

Lawrence Berkeley National Laboratory

Recent Work

Title

Strange hadron production in Au+Au collisions at $\sqrt{s_{NN}} = 7.7, 11.5, 19.6, 27, \text{ and } 39 \text{ GeV}$

Permalink

<https://escholarship.org/uc/item/79f5x4p5>

Journal

Physical Review C, 102(3)

ISSN

2469-9985

Authors

Adam, J
Adamczyk, L
Adams, JR
[et al.](#)

Publication Date

2020-09-01

DOI

10.1103/PhysRevC.102.034909

Peer reviewed

Strange hadron production in Au+Au collisions at $\sqrt{s_{NN}} = 7.7, 11.5, 19.6, 27, \text{ and } 39$ GeV

J. Adam,⁶ L. Adamczyk,² J. R. Adams,³⁹ J. K. Adkins,³⁰ G. Agakishiev,²⁸ M. M. Aggarwal,⁴⁰ Z. Ahammed,⁶⁰ I. Alekseev,^{3,35} D. M. Anderson,⁵⁴ R. Aoyama,⁵⁷ A. Aparin,²⁸ D. Arkhipkin,⁶ E. C. Aschenauer,⁶ M. U. Ashraf,⁵⁶ F. Atetalla,²⁹ A. Attri,⁴⁰ G. S. Averichev,²⁸ V. Bairathi,³⁶ K. Barish,¹⁰ A. J. Bassill,¹⁰ A. Behera,⁵² R. Bellwied,²⁰ A. Bhasin,²⁷ A. K. Bhati,⁴⁰ J. Bielcik,¹⁴ J. Bielcikova,³⁸ L. C. Bland,⁶ I. G. Bordyuzhin,³ J. D. Brandenburg,^{49,6} A. V. Brandin,³⁵ J. Bryslawskij,¹⁰ I. Bunzarov,²⁸ J. Butterworth,⁴⁵ H. Caines,⁶³ M. Calderón de la Barca Sánchez,⁸ D. Cebra,⁸ I. Chakaberia,^{29,6} P. Chaloupka,¹⁴ B. K. Chan,⁹ F-H. Chang,³⁷ Z. Chang,⁶ N. Chankova-Bunzarova,²⁸ A. Chatterjee,⁶⁰ S. Chattopadhyay,⁶⁰ J. H. Chen,¹⁸ X. Chen,⁴⁸ J. Cheng,⁵⁶ M. Cherney,¹³ W. Christie,⁶ H. J. Crawford,⁷ M. Csanád,¹⁶ S. Das,¹¹ T. G. Dedovich,²⁸ I. M. Deppner,¹⁹ A. A. Derevschikov,⁴² L. Didenko,⁶ C. Dilks,⁴¹ X. Dong,³¹ J. L. Drachenberg,¹ J. C. Dunlop,⁶ T. Edmonds,⁴³ N. Elsey,⁶² J. Engelage,⁷ G. Eppley,⁴⁵ R. Esha,⁵² S. Esumi,⁵⁷ O. Evdokimov,¹² J. Ewigleben,³² O. Eyser,⁶ R. Fatemi,³⁰ S. Fazio,⁶ P. Federic,³⁸ J. Fedorisin,²⁸ Y. Feng,⁴³ P. Filip,²⁸ E. Finch,⁵¹ Y. Fisyak,⁶ L. Fulek,² C. A. Gagliardi,⁵⁴ T. Galatyuk,¹⁵ F. Geurts,⁴⁵ A. Gibson,⁵⁹ K. Gopal,²³ D. Grosnick,⁵⁹ A. Gupta,²⁷ W. Guryn,⁶ A. I. Hamad,²⁹ A. Hamed,⁵ J. W. Harris,⁶³ L. He,⁴³ S. Heppelmann,⁸ S. Heppelmann,⁴¹ N. Herrmann,¹⁹ L. Holub,¹⁴ Y. Hong,³¹ S. Horvat,⁶³ B. Huang,¹² H. Z. Huang,⁹ S. L. Huang,⁵² T. Huang,³⁷ X. Huang,⁵⁶ T. J. Humanic,³⁹ P. Huo,⁵² G. Igo,⁹ W. W. Jacobs,²⁵ C. Jena,²³ A. Jentsch,⁶ Y. Ji,⁴⁸ J. Jia,^{6,52} K. Jiang,⁴⁸ S. Jowzaee,⁶² X. Ju,⁴⁸ E. G. Judd,⁷ S. Kabana,²⁹ S. Kagamaster,³² D. Kalinkin,²⁵ K. Kang,⁵⁶ D. Kapukchyan,¹⁰ K. Kauder,⁶ H. W. Ke,⁶ D. Keane,²⁹ A. Kechechyan,²⁸ M. Kelsey,³¹ Y. V. Khyzhniak,³⁵ D. P. Kikoła,⁶¹ C. Kim,¹⁰ T. A. Kinghorn,⁸ I. Kisel,¹⁷ A. Kisel,⁶¹ M. Kocan,¹⁴ L. Kochenda,³⁵ L. K. Kosarzewski,¹⁴ L. Kramarik,¹⁴ P. Kravtsov,³⁵ K. Krueger,⁴ N. Kulathunga Mudiyansele,²⁰ L. Kumar,⁴⁰ R. Kunnawalkam Elayavalli,⁶² J. H. Kwasizur,²⁵ R. Lacey,⁵² J. M. Landgraf,⁶ J. Lauret,⁶ A. Lebedev,⁶ R. Lednicky,²⁸ J. H. Lee,⁶ C. Li,⁴⁸ W. Li,⁵⁰ W. Li,⁴⁵ X. Li,⁴⁸ Y. Li,⁵⁶ Y. Liang,²⁹ R. Licenik,³⁸ T. Lin,⁵⁴ A. Lipiec,⁶¹ M. A. Lisa,³⁹ F. Liu,¹¹ H. Liu,²⁵ P. Liu,⁵² P. Liu,⁵⁰ T. Liu,⁶³ X. Liu,³⁹ Y. Liu,⁵⁴ Z. Liu,⁴⁸ T. Ljubicic,⁶ W. J. Llope,⁶² M. Lomnitz,³¹ R. S. Longacre,⁶ S. Luo,¹² X. Luo,¹¹ G. L. Ma,⁵⁰ L. Ma,¹⁸ R. Ma,⁶ Y. G. Ma,⁵⁰ N. Magdy,¹² R. Majka,⁶³ D. Mallick,³⁶ S. Margetis,²⁹ C. Markert,⁵⁵ H. S. Matis,³¹ O. Matonoha,¹⁴ J. A. Mazer,⁴⁶ K. Meehan,⁸ J. C. Mei,⁴⁹ N. G. Minaev,⁴² S. Mioduszewski,⁵⁴ D. Mishra,³⁶ B. Mohanty,³⁶ M. M. Mondal,²⁶ I. Mooney,⁶² Z. Moravcova,¹⁴ D. A. Morozov,⁴² Md. Nasim,²² K. Nayak,¹¹ J. M. Nelson,⁷ D. B. Nemes,⁶³ M. Nie,⁴⁹ G. Nigmatkulov,³⁵ T. Niida,⁶² L. V. Nogach,⁴² T. Nonaka,¹¹ G. Odyniec,³¹ A. Ogawa,⁶ K. Oh,⁴⁴ S. Oh,⁶³ V. A. Okorokov,³⁵ B. S. Page,⁶ R. Pak,⁶ Y. Panebratsev,²⁸ B. Pawlik,² D. Pawlowska,⁶¹ H. Pei,¹¹ C. Perkins,⁷ R. L. Pintér,¹⁶ J. Pluta,⁶¹ J. Porter,³¹ M. Posik,⁵³ N. K. Pruthi,⁴⁰ M. Przybycien,² J. Putschke,⁶² A. Quintero,⁵³ S. K. Radhakrishnan,³¹ S. Ramachandran,³⁰ R. L. Ray,⁵⁵ R. Reed,³² H. G. Ritter,³¹ J. B. Roberts,⁴⁵ O. V. Rogachevskiy,²⁸ J. L. Romero,⁸ L. Ruan,⁶ J. Rusnak,³⁸ O. Rusnakova,¹⁴ N. R. Sahoo,⁴⁹ P. K. Sahu,²⁶ S. Salur,⁴⁶ J. Sandweiss,⁶³ J. Schambach,⁵⁵ W. B. Schmidke,⁶ N. Schmitz,³³ B. R. Schweid,⁵² F. Seck,¹⁵ J. Seger,¹³ M. Sergeeva,⁹ R. Seto,¹⁰ P. Seyboth,³³ N. Shah,²⁴ E. Shahaliev,²⁸ P. V. Shanmuganathan,³² M. Shao,⁴⁸ F. Shen,⁴⁹ W. Q. Shen,⁵⁰ S. S. Shi,¹¹ Q. Y. Shou,⁵⁰ E. P. Sichtermann,³¹ S. Siejka,⁶¹ R. Sikora,² M. Simko,³⁸ J. Singh,⁴⁰ S. Singha,²⁹ D. Smirnov,⁶ N. Smirnov,⁶³ W. Solyst,²⁵ P. Sorensen,⁶ H. M. Spinka,⁴ B. Srivastava,⁴³ T. D. S. Stanislaus,⁵⁹ M. Stefaniak,⁶¹ D. J. Stewart,⁶³ M. Strikhanov,³⁵ B. Stringfellow,⁴³ A. A. P. Suaide,⁴⁷ T. Sugiura,⁵⁷ M. Sumera,³⁸ B. Summa,⁴¹ X. M. Sun,¹¹ Y. Sun,⁴⁸ Y. Sun,²¹ B. Surrow,⁵³ D. N. Svirida,³ P. Szymanski,⁶¹ A. H. Tang,⁶ Z. Tang,⁴⁸ A. Taranenko,³⁵ T. Tarnowsky,³⁴ J. H. Thomas,³¹ A. R. Timmins,²⁰ D. Tlusty,¹³ T. Todoroki,⁶ M. Tokarev,²⁸ C. A. Tomkiel,³² S. Trentalange,⁹ R. E. Tribble,⁵⁴ P. Tribedy,⁶ S. K. Tripathy,²⁶ O. D. Tsai,⁹ B. Tu,¹¹ Z. Tu,⁶ T. Ullrich,⁶ D. G. Underwood,⁴ I. Upsal,^{49,6} G. Van Buren,⁶ J. Vanek,³⁸ A. N. Vasiliev,⁴² I. Vassiliev,¹⁷ F. Videbæk,⁶ S. Vokal,²⁸ S. A. Voloshin,⁶² F. Wang,⁴³ G. Wang,⁹ P. Wang,⁴⁸ Y. Wang,¹¹ Y. Wang,⁵⁶ J. C. Webb,⁶ L. Wen,⁹ G. D. Westfall,³⁴ H. Wieman,³¹ S. W. Wissink,²⁵ R. Witt,⁵⁸ Y. Wu,²⁹ Z. G. Xiao,⁵⁶ G. Xie,¹² W. Xie,⁴³ H. Xu,²¹ N. Xu,³¹ Q. H. Xu,⁴⁹ Y. F. Xu,⁵⁰ Z. Xu,⁶ C. Yang,⁴⁹ Q. Yang,⁴⁹ S. Yang,⁶ Y. Yang,³⁷ Z. Yang,¹¹ Z. Ye,⁴⁵ Z. Ye,¹² L. Yi,⁴⁹ K. Yip,⁶ I. -K. Yoo,⁴⁴ H. Zbroszczyk,⁶¹ W. Zha,⁴⁸ D. Zhang,¹¹ L. Zhang,¹¹ S. Zhang,⁴⁸ S. Zhang,⁵⁰ X. P. Zhang,⁵⁶ Y. Zhang,⁴⁸ Z. Zhang,⁵⁰ J. Zhao,⁴³ C. Zhong,⁵⁰ C. Zhou,⁵⁰ X. Zhu,⁵⁶ Z. Zhu,⁴⁹ M. Zurek,³¹ and M. Zyzak¹⁷

(STAR Collaboration)

¹Abilene Christian University, Abilene, Texas 79699

²AGH University of Science and Technology, FPACS, Cracow 30-059, Poland

³Alikhanov Institute for Theoretical and Experimental Physics, Moscow 117218, Russia

⁴Argonne National Laboratory, Argonne, Illinois 60439

⁵American University of Cairo, Cairo, Egypt

- ⁶ Brookhaven National Laboratory, Upton, New York 11973
⁷ University of California, Berkeley, California 94720
⁸ University of California, Davis, California 95616
⁹ University of California, Los Angeles, California 90095
¹⁰ University of California, Riverside, California 92521
¹¹ Central China Normal University, Wuhan, Hubei 430079
¹² University of Illinois at Chicago, Chicago, Illinois 60607
¹³ Creighton University, Omaha, Nebraska 68178
¹⁴ Czech Technical University in Prague, FNSPE, Prague 115 19, Czech Republic
¹⁵ Technische Universität Darmstadt, Darmstadt 64289, Germany
¹⁶ Eötvös Loránd University, Budapest, Hungary H-1117
¹⁷ Frankfurt Institute for Advanced Studies FIAS, Frankfurt 60438, Germany
¹⁸ Fudan University, Shanghai, 200433
¹⁹ University of Heidelberg, Heidelberg 69120, Germany
²⁰ University of Houston, Houston, Texas 77204
²¹ Huzhou University, Huzhou, Zhejiang 313000
²² Indian Institute of Science Education and Research (IISER), Berhampur 760010, India
²³ Indian Institute of Science Education and Research, Tirupati 517507, India
²⁴ Indian Institute of Technology, Patna, Bihar, India
²⁵ Indiana University, Bloomington, Indiana 47408
²⁶ Institute of Physics, Bhubaneswar 751005, India
²⁷ University of Jammu, Jammu 180001, India
²⁸ Joint Institute for Nuclear Research, Dubna 141 980, Russia
²⁹ Kent State University, Kent, Ohio 44242
³⁰ University of Kentucky, Lexington, Kentucky 40506-0055
³¹ Lawrence Berkeley National Laboratory, Berkeley, California 94720
³² Lehigh University, Bethlehem, Pennsylvania 18015
³³ Max-Planck-Institut für Physik, Munich 80805, Germany
³⁴ Michigan State University, East Lansing, Michigan 48824
³⁵ National Research Nuclear University MEPhI, Moscow 115409, Russia
³⁶ National Institute of Science Education and Research, HBNI, Jatni 752050, India
³⁷ National Cheng Kung University, Tainan 70101
³⁸ Nuclear Physics Institute of the CAS, Rez 250 68, Czech Republic
³⁹ Ohio State University, Columbus, Ohio 43210
⁴⁰ Panjab University, Chandigarh 160014, India
⁴¹ Pennsylvania State University, University Park, Pennsylvania 16802
⁴² NRC "Kurchatov Institute", Institute of High Energy Physics, Protvino 142281, Russia
⁴³ Purdue University, West Lafayette, Indiana 47907
⁴⁴ Pusan National University, Pusan 46241, Korea
⁴⁵ Rice University, Houston, Texas 77251
⁴⁶ Rutgers University, Piscataway, New Jersey 08854
⁴⁷ Universidade de São Paulo, São Paulo, Brazil 05314-970
⁴⁸ University of Science and Technology of China, Hefei, Anhui 230026
⁴⁹ Shandong University, Qingdao, Shandong 266237
⁵⁰ Shanghai Institute of Applied Physics, Chinese Academy of Sciences, Shanghai 201800
⁵¹ Southern Connecticut State University, New Haven, Connecticut 06515
⁵² State University of New York, Stony Brook, New York 11794
⁵³ Temple University, Philadelphia, Pennsylvania 19122
⁵⁴ Texas A&M University, College Station, Texas 77843
⁵⁵ University of Texas, Austin, Texas 78712
⁵⁶ Tsinghua University, Beijing 100084
⁵⁷ University of Tsukuba, Tsukuba, Ibaraki 305-8571, Japan
⁵⁸ United States Naval Academy, Annapolis, Maryland 21402
⁵⁹ Valparaiso University, Valparaiso, Indiana 46383
⁶⁰ Variable Energy Cyclotron Centre, Kolkata 700064, India
⁶¹ Warsaw University of Technology, Warsaw 00-661, Poland
⁶² Wayne State University, Detroit, Michigan 48201
⁶³ Yale University, New Haven, Connecticut 06520

We present STAR measurements of strange hadron (K_S^0 , Λ , $\bar{\Lambda}$, Ξ^- , $\bar{\Xi}^+$, Ω^- , $\bar{\Omega}^+$, and ϕ) production at mid-rapidity ($|y| < 0.5$) in Au+Au collisions at $\sqrt{s_{NN}} = 7.7 - 39$ GeV from the Beam Energy Scan Program at the Relativistic Heavy Ion Collider (RHIC). Transverse momentum spectra, averaged transverse mass, and the overall integrated yields of these strange hadrons are presented versus the centrality and collision energy. Antibaryon-to-baryon ratios ($\bar{\Lambda}/\Lambda$, $\bar{\Xi}^+/\Xi^-$, $\bar{\Omega}^+/\Omega^-$) are presented

as well, and used to test a thermal statistical model and to extract the temperature normalized strangeness and baryon chemical potentials at hadronic freeze-out (μ_B/T_{ch} and μ_S/T_{ch}) in central collisions. Strange baryon-to-pion ratios are compared to various model predictions in central collisions for all energies. The nuclear modification factors (R_{CP}) and antibaryon-to-meson ratios as a function of transverse momentum are presented for all collision energies. The K_S^0 R_{CP} shows no suppression for p_T up to 3.5 GeV/ c at energies of 7.7 and 11.5 GeV. The $\bar{\Lambda}/K_S^0$ ratio also shows baryon-to-meson enhancement at intermediate p_T (≈ 2.5 GeV/ c) in central collisions at energies above 19.6 GeV. Both observations suggest that there is likely a change of the underlying strange quark dynamics at collision energies below 19.6 GeV.

PACS numbers: 25.75.-q, 25.75.Dw, 25.75.Nq

I. INTRODUCTION

The main motivation of the RHIC Beam Energy Scan (BES) Program is to study the quantum chromodynamics (QCD) phase diagram [1–3]. Systematic analysis of Au+Au collisions from $\sqrt{s_{\text{NN}}} = 39$ GeV down to 7.7 GeV in the RHIC BES Program could help to achieve the following goals: 1) to find the QCD critical point where the first order phase transition at finite baryon chemical potential (μ_B) ends and to identify the phase boundary of the first order phase transition [4–8]; 2) to locate the collision energy where deconfinement begins [9–12].

Strange hadrons are an excellent probe for identifying the phase boundary and onset of deconfinement. Strangeness enhancement in $A+A$ with respect to $p+p$ collisions has long been suggested as a signature of quark-gluon plasma (QGP) [13]. Strangeness has been extensively measured in many experiments at different accelerator facilities [14–41]. Generally, the yields of strange hadrons in nuclear collisions are close to those expected from statistical models [42–45]. The precise measurement of these yields in heavy ion collisions in the BES may lead to a better understanding of strangeness production mechanisms in nuclear collisions and a more constrained extraction of the chemical freeze-out parameters.

The measurement of strange hadrons at high p_T can probe hard parton scatterings in the QGP medium, through the central-to-peripheral nuclear modification factor $R_{\text{CP}} = (\text{yield}/N_{\text{coll}})_{\text{central}}/(\text{yield}/N_{\text{coll}})_{\text{peripheral}}$, where N_{coll} is the average number of binary nucleon-nucleon collisions. It has been observed in Au+Au collisions at $\sqrt{s_{\text{NN}}} = 200$ GeV at RHIC that, at high p_T , R_{CP} of various particles is much less than unity [32, 46], indicating a significant energy loss of the scattered partons in the dense matter. R_{CP} of strange hadrons in the BES, together with other non-strange hadron suppression results, can potentially pin down the beam energy at which energy loss to the medium begins to dominate hard parton interactions [3].

At intermediate p_T (2–5 GeV/ c), as first discovered in central Au+Au events at RHIC [37, 47–49] and later observed at the Large Hadron Collider [39, 50], the p/π and Λ/K_S^0 ratios are larger than unity and much higher than those observed in peripheral $A+A$ and in $p+p$ collisions. These results may indicate different hadronization mechanisms in this p_T range in $A+A$ collisions. There are

recombination/coalescence models which allow soft partons to coalesce into hadrons, or soft and hard partons to recombine into hadrons [51–58]. They naturally reproduce enhanced baryon-to-meson ratios for any quickly falling distribution of parton p_T . Such models rely on recombination or coalescence of constituent quarks, thus existence of a partonic medium. Hence observation of such behavior at hadronization is a fundamental piece of evidence for the formation of the partonic QGP medium. It is also interesting to investigate at which collision energies these phenomena are prevalent [3], in order to locate the energy range over which the onset of the deconfinement happens.

We present strangeness data obtained from Au+Au collisions at $\sqrt{s_{\text{NN}}} = 7.7, 11.5, 19.6, 27,$ and 39 GeV, collected by the STAR experiment during the first phase of the RHIC BES Program in 2010 and 2011.

This paper is organized as follows. Section II briefly describes the experimental setup, the event selection, and the centrality determination. Section III discusses the reconstruction methods of various strange hadrons, the signal extraction methods, the acceptance and efficiency correction factors, the feed-down corrections for Λ hyperons, the extrapolations to low p_T , and the systematic uncertainties. Section IV presents transverse momentum spectra, averaged transverse mass, integrated yields, and various particle ratios of those strange hadrons and comparisons to theory for different centralities and collision energies. Finally, Sec. V is the summary.

II. EXPERIMENTAL SETUP

The Solenoidal Tracker At RHIC (STAR) is a versatile particle detector at the RHIC collider at Brookhaven National Laboratory. A detailed description of its solenoidal magnet and various sub-detectors for tracking, particle identification, and triggering can be found in Ref. [59].

The Time Projection Chamber (TPC) is the main detector at STAR which provides tracking capability covering 2π azimuthal angle in the transverse direction and -1 to 1 in pseudo-rapidity, η [60]. The TPC is immersed in a constant 0.5 Tesla magnetic field parallel to the beam direction, which is generated by the surrounding solenoidal magnet. The track of a charged particle can be reconstructed with a maximum of 45 hit points within the

TPC fiducial radius of $0.5 < r < 2$ m. The location of the primary vertex of a collision event is determined using the reconstructed charged particle tracks. A primary vertex resolution in the transverse plane of $350 \mu\text{m}$ can be achieved with ≈ 1000 tracks. The fitted primary vertex can be included in the track fitting of the charged particles to improve their momentum resolution. With this procedure, a relative momentum resolution for pions of $\approx 2\%$ at $p_T = 1 \text{ GeV}/c$ can be achieved. The TPC also measures the energy loss of charged particles, which allows separation of π and K up to $p_T \simeq 0.7 \text{ GeV}/c$, and statistical proton identification up to $p_T \simeq 1.1 \text{ GeV}/c$ [60].

In 2010, the STAR experiment recorded Au+Au collisions at the nucleon-nucleon center-of-mass energy ($\sqrt{s_{NN}}$) of 7.7, 11.5, and 39 GeV. The data of Au+Au collisions at $\sqrt{s_{NN}} = 19.6$ and 27 GeV were further collected in 2011. The minimum bias trigger at all five energies was defined by the coincidence of the zero-degree calorimeters (ZDC), vertex position detectors (VPD) [61], and/or beam-beam counters (BBC) signals [62]. However, at the lowest beam energies, most of the triggered events are from Au beam nuclei with large emittance that hit the nuclei at rest in the beam pipe. This background can be removed by requiring the primary vertex of an event to be within a radius r of less than 2 cm of the geometrical center of the detector system, which is much less than that of the beam pipe (3.95 cm). The primary vertex position in the beam direction (z -direction) was limited to the values listed in Table I. These values were selected according to the offline z -vertex trigger conditions which were different for different energies. It was further required that at least two tracks from the primary vertex were matched to the cells of the barrel time-of-flight detector (BTOF) [63] in order to remove the pile-up events from different bunch-crossings. Finally, events from bad runs were removed according to an extensive quality assurance analysis of the events (see Ref. [10]). The accepted number of minimum bias events for each of the five energies are also listed in Table I.

TABLE I: The z -vertex acceptance, and the total number of minimum-bias (MB) events used, for different energies.

$\sqrt{s_{NN}}$ (GeV)	z -vertex range (cm)	MB events (10^6)
7.7	[-70, 70]	4.4
11.5	[-50, 50]	12.0
19.6	[-70, 70]	36.3
27	[-70, 70]	72.8
39	[-40, 40]	134.3

The centrality selection of the events was chosen to be 0–80% of the total reaction cross section due to trigger inefficiencies for the most peripheral events. The centrality definition was based on an uncorrected multiplicity distribution and a Glauber Monte Carlo simulation [64], and details can be found in Ref. [10]. By comparison of the Glauber simulation to the measured multiplicity

distribution at each energy, it is possible to determine, for each centrality class, the average number of participant nucleons $\langle N_{\text{part}} \rangle$. The values of $\langle N_{\text{part}} \rangle$ at different centralities and collision energies are listed in Table II.

III. ANALYSIS DETAILS

A. Strange particle reconstruction

The strange hadrons, K_S^0 , $\Lambda(\bar{\Lambda})$, $\Xi^-(\bar{\Xi}^+)$, $\Omega^-(\bar{\Omega}^+)$, and ϕ , have short lifetimes, and decay into a pair of charged particles or into one charged particle plus a $\Lambda(\bar{\Lambda})$. All of them can be reconstructed using the invariant mass technique. The corresponding decay channels and branching ratios are [65]

$$\begin{aligned}
 K_S^0 &\rightarrow \pi^+ + \pi^-, & 69.20\%; \\
 \Lambda(\bar{\Lambda}) &\rightarrow p(\bar{p}) + \pi^-(\pi^+), & 63.9\%; \\
 \Xi^-(\bar{\Xi}^+) &\rightarrow \Lambda(\bar{\Lambda}) + \pi^-(\pi^+), & 99.887\%; \\
 \Omega^-(\bar{\Omega}^+) &\rightarrow \Lambda(\bar{\Lambda}) + K^-(K^+), & 67.8\%; \\
 \phi &\rightarrow K^+ + K^-, & 49.1\%.
 \end{aligned}$$

The truncated mean of the ionization energy loss, $\langle dE/dx \rangle$, measured by the TPC, was used for identification of the charged daughter particles, π^\pm , K^\pm , and $p(\bar{p})$ [66]. Despite the finite statistical precision of the measured $\langle dE/dx \rangle$ for a certain track arising from a limited number of hit points measured by the TPC, the central value of the measured $\langle dE/dx \rangle$ as a function of momentum is well described by the Bichsel function for each particle species [67]. Hence a normalized $\langle dE/dx \rangle$, $n\sigma_{\text{particle}}$, was used in particle identification. It is defined as

$$n\sigma_{\text{particle}} = \frac{1}{\sigma_{\text{particle}}} \log \frac{\langle dE/dx \rangle_{\text{measured}}}{\langle dE/dx \rangle_{\text{particle}}^{\text{Bichsel}}}, \quad (1)$$

where $\langle dE/dx \rangle_{\text{particle}}^{\text{Bichsel}}$ is the expected $\langle dE/dx \rangle$ from the Bichsel function for a certain particle species at a given momentum, and σ_{particle} is the $\langle dE/dx \rangle$ resolution of the TPC for the same particle species at the same momentum. The $n\sigma_{\text{particle}}$ distribution at a given momentum is nearly Gaussian and is calibrated to be centered at zero with a width of unity for each particle species. By default, a loose cut of $|n\sigma_{\text{particle}}| < 4.0$ was used to select all the corresponding charged daughter particles for the reconstruction of K_S^0 , $\Lambda(\bar{\Lambda})$, and $\Xi^-(\bar{\Xi}^+)$. In order to reduce the combinatorial background, a tighter $|n\sigma_p| < 3.0$ was used for selecting the protons in Ω reconstruction, and $|n\sigma_K| < 2.0$ was used for the kaons in ϕ meson reconstruction. In order to improve the average momentum and energy-loss resolution, the charged daughter particle tracks were required to consist of at least 16 TPC hit points for the reconstruction of K_S^0 , $\Lambda(\bar{\Lambda})$, $\Xi^-(\bar{\Xi}^+)$, and $\Omega^-(\bar{\Omega}^+)$, while at least 16 hit points (including the primary vertex) were required for the kaons in ϕ meson reconstruction. The p_T of daughter particles was

TABLE II: The average number of participating nucleons $\langle N_{\text{part}} \rangle$ for various collision centralities in Au+Au collisions at 7.7–39 GeV, determined using the charged particle multiplicity distributions and the Glauber Monte Carlo simulation [10]. The errors represent systematic uncertainties. The inelastic $p + p$ cross-sections used in the simulations are 30.8, 31.2, 32, 33, and 34 mb for $\sqrt{s} = 7.7, 11.5, 19.6, 27,$ and 39 GeV, respectively [64].

$\sqrt{s_{\text{NN}}} \text{ (GeV)}$	0–5%	5–10%	10–20%	20–30%	30–40%	40–60%	60–80%
7.7	337.4 ± 2.1	290.4 ± 6.0	226.2 ± 7.9	160.2 ± 10.2	109.9 ± 11.0	58.4 ± 9.8	20.2 ± 5.3
11.5	338.2 ± 2.0	290.6 ± 6.2	226.0 ± 8.2	159.6 ± 9.5	110.0 ± 10.3	58.5 ± 9.4	20.1 ± 6.7
19.6	338.0 ± 2.3	289.2 ± 6.0	224.9 ± 8.6	158.1 ± 10.5	108.0 ± 10.6	57.7 ± 9.1	19.9 ± 5.9
27	343.3 ± 2.0	299.3 ± 6.2	233.6 ± 9.0	165.5 ± 10.7	114.0 ± 11.3	61.2 ± 10.4	20.5 ± 7.1
39	341.7 ± 2.2	293.9 ± 6.4	229.8 ± 8.7	162.4 ± 10.2	111.4 ± 10.8	59.2 ± 9.7	20.0 ± 6.4

required to be larger than 0.10 GeV/ c for K_S^0 , Λ , and Ξ reconstruction, and larger than 0.15 GeV/ c for Ω and ϕ reconstruction.

Due to the large number of final state particles in Au+Au collisions, there is a significant amount of combinatorial background in the invariant mass distributions of all strange hadrons. The weakly decaying strange hadrons, K_S^0 , $\Lambda(\bar{\Lambda})$, $\Xi^-(\bar{\Xi}^+)$, and $\Omega^-(\bar{\Omega}^+)$, have a typical decay length of $c\tau \approx 2\text{--}7$ cm. Their decay topology can be reconstructed well with their daughter particle tracks measured by the TPC with a precision of ≈ 1 mm. Therefore, a certain set of cuts can be applied to the topological variables in order to significantly reduce the combinatorial background. Such variables include the distance of closest approach (DCA) between the two daughter tracks, the DCA of the daughter tracks to the primary vertex, the DCA of the projected strange hadron path to the primary vertex, the decay length of strange hadrons, and the angles between the spatial vector pointing from the production vertex to the decay vertex and the momentum vector of strange hadrons. These cuts were optimized as a compromise between background reduction and signal efficiency. Table III shows the default topological cuts used for V^0 particle (K_S^0 , Λ , and $\bar{\Lambda}$) reconstruction in this analysis. For the reconstruction of multi-strange hyperons, $\Xi^-(\bar{\Xi}^+)$ and $\Omega^-(\bar{\Omega}^+)$, the Λ candidates reconstructed with p and π daughter tracks are further combined with the “bachelor” tracks — the identified π^\pm for Ξ reconstruction or the identified K^\pm for Ω reconstruction. In order to reduce the combinatorial background, the Λ candidates were required to be inside the invariant mass window of $[M_\Lambda - 0.012 \text{ GeV}/c^2, M_\Lambda + 0.012 \text{ GeV}/c^2]$ and $[M_\Lambda - 0.006 \text{ GeV}/c^2, M_\Lambda + 0.006 \text{ GeV}/c^2]$ for Ξ and Ω reconstruction, respectively, with the known Λ mass $M_\Lambda = 1.115683 \text{ GeV}/c^2$ [65]. The decay topology of multi-strange hyperons is more complicated compared to those of V^0 particles, and hence more topological cuts were used in these hyperon reconstructions. Tables IV and V show the default topological cuts for Ξ and Ω reconstruction, respectively.

The ϕ meson decays strongly at the primary collision vertex and has a short lifetime. Hence its two daughter kaons also appear to originate from the primary vertex. Therefore the primary tracks, which have the primary vertex included in their fit, were used for ϕ meson re-

construction. The DCA of their associated TPC tracks, which exclude primary vertex in their fit, to the primary vertex were required to be less than 3 cm. In order to avoid split tracks, the ratio of the number of hits on a track to the maximum possible number of hits that track may possess was required to be larger than 0.52. Due to the electron/positron contamination in the selected kaon candidates, photon conversions ($\gamma^* \rightarrow e^+e^-$) contribute significantly to the residual background in K^+K^- invariant mass distributions. This contribution can be removed effectively by a cut on the dip angle δ [34, 68–70], which is defined as

$$\delta = \cos^{-1} \left[\frac{p_{T1}p_{T2} + p_{z1}p_{z2}}{p_1p_2} \right], \quad (2)$$

where $p_1, p_2, p_{T1}, p_{T2}, p_{z1}, p_{z2}$ are total, transverse, and longitudinal momenta of the two candidate tracks. By default, the δ was required to be greater than 0.04 radians in this analysis.

After applying the corresponding selection cuts, the resulting invariant mass distributions are shown in Fig. 1 (a) for K_S^0 , (b) for Λ , (c) for Ξ , (d) for Ω , and in Fig. 2 for ϕ . Even with these cuts, some background remains under the mass peak of each hadron. The random combinatorial background was estimated using a rotation method [71] for $K_S^0, \Lambda, \Xi,$ and Ω . In the rotation method, one daughter particle track was picked, π^- , for example, in the case of K_S^0 reconstruction. Then both the two-dimensional (2D) position vector of the track’s first hit (originating from the primary vertex) and its 2D momentum vector in the transverse plane were rotated by an angle of π in this plane. The rotated track was used in the strange hadron reconstruction to break possible correlations between the daughter particles and mimic the random combination contribution. In the Ω analysis, the bachelor K^\pm tracks were rotated by five different angles from $\pi/3$ to $5\pi/3$ to increase the background statistics. For ϕ meson analysis, in order to minimize the statistical errors, the combinatorial background was estimated with the mixed-event technique. The detailed description can be found in Refs. [34, 68–70]. The invariant mass distribution, the mixed-event background, and the background subtracted distribution are shown in Fig. 2.

Besides the combinatorial background, there is also a residual background for each particle. This residual back-

TABLE III: Topological cuts used for V^0 particle (K_S^0 , Λ , and $\bar{\Lambda}$) reconstruction. In this table, \vec{r}_{V^0} and \vec{r}_{PV} denote the V^0 decay vertex position vector and the primary vertex position vector in the STAR coordinate system, respectively. \vec{p}_{V^0} is the reconstructed V^0 momentum vector. Slightly tighter topological cuts, together with a tighter particle identification cut of $|n\sigma_{p(\pi)}| < 3.8$, were used for $\bar{\Lambda}$ reconstruction at $\sqrt{s_{NN}} = 7.7$ GeV to reduce the combinatorial background.

Cut	K_S^0	$\bar{\Lambda}$ (≥ 11.5 GeV), Λ	$\bar{\Lambda}$ (7.7 GeV)
DCA of V^0 to primary vertex	< 0.8 cm	< 0.8 cm	< 0.8 cm
DCA of daughters to primary vertex	> 0.7 cm	> 0.3 cm (p), > 1.0 cm (π)	> 0.5 cm (p), > 1.5 cm (π)
DCA between daughters	< 0.8 cm	< 0.8 cm	< 0.8 cm
V^0 decay length	> 2.5 cm	> 3 cm	> 4 cm
$(\vec{r}_{V^0} - \vec{r}_{PV}) \cdot \vec{p}_{V^0}$	> 0	> 0	> 0

TABLE IV: Topological cuts used for Ξ^- and Ξ^+ reconstruction. In this table, \vec{r}_{Ξ} , \vec{r}_{Λ} and \vec{r}_{PV} denote the Ξ and Λ decay vertex position vectors and the primary vertex position vector in the STAR coordinate system, respectively. \vec{p}_{Λ} and \vec{p}_{Ξ} are the reconstructed Λ and Ξ momentum vectors. Slightly tighter topological cuts, together with a tighter particle identification cut of $|n\sigma_{p(\pi)}| < 3.6$ and a narrower Λ invariant mass window of $[M_{\Lambda} - 0.010 \text{ GeV}/c^2, M_{\Lambda} + 0.010 \text{ GeV}/c^2]$, were used for Ξ reconstruction at $\sqrt{s_{NN}} = 7.7$ and 11.5 GeV to reduce the combinatorial background.

Cut	Ξ^- and Ξ^+ (≥ 19.6 GeV)	Ξ^- and Ξ^+ (≤ 11.5 GeV)
DCA of Ξ to primary vertex	< 0.8 cm	< 0.8 cm
DCA of bachelor π to primary vertex	> 0.8 cm	> 0.8 cm
DCA of Λ to primary vertex	$[0.2, 5.0]$ cm	$[0.2, 5.0]$ cm
DCA of Λ -daughter p to primary vertex	> 0.5 cm	> 0.5 cm
DCA of Λ -daughter π to primary vertex	> 1.0 cm	> 1.5 cm
DCA between Λ and bachelor π	< 0.8 cm	< 0.8 cm
DCA between Λ -daughters	< 0.8 cm	< 0.8 cm
Ξ decay length	> 3.4 cm	> 4.0 cm
Λ decay length	> 5.0 cm	> 5.0 cm
$(\vec{r}_{\Lambda} - \vec{r}_{PV}) \cdot \vec{p}_{\Lambda}$	> 0	> 0
$(\vec{r}_{\Lambda} - \vec{r}_{\Xi}) \cdot \vec{p}_{\Lambda}$	> 0	> 0
$(\vec{r}_{\Xi} - \vec{r}_{PV}) \cdot \vec{p}_{\Xi}$	> 0	> 0
$(\vec{r}_{\Xi} - \vec{r}_{PV}) \times \vec{p}_{\Xi} / \vec{r}_{\Xi} - \vec{r}_{PV} / \vec{p}_{\Xi} $	< 0.2	< 0.12

ground originates from unavoidable particle misidentification. For example, a proton from a Λ decay misidentified as a π^+ may be combined with the π^- daughter of Λ , thereby contributing to the residual background in K_S^0 reconstruction. In Ξ reconstruction, a proton from a real Λ decay can be combined with another random pion to form a fake Λ candidate, which is then combined with the pion daughter of the real Λ (as the bachelor pion) to form a fake Ξ . The bachelor pion of a Ξ can be misidentified as a kaon, thereby contributing to the residual background in Ω reconstruction. In order to remove these kinds of residual background, veto cuts were introduced. In K_S^0 reconstruction, the $\pi^+(\pi^-)$ daughter of a K_S^0 candidate is assumed to be the $p(\bar{p})$ daughter of a $\Lambda(\bar{\Lambda})$ to re-calculate the invariant mass. If it falls inside the invariant mass peak of $\Lambda(\bar{\Lambda})$, then the K_S^0 candidate is rejected. For Ξ reconstruction, the proton daughter will be combined with the pion bachelor to calculate the invariant mass. If it falls inside the Λ invariant mass peak, the Ξ candidate will be rejected. In Ω reconstruction, the bachelor kaon is assumed to be a pion to re-calculate the invariant mass. If it falls inside the invariant mass peak of Ξ , then the Ω candidate will be rejected. There also exists minor residual background in the Λ invariant mass distri-

bution due to the misidentification of the K_S^0 daughters. However, the veto of this residual background would produce a significant drop in Λ reconstruction efficiency due to the large width of K_S^0 invariant mass peak, hence no veto cuts were applied for Λ . The veto cuts for K_S^0 , Ξ , and Ω were applied both in signal reconstruction and in construction of the rotational background. The K_S^0 , Ξ , and Ω invariant mass distributions shown in Fig. 1 were obtained after applying the corresponding veto cuts.

The background distributions estimated via rotational or mixed-event methods for K_S^0 , Λ , Ξ , Ω , and ϕ are subtracted from the corresponding signal distributions. For K_S^0 , Λ , and Ξ , a double or single Gaussian plus polynomial fitting to the resulting invariant mass distribution around the signal peaks was used to determine the signal peak width as well as the shape of the remaining residual background. The signal peak was defined as $[\mu - 4\sigma, \mu + 4\sigma]$ and $[\mu - 4.5\sigma, \mu + 4.5\sigma]$ for V^0 and Ξ , respectively, where μ and σ are the corresponding Gaussian mean and variance parameters. The number of signal candidates was then obtained by subtracting the total background contribution inside the signal peak from the total number of candidates inside the peak. The background was estimated by one of two methods, ei-

TABLE V: Topological cuts used for Ω^- and $\bar{\Omega}^+$ reconstruction. In this table, \vec{r}_Ω , \vec{r}_Λ and \vec{r}_{PV} denote the Ω and Λ decay vertex position vectors and the primary vertex position vector in the STAR coordinate system, respectively. \vec{p}_Λ and \vec{p}_Ω are the reconstructed Λ and Ω momentum vectors. Cuts were optimized for each energy to reduce the combinatorial background.

Cut	Ω^-	$\bar{\Omega}^+$
DCA of Ω to primary vertex	< 0.4 cm; < 0.5 cm (19.6 GeV)	< 0.4 cm; < 0.5 cm (19.6 and 27 GeV); < 0.6 cm (7.7 GeV)
DCA of bachelor K to primary vertex	> 1.0 cm	> 1.0 cm
DCA of Λ to primary vertex	> 0.4 cm; > 0.3 cm (7.7 GeV)	> 0.4 cm; > 0.3 cm (7.7 GeV)
DCA of Λ -daughter p to primary vertex	> 0.6 cm	> 0.6 cm
DCA of Λ -daughter π to primary vertex	> 2.0 cm	> 2.0 cm
DCA between Λ and bachelor K	< 0.7 cm	< 0.7 cm; < 1.0 cm (7.7 GeV)
DCA between Λ -daughters	< 0.7 cm	< 0.7 cm; < 1.0 cm (7.7 GeV)
Ω decay length	> 3.0 cm; > 2.0 cm (7.7 GeV)	> 3.0 cm; > 2.0 cm (7.7 and 11.5 GeV)
Λ decay length	> 5.0 cm; > 4.0 cm (7.7 GeV)	> 5.0 cm; > 4.0 cm (7.7 and 11.5 GeV)
Λ decay length – Ω decay length	> 0	> 0
$(\vec{r}_\Lambda - \vec{r}_\Omega) \cdot \vec{p}_\Lambda$	> 0	> 0
$(\vec{r}_\Omega - \vec{r}_{PV}) \cdot \vec{p}_\Omega$	> 0	> 0
$(\vec{r}_\Omega - \vec{r}_{PV}) \times \vec{p}_\Omega / \vec{r}_\Omega - \vec{r}_{PV} / \vec{p}_\Omega $	< 0.12 ; < 0.15 (7.7 and 11.5 GeV)	< 0.12 ; < 0.15 (7.7 and 11.5 GeV)

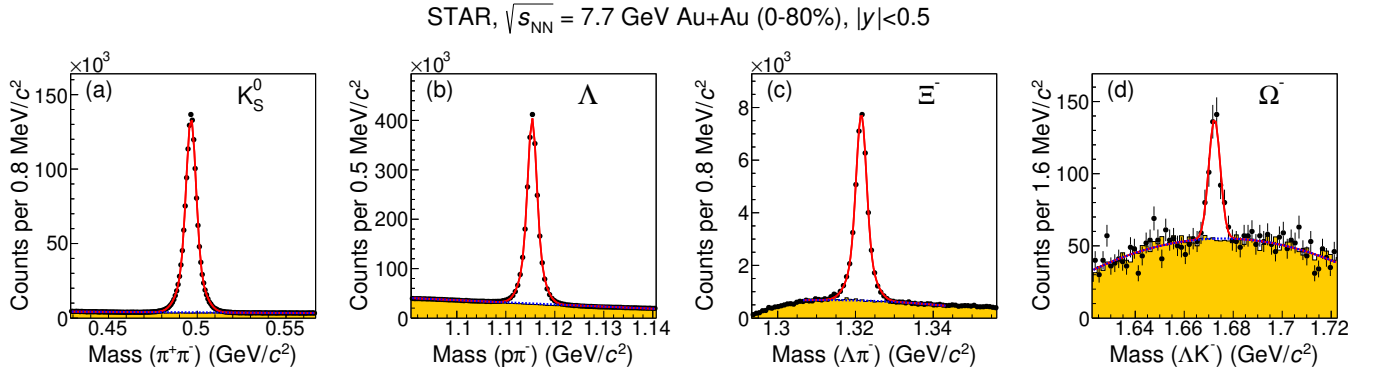


FIG. 1: K_S^0 , Λ , Ξ^- , and Ω^- invariant mass distributions in Au+Au collisions at $\sqrt{s_{NN}} = 7.7$ GeV. The red solid lines represent the function fit results (double Gaussian plus polynomial for K_S^0 , Λ , and Ξ^- ; single Gaussian plus polynomial for Ω^-), and the blue dashed lines are the fitted background contributions. The orange area shows the corresponding rotational background distribution.

ther by integrating the polynomial functions from the fitting, or by using a side-band method at higher p_T bins where a reasonable fitting cannot be achieved due to low statistics. The two side-bands on either side of the signal peak are selected symmetrically. Both have the half-peak width and are $5\sigma(6\sigma)$ away from the mean for $V^0(\Xi)$. For Ω , the total number of signal counts was obtained by the side-band method, with the invariant mass peak position and width determined from the embedding simulation data. For the ϕ meson, the invariant mass distribution was fitted with the non-relativistic Breit-Wigner function for signal plus a polynomial function (up to second order) for the residual background.

B. Acceptance and reconstruction efficiency

Whether a strange hadron can be observed by STAR is determined by the geometrical TPC acceptance for its decay daughter particles, their tracking efficiencies, and the efficiency of subsequent strange hadron reconstruction with the daughter particle tracks. The tracking efficiency and the strange hadron reconstruction efficiency depend on the final state particle multiplicity, which ranges from a few tracks in peripheral collisions to about a thousand tracks in central collisions. Therefore, in STAR, the geometrical acceptance and reconstruction efficiencies for each analyzed particle species were calcu-

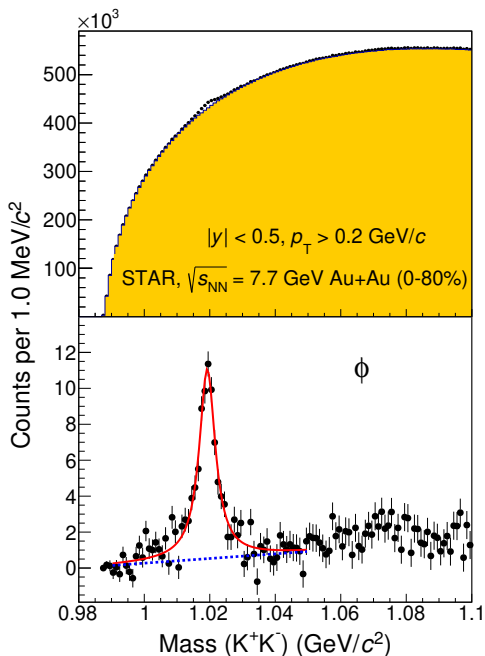


FIG. 2: ϕ invariant mass distributions in Au+Au collisions at $\sqrt{s_{\text{NN}}} = 7.7$ GeV. The upper panel shows the unlike-charge invariant mass distribution (full points) and the mixed-event background (orange area). The lower panel shows the invariant mass distribution after subtracting the background. The red solid line represents the function fit result (Breit-Wigner plus polynomial), and the blue dashed line is the fitted residual background contributions.

lated using an embedding technique, in which the simulated Monte Carlo (MC) particles sampled in a given kinematic range were embedded into real events where their efficiency was studied. The number of embedded particles per event was about 5% of the measured particle multiplicity for a given event. Embedded particles were all taken to originate from the real primary vertex in an event. The subsequent strange hadron propagation through STAR, strange hadron decay, and daughter particle propagation were simulated with the GEANT package [72]. The TPC detector response to the charged daughter particles was simulated with the STAR TPC response simulator (TpcRS). The simulated electronics signals were mixed with those from the real event and processed with the STAR tracking, event reconstruction, and strange hadron reconstruction algorithms. The acceptance and reconstruction efficiency were obtained by dividing the number of reconstructed MC strange hadrons by that of input MC in a certain kinematic range. As an example, the calculated efficiencies for different strange particles in central and peripheral Au+Au collisions at 39 GeV are shown as a function of p_{T} in Fig. 3. Generally, the efficiencies increase toward peripheral collisions and lower collision energies due to decreasing track multiplicities. As shown in Fig. 3, at 39 GeV, from 0–5% to 60–80% collisions, the K_{S}^0 efficiency increases by $\approx 47\%$ at

$p_{\text{T}} \approx 0.3$ GeV/ c and by $\approx 16\%$ at $p_{\text{T}} \approx 5.5$ GeV/ c , while the Ξ^- efficiency increases by $\approx 95\%$ at $p_{\text{T}} \approx 0.8$ GeV/ c and by $\approx 36\%$ at $p_{\text{T}} \approx 4.5$ GeV/ c . With the same analysis cuts, the efficiencies for antibaryons ($\bar{\Lambda}$, $\bar{\Xi}^+$ and $\bar{\Omega}^+$, not shown in Fig. 3 for clarity) are very similar to those of the respective baryons. The efficiency for Ω is much smaller than for Ξ is due to both the relatively tighter analysis cuts in Ω reconstruction and the decay-in-flight of the bachelor kaon.

C. Weak decay feed-down correction for Λ

The reconstructed Λ hyperons with the cuts listed in Table III contain both the prompt components originating from the primary vertex and the secondary components from the weak decays of Ξ , Ξ^0 , and Ω hyperons. The Λ hyperons from the electromagnetic decay of Σ^0 hyperons are also considered to be prompt since they are not experimentally distinguishable from those directly originating from the Au+Au collisions. Naturally, the Λ hyperons from secondary weak-decay vertices have different distributions in the topological cut variables—for example, the DCA of V^0 to the primary vertex. Hence it is mandatory to subtract their contributions to the reconstructed Λ yields before applying the acceptance and reconstruction efficiency corrections described in Sec. III B. As shown in Table III, a tight cut on the DCA of Λ candidates to the primary vertex was used to reduce the secondary contributions. However, some fraction of the secondary Λ hyperons still passed this criterion, especially in the high p_{T} regions. Their contribution to the prompt Λ sample was further evaluated with the help of the Ξ and Ξ^0 MC embedding data. With these MC data, the prompt Λ reconstruction cuts (in Table III) were applied to the reconstruction of the secondary MC Λ particles from the MC Ξ or Ξ^0 decays. Then the total number of reconstructed secondary MC Λ particles was scaled according to the corrected yields of the measured Ξ and Ξ^0 particles (assuming that Ξ^0 has the same yield as Ξ since it cannot be measured by STAR). Those scaled values represent the feed-down contribution and are subtracted from the raw Λ yields. The relative contribution of the secondary Λ from Ξ and Ξ^0 decays was calculated for each p_{T} interval in each collision centrality at each collision energy. For Au+Au collisions at 39 GeV, in 0–5% central collisions, the relative feed-down contribution of Ξ and Ξ^0 to Λ ($\bar{\Lambda}$) ranges from $\approx 23\%$ (30%) at $p_{\text{T}} = 0.5$ GeV/ c to $\approx 6\%$ (15%) at $p_{\text{T}} = 5$ GeV/ c . For Λ , the feed-down contribution from Ω^- decay was not considered since it is expected to be negligible ($< 1\%$) due to the low yield of Ω^- relative to that of Λ in the BES energy range. For $\bar{\Lambda}$, the feed-down contribution from $\bar{\Omega}^+$ cannot be neglected due to the significantly larger ratio of $\bar{\Omega}^+$ to $\bar{\Lambda}$ yield in more central collisions, at lower p_{T} , and especially at lower BES energies. The $\bar{\Omega}^+$ feed-down contribution to $\bar{\Lambda}$ was evaluated with the $\bar{\Omega}^+$ embedding

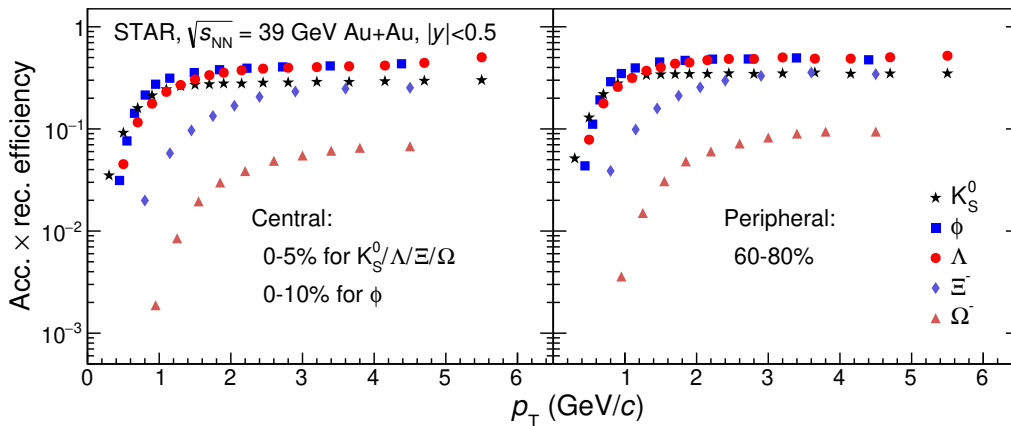


FIG. 3: Geometrical acceptance and reconstruction efficiency of various strange hadrons at mid-rapidity ($|y| < 0.5$) in central (left) and peripheral (right) Au+Au collisions at $\sqrt{s_{NN}} = 39$ GeV. The branching ratios of measured decay channels are not taken into account here.

data and the corrected yields of the measured $\bar{\Omega}^+$ particle at five BES energies, and subtracted from the corresponding raw $\bar{\Lambda}$ yields. In 0–5% central collisions and at $p_T = 0.5$ GeV/c, the relative feed-down contribution of $\bar{\Omega}^+$ to $\bar{\Lambda}$ increases from $\lesssim 3\%$ to $\approx 6\%$ with the collision energy decreasing from 39 to 7.7 GeV.

D. p_T spectra extrapolation at low p_T

The p_T spectra of each strange hadron were obtained by dividing the raw yield in a certain p_T interval by the corresponding acceptance and reconstruction efficiencies presented in Sec. III B. Due to limited detector acceptance at low p_T , and finite statistics at high p_T , the spectra were not measured in these regions and hence needed to be extrapolated to these two regions in order to obtain the p_T integrated yield (dN/dy) as well as the averaged transverse mass ($\langle m_T \rangle - m_0$), where $m_T = \sqrt{p_T^2 + m_0^2}$ is the transverse mass and m_0 is the rest mass. The extrapolation to low p_T is particularly important, since it contributes significantly to both observables, while the extrapolation to high p_T usually provides a much smaller contribution. The blast-wave model [73] can be used for fitting individually the low p_T spectra and extrapolating them to the unmeasured lower p_T region. This model assumes the particles are emitted from a radially expanding thermal source. A common kinetic freeze-out temperature T and a transverse radial flow velocity profile $\beta = \beta_S(r/R)^n$ are used to characterize the source, where β_S is the surface velocity, r/R is the relative radial position in the source, and n is the exponent of flow velocity profile. The p_T distribution of the particles is given by

$$\frac{d^2 N}{2\pi p_T dp_T dy} \propto \int_0^R r dr m_T I_0 \left(\frac{p_T \sinh \rho(r)}{T} \right) \times K_1 \left(\frac{m_T \cosh \rho(r)}{T} \right), \quad (3)$$

where $\rho(r) = \tanh^{-1} \beta$, I_0 and K_1 are the modified Bessel functions. The velocity profile parameter n is set to 1 for all the blast-wave model fitting in this analysis.

The K_S^0 low p_T spectra can be well fitted and hence extrapolated to unmeasured lower p_T regions (< 0.2 GeV/c) with the blast-wave model and other two functions, the exponential function

$$\frac{d^2 N}{2\pi p_T dp_T dy} \propto e^{-\frac{m_T}{T}}, \quad (4)$$

and the Levy function

$$\frac{d^2 N}{2\pi p_T dp_T dy} \propto \left(1 + \frac{m_T - m_0}{nT} \right)^{-n}. \quad (5)$$

In this analysis, all three functions were used to fit the low p_T K_S^0 spectra at all energies. For the blast-wave model, the fit range was [0.2, 1.4] GeV/c. For the exponential function, the fit range was [0.2, 1.8] GeV/c at 7.7 and 11.5 GeV, and [0.2, 1.2] GeV/c at 19.6, 27, and 39 GeV. For the Levy function, the fit range was [0.2, 1.4] GeV/c at 7.7 and 11.5 GeV, and [0.2, 2.0] GeV/c at 19.6, 27, and 39 GeV. The difference between the results from these three functions were considered in the systematic errors for dN/dy and $\langle m_T \rangle - m_0$ due to low p_T extrapolation. The Levy function produced a slightly better fit (lower χ^2) for centralities within 30–80%, while the blast-wave model was better for centralities within 0–30% at all five collision energies. Therefore, the default K_S^0 dN/dy and $\langle m_T \rangle - m_0$ values were calculated with those functions correspondingly. Although the contribution is almost negligible ($< 10^{-5}$), the K_S^0 spectra at intermediate-and-above p_T were fitted separately and extrapolated to unmeasured higher p_T regions (up to 10 GeV/c) with the Levy function for energies above 19.6 GeV or exponential function for energies below 11.5 GeV.

For the Λ and Ξ hyperons, the low p_T spectra can be well fit and hence extrapolated to unmeasured lower p_T

regions (<0.4 GeV/ c for Λ and <0.6 GeV/ c for Ξ) with the blast-wave model. A Boltzmann function

$$\frac{d^2N}{2\pi p_T dp_T dy} \propto m_T e^{-\frac{m_T}{T}}, \quad (6)$$

and the exponential function were used as alternatives for estimation of the systematic error due to low p_T extrapolation. For Λ and $\bar{\Lambda}$, the blast-wave model fit range was $[0.4, 3.0]$ GeV/ c for centralities within 0–60% and $[0.4, 2.0]$ GeV/ c for the centrality of 60–80%. The fit ranges for Boltzmann and exponential functions were both $[0.4, 1.4]$ GeV/ c at all energies and slightly narrowed to $[0.4, 1.2]$ GeV/ c for Λ in central collisions (0–10%) at 7.7 GeV. For Ξ^- and Ξ^+ , the blast-wave model fit range was $[0.6, 3.2]$ GeV/ c for most centralities, and narrowed to $[0.6, 2.2]$ GeV/ c in 60–80% for all energies and in 40–60% for 7.7 and 11.5 GeV. The fit ranges for Boltzmann and exponential functions were $[0.6, 2.2]$ GeV/ c at 11.5 GeV, and $[0.6, 2.6]$ GeV/ c at 19.6, 27, and 39 GeV. At 7.7 GeV, the fit range was $[0.6, 3.2]$ GeV/ c for Ξ^- , and $[0.6, 2.6]$ GeV/ c for Ξ^+ , when data are available. The high p_T extrapolation was done using the Levy function for 19.6 GeV and higher energies, and the exponential function for 7.7 and 11.5 GeV.

For the Ω hyperon, the exponential function was used to fit the spectra over the full measured p_T range (>0.7 GeV/ c for 7.7 GeV and >0.8 GeV/ c for ≥ 11.5 GeV), while the Boltzmann function and the blast-wave model were used alternatively for estimation of the systematic error due to low p_T extrapolation. For the ϕ meson, the Boltzmann function was found to fit better over the full measured p_T range ($[0.4, 1.7]$ GeV/ c at 7.7 GeV, and in the low p_T range ($[0.4, 2.0]$ GeV/ c) at 11.5 GeV and higher energies, and was therefore used to extrapolate the spectra into unmeasured low p_T regions (<0.4 GeV/ c). The Levy function was used as the alternative to fit the full measured p_T spectra for estimation of the systematic error due to low p_T extrapolation.

The integrated dN/dy was obtained by integrating the p_T spectra data in the measured p_T range and the fitted functions in the low p_T and high p_T extrapolated ranges. The contributions of low p_T extrapolation in the integrated dN/dy are ≈ 10 –14% for K_S^0 , ≈ 13 –28% for Λ and $\bar{\Lambda}$, ≈ 23 –46% for Ξ^- and Ξ^+ , ≈ 39 –46% for Ω^- and $\bar{\Omega}^+$, and ≈ 17 –28% for ϕ . These contributions are larger for lower energy and more peripheral collisions due to steeper p_T spectra there.

The $\langle m_T \rangle - m_0$ is obtained by integrating the whole range of the p_T spectra as follows

$$\langle m_T \rangle - m_0 = \frac{\int (m_T - m_0) \frac{dN}{dp_T} dp_T}{\int \frac{dN}{dp_T} dp_T}. \quad (7)$$

The same extrapolation functions used for the integrated dN/dy were used again to calculate the numerator integral of $\langle m_T \rangle - m_0$.

E. Systematic uncertainties

Many possible sources which can contribute to the systematic uncertainties of the p_T spectra were evaluated bin-by-bin in this analysis.

In the signal extraction for K_S^0 and Ξ , the side-band method and the fitting method with double-Gaussian-plus-polynomial functions were used in the estimation of background in the signal peak region. The difference was factored into the systematic uncertainty. The width of the signal peak was first determined with the double-Gaussian-plus-polynomial fitting and then varied to estimate its contribution to the systematic error. The shape of the signal peak in high p_T bins deviated from the symmetric Gaussian shape due to cuts on decay length, which cause a systematic deviation in our signal counting method. The embedding data were used to estimate this deviation, which reaches $\approx 7\%$ (3%) at ≈ 5 GeV/ c for K_S^0 (Ξ). For Λ , the side-band method cannot be used in estimating the systematic error due to the non-linear residual background shape. Hence the fitting ranges have been changed to account for the possible uncertainty in the background shape. Furthermore, the width of the signal peak was varied in the estimation of the systematic uncertainty. The deviation of the signal peak shape from Gaussian was estimated with MC simulation and its contribution to the systematic uncertainty found to be $\approx 8\%$ for p_T at ≈ 5 GeV/ c . Different sources were assumed to be uncorrelated and hence summed quadratically to obtain the total systematic uncertainty in signal extraction, summarized in Table VI.

The daughter particle identification cuts, $|n\sigma|$, were varied from their default value of 4.0 to 3.6. The contribution of this cut to the systematic uncertainty in the p_T spectra is small, as listed in Table VI. The systematic uncertainties due to tracking were estimated by varying the cuts on the minimum number of hit points from the default 16 to 26, the minimum number of hit points used for dE/dx calculation from 1 to 10, and the minimum ratio of the number of hit points to the number of possible hit points from 0.45 to 0.55. The cut values were changed one at a time, and the raw yields and the efficiencies were both re-calculated accordingly to obtain the corrected p_T spectra. The maximum deviations from the default spectra due to these three tracking cuts were accounted for in the systematic uncertainties. The minimum number of daughter hit points contributes the majority of the systematic error in tracking, while the other two cuts contribute minimally. The tracking uncertainties are larger at lower p_T in more central collisions. For K_S^0 and $\bar{\Lambda}$ in the most central Au+Au collisions at 39 GeV, the uncertainties are $\approx 6\%$ and $\approx 7\%$, respectively, at $p_T = 0.5$ GeV/ c , while for Ξ^+ , the uncertainty is $\approx 15\%$ at $p_T = 0.8$ GeV/ c , as listed in Table VI.

The topological cuts were also varied one after another to study the systematic deviations of the p_T spectra. For example, for K_S^0 , the radial decay length cut value was varied in the range of $[2.5, 3.3]$ cm; the DCA of daughters

TABLE VI: Summary of systematic uncertainties for p_T spectra. The range indicates the variation between p_T bins, centralities and energies.

Components	K_S^0	Λ and $\bar{\Lambda}$	Ξ^- and Ξ^+
Signal extraction	0.5–8%	0.5–9%	0.5–3%
Particle identification	< 0.5%	< 0.7%	< 1.1%
Tracking	1.5–7%	1.5–7%	3–15%
Topological reconstruction	1–4%	2–10%	3–8%
Detector uniformity	1–4%	1–7%	1–8%
Weak decay feed-down correction	n.a.	Λ : 0.2–4%; $\bar{\Lambda}$: 0.5–10%	n.a.
Total uncertainty	2–9%	Λ : 3–12%; $\bar{\Lambda}$: 3–13%	4–17%

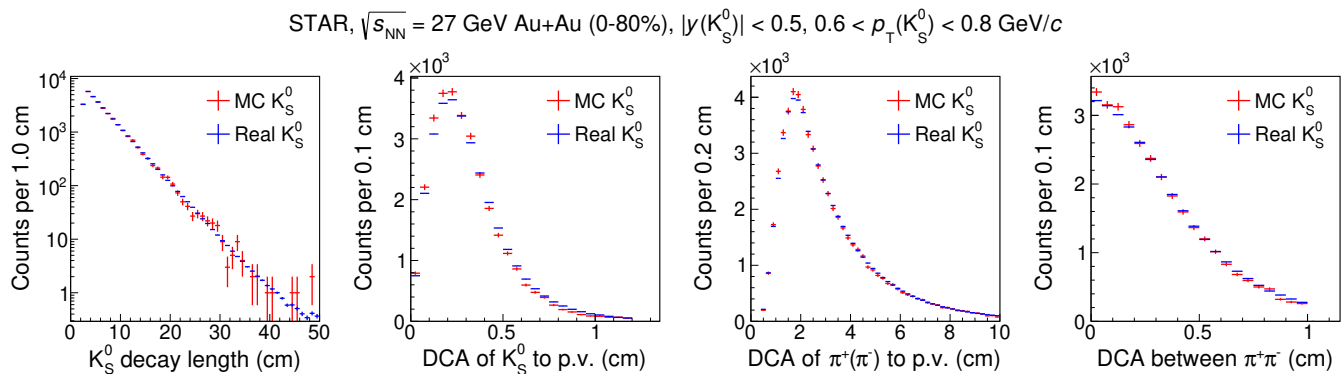


FIG. 4: MC and data comparison on K_S^0 topological variable distributions in Au+Au collisions at $\sqrt{s_{\text{NN}}} = 27$ GeV. A same set of loose selection cuts have been applied to both MC and data. The distributions of combinatorial backgrounds in data were estimated with the rotation method and subtracted, and then the resulting data distributions were scaled down to match with the MC statistics.

in $[0.55, 0.8]$ cm; the DCA of K_S^0 in $[0.6, 1.2]$ cm; and the DCA between daughters in $[0.5, 1.0]$ cm. As listed in Table VI, the systematic uncertainties from these geometric cuts were generally small compared to the tracking uncertainties, except in the most peripheral collisions, where the primary vertex resolution is worse due to low track multiplicity. The same method was applied to the systematic error studies for Ξ and Λ spectra despite there being more topological cuts involved. For these two hyperons, the systematic errors due to these cuts were also small compared to the tracking uncertainties except in the most peripheral collisions. Furthermore, a detailed comparison of the distributions of the topological cut variables from the embedded strange particles and those from strange particles reconstructed in real data was performed. A good agreement between the MC simulation and the data was achieved. As an example, Fig. 4 shows the comparison of K_S^0 MC topological variable distributions and those from the real data at 27 GeV. The systematic uncertainties due to the veto cuts in the K_S^0 and Ξ analyses were also studied and found to be negligible for both particles.

Non-uniformity of the detector acceptance for collisions at different primary vertex positions along the beam direction and detector asymmetry between forward and backward halves may contribute to the systematic uncertainty due to imperfect detector response simulation

and limited real data sample size ($\approx 10^5$ MB events) in embedding simulation. The $|z_{\text{vertex}}|$ cut was varied to study the acceptance uniformity. At 39 GeV, the cut was changed within $[25, 40]$ cm. The resulting change in the p_T spectra and hence the systematic error from this source is negligible. The default rapidity range $[-0.5, 0.5]$ was also divided into a forward half, $[0, 0.5]$, and a backward half, $[-0.5, 0]$. The maximum deviations of the resulting p_T spectra from the default were accounted for in the systematic errors, which can be as large as 8% for Ξ in some p_T bins.

For Λ , the systematic errors due to feed-down corrections were evaluated by propagating the Ξ systematic errors to the Λ raw yield, according to the fractions of those feed-down contributions. For simplicity, the Ξ systematic errors were assumed to be uncorrelated with those for Λ . For $\bar{\Lambda}$, the systematic error due to feed-down contribution from $\bar{\Omega}^+$ was evaluated similarly and directly summed with that from Ξ in each p_T bin.

The final bin-by-bin systematic error for a p_T spectrum was a quadratic sum of all the above sources, assuming that they are fully uncorrelated, and is summarized in Table VI. For the details on the systematic uncertainties for the ϕ and Ω p_T spectra, please refer to Ref. [11].

For the integrated yield, dN/dy , the systematic error in the measured p_T range is simply the sum of the bin-by-bin systematic errors assuming that they are fully cor-

related. In the extrapolated low p_T region, the systematic errors were estimated considering several potential sources. First, if a change in the analysis cuts produces a change in the p_T spectra, then the extrapolation will also change. Second, different fit functions will have different extrapolations even under the same analysis cuts. Both sources were studied in this analysis by changing the analysis cuts and fitting functions in the extrapolation. The final extrapolated systematic errors are a quadratic sum of both contributions. The systematic error of $\langle m_T \rangle - m_0$ was estimated in a similar manner by varying analysis cuts, extrapolation functions, and signal extraction methods. The systematic uncertainties of antibaryon-to-baryon ratios and nuclear modification factors were estimated independently by considering all the sources mentioned above, and hence the possible correlations of the systematic errors in the numerator and the denominator in these observables were consistently treated.

IV. RESULTS AND DISCUSSIONS

A. Transverse momentum spectra

Figures 5, 6, 7, 8, 9, 10, 11, and 12 show the transverse momentum spectra of K_S^0 , Λ , $\bar{\Lambda}$, Ξ^- , $\bar{\Xi}^+$, Ω^- , $\bar{\Omega}^+$, and ϕ at mid-rapidity ($|y| < 0.5$) in different collision centralities from Au+Au collisions at $\sqrt{s_{NN}} = 7.7, 11.5, 19.6, 27, \text{ and } 39$ GeV. The p_T spectra of Ω^- , $\bar{\Omega}^+$ and ϕ are the same as those shown in Ref. [11]. All the p_T spectra shown here have been corrected for geometrical acceptance and reconstruction efficiency, as discussed in Sec. III B. For better visualization, the spectra are scaled by factors of 10 from central to peripheral collisions. The $\Lambda(\bar{\Lambda})$ spectra are corrected for the feed-down contribution from weak decays of Ξ and Ξ^0 baryons using the measured Ξ spectra shown in Fig. 8 or 9. The $\bar{\Lambda}$ spectra are further corrected for the feed-down contribution from weak decays of $\bar{\Omega}^+$ baryons using the measured $\bar{\Omega}^+$ spectra shown in Fig. 11. For more details on feed-down correction, please see Sec. III C. The systematic errors of invariant yields, described in details in Sec. III E, are shown as vertical gray bands in these figures for each p_T bin. They become generally larger toward more central collisions due to larger track multiplicities. The systematic uncertainties at 19.6 and 27 GeV are less than those at the other three energies due to the better data taking conditions in 2011 than in 2010. The default function fit results at low p_T , described in details in Sec. III D, are plotted on top of each p_T spectrum in the corresponding fit range and the low p_T extrapolation range.

B. Averaged transverse mass

The averaged transverse mass, $\langle m_T \rangle - m_0$, can be calculated by Eq. 7 with the measured and extrapolated p_T spectra of a certain particle species. Its energy and centrality dependence reflects the change of the p_T spectra shapes with collision conditions, and hence provides information regarding the reaction dynamics among the constituents of the colliding systems. Figures 13, 14, and 15 show the $\langle m_T \rangle - m_0$ at mid-rapidity ($|y| < 0.5$) for K_S^0 , Λ , $\bar{\Lambda}$, Ξ^- , $\bar{\Xi}^+$, Ω^- , $\bar{\Omega}^+$, and ϕ as a function of $\langle N_{\text{part}} \rangle$ for Au+Au collisions at $\sqrt{s_{NN}} = 7.7, 11.5, 19.6, 27, \text{ and } 39$ GeV. The $\langle m_T \rangle - m_0$ value of each particle species increases with the increasing $\langle N_{\text{part}} \rangle$ at all energies, indicating the gradual development of collective motion with the increasing medium volume. The $\langle m_T \rangle - m_0$ increases faster toward central collisions for hyperons in general, and for Λ in particular, than that of K_S^0 . In contrast, the $\langle m_T \rangle - m_0$ values of multistrange hyperons, Ξ and Ω , seem to be consistent with that of ϕ meson within the uncertainties. Additionally, at lower collision energies and toward central collisions, the $\langle m_T \rangle - m_0$ of anti-hyperons becomes larger than that of hyperons. The difference is most sizable for $\bar{\Lambda}$ and $\bar{\Lambda}$ in the most central Au+Au collisions at $\sqrt{s_{NN}} = 7.7$ GeV. This phenomenon could be explained by the larger possibility for a lower p_T $\bar{\Lambda}$ to be annihilated in a Λ -hyperon-rich medium created in such collisions. Generally, in a thermodynamic system with bulk expansion, the radial flow of a particle is only dependent on its mass. In contrast, the split of $\langle m_T \rangle - m_0$ between antibaryon and baryon indicates that the spectra may not be driven only by the bulk expansion with a common velocity. Hadronic processes, such as baryon-antibaryon annihilations, might also have significant impacts in the final hadron productions at this lower energy region.

Figure 16 shows the $\langle m_T \rangle - m_0$ at mid-rapidity ($|y| < 0.5$) for K_S^0 , Λ , $\bar{\Lambda}$, Ξ^- , and $\bar{\Xi}^+$ from 0–5% central Au+Au collisions at $\sqrt{s_{NN}} = 7.7\text{--}39$ GeV. The previously measured $\langle m_T \rangle - m_0$ for Λ , $\bar{\Lambda}$, Ξ^- , and $\bar{\Xi}^+$ in central Pb+Pb collisions at $\sqrt{s_{NN}} = 6.3\text{--}17.3$ GeV at SPS [24] and in central Au+Au collisions at $\sqrt{s_{NN}} = 130$ GeV at RHIC [29, 31] are also shown for comparison. In general, the STAR BES $\langle m_T \rangle - m_0$ for Λ , $\bar{\Lambda}$, Ξ^- , and $\bar{\Xi}^+$ shows a trend similar to previous measurements but with much smaller uncertainties. The $\langle m_T \rangle - m_0$ values for K_S^0 , Λ , Ξ^- , and $\bar{\Xi}^+$ show an increasing trend with the increasing collision energy. However, the $\langle m_T \rangle - m_0$ value for $\bar{\Lambda}$ at $\sqrt{s_{NN}} = 7.7$ GeV seems to be as large as the value at $\sqrt{s_{NN}} = 11.5$ GeV, and apparently breaks the monotonous increasing trend. This observation again may indicate the significant impact of annihilation processes on antibaryon production in a baryon-rich QCD matter.

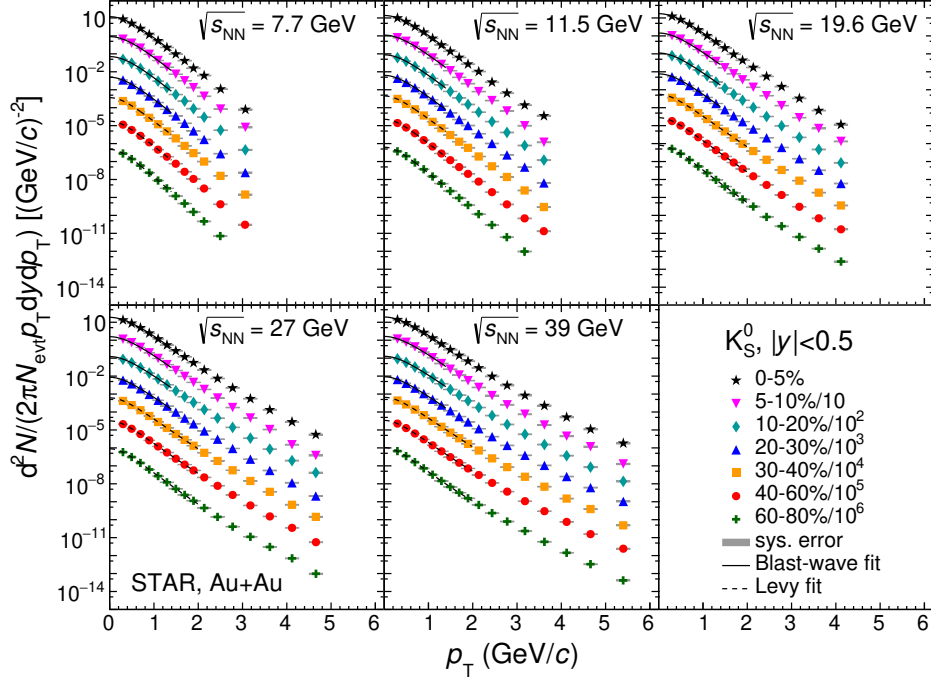


FIG. 5: The transverse momentum spectra of K_S^0 at mid-rapidity ($|y| < 0.5$) from Au+Au collisions at different centralities and energies ($\sqrt{s_{NN}} = 7.7\text{--}39$ GeV). The data points are scaled by factors of 10 from central to peripheral collisions for clarity. The vertical gray bands represent the systematic errors, which are small hence the bands look like horizontal bars. The blast-wave model (or Levy function) fit results are shown in the fit range and the low p_T extrapolation range as solid (dashed) lines for the centrality bins within 0–30% (30–80%).

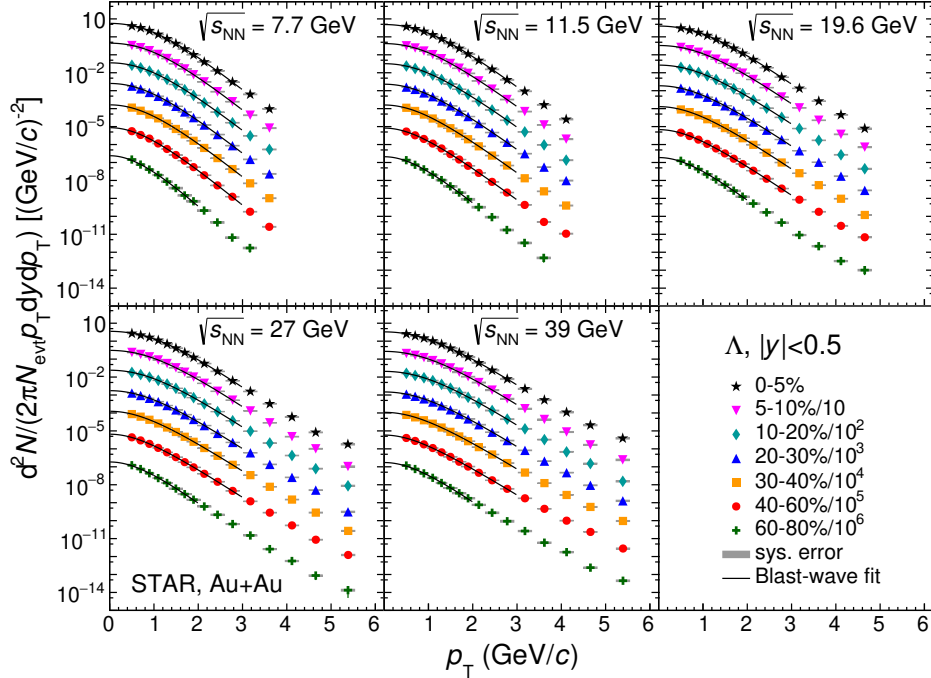


FIG. 6: The transverse momentum spectra of Λ at mid-rapidity ($|y| < 0.5$) from Au+Au collisions at different centralities and energies ($\sqrt{s_{NN}} = 7.7\text{--}39$ GeV). The spectra are corrected for the feed-down of Ξ^- and Ξ^0 decays. The data points are scaled by factors of 10 from central to peripheral collisions for clarity. The vertical gray bands represent the systematic errors, which are small hence the bands look like horizontal bars. The blast-wave model fit results are shown in the fit range and the low p_T extrapolation range as solid lines for all centrality bins.

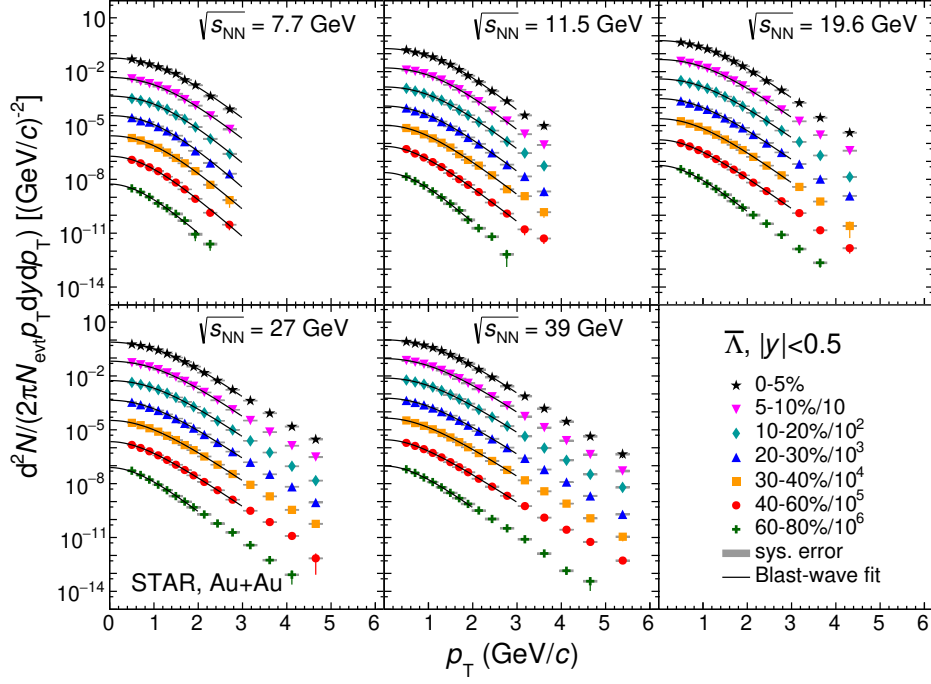


FIG. 7: The transverse momentum spectra of $\bar{\Lambda}$ at mid-rapidity ($|y| < 0.5$) from Au+Au collisions at different centralities and energies ($\sqrt{s_{NN}} = 7.7\text{--}39$ GeV). The spectra are corrected for the feed-down of Ξ^+ , Ξ^0 , and Ω^+ decays. The data points are scaled by factors of 10 from central to peripheral collisions for clarity. The vertical gray bands represent the systematic errors, which are small hence the bands look like horizontal bars. The blast-wave model fit results are shown in the fit range and the low p_T extrapolation range as solid lines for all centrality bins.

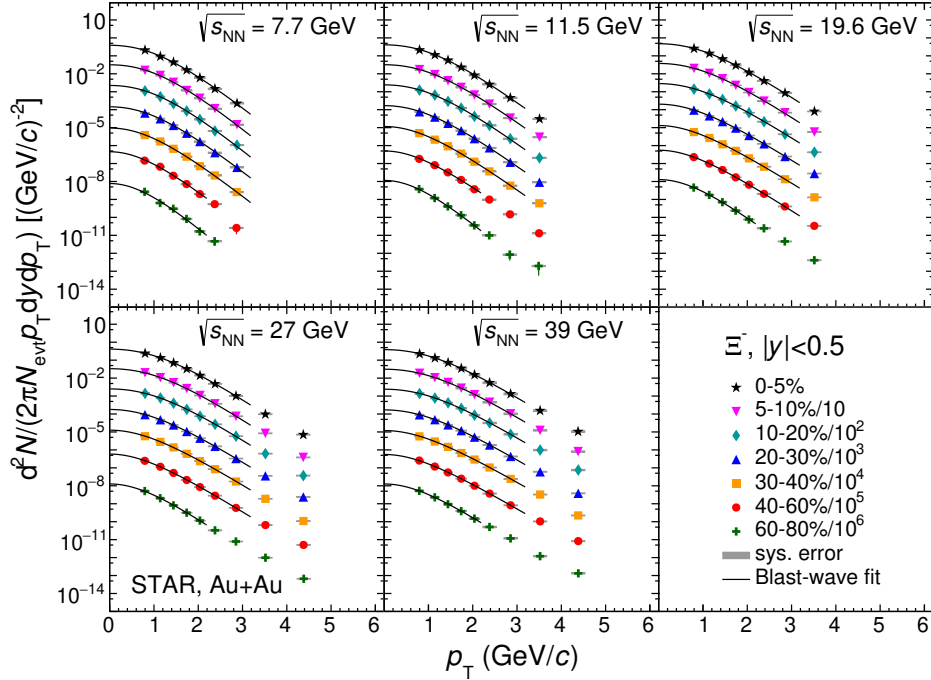


FIG. 8: The transverse momentum spectra of Ξ^- at mid-rapidity ($|y| < 0.5$) from Au+Au collisions at different centralities and energies ($\sqrt{s_{NN}} = 7.7\text{--}39$ GeV). The data points are scaled by factors of 10 from central to peripheral collisions for clarity. The vertical gray bands represent the systematic errors, which are small hence the bands look like horizontal bars. The blast-wave model fit results are shown in the fit range and the low p_T extrapolation range as solid lines for all centrality bins.

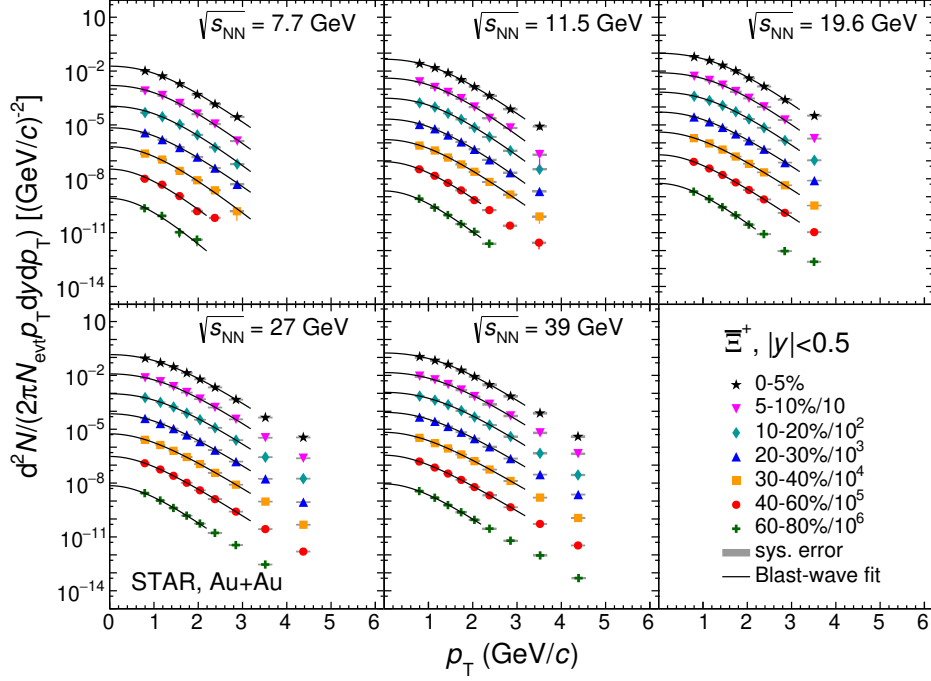


FIG. 9: The transverse momentum spectra of Ξ^+ at mid-rapidity ($|y| < 0.5$) from Au+Au collisions at different centralities and energies ($\sqrt{s_{NN}} = 7.7\text{--}39$ GeV). The data points are scaled by factors of 10 from central to peripheral collisions for clarity. The vertical gray bands represent the systematic errors, which are small hence the bands look like horizontal bars. The blast-wave model fit results are shown in the fit range and the low p_T extrapolation range as solid lines for all centrality bins.

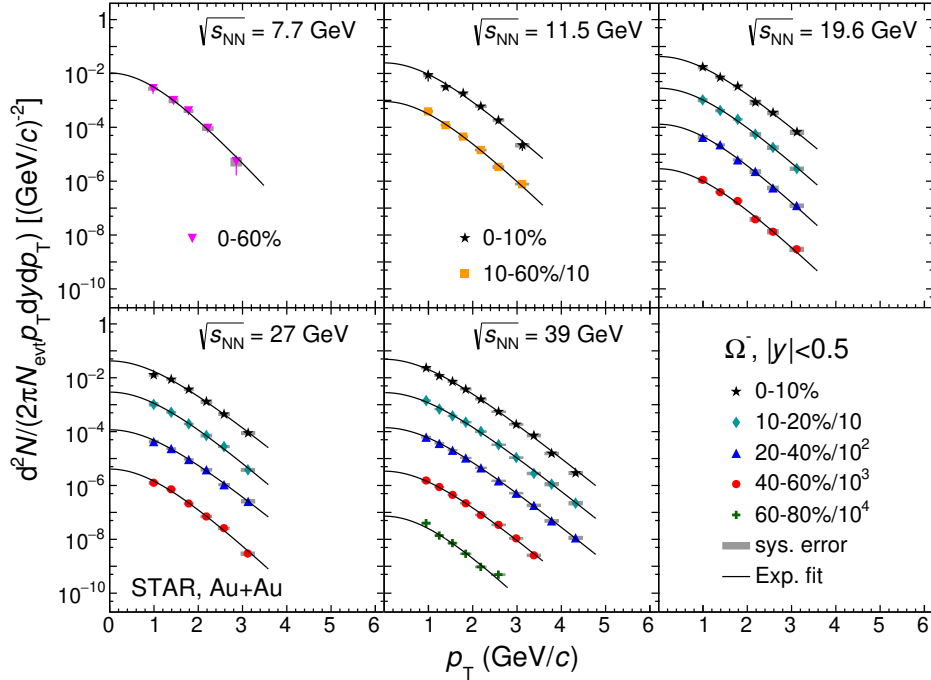


FIG. 10: The transverse momentum spectra of Ω^- at mid-rapidity ($|y| < 0.5$) from Au+Au collisions at different centralities and energies ($\sqrt{s_{NN}} = 7.7\text{--}39$ GeV). The data points are scaled by factors of 10 from central to peripheral collisions for clarity. The vertical gray bands represent the systematic errors, which are small hence the bands look like horizontal bars. The exponential function fit results are shown in the fit range and the low p_T extrapolation range as solid lines for all centrality bins.

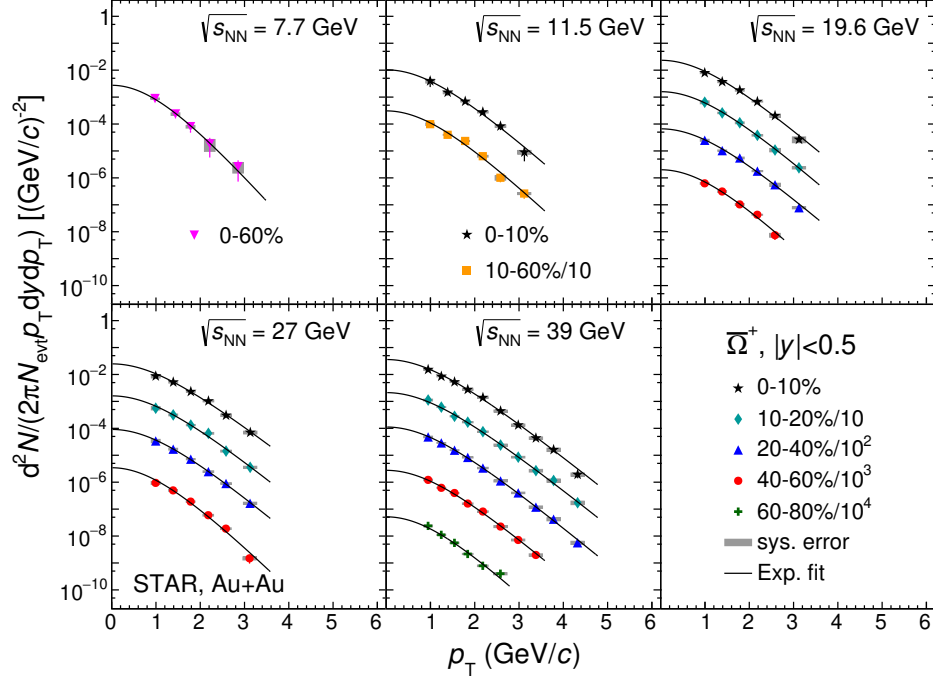


FIG. 11: The transverse momentum spectra of $\bar{\Omega}^+$ at mid-rapidity ($|y| < 0.5$) from Au+Au collisions at different centralities and energies ($\sqrt{s_{NN}} = 7.7\text{--}39$ GeV). The data points are scaled by factors of 10 from central to peripheral collisions for clarity. The vertical gray bands represent the systematic errors, which are small hence the bands look like horizontal bars. The exponential function fit results are shown in the fit range and the low p_T extrapolation range as solid lines for all centrality bins.

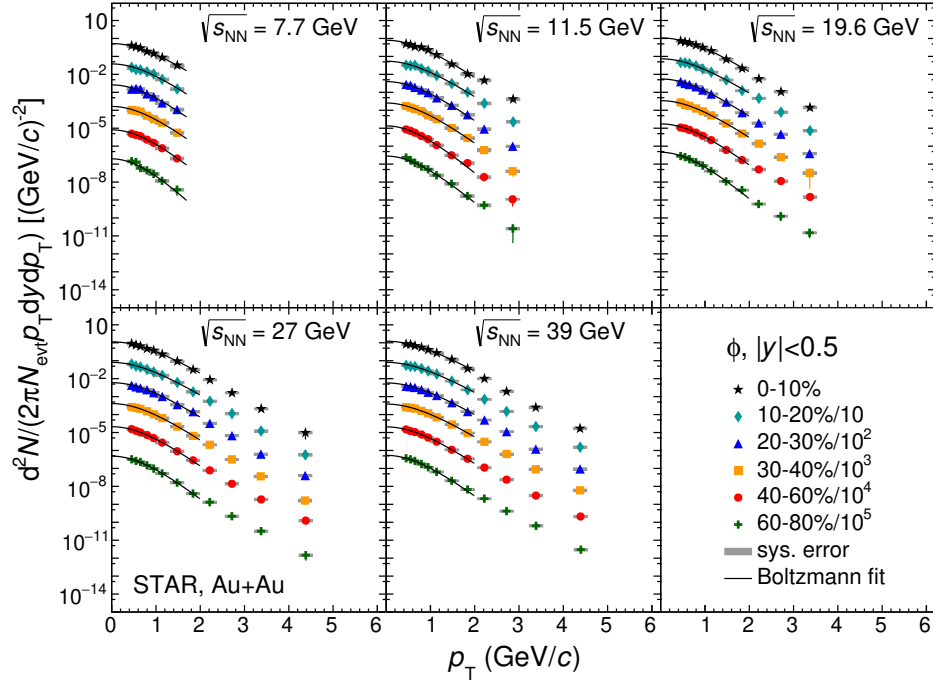


FIG. 12: The transverse momentum spectra of ϕ at mid-rapidity ($|y| < 0.5$) from Au+Au collisions at different centralities and energies ($\sqrt{s_{NN}} = 7.7\text{--}39$ GeV). The data points are scaled by factors of 10 from central to peripheral collisions for clarity. The vertical gray bands represent the systematic errors, which are small hence the bands look like horizontal bars. The Boltzmann function fit results are shown in the fit range and the low p_T extrapolation range as solid lines for all centrality bins.

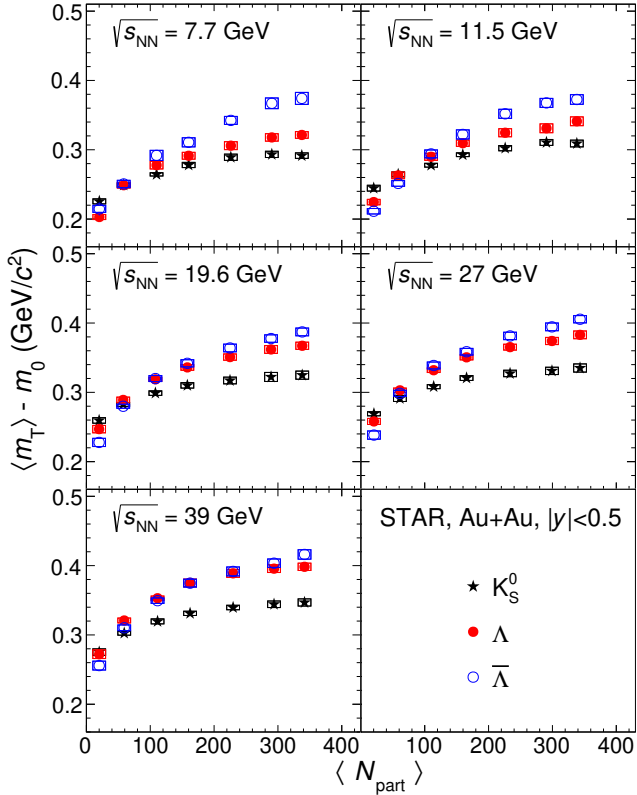


FIG. 13: The averaged transverse mass, $\langle m_T \rangle - m_0$, at mid-rapidity ($|y| < 0.5$) for K_S^0 , Λ , and $\bar{\Lambda}$ as a function of $\langle N_{\text{part}} \rangle$ for Au+Au collisions at $\sqrt{s_{\text{NN}}} = 7.7\text{--}39$ GeV. The box on each data point denotes the systematic error.

C. Particle yields

Figure 17 shows the collision centrality dependence of the integrated yield, dN/dy , per average number of participating nucleon pairs ($\langle N_{\text{part}} \rangle / 2$), of various strange hadrons (K_S^0 , Λ , $\bar{\Lambda}$, Ξ^- , Ξ^+ , Ω^- , $\bar{\Omega}^+$, and ϕ) at mid-rapidity ($|y| < 0.5$) from Au+Au collisions at $\sqrt{s_{\text{NN}}} = 7.7, 11.5, 19.6, 27, \text{ and } 39$ GeV. These normalized yields increase from peripheral to central collisions for all particle species except $\bar{\Lambda}$ at all collision energies. The $\bar{\Lambda}$ normalized yield has weak centrality dependence, and it even slightly decreases toward central collisions at lower collision energies. This is similar to the case of \bar{p} [7], indicating the larger impact of annihilation processes on antibaryon production in more central collisions.

Figures 18 and 19 show the collision energy dependence of the particle yield (dN/dy) at mid-rapidity ($|y| < 0.5$) for K_S^0 , Λ , $\bar{\Lambda}$, Ξ^- , and Ξ^+ from 0–5% central Au+Au collisions at $\sqrt{s_{\text{NN}}} = 7.7, 11.5, 19.6, 27, \text{ and } 39$ GeV, compared to the corresponding data from AGS E896/E891/E917, CERES, NA49, and NA57 in the similar energy range, as well as to the STAR and PHENIX data at higher collision energies. The NA49 and NA57 data are from central Pb+Pb collisions, and have been

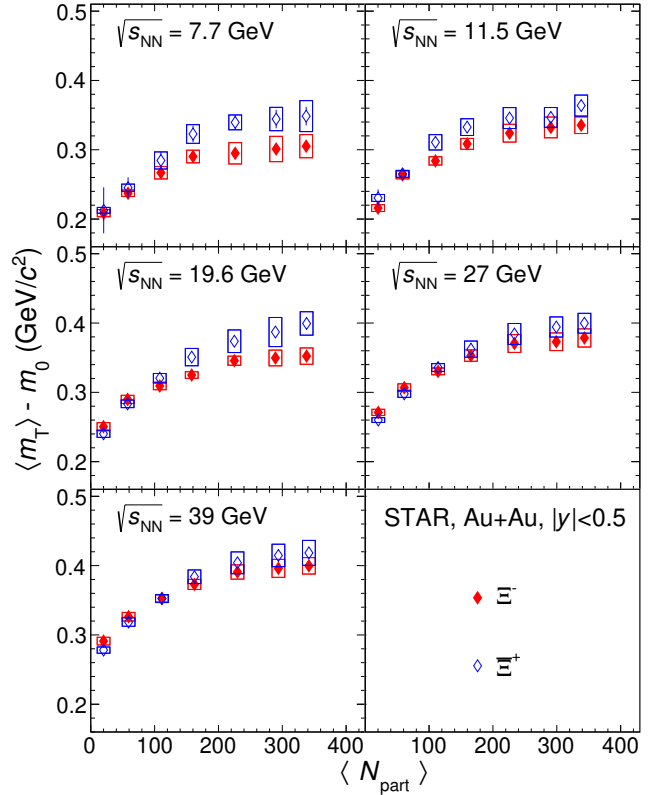


FIG. 14: The averaged transverse mass, $\langle m_T \rangle - m_0$, at mid-rapidity ($|y| < 0.5$) for Ξ^- and Ξ^+ as a function of $\langle N_{\text{part}} \rangle$ for Au+Au collisions at $\sqrt{s_{\text{NN}}} = 7.7\text{--}39$ GeV. The box on each data point denotes the systematic error.

re-scaled according to the estimated numbers of wounded nucleons, $\langle N_W \rangle$. The scale factor is $\langle N_{\text{part}} \rangle / \langle N_W \rangle$, where $\langle N_{\text{part}} \rangle$ is the average number of participants in 0–5% central Au+Au collisions in STAR (see Table II). The E917 data and the STAR K_S^0 and Ξ data at $\sqrt{s_{\text{NN}}} = 130$ GeV have been re-scaled in a similar manner to account for the centrality difference between these measurements and the STAR BES. Figures 18 and 19 show that the STAR BES data lie on a trend established by the corresponding data from AGS, NA49, NA57, CERES and previous STAR data, though there seems to be an obvious non-monotonic energy dependence in the Λ dN/dy when connecting the STAR BES data with the previous STAR measurements at higher energies. In the BES energy range, while the STAR BES data and the NA49 data are consistent within uncertainties in general except the slight difference in Λ yield at $\sqrt{s_{\text{NN}}} = 17.3$ GeV, the NA57 data are significantly higher for all particle species at both energies except the Ξ^+ yield with large uncertainty at $\sqrt{s_{\text{NN}}} = 8.7$ GeV.

As shown in Fig. 19, the yields of anti-hyperons increase rapidly with increasing collision energy. However, there seems to be a non-trivial energy dependence in the Λ and Ξ^- dN/dy . The Ξ^- dN/dy first slightly increases with energy from 7.7 to 19.6 GeV, then remains almost

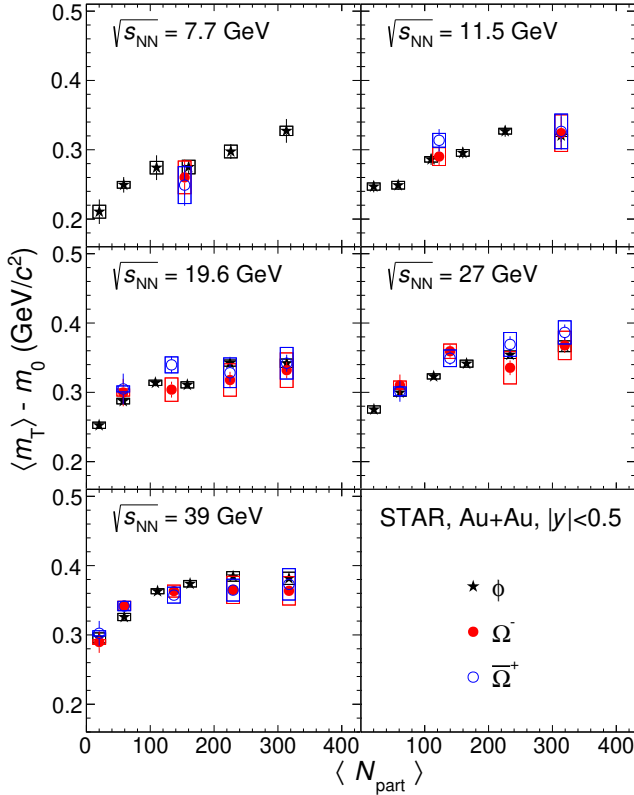


FIG. 15: The averaged transverse mass, $\langle m_T \rangle - m_0$, at mid-rapidity ($|y| < 0.5$) for ϕ , Ω , and $\bar{\Omega}$ as a function of $\langle N_{\text{part}} \rangle$ for Au+Au collisions at $\sqrt{s_{\text{NN}}} = 7.7$ –39 GeV. The box on each data point denotes the systematic error.

constant for energies between 19.6 and 39 GeV, finally rising again toward higher energies. The Λ dN/dy decreases first when energy increases from 7.7 to 39 GeV, then rises up toward higher energies. It should be noted that the proton dN/dy shows a similar minimum at 39 GeV [7]. The proton dN/dy is almost doubled when the collision energy decreases from 39 to 7.7 GeV, reflecting a significant increase in baryon density due to baryon stopping at lower collision energy. Therefore, in terms of hadronic rescatterings, the observed energy dependency of Λ dN/dy in the RHIC BES and higher energies ($\sqrt{s_{\text{NN}}} > 7.7$ GeV) might originate from the interplay between Λ - $\bar{\Lambda}$ pair production, which strongly increases with the increasing collision energy, and the associated production of Λ along with kaons in nucleon-nucleon scatterings [74], which strongly increases with increasing net baryon density and/or decreasing beam energy.

D. Antibaryon-to-baryon ratios

The difference in $\langle m_T \rangle - m_0$ between antibaryons and baryons shown in Fig. 16 might be explained by the absorption of antibaryons due to annihilation at low momentum in a baryon-rich environment. This may result

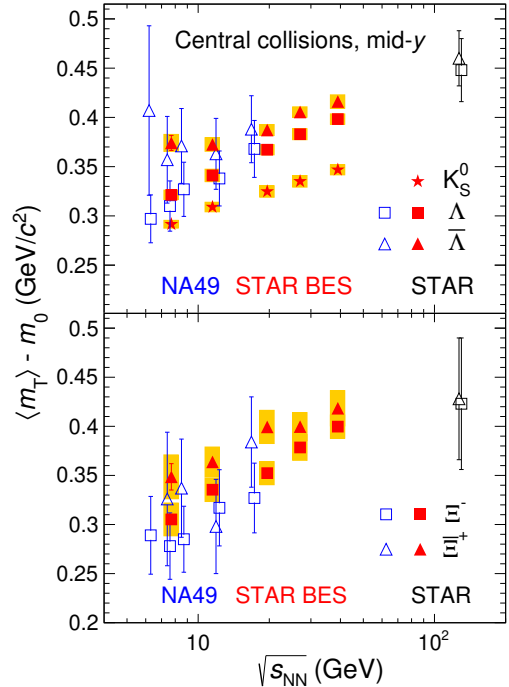


FIG. 16: The averaged transverse mass, $\langle m_T \rangle - m_0$, at mid-rapidity ($|y| < 0.5$) for K_S^0 , Λ , $\bar{\Lambda}$, Ξ^- , and $\bar{\Xi}^+$ as a function of energy from 0–5% central Au+Au collisions at $\sqrt{s_{\text{NN}}} = 7.7$ –39 GeV. For comparison, previous results from central Pb+Pb collisions at $\sqrt{s_{\text{NN}}} = 6.3$ –17.3 GeV at SPS [24] and from central Au+Au collisions at $\sqrt{s_{\text{NN}}} = 130$ GeV at RHIC are shown as open markers [29, 31]. The NA49 and STAR 130 GeV $\bar{\Lambda}$ and $\bar{\Xi}^+$ data points are slightly shifted to the left for clarity. The orange shaded bands on the STAR BES data points represent the systematic errors.

in a decrease of antibaryon yields relative to baryon yields from peripheral to central collisions. Figure 20 shows the antibaryon-to-baryon ratios, $\bar{\Lambda}/\Lambda$, $\bar{\Xi}^+/\Xi^-$, and $\bar{\Omega}^+/\Omega^-$, as functions of $\langle N_{\text{part}} \rangle$ from Au+Au collisions at $\sqrt{s_{\text{NN}}} = 7.7$ –39 GeV. Indeed, the ratios of $\bar{\Lambda}/\Lambda$ and $\bar{\Xi}^+/\Xi^-$ show significant decreases from peripheral to central collisions, especially at lower collision energies. A similar centrality and energy dependence was also observed in the \bar{p}/p ratio [7]. On the other hand, at lower collision energies, strange baryons (not antibaryons) can also be produced in association with kaons through the nucleon-nucleon interactions, which become more important in central collisions due to the increase of binary nucleon-nucleon collisions per participating nucleon pair. This will result in significant baryon stopping at mid-rapidity, but without creating more antibaryons, hence resulting in a decrease of the antibaryon-to-baryon ratio at mid-rapidity with increasing centrality.

To examine more closely how the antibaryon and baryon spectra are different, we plot the $\bar{\Lambda}/\Lambda$ ratio as a function of p_T in Fig. 21 for different centrality bins at $\sqrt{s_{\text{NN}}} = 7.7$ GeV and the normalized $\bar{\Lambda}/\Lambda$ ratio vs p_T

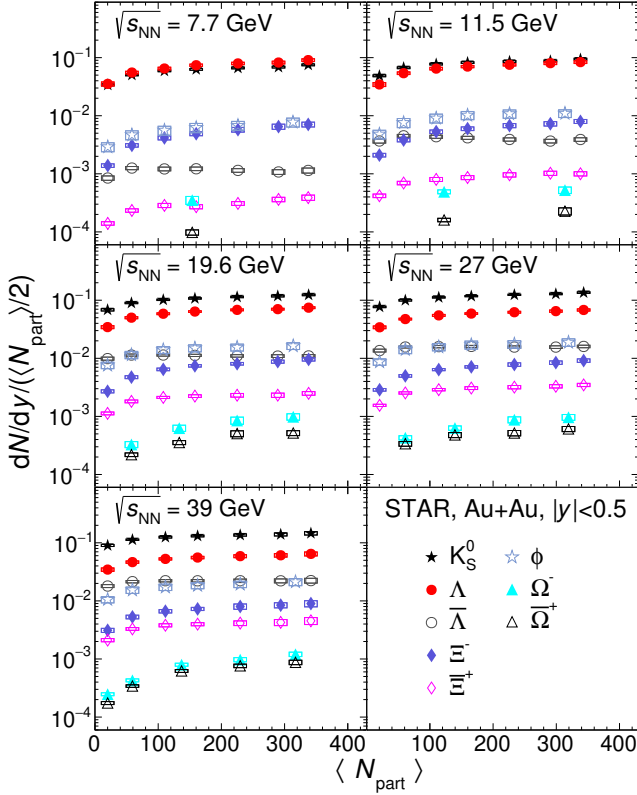


FIG. 17: The integrated yield, dN/dy , per average number of participating nucleon pairs ($\langle N_{\text{part}} \rangle / 2$), of various strange hadrons (K_S^0 , ϕ , Λ , $\bar{\Lambda}$, Ξ^- , $\bar{\Xi}^+$, Ω^- , $\bar{\Omega}^+$) at mid-rapidity ($|y| < 0.5$) as a function of number of participating nucleons, $\langle N_{\text{part}} \rangle$, from Au+Au collisions at $\sqrt{s_{\text{NN}}} = 7.7\text{--}39\text{ GeV}$. The box on each data point denotes the systematic error. For clarity, uncertainties in $\langle N_{\text{part}} \rangle$ are not included.

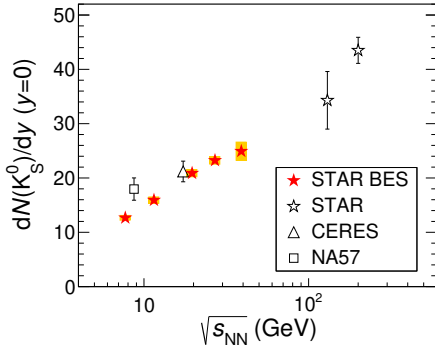


FIG. 18: The K_S^0 integrated yield, dN/dy , at mid-rapidity ($|y| < 0.5$) as a function of collision energy from 0–5% central Au+Au collisions at $\sqrt{s_{\text{NN}}} = 7.7\text{--}39\text{ GeV}$. The orange shaded bands on the STAR BES data points represent the systematic errors. Also shown are the previous mid-rapidity results from 0–5% central Au+Au collisions at $\sqrt{s_{\text{NN}}} = 130$ and 200 GeV ($|y| < 0.5$) from STAR [30, 37], from 0–5% central Pb+Pb collisions at $\sqrt{s_{\text{NN}}} = 8.7\text{ GeV}$ ($|y| < 0.5$) from NA57 [26, 27], and from 0–7% central Pb+Au collisions at $\sqrt{s_{\text{NN}}} = 17.3\text{ GeV}$ from CERES [28]. CERES mid-rapidity data are the extrapolated values based on the measurements at backward rapidity.

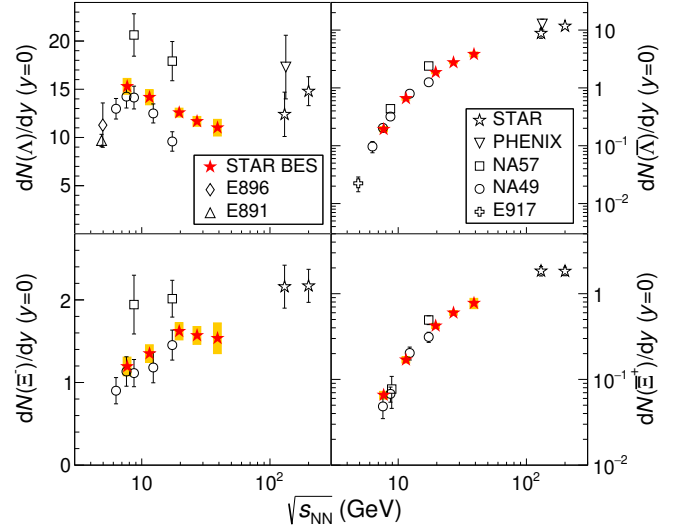


FIG. 19: Collision energy dependence of the Λ , $\bar{\Lambda}$, Ξ^- , and $\bar{\Xi}^+$ integrated yields, dN/dy , at mid-rapidity ($|y| < 0.5$) in 0–5% central Au+Au collisions at $\sqrt{s_{\text{NN}}} = 7.7\text{--}39\text{ GeV}$. Λ ($\bar{\Lambda}$) yields are corrected for weak decay feed-down. The orange shaded bands on the STAR BES data points represent the systematic errors. Also shown are the results from central Au+Au collisions at AGS [14, 15, 17, 18], PHENIX [38] and STAR [29, 31, 32, 37] and central Pb+Pb collisions at NA57 [26, 27] and NA49 [24]. The rapidity ranges are $|y| < 0.5$ for NA57, PHENIX, NA49 Ξ^- ($\bar{\Xi}^+$), and STAR Λ at 130 and 200 GeV, $|y| < 0.75$ for STAR Ξ^- ($\bar{\Xi}^+$), and STAR Λ at 130 and 200 GeV, $|y| < 0.4$ for AGS and NA49 Λ ($\bar{\Lambda}$). The Λ and $\bar{\Lambda}$ results from AGS and PHENIX are inclusive, and those from NA49 and from STAR at higher energies are corrected for weak decay feed-down, while those from NA57 are not significantly affected by weak decay feed-down ($< 5\%$ for Λ and $< 10\%$ for $\bar{\Lambda}$). The E896, PHENIX, and NA57 8.7 GeV $\bar{\Xi}^+$ data points are slightly shifted to the right for clarity.

in Fig. 22 for different energies in central collisions. For $p_T \gtrsim 2\text{ GeV}/c$ at $\sqrt{s_{\text{NN}}} = 39\text{ GeV}$, the ratio decreases with increasing p_T likely due to the semi-hard scattering process dominated by the valence quarks. It is evident that the $\bar{\Lambda}/\Lambda$ ratio at low p_T ($\lesssim 2\text{ GeV}/c$) increases with increasing p_T and energy and decreasing $\langle N_{\text{part}} \rangle$. Hadronic transport model studies could further identify which contributions are dominant: the antibaryon absorption and/or the nucleon-nucleon strangeness association production.

Figure 23 shows the p_T -integrated antibaryon-to-baryon ratios (\bar{B}/B) in central collisions from the STAR Beam Energy Scan in comparison to those from STAR higher energies and NA49. It seems that the STAR BES data are consistent with the NA49 data and fall within the published energy dependence trend. For all energies, the ratios show a hierarchy of $\bar{\Omega}^+/\Omega^- > \bar{\Xi}^+/\Xi^- > \bar{\Lambda}/\Lambda$, which is consistent with the predictions from statistical thermal models [42, 44, 45, 75].

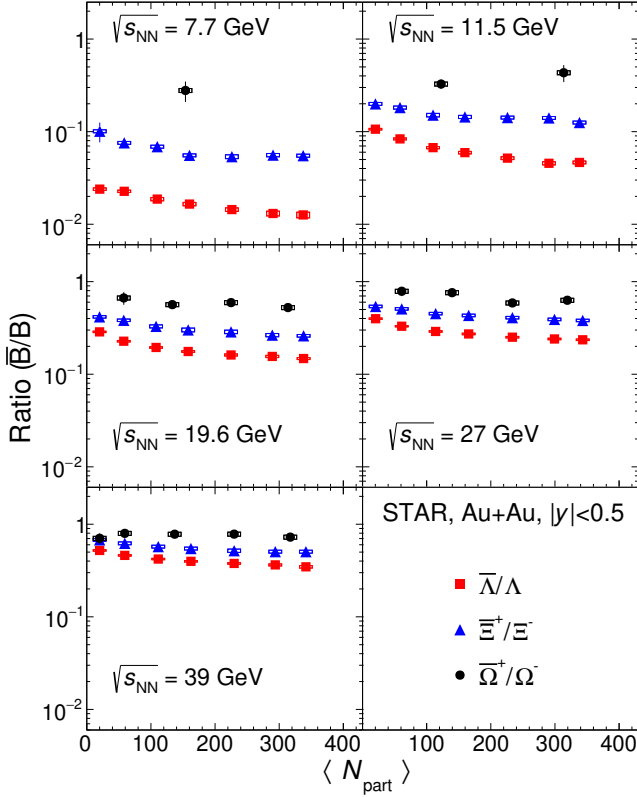


FIG. 20: The antibaryon-to-baryon ratios, $\bar{\Lambda}/\Lambda$, $\bar{\Xi}^+/\Xi^-$, and $\bar{\Omega}^+/\Omega^-$, as functions of $\langle N_{\text{part}} \rangle$ from Au+Au collisions at $\sqrt{s_{\text{NN}}} = 7.7\text{--}39$ GeV. The box on each data point denotes the systematic error.

In heavy-ion collisions, the baryon and antibaryon multiplicities can be described by thermal models [76] with the parameters of particle mass, degeneracy factor, baryon chemical potential (μ_B), strangeness chemical potential (μ_S), charge chemical potential (μ_Q), strangeness saturation factor (γ_s), and chemical freeze-out temperature T_{ch} . By taking the ratio of antibaryon to baryon yield, one obtains

$$\ln(\bar{B}/B) = -2\mu_B/T_{\text{ch}} + \mu_S/T_{\text{ch}} \cdot \Delta S, \quad (8)$$

where ΔS is the difference of strangeness number between antibaryon and baryon. It shows that most parameters can be canceled out in the \bar{B}/B ratios except for μ_S/T_{ch} and μ_B/T_{ch} . These two parameters are properties of the collision system at chemical freeze-out and should be independent of the particle type according to the thermal model, which assumes that all hadrons originate from the same thermal source. With the three measured antibaryon-to-baryon ratios, $\bar{\Lambda}/\Lambda$, $\bar{\Xi}^+/\Xi^-$, and $\bar{\Omega}^+/\Omega^-$, one can test this thermal model assumption by considering that different antibaryon-to-baryon ratios have different strangeness number difference, ΔS . For a certain antibaryon-to-baryon ratio, Eq. 8 is effectively a linear function between μ_B/T_{ch} and μ_S/T_{ch} . With three antibaryon-to-baryon ratios, three straight lines should

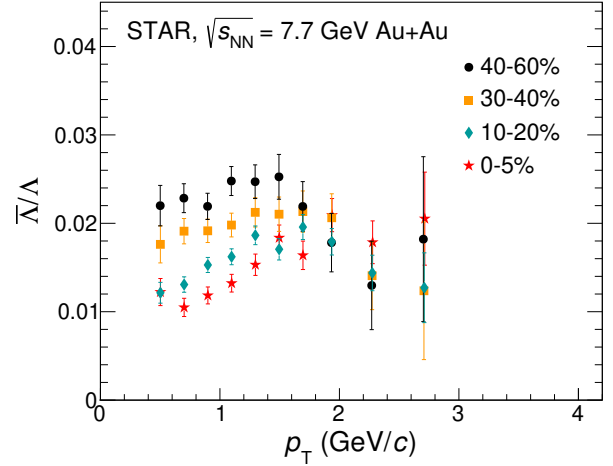


FIG. 21: $\bar{\Lambda}/\Lambda$ ratio as a function of p_T from different centralities of Au+Au collisions at $\sqrt{s_{\text{NN}}} = 7.7$ GeV. The errors are statistical only.

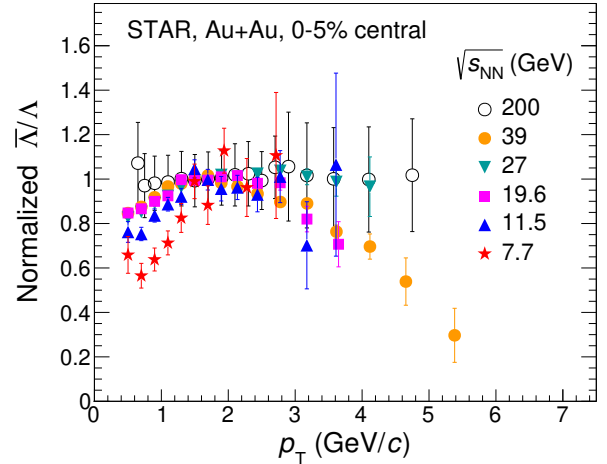


FIG. 22: Normalized $\bar{\Lambda}/\Lambda$ ratio as a function of p_T from 0–5% central Au+Au collisions at different energies. The STAR results at $\sqrt{s_{\text{NN}}} = 200$ GeV [32] are shown as open circles for comparison. The errors are statistical only. All the ratios are normalized according to the average values inside the p_T range of [1.4, 2.0] GeV/c.

cross at the same point on the $(\mu_B/T_{\text{ch}}, \mu_S/T_{\text{ch}})$ plane, which provides a good test for the thermal model assumption. Figure 24 shows the test result for central Au+Au collisions at $\sqrt{s_{\text{NN}}} = 7.7\text{--}39$ GeV, which indicates the validity of this model over the BES energy range. Therefore, the two thermal model parameters, μ_B/T_{ch} and μ_S/T_{ch} , in this collision system, can also be extracted using a linear fit with Eq. 8 to the three measured antibaryon/baryon ratios at each energy, as shown in Fig. 25. The μ_B/T_{ch} and μ_S/T_{ch} parameters in central Au+Au collisions at all five BES energies obtained from the fits are shown in Fig. 26. Also shown

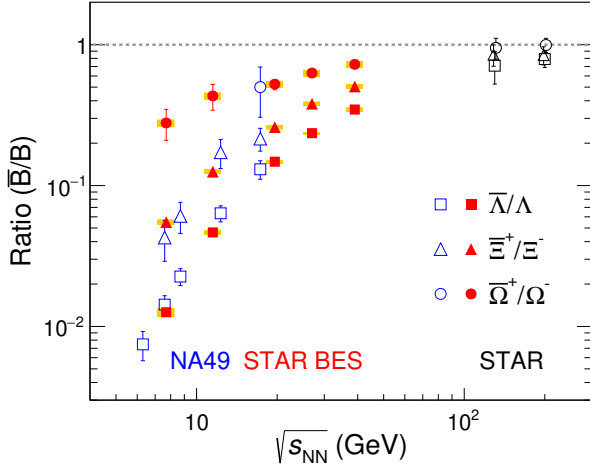


FIG. 23: The collision energy dependence of $\bar{\Lambda}/\Lambda$, $\bar{\Xi}^+/\Xi^-$, and $\bar{\Omega}^+/\Omega^-$ ratios at mid-rapidity ($|y| < 0.5$) in central (0–5% for $\bar{\Lambda}/\Lambda$ and $\bar{\Xi}^+/\Xi^-$; 0–10% for $\bar{\Omega}^+/\Omega^-$ for $\sqrt{s_{NN}} \geq 11.5$ GeV; and 0–60% for $\bar{\Omega}^+/\Omega^-$ for $\sqrt{s_{NN}} = 7.7$ GeV) Au+Au collisions from the STAR Beam Energy Scan (solid symbols). The orange shaded bands represent the systematic errors. The ratios in central Pb+Pb collisions from NA49 [22, 24] and in central Au+Au collisions at higher energies (≥ 130 GeV) from STAR [29, 31, 32, 37] are also shown as open symbols for comparison. The previous STAR $\bar{\Xi}^+/\Xi^-$ and $\bar{\Omega}^+/\Omega^-$ data points are slightly shifted to the left and to the right respectively for clarity.

are the corresponding results from the thermal model (grand-canonical ensemble) fitting to the yields of particles including π , K, p , Λ , Ξ , K_S^0 , and Ω at 39 GeV [7]. We see good agreement between the results from these two methods at this collision energy. Alternatively, the μ_B/T_{ch} and μ_S/T_{ch} parameters can be compared with lattice QCD calculations [77] to further constrain the strangeness chemical freeze-out temperature T_{ch} in these collisions.

E. Baryon-to-meson ratios

Figure 27 shows the ratios of Λ , $\bar{\Lambda}$, Ξ^- , and $\bar{\Xi}^+$ mid-rapidity yields to that of all pions ($1.5(\pi^+ + \pi^-)$) in central Au+Au collisions from the STAR Beam Energy Scan. The existing data from various experiments at different energies are also shown for comparison. The data are compared to the calculations from hadronic transport models (UrQMD 1.3 and HSD [83–86]) and a statistical hadron gas model (SHM [44]). The STAR Beam Energy Scan data are in good agreement with the trend of the existing experimental data. The hadronic models (UrQMD 1.3 and HSD) seem to reproduce the Λ/π data, indicating that the hadronic rescatterings might play an important role in hyperon production in heavy-ion collisions in this energy range. On the other hand,

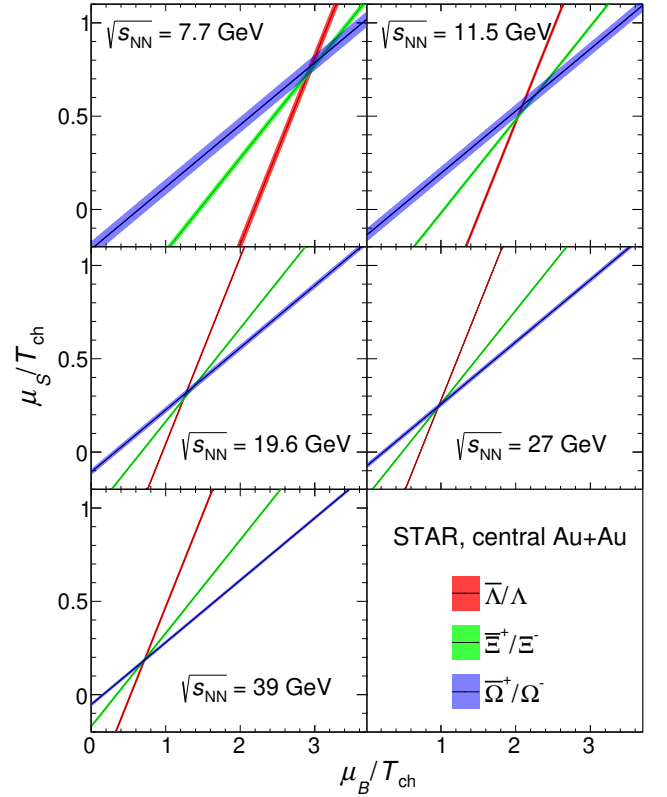


FIG. 24: Testing result of the thermal model in μ_B/T_{ch} and μ_S/T_{ch} parameter space with three strange antibaryon-to-baryon ratios in central Au+Au collisions at $\sqrt{s_{NN}} = 7.7$ –39 GeV. Errors are propagated from the corresponding \bar{B}/B ratios, whose errors are the quadratic sum of statistical and systematic errors.

the SHM model predictions agree well with data across the whole energy range from AGS to top RHIC energies. SHM is based on a grand canonical ensemble and assumes chemical equilibrium. The energy dependence of the parameters T_{ch} and μ_B in the model were obtained with a smooth parametrization of the original fitting parameters to the mid-rapidity particle ratios from heavy ion experiments at SPS and RHIC. The K^+/π^+ [7], Λ/π , and Ξ^-/π ratios all show a maximum at $\sqrt{s_{NN}} \approx 8$ GeV, which seems to be consistent with the picture of maximum net-baryon density at freeze-out at this collision energy [87].

F. Nuclear modification factor

Figure 28 presents the nuclear modification factor, R_{CP} , of K_S^0 , $\Lambda + \bar{\Lambda}$, $\Xi^- + \bar{\Xi}^+$, ϕ and $\Omega^- + \bar{\Omega}^+$ in Au+Au collisions at $\sqrt{s_{NN}} = 7.7$ –39 GeV. R_{CP} is defined as the ratio of particle yield in central collisions to that in peripheral ones scaled by the average number of inelastic

TABLE VII: The average number of binary nucleon-nucleon collisions ($\langle N_{\text{coll}} \rangle$) for various collision centralities in Au+Au collisions at 7.7–39 GeV, determined using the charged particle multiplicity distributions and the Glauber Monte Carlo simulation [10]. The errors represent systematic uncertainties. The inelastic $p+p$ cross-sections used in the simulations are 30.8, 31.2, 32, 33, and 34 mb for $\sqrt{s} = 7.7, 11.5, 19.6, 27,$ and 39 GeV, respectively [64].

$\sqrt{s_{\text{NN}}} \text{ (GeV)}$	0–5%	5–10%	10–20%	20–30%	30–40%	40–60%	60–80%
7.7	774 ± 28	629 ± 20	450 ± 22	283 ± 24	171 ± 23	74 ± 16	19.2 ± 6.3
11.5	784 ± 27	635 ± 20	453 ± 23	284 ± 23	172 ± 22	75 ± 16	19.1 ± 7.8
19.6	800 ± 27	643 ± 20	458 ± 24	285 ± 26	170 ± 23	74 ± 15	18.9 ± 6.9
27	841 ± 28	694 ± 22	497 ± 26	312 ± 28	188 ± 25	82 ± 18	20.0 ± 8.6
39	853 ± 27	687 ± 21	492 ± 26	306 ± 27	183 ± 24	79 ± 17	19.4 ± 7.7

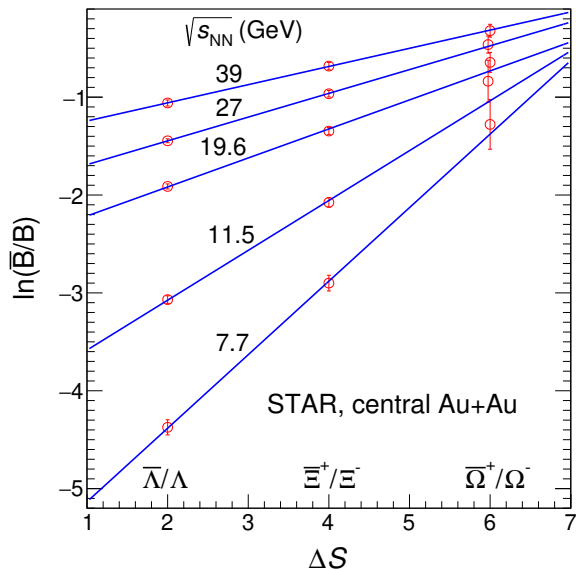


FIG. 25: Thermal model fitting to $\ln(\bar{B}/B)$ vs ΔS with a linear function to determine μ_B/T_{ch} and μ_S/T_{ch} for central Au+Au collisions at $\sqrt{s_{\text{NN}}} = 7.7\text{--}39$ GeV. The $\bar{\Omega}^+/\Omega^-$ data points at 11.5 and 27 GeV are slightly shifted to the left for clarity.

binary collisions N_{coll} , i.e.,

$$R_{\text{CP}} = \frac{[(dN/dp_T)/\langle N_{\text{coll}} \rangle]_{\text{central}}}{[(dN/dp_T)/\langle N_{\text{coll}} \rangle]_{\text{peripheral}}}. \quad (9)$$

Here N_{coll} is determined from Glauber Monte Carlo simulations. See Table VII for the N_{coll} values for Au+Au collisions in the STAR Beam Energy Scan. R_{CP} will be unity if nucleus-nucleus collisions are just simple superpositions of nucleon-nucleon collisions. Deviation of these ratios from unity would imply contributions from nuclear or in-medium effects. For $p_T \approx 4$ GeV/c, one can see from Fig. 28 that the K_S^0 R_{CP} is below unity at $\sqrt{s_{\text{NN}}} = 39$ GeV. This is similar to the observation at top RHIC energy [88] though the lowest R_{CP} value is larger here. Then the K_S^0 R_{CP} at $p_T > 2$ GeV/c keeps increasing with decreasing collision energies, indicating that the partonic energy loss effect becomes less important. Eventually, the cold nuclear matter effect (Cronin effect) [89]

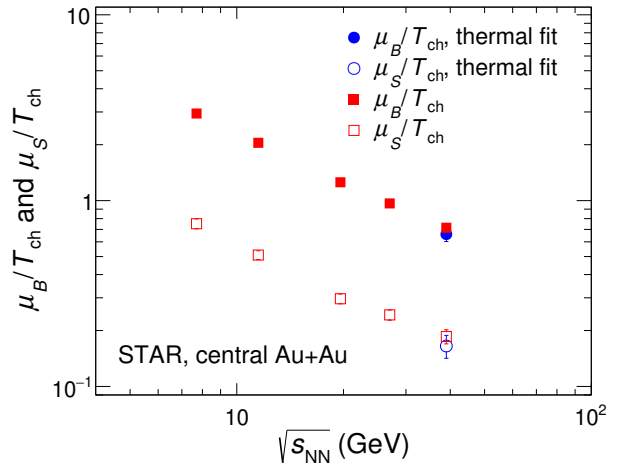


FIG. 26: The μ_B/T_{ch} and μ_S/T_{ch} parameters (red symbols) in central Au+Au collisions at $\sqrt{s_{\text{NN}}} = 7.7\text{--}39$ GeV obtained from the fits shown in Fig. 25. The blue symbols are the corresponding results from the thermal model fit to the yields of π , K, p , Λ , Ξ , K_S^0 , and Ω at $\sqrt{s_{\text{NN}}} = 39$ GeV [7].

starts to take over at $\sqrt{s_{\text{NN}}} = 11.5$ and 7.7 GeV and enhances all the hadron (including K_S^0) yields at intermediate p_T (up to ≈ 3.5 GeV/c). Similar to the observation for identified charged hadrons [12], the energy evolution of strange hadron R_{CP} reflects the decreasing partonic effects with decreasing beam energies. In addition, the particle R_{CP} differences are apparent for $\sqrt{s_{\text{NN}}} \geq 19.6$ GeV. However, the differences become smaller at $\sqrt{s_{\text{NN}}} = 11.5$ GeV and eventually vanish at $\sqrt{s_{\text{NN}}} = 7.7$ GeV, which may also suggest different properties of the system created in Au+Au collisions at $\sqrt{s_{\text{NN}}} = 11.5$ and 7.7 GeV, compared to those in $\sqrt{s_{\text{NN}}} \geq 19.6$ GeV.

G. Baryon enhancement at intermediate p_T

The enhancement of baryon-to-meson ratios at intermediate p_T in central $A+A$ collisions compared to peripheral $A+A$ or $p+p$ collisions at the same energy is interpreted as a consequence of hadron formation through parton recombination and parton collectivity in central

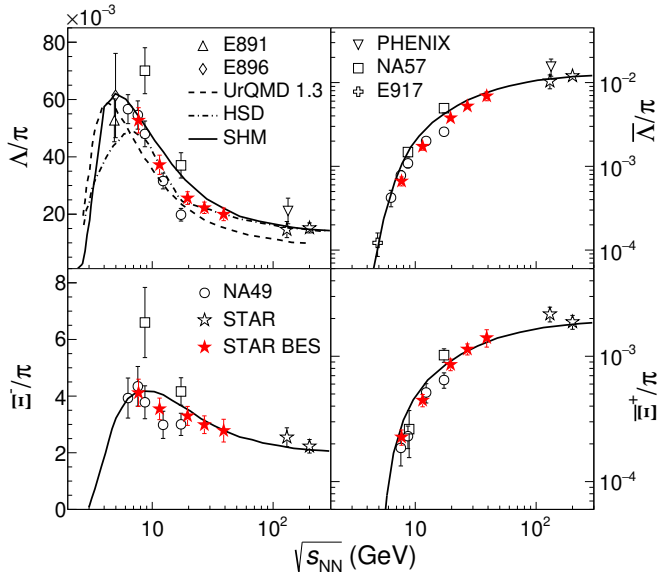


FIG. 27: Energy dependence of Λ , $\bar{\Lambda}$, Ξ^- , and Ξ^+ to pions ($1.5(\pi^+ + \pi^-)$) ratios at mid-rapidity in central Au+Au collisions from STAR Beam Energy Scan (solid symbols). The STAR BES mid-rapidity pion yields are taken from [7]. Errors are the quadratic sum of statistical and systematic errors. Also shown are existing AGS [14, 15, 17, 18, 78], NA49 [20, 23, 24, 79], PHENIX [38, 80] and STAR [29, 31, 32, 37, 81, 82] data as open symbols, as well as calculations from hadronic transport models (UrQMD 1.3 and HSD) [83–86] and a statistical hadron gas model (SHM) [44] as dashed or solid lines. The E896, PHENIX, and NA57 8.7 GeV Ξ^{\pm}/π data points are slightly shifted to the right for clarity.

collisions [51–58, 90]. Therefore, the baryon-to-meson ratios are expected to be sensitive to the parton dynamics of the collision system. The multi-strange baryon-to-meson ratio, Ω/ϕ , has been described in detail in Ref. [11]. Figure 29 shows the $\bar{\Lambda}/K_S^0$ ratio as a function of p_T in different centralities from Au+Au collisions at $\sqrt{s_{NN}} = 7.7\text{--}39$ GeV. The $\bar{\Lambda}$ is chosen instead of Λ , because it is a newly produced baryon in the baryon-rich medium created in lower Beam Energy Scan energies. At $\sqrt{s_{NN}} \geq 19.6$ GeV, the $\bar{\Lambda}/K_S^0$ reaches its maximum value at $p_T \approx 2.5$ GeV/c in central collisions, while in peripheral collisions, the maximum value is significantly lower. This shows that there is baryon enhancement at intermediate p_T for $\sqrt{s_{NN}} \geq 19.6$ GeV similar to that observed at higher energies. For Au+Au collisions at $\sqrt{s_{NN}} \leq 11.5$ GeV, the difference between the values of $\bar{\Lambda}/K_S^0$ in the measured p_T range in 0–5% and 40–60% is much less significant. Unfortunately, the maximum $\bar{\Lambda}/K_S^0$ value in each centrality bin cannot be clearly identified due to limited p_T reach and statistics at $\sqrt{s_{NN}} \leq 11.5$ GeV, hence whether baryon-to-meson enhancement still persists at these energies remains uncertain with the current data.

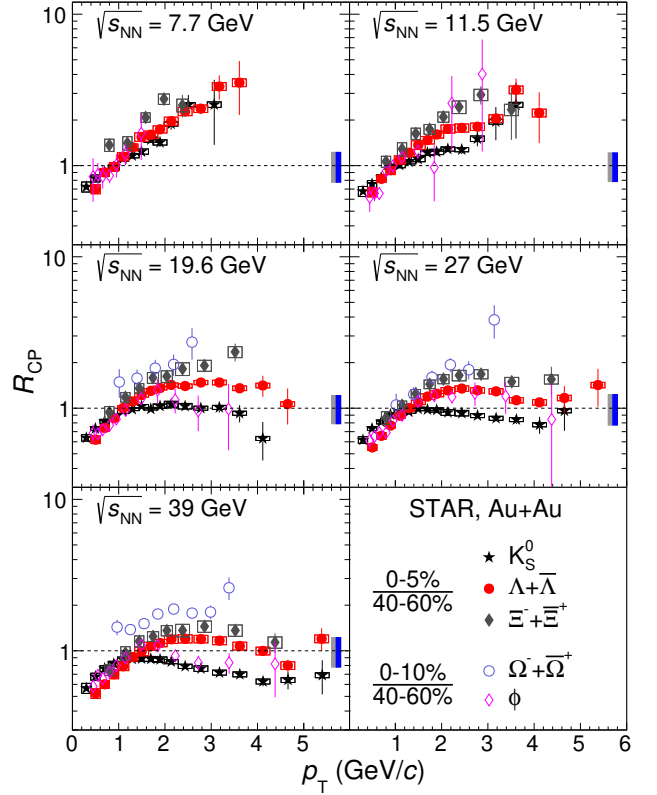


FIG. 28: K_S^0 , $\Lambda + \bar{\Lambda}$, and $\Xi^- + \Xi^+$ $R_{CP}(0\text{--}5\%)/(40\text{--}60\%)$, ϕ and $\Omega^- + \bar{\Omega}^+$ $R_{CP}(0\text{--}10\%)/(40\text{--}60\%)$, at mid-rapidity ($|y| < 0.5$) in Au+Au collisions at $\sqrt{s_{NN}} = 7.7\text{--}39$ GeV. The vertical bars denote the statistical errors. The box on each data point of K_S^0 , Λ , and Ξ denotes the systematic error. There are only statistical errors for Ω and ϕ . For $\sqrt{s_{NN}} \leq 19.6$ GeV, the $\Lambda + \bar{\Lambda}$ R_{CP} excludes the minor contribution from $\bar{\Lambda}$. The gray and blue bands on the right side of each panel represent the normalization errors from N_{coll} for $R_{CP}(0\text{--}5\%)/(40\text{--}60\%)$ and $R_{CP}(0\text{--}10\%)/(40\text{--}60\%)$ respectively.

V. CONCLUSIONS

In this paper we present STAR measurements of strange hadrons (K_S^0 , Λ , $\bar{\Lambda}$, Ξ^- , Ξ^+ , Ω^- , $\bar{\Omega}^+$, and ϕ) production at mid-rapidity ($|y| < 0.5$) in Au+Au collisions at $\sqrt{s_{NN}} = 7.7, 11.5, 19.6, 27,$ and 39 GeV, from data taken in the first phase of the RHIC Beam Energy Scan Program.

The p_T spectra, averaged transverse mass ($\langle m_T \rangle - m_0$), and integrated yield (dN/dy) have been extracted with high precision for all strange hadron species and for all centralities and collision energies. Generally, the STAR BES data follow the trend of the previous measurements from AGS, SPS and RHIC. These measurements also exhibit the following features in strange hadron production in this energy range. **(1)** $\langle m_T \rangle - m_0$ of antibaryons and baryons significantly deviate from each other toward lower collision energies, especially for $\bar{\Lambda}$ and Λ . **(2)** dN/dy of strange baryons (Λ and Ξ^-) show a

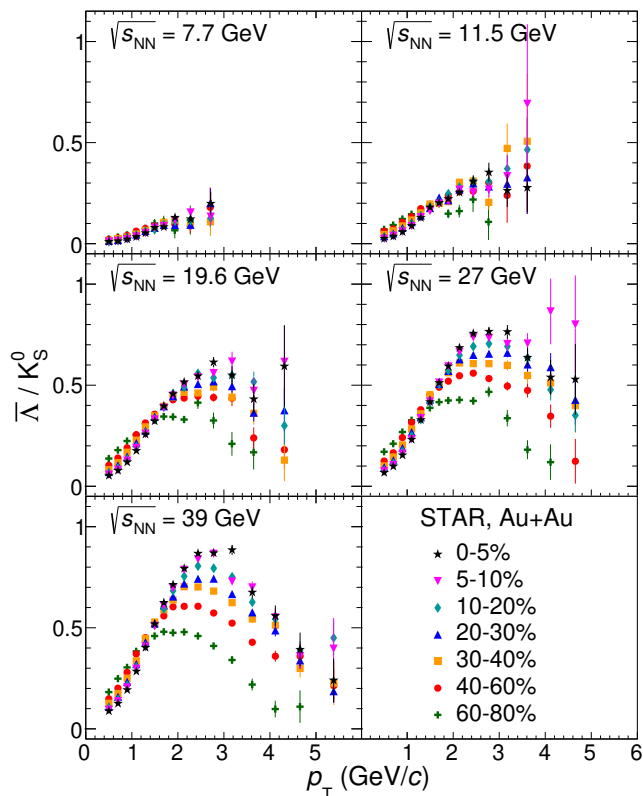


FIG. 29: $\bar{\Lambda}/K_S^0$ ratio as a function of p_T at mid-rapidity ($|y| < 0.5$) in different centralities from Au+Au collisions at $\sqrt{s_{NN}} = 7.7\text{--}39$ GeV. Errors are statistical only.

non-monotonic energy dependence, while the dN/dy of the corresponding antibaryons and K_S^0 and ϕ mesons increases monotonically toward higher energies.

All the antibaryon-to-baryon ratios decrease toward central collisions due to increased baryon stopping over the BES energy range. The $\bar{\Lambda}/\Lambda$ ratio tends to increase significantly with increasing p_T in the low p_T region ($\lesssim 2$ GeV/c) in central collisions, especially for $\sqrt{s_{NN}} \leq 11.5$ GeV. This can be due to either antibaryon absorption or the increase of stopped baryons without pair production. The thermal model has been tested with the measured antibaryon-to-baryon ratios, $\bar{\Lambda}/\Lambda$, $\bar{\Xi}^+/\Xi^-$, and $\bar{\Omega}^+/\Omega^-$, and then two chemical freeze-out parameters, μ_B/T_{ch} and μ_S/T_{ch} , are extracted for central Au+Au collisions at five BES energies. The strange baryon-to-pion ratios are found to be consistent with the calculations of the statistical hadron gas model, and for Λ/π ratio, consis-

tent with hadronic transport models as well.

For intermediate-to-high p_T probes, the nuclear modification factors (R_{CP}) of various strange hadrons and the $\bar{\Lambda}/K_S^0$ ratio have been presented. The $K_S^0 R_{CP}$ shows no suppression for p_T up to 3.5 GeV/c at energies of 7.7 and 11.5 GeV. The particle-type dependence of R_{CP} also becomes smaller at $\sqrt{s_{NN}} \leq 11.5$ GeV. These observations show that the partonic energy loss effect becomes less significant with decreasing collision energy. The cold nuclear matter effect, such as the Cronin effect, starts to take over at $\sqrt{s_{NN}} = 11.5$ and 7.7 GeV and enhances the hadron yields at intermediate p_T . The $\bar{\Lambda}/K_S^0$ ratio data show baryon-to-meson enhancement at intermediate p_T (≈ 2.5 GeV/c) in central collisions at energies above 19.6 GeV. Unfortunately, the precision of current $\bar{\Lambda}/K_S^0$ measurements below 11.5 GeV does not allow us to unambiguously conclude regarding possible baryon-to-meson enhancement at these energies. These measurements point to the beam energy region below 19.6 GeV for further investigation of the deconfinement phase transition.

Acknowledgments

We thank the RHIC Operations Group and RCF at BNL, the NERSC Center at LBNL, and the Open Science Grid consortium for providing resources and support. This work was supported in part by the Office of Nuclear Physics within the U.S. DOE Office of Science, the U.S. National Science Foundation, the Ministry of Education and Science of the Russian Federation, National Natural Science Foundation of China (especially under Contract No. 11890710), Chinese Academy of Science, the Ministry of Science and Technology of China and the Chinese Ministry of Education, the National Research Foundation of Korea, Czech Science Foundation and Ministry of Education, Youth and Sports of the Czech Republic, Hungarian National Research, Development and Innovation Office (FK-123824), New National Excellency Programme of the Hungarian Ministry of Human Capacities (UNKP-18-4), Department of Atomic Energy and Department of Science and Technology of the Government of India, the National Science Centre of Poland, the Ministry of Science, Education and Sports of the Republic of Croatia, RosAtom of Russia, German Bundesministerium für Bildung, Wissenschaft, Forschung und Technologie (BMBF), and the Helmholtz Association.

- [1] M. A. Stephanov, Prog. Theor. Phys. Suppl. **153**, 139 (2004).
 [2] B. Mohanty, Nucl. Phys. A **830**, 899C (2009).
 [3] M. M. Aggarwal *et al.* (STAR Collaboration), arXiv:1007.2613.
 [4] L. Adamczyk *et al.* (STAR Collaboration), Phys. Rev.

- Lett. **112**, 032302 (2014).
 [5] L. Adamczyk *et al.* (STAR Collaboration), Phys. Rev. Lett. **112**, 162301 (2014).
 [6] L. Adamczyk *et al.* (STAR Collaboration), Phys. Rev. Lett. **113**, 092301 (2014).
 [7] L. Adamczyk *et al.* (STAR Collaboration), Phys. Rev. C

- 96**, 044904 (2017).
- [8] L. Adamczyk *et al.* (STAR Collaboration), Phys. Lett. B **785**, 551 (2018).
- [9] L. Adamczyk *et al.* (STAR Collaboration), Phys. Rev. Lett. **110**, 142301 (2013).
- [10] L. Adamczyk *et al.* (STAR Collaboration), Phys. Rev. C **88**, 014902 (2013).
- [11] L. Adamczyk *et al.* (STAR Collaboration), Phys. Rev. C **93**, 021903(R) (2016).
- [12] L. Adamczyk *et al.* (STAR Collaboration), Phys. Rev. Lett. **121**, 032301 (2018).
- [13] J. Rafelski, B. Müller, Phys. Rev. Lett. **48**, 1066 (1982).
- [14] S. Ahmad *et al.* (E891 Collaboration), Phys. Lett. B **382**, 35 (1996), Erratum: Phys. Lett. B **386**, 496 (1996).
- [15] S. Ahmad *et al.*, Nucl. Phys. A **636**, 507 (1998).
- [16] L. Ahle *et al.* (E802 Collaboration), Phys. Rev. C **60**, 044904 (1999).
- [17] B. B. Back *et al.* (E917 Collaboration), Phys. Rev. Lett. **87**, 242301 (2001).
- [18] S. Albergo *et al.* (E896 Collaboration), Phys. Rev. Lett. **88**, 062301 (2002).
- [19] P. Chung *et al.* (E895 Collaboration), Phys. Rev. Lett. **91**, 202301 (2003).
- [20] S. V. Afanasiev *et al.* (NA49 Collaboration), Phys. Rev. C **66**, 054902 (2002).
- [21] T. Anticic *et al.* (NA49 Collaboration), Phys. Rev. Lett. **93**, 022302 (2004).
- [22] C. Alt *et al.* (NA49 Collaboration), Phys. Rev. Lett. **94**, 192301 (2005).
- [23] C. Alt *et al.* (NA49 Collaboration), Phys. Rev. C **77**, 024903 (2008).
- [24] C. Alt *et al.* (NA49 Collaboration), Phys. Rev. C **78**, 034918 (2008).
- [25] T. Anticic *et al.* (NA49 Collaboration), Phys. Rev. C **80**, 034906 (2009).
- [26] F. Antinori *et al.* (NA57 Collaboration), Phys. Lett. B **595**, 68 (2004).
- [27] F. Antinori *et al.* (NA57 Collaboration), J. Phys. G **32**, 427 (2006); <http://wa97.web.cern.ch/WA97/yields.html>.
- [28] J. Milosevic (CERES Collaboration), J. Phys. G **32**, S97 (2006).
- [29] C. Adler *et al.* (STAR Collaboration), Phys. Rev. Lett. **89**, 092301 (2002).
- [30] C. Adler *et al.* (STAR Collaboration), Phys. Lett. B **595**, 143 (2004).
- [31] J. Adams *et al.* (STAR Collaboration), Phys. Rev. Lett. **92**, 182301 (2004).
- [32] J. Adams *et al.* (STAR Collaboration), Phys. Rev. Lett. **98**, 062301 (2007).
- [33] B. I. Abelev *et al.* (STAR Collaboration), Phys. Rev. C **77**, 044908 (2008).
- [34] B. I. Abelev *et al.* (STAR Collaboration), Phys. Lett. B **673**, 183 (2009).
- [35] B. I. Abelev *et al.* (STAR Collaboration), Science **328**, 58 (2010).
- [36] M. M. Aggarwal *et al.* (STAR Collaboration), Phys. Rev. C **83**, 024901 (2011).
- [37] G. Agakishiev *et al.* (STAR Collaboration), Phys. Rev. Lett. **108**, 072301 (2012).
- [38] K. Adcox *et al.* (PHENIX Collaboration), Phys. Rev. Lett. **89**, 092302 (2002).
- [39] B. Abelev *et al.* (ALICE Collaboration), Phys. Rev. Lett. **111**, 222301 (2013).
- [40] B. Abelev *et al.* (ALICE Collaboration), Phys. Lett. B **728**, 216 (2014), Erratum: Phys. Lett. B **734**, 409 (2014).
- [41] J. Chen, D. Keane, Y. G. Ma, A. Tang and Z. Xu, Phys. Rept. **760**, 1 (2018).
- [42] F. Becattini, J. Cleymans, A. Keranen, E. Suhonen and K. Redlich, Phys. Rev. C **64**, 024901 (2001).
- [43] P. Braun-Munzinger, J. Cleymans, H. Oeschler and K. Redlich, Nucl. Phys. A **697**, 902 (2002).
- [44] A. Andronic, P. Braun-Munzinger and J. Stachel, Nucl. Phys. A **772**, 167 (2006).
- [45] K. Redlich and A. Tounsi, Eur. Phys. J. C **24**, 589 (2002).
- [46] J. Adams *et al.* (STAR Collaboration), Phys. Rev. Lett. **91**, 172302 (2003).
- [47] K. Adcox *et al.* (PHENIX Collaboration), Phys. Rev. C **69**, 024904 (2004).
- [48] B. I. Abelev *et al.* (STAR Collaboration), Phys. Lett. B **655**, 104 (2007).
- [49] B. I. Abelev *et al.* (STAR Collaboration), Phys. Rev. Lett. **97**, 152301 (2006).
- [50] B. B. Abelev *et al.* (ALICE Collaboration), Phys. Lett. B **736**, 196 (2014).
- [51] R. C. Hwa and C. B. Yang, Phys. Rev. C **65**, 034905 (2002), Erratum: Phys. Rev. C **67**, 059902 (2003).
- [52] R. C. Hwa and C. B. Yang, Phys. Rev. C **66**, 025205 (2002).
- [53] D. Molnar and S. A. Voloshin, Phys. Rev. Lett. **91**, 092301 (2003).
- [54] R. C. Hwa and C. B. Yang, Phys. Rev. C **75**, 054904 (2007).
- [55] R. J. Fries, B. Muller, C. Nonaka and S. A. Bass, Phys. Rev. Lett. **90**, 202303 (2003).
- [56] R. J. Fries, B. Muller, C. Nonaka and S. A. Bass, Phys. Rev. C **68**, 044902 (2003).
- [57] V. Greco, C. M. Ko and P. Levai, Phys. Rev. Lett. **90**, 202302 (2003).
- [58] V. Greco, C. M. Ko and P. Levai, Phys. Rev. C **68**, 034904 (2003).
- [59] K. H. Ackermann *et al.* (STAR Collaboration), Nucl. Instrum. Meth. A **499**, 624 (2003).
- [60] M. Anderson *et al.*, Nucl. Instrum. Meth. A **499**, 659 (2003).
- [61] W. J. Llope *et al.*, Nucl. Instrum. Meth. A **522**, 252 (2004).
- [62] F. S. Bieser *et al.*, Nucl. Instrum. Meth. A **499**, 766 (2003).
- [63] B. Bonner *et al.*, Nucl. Instrum. Meth. A **508**, 181 (2003).
- [64] M. L. Miller, K. Reygers, S. J. Sanders and P. Steinberg, Ann. Rev. Nucl. Part. Sci. **57**, 205 (2007).
- [65] C. Patrignani *et al.* (Particle Data Group), Chin. Phys. C **40**, 100001 (2016).
- [66] M. Shao, O. Y. Barannikova, X. Dong, Y. Fisyak, L. Ruan, P. Sorensen and Z. Xu, Nucl. Instrum. Meth. A **558**, 419 (2006).
- [67] H. Bichsel, Nucl. Instrum. Meth. A **562**, 154 (2006).
- [68] J. Adams *et al.* (STAR Collaboration), Phys. Lett. B **612**, 181 (2005).
- [69] B. I. Abelev *et al.* (STAR Collaboration), Phys. Rev. Lett. **99**, 112301 (2007).
- [70] B. I. Abelev *et al.* (STAR Collaboration), Phys. Rev. C **79**, 064903 (2009).
- [71] H. Long, PhD thesis, UCLA, (2002).
- [72] V. Fine and P. Nevski, Proc. CHEP **2000**, 143 (2000).
- [73] E. Schnedermann, J. Sollfrank and U. W. Heinz, Phys. Rev. C **48**, 2462 (1993).

- [74] J. Cleymans, A. Forster, H. Oeschler, K. Redlich and F. Uhlig, Phys. Lett. B **603**, 146 (2004).
- [75] F. Becattini, J. Manninen and M. Gazdzicki, Phys. Rev. C **73**, 044905 (2006).
- [76] J. Cleymans, arXiv:nucl-th/9704046.
- [77] A. Bazavov *et al.*, Phys. Rev. Lett. **113**, 072001 (2014).
- [78] L. Ahle *et al.* (E802 Collaboration), Phys. Rev. C **57**, R466 (1998).
- [79] C. Alt *et al.* (NA49 Collaboration), Phys. Rev. C **73**, 044910 (2006).
- [80] K. Adcox *et al.* (PHENIX Collaboration), Phys. Rev. Lett. **88**, 242301 (2002).
- [81] J. Adams *et al.* (STAR Collaboration), Phys. Rev. Lett. **92**, 112301 (2004).
- [82] B. I. Abelev *et al.* (STAR Collaboration), Phys. Rev. C **79**, 034909 (2009).
- [83] S. A. Bass *et al.*, Prog. Part. Nucl. Phys. **41**, 255 (1998).
- [84] M. Bleicher *et al.*, J. Phys. G **25**, 1859 (1999).
- [85] W. Cassing and E. L. Bratkovskaya, Phys. Rept. **308**, 65 (1999).
- [86] E. L. Bratkovskaya, M. Bleicher, M. Reiter, S. Soff, H. Stoecker, M. van Leeuwen, S. A. Bass and W. Cassing, Phys. Rev. C **69**, 054907 (2004).
- [87] J. Randrup and J. Cleymans, Phys. Rev. C **74**, 047901 (2006).
- [88] J. Adams *et al.* (STAR Collaboration), Phys. Rev. Lett. **92**, 052302 (2004).
- [89] J. W. Cronin, H. J. Frisch, M. J. Shochet, J. P. Boymond, P. A. Piroue and R. L. Sumner, Phys. Rev. D **11**, 3105 (1975).
- [90] J. H. Chen, F. Jin, D. Gangadharan, X. Z. Cai, H. Z. Huang and Y. G. Ma, Phys. Rev. C **78**, 034907 (2008).