

 Open access • Proceedings Article • DOI:10.1117/12.788357

## Single-mode LiNbO<sub>3</sub>-based waveguides for L-band interferometry — [Source link](#)

[Guillermo Martin](#), [Eric Anselm](#), [Thibaut Seure](#), [T. Moulin](#) ...+5 more authors

**Institutions:** [Centre national de la recherche scientifique](#), [Autonomous University of Madrid](#)

**Published on:** 14 Jul 2008 - [Proceedings of SPIE](#) (International Society for Optics and Photonics)

**Topics:** [Interferometry](#), [Lithium niobate](#), [Waveguide \(optics\)](#), [Astronomical interferometer](#) and [Ion beam](#)

Related papers:

- [Technological implementation of Bragg grating reflectors in Ti:LiNbO<sub>3</sub> waveguides by proton exchange](#)
- [Miniaturized beamsplitters realized by X-ray waveguides.](#)
- [Micro-Bragg reflectors in photorefractive 3D waveguides](#)
- [Investigation of a Synchronized Flatop AWG Using a Low Coherence Interferometric Method](#)
- [Fabrication and characterization of photonic devices directly written in glass using femtosecond laser pulses](#)

Share this paper:    

View more about this paper here: <https://typeset.io/papers/single-mode-linbo3-based-waveguides-for-l-band-l9ydacemsf>



**HAL**  
open science

## Single-mode LiNbO<sub>3</sub>-based waveguides for L-band interferometry

G. Martin, E. Anselm, T. Seure, T. Moulin, A. Delboulbé, L. Jocou, N. Bodin-Courjal, O. Caballero, J. Olivares

► **To cite this version:**

G. Martin, E. Anselm, T. Seure, T. Moulin, A. Delboulbé, et al.. Single-mode LiNbO<sub>3</sub>-based waveguides for L-band interferometry. SPIE Astronomical Telescopes + Instrumentation, Jun 2008, Marseille, France. 10.1117/12.788357 . hal-00350043

**HAL Id: hal-00350043**

**<https://hal.archives-ouvertes.fr/hal-00350043>**

Submitted on 30 Apr 2021

**HAL** is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers.

L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.



Distributed under a Creative Commons Attribution| 4.0 International License

# Single-mode LiNbO<sub>3</sub>-based waveguides for L-band interferometry

Guillermo Martin<sup>\*a</sup>, Eric Anselm<sup>a</sup>, Thibaut Seure<sup>a</sup>, Thibaut Moulin<sup>a</sup>, Alain Delboulbé<sup>a</sup>, Laurent Jocu<sup>a</sup>, Nadège Bodin Courjal<sup>b</sup>, Olga Caballero<sup>c</sup>, José Olivares<sup>c</sup>

<sup>a</sup>Laboratoire d'Astrophysique de Grenoble, 414 Rue de la Piscine, BP 53, 38400 Saint Martin d'Hères, France ;

<sup>b</sup>FEMTO-ST, Université de Franche-Comté. 16, route de Gray. 25030 Besançon. France ;

<sup>c</sup>Dept. de Física de Materiales. c/ Francisco Tomas y Valiente, 7. 28049 Cantoblanco. España

## ABSTRACT

We present three different techniques for single-mode waveguide realization in Lithium Niobate at the 3.39 $\mu$ m atmospheric transmission band, named L-band. These methods include Titanium diffusion, Ion Beam Implantation and Photo-inscription. After describing the fabrication process and waveguide characterization, we will present an integrated interferometer based on the Young's double slit experiment. From the recorded interferogram we recover information about the source, namely, its peak emission lines.

**Keywords:** Infrared single mode waveguides, Lithium Niobate, Integrated interferometry, L-band

## 1. INTRODUCTION

Infrared observations give access to cool matter studies, where optical (visible) observation is blocked by the absorption of interstellar dust. In particular, the L-band atmospheric transmission window (3.4-4.1  $\mu$ m) presents promising features for astrophysical studies. Cool matter, from dust grains to giant planets, has a temperature range from 3-1500K that makes its blackbody emission fall into the mid-infrared range. As an example, at 800K (temperature of crystallization of some silicates) the thermal emission peaks at 3.6microns, which is at the center of the L-band. Moreover, this band is most appropriate for ground-based detection of hot dust in the habitability regions of stars of the main sequence. [1,2]. Thus, infrared observations are fundamental to study low-temperature regions, such as dusty interstellar clouds or the silicate composition of the dust. Programs like Spitzer are devoted to observation on the 3-180 microns range. Part of the work is devoted to study the mid-IR spectral features of silicates, which in the case of hydrated silicates fall into the L-band (emission lines between 3.1 and 4 microns) [3]. Also, infrared emission of polycyclic aromatic hydrocarbons from circumstellar disks present 3.3 $\mu$ m signatures [4].

Concerning exoplanet detection, the 2.5-5  $\mu$ m window has been identified as an adequate band for planet search science [5]. As an example, the spectrum of a hot Jupiter is centered at 3.8 microns, presenting a strong emission while no atmospheric absorption lines degrade the spectrum [6].

These astrophysical science cases show the interest on the development of instruments devoted to the L-band. Currently, some bulk optics projects are emerging in the L-band (ALADDIN project), but nothing in single mode integrated optics. Indeed, silica and classical glasses cannot reach this band, which explains the lack of instruments in this wavelength range. Thus, an instrument developed for this band will be most welcome, but this requires an initial effort in technological developments as we propose in this paper.

In order to address the L-band we have selected Lithium Niobate (LiNbO<sub>3</sub>). Due to its well-known properties in the near-infrared range (commercial electro-optical modulators at 1550 nm) and large transmission window (450nm-5200nm), LiNbO<sub>3</sub> is an ideal substrate for the realization of infrared beam combiners dedicated to interferometric instruments as the second generation VLTI instruments VSI (VLTi Spectro Imager, where multi-telescope recombination is considered

---

\* guillermo.martin@obs.ujf-grenoble.fr; phone +33 4 76 63 55 09; www.laog.obs.ujf-grenoble.fr

using a single device) or Gravity, another VLTI instrument dedicated to galactic centre observation. It provides also an excellent candidate for realizing precursors for nulling interferometers like PEGASE (spacecraft-instrument dedicated to spectrometry of hot jupiters and brown dwarfs between 2.5-5 microns).

LiNbO<sub>3</sub> is widely used today as a reference material for applications in integrated optics in the telecom window (1.55μm), since it allows obtaining single-mode waveguides by different procedures (Ti or Zn diffusion, proton exchange, ion implantation) with low optical losses (<0.2dB/cm). Whereas these methods are well known for waveguide realization below 1.55 μm, only few laboratories in the world have explored the capacities of LiNbO<sub>3</sub> at higher wavelengths. However, feasibility of single mode waveguides in LiNbO<sub>3</sub> by classical proton exchange techniques has been recently demonstrated [7]. In this paper, we present our results for L-band single mode realization by Titanium diffusion and Ion beam implantation, as well as preliminary results on UV photo-inscription. Finally, we will present some preliminary results on an integrated Young interferometer working in the L-band.

## 2. WAVEGUIDE REALIZATION

In order to achieve single mode waveguides in the L-band, different techniques were developed. Two of them are current methods for waveguide realization, but to the best of our knowledge, none of them was pushed so far as to 3.39microns waveguiding. These techniques are Titanium diffusion and Ion Beam Implantation. A third technique, more prospective, is based on the refractive index modification in LiNbO<sub>3</sub> by UV irradiation. Some recent results have shown the possibility of channel waveguide photo-inscription by doubling an Argon laser (488nm→244nm), producing waveguides with good confinement at 1550nm [8,9]. We have used the same technique in the aim of developing single mode waveguides at 3.39μm.

In the following sections, we describe the different techniques and the characterization results of the waveguides obtained from each.

### 2.1 Titanium diffusion

Titanium diffusion is one of the most expanded techniques for waveguide realization in LiNbO<sub>3</sub>. A thin film (200nm) of Titanium is deposited onto the surface of bulk LiNbO<sub>3</sub>, the sample is placed into an oven, heated at 1020°C and Ti ions diffuse inside the crystal. This produces typically smooth refractive index profiles, which can be modeled by an exponential function distribution (see below).

Two single mode LiNbO<sub>3</sub> planar waveguides at 3.39 microns were obtained with this technique at the FEMTO-ST laboratory, at Université de Franche-Comté. The expected diffusion profiles were calculated by the Galerkin method and the experimental profiles were reconstructed from the angular measurement of dark modes using the WKB inverse method. The dark modes (or m-lines) technique [10], adapted for 3.39 microns, was developed at Laboratoire d'Astrophysique de Grenoble (LAOG). As will be discussed in the characterization section, a first measure at 633nm allows visible characterization of guided modes and further index profile reconstruction. Then, we switch to infrared coupling, obtaining the experimental modes at 3.39 microns and comparing with the theoretically expected ones.

As an example, at 633nm, the experimental profile was given by:

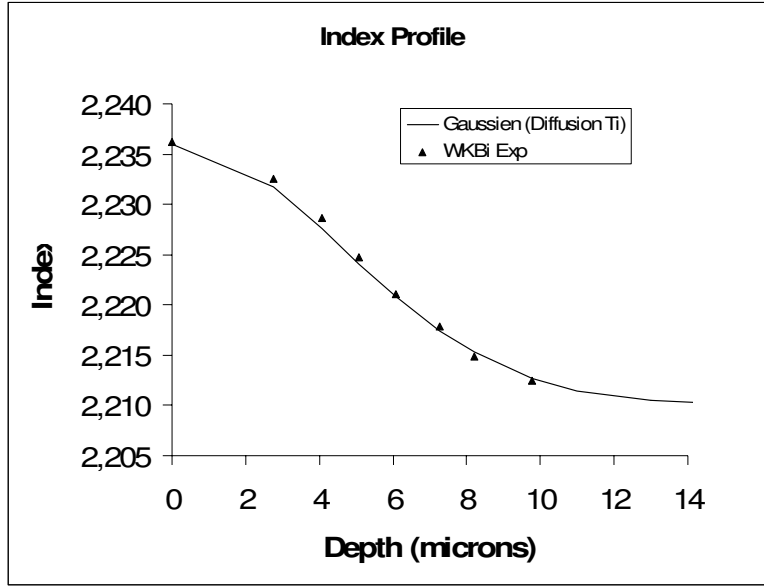


Fig. 1. Refractive index profile reconstructed using the WKB inverse method, for the extraordinary refractive index in a planar waveguide obtained by 50h Ti diffusion over X-cut LiNbO<sub>3</sub> sample.

In order to obtain the characteristics of the Ti-diffused waveguide, we fitted this profile with the expected ion diffusion theoretical formula:

$$n(z) = n_s + \Delta n \cdot e^{-\left(\frac{z}{d}\right)^2} \quad (1)$$

Where  $n_s=2.21$  is the bulk refractive index and  $\Delta n=0.026$  is the birefringence obtained from WKBi method [11]. Fitting the index profile with this formula results in a waveguide depth of  $d=6.5$  microns. Applying the dispersion behaviour of refractive index for LiNbO<sub>3</sub>, we predicted a single mode at 3.39microns, at an output angle of 20.04°. Measurements in the infrared showed indeed a single dark line at 20.7°, that is, an effective index of 2.133 at 3.39µm. This confirms the single mode behavior of the Ti-diffused waveguides obtained for 50h heating at 1020°C.

## 2.2 Ion beam implantation

Ion beam implantation was realized over two different LiNbO<sub>3</sub> samples, one X-cut and one Z-cut. These samples were prepared at the Centro de Microanálisis de Materiales, at the Universidad Autónoma de Madrid using the technique of swift heavy ion-beam irradiation [12]. The planar waveguides were obtained by implantation of 4+ Fluoride ions, at a fluence of  $4 \times 10^{14}$  ions/cm<sup>2</sup> and an energy of 20 MeV.

The experimental measurements of dark modes at 633nm, showed four modes. Using the WKBi method, we reconstructed the experimental refractive index profile in the visible.

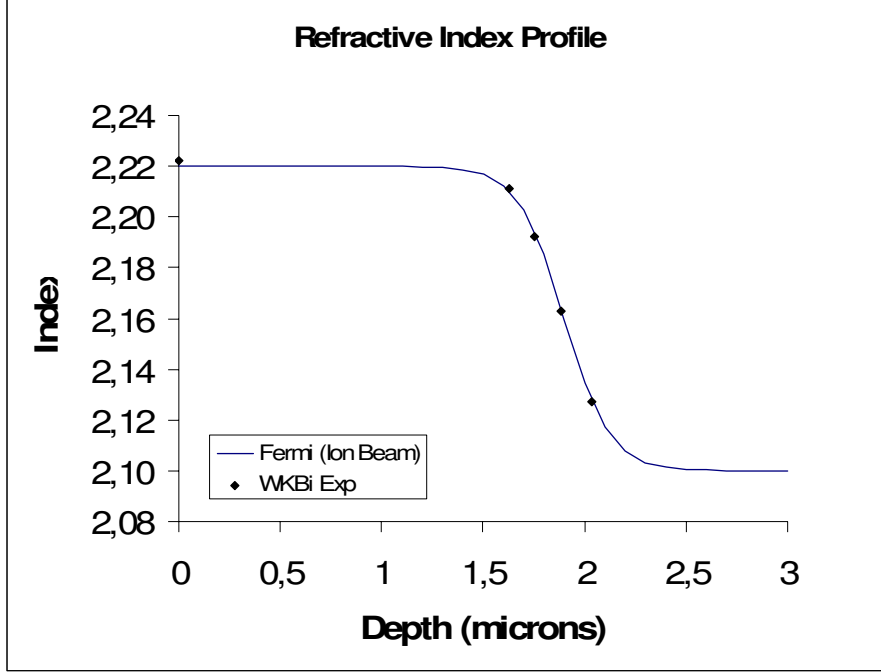


Fig. 2: Extraordinary refractive index profile for the ion-beam irradiated Z-cut sample.

In the case of ion beam implantation, as the profiles are steeper, we used the Fermi theoretical model in order to obtain the waveguide characteristics:

$$n(z) = n_s + \Delta n \cdot \frac{1}{1 + e^{m(z-z_F)}} \quad (2)$$

Where  $n_s$  is the substrate refractive index,  $m$  is the slope of the distribution function,  $\Delta n$  is the induced birefringence and  $z_F$  is the Fermi depth. The two adjustable parameters are  $\Delta n$  and  $z_F$ , while  $m$  is optimized to fit the steepness of the slope to WKBi data. We obtained  $n_s=2.3$ ,  $\Delta n=0.103$ ,  $m=20$  and  $z_F=2.4$ microns for  $n_{ord}=2.4060$  and  $n_s=2.105$ ,  $\Delta n=0.18$ ,  $m=12$  and  $z_F=2.5$  for  $n_{ext}=2.2918$ . These profiles predicted a single mode at  $3.39\mu m$  at  $30.4^\circ$  for  $n_{ext}$  (obtained experimentally at  $27.6^\circ$ ) and  $46^\circ$  for  $n_{ord}$  (obtained experimentally at  $48.9^\circ$ ). The experimental results confirmed the single mode behavior of the waveguides obtained by ion beam implantation with fluoride ions. There is however a discrepancy of about  $3^\circ$  between experimental and theoretical data that must be investigated. Most probably, this difference comes from the behavior of dispersion of the refractive index with wavelength that was not investigated for ion-beam irradiated waveguides.

### 2.3 UV photo-inscription

Inscription of a birefringent area under UV irradiation in  $LiNbO_3$  has been already presented by other authors [8,9], but in the aim of developing waveguides for  $1.55\mu m$ . In this work, we have tried to obtain single mode waveguides at  $3.39$  microns by the same technique. The basic idea of this technique is to use the  $488nm$  emission line of an Argon Laser (40mW CW Lasos60), double the frequency using a non-linear crystal (8mm long BBO crystal from Castech) and focus on the surface of pure  $LiNbO_3$  to obtain ion migration under UV irradiation (the exact mechanisms for photo-induced birefringence are not the scope of this work, and are still not well understood). It's important to note that this effect seems to be only achievable along the extraordinary axis, and not on the ordinary axis. We thus selected an X-cut pure  $LiNbO_3$  substrate and optimized the UV writing parameters: a scanning velocity of 428 mic/s and an incident Argon power of 11mW focalized on the surface with a 4x long working distance microscope objective. The focusing was done before the BBO crystal in order to avoid UV attenuation by the objective lenses.

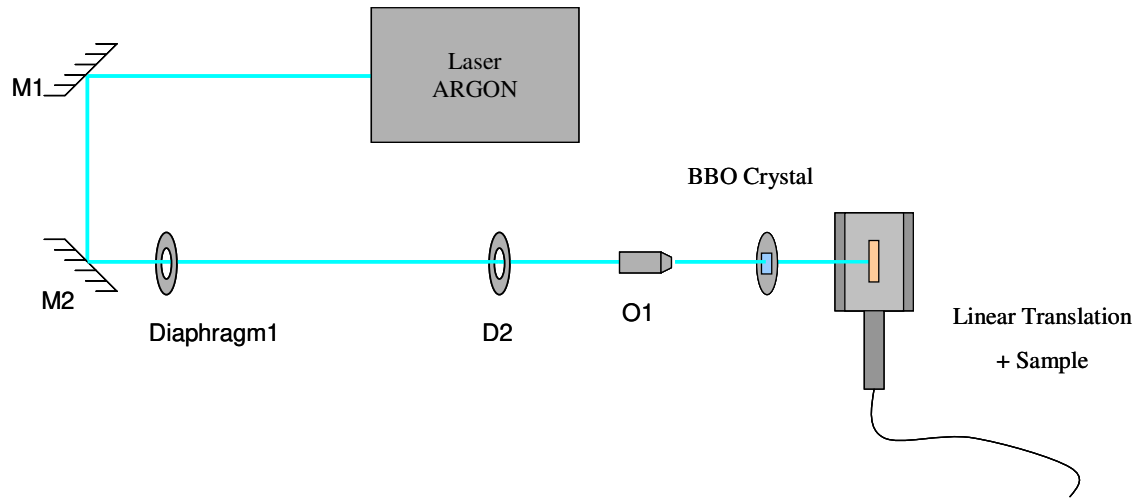


Fig. 3: UV photo-inscription set-up.

### 3. EXPERIMENTAL SET-UP FOR CHARACTERIZATION

#### 3.1 Dark modes set-up

The experimental set-up used for planar waveguide characterization (Ti-diffusion and Ion Beam Implantation) is given in figure 4:

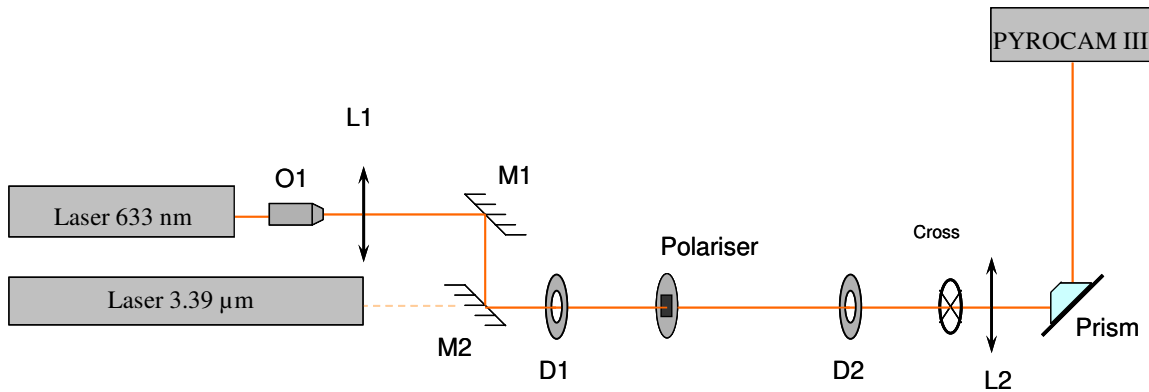


Fig.4: Dark modes characterization set-up for Visible and Infrared co-alignment.

A He-Ne laser, emitting 2mW at 3.39microns is focused onto the surface of the substrate by a ZnSe lens ( $f=75\text{mm}$ ). We used a  $\text{TiO}_2$  prism in contact with  $\text{LiNbO}_3$  waveguides. The output angles corresponding to dark modes are viewed with an infrared camera (Spiricon III). In parallel, a He-Ne 633nm source is co-aligned with the IR laser, using diaphragms D1 and D2, and allowing m-lines measurements in the visible. With these results, we are able to obtain the index profiles for the planar waveguides, for both polarizations, and measure the number of modes at 3.39microns.

### 3.2 Imaging photo-induced channel waveguides

After optimization of the UV writing process, an imaging set-up was designed in order to visualize the output of the channel waveguides and extract information about modal behavior. The optical bench is presented in the following figure:

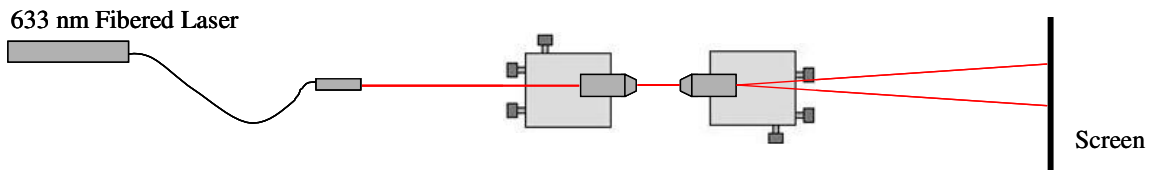


Fig.5 : Optical bench for photo-induced waveguides imaging

A red He-Ne laser is injected by end fire coupling at the interface air-substrate, and scanned along the side of the sample. A gross approach to the irradiated area is done visually, since the diffused photons from end-fire coupling present a darkening near this zone. Then, a fine search of the channel waveguide is done around this area. Eventually, if the irradiation time and the absorbed energy are sufficient to achieve birefringence up to a guiding level, a channel waveguide spot appears when light is efficiently coupled. The experimental observation of such a photo-induced channel waveguide is shown hereunder, in the visible range.

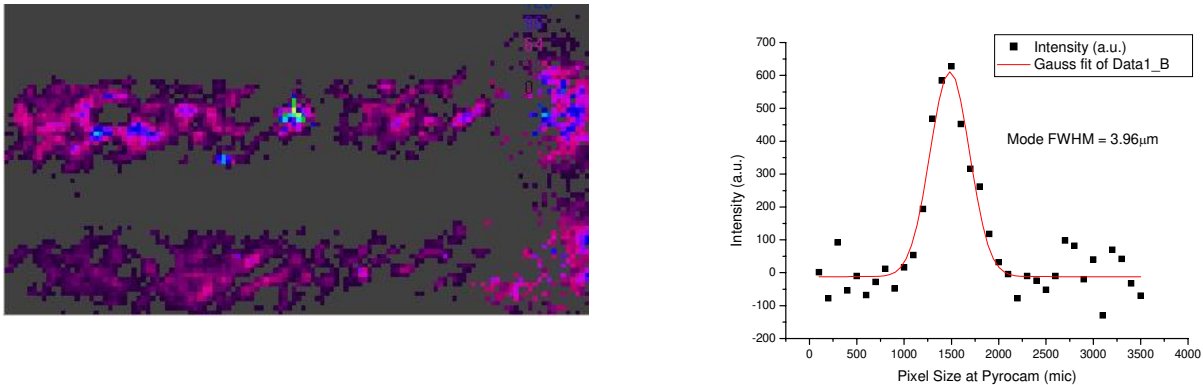


Fig.6: Left: Screen Capture of the Spiricon Camera image of the output of a photo-induced channel waveguide. Right: Transverse plot of intensity and fit to a Gaussian distribution.

By fitting the obtained image of the out-coupling to a Gaussian profile, we obtain a FWHM=3.96microns of the fundamental mode at 633nm. These results confirm that photo-inscription can be used as a method to pattern channel waveguides in  $\text{LiNbO}_3$ ; unfortunately, the birefringence obtained is not sufficient to allow waveguiding at 3.39microns. Indeed, waveguide excitation was done by varying the injection position of the microscope objectives, and no variation of the Gaussian profile at the output was obtained, indicating that only the fundamental mode was guided at 633nm. Thus, there are few chances that 3.39microns wavelength should be guided in the photo-inscribed waveguides presented in this paper, and no experimental results were obtained for this wavelength. New series of UV irradiation are under way in order to increase the induced birefringence and obtain channel waveguides in the L-band.



## 4. INTEGRATED YOUNG'S DOUBLE SLIT INTERFEROMETER

In order to prepare the future integrated optics beam combiners, allowing to achieve complex beam combination schemes [13], a first study was devoted to interference capabilities on single mode LiNbO<sub>3</sub> planar waveguides. The interest in single mode waveguides lies in the fact the high frequency optical aberrations of the injected beam excite high order modes of the waveguide. If the waveguide it's single mode, the optical aberrations are filtered out (as the order above fundamental are radiated) and at the output the distribution intensity of the beam is perfectly Gaussian. Thus, fringe contrast should be improved between multimode and single mode interference patterns obtained in a planar waveguide. In this context, we have developed an integrated optics interferometer based on the Young's Double Slit experiment.

### 4.1 Experimental Set-up

The experimental set-up is shown below. The laser radiation coming from a He-Ne at 3.39 microns is injected directly on two Oxford chalcogenide fibers, transmitting between 1-6 microns. These fibers are set close together by drilling a conic hole in a cylindrical support where the smaller aperture is close to two times the cladding diameter of the fibers (ca. 700microns). At the other end the fibers are set on a V-groove type device, with an inter-core separation of 800microns. Then the LiNbO<sub>3</sub> sample, containing the planar waveguide is gently pressed against the fibers out-put. Finally, a ZnSe lens (f=33.5cm) images the waveguide output on a infrared camera developed at LAOG (COMIC [14]).

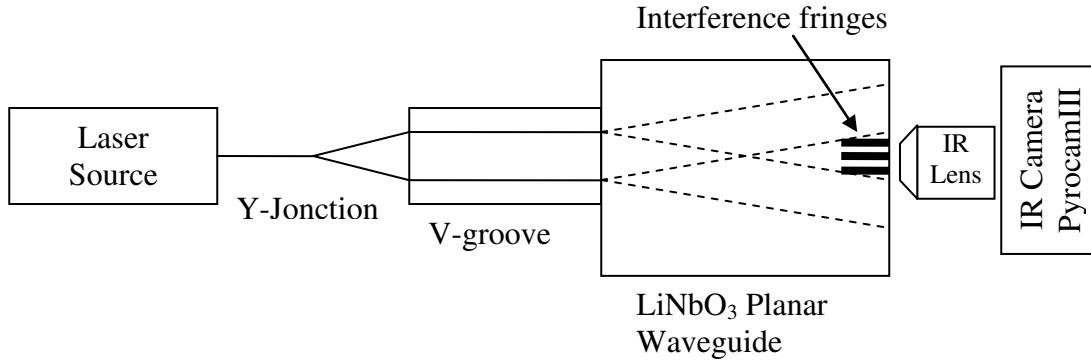


Fig. 7: Schematics of the infrared integrated optic basic interferometer

### 4.2 Theoretical model and expected values

Two chalcogenide fibers collect the photons from the He-Ne 3.39microns source and excite a planar waveguide. At the output, we recover the interference pattern given by:

$$I(Y) = \underbrace{4I_0 \operatorname{sinc}^2\left(\frac{\pi b Y n}{\lambda D}\right)}_{\text{diffraction}} \left[ 1 + \underbrace{e^{\frac{(2Y \Delta \lambda)^2}{\lambda^4}}}_{\text{temporal coherence}} \underbrace{\cos\left(\frac{2\pi a Y n}{\lambda D}\right)}_{\text{interferences}} \right] \quad (3)$$

Where b is the fiber core diameter, D is the planar waveguide length, a is the separation between the sources, n is the effective refractive index of the fundamental guided mode and Y is the lateral position along the output side.  $\Delta \lambda$  represents the spectral width of the source.

In our system, using Oxford Chalcogenide Fibers (core diameter 250 $\mu\text{m}$ , Cladding diameter 350 $\mu\text{m}$ ), we obtain a physical separation on the dedicated V-groove of 700 $\mu\text{m}$ . Using the He-Ne Laser Source at 3.39 $\mu\text{m}$ , considering a quasi-monochromatic source ( $\Delta\lambda\rightarrow 0$ ) and a single mode planar waveguide we obtain a theoretical distribution of fringes at the output given by:

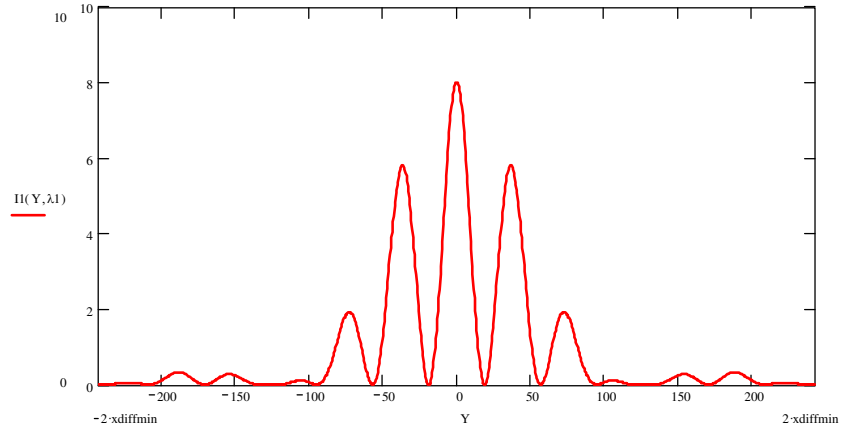


Fig.8: Theoretical distribution of fringes

That is, a diffraction limited pattern, with the fringe period  $i = \lambda D / a \cdot n_{\text{eff}} = 37.81$  microns for a  $D = 19.5$  mm long waveguide and a first maximum of diffraction of width  $w = 242$  microns. Here we used  $a = 0.8$  mm,  $n_{\text{eff}} = 2.185$  for the  $\text{LiNbO}_3$  fundamental mode at  $\lambda = 3.39$   $\mu\text{m}$  in the ordinary axis.

### 4.3 Infrared results

The results obtained for 3.39 $\mu\text{m}$  excitation are shown in figure 9. The experimental period is  $i = 38.31$  mic, in good agreement with the theoretical expected values. It's important to note that pixel size is 50 $\mu\text{m}$ , so that an error of one pixel results in a modification of the measured period of 1 $\mu\text{m}$  (when averaging with the four observed fringes).

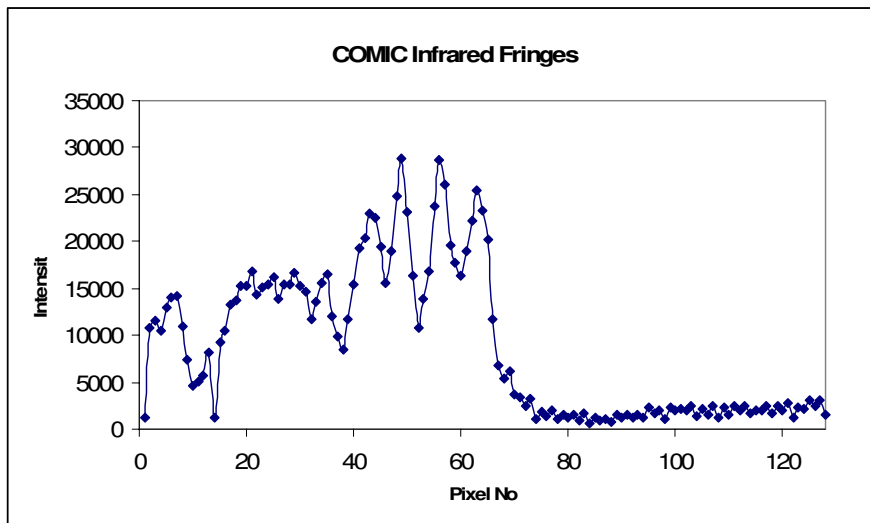


Fig.9: Experimental observation of IR fringes

Implementing a Fast Fourier Transform over the measured fringes, we obtained a peak positioned at  $3.463\mu\text{m}$ , compared with the theoretical emission line at  $3.39\mu\text{m}$ . Although the obtained spectra is weakly resolved, due to the scarce number of fringes and the poor number of pixels available, we can confirm that our interference pattern allows to measure the emission peak wavelength of the source.

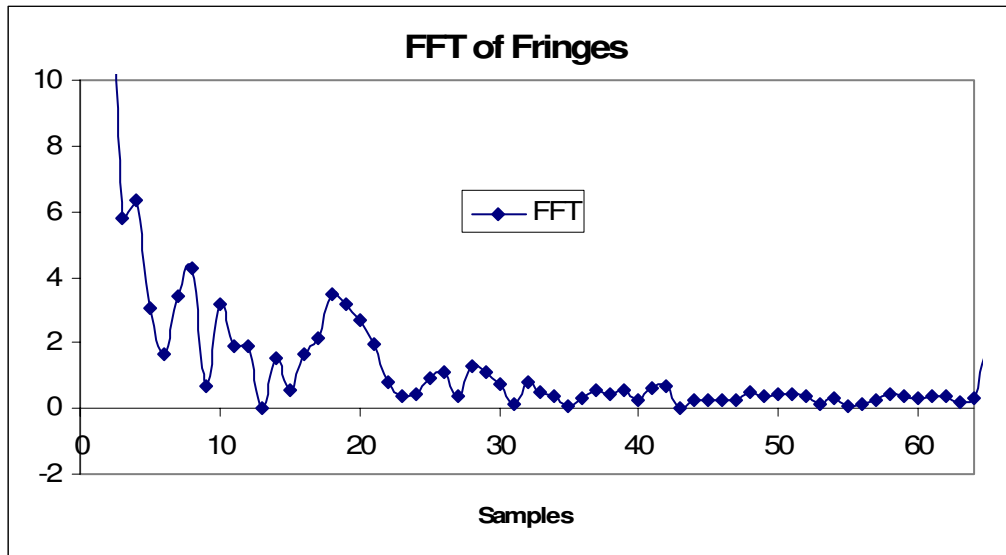


Fig.10: Showing the peak of intensity on pixel no 19. Scaling this value with the fringe period and pixel size, we obtain the position of the peak emission line.

## 5. DISCUSSION

Comparing theoretical and expected interference fringes we observe that the diffraction envelope is hardly seen experimentally, despite the good agreement in the period of the fringes. A Fourier analysis of the fringes allows to obtain the emission wavelength of the source and validate the method, despite the low resolution actually achievable. With the development of more sophisticated infrared cameras, increasing the pixel number and/or diminish their size, we should be able to improve the resolution on the spectrum of emission, and thus open the way to integrated spectroscopy in the L band.

## 6. CONCLUSION

We have presented three techniques with optimal potential for waveguide realization in the infrared range, namely around the L-band, applied to  $\text{LiNbO}_3$ . With Titanium diffusion and Ion-Beam Implantation we were able to obtain single mode planar waveguides. UV photoinscription, despite the ease of use, has not been sufficiently well mastered to allow us to produce channel waveguides in the infrared. It was however proved that single mode waveguides at  $633\text{nm}$  were obtained, and new samples are under study aiming the infrared L band. Finally, we have installed at LAOG an optical bench for L-band integrated optics tests, and shown the possibility of interferometry in single mode waveguides. Optimization of this bench will allow increasing the resolution and give us the possibility to start spectroscopical studies around this infrared band.

## REFERENCES

- [1] O. Absil, R. den Hartog, P. Gongoin, P. Fabry, R. Wilhelm, P. Gitton and F. Puech, "Performance study of ground-based infrared Bracewell interferometers - Application to the detection of exozodiacal dusts with GENIE" *A&A* 448, 2, 787-800 (2006).
- [2] O. Absil, V. Coudé du Foresto, M. Barillot and M.R. Swain, "Nulling interferometry : performance comparison between antarctica and other ground-based sites" *A&A* 475, 3, 1185-1194 (2007).

- [3] Knacke R. F. and Kraetschmer W. *A&A* vol.92, no.3 (1980)
- [4] R. Visser et al. *Molecules in Space and Laboratory*, Paris 2007
- [5] Segransan SPIE 2002.
- [6] D. Sudarsky, A. Burrows and I. Hubeny, "Theoretical Spectra and Atmospheres of Extrasolar Giant Planets" Vol. 588 (2), 1121-1148 (2003)
- [7] G. Li, T. Eckhause, K.A. Winick, J.D. Monnier and J.-P. Berger, "Integrated optic beam combiner in lithium niobate for stellar interferometer" *Proc. Of SPIE*, 6268 (2006).
- [8] Mailis, G., Risiotis, C., Wellington, I.T., Smith, P.G.R., Gawith, C.B.E. and Eason, R.W., "Direct ultraviolet writing of channel waveguides in congruent lithium niobate single crystals" *Opt. Lett.* 28 (16) 1433-1435 (2003).
- [9] Gallo, K., Gawith, C.B.E., Wellington, I.T., Mailis, S., Eason, R.W, Smith, P.G.R., Richardson, D.J. and Kostrikskii, S.M. "Ultraviolet writing of channel waveguides in proton exchanged LiNbO<sub>3</sub>" *J. Appl. Phys.* 101 (2007).
- [10] Tien, P.K. "light Waves in thin films and integrated optics" *Appl. Opt.* 10 (11) 2395 (1971).
- [11] White, J.M. and Heidrich, P.F. "Optical waveguide refractive index profiles determined from measurement of mode indices: simple analysis" *Appl. Optics* 15 (1), 151-155 (1976).
- [12] Olivares, J., Garcia, G., Garcia-Navarro, G., Agullo-Lopez, F., Caballero, O. and Garcia-Cabañes, A. "Generation of high-confinement step-like optical waveguides in LiNbO<sub>3</sub> by swift heavy ion irradiation" *Appl. Phys. Lett.* 86 (2005).
- [13] "VSI: A milliarcsec spectro-imager for the VLTI". F. Malbet et al. arXiv astro (2006)
- [14] COMIC: Lacombe, F., Marco, O., Geoffray, H., Beuzit, J.L., Monin, J.L., Gigan, P., Talureau, B., Feautrier, P., Petmezakis, P., Bonaccini, D. "Adaptative Optics Imaging at 1-5 microns on Large Telescopes: The COMIC Camera for ADONIS", *PASP*, 110, 1087 (1998).