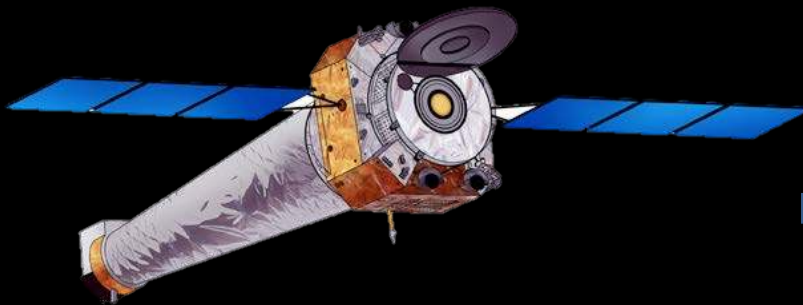
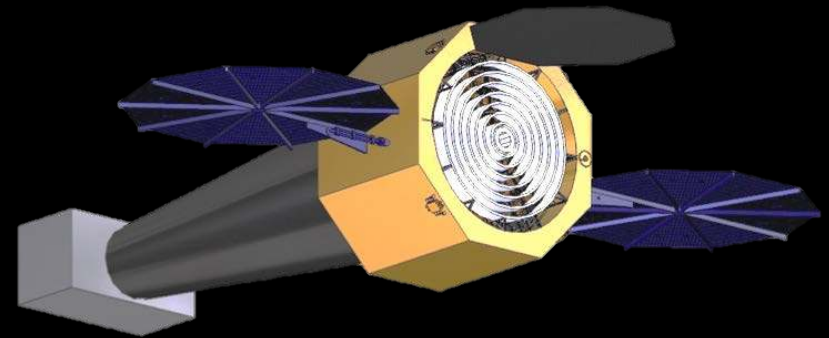


*Adaptive X-ray Optics IV (SPIE 9965)*  
2016 August 28; San Diego, CA (USA)

# Toward large-area sub-arcsecond x-ray telescopes II



*Chandra X-ray Observatory (1999-?)*



*X-Ray Surveyor concept (~2035?)*

Steve O'Dell  
*NASA Marshall Space Flight Center*  
and co-authors

# Authors represent most of the US effort toward sub-arcsecond x-ray telescopes.

Steve O'Dell<sup>a</sup>, Ryan Allured<sup>b</sup>, Andrew Ames<sup>b</sup>, Michael Biskach<sup>c</sup>, David Broadway<sup>a</sup>, Ricardo Bruni<sup>b</sup>, Dave Burrows<sup>d</sup>, Jian Cao<sup>e</sup>, Brandon Chalifoux<sup>f</sup>, Kai-Wing Chan<sup>g</sup>, Yip-Wah Chung<sup>e</sup>, Vincenzo Cotroneo<sup>b</sup>, Ron Elsner<sup>a</sup>, Jessica Gaskin<sup>a</sup>, Mikhail Gubarev<sup>a</sup>, Ralf Heilmann<sup>f</sup>, Ed Hertz<sup>b</sup>, Tom Jackson<sup>d</sup>, Kiran Kilaru<sup>h</sup>, Jeff Kolodziejczak<sup>a</sup>, Ryan McClelland<sup>c</sup>, Brian Ramsey<sup>a</sup>, Paul Reid<sup>b</sup>, Raul Riveros<sup>g</sup>, Jackie Roche<sup>a</sup>, Suzanne Romaine<sup>b</sup>, Timo Saha<sup>i</sup>, Mark Schattenburg<sup>f</sup>, Eric Schwartz<sup>b</sup>, Dan Schwartz<sup>b</sup>, Peter Solly<sup>c</sup>, Susan Trolier-McKinstry<sup>d</sup>, Mel Ulmer<sup>e</sup>, Alexey Vikhlinin<sup>b</sup>, Margeaux Wallace<sup>d</sup>, Xiaoli Wang<sup>e</sup>, David Windt<sup>j</sup>, Youwei Yao<sup>e,f</sup>, Shi Ye<sup>e</sup>, Will Zhang<sup>i</sup>, & Heng Zuo<sup>f</sup>

<sup>a</sup> NASA Marshall Space Flight Center (USA)

<sup>b</sup> Harvard-Smithsonian Center for Astrophysics (USA)

<sup>c</sup> Stinger Ghaffarian Technologies, Inc., Goddard Space Flight Center (USA)

<sup>d</sup> Pennsylvania State University (USA)

<sup>e</sup> Northwestern University (USA)

<sup>f</sup> Massachusetts Institute of Technology (USA)

<sup>g</sup> University of Maryland Baltimore County, Goddard Space Flight Center (USA)

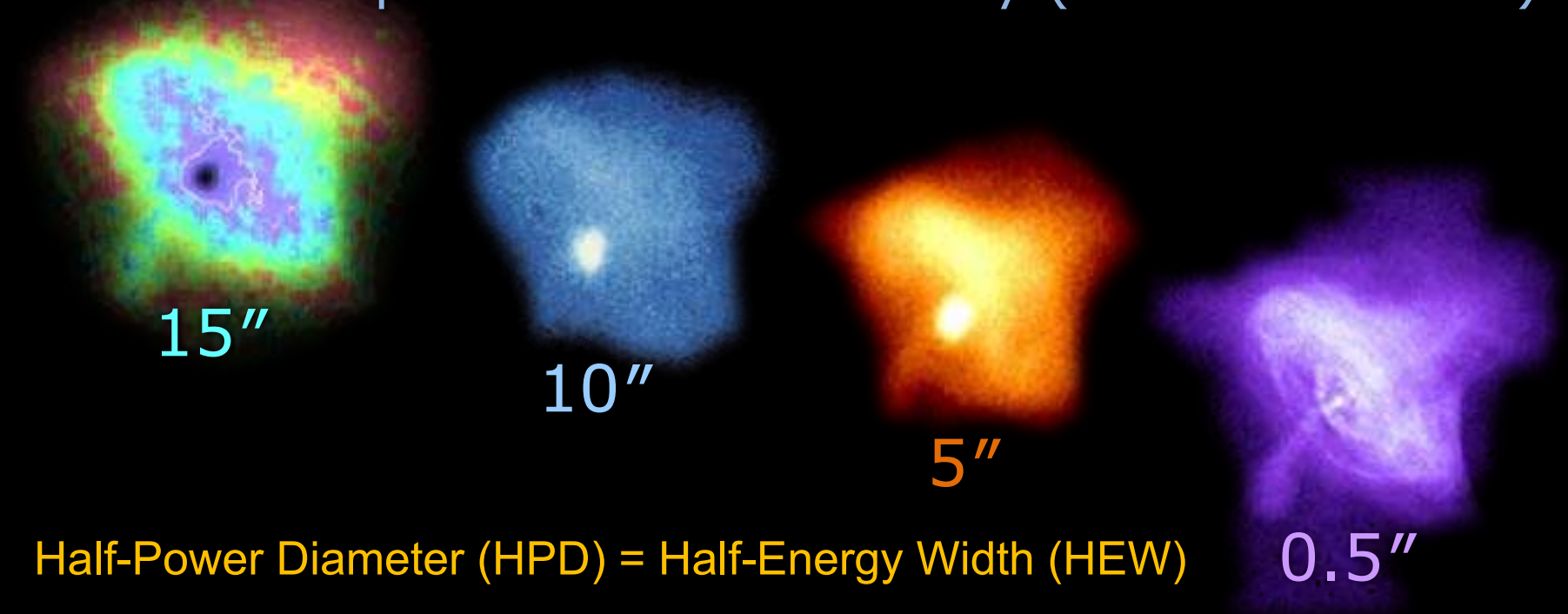
<sup>h</sup> Universities Space Research Association, Marshall Space Flight Center (USA)

<sup>i</sup> NASA Goddard Space Flight Center (USA)

<sup>j</sup> Reflective X-ray Optics LLC (USA)

# Key science metrics of an x-ray telescope are angular resolution and aperture area.

Finer angular resolution improves imaging quality and point-source sensitivity (decreases noise).



Larger aperture area improves sensitivity (increases signal), down to angular-resolution confusion limit.

# Outline

- Motivation and issues
- Categories of potential solutions
- Post-fabrication corrections

# Outline

- Motivation and issues
  - Seek *Chandra* resolution with  $30 \times$  *Chandra* area.
  - Must resolve both technologic and programmatic challenges.
    - Achieve aperture area and imaging performance within constraints of mass, envelope, cost, and schedule.
- Categories of potential solutions
- Post-fabrication corrections

# 2020 Decadal Survey in Astronomy and Astrophysics sets priorities for decade.

- National Research Council (NRC) conducts Survey for relevant Government agencies (NASA & NSF).
  - ❑ Addresses space-based and ground-based astronomy.
  - ❑ NASA typically adopts Decadal Survey's priorities.
  - ❑ NASA will suggest space-based astronomy missions.
    - Four Science and Technology Definition Teams (STDT) are developing facility-class mission concepts for NASA.
      - X-Ray Surveyor
      - Far-IR Surveyor
      - LUVOIR Surveyor
      - Habitable Exoplanet Imaging (HabEx)
    - Only one of these is likely to start in the 2020s.
      - Expected launch date would be mid-to-late 2030s.

# STDT is defining science-driven requirements for X-Ray Surveyor.

- Anticipate scientific need for an x-ray telescope with *Chandra's* 0.5" resolution and 30 × its area.
- There is general agreement on the top-level technologic and programmatic challenges.
  - ❑ Preserve *Chandra's* angular resolution  $\approx 0.5''$  HPD.
  - ❑ Reduce mirror areal mass to *Chandra*/30  $\approx 1.5$  kg/m<sup>2</sup>.
  - ❑ Reduce mirror areal cost to *Chandra*/30  $\approx 1$  M\$/m<sup>2</sup>.
- There is not general agreement on how to solve these challenges.
  - ❑ Stiff optics (rigidly supported over-constrained mirrors)
  - ❑ Static optics (fixed, moderately constrained mirrors)
  - ❑ Active optics (adjustable alignment | figure correction)

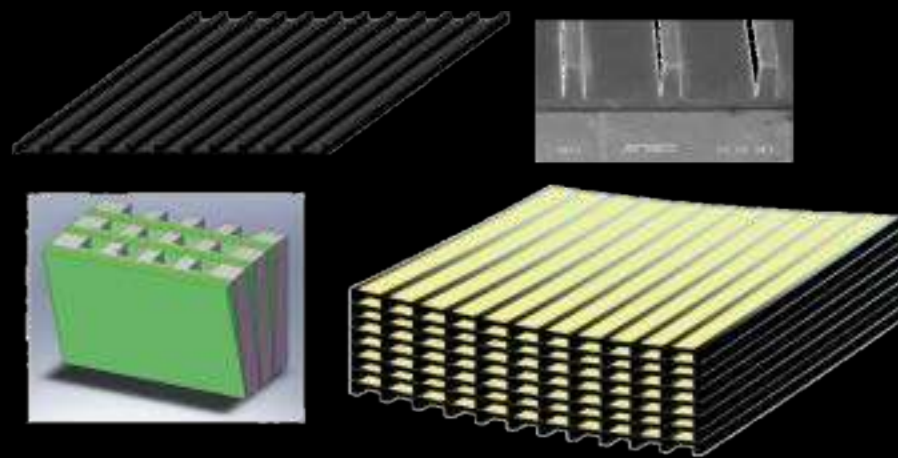
# Outline

- Motivation and issues
- Categories of potential solutions
  - Stiff optics
  - Static optics
  - Active optics
- Post-fabrication corrections



# ESA selected ATHENA as its 2<sup>nd</sup> large-class (L2) mission, for launch in 2028.

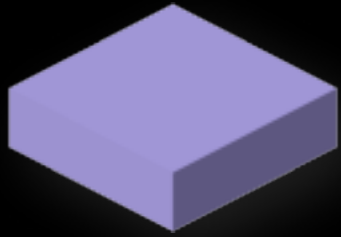
- X-ray mission is in study Phase A, working toward adoption around 2020.
  - Require HEW  $\approx 5''$ .
- ATHENA will use stiff, silicon-pore optics (SPO).
  - Cosine Measurement Systems [NL] leads SPO technology development.
    - Processing and stacking of silicon wafers are highly automated.
    - Resulting rigid x-ray optics units are to be co-aligned.



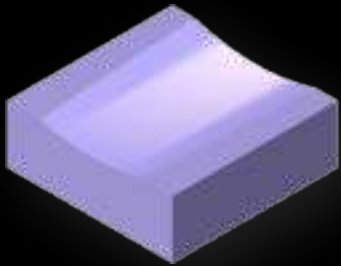
[www.the-athena-x-ray-observatory.eu](http://www.the-athena-x-ray-observatory.eu)

ESA/ Marcos Bavdaz

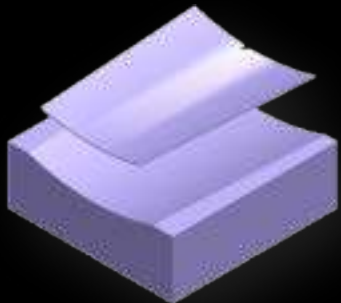
# GSFC uses a novel process for light x-ray mirrors of monocrystalline silicon.



1. Procure single crystalline silicon.
2. Heat and chemically etch to remove all surface/subsurface damage.



1. Wire-EDM machine conical shape.
2. Heat and chemically etch to remove damage.
3. Polish to achieve excellent figure and micro-roughness.



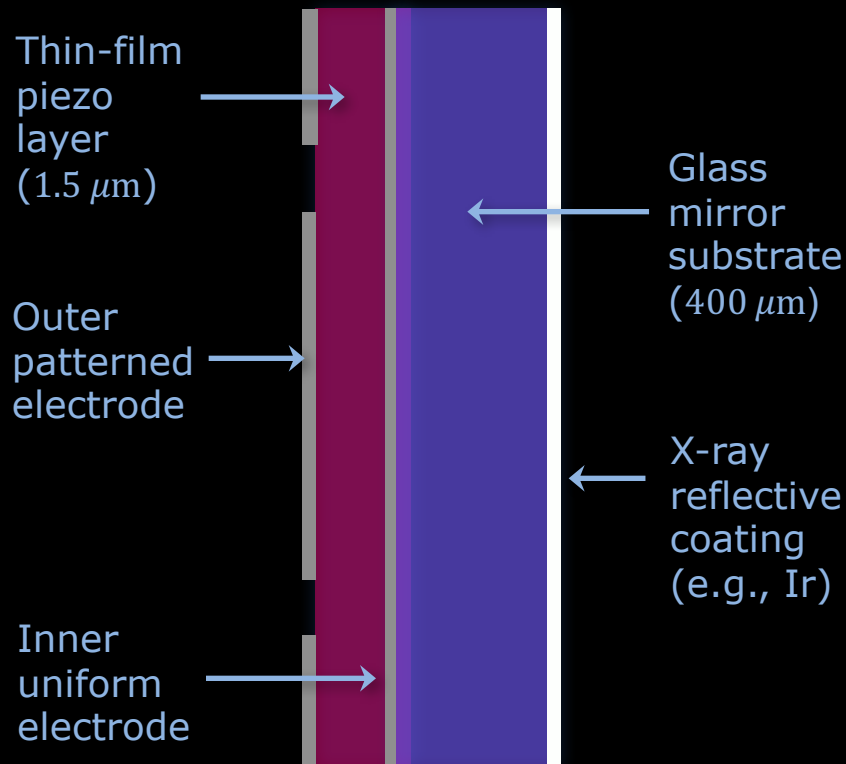
1. Use Wire-EDM to slice off the thin mirror segment.
2. Heat and chemically etch to remove all damage from back and edges.

- Advantages over glass
  - ❑ No internal stress
    - Distortion-free material removal after etching
  - ❑ Ideal material properties
    - High thermal conductivity
    - Low thermal expansion
    - High elastic modulus
  - ❑ Better performance
    - Made mirrors < 3" HEW<sub>2</sub>
    - Aim to build ~ 1" mirror stack by end 2017
    - Aim to build ~ ≥.1" meta-shell by end 2019.

GSFC/ Will Zhang

# SAO is developing lightweight active x-ray optics with thin-film piezo arrays.

- Piezoelectric array corrects figure through surface-parallel actuation

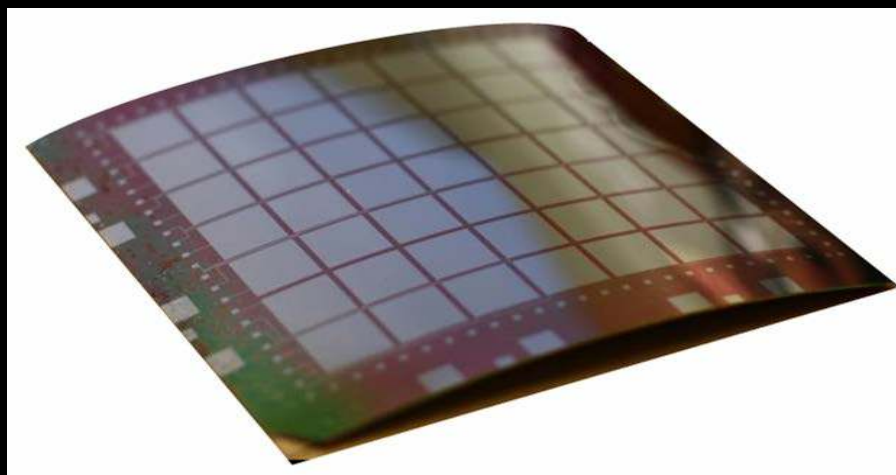


- PSU is fabricating thin-film piezoelectric arrays.
  - ❑ Slumped-glass mirrors from SAO (OAB)
  - ❑ Piezoelectric (PZT) on conductive film, high-T crystallized and annealed
  - ❑ Patterned electrode array with ZnO TFTs
    - Row-column addressing
    - Anisotropic Conductive Film (ACF) connections
  - ❑ Integral T-compensated strain gauges

SAO/ Paul Reid

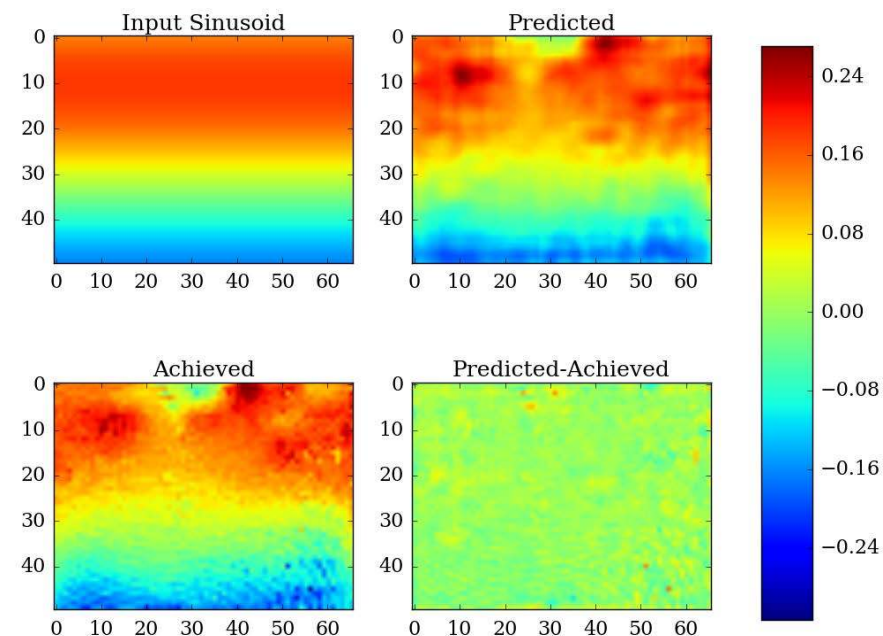
# Correction uses calibrated influence functions to determine array voltages.

- Adjustment methodology
  - ❑ Shack-Hartman wavefront sensor metrology
    - Calibrate influence function
    - Measure mirror figure
  - ❑ Calculate and apply voltages for correction



- Correction matrix

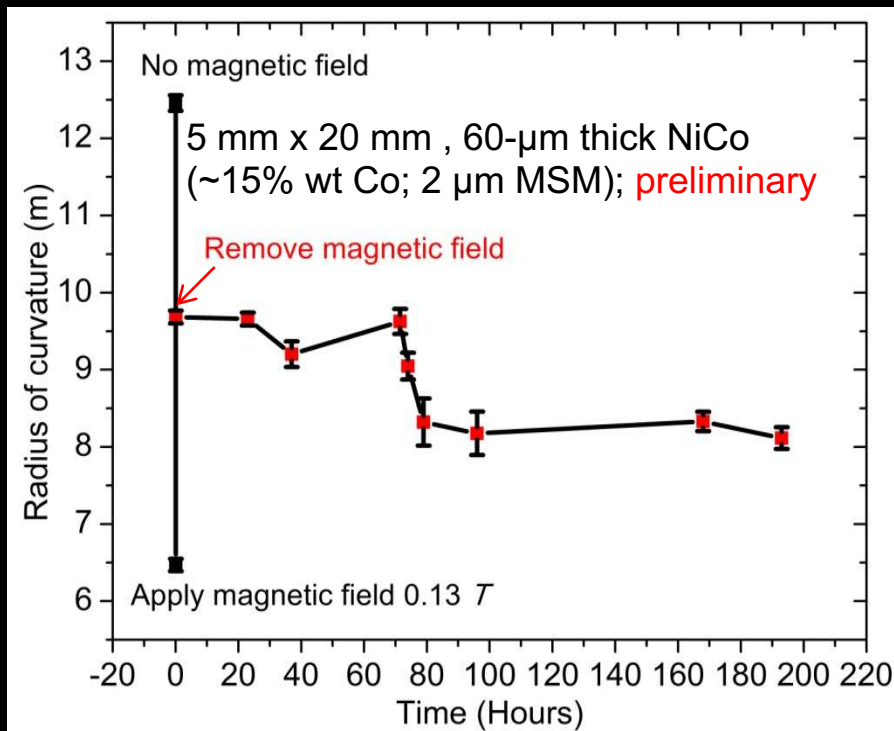
$$\begin{pmatrix} IF_{11} & \dots & IF_{1n} \\ \vdots & \ddots & \vdots \\ IF_{m1} & \dots & IF_{mn} \end{pmatrix} \begin{pmatrix} V_1 \\ \vdots \\ V_m \end{pmatrix} = \begin{pmatrix} D_1 \\ \vdots \\ D_n \end{pmatrix}$$



SAO/ Ryan Allured

# A magnetic smart material MSM provides writable surface-parallel actuation.

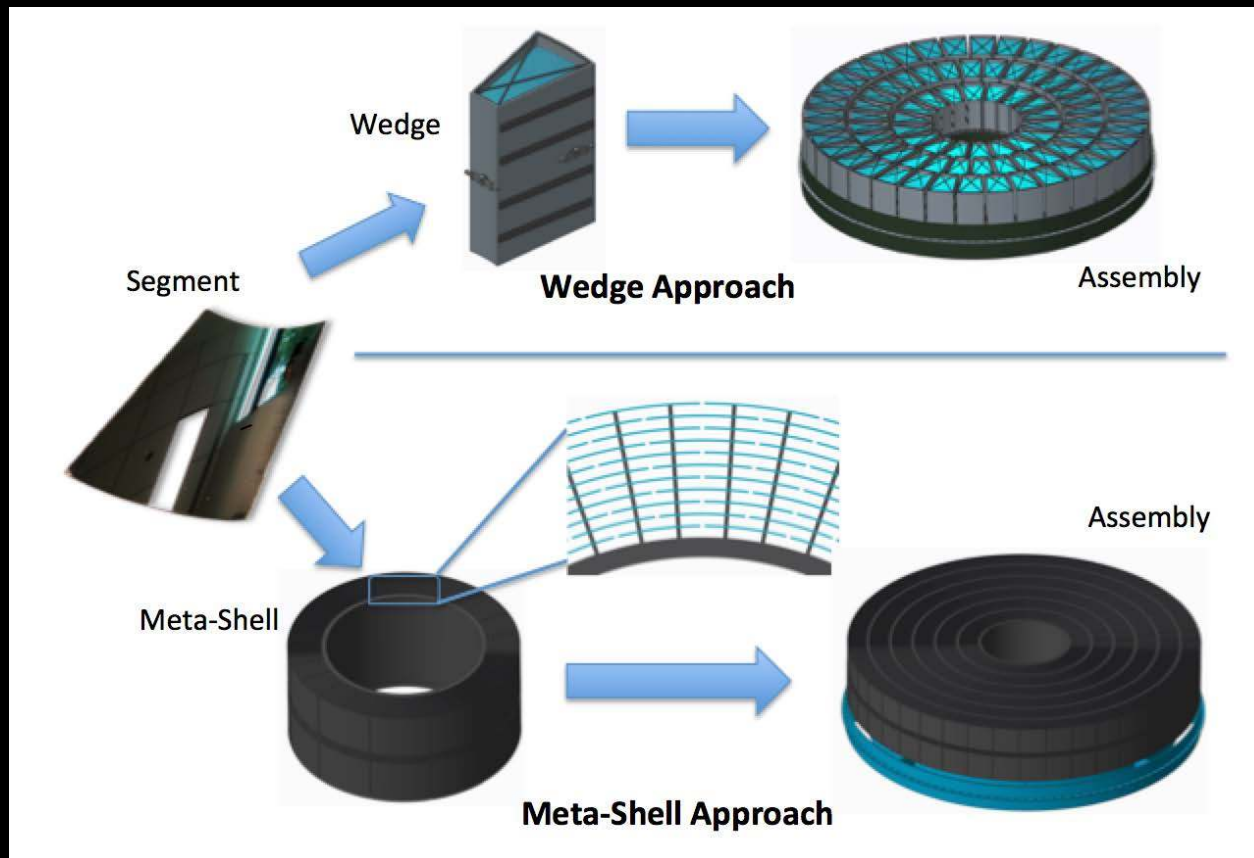
- Use a magnetically hard substrate or coated layer.
- Deposit magnetostrictive (MSM) thin film on back.
- Northwestern Univ. is developing MSM actuators for x-ray optics.
  - ❑ Demonstrated that concept may work.
  - ❑ Must speed up relaxation and stabilization.
  - ❑ Are building models to compare to experiments.
- Can also control coating stress to provide some static figure correction.



NWU/ Mel Ulmer  
NWU/ Xiaoli Wang

# Synthesize large-area nested Wolter-1-like telescope with aligned segments.

- Few-100-m<sup>2</sup> surface area in several 1000 segments



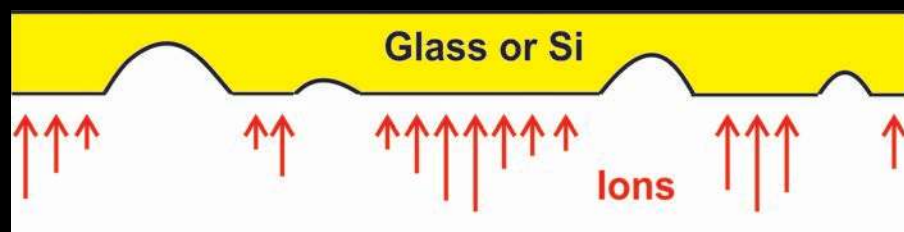
GSFC/ Will Zhang

# Outline

- Motivation and issues
- Categories of potential solutions
- **Post-fabrication corrections**
  - ❑ Differential erosion or deposition
  - ❑ Coating stress manipulation
  - ❑ Ion implantation

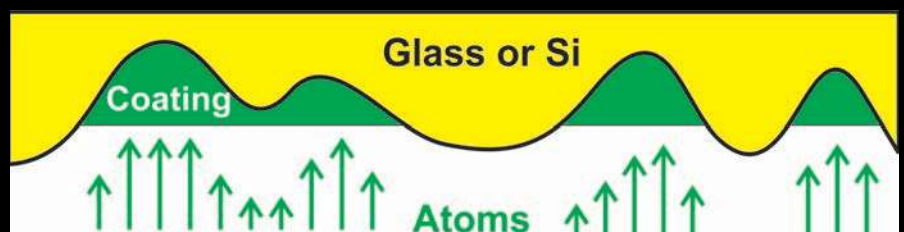
# Low-pressure subtractive or additive machining corrects residual figure errors.

## Differential Erosion

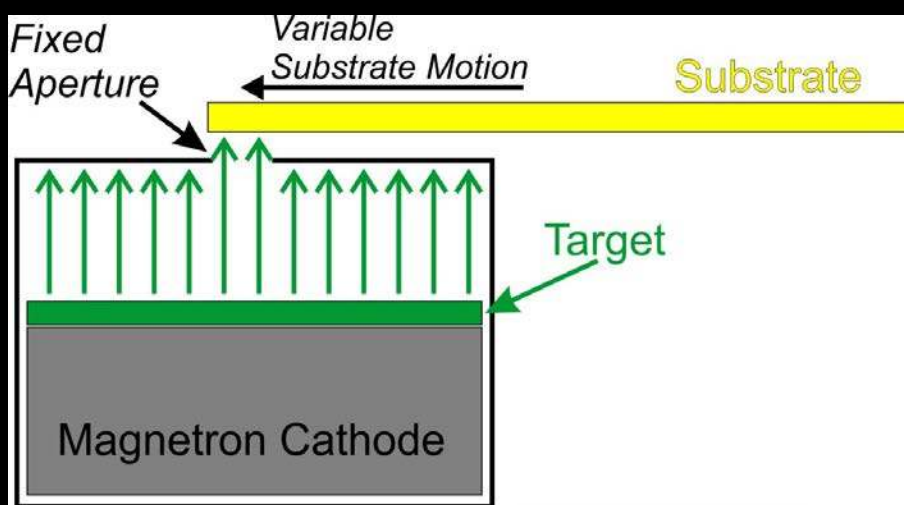
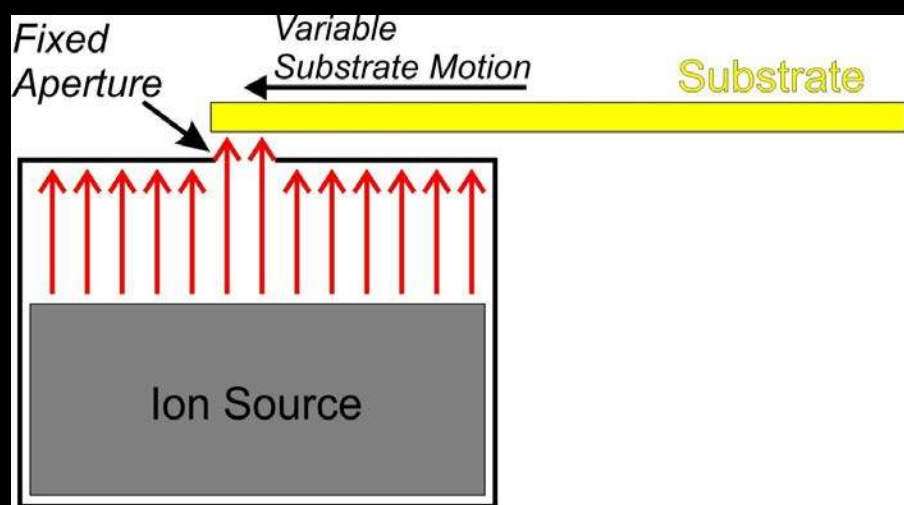


Ion-Beam Figuring

## Differential Deposition



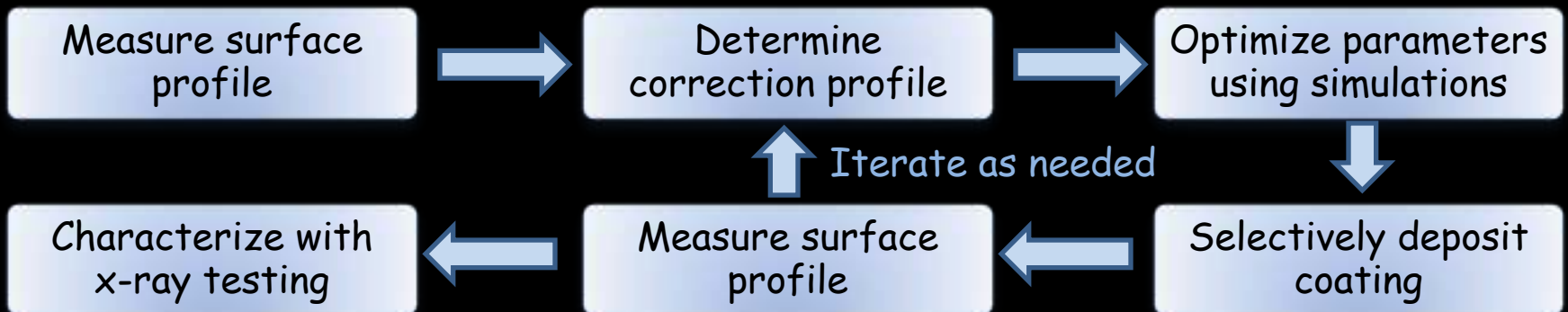
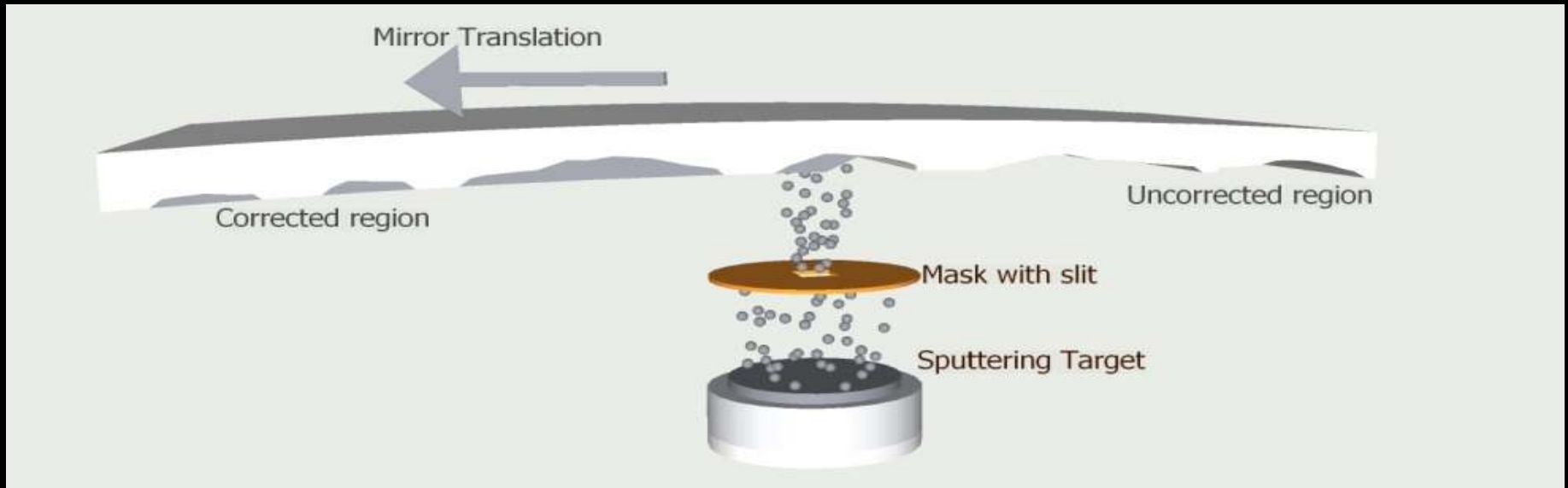
Selective Deposition



RXO/ David Windt

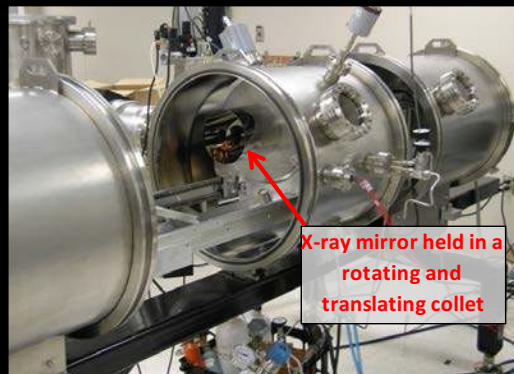


# Differential deposition allows correction of mid-frequency figure errors.

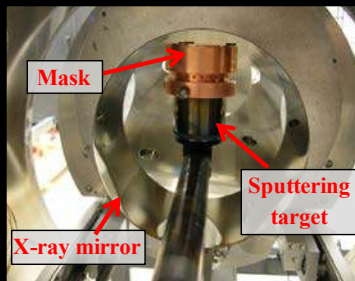


MSFC/USRA/ Kiran Kilaru

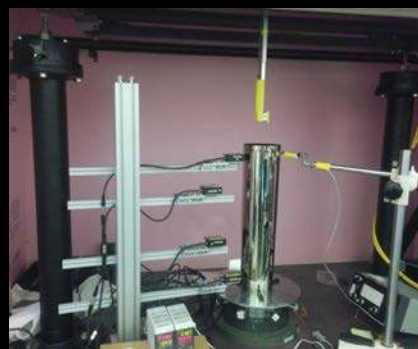
# MSFC is applying differential deposition to correct thin-walled x-ray mirrors.



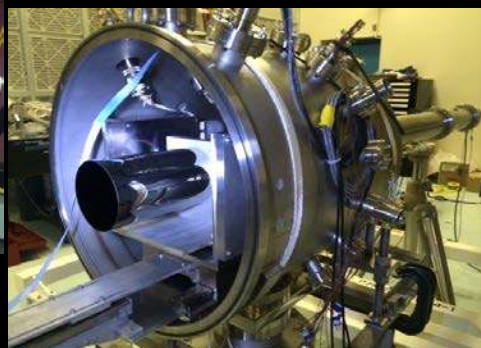
Horizontal differential-deposition chamber



Sputtering head with copper mask positioned inside shell

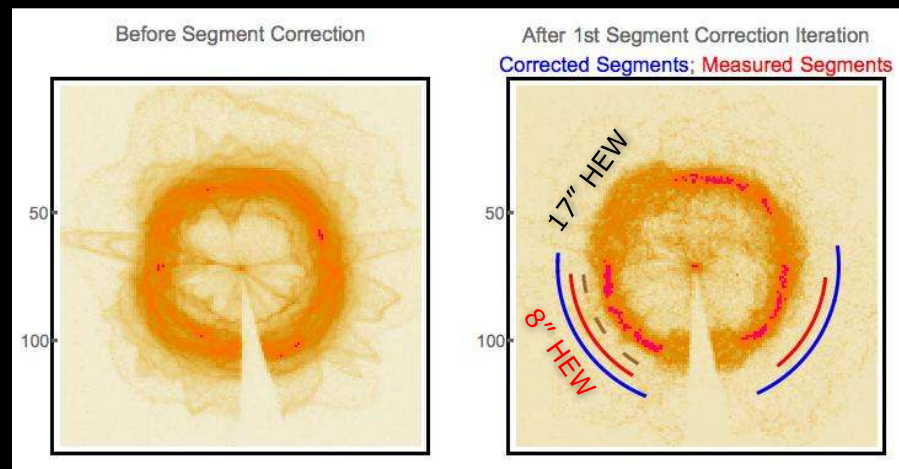


Metrology of shell with MSFC VLTP and circularity test stand



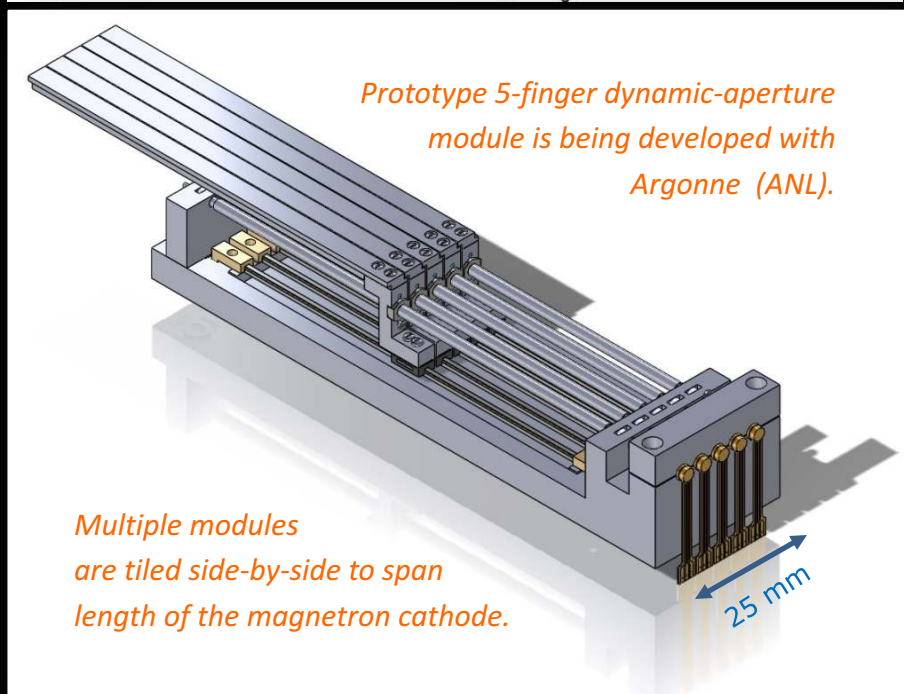
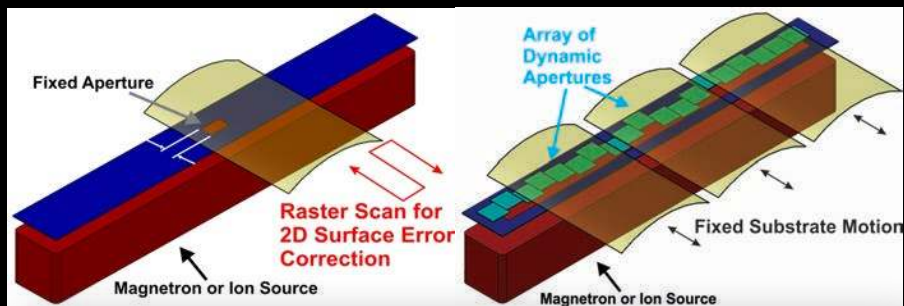
Corrected shell installed in bell housing for x-ray testing at MSFC 100-m beam

- Results are promising.
  - ❑ 2-pass correction of HEW by factor of 3 (metrology)
  - ❑ 1-pass correction of HEW by factor of 2 (x-ray test)
    - 3 corrected azimuthal sections in intrafocal image

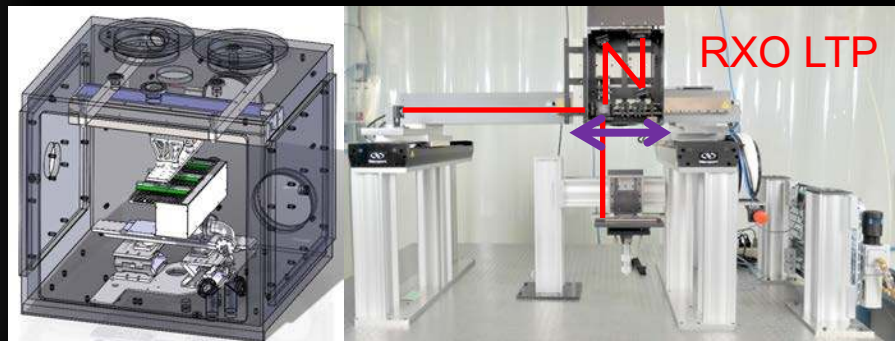


MSFC/USRA/ Kiran Kilaru

# RXO is developing dynamic apertures for 2D differential erosion or deposition.



- Dynamic aperture array
  - ❑ Simultaneously corrects multiple azimuths
    - Constant axial speed
    - Modulated aperture widths
  - ❑ Applications
    - Differential erosion
    - Differential deposition
    - Laterally graded multilayer



RXO/ David Windt

# Coating stress is an issue for sub-arcsecond imaging with thin mirrors.

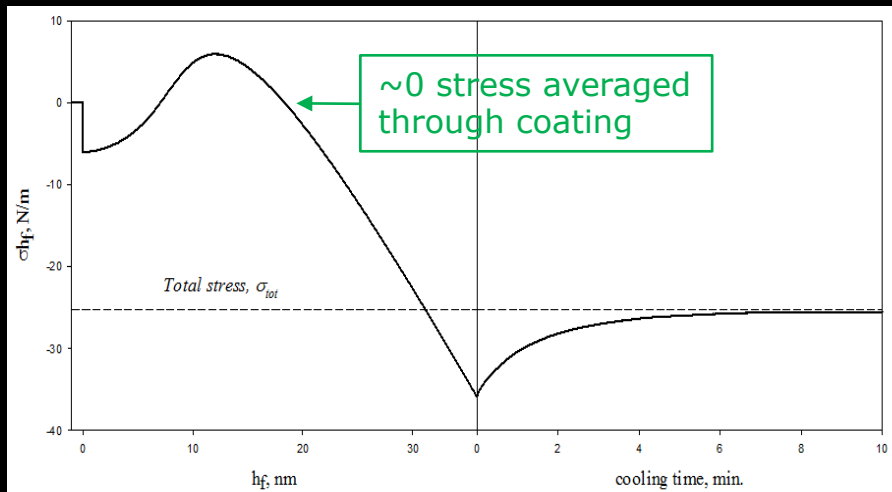
- Astronomical x-ray mirrors use thin-film coatings.
  - High-density coating or multilayers for x-ray reflectivity
  - For active optics, piezoelectric array
- Distortion depends upon film stress  $\sigma_f$  and thickness  $h_f$  and upon substrate thickness  $h_s$ .
  - E.g., Stoney formula:  $\kappa = 6(1 - \nu_s)\sigma_f h_f / (E_s h_s^2)$ 
    - Dependence upon substrate thickness  $h_s$  is quadratic.
    - Key coating parameter is the integrated stress  $\sigma_f h_f$ .
  - Both intrinsic coating stress and temperature-dependent strain (CTE differences) cause distortion.
    - Separating these two effects can be challenging.
    - Annealing or other relaxation may play a role.

# Several groups are investigating various methods for controlling coating stress.

- Monitor integrated stress in situ during sputtering, to take advantage of dynamics of thin-film growth.
  - MSFC/ David Broadway
- Deposit bilayer to tune the net integrated stress of compressive and tensile thin films.
  - RXO/ David Windt; SAO/ Suzanne Romaine
- Deposit coating on front and on back to balance integrated stress.
  - GSFC/ Kai-Wing Chan
- Anneal coating at elevated temperature to relieve stress.
  - GSFC/ Kai-Wing Chan

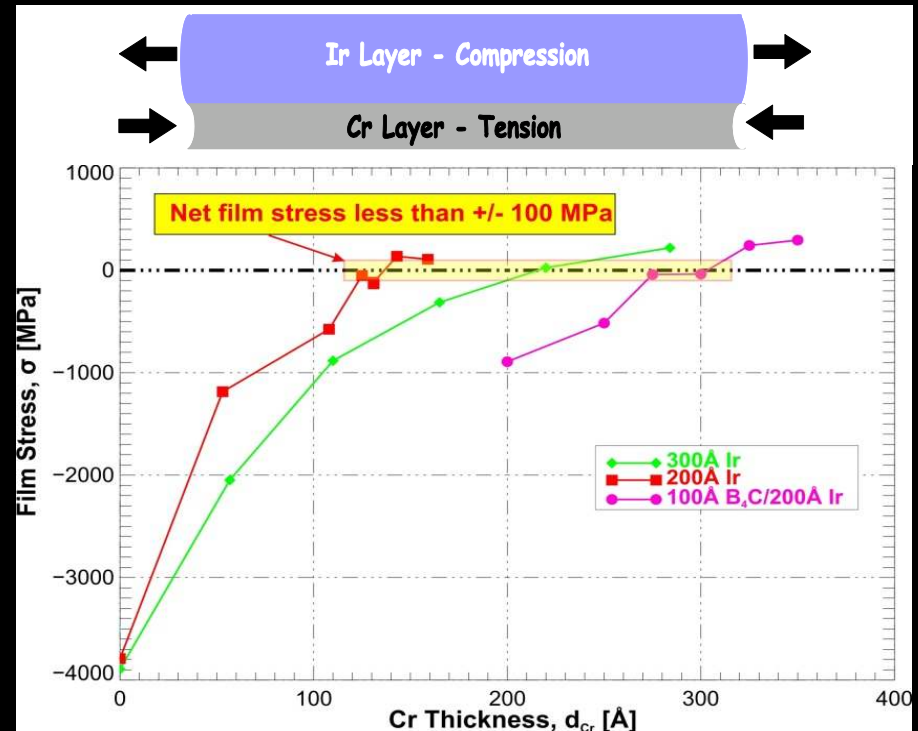
# Methods for controlling deposition stress of thin films (continued)

- Iridium film growth
  - ❑ Tensile through coalescence stage
  - ❑ Very compressive after
- Tune mean stress of film
  - ❑ -3 MPa, 0.5-nm rough



MSFC/ David Broadway

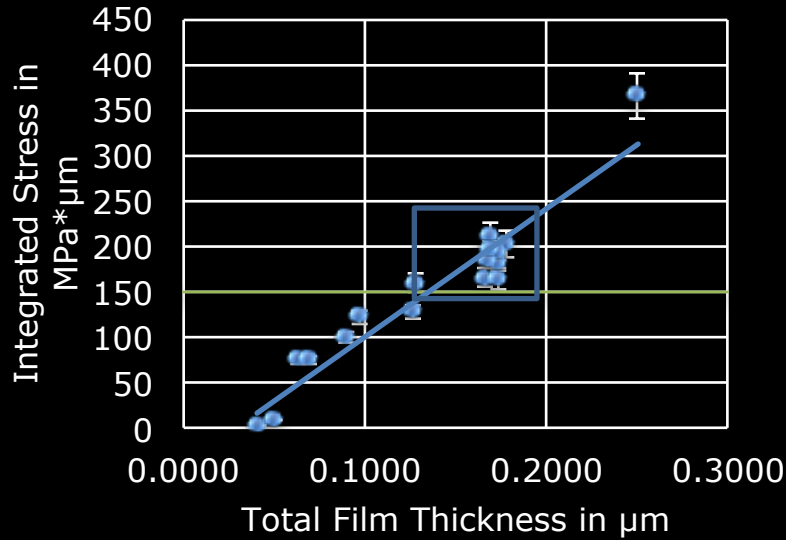
- Bilayer cancelation
  - ❑ Ir compressive stress
  - ❑ Cr tensile stress



RXO/ David Windt

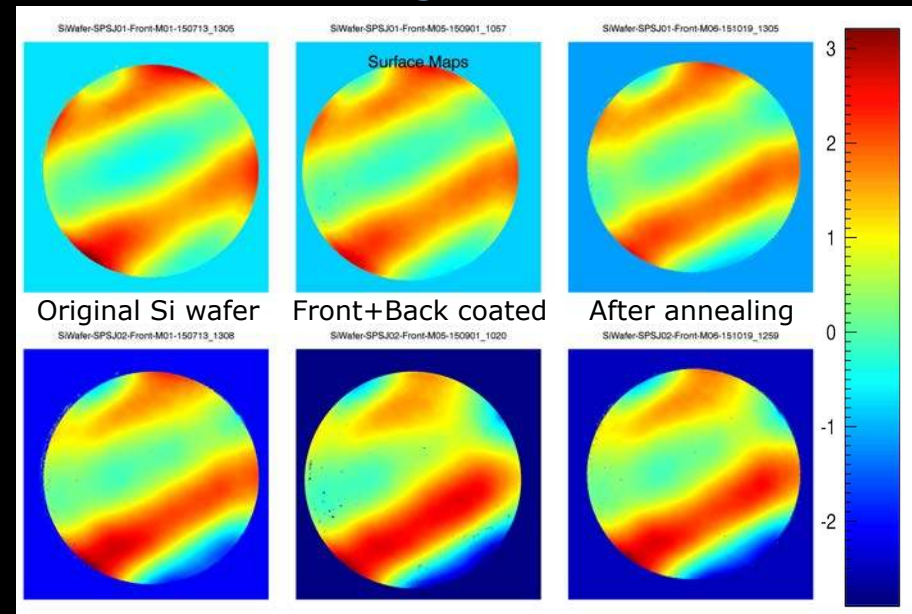
# Methods for controlling deposition stress of thin films (continued)

- PZT compensation
  - ❑ PZT tensile stress on back
  - ❑ Ir/Cr bilayer on front
- Tune net integrated stress
  - ❑ 10-nm Ir + 160 nm Cr



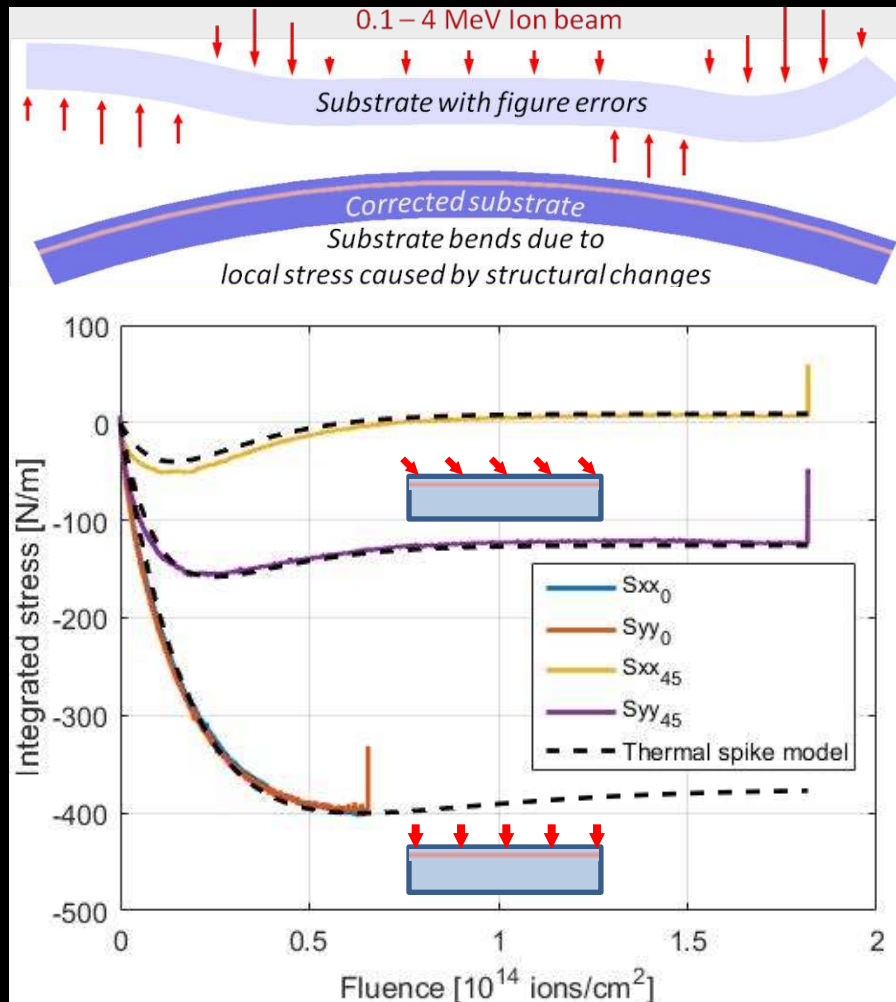
SAO/ Suzanne Romaine

- Multiple approaches
  - ❑ Front and back coating
    - Sputtering
    - Atomic Layer Deposition
  - ❑ Annealing at  $320^\circ\text{C}$



GSFC/ Kai-Wing Chan

# Differential stress from ion implantation allows static correction of figure errors.



- MIT is using its ion beam to develop this approach.
  - ❑ Operates at 1-6 MeV.
    - Implant depth 1 – 4  $\mu\text{m}$
  - ❑ Low surface degradation
  - ❑ Integrated stress measure
    - Thermal spike model
    - Anisotropic ( $\parallel$  vs  $\perp$ ) stress
    - Dose-dependent relaxation



MIT/ Brandon Chalifoux



# Outline

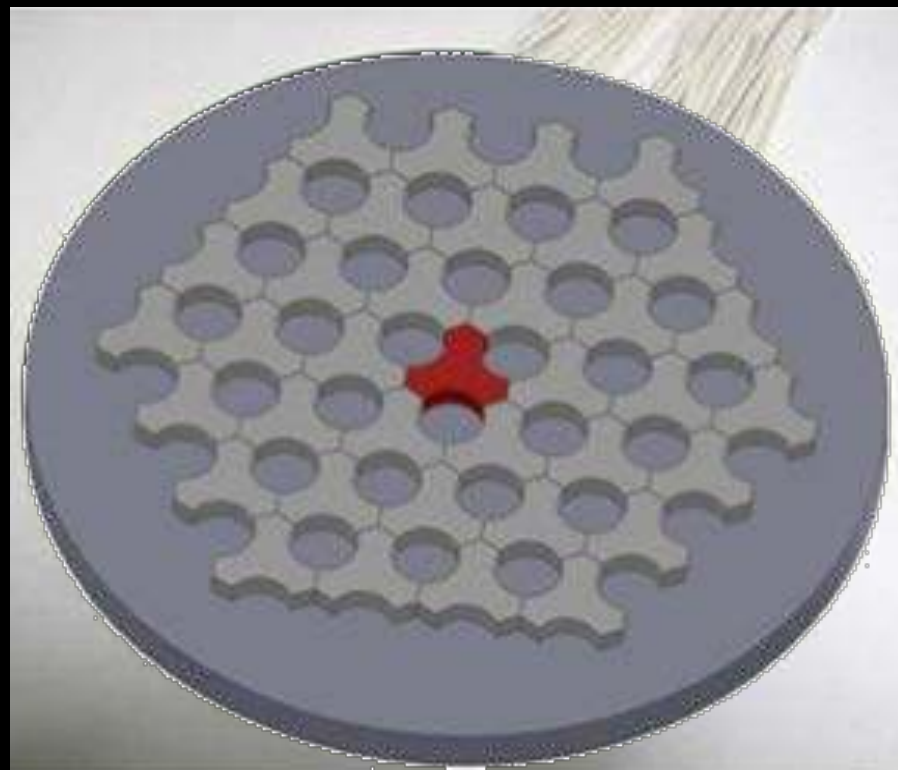
- Motivation and issues
- Categories of potential solutions
- Post-fabrication corrections

# Back-up slides

# An array of electroactive pads provides surface-tangential actuation (STA).



Xinetics deformable mirror (DM) uses  $4 \times 27$  array of electrostrictive (PMN) pads bonded to mirror. Xinetics and ANL characterized DM.

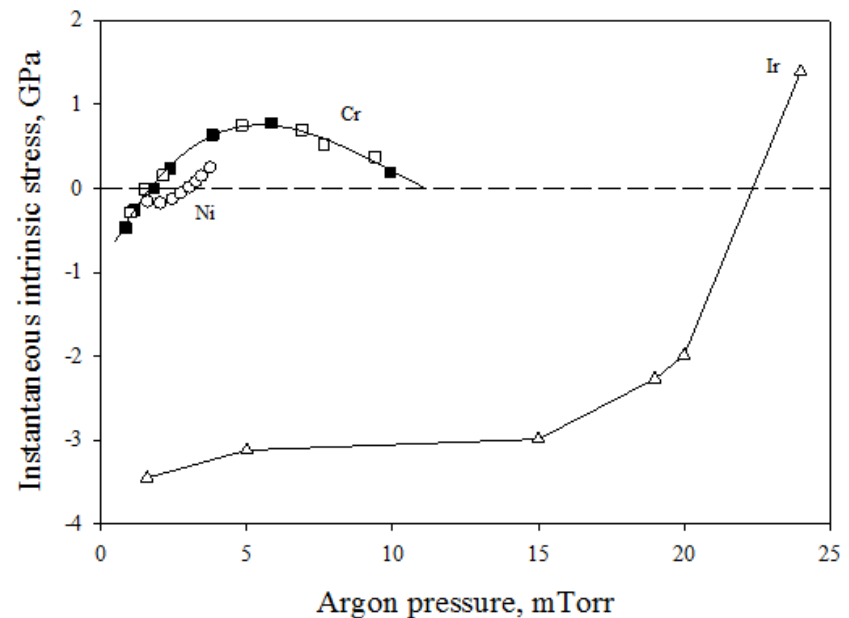
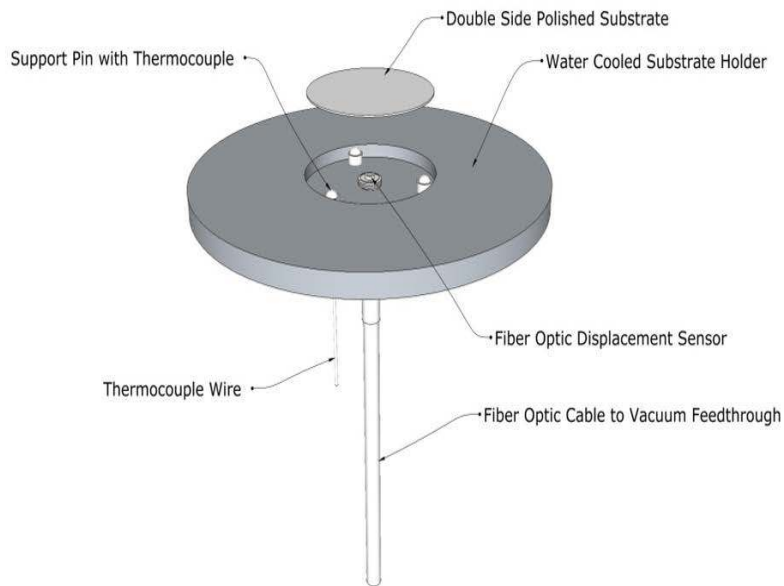


Xinetics prototype uses PMN array bonded to a silicon mirror. Each node is addressable for STA.

NGC/ AOA Xinetics

# MSFC has developed thin-film-stress monitor for in situ measurements.

- In-situ measurement helped identify a mechanism for reducing the stress in sputtered iridium by 3 orders of magnitude



MSFC/ David Broadway