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Stabilization of the electroweak vacuum by a scalar threshold effect

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ABSTRACT: We show how a heavy scalar singlet with a large vacuum expectation value can evade the potential instability of the Standard Model electroweak vacuum. The quartic interaction between the heavy scalar singlet and the Higgs doublet leads to a positive tree-level threshold correction for the Higgs quartic coupling, which is very effective in stabilizing the potential. We provide examples, such as the see-saw, invisible axion and unitarized Higgs inflation, where the proposed mechanism is automatically implemented in well-defined ranges of Higgs masses.

KEYWORDS: Beyond Standard Model, Standard Model

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Contents

1	Introduction	1
2	Stabilizing the Higgs with a scalar singlet	2
2.1	Threshold effect	3
2.2	Case $\lambda_{HS} > 0$	4
2.3	Case $\lambda_{HS} < 0$	7
3	Examples	8
3.1	See-saw	9
3.2	Invisible axion	11
3.3	Unitarized Higgs inflation	12
4	Conclusions	14

1 Introduction

Collider searches have restricted the allowed mass window for the Standard Model Higgs boson to the narrow range between 115.5 and 127 GeV at 95% CL [1, 2]. From the theoretical point of view, this is an interesting range of masses because in this interval the Higgs potential can develop an instability at large field values (for detailed analyses, see refs. [3–20]). This is of course not necessarily a problem, first because the Standard Model is likely to be embedded in a more fundamental theory at high energies which may change the shape of the Higgs potential and, second, because the actual fate of the electroweak vacuum depends on the cosmological history. It should be remarked, however, that an instability caused by a negative quartic coupling is not cured at high temperatures (as opposed to an instability caused by a negative quadratic term) and typically persists in the early universe. Therefore, to avoid potential cosmological constraints on the reheat temperature or the Hubble constant during inflation [21], it may be preferable to cure any Higgs instability at large field values.

There are, of course, infinitely many ways to modify the Higgs potential and raise the instability scale Λ_I . In this paper we point out a very simple and economical mechanism to avoid an instability of the electroweak vacuum. It requires the existence of a new heavy scalar singlet that acquires a large vacuum expectation value (vev) and has a quartic interaction with the ordinary Higgs doublet. The crucial point is that the matching condition of the Higgs quartic coupling, at the scale where the singlet is integrated out, corresponds to a positive shift, as we evolve from low to high energies. Although the stability condition is also modified by the presence of the singlet, a careful analysis shows that, under the conditions specified in section 2, the threshold correction helps to stabilize the potential.

The effect occurs at tree level and thus can be sizable and, in general, dominant over loop contributions. Moreover, the effect does not decouple, in the sense that the size of the shift does not depend on the singlet mass, which could take any value lower than the instability scale Λ_I .

The dynamics of the mechanism is explained in section 2. In section 3 we put the proposed mechanism into context, considering setups in which the existence of a heavy scalar singlet has independent motivations. We present three examples: see-saw origin of neutrino masses, invisible axion, and Higgs inflation. For each example we give the range of Higgs masses where the corresponding scalar field could cure the instability.

2 Stabilizing the Higgs with a scalar singlet

To explore the impact of an additional singlet scalar on the stability of the Higgs potential, we consider a tree-level scalar potential of the form

$$V_0 = \lambda_H \left(H^\dagger H - v^2/2 \right)^2 + \lambda_S \left(S^\dagger S - w^2/2 \right)^2 + 2\lambda_{HS} \left(H^\dagger H - v^2/2 \right) \left(S^\dagger S - w^2/2 \right) . \tag{2.1}$$

Here H is the Higgs doublet, S is a complex scalar field, and V_0 is the most general renormalizable potential that respects a global abelian symmetry under which only S is charged. Although we will consider here a single complex scalar, most of our conclusions remain valid also in the case of multi-Higgs doublets or real singlet fields (with a Z_2 parity replacing the abelian symmetry).

For $\lambda_H, \lambda_S > 0$ and $\lambda_{HS}^2 < \lambda_H \lambda_S$, the minimum of V_0 is at

$$\langle H^\dagger H \rangle = v^2/2, \quad \langle S^\dagger S \rangle = w^2/2 . \tag{2.2}$$

A nonzero vev of S , which is crucial for our mechanism to work, spontaneously breaks the global symmetry (or the Z_2 parity, for a real singlet) giving rise to a potentially dangerous Goldstone boson (or domain walls). Gauging the symmetry of S or explicitly breaking it by (possibly small) terms in V_0 can be used to evade these problems, but does not conceptually modify our results. For simplicity, we restrict our considerations to the potential in eq. (2.1), but generalizations are straightforward.

The presence of the new scalar field S modifies the analysis of the stability conditions of the Higgs potential. One effect, already considered in previous literature is the contribution of the singlet to the renormalization group evolution of the Higgs quartic coupling (for recent analyses, see [22–26] and references therein). The relevant renormalization group equations (RGEs) above the scale $M_S = \sqrt{2\lambda_S}w$ are, at one-loop:

$$\begin{aligned} (4\pi)^2 \frac{d\lambda_H}{d\ln\mu} &= \left(12y_t^2 - 3g'^2 - 9g^2 \right) \lambda_H - 6y_t^4 + \frac{3}{8} \left[2g^4 + (g'^2 + g^2)^2 \right] + 24\lambda_H^2 + 4\lambda_{HS}^2, \\ (4\pi)^2 \frac{d\lambda_{HS}}{d\ln\mu} &= \frac{1}{2} \left(12y_t^2 - 3g'^2 - 9g^2 \right) \lambda_{HS} + 4\lambda_{HS} (3\lambda_H + 2\lambda_S) + 8\lambda_{HS}^2, \\ (4\pi)^2 \frac{d\lambda_S}{d\ln\mu} &= 8\lambda_{HS}^2 + 20\lambda_S^2. \end{aligned} \tag{2.3}$$

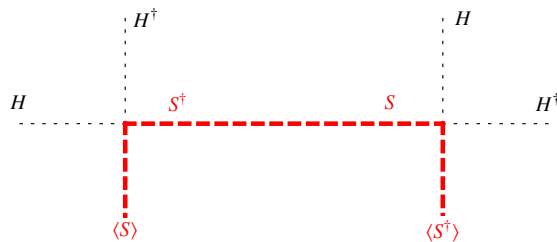


Figure 1. Feynman diagram for the tree-level threshold correction to the Higgs quartic coupling.

If the singlet mass M_S is below the SM instability scale¹ Λ_I and $(\lambda_{HS}/4\pi)^2 \ln(\Lambda_I/M_S)$ is large enough, the positive contribution to the RGE equation for λ_H can prevent it from becoming negative.

2.1 Threshold effect

Besides the loop contribution discussed above, there is a related tree-level effect through which the new singlet can affect the stability bound. Let us consider the limit in which M_S is much larger than the Higgs mass ($w^2 \gg v^2$). At the scale M_S we can integrate out the field S using its equation of motion (neglecting derivatives):

$$S^\dagger S = \frac{w^2}{2} - \frac{\lambda_{HS}}{\lambda_S} \left(H^\dagger H - \frac{v^2}{2} \right). \quad (2.4)$$

Replacing eq. (2.4) in V_0 , we obtain the effective potential below the scale M_S :

$$V_{\text{eff}} = \lambda \left(H^\dagger H - \frac{v^2}{2} \right)^2, \quad \lambda(M_S) = \lambda_H - \frac{\lambda_{HS}^2}{\lambda_S} \Big|_{M_S}. \quad (2.5)$$

This shows that the matching condition at the scale $Q = M_S$ of the Higgs quartic coupling gives a tree-level shift, $\delta\lambda \equiv \lambda_{HS}^2/\lambda_S$, as we go from λ_H just above M_S to λ just below M_S . Figure 1 shows the Feynman diagram that gives rise to such tree-level shift.

To better understand the origin of the shift in the matching condition, let us consider the mass matrix of the fields h and s , corresponding to the real parts of the doublet H (in unitary gauge) and the singlet S , such that $H^\dagger H = h^2/2$ and $S^\dagger S = s^2/2$. At the minimum, the mass matrix is

$$\mathcal{M}^2 = 2 \begin{pmatrix} \lambda_H v^2 & \lambda_{HS} v w \\ \lambda_{HS} v w & \lambda_S w^2 \end{pmatrix}. \quad (2.6)$$

In the limit $\lambda_S w^2 \gg \lambda_H v^2$, the heaviest eigenstate, which is nearly singlet, can be integrated out, leaving behind a “see-saw”-like correction to the lightest eigenvalue

$$m_h^2 = 2v^2 \left[\lambda_H - \frac{\lambda_{HS}^2}{\lambda_S} + \mathcal{O}\left(\frac{v^2}{w^2}\right) \right], \quad (2.7)$$

¹Stabilizing the potential with degrees of freedom heavier than Λ_I requires sizable couplings, see [27, 28].

while $M_S^2 = 2\lambda_S w^2 + 2(\lambda_{HS}^2/\lambda_S)v^2 + \mathcal{O}(v^4/w^2)$. The light state is almost purely h , as the singlet admixture is suppressed by a small mixing angle of order v/w . However, the Higgs mass correction due to the heavy state persists even in the decoupling limit ($w \rightarrow \infty$). The negative sign in the shift of the Higgs mass in eq. (2.7) can be readily understood as coming from the repulsion of mass eigenvalues after turning on the mixing equal to $2\lambda_{HS}vw$.

Naively, as the tree-level shift $\delta\lambda$ corresponds to a larger Higgs quartic coupling above M_S , the chances of keeping it positive seem improved. However, the tree-level conditions for stability change from $\lambda > 0$ in the effective theory below M_S to $\lambda_H > \delta\lambda$ in the full theory above M_S . Thus, it appears that the threshold correction $\delta\lambda$ does not help stability at all. To understand what happens, one has to reexamine the stability conditions more carefully. First of all, remember that the tree-level potential V_0 in eq. (2.1) is a good approximation to the full potential if we evaluate couplings and masses (collectively denoted by λ_i below) at a renormalization scale of the order of the field values of interest. Once we express the scalar potential as $V_0[\lambda_i(\mu = \varphi), \varphi]$, potentially large logarithms of the form $\ln m_i(\varphi)/\mu$ (where $m_i(\varphi) \sim \varphi$ is a typical field-dependent mass) are kept small. Roughly speaking, this means that V_0 with a fixed μ_c will be reliable as long as one examines $\varphi \sim \mu_c$ and restricts field excursions to $|\varphi - \mu_c| < \mu_c e^{8\pi^2\lambda_0/\lambda_1^2}$ (where λ_0 denotes a coupling in the tree-level potential and λ_1 a coupling affecting the radiative corrections, e.g. the top Yukawa coupling squared). By adjusting $\mu \sim \varphi$ one can evaluate reliably the potential at all field values, but the previous estimate tells us when we can use $V_0[\lambda_i(\mu_c), \varphi]$, which has a simpler field dependence.

With the parametrization chosen in eq. (2.1), the EW vacuum corresponds to $V_0 = 0$. Thus the stability condition is $V_0 > 0$ anywhere in field space, away from the EW vacuum. The first stability requirement that we should impose is

$$\lambda_H(\mu) > 0, \quad \lambda_S(\mu) > 0, \tag{2.8}$$

at any renormalization scale μ , or else the potential develops unwanted minima lower than the EW vacuum or is unbounded from below at large field values.

Next, in order to discuss the conditions on the coupling constant λ_{HS} , it is convenient to separate the cases in which λ_{HS} is either positive or negative. This separation is meaningful because λ_{HS} renormalizes multiplicatively (as it is the only coupling that connects H and S), see eq. (2.3), and therefore the RG flow cannot flip its sign.

2.2 Case $\lambda_{HS} > 0$

In this case, V_0 can become negative only when $|S| < w/\sqrt{2}$ (neglecting corrections proportional to v). In this situation, the most dangerous field configuration is well approximated by setting $S = 0$ in eq. (2.1), such that

$$V_0(H, 0) \approx \lambda_H |H|^4 - \frac{\lambda_{HS}}{2\lambda_S} M_S^2 |H|^2 + \frac{M_S^4}{16\lambda_S}. \tag{2.9}$$

The extra stability condition ($V_0 > 0$) is then

$$\lambda_{HS}^2(\mu) < \lambda_H(\mu)\lambda_S(\mu). \tag{2.10}$$

Note that this can be rewritten as $\lambda_H > \delta\lambda = \lambda_{HS}^2/\lambda_S$ and ensures that the light scalar state does not become tachyonic, see eq. (2.7). If this condition were violated at some scale μ_* , it would lead to an instability for field configurations with

$$|S| < \frac{M_S}{2\sqrt{\lambda_S}}, \quad \mu_- < |H| < \mu_+, \quad \mu_{\pm}^2 = \frac{M_S^2 \lambda_{HS}}{4\lambda_H \lambda_S} \left(1 \pm \sqrt{1 - \frac{\lambda_H \lambda_S}{\lambda_{HS}^2}} \right) \Big|_{\mu_*}, \quad (2.11)$$

which could be trusted provided $\mu_- < \mu_* < \mu_+$. Note that, if $\mu_* \gg \mu_{\pm}$, this would not mean that there is an instability to worry about, as it would be located outside the range of validity of the tree-level approximation $V_0(\lambda_i(\mu_*), \varphi)$. Thus, as long as condition (2.10) is satisfied for renormalization scales within a relatively narrow range of energies around M_S (which fixes the mass scale of μ_{\pm}), there is no instability even if this condition were eventually violated at higher scales. Only if parameters happen to lie near a critical point in which at least one of conditions (2.8) or (2.10) is barely satisfied, radiative corrections can become important and invalidate the stability analysis performed with the tree-level potential. In this case one should resort to the one-loop approximation of the potential; otherwise, our analysis is reliable.

We can now better appreciate how the threshold contribution in eq. (2.5) can cure the instability of the SM Higgs potential (provided that $M_S < \Lambda_I$). The correction $\delta\lambda$ has the correct sign to shift the Higgs quartic coupling upwards ($\lambda_H = \lambda + \delta\lambda$), although the stability condition is also shifted upwards by the same amount, becoming $\lambda_H > \delta\lambda$. However, for positive λ_{HS} , the condition $\lambda_H > \delta\lambda$ has to be satisfied only at scales of order M_S , while for larger scales it rapidly reduces to the conventional constraint $\lambda_H > 0$. Moreover, one-loop RG effects (although typically less important than the tree-level matching condition) also help to maintain stability. First, λ_S and λ_{HS} will stay positive once they are positive at M_S . Second, $\beta_{\lambda_H} \equiv d\lambda_H/d\ln\mu$ receives extra positive contributions proportional to λ_{HS}^2 and to λ_H^2 (coupling which is numerically larger after the threshold shift). These two RG effects can reduce (or even overcome) the destabilizing effect from top loops.

To illustrate the situation, we show in figure 2, left panel, how the Higgs quartic coupling runs with the renormalization scale. We consider $M_S = 10^8 \text{ GeV} \ll \Lambda_I^{\text{SM}} = 2 \times 10^9 \text{ GeV}$.² For simplicity we take the couplings of the singlet to be smaller than the SM top and gauge couplings, in order to better isolate the tree-level threshold effect. The same panel also shows the full stability condition, computed numerically by demanding that $V(H, 0) > 0$: we see that at renormalization scales just above M_S the stability condition of eq. (2.10) matters, but at larger field values it rapidly becomes irrelevant and only $\lambda_H > 0$ remains.

To study the efficiency of the stabilization mechanism, we performed a numerical study using the full one-loop effective potential with SM couplings running at two-loops. We limited the evolution of the unknown singlet couplings to the one-loop level, given that the effect we are considering is at tree level. In order to track accurately the large field

²Strictly speaking, this is the scale at which $\lambda = 0$, and corresponds to the instability scale of the tree-level RG-improved potential. The Λ_I that we calculate later on is higher and corresponds to the instability of the one-loop RG-improved potential. For simplicity, in figure 3 we simply plot $\lambda(\mu)$.

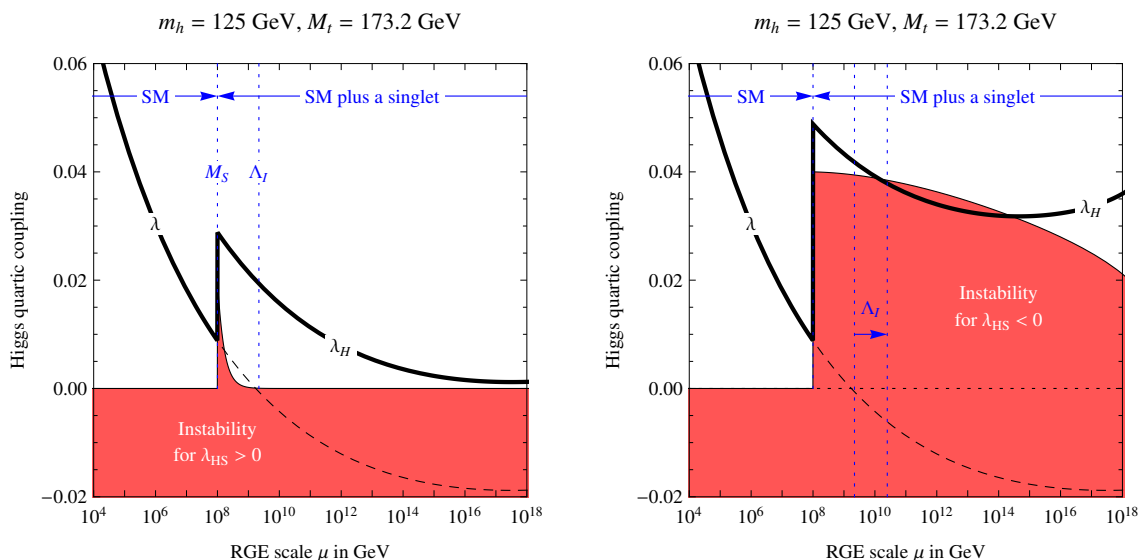


Figure 2. Running of the Higgs quartic coupling in the SM and in the model with a scalar singlet, here assumed to have the mass $M_S = 10^8$ GeV. Left: if $\lambda_{HS} > 0$, thanks to the tree level shift at the singlet mass, the coupling never enters into the instability region, even assuming that singlet contributions to the RG equations are negligible. Right: if $\lambda_{HS} < 0$ the instability can be shifted away or avoided only by singlet contributions to the RG equations.

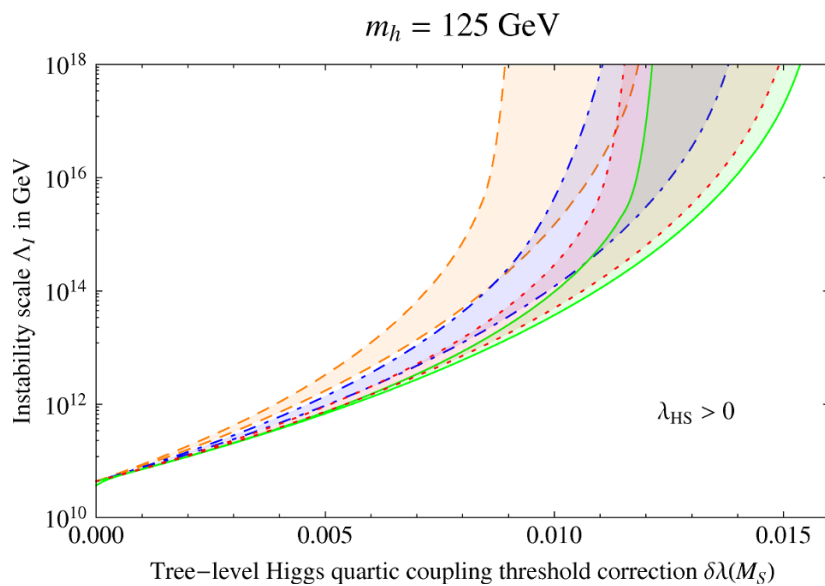


Figure 3. For $m_h = 125$ GeV and $\lambda_{HS} > 0$, bands of the modified instability scale Λ_I versus the threshold correction $\delta\lambda$ to the Higgs quartic coupling due to a scalar singlet with mass $M_S = 10^4, 10^6, 10^8, 10^{10}$ GeV (from left to right). For a fixed M_S value the lowest boundary of the band corresponds to small λ_S, λ_{HS} and the highest boundary to $\lambda_S(M_{Pl}) = 4\pi$.

behaviour of the one-loop potential one can simply include in the running quartic couplings the finite one-loop contributions not captured by RG evolution and impose the stability conditions on these corrected couplings (the shift in the instability scale can be up to one order of magnitude; see [17–19] for further details). The scale at which these one-loop improved running couplings violate the stability condition corresponds then to the field scale at which the potential falls below the EW vacuum. The results are illustrated in figure 3 which shows the new instability scale Λ_I as a function of the threshold shift $\delta\lambda$ for several singlet scalar masses, $M_S = 10^4, 10^6, 10^8, 10^{10}$ GeV below the SM instability scale, $\Lambda_I^{\text{SM}} \simeq 4 \times 10^{10}$ GeV (for $m_h = 125$ GeV, $M_t = 173.2$ GeV and $\alpha_s(M_Z) = 0.1183$). For each value of M_S , there is a band of values for Λ_I due to the freedom in choosing λ_S , once λ_H and $\delta\lambda = \lambda_{HS}^2/\lambda_S$ are fixed. The lower boundary of each band corresponds to $\lambda_S \ll 1$ (and consequently also $\lambda_{SH} \ll 1$ for fixed and small $\delta\lambda$). This case nearly isolates the impact of the tree-level shift on the instability scale (as the running of λ_H above the singlet threshold is SM-like). The upper boundary of each band corresponds to the largest value of λ_S that we allow by requiring $\lambda_S(\mu) < 4\pi$ up to the Planck scale. Large values of λ_S correspond to large λ_{HS} , making the RG effect on λ_H stronger. We conclude that the tree-level shift in λ can have an extremely significant impact in raising the instability scale even for very moderate values of the couplings λ_S and λ_{HS} , and it can easily make the EW vacuum absolutely stable.

2.3 Case $\lambda_{HS} < 0$

In this case V_0 can become negative only for $|S| > w/\sqrt{2}$. In this condition, we can neglect the mass parameters v and w in eq. (2.1) and approximate the potential by keeping only the quartic terms

$$V_0 \approx \lambda_H |H|^4 + \lambda_S |S|^4 + 2\lambda_{HS} |H|^2 |S|^2. \quad (2.12)$$

The stability condition ($V_0 > 0$) is now

$$-\lambda_{HS}(\mu) < \sqrt{\lambda_H(\mu)\lambda_S(\mu)}. \quad (2.13)$$

If this condition is violated at some scale μ_* an instability would develop with

$$|S| > \frac{M_S}{2\sqrt{\lambda_S}}, \quad c_- < \frac{|H|}{|S|} < c_+ \quad c_{\pm}^2 = \frac{-\lambda_{HS}}{\lambda_H} \left(1 \pm \sqrt{1 - \frac{\lambda_H \lambda_S}{\lambda_{HS}^2}} \right) \Big|_{\mu_*}. \quad (2.14)$$

As this determines a direction in field space along which the fields H and S slide towards an unbounded instability, condition (2.13) has to be satisfied at all renormalization scales larger than M_S . Thus the stability condition for negative λ_{HS} is much more constraining than in the case of positive λ_{HS} .

In the case $\lambda_{HS} < 0$, as the stability condition $\lambda_H > \delta\lambda$ must be satisfied at all scales, the tree-level threshold effect is not sufficient to improve the stability. Then one should resort to RG effects to improve the potential stability, as illustrated in figure 2, right panel. By using the RG equations (2.3), we can derive the evolution of the effective Higgs quartic coupling combination $\lambda \equiv \lambda_H - \lambda_{HS}^2/\lambda_S$ above M_S as

$$\frac{d\lambda}{d \ln \mu} = \beta_{\lambda}^{\text{SM}} + \frac{8}{(4\pi)^2} [(\lambda_{HS} - \delta\lambda)^2 + 3\lambda \delta\lambda], \quad (2.15)$$

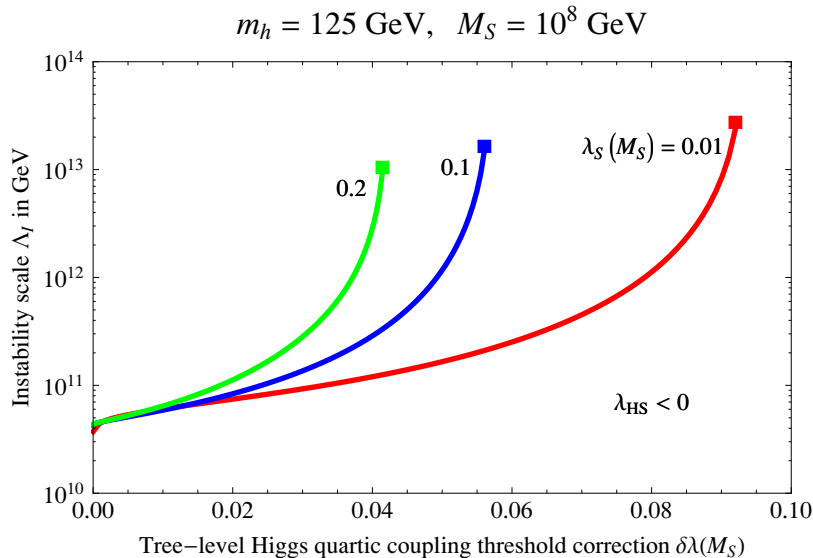


Figure 4. For $m_h = 125$ GeV and $\lambda_{HS} < 0$, the modified instability scale Λ_I versus the threshold correction $\delta\lambda$ to the Higgs quartic coupling due to a scalar singlet with mass $M_S = 10^8$ GeV and $\lambda_S(M_S) = 0.01, 0.1, 0.2$, as indicated.

where $\beta_\lambda^{\text{SM}}$ is the SM beta function for the Higgs quartic coupling and $\delta\lambda = \lambda_{HS}^2/\lambda_S > 0$. We see that the additional term in the beta function of λ is always positive so that RG effects tend to increase the instability scale also in the case $\lambda_{HS} < 0$.

The numerical analysis of the $\lambda_{HS} < 0$ case confirms this expectation. As an illustration, figure 4 shows the instability scale versus the shift $\delta\lambda$ for the same choice of SM parameters as in figure 3 and for the particular case $M_S = 10^8$ GeV with three different values of λ_S as indicated. The end-point of the curves marks the location beyond which (i.e. for larger $\delta\lambda$) the potential becomes completely stable. These end-points occur because λ first decreases as a function of the renormalization scale but, after reaching a minimum, starts increasing at large scales. In comparison with the case $\lambda_{HS} > 0$ (figure 3) we see that larger values of the shift $\delta\lambda$ are now required to have a significant impact on the instability scale.

The stabilization mechanism for $\lambda_{HS} < 0$ we have just described is fragile with respect to possible new contributions to the RGEs that can appear if the singlet couples to other sectors of the theory. In contrast, the stabilization mechanism for $\lambda_{HS} > 0$ is more robust, being based on a tree-level shift. The mechanism is also very effective (because the tree level can be easily large) and economical (because it requires only a heavy scalar singlet). The proposed mechanism can be realized in several situations of physical interest and we now turn to discussing some examples.

3 Examples

In this section, we present some examples in which a heavy scalar singlet has been advocated as a solution to other SM problems such as the smallness of neutrino masses, strong CP problem and the unitarity problem of Higgs inflation.

3.1 See-saw

The see-saw is the conventional mechanism to understand the smallness of neutrino masses. It assumes the existence of heavy right-handed neutrino states N with

$$\mathcal{L}_N = i\bar{N}\not{\partial}N + y_\nu LNH + \frac{M_N}{2}N^2 + \text{h.c.} \quad (3.1)$$

After EW symmetry breaking, nonzero neutrino masses are generated

$$m_\nu = \frac{y_\nu^2 v^2}{M_N}, \quad (3.2)$$

which are naturally small provided $M_N \gg v$.

The impact of the see-saw mechanism on the stability of the Higgs potential has been discussed in the past [19, 29, 30]. The right-handed neutrino Yukawa couplings can play a destabilizing role on β_λ similar to that of the top Yukawa coupling. As they scale like $y_\nu^2 \sim m_\nu M_N$, they become sizable for large M_N and are dangerous for stability only if $M_N \gtrsim 10^{13}$ GeV. For lower M_N the new Yukawas will have a negligible effect on stability.

We do not know what originates the large right-handed Majorana mass, but the simplest idea is to assume that the right-handed neutrinos are coupled to a scalar field carrying two units of lepton number and having a large vev,

$$\frac{\kappa}{2} S N^2 + \text{h.c.} \quad (3.3)$$

The vev of S , which sets the scale of the Majorana mass, $M_N = \kappa \langle S \rangle$ does not necessarily lead to a Goldstone boson because in unified models $B - L$ is usually a gauge symmetry. In this well-motivated realization of the see-saw the scalar field S could naturally reestablish stability of the electroweak vacuum. In this setting the role of the singlet scalar is therefore double. Upon taking a large vev and decoupling, it leaves behind two effects: a Weinberg dimension-5 operator that gives neutrinos a nonzero mass and a threshold effect on the Higgs Yukawa coupling that solves the stability problem of the Higgs potential, as long as the mass of S is smaller than Λ_I .

A lower bound on the lightest right-handed neutrino mass M_1 is derived by assuming that the cosmic baryon asymmetry is explained by thermal leptogenesis.³ In this case, one obtains the bounds [31, 32]:

- $M_1 > 2 \times 10^9$ GeV, if the initial right-handed neutrino density vanishes at high temperature.
- $M_1 > 5 \times 10^8$ GeV, if the initial right-handed neutrino density is thermal at high temperature.
- $M_1 > 2 \times 10^7$ GeV, if the initial right-handed neutrino density dominates the universe at high temperature.

³If neutrinos are nearly degenerate in mass, thermal leptogenesis could operate at much smaller values of M_1 and the following lower bounds do not apply.

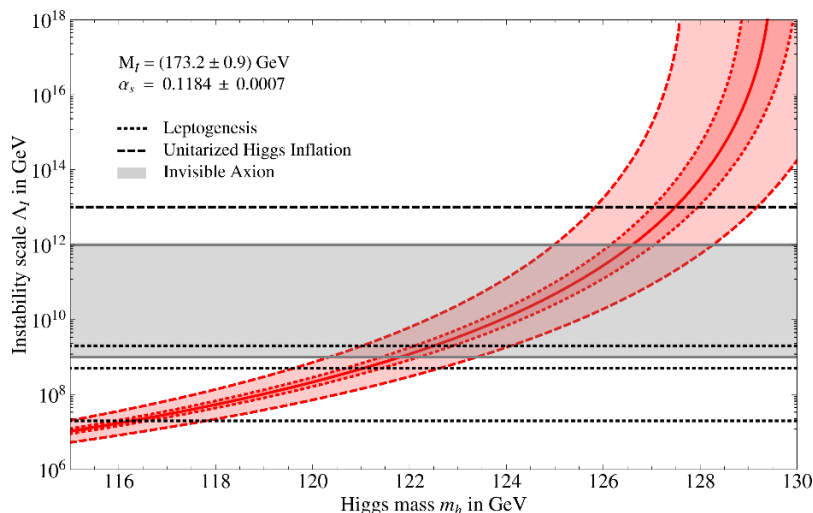


Figure 5. The SM instability scale Λ_I increasing as a function of the Higgs mass. The central line corresponds to $M_t = 173.2 \text{ GeV}$ and $\alpha_s(M_Z) = 0.1184$ and the side-bands to 1 sigma deviations as indicated (with the larger deviation for the top mass uncertainty). The horizontal lines and band mark several values or ranges of interest for Λ_I . The three lowest lines are relevant for the see-saw case and correspond to lower limits on the mass M_1 of the lightest right-handed neutrino N_1 coming from thermal leptogenesis. The bound depends on the initial density ρ_{N_1} : $M_1 > 2 \times 10^7 \text{ GeV}$ for $\rho_{N_1} \sim 0$; $M_1 > 5 \times 10^8 \text{ GeV}$ for thermal ρ_{N_1} and $M_1 > 2 \times 10^9 \text{ GeV}$ for ρ_{N_1} dominating the universe; the shaded band shows the range of singlet masses $10^9 - 10^{12} \text{ GeV}$ relevant for the axion case; and the upper line is the singlet mass 10^{13} GeV relevant for the unitarized Higgs inflation case.

Assuming that the mass of S is equal or smaller than its vev, we can infer the range of Higgs masses for which the scalar setting the see-saw scale could cure any instability of the potential. From figure 5 we can easily read off such Higgs masses.⁴ At 90% CL in M_t and α_s , we find that the see-saw singlet can potentially eliminate the instability of the EW vacuum, as long as $m_h > 120 \text{ GeV}$ (leptogenesis with vanishing initial right-handed neutrino density), $m_h > 119 \text{ GeV}$ (leptogenesis with thermal initial right-handed neutrino density), or $m_h > 115 \text{ GeV}$ (leptogenesis with dominant initial right-handed neutrino density). These limits are compatible with the region of Higgs masses suggested by preliminary ATLAS and CMS data, $m_h = 124 - 126 \text{ GeV}$. The instability scale is raised according to the mechanism discussed in the previous section as long as $M_N \lesssim 10^{13} \text{ GeV}$, since the RG effects of y_ν couplings can be neglected.

We conclude that this simple scenario could comfortably account for the cosmological baryon asymmetry through leptogenesis, for the smallness of neutrino masses and cure the Higgs potential instability. The only drawback of this (beautifully simple but depressing) scenario is that it makes plainly explicit the hierarchy problem: a large singlet vev also gives a tree-level contribution to the Higgs mass term in the Lagrangian which requires a

⁴Figure 5 gives the SM instability scale as a function of m_h as calculated in [19]. Besides the 1-sigma error bands shown, associated with the experimental uncertainties in M_t and α_s , a (conservative) estimate of the higher order radiative corrections not included in the calculation results in a theoretical uncertainty on m_h of $\pm 3 \text{ GeV}$ [19].

large fine-tuning. (This is in contrast with the scenario without the singlet, in which Higgs mass corrections appear at one-loop and are dangerous only when $M_N > 10^7$ GeV [33, 34]).

3.2 Invisible axion

The scalar field S can also be identified with the invisible axion. DFSZ axion models [35, 36] use the SM fermion content and a two-Higgs doublet structure H_u and H_d augmented by a complex scalar S , neutral under SM gauge interactions, with a coupling $\lambda_{HS} S^2 H_u H_d + \text{h.c.}$, analogous to the one in eq. (2.1). This interaction is crucial for the axion mechanism, because it transmits the breaking of the global symmetry triggered by the vev of S to the Higgs sector. One or both of the Higgs doublets can remain light, at the electroweak scale. The presence of an instability is subject to the details of the two-Higgs potential [37–40], but this does not change the essential point. Independently of the model implementation, the field S containing the invisible axion $a = \sqrt{2}\text{Im} S$ with large decay constant $f_a \approx \langle S \rangle$ is a perfect candidate to play the role of the field S in our Higgs stabilization mechanism.

KSVZ axion models [41, 42] use a single Higgs doublet and a complex scalar S coupled to new heavy vector-like fermions Ψ . The Dirac mass term $M \bar{\Psi}\Psi$ is forbidden by imposing the symmetry

$$\Psi_L \rightarrow -\Psi_L, \quad \Psi_R \rightarrow \Psi_R, \quad S \rightarrow -S. \tag{3.4}$$

Then the mass of the heavy fermions comes only from the vev of S :

$$\lambda_\Psi S \bar{\Psi}\Psi + V(H, S). \tag{3.5}$$

The resulting model has a spontaneously broken U(1) global symmetry

$$\Psi \rightarrow e^{i\gamma_5 \alpha} \Psi, \quad S \rightarrow e^{-2i\alpha} S \tag{3.6}$$

which gives rise to a light axion $a = \sqrt{2}\text{Im} S$ with large decay constant $f_a \approx \langle S \rangle$. The scalar potential of the theory is precisely of the form in eq. (2.1), although the coupling λ_{HS} plays no role in axion phenomenology because both $|S|^2$ and $|H|^2$ are separately invariant under the global symmetry.

The decay constant of the axion is allowed to lie in the range

$$10^9 \text{ GeV} < f_a < 10^{12} \text{ GeV} . \tag{3.7}$$

The lower bound comes from non-observation of axion emission from stars and supernovæ. The upper bound comes from requiring that the axion dark matter density

$$\Omega_a \approx 0.15 \left(\frac{f_a}{10^{12} \text{ GeV}} \right)^{7/6} \left(\frac{a_*}{f_a} \right)^2 \tag{3.8}$$

does not exceed the observed value $\Omega_{\text{DM}} \approx 0.23$ under the assumption that the axion vev a_* in the early universe was of the order of f_a [43–45]. The resulting range of singlet mass M_S , which we can roughly take to be the same as the range for f_a in eq. (3.7), overlaps with the range that can stabilize the SM Higgs potential from $m_h \gtrsim 119$ GeV (for $M_S \sim 10^9$ GeV) to $m_h \gtrsim 124$ GeV (for $M_S \sim 10^{12}$ GeV), as can be inferred from figure 5.

3.3 Unitarized Higgs inflation

In the original proposal for SM Higgs inflation [46], a large non-minimal coupling $\xi \sim 10^4$ of the Higgs boson to gravity, $\xi H^\dagger H R$, drives inflation at large Higgs field values. However, due to the same large non-minimal coupling, the unitarity cutoff [47–50] becomes $\Lambda_{\text{SM}} = M_{\text{Pl}}/\xi$ which is not only much lower than the Planck scale but also lower than the field scales at which the inflationary plateau develops. Although the unitarity cutoff during inflation could be larger than the one in the vacuum [51, 52], the perturbative expansion is still questionable and the very existence of an inflationary plateau beyond Λ_{SM} is jeopardized. In addition to this unitarity problem, in all the Higgs mass range allowed by ATLAS and CMS, the potential instability that develops at large field values could also destroy the flat plateau induced by gravitational effects [19].

To fix the unitarity problem, extra dynamical degrees of freedom are required to restore unitarity without ruining the flat plateau. A UV complete model with a real singlet scalar of sigma-model type was proposed in [53] and it is natural to ask if this singlet could also solve simultaneously the instability problem. The Jordan-frame Lagrangian of the model is

$$\begin{aligned} \frac{\mathcal{L}_J}{\sqrt{-g_J}} = & \frac{1}{2} \left(M^2 + \xi \sigma^2 + 2\zeta H^\dagger H \right) R - \frac{1}{2} (\partial_\mu \sigma)^2 - |D_\mu H|^2 \\ & - \frac{1}{4} \lambda_\sigma \left(\sigma^2 - \Lambda^2 + 2 \frac{\lambda_{H\sigma}}{\lambda_\sigma} H^\dagger H \right)^2 - \left(\lambda_H - \frac{\lambda_{H\sigma}^2}{\lambda_\sigma} \right) \left(H^\dagger H - \frac{v^2}{2} \right)^2, \end{aligned} \quad (3.9)$$

where M, Λ and v are mass parameters with $v \ll M, \Lambda$ (so that the σ field is heavy) and ξ, ζ are positive non-minimal couplings with $\xi \gg \zeta$. The scalar potential derived from eq. (3.9) falls in the class of models considered in this paper, once we identify σ with S (although the singlet now is real). In what follows we will also use the effective quartic coupling defined as

$$\lambda \equiv \lambda_H - \frac{\lambda_{H\sigma}^2}{\lambda_\sigma}. \quad (3.10)$$

The large nonzero vev of σ , $\langle \sigma \rangle \simeq \Lambda$, is crucial to make the unitarity cutoff Λ_{UV} larger. It is straightforward to find that

$$\Lambda_{UV} = \left(1 + 6r\xi \right) \frac{M_{\text{Pl}}}{\xi}, \quad (3.11)$$

where the Planck mass is now $M_{\text{Pl}}^2 = M^2 + \xi \Lambda^2$, and we measure the contribution of the σ vev by the ratio $r = \xi \Lambda^2 / M_{\text{Pl}}^2$, which in general can take values from 0 to 1. One can see how for a negligible vev ($r \rightarrow 0$, as was the case of the SM Higgs inflation case) the cutoff is M_{Pl}/ξ while it is pushed up to $r M_{\text{Pl}}$ for moderate values of $r \gtrsim 1/\xi$.

As described in [53], in this scenario the sigma field dominates inflation due to its large non-minimal coupling to R while the Higgs field follows the sigma field along a flat direction, provided $\lambda_{H\sigma} < 0$. In fact, in unitary gauge with $H^T = (0, h)/\sqrt{2}$, for $\sigma, h \gg M/\sqrt{\xi}, \Lambda$, slow-roll inflation takes place along $h \approx \sqrt{(-\lambda_{H\sigma}/\lambda_H)} \sigma$, while the direction orthogonal to the inflaton gets mass of order $\sqrt{-2\lambda_{H\sigma}} M_{\text{Pl}}/\sqrt{\xi}$ (identified with a larger background-dependent cutoff in the Higgs inflationary regime [53]). Here, a negative $\lambda_{H\sigma}$ is crucial for the sigma model to reproduce the Higgs inflation model below the sigma mass and for

the Higgs boson to participate in inflaton dynamics beyond Λ . Then, the Einstein-frame potential along the inflaton field direction $\chi = \sqrt{6}M_{\text{Pl}} \ln(\sqrt{\xi}\sigma/M_{\text{Pl}})$, is given by

$$V(\chi) \simeq \frac{\lambda_\sigma \lambda}{4\lambda_H} \frac{M_{\text{Pl}}^4}{\xi^2} \left(1 - 2e^{-2\chi/\sqrt{6}M_{\text{Pl}}}\right). \quad (3.12)$$

In order to match the COBE normalization of the power spectrum, we obtain the following condition on the parameters,

$$\frac{\sqrt{\lambda_\sigma}}{\xi} = 2 \times 10^{-5} \sqrt{\frac{\lambda_H}{\lambda}}, \quad (3.13)$$

which explains the need of having $\xi \gg 1$.

The mass of the σ field in the vacuum, on the other hand, is given by

$$M_{\bar{\sigma}}^2 = \lambda_\sigma \frac{2rM_{\text{Pl}}^2}{(1+6r\xi)\xi} \simeq \lambda_\sigma \frac{M_{\text{Pl}}^2}{3\xi^2}, \quad (3.14)$$

(where $\bar{\sigma}$ denotes the canonically normalized field, and we have required $r \gtrsim 1/\xi$ for the last expression), which is r -independent and of the same order of Λ_{SM} , the UV cutoff for $r = 0$. The σ field could be lighter than Λ_{SM} at the cost of reducing r below $1/\xi$, but then the unitarity cutoff (3.11) would decrease, back to its SM value. In principle, one could also lower the sigma mass below the unitarity cutoff by choosing a small value of λ_σ . However, the COBE constraint (3.13) precisely fixes the sigma mass in the vacuum to be (taking λ_H/λ of order 1)

$$M_{\bar{\sigma}} \approx 10^{13} \text{ GeV}. \quad (3.15)$$

As we would like the singlet to be lighter than Λ_I in order to cure that instability problem, we are forced to a region of Higgs masses for which $\Lambda_I > 10^{13} \text{ GeV}$. From figure 5 we see that this requires $m_h > 125 \text{ GeV}$ (at 90% CL in M_t and α_s), which is marginally compatible with the ATLAS and CMS hint of $m_h \sim 124\text{--}126 \text{ GeV}$.

As explained above, the scenario requires $\lambda_{H\sigma} < 0$, so that the stability condition to be satisfied (at all scales) is $\lambda = \lambda_H - \lambda_{H\sigma}^2/\lambda_\sigma > 0$. As can be seen from eq. (3.12) this condition amounts to requiring a positive vacuum energy during inflation. From our general analysis we know that in this case the tree-level threshold effect alone would not improve the vacuum stability and RG effects above the sigma mass are necessary for stabilizing the potential. The instability scale Λ_I should be pushed up, if not all the way to the Planck scale, at least sufficiently high to maintain the inflationary plateau stable in an interval that provides enough e-folds of expansion. Such interval is approximately $\Lambda < \sigma < 10\Lambda$, so that it would be sufficient to increase Λ_I by one or two orders of magnitude.

Concerning the RG evolution of couplings above the sigma-field threshold, we follow [54, 55] and take into account the effects of a non-minimal coupling to gravity of the

sigma and Higgs fields through suppression factors in the RGEs, which become:

$$\begin{aligned}
 (4\pi)^2 \frac{d\lambda_H}{d\ln\mu} &= (12y_t^2 - 3g'^2 - 9g^2)\lambda_H - 6y_t^4 + \frac{3}{8} [2g^4 + (g'^2 + g^2)^2] \\
 &\quad + (18c_h^2 + 6)\lambda_H^2 + 2c_\sigma^2\lambda_{H\sigma}^2, \\
 (4\pi)^2 \frac{d\lambda_{H\sigma}}{d\ln\mu} &= \frac{1}{2}\lambda_{H\sigma} [12y_t^2 - 3g'^2 - 9g^2 + 12(1 + c_h^2)\lambda_H + 12c_\sigma^2\lambda_\sigma] + 8c_h c_\sigma \lambda_{H\sigma}^2, \\
 (4\pi)^2 \frac{d\lambda_\sigma}{d\ln\mu} &= 2(3 + c_h^2)\lambda_{H\sigma}^2 + 18c_\sigma^2\lambda_\sigma^2.
 \end{aligned}
 \tag{3.16}$$

(Numerical factors are now appropriate for a real scalar σ , and therefore these equations do not reduce to (2.3) in the limit $c_{\sigma,h} \rightarrow 1$). The suppression factors, c_σ and c_h , are given in terms of the Weyl rescaling factor $\Omega^2 = (M^2/M_{\text{Pl}}^2)(1 + \xi\sigma^2/M^2)$ as $c_\sigma = \Omega^{-2}(\partial\chi/\partial\sigma)^{-2} \simeq (1 + M^2/\xi\sigma^2)/(6\xi) \ll 1$ and $c_h = \Omega^{-2}(\partial\phi/\partial h)^{-2} = 1$, where χ and ϕ are canonical fields. Therefore, the running of λ_H is SM-like, as loops containing the sigma scalar are suppressed for $c_\sigma \ll 1$ and Higgs loops are the same as in the SM. Then, we use the above RG equations to get the corresponding equation for $\lambda \equiv \lambda_H - \lambda_{H\sigma}^2/\lambda_\sigma$, with $\delta\lambda \equiv \lambda_{H\sigma}^2/\lambda_\sigma$, as

$$\frac{d\lambda}{d\ln\mu} \approx \beta_\lambda^{\text{SM}} + \frac{8}{(4\pi)^2} (3\lambda + \delta\lambda) \delta\lambda.
 \tag{3.17}$$

Consequently, as compared to the SM, the sigma scalar gives an extra positive contribution in β_λ . Therefore, the loop effects of the sigma scalar couplings can make the instability scale larger than the SM one for a sizable value of $\delta\lambda$ at the sigma mass scale. In conclusion, unitarized Higgs inflation could still be a viable possibility without an instability problem.

4 Conclusions

The recent LHC Higgs searches have narrowed down the Higgs mass to a range which most likely leads to a large-field instability of the potential, in the context of the SM. One can view this result as an opportunity to learn new information about the early universe through collider experiments [21], or as the need for a UV modification of the SM.

The lack of early discoveries at the LHC is a motivation for considering minimal, albeit technically unnatural, theoretical descriptions of the particle world. In this paper we have analyzed the most minimal UV modification that eliminates the EW vacuum instability, adding one singlet scalar field to the SM degrees of freedom. As the theory is renormalized to low energies and the new singlet is integrated out, the Higgs quartic coupling is shifted downwards, making the Higgs mass smaller. As a result, the physical Higgs mass lies inside the region of instability in the pure SM, although there is no actual instability in the complete UV theory. From this point of view, the instability for m_h in the allowed range 115–127 GeV is just a mirage.

The proposed stabilization mechanism relies on a threshold correction to the Higgs quartic coupling, whose size is independent of the singlet mass. The necessary ingredients are a singlet self-quartic coupling (λ_S), a mixed quartic coupling with the Higgs (λ_{HS}) and a non-zero vev for the singlet. The mechanism can be operative even for a very heavy

singlet, as long as its mass is smaller than the instability scale Λ_I . Occurring at tree level, the effect is sizable and robust.

The analysis of the effect involves some subtleties because, at the singlet threshold, both the Higgs quartic coupling and the stability conditions are shifted by the same amount. We have shown that, when λ_{HS} is positive, the stability conditions become weaker as the field value is increased above the singlet mass. In this situation, the tree-level contribution is very effective in stabilizing the potential. On the other hand, for negative λ_{HS} , the shifts in the Higgs quartic coupling and in the stability condition essentially cancel out, and one has to rely on RG effects. These can help the stabilization, but larger singlet couplings are needed to obtain the desired effect.

The minimal modification of the SM that we have considered, with the addition of one singlet scalar, has motivations that are independent of the stability of the EW vacuum. We have investigated three examples. The new singlet can set the scale of the right-handed neutrino mass in the sea-saw mechanism; or it could play the role of the invisible axion; or finally it could unitarize models with large gravitational non-minimal couplings of the Higgs field invoked for inflationary dynamics. In each case, we were able to define the range of Higgs masses for which the corresponding singlet could also be used to stabilize the SM Higgs potential. We find that stabilization is possible for any Higgs mass allowed by the current LHC limits, in the case of leptogenesis with dominant initial right-handed neutrino density; for $m_h > 119$ GeV, in the case of leptogenesis with thermal initial right-handed neutrino density or in the case of the invisible axion; for $m_h > 120$ GeV, in the case of leptogenesis with vanishing initial right-handed neutrino density; for $m_h > 125$ GeV in the case of unitarized Higgs inflation.

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