

## **Quantum Information Cannot Be Completely Hidden in Correlations: Implications for the Black-Hole Information Paradox**

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Can quantum-information theory shed light on black-hole evaporation? By entangling the in-fallen matter with an external system we show that the black-hole information paradox becomes more severe, even for cosmologically sized black holes. We rule out the possibility that the information about the infallen matter might hide in correlations between the Hawking radiation and the internal states of the black hole. As a consequence, either unitarity or Hawking's semiclassical predictions must break down. Any resolution of the black-hole information crisis must elucidate one of these possibilities.

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In 1917 Vernam invented his one-time pad cipher. In its simplest form, the cipher encodes a message using a random key to determine whether or not to flip each message bit. The original message may be retrieved from the encoded form by anyone who has access to the encoded message as well as the (secret) key. Because the encoded message still contains unflipped bits one might worry whether portions of the original message can be extracted from it. This concern was put to rest by Shannon when he proved that the encoded bit string contained no information of the original message [1]: It was indistinguishable from a random bit string. Where then does the information reside? It is neither in the encoded message nor is it in the key. Instead, all the information has been transformed into pure correlations between these two strings. How does this result apply to the black-hole information paradox? In fact, this classical result has fueled the conjecture that while black-hole information cannot strictly be found within the Hawking radiation [2], it can nonetheless be hidden within correlations between that radiation and something else [3,4].

A direct quantum analogue to the one-time pad would encode an arbitrary quantum state into the correlations between two subsystems, with none of the information about that state accessible from either subsystem alone. Interestingly, such a quantum analogue is impossible for any pure-state encoding into two subsystems [5]. For example, for the mapping

$$\alpha|0\rangle + \beta|1\rangle \rightarrow \alpha(|00\rangle - |11\rangle) + i\beta(|01\rangle + |10\rangle), \quad (1)$$

at least some information about  $\alpha$  and  $\beta$  can be gleaned by looking at either of the two final subsystems alone. That this holds generally is particularly surprising since the one-time pad is often cited as the classical analogue of quantum teleportation [6]. In quantum teleportation, Alice is given an arbitrary quantum state whose details are unknown to her. In addition, she shares one-half of an entangled state with Bob. Alice is allowed to send any classical message to

Bob after which he is to reconstruct the original state. Like the one-time pad, the shared entangled state (analogous to the secret key) contains no information about the original state. Similarly, the message Alice sends Bob (analogous to the one-time pad encoded message) contains no information about the original state. Notwithstanding this close analogy, its impossibility indicates that something must be missing from the above description. In fact, we shall see that a full description of quantum teleportation contains a third subsystem (an "environment") that allows Alice to decohere her measurements thus yielding a classical message.

Consider now an arbitrary quantum state (mixed or potentially entangled to some external reference state) which is encoded into a larger Hilbert space through some unitary process. We prove the "no-hiding theorem" in two steps: First, suppose this encoding process completely hides the information about that state from a particular subsystem of that Hilbert space (i.e., the state of that subsystem shows no dependence on the state being hidden). We prove that the hidden information is wholly encoded in the remainder of Hilbert space with no information stored in the correlations between the two subsystems [5]. Put differently, we prove that, unlike classical information, quantum mechanics allows only one way to completely hide an arbitrary quantum state from one of its subsystems: by moving it to its other subsystems. Second, and more importantly, we prove that this result is robust to imperfections in the hiding process. As more of the original state becomes hidden, it smoothly becomes more accessible in the remainder of Hilbert space in a dimensionindependent manner.

The no-hiding theorem sheds new light on the blackhole information paradox and accentuates the crisis for quantum physics. In particular, it has been speculated that at least some of the information that falls into a black hole (encoded by matter and radiation) may be found in the correlations between the Hawking radiation leaving the black hole and the black hole's internal state or lateepoch radiation [3,4,7–12]. So long as Hawking's semiclassical characterization of the black-hole radiation [2] is accurate, we prove that the quantum information about the in-fallen matter cannot be hidden in these correlations.

Perfect hiding processes.—Consider a process which takes an arbitrary input state  $\rho_I$  from subspace I and encodes it into a larger Hilbert space. This will be a hiding process if there exists some subspace O (the output) whose state  $\sigma_0$  has no dependence on the input state. In other words, our hiding process maps  $\rho_I \rightarrow \sigma_O$  with  $\sigma$  fixed for all  $\rho$ . The remainder of the encoded Hilbert space may be regarded as an ancilla A. Thus, the entire system may be represented in terms of two subsystems O and A. Now for this process to be physical, it must be linear and unitary. By linearity, it is sufficient to study the action on an arbitrary pure state  $\rho_I = |\psi\rangle_{II}\langle\psi|$ . Unitarity allows us to suitably enlarge the ancilla so that the hiding process can be represented as a mapping from pure states to pure states. The hiding process can now be expressed in terms of the Schmidt decomposition of the final state

$$|\psi\rangle_I \to \sum_{k=1}^K \sqrt{p_k} |k\rangle_O \otimes |A_k(\psi)\rangle_A.$$
 (2)

Here  $p_k$  are the K nonzero eigenvalues of  $\sigma$ ,  $\{|k\rangle\}$  are its eigenvectors, and both  $\{|k\rangle\}$  and the ancilla states  $\{|A_k\rangle\}$  are orthonormal sets.

In Eq. (2) we have explicitly allowed for a possible dependence of the ancilla states on  $|\psi\rangle$ . However, the physical nature of this hiding process places a restriction on this dependence. By linearity the ancilla will consist of an orthonormal set of states even for a superposition of inputs  $|A_k(\alpha|\psi\rangle + \beta|\psi_\perp\rangle\rangle\rangle = \alpha|A_k(\psi)\rangle + \beta|A_k(\psi_\perp)\rangle$ , where  $|\psi_\perp\rangle$  denotes any state orthogonal to  $|\psi\rangle$ . Taking the inner product between two such ancilla states yields

$$\alpha^* \beta \langle A_l(\psi) | A_k(\psi_\perp) \rangle + \beta^* \alpha \langle A_l(\psi_\perp) | A_k(\psi) \rangle = 0.$$
 (3)

Thus, for arbitrary complex values of  $\alpha$  and  $\beta$ , all crossterms above must vanish. Given any orthonormal basis  $\{|\psi_j\rangle, j=1,\ldots,d\}$  spanning the input states we may now define an orthonormal set of states,  $|A_{kj}\rangle \equiv |A_k(\psi_j)\rangle$ , spanning a Kd-dimensional Hilbert space that completely describes the reduced state of the ancilla. Unitarity allows us to map any orthonormal set into any other. Thus, we are free to write these as  $|A_{kj}\rangle = |q_k\rangle \otimes |\psi_j\rangle \oplus 0$  where  $\{|q_k\rangle\}$  is an orthonormal set of K states and  $\oplus 0$  means we pad any unused dimensions of the ancilla space by zero vectors. Under this mapping we see that the arbitrary input states  $|\psi\rangle$  are completely encoded within the ancilla and Eq. (2) becomes

$$|\psi\rangle_I \to \sum_k \sqrt{p_k} \, |k\rangle_O \otimes (|q_k\rangle \otimes |\psi\rangle \oplus 0)_A. \tag{4}$$

Since we may swap  $|\psi\rangle$  with any other state in the ancilla using purely ancilla-local operations, we conclude that any information about  $|\psi\rangle$  that is encoded globally is in fact

encoded entirely within the ancilla. No information about  $|\psi\rangle$  is encoded in system-ancilla correlations (nor, in fact, in system-system correlations).

Imperfect hiding processes.—Unlike perfect hiding processes, for which we found it sufficient to consider pure input states, imperfect hiding must allow for some imprecision in the encoding. To fully specify the mapping, we now need to describe its action on entangled states; this further guarantees that the mapping is completely positive and therefore physical.

If the input subsystem I is initially entangled with an (external) reference subsystem I' in state  $|\psi\rangle_{I'I} \equiv \sum_j \sqrt{\lambda_j} |j',j\rangle_{I'I}$  then linearity and Eq. (4) imply that a perfect hiding process on an entangled state has the form

$$|\psi\rangle_{I'I} \to |\Psi^{\text{perfect}}\rangle_{I'OA} \equiv \sum_{jk} \sqrt{\lambda_j p_k} |j'\rangle_{I'} \otimes |k\rangle_O$$
$$\otimes (|q_k\rangle \otimes |j\rangle \oplus 0)_A; \tag{5}$$

i.e., the specification we sought takes the form  $\rho_{I'I} \equiv |\psi\rangle_{I'II'I}\langle\psi| \rightarrow \rho_{I'} \otimes \sigma_O$ , where  $\rho_{I'}$  is the reduced state of the reference subsystem. An imperfect process can be described more generally by  $\rho_{I'I} \rightarrow \rho_{I'O}$  where the output only imprecisely hides the input with

$$\operatorname{Tr} |\rho_{I'O} - \rho_{I'} \otimes \sigma_O| < \epsilon, \tag{6}$$

for some  $\epsilon$ . The choice of trace norm is most appropriate since it places a bound on the probability for any observable to distinguish these states [13]. We can now use the fidelity to quantify the overlap between the global description of imperfectly hidden states and the perfect form given in Eq. (4). Since the fidelity satisfies  $F(\rho, \sigma) \ge 1 - \frac{1}{2} \operatorname{Tr} |\rho - \sigma|$ , we have

$$F(\rho_{I'O}, \rho_{I'} \otimes \sigma_O) \ge 1 - \epsilon/2. \tag{7}$$

By definition, the fidelity is the maximum overlap over all purifications of the pair of states. Equivalently, we may fix one purification and maximize the overlap based on varying the other purification [14]. Let us choose the obvious purification of  $\rho_{I'O}$ , namely, the actual global output which we denote  $|\Psi^{imperfect}\rangle$ . Further, the tensor product  $\rho_{I'}\otimes\sigma_O$  is highly restrictive and it is easy to see that *any* purification thereof must take the form of  $|\Psi^{perfect}\rangle$  (up to some unitary operation on the ancilla). As a consequence, the global state of the imperfect output will strongly overlap with some global state whose form perfectly satisfies the no-hiding theorem

$$\langle \Psi^{\text{imperfect}} | \Psi^{\text{perfect}} \rangle \ge 1 - \epsilon/2,$$
 (8)

or, stated differently,

$$|\Psi^{\text{imperfect}}\rangle = \sqrt{1 - \tilde{\epsilon}} |\Psi^{\text{perfect}}\rangle + \sqrt{\tilde{\epsilon}} |\Psi^{\text{perfect}}_{\perp}\rangle, \quad (9)$$

for some perturbation  $0 \le \tilde{\epsilon} < \epsilon$ . The demonstration of robustness to imperfections completes our proof of the no-hiding theorem.

This result comes as a surprise if we consider another, extensively studied, example of a hiding process—state randomization [15]. There it has been shown that inexact randomization of an arbitrary pure state of dimension d can be performed with an ancilla of dimension  $O(d \log d)$  whereas exact randomization requires an ancilla of dimension at least  $d^2$ . The inexact state randomization therefore cannot be expressed in general as a mere perturbation from the perfect case. By enriching the class of states to be hidden to include states which may be entangled to some reference system, we have demonstrated robustness, with a dimension-independent perturbative degradation. Indeed, this is crucial for any application where the dimensions of the various subspaces involved may be unknown and possibly infinite.

Teleportation revisited.—By the no-hiding theorem, the direct quantum analogue of the one-time pad is impossible for arbitrary input states. Notwithstanding this, quantum states can still be transformed into pure correlations between three or more subsystems [5] (not counting the external reference subsystem). This underscores the analogy between quantum teleportation and the one-time pad. To apply the no-hiding theorem to teleportation, we require a globally quantum description which we obtain by enlarging the ancilla to include the "environment" (or measurement system) [16] used to decohere Alice's Bell state. For a single qubit in an arbitrary pure state  $|\psi\rangle$ , the teleportation protocol reduces to

$$|\psi\rangle \rightarrow \frac{1}{2} \sum_{j,k=0}^{1} |2j+k\rangle_{\text{Alice}} \otimes |2j+k\rangle_{\text{message}} \otimes \sigma_z^j \sigma_x^k |\psi\rangle_{\text{Bob}}.$$
(10)

To complete the protocol, Bob need only use the value of the message to undo the randomizing operations to retrieve  $|\psi\rangle$ . It is easy to check that each of the three subsystems in Eq. (10) is in the maximally mixed state for that space. Thus, the information appears only as intersubsystem correlations. [Relabeling the subsystems of Eq. (10) yields an alternative tripartite analogue to the one-time pad [17]]. However, the above observation does not contradict the nohiding theorem.

In fact, our key result can be recovered by rewriting the teleportation process in terms of a bipartite system. For instance, since the reduced density matrix of Bob's subsystem contains no information about the hidden state  $|\psi\rangle$ , it must lie entirely in the remainder of Hilbert space. Indeed, it is easy to check that the state  $|\psi\rangle$  is completely encoded within the union of the Alice and message subsystems. (The same argument holds for Alice's subsystem or for the message subsystem.) Hence from a purely quantum mechanical perspective, teleportation is consistent with our result. Indeed, this unitary variation of the teleportation protocol could serve as an experimental verification of the no-hiding theorem, where the bipartite systems could be reconstructed separately via quantum-state to-

mography to identify in which subsystem the original qubit was encoded.

Thermodynamics.—The no-hiding theorem offers deep new insights into the nature of quantum information. In particular, it generalizes Landauer's erasure principle [18], according to which any process that erases a bit of information must dump one bit's worth of entropy into the environment. Landauer's principle applies universally to classical or quantum information [18]. However, the no-hiding theorem applies to any process hiding a quantum state, whether by erasure, randomization, thermalization or any other procedure. In this sense, quantum information hiding is equivalent to its erasure, whereas classical information hiding is fundamentally distinct from erasure.

Landauer's principle provides fundamental insight into thermodynamic reasoning, such as in the resolution of Maxwell's demon. In contrast, data hiding provides more insight into the nature of thermalization processes. The terminology used above—input, output and ancillanow takes on thermodynamic interpretations (e.g., initial system, final system, environment, or input system, output radiation, environment, etc.) In the simplest case of a single system and environment, as the state of the system thermalizes, it contracts to a thermal distribution independent of its initial description. Perfect hiding implies complete thermalization, whereas imperfect hiding may shed some light on the approach to an equilibrium state. Either way, as the state vanishes from one subspace, it must appear in the remainder of Hilbert space (i.e., the environment). To apply the no-hiding theorem, we must consider an enlarged purified environment, or superenvironment. As in teleportation, we again find ourselves with the following three subsystems: system, environment, and the remaining supra-environment. We can conclude that the quantum information that vanished is appearing somewhere in the complete environment (including correlations between the two subsystems therein).

Black-hole evaporation.—Having proved the no-hiding theorem in an abstract quantum-information theoretic setting, let us now consider its implications for information flow in and out of black holes.

Hawking's seminal work on black-hole evaporation some 30 years ago [2] precipitated a crisis in quantum physics. Hawking's calculations showed that whatever matter falls into it, a black hole evaporates in a steady stream of ideal featureless radiation. In Hawking's semiclassical analysis this radiation is completely independent of the in-falling matter, at least until the black hole has shrunk to near the Planck mass. For massive black holes (many times a Planck mass) Hawking's analysis should presumably be arbitrarily good. Nothing in Hawking's semiclassical approach changes if the black hole were created, or continued to be fed, with matter whose quantum states are entangled with external (reference) degrees of freedom. However, in such a scenario, we can immediately apply the no-hiding theorem. The in-falling matter would

correspond to subsystem I and the outgoing Hawking radiation would be subsystem O in our formulation. Thus, within the framework of Hawking's semiclassical analysis, the no-hiding theorem implies that no information is carried either within the outgoing radiation or in correlations between the outgoing radiation and anything else. This strong rejection of the correlations option is based on two assumptions alone: unitarity and Hawking's semiclassical analysis of the radiation.

We stress that the exact nature of the Hawking radiation (e.g., whether it is black body or gray body [19]) is irrelevant to our argument—in particular, it need not be thermal—so long as the reduced state of the outgoing radiation field is independent of the detailed state of the in-fallen matter. Furthermore, we note that the state of the in-fallen matter may be subject to a number of superselection rules disallowing certain superpositions. In that sense, the infallen matter is not truly in an arbitrary quantum state. Nonetheless, up to that nuance, any subspaces corresponding to the allowed superpositions must obey the no-hiding theorem.

We now expose the severity of the black-hole information crisis in one specific formulation of the paradox [20]. Suppose one feeds a black hole (with externally entangled states) at the Hawking-emission rate for an arbitrarily long time. Then, Hawking's semiclassical analysis would predict that such a black hole, of a fixed size, could contain an unbounded amount of entropy, associated with the states of the in-fallen matter. This unbounded information density is itself tantamount to a loss of unitarity (at least in our Universe) [21]. This formulation of the black-hole information paradox is particularly instructive as it applies to black holes of arbitrary size.

Naturally, one would always expect some deviations from Hawking's analysis. For instance, although a tiny effect, there should at least be some small scattering of in-falling matter off outgoing Hawking radiation. This is where robustness is key. If various perturbations lead to deviations of size  $\epsilon$  from perfect featureless radiation, then Eq. (9) quantifies the deviation away from the ideal nohiding theorem. Whether this deviation is carrying away information directly or via correlations or through interference with the main contribution is immaterial; the net amount of information that may be carried away in this manner would be  $O(\sqrt{\epsilon})$  or more likely  $O(\epsilon)$ . Since these deviations are believed to be vanishingly small for truly cosmologically sized black holes this route to even a partial resolution to the black-hole information paradox now appears untenable.

The no-hiding theorem provides new insight into the different laws governing classical and quantum information. Unlike classical bits, arbitrary quantum states cannot completely hide in correlations between a pair of subsystems. A robust statement of this result leads to a severe formulation of the black-hole information paradox: Either unitarity fails or Hawking's semiclassical predictions must

break down. The no-hiding theorem rigorously rules out any "third possibility" that the information escapes from the black hole but is nevertheless inaccessible as it is hidden in correlations between semiclassical Hawking radiation and the black hole's internal state. This provides a criterion to test any proposed resolution of the paradox: Any resolution that preserves unitarity must predict a breakdown in Hawking's analysis [2] even for cosmologically sized black holes.

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- [1] C.E. Shannon, Bell Syst. Tech. J. 28, 656 (1949).
- [2] S. W. Hawking, Nature (London) 248, 30 (1974);Commun. Math. Phys. 43, 199 (1975); Phys. Rev. D 14, 2460 (1976).
- [3] D.N. Page, Phys. Rev. Lett. 44, 301 (1980); 71, 1291 (1993); 71, 3743 (1993); Int. J. Mod. Phys. D 3, 93 (1994).
- [4] G. 't Hooft, Nucl. Phys. B256, 727 (1985); Nucl. Phys. B335, 138 (1990).
- [5] R. Cleve, D. Gottesman, and H.-K. Lo, Phys. Rev. Lett. 83, 648 (1999).
- [6] See, e.g., D. Collins and S. Popescu, Phys. Rev. A 65, 032321 (2002).
- [7] F. Wilczek, in *Black Holes, Membranes, Wormholes, and Superstrings*, edited by S. Kalara and D. V. Nanopoulos (World Scientific, Singapore, 1993), p. 1.
- [8] U. H. Danielsson and M. Schiffer, Phys. Rev. D 48, 4779 (1993).
- [9] T.M. Fiola, J. Preskill, A. Strominger, and S.P. Trivedi, Phys. Rev. D 50, 3987 (1994).
- [10] P. Kraus and F. Wilczek, Nucl. Phys. B433, 403 (1995).
- [11] S. Bose, L. Parker, and Y. Peleg, Phys. Rev. D **54**, 7490 (1996)
- [12] N. Obadia and R. Parentani, Phys. Rev. D 67, 024022 (2003).
- [13] M. Nielsen and I. Chuang, *Quantum Computation* (Cambridge University Press, Cambridge, 2000) p. 400.
- [14] B. Schumacher and M.D. Westmoreland, Quant. Info. Proc. 1, 5 (2002).
- [15] P. Hayden, D. Leung, P. W. Shor, and A. Winter, Commun. Math. Phys. 250, 371 (2004).
- [16] S. L. Braunstein, Phys. Rev. A 53, 1900 (1996).
- [17] D. W. Leung, Quantum Inf. Comput. 2, 14 (2002).
- [18] R. Landauer, IBM J. Res. Dev. 3, 183 (1961); C.H. Bennett, Int. J. Theor. Phys. 21, 905 (1982).
- [19] J. D. Bekenstein, Phys. Rev. Lett. 70, 3680 (1993).
- [20] D.N. Page, in *Proceedings of the 9th International Conference on General Relativity and Gravitation*, edited by E. Schmutzer (Friedrich Schiller University, Jena, 1980), p. 582.
- [21] J. Preskill, in *Black Holes, Membranes, Wormholes, and Superstrings*, edited by S. Kalara and D. V. Nanopoulos (World Scientific, Singapore, 1993), p. 22.