

## Zooming in on Heinrich layers

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**Abstract.** Theories explaining the origin of the abrupt, massive discharges of ice-rafted detritus (IRD) into the glacial North Atlantic (the Heinrich layers (HLs)) generally point to the Laurentide ice sheet as the sole source of these events, until it was found that the IRDs also originated from Icelandic and European ice sheets [Bond and Lotti, 1995; Snoeckx et al., 1999; Grousset et al., 2000]. This apparent contradiction must be reconciled as it raises fundamental questions about the mechanism(s) of HL origin. We have analyzed two ~12 cm thick HLs in an ultrahigh-resolution mode (1-2 century intervals) in a mid-Atlantic ridge piston core. The  $\delta^{18}\text{O}$  record (*N. pachyderma* left coiling) reveals strong excursions induced by the melting of the icebergs; these excursions are associated with a strong decrease in the amount of planktic foraminifers and with a 3°C cooling of the surface waters. Counts of coarse detrital grains reveal that IRD are deposited according to a typical sequence (1) volcanic glass, (2) quartz and feldspars, (3) detrital carbonate, that implies a chronology in the melting of the different pan-Atlantic ice sheets. Sr and Nd isotopic composition confirm that in both Heinrich layers H1 and H2, "precursor" IRD came from first Europe/Iceland, followed then by Laurentide-derived IRD. An internal cyclicity can be identified: during H1 and H2, about four to six major, abrupt discharges occurred roughly on a century timescale. The  $\delta^{13}\text{C}$  and  $\delta^{15}\text{N}$  records reveal that dominant inputs of continent-derived organic matter are associated with IRD within the HLs, hiding the plankton productivity signal.

### 1. Introduction

It is well known that most of the typical Heinrich events are characterized by massive abrupt discharges of icebergs released by the periodic collapse of the Laurentide ice sheet. The Laurentide perhaps oscillates according to its internal dynamics with a  $\approx 7.200$  year cycle [McAyeal, 1993; Alley and McAyeal, 1994]. It has been shown, however, that other ice sheets, such as the Icelandic, also contributed to these events [Bond and Lotti, 1995]. Moreover, more rapid cycles ( $\approx 1-2$  kyr) seem to have occurred between the typical Heinrich layers (HLs) [Bond and Lotti, 1995]. Elliot et al. [1998] hypothesized that the European ice sheets could have more likely oscillated with a  $\approx 1-2$  ky cycle, while Laurentide surges would have occurred every  $\approx 7.2$  kyr. The origin of these cycles is not fully understood yet.

More recent work has provided a few lines of evidence that the European ice sheets also contributed to the story [Snoeckx et al., 1999] and that IRDs were first released by these European ice sheets, followed by the Laurentide [Grousset et al., 2000; Scourse et al., 2000]. This observation may open a door to a better understanding of the mechanisms that drove

these events. Whether these precursor European events triggered the major Laurentide surges or the behavior of the different pan-Atlantic ice sheets followed a totally independent cyclicity is still debated. One could probably better understand the origin of the Heinrich events by examining in detail the successive steps of their life.

In this work, we propose to better elucidate the internal structure of the three most recent HLs by analyzing a series of various parameters in an extremely high resolution mode (every 5 mm, i.e., every 1-2 centuries) in a core located in the ice-rafted detritus (IRD) belt [Ruddiman, 1977]: SU90-09. We propose to study several paleoclimatic proxies (foraminifer abundance,  $\delta^{18}\text{O}$ , sea surface temperature (SST)), the geochemical and petrological composition of the IRD (Sr and Nd isotopes and coarse-grained petrology), the isotopic composition of the organic matter ( $\delta^{13}\text{C}$  and  $\delta^{15}\text{N}$ ), and the elemental composition of the bulk sediments. This comprehensive approach should allow us to identify the internal structure of the HLs, which could lead to better understanding their origin.

### 2. Material and Methods

#### 2.1. Material

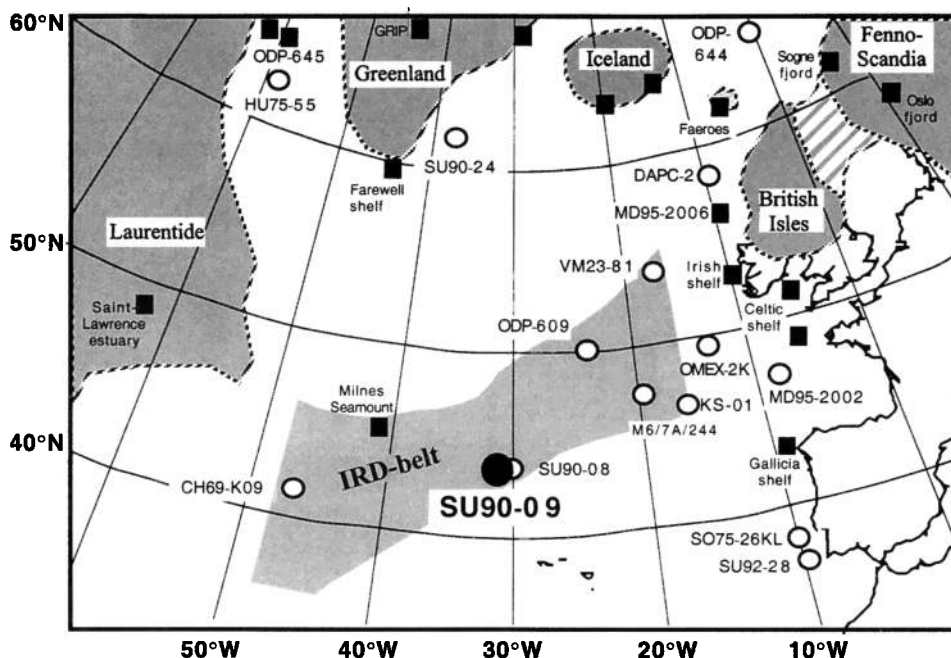
Core SU90-09 was raised from the western side of the Mid-Atlantic Ridge ( $\approx 43^{\circ}05'\text{N}$ ,  $\approx 31^{\circ}05'\text{W}$ , 3.375 m water depth) (Figure 1). This piston core is located on the southern border of the IRD belt [Ruddiman, 1977], 70 km west of the extensively studied core SU90-08 [Grousset et al., 1993; Revel et al., 1996; Cortijo et al., 1997; Vidal et al., 1997; Elliot et al., 1998]. Our present study is focused on the upper 120 cm of SU90-09, which contain the three most recent HLs (H1, H2, and H3).

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**Figure 1** Location of core SU90-09 (solid circle) and of other cores cited in the discussion (open circles). Location of ice sheets is shown by stippled area on land and corresponds to the approximate maximum ice sheet extension at Last Glacial Maximum (LGM). Stippled area in North Atlantic shows position of the ice-rafted detritus (IRD) belt [Ruddiman, 1977]. Possible Source Areas (PSAs) are fingerprinted using the coarse (>63  $\mu\text{m}$ ) fraction of carbonate-free sediments recovered from fjord, estuarine, or continental shelf areas (solid squares).

## 2.2. Methods

Whole-core, low-field magnetic susceptibility was measured with a pass-through Bartington magnetometer. This instrument integrates the sediment section within a small coil (4 cm diameter) to record the changes in magnetic susceptibility along a continuous subsampling of core SU90-09, made using a transparent, PVC, U-shaped channel. Resolution after deconvolution of the signal is of the order of  $\approx 2$  cm. The grey level reflectance of core SU90-09 has been analyzed following Bond *et al.* [1992]. Sedimentary sequences, lithologic facies, and the geometry of physical and biological structures were observed using a digital SCOPIX X-ray device, coupling a high-energy brightness amplifier and a digital camera [Migeon *et al.*, 1999]. Smear slides were fabricated after induration with a polyester resin in order to identify bioturbation, if any. Sediment slices were then sampled at 1 cm intervals and at 0.5 cm intervals through the Heinrich layers, using an electro-osmotic cutter.

Stratigraphy was provided by  $\delta^{18}\text{O}$  isotopic analysis performed at 1 cm intervals on the coarse fraction (150–250  $\mu\text{m}$ ) of planktonic foraminifera *Neogloboquadrina pachyderma* (left coiling (lc), and at 0.5 cm intervals through the Heinrich layers, from 30 to 118 cm. Nine  $^{14}\text{C}$  accelerator mass spectrometry (AMS) ages (B.P.) were obtained on foraminifera *N. pachyderma* (lc) (>150  $\mu\text{m}$ ) and corrected into conventional radiocarbon ages by taking into account  $^{13}\text{C}/^{12}\text{C}$  ratios. These  $^{14}\text{C}$  ages were finally corrected for a 400 year reservoir age [Bard, 1988]. Error bars are  $\leq 100$  years. The  $^{14}\text{C}$  ages have also been transformed into calendar ages, using regression equations derived from Bard *et al.* [1988, 1998] and Stuiver *et al.* [1998] (see Figures 2 and 3).

Carbon and nitrogen stable isotopes ( $\delta^{13}\text{C}$  and  $\delta^{15}\text{N}$ ) were measured by elemental analyzer coupled with an isotope ratio mass spectrometer (EA-IRMS) on bulk organic matter of <50  $\mu\text{m}$  size fractions. Calcite and dolomite removal was previously performed using a careful HCl dissolution procedure extensively described by S. Huon *et al.* (Sources of fine organic matter in the North Atlantic Heinrich layers:  $\delta^{13}\text{C}$  and  $\delta^{15}\text{N}$  isotope tracers, submitted to *Geochimica Cosmochimica Acta*, 2001). The  $\delta^{13}\text{C}$  and  $\delta^{15}\text{N}$  are reported in per mil and refer to Pee dee Belemnite (PDB) and air as standards [Coplen *et al.*, 1983]. Analytical precision is better than  $\pm 0.1\text{‰}$  and  $\pm 0.2\text{‰}$  for  $\delta^{13}\text{C}$  and  $\delta^{15}\text{N}$ , respectively. Bulk organic carbon (TOC) and total nitrogen contents (TN) are reported in  $\text{mg. g}^{-1}$  of sediment (equivalent to per mil) and were determined with a  $\pm 0.1\text{‰}$  constant uncertainty. Owing to the low nitrogen content of the samples (0.1–0.7‰), only a fair precision was obtained with calculated TOC/TN ratios. Therefore they were not used in this study.

Samples were sieved in order to analyze the coarse (>150  $\mu\text{m}$ ) fraction. This grain-size fraction was selected according to Bond *et al.* [1992]. Particles were counted under the binocular microscope, distinguishing quartz, feldspars, micas, rock fragments, detrital carbonate, and volcanic glass. Counts of foraminifera species allowed us to reconstruct sea surface temperature (SST) for both winter and summer seasons, using the modern analog technique [Prell, 1985]. Mean error bars on SST evaluations are about  $\pm 1^\circ\text{C}$ .

Carbonate was dissolved using 0.6 N HCl, and the carbonate-free coarse (>63  $\mu\text{m}$ ) fraction was crushed in a grinder. This grain-size fraction (>63  $\mu\text{m}$ ) was chosen in order to be consistent with previous data obtained on the same

fraction [Snoeckx *et al.*, 1999]. About 10 mg aliquots were dissolved with a mixture of teflon-distilled [HF + HClO<sub>4</sub> + HNO<sub>3</sub>]. Sr and Nd were chemically separated through ionic chromatographic columns in a "class-1000" clean laboratory. We followed chemical and mass spectrometer techniques previously described [Grousset *et al.*, 1998]. Thermo ionisation mass spectrometry analysis of Sr and Nd was done on a Finnigan MAT 261 mass spectrometer. The measured <sup>87</sup>Sr/<sup>86</sup>Sr and <sup>143</sup>Nd/<sup>144</sup>Nd ratios were corrected for mass fractionation by normalizing to <sup>86</sup>Sr/<sup>88</sup>Sr=0.1194 and <sup>146</sup>Nd/<sup>144</sup>Nd=0.7219, respectively. For convenience, <sup>143</sup>Nd/<sup>144</sup>Nd ratios are normalized and reported as follows:  $\epsilon_{Nd(o)} = ((^{143}Nd/^{144}Nd/0.512636) - 1) \times 10^4$ , in which 0.512636 is the chondritic uniform reservoir value [Jacobsen and Wasserburg, 1980]. Sr standard NIST 987 was measured seven times with an average <sup>87</sup>Sr/<sup>86</sup>Sr = 0.710242 ( $\pm 0.000017$ ) versus the certified value of 0.710250, and the French Nd standard "Rennes" was analyzed five times with an average <sup>143</sup>Nd/<sup>144</sup>Nd = 0.511974 ( $\pm 0.000014$ ) versus the recommended value of 0.511962. Blanks were measured (Sr < 10<sup>-9</sup> g and Nd < 2.10<sup>-10</sup> g) and are considered to be negligible in all cases.

### 3. Results

A clear relationship exists between grey scale and magnetic susceptibility records in both cores, at least in the depth distribution of the peaks (Figures 2-3). Both records in core SU90-09 are similar enough to the SU90-08 records of Grousset *et al.* [1993] to infer the position of most of the Heinrich layers (H1-H6) in SU90-09, with only H3 and H6 not well represented. Using the 19 AMS <sup>14</sup>C dates obtained on both cores, the timescale of the core can be then derived. Ages of the H1 and H2 boundaries (Figure 3) are consistent with ages obtained in other cores from North Atlantic [Elliot *et al.*, 1999; Grousset *et al.*, 2000; Scourse *et al.*, 2000]. The average accumulation rate is ~4.3 cm. kyr<sup>-1</sup> but increases up to 8 cm. kyr<sup>-1</sup> during H1. AMS <sup>14</sup>C dates obtained for this core yield the duration of these events: ~1.4 kyr for H1 and ~2.1 kyr for H2. These durations imply accumulation rates of 15.3 and 11.0 g. cm<sup>-2</sup>. kyr<sup>-1</sup> for H1 and H2 respectively. Such fluxes are consistent with those calculated using <sup>230</sup>Th activity in the neighbor core CHN82-31-11PC recovered 80 km away from SU90-09: 19.0 and 11.0 g. cm<sup>-2</sup>. kyr<sup>-1</sup> for H1 and H2, respectively [François and Bacon, 1994].

X-radiographs of thin sediment slices do not reveal any structure, such as hiatus, erosion surface, or any physical perturbation. HLs are clearly identified as dark zones owing to increased IRD and water content decrease; they are characterized by abundant, coarse (millimeter-to-centimeter) particle deposits, alternating with thin (3-5 mm) layers of "regular" glacial sediments, comparable to the ambient sediments between the HLs. Bioturbation (worm burrows) is only visible in the ambient glacial deposits located between the events, as already observed in core KS-01 [Auffret *et al.*, 1996] (see Figure 1 for location). This evidence of unbioturbated, laminated deposits alternating with badly sorted IRD is confirmed by microscopic examination of the HL lithofacies, realized on smear-slides under cross-nicols.

The high-resolution study has been limited to the three last HLs. In the following figures 3, 4, 5, and 7, however, depths

will be expressed in centimeters instead of kiloyears, in order to avoid a strong compression of the records within the HLs due to the high accumulation rates characterizing these events.

#### 3.1. Planktic Foraminifers

HLs are marked by a darkening of the sediment color (Figure 3), reflecting a decrease in the proportion of carbonate plankton. In the HLs, the number of foraminifers is 1 order of magnitude smaller than between these events (Table 1). This decrease is associated with a change in the species proportions. During the HLs, *N. pachyderma* (lc) becomes dominant (up to 80%), and their  $\delta^{18}O$  composition becomes lighter, displaying brief excursions of ~0.6-1.0‰, that parallel similar anomalies already reported for the neighboring core SU90-08 [Cortijo *et al.*, 1997]. These excursions are maximal during the upper part of the HLs. They are associated with a sudden decrease of 3°C of the mean SST, shifting them from 11°C down to 8°C during the summers (Figure 3) and from 6°C down to 3°C during the winters. When combining both SST and  $\delta^{18}O$ , it can be estimated that paleosalinity (SSS) decreased by ~2‰ during both H1 and H2.

#### 3.2. Lithic Grains

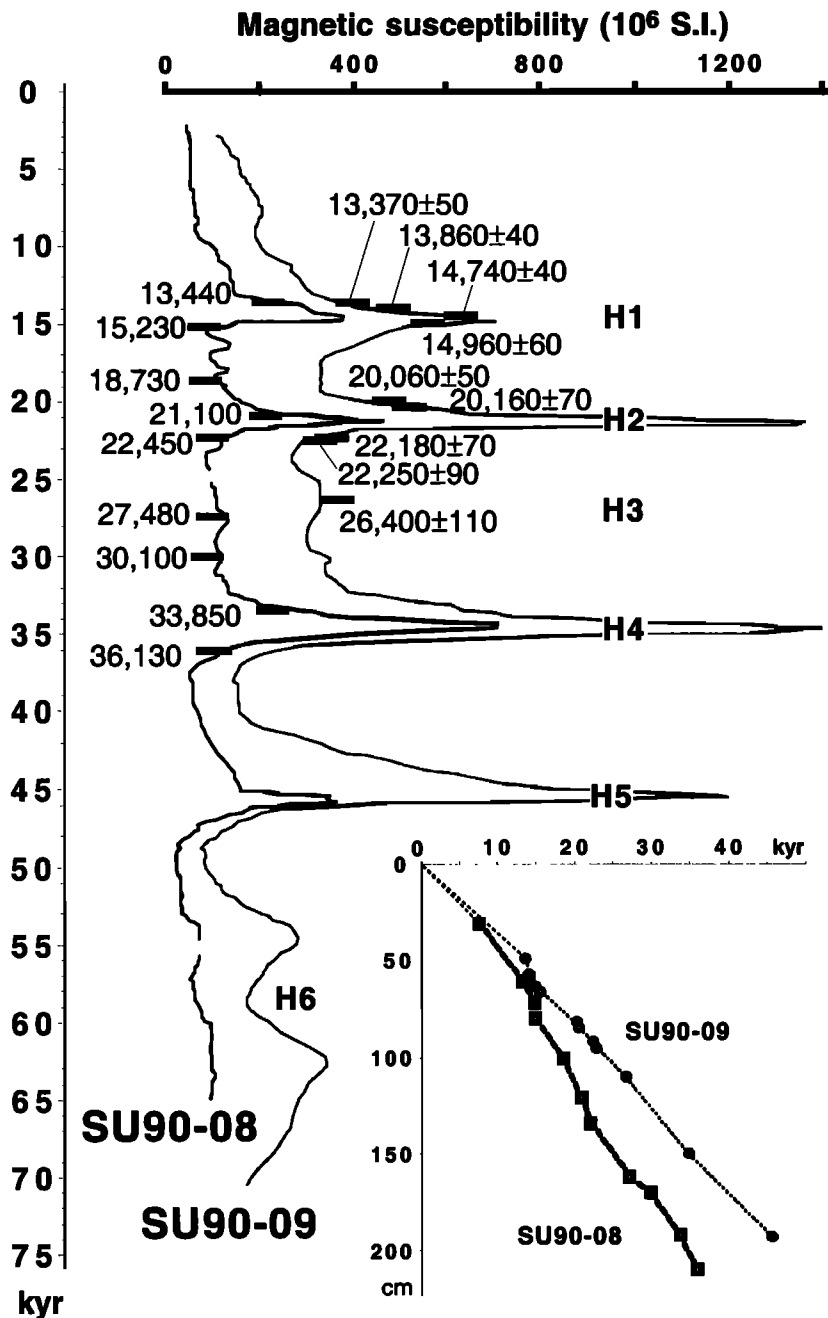
Lithic grains were counted on the coarse (>150  $\mu$ m) fraction under a binocular microscope. Their abundance increases by a factor of 10 during the HLs. A few minerals were distinguished and counted: quartz, feldspars, micas, rock fragments, detrital carbonate, and volcanic glass (Table 2). The most representative are presented in Figure 4 (detrital carbonate, volcanic glass and quartz). All these minerals display strong increases during the HLs, although their maximum occurrence is not always simultaneous. Volcanic grains increase first, at the beginning of the HLs, followed by quartz, and finally by detrital carbonate. This pattern is visible in both H1 and H2. On the contrary, H3 does not display any structure. Within the HLs the mineral abundance is cyclic, displaying centennial-like subevents.

Secondary lithic peaks observed between the HLs in some other cores were attributed to Dansgaard-Oeschger cycles [Bond and Lotti, 1995; Elliot *et al.*, 1998; van Kreveland *et al.*, 2000]. Such secondary events are hardly visible in core SU90-09, probably because this core is located too far south.

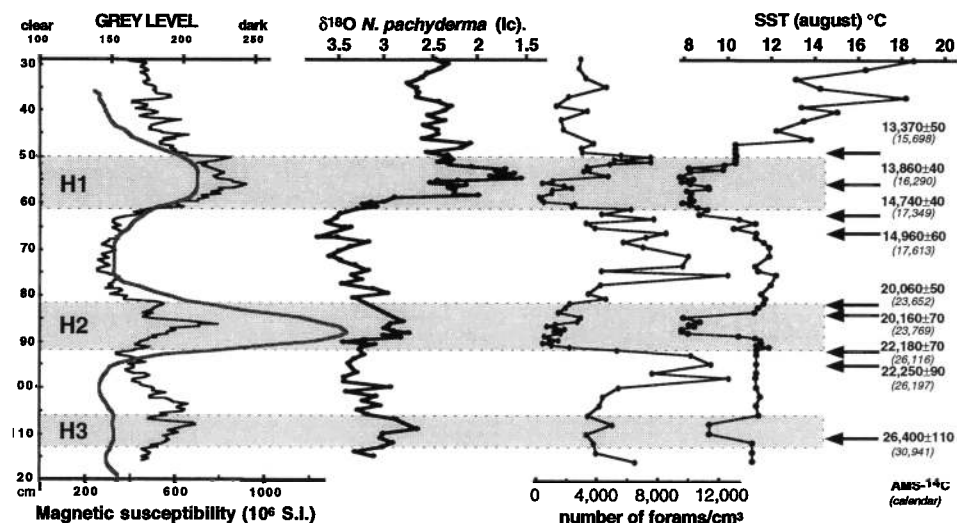
#### 3.3. Isotopic (Sr-Nd) Composition of IRD

The down core variation of the Sr and Nd isotopic composition of the carbonate-free coarse (>63  $\mu$ m) fraction is displayed in Figure 5 and Table 3. Within the Heinrich layer, IRD are characterized by radiogenic Sr ( $0.72 < ^{87}Sr/^{86}Sr < 0.73$ ) and unradiogenic Nd ( $-45 < \epsilon_{Nd(o)} < -20$ ). Within these intervals the isotopic values are characterized by a high-frequency variability. These layers are preceded by a narrow (3-4 cm thick) event displaying much less radiogenic Sr ( $0.715 < ^{87}Sr/^{86}Sr < 0.718$ ) and much more radiogenic Nd ( $-16 < \epsilon_{Nd(o)} < -6$ ). Very similar patterns are observed for both isotopes in both H1 and H2.

In Figure 6a we have plotted <sup>87</sup>Sr/<sup>86</sup>Sr ratios against  $\epsilon_{Nd(o)}$  for all the possible source areas (PSA) from the



**Figure 2:** Cores SU90-08 and SU90-09 are located 70 km from each other, at the same latitude, and can be easily compared. A well constrained timescale has been previously established for core SU90-08 using both the  $\delta^{18}\text{O}$  record of planktic foraminifera [Grousset et al., 1993] and 10 accelerator mass spectrometry (AMS)  $^{14}\text{C}$  dates. The timescale of core SU90-09 has been derived from the other one by matching the magnetic susceptibility records of both cores, using the Analyseries software [Paillard et al., 1996]. "Typical" Heinrich layers (H1, H2, H4, and H5) are marked by strong, sharp, magnetic susceptibility anomalies, whereas "atypical" Heinrich layers (H3 and H6) are hardly visible. Nine  $^{14}\text{C}$  accelerator mass spectrometry ages obtained on core SU90-09 help to further constrain the timescale. These ages (B.P.) were obtained on foraminifera *N. pachyderma* (left coiling (lc)) (150-250  $\mu\text{m}$ ) and corrected into conventional radiocarbon ages by taking into account  $^{13}\text{C}/^{12}\text{C}$  ratios. These  $^{14}\text{C}$  ages were finally corrected for a 400 year reservoir age [Bard, 1988]. Error bars are < 100 years.



**Figure 3.** Correlations of magnetic susceptibility and grey level of bulk sediment in core SU90-09, along with the  $\delta^{18}\text{O}$  isotopic record, abundance of planktic foraminifers, and summer sea surface temperature (SST) ( $^{\circ}\text{C}$ ) calculated using the modern analog technique method [Prell, 1985], between 30 and 120 cm. The  $\delta^{18}\text{O}$  (normalized to Pee dee belemnite standard) isotopic record is from planktic foraminifera *N. pachyderma* (lc). Location of Heinrich layers H1, H2, and H3 is shown by stippled areas. AMS  $^{14}\text{C}$  ages are reported along with calendar ages calculated according to Bard [1998].

continents surrounding the northern Atlantic (fjord, estuarine, and continental shelf samples). Because many of the data from the literature deal with a particular rock sample or geological unit, we consider these samples to be less representative of an entire potential source area than are estuarine sediments or adjacent marine sediments, which are derived from a large area. In order to avoid any problem introduced by a possible grain-size fractionation, we have analyzed the same carbonate-free coarse ( $>63\ \mu\text{m}$ ) fraction in both the PSA and the core SU90-09 samples. It appears that two main domains can be distinguished: first, "field 1" links the volcanic province (Iceland, Faeroes, or Thulea) to the European province (Scandinavia, British Isles, etc); second, "field 2" links the same volcanic unit to the Canada/Greenland province.

In Figure 6b we have plotted the same ratios for all the samples from cores SU90-09. Most of the samples (open circles) fall within field 2 (the Greenland-Canada domain), reflecting the dominance of this source area. A few samples (open crossed circles), corresponding precisely to the volcanic-rich precursor events described above, fall within field 1 (the hyperbolic envelope containing the European domain), and they are close to the Icelandic end-member. We thus consider the Icelandic and the European provinces as the source areas for these specific samples.

### 3.4. The $\delta^{13}\text{C}$ and $\delta^{15}\text{N}$ Composition of Fine-Grained ( $<50\ \mu\text{m}$ ) Organic Matter

The  $\delta^{13}\text{C}$  and  $\delta^{15}\text{N}$  of organic matter decrease to minimum values of  $-27.9$  and  $0.3\text{‰}$  during the phases of enhanced ice rafting corresponding to H1 and H2 (Table 4 and Figure 7). These values contrast sharply with those measured for ambient glacial sediments which average  $-23.5 \pm 0.7$  and  $4.7 \pm 0.9\text{‰}$  for  $\delta^{13}\text{C}$  and  $\delta^{15}\text{N}$ , respectively ( $\pm 1\sigma$ ). No significant deviation from the average values calculated for ambient organic matter is

shown by the samples corresponding to H3 (mean values:  $\delta^{13}\text{C} = -23.4 \pm 0.7\text{‰}$  and  $\delta^{15}\text{N} = 5.0 \pm 0.4\text{‰}$ ). During H1 and H2,  $\delta^{15}\text{N}$  display high-frequency variations (up to 4 ‰), which are not observed for  $\delta^{13}\text{C}$ , and [TOC] and [TN] contents also display a marked contrast from ambient sedimentary organic matter (mean values  $\pm 1\sigma$ : [TOC] =  $2.95 \pm 0.4\ \text{mg}\cdot\text{g}^{-1}$  and [TN] =  $0.50 \pm 0.1\ \text{mg}\cdot\text{g}^{-1}$ ) with maximum decreases to 1.8 and  $0.1\ \text{mg}\cdot\text{g}^{-1}$  for TOC and TN, respectively.

## 4. Discussion

Although it seems obvious now that Heinrich events strongly perturbed the thermohaline circulation and, consequently, the climate [Broecker, 1997], the origin of these abrupt climatic events is not fully understood yet. Since the pioneer papers revealing the Heinrich events [Heinrich, 1988; Bond et al., 1992], it has been considered that they were mostly related to surges from the Laurentide ice sheet that occurred during the last glacial period [Broecker, 1994]. Andrews [1998] still considered that the Hudson Strait was the prime source for sediment delivery to the Atlantic during the HLs but clearly stated that abrupt changes in the delivery of IRD were multisourced. It is clear that "typical" HLs (H1, H2, H4, and H5) are dominated by Laurentide-derived IRD [Gwiazda et al., 1996a; Hemming et al., 1998], whereas "atypical" HLs (H3 and H6) reflect more complex pan-Atlantic IRD origins [Huon and Jantschik, 1993; Grousset et al., 1993; Gwiazda et al., 1996b; Snoeckx et al., 1999]. Recently, more detailed studies demonstrated that HLs were composed of sublayers originating from different sources. Thus a "precursor" event enriched in volcanic IRD has been identified in some cores from the IRD belt, including ODP-609 and VM 23-81 [Bond and Lotti, 1995; Bond et al., 1999], and in the Irminger Basin, in core SU90-24 [Elliot et al., 1998; Bond et al., 1999] (Figure 1). Along the European margin, similar precursors

**Table 1.** Planktic Foraminifer Characteristics: Abundance,  $\delta^{18}\text{O}$  of *N. pachyderma* (lc) and Summer SST<sup>a</sup>

Depth cm	<i>N. pachyderma</i> (lc), $\delta^{18}\text{O}(\text{‰})$ PDB	Planktic Forams per gram	SST (August), °C
30	2.45	2847	18.5
31	2.33		
32	2.41	2748	16.3
33	2.54		
34	2.60	3196	13.1
35	2.74		
36	2.66	4538	14.2
37	2.71		
38	2.66	2039	18.2
39	2.43		
40	2.30	1253	13.4
41	2.36	3282	15.0
42	2.53		
43	2.36	1572	13.5
44	2.58		
45	2.45	1734	12.2
46	2.44		13.8
47	2.60	3707	10.3
48	2.09	2899	10.3
49	2.18	2963	10.3
50	2.54	5492	10.3
50.5	2.33	7422	10.3
51	2.31	5038	10.3
51.5	2.33	7422	10.3
52	2.38	4757	9.8
52.5	2.31	3264	8.2
53	2.08	3450	9.8
53.5	1.72	3020	8.2
54	1.79	4639	7.8
54.5	1.62		8.4
55	1.87	1031	7.9
55.5	1.55	335	8.2
56	2.33	1754	9.1
56.5	2.51	2250	9.1
57	2.14	992	8.1
57.5	2.32	942	8.3
58	2.25	128	8.2
58.5	2.30	303	8.4
59	2.01	452	7.9
59.5	2.89	2432	8.2
60		2319	8.6
60.5	3.15	6153	9.1
61	3.23	4233	8.7
61.5	3.09	7654	10.5
62	3.29	3233	11.3
63	3.46	3798	10.3
64	3.51	8461	11.3
65	3.61	7136	11.3
66	3.39	5630	11.7
67		6936	11.9
68	3.60	9913	11.9
69	3.16	9529	11.2
70	3.36	4241	11.4
71	3.48		

**Table 1** (continued)

Depth cm	<i>N. pachyderma</i> (lc), $\delta^{18}\text{O}(\text{‰})$ PDB	Planktic Forams per gram	SST (August), °C
72	3.58	12456	12.2
73	3.46		
74	3.31	4162	12.0
75	3.17	3354	11.5
76	3.33	4540	11.7
77	3.24		
78	3.48	2142	11.6
79	3.11		
80	2.96	1906	11.4
81	3.33	1418	11.2
82		2899	8.0
83		2664	8.7
84		1238	8.2
85		653	8.5
86	2.79	1817	7.9
86.5	2.93	1131	7.9
87	2.95	1575	8.2
87.5	2.98	522	10.5
88	2.91	803	11.3
88.5	2.74	1420	11.5
89	3.00	408	11.5
89.5	2.82	983	11.3
90	3.20	2165	11.9
90.5	3.43	5253	11.3
91	3.09	10080	11.3
91.5	3.24	11388	11.3
92	3.25	7563	11.3
93	3.14	12534	11.3
94	3.31		
95	3.37	5355	11.3
96			
97	3.28	4358	11.5
98	3.41	4145	11.3
99	3.41		
100	2.93	3361	11.4
101	3.39		
102	3.08	4957	9.2
103	3.21		
104	3.05	3283	9.2
105	3.29		
106	3.17	3769	11.1
107	2.87		
108		3923	11.1
109	2.64		
110	3.00	6481	11.1
111	3.02		
112	2.90	4393	
113	3.01		
114	3.30	6709	
115	3.10		
116		5128	

<sup>a</sup> SST (°C) are calculated using the modern analog technique [Prell, 1985]. PDB is Pee dee belemnite

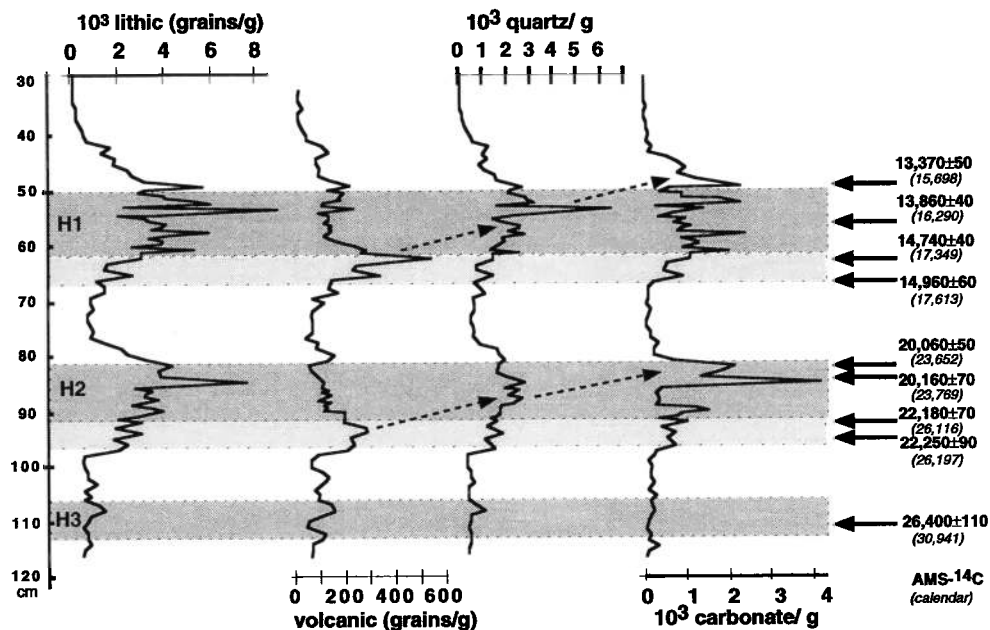
**Table 2.** Abundance of Lithic Grains in the Carbonate-Free Coarse (>150 $\mu$ m) Fraction in Core SU90-09

Depth, cm	Lithic, grains.g-1	Volcanic, grains.g-1	Quartz, grains.g-1	Carbonate, grains.g-1
30	81	4	47	9
31	102	4	55	17
32	119	13	77	9
33	98	13	51	4
34	141	4	85	21
35	115	9	55	13
36	222	26	111	30
37	201	4	145	21
38	226	9	171	13
39	329	13	213	34
40	448	26	303	38
41	687	43	469	77
42	828	47	631	68
43	1826	102	1267	183
44	1510	132	1050	102
45	2007	79	1146	523
46	1888	77	848	801
47	2472	94	1136	1001
48	2767	90	1523	794
49	3236	128	1594	1240
50	5857	209	2716	2362
50.5	3392	159	2182	635
51	3066	123	2081	396
51.5	4191	189	2583	896
52	4273	168	2823	891
52.5	5481	172	2829	1833
53	6137	110	3166	2357
53.5	2381	98	1642	282
54	9095	222	6425	1418
55	3641	83	2379	704
55.5	2231	134	1470	347
56	3299	129	1813	1031
56.5	3538	106	2250	743
57	4284	140	2604	1155
57.5	4119	131	2562	936
58	3604	137	2114	813
58.5	6012	128	2743	2433
59	4273	111	1943	1485
59.5	3504	127	1889	935
60	4122	178	1930	1324
60.5	4037	234	1811	1339
61	2922	258	1360	969
61.5	5518	272	2563	2064
62	3262	261	1434	1165
63	3581	533	1400	1107
64	1746	235	728	555
65	1675	227	783	461
66	3008	335	1408	982
67	1212	145	691	235
68	1540	128	964	119
69	1621	166	1088	111
70	862	64	580	85
71	913	107	640	68



**Table 2** (continued)

Depth, cm	Lithic, grains.g-1	Volcanic, grains.g-1	Quartz, grains.g-1	Carbonate, grains.g-1
72	1033	64	661	128
73	832	55	542	73
74	789	64	525	51
75	853	64	572	47
76	994	68	674	141
77	909	34	606	141
78	1579	90	969	316
79	2270	77	1630	239
80	2566	154	1692	262
81	3375	124	1958	708
82	4341	50	1703	2224
83	3826	59	1598	1894
84	3980	97	2043	1392
85	7681	111	2782	4326
86	2869	99	1938	416
86.5	3657	130	2517	354
87	3330	104	2402	265
87.5	3321	137	2127	348
88	3819	89	2682	370
88.5	3432	138	2298	346
89	2785	109	1817	280
89.5	3644	121	1790	1235
90	4141	108	1768	1576
90.5	3555	196	1805	968
91	3041	199	1566	813
91.5	2242	194	1376	372
92	3203	186	1579	1045
93	2173	273	1186	425
94	3273	280	1852	749
95	2042	214	1056	559
96	2693	229	1339	760
97	2301	202	1535	268
98	720	85	301	231
99	549	58	332	76
100	695	61	348	204
101	828	99	479	155
102	880	131	508	144
103	1150	118	686	227
104	647	74	412	98
105	902	93	435	269
106	746	84	449	126
107	1281	136	777	205
108	1641	150	1064	248
109	962	98	544	139
110	792	50	536	77
111	565	46	409	43
112	851	114	491	161
113	831	104	537	104
114	955	67	469	316
115	640	62	407	86
116	596	55	418	40



**Figure 4.** Abundances of lithic grains, detrital carbonate, volcanic glass, and quartz grains in coarse sediments (>150  $\mu\text{m}$ ) from core SU90-09. Location of Heinrich layers H1, H2, and H3 is shown by stippled areas (light grey band for the European precursors and dark grey band for the Laurentide-derived steps). Arrows emphasize the sequential deposition of (1) volcanic glass, (2) quartz, and (3) detrital carbonate.

made of European continent-derived IRD were recently identified in cores MD95-2002 and SU92-28 [Grousset *et al.*, 2000] (Figure 1). If it can be demonstrated that a sequence of IRD events from different origins can be identified in all HLs, it will tell us more about the timing of the surges (and/or intense iceberg releases) of the different pan-Atlantic ice sheets, and a better understanding of the HL origins could be achieved.

The very high resolution approach that has been conducted on core SU90-09 (every 1-5 centuries) allows us to clearly identify an internal structure (three major steps) of the HLs that is discussed below. Moreover, another mode of variability is superimposed (centennial cycles?) onto this layered structure. Indeed, extremely brief and abrupt events oscillating every 5-10 mm can be observed on different records: grain counts,  $\epsilon_{\text{Nd}}(0)$ ,  $\delta^{15}\text{N}$  (Figures 4, 5, and 7). These two modes of variability are addressed now.

#### 4.1. Sequential Deposition

A clear petrographic sequence can be identified when looking at the vertical distribution of IRD through the two last HLs (Figure 4). The first step of the IRD increase visible at the beginning of H1 and H2 is characterized by a peak of volcanic grains (ashes and basaltic grains), peaking around 22.2 and 14.9 kyr ( $^{14}\text{C}$ ), respectively. In this core, this first step lasted 1-2 centuries (Figure 4). The Sr and Nd isotopic composition of these grains was measured in volcanic debris picked from H1. This composition reflects the volcanic source from where they are derived. The following values have been obtained:  $\epsilon_{\text{Nd}}(0) = +7.4$  and  $^{87}\text{Sr}/^{86}\text{Sr} = 0.703657$  (solid circle in Figure 6a). Such values fit perfectly the composition of the coarse (>63  $\mu\text{m}$ ) Icelandic basalt debris and volcanic

glasses (mean values:  $\epsilon_{\text{Nd}}(0) = +7.6$  and  $^{87}\text{Sr}/^{86}\text{Sr} = 0.70345$ ). The down core distribution of both  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $\epsilon_{\text{Nd}}(0)$  reveals that in both H1 and H2 these volcanic-rich precursor events display the same pattern, characterized by a more radiogenic Nd and less radiogenic Sr (Figure 5). Such an anomaly could be explained by an increase in the proportion of Icelandic glass within the Laurentide-derived IRD. In this case, the isotopic composition of the new mixture would shift within field 2 along hyperbolae linking the Greenland/Canada and the Iceland/Faeroe end-members (Figure 6b). In fact, the isotopic composition of these IRD (open crossed circles) is shifted toward another field of the Sr-Nd plot (field 1), i.e., within the "European-Iceland" isotopic domain, implying that they are made of a complex mixture of European and Icelandic IRD. This is in agreement with the early occurrence of European IRD recognized along the European margins [Grousset *et al.*, 2000]. During H1 it is known that the British ice sheet clearly participated in IRD discharges [McCabe and Clark, 1998]. Moreover, European coccoliths were found as far west as the Labrador Sea [Rahman, 1995], and models simulate a possible input of European icebergs in this region [Schäfer-Neth and Statteger, 1997]. Such a structure is also visible in both cores ODP-609 and VM 23-81 [Bond *et al.*, 1999] and can be considered as the volcanic-rich precursor event. During this first step, European icebergs might have followed the same kind of counterclockwise trajectories as the one suggested for atypical events H3 and H6 [Grousset *et al.*, 1993].

The second step of the structure is characterized by a strong increase in quartz grains (Figure 4), along with a nonnegligible proportion of various crust-derived minerals (feldspars, micas, and igneous rock debris) and, to a lesser extent, detrital carbonates. It corresponds to the paroxysmal

**Table 3.** Sr and Nd Isotopic Composition of the Carbonate-Free Coarse (>63  $\mu\text{m}$ ) Fraction in Core SU90-09 and in the Pan-Atlantic Possible Source Areas (PSAs)

Core Depth, cm	$^{87}\text{Sr}/^{86}\text{Sr}$	$\pm 2\sigma$	$^{143}\text{Nd}/^{144}\text{Nd}$	$\pm 2\sigma$	$\epsilon_{\text{Nd}}(0)$
45	0.723949	6	0.511446	7	-23.2
46	0.723734	7	0.511571	5	-20.8
47	0.723081	7	0.511310	7	-25.9
48	0.723574	8	0.511302	9	-26.0
49	0.722207	22	0.511623	6	-19.8
50	0.721503	14	0.511643	5	-19.4
50.5	0.717195	13	0.511493	9	-22.3
51	0.725139	9	0.511543	10	-21.3
51.5	0.722005	12	0.511470	5	-22.7
52	0.724049	7	0.511404	9	-24.0
52.5	0.723493	7	0.511445	7	-23.2
53	0.722499	7	0.511419	8	-23.7
53.5	0.719022	14	0.511190	8	-28.2
54	0.725038	8	0.511406	9	-24.0
54.5	0.723068	9	0.511244	6	-27.1
55	0.718776	26	0.511362	10	-24.9
55.5	0.722482	13	0.511149	6	-29.0
56			0.511346	7	-25.2
56.5	0.723528	6	0.511068	18	-30.6
57	0.726070	10	0.510541	12	-40.9
57.5	0.725420	7	0.511353	20	-25.0
58	0.726622	11	0.511308	34	-25.9
58.5	0.726567	10	0.511322	14	-25.6
59	0.722638	10	0.510686	6	-38.0
59.5	0.720234	52	0.511367	21	-24.8
60	0.729044	20	0.511469	5	-22.8
61	0.720395	13	0.511615	9	-19.9
61.5	0.722103	10	0.511839	7	-15.6
62	0.716033	19	0.512259	3	-7.4
63	0.718108	14	0.512341	5	-5.8
64	0.717991	6	0.512281	5	-6.9
65	0.716623	9	0.512331	5	-6.0
66	0.719807	7	0.511700	5	-18.3
82	0.721446	6	0.511216	4	-27.7
83	0.724295	9	0.511347	5	-25.1
84			0.511103	3	-29.9
85			0.511159	16	-28.8
86	0.725306	7	0.510808	16	-35.7
86.5	0.724431	7	0.511329	13	-25.5
87	0.722513	7	0.511269	15	-26.7
87.5	0.728609	7	0.511074	13	-30.5
88	0.725773	6	0.510791	17	-36.0
88.5	0.727319	6	0.511193	8	-28.1
89	0.723157	6	0.510436	5	-42.9
89.5	0.723311	6	0.511176	9	-28.5
90	0.723020	8	0.511161	8	-28.8
90.5	0.722538	7	0.511165	8	-28.7
91	0.719380	8	0.511626	5	-19.7
91.5	0.720701	7	0.511969	15	-13.0
92	0.720522	7	0.511855	6	-15.2
93	0.717282	7	0.512085	6	-10.8
94	0.718087	6	0.511783	5	-16.6
95	0.725986	9	0.511671	4	-18.8

**Table 3.** (continued)

Possible Source Areas <sup>a</sup>	<sup>87</sup> Sr/ <sup>86</sup> Sr	±2σ	<sup>143</sup> Nd/ <sup>144</sup> Nd	±2σ	ε <sub>Nd</sub> (o)
Fram Strait	0.718180	37	0.511980	16	-12.8
Fram Strait	0.721620	21	0.512000	15	-12.4
Vøring Plateau	0.722900		0.511903	12	-14.3
Vøring Plateau	0.723500		0.512057	11	-11.3
Oslofjord	0.725039	6	0.512077	8	-10.9
Sogrefjord	0.735614	6	0.511923	3	-13.9
Rockall Trough	0.724352	6	0.512062	14	-11.2
Rockall Trough	0.724012	8	0.512001	6	-12.4
Irish Shelf	0.721308	7	0.511998	5	-12.4
Irish Channel	0.729787	7	0.512011	8	-12.2
Celtic Sea	0.720651	30	0.511965	18	-13.1
British Isles (mean)	0.732519		0.511966		-13.1
Bay of Biscay (mean)	0.730650		0.512030		-11.9
SW Portuguese shelf	0.723874	10	0.512054	6	-11.4
NW Portuguese shelf	0.723285	9	0.511919	11	-14.0
Iceland (mean)	0.703450		0.513924		7.6
Azores	0.704172	23	0.512890	16	4.9
Faeroe Islands	0.704637	11	0.512905	17	5.3
Faeroe Islands	0.705334	29	0.512966	28	6.2
Faeroe Islands	0.703887	10	0.513037	53	7.6
Baffin Bay (b)	0.742870				-33.0
Baffin Bay (b)	0.731560				-27.4
Saint Lawrence	0.720113	8	0.511738	5	-17.5
Milnes Seamount	0.728360	12	0.510991	11	-32.1
GRIP bedrock	0.728800	11	0.510403	20	-43.5
GRIP bedrock	0.731670	7	0.510739	12	-37.0
East Greenland shelf	0.713495	8	0.512094	9	-10.6
Farewell Cape	0.711110	10	0.511483	12	-22.5

a Location given on Figure 1 as solid squares

b Goldstein and Jacobsen [1988]

input of IRD during the Heinrich events, as reflected by the sudden dilution of the planktic foraminifers, the lightening of  $\delta^{18}\text{O}$  of *N. pachyderma* (lc), a 3°C surface water cooling and a ≈2‰ SSS decrease, comparable to the estimation obtained for H4 in the twin core SU90-08 [Cortijo *et al.*, 1997].

The Sr and Nd isotopic composition was measured in quartz grains picked from H1 under the binocular microscope. This composition must reflect the continental source from when they were eroded by ice sheets. Their isotopic composition ( $\epsilon_{\text{Nd}}(\text{o}) = -26$  and  $^{87}\text{Sr}/^{86}\text{Sr} = 0.721341$ ) falls right in the middle of the Greenland/Canada field (solid square in Figure 6b), and is very similar to most of the H1 and H2 samples (open circles on Figure 6b). During this second step the isotopic evolution with time is significantly variable (Figure 5), with  $\epsilon_{\text{Nd}}(\text{o})$  fluctuating between -20 and -45. Despite the presence of negligible Paleozoic outcrops located in northwestern Scotland, the majority of the British Isles outcrops display much more radiogenic  $\epsilon_{\text{Nd}}(\text{o})$  ranging between -9 and -14 [Miller and O'Nions, 1984; this study]. These latter values are consistent with the composition of the coarse (>63 μm) fraction of the British Isles continental shelf sediments (Table 3 and Figure 6a). Moreover, the Fennoscandian ice sheet eroded the old crustal Baltic shield. Quite unradiogenic  $\epsilon_{\text{Nd}}(\text{o})$  values were reported from the Baltic region [Miller and O'Nions, 1985], but they were obtained on bulk rocks. The isotopic composition of the coarse (>63 μm) fraction of the Scandinavian shelf/fjord

sediments (Table 3 and Figure 6a) reveals, indeed, much more radiogenic Nd ( $-14 < \epsilon_{\text{Nd}}(\text{o}) < -11$ ), similar to the British Isles signature. Thus a potential contribution of both the Fennoscandian and British Isles ice sheets would not explain the unradiogenic values observed during the second step ( $-20 > \epsilon_{\text{Nd}}(\text{o}) > -45$ ). Finally, such negative  $\epsilon_{\text{Nd}}(\text{o})$  is characteristic of the Labrador Sea, the Bay of Baffin, across the Canadian shield and Greenland [McCulloch and Wasserburg, 1978; Goldstein and Jacobsen, 1988], and thus step 2 would correspond to the paroxysmal step of the Laurentide surge.

The third step is characterized by an input of dominant detrital carbonate, while crust-derived grains and volcanic glass were disappearing. It is known that the northern Canadian provinces (Ellesmere Islands, Banks and Victoria Islands, and Mackenzie basin) are dominated by limestones [Bischof and Darby, 1997]. Bond *et al.* [1992] suggested that these Paleozoic limestones surrounding the northern part of Baffin Bay and Hudson Bay were the main source of detrital carbonate. Consistently, detrital carbonate is abundant in the HLs from the Labrador Sea-Hudson Bay [Andrews and Tedesco, 1992]. The Baffin Bay area has also been proposed as a possible source area [Dowdeswell *et al.*, 1995]. In order to check this hypothesis, the Sr-Nd isotopic composition was measured in detrital carbonate grains picked up within H1. Their composition should reflect the sea water composition that characterized the ocean, at the time of the parent limestone deposition. The following values were obtained:  $\epsilon_{\text{Nd}}(\text{o}) =$

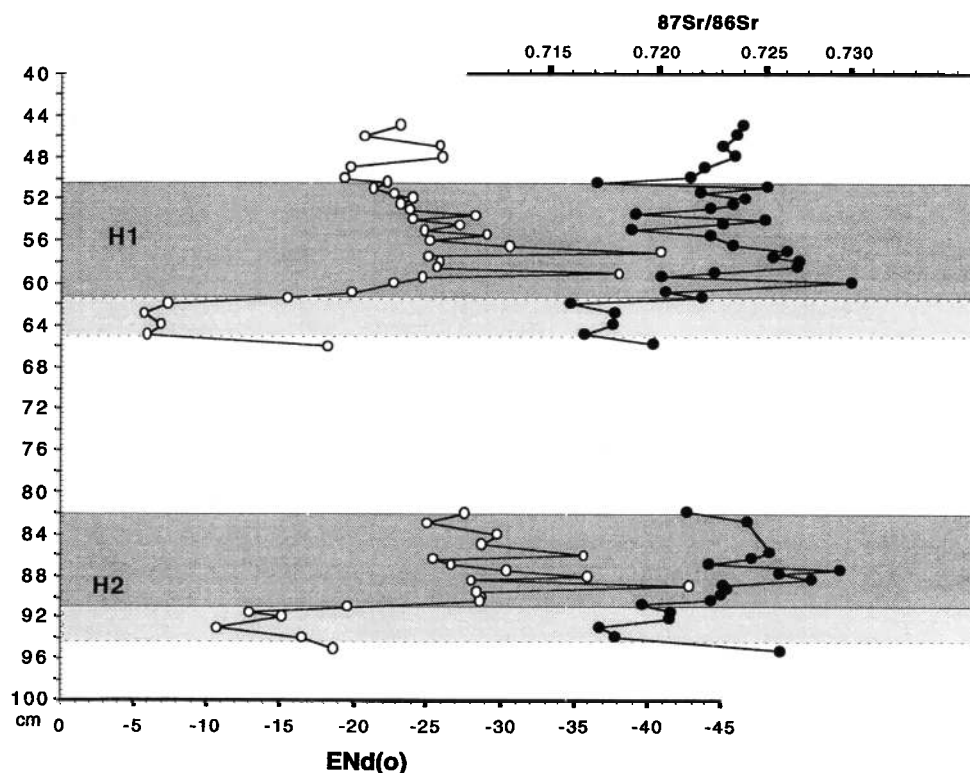


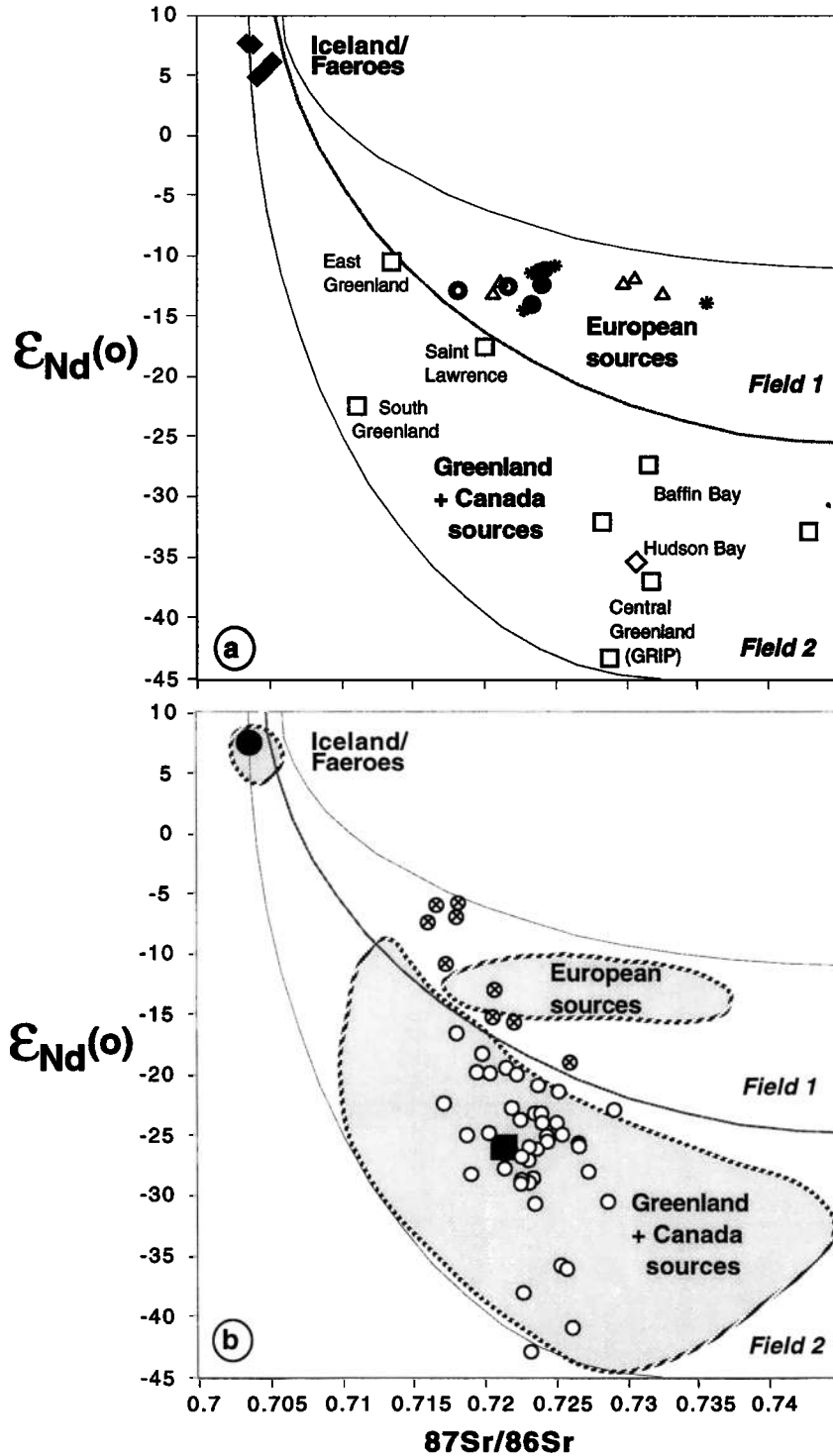
Figure 5. Correlation of  $^{87}\text{Sr}/^{86}\text{Sr}$  with  $^{143}\text{Nd}/^{144}\text{Nd}$  isotopic ratios of the coarse ( $>63\ \mu\text{m}$ ), carbonate-free IRD fraction. The  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios are expressed as  $\epsilon\text{Nd}(o)$ . Error bars are smaller than symbols. Samples are analyzed every centimeter between the HLs and every 5 mm within H1 and H2. Location of Heinrich layers H1 and H2 is shown by stippled areas, dark grey for the Laurentide events and light grey for the European precursor events.

-16.8 and  $^{87}\text{Sr}/^{86}\text{Sr} = 0.708414$ . In the North Atlantic region, such a negative  $\epsilon\text{Nd}(o)$  value is traced to the Labrador Sea-Baffin Bay area, where the modern sea-water imprint ranges from -16 down to -27 [Stordal and Wasserburg, 1986]. These values result from the weathering of the surrounding Archean cratonic blocks as well as the lower Paleozoic carbonate rocks that are abundant in the northern Baffin Bay and the Hudson Bay areas. Concerning the  $^{87}\text{Sr}/^{86}\text{Sr}$  composition of these carbonate grains, the  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio (0.708414) strongly differs from the one that characterizes the recent seawater composition ( $^{87}\text{Sr}/^{86}\text{Sr} = 0.7092$ ). Such a low value is only observed during a short period of the lower Miocene and during the entire Paleozoic period [Elderfield, 1986]. As there is no Miocene formation in the northeastern Canada/Greenland area, this signature can only reflect a Paleozoic origin, which is consistent with the Nd fingerprint and definitely confirms the "Baffin/Labrador" origin of the carbonate grains observed in the HLs. Such a low ratio ( $\approx 0.7084$ ) is not specific to H1. We also found it in H4 from core M6-7A-244/1 [Heinrich, 1988] (Figure 1) ( $^{87}\text{Sr}/^{86}\text{Sr} = 0.70865$ ), reflecting the strong influence of detrital Paleozoic carbonate. On the contrary, the ambient (non-HL) sediment is dominated by the glacial plankton carbonate, displaying an isotopic composition ( $^{87}\text{Sr}/^{86}\text{Sr} = 0.70902$ ) close to the last glacial seawater value ( $^{87}\text{Sr}/^{86}\text{Sr} = 0.7092$ ).

This three-step sequence is observed in both H1 and H2 and may be characteristic of the HL genesis. A similar

sequence is also visible in other parts of the IRD belt, for example, in core VM23-81 (Figure 1), especially during H2 [Bond and Lotti, 1995]. This structure reflects changes in the source of icebergs released in the ocean during a Heinrich event. Icebergs were first released by the European ice sheets (Iceland, British Isles, and Fennoscandia) during a precursor event, followed then by the Laurentide surge.

It seems that during the precursor event, icebergs were released roughly simultaneously by the different European ice sheets. Indeed, before H2, the Icelandic event started just before 22.2 kyr B.P. in core SU90-09 (Figure 3), around 22.0 and 22.4 kyr B.P. in cores VM23-81 and ODP 609, respectively [Bond et al., 1999]. At the same time, the European (British ice sheet IRD) event started along the Celtic margin at 21.85 kyr B.P. in core MD95-2002 [Grousset et al., 2000] and at 22.1 kyr B.P. in core OMEX-2K [Scourse et al., 2000]. Surprisingly, Bischof and Darby [1997, 1999] reported that simultaneously the Innuitian ice sheet released its first icebergs in the Arctic area  $\sim 1.5$  kyr before the Laurentide surge. This early initiation is almost synchronous with the appearance of the atmospheric H2 equivalent observed in the Greenland Ice Sheet Project 2 (GISP2) ice core [Mayewski et al., 1994]. In the same way, before H1, the Icelandic event started around 15 kyr B.P. in cores SU90-09 (Figure 3), VM 23-81 and Ocean Drilling Program (ODP) 609 [Bond et al., 1999]. This synchronicity hypothesis challenges the vision proposed by Dowdeswell et al. [1999], who



**Figure 6.**  $^{87}Sr/^{86}Sr$  ratios versus  $^{143}Nd/^{144}Nd$  ratios, expressed as  $\epsilon_{Nd(o)}$ . These data were obtained on the coarse (>63  $\mu m$ ), carbonate-free fraction on both the core SU90-09 samples and the potential source area (PSA) samples. Error bars are smaller than symbols. (a) The PSA are reported as large solid diamonds (volcanic provinces), open triangles (British Isles sources), stars (Scandinavian sources), closed circles (other European sources), and open squares (North American sources), and their envelopes are contoured by hand, roughly following mixing hyperbolae. The precise location of each PSA sample is provided on Figure 1. Despite the fact that they are bulk samples (as opposed to the other coarse fractions), present-day samples of suspended particulate matter from the Hudson Bay [Goldstein and Jacobsen, 1988] and from the Fram Strait (open circles and open diamond, respectively) are also plotted for comparison. (b) Samples for core SU90-09 (open circles) are represented along with the envelopes of the main source provinces (stippled areas). Volcanic glasses and quartz grains picked from within the European precursor of H1 (level 93 cm) are also plotted (solid circle and solid square, respectively).

**Table 4.** C and N Abundance and Isotopic Composition of the Carbonate-Free Fine (<50  $\mu\text{m}$ ) Fraction

Depth, cm	OC ( $\text{mg}\cdot\text{g}^{-1}$ )	TN ( $\text{mg}\cdot\text{g}^{-1}$ )	OC/TN (C/N)	$\delta^{13}\text{C}$ vs PDB, ‰	$\delta^{15}\text{N}$ vs Air, ‰
40	3.87	0.48	8.06	-21.9	5.41
41	4.03	0.56	7.2	-22.46	5.33
42	3.77	0.55	6.85	-22.67	6.29
43	3.87	0.58	6.67	-22.84	4.27
44	4.0	0.54	7.41	-22.36	5.42
45	3.8	0.52	7.31	-22.24	4.33
46	3.97	0.55	7.22	-22.2	6.33
47	4.05	0.51	7.95	-23.64	5.29
49	3.87	0.58	6.67	-23.44	4.27
50	3.23	0.41	7.88	-24.88	5.18
50.5	3.21	0.38	8.45	-25.3	4.5
51	3.0	0.36	8.33	-25.86	4.36
51.5	3.13	0.46	6.8	-24.98	4.09
52	2.81	0.37	7.59	-25.67	3.93
52.5	2.72	0.32	8.5	-26.25	3.97
53	2.93	0.33	8.88	-27.02	4.35
53.5	2.63	0.28	9.39	-27.27	4.09
54	2.79	0.32	8.72	-26.68	4.63
54.5	2.51	0.28	8.96	-26.78	4.62
55	2.46	0.24	10.25	-27.37	3.27
55.5	2.31	0.24	9.62	-27.43	4.0
56	2.83	0.29	9.76	-27.64	2.72
56.5	2.8	0.24	11.67	-27.76	3.18
57	2.8	0.24	11.67	-27.69	2.15
57.5	2.87	0.27	10.63	-27.49	2.08
58	2.72	0.23	11.83	-27.2	1.58
58.5	2.78	0.22	12.64	-27.99	1.96
59	2.56	0.3	8.53	-27.24	3.21
59.5	2.96	0.34	8.71	-27.3	3.4
60	3.4	0.39	8.72	-26.57	3.63
60.5	2.89	0.34	8.5	-26.55	4.4
61	2.91	0.34	8.56	-26.68	3.5
61.5	3.05	0.42	7.26	-25.81	4.13
62	3.03	0.42	7.21	-24.8	4.5
63	3.22	0.51	6.31	-24.42	5.24
64	3.54	0.56	6.32	-24.56	4.62
65	3.62	0.52	6.96	-24.3	5.4
66	3.61	0.51	7.08	-24.47	6.11
67	3.05	0.51	5.98	-23.41	5.22
68	3.62	0.51	7.1	-24.48	5.47
69	3.53	0.62	5.69	-23.48	4.94
70	4.1	0.61	6.72	-23.72	5.83
71	3.95	0.61	6.48	-23.71	5.48
72	3.94	0.55	7.16	-23.6	5.7
73	3.56	0.61	5.84	-23.18	5.22
74	3.41	0.57	5.98	-22.89	6.51
75	4.56	0.59	7.73	-23.2	6.49
76	4.4	0.62	7.1	-23.14	6.38
77	3.86	0.63	6.13	-23.47	4.94
78	3.42	0.6	5.7	-24.69	4.89
79	3.55	0.57	6.23	-24.9	4.64
80	3.62	0.52	6.96	-24.5	6.03
81	3.85	0.5	7.7	-25.2	5.03
82	3.6	0.46	7.83	-25.04	5.87
83	2.92	0.41	7.12	-25.39	3.45
84	3.36	0.41	8.2	-26.18	3.74
85	3.03	0.29	10.45	-27.36	4.24

Table 4. (continued)

Depth, cm	OC (mg · g <sup>-1</sup> )	TN (mg · g <sup>-1</sup> )	OC/TN (C/N)	δ <sup>13</sup> C vs PDB, ‰	δ <sup>15</sup> N vs Air, ‰
86	3.08	0.29	10.62	-27.11	2.51
86.5	2.8	0.3	9.33	-27.4	4.27
87	2.65	0.27	9.81	-27.41	3.09
87.5	3.53	0.33	10.7	-28.03	3.05
88	3.37	0.31	10.87	-27.67	3.07
88.5	2.76	0.28	9.86	-27.34	2.18
89	3.05	0.27	11.3	-27.65	3.17
89.5	3.4	0.34	10.0	-26.91	3.34
90	4.24	0.34	12.47	-27.0	4.42
90.5	4.19	0.38	11.03	-27.0	2.87
91	3.86	0.47	8.21	-25.58	4.88
91.5	3.93	0.57	6.89	-24.23	5.4
92	4.45	0.73	6.1	-24.16	5.87
93	3.79	0.59	6.42	-23.4	5.2
94	4.1	0.57	7.19	-24.61	5.8
95	4.13	0.6	6.88	-24.73	4.99
96	3.92	0.5	7.84	-25.34	5.31
97	3.1	0.52	5.96	-23.16	5.99
98	3.73	0.69	5.41	-22.47	6.21
99	3.12	0.53	5.89	-22.8	5.34
100	3.3	0.63	5.24	-22.74	5.86
101	3.22	0.59	5.46	-22.9	5.34
102	3.22	0.53	6.08	-23.57	4.99
103	3.49	0.6	5.82	-23.37	6.16
104	3.28	0.62	5.29	-22.98	5.82
105	3.14	0.58	5.41	-23.57	6.05
106	3.1	0.6	5.17	-24.1	5.86
107	3.1	0.6	5.17	-23.0	5.94
108	3.24	0.55	5.89	-23.34	5.18
109	3.23	0.5	6.46	-24.89	5.08
110	3.15	0.49	6.43	-24.65	5.82
111	3.21	0.56	5.73	-23.31	5.17
112	3.66	0.59	6.2	-23.65	5.48
113	3.2	0.5	6.4	-23.3	5.36
114	3.2	0.6	5.33	-22.5	5.86
115	3.29	0.55	5.98	-22.92	5.06

Here vs is versus

considered that deposition of IRD in the Nordic Seas and the North Atlantic Ocean were asynchronous. We do not think, however, that this is contradictory, as their conclusion is based on a low-resolution study of cores, for which a limited number of <sup>14</sup>C ages is provided, thus limiting the control on their timescales. On the contrary, the synchronicity hypothesis is consolidated by consistent observations made in both the Norwegian Sea and the Irminger Basin [M. Elliot *et al.*, manuscript in preparation, 2001].

Along the European margin, the second step (Laurentide-derived IRD) started around 20.4 kyr B.P. in core OMEX-2K [Scourse *et al.*, 2000], at ~20.2 kyr B.P. in core MD95-2002 [Grousset *et al.*, 2000], and 20.5 kyr B.P. in core SO75-26KL [Zahn *et al.*, 1997]. These remote areas were probably reached only by the paroxysmal, late pulse of the Laurentide surge, whereas earlier Laurentide-derived IRD reached the axial part of the IRD belt (core VM23-81) around 21.5 kyr B.P. [Bond *et al.*, 1999]. These Laurentide-derived IRD are made of rock fragments, quartz, and micas, all these minerals being derived from the erosion of the crustal rocks forming the Canadian

shield. They accompany detrital carbonate coming from the Hudson and northern Baffin Bays.

The third step is dominated by detrital carbonate coming from these northwestern regions. These two last steps imply some change may be related to the evolving shape of the Laurentide ice sheet during its surge. The Laurentide ice sheet would have first scraped old crustal formations occurring mostly along the rims of the ice sheet; then, after the major collapse of this ice sheet, icebergs more likely carrying detrital carbonate would come from the central part of the ice sheet (Hudson and northern Baffin Bays).

According to Alley [1998], this sequence of the Heinrich events challenges the idea of a global forcing triggering the HLs. Other processes, such as change in the thermohaline patterns and or destabilization of ice sheet rims were involved, probably implying a multistep story as proposed by Scourse *et al.* [2000]. As already hypothesized, whether or not these precursor events could be associated to the millennial-scale (1-2 kyr) Dansgaard-Oeschger cycles [Grousset *et al.*, 2000] remains an open question, challenged by other interpretations



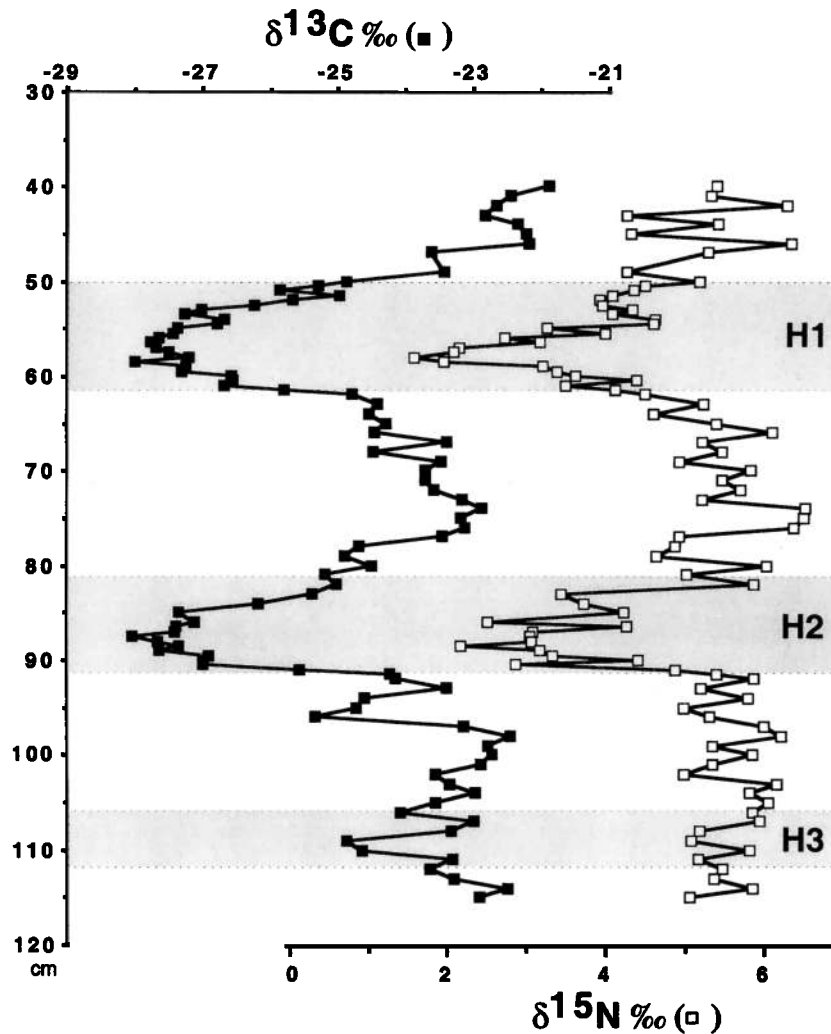


Figure 7. Down core variations in the organic matter composition ( $\delta^{13}\text{C}$  are solid squares, and  $\delta^{15}\text{N}$  are open squares) of the fine ( $<50\ \mu\text{m}$ ) carbonate-free fraction across H1, H2, and H3 in core SU90-09. Location of Heinrich layers H1, H2 and H3 is shown by stippled areas.

[Alley and MacAyeal, 1994; Broecker, 1997; Bond *et al.*, 1999; Scourse *et al.*, 2000]. According to Scourse *et al.* [2000] and Knutz *et al.*, [2001], some other forcing factors must have caused changes in thermocline circulation patterns, which may have triggered the Laurentide collapses. Thus the relationship linking the different ice sheet disequilibria could involve complex feedbacks between sealevel changes, ice sheet dynamics, and variation in the northward heat/moisture transport by the thermohaline circulation.

#### 4.2. High Frequency ( $\approx$ Centennial) Pulses

Cyclic, abrupt oscillations are visible within the HLs. This high-frequency variability is reflected by three independent records: the amount of coarse grains (Figure 4), their Sr-Nd composition (Figure 5) and the organic matter composition (Figure 7). For example, about four to six short oscillations occurred during H1 and about four occurred during H2. Considering the duration of Heinrich events, these oscillations would have occurred on a timescale of 2-5

centuries. As revealed by X-ray imagery, these pulses are separated by horizontal, thin ( $\leq 5\ \text{mm}$  thick) laminations of fine deposits, similar to ambient glacial sediments that were deposited between HLs. Such a fine-scale structure could reflect oscillations in the Laurentide iceberg delivery. This pulsed surge hypothesis would have to be tested by models in order to see if it is consistent with the ice sheet dynamics. Exactly the same structure has been described for H1 in core DAPC-2 (Figure 1) [Knutz *et al.*, 2001]. This observation implies that both the Laurentide and the British ice sheets released their icebergs on a similarly pulsed mode. Knutz *et al.* [2001] suggest that these centennial-scale IRD peaks could correspond to grounding adjustments of the British ice sheet inflicted by a steady rise sea level. This process cannot explain, however, the similar structure observed in core SU90-09 during H2, which implies that more regional changes may have been involved. Another explanation might be the north-south oscillation of the location of the pathway of detritus-laden icebergs, the so-called IRD belt. Ruddiman [1977] originally determined its mean location by averaging the

entire glacial period. In fact, more detailed analysis reveals that for example, the main axis of the IRD belt was centered around 45°N during H2 but was shifted northward up to 47°N during H5 [Grousset *et al.*, 1993]. Such latitudinal oscillations could have generated the observed records. These pulses could be induced by short-term latitudinal shifts of the boundary separating the northern Atlantic cyclonic gyre and the restricted polar clockwise penetration of the North Atlantic drift [Kefffer *et al.*, 1988; Chapman and Maslin, 1999]. Core SU90-09, being located at the southern boundary of the IRD belt, could have been impacted intermittently by the rain of IRD. These hypotheses could be checked by looking at cores located in the center of the IRD belt, where they would have been insensitive to slight latitudinal shifts. Unfortunately, studies already published on such cores were conducted at a lower resolution, and these cores should be revisited at a similar ultrahigh-resolution mode as that used for core SU90-09.

#### 4.3. Continental Versus Marine Sources of Organic Matter Within the HLs

We can make the assumption that marine organic matter originating from cold areas of high latitudes regions represent possible sources of organic matter to deep-sea sediments of the NE Atlantic during Heinrich events. However, their isotopic values generally do not match those measured for H1 and H2 in core SU90-09. For example, present-day  $\delta^{13}\text{C}$  and  $\delta^{15}\text{N}$  averages of  $-21.9 \pm 0.4$  and  $7.6 \pm 0.6\text{‰}$  are observed in the Labrador Sea [Muzuka and Hillaire-Marcel, 1999]. SOM from deep basins of the Bering Sea yields  $\delta^{13}\text{C}$  and  $\delta^{15}\text{N}$  of  $-20.5$  to  $-22.5\text{‰}$  and  $5.5$ - $7.5\text{‰}$ , respectively, for the past 40,000 years B.P. [Nakatsuka *et al.*, 1995]. These isotopic values are much closer to those obtained for ambient sediments ( $\delta^{13}\text{C} = -23.5 \pm 0.7\text{‰}$  and  $\delta^{15}\text{N} = 4.7 \pm 0.9\text{‰}$ ) than for SOM deposited during H1 and H2. An alternative interpretation would be that SOM isotopic composition records an enhanced detrital supply of lithic and/or soil organic matter derived from continental sources of high-latitude regions. Such terrigenous supply may be estimated from studies of glacial sediments from the western Labrador, characterized by minimal isotopic values:  $\delta^{13}\text{C} \approx -27\text{‰}$  and  $\delta^{15}\text{N} \approx 2\text{‰}$  [Muzuka, 1995]. In a more general way, soil and aquatic sources of organic matter located in cold areas should release material derived from "C<sub>3</sub> plant" isotopic signatures [Farquhar *et al.*, 1989]. Glacial erosion of "geological" sources can also provide detrital organic matter with comparable  $\delta^{13}\text{C}$  values, organic matter being represented by various continental and marine remains (i.e., Cambrian-Jurassic records  $-31\text{‰} < \delta^{13}\text{C} < -26\text{‰}$ , [Haynes *et al.*, 1999]). Most of the available  $\delta^{15}\text{N}$  data for arctic and subarctic vegetations point to low values ( $-7.7\text{‰} < \delta^{15}\text{N} < 0.9\text{‰}$ , [Schulze *et al.*, 1994]), which even after various enrichment processes related to soil formation, are consistent with the ones observed for H1 and H2. Other tracers also point to terrigenous sources. Previous analysis from the twin core SU90-08 revealed the presence of this detrital and continental *n*-alkane components in HLs [Villanueva *et al.*, 1998]. Some IRD intervals are characterized by a high content of coal [Heinrich *et al.*, 1989]. Finally, optical examination of the organic matter reveals a large amount of continent-derived

material, including abundant airborne material, such as spores, pollens, and cellulose-rich debris (E. Vergès, personal communication, 2000). Such an airborne particulate enhancement is consistent with the simultaneous increase of dust observed during the Heinrich events in the Greenland ice cores [Mayewski *et al.*, 1994] as already reported [Grousset *et al.*, 2000]. Therefore, considering the important contribution of continental organic matter transported by ice rafting in the IRD belt, the isotopic composition of SOM should not be systematically considered as reflecting a marine signature (Huon *et al.*, submitted manuscript, 2001), but, rather, reflecting a change in the source of SOM.

Using diatom proxies, Sancetta [1992] proposed that in the glacial Northern Pacific and Atlantic Oceans the presence of icebergs supported high production, owing to physical or biochemical mechanisms. This process should be considerably enhanced during HLs. It has not been possible, however, to observe such enhancements during HLs as diatoms totally disappear during these events in core SU90-09 (X. Pichon, personal communication, 2000). This process can be addressed by looking at the  $\delta^{15}\text{N}$  relative increases observed during H1 and H2. A possible interpretation could be that increasing  $\delta^{15}\text{N}$  values are related to more efficient surface water nutrient use [e.g., François *et al.*, 1997]. Surface and bottom water exchanges could then temporarily occur between the discharge maxima observed during H1 and H2. However, this question cannot be solely answered on the basis of stable isotope data as no clear relationship could be drawn between  $\delta^{13}\text{C}$  and  $\delta^{15}\text{N}$  values or between  $\delta^{15}\text{N}$  and paleosalinity estimates recorded by foraminifer  $\delta^{18}\text{O}$ .

## 5. Conclusions

Zooming in on Heinrich layers (HLs) in a high-resolution mode (every 5 mm) permits the extraction of a level of variability that has been poorly appreciated until now. Some conclusions can be drawn from this survey.

H1 and H2 are characterized by a three-step internal structure: during the first step, ice-rafted detritus (IRD) derived from the European ice sheets (Iceland, Fennoscandia, and British Isles) are released into the northern North Atlantic Ocean, displaying a so-called "precursor event" typified by volcanic grains and lasting 1-2 centuries. A second step, which may occur 1.0-1.5 kyr after the beginning of the precursor event, is represented by a major input of Laurentide-derived IRD, dominated by crust-derived particles, typified by a high abundance of quartz. Finally, during a last step, detrital carbonate from the remote Hudson and Baffin Bays are accumulated.

During a Heinrich event, IRD deposition is not a continuous process but reflects a time-transgressive, multisourced release of icebergs, inducing an intermittent sedimentation. These pulses occur on a centennial timescale ( $\approx 2$ -5 centuries) and could be due to a step-by-step collapse of the Laurentide ice sheet or to latitudinal oscillations of the trans-Atlantic IRD belt. Similar short centennial-scale pulses are observed along the European margin as well [Knutz *et al.*, 2001].

The isotopic signature of fine-grained sedimentary organic matter included in H1 and H2 reveals a dominant continental

origin with low values decreasing to  $-27.9$  and  $0.3$  ‰ for  $\delta^{13}\text{C}$  and  $\delta^{15}\text{N}$ , respectively. These values are sharply contrasting with those of ambient sediments ( $\delta^{13}\text{C} = -23.5 \pm 0.7$  ‰ and  $\delta^{15}\text{N} = 4.7 \pm 0.9$  ‰), mainly reflecting the glacial oceanic contribution.

In terms of HL origin, although the Laurentide surge is the major, dominant process characterizing a Heinrich event, the starting point of the entire sequence is a destabilization of the Icelandic and European ice sheets.

**Acknowledgements.** Most of the cores studied here were collected by IFREMER research vessels. IMAGES cores were recovered by the N/V *Marion-Dufresne* (French Polar Institute). We gratefully acknowledge J. Andrews, G. Auffret,

D. Boust, P. Knutz, and P. Martinez for their collegiality in sharing samples with us. P. Pédemay, G. Lavaux, B. Le Coat, J. Tessier, B. Hermann, and M. Roy-Barman provided invaluable technical assistance. Catherine Kissel kindly provided the high-resolution magnetic susceptibility record. We thank J. Andrews, L. Labeyrie, and F. Eynaud for comments and discussions. We thank P.E. Biscaye, M. Maslin, P.C. Knutz, and one anonymous reviewer for friendly and helpful improvements on the manuscript. AMS  $^{14}\text{C}$  ages were measured by Beta Analytic Inc. Financial contributions from the European Community programme ENV4-CT97-0643 and the Centre National de la Recherche Scientifique programmes VARIANTE and ECLIPSE are acknowledged. Thomas Richter was supported through a postdoctoral fellowship by the Deutsche Forschungsgemeinschaft (grant Ri 963/1-1) at DGO-Bodeaux. This is UMR 5805 EPOC contribution 1385.

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(Received June 30, 2000, revised January 10, 2001, accepted January 26, 2001)