 14 DUP-28 UR 1445 14 DUP-28 UR 1446 15 DUP-28 UR 1446 15 DUP-28	s S	2	D	ß	Ð	ß	Ð	D	D	Ð	D	Ð	Ð	Ð	Ð	Ð	ß	Ð	Ð	D	Ð	Ð	Ð	Ð	Ð	5 2	Ð	ß	Ð	Ð	ß	D	D
Sample Road Alt. DSM-29 SM 1375 DSM-29 SM 1375 DSM-29 SM 1375 DSM-29 SM 1375 DUP-35 UR 1365 DUP-35 UR 1405 DUP-35 UR 1405 DUP-31 UR 1435 DUP-33 UR 1436 DUP-33 UR 1436 DUP-33 UR 1420 DUP-31 PE 1386 DUP-33 UR 1406 DUP-31 UR 1420 DUP-31 PE 1420 DUP-33 UR 1450 DUP-34 SM 1465 DUP-37 PE 1420 DUP-38 UR 1450 DUP-39 UR 1450 DUP-38 UR 1450 DUP-38 UR 1450 DUP-38 UR 1450	FIG	10	10	11	1	11	11	11	11	11	11	12	13	13	13	13	14	14	14	14	14	14	15	15	15	15	16	17	18	18	18	19	19
Sample Roac DSM-29 SM DUP-35 UR DUP-36 UR DUP-36 UR DUP-38 UR DUP-31 UR DUP-31 UR DUP-12 PE DUP-13 PL DUP-13 PL DUP-13 PL DUP-13 PR DUP-29 UR DUP-28 UR DUP-28 VR DUP-28 VR DUP-20 PE DUP-28 VR DUP-20 PE DUP-28 VR DUP-28 VR DUP-38 VR DUP-	I Alt. (m)	1375	1405	1365	1405	1420	1435	1385	1390	1400	1295	1420	1450	1465	1375	1420	1505	1520	1490	1445	1450	1355	1540	1540	1465	404	1570	1720	1755	1755	1755	1765	1765
Sample Sample DUP-35 DUP-36 DUP-36 DUP-36 DUP-36 DUP-31 DUP-12 DUP-13 DUP-13 DUP-13 DUP-13 DUP-13 DUP-13 DUP-13 DUP-13 DUP-20 DUP-13 DUP-13 DUP-13 DUP-13 DUP-13 DUP-13 DUP-13 DUP-13 DUP-13 DUP-13 DUP-13 DUP-13 DUP-13 DUP-13 DUP-13 DUP-13 DUP-13 DUP-20 DUP-20 DUP-20 DUP-20 DUP-20 DUP-20 DUP-20 DUP-20 DUP-20 DUP-32 DUP-13 DUP-20 DSM-20 DSM-20 DSM-20 DSM-20 DSM-20 DSM-20 DSM-20 DSM-10 DSM-10 DSSM-10 D	Roac	SM	SM	UR	UR	UR	UR	PE	SM	SM	00	SM	UR	UR	PE	PE	UR	UR	PE	SM	SM	RR	SM	SM	RR	ES	SM						
	Sample	DSM-29	DSM-40	JUP-35	30P-36	JUP-34	DUP-33	DUP-12	JSM-27	DSM-41	JW-24	JSM-24	JUP-32	JUP-31	11-9UC	JUP-13	OUP-29	DUP-28	DUP-20	DSM-22	JSM-21	JRR-12	DSM-17a	dT1-MSC	30-AAC	ES-537	3SM-16	01-MSC	d20-MSC	70-MSC	60-WSC	DSM-04	80-MSC

PEATE et al. HIGH-Ti/Y PARANÁ FLOOD BASALTS AND DUPAL-TYPE MANTLE

Table 1: continued

Sample	Road	Alt. (m)	Flow	SiO2	TiO ₂	AI ₂ 0 ₃	Fe ₂ O ₃ M	о ^г	go Ca	0 Na	0 K20	P ₂ 0	FOI	>	с	īz	Cu	Zn	Ga	dR	-		ž	ä	"
AV-1668	A	- 1151	h/h	51.77	3.58	13.28	13.71 0.	19 3.	90 7·8	30 3-1	3 1-93	3 0-71	n.a.	309	31	22	100	108	25	49 8	29 4	12 2	76 29	-7 65	25
GO-1053	GO	- 137	h/d	50-41	3.91	13.32	14.36 0.	17 4-	31 8-2	20 2.7	3 1.7/	t 0-50	n.a.	342	42	45	132	109	27	36	13 3	87 2	77 26	.7 57	1
RO-1185	RO	- 437	h/d	51.83	3.47	13.40	13.55 0.	19 3.	92 7.E	55 3.2	9 2.0(3 0.75	n.a.	332	n.a.	24	n.a.	n.a.	n.a.	50 7	70 3	39 2	89 26	-7 68	32
DAB-01	AB	1200	dyke	51.40	3.67	13-52	13.02 0.	17 4.	33 7.1	18 3-5	3 2.05	0.51	1.36	345	n.a.	55	152	125	26	51 8	11 3	39 3	09 28	-2 66	51
DSM-26	SM	1435	dyke	52.10	3.41	13.89	12.18 0.	17 4-	39 8.3	39 2.7	7 1-87	0.46	1.40	345	n.a.	65	182	102	23	39 7	83	36 2	95 25	·2 51	6/
DW-22	Aiure	Ι	dyke	50-85	3.57	13.14	14.50 0.	19 4.	47 8.(04 2.9	4 1.78	3 0-53	0.39	387	34	45	138	119	25	55	33	38	79 27	-5 60	00
DRG-6	Aiure	Ι	dyke	50.55	3.66	13.19	14.84 0.	19 4.	50 7.6	37 2.9	0 1-97	0.53	0.38	n.a.	n.a.	39	139	123	25	55	77 4	F0 2	96 27	.7 n.	a.
ST93-1	FD		dyke	51.43	4.13	13.28	13.67 0.	17 4.	15 8.()9 2·5	5 1-92	2 0-61	0.46	360	39	51	190	128	26	40	86 4	12 3	52 31	-5 70	60
ST93-2	FD	I	dyke	51.68	4-03	13.33	13.61 0.	17 4.	12 7.7	16 2.7	0 1-99	9 0-62	0.45	345	38	53	195	120	28	54 8	16 4	12 3	52 31	.6 7,	14
ST93-3	FD	Ι	dyke	51.86	3.68	13.38	12.88 0.	16 4-	36 7.4	t3 2.7	9 1-6(§ 0-50	1.11	346	39	68	164	127	23	48 6	20 4	е 01	18 28	-9	01
ST93-4	FD	I	dyke	52.16	3.66	13.34	12.96 0.	16 4.	31 8.3	35 2.6	5 1.62	2 0.49	0.03	350	39	69	170	128	24	31 7	46 4	01 3	20 27	-9 56	76
ST93-5	FD	I	dyke	54-82	2.63	13.08	13.44 0.	20 2.	39 6:	53 3-0	7 2.5(1.05	1.33	127	25	٢	38	156	27	58	19 6	31 4	33 36	-1 72	21
Low-Ti Pa	lmas rh	yolites																							
DGB-61	GB	1430	unit F	69-21	0.88	12.73	5.66 0.	10 1.	22 2·F	53 2.4	7 4-95	0.26	0.72	n.a.	n.a.	10	53	72	19	188	13 4	14 2	56 22	Ľ.	a.
DGB-59	GB	1475	unit F	67-44	66-0	13.29	6.39 0.	11 1.	37 3.7	76 3-5	0 2.88	3 0-27	2.28	42	n.a.	2	LT	78	19	157 1	56 4	12 2	55 22	22	73
DGB-58	GB	1495	unit F	67-52	0.98	13.29	6.37 0.	11 1.	33 3.6	3-5	3 2.92	0.26	2.21	n.a.	n.a.	œ	72	81	18	171 1	64 4	12 2	65 21	Ľ.	a.
ES-312	ES	629	unit F	67.15	0.98	14.02	5-86 0.	12 1.	33 3.7	6 2.6	9 3.27	0.31	2.39	n.a.	n.a.	œ	54	93	22	75 1	60 4	16 2	78 23	Ľ.	a.
ES-216a	ES	725	unit G	67-80	1.08	12.90	6.43 0.	11 0.	31 3.(01 3.2	3 4.28	3 0.32	0.47	n.a.	n.a.	6	96	17	17	170 1	45 4	17 2	80 23	Ľ.	a.
ES-216b	ES	725	unit G	69.65	1.00	12.12	6.03 0.	10 0.	36 2.9	99 3-2	0 3.7E	0.31	0.47	n.a.	n.a.	œ	96	68	18	147 1	35 4	14 2	53 20	ц.	a.
ES-150	ES	791	unit G	67-83	1-07	12.86	6.36 0.	12 1.	04 3 [.]	19 3-4	3.75	0.31	0.48	n.a.	n.a.	00	98	72	14	160 1	48 4	16 2	69 21	Ľ.	a.
ES-123	ES	818	unit G	68-33	1-03	12.72	6.08 0.	09 1.	11 3.5	29 3.3	4 3.70	0.31	0-69	n.a.	n.a.	00	86	76	19	166 1	49 4	17 2	69 21	Ľ.	a.
ES-105	ES	836	unit G	67-80	1.06	12.83	6.42 0.	11 1.	10 3.5	30 3-4	1 3.67	0.32	0.48	n.a.	n.a.	6	110	76	20	158 1	47 4	16 2	65 21	Ľ.	a.
ES-087	ES	854	unit G	67-61	1.05	12.77	6.39 0.	12 1.	17 3.	18 3-1	9 4.22	0.31	0-52	n.a.	n.a.	10	104	82	19	175 1	31 4	15 2	60 20	Ċ.	a.
X-ray flu X-ray flu plus estii existing profile Ic 49.58°W/ Borehole 51.154°W For defin Anita Ga	oresce mates trigonc ocation c(b/h) 1 /). FD, iltion o iltion o	nce ma of analy metric s: PE, F s: PE, F s: PE, F for analy for analy funit funit funit funit funit funit	ytical L ytical L surve Perico ranco is: AV, J opolis c and u	d trace uncerts y poin (28.14 (28.07 Agua c dyke unit G unit G	e elem ainties its to a °S, 49 be Valk Swarn rhyol ilmas	ent da i, give an esti 66°W .30°W .30°W anca (anca (rhyoli rhyoli	ita. Maj n by Pe. mated); UR, R); RR, R 23.094°S 2010 23.094°S Jnit F rl tes (Ga	or elei ate & Jrubic o Ruf 3, 51.9 1lecteo rland	hent c Hawke ion of i (28.0 no (27 13°W) 13°W) i from <i>s</i> belc <i>s</i> t al.,	arti 188°S, 45 m 8°S, 45 m 8°S, 45 7.91°S, 7.	(1996) (1996) (all al- .62°W 49.89° oncad do Re do Re the Câ	inalized (i). n.a. (i). SM, (); SM, (W); GI (W); GI (W	to 100 , not al s in tab s in tab B, Gua 609°S, to the e to Sul st al. (1	% on a nalysed le and ose-Mc ose-Mc rta-Bon rta-Bon rta-Bon subgro subgro subgro	volat - Alt., - Alt., - Alt., - Alt., - Alt. - Alt. - Alt. - Alt. - Alt. - Alt. - Alt. - Alt. - Alt., - Alt.,	Altitu Altitu a Igre in da io, Gá santa santa	e basis ide me, metres ija (28. a Serra alvão (2 catarir almas	, with asure abov 12°S, (28.3 (28.3 (28.3 (28.3 (28.3 (28.3 (28.3 (28.3)) (28.3	total d usin e sea e sea 49.46 6°S, 4 52, 52 ites, a ites, a	iron a g an <i>a</i> g an <i>a</i> (\\); A •\\); A •\10°V 10 km nd un	s Fe ₂ O Itimet LOI, _ LOI, _ Aiu V); ES, V south it G rl	(t). A (t). A (t). A (t) as c (t) as c (t) as c (t) (t) as c (t) (t) as c (t) (t) as c (t) (t) as c (t) (t) (t). A (t). A (t) as c (t) a	nalytic conjun in ignit irancas 3.35°S, eralda (eralda (es belc	al det ction F (27.9 (27.9 (27.9 (27.9 (27.9 (28.09 (28.09 (28.09) (28.09) (28.09) (28.09) (28.09) (28.09) (28.09) (28.09) (28.09) (28.00) (28.00) (28.00) (28.00) (28.00) (28.00) (28.00) (28.00) (28.00) (28.00) (27.90) (27	ails, with with coad of or coad

Sample	Unit	Flow	Sc	Cs	La	Ce	Nd	Sm	Eu	Tb	Yb	Lu	Hf	Та	Pb	Th	U
DUP-02	В	3	26.8	0.5	43.2	93.8	50.8	11.2	3.58	1.59	3.18	0.50	8.12	1.98	n.a.	4.71	1.15
DSM-36	А	6	n.a.	7.51	n.a.	n.a.											
DSM-35	А	7	34.0	0.4	37.8	80.4	49.3	10.5	3.19	1.44	2.95	0.46	7.08	1.71	5.58	4.16	1.77
DSM-34	А	8	23.8	1.2	39.2	79.3	43.7	9.26	2.91	1.23	2.33	0.38	6.67	1.70	6.81	4.49	1.04
DSM-30	С	10	26.0	n.d.	37.2	83.3	46.3	9.81	3.38	1.47	2.77	0.44	7.55	1.75	5.67	3.77	0.98
DUP-37	С	10	20.1	1.3	37.1	79.4	45.9	10.4	3.21	1.40	2.71	0.42	7.18	1.68	n.a.	3.62	1.27
DUP-35	С	11	28.9	0.3	42.5	90-4	54.3	11.6	3.53	1.54	2.99	0.44	7.97	1.92	n.a.	4.25	1.34
DSM-27	С	11	25.0	1.2	38.0	83-2	47.7	10.2	3.43	1.51	2.76	0.41	7.61	1.78	n.a.	3.86	0.76
DSM-24	Е	12	19.7	0.6	60.5	126	63.5	12.6	4.03	1.72	3.66	0.56	9.32	2.53	9.48	7.39	2.01
DUP-32	D	13	33.0	0.5	49.0	108	60.6	12.8	3.98	1.83	3.29	0.49	8.70	2.30	n.a.	5.20	1.36
DSM-17b	D	15	24.6	1.1	48.6	110	62.4	13.0	4.02	1.74	3.12	0.49	9.30	2.35	n.a.	5.11	1.86
DSM-10	Е	17	25.2	1.7	42.0	92.7	51.6	10.6	3.53	1.44	2.76	0.42	7.78	2.03	6.49	4.78	1.13
DSM-05b	Е	18	27.0	0.6	39.8	87.5	49.0	10.3	3.41	1.47	3.12	0.47	7.30	1.89	7.31	5.02	1.08
DSM-07	Е	18	27.0	0.6	40.7	88.5	50.8	10.2	3.37	1.45	3.11	0.48	7.44	1.91	7.35	5.20	1.37
DSM-04	Е	19	13.8	1.8	79.6	173	82.6	15.3	4.42	1.85	3.63	0.55	12.4	3.32	14.2	12.1	3.00
DSM-08	Е	19	14.1	1.8	81.1	180	85.6	15.9	4.65	2.00	3.92	0.56	12.3	3.20	n.a.	12.8	2.71
DSM-26	—	dyke	24.5	n.d.	39.9	93·1	51.7	10.8	3.56	1.59	2.83	0.44	7.59	1.90	n.a.	4.47	1.40
WS observed	—	_	_	1.03	25.4	57·8	33.5	7.20	2.29	1.14	2.56	0.41	4.95	1.30	n.a.	3.08	0.96
(<i>n</i> = 7) 1σ	_	_	—	0.04	0.8	0.7	1.0	0.24	0.04	0.05	0.08	0.01	0.09	0.02	n.a.	0.06	0.15
WS expected	_	—	—	1.03	25.5	57.5	32.9	7.27	2.25	1.09	2.54	0.39	4.93	1.26	n.a.	3.05	0.90

Table 2: Additional trace element data for Urubici magma type samples

All data by instrumental neutron activation analysis at the Open University, following the procedure of Potts *et al.* (1985), except for Pb, which was analysed by isotope dilution on a VG54E mass spectrometer at the Open University. Precision and accuracy of the neutron activation technique can be estimated from the standard deviations from repeat analyses of WS (in-house standard) and by comparing the observed values for WS with expected values from Potts *et al.* (1985). Isotope dilution measurement of BHVO-1 gave 2.04 ppm Pb. n.a., not analysed; n.d., not detected; *n*, number of analyses.

LOCAL FLOW STRATIGRAPHY Rationale of approach to flow correlation

Exposure is seldom continuous and mainly consists of scattered isolated outcrops with dense vegetation obscuring the areas between. This prevents visual correlation of individual flows between profiles. Physical characteristics of the lavas such as jointing habit, flow thickness and morphology, and petrographic textures, are of little value in flow correlation. Such features are often not unique to an individual flow and they can vary markedly within a single flow over a few kilometres, because they are controlled primarily by lateral variations in the local cooling conditions, palaeotopography, and by distance from the vent source. Most Paraná basalts are subaphyric with broadly similar phenocryst assemblages (Comin-Chiaramonti et al., 1988): even the high-Ti/ Y Urubici and low-Ti/Y Gramado basalts cannot be distinguished petrographically. The only petrographically distinct Urubici flow is the dacite flow (DSM-04,08) which has 30% phenocrysts, but it is found in one profile (the uppermost unit in the SM profile).

Instead, flows are distinguished and correlated from small variations in TiO₂ and MgO (Fig. 6) and from their relative stratigraphic position in each profile. TiO₂ is used because of its relative immobility during posteruption alteration (e.g. Wood et al., 1976) and because Urubici samples have a wide range of TiO₂ contents for a given MgO content (Fig. 6). We recognize a minimum of 19 Urubici flows in the São Joaquim area. Their correlation between profiles is shown in Fig. 7. A thin (<5 m) horizon of aeolian sandstone separates flows 6 and 7 and indicates a persistence of the arid desert climate that typifies the sediments of the underlying Botucatu Formation. Flow contacts are recognized in the field by the presence of a vesicular aa-type flow-top and a sharp transition to the more massive base of the subsequent flow. It is often difficult to determine whether these are true inter-flow contacts or represent an internal contact within a compound-type flow, given the discontinuous nature of the exposure. Although flows 9 and 10 are compositionally very similar, they are clearly two distinct flows because in the UR profile they are separated by a Gramado flow (Fig. 7). Thus, several of the flows

Sample	Unit	Flow	⁸⁷ Sr/ ⁸⁶ Sr measured	⁸⁷ Sr/ ⁸⁶ Sr initial	ε _{sr} initial	¹⁴³ Nd/ ¹⁴⁴ Nd measured	¹⁴³ Nd/ ¹⁴⁴ Nd initial	ε _{Nd} initial	²⁰⁶ Pb/ ²⁰⁴ Pb measured	²⁰⁷ Pb/ ²⁰⁴ Pb measured	²⁰⁸ Pb/ ²⁰⁴ Pb measured
DGB-09	А	1	0.70607	0.70557	14.9	n.a.	n.a.	n.a.	17.591	15.498	38.156
DUP-02	В	3	0.70544	0.70529	10.6	0.512449	0.51234	−2 ·7	17.603	15.494	38.140
DUP-04	А	5	0.70587	0.70561	15.1	0.512360	0.51225	-4.3	17.727	15.514	38.231
DSM-36	А	6	0.70628	0.70590	19.2	0.512276	0.51217	-5.9	17.543	15.515	38.201
DSM-35	А	7	0.70574	0.70550	13.6	0.512380	0.51227	-3.9	17.553	15.488	38-186
DUP-42	А	8	0.70572	0.70541	12.3	0.512395	0.51229	-3.6	17.538	15.464	38.032
DSM-34	А	8	0.70592	0.70572	16.8	0.512383	0.51227	-3.9	17.704	15.535	38.334
DSM-30	С	10	0.70547	0.70518	9.1	0.512403	0.51229	-3.5	17.517	15.486	38.039
DUP-37	С	10	0.70565	0.70531	10.8	0.512438	0.51232	-3.0	17.519	15.477	38.033
DUP-35	С	11	0.70527	0.70506	7.4	0.512407	0.51230	-3.4	17.464	15.458	38.026
DSM-27	С	11	0.70523	0.70502	6.7	0.512455	0.51235	-2.5	17.541	15.490	38.065
DSM-24	E	12	0.70644	0.70575	17.1	0.512328	0.51223	-4.8	18.040	15.543	38.367
DUP-32	D	13	0.70520	0.70488	4.8	0.512452	0.51234	-2·5	17.776	15.505	38.227
DSM-22	D	14	0.70519	0.70483	4.1	0.512469	0.51236	-2·2	17.788	15.496	38.131
DRR-06	D	15	0.70531	0.70499	6.3	0.512459	0.51235	-2·4	17.728	15.501	38.141
DSM-17b	D	15	0.70515	0.70496	5.9	0.512482	0.51237	-1.9	17.749	15.505	38.148
DSM-10	E	17	0.70547	0.70509	7.8	0.512405	0.51230	-3.4	17.960	15.556	38.367
DSM-05b	E	18	0.70593	0.70558	14.8	0.512424	0.51232	-3 ·1	18.138	15.573	38.452
DSM-07	E	18	0.70591	0.70554	14.1	0.512429	0.51233	-2.9	18.131	15.558	38-406
DSM-04	E	19	0.70746	0.70657	28.8	0.512360	0.51226	-4.1	18-253	15.584	38.612
DSM-08	E	19	0.70743	0.70645	27.1	0.512357	0.51226	-4.1	18.202	15.565	38.535
DSM-26	_	dyke	0.70538	0.70511	8.0	0.512455	0.51235	-2·4	17.668	15.504	38.210
DAB-03	_	dyke	0.70584	0.70543	12.6	0.512467	0.51236	-2·2	n.a.	n.a.	n.a.
DGB-59	unit F rhy	olite	0.72335	0.71793	190.0	0.512233	0.51213	-6.8	n.a.	n.a.	n.a.
DGB-61	unit F rhyd	olite	0.72879	0.71987	217.5	0.512166	0.51206	<u>-8</u> .1	n.a.	n.a.	n.a.

Table 3: Radiogenic isotope data for Urubici magma type samples, and selected low-Ti/Y Palmas rhyolites

All data measured on a Finnigan MAT261 mass spectrometer at the Open University. Sr and Nd isotopic fractionation corrections are ${}^{86}Sr/{}^{88}Sr = 0.1194$ and ${}^{144}Nd/{}^{146}Nd = 0.7219$. NBS 987 Sr standard gave a mean value of 0.710228 \pm 17 and Johnson & Mathey (J&M) Nd standard gave a mean value of 0.511864 \pm 14 during this study. Data presented relative to NBS 987 Sr 0.71025 and J&M Nd 0.51185 (equivalent to BCR-1 0.51265). Initial ratios calculated at 132 Ma. Sm/Nd is constant (0.212 \pm 0.006) for the basaltic samples (SiO₂ < 54 wt %), and this value was used in the calculation of initial Nd isotope data if Sm and Nd were not measured directly. Similarly, an average Sm/Nd value (0.209 \pm 0.006: Hawkesworth *et al.*, 1988; Garland *et al.*, 1995) was assumed for the rhyolite data. Measured Pb isotope ratios corrected for an average 0.8% per a.m.u. fractionation relative to recommended values for NBS 981 Pb standard (Todt *et al.*, 1996). Reproducibility of Pb standard data was ~0.07% for ${}^{206}Pb/{}^{204}Pb$ and ${}^{207}Pb/{}^{204}Pb$. Within-run errors (2 SE) on Sr–Nd–Pb sample data are less than or equal to the external uncertainties (2 SD) on the standards. Total procedure blanks were <2 ng for Sr and Nd, and <1 ng for Pb. n.a., not analysed.

recognized might represent multiple eruptions of compositionally similar magma, and this could account for some of the compositional scatter within groups of samples assigned to the same flow (Fig. 6).

Implications of flow correlations

Most flows are found in more than one sampled profile, but exceptions include the upper four flows at the SM profile (flows 16–19), which lie at a higher stratigraphic level than flows in the surrounding profiles, where the higher-level lavas have probably been removed by erosion. The andesitic flow 12 is found only at the SM profile, although from stratigraphic considerations it should also have been found at the RR, UR and PE profiles, and this suggests that it has a relatively restricted spatial extent. Flows 3, 4 and 5, found at the base of the PE profile below a NW–SE trending fault, are correlated with the uppermost flows at the GB profile to the south. These flows are not found in the neighbouring SM and UR profiles, although flows 4 and 5 are found in the AB profile, which has a lower elevation than the surrounding profiles and has been strongly affected by faulting (Figs



Fig. 4. Map of São Joaquim study area, showing the location of the sampled road profiles and the local physiography. The Serra Geral escarpment (dark grey line) marks an abrupt change in elevation between the coastal plain in the east and the high lava plateau in the west. The altitude of the top and base of each road profile are marked in metres above sea level. The basal contact of the lavas with the Botucatu Formation sandstones, if encountered, is indicated by an asterisk.

4 and 7). Flows 1 and 2 are found only near the base of the GB profile. The single Urubici flow found at the RA profile to the south (Bellieni et al., 1984) cannot be correlated with any flows found in the GB profile. At profiles SM and UR, flow 9 overlies three lower Urubici flows (flows 6, 7 and 8), whereas it is the lowest Urubici flow found at the RR profile and it lies directly on the sediments at the CO profile. Thus, the lavas in the São Joaquim region successively overstep the underlying sediments to the north-east, consistent with a pre-eruption topographic relief of at least 400 m (Fig. 7). Palaeotopography was also an important control on flow thickness; for example, the basal Gramado-type flow at the GB profile is up to 60 m thick and fills in an existing valley cut into the underlying sedimentary rocks (de Oliveira, 1956). The thickness of individual flows (average ~20 m) varies non-systematically across the area although significant uncertainties arise in the estimates for many flows because of the lateral distance between exposures: the road profiles covered distances of 5-15 km to sample vertical lava profiles of 400-750 m.

On a regional scale, the elevation of the base of the lava pile rises significantly in a direction parallel to the coast northwards along the Serra Geral escarpment (Peate et al., 1992) from 230 m at the RA profile to 1260 m at the CO profile, 100 km to the north (Fig. 4). It is not clear whether this feature is, to a large extent, an original topographic feature formed before the Paraná magmatism, or the result of subsequent differential uplift. Gallagher et al. (1994) argued from fission-track data for large-scale (i.e. several kilometres) uplift of the whole Serra Geral escarpment, although pre-existing topography was clearly important on a local scale in controlling the emplacement of lavas in the São Joaquim region. The borehole sample ES-537 (elevation 404 m) is compositionally identical to flow 15 (average elevation \sim 1500 m) (Fig. 6a). If they are the same flow, then this implies that the São Joaquim region has been uplifted by ~1 km relative to the ES borehole, which is located ~150 km to the west (Fig. 2). Further evidence for significant post-magmatism differential uplift comes from the correlation of compositionally distinct rhyolite units. The rhyolite flows capping the GB profile (1430–1500 m: Fig. 7) are compositionally similar to the unit F rhyolites of Milner *et al.* (1995*a*), which are found about 500 m lower (950–1000 m) in the RA profile, and the difference in elevation of the base of the lava pile between the two profiles is of a similar magnitude (~500 m: RA 230 m; GB 760 m) (S. Milner, personal comunication, 1997). In



the ES borehole (Fig. 2), rhyolites found at an elevation of ~630 m are similar to the unit F flows found ~350 m higher at the RA profile, and rhyolites from the upper part of the borehole (720–850 m) are compositionally similar to the unit G rhyolites of Milner *et al.* (1995*a*) found at 1080–1160 m in the RA profile (compositional data for the rhyolites from the GB profile and the ES borehole are in Table 1). Thus, post-magmatic uplift on the coastal Serra Geral escarpment, relative to the position of the inland ES borehole, is greatest at the GB profile (~1 km) and less at the RA profile to the south (~0·4 km).

SHALLOW-LEVEL EVOLUTION OF THE URUBICI MAGMA TYPE

Stratigraphic variations

The evolution of the Urubici magmas cannot be described by a single liquid line of descent from a homogeneous parental magma. This is evident both from the range in isotopic composition and from the fact that samples with, for example, 4.5 wt % MgO have a wide range in TiO₂ from 3.4 to 4.3 wt % (Fig. 6). Consideration of sequential variations in flow composition (Fig. 8) might help to unravel the nature of the processes controlling the evolution of the magmas. The Urubici flows are grouped into five units (A–E) to simplify subsequent discussions and to emphasize some of the general stratigraphic variations. The units have similar compositions based on groupings from Fig. 6 (MgO vs TiO₂) and isotope composition, and in general comprise stratigraphically adjacent flows (Fig. 8). A distinction is made between the

Fig. 5. Radiogenic isotope composition of the Urubici magma type, and comparison with the low-Ti/Y Gramado and Esmeralda magma types (Hawkesworth et al., 1988; Peate & Hawkesworth, 1996). Units A-E are groups of compositionally similar, stratigraphically adjacent, Urubici flows, and are a convenient means to simplify later discussions (see text). (a) ϵ_{Nd_i} vs ϵ_{Sr_i} , (b) ${}^{207}Pb/{}^{204}Pb$ vs ${}^{206}Pb/{}^{204}Pb$, (c) ${}^{208}Pb/{}^{204}Pb$ vs ²⁰⁶Pb/²⁰⁴Pb. The isotopic similarity of the unit C and D Urubici samples with the DSDP Site 525A Walvis Ridge samples (Richardson et al., 1982) is striking. NHRL, Northern Hemisphere Reference Line (Hart, 1984). No Urubici samples in this study have the very low 206 Pb/ 204 Pb values (<17.4) found by Hawkesworth et al. (1986). The cause of the discrepancy is uncertain, although the data of Hawkesworth et al. (1986) were determined in a different laboratory (Oxford). Two otherwise compositionally very similar samples from the same flow in the GB profile have very different 206Pb/204Pb (DGB-09 17.59, Table 3; GB13af 17.06, Hawkesworth et al., 1986). Post-eruption alteration can produce variable U/Pb within a single flow. However, the measured difference in ${}^{206}\text{Pb}/{}^{204}\text{Pb}$ of 0.5 would require a large difference in μ of ~25 and, although U and Pb contents were not determined for DGB-09, all the Urubici samples for which U and Pb data exist have relatively low μ $\ll 20$



Fig. 6. (a) Nineteen compositionally distinct Urubici flows are recognized within the São Joaquim sequences from small differences in MgO and TiO₂ contents and their stratigraphical position (Fig. 7). All samples from a particular flow have the same symbol. The line separates samples with ⁸⁷Sr/⁸⁶Sr_i > or < 0.7053. (b) This panel shows data for the high-SiO₂ (>54 wt %) samples (flows 12 and 19, coastal dyke ST93-5), and also the clear compositional gap between the Urubici magmas and the low-Ti/Y Gramado and Esmeralda magmas sampled from the same road profiles (Peate & Hawkesworth, 1996). The grey lines are low-pressure (1 kbar) fractionation paths modelled for two parent magmas [(i) flow 7, (ii) flow 9/10] using the MELTS program (Ghiorso & Sack, 1995) with $f(O_2)$ set at the QFM buffer.

lower units (A, B, C) and the upper units (D, E) from Pb isotope data (Fig. 5). Unit A flows (1, 4–8) have the lowest TiO₂ contents and high ⁸⁷Sr/⁸⁶Sr_i >0·7053. Unit B flows (2, 3) and unit C flows (9–11) have higher TiO₂ and low ⁸⁷Sr/⁸⁶Sr_i <0·7053, whereas the unit D flows (13–15) have very high TiO₂ (>4·0 wt %) and low ⁸⁷Sr/⁸⁶Sr_i <0·7050. Unit E comprises the two evolved flows (12 and 19) and the uppermost basaltic flows (16–18). These flows have high ⁸⁷Sr/⁸⁶Sr_i >0·7053, but are distinguished from the unit A flows by their high ²⁰⁶Pb/²⁰⁴Pb (17·9–18·3).



Fig. 7. Local correlation of Urubici flows from Fig. 6 between the São Joaquim road profiles, which highlights the local importance of preexisting topographic relief in controlling the emplacement of lava flows. $^{40}\text{Ar}^{/39}\text{Ar}$ -dated samples: No. 1, GB40c, 132·4 ± 0·7 Ma; No. 2, DUP-01, 130·2 ± 1·1 Ma; No. 3, DSM-34, 132·3 ± 0·8 Ma; No. 4, DSM-05a, 129·4 ± 1·3 Ma (No. 1: Renne *et al.*, 1992; Nos 2, 3, 4: Turner *et al.*, 1994). No. 5, DSM-08, 132 ± 5 Ma apatite fission-track age (Gallagher *et al.*, 1994).

Major element systematics and fractional crystallization models

Urubici samples plot on the l atm cotectic (ol + pl + cpx + liq) in the CIPW normative Di-Ol-Hy-Qz/Ne diagram of Thompson et al. (1983) (Peate, 1990), which indicates that they equilibrated and crystallized at low pressures in near-surface magma reservoirs. The major element characteristics of the Urubici magmas are consistent with fractionation of an olivineclinopyroxene-plagioclase assemblage (Bellieni et al., 1984). This fractionation leads to near-constant SiO_2 and Al_2O_3 , a decrease of CaO, and the development of the typical tholeiitic differentiation trend of Fe₂O₃ and TiO₂ enrichment, as MgO decreases (Cox, 1980). This is best shown by the low $^{87}\mathrm{Sr}/^{86}\mathrm{Sr}_i$ units (B, C, D). The plausibility of generating this trend by fractional crystallization can be tested with least-squares major element mixing



Fig. 8. Stratigraphical variations in Urubici flow compositions in the São Joaquim area. Flow number vs (a) MgO, (b) ⁸⁷Sr/⁸⁶Sr_i, (c) ²⁰⁶Pb/²⁰⁴Pb. Dashed line indicates approximate division into lower units (A–C) and upper units (D and E). MgO and ⁸⁷Sr/⁸⁶Sr_i data on Urubici samples from GB profile from Mantovani *et al.* (1985) are also plotted.

models. A good model fit (Peate, 1990) is obtained for the composition of the flow with the highest TiO_2 (flow 13: MgO 4·4 wt %, TiO₂ 4·2 wt %) by 15% crystallization of olivine-plagioclase-clinopyroxene in the proportions 15/47/38, starting from the most mafic flow (flow 7: MgO 5.4 wt %, TiO₂ 3.3 wt %). This calculated crystallizing assemblage is similar to the typical tholeiitic extract of Cox (1980), namely, 15/50/35. These trends can be approximately reproduced by low-pressure (1 kbar) fractionation using the MELTS program of Ghiorso & Sack (1995), taking either flow 7 or flow 9/10 as the initial composition and setting $f(O_2)$ conditions at the QFM buffer (Fig. 6). Magnetite joins the fractionating phases for samples with MgO <4.2 wt %, producing a strong decrease in Fe_2O_3 and TiO_2 (Fig. 6), and contributing to the rapid increase in SiO₂. Constant P_2O_5 contents in the andesite and dacite flows reflect crystallization of minor amounts of apatite, which can be found as small euhedral crystals in the porphyritic dacite flow 19.

The high ${}^{87}\text{Sr}/{}^{86}\text{Sr}_i$ (>0.7053) unit A flows are anomalous and define a shallow trend in Fig. 6, with only a slight increase in TiO₂ as MgO decreases. This trend, from flow 7 (MgO 5.4 wt %, TiO₂ 3.3 wt %) to flow 4 (MgO 4.0 wt %, TiO₂ 3.6 wt %) can be modelled by 15% crystallization of an olivine–plagioclase–clinopyroxene–magnetite assemblage in the proportions 10/ 33/53/5. This extract involves an unrealistically high proportion of the Ti-rich phases augite and magnetite, which are required so that the bulk distribution coefficient for Ti is close to one. Successive flows within unit A do not define a simple temporal trend of decreasing MgO that might indicate progressive fractionation. Instead, it is more likely that the shallow trend represents mixing between relatively evolved and relatively primitive liquids. Within unit A, the flows with lower MgO tend to have higher SiO₂ and ⁸⁷Sr/⁸⁶Sr_i, which suggests that the more evolved end-member suffered more crustal interaction.

Crustal assimilation

Crustal assimilation should have a limited effect on Urubici magmas because of their high incompatible trace element contents. Crustal rocks generally have high δ^{18} O relative to the mantle average value of +5.5%, and so oxygen isotope data offer one means by which to assess the extent, if any, of crustal input to these mantle-derived magmas. The sparse δ^{18} O data obtained on whole-rock Urubici samples are primarily controlled by water–rock interactions (Iacumin *et al.*, 1991). Whole-rock samples of the equivalent Khumib magma type in the Etendeka have high δ^{18} O (+8 to +11‰), whereas clinopyroxene mineral separates give lower δ^{18} O (+5.9 to +6.0‰) only slightly elevated above the values expected for clinopyroxenes in equilibrium with uncontaminated mantle-derived basalts (Harris *et al.*, 1989).

Correlations exist between certain incompatible trace element ratios and radiogenic isotope ratios that suggest that some crustal assimilation did take place, although to a much lesser extent than in the low-Ti/Y Gramado magma type (Peate & Hawkesworth, 1996). There is a positive correlation between Th/Ta and ⁸⁷Sr/⁸⁶Sr_i (Fig. 9a) and a negative correlation between La/Ta and ϵ_{Ndi} (Fig. 9b), implying addition of material with high 87 Sr/ 86 Sr_i, low ε_{Nd} and a significant negative Ta–Nb anomaly, all features typical of continental crust. The strong crustal influence seen in the local low-Ti/Y magmas (Gramado basalts; Peate & Hawkesworth, 1996: Palmas rhyolites; Garland et al., 1995) means that mixing of these magmas with an Urubici magma could also mimic many of the compositional effects of crustal assimilation. Th/ Ta and ⁸⁷Sr/⁸⁶Sr_i correlate with increasing SiO₂ going from the basalts to the andesite and dacite flows, but data for the evolved flows do not lie on a mixing line between data fields for the basalts and the low-Ti/Y magmas. This is apparent from Fig. 9c, where the andesite and dacite data plot at higher Zr contents than values for the Urubici basalts and the low-Ti/Y Gramado and Palmas magmas. Instead, these rare, high-SiO₂ (>54 wt %) flows probably result from a localized style of contamination where the amount of assimilation is linked thermally to the extent of crystallization [e.g. AFC (assimilation-fractional crystallization) model of DePaolo (1981)]. The basaltic flow 18, found near the top of the SM profile, is the only Urubici flow that has some compositional features that might indicate that it is a hybrid flow. Samples of this flow plot beneath the main trend on Fig. 9c and have the lowest $(Tb/Yb)_N \sim 2.1$ of the Urubici basalts [where subscript N denotes chondritenormalized using the values of Sun & McDonough (1989)]. It is possible to explain these features, as well as the Sr-Nd-Pb isotope composition, by mixing ~30% of an average low-Ti/Y Gramado magma with a unit D Urubici magma. It is notable though, that the three samples of this flow (DSM-05b, -07, -09), taken over a lateral distance of 1 km, indicate a rather homogeneous composition.

The Urubici basalts show poor correlations between isotope composition and indices of fractionation such as MgO, except for the unit A flows as mentioned in the previous section. On a broad scale, there is a general decrease in 87 Sr/ 86 Sr_i with increasing stratigraphic height for the basalts (Fig. 8). This suggests that the extent of contamination was decreasing with time, perhaps as a result of the more readily mobilized crustal material having been scavenged from the conduit walls by the earlier flows.

There is a general positive correlation between ${}^{87}\text{Sr}/{}^{86}\text{Sr}_i$ and ${}^{206}\text{Pb}/{}^{204}\text{Pb}$ for the Urubici data (Fig. 10). In detail, the lower units (A–C) and the upper units (D and E) form two parallel trends, and each is interpreted as reflecting assimilation of crustal material with high ${}^{87}\text{Sr}/{}^{86}\text{Sr}$ and ${}^{206}\text{Pb}/{}^{204}\text{Pb}$ but starting from two compositionally distinct end-members represented by units C and D, respectively. Two trends are also apparent from the Pb–Pb isotope diagrams (Fig. 5), where the lower and upper units form two en-echelon trends, with increased contamination leading to higher $\Delta 7/4$. The lower and



Fig. 9. Diagrams illustrating the effects of crustal assimilation and mixing on some Urubici flows. (a) ${}^{87}\text{Sr}/{}^{86}\text{Sr}_i$ vs Th/Ta, (b) ϵ_{Nd_i} vs La/Ta. Data from Mantovani *et al.* (1985) are not included in (a) or (b) because the Ta data are ~0.5 ppm higher than the data of this study. (c) Zr vs MgO. The high-SiO₂ Urubici flows are not a mixture of an Urubici basalt with a low-Ti/Y magma because they have higher Zr contents than the Urubici basalts and the low-Ti/Y basalts or rhyolites (Garland *et al.*, 1995; Peate & Hawkesworth, 1996). Basalt flow 18 has low Zr and lies below the main trend for the Urubici basalts. This, together with its REE and isotopic composition, can be explained if it is a hybrid magma between a unit D Urubici magma and an average Gramado magma.



Fig. 10. ²⁰⁶Pb/²⁰⁴Pb vs ⁸⁷Sr/⁸⁶Sr_i diagram for the Urubici basalts, the Tristan plume, Walvis Ridge and South Atlantic MORB. For the Urubici basalts, the lower units (A + B + C) and the upper units (D + E) form two parallel arrays. These are interpreted as reflecting assimilation of high-²⁰⁶Pb/²⁰⁴Pb, high-⁵⁷Sr/⁶⁶Sr_i crustal material, but in each case starting from compositionally distinct end-members represented by units C and D, respectively. Tristan plume and Walvis Ridge data from Richardson *et al.* (1982), le Roex (1985), Weaver *et al.* (1987), le Roex *et al.* (1990), Cliff *et al.* (1991). South Atlantic MORB data from Hanan *et al.* (1986), Castillo & Batiza (1989), and Fontignie & Schilling (1996). Symbols as in Fig. 5.

upper units also form parallel arrays in Fig. 9a. Hawkesworth *et al.* (1986) noted that the Urubici samples define an apparent isochron on a $^{207}\text{Pb}/^{204}\text{Pb}$ vs $^{206}\text{Pb}/^{204}\text{Pb}$ diagram with an age of 1.8 ± 0.4 Ga that is similar to a $^{207}\text{Pb}-^{206}\text{Pb}$ age of 2.2 ± 0.2 Ga for the local Precambrian basement rocks. However, any age significance for the source of the Urubici suite as a whole is probably doubtful, given the effects of crustal contamination and mixing shown in Fig. 7. Units C and D have Nd model ages for separation from 'depleted' MORB-source mantle of 1.2-1.3 Ga.

MANTLE ORIGINS OF THE URUBICI MAGMA TYPE

The aims of this section are twofold: (1) to ascertain which compositional features of the Urubici magmas can be attributed to mantle source characteristics; (2) to discuss the composition of melts derived from the Tristan plume and other magmatic samples of the regional upper mantle (lithosphere and asthenosphere), to assess potential sources for the Urubici magmas.

Isotope and trace element characteristics of uncontaminated Urubici magmas

The Urubici magmas have experienced extensive shallow-level processing. They have evolved compositions

(mg-number <52, Ni <100 ppm) that are far removed from any plausible primary mantle magma {melt in equilibrium with mantle peridotite (mg-number 88-89) has mg-number 68-75 [mg-number is atomic ratio $100 \times \text{Mg/(Mg + Fe}^{2+})$ assuming $\text{Fe}^{3+}/\text{Fe}_{\text{total}} = 0.15$]; Frey et al., 1978}. To see through the effects of shallowlevel crystallization, more reliance is placed on incompatible trace element ratios and isotope compositions. Units C and D are believed to have suffered minimal crustal interaction, by virtue of their low ⁸⁷Sr/⁸⁶Sr_i and Th/Ta, and from their positions anchoring the low ²⁰⁶Pb/ ²⁰⁴Pb ends of the trends in Figs 5 and 10. Thus, these two units provide the best clues to the nature of the mantle source of the Urubici magma type. The unit B flows are similar in many respects to the unit C flows, but only one sample has been analysed for Sr-Nd-Pb isotopes and by neutron activation analysis and it had higher Th/Ta than the unit C flows.

The trace element characteristics of units C and D are shown in a primitive-mantle-normalized diagram (Fig. 11). Two important features are (1) the depletion in Nb relative to La and K [Nb/La ~0.65 vs 1.0 in normal (N-)MORB], which is not a characteristic of typical oceanic basalts, and (2) the marked depletion of heavy relative to middle rare earth elements (REE) [(Tb/ $Yb_{N} \sim 2.5$], which is indicative of the involvement of residual garnet at some stage during the evolution of the magma or its source: such high (Tb/Yb)_N values are rare in tholeiitic flood basalts. Data for units C and D lie at the low $^{87}\mathrm{Sr}/^{86}\mathrm{Sr}_i,$ high ϵ_{Nd_i} end of the Urubici array in Fig. 5a. In detail, unit D flows have lower ⁸⁷Sr/⁸⁶Sr_i and higher ϵ_{Nd_i} and ${}^{206}Pb/{}^{204}Pb$ (${}^{87}Sr/{}^{86}Sr_i$ 0.7048–0.7050, ϵ_{Nd_i} -1.9 to -2.5, $^{206}Pb/^{204}Pb \sim 17.75$) than unit C flows ($^{87}Sr/^{86}Sr_i$ 0.7050–0.7053, $\epsilon_{\rm Nd_i}$ –2.5 to –3.5, $^{206}Pb/^{204}Pb$ \sim 17.50). Minor trace element differences also exist between these two units (e.g. unit D, Tb/Yb_N ~ 2.55 ; unit C, Tb/Yb_N ~ 2.40 implying a temporal change in the composition of the magma being input to the shallowlevel chambers.

Composition of the Tristan mantle plume

The age progressive nature of magmatism along the Walvis Ridge and Rio Grande Rise confirms their interpretation as the fossil magmatic trace of the mantle thermal anomaly, presumably a plume, now located beneath Tristan da Cunha (Fig. 1: O'Connor & Duncan, 1990). The recent (<10 Ma) composition of the plume can be inferred from the magmatism of the Tristan da Cunha island group (Tristan, Inaccessible, Nightingale) and of Gough island, which lies ~500 km to the southeast. Published data (le Roex, 1985; Weaver *et al.*, 1987; le Roex *et al.*, 1990; Cliff *et al.*, 1991) indicate that these islands share many distinctive compositional features that



Fig. 11. Primitive-mantle-normalized trace element diagram [normalizing values from Sun & McDonough (1989)]. The patterns for the Urubici basalts and the Walvis Ridge DSDP Site 525A basalts share many similarities, in particular the low Nb/La, which is in marked contrast to the patterns for recent magmas associated with the Tristan plume (Tristan da Cunha, Gough, Inaccessible). Data sources as for Fig. 10, plus Humphris & Thompson (1983) and Sun & McDonough (1989).

distinguish them from most other oceanic islands; e.g. high Ba/La, Ba/Nb and Th/Ta, low ε_{Nd_i} (-4 to +1) and Dupal-type Pb isotope signatures. In detail, there are inter-island differences (Gough and Nightingale have higher 207 Pb/ 204 Pb for a given 206 Pb/ 204 Pb value than Tristan and Inaccessible) and some intra-island variations (Inaccessible shows a larger range in ε_{Nd_i} than the other islands) that demonstrate some source heterogeneity (Fig. 12).

Dredging and DSDP drilling have recovered samples from the Walvis Ridge and Rio Grande Rise, but comprehensive isotope and trace element data sets are sparse. Nevertheless, most of the compositional variations can be explained by a mixing model involving a 'depleted' MORB-like mantle source and an 'enriched' mantle source that resembles that of Tristan (Humphris & Thompson, 1983). The DSDP Site 525A basalts (~70 Ma) are an exception and their isotopic values plot away from any possible mixing curves. They have low ε_{Nd} (~ -3) , low ²⁰⁶Pb/²⁰⁴Pb (~17.5), and high Ba/Nb (~17) (Fig. 12: Richardson et al., 1982; Humphris & Thompson, 1983). They also have lower Nb/La than primitive mantle, in contrast to other Walvis Ridge samples and recent Tristan plume magmatism (Figs 11 and 13: Milner & le Roex, 1996). Samples similar to the DSDP Site 525A basalts have been dredged from the eastern part of the Walvis Ridge and drilled at DSDP Site 516F on the Rio Grande Rise (Humphris & Thompson, 1983; Hart, 1984).



Fig. 12. Radiogenic isotope composition of Cretaceous continental mafic potassic magmas and carbonatites (Toyoda *et al.*, 1994; Gibson *et al.*, 1995*a*, 1995*b*; Huang *et al.*, 1995; Carlson *et al.*, 1996), interpreted as lithospheric mantle melts, and the Tristan plume (Tristan da Cunha, Inaccessible, Gough, Walvis Ridge DSDP Sites 527 and 528: data sources as for Fig. 10), relative to the Urubici unit C and D lavas. (a) $\epsilon_{\rm Nd_i}$ vs $\epsilon_{\rm Sr_i}$, (b) $^{207}\text{Pb}/^{204}\text{Pb}$ vs $^{206}\text{Pb}/^{204}\text{Pb}$ vs $^{206}\text{Pb}/$



Fig. 13. La/Nb vs ε_{Nd_i} diagram. This illustrates the elevated La/Nb of the Urubici lavas and the Walvis Ridge DSDP Site 525A basalts relative to Tristan plume lavas and other sites on the Walvis Ridge. Furthermore, the Urubici basalts have higher La/Nb than the high-Ti mafic potassic magmas. Data sources as for Fig. 12. OPM is 'oceanic-plume magma' from Gibson *et al.* (1995*a*).

Cretaceous continental alkalic magmatism

The periphery of the Paraná-Etendeka lava field experienced two episodes of compositionally diverse alkalic magmatism during the Cretaceous (124-147 Ma and 55-90 Ma: Ulbrich & Gomes, 1981; Gibson et al., 1995a; Milner et al., 1995b). Most of this Cretaceous alkalic magmatism is compositionally distinct from recent Tristan plume magmatism, but two exceptions have been described from near the continental margin. The coastparallel São Sebastião basanitic dykes on the Brazilian coast between São Paulo and Rio de Janeiro (81 Ma: Regelous, 1993; SJII dykes of Hawkesworth et al., 1992) have Tristan-like incompatible trace element patterns and Sr-Nd-Pb isotope compositions. The presence of the São Sebastião dykes, together with the bulk of the Walvis Ridge data, is strong evidence that the composition of at least one type of mantle in the Tristan plume has not changed significantly over at least the last 90 my. Milner & le Roex (1996) considered the alkalic magmas from Okenyenya, Namibia (Fig. 1: 124-130 Ma) to be compositionally similar to modern Tristan plume magmas, but these magmas have significantly lower ²⁰⁸Pb/²⁰⁴Pb than the main Tristan plume field (Fig. 12).

Several of the Early Cretaceous carbonatite complexes [Jacupiranga, ~250 km north of São Joaquim (Fig. 2), and Anitapolis, 90 km east of São Joaquim (Fig. 4)] have low ²⁰⁶Pb/²⁰⁴Pb compositions similar to the Urubici and Walvis Ridge DSDP Site 525A basalts. The distinctive isotope characteristics of Jacupiranga (Dupal Pb isotopes and Sr and Nd isotope compositions close to estimated Bulk Earth values: ⁸⁷Sr/⁸⁶Sr_i 0.7047–0.7057, ε_{Ndi} + 1.6 to -1.2, ²⁰⁶Pb/²⁰⁴Pb 17:0–17: Toyoda *et al.*, 1994; Huang *et al.*, 1995) are inferred to be a mantle source feature, given the high trace element contents of the samples and the mantle-like δ^{18} O values. Carbonatitic dykes at Anitapolis have slightly higher δ^{18} O, ⁸⁷Sr/⁸⁶Sr_i and ²⁰⁶Pb/²⁰⁴Pb_i values than Jacupiranga, indicating a limited interaction with upper-crustal material (Toyoda *et al.*, 1994).

Gibson et al. (1995a) recognized that the Cretaceous mafic potassic magmatism can be divided into two compositionally and geographically distinct groups: (1) a high-Ti group (TiO₂ ~4.6 wt %) with high ε_{Ndi} (~ -5) and low La/Nb (~1·1) and 87 Sr/ 86 Sr_i (~0·7050), found along the northeast margin of the Paraná lava field, associated with the Proterozoic mobile belts, and (2) a low-Ti group (TiO₂ ~1.8 wt %) with low $\epsilon_{\!Nd_i}\,(\sim$ –13) and high La/Nb (~2.0) and ⁸⁷Sr/⁸⁶Sr_i (~0.7068), found adjacent to the central Paraná lavas along the coast in the east and in Paraguay to the west, associated with cratonic regions. Mafic potassic magmas are widely accepted as representing small degree melts of the lithospheric mantle. Gibson et al. (1995a) concluded that the spatial distribution of high-Ti and low-Ti mafic potassic rocks was analogous to the high-Ti-low-Ti provinciality seen in the Paraná basalts and reflected a major compositional change in the continental lithospheric mantle, as previously suggested on the basis of the basalt data by Erlank et al. (1988) and Hawkesworth et al. (1988). However, this boundary might not be the same as that inferred from the basalts. The low-Ti mafic potassic rocks are spatially associated with tholeiitic dykes that are inferred to be the feeder systems to the volumetrically dominant northern high-Ti/Y flood basalts (Fig. 2), and yet the strike of the tholeiitic dykes is roughly parallel to the inferred lithospheric mantle boundary of Gibson et al. (1995a).

Carlson et al. (1996) showed that the high-Timafic potassic magmas in the Alto Paranaíba region (Fig. 1) generally have $\log^{206} Pb/^{204} Pb_{85Ma}(17.2-18.1)$ and elevated $208 Pb/^{204} Pb_{85Ma}$ (37.8-38.6) (Fig. 12). Some of the lamproitic and kimberlitic samples have extremely low ${}^{187}\text{Os}/{}^{188}\text{Os}_i$ (0.112–0.121), lower than oceanic basalts and abyssal peridotites, and such characteristics are apparently unique to peridotitic lithospheric mantle sources that experienced melt-related Re depletion during the Archaean or Proterozoic (Carlson et al., 1996). Kamafugitic samples have high ¹⁸⁷Os/¹⁸⁸Os_i (0.27-0.31), but the high Os contents rule out the possibility of contamination with old, high-Re/Os crustal material, and instead Carlson et al. (1996) interpreted them as being derived from a pyroxenitic or eclogitic source within the lithospheric mantle. Thus, Carlson et al. (1996) argued that the Sr-Nd-Pb-Os isotopic characteristics of the Alto Paranaíba samples were inherited from diverse sources within the lithospheric mantle and were not compatible with a mantle source that had been modified by melts from the Tristan plume.

A common mantle source for high-Ti/Y and low-Ti/Y magma types?

To understand the mantle origins of the high-Ti/Y Urubici magma type, we should also consider its relationship to the contemporaneous low-Ti/Y magmatism. These two types cannot be related by crustal-level processes such as crystallization or assimilation. Fodor (1987) and Arndt et al. (1993) considered that coexisting highand low-Ti/Y basalts in the Paraná and in other flood basalt provinces could be linked simply by different degrees of melting of a common mantle source, the higher incompatible trace element contents of the high-Ti/Y magmas reflecting lower degrees of melting than for the low-Ti/Y magmas. Fodor (1987) related the melting differences to the proximity of the underlying Tristan plume, with low-Ti/Y basalts over the hot plume axis in the south and high-Ti/Y basalts on the cooler plume periphery to the north. Following McKenzie & Bickle (1988), Arndt et al. (1993) considered lithospheric thickness to be the principal factor in determining the degree and depth of melting, by controlling the extent to which the underlying asthenospheric mantle can decompress and melt; it should be noted that melting of the lithospheric mantle is not the main issue here. Beneath thick lithosphere, the degree of melting will be limited and will occur at high pressure in the presence of residual garnet, thus leading to incompatible-element-enriched basalts with high Ti/Y and Tb/Yb. If similar source material ascends to shallower depths beneath a lithosphere perhaps thinned by extension, the degree of melting will be higher, producing melts with lower concentrations of incompatible elements and without a residual garnet signature (i.e. low Ti/Y and Tb/Yb). The difference in Tb/Yb between the low-Ti/Y and high-Ti/Y magmas is illustrated in Fig. 14b. In some flood basalt provinces, there is a temporal switch from high-Ti/Y and -Tb/Yb lavas to low-Ti/Y and -Tb/Yb lavas (e.g. Ellam, 1992; Fram & Lesher, 1993; White & McKenzie, 1995), consistent with progressive lithospheric thinning during rifting. In the Paraná-Etendeka province, it is important to note that the high-Ti/Y Urubici and low-Ti/Y Gramado magma types are contemporaneous, which would require significant lateral variations in lithospheric thickness along the developing South Atlantic rift in such a model, and yet both types appear to have been erupted from coast-parallel dykes. Any isotopic differences between the high-Ti/Y and low-Ti/Y types, or relative to typical asthenospheric melts, were simply attributed to crustal contamination by Fodor (1987) and Arndt et al. (1993). It remains difficult, though, to reconcile the isotopic variations of the Paraná lavas in terms of melts from a single mantle source without appealing to models that require melts derived from contrasting pressure and/or temperature conditions to also assimilate compositionally different crustal components.



Fig. 14. (a) $(Tb/Yb)_N$ vs $(La/Yb)_N$ diagram illustrating the REE characteristics of the high-Ti and low-Ti mafic potassic magmas (Gibson et al., 1995a, 1995b; Carlson et al., 1996) relative to the high-Ti/Y and low-Ti/Y flood basalts (Table 2; Peate & Hawkesworth, 1996). The large symbols represent average compositions for the high-Ti (filled) and low-Ti (open) mafic potassic magmas. (b) An enlargement of bottom left-hand corner of (a). If the compositions of the high-Ti/Y and low-Ti/Y Paraná basalts represent mixtures of an asthenospheric melt with high-Ti and low-Ti mafic potassic magmas, respectively (Gibson et al., 1995a), then the composition of the asthenospheric melt must lie within the shaded regions. It is clear that, for this model, the asthenosphere-derived end-member must be different for the high-Ti/ Y and low-Ti/Y basalts. The continuous lines show the low (La/Yb)_N portions of mixing trends to the average high-Ti and low-Ti mafic potassic magmas, and the dashed lines represent mixing with the extremes of observed mafic potassic magma compositions.

In certain flood basalt provinces, close compositional similarities exist between some of the flood basalt flows and the recent oceanic magmatism of the associated mantle plume (e.g. Deccan flood basalts and Réunion plume, Mahoney, 1988: Madagascar flood basalts and Marion plume, Mahoney *et al.*, 1991; Storey *et al.*, 1997: Greenland flood basalts and Iceland plume, Thirlwall et al., 1994). However, it has proven more difficult to demonstrate a similar compositional link between the Paraná-Etendeka flood basalts and the youngest magmatic products of the Tristan mantle plume (Hawkesworth et al., 1992; Peate & Hawkesworth, 1996). The high-Ti/Y Urubici lavas have Nb troughs on primitivemantle-normalized trace element plots, which are not a feature of the modern Tristan plume (Fig. 11), and the isotopic differences between these flood basalts and both recent plume-derived magmas and local MORB-source asthenosphere are apparent from comparing the data in Figs 5 and 12. Instead, the Urubici basalts have more compositional features in common with those Cretaceous alkalic magmas that are inferred to be small-degree melts of the continental lithospheric mantle.

A ROLE FOR LITHOSPHERIC MANTLE?

It has been proposed that the lithospheric mantle plays an important role in magma genesis in some, but not all, flood basalt provinces (e.g. Allègre et al., 1982; Hawkesworth et al., 1984, 1988; Ellam & Cox, 1991; Hergt et al., 1991; Ellam et al., 1992; Sweeney et al., 1994; Gibson et al., 1995a; Molzahn et al., 1996; Peate & Hawkesworth, 1996; Turner et al., 1996; Lassiter & DePaolo, 1997). These ideas were influenced primarily by studies of the Mesozoic Gondwanan flood basalt provinces (Paraná-Etendeka, Karoo, Ferrar) and emphasized two important features of these basalts: (1) samples considered to be uncontaminated by crust had trace element and isotope characteristics distinct from typical oceanic basalts, namely, low Nb/La and ε_{Nd_i} and high ⁸⁷Sr/⁸⁶Sr, with a virtual absence (<5%) of any samples with ε_{Nd_i} values higher than Bulk Earth (e.g. Hawkesworth et al., 1984, 1988; Hergt et al., 1991; Peate et al., 1992; Sweeney et al., 1994); (2) high-Ti/Y and low-Ti/Y flood basalts form spatially distinct geochemical provinces that can be traced throughout the former Gondwana supercontinent (Bellieni et al., 1984; Erlank et al., 1988; Hergt et al., 1991). $^{\rm 40}{\rm Ar}{\rm -}^{\rm 39}{\rm Ar}$ ages suggest that the stratigraphic units defined by the Paraná basalt magma types are not chronostratigraphic (Turner *et al.*, 1994), and Peate et al. (1990) showed that they were erupted from different centres. Thus the spectrum of magma compositions cannot represent the temporal evolution of a single mantle source region. The surface distribution of magma types appears to reflect the subcrustal distribution of distinct lithospheric source regions (Hawkesworth et al., 1988), similar to the lithospheric mantle provinciality inferred from the Cretaceous alkalic magmatism (Gibson et al., 1995a).

Assimilation of small-degree lithospheric melts by plume-derived melts?

One mechanism by which the lithospheric mantle could contribute to flood basalt magmatism is by asthenospheric melts assimilating a lithospheric mantle component, either as they migrate up through the lithosphere or via thermal erosion of the base of the lithosphere by an incubating plume head (e.g. McKenzie, 1989; Ellam & Cox, 1991; Saunders et al., 1992). This component is inferred to be in the form of low-melting-point hydrous and carbonated material similar in composition to mafic potassic magmas. Ellam et al. (1992) presented good evidence for such a model to explain the positive correlations between Nd, Pb, and Os isotope compositions in the high-Ti Nuanetsi picrites (Karoo flood basalt province, southern Africa). The low $\epsilon_{\rm Nd_i}$ (~ -10), low $^{206} Pb/^{204} Pb_i$ (~17.4) end-member has low $^{187} Os/^{188} Os_i$ (<0.123), indicative of an incompatible-trace-elementenriched lithospheric mantle component rather than crust.

Gibson et al. (1995a) tried to model the composition of the Paraná high-Ti and low-Ti basalts by mixing between an asthenospheric melt and the Cretaceous high-Ti and low-Ti mafic potassic lithospheric melts, respectively, followed in both cases by extensive crystallization and crustal assimilation. The asthenospheric end-member was assumed to be similar to the incompatible-trace-element-depleted tholeiitic and picritic lavas found on Iceland and on the Caribbean oceanic plateau, which are interpreted as large degree melts, formed by shallow decompression and/or elevated mantle temperatures, of a depleted mantle component within the plume head, referred to by Gibson et al. (1995a) as 'oceanic plume magmas'. It is not clear, though, that lithospheric extension was sufficient to allow decompression to such shallow depths during this stage of the Paraná-Etendeka magmatism (e.g. Turner et al., 1996). Gibson et al. (1995a) suggested that the Sr and Nd isotopic composition of the high-Ti/Y flood basalts was consistent with the addition of ${\sim}50\%$ high-Ti mafic potassic lithospheric-derived melt to an 'oceanic plume magma'.

One problem with this model is highlighted by the REE data (Fig. 14). Both of the axes in this plot $[(Tb/Yb)_N vs (La/Yb)_N]$ have the same denominator and so binary mixing is linear. It is possible to extrapolate mixing lines from the high-Ti mafic potassic magmas through the field of the Urubici magmas to define the region where an asthenospheric end-member must lie in the model of Gibson *et al.* (1995*a*). A similar procedure can be done for the low-Ti mafic potassic magmas and the low-Ti/Y flood basalts. It is clear that the difference in (Tb/Yb)_N between the contemporaneous high-Ti/Y and low-Ti/Y basalts cannot be explained by addition of

mafic potassic melts of different REE composition to a common asthenospheric melt end-member: this problem is not mitigated by the effects of crustal contamination on the Urubici magmas because the contaminated units (A, B, E) have lower (Tb/Yb)_N. The model of Gibson *et al.* (1995*a*) thus requires significant contemporaneous differences in the asthenospheric melting conditions to produce compositionally different parental melts to the Urubici magmas [asthenospheric melt with (Tb/Yb)_N ~2·0] and the Gramado magmas [asthenospheric melt with (Tb/Yb)_N ~1·3].

The least contaminated Urubici units C and D flows have a broadly similar Nd isotope composition ($\varepsilon_{Nd_i} - 2.7 \pm 0.6$) to the high-Ti mafic potassic melts ($\varepsilon_{Nd_i} - 4.8 \pm 1.6$: Gibson *et al.*, 1995*a*, 1995*b*; Carlson *et al.*, 1996), but they have much higher La/Nb (1.6 vs 1.1), a difference attributed by Gibson *et al.* (1995*a*) to crustal assimilation. Given that the local Brazilian crust has $\varepsilon_{Nd_{130Ma}} < -10$, the significant crustal addition that would be required to explain the difference in La/Nb would mean that the uncontaminated Urubici magmas would have $\varepsilon_{Nd_i} > 0$, which then limits the amount of lamproitic addition possible. The available δ^{18} O data argue against significant crustal contamination, and the Pb isotope data for the Urubici magmas are similar to values for the Cretaceous mafic potassic lithospheric melts.

The marked compositional similarities, both in terms of trace elements and radiogenic isotopes, between the Walvis Ridge DSDP Site 525A basalts and the Urubici magma type of the Paraná flood basalts mean that it is difficult to escape the conclusion that ultimately both are derived from a common mantle source. Whichever petrogenetic model is chosen for the origins of the compositional features of the Urubici magma type has important implications for the nature and original location of the enriched mantle component (EM1) seen in the Walvis Ridge DSDP Site 525A basalts, and vice versa, a point also made by Hawkesworth et al. (1986), Toyoda et al. (1994) and Milner & le Roex (1996). Any model that invokes crustal contamination to account for the compositional features of the Urubici magmas must then address the compositional coincidence between the Urubici samples and the DSDP Site 525A Walvis Ridge samples. Another possibility is that the Urubici basalts represent wholesale melting of hydrated Brazilian lithospheric mantle, and the DSDP Site 525A Walvis Ridge lavas sample similar material that was delaminated during break-up and entrained in the asthenosphere.

Melting of hydrated lithospheric mantle peridotite

If melting occurs at the anhydrous peridotite solidus during continental rifting above a mantle plume then >95% of melt produced will be of asthenospheric origin, as proposed by, for example, McKenzie & Bickle (1988) and Arndt & Christensen (1992). However, those workers did not consider the implications of a plume incubating beneath thick lithosphere for several million years or the presence of small amounts of water within the lithospheric mantle. Gallagher & Hawkesworth (1992) and Turner et al. (1996) demonstrated that melting will then preferentially occur at the volatile-enriched solidus within the lithospheric mantle unless there is significant extension to facilitate adiabatic decompression melting of the underlying asthenosphere. A key issue is the thickness of the lithosphere at the time of magmatism. If the lithospheric thickness is >130 km and extension remains low, as appears to be the case in the Paraná (Turner et al., 1996), then melting will be confined to the lithospheric mantle [see also Storey et al. (1997)].

IMPLICATIONS FOR THE LOCATION OF 'DUPAL-ANOMALY' MANTLE AND 'EM1-TYPE' MANTLE IN THE SOUTH ATLANTIC REGION

Hart (1984) defined the Dupal mantle anomaly in the South Atlantic region on the basis of the distinctive isotopic compositions of lavas from some geographically adjacent oceanic islands, aseismic ridges, and seamount chains, which include Tristan, Gough, Walvis Ridge and Rio Grande Rise. These OIB also share several trace element features, such as high Ba/Nb and Ba/La, that are distinct from most other OIB and MORB (e.g. Weaver et al., 1987). The implication is that these OIB all sample a similar, isotopically anomalous, mantle source. This is surprising, given the wide range in, for example, ${}^{206}Pb/{}^{204}Pb$ (17.5–18.8), which cannot easily be explained by variable entrainment of MORB-source asthenospheric material in any plume upwelling (e.g. Fig. 12c). This point is best illustrated by the isotopic data from the Walvis Ridge (Richardson et al., 1982), which, in many publications, are portrayed as a linear array. In fact, as stressed by Milner & le Roex (1996), the Walvis Ridge data form two distinct groups, one broadly similar to data for recent Tristan plume lavas and the other with low 206Pb/204Pb, EM1-type features, but with both groups showing Dupal isotope characteristics. Low ²⁰⁶Pb/²⁰⁴Pb material apparently has a rather restricted distribution offshore, being found only at DSDP Site 525A on the Walvis Ridge (Richardson et al., 1982) and Site 519 on the Rio Grande Rise (Hart, 1984), although more detailed isotopic coverage of the Walvis Ridge and the Rio Grande Rise is clearly required. This material is not involved in the mixing arrays defined by most other Walvis Ridge samples, which reflect interaction between Tristan-like plume material and MORB source asthenosphere (Richardson et al., 1982; Humphris & Thompson, 1983). Similarly, MORB from the nearby Mid-Atlantic Ridge (22°-31°S) do not contain any trace of this low ²⁰⁶Pb/ ²⁰⁴Pb material (e.g. Figs 9 and 12c), although their compositions are consistent with the plume-ridge interaction model outlined by Humphris et al. (1985), Hanan et al. (1986) and Fontignie & Schilling (1996), using data from Tristan and Gough to represent the plume end-member. Where low ²⁰⁶Pb/²⁰⁴Pb material is found in South Atlantic MORB, further south at $48 \cdot 5^{\circ} - 49 \cdot 0^{\circ}$ S and $52 \cdot 3^{\circ} - 52 \cdot 7^{\circ}$ S, it appears to be isolated and not mixed with plume or asthenospheric material (Douglass et al., 1996). An analogous situation is found along the Southwest Indian Ridge at 39°–41°S, where low ²⁰⁶Pb/²⁰⁴Pb compositions are also difficult to ascribe to plume-ridge interactions, in this case involving the Marion plume (Mahoney et al., 1992). Thus, in the South Atlantic region, it is not obvious that the low ²⁰⁶Pb/²⁰⁴Pb, EM1-type mantle is part of the Tristan plume upwelling, unless one resorts to the rather ad hoc explanation that it was present early in the history of the Tristan plume but is now exhausted or just a very minor component.

Since the paper of Hawkesworth et al. (1986) that discussed the possible link between the Paraná high-Ti/ Y flood basalts and the Dupal oceanic magmatism in terms of delaminated lithospheric mantle, data have been published from unambiguous samples of the Brazilian continental mantle lithosphere in the form of Cretaceous small-degree mafic potassic melts (Gibson et al., 1995a; Carlson et al., 1996). From these, it is clear that the distinctive compositional features of the uncontaminated Urubici magmas are also seen in many samples derived from the local Brazilian mantle lithosphere, and this strengthens the case for a lithospheric mantle origin both for these high-Ti/Y flood basalts and, indirectly, the DSDP Site 525A Walvis Ridge samples. The data for the Cretaceous alkalic rocks show that the regional continental lithospheric mantle is isotopically fairly heterogeneous (Fig. 12), and thus is not itself equivalent to the EM1 component as originally defined by Zindler & Hart (1986): its range of 206 Pb/ 204 Pb goes well below the values seen in the Site 525A lavas. In detail, more than one low ²⁰⁶Pb/²⁰⁴Pb 'component' is required, given the significant range in ²⁰⁷Pb/²⁰⁴Pb even at the lowest ²⁰⁶Pb/²⁰⁴Pb, and Mahoney et al. (1996) reached a similar conclusion for Indian Ocean MORB and OIB.

It is likely that material from the base of the continental mantle lithosphere is detached during continental breakup, perhaps enhanced by plume-driven thermal erosion, and is subsequently dispersed within the shallow asthenosphere by secondary convection processes during seafloor spreading (e.g. Hawkesworth *et al.*, 1986; Douglass *et al.*, 1996; Milner & le Roex, 1996). The South Atlantic region is not an isolated example, and a similar model has been proposed to account for the compositional and spatial variations documented within the Madagascar flood basalts–Marion plume–Southwest Ridge Indian MORB system (Mahoney *et al.*, 1991, 1992; Storey *et al.*, 1997). The similarities between this system and the South Atlantic region have been discussed in detail by Milner & le Roex (1996).

The Tristan plume is probably a relatively deep mantle upwelling, and Castillo (1988) linked it to a low seismic velocity anomaly in the lower mantle. The 'enriched' incompatible trace element and isotopic characteristics of the Tristan plume might ultimately also be derived from continental lithospheric mantle material (e.g. le Roex, 1986; Hofmann, 1997), which, in contrast to the DSDP Site 525A source, has been cycled relatively deep in the mantle convection system and brought up as an intrinsic part of the plume. In conclusion, we prefer a model in which the Dupal anomaly in the South Atlantic region, in fact, is two distinct entities [see also Milner & le Roex (1996)]; one shallow in origin (low ²⁰⁶Pb/²⁰⁴Pb, high Ba/Nb delaminated rafts of local lithospheric mantle as sampled at DSDP Site 525A) and the other deep in origin (Tristan plume).

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