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## The $ep \rightarrow e'\rho\eta$ Reaction at and above the $S_{11}(1535)$ Baryon Resonance

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New cross sections for the reaction  $ep \rightarrow ep\eta$  are reported for total center of mass energy  $W = 1.5\text{--}1.86$  GeV and invariant momentum transfer  $Q^2 = 0.25\text{--}1.5$  (GeV/c)<sup>2</sup>. This large kinematic range allows extraction of important new information about response functions, photocouplings, and  $\eta N$  coupling strengths of baryon resonances. Newly observed structure at  $W \sim 1.65$  GeV is shown to come from interference between  $S$  and  $P$  waves and can be interpreted with known resonances. Improved values are derived for the photon coupling amplitude for the  $S_{11}(1535)$  resonance.

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The study of baryon resonances is undergoing a significant rebirth because of new experimental programs at Brookhaven, the Mainz microtron, the Bonn synchrotron, and Jefferson Lab. Continuous, polarized beams and large acceptance detectors are significantly improving experimental accuracy. Older studies found a few dozen states with a variety of total angular momentum, parity, and strangeness [1]. Modern theoretical work examines the microscopic structure in terms of quark and gluon interactions. Empirical constituent quark models (CQM) achieve excellent qualitative agreement with a variety of data [2] and provide important evidence that quark excitations in these states are more important than gluonic excitations. Lattice gauge models simulate full QCD; they presently calculate moderately accurate values for excited state masses [3] and show great promise. These studies require data of much higher quality than was previously available.

Disentangling the wide and overlapping states that populate reaction data is a historical problem. However, reactions involving  $\eta N$  final states couple only to isospin  $\frac{1}{2}$  resonances. Although  $\pi N$  elastic scattering close to  $\eta N$  threshold (total c.m. energy,  $W = 1.485$  GeV) shows no strong signal of a resonance, a prominent peak in the total cross section is seen for  $\eta$  production in  $\gamma N$  and  $\pi N$  experiments. This is widely interpreted as the excitation of a *single* resonance, the spin  $\frac{1}{2}$ , negative parity, isospin  $\frac{1}{2}$  state  $S_{11}(1535)$  [1]. ( $S$  labels the  $\eta N$  orbital angular momentum.) This state has a branching ratio to  $\eta N$  of 30–55% compared to a few percent [1,4] for other states. These unusual features have encouraged alternative theoretical efforts to describe the data in terms of a strong (possibly nonresonant) final state interaction [5].

Most previous experiments used pion beams. An important advantage of electromagnetic experiments is the ability to extract the matrix elements for  $\gamma N \rightarrow N^*$ , commonly called the photon coupling amplitudes. These amplitudes are primarily sensitive to the quark wave function used. They are labeled by the  $\gamma N$  total helicity and the virtual photon polarization and depend on the invariant

momentum transfer to the resonance ( $Q^2$ ). For a spin  $\frac{1}{2}$  resonance, there is one transverse amplitude ( $A_{1/2}$ ) and one longitudinal amplitude ( $C_{1/2}$ ).

Photoproduction experiments ( $Q^2 = 0$ ) have reaffirmed the strong energy dependence and  $S$ -wave (isotropic) character close to threshold [6]. A recent experiment with polarized photons [7] has given new values for  $\eta N$  decay branching ratios of other resonances through interference with the dominant  $S_{11}(1535)$ .

In electroproduction experiments,  $Q^2$  is nonzero and provides additional structure information about the intermediate state. Past  $\eta$  electroproduction experiments [8–11] found an unusually flat  $Q^2$  dependence of  $A_{1/2}$  for the  $S_{11}(1535)$  in contrast to the nucleon form factors and photon coupling amplitudes of other established resonances, e.g.,  $P_{33}(1232)$ . At this time, there is no definitive explanation for this difference. Although previous angular distributions were largely isotropic at all  $Q^2$ , no detailed response functions were extracted because of the poor angular coverage in traditional magnetic spectrometers. Here,  $\eta$  electroproduction is used to study the  $S_{11}(1535)$  over a broad range of  $Q^2$  and  $W$ . At higher  $W$  new interference effects are found that add to our knowledge of  $\eta N$  coupling to higher mass resonances.

The results reported here used the CEBAF Large Acceptance Spectrometer (CLAS) at Jefferson Lab (JLab). It has moderate momentum resolution and excellent solid angle coverage for final state particles produced in collisions of photon or electron beams of up to 5.5 GeV energy with various targets. This is advantageous for  $N^*$  experiments because the resonances decay to multiple particles spread over a large kinematic range.

The CLAS detector [12] measures angles and momenta of charged particles for lab polar angles ( $\theta$ ) in the range of 8°–142°. For this measurement, electron beams with energies of 1.645 GeV [ $0.25 < Q^2 < 0.5$  (GeV/c)<sup>2</sup>] and 2.445 GeV [ $0.5 < Q^2 < 1.5$  (GeV/c)<sup>2</sup>] were incident on a liquid hydrogen target. A full description of these results can be found in Ref. [13]. An electron and proton were identified in the final state. The hardware trigger identified

electrons through threshold Cerenkov detectors and an electromagnetic calorimeter. The proton was identified using the time-of-flight technique. Fiducial cuts were used to restrict particles to detector locations where the single particle detection efficiency is flat. Events for the angular distributions were binned in  $Q^2$ ,  $W$ , and the c.m. decay angles [14]  $\cos\theta_\eta^*$  and  $\phi_\eta^*$ . For the angle integrated cross sections, the same events were binned in  $Q^2$  and  $W$ .

$\eta$  mesons were identified by fitting the missing mass spectrum (see Fig. 1). The fit function is the sum of a peak with a radiative tail and a background function. The background is due to multipion production reactions. Lacking a detailed understanding of the background, a simple function incorporating the proper behavior at the kinematic limit and the CLAS acceptance was used. Acceptance was calculated using a GEANT-based Monte Carlo simulation of the CLAS detector that included bremsstrahlung radiation using the peaking approximation. The maximum acceptance of these data is 54% and no cross section is reported where the acceptance was less than 5%.

A detailed study of potential sources of systematic error was made. The  $ep$  elastic cross section was determined from the same data set used for these results. Agreement within about 5% of previous values was obtained, verifying the efficiency of the hardware trigger to the same level. Other studies estimated errors due to inexact knowledge of the peak shape and background, residual misalignment of the detectors, dependence of the acceptance on the Monte Carlo input distribution, and variations of the fiducial cut edges for the  $e$  and  $p$ . The values varied between 0% and 10%. The total angle-dependent systematic error for each bin was the sum of all the components added in quadrature. Finally, the total error quoted was obtained by adding the error in the  $\eta$  yield, the acceptance error, and the systematic error in quadrature.

Angular distributions were measured as a function of c.m. decay angles  $\cos\theta_\eta^*$  and  $\phi_\eta^*$  for  $W$  for central bin values from 1.5 to 1.83 GeV and for  $Q^2 = 0.375, 0.75$ , and  $1.25$  ( $\text{GeV}/c^2$ ). Sample results for the virtual photon cross section in the center of mass frame are shown in Fig. 2. Al-

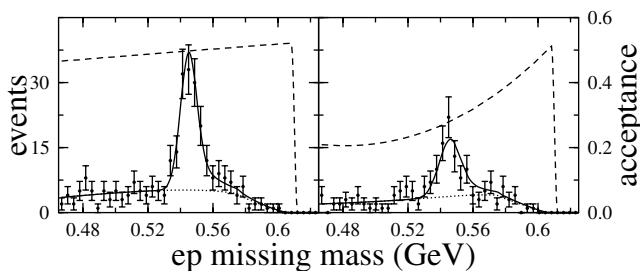


FIG. 1. Missing mass spectra for  $ep \rightarrow epX$ . Bins shown are for  $W = 1.535$  GeV and  $Q^2 = 1.25$  ( $\text{GeV}/c^2$ ). Left plot is for  $\phi_\eta^* = 22.5^\circ$  and  $\cos\theta_\eta^* = -0.4$ ; right plot is for  $\phi_\eta^* = 67.5^\circ$  and  $\cos\theta_\eta^* = 0.0$ . The dashed line (right scale) shows the acceptance. The solid and dotted lines are the full fit function and the background function only.

though all distributions have a significant isotropic component, deviations from isotropy are seen.

Figure 2 also shows predictions by the Mainz [14] and RPI [15] groups. Both models use effective Lagrangians with resonant and nonresonant terms. RPI fits parameters to photoproduction and high  $Q^2$  data [11]. Their results are fairly consistent with the new data at low  $W$ , but have the wrong slope at 1.62 GeV. The Mainz model fits photoproduction data and extends to finite  $Q^2$  using a CQM. They match the new data better at high  $W$  than RPI, but have the wrong magnitude at low  $W$ .

The exact virtual photon cross section is

$$\frac{d^2\sigma}{d\Omega_\eta^*} = \frac{|p_\eta^*|}{K_{cm}} \left[ R_T + \epsilon R_L + R_{LT} \sqrt{\frac{\epsilon}{2}} (\epsilon + 1) \cos\phi_\eta^* + R_{TT} \epsilon \cos 2\phi_\eta^* \right]. \quad (1)$$

The angular distributions were fit to a form,

$$\approx \frac{|p_\eta^*|}{K_{cm}} [A + B \cos\theta_\eta^* + CP_2(\cos\theta_\eta^*) + (D \sin\theta_\eta^* + E \sin\theta_\eta^* \cos\theta_\eta^*) \cos\phi_\eta^* + F \sin^2\theta_\eta^* \cos 2\phi_\eta^*], \quad (2)$$

assuming dominance of the  $S_{11}$  partial wave and truncation to total angular momentum up to  $\frac{3}{2}$  [14]. Here  $\epsilon$  is the polarization parameter and  $K_{cm}$  is the equivalent c.m. photon momentum. The parameters  $A$ ,  $B$ , and  $C$  contain contributions from the longitudinal ( $R_L$ ) and transverse ( $R_T$ ) response functions,  $D$  and  $E$  parametrize the longitudinal-transverse interference response function ( $R_{LT}$ ), and  $F$  contains the transverse-transverse interference response function ( $R_{TT}$ ). Krusche *et al.* [6] used a similar form for photoproduction data, but only  $R_T$  contributes there. Fit results for  $Q^2 = 0.75$  ( $\text{GeV}/c^2$ ) are shown in Fig. 3 along with the Mainz predictions [14].

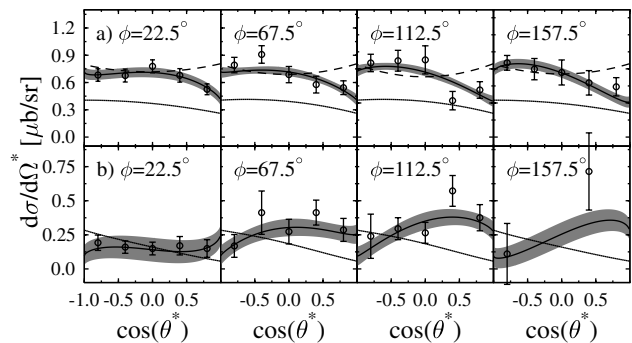


FIG. 2. Differential cross section for  $\gamma_\nu p \rightarrow p\eta$  in the center of mass frame for (a)  $W = 1.53$  GeV and  $Q^2 = 1.25$  ( $\text{GeV}/c^2$ ) and (b)  $W = 1.71$  GeV and  $Q^2 = 0.75$  ( $\text{GeV}/c^2$ ).  $W$  bins are 30 MeV wide and  $Q^2$  bins are  $0.5$   $\text{GeV}^2$ . Values for  $\phi_\eta^*$  symmetric about  $180^\circ$  have been averaged. [No information is lost this way; see Eq. (1).] Solid lines with an error band correspond to the response function fit described in the text. Dashed (solid) lines correspond to the effective Lagrangian calculation of the RPI (Mainz) group. See text for details.

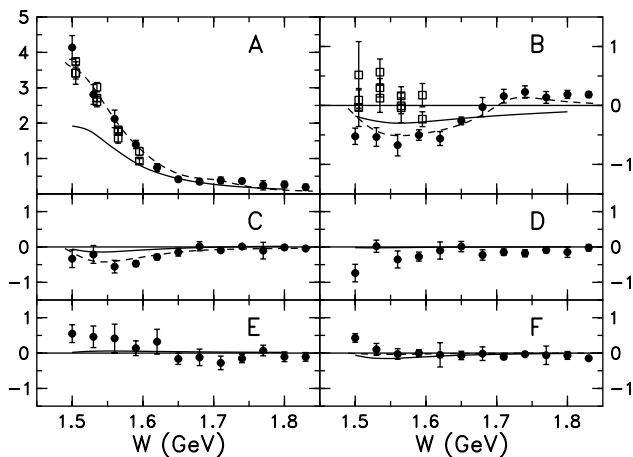


FIG. 3. Results of fitting the  $Q^2 = 0.75$  (GeV/c) $^2$  angular distribution data of this experiment to Eq. (2). See text for details. Open squares are previous data [8]. Contributions from both statistical and systematic sources are displayed. The solid line is the theoretical prediction of the Mainz group and the dashed line is a five resonance fit to A, B, C, and F.

Because of their limited kinematic ranges, previous experiments fit A and B [8] or A only [9–11]. A is mostly due to  $S_{11}(1535)$  and is the largest amplitude. It is probably dominated by the transverse amplitude because the longitudinal contribution is known to be minor [10]. The  $W$  dependence of both A and C are similar to what was seen in Krusche *et al.* [6].

Nonzero values for parameters B–F are evidence for interference with overlapping resonances or nonresonant mechanisms. For the assumptions made, B and D come from interference between  $S_{11}$  and  $P_{11}$  partial waves, and C, E, and F come from interference between  $S_{11}$  and  $D_{13}$  partial waves. More complete partial wave analyses will be required to disentangle these contributions in detail.

The value of B for electroproduction data [8] was positive and poorly determined. In the photoproduction experiment, B was slightly negative. The new result extends to much higher  $W$ . The sign change in B at  $W \sim 1.65$  GeV has not been seen before. Such a rapid change is likely due to resonance effects, perhaps the onset of  $P_{11}(1710)$ . In fact, the  $W$  dependence of A, B, C, and F can be reproduced (see Fig. 3) in a simple isobar model by including  $S_{11}(1535)$ ,  $S_{11}(1650)$ ,  $D_{13}(1520)$ ,  $P_{11}(1440)$ , and  $P_{11}(1710)$  states with standard masses and widths [1]. Neither calculation in Fig. 2 includes the  $P_{11}(1710)$ .

D, E, and F are consistent with zero over almost all of the range of  $Q^2$  and  $W$  covered. This agrees with the Mainz predictions in which  $R_{LT}$  and  $R_{TT}$  are small compared to the dominant transverse amplitude. Because of a lack of statistics, the present data provide only a qualitative test of the predictions.

Angle integrated cross sections were also obtained for events in a given  $Q^2$ ,  $W$  bin using the same methods as for the angular distributions. Distributions in  $W$  for three  $Q^2$  values are shown in Fig. 4. Structure is seen at  $W \sim$

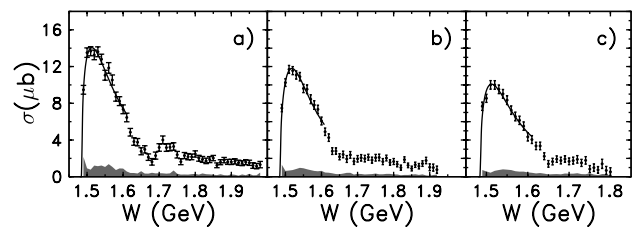


FIG. 4. New integrated cross section data at (a)  $Q^2 = 0.625$  (GeV/c) $^2$ , (b)  $Q^2 = 0.875$  (GeV/c) $^2$ , and (c)  $Q^2 = 1.125$  (GeV/c) $^2$ . The shaded band shows systematic errors. The curves correspond to single-resonance Breit-Wigner fits with an energy-dependent width over the energy range shown.

1.65 GeV. A dip followed by a peak is seen at  $Q^2 = 0.625$  (GeV/c) $^2$  while a significant change in slope is seen at other  $Q^2$ . This  $W$  is where B (see above) changes sign and both probably have the same cause.

Fits to a single Breit-Wigner [ $S_{11}(1535)$  only] shape [14] are also shown in Fig. 4. The nonresonant contribution is ignored in this fit [14,15]. Results are very dependent on the  $W$  range chosen because other contributions become prominent at higher  $W$ . We find the best fit with a maximum  $W$  of 1.62 GeV. These fits give a resonance mass of  $1522 \pm 11$  MeV and full width of  $143 \pm 18$  MeV. A coupled-channels analysis will be required to get the most reliable values. The maximum cross section for the new and all previous experiments was then used to determine  $A_{1/2}$  [14]. For consistent comparison, a full width of 150 MeV and an  $S_{11} \rightarrow \eta N$  branching fraction of 0.55 [4,11] were used. New and reanalyzed old measurements of  $A_{1/2}$  are shown in Fig. 5. A number of values for  $A_{1/2}$  from  $\pi$  and  $\eta$  photoproduction ( $Q^2 = 0$ ) data with significant model dependence have been reported. The PDG value [1] of  $0.09 \pm 0.03$  GeV $^{-1/2}$  reflects this uncertainty.

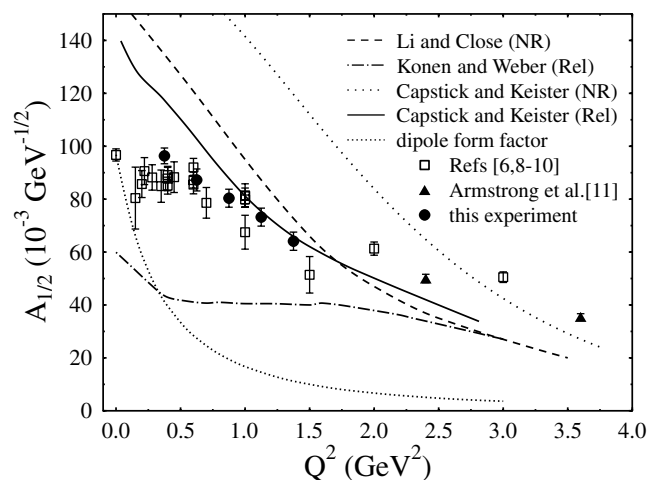


FIG. 5. Values of the photon coupling amplitude,  $A_{1/2}$ , for  $\gamma p \rightarrow S_{11}(1535)$  obtained from the integrated cross section data of this experiment compared to previous data and various CQM calculations [16]. All statistical and systematic errors from data are included.

The theoretical calculations shown in Fig. 5 are nonrelativistic and relativistic CQM predictions [16]. None agree well with the data or each other, an important failure of the CQM.

These and other recent data from JLab [11] have comparable values for the  $S_{11}$  Breit-Wigner width ( $\sim 150$  MeV) for  $Q^2$  between 0.375 and 3.6  $(\text{GeV}/c)^2$ . Photoproduction [6] and Brasse electroproduction values [8] are surprisingly different, 239 and  $\sim 90$  MeV, respectively, in our fits to their data. At high  $Q^2$ , the Brasse data is incompatible with Armstrong [11]; at low  $Q^2$ , either the single resonance interpretation is incorrect or there is a significant change in dynamics with increasing  $Q^2$ .

The eta electroproduction data shown here comprise one of the first results of the CLAS. They cover the region in  $W$  at and above the  $S_{11}(1535)$  resonance in great detail. A consistent picture of the reaction is given over a more extensive kinematic range than any previous data. The values for the interference response functions are small compared to the dominant transverse response function and in qualitative agreement with theoretical predictions [14].  $B$  measures the interference between  $S_{11}$  and  $P_{11}$  partial waves. It is more negative than the photoproduction results [6] at low  $W$  and changes sign at  $W \sim 1.65$  GeV, likely signaling the onset of a strong  $P$ -wave process. At the same  $W$ , a sharp change in slope is seen in the integrated data. The angular distribution data can be described by a simple isobar model using known states. The  $\eta N$  coupling strengths for these states are poorly known, but can be determined using these data. A simple determination of the  $\gamma p \rightarrow S_{11}(1535)$  photon coupling amplitude provides new and far more consistent evidence for its unusually slow falloff with increasing  $Q^2$ . These data along with other recent Jefferson Lab data [11] provide inescapable constraints on models attempting to describe the structure of  $S_{11}(1535)$ .

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*Note added.*—Since this paper was submitted, results from a photoproduction experiment [17] in a similar energy range have appeared. That preprint confirms the  $S$ - $P$  wave interference observed in this paper and finds an  $S_{11}$  width which is compatible with our measurement.

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- [1] D. E. Groom *et al.*, Eur. Phys. J. C **15**, 1 (2000).
  - [2] S. Capstick, Phys. Rev. D **46**, 2864 (1992).
  - [3] D. B. Leinweber *et al.*, Phys. Rev. D **61**, 074502 (2000); S. Sasaki, hep-ph/0004252.
  - [4] T. P. Vrana, S. A. Dytman, and T.-S. H. Lee, Phys. Rep. **328**, 181 (2000).
  - [5] N. Kaiser, T. Waas, and W. Weise, Nucl. Phys. A **612**, 297 (1997).
  - [6] B. Krusche *et al.*, Phys. Rev. Lett. **74**, 3736 (1995).
  - [7] J. Ajaka *et al.*, Phys. Rev. Lett. **81**, 1797 (1998).
  - [8] F. W. Brasse *et al.*, Nucl. Phys. B **139**, 37 (1978); F. W. Brasse *et al.*, Z. Phys. C **22**, 33 (1984).
  - [9] U. Beck *et al.*, Phys. Lett. B **51**, 103 (1974).
  - [10] H. Breuker *et al.*, Phys. Lett. B **74**, 409 (1978).
  - [11] C. S. Armstrong *et al.*, Phys. Rev. D **60**, 052004 (1999).
  - [12] W. Brooks, Nucl. Phys. A **663–664**, 1077 (2000).
  - [13] Richard A. Thompson, Ph.D. thesis, University of Pittsburgh, 2000 (unpublished).
  - [14] G. Knochlein, D. Drechsel, and L. Tiator, Z. Phys. A **352**, 327 (1995); L. Tiator, C. Bennhold, and S. S. Kamalov, Nucl. Phys. A **580**, 455 (1994).
  - [15] M. Benmerrouche, N. C. Mukhopadhyay, and J. F. Zhang, Phys. Rev. D **51**, 3237 (1995); R. Davidson (private communication).
  - [16] S. Capstick and B. D. Keister, Phys. Rev. D **51**, 3598 (1995); W. Konen and H. J. Weber, Phys. Rev. D **41**, 2241 (1990); F. E. Close and Z. Li, Phys. Rev. D **42**, 2194 (1990).
  - [17] F. Renard *et al.*, arXiv: hep-ex/0011098.