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# Global potential energy surfaces for the $\mathrm{H}_{3}^{+}$system. Analytical representation of the adiabatic ground-state $1^{1} A^{\prime}$ potential 

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#### Abstract

Adiabatic global potential energy surfaces, for singlet and triplet states of $A^{\prime}$ and $A^{\prime \prime}$ symmetries, were computed for an extensive grid for a total of 8469 conformations of $\mathrm{H}_{3}^{+}$system at full configuration interaction $a b$ initio level and using an extended basis set that has also been optimized for excited states. An accurate (root-mean-square error lower than $20 \mathrm{~cm}^{-1}$ ) global fit to the ground-state potential is obtained using a diatomics-in-molecules approach corrected by several symmetrized three-body terms with a total of 96 linear parameters and 3 nonlinear parameters. This produces an accurate global potential which represents all aspects of ground-state $\mathrm{H}_{3}^{+}$including the absolute minimum, the avoided crossing and dissociation limits, satisfying the correct symmetry properties of the system. The rovibrational eigenstates have been calculated up to total angular momentum $J=20$ using hyperspherical coordinates with symmetry adapted basis functions. The infrared spectra thus reproduced is within $1 \mathrm{~cm}^{-1}$ with respect to the experimental values for several transitions. © 2000 American Institute of Physics. [S0021-9606(00)00803-5]


## I. INTRODUCTION

Ion-molecule reactions are of great importance in gas phase environments such as molecular hydrogen plasmas occurring in interstellar clouds, planetary ionospheres, ion sources, and thermonuclear experiments. A very important reaction in the interstellar medium is the prototype reaction ${ }^{1}$

$$
\mathrm{H}_{2}^{+}+\mathrm{H}_{2} \rightarrow \mathrm{H}_{3}^{+}+\mathrm{H}
$$

forming $\mathrm{H}_{3}^{+}$, which is a major cation in hydrogen plasmas and plays an important role, due to its simplicity, as a benchmark system for high accuracy ab initio molecular theory and reaction dynamics. Full reviews of the experimental and theoretical work on $\mathrm{H}_{3}^{+}$have been presented by $\mathrm{McNab}^{2}$ and Tennyson. ${ }^{3}$ New interest has arisen in highly rotationally excited states of $\mathrm{H}_{3}^{+}$motivated by the experimental investigations of the extraordinarily complex IR predissociation spectrum-almost 27000 lines over the range 872 to 1094 $\mathrm{cm}^{-1}$, grouped into four separated peaks when considering a pseudo-low-resolution spectrum-by Carrington and Kennedy. ${ }^{4}$ Clary ${ }^{5}$ has suggested that rotational effects play a key role in explaining the temperature dependence of ionmolecule reactions, especially at the low interstellar temperatures.

Very accurate $\mathrm{H}_{3}^{+}$electronic structure calculations in the near equilibrium geometry have been reported. ${ }^{6}$ Local potential energy surfaces (LPES) on the highest level of sophistication have appeared recently, ${ }^{7,8}$ but covering only 69 points on the minimum region of the ground-state potential energy

[^0]surface, based on previous full configuration interaction (FCI) calculations due to Meyer et al. (the so-called MBB LPES). ${ }^{9}$ Dinelli et al., ${ }^{10}$ using high-resolution spectroscopic data for $\mathrm{H}_{3}^{+}, \mathrm{H}_{2} \mathrm{D}^{+}, \mathrm{D}_{2} \mathrm{H}^{+}$, and $\mathrm{D}_{3}^{+}$, have determined effective mass-dependent LPES for each isotopomer. These potential surfaces are expressed as a sum of the massindependent Born-Oppenheimer potential and a massdependent adiabatic correction. Unfortunately, the study of highly rotationally excited states near dissociation, ${ }^{11}$ with a high density of long-lived metastable levels lying in the continuum, above the lowest $\mathrm{H}_{2}+\mathrm{H}^{+}$dissociation limit (and indeed maybe lying above higher dissociation limits corresponding to fragment $\mathrm{H}_{2}$ molecules in higher rotational, or even vibrational, levels), ${ }^{12}$ is impossible to do with LPES because their validity is only for energies below the lowest dissociation limit.

Despite the high symmetry and electronic simplicity-two-electron system-of the $\mathrm{H}_{3}^{+}$ion, only a few $a b$ initio FCI calculations of the global potential energy surfaces (GPES), including the correct behavior as the molecule dissociates and for all possible geometrical configurations, have been reported for ground $\left(1^{1} A^{\prime 13,14}\right)$ and excited $\left(1^{3} \Sigma_{u}^{+},{ }^{15}\right.$ $2{ }^{1} A^{\prime 14}$ ) states and for a maximum of 680 different spatial geometries. This fact contrasts with very accurate GPES obtained for more complex polyatomics such as $\mathrm{H}_{3}$, with more than 8000 different spatial geometries, ${ }^{16} \mathrm{H}_{2} \mathrm{O},{ }^{17,18} \mathrm{H}_{2} \mathrm{~F},{ }^{19}$ or $\mathrm{H}_{4},{ }^{20}$ to name just a few. Moreover, in a recent paper concerning a first attempt on a calculated spectrum for neardissociation $\mathrm{H}_{3}^{+}$, Henderson and Tennyson conclude that, in the absence of a high quality global potential, an attack on the global $\mathrm{H}_{3}^{+}$problem would be very worthwhile and is easily within the range of present methods. ${ }^{21}$ In this paper we
present FCI calculations for the global $\mathrm{H}_{3}^{+}$system including 8469 different spatial geometries and, for each geometry, we compute a total of 36 states of $A^{\prime}$ and $A^{\prime \prime}$ irreducible representations with both singlet and triplet multiplicities.

There are several global analytical representations of the adiabatic ground-state $1{ }^{1} A^{\prime}$ potential for $\mathrm{H}_{3}^{+}$in the literature. ${ }^{13,22,23}$ However, all of these GPES are unsatisfactory for several reasons: the diatomics-in-molecules (DIM) surface of Preston and Tully, ${ }^{22}$ which is the most widely used in trajectory calculations, is qualitatively correct; it was the first study of the avoided crossing due to the presence of two dissociation channels, $\mathrm{H}_{2}\left({ }^{1} \Sigma_{g}^{+}\right)+\mathrm{H}^{+}$and $\mathrm{H}_{2}^{+}\left({ }^{2} \Sigma_{u}^{+}\right)+\mathrm{H}$, but it is not accurate even at low energies; the Schinke et al. GPES ${ }^{13}$ is based on their ab initio calculations but the functional form used has discontinuous derivatives and contains unphysically deep minima for certain regions, probably due to the small number of $a b$ initio points. The most recent GPES of Prosmiti et al. ${ }^{23}$ is a combination of two potential forms using the energy switching approach of Varandas ${ }^{24}$ to connect them. The first potential form, corresponding to the Born-Oppenheimer portion of the Dinelli et al. LPES, ${ }^{10}$ reproduces the spectroscopic measurements with quantitative accuracy and is reliable for the minimum region of the global potential. The second potential form is represented by two terms, a short-range 32 parameters fit to 327 data points of Schinke et al. (with a standard deviation of about $280 \mathrm{~cm}^{-1}$ ) plus a long-range term to describe the charge-induced dipole and charge-quadrupole contributions obtained from perturbation theory. Unfortunately, this latter term does not reflect the symmetry of $\mathrm{H}_{3}^{+}$causing the global potential to contain unphysical behavior with respect to symmetry. In addition, the connection regions between the two potentials of very different accuracy may also be a problem for dynamical calculations.

The construction of an accurate GPES covering the whole configuration space up to and above dissociation, for the ground-state $\mathrm{H}_{3}^{+}$system, remains an important problem. In fact, the most recent dynamical study of this system using a quantum-mechanical approach ${ }^{25}$ is based on the very approximate DIM GPES. ${ }^{22}$ Moreover, in a first attempt to obtain a calculated spectrum for near-dissociation $\mathrm{H}_{3}^{+},{ }^{21}$ Henderson and Tennyson have used the MBB LPES ${ }^{9}$ that lacks any representation of the regions near dissociation. In this paper we also present a global analytical representation of the adiabatic ground-state $1{ }^{1} A^{\prime}$ potential for the $\mathrm{H}_{3}^{+}$system. Moreover, we have calculated the rovibrational levels of the ground-state $\mathrm{H}_{3}^{+}$system, using symmetry adapted basis functions in hyperspherical coordinates. The use of symmetry as well as an iterative Lanczos procedure allowed us to calculate levels up to an energy of about $14000 \mathrm{~cm}^{-1}$ and with high angular momentum $(J=20)$. The results thus obtained are used to check the accuracy of the GPES reported here, by comparison with the experimental infrared spectra as well as with previous theoretical studies of spectroscopic accuracy.

## II. POTENTIAL ENERGY CALCULATIONS

The construction of a reliable GPES faces the problem that the absolute error of the energies, calculated with varia-
tional or perturbative procedures, is large. For the $\mathrm{H}_{3}^{+}$, a full configuration interaction (FCI) solution is trivial and nowadays it is possible to do the computational effort needed to obtain several thousands of FCI energies with an extended basis set. The single best FCI energy obtained for the $\mathrm{H}_{3}^{+}$ minimum ${ }^{9}$ is about $95 \mathrm{~cm}^{-1}$ higher than the very accurate result reported recently. ${ }^{26,7}$

Nevertheless, when constructing a GPES, we are not interested in total energies but in energy differences, which are usually small quantities. In forming the difference between two large quantities of similar size, we must consider that errors at each term of the difference are similar and will cancel to a large extent, with a final result for the difference that may have a lower error (with respect to the exact difference) than the absolute error obtained in the calculated total energy. Usually, the error in the energy differences involving electronic excited states increases with respect to that obtained for the ground state. This is due to the fact that basis sets are optimized, in general, for atomic ground state that correlate with molecular ground state and several molecular excited states. If we compute molecular excited states correlating with atomic excited states, then we need atomic basis sets optimized for both atomic ground state and several excited states.

In order to assess the relative accuracy which can be expected using different basis sets, we begin by comparing results of calculations for $\mathrm{H}_{2}$. We report FCI calculations obtained using the $(10 s 4 p 2 d) /[7 s 4 p 2 d]$ basis of Meyer, Botschwina, and Burton (MBB) ${ }^{9}$ and a new basis set $(11 s 6 p 2 d) /[8 s 6 p 2 d]$ obtained from the MBB basis set which is further augmented by a single $s$ function with exponent 0.0126498 , that has been optimized with respect to the energy of the hydrogen $2 s$ orbital, and two sets of $p$ functions with exponents 0.0455576 and 0.0177738 , that have been optimized with respect to the energy of the hydrogen $2 p$ orbitals. Figures 1 and 2 display errors in energy differences as a function of the internuclear distance (taking as reference the $\mathrm{H}_{2}$ ground-state minimum energy), for singlet states and triplet states, respectively. We have taken as the exact energy differences those obtained from the most accurate potential curves for each state considered ${ }^{27}$ (labeled $X^{1} \Sigma_{g}^{+}, E, F^{1} \Sigma_{g}^{+}, B^{1} \Sigma_{u}^{+}, B^{\prime} \Sigma_{u}^{+}, C^{1} \Pi_{u}$ in Fig. 1 and $a^{3} \Sigma_{g}^{+}, b^{3} \Sigma_{u}^{+}, e^{3} \Sigma_{u}^{+}, i^{3} \Pi_{g}$ in Fig. 2).

From Fig. 1 we can see that errors in energy differences are very similar for the ground state $\left(X^{1} \Sigma_{g}^{+}\right)$using both basis sets (see the upper panel for present results and the lower panel for MBB results). The same behavior is observed in Fig. 2 for errors corresponding to the $b^{3} \Sigma_{u}^{+}$excited state that are very similar for both basis sets. These are the only two states ( $X^{1} \Sigma_{g}^{+}$and $b^{3} \Sigma_{u}^{+}$) of $\mathrm{H}_{2}$ having as asymptotic limits $\mathrm{H}(1 s)+\mathrm{H}(1 s)$. However, if we compare errors in energy differences for any other $\mathrm{H}_{2}$ excited state, with different asymptotic limits, we can see a general worse behavior for MBB basis set in both figures (several states fall out of the limits in both figures for all internuclear distances using the MBB basis set). Moreover, in Fig. 1 we can see that errors corresponding to the $\mathrm{H}_{2}$ ground state $\left(X^{1} \Sigma_{g}^{+}\right)$are lower than $20 \mathrm{~cm}^{-1}$ for distances between approximately 1.0


FIG. 1. Errors in energy differences for different singlet states of the $\mathrm{H}_{2}$ molecule as a function of the internuclear distance. Errors in $\mathrm{cm}^{-1}$ and distances in atomic units. Upper panel corresponds to present results and lower panel corresponds to Meyer et al. (Ref. 9) results.
and 3.5 atomic units, corresponding to the potential well region. From Fig. 1 we can see that the error at dissociation for the $\mathrm{H}_{2}$ ground state ( $X^{1} \Sigma_{g}^{+}$) is about $145 \mathrm{~cm}^{-1}$; this dissociation error affects the remaining error curves as we can clearly see for several triplet states in Fig. 2 (see $a^{3} \Sigma_{g}^{+}$, $b^{3} \Sigma_{u}^{+}$, and $e^{3} \Sigma_{u}^{+}$, upper panel), which show very stable error curves but displaced from zero error in about 145 $\mathrm{cm}^{-1}$. If we compute these error curves with respect to their own minima, the resulting curves are all around the zero value error in energy differences.

We have carried out a similar study for the ground state $\left(X^{2} \Sigma_{g}^{+}\right)$and several excited states $\left(2^{2} \Sigma_{g}^{+}, 1^{2} \Sigma_{u}^{+}, 2^{2} \Sigma_{u}^{+}\right.$, $1{ }^{2} \Pi_{u}, 1^{2} \Pi_{g}$ ) of the $\mathrm{H}_{2}^{+}$. In the upper panel of Fig. 3 the errors in energy differences for all the mentioned doublet states are compared with the exact energy differences. The latter have been obtained from the exact potential energy curves. ${ }^{28}$ The corresponding results using the MBB basis set have been plotted in the bottom panel of Fig. 3. As for the $\mathrm{H}_{2}$ system, there are several states $\left(1^{2} \Pi_{u}, 1^{2} \Pi_{g}\right)$ that fall out of the figure using the same energy scale as that used for our results. Moreover, using the MBB basis set, only two states $\left(X^{2} \Sigma_{g}^{+}\right.$and $1^{2} \Sigma_{u}^{+}$, see the bottom panel in Fig. 3) are in error lower than $20 \mathrm{~cm}^{-1}$ in the region of the potential well, while when using the basis set proposed in this paper all the states except $1^{2} \Pi_{u}$ fulfill this condition. Finally, one important difference with respect to $\mathrm{H}_{2}$ is that for $\mathrm{H}_{2}^{+}$we have a very low error for dissociation.

Using the ( $11 s 6 p 2 d$ ) basis set, with the four innermost


FIG. 2. Errors in energy differences for different triplet states of the $\mathrm{H}_{2}$ molecule as a function of the internuclear distance. Units and panels as in Fig. 1.
$s$ functions contracted to [ $8 s 6 p 2 d$ ], the total energy of $\mathrm{H}_{3}^{+}$ at its ground-state equilibrium geometry (equilateral triangle, equilibrium bond length $R_{e}=1.6500$ a.u.) is -1.343100 a.u., about $161 \mathrm{~cm}^{-1}$ above the exact energy. ${ }^{7,26}$ However, as we have stressed above, this absolute error is not as important as the error in energy differences with respect to a reference zero energy. Since we have computed the very accurate 69 points quoted recently by Cencek et al., ${ }^{7}$ it is possible to obtain a root-mean-square (rms) deviation of our energy differences errors, taken as zero energy value the corresponding energy of the equilibrium geometry both for our energy differences (zero energy at -1.343100 a.u.) as for the exact energy differences (zero energy at -1.343835 a.u.). The resulting rms deviation is less than $17 \mathrm{~cm}^{-1}$. Therefore, the expected rms errors for the $\mathrm{H}_{3}^{+}$ground-state well and its dissociation channels, $\mathrm{H}_{2}\left(X^{1} \Sigma_{g}^{+}\right)+\mathrm{H}^{+}$and $\mathrm{H}(1 s)$ $+\mathrm{H}_{2}^{+}\left(X^{2} \Sigma_{g}^{+}\right)$, are about $20 \mathrm{~cm}^{-1}$. Only the full dissociation region of the $\mathrm{H}_{3}^{+}$ground-state GPES $[\mathrm{H}(1 s)+\mathrm{H}(1 s)$ $+\mathrm{H}^{+}$] is expected to yield a higher error, growing smoothly to a maximum value of $145 \mathrm{~cm}^{-1}$ as the internuclear separations increase.

We use the $(11 s 6 p 2 d) /[8 s 6 p 2 d]$ basis set to compute the final FCI data points for a total of 8469 different $\mathrm{H}_{3}^{+}$ conformations. We have used symmetry group $C_{s}$ for all the geometries and we have computed 36 different states at each point as follows: $9{ }^{1} A^{\prime}, 9{ }^{1} A^{\prime \prime}, 9^{3} A^{\prime}$ and $9^{3} A^{\prime \prime}$. To specify our grids of $\mathrm{H}_{3}^{+}$conformations, we have adopted the coordinates used by Boothroyd et al. ${ }^{16}$ in a refined $\mathrm{H}_{3}$ potential energy surface. These coordinates can be described by the


FIG. 3. Errors in energy differences for different doublet states of the $\mathrm{H}_{2}^{+}$ molecule as a function of the internuclear distance. Units and panels as in Fig. 1.
shortest interatomic distance $r_{1}$, the next-shortest distance $r_{2}$, and the exterior angle between them, $\theta$.

We have used the preliminary grid described by Boothroyd et al. ${ }^{16}$ that comprised 540 conformations. Moreover, we have used also the more comprehensive grid that was specified as follows: $r_{1}$ and $r_{2}$ were chosen from 0.6 to 2.0 $a_{0}$ in increments of $0.1 a_{0}\left(1 a_{0}=0.529177 \AA\right)$, also for $2.15,2.3,2.45,2.6,2.8,3.0,3.2,3.4,3.7,4.0,4.3,4.6,5.0$, and $5.5 a_{0}$, such that $r_{1} \leqslant r_{2} \leqslant r_{3} ; \theta$ ran from $0^{\circ}$ to $90^{\circ}$ inclusive in increments of $10^{\circ}$ and continued from $\theta=90^{\circ}$ to $\theta \leqslant 120^{\circ}$ in one, two, or three equally spaced increments that did not exceed $10^{\circ}$. More conformations have been added to the "comprehensive", grid as described by Boothroyd et al. ${ }^{16}$ totalizing 6548 conformations. Some of them are coincident with the preliminary grid.

We have added conformations corresponding to the $\mathrm{H}_{3}$ van der Waals region, ${ }^{29}$ filling in a region at large $\mathrm{H}-\mathrm{H}_{2}$ distances from 5.5 to $15.0 a_{0}$ with $r_{1}$ chosen from 0.7 to 4.0 $a_{0}$. The total number of different conformations was 7995 including the van der Waals grid.

Finally, we include a grid of conformations that are very close equilateral triangles (close to the $D_{3 h}$ minimum region) that we specify as follows: $r_{1}$ and $r_{2}$ were chosen from 1.61 to $1.69 a_{0}$ in increments of $0.01 a_{0}$ such that $r_{1} \leqslant r_{2} ; \theta$ ran from $56^{\circ}$ to $64^{\circ}$ inclusive in increments of $1^{\circ}$, totalizing 405 different conformations. If we add to all of these conformations the 69 proposed by Meyer et al. ${ }^{9}$ and used by Cencek et al. ${ }^{7}$ in their very accurate $\mathrm{H}_{3}^{+}$ground-state calculations, we obtain a total number of 8469 conformations and 36 FCI
energy values for each one (304884 energy values that we do not quote here for obvious reasons). A file containing the 8469 ground-state $\mathrm{H}_{3}^{+}$data points used to obtain the GPES reported in this paper has been placed in the electronic depository EPAPS. ${ }^{30}$

## III. THE GROUND-STATE $H_{3}^{+}$GLOBAL SURFACE

We write the global potential energy surface corresponding to the $\mathrm{H}_{3}^{+}$ground state $\left(1^{1} A^{\prime}\right)$ as

$$
\begin{equation*}
V_{\mathrm{H}_{3}^{+}}=V_{\mathrm{DIM}}+\sum_{L}^{\mathrm{LMAX}} V_{A B C}^{(3) L}\left(R_{A B}, R_{A C}, R_{B C}\right) \tag{1}
\end{equation*}
$$

where $V_{\text {DIM }}$ in Eq. (1) is the lowest eigenvalue of the symmetric $3 \times 3$ matrix, corresponding to the diatomics-inmolecules approach with neglected overlap, given by

$$
\begin{aligned}
H_{11}= & V_{A B}^{(2)}\left(H_{2},{ }^{1} \Sigma_{g}^{+}\right)-2 V_{H}^{(1)}+\frac{1}{2}\left[V_{A C}^{(2)}\left(H_{2}^{+},{ }^{2} \Sigma_{g}^{+}\right)\right. \\
& +V_{A C}^{(2)}\left(H_{2}^{+},{ }^{2} \Sigma_{u}^{+}\right)+V_{B C}^{(2)}\left(H_{2}^{+},{ }^{2} \Sigma_{g}^{+}\right) \\
& \left.+V_{B C}^{(2)}\left(H_{2}^{+},{ }^{2} \Sigma_{u}^{+}\right)\right], \\
H_{22}= & V_{A C}^{(2)}\left(H_{2},{ }^{1} \Sigma_{g}^{+}\right)-2 V_{H}^{(1)}+\frac{1}{2}\left[V_{A B}^{(2)}\left(H_{2}^{+},{ }^{2} \Sigma_{g}^{+}\right)\right. \\
& +V_{A B}^{(2)}\left(H_{2}^{+},{ }^{2} \Sigma_{u}^{+}\right)+V_{B C}^{(2)}\left(H_{2}^{+},{ }^{2} \Sigma_{g}^{+}\right) \\
& \left.+V_{B C}^{(2)}\left(H_{2}^{+},{ }^{2} \Sigma_{u}^{+}\right)\right], \\
H_{33}= & V_{B C}^{(2)}\left(H_{2},{ }^{1} \Sigma_{g}^{+}\right)-2 V_{H}^{(1)}+\frac{1}{2}\left[V_{A B}^{(2)}\left(H_{2}^{+},{ }^{2} \Sigma_{g}^{+}\right)\right. \\
& +V_{A B}^{(2)}\left(H_{2}^{+},{ }^{2} \Sigma_{u}^{+}\right)+V_{A C}^{(2)}\left(H_{2}^{+},{ }^{2} \Sigma_{g}^{+}\right) \\
& \left.+V_{A C}^{(2)}\left(H_{2}^{+},{ }^{2} \Sigma_{u}^{+}\right)\right], \\
H_{12}= & \frac{1}{2}\left[V_{B C}^{(2)}\left(H_{2}^{+},{ }^{2} \Sigma_{u}^{+}\right)-V_{B C}^{(2)}\left(H_{2}^{+},{ }^{2} \Sigma_{g}^{+}\right)\right], \\
H_{13}= & \frac{1}{2}\left[V_{A C}^{(2)}\left(H_{2}^{+},{ }^{2} \Sigma_{u}^{+}\right)-V_{A C}^{(2)}\left(H_{2}^{+},{ }^{2} \Sigma_{g}^{+}\right)\right], \\
H_{23}= & \frac{1}{2}\left[V_{A B}^{(2)}\left(H_{2}^{+},{ }^{2} \Sigma_{u}^{+}\right)-V_{A B}^{(2)}\left(H_{2}^{+},{ }^{2} \Sigma_{g}^{+}\right)\right],
\end{aligned}
$$

$V_{H}^{(1)}$ being the energy of the ${ }^{2} S$ state of $\mathrm{H}(1 s)$ atom $(-0.5$ a.u. or $-109737 \mathrm{~cm}^{-1}$ ).

The two-body energies $V_{A B}^{(2)}$ (including the nuclear repulsion), may be written as a sum of two terms

$$
\begin{equation*}
V_{A B}^{(2)}=V_{\text {short }}+V_{\text {long }}, \tag{2}
\end{equation*}
$$

where $V_{\text {short }}$ prevails for small internuclear distances (the short range of the potential), which must fulfill the next condition

$$
\lim _{R_{A B} \rightarrow 0} V_{A B}^{(2)}=\lim _{R_{A B} \rightarrow 0} V_{\text {short }} \rightarrow \infty .
$$

A simple alternative is to choose for $V_{\text {short }}$ the shielded Coulomb potential

$$
\begin{equation*}
V_{\text {short }}=c_{0} \frac{e^{-\alpha R_{A B}}}{R_{A B}}, \quad\left(\alpha, c_{0}>0\right) \tag{3}
\end{equation*}
$$

We choose for $V_{\text {long }}$ an expansion of order $I$

$$
\begin{equation*}
V_{\text {long }}=\sum_{i=1}^{I} c_{i} \rho_{A B}^{i}, \tag{4}
\end{equation*}
$$

TABLE I. Two-body ${ }^{\text {a }}$ term $V^{(2)}\left(\mathrm{H}_{2},{ }^{1} \Sigma_{g}^{+}\right)$.

| $i$ | $c_{i}$ |
| :---: | ---: |
| 0 | $0.101395751(+01)$ |
| 1 | $-0.526324017(+01)$ |
| 2 | $0.125893268(+03)$ |
| 3 | $-0.338282397(+04)$ |
| 4 | $0.709290230(+05)$ |
| 5 | $-0.107000112(+07)$ |
| 6 | $0.113271582(+08)$ |
| 7 | $-0.828291974(+08)$ |
| 8 | $0.408176908(+09)$ |
| 9 | $-0.129107831(+10)$ |
| 10 | $0.236346508(+10)$ |
| 11 | $-0.190174039(+10)$ |
| $\alpha_{\text {HH }}$ | $0.192102300(+01)$ |
| $\beta_{\mathrm{HH}}^{(2)}$ | $0.177626800(+01)$ |

${ }^{\mathrm{a}}$ All the coefficients are given in atomic units.
in functions $\rho_{A B}$ fulfilling that $V_{\text {long }}$ tends to zero when the internuclear distance $R_{A B}$ tends to zero or to infinity. Although the first condition is not strictly necessary, it allows us to eliminate possible oscillations of the potential in regions of small internuclear distances. We have found that the functions introduced by Rydberg ${ }^{31}$

$$
\begin{equation*}
\rho_{A B}\left(R_{A B}\right)=R_{A B} e^{-\beta_{A B}^{(N)} R_{A B}}, \quad\left(\beta_{A B}^{(N)}>0\right), \tag{5}
\end{equation*}
$$

yield high accuracy fits of potential energy curves ${ }^{32}$ for diatomic molecules.

The linear parameters $c_{i}, i=0,1, \ldots, I$ and the nonlinear parameters $\alpha$ and $\beta_{A B}^{(2)}$ are determined by fitting the ab initio energies of the diatomic fragments for all the states considered in Eq. (1), computed using the same hydrogen basis set as for the triatomic system and using the same $a b$ initio FCI procedure. In Tables I-III we report the parameters corresponding to the diatomic potentials for all the states needed to construct the DIM surface $\left(\mathrm{H}_{2}: X^{1} \Sigma_{g}^{+}, \mathrm{H}_{2}^{+}: X^{2} \Sigma_{g}^{+}\right.$and $\left.1^{2} \Sigma_{u}^{+}\right)$. Moreover, it is also feasible to construct higher order DIM matrices by using more diatomic potentials, corresponding to excited states of $\mathrm{H}_{2}$ and $\mathrm{H}_{2}^{+}$, that have been calculated in the preceding section. This possibility has the advantage of a larger initial approximation to the GPES,

TABLE II. Two-body ${ }^{\text {a }}$ term $V^{(2)}\left(\mathrm{H}_{2}^{+},{ }^{2} \Sigma_{g}^{+}\right)$.

| $i$ | $c_{i}$ |
| :---: | ---: |
| 0 | $0.102153115(+01)$ |
| 1 | $-0.454659563(+00)$ |
| 2 | $-0.897629388(+00)$ |
| 3 | $0.247689483(+02)$ |
| 4 | $-0.265726684(+03)$ |
| 5 | $0.172217757(+04)$ |
| 6 | $-0.659965030(+04)$ |
| 7 | $0.120998241(+05)$ |
| 8 | $0.530127306(+04)$ |
| 9 | $-0.704964844(+05)$ |
| 10 | $0.125717969(+06)$ |
| 11 | $-0.763708295(+05)$ |
| $\alpha_{\mathrm{HH}^{+}}^{(2)}$ | $0.167592900(+01)$ |
| $\beta_{\mathrm{HH}^{+}}$ | $0.911150800(+00)$ |

${ }^{\mathrm{a}}$ All the coefficients are given in atomic units.

TABLE III. Two-body ${ }^{\text {a }}$ term $V^{(2)}\left(\mathrm{H}_{2}^{+},{ }^{2} \Sigma_{u}^{+}\right)$.

| $i$ | $c_{i}$ |
| :---: | ---: |
| 0 | $0.365653648(+03)$ |
| 1 | $0.139647152(+01)$ |
| 2 | $-0.161982515(+02)$ |
| 3 | $0.606990392(+03)$ |
| 4 | $-0.133254221(+05)$ |
| 5 | $0.189935494(+06)$ |
| 6 | $-0.177898791(+07)$ |
| 7 | $0.110321235(+08)$ |
| 8 | $-0.447664035(+08)$ |
| 9 | $0.114117310(+09)$ |
| 10 | $-0.165634015(+09)$ |
| 11 | $0.104351423(+09)$ |
| $\alpha_{\text {H }^{+}}^{(2)}$ | $0.938759600(+01)$ |
| $\beta_{\mathrm{HH}^{+}}$ | $0.112172600(+01)$ |

${ }^{\mathrm{a}}$ All the coefficients are given in atomic units.
mainly for the dissociation channels including the long-range behavior, and the disadvantage of a longer time consumed in calculating points of the final GPES.

For the three-body terms of the global potential, $V_{A B C}^{(3) L}$ in Eq. (1), we choose an expansion of order $K$ in product functions that decays exponentially with the distance. Therefore, the $V_{A B C}^{(3) L}$ is neglected at all the dissociation limits and when the internuclear distances tends to zero

$$
\begin{equation*}
V_{A B C}^{(3) L}\left(R_{A B}, R_{A C}, R_{B C}\right)=\sum_{i j k}^{K} d_{i j k}^{L} \rho_{A B}^{i} \rho_{A C}^{j} \rho_{B C}^{k} . \tag{6}
\end{equation*}
$$

The variables $\rho_{A B}, \rho_{A C}$, and $\rho_{B C}$ are the modified form of the Rydberg functions used to represent the long-range term of the two-body potential [see Eq. (5)], but with different exponential parameters

$$
\begin{equation*}
\rho_{A B}\left(R_{A B}\right)=R_{A B} e^{-\beta_{A B}^{(3) L} R_{A B}}, \quad\left(\beta_{A B}^{(3) L}>0\right) \tag{7}
\end{equation*}
$$

In order to avoid the inclusion of terms that depend on only one interatomic distance in Eq. (7), which have been included in the two-body contributions, one must impose several constraints. ${ }^{32}$

If the system under consideration has three identical atoms, further constraints in the linear $d_{i j k}^{L}$ [Eq. (6)] and nonlinear $\beta_{A B}^{(3) L}$ [Eq. (7)] parameters must be added to ensure that the global potential is invariant with respect to permutations of all the equivalent nuclei. ${ }^{32}$

In Table IV we present the rms values for different fits of the global $\mathrm{H}_{3}^{+}$ground-state potential using one or several three-body terms as indicated by the LMAX value in the corresponding column in this table [see also Eq. (1); in fact, we fit the three-body terms to $V_{\mathrm{H}_{3}^{+}}-V_{\mathrm{DIM}}$, where $V_{\mathrm{H}_{3}^{+}}$are the 8469 calculated data points]. We can see that, when only one three-body term is considered (see $L M A X=1$ column), the accuracy of the fit cannot reach the accuracy of the data points (which we have estimated at about $20 \mathrm{~cm}^{-1}$ ), even for high order expansions and a great number of linear parameters, the convergence being very slow. The reason for this behavior is that the functional form of the three-body term is unable to reproduce the long-range part of the global potential. However, when the accuracy of the data points is

TABLE IV. Accuracy vs the order of the fit.

| K | $L M A X=1^{\text {a }}$ |  | $L M A X=2^{\text {a }}$ |  | $L M A X=3^{\text {a }}$ |  | $L M A X=4^{\text {a }}$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $n_{\text {par }}{ }^{\text {b }}$ | $\mathrm{rms}^{\text {c }}$ | $n_{\text {par }}{ }^{\text {b }}$ | $\mathrm{rms}^{\text {c }}$ | $n_{\text {par }}{ }^{\text {b }}$ | $\mathrm{rms}^{\text {c }}$ | $n_{\text {par }}{ }^{\text {b }}$ | $\mathrm{rms}^{\text {c }}$ |
| 3 | 3 | 2612.32 | 6 | 1311.93 | 9 | 509.59 | 12 | 453.98 |
| 4 | 6 | 2536.07 | 12 | 352.55 | 18 | 217.55 | 24 | 115.42 |
| 5 | 10 | 800.24 | 20 | 211.60 | 30 | 104.58 | 40 | 50.71 |
| 6 | 16 | 741.13 | 32 | 121.71 | 48 | 38.47 | 64 | 30.85 |
| 7 | 23 | 397.67 | 46 | 83.24 | 69 | 27.63 | 92 | 18.51 |
| 8 | 32 | 367.59 | 64 | 62.26 | 96 | 18.56 | 128 | 14.21 |
| 9 | 43 | 283.65 | 86 | 57.01 | 129 | 15.04 | 172 | 12.45 |
| 10 | 56 | 265.81 | 112 | 51.06 | 168 | 12.94 | 224 | 10.51 |
| 11 | 71 | 249.03 | 142 | 44.77 | 213 | 11.28 |  |  |
| 12 | 89 | 226.64 | 178 | 38.82 |  |  |  |  |
| 13 | 109 | 207.40 | 218 | 32.88 |  |  |  |  |
| 14 | 132 | 178.03 |  |  |  |  |  |  |
| 15 | 158 | 165.43 |  |  |  |  |  |  |
| 16 | 187 | 144.80 |  |  |  |  |  |  |
| 17 | 219 | 131.16 |  |  |  |  |  |  |

${ }^{\mathrm{a}} L M A X$ is the number of nonlinear parameters of the fit.
${ }^{\mathrm{b}} n_{\text {par }}$ is the number of linear parameters of the fit.
${ }^{\mathrm{c}}$ Root-mean-square (rms) errors in $\mathrm{cm}^{-1}$.
yexpected to be greater than $300 \mathrm{~cm}^{-1}$ (as usually occurs for the published GPES), this is not an important problem, because with only one three-body term we are able to obtain rms errors lower than the accuracy of the data. In this case we have estimated a better accuracy of the data points. Increasing the number of three-body terms, a better global fit can be reached with much fewer linear parameters. In fact, values of the rms lower than that of the data points can be achieved, as illustrated in Table IV ( $L M A X=2,3,4$ columns). We can also see that the convergence is faster in this case. Moreover, when we go from $L M A X=3$ to $L M A X$ $=4$, we can see that, for a given number of linear parameters, we obtain a similar rms indicating that the process is also convergent with respect to the addition of more threebody terms. This result indicates that a linear combination of different three-body terms is able to reproduce the longrange part of the global potential. Therefore, we select as the final fit that underlined in Table IV $(K=8, L M A X=3)$, corresponding to 96 linear parameters and three nonlinear parameters with an rms value of $18.56 \mathrm{~cm}^{-1}$ that is similar to that estimated for the data points. In Table V we collect the parameters corresponding to this 'final', fit. However, as we can see from Table IV, the procedure presented here is able to attain lower rms errors for the global fit. This is a very important result because if a great number of very accurate data points are obtained, ${ }^{33}$ such a procedure would be needed to produce a very accurate analytical GPES. A FORTRAN program to generate the final GPES is available from the authors upon request. ${ }^{34}$

To verify the accuracy of the final GPES we have compared some points with the exact Born-Oppenheimer ones given by Röhse et al. ${ }^{35}$ The exact energy for dissociation into $\mathrm{H}_{2}+\mathrm{H}^{+}$is $37170 \mathrm{~cm}^{-1}$, just $15 \mathrm{~cm}^{-1}$ higher than our potential energy ( $37155 \mathrm{~cm}^{-1}$ ). The exact dissociation limit corresponding to an elongated $\mathrm{H}_{2}\left(R=2.50 a_{0}\right)+\mathrm{H}^{+}$is $54845 \mathrm{~cm}^{-1}$, just $16 \mathrm{~cm}^{-1}$ higher than our potential energy ( $54829 \mathrm{~cm}^{-1}$ ). However, we must stress that the preceding exact value corresponds to the crossing point between the $\mathrm{H}_{2}$
$\left(R=2.50 a_{0}\right)+\mathrm{H}^{+} \quad$ and $\mathrm{H}_{2}^{+} \quad\left(R=2.50 a_{0}\right)+\mathrm{H}(1 s)$ curves, while in our potential we obtain the corresponding crossing point at $R=2.49 a_{0}$. Therefore, our dissociation limit ( $54711 \mathrm{~cm}^{-1}$ ) corresponding to an elongated $\mathrm{H}_{2}^{+}(R$ $\left.=2.50 a_{0}\right)+\mathrm{H}(1 s)$ underestimates the exact value given above by $134 \mathrm{~cm}^{-1}$. The exact barrier to linearization is $14299 \mathrm{~cm}^{-1}$, which is overestimated by our potential by 2 $\mathrm{cm}^{-1}\left(14301 \mathrm{~cm}^{-1}\right)$. Finally, our final GPES underestimates the $\mathrm{H}(1 s)+\mathrm{H}(1 s)+\mathrm{H}^{+}$dissociation energy by about $160 \mathrm{~cm}^{-1}$, corresponding to the region of very high energies ( $75301 \mathrm{~cm}^{-1}$ ).

In Fig. 4 potential energy contours of the $\mathrm{H}_{3}^{+}$groundstate GPES have been plotted using Jacobi coordinates in which $\mathbf{r}$ is the $\mathrm{H}_{2}$ internuclear vector, $\mathbf{R}$ is the vector joining the center of mass of $\mathrm{H}_{2}$ to the remaining H , atom and $\Theta$ is the angle between them. The GPES corresponding to the present results have been plotted for three different $\Theta$ values $\left(0^{\circ}, 45^{\circ}\right.$, and $90^{\circ}$, see left-hand panels in Fig. 4). We have also plotted, in the right-hand panels of the same figure and for the same $\Theta$ angles, the $\mathrm{H}_{3}^{+}$ground-state GPES obtained previously by Prosmiti et al. ${ }^{23}$ As we can see from the comparison of both GPES in Fig. 4, in both cases the minimum regions, corresponding to short $\mathbf{r}$ and $\mathbf{R}$ distances, are identical. However, if we fix our attention to long $\mathbf{r}$ and $\mathbf{R}$ distances, we can observe two very important differences. The first one is the presence of undesirable structures corresponding to nonsmooth contours that can produce discontinuous derivatives of the Prosmiti et al. GPES. ${ }^{23}$ The second one is the lack of symmetry produced in the Prosmiti et al. GPES, ${ }^{23}$ as we can see clearly in the right-hand upper panel in Fig. 4 if we pay attention to the different appearance of the dissociation channels, that should be identical to the present results (see left-hand upper panel in Fig. 4).

Furthermore, in Fig. 5 we have plotted the same GPES as in Fig. 4, but now using a "stereographic projection" in hyperspherical coordinates. ${ }^{36}$ The three hyperspherical coordinates are $\rho, \theta$, and $\phi_{\tau} .{ }^{36}$ The coordinate $\rho$ can be said to

TABLE V. Parameters of the three-body ${ }^{\text {a }}$ terms $V^{(3) L}$.

|  | $L=1$ | $L=2$ | $L=3$ |
| :---: | :---: | :---: | :---: |
| $i j k$ | $d_{i j k}^{1}$ | $d_{i j k}^{2}$ | $d_{i j k}^{3}$ |
| 110 | $0.121717530(+00)$ | $0.346021655(+01)$ | $0.466629318(+02)$ |
| 111 | -0.221 $211994(+01)$ | $-0.907685339(+02)$ | -0.190 $559594(+04)$ |
| 210 | $-0.215351596(+01)$ | $-0.123551634(+02)$ | $-0.285432826(+03)$ |
| 211 | $0.164527455(+02)$ | $0.205302095(+03)$ | $0.141983537(+05)$ |
| 220 | $0.457345802(+00)$ | $0.230134235(+03)$ | $0.481335752(+04)$ |
| 310 | $0.684755535(+01)$ | $-0.946613730(+01)$ | $-0.378444681(+04)$ |
| 221 | $0.323710746(+02)$ | $-0.219021561(+03)$ | -0.926 $114214(+05)$ |
| 311 | $-0.107749346(+03)$ | $-0.637342471(+03)$ | -0.581 $074306(+05)$ |
| 320 | $-0.403023186(+01)$ | $-0.716554161(+03)$ | $-0.817305389(+04)$ |
| 410 | $0.290457775(+01)$ | $0.227824886(+03)$ | $0.686126620(+05)$ |
| 222 | $0.100328131(+03)$ | $0.125357893(+04)$ | $0.613774043(+06)$ |
| 321 | $-0.784727771(+02)$ | $0.676226817(+02)$ | $0.234769592(+06)$ |
| 330 | $-0.355883520(+02)$ | $0.201090912(+04)$ | $0.125170847(+06)$ |
| 411 | 0.342774 866(+03) | $0.185046740(+04)$ | $0.252158250(+06)$ |
| 420 | $0.153901234(+01)$ | $0.904243255(+03)$ | -0.174 $950585(+06)$ |
| 510 | $-0.507373537(+02)$ | $-0.760782378(+03)$ | $-0.491418372(+06)$ |
| 322 | -0.781 $143074(+02)$ | -0.158 $269789(+04)$ | -0.112 $789831(+07)$ |
| 331 | $0.147626200(+03)$ | $0.529381639(+03)$ | -0.488 917 697(+06) |
| 421 | $0.366221664(+02)$ | $0.175945920(+03)$ | -0.398 080 764(+06) |
| 430 | $0.425337721(+02)$ | $-0.209129837(+04)$ | -0.153180 476(+06) |
| 511 | $-0.485699297(+03)$ | $-0.322417947(+04)$ | -0.828 $798279(+06)$ |
| 520 | $0.376517001(+02)$ | $-0.350892308(+03)$ | $0.117875724(+07)$ |
| 610 | $0.866519742(+02)$ | $0.114901910(+04)$ | $0.175662995(+07)$ |
| 332 | 0.302965 201(+03) | $0.125881368(+04)$ | 0.137477 152(+07) |
| 422 | $-0.374046089(+03)$ | $0.170029500(+03)$ | $0.604811338(+06)$ |
| 431 | $-0.115175393(+03)$ | $-0.478735978(+03)$ | $0.519369739(+06)$ |
| 440 | $0.130307975(+03)$ | $0.150614359(+04)$ | -0.244 769 673(+06) |
| 521 | $0.175553953(+03)$ | $0.967447169(+02)$ | $0.293758736(+06)$ |
| 530 | -0.138 $784667(+03)$ | $0.441133792(+03)$ | $0.153034935(+06)$ |
| 611 | $0.163275538(+03)$ | $0.209375703(+04)$ | $0.136008552(+07)$ |
| 620 | $-0.521422172(+01)$ | $0.266104075(+01)$ | -0.259 $383146(+07)$ |
| 710 | -0.457545705(+02) | $-0.696999913(+03)$ | -0.251569 755(+07) |
| $\beta^{(3) L}$ | $0.652302250(+00)$ | $0.868656020(+00)$ | $0.218844970(+01)$ |

${ }^{\mathrm{a}}$ All the coefficients are given in atomic units.
/ describe the overall size of the system, and $\theta$ and $\phi_{\tau}$ describe its shape. Pack and Parker ${ }^{36}$ have noted that it is often advantageous to view the surface of the internal sphere as functions of $\theta$ and $\phi_{\tau}$ with $\rho$ fixed. The stereographic projection has $X$ and $Y$ defined as

$$
X=\tan \left(\frac{\theta}{2}\right) \cos \phi_{\tau}, \quad Y=\tan \left(\frac{\theta}{2}\right) \sin \phi_{\tau} .
$$

The three internal coordinates, $\rho, \theta$, and $\phi_{\tau}$, are easily related to Jacobi coordinates $r_{\tau}, R_{\tau}$, and $\Theta_{\tau}$, with $\tau=A, B, C$ ( $A, B, C$ denoting the three particles of interest), through the expressions ${ }^{36}$

$$
\begin{align*}
& R_{\tau}=\frac{\rho}{d_{\tau} \sqrt{2}} \sqrt{1+\sin \theta \cos \left(2 \phi_{\tau}\right)}, \\
& r_{\tau}=\frac{\rho d_{\tau}}{\sqrt{2}} \sqrt{1-\sin \theta \cos \left(2 \phi_{\tau}\right)},  \tag{8}\\
& \cos \Theta_{\tau}=\frac{\rho}{2} \frac{\sin \left(2 \phi_{\tau}\right) \sin \theta}{r_{\tau} R_{\tau}}
\end{align*}
$$

with $d_{\tau}=\sqrt{m_{\tau}\left(M-m_{\tau}\right) /(\mu M)}$ being the mass-scaling factor, $\mu=\sqrt{m_{\tau} m_{\tau+1} m_{\tau+2} / M}$ being the reduced mass of the triatomic system, and where $M=m_{\tau}+m_{\tau+1}+m_{\tau+2}$ is the total mass of the system.

In all panels of Fig. 5, six arrangement channels appear instead of the expected three because of the inversion symmetry in the $\phi_{\tau}$ coordinate, causing each channel to be repeated twice. Some of the differences and similarities of the present GPES with respect to the Prosmiti et al. GPES ${ }^{23}$ are more evident from Fig. 5. In this figure we have selected three fixed $\rho$ values. One of them corresponds to the absolute minimum position ( $\rho=2.1715 a_{0}$ ), with zero energy represented by the central point in the upper panels of Fig. 5. As we can see from these plots, the minimum region is practically identical and all the dissociation channels are also identical as corresponds to three identical nuclei, for both GPES. However, as we can see in the middle and bottom panels of Fig. 5, when we enlarge the $\rho$ coordinate, a clear breakage of symmetry occurs in the Prosmiti et al. GPES ${ }^{23}$ (right panels), while the GPES reported herein has the correct symmetry properties (left panels).

## IV. ROVIBRATIONAL ANALYSIS AND INFRARED SPECTRA OF THE H ${ }_{3}^{+}$

The infrared spectra of $\mathrm{H}_{3}^{+}$(and its isotopic variants), first observed in 1980 by Oka, ${ }^{37}$ have been extensively studied, ${ }^{23,38-48}$ and it is nowadays rather well assigned (see, for example, Ref. 48). It should be noted that there are sev-


FIG. 4. Contours of the ground-state $\mathrm{H}_{3}^{+}$GPES in Jacobi coordinates $\mathbf{r}, \mathbf{R}$, and $\Theta$. For each contour map the $\Theta$ angle is fixed $\left(0^{\circ}\right.$, upper panels; $45^{\circ}$, intermediate panels, and $90^{\circ}$, bottom panels). The left-hand panels correspond to the present results; the right-hand panels correspond to the Prosmiti et al. GPES (Ref. 23). The solid curves are contours of the interaction potential. The contours range from 5000 to $75000 \mathrm{~cm}^{-1}$ for the bottom panels, from 10000 to $75000 \mathrm{~cm}^{-1}$ for the intermediate panels, and from 15000 to $75000 \mathrm{~cm}^{-1}$ for the top panels, always in steps of $5000 \mathrm{~cm}^{-1}$. Distances are given in atomic units.
eral LPES's, ${ }^{47,48}$ fitted in the region of the well, which yield results within the experimental resolution, of the order of $0.001 \mathrm{~cm}^{-1}$. However, the potential reported here is designed to describe the entire configuration space, including asymptotic regions as well as the existing conical intersections; this GPES does not have the same spectroscopic accuracy, but a reasonable global accuracy for dynamical calculations. The study of the infrared spectra, which we shall present below, only pretends to check the quality of the GPES presented in this work, in the region of the well, maintaining that we assume a similar accuracy for all the dissociation channels. Moreover, we are also interested in calculating higher rovibrational levels than those previously reported.

The rovibrational energy levels and transitions of $\mathrm{H}_{3}^{+}$ have mainly been studied using Jacobi coordinates ${ }^{23,47-49}$ and hyperspherical coordinates. ${ }^{50-55}$ In this work we shall use the hyperspherical coordinates of Pack and Parker, ${ }^{56}$ (denoted by APHJ), which are closely related to those described by Smith ${ }^{57,58}$ and Johnson. ${ }^{59,60}$ The body-fixed frame chosen corresponds to the principal axis system with the $z$-axis being perpendicular to the plane of the molecule, and it is related to the space-fixed frame through the $\alpha, \beta, \gamma$ Eulerian angles.


FIG. 5. Stereographic projection of contour plots of the ground-state $\mathrm{H}_{3}^{+}$ GPES in hyperspherical coordinates $\rho, \theta$, and $\phi_{\tau}$ (see the text for more details on these coordinates and for the definition of X and Y ). For each contour map the $\rho$ distance is fixed ( $2.1715 a_{0}$, upper panels; $6 a_{0}$, intermediate panels; and $9 a_{0}$, bottom panels). The left-hand panels correspond to the present results; the right-hand panels correspond to the Prosmiti et al. GPES (Ref. 23). The solid curves are contours of the interaction potential. The contours go from 64000 to $72000 \mathrm{~cm}^{-1}$ for the bottom panels and from 49000 to $58000 \mathrm{~cm}^{-1}$ for the intermediate panels; in both cases the minimum contours is nearest to the equatorial region and the increment from each contour to the closest one is $1000 \mathrm{~cm}^{-1}$. Finally, for the top panels the contours range from $0 \mathrm{~cm}^{-1}$ (the central point in both panels) to 26000 $\mathrm{cm}^{-1}$ in steps of $2000 \mathrm{~cm}^{-1}$.

The three internal coordinates, $\rho, \theta$, and $\phi_{\tau}$, are easily related to the Jacobi coordinates $r_{\tau}, R_{\tau}$, and $\Theta_{\tau}{ }^{36}$ [see Eq. 8)].

In these coordinates, the Hamiltonian takes the form ${ }^{56}$

$$
\begin{equation*}
H=T_{\rho}+T_{h}+T_{R}+T_{C}+V\left(\rho, \theta, \phi_{\tau}\right) \tag{9}
\end{equation*}
$$

where

$$
\begin{align*}
& T_{\rho}=-\frac{\hbar^{2}}{2 \mu \rho^{5}} \frac{\partial}{\partial \rho} \rho^{5} \frac{\partial}{\partial \rho},  \tag{10}\\
& T_{h}=-\frac{\hbar^{2}}{2 \mu \rho^{2}}\left[\frac{4}{\sin (2 \theta)} \frac{\partial}{\partial \theta} \sin (2 \theta) \frac{\partial}{\partial \theta}+\frac{1}{\sin ^{2} \theta} \frac{\partial^{2}}{\partial \phi_{\tau}^{2}}\right] \\
& T_{R}=\frac{\hat{J}_{x}^{2}}{\mu \rho^{2}(1-\sin \theta)}+\frac{\hat{J}_{y}^{2}}{\mu \rho^{2}(1+\sin \theta)}+\frac{\hat{J}_{z}^{2}}{2 \mu \rho^{2} \sin ^{2} \theta},
\end{align*}
$$

$$
T_{C}=-\frac{i \hbar}{\mu \rho^{2}} \frac{\cos \theta}{\sin ^{2} \theta} \hat{J}_{z} \frac{\partial}{\partial \phi_{\tau}}
$$

where $\hat{J}_{x}, \hat{J}_{y}, \hat{J}_{z}$ are the components of the total angular momentum operator in the body-fixed frame.

The eigenstates of this Hamiltonian are expanded as

$$
\begin{align*}
\Phi_{i}^{J M}= & 4 \rho^{-5 / 2} \sum_{v, k, n, \Omega} C_{v, k, n, \Omega}^{J M i} W_{\Omega, n}^{J M}\left(\alpha, \beta, \gamma, \phi_{\tau}\right) \\
& \times F_{k}^{J, \Omega, n}(\theta) \varphi_{v}(\rho), \tag{11}
\end{align*}
$$

where the angular functions $W_{\Omega, n}^{J M}$ are of the form

$$
\begin{equation*}
W_{\Omega, n}^{J M}\left(\alpha, \beta, \gamma, \phi_{\tau}\right)=\sqrt{\frac{2 J+1}{8 \pi^{2}}} D_{M \Omega}^{J *}(\alpha, \beta, \gamma) \frac{e^{i n \phi_{\tau}}}{\sqrt{2 \pi}} \tag{12}
\end{equation*}
$$

the $D_{M \Omega}^{J *}$ being Wigner rotation matrices. ${ }^{61}$ In Eq. (11), the $F_{k}^{J, \Omega, n}(\theta)$ are related to the $G_{k}(p, q, x)$ Jacobi polynomials ${ }^{62}$ as

$$
\begin{equation*}
F_{k}^{J, \Omega, n}(\theta)=\frac{(\sin \theta)^{a}(\cos \theta)^{b}}{\sqrt{2}} \sqrt{\frac{(2 k+a+b+1)[\Gamma(2 k+a+b+1)]^{2}}{k!\Gamma(k+a+b+1) \Gamma(k+a+1) \Gamma(k+b+1)}} G_{k}\left(b+1, a+b+1, \cos ^{2} \theta\right), \tag{13}
\end{equation*}
$$

with $a=|n+\Omega| / 2$ and $b=\sqrt{\left[J(J+1)-\Omega^{2}\right] / 2}$. This functions are orthonormalized in the interval $0 \leqslant \theta \leqslant \pi / 2, F_{k}^{J, \Omega, n}$ $=F_{k}^{J,-\Omega,-n}$ and satisfy

$$
\begin{equation*}
\left\{-\frac{4}{\sin (2 \theta)} \frac{d}{d \theta} \sin (2 \theta) \frac{d}{d \theta}+\frac{n^{2}+2 n \Omega+\Omega^{2}}{\sin ^{2} \theta}+\frac{2 J(J+1)-2 \Omega^{2}}{\cos ^{2} \theta}-\mathcal{K}(\mathcal{K}+4)\right\} F_{k}^{J, \Omega, n}(\theta) \tag{14}
\end{equation*}
$$

with $\mathcal{K}=4 k+2 a+2 b$.
Finally, the $\varphi_{v}(\rho)$ functions appearing in Eq. (11) are the solutions of the one-dimensional differential equation

$$
\begin{equation*}
\left\{-\frac{\hbar^{2}}{2 \mu} \frac{d^{2}}{d \rho^{2}}+V_{\text {ref }}(\rho)-E_{v}\right\} \varphi_{v}(\rho)=0 \tag{15}
\end{equation*}
$$

where $V_{\text {ref }} \equiv V\left(\rho, \theta=0, \phi_{\tau}=0\right)$ in this case. Equation (15) is solved numerically in a large grid of equispaced points in $\rho$.
The Hamiltonian matrix elements in this basis set representation are of the form

$$
\begin{align*}
\left\langle\Phi_{v} F_{k}^{J \Omega n} W_{\Omega n}^{J M}\right| H\left|\Phi_{v^{\prime}} F_{k^{\prime}}^{J \Omega^{\prime} n^{\prime}} W_{\Omega^{\prime} n^{\prime}}^{J M}\right\rangle= & \delta_{\Omega \Omega^{\prime}} \delta_{n n^{\prime}} \delta_{k k^{\prime}}\left\{\left\langle\varphi_{v}\right|-\frac{\hbar^{2}}{2 \mu} \frac{d^{2}}{d \rho^{2}}\left|\varphi_{v^{\prime}}\right\rangle+\left\langle\varphi_{v}\right| \frac{\hbar^{2}}{2 \mu \rho^{2}}\left|\varphi_{v^{\prime}}\right\rangle\left[\frac{15}{4}+\mathcal{K}(\mathcal{K}+4)\right]\right\} \\
& +\delta_{\Omega \Omega^{\prime},} \delta_{n n^{\prime}}\left\langle\varphi_{v}\right| \frac{\hbar^{2}}{2 \mu \rho^{2}}\left|\varphi_{v^{\prime}}\right\rangle\left\langle F_{k}^{J, \Omega, n}\right| \frac{\cos \theta-1}{\sin ^{2} \theta}\left|F_{k^{\prime}}^{J, \Omega, n}\right\rangle 2 n \Omega \\
& +\delta_{\Omega \Omega^{\prime} \pm 2} \delta_{n n^{\prime}}\left\langle\varphi_{v}\right| \frac{\hbar^{2}}{2 \mu \rho^{2}}\left|\varphi_{v^{\prime}}\right\rangle\left\langle F_{k}^{J, \Omega, n}\right| \frac{\sin \theta}{\cos ^{2} \theta}\left|F_{k^{\prime}}^{J, \Omega, n}\right\rangle \\
& \times \sqrt{J(J+1)-\Omega^{\prime}\left(\Omega^{\prime} \pm 1\right)} \sqrt{J(J+1)-\left(\Omega^{\prime} \pm 1\right) \Omega} \\
& +\delta_{\Omega \Omega^{\prime}} \delta_{n^{\prime}-n, \lambda}\left\langle\varphi_{v} F_{k}^{J, \Omega, n}\right| V_{\lambda}(\rho, \theta)\left|\varphi_{v^{\prime}} F_{k^{\prime}}^{J, \Omega, n^{\prime}}\right\rangle \tag{16}
\end{align*}
$$

where the $V_{\lambda}$ are the coefficients of the expansion of the potential as
$V\left(\rho, \theta, \phi_{\tau}\right)=\sum_{\lambda} V_{\lambda}(\rho, \theta) e^{i \lambda \phi_{\tau}} \quad$ with $\quad \lambda=0, \pm 6, \pm 12, \ldots$.

The integrals involved in the Hamiltonian matrix are evaluated numerically. A trapezoidal integration is used for $\rho$
while the integral on $\theta$ is carried out by means of a GaussLegendre quadrature with a large number of points ( $\approx 300$ ).

The permutational group of three identical particles is isomorphic with the $D_{3}$ group and when adding the symmetry under the inversion of spatial coordinates (isomorphic with the $C_{i}$ group) the group of all symmetry operations becomes isomorphic with the $D_{3 h}$ group. Hyperspherical coordinates are particularly well suited for treating the permu-

TABLE VI. Rovibrational eigenvalues of $\mathrm{H}_{3}^{+}$for $J=0$ for the present GPES (with zero-point energy, ZPE $=4362.0841 \mathrm{~cm}^{-1}$ ), for the GPES of Ref. 23 (with ZPE $=4362.7801 \mathrm{~cm}^{-1}$ ) and for the LPES of Ref. 7 (with $\mathrm{ZPE}=4361.44 \mathrm{~cm}^{-1}$ ). The asterisk mark is used to indicate that we obtain a different assignment than in previous works.

| $i(\Gamma, J)$ | $\left(v_{1}, v_{2}^{l}\right)$ | $\Gamma$ | Present GPES | GPES of Ref. 23 | LPES of Ref. 7 | Exp+fit (Ref. 48) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | $\left(0,0^{0}\right)$ | $A_{1}^{\prime}$ | 0.0000 | 0.0000 | 0.00 | 0.00 |
| 2 | $\left(1,0^{0}\right)$ ) | $A_{1}^{\prime}$ | 3179.0807 | 3178.7480 | 3178.15 |  |
| 3 | $\left(0,2^{0}\right)$ | $A_{1}^{\prime}$ | 4777.6088 | 4777.1021 | 4778.01 |  |
| 4 | ( $2,0^{0}$ ) | $A_{1}^{\prime}$ | 6263.5741 | 6260.4428 | 6261.81 |  |
| 5* | $\left(0,3^{3}\right)$ | $A_{1}^{\prime}$ | 7284.4716 | 7283.1507 | 7285.32 |  |
| 6 | $\left(1,2^{0}\right)$ | $A_{1}^{\prime}$ | 7769.2711 | 7766.7933 | 7768.84 |  |
| 7 | $\left(0,4^{0}\right)$ | $A_{1}^{\prime}$ | 8999.8244 | 8995.3972 | 9001.36 |  |
| 8 | (3, $0^{0}$ ) | $A_{1}^{\prime}$ | 9253.9503 | 9245.7749 | 9251.42 |  |
| 9 | $\left(1,3^{3}\right)$ | $A_{1}^{\prime}$ | 9967.8366 | 9963.1895 | 9968.34 |  |
| 10 | $\left(2,2^{0}\right)$ | $A_{1}^{\prime}$ | 10593.5900 | 10586.2089 | 10592.50 |  |
| 11 | $\left(0,5^{3}\right)$ | $A_{1}^{\prime}$ | 10920.0617 | 10910.9288 | 10923.49 |  |
| 12 | $\left(1,4^{0}\right)$ | $A_{1}^{\prime}$ | 11813.2272 | 11801.1910 | 11814.16 |  |
| 13 | $\left(4,0^{0}\right)$ | $A_{1}^{\prime}$ | 12149.0286 | 12133.5773 |  |  |
| 1 | $\left(0,3^{3}\right)$ | $A_{2}^{\prime}$ | 7492.4355 | 7491.4193 | 7492.61 | 7493.113 |
| 2 | $\left(1,3^{3}\right)$ | $A_{2}^{\prime}$ | 10209.5390 | 10206.0906 | 10209.65 |  |
| 3 | $\left(0,5^{3}\right)$ | $A_{2}^{\prime}$ | 11527.6239 | 11522.6036 | 11529.15 |  |
| 4 | $\left(2,3^{3}\right)$ | $A_{2}^{\prime}$ | 12831.5467 | 12820.0798 |  |  |
| 5 | $\left(0,6^{3}\right)$ | $A_{2}^{\prime}$ | 13752.4209 | 13739.8937 |  |  |
| 1 | $\left(0,1^{1}\right)$ | $E^{\prime}$ | 2521.0814 | 2520.9840 | 2521.20 | 2521.422 |
| 3 | (0, $2^{2}$ ) | $E^{\prime}$ | 4997.5292 | 4996.9634 | 4997.73 | 4998.058 |
| 5 | $\left(1,1^{1}\right)$ | $E^{\prime}$ | 5554.5911 | 5554.2073 | 5553.95 | 5554.274 |
| 7 | (0, ${ }^{1}$ ) | $E^{\prime}$ | 7004.9686 | 7003.2497 | 7005.81 | 7006.187 |
| 9 | $\left(1,2^{2}\right)$ | $E^{\prime}$ | 7870.0238 | 7868.5633 | 7869.09 | 7870.664 |
| 11 | $\left(2,1^{1}\right)$ | $E^{\prime}$ | 8488.8693 | 8485.4081 | 8487.53 |  |
| 13 | (0, $4^{2}$ ) | $E^{\prime}$ | 9111.1184 | 9107.2931 | 9112.90 |  |
| 15 | $\left(1,3^{1}\right)$ | $E^{\prime}$ | 9652.5387 | 9649.7530 | 9653.44 |  |
| 17 | ( $0,4^{4}$ ) | $E^{\prime}$ | 9996.9482 | 9990.8078 | 9996.67 |  |
| 19 | $\left(2,2^{2}\right)$ | $E^{\prime}$ | 10645.3620 | 10639.1250 | 10644.59 |  |
| 21* | $\left(1,4^{2}\right)$ | $E^{\prime}$ | 10859.6242 |  | 10862.46 |  |
| 23 | $\left(3,1^{1}\right)$ | $E^{\prime}$ | 11324.3941 |  | 11322.31 |  |
| 25 | $\left(1,4^{2}\right)$ | $E^{\prime}$ | 11656.2449 |  | 11657.69 |  |
| 27* | $\left(0,5^{1}\right)$ | $E^{\prime}$ | 12077.4622 |  | 12078.43 |  |

tational symmetry of triatomic systems with three identical nuclei, which may yield a significant reduction in the size of the Hamiltonian matrices for a particular irreducible representation. Knowing the effect of the different symmetry operators on the hyperspherical coordinates, ${ }^{51,63-65}$ the symmetry adapted basis functions for the $\Gamma$ irreducible representation of the $D_{3 h}$ group are written as linear combination of the functions of Eq. (12) as

$$
\begin{equation*}
\mathcal{W}_{\Omega n}^{J M \Gamma}\left(\alpha, \beta, \gamma, \phi_{\tau}\right)=A_{\Omega n}^{J \Gamma} W_{\Omega n}^{J M}+B_{\Omega n}^{J \Gamma} W_{-\Omega-n}^{J M} \tag{18}
\end{equation*}
$$

where the $A_{\Omega n}^{J \Gamma}$ and $B_{\Omega n}^{J \Gamma}$ coefficients are obtained using the corresponding projection operator, taking the values

$$
\begin{align*}
A_{\Omega n}^{J \Gamma} \propto\{ & \chi^{\Gamma}(E)+\chi^{\Gamma}\left(E^{*}\right)(-1)^{\Omega}+\left[\chi^{\Gamma}\left(C_{3}\right)\right. \\
& \left.\left.+\chi^{\Gamma}\left(S_{3}\right)(-1)^{\Omega}\right] 2 \cos (4 n \pi / 3)\right\} \\
B_{\Omega n}^{J \Gamma} \propto & {[1+2 \cos (4 n \pi / 3)]\left\{\chi^{\Gamma}\left(C_{2}\right)(-1)^{J+\Omega+n}\right.}  \tag{19}\\
& \left.+\chi^{\Gamma}\left(\sigma_{v}\right)(-1)^{J+n}\right\}
\end{align*}
$$

for $\Omega \neq 0$ and/or $n \neq 0$, while for $n=\Omega=0$ these coefficients are given by

$$
\begin{align*}
A_{\Omega n}^{J \Gamma} \propto & \chi^{\Gamma}(E)+\chi^{\Gamma}\left(E^{*}\right)+\left[\chi^{\Gamma}\left(C_{2}\right)+\chi^{\Gamma}\left(\sigma_{v}\right)\right](-1)^{J} \\
& \times[1+2 \cos (4 n \pi / 3)] \\
& +\left[\chi^{\Gamma}\left(C_{3}\right)+\chi^{\Gamma}\left(S_{3}\right)\right] 2 \cos (4 n \pi / 3), \\
B_{\Omega n}^{J \Gamma}= & 0, \tag{20}
\end{align*}
$$

where $\chi^{\Gamma}(\mathcal{C})$ is the character of the symmetry class $\mathcal{C}$ for the $\Gamma$ irreducible representation.

The only good quantum numbers associated with each eigenstate are the total angular momentum, $J$, and the symmetry, characterized by the $\Gamma$ irreducible representation of the $D_{3 h}$ group. Traditionally, for the classification of the eigenstates several approximated quantum numbers are used, ${ }^{66}\left(v_{1}, v_{2}^{l}\right)$ and $G$ and $U . v_{1}$ and $v_{2}$ correspond to the symmetric and antisymmetric vibrations, respectively, while $l$ labels the bending. Since in the present treatment the $\varphi_{v}(\rho)$ functions correspond to a prediagonalization of the symmetric stretch at the equilibrium values of $\theta$ and $\phi_{\tau}, v_{1}$ corresponds to the dominant $v$ in the expansion of Eq. (11). The asymmetric stretch, associated with the $\theta$ coordinate, is analyzed by comparing the full eigenstates with the monodimensional solutions of the problem for $\rho^{\mathrm{eq}}, \phi_{\tau}^{\mathrm{eq}}$ and $J=0$. The $l$

TABLE VII. Root-mean-square (rms) error of the energy levels calculated with the GPES reported in this work with respect to most accurate energy levels of Ref. 48, listing the number of levels used in the statistics as well as the maximum total angular momentum, $J_{\max }$, considered in the comparison.

| $\left(v_{1}, v_{2}^{l}\right)$ | Number of levels | $J_{\max }$ | rms error |
| :---: | :---: | :---: | :---: |
| $\left(0,0^{0}\right)$ | 60 | 10 | 0.243 |
| $\left(0,1^{1}\right)$ | 122 | 10 | 0.261 |
| $\left(1,0^{0}\right)$ | 60 | 10 | 0.946 |
| $\left(0,2^{0}\right)$ | 54 | 10 | 0.832 |
| $\left(0,2^{2}\right)$ | 105 | 10 | 0.474 |
| $\left(1,1^{1}\right)$ | 106 | 10 | 0.606 |
| $\left(2,0^{0}\right)$ | 45 | 10 | 1.221 |
| $\left(0,3^{1}\right)$ | 71 | 10 | 1.227 |
| $\left(0,3^{3}\right)$ | 63 | 10 | 1.029 |
| $\left(1,2^{0}\right)$ | 27 | 10 | 0.539 |
| $\left(1,2^{2}\right)$ | 47 | 10 | 1.414 |
| $\left(2,1^{1}\right)$ | 18 | 10 | 0.383 |

is associated with the $\phi_{\tau}$ variable and is obtained from the dominant $n$ and $\Omega$ values, $n_{\max }$ and $\Omega_{\max }$, respectively, in the expansion of Eq. (11) as $l=\left|\Omega_{\max }+n_{\max }\right| / 2$. Similarly, the $G$ number is also obtained from $n_{\max }$ and $\Omega_{\max }$ as $G$ $=\left|n_{\max }+3 \Omega_{\max }\right| / 2$. Finally, $U$ is equal, in absolute value, to $l$, and its sign allows us to distinguish between the $A_{1}$ and $A_{2}$ components. Since we are using symmetry adapted functions, this latter number is not required to specify the character of the states, and in what follows we shall label the states by $\left(v_{1}, v_{2}^{l}\right), J, G, \Gamma$. It should be noted that the assignment is not a simple task, especially at high energies, and we have also used the data from previous assignments. ${ }^{7,48}$

In Table VI the eigenvalues for $J=0$ are compared with those of previous works. The LPESs of Refs. 48 and 7 describe the region of the well, yielding results in very good agreement with the experimental transitions, within an error typically less than $0.1 \mathrm{~cm}^{-1}$. However, these LPESs do not correctly describe the asymptotic regions and, therefore, they are not well suited for dynamical studies. The present GPES and that of Ref. 23 describe the entire configuration space and none of them includes adiabatic terms. The spectroscopic accuracy of these two GPES is lower than that of the previously LPES mentioned. However, the present GPES seems to yield results in better agreement with respect to the spectroscopically accurate LPESs of Refs. 48 and 7 than that of Ref. 23. In particular, the maximum difference between the rovibrational levels obtained with this GPES and that of Ref. 7 is of the order of $1-2 \mathrm{~cm}^{-1}$ for energies of the order of $16000 \mathrm{~cm}^{-1}$ above the minimum. This difference is considered to be reasonably good for spectroscopy, especially because the adiabatic corrections have not been introduced.

In order to determine more precisely the spectroscopic accuracy of the present GPES, the eigenstates up to $\approx 14000$ $\mathrm{cm}^{-1}$ have been calculated for total angular momentum in the range $0 \leqslant J \leqslant 20$. The Hamiltonian matrix size increases very rapidly with $J$ and it is therefore impossible to apply a variational method for such large $J$ s. Instead, we use an iterative procedure based on the Lanczos algorithm ${ }^{67}$ in two steps. The eigenvalues are obtained with a nonorthogonal Lanczos procedure following the method of Cullum and Willoughby. ${ }^{68}$ The eigenstates are then obtained iteratively
using the conjugate gradient method ${ }^{69,70}$ in a very efficient way. The nuclear spin of the nuclei is $1 / 2$ and the total wave functions, including nuclear spin, must be antisymmetric under exchange of any pair of nuclei, according to the FermiDirac statistic. Following Watson, ${ }^{66}$ this implies that the wave function (without the nuclear spin part) must be $A_{2}^{\prime}$ or $A_{2}^{\prime \prime}$ for total nuclear spin $I=3 / 2\left(\right.$ ortho $\left.\mathrm{H}_{3}^{+}\right)$and $E^{\prime}$ or $E^{\prime \prime}$ for $I=1 / 2\left(\right.$ para $\left.\mathrm{H}_{3}^{+}\right)$. Up to $J=10$, the allowed energy levels calculated following this method are in very good agreement with those reported by Dinelli et al., ${ }^{48}$ which are of the highest spectroscopic accuracy available for this system; the typical rms deviations for 778 levels with $J \leqslant 10$ are listed in Table VII. As for $J=0$, the typical error is usually lower than $1 \mathrm{~cm}^{-1}$ and it should be noted that the larger errors occur for $J>10$ values, as can be seen in Table VIII, where the energy levels for some selected $J$ s are shown together with those of Dinelli et al. ${ }^{48}$ As noted by Dinelli et al., ${ }^{48}$ their calculations for high $J$ s are not well converged, which can be the reason of the discrepancy. However, the method described above allows us the calculation of energy levels for high $J$ values in a very efficient way, and in Table VIII the eigenvalues for $J=15$ and $J=20$ are also shown. Moreover, the use of symmetry adapted functions and hyperspherical coordinates yields a simple and nearly automatic procedure to assign the energy levels, almost always in good agreement with the previous assignment. ${ }^{48}$

Finally, in Table IX, the deviation of the present results with respect to some experimental transitions ${ }^{41,43,44,46,71,72,73}$ is shown and compared to that obtained using the levels of Dinelli et al. ${ }^{48}$ The typical error of the present GPES is of the order of tenths of $\mathrm{cm}^{-1}$. These are larger errors than those obtained by Dinelli et al. ${ }^{48}$ Nevertheless, it should be noted that in the present potential, of global nature, adiabatic corrections have not been included. As has been recently discussed,,${ }^{8,74}$ in order to reproduce the experimental data for $\mathrm{H}_{3}^{+}$to within a few hundredths of $\mathrm{cm}^{-1}$ adiabatic corrections to the Born-Oppenheimer approximation should be included, for example using different masses for vibrational and rotational motions.

## V. CONCLUSIONS

In this paper we have reported a new global potential energy surface for the ground-state of the $\mathrm{H}_{3}^{+}$system, based on a huge number of full configuration interaction energies, covering all the regions of the potential surface. The rms error of this GPES has been estimated to be lower than 20 $\mathrm{cm}^{-1}$. The global fit is totally symmetric with respect to permutations of the hydrogen atoms and cover all regions of the GPES. To test the quality of this GPES at the absolute minimum region, we have also reported calculations of rovibrational levels up to higher $J$ values than those reported previously and we have reproduced the infrared spectra within $1 \mathrm{~cm}^{-1}$ with respect to the experimental ones. We therefore conclude that the accuracy of the present GPES is very high, especially taking into account its "global character,", and very well suited for dynamical calculations, due to the accuracy and symmetrical behavior of the dissociation channels.

TABLE VIII. Rovibrational eigenvalues ( $E$ given in $\mathrm{cm}^{-1}$ ) of $\mathrm{H}_{3}^{+}$for $J=5,10,15,20$, for the present GPES compared with the very accurate values of Dinelli et al. (Ref. 48).

| $\left(v_{1}, v_{2}^{l}\right)$ | $J$ | $G$ | $\Gamma$ | $E$ | $E$ (Ref. 48) | $\left(v_{1}, v_{2}^{l}\right)$ | $J$ | $G$ | $\Gamma$ | $E$ | $E$ (Ref. 48) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| (0, $0^{0}$ ) | 5 | 0 | $A_{2}^{\prime}$ | 1271.407 | 1271.269 | (0, $2^{2}$ ) | 5 | 7 | $E^{\prime \prime}$ | 5363.415 | 5363.836 |
| $\left(0,1^{1}\right)$ | 5 | 3 | $A_{2}^{\prime}$ | 3553.086 | 3553.340 | $\left(0,2^{0}\right.$ ) | 5 | 5 | $E^{\prime \prime}$ | 5459.792 | 5460.465 |
| (0, 1 ${ }^{1}$ ) | 5 | 3 | $A_{2}^{\prime}$ | 3673.833 | 3673.964 | (0, $2^{2}$ ) | 5 | 5 | $E^{\prime \prime}$ | 5899.067 | 5899.400 |
| $\left(1,0^{0}\right)$ | 5 | 0 | $A_{2}^{\prime}$ | 4420.118 | 4419.147 | (0, $2^{0}$ ) | 5 | 1 | $E^{\prime \prime}$ | 6002.574 | 6003.163 |
| (0, $2^{2}$ ) | 5 | 6 | $A_{2}^{\prime}$ | 5658.859 | 5659.225 | (0, $2^{2}$ ) | 5 | 1 | $E^{\prime \prime}$ | 6276.403 | 6276.731 |
| $\left(0,2^{0}\right)$ | 5 | 0 | $A_{2}^{\prime}$ | 6022.482 | 6023.076 | (0, $2^{2}$ ) | 5 | 1 | $E^{\prime \prime}$ | 6376.335 | 6376.537 |
| $\left(0,2^{2}\right)$ | 5 | 0 | $A_{2}^{\prime}$ | 6391.661 | 6391.893 | $\left(1,1^{1}\right)$ | 5 | 4 | $E^{\prime \prime}$ | 6410.824 | 6410.568 |
| $\left(1,1^{1}\right)$ | 5 | 3 | $A_{2}^{\prime}$ | 6568.640 | 6568.277 | $\left(1,1^{1}\right)$ | 5 | 4 | $E^{\prime \prime}$ | 6529.765 | 6529.265 |
| $\left(1,1^{1}\right)$ | 5 | 3 | $A_{2}^{\prime}$ | 6679.769 | 6679.233 | $\left(1,1^{1}\right)$ | 5 | 2 | $E^{\prime \prime}$ | 6673.165 | 6672.774 |
| (2, $0^{0}$ ) | 5 | 0 | $A_{2}^{\prime}$ | 7476.045 | 7474.726 | $\left(1,1^{1}\right)$ | 5 | 2 | $E^{\prime \prime}$ | 6793.077 | 6792.501 |
| $\left(0,3^{1}\right)$ | 5 | 3 | $A_{2}^{\prime}$ | 7962.296 | 7963.536 | (2, $0^{0}$ ) | 5 | 5 | $E^{\prime \prime}$ | 6955.032 | 6953.812 |
| $\left(0,3^{1}\right)$ | 5 | 3 | $A_{2}^{\prime}$ | 8228.598 | 8229.624 | $\left(2,0^{0}\right)$ | 5 | 1 | $E^{\prime \prime}$ | 7456.048 | 7454.731 |
| $\left(1,2^{2}\right)$ | 5 | 6 | $A_{2}^{\prime}$ | 8539.626 | 8539.900 | $\left(0,3^{3}\right)$ | 5 | 8 | $E^{\prime \prime}$ | 7637.726 | 7638.501 |
| $\left(0,3^{3}\right)$ | 5 | 3 | $A_{2}^{\prime}$ | 8682.235 | 8683.163 | $\left(0,3^{1}\right)$ | 5 | 4 | $E^{\prime \prime}$ | 7766.665 | 7767.965 |
| (0, ${ }^{3}$ ) | 5 | 3 | ${ }_{2}$ |  |  | $\left(0,3^{1}\right)$ | 5 | 4 | $E^{\prime \prime}$ | $8052.009$ | $8053.102$ |
| $\left(0,0^{0}\right)$ | 5 | 4 | $E^{\prime}$ | 929.075 | 928.975 | (0, $3^{1}$ ) | 5 | 2 | $E^{\prime \prime}$ | 8089.003 | 8090.127 |
| (0, $0^{0}$ ) | 5 | 2 | $E^{\prime}$ | 1187.237 | 1187.108 | $\left(1,2^{2}\right)$ | 5 | 7 | $E^{\prime \prime}$ | 8256.267 | 8256.630 |
| (0, 1 ${ }^{1}$ ) | 5 | 5 | $E^{\prime}$ | 3299.915 | 3300.131 | $\left(0,3^{1}\right)$ | 5 | 2 | $E^{\prime \prime}$ | 8267.694 | 8269.095 |
| (0, 1 ${ }^{1}$ ) | 5 | 1 | $E^{\prime}$ | 3722.399 | 3722.638 | $\left(0,3^{3}\right)$ | 5 | 2 | $E^{\prime \prime}$ | 8396.024 | 8396.945 |
| (0, 1 ${ }^{1}$ ) | 5 | 1 | $E^{\prime}$ | 3863.331 | 3863.413 | $\left(0,3^{3}\right)$ | 5 | 4 | $E^{\prime \prime}$ | 8619.566 | 8620.290 |
| $\left(1,0^{0}\right)$ | 5 | 4 | $E^{\prime}$ | 4085.649 | 4084.721 | $\left(1,2^{0}\right)$ | 5 | 5 | $E^{\prime \prime}$ | 8634.387 | 8634.809 |
| $\left(1,0^{0}\right)$ | 5 | 2 | $E^{\prime}$ | 4337.991 | 4337.030 | $\left(1,2^{2}\right)$ | 5 | 5 | $E^{\prime \prime}$ | 8774.061 | 8774.309 |
| $\left(0,2^{0}\right)$ | 5 | 4 | $E^{\prime}$ | 5690.234 | 5690.840 | $\left(0,3^{3}\right)$ | 5 | 2 | $E^{\prime \prime}$ | 8792.310 | 8793.047 |
| $\left(0,2^{0}\right)$ | 5 | 2 | $E^{\prime}$ | 5939.093 | 5939.691 | (0, $0^{0}$ ) | 10 | 6 | $A_{2}^{\prime}$ | 3726.944 | 3726.549 |
| (0, $2^{2}$ ) | 5 | 4 | $E^{\prime}$ | 6089.518 | 6089.819 | $\left(0,1^{1}\right)$ | 10 | 9 | $A_{2}^{\prime}$ | 5198.084 | 5198.463 |
| (0, $2^{2}$ ) | 5 | 2 | $E^{\prime}$ | 6169.134 | 6169.466 | (0, $1^{1}$ ) | 10 | 9 | $A_{2}^{\prime}$ | 5454.525 | 5454.531 |
| (0, $2^{2}$ ) | 5 | 2 | $E^{\prime}$ | 6327.764 | 6327.961 | $\left(0,1^{1}\right)$ | 10 | 3 | $A_{2}^{\prime}$ | 6539.929 | 6539.917 |
| $\left(1,1^{1}\right)$ | 5 | 5 | $E^{\prime}$ | 6346.639 | 6346.289 | (0, $2^{2}$ ) | 10 | 12 | $A_{2}^{\prime}$ | 6668.748 | 6671.817 |
| $\left(1,1^{1}\right)$ | 5 | 1 | $E^{\prime}$ | 6733.647 | 6733.247 | $\left(1,0^{0}\right)$ | 10 | 6 | $A_{2}^{\prime}$ | 6805.429 | 6804.443 |
| $\left(1,1^{1}\right)$ | 5 | 1 | $E^{\prime}$ | 6860.444 | 6859.851 | $\left(0,1^{1}\right)$ | 10 | 3 | $A_{2}^{\prime}$ | 6959.550 | 6959.027 |
| (2, $0^{0}$ ) | 5 | 4 | $E^{\prime}$ | 7148.211 | 7146.931 | $\left(1,1^{1}\right)$ | 10 | 9 | $A_{2}^{\prime}$ | 8108.049 | 8109.734 |
| (2, $0^{0}$ ) | 5 | 2 | $E^{\prime}$ | 7395.690 | 7394.378 | $\left(1,0^{0}\right)$ | 10 | 6 | $A_{2}^{\prime}$ | 8259.398 | 8259.969 |
| (0, $3^{1}$ ) | 5 | 5 | $E^{\prime}$ | 7829.274 | 7830.292 | $\left(1,1^{1}\right)$ | 10 | 9 | $A_{2}^{\prime}$ | 8445.726 | 8445.703 |
| (0, $3^{3}$ ) | 5 | 7 | $E^{\prime}$ | 7962.996 | 7963.769 | (0, ${ }^{2}$ ) | 10 | 6 | $A_{2}^{\prime}$ | 8690.943 | ... |
| (0, $3^{1}$ ) | 5 | 1 | $E^{\prime}$ | 8151.162 | 8152.309 | (0, ${ }^{2}$ ) | 10 | 6 | $A_{2}^{\prime}$ | 9018.072 | $\ldots$ |
| $\left(0,3^{1}\right)$ | 5 | 1 | $E^{\prime}$ | 8403.942 | 8404.998 | $\left(0,2^{2}\right)$ | 10 | 0 | $A_{2}^{\prime}$ | 9317.968 | $\ldots$ |
| $\left(0,3^{3}\right)$ | 5 | 5 | $E^{\prime}$ | 8484.773 | 8485.553 | $\left(1,2^{2}\right)$ | 10 | 12 | $A_{2}^{\prime}$ | 9323.874 | $\ldots$ |
| $\left(0,3^{3}\right)$ | 5 | 1 | $E^{\prime}$ | 8556.215 | 8557.124 | $\left(0,2^{2}\right)$ | 10 | 0 | $A_{2}^{\prime}$ | 9555.467 | $\ldots$ |
| (0, $3^{3}$ ) | 5 | 1 | $E^{\prime}$ | 8758.187 | 8759.116 | $\left(1,2^{2}\right)$ | 10 | 12 | $A_{2}^{\prime}$ | 9590.213 | 9598.235 |
| $\left(1,2^{0}\right)$ | 5 | 4 | $E^{\prime}$ | 8793.437 | 8793.787 |  | 10 | 12 | ${ }_{2}$ | 950.213 | 959.235 |
| $\left(1,2^{0}\right)$ | 5 | 2 | $E^{\prime}$ | 8937.686 | 8938.016 | $\left(0,0^{0}\right)$ | 10 | 10 | $E^{\prime}$ | 2451.768 | 2451.609 |
| (0, $0^{0}$ ) | 5 | 3 | $A_{2}^{\prime \prime}$ | 1080.607 | 1080.490 | (0, $0^{0}$ ) | 10 | 8 | $E^{\prime}$ | 3197.225 | 3196.903 |
| $\left(0,1^{1}\right)$ | 5 | 6 | $A_{2}^{\prime \prime}$ | 3047.126 | 3047.387 | $\left(0,0^{0}\right)$ | 10 | 4 | $E^{\prime}$ | 4086.839 | 4086.428 |
| $\left(0,1^{1}\right)$ | 5 | 0 | $A_{2}^{\prime \prime}$ | 3742.949 | 3743.187 | $\left(0,0^{0}\right)$ | 10 | 2 | $E^{\prime}$ | 4297.048 | 4296.621 |
| $\left(1,0^{0}\right)$ | 5 | 3 | $A_{2}^{\prime \prime}$ | 4233.648 | 4232.699 | $\left(0,1^{1}\right)$ | 10 | 11 | $E^{\prime}$ | 4539.163 | 4539.952 |
| $\left(0,2^{0}\right)$ | 5 | 3 | $A_{2}^{\prime \prime}$ | 5829.813 | 5830.419 | $\left(1,0^{0}\right)$ | 10 | 10 | $E^{\prime}$ | 5559.543 | 5559.156 |
| (0, $2^{2}$ ) | 5 | 3 | $A_{2}^{\prime \prime}$ | 5970.939 | 5971.239 | $\left(0,1^{1}\right)$ | 10 | 7 | $E^{\prime}$ | 5842.673 | 5842.782 |
| $\left(1,1^{1}\right)$ | 5 | 6 | $A_{2}^{\prime \prime}$ | 6129.772 | 6129.561 | ( $0,1^{1}$ ) | 10 | 7 | $E^{\prime}$ | 6145.484 | 6145.235 |
| (0, $2^{2}$ ) | 5 | 3 | $A_{2}^{\prime \prime}$ | 6213.395 | 6213.702 | ( $0,1^{1}$ ) | 10 | 5 | $E^{\prime}$ | 6227.051 | 6226.789 |
| $\left(1,1^{1}\right)$ | 5 | 0 | $A_{2}^{\prime \prime}$ | 6753.528 | 6753.128 | $\left(0,1^{1}\right)$ | 10 | 1 | $E^{\prime}$ | 6327.058 | 6326.432 |
| (2, $0^{0}$ ) | 5 | 3 | $A_{2}^{\prime \prime}$ | 7293.412 | 7292.111 | $\left(0,1^{1}\right)$ | 10 | 5 | $E^{\prime}$ | 6628.948 | 6628.627 |
| $\left(0,3^{1}\right)$ | 5 | 6 | $A_{2}^{\prime \prime}$ | 7551.059 | 7552.148 | $\left(0,1^{1}\right)$ | 10 | 1 | $E^{\prime}$ | 6666.099 | 6666.069 |
| $\left(0,3^{1}\right)$ | 5 | 0 | $A_{2}^{\prime \prime}$ | 8136.357 | 8137.599 | (0, $2^{0}$ ) | 10 | 10 | $E^{\prime}$ | 7034.483 | 7036.309 |
| $\left(0,3^{3}\right)$ | 5 | 6 | $A_{2}^{\prime \prime}$ | 8276.489 | 8277.121 | (0, $1^{1}$ ) | 10 | 1 | $E^{\prime}$ | 7055.797 | 7055.249 |
| $\left(0,3^{3}\right)$ | 5 | 0 | $A_{2}^{\prime \prime}$ | 8650.563 | 8651.706 | (1, $0^{0}$ ) | 10 | 4 | $E^{\prime}$ | 7192.905 | 7191.838 |
| $\left(1,2^{2}\right)$ | 5 | 3 | $A_{2}^{\prime \prime}$ | 8807.236 | 8808.113 | $\left(1,0^{0}\right)$ | 10 | 2 | $E^{\prime}$ | 7382.878 | 7381.697 |
| $\left(1,2^{0}\right)$ | 5 | 3 | $A_{2}^{\prime \prime}$ | 8886.355 | 8886.702 | $\left(1,1^{1}\right)$ | 10 | 11 | $E^{\prime}$ | 7474.786 | 7477.810 |
|  |  |  |  |  |  | (0, $2^{2}$ ) | 10 | 10 | $E^{\prime}$ | 7686.487 | 7687.224 |
| $\left(0,0^{0}\right)$ | 5 | 5 | $E^{\prime \prime}$ | 729.089 | 729.014 | $\left(0,2^{0}\right)$ | 10 | 8 | $E^{\prime}$ | 7785.850 | 7787.014 |
| (0, $0^{0}$ ) | 5 | 1 | $E^{\prime \prime}$ | 1250.447 | 1250.311 | (0, $2^{2}$ ) | 10 | 8 | $E^{\prime}$ | 8006.069 | 8006.515 |
| (0, 1 ${ }^{1}$ ) | 5 | 4 | $E^{\prime \prime}$ | 3396.264 | 3396.538 | (0, $2^{2}$ ) | 10 | 8 | $E^{\prime}$ | 8443.772 | 8443.842 |
| (0, 1 ${ }^{1}$ ) | 5 | 4 | $E^{\prime \prime}$ | 3509.980 | 3510.150 | (0, $0^{0}$ ) | 10 | 9 | $A_{2}^{\prime \prime}$ | 2857.013 | 2856.729 |
| (0, 1 ${ }^{1}$ ) | 5 | 2 | $E^{\prime \prime}$ | 3660.107 | 3660.349 | (0, $0^{0}$ ) | 10 | 3 | $A_{2}^{\prime \prime}$ | 4215.669 | 4215.239 |
| (0, $1^{1}$ ) | 5 | 2 | $E^{\prime \prime}$ | 3792.938 | 3793.033 | $\left(1,0^{0}\right)$ | 10 | 9 | $A_{2}^{\prime \prime}$ | 5945.606 | 5944.843 |
| $\left(1,0^{0}\right)$ | 5 | 5 | $E^{\prime \prime}$ | 3889.573 | 3888.684 | $\left(0,1^{1}\right)$ | 10 | 6 | $A_{2}^{\prime \prime}$ | 6087.601 | 6087.546 |
| $\left(1,0^{0}\right)$ | 5 | 1 | $E^{\prime \prime}$ | 4399.673 | 4398.704 | $\left(0,1^{1}\right)$ | 10 | 6 | $A_{2}^{\prime \prime}$ | 6412.622 | 6412.298 |

TABLE VIII. (Continued.)

| $\left(v_{1}, v_{2}^{l}\right)$ | $J$ | $G$ | $\Gamma$ | $E$ | $E$ (Ref. 48) | $\left(v_{1}, v_{2}^{l}\right)$ |  | $G$ | I | $E$ | $E$ (Ref. 48) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\left(0,1^{1}\right)$ | 10 | 0 | $A_{2}^{\prime \prime}$ | 7080.879 | 7080.385 | $\left(0,0^{0}\right)$ | 15 | 3 | $A_{2}^{\prime \prime}$ | 8539.756 | ... |
| $\left(1,0^{0}\right)$ | 10 | 3 | $A_{2}^{\prime \prime}$ | 7317.492 | 7316.346 | (0, 1 ${ }^{1}$ ) | 15 | 12 | $A_{2}^{\prime \prime}$ | 8773.773 | 8785.042 |
| (0, $2^{0}$ ) | 10 | 9 | $A_{2}^{\prime \prime}$ | 7506.638 | ... | (0, 1 ${ }^{1}$ ) | 15 | 12 | $A_{2}^{\prime \prime}$ | 9249.313 | ... |
| (0, $2^{2}$ ) | 10 | 9 | $A_{2}^{\prime \prime}$ | 8091.794 | $\ldots$ | $\left(1,0^{0}\right)$ | 15 | 15 | $A_{2}^{\prime \prime}$ | 9470.135 | $\cdots$ |
| (1, $0^{0}$ ) | 10 | 3 | $A_{2}^{\prime \prime}$ | 8711.547 | $\ldots$ | $\left(0,1^{1}\right)$ | 15 | 6 | $A_{2}^{\prime \prime}$ | 10261.999 | $\ldots$ |
| (0, $2^{0}$ ) | 10 | 9 | $A_{2}^{\prime \prime}$ | 8932.027 | $\ldots$ | (0, $2^{2}$ ) | 15 | 15 | $A_{2}^{\prime \prime}$ | 10334.975 | $\ldots$ |
| $\left(1,1^{1}\right)$ | 10 | 6 | $A_{2}^{\prime \prime}$ | 9057.381 | 9057.019 | (2, $0^{0}$ ) | 15 | 9 | $A_{2}^{\prime \prime}$ | 10550.716 | $\ldots$ |
| (0, $2^{2}$ ) | 10 | 3 | $A_{2}^{\prime \prime}$ | 9268.059 | ... | $\left(0,1^{1}\right)$ | 15 | 0 | $A_{2}^{\prime \prime}$ | 10788.063 | $\ldots$ |
| $\left(1,1^{1}\right)$ | 10 | 6 | $A_{2}^{\prime \prime}$ | 9347.706 | 9347.039 | $\left(0,3^{3}\right)$ | 15 | 18 | $A_{2}^{\prime \prime}$ | 10907.195 | $\ldots$ |
| (0, $3^{3}$ ) | 10 | 12 | $A_{2}^{\prime \prime}$ | 9403.539 |  |  |  |  |  |  |  |
| $\left(0,2^{2}\right)$ | 10 | 3 | $A_{2}^{\prime \prime}$ | 9564.754 |  | (0, $0^{0}$ ) | 15 | 13 | $E^{\prime \prime}$ | $6154.903$ | $\ldots$ |
| $\left(1,1^{1}\right)$ | 10 | 0 | $A_{2}^{\prime \prime}$ | 10006.830 |  | $\left(0,0^{0}\right)$ | 15 | 11 | $E^{\prime \prime}$ | $6943.399$ | $\ldots$ |
|  |  |  |  |  |  | $\left(0,1^{1}\right)$ | 15 | 16 | $E^{\prime \prime}$ | 6954.510 | $\cdots$ |
| (0, $0^{0}$ ) | 10 | 5 | $E^{\prime \prime}$ | 3485.134 | 3484.774 | $\left(0,1^{1}\right)$ | 15 | 14 | $E^{\prime \prime}$ | 7833.973 | $\cdots$ |
| (0, $0^{0}$ ) | 10 | 5 | $E^{\prime \prime}$ | 3926.570 | 3926.171 | $\left(0,0^{0}\right)$ | 15 | 7 | $E^{\prime \prime}$ | 8018.194 | $\ldots$ |
| (0, $0^{0}$ ) | 10 | 1 | $E^{\prime \prime}$ | 4348.792 | 4348.363 | (0, 1 ${ }^{1}$ ) | 15 | 14 | $E^{\prime \prime}$ | 8267.157 | $\ldots$ |
| (0, 1 ${ }^{1}$ ) | 10 | 10 | $E^{\prime \prime}$ | 5026.095 | 5026.286 | (0, $0^{0}$ ) | 15 | 5 | $E^{\prime \prime}$ | 8353.640 | $\ldots$ |
| (0, 1 ${ }^{1}$ ) | 10 | 8 | $E^{\prime \prime}$ | 5555.350 | 5555.503 | (0, $0^{0}$ ) | 15 | 1 | $E^{\prime \prime}$ | 8692.435 | $\ldots$ |
| (0, 1 ${ }^{1}$ ) | 10 | 8 | $E^{\prime \prime}$ | 5827.898 | 5827.729 | (0, $2^{2}$ ) | 15 | 17 | $E^{\prime \prime}$ | 8918.178 | $\ldots$ |
| (0, 1 ${ }^{1}$ ) | 10 | 4 | $E^{\prime \prime}$ | 6401.130 | 6401.110 | $\left(1,0^{0}\right)$ | 15 | 13 | $E^{\prime \prime}$ | 9133.420 | $\cdots$ |
| (1, $0^{0}$ ) | 10 | 7 | $E^{\prime \prime}$ | 6580.702 | 6579.774 | (0, 1 ${ }^{1}$ ) | 15 | 10 | $E^{\prime \prime}$ | 9465.563 | $\ldots$ |
| (0, 1 ${ }^{1}$ ) | 10 | 2 | $E^{\prime \prime}$ | 6612.470 | 6612.443 | (0, $0^{0}$ ) | 20 | 18 | $A_{2}^{\prime}$ | 9750.216 | $\ldots$ |
| (0, 1 ${ }^{1}$ ) | 10 | 4 | $E^{\prime \prime}$ | 6812.167 | 6811.731 | (0, $0^{0}$ ) | 20 | 18 | $A_{2}^{\prime}$ | 10283.596 | $\cdots$ |
| $\left(1,0^{0}\right)$ | 10 | 5 | $E^{\prime \prime}$ | 6968.083 | 6967.289 | (0, $0^{0}$ ) | 20 | 12 | $A_{2}^{\prime}$ | 12089.177 | $\ldots$ |
| (0, $1^{1}$ ) | 10 | 2 | $E^{\prime \prime}$ | 7073.194 | 7072.408 | $\left(1,0^{0}\right)$ | 20 | 18 | $A_{2}^{\prime}$ | 12577.884 | $\ldots$ |
| (0, $2^{2}$ ) | 10 | 11 | $E^{\prime \prime}$ | 7219.631 | 7220.652 | $\left(1,0^{0}\right)$ | 20 | 18 | $A_{2}^{\prime}$ | 13063.787 | $\ldots$ |
| $\left(1,0^{0}\right)$ | 10 | 1 | $E^{\prime \prime}$ | 7430.491 | 7429.303 | (0, $1^{1}$ ) | 20 | 21 | $A_{2}^{\prime}$ | 13113.814 | $\ldots$ |
| (0, $1^{1}$ ) | 10 | 2 | $E^{\prime \prime}$ | 7921.086 | ... | $\left(0,0^{0}\right)$ | 20 | 6 | $A_{2}^{\prime}$ | 13314.547 | $\ldots$ |
| $\left(0,2^{0}\right)$ | 10 | 7 | $E^{\prime \prime}$ | 8108.997 | 8109.399 | (0, $0^{1}$ ) | 20 | 21 | $A_{2}^{\prime}$ | 13617.513 | $\cdots$ |
| $\left(1,1^{1}\right)$ | 10 | 8 | $E^{\prime \prime}$ | 8379.827 | 8380.666 | $\left(0,0^{0}\right)$ | 20 | 20 | $E^{\prime}$ | 8541.217 | $\ldots$ |
| ( $0,0^{0}$ ) | 15 | 12 | $A_{2}^{\prime}$ | 6575.316 | $\ldots$ | $\left(0,0^{0}\right)$ | 20 | 16 | $E^{\prime}$ | 10700.847 | $\ldots$ |
| $\left(0,1^{1}\right)$ | 15 | 15 | $A_{2}^{\prime}$ | 7674.211 | $\ldots$ | (0, $1^{1}$ ) | 20 | 19 | $E^{\prime}$ | 11193.284 | $\cdots$ |
| (0, $0^{0}$ ) | 15 | 6 | $A_{2}^{\prime}$ | 8198.855 | $\ldots$ | (0, $0^{0}$ ) | 20 | 20 | $E^{\prime}$ | 11449.958 | $\ldots$ |
| (0, $0^{0}$ ) | 15 | 0 | $A_{2}^{\prime}$ | 8705.218 | ... | $\left(1,0^{0}\right)$ | 20 | 20 | $E^{\prime}$ | 11490.210 | $\ldots$ |
| $\left(1,0^{0}\right)$ | 15 | 12 | $A_{2}^{\prime}$ | 9516.220 | $\ldots$ | $\left(0,1^{1}\right)$ | 20 | 19 | $E^{\prime}$ | 11737.466 | $\ldots$ |
| (0, 1 ${ }^{1}$ ) | 15 | 9 | $A_{2}^{\prime}$ | 9760.939 | $\ldots$ | $\left(0,0^{0}\right)$ | 20 | 14 | $E^{\prime}$ | 12198.864 | $\ldots$ |
| $\left(0,1^{1}\right)$ | 15 | 9 | $A_{2}^{\prime}$ | 10308.177 | $\ldots$ | $\left(1,1^{1}\right)$ | 20 | 19 | $E^{\prime}$ | 12220.050 | $\ldots$ |
| $\left(1,1^{1}\right)$ | 15 | 15 | $A_{2}^{\prime}$ | 10535.633 | $\ldots$ | $\left(0,0^{0}\right)$ | 20 | 10 | $E^{\prime \prime}$ | 12591.204 | $\ldots$ |
| $\left(0,1^{1}\right)$ | 15 | 3 | $A_{2}^{\prime}$ | 10591.474 | $\ldots$ | $\left(0,0^{0}\right)$ | 20 | 15 | $A_{2}^{\prime \prime}$ | 11112.254 | $\ldots$ |
| $\left(1,1^{1}\right)$ | 15 | 15 | $A_{2}^{\prime}$ | 10894.706 | $\cdots$ | $\left(0,1^{1}\right)$ | 20 | 18 | $A_{2}^{\prime \prime}$ | 11933.578 | $\ldots$ |
| (0, $0^{0}$ ) | 15 | 14 | $E^{\prime}$ | 5680.067 | $\ldots$ | $\left(0,0^{0}\right)$ | 20 | 18 9 | $A_{2}^{\prime \prime}$ $A_{2}^{\prime \prime}$ | 12413.818 12803.849 | $\ldots$ |
| (0, $0^{0}$ ) | 15 | 10 | $E^{\prime}$ | 7268.264 | $\cdots$ | $\left(1,1^{1}\right)$ | 20 | 18 | $A_{2}^{\prime \prime}$ | 13219.849 | $\ldots$ |
| (0, $0^{0}$ ) | 15 | 8 | $E^{\prime}$ | 7802.829 | $\cdots$ | $\left(0,0^{0}\right)$ | 20 | 3 | $A_{2}^{\prime \prime}$ | 13739.397 | $\cdots$ |
| $\left(0,1^{1}\right)$ | 15 | 13 | $E^{\prime}$ | 8378.518 | 8385.960 | $\left(1,1^{1}\right)$ | 20 | 18 | $A_{2}^{\prime \prime}$ | 13803.937 | $\ldots$ |
| (0, $0^{0}$ ) | 15 | 4 | $E^{\prime}$ | 8473.833 | ... | $\left(1,0^{0}\right)$ | 20 | 15 | $A_{2}^{\prime \prime}$ | 14135.544 | $\ldots$ |
| (0, $0^{0}$ ) | 15 | 2 | $E^{\prime}$ | 8655.538 | $\ldots$ | (0, $0^{0}$ ) | 20 | 19 | $E^{\prime \prime}$ | 9284.714 | $\ldots$ |
| (1, $0^{0}$ ) | 15 | 14 | $E^{\prime}$ | 8693.101 | $\cdots$ |  |  |  |  |  |  |
| (0, 1 ${ }^{1}$ ) | 15 | 13 | $E^{\prime}$ | 8795.509 | $\ldots$ | (0, $0^{0}$ ) | 20 | 17 | $E^{\prime \prime}$ | 10284.461 | $\ldots$ |
| $\left(0,1^{1}\right)$ | 15 | 11 | $E^{\prime}$ | 9143.916 | $\ldots$ | $\left(0,1^{1}\right)$ | 20 | 20 | $E^{\prime \prime}$ | 11199.609 | $\ldots$ |
| (0, 1 ${ }^{1}$ ) | 15 | 11 | $E^{\prime}$ | 9652.220 | ... | (0, $0^{0}$ ) | 20 | 13 | $E^{\prime \prime}$ | 11795.800 | $\cdots$ |
| (0, $2^{2}$ ) | 15 | 16 | $E^{\prime}$ | 9732.107 | $\ldots$ | (1, $0^{0}$ ) | 20 | 19 | $E^{\prime \prime}$ | 12194.310 | $\ldots$ |
| (0, $0^{0}$ ) | 15 | 15 | $A_{2}^{\prime \prime}$ | 5092.286 | $\ldots$ | (0, $0^{0}$ ) | 20 | 11 | $E^{\prime \prime}$ | 12354.754 | $\ldots$ |
| (0, $0^{0}$ ) | 15 | 9 | $A_{2}^{\prime \prime}$ | 7553.581 | $\ldots$ | (0, $0^{0}$ ) | 20 | 13 | $E^{\prime \prime}$ | 12754.707 | $\ldots$ |
| (1, $0^{0}$ ) | 15 | 15 | $A_{2}^{\prime \prime}$ | 8124.367 | $\ldots$ |  |  |  |  |  |  |

The calculations have also been extended to several excited states using the same spatial conformations (8469 data points for each state). We plan to obtain analytical global potential energy surfaces and rovibrational analysis for excited states
of the $\mathrm{H}_{3}^{+}$system using these data and the same procedure presented here for the ground state. Moreover, we plan to study the prototype reaction $\mathrm{H}_{2}^{+}+\mathrm{H}_{2} \rightarrow \mathrm{H}_{3}^{+}+\mathrm{H}$, for which the present $\mathrm{H}_{3}^{+}$study is a necessary step to obtain the $\mathrm{H}_{4}^{+}$GPES.

TABLE IX. Comparison of observed (Refs. 41,43,71,44,46,72, and 73) and calculated transitions (in $\mathrm{cm}^{-1}$ ) in different absorption bands of $\mathrm{H}_{3}^{+}$, for the present calculations and using the energy levels of Ref. 48.

| Band | Number of transitions <br> considered | rms error <br> present GPES | rms error <br> Exp+Fit (Ref. 48) |
| :---: | :---: | :---: | :---: |
| $\nu_{2} \leftarrow 0:\left(0,1^{1}\right) \leftarrow\left(0,0^{0}\right)$ | 92 | 0.282 | 0.009 |
| $\nu_{1} \leftarrow 0:\left(1,0^{0}\right) \leftarrow\left(0,0^{0}\right)$ | 9 | 0.693 | 0.016 |
| $2 \nu_{2} l_{0} \leftarrow \nu_{2}:\left(0,2^{0}\right) \leftarrow\left(0,1^{1}\right)$ | 14 | 0.371 | 0.268 |
| $2 \nu_{2} l_{2} \leftarrow \nu_{2}:\left(0,2^{2}\right) \leftarrow\left(0,1^{1}\right)$ | 77 | 0.145 | 0.008 |
| $\nu_{1}+\nu_{2} \leftarrow \nu_{1}:\left(1,1^{1}\right) \leftarrow\left(1,0^{0}\right)$ | 21 | 0.477 | 0.009 |
| $\nu_{2}+\nu_{1} \leftarrow \nu_{2}:\left(1,1^{1}\right) \leftarrow\left(0,1^{1}\right)$ | 21 | 0.630 | 0.125 |
| $3 \nu_{2} l_{1} \leftarrow 0:\left(0,3^{1}\right) \leftarrow\left(0,0^{0}\right)$ | 15 | 1.242 | 0.056 |

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