## Solid Iron Compressed Up to 560 GPa

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Dynamic compression by multiple shocks is used to compress iron up to 560 GPa (5.6 Mbar), the highest solid-state pressure yet attained for iron in the laboratory. Extended x-ray absorption fine structure (EXAFS) spectroscopy offers simultaneous density, temperature, and local-structure measurements for the compressed iron. The data show that the close-packed structure of iron is stable up to 560 GPa, the temperature at peak compression is significantly higher than expected from pure compressive work, and the dynamic strength of iron is many times greater than the static strength based on lower pressure data. The results provide the first constraint on the melting line of iron above 400 GPa.

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Material properties at high-energy-density conditions are of broad interest across several scientific fields [1,2], from geophysics, planetary science, to laboratory astrophysics as well as inertial confinement fusion. Iron (Fe) is a key constituent of terrestrial planets and exoplanets and the sixth most abundant element in the Universe and, thus, is one of the most studied materials under extreme conditions. Theoretical calculations have made many predictions on Fe structures [3–8], melting line [9,10], elasticity [11–17], and conductivity [18] at elevated pressures and temperatures. Recently, the theoretical phase diagram of Fe has been extended up to 100 TPa and 40000 K [19]. However, experimental data on solid Fe are reported only up to 370 GPa [20–23]. Measurements at higher pressures represent a challenging frontier in material science, which can impact the development of evolution models for exoplanets [24]. The lack of data has left the theoretical models and predictions untested in the pressure regime beyond Earth's core conditions.

In this Letter we report measurements of solid Fe compressed up to 560 GPa, providing the highest-pressure data to date for constraining these models. The data demonstrate a close-packed structure of solid Fe up to 560 GPa, and provide the first constraint on the melting line of Fe above 400 GPa. The temperature at peak compression is substantially higher than expected from compressive work, indicating an enhanced strength of Fe under our experimental conditions.

The record pressure on Fe is achieved by dynamic ramp compression, where the pressure history is tuned to follow a particular path. This technique was recently used to explore solid diamond to 800 GPa [25]. In contrast to single-shock compression where the states are all along the Hugoniot, ramp or multishock compression can achieve off-Hugoniot states with lower temperatures. Hence, materials can maintain a solid structure while being compressed into the TPa regime.

Diagnosing the material properties under extreme conditions is as important as the creation of high-pressure states. We have performed EXAFS (extended x-ray absorption fine structure) measurements on the compressed Fe, providing in situ density, temperature, and localstructure data. EXAFS refers to the oscillatory modulations in x-ray absorption spectra above an absorption edge, which are generated by interference between photoelectron waves and scattering by neighbor atoms [26]. The period of the modulations depends primarily on the nearest-neighbor distance, from which the density can be determined. The decay of the modulations at higher x-ray energies is set by the Debye-Waller factor (DWF) [27], from which the temperature can be determined. EXAFS spectroscopy has been employed to study dynamically compressed Fe up to 35 GPa [28]. Here we extend the measurements to much higher pressures by taking into account anharmonic effects [29–31].

The experiments were performed on the OMEGA laser system at the Laboratory for Laser Energetics (LLE). As shown in Fig. 1(a), a broadband x-ray backlighter was generated by a spherical implosion [28,32]. The target was a  $4\mu$ m-thick Fe foil (purity better than 99.99%) sandwiched between two diamond plates, each 35  $\mu$ m thick. The diamond played a similar role as in the diamond anvil cell: confining the sample and maintaining the pressure, thus creating a spatially uniform compression state in Fe. The driver consists of five stacked laser pulses, and a typical pulse shape is plotted in Fig. 1(b). The corresponding stress history in Fe, obtained by velocimetry as described below, is plotted in Fig. 1(c). Both the laser energy and the delay between the driver and the backlighter were varied in order to probe different pressures and temperatures.

The stress in Fe is obtained from measurements of the free surface velocity of diamond using a line-imaging VISAR (velocity interferometer system for any reflector)

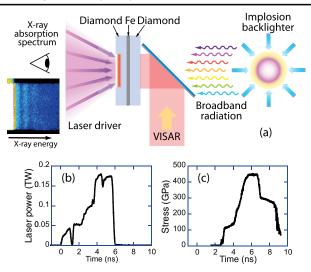


FIG. 1 (color online). (a) Experimental schematic. A raw image of a typical x-ray absorption spectrum is displayed at the leftlower corner, showing the intensity modulations above the K edge. (b) A typical laser pulse shape of the drive formed by five 1 ns laser pulses stacked in time. (c) The stress in Fe obtained from velocimetry measurements for the pulse shape in (b).

[33]. The 532-nm VISAR probe was directed onto the target by a mirror integrated into the target package. The mirror also shields the target from the implosion, and its presence has minimal effect on the EXAFS measurements. The method of characteristics [34] was used to backwards propagate the free surface velocity to determine the history of normal stress at the iron-diamond interface using a measured diamond equation of state [25]. The Fe thickness was substantially thinner than the diamond sandwich such that the pressure in the Fe sample equilibrates with that of the diamond. The pressure equilibration is confirmed by hydrodynamic simulations using LASNEX [35]. The short duration of the x-ray backlighter ( $\sim$ 120 ps) [32] ensures little temporal variation in the state of Fe during EXAFS measurements.

The normalized EXAFS signal  $\chi(k) = \mu(k)/\mu_0(k) - 1$ for undriven and compressed Fe is shown in Figs. 2(a) and 2(b), respectively ( $\mu_0$  is the smooth atomic like absorption coefficient, k is the wave number). At ambient conditions, Fe has a body-centered cubic (bcc) structure and the data agree well with both synchrotron measurements and calculations using FEFF code [26] [see Fig. 2(a)]. Fe is known to undergo a phase transition from bcc to hcp (hexagonal close packed) at pressures above 13 GPa [36,37]. The disappearance of the "w" peak is a signature of the phase transition [28] as shown in Fig. 2(b). The spectra of compressed Fe are fitted well with FEFF calculations of a hcp lattice (c/a = 1.60) using the IFEFFIT program [38] including the third cumulant to take into account the anharmonic effect. Variation of the c/a ratio in the range of 1.59–1.63 as determined in diffraction measurements [20] has negligible effect on the inferred density and temperature.

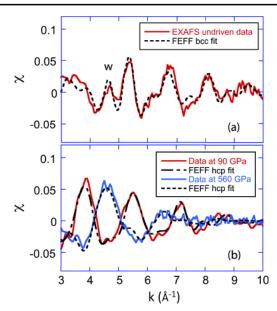


FIG. 2 (color online). (a) EXAFS data at ambient conditions and FEFF bcc fit. (b) EXAFS data at 90 GPa and 560 GPa, with corresponding FEFF fitting results.

Since simulations [3-8,19] suggest the existence of hcp, bcc, or fcc phases at high pressures and high temperatures, we performed first-principles quantum molecular dynamics (QMD) simulations to create bcc, hcp, fcc, as well as liquid atomic configurations of Fe to compare the difference in EXAFS spectra (see the Supplemental Material [39]). EXAFS measurements will not discriminate between close-packed phases including hcp, fcc, and dhcp as they display similar EXAFS signals [8]. In the following data analysis, we use the hcp phase to represent close-packed structures. The pair distribution function g(r)from the QMD simulations, shown in Fig. 3(a), indicates that at high temperatures, the first two shells of the bcc lattice are merged into a broader peak, which resembles the first peak in g(r) of a hcp lattice. However, at the same density the nearest-neighbor distance  $R_1$  in the bcc lattice is  $\sim 3\%$  shorter than that in the hcp lattice, and thus the resulting EXAFS signal is shifted for the hot bcc lattice relative to the hcp lattice. Therefore, fitting the same data with the hot bcc lattice results in an  $\sim 10\%$  reduction in density compared to hcp fitting. The liquid phase, characterized by an even broader and more shifted first peak in g(r), leads to even lower densities than the bcc phase in the data fitting.

The compression obtained from EXAFS data as a function of stress is plotted in Fig. 3(b). The density is determined from  $R_1$  which is obtained by fitting EXAFS data using four types of lattice structures: harmonic hcp, anharmonic hcp, bcc, and liquid generated by the QMD simulations. In the same figure the Fe Hugoniot [40] and principle isentrope from the LEOS library [41] are also shown for comparison. For the multishock compression, the density should lie between the Hugoniot and the

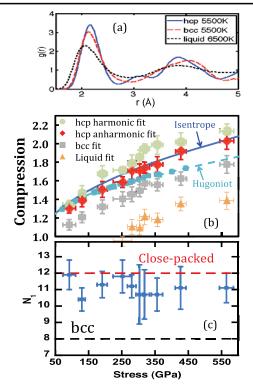


FIG. 3 (color online). (a) g(r) from QMD simulations for bcc, hcp, and liquid phases at  $1.575 \times \text{compression}$ . (b) Compression as a function of stress for four types of FEFF fitting of EXAFS data, together with Hugoniot data [40] and the principle isentrope [41]. The fitting of undriven data gives a compression of  $1.01 \pm 0.02$ . (c) Best-fit first-shell coordination number using GNXAS.

isentrope. Clearly the anharmonic correction is necessary to obtain the correct density for the hcp phase. The bcc and liquid fittings give lower density at the same stress as explained above, which is too low compared to the Hugoniot for all the data. Therefore, the fitting results indicate that our EXAFS data are inconsistent with the bcc and liquid phases.

The experimental EXAFS data were also analyzed using an independent approach implemented in the GNXAS package [42]. This approach is based on relating g(r) to  $\chi(k)$ , a well-established procedure developed for disordered systems. Coordination numbers of the first shell obtained from data fitting are shown in Fig. 3(c) as a function of stress. From 90 to 560 GPa, the average value of the best-fit  $N_1$  is  ${\sim}11\pm1,$  close to the expected coordination number of a close-packed structure ( $N_1 = 12$ , red dashed line) but far from the bcc lattice ( $N_1 = 8$ , black dashed line). Therefore, both FEFF and GNXAS fitting of EXAFS data are consistent with a close-packed structure rather than the bcc phase. The fact that the observed coordination number is slightly less than the ideal value for a close-packed structure could be due to microstructural defects [43] or significant thermal vibrations at such high temperatures. The  $\sigma^2$  values obtained from these two independent methods are also in good agreement (see the Supplemental Material [39]).

The temperature dependence of EXAFS measurements occurs through the DWF  $\sigma^2$  [43–46]. Since EXAFS spectroscopy measures the interatom distances, the correlation effect in atomic motion should be taken into account [27]. The correlated Einstein model and Debye model with harmonic approximation are commonly used to calculate  $\sigma^2$ , and generally produce similar results [31]. Taking into account the anharmonic effect in the Einstein model to the lowest order, in the high-temperature limit we get (see the Supplemental Material [39])

$$\sigma^{2} = \sigma_{E}^{2}(T) + (\sigma_{3}/\sigma_{E}^{2}(T))^{2}, \qquad (1)$$

where  $\sigma_3$  is the third cumulant, and  $\sigma_E^2$  is from the harmonic Einstein model  $\sigma_E^2 = \hbar/(M\omega_E) \coth(\hbar\omega_E)/(2k_BT)$ (*M* is the atomic mass,  $\omega_E$  is the Einstein frequency,  $\hbar$  is the Planck constant,  $k_B$  is the Boltzmann constant). The second term in Eq. (1) represents the lowest-order anharmonic contribution. The Einstein temperature is given by  $T_E = (3/4)T_{\text{Debye}}$  for the first shell [27], with  $T_{\text{Debye}}$ extrapolated from experimental measurements of Fe up to 300 GPa [21]. The validity of Eq. (1) has been confirmed using synthetic EXAFS data from QMD simulations (see the Supplemental Material [39]).

The temperature obtained from the DWF in EXAFS data using Eq. (1) is plotted in Figs. 4(a) and 4(b) as a function of stress. The relatively large error bars are due to the fact that at such high temperatures, the EXAFS signal is damped strongly at higher k, and hence only a limited number of peaks are observable as shown in Fig. 2(b). Nevertheless, the data clearly indicate that off-Hugoniot states have been achieved up to 560 GPa and 8000 K, which are the highest stress and temperature ever reached for solid iron. The data fall into three groups based on the compression history: single shock (blue diamonds), first shock stress  $P_1 = 80-100$  GPa shown in Fig. 4(a) (red squares), and  $P_1 = 130-150$  GPa shown in Fig. 4(b) (black triangles). The two single-shock points agree well with the calculated Hugoniot. The multishock data show the expected higher temperature at higher  $P_1$ , but compared to the isentropes calculated following a 100 and 150 GPa initial shock as plotted in Figs. 4(a) and 4(b), respectively (solid lines), our inferred temperatures are 1000-4000 K higher.

The temperature can result from processes such as x-ray preheat by either the laser driver or the backlighter, thermal conduction from diamond to Fe, shock heating, and heating due to work against the strength of Fe during the high strain rate compression (plastic work). EXAFS measurements of undriven targets as plotted in Fig. 2(a) give a temperature of  $310 \pm 70$  K, indicating that heating by the backlighter is negligible. We have performed radiation hydrodynamic simulations to estimate the x-ray heating by the driver and the thermal conduction. Both are much smaller than shock heating and plastic work. The contribution due to shock heating after the initial shock is also expected to be small because the subsequent compression

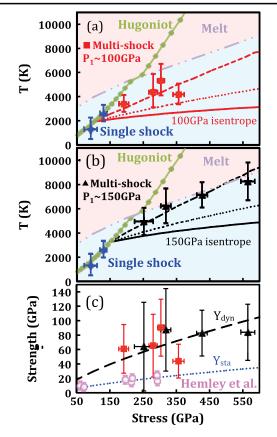


FIG. 4 (color online). (a) and (b) Temperature inferred from the DWF in EXAFS data using Eq. (1) as a function of stress for the two groups of multishock data with an initial shock of ~100 GPa and ~150 GPa, respectively. The single-shock data are also shown (blue diamonds). The melting curve [9] (dot-dot-dashed lines) and the Hugoniot [41] (green solid lines with dots) are plotted for comparison. Also shown are isentropes (solid lines), isentropes with the temperature increase calculated using  $Y_{sta}$  (dotted lines) and  $Y_{dyn} = 3Y_{sta}$  (dashed lines). (c) Fe strength vs stress from our data (red squares and black triangles) and static data (circles) [22].

is either shockless or consists of weak shocks with little additional entropy. The major contribution to the temperature is thus mainly due to heating by the first shock and the plastic work. The latter can be expressed as [47]

$$\Delta T_{\text{plastic}} = \frac{V_0}{C_V} \int_{\eta_1} Y(\eta) \frac{d\varepsilon}{d\eta} d\eta, \qquad (2)$$

where  $\varepsilon$  is the plastic strain,  $\varepsilon = (2/3)[\ln(\eta) - Y/(2G)]$ , *Y* is the strength, *G* is the shear modulus,  $C_V$  is the specific heat,  $\eta = \rho/\rho_0$ ,  $\rho$  is the density,  $V_0$  and  $\rho_0$  are the volume and the density at ambient conditions, and  $\eta_1$  is the compression by the first shock. The Dulong-Petit limit is used for the heat capacity  $C_V \simeq 3k_B$  as the temperature already reaches above 3000 K for off-Hugoniot data after the first shock. We used a density- and temperature-dependent shear modulus *G* based on Ref. [11], but the choice of *G* [11–17] does not significantly affect the results because in the expression of  $d\varepsilon/d\eta$ ,  $d/d\eta(Y/2G) \ll (d/d\eta) \ln(\eta)$ under our conditions. For the strength of Fe, the published data up to 300 GPa by static compression ( $Y_{sta}$ ) are generally considered as the lower limit due to deformation of diamond anvils [22]. The data trend is extrapolated to 600 GPa as shown in Fig. 4(c) (dotted line). The plastic work calculated using  $Y_{sta}$  leads to an ~500–1500 K temperature increase as shown in Figs. 4(a) and 4(b) (dotted line), which is close to the lower limit of our data.

This substantial plastic heating indicates that the strength of Fe is much higher than expected from lowerpressure static data. The upper limit of the dynamic strength  $Y_{dyn}$  for our compression path can be estimated by attributing the difference between the experimental temperature and the calculated isentrope temperature to the plastic work. The temperature rise using an assumed dynamic strength  $Y_{dyn} = 3Y_{sta}$  is shown by the dashed lines in Figs. 4(a) and 4(b), which, on the average, is consistent with both groups of data. Comparison of  $Y_{dyn}$  from our multishock data and  $Y_{sta}$  from static data is shown in Fig. 4(c). Besides the fact that the static experiments give a lower limit on the strength due to anvil deformation, the difference between  $Y_{dyn}$  and  $Y_{sta}$  can also arise from the commonly observed increase in strength at high strain rates [48]. One uncertainty in setting  $Y_{dyn}$  lies in the fraction of plastic work converted into heat, the Taylor-Quinney factor  $\beta$  [49]. In the above calculations,  $\beta$  is assumed to be 100%. Recent experiments have shown that as the strain rate is increased from  $3.8 \times 10^3$  to  $8.4 \times 10^3$  s<sup>-1</sup>,  $\beta$  rises from  $\sim$ 50% to  $\sim$ 100% [50]. However, there are no measurements or calculations on  $\beta$  so far for strain rates on the order of  $10^6 - 10^7$  s<sup>-1</sup> as in our experiments; thus, our data effectively constrain the upper limit of the product  $\beta Y$ .

To summarize, EXAFS measurements have provided density, temperature, and local-structure data for Fe compressed up to 560 GPa. By comparing with QMD simulations, the data are consistent with a close-packed solid structure rather than bcc or liquid phases. The high temperature at peak compression is interpreted to be due to the dynamic strength of Fe, which, for the compression history explored here, is several times greater than expected from lower pressure static data. The platform developed for the EXAFS study of Fe can be applied to many solids at pressures well into the TPa regime, which is accessible by ramp or multishock compression using the National Ignition Facility [51].

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- National Research Council (U.S.) Committee on High Energy Density Physics, Frontiers in High Energy Density Physics: The X-Games of Contemporary Science (National Academies Press, Washington, DC, 2003).
- [2] Report of the Workshop on High Energy Density Laboratory Physics Research Needs, DOE Office of Science and NNSA, 2009.
- [3] A.B. Belonoshko, R. Ahuja, and B. Johansson, Nature (London) **424**, 1032 (2003).
- [4] Y. Kuwayama, K. Hirose, N. Sata, and Y. Ohishi, Earth Planet. Sci. Lett. 273, 379 (2008).
- [5] T. Sakai, E. Ohtani, N. Hirao, and Y. Ohishi, Geophys. Res. Lett. 38, L09302 (2011).
- [6] L. Vocadlo, I. G. Wood, D. Alfe, and G. D. Price, Earth Planet. Sci. Lett. 268, 444 (2008).
- [7] A. S. Mikhaylushkin, S. Simak, L. Dubrovinsky, N. Dubrovinskaia, B. Johansson, and I. Abrikosov, Phys. Rev. Lett. 99, 165505 (2007).
- [8] K. Kadau, T. Germann, P. Lomdahl, R. Albers, J. Wark, A. Higginbotham, and B. Holian, Phys. Rev. Lett. 98, 135701 (2007).
- [9] D. Alfe, G. D. Price, and M. J. Gillan, Phys. Rev. B 65, 165118 (2002); D. Alfe, Phys. Rev. B 79, 060101 (2009).
- [10] A. B. Belonoshko, R. Ahuja, and B. Johansson, Phys. Rev. Lett. 84, 3638 (2000).
- [11] X. Sha and R. E. Cohen, Phys. Rev. B 81, 094105 (2010).
- [12] R. E. Cohen, L. Stixrude, and E. Wasserman, Phys. Rev. B 56, 8575 (1997).
- [13] A. Laio, S. Bernard, G. L. Chiarotti, S. Scandolo, and E. Tosatti, Science 287, 1027 (2000).
- [14] G. Steinle-Neumann, L. Stixrude, R.E. Cohen, and O. Gulseren, Nature (London) 413, 57 (2001).
- [15] D. Alfe, G. D. Price, and M. J. Gillan, Phys. Rev. B 64, 045123 (2001).
- [16] L. Vocadlo, D. P. Dobson, and I. G. Wood, Earth Planet. Sci. Lett. 288, 534 (2009).
- [17] A.B. Belonoshko, T. Bryk, and A. Rosengren, Phys. Rev. Lett. 104, 245703 (2010).
- [18] M. Pozzo, C. Davis, D. Gubbins, and D. Alfe, Nature (London) 485, 355 (2012).
- [19] L. Stixrude, Phys. Rev. Lett. 108, 055505 (2012).
- [20] S. Tateno, K. Hirose, Y. Ohishi, and Y. Tatsumi, Science 330, 359 (2010).
- [21] L.S. Dubrovinsky, S.K. Saxena, N.A. Dubrovinskaia, S. Rekhi, and T. Le Bihan, Am. Mineral. 85, 386 (2000).
- [22] R.J. Hemley, H. Mao, G. Shen, J. Badro, P. Gillet, M. Hanfland, and D. Hausermann, Science 276, 1242 (1997).
- [23] S. Anzellini, A. Dewaele, M. Mezouar, P. Loubeyre, and G. Morard, Science 340, 464 (2013).
- [24] D.C. Swift, J.H. Eggert, D.G. Hicks, S. Hamel, K. Caspersen, E. Schwegler, G. W. Collins, N. Nettelmann, and G.J. Ackland, Astrophys. J. 744, 59 (2012).
- [25] D. K. Bradley, J. Eggert, R. Smith, S. Prisbrey, D. Hicks, D. Braun, J. Biener, A. Hamza, R. Rudd, and G. Collins, Phys. Rev. Lett. **102**, 075503 (2009).
- [26] J. J. Rehr and R. C. Albers, Rev. Mod. Phys. 72, 621 (2000).
- [27] E. Sevillano, H. Meuth, and J.J. Rehr, Phys. Rev. B 20, 4908 (1979).

- [28] B. Yaakobi, T. Boehly, D. Meyerhofer, T. Collins, B. Remington, P. Allen, S. Pollaine, H. Lorenzana, and J. Eggert, Phys. Rev. Lett. 95, 075501 (2005).
- [29] J. M. Tranquada and R. Ingalls, Phys. Rev. B 28, 3520 (1983).
- [30] E. A. Stern, P. Livins, and Z. Zhang, Phys. Rev. B 43, 8850 (1991).
- [31] K.D. Machadoa, J. Chem. Phys. 134, 064503 (2011).
- [32] B. Yaakobi, F. J. Marshall, T. R. Boehly, R. P. J. Town, and D. D. Mayerhofer, J. Opt. Soc. Am. B 20, 238 (2003).
- [33] P. M. Celliers, D. K. Bradley, G. W. Collins, D. G. Hicks, T. R. Boehly, and W. J. Armstrong, Rev. Sci. Instrum. 75, 4916 (2004).
- [34] S. D. Rothman and J. Maw, J. Phys. IV (France) 134, 745 (2006).
- [35] G.B. Zimmerman and W.L. Kruer, Comments Plasma Phys. Contr. Fusion **2**, 51 (1975).
- [36] D. H. Kalantar et al., Phys. Rev. Lett. 95, 075502 (2005).
- [37] T. Takahashi and W. A. Bassett, Science 145, 483 (1964);
  W. A. Bassett and E. Huang, Science 238, 780 (1987).
- [38] B. Ravel and M. J. Newville, Synchrotron Radiation News 12, 537 (2005).
- [39] See Supplemental Material at http://link.aps.org/ supplemental/10.1103/PhysRevLett.111.065501 for details of VISAR measurements, QMD simulations, EXAFS data analysis, and validation of Eq. (1).
- [40] J. M. Brown, J. N. Fritz, and R. S. Hixson, J. Appl. Phys. 88, 5496 (2000).
- [41] The LEOS data library is based on a QEOS-like model described in R. M. More, K. H. Warren, D. A. Young, and G. B. Zimmerman, Phys. Fluids **31**, 3059 (1988); D. A. Young and E. M. Corey, J. Appl. Phys. **78**, 3748 (1995).
- [42] A. Filipponi and A. Di Cicco, Phys. Rev. B 52, 15135 (1995); A. Filipponi, J. Phys. Condens. Matter 6, 8415 (1994).
- [43] A. Di Cicco, M. Berrettoni, S. Stizza, E. Bonetti, and G. Cocco, Phys. Rev. B 50, 12386 (1994).
- [44] X. Zhu, R. Birringer, U. Herr, and H. Gleiter, Phys. Rev. B 35, 9085 (1987).
- [45] A. Higginbotham, R.C. Albers, T.C. Germann, B.L. Holian, K. Kadau, P.S. Lomdahl, W.J. Murphy, B. Nagler, and J.S. Wark, High Energy Density Phys. 5, 44 (2009).
- [46] Generally the DWF has contributions from both the configurational disorder (static) and thermal disorder (dynamic). The static contribution has been shown to be negligible in either nanocrystalline Fe [43,44] or shockcompressed Fe [45] because the interfacial regions do not have any short-range order. Therefore the DWF in Fe EXAFS data is dominated by the thermal contribution.
- [47] G.R. Fowles, J. Appl. Phys. 32, 1475 (1961).
- [48] R.F. Smith, J.H. Eggert, R.E. Rudd, D.C. Swift, C.A. Bolme, and G.W. Collins, J. Appl. Phys. **110**, 123515 (2011).
- [49] G.I. Taylor and H. Quinney, Proc. R. Soc. A 143, 307 (1934).
- [50] D. Rittel, G. Ravichandran, and A. Venkert, Mater. Sci. Eng. A 432, 191 (2006).
- [51] G.H. Miller, E.I. Moses, and C.R. Wuest, Opt. Eng. (Bellingham, Wash.) 43, 2841 (2004).