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Stalagmite-inferred variability of the Asian summer monsoon during the penultimate glacial–interglacial period

T.-Y. Li^{1,2,3}, C.-C. Shen³, L.-J. Huang³, X.-Y. Jiang⁴, X.-L. Yang¹, H.-S. Mii⁵, S.-Y. Lee⁶, and L. Lo³

¹Key Laboratory of Eco-environments in Three Gorges Reservoir Region, Ministry of Education,

School of Geographical Sciences, Southwest University, Chongqing 400715, China

²State Key Laboratory of Loess and Quaternary Geology, Institute of Earth Environment, CAS, Xi'an 710075, China

³High-Precision Mass Spectrometry and Environment Change Laboratory (HISPEC), Department of Geosciences,

National Taiwan University, Taipei 10617, Taiwan, ROC

⁴College of Geographical Science, Fujian Normal University, Fuzhou 350007, China

⁵Department of Earth Sciences, National Taiwan Normal University, Taipei 11677, Taiwan, ROC

⁶Research Center for Environmental Changes, Academia Sinica, Taipei 11529, Taiwan, ROC

Correspondence to: C.-C. Shen (river@ntu.edu.tw)

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Abstract. The orbital-timescale dynamics of the Quaternary Asian summer monsoons (ASM) are frequently attributed to precession-dominated northern hemispheric summer insolation. However, this long-term continuous ASM variability is inferred primarily from oxygen isotope records of stalagmites, mainly from Sanbao cave in mainland China, and may not provide a comprehensive picture of ASM evolution. A new spliced stalagmite oxygen isotope record from Yangkou cave tracks summer monsoon precipitation variation from 124 to 206 thousand years ago in Chongqing, southwest China. Our Yangkou record supports that the evolution of ASM was dominated by the North Hemisphere solar insolation on orbital timescales. When superimposed on the Sanbao record, the precipitation time series referred from Yangkou cave stalagmites supports the strong ASM periods at marine isotope stages (MIS) 6.3, 6.5, and 7.1 and weak ASM intervals at MIS 6.2, 6.4, and 7.0. This consistency confirms that ASM events affected most of mainland China. Except for the solar insolation forcing, the large amplitude of minimum δ^{18} O values in Yangkou record during glacial period, such as MIS 6.5, could stem from the enhanced prevailing Pacific trade wind and/or continental shelf exposure in the Indo-Pacific warm pool.

1 Introduction

Climate in East Asia, the most densely populated region in the world, is profoundly influenced by the Asian monsoon (AM), which includes the Indian monsoon and East Asian monsoon sub-systems. Asian summer monsoon (ASM) precipitation strongly governs regional vegetation, agriculture, culture, and economies (e.g., Cheng et al., 2012a), and even affected the stability of Chinese dynastic rule (Zhang et al., 2008; Tan et al., 2011).

Our current understanding of ASM variation over the past 500 kyr BP (before AD 1950) has been reconstructed using oxygen isotope records of Chinese stalagmites (Wang et al., 2008; Cheng et al., 2012b) with the advantages of absolute and high-precision chronologies (e.g., Cheng et al., 2000, 2013; Shen et al., 2002, 2012). Stalagmite-inferred orbital-scale ASM intensity closely follows the change in precession-dominated northern hemispheric (NH) summer insolation (NHSI) (Wang et al., 2008; Cheng et al., 2012b). However, these 100s kyr records were mainly from a single cave, namely Sanbao cave, located in Hubei Province, China (Fig. 1; Wang et al., 2008; Cheng et al., 2012b). Utilizing only one site leads to uncertainties in the spatial extent of Quaternary ASM evolution. These uncertainties stem from differences in local or regional climatic and environmental conditions (Lachniet, 2009), hydrological variability



Figure 1. (A) Map of precipitation anomaly $(mm day^{-1})$ in June, July, and August (JJA) of AD 1998–2000 during a La Niña event from July 1998 to April 2001 (http://www.cpc.ncep.noaa. gov/products/analysis_monitoring/ensostuff/ensoyears.shtml) compared with the averaged state of JJA from 1980 to 2010. Triangle symbols denote cave sites of Yangkou (this study), Sanbao (Wang et al., 2008), and Hulu (Cheng et al., 2006). Solid circles indicate marine sediment cores of ODP806B and TR163-19 (Lea et al., 2000). Arrows depict present ground wind directions of the ISM and EASM and also trade wind in the equatorial Pacific. Summer precipitation intensity in eastern and southern China was enhanced during the La Niña event. (B) An enlarged map of precipitation anomaly with cave sites of Yangkou, Sanbao, and Hulu.

of monsoonal sources (e.g., Dayem et al., 2010; Clemens et al., 2010; Pausata et al., 2011), and interactions between climatic subsystems (e.g., Maher and Thompson, 2012; Tan, 2014).

Sanbao records, for example, show distinct ASM events at marine isotope stages (MIS) 6.3 and 6.5 during the penultimate glacial time and a weaker summer monsoon during the penultimate glacial maximum (PGM) at MIS 6.2 (Fig. 1 of Wang et al., 2008). To clarify whether this combination of weak PGM ASM intensities and strong ASM events during the penultimate glacial–interglacial (G–IG) period are local effects, we built an integrated stalagmite oxygen stable isotope record from Yangkou cave, Chongqing, China, covering 124–206 kyr BP (Fig. 1). Through comparison with records from other Chinese caves (Cheng et al., 2006, 2009; Wang et al., 2008) confirms the fidelity of Sanbao cave-inferred ASM intensities.

2 Material and methods

2.1 Regional settings and samples

Stalagmites were collected from Yangkou cave $(29^{\circ}02' \text{ N}, 107^{\circ}11' \text{ E}; altitude: 2140 m; length: 2245 m), located at Jinfo Mountain National Park, Chongqing City, southwestern China (Fig. 1) during two field trips in October 2010 and July 2011. The cave, developed in Permian limestone bedrock, is$



Figure 2. Photographs of the five stalagmites collected from Yangkou cave. Brown dashed curves show hiatuses. Straight lines represent subsampling routes for oxygen isotope measurement. Yellow curves denote drilled subsamples for U-Th dating. White dots are the subsamples collected for Hendy test (Hendy, 1971).

400 km southwest of Sanbao cave $(31^{\circ}40' \text{ N}, 110^{\circ}26' \text{ E})$ in Hubei Province (Wang et al., 2008). The cave air temperature is 7.5 °C and the average relative humidity is > 80 % (October 2011–October 2013). The regional climate is influenced by both the Indian summer monsoon (ISM) and East Asian summer monsoon (EASM). Annual rainfall is 1400– 1500 mm, 83 % from April to October (Zhang et al., 1998). Five stalagmites, YK05, YK12, YK23, YK47, and YK61, which formed within a time interval of 124–206 kyr BP, were halved and polished for U-Th dating and oxygen stable isotope analysis.

2.2 U-Th dating

Chemistry and instrumental analysis were conducted in the High-Precision Mass Spectrometry and Environment Change Laboratory (HISPEC), Department of Geosciences, National Taiwan University. Fifty three powdered subsamples, 60-80 mg each, were drilled from the polished surface along the deposit lamina of the five stalagmites (Fig. 2, Table 1), on a class-100 bench in a class-10000 subsampling room. U-Th chemistry (Shen et al., 2003) was performed in a class-10 000 clean room with independent class-100 benches and hoods (Shen et al., 2008). A multi-collector inductively coupled plasma mass spectrometer (MC-ICP-MS), Thermo Fisher Neptune with secondary electron multiplier protocols was used for the determination of U-Th isotopic contents and compositions (Shen et al., 2012). The decay constants used are $9.1577 \times 10^{-6} \text{ yr}^{-1}$ for ²³⁰Th, $2.8263 \times 10^{-6} \text{ yr}^{-1}$ for 234 U (Cheng et al., 2000), and 1.55125×10^{-10} yr⁻¹ for ²³⁸U (Jaffey et al., 1971). All errors of U-Th isotopic data and U-Th dates are two standard deviations (2σ) unless otherwise noted. Age (before AD 1950) corrections were made using an 230 Th / 232 Th atomic ratio of 4 ± 2 ppm, which are the values for material at secular equilibrium with the crustal ²³²Th/²³⁸U value of 3.8 (Taylor and McLennan, 1995) and an arbitrary uncertainty of 50 %.

Subsample ID		Depth	238U	²³² Th	δ^{234} U	$[^{230}\text{Th}/^{238}\text{II}]$	$[^{230}\text{Th}/^{232}\text{Th}]$	Age (kyr)	Age (kyr BP)	δ ²³⁴ Uinitial
		Depui			19			rige (kyi)	11ge (ky1, D1)	i ih
		(mm)	(ppb)	(ppt)	measured"	activity	(ppm) ^a	uncorrected	corrected	corrected
Stalagmite: YK5	VV5 01	2.0	8720 ± 12	5520171	215.9 2.1	1 0102 + 0 0024	265 626 + 2445	170 7 1 2	170.6 ± 1.2	259 5 1 2 7
	1K3-01	5.0	$8/30 \pm 13$	555.0 ± 7.1	213.8 ± 2.1	1.0192 ± 0.0024	203020 ± 3443	$1/9.7 \pm 1.5$	$1/9.0 \pm 1.5$	558.5 ± 5.7
	YK5-02	24.0	7335 ± 14	263.1 ± 7.1	218.4 ± 2.7	1.0235 ± 0.0027	471128 ± 12563	180.4 ± 1.6	180.4 ± 1.6	363.6 ± 4.8
	YK5-03	57.0	4322.4 ± 7.6	5997 ± 17	192.9 ± 2.3	1.0002 ± 0.0024	11903 ± 39	181.2 ± 1.4	181.1 ± 1.4	321.9 ± 4.1
	YK5-04	79.0	5041 ± 10	500.2 ± 5.7	1877 + 29	0.9997 ± 0.0026	166348 ± 1928	1832 ± 17	1832 ± 17	315.1 ± 5.0
	VV5 05	88.0	5720 6 1 0 4	256.1 ± 5.1	101.1 ± 2.9	0.0096 ± 0.0020	265267 ± 2914	103.2 ± 1.7 194.2 ± 1.6	103.2 ± 1.7 104.1 ± 1.6	210.6 ± 4.2
	IK3-05	88.0	5729.0 ± 9.4	550.1 ± 5.1	164.0 ± 2.4	0.9980 ± 0.0027	203207 ± 3814	164.2 ± 1.0	164.1 ± 1.0	510.0 ± 4.2
	YK5-06	103.0	5375.3 ± 9.9	593.2 ± 5.0	202.1 ± 2.6	1.0161 ± 0.0022	152028 ± 1290	184.2 ± 1.5	184.1 ± 1.5	340.1 ± 4.7
	YK5-07	128.0	4986.2 ± 8.8	137.6 ± 5.8	201.6 ± 2.3	1.0175 ± 0.0023	608876 ± 25827	185.1 ± 1.4	185.0 ± 1.4	340.0 ± 4.1
	YK5-08	149.0	6076 ± 14	269.0 ± 5.2	205.0 ± 3.0	1.0259 ± 0.0028	382639 ± 7471	187.2 ± 1.8	187.2 ± 1.8	348.1 ± 5.3
	VK5 00	177.0	0070 ± 14	1102.7 ± 7.2	205.0 ± 5.0	1.0257 ± 0.0020	12((00 + 000	107.2 ± 1.0	107.2 ± 1.0	$3+0.1 \pm 3.5$
	1K3-09	177.0	8808 ± 11	1103.7 ± 7.2	215.0 ± 1.9	$1.03/4 \pm 0.0016$	130099 ± 889	$18/.9 \pm 1.1$	$18/.8 \pm 1.1$	303.7 ± 3.3
	YK5-10	188.0	12100 ± 19	168.3 ± 6.1	210.0 ± 2.5	1.0368 ± 0.0027	1230671 ± 44610	189.9 ± 1.7	189.8 ± 1.7	359.2 ± 4.7
Stalagmite: YK12	10110.01	2.4	(2(2))	2005 1 24	200 4 1 2	0.0000 1.0.0015	05540 + 464	100 56 1 0 16	122 (0 1 0 1(151 0 1 1 0
	YK12-01	3.6	6262.6 ± 4.1	3895 ± 24	309.6 ± 1.2	0.9620 ± 0.0015	25540 ± 164	133.76 ± 0.46	133.69 ± 0.46	451.8 ± 1.9
	YK12-02	10.5	5016.7 ± 2.5	12393 ± 25	296.1 ± 1.2	0.9590 ± 0.0017	6410 ± 17	135.88 ± 0.51	135.78 ± 0.51	434.7 ± 1.8
	YK12-03	21.5	6384.1 ± 3.6	1050 ± 21	296.2 ± 1.1	0.9796 ± 0.0014	98334 ± 1947	141.43 ± 0.46	141.36 ± 0.46	441.8 ± 1.7
	VK12-04	40.0	56753 ± 58	9675 ± 32	273.0 ± 1.6	0.9792 ± 0.0017	9483 ± 34	147.07 ± 0.67	146.98 ± 0.67	4137 ± 26
	1K12-04	40.0	12214 + 12	1400 + 21	275.0 ± 1.0	0.9792 ± 0.0017	145202 + 2004	152.00 ± 0.07	140.90 ± 0.07	$+15.7 \pm 2.0$
	YK12-05	57.5	13314 ± 13	1488 ± 21	259.4 ± 1.5	0.9840 ± 0.0015	145382 ± 2094	152.20 ± 0.62	152.14 ± 0.62	398.9 ± 2.4
	YK12-06	78.0	1746.6 ± 5.5	1425 ± 24	253.54 ± 0.90	0.9852 ± 0.0013	134061 ± 2272	154.30 ± 0.49	154.24 ± 0.49	392.1 ± 1.5
	YK12-07	80.0	8830.3 ± 5.3	38573 ± 98	212.8 ± 1.2	0.9796 ± 0.0027	3702 ± 14	165.3 ± 1.1	165.1 ± 1.1	339.4 ± 2.2
	VK12-08	92.0	71066 ± 36	7546 ± 25	100.70 ± 0.89	0.9823 ± 0.0014	15274 ± 55	171.08 ± 0.64	170.99 ± 0.64	323.9 ± 1.5
	1K12-00	101.0	7100.0±5.0	1402 + 22	1)).70 ± 0.0)	0.9025 ± 0.0014	15274 ± 55	171.00 ± 0.04	176.77 ± 0.04	323.7 ± 1.3
	YK12-09	101.0	9513.1 ± 6.5	4483 ± 23	203.4 ± 1.1	$0.99/6 \pm 0.0013$	34954 ± 182	$1/5.80 \pm 0.72$	$1/5.73 \pm 0.72$	334.3 ± 2.0
	YK12-10	105.0	5118.6 ± 6.7	2378 ± 21	185.4 ± 1.9	0.9924 ± 0.0018	35265 ± 317	181.0 ± 1.1	180.9 ± 1.1	309.3 ± 3.3
	YK12-11	109.5	6109.1 ± 3.8	572 ± 18	178.4 ± 1.2	0.9875 ± 0.0013	174125 ± 5633	181.93 ± 0.77	181.87 ± 0.77	298.4 ± 2.1
e: YK23	YK23-01	2.4	2893.2 ± 2.3	13899 ± 26	102.8 ± 1.5	0.8935 ± 0.0018	3070.9 ± 8.0	172.8 ± 1.0	172.6 ± 1.0	167.6 ± 2.4
	YK23-02	9.6	2608.9 ± 1.7	13210 ± 23	99.6 ± 1.1	0.9008 ± 0.0016	2937.3 ± 7.1	177.70 ± 0.95	177.53 ± 0.95	164.5 ± 1.9
		Histus								
	VIZ22 02	11.2	2705 2 1 1 2	1270 17	50 55 1 0 01	0.0700 + 0.0016	29 (92 + 255	197.2 1.0	107.2 1.0	101.1 + 1.6
	YK25-05	11.2	$2/05.2 \pm 1.5$	$13/0 \pm 1/$	59.55 ± 0.91	$0.8/99 \pm 0.0016$	28083 ± 333	$18/.3 \pm 1.0$	187.3 ± 1.0	101.1 ± 1.0
	YK23-04	14.8	2541.1 ± 1.2	10313 ± 20	60.06 ± 0.89	0.8830 ± 0.0015	3592.3 ± 8.9	188.73 ± 0.98	188.57 ± 0.98	102.4 ± 1.5
		Hiatus								
nit	VK23-05	16.8	32555 + 20	1365 ± 14	325 ± 11	0.8632 ± 0.0012	33986 ± 363	103.47 ± 0.00	103.40 ± 0.00	56.1 ± 1.8
Stalagı	VK22-05	27.6	2094.7 ± 1.5	1305 ± 14	52.5 ± 1.1	0.0052 ± 0.0012	$197(4 \pm 112)$	105.47 ± 0.00	105.70 ± 0.00	50.1 ± 1.0
	YK25-00	27.0	3084.7 ± 1.5	2354 ± 14	32.53 ± 0.92	$0.80/1 \pm 0.0012$	$18/04 \pm 112$	195.87 ± 0.93	195.79 ± 0.93	30.0 ± 1.0
	YK23-07	35.6	2208.7 ± 1.3	2343 ± 15	47.1 ± 1.0	0.8848 ± 0.0014	13768 ± 89	197.5 ± 1.1	197.5 ± 1.1	82.2 ± 1.8
	YK23-08	42.4	1917.04 ± 0.90	4503 ± 17	39.3 ± 1.1	0.8795 ± 0.0013	6182 ± 25	199.3 ± 1.1	199.2 ± 1.1	68.9 ± 1.9
		Histus								
	VIZ22.00	12.0	2720 4 1 1 5	1100 14	21.22 ± 0.00	0.9(22 + 0.0012	24260 + 420	201.0 ± 1.1	200.0 ± 1.1	275117
	YK25-09	43.0	$2/20.4 \pm 1.5$	1128 ± 14	21.23 ± 0.90	0.8033 ± 0.0013	34309 ± 430	201.0 ± 1.1	200.9 ± 1.1	$3/.3 \pm 1.7$
	YK23-10	62.4	3355.3 ± 2.2	698 ± 23	16.2 ± 1.0	0.8657 ± 0.0014	68753 ± 2263	206.2 ± 1.2	206.1 ± 1.2	29.0 ± 1.8
	YK23-11	77.2	2262.6 ± 1.5	899 ± 19	15.0 ± 1.1	0.8655 ± 0.0015	35976 ± 777	206.9 ± 1.3	206.8 ± 1.3	26.9 ± 2.1
ite	YK47-01	118.8	812.37 ± 0.81	6437 ± 11	395.2 ± 1.8	1.0173 ± 0.0022	2120.0 ± 6.0	130.19 ± 0.61	129.99 ± 0.61	570.7 ± 2.8
lagm YK47	VK47 02	137.5	765.06 ± 0.70	2007.5 ± 7.6	308.0 ± 1.8	1.0205 ± 0.0010	4343 ± 13	132.27 ± 0.57	132.14 ± 0.57	570.7 ± 2.8
	11(47-02	157.5	105.70 ± 0.10	2771.5 ± 1.0	570.7 ± 1.0	1.02/5 ± 0.001/	4545 ± 15	152.27 ± 0.57	152.14 ± 0.57	517.1 ± 2.0
ota										
Stalagmite: YK61	YK01-01	13.6	3421.4 ± 2.1	$13/36 \pm 25$	295.8 ± 1.2	$0.91/2 \pm 0.0019$	3779 ± 10	125.39 ± 0.51	125.20 ± 0.51	421.5 ± 1.8
	YK61-02	15.5	3636.8 ± 1.9	4502 ± 12	275.4 ± 1.2	0.9027 ± 0.0013	12039 ± 37	125.80 ± 0.41	125.72 ± 0.41	393.0 ± 1.8
	YK61-03	17.0	3974.8 ± 2.4	4663 ± 10	261.5 ± 1.2	0.8936 ± 0.0013	12577 ± 32	126.29 ± 0.41	126.21 ± 0.41	373.6 ± 1.8
	YK61-04	20.0	34186 ± 37	1271.0 ± 8.9	302.6 ± 1.8	0.9278 ± 0.0013	41205 ± 291	126.64 ± 0.48	126.58 ± 0.48	432.9 ± 2.6
	VK(1.07	20.0	1520 4 1 2 4	2(27 + 22	240.2 1.0	0.0610 ± 0.0013	((50 + (2)	127.01 ± 0.70	127.50 ± 0.70	102.7 ± 2.0
	1 K01-05	22.4	1520.4 ± 2.4	3027 ± 33	540.2 ± 2.4	0.9619 ± 0.0024	6658 ± 65	$12/.00 \pm 0.72$	127.50 ± 0.72	$48/.8 \pm 3.3$
	YK61-06	26.0	2414.5 ± 4.3	2217 ± 29	315.2 ± 2.4	0.9448 ± 0.0027	16993 ± 229	128.33 ± 0.80	128.25 ± 0.80	453.0 ± 3.6
	YK61-07	28.3	4454.4 ± 4.8	801.0 ± 8.8	313.7 ± 1.7	0.9452 ± 0.0013	86784 ± 959	128.70 ± 0.47	128.63 ± 0.47	451.4 ± 2.5
	YK61-08	30.1	24344 + 23	6574 + 86	3145 ± 16	0.9479 ± 0.0012	57958 + 756	12921 ± 043	129.15 ± 0.43	4531 + 23
	1 K01-00	10.0	$2+3+.+ \pm 2.5$	007.4 ± 0.0	314.5 ± 1.0	0.9479 ± 0.0012	37,550 ± 750	129.21 ± 0.43	129.15 ± 0.45	$+35.1 \pm 2.5$
	1 K01-09	40.8	3033.5 ± 4.6	207 ± 25	302.5 ± 2.1	0.9389 ± 0.0019	$2/130/\pm 32442$	129.37 ± 0.64	129.31 ± 0.64	430.1 ± 3.2
	YK61-10	47.8	3140.5 ± 3.0	132.3 ± 7.0	305.6 ± 1.6	0.9459 ± 0.0013	370865 ± 19563	130.52 ± 0.45	130.46 ± 0.45	441.9 ± 2.3
	YK61-11	61.3	5420.5 ± 6.6	3648 ± 10	306.2 ± 1.8	0.9502 ± 0.0016	23311 ± 67	131.47 ± 0.55	131.39 ± 0.55	443.9 ± 2.7
		Histus								
	VV61 12	62 1	2207.2 + 1.0	1047 5 1 9 2	202.0 ± 1.2	0.0001 + 0.0012	10171 + 94	120 78 1 0 45	120 70 1 0 45	451.0 1.2.0
	1 K01-12	05.1	2307.3 ± 1.8	1947.3 ± 8.3	303.9 ± 1.3	0.9601 ± 0.0012	$191/1 \pm 84$	139.70 ± 0.45	139.70 ± 0.45	431.0 ± 2.0
	YK61-13	74.0	5853.2 ± 7.4	3435 ± 11	287.2 ± 1.7	0.9743 ± 0.0017	27409 ± 90	142.09 ± 0.63	142.01 ± 0.63	429.2 ± 2.7
	YK61-14	88.0	3614.8 ± 7.1	352 ± 20	321.2 ± 2.9	1.0365 ± 0.0027	175586 ± 9727	151.4 ± 1.1	151.3 ± 1.1	492.7 ± 4.7
	YK61-15	110.0	47053 ± 85	672 ± 16	3203 ± 26	1.0476 ± 0.0026	121109 ± 2076	154.9 ± 1.1	154.9 ± 1.1	496.2 ± 4.4
	VV61 16	120.0	$\pm 103.3 \pm 0.3$	672 ± 10	202.7 ± 2.0	$1.0+70 \pm 0.0020$	121177 ± 2770 120661 ± 2762	107.7 ± 1.1	107.7 ± 1.1	177612
	1 K01-10	130.0	$51/5.2 \pm 8.0$	040 ± 18	303.7 ± 2.3	1.0495 ± 0.0022	$138001 \pm 3/63$	100.25 ± 0.98	100.18 ± 0.98	$4/1.0 \pm 3.8$
	YK61-17	137.8	6174.8 ± 8.5	405.3 ± 7.9	299.4 ± 2.0	1.0514 ± 0.0019	264459 ± 5140	162.16 ± 0.87	162.10 ± 0.87	473.5 ± 3.4
	YK61-18	167.8	4766.3 ± 5.3	347.8 ± 7.3	274.1 ± 1.7	1.0478 ± 0.0014	237115 ± 4998	169.06 ± 0.77	168.99 ± 0.77	441.9 ± 3.0
	YK61-19	185.8	2984.1 ± 2.9	1897.4 + 9.4	239.0 ± 1.7	1.0238 ± 0.0015	26585 ± 135	172.56 ± 0.84	172.49 ± 0.84	3892 + 20
	11101-19	105.0	2701.1 ± 2.7	1071.7 ± 7.4	207.0 ± 1.1	1.0250 ± 0.0015	20000 ± 100	. / 2.20 ± 0.04		507.2 ± 2.7

Chemistry was performed during 2011–2012 (Shen et al., 2003) and instrumental analyses on MC-ICP-MS (Shen et al., 2012). Analytical errors are 2σ of the mean. a δ^{234} U = ($[2^{234}$ U/ 2^{238} U]_{activity} - 1) · 1000.

a δ^{-3N} U = ($[t^{-N}U]/t^{-2N}U]_{activity} = 1$). 1000. b δ^{234} U initial corrected was calculated based on ²³⁰Th age (*T*), i.e., $\delta^{234}U_{initial} = \delta^{234}U \cdot e^{\lambda_{234} \cdot T}$, and *T* is corrected age. c $[2^{30}Th/2^{38}U]_{activity} = 1 - e^{-\lambda_{230}T} + (\delta^{234}U/1000)[\lambda_{230}/(\lambda_{230} - \lambda_{234})](1 - e^{-(\lambda_{230} - \lambda_{234})T})$, where *T* is the age. Decay constants used are available in Cheng et al. (2000). d The degree of detrital ²³⁰Th contamination is indicated by the $[2^{30}Th/2^{32}Th]$ atomic ratio instead of the activity ratio. e Age [yr BP (before AD 1950)] corrections were made using an ²³⁰Th/2^{32}Th atomic ratio of 4 ± 2 ppm.

Those are the values for material at secular equilibrium, with the crustal 232 Th/ 238 U value of 3.8. The errors are arbitrarily assumed to be 50 %.

2.3 Stable isotopes

Five-to-seven coeval subsamples, $60-120 \mu g$ each, were drilled from one layer per stalagmite to measure the oxygen and carbon isotopic compositions as part of the so-called "Hendy test" (Hendy, 1971). To obtain oxygen time series, 604 subsamples, $60-120 \mu g$ each, were drilled at 0.5-3.0 mm intervals along the maximum growth axis. Measurement of oxygen stable isotopes was performed by two isotope ratio mass spectrometers, including a Finnigan Delta V Plus in the Southwest University, China, and a Micromass IsoPrime instrument at the National Taiwan Normal University. Oxygen isotope values were reported as $\delta^{18}O$ (‰) with respect to the Vienna Pee Dee Belemnite standard (V-PDB). An international standard, NBS-19, was used in both laboratories to confirm that the 1σ standard deviation of $\delta^{18}O$ was better than $\pm 0.1\%$.

3 Results and discussion

3.1 Chronology

U-Th isotopic and concentration data and dates of all stalagmite subsamples are given in Table 1. High uranium levels range from 0.8 to 13 ppm and relatively low thorium contents from 100 s to 10 000 ppt. Corrections for initial ²³⁰Th are less than 90 years, much smaller than dating uncertainties of 400–1800 years that are common for stalagmites with these ²³⁰Th ages (Table 1). Determined age intervals are 179.6– 189.8, 133.7–181.9, 172.6–206.8, 130.0–132.1, and 97.2– 172.5 kyr BP for stalagmites YK05, YK12, YK23, YK47, and YK61, respectively (Fig. 3). One to two hiatuses are observed for stalagmites YK12, YK23, and YK61 (Figs. 2, 3). The chronology of each stalagmite was developed using linear interpolation between U-Th dates, which are all in stratigraphic order (Fig. 3).

3.2 Yangkou oxygen isotope data

The well-known Hendy test has been taken as an essential requirement when assessing the ability of stalagmites to serve as paleoclimate archives (Hendy, 1971) (Fig. 4). Despite relative large δ^{13} C variations of 0.1–0.4 % (1 σ) for coeval subsamples on the five selected layers (Fig. 4a), only a small variations in δ^{18} O of $\pm 0.1 - 0.2 \% (1\sigma)$ are observed on individual horizons of coeval subsamples (Fig. 4b). There is no relationship (0.01 < r^2 < 0.36) between δ^{18} O and δ^{13} C values for coeval subsamples of four layers (Fig. 4c), which is an additional part of the Hendy test. Although an apparent high correlation ($r^2 = 0.94$) for the plot of δ^{18} O versus δ^{13} C is expressed for the depth of 134.3 mm of stalagmite YK61 (Fig. 4c), the δ^{18} O values, from -8.2% to -8.4%, change only 0.2 %. The absence of a clear increasing δ^{18} O trend outward on the same layer (Fig. 4b) also suggests an insignificant effect of kinetic fractionation. The replication of the



Figure 3. Age models of Yangkou stalagmites, established with U-Th dates with 2σ precisions of $\pm 0.3 - 1.0 \%$ (horizontal error bars).



Figure 4. Hendy test on the arbitrarily selected laminae of five stalagmites with coeval data of (**A**) δ^{13} C and (**B**) δ^{18} O. (**C**) Plots of δ^{18} O versus δ^{13} C for coeval subsamples.

 δ^{18} O records both within Yangkou cave (Fig. 5) and between Chinese caves (Fig. 6), as well as successful Hendy tests, indicates that the stalagmites formed under an oxygen isotopic equilibrium condition. The Yangkou stalagmite δ^{18} O data therefore represent rainfall oxygen isotopic change, which is a reflection of regional hydrological variability in the AM territory (e.g., Wang et al., 2001, 2008; Cheng et al., 2009; Li et al., 2011).

The oxygen isotope sequences for all of the Yangkou stalagmites are illustrated in Fig. 5a. The spliced record covers a time interval from 124 to 206 kyr BP, with three narrow hiatuses at 132.1–133.5, 190.4–193.2, and 200.3–200.9 kyr BP. This δ^{18} O record varies from -10% to -4%. The highest δ^{18} O data of -5% ~-4% occurs at 128–136 kyr BP, the PGM.



Figure 5. Cave stalagmite oxygen isotope records of (**A**) Yangkou (this study), (**B**) Sanbao (Wang et al., 2008; Cheng et al., 2009), and (**C**) Hulu (Cheng et al., 2006). U-Th ages and 2σ errors were color-coded by stalagmite. Numbers of MIS 5.5–7.3 are given by Sanbao record. Gray line is NHSI on 21 July at 30° N.

3.3 Comparison with other Chinese stalagmite records

The new spliced stalagmite δ^{18} O sequence from Yangkou cave over the time period of 124–206 kyr BP shows four strong ASM intervals at MIS 5.5, 6.3, 6.5, and 7.1 and four weak ASM intervals corresponding to MIS 6.2, MIS 6.4, MIS 7.0, and MIS 7.2 (Fig. 5a). This variation of stalagmite-inferred ASM recorded in Yangkou cave is aligned with previous ASM changes from other Chinese caves, such as Sanbao (Wang et al., 2008; Cheng et al., 2009) and Hulu (32°30' N, 119°10' E) (Cheng et al., 2006), from MIS 5.5 to 7.2 (Fig. 5).

The onsets of strong ASM intervals at MIS 5.5, 6.5, and 7.1 are at 128.3 ± 0.8 , 179.9 ± 0.9 , and 201.5 ± 1.1 kyr BP, respectively, in the Yangkou record and concurrent with their counterparts in Sanbao (Wang et al., 2008; Cheng et al., 2009) and Hulu (Cheng et al., 2006). Transients from strong to weak ASM states occur at 135-136 kyr BP during MIS 6.2–6.3, and 164–165 kyr BP during MIS 6.4–6.5. These also match changes in the Sanbao and Hulu records.

Over the past 200 kyr BP, the weakest ASM interval has been suggested to be at MIS 6.2 in the Sanbao records (Wang et al., 2008). For example, the δ^{18} O data are 1% higher than those at weak ASM intervals of MIS 6.4, 7.0, and 7.2 (Fig. 5). Concurrence between ASM records and ice-rafted debris events in the North Atlantic supports the hypothesis



Figure 6. Comparison of Chinese cave δ^{18} O records of (**A**) Yangkou and (**B**) Sanbao (Wang et al., 2008; Cheng et al., 2009) with (**C**) reconstructed SST records in the WPWP (core ODP806B) and EEP (core TR163-19) (Lea et al., 2000), and (**D**) a global stack benthic foraminifer δ^{18} O sequence LR04 (Lisiecki and Raymo, 2005). Numbers of MIS 5.5–8 are given by LR04 record. Gray line is NHSI on 21 July at 30° N. Vertical bars denote high insolation intervals.

of a NH high-latitude forcing of the ASM (Cheng et al., 2009). δ^{18} O values at MIS 6.2 in Yangkou record are 1.5–2% higher than those at MIS 6.4, 7.0, and 7.2 (Fig. 5). This large difference suggests that this event in Chongqing may have been relatively intensified through NH forcing as compared with the Hubei regions during the PGM.

The Sanbao record indicates that the strongest ASM condition over the past 500 kyr BP occurs at MIS 6.5 (Cheng et al., 2012b). This ASM event, lasting 13 kyr, is 3 kyr longer than a comparable event (in terms of intensity) at interglacial MIS 5.3, and was stronger than at any time during MIS 1, 5.5, 7.3, 9.5, and 11.3, which experienced higher sea level and NH insolation (Fig. 1 of Cheng et al., 2012b). The lowest contemporaneous δ^{18} O data in the Yangkou record (Fig. 5) show a similar ASM intensity at MIS 6.5 in southwest China.

During the MIS 5, the variations of Chinese stalagmite δ^{18} O records are not consistent among caves (Cheng et al., 2012). In Sanbao record (Wang et al., 2008), the δ^{18} O

minimum at MIS 5.3 is more depleted than at MIS 5.5. This phenomenon is seemingly illustrated in Yangkou records (Fig. 5a). However, Dongge (Kelly et al., 2006) and Tianmen (Cai et al., 2010a) stalagmite records are characterized by the most depletion in ¹⁸O at MIS 5.5 (Fig. 2 of Cai et al., 2010a). This discrepancy may be attributable to different hydrological conditions at MIS 5. Long time series from more Chinese caves are required to derive a clear picture of amplitude changes in relation to orbital forcing at MIS 5.

Overall, consistency of the stalagmite δ^{18} O sequences between Yangkou and other Chinese caves supports the idea that ASM intensity primarily follows NHSI on orbital timescales and is driven by precessional forcing and is punctuated by NH high-latitude climatic fluctuations (e.g., Wang et al., 2001, 2008; Cheng et al., 2009). Agreement in the amplitude and the transition of δ^{18} O dynamics during different MIS also confirms that the Sanbao stalagmite-inferred ASM events at MIS 6, including a very weak one at MIS 6.2 and the strongest one at MIS 6.5, are likely predominant over the entire mainland during the penultimate G–IG cycles (Cheng et al., 2012a) (Fig. 6).

3.4 Forcings for the abnormally strong ASM at MIS 6.5

The extraordinarily strong ASM condition at MIS 6.5 during the penultimate glacial period is one of the most striking features revealed by stalagmite records from three different Chinese caves (Fig. 5). This strong summer monsoon event is also observed in Chinese Loess plateau record (Rousseau et al., 2009). Modeling experiments suggest this increased monsoon intensity is primarily attributed to high NH insolation (Masson et al., 2000).

Wang et al. (2008) found a correlation between the stalagmite-inferred ASM intensity and the atmospheric δ^{18} O records from Antarctic Vostok ice-core O₂ bubbles (Sowers et al., 1991; Petit et al., 1999), and suggested that the Dole effect (Dole, 1936; Bender et al., 1994) can explain this similarity. A minimum atmospheric δ^{18} O (δ^{18} O_{atm}) peak at 170 kyr BP in the Vostok ice core (Petit et al., 1999), for example, matches the strong-ASM period at MIS 6.5.

The evolution of $\delta^{18}O_{atm}$ inferred from the Vostok ice core most likely results from changes in summer insolation and precipitation in NH, where land provides space for the growth of vegetation and photosynthesis during glacial periods (Sun et al., 2000). However, the summer insolation at MIS 6.5 is less than the interglacial periods at MIS 5.5 and 7.3 (Fig. 5), suggesting that the minimal stalagmite δ^{18} O values at MIS 6.5 could also be associated with additional secondary forcing(s).

Climate conditions around Yangkou and Sanbao caves are influenced by the Indian summer monsoon (ISM) and East Asian summer monsoon (EASM) (Fig. 1). The ISM is primarily driven by a south–north land–sea thermal gradient; instead, the EASM is controlled by both south–north and east– west land–sea gradients (Wang and Lin, 2002). The EASM precipitation is influenced by the northwestern Pacific tropical high, developed by the mainland-Pacific thermal gradient (Wang et al., 2003). The Pacific climatic variability can, therefore, affect EASM precipitation (Tan, 2014).

Cai et al. (2010b) and Jiang et al. (2012) argued for a significant impact of the western tropical Pacific sea surface temperature (SST) on the EASM precipitation. They proposed that the evolution and spatial asynchroneity of stalagmite-inferred Holocene precipitation histories at different AM regions could be attributed to SST changes in the western Pacific. Planktonic foraminiferal-inferred SST records of the marine sediment core ODP806B (0°19' N, 159°22'E) in the western Pacific warm pool (WPWP) and TR163-19 (2°16' N, 90°57' W) in the eastern equatorial Pacific (EEP) (Lea et al., 2000) are plotted in Fig. 6, along with the LR04 stacked benthic δ^{18} O sequence (Lisiecki and Raymo, 2005) and Yangkou and Sanbao cave time series. A SST gradient between the WPWP and EEP during the glacial periods of MIS 6 and 8 is 2°C, larger than the 0.5–1.5°C gradient during the warm interglacial windows of MIS 5.5 and 7 (Fig. 6). Combined with salinity gradient data, Lea et al. (2000) suggested that the transport of water vapor to the western Pacific was enhanced during glacial times. This large SST gradient could result in an enhanced Walker circulation in the Pacific, similar to the modern La Niña state, which moves the rainfall zone westward and intensifies EASM precipitation (Clement et al., 1999) (Fig. 1). Under a weak Walker circulation, analogous to present El Niño conditions, the rainfall zone in the Pacific migrated eastward and EASM precipitation was reduced (Clement et al., 1999). We speculate that the extremely strong EASM precipitation at MIS 6.5 was not only governed by high NHSI, but also partially affected by the Pacific SST gradient.

This speculation is supported by modern meteorological observations (e.g., Xue et al., 2007; Tan, 2014) and resolved decadal marine records (Oppo et al., 2009). La Niña years accompany precipitation probabilities above normal in mainland China (Tan, 2014, and references therein). However, comparison of SST histories in the South China Sea and eastern equatorial Pacific SST suggests an El Niño-like condition for the last glacial time (Koutavas et al., 2002), opposite to the findings by Lea et al. (2000). The study by Koutavas et al. (2002) does not support our argument at MIS 6.5.

Sea level change could be one of the secondary factors. Marine proxy records and model simulations show that the exposure of the Sunda shelf at the Last Glacial Maximum (LGM) associated with a low sea level condition can alters regional hydrologic pattern in Southeast Asia (DiNezio and Tierney, 2013). During the LGM, the strong Pacific equatorial SST gradient could strengthen the Pacific Walker circulation and increase rainfall in the west tropical Pacific. As pointed out by DiNezio and Tierney (2013), both of the proxies and model simulations are highly uncertain renditions of climate history, and thus multi-proxy records and high precise models are critical to understand paleoclimate.

3.5 Abrupt ASM changes

One prominent feature of ASM dynamics is the occurrence of sudden δ^{18} O shifts at about the midpoint of precessiondominated NHSI change expressed in all Chinese caves over the study time window (Kelly et al., 2006; Cai et al., 2010a; Wang et al., 2008; Cheng et al., 2012a) (Fig. 5). For example, the jumps from weak to strong ASM states lasted < 100 years from MIS 6.2 to 5.5 and 500 years from MIS 7.2 to 7.1 (this study; Wang et al., 2008; Cheng et al., 2009). Climate in Hulu Cave is primarily dominated by EASM; on the other hand, Yangkou and Sanbao caves are located in a region influenced by both EASM and ISM. This agreement of local abrupt δ^{18} O changes supports the synchroneity of both monsoon sub-system variations on precessional timescale (e.g., Cheng et al., 2012a) and confirms the robustness and regionality of these abrupt transitions in the vast ASM territory. Yangkou records also support the phase lag between ASM and NHSI (Cheng et al., 2009, 2012a). This phase lag could be attributed to the influence of millennial-scale abrupt climate change in NH high latitudes (Porter and An, 1995; Sun et al., 2012), which delayed the response of ASM to the rising NHSI (Ziegler et al., 2010; Cheng et al., 2012a).

4 Conclusions

In this study, our new spliced δ^{18} O record of five stalagmites from Yangkou cave, Chongqing, exhibits ASM variability over the time period during 124–206 kyr BP. The prominent consistency between the Yangkou and previous Chinese cave δ^{18} O sequences confirms the duration and intensity of the encompassed ASM events in the entire mainland. Our data supports the hypothesis that the ASM change primarily follows NHSI on a precessional timescale. The weakest ASM condition during low-insolation MIS 6.2 was influenced by forcing originating from the North Atlantic. The strongest ASM intensity at MIS 6.5 over the past 500 kyr BP (Cheng et al., 2012b) was presumably partially related to zonal forcing and/or sea level change associated with G-IG dynamics of Walker circulation in the Pacific. More robust geological archives and model simulations are needed to decipher detailed mechanism and forcings for G-IG ASM evolution.

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