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Rapid directivity detection by azimuthal amplitude spectra inversion

Simone Cesca · Sebastian Heimann ·
Torsten Dahm

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Abstract An early detection of the presence of rupture directivity plays a major role in the correct estimation of ground motions and risks associated to the earthquake occurrence. We present here a simple method for a fast detection of rupture directivity, which may be additionally used to discriminate fault and auxiliary planes and have first estimations of important kinematic source parameters, such as rupture length and rupture time. Our method is based on the inversion of amplitude spectra from P-wave seismograms to derive the apparent duration at each station and on the successive modelling of its azimuthal behaviour. Synthetic waveforms are built assuming a spatial point source approximation, and the finite apparent duration of the spatial point source is interpreted in terms of rupture directivity. Since synthetic seismograms for a point source are calculated very quickly, the presence of directivity may be detected within few seconds, once a focal mechanism has been derived. The method is here first tested using synthetic datasets, both for linear and planar sources, and then successfully applied to recent Mw 6.2–6.8 shallow earthquakes in Peloponnese, Greece. The method is suitable for

automated application and may be used to improve kinematic waveform modelling approaches.

Keywords Directivity · Earthquake source · Kinematic model · Amplitude spectra

1 Introduction 31

Tectonically driven shallow earthquake sources are generally explained by means of shear cracks occurring along a limited, almost planar region, we refer as the focal region. A point source representation is a common first approximation, which is valid when treating far-field low frequency seismic waveform, using wavelengths larger than the rupture size. Higher frequencies seismograms and spectra contain information which can be related to the finiteness of the rupture process and thus can be used to determine parameters describing the finite source. Size and shape of the rupture area, rupture velocity and preferential rupture directions, an effect known as rupture directivity, are some of the parameters which can be retrieved by the analysis of high-frequency waveforms. In particular, we are interested here in discussing the problem of early detection of rupture directivity, distinguishing between a prominent or partial unilateral rupture (a case which will be further referred as asymmetric bilateral rupture), and a

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53 bilateral one, with rupture nucleating at the centre
54 of the rupture area and propagating toward its
55 edges. The azimuthal dependency of amplitudes
56 and durations of different seismic phases is a first
57 indicator of directivity effects and is consequence
58 of the characteristics of the finite rupture process
59 along the fault plane, specifically the main di-
60 rection and speed of the rupture front propaga-
61 tion. Directivity has been often observed and has
62 been modelled for several earthquakes in the past,
63 with several studies treating specific earthquakes
64 or limited datasets (e. g., McGuire et al. 2002;
65 Warren and Shearer 2006; Caldeira et al. 2009). A
66 quick detection of directivity effects is important
67 towards a correct estimation of ground motions,
68 stress field perturbations and tsunamogenic risks
69 and consequently to mitigate earthquake effects.
70 These considerations provide important reasons
71 to further investigate and develop specific tools
72 for stable, rapid and automated directivity de-
73 tection, which can be used within early warning
74 systems.

75 Several methods have been applied in the past
76 to detect and classify earthquake source direc-
77 tivity. A common approach is the identification
78 of predominant unilateral ruptures from the time
79 duration and spectral analysis of body wave pulses
80 (e.g. Boore and Joyner 1978; Beck et al. 1995;
81 Warren and Shearer 2006; Caldeira et al. 2009).
82 Pulse lengths at different stations are interpreted
83 in terms of the apparent duration of the source
84 time function (STF), and their variation in depen-
85 dence on azimuth and incidence angle is inter-
86 preted to detect directivity: similarly to a Doppler
87 effect in classical physics, shorter STFs would in-
88 dicate a rupture propagating towards the consid-
89 ered station, while longer pulses indicate a rupture
90 propagation in the opposite direction. Directiv-
91 ity effects may also be revealed based on the
92 analysis of surface waves at different azimuths
93 (Ben-Menahem 1961; Pro et al. 2007). Whereas
94 time domain methods remain more common, a
95 significant contribution within this type of inver-
96 sion methods was provided by the spectral ap-
97 proach discussed in Warren and Shearer (2006).
98 This method is based on the spectral estimation
99 of the pulse broadening and accounts for the az-
100 imuthal and incidence angle dependencies; it is
101 well suited for the analysis of intermediate and

102 deep focus earthquakes and was successfully ap-
103 plied to several events. The main limits of this
104 class of methods are related to the fact that wave
105 propagation and the superposition of different
106 seismic phases are not accounted, since wave
107 propagation effects between source and receiver
108 (Green's functions) are limited to the estimation
109 of the incidence angle of given seismic phases.
110 Another possible limitation is the requirement of
111 several stations with good azimuthal coverage in
112 order to ensure reliable results. A second range
113 of applications, which on the contrary accounts
114 precisely for the effects of the earth's model on
115 the observed waveforms, is based on empirical
116 Green's functions technique (Hartzell 1978; Li
117 and Toksöz 1993; Velasco et al. 1994; Cassidy
118 1995; Müller 1985; Velasco et al. 2004; Vallée
119 2007). In this case, an aftershock with common
120 hypocenter and focal mechanism of the studied
121 event can be used to remove path effects, and iso-
122 late finite source apparent durations at different
123 stations. Evidently, the application of these tech-
124 niques is strongly limited by the availability of
125 a proper aftershock. Brüstle and Müller (1987),
126 and Imanishi and Takeo (2002) have investigated
127 the adoption of master-event techniques to detect
128 directivity: the identification of stopping phases
129 (Madariaga 1977, 1983; Bernard and Madariaga
130 1984; Spudich and Frazer 1984) at different sta-
131 tions was used there to determine the main di-
132 rection of rupture propagation, besides other
133 source properties. Stopping phases identification
134 (Imanishi and Takeo 1998, 2002) typically re-
135 quires a careful waveform analysis, which may be
136 hardly implemented within automated routines.
137 A third group of techniques are based on com-
138 plete kinematic waveform inversion, with the aim
139 of retrieving a most detailed image of the finite
140 rupture process, not limited to the identification
141 of directivity. The range of methods and appli-
142 cations is very wide, including higher order mo-
143 ment tensor analysis (Dahm and Krüger 1999;
144 McGuire et al. 2001, 2002), detailed slip map
145 approaches (e.g. Olson and Apsel 1982; Hartzell
146 and Helmberger 1982; Hartzell and Heaton 1983;
147 Beroza and Spudich 1988), and inversion meth-
148 ods adopting constrained and simplified kinematic
149 models (Dreger and Kaverina 2000; Vallée and
150 Bouchon 2004; Galovic et al. 2009; Cesca et al.

151 2010). All these methods have a significant poten-
 152 tial for a stable determination of directivity but
 153 their adoption towards its very fast detection is
 154 limited, often requiring time consuming computa-
 155 tion of synthetic seismograms for several extended
 156 source models. Methods developed by Dreger and
 157 Kaverina (2000) and following Cesca et al. (2010)
 158 have shown a good performance and have been
 159 tested for near real-time applications, but they are
 160 still based on extended source representations and
 161 thus require heavier computations with respect to
 162 our method. Finally, recent results by Zahradnik
 163 et al. (2008) showed the possibility of discrimi-
 164 nating the true fault plane on the base of spatial
 165 offsets between epicentre and centroid locations.
 166 However, the method has been currently applied
 167 only to a limited number of earthquakes, with
 168 variable results, and the determination of directiv-
 169 ity may be beyond its possibilities, for example for
 170 symmetric bilateral ruptures.

171 We present here a simple alternative method to
 172 quickly detect directivity for shallow earthquakes
 173 and discuss it with the aid of a set of applica-
 174 tions, including both synthetic datasets and obser-
 175 vations from recent earthquakes in Greece. Our
 176 method is based on a point source representation,
 177 which drastically reduces computational require-
 178 ments and makes it feasible for early detection.
 179 Directivity is detected on the basis of a frequency
 180 domain inversion of the apparent duration at
 181 each station and the further interpretation of its
 182 azimuthal variation. Main strength points of the
 183 proposed method include the adoption of a com-
 184 mon dataset and modelling tools for focal mech-
 185 anism and directivity determination, the inclusion
 186 of Green's functions accounting for wave propa-
 187 gation through the chosen earth models without
 188 needing specific aftershocks, and the simplicity
 189 and quickness of the inversion process.

190 2 Directivity and amplitude spectra inversion

191 We make here use of the recently developed
 192 Kiwi tools (Heimann 2010; Cesca et al. 2010;
 193 <http://kinherd.org>), which provide a flexible in-
 194 strument to generate synthetic seismograms for
 195 point and extended sources and to invert different
 196 earthquake source parameters, allowing the se-

lection of different waveform tapers, frequency 197
 filters, inversion domains and misfit functions. 198
 Cesca et al. (2010) showed successful applications 199
 to shallow earthquake at regional distances, and 200
 was able to derive both point source (best double 201
 couple, DC, model, scalar moment and centroid 202
 depth) and extended source (fault plane discrim- 203
 ination, rupture size, rupture time, rupture nucle- 204
 ation) parameters. 205

The first inversion step follows the approach 206
 described in Cesca et al. (2010), to obtain the focal 207
 mechanism, scalar moment and centroid depth: 208

1. Focal mechanism. We invert amplitude spec- 209
 tra of full waveforms, according to Cesca et al. 210
 (2010), to derive a point source focal mech- 211
 anism (DC, depth and scalar moment); the 212
 source epicentral location is assumed to be 213
 originally known. 214
2. Polarities. The focal mechanism presents a 215
 polarity ambiguity, which can be solved by 216
 comparing observed displacements and syn- 217
 thetic seismograms for the two possible polar- 218
 ity configurations; however, the detection of 219
 the true polarity is here not strictly required, 220
 as the whole inversion process is carried out 221
 in the frequency domain, and only amplitude 222
 spectra are involved in the fitting procedure. 223

The source representation through the Kiwi tools 224
 allows the adoption of different rise times. For a 225
 spatially extended source model, where the rup- 226
 ture region is discretised into a number of spatial 227
 point sources, the rise time represent the time 228
 during which each point source radiates seismic 229
 energy. The duration of the whole rupture process 230
 is related to rise and rupture times. If we adopt 231
 a point source representation, the rise time will 232
 represent the duration of the source time function. 233
 This parameter was used in Cesca et al. (2010) 234
 to have a first, rough, estimate of rupture times 235
 and to choose a proper rise time during kinematic 236
 source modelling. We proceed here differently: in- 237
 stead of determining true duration of the rupture 238
 process, we investigate apparent durations as seen 239
 by individual stations. 240

In detail, during the second inversion step, we 241
 proceed as follows (Fig. 1 illustrates an example 242
 of the main steps, relative to selected seismic 243

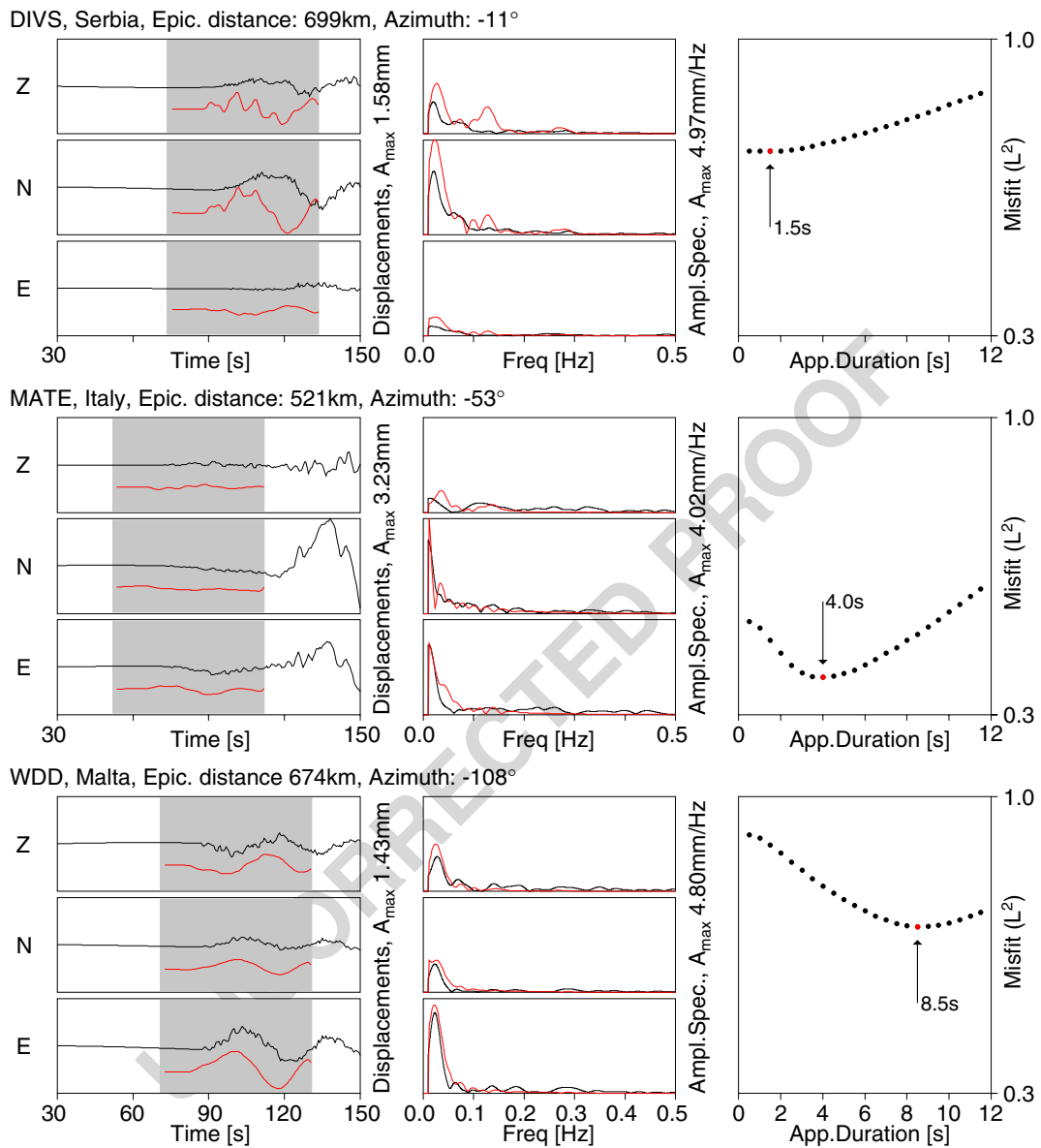


Fig. 1 Example of the procedure followed to derive the apparent source duration at different stations. Selected waveforms, spectra and amplitude spectra inversion results refer to an application to the Andravida earthquake, Greece, which is further extensively discussed in this study. *Left*: filtered displacements (*black lines*) and synthetic seismograms (*red lines*) for the chosen point source model are

tapered to select P waves time windows (*grey intervals*). *Centre*: amplitude spectra comparison (*red lines* correspond to the best fitting synthetic spectra, after comparing several source durations). *Right*: comparison of amplitude spectra misfit values for different source durations (best solutions for each station are identified by *red circles*)

244 waveforms from the Andravida earthquake,
245 which is later discussed in the text):

246 1. Waveform selection. We use all available spatial
247 components, preferably using North, East

and vertical orientations, rather than rotated 248
traces, in order to have P wave energy on all 249
traces (which is theoretically null on transversal 250
components); the presence of more traces 251
for each station has a smoothing effect; after 252

253 testing with different datasets, we found that
 254 more components provide more stability. We
 255 perform a deconvolution of the instrumental
 256 response from the data, and conversion to
 257 displacements.

258 2. Tapering. We limit the inversion process to P-
 259 wave time windows, which are automatically
 260 selected on the base of the source-receiver
 261 geometry and theoretical arrival time for the
 262 earth model used during the inversion (an ar-
 263 rival time database is calculated in advance, to
 264 reduce computational effort at the time of the
 265 inversion); for the case studies here described
 266 we use 60 s length time windows, starting 15 s
 267 before theoretical first P arrival, and apply
 268 a bandpass filter in the range 0.01–0.5 Hz
 269 (these parameters may be modified depending
 270 on the earthquake size, the source depth, the
 271 average duration, and the range of epicen-
 272 tral distances where waveforms are inverted).
 273 Tapers should be chosen in order to resolve
 274 directivity effects. A minimum length should
 275 account at least for two times the average rup-
 276 ture duration and for different periods at the
 277 frequency range used for the inversion. For
 278 stations located at small epicentral distances,
 279 with minor delay between S and P phases,
 280 tapers may be modified to avoid S waves.

281 3. Scalar moment inversion. Since the estima-
 282 tion of the scalar moment may slightly vary
 283 depending on the inversion approach (e.g.,
 284 full waveform or body waves, time domain or
 285 amplitude spectra inversion, etc.), we mention
 286 here the possibility to perform a specific in-
 287 version using an approach consistent with the
 288 following directivity inversion. Traces from
 289 all seismic stations would be used to invert
 290 the scalar moment (e.g. by amplitude spectra
 291 inversion, using a Levenberg–Marquardt ap-
 292 proach and an L^2 norm misfit function). In
 293 the following applications this step is not per-
 294 formed, as we count with stable estimations of
 295 the scalar moments, provided by the fit of low
 296 frequency amplitude spectra from the whole
 297 waveforms.

298 4. Apparent duration inversion. For each of the
 299 stations, we perform an amplitude spectra in-
 300 version to derive the apparent source duration
 301 at that station; the frequency domain inver-

302 sion approach is less sensitive to unmodelled
 303 structural heterogeneities; we perform here
 304 a grid search for possible durations (for the
 305 following case studies, tested durations varies
 306 up to 30 s, with an increment of 0.5 s); in
 307 general we observe smooth single-minimum
 308 curves of misfit versus apparent durations,
 309 and tests with different inversion approaches
 310 (e.g. gradient methods) have shown very con-
 311 sistent results with respect to the grid walk
 312 procedure.

The apparent source time function durations can
 313 be then quickly interpreted in term of simplified
 314 laws for finite rupture models. With the aid of
 315 synthetic tests and application to selected earth-
 316 quake datasets, we will show that, often, it is not
 317 necessary to have a complex rupture model to
 318 fit the azimuthal distribution of apparent dura-
 319 tions. For simple extended source model, such as
 320 a one-dimensional linear source or a Haskell bi-
 321 dimensional rupture model (Haskell 1964), the
 322 effects of directivity can be treated analytically. A
 323 unilateral rupture along a horizontal linear source
 324 will produce theoretical P-wave pulses of shorter
 325 duration for stations located toward the rupture
 326 propagation, and larger duration for stations in
 327 the opposite direction. Bilateral ruptures result
 328 in a minor azimuthal variation of the apparent
 329 source time function. A range of asymmetrically
 330 bilateral rupture models exists in between. Effects
 331 of oblique and vertical rupture propagations may
 332 also be modelled but are more difficult to reveal
 333 (Beck et al. 1995) and have been more rarely
 334 observed (e.g. Eshghi and Zare 2003; Nadim et al.
 335 2004). 336

We originally focus on the two-dimensional
 337 problem, with source and observer laying on the
 338 same plane. Let us assume a horizontal linear
 339 source model of length L , with the rupture starting
 340 at one edge (A) and propagating unilaterally till
 341 the other edge (B). The rupture time t_R is the time
 342 required for the rupture front to propagate along
 343 the entire rupture length, from A to B, at a rupture
 344 velocity v_R , which is assumed to be constant. The
 345 rise time t_r , defined as the duration of seismic
 346 source emission from a point along the source, is
 347 here assumed to be constant, according to healing
 348 front theory (Nielsen and Madariaga 2003) and
 349

350 will be further considered negligible with respect
 351 to the rupture time. Finally, v_P is the average P
 352 wave velocity at the focal region. Typically, rup-
 353 ture propagates with a velocity slightly below the
 354 shear wave velocity at the focal region, which also
 355 shows a common scale with compressional wave
 356 velocity in seismogenic regions. Then, according
 357 to Ben-Menahem and Singh (1981), and including
 358 the rise time, for a receiver located at azimuth φ
 359 (defined with respect to the direction of rupture
 360 propagation) the apparent source duration $\Delta t(\varphi)$
 361 will be given by:

$$\Delta t(\varphi) = t_r + \frac{L}{v_R} - \frac{L}{v_P} \cos(\varphi). \tag{1a}$$

362 In view of a more general formulation, also ac-
 363 counting for asymmetric and pure bilateral rup-
 364 ture, the rupture length L is divided into two
 365 segments L_1 and L_2 , with the following expression
 366 for the apparent source duration $\Delta t(\varphi)$:

$$\Delta t(\varphi) = \text{Max} \left[t_r + L_1/v_R - (L_1/v_P) \cos(\varphi), \right. \\ \left. t_r + L_2/v_R + (L_2/v_P) \cos(\varphi) \right] \tag{1b}$$

367 We can then introduce the following non-
 368 dimensional variables: $\tau(\varphi)$ is the ratio between
 369 the apparent source duration $\Delta t(\varphi)$ and the rup-
 370 ture time t_R , $t_{r/R}$ is the ratio between rise and
 371 rupture time, $v_{R/P}$ is the ratio between rupture
 372 velocity and P wave velocity at the source; L_1 and
 373 L_2 ($L_1 \geq L_2$) are expressed as $(1-\chi)L$ and χL ,
 374 respectively, χ being the ratio between the short-
 375 est segment and the entire rupture length (χ may
 376 range from 0, for a pure unilateral rupture, to 0.5,
 377 for a pure bilateral one). The azimuthal depen-
 378 dency of $\tau(\varphi)$, making use of the non-dimensional
 379 notation is the following ($t_{r/R}$ can be in general
 380 neglected):

$$\tau(\varphi) = \text{Max} \left[t_{r/R} + 1 - v_{R/P} \cos(\varphi), \right. \\ \left. t_{r/R} + L_2/L_1 + L_2/L_1 \cos(\varphi) \right]. \tag{2}$$

381 The radiation pattern for three significant cases
 382 (pure unilateral, pure bilateral and asymmetric
 383 bilateral) is shown in Fig. 2, where we have chosen
 384 $v_{R/P} = 0.5$ ($v_{R/P}$ equal to 0.25 and 1.0 for the slow
 385 and fast cases respectively), χ is equal to 0, 1/3
 386 and 1/2 for the three considered cases. We choose
 387 different rupture lengths L , in order to have a
 388 constant length of the largest rupture segment.

Symmetries of apparent duration radiation pat- 389
 terns can be observed, with a one lobe shape for 390
 a pure unilateral rupture and a two lobe shape for 391
 a pure bilateral one. The theoretical curve of the 392
 apparent rupture duration (Fig. 2, right) for the 393
 unilateral case range from $t_r + t_R - t_P$ (azimuth of 394
 rupture direction) to $t_r + t_R + t_P$ (opposite direc- 395
 tion); its average value is equal to $t_r + t_R$ (for a 396
 unilateral rupture model t_R and t_P are the rupture 397
 time and P wave travel time along the entire 398
 rupture length). For the pure bilateral rupture, the 399
 apparent duration varies between $t_r + t_R$ (perpen- 400
 dicular to rupture direction) to $t_r + t_R + t_P$ (paral- 401
 lel to rupture direction), with t_R and t_P referring 402
 here to half of the rupture length. The larger vari- 403
 ation of the apparent duration for the unilateral 404
 case, with respect to the bilateral, explains the 405
 major difficulties in observing directivity for the 406
 second case. Less known is the behaviour of asym- 407
 metric bilateral ruptures, although this model is 408
 the most general. In this case the radiation pat- 409
 tern (Fig. 2, bottom) present a deformed one- 410
 lobe shape, with the minimum observed apparent 411
 rupture duration at about 45° from the rupture di- 412
 rection of the largest rupture length. The azimuth 413
 α , where a cusp-like minimum in the apparent 414
 duration may be observed, can be obtained by 415
 equalizing the two right terms in Eq. 2: 416

$$\alpha = a \cos \left[\frac{(1 - 2\chi)}{v_{R/P}} \right]. \tag{3}$$

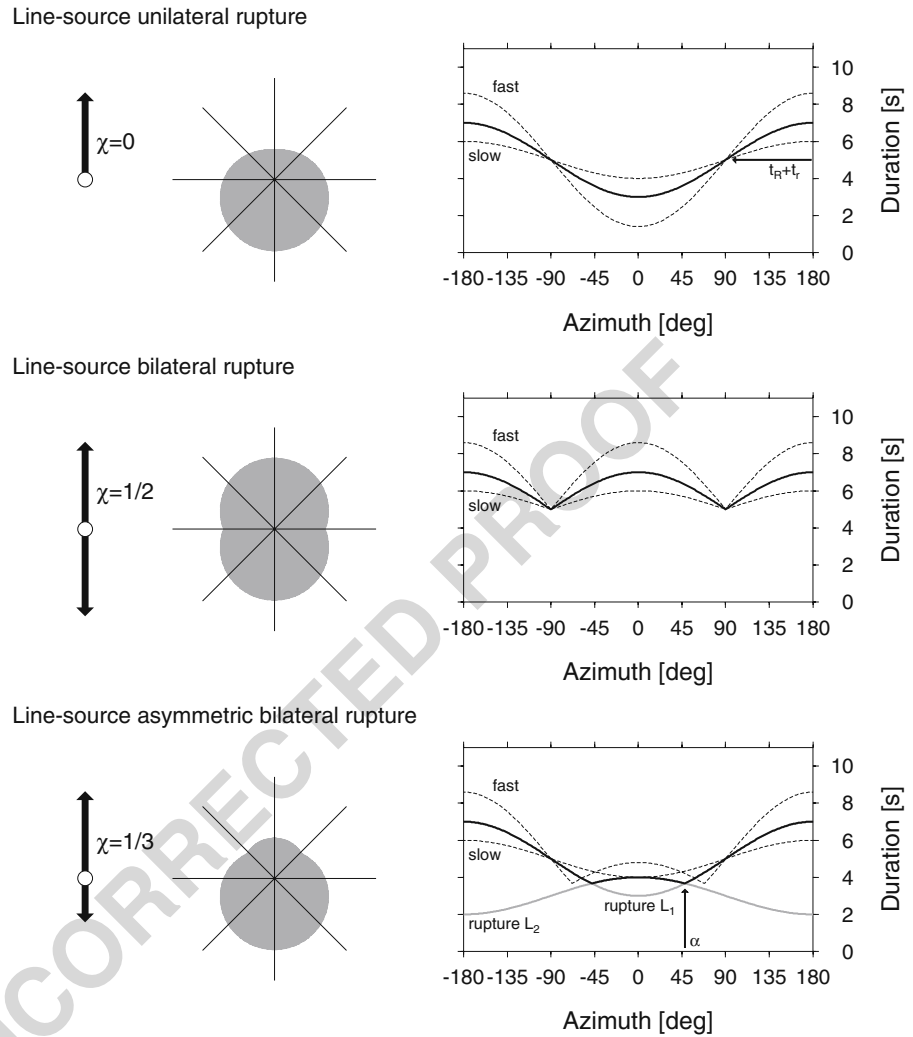
Since v_R pertains to $[0, v_P]$, it follows that for a 417
 pure unilateral rupture ($\chi = 0$) we have a single 418
 minimum in direction of rupture propagation, 419
 while cusp-like minima are not observed. For a 420
 pure bilateral rupture ($\chi = 0.5$), $\alpha = \pm\pi/2$ always. 421
 For the intermediate case of asymmetric bilateral 422
 ruptures ($0 < \chi < 0.5$), two cusp-like minima are 423
 observed if the following condition is met: 424

$$\chi > \frac{(1/v_{R/P})}{2}. \tag{4}$$

This means that the observation of two minima in 425
 the apparent duration curve indicate a dominant 426
 bilateral rupture processes. 427

Figure 2 additionally shows how larger vari- 428
 ations in the apparent duration estimations are 429
 found, when increasing the rupture velocity (in 430
 the figure, effects for an extreme case of $v_R =$ 431

Fig. 2 Theoretical models for pure unilateral (*top*), pure bilateral (*centre*) and asymmetric bilateral (*bottom*) line sources. The ruptures start at a nucleation point (*white circles, left plots*) and propagate along segments L_1 and L_2 (*black arrows*), producing different apparent duration radiation patterns (*grey regions, central plots*) which have an azimuthal dependence. *Right plots* present the curves of the apparent duration versus azimuth: *thick black lines* represent an average behaviour ($v_R = 0.5v_P$), while *dashed lines* represent fast ($v_R = v_P$) and slow ($v_R = 0.25v_P$) rupture cases. In the case of a partial unilateral rupture, the curve of apparent duration for the average case (*thick black line*) is given by the maximum of the curves associated to unilateral ruptures along the two segments (*thick grey lines*)



432 v_P are considered). On the other hand, a slower
 433 rupture tends to behave similarly to a spatial point
 434 source model (with a finite time duration); Fig. 2
 435 shows the case of $v_R = 0.25v_P$, a proportion which
 436 is not unrealistic and can be proper for shallow
 437 earthquake in sediment layers (e.g. Selby et al.
 438 2005; Dahm et al. 2007).

439 For a 3D earth model, directivity effects do not
 440 depend only on station azimuth but also on take-
 441 off angles. Equations 1a, 1b and 2 will then depend
 442 on the angle between the rupture direction and
 443 the ray direction at the source, instead than on sta-
 444 tion azimuth. These considerations opened space
 445 for specific detailed inversion approaches, such
 446 as the successful study by Warren and Shearer
 447 (2006), which accounted for take off angles and

required a dense station distribution both in terms 448
 of azimuthal coverage and range of epicentral dis- 449
 tances. We discuss here the problem for the case 450
 of unilateral rupture. A first observation is that 451
 even for the same source-receiver configuration, 452
 different rays travelling along different paths, with 453
 different take-off angles, will present different 454
 directivity effects. For any considered bodywave 455
 with take-off angle θ , we predict theoretically a 456
 minor variability of the apparent duration with 457
 respect to the one modelled for a planar case. For 458
 example, the maximal apparent duration, at the 459
 azimuth opposite to the direction of the rupture 460
 propagation, will be equal to $t_r + t_R + t_P \cos(\theta)$, 461
 which differs from Eq. 1a. Additionally, we have 462
 the superposition of different P wave arrivals, with 463

464 different take-off angles. In any case, the interpre-
 465 tation of the maximal apparent duration following
 466 Eq. 1a and neglecting take-off angles and rise time
 467 effects will lead to an overestimation of the terms
 468 t_R and t_P and consequently an overestimation of
 469 the rupture length. The value derived in such ap-
 470 proximation can be safely considered as an upper
 471 bound to the real rupture length.

472 Our approach here is to model only the az-
 473 imuthal variation of directivity effects. The ap-
 474 proximation is limited by different conditions.
 475 First, we will focus on shallow earthquakes and
 476 use only regional distances seismograms (epicen-
 477 tral distances in the range 200–1,000 km), basing
 478 our inversion on the fit of time windows centred at
 479 the first P wave arrivals. Additionally, we will con-
 480 sider only the case of horizontal to sub-horizontal
 481 ruptures (rupture propagating along direction dip-
 482 ping at most 20°). In these circumstances, we will
 483 show that the main rupture propagation can be
 484 detected even with the approximated approach
 485 here described. The interpretation of additional
 486 rupture parameters is strongly limited by neglect-
 487 ing take-off angles and rise time effects, and the
 488 derived rupture length should be considered as an
 489 upper bound to its real value.

490 On the base of the previous discussion, we can
 491 now explain the last step of the inversion ap-
 492 proach, where the distribution of apparent source
 493 time durations is interpreted in terms of simplified
 494 rupture models:

- 495 1. Point source model. If we assume a spatial
 496 point source model, with a given source time
 497 function of duration t_r , the curve of appar-
 498 ent duration will be a straight line, with no
 499 azimuthal dependence. This is the particular
 500 case, which can also be described by the pre-
 501 vious equations, with $L = 0$, which leads to
 502 $t_R = t_P = 0$. The model has a unique unknown,
 503 t_r (by definition, rise time), which represents
 504 the apparent rupture time everywhere.
- 505 2. Pure/predominant unilateral rupture. In this
 506 case, the azimuthal distribution of apparent
 507 durations is expected to follow a sinusoidal be-
 508 haviour. The fit is expected to be larger in case
 509 of a pure unilateral rupture, but the model can
 510 still well reproduce data also for asymmetric
 511 bilateral ruptures. In both cases, the minimum

of the curve will indicate the main rupture 512
 direction (φ_0). The model is represented by 513
 the curve of the apparent duration $\Delta t(\varphi)$: 514

$$\Delta t(\varphi) = -A \cos(\varphi - \varphi_0) + B \quad (5)$$

with A and B positive, describing in first approx- 515
 imation the travel time of P waves along the rup- 516
 ture length (t_P) and the sum of rise and rupture 517
 times ($t_r + t_R$). 518

3. Pure bilateral rupture. In case of a pure bilat- 519
 eral rupture, we will use the following curve, 520
 instead: 521

$$\Delta t(\varphi) = A |\cos(\varphi - \varphi_0)| + B. \quad (6)$$

4. Comparison of directivity models. Since 522
 Eqs. 5 and 6 are dependent on three unknown 523
 parameters, while the standard average rup- 524
 ture time has only one (the average earth- 525
 quake duration or duration of the common 526
 source time function), they will always pro- 527
 vide a better fit with respect to the common 528
 rupture duration model. An F test can than 529
 be used to evaluate the misfit improvement 530
 versus the increase of degrees of freedom. F 531
 values above 0.5 will be used here to prefer 532
 unilateral or bilateral models with respect to 533
 the point source solution. The modelling of 534
 asymmetric bilateral rupture requires more 535
 free parameters, being the superposition of 536
 two functions as in Eq. 5. Chances for the im- 537
 plementation of such a model are highly lim- 538
 ited by several factors, including data quality, 539
 focal mechanism, epicentral azimuthal cover- 540
 age and local structural heterogeneities. We 541
 suggest here the adoption of this model only 542
 for specific cases rather than its implementa- 543
 tion within the automated processing. 544

The major advantages of the presented approach 545
 are that only point source synthetic seismograms 546
 are used, so that forward modelling and inversion 547
 are fast. A second advantage lies in the intrinsic 548
 simplicity of this approach, which provides simple 549
 plots of easy interpretation and suggest its imple- 550
 mentation for automated routines. In the worse 551
 case, if no clear pattern is detected, the retrieved 552
 information can still be used to better focus a 553
 more detailed full waveform kinematic inversion 554

(e.g., as in Cesca et al. 2010), by limiting the range of rupture times and/or spatial extension to be tested. A last important advantage resides in the coherency of the inversion approach: we use the same tools and the same data to derive first the focal mechanism (or moment tensor) and to detect directivity effects, instead of relying on an externally calculated focal mechanism solution, which may be biased by a specific selection of stations, earth model, and processing routines. An important limitation of the method applicability is represented by a specific range of earthquake magnitudes. For small earthquakes, with the rupture process occurring in less than 2 s, the variation of the apparent rupture duration could be only detected by fitting high frequency spectra (up to 0.5 Hz or above), which requires a detailed knowledge of the crustal structure, and possibly the adoption of 3D earth models. On the opposite side, large earthquakes may present more complex rupture processes, with different slip patches or asperities, breaking at different times. Large earthquakes also present significant discrepancies between hypocentral and centroid locations. Both effects are not considered in our model and can then lead to erroneous interpretations. As a rule of thumb, we believe that earthquakes with magnitudes in a range M_w 5.5–7.0 may be the best suited for a successful application.

3 Synthetic tests for linear and planar sources

Before applying the method to real data, and in order to assess the method performance, we carry out a set of inversions using synthetic datasets. We generate synthetic seismograms first for linear sources and consequently to planar ones. Considered source models present different significant focal mechanisms and directions of rupture propagation. Both synthetics for linear and planar sources are generated using the Kiwi tools (<http://kinherd.org>; Heimann 2010), assuming the PREM model (Dziewonski and Anderson 1981). The inversion is then carried out as described in the previous paragraph, assuming a spatial point source. Focal mechanisms, scalar moment and depth are retrieved at a first stage, by using the method described in Cesca et al. (2010), which has

been here specifically modified in order to include the retrieval of the apparent rupture time at each station. In the synthetic tests, a dense grid of 154 stations is considered, in order to plot apparent duration contours. Stations location accomplish to the chosen conditions in terms of epicentral distances.

Three line source mechanisms (strike-slip (SS) strike $\varphi = 30$, dip $\delta = 90$, rake $\lambda = 0$; normal fault (NF) $\varphi = 30$, $\delta = 45$, $\lambda = 90$; thrust fault (TF) $\varphi = 30$, $\delta = 20$, $\lambda = -90$) and three rupturing models (pure unilateral, asymmetric bilateral, pure bilateral) are considered at first, thus providing a set of nine source models. Ruptures propagate horizontally for the strike-slip and normal fault, in direction NNE–SSW, and toward ESE along the low-angle dipping plane, for the thrust fault. The source model centroid is always located at a depth of 20 km, the source length is 30 km, rupture velocity is 3.5 km/s. Pure unilateral ruptures start at the southern (SS, NF) or western (TF) edge. The same main rupture direction is used for asymmetric bilateral ruptures, the two ruptured segments having lengths $L_1 = 22.5$ and $L_2 = 7.5$ km. Rise time is fixed to 2 s in all cases. The inversion is carried out using 40 s time windows, starting 10 s before the theoretical arrival of P phases. A frequency bandpass, between 0.01 and 0.5 Hz, is used to filter Green's functions and data. Inversion results are shown in Fig. 3. Coloured surface, representing the inverted apparent source duration, highlight the radiation patterns of apparent duration. The azimuthal distribution of apparent duration clearly shows directivity effect and its minor dependence on epicentral distance using our approach and proof that a good quality fit can be achieved using the simplified azimuthal dependent curves. In particular, the fit of the cosine curve described by Eq. 5 can be used to detect unilateral or asymmetric bilateral ruptures, while the curve from Eq. 6 to detect pure bilateral ruptures. Rupture directivity is correctly detected in all cases, with the exception of the bilateral rupture along the low-dipping angle plane of the thrust mechanism (Fig. 3, bottom right), where unilateral rupture is incorrectly estimated. The reason for this discrepancy can be described as follows. Since the extended source is not horizontal, the segment toward ESE and WNW are located below

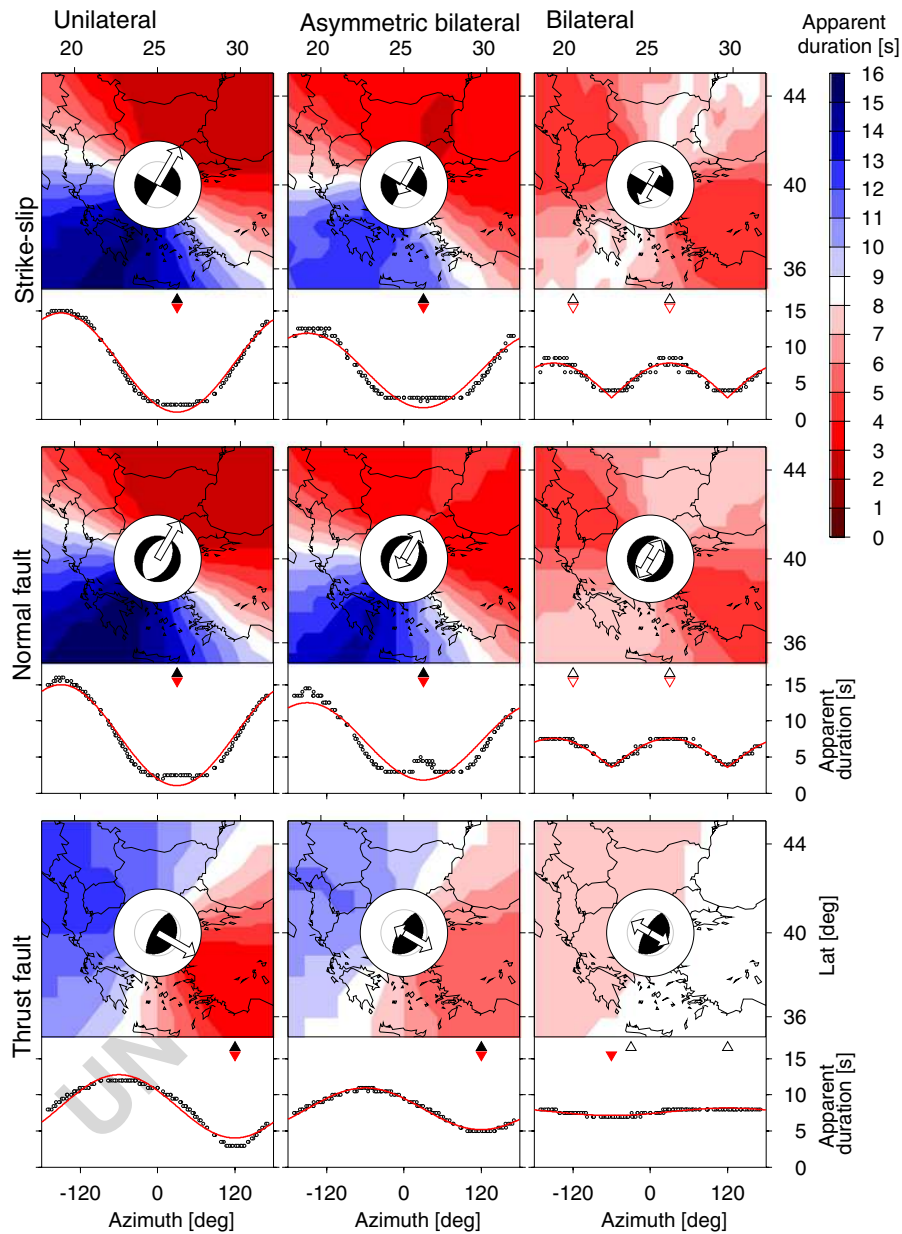


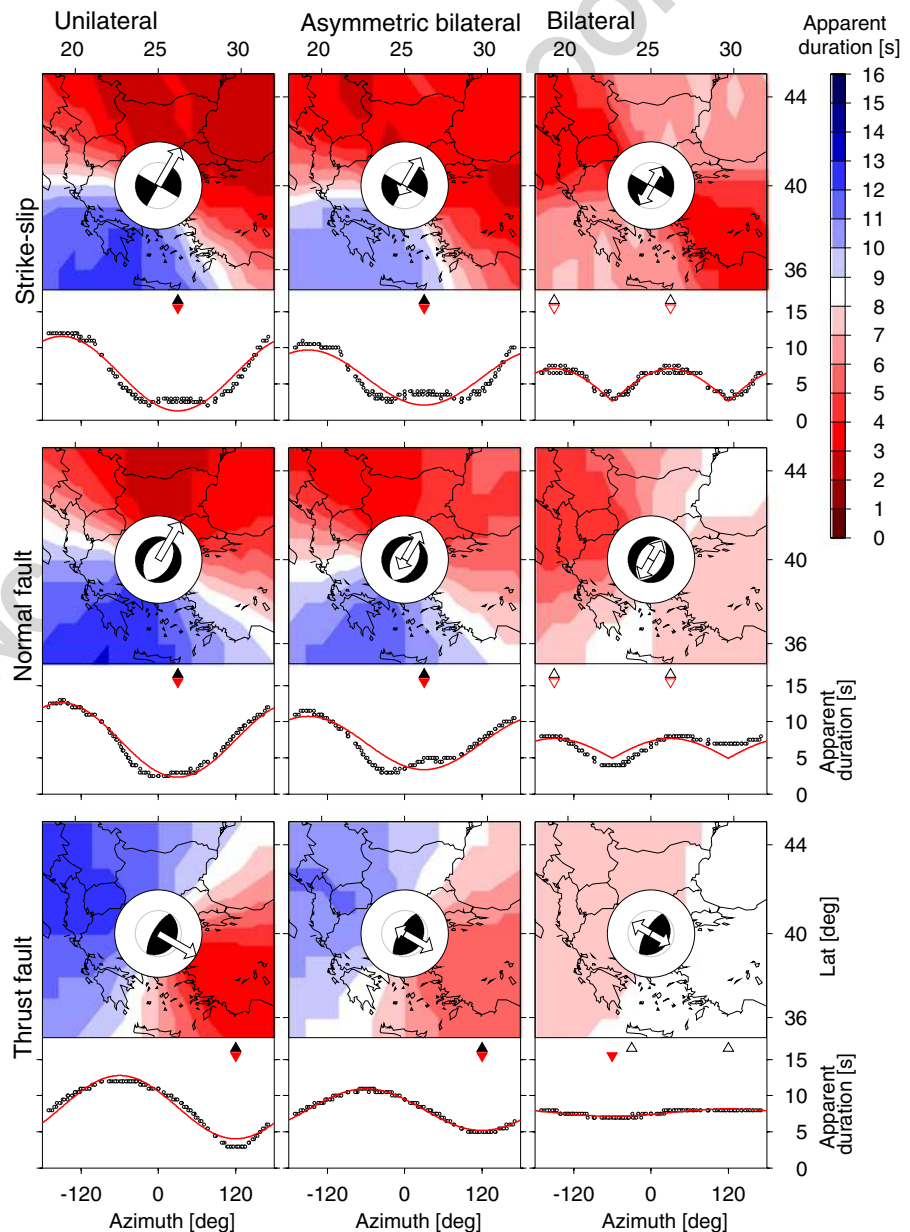
Fig. 3 Inversion results for linear sources. We consider three focal mechanisms (strike-slip, *top*; normal fault, *centre*; thrust fault, *bottom*) and three rupture processes (pure unilateral, *left*; asymmetric bilateral, *centre*; pure bilateral, *right*). For each case, we show colour plots representing the inverted apparent duration at a dense grid of station around the epicentre (red to blue scale represents increasingly longer apparent durations). The focal mechanisms are shown at the epicentral location, together with *white arrows* describing rupture directions (*arrow*

sizes are proportional to rupture lengths). *Graphics below each colour plot* represent the apparent duration versus azimuth (*dots*) and the best fitting model (*red curves*). *Upper triangles* represent the correct solution (a *single black triangle* is plotted at the proper azimuth for unilateral and asymmetric bilateral ruptures; *two white triangles* indicate rupture directions for pure bilateral rupture cases). *Inverted triangles* represent inversion results (*single red triangle* for unilateral and asymmetric bilateral ruptures, *two white triangles* for pure bilateral ruptures)

650 and above the centroid depth, respectively. Since
 651 we use a layered model, the frequency content
 652 of synthetic seismogram varies from shallower to
 653 deeper sources, as well as take off angles. These
 654 effects, which cannot be reproduced by a point
 655 source located at the centroid depth, results larger
 656 than those related to the bilateral rupture, thus
 657 explaining the detection of an apparent directivity
 658 towards dip direction.

Given the successful application to line sources, 659
 we simulate now more realistic rupture processes, 660
 generating synthetic seismograms for the eikonal 661
 source model (Heimann 2010; Cesca et al. 2010). 662
 We use here circular faults, with rupture propagat- 663
 ing with a variable velocity, scaling by a coefficient 664
 0.9 with shear wave velocity in the crustal model. 665
 Given the adoption of the PREM model, the 666
 source depth and its extension, rupture velocity 667

Fig. 4 Inversion results for circular eikonal sources. For each of the nine considered source models, we show colour plots representing the inverted apparent duration around the epicentre (*red to blue scale* represents increasingly longer apparent durations). Graphics below each coloured plot represent the apparent duration versus azimuth (*dots*) and the best fitting model (*red curves*). We use the same symbol convention as in Fig. 3



668 range between 2.9 and 4.0 km/s. Inversion is carried
669 out using the same approach and parameters
670 as for the previous test and results summarized in
671 Fig. 4. The main characteristics of the azimuthal
672 patterns of the apparent rupture duration are pre-
673 served. Directivity effects can be detected and
674 modelled for all pure unilateral and asymmetric
675 bilateral ruptures, although the fit quality results
676 in some cases significantly poorer than for line
677 source cases. The adoption of bi-dimensional rup-
678 tures, and specifically the inclusion of sources at
679 different depths, slightly modifies the apparent
680 duration pattern, as can be seen by the compari-
681 son of plots in Figs. 3 and 4 relative to normal fault
682 mechanism. The modification of the apparent
683 rupture radiation pattern has similar causes than
684 those described for linear sources. Additionally,
685 it is here also depending on the variable rupture
686 velocity (which, according to the crustal model,
687 is faster for deeper sources than for shallower
688 ones). These effects result critical for the inversion
689 of a bilateral rupture for the case of a thrust
690 fault mechanism, which is erroneously interpreted
691 as unilateral (with a rupture propagation in dip
692 direction). On the other hand, the discrimination
693 between pure and asymmetric bilateral rupture is
694 not always possible; the observation of the charac-
695 teristic lobe associated to the asymmetric bilateral
696 rupture may indicate such a rupture process, but
697 could not be sufficient, alone, to distinguish this
698 case to a pure unilateral rupture.

699 A variation of the scalar moment will lead to a
700 scaling of synthetic seismograms and modify their
701 amplitude spectra. As a consequence, the inver-
702 sion of apparent durations may lead to slightly
703 different results. In order to investigate these
704 effects, we have perturbed the scalar moments
705 used for synthetic tests, and analyse inversion
706 results. While the uni- or bilateral mode of the
707 rupture and the main rupture direction are not
708 influenced by a variation of the scalar moment and
709 are always correctly retrieved, the rupture time
710 and the following estimation of rupture lengths
711 suffer slight changes. A perturbation of 10% of
712 the correct scalar moment always led to uncer-
713 tainties below 5% in terms of rupture time and
714 rupture length. Synthetic tests suggest the imple-
715 mentation of the method here proposed towards a
716 rapid detection of directivity effects. Additionally,

717 these tests point out specific cases, where the in-
718 version approach results more critical, and where
719 a careful discussion of results is suggested. In gen-
720 eral, horizontal ruptures and pure or partially uni-
721 lateral ruptures are more easily detected, whereas
722 pure bilateral sources and rupture propagating
723 along dipping directions may be more problematic
724 to resolve.

4 The Andravida 8.6.2008 earthquake

725 On June 8th, 2008, a magnitude Mw 6.4 726
727 earthquake struck NW Peloponnese, Greece. 728
729 The earthquake, here further referred as the 730
731 Andravida earthquake, produced two casual- 732
733 ties, about 100 injuries and several damages 734
735 (Chouliaras 2009). A wide number of studies cov- 736
737 ered the earthquake source and its effects. A pure 738
739 strike-slip focal mechanisms was unanimously 740
741 provided by several institutions and catalogues 742
743 surveying regional and global seismicity, includ- 744
745 ing National Observatory in Athens (NOA), 746
747 Aristotle University of Thessaloniki (AUTH), 748
749 INGV European-Mediterranean RCMT Cata- 750
751 logue (INGV-RCMT), Swiss Federal Institute 752
753 of Technology, United States Geological Sur- 754
755 vey (USGS) and Global CMT Catalogue (CMT). 756
757 According to these models, fault planes are al- 758
759 most vertical and oriented NNE–SSW and WNW– 760
761 ESE. Source depth estimations showed some 762
763 variability, ranging between 10 and 38 km, and 764
765 magnitudes Mw ranged between 6.3 and 6.5. 766
767 The epicentral locations of the earthquake after- 768
769 shocks (Ganas et al. 2009; Gallovic et al. 2009; 770
771 Kostantinou et al. 2009), which are distributed 772
773 within a narrow strip extending NNE–SSW, pro- 774
775 vide a convincing image of the rupture orienta- 776
777 tion. The cloud of aftershocks elongates for about 778
779 30–35 km, providing a first rough estimation of 780
781 rupture size. The aftershock distribution is denser 782
783 towards the Northern edge. To the south, epi- 784
785 central locations may indicate a minor bending 786
787 of the rupture area to a slightly larger strike. 788
789 All published source models are consistent with 790
791 the identification of the NNE–SSW striking fault 792
793 plane. Sokos et al. (2008), using hypocentral- 794
795 centroid relative location method, identified the 796
797 same plane; since centroids locations are generally 798

763 located North of hypocentral locations, some indi-
 764 cation for a dominant propagation towards North
 765 may arise from this study. Even more convincing,
 766 with respect to the detection of directivity, are
 767 the studies of Kostantinou et al. (2009), Gallovic
 768 et al. (2009) and Cesca et al. (2010). The first
 769 authors derived a finite source model, also con-
 770 sistent with their aftershock relocations, finding
 771 a rupture length of 22.4 km and an asymmetric
 772 bilateral rupture, with a major rupture along the
 773 NE branch. Gallovic et al. (2009) used a conjugate
 774 gradient method to detect the spatio-temporal
 775 evolution of the rupture process. Results indicate
 776 a predominantly unilateral rupture, propagating
 777 along a main slip patch, with a rupture length
 778 of about 20 km and a rupture velocity of about
 779 3 km/s. Cesca et al. (2010), based on amplitude
 780 spectra inversion of full waveform and assuming

an eikonal source model, detected an asymmetric 781
 rupture propagation, with a predominance of 782
 rupture propagation towards NNE; the rupture 783
 length was estimated 40 km, while the average 784
 rupture velocity was fixed to about 3.2 km/s. The 785
 consistency of these results (see Fig. 5 bottom), 786
 obtained with different methods and datasets, 787
 offer a serious reference to our study in terms of 788
 fault plane identification and rupture directivity. 789

In our inversion, we assume the epicentral loca- 790
 tion provided on the EMSC-CSEM webpage 791
 and the focal mechanism determined by Cesca 792
 et al. (2010), which is based on the fit of full 793
 waveform amplitude spectra. We observed that 794
 a new estimation of the scalar moment, based 795
 on the fit of P wave spectra only, would present 796
 minor variation with respect to the assumed value 797
 (5.97×10^{18} instead of 6.07×10^{18} Nm). Then, we invert 798

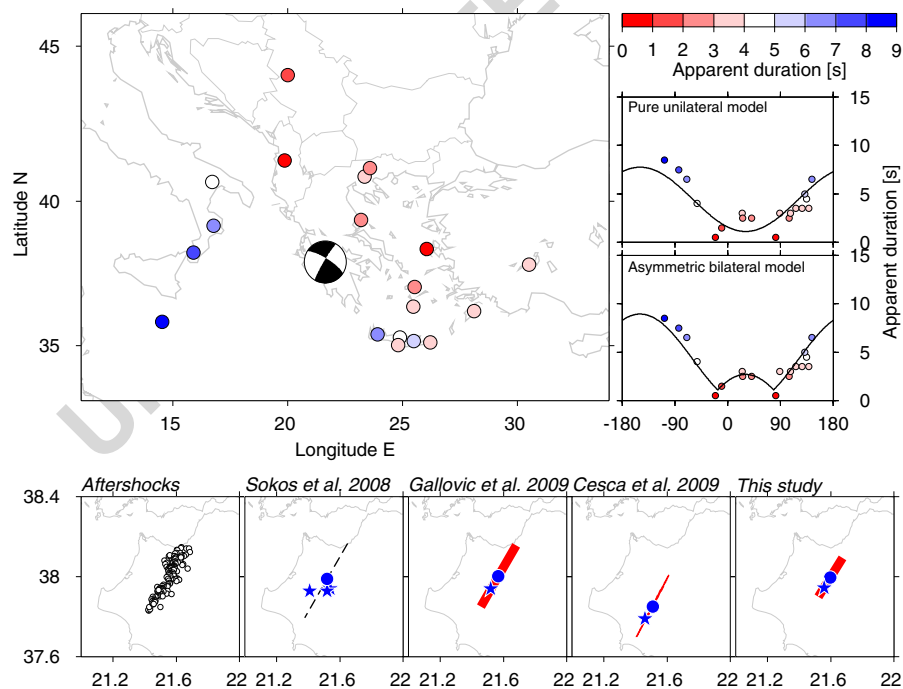


Fig. 5 Inversion results for the Mw 6.4 Andravida (NW Peloponnese) earthquake (*top*) and comparison with published source models (*bottom*). *Top*: coloured dots represent the inverted apparent duration at the stations used, according to the given colour scale; the azimuthal distribution of apparent durations may be fitted (*top right*) assuming a pure unilateral or, better, a partially unilateral rupture (*thick lines*). *Bottom*: the comparison of aftershocks distribution (after Gallovic et al. 2009) identifies

the NNE–SSW rupture plane in agreement with centroid-hypocentral technique (Sokos et al. 2008), adjoint method (Gallovic et al. 2009), full waveform kinematic inversion (Cesca et al. 2010) and our results; the last four methods consistently detect a partially unilateral rupture towards NNE. Stars and blue circles represent here nucleation points and centroids respectively; the rupture area is plotted in red

799 for the apparent duration at each station sep- 848
 800 arately and plot resulting values in function of 849
 801 station azimuth (Fig. 5), according to the discussed 850
 802 methodology. As a first approximation, we try to 851
 803 fit apparent durations by means of pure unilateral 852
 804 and pure bilateral rupture models, assuming both 853
 805 fault planes. Since fault planes are almost vertical, 854
 806 only horizontal directions of the rupture velocity 855
 807 along these planes are considered. The unilateral 856
 808 rupture model with rupture propagation towards 857
 809 NNE provide a very good fit to the apparent du- 858
 810 ration data, and is preferred to remaining models 859
 811 on the base of the F test. This result provides a 860
 812 clear indication for a rupture propagating towards 861
 813 NNE, and thus can be used to discriminate the 862
 814 true fault plane (NNE–SSW) from the auxiliary
 815 one (WNW–ESE). Based on the good fit, we try
 816 to refine our solution by investigating asymmetric
 817 bilateral ruptures. Results provide an even more
 818 convincing fit (Fig. 5, bottom right), when an
 819 asymmetric bilateral source model with rupture
 820 propagating mostly Northward is assumed, as the
 821 curve account for the two symmetric minima at
 822 about -21.5 and 82.5° and the internal charac-
 823 teristic lobe of asymmetric rupture (see Fig. 2,
 824 bottom). On the other side, this result is in very
 825 good agreement with published models discussed
 826 before. According to the previous discussion for
 827 a simplified bidimensional case, the maxima of
 828 the two cosine curves associated to the rupturing
 829 of two segments of the fault are equal to $t_P +$
 830 $t_R + t_r$, where these terms refer to the P wave
 831 propagation, rupture and rise time related to each
 832 segment. The maxima of the apparent duration
 833 curve are equal to 8.9 s (segment L_1 toward NNE)
 834 and 2.7 s (segment L_2 toward SSW). Assuming
 835 an average P wave velocity of 8 km/s (consistent
 836 with the used velocity model at the hypocentral
 837 depth), a rupture velocity of 3 km/s (consistent
 838 with Gallovic et al. 2009), and considering the rise
 839 time negligible with respect to rupture time, we
 840 obtain rupture lengths of about 19 and 6 km. The
 841 total length of about 25 km for the main patch
 842 is in general agreement with most of published
 843 results. We observe a discrepancy with the rupture
 844 size of about 40 km determined in Cesca et al.
 845 (2010). This last value might be overestimated,
 846 as the adopted full waveform kinematic inversion
 847 may in some cases be affected by a trade-off be-

tween different source parameters. The observed 848
 discrepancy may also indicate a different response 849
 of the two approaches to the rupture process and 850
 energy emission, with the full waveform inver- 851
 sion detecting the largest rupture length, and the 852
 directivity inversion identifying the main rupture 853
 patch. The upper limit value we found here is 854
 slightly larger than the length estimated by stan- 855
 dard empirical relations (according to Wells and 856
 Coppersmith 1994, the average rupture length for 857
 a Mw 6.4 is about 14 km). We remark that the 858
 interpretation of inversion results to this extent 859
 should be carried out only in best conditions, 860
 where the fitting of the apparent duration curve 861
 is good enough to further interpret it. 862

5 The SW Peloponnese 14–20.2.2008 seismic 863 sequence 864

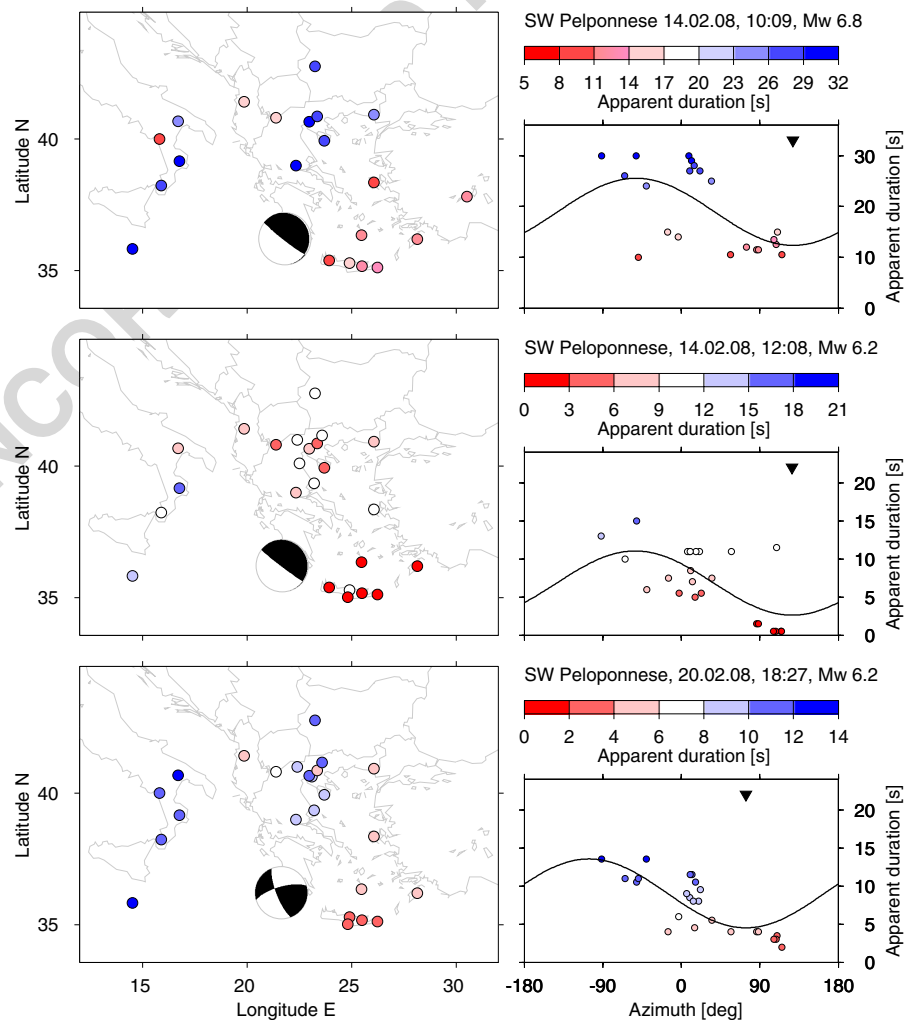
A seismic sequence struck the region offshore 865
 SW Peloponnese, Greece, in the days following 866
 February 14th, 2008, with three major earth- 867
 quakes occurring within a week. On February 868
 14th, a first Mw 6.8 (magnitude estimated by 869
 EMSC-CSEM and Cesca et al. 2010) event struck 870
 the region at 10:09 UTC. Two hours later, at 871
 12:08, UTC, a Mw 6.2 aftershock occurred. Fi- 872
 nally, on February 20th (18:27 UTC), a Mw 6.2 873
 event took place. Focal mechanisms (EMSC- 874
 CSEM webpage) indicate thrust faulting for the 875
 first two events, while the last one has a different, 876
 strike-slip mechanism, with fault planes striking 877
 ENE and NNW. Source depths, according to 878
 EMSC-CSEM catalogue, were 30, 20 and 25 km, 879
 respectively, for the three earthquakes. Whereas 880
 different institutions (e.g. NOA, AUTH, Uni- 881
 versity of Patras UPSL, INGV-RCMT, USGS, 882
 CMT) provided point source solutions for these 883
 events, few trials has been carried out so far to 884
 interpret rupture kinematics (Roumelioti et al. 885
 2009; Cesca et al. 2010). A strongly uneven sta- 886
 tion distribution and large epicentral gaps toward 887
 SW have possibly limited source modelling until 888
 now. Based on full waveform inversion, Cesca 889
 et al. (2010) identified the ENE–WSW plane for 890
 the strike-slip earthquake of February 20th, and 891
 a partial unilateral rupture towards the coast was 892
 found. For the first two earthquakes, the low angle 893

894 planes dipping toward NW were preferred, but the
 895 inversion results were not completely satisfacto-
 896 rily. Roumelioti et al. (2009) adopted an empirical
 897 Green's functions approach, using the Mw 6.2
 898 aftershock to model the finite fault of the largest
 899 earthquake; their results support the identification
 900 of the low angle dipping plane as well as directivity
 901 towards SSW. Finally, the identification of low
 902 dip angle rupture planes in this region for thrust
 903 earthquakes would agree with local tectonics, as-
 904 sociating the earthquake occurrence to oceanic
 905 subduction (Underhill 1999).

906 Figure 6 summarizes our inversion results.
 907 Differently from the application to the Andravida
 908 earthquake, apparent durations for the earth-

909 quakes occurring on February 14th show a major
 910 spreading and are worse fitted by the simplified
 911 cosine function we associated to pure unilateral
 912 ruptures. A better fit, and minor spreading, is
 913 observed for the main event, with respect to its
 914 aftershock. However, for both two earthquakes, a
 915 general trend can be detected, indicating a mini-
 916 mum of the cosine curve for an azimuth of about
 917 135°, which suggests a main direction of the rup-
 918 ture propagation towards SE. These results are
 919 unable by themselves to provide further informa-
 920 tions about the true fault plane, as they may be
 921 modelled either assuming the low-angle and the
 922 steep dipping plane. A better fit is obtained for the
 923 February 20th earthquake: clear directivity effect

Fig. 6 Inversion results for the February 14th, 2008, Mw 6.8 (top), the February 14th, 2008, Mw 6.2 (centre) and the February 20th, 2005, Mw 6.2 (bottom) earthquakes, offshore SW Peloponnese. For each earthquake, coloured dots represent the inverted apparent duration at the stations used, according to the colour scale given for each case; the azimuthal distribution of apparent durations (right) may be fitted assuming unilateral ruptures (thick black lines). Inverted triangles indicate the retrieved rupture directions



924 is here retrieved, indicating a rupture mostly prop-
 925 agating towards ENE, thus along the WSW–ENE
 926 fault plane. The differentiation between pure or
 927 partial unilateral rupture may be here rewarded
 928 as beneath the limit of a safe data interpretation,
 929 but we observed how the general result is in well
 930 agreement with previous results by Cesca et al.
 931 (2010), where the source model was derived by
 932 the fit of high-frequency (up to 0.1 Hz) amplitude
 933 spectra from the whole waveforms.

934 Finally, we investigate effects of the assumption
 935 of imprecise point source parameters on the esti-
 936 mation of rupture directivity. The effects of anom-
 937 alous source depth estimation are studied for the
 938 February 20th aftershock, as different Institutions
 939 have provided a range of different values ranging
 940 from 8 to 25 km. We repeated the inversion using
 941 a different point source solution (strike 249°, dip
 942 88°, rake −12°, depth 12 km), as provided by the
 943 INGV European-Mediterranean RCMT Catalog;
 944 this focal mechanism is similar to our solution and
 945 major differences concern centroid depth, which is
 946 now shallower. Directivity inversion remains very
 947 stable, showing a consistent identification of uni-
 948 lateral rupture direction toward SE, and indicates
 949 that a source depth variation of about 10km does
 950 not result in any significant variation in the radi-
 951 ation pattern of apparent duration. In a similar way,
 952 we tested slightly different focal mechanisms for
 953 both earthquakes of February 14th: even if our fo-
 954 cal mechanisms are in relatively good agreement
 955 with other published solutions, some difference
 956 can be observed. For example, the INGV-RCMT
 957 catalogue indicates strike angles of 333° and 298°,
 958 for these earthquakes, which differ from our so-
 959 lution (347° and 341° respectively). Even in this
 960 case, after adopting the source parameters pro-
 961 vided by INGV-RCMT, the inversion results are
 962 stable, with the detection of main rupture direc-
 963 tions pointing towards SE–ESE.

964 6 Conclusions

965 We propose here a new method for a quick detec-
 966 tion of directivity effects for shallow earthquakes
 967 at regional distances. Among the most important

features of the method, we highlight here the 968
 following ones: 969

- Rapid inversion 970

The assumption of spatial point source allows 971
 an extremely rapid generation of synthetic seis- 972
 mograms and point source parameters inversion, 973
 thus offering a tool to early detect directivity; to 974
 quantify such improvement, on a standard single 975
 processor PC the inversion of directivity is here 976
 carried out within a minute, about 20 times faster 977
 than the full kinematic inversion for the same 978
 event, using the approach described in Cesca et al. 979
 (2010). 980

- Coherent inversion 981

The use of the Kiwi tools for data processing 982
 and inversions improves significantly the consis- 983
 tency of our methodology: for example, the same 984
 dataset and the same inversion tools can be used 985
 to first derive the focal mechanism, then the scalar 986
 moment and finally the apparent duration; we 987
 believe this consistency between data used for 988
 different inversions significantly improve the co- 989
 herency of the inversion approach. 990

- Accounting for wave propagation 991

The method is based on amplitude spectra inver- 992
 sion, using theoretical Green's functions for the 993
 chosen earth model. In this way, we account for 994
 wave propagation effect, an improvement with 995
 respect to standard methods based on pulse length 996
 estimations. 997

- No requirements of specific aftershocks 998

Avoiding the use of empirical Green's function, 999
 the method is not limited by the existence of 1000
 a proper aftershock, nor need to wait for its 1001
 occurrence. 1002

- Automation 1003

The adoption of the Kiwi tools and the simplicity 1004
 of the inversion approach made possible the im- 1005
 plementation of the method as automated routine. 1006

In this manuscript we have demonstrated the 1007
 method performance, both with a range of syn- 1008
 thetic tests and with observed data for different 1009

1010 shallow earthquakes recently occurred. These ap-
 1011 plications offer indications about the quality and
 1012 extent of inversion results. The retrieval of pure
 1013 unilateral and pure bilateral ruptures is in gen-
 1014 eral better resolved than asymmetric ruptures,
 1015 although the application to the June 8th, 2008,
 1016 Andravida earthquake showed that this case can
 1017 also be detected, in favourable conditions. In
 1018 general, directivity effects are better resolved for
 1019 strike slip earthquakes, with respect to normal or
 1020 thrust faulting. Directivity detection offers often a
 1021 chance to identify the rupture plane, discriminat-
 1022 ing it from the auxiliary one. The determination
 1023 of rupture time, rise time and rupture velocity on
 1024 the base of the proposed method is beyond its
 1025 purposes and should require a careful supervision.
 1026 We have here focused to earthquake with magni-
 1027 tudes of Mw 6 to 7 and shallow hypocentres. The
 1028 extension of this inversion approach for the study
 1029 of other range of magnitudes or deeper sources
 1030 may be investigated in future.

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